



Weakening of Antarctic Stratospheric Planetary Wave Activities in Early Austral Spring Since the Early 2000s: A Response to Sea Surface Temperature Trends

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1 **Abstract**

2 Using multiple reanalysis datasets and modeling simulations, the trends of
3 Antarctic stratospheric planetary wave activities in early austral spring since the early
4 2000s are investigated in this study. We find that the stratospheric planetary wave
5 activities in September have weakened significantly since 2000, which is related to the
6 weakening of the tropospheric wave sources in the extratropical southern hemisphere.
7 Further analysis indicates that the trend of September sea surface temperature (SST)
8 over 20°N-70°S is statistically linked to the weakening of stratospheric planetary wave
9 activities. Numerical simulations support the result that the SST trend in the
10 extratropical southern hemisphere (20°S-70°S) and the tropics (20°N-20°S) induce the
11 weakening of wave-1 component of tropospheric geopotential height in the
12 extratropical southern hemisphere, which subsequently leads to the decrease in
13 stratospheric wave flux. The responses of stratospheric wave activities in the southern
14 hemisphere to stratospheric ozone recovery is not significant in simulations. In addition,
15 both reanalysis data and numerical simulations indicate that the Brewer-Dobson
16 circulation (BDC) related to wave activities in the stratosphere has also been weakening
17 in early austral spring since 2000 due to the trend of September SST in the tropics and
18 extratropical southern hemisphere.

19

20 **Key words:** *Antarctic; Stratospheric planetary wave activities; Tropospheric wave*
21 *sources; Sea surface temperature*

22



23 1. Introduction

24 The stratospheric planetary wave activities have important influences on
25 stratospheric temperature (e.g., Hu & Fu, 2009; Lin et al., 2009; Li & Tian, 2017; Li et
26 al., 2018), polar vortex (e.g., Kim et al., 2014; Zhang et al., 2016; Hu et al., 2018) and
27 distribution of chemical substances (e.g., Gabriel et al., 2011; Ialongo et al., 2012;
28 Kravchenko et al., 2011; Zhang et al., 2019a). Meanwhile, the stratospheric circulation
29 modulated by planetary waves can exert impacts on tropospheric weather and climate
30 (e.g., Haigh et al., 2005; Zhang et al., 2019b) through downward control processes
31 (Haynes et al., 1991), which is useful for extended forecast by using preceding signals
32 in the stratosphere (e.g., Baldwin et al., 2001; Wang et al., 2020).

33 The planetary perturbations generated by large-scale topography, convection and
34 continent-ocean heating contrast can propagate from the troposphere to the stratosphere
35 (Charney & Drazin, 1961) and form stratospheric planetary waves. As the land-sea
36 thermal contrast in the northern hemisphere is larger than that in the southern
37 hemisphere and produces stronger zonal forcing for the genesis of stratospheric waves,
38 the majority of attention has been given to wave activities and their impacts on weather
39 and climate in the northern hemisphere (e.g., Kim et al., 2014; Zhang et al., 2016; Hu
40 et al., 2018). However, planetary wave activities in the southern hemisphere also play
41 an important role in heating the stratosphere dynamically (e.g., Hu & Fu, 2009; Lin et
42 al., 2009), which suppresses Polar Stratospheric Clouds (PSCs) formation and ozone
43 depletion (e.g., Shen et al., 2020a; Tian et al., 2018). The Antarctic sudden stratospheric
44 warming (SSW) that occurred in 2002 (e.g., Baldwin et al., 2003; Nishii & Nakamura,



45 2004; Newman & Nash, 2005) and 2019 (e.g., Yamazaki et al., 2020; Shen et al., 2020a;
46 Shen et al., 2020b) were associated with significant upward propagation of wave flux.
47 Such episodes are extraordinarily rare in the history, and the one in 2019 contributed to
48 the formation of the smallest Antarctic ozone hole on record (WMO, 2019). In addition,
49 some studies reported that wildfires in Australia at the end of 2019 are related to
50 negative phase of the Southern Annular Mode (SAM), which was induced by the
51 extended influence of the SSW event that occurred in September (Lim et al., 2019; Shen
52 et al., 2020b). In a word, the Antarctic planetary wave activities are important for the
53 stratosphere-troposphere interactions and climate system in the southern hemisphere.

54 Long-term observations in the Antarctic stratosphere show a significant ozone
55 decline from the early 1980s to the early 2000s due to anthropogenic emission of
56 chlorofluorocarbons (CFCs) (WMO, 2011) and a recovery signal since 2000s because
57 of phasing out CFCs in response to Montreal Protocol (e.g., Angell and Free, 2009;
58 Krzyścin, 2012; Zhang et al., 2014; Banerjee et al., 2020). The Antarctic stratospheric
59 ozone depletion and recovery have important impacts on climate in the southern
60 hemisphere. The ozone depletion cools the Antarctic stratosphere through reducing
61 absorption of radiation and leads to the strengthening of Antarctic polar vortex during
62 austral spring (e.g., Randel & Wu, 1999; Solomon et al., 1999; Thompson et al., 2011).
63 The anomalous circulation in the Antarctic stratosphere during austral spring exerts
64 impacts on tropospheric circulations (e.g., intensification of SAM index, poleward shift
65 of tropospheric jet position and expansion of the Hadley cell edge) in the subsequent
66 months (e.g., Thompson et al., 2011; Swart & Fyfe, 2012; Son et al., 2018; Banerjee et



67 al., 2020) and influences the distribution of precipitation and dry zone in the southern
68 hemisphere (e.g., Thompson et al., 2011; Barnes et al., 2013; Kang et al., 2011).
69 Following the healing of ozone loss in the Antarctic ozone hole since 2000s (e.g.,
70 Solomon et al., 2016; Susan et al., 2019), great attention has been paid on possible
71 impacts of ozone recovery on climate system in the southern hemisphere (e.g., Son et
72 al., 2008; Barnes et al., 2013; Xia et al., 2020; Banerjee et al., 2020). Son et al. (2008)
73 implemented the Chemistry-Climate Model Validation (CCMVal) models to predict the
74 response of the southern hemisphere westerly jet to stratospheric ozone recovery. Based
75 on the Phase 5 of Coupled Model Intercomparison Projects (CMIP5) models, Barnes et
76 al. (2013) proposed that the tropospheric jet and dry zone edge no longer shift poleward
77 during austral summer since the early 2000s due to ozone recovery. Banerjee et al.
78 (2020) analyzed observations and reanalysis datasets. They found that following the
79 ozone recovery after 2000, the increase of SAM index and the poleward shifting of
80 tropospheric jet position as well as the Hadley cell edge all experienced a pause. Their
81 results suggest that ozone depletion and recovery have made important contributions to
82 the climate shift that occurred around 2000 in the southern hemisphere.

83 However, some previous studies have reported zonally asymmetric warming
84 patterns in Antarctic stratosphere, which are generated by increased planetary wave
85 activities during austral spring from the early 1980s to the early 2000s (Hu & Fu, 2009;
86 Lin et al., 2009). Note that the Antarctic stratosphere was experiencing radiative cooling
87 in the same period due to ozone depletion (e.g., Randel & Wu, 1999; Solomon et al.,
88 1999; Thompson et al., 2011). The increase in stratospheric planetary wave activities



89 cannot be explained by ozone decline, because the acceleration of stratospheric
90 circumpolar wind caused by radiative cooling induces more wave energy to be reflected
91 back to the troposphere (e.g., Andrews et al., 1987; Holton et al., 2004). Hu & Fu (2009)
92 attributed the increase in Antarctic stratospheric wave activities to the SST trend from
93 the 1980s to the 2000s. Their results indicate that in addition to ozone change, other
94 factors such as SST trend also contribute to climate change in the southern hemisphere.
95 Moreover, the phase of Interdecadal Pacific Oscillation (IPO) also changed at around
96 2000 (e.g., Trenberth et al., 2013). SST variation influences Rossby wave propagation
97 and tropospheric wave sources, and thereby indirectly affects stratospheric wave
98 activities (e.g., Lin et al., 2012; Hu et al., 2018; Tian et al., 2018). The questions here
99 are: (1) Has the stratospheric planetary wave activity trend in the southern hemisphere
100 been shifting since the 2000s? (2) What are the factors responsible for the trend of
101 Antarctic stratospheric planetary wave activity since the 2000s?

102 In this study, we reveal the trend of Antarctic planetary wave activity in early
103 austral spring since the 2000s based on multiple reanalysis datasets. We also conduct
104 sensitive experiments forced by linear increments of ozone and SST fields since the
105 2000s to investigate the response of Antarctic planetary activity to above two factors.
106 The remainder of the paper is organized as follows. Section 2 describes the data,
107 methods and configurations of model simulations. Section 3 presents the trends of
108 stratospheric and tropospheric wave activities in early austral spring. Section 4
109 investigates the connections between the trends of SST and stratospheric wave activities.
110 Section 5 discusses the responses of tropospheric wave source and stratospheric wave



111 activity to SST trend based on simulations. Major conclusions and discussion are
112 presented in Section 6.

113 **2. Datasets, methods and experimental configurations**

114 a. Datasets

115 In this study, daily and monthly mean data extracted from the Modern-Era
116 Retrospective analysis for Research and Applications Version 2 (MERRA-2;
117 Bosilovich et al., 2015) dataset are used to calculate trends of zonally averaged zonal
118 wind and temperature, BDC, tropospheric wave sources, and the Elisassen-Palm (E-P)
119 flux and its divergence in September. To verify the trend of stratospheric E-P flux, we
120 also refer to the results derived from the European Centre for Medium-range Weather
121 Forecasting (ECMWF) Interim Reanalysis (ERA-Interim; Dee et al., 2011) dataset, the
122 Japanese 55-year Reanalysis (JRA-55; Kobayashi et al., 2015) dataset and the National
123 Centers for Environmental Prediction-Department of Energy Global Reanalysis 2
124 (NCEP-2; Kanamitsu et al., 2002) dataset.

125 SST data are extracted from the Extended Reconstructed Sea Surface Temperature
126 (ERSST) dataset, which is a global monthly mean sea surface temperature dataset
127 derived from the International Comprehensive Ocean-Atmosphere Dataset (ICOADS).
128 The ERSST is on global $2^{\circ} \times 2^{\circ}$ grid and covers the period from January 1854 to the
129 present. We use the newest version (version 5, i.e., v5) dataset to calculate trends and
130 correlations, and produce SST forcing field for model simulations. More details about
131 this version of ERSST can be found in Huang et al. (2017).

132 In addition, the unfiltered Interdecadal Pacific Oscillation (IPO) index derived



133 from the ERSST v5 dataset is also used in this study. The IPO index is available at
 134 <https://psl.noaa.gov/data/timeseries/IPOTPI/tpi.timeseries.ersstv5.data> and more
 135 detailed information about the index can be found in Henley et al. (2015).

136 b. Diagnosis of wave activities and Brewer-Dobson circulation

137 Planetary wave activities are measured by E-P flux ($\vec{F} \equiv (0, F^{(\phi)}, F^{(z)})$) and its
 138 divergence D_F . Their algorithms are expressed by Eqs. (1)-(3) (Andrews et al., 1987):

139
$$F^{(\phi)} = \rho_0 a \cos \phi (\overline{u_z v' \theta'} / \overline{\theta_z} - \overline{v' u'}) \quad (1)$$

140
$$F^{(z)} = \rho_0 a \cos \phi \{ [f - (a \cos \phi)^{-1} (\overline{u \cos \phi})_\phi] \overline{v' \theta'} / \overline{\theta_z} - \overline{w' u'} \} \quad (2)$$

141
$$D_F = \frac{\nabla \cdot \vec{F}}{\rho_0 a \cos \phi} = \frac{(a \cos \phi)^{-1} \frac{\partial}{\partial \phi} (F^{(\phi)} \cos \phi) + \frac{\partial F^{(z)}}{\partial z}}{\rho_0 a \cos \phi} \quad (3)$$

142 where u, v represent zonal and meridional components of horizontal wind, w is
 143 vertical velocity, θ is potential temperature, a is the Earth radius, f is the Coriolis
 144 parameter, z is geopotential height, ϕ is latitude, ρ_0 is the background air density.

145 The quasi-geostrophic refractive index (RI) is used to diagnose the environment
 146 of wave propagation (Chen & Robinson, 1992). Its algorithm is written as Equation (4):

147
$$RI = \frac{\overline{q}_\phi}{\overline{u}} - \left(\frac{k}{a \cos \phi} \right)^2 - \left(\frac{f}{2NH} \right)^2 \quad (4)$$

148 where the zonal-mean potential vorticity meridional gradient \overline{q}_ϕ is

149
$$\overline{q}_\phi = \frac{2\Omega}{a} \cos \phi - \frac{1}{a^2} \left[\frac{(\overline{u \cos \phi})_\phi}{a \cos \phi} \right]_\phi - \frac{f^2}{\rho_0} \left(\rho_0 \frac{\overline{u_z}}{N^2} \right)_z \quad (5)$$

150 q , k , N^2 and Ω are the potential vorticity, zonal wave number, buoyancy
 151 frequency, and Earth's angular frequency, respectively.

152 The Brewer-Dobson circulation driven by wave breaking in the stratopause is
 153 closely related to stratospheric wave activities. Its meridional and vertical components



154 (\bar{v}^*, \bar{w}^*) and stream function $(\psi^*(p, \phi))$ are expressed by Eqs. (4)-(6) (Andrews et al.,
155 1987; Birner & Bönisch, 2011) :

$$156 \quad \bar{v}^* \equiv \bar{v} - \rho_0^{-1}(\rho_0 \overline{v'\theta'}) / \overline{\theta_z}_z \quad (6)$$

$$157 \quad \bar{w}^* \equiv \bar{w} + (a \cos \phi)^{-1}(\cos \phi \cdot \overline{v'\theta'}) / \overline{\theta_z}_\phi \quad (7)$$

$$158 \quad \psi^*(p, \phi) = \int_0^p \frac{-2\pi a \cdot \cos \phi \cdot \bar{v}^*(p'', \phi)}{g} dp'' \quad (8)$$

159 where p is the air pressure, π is the circular constant, g is the gravitational
160 acceleration.

161 In Eqs. (1)-(8), the overbar and prime denote zonal mean and departure from zonal
162 mean, respectively. The subscripts denote partial derivatives. The Fourier
163 decomposition is used to obtain components of Eqs. (1)-(3) with different zonal wave
164 numbers. Meanwhile, the Fourier decomposed components of geopotential height zonal
165 deviations are also used to determine tropospheric wave sources.

166 c. Statistical methods

167 The trend is measured by the slope of linear regression based on the least square
168 estimation. The correlation is used to analyze statistical links between different
169 variables. In this paper, all the time series have been linearly detrended before
170 calculating correlation coefficients (r) and their corresponding significances.

171 The change-point testing (e.g. Banerjee et al., 2020) is used to make sure the
172 significance of trend or correlation coefficient is not unduly influenced by some
173 particular beginning or ending years, and thereby confirm that the trend exists
174 objectively.

175 We use two-tailed student's t test to calculate the significances of trend, correlation



176 coefficient or mean difference. The result of significance test is measured by p value or
177 confidence intervals in this paper. $p \leq 0.1$, $p \leq 0.05$ and $p \leq 0.01$ suggest the trend,
178 correlation coefficient or mean difference is significant at/above the 90%, 95% and 99%
179 confidence levels, respectively. The confidence interval of trend is shown in (7):

$$180 \quad [\hat{b} - t_{1-\alpha/2}(n-2)\hat{\sigma}_b, \hat{b} + t_{1-\alpha/2}(n-2)\hat{\sigma}_b] \quad (7)$$

181 where \hat{b} is estimated value of slope, $\hat{\sigma}_b$ is standard error of slope and it is written

182 as: $\hat{\sigma}_b = \hat{b} \cdot \sqrt{\frac{1}{n-2} - \frac{r^2}{n-2}}$, $t_{1-\alpha/2}(n-2)$ denotes the value of t-distribution with the degree

183 of freedom equal to $n-2$ and the two-tailed confidence level equal to $1-\alpha$
184 ($\alpha = 0.90, 0.95$ or 0.99). The confidence interval of mean difference is expressed
185 by Eq. (8):

$$186 \quad [\bar{X} - \bar{Y} - t_{1-\alpha/2}(M+N-2) \cdot S_w \cdot \sqrt{\frac{1}{M} + \frac{1}{N}}, \bar{X} - \bar{Y} + t_{1-\alpha/2}(M+N-2) \cdot S_w \cdot \sqrt{\frac{1}{M} + \frac{1}{N}}] \quad (8)$$

187 where

$$188 \quad S_w = \sqrt{\frac{1}{M+N-2} \left[\sum_{i=1}^M (X_i - \bar{X})^2 + \sum_{j=1}^N (Y_j - \bar{Y})^2 \right]} \quad (9)$$

189 Here, \bar{X} and \bar{Y} are the sample averages, M and N are the numbers of sample
190 sizes with two populations, $t_{1-\alpha/2}(M+N-2)$ denotes the value of t-distribution with
191 the degree of freedom equal to $M+N-2$ and the two-tailed confidence level equal
192 to $1-\alpha$.

193 Previous studies have indicated that SST impact on the stratosphere shows a
194 spatial dependence (e.g. Xie et al., 2020). To find out a robust relationship between the
195 trend of SST in a specific region and the trend of stratospheric wave activities, we divide



196 the global ocean into three regions: SH (the extratropical southern hemisphere, 70°S-
197 20°S), TROP (the tropics, 20°S-20°N) and NH (the extratropical northern hemisphere,
198 20°N-70°N). Since the impacts in different regions might be combined, we also
199 consider three combined regions named as SHtrop (the extratropical southern
200 hemisphere and the tropics, 70°S-20°N), NHtrop (the extratropical northern hemisphere
201 and the tropics, 20°S-70°N) and the Globe (70°S-70°N). To find statistical connections
202 between the trend of SST and that of stratospheric wave activities, we examine the first
203 three leading patterns (EOF1, EOF2, EOF3) and principal components (PC1, PC2, PC3)
204 of SST in above six regions obtained from Empirical Orthogonal Function (EOF)
205 analysis. In all the six regions, there is always one EOF modes that shows great
206 similarity to the spatial pattern of trend (not shown) as we do not detrend SST time
207 series when the EOF analysis is carried out. Thus, the significance of the correlation
208 between the PC time series of that EOF mode and time series of stratospheric E-P flux
209 can be used as the criterion to determine the statistical connection between the trend of
210 SST and the trend of stratospheric wave activities.

211 d. The model and experiment configurations

212 The FWSC component in the Community Earth System Model (CESM; version
213 1.2.0) is used to verify the impact of SST and ozone recovery trends on tropospheric
214 wave sources and stratospheric wave activities in early austral spring. The FWSC
215 component is the Whole Atmosphere Community Climate Model version 4 (WACCM4)
216 with specified chemistry forcing fields (such as ozone, greenhouse gases (GHG),
217 aerosols and so on), which have fixed values in 2000 by default. The WACCM4



218 includes active atmosphere, data ocean (run as a prescribed component, simply reading
219 SST forcing data instead of running ocean model), land and sea ice. Important physics
220 schemes in the WACCM4 are based on those in the Community Atmospheric Model
221 version 4 (CAM4; Neale et al., 2013). The WACCM4 uses a finite-volume dynamic
222 framework and extends from the ground to approximately 145 km (5.1×10^{-6} hPa)
223 altitude in the vertical with 66 vertical levels. The simulations presented in this paper
224 are conducted at a horizontal resolution of $1.9^\circ \times 2.5^\circ$. More information about the
225 WACCM can be found in Marsh et al. (2013).

226 Control experiments and sensitive experiments are conducted to investigate
227 responses of Antarctic stratospheric wave activities to SST trend and the ozone recovery
228 trend in early austral spring. For the experiments of SST trends, monthly mean global
229 SST during 1980-2000 derived from the ERSST v5 dataset is used as SST forcing field
230 in the control experiment (sstctrl). For the four sensitive experiments (sstNH, sstSH,
231 ssttrop, sstSHtrop), linear increments of SST in different regions in September during
232 2000-2017 are used as the forcing field. Ozone, aerosols and greenhouse gases (GHG)
233 in the control experiment and the four sensitive experiments all have the fixed values
234 in 2000. For the experiments of ozone recovery trend, monthly mean three-dimensional
235 global ozone during 1980-2000 derived from the MERRA-2 dataset is used as the ozone
236 forcing field in the control experiment (O3ctrl). The sensitive experiment (O3sen) is
237 forced by linear increments of ozone in September during 2001-2017. The SSTs in
238 O3ctrl and O3sen both are monthly mean global SST during 1980-2000. The aerosol
239 and greenhouse gases values in 2000 are used. These experiment configurations are



240 summarized and listed in Table 1 and Table 2.

241 First, we run the FWSC component to generate randomly different initial
242 conditions for 120 years with free run. Then, each experiment includes 100 ensemble
243 members that run from July to September forced by these initial conditions from the
244 21st year to the 120th year in July. The forcing fields of SST and ozone are only
245 superposed from July to September. July and August are taken as spin-up time and
246 simulations during this period are discarded. The ensemble mean in September derived
247 from these 100 ensemble members is regarded as the final result of each experiment. A
248 similar approach is implemented for sensitive experiments, in which the forcing fields
249 superposed only in certain months. The same approach has been used in previous
250 studies (e.g., Zhang et al., 2018).

251 **3. Trend of planetary wave activities in early austral spring**

252 Figure 1 shows the trends of stratospheric planetary wave activities in the southern
253 hemisphere September during 1980-2000 and 2000-2017, respectively. Note that the
254 vertical E-P flux entering into the stratosphere over 50°S-70°S in September has been
255 increasing during 1980-2000, accompanied by intensified wave flux convergence in the
256 upper stratosphere (Fig. 1a) that is mainly contributed by the wave-1 component (Fig.
257 1b). This feature implies that the stratospheric planetary wave activities have
258 strengthened in early austral spring during 1980-2000. A similar result has been
259 reported in previous studies (Hu & Fu, 2009; Lin et al., 2009). During 2000-2017,
260 however, vertical transport of stratospheric E-P flux weakened over the subpolar region
261 of the southern hemisphere, which was accompanied by intensified wave flux



262 divergence in the upper stratosphere (Fig. 1d) mainly contributed by the wave-1
263 component (Fig. 1e) while the wave-2 component also made certain contributions (Fig.
264 1f). Similar features also appear in August, but not as significant as that in September
265 (Fig. S1). For this reason, hereafter we focus on the features in September.

266 The SSW that occurred in 2002 was accompanied with large upward wave fluxes
267 in the stratosphere, which is extremely rare in history and has been studied in numerous
268 previous studies (e.g., Baldwin et al., 2003; Nishii & Nakamura, 2004; Newman &
269 Nash, 2005). Since the period with a negative trend of stratospheric vertical wave flux
270 is short, it is necessary to further investigate whether such a negative trend is artificially
271 influenced by the single year of 2002. Therefore, following Banerjee et al. (2020), we
272 use a change-point method to test the significance of the trend during various periods
273 based on four reanalysis datasets (ERA-Interim, MERRA-2, JRA-55, NCEP-2).
274 Figures 2a (including the year 2002) and 2b (excluding the year 2002) display the time
275 series (F_z) of area-weighted vertical stratospheric wave flux over the southern
276 hemisphere subpolar region obtained from different reanalysis datasets. Note that the
277 wave flux time series obtained from the four reanalysis datasets all present a positive
278 trend from the early 1980s to the early 2000s and a negative trend from the early 2000s
279 to present, regardless of whether the extreme value in 2002 is removed or not. The
280 correlation coefficients of the time series between these reanalysis datasets are above
281 0.9 and statistically significant (Table 3), suggesting that the time series derived from
282 different datasets are consistent with each other. Figures 2c-f show the trends and
283 corresponding confidence intervals calculated with four different beginning years (1980,



284 1981, 1982, 1983), four different ending years (2015, 2016, 2017, 2018), and change-
285 point years from 1998 to 2013. The trends and confidence intervals in Figures 2g-j are
286 the same as that in Figures 2c-f, except that the extreme value in 2002 is removed. The
287 positive trend from the early 1980s to the 21st century remains significant regardless of
288 different beginning years and ending change-point years (Figs. 2c-j). However, Figures
289 2c-f and Figures 2g-j indicate that the positive value of the trend is decreasing gradually
290 when the period is prolonged, which is apparently attributed to the negative trend with
291 the beginning change-point year of around 2000. Although the negative trend from the
292 change-point year to ending year becomes less significant when the value in 2002 is
293 removed, it remains significant in some periods, which are also illustrated on diagrams
294 of latitude-pressure profiles (Fig. S2). Therefore, the weakening of stratospheric wave
295 activities in early austral spring since the early 2000s is robust. In this paper, we take
296 the year 2000 as the beginning year of the weakening trend to simplify descriptions in
297 the following discussion.

298 Figure 3 shows the trends of tropospheric wave sources in September since 2000.
299 There is a significant positive trend of the wave-1 component in 500 hPa geopotential
300 height over the southern Indian ocean and a significant negative trend over the southern
301 Pacific, which form an out-of-phase superposition on its climatology (Fig. 3b). The
302 trend pattern of wave-2 component is also out-of-phase with its climatology, although
303 it is not significant (Fig. 3c). The above features still maintain when the values in 2002
304 are removed (Fig. S3b, c), implying that the southern hemispheric tropospheric wave
305 sources in early austral spring have weakened since 2000, which is also reflected in the



306 decrease of tropospheric vertical wave flux (Fig. 3d, e; Fig. S3d, e).

307 **4. Role of SST trends in the weakening of Antarctic stratospheric**
308 **wave activities**

309 In this section, we further explore factors that lead to the weakening of
310 tropospheric wave sources and stratospheric wave activities since the early 2000s in
311 early austral spring. Numerous studies reported that the variations in sea surface
312 temperature can affect stratospheric climate (e.g., Li, 2009; Hurwitz et al., 2011; Lin et
313 al., 2012; Hu et al., 2014; Hu et al., 2018; Tian et al., 2018; Xie et al., 2020). Hu & Fu
314 (2009) also attributed the strengthened stratospheric wave activities in the southern
315 hemisphere to SST trend from the early 1980s to the early 2000s. Furthermore, global
316 SST in September during 2000-2017 also has a significant trend. The significant
317 warming pattern is mainly found over the southern Indian ocean, the southern Atlantic
318 ocean, the eastern and western equatorial Pacific, the western equatorial and Northern
319 Atlantic ocean (Fig. 4b). A significant cooling pattern is located over the southeast
320 Pacific (Fig. 4b). In a word, the spatial pattern of SST trend during 2000-2017 is
321 obviously different from that during 1980-2000 (Fig. 4a, b). Thus, it is necessary to
322 analyze the connection between SST trend and wave activity trend since the early 2000s.

323 Figure 5 shows the significance of principle component (PC) trends (Figs. 5a-f) of
324 SST in different regions, and the significance of correlations (Figs. 5g-l) between the
325 PC time series and Fz during various periods in September. The trend of PC1 time series
326 in SH region is significant during several periods (Fig. 5a), while the correlation between
327 PC1 and Fz is only significant with the particular ending year of 2015 (Fig. 5g). This



328 feature suggests that the connection between the SST trend in SH region and the trend
329 of stratospheric wave activity is not robust. The correlation between trend of
330 stratospheric wave activity and that of SST in TROP or NH region is also weak (Fig.
331 5e, f). As for the combined regions, note that the PC2 time series in SHtrop region has
332 a significant trend (Fig. 5d) and the correlation between the PC2 time series in SHtrop
333 and Fz with the beginning year of around 2000 is also significant (Fig. 5j) regardless of
334 different ending years. This feature implies that the extratropical southern hemisphere
335 and tropical SST has a robust connection with stratospheric wave activities in early
336 austral spring since the early 2000s. The correlations between Fz and all PC time series
337 in NHtrop (Fig. 5k) and Globe (Fig. 5l) region are not as robust as that between Fz and
338 PC2 time series in SHtrop region (Fig. 5j), indicating that the connection between SST
339 trend in extratropical northern hemisphere and the trend of stratospheric wave activity
340 is weak.

341 Figure 6 shows the first three EOF modes of September SST in SHtrop region
342 during 2000-2017. The second mode (Fig. 6b) shows a great similarity to the spatial
343 pattern of SST trend (Fig. 4b), and the corresponding PC2 time series also has a
344 significant trend (slope=1.71, $p < 0.01$). The correlation between PC2 and the Fz is
345 significant ($r = -0.56$, $p = 0.016$) and the correlation coefficient remains significant ($r = -$
346 0.46 , $p = 0.065$) at the 90% confidence level when the value in 2002 is removed. This
347 result suggests that the SST trend in SHtrop region is closely related to the recent
348 weakening of stratospheric wave activities. The first EOF mode is similar to IPO (Fig.
349 6a) and its corresponding principal component is highly significantly correlated ($r = -$



350 0.98, $p < 0.01$) with the unfiltered IPO index. However, it shows no significant trend (Fig.
351 6d) and has no significant correlation (Fig. 6g) with stratospheric wave flux, implying
352 that the linkage between the IPO phase change at around 2000 (e.g. Trenberth et al.,
353 2013) and the weakening of Antarctic stratospheric wave activities is weak. The
354 correlation between PC3 and Fz is also not significant (Fig. 6i). Therefore, it is possible
355 that the combined effect of SST trend (the second EOF mode) in the tropical and
356 extratropical southern hemisphere leads to the weakening of stratospheric wave
357 activities in early austral spring since the early 2000s.

358 **5. Simulated changes in Antarctic stratospheric wave activities forced** 359 **by SST trends**

360 The analysis in Section 4 suggests that the SST trend in SHtrop region may
361 contribute to the weakening of the southern hemispheric stratospheric wave activities.
362 Here, numerical experiments sstNH, sstSH, ssttrop and sstSHtrop forced by linear
363 increments of SST in September during 2000-2017 (Fig. 7; more details can be found
364 in Section 2) are conducted to verify the results discussed in Section 4.

365 Figure 8 shows the simulated response of 500 hPa geopotential height to SST
366 changes in different regions. The climatological distributions of the wave-1 component
367 (Figs. 8b, e, h, k) and the wave-2 component (Figs. 8c, f, i, l) from the simulations are
368 consistent with that from reanalysis dataset (Figs. 4b, c), indicating that the model can
369 well capture spatial distributions of the atmospheric waves. Note that the Fourier
370 component (wave-1 and wave-2) anomalies simulated with SST changes in SH, TROP
371 and SHtrop all are significant. They superpose on the corresponding climatological



372 patterns in an out-of-phase style (Figs. 8e, f, h, i, k, l), indicating that the SST trends in
373 SH, TROP and SHtrop lead to a weakening of tropospheric wave sources in the
374 extratropical southern hemisphere. However, the 500 hPa geopotential height anomaly
375 of the predominate wave-1 component in the extratropical southern hemisphere forced
376 by the experiment with NH SST change is relatively weak (Fig. 8b). This feature
377 suggests that the SST trend in extratropical northern hemisphere is incapable of
378 inducing a robust response of tropospheric wave sources in the extratropical southern
379 hemisphere.

380 Figure 9 shows the simulated responses of stratospheric wave activities in the
381 southern hemisphere to SST changes in different regions. It is found that the
382 experiments with SST changes in SH, TROP and SHtrop show significantly weakened
383 stratospheric wave activities (Figs. 9d, g, j), which are mainly attributed to the responses
384 of the wave-1 component (Figs. 9e, h, k). These results are consistent with the responses
385 of tropospheric wave sources (Figs. 8d, e, g, h, j, k). However, there are no significant
386 anomalies of stratospheric wave flux in the subpolar region as exhibited in Figures 9a
387 and 9b, which is consistent with the response of corresponding tropospheric wave
388 sources (Fig. 8a, b) and the weak correlation between Fz and PC time series of SST in
389 NH region (Fig. 5i). It suggests that the response of southern hemisphere stratospheric
390 wave activities to SST trend in NH region is weak.

391 Results of all these experiments are summarized and displayed in Figure 10, which
392 is quantified by the frequency distribution of southern hemisphere stratospheric vertical
393 wave flux derived from the 100 ensemble members of each experiment. Compared to



394 the blue fitting curves, the red fitting curves shift to the left as shown in Figs. 10b, 10c
395 and 10d, suggesting that the SST changes in SH, TROP and SHtrop regions weaken the
396 upward propagation of stratospheric wave flux. The area-weighted anomalies of
397 vertical E-P flux in the subpolar region of the southern hemisphere induced by SST
398 changes in SH, TROP and SHtrop regions are $-0.084 \times 10^5 \text{ kg} \cdot \text{s}^{-2}$, $-0.12 \times 10^5 \text{ kg} \cdot \text{s}^{-2}$ and
399 $-0.13 \times 10^5 \text{ kg} \cdot \text{s}^{-2}$, respectively. The sum of the anomalies forced by sstSH and ssttrop is
400 not equal to the anomaly forced by sstSHtrop, which may be resulted from non-linear
401 interactions between the responses of wave activities to SST trends in SH region and
402 TROP region. The weakening of stratospheric wave activities forced by SST increment
403 in the tropical region is more obvious and more significant than that in extratropical
404 southern hemisphere (Figs. 10b, c, e), implying that the SST trend in the tropical region
405 contributes more to the weakening of stratospheric wave activities since 2000.
406 Meanwhile, it is apparent that the weakening of the southern hemisphere stratospheric
407 wave activities forced by sstSHtrop is the most significant among all the sensitive
408 experiments (Fig. 10e). The reduction of vertical E-P flux over (50°S - 70°S , 200 hPa-
409 10 hPa) forced by sstSHtrop is approximately 12%. These simulation results indicate
410 that the weakening of the Antarctic stratospheric wave activities in September since
411 2000 is induced by the combined effects of SST trends in the tropical and extratropical
412 southern hemisphere. It also explains why the independent correlation between Fz and
413 PC obtained for SH or TROP region is not as significant as that between Fz and PC
414 obtained for SHtrop region (Figs. 5g, h, j). Moreover, the mean linear increment of area-
415 weighted vertical E-P flux from 200 hPa to 10 hPa over 70°S - 50°S in September during



416 2000-2017 derived from four reanalysis datasets is about $-0.38 \times 10^5 \text{ kg} \cdot \text{s}^{-2}$. Therefore,
417 the contribution of SST trend over 20°N - 70°S (the SHtrop region) to the weakening of
418 stratospheric activities is approximately 34%.

419 **6. Conclusions and Discussion**

420 This study analyzes the trend of Antarctic stratospheric planetary wave activities
421 in early austral spring since the early 2000s based on various reanalysis datasets. Using
422 the change-point method, we find that the Antarctic stratospheric wave activities in
423 September have been weakening significantly since 2000, which means the intensified
424 trend of wave activities noted in previous researches (Hu & Fu, 2009; Lin et al., 2009)
425 are reversed after 2000 in early austral spring. Further analysis suggests that the
426 weakening of stratospheric wave activities is related to the weakening of tropospheric
427 wave sources in extratropical Southern Hemisphere, which is mainly contributed by the
428 wave-1 component. Moreover, EOF analysis and correlation analysis indicate that the
429 stratospheric wave activities in early austral spring during 2000-2017 are related to PC2
430 of SST over 20°N - 70°S (i.e., the SHtrop region). The corresponding EOF2 mode also
431 shows a great similarity to the spatial pattern of SST trend, suggesting that the
432 weakening of stratospheric wave activities is connected to the trend of SST in SHtrop
433 region. Meanwhile, the linkage between the SST trend in NH region and the weakening
434 of stratospheric wave activities is weak. Finally, the model simulations support the
435 conclusion that the SST changes in SHtrop region lead to the weakening of tropospheric
436 wave sources and stratospheric wave activities. The contribution of SST trend in
437 tropical region to the weakening of stratospheric wave activities is larger than that in



438 the extratropical southern hemisphere. However, the response of tropospheric wave
439 sources and stratospheric wave activities to SST trend in NH region is not significant.
440 The contribution of SST trend over SHtrop region to the weakening of stratospheric
441 wave activities is about 34%.

442 The question that remains answered is whether the ozone recovery trend also
443 contributes to the weakening of stratospheric wave activities in September since the
444 early 2000s. As described in Section 2, a control experiment (O3ctrl) forced by
445 climatological ozone and a sensitive experiment forced by the linear increment of
446 global ozone in September during 2001-2017 are conducted to address the above
447 question. The pattern of ozone forcing field is similar to its trend pattern (Figs. S4c, d;
448 Fig. S5). We choose the period of 2001-2017 because we notice that the ozone recovery
449 trend derived from MERRA-2 in September with the beginning year of 2000 is not
450 significant (Fig. S4a, b). Meanwhile, as the SSW in 2002 induces poleward transport
451 of large amounts of ozone, the data in 2002 are removed when linear increments are
452 calculated. Other details about these two experiments have been given in Section 2 and
453 Table 2. The simulated results indicate that there is no significant response of wave flux
454 (Fig. 11a, d) as well as its Fourier decomposed components (Fig. 11b, c) over southern
455 hemisphere subpolar region in the stratosphere, suggesting that the prescribed ozone
456 recovery is incapable of inducing the weakening of stratospheric wave activities.

457 Many researchers claimed that the climate transition around 2000 in the southern
458 hemisphere is related to ozone depletion and recovery (e.g., Barnes et al., 2013;
459 Banerjee et al., 2020). Note that there is no contradiction between our results and these



460 previous studies. Firstly, the southern hemisphere tropospheric circulation (i.e., the
461 SAM index, the tropospheric jet position and the Hadley cell edge) transition related to
462 ozone depletion and recovery reported in these previous studies basically occurred in
463 austral summer (e.g., Son et al., 2008; Thompson et al., 2011; Barnes et al., 2013;
464 Banerjee et al., 2020). These tropospheric circulation transitions are induced by
465 downward coupling of circulation anomalies in the stratosphere (e.g., Thompson et al.,
466 2011) during October and November, when solar radiation covers the entire Antarctic
467 and causes radiative heating effects. However, we focus on September in the present
468 study. The Antarctic stratospheric circulation response to ozone variation in September
469 is not as strong as that in October or November (e.g., Thompson et al., 2011, Fig. 1b, d)
470 because solar radiation can only reach part of the Antarctic stratosphere during a
471 majority period of September. This fact implies that the response of wave propagation
472 environment in the Antarctic stratosphere to ozone trend is also not significant (Fig. S6).
473 Secondly, the FWSC component used in this study is an atmospheric module with
474 prescribed SST and gases. Therefore, the model results only indicate that the weakening
475 of stratospheric wave activities can be attributed to SST trends, while the impact of
476 ozone depletion and recovery trend in the tropics and mid-latitudes on the shift of SST
477 trend pattern cannot be determined based on the model simulations. This is an issue
478 beyond the scope of this study and further investigation is necessary using a fully
479 coupled earth system model.

480 In addition, the reanalysis datasets show that the Brewer-Dobson circulation
481 related to wave activities in the stratosphere weakened significantly in early austral



482 spring during 2000-2017 (Fig. 12b), which is contrary to the intensified trend during
483 1980-2000 (Fig. 12a). The transition of BDC around 2000 is believed to be associated
484 with ozone depletion and recovery (e.g., Polvani et al., 2017; Polvani et al., 2018).
485 However, our modeling results suggest that the SST trend is responsible for the
486 weakening of BDC in September since 2000 (Fig. 12d, e, f), The response of BDC to
487 ozone recovery is not significant (Fig. 12c), especially for the branch near the Antarctic.
488 These results indicate that the SST trend should be taken into consideration when
489 exploring the mechanism for the climate transition in the southern hemispheric
490 stratosphere around 2000.

491

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501

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705

706 **Table 1.** Configurations of experiments for SST trends.

Experiments	Descriptions
sstctrl	Control run. Seasonal cycle of monthly mean global SST data over 1980-2000 is derived from the ERSST v5 dataset. Fixed values of ozone greenhouse gases and aerosol fields in 2000 are used.
sstNH	As in sstctrl, but with linear increments of SST in September over 2000-2017 in NH (20°N-70°N). The applied global SST anomalies are shown in Fig. 7a.
sstSH	As in sstctrl, but with linear increments of SST in September over 2000-2017 in SH (20°S-70°S). The applied global SST anomalies are shown in Fig. 7b.
ssttrop	As in sstctrl, but with linear increments of SST in September over 2000-2017 in the tropics (20°S-20°N). The applied global SST anomalies are shown in Fig. 7c.
sstSHtrop	As in sstctrl, but with linear increments of SST in September over 2000-2017 in SHtrop (20°N-70°S). The applied global SST anomalies are shown in Fig. 7d.

707 **Table 2.** Configurations of experiments for the ozone recovery trend.

Experiments	Descriptions
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O3ctrl	Control run. The seasonal cycle of monthly averaged global SST data over 1980-2000 is derived from ERSST v5 dataset. The seasonal cycle of monthly mean three-dimensional global ozone over 1980-2000 is derived from MERRA-2 dataset. The GHGs and aerosol fields are specified to be fixed values in 2000.
O3sen	As in O3ctrl, but superposed with linear increments of global ozone in September over 2001-2017. The ozone data in 2002 are removed when the linear increments are calculated. The applied ozone anomalies in Southern Hemisphere are shown in Fig. S5.

708 **Table 3.** Correlations of stratospheric vertical wave flux time series (area-weighted
 709 from 100 hPa to 30 hPa over 70°S-50°S) between different reanalysis dataset.

	ERA-Interim	JRA-55	MERRA-2	NCEP-2
ERA-Interim	1.00 (p=0.00)	0.99 (p<0.01)	0.98 (p<0.01)	0.93 (p<0.01)
JRA-55		1.00 (p=0.00)	0.98 (p<0.01)	0.93 (p<0.01)
MERRA-2			1.00 (p=0.00)	0.94 (p<0.01)
NCEP-2				1.00 (p=0.00)

710

711 **Figure captions:**

712 **FIG. 1.** Trends of southern hemisphere (a, d) stratospheric E-P flux (arrows, units of
 713 horizontal and vertical components are 10^5 and $10^3 \text{ kg} \cdot \text{s}^{-2}$ per year, respectively) and its
 714 divergence (shadings) with their (b, e) wave-1 components and (c, f) wave-2
 715 components over (a, b, c) 1980-2000 and (d, e, f) 2000-2017 in September derived from
 716 MERRA-2 dataset. The stippled regions indicate the trend of E-P flux divergence
 717 significant at/above the 90% confidence level. The green contours from outside to
 718 inside (corresponding to $p=0.1, 0.05$) indicate the trend of vertical E-P flux significant



719 at the 90% and 95% confidence level, respectively.

720 **FIG. 2.** (a) The mean time series (solid line) and piecewise (during 1980-2000 and
721 2000-2018) linear regressions (dashed lines) of vertical E-P flux area-weighted from
722 100 hPa to 30 hPa over 70°S-50°S in September during 1980-2018 derived from ERA-
723 Interim (yellow), MERRA-2 (blue), JRA-55 (red) and NCEP-2 (green). Figure (b) is
724 the same as Figure (a), except for that the data in 2002 are removed. (c, d, e, f) The
725 trends (dots) and uncertainties (error bars) calculated during various periods using the
726 change-point method with different beginning and ending years (titles). Circles and
727 squares in Figures (c, d, e, f) represent positive trends from beginning years to change-
728 point years (x-axes) and negative trends from change-point years to ending years,
729 respectively. Different colors of dots and error bars in Figures (c, d, e, f) correspond to
730 colors in Figure (a), which represent trends and uncertainties derived from different
731 datasets. The long and short error bars in same color reflect the 95% and 90%
732 confidence intervals calculated by two-tailed t test. The error bar is omitted when the
733 significance of trend is lower than corresponding confidence level. Negative trends and
734 corresponding uncertainties with the beginning change-point years after 2005 are also
735 omitted, since the trend value shows large fluctuation with shortening of time series.
736 Figures (g, h, i, j) are the same as Figures (c, d, e, f), except that the data in 2002 are
737 removed when calculating trends and uncertainties.

738 **FIG. 3.** Trends (shadings) and climatological distributions (contours with an interval
739 of 20 gpm, positive and negative values are depicted by solid and dashed lines
740 respectively, zeroes are depicted by thick solid lines) of southern hemispheric (a) 500



741 hPa geopotential height zonal deviations with their (b) wave-1 component and (c)
742 wave-2 component in September during 2000–2017 derived from MERRA-2 dataset.
743 Trends of southern hemispheric (d) tropospheric E-P flux (arrows, units of horizontal
744 and vertical components are 3×10^5 and 3×10^3 kg s⁻² per year, respectively) and its
745 vertical component (shading) with their (e) wave-1 component and (f) wave-2
746 component in September during 2000–2017 derived from MERRA-2 dataset. The
747 stippled regions represent the trend significant at/above the 90% confidence level.

748 **FIG. 4.** Trends of SST in September over (a) 1980-2000 and (b) 2000-2017 derived
749 from ERSST v5 dataset. The stippled regions represent the trends significant at/above
750 the 90% confidence level.

751 **FIG. 5.** Trend significance of the first three SST principal components (PCs) in (a) the
752 extratropical southern hemisphere (SH, 70°S-20°S), (b) the tropics (TROP, 20°S-20°N),
753 (c) the extratropical northern hemisphere (NH, 20°N-70°N), (d) the extratropical
754 southern hemisphere and the tropics (SHtrop, 70°S-20°N), (e) the extratropical northern
755 hemisphere and the tropics (NHtrop, 20°S-70°N), (f) the globe (70°S-70°N) and the
756 corresponding (g, h, i, j, k, l) correlation significances between them and vertical E-P
757 flux (F_z , area-weighted from 100 hPa to 30 hPa over 70°S-50°S) during different
758 beginning years (x-axes) and ending years (y-axes). The red and blue dots indicate
759 positive and negative trend or correlation coefficient are significant, respectively. The
760 black dots indicate the trends or correlation coefficients are not significant. The stars
761 indicate that the trends and the corresponding correlation coefficients are both
762 significant. Each panel is divided into three regions from bottom to top, corresponding



763 to the first, the second and the third principal components, respectively. The criterion
764 to distinguish whether the trends and correlations are significant or not is the 90%
765 confidence level.

766 **FIG. 6.** (a, b, c) The first three EOF patterns of SST in SHtrop region. (d, e, f) The
767 original time series of the first three principle components (PCs, blue solid lines
768 correspond to left inverted y-axes) and stratospheric vertical E-P flux (Fz, area-
769 weighted from 100 hPa to 30 hPa over 70°S-50°S, red solid lines correspond to right y-
770 axes) in September during 2000-2017. The blue and red dashed lines in (d, e, f)
771 represent the linear regressions of PC time series and Fz time series, respectively. The
772 meaning of (g, h, i) are the same as (d, e, f) correspondingly, except the detrended time
773 series. The unbracketed and bracketed numbers in (g, h, i) represent the correlation
774 coefficients between detrended PC time series and Fz time series and the corresponding
775 p values calculated by two-tailed t test, respectively.

776 **FIG. 7.** Differences in SST forcing field between sensitive experiments ((a) sstNH; (b)
777 sstSH; (c) ssttrop; (d) sstSHtrop) and the control experiment (sstctrl).

778 **FIG. 8.** Differences (shadings) of (a, d, g, j) 500 hPa geopotential height zonal
779 deviations with their (b, e, h, k) wave-1 component and (c, f, i, l) wave-2 component
780 between sensitive experiments ((a, b, c) sstNH; (d, e, f) sstSH; (g, h, i) ssttrop; (j, k, l)
781 sstSHtrop) and the control experiment (sstctrl). The mean distributions (contours with
782 an interval of 20 gpm, positive and negative values are depicted by solid and dashed
783 lines respectively, zeroes are depicted by thick solid lines) of them are derived from the
784 control experiment. The stippled regions represent the mean difference significant



785 at/above the 90% confidence level.

786 **FIG. 9.** Differences of (a, d, g, j) stratospheric E-P flux (arrows, units in horizontal and
787 vertical components are 0.05×10^7 and $0.05 \times 10^5 \text{ kg} \cdot \text{s}^{-2}$, respectively) and its divergence
788 (shadings) with their (b, e, h, k) wave-1 component and (c, f, i, l) wave-2 component
789 between sensitive experiments ((a, b, c) sstNH; (d, e, f) sstSH; (g, h, i) ssttrop; (j, k, l)
790 sstSHtrop) and the control experiment (sstctrl). The stippled regions represent the mean
791 differences of E-P flux divergence significant at/above the 90% confidence level. The
792 green contours from outside to inside (corresponding to $p=0.1, 0.05$) represent the mean
793 differences of vertical E-P flux significant at the 90% and 95% confidence levels,
794 respectively.

795 **FIG. 10.** (a, b, c, d) Frequency distributions (pillars, blue for control experiment and
796 orange for sensitive experiments) of vertical E-P flux (F_z , area-weighted from 200 hPa
797 to 10 hPa over 70°S - 50°S) and its 5-point low-pass filtered fitting curves (solid lines,
798 blue for control experiment and red for sensitive experiments) derived from 100
799 ensemble members of the control experiment (sstctrl) and sensitive experiments ((a)
800 sstNH; (b) sstSH; (c) ssttrop; (d) sstSHtrop), respectively. (e) Mean differences (grey
801 pillars) and corresponding uncertainties (error bars) of F_z between sensitive
802 experiments and the control experiment. The blue and red error bars reflect the 90%
803 and 95% confidence levels calculated by two-tailed t test, respectively. The error bar is
804 omitted when the significance of mean difference is lower than the corresponding
805 confidence level.

806 **FIG. 11.** Differences of (a) stratospheric E-P flux (arrows, units in horizontal and



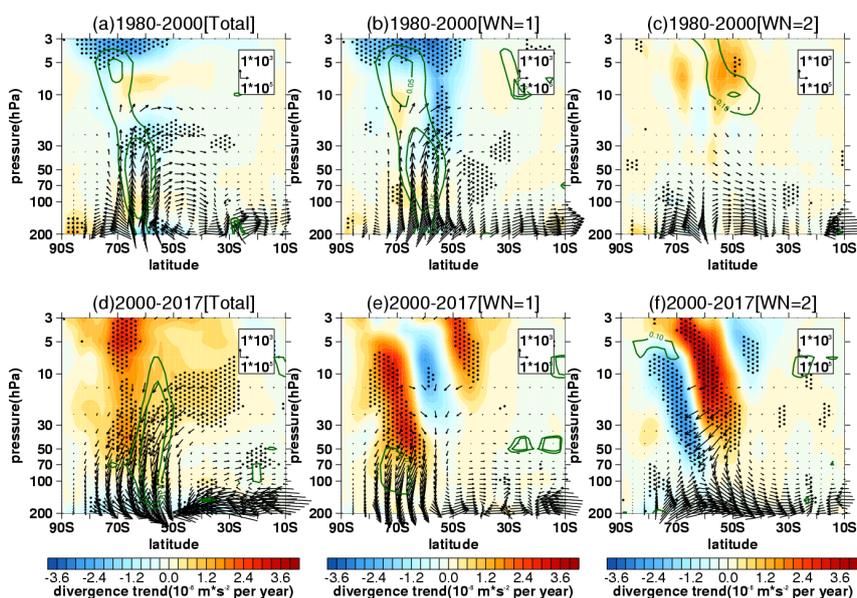
807 vertical components are 0.02×10^7 and $0.05 \times 10^5 \text{ kg} \cdot \text{s}^{-2}$, respectively) and its divergence
808 (shadings) with their (b) wave-1 component and (c) wave-2 component between the
809 sensitive experiment (O3sen) and the control experiment (O3ctrl). The stippled regions
810 represent the mean differences of E-P flux divergence significant at/above the 90%
811 confidence level. The green contours from outside to inside (corresponding to $p=0.1$,
812 0.05) represent the mean differences of vertical E-P flux significant at the 90% and 95%
813 confidence levels, respectively. (d) Frequency distributions (pillars, blue for O3ctrl and
814 orange for O3sen) of vertical E-P flux (F_z , area-weighted from 200 hPa to 10 hPa over
815 70°S - 50°S) and its 5-point low-pass filtered fitting curves (solid lines, blue for O3ctrl
816 and red for O3sen) derived from 100 ensemble members.

817 **FIG. 12.** (a) Trends of southern hemispheric Brewer-Dobson circulation (arrows, units
818 in horizontal and vertical components are 0.2×10^{-2} and $0.2 \times 10^{-4} \text{ m} \cdot \text{s}^{-1}$ per year,
819 respectively) and its stream function (shadings) in September during (a) 1980-2000 and
820 (b) 2000-2017 derived from MERRA-2 dataset. Data in 2002 are removed when trends
821 are calculated in Figure (b). (c) Differences of Brewer-Dobson circulation (arrows,
822 units in horizontal and vertical components are 10^{-2} and $10^{-4} \text{ m} \cdot \text{s}^{-1}$, respectively) and its
823 stream function (shadings) between the O3ctrl and O3sen. (d, e, f) Differences of
824 Brewer-Dobson circulation and its stream function between the control experiment
825 (sstctrl) and various sensitive experiments ((d) sstSH; (e) ssttrop; (f) sstSHtrop) with
826 SST changes. The stippled regions represent the trends or differences of the stream
827 function significant at/above the 90% confidence level. The green contours from
828 outside to inside (corresponding to $p=0.1$, 0.05) represent the trends or differences of



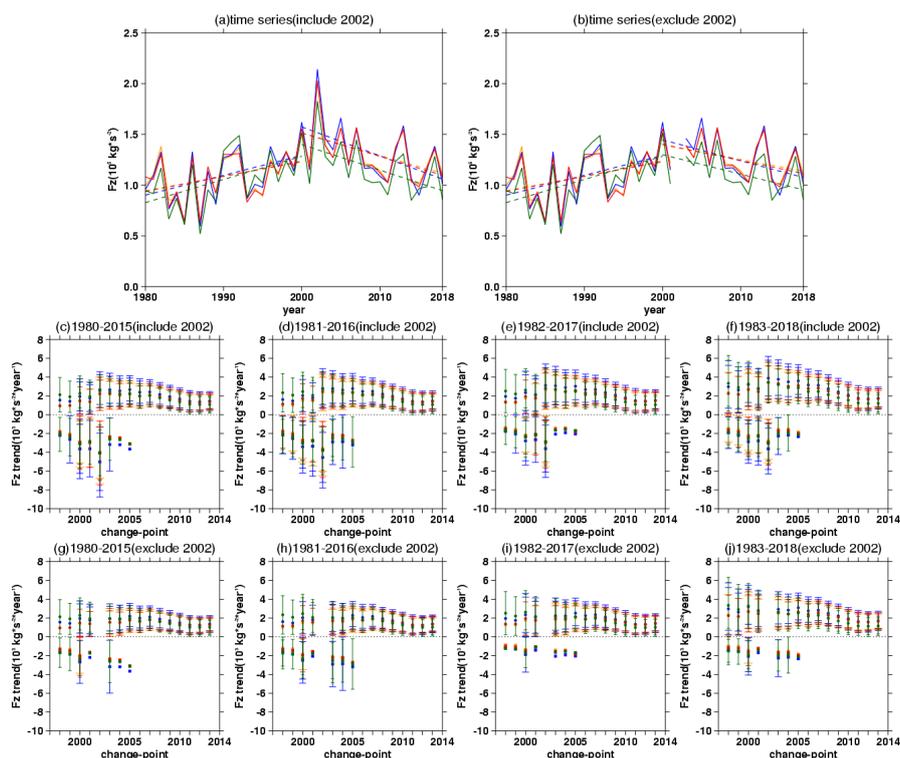
829 the vertical components significant at the 90% and 95% confidence levels, respectively.

830



831

832 **FIG. 1.** Trends of southern hemisphere (a, d) stratospheric E-P flux (arrows, units of
833 horizontal and vertical components are 10^5 and 10^3 $\text{kg}\cdot\text{s}^{-2}$ per year, respectively) and its
834 divergence (shadings) with their (b, e) wave-1 components and (c, f) wave-2
835 components over (a, b, c) 1980-2000 and (d, e, f) 2000-2017 in September derived from
836 MERRA-2 dataset. The stippled regions indicate the trend of E-P flux divergence
837 significant at/above the 90% confidence level. The green contours from outside to
838 inside (corresponding to $p=0.1$, 0.05) indicate the trend of vertical E-P flux significant
839 at the 90% and 95% confidence level, respectively.

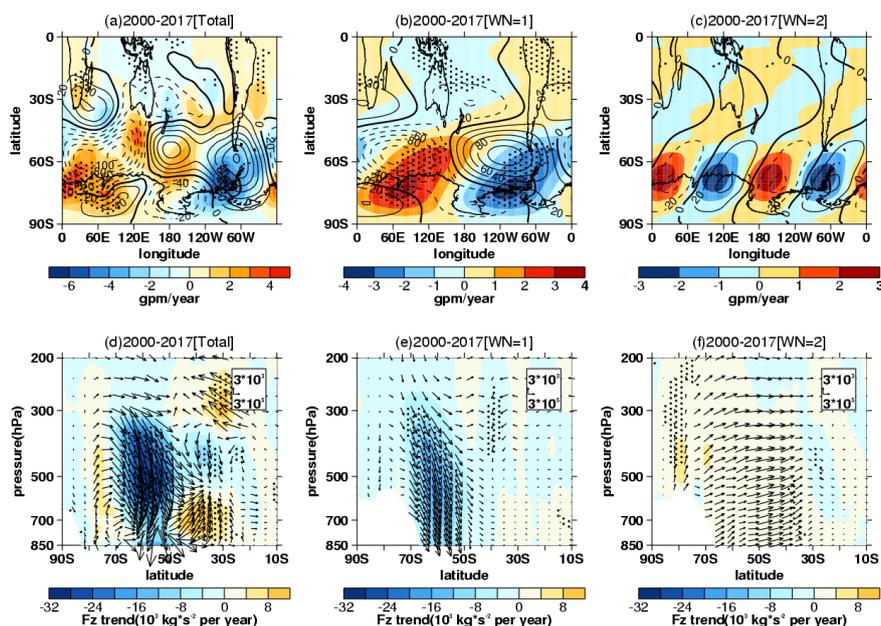


840

841 **FIG. 2.** (a) The mean time series (solid line) and piecewise (during 1980-2000 and
842 2000-2018) linear regressions (dashed lines) of vertical E-P flux area-weighted from
843 100 hPa to 30 hPa over 70°S-50°S in September during 1980-2018 derived from ERA-
844 Interim (yellow), MERRA-2 (blue), JRA-55 (red) and NCEP-2 (green). Figure (b) is
845 the same as Figure (a), except for that the data in 2002 are removed. (c, d, e, f) The
846 trends (dots) and uncertainties (error bars) calculated during various periods using the
847 change-point method with different beginning and ending years (titles). Circles and
848 squares in Figures (c, d, e, f) represent positive trends from beginning years to change-
849 point years (x-axes) and negative trends from change-point years to ending years,
850 respectively. Different colors of dots and error bars in Figures (c, d, e, f) correspond to
851 colors in Figure (a), which represent trends and uncertainties derived from different



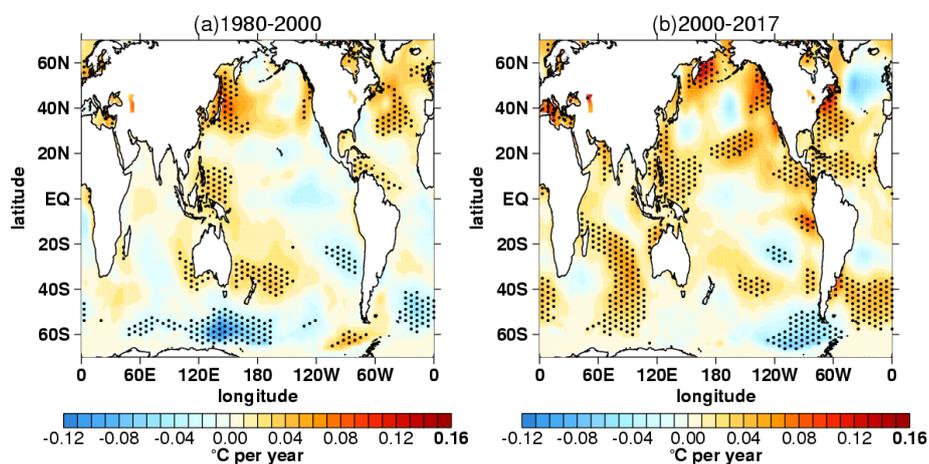
852 datasets. The long and short error bars in same color reflect the 95% and 90%
853 confidence intervals calculated by two-tailed t test. The error bar is omitted when the
854 significance of trend is lower than corresponding confidence level. Negative trends and
855 corresponding uncertainties with the beginning change-point years after 2005 are also
856 omitted, since the trend value shows large fluctuation with shortening of time series.
857 Figures (g, h, i, j) are the same as Figures (c, d, e, f), except that the data in 2002 are
858 removed when calculating trends and uncertainties.



859 **FIG. 3.** Trends (shadings) and climatological distributions (contours with an interval
860 of 20 gpm, positive and negative values are depicted by solid and dashed lines
861 respectively, zeroes are depicted by thick solid lines) of southern hemispheric (a) 500
862 hPa geopotential height zonal deviations with their (b) wave-1 component and (c)
863 wave-2 component in September during 2000–2017 derived from MERRA-2 dataset.
864 Trends of southern hemispheric (d) tropospheric E-P flux (arrows, units of horizontal
865

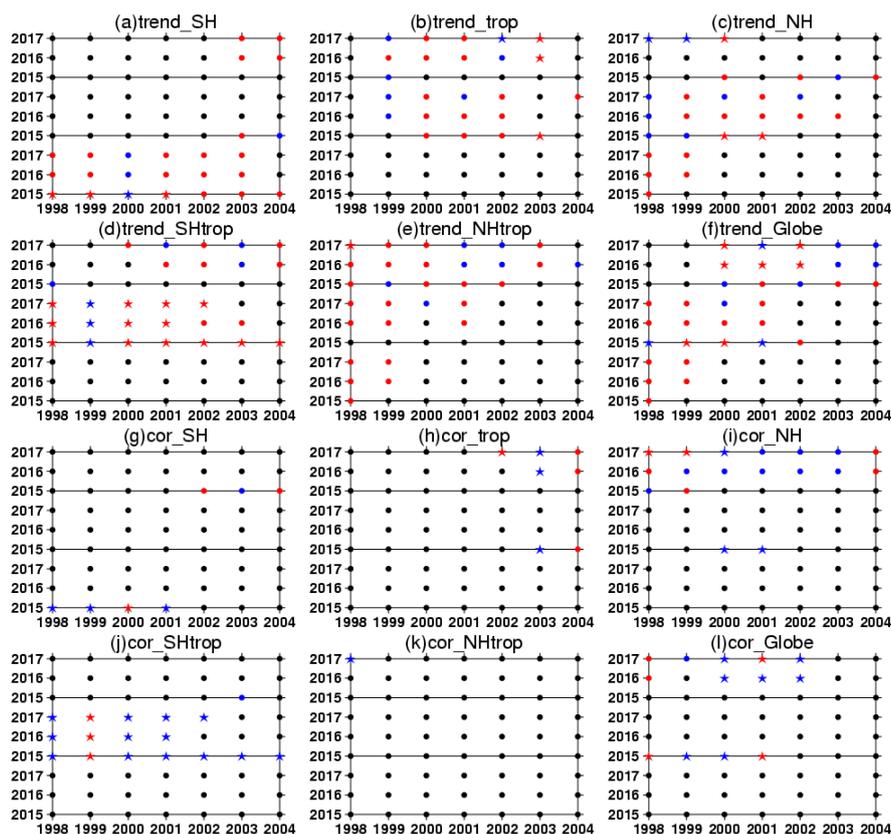


866 and vertical components are 3×10^5 and 3×10^3 kg s⁻² per year, respectively) and its
867 vertical component (shading) with their (e) wave-1 component and (f) wave-2
868 component in September during 2000–2017 derived from MERRA-2 dataset. The
869 stippled regions represent the trend significant at/above the 90% confidence level.



870

871 **FIG. 4.** Trends of SST in September over (a) 1980-2000 and (b) 2000-2017 derived
872 from ERSST v5 dataset. The stippled regions represent the trends significant at/above
873 the 90% confidence level.

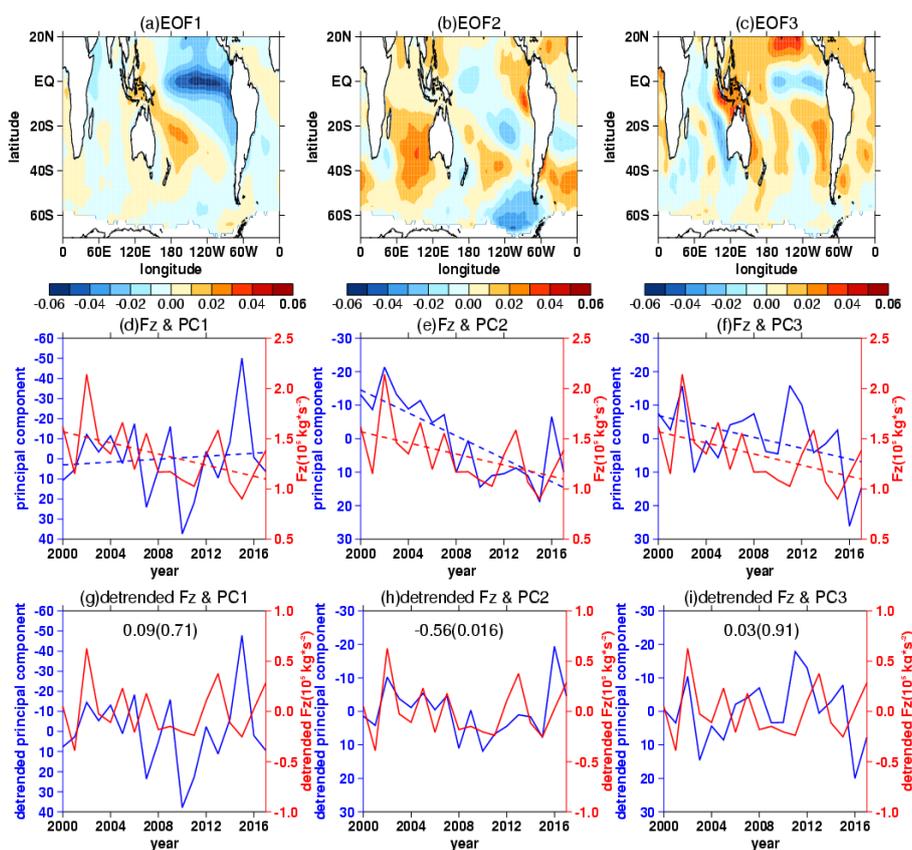


874

875 **FIG. 5.** Trend significance of the first three SST principal components (PCs) in (a) the
876 extratropical southern hemisphere (SH, 70°S-20°S), (b) the tropics (TROP, 20°S-20°N),
877 (c) the extratropical northern hemisphere (NH, 20°N-70°N), (d) the extratropical
878 southern hemisphere and the tropics (SHtrop, 70°S-20°N), (e) the extratropical northern
879 hemisphere and the tropics (NHtrop, 20°S-70°N), (f) the globe (70°S-70°N) and the
880 corresponding (g, h, i, j, k, l) correlation significances between them and vertical E-P
881 flux (F_z , area-weighted from 100 hPa to 30 hPa over 70°S-50°S) during different
882 beginning years (x-axes) and ending years (y-axes). The red and blue dots indicate
883 positive and negative trend or correlation coefficient are significant, respectively. The
884 black dots indicate the trends or correlation coefficients are not significant. The stars



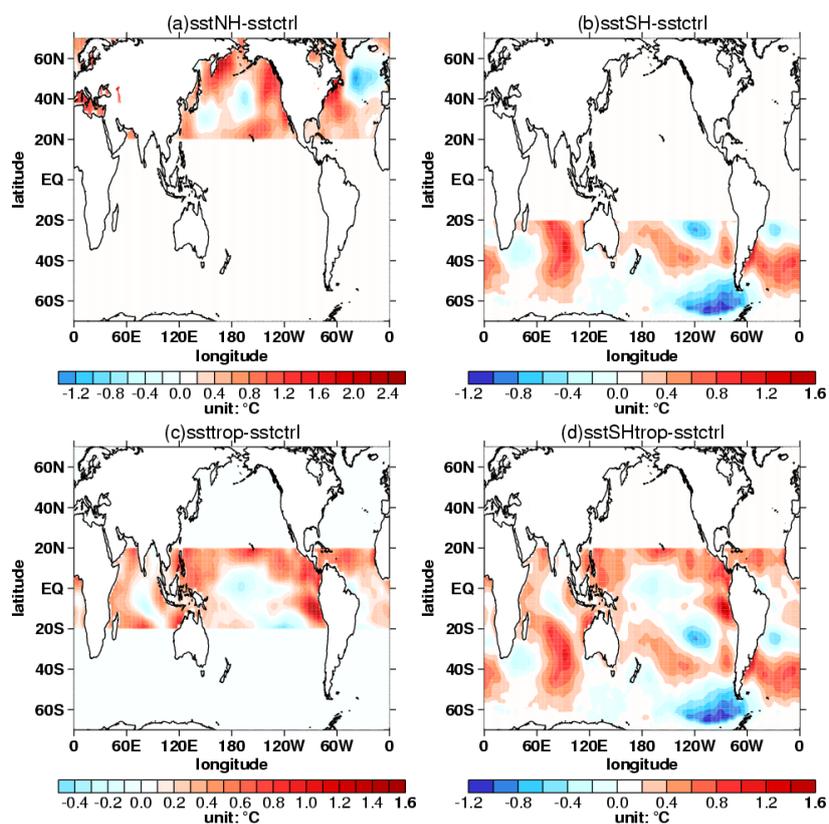
885 indicate that the trends and the corresponding correlation coefficients are both
886 significant. Each panel is divided into three regions from bottom to top, corresponding
887 to the first, the second and the third principal components, respectively. The criterion
888 to distinguish whether the trends and correlations are significant or not is the 90%
889 confidence level.



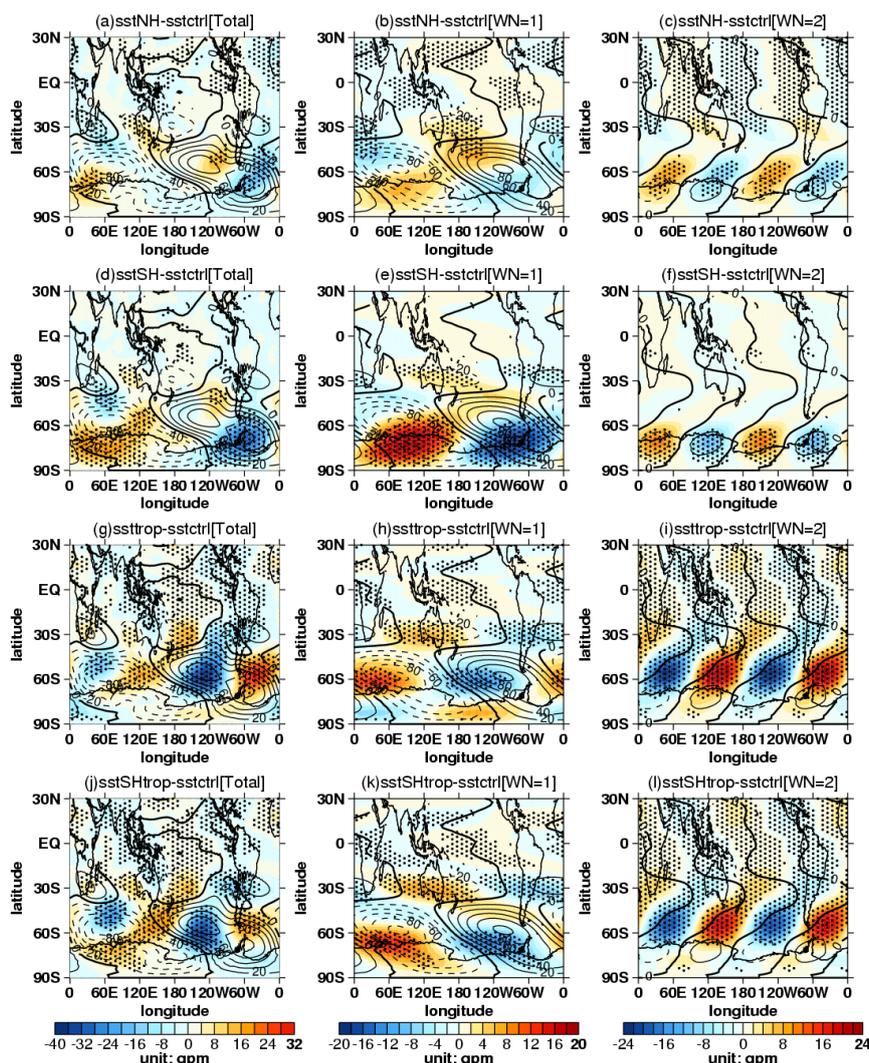
890
891 **FIG. 6.** (a, b, c) The first three EOF patterns of SST in SHtrop region. (d, e, f) The
892 original time series of the first three principle components (PCs, blue solid lines
893 correspond to left inverted y-axes) and stratospheric vertical E-P flux (Fz, area-
894 weighted from 100 hPa to 30 hPa over 70°S-50°S, red solid lines correspond to right y-
895 axes) in September during 2000-2017. The blue and red dashed lines in (d, e, f)



896 represent the linear regressions of PC time series and Fz time series, respectively. The
897 meaning of (g, h, i) are the same as (d, e, f) correspondingly, except the detrended time
898 series. The unbracketed and bracketed numbers in (g, h, i) represent the correlation
899 coefficients between detrended PC time series and Fz time series and the corresponding
900 p values calculated by two-tailed t test, respectively.



901
902 **FIG. 7.** Differences in SST forcing field between sensitive experiments ((a) sstNH; (b)
903 sstSH; (c) ssttrop; (d) sstSHtrop) and the control experiment (sstctrl).

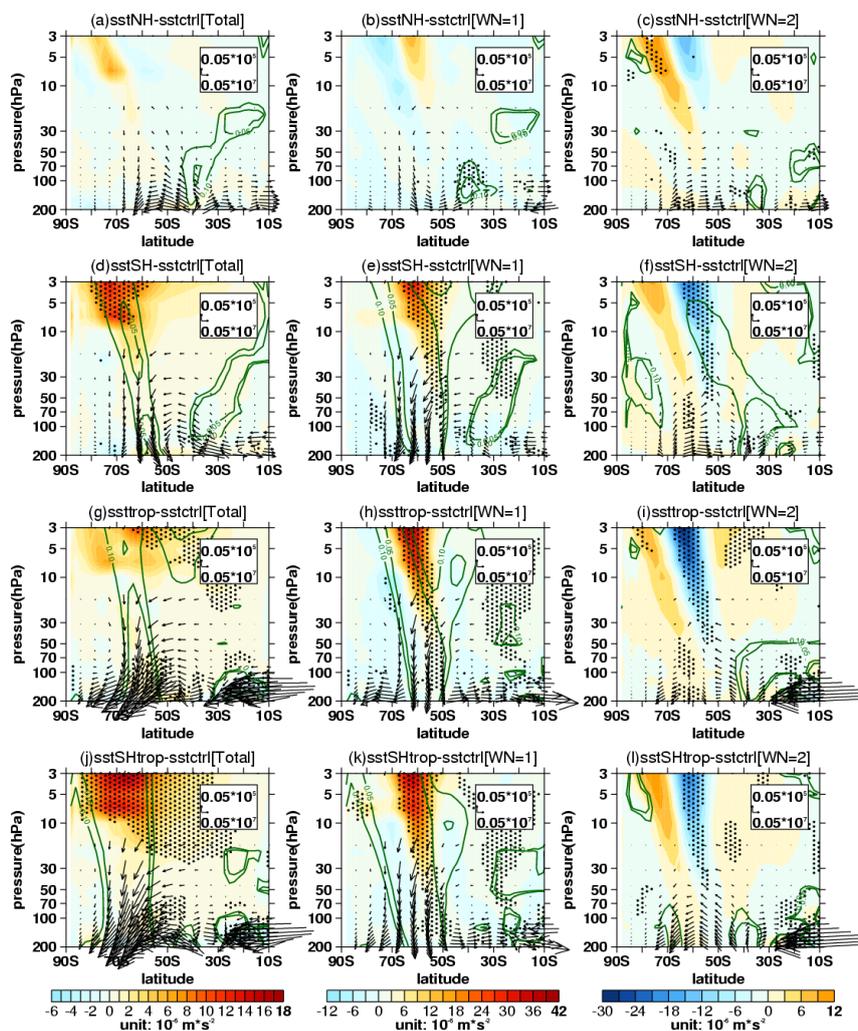


904

905 **FIG. 8.** Differences (shadings) of (a, d, g, j) 500 hPa geopotential zonal
 906 deviations with their (b, e, h, k) wave-1 component and (c, f, i, l) wave-2 component
 907 between sensitive experiments ((a, b, c) sstNH; (d, e, f) sstSH; (g, h, i) ssttrop; (j, k, l)
 908 sstSHtrop) and the control experiment (sstctrl). The mean distributions (contours with
 909 an interval of 20 gpm, positive and negative values are depicted by solid and dashed
 910 lines respectively, zeroes are depicted by thick solid lines) of them are derived from the



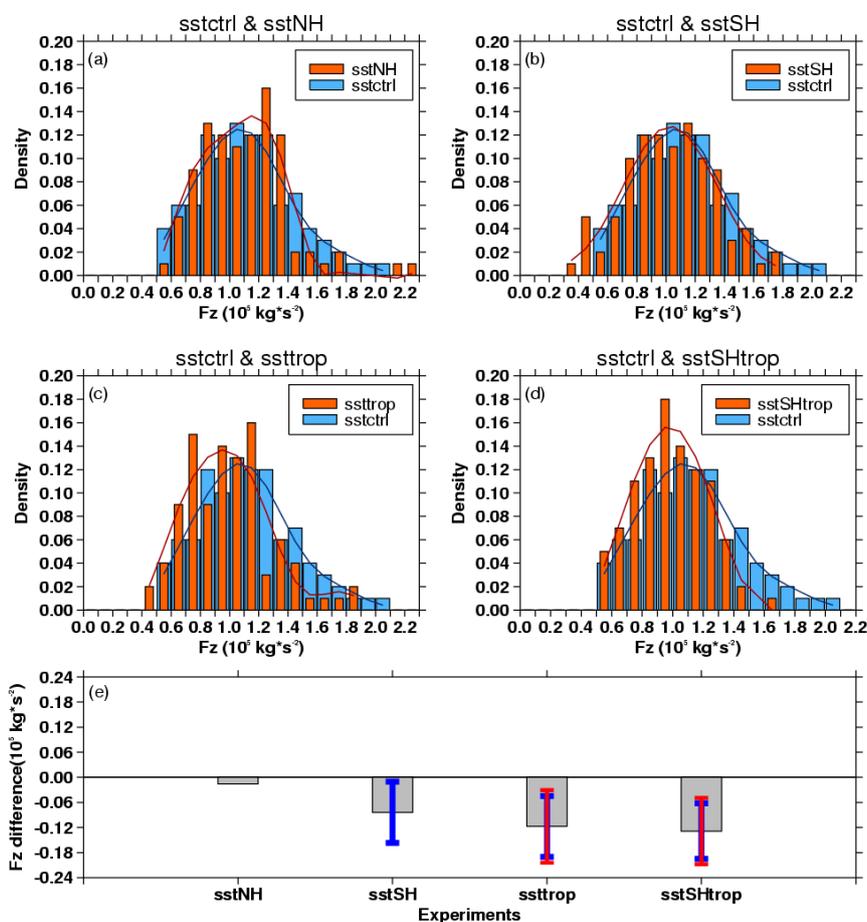
911 control experiment. The stippled regions represent the mean difference significant
 912 at/above the 90% confidence level.



913
 914 **FIG. 9.** Differences of (a, d, g, j) stratospheric E-P flux (arrows, units in horizontal and
 915 vertical components are 0.05×10^7 and $0.05 \times 10^5 \text{ kg} \cdot \text{s}^{-2}$, respectively) and its divergence
 916 (shadings) with their (b, e, h, k) wave-1 component and (c, f, i, l) wave-2 component
 917 between sensitive experiments ((a, b, c) sstNH; (d, e, f) sstSH; (g, h, i) ssttrop; (j, k, l)
 918 sstSHtrop) and the control experiment (sstctr). The stippled regions represent the mean



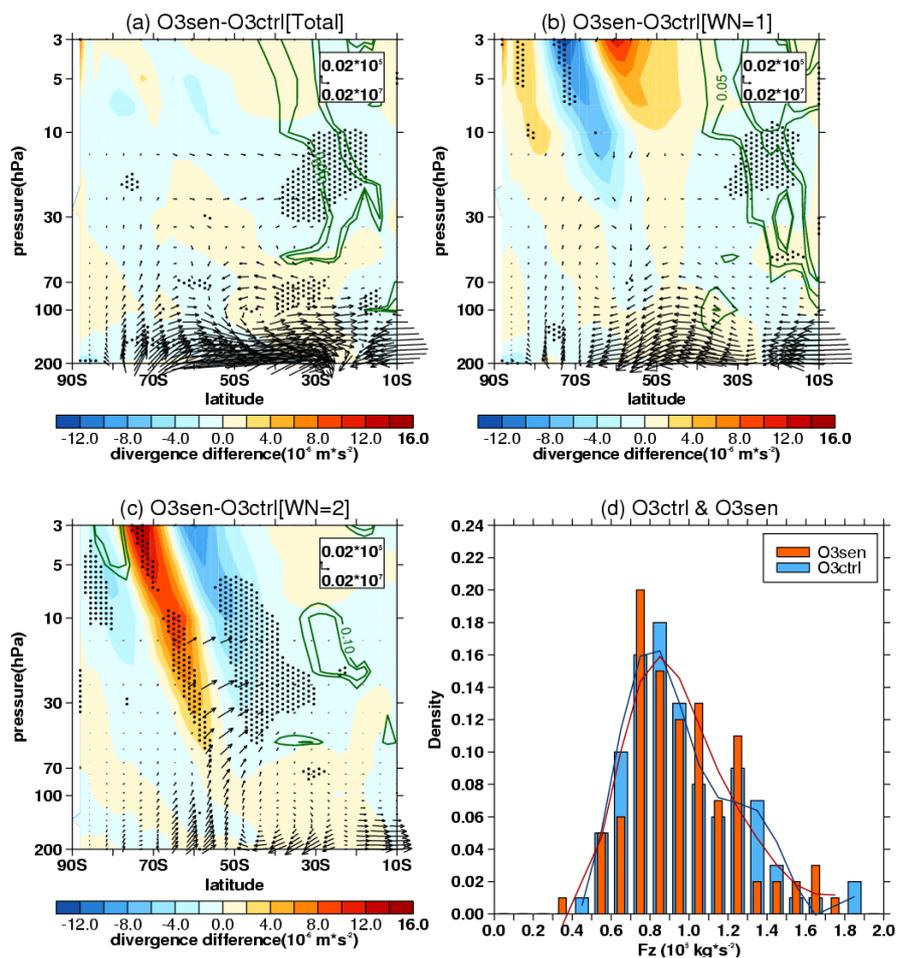
919 differences of E-P flux divergence significant at/above the 90% confidence level. The
 920 green contours from outside to inside (corresponding to $p=0.1, 0.05$) represent the mean
 921 differences of vertical E-P flux significant at the 90% and 95% confidence levels,
 922 respectively.



923
 924 **FIG. 10.** (a, b, c, d) Frequency distributions (pillars, blue for control experiment and
 925 orange for sensitive experiments) of vertical E-P flux (F_z , area-weighted from 200 hPa
 926 to 10 hPa over 70°S - 50°S) and its 5-point low-pass filtered fitting curves (solid lines,
 927 blue for control experiment and red for sensitive experiments) derived from 100



928 ensemble members of the control experiment (sstctrl) and sensitive experiments ((a)
 929 sstNH; (b) sstSH; (c) ssttrop; (d) sstSHtrop), respectively. (e) Mean differences (grey
 930 pillars) and corresponding uncertainties (error bars) of F_z between sensitive
 931 experiments and the control experiment. The blue and red error bars reflect the 90%
 932 and 95% confidence levels calculated by two-tailed t test, respectively. The error bar is
 933 omitted when the significance of mean difference is lower than the corresponding
 934 confidence level.

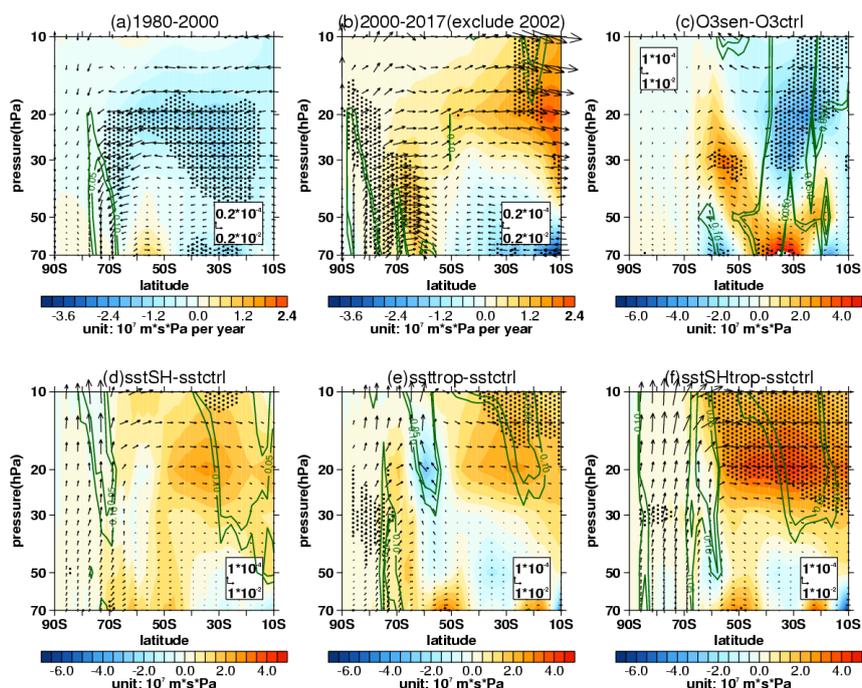


935

936 **FIG. 11.** Differences of (a) stratospheric E-P flux (arrows, units in horizontal and



937 vertical components are 0.02×10^7 and $0.05 \times 10^5 \text{ kg} \cdot \text{s}^{-2}$, respectively) and its divergence
 938 (shadings) with their (b) wave-1 component and (c) wave-2 component between the
 939 sensitive experiment (O3sen) and the control experiment (O3ctrl). The stippled regions
 940 represent the mean differences of E-P flux divergence significant at/above the 90%
 941 confidence level. The green contours from outside to inside (corresponding to $p=0.1$,
 942 0.05) represent the mean differences of vertical E-P flux significant at the 90% and 95%
 943 confidence levels, respectively. (d) Frequency distributions (pillars, blue for O3ctrl and
 944 orange for O3sen) of vertical E-P flux (F_z , area-weighted from 200 hPa to 10 hPa over
 945 70°S - 50°S) and its 5-point low-pass filtered fitting curves (solid lines, blue for O3ctrl
 946 and red for O3sen) derived from 100 ensemble members.



947

948 **FIG. 12.** (a) Trends of southern hemispheric Brewer-Dobson circulation (arrows, units
 949 in horizontal and vertical components are 0.2×10^{-2} and $0.2 \times 10^{-4} \text{ m} \cdot \text{s}^{-1} \text{ Pa}$ per year,



950 respectively) and its stream function (shadings) in September during (a) 1980-2000 and
951 (b) 2000-2017 derived from MERRA-2 dataset. Data in 2002 are removed when trends
952 are calculated in Figure (b). (c) Differences of Brewer-Dobson circulation (arrows,
953 units in horizontal and vertical components are 10^{-2} and 10^{-4} $\text{m}\cdot\text{s}^{-1}$, respectively) and its
954 stream function (shadings) between the O3ctrl and O3sen. (d, e, f) Differences of
955 Brewer-Dobson circulation and its stream function between the control experiment
956 (sstctrl) and various sensitive experiments ((d) sstSH; (e) ssttrop; (f) sstSHtrop) with
957 SST changes. The stippled regions represent the trends or differences of the stream
958 function significant at/above the 90% confidence level. The green contours from
959 outside to inside (corresponding to $p=0.1, 0.05$) represent the trends or differences of
960 the vertical components significant at the 90% and 95% confidence levels, respectively.