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4 **Measurement Report: Interpretation of Wide Range**
5 **Particulate Matter Size Distributions in Delhi**
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40 **ABSTRACT**

41 Delhi is one of the world's most polluted cities, with very high concentrations of airborne
42 particulate matter. However, little is known on the factors controlling the characteristics of wide
43 range particle number size distributions. Here, new measurements are reported from three field
44 campaigns conducted in winter, pre-monsoon and post-monsoon seasons on the Indian Institute of
45 Technology campus in the south of the city. Particle number size distributions were measured
46 simultaneously using a Scanning Mobility Particle Sizer and a Grimm optical particle monitor,
47 covering 15 nm to >10 μm diameter. The merged, wide-range size distributions were categorised
48 into five size ranges: nucleation (15-20 nm), Aitken (20-100 nm), accumulation (100 nm-1 μm),
49 large fine (1-2.5 μm) and coarse (2.5-10 μm) particles. The ultrafine fraction (15-100 nm) accounts
50 for about 52 % of all particles by number (PN₁₀-total particle number from 15 nm to 10 μm), but
51 just 1 % by PM₁₀ volume (PV₁₀- total particle volume from 15 nm to 10 μm). The measured size
52 distributions are markedly coarser than most from other parts of the world, but are consistent with
53 earlier cascade impactor data from Delhi. Our results suggest substantial aerosol processing by
54 coagulation, condensation and water uptake in the heavily polluted atmosphere, which takes place
55 mostly at nighttime and in the morning hours. Total number concentrations are highest in winter,
56 but the mode of the distribution is largest in the post-monsoon (autumn) season. The accumulation
57 mode particles dominate the particle volume in autumn and winter, while the coarse mode
58 dominates in summer. Polar plots show a huge variation between both size fractions in the same
59 season and between seasons for the same size fraction. The diurnal pattern of particle numbers is
60 strongly reflective of a road traffic influence upon concentrations, especially in autumn and winter,
61 although other sources such as cooking and domestic heating may influence the evening peak.
62 There is a clear influence of diesel traffic at nighttime when it is permitted to enter the city, and also
63 indications in the size distribution data of a mode <15 nm, probably attributable to CNG/LPG
64 vehicles. New particle formation appears to be infrequent, and in this dataset is limited to one day

65 in the summer campaign. Our results reveal that the very high emissions of airborne particles in
66 Delhi, particularly from traffic, determine the variation of particle number size distributions.
67

68 1. INTRODUCTION

69 Air pollution in Delhi has been studied for many years, and the authorities have implemented several
70 interventions designed to limit the concentrations. The sulphur content of diesel and petrol fuels was
71 reduced to 50 ppm during 1996-2010, more than 1300 industries were shut down due to hazardous
72 emissions, commercial vehicles older than 15 years were gradually taken out of the traffic fleet, and
73 public transport vehicles and auto-rickshaws were converted to compressed natural gas (CNG) fuel
74 (Narain and Krupnick, 2007). An odd–even vehicle number plate restriction has been applied during
75 working days (Chowdhury et al., 2017). Although these measures have reduced gaseous pollutants
76 (SO₂ and CO) and primary particulate matter, in recent years, several studies have reported that the
77 PM_{2.5} concentrations have been constant or slowly increasing in India, especially in the winter and
78 autumn seasons (Babu et al., 2013; Balakrishnan et al., 2019; Dandona et al. 2017, Kumar et al.,
79 2017), except in 2020. In 2020, the PM_{2.5} level decreased by approximately 40 %, due to Covid-19
80 measures (Rodríguez-Urrego and Rodríguez-Urrego 2020; Mahato et al., 2020). Although the overall
81 emission sources in India are dominated by traffic, industry, construction, and local biomass burning,
82 haze pollution events in Delhi are frequently related to the large-scale open burning of post-harvest
83 crop residues/wood during the crop burning season in nearby rural regions (Cusworth et al. 2018;
84 Bikkina et al. 2019; Kanawade et al., 2020). Furthermore, the sources of particles are mostly local
85 (Hama et al., 2020), meteorological factors play an important role in influencing concentrations of
86 air pollution (Tiwari et al., 2014; Yadav et al., 2016; Guo et al., 2017; Dumka et al. 2019; Kumar et
87 al. 2020).

88

89 Annual average PM_{2.5} levels range between 81 and 190 µg/m³ in Delhi and are clearly higher than
90 the WHO guideline value (5 µg/m³) and Indian national limit value (40 µg/m³) (Hama et al., 2020).
91 To the best of our knowledge, most studies in India have focussed on the source apportionment from
92 chemical profiles of particles (Pant and Harrison, 2012; Jain et al. 2020; Bhandari et al., 2020; Rai
93 et al., 2020). Mostly they have reported that biomass burning contributes greatly to PM_{2.5} mass while

94 traffic contributes heavily to PM₁₀ mass in Delhi. Residential energy use contributes 50 % of the
95 PM_{2.5} mass concentration and the construction sectors are also considered an important source of
96 particle mass (Guttikunda et al., 2014; Butt et al., 2016; Conibear et al., 2018). Furthermore, it is
97 particularly important to understand the absolute contribution and sources of different sizes of
98 particles within PM_{2.5}. A recently published paper by Das et al. (2021) highlighted that <250 nm
99 particles contribute a significant proportion of the total PM_{2.5} mass and are a potentially important
100 link with human health.

101

102 The Particle Number Size Distribution (PNSD) can provide air pollution source apportionment with
103 high time resolution compared to use of chemical species, and influences the aerosol transport and
104 transformation profiles in the urban atmosphere and toxicological effects on humans (Wu and Boor,
105 2021). Many PNSD studies have been conducted in urban, traffic and background sites over the past
106 decades and three review studies have been published (Vu et al., 2015; Azimi et al., 2014; Wu and
107 Boor, 2021). There are some studies evaluating the number or mass particle size distribution (PSD)
108 in Delhi (Mönkkönen et al., 2005; Chelani et al., 2010; Gupta et al., 2011; Pant et al., 2016; Gani et
109 al., 2020). Harrison (2020) compared PNSDs from Delhi, Beijing and London and reported that the
110 particles from Delhi are far greater in number with a much larger modal diameter, close to 100 nm.
111 In a recent paper, Gani et al. (2020) has investigated the PNSD up to 0.5 µm sizes from 2017 to 2018
112 and reported that rapid coagulation is an important process in Delhi.

113

114 The wide range PNSD is important to describe all sources of inhalable particles (<10 µm). It is not
115 easy to separately identify particles arising from resuspension, sea salt and construction, or from brake
116 wear and combustion or vehicle exhaust, using only the <0.5 µm particle size range. Harrison et al.
117 2011 reported that using wide range particle sizes in source apportionment was extremely successful
118 in identifying the separate contributions of on-road emission including brake wear and resuspension.
119 Although there are a few studies of wide range particle characterization in Beijing (Jing et al., 2014)

120 and source apportionment in Venice, Italy (Masiol et al., 2016), there has been no previous wide
121 range PNSD study in Delhi. In this study, we aimed to interpret particulate matter size distributions
122 over a wide range (15 nm to 10 μm) in the winter, post-monsoon and pre-monsoon seasons in Delhi.
123 Future studies will look at two-step receptor modelling of wide range particulate matter size
124 distributions and chemical composition in Delhi.

125

126 **2. METHODS**

127 **2.1 Study Area**

128 The measurements were part of the NERC/MoES Air Pollution and Human Health in an Indian mega-
129 city (APHH-Delhi, www.urbanair-india.org) study, a joint UK-India project addressing air pollution
130 in Delhi. The sampling location was ~15 m above ground level on the 5th floor of the Civil
131 Engineering Department at the Indian Institute of Technology Delhi (IIT Delhi) campus, located in
132 New Delhi, representative of an urban background environment (28.545 N, 77.193 E) (Figure S1).
133 The measurement station is at a 120 m distance from a major arterial road. As part of APHH-Delhi,
134 there were three field campaigns: (i) Jan-Feb 2018 (winter), (ii) May-June 2018 (summer; pre-
135 monsoon) and (iii) Oct-Nov 2018 (autumn; post-monsoon). In all field campaigns, a suite of gas and
136 particulate phase instrumentation was deployed within a temperature controlled laboratory.

137

138 These sampling periods were representative of conditions for PM and gases during these seasons in
139 Delhi. We found the average $\text{PM}_{2.5}$ concentration to be approximately 180 $\mu\text{g}/\text{m}^3$, 220 $\mu\text{g}/\text{m}^3$ and 120
140 $\mu\text{g}/\text{m}^3$ for winter, autumn (excluding Diwali) and summer, respectively measured by a TEOM-FDMS
141 (TEOM-Filter Dynamic Measurement System). Hama et al. (2020) studied the long term (from 2014
142 to 2017) trends of air pollution in Delhi at 6 stations (residential, commercial, and industrial sites)
143 and reported that the mean $\text{PM}_{2.5}$ concentrations ranged between 147 – 248 $\mu\text{g}/\text{m}^3$, 147 – 248 $\mu\text{g}/\text{m}^3$
144 and 76 – 135 $\mu\text{g}/\text{m}^3$ for winter, autumn and summer, respectively, and a good correlation between

145 sites within Delhi. This gives reassurance that the PM_{2.5} concentrations measured at our site are within
146 the typical range of those observed in Delhi.

147

148 **2.2 Measurements**

149 Aerosol particle sizes in the atmosphere span a very wide range from a few nanometers at the lower
150 end to some tens of micrometers at the upper end. Because of this very wide range of sizes, particle
151 properties vary considerably across the size spectrum with the behaviour of the smaller particles being
152 determined by their high mobility and hence diffusivity, whilst at the coarse end of the size
153 distribution inertial properties are especially important. Due to this divergence in behaviour, no
154 instrument is capable of measurement of the whole range of particle sizes.

155 To measure the particle size range used in this study, two particle instruments were used to collect
156 number size distributions (NSD). For the range 15-640nm, a TSI- Scanning Mobility Particle Sizer
157 (SMPS) 3936 was used, consisting of a TSI 3080 Electrostatic Classifier, TSI 3081 DMA and TSI
158 3775 CPC. To extend this range into the coarse mode a Portable Laser Aerosol Spectrometer and
159 Dust Monitor (GRIMM 1.108) were used alongside the SMPS.

160

161 **2.3 Merging Process**

162 Merging procedures have usually been reported for merging SMPS and APS (Aerosol Particle Sizer)
163 data, but here Grimm optical spectrometer (OP) data is merged with SMPS data. For a complete
164 particle size distribution, simultaneously collected, paired hourly averaged particle number size
165 distributions collected from the SMPS and Grimm were merged. The merging procedure is based on
166 the principle of converting the diameters of the Grimm-derived data to a diameter matching the
167 SMPS-derived data, in the region where the size distribution measurements overlap. The Grimm
168 measures the optical diameter d_b^t whereas the SMPS measures the mobility diameter d_a^t of the
169 particles. Comprehensive descriptions of the procedure and mathematics are given by DeCarlo et al.
170 (2004) and Schmid et al. (2007). The Grimm NSD are translated onto the extended electrical mobility

171 diameter axis of the SMPS using equation (R1) (Beddows et al. 2010; Liu et al., 2016; Ondracek et
172 al., 2009).

173

$$174 \quad d_b^t = \frac{d_a^t}{X} \sqrt{\frac{c(a_a^t)}{c(a_b^t)}} \quad (\text{R1})$$

175

176 The Cunningham slip correction factor is given by C and the unknown variables such as the shape
177 factor of the particles are accounted for by a free parameter X (given by equation R2) which is adjusted
178 until the tails of the SMPS and Grimm NSD overlap each other giving a continuous NSD across the
179 particle size bins measured by the two instruments.

180

$$181 \quad X = \sqrt{\frac{\rho_e^t}{\rho_o}} \quad (\text{R2})$$

182

183 The estimated transition-regime effective density ρ_e^t (normalised by the unit density, ρ_o) typically
184 ranges from 0.77 to 2.56 g/cm³ when aerodynamic diameter is used in merging. Detailed
185 information upon the effective particle density based on the geographical regions is seen in the Wu
186 and Boor (2021) study.

187

188 The merging algorithm (originally programmed in CRAN R) was implemented using Excel
189 spreadsheets and the solver tool minimised the separation between the tails of the overlapping SMPS
190 and Grimm. Due to the imperfect nature of the data, each of the merges was allocated a factor
191 indicating quality based on whether: (i) there is a successful fit; (ii) the scatter of the data across the
192 overlapping tails; (iii) the fraction of points on the tail falling onto the fitted curve; and (iv) how
193 smooth the overlap is (Table S1). The size bins overlap (300-700 nm) between Grimm and SMPS.
194 This process was repeated for the winter, summer and autumn data sets and any results failing the test

195 were either repeated or the data removed from the analysis. In all, only 8 samples from 1117 failed
196 to give an acceptable fit in the merge procedure.

197

198 **2.4 Data and Quality Management**

199 Data from SMPS and Grimm were measured with 1-min resolution and converted to hourly averages.

200 The seasons were categorized as winter, autumn and summer. The measurements were taken in

201 winter from 12 January 16:00 to 11 February 04:00, in autumn from 24 October 16:00 to 11

202 November 10:00, in summer 16 May 19:00 to 05 June 15:00 in 2018. There were 709, 403 and 477

203 total pairs (hours) in the data sets in winter, autumn and summer, respectively. But 172, 43 and 257

204 pairs in winter, autumn and summer, respectively were excluded because of the non-availability of

205 data at that time. Data coverage is 76 % for winter, 95 % for autumn and 46 % for summer. Figure

206 S2 in the Supplementary shows hourly mean values of total particle counts for three seasons. In order

207 to evaluate day and night time PNC (particle number concentration) differences, the day and night

208 were defined as 07:00-19:00 and 19:00 – 07:00, respectively. All times reported are local times

209 recorded in Indian Standard Time (IST; GMT+05:30).

210

211 R version 3.1.2 was used to analyse the data (R Core Team, 2015). Firstly, all data were checked for

212 clean-up of the robustness of the data sets, to detect anomalous records and take out the extreme

213 values. Data greater than the 99.5th percentile were deleted. Diwali time in 2018 (7th of November

214 2018 from 16:00 to 23:00) was taken out the data set in order to exclude its extreme effect on PNSD

215 values. Particle number concentrations during Diwali time are given in the Supplementary, Figure

216 S3. There were some single gaps in the data matrixes. These missing data were replaced by linearly

217 interpolated values from the nearest bins to those samples.

218

219 In the literature, PNCs measured below 1 μm are frequently split into three ranges: nucleation, Aitken

220 and accumulation (Gani et al., 2020). Nucleation size ranges have variously been described as below

221 30 nm (Masiol et al. 2016) or below 25 nm (Gani et al., 2020) or below 20 nm (Wu and Boor, 2021).
222 Some studies have evaluated wide range PNSDs split into 4 ranges (nucleation, Aitken, accumulation
223 and coarse) (Masiol et al. 2016; Harrison et al., 2011). In this study, the modes have been aggregated
224 into five size groups: nucleation (15-20 nm), Aitken (20 -100 nm), accumulation (100 nm – 1 µm),
225 large fine (1 µm – 2.5 µm) and coarse (2.5 µm – 10 µm) based on merged. Ultrafine particles (UFP)
226 are considered to be total PN counts of Nucleation and Aitken modes (<100 nm).

227

228 The particle mass was calculated for the SMPS+OP merged data, assuming a density of 1.6 g cm^{-3}
229 (Gani et al., 2020). Estimation of particle density as a function of size is extremely difficult, and there
230 are few data for particle density from Delhi. Since Gani et al (2020) used the density of PM at the
231 same location as in our study, we used the same density value to convert PN to PM mass. Figure S4
232 shows the comparison of $\text{PM}_{2.5}$ measured by SMPS+OP and TEOM-FDMS in Delhi for the three
233 seasons. Figure S5 shows the comparison of $\text{PM}_{2.5}$ with relative humidity measured by SMPS+OP
234 and TEOM in Delhi for the three seasons. A good correlation of the estimated particle mass with
235 independent measurements with a co-located TEOM-FDMS was observed, except in summer.

236

237 The cumulative frequency of observations as a function of particle size was calculated for each hour
238 of the day. Standard central measures from the cumulative frequency plots were represented by the
239 geometric mean diameter (GMD) for each size distribution. They were used to examine particle
240 growth processes. Firstly, the growth of GMD was estimated visually from the diurnal GMD data
241 plot (Fig. 8). The minimum growth time used for estimation of the growth rate (GR) was selected as
242 three hours, and if the growth lasted for long enough, the GR was estimated. The observed growth of
243 the GMD of the particle was quantified by fitting the GMD of particles during the growth process
244 event over a period of time 't' (equation R3). Detailed information on the method can be found in
245 Sarangi et al. (2015; 2018).

246

247 Growth rates (nm/hour)-GR =dGMD/dt (R3)

248

249 3. RESULTS

250 3.1 Particle Number and Size

251 Table S2 gives the descriptive statistics of particle number counts ($\#/cm^3$) calculated using every 1-
252 hour measurements for the nucleation, Aitken, accumulation, large fine and coarse modes between
253 15 nm and 10 μm in all seasons. Time series of total particle number counts are presented in Figure
254 S2. The average total PN levels were $36.73 \times 10^3 cm^{-3}$ in winter, $29.35 \times 10^3 cm^{-3}$ in autumn and
255 $18.91 \times 10^3 cm^{-3}$ in summer. Generally, the wintertime PN levels were higher than the other seasons.
256 The wintertime PN levels of nucleation, Aitken and accumulation modes were ~ 1.5 , 1.8 and 2.2 times
257 higher than in summer, respectively. Similar ratios were obtained by Guttikunda and Gurjar (2012)
258 in Delhi for particulate matter concentrations. This is attributed to the unfavorable dispersion
259 conditions, including low wind speed and low mixing height during the winter season. The autumn
260 PN levels of nucleation, Aitken and accumulation modes were ~ 1.5 , 1.3 and 1.9 times higher than in
261 summer, respectively. The wintertime and autumn average PN levels are similar except for the Aitken
262 mode for which winter is 1.4 times higher than in autumn. However, for the large fine and coarse
263 modes the PN level was not markedly different between winter, autumn, and summer. Gani et al.
264 (2020) reported that the average PN levels were $52.50 \times 10^3 cm^{-3}$ in winter, $43.40 \times 10^3 cm^{-3}$ in
265 summer, and $38.00 \times 10^3 cm^{-3}$ in autumn in Delhi measured in 2017. The differences in the magnitude
266 of number counts between the two studies are potentially explained by the difference in the sampling
267 period and changes in emissions.

268

269 Figure 1 shows a comparison of average particle number and volume and the contribution to total
270 PN. The average PV (particle volume) levels indicate that PV of the Aitken mode is highest in winter,
271 while the accumulation mode is highest in autumn and the coarse mode is highest in summer. The
272 contribution of UFP to numbers is highest in summer (57 %) but their contribution to volume is the
273 lowest in autumn and summer (<1 %). The contribution to both number and volume of the

274 accumulation mode is highest in autumn with 51 % and 75 %, respectively. UFP contributions to total
275 PV are below 1 % in Delhi. Furthermore, it can be seen clearly that the coarse fraction of particles
276 dominates in summer, while the accumulation mode dominates in autumn and winter.

277

278 Wu and Boor (2021) analysed the PNSD observations made between 1998 and 2017 in 114 cities in
279 43 countries around the globe. They reported that there are significant variations in the magnitude of
280 urban aerosol PNSD among different geographical regions. The main finding of their study is that the
281 PNSD in Europe, North America, Australia, and New Zealand are dominated by nucleation- and
282 Aitken-mode particles while in Central, South, Southeast and East Asia they are dominated by the
283 substantial contribution from the accumulation mode, which is consistent with our finding. Pant et
284 al. (2016) report mass size distributions for particulate matter sampled by cascade impactor in Delhi
285 in winter. The dominant modes appear at around 3-4 μm and 0.6 μm , with a lesser peak at 0.2 μm
286 aerodynamic diameter. These are respectively in the coarse (former mode) and accumulation (latter
287 two modes) ranges as classified in the current study. The largest component of mass was in the
288 accumulation mode, and the distribution fits well with the pattern of data seen in Figure 1. Major
289 components of the coarse fraction were Al, Si, Ca and Fe (Pant et al., 2016), suggestive of soil and
290 street dust as major contributors. The elements most notably in the accumulation fraction were Cu,
291 Zn, Pb and Sb, indicative of non-exhaust traffic emissions and metallurgical sources, and S, which
292 showed a major peak due to sulphate, peaking at 0.9 μm (Pant et al., 2016).

293

294

295 **3.2 Diurnal Change**

296 Figure 2 shows the diurnal variation of particle number concentrations and of $\text{PM}_{2.5}$, BC, NO and
297 NO_2 for each season (excluding the day of Diwali), and the normalized time variations of all particle
298 fractions are given in Figure S6. Figure S7 represents the diurnal variation of meteorological
299 parameters. In general, there are large differences of PN levels between cold seasons (winter and
300 autumn) and warm season (summer) for nucleation size particles. Coarse mode particle numbers in

301 the summer are higher than in winter and autumn, except in the evening time. For autumn and winter,
302 particle counts are similar from 7 am to 7 pm (daytime). However, from 7 pm to 7 am particle counts
303 in winter are higher than in autumn. The lowest levels for all modes were present during the afternoon
304 in all seasons (2-4 pm), followed by highest levels during the night in winter (after 8 pm). The winter
305 and autumn diurnal profiles had two peaks for below 1 μm particle size in the morning and evening
306 corresponding to the traffic rush hours. But in the summer the same peaks for nucleation, Aitken and
307 accumulation modes are seen although of smaller magnitude, and one hour earlier comparing to the
308 winter and autumn. Pant et al. (2016) reported the diurnal variation of traffic at one of the major
309 arterial roads in Delhi and Dhyani et al. (2019) reported on traffic-related emission. Figure S9 shows
310 the diurnal variation in traffic at a major road intersection in Delhi. Cars, two/three wheelers, bus and
311 LCV (light commercial vehicle) fleet numbers increase in the morning, persist throughout daytime
312 and start to decrease at 22:00. Due to the prohibition of access for heavy-duty diesel vehicles to central
313 Delhi from 6:00 am until 11:00 pm in the night, during the daytime including the traffic rush hours
314 the HCV (heavy commercial vehicles) number is at its lowest level (Figure S9 or Dhyani et al. 2019).
315 While road traffic clearly influences the diurnal pattern in PN, other sources including cooking and
316 domestic combustion are likely to contribute. Small midday PN peaks were observed during the
317 summer in the nucleation, Aitken and coarse modes. Another study conducted in Delhi reported the
318 same midday peaks in the warm season and the highest levels in the cold season (Gani et al. 2020),
319 which may be related to bus and LCV emissions at midday.

320

321 Figure 3 shows the differences in diurnal variations of total PN levels between the weekday and
322 weekend. These are based upon a small dataset, and hence the rather small differences within a season
323 may not be meaningful. In winter the PN levels on Saturday and Sunday are higher than on the
324 weekdays during the night (from 8 pm to 10 am the next day). However, after the morning rush hour
325 peaks, during the daytime the PN levels are the same for all days. The diurnal variation of PN in the
326 autumn shows no significant differences among the days with the same main peaks in the morning

327 for all days, although highest on Saturday. There is a flattened peak (from 8 to 10 am) in the morning
328 rush hour for the weekday while there are pointed peaks at approximately 9 am on Saturday and
329 Sunday in winter and autumn. Measurements made during the summer period are very limited. Due
330 to there being only 4 full days and 9 half days of measurements, it is very hard to draw any
331 conclusions. Even so, there are indications of a weekday traffic effect upon the PN levels in summer.
332 There is only one day of measurements on a Sunday (3rd June 2018) and it shows the midday peaks.
333 Overall, despite seasonal differences, there appears to be a strong influence of light duty road vehicles
334 upon the diurnal profiles, reflecting traffic volumes, with an impact of heavy duty vehicles upon
335 nighttime concentrations of all particle fractions.

336

337 NPF events present variable seasonality for different areas, though in most cases they appear to be
338 more frequent during spring or summer (Salvador et al., 2021). Gani et al. 2020 studied long term
339 PNSD in Delhi and have stated that they did not see any NPF during the winter or autumn seasons in
340 Delhi. In this study, the identification of NPF events was conducted manually using the criteria set
341 by Dal Maso et al. (2005) and used by Bousiotis et al (2019; 2021). The data were analysed visually
342 on a day-to-day basis: each 24-hour period, from midnight to midnight. According to these criteria, a
343 NPF event is considered when: a distinctly new mode of particles appears in the nucleation mode size
344 range, prevails for some hours, and shows signs of growth. These are the initial criteria used in
345 identifying the events. Following that, as the dataset starts from a rather large size (15 nm), to be
346 more confident about the events and not to confuse them with pollution events, high time resolution
347 data for NO_x as well as the fluctuations of the condensation sink were also used to identify pollution
348 events affecting particle concentrations which were not considered. Hence, while we checked the
349 particle size distributions for the NPF events, we also looked at the levels of pollutants to ensure that
350 what was attributed to a NPF event was not particles from pollution / direct emissions. By considering
351 the pollution levels and condensation sink we can reduce the possibility of including particle
352 formation events that are not associated with secondary formation. After analysing all data,

353 measurements from only one day during the measurement campaign were compliant with the criteria
354 set as a class Ia NPF event. Figure 4 presents the contour plots of average diurnal variation for all
355 seasons and for the NPF event on 3rd June. NPF may be suppressed due to very high pre-existing
356 aerosol concentrations (Kanawade et al., 2020; Gani et al. 2020) during severe air pollution episodes
357 in Delhi. This suppression effect has also been observed in European cities (Bousiotis et al., 2019;
358 2021).

359

360 A new study by Sebastian et al. (2021) analysed PNSD and the frequency of NPF at six different
361 locations in India. The Delhi observation site is in an urban area and located at CSIR-National
362 Physical Laboratory (NPL), approximately 8 km from the IIT location described as urban background
363 in our study. They found that the NPF frequently occurs in the spring season, but is least common in
364 autumn and winter due to air pollution episodes suppressing the NPF. They also stated that the highest
365 concentration and frequency of occurrence of NPF events in was Delhi as compared to other sites. As
366 in other studies (such as Bousiotis et al., 2021), this study also emphasized that the increased
367 concentrations of precursor gases are important for the occurrence of NPF in urban areas.

368

369 **3.3 Day and Night Time Differences in PN and PV**

370 Table S3 presents the summary statistics of the particle number and mass levels derived from merged
371 particle number data and BC, NO_x and PM_{2.5} at night and day for each season, excluding Diwali.
372 Figure 5 shows the particle number comparison of all modes at night and day seasonally. In both
373 night and day, the nucleation counts are approximately the same in autumn and summer (N/D=1.1
374 and 1.0), and a little higher at night in winter (N/D=1.3). But in the night, Aitken and accumulation
375 counts are higher than in the day by factors of 1.4 and 1.5 times in summer, 1.2 and 1.5 times in
376 autumn, respectively and approximately 2 in winter. While the coarse mode PN counts are
377 approximately the same for all seasons and day / night, the large fine PN level in the nighttime are
378 significantly higher (1.7) than in the daytime in summer. It seems that in the nighttime high PM
379 concentrations are due to the increasing Aitken and accumulation modes occurring from coagulation

380 of nucleation mode particles, condensation of low volatility species or hygroscopic growth. In
381 addition, biomass burning and older diesel vehicles can contribute significantly to particles in these
382 fractions (Kumar et al., 2013; Chen et al., 2017; Gani et al., 2020). Meteorological factors can also
383 profoundly affect the PN levels in daytime and nighttime. The differences of wind speed between day
384 and night in summer are lower than in winter and autumn (Figure S7). Higher wind speed, and lower
385 humidity, may favour the resuspension of coarse dust as a dominant mechanism in the summer.
386 Seasonal changes in mixing depths are surprisingly small (Figure S7) and hence unlikely to have a
387 major influence. However, the major increases and decreases in the diurnal plots of pollutants (Figure
388 2) are consistent with the diurnal plots of MLH (Figure S7). Autumn and winter also have longer
389 periods with low mixing heights, also seen in Figure S7. TPN showed a negative exponential
390 dependence upon MLH, which became more scattered at lower values of MLH, probably reflecting
391 the larger relative errors in MLH estimates at smaller values.

392

393 Figure S8 represents the relation between the ventilation coefficients ($VC = MLH \times \text{wind speed}$) and
394 TPN as a function of hour and month. Gani et al. (2019) reported that the VC is 4–6 times smaller for
395 the wintertime compared to the summer in Delhi. In this study, the VC is 1.8 and 1.6 times higher in
396 summer (mean 2732 m²/s) compared to the winter (mean 1491 m²/s) and autumn (mean 1702 m²/s),
397 respectively. The daytime hourly TPN levels are lower as related to the higher VC and the lower VC
398 in colder months gives higher TPN. Although there are not enough daily data (especially for summer)
399 to give more detail, we can see the same trend as comparing to the weekly data from Gani et al (2019)
400 study.

401

402 Overall, for the daytime for all seasons, hourly averaged UFP concentrations are usually less than
403 the nighttime, however the UFP contribution to the PN₁ (55 % in day, 50 % in night for winter; 52
404 % in day, 45 % in night for autumn; 58 % in day, 56 % in night for summer) and PN₁₀ (38 % in day,
405 38 % in night for winter; 40 % in day, 33 % in night for autumn; 36 % in day, 33 % in night for

406 summer) are mostly slightly higher in the daytime. Similarly, Gani et al. (2020) have reported the
407 highest contribution (of UFP to PNC) in the daytime compared with the nighttime in Delhi. Due to
408 the difference of PN size range (they measured down to 12nm), they found the UFP contribution to
409 PNC higher than in the present study.

410

411 **3.4 Size Distributions**

412 Figure 6 shows the average PNSD in three seasons in Delhi. Volume and Area distributions are shown
413 in the Supplementary Materials in Figure S10. The highest number concentrations are seen in winter,
414 followed by autumn, and then summer. Although the number concentrations of particles below 200
415 nm are far greater in winter those between 200 and 600nm are greater in autumn, within the
416 accumulation mode. The winter and summer PNSD show modes at approximately 100 nm but the
417 autumn PNSD shows the mode at approximately 200 nm. This could be due to changing sources of
418 particles in Delhi between seasons (Jain et al. 2020), in addition to (differing) aerosol dynamical
419 processes. The Delhi atmosphere is more polluted comparing with most other cities based on particle
420 number and mass (Harrison, 2020). This will cause a tendency for particles to grow more rapidly by
421 coagulation and condensation (Harrison et al., 2018), but this might be expected to occur in all
422 seasons.

423

424 As described above, in Delhi the nighttime particle concentrations are markedly higher than the
425 daytime concentrations. The PNSD changes for each hour of the day across all three seasons were
426 analysed (Figure S11) and categorized. Figure 7 presents the PNSD differences between daytime and
427 nighttime and shows the variation in PNSDs within the day in all seasons. The main difference
428 between day and night in winter is only the number concentration, with little change in the mode size
429 between day and night, while the PNSDs in summer and autumn show bimodal distributions with
430 modes at approximately 30 and 140 nm in summer, and 35 and 200 nm in autumn. When we focus
431 on PNSD during the daytime, it can be clearly seen that the modes are manifest at different times: In

432 winter, while the PNSD shows the same mode at approximately 100 nm from 8 am to 2 pm, the mode
433 in the afternoon (from 2 pm to 6 pm) drops slightly in size (70 nm). In the morning and afternoon
434 there are two small peaks at 60 nm and 40 nm for the Aitken fraction and 170 and 130 nm for the
435 accumulation fraction in autumn. During the day in summer, there are two peaks at approximately 30
436 nm from 10 am to 6 pm. This may be associated with summer nucleation events and NPF on 3rd June
437 2018 (Figure 4). Furthermore it may be related to the growth of particles from 10 am to 2 pm in
438 autumn and summer. The full reasons for these changing PNSDs are not clear, and it would be unwise
439 to attempt a detailed interpretation of a very small dataset.

440

441 Figure 8 shows the average geometric mean diameter (GMD) change with hour of the day. Two
442 overall periods of GMD increase are observed. One of them is in nighttime in all seasons with GMD
443 growing at between 4.6 nm/hour in summer and 6.2 nm/hour in winter. The particle growth in autumn
444 is predominantly (when compared to the winter and summer) both late in the night (from 0 am to 5
445 am) and in the morning (from 8 am to 12 pm). Considering the PNSD trend in autumn (Figure 7), the
446 GMD rises at 9 nm/hour from morning to noon. Similar results were obtained in the USA (Kuang et
447 al., 2012), Canada (Jeong et al., 2010; Iida et al. 2008), Italy (Hamed et al., 2007) and Japan (Han et
448 al. 2013). However the calculated GMD growth rate is smaller than that calculated by Sarangi et al.,
449 (2015; 2018) in Delhi, by Kalafut-Pettibone et al. (2011) in Mexico City and by Zhang et al. (2011)
450 in Beijing. The changing GMD with time in Delhi could be the result of changing sources, and/or of
451 dynamics. Nocturnal growth may be the result of reducing temperatures and increasing RH causing
452 vapour condensation (Sarangi et al., 2018). Morning growth may be due to oxidation processes
453 leading to production of less volatile vapours which then condense onto the particles (Sarangi et al.
454 2018).

455

456 Figure 9 gives the average particle number, volume, area and mass size distribution for all seasons.
457 While the number size distributions have one mode, two peaks are observed in volume distributions,

458 centered at 0.5 μm and 6 μm . These relate to two different main sources, which might be secondary
459 aerosol (such as sulphate at high RH) in the fine mode and road dust resuspension, soil or construction
460 dust for the coarse mode (Pant et al. 2016). In winter and autumn fine mode particle volumes are
461 higher than the coarse mode. However, in summer the coarse mode particle volumes are higher than
462 the fine particle level. In a recent paper, Thamban et al. (2021) show that modes in the mass size
463 distributions of hydrocarbon organic aerosol (HOA), Semi-volatile oxygenated organic aerosol
464 (SVOOA), biomass burning organic aerosol (BBOA) and low-volatile oxygenated organic aerosol
465 (LVOOA) measured by aerosol mass spectrometry are typically in the range 300-600nm vacuum
466 aerodynamic diameter, very consistent with the peaks seen in the mass distributions in Figure 9.

467

468 Hama et al. (2020) obtained the spatiotemporal characteristics of daily-averaged air pollutants and
469 concluded that the particulate matter mass (PM_{10} and $\text{PM}_{2.5}$) is dominated by local sources across
470 Delhi. The main local air pollutant sources in Delhi include traffic, construction, resuspension of dust,
471 diesel generators, power plants, industries and biomass burning (Kumar et al., 2013; Nagpure et al.,
472 2015; Hama et al., 2020).

473

474 All average PNSD graphs show an increasing trend in PNC at particle sizes below 19 nm particle
475 diameter. SMPS measurements in this study were conducted only above 15 nm. So, the peak particle
476 size within this size range cannot be seen. However, the clear increase in particle number below 19
477 nm indicates that another source may be important in Delhi. This small mode and bimodal PNSD
478 during the day (Figure 7) may be associated with the road transport vehicle types in Delhi. Despite
479 the diesel restriction during the rush hours and conversion of the public transport vehicles to CNG,
480 several studies have reported that $\text{PM}_{2.5}$ concentrations have been remaining steady or are slowly
481 increasing in India, especially in the winter and autumn seasons (Babu et al., 2013; Balakrishnan et
482 al., 2019; Dandona et al. 2017; Kumar et al., 2017).

483

484 The fuels used in Delhi's traffic fleet are petrol, diesel, CNG and LPG. Legislation limits the sulphur
485 content of the fuel to 50 ppm in diesel as per Bharat Stage IV. The diesel vehicles are not required to
486 be fitted with particle traps. The technology of the gasoline vehicle fleet varies as vehicle engine
487 capacity changes. Cars, two/three wheelers, bus and LCV fleet volumes are high during the day. Due
488 to the time restrictions on trucks/heavy good vehicles entering the city, during the daytime the HCV
489 number is at its lowest level (Figure S9).

490

491 Previous published studies indicate that emissions of particles from CNG vehicles (Euro 4, 5, 6) with
492 diameter greater than 23 nm are as low as a diesel particle filter equipped vehicle, and an order of
493 magnitude lower than gasoline vehicles (Kontses et al. 2020; Giechaskiel, et al. 2019; Magara-Gomez
494 et al. 2014; Schreiber et al. 2007), and CNG vehicles mainly emit nuclei-mode particles (Zhu et a.,
495 2014; Toumasatos et al. (2020). Zhu et al. (2014) calculated size-resolved particle emission factors
496 from on-road diesel buses and CNG buses and reported that the PNSD of diesel buses dominate the
497 accumulation mode diameters of 74-87 nm while the PNSD of CNG buses dominated the nucleation
498 mode with modes at 21-24 nm. Total PN emissions of diesel buses per vehicle were 4 times higher
499 than the level of CNG buses. However, the PN level in the nucleation mode (15-25 nm) of CNG buses
500 was 1.7 times higher than from the diesel buses in the nucleation mode. Toumasatos et al. (2020)
501 studied the particle emission performance of the Euro 6 CNG and gasoline vehicles and discussed the
502 current EU cut-off solid PN size threshold of 23 nm. The results revealed that $PN > 23$ nm represented
503 43 % of $PN > 10$ nm and 8 % of $PN > 2.5$ nm for gasoline vehicles and 7 % of $PN > 10$ nm and 1 % of
504 $PN > 2.5$ nm for CNG vehicles respectively. These studies of emission PNSDs show that a significant
505 number of particles reside below the EU lower measurement limit of of 23 nm, and many are even
506 smaller than 10 nm. These probably contribute to the mode seen just appearing at the extreme small
507 particle limit of Figure 9.

508

509 When the PNSD results measured in Delhi are compared with the main emission categories in the
510 literature (Kumar et al., 2013; Vu et al., 2015), it seems that the average size distributions measured
511 in the atmosphere in Delhi are much coarser, which is presumably due to condensation and
512 coagulation, or it could be that secondary particles dominate over the primary emissions. Pant et al.
513 (2016) hypothesised that the main accumulation mode peak in their winter measurements arose from
514 aqueous droplet evaporation, although this mechanism would be unlikely to explain the mode seen
515 in the summer data. Thamban et al. (2021) have also reported particle growth in the Delhi atmosphere
516 from condensation of organic compounds formed from oxidation processes.

517

518 Previous studies have attempted to quantify the relative contribution of primary and secondary
519 sources to the total and mode-segregated particle number concentrations (Kulmala et al. 2021;
520 Casquero-Vera et al. 2021; Hama et al. 2017; Kulmala et al. 2016; Rodríguez, & Cuevas, 2007).
521 Rodríguez and Cuevas (2007) first presented the methodology for the separation of traffic related
522 primary aerosol particles from the total using the BC as the main tracer of traffic. The method was
523 tested in this study, but did not prove appropriate as the BC sources in Delhi are more complex, and
524 arise not only from traffic. The BC diurnal trend (Figure 2) does not show the rush hour peaks, and
525 reflects mostly the combustion activity at night, presumably including the heavy duty diesel
526 emissions. A recent study by Kulmala et al. (2021) used NO_x as a tracer of primary sources. Figure 2
527 shows that only the NO₂ diurnal trend in autumn is clearly related to traffic sources. Furthermore the
528 sources of BC and NO_x are largely the same, as judged from the high similarity between BC and NO_x
529 diurnal trends (Figure 2).

530

531 **3.5 Correlations of PN with NO₂, NO, and BC**

532 Figure 10 shows the correlation coefficients between the hourly average PNs of five particle size
533 fractions and NO, NO₂, and BC measured in Delhi. Nucleation mode PN is better correlated with the
534 Aitken mode PN in winter and summer despite the lower correlation in autumn. The correlations

535 among $>1 \mu\text{m}$ size fractions are higher in summer than winter and autumn. Tyagi et al. (2016) stated
536 that the major source of NO_x emissions is vehicle exhaust and power plants in Delhi. Furthermore,
537 studies have reported that approximately 80-90 % of NO_x and CO are produced from the transport
538 sector in Delhi (Gurjar et al., 2004; Gulia et al., 2015; Tyagi et al., 2016; Hama et al., 2020). As seen
539 in Figure 2, the NO_2 diurnal trend is very similar to nucleation and Aitken particle trends, especially
540 in the autumn. NO_2 peaks in autumn in the traffic rush hours are larger than in winter and summer.
541 In addition, there are no significant correlations between NO_2 and NO or BC in autumn (0.02 for NO,
542 0.03 for BC) compared to the summer (0.73 for NO, 0.61 for BC) and winter (0.37 for NO and 0.28
543 for BC) (Figure S12). NO and BC diurnal trends show the same higher level in the night (Figure 2)
544 and also, they have higher correlation coefficients (0.78 in winter, 0.77 in summer, 0.72 in autumn)
545 for all seasons, similar to the accumulation mode particle counts (Figure 2, Figure 10, Table S2). NO_2
546 and $<100 \text{ nm}$ particles may be associated with traffic sources, while the NO and BC and $<1 \mu\text{m}$
547 particles could be associated with biomass burning, industry, (small generator) power generation, or
548 possibly also with diesel vehicles.

549

550 **3.6 Wind Effects**

551 Figure S13 represents polar plots of BC, NO and NO_2 measured in Delhi. This shows no consistent
552 pattern. There are differences between the pollutants in terms of directional and wind speed
553 associations, and for each pollutant / season. There is no obvious indication of a strong local source
554 influence, typically manifest as an intense area in the very centre of the plot circle. The plots for the
555 particle size fractions (Figure 11) also show little consistency between seasons for a given size
556 fraction. Within a season, however, adjacent size fractions often show a similarity of behaviour
557 (consistent with their correlations, see above) but this similarity does not extend across all size ranges
558 within a season.

559

560

561 **4. CONCLUSIONS**

562 This study serves to highlight the remarkable complexity of airborne particulate matter in Delhi. The
563 size distributions show marked seasonal changes, with coarse particles dominant in summer, but not
564 in the cooler seasons, when the accumulation mode dominates. The measured size distributions show
565 a fine mode aerosol with a considerably larger modal diameter than that typically seen in western
566 countries, and larger than the modal emission size from major source categories. It appears that
567 the high particle concentrations and chemically reactive atmosphere are promoting rapid coagulation
568 and condensational growth of particles, and therefore the measured size distributions are driven more
569 by aerosol dynamical processes than source characteristics. Growth via a liquid droplet phase in the
570 cooler months may also occur. There is little evidence for a contribution of new particle formation
571 (although the summer season dataset is small), consistent with earlier work by Gani et al. (2020).
572 Another notable feature is the apparent complexity and seasonal variability of sources of NO, NO₂
573 and BC, pollutants which can often be used to identify or locate sources of emissions. This is reflected
574 in the various particle fractions, which generally correlate poorly with the other pollutants and with
575 other than proximate size fractions.

576

577 The diurnal variation of all particle fractions is strongly suggestive of a road traffic influence,
578 especially in the winter campaign. This appears strongly influenced by the emissions of heavy duty
579 diesel traffic which is only able to access central Delhi at night. A size mode of <15 nm may well
580 be attributable to vehicles using LPG/CNG fuels. However, the seasonal variability of the geographic
581 distribution and wind speed dependence of sources revealed by the polar plots is strongly indicative
582 of many other sources also contributing to all size fractions of particles.

583

584 **DATA ACCESSIBILITY**

585 Data supporting this publication are openly available from the UBIRA eData repository at

586 <https://doi.org/10.25500/edata.bham.00000730>

587 **AUTHOR CONTRIBUTIONS**

588 This study was conceived by RMH. WJB managed the research programme, and MSA and LRC
589 collected the data. DCB and UAS led the data analysis with contributions from JB and DB. UAS
590 and RMH co-authored the first draft. ZS and all co-authors provided comments and revisions.

591

592 **COMPETING INTERESTS**

593 The authors declare that they have no conflict of interest.

594

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597

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604

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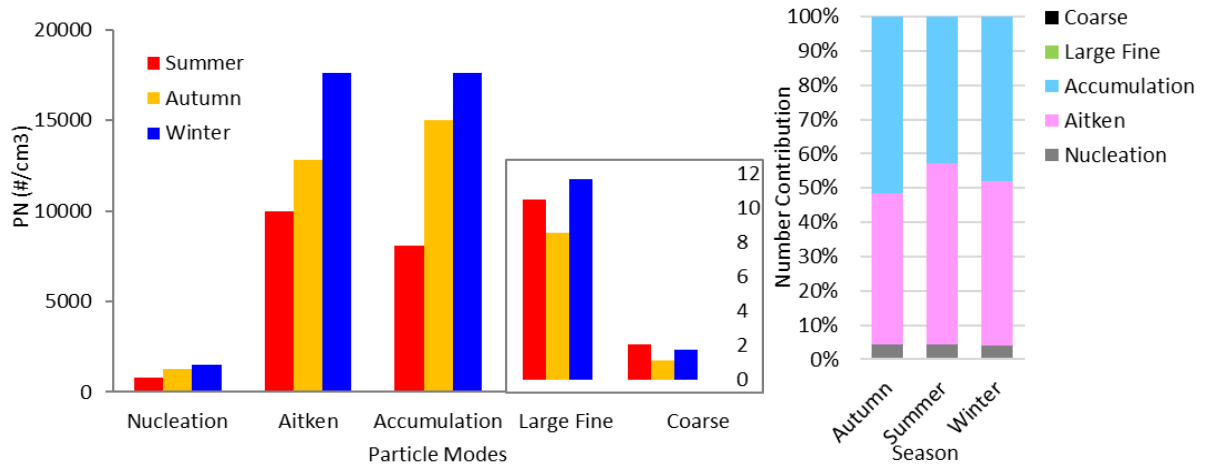
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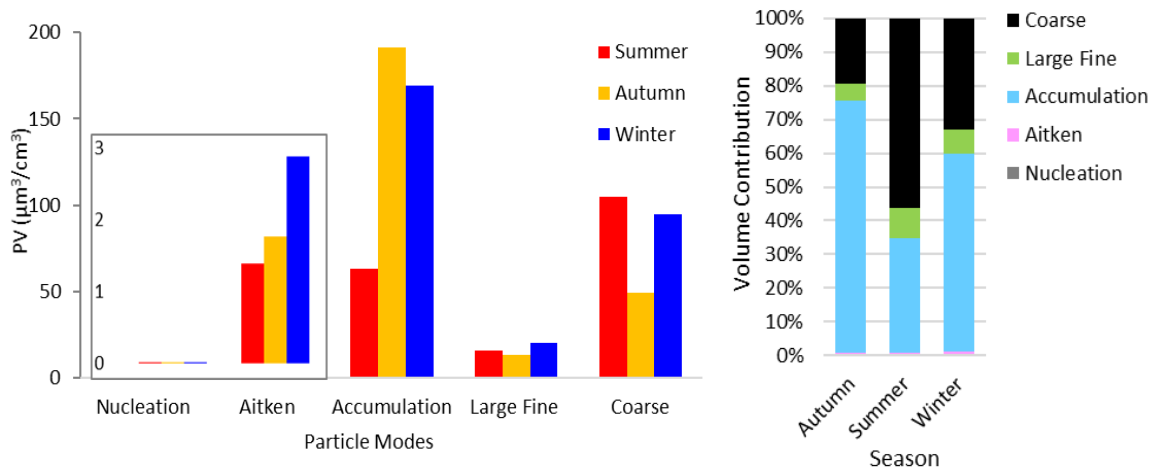
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- 969 **FIGURE CAPTIONS:**
970
971 **Figure 1.** Comparison of average particle number counts ($\#/cm^3$) for nucleation, aitken,
972 accumulation, large fine and coarse modes of PM between 15 nm and 10 μm in all
973 seasons, and the volume contributions for comparison.
974
975 **Figure 2.** Average diurnal variation of particle number counts for nucleation, Aitken,
976 accumulation, large fine and coarse modes and PM_{2.5}, BC, NO and NO₂ in autumn,
977 summer and winter.
978
979 **Figure 3.** Average diurnal variation of total particle number counts (between 15 and 10 μm) for
980 weekday average (Monday to Friday), Saturday and Sunday in Delhi. (Summer data
981 are very limited. There are no data on Friday afternoon, night and Saturday early
982 morning (Figure S6)).
983
984 **Figure 4.** Diurnal contour plots for particles derived by SMPS between 15 and 660 nm averaged
985 for each season (a: winter, b: Autumn and c: Summer) and for 3 June 2018 data when
986 a NPF event probably occurred (d), the solid line showing the NO_x mixing ratio.
987
988 **Figure 5.** The hourly average of day and night particle numbers for all modes from the wide
989 range particle sizes derived from the merged data. UFP = Nucleation +Aitken, PN₁ =
990 UFP+Accumulation, PN₁₀ = PN₁+Large Fine+Coarse.
991
992 **Figure 6.** Seasonal average (line) and standard deviation (shadow) of particle number size
993 distributions.
994
995 **Figure 7.** Hourly average day and night (left side) and during day hours (right side) particle
996 number distributions in autumn, summer and winter in Delhi.
997
998 **Figure 8.** Diurnal change of the geometric mean diameter (GMD) calculated for winter, autumn
999 and summer seasons. Growth rates (nm/hour) are calculated from $dGMD/dt$.
1000
1001 **Figure 9.** Hourly average particle number, volume and area distributions in the winter (a),
1002 autumn (b) and summer (c) in Delhi.
1003
1004 **Figure 10.** Correlation coefficient (R) between the hourly average PNs of five particle size
1005 fractions (left side) and NO, NO₂, BC (right side).
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1007 **Figure 11.** Polar plots of PNs ($\#/cm^3$) for five particle size fractions in winter, autumn and
1008 summer in Delhi.
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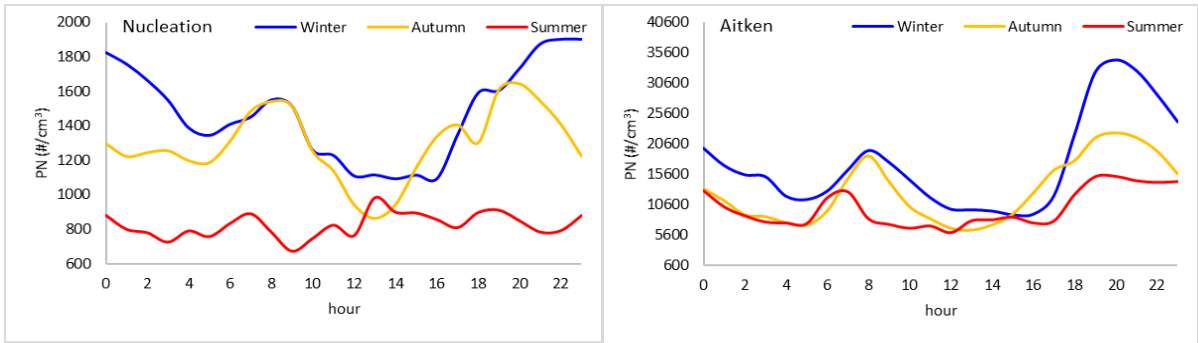


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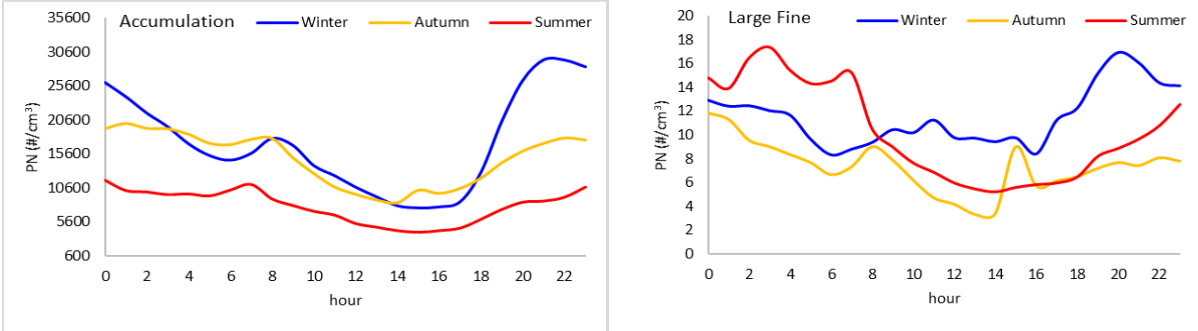
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1016 **Figure 1.** Comparison of average particle number counts ($\#/cm^3$) for nucleation, aitken,
 1017 accumulation, large fine and coarse modes of PM between 15 nm and 10 μm in all seasons, and the
 1018 volume contributions for comparison.
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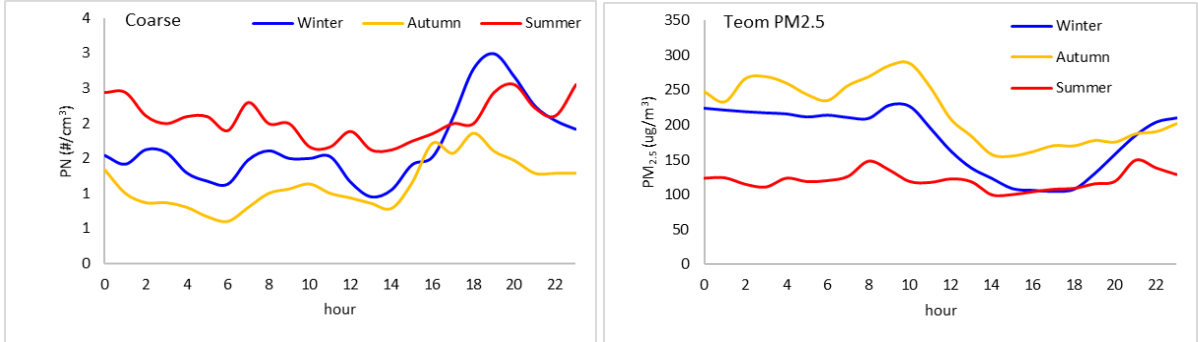
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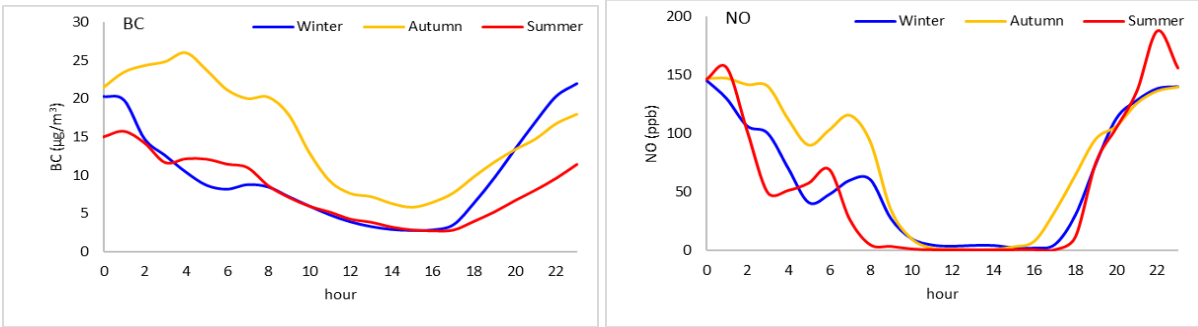
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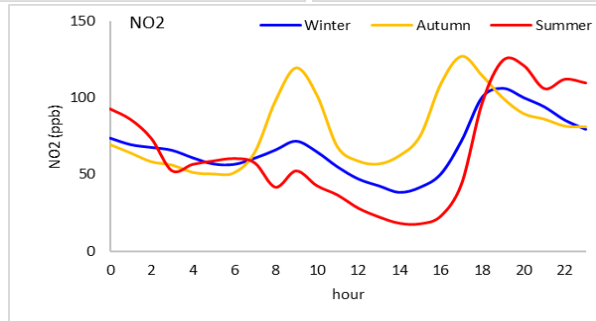
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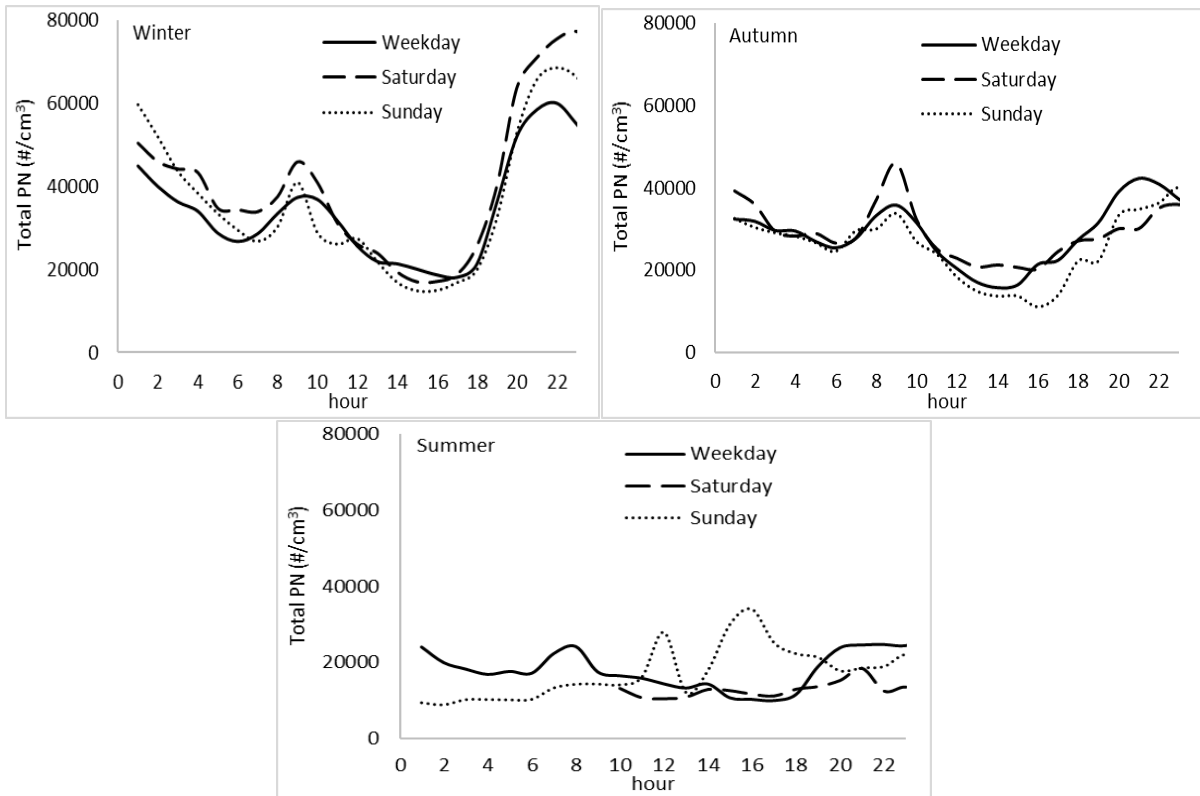
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Figure 2. Average diurnal variation of particle number counts for nucleation, Aitken, accumulation, large fine and coarse modes and PM_{2.5}, BC, NO and NO₂ in autumn, summer and winter.

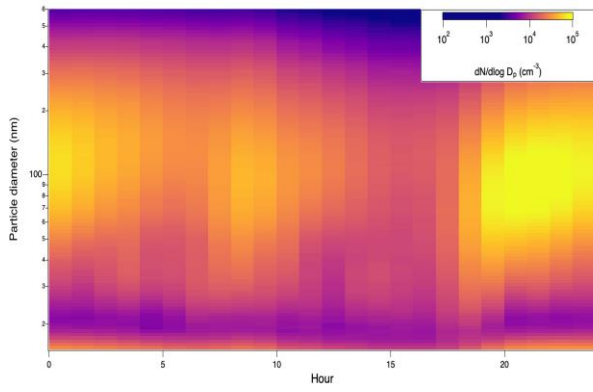
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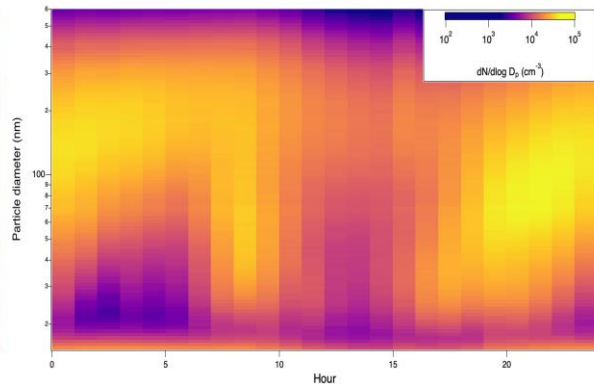
1032 **Figure 3.** Average diurnal variation of total particle number counts (between 15 and 10 μm) for
1033 weekday average (Monday to Friday), Saturday and Sunday in Delhi. (Summer data are very
1034 limited. There are no data on Friday afternoon, night and Saturday early morning (Figure S6)).
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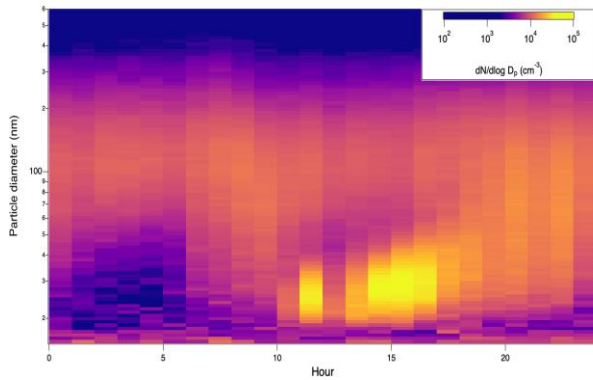
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a) Winter



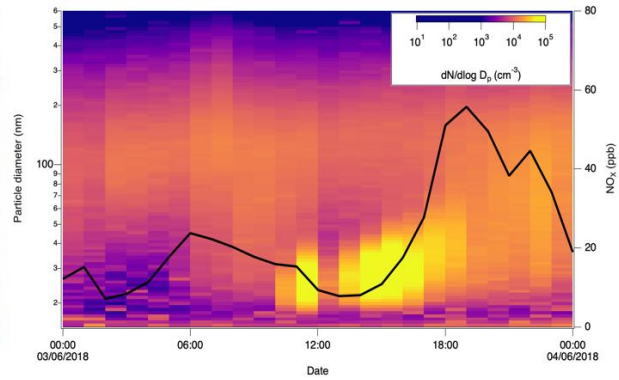
b) Autumn



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c) Summer



d) 3 June 2018

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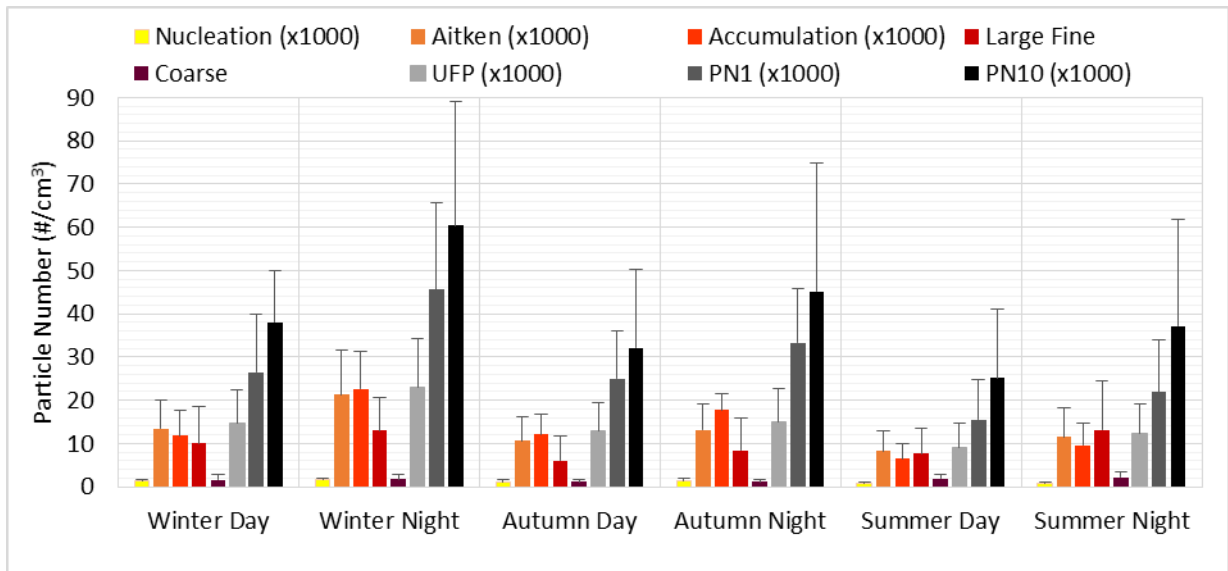
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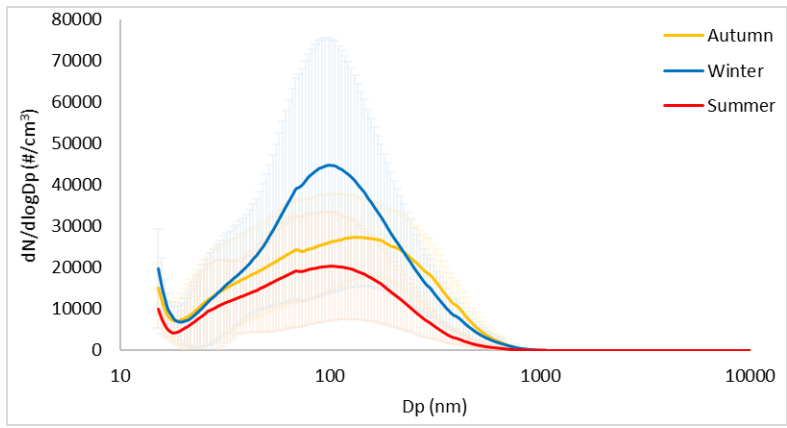
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Figure 4. Diurnal contour plots for particles derived by SMPS between 15 and 660 nm averaged for each season (a: winter, b: Autumn and c: Summer) and for 3 June 2018 data when a NPF event probably occurred (d), the solid line showing the NO_x mixing ratio.



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Figure 5. The hourly average of day and night particle numbers for all modes from the wide range particle sizes derived from the merged data. UFP = Nucleation +Aitken, PN_1 = UFP+Accumulation, PN_{10} = PN_1 +Large Fine+Coarse.

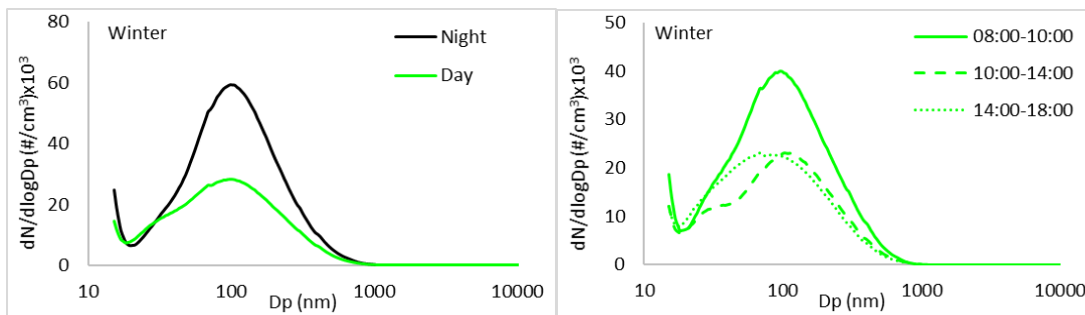


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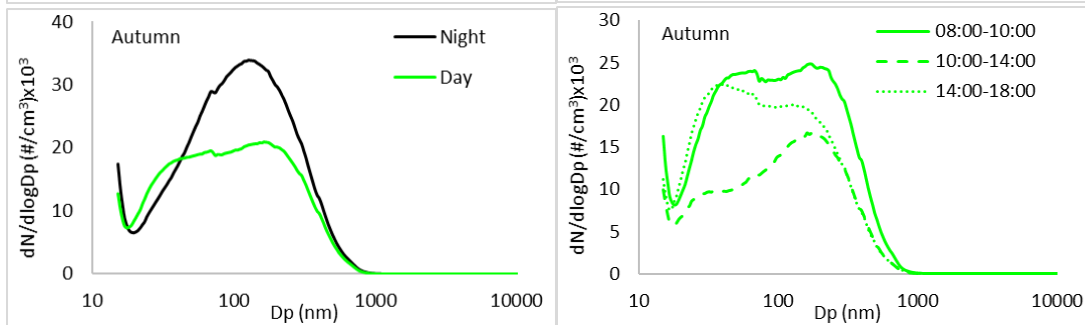
1052 **Figure 6.** Seasonal average (line) and standard deviation (shadow) of particle number size
1053 distributions.

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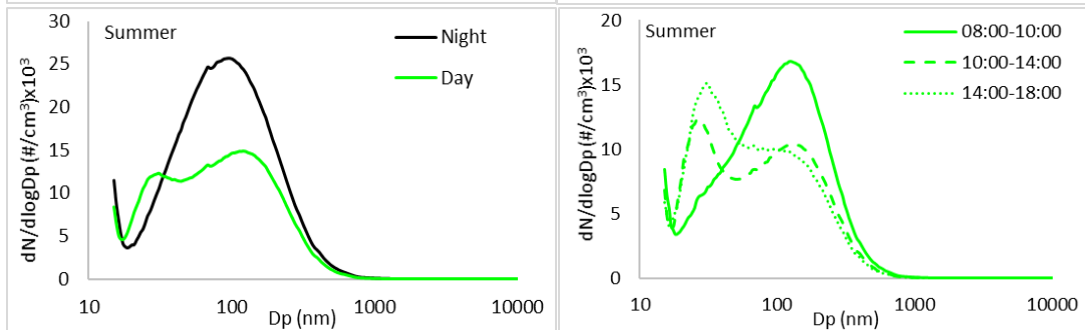
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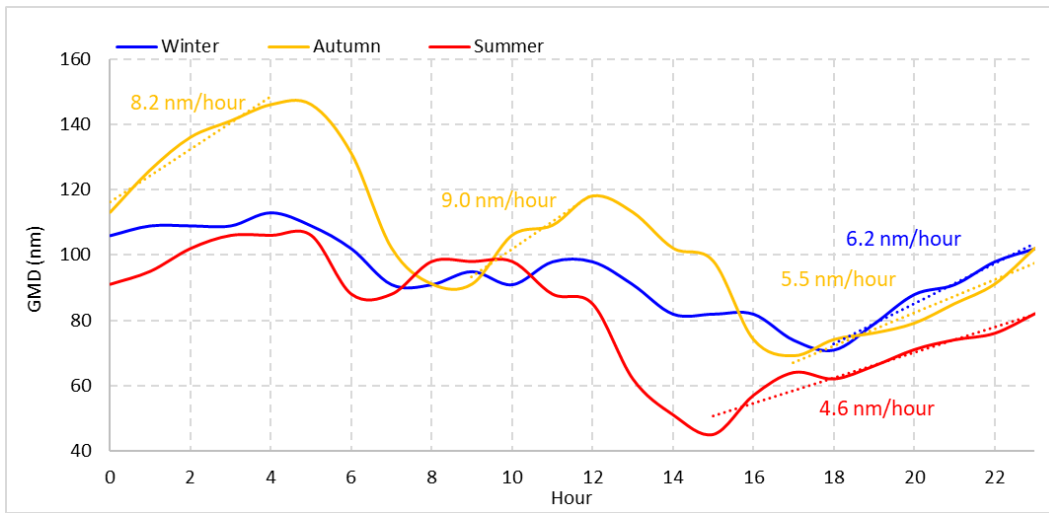


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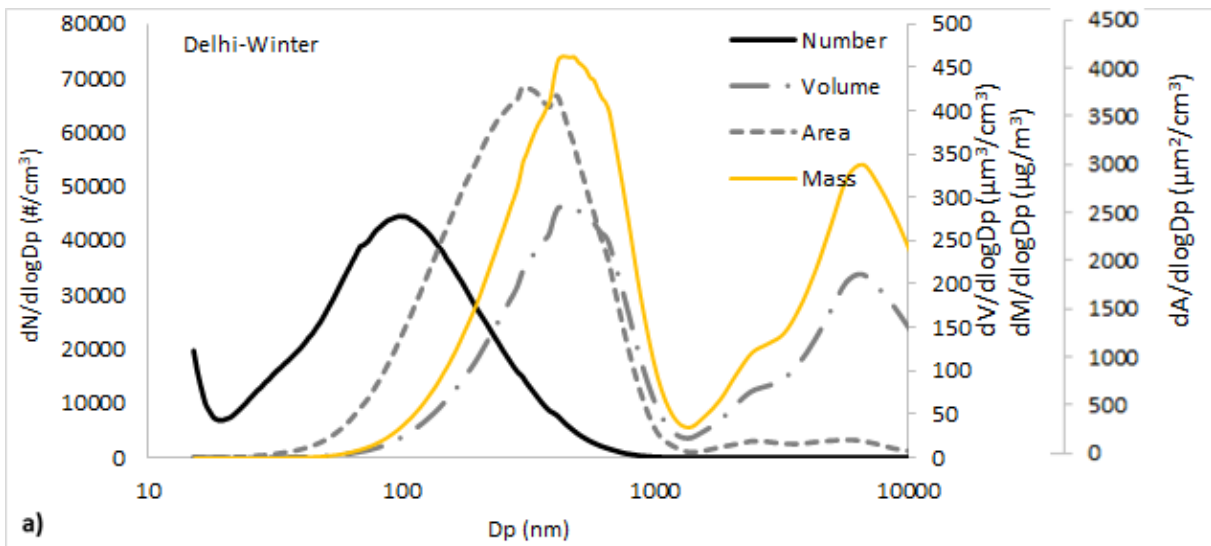
Figure 7. Hourly average day and night (left side) and during day hours (right side) particle number distributions in autumn, summer and winter in Delhi.



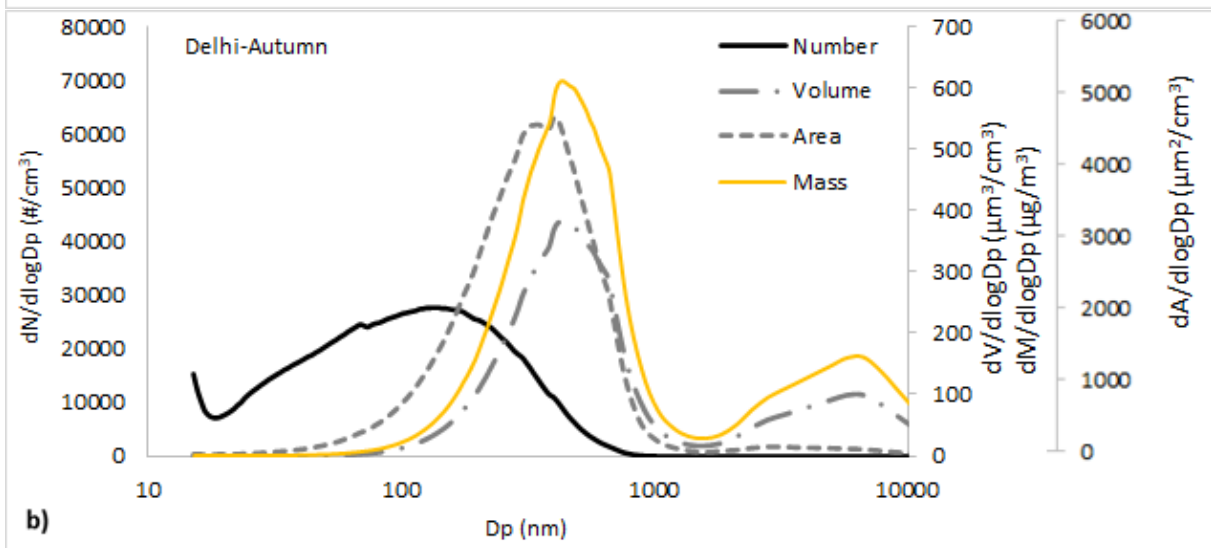
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1062 **Figure 8.** Diurnal change of the geometric mean diameter (GMD) calculated for winter, autumn
 1063 and summer seasons. Growth rates (nm/hour) are calculated from $dGMD/dt$.
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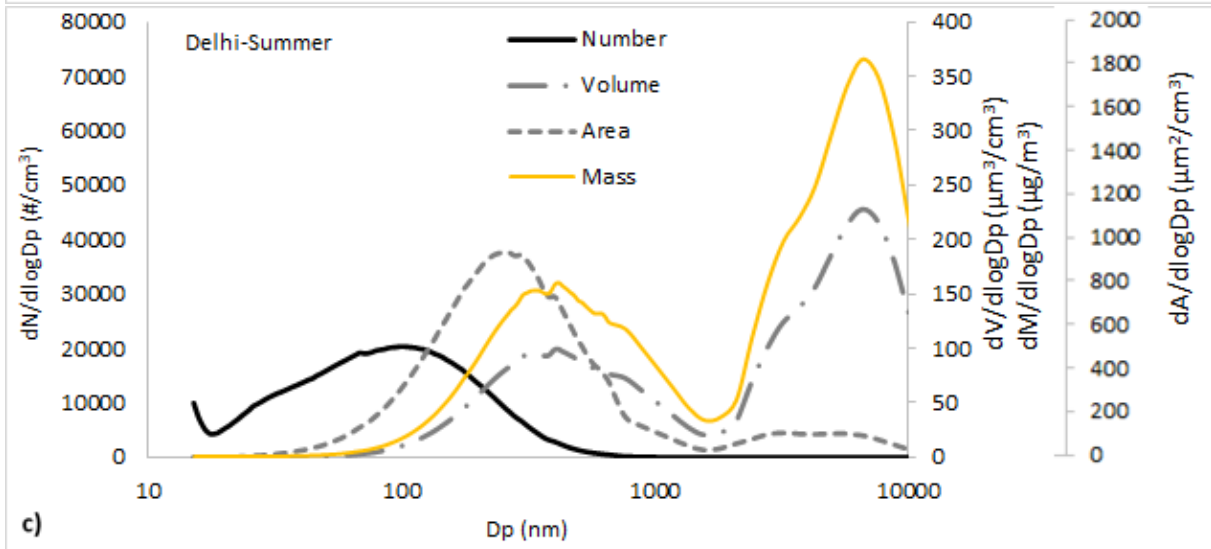
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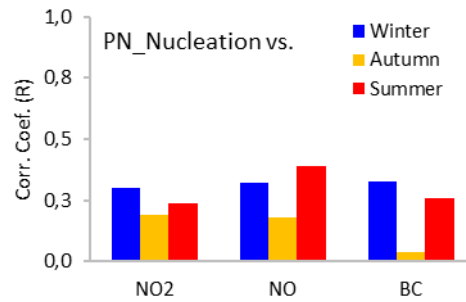
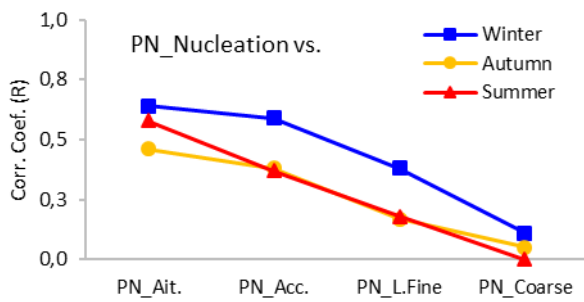


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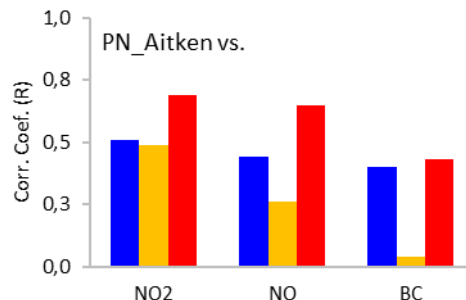
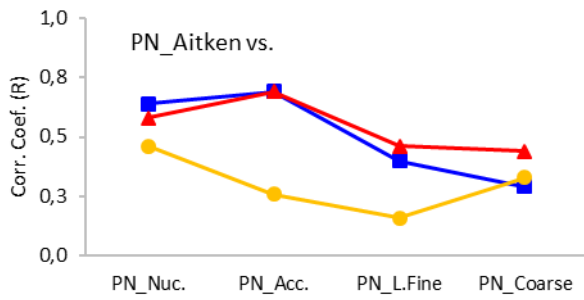


1068 **Figure 9.** Hourly average particle number, volume and area distributions in the winter (a), autumn
1069 (b) and summer (c) in Delhi.

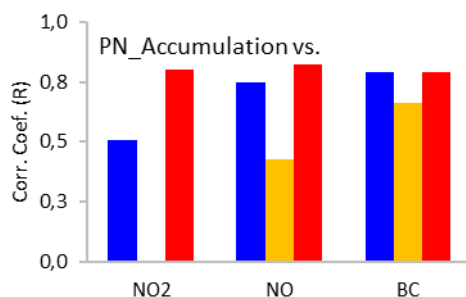
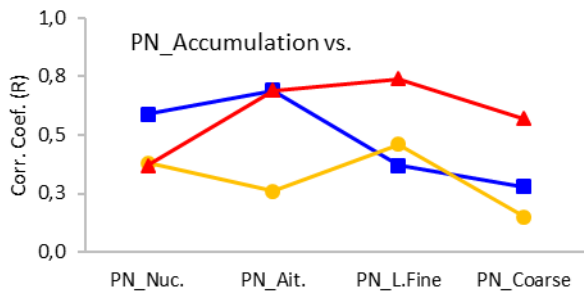
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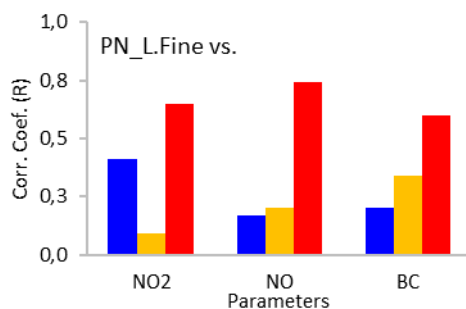
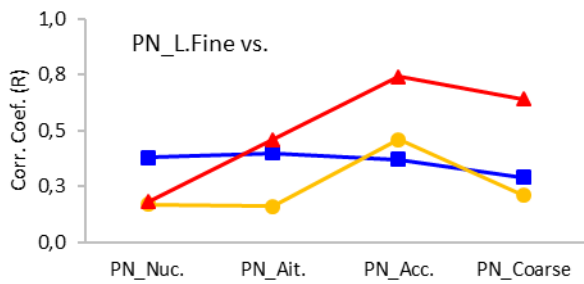
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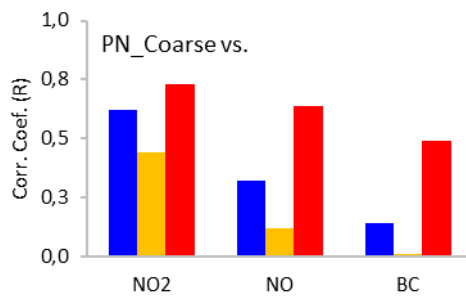
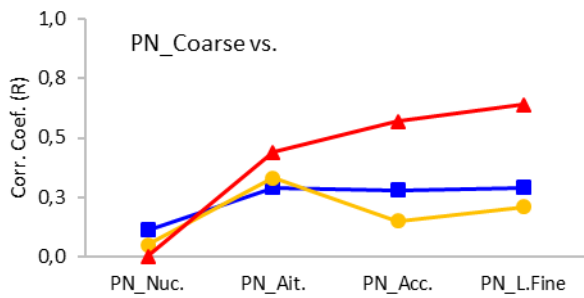
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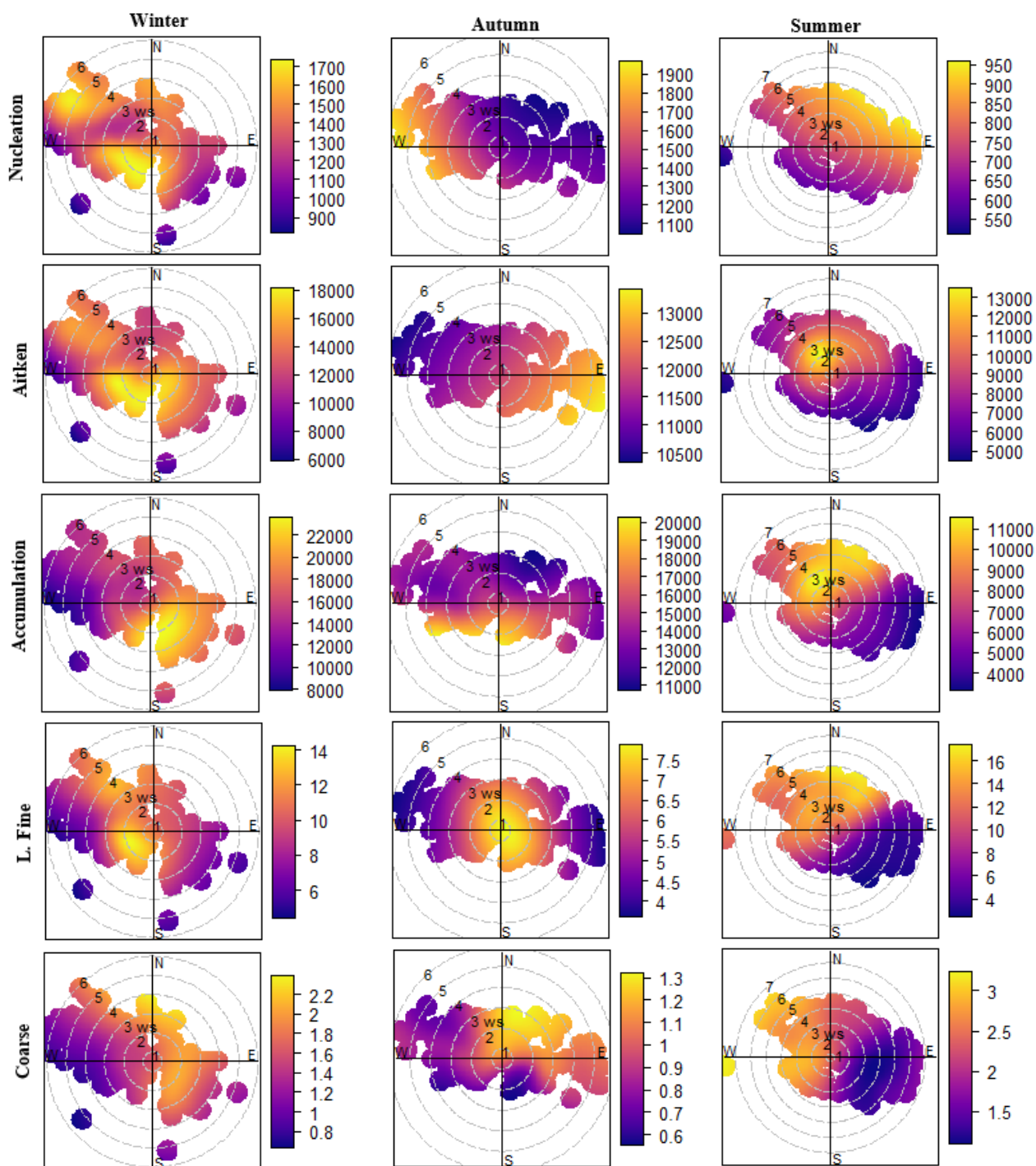
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1077 **Figure 10.** Correlation coefficient (R) between the hourly average PNs of five particle size
 1078 fractions (left side) and NO, NO₂, BC (right side).
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1082 **Figure 11.** Polar plots of PNs ($\#/cm^3$) for five particle size fractions in winter, autumn and summer
 1083 in Delhi.
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