



## Evidence for mass independent fractionation of even mercury isotopes in the troposphere

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**Abstract:** Mass independent fractionation (MIF) of even mercury (Hg) isotopes has long been observed in atmospheric related samples and is confirmed to be generated in the atmosphere, but its exact mechanism is covered up by the Hg sources and atmospheric transformations and stays unclear. Here, we present the first Hg isotope compositions of particulate bound mercury (PBM) in the Northwest Pacific and observe highly positive  $\Delta^{200}\text{Hg}$  values (up to 0.42‰). The MIF signatures are mainly controlled by photoreduction, gaseous elemental mercury (GEM) oxidation, and even-MIF dominated oxidation processes. Mercury in a small part of samples influenced by anthropogenic emissions is recognized by Hg concentrations and  $\Delta^{199}\text{Hg}$  signatures. The correlation between  $\Delta^{200}\text{Hg}$  and light conditions confirms that even-MIF is linked to photochemical reactions. The correlation between  $\Delta^{200}\text{Hg}$  and altitudes



24 suggests that a max even-MIF signatures existed in the troposphere. We  
25 use  $\Delta^{199}\text{Hg}/\Delta^{200}\text{Hg}$  ratios and ternary isotopic mixing model to estimate the  
26 contributions of photoreduction, GEM oxidation and even-MIF dominated  
27 oxidation. Our results demonstrate that atmospheric transformations are far  
28 more important than Hg sources in shifting Hg isotope compositions of  
29 PBM samples, especially in the marine boundary layer of the open ocean,  
30 which is characterized by less anthropogenic influences and has  
31 implications for our understanding of the mechanism of even-MIF and  
32 subsequently Hg behaviors in the atmosphere.

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## 1. Introduction

Mercury (Hg) mainly exists in three forms in the atmosphere: gaseous elemental mercury (GEM), gaseous oxidized mercury (GOM), and particulate bound mercury (PBM). GOM and PBM both exhibit a short residence time in the atmosphere and are readily deposited near Hg sources. In contrast, GEM is subject to long range transport and can be deposited significant distances from emission sources, including in relatively remote regions of the planet <sup>1</sup>. Atmospheric Hg continuously undergoes complicated physical and chemical transformations, including photoreactions, dark abiotic redox reaction, adsorption, and desorption, before being incorporated into the underlying surface <sup>2</sup>. Extensive efforts have been made to measure and model Hg in the atmosphere <sup>3</sup>. Recent advances in Hg stable isotopes have greatly improved our understanding of Hg cycling in the environment, including within the atmosphere <sup>4-6</sup>.

Mercury has seven stable isotopes (<sup>196</sup>Hg, <sup>198</sup>Hg, <sup>199</sup>Hg, <sup>200</sup>Hg, <sup>201</sup>Hg, <sup>202</sup>Hg, and <sup>204</sup>Hg). Mass dependent fractionation (MDF) of Hg isotopes is widely observed in natural environments, whereas significant mass independent fractionation of odd mass Hg isotopes (odd-MIF) is mainly induced by photochemical reactions <sup>7</sup>. Significant MIF of even mass Hg isotopes (even-MIF) has been measured in atmospheric samples or samples that relates to atmospheric Hg source <sup>8-14</sup>. However, the exact mechanism triggering even-MIF is unclear <sup>15</sup>. It is generally speculated that specific



58 oxidation reactions transforming Hg(0) to Hg(II) induced even-MIF. Fu et  
59 al. attributed even-MIF to surface-mediated Hg photoreduction involving  
60 halogens that can as of yet not explain <sup>16</sup>. Nonetheless, even-MIF  
61 signatures have become a useful tracer to constrain atmospheric Hg  
62 deposition pathways, reservoir size, and atmospheric conditions. Both  
63 MDF and MIF provide useful information to trace Hg sources and identify  
64 Hg transformation processes <sup>13,16-18</sup>.

65 Isotopic compositions of PBM in urban and remote areas reflect  
66 differing controls including anthropogenic emissions and atmospheric  
67 transformations <sup>19-23</sup>. A recent study proposed that atmospheric  
68 transformations could induce more positive odd-MIF values in PBM than  
69 that from anthropogenic Hg sources <sup>24</sup>. Of marine boundary layer (MBL),  
70 most PBM data have been collected from ground platforms <sup>2</sup> and show that  
71 PBM plays an important role in the geochemical cycling of Hg. The MBL  
72 is the largest transport layer and reaction vessel of atmospheric Hg due to  
73 its relatively high humidity, sufficient sunlight, and abundant atmospheric  
74 oxidants <sup>25,26</sup>. To date, several studies have been conducted in the MBL  
75 over the open ocean and suggested that continental anthropogenic  
76 emissions have little contribution to Hg in the open seas MBL and GEM  
77 oxidation and photoreduction greatly shift isotopic compositions of Hg  
78 species (i.e., GEM, PBM) in the MBL <sup>17,27</sup>. Hence, the Hg isotope  
79 signatures in the MBL is of great significance to understand the



80 transformation of atmospheric Hg. However, the relative contributions of  
81 GEM oxidation, photoreduction, and other atmospheric Hg transformation  
82 processes is rarely quantified and reported.

83 This study investigated the distribution and source of PBM in the  
84 MBL over the Northwest Pacific Ocean and the associated controlling  
85 factors on even-MIF. PBM samples were collected during three cruises and  
86 subsequently analyzed for selected Hg isotopes. The isotopic compositions  
87 of the PBM were then combined with the Hybrid Single-Particle  
88 Lagrangian Integrated Trajectory (HYSPLIT) model <sup>28,29</sup> to identify the  
89 sources of PBM. Atmospheric transformations were identified by odd-MIF  
90 and even-MIF. We proposed two ways to quantify contributions from  
91 photoreduction, GEM oxidation, and even-MIF dominated photochemical  
92 reactions in troposphere using Hg isotopic compositions.

93

## 94 2. Methods

95 **2.1 Study sections.** Three Cruises were conducted in the Northwest Pacific  
96 during the periods from August to November 2019 (denoted as Cruise A),  
97 May to June 2018 (denoted as Cruise B), and August to September 2019  
98 (denoted as Cruise C) (**Fig. 1**). Cruise A circumnavigated the southern  
99 portion of the Northwest Pacific (7.00–24.74°N), whereas Cruise B  
100 primarily navigated the western North Pacific around 30°N (21.00–  
101 37.02°N). Cruise C was extended from the northwest marginal sea of the



102 North Pacific to the Bering Strait (34.52–73.34°N). Based on atmospheric  
103 circulation and the results of 120 h back-trajectory analysis of air masses  
104 from the HYSPLIT model, the northeast trade wind prevails at 0°N–30°N  
105 and the westerly wind prevails at 30°N–60°N in the North Pacific. Polar  
106 easterly winds prevail at high latitudes. The area of the South China Sea  
107 and Western Pacific sampled during Cruise A is affected by the southwest  
108 monsoon, while Northeast Asia, sampled during Cruise B, is controlled by  
109 the northwest monsoon.

110

111 **2.2 Sample collection.** Two ARA N-FRM samplers (ARA Instruments,  
112 USA) were deployed on the compass deck of the research vessel at a height  
113 of approximately 15 m above mean sea level (AMSL) and at about 10 m  
114 upstream of the exhaust outlet. To reduce the potential for contamination  
115 from the ship's exhaust plume, sampling was stopped during station work  
116 and when bad weather was encountered. Quartz fiber membranes (Grade  
117 QMA, 47 mm, Whatman) were used for the collection of PBM as outlined  
118 in a previous study <sup>21</sup>. The sampling time of each sample lasted for 48–72  
119 h at a flow rate of about 20 L/min.

120

121 **2.3 Sample preconcentration.** PBM on the membrane was released via a  
122 dual-stage tube furnace combustion protocol <sup>21,30</sup>. A KMnO<sub>4</sub> trapping  
123 solution (0.1% KMnO<sub>4</sub> (m/v) + 10% H<sub>2</sub>SO<sub>4</sub> (v/v)) was utilized to capture



the released Hg<sup>21</sup>, after which it was preserved in the dark at 4°C until analyzed for Hg concentration and relative isotopic abundances. Mercury concentrations in the trapping solutions were measured by cold vapor atomic fluorescence spectrometry (CVAFS, MERX, Brooks Rand Instruments, USA) following US EPA Method 1631. PBM levels were so low that Hg concentrations in the trapping solutions were lower than 1.0 ng/L. Hg standard solutions (NIST 3133) were therefore added to ensure sufficient Hg mass (approximately 5 ng) for isotope analysis following our previous study (**Note S1**)<sup>31</sup>.

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**2.4 Hg stable isotope analyses.** Mercury isotopic compositions of the solutions were determined with a Nu Plasma Multi-Collector Inductively Coupled Plasma Mass Spectrometer (MC-ICPMS) housed at the State Key Laboratory of Marine Environmental Science at Xiamen University. A modified cold-vapor generator and an Aridus III desolvating nebulizer system (CETAC Technologies, USA) were used for Hg and thallium (Tl) introduction, respectively, following previously published methods<sup>32</sup>. The Hg isotopic compositions are reported in  $\delta$  (‰) and  $\Delta$  (‰) notation, which represents the MDF and MIF of the Hg isotopes, respectively<sup>33</sup>, such that:

$$\delta^{\text{xxx}}\text{Hg}_{\text{sample}} = \left( \left( \text{xxxHg}_{\text{sample}} / ^{198}\text{Hg}_{\text{sample}} \right) / \left( \text{xxxHg}_{\text{NIST3133}} / ^{198}\text{Hg}_{\text{NIST3133}} \right) - 1 \right) \times 1000 \quad (5)$$

$$\Delta^{\text{xxx}}\text{Hg} = \delta^{\text{xxx}}\text{Hg} - \beta \times \delta^{202}\text{Hg} \quad (6)$$



146 where xxx refers to the mass of each Hg isotope with atomic mass units  
147 (amu) of 199, 200, 201, and 202. The fractionation factor  $\beta$  is 0.2520,  
148 0.5024, and 0.7520 for  $^{199}\text{Hg}$ ,  $^{200}\text{Hg}$ , and  $^{201}\text{Hg}$ , respectively.

149 The repeated measurements of NIST 8610 gave long-term average  
150  $\delta^{202}\text{Hg}$ ,  $\Delta^{199}\text{Hg}$ , and  $\Delta^{200}\text{Hg}$  values of  $-0.52 \pm 0.12\text{‰}$  (2SD, n=5),  
151  $-0.02 \pm 0.04\text{‰}$  (2SD, n=5), and  $0.00 \pm 0.02\text{‰}$  (2SD, n=5), respectively.  
152 These values are in accordance with previous studies<sup>12,34,35</sup>. Since the  
153 samples were measured only once, the 2SD of the isotopic compositions  
154 for each sample were selected as the 2SD of NIST 8610 (**Table S1**). The  
155 isotopic compositions of PBM and the corresponding uncertainties (see  
156 **Table S2**) were calculated by the method we previously developed<sup>31</sup>.  
157 Additional details pertaining to the approach are provided in the **Note S2**.

158

159 **2.5 Quality assurance and quality control.** Any unused pre-cleaned  
160 quartz filters were placed in the closed sampling systems for 2 d to obtain  
161 field blanks of PBM. In the process of sample preconcentration, unused  
162 pre-cleaned quartz filters were periodically combusted with the samples to  
163 obtain method blanks. For Hg isotopic analyses, the sample introduction  
164 system was rinsed between samples with 3%  $\text{HNO}_3$  solution until the signal  
165 intensity dropped to background levels to avoid memory effects. Generally,  
166 the blanks accounted for <3% in all the trapping solutions. The dual-stage  
167 tube furnace combustion protocol was tested by adding NIST 3133 to



168 unused pre-cleaned quartz filters. The recovery was  $101 \pm 8\%$  ( $n=6$ , 1SD).

169 Standard solutions (NIST 8610) were used as a substitute for samples  
170 and mixed with NIST 3133 in different proportions (33.3%, 50%, and  
171 66.7%) to evaluate the isotopic compositions of PBM. The calculated  
172 isotopic compositions of NIST 8610 were  $-0.57 \pm 0.10\text{‰}$  (1SD,  $n=8$ ),  
173  $-0.07 \pm 0.09\text{‰}$  (1SD,  $n=8$ ), and  $0.01 \pm 0.04\text{‰}$  (1SD,  $n=8$ ) for  $\delta^{202}\text{Hg}$ ,  $\Delta^{199}\text{Hg}$ ,  
174 and  $\Delta^{200}\text{Hg}$ , respectively (**Fig. S1**).

175

176 **2.6 Backward trajectories of air masses and identification of potential**  
177 **source regions.** Air mass transport to the sampling area was assessed using  
178 the NOAA Air Resources Laboratory GDAS  $1^\circ$  data archive and the  
179 HYSPLIT model. Trajectory frequencies were analyzed to identify the  
180 potential source as well as represent air mass transport. The trajectory  
181 frequency started a trajectory from a single location and height every 6 h  
182 and then summed the frequency that the trajectory passed over a grid cell.  
183 This data was then normalized by the total number of trajectories (**Fig. S2**).  
184 The parameter settings are shown in **Table S3**.

185 Locations of each sampling site are reported as the mean longitude  
186 and latitude of the starting and ending points of sampling. The start time  
187 used to run the model was chosen as the end date of sampling for each  
188 sample (**Table S2**). The backward trajectories were calculated at a height  
189 of 500 m AMSL and are representative of the approximate height of the



190 MBL where atmospheric pollutants are well mixed.

191

## 192 **2.7 Analyses of height of backward trajectories of air mass.** 315 h

193 backward trajectories of air mass arrived at 15 m height AMSL at each  
 194 sampling site were calculated at 1 h intervals using HYSPLIT model. Three  
 195 backward trajectories were output for each sampling site at starting point,  
 196 intermediate point, and end point. A total of 948 height values were  
 197 obtained for each sample. The frequency of trajectory points with height  
 198 higher than 500 m, 1500 m, and 3000 m AMSL ( $f_{500}$ ,  $f_{1500}$ , and  $f_{3000}$ ) were  
 199 calculated, respectively (**Table S4**). Furthermore, limit condition that  
 200 trajectory points fall in day light ( $f_{500\_light}$ ,  $f_{1500\_light}$ , and  $f_{3000\_light}$ ) was added  
 201 to represent air mass originated from upper atmosphere and suffered from  
 202 sunlight.

203

## 204 **2.8 Estimation of the max $\Delta^{200}\text{Hg}$ values produced in the troposphere.**

205 For convenience, only three heights ( $F_{\text{position}} < 1/18$ ,  $1/18 < F_{\text{position}} < 1/6$ ,  
 206  $F_{\text{position}} > 1/6$ ) were chose to calculate the max  $\Delta^{200}\text{Hg}$  value based on the  
 207 following equations:

$$208 \quad f_a \times \Delta_a^{200}\text{Hg}_i + f_b \times \Delta_b^{200}\text{Hg}_i + f_c \times \Delta_c^{200}\text{Hg}_i = \Delta_{\text{sam}_{iv}}^{200}\text{Hg} \quad (7)$$

$$209 \quad f_a + f_b + f_c = 1 \quad (8)$$

210 where  $f_a$ ,  $f_b$ , and  $f_c$  represent frequencies of trajectory points at a height of  
 211  $F_{\text{position}} < 1/18$ ,  $1/18 < F_{\text{position}} < 1/6$ , and  $F_{\text{position}} > 1/6$ , respectively.



212  $\Delta_a^{200}Hg_i$ ,  $\Delta_b^{200}Hg_i$ ,  $\Delta_c^{200}Hg_i$  represent  $\Delta^{200}Hg$  values at the three  
213 heights and modeled by Monte Carlo simulation approach through the  
214 pseudorandom number generation ( $i=1:10000$ ).  $\Delta_{sam_{iv}}^{200}Hg$  represent  
215  $\Delta^{200}Hg$  values of PBM samples of type v.

216

217

### 218 3. Results

219 **3.1 PBM concentrations.** PBM concentrations exhibited mean values of  
220  $16.1 \pm 9.2$  pg/m<sup>3</sup> ( $n=33$ , 2SD) and  $16.0 \pm 10.5$  pg/m<sup>3</sup> ( $n=7$ , 2SD) during  
221 Cruises A and B, respectively (**Fig. 2a**). Concentrations were lower than  
222 continental boundary layer PBM in most areas, especially in East Asia, but  
223 higher than those in the MBL of the Pacific ocean (typically  $<10$  pg/m<sup>3</sup>)<sup>2</sup>.  
224 Long-term observations have shown that PBM concentrations are usually  
225 lower than 30 pg/m<sup>3</sup> in the free troposphere<sup>36</sup>, and thus, similar to that  
226 found herein for the MBL.

227 Highly variable PBM concentrations ( $11.5$ – $63.4$  pg/m<sup>3</sup>) were found in  
228 samples from Cruise C; the mean concentration was  $31.8 \pm 32.8$  pg/m<sup>3</sup>  
229 ( $n=12$ , 2SD). The lowest concentration was observed at station 41, which  
230 was in the Sea of Japan, while the highest concentration was observed at  
231 station 47, which was in the Bering Sea. Compared to Cruises A and B,  
232 Cruise C was closer to the coast. The elevated PBM concentrations in  
233 Cruise C may thus be due to anthropogenic influences.



234 **3.2 PBM isotope compositions.** The isotope compositions of PBM  
235 collected from the MBL exhibited negative  $\delta^{202}\text{Hg}$  ( $-0.98 \pm 1.46\text{‰}$ ,  $n=52$ ,  
236 2SD) and variable  $\Delta^{199}\text{Hg}$  values ( $-0.08 \pm 0.56\text{‰}$ ,  $n=52$ , 2SD). The isotopic  
237 signatures overlapped with previously reported PBM isotope data. As  
238 shown in **Fig. 3**, the PBM collected from urban regions exhibited large  
239 variations in  $\Delta^{199}\text{Hg}$  values ( $-0.02 \pm 0.61\text{‰}$ ,  $n=205$ , 2SD) <sup>19-23,37,38</sup>. More  
240 positive  $\Delta^{199}\text{Hg}$  values ( $0.40 \pm 0.71\text{‰}$ ,  $n=128$ , 2SD) have been observed for  
241 PBM collected from areas of high altitude (from 50 m to 3816 m) <sup>23,24</sup>. The  
242 PBM collected from polar regions displayed positive  $\delta^{202}\text{Hg}$  and negative  
243  $\Delta^{199}\text{Hg}$  values <sup>11,13</sup>. In contrast, the isotopic compositions of PBM collected  
244 from specific emission sources, such as vehicles, industry, waste, flue gas,  
245 and volcanos, have been characterized by negative  $\delta^{202}\text{Hg}$  and near-zero  
246  $\Delta^{199}\text{Hg}$  <sup>5,39,40</sup>. In addition, significant positive  $\Delta^{200}\text{Hg}$  values ( $0.14 \pm 0.30\text{‰}$ ,  
247  $n=52$ , 2SD) were observed in the current study (**Fig. 2d**).

248 Different isotopic signatures were found among different PBM  
249 samples within the same cruise and between different cruises (**Fig. 2 & Fig.**  
250 **S3**). Lower  $\delta^{202}\text{Hg}$  values were observed near Micronesia and the Bering  
251 Sea during Cruises A and C, while higher  $\delta^{202}\text{Hg}$  values were found during  
252 Cruise B (**Fig. 2b**). The spatial distribution of  $\Delta^{199}\text{Hg}$  of PBM was more  
253 complicated (**Fig. 2c**). For Cruises A and B, positive  $\Delta^{199}\text{Hg}$  values were  
254 mainly distributed in areas of open ocean in the Northwest Pacific, while  
255 negative  $\Delta^{199}\text{Hg}$  values were dispersed. Positive and negative  $\Delta^{199}\text{Hg}$



values were observed during Cruise C in the marginal sea of Northeast Asia and the Bering Sea, respectively.

#### 4. Discussion

According to the  $\Delta^{199}\text{Hg}$  values, PBM samples could be divided into two groups (**Fig. S4**): 1) samples characterized by positive  $\Delta^{199}\text{Hg}$  values, which were related to the air mass from areas of open ocean (e.g., the Northwest Pacific and the marginal sea in Northeast Asia), and 2) samples possessing negative  $\Delta^{199}\text{Hg}$  values related to the air mass from the South China Sea, Micronesia, Hawaii, Japan, Russia, and Bering Sea. Air mass source analysis was predicted by HYSPLIT model (see Methods).

Long-range transport from anthropogenic emissions in mid-latitude regions and GEM oxidation are two possible sources for the negative  $\Delta^{199}\text{Hg}$  values of PBM<sup>13</sup>. The  $\Delta^{199}\text{Hg}$  values are plotted against Hg concentrations in the PBM in **Fig. 4a**. Positive correlation are observed in PBM samples with negative  $\Delta^{199}\text{Hg}$  values (**Fig. 4b**). The samples collected during Cruises B and C possess a stronger correlation ( $r^2=0.61$ ,  $p<0.05$ ) than the samples collected during Cruise A ( $r^2=0.04$ ,  $p>0.1$ ). The higher Hg concentrations in PBM are most likely caused by the long-range transport from anthropogenic emissions, which produced the slightly negative  $\Delta^{199}\text{Hg}$  values. However, the significantly negative  $\Delta^{199}\text{Hg}$  values associated with lower Hg concentrations in the PBM presumably result from the *in situ* GEM oxidation. A few data points with slightly negative



278  $\Delta^{199}\text{Hg}$  deviate positively from the regression lines in **Fig. 4b**. These  
279 samples was likely influenced by the photoreduction of  $\text{Hg(II)}$ , which  
280 induced positive  $\Delta^{199}\text{Hg}$  values in residual  $\text{Hg(II)}$ . Noted that the  
281 background PBM pool is characterized by significant positive  $\Delta^{199}\text{Hg}$   
282 values, which is also associated with photoreduction <sup>24</sup>. Hence, samples  
283 related to the open ocean with highly positive  $\Delta^{199}\text{Hg}$  values are mainly  
284 controlled by photoreduction.

285 A large variation range of  $\delta^{202}\text{Hg}$  values are observed for PBM  
286 samples with both positive and negative  $\Delta^{199}\text{Hg}$  values (**Fig. 4c**).  
287 According to previous study, strong photoreduction of  $\text{Hg(II)}$  induced  
288 positive shift of  $\delta^{202}\text{Hg}$  values in reactants (e.g., PBM) <sup>27,41</sup>. Although  
289 inverse kinetic isotope effect was observed in  $\text{Hg(0)}$  oxidation by Cl and  
290 Br atoms <sup>42</sup>, the subsequent adsorption of  $\text{Hg(II)}$  on particulate surfaces  
291 would lead to negatively shift of  $\delta^{202}\text{Hg}$  values in PBM. Furthermore, GEM  
292 originated from dissolved gaseous mercury displayed more negative  
293  $\delta^{202}\text{Hg}$  ( $-2.98$  to  $-0.99\text{‰}$ ) <sup>14</sup> than that in the upper atmosphere ( $-0.02$  to  
294  $1.64\text{‰}$ ) <sup>43,44</sup>, indicating that it is difficult to identify Hg source with  $\delta^{202}\text{Hg}$   
295 values. In addition to photochemical and redox reactions, gas-particle  
296 partitioning, dissolution, and evaporation would also affect  $\delta^{202}\text{Hg}$  values  
297 in PBM. Moreover, no correlation was found between  $\delta^{202}\text{Hg}$  values and  
298 Hg concentrations ( $p>0.1$ , **Fig. 4d**). Therefore, the  $\delta^{202}\text{Hg}$  values cannot be  
299 used as a diagnostic tool for atmospheric transformation processes.



300 However, the PBM samples with high Hg concentrations are characterized  
301 by negative  $\delta^{202}\text{Hg}$  values, confirming the contribution from anthropogenic  
302 sources. These samples are related to the air mass from Bering Sea,  
303 coincide with enhanced levels of TGM observed in Bering Sea <sup>45</sup>.

304 In addition to odd-MIF, the even-MIF is an important indicator for Hg  
305 sources and transformation processes in the atmosphere <sup>16</sup>. As shown in  
306 **Fig. 5**, the  $\Delta^{200}\text{Hg}$  values were considered to be contributed by  
307 photochemical reactions in the upper troposphere. GEM oxidation could  
308 also induce even-MIF at low altitudes, but limited to specific oxidants (e.g.,  
309 Br, Cl). This process could be identified by backward air mass trajectory  
310 as discussed below. The GEM oxidation processes without even-MIF, the  
311 resulting PBM generally inherited even-MIF signatures of GEM and RGM  
312 and was characterized by near-zero  $\Delta^{200}\text{Hg}$  values. Here, we define PBM  
313 dominated by troposphere photochemical reactions are characterized by  
314 odd-MIF and significant even-MIF ( $\Delta^{200}\text{Hg} \geq 0.15\text{‰}$ ) and PBM dominated  
315 by GEM oxidation without even-MIF at low altitudes are characterized by  
316 negative odd-MIF and near-zero even-MIF at PBM. Therefore, except for  
317 samples with high Hg concentrations (anthropogenic emissions), most  
318 PBM samples were related with photoreduction (+odd-MIF), GEM  
319 oxidation (−odd-MIF), and troposphere photochemical reactions (+even-  
320 MIF), which could be identified by  $\Delta^{199}\text{Hg}$  and  $\Delta^{200}\text{Hg}$  values. Samples  
321 related with GEM oxidation are overlapped with PBM collected from polar



region, which is also similar to GEM in terms of  $\Delta^{199}\text{Hg}$  and  $\Delta^{200}\text{Hg}$  values. According to previous study, these samples are primarily sourced from near-complete GEM oxidation<sup>11,13</sup>. As photoreduction induces no even-MIF and troposphere photochemical reactions are considered to induce even-MIF, samples with positive  $\Delta^{199}\text{Hg}$  and near-zero  $\Delta^{200}\text{Hg}$  values are believed to experience photoreduction rather than photo-oxidation and vice versa. Thus, the PBM samples could be classified into 6 types of different sources and atmospheric transformations based on Hg concentrations and isotope signatures (**Table 1**).

The ratio of  $\Delta^{199}\text{Hg}$  to  $\Delta^{201}\text{Hg}$  is commonly used to identify fractionation processes. The  $\Delta^{199}\text{Hg}/\Delta^{201}\text{Hg}$  ratio was 0.75 and 0.81 for the groups with positive and negative  $\Delta^{199}\text{Hg}$  values, respectively (**Fig. S5a**). Both are lower than those observed for Hg(0) photo-oxidation (1.64–1.89)<sup>42</sup> and Hg(II) photo-reduction (1.00–1.31)<sup>41,46</sup>, but higher than that in TGM from remote areas (0.73)<sup>10,23,47</sup> and coastal areas (0.55)<sup>48,49</sup>. The ratio lower than 1.00 has been presented in oxic experiments during photochemical reduction of Hg(II)<sup>50</sup>. These ratios indicated that the odd-MIF of PBM isotopes was induced by multiple processes rather than a single oxidation or reduction process. For each cruise, the ratio increased in the order Cruise A > Cruise B > Cruise C (**Fig. S5b**). With regards to the variations in latitude among the three cruises (**Table S2**), the ratios have a statistical correlation with mean latitude of each cruise ( $R^2=0.99$ ,  $p<0.01$ ).



344 However, no correlations between the ratios and the corresponding mean  
345  $\Delta^{199}\text{Hg}$  value was found, which has been observed in PBM samples with  
346 highly positive  $\Delta^{199}\text{Hg}$  from megacities<sup>51</sup>. Also, no correlations between  
347 odd-MIF and latitude for PBM were found ( $p>0.05$ ). Similarly, no  
348 correlations between odd-MIF and radiation duration for PBM were  
349 observed ( $p>0.05$ , the radiation duration was represented by astronomical  
350 twilight which could be obtained from websites [www.wunderground.com](http://www.wunderground.com)),  
351 which is inconsistent with a previous study<sup>20</sup> and implies potential impact  
352 of oxidations on the odd-MIF. It has been previously suggested that  
353 troposphere photochemical reactions that caused even-MIF was possibly  
354 accompanied by odd-MIF, although the magnitude is smaller than that  
355 induced by photoreduction<sup>16</sup>. Another reason maybe is that this study used  
356 ship-based sampling rather than fixed-point sampling, the baseline  
357 (background) PBM isotope signatures varied with sampling sites.  
358 Additionally, latitude was closely related with air temperature and light  
359 intensity, which may affect the MIF of PBM, but air temperature and light  
360 intensity were absent in this study. Nonetheless, the ratio for all PBM  
361 samples was 1.02 in this study, suggested odd-MIF was mainly caused by  
362 photochemical reactions and  $\Delta^{199}\text{Hg}$  was a comprehensive result of  
363 different transformations.

364 The mechanism of even-MIF has been discussed for nearly a decade.  
365 Most scientists support the view that even-MIF is produced in the upper



atmosphere and could be used to estimate the contribution of Hg from the upper troposphere<sup>11,12,15,24</sup>. Recent studies, however, propose a different mechanism in which small but significant  $\Delta^{200}\text{Hg}$  anomalies can originate from *in situ* photooxidation of Hg(0) by UV light and oxidants at low altitudes<sup>16,24,42</sup>. Thus, we traced the air mass height of PBM samples during the past 13 days in comparison to the moment of sample collection (**Fig. S6**). This timeframe was selected because the global mean lifetime of Hg(II) before photoreduction in the troposphere is approximately 13 days<sup>52</sup>. The back trajectory model showed that air masses mainly (>90%) originated from below 1.5 km AMSL before arriving at the sampling sites in 21 out of 52 of the PBM samples, while air masses were completely derived from below 500 m AMSL in only 3 PBM samples (see **Methods; Table S4**). Highly positive  $\Delta^{200}\text{Hg}$  values ( $\geq 0.15\text{‰}$ ) were observed in 24 PBM samples. Discrete data points exhibited  $\Delta^{200}\text{Hg}$  values that ranged from  $-0.11\text{‰}$  to  $0.14\text{‰}$  ( $n=28$ ). Here, we analyzed the relationships between  $\Delta^{200}\text{Hg}$  values and frequencies of trajectory points when heights were higher than 500 m, 1500 m, and 3000 m AMSL ( $f_{500}$ ,  $f_{1500}$ , and  $f_{3000}$ ) (**Fig. 6a, b, and c**). Air masses mainly originated from low altitudes during Cruise A, because the mean frequencies were higher in Cruises B and C (48.08% for  $f_{500}$  and 27.17% for  $f_{1500}$ ) than in Cruise A (25.33% for  $f_{500}$  and 11.41% for  $f_{1500}$ ). Consequently, the latter mechanism in which even-MIF produced at low altitudes was dominant in shaping  $\Delta^{200}\text{Hg}$  values of the



PBM samples is supported by three lines of evidence: (1) highly positive  $\Delta^{200}\text{Hg}$  values were positively correlated with multiple frequencies (i.e.,  $f_{500}, f_{1500}, f_{3000}$ ) during Cruise A, and the correlations were more significant at a lower altitude; (2) the positive relationships between highly positive  $\Delta^{200}\text{Hg}$  values and frequencies  $f_{500}$  and  $f_{1500}$  were stronger in Cruise A than in Cruises B and C; and (3) significant negative relationships were observed between highly positive  $\Delta^{200}\text{Hg}$  values and  $f_{3000}$  in Cruises B and C with more air masses from high altitudes. Moreover, the highly positive  $\Delta^{200}\text{Hg}$  values in 6 PBM samples displayed  $f_{1500} = 0$ , indicating that these samples were associated with an air mass from low altitudes ( $<1.5$  km) as the even-MIF in these samples was mainly produced at low altitudes. The mechanism could be due to vertical mixing of air masses from high and low altitudes, or photo-oxidation of  $\text{Hg}(0)$  at low altitudes. We also analyzed the frequencies as a function of height and light conditions; that is, when trajectory points occurred during day light ( $f_{500\_light}, f_{1500\_light}$ , and  $f_{3000\_light}$ ) (**Fig. 6d, e, and f**). The results showed that light played an important role in even-MIF, as  $\Delta^{200}\text{Hg}$  and  $f_{500\_light}$  ( $f_{1500\_light}, f_{3000\_light}$ ) exhibited an analogous correlation to  $\Delta^{200}\text{Hg}$  and  $f_{500}$  ( $f_{1500}, f_{3000}$ ) ( $p > 0.05$ , paired  $t$  test).

There would be a height that produce the max  $\Delta^{200}\text{Hg}$  value in terms of the negative relationship between  $f_{3000}$  and  $\Delta^{200}\text{Hg}$ , although the trend between  $f_{3000}$  and  $\Delta^{200}\text{Hg}$  is completely opposite in Cruise A and in Cruise



B and C, that is, the trend is opposite at low and high latitudes (**Fig. 6c**). In view of the fact that the tropospheric height decreases with the increase of latitude, the position of the same altitude relative to the tropopause ( $F_{\text{position}}$ ) is different at high and low latitudes. For examples, we set the tropopause height at high and low latitudes to 9 km and 18 km, the position of 3000 m is 1/3 and 1/6 of the tropopause, respectively. Hence, we could obtain  $F_{\text{position}} = \text{height}/9000$  at high latitudes and  $F_{\text{position}} = \text{height}/18000$  at low latitudes, respectively. Fit  $F_{\text{position}}$  to the slope of  $\Delta^{200}\text{Hg}$  and  $f_{500}$  ( $f_{1500}, f_{3000}$ ) linearly (**Fig. S7**), the slope showed a negative linear correlation of the  $F_{\text{position}}$ , indicating that, as the height increases, the  $\Delta^{200}\text{Hg}$  values first increases and then decreases. Hence, there should be a maximum  $\Delta^{200}\text{Hg}$  values at a certain height. The fitting curve ( $r^2=0.95, p<0.01$ ) intersects the x axis at 0.18, indicating that the max  $\Delta^{200}\text{Hg}$  value would be produced at  $F_{\text{position}} = 0.18$ , that is, the max  $\Delta^{200}\text{Hg}$  value would be produced at height of 3240 m and 1620 m at low and high latitudes, respectively. Thus, the results confirmed the opposite trend between  $f_{3000}$  and  $\Delta^{200}\text{Hg}$  at low and high latitudes. The max  $\Delta^{200}\text{Hg}$  value was estimated by a Monte Carlo simulation approach and isotopic mixing model using data of type v which is dominated by troposphere photochemical reactions (see **Methods**). The estimated value of the max  $\Delta^{200}\text{Hg}$  of  $1.10\pm0.58$  for PBM is overall similar to the maximum value ( $1.24 \pm 0.08$ ) currently observed in atmospheric environment<sup>8</sup>.



432 The precipitation samples are characterized by more positive  $\Delta^{200}\text{Hg}$   
433 values than PBM and GOM samples<sup>16,17,27,53</sup>, suggesting that even-MIF  
434 produced in the MBL over the open ocean tend to be accumulated in  
435 aqueous phase and brought into seawater. The open ocean seawater shows  
436 positive  $\Delta^{200}\text{Hg}$  values<sup>54,55</sup>. Latitudinal variations of  $\Delta^{200}\text{Hg}$  were observed  
437 in seawater and precipitation, although negative correlation was found in  
438 seawater  $\Delta^{200}\text{Hg}$  and latitudes and positive correlation was found in  
439 precipitation  $\Delta^{200}\text{Hg}$  and latitudes, respectively<sup>55,56</sup>. As suggested,  
440 seawater  $\Delta^{200}\text{Hg}$  are controlled by both Hg(0) and Hg(II) and larger ocean  
441 Hg(0) uptake at high latitudes results in low seawater  $\Delta^{200}\text{Hg}$ <sup>55</sup>. While for  
442 precipitation  $\Delta^{200}\text{Hg}$ , max  $\Delta^{200}\text{Hg}$  values in the troposphere are estimated  
443 to be generated in lower altitudes at high latitudes,  $\Delta^{200}\text{Hg}$  accumulated in  
444 precipitations would be less affected before reaching the ground (being  
445 collected).

446 As discussed above, photoreduction, troposphere photochemical  
447 reactions, and GEM oxidation are the three major processes that trigger  
448 odd-MIF and even-MIF (**Table 1**). Noted that adsorption and desorption  
449 between PBM and GOM induced no odd-MIF and the sources of Hg in the  
450 atmosphere are accompanied with near-zero odd-MIF<sup>5,39,40</sup>. Moreover, the  
451  $\Delta^{200}\text{Hg}$  values for GEM, GOM, and PBM are near-zero in source materials  
452 and therefore emissions<sup>57</sup>. Therefore, the  $\Delta^{200}\text{Hg}$  and  $\Delta^{199}\text{Hg}$  values could  
453 be used to estimate the contributions of photoreduction, photo-oxidation,



454 and GEM oxidation. Here, we proposed two ways to quantify contributions  
455 from atmospheric transformations: (a) specific diagnostic ratios of  
456  $\Delta^{199}\text{Hg}/\Delta^{200}\text{Hg}$ ; (b) ternary isotopic mixing model.

457 (a) Even-odd patterns of Hg isotope fractionation factors have been  
458 proposed to be a better indicator of MIF mechanisms than  $\Delta^{199}\text{Hg}/\Delta^{201}\text{Hg}$   
459 slopes. According to previous study, a consistent pattern between  $\Delta^{199}\text{Hg}$   
460 and  $\Delta^{200}\text{Hg}$  has been observed with large variation range (1.1-3.3) on GEM  
461 and oxidized Hg phases (i.e., reactive Hg and precipitation Hg)<sup>16,48,58</sup>. As  
462 listed in Table 1, samples in type i could be used to represent near-complete  
463 GEM oxidation and inherit the isotope composition of GEM. We then built  
464  $\Delta^{199}\text{Hg}/\Delta^{200}\text{Hg}$  slopes between type i and type ii, iii, and v, respectively.  
465 Based on Williamson-York bivariate linear regression method<sup>59</sup>, the  
466 observed fitted curve shaped a slope of  $1.81 \pm 0.27$  ( $r^2=0.57$ ,  $p<0.01$ ),  
467  $4.90 \pm 0.82$  ( $r^2=0.31$ ,  $p<0.05$ ), and  $0.03 \pm 0.16$  ( $r^2=0.00$ ,  $p=0.94$ ), respectively.  
468 We suggest that the slope was primarily determined by the isotope  
469 signatures of oxidized Hg phases, because GEM varies in a narrow range  
470 of  $\Delta^{199}\text{Hg}$  and  $\Delta^{200}\text{Hg}$  with  $-0.20 \pm 0.17\text{‰}$  and  $-0.06 \pm 0.12\text{‰}$  (2SD,  $n=208$ ),  
471 respectively<sup>9,10,13,16,44,48,58,60-63</sup>. Thus, we hypothesized that the ratios of  
472  $\Delta^{199}\text{Hg}/\Delta^{200}\text{Hg}$  could be used as a diagnostic for proportions of  
473 photoreduction and troposphere photochemical reactions despite of GEM  
474 oxidation. To gain a more representative result, we integrate our data with  
475 a subset of the published PBM data (**Fig. 7**). The resulting ratios of



$\Delta^{199}\text{Hg}/\Delta^{200}\text{Hg}$  were  $17.71 \pm 1.42$  ( $r^2=0.27$ ,  $p<0.01$ ) and  $0.61 \pm 0.06$  ( $r^2=0.43$ ,  $p<0.01$ ) for photoreduction-dominated and troposphere-oxidation-dominated PBM samples, respectively. The proportions of photoreduction and photo-oxidation could be calculated as the equation below:

$$f_{\text{red}} = \frac{\arctan(k) - \arctan 0.61}{\arctan 17.71 - \arctan 0.61} \times 100\% = \frac{\arctan(k) - 0.548}{0.966} \times 100\% \quad (1)$$

where  $f_{\text{red}}$  represents proportions of photoreduction and  $k$  represent ratios of  $\Delta^{199}\text{Hg}/\Delta^{200}\text{Hg}$ . Following equation (1), we suggested that  $66 \pm 16\%$  of Hg was photoreduced in precipitation and  $75 \pm 16\%$  Hg was photoreduced in oxidized Hg phases at high altitudes (**Table S6**). In this study, the  $f_{\text{red}}$  was calculated for sample types ii, iii, and v and listed in **Table 1**. It should be noted again that the  $f_{\text{red}}$  was calculated without considering the influence of GEM oxidation. If GEM oxidation effects greatly, such as the observed  $\Delta^{199}\text{Hg}$  values are near-zero or negative, the ratios would be useless. For example, the calculated  $f_{\text{red}}$  for sample types v was -0.58 and meaningless.

(b) As shown in **Fig. 7**, the three end-member could be used to calculate contributions from photoreduction, troposphere photochemical reactions, and GEM oxidation through ternary isotopic mixing model.

$$X \cdot \Delta^{200}\text{Hg}_{\text{red}} + Y \cdot \Delta^{200}\text{Hg}_{\text{oxi}} + Z \cdot \Delta^{200}\text{Hg}_{\text{GEMo}} = \Delta^{200}\text{Hg}_{\text{sam}} \quad (2)$$

$$X \cdot \Delta^{199}\text{Hg}_{\text{red}} + Y \cdot \Delta^{199}\text{Hg}_{\text{oxi}} + Z \cdot \Delta^{199}\text{Hg}_{\text{GEMo}} = \Delta^{199}\text{Hg}_{\text{sam}} \quad (3)$$

$$X + Y + Z = 1 \quad (4)$$

where X, Y, and Z represent the contribution of the three atmospheric transformations of the photoreduction, troposphere photochemical



reactions, and GEM oxidation, respectively. The Hg isotope signatures of the three atmospheric transformations are the mean values of the three end-member as mentioned above (**Table S5**). The results of the mixing model calculation showed that the contributions of photoreduction, troposphere photochemical reactions, and GEM oxidation varied within -12% to 77%, -12% to 35%, and 2% to 97%, respectively (**Table 1**). The results showed similar contribution from photoreduction with estimations from ratios of  $\Delta^{199}\text{Hg}/\Delta^{200}\text{Hg}$ . The estimated highly contributions ( $53\pm22\%$ , 1SD) of GEM oxidation is consistent with previous suggestion that GEM oxidation played a major role ( $\sim47\pm22\%$ ) in the formation of Hg(II) in PBM<sup>17,51</sup>. The weak contributions ( $10\pm12\%$ , 1SD) of troposphere photochemical reactions indicate strong dilutions of near-zero  $\Delta^{200}\text{Hg}$  from Hg sources and other atmospheric transformations.

511

## 5. Conclusion

The isotopic compositions of PBM in the MBL of the Northwest Pacific suggest that the even-MIF and odd-MIF signatures are useful tracers for identifying atmospheric transformations. The  $\delta^{202}\text{Hg}$  signature was significantly shifted by Hg sources and *in situ* transformations of atmospheric Hg. Strong correlations between the even-MIF of PBM and height of the air mass provide additional support for the occurrence of even-MIF in the troposphere, which may link to photochemical reactions



520 in the atmosphere. The mechanism driving the highly positive  $\Delta^{200}\text{Hg}$   
521 signatures of PBM in this study is probably the oxidation of  $\text{Hg}(0)$  at low  
522 altitudes. The MBL over the open ocean may promote the intrusion of  
523 even-MIF signatures and its subsequent recording by seawater and marine  
524 sediments<sup>55,64,65</sup>. It is highly uncertain how strong even-MIF signatures  
525 could be generated in the troposphere, the estimated max  $\Delta^{200}\text{Hg}$  value  
526 provides data support for future research. The quantification of  
527 contributions of atmospheric transformations by the proposed methods  
528 have implications for the study of Hg behaviors in the atmosphere.

529

530 **Description of statistical analysis.** All the statistical analyses were  
531 performed using Origin 2019b and Excel 2019. The Paired *t*-test, Pearson's  
532 R-Square and *P*-value are calculated by algorithms of the software.

533

534 **Data availability.** The authors declare that the main data supporting the  
535 findings of this study are available within the paper and its supplementary  
536 information files. Extra data are available from the corresponding author  
537 upon request.

538

539 **Author contributions**

540 S.H., and Y.Z.\* conceived and designed this project, S.H., Y.H., and H.S.  
541 conducted the field sampling, S.H., S.L., Y.Z., and Y.C. carried out all the



542 measurements, and S.H., K.L., and Y.Z.\* wrote the draft paper.

543

### 544 **Competing interests**

545 The authors declare no competing interests.

546

### 547 **Additional information**

548 Correspondence and requests for materials should be addressed to Y.Z.

549

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742 Table 1 The possible Hg sources and atmospheric transformations for PBM  
 743 samples and the estimated contributions

ID	X	Y	Z	$f_{\text{red}}^*$ (%)	$f_{\text{red}}^{**}$ (%)
Type i. GEM oxidation (negative $\Delta^{199}\text{Hg}$ and near-zero $\Delta^{200}\text{Hg}$ )					
16	0.13	-0.02	0.89	N.A.	N.A.
25	0.14	-0.01	0.87		
26	0.14	-0.09	0.95		
37	0.15	-0.09	0.94		
Type ii. Photoreduction & troposphere photochemical reactions (positive $\Delta^{199}\text{Hg}$ and $\Delta^{200}\text{Hg}$ )					
5	0.50	0.17	0.33	54±29	70±11
6	0.46	0.18	0.36		
7	0.34	0.33	0.33		
14	0.58	0.12	0.30		
19	0.65	0.10	0.25		
20	0.57	0.19	0.24		
23	0.77	0.21	0.02		
28	0.26	0.19	0.55		
32	0.61	0.27	0.13		
34	0.31	0.31	0.38		
35	0.37	0.25	0.38		
36	0.52	0.17	0.31		
39	0.42	0.10	0.47		
42	0.36	0.24	0.40		
43	0.50	0.16	0.34		
Type iii. Photoreduction (positive $\Delta^{199}\text{Hg}$ and near-zero $\Delta^{200}\text{Hg}$ )					
4	0.64	-0.03	0.39	85±14	100±14
8	0.54	-0.02	0.48		
9	0.45	-0.08	0.63		
10	0.48	0.08	0.44		
13	0.54	-0.06	0.53		
15	0.34	0.10	0.55		
21	0.38	0.09	0.53		
22	0.55	0.04	0.41		
24	0.74	-0.10	0.35		
27	0.50	-0.12	0.61		
30	0.45	0.09	0.46		
31	0.59	0.04	0.37		
41	0.57	0.03	0.40		
44	0.62	0.02	0.37		
45	0.74	-0.03	0.29		
50	0.53	-0.01	0.48		
51	0.58	0.00	0.42		



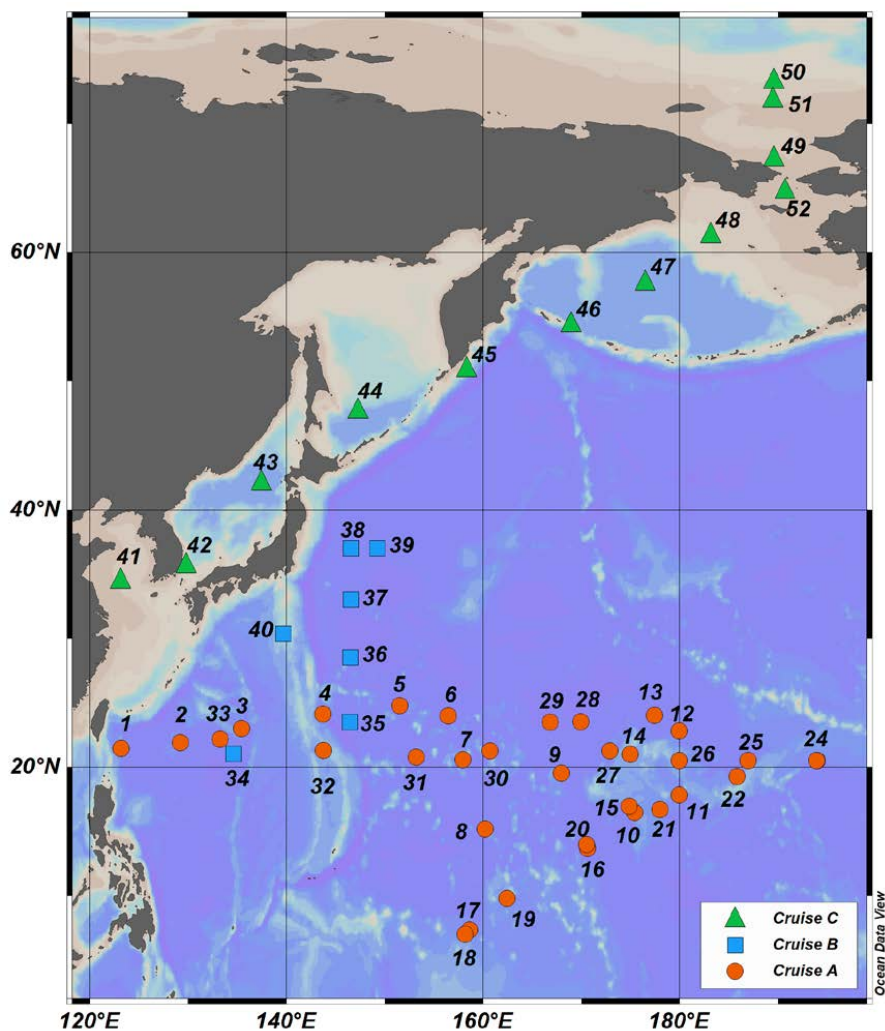
Type iv. GEM oxidation & photoreduction (slightly negative $\Delta^{199}\text{Hg}$ and near-zero $\Delta^{200}\text{Hg}$ )					
2	0.31	0.04	0.65		
11	0.28	0.09	0.63	N.A.	N.A.
33	0.31	-0.01	0.71		
38	0.25	0.08	0.67		
Type v. troposphere photochemical reactions (negative $\Delta^{199}\text{Hg}$ and positive $\Delta^{200}\text{Hg}$ )					
1	-0.12	0.34	0.78		
3	0.11	0.19	0.69		
12	-0.07	0.35	0.72		
17	0.12	0.24	0.64		
18	-0.10	0.13	0.97	-58±41	-37±128
29	0.00	0.27	0.73		
40	-0.06	0.23	0.83		
46	0.15	0.15	0.70		
52	0.07	0.21	0.72		
Type vi. Anthropogenic emissions (slightly negative $\Delta^{199}\text{Hg}$ and high Hg concentrations)					
47	0.29	-0.03	0.73		
48	0.26	0.04	0.70	N.A.	N.A.
49	0.23	0.06	0.70		
mean	0.37	0.10	0.53		
sd	0.23	0.12	0.22		

\*The  $f_{\text{red}}$  was calculated from ratios of  $\Delta^{199}\text{Hg}/\Delta^{200}\text{Hg}$ .

\*\* The  $f_{\text{red}}$  was calculated from  $X/(X+Y)$



## 747 Figures



748 Figure 1. Map of sampling locations. The points represent the locations  
 749 at the “midpoint” moment during the sampling time of PBM samples.  
 750 (Ocean Data View 2020)

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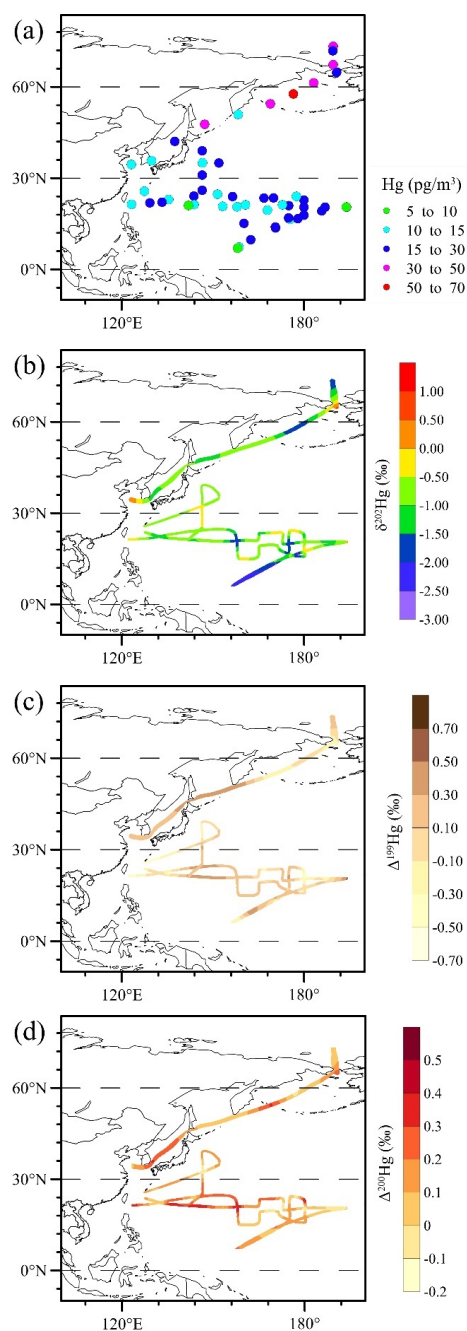


Figure 2. Spatial distributions of (a) Hg concentrations, (b)  $\delta^{202}\text{Hg}$ , (c)  $\Delta^{199}\text{Hg}$ , and (d)  $\Delta^{200}\text{Hg}$  of PBM in the MBL of the Northwest Pacific.

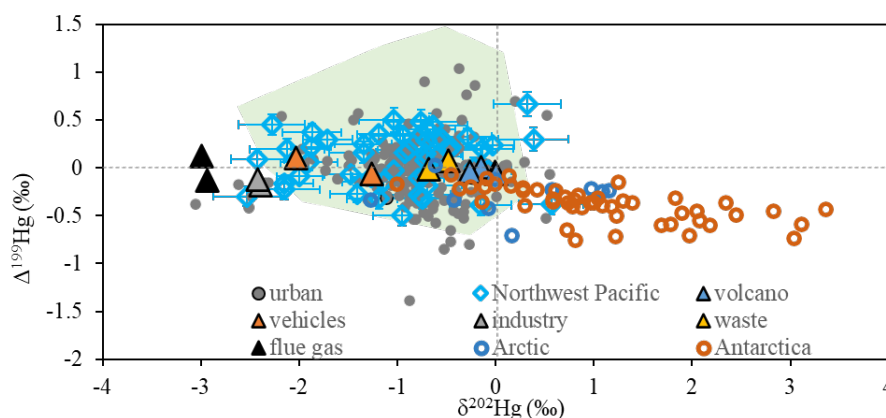
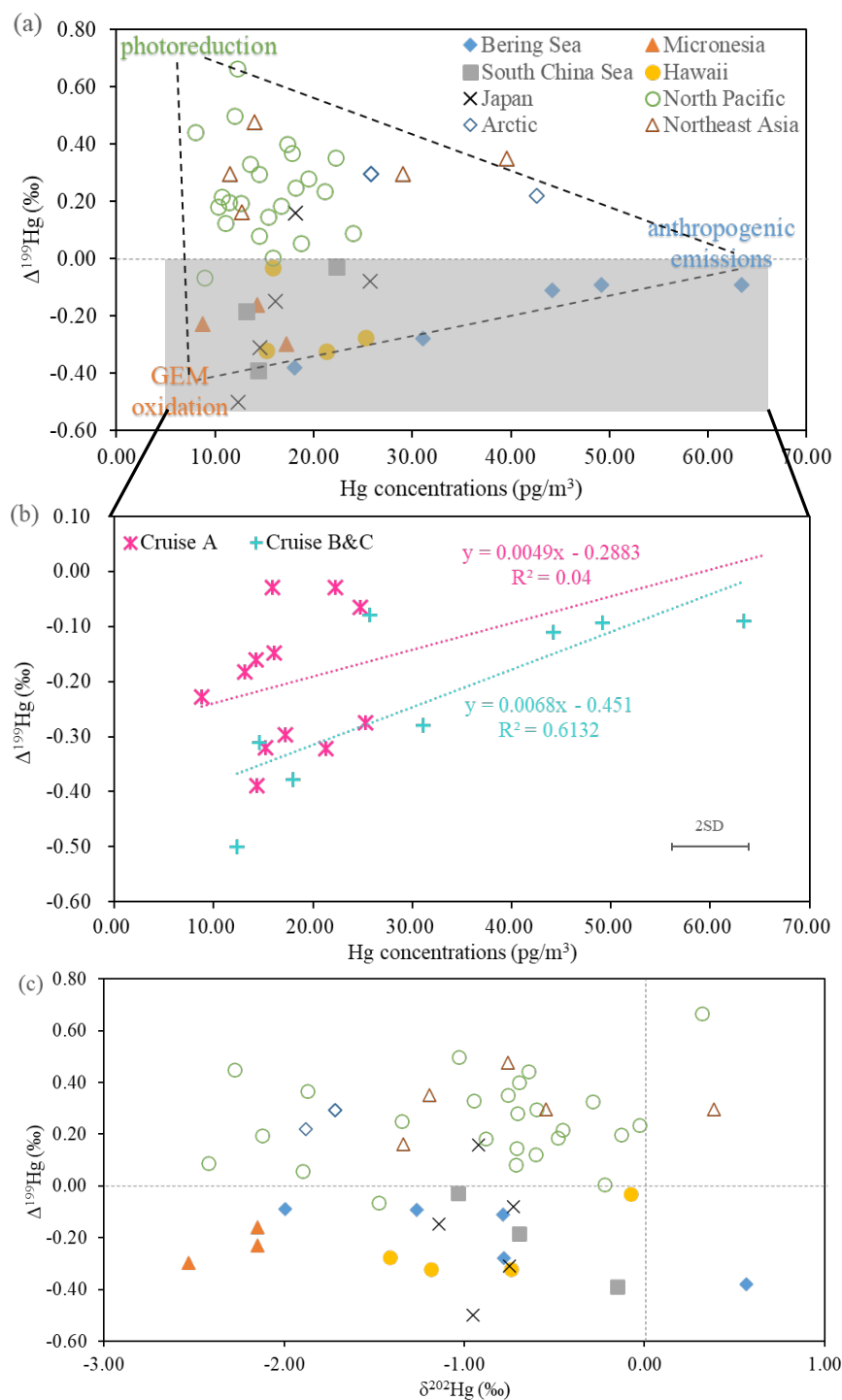


Figure 3. Plot of  $\Delta^{199}\text{Hg}$  vs.  $\delta^{202}\text{Hg}$  for PBM in the Northwest Pacific from this study, along with isotope data for PBM reported in the literature. The references for the literature data are: urban<sup>19-23,37,38</sup>, volcano<sup>40</sup>, traffic/vehicles<sup>5</sup>, industry<sup>5</sup>, waste<sup>5</sup>, flue gas<sup>39</sup>, arctic<sup>13</sup>, antarctica<sup>11</sup>, and high altitude areas (the green zone)<sup>23,24</sup>.



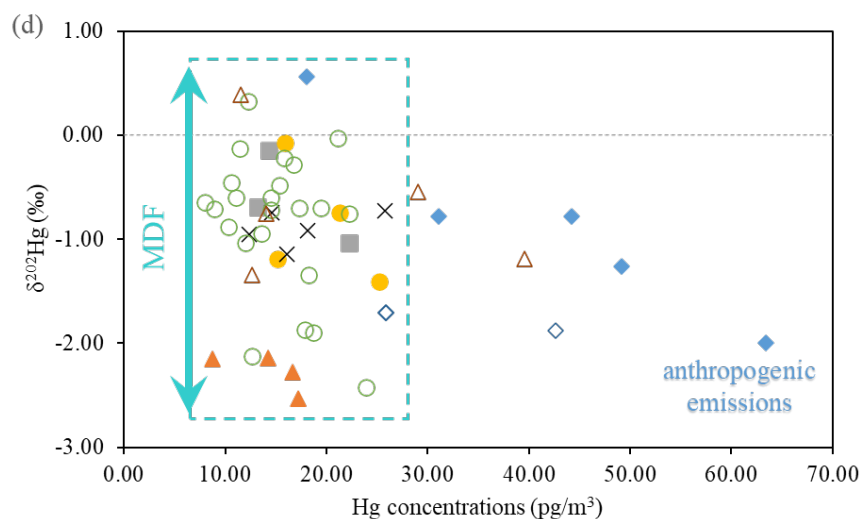
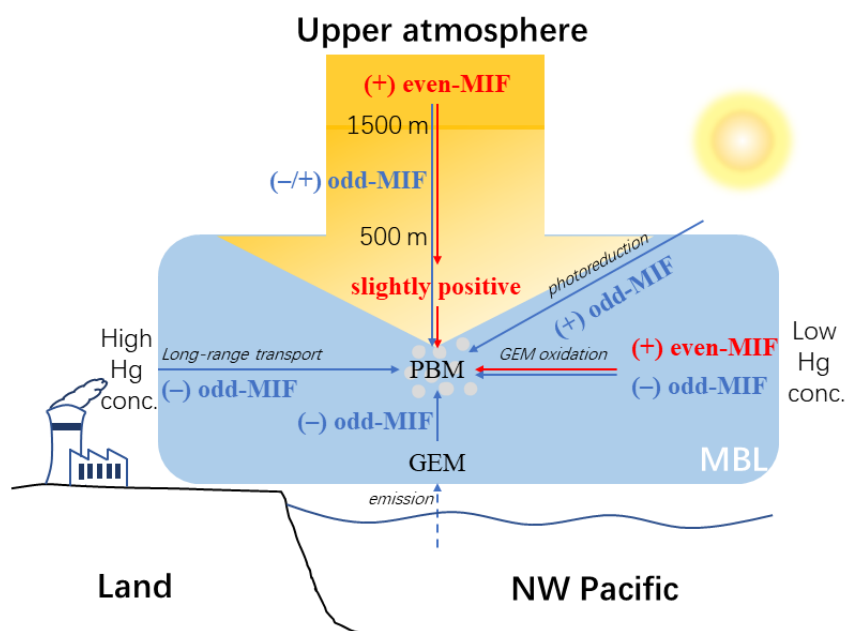


Figure 4. (a) Plot of  $\Delta^{199}\text{Hg}$  vs. Hg concentration for PBM samples. (b) Plot of  $\Delta^{199}\text{Hg}$  vs. Hg concentration for negative  $\Delta^{199}\text{Hg}$  values of PBM during the different cruises. (c) Plot of  $\Delta^{199}\text{Hg}$  vs.  $\delta^{202}\text{Hg}$  for PBM samples. (d) Plot of  $\delta^{202}\text{Hg}$  vs. Hg concentration for PBM samples. The error bar represents  $\pm 2\text{SD}$  for Hg concentrations. The typical 2SD analytic uncertainty of PBM samples was 0.22‰ to 0.42‰ and 0.07‰ to 0.13‰ for  $\delta^{202}\text{Hg}$  and  $\Delta^{199}\text{Hg}$ , respectively, and listed in **Table S2**.



769 Figure 5. Schematic diagram of possible Hg sources and transformation  
 770 processes for PBM over the MBL.

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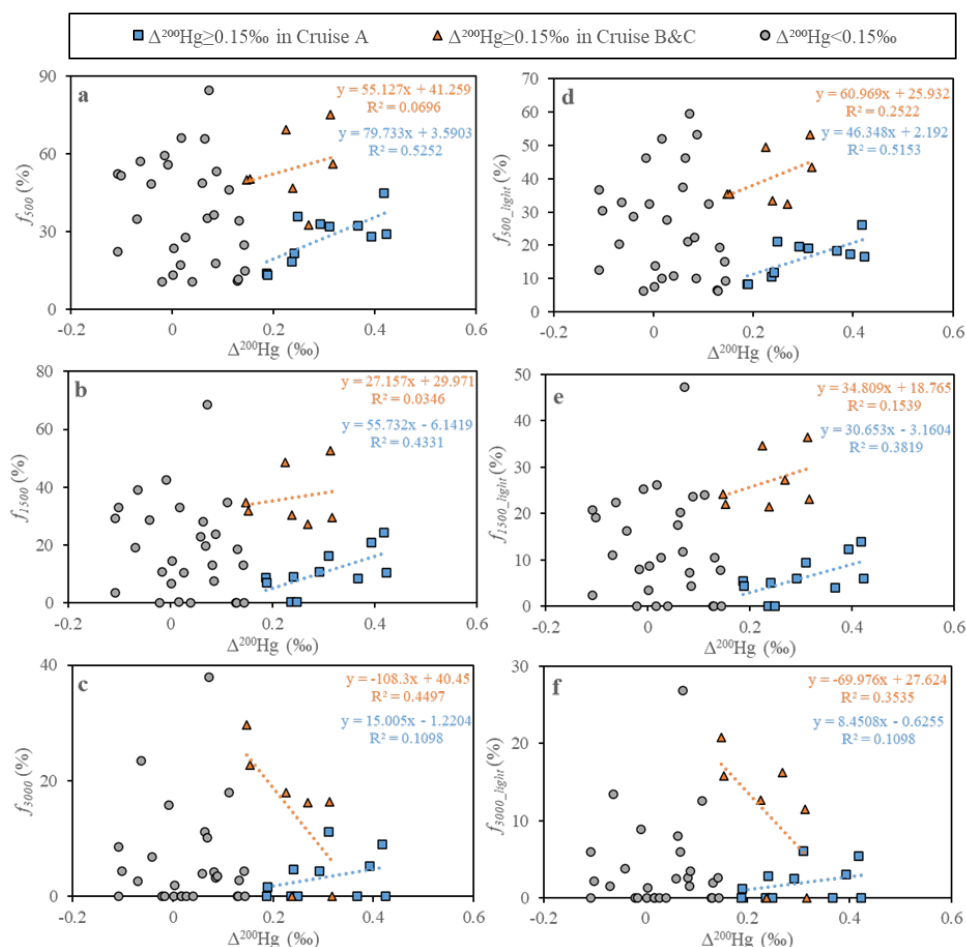


Figure 6. Plots of (a)  $f_{500}$ , (b)  $f_{1500}$ , (c)  $f_{3000}$ , (d)  $f_{500\_light}$ , (e)  $f_{1500\_light}$ , and (f)  $f_{3000\_light}$  vs.  $\Delta^{200}\text{Hg}$ . The square and triangle represent PBM samples during Cruise A and Cruises B&C, respectively. The circles represent PBM samples with  $\Delta^{200}\text{Hg}$  values lower than 0.15‰.

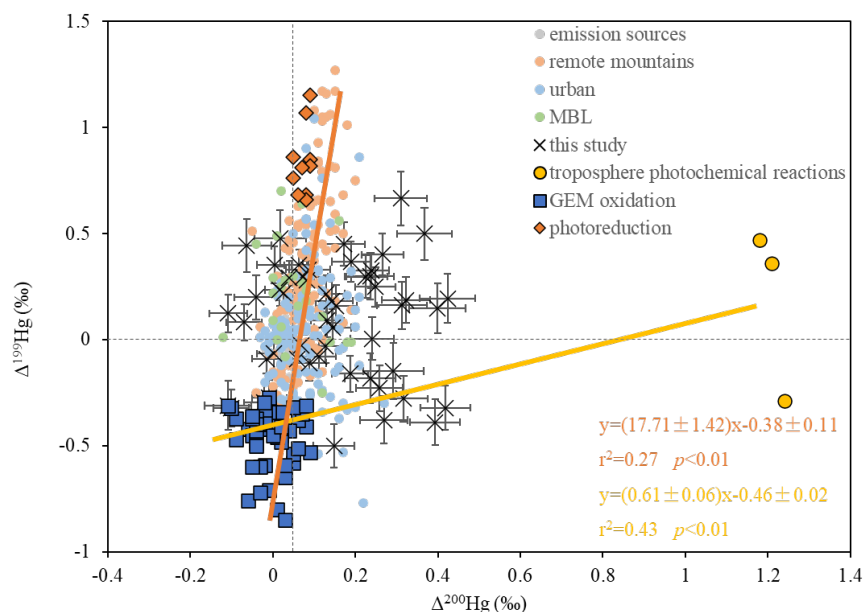


Figure 7. Plot of  $\Delta^{199}\text{Hg}$  vs.  $\Delta^{200}\text{Hg}$  for PBM samples collected from emission sources<sup>5,39,40</sup>, remote mountains<sup>23,24,66</sup>, urban<sup>19-23,38</sup>, MBL<sup>17,27</sup>, and this study. The three end-members of troposphere photochemical reactions, GEM oxidation, and photoreduction are compiled from the literature (**Table S5**). The criteria for the selection of the published Hg isotope data are following: samples (mainly in remotd mountains) with highly positive  $\Delta^{199}\text{Hg}$  ( $>0.66\text{‰}$ ) and near-zero  $\Delta^{200}\text{Hg}$  values are supposed to be dominated by photoreduction<sup>24</sup>; samples (mainly in precipitation) with highly positive  $\Delta^{200}\text{Hg}$  ( $>1.10\text{‰}$ ) values are supposed to be dominated by troposphere photochemical reactions<sup>8</sup>; samples (mainly in polar regions) with negative  $\Delta^{199}\text{Hg}$  ( $<-0.30\text{‰}$ ) and near-zero  $\Delta^{200}\text{Hg}$  values are supposed to be dominated by GEM oxidation<sup>11,13,20,21,27</sup>.



789 Error bars represent the 2SD uncertainties of individual samples.