

1 **Measurement report: Changes in light absorption and molecular**  
2 **composition of water-soluble humic-like substances during a winter**  
3 **haze bloom-decay process in Guangzhou, China**

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17

18 **Abstract**

19 Water-soluble humic-like substances (HULIS) absorb light in near-UV and visible  
20 wavelengths and exert significant influence on the atmospheric environment and climate.  
21 However, knowledge on HULIS evolution during haze bloom-decay process is limited.  
22 Herein, PM<sub>2.5</sub> samples were obtained during a winter haze event in Guangzhou, China,  
23 and light absorption and molecular composition of HULIS were investigated by UV-vis  
24 spectrophotometry and ultrahigh-resolution mass spectrometry. Compared with HULIS  
25 in clean days, the absorption coefficients (Abs<sub>365</sub>) of HULIS in haze days were  
26 significantly higher but the mass absorption efficiencies (MAE<sub>365</sub>) were relatively lower,  
27 suggesting diverse and dynamic absorption properties of HULIS during haze episodes.  
28 The CHO and CHON compounds were the most abundant components in HULIS,  
29 followed by CHOS, CHONS, and CHN. Haze HULIS presented comparatively higher  
30 molecular weight, lower aromaticity index (AI<sub>mod</sub>), and higher O/C<sub>w</sub>, O/N<sub>w</sub>, and O/S<sub>w</sub>  
31 ratios, indicating that HULIS fractions undergo relatively higher oxidation during haze  
32 days than clean days. Moreover, CHON and CHO compounds with high AI<sub>mod</sub> were the  
33 major potential chromophores in HULIS and significantly contributed to HULIS light  
34 absorption. It's worth noting that the proportions of these chromophores were decreased  
35 during haze event, mainly owing to their higher oxidation during haze episode. Besides,  
36 accumulated contribution of organic compounds emitted from vehicles and formed from  
37 reactions of bio-VOCs also diluted light-absorbing compounds in haze HULIS. These  
38 findings help to understand HULIS evolution during haze bloom-decay process in the  
39 subtropic region of China.

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## 41 **1. Introduction**

42 Water-soluble humic-like substances (HULIS), belonging to a class of highly  
43 complex organic compounds with physical/chemical properties similar to humic  
44 substances in natural aquatic/soil environments, which constitute 30%–70% of water-  
45 soluble organic compounds in ambient aerosols and are responsible for > 70% of light  
46 absorption (at 365 nm) in water-soluble brown carbon (BrC) (Graber and Rudich, 2006;  
47 Laskin et al., 2015; Huang et al., 2018). They are thought to be comprised of aromatic  
48 structures containing aliphatic side chains and oxygenated functional groups such as  
49 hydroxyl, carboxyl, nitrate, and organosulfate groups (Lin et al., 2012; Song et al., 2018;  
50 Zeng et al., 2020). HULIS are ubiquitously identified in atmospheric aerosols, fog, cloud,  
51 and rain water, and have been demonstrated to play significant effects on both  
52 atmospheric environment and climate (Bianco et al., 2018; Wu et al., 2018; Zhan et al.,  
53 2022). In addition, HULIS exert adverse health effects because they can enhance the  
54 oxidative potential of organic aerosols (Ma et al., 2019; Cao et al., 2021).

55 In recent years, severe particulate pollution (i.e., haze events) frequently occur in  
56 some developing country such as China, which has drawn extensive public and scientific  
57 concerns (An et al., 2019; Zhang et al., 2020). According to An et al. (2019),  
58 contributions of organic aerosols, including primary organic aerosols and secondary  
59 organic aerosols (SOA), are significant for severe haze events; in particular, the  
60 contribution of SOA in China is expected to continuously increase because of stronger  
61 chemical reactions in the atmosphere. HULIS are an important component in organic  
62 aerosols, which originate from a variety of primary emissions (e.g., biomass burning  
63 (BB), coal combustion, off-road engine emission) (Fan et al., 2016; Cui et al., 2019; Tang

64 et al., 2020) and secondary chemical oxidation of biogenic and anthropogenic volatile  
65 organic compounds (VOCs) (Yu et al., 2016; Tomaz et al., 2018) and soot (Fan et al.,  
66 2020). During the haze episode, a number of chemical processes occur in aqueous phase  
67 (Wong et al., 2017; Wu et al., 2018) and gas phase (Sumlin et al., 2017), which lead to  
68 significant changes in chemical composition and light absorption properties of HULIS.  
69 For instance, recent studies on oxidation of BB-derived BrC have indicated that although  
70 both enhancement and bleaching of BrC occur during aging, bleaching of BrC becomes  
71 dominant over a long period (Fan et al., 2020; Wong et al., 2017; Ni et al., 2021).  
72 However, multiphase reaction between carbonyl and amine has demonstrated rapid  
73 formation of light-absorbing organic compounds (Kampf et al., 2016). Nevertheless, it  
74 should be noted that these results were mainly obtained from laboratory experiments and  
75 may not reflect the complex evolution behavior of BrC in atmospheric environment.

76 High concentrations of HULIS have been determined during typical haze episodes  
77 in northern, eastern, and southern China (Fan et al., 2016; Zhang et al., 2020; Wang et al.,  
78 2020), and have been demonstrated to significantly influence atmospheric visibility,  
79 environment, and photochemical process. Guangzhou is the biggest city in the Pearl  
80 River Delta (PRD), one of the most developed regions in China, and is located in the  
81 subtropical zone with a population of over 18 million people (Yu et al., 2017). Although  
82 a remarkable decline in atmospheric particulate matter (PM<sub>2.5</sub>) pollution has been  
83 observed in recent years owing to strict regulatory controls, O<sub>3</sub> and VOCs still remain at  
84 higher levels and severe haze pollution caused by fine particulate matter frequently occur  
85 in winter (An et al., 2019; Li et al., 2019; Yang et al., 2022). Several studies have  
86 investigated the optical, chemical, and molecular properties of HULIS in the PRD region

87 (Lin et al., 2010, 2012; Fan et al., 2016; Liu et al., 2018; Jiang et al., 2020, 2021). For  
88 example, the studies on the temporal variations of water-soluble HULIS in Guangzhou  
89 indicated that HULIS had higher concentrations and mass absorption efficiencies  
90 ( $MAE_{365}$ ) in the winter, which were attributed to the increasing contribution of BB and  
91 secondary nitrate formation in the winter monsoon period (Fan et al., 2016; Jiang et al.,  
92 2020, 2021). In addition, the molecular composition of HULIS (and BrC) in the PRD  
93 region were also investigated and demonstrated that the levels of unsaturated and  
94 aromatic structures are the important factor influencing their light absorption properties  
95 (Jiang et al., 2020). However, detailed information regarding the evolution of light  
96 absorption and molecular composition of HULIS during haze events is still scarce.

97       Recently, ultrahigh-resolution Fourier transform ion cyclotron resonance mass  
98 spectrometry (FT-ICR MS) coupled with electrospray ionization (ESI) sources has been  
99 frequently employed to investigate the molecular characteristics of HULIS in ambient  
100 aerosols (Song et al., 2018, 2022; Tang et al., 2020; Zeng et al., 2021). Owing to its  
101 extremely high mass resolution and accuracy, this technique allows further exploration of  
102 the evolution of HULIS during haze event. The present study performed comprehensive  
103 characterization of HULIS in  $PM_{2.5}$  collected during a haze event in Guangzhou, China.  
104 The abundances and light absorption properties of HULIS were first measured, and  
105 carbonaceous fractions, water-soluble ions, and levoglucosan (Lev) were determined.  
106 Subsequently, four HULIS samples collected during different haze stages were analyzed  
107 using FT-ICR MS operated in both ESI<sup>-</sup> and ESI<sup>+</sup> modes. To the best of our knowledge,  
108 the present study is the first to apply a combination of optical properties and molecular  
109 characterization by FT-ICR MS to investigate HULIS in a haze event in the subtropical

110 zone of China. The results obtained provide novel insights into the evolution of HULIS  
111 during haze event, and are important for predicting the environmental and climatic effects  
112 of HULIS in South China.

## 113 **2. Material and Methods**

### 114 **2.1. Aerosol sampling**

115 The PM<sub>2.5</sub> samples were collected on the campus of Guangzhou Institute of  
116 Geochemistry, Chinese Academy of Sciences, Guangzhou, China (23.14N, 113.35E),  
117 which is an academic and residential region. Traffic emissions and residential activities  
118 are the potential pollution sources in the sampling area. The 24-h PM<sub>2.5</sub> sampling was  
119 conducted using a high-volume sampler (Tianhong Intelligent Instrument Plant, Wuhan,  
120 China, with a flow rate of 1.0 m<sup>3</sup> min<sup>-1</sup>) during 7 to 30 January of 2018, and a total of 24  
121 samples were collected on the prebaked quartz filters (20.3 × 25.4 cm<sup>2</sup>, Whatman,  
122 Maidstone, UK). Field blank samples were collected by keeping a blank filter in the  
123 sampler without pumping air. Before sampling, the filters were wrapped in aluminum foil  
124 and prebaked at 450°C for 6 h to remove carbonaceous impurities. Before and after  
125 sampling, the filters were weighed at 25°C and 50% RH on a microbalance (Sartorius  
126 Model BP210D). The PM<sub>2.5</sub> concentrations were determined by weighing the filters  
127 before and after collection. Finally, all filter samples were stored in a refrigerator at -20 °C  
128 until analysis. The mass accuracy achieved was < 2% based on triplicate analyses of filter  
129 samples. Meteorological data (<http://www.wunderground.com/history/airport/ZGGG>),  
130 including wind speed, temperature, relative humidity, and concentrations of SO<sub>2</sub>, O<sub>3</sub>, and  
131 NO<sub>2</sub>, for the sampling days are presented in Figure 1 and Table S1.

## 132 **2.2. Isolation of HULIS**

133 HULIS were isolated using a water extraction and solid-phase extraction (SPE)  
134 procedure as described previously (Zou et al., 2020). This method has been used in most  
135 previous studies because of its easy operation and high reliability and reproducibility and  
136 low limit of detection (Fan et al., 2002), therefore, it was also used in this study. Briefly,  
137 portions of the PM<sub>2.5</sub> samples (100 cm<sup>2</sup>) were ultrasonically extracted with 50 mL of  
138 ultrapure water for 30 min. The extracts were filtered through a 0.22- $\mu$ m PTFE syringe  
139 filter to remove the suspended insoluble particles. About 50 mL of water extracts were  
140 obtained from each sample, of which 20 mL was used for the isolation and analysis of  
141 HULIS, 20 mL for analysis of water-soluble organic carbon (WSOC), and the remained  
142 extracts for the analysis of inorganic ions, respectively. Then, the 20 mL water extracts  
143 were adjusted to pH of 2 with HCl, and loaded on a preconditioned SPE cartridge (Oasis  
144 HLB, 200 mg/6 mL, Waters, USA). The hydrophilic fraction (i.e., inorganic ions, high-  
145 polar organic acids, etc) was removed with ultrapure water, whereas the relatively  
146 hydrophobic HULIS fraction was retained and eluted with 2% (v/v) ammonia/methanol.  
147 Finally, HULIS solution was evaporated to dryness with a gentle N<sub>2</sub> stream and  
148 redissolved with ultrapure water for the analysis.

149 It is noted that the HULIS here is the hydrophobic portion of water-soluble organic  
150 matter, which can be isolated with different types of SPE columns (e.g., HLB, C-18,  
151 DEAE, XAD-8, and PPL) (Fan et al., 2012, 2013; Lin et al., 2012; Zou et al., 2020; Jiang  
152 et al., 2020). Although each resin type has its special chemical properties, the  
153 hydrophobic HULIS isolated with different sorbents were similar in chemical, molecular  
154 properties based on previous studies (Fan et al., 2012, 2013; Zou et al., 2020). Therefore,

155 for better comparison with other studies, the hydrophobic fractions isolated by SPE  
156 methods were all termed as HULIS in the present paper.

### 157 **2.3. Light absorption analysis**

158 The absorption spectra of the WSOC and HULIS fractions were measured by a UV-  
159 vis spectrophotometer (UV-2600, Shimadzu) between 200 to 700 nm. Each spectrum was  
160 corrected for the filter blanks. The light absorption coefficients, absorption Ångström  
161 exponent (AAE) and mass absorption efficiency ( $MAE_{\lambda}$ ) were calculated and the detailed  
162 methods are presented in the Supporting Information (SI).

### 163 **2.4. Chemical analysis**

164 For FT-ICR MS analysis, the HULIS samples were isolated from  $PM_{2.5}$  collected  
165 during four periods: before haze days (clean-I days, 7–12 January), haze bloom days  
166 (haze-I days, 13–18 January), haze decay days (haze-II days, 19–24 January), and after  
167 haze days (clean-II days, 25–30 January). A filter punch (18 cm in diameter) was taken  
168 from every sample, and all the six samples in each period was combined for the isolation  
169 of HULIS fractions. The obtained HULIS samples were measured with an ESI FT-ICR  
170 MS (Bruker Daltonik GmbH, Bremen, Germany) equipped with a 9.4 T refrigerated  
171 actively shielded superconducting magnet. The system was operated in both ESI<sup>-</sup> and  
172 ESI<sup>+</sup> modes. The scan range was set to  $m/z$  from 100 to 1000, with a typical mass-  
173 resolving power  $>450,000$  at  $m/z$  319 with  $<0.2$  ppm absolute mass error. The mass  
174 spectra were calibrated externally with arginine clusters and internally recalibrated with  
175 typical  $O_5$ -class species peaks in DataAnalysis 4.4 (Bruker Daltonics). Due to the  
176 inherent differences in the ionization mechanisms between ESI<sup>-</sup> and ESI<sup>+</sup> modes, the



177 data detected by the two ionization modes can provide complementary information on the  
178 molecular composition of atmospheric HULIS (Lin et al., 2012; Lin et al., 2018). The  
179 details of data analysis are provided in the SI.

180 The amounts of organic carbon (OC) and elemental carbon (EC) were determined by  
181 a OC/EC analyzer (Sunset Laboratory Inc., USA) (Mo et al., 2018). The concentrations  
182 of WSOC and HULIS were determined by a TOC analyzer (Shimadzu TOC\_VCPH,  
183 Kyoto, Japan). The water-soluble inorganic species ( $\text{NO}_3^-$ ,  $\text{SO}_4^{2-}$ ,  $\text{Cl}^-$ ,  $\text{NH}_4^+$ ,  $\text{K}^+$ ,  $\text{Na}^+$ ,  
184  $\text{Ca}^{2+}$ ,  $\text{Mg}^{2+}$ ,) were measured with a Dionex ICS-900 ion chromatography system (Thermo  
185 Fisher Scientific, USA) as described previously (Huang et al., 2018). The concentrations  
186 of Lev were analyzed with a gas chromatography–MS after derivatization with BSTFA  
187 and pyridine at 70 °C for 3 h (Huang et al., 2018). Detailed information regarding these  
188 measurements is provided in the SI.

### 189 **3. Results and Discussion**

#### 190 **3.1. Abundance and chemical composition of PM<sub>2.5</sub>**

191 Figure 1 shows the meteorological conditions, PM<sub>2.5</sub> concentration, and  
192 concentrations of major chemical constituents, including carbon fractions and water-  
193 soluble inorganic ions in PM<sub>2.5</sub> samples obtained during a haze bloom-decay process.  
194 Based on the variation in PM<sub>2.5</sub> concentration, these samples were categorized into four  
195 groups: clean-I days (before haze, 14–24  $\mu\text{g m}^{-3}$ ), haze-I days (haze bloom, 45–114  $\mu\text{g}$   
196  $\text{m}^{-3}$ ), haze-II days (haze decay, 58–115  $\mu\text{g m}^{-3}$ ), and clean-II days (after haze, 9–35  $\mu\text{g}$   
197  $\text{m}^{-3}$ ). As indicated in Table S1 and Figure 1, the PM<sub>2.5</sub> concentrations increased from 18  
198  $\pm 3.3 \mu\text{g m}^{-3}$  in clean-I days to  $82 \pm 26$  and  $84 \pm 22 \mu\text{g m}^{-3}$  in haze-I and haze-II days,

199 respectively, and then decreased to  $21 \pm 10 \mu\text{g m}^{-3}$  in clean-II days. This finding  
200 obviously indicated that the average  $\text{PM}_{2.5}$  concentrations during the examined haze  
201 episode are higher than the second-grade national ambient air quality standard in China  
202 ( $75 \mu\text{g m}^{-3}$ , 24 h), whereas those during clean days are lower than the first-grade national  
203 ambient air quality standard in China ( $35 \mu\text{g m}^{-3}$ , 24 h). However, the average  $\text{PM}_{2.5}$   
204 concentrations during the haze event are lower than those in the cities in winter haze,  
205 including Shenyang ( $108 \mu\text{g m}^{-3}$ ) (Zhang et al., 2020), and Nanjing ( $123 \pm 28.5 \mu\text{g m}^{-3}$ )  
206 (Li et al., 2020), Beijing ( $158 \mu\text{g m}^{-3}$ ) and Xi'an ( $345 \mu\text{g m}^{-3}$ ) (Zhang et al., 2018).

207 As shown in Table S1, the average concentrations of OC and EC were 2.2–15 and  
208  $0.36\text{--}2.7 \mu\text{gC m}^{-3}$  in the four stages, respectively, implying that the distinct changes in  
209 OC and EC were higher during haze episodes than those in clear days. During the entire  
210 study period, WSOC concentration ranged from 0.5 to  $12.5 \mu\text{gC m}^{-3}$  ( $4.3 \pm 1.2 \mu\text{gC m}^{-3}$ ),  
211 which contributed to 53%–57% of OC in  $\text{PM}_{2.5}$ . The HULIS concentration noted in the  
212 present study ranged from 0.15 to  $6.1 \mu\text{gC m}^{-3}$  ( $2.2 \pm 1.9 \mu\text{gC m}^{-3}$ ), which was  
213 comparable to those observed in the PRD region, such as Hong Kong ( $2.38 \pm 1.62 \mu\text{gC}$   
214  $\text{m}^{-3}$ ) (Ma et al., 2019), Guangzhou ( $2.4 \pm 1.6 \mu\text{gC m}^{-3}$ ) (Fan et al., 2016), and Heshan  
215 ( $2.08 \pm 1.16 \mu\text{gC m}^{-3}$ ) (Jiang et al., 2020), but lower than those in northern cities of  
216 China, such as Xi'an ( $12.4 \pm 6.5 \mu\text{gC m}^{-3}$ ) (Huang et al., 2020), Beijing ( $3.79 \pm 3.03$   
217  $\mu\text{gC m}^{-3}$ ) (Mo et al., 2018), and Lanzhou ( $4.7 \mu\text{gC m}^{-3}$ ) (Tan et al., 2016). As shown in  
218 Figure 1, HULIS also exhibited obvious variations during the entire sampling period. The  
219 average HULIS concentration was  $0.46 \pm 0.22 \mu\text{gC m}^{-3}$  in clean-I days, which sharply  
220 increased to  $4.5 \pm 1.2 \mu\text{gC m}^{-3}$  in haze-I days, then decreased to  $3.1 \pm 1.2 \mu\text{gC m}^{-3}$  in  
221 haze-II days, and rapidly declined to  $0.75 \pm 0.52 \mu\text{gC m}^{-3}$  in clean-II days. This result

222 was consistent with the changing trend of WSOC, OC, and EC. In addition, the  
223 HULIS/WSOC ratios were about  $0.50 \pm 0.13$  in the  $PM_{2.5}$  samples, which are in broad  
224 agreement with other studies showing that HULIS is the major fraction of WSOC (Fan et  
225 al., 2016; Ma et al., 2019; Jiang et al., 2020).

226 As illustrated in Figure 1, obvious variations in chemical compositions were also  
227 observed in these  $PM_{2.5}$  samples. Secondary inorganic aerosols (SIA) (i.e.,  $SO_4^{2-}$ ,  $NO_3^-$ ,  
228 and  $NH_4^+$ ), OC, and EC exhibited a similar variation during the entire study period, and  
229 their contents sharply increased from 10 January in clean-I days to 13–18 January in  
230 haze-I days, then slowly decreased in haze-II days, and finally reached lower levels in  
231 clean-II days. It must be noted that the increasing rate of EC was similar to that of SIA in  
232 haze-I days, indicating that direct emissions and atmospheric reactions may play similar  
233 roles in  $PM_{2.5}$  increase during this haze bloom period. As indicated in Figure 1f, the  
234 highest values of  $NO_3^-/SO_4^{2-}$  were observed in haze-I days, implying the important  
235 influence of traffic exhausts in the haze bloom period (Mo et al., 2018). In addition, the  
236 high  $NO_2$  and  $O_3$  concentrations and the stable meteorological condition with high  
237 temperature also led to the outburst of fine particulate pollution in this period. During  
238 haze-II days, the SIA and OM contents in  $PM_{2.5}$  slowly decreased, whereas the  
239 concentrations of  $Na^+$ ,  $Cl^-$ , and unidentified materials in  $PM_{2.5}$  increased (Figure 1e,h),  
240 suggesting that local contribution weakened and regional contribution via sea salt became  
241 more important (Jiang et al., 2021). This phenomenon was also observed to be consistent  
242 with the changes in the pollutant sources transported by air masses. As indicated in  
243 Figure S1, the  $PM_{2.5}$  samples in haze-II days included some contributors transported from

244 coastal area of eastern Guangdong Province and Fujian Province, and the  $PM_{2.5}$  are likely  
245 to be enriched with sea salt materials and mineral dusts.

### 246 **3.2. Light absorption**

247 The light absorption properties of WSOC and HULIS (Figure 1d, i, j and Table S2)  
248 exhibited obvious temporal variations during the sampling period. The AAE values of  
249 WSOC and HULIS ranged from 4.1 to 6.4 and 5.6 to 6.6, respectively. The AAE values  
250 for HULIS were obviously higher than those for WSOC in the same sample (Figure 1i),  
251 indicating that light absorption of HULIS is more wavelength-dependent than that of  
252 WSOC. Similar results were also observed in previous studies (Park et al., 2018; Jiang et  
253 al., 2020; Cao et al., 2021). These differences could be attributed to the differences in  
254 chemical composition of chromophores in WSOC and HULIS. As shown in Table S2, the  
255  $E_{250}/E_{365}$  values of HULIS (5.3–5.6) are higher than that (4.4–5.1) of WSOC, suggesting  
256 that the light-absorbing species in HULIS may have relatively lower aromaticity and/or  
257 lower molecular weight than those in WSOC (Chen et al., 2016; Li and Hur, 2017).  
258 Therefore, the HULIS fractions exhibit relatively higher absorption at UV and short  
259 visible wavelengths and relatively lower absorption at long visible wavelengths, which  
260 results in relatively higher AAE values. Moreover, the AAE values of HULIS did not  
261 present significant variation during the entire haze process, which could be related to the  
262 evolution of HULIS chromophores at different stages (Jiang et al., 2020; Huang et al.,  
263 2018; Deng et al., 2022). At first, the enhanced oxidation of aromatic species in haze  
264 days could lead to the bleaching or degradation of chromophores (a detailed explanation  
265 was provided in Section 3.3), thus a less wavelength dependence (Forrister et al., 2015;  
266 Zhan et al., 2022). In contrast, the outburst of secondary organic aerosols and the

267 photolysis of organic aerosols in haze days both resulted in the higher AAE values (Saleh  
268 et al., 2013; Dasari et al., 2019). Consequently, the different trends in AAE were  
269 counterbalanced during the haze days, which resulted in no significant AAE variations  
270 for HULIS in the entire sampling process. This is also consistent with the trends of the  
271  $E_{250}/E_{365}$  ratios of HULIS in the four stages (Table S2).

272 Light absorption at 365 nm ( $Abs_{365}$ ) for WSOC and HULIS were  $2.5 \pm 2.0$  and  $1.8 \pm$   
273  $1.6 \text{ M m}^{-1}$ , respectively (Table S2). HULIS contributed to about 72% of light absorption  
274 coefficients by WSOC, implying that they enriched the major light-absorbing  
275 components in WSOC. As shown in Figure 1d, the  $Abs_{365}$  values for HULIS presented  
276 obvious temporal variations. The  $Abs_{365,HULIS}$  value was  $0.55 \pm 0.06 \text{ M m}^{-1}$  in clean-I  
277 days, which first increased to  $3.4 \pm 1.5 \text{ M m}^{-1}$  in haze-I days and then slowly decreased to  
278  $2.6 \pm 0.85 \text{ M m}^{-1}$  in haze-II days, and finally rapidly declined to  $0.64 \pm 0.32 \text{ M m}^{-1}$  in  
279 clean-II days. This result was similar to the variations in the mass concentration of  
280 HULIS. Furthermore, the  $Abs_{365}$  values for HULIS in Guangzhou were found to be  
281 higher than those observed in southeastern Tibetan Plateau ( $0.38\text{--}1.0 \text{ M m}^{-1}$ ) (Zhu et al.,  
282 2018), but obviously lower than those in Xi'an ( $7.6\text{--}36 \text{ M m}^{-1}$ ) (Shen et al., 2017) and  
283 Beijing, ( $3.7\text{--}10.1 \text{ M m}^{-1}$ ) (Du et al., 2014).

284 In general,  $MAE_{365}$  value can be used to assess the light absorption capacity of target  
285 organic compounds (Li et al., 2019). As shown in Figure 1j and Table S2, the average  
286  $MAE_{365}$  value for WSOC was  $1.0 \pm 0.21 \text{ m}^2 \text{ gC}^{-1}$  ( $0.68\text{--}1.3 \text{ m}^2 \text{ gC}^{-1}$ ), nearly same to  
287  $1.1 \pm 0.27 \text{ m}^2 \text{ gC}^{-1}$  ( $0.77\text{--}1.8 \text{ m}^2 \text{ gC}^{-1}$ ) for HULIS, during the entire sampling period.  
288 Moreover, the  $MAE_{365}$  values for HULIS measured in the present study were noted to be  
289 dropped in the ranges of those determined in Beijing ( $1.43 \pm 0.33 \text{ m}^2 \text{ g C}^{-1}$ ) (Mo et al.,

290 2018), Xi'an ( $0.91\text{--}1.85\text{m}^2 \text{ g C}^{-1}$ ) (Yuan et al., 2021), and Hong Kong ( $1.84 \pm 0.77 \text{ m}^2$   
291  $\text{gC}^{-1}$ ) (Ma et al., 2019). The average  $\text{MAE}_{365}$  values for HULIS exhibited some temporal  
292 variations. The  $\text{MAE}_{365}$  values for HULIS were  $0.91\pm 0.03$  and  $0.95 \pm 0.11 \text{ m}^2 \text{ gC}^{-1}$  in  
293 haze-I and haze-II days, respectively, which were lower than those ( $1.3 \pm 0.22$  and  $1.3 \pm$   
294  $0.27 \text{ m}^2 \text{ gC}^{-1}$ , respectively) observed in clean-I and clean-II days, suggesting that HULIS  
295 have a relatively weaker light absorption capability in haze days. This finding is  
296 consistent with the results reported by Zhang et al. (2017), who found that the  $\text{MAE}_{365}$   
297 values in the heating or non-heating seasons during hazy days were lower than those in  
298 clean days. These differences in  $\text{MAE}_{365}$  values may potentially contribute to the  
299 enhanced oxidation reaction that was derived by the increased  $\text{O}_3$  levels and high  
300 temperature and relative humidity (RH) during haze days (Figure 1). This oxidation  
301 process would lead the chromophores containing C=C unsaturated bond to be severely  
302 degraded (Wang et al., 2017a; Zhang et al., 2017). Besides, an increase in additional  
303 sources for HULIS in the study area, such as weaker or non-light-absorbing compounds  
304 formed by atmospheric oxidation, could also result in weaker light absorption of HULIS  
305 during the haze episode (Liu et al., 2018).

### 306 **3.3. Molecular evolution of HULIS during the haze process**

307 For an in-depth understanding of the variation in HULIS at molecular level during  
308 the haze process, the four HULIS samples collected in different stages of the haze  
309 process were analyzed by ESI FT-ICR MS in both negative and positive modes. As  
310 shown in Figure 2, thousands of peaks were detected in the mass range between  $m/z$  100  
311 and  $m/z$  700, with the high intensity ions noted within  $m/z$  150–400. It is obvious that  
312 some organic compounds with stronger arbitrary abundance were labeled, and their

313 formulas, double bond equivalent (DBE), modified aromaticity index ( $AI_{mod}$ ), and  
314 potential sources were listed in Table S3. Compounds a ( $C_7H_7NO_3$ ) and b ( $C_8H_6O_4$ ), both  
315 have high DBE values, which might be assigned to aromatics such as methylnitrophenol  
316 and phthalic acid, whereas compound d ( $C_8H_{18}O_4S$ ) with low DBE value and high O/S  
317 ratio was probably aliphatic organosulfate. According to previous studies, these organic  
318 molecules might be derived from BB and diesel fuel and thereby these results suggested  
319 that both BB and vehicular emissions are important sources of BrC in ambient aerosols  
320 (Mohr et al., 2013; Riva et al., 2015; Blair et al., 2017). Furthermore, compound e  
321 ( $C_{10}H_{17}NO_7S$ ) and compound f ( $C_{10}H_{18}N_2O_{11}S$ ) in Table S3 were found to be identical to  
322 the oxidation products of monoterpenes, suggest that biogenic sources could contribute to  
323 the formation of HULIS (Surratt et al., 2008; Wang et al., 2019). Thus, HULIS could be  
324 affected by multiple sources during the haze process, possibly including BB, biogenic  
325 sources, and anthropogenic emissions.

326 The identified formulas could be divided into seven compound categories, namely,  
327  $CHO^-$ ,  $CHON^-$ ,  $CHOS^-$ , and  $CHONS^-$  detected in  $ESI^-$  mode and  $CHO^+$ ,  $CHN^+$ , and  
328  $CHON^+$  detected in  $ESI^+$  mode. As illustrated in Figure 2, the  $CHO$  compounds were the  
329 most abundant group in all the HULIS, accounting for 43%–50% and 51%–57% of the  
330 overall compounds detected in the  $ESI^-$  and  $ESI^+$  modes, respectively. It must be noted  
331 that relatively lower contents of  $CHO^-$  were detected during the haze episode (haze-I and  
332 haze-II days) and  $CHO^+$  molecules in haze-I HULIS. The  $CHON$  compounds were the  
333 second most abundant group in all the HULIS. As shown in Figure 2, the relative content  
334 of  $CHON^-$  was 23% in clean-I days, which slightly increased to 24%–25% in haze  
335 episode, and then decreased to 23% in clean-II days. In contrast, the relative content of

336 CHON<sup>+</sup> compounds was 41% in clean-I days, which increased to 45% in haze-I days,  
337 then fell to 42% in haze-II days and 41% in clean-II days. Both CHOS<sup>-</sup> and CHONS<sup>-</sup>  
338 compounds were identified in all the four HULIS, accounting for 19%–22% and 8%–11%  
339 of the total identified compounds, respectively. The CHN<sup>+</sup> compounds were the least  
340 abundant (1.3%–3.6%) in the four HULIS samples, and were relatively higher during the  
341 haze episode, especially in haze-I days.

342 Tables S4 and S5 show the relative abundance weighted elemental ratios, molecular  
343 weight (MW), DBE,  $AI_{mod}$ , and carbon oxidation state ( $OS_C$ ) for the identified  
344 compounds in HULIS. The  $MW_w$  values for HULIS determined in the ESI<sup>-</sup> mode in  
345 haze-I and haze-II days were 302 and 283, respectively, which were higher than those in  
346 clean-I and clean-II days (266 and 264, respectively). Similar variation was also observed  
347 for  $MW_w$  for HULIS detected in ESI<sup>+</sup> mode (Table S5). These results clearly indicated  
348 that more higher MW compounds constituted HULIS obtained during the haze episode.  
349 Furthermore, the molecular properties of HULIS in different stages of haze process also  
350 exhibited some observable differences. As shown in Table S4, the HULIS samples in  
351 haze episode detected by ESI<sup>-</sup> mode presented relatively lower  $AI_{mod,w}$  values and  
352 relatively higher  $O/C_w$ ,  $O/N_w$ , and  $O/S_w$  ratios than those in clean days, indicating that  
353 haze HULIS exhibited relatively lower aromaticity and higher oxidation degree than  
354 clean HULIS. These differences can be attributed to the enhanced oxidation degradation  
355 of aromatic compounds (e.g., phenols, nitroaromatic compounds and polycyclic aromatic  
356 hydrocarbons (PAHs)) during the haze process. In addition, increased contribution from  
357 traffic emission and secondary reactions of bio-VOCs also decreased the aromaticity and  
358 increased the oxidation degree of HULIS (Liu et al., 2016; Tang et al., 2020). These



359 changes in HULIS compounds led to the decrease in their MAE<sub>365</sub> values during the haze  
360 episode, as described above ( Zhong and Jang, 2014; Song et al., 2019).

### 361 **3.3.1. CHO Compounds**

362 The CHO compounds bear O-containing functional groups, and have been  
363 frequently detected in ambient aerosols. As shown in Figure 2, the CHO compounds were  
364 the predominant component in the four HULIS samples, and the MW<sub>w</sub> values for CHO-  
365 and CHO+ compounds were 247–288 and 236–272, respectively, with relatively higher  
366 MW<sub>w</sub> values observed for the CHO group (CHO- and CHO+) in haze HULIS, especially  
367 in haze-I samples. This finding may be related to the stronger oxidation of HULIS during  
368 haze days, because the aqueous oxidation of biomass burning aerosols was found to yield  
369 high MW of organic products (Tomaz et al., 2018; Yu et al., 2016).

370 The OS<sub>C</sub> is often used to describe the degree of oxidation of organic species in the  
371 atmosphere (Kroll et al., 2011; Wang et al., 2017b). Figure 3 shows plots of OS<sub>C</sub> versus  
372 carbon number for the CHO compounds. As indicated in the figure, CHO compounds  
373 exhibited OS<sub>C</sub> from -2 to +1 with up to 40 carbon atoms. Kroll et al. (2011) proposed  
374 that compounds with OS<sub>C</sub> between -0.5 and +1 and < 18 carbon atoms can be attributed  
375 to semi-volatile and low-volatile oxidized organic aerosols (SV-OOA and LV-OOA),  
376 which are mainly formed by complex oxidation reactions in atmosphere. Compounds  
377 with OS<sub>C</sub> between -0.5 and -1.5 and 6–23 carbon atoms are related to primary biomass  
378 burning organic aerosol (BBOA). In addition, compounds with OS<sub>C</sub> between -1 and -2  
379 and ≥18 carbon atoms have been suggested to be hydrocarbon-like organic aerosols  
380 (HOA), which are regarded as primary combustion surrogate (Zhang et al., 2005; Kroll et  
381 al., 2011; Wang et al., 2017b).

382 As illustrated in Figure 3 and Table S6, most of the CHO<sup>-</sup> compounds clustered in  
383 the BBOA region, accounting for 40%–46% of the total CHO<sup>-</sup> compounds, thus  
384 suggesting that BB may be a major contributor to CHO compounds in HULIS. Figure 3  
385 clearly indicates that the majority of aromatic and condensed aromatic compounds  
386 produced signals in the OS<sub>C</sub> region between -0.5 and 1.0 and carbon number of 3–18  
387 (Figure 3), which corresponded to SV-OOA and LV-OOA. The proportions of SV-OOA  
388 and LV-OOA accounted for 23%–28% and 1.9%–2.4% of the total CHO<sup>-</sup> compounds,  
389 respectively, and presented no significant variation. In contrast, the HOA components in  
390 haze-I days showed the highest abundance (18%), which were much higher than those  
391 (3.5%–4.5%) in haze-II, clean-I, and clean-II days. This finding indicated that the  
392 increase in the primary source is associated with fossil fuel combustion such as vehicle  
393 emissions during the haze bloom period (Zhang et al., 2005).

394 As shown in Figure 3, CHO<sup>+</sup> compounds presented lower OS<sub>C</sub> (from -2.0 to 1.0)  
395 than CHO<sup>-</sup> compounds. Most of the CHO<sup>+</sup> compounds occurred in the BBOA region in  
396 all four HULIS samples, making up to 60%–72% of the total CHO<sup>+</sup> compounds, which  
397 again suggesting that BB is an important contributor to CHO compounds in HULIS. The  
398 HOA among CHO<sup>+</sup> compounds showed the same changing trends as those among CHO<sup>-</sup>  
399 compounds, and higher HOA abundance was observed during haze-I days. In addition,  
400 some high AI<sub>mod</sub> values of aromatics were found in the regions A1+ and A2+ (Figure 3),  
401 which implied that the highest AI<sub>mod</sub> values (AI ≥ 0.67) with DBE ≥ 22 were only  
402 detected during the haze days possibly owing to soot-derived materials or oxidized PAHs  
403 (Decesari et al., 2002; Kuang and Shang, 2020). It must be noted that the sampling site in  
404 the present study is influenced by traffic sources, the enhanced oxidation of vehicle-

405 exhausted soot also results in the accumulation of water-soluble high aromatic organic  
406 species (Decesari et al., 2002).

### 407 **3.3.2. CHON Compounds**

408 In the present study, 1379–2217 and 2008–2943 formulas were assigned to CHON  
409 compounds identified in the ESI<sup>-</sup> and ESI<sup>+</sup> spectra, respectively, which accounted for  
410 23%–25% (ESI<sup>-</sup>) and 41%–45% (ESI<sup>+</sup>) of total identified compounds, respectively.  
411 Relatively higher contents of CHON<sup>-</sup> compounds were obviously detected in HULIS  
412 samples obtained during haze-I days, suggesting the occurrence of more N-containing  
413 components in HULIS during haze bloom days. As shown in Tables S4 and S5, the  
414 average MW<sub>w</sub> values for CHON<sup>-</sup> and CHON<sup>+</sup> compounds were 328 and 317 in haze-I  
415 days, respectively, which were slightly higher than those determined in haze-II days and  
416 all higher than those observed in clean-I and clean-II days. Meanwhile, the AI<sub>mod,w</sub> values  
417 for CHON<sup>-</sup> in haze days were 0.31–0.34, which were slightly lower than those in clean  
418 days (0.37 and 0.40). These findings indicated that more high MW CHON compounds  
419 with lower aromatic structures were formed during the haze episode.

420 The O/N<sub>w</sub> ratios for CHON<sup>-</sup> and CHON<sup>+</sup> during haze-I and haze-II days were 5.3–  
421 5.7 and 3.8, respectively, which were higher than those determined during the two clean  
422 periods, confirming that these compounds were highly oxidized during the haze episode  
423 (Tables S4 and S5). In general, compounds with O/N ≥ 3 may indicate oxidized N groups  
424 such as nitro (–NO<sub>2</sub>) or nitrooxy (–ONO<sub>2</sub>), whereas compounds with O/N < 3 may denote  
425 the reduced N compounds (i.e., amines) (Lin et al., 2012; Song et al., 2018). In the  
426 present study, most of the CHON compounds (79%–91% of CHON<sup>-</sup> compounds and  
427 61%–64% of CHON<sup>+</sup> compounds) exhibited O/N ≥ 3, suggesting that high

428 concentrations of nitro compounds or organonitrates were contained in the CHON  
429 compounds. Moreover, these compounds were more abundant in the CHON<sup>-</sup> group  
430 during the haze episode (87%–91%), when compared with those during clean-I and  
431 clean-II days (79%–82%), again implying that CHON<sup>-</sup> compounds undergo relatively  
432 higher oxidization during the haze episode. As indicated in Figure 1, the increase in NO<sub>2</sub>  
433 was consistent with increased production of highly oxidized N-containing organic  
434 compounds (NOCs) during the haze episode, which suggested the significant contribution  
435 of NO<sub>3</sub>-related multigenerational chemistry to organonitrate aerosol formation  
436 (Berkemeier et al., 2016).

437 The majority of aromatics and condensed aromatics produced clear signals in  
438 regions associated with SV-OOA and LV-OOA (Figure 4). BBOA also constituted a  
439 significant proportion (33%–39%) in the CHON<sup>-</sup> group, and a relatively lower BBOA  
440 content was observed in haze-I days. The abundance of HOA was relatively lower,  
441 accounting for 2.3%–7.8% of the total CHON compounds, and the relative abundance of  
442 HOA in haze-I days was much higher than that in haze-II, clean-I, and clean-II days,  
443 suggesting the accumulation of primary fossil fuel combustion during haze-I days.

444 The CHON<sup>+</sup> compounds mainly occurred at the range of  $-2.0 < OS_C < 1.5$ , with  
445 average OS<sub>C</sub> values of around  $-1.0$  for each sample, clearly indicating that CHON<sup>+</sup>  
446 compounds were relatively lower than CHON<sup>-</sup> compounds. Most of the CHON<sup>+</sup>  
447 compounds were detected in the BBOA region, accounting for 60%–76% of the total  
448 CHON<sup>+</sup> compounds. The relative contribution of BBOA in haze-I days was lower than  
449 that in haze-II and clean days. Moreover, a large number of aromatic species were  
450 observed at the region B1<sup>+</sup> (Figure 4), demonstrating that higher aromatic compounds

451 were only detected in haze-I days, which may be related to soot or BC. Similar trend was  
452 also exhibited by CHO+ compounds, indicating the contribution of local combustion  
453 sources (e.g., traffic emission) during haze-I days.

### 454 **3.3.3. CHOS and CHONS Compounds**

455 In this study, 478–696 CHOS compounds and 306–589 CHONS compounds were  
456 identified in ESI– mode (Table S4). Among these S-containing compounds, >86% of the  
457 CHOS compounds had O/S ratios >4, whereas > 89% of the CHONS compounds  
458 presented O/S ratios >7, suggesting that these S-containing compounds were possibly  
459 organosulfates and nitrooxyorganosulfates. As listed in Table S4, the  $AI_{mod,w}$  values for  
460 CHOS and CHONS were about 0.02 and 0.01 in the HULIS fraction, which were much  
461 lower than those for CHO and CHON. Almost 99% of the CHOS and CHONS  
462 compounds in the HULIS fraction had  $AI_{mod}$  values <0.5, while >93% of the CHONS  
463 compounds had  $AI_{mod} = 0$ , indicating that they were mainly comprised of aliphatic and  
464 olefinic organosulfates. These results are consistent with the previous findings that the  
465 major S-containing compounds among organic aerosols in Guangzhou are organosulfates  
466 formed by secondary oxidation reaction of long-chain alkenes/fatty acids with  $SO_2$  (Jiang  
467 et al., 2020), which generally possessed long aliphatic carbon chains and a higher degree  
468 of oxidation. However, these compounds are different from the S-containing compounds  
469 detected during the hazy days in Beijing (Jiang et al., 2016; Mo et al., 2016), which were  
470 determined to be aliphatic organosulfates with low degree of oxidation and higher  
471 amounts of aromatics and PAH-derived organosulfates, having a strong correlation with  
472 anthropogenic emissions.

473 As described earlier, CHOS<sup>-</sup> and CHONS<sup>-</sup> compounds might be related to  
474 organosulfates or nitrooxyorganosulfates, which have been observed to be derived from  
475 atmospheric reactions of bio-VOCs such as  $\alpha$ -pinene, limonene, and isoprene (Huang et  
476 al., 2018; Surratt *et al.*, 2008) and fossil fuel combustion including coal combustion, off-  
477 road engine emissions (Song et al., 2018, 2019; Cui et al., 2019). In the present study, the  
478 relative contents of S-containing compounds (CHOS+CHONS) in the HULIS fraction in  
479 haze days were all higher than those in clean days (Figure 2). Moreover, the CHOS and  
480 CHONS compounds in haze HULIS always have relatively high relatively O/S ratios  
481 than those in clean HULIS. These findings suggested the relatively higher contribution of  
482 SO<sub>2</sub>-related chemical oxidation during the haze event.

#### 483 **3.3.4. CHN Compounds**

484 The N-bases (CHN) are usually identified in ambient aerosols and smokes from BB.  
485 In the present study, 110–165 CHN<sup>+</sup> compounds were identified in ESI<sup>+</sup> mode, with  
486 most of them (>86%) presenting DBE  $\geq$  2, suggesting that they might be nitrile and  
487 amine species (Lin et al., 2012). As shown in Figure 2, the abundances of CHN<sup>+</sup>  
488 compounds were 2.0%–3.6% in the haze days, which were much higher than those noted  
489 in clean days (1.3%–1.4%), indicating higher contribution of CHN<sup>+</sup> compounds to the  
490 HULIS fraction during the haze episode. The MW<sub>w</sub> values for CHN<sup>+</sup> compounds were  
491 204–223, which were lower than those for the other groups (i.e., CHO<sup>+</sup>, CHON<sup>+</sup>) (Table  
492 S5). However, the average AI<sub>mod</sub> values for N-bases (0.37–0.48) detected in the ESI<sup>+</sup>  
493 mode were much higher than those for CHO<sup>+</sup> (0.11–0.12) and CHON<sup>+</sup> (0.20–0.22)  
494 compounds, implying that these reduced CHN<sup>+</sup> compounds exhibited more unsaturated  
495 or aromatic structures.

496 To further understand the molecular distribution of CHN<sup>+</sup> compounds during the  
497 haze process, van Krevelen (VK) diagrams were constructed by plotting the H/C ratio  
498 versus N/C ratio (Figure S2). It was obvious that this plot could separate the compound  
499 classes with different degree of AI. As shown in Figure S2, compounds (denoted in black  
500 color) in the upper region of the VK diagram had one N atom with DBE = 0, indicating  
501 that they are aliphatic amines. It can be noted from Table S7 that the aliphatic group  
502 presented the lowest abundance in all the samples, suggesting that the CHN<sup>+</sup> compounds  
503 possessed comparatively lower aliphatic structures. Olefinic compounds showed the  
504 highest abundance in the four samples, which accounted for 37%–51% of the total CHN<sup>+</sup>  
505 compounds. Importantly, a large proportion of the compounds (>39%) exhibited high  
506 degree of AI (AI > 0.5) (Figure S2 and Table S7), suggesting a large amounts of aromatic  
507 structure and N-heterocyclic ring in HULIS. Moreover, the CHN<sup>+</sup> compounds in haze-I  
508 days presented obviously lower content of aromatic structures than those in haze-II,  
509 clean-I, and clean-II days, signifying the relatively high contribution of fossil fuel  
510 combustion (which generally emits more low-aromatic CHN compounds) during the haze  
511 bloom episode(Song et al., 2022). In addition, the CHN<sup>+</sup> group also constituted a large  
512 proportion of BBOA (Table S6), which indicated the significant contribution of BB.  
513 However, it must be noted that a relatively lower content of BBOA was detected during  
514 haze-I days, which was consistent with the changing trends of CHON<sup>-</sup> or CHON<sup>+</sup>  
515 compounds during the haze episode. These results suggested the relatively lower  
516 contribution of BB during haze-I days, because quiet and stable weather conditions can  
517 prevent regional transport of BB sources during this stage (Wu et al., 2018).

518 **3.4. Factors influencing light absorption and molecular characteristics of HULIS**  
519 **during the haze bloom-decay process**

520 As described earlier, the light absorption properties of HULIS exhibited obvious  
521 variation during the haze bloom-decay process. The average  $\text{Abs}_{365}$  value for HULIS was  
522  $0.55 \pm 0.06 \text{ M m}^{-1}$  in clean-I days, which first increased to  $3.4 \pm 1.5 \text{ M m}^{-1}$  in haze-I days,  
523 then slowly decreased to  $2.6 \pm 0.85 \text{ M m}^{-1}$  in haze-II days, and finally rapidly declined to  
524  $0.64 \pm 0.32 \text{ M m}^{-1}$  in clean-II days. In general, the light absorption of HULIS can be  
525 related to their chemical and molecular properties that are influenced by factors such as  
526 sources, secondary formation, and aging process. The results of principal component  
527 analysis (PCA) obviously showed a positive loading for principal component 1 (PC1),  
528 and the  $\text{Abs}_{365}$  values for HULIS were clustered with EC,  $\text{K}_{\text{bb}}^+$ , Lev,  $\text{NH}_4^+$ , and  $\text{NO}_3^-$   
529 (Figure 5). These results suggested that BB and other sources such as new particle  
530 formation could contribute to light absorption of HULIS (An et al., 2019; Song et al.,  
531 2019). Similarly, the findings of Pearson correlation coefficient analysis revealed that the  
532  $\text{Abs}_{365}$  values for HULIS exhibited significant positive correlations with  $\text{K}_{\text{bb}}^+$  ( $r = 0.728$ ,  
533  $p < 0.01$ ) and Lev ( $r = 0.800$ ,  $p < 0.01$ ) (Table S8). As Lev and  $\text{K}_{\text{bb}}^+$  are generally  
534 considered as tracers derived from BB, these results suggested the significant  
535 contribution of BB to light absorption of HULIS. This observation was also supported by  
536 the abundance of BBOA compounds detected in all the four HULIS samples (Table S6).  
537 The significant positive relationships between the  $\text{Abs}_{365}$  values for HULIS and  
538 secondary ions (i.e.,  $\text{NO}_3^-$  ( $r = 0.702$ ,  $p < 0.01$ ),  $\text{SO}_4^{2-}$  ( $r = 0.554$ ,  $p < 0.05$ ), and  $\text{NH}_4^+$  ( $r$   
539  $= 0.899$ ,  $p < 0.01$ )) indicated the important impact of secondary formation on the light  
540 absorption of HULIS. Besides, the  $\text{Abs}_{365}$  values for HULIS were also strongly correlated



541 with NO<sub>2</sub>, O<sub>3</sub>, and NO<sub>2</sub>, which confirmed the important impact of atmospheric oxidation  
542 reactions on the light absorption of HULIS.

543 It must be noted that MAE<sub>365</sub> is a key parameter signifying the light absorption  
544 ability of HULIS. As listed in Table S2, the MAE<sub>365</sub> values for HULIS varied in different  
545 stages, and were lower in haze days owing to the variation in the chemical and molecular  
546 composition of HULIS during the haze bloom-decay process. Furthermore, the AI<sub>mod</sub>  
547 values for HULIS varied in different stages (Tables S4), and were relatively lower in  
548 haze days, indicating that haze HULIS have comparatively lower degree of conjugation  
549 or aromaticity. This finding suggested that the HULIS compounds may undergo higher  
550 oxidation during the haze episode, causing a decline in chromophores and reduction in  
551 the light absorption capacity of HULIS (Lin et al., 2017). Besides, the accumulated  
552 contribution of organic compounds from vehicle emission and secondary chemical  
553 reactions of bio-VOCs may also dilute light-absorbing compounds in haze HULIS (Tang  
554 et al., 2020; Liu et al., 2016).

555 Lin et al. (2018) reported that potential light-absorbing chromophores can be  
556 determined in the region between DBE = 0.5 × C (linear conjugated polyenes C<sub>x</sub>H<sub>y</sub>C<sub>2</sub>)  
557 and DBE = 0.9 × C (fullerene-like hydrocarbons). In the present study, most of the high-  
558 intensity CHON, CHO, and CHN compounds with high AI values were clustered in  
559 potential BrC chromophore region (Figures S3 and S4), which mainly comprised CHON  
560 (46%–50% in ESI<sup>-</sup> mode and 56%–62% in ESI<sup>+</sup> mode, respectively) and CHO (44%–48%  
561 in ESI<sup>-</sup> mode and 29%–38% in ESI<sup>+</sup> mode, respectively) compounds (Table 1).  
562 Although the contribution of CHN<sup>+</sup> compounds to BrC was relatively lower, the content  
563 of potential chromophores among the total CHN<sup>+</sup> compounds was higher than those in

564 CHON<sup>+</sup> and CHO<sup>+</sup> compounds. Therefore, these three groups of light-absorbing  
565 compounds (i.e., CHON<sup>+</sup>, CHN<sup>+</sup>, and CHO<sup>+</sup> compounds) were further examined. As  
566 shown in Table 1, the Int<sub>C</sub>/Int<sub>BrC</sub> values of CHO<sup>-</sup> (i.e., content of CHO<sup>-</sup> chromophores  
567 in the total chromophores) decreased from 48% to 44% whereas the Int<sub>C</sub>/Int<sub>BrC</sub> values of  
568 CHON<sup>-</sup> increased from 46% to 50% during the haze bloom process. These findings  
569 indicated that more NOCs chromophores were formed during this stage in which higher  
570 NO<sub>2</sub> concentration may be preferred for the formation of N-containing chromophores  
571 such as nitrophenols. However, it must be noted that the proportions of both CHO<sup>-</sup> and  
572 CHON<sup>-</sup> chromophores among the total identified compounds decreased from clean-I to  
573 haze-I days, suggesting the occurrence of stronger photo-bleaching process during the  
574 haze bloom stage (Zeng et al., 2020). Likewise, both CHO<sup>+</sup> and CHON<sup>+</sup> compounds  
575 presented similar variation during the entire study period. In addition, the CHN<sup>+</sup>  
576 compounds also exhibited higher Int<sub>C</sub>/Int<sub>BrC</sub> values during the haze bloom process and  
577 suggesting the accumulated contribution from local combustion process. Furthermore, the  
578 proportion of CHON<sup>+</sup> chromophores in the total CHON<sup>+</sup> compounds increased with the  
579 decreasing content of CHN<sup>+</sup> chromophores, may implying that some aromatic CHN<sup>+</sup>  
580 compounds were transformed to CHON<sup>+</sup> compounds during the aging process.

581

#### 582 **4. Conclusions**

583 This study investigated the evolution of light absorption and molecular properties of  
584 HULIS during a winter haze bloom-decay process, and examined the key factors  
585 affecting the light absorption of HULIS in Guangzhou, China. The results showed that  
586 HULIS exhibited significant variation in light absorption during the haze bloom-decay

587 process. First, higher Abs<sub>365</sub> values were observed in haze days, indicating the presence  
588 of significant amounts of light-absorbing organic compounds during the haze episode.  
589 However, the MAE<sub>365</sub> values for HULIS in haze days were relatively lower than those in  
590 clean days, suggesting the light absorption capabilities of HULIS were weakened during  
591 the haze event. Furthermore, CHON and CHO compounds, exhibiting relatively higher  
592 degree of conjugated structure, were the most abundant groups in all the HULIS samples,  
593 and were also the major contributors to light absorption capacity of HULIS. Importantly,  
594 the molecular properties of HULIS dynamically varied during the entire haze episode.  
595 When compared with HULIS in clean days, those in haze days presented relatively lower  
596 AI<sub>mod</sub> values and higher O/C<sub>w</sub>, O/N<sub>w</sub>, and O/S<sub>w</sub> ratios, suggesting the predominance of  
597 compounds with low aromaticity and higher oxidation in HULIS during haze episode.  
598 These results indicated that HULIS compounds undergo relatively stronger oxidation  
599 during the haze days. Moreover, PCA and Pearson correlation analysis revealed that BB  
600 and secondary chemical formation both contributed to the variation in the light absorption  
601 properties of HULIS. Both primary sources (such as accumulated contribution of organic  
602 compounds formed from local traffic emission) and secondary sources (such as stronger  
603 chemical reactions) led to the rapid increase in HULIS during the haze bloom days.  
604 However, stronger oxidation of HULIS compounds were observed during the haze  
605 episode, and some potential BrC chromophores were degraded. In addition, the chemical  
606 reactions of bio-VOCs such as isoprene also diluted the light-absorbing compounds in  
607 HULIS.

608 Thus, the present study provides novel insights into the light and molecular  
609 evolution of HULIS during haze event, which are important for predicting the

610 environmental and climatic effects of HULIS. However, as this study examined only one  
611 haze bloom-decay process in winter in Guangzhou, the results obtained may be not  
612 adequate for understanding all the haze episodes in South China. Therefore, there is a  
613 need for a comprehensive investigation of haze episode in different seasons and regions  
614 in future.

615

#### 616 **Data availability**

617 The research data are available in the Harvard Dataverse  
618 (<https://doi.org/10.7910/DVN/DYGYQT>, Song, 2022).

619

620 **Author contributions.** J. Song and P. Peng designed the research together. C, Zou, T.  
621 Cao, and M. Li carried out the PM<sub>2.5</sub> sampling experiments. C, Zou and T. Cao extracted  
622 and analyzed the WSOC and HULIS samples. B. Jiang analyzed the HULIS samples by  
623 FT-ICR MS. C. Zou and J. Song wrote the paper. J. Li, X. Ding, Z Yu, and G. Zhang  
624 commented and revised the paper.

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627

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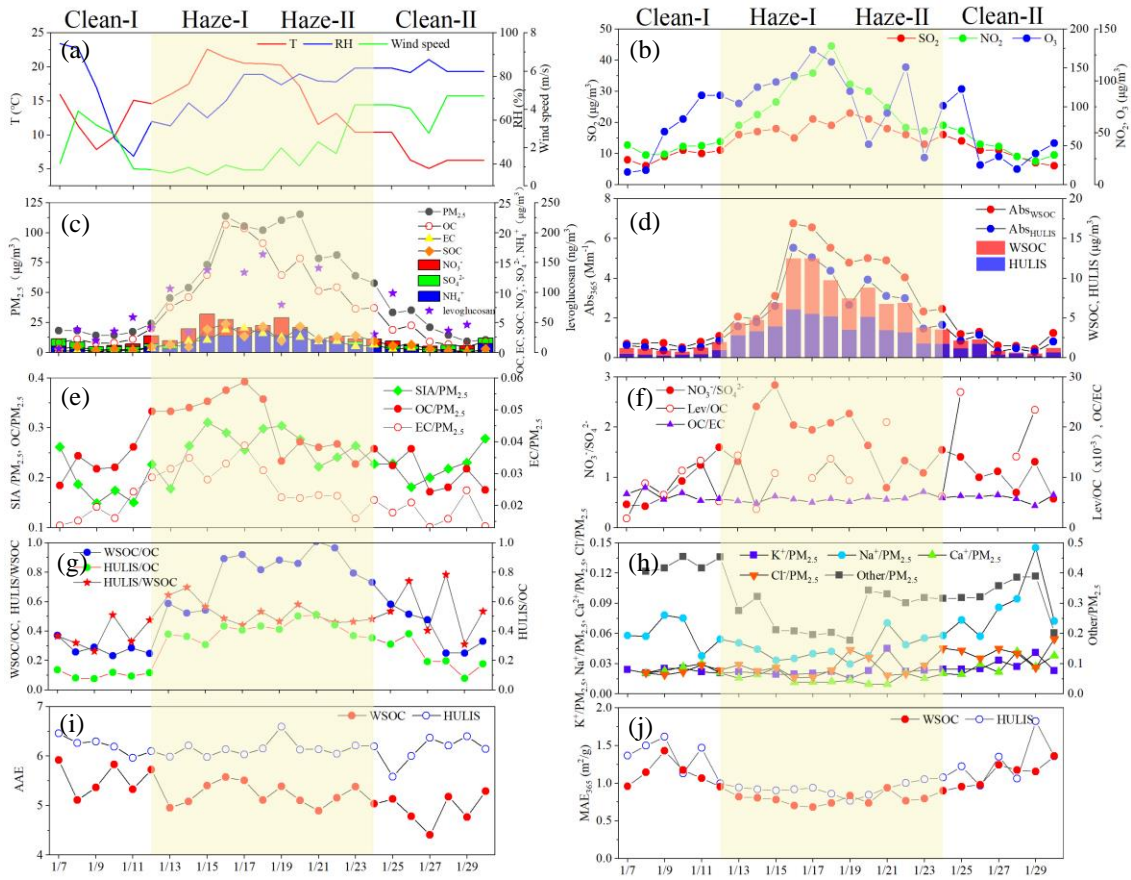
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**Figure 1.** Temporal variation in meteorological parameters, concentrations of chemical composition, and optical properties ( $Abs_{365}$ ,  $MAE_{365}$ , and  $AAE$ ) of water-soluble BrC in the  $PM_{2.5}$  samples.

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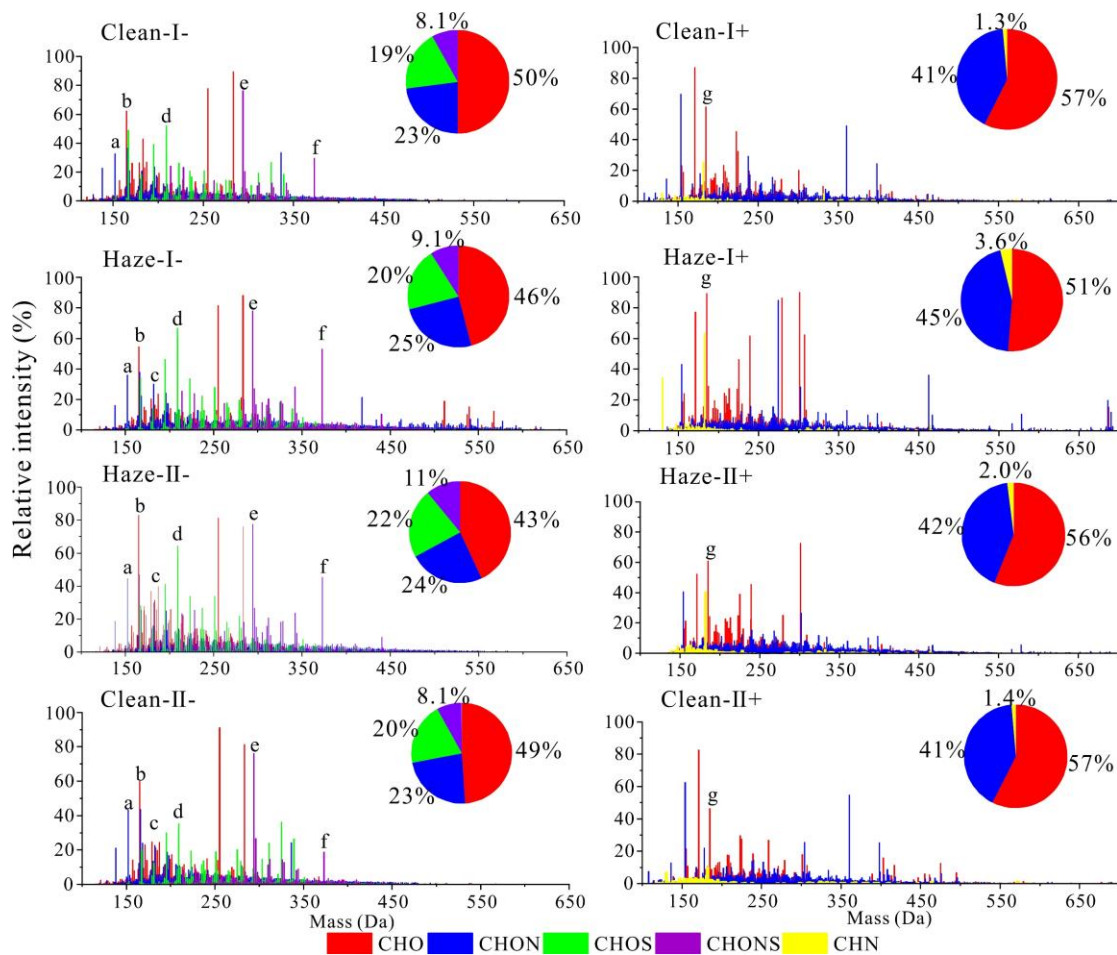
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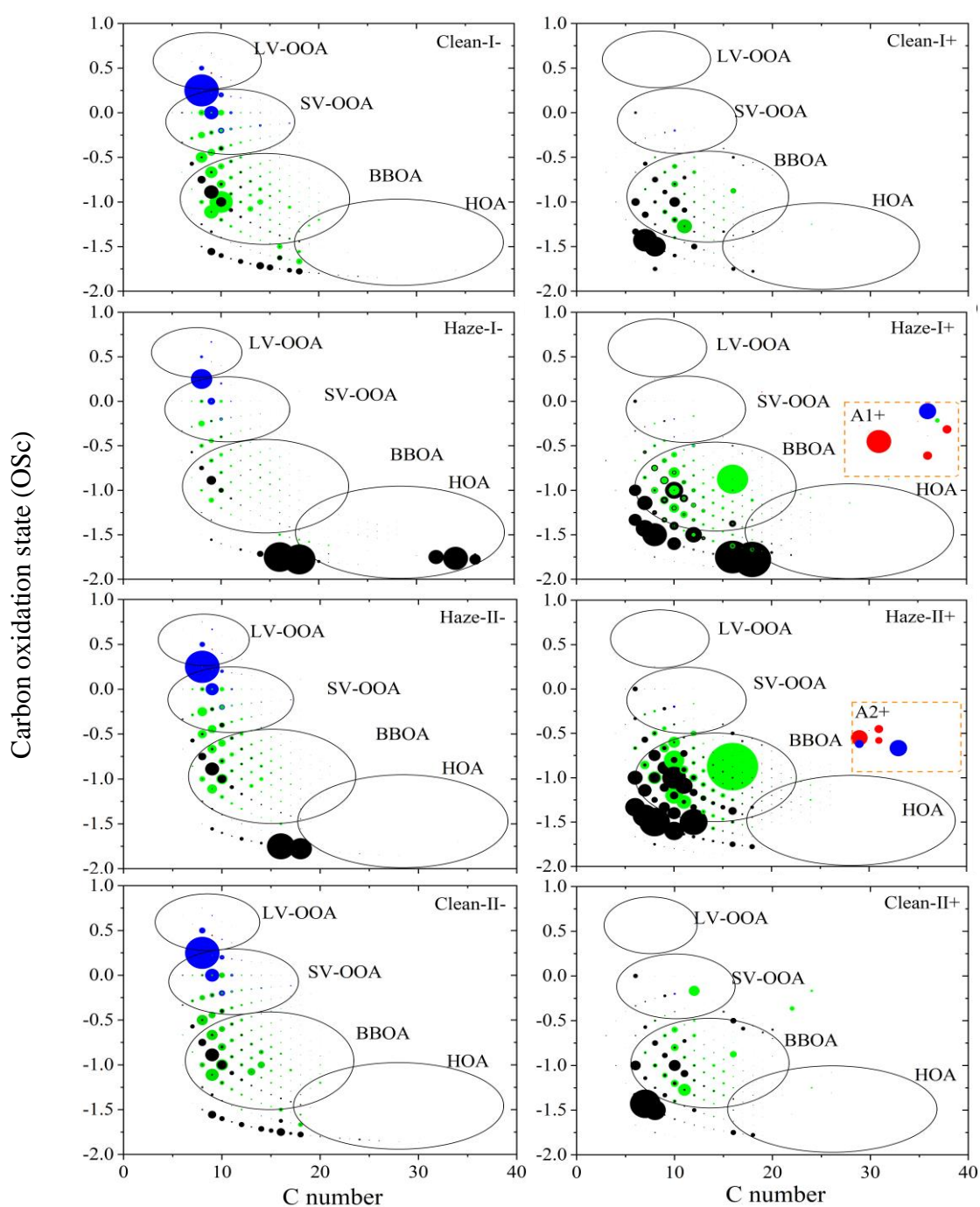
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**Figure 2.** Mass spectra of HULIS detected in ESI- and ESI+ modes during the haze process. The pie charts represent the intensity percent of different compound groups.

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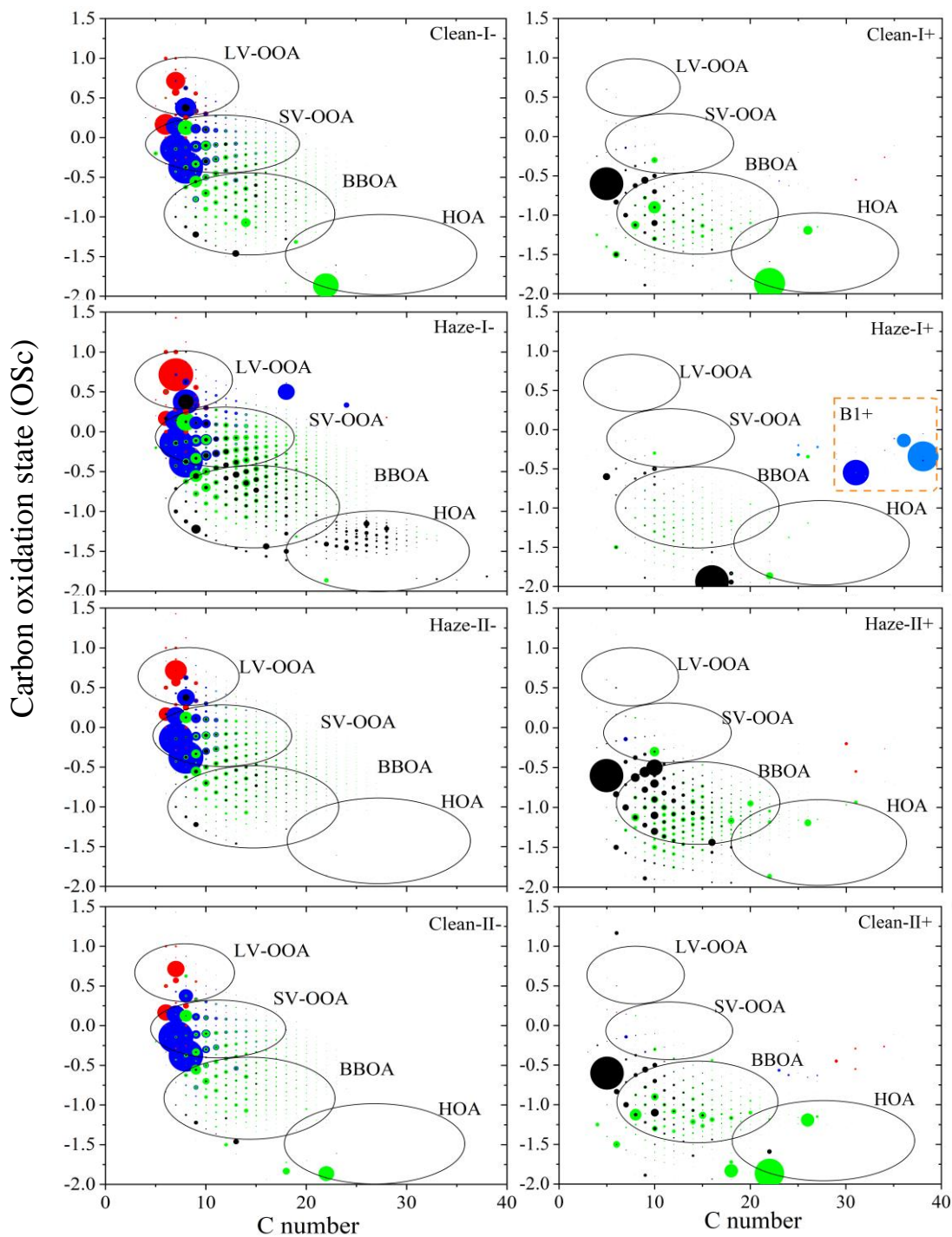
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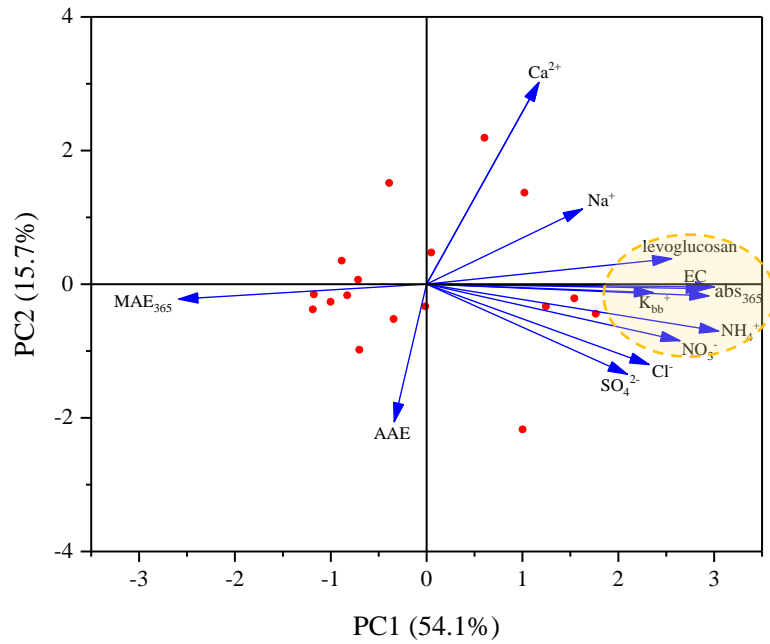
1010 **Figure 3.** Carbon oxidation state (OSc) plots for CHO<sup>-</sup> and CHO<sup>+</sup>. Formulas with black,
 1011 green, blue, and red are assigned to aliphatic (AI = 0), olefinic (0 < AI < 0.5), aromatic
 1012 (0.5 ≤ AI < 0.67), and condensed aromatic (AI ≥ 0.67) species (Koch and Dittmar, 2006),
 1013 respectively.

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1046 **Figure 4.** Carbon oxidation state (OSc) plots for CHON- and CHON+. Formulas with  
 1047 black, green, blue, and red are assigned to aliphatic (AI = 0), olefinic ( $0 < AI < 0.5$ ),  
 1048 aromatic ( $0.5 \leq AI < 0.67$ ), and condensed aromatic ( $AI \geq 0.67$ ) species (Koch and Dittmar,  
 1049 2006), respective.

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1067 **Figure 5.** Principal component analysis results for the optical properties of HULIS and  
1068 chemical compositions of PM<sub>2.5</sub>.  
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