

# Quantification of CH<sub>4</sub> emissions from waste disposal sites near the city of Madrid using ground- and space-based observations of COCCON, TROPOMI and IASI

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**Abstract.** The objective of this study is to derive methane (CH<sub>4</sub>) emissions [from three landfills, which are the most significant CH<sub>4</sub> sources in the metropolitan city Madrid, Spain.](#) ~~of the metropolitan city Madrid Spain from the~~ [The CH<sub>4</sub> emissions are derived based on the CH<sub>4</sub> enhancements observed](#) ~~seen~~ by the space-borne and the ground-based instruments. To this end, we apply satellite-based measurements from the Tropospheric Monitoring Instrument (TROPOMI) and the Infrared Atmospheric Sounding Interferometer (IASI) together with measurements from the ground-based Collaborative Carbon Column Observing Network (COCCON) instruments.

In 2018, a two-week field campaign for measuring the atmospheric concentrations of greenhouse gases was performed in Madrid in the framework of Monitoring greenhouse Gas Emissions of Madrid city (MEGEI-MAD) project. Five COCCON instruments were deployed at different locations around the Madrid city center enabling the observation of total column averaged CH<sub>4</sub> mixing ratios (XCH<sub>4</sub>). Using available wind data, the differences between CH<sub>4</sub> columns observed at these locations allow to estimate the emissions emerging from the surrounded area. In addition, based on the dominating wind direction in the Madrid region, we calculate the difference of the satellite data maps for two opposite wind regimes (northeast – southwest, NE – SW). In the following, we refer to the resultant signal as the wind-assigned anomaly. We use TROPOMI tropospheric nitrogen dioxide (NO<sub>2</sub>) observations as a test to verify our method of wind-assigned anomaly and its implementation, taking advantage of the much better detectability of the plume due to the short lifetime and low background

35 concentrations of NO<sub>2</sub>. Pronounced bipolar plumes are found along NE and SW wind direction, which implies that our method of wind-assigned anomaly is working as expected. The wind-assigned TROPOMI XCH<sub>4</sub> anomaly shows much weaker symmetric plumes than the NO<sub>2</sub> due to the long lifetime of CH<sub>4</sub> and in consequence a high accumulated background of CH<sub>4</sub> in the atmosphere. [For the CH<sub>4</sub> emission estimation, the wind-assigned plume method is applied to the lower tropospheric CH<sub>4</sub>/dry air column ratio \(TXCH<sub>4</sub>\) of the combined TROPOMI+IASI product. TXCH<sub>4</sub> is derived from a-posteriori merged](#)

40 [IASI profile and the TROPOMI total column data.](#) ~~The wind-assigned plume method is also applied to the tropospheric and upper tropospheric/stratospheric column-averaged CH<sub>4</sub>-mixing-ratio products (in the following referred to as TXCH<sub>4</sub> and UTSXCH<sub>4</sub>) derived from a-posteriori merged Infrared Atmospheric Sounding Interferometer (IASI) profile and TROPOMI total column data.~~

Based on the NE and SW wind fields, we developed a simple plume model locating the source at three waste disposal sites ~~east of Madrid for CH<sub>4</sub>.~~ As CH<sub>4</sub> emission strength we estimate  $7.4 \times 10^{25} \pm 6.4 \times 10^{24}$  molec s<sup>-1</sup> from the TROPOMI XCH<sub>4</sub> data and  $7.1 \times 10^{25} \pm 1.0 \times 10^{25}$  molec s<sup>-1</sup> from the TROPOMI&IASI merged TXCH<sub>4</sub> data. ~~The COCCON observations indicate a weaker CH<sub>4</sub> emission strength of around  $3.7 \times 10^{25}$  molec s<sup>-1</sup> from local source (near to the Valdemingómez waste plant) in accordance with observations in a single day and.~~ [COCCON observations are used to estimate the local source as an independent method. The COCCON observations indicate a weaker CH<sub>4</sub> emission strength of around  \$3.7 \times 10^{25}\$  molec s<sup>-1</sup> from](#)

50 [local source \(near to the Valdemingómez waste plant\) in accordance with observations in a single day. That this strength is lower than the one derived from the satellite observations is a plausible result, because the analysis of the satellite data refers to a larger area, covering further emission sources in the study region, whereas the signal observed by COCCON is generated by a nearby local source.](#) All emission rates estimated from the different observations are significantly larger than the emission rates provided via the official Spanish Register of Emissions and Pollutant Sources.

## 55 1 Introduction

Methane (CH<sub>4</sub>) is the second most important anthropogenic greenhouse gas (GHG) after carbon dioxide (CO<sub>2</sub>) and contributes about 23.4% to the radiative forcing by long-lived GHGs in the atmosphere (Etminan et al., 2016). The amount of atmospheric CH<sub>4</sub> increased to 1880 ppb in 2019, corresponding to 260% of the pre-industrial level (World Meteorological Organization, 2020). The global atmospheric CH<sub>4</sub> sources are approximately 40% of natural sources (e.g. wetlands and termites) and about

60 60% of anthropogenic sources (e.g. agriculture and fossil fuels, Saunois et al., 2020). CH<sub>4</sub> is primarily removed through the reaction with the hydroxyl radicals (OH), mostly in the troposphere, which accounts for about 90% of the global CH<sub>4</sub> sink (Kirschke et al., 2013). Consequently, small changes in OH can lead to considerable variability in CH<sub>4</sub> amounts (Dlugokencky et al., 2011). It is therefore important to increase our knowledge on how the different sources and sinks affect the CH<sub>4</sub> amount in the atmosphere. Understanding the sources and sinks of CH<sub>4</sub> is also of importance for future climate emission scenarios.

65 However, CH<sub>4</sub> sources and sinks are still not fully understood. Although a lot of research studies had their focus on the global OH sink of CH<sub>4</sub> and the sum of CH<sub>4</sub> sources, which are relatively well known, large uncertainties still remain in each of the

individual CH<sub>4</sub> sources (De Wachter et al., 2017). ~~The CH<sub>4</sub> emitted from wetlands and other inland waters are the most important uncertainty components of the CH<sub>4</sub> budget (Saunois et al., 2016).~~

Satellite observations of CH<sub>4</sub> started with the launch of the Interferometric Monitor for Greenhouse gases (IMG) aboard the ADEOS satellite in August 1996 (Clerbaux et al., 2003). From 2002-2012 the SCIAMACHY (Scanning Imaging Absorption Spectrometer for Atmospheric Cartography) on board the European Envisat satellite performed measurements of total column CH<sub>4</sub> (Frankenberg et al., 2006). The Greenhouse Gases Observing Satellite (GOSAT) launched in 2009 is the first satellite dedicated to the monitoring of atmospheric GHGs and is still in operation (Kurze et al., 2009). Atmospheric CH<sub>4</sub> measurements from satellite instruments have been used to study CH<sub>4</sub> hotspot emission (e.g. the anomalous CH<sub>4</sub> emission source regions (Kort et al., 2014), anthropogenic emissions (Marais et al., 2014), and tropical wetlands (Lunt et al., 2019). A lot of research has also been carried out to map emission trends (Schneising et al., 2014; Maasakkers et al., 2019) and to estimate regional emissions (Monteil et al., 2013; Turner et al., 2015; Kuze et al., 2020; Tunncliffe et al., 2020). The current global GOSAT observations are of high quality but have sparse spatial and temporal coverage, limiting the capability to estimate the changes in daily emissions on small scales (Lorente et al., 2021).

Launched in October 2017, the TROPOspheric Measuring Instrument (TROPOMI) on board the Copernicus Sentinel-5 Precursor satellite provides complete daily global coverage of CH<sub>4</sub> with an unprecedented resolution of  $7 \times 7$  km<sup>2</sup>. The resolution was upgraded to  $5.5 \times 7$  km<sup>2</sup> in August 2019. Compared to previous satellite instruments, the TROPOMI instrument is therefore able to map the CH<sub>4</sub> enhancements due to emissions on fine scale and to detect large point sources (Varon et al., 2019). Borsdorff et al. (2020) also investigated the CO emissions of the metropolis Mexico City using TROPOMI observations and the study showed that TROPOMI has the potential to constrain the emission strengths on regional area. TROPOMI CH<sub>4</sub> data show an excellent agreement with the measurements from the validated GOSAT (Hu et al., 2018) and the ground-based Total Carbon Column Observing Network (TCCON) (Lorente et al., 2021).

Satellite retrievals using thermal infrared nadir spectra as observed for instance by IASI ([Infrared Atmospheric Sounding Interferometer](#)) or TES (Tropospheric Emission Spectrometer) are especially sensitive to CH<sub>4</sub> concentrations between the middle troposphere and the stratosphere (e.g. Siddans et al., 2017; García et al., 2018; De Wachter et al., 2017; Kulawik et al., 2021; Schneider et al., 2021a). The IASI sensors are currently orbiting aboard of three different Metop (Meteorological operational) satellites and offer twice daily global coverage with high horizontal resolution (ground pixel diameter at nadir is 12 km). The IASI CH<sub>4</sub> products have a particular good quality and sensitivity between the middle troposphere and the stratosphere as documented in different validation studies (e.g. Siddans et al., 2017; De Wachter et al., 2017; García et al., 2018; Schneider et al., 2021a).

TCCON measures solar absorption spectra in the near infrared region by using high-resolution Fourier Transform Infrared (FTIR) spectrometers (Bruker 125HR, Washenfelder et al., 2006), and is primarily designed to provide accurate and long-lasting time series of column-average dry-air molar fractions of GHGs and other atmospheric constituents (Wunch et al., 2011). Therefore, TCCON 125HR provides crucial validation resources for satellite greenhouse gas data, showing TROPOMI CH<sub>4</sub> to be of good quality (Hasekamp et al., 2019; Lorente et al., 2021). ~~However, there are to our knowledge no studies on using~~

~~TROPOMI together with ground-based portable FTIR spectrometer to derive CH<sub>4</sub> emission from metropolitan city centers (Pandey et al., 2019; Varon et al., 2019; Gouw et al., 2020).~~ Recently, TCCON 125HR GHG observations have been extended by the Collaborative Carbon Column Observing Network (COCCON, Frey et al., 2019), which is a research infrastructure using well-calibrated low-resolution FTIR spectrometers (EM27/SUN, Gisi et al, 2012) and a common data analysis scheme.

Due to the ruggedness of the portable devices used and simple operability, COCCON is well suited for implementing arrays of spectrometers for the quantification of local GHG sources (Hase et al., 2015; Luther et al., 2019; Vogel et al., 2019; Dietrich et al., 2021). Many works have used satellite observations to study CH<sub>4</sub> emission from nature gas/oil leakage (Pandey et al., 2019; Varon et al., 2019; Gouw et al., 2020) but however, there are to our knowledge no studies on using TROPOMI together with ground-based portable FTIR spectrometer to derive CH<sub>4</sub> emission from landfills in metropolitan city centers. The metropolitan cities are continuously growing due to population movements, industries, etc., and thus, more and more cities incorporate landfills (and other CH<sub>4</sub> potential sources) into their limits/influential areas. These landfills become one of the main CH<sub>4</sub> source in the metropolitan cities.

In this study we analyze nearly three years of TROPOMI total column-average dry-air molar fraction of CH<sub>4</sub> (XCH<sub>4</sub>) measurements together with COCCON spectrometer observations in the framework of the Monitoring greenhouse Gas Emissions of Madrid (MEGEI-MAD) project (García et al., 2019), in an attempt of quantifying the CH<sub>4</sub> emissions from major emission sources – namely three landfills in Madrid, the most important metropolitan area of Spain. In Section 2 our methodology is described, which is as follows: we calculate the difference of the satellite data maps for two opposite wind regimes (we refer to the resulting signals as wind-assigned anomalies). A simple plume model is then applied to predict the wind-assigned anomalies for a chosen position and strength of a source.~~in an attempt of quantifying the CH<sub>4</sub> emissions of the most important metropolitan area of Spain. Section 2 describes our methodology.~~ The results of our study are presented and discussed in Section 3. The conclusions of this study drawn from these results are given in Section 4.

## 2 Method

### 2.1 Ground-based and space-borne instrumentations

#### 2.1.1 COCCON XCH<sub>4</sub> data set

The Bruker EM27/SUN is a robust and portable FTIR spectrometer, operating at a medium spectral resolution of 0.5 cm<sup>-1</sup>. The EM27/SUN FTIR spectrometer has been developed by the Karlsruhe Institute of Technology (KIT) in cooperation with Bruker Optics GmbH for measuring GHG concentrations (Gisi et al., 2012; Hase et al., 2016). An InGaAs (Indium-Gallium-Arsenide) photodetector is used as the primary detector, covering a spectral range of 5500 – 11000 cm<sup>-1</sup>. A decoupling mirror reflects 40% of the incoming converging beam to an extended InGaAs photodetector element, covering the spectral range of 4000 – 5500 cm<sup>-1</sup> for simultaneous carbon monoxide (CO) observations. The recording time, for a typical measurement consisting of five forward and five backward scans, is about 58 seconds in total.

Several successful field campaigns and long-term deployments have demonstrated that the Bruker EM27/SUN FTIR spectrometer is an excellent instrument with good quality, robustness and reliability and its performance offers the potential to support TCCON (Frey et al., 2015 and 2019; Klappenbach et al., 2015; Chen et al., 2016; Butz et al., 2017; Sha et al., 2019; Jacobs et al., 2020; Tu et al., 2020a and 2020b; Dietrich et al., 2021). The Bruker EM27/SUN spectrometers have become commercially available from April 2014 onwards and currently about 70 spectrometers are operated by different working groups in Germany, France, Spain, Finland, Romania, USA, Canada, UK, India, Korea, Botswana, Japan, China, Mexico, Brazil, Australia and New Zealand. The development of the COCCON (<https://www.imk-asf.kit.edu/english/COCCON.php>) became possible by continued European Space Agency (ESA) support. COCCON intends to become a supporting infrastructure for GHG measurements based on common standards and data analysis procedures for the EM27/SUN (and spectrometers of comparable characteristics) (Frey et al., 2019).

Every COCCON instrument is checked (alignment and instrument line shape) and calibrated with respect to a co-located TCCON spectrometer and the primary Bruker EM27/SUN spectrometer unit operated permanently at KIT before deployment (Frey et al., 2019). For the purpose of COCCON data analysis procedures, a preprocessing tool (PREPROCESS) is applied to the raw interferograms for the generation of spectra and a non-linear least-squares fitting algorithm (PROFFAST) is used for the determination of the desired trace gas abundances from pre-generated spectra. These data processing and analysis tools are open source and freely available (<https://www.imk-asf.kit.edu/english/COCCON.php>). The ESA provided support for the code development in the framework of two projects (COCCON-PROCEEDS and COCCON-PROCEEDS II) and also supports in this framework the buildup of a central facility of the network. The demonstration of central facility functionalities is performed by KIT in cooperation with other European partners (e.g. the Belgian Institute for Space Aeronomy and Norwegian Institute for Air Research).

All the Bruker EM27/SUN spectrometers used in the MEGEI-MAD project were operated in accordance with COCCON requirements. The resulting  $XCH_4$  data used in this work were generated by the central facility operated by KIT for demonstrating a centralized data retrieval for the COCCON network. For these reasons, we refer to the Bruker EM27/SUN spectrometers as COCCON spectrometers in the following. The COCCON  $XCH_4$  data product is derived from the co-observed total column amounts of  $CH_4$  and oxygen ( $O_2$ ), and the assumed dry-air molar fraction of  $O_2$  (0.2095) (Wunch et al., 2015).

$$XCH_4 = \frac{column_{CH_4}}{column_{O_2}} \times 0.2095 \quad \text{Eq. 1}$$

### 2.1.2 TROPOMI $XCH_4$ data set

The TROPOMI data processing deploys the RemoTeC algorithm (Butz et al., 2009, 2011; Hasekamp and Butz, 2008) to retrieve  $XCH_4$  from TROPOMI measurements of sunlight backscattered by the Earth's surface and atmosphere in the near-infrared (NIR) and shortwave-infrared (SWIR) spectral bands (Hu et al., 2016, 2018; Hasekamp et al., 2019; Landgraf et al., 2019). This algorithm has been extensively used to derive  $CH_4$  and  $CO_2$  from GOSAT (Butz et al, 2011; Guerlet et al., 2013). The TROPOMI  $XCH_4$  is calculated from the  $CH_4$  vertical sub-columns  $x_i$  and the dry-air column obtained from the surface

pressure from European Centre for Medium-Range Weather Forecasts (ECMWF), and the altitude from the Shuttle Radar Topography Mission (SRTM) (Farr et al., 2007) digital elevation map with a resolution of 15 arcsec (Lorente et al., 2021)

165 analysis:

$$XCH_4 = \sum_{i=0}^n \frac{x_i}{column_{dryair}} \quad \text{Eq. 2}$$

This study uses the TROPOMI data set of  $XCH_4$  [from Lorente et al., \(2021\)](#) between April 30, 2018 and December 30, 2020 within the rectangular area of 39.5°N – 41.5°N and 4.5°W – 3.0°W (125 km × 220 km) over Madrid. In this study we apply a strict quality control to TROPOMI  $XCH_4$  (quality value  $q = 1.0$ ) to exclude data of questionable quality and to assure data under clear-sky and low-cloud atmospheric conditions (Lorente et al., 2021).

### 170 2.1.3 IASI $CH_4$ data and its synergetic combination with TROPOMI data

Here we use the IASI  $CH_4$  product as generated by the latest MUSICA IASI processor version (Schneider et al., 2021b). Combing these IASI profile data with the TROPOMI total column data causes strong synergies. Schneider et al. (2021a) developed an a posteriori method for such synergetic combination and documented the possibility to detect tropospheric partial column averaged dry-air molar fractions of  $CH_4$  ( $TXCH_4$ ) independently from the upper tropospheric/stratospheric dry-air molar fractions of  $CH_4$  ( $UTSXCH_4$ ). This is not possible by either the TROPOMI or IASI product individually. In this study we use a tropospheric product averaged from ground to 7 km a.s.l., and an upper tropospheric/stratospheric product averaged from 7 to 20 km a.s.l..

## 2.2 COCCON Madrid campaign

180 Madrid has almost 3.3 million inhabitants with a metropolitan area population of approximately 6.5 million. Madrid is located on the southern Meseta Central and 60 km south of the Guadarrama mountains with a considerable altitude difference across the city, ranging from 570 to 700 m a.s.l.

This work was made in the framework of the MEGEI-MAD project (García et al., 2019), which aimed to measure atmospheric concentrations of GHGs in an urban environment combing FTIR instruments and ground-level analyzers. Another objective of MEGEI-MAD was to analyze the possible use of portable COCCON instruments to shape an operational network for Madrid in the future. The MEGEI-MAD project was initiated by the Izaña Atmospheric Research Center (AEMet), in cooperation with two German research groups – the Karlsruhe Institute of Technology and the University of Heidelberg, and two Spanish research groups – the Autonomous University of Barcelona and the University of Valladolid.

190 Within MEGEI-MAD, a two-week field campaign was carried out from September 24 to October 7, 2018 in Madrid, where five COCCON instruments were located at five different places circling the metropolitan area (see Figure 1). Table 1 summaries the coordinates, altitudes of the COCCON locations and auxiliary meteorological data collected for data analysis of the observations. The locations have been chosen by considering the prevailing winds and the emission sources of  $CO_2$  and  $CH_4$ , as well as other technical and logistic criteria (García et al., 2019; García et al., 2021, in preparation).

Though every COCCON instrument is already calibrated at KIT, long-distance transportation and long-term usage under different conditions might cause some drifts of instrumental characteristics. To minimize any systematic errors due to drifts of the calibration, a pre-campaign side-by-side comparison of the five COCCON instruments used in MEGEI-MAD was carried out at the AEMet Headquarter between September 17 and September 20, 2018, to obtain updated calibration factors (Table 2). The calibration factors are computed as ratios between the observations from each instrument and the ensemble average. These XCH<sub>4</sub> calibration factors are in good agreement with previous calibration results for these instruments obtained during side-by-side comparisons at KIT (empirical standard deviation of the ratios between previous calibration and campaign results  $\approx$  0.17%). These small changes among the COCCON instruments demonstrate the excellent characteristic of stability. The set of five COCCON spectrometers used in this work is calibrated based on the inter-comparison measurements and the solar zenith angle (SZA) range is limited to 75° to reduce airmass-dependent effects. The instrument-specific calibration factors systematic differences are within  $\pm 0.06$  % for XCH<sub>4</sub>.

**Table 1: Locations of the five COCCON instruments and meteorological records for the MEGEI-MAD field campaign during September 24 – October 7, 2018.**

Station	EM27/SUN	Latitude (°N)	Longitude (°W)	Altitude (m a.s.l.)	Meteorological Records
Tres Olivos	KIT SN53	40.499	3.689	736	Datalogger from AEMet
Barajas	AEMet SN85	40.465	3.581	637	Barajas Airport
Jose Echegaray	DLR SN69	40.379	3.613	633	Datalogger from DLR
Cuatro Vientos	KIT SN52	40.368	3.780	703	Cuatro Vientos Airport
AEMeT	KIT SN81	40.452	3.724	685	AEMeT Headquarter



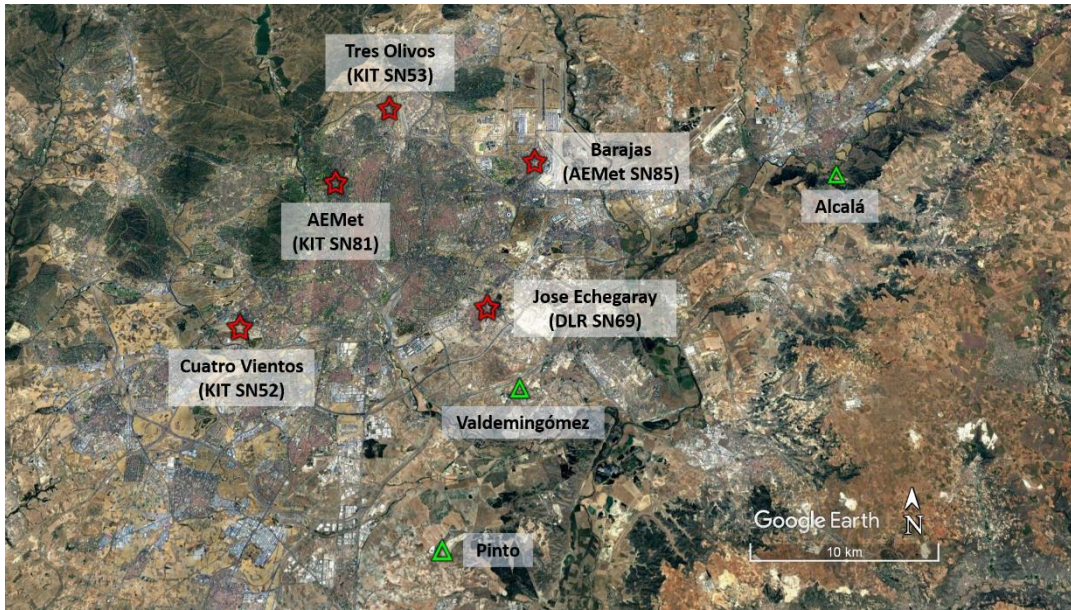


Figure 1: Locations of the five COCCON instruments used in the Madrid field campaign during September 24 – October 7, 2018, represented with red stars and locations of three waste treatment and disposal plants, represented with the green triangles.

Table 2: XCH<sub>4</sub> calibration factors for the five COCCON instruments from the pre-campaign inter-comparison measurements at the AEMet Headquarter.

	KIT SN53	AEMeT SN85	DLR SN69	KIT SN52	KIT SN81
side-by-side at AEMet	0.999212	1.000719	1.000594	1.000077	0.999398

### 2.3 Emission strength calculation using a simple plume model

The daily plume is modelled as a function of wind direction and wind speed. The schematic dispersion model for describing emissions assumes an expanding cone-shaped plume with the tip at the plume source at location (0,0). The plume cone has an opening angle of size  $\alpha$  and any grid cell within the cone is affected by the emission (see Figure 2). The angle  $\alpha$  is a technical parameter to schematically describe a spreading of the plume and is empirically adjusted to a value of 60°. Different opening angles are modelled and presented in Figure A- 1. The modelled plume has the most similar shape compared to the TROPOMI measured NO<sub>2</sub> plume (see Section 3.2) when  $\alpha \geq 60^\circ$ . If the grid cell  $(x, y)$  locates inside the cone, the column enhancement for this cell can be calculated by:

$$\Delta column_{(x,y)} = \frac{\varepsilon}{v \cdot d_{(x,y)} \cdot \alpha} \quad \text{Eq. 3}$$

where  $\varepsilon$  is the emission strength at the source point in molec s<sup>-1</sup>,  $v$  is the wind speed in m s<sup>-1</sup>,  $d$  is the distance between the downwind point and the source,  $\alpha$  is the opening angle of the plume in rad (here assumed to be 60°).

The distance from a general grid cell  $(x, y)$  from the source is:



$$d(x, y) = \sqrt{x^2 + y^2} \quad \text{Eq. 4}$$

The enhanced dry-air volume mixing ratio for target species ( $\Delta XVMR$ ) at the center of the grid cell  $(x, y)$  can then be calculated by dividing the column enhancement by the total column of dry air ( $column_{dryair}$ ):

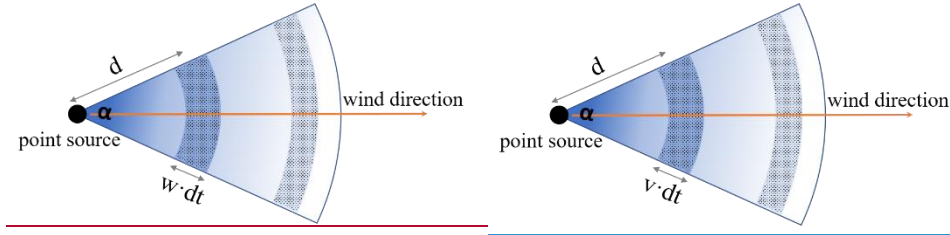
$$\Delta XVMR = \frac{\Delta column_{(x,y)}}{column_{dryair}} \quad \text{Eq. 5}$$

225 The  $column_{dryair}$  is computed from the surface pressure:

$$column_{dryair} = \frac{P_s}{m_{dryair} \cdot g(\varphi)} - \frac{m_{H_2O}}{m_{dryair}} \cdot column_{H_2O} \quad \text{Eq. 6}$$

where  $P_s$  is the surface pressure,  $m_{dryair}$  and  $m_{H_2O}$  are the molecular masses of dry air ( $\sim 28.96 \text{ g} \cdot \text{mol}^{-1}$ ) and water vapor ( $\sim 18 \text{ g} \cdot \text{mol}^{-1}$ ), respectively.  $column_{dryair}$  and  $column_{H_2O}$  are the total column amount of dry air and water vapor, and  $g(\varphi)$  is the latitude-dependent surface acceleration due to gravity.

230 In this study, each individual landfill is considered as an individual point source. The daily plumes from the individual landfills are super-positioned to have a total daily plume. The averaged enhancement of XVMR (plume) over the study area is computed for the selected wind sector. The plume for the opposite wind regime is also constructed in the same manner. The differences between these two data sets are therefore the wind-assigned anomalies (see Sect. 3.3). By fitting the modelled wind-assigned anomalies to the anomalies as observed by the satellite, we can estimate the actual emission strength (see Sect. [B.2 2-5](#)). Note that the applied calculation scheme would also be extendible to areal sources by superimposing such calculations using different locations of the origin.



240 **Figure 2: Sketch of the simple plume model used to explain the  $\text{CH}_4$  emission estimation method.** The methane at the point source is distributed along the wind direction (wind speed:  $v$ ) in the cone-shaped area with an opening angle of  $\alpha$ . The point source emits the methane at an emission rate of  $\epsilon$ . We assumed the methane molecules are evenly distributed in the dotted area A, and the distance from area A to the point source is  $d$ . Therefore, the emitted methane in  $dt$  time period equals to the amount of methane in the area A. It yields the equation  $\epsilon \times dt \approx \Delta column \times \frac{\alpha}{\pi} \times \pi \times d \times v \times dt$ .

## 2.4 $\text{CH}_4$ background signal

The satellite data can be written as a vector  $y$ , where each element corresponds to an individual satellite data point. This signal is caused by a  $\text{CH}_4$  background signal and the  $\text{CH}_4$  plume due to the emissions from the waste disposal sites near Madrid:

$$y = y_{bg} + y_{plume} \quad \text{Eq. 7}$$

245 It is of great importance to adequately separate both components for estimating the emission strength from the satellite data. For determining the background signal ( $y_{bg}$ ), we setup a background model:

$$\mathbf{M}_{BG} = \mathbf{y}_{BG} - \mathbf{K}_{BG} \mathbf{x}_{BG} \quad \text{Eq. 8}$$

250 The matrix  $\mathbf{K}_{BG}$  is a Jacobian matrix that allows to reconstruct the background according to a few background model coefficients (the elements of the vector  $\mathbf{x}_{BG}$ ). We also create a Jacobian  $\mathbf{K}_{BG}^*$ , which is the same as  $\mathbf{K}_{BG}$  but set to zero for observations where the wind data suggest a significant impact of the  $\text{CH}_4$  plume on the satellite data. The calculations of the plume  $\text{CH}_4$  signals are made according to Sect. 2.3. With the use of  $\mathbf{K}_{BG}^*$  we make sure that the estimated background signal is not affected by the  $\text{CH}_4$  plume.

255 The Jacobian matrix  $\mathbf{K}_{BG}$  considers a smooth background, which is a constant  $\text{CH}_4$  value, a linear increase with time and a seasonal cycle described by the amplitude and phase of the three frequencies 1/year, 2/year, and 3/year. Furthermore, we fit a daily anomaly, which is the same for all data measured during a single day and a horizontal anomaly, which is the same for any time but dependent on the horizontal location. For the latter we use a  $0.1^\circ \times 0.135^\circ$  (latitude  $\times$  longitude) grid.

We invert the problem in order to estimate the background model coefficients (elements of the vector  $\mathbf{x}_{BG}$ ):

$$\hat{\mathbf{x}}_{BG} = \mathbf{G}_{BG} \mathbf{y}_{BG} \quad \text{Eq. 9}$$

With  $\mathbf{G}_{BG}$  being the so-called gain matrix

$$\mathbf{G}_{BG} = (\mathbf{K}_{BG}^{*T} \mathbf{S}_{y,n}^{-1} \mathbf{K}_{BG}^* + \mathbf{S}_a^{-1})^{-1} \mathbf{K}_{BG}^{*T} \mathbf{S}_{y,n}^{-1} \quad \text{Eq. 10}$$

260 The matrix  $\mathbf{S}_{y,n}$  stands for the noise covariance of the satellite data. For constraining the problem, we use a diagonal  $\mathbf{S}_a^{-1}$  (no constraint between different coefficients) with a very low constraint value for the coefficient determining the constant and higher constraint values for the other coefficients. Here we use  $\mathbf{S}_{y,BG}$  as the diagonal matrix with the mean square value of the difference  $\mathbf{y}_{BG} - \mathbf{K}_{BG}^* \hat{\mathbf{x}}_{BG}$  being the diagonal elements. In this context,  $\mathbf{S}_{y,BG}$  considers the deficits of the background model.

The uncertainty of the background model coefficients can be calculated as:

$$\mathbf{S}_{\hat{\mathbf{x}}_{BG}} = \mathbf{G}_{BG} \mathbf{S}_{y,BG} \mathbf{G}_{BG}^T \quad \text{Eq. 11}$$

For each day there is an uncertainty in the background coefficients and the uncertainty is correlated with the uncertainty at other days. All this information is provided in the uncertainty covariance  $\mathbf{S}_{\hat{\mathbf{x}}_{BG}}$ .

265 With the full Jacobian  $\mathbf{K}_{BG}$  we can now model the background for the measurement state (also for the measurements that are assumed to be affected by the  $\text{CH}_4$  waste disposal plume):

$$\mathbf{y}_{BG} = \mathbf{K} \hat{\mathbf{x}}_{BG} \quad \text{Eq. 12}$$

and calculate the plume signal according to Eq. 7 as:

$$\mathbf{y}_{\text{plume}} = \mathbf{y} - \mathbf{K} \hat{\mathbf{x}}_{BG} \quad \text{Eq. 13}$$

The uncertainty of these plume signal is the sum of the uncertainties of the satellite data  $\mathbf{S}_{y,n}$  and the uncertainty of the estimated background:

$$\mathbf{S}_{y,\text{plume}} = \mathbf{S}_{y,n} + \mathbf{K} \mathbf{S}_{\hat{\mathbf{x}}_{BG}} \mathbf{K}^T \quad \text{Eq. 14}$$

## 270 2.5 Fitting of CH<sub>4</sub> emission rates

Because the CH<sub>4</sub> plume signal is rather weak compared to the CH<sub>4</sub> background uncertainty and the noise level of the satellite data, we have to work with averages in order to reduce the data noise. The averaging is made by classifying the observation in two predominant wind categories. We calculate the average plume maps for the southwest and northeast wind situations (see Figure 6 and Figure 8). Then we calculate the difference between the south-west and north-east plume maps (the wind assigned anomalies or  $\Delta$  maps). All the calculations are made by binning all observations that fall within a certain  $0.135^\circ \times 0.1^\circ$  (longitude  $\times$  latitude) area. In order to significantly reduce the data noise, we only consider averages for the  $0.135^\circ \times 0.1^\circ$  areas based on at least 25 individual observations made under southwest wind conditions and 25 individual observations made under northeast wind conditions. The binning, the averaging, the wind assigned  $\Delta$  maps calculations, and the data number filtering is achieved by operator  $\mathbf{D}$ , and we can write:

$$\Delta \mathbf{y}_{\text{plume}} = \mathbf{D} \mathbf{y}_{\text{plume}} \quad \text{Eq. 15}$$

280 and

$$\Delta \mathbf{S}_{\text{y,plume}} = \mathbf{D} \mathbf{S}_{\text{y,plume}} \mathbf{D}^T \quad \text{Eq. 16}$$

Here  $\Delta \mathbf{y}_{\text{plume}}$  is a column vector whose elements capture the different signal of the two wind directions at the different locations and  $\Delta \mathbf{S}_{\text{y,plume}}$  is the corresponding uncertainty covariance.

—For modelling the plume signals we use a priori knowledge of CH<sub>4</sub> emission locations, i.e. assuming a repartition of the emissions between the three waste disposal sites according to Table 3 (see Sect. 3.1). Together with information from the wind, we then model the CH<sub>4</sub> plume's wind assigned anomaly signal  $\Delta \mathbf{y}_{\text{plume}}$ :

$$\Delta \mathbf{y}_{\text{plume}} = \Delta \mathbf{k} \mathbf{x} \quad \text{Eq. 17}$$

Here the Jacobian  $\Delta \mathbf{k}$  (a column vector) represents the wind assigned anomaly model as described in Sect. 2.3. It describes how an emission at the waste disposal sites according to Table 3 would be seen in the difference signal. We are interested in the coefficient  $\mathbf{x}$  (a scalar describing how the assumed emissions from Table 3 have to be scaled by a common factor in order to achieve the best agreement with the observed plume).

290 Similar to Eq. 9 and Eq. 10 we write:

$$\hat{\mathbf{x}} = \mathbf{g}^T \Delta \mathbf{y}_{\text{plume}} \quad \text{Eq. 18}$$

with the row vector

$$\mathbf{g}^T = (\Delta \mathbf{k}^T \Delta \mathbf{S}_{\text{y,plume}}^{-1} \Delta \mathbf{k})^{-1} \Delta \mathbf{k}^T \Delta \mathbf{S}_{\text{y,plume}}^{-1} \quad \text{Eq. 19}$$

This fitting of the emission rate correctly considers the respective uncertainty of the difference signals at the different locations.

Because of the small plume signals, it is important to estimate the reliability of the fitted emission rate. The uncertainty of  $\mathbf{x}$  due to the background uncertainty and the noise in the satellite data can be estimated as:

$$\epsilon_{\text{BG}} = \sqrt{\mathbf{g}^T \mathbf{D} \mathbf{K} \mathbf{S}_{\text{x,BG}} \mathbf{K}^T \mathbf{D}^T \mathbf{g}} \quad \text{Eq. 20}$$

295 and

$$\epsilon_{\#} = \sqrt{g^T D S_{y,\#} D^T g}$$

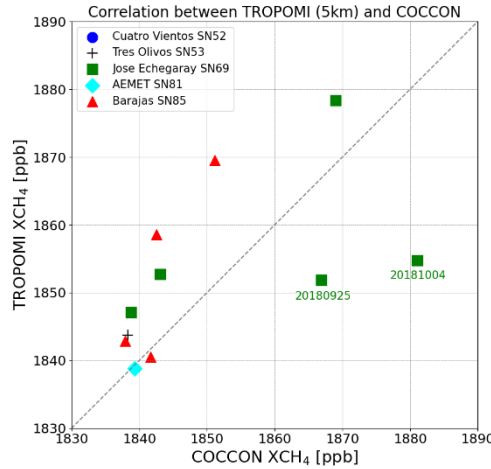
Eq. 21

respectively.

### 3 Results and discussion

#### 3.1 Intercomparison of TROPOMI and COCCON XCH<sub>4</sub> measurements. Validation of TROPOMI using COCCON measurements

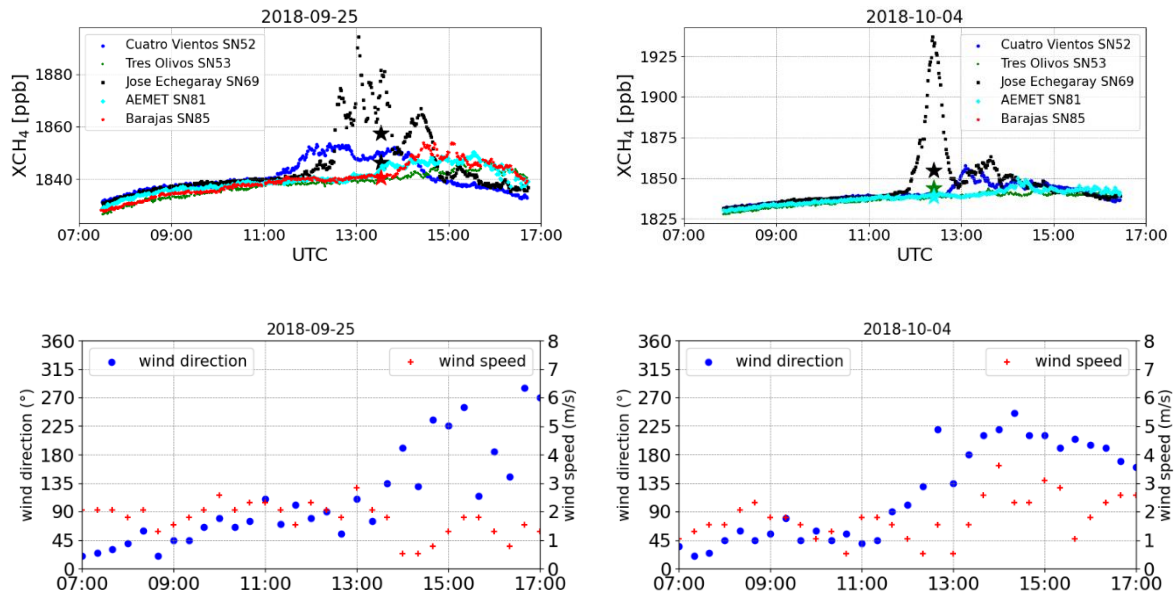
To detect whether TROPOMI is capable of measuring XCH<sub>4</sub> precisely in Madrid area, we perform intercomparison between TROPOMI and COCCON XCH<sub>4</sub> measurements. Figure 3 shows the correlation between COCCON and TROPOMI measurements. The TROPOMI data are the mean value of observations collected within a radius of 5 km around each COCCON station. The coincident COCCON XCH<sub>4</sub> is the mean value of the measurements within 30 minutes before or after the TROPOMI overpass. The distance between two stations ranges between 6 km and 14.2 km. The TROPOMI data within a circle with a large radius might cover the information from other stations nearby, which brings error in the correlation between the coincident data. Therefore, we choose 5 km as the radius of collection circle for TROPOMI and the coincident data at each station show generally good agreement. It shows that TROPOMI data have good quality and a low bias. Note that there are 1 to 2 TROPOMI measurements located within a circle of 5 km radius around each station.



**Figure 3: Correlation plot between TROPOMI observations collected within 5 km radius around each COCCON station and coincident COCCON (within 30 minutes) measurements at five stations in 2018.**

The coincident data on September 25, 2018 and October 4, 2018 show large biases at Jose Echegaray station where the SN69 COCCON instrument is located. Due to its coarser spatial resolution the TROPOMI XCH<sub>4</sub> observations do not capture the local enhancements detected by the COCCON instrument in the vicinity of the assumed source. Figure 4 illustrates the two exemplary days of the time series of COCCON SN69 and coincident TROPOMI observations. Obvious enhancements are

observed at around 13:00 UTC [by the COCCON instrument in the downwind site](#) on September 25 and at around 12:30 on October 4, 2018 (see Figure A- 2 for the other days). It is noted that the XCH<sub>4</sub> enhancements can also be observed by the other stations when the CH<sub>4</sub> plume passes over Madrid. We only discuss the two representative days here, as we focus on the specific source near the Jose Echegaray station. The Valdemingómez and Pinto waste plants are located nearby with a distance of 4.5 km and 12 km, respectively. [These five COCCON stations can serve as an independent source of information for constraining the wind speed. For example, the distance between the Jose Echegaray and Barajas is about 10 km. The highest anomalies of XCH<sub>4</sub> arrived around 1.5 hours later at Barajas station than it appeared at the Jose Echegaray station on 25 September 2018, which indicates an averaged wind speed of 1.8 m/s. This value fits well to the ERA5 model wind velocity.](#)



**Figure 4: Time series of COCCON measurements at five stations on two days in 2018. Star symbols represent the averaged TROPOMI observations within a radius of 5 km around each station. Lower panels show the wind direction and wind speed measured at the Cuatro Vientos Airport.**

TROPOMI detected 10 ppb higher XCH<sub>4</sub> at Jose Echegaray station than at Barajas station on September 25, 2018. However, COCCON observed much a higher amount of XCH<sub>4</sub> (53 ppb) at Jose Echegaray station than at Barajas station (and other stations) at around 13:00 UTC. The delayed enhancement at AEMet and Barajas stations at the downwind direction is found after the wind direction changed from north more towards south direction. Another obvious enhancement of XCH<sub>4</sub> is observed at Jose Echegaray station by the COCCON SN69 instrument at around 12:30 on October 4, 2018, with about 97 ppb higher XCH<sub>4</sub> than COCCON measurements at the other four stations. However, TROPOMI only measured about 13 ppb higher XCH<sub>4</sub> at Jose Echegaray station compared to the TROPOMI measurements at the other stations. These considerable enhancements at Jose Echegaray station observed by the COCCON instrument are likely due to the local source (the nearby Valdemingómez waste plant). The plume is in close vicinity to the source narrower than the pixel scale of the satellite, and therefore is only detected as attenuated signal by TROPOMI. The full width at the half maximum (FWHM) of the enhancement peak on October

4, 2018 roughly covers a time period of 30min, with a corresponding wind direction change of  $22.5^\circ$  ( $\sim 0.4$  rad) and an averaged  
 340 wind speed of  $1.0 \text{ m s}^{-1}$ . The distance between the COCCON SN69 to the Valdemingómez waste plant is about 4500 m. The  
 97 ppb enhancement measured by COCCON SN69 instrument yields an estimated emission strength of  $3.7 \times 10^{25} \text{ molec s}^{-1}$ .

According to the Spanish Register of Emissions and Pollutant Sources (PRTR, <http://www.en.prtr-es.es/>, last access: 20  
 February, 2021), more than 95% of total  $\text{CH}_4$  emissions are from three waste treatment and disposal plants in the Madrid  
 region (locations showed in Figure 1). The annual  $\text{CH}_4$  emission rates from the PRTR for each plant are listed in Table 3. The  
 345 total emission strength for each plant is about  $2.5 \times 10^{25} \text{ molec s}^{-1}$ . This value only considers the "cells" in production, i.e. those  
 where the waste is not yet covered with soil. The emissions from sealed cells are not included in the total emissions, but they  
 still emit  $\text{CH}_4$  during some years after sealing. So, the estimated emission rates from the inventories is expected to  
 underestimate the true emissions, which fits reasonably with the estimated emission rate derived from COCCON  
 measurements. The COCCON instruments show a very good ability to detect the source. Based on this evidence we investigate  
 350 the potential of the TROPOMI and IASI  $\text{CH}_4$  products for detecting  $\text{CH}_4$  sources in the following.

**Table 3:  $\text{CH}_4$  emission rates in three waste treatment and disposal plants in Madrid from PRTR.**

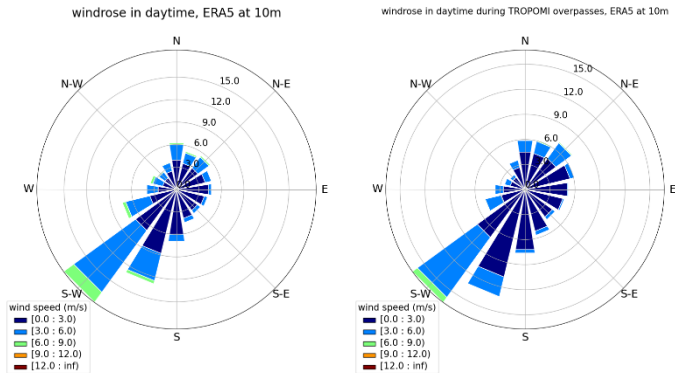
Waste treatment and disposal plants	Valdemingómez ( $\text{molec s}^{-1}$ )	Pinto ( $\text{molec s}^{-1}$ )	Alcalá ( $\text{molec s}^{-1}$ )	Total ( $\text{molec s}^{-1}$ )
2017	$7.4 \times 10^{24}$	$1.2 \times 10^{25}$	$2.1 \times 10^{24}$	$2.2 \times 10^{25}$
2018	$7.4 \times 10^{24}$	$1.3 \times 10^{25}$	$2.1 \times 10^{24}$	$2.2 \times 10^{25}$
2019	$9.8 \times 10^{24}$	$1.4 \times 10^{25}$	$9.4 \times 10^{23}$	$2.5 \times 10^{25}$

### 3.2 Predominant wind

To better representing the whole area of Madrid, the hourly ERA5 model wind at a height of 10m around Madrid is used.  
 ERA5 is the fifth generation climate reanalysis produced by the European Centre for Medium-range Weather Forecasts  
 355 (ECMWF) (Copernicus Climate Change Service, 2017). The TROPOMI overpasses over Madrid cover the time range from  
 12:00 UTC – 14:30 UTC (IASI overpasses are typically from 09:30 UTC – 10:30 UTC), but the dispersion of emitted  $\text{CH}_4$  is  
 influenced by the ground conditions (e.g. wind speed and wind direction) over a wider time range (Delkash et al., 2016; Rachor  
 et al., 2013). Therefore, the wind information between daytime (08:00 UTC – 18:00 UTC) is chosen to define the predominant  
 wind direction for each day. Figure 5 presents the wind roses for daytime between 10 November 2017 and 10 October 2020  
 360 (the first and last day with valid TROPOMI data). The dominating wind direction was southwesterly. To the northwest of  
 Madrid are the Guadarrama mountains located and the Jarama and Manzanares river basins, which influence the air flow.  
 Therefore, we use a wider wind range for the specific wind area in this study to cover the dominant wind directions, i.e. SW  
 for the range of  $135^\circ$  –  $315^\circ$  and NE for the remaining direction. If a wind direction dominates 60% of records for one day,  
 i.e., if the wind direction belongs to one specific area more than 60% of the daytime (08:00 UTC – 19:00 UTC), then this  
 365 predominant wind direction is selected for that day. The SW and NE wind fields are used for constructing wind-assigned



anomalies in this study and we will demonstrate this construction by using TROPOMI nitrogen dioxide (NO<sub>2</sub>) data in the next section. Table 4 summaries the number of days and wind speed for each specific wind area. The wind direction during the TROPOMI overpasses was 61.8% in SW wind field and 28.4% in NE wind field and their averaged wind speed is similar.



370 **Figure 5: Wind roses for daytime (08:00 UTC – 19:00 UTC) from 10 November 2017 to 10 October 2020 for the ERA5 model wind. The left panel covers all days and the right panel covers the days with TROPOMI overpasses.**

**Table 4: Number of days and the averaged ERA5 wind speed ( $\pm$  standard deviation) per specific wind area in daytime (08:00 UTC – 18:00 UTC) from 10 November 2017 to 10 October 2020. Columns 2 and 3 are for all days, and columns 4 and 5 are for days with TROPOMI overpass.**

Wind direction range	Number of days in total (%)	Averaged wind speed $\pm$ standard deviation (m s <sup>-1</sup> )	TROPOMI overpass	
			Number of days in total (%)	Averaged wind speed $\pm$ standard deviation (m s <sup>-1</sup> )
NE / >315° or <135°	30.4	2.6 $\pm$ 1.5	28.4	2.3 $\pm$ 1.2
SW / 135° – 315°	68.4	2.8 $\pm$ 1.7	61.8	2.3 $\pm$ 1.4

375 **3.3 Illustration and validation of the wind-assigned anomaly method**

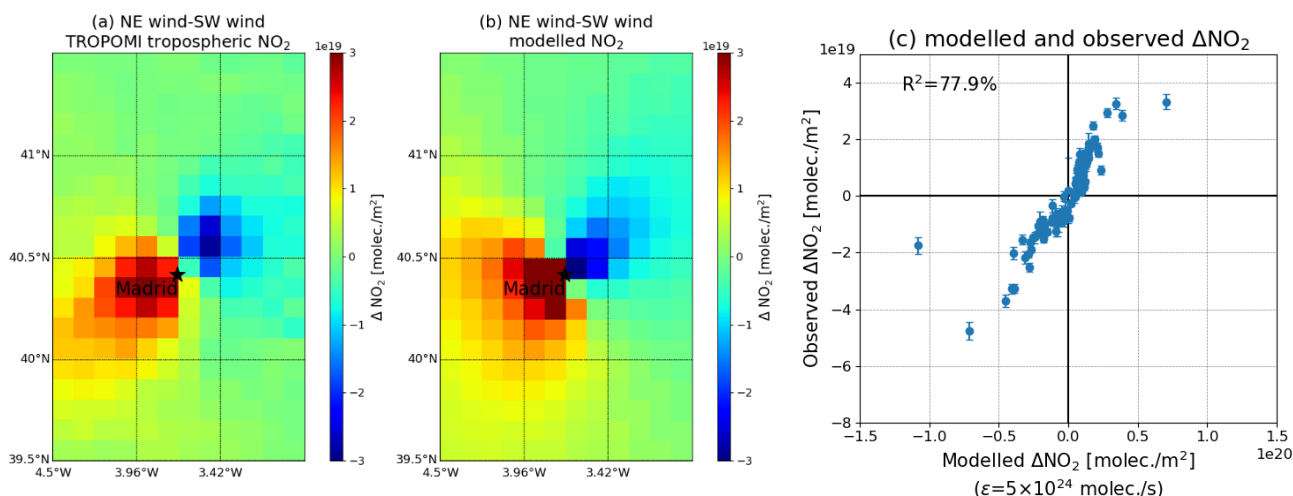
When fossil fuels are burned, nitrogen monoxide (NO) is formed and emitted into the atmosphere. NO reacts with O<sub>2</sub> to form NO<sub>2</sub> and with ozone (O<sub>3</sub>) to produce O<sub>2</sub> and NO<sub>2</sub>. NO<sub>2</sub> is an extremely reactive gas with a short lifetime of a couple of hours and has lower background levels than CH<sub>4</sub> (Kenagy et al, 2018; Shah et al., 2020). It is measured by TROPOMI with excellent quality. Therefore, it is a suitable proxy for demonstrating the method developed for the wind-assigned anomaly.

380 TROPOMI offers simultaneous observations of NO<sub>2</sub> columns with a recommended quality value for the analysis of TROPOMI NO<sub>2</sub> columns of qa>0.75 (<http://www.tropomi.eu/sites/default/files/files/publicSentinel-5P-Level-2-Product-User-Manual-Nitrogen-Dioxide.pdf>). Based on the predominant wind direction in Madrid (see section 3.2), the averaged wind-assigned anomalies are defined here as the difference of the mean TROPOMI NO<sub>2</sub> column under the wind direction from NE and the mean TROPOMI NO<sub>2</sub> column under the predominant wind direction of SW in Madrid.

385 Figure 6 (a) illustrates the wind-assigned anomalies of TROPOMI NO<sub>2</sub> ( $\Delta$ NO<sub>2</sub>) on a 0.1°  $\times$  0.135° latitude/longitude grid during 2018 – 2019. Pronounced fusiform-shape plumes are observed along NE – SW wind direction as expected. Figure 6 (b) shows the wind-assigned anomalies derived from the simple model introduced in Sect. 2.3, using Madrid city center as the

point source and an assumed emission rate ( $\varepsilon$ ) of  $5.0 \times 10^{24}$  molec  $s^{-1}$  and using ERA5 10 m wind data. The similar symmetrical positive and negative plumes to those in Figure 6 (a) imply that our method of wind-assigned anomaly is working nicely, and that the ERA5 10 m data are indeed representative for the area and that the implementation of the satellite data analysis is correct. Figure 6 (c) shows the strong correlation between the wind-assigned anomalies derived from the TROPOMI measurements and the simple plume model ( $\varepsilon = 5.0 \times 10^{24}$  molec  $s^{-1}$ ). Using the fitting method as described in Sect. B.2-2.5, we estimate an emission rate of  $3.5 \times 10^{24}$  molec  $s^{-1} \pm 3.9 \times 10^{22}$  molec  $s^{-1}$ . Here the uncertainty is due to noise of the observations and is calculated according to Eq. 21. This estimated source strength is weaker than the strength obtained by Beirle et al. (2011), where the reported  $NO_x$  emission is around  $150$  mol  $s^{-1}$  in Madrid, corresponding to a  $NO_2$  emission of  $6.8 \times 10^{25}$  molec  $s^{-1}$ . It is because our model does not consider the decay of  $NO_2$ , which results in a lower emission rate.

The result of this test using  $NO_2$  also allows the used angular spread parameter in the plume model to be adjusted (see Section 2.3 and equation (3)). As it can be seen from Figure A- 1, assuming an angular spread of  $60^\circ$  reasonably reproduces the shape of the plume.



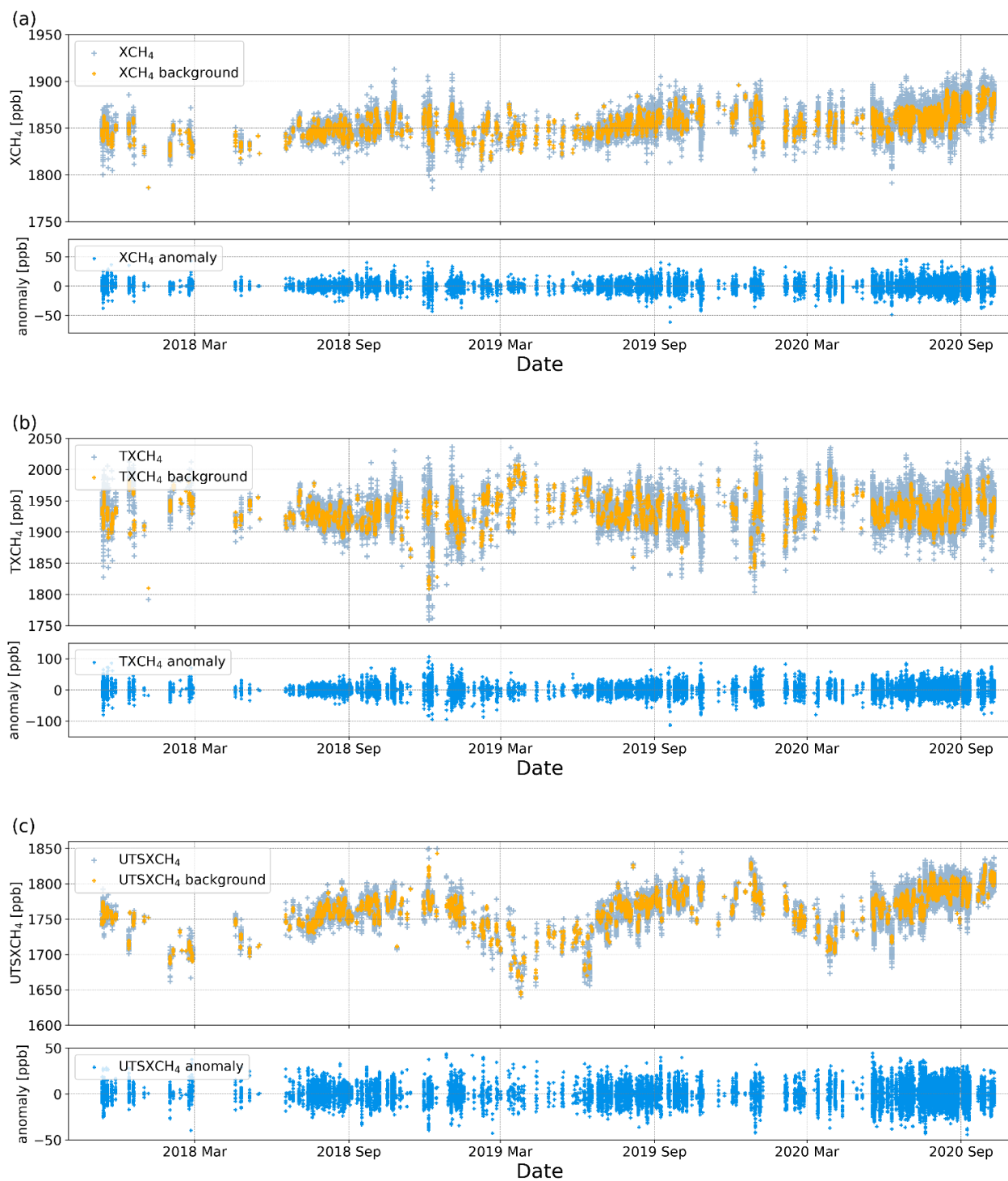
**Figure 6: Wind-assigned anomalies derived from (a) TROPOMI tropospheric  $NO_2$  column, derived from (b) our simple plume model ( $\varepsilon = 5 \times 10^{24}$  molec  $s^{-1}$ ) over Madrid in NE - SW direction on a  $0.1^\circ \times 0.135^\circ$  latitude/longitude grid during 2018 – 2020, and (c) shows the correlation plot between observed  $\Delta NO_2$  and modelled  $\Delta NO_2$  ( $\varepsilon = 5 \times 10^{24}$  molec  $s^{-1}$ ) during 2018-2019.**

### 3.4 $XCH_4$ and $TXCH_4$ anomaly

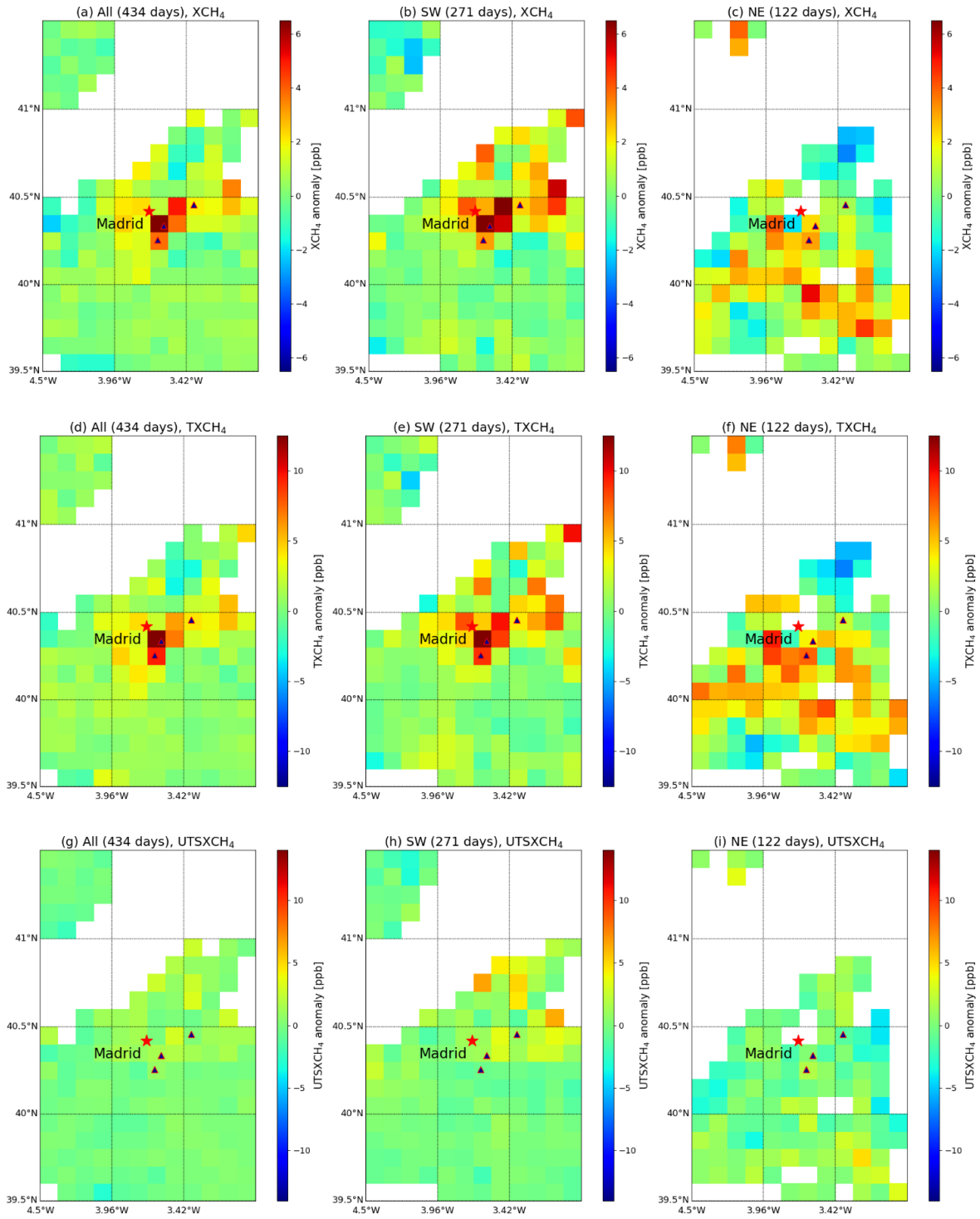
$CH_4$  has a relatively longer lifetime than  $NO_2$  and its background in the atmosphere is high. An increasing trend with obvious seasonality and strong day-to-day signals for  $XCH_4$  is seen in Figure 7 (upper panels). Therefore, these background signals need to be removed before simulating the wind-assigned anomalies (see Sect. B.1-2.4). After removing the background, the anomalies (raw data - background) represent more or less the emission from local area (Figure 7 lower panels).

Figure 8 illustrates the anomalies of  $XCH_4$ ,  $TXCH_4$  and  $UTSXCH_4$  for all measurement days, days under SW wind field and days under NE wind field. The distributions over the whole area for  $XCH_4$  and  $TXCH_4$  are similar and no obvious

enhancement is observed in UTSXCH<sub>4</sub>, as expected, since CH<sub>4</sub> abundances dominate in the troposphere. The areas where the three waste plants are located show obvious high anomalies [in the all-day-averaged figures \(Figure 8a and d\)](#) and in downwind direction, demonstrating that our method of removing the background works well and the satellite products can detect the local pollution sources after removing the background. Enhanced plumes of XCH<sub>4</sub> and TXCH<sub>4</sub> are better visible in the downwind side of SW than in the downwind side of NE wind field. Because SW is the most dominant wind direction, the SW plume signal is based on a higher number of data and thus less noise.



445 **Figure 7: Time series of (a)  $XCH_4$ , (b)  $TXCH_4$  and (c) UTS  $XCH_4$ , showing raw data and background in each upper panel and anomalies in each corresponding lower panel.**



450 **Figure 8: (a-c)  $\text{XCH}_4$ , (d-f)  $\text{TXCH}_4$  and (g-i)  $\text{UTSXCH}_4$  anomalies for all days, days with SW wind and NE wind directions. The triangle symbols represent the location of waste plants.**

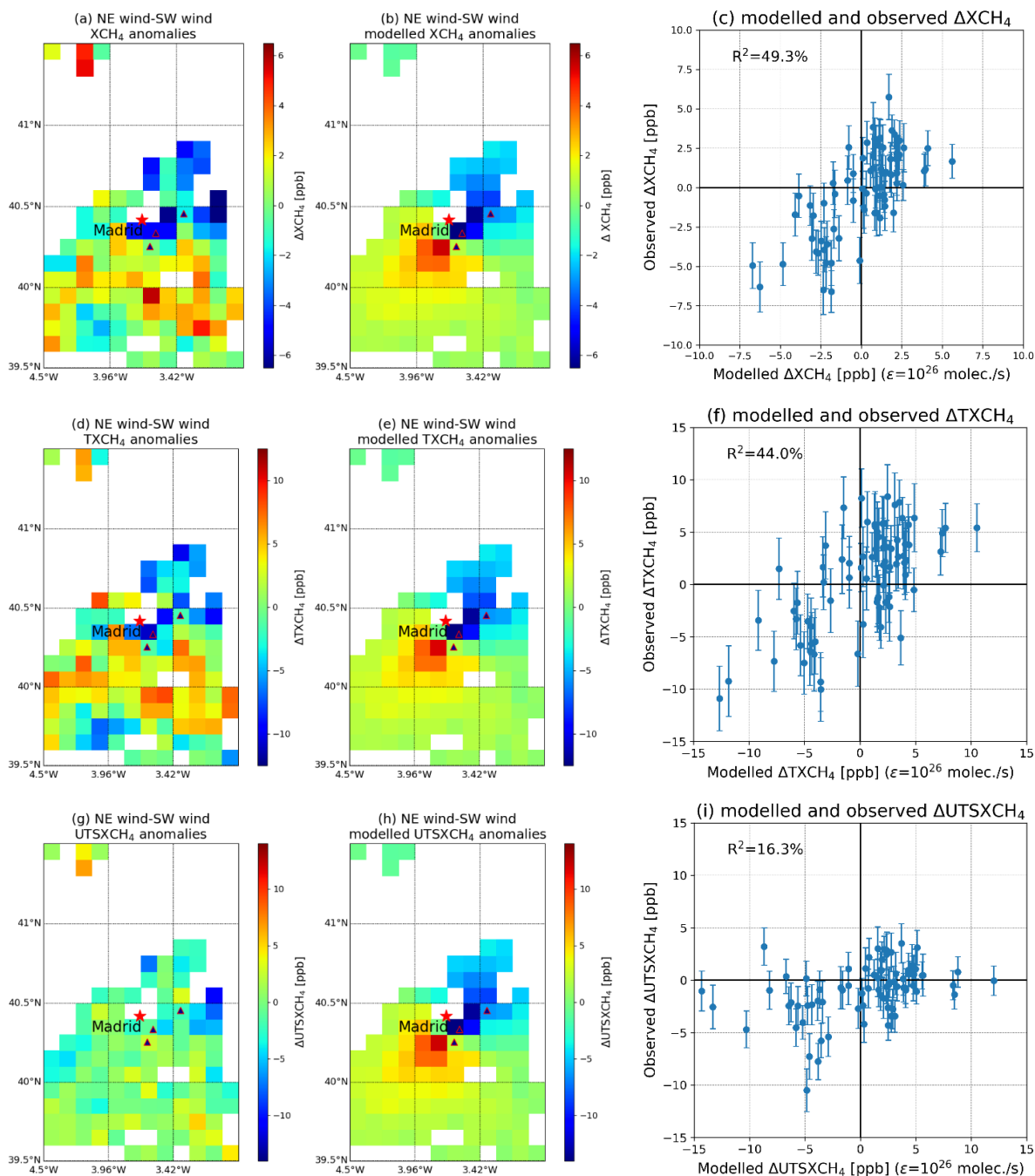
### 3.5 Estimation of CH<sub>4</sub> emission strengths from satellite data sets

The wind-assigned anomalies derived from XCH<sub>4</sub> anomalies and TXCH<sub>4</sub> anomalies on a  $0.1^\circ \times 0.135^\circ$  latitude/longitude grid are presented in Figure 9. The XCH<sub>4</sub> and TXCH<sub>4</sub> wind-assigned anomalies show similar bipolar plumes but are more disturbed compared to those derived from NO<sub>2</sub>. This is because the CH<sub>4</sub> signal is weak compared to the background concentration, so the noise level of the measurement and the imperfect elimination of the background are significant disturbing factors.

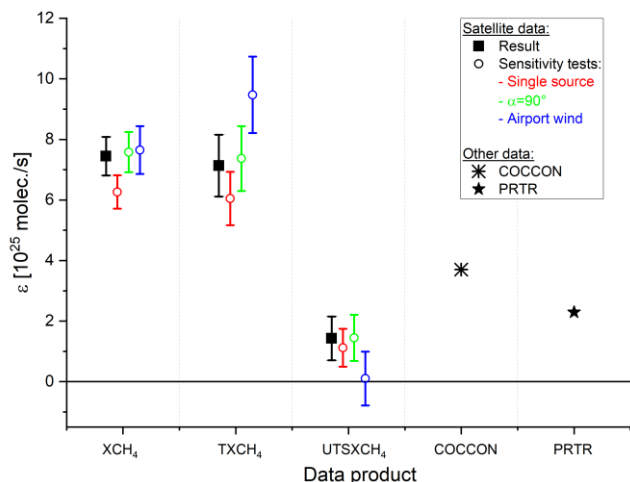
Based on the a priori knowledge of the locations of the three waste plants, we choose their locations as point sources to model the enhanced XCH<sub>4</sub> according to the wind information. The initial emission strength is  $1 \times 10^{26}$  molec s<sup>-1</sup> in total and the emission rate at each point source is repartitioned among these three sites according to Table 3. The modelled and observed wind-assigned anomalies show a reasonable linear correlation (coefficient of determination R<sup>2</sup> of about 49% and 44% for XCH<sub>4</sub> and TXCH<sub>4</sub>, respectively) with observed  $\Delta$ XCH<sub>4</sub>. Based on Eq. 18, we obtained an estimated emission rate of  $7.4 \times 10^{25} \pm 6.4 \times 10^{24}$  molec s<sup>-1</sup> for XCH<sub>4</sub> and  $7.1 \times 10^{25} \pm 1.0 \times 10^{25}$  molec s<sup>-1</sup> for TXCH<sub>4</sub>. The uncertainty values given here are the square root sum of the uncertainty due to the background signal and the data noise, which are calculated according to Eq. 20 and 21. Figure 9 (g), (h) and (i) show the wind-assigned anomalies for UTSXCH<sub>4</sub>. For the modelled UTSXCH<sub>4</sub> anomalies we assume here the CH<sub>4</sub> enhancement to occur at altitudes between 7 and 20 km a.s.l. As expected, the fit of these model data to the observed UTSXCH<sub>4</sub> data yields emission rates of close to zero ( $1.4 \times 10^{25} \pm 7.2 \times 10^{24}$  molec s<sup>-1</sup>), revealing that there is no significant plume signal above 7 km a.s.l. The fact that for TXCH<sub>4</sub> we obtain practically the same emission rates as for XCH<sub>4</sub> and that in the UTSXCH<sub>4</sub> data we see almost no plume nicely proves the quality of our careful background treatment method and the low level of cross-sensitivity between the TXCH<sub>4</sub> and UTSXCH<sub>4</sub> data products. The applied background treatment allows detecting the surface-near emission signal consistently in the total column XCH<sub>4</sub> data and in the tropospheric TXCH<sub>4</sub> data.

Figure 10 illustrates the estimated emission strengths for the different products. The emission strengths derived from the satellites are higher than the ones derived from COCCON measurements, as TROPOMI covers a larger area while COCCON measurements are only ~~sensitive to measured~~ local sources [from the nearby waste plant](#). Table 5 presents some estimated emissions for CH<sub>4</sub> from other studies. Our results fit well with the Indianapolis emissions and is reasonably comparable to the emissions from another European metropolitan cities of similar characteristics, like e.g. Berlin. [The PRTR inventory only lists the active landfill cells and does not include the closed ones in Madrid, which probably still emit for many years \(Sánchez et al., 2019\). This reference table is an additional argument for the relevance of the kind of work we are presenting here.](#)





**Figure 9: Wind-assigned XCH<sub>4</sub> plume derived from (a) TROPOMI XCH<sub>4</sub> anomalies, (d) synergetic TXCH<sub>4</sub> anomalies and (g) UTSXCH<sub>4</sub> anomalies and their corresponding modeled plume (b, e, h) over Madrid in NE – SW direction on a 0.1° × 0.135° latitude/longitude grid. The correlation plots between observed ΔXCH<sub>4</sub> and modelled ΔXCH<sub>4</sub> ( $\epsilon=1 \times 10^{26}$  molec s<sup>-1</sup>) for different products (c, f, i). Here we use the three waste plants as the point sources (blue triangle with red edge color). The initial emission rate in the plume model is  $1 \times 10^{26}$  molec s<sup>-1</sup> and proportionally distributed into three point sources based on the a priori knowledge of emission rate in each waste plant. For the modelled UTSXCH<sub>4</sub> anomalies we assume the CH<sub>4</sub> enhancements to occur at altitudes between 7 and 20 km a.s.l.**



**Figure 10: Emission strengths for different products with sensitivity tests. Also included are the COCCON observations which characterize the Valdemingómez waste plant contribution and the total of all three sources according to the PRTR inventory.**

**Table 5: CH<sub>4</sub> estimated emissions in some studies.**

Reference	Location (specific source)	Estimated emissions (molec s <sup>-1</sup> )
Cambaliza et al., 2015	Indianapolis	$8.1 \times 10^{25}$
Luther et al., 2019	Upper (coal mining)	$7.1 \times 10^{24} - 1.3 \times 10^{26}$
Klausner et al., 2020	Berlin	$2.0 \times 10^{26}$
Alvarez et al., 2018	U.S. (oil and natural gas supply chain)	$1.6 \times 10^{28}$

### 3.6 Sensitivity study for emission strength estimates

The point sources and their proportion in the total emission rate in this study are based on the a priori knowledge of three different waste plant locations. If we use a single source located at the Pinto waste disposal site only, it yields an emission rate of  $6.3 \times 10^{25}$  molec s<sup>-1</sup>, ~15% lower than that of the three-point sources for ~~TX~~CH<sub>4</sub> and  $6.0 \times 10^{25}$  molec s<sup>-1</sup> (-15%) for ~~TX~~ tropospheric CH<sub>4</sub> (see Figure 10). The opening angle ( $\alpha$ ) is experimentally selected based on the comparison between the TROPOMI measured and modelled NO<sub>2</sub> plume, which results in some uncertainties as well. Using 90° instead of 60° for  $\alpha$  in the plume model results in an emission strength of  $7.6 \times 10^{25}$  molec s<sup>-1</sup> (+3% change) for ~~TX~~CH<sub>4</sub> and of  $7.4 \times 10^{25}$  molec s<sup>-1</sup> (+4% change) for ~~TX~~ tropospheric CH<sub>4</sub>.

The surface wind can be influenced by the topography and the actual transport pathway from emission source to the measurement station is difficult to know (Chen et al., 2016; Babenhauserheide et al., 2020). To study the wind sensitivity, the hourly wind information measured at the Cuatro Vientos Airport at 10m height is used instead of the ERA5 10 m wind. There are other in situ measurements available but not used here, as the AEMet Headquarter station is affected by nearby buildings and the Barajas Airport station is very close to a secondary river (Jarama) that determines a specific wind pattern. The wind

measured at the Cuatro Vientos Airport is quite different compared to the ERA5 wind, as in situ measured NE wind becomes dominant as well and the wind speed in SW wind field increases by ~50% compared to that of ERA5 wind (Figure A- 3, Figure A- 4 and Table A- 1). Using the wind measured at the Cuatro Vientos Airport results in an emission rate of  $7.7 \times 10^{25} \text{ molec s}^{-1}$  (+4%) for  $\text{XCH}_4$  and  $9.5 \times 10^{25} \text{ molec s}^{-1}$  (+34%) for  $\text{TXCH}_4$ .

In summary, the uncertainties derived from source location, opening angle or wind cannot be ignored, but nevertheless the emission rates estimated from the space-borne observations are clearly larger than the values reported in Table 3 and are larger than the ones estimated from the COCCON SN69 observations in October 2018.

## 4 Conclusions

The present study analyzes TROPOMI  $\text{XCH}_4$  and IASI  $\text{CH}_4$  retrievals over an area around Madrid for more than 400 days within a rectangle of  $39.5^\circ\text{N} - 41.5^\circ\text{N}$  and  $4.5^\circ\text{W} - 3.0^\circ\text{W}$  ( $125 \text{ km} \times 220 \text{ km}$ ) from 10 November 2017 until 10 October 2020. During this time period, a two-week field campaign was conducted in September 2018 in Madrid, in which five ground-based COCCON instruments were used to measure  $\text{XCH}_4$  at different locations around the city center.

First, TROPOMI  $\text{XCH}_4$  is compared with co-located COCCON data from the field campaign, showing a generally good agreement, even though the radius of the collection circle for the satellite measurements is as low as 5 km. However, there are six days when obvious enhancements due to local sources were observed by COCCON around noon at the most southeast station (Jose Echegaray), which were underestimated by TROPOMI. The ground-based COCCON observations indicate a local source strength of  $3.7 \times 10^{25} \text{ molec s}^{-1}$  from observations at Jose Echegaray station on October 4, 2018, which is reasonable compared to the emissions assumed for nearby waste plants. The waste plant locations are later used as the point sources to model the emission strength for  $\text{CH}_4$ .

According to the ERA5 model wind at 10 m height, SW ( $135^\circ - 315^\circ$ ) winds (NE covering the remaining wind field) are dominant in the Madrid city center in the time range from November 2017 to October 2020. Based on this wind information, the wind-assigned anomalies are defined as the difference of satellite data between the conditions of NE wind field and SW wind field. We use the simultaneously measured tropospheric  $\text{NO}_2$  column amounts from TROPOMI as a proxy to evaluate the wind-assigned anomaly approach due to its short lifetime, and clear plume shape, by using ERA5 model wind. Pronounced and bipolar  $\text{NO}_2$  plumes are observed along the NE – SW wind direction and a tropospheric  $\text{NO}_2$  emission strength of  $3.5 \times 10^{24} \pm 3.9 \times 10^{22} \text{ molec s}^{-1}$  is estimated. This implies that our method of wind-assigned anomaly is working reliably, and that the ERA5 wind data used are indeed representative for the area and the implementation of the satellite data analysis is correct.

$\text{CH}_4$  has a long lifetime and so there are strong  $\text{CH}_4$  background signals in the atmosphere. Therefore, this background values are needed to be removed and the anomalies have to be determined before calculating emission strengths. In this study, the background values include linear increase, seasonal cycle, daily variability and horizontal variability. The areas where the three waste plants are located show obvious high anomalies, demonstrating that the satellite products can detect the local

sources after removing the background. Enhanced plumes are pronounced in the downwind side of SW, whereas the observed downwind plume signal for NE wind is noisier, partly due to the lower number of NE wind situations.

540 The wind-assigned TROPOMI XCH<sub>4</sub> anomalies show a less clear bipolar plume than NO<sub>2</sub>. This is because CH<sub>4</sub> has a long lifetime and its high background is difficult to be totally removed. Based on the wind-assigned method, the emission strength estimated from the TROPOMI XCH<sub>4</sub> data is  $7.4 \times 10^{25} \pm 6.4 \times 10^{24}$  molec s<sup>-1</sup>. In addition, this method is applied to the tropospheric partial column averaged (ground – 7 km a.s.l.) dry-air molar fractions of methane (TXCH<sub>4</sub>, obtained by combining TROPOMI and IASI products) yielding an emission strength of  $7.1 \times 10^{25} \pm 1.0 \times 10^{25}$  molec s<sup>-1</sup>. We show that in the upper  
545 troposphere/stratosphere there is no significant plume signal ( $1.4 \times 10^{25} \pm 7.2 \times 10^{24}$  molec s<sup>-1</sup>). The estimation of very similar emission rates from XCH<sub>4</sub> and TXCH<sub>4</sub> together with the estimated negligible emission rates when using data representing the upper troposphere/stratosphere proves the robustness of our method. The emission rates derived from satellites (XCH<sub>4</sub> and TXCH<sub>4</sub>) are higher than that derived from COCCON observations, as satellites cover larger areas with other CH<sub>4</sub> sources and COCCON likely measures local sources.

550 The surface wind is easily influenced by the topography, which introduces uncertainties in the estimated emission strengths. Using in situ measured wind at the Cuatro Vientos Airport instead of ERA5 model wind results in an estimated emission rate of  $7.7 \times 10^{25}$  molec s<sup>-1</sup> (+4%) for ~~TX~~CH<sub>4</sub> and  $9.5 \times 10^{25}$  molec s<sup>-1</sup> (+34%) for ~~TX~~ [tropospheric](#) CH<sub>4</sub>. Uncertainties can as well be caused by the choice of the opening angle in the plume model. The estimated emission rates with  $\alpha=90^\circ$  are  $7.6 \times 10^{25}$  molec s<sup>-1</sup> (+3%) for ~~TX~~CH<sub>4</sub> and of  $7.4 \times 10^{25}$  molec s<sup>-1</sup> (+4%) for ~~TX~~ [tropospheric](#) CH<sub>4</sub>. When using a single source located in the Madrid  
555 city center, the emission strengths are  $6.3 \times 10^{25}$  molec s<sup>-1</sup> (-15%) for ~~TX~~CH<sub>4</sub> and  $6.0 \times 10^{25}$  molec s<sup>-1</sup> (-15%) for ~~TX~~ [tropospheric](#) CH<sub>4</sub>.

In summary, in this study for the first time TROPOMI observations are used together with IASI observations and the ground-based COCCON observations to investigate CH<sub>4</sub> emissions from [landfills in](#) an important metropolitan area like the Madrid city. The COCCON instruments show a promising potential for satellite validation and an excellent ability for observation of  
560 local sources. The data presented here are provided with the confidence that TROPOMI is able to detect the tropospheric NO<sub>2</sub> and XCH<sub>4</sub> anomalies over metropolitan areas with support from meteorological wind analysis data. As outlook, this methodology could also be applied to other source regions, space-based sensors and sources of CO<sub>2</sub>.

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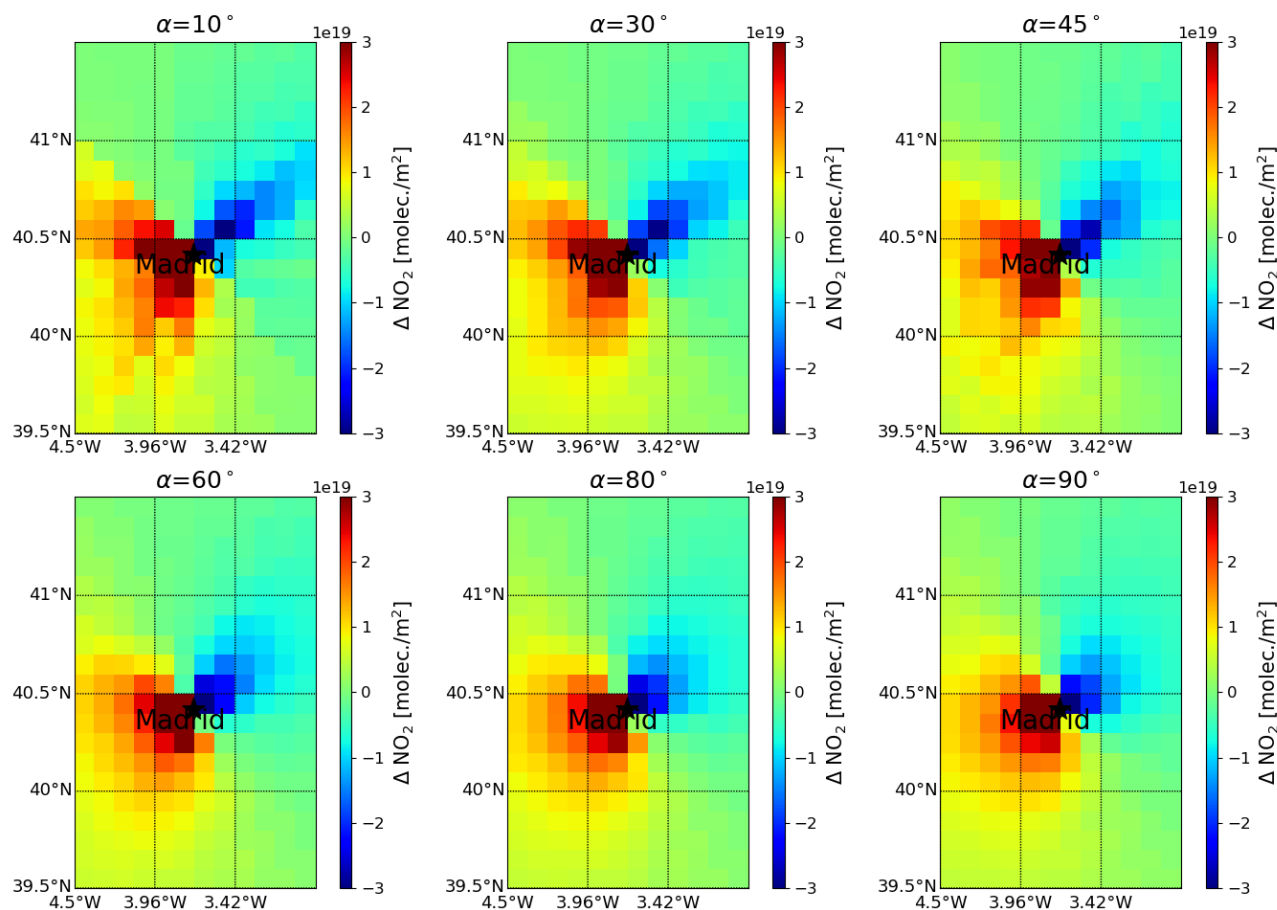
*Data availability.* The data are accessible by contacting the corresponding author (qiansi.tu@kit.edu). The SRON S5P-RemoTeC scientific TROPOMI CH<sub>4</sub> dataset from this study is available for download at <https://doi.org/10.5281/zenodo.4447228> (Lorente et al., 2021, last access: 06 May 2021). The TROPOMI data set is publicly available from <https://scihub.copernicus.eu/> (last access: 06 May 2021; ESA, 2020). The access and use of any Copernicus Sentinel data available through the Copernicus Open Access Hub are governed by the legal notice on the use of Copernicus Sentinel Data and Service Information, which is given here: [https://sentinels.copernicus.eu/documents/247904/690755/Sentinel\\_Data\\_Legal\\_Notice](https://sentinels.copernicus.eu/documents/247904/690755/Sentinel_Data_Legal_Notice) (last access: 06 May 2021; European Commission, 2020). The MUSICA IASI data set is available for download via <https://doi.org/10.35097/408> (Schneider et al. 2021c).

*Author contributions.* Qiansi Tu, Frank Hase and Omaira García developed the research question. Qiansi Tu wrote the manuscript and performed the data analysis with input from Frank Hase, Omaira García, Matthias Schneider and Farahnaz Khosrawi. Frank Hase suggested the method of constructing wind-assigned anomalies for source quantification. Matthias Schneider suggested the method for calculating the anomalies, for fitting the emission rates and estimating the uncertainty. Omaira García provided the COCCON and meteorological data and helped to interpret it. Tobias Borsdorff and Alba Lorente supported technically for the TROPOMI data analysis. Matthias Schneider, Benjamin Ertl and Christopher J. Diekmann provided the combined (MUSICA IASI + TROPOMI) data and supported technically for the analysis of these data. All other coauthors participated in the field campaign and provided the data. All authors discussed the results and contributed to the final manuscript.

*Competing interests.* The authors declare that they have no conflict of interest.

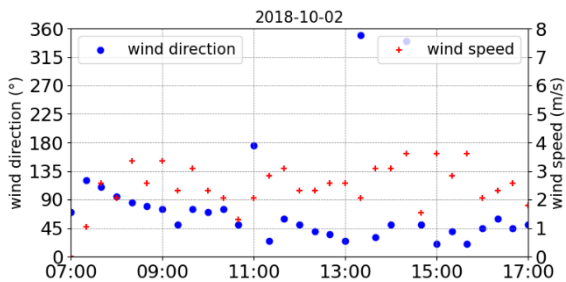
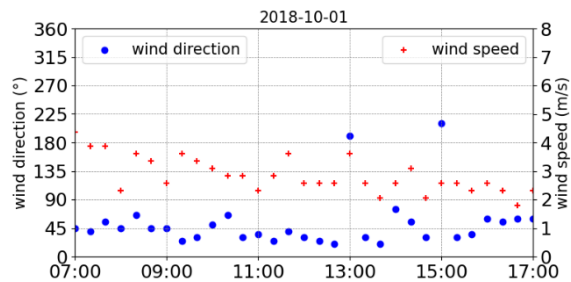
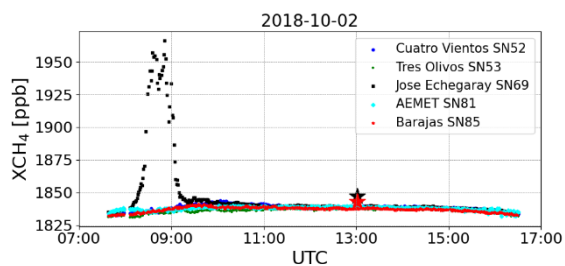
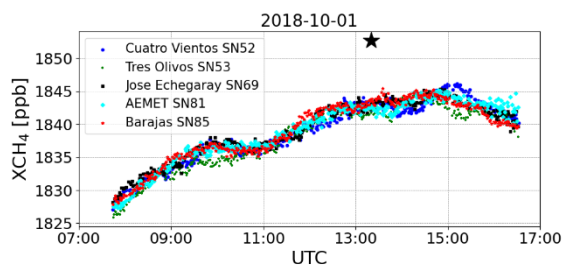
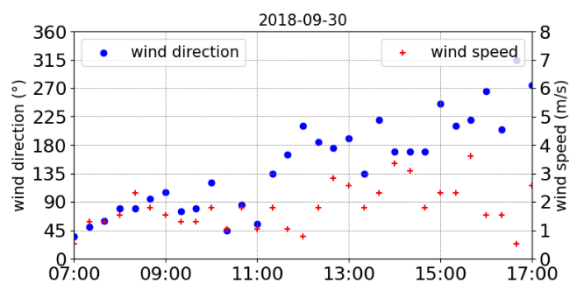
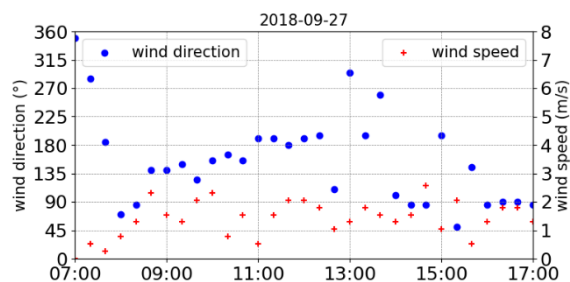
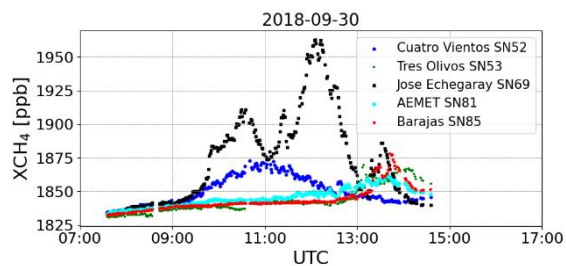
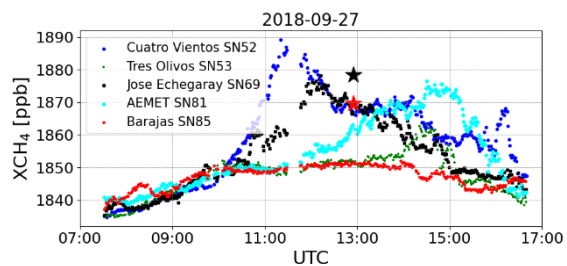
*Acknowledgements.* We acknowledge ESA support through the COCCON-PROCEEDS and COCCON-PROCEEDS II projects. In addition, this research was funded by the Ministerio de Economía y Competitividad from Spain through the INMENSE project (CGL2016-80688-P). This research has largely benefit from funds of the Deutsche Forschungsgemeinschaft (provided for the two projects MOTIV and TEDDY with IDs/290612604 and 416767181, respectively). Important part of this work was performed on the supercomputer ForHLR funded by the Ministry of Science, Research and the Arts Baden-Württemberg and by the German Federal Ministry of Education and Research.

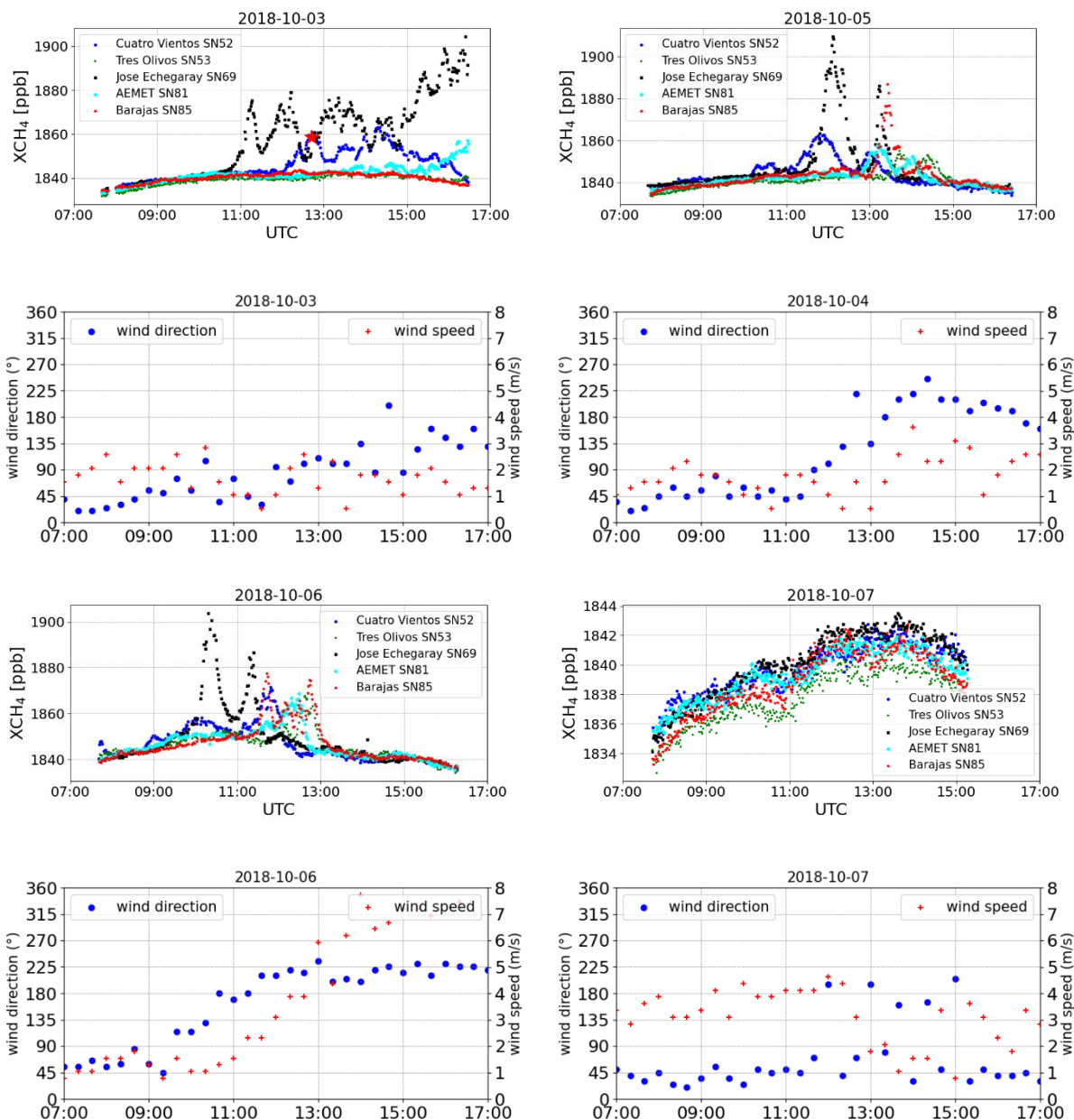
We acknowledge the support by the Deutsche Forschungsgemeinschaft and the Open Access Publishing Fund of the Karlsruhe Institute of Technology.



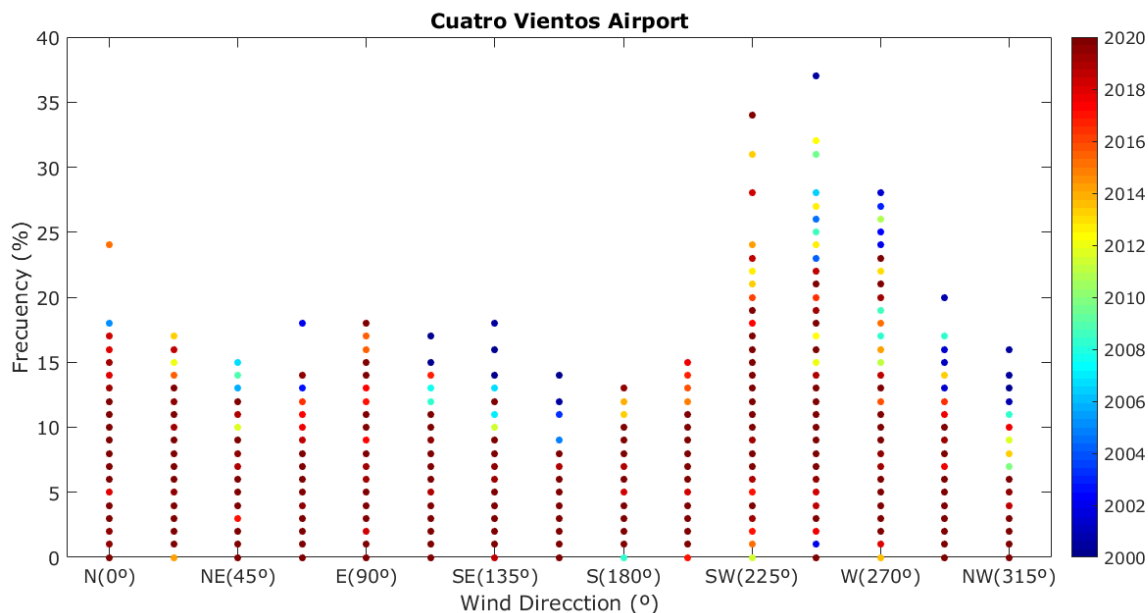
**Figure A- 1:** Examples of wind-assigned  $\text{NO}_2$  plume based on the simple plume model ( $\varepsilon = 5.0 \times 10^{24}$  molec  $\text{s}^{-1}$ ) using Madrid as the point source in NE – SW direction on a  $0.1^\circ \times 0.135^\circ$  latitude/longitude grid with different opening angle ( $\alpha$ ) from  $10^\circ$  to  $90^\circ$ .



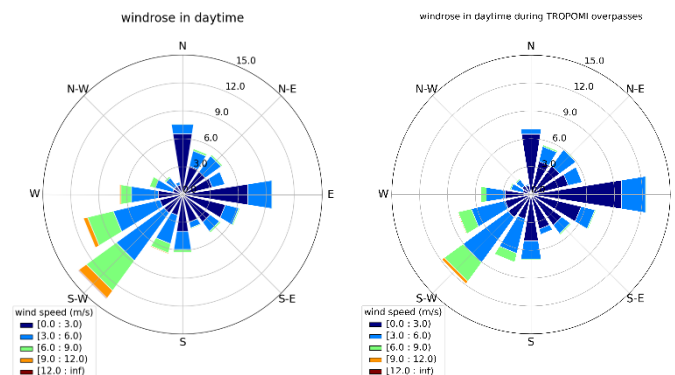




**Figure A- 2: Time series of COCCON measurements at five stations and corresponding time series of wind fields (direction and speed) measured at the Cuatro Vientos Airport on eight days during MEGEI-MAD campaign in 2018. Star symbols represent the TROPOMI observations within a radius of 5 km around each station.**



**Figure A- 3: Percentage of occurrence for wind direction measured at the Cuatro Vientos Airport between 2000 and 2020. The predominant wind direction is southwest and up to 35% of time (personal communication of Omaira García).**



**Figure A- 4: Wind roses for daytime (08:00 UTC – 19:00 UTC) from 10 November 2017 to 11 September 2020 from the wind measurements at the Cuatro Vientos Airport. The left panel covers all days and the right panel covers the days with TROPOMI overpasses.**

**Table A- 1 Number of days and the averaged wind speed ( $\pm$  standard deviation) per specific wind area in daytime (08:00 UTC – 18:00 UTC) from 10 November 2017 to 11 September 2020 measured at the Cuatro Vientos Airport. Columns 2 and 3 are for all days, and columns 4 and 5 are for days with TROPOMI overpass.**

Wind direction range	Number of days in total (%)	Averaged wind speed $\pm$ standard deviation ( $\text{m s}^{-1}$ )	TROPOMI overpass	
			Number of days in total (%)	Averaged wind speed $\pm$ standard deviation ( $\text{m s}^{-1}$ )
NE / $>315^\circ$ or $<135^\circ$	35.4	$2.4 \pm 1.5$	36.0	$2.2 \pm 1.3$
SW / $135^\circ - 315^\circ$	49.3	$4.2 \pm 2.5$	44.4	$3.4 \pm 2.1$

**B.1 CH<sub>4</sub> background signal**

The satellite data can be written as a vector  $\mathbf{y}$ , where each element corresponds to an individual satellite data point. This signal is caused by a CH<sub>4</sub> background signal and the CH<sub>4</sub> plume due to the emissions from the waste disposal sites near Madrid:

$$\mathbf{y} = \mathbf{y}_{BG} + \mathbf{y}_{plume} \quad \text{Eq. 7}$$

It is of great importance to adequately separate both components for estimating the emission strength from the satellite data.

635 For determining the background signal ( $\mathbf{y}_{BG}$ ), we setup a background model:

$$\mathbf{M}_{BG} = \mathbf{y}_{BG} = \mathbf{K}_{BG} \mathbf{x}_{BG} \quad \text{Eq. 8}$$

The matrix  $\mathbf{K}_{BG}$  is a Jacobian matrix that allows to reconstruct the background according to a few background model coefficients (the elements of the vector  $\mathbf{x}_{BG}$ ). We also create a Jacobian  $\mathbf{K}_{BG}^*$ , which is the same as  $\mathbf{K}_{BG}$  but set to zero for observations where the wind data suggest a significant impact of the CH<sub>4</sub> plume on the satellite data. The calculations of the plume CH<sub>4</sub> signals are made according to Sect. 2.3. With the use of  $\mathbf{K}_{BG}^*$  we make sure that the estimated background signal is not affected by the CH<sub>4</sub> plume.

640 The Jacobian matrix  $\mathbf{K}_{BG}$  considers a smooth background, which is a constant CH<sub>4</sub> value, a linear increase with time and a seasonal cycle described by the amplitude and phase of the three frequencies 1/year, 2/year, and 3/year. Furthermore, we fit a daily anomaly, which is the same for all data measured during a single day and a horizontal anomaly, which is the same for any time but dependent on the horizontal location. For the latter we use a  $0.1^\circ \times 0.135^\circ$  (latitude  $\times$  longitude) grid.

645 We invert the problem in order to estimate the background model coefficients (elements of the vector  $\mathbf{x}_{BG}$ ):

$$\hat{\mathbf{x}}_{BG} = \mathbf{G}_{BG} \mathbf{y} \quad \text{Eq. 9}$$

With  $\mathbf{G}_{BG}$  being the so-called gain matrix

$$\mathbf{G}_{BG} = (\mathbf{K}_{BG}^{*T} \mathbf{S}_{y,n}^{-1} \mathbf{K}_{BG}^* + \mathbf{S}_a^{-1})^{-1} \mathbf{K}_{BG}^{*T} \mathbf{S}_{y,n}^{-1} \quad \text{Eq. 10}$$

The matrix  $\mathbf{S}_{y,n}$  stands for the noise covariance of the satellite data. For constraining the problem, we use a diagonal  $\mathbf{S}_a^{-1}$  (no constraint between different coefficients) with a very low constraint value for the coefficient determining the constant and higher constraint values for the other coefficients. Here we use  $\mathbf{S}_{y,BG}$  as the diagonal matrix with the mean square value of the difference  $\mathbf{y}_{BG} - \mathbf{K}_{BG}^* \hat{\mathbf{x}}_{BG}$  being the diagonal elements. In this context,  $\mathbf{S}_{y,BG}$  considers the deficits of the background model.

650 The uncertainty of the background model coefficients can be calculated as:

$$\mathbf{S}_{\hat{\mathbf{x}}_{BG}} = \mathbf{G}_{BG} \mathbf{S}_{y,BG} \mathbf{G}_{BG}^T \quad \text{Eq. 11}$$

For each day there is an uncertainty in the background coefficients and the uncertainty is correlated with the uncertainty at other days. All this information is provided in the uncertainty covariance  $\mathbf{S}_{\hat{\mathbf{x}}_{BG}}$ .

655 With the full Jacobian  $\mathbf{K}_{BG}$  we can now model the background for the measurement state (also for the measurements that are assumed to be affected by the CH<sub>4</sub> waste disposal plume):

$$\mathbf{y}_{BG} = \mathbf{K}\hat{\mathbf{x}}_{BG} \quad \text{Eq. 12}$$

and calculate the plume signal according to Eq. 7 as:

$$\mathbf{y}_{plume} = \mathbf{y} - \mathbf{K}\hat{\mathbf{x}}_{BG} \quad \text{Eq. 13}$$

The uncertainty of these plume signal is the sum of the uncertainties of the satellite data  $\mathbf{S}_{y,n}$  and the uncertainty of the estimated background:

$$\mathbf{S}_{y,plume} = \mathbf{S}_{y,n} + \mathbf{K}\mathbf{S}_{\hat{\mathbf{x}}_{BG}}\mathbf{K}^T \quad \text{Eq. 14}$$

## B.2 Fitting of CH<sub>4</sub> emission rates

660 Because the CH<sub>4</sub> plume signal is rather weak compared to the CH<sub>4</sub> background uncertainty and the noise level of the satellite data, we have to work with averages in order to reduce the data noise. The averaging is made by classifying the observation in two predominant wind categories. We calculate the average plume maps for the southwest and northeast wind situations (see Figure 6 and Figure 8). Then we calculate the difference between the south-west and north-east plume maps (the wind-assigned anomalies or  $\Delta$ -maps). All the calculations are made by binning all observations that fall within a certain  $0.135^\circ \times 0.1^\circ$  (longitude  $\times$  latitude) area. In order to significantly reduce the data noise, we only consider averages for the  $0.135^\circ \times 0.1^\circ$  areas based on at least 25 individual observations made under southwest wind conditions and 25 individual observations made under northeast wind conditions. The binning, the averaging, the wind-assigned  $\Delta$ -maps calculations, and the data number filtering is achieved by operator  $\mathbf{D}$ , and we can write:

$$\Delta\mathbf{y}_{plume} = \mathbf{D}\mathbf{y}_{plume} \quad \text{Eq. 15}$$

and

$$\Delta\mathbf{S}_{y,plume} = \mathbf{D}\mathbf{S}_{y,plume}\mathbf{D}^T \quad \text{Eq. 16}$$

670 Here  $\Delta\mathbf{y}_{plume}$  is a column vector whose elements capture the different signal of the two wind directions at the different locations and  $\Delta\mathbf{S}_{y,plume}$  is the corresponding uncertainty covariance.

For modelling the plume signals we use a priori knowledge of CH<sub>4</sub> emission locations, i.e. assuming a repartition of the emissions between the three waste disposal sites according to Table 3 (see Sect. 3.1). Together with information from the wind, we then model the CH<sub>4</sub> plume's wind-assigned anomaly signal  $\Delta\mathbf{y}_{plume}$ :

$$\Delta\mathbf{y}_{plume} = \Delta\mathbf{k}\mathbf{x} \quad \text{Eq. 17}$$

675 Here the Jacobian  $\Delta\mathbf{k}$  (a column vector) represents the wind-assigned anomaly model as described in Sect. 2.3. It describes how an emission at the waste disposal sites according to Table 3 would be seen in the difference signal. We are interested in the coefficient  $\mathbf{x}$  (a scalar describing how the assumed emissions from Table 3 have to be scaled by a common factor in order to achieve the best agreement with the observed plume).

Similar to Eq. 9 and Eq. 10 we write:

$$\hat{\mathbf{x}} = \mathbf{g}^T \Delta\mathbf{y}_{plume} \quad \text{Eq. 18}$$

680 with the row vector

$$\mathbf{g}^T = (\Delta \mathbf{k}^T \Delta \mathbf{S}_{y,\text{plume}}^{-1} \Delta \mathbf{k})^{-1} \Delta \mathbf{k}^T \Delta \mathbf{S}_{y,\text{plume}}^{-1} \quad \text{Eq. 19}$$

This fitting of the emission rate correctly considers the respective uncertainty of the difference signals at the different locations.

Because of the small plume signals, it is important to estimate the reliability of the fitted emission rate. The uncertainty of  $x$  due to the background uncertainty and the noise in the satellite data can be estimated as:

$$\epsilon_{BG} = \sqrt{\mathbf{g}^T \mathbf{D} \mathbf{K} \mathbf{S}_{\hat{x},BG} \mathbf{K}^T \mathbf{D}^T \mathbf{g}} \quad \text{Eq. 20}$$

and

$$\epsilon_n = \sqrt{\mathbf{g}^T \mathbf{D} \mathbf{S}_{y,n} \mathbf{D}^T \mathbf{g}} \quad \text{Eq. 21}$$

685 respectively.



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