# Analysis of CO<sub>2</sub>, CH<sub>4</sub> and CO surface and column concentrations observed at Reunion Island by assessing WRF-Chem simulations

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Abstract. Reunion Island is situated in the Indian Ocean and holds one of the very few atmospheric observatories in the tropical Southern Hemisphere. Moreover, it hosts experiments providing both ground-based surface and column observations of  $CO_2$ ,  $CH_4$  and CO atmospheric concentrations. This work presents a comprehensive study of these observations made in the capital Saint-Denis and at the high-altitude Maïdo Observatory. We used simulations of the Weather Research and Forecasting

5 model coupled with Chemistry (WRF-Chem), in its passive tracer option (WRF-GHG), to gain more insight in the factors that determine the observed concentrations. Additionally, this study provides an evaluation of the WRF-GHG performance in a region of the globe where it has not yet been applied.

A comparison of the basic meteorological fields near the surface and along atmospheric profiles showed that WRF-GHG has decent skill in reproducing these meteorological measurements, especially temperature. Furthermore, a distinct diurnal  $CO_2$ 

- 10 cycle with values up to 450 ppm was found near the surface in Saint-Denis, driven by local anthropogenic emissions, boundary layer dynamics and accumulation due to low wind speed at night. Due to an overestimation of local wind speed, WRF-GHG underestimates this nocturnal buildup. At Maïdo, a similar diurnal cycle is found but with much smaller amplitude. There, surface CO<sub>2</sub> is essentially driven by the surrounding vegetation. The hourly column-averaged mole fractions of CO<sub>2</sub> (XCO<sub>2</sub>) of WRF-GHG and the corresponding TCCON observations were highly correlated with a Pearson correlation coefficient of 0.90.
- 15 These observations represent different air masses than those near the surface, they are influenced by processes from Madagascar, Africa and further away. The model shows contributions from fires during the Southern Hemisphere biomass burning season, but also biogenic enhancements associated with the dry season. Due to a seasonal bias in the boundary conditions, WRF-GHG fails to accurately reproduce the CH<sub>4</sub> observations at Reunion Island. Further, local anthropogenic fluxes are the largest source influencing the surface CH<sub>4</sub> observations. However, these are likely overestimated. Further, WRF-GHG is ca-
- 20 pable of simulating CO levels on Reunion Island with a high precision. As to the observed CO column (XCO), we confirmed that biomass burning plumes from Africa and elsewhere are important for explaining the observed variability. The in situ observations at the Maïdo Observatory can characterize both anthropogenic signals from the coastal regions and biomass burning

enhancements from afar. Finally, we found that a high model resolution of 2 km is needed to accurately represent the surface observations. At Maïdo an even higher resolution might be needed because of the complex topography and local wind patterns.

25 To simulate the column FTIR observations on the other hand, a model resolution of 50 km might already be sufficient.

#### 1 Introduction

Major greenhouse gases such as carbon dioxide (CO<sub>2</sub>) and methane (CH<sub>4</sub>) have a direct impact on the radiative forcing of the atmosphere. They are the main drivers of climate change, since their global mean concentrations have increased over the industrial era by about 47 % and 156 %, for CO<sub>2</sub> and CH<sub>4</sub> respectively, as a result of human activities (Masson-Delmotte

30 et al., 2021). Carbon monoxide (CO) on the other hand is not a greenhouse gas but indirectly affects the lifetime of  $CH_4$  in the atmosphere through its competing reaction with OH. Additionally it plays a major role in air pollution as it participates in the formation of tropospheric ozone and urban smog.

The importance of these gases, hereafter all referred to as greenhouse gases (GHG), has led to the establishment of global observation networks to monitor their trends and variability. Ground-based remote sensing networks such as the Network for

- 35 the Detection of Atmospheric Composition Change (NDACC) and the Total Carbon Column Observing Network (TCCON) are known for their long time series of accurate column observations (De Mazière et al., 2018; Wunch et al., 2011). The Fourier Transform infrared (FTIR) spectrometer observations carried out in these networks use direct sunlight to measure the absorption of atmospheric trace gases along the line-of-sight and provide precise information on the total column abundance or vertical profile of GHG and other species. They are used by scientists worldwide to detect changes in the atmospheric composition,
- 40 to improve our understanding of the carbon cycle, or to provide validation for space-based measurements. Recently, this kind of observations from mobile low-cost FTIR spectrometers within the Collaborative Carbon Column Observing Network (COCCON) has been used to constrain fluxes in urban regions (Hase et al., 2015; Vogel et al., 2019; Makarova et al., 2021). In addition to FTIR observations, surface in situ observations of these gases are carried out to better constrain sources and sinks on an even smaller scale. Both observation types contain valuable information on the emissions and transport of these species
- 45 and are complementary.

Reunion Island (55 °E, 21 °S) is a French island in the Indian Ocean, situated about 550 km east of Madagascar. It hosts one of the very few atmospheric observatories in the tropical Southern Hemisphere, which provides both ground-based in situ and FTIR observations of GHG, contributing to the Integrated Carbon Observation System (ICOS) and NDACC and TCCON, respectively. GHG observations at Reunion Island are made at two sites: in the capital Saint-Denis and at the high-

- <sup>50</sup> altitude Maïdo Observatory (Baray et al., 2013). Several studies already investigated the factors influencing the observations at Reunion Island. Zhou et al. (2018) analyzed the trends and seasonal cycles of  $CH_4$  and CO by comparing the groundbased remote sensing and in situ observations. They noticed a distinct seasonal cycle in the column-averaged dry-air mole fractions of CO (XCO) with peak values between September and November, linked to the biomass burning season in Africa and South America, which confirmed earlier work from Duflot et al. (2010). Furthermore, backward trajectory simulations
- 55 revealed different origins of air masses observed at Reunion Island near the surface and higher up, resulting in surface CO

concentrations that are systematically lower than XCO. Near the surface, air masses generally originate in the Indian Ocean, while higher up the air comes from Africa and South America. The ability to detect biomass burning plumes at Reunion Island was also reported by Vigouroux et al. (2012). The available  $XCO_2$  timeseries has however not yet been investigated. Additionally, the Maïdo Observatory hosts a wide range of instruments of which the measurements have already been used by

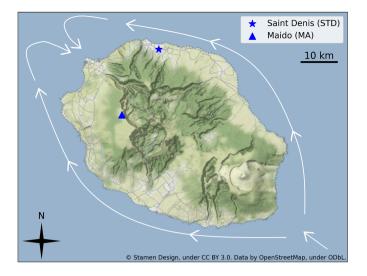
a variety of scientists to characterize the processes that occur at this particular location (Guilpart et al., 2017; Foucart et al., 2018; Duflot et al., 2019; Verreyken et al., 2021). However, the in situ observations at the Maïdo Observatory of the longer-lived species  $CO_2$  and  $CH_4$  have not yet been studied in detail, and this applies also to the available surface measurements at Saint-Denis.

Therefore the aim of the current work is to make a comprehensive description and analysis of in situ and column observations

- of  $CO_2$ ,  $CH_4$  and CO at Reunion Island, both at Saint-Denis and Maïdo. To gain more insight into the factors that influence the observed concentrations, we will rely on simulations of the widely used Weather Research and Forecasting model coupled with chemistry (WRF-Chem; Skamarock et al. (2021)), in its passive tracer option called WRF-GHG (Beck et al., 2013). This regional atmospheric model simulates 4D fields of  $CO_2$ ,  $CH_4$  and CO, resulting from their sources, sinks and transport in the troposphere, without interaction with other species, while accounting for the meteorology. It makes it possible to separate
- reach chemical compound into several tracers representing the contributions of different emissions sources within the model domain such as anthropogenic, biogenic, biomass burning and oceanic. Moreover it supports the online calculation of biogenic CO<sub>2</sub> fluxes following the Vegetation Photosynthesis and Respiration Model (VPRM; Mahadevan et al. (2008)). Thus far, applications of WRF-GHG have mainly focused on CO<sub>2</sub> to study city emissions (Pillai et al., 2016; Feng et al., 2016; Park et al., 2018; Zhao et al., 2019), or to evaluate the VPRM model (Ahmadov et al., 2007; Jamroensan, 2013; Dayalu et al., 2018;
- 75 Hu et al., 2020; Park et al., 2020). It has also been used in combination with in situ and column observations, flux towers and satellite measurements to better understand the carbon cycle (Pillai et al., 2010, 2012; Liu et al., 2018; Li et al., 2020). The model was shown to be an excellent tool for studying regional carbon budgets and is therefore very well suited to our needs. Few studies have used WRF-GHG to simulate CH<sub>4</sub> and CO, and these studies focused on explaining enhancements identified by satellite instruments (Beck et al., 2013; Dekker et al., 2017; Tsivlidou, 2018; Borsdorff et al., 2019; Dekker et al., 2019;
- 80 Verkaik, 2019). Hence, this work additionally aims at evaluating the model performance for these species in a region where it has not yet been applied. This might potentially draw attention to shortcomings in the model, allowing and motivating the model community to improve it.

This manuscript focuses on the factors that influence the observed GHG concentrations and their variations at Saint-Denis and the Maïdo Observatory. In particular, it addresses the following questions: (1) To what extent are the observations influenced

- by local and nearby sources and sinks, or long-range transport of emitted gases? (2) What are the different contributions (of anthropogenic, biogenic, biomass burning, ocean fluxes) to the observed concentrations, both at the surface and in the total column? (3) How accurate is WRF-GHG in simulating the different observation types of the three gases (CO<sub>2</sub>, CH<sub>4</sub> and CO) in the Southern Indian Ocean region, in particular at Saint-Denis and at the Maïdo Observatory? What are its strengths and weaknesses?
- 90 The structure of this document is as follows. Section 2 describes the location of the observation sites at Reunion Island, the



**Figure 1.** Map of Reunion Island, indicating the location of the two measurement sites: Saint-Denis (star) and Maïdo (triangle). The white arrows roughly illustrate the local wind patterns, generated by the trade winds and the orography of the island.

general transport patterns, the GHG-measuring instruments and the data sets used in this study. Details on the model set-up and input inventories are described in Sect. 3. Section 4 constitutes the main part of this manuscript. First, the model performance is evaluated with regard to meteorological fields, both at the surface and higher up, in subsection 4.1. The model assessment and data analysis at Saint-Denis and Maïdo are discussed in subsection 4.2 and 4.3 respectively. Finally, the impact of model resolution is discussed in Sect. 5 and conclusions are drawn in Sect. 6.

#### 2 Observations at Reunion Island

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The data used in this study are coming from two observation sites on Reunion Island: Saint-Denis (referred to as STD from now on; 20.9014 ° S, 55.4848 ° E; 85 m above sea level) which is the capital city and is situated close to the northern coast, and the Maïdo Observatory (referred to as MA from now on; 21.0796 ° S, 55.3841 ° E; 2155 m a.s.l.) which is close to the 100 top of a mountain ridge on the northwest side of the island. Currently, each site is equipped with a Fourier Transform InfraRed (FTIR) spectroscopy instrument and an in situ cavity ring-down spectroscopy (CRDS) analyzer, which are described more in detail below. These instruments measure the column-averaged dry-air mole fractions and local near-surface mole fractions, respectively. The locations of both sites on the island are shown in Fig. 1.

#### 105 2.1 Climate and transport patterns

The atmospheric transport around Reunion Island is controlled by the position of the Intertropical convergence zone (ITCZ) and the south Hadley cell (Baldy et al., 1996; Foucart et al., 2018). During a large part of the year, a strong subtropical high induces

steady southeasterly trade winds near the surface and westerlies aloft. Hence, the air above Reunion Island is characterized by a wind (and temperature) inversion causing generally clear skies which are common during the dry season (Baldy et al., 1996;

110 Lesouëf et al., 2011; Baray et al., 2013). Typically, this (colder) dry season lasts from May to November (Foucart et al., 2018).
In Austral summer (January to March) the ICTZ moves South, sometimes reaching Reunion Island. This results in weaker trade winds and often heavy rains, resulting in the (warmer) wet season in those months (Baldy et al., 1996; Foucart et al., 2018).

With its high altitudes (up to 3 000 m above sea level), Reunion Island represents a sudden obstacle for the stable southeasterly trade winds. In combination with the inversion laver, this causes a blocking on the windward side and wind flow splitting (and

- 115 trade winds. In combination with the inversion layer, this causes a blocking on the windward side and wind flow splitting (and accelerating) around the island, to form counter-flowing vortices on the northwestern (lee) side (Lesouëf et al., 2011). This is illustrated by the white arrows in Fig. 1. Moreover, the split flow is under influence of thermally driven circulations (so-called trade breathing): nighttime downslope and land-breezes push the trade wind offshore whereas daytime upslope and sea-breezes allow the wind to pass over coastal areas (Lesouëf et al., 2011; Lesouëf et al., 2013). These circulations are the dominant (daily)
- 120 wind pattern on the northwestern side of Reunion Island (where MA is situated) which is sheltered of the trade winds (Lesouëf et al., 2011; Baray et al., 2013; Guilpart et al., 2017; Verreyken et al., 2021).

#### 2.2 Saint-Denis

Saint-Denis is the capital of Reunion Island, located by the coast in the Northern part of the island. As of 2018, there were 309 635 inhabitants in the metropolitan area of Saint-Denis, with a population density of about 1 100 per square kilometer. The city

- 125 lies on a slope between the ocean and the nature reserve of La Roche Écrite (ultimately reaching a top of 2 276 m). The observations at STD are made on top of a building at the University of Reunion Island (85 m above sea level). In situ mole fractions of  $CO_2$  and  $CH_4$  are measured by a CRDS analyzer (Picarro G1301) since August 2010, in collaboration with the Laboratoire de l'Atmosphère et des Cyclones (LACy), the Observatoire des Sciences de l'Univers de la Réunion (OSU-R) and the Laboratoire des Sciences du Climat et de l'Environnement (LSCE). The measurements are available with a time frequency
- 130 of 1 min and the uncertainty on the measured mole fractions are about 0.1 ppm and 2 ppb, for  $CO_2$  and  $CH_4$  respectively. In September 2011, BIRA-IASB installed a high-resolution Bruker IFS 125HR FTIR at STD, next to the PICARRO analyzer. This instrument is primarily dedicated to measuring the near-infrared (NIR: 4000 - 16000 cm<sup>-1</sup>) spectra and contributes to TCCON (Wunch et al., 2011). The solar spectra are used to retrieve the total column-averaged dry-air mole fraction of  $CO_2$ ,  $CH_4$  and CO (De Maziere et al., 2017). The standard TCCON retrieval algorithm, called GGG2014, applies a profile scaling,
- 135 therefore deriving information on the total column only and not on the vertical profile. TCCON measurements have been calibrated to WMO standards, so it is assumed that there are no systematic biases compared to in situ measurements (Wunch et al., 2010). More detail on both instruments can be found in Zhou et al. (2018).

#### 2.3 Maïdo

The Maïdo Observatory (2155 m a.s.l.) is located close to the summit of a mountain with the same name which has an altitude of about 2 200 m.a.s.l. and is situated in the western part of the island. The Observatory is devoted to long-term atmospheric monitoring in the tropical region of the Southern Hemisphere and houses a variety of atmospheric measurement instruments such as lidar systems, spectro-radiometers, and in situ gas and aerosol analyzers (Baray et al., 2013). To the west of MA is a gentle slope reaching the coastal areas and the ocean, while the summit lies to the east of the site, followed by a cliff leading to the caldera of Cirque de Mafate. The area around MA is covered by mountain shrubs and heathlands (Duflot et al., 2019).

145 The mole fractions of all three gases (CO<sub>2</sub>, CH<sub>4</sub> and CO) have been collected by a CRDS analyzer (PICARRO G2401) at MA since December 2014 and have been certified as Integrated Carbon Observation System (ICOS) atmospheric data in late 2019 (De Mazière et al., 2021). The measurements are available at a time resolution of 1 min and the uncertainties are about 50 ppb, 1 ppb and 2 ppb for CO<sub>2</sub>, CH<sub>4</sub> and CO, respectively.

In March 2013, BIRA-IASB started operating a second Bruker IFS 125HR FTIR spectrometer, besides the one at STD, but observing the solar spectra in the mid-infrared (MIR) range from 600 to  $4500 \text{ cm}^{-1}$  (Baray et al., 2013). These FTIR measurements are affiliated with NDACC. Gas mole fractions of CH<sub>4</sub> and CO are retrieved from the FTIR solar spectra by the SFIT4 algorithm which is based on the optimal estimation method of Rodgers (2000). More information about the specific methods used can be found in Zhou et al. (2018). The final data consist of the retrieved vertical profiles, expressed as volume mixing ratio (VMR) profiles on a vertical altitude grid.

#### 155 2.4 Meteorological measurements

The quality of the WRF-GHG simulations is evaluated against meteorological fields that are being measured in parallel at the two observation sites. More specifically, there are in situ measurements of 2 m temperature, 10 m wind direction and wind speed. These fields are measured by the Vaisala Weather Transmitter (model WXT510 at STD and model WXT520 at MA) every 3 s.

160 Additionally we will compare the WRF-GHG output with vertical profiles from operational daily meteorological Meteomodem M10 radiosonde launches performed by Météo-France at 12:00UTC at the Gillot airport (4 km away from STD). The Meteomodem M10 radiosondes provide measurements of temperature, pressure and relative humidity with respect to water and zonal and meridional winds. Detailed description of this sensor can be found in Dupont et al. (2020).

#### 3 WRF-GHG model

- 165 WRF-GHG is an abbreviation for the Weather Research and Forecast model coupled with Chemistry (WRF-Chem) in its passive tracer option (Skamarock et al., 2021; Beck et al., 2011). WRF-Chem simulates the emission, transport, mixing, and chemical transformation of trace gases and aerosols simultaneously with the meteorology. In WRF-GHG only CO<sub>2</sub>, CO and CH<sub>4</sub> are transported and there are no chemical reactions simulated. Separate tracers for each compound represent the contribution from the fluxes within the model domains (d01-d03) from different categories: anthropogenic, biomass burning,
- 170 biogenic (for  $CO_2$  and  $CH_4$  (termites)), ocean (for  $CO_2$  only) and wetlands (for  $CH_4$  only). Additionally, there is a so-called background tracer which represents the contribution of the initial and lateral boundary conditions. The sum of all tracers for a species is equal to the total modeled mole fractions. In this study, WRF-Chem version 4.1.5 is used.

Two time periods have been simulated: from 1 August 2015 until 1 May 2016 and from 1 July 2016 until 15 July 2017. These periods have been selected because then, quite complete datasets are available from all considered instruments. The first 14

175 days in each period are regarded as spin-up period and are not used in the model-data comparisons. The model provides 3D

## 3.1 Emissions, initial and boundary conditions

fields of CO<sub>2</sub>, CH<sub>4</sub>, CO and meteorological fields every hour.

An overview of the data that are used as input to the WRF-GHG model are given in Table 1 and described hereafter. The hourly meteorological initial and lateral boundary conditions (IC-BCs) are obtained from the European Centre for Medium-Range
180 Weather Forecasts (ECMWF) global ERA5 reanalysis dataset (0.25 °x 0.25 °) (Hersbach et al., 2018a, b), while the chemical IC-BCs are imported from the CAMS global reanalysis for greenhouse gases (EGG4, for CO<sub>2</sub> and CH<sub>4</sub>) and reactive gases (EAC4, for CO, Inness et al. (2019)). The data for CO<sub>2</sub> and CH<sub>4</sub> are available every 3 h, while for CO every 6 h. These fields from the CAMS reanalysis are used to drive the background tracers. The IC-BCs of the tracers corresponding with the contribution from surface fluxes are set to zero.

- The anthropogenic emissions for  $CH_4$  and CO are taken from the Emission Database for Global Atmospheric Research (EDGAR): v5.0 Global Greenhouse Emissions product (Crippa et al., 2019b) for  $CH_4$  and v5.0 Global Air Pollutant Emissions product (Crippa et al., 2019a) for CO. Further, we performed simulations over a short period of a couple days to test alternative inventories for anthropogenic  $CO_2$  and CO fluxes. We concluded that the Open-Data Inventory for Anthropogenic Carbon dioxide (ODIAC2020, Oda and Maksyuto (2015, 2011); Oda et al. (2018)) was more representative for the anthropogenic  $CO_2$
- 190 emissions, probably due to its much higher spatial resolution (1 km) compared to EDGAR (0.1 °). Similarly, we use a CO surface emission inventory at a resolution of 500 m, based on the posterior estimates of a mesoscale inverse model (Jérôme Brioude, personal communication), but only in the innermost domain d03. The atmospheric transport of the inverse model was calculated using the FLEXPART Lagrangian dispersion model (Verreyken et al., 2019) coupled with the MESO-NH mesoscale model (Lac et al., 2018) at a resolution of 500 m and 60 vertical levels. FLEXPART-MESO-NH was
- 195 run backward in time to calculate the source-receptor relationships between MA and the surface sources from the CO measurements at MA from April 4th to May 3rd 2019, during the BIO-MAIDO campaign (Dominutti et al., 2022). The ODIAC CO<sub>2</sub> emission inventory was used as a priori to benefit from its native spatial resolution of major urban areas. A scaling factor, based on the ratio between the mean CO enhancement above background and mean CO<sub>2</sub> enhancement above background, was applied on the CO<sub>2</sub> fluxes to obtain a priori surface CO fluxes. A temporal resolution of one hour was used for the observed
- 200 and simulated CO mixing ratios at MA. A lognormal distribution was assumed for the observation and surface flux errors (Brioude et al., 2012, 2013). Such assumption better matches the CO distribution in the atmosphere, and prevent the inversion to calculate negative fluxes.

The anthropogenic fluxes used within WRF-GHG are combined with a temporal emission factor from Nassar et al. (2013). Note that these factors are representative for  $CO_2$  and might be less accurate for CO and  $CH_4$ .

205 Daily biomass burning emissions for all three gases are obtained from the Fire INventory from NCAR (FINN v1.5) (Wiedinmyer et al., 2011). The biogenic  $CH_4$  flux from wetlands is obtained from the WetCHARTs v1.0 ensemble (Bloom et al., 2017), while the biogenic  $CO_2$  flux from oceans is taken from the observation-based global monthly gridded sea surface p $CO_2$ climatology by Landschützer et al. (2017), which also provides air-sea CO<sub>2</sub> fluxes. Finally, the biogenic CO<sub>2</sub> flux from the vegetation is simulated online using the VPRM model (Mahadevan et al., 2008; Ahmadov et al., 2007). This model uses the

- 210 2 m temperature and downward shortwave radiation calculated by WRF-GHG in combination with surface reflectance data from the Moderate Resolution Imaging Spectroradiometer (MODIS). Further, it uses the global SYNMAP land cover data of 1 km resolution by Jung et al. (2006). Additionally, the VPRM requires a set of four model parameters for each vegetation class, dependent on the region of interest. Ideally, these parameters are optimized using a network of eddy flux towers. Since this is not available at Reunion Island, we use the set of parameters optimized by Botía et al. (2021), based on measurements from nine sites in the Amazon region in Brazil, created in the context of the Large Scale Biosphere-Atmosphere Experiment
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	Species	Source	Time and spatial resolution
Initial and lateral	$CO_2, CH_4$	CAMS reanalysis for greenhouse gases	3-hourly, 0.75 $^\circ$
boundary conditions	СО	CAMS reanalysis for reactive gases	6-hourly, 0.75 $^\circ$
Anthropogenic flux	$CO_2$	ODIAC2020	Monthly, 1 km (land), 1 °(ocean)
(multiplied with	$CH_4$	EDGAR v5.0 Global Greenhouse Emissions	Yearly, 0.1 $^{\circ}$
temporal factors of	СО	EDGAR v5.0 Global Air Pollutant Emissions (d01,	Yearly, 0.1 $^\circ$
Nassar et al. (2013))		d02)	
		Brioude (d03)	Yearly, 500 m
Biomass burning flux	$CO_2, CH_4, CO$	FINN v1.5	Daily, 1 km
Biogenic flux	CO <sub>2</sub>	online (VPRM)	Hourly, model resolution
	$CH_4$	online, WetCHARTs v1.0	Monthly, 0.5 $^\circ$
Ocean flux	$CO_2$	Observation-based global monthly gridded sea surface	Monthly, 1 $^{\circ}$
		pCO <sub>2</sub> climatology	

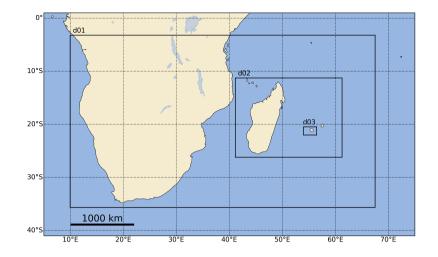
Table 1. Overview of data sets used as input for the WRF-GHG simulations.

(LBA-ECO). Exact parameter values are given in Table A1 of appendix A.

#### 3.2 Settings

To achieve a high-resolution model grid over Reunion Island, a configuration of three nested domains was established, going from a larger domain with lower resolution to a smaller domain with higher resolution. The domains are shown in Fig. 2. Their

- 220 respective resolutions are 50 km, 10 km and 2 km. The innermost domain d03 covers Reunion Island and the two measurement sites completely. WRF-GHG uses a hybrid vertical coordinate, which is a coordinate that is terrain following near the ground and becomes isobaric higher up. In all our domains the model has 60 vertical levels extending from the surface up to 50 hPa. The following physical parameterization options are used: the Morrison 2-moment Scheme (Morrison et al., 2009) for micro physics, the RRTMG Shortwave and Longwave Schemes (Iacono et al., 2008) for shortwave and longwave radiation. The Eta
- 225 Similarity Scheme (Janjić, 1994) for surface layer processes and the Unified Noah Land Surface Model (Tewari et al., 2004) for



**Figure 2.** Location of the WRF-GHG domains, with horizontal resolutions of 50 km (d01), 10 km (d02) and 2 km (d03). All domains have 60 (hybrid) vertical levels extending from the surface up to 50 hPa.

the land surface. To choose between the diverse parameterization schemes for cumulus parameterization and planetary boundary layer (PBL) physics, several model test runs were made for a short simulation period of a couple of days and compared with the observed meteorology. As a result, the University of Washington (TKE) Boundary Layer Scheme (Bretherton and Park, 2009) for PBL physics and the Grell–Freitas Ensemble Scheme (Grell and Freitas, 2014) for cumulus parameterization, but only in the largest domain (d01), were chosen for this study.

## 3.3 Data handling

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The various observation types are dealt with in different ways for comparison with the model.

The surface observations (both meteorological fields and GHGs) are averaged over a period of 30 min around the hourly model time step. At STD, we compare these data with the lowermost level of the model grid cell whose center is closest to the location of the instrument. Because of the complex topography, the cell covering MA is less representative for the Observatory as its

center is located behind the summit, below in the caldera of Cirque de Mafate. Model-data comparisons of the surrounding cells showed that the cell to the West of MA is more representative. Therefore, this alternative model grid cell, of which the center is only 1.3 km away from the Observatory, is used in the analysis.

In order to compare similar quantities, the total-column-averaged dry-air mole fractions from TCCON and NDACC are truncated to the same atmospheric column that is simulated by WRF-GHG, e.g. from surface up to 50 hPa. This is needed because the FTIR data represent the total atmospheric column whereas the WRF-GHG upper limit lies at around 21 km.

As NDACC additionally provides volume mixing ratio profiles, the column-averaged mole fractions are re-calculated taking only those layers below the model upper limit. For TCCON, only information on the total column is retrieved. Therefore, we multiply the TCCON data with a factor representing the ratio between the column-averaged mole fraction of the smaller column

- 245 (up to 50 hPa) to that of the total column. This ratio is calculated from the a priori information. In the rest of this manuscript, all dry-air column-averaged mole fractions (so-called Xgas) mentioned in this article refer to this reduced atmospheric column only (surface up to 50 hPa). Due to the specific profile of the respective gases in the atmosphere, this scaling is more significant for XCH<sub>4</sub>, than it is for XCO or XCO<sub>2</sub>: the values generally increase after scaling by about 27-35 ppb for XCH<sub>4</sub>, 3 ppb for XCO and 0.25 ppm for XCO<sub>2</sub>.
- 250 To compare with the hourly WRF-GHG outputs, the scaled mole fractions are averaged over a period of 30 min around the model time step. Furthermore, a smoothing is applied to the WRF-GHG profiles, according to Rodgers and Connor (2003). Because of the different characteristics of the TCCON and NDACC observing systems, this smoothing procedure is slightly different at the two sites. Technical details on how the smoothed dry-air column-averaged mole fractions of WRF-GHG are calculated at STD and MA can be found in Appendix B.

#### 255 4 Results

#### 4.1 Meteorological evaluation

#### 4.1.1 Surface measurements

To assess the general model performance, the hourly model output of d03 near the surface is compared with local measurements of 2 m temperature and 10 m wind direction and speed at both sites. Table 2 gives the root mean square error (RMSE), mean bias error (MBE) and Pearson correlation coefficient (CORR) of the model-data comparison over the complete time series (13 583 paired data points at STD, 14 031 at MA).

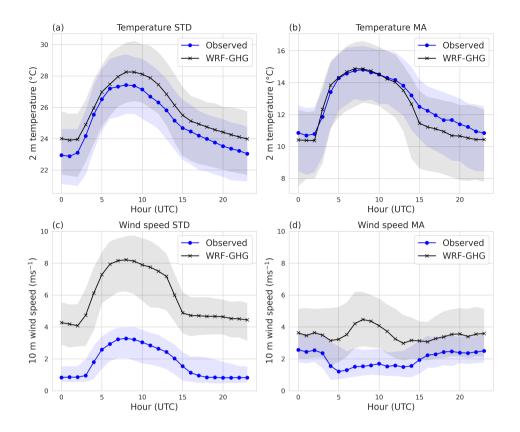
		2 m temperature (C) STD MA		10 m wir	d direction ( $^{\circ}$ )	$10 \text{ m wind speed } (\text{ms}^{-1})$		
				STD	MA	STD	MA	
RI	MSE	1.33	1.94	52.33	66.80	4.29	2.93	
N	1BE	0.74	-0.35	/	/	3.83	1.59	
C	ORR	0.93	0.83	0.72	0.76	0.73	0.27	

Table 2. Overview of the meteorological evaluation of surface measurements at the two sites.

The 2 m temperature is well simulated by WRF-GHG at both sites with very high correlation coefficients of 0.93 at STD and 0.83 at MA, and RMSE between 1 ° and 2 ° Celsius (1.33 at STD, 1.94 at MA). Figure 3a-b compares the median diurnal cycle at both sites, which is very well reproduced by the model. Overall, higher temperatures are measured at STD compared to MA because of the large difference in altitude between the sites (85 masl compared to 2 155 masl).

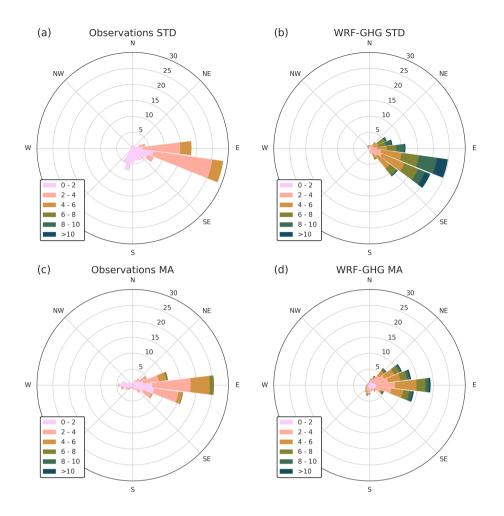
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The windroses in Fig. 4 show the most occurring 10 m wind direction and their corresponding wind speed. The 10 m wind direction of WRF-GHG correlates well with the measurements at both sites (correlation coefficients of 0.72 and 0.76). There is a larger error of the wind direction at MA (RMSE of 66.80 °) compared to STD (RMSE of 52.33 °). At STD, the wind is mainly from east or southeast (trade winds), however for calmer wind speeds (< 2 ms<sup>-1</sup>) the wind can also come from



**Figure 3.** (a)-(b) Diurnal cycle of the 2 m temperature and (c)-(d) 10 m wind speed at both STD and MA. The blue and black lines show the median values for every hour of the measurements and simulations, respectively. The shaded blue and grey areas indicate the corresponding interquartile ranges of the measurements and simulations, respectively. Hours are given in UTC (local time at Reunion Island is UTC+4). Remark that the temperature plots have different y-axes.

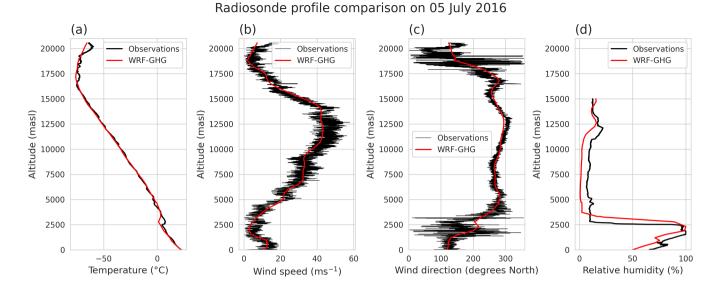
- 270 the south(west). WRF-GHG captures the dominant southeastern winds, but does not simulate winds from the south. It highly overestimates the wind speed, with a mean bias error of  $3.83 \text{ ms}^{-1}$  and a RMSE of  $4.29 \text{ ms}^{-1}$ . There is a clear diurnal cycle of the wind speed at STD, shown in Fig. 3c, with stronger winds during the day and calmer conditions at night. As WRF-GHG follows the observed pattern, the correlation coefficient is still quite high (0.76). The overestimation might be caused by an underestimation of the surface roughness of the city within WRF-GHG. Besides the unified NOAH land surface model (see
- 275 Sec. 3.2), no additional urban surface model was included in the simulations. Other studies using the WRF model often show wind speed overestimation above urban areas (Feng et al., 2016; Barlage et al., 2016; Zhang et al., 2009; Kim et al., 2013). Additionally, there is a large gradient in the surface wind speed near STD, caused by the presence of the strong trade winds. Therefore, an insufficient high model resolution might also be the cause for the wind speed overestimation.
- At MA on the other hand, the most common wind direction is east with some occurrences of west winds. This points to the typical thermally induced circulations during the day, whereby wind is driven from the coast upwards and sometimes reaches MA (Duflot et al., 2019). The prominent east winds illustrate the presence of overflowing trade winds. The simulated winds



**Figure 4.** Windrose from observations and WRF-GHG simulations at STD ((a)-(b)) and at MA ((c)-(d)). The colors indicate the associated wind speed (in  $ms^{-1}$ ), while the lengths of the bars show the frequency of any wind direction binned by 15°, given in percentage.

from WRF-GHG are mainly from east, indicating that the larger errors at MA might be linked to the missing western wind components. This is likely due to the complex topography around the Maïdo Observatory and the model resolution (of 2 km) which might be insufficient to resolve these very local wind dynamics.

As to the wind speed at MA, the bias and RMSE are smaller  $(1.59 \text{ and } 2.93 \text{ ms}^{-1}$ , respectively) than at STD but the model is still overestimating the wind speed. Moreover, the correlation is very low at this site. The daily 10 m wind speed cycle at MA is less distinct than at STD, however at night the wind is more often faster (> 2 ms^{-1}) than during the day. This could be linked with the local wind dynamics around MA where during the day calmer upslope winds from the west often reach MA while at night the Observatory is generally in the free troposphere under influence of the faster trade winds (Guilpart et al., 2017).



**Figure 5.** Example of the radiosonde data at Gillot airport on 5 July 2016, compared with the model for (a) temperature, (b) wind speed, (c) wind direction and (d) relative humidity. The black line represents the measured values. The red line is the corresponding WRF-GHG data.

#### 290 4.1.2 Radiosonde profile measurements

Daily radiosonde profiles of air temperature, wind direction, wind speed and relative humidity are compared with the model data to assess the accuracy of WRF-GHG on all levels of the troposphere. The profiles were matched as follows: for every data point measured by the radiosonde, the grid cell corresponding to its coordinate is selected. Next, the model profile (consisting of the meteorological field in the complete vertical column above the selected grid cell) is interpolated to the altitude of the measurement. This interpolated value is then paired with the value of the measurement. This results in a paired model-data

295 measurement. This interpolated value is then paired with the value of the measurement. This results in a paired model-data profile for the four variables once every day. An example of such a paired profile on 5 July 2016 is shown in Fig. 5. For every paired profile, the RMSE, MBE and CORR statistics are calculated. This is done for a total of 267 days in the year 2016.

For the temperature, correlation coefficients are very high on all days (median is 0.99). Moreover, the RMSE values are 300 quite small on most days (median is 1.07 °C) indicating that WRF-GHG can simulate the temperature in the troposphere quite accurately.

There is a good correlation for the wind direction and speed profiles as well. Half of the days in 2016 have a correlation coefficient higher than 0.87 for wind direction and 0.83 for wind speed. When calculating the RMSE of wind direction along the daily profiles, we find a median of  $48.18^{\circ}$  while the median RMSE of wind speed is  $3.87 \text{ ms}^{-1}$ . On most days, WRF-GHG

is slightly underestimating the wind speed (median bias error is  $-1 \text{ ms}^{-1}$ ), which is in contrast with the overestimation found at the surface sites.

The profiles of relative humidity are analysed up to an altitude of 15 km because the measurements are less accurate higher up.

WRF-GHG correlates well with these profiles, with a correlation coefficient higher than 0.87 on 75 % of the days in 2016. The median RMSE on the daily profiles of relative humidity is only 11.6 %, showing a decent model performance.

310 Overall we can conclude that the simulations of basic meteorological parameters are quite accurate along vertical profiles, where near the surface wind speed and direction agree less well with the observations.

#### 4.2 GHG data at Saint-Denis

At STD, the in situ surface mole fractions of  $CO_2$  and  $CH_4$  are measured together with the TCCON column-averaged dry-air mole fractions of CO<sub>2</sub>, CH<sub>4</sub> and CO. The comparison with the WRF-GHG simulations will be described in detail below, for each species and measurement type separately. Full time series of the observed and modeled data can be found in appendix C. An overview of the statistics of the comparisons is given in Table 3.

		STD		MA					
	(X)CO <sub>2</sub>	$(X)CH_4$	(X)CO	(X)CO <sub>2</sub>	$(X)CH_4$	(X)CO			
	(ppm)	(ppb)	(ppb)	(ppm)	(ppb)	(ppb)			
			in situ						
RMSE	9.17	18.51	/	1.95	19.33	10.99			
MBE	-5.39	9.04	/	-0.15	14.09	5.51			
CORR	0.62	0.35	/	0.75	0.30	0.83			
			FTIR						
RMSE	0.75	10.26	8.08	/	10.80	7.37			
MBE	-0.37	5.69	5.07	/	-5.65	1.81			
CORR	0.90	0.31	0.89	/	0.37	0.90			

Table 3. Overview of the WRF-GHG performance for hourly in situ and column observations of GHG at Reunion Island. Comparison with the column observations is based on the smoothed model profiles. There are no in situ CO data available at STD (Saint-Denis) and no XCO<sub>2</sub> data at MA (Maïdo).

### 4.2.1 Surface CO<sub>2</sub>

The model-data comparison of the surface data shows a moderate correlation coefficient of 0.62 together with a relative large error of 9.17 ppm and a model underestimation of 5.39 ppm. The scatter plot in Fig. 6a indicates that these discrepancies arise

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from a model underestimation of the higher  $CO_2$  mole fractions. The lower  $CO_2$  concentrations are in general much better reproduced.

The  $CO_2$  measurements at STD show a clear diurnal cycle (see Fig. 7a), with lower values during the day and higher values during the night. The diurnal cycle of WRF-GHG reproduces this pattern but with much lower nighttime concentrations, leading to the moderate correlation found in Table 3.

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As shown in the diurnal cycle in Fig. 7b, the main contributors to the total  $CO_2$  signal in WRF-GHG, in addition to the

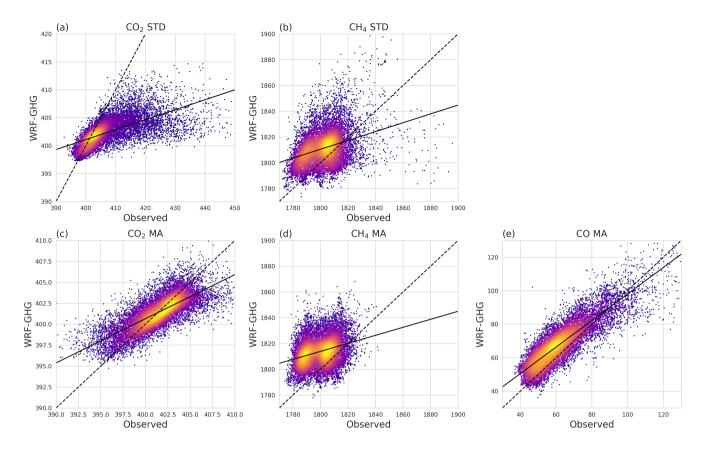


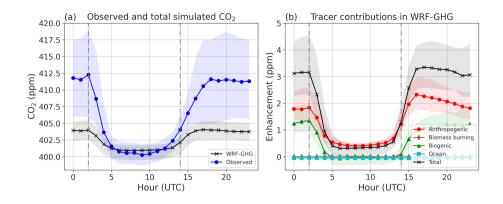
Figure 6. Scatter plot of hourly observed and modeled in situ greenhouse gases at STD (a,b) and MA (c,d,e). The colors indicate the point density.

background signal, are the anthropogenic and biogenic tracers. They correspond with anthropogenic and biogenic fluxes within the model domains (d01-d03) and show similar diurnal patterns with maxima at night and minima during the day. The influence of biomass burning or ocean fluxes are negligible at STD.

In urban areas, anthropogenic pollution is generally trapped in and around the city creating a so-called urban  $CO_2$  dome (Idso et al., 2002). The strength of this dome is primarily dependent on the local emissions and variations in the boundary layer. In calm weather, near-surface air temperature inversions at night trap anthropogenic pollution near the ground in the shallow nocturnal boundary layer, leading to strongly enhanced  $CO_2$  mixing ratios. During the day, solar radiation causes convective mixing of the air creating a deep planetary boundary layer (PBL). The near-surface  $CO_2$  concentrations are then diluted by this

However, wind speed and direction can alter the strength of this urban  $CO_2$  enhancement: at higher wind speeds (from 2 ms<sup>-1</sup>), ventilation processes prevent strong  $CO_2$  accumulation, while winds from rural areas could bring pristine air to the city (Idso et al., 2002; Rice and Bostrom, 2011; Massen and Beck, 2011; García et al., 2012; Xueref-Remy et al., 2018).

thorough mixing of air, and the urban dome extends to greater heights.

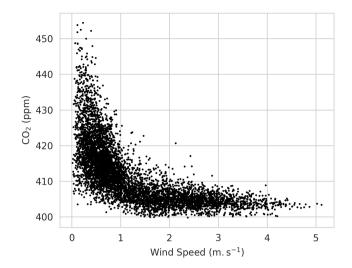


**Figure 7.** Diurnal cycle of (a) in situ CO<sub>2</sub> at STD (b) and model tracer contributions. The black line in (a) represents the median hourly concentrations of WRF-GHG, while the blue line represents the observed values. The shaded areas cover the interquartile ranges. The gray dotted vertical lines at 2 h and 14 h UTC indicate the approximate times of sunrise and sunset. The colored lines in (b) represent the different tracer deviations from the background concentration in WRF-GHG: anthropogenic (red), biogenic (green), ocean (cyan) and biomass burning (brown) tracer. The black line is the sum of all tracers, except the background.

Within WRF-GHG, the main contributors to the simulated CO<sub>2</sub> in the grid cells around STD are anthropogenic and peak during the day. The biogenic CO<sub>2</sub> flux at the grid cell of STD is zero because the corresponding VPRM vegetation class is 100 %
"Barren, urban and built-up": the model assumes that there is no vegetation within the city. Given that Saint-Denis is the capital city of Reunion Island and has plenty of anthropogenic activities, the impact of local vegetation is probably very small and these WRF-GHG fluxes appear realistic. The nighttime peak of CO<sub>2</sub> mixing ratios is therefore attributed to PBL dynamics and regional transport.

Figure 7a shows that the observed interquartile range at night is wide, indicating a large variability in the CO<sub>2</sub> accumulation.

- We find a negative correlation between 10 m wind speed and in situ CO<sub>2</sub> concentrations at night for the observations at STD (see Fig. 8). At low wind speeds, a large variability in CO<sub>2</sub> mixing ratios is observed: from 400 ppm up to 450 ppm. For wind speeds above  $2 \text{ ms}^{-1}$  on the other hand, values higher than 410 ppm are rarely found, which indicates that ventilation processes take place. At the same time, there is a large overestimation of the wind speed within WRF-GHG (Sec. 4.1.1, Table 2). At night, 91.7 % of the simulated hours has a wind speed of more than  $2 \text{ ms}^{-1}$ , compared to only 23.2 % of the nocturnal observations,
- 350 leading to a wind speed MBE of  $3.23 \text{ ms}^{-1}$  at night. Note that the mean wind speed at night within WRF-GHG is  $4.4 \text{ ms}^{-1}$  (see also Fig. 3c), while the nighttime CO<sub>2</sub> concentration in WRF-GHG is on average 403.8 ppm (see also Fig. 7a). Looking at Fig. 8, these values follow the pattern as found in the observations: a nocturnal wind speed of more than  $4 \text{ ms}^{-1}$  corresponds with CO<sub>2</sub> mole fractions of about 403-404 ppm. Therefore the model is likely underestimating the in situ CO<sub>2</sub> observations at STD because of an overestimation of the surface wind speed.
- Further, we examine the relation of nighttime  $CO_2$  concentrations and 10 m wind direction at STD. Figure 9a shows that the dominant observed wind direction at night is east-southeast (ESE), followed by south (S). The ESE winds generally correspond



**Figure 8.** Scatter plot of nighttime wind speed at STD against hourly observed in situ CO<sub>2</sub>. Nighttime hours are defined as those between 14 UTC and 2 UTC.

with higher wind speeds (>  $2 \text{ ms}^{-1}$ ) and lower CO<sub>2</sub> concentrations (generally below 410 ppm), whereas observations with south winds generally coincide with very low wind speeds and CO<sub>2</sub> accumulation (Fig. 9a and Fig. 9b). The region in the east-southeast of STD is a rural area dominated by agricultural activities. Therefore these stronger ESE winds would generally bring lower-CO<sub>2</sub>-content air to STD. WRF-GHG, on the other hand, overestimates the wind speed and almost consistently

simulates ESE winds and lower  $CO_2$  concentrations (Fig. 9c and Fig. 9d). As such, the model-data mismatch is likely caused by a combination of both wind speed overestimation in WRF-GHG and discrepancies as to the wind direction, which are interrelated.

### 4.2.2 XCO<sub>2</sub>

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365 A higher correlation coefficient of 0.90 is found when comparing the daytime hourly-averaged TCCON XCO<sub>2</sub> data with the smoothed XCO<sub>2</sub> from WRF-GHG, see Table 3. Moreover, the RMSE along the time series is 0.75 ppm and there is a model underestimation of 0.37 ppm. The hourly relative model-data errors are below 0.5 % (Fig. 10). This is however slightly larger than the TCCON standard deviation of the hourly averages, which is around 0.1 %.

WRF-GHG provides a separation into different tracer contributions, which are shown in Fig. 11. Monthly averages of these

370 contributions for XCO<sub>2</sub> show a large (positive) biogenic enhancement in the months August to December. The biogenic tracer in WRF-GHG is driven by the online biogenic CO<sub>2</sub> fluxes calculated through the VPRM module as the sum of the gross ecosystem exchange (GEE) and respiration (Mahadevan et al., 2008). A positive biogenic tracer suggests that the respiration accumulates more CO<sub>2</sub> than the ecosystem can capture during the day by photosynthesis. Indeed, in the Southern Hemisphere, the dry season is generally from May until November leading to a decrease in GEE in some ecosystems (Quansah et al., 2015;

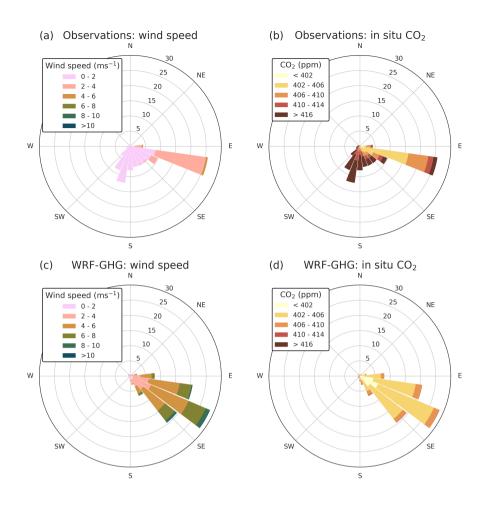


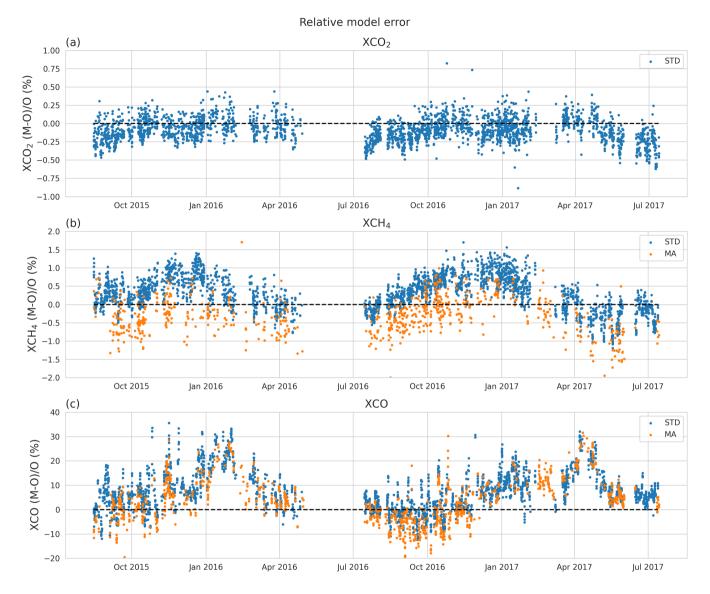
Figure 9. Wind rose of hourly nighttime data at STD, where the night is defined between 14 UTC and 2 UTC. (a) and (b) show the distribution of the observed wind speed per wind direction and near-surface  $CO_2$  concentration per wind direction, respectively. (c) and (d) show the same for WRF-GHG simulated data. The lengths of the bars show the frequency of occurrence in percentage.

Räsänen et al., 2017). Moreover, this carbon source was higher in 2016 because of a strong El Niño event leading to higher temperatures and less precipitation in the Tropics (Yue et al., 2017).
 The anthropogenic enhancement is relatively constant throughout the year. There is also a small biomass burning component

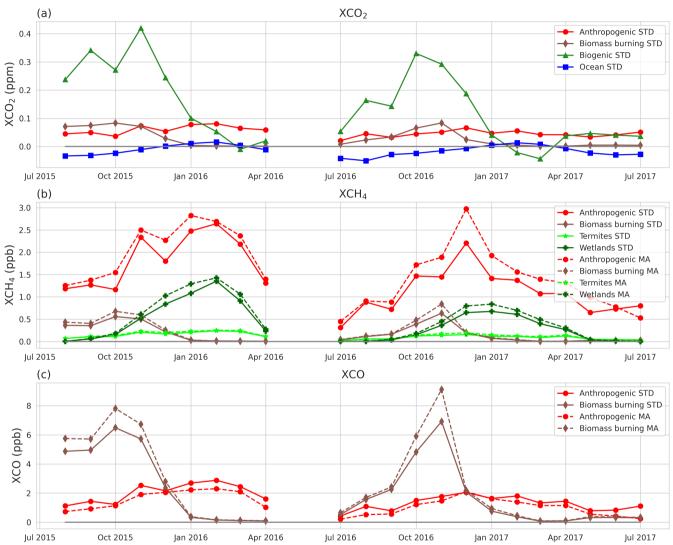
modeled in  $XCO_2$  in the months August to December, which corresponds to the biomass burning (BB) season. During these months, frequent fires occur in southern Africa and America. Duflot et al. (2010) showed that these polluted air masses can

380 be transported to Reunion Island and detected by FTIR observations such as XCO. Note that the  $XCO_2$  enhancements due to biomass burning coincides with the biogenic enhancements because, especially in the tropics, the occurrence and the duration of the BB season are linked to the dry season (Giglio et al., 2006).

A recent study using a COCCON spectrometer at Gobabeb in Namibia showed that the African biosphere can impact the



**Figure 10.** Relative percentage differences between hourly (smoothed) WRF-GHG and FTIR observations of XCO<sub>2</sub>, XCO and XCH<sub>4</sub>. The blue dots represent the data at STD (WRF - TCCON), while the orange data are from MA (WRF - NDACC).



Monthly mean tracer deviations from background

**Figure 11.** Monthly mean tracer contributions to the column-averaged mole fractions of (a)  $CO_2$ , (b) CO and (c)  $CH_4$ . The different colours represent different tracers: anthropogenic (red), biogenic (green), ocean (blue), wetlands (dark green), termites (light green) or biomass burning (brown). The solid lines are the mean monthly contributions at STD, while the dashed lines are for MA.

observed XCO<sub>2</sub> signal there due to medium and long-range transport (Frey et al., 2021). More specifically, they demonstrated

that the carbon sink of the African biosphere during Austral Summer can be observed in their XCO<sub>2</sub> measurements while it is not (or to a lesser extent) visible in the timeseries at Reunion. Backward trajectories for one specific day in February 2017 revealed that this contrast comes from the sampling of different air masses. This might suggest that the influence of fluxes from the African continent on the air above Reunion is seasonally dependent: a high impact during the dry season and a lower impact during wet season. Backward trajectory simulations over a longer time period performed by Zhou et al. (2018) might suggest this as well, however more research is needed to proof this statement.

As expected, the column observations of CO<sub>2</sub> are determined by different processes than the in situ CO<sub>2</sub> concentrations. Where the variation of XCO<sub>2</sub> is mainly driven by fluxes on the African continent, the surface CO<sub>2</sub> mole fractions are heavily influenced by local sources and PBL dynamics. This also agrees with the trajectory calculations by Zhou et al. (2018), showing that surface air mainly originates in the Indian Ocean while free tropospheric air is mainly coming from Africa and South America.

#### 4.2.3 Surface CH<sub>4</sub>

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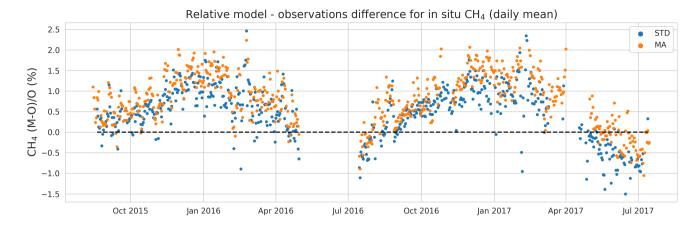
The model-data comparison for  $CH_4$  at STD (Table 3) shows only a weak correlation (0.35) between modeled  $CH_4$  and observations at STD. WRF-GHG shows an overestimation of about 9 ppb. Figure 6(b) and (d) show two apparent distributions in the scatter-plot. This is linked with the seasonal cycle of  $CH_4$ : observations show minimum values in December–February

- 400 and maximum values in August–September. The errors between the observations and WRF-GHG are not constant over time: in Fig. 12 a seasonal bias is found, with larger errors between December and February, which is Austral summer. This is a known weakness in the CAMS reanalysis, used as boundary information, as pointed out in the most recent validation report by Ramonet et al. (2020): "For all observations  $CH_4$  shows a seasonality in the relative difference between observations and CAMS simulations, which is increasing in the Southern Hemisphere after 2008. ... The seasonal dependence, which needs to
- 405 be investigated in more detail, may be related to the representation of OH in the model, or/and to errors in the seasonal cycle of surface emissions (mainly from agriculture and wetlands)." This demonstrates the importance of accurate lateral boundary conditions to simulate long-lived tracers with regional models such as WRF-GHG.

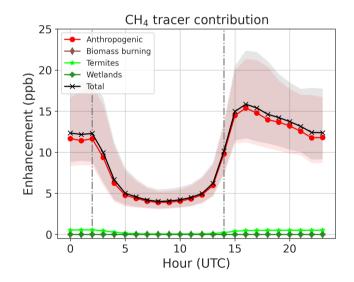
The diurnal cycle of the  $CH_4$  tracer contributions in Fig. 13 shows that the modeled  $CH_4$  consists almost entirely of the background signal and an anthropogenic enhancement, whereby both factors can add to the model-data mismatch.

410 The diurnal cycle is less pronounced in the observations (not shown). Nighttime values are on average only slightly larger than during the day, with a mean difference of only 3.09 ppb ( $\sigma^2$ =6.35), indicating that a nocturnal accumulation as identified for CO<sub>2</sub> in Saint-Denis is less evident for CH<sub>4</sub>. Moreover, the overestimation of the daily amplitude in WRF-GHG suggests an overestimation of the local anthropogenic CH<sub>4</sub> fluxes from EDGAR.

In contrast to  $CO_2$ , the  $CH_4$  mole fractions near the surface are less impacted by PBL dynamics and more by the background concentration.



**Figure 12.** Time series of daily mean relative percentage differences between model and in situ observations of  $CH_4$  at STD (orange) and MA (blue). The black dashed line indicates zero difference.



**Figure 13.** Diurnal cycle of in situ  $CH_4$  tracer contributions of WRF-GHG at STD. The black crosses represent the median hourly enhancements above background for the sum of all tracers. The separate tracer contributions are given in red (anthropogenic), dark green (wetlands), light green (termites) or brown (biomass burning). The shaded areas cover the interquartile ranges. The gray dotted vertical lines at 2 h and 14 h UTC indicate the approximate times of sunrise and sunset.

#### 4.2.4 XCH<sub>4</sub>

Incorrect boundary values and hence background concentrations have an impact on all  $CH_4$  simulations at Reunion Island. Therefore the statistics for the column-averaged mole fraction of  $XCH_4$  at STD (part of TCCON) are worse than for  $CO_2$ . A weak correlation is found (0.31, Table 3) and the model overestimates TCCON  $XCH_4$  by 5.69 ppb, which is slightly less than the bias with respect to in situ data (Sec. 4.2.2)

420 the bias with respect to in situ data (Sec. 4.2.3).

Figure 10c shows the relative differences between WRF-GHG and observational data, which have the same seasonal pattern as for the in situ comparisons (Fig. 12) caused by the reported seasonal bias for  $CH_4$  in the CAMS reanalysis data. The relative differences are below 2 %, but due to this seasonality in the errors, very little correlation is found.

- Even though the model fails at reproducing the measured time series, it is still interesting to examine the different modelled 425 tracer contributions to  $XCH_4$  (Fig. 11b). The tracers contribute only a couple of ppb to the total signal, with the anthropogenic being dominant throughout the year. Further, small peaks in biomass burning enhancements are found during the BB season as for the other species. The biogenic tracers for  $CH_4$  in WRF-GHG are generated by emissions from termites and wetlands. The termite signal is however very small and thus not relevant for this region. The signal from wetlands is larger, especially in Austral summer. This roughly coincides with the rain season, causing greater wetland extent (Lunt et al., 2019).
- 430 In the same way as for  $CO_2$ , the surface  $CH_4$  mole fractions at STD are influenced by local sources at Reunion Island, while fluxes from Africa and Madagascar are detected in the column observations because of the different air masses they sample.

#### 4.2.5 XCO

At STD, CO is only available as a column-averaged mole fraction (part of TCCON). As seen in Table 3, a very high correlation (0.89) is found for the hourly averaged paired data. WRF-GHG slightly overestimates the observed XCO (MBE: 5.07 ppb).

435 Figure 10c shows that the relative error between WRF-GHG and the XCO observations from TCCON is often below 20 % but not constant: larger errors up to 30 % are found from January until May.

As for the other species, large contributions of biomass burning (BB) emissions are found in the months August to December (Fig. 11c). Duflot et al. (2010) showed that XCO values during the BB season can reach up to twice the CO background concentration from other months. The rather limited BB enhancement found in WRF-GHG suggests that a substantial amount

- 440 of the XCO increase in the BB season is already included in the background tracer. This suggests that fires outside of the large domain can also be detected at Reunion Island, such as those from South America, which would confirm Duflot et al. (2010). The anthropogenic contribution is more constant throughout the year and the dominant contribution outside BB season. However it remains rather small compared to background, which appears to be the main driver behind the simulated XCO values at Reunion Island.
- 445 The larger model overestimation in January 2016 and April 2017 are thus likely linked to the background tracer, which is based on the CAMS global reanalysis for reactive gases. The corresponding CAMS validation report (Errera et al., 2021) mentions no known biases, but shows similar relative errors in those months in the Southern Hemisphere and in particular at MA (visible in Figure S.6 on p10). Again this demonstrates the importance of accurate boundary conditions for simulating XCO in this

region, but additionally points to the large influence of remote regions (outside domain d01) on the observed XCO time series 450 at Reunion Island.

#### 4.3 GHG data at Maïdo

At MA, the surface mole fractions of all three gases ( $CO_2$ ,  $CH_4$  and CO) are measured together with the column-averaged mole fractions of  $CH_4$  and CO that are part of NDACC. The results of each species are given in the sections below. Again, full time series of the observed and modeled data can be found in appendix C and statistical metrics of the model-data comparison

455 in Table 3. Remark that all statistical analyses of in situ observations at MA are performed on the complete dataset. Studies at other high-altitude stations often filter only those measurements which are representative for the free troposphere (Sepúlveda et al., 2014). However, analyses comparing only day- or nighttime data at MA showed no significant differences in the results.

#### 4.3.1 Surface CO<sub>2</sub>

At MA, the in situ CO<sub>2</sub> observations by the PICARRO instrument are well reproduced by WRF-GHG, resulting in a correlation coefficient of 0.75 and a very small mean bias error of -0.15 ppm (see Table 3). As at STD, the diurnal CO<sub>2</sub> cycle at MA shows a daytime minimum and a nighttime maximum (see Fig. 14a), however the amplitude is much smaller. This pattern is caught by WRF-GHG, although the amplitude is slightly underestimated. During the day, WRF-GHG shows a small overestimation of the CO<sub>2</sub> measurements of about 0.9 ppm, while at night a slight underestimation (about 0.4 ppm) is found. As seen on Fig. 14b, the modeled diurnal variation is almost entirely produced by the biogenic tracer, indicating that the biogenic flux calculated by

465 VPRM might be the reason for the model-observations discrepancies. The VPRM parameters used in the model are based on model tests in the Amazon region.

Foucart et al. (2018) showed that due to surface radiative cooling, the observatory is primarily situated in the free troposphere at night. The air at MA is then disconnected from local pollution sources and air from remote regions can be sampled. Indeed, no anthropogenic contribution is detected during the nighttime. However, the observed and simulated diurnal cycle of  $CO_2$ 

- 470 (Fig. 14a and b), shows that nighttime measurements at MA are still influenced by respiration of the local vegetation. The anthropogenic contribution at MA is very minor in WRF-GHG, which is expected because of the remote location of the Observatory. A very small enhancement is identified during the day. Since the local grid cell used for the model comparison does not include any anthropogenic flux, this enhancement is advected from elsewhere. It has been shown that orographic lifting can bring polluted air from coastal areas in the West towards MA during the day (Foucart et al., 2018; Duflot et al., 2019).
- 475 Despite that these westerly winds during the day were not reproduced by WRF-GHG (see Sect. 4.1.1), a daytime anthropogenic enhancement is found in the simulations. The model components representing biomass burning and ocean fluxes at MA are negligible.

So, according to WRF-GHG the main contribution (above the background) to the  $CO_2$  signal at MA is coming from the local vegetation and its photosynthesis and respiration, leading to a distinct diurnal cycle. The importance of the surrounding

480 biosphere for the surface observations at MA was also found by Verreyken et al. (2021) for volatile organic compounds. Even

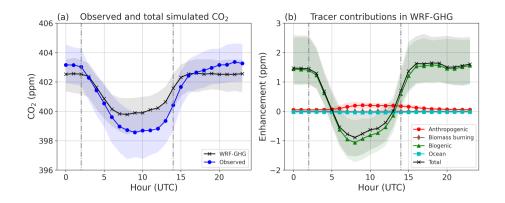


Figure 14. Same as Figure 7 but for Maïdo.

though the diurnal cycles of  $CO_2$  at STD and MA display similar patterns of minima during the day and maxima at night, they are caused by entirely different mechanisms.

#### 4.3.2 Surface CH<sub>4</sub>

Because of the importance of accurate background concentrations, the model performance at simulating in situ  $CH_4$  concentrations

- trations at MA is very similar compared to STD: the correlation is low (0.30) and the model overestimates the observations by circa 19 ppb. The modeled signal consists almost entirely out of the anthropogenic tracer (additional to the background signal), see Fig. 15a. Since the errors at MA follow the same pattern (Fig. 12) as at STD, and because inaccurate background information affects all CH<sub>4</sub> simulations, the seasonal bias in the CAMS reanalysis is also the cause for the weak model performance at MA. The errors at MA are larger than those at STD, likely due to the relatively low resolution of the EDGAR
  inventory (0.1 °) used for anthropogenic CH<sub>4</sub> emissions, leading to horizontal dilution: the concentration difference between
- high emission areas and their surroundings becomes smaller leading to an overestimation of the emissions in the low emission areas. In contrast with the model results, no diurnal cycle could be detected in the observations (not shown).

#### 4.3.3 XCH<sub>4</sub>

The model-data comparison for the NDACC data shows a very weak correlation (0.37) and a model underestimation of NDACC 495 XCH<sub>4</sub> (-5.65 ppb). This mean bias error has an opposite sign compared to TCCON CH<sub>4</sub>. Zhou et al. (2018) showed that NDACC XCH<sub>4</sub> is generally about 10 ppb lower than TCCON XCH<sub>4</sub> at Reunion Island due to their difference in vertical sensitivity. This pattern in the bias is the same as the one found by Ramonet et al. (2020) in comparisons of the CAMS reanalysis with NDACC and TCCON XCH<sub>4</sub>. Again, a seasonal pattern is found in the relative differences (Fig. 10b), caused by the reported bias for CH<sub>4</sub> in the CAMS reanalysis data.

500 The tracer contributions to the  $XCH_4$  signal at MA in WRF-GHG are very similar to those at STD, with seasonal enhancements from biomass burning and wetlands from Africa, alongside a more constant anthropogenic part (Fig. 11b). Remark that the

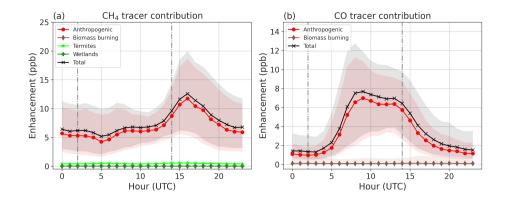


Figure 15. Diurnal cycle of tracer contributions for (a) CH<sub>4</sub> and (b) CO in situ surface concentration at MA.

contributions at MA seem to be slightly larger than those at STD. This is because the atmospheric column above the highaltitude station of MA is smaller than the one above STD and because the enhancements are transported from Africa and Madagascar by the westerlies higher up in the troposphere. The relative contributions averaged over the column are then higher at MA than at STD.

#### 4.3.4 Surface CO

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WRF-GHG captures the in situ surface CO time series at MA quite well: there is a high correlation of 0.83 (Table 3, Fig. 6e). The RMSE is about 11 ppb and there is a small model overestimation of 5.51 ppb. During the day, a small anthropogenic enhancement is found (see Fig. 15b), as for CO<sub>2</sub>. As already noticed, thermal contrasts make air masses from the coastal

510 areas arise during the day along the mountain slope before reaching the Maïdo Observatory. This air contains anthropogenic pollution. At the observatory itself no anthropogenic CO fluxes are implemented, so the model is representing this daytime advection to some extent.

Another contributor is the BB signal from August to December. The contribution is not very visible in the diurnal cycle due to its seasonal nature, but daily enhancements of up to 40 ppb are simulated by WRF-GHG. BB contributions from the African

515 continent and Madagascar are highest during the night, when MA is generally located in the free troposphere and transport from distant regions is detected (Baray et al., 2013).

### 4.3.5 XCO

A very high correlation (0.90) is found for the hourly averaged paired column data of NDACC and WRF-GHG (Table 3).

In general WRF-GHG slightly overestimates the observed XCO (MBE of 1.81 ppb). Remark that the errors are larger (the overestimation is larger) for the TCCON data compared to the NDACC data (also Fig. 10b). This is probably linked to biases between the TCCON and NDACC data sets. Zhou et al. (2019) showed that there is a bias of 2.5 % between TCCON XCO and NDACC XCO at Reunion Island, due to differences in the retrieval algorithm and data corrections.

As for  $XCH_4$ , the average monthly tracer contributions of XCO at MA are very similar as those at STD (Fig. 11b), with large contributions of BB emissions in the months August to December. Because of the unique location of MA, both the in situ observations and the column-averaged XCO are sensitive to these large seasonal events.

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#### 5 Model resolution

The above analysis was done using the WRF-GHG simulations from the innermost domain d03 (Fig. 2), which has a horizontal resolution of 2 by 2 km. As surface in situ observations are heavily influenced by local fluxes and dynamics, this high resolution is necessary to represent these measurements accurately, especially in regions with complex topography. As the ground-based

- 530 remote-sensing FTIR observation sample a much larger volume of air, a lower model resolution is likely sufficient to catch the fluxes and processes that influence them. Therefore, the model-data comparison for domains d01 and d02, with a horizontal resolution of 50 km and 10 km, respectively, is given in this section. Table 4 gives the statistical metrics for the FTIR observations at both sites, for all model domains.
- The results are very similar among the different model resolutions, indicating that even a horizontal resolution of 50 km (as in d01) could be sufficient to simulate the FTIR observations at Reunion Island. In addition to the larger sampling volume, this can be explained by the fact that the most important contributions to the column are coming from remote areas such as Africa and Madagascar, situated in d01 and d02. The added value of high-resolution transport in d03 is negligible for the FTIR observations.

	STD								МА						
	XCO <sub>2</sub> (ppm)			XCH <sub>4</sub> (ppb)		XCO (ppb)		XCH <sub>4</sub> (ppb)			XCO (ppb)				
	d01	d02	d03	d01	d02	d03	d01	d02	d03	d01	d02	d03	d01	d02	d03
RMSE	0.66	1.26	0.75	12.45	11.89	10.26	8.05	8.01	8.08	10.55	11.03	10.80	7.64	7.24	7.37
MBE	0.12	-0.24	-0.37	8.69	6.89	5.69	5.06	5.02	5.07	-5.22	-6.00	-5.65	1.92	1.46	1.81
CORR	0.90	0.75	0.90	0.27	0.34	0.31	0.88	0.89	0.89	0.36	0.37	0.37	0.89	0.89	0.89

**Table 4.** Overview of WRF-GHG performance of simulating hourly FTIR observations of GHG at Reunion Island, for all model domains.

 Comparison with the columns observations is done using the smoothed model profiles.

#### 6 Conclusions

We studied the variability of  $CO_2$ ,  $CH_4$  and CO surface and column observations at Reunion Island and evaluated possible factors influencing their observed mole fractions. This was achieved by comparing the available data sets with simulations of the WRF-GHG model over two periods between 2015 and 2017, totalling 20 months. The model performance was first evaluated for basic meteorological fields both near the surface as well as along atmospheric profiles. WRF-GHG shows good skill in reproducing these measurements, especially temperature. However, the local wind speed in Saint-Denis is overestimated

- 545 by almost 4 ms<sup>-1</sup> and also at Maïdo, there are some discrepancies in the wind speed and direction, which are likely linked to the complex topography and the model resolution of 2 km not being sufficient to represent very local dynamical processes. Nevertheless, the results enable us to answer the scientific questions posed in the introduction.
  - (1) To what extent are the observations influenced by local and nearby sources and sinks, or long-range transport of emitted gases?
- At both Saint-Denis and Maïdo, the in situ observations are heavily influenced by local and nearby sources and sinks, especially for CO<sub>2</sub>. However, the in situ observations at Maïdo can detect both signals from the coastal regions and from afar at night, when the Observatory is located in the boundary layer and the free troposphere, respectively. On the other hand, the column-averaged mole fractions describe different air masses than those near the surface and are not or only very slightly influenced by local activities. As shown by previous studies and confirmed here, these measurements at Reunion Island are influenced by processes from distant areas such as Africa and Madagascar. This is further evidenced by the fact that a model resolution of 50 km appears to be sufficient to simulate these observations.
- (2) What are the different tracer contributions to the observed concentrations, both at the surface and in the total column? The surface  $CO_2$  mole fractions in Saint-Denis follow a distinct diurnal cycle with values up to 450 ppm at night, driven by local anthropogenic emissions, planetary boundary layer dynamics and accumulation due to low wind speeds. 560 Additionally, the signal includes respiration from vegetation that is carried by eastern winds from more rural regions. At the Maïdo Observatory on the other hand, a similar diurnal cycle of CO<sub>2</sub> is found but with much smaller amplitude. There, the surface CO<sub>2</sub> mole fractions are essentially driven by the surrounding vegetation that take up CO<sub>2</sub> during the day and release  $CO_2$  during the night through respiration. The different model tracers of XCO<sub>2</sub> show contributions from fire emissions during the biomass burning season, but also positive biogenic enhancements associated with the dry 565 season. For CH<sub>4</sub>, tracer contributions reveal that the emission sources within the model domain have only a minimal effect on the overall signal. Besides the background, local anthropogenic fluxes are the major source influencing the in situ CH<sub>4</sub> observations at Reunion Island. However, the comparisons between the model fields and observations at Saint-Denis show that the anthropogenic emissions from EDGAR are likely largely overestimated; this is even more evidenced at Maïdo. Some (minor) impacts from Africa and Madagascar can be seen in the XCH<sub>4</sub> observations: fire plumes during 570 the biomass burning season and wetland emissions during the rainy season. For XCO, the importance of biomass burning plumes from Africa and elsewhere for the observed variability is confirmed. These plumes can also be detected by the in situ observations at Maïdo at night, while local anthropogenic signals are the main influence during the day.
  - (3) How accurate is WRF-GHG in simulating the different observation types of the three gases (CO<sub>2</sub>, CH<sub>4</sub> and CO) in the Southern Indian Ocean region, in particular at Saint-Denis and at the Maïdo Observatory? What are its strengths and weaknesses?

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In general, WRF-GHG shows great skill in simulating the different in situ surface and column observations of GHG. The simulations of  $XCO_2$  and XCO show a high correlation with the TCCON data with coefficients of 0.9 and 0.89, respectively. Similarly, a Pearson correlation coefficient of 0.9 and low errors are found between the model and NDACC

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	Evergreen forest	Deciduous forest	Mixed forest	Shrubs	Savanna	Crops	Grasses
$PAR_0$	993.9	324.0	206.0	303.0	6860.7	2329.0	15475.5
$\lambda$	0.1096	0.1729	0.2555	0.0874	0.0277	0.0417	0.0568
$\alpha$	0.2114	0.3258	0.3422	0.0239	-0.2535	-0.0814	-0.3122
β	1.8187	0.	0.	0.	7.1125	3.6716	7.3377

Table A1. VPRM parameters used within WRF-GHG

XCO time series. Further, WRF-GHG is able to adequately reproduce the in situ CO observations at Maïdo, and consequently to some extent the anabatic winds that are typical for the northwest part of the island, despite the differences in modeled and observed wind directions. The high model resolution of 2 km is needed to accurately represent local fluxes and small-scale processes that affect the in situ observations. However, because of the complex topography and the unique local wind patterns, an even higher resolution might be needed to simulate more precisely the observations at Maïdo. In addition, certain model flaws were discovered in this study. Due to an overestimation of local wind speeds in the capital, WRF-GHG underestimates the nocturnal  $CO_2$  buildup, leading to a correlation coefficient of only 0.62 between the model and surface  $CO_2$  measurements at Saint-Denis. Further, we found a small model underestimation of the amplitude of the diurnal cycle of surface  $CO_2$  at Maïdo which might indicate that the VPRM parameters could be improved for this region. Finally, WRF-GHG fails to reproduce accurately the different  $CH_4$  observations at Reunion Island due to a seasonal bias in the background arising from the CAMS reanalysis.

590 This study showed an application of the WRF-GHG model in a region of the globe where it had not yet been run before. It demonstrated that WRF-GHG had great skill in simulating the meteorological fields and different in situ surface and column observations of GHG. However results are highly dependent on accurate boundary conditions and the availability of high resolution emission inventories.

#### **Appendix A: VPRM parameters**

595 As mentioned in Sec. 3.1, this study uses the VPRM parameter set that was optimized by Botía et al. (2021) for the Amazon region in Brazil. Table A1 gives the exact values for every vegetation class.

#### Appendix B: Smoothing model data

A smoothing correction is applied when comparing the model data with the TCCON and NDACC data. Retrieved columnaveraged mole fractions are affected by the observing system characteristics and therefore Rodgers and Connor (2003) suggest

600 to take into account the a priori information and averaging kernels of the retrieval when calculating the Xgas of the model. The different steps undertaken to calculate this are explained hereafter for TCCON and NDACC separately.

Generally, the smoothed Xgas from WRF-GHG is calculated as:

$$X_{\text{gas},s} = \frac{TC_s^{\text{gas}}}{TC_{\text{air}}} = \frac{\sum_i PC_s^{\text{gas},i}}{TC_{\text{air}}} = \frac{\sum_i x_s^i PC_{\text{air}}^i}{TC_{\text{air}}},\tag{B1}$$

where  $PC_{air}^i$  is the partial column number density of dry air in layer *i* and  $x_s^i$  is the volume mixing ratio with respect to dry air 605 in layer *i* of the smoothed model profile. In the following, all parameters indicating a volume mixing ratio or column number density are also with respect to dry air, however for brevity it will not be specified any more.

#### **B1 TCCON**

Equation (B1) requires a smoothed vertical profile  $(x_s^i)$ . Since TCCON does not provide profile retrievals, we can not calculate this. Instead, we use the following smoothing equation for TCCON:

610 
$$X_{\text{gas},s} = \frac{TC_{\text{apriori}}^{\text{gas}}}{TC_{\text{TCCON}}} + \boldsymbol{a} \cdot \left(\frac{\boldsymbol{P}C_{\text{WRF,regrid}}^{\text{gas}}}{TC_{\text{WRF,regrid}}} - \frac{\boldsymbol{P}C_{\text{apriori}}^{\text{gas}}}{TC_{\text{TCCON}}}\right),\tag{B2}$$

where  $TC_{\text{TCCON}}$  is the total column (number density) of dry air from TCCON (for an atmospheric column up to 50 hPa). Similarly,  $TC_{\text{apriori}}^{\text{gas}}$  is the total column of the a priori mole fraction from TCCON calculated as the sum of the partial columns  $PC_{\text{apriori}}^{\text{gas},i}$  over those layers *i* that are below 50 hPa. Further, *a* is the vector with the column averaging kernels of TCCON. The regridded partial column profile of WRF-GHG ( $PC_{\text{WRF,regrid}}^{\text{gas}}$ ) and the total column of dry air from WRF ( $TC_{\text{WRF,regrid}}$ ) are

- 615 calculated in a few steps which are explained below. By including the total column of dry air from WRF in eq. (B2), we want to eliminate potential differences in air between TCCON and the model. As such, the priority is given to the volume mixing ratio profiles (instead of the calculation of dry air).
  - 1. Extend the WRF-GHG atmospheric profiles (gas mole fraction, pressure, temperature and water vapour) above the model limit (50 hPa) using information of the TCCON a priori profiles.
- 620 2. Calculate the dry-air partial column in layer *i* using the ideal gas law:

$$PC_{\rm air}^i = \frac{P^i}{RT^i} \frac{\tau^i}{1 + 1.6075 q^i},$$

with P atmospheric pressure, T air temperature, R ideal gas constant, q mass mixing ratio of water vapour and  $\tau$  layer thickness.

- 3. Calculate the gas number density partial columns as  $PC_{gas}^i = x_{gas}^i PC_{air}^i$ , with  $x_{gas}^i$  the gas mole fraction in layer *i*.
- 4. Regrid these partial column profiles to the full TCCON grid using a transformation matrix D as in Langerock et al. (2015):

 $PC_{\text{WRF,regrid}}^{\text{gas}} = \mathbf{D} \cdot PC_{\text{gas}} \text{ and } PC_{\text{WRF,regrid}} = \mathbf{D} \cdot PC_{\text{air}}$ 

5. Finally,  $TC_{\text{WRF,regrid}} = \sum_{i} PC_{\text{WRF,regrid}}^{i}$  where the sum is taken over all layers *i* below 50 hPa.

#### **B2 NDACC**

630 The smoothed Xgas from WRF-GHG at MA is calculated slightly different than at STD, as for NDACC volume mixing ratio profiles are provided. The smoothing equation can be written as:

$$\boldsymbol{x}_{\text{gas},\boldsymbol{s}} = \boldsymbol{x}_{\text{apriori}}^{\text{gas}} + \mathbf{A} \cdot \left( \boldsymbol{x}_{\text{WRF, regrid}}^{\text{gas}} - \boldsymbol{x}_{\text{apriori}}^{\text{gas}} \right), \tag{B3}$$

where  $x_{apriori}^{gas}$  is the volume mixing ratio (vmr) a priori profile from NDACC and A is the NDACC vmr averaging kernel matrix. Similar as for TCCON, a few steps need to be made to make the WRF-GHG data fit in eq. (B3). Steps 1-4 as described above should be followed, but using NDACC information instead of TCCON (a priori vmr profile, temperature, and water vapour

635

profiles, vertical grid). Then the regridded vmr profile from WRF-GHG is calculated as  $x_{\text{WRF, regrid}}^{\text{gas}} = \frac{PC_{\text{WRF, regrid}}^{\text{gas}}}{PC_{\text{WRF, regrid}}}$ . Finally, the smoothed dry-air mole fraction at MA is given by:

$$X_{\text{gas},s} = \frac{\sum_{i} x_{\text{gas},s}^{i} P C_{\text{WRF, regrid}}^{i}}{\sum_{i} P C_{\text{WRF, regrid}}^{i}},$$

where the sum is taken over all layers *i* below 50 hPa.

#### **Appendix C:** Time series 640

The full time series of both the observed and modeled concentrations at STD and MA are given in the figures hereafter. Figure C1 and C2 show the time series of the in situ data at STD and MA, respectively. Similarly, Fig. C3 and C4 show the comparison of the FTIR data at STD (TCCON) and MA (NDACC).

Author contributions. SC set up the model simulations, made the analysis and wrote the manuscript. JB provided the anthropogenic CO fluxes at Reunion Island and its description. VD provided the meteorological radiosonde profiles. BL provided expertise on the FTIR data 645 and model comparison techniques. DF assisted in setting up the model and together with JFM helped with correctly interpreting the results. JMM, CH, NK, MR and ML provided the FTIR and in situ measurements on Reunion Island. EM and MDM provided general guidance and support during the analysis and revised and edited the manuscript. All authors reviewed and commented on the manuscript.

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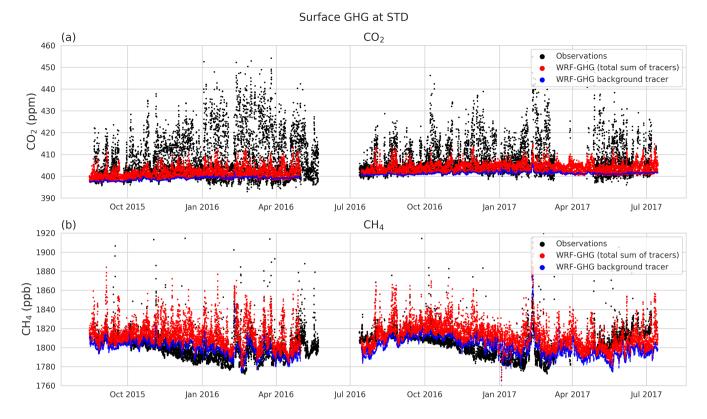


Figure C1. Time series of all observed (black) and modeled (red) in situ concentrations at STD of (a)  $CO_2$  and (b)  $CH_4$ . The blue dots represent the modeled background tracer.

input for the WRF-GHG simulations, were downloaded from the Copernicus Climate Change Service (C3S) Climate Data Store. The results
 contain modified Copernicus Atmosphere Monitoring Service information 2020-2021. Neither the European Commission nor ECMWF is responsible for any use that may be made of the Copernicus information or data it contains.

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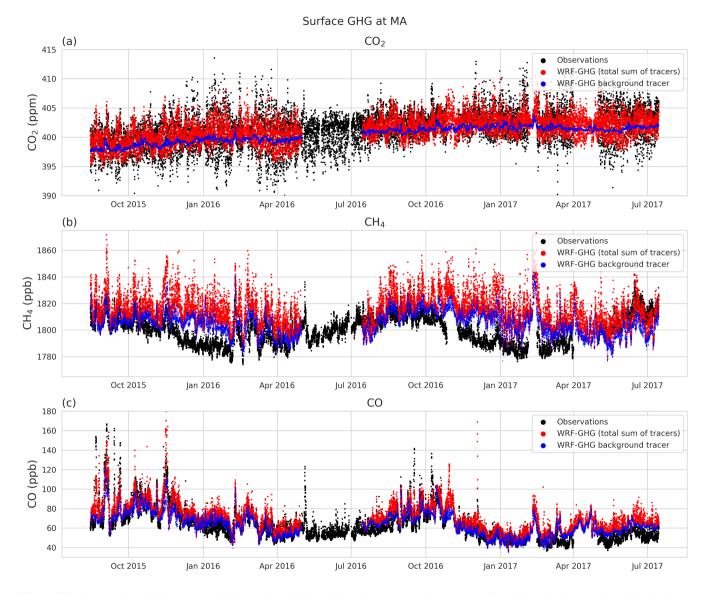


Figure C2. Time series of all observed (black) and modeled (red) in situ concentrations at MA of (a)  $CO_2$ , (b)  $CH_4$  and (c) CO. The blue dots represent the modeled background tracer.

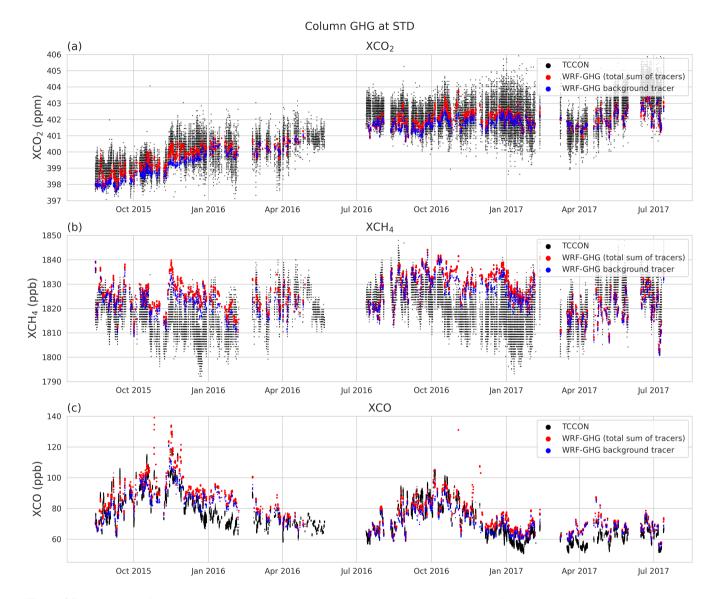


Figure C3. Time series of all observed (black) and modeled (red) column concentrations at STD of (a)  $CO_2$ , (b)  $CH_4$  and (c) CO. The blue dots represent the modeled background tracer. The modeled data is hourly and smoothed. The observed data is scaled to the atmospheric column until 50 hPa and all available measurements are shown.

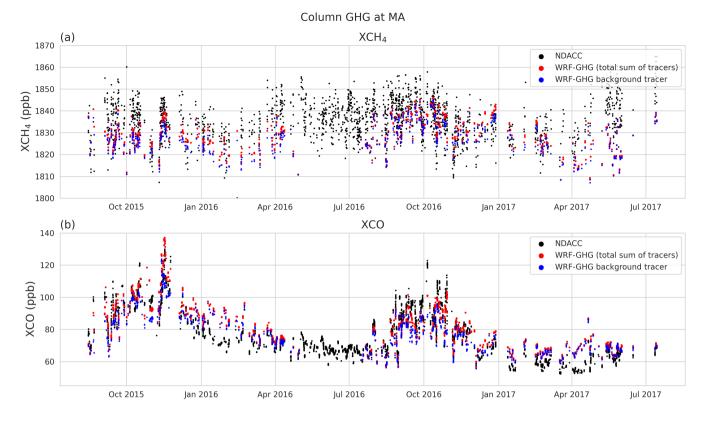


Figure C4. Time series of all observed (black) and modeled (red) column concentrations at MA of (a)  $CH_4$  and (b) CO. The blue dots represent the modeled background tracer. The modeled data is hourly and smoothed. The observed data is scaled to atmospheric column until 50 hPa and all available measurements are shown.

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