

LES study on turbulent ~~particledust~~ deposition and its dependence on atmospheric boundary-layer stability

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Abstract. It is increasingly recognized that atmospheric boundary-layer stability (ABLS) plays an important role in aeolian processes. While the effects of ABLS on ~~particledust~~ emission have attracted much attention and been investigated in several studies, have-been-documented-in-several-studies, those on ~~particledust~~ deposition are so far less-well studied-are-less-well-studied. By means of large-eddy simulation, we investigate how ABLS influences the probability distribution of surface shear stress and hence ~~particledust~~ deposition. Statistical analysis of the model results reveals that the shear stress can be well approximated by using a Weibull distribution and the ABLS influences on ~~particledust~~ deposition can be estimated by considering the shear stress fluctuations. The model-simulated ~~particledust~~ depositions are compared with the predictions of a ~~particledust~~-deposition scheme and measurements, and the findings are then used to improve the ~~particledust~~-deposition scheme. This research represents a further step towards developing ~~dust deposition~~ schemes that account for the stochastic nature of ~~particledust~~ processes.

Keywords: ~~ParticleDust~~ deposition, Atmospheric boundary-layer stability, Surface shear stress, Weibull distribution, Stochastic ~~particledust~~ process

1 Introduction

Dry deposition is the removal of particulates and gases at the air-surface interface by turbulent transfer and gravitational settling (Schmel, 1980; Droppo, 2006; Hicks et al., 2016). Because it is the only process for the removal of particles from the atmosphere in the absence of precipitation, developing reliable methods for estimating dry deposition of particles has attracted much interest since the early 1940s ~~(Gregory, 1945; Gregory, 1945; Chamberlain, 1953; Slinn and Slinn, 1980; Slinn, 1982; Walcek et al., 1986; Zhang et al., 2001; Petroff and Zhang, 2010; Zhang and Shao, 2014; Seinfeld et al., 2016)~~. Several particle-deposition schemes have been proposed ~~(Slinn, 1982; Walcek et al., 1986; Zhang and Shao, 2014; Zhang et al., 2001)~~ ~~(e.g., Slinn, 1982; Zhang and Shao, 2014)~~ for regional/global models, which are driven by using several environmental parameters, including the Reynolds surface shear stress (typically averaged over 15-30 min). However, field observations indicate that the use of Reynolds stress as the only wind-related parameter in such schemes may not be sufficient to achieve accurate estimates of particle deposition, because of the nonlinear relationship between deposition velocity and wind shear. The observations using

the eddy correlation method show that particle-deposition velocity has strong spatiotemporal variations associated with the fluctuations of wind speed (Connan et al., 2018; Damay et al., 2009; Lamaud et al., 1994; Wesely et al., 1983, 1985). It is also observed that when the background wind speeds are similar, dry deposition velocities under convective conditions are larger than under neutral and stable conditions (Fowler et al., 2009). Pellerin et al. (2017) suggested that cospectral similarities exist between heat and particle-deposition fluxes and that atmospheric turbulence plays a role in ~~particledust~~ deposition. It is therefore necessary to find a link between instantaneous wind and particle deposition and to correctly represent this link in particle-deposition schemes, i.e., to introduce and account for the effect of turbulence on particle deposition.

~~Some aeolian processes~~Models for, e.g., turbulent ~~particledust~~ emission (Klose and Shao, 2012, 2013) and ~~intermittent sand~~ saltation (Li et al., 2020; Liu et al., 2018; Rana et al., 2020)(~~Liu et al., 2018; Li et al., 2020; Rana et al., 2020~~), have ~~been under development~~ been developed, but ~~To the best of~~ to the best of our knowledge, although turbulent ~~particledust~~ deposition is now perceived to be important, a scheme is yet to be constructed for its quantitative estimate.

The turbulent wind flow in a particle-deposition scheme is reflected in the turbulent shear stress (or vertical momentum flux) (Fowler et al., 2009; ~~Petroff and Zhang, 2010; Slinn, 1982; Zhang et al., 2013; Zhang and Shao, 2014~~). It is well-known that apart from gravitational settling, particle deposition is driven by turbulent diffusion which is intimately related to the vertical momentum transfer in the atmospheric boundary layer (ABL) (Wyngaard, 2010). Based on the Prandtl mixing-length theory, the shear stress can be parameterized in neutral conditions. However, it is known that for a given mean wind speed (at a reference height) in the ABL, both the mean value and the perturbations of shear stress depend on the atmospheric boundary-layer stability (ABLS), for instance, shear stress shows generally larger fluctuations in convective ABLS. (Klose and Shao, 2013) ~~Klose and Shao (2012)~~ pointed out that:

~~In a~~ In a convective atmospheric boundary layer conditions, large eddies have coherent structures of dimensions comparable to boundary-layer depth, which are efficient entities in generating localized momentum fluxes to the surface. Although the eddies only occupy fractions of time and space, the momentum fluxes to these fractions can be many times the average. (p. 49) Hicks et al. (2016) mentioned that ABLS is of immediate concern in the micrometeorological community, because of its influences on the intermittency, gustiness and diurnal cycle of particle deposition. Similar to ~~turbulent~~ dust emission and ~~intermittent~~ sand saltation, intermittent ~~particledust~~ deposition also occurs as a result of fluctuating surface shear stress. The current particle-deposition schemes only consider the mean behavior of wind (Petroff and Zhang, 2010; Slinn, 1982; Zhang and Shao, 2014; Zhang et al., 2001), and how this mean behavior varies with ABLS via the Monin-Obukhov similarity theory (Monin et al., 2007; Monin and Obukhov, 1954), but not the fluctuations of the associated shear stress and how they vary with ABLS.

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We argue that focusing only on the effects of ABLs on mean wind is insufficient to accurately model particle deposition-~~accurately~~. In this study, we explore the influences of ABLs on the turbulent behavior of particle deposition and attempt to improve an existing particle deposition scheme. A large-eddy simulation (LES) model is used here to simulate turbulence and particle deposition under various ABLs conditions, and parts of the study design follow Klose and Shao (2013). The ~~dust~~-particle depositions simulated using the LES model and predicted using the particle-deposition scheme of Zhang and Shao (2014, ZS14 hereafter) are compared with each other and with measurements. Specifically, Here, we address the following three issues: (1) How ABLs affects the probability distribution of surface shear stress; (2) How ABLs impacts on particle deposition; and (3) How the ZS14 scheme can be improved to account for the ABLs effect. On this basis, an improvement to the ZS14 scheme (also applicable to other schemes) is proposed. The remaining part of the paper is organized as follows: Sect. 2 gives a brief description of the Weather Research and Forecast – Large-Eddy Simulation Model with Dust module (WRF-LES/D), the ZS14 scheme, and the design of the numerical experiments. Sect. 3 discusses the findings of the numerical simulations and the improvement to the ZS14 scheme. The concluding remarks are given in Sect. 4.

2. Model/Method

2.1 WRF-LES/D

The WRF-LES/D used here is initially developed by Shao et al. (2013) and Klose and Shao (2013) by coupling the WRF-LES (Moeng et al., 2007; Skamarock et al., 2008) with a land-surface module and dust module. As demonstrated in the earlier studies, WRF-LES/D is a reasonably-well-established system for applications of-to simulating turbulence, turbulent particledust emission and transport for various ABLs conditions. WRF-LES is a three-dimensional and non-hydrostatic model for fully compressible flow. The model separates the turbulent flow into a grid-resolved component and a subgrid component. The $k-l$ subgrid closure (Deardorff, 1980) together with the turbulent kinetic energy TKE(TKE) equation (Skamarock et al., 2008) is based on nonlinear backscatter and anisotropic (Kosović, 1997; Mirocha et al., 2010) are used here. The governing equations in WRF-LES/D include the equations of motion, continuity equation, enthalpy equation, equation of state and the particledust conservation equation, as shown below

$$\frac{\partial u_i}{\partial t} + \frac{\partial u_i u_j}{\partial x_j} = -\delta_{i3} g + \varepsilon_{ij3} f u_j - \frac{1}{\rho_a} \frac{\partial p}{\partial x_i} + \nu \frac{\partial^2 u_i}{\partial x_j \partial x_j} - \frac{1}{\rho_a} \frac{\partial \tau_{ij}}{\partial x_j}$$

(1)

$$\frac{\partial \rho_a}{\partial t} + \frac{\partial (\rho_a u_j)}{\partial x_j} = 0$$

(2)

$$\frac{\partial c_p T}{\partial t} + \frac{\partial (c_p u_j T)}{\partial x_j} = \frac{\partial H_j}{\partial x_j} + s_r$$

(3)

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$$p = \rho_a R_a T \quad (4)$$

$$\frac{\partial c}{\partial t} + u_j \frac{\partial c}{\partial x_j} - w_i \frac{\partial c}{\partial z} = -\frac{\partial F_j}{\partial x_j} + s_r \quad (5)$$

where u_i (u, v, w) is the grid-resolved flow velocity along x_i (x, y, z) refer to the streamwise, spanwise, and vertical directions, respectively; g is the acceleration due to gravity; ρ_a is the air density; f is the Coriolis parameter; p is the air pressure; τ_{ij} is the subgrid stress tensor modeled using an eddy viscosity approach where the eddy viscosity is represented as the product of a length scale and a velocity scale characterizing the subgrid-scale (SGS) turbulent eddies (Dupont et al., 2013), with the velocity scale being derived from the SGS TKE and the length scale from the grid spacing (Skamarock et al., 2008). ν is the kinematic viscosity; δ_{ij} is the Kronecker operator and ε_{ijk} is the alternating operator; c_p is the specific heat of air at constant pressure; T is air temperature; H_j is the j th component of subgrid heat flux; R_a is the specific gas constant of air; c is the particledust concentration; w_i is the particle terminal velocity; F_j is the j th component of subgrid particledust flux; s_T and s_r are the source or sink terms for heat and particles, respectively. The subgrid eddy diffusivity is set to be obtained using subgrid eddy viscosity divided by dividing the Prandtl number. For the surface layer, an important parameterization to solve the governing equations for high-Reynolds-number turbulence is embedded in the surface boundary condition, which computes the instantaneous local surface shear stress using the bulk transfer method (Kalitzin et al., 2008; Kawai and Larsson, 2012; Piomelli et al., 2002; Zheng et al., 2020) as follows,

$$\tau = \rho_a K_m \frac{\partial V}{\partial z} \quad (6)$$

with

$$K_m = \frac{ku_* z}{\varphi_m} \quad (7)$$

where K_m is the eddy viscosity and φ_m is the MOST stability function; $V = \sqrt{u^2 + v^2}$. Even though Shao et al. (2013) questioned the application of the MOST in LES, it is still used here, as our emphasis is on the variance of shear stress in the simulation domain. Several land-surface models (LSMs) can be selected (e.g., Chen and Dudhia, 2000; Pleim and Xiu, 2003) in WRF-LES/D, and the 5-layer thermal diffusion (Dudhia, 1996) is used in this study. Surface heat flux in this study is artificially given. In addition, we denote the surface heat flux as H_0 , is specified. The and dust-dry deposition flux to the ground for each grid, denoted as F_{ds} , is obtained by multiplying the deposition velocity V_d and particle concentration c in the lowest layer, and V_d is estimated using the ZS14 deposition scheme.

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2.2 Particle-deposition scheme of ZS14

The ~~dust~~ particle deposition on the surface is more complicated than the momentum flux as the change of particledust concentration ~~changing~~ close to the surface is unclear. To solve the particledust conservation equation, (Eq. (5)), the emission and deposition fluxes at the surface need to be specified. The problem of particledust emission has been dealt with elsewhere (e.g., (Shao, (2004) focuses on the particle emission without turbulence effects; (Klose et al., (2014); and Klose and Shao, (2012)) emphasize the turbulent particle emission) (Shao, 2004; Klose and Shao, 2013) and is not considered here. For our purpose, particledust emission is assumed to be zero. This section gives the parameterization scheme ~~of surface~~ settlement proposed by ZS14. The detail of the scheme is as described in ZS14, only the main results are given here for completeness. In general, we can express particledust deposition flux F_d as

$$F_d = -\left(K_p + k_p\right) \frac{\partial c}{\partial z} - w_t \cdot c \quad (8)$$

where K_p and k_p are the eddy diffusivity and the molecular diffusivity, respectively. By analogy with the bulk-transfer formulation of scalar fluxes in ABL, F_d can be parameterized as

$$F_d = -V_d(z) \cdot c(z) \quad (9)$$

where $c(z)$ is the particledust concentration at height z (the center height of the lowest model level in this study), $V_d(z)$ is the corresponding dry deposition velocity.

The surface layer is divided into an inertial layer and a roughness layer. Integrating Eq. (8) in the inertial layer and substituting Eq. (9) into it, $V_d(z)$ is obtained as follows:

$$V_d(z) = \left(r_g + \frac{r_s - r_g}{\exp(r_a / r_g)} \right)^{-1} \quad (10)$$

With r_g being the gravitational resistance, r_s being the collection resistance, and r_a being the aerodynamic resistance for the inertial layer.

The gravitational resistance r_g is defined as the reciprocal of the gravitational settling velocity w_t and depends mainly on particle size and density. A free-falling particle is subject to gravitational and aerodynamic drag forces. When these forces are in equilibrium, the gravitational settling velocity of the particle smaller than $20 \mu\text{m}$ can be reasonably accurately calculated according to the Stokes formula (Malcolm and Raupach, 1991; Seinfeld and Pandis, 2006). When these forces are in equilibrium, the gravitational settling velocity reaches the terminal velocity given by the Stokes formula

$$w_t = \frac{C_u \rho_p D_p^2 g}{18 \mu_a} = r_g^{-1} \quad (12)$$

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where D_p is the particle diameter, ρ_p is the particle density, μ_a is the air dynamic viscosity, C_u is the Cunningham correction factor that accounts for the slipping effect affecting the fine particles.

Using the MOST, the aerodynamic resistance is calculated as we have

$$r_a = \frac{S_{CT}}{\kappa u_*} \left[\ln \left(\frac{z - z_d}{h - z_d} \right) - \psi_m \right]$$

(11)

where z_d is the displacement height, h is the height of roughness element, ψ_m is the integral of stability function in the inertial layer, $S_{CT} = K_m/K_p$ (Csanady, 1963), and κ is the von Karman constant. The gravitational resistance r_g is defined as the reciprocal of the gravitational settling and depends mainly on particle size and density. A free-falling particle is subject to gravitational and aerodynamic drag forces. When these forces are in equilibrium, the gravitational settling velocity reaches the terminal velocity given by the Stokes formula

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where D_p is the particle diameter, ρ_p is the particle density, μ_a is the air dynamic viscosity, C_u is the Cunningham correction factor that accounts for the slipping effect affecting the fine particles.

In the roughness layer, the collection process is reflected in collection resistance, defined by $r_s = -\frac{c(h)}{F_d}$ with an assumption of that particulate concentration is zero on roughness elements or ground. In addition to the meteorological factors and land-use category, Zhang and Shao (2014) established a relationship between aerodynamic and surface-collection processes by using an analogy between drag partition and deposition flux partition, which can describe surface heterogeneity.

$$r_s^{-1} = R \cdot \frac{\tau}{\rho_a u_h} \left(\frac{E}{C_d} \frac{\tau_c}{\tau} + \left(1 + \frac{\tau_c}{\tau} \right) S_c^{-1} + 10 \frac{3}{\hat{T}} \right) + w_t \quad (13)$$

where Here, R is the reduction in collection caused by particle rebound, u_h is the wind speed at the top of roughness layer, E is the collection coefficient of roughness elements and it includes the collection efficiency from Brownian motion (E_B), impaction (E_{im}) and interception (E_{in}), C_d is the drag coefficient for isolated roughness element, τ_c/τ describes the drag partition with τ_c being the pressure drag (the force exerted on roughness elements), R is the reduction of collection caused by particle rebound, E is the collection coefficient of the roughness elements, which includes the contributions of Brownian motion, impaction and interception, S_c is the Schmidt number which is the ratio of air viscosity to molecular diffusion, u_h is the wind speed at the top of roughness layer, C_d is the drag coefficient for isolated roughness

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element, $10^{-\frac{3}{\hat{T}}}$ represents the turbulent impactation efficiency with \hat{T} being the dimensionless particle relaxation time. E_B , E_{im} , E_{in} , and R are expressed as

$$E_B = C_B S_c^{-2/3} Re^{n_B-1} \quad (14)$$

$$E_{im} = \left(\frac{S_t}{0.6 + S_t} \right)^2 \quad (15)$$

$$E_{in} = u_* 10^{-S_t} \frac{2D_p}{d_c} \quad (16)$$

$$R = \exp(-\sqrt{S_t}) \quad (17)$$

where Re is the roughness element Reynolds number, C_B and n_B are parameters depending on Re , and d_c is the diameter of the roughness element, and S_t is the Stokes number.

The ratio τ_c/τ The ratio τ_c/τ describes the drag partition with τ_c being the pressure drag (the force exerted on roughness elements) and can be calculated according to Yang and Shao (2006), as follows: as

$$\frac{\tau_c}{\tau} = \frac{\beta_1 \lambda_e}{1 + \beta_1 \lambda_e}$$

(18)

and

$$\lambda_e = \frac{\lambda}{(1-\eta)^6} \cdot \exp\left(-\frac{\lambda}{10 \cdot (1-\eta)^6}\right)$$

(19)

with β_1 (≈ 200) being the ratio of the pressure-drag coefficient for isolated roughness element to that of for bare surface, λ being the frontal area index of the roughness elements, and η being the basal area index or the fraction of cover.

From Eqs. (10) to (19), it can be seen that V_d and τ are nonlinearly related. As example, for the particles with a diameter of 1 μm , analysis shows that when τ is small, V_d is dominated by w_t when τ is small. As τ increases, w_t and τ are both important to V_d . With τ increasing further, the effect of τ becomes much greater than gravity settling, thus the V_d is mainly determined by τ .

2.3 Simulation Set-up

Numerical experiments are carried out with WRF-LES/D for various atmospheric stability and background-wind conditions for two different roughness lengths (Table 1). Similar to (Klose and Shao, (2013), the domain of the simulation is $2000 \times 2000 \times 1500 \text{ m}^3$ and the number of grid points is $200 \times 200 \times 90$ corresponding to a horizontal resolution $\Delta x = \Delta y = 10 \text{ m}$. The Arakawa-C staggered grid is used. The depth

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of the lowest model layer is 1 m and the grid above is stretched following a logarithmic function of z . The simulation time is 90 minutes with a time step of 0.05 s and the output interval is 10 s. The first 30 minutes of the simulation is the model spin-up time and the data of the remaining 60 minutes are used for the analysis.

For model initialization, the wind and dust concentration (Chamberlain, 1967; Monin, 1970; Kind, 1992) are assumed to be logarithmic in the vertical and uniform in the horizontal direction. For each experiment, a constant surface heat flux is specified. A 300 m deep Rayleigh damping layer is used at the upper boundary with a damping coefficient of 0.01. The wind speed at the top boundary U_{z_s} is given in Table 1. The surface heat flux, H_0 , increases from -50 to 600 W m⁻², and for each surface heat flux, the top wind speed conditions increase from 4 to 16 m s⁻¹ in Exp (1-20) and from 5.44 to 18.12 m s⁻¹ in Exp (21-35). The roughness length z_0 for sand surface used in Exp (1-20) is 0.153 mm following wind tunnel experiment (Zhang and Shao, 2014) but 0.76 mm in Exp (21-35) according to field observation (Bergametti et al., 2018). The lateral boundary conditions are periodic, which allows the simulation of a well-developed boundary layer. The vertical scaling velocity is estimated using heat flux,

$$w_* = \left(\frac{g}{\bar{\theta}} \frac{H_0}{\rho_a c_p} z_l \right)^{1/3}, \text{ with } \bar{\theta} \text{ being the mean potential temperature and } z_l = 1000 \text{ m is the boundary layer}$$

inversion height. Usually, w_* is not used for stable ABLs, but used here as an indicator for the suppression of turbulence by negative buoyancy.

Table 1: List of numerical experiments with $z_0 = 0.153$ mm for Exp (1-20) in wind tunnel experiments (Zhang and Shao, 2014) and $z_0 = 0.76$ mm for Exp (21-35) in field observation (Bergametti et al., 2018) for the sand surface.

$z_0 = 0.153$ mm		$z_0 = 0.76$ mm		H_0 (W m ⁻²)	w_* (m s ⁻¹)
NAME	U (m s ⁻¹)	NAME	U (m s ⁻¹)		
EXP1	4	EXP21	5.44	-50	-1.12
EXP2	8	EXP22	10.87	-50	-1.12
EXP3	12	EXP23	18.12	-50	-1.12
EXP4	16	--	--	-50	-1.12
EXP5	4	EXP24	5.44	0	0
EXP6	8	EXP25	10.87	0	0
EXP7	12	EXP26	18.12	0	0
EXP8	16	--	--	0	0
EXP9	4	EXP27	5.44	200	1.77
EXP10	8	EXP28	10.87	200	1.77
EXP11	12	EXP29	18.12	200	1.77
EXP12	16	--	--	200	1.77
EXP13	4	EXP30	5.44	400	2.23
EXP14	8	EXP31	10.87	400	2.23
EXP15	12	EXP32	18.12	400	2.23
EXP16	16	--	--	400	2.23
EXP17	4	EXP33	5.44	600	2.55
EXP18	8	EXP34	10.87	600	2.55
EXP19	12	EXP35	18.12	600	2.55
EXP20	16	--	--	600	2.55

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3. Results

3.1 Turbulent shear stress

In the first set of the analysis, we examine the impact of atmospheric stability on shear stress fluctuations. Early particulate deposition studies considered only the time average of surface shear stress, τ_r , with the assumption that shear stress is horizontally homogeneous. In WRF-LES/D, the corresponding mean resultant shear stress τ_r can be obtained as below:

$$\tau_r = \sqrt{\bar{\tau}_{xz}^2 + \bar{\tau}_{yz}^2} \quad (2016)$$

The shorthand notation $\bar{f} = \frac{1}{N_x N_y N_t} \sum_{n_x, n_y, n_t} f(n_x, n_y, n_t)$ is introduced to represent the space and time average over the simulation domain and time period (hereafter ensemble mean) with N_x (=200) and N_y (=200) are the numbers of grid points in the x - and y -direction, respectively, and N_t (=360) the time steps of model output.

Figure 1a-c show the instantaneous shear stress, τ , of a sample grid ($n_x = 198$, $n_y = 41$) over a one-hour period for the runs with $z_0 = 0.153$ mm, $U = 4$ m s⁻¹ and various ABL stabilities ($H_0 = 0, 200, 600$ W m⁻²). Figure 1d-f is same as Fig. 1a-c, but for $U = 16$ m s⁻¹. The panel shows that τ is not a constant, and the mean resultant shear stress, as well as the shear stress fluctuations, increase with increasing atmospheric instability. In addition, the insert plots in Fig. 1 show that the autocorrelation function, ACF, is oscillated during decrease. The oscillation periodicity is longer under weak wind conditions (Fig. 1a-c) than strong wind (Fig. 1d-f). The ACF in neutral conditions decreases more rapidly than in convective conditions. Recall the definition of coherent motion given by Robinson (1991) - the correlation of variables over a range of long time larger than the smallest scales of flow is an evidence of coherent oscillating motion. Thus, the regular oscillation and a long-time correlation of τ are closely related to the evolution of the coherent structure. This indicates that in a convective ABL, stronger large-scale coherent structures exist even under weak wind conditions.

To gain insight into the behavior of the unsteady shear stress field, we introduce the turbulence intensity of surface shear stress (TI-S) defined as the ratio of the standard deviation of fluctuating surface shear stress, σ_τ , to the mean resultant stress τ_r , i.e., σ_τ/τ_r . Analysis shows that σ_τ/τ_r increases as atmospheric conditionss become more unstable and decreases with increasing wind speed (e.g., Fig. 1). High wind speeds tend to force the ratio to be more similar to that in neutral ABLs, as the mean-wind induced shear stress becomes dominant over the large-eddy induced shear-stress fluctuations. For a weak TI-S, τ is dominated by τ_r and the stress fluctuations are small compared to τ_r . As TI-S increases, the contribution of momentum transport by large eddies becomes significant because in unstable ABLs,

buoyancy generated large eddies penetrate to high levels and intermittently enhance the momentum transfer to the surface.

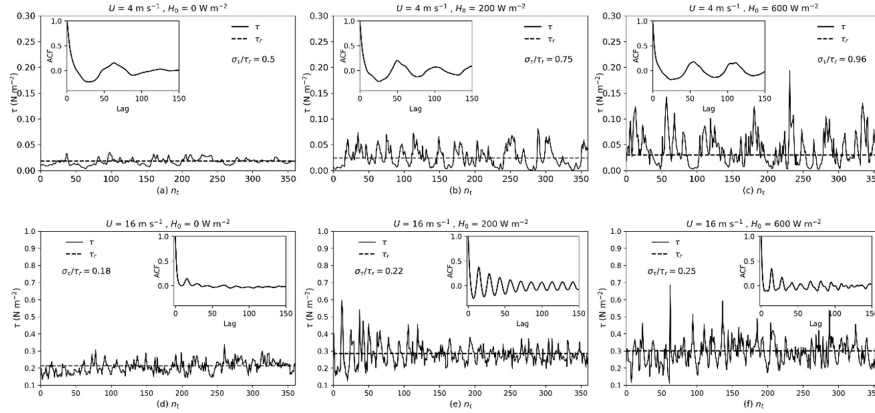


Figure 1. Time evolutions of surface shear stress τ with different H_0 values and $z_0 = 0.153$ mm at the grid point $n_x = 198$ and $n_y = 41$ (a-c) for $U = 4$ m s⁻¹; (d-f) for $U = 16$ m s⁻¹; the insert plots are the autocorrelation functions of τ .

The intermittent surface shear stress can directly cause localized particle dust-deposition. Therefore, particle dust-deposition is also intermittent in space and time. However, to our knowledge, in existing particledust-deposition schemes (e.g., ZS14 used here), the particledust-deposition velocity is calculated using only the mean resultant shear stress τ_r , instead of the instantaneous shear stress. We denote this deposition velocity as V_{d,τ_r} . The mean deposition velocity simulated by WRF-LES/D, denoted as $V_{d,LES}$, is estimated via the ratio of the ensemble mean of particledust deposition flux and the ensemble mean of particledust concentration:

$$V_{d,LES} = -\frac{\bar{F}_d}{\bar{C}} \quad (21)$$

which is consistent with the methods commonly used in field observations and wind-tunnel experiments.

Figures 2a and 2b, with the same wind conditions and surface heat fluxes as for as in Fig.1a-c and 1d-f, show the time evolution of the instantaneous deposition velocity V_d for particles with a diameter of 1.46 μm and surface heat flux $H_0 = 600$ W m⁻². This size is chosen because it is the most sensitive to turbulent diffusion compared to the other four sizes (2.8, 4.8, 9, 16 μm) used in Exp (1-20). As shown, the fluctuating behavior of V_d is consistent with that of τ . Moreover, Fig. 2a shows a substantial difference between $V_{d,LES}$ and V_{d,τ_r} , while Fig. 2b shows V_{d,τ_r} is similar with $V_{d,LES}$. This suggests that the ZS14 scheme can more accurately estimate the deposition velocity for weak TI-S but underestimates the deposition velocity for strong TI-S. The reason for this is that in the case of strong TI-S, particledust

deposition caused by the gusty wind plays an important role as V_d and τ are non-linearly related, which is not reflected in V_{d,τ_r} . Since τ fluctuates and sometimes strongly, a bias always exists in conventional **particledust**-deposition schemes and the magnitude of the bias depends on turbulence intensity. Therefore, in order to estimate **particledust** deposition accurately, we need to first describe and parameterize the shear stress.

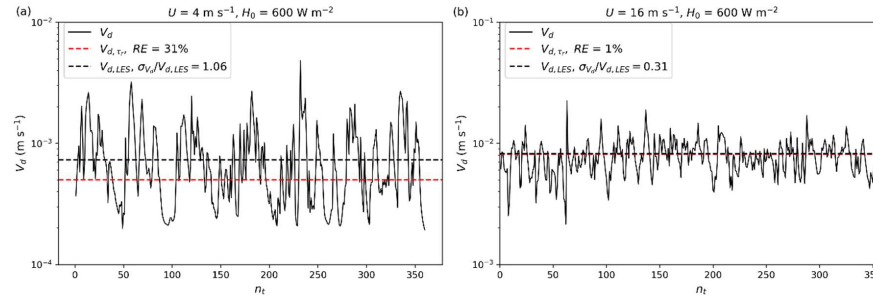


Figure 2. Time evolutions of deposition velocity V_d at grid point $n_x = 198$, $n_y = 41$ when $H_0 = 600 \text{ W m}^{-2}$, $z_0 = 0.153 \text{ mm}$ and (a) $U = 4 \text{ m s}^{-1}$ and (b) $U = 16 \text{ m s}^{-1}$. $RE = \left| \frac{V_{d,LES} - V_{d,\tau_r}}{V_{d,LES}} \right| \times 100\%$ is the relative error between V_{d,τ_r} and $V_{d,LES}$, $\sigma_{V_d}/V_{d,LES}$ is the ratio of the standard deviation of simulated instantaneous deposition velocity V_d and mean deposition velocity, $V_{d,LES}$.

As a main predisposing factor for aeolian processes, turbulent shear stress has attracted **increasing much** attention **in recent years** (e.g., Klose et al., 2014; Li et al., 2020; Liu et al., 2018; Rana et al., 2020; Zheng et al., 2020) (e.g., Klose et al., 2014; Li et al., 2020a; Liu et al., 2018; Rana et al., 2020; Zheng et al., 2020). Similar to previous studies, we use the probability density function $p(\tau)$ to characterize the stochastic variable τ . Figure 3 shows that the variability of τ increases as atmospheric instability increases **in different** **wind conditions**. The statistic moments of τ , including its mean resultant value τ_r , standard deviation σ_τ , skewness γ_1 of Exp (1-20) are listed in Table 2. σ_τ and τ_r **increases** with increased instability, and the distribution is positively skewed. Positive skewness is characterized by the distribution having a longer positive tail as compared with the negative tail and the distribution appears as a left-leaning (i.e., tends toward low values) curve. This indicates that large negative fluctuations are not as frequent as large positive fluctuations. The data also shows γ_1 generally shows a downward trend as TI-S decreases, which is consistent with (Monahan, 2006), i.e., as TI-S decreases, $p(\tau)$ becomes increasingly Gaussian.

Table 2. Statistics of shear stress for numerical experiments Exp (1-20).

NAME	H_0	U	τ_r	σ_τ	σ_τ/τ_r	γ_1	α	β	λ/L_0
EXP1	-50	4	0.0156	0.0086	0.554	1.902	2.026	0.011	0.475
EXP2	-50	8	0.0295	0.0096	0.327	1.573	3.154	0.023	0.153
EXP3	-50	12	0.0524	0.0115	0.22	1.029	3.923	0.044	0.06
EXP4	-50	16	0.1009	0.0158	0.157	0.835	4.819	0.09	0.02
EXP5	0	4	0.0185	0.0093	0.5	1.896	3.049	0.017	0

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EXP6	0	8	0.0604	0.0151	0.25	1.142	5.004	0.055	0
EXP7	0	12	0.1315	0.0266	0.202	0.166	5.383	0.122	0
EXP8	0	16	0.2136	0.038	0.178	0.087	6.191	0.196	0
EXP9	200	4	0.024	0.018	0.75	1.142	1.56	0.025	-0.696
EXP10	200	8	0.0812	0.0325	0.4	1.02	3.022	0.076	-0.11
EXP11	200	12	0.1676	0.0451	0.269	0.512	4.078	0.156	-0.037
EXP12	200	16	0.2848	0.0624	0.219	0.766	5.214	0.259	-0.017
EXP13	400	4	0.026	0.0248	0.955	1.127	1.302	0.03	-1.258
EXP14	400	8	0.0825	0.0372	0.451	0.646	2.513	0.081	-0.216
EXP15	400	12	0.1728	0.0522	0.302	0.677	3.776	0.160	-0.071
EXP16	400	16	0.2992	0.0646	0.216	0.289	5.214	0.278	-0.031
EXP17	600	4	0.0299	0.0287	0.96	1.083	1.303	0.035	-1.575
EXP18	600	8	0.0894	0.0424	0.474	0.715	2.472	0.089	-0.29
EXP19	600	12	0.1767	0.0604	0.342	0.614	3.252	0.167	-0.103
EXP20	600	16	0.3003	0.0739	0.246	0.511	4.493	0.277	-0.046

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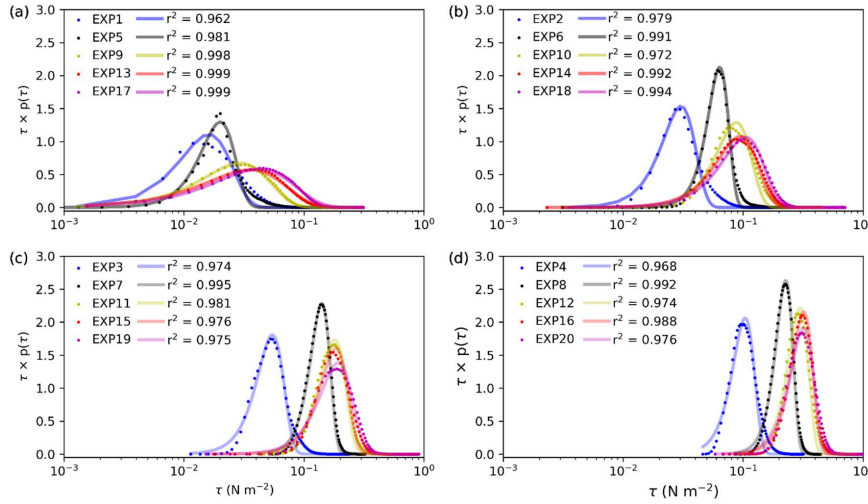


Figure 3. Probability density functions derived from WRF-LES/D simulated surface shear stress (dots) and the corresponding fitted Weibull density functions (solid lines, r^2 is the coefficients of determination) for different surface heat fluxes and different wind speeds: (a) $U = 4 \text{ m s}^{-1}$, (b) $U = 8 \text{ m s}^{-1}$, (c) $U = 12 \text{ m s}^{-1}$, (d) $U = 16 \text{ m s}^{-1}$ with $z_0 = 0.153 \text{ mm}$.

The parameterization of surface shear stress has attracted intense interests, [for example, \(Klose et al., \(2014\). Klose et al. \(2014\)](#) reported that τ in unstable conditions is Weibull distributed based on large-eddy simulations. Shao et al. (2020) found that $p(\tau)$ is skewed to small τ values (i.e., positively skewed) based on field observations. [Li et al., \(2020\) Li et al. \(2020\)](#) suggested that τ in neutral conditions is Gauss distributed based on a wind-tunnel experiment. Colella and Keith (2003) explained that in turbulent shear flows, the non-linear interaction between the eddies gives rise to a departure from Gaussian behavior. Our

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results show that the Gaussian approximation is inadequate in representing the skewed $p(\tau)$, especially for the conditions of strong turbulence intensity (e.g., unstable cases in Fig. 3a). Therefore, $p(\tau)$ here is approximated using a Weibull distribution, i.e.,

$$p(\tau) = \frac{\alpha}{\beta} \left(\frac{\tau}{\beta} \right)^{\alpha-1} \exp\left(-(\tau/\beta)^\alpha\right)$$

(22)

where α and β are the shape and scale parameters, respectively. The values of α and β values offer the numerical experiments Exp (1-20) are listed in Table 2. It can be seen that both α and β depend on wind speed and atmospheric stability. However, β is mainly determined by wind conditions when the wind is strong, while it is affected by ABL stability when the wind is weak. The behavior of α and β are shown in Fig. 4. $|1/L_o|$ is the absolute value of the reciprocal of the Obukhov length L_o which can be calculated by using

$$L_o = - \frac{\overline{\theta} u_*^3}{kg \frac{H_0}{\rho_a c_p}} \quad (23)$$

In both stable and unstable atmospheric conditions, analysis shows that the scale parameter α is related to ABL stability as the power of $|1/L_o|$ where L_o is the Monin-Obukhov length. Fig. 4a shows that α decreases with the $|1/L_o|$, satisfying approximately Eq. (24). For neutral conditions, L_o goes to infinity, Eq. (24) no longer applies. Therefore, the shape parameter obtained by the fitting was directly used for pdf reproduction for the neutral cases instead of the approximated α used for stable and unstable conditions. As Fig. 4b shows, the β parameter increases almost linearly with $u_*^2 + 0.001 \cdot w_*^2$ but can be best approximated using Eq. (25) with $u_{*r} = \sqrt{\tau_r/\rho_a}$.

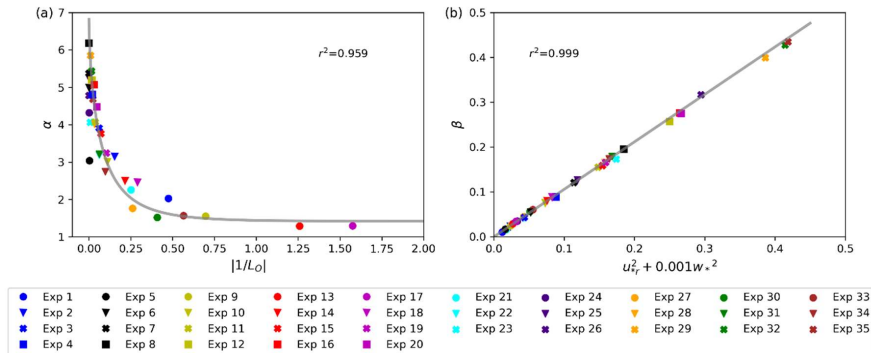


Figure 4. (a) Dependency of the shape parameter α on L_o^{-1} for all numerical experiments Exp (1-35); (b) Dependency of scaling parameter β on $(u_*^2 + 0.001w_*^2)$ for Exp (1-35).

$$\alpha = 5.39 \cdot \exp \left(-5.43 \left(\frac{1}{L_o} \right)^{2/3} \right) + 1.42$$

(1924)

$$\beta = 1.058 \cdot (u_{*r}^2 + 0.001w_*^2)$$

(2025)

Using Eqs. (1822)-(2025), we can approximately describe the turbulent surface shear stress in non-neutral cases.

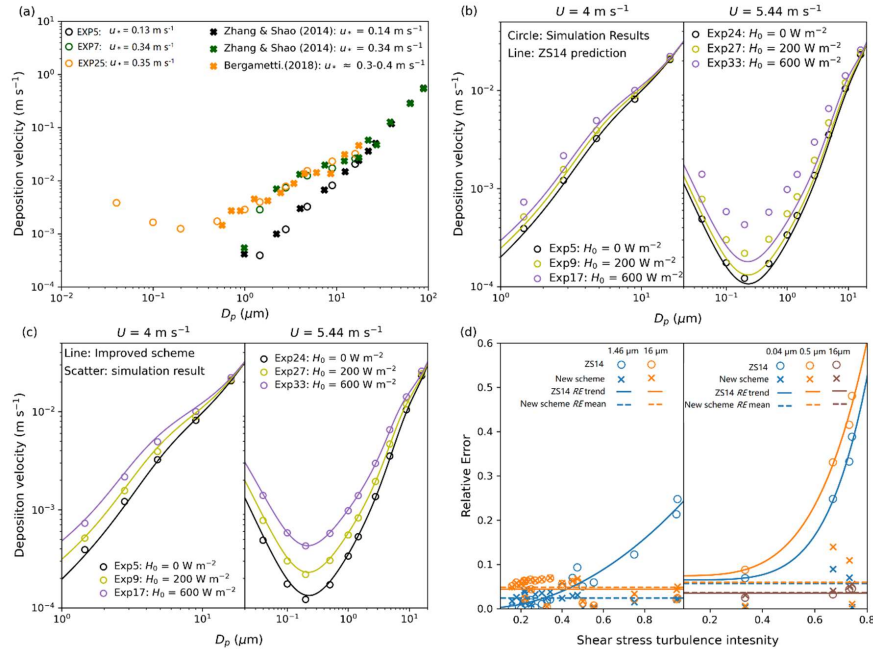
3.2 Improvement to ~~particledust~~ deposition scheme

Figure 5a shows the performances of WRF-LES/D by comparing the simulated deposition velocity, $V_{d,LES}$, with wind tunnel experiments (Zhang and Shao, 2014) and field observation (Bergametti et al., 2018). The observed data are measured under neutral conditions and similar wind flow. As shown, the simulation results agree well with the observed data. On this basis, ~~by we~~ further evaluating the performance of the ZS14 scheme, ~~weand foundshow~~ that the accuracy of the ZS14 scheme decreases ~~with increasingas~~ instability ~~increases~~. ForAs examples, Fig. 5b ~~compareseompared the deposition velocities $V_{d,LES}$ of Exp (5, 9, 17) and Exp (24, 27, 33).~~ $V_{d,LES}$ with ~~those calculated by the ZS14 scheme~~ ~~result V_{d,τ_r} using which is calculated using τ_r from the corresponding experiments, V_{d,τ_r} .~~ It shows that under weak wind conditions, V_{d,τ_r} predicts the deposition well under neutral conditions and underestimates the deposition under convective conditions, especially for particles that are not dominated by molecular diffusion and gravity, and the underestimation increases with the atmospheric instability. To predict the deposition velocity more accurately for convective conditions, we need to account for the effect of shear-stress fluctuations, i.e., the instantaneous shear stress distribution. Thus, the dry deposition scheme can be improved as

$$V_{d,\tau} = \int_0^{\infty} V_d(\tau) p(\tau) d\tau$$

(2126)

with $p(\tau)$ is as given by Eqs. (2218)-(250). As Fig. 5c shows, the improved scheme results $V_{d,\tau}$ and the simulation value $V_{d,LES}$ are shown a remarkable congruence.



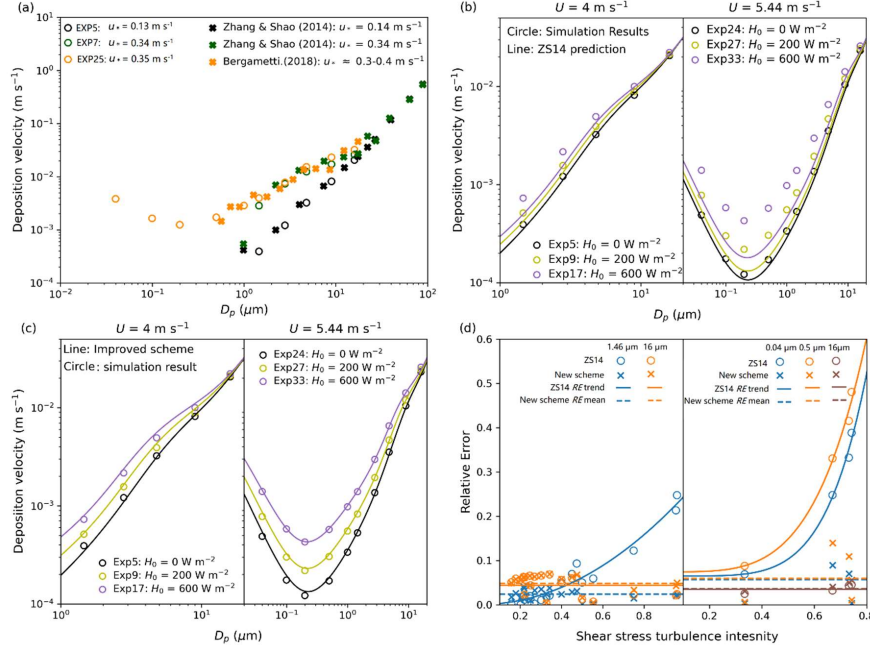


Figure 5. (a) Validation of the simulated deposition velocity from WRF-LES/D (circles) by comparing with the observation data (crosses). (b) the comparison of the predicted result by ZS14 scheme (lines) with the simulated value (circles) of Exp (5, 9, 17) (left) and Exp (24, 27, 33) (right). (c) the comparison of the predicted result by the improved scheme (lines) with the simulated value (circles) of Exp (5, 9, 17) (left) and Exp (24, 27, 33) (right). (d) Comparison of relative error as a function of shear stress turbulence intensity (TI-S), estimated by ZS14 scheme (circles) and the improved scheme (crosses) for Exp (1-20) (Left) and Exp (24, 27, 30, 33) (right).

To make the comparison more clear, the relative errors (RE) of the predicted deposition velocity by ZS14 scheme and improved scheme are compared with the WRF-LES/D simulation value and are calculated as below

$$RE = \left| \frac{V_{d,LES} - V_{d,\tau_r} \text{ (or } V_{d,\tau})}{V_{d,LES}} \right| \times 100\%$$

(2227)

Analysis shows that the value of relative error, RE , depends on surface conditions, wind conditions, atmospheric stabilities, and particle sizes. It increases obviously with increased atmospheric instability under weak wind conditions, while it becomes less sensitive to stability when the wind is strong. Through the analysis, we find that the RE of the ZS14 scheme generally increases with the shear stress turbulence intensity, TI-S, and the value depends on particle size, as shown in Fig. 5d (left). Thus, we compared the RE of some different sized particles to investigate that the particle in which size range is strongly affected

(Fig. A2). The result shows that RE first increases and then decreases with increasing particle size, and the particles with size normally in the range of 0.01 to 540 are strongly affected by turbulent shear stress and $p(\tau)$ needs to be considered. After modification, the errors are limited to less or about 10%. For example, the relative error of Exp (17, i.e., $U = 4 \text{ m s}^{-1}$ and $H_0 = 600 \text{ W m}^{-2}$) for particles of $1.46 \mu\text{m}$ is reduced from $\sim 25\%$ to $\sim 3\%$. The relative error of Exp (33, i.e., $U = 5.44 \text{ m s}^{-1}$ and $H_0 = 600 \text{ W m}^{-2}$) for particles of $0.5 \mu\text{m}$ is reduced from $\sim 50\%$ to $\sim 12\%$.

To further analyze if the RE of ZS14 in unstable conditions is dominated by kinetic instability or dynamic instability, the Richardson number is calculated. Analysis shows that TI-S is positively correlated to gradient Richardson number Ri (Eq. A1). Under unstable conditions associated with strong vertical motion and weak winds, and the RE of ZS14 increases with the increasing magnitude of Richardson number Ri under convection-predominant unstable condition associating weak winds and strong vertical motion (Fig. A3). The relationship between Ri and TI-S needs further study. Consequently, the results illustrate that the modified scheme $V_{d,\tau}$ tends to be more accurate than the unmodified scheme V_{d,τ_r} .

4. Conclusion

The present study was designed to determine the effect of ABL stability on dust particle deposition. For this purpose, the WRF-LES/D was used to model atmospheric boundary-layer turbulence under the presence of atmospheric stability effects to recover statistics of shear stress variability. We then presented an improved particledust-deposition scheme with the consideration of turbulent shear stress. While ABLS can broadly represent levels of atmospheric turbulence, its effect on particledust deposition is wind speed dependent. Through a series of numerical experiments, we have shown the turbulent characteristics of particledust deposition velocity caused by the turbulent wind flow and pointed out the existing dust-deposition shortcomings of the scheme ZS14s scheme have deficiencies in representing particledust deposition under convective conditions. The relative error RE increases as the ABL instability increases for low wind conditions, i.e., RE increases with shear stress turbulence intensity, especially for a certain size range of particles.

Since the dependency of particledust deposition on micrometeorology is imbedded in the application of the surface shear stress, we believe that the dependency of particledust deposition on ABL stability is ultimately attributed to the statistical behavior of shear stress τ . Therefore, in this study, a model including the effects of surface shear fluctuations is proposed and validated by numerical experiments.

Additionally, the fluctuations of surface shear caused by turbulence are available to estimate by can be approximated with a Weibull distribution function. The shape parameter decreases exponentially with the reciprocal of Monin-Obukhov length, and the scale parameter increases linearly with $u_{*r}^2 + 0.001w_*^2$. After statistically revising the original scheme, an improved model is obtained. Using the modified model, the

deposition velocity tends towards numerical experimental results.

The project is the first comprehensive investigation of the turbulent characters of particledust deposition and the findings will be of interest to improve the accuracy of particledust-deposition predictions on regional or global scales. One source of weakness in this study is the variation of τ may be changed by with surface roughness and needs further study, as the roughness length does not fully reflect the effect of the surface topography on the turbulence structure. In spite of this limitation, the study adds to our understanding of the influence caused by ABLs on particle deposition.

Appendix

Figure A1 shows the probability density distribution of surface shear stress for experiments (21-35); Figure A2 shows the changing of relative error with particle size; Figure A3 shows the variation of relative error (RE) of the ZS14 scheme (Eq. (10)) and improved scheme (Eq. (264)) with gradient Richardson number Ri .

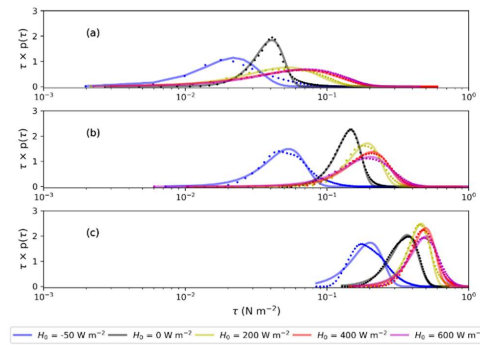


Figure A1. Probability distributions of simulated surface shear stress (dots) and the corresponding fitted Weibull density distribution (solid lines) with different surface heat flux for different wind conditions: **(a)** $U = 5.44 \text{ m s}^{-1}$, **(b)** $U = 10.87 \text{ m s}^{-1}$, **(c)** $U = 18.12 \text{ m s}^{-1}$.

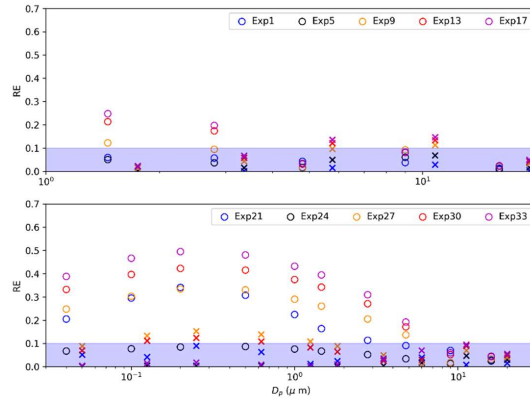


Figure A2. *RE* changes with particle size under weak wind conditions.

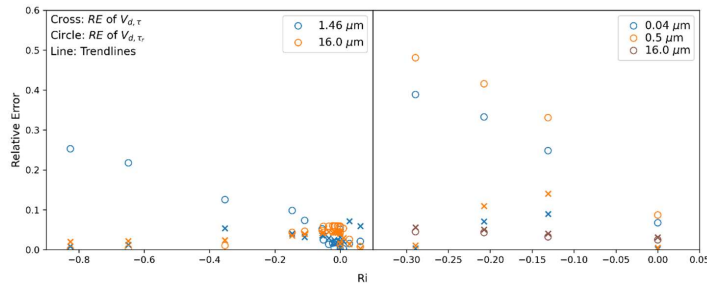


Figure A3. Comparison of relative error as a function of *Ri*, estimated by ZS14 scheme (circles) and the improved scheme (crosses) for Exp (1-20) (left) and Exp (24, 27, 30, 33) (right).

$$R_i = -\frac{g}{\theta} kz \frac{\varphi_\theta}{\varphi_m^2} \frac{H_0}{\rho_a c_p u_*^3} \quad (A1)$$

where z is the center height of the lowest layer, $\bar{\theta}$ is the potential temperature of the lowest layer,

Code and data availability

The source code used in this study is the WRF-chem version 3.7 in the LES mode coupled with a new deposition scheme. WRF-LES model can be downloaded at https://www2.mmm.ucar.edu/wrf/users/download/get_sources.html. The code of the coupled deposition scheme and data set obtained by the simulation are available online at <https://github.com/YinXin2021/WRF-LES-DustDepositionScheme>.

Author contributions

XY, YPS and JZ were responsible for the formal analysis, Methodology. XY and CJ were responsible for the data curation, software, validation and visualization. YPS, JZ and NH were responsible for the supervision, project administration and funding acquisition. XY was responsible for investigation and Writing - original draft preparation. XY, YPS, and JZ and CJ were responsible for the Writing – review & editing.

Competing interests

The authors declare that they have no conflict of interest.

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