

#### 1 Wet deposition in the remote western and central

#### 2 Mediterranean as a source of trace metals to surface seawater

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- 30
- 31 Abstract. This study reports the only recent characterisation of two contrasted wet deposition events
- 32 collected during the PEACETIME cruise in the open\_Mediterranean Seaopen seawater, and their

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33 impact on trace metals (TMs) marine stocks. Rain samples were analysed for Al, 12 trace metals (TMs 34 hereafter, including( Co, Cd, Cr, Cu, Fe, Mn, Mo, Ni, Pb, Ti, V and Zn) and nutrients (N, P, DOC) 35 concentrations. The first rain sample collected in the Ionian Sea (rain-Rain ION) was a wet-typical 36 regional background wet deposition event whereas the second rain sample, collected in the Algerian 37 Basin (Rrain FAST), was a Saharan dust wet deposition event. The concentrations of TMs in the two 38 rain samples were significantly lower compared to concentrations in rains collected at coastal sites 39 reported in the literature, suggesting either less anthropogenic influence in the remote Mediterranean 40 environmentSea, or decreased anthropogenic emissions during the last-preceding decades-in the 41 Mediterranean Sea. The TMs inventories in the surface microlayer and mixed layer (0-20\_m) at ION 42 and FAST stations before and after the events were; compared to atmospheric fluxes which, showed 43 that the atmospheric inputs were a significant source of particulate TMs for both layers. At the scale 44 of the western and central Mediterranean, the atmospheric inputs were of the same order of magnitude 45 as marine stocks within the ML for dissolved Fe, Co and Zn, underlining highlighting the role of the

- 46 atmosphere in their biogeochemical cycles in the stratified Mediterranean Sea (state season) under
- 47 <u>background conditions (?)</u>. In case of Our results suggest that -intense dust-rich wet dust-deposition
- 48 events; are an important source the contribution of atmospheric inputs could be critical for of dissolved
- 49 stocks of the majority of TMs to the study region.

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#### 51 1. Introduction

52 Atmospheric deposition of continental aerosol has long been recognized to influence trace element 53 (TE) concentrations in remote oceanic surface waters (Buat-Ménard and Chesselet, 1979; Hardy, 54 1982; Buat-Ménard, 1983). In particular, the Mediterranean Sea (Med Sea) is an oligotrophic 55 environment where marine biosphere growth is nutrient-limited during the long Mediterranean 56 summer season, which is characterized by a strong thermal stratification of surface waters (The 57 Mermex Group, 2011). The Mediterranean atmosphere is characterized by the permanent presence of 58 anthropogenic aerosols from industrial and domestic activities around the basin (e.g., Sciare et al., 59 2003; Kanakidou et al., 2011). In addition to this anthropogenic background, the Mediterranean basin 60 is also subject to seasonal contributions of particles from biomass fires in summer (Guieu et al., 2005) 61 and to intense sporadic Saharan dust inputs (e.g., Loÿe-Pilot and Martin, 1996; Vincent et al., 2016). 62 Several studies have emphasized that the atmospheric deposition of aerosols, notably through wet 63 deposition, plays a significant role in the marine cycles of both-nutrients, such as nitrogen (N) and 64 phosphorus (P) (e.g. Pulido-Villena et al., 2010; Richon et al., 2018 a and b; Violaki et al., 2018) and 65 micronutrients, such as iron (Fe) (Bonnet and Guieu, 2006). Recently, atmospheric dust inputs were 66 identified to have a fertilizing effect on the planktonic stocks and fluxes, even in the presence of 67 relatively high dissolved (/)nutrients N, P and Fe marine concentrations (Ridame et al., 2011; reviewed 68 in Guieu and Ridame, 2021). Mackey et al. (2012) showed that TMs provided by dust deposition could 69 explain this fertilizing effect. Indeed, some-many\_TMs, including Mn, Co, Ni (Mackey et al., 2012), 70 Cu (Annett et al., 2008) and Zn (Morel et al., 1991), play physiological roles for phytoplanktonic 71 organisms. These TMs are present in very low concentrations in oligotrophic systems such as the Med 72 Sea, possibly limiting (or co-limiting) the phytoplankton growth (Pinedo-Gonzàles et al. 2015), and 73 implying the rolepointing to the importantce of dust deposition as source of TMs for planktonic 74 communities. On the other hand, atmospheric deposition of European aerosol particless was identified 75 to have a negative effect on chlorophyll concentrations (Gallisai et al. 2014), by providing trace 76 tals, as Cu, at toxic levels (Jordi et al. 2012).

77 The atmospheric deposition of TMs in the Mediterranean is related to both dust and anthropogenic 78 aerosole deposition (Desboeufs et al., 2018). The role of dust deposition as a source of TMs was 79 observed from the correlation between the atmospheric deposition of mineral dust, and the enrichment 80 of dissolved TMs (Cd, Co, Cu, Fe) in the Mediterranean Sea surface microlayer (Tovar Sánchez et al., 81 2014). For the water column, the adding of dissolved Fe and Mn was emphasized in mesocosm 82 experiments after dust addition mimicking intense wet dust deposition (Wuttig et al., 2013). Yet, the 83 direct impact of wet deposition events on TMs concentrations in surface seawater has never not been 84 studied and reported in situ, where TM concentrations were determined in both rainwater and seawater 85 samples collected from the same location before the present study, whether in the Mediterranean Sea Commented [RS(-S5]: Define/give examples

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Commented [RS(-S7]: Which nutrients?



86 or in other oceanic regions. Moreover, the two key criteria used to assess the potential impact of TMs

87 and nutrients wet deposition are and their respective concentrations (or fluxes) and solubility, i.e., the

88 partitioning between dissolved and total concentrations in rainwaters. Indeed, it is considered that the

89 dissolved fraction of nutrients and TMs can be directly assimilated by the phytoplankton (ref). <u>A f</u>Few

90 studies <u>have focussed onreported</u> concentrations of

91 TMs in rainwater samples was-collected around the Mediterranean basin: Al-Momani et al., (1998),

92 Kanellopoulou, (2001) and Özsoy and Örnektekin, (2009) in the eastern basin, and Losno (1989),

Guieu et al. (1997), Frau et al., (1996), Chester et al., (1997) and Guerzoni et al. (1999b) in the western

basin. These studies led to highly variable TMs concentrations and solubility, illustrating a more

generalthe large variability of TMs inputs during wet deposition <u>events</u> in the Mediterranean Sea
 (reviewed in Desboeufs, 2021). All these studies were performed at coastal sites. Offshore samples of

96 (reviewed in Desboeufs, 2021). All these studies were performed at coastal sites. Offshore samples of
 97 rainwater have rarely been <u>studied\_reported\_in</u> the literature-<u>so\_far</u>. In the Mediterranean, to our

98 knowledge, trace element concentrations from only three rain samples collected at sea in April 1981

99 have been reported in a PhD thesis (Dulac, 1986). However, due to the continental and local sources

100 of pollution and the variety of anthropogenic aerosol sources (Amato et al., 2016), the TM<sub>9</sub> rain

101 composition of the coastal zone could may not be representative of atmospheric deposition in to the

102 remote Mediterranean.

103 The PEACETIME cruise (ProcEss studies at the Air-sEa Interface after dust deposition in the 104 MEditerranean Sea) performed in spring 2017 aimed at-to studying the impacts of atmospheric 105 deposition, in particular Saharan dust events, on the physical, chemical and biological processes in 106 this marine oligotrophic environment (Guieu et al., 2020). We investigated here the concentration and 107 solubility of TMs and nutrients of from two rain events waters sampled in the open Mediterranean 108 open-See during the cruise. We compare our results on-of TMs concentrations and solubility in 109 rainwater with previous studies based on rainwater samples collected at coastal sites to investigate 110 potential differences with the open sea. Additionally, to assess the impact of wet deposition on the 111 surface TMs concentrations, surface seawater, including the surface micro-layer, and rain water was 112 concurrently collected. To the best of our knowledge, this is the first time TM data for these marine

113 <u>compartments have all been presented at the same time.</u> for the first time with rain samples.

114 **2.** Sampling and methods

#### 115 2.1 Sampling and chemical analysis of rainwater

116 The PEACETIME oceanographic campaign (<u>https://doi.org/10.17600/17000300</u>) took place in the

117 western and central Mediterranean Sea on-board the French research vessel (R/V) Pourquoi Pas ?

118 between 11 May and 10 June 2017, i.e. at the beginning of the Mediterranean stratification season

119 (Guieu et al., 2020). The rain collector was installed on the upper deck (22 m above sea level) where

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120 no on-board activities were taking place to avoid contamination. A-The rain collector was equipped 121 with an on-line filtration system to directly-separate the dissolved and particulate fractions at the time 122 of collection (details of the filtration system are available in Heimburger et al., 2012) allowing for the 123 calculation of solubility of TMs in the rainwater at the time of collection. The filtration device was 124 equipped with a Nuclepore® polycarbonate membrane (PC)-filter (porosity: 0.2 µm, diameter: 47 125 mm).T, and the diameter of the funnel of the collector was 24 cm. The rain collectors were was 126 installed opened only when rain was expected and kept elosed withincovered by an acid-washed, 127 sealed plastic film until the rainfall began when not in use. All the sampling materials were thoroughly 128 acid-washed in the laboratory prior to the cruise departure (washing protocol described in Heimburger 129 et al., 2012). No stabilizing setup-device was used to keep the funnel level during the pitch and roll of 130 the ship, preventing a precise assessment of the height of rainfall from the collected water volume. 131 During the rain sampling, the ship was always facing the wind to avoid contamination by the smoke 132 ship's exhaust of the ship itself, as the chimney was situated on the lower deck and behind the 133 collector.

134 Immediately after sampling, the collector was disassembled under the a laminar flow hood inside an 135 on-board clean-room on board-laboratory. The dissolved fraction was separated into 4-four aliquots dedicated to i) dissolved organic carbon (DOC) determination by high-temperature catalytic 136 137 oxidation (HTCO) on a Shimadzu total organic carbon analyzer (as described in Van Wambeke et al 138 2021), ii) major ions by ion chromatography (IC), iii) metals analysis TM determination by inductively 139 coupled plasma <u>coupled</u> methods (ICP), and iv) pH measurement. For ICP measurements, the sample was acidified immediately to 1% by volume of ultra-pure nitric acid: 67-69%, Ultrapur, Normatom®, 140 141 VWR. For DOC analysis... For IC analysis, the filtrate sample (dissolved fraction) was immediately 142 frozen. The filter (particulate fraction) was dried under the laminar flow hood, and then put in a storage 143 box and packed with a plastic bag to avoid contamination. After returning to the laboratory, filters 144 were acid digested by using the adapted protocol adapted from Heimburger et al.

145 (2012) as follows: filters were placed in tightly capped Savillex<sup>TM</sup> PFA digestion vessels with 4 mL

146 of a mixture of HNO3 (67-69%, Ultrapur, Normatom®, VWR), H2O and HF acids (40%, Ultrapur,

147 Normatom<sup>®</sup>, VWR) in a proportion <u>of (3: 1: 0.5)</u>, then heated in an oven at 130°C for 14 hours. After

cooling, <u>the acid solution</u> was completely evaporated on a heater plate (ANALAB, 250, A4) at 140°C
for about

150 2h, then 0.5 mL of  $H_2O_2$  (30-32%, Romil-UpA<sup>TM</sup>) and 1 mL of the acidified water (2% HNO<sub>3</sub>) was 151 added to the vessels and heated <u>during for 30</u> min. to dissolve the dry residue in the bottom of the 152 vessels; finally, 12 mL of acidified water (1% HNO<sub>3</sub>) was added to obtain 13.5 mL of solution in a 153 tube for ICP-MS analyses.

154 The dissolved fraction was analysed by IC (IC 850 Metrohm) for the inorganic and organic anions

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**Commented [RS(-S14]:** This should be number 1 as it is the focus of the paper. List = i) TMs, ii) DOC, iii) major seawater and atmospheric ions, iv) pH. Then a description of the methods used for each, in the order they are listed



155 (NO2<sup>-</sup>, NO3<sup>-</sup>, PO4<sup>3-</sup>, SO4<sup>2-</sup>, F<sup>-</sup>, Cl<sup>-</sup>, Br<sup>-</sup>, HCOO<sup>-</sup>, CH<sub>3</sub>COO<sup>-</sup>, C<sub>2</sub>H<sub>5</sub>COO<sup>-</sup>, MSA, C<sub>2</sub>O4<sup>2-</sup>) and for the cation NH4<sup>+</sup> (Mallet et al., 2017). On the other hand, the dissolved fraction and solution from the digestion 156 157 of the particulate fraction of the rain samples s were analysed by ICP-AES (Inductively Coupled 158 Plasma Atomic Emission Spectrometry, Spectro ARCOS Ametek®) for major elements (Al, Ca, K, 159 Mg, Na, S) (Desboeufs et al., 2014) and by HR-ICP-MS (High Resolution Inductively Coupled Plasma 160 Mass 161 Spectrometry, Neptune Plus TM at Thermo Scientific TM) for TMs: Cd, Co, Cr, Cu, Fe, Mn, Mo, Ni, P, 162 Pb, Ti, V and Zn. The speciation of dissolved P was estimated by determining the dissolved inorganic 163 phosphorus (DIP) from phosphate concentrations expressed as P and the dissolved organic phosphorus 164 (DOP) from the difference between total dissolved phosphorus (TDP), obtained by ICP-MS, and DIP, 165 obtained by IC. 166 In order to estimate the contamination of sampling and analytical protocols, 3-three blanks of rain 167 samples (collected on-board during the cruise with the same protocol without rain events) were used 168 and processed. The procedural limit of detections (LoD) were defined as 3 x standard deviation of 169 blank samples both for dissolved and particulate fractions estimated after acid digestion. All samples 170 dissolved and particulate sample concentrations were higher than LoD, except for NO2 in the two rain 171 samples. The blank concentrations represented 10.2% oin average for TMs and were typically lower 172 than 20% of the sample concentrations, except for Cd (52%) and Mo (43%) in the dissolved fraction. 173 For concentration computation subtracted these 174 eoncentrations obtained in rain samples. Blank concentrations were subtracted from all sample 175 concentrations.

#### 176 2.2 Atmospheric ancillary measurements

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The PEGASUS (Portable Gas and Aerosol Sampling UnitS, www.pegasus.cnrs.fr) mobile platform 178 179 of LISA is a self-contained facility based on two standard 20-feet containers, adapted with air conditioning, rectified power, air intake and exhausts for sampling and online measurements of 180 181 atmospheric aerosols and gaseous compounds, and their analysis (Formenti et al., 2019). During the 182 PEACETIME cruise, only the sampling module of the facility was deployed on the starboard side of 183 deck 7 of the R/V. The PEGASUS instrumental payload of relevance to this paper included 184 measurements of the major gases such as NOx, SO<sub>2</sub>, O<sub>3</sub> and CO by online analysers (Horiba APNA, 185 APSA, APOA and PICARRO respectively; 2-min resolution, detection limit for all analytes was 0.5 186 ppb and 1 ppb for CO). These gases, that were used to estimate the origins of collected air masses the 187 sampled air.

From the first of lst June 2017 (?) (not operational before), additional measurements by an ALS450<sup>®</sup>
Rayleigh-

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**Commented [RS(-S17]:** More detail needed. Not clear what the rain blanks were. What were the blanks for the dissolved fraction? Empty bottles swished out with MQ or acidified MQ and a filter? Blanks and LoDs can be found where?

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190 Mie lidar (Leosphere TM; Royer et al., 2011) was-were used to monitor the vertical distribution of 191 aerosols over time and the associated integrated columns. The vertical lidar profiles were analysed to 192 yield the apparent backscatter coefficient (ABC) corrected from the molecular transmission, as well 193 as the volume depolarisation ratio (VDR). The inversion procedure (Chazette et al., 2016; 2019) to 194 retrieve the aerosol extinction coefficient (unit km<sup>-1</sup>) uses a vertical-dependent lidar ratio that takes 195 into account two aerosol layers. The first layer corresponds to marine aerosols in the marine boundary 196 layer (MBL), the second to a desert aerosol layer that can extend between ~1 and 6 km above mean sea level (amsl). In accordance with Chazette et al. (2016), for the same wavelength and region, the 197 198 lidar ratios were set to 25 and 55 sr, respectively. The vertical profile of the aerosol extinction 199 coefficient was retrieved from 0.2 km amsl on-upwards with a vertical resolution of 15 m. Based on 200 these profiles, the integrated column content of dust aerosols was estimated using a specific extinction 201 cross-section of 1.1 m<sup>2</sup> g<sup>-1</sup> as proposed by Raut and Chazette (2009).

In addition, detailed meteorological data such as air and sea temperature, atmospheric pressure, relative humidity, atmospheric pressure, heat flux and wind speed and direction were provided on a 30 seconds time step basis by the ship's permanent instrumentation.

#### 205 2.3 Sampling and analysis of dissolved TMs in seawater

206 Before and after each rain, seawater samples were collected in the surface microlayer (SML: <1 mm) 207 and subsurface seawater (SSW: <1 -m depth) (Tovar-Sánchez et al., 2020, this special-issue). SML 208 and SSW samples were collected from a pneumatic boat 0.5-1 nautical miles away from the R/V in 209 order to avoid any potential contamination. SML samples were collected using a silicate glass plate 210 sampler (Stortini et al., 2012; Tovar-Sánchez et al., 2019) which had previously been acid-cleaned 211 with acid-overnight and rinsed thoroughly with ultrapure water ( $\frac{MQ-water 18 \text{ m}\Omega \text{ cm}-1}{100 \text{ m}^2}$ ). The 39 x 25 212 cm silicate glass plate had an effective sampling surface area of 1950 cm<sup>2</sup> considering both sides. In 213 order to check for procedural contamination, SML blanks were collected at some stations on board of 214 the pneumatic boat by rinsing the glass plate with ultra-pure water and collecting 0.5 L of ultra-pure 215 water using the glass plate system. The surface microlayer thickness was calculated following the 216 formula of Wurl (2009). SSW was collected by using an-acid-washed Teflon tubing connected to a 217 peristaltic pump. The total fraction (i.e. T-SML) was directly collected from the glass plate system 218 without filtration into a 0.5 L acid cleaned LDPE bottles, while the dissolved fraction in the SML (i.e. 219 D-SML) and SSW (i.e. D-SSW) was filtered in situ through an acid-cleaned polypropylene cartridge 220 filter (0.22 µm; MSI, Calyx®).

221 TMs samples were also collected in the water column using the <u>a titanium</u> trace metals clean (TMC)

222 rosette (mounted with 24 teflon-coated Go-Flo bottles) before and after the rain eventse (Bressac et

223 al., 2021). Although rosette deployments were performed over the whole water column, we focus here

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224	on the 0-20 m marine-mixed-layer (ML). The water column was sampled using the TMC titanium
225	rosette mounted with 24 teflon coated Go Flo bottles.
226	Immediately after recovery, the Go-Flo bottles were transferred inside a class-100 clean laboratory
227	container. Seawater samples were directly filtered from the bottles through acid-cleaned $0.2\_\mu m$
228	capsule filters (Sartorius Sartobran-P-capsule $0.45/0.2$ - $\mu$ m). All samples were acidified on board to

229 pH <\_2 with Ultrapure-grade HCl under a class-100 HEPA laminar flow hood. Metals (namely Cd, 230 Co, Cu, Ni, Mo, V, Zn and Pb) were pre-concentrated using an organic extraction method (Bruland et 231 al., 1979) and quantified by ICP-MS (Perkin Elmer ELAN DRC-e) in the home laboratory. In order 232 to breakdown metal\_organic complexes and remove organic matter (Achterberg et al., 2001; Milne et 233 al., 2010), the total fraction samples (i.e. T-SML) were digested-UV-treated prior to the pre-234 concentration step using a UV system consisting of one UV (80 W) mercury lamp that irradiated the 235 samples (contained in quartz bottles) during for 30 min. The accuracy of the pre-concentration method 236 and analysis for TMs was established using Seawater Reference Material (CASS 6, NRC-CNRC) with 237 recoveries ranging from 89% for Mo to 108% for Pb. Due to the complexity of the analytical method, 238 all the TMC samplings were not analysed for these metals. Overall, 1 or 2 depths were obtained in the 239 mixed layer (0-20 m). Dissolved Fe and Al concentrations were also measured on board. Dissolved 240 Fe concentrations were measured using an automated Flow Injection Analysis (FIA) with online 241 preconcentration and chemiluminescence detection (Bonnet and Guieu, 2006), and dissolved Al 242 concentrations using the fluorometric method described by Hydes and Liss (1976). Sampling and 243 analysis for dissolved Fe and Al concentrations are fully described in Bressac et al. (2021), and covered at least 4-four depths in the 0-20 m mixed layer. 244

#### 245 2.4 Enrichment factor and solubility

246 In order to better constrain the origin of TMs in the rain samples, their enrichment factor s(EF; Rahn 247 1976) relative to the Earth's crust was-were calculated based on their total concentrations (dissolved 248 + particulate fractions) as:

$$EF = \frac{([X]/[Al])sample}{([X]/[Al])crust}$$
(1)

250 where [X]/[Al] is the ratio between an element X and Al concentrations in rainwater samples (at the 251 numerator), and in the Earth's crust (denominator) from Rudnick and Gao (2003). Aluminium is 252 currently used as a reference element as it only has a crustal origin. For a given TM, EF >1 indicates 253 an enrichment with respect to the average composition of the Earth's crust. To account for the soil 254 composition variability of mineral dust atmospheric-sources, TMs with an EF value >10 are 255 considered significantly enriched, which points to a derived from-non-crustal sources (Rahn, 1976).

256 The relative solubility of TMs in the two rainwater events was calculated as: Commented [RS(-S19]: Which metals? Mo and Pb or not all samples were preconcentrated for all metals listed?

Commented [RS(-S20]: Not exactly but the vast majority of Al is crustal. Therefore, it is currently accepted as one of the best proxy elements. It is predominantly crustal



	_[X]disssolved											
257	$S_X\%$	[X]total	x 100	(2)								
250	where $\Gamma \theta$ is the relative colubility (in $\theta$	() of an al	mant V i	n tha minur								

258 where  $S_X\%$  is the relative solubility (in %) of an element X in the rainwater,  $[X]_{dissolved}$  and  $[X]_{total}$  are

259 its soluble and total concentration, respectively.

#### 260 2.5 Atmospheric deposition fluxes

261 Impacts on biogeochemical cycles and ecosystem functioning after a rain event occur on time scales 262 of a few days (2-3), and space scales of tens of km (about 20-50 km within the radius of the ship's 263 position). In the specific context of oceanographic cruising, the documentation of these impacts is 264 restricted to the vertical dimension at the prescribed temporal scale. In this vertical dimension, the 265 exchange of TMs across into the ML was controlled both by atmospheric inputs over the R/V position 266 and by advection from surrounding water masses that may have been impacted by surrounding 267 rainfall. Therefore, we had to consider this process in our estimation of the atmospheric fluxes 268 contributions. For this purpose, the atmospheric fluxes have to be integrated to the extent of the rain 269 area that can impact the marine surface layers. We derived wet deposition fluxes by considering the 270 total precipitation accumulated during the duration of the rain over the area around the R/V location. 271 Thus, the wet deposition fluxes in our rain samples were calculated by multiplying the volume 272 weighted mean rainfall concentration by the total precipitation (equation and explanation for using 273 <u>VWM rainfall here</u>). The total precipitation of the rain events was issued from the hourly total 274 precipitation accumulated during the rain events over the region from ERA5 ECMWF reanalysis 275 (Herbasch et al., 2018) and from the rain rate composite radar products from the European OPERA 276 database (Saltikoff et al., 2019), when it wais possible. Although subject to uncertainties (Morin et al., 277 2003), a surface-based weather radar is probably the best tool to estimate rainfall in the surroundings 278 of the R/V because ...... However, the OPERA database does not include Italian radars, which anyway 279 did not cover the central area of the Ionian Sea during the cruise anyway. ERA5 data are available on 280 regular latitude-longitude grids at 0.25° x 0.25° resolution. The accumulated precipitation was taken 281 from the grid-points spanning the ship's location, more or lessapproximately 0.25° around the central 282 grid-point for integrating the regional variability. Surface rain rate radar composite images were 283 available every 15 minutes with a spatial resolution of 2 km x 2 km. The accumulated precipitation 284 was the sum of integrarated (?)+ rain rates during the rain duration averaged over the radar pixels 285 spanning the ship's location within a radius of about 25 km around the ship location.

#### 286 2.6 Stocks in the surface seawater

For the surface microlayer (SML), stocks of TMs were estimated from the integration of TMs
 concentration over the thickness of the layerSML. The thickness ranged from 32 to 43 μm and from

289 26 to 43 μm at ION and FAST, respectively (Tovar-Sanchez et al., 2020).

**Commented [RS(-S21]:** Sometimes ship, sometimes R/V, perhaps best to stick with one



290 The trace metals stocks within the ML were calculated by trapezoidal integrations of marine 291 concentrations from SSW and TMC rosette samplings. The upper water column was stratified along 292 the cruise transect (Taillandier et al., 2020), with a -ML depth ranging from 7 to 21 m (11 to 21 m at 293 ION station and 11 to 19 m at FAST station (Van Wambeke et al., 2020). The ML depth (MLD) 294 fluctuations, for example-due to wind peaks associated with rain events for example, could create 295 rapidly changing conditions of vertical advection from deeper waters. However, with no significant 296 increase in TMs concentrations being observed below the ML down to about 50 m (not shown), the 297 enrichment observed in the ML after rainfalls could not be attributed to any mixing with deeper water 298 due to high wind. In consequence, stocks in the ML have been integrated over a constant depth range 299 of 0-20 m for comparison, as-in accord with Bressac et al. (2021). For Cu, Fe, Ni and, Zn, stocks were 300 estimated both for the dissolved and particulate fractions in the SML and ML, for Co, Cd, Mo, Pb and 301 V for the dissolved fraction only in the ML and for both fractions in the SML and for Mn and Ti only 302 for the particulate fraction in the ML.

The partitioning coefficient between the particulate and dissolved phases (Kd
 =[particulate]/[dissolved]) was used to investigate exchanges between the dissolved and particulate
 pools of TMs.(ref).

#### 306 3. Results

#### 307 3.1 General conditions

308 The general meteorological conditions during the cruise indicated that the ION and FAST stations

309 were highly affected by cloudy weather conditions. During these periods, 2-two significant rain

eventss occurred overon the R/V's position and have been collected were sampled: The first rain
 sample (hereafter Rain ION) was collected during the 4-day ION station occupation in the Ionian Sea
 in the early morning of 29/05 at 03:08

313 UTC<sub>2</sub>, and tThe second rain event (hereafter, Rain FAST) occurred during the 5-day "Fast action"

314 station, (hereafter 'FAST') in the Algerian Basin during the night of 05/06 at 00:36 UTC (Table 1).

315 The two rain sample <u>collectionss</u> coincided with peaks in relative humidity and wind speed, and

316 minima in air temperature (not shown).

#### 317 Table 1: Information regarding the two rains collected during the PEACETIME cruise.

Sample	Sampling time	Station name (dates) and rain location	Estimated total precipitation			
Rain ION	29 May 2017, 03:08-04:00 (UTC) 05:08–06:00 (local time)	ION (25–29 May) 35.36°N, 19.92°E	3.5 ±1.2 mm			

**Commented [RS(-S22]:** ML, ML depth or MLD - consistency

**Commented [RS(-S23]:** Why was this? Because of concs below the LoD?

**Commented [RS(-S24]:** 09/05/2017. Date here is numeric, earlier it is longhand 1<sup>st</sup> September

**Commented [RS(-S25]:** Unnecessary – you've been referring to them already



Rain FAST	05 June 2017, 00:36-01:04 (UTC) 02:36–03:04 (local time)	FAST (2–7 June) 37.94°N, 2.91°E	6.0 ±1.5 mm
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### 318

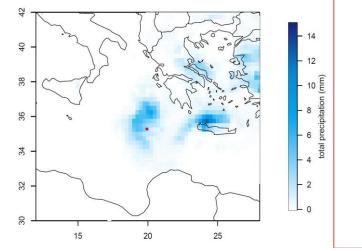
#### 319 3.1.1. Rain ION

320 The ERA 5 data reanalysis shows 2-two periods of precipitation in the surrounding of yicinity of the 321 ship's position, i.e. in the morning and evening of June 26 (not shown) and in the night between June 322 28 and 29, in agreement with on-board visual observations. The rain event collected at ION was the 323 product of a large cloud system, covering an area of about 90 000 km<sup>2</sup> around the R/V position, 324 spreading over the Ionian and Aegean Seas (Fig. 1). As nNo radar measurements being were available 325 for this area, the accumulated rate  $(3.5 \pm 1.2 \text{ mm})$  was estimated from ERA 5 data reanalysis on the 326 grid-point corresponding to the ION station and was  $3.5 \pm 1.2 \text{ mm}(\pm 0.25^{\circ})$ -around the R/V position 327  $(\pm 0.25^{\circ})$ . The wash-out of the atmospheric particles was revealed by the decrease in aerosol number 328 concentrations monitored onboard from about 1900 to 300 part.cm<sup>-3</sup> (supplementary material Fig. S1). 329 Air mass back-trajectories showed that the scavenged air masses came from Greece both in the marine 330 boundary layer and in the free troposphere (Fig. S1). The satellite observations also showed low 331 aerosol optical thickness during this period (not shown), meaning low amounts of aerosols in the 332 atmospheric column. No significant European pollution influence was monitored by on-board 333 measurements during this event, with major gas mixing ratios and aerosol concentrations in-close to 334 the average values of the cruise (Fig. S1) and typical of clean atmospheric concentrations, i.e. under 335 detection limit for NOx, 1.2 ppb for SO<sub>2</sub>, 51 ppb for O<sub>3</sub>, 80 ppb for CO and 3000 part.cm<sup>-3</sup>. On this 336 basis, this wet event was representative of a Mediterranean background marine rain event (ref for 337 these typical background concs).

Commented [RS(-S26]: Rate requires a time dimension

**Commented** [RS(-S27]: Onboard is one word – several cases need correcting





**Commented [RS(-S28]:** It would be nice to have S3 in this figure for comparison instead of in the SI

#### 338

# 339Figure 1: Total precipitation (mm) between 28 May 201728 at 20:00 UTC and 29 May 2017 at 10:00340UTC from ERA5 ECMWF reanalysis. The red circle indicates the R/V position.

#### 341 3.1.2. Rain FAST

342 As detailed in Guieu et al. (2020), the FAST station position was decided on the basis of regional 343 model forecast runs and satellite observations, in for the purpose to of catching a wet dust deposition 344 event. Significant dust emissions were observed from NASCube (http://nascube.univ-lille1.fr/, 345 Gonzales and Briottet, 2017) over North Africa from the night of 30-31 May, then new dust emissions 346 in the night from 3 to 4 June in Algeria and southern Morocco associated with a northward atmospheric 347 On 30 May, the SEVIRI AOD satellite product (https://www.icare.univflux. 348 lille.fr/dataaccess/browse-images/geostationary-satellites/, Thieuleux et al., 2005) confirmed the 349 presence of atmospheric dust in a cloudy air mass over the western part of the Mediterranean, and 350 from 2 June the export of a dust plume from North Africa south of the Balearic Islands with high 351 AOD (>0.8) on the Alboran Sea was observed (Fig. S2). The dust plume was transported to the NE 352 up to Sardinia on June 4, with AOD\_<0.5-in all the area and c. Clear sky with low AOD was left 353 observed west of 4°E on June

354 5.

355On-board lidar measurements (Fig. 2 a,b,c) showed that the aerosol plume was present over the ship's<br/>position from 2 June at 21:00 UTC until the rain event, and corresponded to a dust aerosol layer well<br/>highlighted by the high depolarization. The dust plume was concentrated between 3 and 4 km at the<br/>beginning of the station occupation, then expanded down to the marine boundary layer (about 500 m

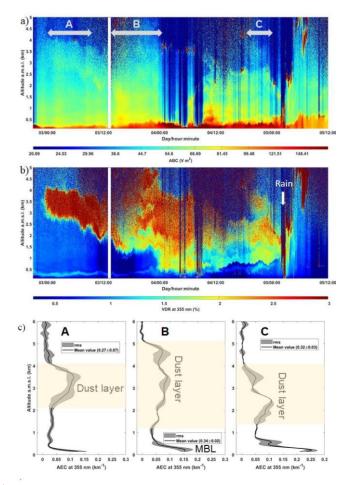


359 amsl) byto the end of the day on 3 June 20173 down to the marine boundary layer (about 500 m amsl).

360 The mass integrated contents-concentration of dust aerosols derived from the profiles of aerosol

extinction ranged from a minimum of 0.18  $\pm$ 0.005 g m<sup>-2</sup> just before the rain to a maximum of 0.24 361

362  $\pm 0.009$  g m<sup>-2</sup>, where standard deviations indicate the temporal variability (1 sigma).



#### 363 364

Figure 2: On-board lidar-derived a) Apparent backscatter coefficient (ABC), (b) Temporal evolution 365 (in Local-local Timetime) of the lidar-derived volume depolarization ratio (VDR) where the dust 366 plume is highlighted for values higher than ~1.7 (yellow to red colours) and the rain by values higher 367 than 3 (indicated by the arrow), and c) vertical profiles of the aerosol extinction coefficient (AEC) in 368 cloud free condition, integrated over 3 periods along the dust plume event, noted A, B and C in

**Commented [RS(-S29]:** Can these panels be relabelled? You have panels a-c and labelling within panel a of arrows a-c, and in panel c of box a-c. All these a-cs get a bit confusing



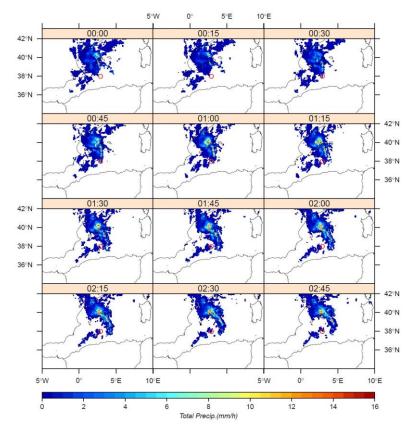
369 panelfigure a (top). The grey shade represents the root mean square (rms) variability along the time
370 of the measurement. The dust layer is highlighted on the profiles. The mean aerosol optical thickness
371 is given in the boxed legend with its temporal variability (1 sigma). The location of the marine
372 boundary layer (MBL) is also pointed.

373 Rainfalls were-was observed by weather radar images in the neighbouring-area of-the R/V 374 neighbouring from 3 June at 7:00 UTC. The rainfalls recorded around the FAST station were was 375 associated with 2-two periods of rain: the 03/06 from 07:00 to 14:00 UTC on 03/06, and from the 376 04/06 at 16:00 UTC on 04/06 to 05/06 at 06:00 UTC on 05/06. For this latter case, a rain front (100 377 000 km<sup>2</sup>), moving eastward from Spain and North Africa regions, reached the FAST station the night 378 between the 4 and 5 June (Fig. 3). Wet deposition between the 4 and early 5 June in the FAST station 379 area were confirmed by radar imagery, showing several other instances of rain spots-around the R/V 380 position before and after the rain sampling (Fig. 3). Continuous on-board lidar measurements 381 confirmed the below-cloud deposition during the rain event of early 5 June (Fig. 2b). Rain FAST was 382 a wet deposition event occurring at the end of an episode of transport of Saharan dust, whereas 383 precipitation of on the 3 June occurred during the maximum of the dust plume (Fig. 2b and S2). The 384 surface concentrations of gas and particles, measured on-board, suggest no clear dust or anthropogenic 385 influence in the atmospheric boundary layer during the-this period of wet deposition, in agreement 386 with back trajectories of low altitude air masses (Fig S2.), presuming no local mixing between dust 387 and anthropogenic particles into rain samples. The total precipitation estimated from radar rainfall 388 estimates yield an accumulated precipitation of  $6.0 \pm 1.5$  mm ( $\pm 25$  km around the ship's position), in 389 agreement with ECMWF reanalysis ERA5 (Fig. S2) for the wet deposition on the night of 4-5 June

390 (5.7  $\pm$ 1.4 mm in the grid-point spanning the R/V position, i.e.  $\pm$ 0.25°-around).

Commented [RS(-S30]: repetition





391

Figure 3: Rain rates (mm/h) during the night between the 4 and 5 June, when Rain Fast-FAST whas
 been-collected on-board, issued from European rain radar composites (OPERA programme) of 5
 June 5 between 00:00 and 02:45 UTC.

#### 395 3.2. Chemical composition of rains

396 Dissolved and total concentrations of nutrients and TMs in the rain samples are presented in Table 2.

397 Among all measured dissolved concentrations,  $NO_3$  was the most abundant nutrient, followed by

398 ammonium (Table 2). The nitrite concentration was <u>under below</u> the limit of detection for the two rain

399 samples. Regarding TMs in rain, Fe and Zn presented the highest concentrations in rain samples with

400 the same order of magnitude (10 to 25  $\mu g$  L  $^{\text{-1}}$ ). Co, Cd and Mo had the lowest concentrations

**Commented [RS(-S31]:** Use ionic formula or name, not both in the same sentence



**Output** 380 (<0.1 μg L-1 in both events), whereas the other TMs concentrations ranged Discussions between 0.1 and 10 μg 381 L<sup>-1</sup> in both rain <u>sampless</u> (Table 2). Concentrations of nutrients and the majority of TMs were higher in the 382 dust-rich rain, except <u>for</u> Pb (similar concentrations in both rain <u>samples</u>) and Cr (3 times higher concentration 383 in Rain ION relative to Rain FAST).

384 Table 2: Dissolved and total concentrations of nutrients and TMs in the two rain <u>samples</u> collected during the 385 PEACETIME cruise in µg.L<sup>-1</sup> or ng.L<sup>-1</sup> and µmol.L<sup>-1</sup> or nmol.L<sup>-1</sup> in the parentheses (sd = standard deviation 386 from three replicates).

			Rain ION							Rain FAST							
			Dissolved			Total			Dissolved				Total				
Nutrients		μg L <sup>-1</sup> (μm ol L <sup>-1</sup> )	concentrations ±sd		sd	concentrations		±	sd	concentrations		±sd		concentrations		±sc	
	s NO3		1185	(19.1)	71	(1.1)					3694	(59.6	222	(3.6			
	NH4 <sup>+</sup>	$\mu gL^{\text{-l}}(\mu m\text{ol}L^{\text{-l}})$	366	(20.3)	11	(0.6)					654	(36.3	19	(1.1			
	DIN	$\mu gL^{\text{-}1}(\mu m\text{ol}L^{\text{-}1})$	552	(39.4)	82	(1.7)					1343	(96)	241	(17			
	PO43-	$\mu g  L^{\text{-}1} (\text{nm ol}   L^{\text{-}1})$	18.1	(189)	0.5	(6)					19.0	(200)	0.6	(6)			
	DIP	$\mu g  L^{\text{-}1}(\text{nm ol}   L^{\text{-}1})$	5.87	(189)	0.18	(6)					6.20	(200)	0.19	(6)			
	DOP	$\mu g  L^{\text{-}1}(\text{nm ol}   L^{\text{-}1})$	8.63	(278)	1.94	(75)					4.91	(158)	1.56	(57)			
	TP	$\mu g  L^{\text{-}1}(\text{nm ol}   L^{\text{-}1})$	14.51	(468)	2.52	(81)	16.6	(536)	1.0	(33)	11.11	(358)	1.95	(63)	58.7	(1894)	3.5
	DIN/DIP			(208)								(480)					
	DOC	(µm ol L <sup>-1</sup> )		(105.7)		(2.2)						(95.5)		(1.2)			
Metals	Al	$\mu g  L^{\text{-l}}  (\text{nm ol}   L^{\text{-l}})$	13.0	(480)	0.8	(30)	14.6	(540)	0.9	(32)	23.4	(867)	0.7	(24)	440	(16308)	7
	Cu	$\mu g  L^{\text{-l}}  (\text{nm ol}   L^{\text{-l}})$	0.71	(11.1)	0.02	(0.3)	0.73	(11.5)	0.02	(0.3)	1.15	(18.0)	0.04	(0.6)	1.63	(25.7)	0.06
	Fe	$\mu g  L^{\text{-}1}  (\text{nm ol}   L^{\text{-}1})$	15.1	(270)	0.4	(6)	17.9	(321)	0.6	(11)	19.2	(344)	0.1	(2)	231	(4140)	7
	Mn	$\mu g  L^{\text{-}1} (\text{nm ol}   L^{\text{-}1})$	0.55	(10.0)	0.02	(0.3)	0.60	(10.9)	0.02	(0.4)	3.17	(57.8)	0.07	(1.2)	5.26	(95.7)	0.12
	Ni	$\mu g  L^{\text{-}1}  (\text{nm ol}   L^{\text{-}1})$	0.52	(8.8)	0.02	(0.3)	0.67	(11.4)	0.02	(0.4)	0.59	(10.1)	0.02	(0.4)	0.84	(14.3)	0.03
	Ti	$\mu g  L^{\text{-}1}  (\text{nm ol}   L^{\text{-}1})$	0.48	(10.0)	0.04	(0.8)	0.65	(13.6)	0.48	(3.2)	0.22	(4.7)	0.01	(0.1)	33.36	(697)	0.51
	v	$\mu g  L^{\text{-l}} (\text{nm ol}   L^{\text{-l}})$	0.37	(7.4)	0.01	(0.2)	0.38	(7.42)	0.01	(0.25)	1.37	(26.9)	0.03	(0.5)	2.02	(39.7)	0.04
	Zn	$\mu g  L^{\text{-l}} (\text{nm ol}   L^{\text{-l}})$	24.8	(379)	0.8	(12)	25.3	(387)	0.8	(12)	22.7	(347)	0.6	(8)	26.3	(402)	0.7
	Cd	$ng L^{-1}(pmol L^{-1})$	12.9	(115)	6.4	(57)	13.1	(117)	6.3	(56)	20.2	(180)	10.3	(92)	23.7	(210)	6.8
	Co	$ng L^{-1}(pmol L^{-1})$	44	(749)	13	(229)	47.4	(804)	14.5	(246)	82	(1386)	20	(347)	157	(2661)	28
	Cr	$ng L^{-1}(pmol L^{-1})$	241	(4636)	16	(300)	628	(12079)	5	(95)	79	(1522)	14	(260)	443	(8514)	43
	Mo	$ng L^{\cdot 1}(pmol L^{\cdot 1})$	28	(288)	10	(106)	4.1	(43)	1.4	(14)	82	(855)	11	(113)	92.1	(960)	16.2
	Pb	$ng L^{-1} (pmol L^{-1})$	170	(822)	11	(54)	175.1	(845)	1.4	(7)	166	(801)	9	(41)	604	(2917)	19

387

#### 388 3.3. Marine concentrations and stocks

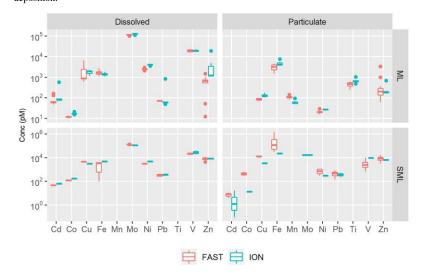
All the TMs had significantly higher concentrations in the ML compared to deep water masses, in agreement with a stratified profile associated with atmospheric input. The particulate and dissolved 391 trace metal concentrations within the ML (0-20 m) and the SML are displayed in Fig. 4.392 Concentrations were of the same order of magnitude in-at the two studied stations ION and FAST, except for the 393 particulate phase in the SML where the concentrations of Cu and Co were significantly lower (values) at 394 ION station. The TMs were mainly in dissolved forms in the ML (range of values), except for Fe, whose dissolved 395 and particulate concentrations were on the same order of magnitude (values). On the contrary, the particulate 396 phase contribution dominated for TMs in the SML, in particular at the ION station. At both stations, 397 the highest TMs concentrations in the surface seawater were found for Mo in the dissolved fraction (values),

**Commented [RS(-S32]:** Are you really talking about deep water masses or just the water below the ML?

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397 in agreement with the abundance of dissolved Mo in seawater (~ 107 nM, Smedley and Kinniburgh, 398 2017), and Fe in the particulate fraction (value). All the particulate and dissolved TMs concentrations 399 measured during the cruise were representative within the range previously published for the 400 Mediterranean Sea (Sherrell and Boyle, 1988; Saager et al., 1993; Morley et al., 1997; Yoon et al., 401 1999; Wuttig et al., 2013; Baconnais et al., 2019; Migon et al., 2020). Zn presented the largest range 402 of concentrations within the ML both in the particulate and dissolved phases (values), due to some 403 high concentrations. However, the concentrations stayed in the typical range of values found-reported 404 in for the Mediterranean Sea (Bethoux et al., 1990, Yoon et al., 1999). The concentrations In-in the 405 SML<sub>5</sub> the concentrations were lower than in the ML and Pb dominated both in dissolved and 406 particulate phases. Tovar-Sanchez et al. (2020) showed that the TM# concentrations in the SML during 407 the PEACETIME campaign were generally lower than those previously measured in the 408 Mediterranean Sea, except in for the particulate phase during at the FAST station after following dust 409 deposition.



**Commented [RS(-S33]:** Did any of these studies determine pMo and/or dMo? If not, say that yours are the first

**Commented [RS(-S34]:** Was contamination suspected?

Commented [RS(-S35]: I don't understand this statement

410

411 Figure 4: Box\_plots of dissolved (left panels) and particulate (right panels) marine concentrations

412 (pM) for the different TMs within the ML (upper panels) and the SML (lower panels) at ION (green)

- 413 and FAST (red) stations. In the box plots, the box indicates the interquartile range, i.e. the 25th and
- 414 the 75-th percentile, and the line within the box marks the median. The whiskers indicate the quartiles

**Commented [RS(-S36]:** Could these be plotted on the same scale to make it easier to compare the SML and ML concs?



415 ±1.5 times the interquartile range. Points above and below the whiskers indicate outliers outside the

- 416 **10-th and 90th percentile.**
- 417 **4. Discussion**
- 418 4.1. Composition of rain collected over the remote Mediterranean Sea

419 4.1.1. Concentrations

420 Regarding nutrients, nitrogen species concentrations in rain samples were in good agreement with

421 those reported in Mediterranean rain samples, ranging from 1130 to  $5100 \,\mu g \, L^{-1}$  for NO<sub>3</sub><sup>-</sup> and between

 $422 \qquad 207 \mbox{ and } 1200 \mbox{ } \mu g \ L^{\text{-1}} \mbox{ for } NH_4^{\text{+}} \mbox{ (Loye-Pilot et al., 1990; Avila et al., 1997; Al Momani et al., 1998; }$ 

423 Herut et al., 1999; Violaki et al., 2010; Izquieta-Rojano et al., 2016; Nehir and Koçak, 2018). The

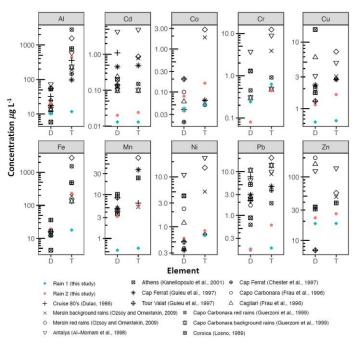
424 FAST rain concentrations were <u>with</u>in the <u>average valuespublished range</u>, whereas the ION rain was

- in the low range, confirming a background signature at this station. The rainwater samples presented
- 426 a large dominance of N in comparison to P, as observed from the N/P ratio derived from DIN/DIP
- (Table 2) ranging from 208 at ION to 480 at FAST. Previous observations showed a predominance of
  N relative to P in the atmospheric bulk deposition over the Mediterranean coast, with ratios higher
- 429 than the Redfield ratio (Markaki et al., 2010, Desboeufs et al., 2018). The highest ratio observed
- 430 reached 1200 (in <u>the case of DIN/TDP</u>), but were on-averaged around about 100 in bulk deposition
- 431 (unfiltered rain). The highest ratio could be linked to a washout effect of the gaseous N species (as
- 432 NOx and NH<sub>3</sub>) by rain (Ochoa-Hueso et al., 2011). At the two stations, no high NOx concentrations 433 were observed in the boundary layer before wet deposition (range). The presence of nitrate and
- were observed in the boundary layer before wet deposition <u>(range)</u>. The presence of nitrate and
   ammonium in the background aerosols has been <u>emphasized-observed</u> during recent campaigns in the
- 435 remote Mediterranean atmosphere (e.g. Mallet et al., 2019). To our knowledge, no data are available
- 436 on both P and N concentrations in Mediterranean aerosols. The lowest concentrations of P relative to
- 437 N in aerosol particles in <u>the Mediterranean have beenwere</u> observed during the cruise (<u>value</u>, Fu et
- 438 al., in prep.). The TDP concentrations were consistent with the average value of  $8.4 \ \mu g \ L^{-1}$  measured
- 439 in African dust rain samples collected in Spain over the 1996-2008 period (Izquierdo et al., 2012).
- 440 Inorganic phosphorus predominated in the dust-rich rain, whereas organic P was dominant in the441 background rain as the contribution of DOP to the TDP was 60% and 44% in Rain ION and Rain
- 442 FAST, respectively. The DOP/TDP ratio presents a very large range in Mediterranean rains, spanning
- 443 from 6% in Spanish dusty rain samples (Izquierdo et al., 2012) to 75-92% in rains from Crete Islands
- (Violaki et al., 2018). A reason for this wide range could be that Mediterranean European aerosols, as
- 445 opposed to Saharan dust particles, are dominated by organic phosphorus compounds associated with

446 bacteria (Longo et al., 2014).

Commented [RS(-S37]: Wet +dry?





447

Figure 5: Comparison of dissolved (D) and total (T) TMs concentrations along with data from 14 former studies carried
 out in to previous studies in the eastern and western Mediterranean Sea.

450 The dissolved and total TMs concentrations in the PEACETIME rains were lower than those reported 451 in coastal areas (eastern Basin: Özsoy and Örnektekin, 2009; Al-Momani et al., 1998; Kanellopoulou 452 et al, 2001 and western Basin: Guieu et al., 1997; Guerzoni et al., 1999b; Chester et al., 1997; Losno, 453 1989; Frau et al., 1996) (Fig. 5), notably especially for the background Rain ION. This suggests the 454 probable effect of both local anthropogenic influence at coastal sites due to higher aerosol 455 concentrations in comparison to the remote Mediterranean (Fu et al., in prep.) and due to the reduction 456 of anthropogenic emission for some elements since most of the referenced works on coastal rainwaters 457 date from the late 1990s, This is particularly true for Cd and Pb and Cd whose emissions have strongly 458 decreased over the last decades, notably due to removal of lead in gasoline and reduction of coal 459 combustion (Pacyna et al., 2007). This has resulted in a clear decrease in the particulate concentrations 460 of these metals in the Mediterranean atmosphere (Migon et al., 2008), consistent with the fact that 461 concentrations in the PEACETIME rains are one to two orders of magnitude lower than reported in

**Commented [RS(-538]:** Are you comparing like with like here? You state that concentrations likely decrease offshore and, although emissions of TMs may have decreased, if the literature data is from coastal sites you don't have direct evidence for this decline over the open ocean. However, you could infer it from reductions in leaded fuel and coal combustion. Is the decreasing atmospheric input evident in the full depth water column samples? Perhaps this is what the Fu et al data shows but the reader can't check this as the reference is 'in prep'.

Later you mention the Dulac thesis as evidence. This paragraph needs rewriting and making more concise in order to strengthen your final, important statement. As it stands, this section is too speculative.



the literature. For these metals, the discrepancy is also observed between the concentrations in our

463 open-sea rain samples and the concentrations (state which elements) measured in three rains collected

464 at sea in April 1981 (Dulac, 1986), confirming that the large decrease of concentrations could be

465 related with to the decrease of in anthropogenic emissions. Thus, our results show the former

466 <u>literaturedata from before [year] cannot should not be used as a current reference about for coastal</u>

467 rain composition due to recent environmental mitigation on of metals <u>TM</u> emissions.

468 4.1.2. Enrichment factors

469 Enrichment Factors (EF) and solubility values of TMs and P observed during the two rain events were 470 very contrasted (Fig. 6). In Rain ION, almost all elements were significantly enriched relative to the 471 earthupper continental crust (EF >10, and up to ~103 for Cd and Zn), whereas in Rain FAST, only Zn 472 (73), Cd (48) and Mo (15) were slightly enriched. Only Ti, Fe, and Mn did not present a significant 473 enrichment (EF < 10) in Rain ION, in agreement with previous studies in the Mediterranean 474 environment showing that these metals are mainly associated with mineral dust in -atmospheric 475 deposition (e.g., Guieu et al., 2010; Desboeufs et al., 2018). For bBoth rainsrain samples, the EF of 476 Zn was oin average five times higher than the EF found in the rains previously studied inrain samples 477 from coastal sites (?) in the Mediterranean region (Özsoy and Örnektekin, 2009; Al-Momani et al., 478 1998; Losno, 1989). However, extremely high enrichments of Zn in rainwater have also been reported 479 from island sites in the Mediterranean Sea, for example, by Frau et al. (1996) reported with a 480 geometric mean EFs of about 6500 in both crust-rich and crust-poor rains from two sites in southern 481 Sardinia, and- Fu et al. (2017) also-reported EF higher than 1000 for Zn in atmospheric bulk 482 deposition oin Lampedusa Island. The Zn EF at station ION is the same order of magnitude as at these 483 island sites which suggests - something about the anthro background signal being high in the open 484 Med. //The anthropogenic origin of particulate TMs and P concentrations in seawater have been reported 485

486 by several studies on atmospheric deposition in the western Mediterranean (e.g., Guieu et al., 2010; 487 Sandroni and Migon, 2002; Desboeufs et al., 2018). For example, Desboeufs et al. (2018) showed that 488 there is a large contribution of anthropogenic combustion sources to the in-P, Cr, V and Zn background 489 deposition fluxes. Aerosol composition monitoring over the Mediterranean coastal area showed the 490 role of land-based sources and ship traffic sources on TMs contents (Bove et al., 2016; Becagli et al., 491 2017). However, all these sampling sites were located in coastal areas, where it was difficult to 492 discriminate the potential local influences. Here, even if the on-board atmospheric gas and particle 493 measurements did not show a specific anthropogenic influence during the period of Rain ION, the 494 particles scavenged by this rain presented a clearsuggest an anthropogenic signature for all TMs **Commented [RS(-S39]:** Your rain was open ocean not coastal so doesn't really show that historical coastal TM concentrations are higher. The comparison is with the thesis. You could say that coastal values are not representative of open ocean values generally

**Commented [RS(-S40]:** This section jumps around a bit, Rain ION then Rain FAST, the Rain ION again, then both, etc. It would be easier to read if one station was discussed and the other one contrasted to it

**Commented [RS(-S41]:** Change to dust-rich and dust-poor

**Commented [RS(-S42]:** Do you mean atmos bulk wet depo here? If so, for clarity best to just say rain. If bulk dry you need to state that too

**Commented [RS(-S43]:** What do you mean here? To resolve the contribution of various local industries and shipping or local inputs from long range transport



495 except Ti, Fe and Mn.; <u>Hh</u>owever for Fe and Mn, <u>an-the</u> influence of non-crustal sources in Rain ION

496 is visible through a clear increase in the EF values compared to FAST (Fig. 6). This means that even

497 over remote the Med Sea, the chemical composition of background aerosol particles is likely

498 continuously impacted by anthropogenic sources.

499 Moreover, the EF values of TMs for Rain FAST were significantly lower than for Rain ION (Fig. 6)

500 but similar to Saharan rains (Guerzoni et al., 1999b; Özsoy and Örnektekin, 2009). The comparison

501 between dust-rich and background rains generally reveals a net difference of concentrations (at least

502 higher by a factor 3 in dust-rich), notably for Al, Fe, Mn and Cr (Guerzoni et al., 1999b; Özsoy and

503 Örnektekin, 2009). Such contrast was indeed observed for Al, Fe and Mn in the PEACETIME rains

504 (Fig. 5), but also for Cu and Pb. The combination of higher concentrations and EF values <10 found

505 in Rain FAST confirm that the dust contribution was important on deposition fluxes of many TMs.

#### 506 4.1.3. Solubility

507 The solubility values were also larger higher in Rain ION than in the dusty Rain FAST, except for Mo 508 for which the difference between both rain samples wais not significant (Fig. 6). For Rain ION, TMs 509 and P presented solubility higher than 78%, except for Cr (38%). In Rain FAST, solubility values 510 <10% were observed for Al and Fe, more than 10 times lower than in Rain ION. For the other TMs, 511 the highest difference in solubility was observed for Pb whose solubility decreased from 97% in Rain 512 ION to 27% in Rain FAST. In a review on TMs solubility in Mediterranean rainwaters collected in 513 coastal areas, Desboeufs (2021) emphasize the large range of solubility for all the TMs: Fe (0.8-41%), 514 Cr (6-80%), Pb (5-90%), Ni (22-93%), Mn (16-95%), Cu (22-96%), Zn (14-99%), V (35\_99%) and 515 Cd (72-99%). The solubility ranges found in this study were generally consistent with those reviewed 516 by Desboeufs (2021). In particular, the Mn solubility values in FAST (60%) and ION (92%) rains are 517 close to those reported by Dulac (1986) from a dust-rich (57%) and an anthropogenic (83%) rain 518 collected at sea in the Ligurian Sea and west of Sardinia in April 1981, respectively. Only Fe solubility 519 (84%) found in Rain ION was higher than the average values previously reported. In the Rain FAST, 520 Fe solubility was 8%, this is 10 times lower than the average Fe solubility in 10 dust-rich rains 521 collected on the southeastern coast of Sardinia by Guerzoni et al. (1999b), but consistent with Saharan 522 dust wet deposition collected in the Atlantic Ocean (Chance et al. 2015; Powell et al., 2015; Baker et 523 al., 2017).

**Commented [RS(-S44]:** These two sentences disagree with each other. Need to be reworded.

**Commented [RS(-S45]:** How does this study compare to the factor of three increase in concs between background and dust influenced rains – use your values – and state how this would impact EFs – more Al drives down EFs, Al predominantly from dust, etc.

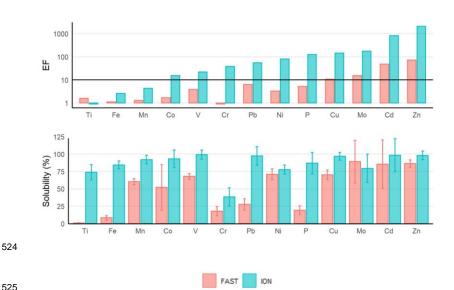
**Commented [RS(-S46]:** Vague, just say that concs were higher and EFs suggest the reason for this was the impact of the dust plume. The EFs don't really tell us this. This section needs tightening up.

Commented [RS(-S47]: Big target to hit!

**Commented** [RS(-S48]: Can you suggest a reason for the difference?

525





526 Figure 6: Enrichment Factors (EF, upper panel) and solubility (%, bottom panel) of phosphorus (P) and TMs ordered 527 by increasing EF in the two rainwater samples.

528 Few studies have compared TMs solubility between dust-rich and background rains in the Mediterranean. In Sardinia, Guerzoni et al. (1999b) observed an increase in solubility values from 529 530 dust-rich to background rains for Al, Cr, Fe and Pb (and hardlybut only a slight increase for Cd), and 531 reported an inverse relationship between the particle concentration and solubility of Al, Fe, Pb and 532 Cd. Similarly, Theodosi et al. (2010) showed a decrease in TMs solubility with the increase in dust 533 load in rains collected on Crete Island. In those two studies, this the magnitude of the decrease was dependent on the considered metalspecific to the TM, with Pb presenting the highest decrease in 534 535 solubility. The decrease in solubility from background to dust\_rich rains was also-observed for P in 536 Spain (where?) by Izquierdo et al. (2012), with values of solubility decreasing from 25% to 7%, and 537 for Mn in offshore rains as mentioned above (Dulac, 1986).

538 //Metal partitioning in rainwater can be influenced by a number of parameters, such as pH, presence

- 539 of dissolved organic complexing ligands, in-cloud processinges, particle origin and load (Desboeufs
- 540 et al., 1999; Bonnet and Guieu, 2004; Paris and Desboeufs, 2013; Heimburger et al., 2013). However,
- 541 the particulate desert-mineral dust load, reflecting the dust vs-versus anthropogenic signature, is the
- 542 main control of TMs solubility in the Mediterranean rainwater (Özsoy and Örnektekin, 2009;

Commented [RS(-S49]: Could you make the two panels the same size and the bars the same width?

Commented [RS(-S50]: Could draw parallels with aero samples here, e.g. Jickells et al. (2016)

Commented [RS(-S51]: Dust loading is only one proposed control and may not be a direct control - provide explanation e.g. perhaps the CaCO3 concentration is or the reduced impact of acidic gases, low RH, etc is - ref Baker et al., 2021



543 Theodosi et al., 2010). A much lower solubility of TMs in Rain FAST than in Rain ION (except for

544 Mo) is consistent with the EFs in Rain FAST (Fig. 6), although no correlation between solubility and

545 EF values could be observed. The case of Mo is unique, since its solubility was comparable in Rains

546 ION and FAST despite a >20 times higher EF in Rain ION. As Mo solubility is seldom studied in the

547 literature, we could not conclude on the reason for this particular outcome. It is interesting to note

that despite the desert signature, the majority of metals have solubility greater than 50% in rain FAST.

#### 549 4.2. Atmospheric wet deposition as a source of TMs to the surface seawater

550 4.2.1. Atmospheric fluxes

As mentioned before, the two collected rains were part of large rain systems, associated with patchy rainfalls that lasted several hours or days (section 3.1). This spatio-temporal variability led to

553 heterogeneity in both rainwater concentrations and accumulated precipitation across the studied

region. Such spatial variability has been observed by Chance et al. (2015) in the Atlantic Ocean.

555 Moreover, even weak lateral advection can transfer surface water impacted by intense precipitation in

the vicinity of the vessel. On this basis, the spatial extrapolation of wet deposition fluxes seems subject

557 to a large uncertaintiesy when the rain samples are not collected across the rain area (Chance et al.,

2015). To best counteract this effect, spatial variability was taken into account to quantify the total precipitation i.e.  $3.5 \pm 1.2$  mm for rain ION and  $6.0 \pm 1.5$  mm for rain FAST (see 3.1) in order to

560 quantify the wet deposition fluxes.

561 From the total (dissolved + particulate) Al concentration measured in the R+ain FAST sample, we 562 estimated the wet mineral dust deposition flux at  $65 \pm 18$  mg m<sup>-2</sup>, assuming 7.1% Al in dust (Guieu et 563 al., 2002). The vertical distribution of dust particles (Fig. 2b) and the absence of high Al 564 concentrations close to the sea surface (Fu et al., in prep.) indicate that dust dry deposition can be 565 neglected. Based on the increase in total Al in the upper 20 m of the column water following the 566 deposition events, Bressac et al. (2021) derived an average dust deposition flux of ~55 mg m<sup>-2</sup> at 567 FAST station which is, comparable to our estimate. Although low compared to deposition fluxes 568 reported in the western Mediterranean (Bergametti et al., 1989; Loÿe-Pilot and Martin, 1996; Ternon 569 et al., 2010), such our flux values estimates are among similar to the most intense weekly dust 570 deposition fluxes recorded more recently infrom Corsica between 2011 and 2013 (range) and is

571 equivalent to the mean weekly flux (value) observed inreported for Majorca Island during the same

- 572 period (Vincent et al., 2016). The aerosol-columnar aerosol concentration during the dust event at the
- 573 FAST station being-was estimated to be between 0.18 and 0.24 g m<sup>-2</sup> (see Section 3.1.), the expected
- 574 maximum values of atmospheric dust flux could be in this range. The comparison with the estimated

**Commented [RS(-S52]:** Therefore, your data does not support this argument. Disagreement within statement. The EFs close to crustal TMs in RainFAST combined with the higher concentrations point to the role of mineral dust in reducing the fractional solubility of TMs in rainwater – or the presence of mineral dust overwhelms the background signal resulting in a net decrease in fractional solubility – although the net effect is the same, I favour the latter explanation

**Commented [RS(-S53]:** Mo has a high conc in SW. If it has a predominantly marine source this could explain your uniform solubility. It would not explain your increase in EF necessarily but if there was an increase in seasalt aerosols it could do. You have the data to test this hypothesis

**Commented [RS(-S54]:** Marine source? How reliable were the solubility data? Close to LoD and blanks?

#### Commented [RS(-S55]: How large?

**Commented [RS(-S56]:** This is different than the Al proportion used in your EF calculations (Rudnick and Gao, 2003). Your EFs would be lower if you used concentrations based on 7.1% Al (Guieu et al., 2002) or your fluxes would be higher if you used the Rudnick and Gao(~ 8 %). It would be worth noting this variability. There is certainly an argument for using Saharan elemental ratios in your EF calcs (as discussed in Shelley et al. (2105)

**Commented [RS(-S57]:** Fluxes aren't measured directly so reported, calculated or estimated are better words

**Commented [RS(-S58]:** This is x3-4 higher than your flux estimates. A sentence to clarify why you support this view.



flux indicates that the atmospheric column was probably not totally washed-out by the short rain event.

576 Indeed, Fig. 2b shows that a significant depolarization was observed immediately after the rain ended

577 on the ship, before atmospheric advection could have brought dusty air possibly not affected by rain.

578 Satellite products (Fig. S2) confirm that on 5 June, the dusty air mass was transported farther to the

579 north-east from the station where it was replaced by clear air.



**Commented [RS(-S59]:** It would be useful to see the Al wet deposition flux as this is what you are using as the basis to estimate the bulk wet depo flux. It would also be nice to have a third panel showing the ratio or the Kd of dissolved to particulate TMs

580

Figure 7: Dissolved (upper panels) and particulate (lower panels) wet deposition fluxes (µmol m<sup>-2</sup>) for
 the different TMs estimated from the two rains sampled on-board, considering the standard deviation
 on the TMs concentrations and the spatial variability of total precipitation over the area of sampling
 (Rain ION in blue and Rain FAST in red). Note different scales on the y axes.

585 The atmospheric dissolved and particulate wet deposition fluxes of TMs, derived from the chemical 586 composition of rain samples and total precipitation, are presented in Fig. 7. Co, Mo and Cd presented 587 the lowest fluxes in the two rainfalls. Zn and Fe fluxes were oin the same order of magnitude and were 588 the highest dissolved fluxes compared to the other TMs in the two rains. The comparison showed that 589 almost all the dissolved TMs fluxes were higher in the dusty rain, except Cr and Ti. For the particulate 590 phase, the fluxes were mainly increased by the presence of mineral dust deposition for Co, Fe, Mn, 591 Pb, Ti and V. Our results emphasized support previous studies that report that the presence of mineral 592 dust-deposition, even here in the case of a-the moderate deposition input flux reported here, enabled 593 resulted in higher atmospheric inputs of TMs than from a-low perturbed, anthropogenic background 594 rain. We found this to most This is notably be the case for dissolved Cd, Cu, Mn, V and Zn and for 595 particulate Fe and Ti with more than one an order of magnitude fluxes difference in input fluxes 596 between theestimated from the two rain eventss. The orders of magnitude found in this study could 597 be used as "typical a benchmark atmospheric fluxes" to estimate atmospheric inputs of TMs by a rain

**Commented [RS(-S60]:** Explain how these two parameters differ. Are they not the same thing here?

**Commented [RS(-S61]:** Explain why. They also have the greatest uncertainties – low concentrations and high blank contributions and >100% CRM recovery for Mo ?

**Commented [RS(-S62]:** Because these elements were primarily found in the particulate phase?

**Commented [RS(-S63]:** Emphasised is stretching it for data from just two samples

**Commented [RS(-S64]:** With the exception of Mn, these are predominantly considered pollution derived elements. It would be worth mentioning this.

**Commented [RS(-S65]:** Poorly soluble, lithogenic elements. Al should also be reported here. I'm assuming it showed the same behaviour as Fe and Ti?

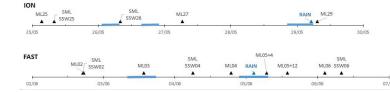
Commented [RS(-S66]: Do you mean the range?



event-from wet deposition to the western Med Sea. However, wWe must keep in mind, however, that
annual and long-term deposition fluxes of dust-related elements (e.g. Ternon et al., 2010), but also
nitrogen species (e.g., Richon et al., 2018b), in the Med Sea, are dominated by a few atypical, intense
deposition events in the Med Seawhen they occurred, as is the case in many other oceanic regions
(Duce et al., 1991),-

603 4.2.2. Comparison between TM<del>s</del> wet deposition inputs and marine stocks at ION and FAST stations

604 Our sampling strategy of Marine sampling sequences carried outcollecting seawater before and after 605 the rains were, to the best of our knowledge, the first direct observations intended to trace the fate of 606 atmospheric metals TMs and nutrients in the water column after wet deposition events. The time chart 607 of the sampling of rain and column water (surface microlayer, subsurface seawater and mixed layer) 608 is presented in Fig. 8. The impact of the two wet deposition events on nutrients stocks in the 609 Mediterranean surface waters is discussed in detail in van Wanbeke et al. (2020) and Pulido-Villena 610 et al. (2021). To briefly summarise, bBoth nitrate and DIP increased in the ML following the rains. 611 Although the closure of the N and P budgets had to necessarily take into account post-deposition 612 processes such as new nutrient transfer through the microbial food web (uptake, remineralisation, and 613 adsorption/desorption processes on sinking particles), it was shown that wet deposition was a 614 significant source of nutrients for ML during the cruise (roughly what %?). We focus here on the role 615 of TMswet deposition as a source of metals TMs to the column water. To do so, we estimated the 616 potential enrichment of the SML and ML from the rain by calculating the difference (delta) in TMs 617 stocks before and after rains.



618

619Figure 8: Sampling chronology during the ION and FAST stations for SML, SSW and ML. The blue620periods correspond to rainfall in the station area (after ERA 5 reanalysis and radar imagery, see621section 3.1). Samplings were-was performed 4 days and 2 days before and 2 h after Rain ION, and at622a higher frequency at the FAST station: 57, 37 and 7.5 hours before and 4.5, 12, 24 hours after Rain623FAST. SML and SSW samples could not be collected immediately before and after the rains because624of bad weather conditions, and were collected 3 and 4 days before Rain ION, and 57 and 20 h before625and 30 h after Rain FAST.

**Commented [RS(-S67]:** What does this mean? All elements discussed or only lithogenic elements?

**Commented [RS(-S68]:** I'm wondering if concentrations or partitioning are a better choice of word than stocks (this applies throughout the paper)

Commented [RS(-S69]: ... in an open ocean setting

Commented [RS(-S70]: Move to Methods

Commented [RS(-S71]: Move to methods

**Commented [RS(-S72]:** Given how quickly the SML TM concentrations respond to atmospheric inputs and their relatively short residence times in the SML (hours – Ebling and Landing, 2015), how can you be sure that your SML sampling resolution was capturing the impact of the wet deposition events of the SML, especially given that wind speed (and sea state) is a critical factor in determining the integrity of the SML?



626 At ION, no SML sampling was done after rain, preventing the study of the rain effect. For the ML, 627 the large variability in total and dissolved stocks between the two casts ML25 and ML27 before the

628 rain makes the establishment of a background concentration levels before rain difficult. ML27 was

629 used for theas-initial conditions since as it wais the closest sampling from-point to the post-rain

630 measurements sample collection (ML29). As mentioned previously, dusty rain deposition over the

631 FAST station area started on 3 June. Bressac et al. (2021) showed that the dust signature, traced by

632 changes in Al and Fe stocks in the ML, was already visible from the ML03 sampling. We defined the

633 enrichment of seawater layers as the difference between the maximum stocks after rains (from SML

634 04 or 06 and ML05+4) and the initial seawater stocks (SML02 and ML02).

635 At ION, only particulate Cu (+27%) and Zn (+44%) stocks increased in the ML after the rain. Even if

636 the dissolved forms of Cu and Zn predominated in the ML, this increase was accompanied by 637 increasing Kd values, i.e. in the particulate/dissolved partitioning (0.07 vs 0.12 for Cu and 0.14 vs 0.2

638 for Zn). This was also the case for Fe (Kd increased from 2.6 to 4.3), although no significant difference 639

(<5%) was observed on the particulate Fe stock. The Kd values in the ION rain sample being higher

than in the marine stock before the rain, that suggests the wet deposition at ION is mainly an additional 640

641 source of particulate TMs.

642 At FAST, the ML stocks increased in the particulate phase for Fe (+61%), Mn (+15%) and Zn (+9%)

643 and in the dissolved phase for Cu (+9%), Fe (+46%), Pb (+8%) and Zn (+15%) (Fig. 9). In addition

644 of to the marine inventories, the particulate TMs inputs by rain was also observed on Kd values and

645 total X/Al in the ML. For example, Kd(Fe) increased from 0.14 to 0.17 in ML and its Kd was 0.25 in

646 the rain. Even for  $Ni_{a}$  for which no change in stock could be evidenced, Kd(Ni) decreased from 0.1 to 647

0.07 and its Kd in the rain was 0.006. For Mn/Al, the Kd which fell from 0.27 before the rain to 0.008 648 after the rain (ML05+4), in accordance with the rain ratio (0.004). In the SML, the dissolved and

649 particulate stocks increased following rains for all TMs, from a factor 1.5 (Mo) to 10 (Fe) for the

650 dissolved phase and from a factor 1.6 (Ni) to 67 (Fe) for the particulate phase (Fig. 9). Commented [RS(-S73]: As it reforms so quickly and rain disrupts it, how can you be sure that the SML sampled after the rain at FAST was in contact with the atmosphere and accumulating RW TMs?

Commented [RS(-S74]: Suggesting that there was a surface advective current?

**Commented [RS(-S75]:** Dissolved or particulate or both?

Commented [RS(-S76]: I wonder if this is because of the mismatch between SML sampling and rosette casts.

Commented [RS(-S77]: Additional to what?

Commented [RS(-S78]: Particles scavenging dissolved TMs?

Commented [RS(-S79]: Again, the SML samples may not have been representative of the wet depo inputs. In contrast, the ML likely was



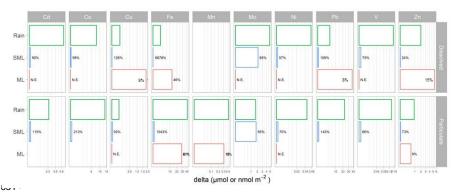


Figure 9: Comparison between TMs wet deposition fluxes (in green) and TMs marine stock deltas
(before and after the rain) in the SML (in blue) and in the ML (in red) at FAST. Dissolved = upper
panels and particulate = lower panels. Marine stocks increases are expressed in absolute values (Cd,
Co and Pb stocks in nmol m<sup>-2</sup>, and the other TMs in µmol m<sup>-2</sup>) and in relative values (%). N.E.t.\_= not
enhanced (increase <5%).</li>

657 The comparison between the observed enhancements in the SML stocks and the rain inputs at FAST 658 (Fig. 9) indicates that the atmospheric fluxes inputs can explain support all observed deltas. Indeed, 659 the atmospheric particulate and dissolved fluxes of TMs were 2 to 4 orders of magnitude higher than 660 the mean stocks present in the SML, except Mo which was oin the same order of magnitude. In the 661 ML, the magnitude of atmospheric particulate inputs was higher or similar to the particulate marine 662 delta of Fe, Mn and Zn. For Cu, Fe, Pb and Zn, the increase in dissolved stocks within the ML was 2 663 to 10 times higher than what could be provided from the atmospheric inputs. As described in Guieu 664 et al. (2020), marine dynamical conditions at FAST were favourable to observe any change in the 665 water masses strictly attributed to external inputs coming from the atmosphere on a short time scale. 666 However, a cumulative effect of previous and surrounding wet deposition events could explain this 667 difference between atmospheric inputs from rain FAST and increase of marine dissolved stocks. We 668 cannot exclude the possibility of lateral transport of metals-TM from surrounding waters being 669 enriched by the rain events of 3 June 3 for example, as revealed by the increase in the 0-20 m Fe and 670 Al inventories (Bressac et al., 2021). Another hypothesis to explain that higher stock increases of 671 metals in the ML than the one derived from atmospheric deposition is related to post-deposition 672 processes. Indeed, once deposited, the atmospheric particulate fraction could still be partly solubilized 673 in seawater, as the solubilisation of TMs (e.g. Fe) could occur over several hours or days (Wagener et 674 al., 2008; Wuttig et al., 2013; Desboeufs et al., 2014). This could lead to an underestimate of the

**Commented [RS(-S80]:** If we assume that the SML samples were representative of this layer before and after rain (and even if we don't), the similarity in Mo concentrations suggests to me that this element is not primarily delivered by atmospheric inputs

**Commented [RS(-S81]:** From rain or total atmos depo?

Commented [RS(-S82]: Days?

**Commented [RS(-S83]:** Some ML residence time calculations could be of use here

**Commented [RS(-S84]:** Many people have shown a delayed response of the dissolved pool to atmospheric inputs



675 dissolved TMs from atmospheric inputs. Moreover, the time lag between the rain and the first SML 676 sampling (1 day) does not allow us to conclude on the role played by SML as a "trap" of the added 677 dissolved metals by rain. However, results showed clearly the increase of dissolved TMs in the SML 678 even 24 h after the rain (Fig. 9). Even if this increase could be due to dissolution processes (Tovar-679 Sanchez et al., 2020), we cannot exclude that the residence time of dissolved atmospheric TMs in the 680 SML was sufficient to mask the atmospheric inputs in the ML05+4 sample. It is also known that 681 dissolved concentrations in the ML are subject to various biological processes such as phytoplankton 682 uptake (Morel et al., 2003). The comparison between ML05+4 and ML05+12 samples performed after 683 the rain shows that the dissolved and particulate stocks decreased quickly for all the TMs (not shown), 684 in agreement with the predominance of removal processes (sedimentation, biological transfer, 685 adsorption) on these stocks. However, the rate of decrease depended on the TMs, showing that some 686 removal processes predominate over others depending on the metal. For example, the dissolved metals 687 TM\_decreases could correspond to scavenging onto particles, which is a common physicoel chemical 688 process occurring in the ocean for Fe (Wagener et al., 2010; Bressac and Guieu., 2013) or, -Co (Migon 689 et al., 2020) and many other TMs.

690 <u>//</u>Finally, our results show that dust<u>y</u> wet deposition was a net source of all the studied trace metals

for SML both in the dissolved and particulate fraction. For ML, atmospheric dust inputs were also a net source of particulate Fe, Mn, and Zn, and dissolved Cu, Fe, Pb and Zn. Due to various marine

692 net source of particulate Fe, Mn, and Zn, and dissolved Cu, Fe, Pb and Zn. Due to various marine 693 post-deposition processes, it was is more complicated to observe the effect of wet deposition on

693 post-deposition processes, it <u>was-is</u> more complicated to observe the effect of wet deposition on 694 dissolved stocks, explaining why the SML and ML particulate stocks were more impacted by rain

695 than the dissolved stocks. On a timescale of hours, the Fe inventory was the most impacted by the

696 dusty rain input, both in dissolved and particulate phases, confirming that the dust-rich rains are a net

697 source of Fe to the surface Mediterranean Sea (Bonnet and Guieu, 2006, Bressac et al., 2021).

698 4.2.3. Comparison between TM# wet atmospheric inputs and marine stocks at the scalein the Western
 699 western and Central Central Mediterranean Sea

As observed from dissolved TMs stocks measured before and after the rains, a large part of the

701 uncertainties in the data analysis results from various removal processes of TMs after wet deposition,

702 which could have time resolution shorter than the sampling step. In order to limit the effect on of these

- 703 potential processes in-on\_data analysis, here we further study the role of wet deposition by comparing
- atmospheric dissolved fluxes to marine dissolved stocks by using TMs profiles in the ML at all 13
- marine stations, i.e. 22 ML samplings, throughout the whole cruise (Fig. 10). Indeed, considering that
   the collected rains were originating from large rain systems covering more than 50 000 km<sup>2</sup> around
- the sampling zone and were typical of Mediterranean wet deposition, we hypothesized that they could

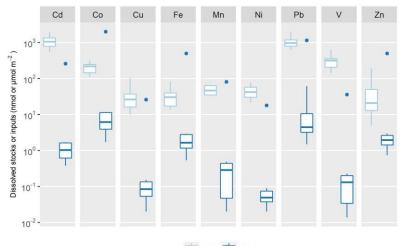
**Commented [RS(-S85]:** It would be useful to see this – put in Supplementary Materials

**Commented [RS(-S86]:** I'm not completely convinced it does because of the resolution problem. It certainly suggests it might given Tovar-Sanchez's findings of increases 24 h after rain but the signal is likely to be (significantly) diminished after this time.



708 have occurred in any of the explored areas during the cruise. Exceptional intense dust deposition 709 events have been recorded in the Mediterranean, reaching 20 g m<sup>-2</sup> (Bonnet et al., 2006). Sporadic and 710 intense wet dust deposition higher than 1 g m-2 are regularly observed in the spring in the western 711 Mediterranean basin (e.g., Vincent et al., 2016). At the beginning of the cruise, an intense wet dust 712 deposition event (not collected) occurred over the South of Sardinia and over the Tyrrhenian Sea with 713 fluxes reaching about 9 g m<sup>-2</sup> (Bressac et al., 2021). In order to take into account the effect of such 714 an event, we also estimated the atmospheric fluxes of dissolved metals-TMs based on a 9 g m<sup>-2</sup> wet 715 dust deposition event eonsidering using solubility values found in the estimated from rain-Rain FAST 716 (Fig. 10). The As TM metal solubility decreasing decreases with increasing dust load (Theodosi et al., 717 2010), this estimation constitutes probably a maximum value of the dissolved inputs of trace metals 718 by such a dust deposition event. In addition to removal processes, the impact of rain inputs on TMs 719 marine stocks is also controlled by MLD fluctuations that which we ignored in the this work described 720 above by using a fixed ML depth for the FAST and ION stations. As However, s the variability of this 721 MLD (7-21 m during the cruise, typical of Mediterranean thermal stratified period) could change the 722 marine budgets by a factor of 3....So-we considered-used the measured MLD (Van Wambeke et al.,

723 2020) <u>at each station for calculating the marine budgets of TM water budgetss at each station</u>.



📋 ML ᄇ Rain

**Commented [RS(-S87]:** Which ones? Theodosi reports this for the TMs he studied but not your full suite. There are some that do, some that don't. See Jickells et al., 2016 and Baker et al., 2020. I appreciate these papers report data for the Atlantic rather than the Med but as they report data from Saharan and European air masses they are relevant.



Figure 10: Comparison of marine stocks in the ML at all the stations occupied during the PEACETIME cruise with atmospheric inputs estimated (1) from ION and FAST rains (Boxesboxes) and (2) from an intense wet dust deposition event of 9 g m<sup>-2</sup> (blue dots). Cd, Co and Pb stocks are in nmol m<sup>2</sup>, and the other TMs in µmol m<sup>-2</sup>. For Mn, marine stocks are derived from surface concentrations close to Corsica coasts (Wuttig et al., 2013: samples OUT at 0, 5 and 10 m) and in the Ionian Sea (Saager et al., 1993: Bannock basin at 0, 10, and 15 m), as no measurement is available

- 731 from the PEACETIME cruise. Boxes and whiskers as in Fig. 4.
- 732 Applying to the whole transect, the atmospheric inputs, obtained from our rain composition, were at
- 733 least 100-fold smaller than the dissolved stocks in the mixed layer, except for Co, Fe and Zn. The
- atmospheric inputs represented more than 30% of the dissolved Zn stocks and 10 to 18% for Fe. This
- 735 significant input of dissolved Fe is in agreement with our field observations in the ML. For

736 Co, the maximum atmospheric fluxes estimated during the cruise represented >10% of stocks. Here the

comparison is based only on dissolved TMs in rain water, yet as discussed previously, the solubilisation
 post-deposition of atmospheric particles in the water column could further enrich the marine dissolved

- range stocks. In the case of the intense dust deposition event, the dissolved inputs are of the same order of
- 740 magnitude as marine stocks for Co, Fe, Mn, Pb and Zn. The enrichment in dissolved Fe and Mn was
- 741 previously observed by Wuttig et al. (2013) after artificial dust seeding in large mesocosms (simulating
- 742 a wet deposition event of 10 g m<sup>-2</sup>). The surface seawater could be significantly affected by the deposition
- 743 of these dissolved elements in the case of wet dust deposition. As tThe marine TMs concentrations
- 744 measured during the cruise being were typical of Mediterranean surface seawater concentrations, we
- 745 conclude from these comparisons that wet deposition events, notably wet dust deposition events, prove

to bewere an external source of dissolved TMs for the Mediterranean Sea during the period of thermal
 stratification.

#### 748 5. Conclusions

749 This study provides both the dynamical properties and chemical characterization of two rainwaters 750 rain events collected in the open Mediterranean Sea, concurrently with TM (list) s marine stocks in 751 surface seawater. Our results are the only recent report of TM\_concentrations, EFs and fractional 752 solubility values for TMs in rain samples collected in the remote Mediterranean Sea. By highlighting 753 the discrepancy between TMs concentrations with the previous offshore and coastal rain studies, this 754 work demonstrates the need to provide a new and recent database on metal-TM composition in 755 Mediterranean rains in order to estimate the role of atmospheric TMs deposition. We showed-have 756 shown the representativeness of rain-Rain FAST as typical of Saharan dust wet deposition as well-in 757 its chemical composition as well as in its magnitude and extent, whereas Rain ION is a more-typical Commented [RS(-S88]: Only dissolved or dissolved and particulate?



of an low perturbed anthropogenic background rain of for the remote Mediterranean Sea. On this
 basis, we suggest to useusing the chemical composition of PEACETIME rains as a new reference for
 the studies of TMs on wet deposition in the Mediterranean Sea.

761 Since atmospheric TMs have been identified as critical oligo-nutrients for the marine biosphere, it is

762 important to study the response of the receiving waters to atmospheric inputs. our This study is the

763 first to provide in situ evidence that atmospheric wet deposition constitutes a significant external

764 source for some of these elements to surface stratified Mediterranean seawater. Our results show We

765 recommend that the original approach developed here is very relevant in this purpose and could beis

766 used in other parts of the world where atmospheric wet deposition is suspected thought to impact the

767 marine biosphere, such as in HNLC areas.

768

769

Data availability. Guieu et al., Biogeochemical dataset collected during the PEACETIME cruise.
SEANOE. https://doi.org/10.17882/75747 (2020). Atmospheric Data are accessible on http://www.obsvlfr.fr/proof/php/PEACETIME/peacetime.php.

773 Author contributions. KD and FF designed the study and wrote the manuscript; FF, ST, JFD, Ch.G made

774 the on-board atmospheric measurements and sampling during the cruise; FF, ST and JD analysed the rain

775 samples; MB, ATS, and ARR made the marine TMs sampling and analyses; PF was the reference scientist

776 of PEGASUS, AF and FM managed all the technical preparation of atmospheric samplings, PC analysed

the lidar data; KD, FD and Ce.G designed the cruise strategy; KD and Ce.G coordinated the PEACETIME

778 project, FD coordinated the ChArMEx funding request, and near-real time and forecast survey of

atmospheric conditions during the cruise; all the authors commented on the manuscript and contributed to

780 its improvement.

781 Competing interests. The authors declare that they have no conflict of interest.

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## 1087 Figure Captions:1088

1089 Figure 1: Total precipitation (mm) between May 28 at 20:00 UTC and 29 at 10:00 UTC from ERA5 ECMWF 1090 reanalysis. The red circle indicates the R/V position.

1091 Figure 2: On-board lidar-derived a) Apparent backscatter coefficient (ABC), (b) Temporal evolution (in 1092 Local Time) of the lidar-derived volume depolarization ratio (VDR) where the dust plume is highlighted for 1093 values higher than ~1.7 (yellow to red colors) and the rain by values higher than 3 (indicated by the arrow), 1094 and c) vertical profiles of the aerosol extinction coefficient (AEC) in cloud free condition, integrated over 3 1095 periods along the dust plume event, noted A, B and C in figure a. The grey shade represents the root mean

- 1096 square (rms) variability along the time of the measurement. The dust layer is highlighted on the profiles. The
- 1097 mean aerosol optical thickness is given in the boxed legend with its temporal variability (1 sigma). The 1098 location of the marine boundary layer (MBL) is also pointed.

1099 Figure 3: Rain rates (mm/h) during the night between the 4 and 5 June, when Rain Fast was collected on1100 board, issued from European rain radar composites (OPERA programme) of June 5 between 00:00 and 02:451101 UTC.

1102 Figure 4: Boxplots of dissolved (left panels) and particulate (right panels) marine concentrations (pM) for 1103 the different TMs within the ML (upper panels) and the SML (lower panels) at ION and FAST stations. In 1104 the box plots, the box indicates the interquartile range, i.e. the 25<sup>th</sup> and the 75<sup>th</sup> percentile, and the line within 1105 the box marks the median. The whiskers indicate the quartiles ±1.5 times the interquartile range. Points above 1106 and below the whiskers indicate outliers outside the 10<sup>th</sup> and 90<sup>th</sup> percentile.

1107 Figure 5: Comparison of dissolved (D) and total (T) TMs concentrations along with data from 14 former 1108 studies carried out in the eastern and western Mediterranean Sea.

1109 Figure 6: Enrichment Factors (EF, upper panel) and solubility (%, bottom panel) of phosphorus (P) and TMs 1110 ordered by increasing EF in the two rainwater samples.

1111 Figure 7: Dissolved (upper panels) and particulate (lower panels) wet deposition fluxes ( $\mu$ mol m<sup>2</sup>) for the 1112 different TMs estimated from the two rains sampled on-board, considering the standard deviation on the TMs 1113 concentrations and the spatial variability of total precipitation over the area of sampling (Rain ION in blue 1114 and Rain FAST in red).

1115 Figure 8: Sampling chronology during the ION and FAST stations for SML, SSW and ML. The blue periods 1116 correspond to rainfall in the station area (after ERA 5 reanalysis and radar imagery, see section 3.1). 1117 Samplings were performed 4 days and 2 days before and 2 h after Rain ION, and at a higher frequency at the 1118



FAST station: 57, 37 and 7.5 hours before and 4.5, 12, 24 hours after Rain FAST. SML and SSW samples 1119 could not be collected immediately before and after the rains because of bad weather conditions, and were 1120 collected 3 and 4 days before Rain ION, and 57 and 20 h before and 30 h after Rain FAST.

1121 Figure 9: Comparison between TMs wet deposition fluxes (in green) and TMs marine stock delta (before 1122 and after the rain) in the SML (in blue) and in the ML (in red) at FAST. Dissolved = upper panels and 1123 particulate = lower panels. Marine stocks increase are expressed in absolute values (Cd, Co and Pb stocks in 1124 nmol m<sup>-2</sup>, and the other TMs in  $\mu$ mol m<sup>-2</sup>) and in relative values (%). N.E.: not enhanced (increase <5%).

1125 Figure 10: Comparison of marine stocks in the ML at all the stations occupied during the PEACETIME 1126 cruise with atmospheric inputs estimated (1) from ION and FAST rains (Boxes) and (2) from an intense wet 1127 dust deposition event of 9 g m<sup>-2</sup> (blue dots). Cd, Co and Pb stocks are in nmol m<sup>-2</sup>, and the other TMs in 1128  $\mu$ mol m<sup>-2</sup>. For Mn, marine stocks are derived from surface concentrations close to Corsica coasts (Wuttig et 1129 al., 2013: samples OUT at 0, 5 and 10 m) and in the Ionian Sea (Saager et al., 1993: Bannock basin at 0, 10, 1130 and 15 m), as no measurement is available from the PEACETIME cruise. Boxes and whiskers as in Fig. 4.

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