# Particle number concentrations and size distributions in the stratosphere: Implications of nucleation mechanisms and particle microphysics

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Abstract. While formation and growth of particles in the troposphere have been extensively 22 23 studied in the past two decades, very limited efforts have been devoted to understanding these in the stratosphere. Here we use both Cosmics Leaving OUtdoor Droplets (CLOUD) laboratory 24 measurements taken under very low temperatures (205-223K) and Atmospheric Tomography 25 26 Mission (ATom) in-situ observations of particle number size distributions (PNSD) down to 3 nm 27 to constrain nucleation mechanisms and to evaluate model simulated particle size distributions in 28 the lowermost stratosphere (LMS). We show that the binary homogenous nucleation (BHN) scheme used in most of the existing stratospheric aerosol injection (a proposed method of solar 29 radiation modification) modeling studies overpredict the nucleation rates by 3-4 orders of 30

magnitude (when compared to CLOUD data) and particle number concentrations in the 31 background LMS by a factor  $\sim 2-4$  (when compared to ATom data). Based on a recently developed 32 33 kinetic nucleation model, which gives rates of both ion-mediated nucleation (IMN) and BHN at low temperatures in good agreement with CLOUD measurements, both BHN and IMN occur in 34 the stratosphere. However, IMN rates are generally more than one order of magnitude higher than 35 BHN rates and thus dominate nucleation in the background stratosphere. In the Southern 36 Hemisphere (SH) LMS with minimum influence of anthropogenic emissions, our analysis shows 37 that ATom measured PNSDs generally have four apparent modes. The model captures reasonably 38 well the two modes (Aitken mode and the first accumulation mode) with the highest number 39 concentrations and the size-dependent standard deviations. However, the model misses an apparent 40 second accumulation mode peaking around 300–400 nm, which is in the size range important for 41 aerosol direct radiative forcing. The bi-modal structure of accumulation mode particles has also 42 been observed in the stratosphere well above tropopause and in the volcano-perturbed stratosphere. 43 We suggest that this bi-modal structure may be caused by the effect of charges on coagulation and 44 growth, which is not yet considered in any existing models and may be important in the 45 46 stratosphere due to high ionization rates and long lifetime of aerosols. Considering the importance of accurate PNSDs for projecting realistic radiation forcing response to stratospheric aerosol 47 48 injection (SAI), it is essential to understand and incorporate such potentially important processes in SAI model simulations. 49

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### 53 1. Introduction

54 Solar radiation modification (also known as solar geoengineering) approaches are being developed in response to the climate crisis (IPCC, 2021). They would temporarily offset climate 55 change by reducing incoming sunlight, augmenting (currently inadequate) mitigation efforts and 56 buying time to reduce atmospheric levels of CO<sub>2</sub>, which is the root cause of the climate crisis. A 57 recent report by the National Academies of Sciences, Engineering and Medicine (NASEM) 58 emphasizes the urgent need to have a comprehensive understanding of the feasibility and potential 59 risks/benefits of solar climate intervention approaches (NASEM, 2021). Stratospheric aerosol 60 injection (SAI) has demonstrated the most promise as proximately engineerable (Shepherd et al., 61 2009; Lockley et al., 2020; IPCC, 2021) and has been extensively studied using models (e.g., 62 GeoMIP: Kravitz et al., 2011; GLENS: Mills et al., 2017; Richter et al., 2022). The NASEM report 63 (NASEM, 2021) pointed out that "the overall magnitude and spatial distribution of the forcing 64 produced by SAI depends strongly on the aerosol size distribution" and "One of the research 65 priorities for SAI is thus to address critical gaps in knowledge about the evolution of the aerosol 66 particle size distribution". In the stratosphere, sulfate aerosols are formed by nucleation, followed 67 by condensational growth and coagulation, and lost by evaporation in the upper stratosphere and 68 downward sedimentation into the troposphere (Turco et al., 1982). New particle formation (NPF) 69 (or nucleation) affects not only the number abundance but also the size distributions of 70 stratospheric particles (e.g., Brock et al., 1995; Lee et al., 2003). There is increasing evidence 71 72 (Weisenstein et al., 2022, Laakso et al., 2022) that a careful treatment of microphysical processes is necessary for projecting realistic radiative forcing response to SAI. 73

The process of NPF under tropospheric conditions has been extensively explored over the last 74 two decades through laboratory and field measurements, theoretical studies, and numerical 75 simulations (e.g., Yu and Turco, 2000; Vehkamäki et al., 2002; Kulmala et al., 2004; Kirkby et al., 76 2011; Dawson et al., 2012; Zhang et al., 2012; Kürten et al., 2016; Yu et al., 2018; Kerminen et 77 78 al., 2018; Lee et al., 2019). Although some of the advances in our understanding of nucleation gained in the last two decades can be applied to stratospheric conditions, focused studies 79 specifically examining the mechanisms of NPF under stratospheric conditions are quite limited. 80 Indeed, the H<sub>2</sub>SO<sub>4</sub>-H<sub>2</sub>O binary homogenous nucleation (BHN) parameterization developed two 81 decades ago by Vehkamäki et al. (2002) (named BHN V2002 thereafter) has been widely used in 82 SAI modeling studies when nucleation process is explicitly considered (e.g., Tilmes et al., 2015; 83 Jones et al., 2021; Weisenstein et al., 2022). Tilmes et al. (2015) described a Geoengineering 84 85 Model Intercomparison Project (GeoMIP) experiment designed for climate and chemistry models, using the stratospheric aerosol distribution derived from the ECHAM5-HAM microphysical model 86 (Stier et al., 2005) which calculated nucleation rates with the BHN V2002 scheme. Both models 87 (UKESM1 and CESM2-WACCM6) employed for a recent GeoMIP G6sulfur study (Jones et al., 88 2021) used the BHN V2002 scheme. In another recent SAI study based on three interactive 89 stratospheric aerosol microphysics models (Weisenstein et al., 2022), two models (MAECHAM5-90 HAM and SOCOL-AER) used BHN V2002 scheme while the other (CESM2-WACCM) used an 91 empirical nucleation scheme to calculate nucleation rate as a function of sulfuric acid concentration 92 only (i.e, no dependence on temperature and relative humidity). To our knowledge, the 93 performance of this widely used BHN V2002 under stratospheric conditions has not been 94 carefully examined, probably due to the lack of suitable in situ measurements of freshly nucleated 95 particles in the stratosphere for constraining the scheme. In this regard, particle size distributions 96 down to 3 nm measured in-situ during the NASA Atmospheric Tomography Mission (ATom) in 97

98 the lowermost stratosphere (LMS) of both SH and NH in four different seasons (Williamson et al.,

- 99 2019, 2021; Kupc et al., 2020; Brock et al., 2021) provide much-needed data to constrain our
- 100 understanding of the nucleation and particle microphysics in the stratosphere. In addition, well-
- 101 controlled CLOUD experiments taken under low temperature (within the range of stratosphere)
- 102 can also be used to assess the performance of nucleation schemes under stratospheric conditions.
- Another important issue related to stratospheric particles is the role of ionization in nucleation. It
  is well established that nucleation of H<sub>2</sub>SO<sub>4</sub>-H<sub>2</sub>O on ions is favored over homogenous nucleation
  (Hamill et al., 1982; Yu and Turco, 2000; Lovejoy et al., 2004; Kirkby et al. 2011; Yu et al., 2018)
  but the role of ionization in NPF in the stratosphere has not been considered in any previous SAI
  studies (to our knowledge) in spite of the very high ionization rates in the stratosphere.
- In this study, we use both CLOUD laboratory measurements taken under very low stratospheric temperatures and ATom PNSD measurements in LMS to constrain nucleation mechanisms and model simulated particle size distributions. For 3-D simulation of size-resolved stratospheric aerosols, we use the GEOS-Chem with the unified tropospheric-stratospheric chemistry-transport model with the size-resolved advanced particle microphysics (APM) package.
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# 114 **2. Model and data**

# 115 2.1 GEOS-Chem/APM

The GEOS-Chem model is a global 3-D model of atmospheric composition (e.g., Bey et al., 116 2001) and is continuously being improved (e.g., Luo et al., 2020; Holmes et al., 2019; Keller et al., 117 2014; Murray et al., 2012; Pye and Seinfeld, 2010; van Donkelaar et al., 2008; Evans and Jacob, 118 2005; Martin et al., 2003). The GEOS-Chem tropospheric-stratospheric unified chemistry 119 extension (UCX; Eastham et al., 2014), now the standard GEOS-Chem configuration, implements 120 stratospheric chemistry, calculation of J-values for shorter wavelengths, and improved modeling 121 122 of high-altitude aerosols. Extension of the chemistry mechanism to include reactions relevant to the stratosphere enables the capturing of stratospheric responses and troposphere-stratosphere 123 coupling. UCX adds 28 species and 104 kinetic reactions, including 8 heterogeneous reactions, 124 along with 34 photolytic decompositions. Atomic oxygen [both  $O(^{3}P)$  and  $O(^{1}D)$ ] is explicitly 125 modeled; although also of short lifetime in the stratosphere, these species are important in correctly 126 modeling stratospheric chemistry. Photochemistry is extended up to the stratopause to high-energy 127 photons (177 nm) using the Fast-JX model, which includes cross-section data for many species 128 129 relevant to the troposphere and stratosphere. Photolysis rates respond to changes in the stratospheric ozone layer. Additional heterogeneous reactions (Kirner et al., 2011, Rotman et al., 130 2001, Shi et al., 2001) are included to capture seasonal ozone depletion. H<sub>2</sub>O is treated as a 131 chemically-active advected tracer within the stratosphere. These permit chemical feedbacks 132 between stratospheric ozone and aerosols and tropospheric photochemistry. The improved GEOS-133 Chem with coupled stratospheric-tropospheric responses has been evaluated with sonde and 134 satellite measurements of O<sub>3</sub>, HNO<sub>3</sub>, H<sub>2</sub>O, HCl, ClO, NO<sub>2</sub> and stratospheric intrusions (Eastham 135 et al., 2014; Gronoff et al., 2021; Knowland et al., 2022). Yu and Luo (2009) incorporated a size-136 resolved (sectional) APM package into GEOS-Chem, henceforth referred to as GC-APM. The 137 APM separates secondary particles from primary particles, uses 40 bins to represent secondary 138 particles with high size resolution for the size range important for the growth of nucleated particles 139 to accumulation mode sizes, and contains options to calculate nucleation rates based on different 140

nucleation schemes. In GEOS-Chem/APM, nucleation is calculated before condensation using a 141 142 time-splitting technique. Therefore, no competition between nucleation and condensation for 143 sulfuric acid vapor is considered. In most conditions, nucleation consumes only a very small fraction (<1%) of sulfuric acid vapor in the air and the time splitting does not affect the results. 144 When nucleation rate is high, reduced time step for nucleation and growth is used to ensure that 145 146 the fraction of sulfuric acid vapor consumed by nucleation each time step is small. The GC-APM uses a semi-implicit scheme to calculate sulfuric acid condensation together with sulfuric acid gas 147 phase production to ensure that the change of sulfuric acid vapor concentration is smooth. APM is 148 fully coupled with GEOS-Chem in both the troposphere and stratosphere, and is employed for the 149 present study. 150

In the present study we have carried out GEOS-Chem-UCX/APM global simulations from 151 01/2015 to 05/2018, with the first 17 months as spin-up and the remaining period covering ATom 152 1-4 periods (06/2016-05/2018). The horizontal resolution is  $4^{\circ} \times 5^{\circ}$  and there are 72 vertical layers. 153 Emissions from different sources, regions, and species are computed via the Harvard-NASA 154 Emissions Component (HEMCO) on a user-defined grid (Keller et al., 2014). Historical global 155 anthropogenic emissions are based on the Community Emissions Data System (CEDS) inventory 156 (Hoesly et al., 2018). Regional anthropogenic emissions over the United States, Canada, Europe, 157 and East Asia are replaced by regional emission inventories of the National Emissions Inventory 158 https://www.epa.gov/air-emissions-inventories/2017-national-emissions-inventory-nei-159 (NEI, data), the Air Pollutant Emission Inventory (APEI, https://www.canada.ca/en/environment-160 climate-change/services/pollutants/air-emissions-inventory-overview.html), the Co-operative 161 162 Programme for Monitoring and Evaluation of the Long-range Transmission of Air Pollutants in Europe (EMEP, https://www.emep.int/index.html), and the MIX Asian emission inventory (Li et 163 al., 2017), respectively. Monthly mean aircraft emissions are generated based on the Aviation 164 Emissions Inventory v2.0 (Stettler et al., 2011). The aircraft particle emissions include nucleation 165 mode sulfate particles (Emission index =  $2 \times 10^{17}$  /kg-fuel, mean diameter = 9 nm, based on Kärcher 166 et al., 2000), and black carbon and primary organic carbon (POC) particles. Global biomass 167 burning is taken from Global Fire Emissions Database version 4 (van der Werf et al., 2017). The 168 volcanic emissions of SO<sub>2</sub> are taken from AeroCom point-source data (Carn et al., 2015). Fixed 169 global surface boundary conditions are applied for N<sub>2</sub>O, CFCs, HCFCs, halons, OCS and long-170 171 lived organic chlorine species (Eastham et al., 2014).

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### 173 2.2 Airborne ATom measurements of PNSD

Measurements are essential in advancing our understanding of stratospheric aerosol properties 174 and the fundamental processes governing these properties. NASA's Atmospheric Tomography 175 Mission (ATom; Wofsy et al., 2021; Thompson et al., 2022) is a multi-agency effort that provides 176 global in situ aircraft observations of the vertical structure of aerosols from near surface to ~12 km 177 altitude. PNSDs are measured using the NOAA Aerosol Microphysical Properties (AMP) package 178 (Brock et al., 2019) comprising nucleation-mode aerosol size spectrometer(s) (NMASS) 179 (Williamson et al. 2018), ultra-high-sensitivity aerosol spectrometer(s) (UHSAS) (Kupc et al. 180 2018), and a laser aerosol spectrometer (LAS) covering aerosol sizes from 3 nm to 4.5 µm. The 181 aerosol number abundance can be obtained by integrating the PNSD measurements. 182 183

### 184 2.3 The CLOUD (Cosmics Leaving OUtdoor Droplets) measurements

Laboratory measurements of nucleation rates as a function of key controlled parameters have been carried out in a 26.1 m<sup>3</sup> stainless steel cylinder chamber at the European Organization for Nuclear Research (CERN), in the framework of the CLOUD experiment (Cosmics Leaving OUtdoor Droplets) (e.g., Kirkby et al., 2011; Kürten et al., 2016; Dunne et al., 2016). Some of these experiments were conducted at the temperature in the range of those in the stratosphere (Kirkby et al., 2011; Dunne et al., 2016) which are used in this study to evaluate nucleation schemes under stratospheric conditions.

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### 193 **3. Results**

# 3.1 H<sub>2</sub>SO<sub>4</sub>-H<sub>2</sub>O binary homogeneous nucleation (BHN) and binary ion-mediated nucleation (BIMN) under stratospheric conditions

Nucleation is one of the microphysical processes influencing particle size distributions in the 196 stratosphere (Turco et al., 1982) The CLOUD measurements under a wide range of well-controlled 197 conditions (Kirkby et al., 2011; Dunne et al., 2016) provide a unique set of data to evaluate the 198 nucleation theories. Yu et al. (2020) compared nucleation rates calculated based on a number of 199 commonly used aerosol nucleation parameterizations with the CLOUD measurements. Here we 200 specifically examine the comparison under stratospheric conditions where temperature is below ~ 201 230 K. Since ammonia concentrations in the stratosphere are generally negligible, we focus on 202 binary nucleation in the present study. The contribution of organics to particle formation, growth, 203 and compositions in the upper troposphere and LMS has been investigated in several studies (Kupc 204 et al., 2020; Murphy et al., 2021; Williamson et al., 2021). Because of the lack of information with 205 206 regard to the low volatile gaseous organic species, the possible role of organics in new particle formation in LWS is not considered in the present study. 207





Figure 1. Comparison of nucleation rates based on three different schemes with CLOUD measurements within the low temperature range (T = 205-223 K) as that in the stratosphere for (a) binary homogeneous nucleation (no ionization) and (b) ion nucleation (at the presence of ionization rates 2.51 – 110 ion-pairs cm<sup>-3</sup>s<sup>-1</sup>). The different nucleation schemes shown are: BHN of Vehkamäki et al. (2002) (BHN\_V2002), BHN of Yu et al. (2020) (BHN\_Y2020), and BIMN

of Yu et al. (2020) (BIMN\_Y2020). For comparison, under binary condition of (a), BIMN rates at Q = 20 ion-pairs cm<sup>-3</sup>s<sup>-1</sup> are given while under binary ion nucleation condition of (b), BHN rates are also given. Values of H<sub>2</sub>SO<sub>4</sub> vapor concentration ([H<sub>2</sub>SO<sub>4</sub>]) range from 10<sup>6</sup> to 3×10<sup>7</sup> cm<sup>-3</sup> and are separated into four groups in the plots (Circles: 10<sup>6</sup> – 5×10<sup>6</sup> cm<sup>-3</sup>; triangles: 5×10<sup>6</sup> – 10<sup>7</sup> cm<sup>-3</sup>; Squares: 10<sup>7</sup> – 1.5×10<sup>7</sup> cm<sup>-3</sup>; Diamonds: 1.5×10<sup>7</sup> – 3×10<sup>7</sup> cm<sup>-3</sup>).

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221 Figure 1 compares nucleation rates based on the following three different schemes with CLOUD measurements under stratospheric temperature range (T = 205-223 K): BHN of 222 223 Vehkamäki et al. (2002) (BHN V2002), BHN of Yu et al. (2020) (BHN Y2020), and BIMN of Yu et al. (2020) (BIMN Y2020). BHN V2002 and BHN\_Y2020 differ in term of thermodynamic 224 data and nucleation approach used (Yu et al., 2020). To show the relative importance of 225 homogeneous versus ion nucleation, BIMN rates at Q = 20 ion-pairs cm<sup>-3</sup>s<sup>-1</sup> were given under 226 binary homogeneous condition in Fig. 1a and BHN rates were also given under binary ion 227 nucleation condition in Fig. 1b. Nucleation rates based on BHN V2002 are consistently 3-5 orders 228 of magnitude higher than those observed under H<sub>2</sub>SO<sub>4</sub>-H<sub>2</sub>O binary nucleation conditions without 229 (Fig. 1a) and with (Fig. 1b) the effect of ionizations, while those based on BHN Y2020 and 230 BIMN Y2020 are close to the observed values. It should be noted that similar to the CLOUD 231 measurements with the effect of ionization, BHN rates are included in the BIMN rates (Yu et al., 232 2018) and the difference between BIMN and BHN rates indicates the contribution of ion mediated 233 or induced nucleation. Under the conditions of Fig. 1a, assuming ionization rate of 20 ion-pairs 234 235  $cm^{-3}s^{-1}$  (within the range of its typical value in the stratosphere) the BIMN rates are about one order of magnitude higher than BHN rates when the nucleation rates are below  $\sim 5 \text{ cm}^{-3}\text{s}^{-1}$  but 236 close to BHN rates when nucleation rates are above ~ 5 cm<sup>-3</sup>s<sup>-1</sup>. Similar difference between 237 BHN Y2020 and BIMN Y2020 can also be seen in Fig. 1b, indicating the importance of ion 238 nucleation at relatively lower nucleation rates (mostly associated with relatively lower [H<sub>2</sub>SO<sub>4</sub>]) 239 and dominance of homogeneous nucleation at higher nucleation rates (associated with larger 240 [H<sub>2</sub>SO<sub>4</sub>]). As we show next, [H<sub>2</sub>SO<sub>4</sub>] in the background stratosphere is generally quite low and 241 242 thus ion nucleation dominates but BHN can become important in the SO<sub>2</sub> plumes injected into the stratosphere. 243

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### **3.2** Nucleation rates and particle number concentrations in the stratosphere

Figure 2 shows the zonal mean SO<sub>2</sub> emission (SO<sub>2</sub> emit), particle number emitted by aviation 246 (PN aviation), temperature (T), relative humidity (RH), ionization rate (Q), and [H<sub>2</sub>SO<sub>4</sub>] averaged 247 during the two-year period (06/2016–05/2018) covering ATom 1-4. To focus on lower stratosphere 248 (LS), only the values of these variables in the stratosphere (grid boxes with more than 50% time 249 above tropopause) are shown. The SO<sub>2</sub> emissions include all sources including volcanos and 250 aviation. During this period, there was one relatively strong volcanic event, the Bezymianny 251 252 volcano (55.98°N, 160.59°E), on December 20, 2017 that injected  $5 \times 10^6$  kg S into an altitude of ~ 253 14–18 km (Carn et al., 2015). Aviation emission is generally limited to below ~ 12.5 km altitude. Based on MERRA2 meteorology data, which is used to drive GEOS-Chem, almost all of grid 254 boxes at 12 km are under the troppause in the tropics (30°N-30°S), most of grid boxes at 12 km 255 in the high latitude regions (60°N–90°N, 60°S–90°S) are above tropopause, and some fractions of 256 257 grid boxes at 12 km in the middle latitude regions (30°N–60°N, 30°S–60°S) are above tropopause.

As can be seen from Fig. 2b, some of aviation emissions in the middle and high latitude regions

are in the LMS, and the amount emitted into NH LMS is much higher (by several orders of 259 magnitude) than that in SH. The temperature in the LS ranges from 190-225K, with the lowest 260 261 value in the region just above tropical tropopause (Fig. 2c). RH in LS has highest values near tropopause but drops quickly with increasing altitude, from  $\sim 30-50\%$  near tropopause to  $\sim 0.1-1\%$ 262 at  $\sim 25$  km in the tropical and middle latitudes (Fig. 2d). The spatial variations of T and RH have 263 important effects on nucleation in LS. The cosmic ray induced ionization rate in LS has large 264 latitudinal gradient, ranging from ~40–100 ion-pair std. cm<sup>-3</sup>s<sup>-1</sup> (here "std. cm<sup>-3</sup>" refers to per cubic 265 centimeter at standard temperature and pressure, 273 K and 1013 hPa respectively) in the tropics 266 to 100–400 ion-pair std.  $cm^{-3}s^{-1}$  in middle and high latitude region (Fig. 2e). The high ionization 267 rates may have important implication for particle microphysics in LS, which will also be discussed 268 in Section 3.3. H<sub>2</sub>SO<sub>4</sub> is the most important aerosol precursor in LS and its concentration depends 269 on SO<sub>2</sub> concentrations and oxidation, condensation sink, and its vapor pressure that depends on T270 and RH. The annual mean [H<sub>2</sub>SO<sub>4</sub>] (Fig. 2f) has large spatial variations, ranging from a minimum 271 of ~  $1-2 \times 10^5$  std. cm<sup>-3</sup> at altitudes of ~ 12-15 km in polar regions to ~  $4-20 \times 10^5$  std. cm<sup>-3</sup> close 272 to the tropopause. From  $\sim 18-25$  km (well above the ATom measurement altitude). [H<sub>2</sub>SO<sub>4</sub>] 273 increases with altitude, mainly due to the increasing H<sub>2</sub>SO<sub>4</sub> vapor pressure associated with vertical 274 changes of T (Fig. 2c) and RH (Fig. 2d). 275





Figure 2. Zonal mean SO<sub>2</sub>\_emit, PN\_Emit, *T*, RH, *Q*, and [H<sub>2</sub>SO<sub>4</sub>] averaged during the two-year period (06/2016-05/2018) covering ATom 1-4. To focus on the lower stratosphere, only the values of these variables in grid boxes with more than 50% time above tropopause and below 25 km are shown.

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To demonstrate the effect of nucleation schemes on simulated aerosol properties, we compare 284 in Fig. 3 zonal mean and vertical profiles of nucleation rates (J) and number concentrations of 285 condensation nuclei larger than 3 nm (CN3) simulated based on the three nucleation schemes: 286 BHN V2002, BHN Y2020, and BIMN Y2020. In all three schemes, the aviation emissions of 287 both SO<sub>2</sub> (Fig. 2a) and particle numbers (Fig. 2b) are the same. The model simulations indicate 288 289 that NPF occurs in the lower stratosphere but is mostly confined to LMS except in the area of volcano injection (for example, above ~ 14 km around ~  $52^{\circ}$ N). There exist large differences in 290 the nucleation rates predicted by the three schemes (noting the logarithmic color scale), with 291 BHN V2002 rates generally 1-4 orders of magnitude higher while BIMN Y2020 rates ~ one 292 order of magnitude higher than those based on BHN-Y2020. The difference between 293 BIMN Y2020 and BHN V2002 rates are smaller in the LMS over tropics (30°S-30°N) where 294 temperature is the lowest (see Fig. 2c). The magnitudes of differences are consistent with 295 comparisons with CLOUD measurements (Fig. 1). The difference in nucleation rates leads to 296 substantial difference in CN3 in LMS, with those based on BHN V2002 a factor 2–5 higher than 297 those based on BHN Y2020 in LMS. LMS CN3 based on BIMN Y2020 is about 50% higher than 298 that of BHN Y2020. Compared to the difference in nucleation rates, the differences in CN3 is 299 much smaller. This is expected because on one hand only a small fraction of nucleated particles 300 survive the coagulation scavenging and grow beyond 3 nm, and on the other hand direct emission 301 of particle numbers from aviation (Fig. 2b; treated as direct emission but most of these are actually 302 nucleated on chemi-ions in the exhaust plume shortly after emission) (Brock et al., 2000) and 303 transport provide substantial amount of CN3 even without nucleation. Nevertheless, nucleation is 304 still significant enough to affect the CN3. It is interesting to note that CN3 based on BIMN Y2020 305 is higher at altitudes  $> \sim 22$  km (Fig. 3h), which is associated with higher nucleation rates based 306 on BIMN Y2020 than those based on BHN V2002 and BHN Y2020 within the altitude range of 307 35-55 km. Another interesting point is that there is a much smaller vertical gradient in 308 BHN V2002 nucleation rates in the tropical region (30S-30N) within ~ 17-20 km (see Figs. 3a 309 and 3d), likely a result of different dependences of nucleation rates based on different schemes on 310 T, RH, and [H<sub>2</sub>SO<sub>4</sub>] which have large vertical variations (see Fig. 2). It can be seen from Fig. 3 311 that the simulations based on three nucleation schemes all show large hemispheric difference in 312 particle number concentrations (by a factor of ~3-6) in LMS at middle and high latitudes, 313 consistent with the ATom measurements (Williamson et al., 2021). Our sensitivity study (by 314 turning off aviation emission, not shown, to be reported in a separate study) indicates this large 315 hemispheric difference is largely caused by aviation emissions, confirming the analysis of 316 Williamson et al. (2021). 317



Figure 3. Model simulated zonal mean and vertical profiles of nucleation rates (J; upper panels) and number concentrations of particles larger than 3 nm (CN3; lower panels) in the stratosphere during the two-year period covering ATom 1-4 (06/2016- 05/2018), based on three nucleation schemes (a&e: BHN\_V2002, b&f: BHN\_Y2020, and c&g: BIMN\_Y2020). The vertical profiles in (d) and (h) are averaged for three latitude zones (90S-30S, 30S-30N, and 30N-90N). The values for those grids with at least 50% of time above the tropopause are shown.

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While it is difficult to observe nucleation rates in the stratosphere, the measurement of freshly 327 nucleated nanoparticles can be used to constrain nucleation schemes. Figure 4a compares the 328 model simulated CN3 (all particles with diameter larger than 3 nm, with the upper size limit of 12 329 µm corresponding to the size of last model bin) based on the three nucleation schemes at altitudes 330 of around 12 km in SH middle and high altitudes during four seasons with the corresponding ATom 331 1-4 observations. As an example, Figures 4b-d show the model simulated horizontal distributions 332 of CN3 at 12 km altitude during ATom 4 with the values and locations of ATom4 CN3 data 333 overlaid. We choose SH for comparison, as it represents the background stratosphere with 334 minimum influence of anthropogenic emissions (i.e., aviation) (Fig. 2b), to avoid the uncertainty 335 associated with aviation emissions. In Figure 4, the model results are two-month average 336 corresponding to the flight months of each ATom campaign while the measurement data points 337 shown are those sampled within the altitudes range of 11.5–12.5 km, in the stratosphere 338 (ozone>250 ppbv and RH<10%, following the same stratosphere definitions as in Murphy et al. 339 (2021) and Williamson et al. (2021)), and averaged to a 4°x5° gridbox for comparison with 340 modeled results. The impact of nucleation scheme on CN3 can be clearly seen: BHN V2002 341 overpredicted CN3 by a factor of 2-4, BHN Y2020 slightly underpredicted CN3, and 342 BIMN Y2020 slightly overpredicted CN3. The larger vertical spread in CN3 from BHN V2002 343 is caused by the large CN3 latitude gradient associated with higher nucleation near tropopause 344

345 (Fig. 3). The comparisons above show that the ATom measurements provide a good constraint on

our understanding of the processes controlling CN3 in the LMS at mid-high latitudes.



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Figure 4. CN3 at altitudes of around 12 km in SH middle and high latitudes: (a) Model simulated
versus observed during ATom 1-4 (Circles: ATom1; Triangles: ATom2; Sqaures: ATom3;
Diamonds: ATom4); (b-d) model simulated horizontal distributions corresponding to ATom 4
based on three different nucleation schemes (BHN\_V2002, BHN\_Y2020, and BIMN\_Y2020),
with the values and locations of ATom 4 CN3 measurements shown in the circles.

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### 355 **3.3 PNSDs in the stratosphere**

Figure 5 shows the model simulated evolution of PNSDs at an altitude of 12 km over a site in 356 SH (70°S, 60°W) during the two-year ATom period based on the three different nucleation 357 schemes. The PNSDs shown in Fig. 5 are averaged into four different seasons corresponding to 358 the months of ATom 1-4 field campaigns and are presented in Fig. 6 for comparison with the 359 observed mean PNSDs in SH LMS (Williamsons et al., 2021). It should be noted that modeled 360 PNSDs in Fig. 6 are two-month average at one fixed site at an altitude of 12 km (in the region 361 where many of SH LMS measurements were taken, see Fig. 4) while the observed ones are 362 averaged over all SH LMS air mass sampled during the corresponding ATom campaign. While 363 the comparison in Fig. 6 is not exactly coterminous, it allows us to make quantitative comparisons 364 of modeled and observed PNSDs. To take into account the variations in both model and observed 365 PNSDs, standard deviations are shown as error bars in the measured and modeled curves based on 366 BIMN Y2020. 367



Figure 5. Model simulated evolution of PNSDs at a site in SH (70S, 60 W) at altitude of 12 km
based on three nucleation schemes (BHN\_2002, BHN\_Y2020, and BIMN\_Y2020).

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Figure 6 shows that PNSDs measured in the background LMS have multiple modes: a 373 nucleation mode (NuclM:  $\leq 10$  nm), an Aitken mode (AitkenM: ~10 - 80 nm), and two 374 accumulation modes (AccuM1:  $\sim 80 - 250$  nm and AccuM2:  $\sim 250 - 700$  nm). It should be noted 375 that these modes are not the same size limits as those presented in the public ATom dataset. As 376 377 shown in Figures 5 and 6, the model based on all three nucleation schemes generally captures the AitkenM and AccuM1 and the existence of a minimum in PNSDs around 80 nm, although there 378 exist differences. Interestingly, the relative height (or peak values of dN/dlogD<sub>p</sub>) of AitkenM and 379 AccuM1 has strong seasonal variations. The model captures a relatively higher AitkenM in SH 380 Summer and Fall and a higher AccuM1 in SH Spring. The model simulated PNSDs also agree 381

well with the measurements in term of the size-dependent normalized standard deviation ( $\sigma_N$ , i.e., 382 383 the standard deviation  $\sigma$  divided by the mean): relatively smaller  $\sigma_N$  for AccuM1 and larger size 384 part of AitkenM and much larger  $\sigma_N$  for NuclM, smaller size part of AitkenM, and AccuM2. While the larger  $\sigma_N$  for NuclM is understandable because of NPF, it is surprising for AccuM2. The 385 AccuM2 particles have relatively long lifetime and are expected to be well-mixed (and thus have 386 387 small variations) in LS. The transport of AccuM2 particles from UT may contribute to the larger variations. Murphy et al. (2021) showed the chemical signature of this transported mode, and here 388 we show that the variation in the size distribution may also contain information about the mixing 389 of UT particles into LMS. Compared to the observations, the model simulated AccuM2  $\sigma_N$  are 390 larger in SH Winter and Spring but are smaller in SH Summer and Fall. The possible reasons for 391 the large variations of AccuM2 in LMS and the differences between model simulations and 392 measurements remain to be studied. 393



Figure 6. Model simulated seasonal mean PNSDs at a site in SH (70°S, 60°W) at altitude of 12 km based on three nucleation schemes and comparisons with the corresponding ATom measurements (a: SH Spring 09–10/2017, b: SH Summer 01–02/2017, c: SH Fall 04–05/2018, and

d: SH Winter 06–07/2016). To take into account the variations in both model and observed PNSDs,
standard deviations are shown as error bars in the measured and modeled curves based on
BIMN\_Y2020. Three vertical dashed lines at 10 nm, 80 nm, and 250 nm are drawn in (a) to guide
the eye to the four modes discussed in the text.

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404 The large impacts of nucleation schemes on PNSDs, especially those smaller than 100 nm, can 405 be seen in Fig. 6. The formation rates and concentrations of nucleation mode particles are very high based on BHN V2002 (peak dNdlogDp values reaching well above 10<sup>3</sup> std. cm<sup>-3</sup>), negligible 406 based on BHN Y2020 (dNdlogDp values for particles <10 nm are generally below 1 std. cm<sup>-3</sup>), 407 408 and moderate based on BIMN Y2020. When compared to the observed values, the number 409 concentrations of particles within 3-10 nm based on BHN V2002 are 1-2 orders of magnitude too high but those based on BHN Y2020 are 1-2 orders of magnitudes too low, while those based 410 on BIMN Y2020 are of the same order of magnitude. The impact of nucleation schemes on NuclM 411 propagates into the AitkenM and AccuM1, with BHN Y2020 giving the lowest number 412 concentrations while BHN V2002 gives the highest AitkenM and BIMN Y2020 gives the highest 413 AccuM1. It should be noted that, while the line of BHN-Y2020 is lower than that of BIMN and 414 BHN-V2002 for particles of smaller sizes ( $\sim 300$  nm), it is slightly higher for larger particles ( $\geq \sim$ 415 300 nm). This is consistent with the competition of sulfuric acid gas between pre-existing larger 416 particles and nucleated smaller particles. It is interesting to note that AccuM1 based on 417 BIMN Y2020 is higher than that based on BHN V2002 although BHN V2002 predicts higher 418 NuclM and AitkenM, indicating a non-linear interaction among nucleation, growth, and 419 coagulation. The competition between nucleation and condensation for available sulfuric acid gas 420 has been shown to be important for SAI studies (Laakso et al., 2022). 421

422 There exist a number of differences in the simulated and observed PNSDs. Firstly, measurements indicate a slight increase of dNdlogD<sub>p</sub> with decreasing sizes for particles < 10 nm 423 but the simulated PNSDs based on BIMN Y2020, the scheme mostly consistent with CLOUD 424 measurements and predicting NuclM concentrations closest to those observed, decreases with 425 decreasing sizes for particles < 10 nm. The possible reasons of the difference remain to be 426 investigated but probably are associated with uncertainty in nucleation rates and size-dependent 427 growth rates of freshly nucleated particles, and/or the fact that ATom observations are bias towards 428 daytime. In addition, the small number of particles in this mode is likely within the uncertainty in 429 the ATom measurements (about 7% of the total number of particles), so that this measured mode 430 may not be significant. Secondly, the model appears to overpredict the smaller size part (~10-40 431 nm) of AitkenM although it is close to the larger part of the mode (~40-80 nm). The overprediction 432 may be a result of the underestimated growth rates or coagulation scavenging rates of these 433 particles or overpredicted growth rates of NuclM particles. Thirdly, the model generally 434 overpredicts the mean mode sizes of AccuM1 and underpredicts the concentrations of the mode 435 except in SH Spring. The nucleation schemes have observable effects on the concentrations and 436 mean sizes of AccuM1 and overall the simulations based on BIMN Y2020 are in stronger 437 agreement with measurements. Finally, the observed PNSDs show a clear AccuM2 in all seasons 438 439 except Fall but such a mode cannot be clearly seen in the model simulated PNSDs, indicating that the model underpredicts the concentrations of AccuM2 mode particles. AccuM2 particles are 440

within the size range with most efficient scattering of solar radiation and thus are important forSAI. It is therefore necessary to identify the sources of this difference and to improve the model.

443 As pointed out earlier, the comparison in Fig. 6 does not exactly match in terms of time and location, which likely contributes to some of the differences shown in Fig. 6. Some of the 444 differences can also be caused by the uncertainties in the model in term of emissions, transport, 445 chemistry, aerosol microphysics, and deposition. Nevertheless, some of these differences, 446 especially the shape of PNSDs (AccuM2, NuclM, etc.), are unlikely to be fully accounted for by 447 the above-mentioned possible mismatch or model uncertainties and thus may indicate that some 448 fundamental processes are not represented in the model. One possible cause of the differences is 449 that the transport of organic-sulfate particles from UT (Murphy et al., 2014, 2021) is not properly 450 simulated by the model. Based on size-resolved particle composition measurements, Murphy et al. 451 (2021) showed that the LMS accumulation mode particles (diameter  $\sim 0.1$  and 1.0  $\mu$ m) have at 452 453 least two modes: the larger mode consists mostly of sulfuric acid particles produced in the stratosphere, and the smaller mode consists mostly of organic-sulfate particles transported from 454 the troposphere. Murphy et al. (2014) showed that the fraction of organic-sulfate aerosols above 455 tropopause decreases quickly with altitudes. While the organic-sulfate mode aerosols from UT 456 may contribute to the bi-modal structure of accumulation mode particles in the LMS observed 457 during ATom, it is unlikely to contribute to the bi-modal structure of particles larger than ~200 nm 458 observed at altitude above ~20 km both in the background and in volcano perturbed stratosphere 459 (Deshler et al., 2013, 2019; also see Fig. 7). Here, we suggest that the role of charges on 460 coagulation and growth of particles in the stratosphere could be another process causing the bi-461 modal of large particles in the stratosphere. 462

As shown Fig. 2e, ionization rates are high in LS, ranging from ~ 40–100 ion-pair std. cm<sup>-3</sup>s<sup>-1</sup>. 463 Due to their low number concentrations (~100–1000 std. cm<sup>-3</sup>) but long lifetime, particles in the 464 stratosphere are expected to be in charge equilibrium. Figure 7 shows mean particle number size 465 distribution (PNSD) and particle volume size distribution (PVSD) observed during ATom 1-4 in 466 SH LMS and measured within 20-25 km altitude over Lararie WY in 1992, and fraction of particles 467 carrying *n* charges based on the modified Boltzmann equilibrium equation (Clement and Harrison, 468 1992). The bi-modal structure of accumulation mode particles can be clearly seen in both 469 background and volcano perturbed stratosphere. It should be noted that while the smaller mode 470 generally dominates the number concentrations, the larger mode dominates mass concentrations. 471 Under equilibrium more particles are charged (i.e.,  $1-f_0 > 50\%$ ) than neutral (f<sub>0</sub>) for particles with 472 473 diameter larger than ~80 nm and a significant fraction (> 25%) of particles larger than 300 nm carrying multiple charges. While the equilibrium charge fraction is small for NuclM particles ( $\lesssim$ 474 10 nm), this fraction can be much larger when nucleation on ions occurs, which is consistent with 475 476 the observed overcharging of freshly nucleated particles (Laakso et al. 2007; Yu and Turco, 2008). 477 Particle coagulation rates are influenced by forces exerted between colliding particles, including van der Waals and electrostatic forces, which can modify the effective collision cross section and 478 479 sticking coefficient. The van der Waals force has been shown to be important in the stratosphere (English et al., 2011, 2012) and has been considered in the simulations shown above. The effects 480 481 of charges on coagulation and implications for PNSDs in the stratosphere have not yet been studied 482 (to our knowledge). Since coagulation is a dominant process for the growth of accumulation mode particles in the stratosphere, we hypothesize that differential coagulation rates for neutral and 483

charged particles in accumulation modes can potentially act as a physical process separating the 484 485 modeled single accumulation mode (Fig. 6) into two modes (AccuM1 and AccuM2) as observed. 486 Further research is needed to test this hypothesis. In addition to affecting coagulation, charge on small particles can also enhance the growth rate due to ion-dipole interactions of condensing 487 molecules with charged particles (Nadykto and Yu, 2005). This enhancement is expected to be 488 489 stronger in the stratosphere because of lower temperature (Nadykto and Yu, 2005). Beside these, 490 Svensmark et al. (2020) showed that the condensation of ion clusters can enhance particle growth rates. How much the enhanced coagulation and growth rates of charged particles may shape 491 PNSDs and modes in the stratosphere remains to be investigated. 492



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**Figure 7.** ATom 1-4 mean observed particle number size distribution (PNSD, or dN/dlogDp) and particle volume size distribution (PVSD, or dV/dlogDp) in SH LMS, balloon-borne measured mean PNSD and PVSD within 20-25 km altitude over Lararie WY in 1992, and fraction of particles carrying n (n = 0, 1, 2, 3, 4, and 5) charges based on the modified Boltzmann equilibrium equation (Clement and Harrison, 1992). Note that  $f_n$  with  $n \ge 1$  including both positive and negative charges, i.e., for example, half of  $f_1$  carrying one negative charge while the other half positive.

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### 502 4. Summary and Discussions

Interest in stratospheric aerosols has been increasing in recent years, due to the ongoing discussion about the plausibility, potential benefits and risks of offsetting climate change through stratospheric aerosol injection (SAI) to buy time for reduction of CO<sub>2</sub> in the atmosphere. Recent studies indicate the dependence of SAI radiative efficacy (Dai et al., 2018) on the particle size distribution (NASEM, 2021) and thus it is critical to improve foundational understanding and model representation of aerosol microphysics processes controlling the evolution of stratospheric
 aerosols, both under background conditions and perturbed scenarios. While formation and growth
 of particles in the troposphere have been extensively studied in the past two decades, very limited
 efforts have been devoted to understanding these in the stratosphere.

In the present study we use both CLOUD laboratory measurements taken under very low 512 513 stratospheric temperatures and ATom in-situ observations of particle number size distributions (PNSD) down to 3 nm to constrain nucleation schemes and model-simulated particle size 514 distributions in the lowermost stratosphere (LMS). We show that the binary homogenous 515 nucleation scheme used in most of the existing SAI modeling studies overpredicts the nucleation 516 rates by 3-4 orders of magnitude (when compared to CLOUD data), leading to significant 517 overprediction of particle number concentrations in the background stratosphere (by a factor of 2-518 4 in SH LMS, compared to ATom data). Based on a recently developed kinetic nucleation model 519 520 which provides rates of both ion-mediated nucleation (IMN) and BHN at low temperatures in good agreement with CLOUD measurements, both BHN and IMN occur in the stratosphere but IMN 521 rates are generally more than one order of magnitude higher than BHN rates and thus dominate 522 nucleation in the background stratosphere. 523

In the SH LMS that has minimal influences from anthropogenic emissions, our analysis shows 524 that ATom-measured PNSDs generally have four apparent modes: a nucleation mode (NuclM: ≤ 525 10 nm), which may not be statistically significant, an Aitken mode (AitkenM: ~10-80 nm), and 526 527 two accumulation modes (AccuM1: ~ 80-250 nm and AccuM2: ~ 250-700 nm). The model generally captures the AitkenM and AccuM1 and the existence of a minimum in PNSDs at  $\sim 80$ 528 nm, although there are differences. The model captures a relatively higher AitkenM in SH Summer 529 and Fall and a higher AccuM1 in SH Spring. The model simulated PNSDs also agree well with 530 the measurements in term of the size-dependent standard deviations: relatively smaller standard 531 deviations for AccuM1 and larger size part of AitkenM and much larger standard deviations for 532 NuclM, smaller size part of AitkenM, and AccuM2. 533

A detailed comparison indicates the existence of a third PNSD mode peaking around 300–400 534 nm in the ATom measurements that are not captured by the model. Compared to the observations, 535 the model-simulated AccuM2 standard deviations are larger in SH Winter and Spring but are 536 smaller in SH Summer and Fall. In addition, the model overpredicts the number concentration of 537 particles in the size range of 10-50 nm. These differences may indicate that, in addition to 538 nucleation, the model may be missing some fundamental microphysical processes of stratospheric 539 aerosols. Our analysis shows that, in the stratosphere, more particles are charged (positive + 540 negative) than neutral for particles with diameter larger than ~80 nm and a significant fraction (> 541 25%) of particles larger than 300 nm carrying multiple charges. We propose that the role of charges 542 on coagulation and growth of particles in the stratosphere, where ionization rates are high and 543 544 particles have very long lifetime, is likely one of such processes. Considering the importance of accurate particle size distributions (especially the accumulation mode particles) for projecting 545 546 realistic radiative forcing response to stratospheric aerosols, it is essential to understand and incorporate such potentially important processes in model simulations of future changes in the 547 548 stratosphere. It should be noted that the ATom measurement period does not have a high 549 stratospheric aerosol loading (i.e., no major volcano eruptions). It remains to be investigated if previous assessments of volcanic aerosol microphysics missed something important. We expect 550

the uncertainties in the nucleation schemes and unknown cause of the bi-modal structure of accumulation mode particles will affect particle optical properties and surface area and thus radiative forcing or chemistry. The present work highlights the importance of advancing scientific understanding of processes controlling properties of stratospheric particles as well as further development, improvement, and validation of models for reducing uncertainties of SAI simulations (e.g., Golja et al., 2021, Sun et al., 2022).

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- **558 Conflict of interest**: The authors declare that they have no conflict of interest.
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Data availability. The GEOS-Chem model is available to the public at https://geos-564 chem.seas.harvard.edu/. in analysis available 565 Simulation output this is at https://doi.org/10.5281/zenodo.6909944. The ATom dataset is published as Wofsy et al., (2021, 566 https://doi.org/10.3334/ORNLDAAC/1925) and is also available 567 at https://espoarchive.nasa.gov/archive/browse/atom (last access: June 2022). 568

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