

1 **Radiation fog properties in two consecutive events under polluted and clean**
2 **conditions in the Yangtze River Delta, China: A simulation study**

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15 **Abstract.** Aerosol–cloud interaction (ACI) in fog and planetary boundary layer (PBL)
16 conditions play critical roles in the fog life cycle. However, it is not clear how ACI in the first
17 fog (Fog1) affects **the PBL**, and subsequently affects ACI in the second fog (Fog2), which is
18 important to understand the interaction between ACI and **the PBL** as well as their effects on fog
19 properties. To fill this knowledge gap, we simulate two successive radiation fog events in the
20 Yangtze River Delta, China, using the Weather Research and Forecasting model coupled with
21 Chemistry (WRF-Chem). Our simulations indicate that **the PBL conditions conducive to Fog2**
22 **formation** are affected by ACI with high aerosol loading in Fog1; subsequently, **the PBL**
23 promotes ACI in Fog2, resulting in a higher liquid water content, higher droplet number
24 concentration, smaller droplet size, larger fog optical depth, wider fog distribution, and longer
25 fog lifetime in Fog2 than in Fog1. This phenomenon is related to the following physical factors.
26 The first factor involves **meteorological conditions conducive to Fog2 formation**, including low
27 temperature, high humidity, and high stability. The second factor is the feedbacks between

28 microphysics and radiative cooling. A higher fog droplet number concentration increases the
29 liquid water path and fog optical depth, thereby enhancing long-wave radiative cooling and
30 condensation near the fog top. The third factor is the feedbacks between macrophysics,
31 radiation, and turbulence. A higher fog top presents stronger long-wave radiative cooling near
32 the fog top than near the fog base, which weakens temperature inversion and strengthens
33 turbulence, ultimately increasing the fog-top height and fog area.

34 In summary, under polluted conditions, ACI postpones the dissipation of Fog1 owing to
35 these two feedbacks and generates PBL meteorological conditions that are more conducive to
36 the formation of Fog2 than those prior to Fog1. These conditions promote the earlier formation
37 of Fog2, further enhancing the two feedbacks and strengthening the ACI in Fog2. Our findings
38 are critical for studying the interaction between aerosols, fog, and the PBL; moreover, they
39 shed new light on ACI.

40 **1 Introduction**

41 Fog comprises water droplets or ice crystals suspended above the ground (WMO, 1992). This
42 results in low visibility, which affects the human health, transportation, and power systems (Niu
43 et al., 2010). Uncertainties exist in fog forecasting (Zhou and Du, 2010; Zhou et al., 2011). An
44 important reason is that the physical processes of fog remain unclear because many processes
45 (aerosol activation, condensation, radiation, and turbulence) occur simultaneously and interact
46 with each other nonlinearly (Haefelin et al., 2010), which affects fog properties (Mazoyer et
47 al., 2022) and impedes related parameterisation (Poku et al., 2021). To better understand the
48 physical processes of fog, Comprehensive studies based on observations and simulations have
49 been conducted to better understand the physical processes of fog (Fernando et al., 2021;
50 Gultepe et al., 2014; Guo et al., 2015; Hammer et al., 2014; Liu et al., 2011; Price et al., 2018;
51 Shen et al., 2018; Wang et al., 2021). The critical roles of aerosols and planetary boundary layer

52 (PBL) in these processes have been shown (Boutle et al., 2018; Niu et al., 2011; Quan et al.,
53 2021).

54 Since fog is a special type of cloud (Guo et al., 2021; Kim and Yum, 2010, 2013; Wang et
55 al., 2023), aerosol–fog interaction is expected to share similarities with aerosol–cloud
56 interaction (ACI). The ACI in fog reflects the response of fog properties to changes in aerosol
57 loading. Studies on ACI revealed that increasing aerosol loading increased cloud droplet
58 concentration, thereby increasing the cloud optical depth under a constant liquid water content
59 (LWC) (Garrett and Zhao, 2006; Twomey, 1977; Wang et al., 2013; Wang et al., 2018; Zhao
60 and Garrett, 2015). Various continental fog observation projects showed that fog microphysical
61 properties were significantly affected by aerosol loading (Mazoyer et al., 2019; Niu et al., 2011;
62 Quan et al., 2011; Wang et al., 2021). For instance, in polluted fog observations, Quan et al.
63 (2011) found that the fog droplet number concentration (N_d) was higher than $1,000\text{ cm}^{-3}$ and
64 effective radius (R_e) was approximately $7\text{ }\mu\text{m}$ in the North China Plain. In clean fog
65 observations, Wang et al. (2021) showed that N_d was smaller than 100 cm^{-3} and R_e was
66 approximately $9\text{ }\mu\text{m}$ in the tropical rainforest in Xishuangbanna, China. Several simulation
67 studies reproduced these observations, and demonstrated the complex impact of ACI on fog
68 micro- and macrophysics (Jia et al., 2019; Maalick et al., 2016; Stolaki et al., 2015; Yan et al.,
69 2020). Regarding fog microphysics, increasing aerosol loading in the simulations increased N_d
70 and LWC due to increased activation and condensation (Jia et al., 2019; Stolaki et al., 2015;
71 Yan et al., 2020). Regarding fog macrophysics, several modelling studies reported that
72 increasing aerosol loading increased the fog-top height (Jia et al., 2019; Stolaki et al., 2015)
73 and prolonged the fog lifetime by delaying its dissipation (Quan et al., 2021; Yan et al., 2021).

74 Furthermore, previous studies found that meteorological conditions played crucial roles in
75 aerosol–cloud interaction as well as cloud macro- and microphysics (Ackerman et al., 2004;
76 Kumar et al., 2017; Kumar et al., 2021; Liu et al., 2019; Liu et al., 2020; Toll et al., 2019).

77 Similarly, studies on fog showed that ACI was affected by meteorological conditions in the
78 PBL (e.g., radiation, thermodynamics, and dynamics), which further affected fog micro- and
79 macrophysics (Haeffelin et al., 2010). Previous studies showed that radiative cooling was an
80 important factor in temperature inversion that provided stable conditions for fog formation
81 (Fitzjarrald and Lala, 1989; Holets and Swanson, 1981; Roach et al., 1976). According to Zhou
82 and Ferrier (2008), turbulence may suppress or deepen the fog-top height, which was related to
83 the critical turbulence coefficient. The critical turbulence coefficient was the turbulence
84 threshold for diagnosing whether turbulence suppressed fog or not. When the turbulence
85 intensity within the fog did not exceed the critical turbulence coefficient, the fog persisted;
86 however, when it surpassed its threshold, the fog dissipated (Zhou and Ferrier, 2008). When
87 temperature inversion was weak, excessive vertical turbulent mixing delayed fog formation
88 (Maronga and Bosveld, 2017). However, when temperature inversion was sufficiently strong,
89 vertical turbulent mixing at the middle and fog base increased the fog-top height, as suggested
90 by observations (Ye et al., 2015) and simulations (Porson et al., 2011). Consequently,
91 turbulence may impact fog macrophysics. Moreover, aerosols were reported to affect
92 turbulence, thereby impacting fog macrophysics (Jia et al., 2019; Quan et al., 2021). A
93 qualitative analysis, conducted in a previous study, revealed that aerosols promoted turbulence
94 and horizontal distribution because of weaker temperature inversion (Jia et al., 2019).

95 Previous studies typically focused on either a single fog event or analysed multiple fog
96 events statistically; however, several studies noted that LWC, N_d , and liquid water path (LWP)
97 in a latter fog event exhibited larger values compared to those for the preceding fog event (Quan
98 et al., 2011; Wærsted et al., 2017). What are the physical mechanisms behind the property
99 changes during two successive fog events? Furthermore, which fog event has macro- and
100 microphysical properties that are more sensitive to aerosol loading, i.e., experiencing a stronger
101 ACI? What are the mechanisms underlying the interactions between ACI and the PBL? To

102 answer these questions, two successive radiation fog events in the Yangtze River Delta (YRD)
103 region of China are simulated in this article using the Weather Research and Forecasting model
104 coupled with Chemistry (WRF-Chem). The two fog events provide an excellent opportunity to
105 investigate ACI under polluted conditions as a chain. This involves analysing how high aerosol
106 loading affects properties in the first fog event, how the properties in the first polluted fog event
107 affect radiation and PBL structure, and finally, how radiation and the PBL affect properties and
108 ACI in the second fog event under polluted conditions. Additionally, since fog is a special type
109 of cloud near the ground, studying the evolution of ACI in fog aids in examining the progression
110 of ACI in cloud, which is critical for climate prediction (Boutle et al., 2018; Vautard et al.,
111 2009) .

112 The rest of the article is organized as follows. Section 2 presents descriptions of the two
113 successive fog events, experimental design, and data source. Section 3 presents simulation
114 evaluation. Section 4 shows larger aerosol-induced changes in Fog2 than in Fog1. Section 5
115 presents the physical mechanisms underlying the larger aerosol-induced changes in Fog2 than
116 in Fog1. Finally, Section 6 summarises the conclusions of this study.

117 **2 Experimental design and data source**

118 Here, we study how radiation fog properties are affected by high aerosol loading and PBL
119 meteorological conditions in two successive events in the YRD region. Before fog events in the
120 YRD, the PM_{2.5} mass concentration was over 100 $\mu\text{g m}^{-3}$ due to anthropogenic emissions (Zhu
121 et al., 2019). On 26 and 27 November 2018, two successive radiation fog events occurred in
122 northern YRD. The first fog event is called Fog1, and the second is called Fog2. Ground-based
123 observations at the Nanjing site (32.2 °N 118.7 °E) show that two fog events (visibility < 1,000
124 m) are accompanied by high relative humidity, low temperature, and weak wind speed (Fig. 1).
125 As shown in Fig. S1, the surface is controlled by a high–pressure system with cold and moist

126 air in northern YRD at 20:00 local standard time (LST) (LST = Universal Time Coordinated +
127 8 h) on 26 and 27 November 2018. WRF-Chem (version 4.1.3) is used to simulate the two
128 successive radiation fog events. WRF-Chem couples physical and chemical processes;
129 therefore, it has been widely used to study ACI (Jia et al., 2019; Lee et al., 2016; Yan et al.,
130 2020; Yan et al., 2021). The model is integrated from 14:00 LST on 24 November 2018 to
131 14:00 LST on 27 November 2018, with the first 24 hours regarded as the spin-up time. As
132 shown in Fig. S2, the model is configured using three nested domains, and all domain centres
133 are located in Nanjing. The three nested domains are 90×122 , 118×142 , and 130×154 grid
134 cells with resolutions of 27, 9, and 3 km, respectively. The simulation area covers the major
135 weather system affecting the YRD. The model includes 36 vertical levels, of which 17 layers
136 are located at the lowest 500 m above the ground level. Moreover, Yang et al. (2019) noted a
137 better fog simulation performance when the bottom layer was 8 m above the ground since this
138 layer affected the fog and surface flux interaction. Consequently, in this study, we set the bottom
139 layer of the model to 8 m. The model is driven by the National Centre for Environmental
140 Prediction (NCEP) Final (FNL) $1^\circ \times 1^\circ$ reanalysis data (<https://rda.ucar.edu/datasets/ds083.2/>)
141 (Ding et al., 2019; Jia et al., 2019). The Multiresolution Emission Inventory for China (MEIC)
142 database (<http://meicmodel.org>) is used for anthropogenic emissions in the model (Li et al.,
143 2017a; Zheng et al., 2018).

144 Table 1 lists the parameterisation schemes of physical processes used in this study. The
145 microphysics scheme is Morrison (Morrison et al., 2005) coupled with the activation scheme
146 (Abdul-Razzak, 2002). The PBL scheme is MYNN2.5 (Nakanishi and Niino, 2009).
147 Turbulence is parameterised in the MYNN2.5 scheme and there is also a sub-grid cloud
148 parameterisation (Chaboureau and Bechtold, 2002) in the MYNN2.5 scheme. The radiation
149 schemes are coupled with the aerosol–cloud–radiation interactions. The long- and short-wave
150 radiation schemes are RRTMG (Iacono et al., 2008) and Goddard (Matsui et al., 2020),

151 respectively. The cumulus scheme is Grell 3D (Grell and Dévényi, 2002). The chemistry
152 schemes are MOSAIC-4 bins (Zaveri et al., 2008) and CBMZ (Zaveri and Peters, 1999).

153 For model **evaluation**, meteorological data are retrieved from the China Meteorological
154 Administration (<http://www.nmic.cn/>). **The cloud product (level 2 full-disk cloud property data)**
155 **from the Himawari-8 geostationary satellite is used (Bessho et al., 2016; Iwabuchi et al., 2018)**
156 **(<https://www.eorc.jaxa.jp/tree/index.html>).** The quality of the Himawari cloud product is
157 **reliable because this product has been evaluated against the Moderate Resolution Imaging**
158 **Spectroradiometer (MODIS) (Bessho et al., 2016; Letu et al., 2020) and cloud profiles from**
159 **aircraft measurements (Zhao et al., 2020). Spatial resolution of the Himawari cloud product is**
160 **$0.05^{\circ} \times 0.05^{\circ}$ (Yang et al., 2020).** PM_{2.5} mass concentration data are obtained from the Ministry
161 of Environmental Protection (<https://quotsoft.net/air/>).

162 To investigate the aerosol-induced changes in fog macro- and microphysics, one control
163 run and two sensitivity tests are conducted: EXP1, EXP2, and EXP3, respectively. High and
164 low emissions indicate polluted and clean conditions, respectively. The differences indicate the
165 aerosol effect on fog properties. **In EXP1**, the emission intensity is obtained directly from the
166 MEIC database to simulate fog under polluted conditions. **In EXP2**, the emission intensity is
167 multiplied by 0.05 to simulate fog under clean conditions, as described by Jia et al. (2019) and
168 Yan et al. (2021). **In EXP3**, Fog1 **occurs under clean conditions** (5% of emission from the MEIC
169 database) and Fog2 **occurs under polluted conditions** (the default emission from the MEIC
170 database). **According to Fog1 dissipation time, clean conditions change to polluted conditions**
171 **at 12:00 LST on 26 November 2018.** Compared with the difference between EXP1 and EXP2,
172 the difference between EXP3 and EXP2 reveals whether the fog properties and ACI with higher
173 aerosol loading in Fog1 affects those in Fog2.

174 3 Simulation evaluation

175 Simulation evaluations for temperature, relative humidity, and wind speed are shown in Fig. 2.
176 The correlation coefficients of 2 m temperature (T_{2m}), 2 m relative humidity (RH_{2m}), and 10 m
177 wind speed (WS_{10m}) between the simulations and observations are 0.9, 0.9, and 0.6, respectively,
178 passing the significance test at 99%. Therefore, the simulations are generally consistent with
179 the observations. The mean bias (MB) of T_{2m} , RH_{2m} , and WS_{10m} between the simulations and
180 observations are 1.0 °C, 2.7%, and 0.4 m s⁻¹, respectively, consistent with evaluation results in
181 studies by Hu et al. (2021), Gao et al. (2016), and Yang et al. (2022). Figure 3 shows the
182 evaluation of PM_{2.5} distribution, and Table 2 summarises statistics of the mean mass
183 concentration of PM_{2.5} based on the method proposed by Boylan and Russell (2006). The
184 normalised mean bias (NMB), normalised mean error (NME), mean fractional bias (MFB), and
185 mean fractional error (MFE) between the simulations and observations are 25%, 30%, 24%,
186 and 28%, respectively (Eqs. S3–S6 in the supplement). Although the PM_{2.5} mass concentration
187 is overestimated, it remains within a reasonable range (Shu et al., 2021; Yang et al., 2022; Zhai
188 et al., 2018).

189 Figure 4 shows the evaluation of fog spatial distribution. The simulated fog optical depth
190 distribution is compared with the Himawari-8 cloud optical depth products at 08:00 LST on 26
191 and 27 November 2018, respectively. To identify observed fog at ground-based stations (the
192 black circles in Fig. 4), we apply two criteria: visibility less than 1 km and relative humidity
193 greater than 90% (Yan et al., 2020). Qualitatively, the value of fog optical depth and the fog
194 spatial distribution in the simulation are roughly similar to those observed by the Himawari
195 satellite and at ground-based stations. Likewise, Lee et al. (2016) evaluated fog distribution
196 simulations against satellite-derived cloud optical depth from satellite and concluded that the
197 distributions of simulations and observations were generally comparable to each other.

198 Further, to quantitatively evaluate the simulation, the Heidke skill score (HSS) is
199 calculated as follows (Barnston, 1992):

$$200 \quad HSS = \frac{2(ad - bc)}{(a + c)(c + d) + (a + b)(b + d)} \quad (1)$$

201 Elements $a-d$ are determined by the occurrence of fog at observation stations located in domain
202 03 and the closest model grids to those observations, as shown in Table 3. If fog events are both
203 observed at stations and simulated at the closest model grids, we recognize those as "hits" and
204 a in Eq.1 represents the total number of "hits" during the entire fog event. Similarly, d represents
205 the number of "correct negatives" for the correct non-event simulations. On the other hand, if
206 fog events are simulated but not observed, we recognize those as "false alarms" and b represents
207 the total number of "false alarms" during the entire fog event. Conversely, c represents the total
208 number of "misses", which indicates that fog events are observed but not simulated. The criteria
209 of observed fog are shown in the last paragraph. Simulated foggy grids are classified based on
210 three criteria (Jia et al., 2019; Zhao et al., 2013): fog water mixing ratio over 0.01 g kg^{-1} , N_d
211 greater than 1 cm^{-3} , and fog base touching the ground. The perfect HSS score is 1.0, indicating
212 that simulations are identical to observations. Here, the HSS score are 0.34 and 0.36 in Fog1
213 and Fog2, respectively, which are close to previous reports (Mecikalski et al., 2008; Xu et al.,
214 2020; Yamane et al., 2010). Therefore, the model generally captures the fog spatial distribution.

215 **4 Larger aerosol-induced changes in Fog2 than in Fog1**

216 Here, we analyse the fog macro- and microphysical characteristics under clean and polluted
217 conditions (Fig. 5). To ensure sufficient sample size for statistical analysis, only data with the
218 fog area fraction larger than 5% are analysed. The fog area fraction is calculated as the number
219 of foggy grid cells divided by the total number of grids in domain 03. We also test other

220 thresholds, 1%, 2.5%, 7.5%, and 10% (Fig. S3). The results are similar to those based on the
221 threshold of 5%.

222 The ratios of changes between the polluted and clean conditions reveal that high aerosol
223 loading affects fog macro- and microphysical properties in Fog1 and Fog2 (Fig. 5a). Compared
224 to fog microphysics under clean conditions, N_d and LWC in Fog1 increase by respectively 463.0%
225 and 81.7%; however, R_e decreased by 32.1% under polluted conditions. Furthermore, because
226 of the ACI, N_d and LWC in Fog2 increase by respectively 672.4% and 113.5%; however R_e
227 decreases by 40.0%. Therefore, aerosol-induced changes in Fog2 are larger than those in Fog1,
228 as shown in Fig. 5a (N_d : 209.5%, LWC: 31.8%, and R_e : -6.9%). Similarly, aerosol-induced
229 changes in fog macrophysics are larger in Fog2. Compared with values under clean conditions,
230 the fog area, fog-top height, and duration in Fog1 increase by respectively 23.1%, 109.6%, and
231 20.0% under polluted conditions; the corresponding values in Fog2 are larger (34.9%, 350.5%,
232 and 25.0%, respectively). In addition, LWP and fog optical depth (τ_c) exhibit similar trends.
233 With a similar trend between observation and simulation, Figure S4 shows that aerosol mass
234 concentration is similar before Fog1 and Fog2 formation, and aerosol number concentration
235 before Fog2 is less than that before Fog1 formation. Therefore, changes in aerosol concentration
236 are not the main reason for the increase in aerosol-induced changes in the two fog properties.
237 To determine whether ACI under polluted conditions leads to an increase in aerosol-induced
238 changes in Fog1 and Fog2, we design a sensitivity test called EXP3, as mentioned above.
239 Furthermore, to quantitatively evaluate the strength of ACI in the two fog events, we examine
240 the responses of τ_c to changes in N_d (Eq. 2) (Ghan et al., 2016):

$$241 \quad \frac{\Delta \ln \tau_c}{\Delta \ln N_d} = \frac{\Delta \ln LWP}{\Delta \ln N_d} - \frac{\Delta \ln R_e}{\Delta \ln N_d} \quad (2)$$

242 Based on the similar aerosol concentration background (Fig. S4), the responses of τ_c to
243 changes in N_d quantitatively confirm which fog has **stronger ACI**. As shown in Table 4, the

244 strength of ACI in Fog2 (1.32) is larger than that in Fog1 (0.98). If Fog1 occurs under clean
245 conditions and Fog2 occurs under polluted conditions (EXP3), ACI in Fog2 is 1.17, which is
246 lower than that in EXP1 (1.32). This implies that high aerosol loading in Fog1 enhances ACI
247 in Fog2. Relative changes in the above properties between Fog1 and Fog2 are calculated as
248 $(\text{Fog2} - \text{Fog1})/\text{Fog1}$. The values of $\Delta \ln \tau_c / \Delta \ln N_d$, $\Delta \ln LWP / \Delta \ln N_d$, as well as $-\Delta \ln R_e / \Delta \ln N_d$ are
249 34.7%, 42.1%, and 9.1% larger in Fog2 than in Fog1, respectively. These numbers
250 quantitatively confirm stronger ACI in Fog2 and indicate that LWP is the dominant factor for
251 enhancing ACI. LWP depends on the fog-top height and LWC. As shown in Fig. 5a, when
252 aerosol loading changes from clean to pollution, the rate of increase in fog-top height in Fog2
253 (350.5%) is much larger than that in Fog1 (109.6%). Although the increase in LWC in Fog2
254 (113.5%) is also larger than that in Fog1 (81.7%), the magnitude of increase in LWC is smaller
255 than that increase in fog-top height, indicating that ACI are more sensitive to fog-top height
256 than to LWC.

257 Fog duration is determined by the time of fog formation and dissipation, which is primarily
258 extended because high aerosol loading delays fog dissipation, as reported previously (Jia et al.,
259 2019; Quan et al., 2021). In this study, high aerosol loading not only delays fog dissipation but
260 also promotes earlier fog formation, particularly during Fog2 (Fig. 5b). Fog formation is related
261 to the PBL conditions which are affected by ACI. To investigate the aerosol effect on the Fog2
262 formation stage, fog spatial distribution at the formation stage from 19:00 LST to 21:00 LST
263 on 26 November is examined, as shown in Figure 6. The fog area is rather small at 19:00 LST
264 under both polluted and clean conditions. At 20:00 LST, fog formation is similar under both
265 polluted and clean conditions in grid cells located outside the black box. However, inside the
266 black box, there are several foggy grid cells under polluted conditions. At 21:00 LST, fog area
267 in the black box further expands under polluted conditions. However, there is almost no fog in
268 the black box at 20:00 and 21:00 LST under clean conditions. Therefore, high aerosol loading

269 promotes earlier formation of Fog2, which is primarily caused by meteorological conditions in
270 the PBL inside the black box. In addition, the fog area outside the black box is larger under
271 polluted conditions than under clean conditions, which is mainly related to the stronger
272 turbulence diffusion under polluted conditions. A detailed analysis is presented in Sect. 5.

273 **5 Physical mechanisms underlying the larger aerosol-induced changes in Fog2 than in** 274 **Fog1**

275 **5.1. More conducive meteorological conditions to Fog2 formation**

276 Meteorological conditions in the PBL affect fog formation time and ACI during fog events. As
277 shown in Table 5, under clean conditions, RH_{2m} before Fog2 formation is higher and PBL
278 height (PBLH) is lower than those before Fog1 formation in domain 03. Polluted conditions
279 yield similar results. Furthermore, compared with the difference in aerosol-induced changes in
280 RH_{2m} and PBLH before fog formation, RH_{2m} increases by 6% and PBLH decreases by 92 m
281 under polluted conditions, which is larger than those (RH_{2m} : 4% and PBLH: -59 m) under clean
282 conditions. Therefore, high aerosol loading generates meteorological conditions **more**
283 **conducive to Fog2 formation** during the two successive fog events.

284 To further analyse how high aerosol loading promotes Fog2 formation, we focus on the
285 black box in Fig. 6, as described in Sect. 4 and by Yan et al. (2021). The regional average
286 differences in the total optical depth (τ_t), downwelling short-wave radiation (SW) at the ground,
287 T_{2m} , PBLH, RH_{2m} , and water vapour mixing ratio ($Q_{v_{bot}}$) at the model bottom layer (8 m) in
288 the black box between polluted and clean conditions are calculated (Fig. 7). **Compared with**
289 **clean conditions, the larger τ_t (mainly due to larger τ_c) and delayed fog dissipation in polluted**
290 **conditions reduce short-wave radiation reaching the ground (from -46 W m^{-2} to -121 W m^{-2})**
291 **during the Fog1 dissipation time. This leads to a decrease in T_{2m} (from -0.2 °C to -1 °C) and**

292 PBLH (from -42 m to -118 m), which further prolongs fog duration (Fig. 7). Notably, Q_{vbot}
293 under polluted conditions is lower than that under clean conditions before the complete
294 dissipation of Fog1, because of reduced fog water evaporation. When the fog dissipates
295 completely, the lower PBLH accumulates more water vapour, increasing Q_{vbot} and RH_{2m} . The
296 positive feedbacks between ACI and PBL are similar to the feedbacks between high aerosol
297 loading and the PBL reviewed by Li et al. (2017b). Furthermore, the feedback mechanism
298 between high aerosol loading and PBL introduced by Zhong et al. (2018) supports the daytime
299 feedbacks between fog and the PBL in the present study. Additionally, aerosol extinction is also
300 considered in τ_t . Whether aerosol optical depth (AOD) affects PBL significantly should also be
301 discussed. As shown in Table 5, RH_{2m} and PBLH before Fog1 on 25 November under clean
302 conditions are 76% and 669 m, respectively, similar to those under polluted conditions (76%
303 and 670 m, respectively). Therefore, it is unlikely that aerosol-meteorology interaction leads to
304 the meteorological differences in Fig. 7. In addition, a previous study (Yan et al., 2021) also
305 noted that aerosol–fog interaction was more remarkable than aerosol–radiation interaction.
306 Therefore, lower temperature, higher relative humidity, and higher stability result from ACI in
307 Fog1 under polluted conditions, contributing to the earlier formation of Fog2.

308 Larger τ_c and delayed dissipation result in lower temperature, higher relative humidity, and
309 higher stability by affecting solar radiation during the daytime. How are these conducive
310 conditions maintained after the sunset around 17:00 LST? Figure 8a shows that cold advection
311 is the major reason for the difference in temperature between polluted and clean conditions. We
312 further seek to unveil the reason cold advection is stronger under polluted conditions. Figure
313 8b shows a cold centre, with wind diverging outwards. The cold centre is related to lower
314 temperature under polluted conditions due to larger τ_c and longer duration in Fog1. Likewise,
315 Steeneveld and De Bode (2018) noted that fog appeared earlier with cold advection. In addition,

316 lower PBLH induced by high aerosol loading promote the maintenance of higher humidity and
317 stability.

318 Overall, due to ACI at the Fog1 dissipation stage, the meteorological conditions are more
319 conducive for promoting Fog2 formation. Furthermore, this interaction enhances the feedbacks
320 in the fog physical processes, thus leading to a stronger ACI in Fog2. Details are discussed in
321 Sect. 5.2 and 5.3.

322

323 5.2. Feedbacks between microphysics and long-wave cooling

324 Section 5.1 reveals the mechanism through which ACI in Fog1 leads to meteorological
325 conditions more conducive to Fog2 formation. In Sect. 5.2, we demonstrate how conducive
326 meteorological conditions play a fundamental role in promoting feedbacks between
327 microphysics and long-wave cooling, resulting in a stronger ACI in Fog2.

328 As shown in Fig. 5a, aerosol-induced changes in N_d and LWC during Fog2 are larger than
329 those during Fog1 because lower temperature and higher humidity are more conducive for
330 aerosol activation and fog condensation (Petters and Kreidenweis, 2007; Simmel and Wurzler,
331 2006). Owing to competition for available water vapour (Mazoyer et al., 2022; Yum and
332 Hudson, 2005), R_e in Fog2 is smaller than that in Fog1. As shown in Fig. 9a, LWP is larger
333 under polluted conditions than that under clean conditions, particularly for Fog2. The average
334 LWP in Fog1 and Fog2 under polluted conditions are 11.6 g m^{-2} and 24.3 g m^{-2} , respectively.
335 When LWP is less than 20 g m^{-2} , vertically integrated long-wave cooling and short-wave
336 heating are stronger under polluted conditions than those under clean conditions (Fig. 9b). This
337 is similar to the results from Petters et al. (2012) and Prabhakaran et al. (2023). Because N_d
338 shows a similar trend with LWP (Fig. S5), the dependence of heating and cooling rates on
339 droplet concentration is consistent with the results based on LWP. Additionally, increased τ_c

340 in Fog2 triggers stronger positive feedbacks between microphysics and long-wave cooling,
341 further enhancing cooling, activation, and condensation and thereby increasing N_d and LWC.
342 Jia et al. (2019) emphasised that high aerosol loading promoted these positive feedbacks. This
343 study further highlights the synergistic effects of high aerosol loading and meteorological
344 conditions on the enhancement of positive feedbacks, which promotes ACI in Fog2.

345 To better understand how the above positive feedbacks affect ACI, Fig. 10 presents the
346 average extinction coefficient through the fog, that is, τ_c at per unit height ($\tau_c/\Delta h$), radiative
347 cooling rate (T_{LW}), condensational growth rate (LWC_{COND}), and LWC tendency due to vertical
348 mixing (LWC_{mixing}) in the two successive fog events. Radiative cooling is the strongest near the
349 fog top and weakest near the fog base (Ducongé et al., 2020; Mazoyer et al., 2017; Wærsted et
350 al., 2017). Consequently, LWC_{COND} and LWC_{mixing} both follow similar profiles in response to
351 radiative cooling. Therefore, if the vertical profiles of the three terms use absolute height, they
352 will be distorted. To overcome this problem, physical quantities are normalised by the fog-top
353 height. Compared with those in Fog1, larger extinction coefficient (Fig. 10a-b), stronger long-
354 wave radiative cooling (Fig. 10c-d), and more condensation (Fig. 10e-f) near the fog top are
355 noted in Fog2 because of the conducive conditions to Fog2 formation, which further increases
356 LWC, fog-top height in Fog2 (black and purple lines) as well as LWP. Enhancement of these
357 parameters indicate that the feedbacks between microphysics and long-wave cooling are
358 stronger in Fog2 than in Fog1. As a result, ACI is stronger in Fog2 than in Fog1, due to
359 favourable PBL conditions caused by ACI in Fog 1. In addition, as shown in Fig. 10g-h, vertical
360 mixing transports fog water from the fog top to the fog base, and the strength of this
361 transportation is stronger in Fog2 than in Fog1, because of stronger turbulent kinetic energy
362 (TKE) in Fog2. The effect of TKE on fog is analysed in Sect. 5.3.

363 **5.3. Feedbacks between macrophysics, radiation, and turbulence**

364 Section 5.2 analyses the microphysics-related mechanisms underlying a stronger ACI in Fog2.
365 This subsection not only focuses on macrophysics and its feedbacks with radiation and
366 turbulence but also discusses how the combined effects of high aerosol loading and
367 meteorological conditions impact the feedbacks and enhance ACI in Fog2, compared with those
368 in Fog1. Briefly, fog macrophysics involves duration and distribution. The reason why the
369 duration of Fog2 is longer than that of Fog1 is due to the earlier formation of Fog2, which is
370 induced by meteorological conditions more conducive to Fog2 formation, as discussed in Sect.
371 5.1. The reason for the wider distribution (fog-top height and fog area) is discussed here.

372 **5.3.1 Effects of macrophysics on radiation**

373 Meteorological conditions more conducive to Fog2 formation and ACI promote condensation
374 near the fog top (Fig. 10d, f), thereby raising the fog-top height in Fog2 compared with that in
375 Fog1 (black and purple lines in Fig. 10). Therefore, both fog-top height and τ_c in Fog2 are higher
376 than those in Fog1. Compared with that in Fog1, the higher τ_c in Fog2 enhances cooling near
377 the fog top and downwelling long-wave radiation, weakening the cooling at the fog base than
378 near the fog top (Fig. 10c). Additionally, the horizontal distribution of Fog2 is wider than that
379 of Fog1 (Fig. 5b). Therefore, more foggy grid cells show more radiative cooling near the fog
380 top and downwelling long-wave radiation at the fog base in Fog2.

381 **5.3.2 Effects of radiation on turbulence**

382 The above analysis reveals the mechanism underlying the effects of meteorology and ACI on
383 radiation in fog. How does radiation affect stability and turbulence (i.e., TKE)? To answer this
384 question, we must know the dominant factors contributing to TKE, as described in the following
385 TKE budget equation:

$$386 \quad \frac{\Delta TKE}{\Delta t} = TKE_{\text{shear}} + TKE_{\text{buoy}} - TKE_{\text{diss}} + TKE_{\text{mixing}} \quad (3)$$

387 where $\Delta TKE/\Delta t$ is the TKE tendency with time (Fig. 11b), and the four terms on the right side
 388 of Eq. (3) are contributors to TKE, including wind shear (Fig. 11c), buoyancy (Fig. 11d),
 389 dissipation (Fig. 11e), and vertical mixing (Fig. 11f). Detailed equations of these contributions
 390 to TKE are provided in supplementary information (Eqs. S5-S8) (Nakanishi and Niino (2009)).

391 As shown in Fig. 11a, TKE in Fog2 is stronger than that in Fog1, particularly under
 392 polluted conditions. Since the vertical mixing term is one order smaller than the others, it is
 393 negligible (Fig. 11f). At night, only the shear term is positive and, therefore, the main
 394 contributor to TKE (Fig. 11c), consistent with the speculations of Kim and Yum (2012).
 395 However, the dominant term driving the differences in TKE between polluted and clean
 396 conditions is buoyancy (Fig. 11d). As shown in Fig. 11b, $\Delta TKE/\Delta t$ is larger under polluted
 397 conditions than under clean conditions. Meanwhile, the shear term is smaller but the buoyancy
 398 term is larger under polluted conditions than under clean conditions, and the dissipation term is
 399 similar between the two conditions. Therefore, the buoyancy term is the main factor that
 400 increase TKE under polluted conditions, corroborating the qualitative speculations by Jia et al.
 401 (2019). This is particularly true for Fog2. In addition, at daytime, $\Delta TKE/\Delta t$ is weaker under
 402 polluted conditions, because higher τ_c reduces short-wave radiation reaching the surface. These
 403 results are consistent with the higher stability during the dissipation stage under polluted
 404 conditions, as described in Sect. 5.1.

405 After confirming the importance of the buoyancy term, we analyse the effect of radiation
 406 on buoyancy and then on TKE. Buoyancy contributions to TKE are determined by temperature
 407 inversion in the PBL at the night time. As shown in Fig. 12a-b, temperature inversion is close
 408 to the surface. With the effect of ACI, much stronger radiative cooling leads to a more rapid
 409 temperature drop at the fog top than at the fog base (Fig. 12c), thereby causing weaker

410 temperature inversion under polluted conditions. Therefore, stability is weaker and TKE is
411 larger under polluted conditions, particularly in Fog2.

412 **5.3.3 Effects of turbulence on macrophysics**

413 Previous observations (Liu et al., 2010; Román-Cascón et al., 2016) and large eddy simulations
414 (Bergot, 2013; Mazoyer et al., 2017; Nakanishi, 2000) showed that turbulence could increase
415 the fog-top height. In this study, we note that increasing TKE increases fog-top height (black
416 and purple lines in Fig. 10) and fog area (Fig. 5b), which is consistent with observations of Jia
417 et al. (2019) and Quan et al. (2021). The increased fog-top height increases TKE by promoting
418 radiative cooling near the fog top and weakening temperature inversion. This reflects the
419 feedbacks between macrophysics, radiation, and turbulence. Overall, owing to meteorological
420 conditions more conducive to Fog2 formation, the feedbacks are stronger in Fog2 than in Fog1.

421 **6 Conclusion**

422 To explore the interactions between the PBL and ACI, as well as their effects on fog properties,
423 WRF-Chem 4.1.3 is used to simulate two successive radiation fog events that occurs in the
424 northern YRD region in China on 26 and 27 November 2018. Two fog events simulation (Fog1
425 and Fog2) well reproduces the observed results.

426 The results show higher LWC, higher N_d , smaller R_e , higher fog-top height, longer duration,
427 wider spatial distribution, higher LWP, and higher τ_c under polluted conditions than under clean
428 conditions. Aerosol-induced changes in micro and macro-physical properties are more
429 significant in Fog2 than in Fog1. When Fog1 occurs under clean conditions, the response of
430 Fog2 to high aerosol loading becomes weaker. Therefore, ACI with high aerosol loading in
431 Fog1 promotes aerosol-induced changes in Fog2. A conceptual diagram is proposed to describe

432 the mechanism of fog property changes as well as ACI evolution during two successive
433 radiation fog events (Fig. 13). Moreover, the mechanisms of changes in fog properties and ACI
434 evolution are discussed based on the synergistic effects of aerosols and meteorological
435 conditions. The microphysics–radiation feedbacks and macrophysics–radiation–turbulence
436 feedbacks delay Fog1 dissipation, generating more conducive conditions for promoting the
437 earlier formation of Fog2. Furthermore, the microphysics–radiation feedbacks and
438 macrophysics–radiation–turbulence feedbacks are strengthened in Fog2 due to the conditions
439 more conducive to Fog2 formation, enhancing ACI in Fog2 compared with those in Fog1.
440 Detailed mechanisms are summarised below, including meteorological conditions and the two
441 types of feedbacks.

442 First, meteorological conditions before Fog2 formation are more conducive than those
443 before Fog1 formation, which play fundamental roles in changing fog properties and enhancing
444 ACI during two fog events. This is related to the delayed dissipation of Fog1 induced by τ_c .
445 During Fog1 dissipation (daytime), the cooling effect caused by the higher τ_c contributes to the
446 lower temperature, higher relative humidity, and higher stability. At night, cold advection near
447 the ground is enhanced. Meanwhile, the temperature remains low, forming a cold centre, due
448 to low daytime temperature. Moreover, the surface wind diverges outward from the cold centre,
449 strengthening the cold advection. Ultimately, the meteorological conditions induced by high
450 aerosol loading are more conducive for promoting the earlier formation as well as a longer
451 duration of Fog2 than of Fog1.

452 Second, the positive feedbacks between microphysics and radiative cooling are crucial
453 physical mechanisms for changing fog properties and enhancing ACI. In Fog2, high aerosol
454 loading and more conducive meteorological conditions synergistically promote fog
455 microphysics. Lower temperature and higher relative humidity promote aerosol activation and
456 condensation. Consequently, N_d , LWP, and τ_c are higher, whereas R_e is smaller, in Fog2 than in

457 Fog1. Radiative cooling and heating within the fog layer depend on LWP and N_d . When LWP
458 in fog is less than 20 g m^{-2} , and higher aerosol loading enhances vertically integrated cooling
459 and heating in optically thin fog. These variations in microphysics lead to stronger long-wave
460 radiative cooling and condensational growth near the top of Fog2. Therefore, the positive
461 feedbacks between microphysics and radiation are stronger in Fog2, which further promote
462 stronger ACI.

463 Finally, the feedbacks between fog macrophysics, radiation, and turbulence affect fog
464 properties. Under polluted conditions, the higher fog top strengthens the fog-top long-wave
465 radiative cooling and then reduces the strength of temperature inversion near the surface and
466 enhances turbulence. Stronger turbulence further increases the fog-top height and fog area.
467 Because of meteorological conditions more conducive to Fog2 formation, the feedbacks are
468 stronger in Fog2 than in Fog1, contributing to the enhancement of ACI.

469 This study focuses on a two-day radiation fog event in the Yangtze River Delta, China,
470 which has a large population. The conclusions are expected to be applicable to radiation fog
471 events in this region and other regions with similar human activities. It would be interesting to
472 see if similar conclusions can be found in other fog types (e.g., advection fog) in other regions
473 (e.g., ocean). Furthermore, there are large uncertainties in the aerosol–cloud interaction (Fan et
474 al., 2016; Guo et al., 2018; Rosenfeld et al., 2019; Seinfeld et al., 2016; Zhu and Penner, 2020;
475 Zhu et al., 2019). The findings of our study offer novel insights into the potential involvement
476 of mechanisms responsible for evolution of ACI, particularly for stratus, which is similar to fog.

477 Data and code availability. The data repositories have been listed in Sect. 2. Codes are accessed
478 by contacting Chunsong Lu via luchunsong110@gmail.com.

479

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487

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490

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Table 1. Summary of major parameterisation schemes.

Scheme	Option
Microphysics	Morrison
Boundary layer	MYNN
Short-wave radiation	Goddard
Long-wave radiation	RRTMG
Cumulus	Grell 3D
Aerosol chemistry	MOSAIC (4 bins)
Gas phase chemistry	CBMZ

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799 **Table 2.** Evaluation of PM_{2.5} mass concentration. NMB, NME, MFB, and MFE stand for
800 normalised mean bias, normalised mean error, mean fractional bias, and mean fractional error,
801 respectively. Time ‘2514’ (DateHour) indicates 14:00 local standard time (LST) (LST =
802 Universal Time Coordinated + 8 h) on 25 November 2018. The other time expressions follow
803 the same logic.

DateHour	NMB (%)	NME (%)	MFB (%)	MFE (%)
2514-2614	13	25	13	24
2614-2714	38	42	35	38
Total	25	30	24	28

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Table 3. Elements a–d in the Heidke skill score calculation

	Fog observed	No fog observed
Fog simulated	a	b
No fog simulated	c	d

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832 **Table 4.** Quantitative estimation of ACI strength in two fog events (Fog1 and Fog2), including
833 the responses of fog optical depth (τ_c), liquid water path (LWP), and fog effective radius (R_e)
834 to the changes in fog droplet number concentration (N_d). In EXP1, both fog events occur under
835 polluted conditions, and fog events in EXP2 occur under clean conditions. In EXP3, Fog1
836 occurs under clean conditions and Fog2 occurs under polluted conditions. The ratio represents
837 the relative change between Fog1 and Fog2, calculated as $(\text{Fog2} - \text{Fog1})/\text{Fog1}$. In the fourth
838 and sixth columns, Fog1 in both EXP2 and EXP3 occur under clean conditions.

	EXP1 vs EXP2			EXP3 vs EXP2		
	Fog1	Fog2	Ratio	Fog1	Fog2	Ratio
$\Delta \ln \tau_c / \Delta \ln N_d$	0.98	1.32	34.7%	–	1.17	–
$\Delta \ln LWP / \Delta \ln N_d$	0.76	1.08	42.1%	–	1.00	–
$-\Delta \ln R_e / \Delta \ln N_d$	0.22	0.24	9.1%	–	0.17	–

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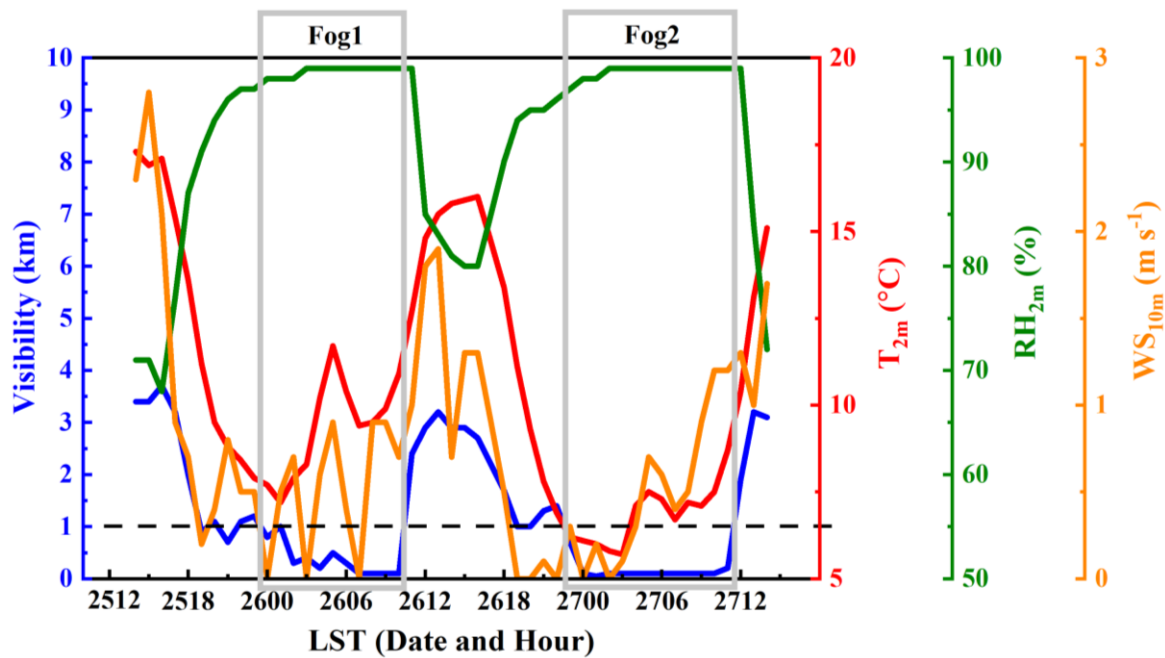
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845 **Table 5.** Average 2 m relative humidity (RH_{2m}) and planetary boundary layer height (PBLH)
 846 above the ground in domain 03 during 12:00–20:00 local standard time (LST) (LST = Universal
 847 Time Coordinated + 8 h) on 25 and 26 November 2018 under clean and polluted conditions.
 848 DIF is the difference in each property between 25 and 26 November.

	Clean			Polluted		
	25 Nov	26 Nov	DIF	25 Nov	26 Nov	DIF
RH _{2m} (%)	76	80	4	76	82	6
PBLH (m)	669	610	−59	670	578	−92

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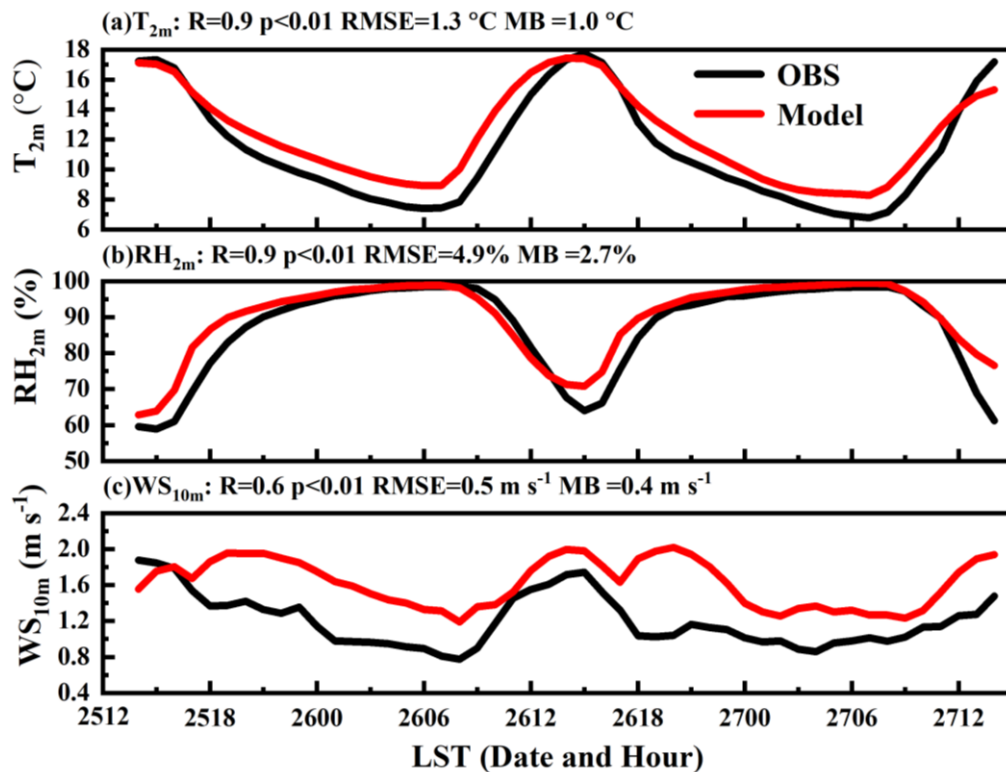
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852 **Figure 1.** The timeseries of visibility, 2 m temperature (T_{2m}), 2 m relative humidity (RH_{2m}),
 853 and 10 m wind speed (WS_{10m}) above the ground at the Nanjing observation site ($31.93^{\circ}N$,
 854 $118.9^{\circ}E$). Fog1 and Fog2 in the light grey box are the two fog events. Time ‘2512’ indicates
 855 12:00 local standard time (LST) (LST = Universal Time Coordinated + 8 h) on 25 November
 856 2018. The other time expressions follow the same logic.

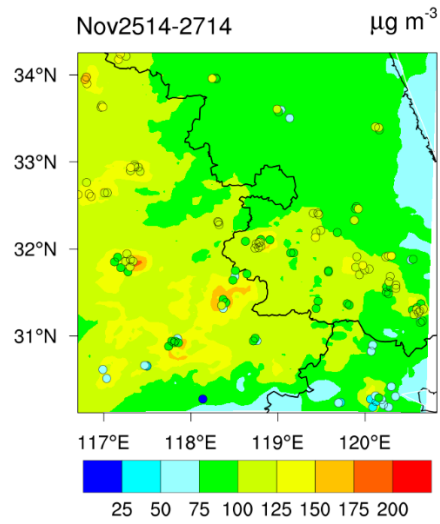
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859 **Figure 2.** Hourly variations in observed (black lines) and simulated (red lines) meteorological
 860 properties, including (a) 2 m temperature (T_{2m}), (b) 2 m relative humidity (RH_{2m}), and (c) 10 m
 861 wind speed (WS_{10m}) above the ground, averaged over 104 meteorological stations in domain
 862 03 from 14:00 local standard time (LST) (LST = Universal Time Coordinated + 8 h) on 25
 863 November to 14:00 LST on 27 November 2018. R, p, RMSE, and MB indicate the correlation
 864 coefficient, significance level, root-mean-square error, and mean bias, respectively. **The**
 865 **equations for RMSE and MB (Eq. S1-S2) are given in the supplement.** Time ‘2512’ indicates
 866 12:00 LST on 25 November 2018. The other time expressions follow the same logic.

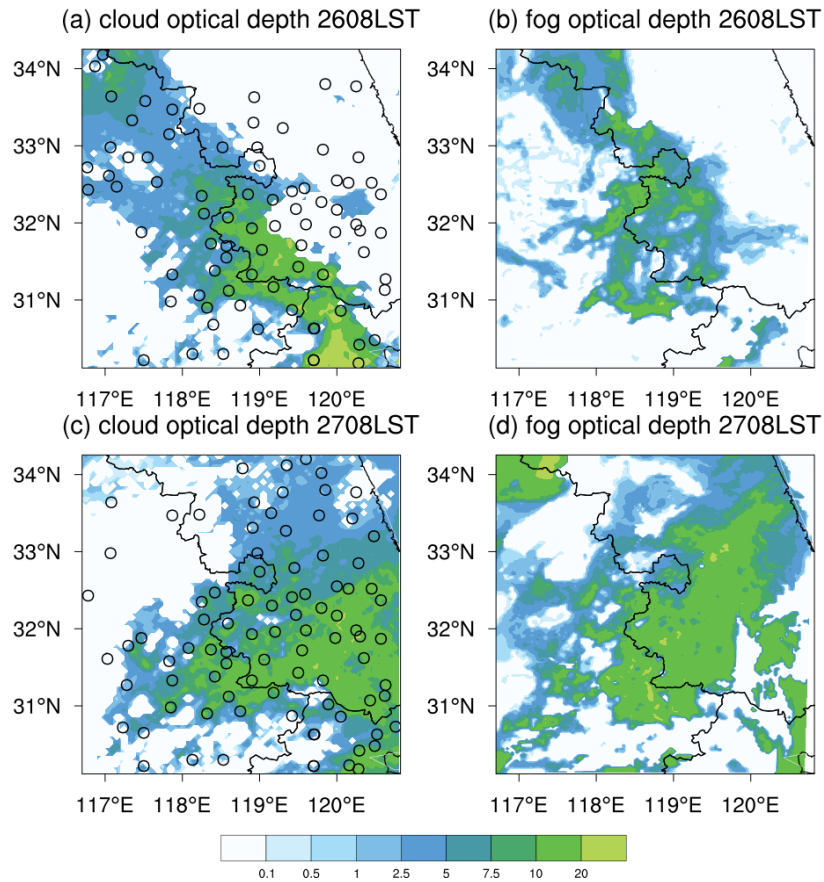
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869 **Figure 3.** Simulated (shaded area) and observed (coloured dots) average distributions of PM_{2.5}
 870 concentration ($\mu\text{g m}^{-3}$) from 14:00 local standard time (LST) (LST = Universal Time
 871 Coordinated + 8 h) on 25 November to 14:00 LST on 27 November 2018.

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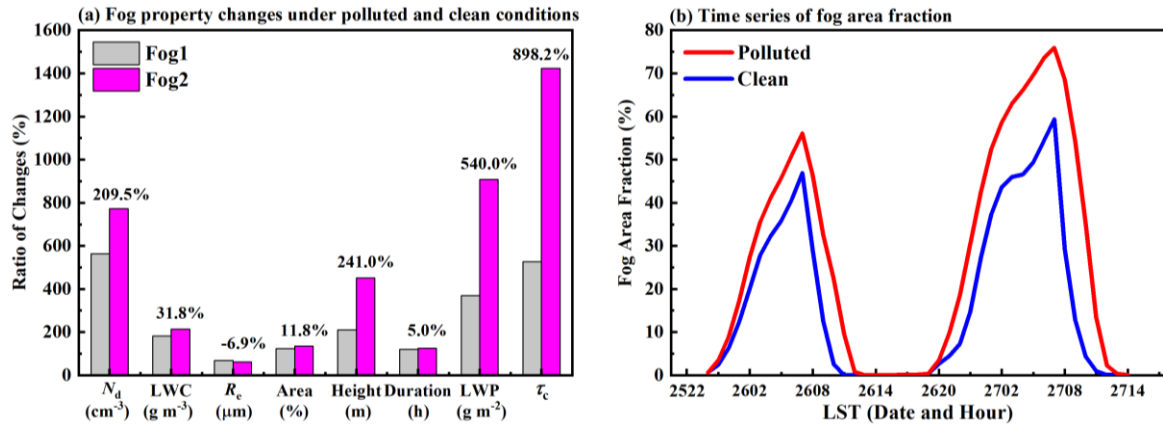
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874 Figure 4. (a, c) Distributions of ground-based fog observations (the black circular points) and
 875 cloud optical depth from Himawari-8 products at 08:00 LST on 26 and 27 November 2018. (b,
 876 d) Simulated fog optical depth distributions in domain 03 at the corresponding time of
 877 observations. Time ‘2608LST’ indicates 08:00 local standard time (LST) (LST = Universal
 878 Time Coordinated + 8 h) on 26 November 2018. The other time expressions follow the same
 879 logic.

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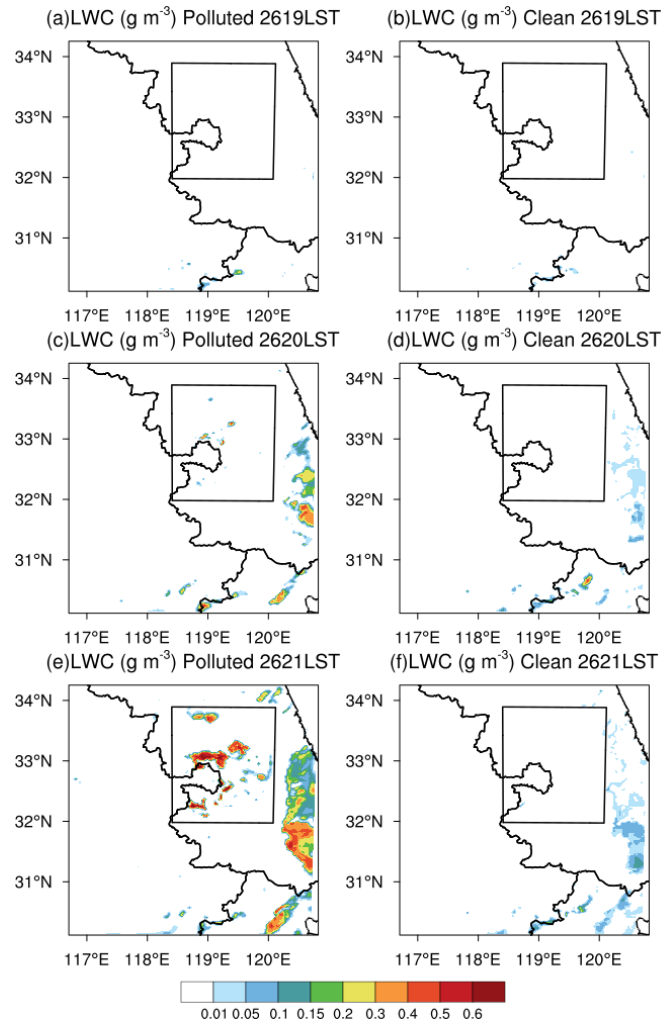
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884 **Figure 5.** (a) Aerosol-induced changes in macro- and microphysical properties during the first
 885 fog (Fog1) and the second fog (Fog2) events under polluted and clean conditions. (b) Temporal
 886 evolution of fog area fraction under clean and polluted conditions. N_d , LWC, R_e , Area, Height,
 887 Duration, LWP, and τ_c indicate fog droplet number concentration, liquid water content,
 888 effective radius, fog area fraction, fog-top height, liquid water path, and fog optical depth,
 889 respectively. The ratios of changes are calculated by Polluted/Clean in Fig. 5a which reveal the
 890 aerosol-induced changes. The numbers above the bars in Fig. 5a represent the difference in
 891 those ratios of changes between Fog1 and Fog2 (calculated by Fog2–Fog1). Time ‘2522’ in Fig.
 892 5b indicates 22:00 local standard time (LST) (LST = Universal Time Coordinated + 8 h) on 25
 893 November 2018. The other time expressions follow the same logic.

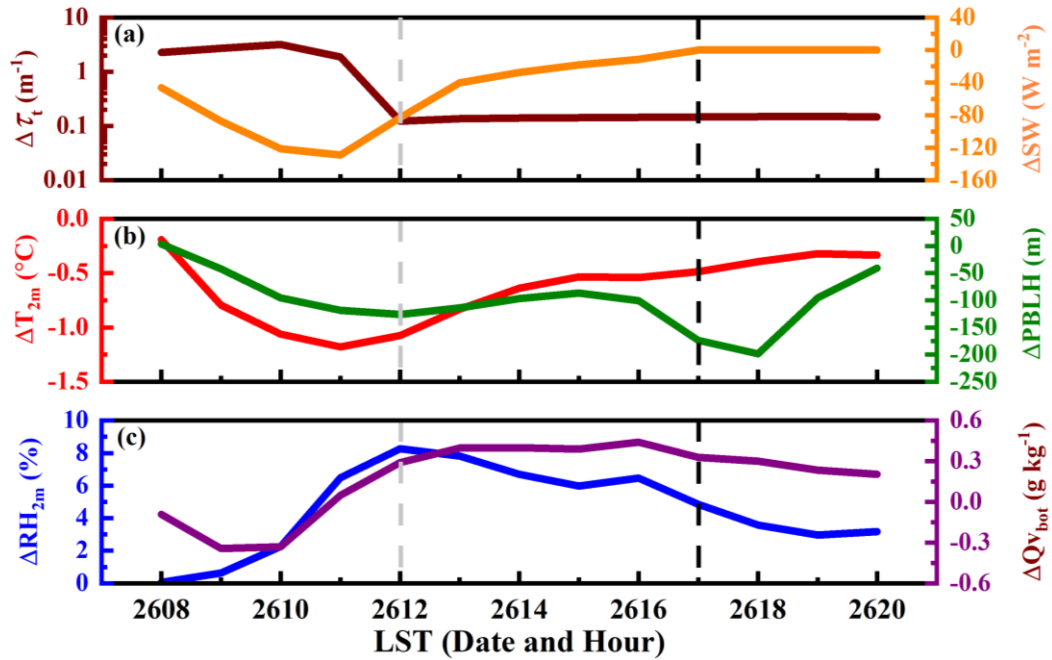
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896 **Figure 6.** Liquid water content (LWC) distribution at the bottom layer from 19:00-21:00 local
 897 standard time (LST) (LST = Universal Time Coordinated + 8 h) on 26 November 2018 under
 898 (a, c, e) polluted and (b, d, f) clean conditions. The black box is the area in which Fog2 formed
 899 earlier under polluted conditions. Time ‘2619LST’ indicates 19:00 LST on 26 November 2018.
 900 The other time expressions follow the same logic.

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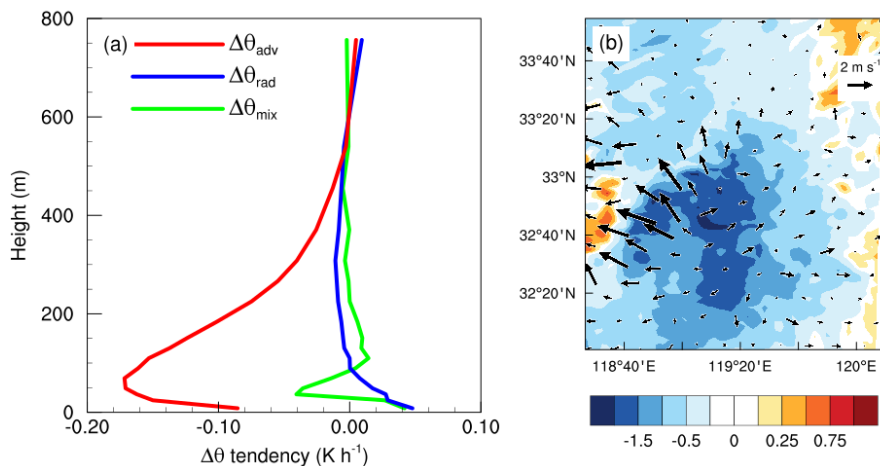


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903 **Figure 7.** Differences in properties between polluted and clean conditions in the black box in
 904 Fig. 6, including (a) total optical depth (τ_t), surface downwelling short-wave radiation (SW), (b)
 905 2 m temperature (T_{2m}), planetary boundary layer height (PBLH), (c) 2 m relative humidity
 906 (RH_{2m}), and water vapour mixing ratio at the bottom of the model ($Q_{v\text{bot}}$), where $\tau_t = \tau_c$ (fog
 907 optical depth) + AOD (aerosol optical depth). The grey dashed line is the time of complete
 908 evaporation of Fog1 under polluted conditions. The black dashed line is the time of sunset.
 909 Time '2608' indicates 08:00 local standard time (LST) (LST = Universal Time Coordinated +
 910 8 h) on 26 November 2018. The other time expressions follow the same logic.

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Average during 17:00-19:00 LST before fog formation



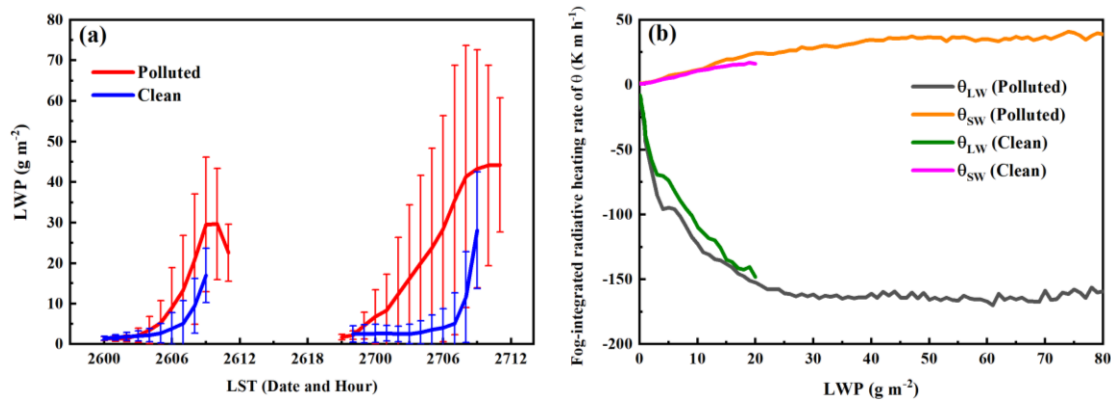
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913 **Figure 8.** (a) Differences (Polluted – Clean) in terms contributing to the potential temperature
914 tendency, including radiation (θ_{rad}), vertical mixing (θ_{mix}), and advection (θ_{adv}) in the black box
915 in Fig. 6 before fog formation (17:00–19:00 local standard time [LST = Universal Time
916 Coordinated + 8 h]). (b) The shaded area represents the mean temperature difference (Polluted
917 – Clean), and vectors represent the mean wind vector difference (Polluted – Clean) at the
918 bottom of the model.

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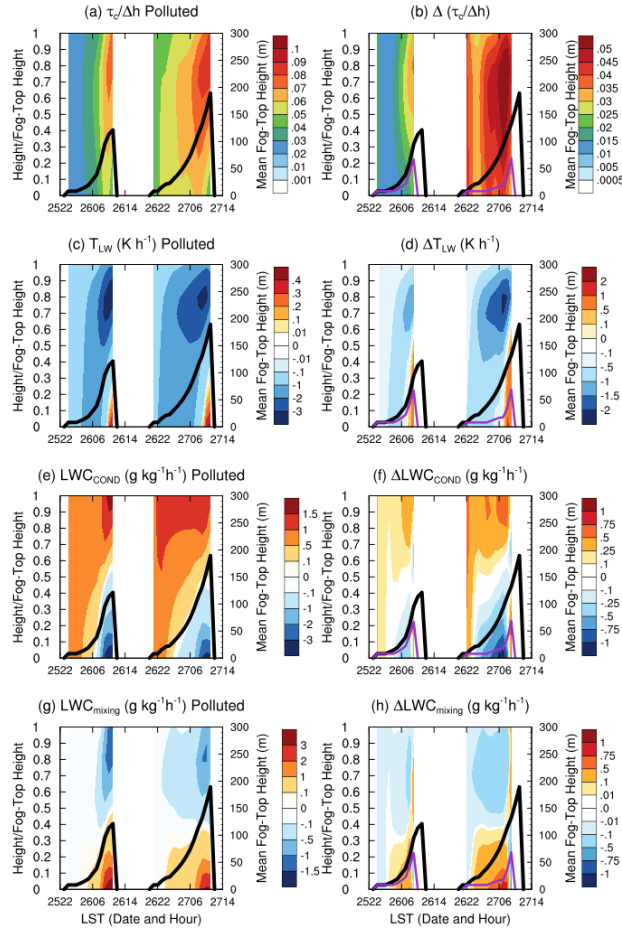
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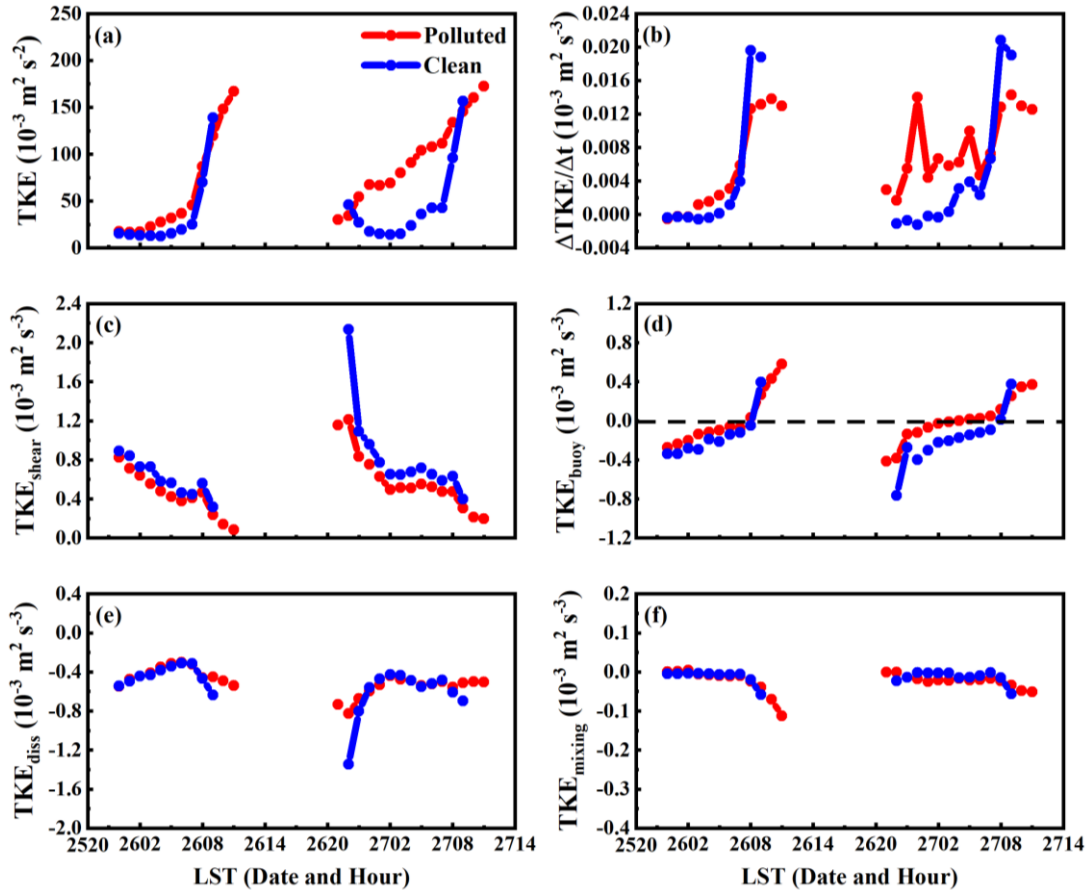
923 **Figure 9.** (a) The timeseries of liquid water path (LWP) under polluted and clean conditions.
 924 The length of the bar represents standard deviation. (b) Dependence of fog-integrated radiative
 925 cooling or heating with LWP under polluted and clean conditions. θ_{LW} and θ_{SW} represent
 926 vertically integrated heating rate of potential temperature (θ) within the fog layer due to long-
 927 wave radiation and short-wave radiation, respectively. Time '2512' indicates 12:00 local
 928 standard time (LST) (LST = Universal Time Coordinated + 8 h) on 25 November 2018. The
 929 other time expressions follow the same logic.

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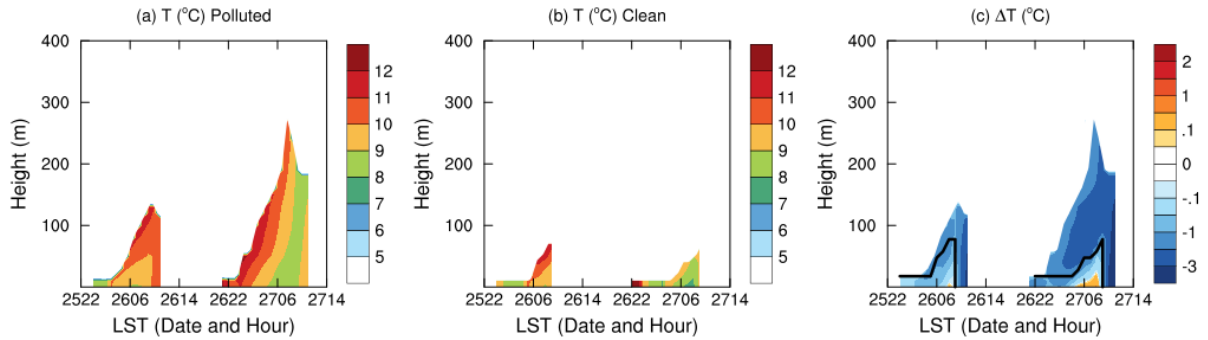
932 **Figure 10.** Time-height profiles of (a-b) average extinction coefficient through the fog layers,
 933 which is fog optical depth (τ_c) at per unit height ($\tau_c/\Delta h$), (c-d) radiative cooling rate (T_{LW}), (e-f)
 934 condensation growth rate (LWC_{COND}), and (g-h) liquid water content tendency due to vertical
 935 mixing (LWC_{mixing}). Heights on the left axes are normalised by the fog-top heights and the left
 936 axes are mean fog-top heights. The left column represents polluted conditions and the right
 937 column represents the difference (Polluted – Clean). Black and purple lines are the mean fog
 938 top heights under polluted and clean conditions, respectively. Time ‘2522’ indicates 22:00 local
 939 standard time (LST) (LST = Universal Time Coordinated + 8 h) on 25 November 2018. The
 940 other time expressions follow the same logic.



942

943 **Figure 11.** (a) Temporal evolution of turbulent kinetic energy (TKE), (b) TKE tendency, (c)
 944 wind shear term (TKE_{shear}), (d) buoyancy term (TKE_{buoy}), (e) dissipation term (TKE_{diss}), and (f)
 945 vertical mixing terms (TKE_{mixing}) under polluted and clean conditions. The dashed line
 946 represents the zero line for TKE_{buoy} . Time ‘2522’ indicates 22:00 local standard time (LST)
 947 (LST = Universal Time Coordinated + 8 h) on 25 November 2018. The other time expressions
 948 follow the same logic.

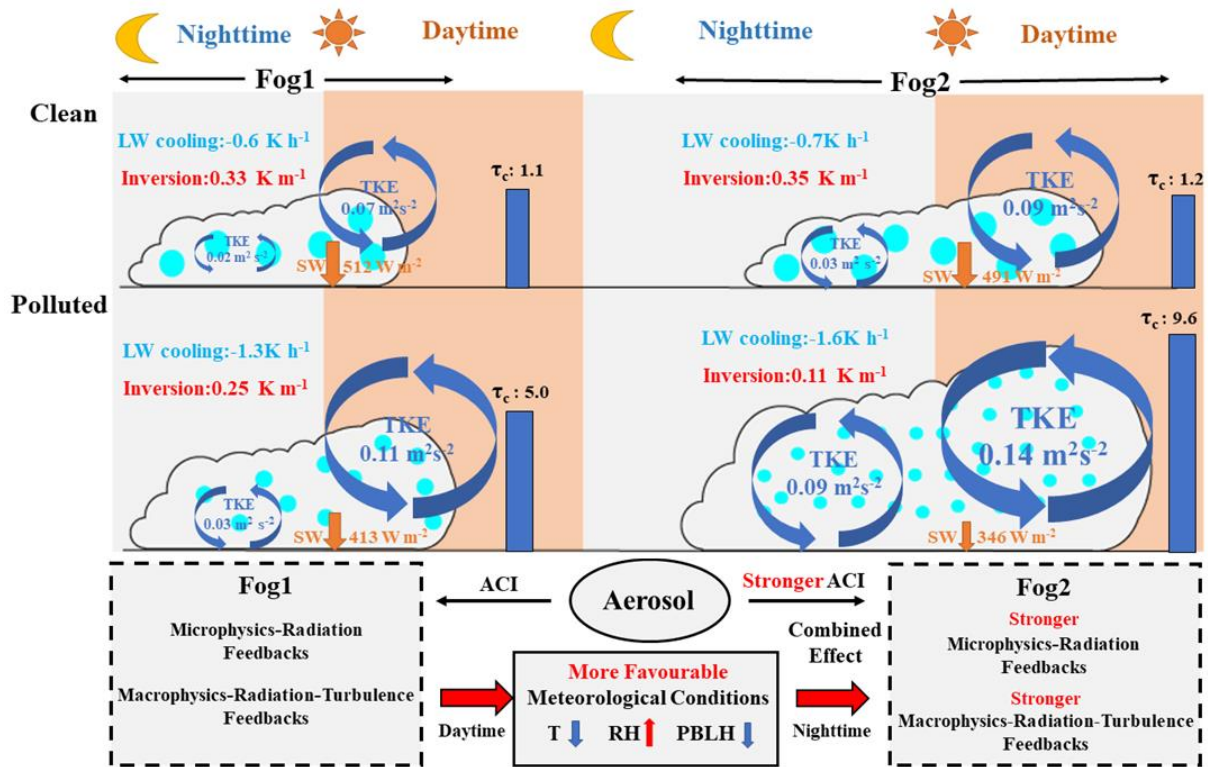
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950

951 **Figure 12.** Time-height profiles of in-fog temperature (T) under (a) polluted and (b) clean
 952 conditions. (c) Difference between polluted and clean conditions. The Black line on the right
 953 side represents the maximal fog-top height under clean conditions. Time '2522' indicates 22:00
 954 local standard time (LST) (LST = Universal Time Coordinated + 8 h) on 25 November 2018.
 955 The other time expressions follow the same logic.

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957

958 **Figure 13.** Conceptual image of interactions between aerosol–fog interaction (ACI) and
 959 planetary boundary layer (PBL). τ_c , SW, LW, TKE, T, RH, and PBLH stand for fog optical
 960 depth, short-wave radiation, long-wave radiation, turbulent kinetic energy, temperature, relative
 961 humidity, and planetary boundary layer height, respectively. LW and inversion are calculated
 962 at night time, and τ_c is calculated at daytime.

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