

1 Numerical modeling of relative contribution of planetary waves to the 2 atmospheric circulation

3 Andrey V. Koval^{1,2}, Olga N. Toptunova², Maxim A. Motsakov², Ksenia A. Didenko^{1,2}, Tatiana S.
4 Ermakova^{1,2}, Nikolai M. Gavrilov¹, Eugene V. Rozanov³

5 ¹Atmospheric Physics Department, Saint-Petersburg State University, Saint Petersburg, 199034, Russia

6 ²Department of Meteorological Forecasts, Russian State Hydrometeorological University, 195196 Saint-Petersburg, Russia

7 ³Physikalisch-Meteorologisches Observatorium, Davos World Radiation Centre, Davos Dorf, 7260, Switzerland

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9 *Correspondence to:* Eugene V. Rozanov (Eugene.Rozanov@pmodwrc.ch)

10 **Abstract.** Using the general circulation model of the middle and upper atmosphere (MUAM), a number of numerical scenarios
11 were implemented to study the impact of individual planetary waves (PWs) on the global atmospheric circulation, including
12 zonal wind, temperature, and residual meridional circulation. The calculations were performed for the winter conditions of the
13 Northern Hemisphere (January–February). The contribution to the formation of the dynamic and temperature regimes of the
14 middle and upper atmosphere made by equatorial Kelvin waves propagating to the east, as well as atmospheric normal modes
15 with periods from 4 to 16 days is shown. In particular, it is demonstrated that the impact of a 5-day PW and an ultrafast Kelvin
16 wave can change the speed of circulation flows by up to ~~56~~% in the areas of their amplitude maxima. At the same time, this
17 effect can be significantly enhanced in certain periods of time. The presented research results are important for a deeper
18 understanding of the mechanisms of large-scale atmospheric interactions. Despite the obviousness and simplicity of the
19 problem, such work has not been carried out at the moment.

20 **Keywords:** planetary waves, normal atmospheric modes, residual meridional circulation, numerical simulation, atmospheric
21 dynamics

22 1 Introduction

23 Planetary waves (PWs, known as Rossby waves) are large-scale variations in the hydrodynamic parameters of the
24 atmosphere (wind, temperature, density), which are formed due to the potential vorticity conservation. The horizontal
25 distribution of PWs is determined by the counteraction of the meridional gradient of the Coriolis force and the meridional
26 displacements of the jet streams. According to the classic theory (e.g., Holton, 1975), a number of waves fit along the latitude
27 circle, ~~which is determined by~~determining the zonal wave number. The amplitudes of PWs increase due to a decrease in the
28 density of the atmosphere, when they propagate from their sources in the troposphere. In the middle and upper atmosphere
29 these disturbances become an important driver of the atmospheric circulation. One of the important features of planetary waves

30 is their active interaction with the mean flow causing transfer of energy and momentum. This feature was reflected in the
31 formulation of the generalized Eliassen Palm theorem (Eliassen and Palm, 1961). PWs can provide a significant acceleration
32 of the background flow in the middle atmosphere when dissipating. This acceleration is comparable to the acceleration
33 associated with gravity waves and atmospheric tides (e.g., Pogoreltsev, 1999).

34 Another important feature of PWs, which explains the need for their comprehensive study, is that they are a link
35 between different atmospheric layers and regions. The PWs can contribute to the signal propagation from the quasi-biennial
36 oscillation (QBO) of the equatorial zonal wind into the thermosphere (Koval et al., ~~2022~~2022a,b) and from the equatorial
37 region to the extratropical region (Holton & Tan, 1980). The ability of PWs to be reflected downward at the heights of the
38 lower thermosphere, due to changes in vertical temperature gradients associated with solar activity cycle, can also have a
39 significant effect on the dynamic and temperature regimes of the middle atmosphere (Koval et al., 2018a).

40 According to the so-called “downward control principle” (Haynes et al., 1991), PWs are the main driving force of
41 meridional extratropical circulation (see also Holton et al., 1995). Due to its global nature, meridional circulation is considered
42 to be the most important mechanism of dynamic interaction between different layers and regions of the atmosphere, affecting
43 the transport of aerosol, atmospheric gases and, consequently, the composition of the atmosphere. Changes in the meridional
44 circulation can affect the ozone layer behavior. The state of the ozone layer has attracted increased attention due to global
45 ozone depletion (e.g., Newman et al, 2009). PWs are the main factor in the development of sudden stratospheric warming
46 (Schoeberl, 1978; Nath et al., 2016).

47 A lot of studies are currently dedicated to the PWs having different periods and zonal wavenumbers. For example,
48 numerical simulations of PWs influence were discussed in Liu et al. (2004); Chang et al. (2014); Wang et al. (2017); Forbes
49 et al. (2018; 2020); He et al. (2020) and many others. Ground based radar measurements were presented by Clark et al. (2002);
50 Jiang et al. (2008); Pancheva et al., (2008) and satellite measurements by Day et al. (2011); Forbes & Zhang (2017); Pancheva
51 et al., (2018); Merzlyakov et al. (2013), as well as processing of reanalysis data/weather forecasting system by Sassi et al.
52 (2011); Qin et al. (2021), etc.

53 In this paper we considered the relative contribution of various PW modes to the formation of the global atmospheric
54 circulation using the unique opportunity that numerical modeling gives us. In order to further understand the nature of large-
55 scale atmospheric dynamics, we carried out a number of numerical experiments to quantify the sensitivity of the zonal wind
56 and temperature fields, as well as meridional circulation components to the switching on/off sources of various PW modes in
57 the model. Despite the obviousness and simplicity of the problem, such work has not been carried out at the moment.
58 Unfortunately, there is no universal way to study the impact of all Rossby waves, each wave has its own characteristics,
59 depending, in particular, on the season, the impact of large-scale processes such as quasi-biennial oscillation of the equatorial
60 zonal wind, El-Nino southern oscillation, etc. Therefore, we have chosen only a part of the PW spectrum, the amplitudes of
61 which are maximized during the boreal winter.

62 2 Methodology

63 **The MUAM model.** Planetary waves are studied using the Middle and Upper Atmosphere Model (MUAM,
64 Pogoreltsev et al., 2007). MUAM is a three-dimensional nonlinear mechanistic model of the general atmospheric circulation
65 at heights from the surface to the F2 ionospheric layer (up to 300-400 km). This is one of the most promising and modern
66 models of atmospheric wave dynamics, which makes it possible study the processes in the middle and upper atmosphere, as
67 well as their interaction with lower levels (see, for example, Gavrilov et al., 2018; Ermakova et al., 2019; Koval et al., 2018a,
68 b; [2022a,b](#); Medvedeva et al., 2019). One of the advantages of MUAM is that it allows us not only to analyze the
69 amplitudes of planetary waves, but also to associate them with various generating sources. The log-isobaric height $x = -$
70 $\frac{H}{p_s} \ln(p/p_s)$ is used as the vertical coordinate in MUAM, where p is the pressure in hPa, p_s is the surface pressure, and H is
71 the pressure scale height. The latitude and longitude spacing of horizontal grid of the model is $5.625^\circ \times 5^\circ$. A version of the
72 model with 56 vertical levels is used, covering a vertical range from the Earth surface to about 300 km. The time integration
73 step is 225 s.

74 The MUAM radiation module takes into account atmospheric net radiative heating due to solar and infrared irradiance.
75 The thermosphere includes parameterization of heating in the extreme ultraviolet band. Ion drag, molecular and turbulent
76 viscosity and thermal conductivity are included as well. The model provides the possibility of planetary waves' excitation near
77 the Earth's surface. The possibility of changing the albedo of the underlying surface is available. Weather changes and
78 cloudiness in the troposphere are not simulated. The MUAM uses three parametrizations of gravity waves with different phase
79 velocities, including orographic waves. For further description of the processes involved in the current version of the model,
80 please refer to Koval et al. ([2022a](#)).

81 The main parameters simulated by the MUAM include 4-dimensional fields of the zonal, meridional and vertical
82 velocity components, geopotential height, and temperature with time step of 2 h. By the MUAM initialization, zonal mean
83 climatological distributions of the geopotential height and temperature are set with the lower boundary conditions at the 1000
84 hPa isobaric level. These distributions were obtained using the reanalysis MERRA-2 data (Gelaro, et al., 2017) and averaged
85 over 20 years (from 2000 to 2019) for January-February.

86 Since the MUAM does not reproduce tropospheric weather, the sources of the westward propagating PWs
87 (atmospheric normal modes, NMs) and the eastward PWs (Kelvin waves) in the MUAM are specified using additional terms
88 in the heat balance equation, having the form of time-dependent sinusoidal harmonics with zonal wavenumbers $m = 1$ and m
89 $= 2, 3$, and periods matching to simulated PWs. To specify the latitudinal structure of the PW components, the corresponding
90 Hough functions obtained using the method described by Swarztrauber and Kasahara (1985) are used. PW periods are equal
91 to the resonant response of the atmosphere to the wave action at the lower boundary (Pogoreltsev, 1999). Westward
92 propagating NMs (1.1), (1.2), (1.3), and (2.1), (2.2) in the classification proposed by Longuet-Higgins (1968) are considered.
93 They have periods of about 5, 10, 16 days with a zonal number of 1, and about 4 and 7 days with a zonal number of 2. In
94 addition, eastward propagating ultrafast Kelvin wave (UFWK, having period of about 3.5 days, a zonal number of 1) are

95 studied. In addition to the mentioned PWs, MUAM also includes sources of slow and fast Kelvin waves ($m=1$), and quasi-
96 two-day wave ($m=3$). However, their amplitudes and contribution to the global circulation during the boreal winter are weak,
97 so they are beyond the scope of this study.

98 The spatial resolution of the model is relatively coarse, however, as the previous studies have shown, this resolution
99 is more than enough to resolve global atmospheric oscillations, including tides (e.g., Suvorova & Pogoreltsev, 2011; Shevchuk
100 et al. 2018; Didenko et al., 2021) and planetary waves (e.g., Gavrilov et al., 2018; Koval et al., 2018a,b; 2022a,b and references
101 therein). Very important drivers of the atmospheric circulation are gravity waves (GWs). Naturally, the GWs (of orographic
102 and non-orographic origin) cannot be resolved by the MUAM, so parameterizations are used to involve their dynamic and
103 thermal effects. There are three of them in model. For GWs having small phase speeds (5-30 m/s) a parameterization by
104 Lindzen (Lindzen, 1981) is implemented. For faster waves with phase speeds of 30-125 m/s, which are in particular important
105 in the thermosphere, a version of the spectral parameterization proposed by Yigit and Medvedev (2009) is applied. The
106 parameterization uses 15 GW spectral components uniformly distributed within the period range from 40 min to 3 h. A third
107 parameterization implemented into the MUAM is responsible for accounting of stationary GWs of orographic origin (Gavrilov
108 and Koval, 2013).

109 **Residual meridional circulation.** A significant problem when considering meridional flows in the framework of the
110 classical Eulerian approach (i.e., with zonal averaging of meridional and vertical circulation flows) is that, in the equations of
111 dynamics, the wave sources of momentum and heat are compensated by advective flows of momentum and heat (Charney and
112 Drazin, 1961). This feature does not allow one to isolate and analyze the wave action on the mean flow. At the same time, in
113 the continuity equation for long-lived gas components, there is a compensation of wave and mean flows. Thus, the use of the
114 Eulerian mean meridional circulation is inefficient for calculating mass transfer and long-lived gas species: and analysing
115 wave-mean flow interaction. A thorough analysis of this topic was made by Butchart (2014). In this study, the Transformed
116 Eulerian Mean (TEM) approach, introduced by Andrews and McIntyre (1976), was used to diagnose the impact of PW on the
117 mean flow. The TEM approach is based on consideration of the components of the mean residual meridional circulation
118 (RMC), which is a superposition of eddy and advective mean transport. Formulas for calculating the RMC components are
119 presented, for example, by Koval et al. (~~2022~~2022a). The time-averaged RMC represents the net average movement of air
120 masses and, therefore, in contrast to the conventional mean Eulerian circulation, it approximates of the average advective
121 movement of atmospheric species.

122 **Scenarios of model experiments.** A series of numerical experiments (model runs) was carried out for January-
123 February to identify the influence of various wave components on the variability of the global circulation and the RMC. The
124 scenarios of the model runs are presented in Table 1: a reference run of the model (#1) was carried out to calculate the
125 atmospheric circulation with the inclusion of all sources of the considered PWs, and other runs were performed with the sources
126 of individual waves turned off. Designations of 4DW, 5DW... mean PWs having periods of 4, 5 days and others. UFKW
127 means ultrafast Kelvin wave. The PW amplitudes were obtained using the longitude-time Fourier expansion into the first 4
128 harmonics applied to the geopotential height fields. Next, an approximation was carried out using the least squares method to

129 the given oscillation periods.

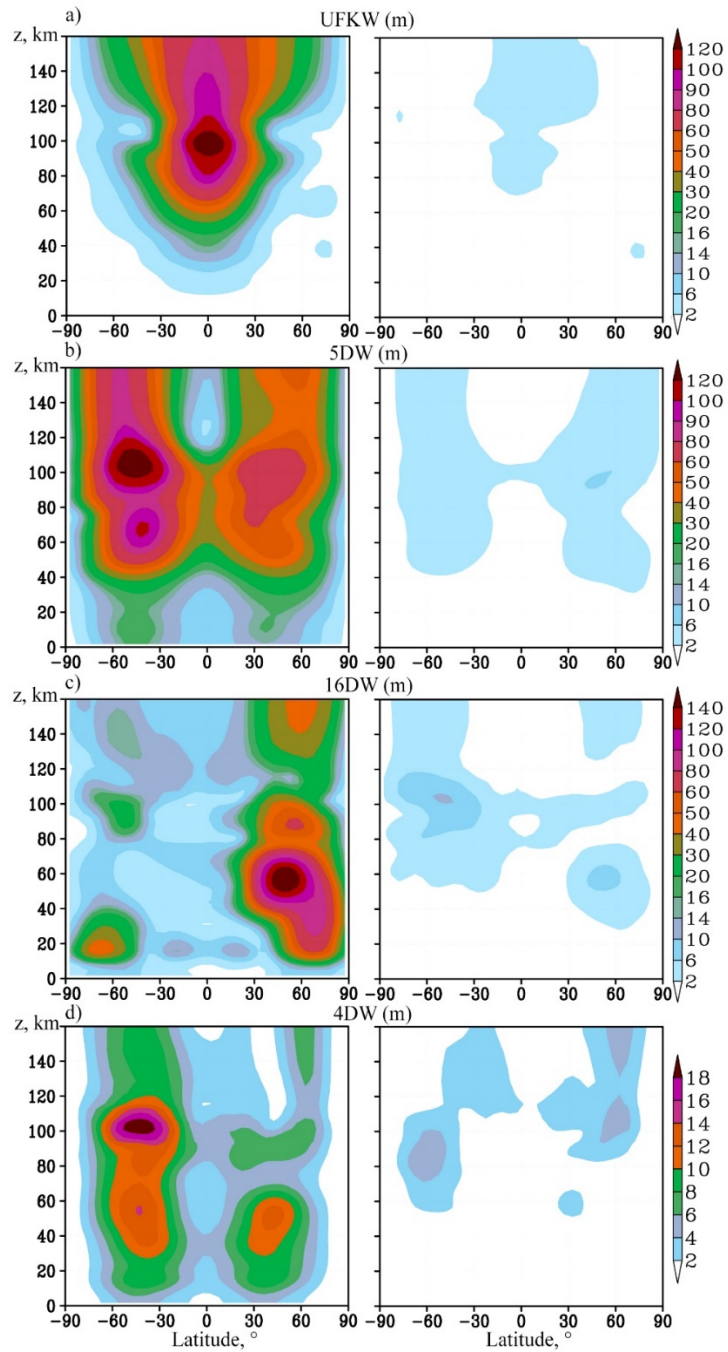
130 **Table 1. Scenarios of model calculations, including different PWs.**

runs	4DW	5DW	7DW	10DW	16DW	UFW
1	+	+	+	+	+	+
2	+	+	+	+	+	
3	+		+	+	+	+
4	+	+	+		+	+
5	+	+	+	+		+
6		+	+	+	+	+
7	+	+		+	+	+

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132 **3 Amplitudes of planetary waves**

133 Fig. 1 shows the amplitudes of geopotential height variations due to the observing planetary waves for January-
134 February. The wave amplitude according to the results of the initial model simulation with the inclusion of sources of all
135 considered PWs (run #1) is presented on the left side. For comparison, the right panels show the amplitudes of these waves for
136 the model simulations with each wave source turned off (see scenarios in Table 1).



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Figure 1. Amplitudes of variations of geopotential height (m) with the source of the respective PW in the MUAM been turned on (left panels) and off (right panels) for the following PW modes: a) Ultrafast Kelvin wave, b) 5-day PW, c) 16-day (all with a zonal wave number $m=1$); d) 4-day (with $m=2$). Note that the color scale is unevenly different for different panels.

141 The amplitude of eastward propagating UFKW (a period of about 3.5 days) is shown in Fig. 1a. Kelvin waves are
142 localized in the low latitude region unlike classic atmospheric NMs, the horizontal structure of which is caused primarily by
143 the action of the Coriolis force weakening ~~hemthem~~ near the equator. The UFKW is mainly excited by the tropospheric source
144 specified in the MUAM. Its generation by internal atmospheric interactions is relatively weak (compare the left and right
145 panels of Fig.1a). The westward propagating NMs, shown in Fig. 1b-d, have maxima in the middle latitudes of both
146 hemispheres. Waves with larger phase velocities (4-d and 5-d NMs) can propagate in both hemispheres (Fig. 1b and 1d), while
147 slower waves predominantly propagate through the eastward wind structures of the winter (in our case – the Northern)
148 hemisphere (Fig.11c). This is due to propagation barriers of these waves occurring when their phase velocity is less than the
149 westward zonal jet stream in the summer stratosphere and mesosphere (see, for example, Charney and Drazin, 1961). The
150 presence of these barriers is also confirmed by the calculation of the refractive index of the atmosphere for the PWs considered.
151 According to Matsuno (1970), PWs propagate along waveguides: regions of positive refractive index. Our calculations showed
152 that in the Southern Hemisphere the waveguide for 10- and 16-day waves is interrupted, preventing their direct upward
153 propagation. These waves propagate to the Southern Hemisphere from the Northern one, crossing the equator in the
154 stratosphere, as was shown, for example, in the study by Koval et al. (2018a).

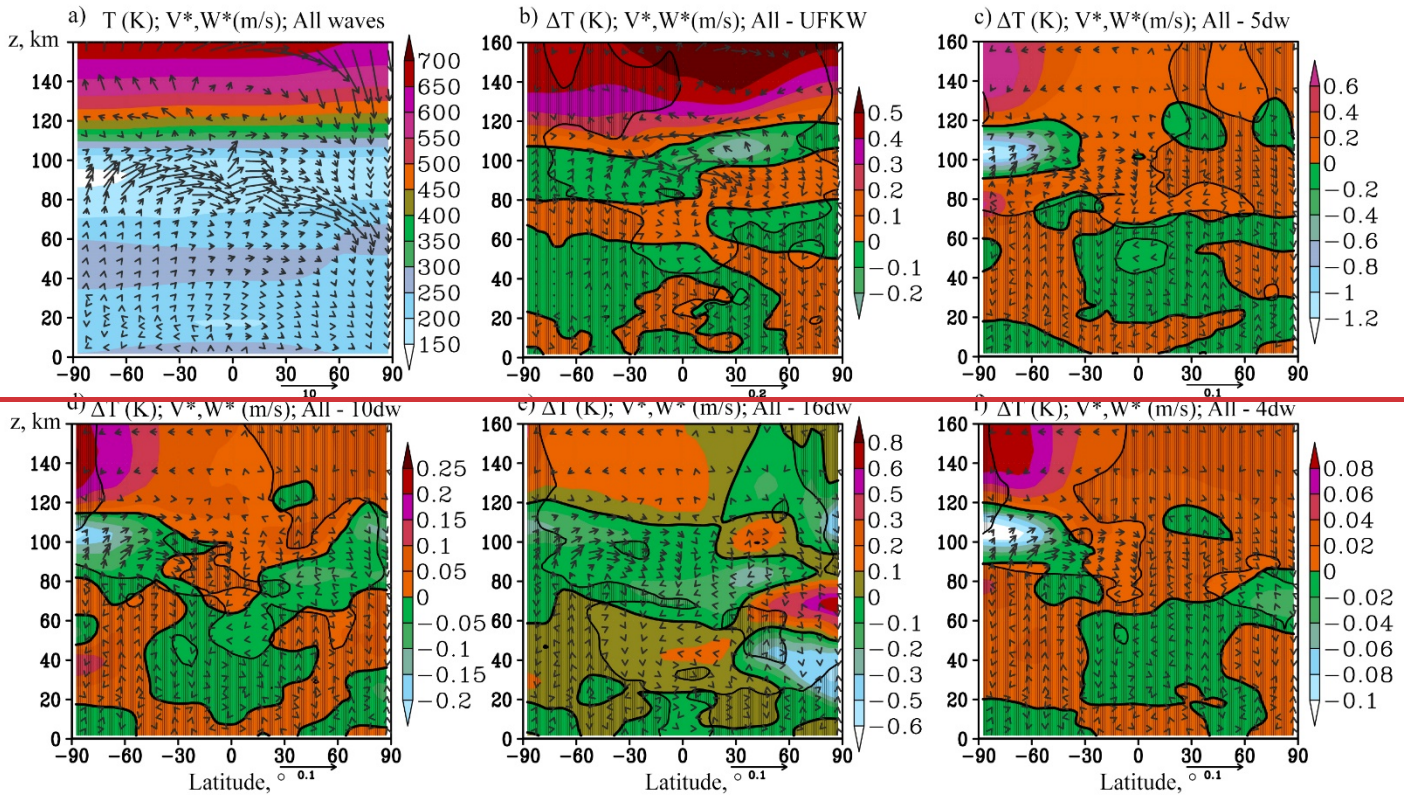
155 Fig. 1 shows the deficiency of waves generation in the middle atmosphere inside the model, and the PW amplitudes
156 with the sources turned off (right panels) do not exceed a numerical noise level. An exception is the maximum amplitude of
157 16-day PW in the right Fig. 1c, which is formed at latitude near 60° S and altitude of about 100 km. When the tropospheric
158 source is turned off, this maximum of geopotential height reaches 15 m in the right panel of Fig. 1c, whereas it is about 24m
159 for the turned-on wave source (the left panel of Fig.1c). This reveals an interesting effect of 16-day PW generating by internal
160 atmospheric sources was discovered. The main source of the 16-day wave generating in the southern lower thermosphere in
161 the MUAM may be elucidated by the nonlinear interaction of the 5- and 4-day waves, whose amplitudes have maxima in the
162 same latitude-altitude region in the left panels of Fig. 1b and 1d. Therefore, further study of this phenomenon is required.

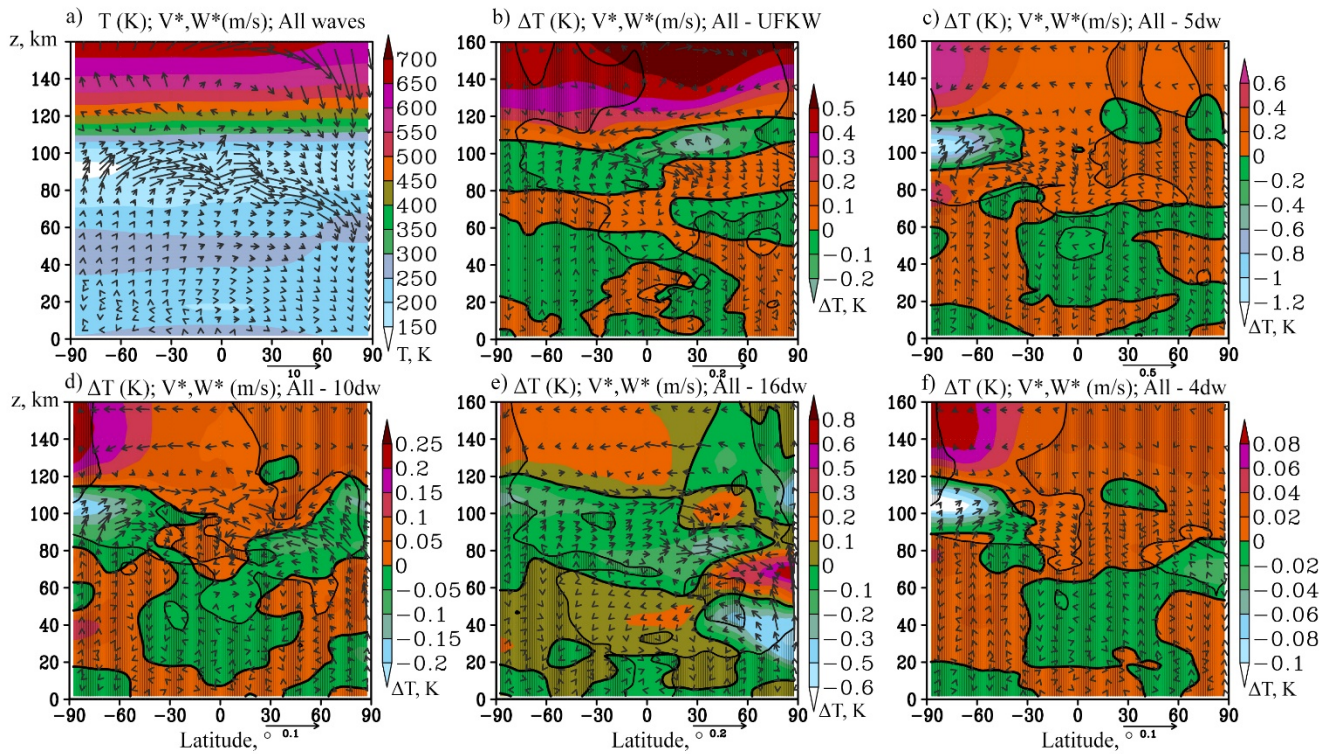
163 A detailed comparison of the MUAM-simulated PW amplitudes for January-February with satellite and radar
164 observations, also with reanalysis data was carried out. For example, the amplitudes of PWs in the geopotential field calculated
165 according to NCEP/NCAR reanalysis data at 10 and 30 hPa pressure levels were presented in the study by Pancheva et al.,
166 (2008). The values of these amplitudes agree with our results. The calculated PW amplitudes in geopotential height according
167 to the MERRA-2 reanalysis data and averaged over the years used for the initialization of MUAM have also similar value and
168 structure to the simulated one's. Additionally, Yamazaki et al. (2021) presents the distributions of 4-day PW amplitudes
169 according to measurements of geopotential height using Microwave Limb Sounder on Aura satellite-microwave sensing, the
170 structure of which corresponds to our calculations. Whereas, the presented values of the PW amplitudes may differ
171 significantly, which is primarily due to the fact that the data for individual specific days are presented in the specified article.
172 The data from the global numerical weather forecasting system (NOGAPS-ALPHA) is used by Sassi et al., (2012) to calculate
173 structures of geopotential height variations by atmospheric NMs. These structures are similar to our distributions. In addition,
174 the 5-day wave amplification in the southern mesosphere similar to the one demonstrated in the left Fig. 1a is shown. For a

175 more detailed analysis of the simulated PWs, in order to compare with the published data, the amplitudes of temperature
176 variations by PWs were also calculated. The simulated 5-day PW and UFKW in temperature field were compared, in particular,
177 with the wave amplitudes calculated from TIMED/SABER temperature data (Pancheva et al., 2010). The amplitude values
178 accordance (up to 6 K at the MLT height for January for 5-day PW at the mid-latitudes of both hemispheres, for UFKW - at
179 the equator) and the spatial distribution accordance of PW across latitudes were found. Moreover, the simulated PW amplitudes
180 correlate in magnitude and spatial distribution with the respective waves obtained in a number of studies (Pancheva et al.,
181 2008, 2009; Forbes et al., 2017; Pedatella & Forbes, 2009; Huang et al., 2017).

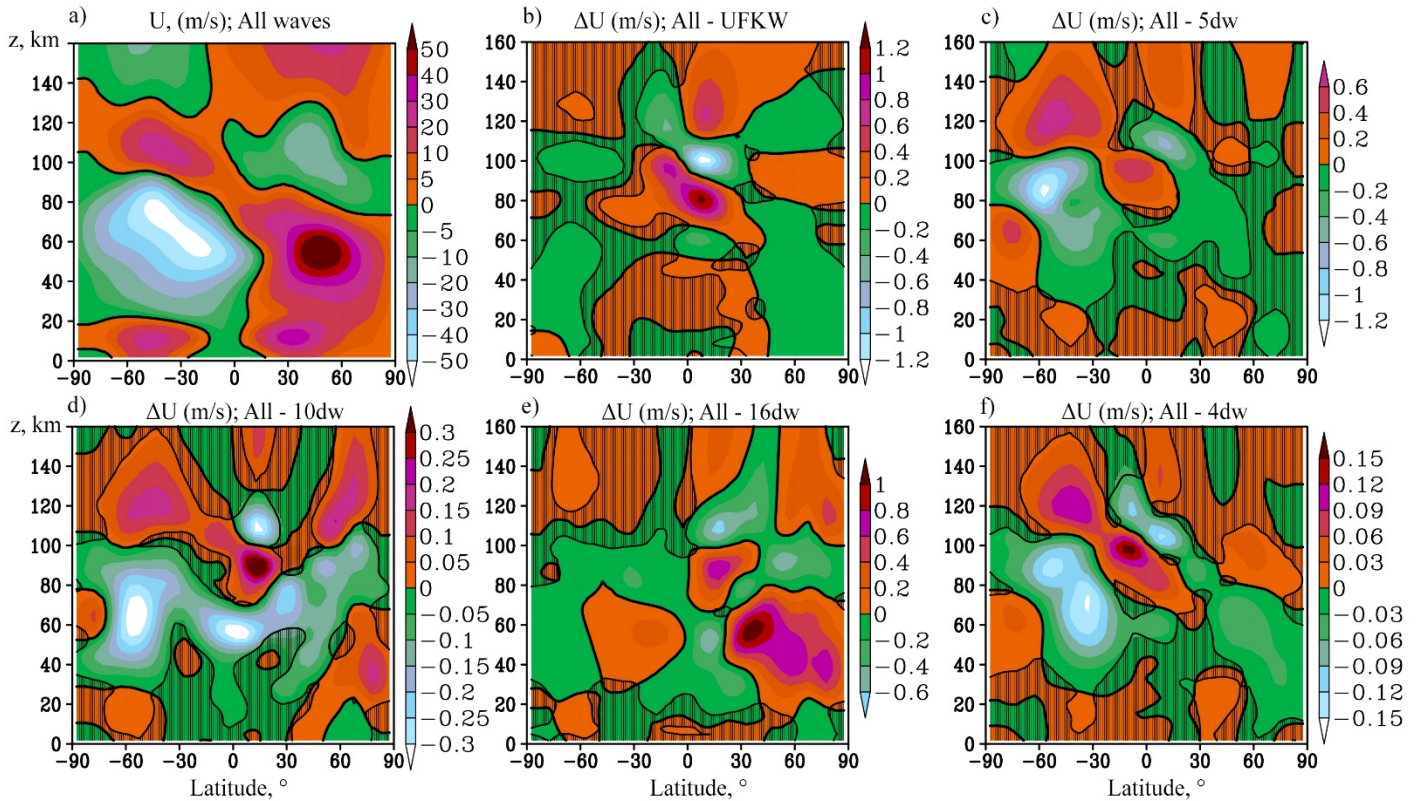
182 **4 Relative PW contribution to the general atmospheric circulation**

183 The residual meridional circulation (RMC) was calculated to analyze the changes in atmospheric circulation caused
184 by various PWs for each MUAM simulation scenario presented in Table 1, with all PW sources turned on for comparisons
185 with model runs at turned-off sources of particular wave modes. The RMC structure should be sensitive to the PW impact as
186 it is a combination of advective and wave-induced eddy components. The latter is driven primarily to PWs according to the
187 “Downward control principle” (Haynes et al., 1991). Fig. 2 shows the RMC components and temperature averaged over
188 January-February for model calculation No. 1 (all PW sources included) and differences in these fields due to turning off each
189 of analysed PW mode. Respective zonal-mean zonal wind increments are shown in Fig. 3. Simulated zonal-mean wind (Fig.
190 3a) and temperature (Fig. 2a) correlate with those obtained with the empirical models HWM-14 (Drob et al., 2015) and
191 NRLMSIS 2.0 (Emmert et al., 2020), also with a semiempirical wind model by Jacobi et al. (2009) and with the MERRA-2
192 reanalysis data.





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 195 **Figure 2. a) RMC components (arrows, m/s, vertical component multiplied by 200) and mean zonal temperature components**
 196 **(colours, K) for January-February with all PW sources turned on; b-f) increments in RMC and temperature due to switching off**
 197 **sources of PW: UFKW, 5-, 10-, 16- and 4-day waves, respectively. Shaded areas show insignificant temperature and/or RMC**
 198 **increments at 95%.**



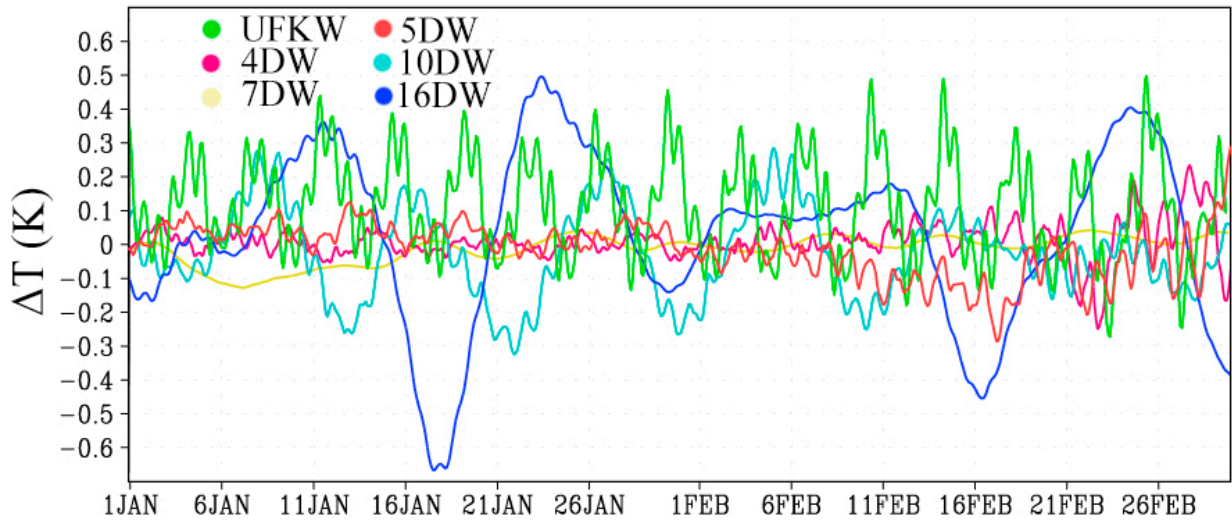
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 200 **Figure 3. a) zonal wind components (colours, m/s) for January-February with all PW sources turned on; b-f) increments in zonal**
 201 **wind due to switching off sources of PW: UFKW, 5-, 10-, 16- and 4-day waves. Shaded areas show insignificant wind increments at**
 202 **95%.**

203 Fig. 2 and 3 show ~~relatively small~~ influence of turning off each individual PW to the zonal-mean temperature and
 204 zonal wind. The main impacts are usually localized in the regions of maximum PW amplitudes. The greatest contribution to
 205 the circulation change is made by 5-day PW. The main differences in Fig. 2c occur in the southern lower thermosphere, which
 206 correspond to a RMC strengthening in a layer between 80 and 120 km after switching on 5-d PW tropospheric source. The
 207 acceleration of zonal wind (eastward above 100 km, and westward below) is observed in the same region in Fig. 3c. This effect
 208 is primarily explained by the convergence of the Elissen-Palm flux (EP) in this region. The acceleration of the RMC there
 209 leads to the lifting up of a warmer air and warming of the atmospheric layer between 60 and 90 km, as well as to the acceleration
 210 of air transport from the coldest region of the atmosphere (about 90 km, at latitudes from the South Pole to 60° N), which leads
 211 to the cooling of the atmosphere above this layer. In addition, in the circumpolar southern stratosphere, at a level of about 60
 212 km, there is deceleration of the zonal wind, which, on the contrary, is associated with the EP flux divergence. The described
 213 changes in RMC and zonal wind between 60 and 120 km can reach values up to 6% forming a significant contribution to the
 214 atmospheric circulation from only one wave. Relative changes in RMC components and zonal wind are presented in Figs.S1b-
 215 S3b in the supplemental information.

216 The maximum UFKW amplitude is located at 100 km in the equatorial region (see Fig. 1b). Then the wave propagates
217 higher, gradually attenuating. Its contribution to the circulation flows changes is also maximized in this region and exerted
218 mainly in the strengthening of the zonal wind (Fig. 3b) and the RMC (Fig. 2b). Similar to 5-day PW, the ~~circulation fluxes~~RMC
219 increments can reach up to 5%–6% as it is shown in Figs. Figs.S1a and S2a. Fig S3a shows that zonal mean wind changes in
220 the equatorial region, between 80 and 120 km can exceed 10% in areas where wind values are greater than 5 m/s. The UFKW
221 impact in the 100-120 km layer leads to cooling in the Northern ~~hemisphere~~Hemisphere caused by a slowdown in meridional
222 transport and additional updrafts causing adiabatic cooling.

223 The impact of the 16-day wave on the circulation, as shown above (Fig. 1e), Figs. 2e and 3e), is comparable in value
224 with 5-day PW and UFKW, however in has a maximum different structure. Maximum PW amplitude occurs in the stratosphere
225 of the Northern ~~hemisphere~~Hemisphere, and its contribution to atmospheric circulation is observed in this region. Figure 2d2e
226 shows that introduction of 16-day wave leads to cooling of the layer below 50 km and heating of the overlying layer. The
227 temperature changes here are explained by the change in the RMC components: in particular, the acceleration and weakening
228 of the RMC descending branch contributes to adiabatic heating and cooling, respectively. This is accompanied by acceleration
229 of the zonal wind (Fig. 3e), directed in this region to the east (Fig. 3a). Statistically significant changes in circulation
230 components may reach 6% in the high-latitude stratosphere as shown in Figs. S1d and S2d. Below, in Figs. 5 and 6 it is shown
231 that action of the 16-day PW may be stronger than 5-day PW and UFKW at certain points in time.

232 10- and 4-day PW make a smaller contribution to the dynamic and thermal regime of the atmosphere. Specifically,
233 the structure of the 10-day wave in the middle atmosphere is similar to the structure of the 16-day one: the amplitude maximum
234 is observed in the northern stratosphere, but due to the higher phase velocity, its waveguide in the southern middle atmosphere
235 is wider. Propagating in the Southern ~~hemisphere~~Hemisphere, it contributes to the zonal wind acceleration up to heights of
236 140 km (Fig. 3d) and to the respective temperature changes. A faster 7-day wave, like 5-day wave, is able to propagate along
237 waveguides in both hemispheres. Generally, the 10- and 7-day PW contributions cause the same effects as the 5-day one
238 described above, although they are much weaker in this region.



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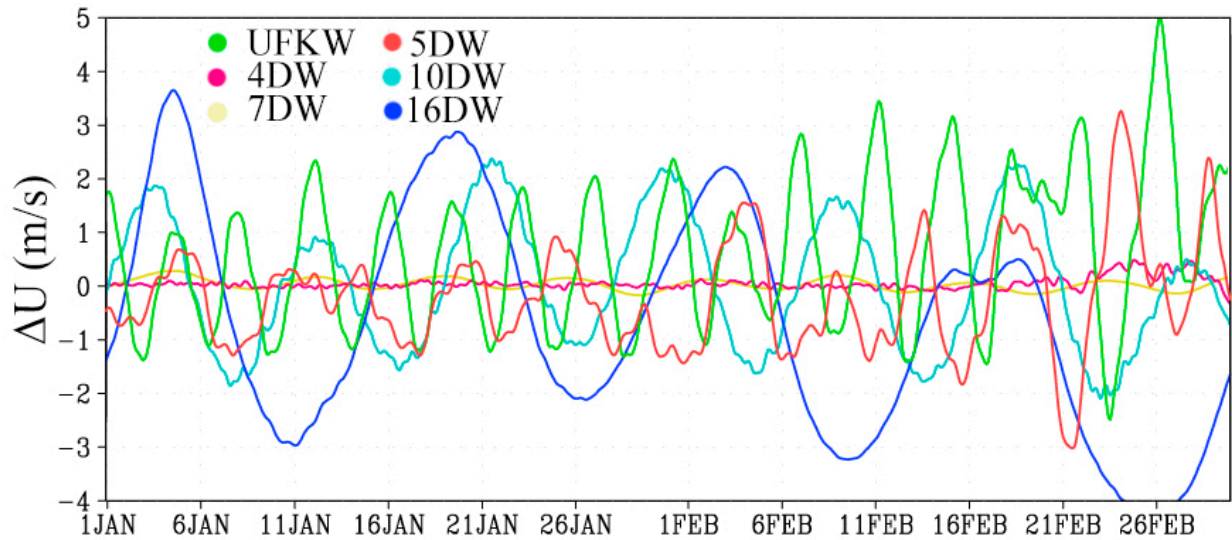
Figure 4. Time series of mean zonal temperature variations due to the inclusion of tropospheric sources of various PW in the regions of their maximum amplitudes in the MUAM.

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The relatively weak increments, examined in Fig. 2 and 3, require an assessment of statistical significance. Such an assessment was carried out using the Student's paired t-test applied to 45312 pairs of samples in each of the latitude-altitude grid node (64 longitude points \times 708 time points for January-February with a 2-hour model output). Statistically insignificant increments at the 95% significance level are marked with shading. In Fig. 4 shading indicates statistically insignificant data on either temperature or RMC.

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For a more detailed analysis of the PW effects on atmospheric circulation, the time series of zonal-mean temperature and zonal wind variations due to the considered PW effects were observed – Fig. 4 and 5, respectively. Latitudes and heights corresponding to the maxima of the PW amplitudes were selected: the equator, 100 km is for the UFKW; 5-day wave is considered at 50° S and 105 km; 7-, 10- and 16-day waves: 50° N and 55 km; 4-day wave: 45° S and 105 km.



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Figure 5. Time series of zonal-mean zonal wind variations due to the inclusion of tropospheric sources of various PW in the regions of their maximum amplitudes in the MUAM.

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In all cases, especially for the zonal wind (Fig. 5), the wave structure of increments with a period corresponding to the period of the considered PW is observed. In particular, wind changes, which significantly exceed the averaged data for January and February (presented in Fig. 3) can be seen in this figure. Specifically, the inclusion of 16-day wave and the UFKW can cause the wind speed changes up to 4 m/s, and up to 5 m/s, respectively. PWs with zonal number 2 (4- and 7-day) make much smaller changes to the zonal flow, while, the weakening of the zonal flow is accompanied by the increase of these waves, as well as the 5-day wave and the UFKW by the end of February. Temperature variations in Fig. 4 have a more complex structure since temperature variations are affected not only by pressure fluctuations, but also by meridional circulation fluctuations.

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5 Conclusion and summary

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A number of model simulation have been carried out for January-February, using a 3-dimensional nonlinear mechanistic numerical model of the general circulation of the middle and upper atmosphere MUAM, to estimate the sensitivity of the atmosphere dynamic and thermal regime to the various planetary waves impact. The MUAM model allows to include selectively sources of various PW modes, which gives the opportunity to deeper study the contribution of each PW to the atmospheric circulation structure. Moreover, for a more detailed diagnostics of the PW effect on the mean flow, the transformed Eulerian mean approach was used, implying the calculation of the residual mean meridional circulation, which is a superposition of eddy and advective mean transport.

The amplitudes of the simulated PWs are consistent with the ground-based, satellite observations data, as well as with

272 the reanalysis and assimilation of meteorological data. The obtained increments of hydrometeorological parameters are
273 maximal, as a rule, in the regions of maximum amplitudes of the considered PWs. In particular, the inclusion of 5-day PW and
274 an UFKW can transform the ~~speed of the background wind and the component~~ components of the residual meridional
275 circulation up to ~~5%-6% each forming a significant contribution to the atmospheric circulation. The impact of the 16-day wave~~
276 ~~on the circulation is comparable in value with 5-day PW and UFKW, however in has different structure. Changes in circulation~~
277 ~~components occur in the high-latitude stratosphere and may reach up to 6%.~~ In turn, ~~all the above mentioned~~ changes in the
278 meridional circulation, especially its vertical component, as well as a ~~variation~~ variations of wave activity fluxes, can cause
279 variations in the background temperature of more than 1 K. ~~At the same time, at certain moments, this effect is much stronger.~~
280 ~~In addition, the waves can be superimposed on one another, and their effect can be summarized. I.e., the cumulative effect of~~
281 ~~the considered waves can significantly increase at certain moments of time.~~

282 The effect of 16-day PW generation by an internal atmospheric source in the southern lower thermosphere,
283 independent of the tropospheric PW sources specified in the model, was found. Most probably, the point is that 4-day PW with
284 a wave number 2 interacts nonlinearly with a 5-day PW with a wave number 1 causing a secondary wave excitation. Such
285 mechanism is described, e.g., by Pogoreltsev (2001): when two waves having frequencies ω and zonal numbers m interact, a
286 new (secondary) wave arises, in which the frequency and wave number are the sum or difference of the corresponding values
287 of the primary waves. ~~Hence, the direct effect of the PWs can be enhanced due to their nonlinear interactions. Finally, this~~
288 ~~causes deceleration of the mean flow, creating better conditions for the SSW onset (e.g., Pogoreltsev et al., 2014).~~ However,
289 additional calculations are required to confirm this theory.

290 In addition, it should also be noted that for proper modelling of large-scale atmospheric dynamics, all models of the
291 general atmospheric circulation should be tested for the ability to reproduce the global resonant properties of the atmosphere
292 (the so-called atmospheric normal modes). This possibility has been repeatedly described in MUAM (e.g., Pogoreltsev, 2007,
293 Koval et al., 2021), which underlines the reliability of the results obtained.

294
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300
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