

**Weakening of Antarctic Stratospheric Planetary Wave  
Activities in Early Austral Spring Since the Early 2000s: A  
Response to Sea Surface Temperature Trends**

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1 **Abstract**

2 Using multiple reanalysis datasets and modeling simulations, the trends of  
3 Antarctic stratospheric planetary wave activities in early austral spring since the early  
4 2000s are investigated in this study. We find that the stratospheric planetary wave  
5 activities in September have weakened significantly since 2000, which is mainly related  
6 to the weakening of the tropospheric wave sources in the extratropical southern  
7 hemisphere. As the Antarctic ozone also shows clear shift around 2000, the impact of  
8 ozone recovery on Antarctic planetary wave activity is also examined through  
9 numerical simulations. Significant ozone recovery in lower stratosphere changes the  
10 atmospheric state for wave propagation to some extent, inducing a slight decrease of  
11 vertical wave flux in upper troposphere and lower stratosphere (UTLS). However, the  
12 changes of wave propagation environment in middle and upper stratosphere over  
13 subpolar region are not significant. The ozone recovery has minor contribution to the  
14 significant weakening of stratospheric planetary wave activity in September. Further  
15 analysis indicates that the trend of September sea surface temperature (SST) over 20°  
16 N-70°S is well linked to the weakening of stratospheric planetary wave activities. The  
17 model simulations reveal that the SST trend in the extratropical southern hemisphere  
18 (20°S-70°S) and the tropics (20°N-20°S) induce a weakening of wave-1 component of  
19 tropospheric geopotential height in the extratropical southern hemisphere, which  
20 subsequently leads to a decrease in stratospheric wave flux. In addition, both reanalysis  
21 data and numerical simulations indicate that the Brewer-Dobson circulation (BDC)  
22 related to wave activities in the stratosphere has also been weakening in early austral

23 spring since 2000 due to the trend of September SST in the tropics and extratropical  
24 southern hemisphere.

25

26 **Key words:** *Antarctic; Stratospheric planetary wave activities; Tropospheric wave*  
27 *sources; Sea surface temperature*

28

## 29 **1. Introduction**

30 The stratospheric planetary wave activities have important influences on  
31 stratospheric temperature (e.g., Hu & Fu, 2009; Lin et al., 2009; Li & Tian, 2017; Li et  
32 al., 2018), polar vortex (e.g., Kim et al., 2014; Zhang et al., 2016; Hu et al., 2018) and  
33 distribution of chemical substances (e.g., Gabriel et al., 2011; Ialongo et al., 2012;  
34 Kravchenko et al., 2011; Zhang et al., 2019a). Meanwhile, the stratospheric circulation  
35 modulated by planetary waves can exert impacts on tropospheric weather and climate  
36 (e.g., Haigh et al., 2005; Zhang et al., 2019b) through downward control processes  
37 (Haynes et al., 1991), which is useful for extended forecast by using preceding signals  
38 in the stratosphere (e.g., Baldwin et al., 2001; Wang et al., 2020).

39 The planetary perturbations generated by large-scale topography, convection and  
40 continent-ocean heating contrast can propagate from the troposphere to the stratosphere  
41 (Charney & Drazin, 1961) and form stratospheric planetary waves. As the land-sea  
42 thermal contrast in the northern hemisphere is larger than that in the southern  
43 hemisphere and produces stronger zonal forcing for the genesis of stratospheric waves,  
44 the majority of attention has been given to wave activities and their impacts on weather

45 and climate in the northern hemisphere (e.g., Kim et al., 2014; Zhang et al., 2016; Hu  
46 et al., 2018). However, planetary wave activities in the southern hemisphere also play  
47 an important role in heating the stratosphere dynamically (e.g., Hu & Fu, 2009; Lin et  
48 al., 2009), which suppresses Polar Stratospheric Clouds (PSCs) formation and ozone  
49 depletion (e.g., Shen et al., 2020a; Tian et al., 2017). The Antarctic sudden stratospheric  
50 warming (SSW) that occurred in 2002 (e.g., Baldwin et al., 2003; Nishii & Nakamura,  
51 2004; Newman & Nash, 2005) and 2019 (e.g., Yamazaki et al., 2020; Shen et al., 2020a;  
52 Shen et al., 2020b) was associated with significant upward propagation of wave flux.  
53 Such episodes are extraordinarily rare in the history, and the one in 2019 contributed to  
54 the formation of the smallest Antarctic ozone hole on record (WMO, 2019). In addition,  
55 some studies reported that wildfires in Australia at the end of 2019 are related to  
56 negative phase of the Southern Annular Mode (SAM), which was induced by the  
57 extended influence of the SSW event that occurred in September (Lim et al., 2019; Shen  
58 et al., 2020b). In a word, the Antarctic planetary wave activities are important for the  
59 stratosphere-troposphere interactions and climate system in the southern hemisphere.

60 Long-term observations in the Antarctic stratosphere show a significant ozone  
61 decline from the early 1980s to the early 2000s due to anthropogenic emission of  
62 chlorofluorocarbons (CFCs) (WMO, 2011) and a recovery signal since 2000s because  
63 of phasing out CFCs in response to Montreal Protocol (e.g., Angell and Free, 2009;  
64 Krzyścin, 2012; Zhang et al., 2014; Banerjee et al., 2020). The Antarctic stratospheric  
65 ozone depletion and recovery have important impacts on climate in the southern  
66 hemisphere. The ozone depletion cools the Antarctic stratosphere through reducing

67 absorption of radiation and leads to the strengthening of Antarctic polar vortex during  
68 austral spring (e.g., Randel & Wu, 1999; Solomon, 1999; Thompson et al., 2011). The  
69 anomalous circulation in the Antarctic stratosphere during austral spring exerts impacts  
70 on tropospheric circulations (e.g., intensification of SAM index, poleward shift of  
71 tropospheric jet position and expansion of the Hadley cell edge) in the subsequent  
72 months (e.g., Thompson et al., 2011; Swart & Fyfe, 2012; Son et al., 2018; Banerjee et  
73 al., 2020) and influences the distribution of precipitation and dry zone in the southern  
74 hemisphere (e.g., Thompson et al., 2011; Barnes et al., 2013; Kang et al., 2011).  
75 Following the healing of ozone loss in the Antarctic ozone hole since 2000s (e.g.,  
76 Solomon et al., 2016; Susan et al., 2019), great attention has been paid on possible  
77 impacts of ozone recovery on climate system in the southern hemisphere (e.g., Son et  
78 al., 2008; Barnes et al., 2013; Xia et al., 2020; Banerjee et al., 2020). Son et al. (2008)  
79 implemented the Chemistry-Climate Model Validation (CCMVal) models to predict the  
80 response of the southern hemisphere westerly jet to stratospheric ozone recovery. Based  
81 on the Phase 5 of Coupled Model Intercomparison Projects (CMIP5) models, Barnes et  
82 al. (2013) proposed that the tropospheric jet and dry zone edge no longer shift poleward  
83 during austral summer since the early 2000s due to ozone recovery. Banerjee et al.  
84 (2020) analyzed observations and reanalysis datasets. They found that following the  
85 ozone recovery after 2000, the increase of SAM index and the poleward shifting of  
86 tropospheric jet position as well as the Hadley cell edge all experienced a pause. Their  
87 results suggest that ozone depletion and recovery have made important contributions to  
88 the climate shift that occurred around 2000 in the southern hemisphere.

89        However, some previous studies have reported zonally asymmetric warming  
90 patterns in Antarctic stratosphere, which are generated by increased planetary wave  
91 activities during austral spring from the early 1980s to the early 2000s (Hu & Fu, 2009;  
92 Lin et al., 2009). Note that the Antarctic stratosphere was experiencing radiative cooling  
93 in the same period due to ozone depletion (e.g., Randel & Wu, 1999; Solomon, 1999;  
94 Thompson et al., 2011). The increase in stratospheric planetary wave activities cannot  
95 be explained by ozone decline, because the acceleration of stratospheric circumpolar  
96 wind caused by radiative cooling induces more wave energy to be reflected back to the  
97 troposphere (e.g., Andrews et al., 1987; Holton  
98 et al., 2004). Hu & Fu (2009) attributed the increase in Antarctic stratospheric wave  
99 activities to the SST trend from the 1980s to the 2000s. Their results indicate that in  
100 addition to ozone change, other factors such as changes in SST also contribute to  
101 climate change in the southern hemisphere. Moreover, the phase of Interdecadal Pacific  
102 Oscillation (IPO) also changed at around 2000 (e.g., Trenberth et al., 2013). SST  
103 variation influences Rossby wave propagation and tropospheric wave sources, and  
104 thereby indirectly affects stratospheric wave activities (e.g., Lin et al., 2012; Hu et al.,  
105 2018; Tian et al., 2017). The questions here are: (1) Has the trend of stratospheric  
106 planetary wave activity in the southern hemisphere been shifting since the 2000s? (2)  
107 What are the factors responsible for the trend of Antarctic stratospheric planetary wave  
108 activity since the 2000s?

109        In this study, we reveal the trend of Antarctic planetary wave activity in early  
110 austral spring since the 2000s based on multiple reanalysis datasets. We also conduct

111 sensitive experiments forced by linear increments of ozone and SST fields since the  
112 2000s to investigate the response of Antarctic planetary activity to above two factors.  
113 The remainder of the paper is organized as follows. Section 2 describes the data,  
114 methods and configurations of model simulations. Section 3 presents the trends of  
115 stratospheric and tropospheric wave activities in early austral spring. Section 4  
116 examines the impact of ozone recovery on Antarctic stratospheric planetary wave  
117 activity. Section 5 investigates the connections between the trends of SST and  
118 stratospheric wave activities. Sections 6 discusses the responses of tropospheric wave  
119 sources and stratospheric wave activities to SST changes based on model simulations.  
120 Major conclusions and discussion are presented in Section 7.

## 121 **2. Datasets, methods and experimental configurations**

### 122 a. Datasets

123 In this study, daily and monthly mean data extracted from the Modern-Era  
124 Retrospective analysis for Research and Applications Version 2 (MERRA-2;  
125 Bosilovich et al., 2015) dataset are used to calculate trends of Brewer-Dobson  
126 circulation (BDC), tropospheric wave sources, and the Elisassen-Palm (E-P) flux and  
127 its divergence in September. To verify the trend of stratospheric E-P flux, we also refer  
128 to the results derived from the European Centre for Medium-range Weather Forecasting  
129 (ECMWF) Interim Reanalysis (ERA-Interim; Dee et al., 2011) dataset, the Japanese  
130 55-year Reanalysis (JRA-55; Kobayashi et al., 2015) dataset and the National Centers  
131 for Environmental Prediction-Department of Energy Global Reanalysis 2 (NCEP-2;  
132 Kanamitsu et al., 2002) dataset.

133 The observed total column ozone (TCO) data are extracted from SBUV v8.6  
 134 satellite dataset, which is a monthly and zonal mean dataset on 5° grid. Ozone data  
 135 derived from MERRA-2 dataset are also used to calculate TCO.

136 SST data are extracted from the Extended Reconstructed Sea Surface Temperature  
 137 (ERSST) dataset, which is a global monthly mean sea surface temperature dataset  
 138 derived from the International Comprehensive Ocean-Atmosphere Dataset (ICOADS).  
 139 The ERSST is on global 2°×2° grid and covers the period from January 1854 to the  
 140 present. We use the latest version (version 5, i.e., v5) dataset to calculate trends and  
 141 correlations, and produce SST forcing field for model simulations. More details about  
 142 this version of ERSST can be found in Huang et al. (2017).

143 In addition, the unfiltered Interdecadal Pacific Oscillation (IPO) index derived  
 144 from the ERSST v5 dataset is also used in this study. More detailed information about  
 145 the index can be found in Henley et al. (2015).

#### 146 b. Diagnosis of wave activities and Brewer-Dobson circulation

147 Planetary wave activities are measured by E-P flux ( $\vec{F} \equiv (0, F^{(\phi)}, F^{(z)})$ ) and its  
 148 divergence  $D_F$ . Their algorithms are expressed by Eqs. (1)-(3) (Andrews et al., 1987):

$$149 \quad F^{(\phi)} = \rho_0 a \cos \phi (\overline{u_z v' \theta'} / \overline{\theta_z} - \overline{v' u'}) \quad (1)$$

$$150 \quad F^{(z)} = \rho_0 a \cos \phi \{ [f - (a \cos \phi)^{-1} (\overline{u \cos \phi})_\phi] \overline{v' \theta'} / \overline{\theta_z} - \overline{w' u'} \} \quad (2)$$

$$151 \quad D_F = \frac{\nabla \cdot \vec{F}}{\rho_0 a \cos \phi} = \frac{(a \cos \phi)^{-1} \frac{\partial}{\partial \phi} (F^{(\phi)} \cos \phi) + \frac{\partial F^{(z)}}{\partial z}}{\rho_0 a \cos \phi} \quad (3)$$

152 where  $\mathbf{u}$ ,  $\mathbf{v}$  represent zonal and meridional components of horizontal wind,  $w$   
 153 is vertical velocity,  $\theta$  is potential temperature,  $a$  is the Earth radius,  $f$  is the



154 Coriolis parameter,  $z$  is geopotential height,  $\phi$  is latitude,  $\rho_0$  is the background air  
 155 density.

156 The quasi-geostrophic refractive index (RI) is used to diagnose the environment  
 157 of wave propagation (Chen & Robinson, 1992). Its algorithm is written as Eq. (4):

$$158 \quad RI = \frac{\bar{q}_\phi}{\bar{u}} - \left(\frac{k}{a \cos \phi}\right)^2 - \left(\frac{f}{2NH}\right)^2 \quad (4)$$

159 where the zonal-mean potential vorticity meridional gradient  $\bar{q}_\phi$  is

$$160 \quad \bar{q}_\phi = \frac{2\Omega}{a} \cos \phi - \frac{1}{a^2} \left[ \frac{(\bar{u} \cos \phi)_\phi}{a \cos \phi} \right]_\phi - \frac{f^2}{\rho_0} \left( \rho_0 \frac{\bar{u}_z}{N^2} \right)_z \quad (5)$$

161  $H$ ,  $q$ ,  $k$ ,  $N^2$  and  $\Omega$  are the scale height, potential vorticity, zonal wave number,  
 162 buoyancy frequency, and Earth's angular frequency, respectively.

163 The Brewer-Dobson circulation driven by wave breaking in the stratosphere is  
 164 closely related to stratospheric wave activities. Its meridional and vertical components  
 165  $(\bar{v}^*, \bar{w}^*)$  and stream function  $(\psi^*(p, \phi))$  are expressed by Eqs. (4)-(6) (Andrews et al.,  
 166 1987; Birner & Bönisch, 2011) :

$$167 \quad \bar{v}^* \equiv \bar{v} - \rho_0^{-1} (\rho_0 \bar{v}' \theta' / \bar{\theta}_z)_z \quad (6)$$

$$168 \quad \bar{w}^* \equiv \bar{w} + (a \cos \phi)^{-1} (\cos \phi \cdot \bar{v}' \theta' / \bar{\theta}_z)_\phi \quad (7)$$

$$169 \quad \psi^*(p, \phi) = \int_0^p \frac{-2\pi a \cdot \cos \phi \cdot \bar{v}^*(p'', \phi)}{g} dp'' \quad (8)$$

170 where  $p$  is the air pressure,  $\pi$  is the circular constant,  $g$  is the gravitational  
 171 acceleration.

172 In Eqs. (1)-(8), the overbar and prime denote a zonal mean and departure from the  
 173 zonal mean, respectively. The subscripts denote partial derivatives. The Fourier  
 174 decomposition is used to obtain components of Eqs. (1)-(3) with different zonal wave

175 numbers. Meanwhile, the Fourier decomposed components of geopotential height zonal  
176 deviations are also used to determine tropospheric wave sources.

177 c. Statistical methods

178 The trend is measured by the slope of linear regression based on the least square  
179 estimation. The correlation is used to analyze statistical links between different  
180 variables. In this paper, all the time series have been linearly detrended before  
181 calculating correlation coefficients ( $r$ ) and their corresponding significances.

182 The change-point testing (e.g. Banerjee et al., 2020) is used to make sure the  
183 significance of trend or correlation coefficient is not unduly influenced by some  
184 particular beginning or ending years, and thereby confirm that the trend exists  
185 objectively.

186 We use two-tailed student's t test to calculate the significances of trend, correlation  
187 coefficient or mean difference. The result of significance test is measured by p value or  
188 confidence intervals in this paper.  $p \leq 0.1$ ,  $p \leq 0.05$  and  $p \leq 0.01$  suggest the trend,  
189 correlation coefficient or mean difference is significant at/above the 90%, 95% and 99%  
190 confidence levels, respectively. The confidence interval of trend is shown in (7) (Shirley  
191 et al., 2004):

$$192 \quad [\hat{b} - t_{1-\alpha/2}(n-2)\hat{\sigma}_b, \hat{b} + t_{1-\alpha/2}(n-2)\hat{\sigma}_b] \quad (7)$$

193 where  $\hat{b}$  is estimated value of slope,  $\hat{\sigma}_b$  is standard error of slope and it is written as:

$$194 \quad \hat{\sigma}_b = \hat{b} \cdot \sqrt{\frac{1}{n-2} - \frac{r^2}{n-2}}, \quad t_{1-\alpha/2}(n-2) \text{ denotes the value of t-distribution with the degree of}$$

195 freedom equal to  $n-2$  and the two-tailed confidence level equal to  $1-\alpha$  ( $\alpha = 0.90$ ,

196 0.95 or 0.99). The confidence interval of mean difference is expressed by Eq. (8)

197 (Shirley et al., 2004):

$$198 \quad [\bar{X} - \bar{Y} - t_{1-\alpha/2}(M+N-2) \cdot S_w \cdot \sqrt{\frac{1}{M} + \frac{1}{N}}, \bar{X} - \bar{Y} + t_{1-\alpha/2}(M+N-2) \cdot S_w \cdot \sqrt{\frac{1}{M} + \frac{1}{N}}] \quad (8)$$

199 where

$$200 \quad S_w = \sqrt{\frac{1}{M+N-2} \left[ \sum_{i=1}^M (X_i - \bar{X})^2 + \sum_{j=1}^N (Y_j - \bar{Y})^2 \right]} \quad (9)$$

201 Here,  $\bar{X}$  and  $\bar{Y}$  are the sample averages,  $M$  and  $N$  are the numbers of sample  
202 sizes with two populations,  $t_{1-\alpha/2}(M+N-2)$  denotes the value of t-distribution with  
203 the degree of freedom equal to  $M+N-2$  and the two-tailed confidence level equal to  
204  $1-\alpha$ .

205 Previous studies have indicated that SST impact on the stratosphere shows a  
206 spatial dependence (e.g., Xie et al., 2020). To find out a robust relationship between the  
207 trend of SST in a specific region and the trend of stratospheric wave activities, we divide  
208 the global ocean into three regions: SH (the extratropical southern hemisphere, 70°S-  
209 20°S), TROP (the tropics, 20°S-20°N) and NH (the extratropical northern hemisphere,  
210 20°N-70°N). Since the impacts in different regions might be combined, we also  
211 consider three combined regions named as SHtrop (the extratropical southern  
212 hemisphere and the tropics, 70°S-20°N), NHtrop (the extratropical northern hemisphere  
213 and the tropics, 20°S-70°N) and the Globe (70°S-70°N). To find statistical connections  
214 between the trend of SST and that of stratospheric wave activities, we examine the first  
215 three leading patterns (EOF1, EOF2, EOF3) and principal components (PC1, PC2, PC3)  
216 of SST in above six regions obtained from Empirical Orthogonal Function (EOF)

217 analysis. In all the six regions, there is always one EOF mode that shows great similarity  
218 to the spatial pattern of trend (not shown) as we do not detrend SST time series when  
219 the EOF analysis is carried out. Thus, the significance of the correlation between the  
220 PC time series of that EOF mode and time series of stratospheric E-P flux can be used  
221 as the criterion to determine the statistical connection between the trend of SST and the  
222 trend of stratospheric wave activities.

#### 223 d. The model and experiment configurations

224 The F\_2000\_WACCM\_SC (FWSC) component in the Community Earth System  
225 Model (CESM; version 1.2.0) is used to verify the impacts of SST and ozone recovery  
226 on tropospheric wave sources and stratospheric wave activities in early austral spring.  
227 The FWSC component is the Whole Atmosphere Community Climate Model version 4  
228 (WACCM4) with specified chemistry forcing fields (such as ozone, greenhouse gases  
229 (GHG), aerosols and so on), which have fixed values in 2000 by default. The WACCM4  
230 includes active atmosphere, data ocean (run as a prescribed component, simply reading  
231 SST forcing data instead of running ocean model), land and sea ice. Physics schemes  
232 in the WACCM4 are based on those in the Community Atmospheric Model version 4  
233 (CAM4; Neale et al., 2013). The WACCM4 uses a finite-volume dynamic framework  
234 and extends from the ground to approximately 145 km ( $5.1 \times 10^{-6}$  hPa) altitude in the  
235 vertical with 66 vertical levels. The simulations presented in this paper are conducted  
236 at a horizontal resolution of  $1.9^\circ \times 2.5^\circ$ . More information about the WACCM can be  
237 found in Marsh et al. (2013).

238 Control experiments and sensitive experiments are conducted to investigate

239 responses of Antarctic stratospheric wave activities to SST trends and the ozone  
240 recovery trend in early austral spring. For the experiments of SST trends, monthly mean  
241 global SST during 1980-2000 derived from the ERSST v5 dataset is used as SST  
242 forcing field in the control experiment (sstctrl). For the four sensitive experiments  
243 (sstNH, sstSH, ssttrop, sstSHtrop), linear increments of SST in different regions in  
244 September during 2000-2017 are used as the forcing field. Ozone, aerosols and  
245 greenhouse gases (GHG) in the control experiment and the four sensitive experiments  
246 all have the fixed values in 2000. For the experiments of ozone recovery trend, monthly  
247 mean three-dimensional global ozone during 1980-2000 derived from the MERRA-2  
248 dataset is used as the ozone forcing field in the control experiment (O3ctrl). The  
249 sensitive experiment (O3sen) is forced by linear increments of ozone in September  
250 during 2001-2017. The SSTs in O3ctrl and O3sen both are monthly mean global SST  
251 during 1980-2000. The aerosol and greenhouse gases values in 2000 are used. These  
252 experiment configurations are summarized and listed in Table 1 and Table 2.

253 Firstly, we run the FWSC component to generate randomly different initial  
254 conditions for 120 years with free run. Then, each experiment includes 100 ensemble  
255 members that run from July to September forced by these initial conditions from the  
256 21st year to the 120th year in July. The forcing fields of SST and ozone are only  
257 superposed from July to September. July and August are taken as spin-up time and  
258 simulations during this period are discarded. The ensemble mean in September derived  
259 from these 100 ensemble members is regarded as the final result of each experiment. A  
260 similar approach is implemented for sensitive experiments, in which the forcing fields

261 superposed only in certain months. The same approach has been used in previous  
262 studies (e.g., Zhang et al., 2018).

### 263 **3. Trend of planetary wave activities in early austral spring**

264 Figure 1 shows the trends of stratospheric planetary wave activities in the southern  
265 hemisphere September during 1980-2000 and 2000-2017, respectively. Note that the  
266 vertical E-P flux entering into the stratosphere over 50°S-70°S in September has been  
267 increasing during 1980-2000, accompanied by intensified wave flux convergence in the  
268 upper stratosphere (Fig. 1a) that is mainly contributed by the wave-1 component (Fig.  
269 1b). This feature implies that the stratospheric planetary wave activities have  
270 strengthened in early austral spring during 1980-2000. A similar result has been  
271 reported in previous studies (Hu & Fu, 2009; Lin et al., 2009). During 2000-2017,  
272 however, vertical propagation of stratospheric E-P flux weakened over the subpolar  
273 region of the southern hemisphere, which was accompanied by intensified wave flux  
274 divergence in the upper stratosphere (Fig. 1d) mainly contributed by the wave-1  
275 component (Fig. 1e) while the wave-2 component also made certain contributions (Fig.  
276 1f). Similar features also appear in August, but not as significant as that in September  
277 (Fig. S1). For this reason, hereafter we focus on the features in September.

278 The SSW that occurred in 2002 was accompanied with large upward wave fluxes  
279 in the stratosphere, which is extremely rare in history and has been studied extensively  
280 in previous studies (e.g., Baldwin et al., 2003; Nishii & Nakamura, 2004; Newman &  
281 Nash, 2005). Since the period with a negative trend of stratospheric vertical wave flux  
282 is short, it is necessary to further investigate whether such a negative trend is artificially

283 influenced by the single year of 2002. Therefore, following Banerjee et al. (2020), we  
284 use a change-point method to test the significance of the trend during various periods  
285 based on four reanalysis datasets (ERA-Interim, MERRA-2, JRA-55, NCEP-2).  
286 Figures 2a (including the year 2002) and 2b (excluding the year 2002) display the time  
287 series of area-weighted vertical stratospheric wave flux ( $F_z$ ) over the southern  
288 hemisphere subpolar region obtained from different reanalysis datasets. Note that the  
289 wave flux time series obtained from the four reanalysis datasets all present a positive  
290 trend from the early 1980s to the early 2000s and a negative trend from the early 2000s  
291 to present, regardless of whether the extreme value in 2002 is removed or not. The  
292 correlation coefficients of the time series between these reanalysis datasets are above  
293 0.9 and statistically significant (Table 3), suggesting that the time series derived from  
294 different datasets are consistent with each other. Figures 2c-f show the trends and  
295 corresponding confidence intervals calculated with four different beginning years (1980,  
296 1981, 1982, 1983), four different ending years (2015, 2016, 2017, 2018), and change-  
297 point years from 1998 to 2013. The trends and confidence intervals in Figures 2g-j are  
298 the same as that in Figures 2c-f, except that the extreme value in 2002 is removed. The  
299 positive trend from the early 1980s to the 21st century remains significant regardless of  
300 different beginning years and ending change-point years (Figs. 2c-j). However, Figures  
301 2c-f and Figures 2g-j indicate that the positive value of the trend is decreasing gradually  
302 when the period is prolonged, which is apparently attributed to the negative trend with  
303 the beginning change-point year of around 2000. Although the negative trend from the  
304 change-point year to ending year becomes less significant when the value in 2002 is

305 removed, it remains significant in some periods, which are also illustrated on diagrams  
306 of latitude-pressure profiles (Fig. S2). Therefore, the weakening of stratospheric wave  
307 activities in early austral spring since the early 2000s is robust. In this paper, we take  
308 the year 2000 as the beginning year of the weakening trend to simplify descriptions in  
309 the following discussion.

310 Figure 3 shows the trends of tropospheric wave sources in September since 2000.  
311 There is a significant positive trend of the wave-1 component in 500 hPa geopotential  
312 height over the southern Indian ocean and a significant negative trend over the southern  
313 Pacific, which form an out-of-phase superposition on its climatology (Fig. 3b). The  
314 trend pattern of wave-2 component is also out-of-phase with its climatology, although  
315 it is not significant (Fig. 3c). The above features still maintain when the values in 2002  
316 are removed (Figs. S3b, c), implying that the southern hemispheric tropospheric wave  
317 sources in early austral spring have weakened since 2000, which is also reflected in the  
318 decrease of tropospheric vertical wave flux (Figs. 3d, e; Figs. S3d, e).

#### 319 **4. Response of Antarctic stratospheric wave activity to ozone recovery**

320 Previous studies have suggested that ozone depletion and recovery are important  
321 to climate shift that occurred around 2000 in the southern hemisphere during austral  
322 summer (e.g., Son et al., 2008; Thompson et al., 2011; Barnes et al., 2013; Banerjee et  
323 al., 2020). The impacts of stratospheric ozone changes on Antarctic wave propagation  
324 during austral summer has also been examined in previous studies (e.g., Hu et al., 2015).  
325 However, whether ozone recovery in September explains the weakening of  
326 stratospheric planetary waves at the same month remains unclear. The correlation



327 between detrended time series of September Antarctic total column ozone (TCO)  
328 derived from SBUV and stratospheric vertical wave flux (Fz) is 0.70 ( $p=0.0011$ ) during  
329 2000-2017. The increase of wave activity in polar stratosphere causes heating effects  
330 and suppresses the formation of PSCs, and hence, slow down the ozone depletion (e.g.,  
331 Shen et al. 2020a). Therefore, the Antarctic ozone and stratospheric wave activity show  
332 statistically significant positive correlation. Theoretically, heating effects caused by  
333 ozone recovery in Antarctic stratosphere may also decelerate the Antarctic stratospheric  
334 polar vortex and induce more waves to propagate into stratosphere (Andrews et al.,  
335 1987; Holton et al., 2004). These preliminary analysis cannot verify that the ozone  
336 recovery is responsible for weakening of stratospheric wave activity. The role of ozone  
337 recovery in stratospheric wave changes needs to be further explored by model  
338 simulations. In this section, we use a group of time-slice experiments (O3ctrl and O3sen)  
339 to address this issue.

340 Figure 4 shows the time series and piecewise trends of September TCO in the  
341 Antarctic during 1980-2017. As reported by previous studies (e.g., Angell and Free,  
342 2009; Banerjee et al., 2020; Krzyścin, 2012; Solomon et al., 2016; WMO, 2011; Zhang  
343 et al., 2014), the Antarctic ozone show a significant decline during 1980-2000 (Figs. 4a,  
344 b, c) and a slight recovery during 2001-2017 (Figs. 4a, d, e). The recovery trend is  
345 calculated with data in 2002 removed because the large poleward transport induced by  
346 SSW in 2002 leads to extreme values of ozone (e.g. Solomon et al., 2016). In addition,  
347 the correlation of TCO between MERRA-2 and SBUV datasets is 0.61 ( $p=4.5\times 10^{-5}$ ),  
348 suggesting the changes of TCO derived from the reanalysis dataset and the observations

349 have a good consistency. Thus, in order to get three-dimensional structure of ozone  
350 changes, the ozone data from MERRA-2 are used to make forcing fields for CESM. As  
351 described in Section 2, a control experiment (O3ctrl) forced by climatological ozone  
352 and a sensitive experiment forced by the linear increment of global ozone in September  
353 during 2001-2017 are conducted to explore the impacts of ozone recovery. The pattern  
354 of ozone forcing fields is similar to its trend patterns (Figs. 4d, e; Figs. 5a, b). Other  
355 details of these two experiments have been given in Section 2 and Table 2.

356 Fig. 6 and Fig. 7 show the responses of wave activity and wave propagation  
357 environment forced by O3sen. Note that the significant ozone recovery over south pole  
358 mainly appears in lower stratosphere (about 200 hPa to 50 hPa) (Fig. 4e). In most  
359 southern polar regions from 50 hPa to 3 hPa, the ozone recovery is not significant (Fig.  
360 4e). The features are attributed to limitation of ODSs emission and reduction of  
361 heterogeneous reaction on PSCs, which mainly distribute in lower stratosphere (e.g.,  
362 Solomon, 1999). Ozone recovery in polar lower stratosphere absorbs more ultraviolet  
363 radiation and causes cooling in Antarctic troposphere (Fig. 6b). To maintain thermal  
364 balance, zonal wind accelerates below 200 hPa over 60°S-70°S (Fig. 6a).

365 The changes of zonal wind and temperature forced by ozone recovery induce  
366 changes in wave propagation environment. The refractive index (RI) is a good metric  
367 to reflect the atmosphere state for wave propagation. Theoretically, planetary waves  
368 tend to propagate into large RI regions (Andrews et al., 1987). The responses of RI and  
369 its terms are shown in Figs. 6c-f. Note that the second term of RI does not change with  
370 atmospheric state and the third term of RI is insignificant compared to the first term

371 (Hu et al., 2019). Previous studies indicate that changes in zonal mean potential  
 372 vorticity meridional gradient  $\bar{q}_\varphi$  could explain the changes in RI in middle and high  
 373 latitudes (e.g. Hu et al., 2019; Simpson et al., 2009). Consistent with these studies, the  
 374 pattern of  $\bar{q}_\varphi$  show some similarity with pattern of RI (Figs. 6c, d), especially in lower  
 375 stratosphere over subpolar regions (Figs. 6c, d). According to the Eq. (5), the first term  
 376 of  $\bar{q}_\varphi$  does not change with atmospheric state. Therefore, the second term  
 377  $(-\frac{(\bar{u} \cos \varphi)_\varphi}{\cos \varphi}]_\varphi$ , hereafter uyy term or barotropic term) and the third term  
 378  $(-\frac{f^2}{\rho_0}(\rho_0 \frac{\bar{u}_z}{N^2})_z)$ , hereafter uzz term or baroclinic term) are investigated. Note that the  
 379 pattern of responses in baroclinic term is similar with  $\bar{q}_\varphi$  (Figs. 6d, f). The uzz term  
 380 also can be written as  $(\frac{f^2}{HN^2} + \frac{f^2}{N^4} \frac{dN^2}{dz})\bar{u}_z - \frac{f^2}{N^2}\bar{u}_{zz}$ . Meanwhile, zonal wind  
 381 acceleration in upper troposphere weakens the vertical shear of u ( $\bar{u}_z$ ) around 200 hPa  
 382 over subpolar regions, inducing a decrease of baroclinic term and RI in upper  
 383 troposphere and lower stratosphere (UTLS) over 60°S-70°S (Figs. 6d, f). The response  
 384 of RI induces a slight decrease of vertical wave flux in UTLS over subpolar regions  
 385 (Fig. 7a), which is mainly contributed by its wave-1 component (Fig. 7b). However, the  
 386 changes of wave activity in UTLS are not significant in ensemble mean of simulations  
 387 (Figs. 7a, b, c). Meanwhile, note that the responses of zonal wind and temperature to  
 388 ozone recovery are not significant above 50 hPa over subpolar regions (Figs. 6a, b),  
 389 inducing negligible changes of wave propagation environment (Fig. 6c) and wave  
 390 activity (Fig. 7) in middle and upper stratosphere.

391 In a word, the significant ozone recovery in Antarctic lower stratosphere changes

392 wave propagation in upper troposphere and lower stratosphere to some extent. However,  
393 these weak responses still cannot explain the significant decrease of stratospheric wave  
394 flux in September.

## 395 **5. Role of SST trends in the weakening of Antarctic stratospheric** 396 **wave activities**

397 In this section, we further explore factors responsible for the weakening of  
398 tropospheric wave sources and stratospheric wave activities since the early 2000s in  
399 early austral spring. Many studies reported that SST variations can affect stratospheric  
400 climate (e.g., Li, 2009; Hurwitz et al., 2011; Lin et al., 2012; Hu et al., 2014; Hu et al.,  
401 2018; Tian et al., 2017; Xie et al., 2020). Hu & Fu (2009) also attributed the  
402 strengthened stratospheric wave activities in the southern hemisphere to SST trend from  
403 the early 1980s to the early 2000s. Furthermore, global SST in September during 2000-  
404 2017 also has a significant trend. The significant warming pattern is mainly found over  
405 the southern Indian ocean, the southern Atlantic ocean, the eastern and western  
406 equatorial Pacific, the western equatorial and Northern Atlantic ocean (Fig. 8b). A  
407 significant cooling pattern is located over the southeast Pacific (Fig. 8b). In addition,  
408 the transitions around 2000 exist in SST time series over some regions. In the southern  
409 Indian ocean, SST shows insignificant trend during 1980-2000 and significant warming  
410 trend during 2000-2017 (Fig. 8c). The subtropical Pacific ocean in east of Australia is  
411 linked with the Pacific-Southern America (PSA) wave train (e.g. Shen et al., 2020b),  
412 and the SST there shows significant warming trend during 1980-2000 and insignificant  
413 trend during 2000-2017. The SST in southeast Pacific shows insignificant trend during

414 1980-2000 and significant cooling during 2000-2017 (Fig. 8e). Trends of SST in  
415 southern Atlantic ocean are opposite during these two piecewise periods, showing  
416 significant cooling during 1980-2000 and significant warming during 2000-2017. It is  
417 apparent that the spatial pattern of SST trend during 2000-2017 is obviously different  
418 from that during 1980-2000 (Fig. 8a, b), which may affect the tropospheric wave  
419 sources. Thus, it is necessary to analyze the connection between SST trend and wave  
420 activity trend since the early 2000s.

421 Figure 9 shows the significance of the trend of principle component (PC) time  
422 series of SST in different regions (Figs. 9a-f), and the significance of correlations (Figs.  
423 9g-l) between the PC time series and Fz in September during various periods. The trend  
424 of PC1 time series in SH region is significant during several periods (Fig. 9a), while the  
425 correlation between PC1 and Fz is only significant with the particular ending year of  
426 2015 (Fig. 9g). This feature suggests that the connection between the SST trend in SH  
427 region and the trend of stratospheric wave activity is not robust. The correlation  
428 between trend of stratospheric wave activity and that of SST in TROP or NH region is  
429 also weak (Fig. 9e, f). As for the combined regions, note that the PC2 time series in  
430 SHtrop region has a significant trend (Fig. 9d) and the correlation between the PC2 time  
431 series in SHtrop and Fz with the beginning year of around 2000 is also significant (Fig.  
432 9j) regardless of different ending years. This feature implies that the extratropical  
433 southern hemisphere and tropical SST has a robust connection with stratospheric wave  
434 activities in early austral spring since the early 2000s. The correlations between Fz and  
435 all PC time series in NHtrop (Fig. 9k) and Globe (Fig. 9l) region are not as robust as

436 that between Fz and PC2 time series in SHtrop region (Fig. 9j), indicating that the  
437 connection between SST trend in extratropical northern hemisphere and the trend of  
438 stratospheric wave activity is weak.

439 Figure 10 shows the first three EOF modes of September SST in SHtrop region  
440 during 2000-2017. The second mode (Fig. 10b) shows a great similarity to the spatial  
441 pattern of SST trend (Fig. 8b), and the corresponding PC2 time series also has a  
442 significant trend (slope=1.71,  $p<0.01$ ). The correlation between PC2 and Fz is  
443 significant ( $r=-0.56$ ,  $p=0.016$ ) and the correlation coefficient remains significant ( $r=-$   
444  $0.46$ ,  $p=0.065$ ) at the 90% confidence level when the value in 2002 is removed. This  
445 result suggests that the SST trend in SHtrop region is closely related to the recent  
446 weakening of stratospheric wave activities. The first EOF mode is similar to IPO (Fig.  
447 10a) and its corresponding principal component is significantly correlated ( $r=-0.98$ ,  
448  $p<0.01$ ) with the unfiltered IPO index. However, it shows no significant trend (Fig. 10d)  
449 and has no significant correlation (Fig. 10g) with stratospheric wave flux, implying that  
450 the linkage between the IPO phase change at around 2000 (e.g. Trenberth et al., 2013)  
451 and the weakening of Antarctic stratospheric wave activities is weak. The correlation  
452 between PC3 and Fz is also not significant (Fig. 10i). Therefore, it is possible that the  
453 combined effect of SST trend (the second EOF mode) in the tropical and extratropical  
454 southern hemisphere leads to the weakening of stratospheric wave activities in early  
455 austral spring since the early 2000s.

## 456 **6. Simulated changes in Antarctic stratospheric wave activities forced** 457 **by SST trends**

458 The analysis in Section 5 suggests that the SST changes in SHtrop region may  
459 contribute to the weakening of the southern hemispheric stratospheric wave activities.  
460 Here, numerical experiments sstNH, sstSH, ssttrop and sstSHtrop forced by linear  
461 increments of SST in September during 2000-2017 (Fig. 11; more details can be found  
462 in Section 2) are conducted to verify the results presented in Section 5.

463 Figure 12 shows the simulated response of 500 hPa geopotential height to SST  
464 changes in different regions. The climatological distributions of the wave-1 component  
465 (Figs. 12b, e, h, k) and the wave-2 component (Figs. 12c, f, i, l) from the simulations  
466 are consistent with that from reanalysis dataset (Figs. 3b, c), indicating that the model  
467 can well capture spatial distributions of the atmospheric waves. Note that the wave-1  
468 and wave-2 anomalies simulated with SST changes in SH, TROP and SHtrop are all  
469 significant. They superpose on the corresponding climatological patterns in an out-of-  
470 phase style (Figs. 12e, f, h, i, k, l), indicating that the changes in SST in SH, TROP and  
471 SHtrop lead to a weakening of tropospheric wave sources in the extratropical southern  
472 hemisphere. However, the predominate wave-1 component of the 500 hPa geopotential  
473 height anomaly in the extratropical southern hemisphere forced by the experiment with  
474 NH SST change is relatively weak (Fig. 12b). This feature suggests that the SST  
475 changes in extratropical northern hemisphere are incapable of inducing a robust  
476 response of tropospheric wave sources in the extratropical southern hemisphere.

477 Figure 13 shows the simulated responses of stratospheric wave activities in the  
478 southern hemisphere to SST changes over different regions. It is apparent that the  
479 experiments with SST changes in SH, TROP and SHtrop show significantly weakened

480 stratospheric wave activities (Figs. 13d, g, j), which are mainly attributed to the  
481 responses of the wave-1 component (Figs. 13e, h, k). These results are consistent with  
482 the responses of tropospheric wave sources (Figs. 12d, e, g, h, j, k). However, there are  
483 no significant anomalies of stratospheric wave flux in the subpolar region in Figures  
484 13a and 13b, which is consistent with the response of corresponding tropospheric wave  
485 sources (Figs. 12a, b) and the weak correlation between Fz and PC time series of SST  
486 in NH region (Fig. 9i). The result here suggests that the response of southern  
487 hemisphere stratospheric wave activities to SST trend in NH region is weak.

488       The results of stratospheric vertical wave flux over 50°S-70°S derived from the  
489 100 ensemble members of each experiment are shown in Figure S4, and the frequency  
490 distributions of them are displayed in Figure 14. Compared to the blue fitting curves,  
491 the red fitting curves shift to the left as shown in Figs. 14b, 14c and 14d, suggesting  
492 that the SST changes in SH, TROP and SHtrop regions weaken the upward propagation  
493 of stratospheric wave flux. The area-weighted anomalies of vertical E-P flux in the  
494 subpolar region of the southern hemisphere induced by SST changes in SH, TROP and  
495 SHtrop regions are  $-0.084 \times 10^5 \text{ kg} \cdot \text{s}^{-2}$ ,  $-0.12 \times 10^5 \text{ kg} \cdot \text{s}^{-2}$  and  $-0.13 \times 10^5 \text{ kg} \cdot \text{s}^{-2}$ ,  
496 respectively. The sum of the anomalies forced by sstSH and ssttrop is not equal to the  
497 anomaly forced by sstSHtrop, which may be resulted from non-linear interactions  
498 between the responses of wave activities to SST trends in SH region and TROP region.  
499 The weakening of stratospheric wave activities forced by SST increment in the tropical  
500 region is more significant than that in extratropical southern hemisphere (Figs. 14b, c,  
501 e), implying that the SST trend in the tropical region contributes more to the weakening



502 of stratospheric wave activities since 2000. Meanwhile, it is apparent that the  
503 weakening of the southern hemisphere stratospheric wave activities forced by  
504 sstSHtrop is the most significant among all the sensitive experiments (Fig. 14e). The  
505 reduction of vertical E-P flux over (50°S-70°S, 200 hPa-10 hPa) forced by sstSHtrop is  
506 approximately 12%. These modeling results indicate that the weakening of the  
507 Antarctic stratospheric wave activities in September since 2000 is induced mainly by  
508 the combined effects of SST trends in the tropical and extratropical southern  
509 hemisphere. It also explains why the independent correlation between Fz and PC time  
510 series obtained over SH or TROP region is not as significant as that between Fz and PC  
511 time series obtained over SHtrop region (Figs. 9g, h, j). Moreover, the mean linear  
512 increment of area-weighted vertical E-P flux from 200 hPa to 10 hPa over 70°S-50°S  
513 in September during 2000-2017 derived from four reanalysis datasets is about -  
514  $0.38 \times 10^5 \text{ kg} \cdot \text{s}^{-2}$ . Therefore, the contribution of SST trend over 20°N-70°S (the SHtrop  
515 region) to the weakening of stratospheric activities is approximately 34%.

516 In addition, the reanalysis datasets show that the Brewer-Dobson circulation  
517 related to wave activities in the stratosphere weakened significantly in early austral  
518 spring during 2000-2017 (Fig. 15b), which is contrary to the intensified trend during  
519 1980-2000 (Fig. 15a). The transition of BDC around 2000 is believed to be associated  
520 with ozone depletion and recovery (e.g., Polvani et al., 2017; Polvani et al., 2018).  
521 However, our modeling results suggest that the SST trend is responsible for the  
522 weakening of BDC in September since 2000 (Figs. 15d, e, f), The response of BDC to  
523 ozone recovery is not significant (Fig. 15c) in September, especially for the branch near

524 the Antarctic. These results indicate that apart from the ozone depletion and recovery  
525 the SST trend should also be taken into consideration when exploring the mechanism  
526 for the climate transition in the southern hemispheric stratosphere around 2000.

527 Previous studies reported that there is usually a time lag for tropic SST to affect  
528 extratropical circulation (e.g., Shaman & Tziperman, 2011). Thus, the impact of tropical  
529 SST change before September needs to be further examined. Our simulations indicate  
530 that the tropical SST trend in September plays a dominate role in weakening of  
531 stratospheric wave activity at the same month, and the effect of tropical SST change  
532 before September is negligible compared to that in September (The detailed evidences  
533 to address this issue are shown in the appendix).

## 534 **7. Conclusions and Discussions**

535 This study analyzes the trend of Antarctic stratospheric planetary wave activities  
536 in early austral spring since the early 2000s based on various reanalysis datasets and  
537 model simulations. Using the change-point method, we find that the Antarctic  
538 stratospheric wave activities in September have been weakening significantly since  
539 2000, which means the intensified trend of wave activities noted in previous researches  
540 (Hu & Fu, 2009; Lin et al., 2009) are reversed after 2000 in early austral spring. Further  
541 analysis suggests that the weakening of stratospheric wave activities is related to the  
542 weakening of tropospheric wave sources in extratropical southern hemisphere, which  
543 is mainly contributed by the wave-1 component.

544 As the Antarctic ozone also shows clear shift around the 2000, we firstly examine  
545 the impact of ozone recovery on Antarctic stratospheric planetary wave activity. Our

546 simulation results indicate that significant ozone recovery in lower stratosphere changes  
547 the atmospheric state for wave propagation to some extent, inducing a slight decrease  
548 of vertical wave flux over UTLS region in subpolar southern hemisphere. Meanwhile,  
549 the changes of wave activity in middle and upper stratosphere over subpolar region  
550 induced by ozone recovery are not significant. Therefore, the ozone recovery has minor  
551 contribution to the significant weakening of stratospheric planetary wave activity in  
552 September.

553 EOF analysis and correlation analysis indicate that the stratospheric wave  
554 activities in early austral spring during 2000-2017 are related to PC2 of SST over 20°  
555 N-70°S (i.e., the SHtrop region). The corresponding EOF2 mode also shows a good  
556 similarity to the spatial pattern of SST trend, suggesting that the weakening of  
557 stratospheric wave activities is connected to the trend of SST in SHtrop region.  
558 Meanwhile, the linkage between the SST trend in NH region and the weakening of  
559 stratospheric wave activities is weak. The model simulations also support that the SST  
560 changes in SHtrop region lead to a weakening of tropospheric wave sources and  
561 stratospheric wave activities. The contribution of SST trend in tropical region to the  
562 weakening of stratospheric wave activities is larger than that in the extratropical  
563 southern hemisphere. However, the response of tropospheric wave sources and  
564 stratospheric wave activities to SST trend in NH region is not significant. The  
565 contribution of SST trend over SHtrop region to the weakening of stratospheric wave  
566 activities is about 34%. Finally, both reanalysis datasets and numerical simulations  
567 indicate that the Brewer-Dobson circulation related to stratospheric wave activity has

568 also been weakening in early austral spring since 2000, which is also attributed to the  
569 changes of September SST in tropics and extratropical southern hemisphere.

570 Although many researchers claimed that the climate transition around 2000 in  
571 southern hemisphere is related to ozone depletion and recovery (e.g., Barnes et al., 2013;  
572 Banerjee et al., 2020), there is no contradiction between our results and these previous  
573 studies. Firstly, the southern hemisphere tropospheric circulation (i.e., the SAM index,  
574 the tropospheric jet position and the Hadley cell edge) shifts related to ozone changes  
575 in these previous studies basically occurred in austral summer (e.g., Son et al, 2008;  
576 Thompson et al., 2011; Barnes et al, 2013; Banerjee et al., 2020). These tropospheric  
577 circulation changes are induced by downward coupling of circulation anomalies in the  
578 stratosphere (e.g., Thompson et al., 2011) during October and November, when solar  
579 radiation covers the entire Antarctic and causes heating effects. However, the Antarctic  
580 stratospheric circulation response to ozone variation in September is not as strong as  
581 that in October or November (e.g., Thompson et al., 2011, Figs. 1b, d) because solar  
582 radiation can only reach part of Antarctic stratosphere during a majority period of  
583 September. This implies that the response of atmospheric state in September to  
584 Antarctic stratospheric ozone change is not significant. Secondly, the FWSC  
585 component used in this study is an atmospheric module with prescribed SST and forcing  
586 gases. Therefore, our model results only indicate that the weakening of stratospheric  
587 wave activity can be attributed to SST changes, while the impact of ozone change in  
588 middle and low latitudes on SST cannot be determined based on these simulations.  
589 Whether the transition signal of Antarctic stratospheric ozone is stored in the ocean

590 needs more efforts to explore. This is an issue beyond the scope of this study and further  
591 investigation is necessary by using a fully coupled earth system model.

592 The southern hemisphere stratospheric wave activity trend from the early 1980s to  
593 the early 2000s has been investigated by Hu and Fu (2009) (hereafter HF2009) and  
594 hence is not discussed in detail in the above. HF2009 attributed the strengthening of  
595 stratospheric wave activity in austral spring during 1979-2006 to the SST trends as well,  
596 however, they gave no more details about the trends of tropospheric wave sources. In  
597 this study, trends of tropospheric wave sources in September during 1980-2000 derived  
598 from MERRA-2 data is analyzed, and we also conducted an experiment (sstSHtrop)  
599 forced by the changes in September SST during 1980-2000 over 20°N-70°S (see Fig.  
600 S9 for applied SST anomalies). The model result indicates that the SST changes over  
601 20°N-70°S contribute to intensification of wave-2 component of tropospheric wave  
602 sources (Fig. S10f) and weakening of the wave-1 component (Fig. S10e), which is  
603 overall analogous to the trends derived from MERRA-2 data (Figs. S10b, c).  
604 Accordingly, the simulated wave-2 component of wave flux increases significantly in  
605 the stratosphere (Fig. S10h), while the response of the wave-1 component is not  
606 significant (Fig. S10i). In a word, the results from sstSHtrop80 suggest that the SST  
607 changes over 20°N-70°S induce a strengthening of stratospheric wave activity in  
608 September during 1980-2000. But it cannot explain the intensified wave-1 component  
609 of the stratospheric wave activity shown in Fig. 1b. A more detailed attribution of the  
610 trend of Antarctic stratospheric wave activity during 1980-2000 needs much more  
611 efforts.

612 The simulated stratospheric eddy heat flux (Fig. 11b in HF2009) forced by  
613 observed time-varying SST in HF2009 is relatively weak compared to that derived from  
614 reanalysis data (Fig. 6b in HF2009). Similarly, Wang and Waugh (2012) (hereafter  
615 WW2012) used stratosphere-resolving chemistry-climate model forced by time-  
616 varying factors to evaluate the trends of stratospheric temperature, residual circulation  
617 as well as wave activity during recent decades, and the trend of cumulative eddy heat  
618 flux shown in their paper is not significant (Fig. 6 in WW2012). Additionally, Polvani  
619 et al. (2018) used time-varying ODSs that cover the period from 1960s to 2080s to  
620 simulate Brewer-Dobson circulation and attained an obvious trend transition around  
621 2000. We had also tried to conduct transient experiments forced by time-varying SST  
622 derived from ERSST v5 with different initial conditions, however, the trends of wave  
623 activities in the transient simulations are so weak, though the opposite trend signs exist  
624 during 1980-2000 and 2000-2018 (Table S2, Fig. S11). The significance of simulated  
625 trend may be related to model performance and the length of simulating period. As the  
626 period we focus is relatively short and our purpose is attribution rather than generating  
627 a real trend, we perform the ensemble time-slice experiments in this study, which are  
628 also used in many other previous researches (e.g., Hu et al., 2018; Kang et al., 2011;  
629 Zhang et al., 2016) to attribute trends in the atmosphere. In addition, most of the current  
630 climate models cannot generate a realistic wave activity trend as waves in the  
631 atmosphere are linked with various processes and factors (e.g., Baldwin & Dunkerton,  
632 2005; Garcia & Randel, 2008; Labitzke, 2005; Shindell et al., 1999; Shu et al., 2013;  
633 Xie et al., 2008).

634

635 **Data availability:**

636 The ERA-Interim is available at: <https://apps.ecmwf.int/datasets/data/interim->  
637 [full-daily/levtype=sfc/](https://apps.ecmwf.int/datasets/data/interim-full-daily/levtype=sfc/). The MERRA-2 is available at: [https://disc.gsfc.nasa.gov/d](https://disc.gsfc.nasa.gov/datasets?keywords=%22MERRA-2%22&page=1&source=Models%2FAnalyses%20MERRA-2)  
638 [atasets?keywords=%22MERRA-2%22&page=1&source=Models%2FAnalyses%20M](https://disc.gsfc.nasa.gov/datasets?keywords=%22MERRA-2%22&page=1&source=Models%2FAnalyses%20MERRA-2)  
639 [ERRA-2](https://disc.gsfc.nasa.gov/datasets?keywords=%22MERRA-2%22&page=1&source=Models%2FAnalyses%20MERRA-2). The JRA-55 is available at: [https://jra.kishou.go.jp/JRA-55/index\\_en.ht](https://jra.kishou.go.jp/JRA-55/index_en.html#download)  
640 [ml#download](https://jra.kishou.go.jp/JRA-55/index_en.html#download). The NCEP-2 is available at: [http://www.cpc.ncep.noaa.gov/product](http://www.cpc.ncep.noaa.gov/products/wesley/reanalysis2/)  
641 [s/wesley/reanalysis2/](http://www.cpc.ncep.noaa.gov/products/wesley/reanalysis2/). The ERSST v5 dataset is available at: [https://www1.ncdc.](https://www1.ncdc.noaa.gov/pub/data/cmb/ersst/v5/netcdf/)  
642 [noaa.gov/pub/data/cmb/ersst/v5/netcdf/](https://www1.ncdc.noaa.gov/pub/data/cmb/ersst/v5/netcdf/). The observations of TCO from SBUV v  
643 8.6 satellite dataset are available at: [https://acd-ext.gsfc.nasa.gov/Data\\_services/m](https://acd-ext.gsfc.nasa.gov/Data_services/merged/data/sbu_v86_mod.int_1yr.70-18.za.r7.txt)  
644 [erged/data/sbu\\_v86\\_mod.int\\_1yr.70-18.za.r7.txt](https://acd-ext.gsfc.nasa.gov/Data_services/merged/data/sbu_v86_mod.int_1yr.70-18.za.r7.txt). The unfiltered IPO index derived  
645 from ERSST v5 dataset is available at: [https://psl.noaa.gov/data/timeseries/IPOT](https://psl.noaa.gov/data/timeseries/IPOTPI/tpi.timeseries.ersstv5.data)  
646 [PI/tpi.timeseries.ersstv5.data](https://psl.noaa.gov/data/timeseries/IPOTPI/tpi.timeseries.ersstv5.data).

647 **Author contributions:**

648 Yihang Hu conducted experiments, produced figures and tables, organized and  
649 wrote the manuscript. Wenshou Tian, Jiankai Zhang and Tao Wang contributed to revise  
650 the manuscript. Mian Xu helped to design experiments.

651 **Competing interests:**

652 The authors declare that they have no competing interest.

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663

664

## APPENDIX

### 665 **Analysis of time lag for tropical SST affects Antarctic stratospheric** 666 **wave activity**

667 As stated in the Section 2, the tropical SST anomalies (the linear increments) in  
668 experiment ssttrop are also applied in July and August (Fig. S5a, b) to avoid abrupt SST  
669 variations from month to month, and the two months are taken as spin-up time.  
670 Therefore, whether the SST forcing in July and August also contribute to the weakening  
671 of Antarctic stratospheric wave activity in September or not cannot be justified based  
672 on the experiment ssttrop only. Here, we performed an additional experiment  
673 ssttropAug without September SST anomalies (Fig. S5f) to clarify whether the  
674 weakening of Antarctic stratospheric wave activity is induced by the tropical SST trend  
675 at the same month. Like other numerical experiments described in Table 1, the  
676 ssttropAug also includes 100 ensemble members that run from July to September forced  
677 by the same initial conditions from the 21st year to the 120th year in July generated by



678 free run. The detailed descriptions of ssttropAug and other relevant experiments in the  
 679 manuscript are displayed together in the Table S1 for comparison. Figure S5 shows the  
 680 applied global SST anomalies in ssttrop and ssttropAug from July to September.

681 The responses of tropospheric wave sources and stratospheric wave activities in  
 682 ssttropAug are shown in Figs. S6a-c and Figs. S6d-f, respectively. Note that the  
 683 anomalies of subpolar tropospheric geopotential height in September forced by changes  
 684 in tropical SST in August does not superpose on their climatological patterns in an  
 685 evident out-of-phase style (Figs. S6a-c). The anomaly of wave-1 component of  
 686 geopotential height shows a slight in-phase overlap with its climatology over subpolar  
 687 region (Fig. S6b). Accordingly, the responses of stratospheric wave activities over  
 688 subpolar of southern hemisphere are not significant (Figs. S6d-f). The results here  
 689 suggest that, the decrease of September vertical wave flux induced by SST changes in  
 690 August is negligible comparing to that in the experiment with anomalous SST forcing  
 691 in September (Figs. S6g), and the tropical SST trend in September plays a dominate  
 692 role in weakening of stratospheric wave activity at the same month.

693 Furthermore, we also use a linear barotropic model (LBM) (e.g., Shaman &  
 694 Tziperman, 2007; Shaman & Tziperman, 2011) to quantify the time scale for  
 695 propagation of tropical anomalies to high latitudes. The LBM are developed to solve  
 696 the barotropic vorticity equation, which is given as Eq. (A1):

$$697 \quad J(\bar{\psi}, \nabla^2 \psi') + J(\psi', \nabla^2 \bar{\psi} + f) + \alpha \nabla^2 \psi' + K \nabla^4 \nabla^2 \psi' = R \quad (\text{A1})$$

698 where the Jacobian  $J(A, B)$  is

$$699 \quad J(A, B) = \frac{1}{r^2} \left( \frac{\partial A}{\partial \lambda} \frac{\partial B}{\partial \mu} - \frac{\partial A}{\partial \mu} \frac{\partial B}{\partial \lambda} \right) \quad (\text{A2})$$

700 the forcing function  $R$  is

$$701 \quad R = -(f + \nabla^2 \bar{\psi})D \quad (A3)$$

702  $\psi$  is the streamfunction,  $f$  is the Coriolis force,  $\alpha$  is the Rayleigh coefficient,  $K$   
703 is the diffusion coefficient,  $\lambda$  is the longitude,  $\mu = \sin(\theta)$ ,  $\theta$  is the latitude,  $r$  is  
704 the earth's radius and  $D$  is the divergence.

705 We use the wave-1 component of streamfunction derived from ensemble mean of  
706 sstctrl as the background field. In LBM, the initial anomaly is given by the divergence.  
707 The divergence forcing field is limited in 40°E-140°W, 10°S-0° (Fig. S7) to ensure that  
708 the tropical initial anomaly of streamfunction superpose on its background field in an  
709 out-of-phase style. We set  $D = -7.9 \times 10^{-7} \text{ s}^{-1}$ , which is the mean divergence over the  
710 forcing region. The LBM simulated streamfunction anomalies are shown in Figs. S8b-  
711 i. Note that the anomalies in tropics only take a few days to arrive the high latitudes in  
712 southern hemisphere. After about four days, a stable anti-phase superposition of  
713 streamfunction is well established in extratropical southern hemisphere (Figs. S8f-i).  
714 These results are supported by previous studies (e.g., Shaman & Tziperman, 2011),  
715 which also indicate that the horizontal propagation of anomaly in atmosphere takes a  
716 few days.

717 Previous studies also reported that it takes about 4 days for wave-1 to propagate  
718 from troposphere into stratosphere and 1-2 days for wave-2 (e.g., Randel, 1987). Thus,  
719 the tropical oceans affect the stratosphere at mid-high latitudes with a lag of several  
720 days. However, the SST forcing field applied in CESM is on monthly scale. It is  
721 reasonable to use September SST trend to drive and explain the trends of extratropical

722 circulation and wave activity at the same month.

723

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939  
940 **Table 1.** Configurations of experiments for SST trends.

Experiments	Descriptions
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sstctrl	Control run. Seasonal cycle of monthly mean global SST data over 1980-2000 is derived from the ERSST v5 dataset. Fixed values of ozone greenhouse gases and aerosol fields in 2000 are used.
sstNH	As in sstctrl, but with linear increments of SST in September over 2000-2017 in NH (20°N-70°N). The applied global SST anomalies are shown in Fig. 7a.
sstSH	As in sstctrl, but with linear increments of SST in September over 2000-2017 in SH (20°S-70°S). The applied global SST anomalies are shown in Fig. 7b.
ssttrop	As in sstctrl, but with linear increments of SST in September over 2000-2017 in the tropics (20°S-20°N). The applied global SST anomalies are shown in Fig. 7c.
sstSHtrop	As in sstctrl, but with linear increments of SST in September over 2000-2017 in SHtrop (20°N-70°S). The applied global SST anomalies are shown in Fig. 7d.

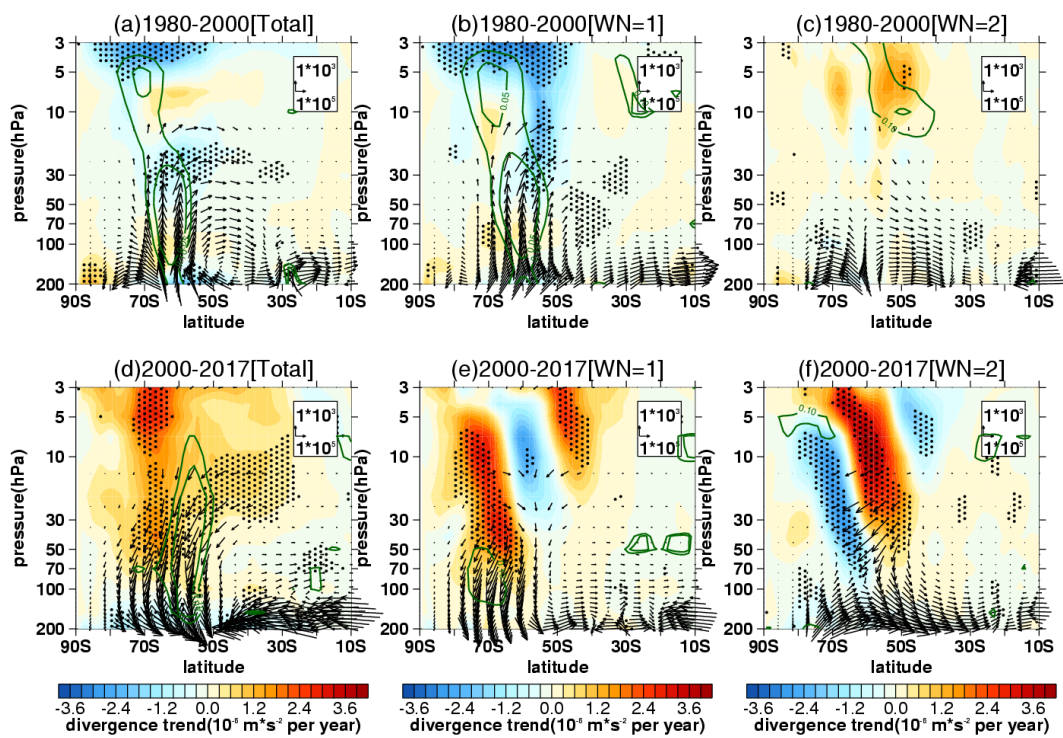
941 **Table 2.** Configurations of experiments for the ozone recovery trend.

Experiments	Descriptions
O3ctrl	Control run. The seasonal cycle of monthly averaged global SST data over 1980-2000 is derived from ERSST v5 dataset. The seasonal cycle of monthly mean three-dimensional global ozone over 1980-2000 is derived from MERRA-2 dataset. The GHGs and aerosol fields are specified to be fixed values in 2000.
O3sen	As in O3ctrl, but superposed with linear increments of global ozone in September over 2001-2017. The ozone data in 2002 are removed when the linear increments are calculated. The applied ozone anomalies in Southern Hemisphere are shown in Fig. 5.

942 **Table 3.** Correlations of stratospheric vertical wave flux time series (area-weighted  
943 from 100 hPa to 30 hPa over 70°S-50°S) between different reanalysis dataset.

	ERA-Interim	JRA-55	MERRA-2	NCEP-2
ERA-Interim	1.00 (p=0.00)	0.99 (p<0.01)	0.98 (p<0.01)	0.93 (p<0.01)
JRA-55		1.00 (p=0.00)	0.98 (p<0.01)	0.93 (p<0.01)
MERRA-2			1.00 (p=0.00)	0.94 (p<0.01)
NCEP-2				1.00 (p=0.00)

944

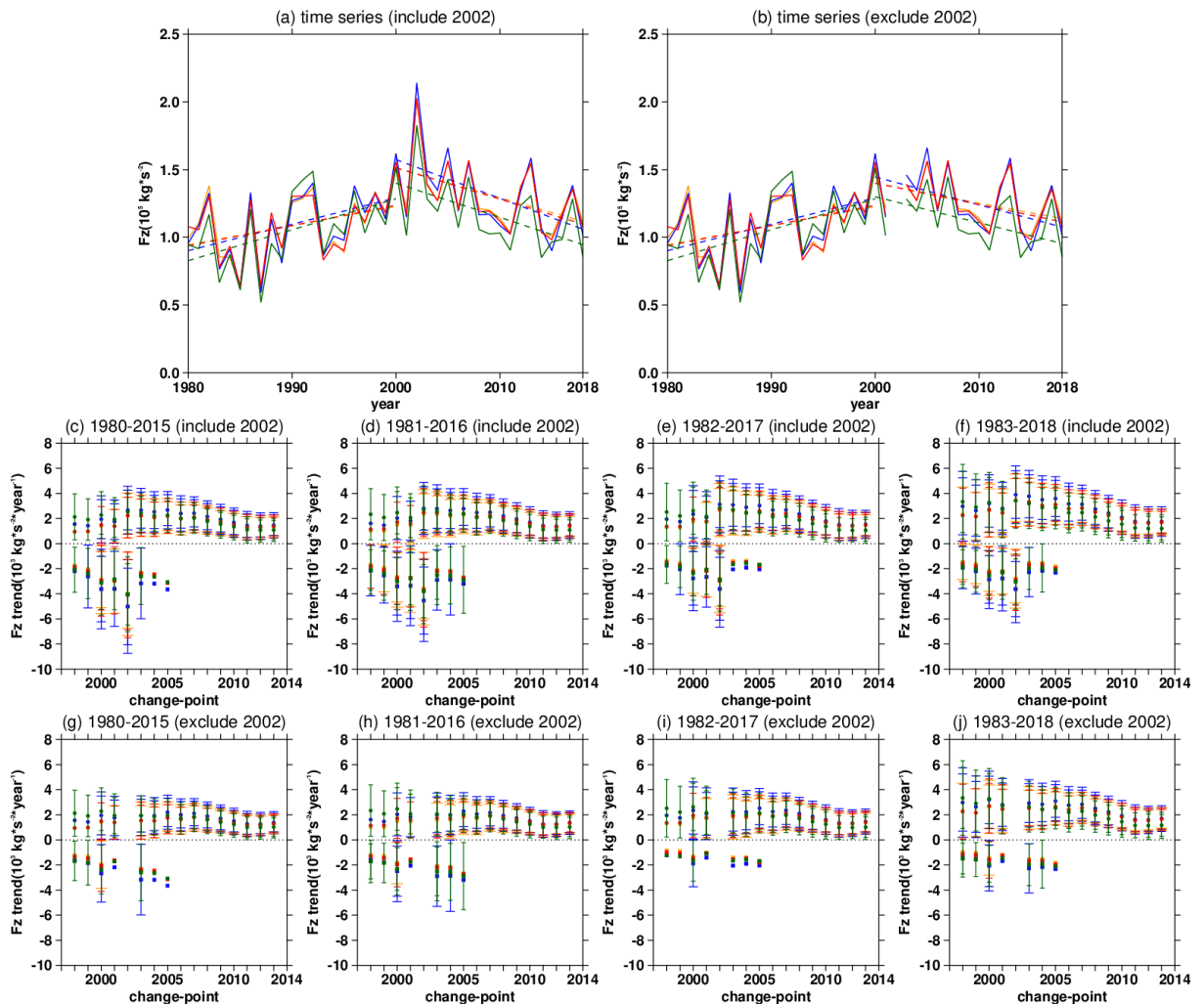


945

946 **FIG. 1.** Trends of southern hemisphere (a, d) stratospheric E-P flux (arrows, units of  
 947 horizontal and vertical components are  $10^5$  and  $10^3$   $\text{kg}\cdot\text{s}^{-2}$  per year, respectively) and its  
 948 divergence (shadings) with their (b, e) wave-1 components and (c, f) wave-2  
 949 components over (a, b, c) 1980-2000 and (d, e, f) 2000-2017 in September derived from  
 950 MERRA-2 dataset. The stippled regions indicate the trend of E-P flux divergence  
 951 significant at/above the 90% confidence level. The green contours from outside to  
 952 inside (corresponding to  $p=0.1$ ,  $0.05$ ) indicate the trend of vertical E-P flux significant

953 at the 90% and 95% confidence level, respectively.

954



955

956 **FIG. 2.** (a) The mean time series (solid lines) and piecewise (during 1980–2000 and

957 2000–2018) linear regressions (dashed lines) of vertical E-P flux area-weighted from

958 100 hPa to 30 hPa over 70°S–50°S in September during 1980–2018 derived from ERA-

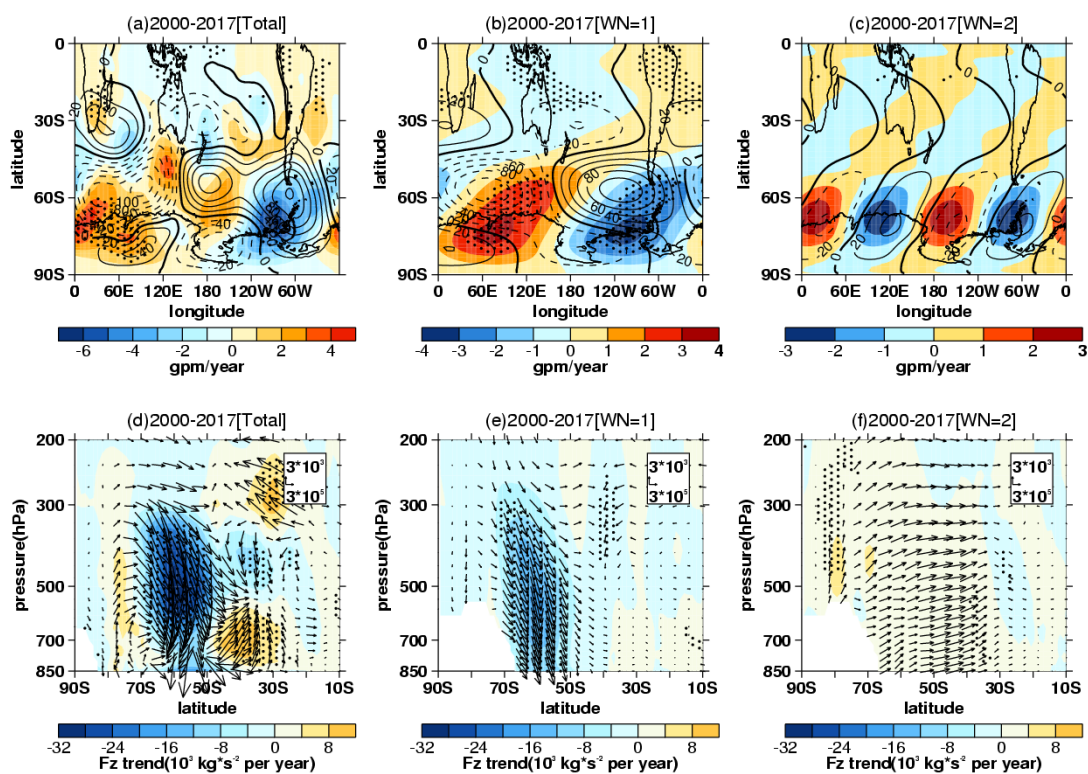
959 Interim (yellow), MERRA-2 (blue), JRA-55 (red) and NCEP-2 (green). Figure (b) is

960 the same as Figure (a), except for that the data in 2002 are removed. (c, d, e, f) The

961 trends (dots) and uncertainties (error bars) calculated during various periods using the

962 change-point method with different beginning and ending years (titles). Circles and

963 squares in Figures (c, d, e, f) represent positive trends from beginning years to change-  
 964 point years (x-axes) and negative trends from change-point years to ending years,  
 965 respectively. Different colors of dots and error bars in Figures (c, d, e, f) correspond to  
 966 colors in Figure (a), which represent trends and uncertainties derived from different  
 967 datasets. The long and short error bars in same color reflect the 95% and 90%  
 968 confidence intervals calculated by two-tailed t test. The error bar is omitted when the  
 969 significance of trend is lower than corresponding confidence level. Negative trends and  
 970 corresponding uncertainties with the beginning change-point years after 2005 are also  
 971 omitted, since the trend value shows large fluctuation with shortening of time series.  
 972 Figures (g, h, i, j) are the same as Figures (c, d, e, f), except that the data in 2002 are  
 973 removed when calculating trends and uncertainties.

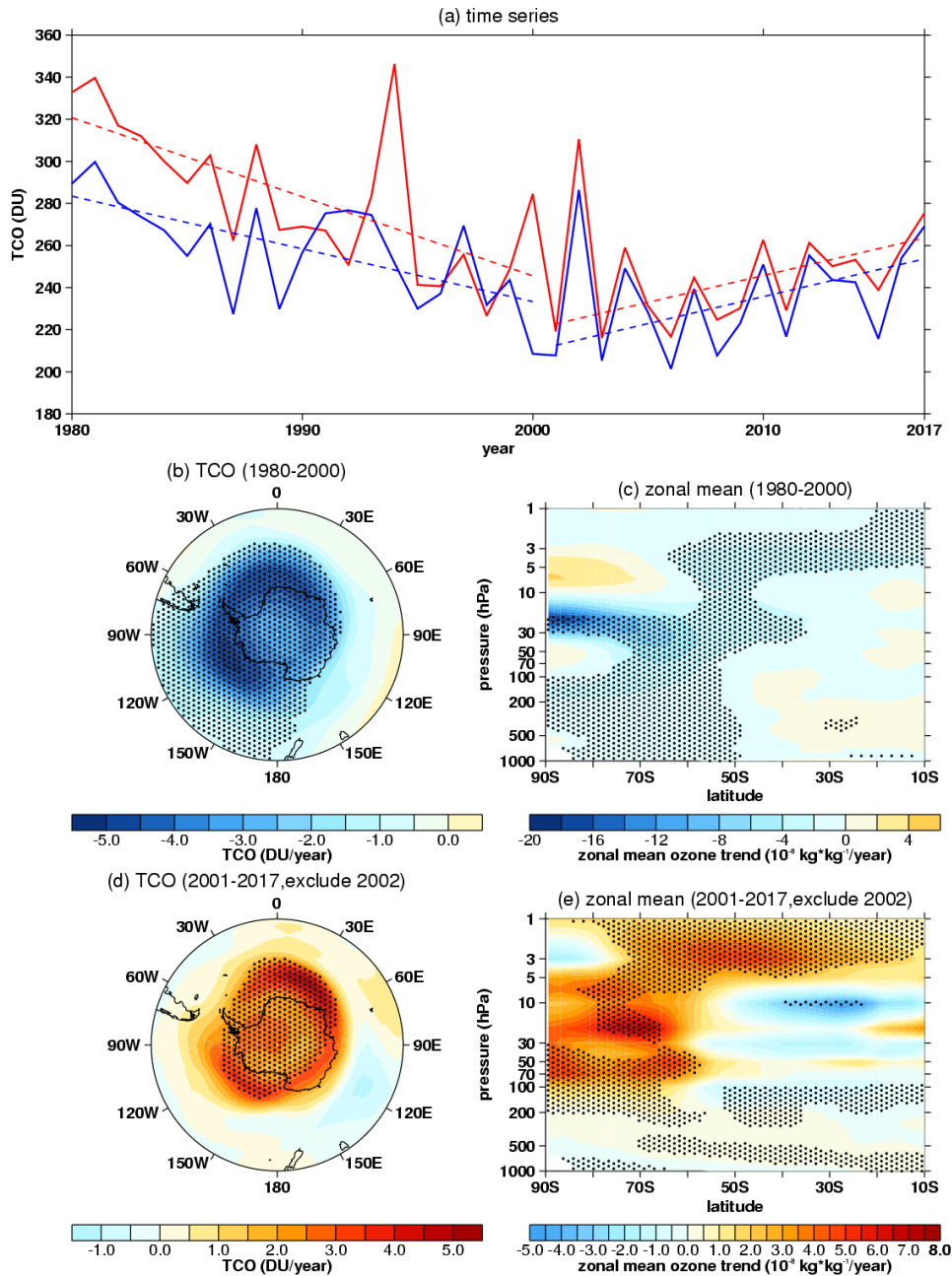


974

975 **FIG. 3.** Trends (shadings) and climatological distributions (contours with an interval  
 976 of 20 gpm, positive and negative values are depicted by solid and dashed lines



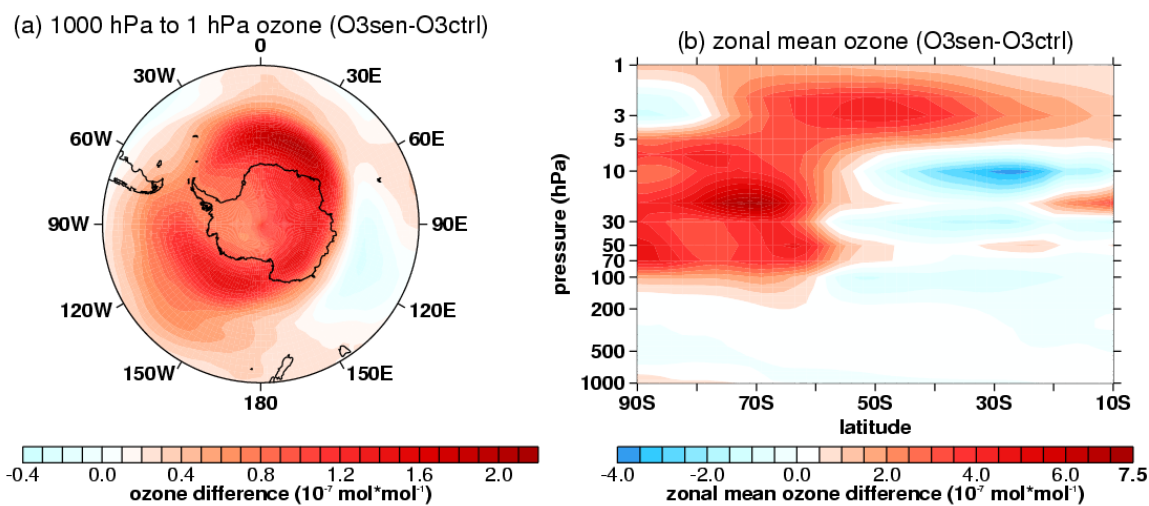
977 respectively, zeroes are depicted by thick solid lines) of southern hemispheric (a) 500  
978 hPa geopotential height zonal deviations with their (b) wave-1 component and (c)  
979 wave-2 component in September during 2000–2017 derived from MERRA-2 dataset.  
980 Trends of southern hemispheric (d) tropospheric E-P flux (arrows, units of horizontal  
981 and vertical components are  $3 \times 10^5$  and  $3 \times 10^3$  kg s<sup>-2</sup> per year, respectively) and its  
982 vertical component (shading) with their (e) wave-1 component and (f) wave-2  
983 component in September during 2000–2017 derived from MERRA-2 dataset. The  
984 stippled regions represent the trend significant at/above the 90% confidence level.



985

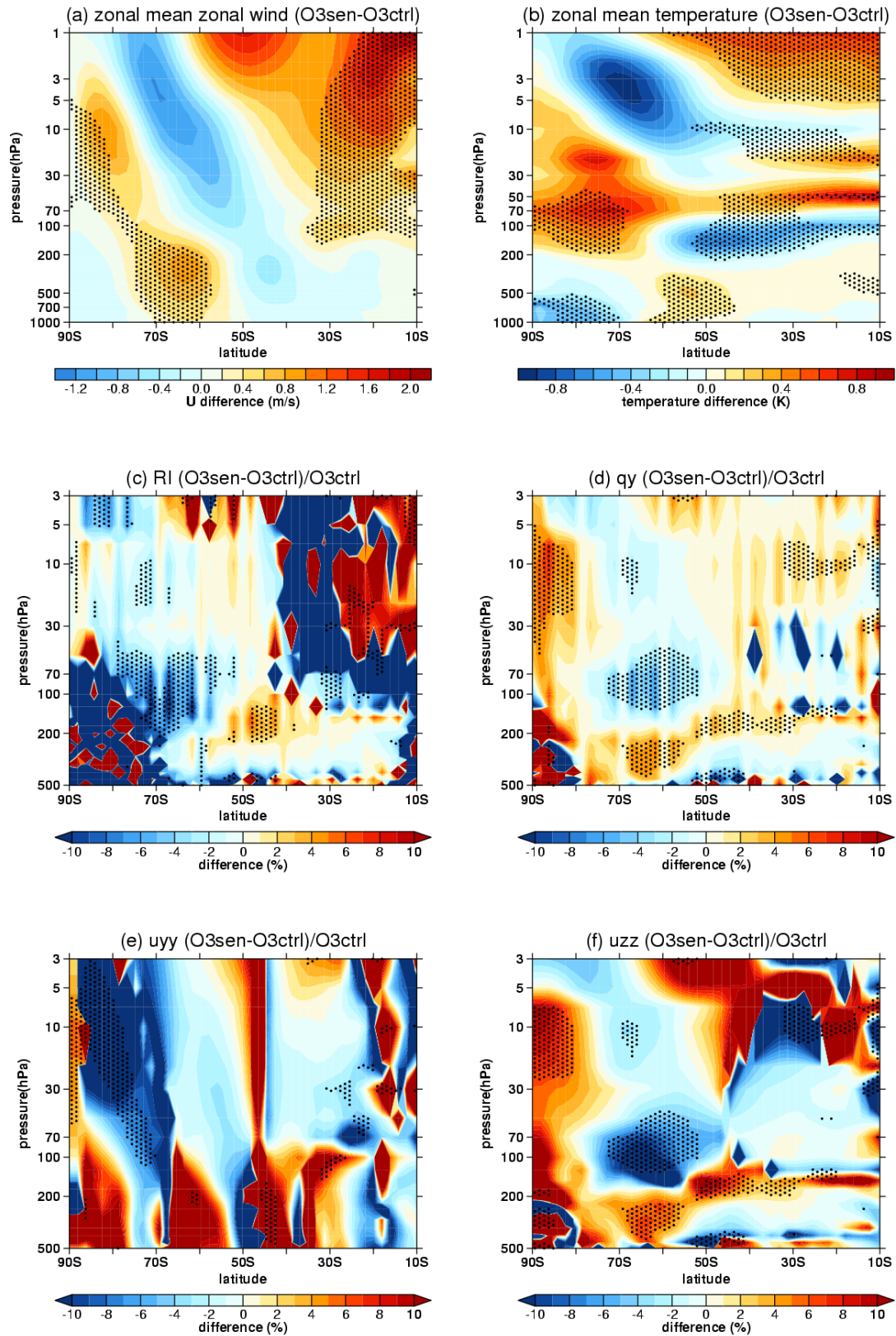
986 **FIG. 4.** (a) Time series (solid lines) of aera-weighted total column ozone (TCO) over  
 987 60°S to 90°S derived from MERRA-2 (red) and SBUV (blue) datasets. The dashed lines  
 988 represent linear regression of TCO. (b, d) The TCO trends in September during 1980-  
 989 2000 (b) and 2001-2017 (d) derived from MERRA-2 dataset. The outermost latitudes  
 990 in Figs. 4c, d are both 40°S. (c, e) The zonal mean ozone trends on latitude-pressure  
 991 profile in September during 1980-2000 (c) and 2001-2017 (e) derived from MERRA-2

992 dataset. The stippled regions in Figs. 4b-e represent trends significant at/above the 90%  
993 confidence level. Data in 2002 are removed when trends, regressions and significances  
994 are calculated in Fig. 4.



995

996 **FIG. 5.** (a) Difference of horizontal ozone forcing field averaged from 1000 hPa to 1  
997 hPa between  $O3_{sen}$  and  $O3_{ctrl}$ . The outermost latitude in Fig. 5a is  $40^\circ \text{S}$ . (b) Zonal  
998 mean difference of ozone forcing fields on latitude-pressure profile in the southern  
999 hemisphere between  $O3_{sen}$  and  $O3_{ctrl}$ .



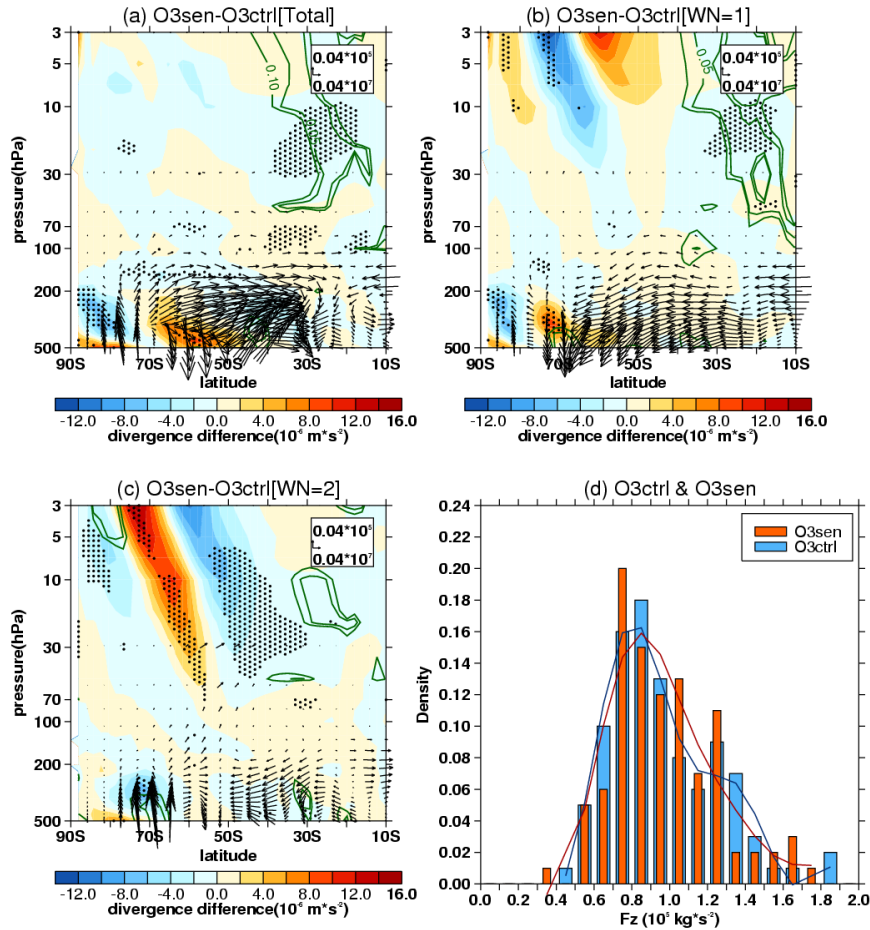
1000

1001 **FIG. 6.** Difference of (a) zonally averaged zonal wind, (b) zonally averaged

1002 temperature, (c) refractive index, (d)  $a^2 \bar{q}_\varphi$ , (e)  $-\left[\frac{(\bar{u} \cos \varphi)_\varphi}{\cos \varphi}\right]_\varphi$  (uyy term), (f)

1003  $-\frac{a^2 f^2}{\rho_0} (\rho_0 \frac{\bar{u}_z}{N^2})_z$  (uzz term) between O3sen and O3ctrl. The stippled regions represent

1004 the difference significant at/above 90% confidence level.



1005

1006 **FIG. 7.** Differences of (a) stratospheric E-P flux (arrows, units in horizontal and vertical

1007 components are  $0.04 \times 10^7$  and  $0.04 \times 10^5$  kg s<sup>-2</sup>, respectively) and its divergence

1008 (shadings) with their (b) wave-1 component and (c) wave-2 component between the

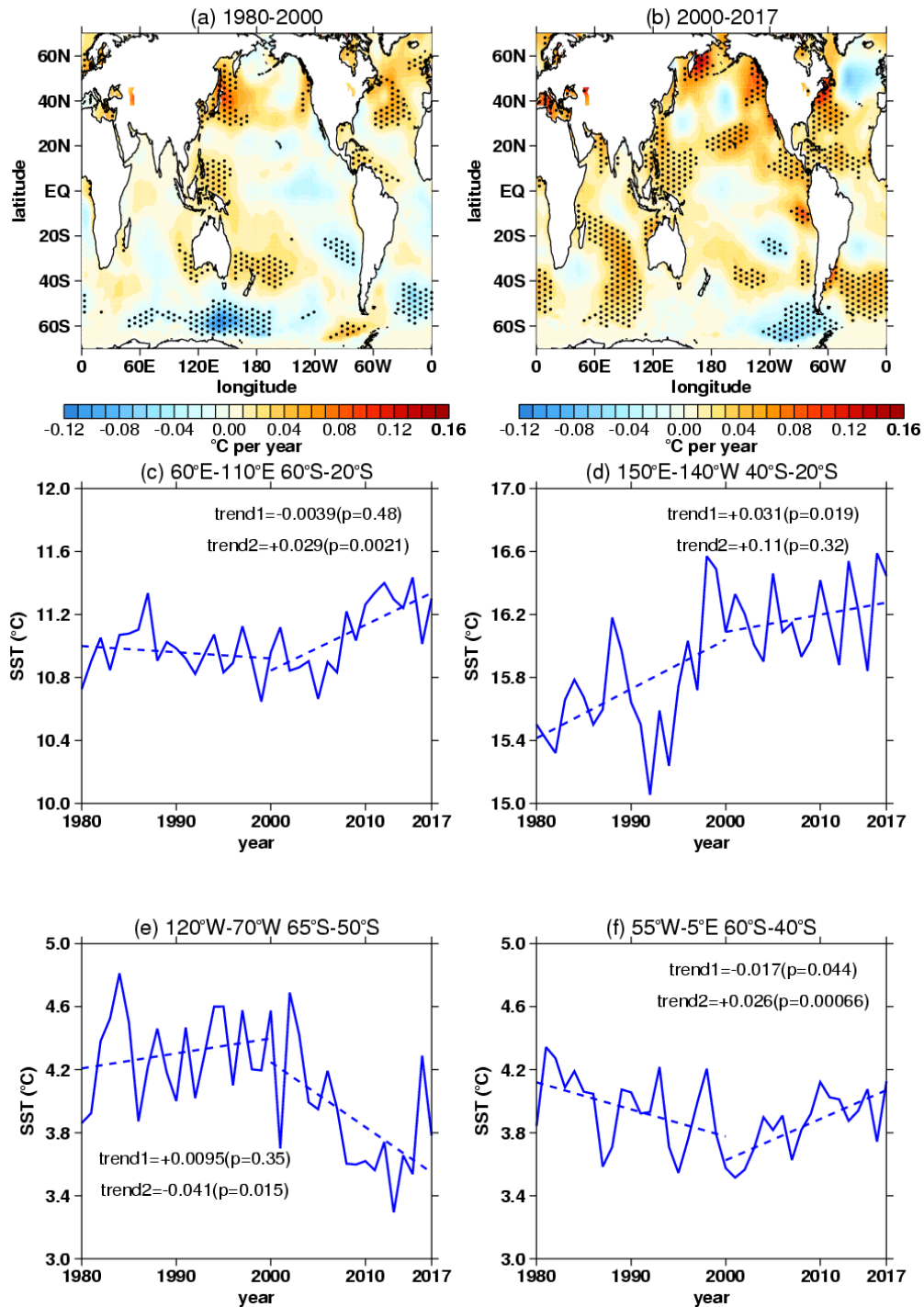
1009 sensitive experiment (O3sen) and the control experiment (O3ctrl). The stippled regions

1010 represent the mean differences of E-P flux divergence significant at/above the 90%

1011 confidence level. The green contours from outside to inside (corresponding to p=0.1,

1012 0.05) represent the mean differences of vertical E-P flux significant at the 90% and 95%

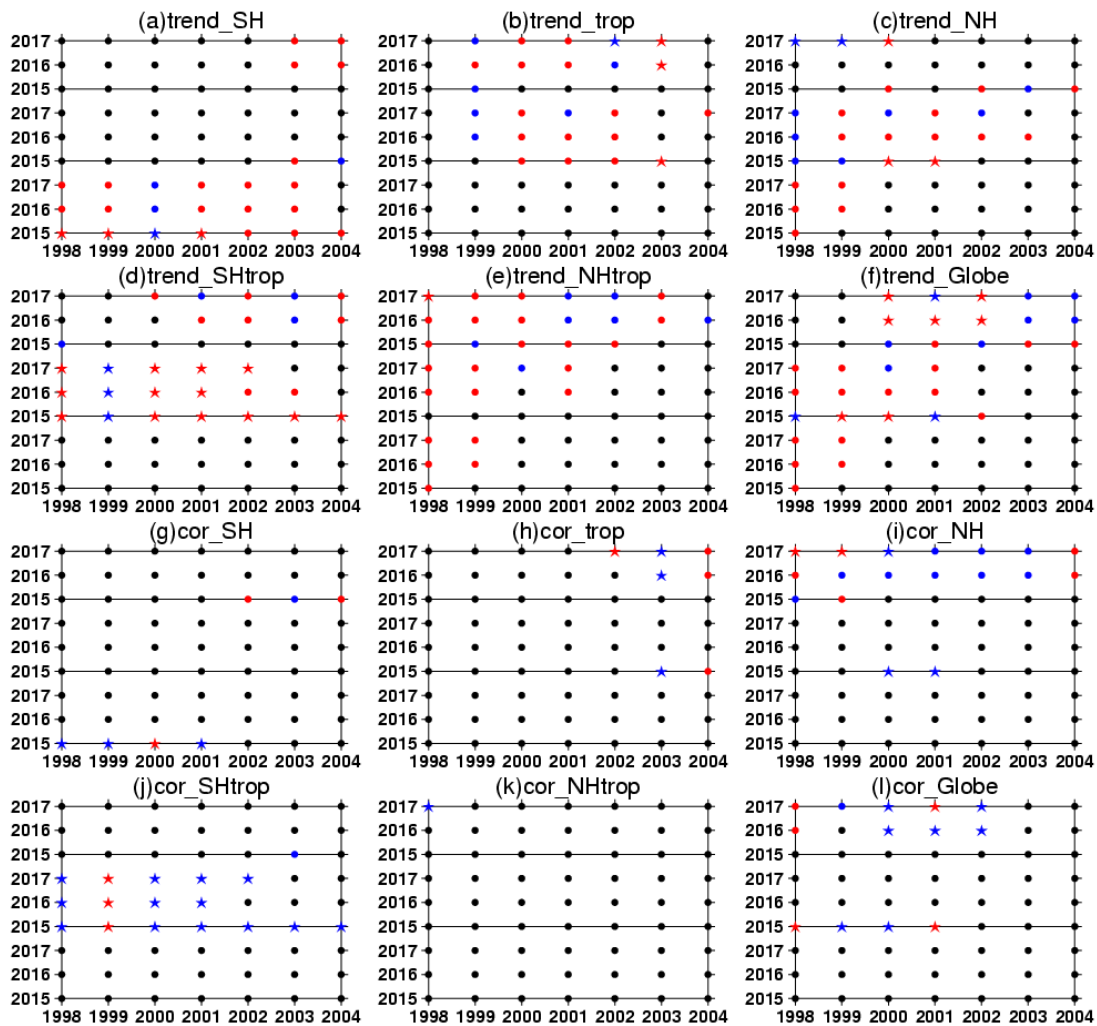
1013 confidence levels, respectively. (d) Frequency distributions (pillars, blue for O3ctrl and  
 1014 orange for O3sen) of vertical E-P flux (Fz, area-weighted from 200 hPa to 10 hPa over  
 1015 70°S-50°S) and its 5-point low-pass filtered fitting curves (solid lines, blue for O3ctrl  
 1016 and red for O3sen) derived from 100 ensemble members.



1017

1018 **FIG. 8.** Trends of SST in September over (a) 1980-2000 and (b) 2000-2017 derived

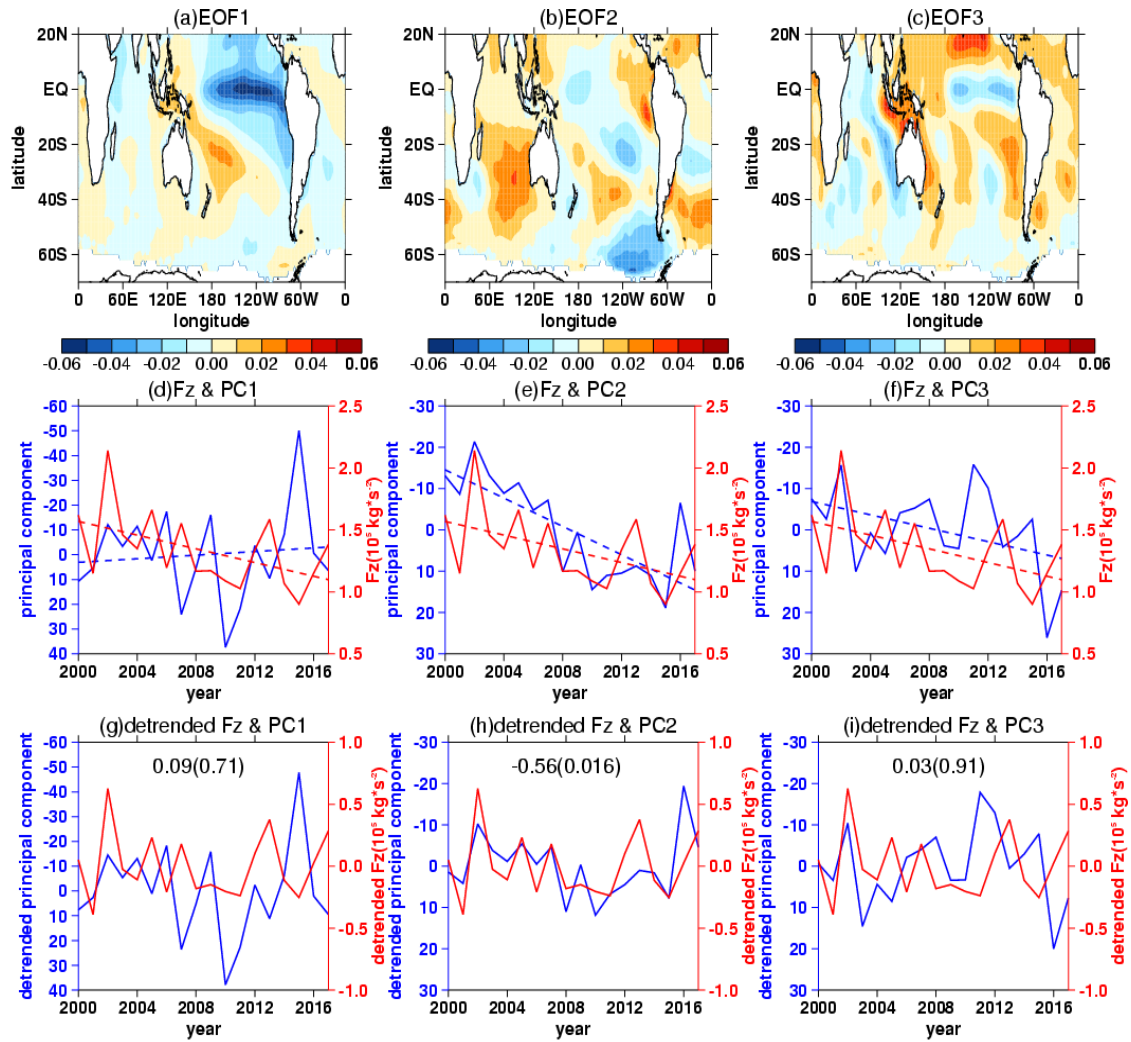
1019 from ERSST v5 dataset. The stippled regions represent the trends significant at/above  
 1020 the 90% confidence level. (c-f) Time series (blue solid lines) of SST during 1980-2017  
 1021 over different regions (titles). The dashed lines represent linear regressions of SST time  
 1022 series on piecewise periods (1980-2000 and 2000-2017). The “trend1” and “trend2”  
 1023 labeled in Figs. 8c-f represent the trend coefficients and the corresponding significances  
 1024 (bracketed) over 1980-2000 and 2000-2017, respectively.



1025  
 1026 **FIG. 9.** Trend significance of the first three SST principal components (PCs) in (a) the  
 1027 extratropical southern hemisphere (SH, 70°S-20°S), (b) the tropics (TROP, 20°S-20°N),  
 1028 (c) the extratropical northern hemisphere (NH, 20°N-70°N), (d) the extratropical  
 1029 southern hemisphere and the tropics (SHtrop, 70°S-20°N), (e) the extratropical northern

1030 hemisphere and the tropics (NHtrop, 20°S-70°N), (f) the globe (70°S-70°N) and the  
1031 corresponding (g, h, i, j, k, l) correlation significances between them and vertical E-P  
1032 flux (Fz, area-weighted from 100 hPa to 30 hPa over 70°S-50°S) during different  
1033 beginning years (x-axes) and ending years (y-axes). The red and blue dots indicate  
1034 positive and negative trend or correlation coefficient are significant, respectively. The  
1035 black dots indicate the trends or correlation coefficients are not significant. The stars  
1036 indicate that the trends and the corresponding correlation coefficients are both  
1037 significant. Each panel is divided into three regions from bottom to top, corresponding  
1038 to the first, the second and the third principal components, respectively. The criterion  
1039 to distinguish whether the trends and correlations are significant or not is the 90%  
1040 confidence level.





1041

1042 **FIG. 10.** (a, b, c) The first three EOF patterns of SST in SHtrop region. (d, e, f) The

1043 original time series of the first three principle components (PCs, blue solid lines

1044 correspond to left inverted y-axes) and stratospheric vertical E-P flux (Fz, area-

1045 weighted from 100 hPa to 30 hPa over 70°S-50°S, red solid lines correspond to right y-

1046 axes) in September during 2000-2017. The blue and red dashed lines in (d, e, f)

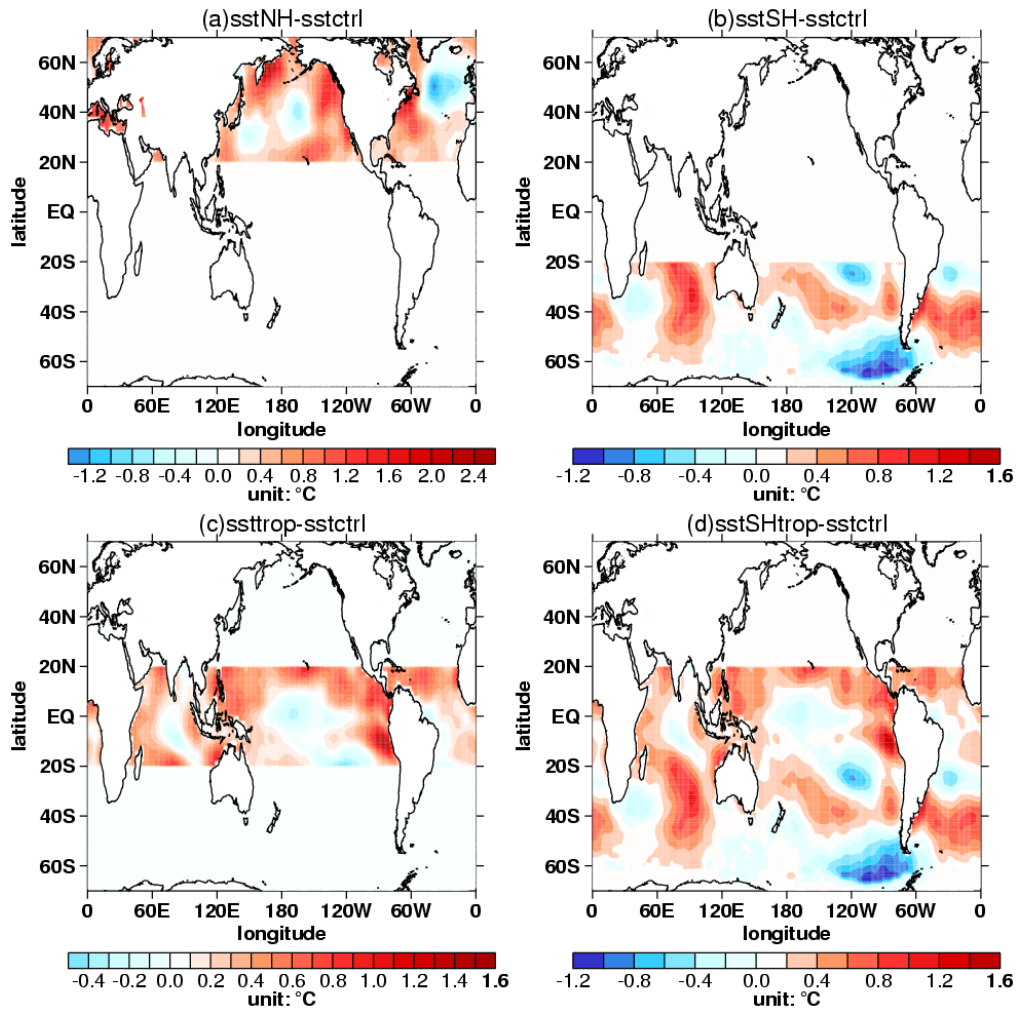
1047 represent the linear regressions of PC time series and Fz time series, respectively. The

1048 meaning of (g, h, i) are the same as (d, e, f) correspondingly, except the detrended time

1049 series. The unbracketed and bracketed numbers in (g, h, i) represent the correlation

1050 coefficients between detrended PC time series and Fz time series and the corresponding

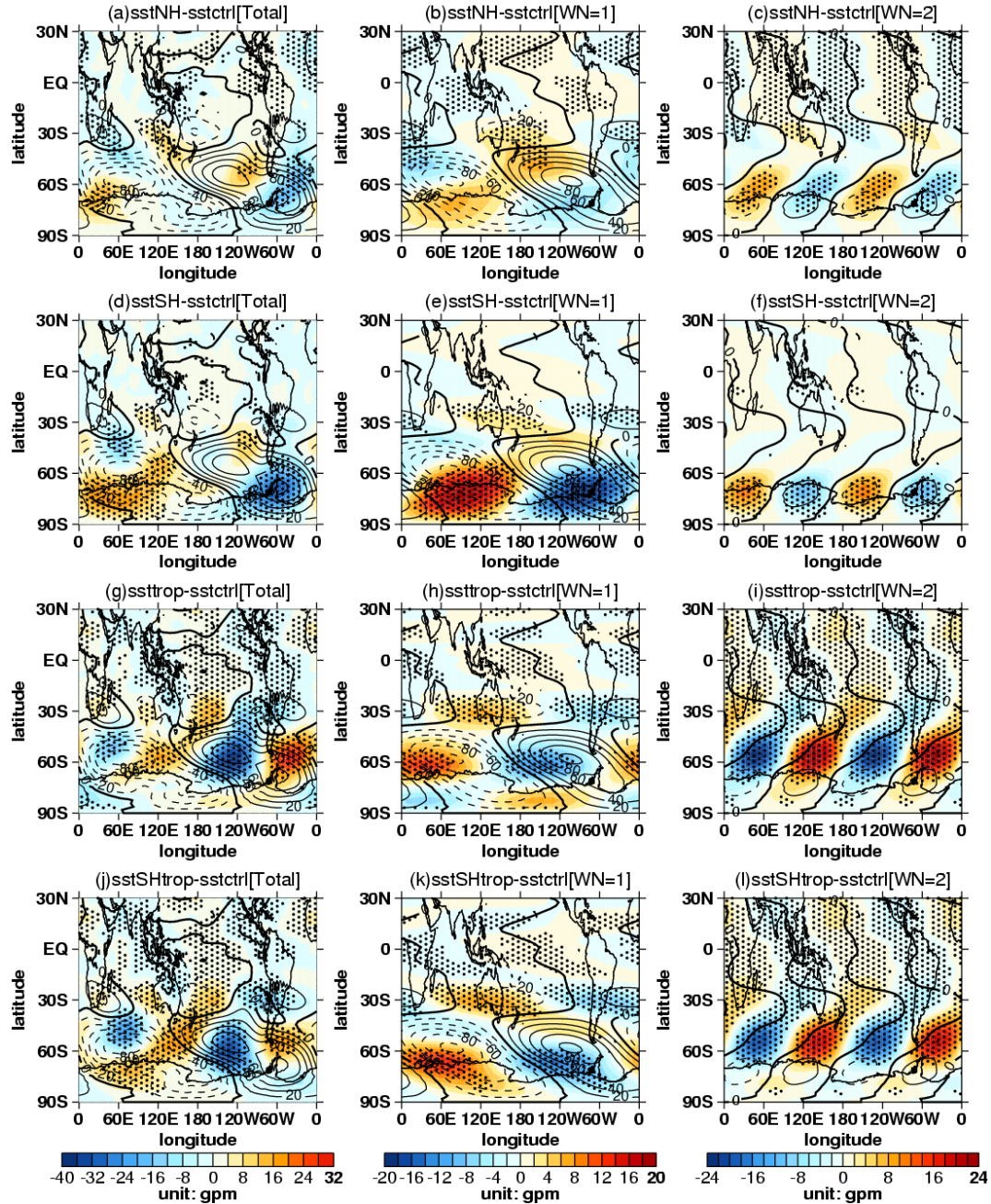
1051 p values calculated by two-tailed t test, respectively.



1052

1053 **FIG. 11.** Differences in SST forcing field between sensitive experiments ((a) sstNH; (b)

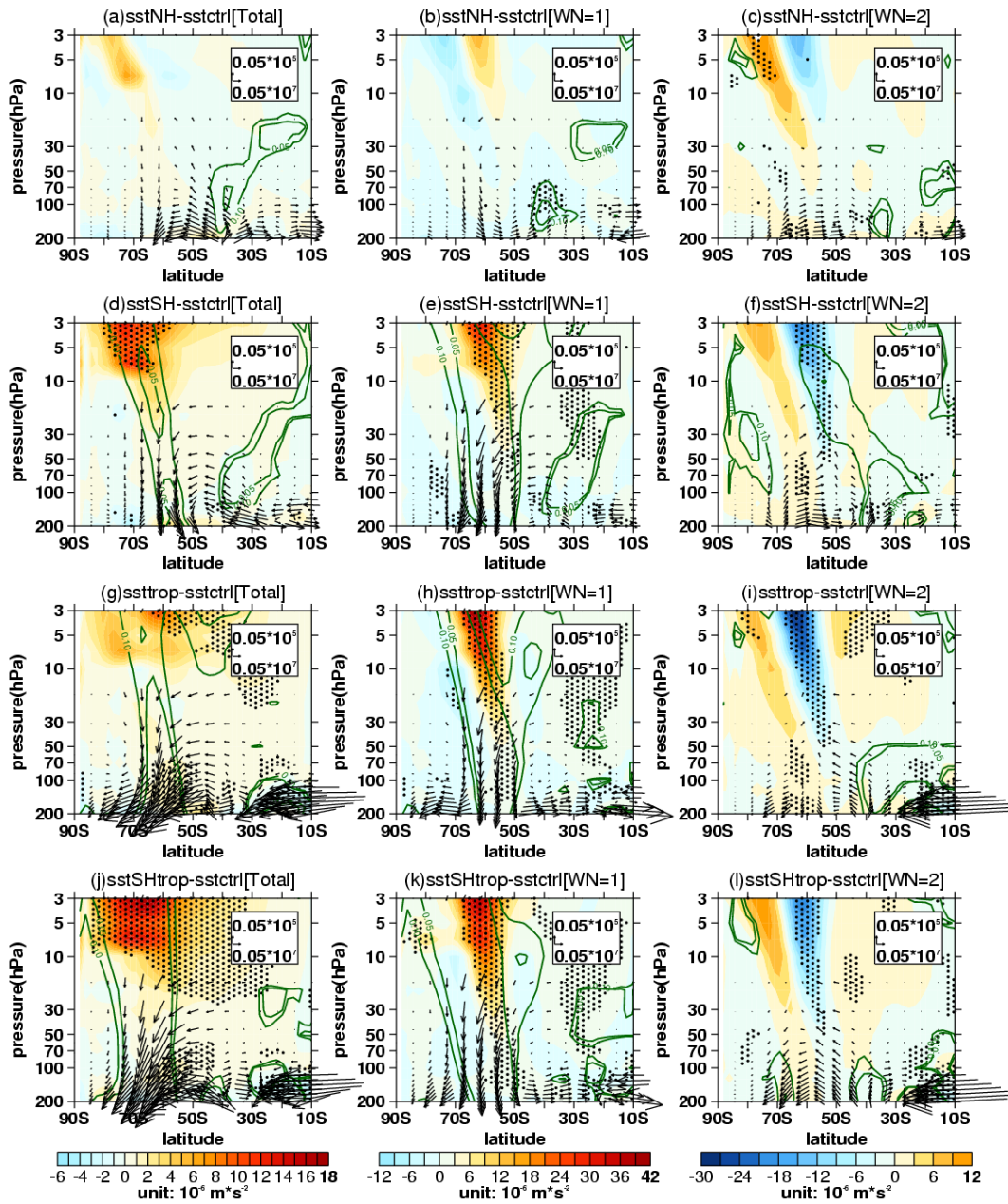
1054 sstSH; (c) ssttrop; (d) sstSHtrop) and the control experiment (sstctrl).



1055

1056 **FIG. 12.** Differences (shadings) of (a, d, g, j) 500 hPa geopotential height zonal  
 1057 deviations with their (b, e, h, k) wave-1 component and (c, f, i, l) wave-2 component  
 1058 between sensitive experiments ((a, b, c) sstNH; (d, e, f) sstSH; (g, h, i) ssttrop; (j, k, l)  
 1059 sstSHtrop) and the control experiment (sstctrl). The mean distributions (contours with  
 1060 an interval of 20 gpm, positive and negative values are depicted by solid and dashed  
 1061 lines respectively, zeroes are depicted by thick solid lines) of them are derived from the

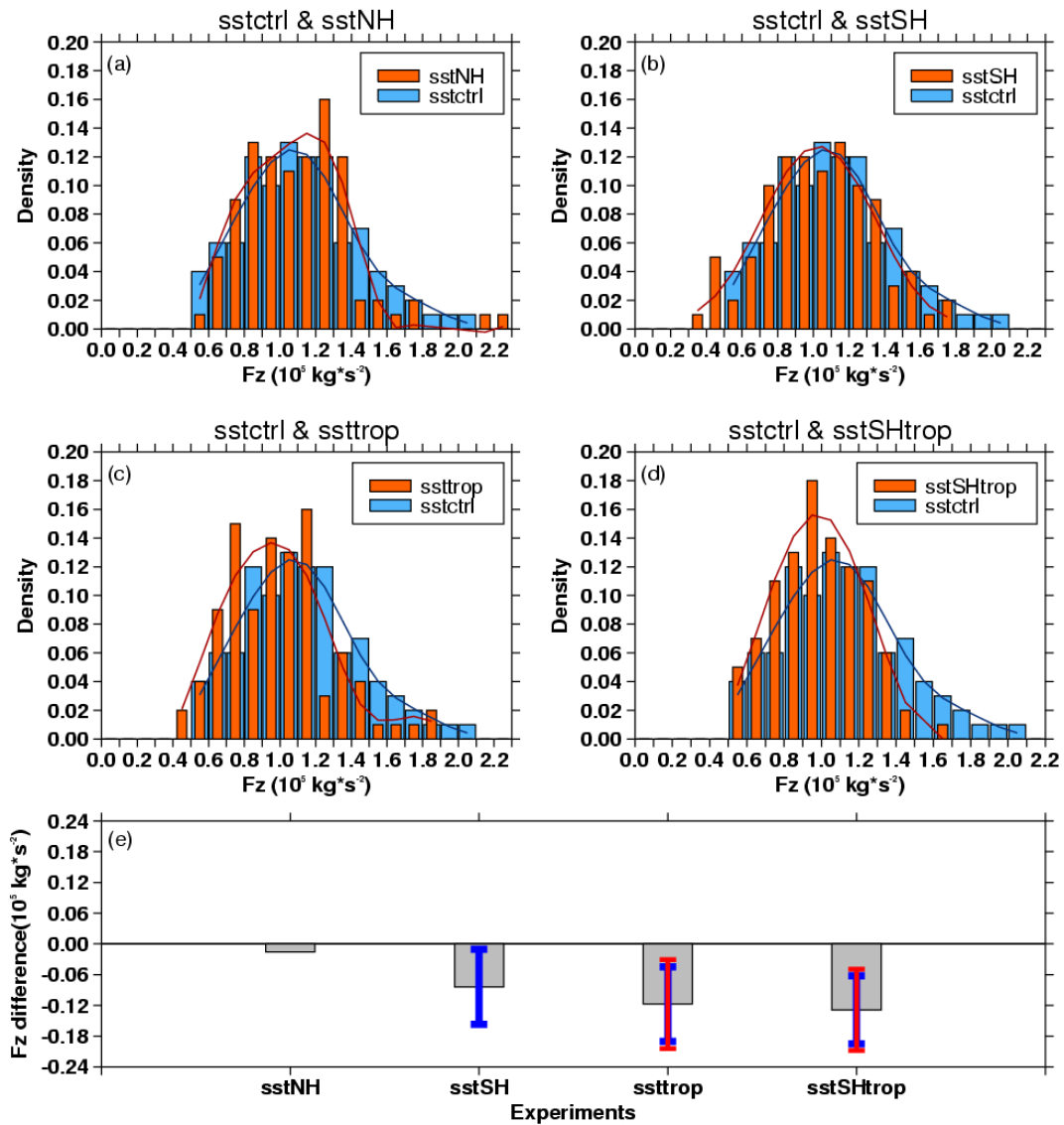
1062 control experiment. The stippled regions represent the mean difference significant  
 1063 at/above the 90% confidence level.



1064

1065 **FIG. 13.** Differences of (a, d, g, j) stratospheric E-P flux (arrows, units in horizontal  
 1066 and vertical components are  $0.05 \times 10^7$  and  $0.05 \times 10^5 \text{ kg} \cdot \text{s}^{-2}$ , respectively) and its  
 1067 divergence (shadings) with their (b, e, h, k) wave-1 component and (c, f, i, l) wave-2  
 1068 component between sensitive experiments ((a, b, c) sstNH; (d, e, f) sstSH; (g, h, i)  
 1069 ssttrop; (j, k, l) sstSHtrop) and the control experiment (sstctrl). The stippled regions

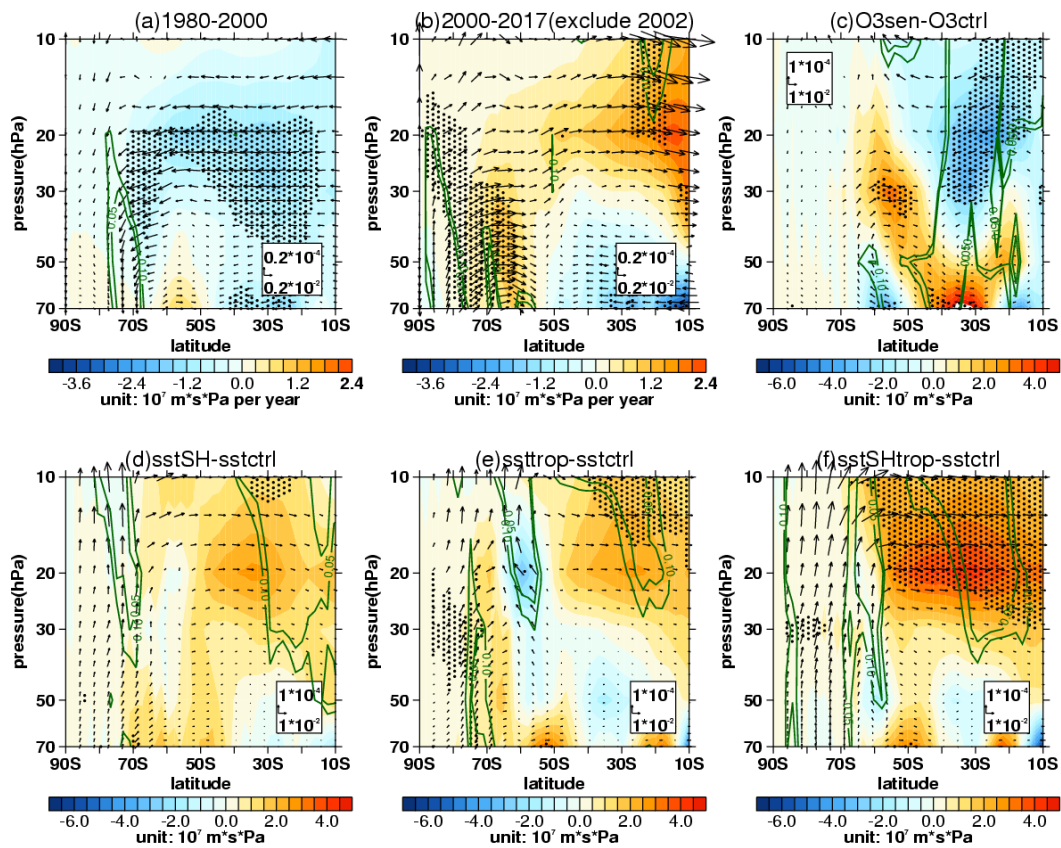
1070 represent the mean differences of E-P flux divergence significant at/above the 90%  
 1071 confidence level. The green contours from outside to inside (corresponding to  $p=0.1$ ,  
 1072 0.05) represent the mean differences of vertical E-P flux significant at the 90% and 95%  
 1073 confidence levels, respectively.



1074  
 1075 **FIG. 14.** (a, b, c, d) Frequency distributions (pillars, blue for control experiment and  
 1076 orange for sensitive experiments) of vertical E-P flux ( $F_z$ , area-weighted from 200 hPa  
 1077 to 10 hPa over  $70^\circ\text{S}$ - $50^\circ\text{S}$ ) and its 5-point low-pass filtered fitting curves (solid lines,  
 1078 blue for control experiment and red for sensitive experiments) derived from 100

1079 ensemble members of the control experiment (sstctrl) and sensitive experiments ((a)  
 1080 sstNH; (b) sstSH; (c) ssttrop; (d) sstSHtrop), respectively. (e) Mean differences (grey  
 1081 pillars) and corresponding uncertainties (error bars) of Fz between sensitive  
 1082 experiments and the control experiment. The blue and red error bars reflect the 90%  
 1083 and 95% confidence levels calculated by two-tailed t test, respectively. The error bar is  
 1084 omitted when the significance of mean difference is lower than the corresponding  
 1085 confidence level.

1086



1087

1088 **FIG. 15.** (a) Trends of southern hemispheric Brewer-Dobson circulation (arrows, units  
 1089 in horizontal and vertical components are  $0.2 \times 10^{-2}$  and  $0.2 \times 10^{-4} \text{ m} \cdot \text{s}^{-1}$  per year,  
 1090 respectively) and its stream function (shadings) in September during (a) 1980-2000 and  
 1091 (b) 2000-2017 derived from MERRA-2 dataset. Data in 2002 are removed when trends

1092 are calculated in Figure (b). (c) Differences of Brewer-Dobson circulation (arrows,  
1093 units in horizontal and vertical components are  $10^{-2}$  and  $10^{-4} \text{ m}\cdot\text{s}^{-1}$ , respectively) and its  
1094 stream function (shadings) between the O3ctrl and O3sen. (d, e, f) Differences of  
1095 Brewer-Dobson circulation and its stream function between the control experiment  
1096 (sstctrl) and sensitive experiments ((d) sstSH; (e) ssttrop; (f) sstSHtrop) with SST  
1097 changes. The stippled regions represent the trends or differences of the stream function  
1098 significant at/above the 90% confidence level. The green contours from outside to  
1099 inside (corresponding to  $p=0.1, 0.05$ ) represent the trends or differences of the vertical  
1100 components significant at the 90% and 95% confidence levels, respectively.