

# Atmospheric Measurements at Mt. Tai - Part II: HONO Budget and Radical (RO<sub>x</sub> + NO<sub>3</sub>) Chemistry in the Lower Boundary Layer

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## Abstract

In the summer of 2018, a comprehensive field campaign, with measurements on HONO and related parameters, was conducted at the foot (150 m a.s.l.) and the summit of Mt. Tai (1534 m a.s.l.) in the central North China Plain (NCP). With the implementation of a 0-D box model, the HONO budget with six additional sources and its role in radical chemistry at the foot station were explored. We found that the model default source,  $\text{NO} + \text{OH}$ , could only reproduce 13% of the observed HONO, leading to a strong unknown source strength of up to 3 ppbv  $\text{h}^{-1}$ . Among the additional sources, the  $\text{NO}_2$  uptake on the ground surface dominated (~70%) night-time HONO formation, and its photo-enhanced reaction dominated (~80%) daytime HONO formation. Their contributions were sensitive to the mixing layer height (MLH) used for the parameterizations, highlighting the importance of a reasonable MLH for exploring ground-level HONO formation in 0-D models and the necessity of gradient measurements. A  $\Delta\text{HONO}/\Delta\text{NO}_x$  ratio of 0.7% for direct emissions from vehicle exhaust was inferred and a new method to quantify its contribution to the observations was proposed and discussed. Aerosol-derived sources, including the  $\text{NO}_2$  uptake on the aerosol surface and the particulate nitrate photolysis, did not lead to significant HONO formation, with their contributions lower than  $\text{NO} + \text{OH}$ .

HONO photolysis in the early morning initialized the daytime photochemistry at the foot station. It was also a substantial radical source throughout the daytime, with contributions higher than  $\text{O}_3$  photolysis to OH initiation. Moreover, we found that OH dominated the atmospheric oxidizing capacity in the daytime, while modeled  $\text{NO}_3$  appeared to be significant at night. Peaks of modeled  $\text{NO}_3$  time series and average diurnal variation reached 22 and 9 pptv, respectively.  $\text{NO}_3$  induced reactions contribute 18% of nitrate formation potential ( $\text{P}(\text{HNO}_3)$ ) and 11% of the isoprene ( $\text{C}_5\text{H}_8$ ) oxidation throughout the whole day.

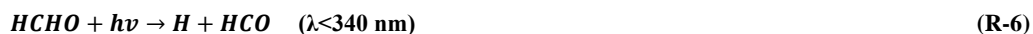
At night,  $\text{NO}_3$  chemistry led to 51% and 44% of  $\text{P}(\text{HNO}_3)$  or the  $\text{C}_5\text{H}_8$  oxidation, respectively, implying that  $\text{NO}_3$  chemistry could significantly affect night-time secondary organic and inorganic aerosol formation in this high- $\text{O}_3$  region. Considering the severe  $\text{O}_3$  pollution in the NCP and the very limited  $\text{NO}_3$  measurements, we suggest that besides direct measurements of  $\text{HO}_x$  and primary  $\text{HO}_x$  precursors ( $\text{O}_3$ , HONO, alkenes, etc.),  $\text{NO}_3$  measurements should be conducted to understand the atmospheric oxidizing capacity and air pollution formation in this and similar regions.

## 1 Introduction

Numerous field campaigns coupled with model simulations have been conducted worldwide to understand the summertime atmospheric chemistry as it is linked to the regional air quality and global climate (Alicke et al., 2003; Elshorbany et al., 2012; Heard et al., 2004; Kanaya et al., 2009, 2013; Michoud et al., 2012; Ren et al., 2003; Rohrer and Berresheim, 2006; Tan et al., 2017; Travis et al., 2020). One of the key issues is the level and the production/loss paths of the atmospheric oxidizing capacity governed by radicals (OH,  $\text{NO}_3$ , etc.). This is also essential in converting primary to secondary pollutants and the removal of greenhouse gases (Lu et al., 2019; Seinfeld and Pandis, 2016).

On a global scale, OH controls atmospheric oxidation. As the detergent in the troposphere, OH can oxidize most trace gases, including inorganic ( $\text{SO}_2$ ,  $\text{NO}_2$ , etc.) and organic compounds (VOCs, etc.), and determines the lifetime of greenhouse gases

(e.g., CH<sub>4</sub>). Besides the fast conversion of HO<sub>2</sub> to OH (R-1) as part of the radical propagation cycle, primary OH (radical initiation) mainly originates from photolysis reactions, including O<sub>3</sub> ((R-2) to (R-4)), HONO (R-5), HCHO ((R-6) to (R-9), and (R-1)), and H<sub>2</sub>O<sub>2</sub> (R-10), and the ozonolysis of alkenes (not shown in detail here). In particular, HONO photolysis is reported to be an important or even the major OH source in the lower atmosphere of polluted regions, with a contribution of 20 – 90% (Alicke et al., 2003; Elshorbany et al., 2009; Kleffmann et al., 2005; Platt et al., 1980; Slater et al., 2020; Whalley et al., 2021; Xue et al., 2020). However, this process still needs more global quantifications due to the incomplete understanding of HONO formation and its vertical distribution in the atmosphere (Kleffmann, 2007). A state-of-art summary of the reported HONO sources can be found in our recent study (Xue et al., 2020).



Besides, other oxidants can also be of importance on a regional scale. For example, the NO<sub>3</sub> radical could be a major oxidant in forests (vegetation shadows slow down its photolysis) or in the nocturnal boundary layer at high O<sub>3</sub> regions (Brown and Stutz, 2012). Formed by the reaction of NO<sub>2</sub> + O<sub>3</sub> (R-11), high NO<sub>3</sub> levels usually occur at night, concerning its very rapid photolysis during the daytime (R-12) and (R-13). Moreover, high NO<sub>3</sub> concentrations are only observed for high O<sub>3</sub> and medium NO<sub>x</sub> concentrations in the absence of significant levels of NO caused by reaction (R-15). Like OH, NO<sub>3</sub> also has high reactivity with various trace gases (Brown and Stutz, 2012; Mellouki et al., 2021). For example, NO<sub>3</sub> reacts with NO<sub>2</sub> to form N<sub>2</sub>O<sub>5</sub> (R-14), which can undergo hydrolysis on wet surfaces or clouds to produce HNO<sub>3</sub> (or NO<sub>3</sub><sup>-</sup>) (R-16) or decomposition back to NO<sub>3</sub> + NO<sub>2</sub> (R-17). NO<sub>3</sub> can also react with various organic compounds to form secondary organic aerosol (SOA). For instance, NO<sub>3</sub> reacts with isoprene (C<sub>5</sub>H<sub>8</sub>), leading to significant organic nitrates (e.g., alkyl nitrates) production (Rollins et al., 2009).





In the past decade, particle pollution, such as PM<sub>2.5</sub>, is going down while O<sub>3</sub> pollution is increasing in many cities of China (Han et al., 2020; Li et al., 2019; Sun et al., 2016, 2019). Especially in the North China Plain (NCP), air pollution becomes a major environmental risk for the public health of >330 million people living in this region. This raises efforts in exploring the NO<sub>x</sub>-VOCs-O<sub>3</sub> chemistry. Meanwhile, high O<sub>3</sub> indicates an enhanced atmospheric oxidizing capacity; that is, elevated OH and NO<sub>3</sub> levels are expected. However, very few observations have been reported of OH and NO<sub>3</sub> levels, as well as their production (e.g., HONO photolysis or NO<sub>2</sub> + O<sub>3</sub>) or loss (e.g., to oxidize primary pollutants), have been reported in the high-O<sub>3</sub> region of the NCP by far (Lu et al., 2019; Suhail et al., 2019). Herein, we provide the first HONO measurements at the foot of Mt. Tai (in Tai'an city, a typical urban site), followed by measurements at the summit of Mt. Tai in the summer of 2018. Data from the summit station is presented in the companion paper, in which daytime HONO formation and its role in the atmospheric oxidizing capacity at the summit level are studied. In this paper, coupled with the box model, the HONO budget and the radical chemistry at the ground level are explored and discussed.

## 105 2 Methods

### 2.1 Field Campaign

#### 2.1.1 Measurement Sites

In the summer (from late May to July) of 2018, a comprehensive field campaign was conducted to understand the atmospheric oxidizing capacity and O<sub>3</sub> pollution in Tai'an, a city in the middle of the NCP. Tai'an is located nearly in the middle between Beijing and Shanghai. The city has a population of about 5.6 million and is about 60 km south of Jinan city (the capital city of Shandong province, population: ~8.7 million). Measurements were conducted both at the ground level (the foot of Mt. Tai, 150 m a.s.l.) and the summit level (the summit of Mt. Tai, 1534 m a.s.l., 36.23°N, 117.11°E). The foot station was inside Shandong College of Electric Power (36.18°N, 117.11°E) in the Taishan district of Tai'an city. There are no industrial activities around this site, which is surrounded by the campus, residential area, and a business district. The 801<sup>st</sup> province road is in the northeast of this typical urban site. Inside the campus (about 50 ha) frequent traffic was not observed, but it sometimes occurred on the urban roads nearby. Mt. Tai is located in the north part of Tai'an city. Locations of these two stations on the map can be found in the companion ACP paper (entitled "Atmospheric Measurements at Mt. Tai - Part I: HONO Formation and Its Role in the Oxidizing Capacity of the Upper Boundary Layer", Xue et al. (2021a)).

Measurements at these two stations allow us to study HONO formation and its role in the atmospheric oxidizing capacity of the lower (the foot study) and the upper boundary layer (the summit study). Briefly, in the summit study, we found remarkably high daytime HONO levels as well as high unknown HONO source strength, which was mainly caused by rapid vertical transport from the ground to the summit levels driven by mountain winds.

### 2.1.2 Instrumentation

HONO mixing ratios were continuously measured by the LOPAP technique (LOPAP-03, QUMA GmbH, Germany) (Heland et al., 2001; Kleffmann et al., 2006) from 29<sup>th</sup> May to 8<sup>th</sup> July 2018 at the foot station, followed by measurements at the summit station from 9<sup>th</sup> to 31<sup>st</sup> July 2017. At the foot station, NO-NO<sub>2</sub>-NO<sub>x</sub>, O<sub>3</sub>, CO, and SO<sub>2</sub> were online measured by a series of Thermo Fisher Scientific instruments (42i, 49i, 48i, and 43i, respectively). Because chemiluminescence techniques using molybdenum converters were reported to overestimate the NO<sub>2</sub> level caused by interferences of other NO<sub>y</sub> species, the measured NO<sub>2</sub> was corrected with a family constraint in a model run (see Section 3.1.2). VOCs (56 species) and OVOCs (15 species) were measured by a homemade GC-FID instrument (Liu et al., 2016) and the USEPA DNPH-HPLC method (Wang et al., 2020), respectively. Gas-phase H<sub>2</sub>O<sub>2</sub> was measured by a monitor based on the wet chemical method (AL2021, Aerolaser GmbH, Germany), details of which can be found in Ye et al. (2018). Water-soluble inorganic particle species (i.e., NO<sub>3</sub><sup>-</sup>, SO<sub>4</sub><sup>2-</sup>, Cl<sup>-</sup>, Na<sup>+</sup>, K<sup>+</sup>, Ca<sup>2+</sup>, etc.) of PM<sub>2.5</sub> were collected on Teflon filters every two hours at a sampling flow of 100 L min<sup>-1</sup> and analyzed by an ion chromatograph (Liu et al., 2020).

Meteorological parameters, including atmospheric temperature (T), pressure (p), relative humidity (RH), wind direction (WD), wind speed (WS), and solar irradiance (Ra) were continuously measured by an auto meteorological station. J(NO<sub>2</sub>) was measured by a 4- $\pi$  filter radiometer (Metcon GmbH, Germany). 10-min average data (except for PM<sub>2.5</sub>) were used for the analysis of time series and statistic descriptions of the data. In contrast, hourly data were used for model simulations. PM<sub>2.5</sub> measurements were obtained from the Tai'an monitoring station (200 m east of the foot station), and only hourly average data were available. Other J-values used in this study, including J(HONO), J(O(<sup>1</sup>D)), J(H<sub>2</sub>O<sub>2</sub>), J(HCHO)<sub>rad</sub>, and J(HNO<sub>3</sub>), are calculated by the box model based on trigonometric SZA functions (MCM default photolysis frequency calculation, see Jenkin et al. (1997)) and scaled by the measured J(NO<sub>2</sub>). For instance,  $J(\text{HONO}) = J(\text{HONO})_{\text{model}} * J(\text{NO}_2)_{\text{measured}} / J(\text{NO}_2)_{\text{model}}$ .

## 2.2 Model Description

### 2.2.1 Box Model and Constraints

The Framework for 0-D Atmospheric Modeling, F0AM v4.0 (available at <https://github.com/AirChem/F0AM>) developed by Wolfe et al. (2016) was used to explore the HONO budget and the radical chemistry. The MCM v3.3.1 chemical mechanism was obtained from <http://mcm.leeds.ac.uk/MCMv3.3.1/home.htm>. Note that the present F0AM model could also be run with family constraints (see details in Wolfe et al. (2016)), such as the NO<sub>y</sub> and Cl<sub>y</sub> families, etc. Hence, it allows us to correct for interferences in the NO<sub>2</sub> measurements by the chemiluminescence method (see Section 3.1.2).

150 The model was constrained by the measured J(NO<sub>2</sub>), T, RH, P, VOCs, OVOCs, and all the other measured inorganic species, including the corrected NO<sub>2</sub> by the family constraint. Continuous VOCs measurements were available from 12<sup>th</sup> June to 6<sup>th</sup> July, and hence box model simulations were performed during this period. While J(NO<sub>2</sub>) measurement was available from 16<sup>th</sup> June, J(NO<sub>2</sub>) from 12<sup>th</sup> to 16<sup>th</sup> June was estimated through the high quadratic correlation ( $R^2 = 0.96$ , Figure S1) between J(NO<sub>2</sub>) and solar irradiance.

155 **2.2.2 Model Scenarios**

Table 1 shows the description of different model scenarios. A base case (Sce-0) with all the measured parameters as constraints was run to simulate radical concentrations and their production/loss rates. The family constraint was used in this scenario to correct for interferences of NO<sub>2</sub> measurements (Section 3.1.2). Meanwhile, the role of HONO in radical chemistry was also explored by several model sensitivity tests with reducing or increasing the constrained HONO.

160 With the simulated OH and the corrected NO<sub>2</sub> from the base case, we could further explore the HONO budget. Three model scenarios were conducted to assess the potential contributions of different HONO sources, including one with only the default model source (Sce-1), and one with all the six additional sources, including direct emission, dark and photo-enhanced NO<sub>2</sub> uptake on aerosol and ground surfaces and nitrate photolysis (Sce-2). In Sce-3, photo-enhanced NO<sub>2</sub> uptake on the ground surface was reduced by a factor of 10, aerosol-derived sources (NO<sub>2</sub> uptake on the aerosol surface or particulate nitrate

165 photolysis) were significantly enhanced (see Section 3.2.3) to test whether the aerosol-derived sources could well explain the observations.

**Table 1: Description of different model scenarios.**

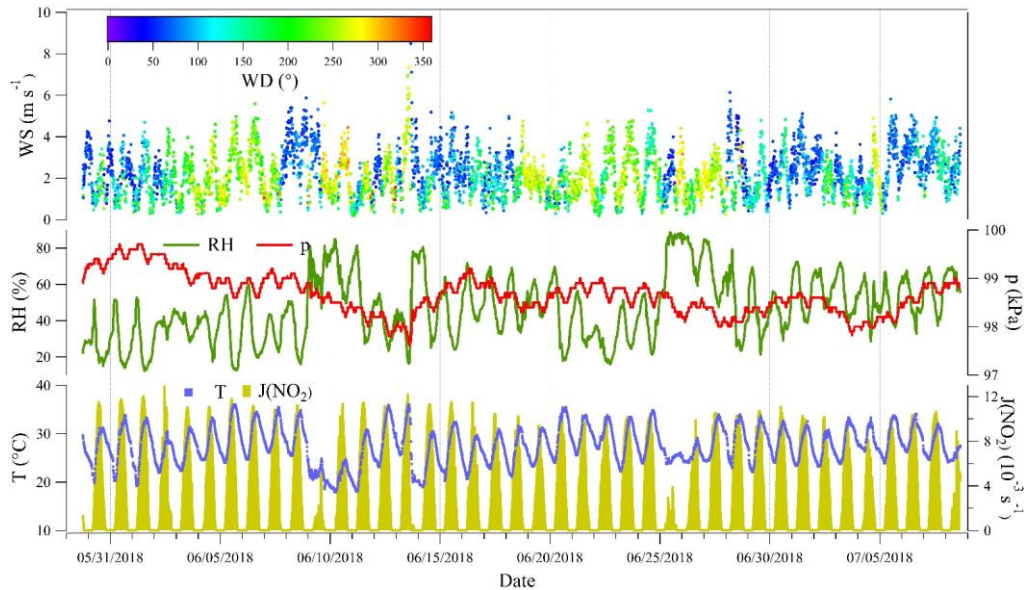
Scenarios	Constraints	Objectives
Sce-0	All measurements; NO <sub>z</sub> family constraint	NO <sub>2</sub> correction; radical concentration and chemistry
Sce-1	All measurements + corrected NO <sub>2</sub> and simulated radicals from Sce-0	HONO simulation with NO+OH
Sce-2	Same as Sce-1	HONO simulation with additional sources
Sce-3	Same as Sce-1, but with reduced ground NO <sub>2</sub> uptake and enhanced aerosol-derived sources	Testing the performance of aerosol-derived sources

3. Results and Discussion

3.1 Overview of the Measurements and Potential Interferences

170 3.1.1 Overview of the Measurements

Figure 1 shows the meteorological parameters measured at the foot station. During the campaign, it was generally sunny except slightly rainy (<10 mm) on 9<sup>th</sup>, 10<sup>th</sup>, 13<sup>th</sup>, and 28<sup>th</sup> and heavy rainy (~100 mm) at night of 25<sup>th</sup>/26<sup>th</sup> June. Ambient temperature was normally around 25 °C at night and around 30 °C during the daytime, except for rainy days when the temperature was relatively low. The relative humidity (RH) was high (up to 80%) on those rainy or cloudy days and low (around 40%) on other days. Campaign-averaged temperature and RH were 27.5 °C and 46.6%, respectively (Table 2). Air masses observed at this site originated from multiple directions, including west, south, and east, which are shown in the wind rose plot in Figure S2A. Wind speed was generally low, with an average of about 2 m s<sup>-1</sup>. In addition, the wind rose results generally agree well with HYSPLIT trajectory results in Figures S2B and S2C.



180 **Figure 1: Meteorological parameters measured at the foot of Mt. Tai during the campaign.**

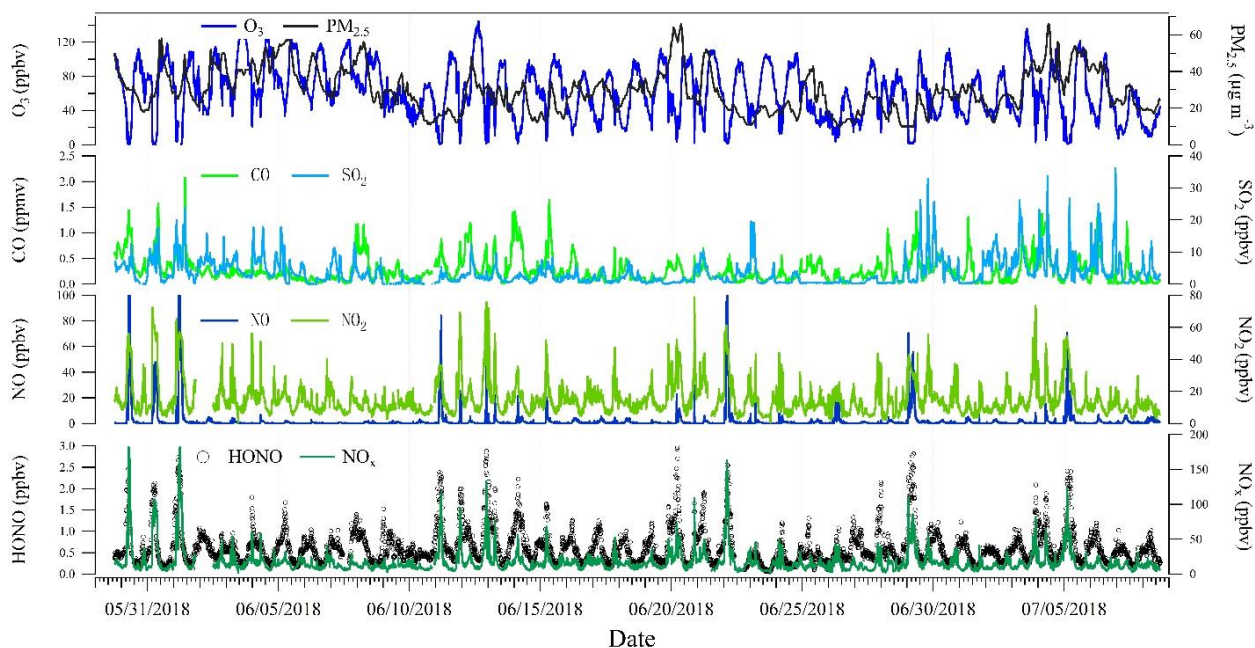
**Table 2: Statistic summary of meteorological parameters and the measured species. Note that the observation data point number (Obs) of hourly PM<sub>2.5</sub> is about 1/6 of others (10 min time resolution). The measured rather than the corrected NO<sub>2</sub> was used here. SD: standard deviation; Min: minimum; Max: maximum.**

Parameters	Mean	SD	Median	Min	Max	Obs
WS (m s <sup>-1</sup> )	2.2	1.1	2.1	0.2	9.7	5749
RH (%)	46.6	17.5	44.9	12.2	88.7	5749
P (kPa)	98.7	0.4	98.6	97.6	99.7	5749
T (°C)	27.5	3.8	27.4	17.9	36.1	5749
J(NO <sub>2</sub> ) (s <sup>-1</sup> )	3.2E-03	3.7E-03	1.0E-03	0	1.1E-02	3183

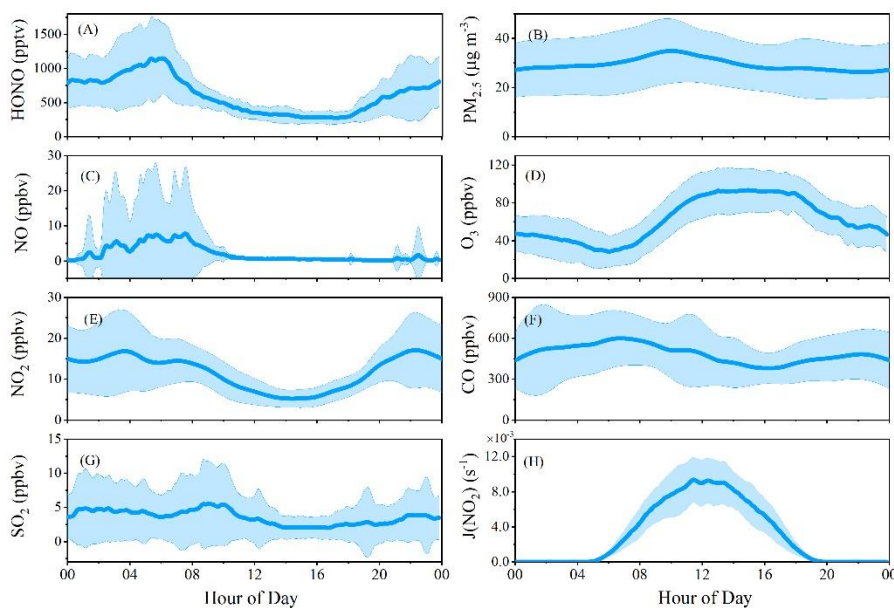
O <sub>3</sub> (ppbv)	63	31	62	0.1	145	5727
PM <sub>2.5</sub> (μg m <sup>-3</sup> )	29	12	28	10	66	959
CO (ppmv)	0.28	0.25	0.20	0.01	2.08	5717
SO <sub>2</sub> (ppbv)	3.6	4.0	2.2	0	36.2	5648
NO (ppbv)	2.0	8.3	0.3	0	126.0	5749
NO <sub>2</sub> (ppbv)	15.2	10.8	12.3	0	78.8	5601
HONO (ppbv)	0.62	0.42	0.52	0.05	2.97	5423

In Figure 2 and ~~Figure S3~~ Figure S3 the measured HONO and related species at the foot station and their diurnal variation profiles are presented, respectively. The measured HONO showed a typical diurnal variation with accumulation after sunset and decay after sunrise. Mixing ratios of the measured HONO varied from 0.05 to about 3 ppbv, with an average of  $0.62 \pm 0.42$  ppbv. The measured NO<sub>2</sub> showed a very similar variation to HONO, and their correlation was high ( $R = 0.73$ ), indicating a potential role of NO<sub>2</sub> in HONO formation. Severe O<sub>3</sub> pollution (maximum: 145 ppbv) was observed at this site, with O<sub>3</sub> levels frequently exceeding the Class-1 limit value (1-h 160 μg m<sup>-3</sup>, equivalent to 82 ppbv at 298 K and 101 kPa) of the National Ambient Air Quality Standard of China (GB3095-2012), while the NO<sub>x</sub> level was typically lower than O<sub>3</sub>. Consequently, a relatively low NO was frequently found, whose concentration was generally lower than 1 ppbv, except for some fresh plumes with higher NO concentrations inside. The two primary pollutants, CO and SO<sub>2</sub>, were generally lower than 0.5 ppmv and 5 ppbv, respectively, except for several polluted events, within which CO and SO<sub>2</sub> reached around 2 ppmv and around 35 ppbv, respectively. However, all the primary pollutants, including NO, CO, and SO<sub>2</sub>, showed poor correlations with HONO ( $R = 0.49, 0.44$ , and  $0.13$ , respectively), implying the minor role of direct emission in HONO formation. The measured hourly PM<sub>2.5</sub> varied from 10 to 66 μg m<sup>-3</sup>, with an average of 29 μg m<sup>-3</sup>. The correlation of PM<sub>2.5</sub> and HONO was very low ( $R = 0.06$ ), suggesting a minor role of aerosol-derived sources in HONO formation. More discussion on the HONO budget is presented in Section 3.2.





**Figure 2: HONO and related species measured during the campaign.**



**Figure 3: Average diurnal profiles of HONO and related species measured during the campaign. Note that some of these data were also shown in the companion ACP paper for comparison between measurements at the foot and the summit stations.**

Compared to other summertime measurements worldwide (Table 3), the measured HONO level at this site is similar to some measurements in China, such as Beijing in 2007 (Hendrick et al., 2014), Beijing in 2008 and 2017 (Crimley et al., 2019;

Hendrick et al., 2014) and Guangzhou in 2006 (Yang et al., 2014); in Europe, such as Milan in 1998 (Alicke et al., 2002) and Rome in 2001 (Acker et al., 2006); and in North America, such as New York in 2001 (Ren et al., 2003) and Colorado in 2011 (Vandenboer et al., 2013). Besides, it is lower than measurements in cities during polluted periods, such as Jinan in 2016 (Li et al., 2018), Santiago de Chile in 2005 (Elshorbany et al., 2009), Santiago de Chile in 2009 (Villena et al., 2011), and Mexico in 2003 (Volkamer et al., 2010), but higher than recent measurements near European cities, including Forschungszentrum Karlsruhe (Kleffmann et al., 2003), Forschungszentrum Jülich (Elshorbany et al., 2012), suburban Paris (Michoud et al., 2014), and Cyprus (Meusel et al., 2016). It is noteworthy that the measured HONO at the foot station is significantly higher than that observed at the summit station in the same summer, indicating possibly different roles and formation paths of HONO at these two stations.

**Table 3: Examples of worldwide HONO measurements at the ground level.**

Location	Period	Techniques	Mean (pptv)	Range (pptv)	HONO/NO <sub>x</sub> %	Reference
<b>Europe</b>						
Milan	May-Jun 1998	DOAS	920 <sup>a</sup> /140 <sup>b</sup>	<4400		(Alicke et al., 2002)
Pabstthum	Jul-Aug 1998	DOAS	330 <sup>a</sup> /70 <sup>b</sup>	<1200	1.4 <sup>a</sup> /1.0 <sup>b</sup>	(Alicke et al., 2003)
Rome	May-Jun 2001	LP-DOAS, DNP-HPLC, WEDD	580	<2000	3 <sup>*</sup>	(Acker et al., 2006)
Forschungszentrum Karlsruhe	Oct 2001	LOPAP	400	180-1100	1-6 <sup>a</sup>	(Kleffmann et al., 2003)
Forschungszentrum Jülich	Jun-Jul 2005	LOPAP	220	50-1100	2 (0.6-12)	(Elshorbany et al., 2012)
Suburban Paris	Jul 2009	Ni-troMAC	~150	10-500		(Michoud et al., 2014)
Cyprus	Jul-Aug 2014	LOPAP	35	<300	33	(Meusel et al., 2016)
<b>South America</b>						
Santiago de Chile	Mar 2005	LOPAP	2500 <sup>a</sup> /2300 <sup>b</sup>	670-7100	3.9 (1.3-9.2)	(Elshorbany et al., 2009)
Santiago de Chile	Nov 2009	LOPAP	1500/1100 <sup>**</sup>	220-3800/150-4600 <sup>**</sup>	2.0 (0.7-5.9)	(Villena et al., 2011)
<b>North America</b>						
California	Aug-Sep 1979	DOAS	1090 <sup>a*</sup> / <sup>&lt;</sup> 280 <sup>b</sup>	<4100		(Platt et al., 1980)
New York	Jul-Aug 2001	HPLC	660	400-1400		(Ren et al., 2003)
Colorado	Feb-Mar 2011	NI-PT-CIMS	500 <sup>a*</sup> /100 <sup>b</sup>	<2000	3.5-7.6 <sup>a</sup>	(Vandenboer et al., 2013)
Mexico	Mar-May 2003	LP-DOAS	1200 <sup>*</sup>	<3000		(Volkamer et al., 2010)
<b>China</b>						
Beijing	Aug-Sep 2004	DOAS		<6100	8.4 <sup>c</sup>	(Qin et al., 2006)
Guangzhou	Jun 2006	LOPAP	950 <sup>a</sup> /240 <sup>b</sup>	10-5000	4.3 <sup>a</sup> /4.5 <sup>b</sup>	(Li et al., 2012)
Yufa	Jul-Aug 2006	LOPAP	890 <sup>a</sup> /430 <sup>b</sup>	30-3600	4.6 <sup>a</sup> /4.8 <sup>b</sup>	(Yang et al., 2014)
Beijing	Aug 2007	DOAS	1450	440-2900	5 <sup>c</sup>	(Spataro et al., 2013)
Beijing	Jul 2008	DOAS	180 <sup>d</sup>	100-800	0.8 <sup>cd</sup>	(Hendrick et al., 2014)
Xianghe	Jun 2012	DOAS	90 <sup>d</sup>	100-700	1.7 <sup>cd</sup>	(Hendrick et al., 2014)
Jinan	Jun-Aug 2016	LOPAP	1200 <sup>a</sup> /1010 <sup>b</sup>	<6000	6 <sup>a</sup> /5 <sup>b</sup>	(Li et al., 2018)
Wangdu	Aug 2016	SCIC	230	30-1140	2.1 <sup>b</sup>	(Xue et al., 2021c)
Wangdu	Jun 2017	LOPAP	360 <sup>e</sup> /720 <sup>f</sup> /1360 <sup>g</sup>	260-860 <sup>e</sup> /320-1490 <sup>f</sup> /400-3130 <sup>g</sup>	2.6 <sup>e</sup> /5.4 <sup>f</sup> /11.8 <sup>g</sup>	(Xue et al., 2021c)
Beijing	May-Jun 2017	Intercomp. <sup>h</sup>		~100-10000		(Crilley et al., 2019)
Tai'an	May-Jul 2018	LOPAP	620	50-2970	4.2	This study
Mt. Tai Summit	Jul 2018	LOPAP	133	1880	6.4	This study

<sup>a</sup>: night-time, <sup>b</sup>: daytime, <sup>c</sup>: HONO/NO<sub>2</sub>, <sup>d</sup>: noontime, <sup>e</sup>: non-fertilization period, <sup>f</sup>: pre-fertilization period, <sup>g</sup>: intensive fertilization period, Intercomp<sup>h</sup>: intercomparison of 5 instruments.

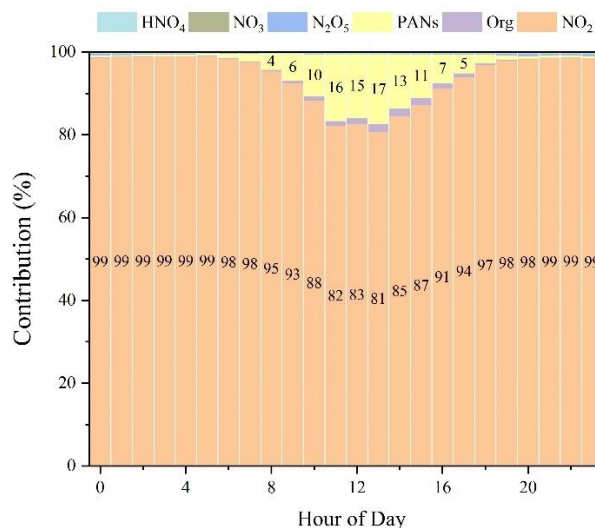
\*: half of the diurnal maximum.

\*\*: 3<sup>rd</sup> and 21<sup>st</sup> floors, respectively.

### 220 3.1.2 NO<sub>2</sub> Interference and Correction

As the most important HONO precursor, accurate measurement of NO<sub>2</sub> plays a key role in analyzing HONO formation. The NO<sub>x</sub> monitor used in this study could specifically detect NO. To measure NO<sub>2</sub>, a molybdenum converter is used to convert NO<sub>2</sub> to NO. However, this chemical conversion process suffers from the interference of other NO<sub>y</sub> species (Villena et al., 2012), primarily including inorganic species such as (measured) HONO, (non-measured) HNO<sub>3</sub>, HNO<sub>4</sub>, N<sub>2</sub>O<sub>5</sub>, and NO<sub>3</sub>,  
225 peroxyacyl nitrates (PANs, RC(O)OONO<sub>2</sub>), organic nitrates (RONO<sub>2</sub>), and peroxy nitrates (ROONO<sub>2</sub>), etc. In this study, we defined the sum of RONO<sub>2</sub> and ROONO<sub>2</sub> as organic nitrates\*. Hence, the measured NO<sub>2</sub> is the sum of real NO<sub>2</sub> and those interfering species. HONO was measured and subtracted from the measured NO<sub>2</sub>, and we defined NO<sub>2</sub>\* = the measured NO<sub>2</sub> – HONO. As NO<sub>2</sub> is the most important HONO precursor, we used the family constraint (NO<sub>2</sub>\* = NO<sub>2</sub> + HNO<sub>4</sub> + 2N<sub>2</sub>O<sub>5</sub> + NO<sub>3</sub> + PANs + organic nitrates\*) in the base case (Sce-0) to separate each species from NO<sub>2</sub>\*. To describe the PANs class,  
230 PAN, PPN, and MPAN (MCM names, see their structures at <http://mcm.leeds.ac.uk/MCMv3.3.1/>) were considered. In the organic nitrates\* class, CH<sub>3</sub>NO<sub>3</sub>, C<sub>2</sub>H<sub>5</sub>NO<sub>3</sub>, NC<sub>3</sub>H<sub>7</sub>NO<sub>3</sub>, IC<sub>3</sub>H<sub>7</sub>NO<sub>3</sub>, TC<sub>4</sub>H<sub>9</sub>NO<sub>3</sub>, NOA, ISOP<sub>34</sub>NO<sub>3</sub>, ISOPANO<sub>3</sub>, ISOPDNO<sub>3</sub>, ISOPCNO<sub>3</sub>, and ISOPBNO<sub>3</sub> (MCM names) were included. Considering that HNO<sub>3</sub> is very sticky, we expect HNO<sub>3</sub> was mostly absorbed by the filter and/or sampling tubes before the converter rather than being converted to NO by the converter. Therefore, HNO<sub>3</sub> was generally not included in the family constraint and only considered for the uncertainty analysis.  
235 Figure 4 shows the model results of the relative contribution of each NO<sub>2</sub>\* species to NO<sub>2</sub>\*. At night with the absence of photochemistry, the real NO<sub>2</sub> dominated NO<sub>2</sub>\* components, with a contribution of >95%, suggesting a small interference on the NO<sub>2</sub> measurement. However, the contribution of real NO<sub>2</sub> was found to decrease during the daytime due to the increasing interference. For example, at 11:00, the real NO<sub>2</sub> contributed 82% of the NO<sub>2</sub>\*, indicating that the interferences could be as high as +22% (calculated from 18%/82%). In particular, at 13:00, PANs alone caused the highest interference by +21%  
240 (calculated from 17%/81%).

The variations of the simulated PANs and NO<sub>3</sub> and their ratios to NO<sub>2</sub> were similar to previous observations (Brown and Stutz, 2012; Roberts et al., 1998; Su et al., 2008; Villena et al., 2012; Xue et al., 2011), indicating that the uncertainty of the method is small. For the following model simulations and analysis, only the corrected NO<sub>2</sub> was used. Besides, Figure S34 exhibits the parallel test results, in which HNO<sub>3</sub> was included in the family constraint. It can be found that the interference became more  
245 significant; for instance, the interference could be as high as +75% (calculated from 43%/57%, at 11:00). This represents the upper limit of the interferences, if the sampling tubes and the inlet filter are heated so that HNO<sub>3</sub> could have reached the converter.



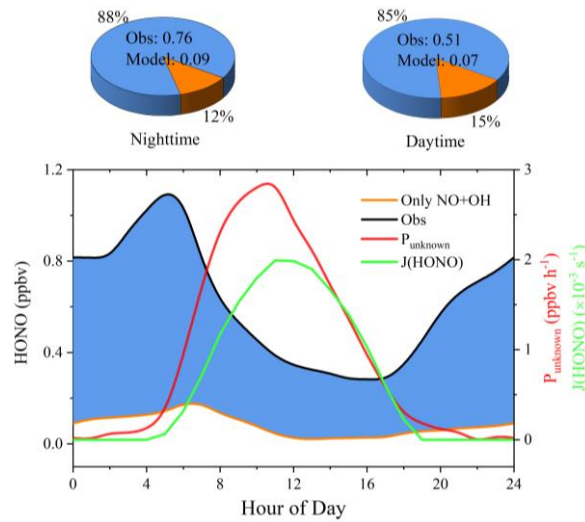
**Figure 4: Relative contribution of each  $\text{NO}_2^*$  species. PANs = PAN + PPN + MPAN. Org represents organic nitrates\* ( $\text{RONO}_2 + \text{ROONO}_2$ ).**

Additionally, as shown in Figure S45, the simulated  $\text{HNO}_4$  showed 1) a different diurnal variation from, 2) generally 1 – 2 orders of magnitude lower concentrations than, and 3) a very poor correlation ( $R^2 = 0.06$ ) with the observed HONO, indicating its negligible interference on the HONO measurement by the LOPAP technique (Legrand et al., 2014). It is worth noting that for the description of  $\text{O}_3$  formation in the polluted atmosphere, accurate measurements of  $\text{d NO}_x$  and HONO are necessary.

## 3.2 HONO Sources and Budget

### 3.2.1 Model Default Source ( $\text{NO} + \text{OH}$ ) and Unknown Source Strength

The homogeneous reaction of  $\text{NO}$  and  $\text{OH}$  has been adopted as the default HONO source in atmospheric chemistry models, including MCM. Model results from Sce-1 that only contains the homogeneous source with the modeled  $\text{OH}$  from Sce-0 are shown in Figure 5. Apparently, the source of  $\text{NO} + \text{OH}$  is too small to explain the observed HONO as the simulated one is almost one order of magnitude lower than observations. Its contributions to the measured daytime and night-time HONO are 15% and 12%, respectively.



**Figure 5: Simulated HONO by the default mechanism (Sce-1, left axis) compared with the observations (Obs, left axis), unknown source strength ( $P_{\text{unknown}}$ , right axis), HONO photolysis frequency ( $J(\text{HONO})$ , right axis). The relative contributions of NO + OH to the observations at night (19:00–4:50) and day (5:00–18:50) were shown in the left and the right pie charts, respectively. The shaded area in blue in the plot and the pie charts represents the difference between the observation and modeled values. The relative contributions of NO + OH to the observations at night (19:00–4:50) and day (5:00–18:50) are shown in the left and the right pie charts, respectively.**

Then we calculated the unknown source strength ( $P_{\text{unknown}}$ ) based on the following equation (Sörgel et al., 2011; Su et al., 2008).

$$P_{\text{unknown}} = \frac{\Delta \text{HONO}}{\Delta t} + L(\text{HONO})_{\text{pho}} + L(\text{HONO})_{\text{HONO}+\text{OH}} - P(\text{HONO})_{\text{NO}+\text{OH}} \quad (\text{Eq-1})$$

The HONO loss rates through photolysis ( $L(\text{HONO})_{\text{pho}}$ ) and reaction with OH ( $L(\text{HONO})_{\text{HONO}+\text{OH}}$ ) and production rate from NO + OH were obtained from the base model scenario (Sce-0 with a constraint of the measured HONO). HONO mixing ratio differences within a one-hour interval,  $\frac{\Delta \text{HONO}}{\Delta t}$ , were calculated by the measurements, and are compared with  $P_{\text{unknown}}$  in Figure S56. As shown in Figure 5,  $P_{\text{unknown}}$  rapidly increased in the morning and peaked at nearly 3 ppbv h<sup>-1</sup> at 11:00, followed by a decrease, revealing a photo-enhanced source. Note that the profile of  $P_{\text{unknown}}$  ~~was-is~~ asymmetric around ~~11:12:00~~ 12:00 solar noon, indicating the unknown source is not simply photolytic but also includes its precursors. For instance, higher NO<sub>2</sub> levels were observed in the morning than that in the afternoon (Figure 3E), which preliminarily implies the importance of NO<sub>2</sub>-to-HONO conversion that also have an asymmetric distribution (e.g., NO<sub>2</sub>, Figure S3). The possible additional HONO sources that are responsible for  $P_{\text{unknown}}$  are discussed in the following section.

### 3.2.2 Additional Sources vs. $P_{\text{unknown}}$

#### 3.2.2.1 Direct Emission: $\Delta \text{HONO}/\Delta \text{NO}_x$ Ratio

The  $\Delta \text{HONO}/\Delta \text{NO}_x$  ratio for the direct emission was determined from fresh plumes, which reached the following requirements:

1) at night when photolysis was absent, 2) rapid NO increase by >10 ppbv within 10 min. Only 17 cases were obtained throughout the campaign due to the persistent high O<sub>3</sub> and the fast NO-to-NO<sub>2</sub> conversion. If NO is completely titrated by

285 O<sub>3</sub>, there is a risk that the plume may not be fresh and that the inferred  $\Delta\text{HONO}/\Delta\text{NO}_x$  might be overestimated due to heterogeneous secondary HONO sources. In Table 4 the obtained  $\Delta\text{HONO}/\Delta\text{NO}_x$  was shown, varying from 0.18% to 1.86%, with an average of 0.98% and a median of 0.90%. The inferred value might be larger than the real one as potential NO<sub>2</sub>-to-HONO conversion leads to an overestimation. This is consistent with the observation that the inferred  $\Delta\text{HONO}/\Delta\text{NO}_x$  is generally higher in high RH conditions due to the potential NO<sub>2</sub>-to-HONO conversion (Figure 6). More importantly, we found  
 290 that the observed HONO/NO<sub>x</sub> is convergent as NO/NO<sub>2</sub> increases (Figure 6), which allows a further correction of  $\Delta\text{HONO}/\Delta\text{NO}_x$ . The reported NO/NO<sub>2</sub> ratios from combustion processes show significant variations, e.g., 6.7 in Wuppertal (Kurtenbach et al., 2012), ~18 in Denver (Wild et al., 2017), 5 – 30 in London (Carslaw and Beevers, 2005), and 13 – 43 from China IV/V vehicles (He et al., 2020). Furthermore, in the emission inventory, the NO/NO<sub>2</sub> emission ratio in the NCP is about 9 (Zhang et al., 2009). However, the measured night-time NO/NO<sub>2</sub> ratios were less than 3 (Figure 6), much lower compared  
 295 to direct emission measurements, indicating that the observed plumes were not fresh enough to directly obtain the primary  $\Delta\text{HONO}/\Delta\text{NO}_x$  emission ratio. By using a typical NO/NO<sub>2</sub> ratio of 10 from car exhaust, the calculated  $\Delta\text{HONO}/\Delta\text{NO}_x$  through the convergent function (Figure 6) is 0.7%, similar to that obtained from laboratory or tunnel experiments (Kirchstetter et al., 1996; Kramer et al., 2020; Kurtenbach et al., 2001; Liu et al., 2017a).

Caused by its fast daytime photolysis, direct HONO emissions (HONO<sub>emi</sub>) are likely significantly overestimated when a  
 300 constant  $\Delta\text{HONO}/\Delta\text{NO}_x$  ratio is considered, as done in former studies (Kramer et al., 2020; Liu et al., 2017b). Due to the much shorter lifetime of HONO ( $\tau(\text{HONO})$ ) compared to NO<sub>x</sub> ( $\tau(\text{NO}_x)$ ), a much smaller fraction of the emitted HONO compared to NO<sub>x</sub> will survive in the daytime atmosphere. Accordingly, for the first time, we calculated  $\tau(\text{HONO})$  and  $\tau(\text{NO}_x)$  and considered them in the parameterization of HONO<sub>emi</sub> (see the method in Section 1 of the Supporting Information). As shown in Figure S67A, daytime  $\tau(\text{NO}_x)$  was typically one order of magnitude higher than  $\tau(\text{HONO})$  (Figure S67A), indicating a  
 305 remarkable overestimation of the contribution of HONO<sub>emi</sub> to the measured HONO when using a constant  $\Delta\text{HONO}/\Delta\text{NO}_x$  emission ratio (Figure S67B). Hence, HONO<sub>emi</sub> was quantified by the following equations:

$$\text{HONO}_{emi} = 0.7\% \times [\text{NO}_x] \text{ (night-time)} \quad (\text{Eq-2})$$

$$\text{HONO}_{emi} = 0.7\% \times [\text{NO}_x] \times \frac{\tau(\text{HONO})}{\tau(\text{NO}_x)} \text{ (daytime)} \quad (\text{Eq-3})$$

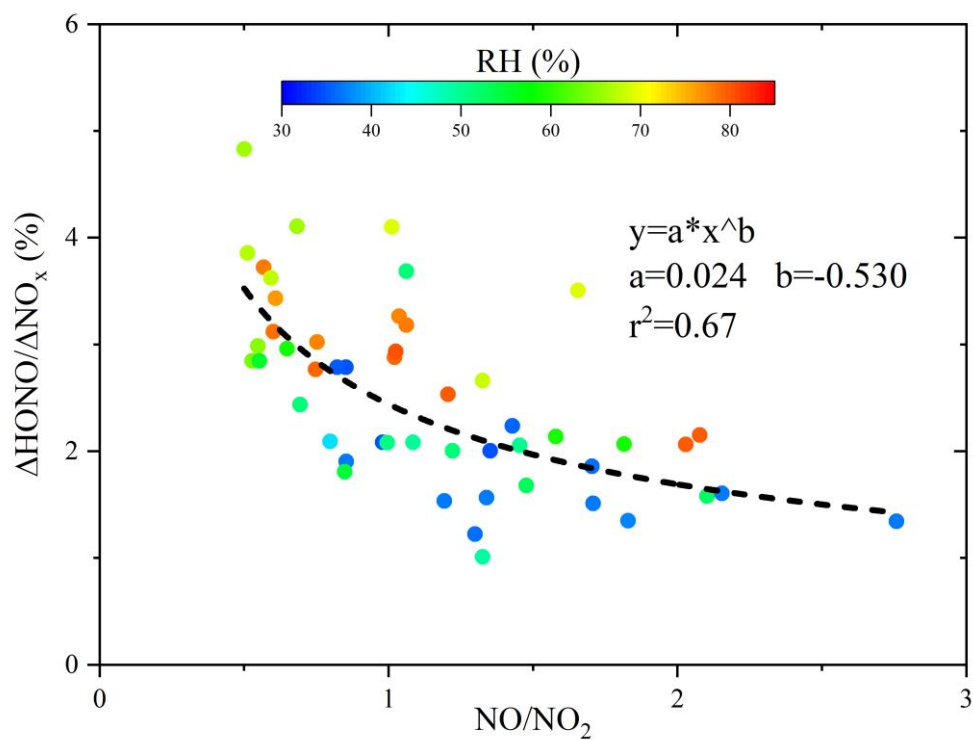
In summary, direct emission contributed about 1 – 26% of the measured HONO, with an average of 13%. Moreover, the new  
 310 method developed here may still have some uncertainties but largely reduces the significant overestimation of the contribution of HONO<sub>emi</sub> to the observations during daytime compared to using a constant  $\Delta\text{HONO}/\Delta\text{NO}_x$  (Figure S67B).

**Table 4:  $\Delta\text{HONO}/\Delta\text{NO}_x$  ratios determined from night-time (19:00 – 4:50) data of the campaign. Concentrations (ppbv) of HONO/NO/NO<sub>x</sub> at the start and the end of each plume of fresh emissions are also shown.**

Date	Period	HONO		NO		NO <sub>x</sub>		$\Delta\text{HONO}/\Delta\text{NO}_x$
1 <sup>st</sup> June	2:20-2:30	1.58	1.77	3.1	13.5	64.4	75.8	1.67%
	4:20-4:30	2.67	2.74	0.6	38.2	36.0	74.2	0.18%
11 <sup>th</sup> June	3:00-3:10	1.32	1.71	0.2	18.9	24.9	50.0	1.55%
	3:50-4:20	1.68	2.21	15.7	71.7	51.3	107.0	0.95%

	4:30-4:40	1.91	2.31	41.2	72.4	75.3	107.3	1.25%
	22:00-22:10	1.41	1.69	1.0	16.6	41.1	78.4	0.75%
	23:30-23:40	1.74	1.99	1.7	17.0	53.2	71.1	1.40%
12 <sup>th</sup> June	22:10-22:30	2.64	2.78	0.8	59.0	68.7	132.9	0.22%
14 <sup>th</sup> June	3:00-3:10	1.43	1.76	6.6	21.2	38.8	56.5	1.86%
20 <sup>th</sup> June	20:50-21:10	0.94	1.65	0.9	29.7	30.0	108.5	0.90%
22 <sup>nd</sup> June	2:10-2:20	1.29	1.98	4.6	95.1	51.8	147.2	0.72%
	2:40-3:00	1.58	2.20	38.3	120.1	83.1	163.7	0.77%
29 <sup>th</sup> June	1:00-1:20	0.96	2.26	3.1	70.5	24.3	109.4	1.53%
	2:20-2:40	0.34	0.40	7.0	21.1	45.6	59.6	0.43%
	4:20-4:30	1.79	1.97	11.7	42.2	44.0	74.0	0.60%
5 <sup>th</sup> July	1:10-1:30	0.77	1.38	0.1	18.2	33.8	70.6	1.66%
	2:10-2:30	1.09	1.26	6.4	71.0	56.2	124.6	0.25%
Mean								0.98%

315



**Figure 6: The inferred direct emission ratio ( $\Delta\text{HONO}/\Delta\text{NO}_x$ ) and  $\text{NO}/\text{NO}_2$  colored by RH. Only data with  $\text{NO}/\text{NO}_2 > 0.5$  was shown as lower ones indicate much aged rather than fresh plumes.**



### 3.2.2.2 NO<sub>2</sub> Uptake on the Aerosol Surface

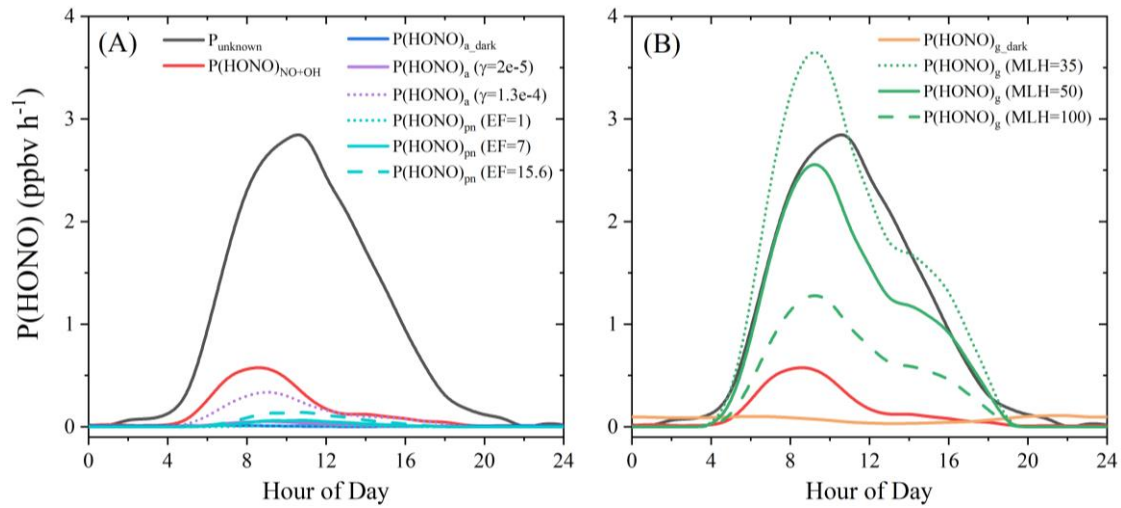
320 Parameterizations of HONO formation from the NO<sub>2</sub> uptake on the aerosol surface without and with photo-enhanced effects (P(HONO)<sub>a\_dark</sub> and P(HONO)<sub>a</sub>) are described by (Eq-4) and (Eq-5), respectively. In (Eq-4) and (Eq-5), HONO yields of 50% and 100% were considered for the dark and the photo-enhanced NO<sub>2</sub> conversion, respectively (Finlayson-Pitts et al., 2003; George et al., 2005). A relatively large NO<sub>2</sub> uptake coefficient  $\gamma_{a\_dark}$  of  $1 \times 10^{-5}$  was used here to represent its upper limit. Its overestimation should not cause significant uncertainties as P(HONO)<sub>a\_dark</sub> was negligible to HONO formation (see the  
325 following discussion). For the photo-enhanced NO<sub>2</sub> uptake coefficient  $\gamma_a$  values of  $1.3 \times 10^{-4}$  (upper limit derived from the summit measurements [in our companion paper](#), see Xue et al. (2021b)) and  $2 \times 10^{-5}$  (more realistic value derived from laboratory experiments, see Stemmler et al. (2006 and 2007)) were used in (Eq-5) to constrain the upper limit and a more realistic value of P(HONO)<sub>a</sub>.

$$P(HONO)_{a\_dark} = \frac{v(NO_2) \times S_a \times [NO_2]}{8} \times \gamma_{a\_dark}, \quad (Eq-4)$$

$$330 \quad P(HONO)_a = \frac{v(NO_2) \times S_a \times [NO_2]}{4} \times [\gamma_a \times \frac{J(NO_2)_{measured}}{0.005 s^{-1}}], \quad (Eq-5)$$

$v(NO_2)$ ,  $S_a$ ,  $[NO_2]$ , and  $J(NO_2)_{measured}$  denote the average NO<sub>2</sub> molecular speed ( $m s^{-1}$ ), aerosol surface density ( $m^{-1}$ ), NO<sub>2</sub> concentration (ppbv), and the measured NO<sub>2</sub> photolysis frequency ( $s^{-1}$ ), respectively. As aerosol size distribution measurements were not available at the foot station, we estimated  $S_a$  based on the measured PM<sub>2.5</sub> concentrations as they were highly correlated. For instance, measurements at the summit station during this campaign and other sites in the NCP found high  
335 correlations between PM<sub>2.5</sub> and  $S_a$  (derived from particle size distribution measurement) with a  $S_a/PM_{2.5}$  ratio of about  $8 \times 10^{-6} - 1.3 \times 10^{-5} m^2 \mu g^{-1}$  (Wu et al., 2008; Xue et al., 2020). Here a  $S_a/PM_{2.5}$  ratio of  $1.0 \times 10^{-5} m^2 \mu g^{-1}$  was used, and its uncertainty will not cause significant changes in the HONO simulations because of its small contribution (see the following discussion). Diurnal variations of P(HONO)<sub>a\_dark</sub> and P(HONO)<sub>a</sub>, in comparison with  $P_{unknown}$  and P(HONO)<sub>NO+OH</sub>, are shown in Figure 7A. Clearly, both P(HONO)<sub>a\_dark</sub> and P(HONO)<sub>a</sub> were negligible compared to daytime  $P_{unknown}$ . P(HONO)<sub>a</sub> increased with  $\gamma_a$ , but  
340 even when using the upper limit of  $\gamma_a = 1.3 \times 10^{-4}$ , it was still too small to be comparable to P(HONO)<sub>NO+OH</sub> and far from explaining  $P_{unknown}$ . This observation reveals minor impacts of heterogeneous HONO formation on particle surfaces to the missing HONO source, particularly during daytime.





**Figure 7: HONO production rates from different ((A): aerosol-derived, (B): ground-derived) sources and unknown HONO source strength.**

### 3.2.2.3 pNO<sub>3</sub> Photolysis

Parameterization of HONO formation from particulate nitrate photolysis ( $P(\text{HONO})_{\text{pn}}$ ) is presented in (Eq-6). Recent studies found that EF values were generally lower than 10, for instance, a value of 7 from a recent field study (Romer et al. 2018). In laboratory studies, even lower values of ~1 were observed (Laufs and Kleffmann, 2016; Shi et al., 2021; Wang et al., 2021). Hence an EF value of 7 was used in the  $P(\text{HONO})_{\text{pn}}$  calculation, and values of 1 and 15.6 were also used to test the sensitivities (the upper limit value was derived from the summit measurement, see Xue et al. (2021a)).

$$P(\text{HONO})_{\text{pn}} = p\text{NO}_3 * J(\text{HNO}_3) * EF, \quad (\text{Eq-6})$$

where  $p\text{NO}_3$  and  $J(\text{HNO}_3)$  represent the measured particulate nitrate (with unit converted from  $\mu\text{g m}^{-3}$  to ppbv) and the photolysis frequency of gas-phase  $\text{HNO}_3$  ( $\text{s}^{-1}$ ), respectively.

Diurnal variations of  $P(\text{HONO})_{\text{pn}}$  with different EF values are shown in Figure 7A. With EF varying from 1 to 7,  $P(\text{HONO})_{\text{pn}}$  was 1 – 2 orders of magnitude lower than  $P_{\text{unknown}}$ . Even using a high EF = 15.6,  $P(\text{HONO})_{\text{pn}}$  was still significantly less than half of  $P(\text{HONO})_{\text{NO+OH}}$ . Therefore, model results constrained by field measurements and recent kinetics suggested that the two aerosol-derived sources ( $\text{NO}_2$  conversion and nitrate photolysis) may not have significant impacts on daytime HONO formation, with their contributions significantly lower than half of  $P(\text{HONO})_{\text{NO+OH}}$ .

### 3.2.2.4 NO<sub>2</sub> Uptake on the Ground Surface

Parameterizations of HONO production from the  $\text{NO}_2$  uptake on the ground surface without and with photo-enhanced effects ( $P(\text{HONO})_{\text{g\_dark}}$  and  $P(\text{HONO})_{\text{g}}$ ) are demonstrated in (Eq-7) and (Eq-8), respectively.  $\text{NO}_2$  uptake coefficients of  $\gamma_{\text{g\_dark}}$  and  $\gamma_{\text{g}}$

were set to  $1.6 \times 10^{-6}$  and  $2 \times 10^{-5}$  (Han et al., 2016; Stemmler et al., 2006, 2007), respectively. The photo-enhancement effect was described by  $\frac{J(NO_2)_{measured}}{0.005 \text{ s}^{-1}}$  (Vogel et al., 2003; Wong et al., 2013; Xue et al., 2020).

$$P(HONO)_{g\_dark} = \frac{v(NO_2) \times [NO_2]}{8 \times MLH} \times \gamma_{g\_dark}, \quad (\text{Eq-7})$$

$$P(HONO)_g = \frac{v(NO_2) \times [NO_2]}{4 \times MLH} \times \gamma_g \times \frac{J(NO_2)_{measured}}{0.005 \text{ s}^{-1}}, \quad (\text{Eq-8})$$

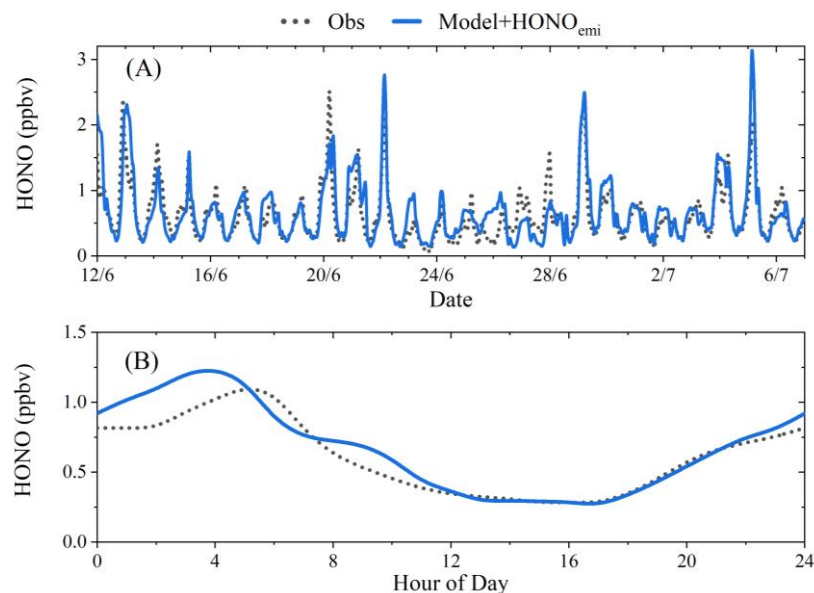
As shown above, one of the most important parameters for calculating ground HONO formation in a box model is the mixing layer height (MLH) as it is part of the denominators in both (Eq-7) and (Eq-8). MLH for HONO should be significantly lower than the boundary layer height (BLH) due to its formation on the ground level and its short lifetime, which could be confirmed by gradient measurements (Kleffmann et al., 2003; Meng et al., 2020; Vandenboer et al., 2013; Vogel et al., 2003; Wang et al., 2019; Wong et al., 2012; Xing et al., 2021; Ye et al., 2018b) (Kleffmann et al., 2003; Meng et al., 2020; Vogel et al., 2003; Wang et al., 2019; Wong et al., 2012; Xing et al., 2021; Ye et al., 2018b). For instance, a recent gradient HONO study by the MAX-DOAS technique in southwest China found a steep HONO gradient from 0 to 4 km (Xing et al., 2021). When considering their measurement at 17:00 (UTC+8) as an example, HONO levels rapidly decreased from 4.8 ppbv at the ground level (~4 m above the ground surface) to 1.6, 0.7, 0.3, 0.2, and 0.1 ppbv averaged in the layers of 0 – 100, 100 – 200, 200 – 300, 300 – 400, and 400 – 500 m above the ground level, respectively. In contrast, both  $NO_2$  and aerosol extinction remarkably increased from the ground level to about 200 m and then decreased with altitude (>200 m). These observations indicate that 1) ground-derived sources dominated daytime HONO formation; 2) the MLH for HONO was much less than 100 m, and 3) significant overestimation, i.e., by a factor of >3 in Xing et al. (2021), could be expected if using measurements on the ground surface to represent the average HONO within an MLH higher than 100 m. A similar phenomenon could also be found in tower-based vertical measurements in Germany and USA. For instance, from the ground level (4-10 m) to 100 m above the ground surface, Vogel et al. (2003) and Vandenboer et al. (2013) observed similarly declining HONO levels from ~0.6 to 0.3 (representative cases from Figure 4 in Vogel et al. (2003) and Figure 8 in Vandenboer et al. (2013)), respectively. All of those cases suggest that near-ground surface measurements were more weighted by ground-derived sources. Moreover, this phenomenon was observed during their whole campaigns including daytime and nighttime, suggesting a similar level of MLH. Hence, the maximum of MLH could be furtherly scaled to 100 m for near-ground surface measurements. A minimum MLH of 35 m was derived based on the assumption that all the  $P_{unknown}$  could be wholly explained by photosensitized heterogeneous  $NO_2$  reaction on the ground surface in our recent study (Xue et al., 2021c). Therefore, in the present study, the MLH for HONO was set at a constant height of 50 m, with sensitivity tests performed with the MLH set at 35 and 100 m. In general, ~50 m level could represent a general MLH for OD models to study HONO measurements near the ground surface. Similar values (25 – 100 m) were also used in previous box model studies (Harrison et al., 1996; Lee et al., 2016; Xue et al., 2020, 2021c) (Lee et al., 2016; Xue et al., 2020, 2021e). Nevertheless, It-it should be highlighted that a box model as used in the present study is not an ideal tool for studying a ground source when comparing with near-ground surface measurements in the atmosphere. For the future, gradient measurements are recommended, which should be compared with 1-D model simulations.

395 Diurnal variations of  $P(\text{HONO})_{\text{g\_dark}}$  and  $P(\text{HONO})_{\text{g}}$ , in comparison with  $P_{\text{unknown}}$  and  $P(\text{HONO})_{\text{NO+OH}}$ , are shown in Figure 7B.  $P(\text{HONO})_{\text{g\_dark}}$  is the largest HONO source during the night-time, while it was negligible during the daytime, which is consistent with many previous studies (Li et al., 2010; Liu et al., 2019; Vogel et al., 2003; Xue et al., 2020; Zhang et al., 2019b, 2019a, 2016). Considering the photo-enhanced formation,  $P(\text{HONO})_{\text{g}}$  shows a similar shape and a similar level to daytime  $P_{\text{unknown}}$ , indicating the potential dominance of  $P(\text{HONO})_{\text{g}}$  in the daytime HONO formation. When changing MLH to 100 (or  
400 35) m, the level of  $P(\text{HONO})_{\text{g}}$  becomes much lower (or higher) than  $P_{\text{unknown}}$ . Due to this, sensitivity tests on MLH were conducted but in Sce-2 a constant MLH of 50 m was used. Small differences in the shapes of measured and modeled results may be also caused by the variable MLH induced by variable vertical mixing in the atmosphere and the variable photolytic lifetime of HONO during the daytime.

### 3.2.3 HONO Budget

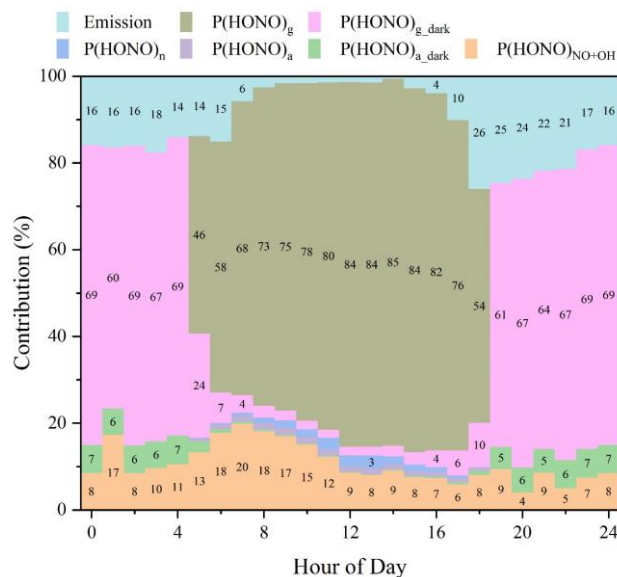
405 Along with the previous discussion, we conducted a model run (Sce-2) with all the discussed HONO sources. As shown in Figure 8, this model performed well in predicting HONO, indicating reasonable parameterizations of the HONO sources. Here, the time series of the modeled HONO were very consistent with those of the observations both for the variations and the absolute levels. ~~An~~ The only major exception was a period of only the heavy rain ~~period~~ from 25<sup>th</sup> to 28<sup>th</sup> June (see the next section). In particular, the model exhibited very high performance in predicting noontime (10:00 – 16:00) HONO as the  
410 modeled HONO was very close to the observed HONO (Figure 8B). Moreover, in Sce-3 we reduced  $\gamma_{\text{g}}$  by a factor of 10 and enlarged  $\gamma_{\text{a}}$  from  $2 \times 10^{-5}$  to  $1.2 \times 10^{-3}$  or EF from 7 to 400. While ~~that~~ the modified model could also generally predict the observed HONO levels (Figures S78A and S89A), ~~but~~ it largely failed to reproduce the noontime observations in levels and variations (Figures S78B and S89B) including its asymmetry as mentioned in Section 3.2.1. This observation reinforces our conclusion that aerosol-derived sources play only a minor role in daytime HONO formation.

415



**Figure 8: Modeled HONO mixing ratios (Model, in blue) in comparison with observations (Obs, in black). (A): time series; (B): average diurnal variations. Model+HONO<sub>emi</sub> represents the sum of the modeled HONO and HONO<sub>emi</sub>.**

Figure 9 displays the diurnal relative contributions of different HONO sources. It clearly shows that dark NO<sub>2</sub> uptake on the ground surface dominated (~70%) night-time HONO formation while photo-enhanced NO<sub>2</sub> uptake on the ground surface dominated (~80%) daytime HONO formation. P(HONO)<sub>NO+OH</sub> played a moderate role throughout the whole day, with a contribution of 5 – 15% except for a relatively larger contribution (~20%) in the early morning due to high NO levels. Direct emissions made moderate contributions of 15 – 25% at night but negligible ones during daytime. Contributions of P(HONO)<sub>a\_dark</sub>, P(HONO)<sub>a</sub>, and P(HONO)<sub>pn</sub> were always lower than 10%, and their contributions could be even smaller when using realistic kinetic parameters derived in recent studies. Therefore, aerosol-derived HONO sources may not significantly contribute to HONO formation for near-ground surface measurements (Chen et al., 2019; Neuman et al., 2016; Sarwar et al., 2008; Vogel et al., 2003; Wong et al., 2013; Xue et al., 2020; Zhang et al., 2016, 2019b).



**Figure 9: Relative contributions of different HONO sources. Note that the contribution from direct emission was calculated based on the ratio of  $\text{HONO}_{\text{emi}}$  to the observed HONO.**

### 3.2.4 Other Potential Sources

As discussed before, the model (Sce-2) could generally well reproduce most observations except for the period from 25<sup>th</sup> to 27<sup>th</sup> June. A significant overestimation occurred from midday of 25<sup>th</sup> to the morning of 26<sup>th</sup>, which was caused by the enhanced wet/dry deposition due to the heavy rain (>100 mm, Figure S249) during the night 24<sup>th</sup>/25<sup>th</sup>. In contrast, from midday of 26<sup>th</sup> to the night of 27<sup>th</sup>/28<sup>th</sup>, a significant underestimation by the model was obtained. Besides, an elevation of  $\text{HONO}/\text{NO}_x$  was found during this period (Figure S249). This might be caused by 1) the enhanced HONO emission from urban soil or 2) the enhanced  $\text{NO}_2$  uptake on the ground surface. HONO emissions from soils may occur through biological processes observed in the laboratory experiments or field measurements over the agricultural fields (Oswald et al., 2013; Scharko et al., 2015; Tang et al., 2019; Xue et al., 2019), while evidence for its occurrence on urban soil surfaces after the rain is still not uncertain.

At 13:00 on 26<sup>th</sup> and 27<sup>th</sup> June, the model predicted lower HONO by almost a factor of 2 – 4 (observation: 0.45 ppbv on both days; model: 0.13 and 0.21 ppbv), which needs an enhancement of at least 2 – 4 in  $\gamma_g$  if using  $\text{NO}_2$  uptake on the ground surface to explain the underestimation. Current laboratory studies have studied the enhancement effect of atmospheric RH (in the range of 10 – 70%) on the  $\text{NO}_2$  uptake coefficient on the humic acid surface with enhancement factors of less than 3 (Han et al., 2016; Stemmler et al., 2006, 2007). Campaign averages of the measured  $\text{NO}_2$  and RH at 13:00 were 7.4 ppbv and 35.5%, respectively. At 13:00 on 26<sup>th</sup> (27<sup>th</sup>) June, the measured  $\text{NO}_2$  of 7.9 (4.3) ppbv was similar to (or lower than) the campaign average, but the RH of 67.6% (53.1%) was significantly higher than the campaign average. But the RH was still in the range (10 – 70%) for which RH showed a small enhancement effect on  $\gamma_g$  (less than 3). Hence, after rain, an enhanced  $\text{NO}_2$  uptake may be responsible for the underestimation. In addition, soil HONO emission may co-exist, but more evidence for urban regions was needed. However, the impact of highly humid surfaces (e.g., rainwater on the ground surface) on different HONO

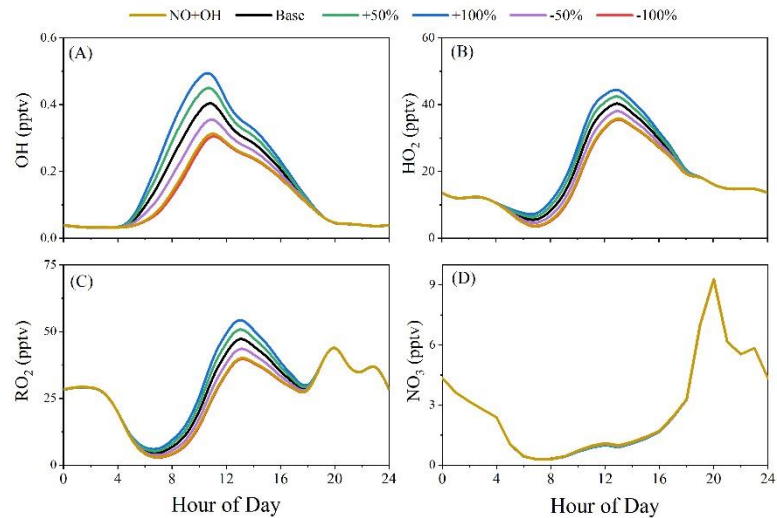
450 sources and sinks was still unclear. Further studies on the impact of rain on urban soil surface processes are necessary, such as field studies of soil HONO emission fluxes and laboratory studies of NO<sub>2</sub> uptake kinetics on relevant surfaces.

### 3.3 Radical Chemistry

Comprehensive field measurements in comparison to model studies allow studying the role of HONO in the radical chemistry of the atmosphere. HONO is expected to strongly impact OH levels in the lower atmosphere due to strong daytime HONO  
 455 sources and due to its fast photolysis. In addition, considering high O<sub>3</sub> levels at the present field site, NO<sub>3</sub> chemistry could also be important particularly during night-time, which will also be discussed in this section.

#### 3.3.1 Role of HONO in Radical Concentrations

Figure 10 shows the simulated radical concentrations in different model scenarios, in which their sensitivities to the constrained HONO were tested. It can be obtained that RO<sub>x</sub> radicals (OH, HO<sub>2</sub>, and RO<sub>2</sub>) were significantly affected by the constrained HONO, implying the vital role of HONO in the RO<sub>x</sub> budget. For instance, the peak OH concentration in the base case was  
 460 0.42 pptv (equivalent to 1.0×10<sup>7</sup> molecules cm<sup>-3</sup>). It decreased to 0.37 (or 0.32) pptv when HONO was reduced by 50% (or 100%) and increased to 0.46 (or 0.51) pptv when HONO was enlarged by 50% (or 100%). In contrast, modeled NO<sub>3</sub> concentrations showed very small variations whether HONO was reduced or enlarged. NO<sub>3</sub> concentrations are mainly governed by the levels of O<sub>3</sub> + NO<sub>2</sub> during night-time, when HONO has no impact on radical levels, due to its missing  
 465 photolysis. The almost same radical concentrations in the model cases NO + OH and case -100% indicate the minor role of NO + OH in the radical budget as this OH sink is exactly compensated by the OH production through (R-5).

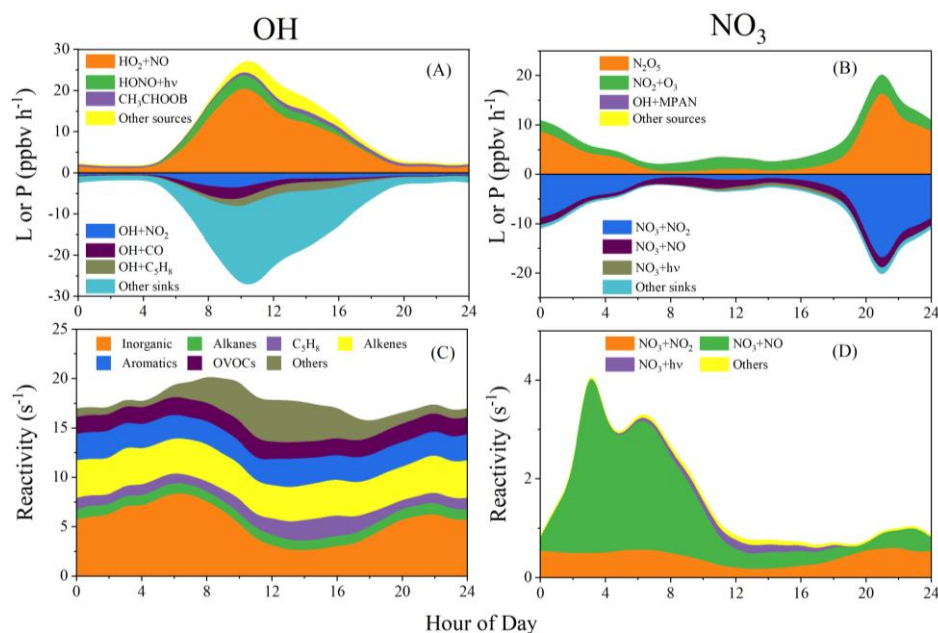


470 **Figure 10: Simulated concentrations of (A): OH, (B): HO<sub>2</sub>, (C): RO<sub>2</sub>, and (D): NO<sub>3</sub>. Different colored lines show results in different scenarios (NO + OH: only with the homogeneous source; Base: constrained by the observed HONO; +50%: constrained by the observed HONO×1.5; +100%: constrained by the observed HONO×2; -50%: constrained by the observed HONO×0.5; -100%: constrained by HONO = 0).**

### 3.3.2 Radical Production/Loss Rates and Reactivity

Figure 11A and Figure 11B illustrate the production/loss rates of OH and NO<sub>3</sub>, respectively. The total production rates of these radicals were similar to their loss rates due to their short lifetimes and high reactivities. For OH (Figure 11A), its largest source was the reaction of HO<sub>2</sub> + NO (average contribution: 70%), which is part of the propagation cycle and which is not a radical initiation source (Elshorbany et al., 2010). HONO photolysis was the second-largest OH source (average contribution: 11%), and it is expected to be the largest primary OH source after subtracting OH loss through HONO + OH and NO + OH (see Section 3.3.4). Reactions with NO<sub>2</sub>, CO, and C<sub>5</sub>H<sub>8</sub> acted as the top three OH sinks but did not dominate OH loss due to high OH reactivity caused by various other reactions, particularly those with other VOCs (see below).

Figure 11C and S104A show the OH reactivity with different classes of pollutants and their relative contributions, respectively. Total OH reactivity showed a small peak of 20 s<sup>-1</sup> in the morning and then kept almost constant around 17 s<sup>-1</sup>. Among different classes of pollutants, the measured inorganics (ordered by OH reactivity contribution: NO<sub>2</sub> > CO > NO > O<sub>3</sub> > HONO > SO<sub>2</sub> > H<sub>2</sub>O<sub>2</sub>) contributed the largest to the OH reactivity with values in the range of 2.6 – 8.4 s<sup>-1</sup>. Their total contribution was larger in the morning (43%) due to high NO, NO<sub>2</sub>, and CO levels (Figure 2) and decreased to 15% at noontime. Reactivities with the measured alkanes, alkenes, aromatics, and OVOCs were 0.95 – 1.2 s<sup>-1</sup>, 3.3 – 3.9 s<sup>-1</sup>, 2.2 – 2.9 s<sup>-1</sup>, and 1.65 – 1.9 s<sup>-1</sup>, leading to relative contributions of around 5 – 7%, 18 – 22%, 11 – 17%, and 9 – 12% throughout the whole day, respectively. Likewise, C<sub>5</sub>H<sub>8</sub> alone contributed 4% of the OH reactivity in the early morning (0.85 s<sup>-1</sup>), and its contribution increased to 12% at noontime (2.1 s<sup>-1</sup>) as a result of high levels of C<sub>5</sub>H<sub>8</sub> at noontime.



**Figure 11: Production rates (P), loss rates (L) and reactivities of radicals. (A): L and P of OH; (B): L and P of NO<sub>3</sub>; (C): Reactivities of OH and (D): Reactivities of NO<sub>3</sub>. In (A) and (B), the top-3 sources or sinks are shown, and all the others are summarized in “Other sources” or “Other sinks”. In (C), OH reactivities with different families of the measured species are shown**



and reactivities with all the unmeasured species are summarized in “Others”. In (D), NO<sub>3</sub> reactivities from top-3 reactions are shown and all the others are summarized in “Others”.

495 Figure 11D and S104B show NO<sub>3</sub> reactivity with different pollutant classes and their relative contributions, respectively. Compared with the total OH reactivity, the total NO<sub>3</sub> reactivity exhibited lower values and a different variation profile. It showed a minimum of 1 s<sup>-1</sup> at noontime and increased to around 4 s<sup>-1</sup> at 2:00. The reaction NO<sub>2</sub> + O<sub>3</sub> (R-11) is the most important net NO<sub>3</sub> source. In contrast, NO<sub>3</sub> formation by the N<sub>2</sub>O<sub>5</sub> deposition (R-17) is almost compensated by the same amount of NO<sub>3</sub> loss during N<sub>2</sub>O<sub>5</sub> production (R-14). NO<sub>3</sub> loss was dominated by its photolysis and its reactions with NO during the daytime and reactions with NO<sub>2</sub> at night. More discussion on the NO<sub>3</sub> chemistry is presented in the following section.

### 3.3.3 NO<sub>3</sub> Chemistry

As shown in Figure 10D, high NO<sub>3</sub> levels (diurnal peak: 9.3 pptv, time-series peak: 22 pptv) were simulated by the model. High NO<sub>3</sub> concentrations, as well as high NO<sub>3</sub> reactivities (Figures 10D), generally appeared at night (18:00 to 4:00 in the next day) when OH was very low and NO<sub>3</sub> was not lost by photolysis, indicating that the NO<sub>3</sub>-initialized chemistry may play an important role in night-time chemistry at this site. To verify this assumption, we compared the C<sub>5</sub>H<sub>8</sub> oxidation and nitrate formation through NO<sub>3</sub>-induced reactions with other reactions.

#### 3.3.3.1 C<sub>5</sub>H<sub>8</sub> Oxidation

Figure 12 shows (A): the C<sub>5</sub>H<sub>8</sub> loss rates (L(C<sub>5</sub>H<sub>8</sub>)) through different oxidation paths and (B): their relative contributions. L(C<sub>5</sub>H<sub>8</sub>) through O<sub>3</sub> was generally in the range of 0.04-0.12 ppbv h<sup>-1</sup>. L(C<sub>5</sub>H<sub>8</sub>) through OH showed high values in the daytime and low ones during night. On the contrary to OH, loss rates L(C<sub>5</sub>H<sub>8</sub>) through NO<sub>3</sub> were low in the daytime and high during night. On average, L(C<sub>5</sub>H<sub>8</sub>) through OH, O<sub>3</sub>, and NO<sub>3</sub> oxidation were 1.3, 0.07, and 0.16 ppbv h<sup>-1</sup>, with relative contributions of 84%, 5%, and 11%, respectively. During the daytime, L(C<sub>5</sub>H<sub>8</sub>) through OH oxidation was generally one order of magnitude higher than those through NO<sub>3</sub> or O<sub>3</sub> oxidation, leading to a dominated C<sub>5</sub>H<sub>8</sub> loss contribution of generally >90% through OH oxidation (Figure 12B). However, at night, OH was much lower and NO<sub>3</sub> was higher due to the absence of photochemistry, resulting in an increasing contribution of L(C<sub>5</sub>H<sub>8</sub>) through NO<sub>3</sub> oxidation (Figure 12B). Average L(C<sub>5</sub>H<sub>8</sub>) through night-time NO<sub>3</sub> oxidation increased to 0.30 ppbv h<sup>-1</sup>, but L(C<sub>5</sub>H<sub>8</sub>) through OH oxidation decreased to 0.33 ppbv h<sup>-1</sup>, resulting in a relatively high contribution of NO<sub>3</sub> oxidation (32 – 57%). NO<sub>3</sub> oxidation contributed to 44% of the night-time C<sub>5</sub>H<sub>8</sub> loss, which is comparable to OH oxidation (48%) and much higher than O<sub>3</sub> oxidation (8%). C<sub>5</sub>H<sub>8</sub> is an important common hemiterpene emitted by many species of vegetation and its oxidation plays a key role in secondary organic aerosol (SOA) formation. The results of the present study show that daytime OH-induced C<sub>5</sub>H<sub>8</sub> oxidation is the most important, while NO<sub>3</sub>-induced oxidation of C<sub>5</sub>H<sub>8</sub> may also significantly affect the SOA formation during the night-time (Brown and Stutz, 2012; Mellouki et al., 2021).



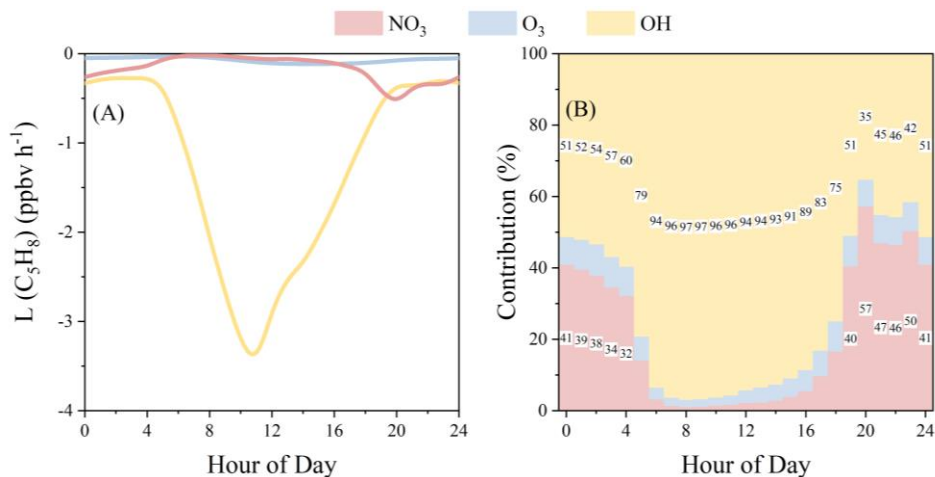


Figure 12: (A): Diurnal loss rate of C<sub>5</sub>H<sub>8</sub> by NO<sub>3</sub>, O<sub>3</sub> and OH oxidation and (B): their diurnal relative contributions.

### 3.3.3.2 HNO<sub>3</sub> Formation

As an important component of particulate matter, inorganic nitrate (pNO<sub>3</sub>) is produced by the partitioning of gas-phase HNO<sub>3</sub> and by heterogeneous uptake of N<sub>2</sub>O<sub>5</sub>. HNO<sub>3</sub> formation from OH+NO<sub>2</sub> and from N<sub>2</sub>O<sub>5</sub> hydrolysis on the aqueous aerosol surface are also the major daytime and night-time sinks for removing NO<sub>x</sub> from the atmosphere. Hence, the production of HNO<sub>3</sub>, defined as P(HNO<sub>3</sub>) = P(HNO<sub>3</sub>)<sub>OH</sub> + P(HNO<sub>3</sub>)<sub>NO<sub>3</sub></sub>, represents the upper limits of pNO<sub>3</sub> production. P(HNO<sub>3</sub>)<sub>OH</sub> denotes the HNO<sub>3</sub> production through (R-18) in the model (Sce-0). For calculation of P(HNO<sub>3</sub>)<sub>NO<sub>3</sub></sub>, both HNO<sub>3</sub> formation through N<sub>2</sub>O<sub>5</sub> heterogeneous uptake on the aerosol surface (R-16) and other NO<sub>3</sub>-induced reactions were considered (the former is the dominant one).



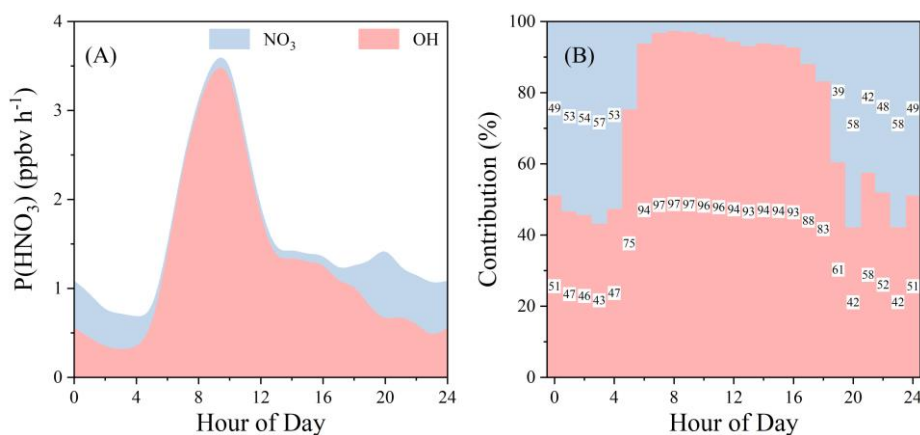
In the model, parameterization for the heterogeneous N<sub>2</sub>O<sub>5</sub> uptake is described by (Eq-9).

$$P(\text{HNO}_3)_{\text{N}_2\text{O}_5} = \frac{v(\text{N}_2\text{O}_5) \times S_a \times [\text{N}_2\text{O}_5]}{4} \times \gamma_{\text{N}_2\text{O}_5} \times 2, \quad (\text{Eq-9})$$

where  $v(\text{N}_2\text{O}_5)$ ,  $[\text{N}_2\text{O}_5]$ , and  $\gamma_{\text{N}_2\text{O}_5}$  represent the molecular speed, concentration, and heterogeneous uptake coefficient of N<sub>2</sub>O<sub>5</sub>, respectively.  $\gamma_{\text{N}_2\text{O}_5}$  was typically set as 0.1 reported in previous studies (Brown and Stutz, 2012; Wang et al., 2017).

As shown in Figure 13, the overall P(HNO<sub>3</sub>) was high during the daytime and low during the night. During the daytime, P(HNO<sub>3</sub>)<sub>NO<sub>3</sub></sub> was generally much lower than P(HNO<sub>3</sub>)<sub>OH</sub>, leading to high contributions of P(HNO<sub>3</sub>)<sub>OH</sub> (>90%). However, during the night, P(HNO<sub>3</sub>)<sub>OH</sub> and its relative contribution to P(HNO<sub>3</sub>) remarkably increased. On average throughout all day, P(HNO<sub>3</sub>)<sub>NO<sub>3</sub></sub> contributed 18%, significantly lower than P(HNO<sub>3</sub>)<sub>OH</sub> (82%). However, at night, P(HNO<sub>3</sub>)<sub>NO<sub>3</sub></sub> contribution increased to 51%, slightly higher than P(HNO<sub>3</sub>)<sub>OH</sub> (49%). Model results may have uncertainties but shed light on the atmospheric chemistry in this polluted region. By far, very few NO<sub>3</sub> measurements are available in China (Lu et al., 2019; Suhail et al., 2019), while its high concentration and important role in chemical oxidation presented in this study indicate the

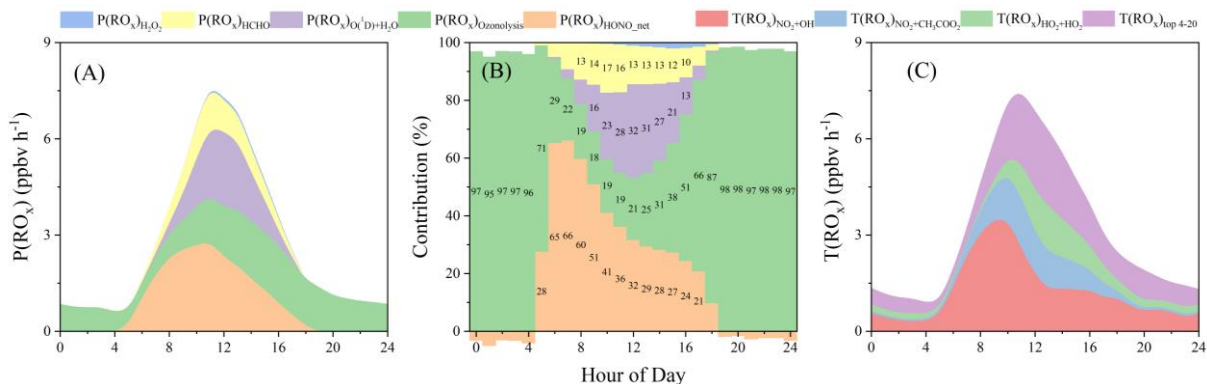
545 necessity of direct  $\text{NO}_3$  (as well as related species such as  $\text{N}_2\text{O}_5$ ,  $\text{ClNO}_2$ , etc.) measurements in the NCP, where summertime  $\text{O}_3$  levels are substantially increasing (Han et al., 2020; Li et al., 2019; Sun et al., 2016, 2019).



550 **Figure 13: (A):  $\text{HNO}_3$  production ( $\text{P}(\text{HNO}_3)$ ) from  $\text{NO}_3$ - or  $\text{OH}$ - induced reactions and (B): their relative contribution at each hour of the day.  $\text{NO}_3$ -induced reactions include heterogeneous uptake of  $\text{N}_2\text{O}_5$  on the aerosol surface and all the other  $\text{NO}_3$  reactions that produce  $\text{HNO}_3$ .**

### 3.3.4 Radical Initiation vs. Termination

Measurements of other radical precursors, such as  $\text{H}_2\text{O}_2$  (through photolysis to produce  $\text{OH}$ ),  $\text{HCHO}$  (through photolysis to produce  $\text{HO}_2$ ), and alkenes (through ozonolysis via Criegee intermediate to produce  $\text{OH}$ ,  $\text{HO}_2$ , and  $\text{RO}_2$ ), were available, which allows a comparison of radical initiation (primary production) and termination ( $\text{T}(\text{RO}_x)$ ). As shown in Figure 14, the overall radical initiation and termination showed similar variations and levels. Both, the sum of the  $\text{RO}_x$  initiation and termination showed peaks of about  $7.6 \text{ ppbv h}^{-1}$  at noon, which are in the range of  $2.5\text{--}12.2 \text{ ppbv h}^{-1}$  reported in previous studies (Elshorbany et al., 2010, 2012; Hofzumahaus et al., 2009; Kukui et al., 2014; Liu et al., 2012; Ren et al., 2003). During the daytime, it is evident that HONO photolysis ( $\text{P}(\text{RO}_x)_{\text{HONO}_{\text{net}}}$ ) made the largest contribution (20 – 70%, Figure 14B) to  $\text{RO}_x$  initiation, with an average of 37% (or 32% for all-day, Figure S112), followed by ozonolysis (29%),  $\text{O}_3$  photolysis (21%),  $\text{HCHO}$  photolysis (13%), and  $\text{H}_2\text{O}_2$  photolysis (1%). In particular,  $\text{RO}_x$  production from the ozonolysis of alkenes was significantly lower than that from HONO during 6:00 – 14:00 until later after 17:00 when it started to dominate  $\text{RO}_x$  production. At night with the absence of photochemistry, ozonolysis was the major source for primary  $\text{RO}_x$ . Due to the also high production of radicals during the daytime, ozonolysis played an important role in primary  $\text{RO}_x$  production (39% for all day). The termination of radicals  $\text{T}(\text{RO}_x)$  was dominated by  $\text{NO}_2 + \text{OH} \rightarrow \text{HNO}_3$ ,  $\text{NO}_2 + \text{CH}_3\text{COO}_2 \rightarrow \text{PAN}$ , and  $\text{HO}_2 + \text{HO}_2 \rightarrow \text{H}_2\text{O}_2$  (Elshorbany et al., 2010, 2012; Hofzumahaus et al., 2009; Kukui et al., 2014; Liu et al., 2012; Stone et al., 2012).



**Figure 14: Primary RO<sub>x</sub> production and net RO<sub>x</sub> loss. (A): Production from different sources and (B): their relative contributions at different hours of the day. (C): the top-20 RO<sub>x</sub> loss rates. Note that the top-20 net radical loss paths were summarized here. It could represent the majority of total T(RO<sub>x</sub>) as others (<0.03 ppbv h<sup>-1</sup>) were at least 2 orders of magnitude lower than the sum of top-20. Night-time P(RO<sub>x</sub>)<sub>HONO\_net</sub> was negative (a net sink for OH) so that its contribution was also negative at night. The same amounts of radical loss or production from equilibrium reactions (e.g., HO<sub>2</sub> + NO<sub>2</sub> ↔ HNO<sub>4</sub>; CH<sub>3</sub>COO<sub>2</sub> + NO<sub>2</sub> ↔ PAN) was excluded from radical initiation or termination. T(RO<sub>x</sub>)<sub>NO<sub>2</sub>+CH<sub>3</sub>COO<sub>2</sub></sub> represents the net PAN formation.**

### 3.4 Implications for Further Studies

For the first time, we considered HONO and NO<sub>x</sub> lifetimes to quantify the contribution of direct emission to daytime HONO formation. The method developed here remarkably reduced the overestimation of contribution from direct emission. It is universal and should be used for all ground measurements to quantify the contribution of direct emission to daytime HONO formation.

In the present study, we also conclude that heterogeneous NO<sub>2</sub> reaction on the ground surfaces is the major HONO source. Constraints on aerosol-derived sources, including NO<sub>2</sub> uptake on the aerosol surfaces and particulate nitrate photolysis, are conducted by our measurements at the summit of Mt. Tai and recent laboratory studies. Therefore, it could be expected that similar conclusions can be found in other studies when considering ground measurements. Additionally, parametrizations of HONO sources used for box model simulations are applicable for other studies. The values of some parameters, such as NO<sub>2</sub> uptake coefficients, MLH and particulate nitrate photolysis frequency, are obtained or derived from laboratory or field studies. Further studies may improve the understanding of the variation of those parameters, for instance, NO<sub>2</sub> uptake coefficients may vary with locations that have different landscapes. In particular, based on three vertical measurements in Germany, the USA, and China, similar levels of MLH (<100 m) were derived, indicating the potential application of this method to ground measurements worldwide. This could significantly reduce the underestimation of HONO formation from ground-derived sources compared to models in which the BLH was used. Meanwhile, sensitivity tests should be conducted and uncertainties should be discussed accordingly.

It has been recognized that HONO photolysis could initiate daytime atmospheric chemistry in the early morning and also acts as a substantial OH source during the daytime in polluted regions. The significant contribution and impact of HONO on radical levels could motivate studies by chemistry-transport models, most of which have currently not included HONO chemistry.

595 Furthermore, O<sub>3</sub> pollution is becoming a key environmental issue in China. While high levels of O<sub>3</sub> are present, moderate levels of NO<sub>x</sub> are frequently accompanied. In this case, NO<sub>3</sub> chemistry could make a considerable contribution to atmospheric oxidization capacity, especially during the nighttime. Its follow-up impacts on atmospheric composition, such as the formation of nitrate and SOA, need further field measurements and model quantifications.

#### 4. Summary

600 Atmospheric HONO and related parameters (VOCs, NO<sub>x</sub>, PM<sub>2.5</sub>, J(NO<sub>2</sub>), etc.) were measured at the foot and the summit of Mt. Tai in the summer of 2018. The present study was conducted mainly based on measurements at the foot station. The observed HONO varied from 0.05 to about 3 ppbv, with an average of  $0.62 \pm 0.42$  ppbv. With the implementation of a 0-D box model (F0AM) coupled with the Master Chemical Mechanism (MCM v3.3.1), the HONO budget and the radical (RO<sub>x</sub> + NO<sub>3</sub>) chemistry were explored.

605 The main conclusions are summarized as follows:

1. The default HONO source, NO + OH, significantly underestimated the observed HONO concentrations. This reaction could only account for 13% of the observed HONO, revealing a strong unknown source (P<sub>unknown</sub>). The diurnal profile of P<sub>unknown</sub> rapidly increased in the morning and peaked nearly 3 ppbv h<sup>-1</sup> at ca. 1 hour before noon, suggesting additional photo-enhanced HONO formation processes.
- 610 2. A HONO/NO<sub>x</sub> ratio of 0.7% was derived for direct emission, and its contribution (15 – 25% at night but negligible during the daytime) was quantified by a new method developed in this study. Based on the constraints on the aerosol-derived HONO sources (NO<sub>2</sub> uptake on the aerosol surface and nitrate photolysis) obtained from the summit measurement (see the companion paper) and from recent laboratory studies, we found that the aerosol-derived HONO sources may not significantly contribute to HONO formation at the ground level. Their contributions to HONO  
615 formation were always smaller than NO + OH. Heterogeneous NO<sub>2</sub> conversion on the ground surface made the largest contribution to P<sub>unknown</sub>, but it was sensitive to the MLH used for its parameterization. This addressed the importance of a reasonable MLH for exploring ground-level HONO formation in 0-D models and the necessity of vertical measurements.
- 620 3. HONO played an important role in RO<sub>x</sub> but a negligible role in NO<sub>3</sub> concentrations. OH dominated the atmospheric oxidizing capacity in the daytime, while NO<sub>3</sub> appeared to be significant at night. Peaks of NO<sub>3</sub> time series and diurnal variation reached 22 and 9 pptv, respectively. NO<sub>3</sub> induced reactions contribute 18% of nitrate formation potential and 11% of the C<sub>5</sub>H<sub>8</sub> oxidation throughout the whole day. While at night, NO<sub>3</sub> chemistry led to 51% and 44% of the nitrate formation potential and the C<sub>5</sub>H<sub>8</sub> oxidation, respectively. NO<sub>3</sub> chemistry may significantly affect night-time secondary organic and inorganic aerosol formation in this high O<sub>3</sub> region. Hence, the direct measurement of NO<sub>3</sub>  
625 (along with HO<sub>x</sub>, N<sub>2</sub>O<sub>5</sub>, ClNO<sub>2</sub>, etc.) in this region should be conducted.

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**Data Availability:** All the data used in this study is available upon request from the corresponding authors.

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