

**Weakening of Antarctic Stratospheric Planetary Wave
Activities in Early Austral Spring Since the Early 2000s: A
Response to Sea Surface Temperature Trends**

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1 **Abstract**

2 Using multiple reanalysis datasets and modeling simulations, the trends of
3 Antarctic stratospheric planetary wave activities in early austral spring since the early
4 2000s are investigated in this study. We find that the stratospheric planetary wave
5 activities in September have weakened significantly since 2000, which is mainly related
6 to the weakening of the tropospheric wave sources in the extratropical southern
7 hemisphere. As the Antarctic ozone also shows clear shift around 2000, the impact of
8 ozone recovery on Antarctic planetary wave activity is also examined through
9 numerical simulations. Significant ozone recovery in lower stratosphere changes the
10 atmospheric state for wave propagation to some extent, inducing a slight decrease of
11 vertical wave flux in upper troposphere and lower stratosphere (UTLS). However, the
12 changes of wave propagation environment in middle and upper stratosphere over
13 subpolar region are not significant. The ozone recovery has minor contribution to the
14 significant weakening of stratospheric planetary wave activity in September. Further
15 analysis indicates that the trend of September sea surface temperature (SST) over 20°
16 N-70°S is well linked to the weakening of stratospheric planetary wave activities. The
17 model simulations reveal that the SST trend in the extratropical southern hemisphere
18 (20°S-70°S) and the tropics (20°N-20°S) induce a weakening of wave-1 component of
19 tropospheric geopotential height in the extratropical southern hemisphere, which
20 subsequently leads to a decrease in stratospheric wave flux. In addition, both reanalysis
21 data and numerical simulations indicate that the Brewer-Dobson circulation (BDC)
22 related to wave activities in the stratosphere has also been weakening in early austral

23 spring since 2000 due to the trend of September SST in the tropics and extratropical
24 southern hemisphere.

25

26 **Key words:** *Antarctic; Stratospheric planetary wave activities; Tropospheric wave*
27 *sources; Sea surface temperature*

28

29 **1. Introduction**

30 The stratospheric planetary wave activities have important influences on
31 stratospheric temperature (e.g., Hu & Fu, 2009; Lin et al., 2009; Li & Tian, 2017; Li et
32 al., 2018), polar vortex (e.g., Kim et al., 2014; Zhang et al., 2016; Hu et al., 2018) and
33 distribution of chemical substances (e.g., Gabriel et al., 2011; Ialongo et al., 2012;
34 Kravchenko et al., 2011; Zhang et al., 2019a). Meanwhile, the stratospheric circulation
35 modulated by planetary waves can exert impacts on tropospheric weather and climate
36 (e.g., Haigh et al., 2005; Zhang et al., 2019b) through downward control processes
37 (Haynes et al., 1991), which is useful for extended forecast by using preceding signals
38 in the stratosphere (e.g., Baldwin et al., 2001; Wang et al., 2020).

39 The planetary perturbations generated by large-scale topography, convection and
40 continent-ocean heating contrast can propagate from the troposphere to the stratosphere
41 (Charney & Drazin, 1961) and form stratospheric planetary waves. As the land-sea
42 thermal contrast in the northern hemisphere is larger than that in the southern
43 hemisphere and produces stronger zonal forcing for the genesis of stratospheric waves,
44 the majority of attention has been given to wave activities and their impacts on weather

45 and climate in the northern hemisphere (e.g., Kim et al., 2014; Zhang et al., 2016; Hu
46 et al., 2018). However, planetary wave activities in the southern hemisphere also play
47 an important role in heating the stratosphere dynamically (e.g., Hu & Fu, 2009; Lin et
48 al., 2009), which suppresses Polar Stratospheric Clouds (PSCs) formation and ozone
49 depletion (e.g., Shen et al., 2020a; Tian et al., 2017). The Antarctic sudden stratospheric
50 warming (SSW) that occurred in 2002 (e.g., Baldwin et al., 2003; Nishii & Nakamura,
51 2004; Newman & Nash, 2005) and 2019 (e.g., Yamazaki et al., 2020; Shen et al., 2020a;
52 Shen et al., 2020b) was associated with significant upward propagation of wave flux.
53 Such episodes are extraordinarily rare in the history, and the one in 2019 contributed to
54 the formation of the smallest Antarctic ozone hole on record (WMO, 2019). In addition,
55 some studies reported that wildfires in Australia at the end of 2019 are related to
56 negative phase of the Southern Annular Mode (SAM), which was induced by the
57 extended influence of the SSW event that occurred in September (Lim et al., 2019; Shen
58 et al., 2020b). In a word, the Antarctic planetary wave activities are important for the
59 stratosphere-troposphere interactions and climate system in the southern hemisphere.

60 Long-term observations in the Antarctic stratosphere show a significant ozone
61 decline from the early 1980s to the early 2000s due to anthropogenic emission of
62 chlorofluorocarbons (CFCs) (WMO, 2011) and a recovery signal since 2000s because
63 of phasing out CFCs in response to Montreal Protocol (e.g., Angell and Free, 2009;
64 Krzyścin, 2012; Zhang et al., 2014; Banerjee et al., 2020). The Antarctic stratospheric
65 ozone depletion and recovery have important impacts on climate in the southern
66 hemisphere. The ozone depletion cools the Antarctic stratosphere through reducing

67 absorption of radiation and leads to the strengthening of Antarctic polar vortex during
68 austral spring (e.g., Randel & Wu, 1999; Solomon, 1999; Thompson et al., 2011). The
69 anomalous circulation in the Antarctic stratosphere during austral spring exerts impacts
70 on tropospheric circulations (e.g., intensification of SAM index, poleward shift of
71 tropospheric jet position and expansion of the Hadley cell edge) in the subsequent
72 months (e.g., Thompson et al., 2011; Swart & Fyfe, 2012; Son et al., 2018; Banerjee et
73 al., 2020) and influences the distribution of precipitation and dry zone in the southern
74 hemisphere (e.g., Thompson et al., 2011; Barnes et al., 2013; Kang et al., 2011).
75 Following the healing of ozone loss in the Antarctic ozone hole since 2000s (e.g.,
76 Solomon et al., 2016; Susan et al., 2019), great attention has been paid on possible
77 impacts of ozone recovery on climate system in the southern hemisphere (e.g., Son et
78 al., 2008; Barnes et al., 2013; Xia et al., 2020; Banerjee et al., 2020). Son et al. (2008)
79 implemented the Chemistry-Climate Model Validation (CCMVal) models to predict the
80 response of the southern hemisphere westerly jet to stratospheric ozone recovery. Based
81 on the Phase 5 of Coupled Model Intercomparison Projects (CMIP5) models, Barnes et
82 al. (2013) proposed that the tropospheric jet and dry zone edge no longer shift poleward
83 during austral summer since the early 2000s due to ozone recovery. Banerjee et al.
84 (2020) analyzed observations and reanalysis datasets. They found that following the
85 ozone recovery after 2000, the increase of SAM index and the poleward shifting of
86 tropospheric jet position as well as the Hadley cell edge all experienced a pause. Their
87 results suggest that ozone depletion and recovery have made important contributions to
88 the climate shift that occurred around 2000 in the southern hemisphere.

89 However, some previous studies have reported zonally asymmetric warming
90 patterns in Antarctic stratosphere, which are generated by increased planetary wave
91 activities during austral spring from the early 1980s to the early 2000s (Hu & Fu, 2009;
92 Lin et al., 2009). Note that the Antarctic stratosphere was experiencing radiative cooling
93 in the same period due to ozone depletion (e.g., Randel & Wu, 1999; Solomon, 1999;
94 Thompson et al., 2011). The increase in stratospheric planetary wave activities cannot
95 be explained by ozone decline, because the acceleration of stratospheric circumpolar
96 wind caused by radiative cooling induces more wave energy to be reflected back to the
97 troposphere (e.g., Andrews et al., 1987; Holton
98 et al., 2004). Hu & Fu (2009) attributed the increase in Antarctic stratospheric wave
99 activities to the SST trend from the 1980s to the 2000s. Their results indicate that in
100 addition to ozone change, other factors such as changes in SST also contribute to
101 climate change in the southern hemisphere. Moreover, the phase of Interdecadal Pacific
102 Oscillation (IPO) also changed at around 2000 (e.g., Trenberth et al., 2013). SST
103 variation influences Rossby wave propagation and tropospheric wave sources, and
104 thereby indirectly affects stratospheric wave activities (e.g., Lin et al., 2012; Hu et al.,
105 2018; Tian et al., 2017). The questions here are: (1) Has the trend of stratospheric
106 planetary wave activity in the southern hemisphere been shifting since the 2000s? (2)
107 What are the factors responsible for the trend of Antarctic stratospheric planetary wave
108 activity since the 2000s?

109 In this study, we reveal the trend of Antarctic planetary wave activity in early
110 austral spring since the 2000s based on multiple reanalysis datasets. We also conduct

111 sensitive experiments forced by linear increments of ozone and SST fields since the
112 2000s to investigate the response of Antarctic planetary activity to above two factors.
113 The remainder of the paper is organized as follows. Section 2 describes the data,
114 methods and configurations of model simulations. Section 3 presents the trends of
115 stratospheric and tropospheric wave activities in early austral spring. Section 4
116 examines the impact of ozone recovery on Antarctic stratospheric planetary wave
117 activity. Section 5 investigates the connections between the trends of SST and
118 stratospheric wave activities. Sections 6 discusses the responses of tropospheric wave
119 sources and stratospheric wave activities to SST changes based on model simulations.
120 Major conclusions and discussion are presented in Section 7.

121 **2. Datasets, methods and experimental configurations**

122 a. Datasets

123 In this study, daily and monthly mean data extracted from the Modern-Era
124 Retrospective analysis for Research and Applications Version 2 (MERRA-2;
125 Bosilovich et al., 2015) dataset are used to calculate trends of Brewer-Dobson
126 circulation (BDC), tropospheric wave sources, and the Elisassen-Palm (E-P) flux and
127 its divergence in September. To verify the trend of stratospheric E-P flux, we also refer
128 to the results derived from the European Centre for Medium-range Weather Forecasting
129 (ECMWF) Interim Reanalysis (ERA-Interim; Dee et al., 2011) dataset, the Japanese
130 55-year Reanalysis (JRA-55; Kobayashi et al., 2015) dataset and the National Centers
131 for Environmental Prediction-Department of Energy Global Reanalysis 2 (NCEP-2;
132 Kanamitsu et al., 2002) dataset.

133 The observed total column ozone (TCO) data are extracted from SBUV v8.6
 134 satellite dataset, which is a monthly and zonal mean dataset on 5° grid. Ozone data
 135 derived from MERRA-2 dataset are also used to calculate TCO.

136 SST data are extracted from the Extended Reconstructed Sea Surface Temperature
 137 (ERSST) dataset, which is a global monthly mean sea surface temperature dataset
 138 derived from the International Comprehensive Ocean-Atmosphere Dataset (ICOADS).
 139 The ERSST is on global 2°×2° grid and covers the period from January 1854 to the
 140 present. We use the latest version (version 5, i.e., v5) dataset to calculate trends and
 141 correlations, and produce SST forcing field for model simulations. More details about
 142 this version of ERSST can be found in Huang et al. (2017).

143 In addition, the unfiltered Interdecadal Pacific Oscillation (IPO) index derived
 144 from the ERSST v5 dataset is also used in this study. More detailed information about
 145 the index can be found in Henley et al. (2015).

146 b. Diagnosis of wave activities and Brewer-Dobson circulation

147 Planetary wave activities are measured by E-P flux ($\vec{F} \equiv (0, F^{(\phi)}, F^{(z)})$) and its
 148 divergence D_F . Their algorithms are expressed by Eqs. (1)-(3) (Andrews et al., 1987):

$$149 \quad F^{(\phi)} = \rho_0 a \cos \phi \overline{(u_z v' \theta' / \theta_z - v' u')} \quad (1)$$

$$150 \quad F^{(z)} = \rho_0 a \cos \phi \{ [f - (a \cos \phi)^{-1} \overline{(u \cos \phi)_\phi}] \overline{v' \theta' / \theta_z} - \overline{w' u'} \} \quad (2)$$

$$151 \quad D_F = \frac{\nabla \cdot \vec{F}}{\rho_0 a \cos \phi} = \frac{(a \cos \phi)^{-1} \frac{\partial}{\partial \phi} (F^{(\phi)} \cos \phi) + \frac{\partial F^{(z)}}{\partial z}}{\rho_0 a \cos \phi} \quad (3)$$

152 where \mathbf{u} , \mathbf{v} represent zonal and meridional components of horizontal wind, w
 153 is vertical velocity, θ is potential temperature, a is the Earth radius, f is the

154 Coriolis parameter, z is geopotential height, ϕ is latitude, ρ_0 is the background air
 155 density.

156 The quasi-geostrophic refractive index (RI) is used to diagnose the environment
 157 of wave propagation (Chen & Robinson, 1992). Its algorithm is written as Eq. (4):

$$158 \quad RI = \frac{\bar{q}_\phi}{\bar{u}} - \left(\frac{k}{a \cos \phi}\right)^2 - \left(\frac{f}{2NH}\right)^2 \quad (4)$$

159 where the zonal-mean potential vorticity meridional gradient \bar{q}_ϕ is

$$160 \quad \bar{q}_\phi = \frac{2\Omega}{a} \cos \phi - \frac{1}{a^2} \left[\frac{(\bar{u} \cos \phi)_\phi}{a \cos \phi} \right]_\phi - \frac{f^2}{\rho_0} \left(\rho_0 \frac{\bar{u}_z}{N^2} \right)_z \quad (5)$$

161 H , q , k , N^2 and Ω are the scale height, potential vorticity, zonal wave number,
 162 buoyancy frequency, and Earth's angular frequency, respectively.

163 The Brewer-Dobson circulation driven by wave breaking in the stratosphere is
 164 closely related to stratospheric wave activities. Its meridional and vertical components
 165 (\bar{v}^*, \bar{w}^*) and stream function $(\psi^*(p, \phi))$ are expressed by Eqs. (4)-(6) (Andrews et al.,
 166 1987; Birner & Bönisch, 2011) :

$$167 \quad \bar{v}^* \equiv \bar{v} - \rho_0^{-1} (\rho_0 \bar{v}' \theta' / \bar{\theta}_z)_z \quad (6)$$

$$168 \quad \bar{w}^* \equiv \bar{w} + (a \cos \phi)^{-1} (\cos \phi \cdot \bar{v}' \theta' / \bar{\theta}_z)_\phi \quad (7)$$

$$169 \quad \psi^*(p, \phi) = \int_0^p \frac{-2\pi a \cdot \cos \phi \cdot \bar{v}^*(p'', \phi)}{g} dp'' \quad (8)$$

170 where p is the air pressure, π is the circular constant, g is the gravitational
 171 acceleration.

172 In Eqs. (1)-(8), the overbar and prime denote a zonal mean and departure from the
 173 zonal mean, respectively. The subscripts denote partial derivatives. The Fourier
 174 decomposition is used to obtain components of Eqs. (1)-(3) with different zonal wave

175 numbers. Meanwhile, the Fourier decomposed components of geopotential height zonal
176 deviations are also used to determine tropospheric wave sources.

177 c. Statistical methods

178 The trend is measured by the slope of linear regression based on the least square
179 estimation. The correlation is used to analyze statistical links between different
180 variables. In this paper, all the time series have been linearly detrended before
181 calculating correlation coefficients (r) and their corresponding significances.

182 The change-point testing (e.g. Banerjee et al., 2020) is used to make sure the
183 significance of trend or correlation coefficient is not unduly influenced by some
184 particular beginning or ending years, and thereby confirm that the trend exists
185 objectively.

186 We use two-tailed student's t test to calculate the significances of trend, correlation
187 coefficient or mean difference. The result of significance test is measured by p value or
188 confidence intervals in this paper. $p \leq 0.1$, $p \leq 0.05$ and $p \leq 0.01$ suggest the trend,
189 correlation coefficient or mean difference is significant at/above the 90%, 95% and 99%
190 confidence levels, respectively. The confidence interval of trend is shown in (7) (Shirley
191 et al., 2004):

$$192 \quad [\hat{b} - t_{1-\alpha/2}(n-2)\hat{\sigma}_b, \hat{b} + t_{1-\alpha/2}(n-2)\hat{\sigma}_b] \quad (7)$$

193 where \hat{b} is estimated value of slope, $\hat{\sigma}_b$ is standard error of slope and it is written as:

$$194 \quad \hat{\sigma}_b = \hat{b} \cdot \sqrt{\frac{1}{n-2} - \frac{r^2}{n-2}}, \quad t_{1-\alpha/2}(n-2) \text{ denotes the value of t-distribution with the degree of}$$

195 freedom equal to $n-2$ and the two-tailed confidence level equal to $1-\alpha$ ($\alpha=0.90$,

196 0.95 or 0.99). The confidence interval of mean difference is expressed by Eq. (8)
 197 (Shirley et al., 2004):

$$198 \left[\bar{X} - \bar{Y} - t_{1-\alpha/2}(M+N-2) \cdot S_w \cdot \sqrt{\frac{1}{M} + \frac{1}{N}}, \bar{X} - \bar{Y} + t_{1-\alpha/2}(M+N-2) \cdot S_w \cdot \sqrt{\frac{1}{M} + \frac{1}{N}} \right] \quad (8)$$

199 where

$$200 S_w = \sqrt{\frac{1}{M+N-2} \left[\sum_{i=1}^M (X_i - \bar{X})^2 + \sum_{j=1}^N (Y_j - \bar{Y})^2 \right]} \quad (9)$$

201 Here, \bar{X} and \bar{Y} are the sample averages, M and N are the numbers of sample
 202 sizes with two populations, $t_{1-\alpha/2}(M+N-2)$ denotes the value of t-distribution with
 203 the degree of freedom equal to $M+N-2$ and the two-tailed confidence level equal to
 204 $1-\alpha$.

205 Previous studies have indicated that SST impact on the stratosphere shows a
 206 spatial dependence (e.g., Xie et al., 2020). To find out a robust relationship between the
 207 trend of SST in a specific region and the trend of stratospheric wave activities, we divide
 208 the global ocean into three regions: SH (the extratropical southern hemisphere, 70°S-
 209 20°S), TROP (the tropics, 20°S-20°N) and NH (the extratropical northern hemisphere,
 210 20°N-70°N). Since the impacts in different regions might be combined, we also
 211 consider three combined regions named as SHtrop (the extratropical southern
 212 hemisphere and the tropics, 70°S-20°N), NHtrop (the extratropical northern hemisphere
 213 and the tropics, 20°S-70°N) and the Globe (70°S-70°N). To find statistical connections
 214 between the trend of SST and that of stratospheric wave activities, we examine the first
 215 three leading patterns (EOF1, EOF2, EOF3) and principal components (PC1, PC2, PC3)
 216 of SST in above six regions obtained from Empirical Orthogonal Function (EOF)

217 analysis. In all the six regions, there is always one EOF mode that shows great similarity
218 to the spatial pattern of trend (not shown) as we do not detrend SST time series when
219 the EOF analysis is carried out. Thus, the significance of the correlation between the
220 PC time series of that EOF mode and time series of stratospheric E-P flux can be used
221 as the criterion to determine the statistical connection between the trend of SST and the
222 trend of stratospheric wave activities.

223 d. The model and experiment configurations

224 The F_2000_WACCM_SC (FWSC) component in the Community Earth System
225 Model (CESM; version 1.2.0) is used to verify the impacts of SST and ozone recovery
226 on tropospheric wave sources and stratospheric wave activities in early austral spring.
227 The FWSC component is the Whole Atmosphere Community Climate Model version 4
228 (WACCM4) with specified chemistry forcing fields (such as ozone, greenhouse gases
229 (GHG), aerosols and so on), which have fixed values in 2000 by default. The WACCM4
230 includes active atmosphere, data ocean (run as a prescribed component, simply reading
231 SST forcing data instead of running ocean model), land and sea ice. Physics schemes
232 in the WACCM4 are based on those in the Community Atmospheric Model version 4
233 (CAM4; Neale et al., 2013). The WACCM4 uses a finite-volume dynamic framework
234 and extends from the ground to approximately 145 km (5.1×10^{-6} hPa) altitude in the
235 vertical with 66 vertical levels. The simulations presented in this paper are conducted
236 at a horizontal resolution of $1.9^\circ \times 2.5^\circ$. More information about the WACCM can be
237 found in Marsh et al. (2013).

238 Control experiments and sensitive experiments are conducted to investigate

239 responses of Antarctic stratospheric wave activities to SST trends and the ozone
240 recovery trend in early austral spring. For the experiments of SST trends, monthly mean
241 global SST during 1980-2000 derived from the ERSST v5 dataset is used as SST
242 forcing field in the control experiment (sstctrl). For the four sensitive experiments
243 (sstNH, sstSH, ssttrop, sstSHtrop), linear increments of SST in different regions in
244 September during 2000-2017 are used as the forcing field. Ozone, aerosols and
245 greenhouse gases (GHG) in the control experiment and the four sensitive experiments
246 all have the fixed values in 2000. For the experiments of ozone recovery trend, monthly
247 mean three-dimensional global ozone during 1980-2000 derived from the MERRA-2
248 dataset is used as the ozone forcing field in the control experiment (O3ctrl). The
249 sensitive experiment (O3sen) is forced by linear increments of ozone in September
250 during 2001-2017. The SSTs in O3ctrl and O3sen both are monthly mean global SST
251 during 1980-2000. The aerosol and greenhouse gases values in 2000 are used. These
252 experiment configurations are summarized and listed in Table 1 and Table 2.

253 Firstly, we run the FWSC component to generate randomly different initial
254 conditions for 120 years with free run. Then, each experiment includes 100 ensemble
255 members that run from July to September forced by these initial conditions from the
256 21st year to the 120th year in July. The forcing fields of SST and ozone are only
257 superposed from July to September. July and August are taken as spin-up time and
258 simulations during this period are discarded. The ensemble mean in September derived
259 from these 100 ensemble members is regarded as the final result of each experiment. A
260 similar approach is implemented for sensitive experiments, in which the forcing fields

261 superposed only in certain months. The same approach has been used in previous
262 studies (e.g., Zhang et al., 2018).

263 **3. Trend of planetary wave activities in early austral spring**

264 Figure 1 shows the trends of stratospheric planetary wave activities in the southern
265 hemisphere September during 1980-2000 and 2000-2017, respectively. Note that the
266 vertical E-P flux entering into the stratosphere over 50°S-70°S in September has been
267 increasing during 1980-2000, accompanied by intensified wave flux convergence in the
268 upper stratosphere (Fig. 1a) that is mainly contributed by the wave-1 component (Fig.
269 1b). This feature implies that the stratospheric planetary wave activities have
270 strengthened in early austral spring during 1980-2000. A similar result has been
271 reported in previous studies (Hu & Fu, 2009; Lin et al., 2009). During 2000-2017,
272 however, vertical propagation of stratospheric E-P flux weakened over the subpolar
273 region of the southern hemisphere, which was accompanied by intensified wave flux
274 divergence in the upper stratosphere (Fig. 1d) mainly contributed by the wave-1
275 component (Fig. 1e) while the wave-2 component also made certain contributions (Fig.
276 1f). Similar features also appear in August, but not as significant as that in September
277 (Fig. S1). For this reason, hereafter we focus on the features in September.

278 The SSW that occurred in 2002 was accompanied with large upward wave fluxes
279 in the stratosphere, which is extremely rare in history and has been studied extensively
280 in previous studies (e.g., Baldwin et al., 2003; Nishii & Nakamura, 2004; Newman &
281 Nash, 2005). Since the period with a negative trend of stratospheric vertical wave flux
282 is short, it is necessary to further investigate whether such a negative trend is artificially

283 influenced by the single year of 2002. Therefore, following Banerjee et al. (2020), we
284 use a change-point method to test the significance of the trend during various periods
285 based on four reanalysis datasets (ERA-Interim, MERRA-2, JRA-55, NCEP-2).
286 Figures 2a (including the year 2002) and 2b (excluding the year 2002) display the time
287 series of area-weighted vertical stratospheric wave flux (F_z) over the southern
288 hemisphere subpolar region obtained from different reanalysis datasets. Note that the
289 wave flux time series obtained from the four reanalysis datasets all present a positive
290 trend from the early 1980s to the early 2000s and a negative trend from the early 2000s
291 to present, regardless of whether the extreme value in 2002 is removed or not. The
292 correlation coefficients of the time series between these reanalysis datasets are above
293 0.9 and statistically significant (Table 3), suggesting that the time series derived from
294 different datasets are consistent with each other. Figures 2c-f show the trends and
295 corresponding confidence intervals calculated with four different beginning years (1980,
296 1981, 1982, 1983), four different ending years (2015, 2016, 2017, 2018), and change-
297 point years from 1998 to 2013. The trends and confidence intervals in Figures 2g-j are
298 the same as that in Figures 2c-f, except that the extreme value in 2002 is removed. The
299 positive trend from the early 1980s to the 21st century remains significant regardless of
300 different beginning years and ending change-point years (Figs. 2c-j). However, Figures
301 2c-f and Figures 2g-j indicate that the positive value of the trend is decreasing gradually
302 when the period is prolonged, which is apparently attributed to the negative trend with
303 the beginning change-point year of around 2000. Although the negative trend from the
304 change-point year to ending year becomes less significant when the value in 2002 is

305 removed, it remains significant in some periods, which are also illustrated on diagrams
306 of latitude-pressure profiles (Fig. S2). Therefore, the weakening of stratospheric wave
307 activities in early austral spring since the early 2000s is robust. In this paper, we take
308 the year 2000 as the beginning year of the weakening trend to simplify descriptions in
309 the following discussion.

310 Figure 3 shows the trends of tropospheric wave sources in September since 2000.
311 There is a significant positive trend of the wave-1 component in 500 hPa geopotential
312 height over the southern Indian ocean and a significant negative trend over the southern
313 Pacific, which form an out-of-phase superposition on its climatology (Fig. 3b). The
314 trend pattern of wave-2 component is also out-of-phase with its climatology, although
315 it is not significant (Fig. 3c). The above features still maintain when the values in 2002
316 are removed (Figs. S3b, c), implying that the southern hemispheric tropospheric wave
317 sources in early austral spring have weakened since 2000, which is also reflected in the
318 decrease of tropospheric vertical wave flux (Figs. 3d, e; Figs. S3d, e).

319 **4. Response of Antarctic stratospheric wave activity to ozone recovery**

320 Previous studies have suggested that ozone depletion and recovery are important
321 to climate shift that occurred around 2000 in the southern hemisphere during austral
322 summer (e.g., Son et al., 2008; Thompson et al., 2011; Barnes et al., 2013; Banerjee et
323 al., 2020). The impacts of stratospheric ozone changes on Antarctic wave propagation
324 during austral summer has also been examined in previous studies (e.g., Hu et al., 2015).
325 However, whether ozone recovery in September explains the weakening of
326 stratospheric planetary waves at the same month remains unclear. The correlation

327 between detrended time series of September Antarctic total column ozone (TCO)
328 derived from SBUV and stratospheric vertical wave flux (F_z) is 0.70 ($p=0.0011$) during
329 2000-2017. The increase of wave activity in polar stratosphere causes heating effects
330 and suppresses the formation of PSCs, and hence, slow down the ozone depletion (e.g.,
331 Shen et al. 2020a). Therefore, the Antarctic ozone and stratospheric wave activity show
332 statistically significant positive correlation. Theoretically, heating effects caused by
333 ozone recovery in Antarctic stratosphere may also decelerate the Antarctic stratospheric
334 polar vortex and induce more waves to propagate into stratosphere (Andrews et al.,
335 1987; Holton et al., 2004). These preliminary analysis cannot verify that the ozone
336 recovery is responsible for weakening of stratospheric wave activity. The role of ozone
337 recovery in stratospheric wave changes needs to be further explored by model
338 simulations. In this section, we use a group of time-slice experiments (O3ctrl and O3sen)
339 to address this issue.

340 Figure 4 shows the time series and piecewise trends of September TCO in the
341 Antarctic during 1980-2017. As reported by previous studies (e.g., Angell and Free,
342 2009; Banerjee et al., 2020; Krzyścin, 2012; Solomon et al., 2016; WMO, 2011; Zhang
343 et al., 2014), the Antarctic ozone show a significant decline during 1980-2000 (Figs. 4a,
344 b, c) and a slight recovery during 2001-2017 (Figs. 4a, d, e). The recovery trend is
345 calculated with data in 2002 removed because the large poleward transport induced by
346 SSW in 2002 leads to extreme values of ozone (e.g. Solomon et al., 2016). In addition,
347 the correlation of TCO between MERRA-2 and SBUV datasets is 0.61 ($p=4.5\times 10^{-5}$),
348 suggesting the changes of TCO derived from the reanalysis dataset and the observations

349 have a good consistency. Thus, in order to get three-dimensional structure of ozone
350 changes, the ozone data from MERRA-2 are used to make forcing fields for CESM. As
351 described in Section 2, a control experiment (O3ctrl) forced by climatological ozone
352 and a sensitive experiment forced by the linear increment of global ozone in September
353 during 2001-2017 are conducted to explore the impacts of ozone recovery. The pattern
354 of ozone forcing fields is similar to its trend patterns (Figs. 4d, e; Figs. 5a, b). Other
355 details of these two experiments have been given in Section 2 and Table 2.

356 Fig. 6 and Fig. 7 show the responses of wave activity and wave propagation
357 environment forced by O3sen. Note that the significant ozone recovery over south pole
358 mainly appears in lower stratosphere (about 200 hPa to 50 hPa) (Fig. 4e). In most
359 southern polar regions from 50 hPa to 3 hPa, the ozone recovery is not significant (Fig.
360 4e). The features are attributed to limitation of ODSs emission and reduction of
361 heterogeneous reaction on PSCs, which mainly distribute in lower stratosphere (e.g.,
362 Solomon, 1999). Ozone recovery in polar lower stratosphere absorbs more ultraviolet
363 radiation and causes cooling in Antarctic troposphere (Fig. 6b). To maintain thermal
364 balance, zonal wind accelerates below 200 hPa over 60°S-70°S (Fig. 6a).

365 The changes of zonal wind and temperature forced by ozone recovery induce
366 changes in wave propagation environment. The refractive index (RI) is a good metric
367 to reflect the atmosphere state for wave propagation. Theoretically, planetary waves
368 tend to propagate into large RI regions (Andrews et al., 1987). The responses of RI and
369 its terms are shown in Figs. 6c-f. Note that the second term of RI does not change with
370 atmospheric state and the third term of RI is insignificant compared to the first term

371 (Hu et al., 2019). Previous studies indicate that changes in zonal mean potential
 372 vorticity meridional gradient \bar{q}_φ could explain the changes in RI in middle and high
 373 latitudes (e.g. Hu et al., 2019; Simpson et al., 2009). Consistent with these studies, the
 374 pattern of \bar{q}_φ show some similarity with pattern of RI (Figs. 6c, d), especially in lower
 375 stratosphere over subpolar regions (Figs. 6c, d). According to the Eq. (5), the first term
 376 of \bar{q}_φ does not change with atmospheric state. Therefore, the second term
 377 ($-\left[\frac{(\bar{u} \cos \varphi)}{\cos \varphi}\right]_\varphi$, hereafter uyy term or barotropic term) and the third term
 378 ($-\frac{f^2}{\rho_0}(\rho_0 \frac{\bar{u}_z}{N^2})_z$, hereafter uzz term or baroclinic term) are investigated. Note that the
 379 pattern of responses in baroclinic term is similar with \bar{q}_φ (Figs. 6d, f). The uzz term
 380 also can be written as $(\frac{f^2}{HN^2} + \frac{f^2}{N^4} \frac{dN^2}{dz})\bar{u}_z - \frac{f^2}{N^2}\bar{u}_{zz}$. Meanwhile, zonal wind
 381 acceleration in upper troposphere weakens the vertical shear of u (\bar{u}_z) around 200 hPa
 382 over subpolar regions, inducing a decrease of baroclinic term and RI in upper
 383 troposphere and lower stratosphere (UTLS) over 60°S-70°S (Figs. 6d, f). The response
 384 of RI induces a slight decrease of vertical wave flux in UTLS over subpolar regions
 385 (Fig. 7a), which is mainly contributed by its wave-1 component (Fig. 7b). However, the
 386 changes of wave activity in UTLS are not significant in ensemble mean of simulations
 387 (Figs. 7a, b, c). Meanwhile, note that the responses of zonal wind and temperature to
 388 ozone recovery are not significant above 50 hPa over subpolar regions (Figs. 6a, b),
 389 inducing negligible changes of wave propagation environment (Fig. 6c) and wave
 390 activity (Fig. 7) in middle and upper stratosphere.

391 In a word, the significant ozone recovery in Antarctic lower stratosphere changes

392 wave propagation in upper troposphere and lower stratosphere to some extent. However,
393 these weak responses still cannot explain the significant decrease of stratospheric wave
394 flux in September.

395 **5. Role of SST trends in the weakening of Antarctic stratospheric** 396 **wave activities**

397 In this section, we further explore factors responsible for the weakening of
398 tropospheric wave sources and stratospheric wave activities since the early 2000s in
399 early austral spring. Many studies reported that SST variations can affect stratospheric
400 climate (e.g., Li, 2009; Hurwitz et al., 2011; Lin et al., 2012; Hu et al., 2014; Hu et al.,
401 2018; Tian et al., 2017; Xie et al., 2020). Hu & Fu (2009) also attributed the
402 strengthened stratospheric wave activities in the southern hemisphere to SST trend from
403 the early 1980s to the early 2000s. Furthermore, global SST in September during 2000-
404 2017 also has a significant trend. The significant warming pattern is mainly found over
405 the southern Indian ocean, the southern Atlantic ocean, the eastern and western
406 equatorial Pacific, the western equatorial and Northern Atlantic ocean (Fig. 8b). A
407 significant cooling pattern is located over the southeast Pacific (Fig. 8b). In addition,
408 the transitions around 2000 exist in SST time series over some regions. In the southern
409 Indian ocean, SST shows insignificant trend during 1980-2000 and significant warming
410 trend during 2000-2017 (Fig. 8c). The subtropical Pacific ocean in east of Australia is
411 linked with the Pacific-Southern America (PSA) wave train (e.g. Shen et al., 2020b),
412 and the SST there shows significant warming trend during 1980-2000 and insignificant
413 trend during 2000-2017. The SST in southeast Pacific shows insignificant trend during

414 1980-2000 and significant cooling during 2000-2017 (Fig. 8e). Trends of SST in
415 southern Atlantic ocean are opposite during these two piecewise periods, showing
416 significant cooling during 1980-2000 and significant warming during 2000-2017. It is
417 apparent that the spatial pattern of SST trend during 2000-2017 is obviously different
418 from that during 1980-2000 (Fig. 8a, b), which may affect the tropospheric wave
419 sources. Thus, it is necessary to analyze the connection between SST trend and wave
420 activity trend since the early 2000s.

421 Figure 9 shows the significance of the trend of principle component (PC) time
422 series of SST in different regions (Figs. 9a-f), and the significance of correlations (Figs.
423 9g-l) between the PC time series and Fz in September during various periods. The trend
424 of PC1 time series in SH region is significant during several periods (Fig. 9a), while the
425 correlation between PC1 and Fz is only significant with the particular ending year of
426 2015 (Fig. 9g). This feature suggests that the connection between the SST trend in SH
427 region and the trend of stratospheric wave activity is not robust. The correlation
428 between trend of stratospheric wave activity and that of SST in TROP or NH region is
429 also weak (Fig. 9e, f). As for the combined regions, note that the PC2 time series in
430 SHtrop region has a significant trend (Fig. 9d) and the correlation between the PC2 time
431 series in SHtrop and Fz with the beginning year of around 2000 is also significant (Fig.
432 9j) regardless of different ending years. This feature implies that the extratropical
433 southern hemisphere and tropical SST has a robust connection with stratospheric wave
434 activities in early austral spring since the early 2000s. The correlations between Fz and
435 all PC time series in NHtrop (Fig. 9k) and Globe (Fig. 9l) region are not as robust as

436 that between Fz and PC2 time series in SHtrop region (Fig. 9j), indicating that the
437 connection between SST trend in extratropical northern hemisphere and the trend of
438 stratospheric wave activity is weak.

439 Figure 10 shows the first three EOF modes of September SST in SHtrop region
440 during 2000-2017. The second mode (Fig. 10b) shows a great similarity to the spatial
441 pattern of SST trend (Fig. 8b), and the corresponding PC2 time series also has a
442 significant trend (slope=1.71, $p<0.01$). The correlation between PC2 and Fz is
443 significant ($r=-0.56$, $p=0.016$) and the correlation coefficient remains significant ($r=-$
444 0.46 , $p=0.065$) at the 90% confidence level when the value in 2002 is removed. This
445 result suggests that the SST trend in SHtrop region is closely related to the recent
446 weakening of stratospheric wave activities. The first EOF mode is similar to IPO (Fig.
447 10a) and its corresponding principal component is significantly correlated ($r=-0.98$,
448 $p<0.01$) with the unfiltered IPO index. However, it shows no significant trend (Fig. 10d)
449 and has no significant correlation (Fig. 10g) with stratospheric wave flux, implying that
450 the linkage between the IPO phase change at around 2000 (e.g. Trenberth et al., 2013)
451 and the weakening of Antarctic stratospheric wave activities is weak. The correlation
452 between PC3 and Fz is also not significant (Fig. 10i). Therefore, it is possible that the
453 combined effect of SST trend (the second EOF mode) in the tropical and extratropical
454 southern hemisphere leads to the weakening of stratospheric wave activities in early
455 austral spring since the early 2000s.

456 **6. Simulated changes in Antarctic stratospheric wave activities forced** 457 **by SST trends**

458 The analysis in Section 5 suggests that the SST changes in SHtrop region may
459 contribute to the weakening of the southern hemispheric stratospheric wave activities.
460 Here, numerical experiments sstNH, sstSH, ssttrop and sstSHtrop forced by linear
461 increments of SST in September during 2000-2017 (Fig. 11; more details can be found
462 in Section 2) are conducted to verify the results presented in Section 5.

463 Figure 12 shows the simulated response of 500 hPa geopotential height to SST
464 changes in different regions. The climatological distributions of the wave-1 component
465 (Figs. 12b, e, h, k) and the wave-2 component (Figs. 12c, f, i, l) from the simulations
466 are consistent with that from reanalysis dataset (Figs. 3b, c), indicating that the model
467 can well capture spatial distributions of the atmospheric waves. Note that the wave-1
468 and wave-2 anomalies simulated with SST changes in SH, TROP and SHtrop are all
469 significant. They superpose on the corresponding climatological patterns in an out-of-
470 phase style (Figs. 12e, f, h, i, k, l), indicating that the changes in SST in SH, TROP and
471 SHtrop lead to a weakening of tropospheric wave sources in the extratropical southern
472 hemisphere. However, the predominate wave-1 component of the 500 hPa geopotential
473 height anomaly in the extratropical southern hemisphere forced by the experiment with
474 NH SST change is relatively weak (Fig. 12b). This feature suggests that the SST
475 changes in extratropical northern hemisphere are incapable of inducing a robust
476 response of tropospheric wave sources in the extratropical southern hemisphere.

477 Figure 13 shows the simulated responses of stratospheric wave activities in the
478 southern hemisphere to SST changes over different regions. It is apparent that the
479 experiments with SST changes in SH, TROP and SHtrop show significantly weakened

480 stratospheric wave activities (Figs. 13d, g, j), which are mainly attributed to the
481 responses of the wave-1 component (Figs. 13e, h, k). These results are consistent with
482 the responses of tropospheric wave sources (Figs. 12d, e, g, h, j, k). However, there are
483 no significant anomalies of stratospheric wave flux in the subpolar region in Figures
484 13a and 13b, which is consistent with the response of corresponding tropospheric wave
485 sources (Figs. 12a, b) and the weak correlation between Fz and PC time series of SST
486 in NH region (Fig. 9i). The result here suggests that the response of southern
487 hemisphere stratospheric wave activities to SST trend in NH region is weak.

488 The results of stratospheric vertical wave flux over 50°S-70°S derived from the
489 100 ensemble members of each experiment are shown in Figure S4, and the frequency
490 distributions of them are displayed in Figure 14. The results of all these experiments
491 are summarized and displayed in Figure 14, which are quantified by the frequency
492 distribution of southern hemisphere stratospheric vertical wave flux derived from the
493 100 ensemble members of each experiment. Compared to the blue fitting curves, the
494 red fitting curves shift to the left as shown in Figs. 14b, 14c and 14d, suggesting that
495 the SST changes in SH, TROP and SHtrop regions weaken the upward propagation of
496 stratospheric wave flux. The area-weighted anomalies of vertical E-P flux in the
497 subpolar region of the southern hemisphere induced by SST changes in SH, TROP and
498 SHtrop regions are $-0.084 \times 10^5 \text{ kg} \cdot \text{s}^{-2}$, $-0.12 \times 10^5 \text{ kg} \cdot \text{s}^{-2}$ and $-0.13 \times 10^5 \text{ kg} \cdot \text{s}^{-2}$,
499 respectively. The sum of the anomalies forced by sstSH and ssttrop is not equal to the
500 anomaly forced by sstSHtrop, which may be resulted from non-linear interactions
501 between the responses of wave activities to SST trends in SH region and TROP region.

502 The weakening of stratospheric wave activities forced by SST increment in the tropical
503 region is more significant than that in extratropical southern hemisphere (Figs. 14b, c,
504 e), implying that the SST trend in the tropical region contributes more to the weakening
505 of stratospheric wave activities since 2000. Meanwhile, it is apparent that the
506 weakening of the southern hemisphere stratospheric wave activities forced by
507 sstSHtrop is the most significant among all the sensitive experiments (Fig. 14e). The
508 reduction of vertical E-P flux over (50°S-70°S, 200 hPa-10 hPa) forced by sstSHtrop is
509 approximately 12%. These modeling results indicate that the weakening of the
510 Antarctic stratospheric wave activities in September since 2000 is induced mainly by
511 the combined effects of SST trends in the tropical and extratropical southern
512 hemisphere. It also explains why the independent correlation between Fz and PC time
513 series obtained over SH or TROP region is not as significant as that between Fz and PC
514 time series obtained over SHtrop region (Figs. 9g, h, j). Moreover, the mean linear
515 increment of area-weighted vertical E-P flux from 200 hPa to 10 hPa over 70°S-50°S
516 in September during 2000-2017 derived from four reanalysis datasets is about -
517 $0.38 \times 10^5 \text{ kg} \cdot \text{s}^{-2}$. Therefore, the contribution of SST trend over 20°N-70°S (the SHtrop
518 region) to the weakening of stratospheric activities is approximately 34%.

519 In addition, the reanalysis datasets show that the Brewer-Dobson circulation
520 related to wave activities in the stratosphere weakened significantly in early austral
521 spring during 2000-2017 (Fig. 15b), which is contrary to the intensified trend during
522 1980-2000 (Fig. 15a). The transition of BDC around 2000 is believed to be associated
523 with ozone depletion and recovery (e.g., Polvani et al., 2017; Polvani et al., 2018).

524 However, our modeling results suggest that the SST trend is responsible for the
525 weakening of BDC in September since 2000 (Figs. 15d, e, f), The response of BDC to
526 ozone recovery is not significant (Fig. 15c) in September, especially for the branch near
527 the Antarctic. These results indicate that apart from the ozone depletion and recovery
528 the SST trend should also be taken into consideration when exploring the mechanism
529 for the climate transition in the southern hemispheric stratosphere around 2000.

530 Previous studies reported that there is usually a time lag for tropic SST to affect
531 extratropical circulation (e.g., Shaman & Tziperman, 2011). Thus, the impact of tropical
532 SST change before September needs to be further examined. Our simulations indicate
533 that the tropical SST trend in September plays a dominate role in weakening of
534 stratospheric wave activity at the same month, and the effect of tropical SST change
535 before September is negligible compared to that in September (The detailed evidences
536 to address this issue are shown in the appendix).

537 **7. Conclusions and Discussions**

538 This study analyzes the trend of Antarctic stratospheric planetary wave activities
539 in early austral spring since the early 2000s based on various reanalysis datasets and
540 model simulations. Using the change-point method, we find that the Antarctic
541 stratospheric wave activities in September have been weakening significantly since
542 2000, which means the intensified trend of wave activities noted in previous researches
543 (Hu & Fu, 2009; Lin et al., 2009) are reversed after 2000 in early austral spring. Further
544 analysis suggests that the weakening of stratospheric wave activities is related to the
545 weakening of tropospheric wave sources in extratropical southern hemisphere, which

546 is mainly contributed by the wave-1 component.

547 As the Antarctic ozone also shows clear shift around the 2000, we firstly examine
548 the impact of ozone recovery on Antarctic stratospheric planetary wave activity. Our
549 simulation results indicate that significant ozone recovery in lower stratosphere changes
550 the atmospheric state for wave propagation to some extent, inducing a slight decrease
551 of vertical wave flux over UTLS region in subpolar southern hemisphere. Meanwhile,
552 the changes of wave activity in middle and upper stratosphere over subpolar region
553 induced by ozone recovery are not significant. Therefore, the ozone recovery has minor
554 contribution to the significant weakening of stratospheric planetary wave activity in
555 September.

556 EOF analysis and correlation analysis indicate that the stratospheric wave
557 activities in early austral spring during 2000-2017 are related to PC2 of SST over 20°
558 N-70°S (i.e., the SHtrop region). The corresponding EOF2 mode also shows a good
559 similarity to the spatial pattern of SST trend, suggesting that the weakening of
560 stratospheric wave activities is connected to the trend of SST in SHtrop region.
561 Meanwhile, the linkage between the SST trend in NH region and the weakening of
562 stratospheric wave activities is weak. The model simulations also support that the SST
563 changes in SHtrop region lead to a weakening of tropospheric wave sources and
564 stratospheric wave activities. The contribution of SST trend in tropical region to the
565 weakening of stratospheric wave activities is larger than that in the extratropical
566 southern hemisphere. However, the response of tropospheric wave sources and
567 stratospheric wave activities to SST trend in NH region is not significant. The

568 contribution of SST trend over SHtrop region to the weakening of stratospheric wave
569 activities is about 34%. Finally, both reanalysis datasets and numerical simulations
570 indicate that the Brewer-Dobson circulation related to stratospheric wave activity has
571 also been weakening in early austral spring since 2000, which is also attributed to the
572 changes of September SST in tropics and extratropical southern hemisphere.

573 Although many researchers claimed that the climate transition around 2000 in
574 southern hemisphere is related to ozone depletion and recovery (e.g., Barnes et al., 2013;
575 Banerjee et al., 2020), there is no contradiction between our results and these previous
576 studies. Firstly, the southern hemisphere tropospheric circulation (i.e., the SAM index,
577 the tropospheric jet position and the Hadley cell edge) shifts related to ozone changes
578 in these previous studies basically occurred in austral summer (e.g., Son et al, 2008;
579 Thompson et al., 2011; Barnes et al, 2013; Banerjee et al., 2020). These tropospheric
580 circulation changes are induced by downward coupling of circulation anomalies in the
581 stratosphere (e.g., Thompson et al., 2011) during October and November, when solar
582 radiation covers the entire Antarctic and causes heating effects. However, the Antarctic
583 stratospheric circulation response to ozone variation in September is not as strong as
584 that in October or November (e.g., Thompson et al., 2011, Figs. 1b, d) because solar
585 radiation can only reach part of Antarctic stratosphere during a majority period of
586 September. This implies that the response of atmospheric state in September to
587 Antarctic stratospheric ozone change is not significant. Secondly, the FWSC
588 component used in this study is an atmospheric module with prescribed SST and forcing
589 gases. Therefore, our model results only indicate that the weakening of stratospheric

590 wave activity can be attributed to SST changes, while the impact of ozone change in
591 middle and low latitudes on SST cannot be determined based on these simulations.
592 Whether the transition signal of Antarctic stratospheric ozone is stored in the ocean
593 needs more efforts to explore. This is an issue beyond the scope of this study and further
594 investigation is necessary by using a fully coupled earth system model.

595 The southern hemisphere stratospheric wave activity trend from the early 1980s to
596 the early 2000s has been ~~investigated~~ ~~examined~~ by Hu and Fu (2009) (~~hereafter HF2009~~)
597 and hence is not ~~analyzed in present study~~ ~~discussed in detail in the above.~~ ~~HF2009~~
598 ~~attributed the strengthening of stratospheric wave activity in austral spring during 1979-~~
599 ~~2006 to the SST trends as well, however, they gave no more details about the trends of~~
600 ~~tropospheric wave sources. In this study, trends of tropospheric wave sources in~~
601 ~~September during 1980-2000 derived from MERRA-2 data is analyzed, and we also~~
602 ~~conducted an experiment (sstSHtrop) forced by the changes in September SST during~~
603 ~~1980-2000 over 20°N-70°S (see Fig. S9 for applied SST anomalies). The model result~~
604 ~~indicates that the SST changes over 20°N-70°S contribute to intensification of wave-2~~
605 ~~component of tropospheric wave sources (Fig. S10f) and weakening of the wave-1~~
606 ~~component (Fig. S10e), which is overall analogous to the trends derived from MERRA-~~
607 ~~2 data (Figs. S10b, c). Accordingly, the simulated wave-2 component of wave flux~~
608 ~~increases significantly in the stratosphere (Fig. S10h), while the response of the wave-~~
609 ~~1 component is not significant (Fig. S10i). In a word, the results from sstSHtrop80~~
610 ~~suggest that the SST changes over 20°N-70°S induce a strengthening of stratospheric~~
611 ~~wave activity in September during 1980-2000. But it cannot explain the intensified~~

612 wave-1 component of the stratospheric wave activity shown in Fig. 1b. A more detailed
613 attribution of the trend of Antarctic stratospheric wave activity during 1980-2000 needs
614 much more efforts.

615 The simulated stratospheric eddy heat flux (Fig. 11b in HF2009) forced by
616 observed time-varying SST in HF2009 is relatively weak compared to that derived from
617 reanalysis data (Fig. 6b in HF2009). Similarly, Wang and Waugh (2012) (hereafter
618 WW2012) used stratosphere-resolving chemistry-climate model forced by time-
619 varying factors to evaluate the trends of stratospheric temperature, residual circulation
620 as well as wave activity during recent decades, and the trend of cumulative eddy heat
621 flux shown in their paper is not significant (Fig. 6 in WW2012-Wang and Waugh
622 (2012)). Additionally, ~~In addition,~~ Polvani et al. (2018) used time-varying ODSs that

623 cover the period from 1960s to 2080s to simulate Brewer-Dobson circulation and
624 attained an obvious trend transition around 2000. ~~Their simulations cover from 1960s~~
625 ~~to 2080s. We had also tried to conduct transient experiments forced by time-varying~~
626 SST derived from ERSST v5 with different initial conditions, however, the trends of
627 wave activities in the transient simulations are so weak, though the opposite trend signs
628 exist during 1980-2000 and 2000-2018 (Table S2, Fig. S11). The significance of
629 simulated trend may be related to model performance and the length of simulating
630 period. As the period we focus is relatively short and our purpose is attribution rather
631 than generating a real trend, we perform the ensemble time-slice experiments in this
632 study, which are also used in many other previous researches (e.g., Hu et al., 2018;
633 Kang et al., 2011; Zhang et al., 2016) to attribute trends in the atmosphere. In addition,

634 most of the current climate models cannot generate a realistic wave activity trend as
635 waves in the atmosphere are linked with various processes and factors (e.g., Baldwin
636 & Dunkerton, 2005; Garcia & Randel, 2008; Labitzke, 2005; Shindell et al., 1999; Shu
637 et al., 2013; Xie et al., 2008).

638

639 **Data availability:**

640 The ERA-Interim is available at: <https://apps.ecmwf.int/datasets/data/interim->
641 [full-daily/levtype=sfc/](https://apps.ecmwf.int/datasets/data/interim-full-daily/levtype=sfc/). The MERRA-2 is available at: [https://disc.gsfc.nasa.gov/d](https://disc.gsfc.nasa.gov/datasets?keywords=%22MERRA-2%22&page=1&source=Models%2FAnalyses%20MERRA-2)
642 [atasets?keywords=%22MERRA-2%22&page=1&source=Models%2FAnalyses%20M](https://disc.gsfc.nasa.gov/datasets?keywords=%22MERRA-2%22&page=1&source=Models%2FAnalyses%20MERRA-2)
643 [ERRA-2](https://disc.gsfc.nasa.gov/datasets?keywords=%22MERRA-2%22&page=1&source=Models%2FAnalyses%20MERRA-2). The JRA-55 is available at: [https://jra.kishou.go.jp/JRA-55/index_en.ht](https://jra.kishou.go.jp/JRA-55/index_en.html#download)
644 [ml#download](https://jra.kishou.go.jp/JRA-55/index_en.html#download). The NCEP-2 is available at: [http://www.cpc.ncep.noaa.gov/product](http://www.cpc.ncep.noaa.gov/product/s/wesley/reanalysis2/)
645 [s/wesley/reanalysis2/](http://www.cpc.ncep.noaa.gov/product/s/wesley/reanalysis2/). The ERSST v5 dataset is available at: [https://www1.ncdc.](https://www1.ncdc.noaa.gov/pub/data/cmb/ersst/v5/netcdf/)
646 [noaa.gov/pub/data/cmb/ersst/v5/netcdf/](https://www1.ncdc.noaa.gov/pub/data/cmb/ersst/v5/netcdf/). The observations of TCO from SBUV v
647 8.6 satellite dataset are available at: [https://acd-ext.gsfc.nasa.gov/Data_services/m](https://acd-ext.gsfc.nasa.gov/Data_services/merged/data/sbu_v86_mod.int_lyr.70-18.za.r7.txt)
648 [erged/data/sbu_v86_mod.int_lyr.70-18.za.r7.txt](https://acd-ext.gsfc.nasa.gov/Data_services/merged/data/sbu_v86_mod.int_lyr.70-18.za.r7.txt). The unfiltered IPO index derived
649 from ERSST v5 dataset is available at: [https://psl.noaa.gov/data/timeseries/IPOT](https://psl.noaa.gov/data/timeseries/IPOTPI/tpi.timeseries.ersstv5.data)
650 [PI/tpi.timeseries.ersstv5.data](https://psl.noaa.gov/data/timeseries/IPOTPI/tpi.timeseries.ersstv5.data).

651 **Author contributions:**

652 Yihang Hu conducted experiments, produced figures and tables, organized and
653 wrote the manuscript. Wenshou Tian, Jiankai Zhang and Tao Wang contributed to revise
654 the manuscript. Mian Xu helped to design experiments.

655 **Competing interests:**

656 The authors declare that they have no competing interest.

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667

668

APPENDIX

669 **Analysis of time lag for tropical SST affects Antarctic stratospheric** 670 **wave activity**

671 As stated in the Section 2, the tropical SST anomalies (the linear increments) in
672 experiment ssttrop are also applied in July and August (Fig. [S4S5a, b](#)) to avoid abrupt
673 SST variations from month to month, and the two months are taken as spin-up time.
674 Therefore, whether the SST forcing in July and August also contribute to the weakening
675 of Antarctic stratospheric wave activity in September or not cannot be justified based
676 on the experiment ssttrop only. Here, we performed an additional experiment
677 ssttropAug without September SST anomalies (Fig. [S4S5f](#)) to clarify whether the

678 weakening of Antarctic stratospheric wave activity is induced by the tropical SST trend
679 at the same month. Like other numerical experiments described in Table 1, the
680 ssttropAug also includes 100 ensemble members that run from July to September forced
681 by the same initial conditions from the 21st year to the 120th year in July generated by
682 free run. The detailed descriptions of ssttropAug and other relevant experiments in the
683 manuscript are displayed together in the Table S1 for comparison. Figure [S4-S5](#) shows
684 the applied global SST anomalies in ssttrop and ssttropAug from July to September.

685 The responses of tropospheric wave sources and stratospheric wave activities in
686 ssttropAug are shown in Figs. [S5a-S6a-c](#) and Figs. [S5d-S6d-f](#), respectively. Note that the
687 anomalies of subpolar tropospheric geopotential height in September forced by changes
688 in tropical SST in August does not superpose on their climatological patterns in an
689 evident out-of-phase style (Figs. [S5a-S6a-c](#)). The anomaly of wave-1 component of
690 geopotential height shows a slight in-phase overlap with its climatology over subpolar
691 region (Fig. [S5b-S6b](#)). Accordingly, the responses of stratospheric wave activities over
692 subpolar of southern hemisphere are not significant (Figs. [S5d-S6d-f](#)). The results here
693 suggest that, the decrease of September vertical wave flux induced by SST changes in
694 August is negligible comparing to that in the experiment with anomalous SST forcing
695 in September (Figs. [S5g-S6g](#)), and the tropical SST trend in September plays a dominate
696 role in weakening of stratospheric wave activity at the same month.

697 Furthermore, we also use a linear barotropic model (LBM) (e.g., Shaman &
698 Tziperman, 2007; Shaman & Tziperman, 2011) to quantify the time scale for
699 propagation of tropical anomalies to high latitudes. The LBM are developed to solve

700 the barotropic vorticity equation, which is given as Eq. (A1):

$$701 \quad J(\bar{\psi}, \nabla^2 \psi') + J(\psi', \nabla^2 \bar{\psi} + f) + \alpha \nabla^2 \psi' + K \nabla^4 \nabla^2 \psi' = R \quad (\text{A1})$$

702 where the Jacobian $J(A, B)$ is

$$703 \quad J(A, B) = \frac{1}{r^2} \left(\frac{\partial A}{\partial \lambda} \frac{\partial B}{\partial \mu} - \frac{\partial A}{\partial \mu} \frac{\partial B}{\partial \lambda} \right) \quad (\text{A2})$$

704 the forcing function R is

$$705 \quad R = -(f + \nabla^2 \bar{\psi})D \quad (\text{A3})$$

706 ψ is the streamfunction, f is the Coriolis force, α is the Rayleigh coefficient, K
707 is the diffusion coefficient, λ is the longitude, $\mu = \sin(\theta)$, θ is the latitude, r is
708 the earth's radius and D is the divergence.

709 We use the wave-1 component of streamfunction derived from ensemble mean of
710 sstctrl as the background field. In LBM, the initial anomaly is given by the divergence.
711 The divergence forcing field is limited in 40°E-140°W, 10°S-0° (Fig. S6S7) to ensure
712 that the tropical initial anomaly of streamfunction superpose on its background field in
713 an out-of-phase style. We set $D = -7.9 \times 10^{-7} \text{ s}^{-1}$, which is the mean divergence over
714 the forcing region. The LBM simulated streamfunction anomalies are shown in Figs.
715 S7bS8b-i. Note that the anomalies in tropics only take a few days to arrive the high
716 latitudes in southern hemisphere. After about four days, a stable anti-phase
717 superposition of streamfunction is well established in extratropical southern
718 hemisphere (Figs. S7fS8f-i). These results are supported by previous studies (e.g.,
719 Shaman & Tziperman, 2011), which also indicate that the horizontal propagation of
720 anomaly in atmosphere takes a few days.

721 Previous studies also reported that it takes about 4 days for wave-1 to propagate

722 from troposphere into stratosphere and 1-2 days for wave-2 (e.g., Randel, 1987). Thus,
723 the tropical oceans affect the stratosphere at mid-high latitudes with a lag of several
724 days. However, the SST forcing field applied in CESM is on monthly scale. It is
725 reasonable to use September SST trend to drive and explain the trends of extratropical
726 circulation and wave activity at the same month.

727

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943

944 **Table 1.** Configurations of experiments for SST trends.

Experiments	Descriptions
sstctrl	Control run. Seasonal cycle of monthly mean global SST data over 1980-2000 is derived from the ERSST v5 dataset. Fixed values of ozone greenhouse gases and aerosol fields in 2000 are used.
sstNH	As in sstctrl, but with linear increments of SST in September over 2000-2017 in NH (20°N-70°N). The applied global SST anomalies are shown in Fig. 7a.
sstSH	As in sstctrl, but with linear increments of SST in September over 2000-2017 in SH (20°S-70°S). The applied global SST anomalies are shown in Fig. 7b.
ssttrop	As in sstctrl, but with linear increments of SST in September over 2000-2017 in the tropics (20°S-20°N). The applied global SST anomalies are shown in Fig. 7c.
sstSHtrop	As in sstctrl, but with linear increments of SST in September over 2000-2017 in SHtrop (20°N-70°S). The applied global SST anomalies are shown in Fig. 7d.

945 **Table 2.** Configurations of experiments for the ozone recovery trend.

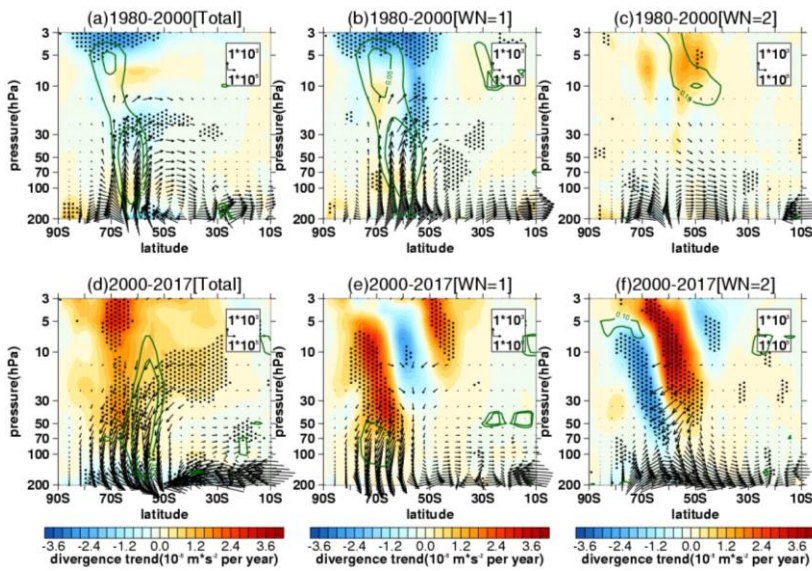
Experiments	Descriptions
O3ctrl	Control run. The seasonal cycle of monthly averaged global SST data over 1980-2000 is derived from ERSST v5 dataset. The seasonal cycle of monthly mean three-dimensional global ozone over 1980-2000 is derived from MERRA-2 dataset. The GHGs and aerosol fields are specified to be fixed values in 2000.

O3sen As in O3ctrl, but superposed with linear increments of global ozone in September over 2001-2017. The ozone data in 2002 are removed when the linear increments are calculated. The applied ozone anomalies in Southern Hemisphere are shown in Fig. S5.

946 **Table 3.** Correlations of stratospheric vertical wave flux time series (area-weighted
947 from 100 hPa to 30 hPa over 70°S-50°S) between different reanalysis dataset.

	ERA-Interim	JRA-55	MERRA-2	NCEP-2
ERA-Interim	1.00 (p=0.00)	0.99 (p<0.01)	0.98 (p<0.01)	0.93 (p<0.01)
JRA-55		1.00 (p=0.00)	0.98 (p<0.01)	0.93 (p<0.01)
MERRA-2			1.00 (p=0.00)	0.94 (p<0.01)
NCEP-2				1.00 (p=0.00)

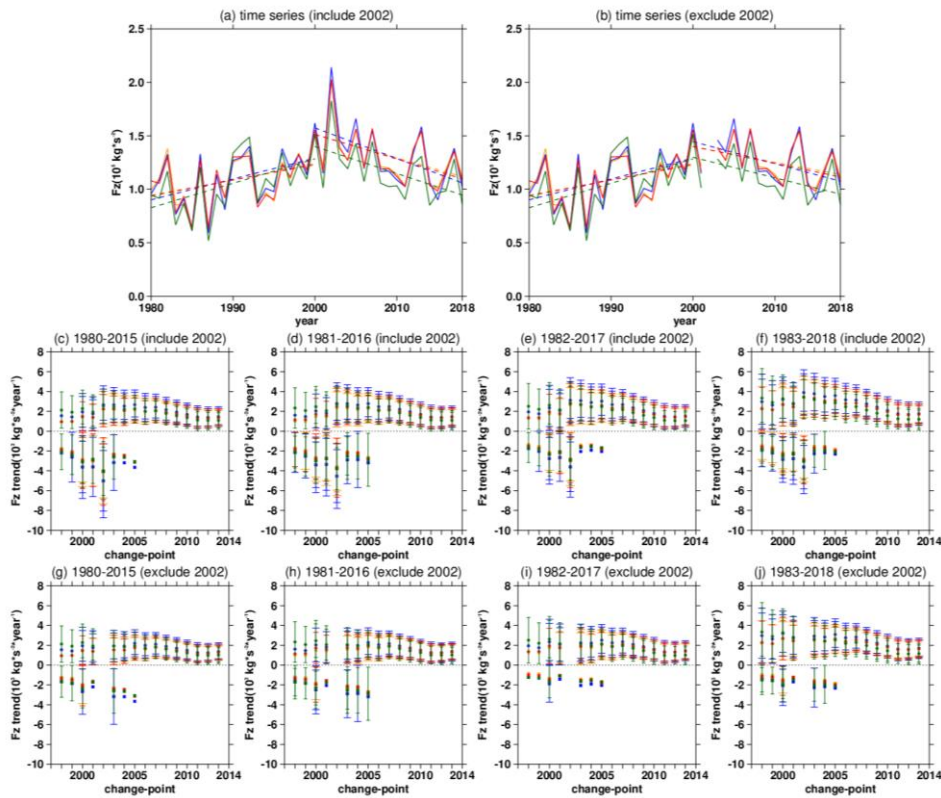
948



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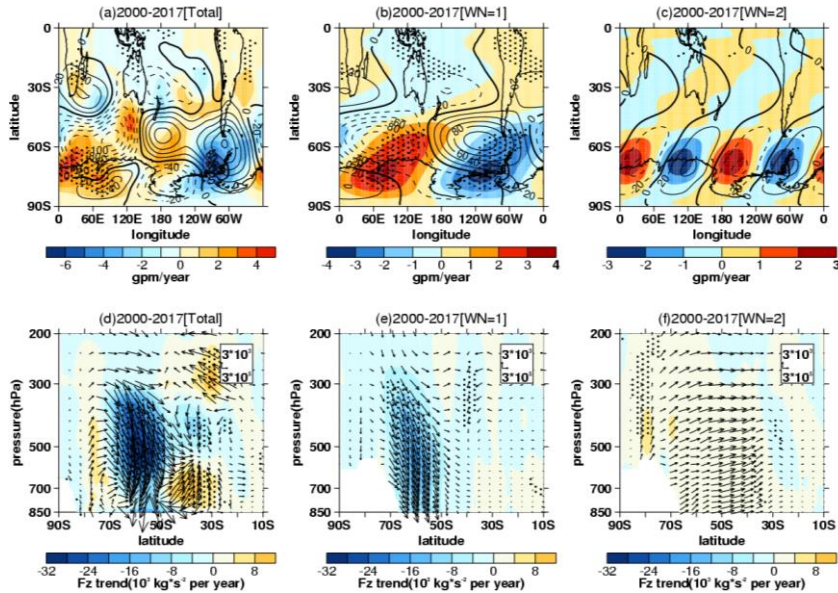
950 **FIG. 1.** Trends of southern hemisphere (a, d) stratospheric E-P flux (arrows, units of
951 horizontal and vertical components are 10^5 and 10^3 $\text{kg}\cdot\text{s}^{-2}$ per year, respectively) and its

952 divergence (shadings) with their (b, e) wave-1 components and (c, f) wave-2
 953 components over (a, b, c) 1980-2000 and (d, e, f) 2000-2017 in September derived from
 954 MERRA-2 dataset. The stippled regions indicate the trend of E-P flux divergence
 955 significant at/above the 90% confidence level. The green contours from outside to
 956 inside (corresponding to $p=0.1$, 0.05) indicate the trend of vertical E-P flux significant
 957 at the 90% and 95% confidence level, respectively.
 958



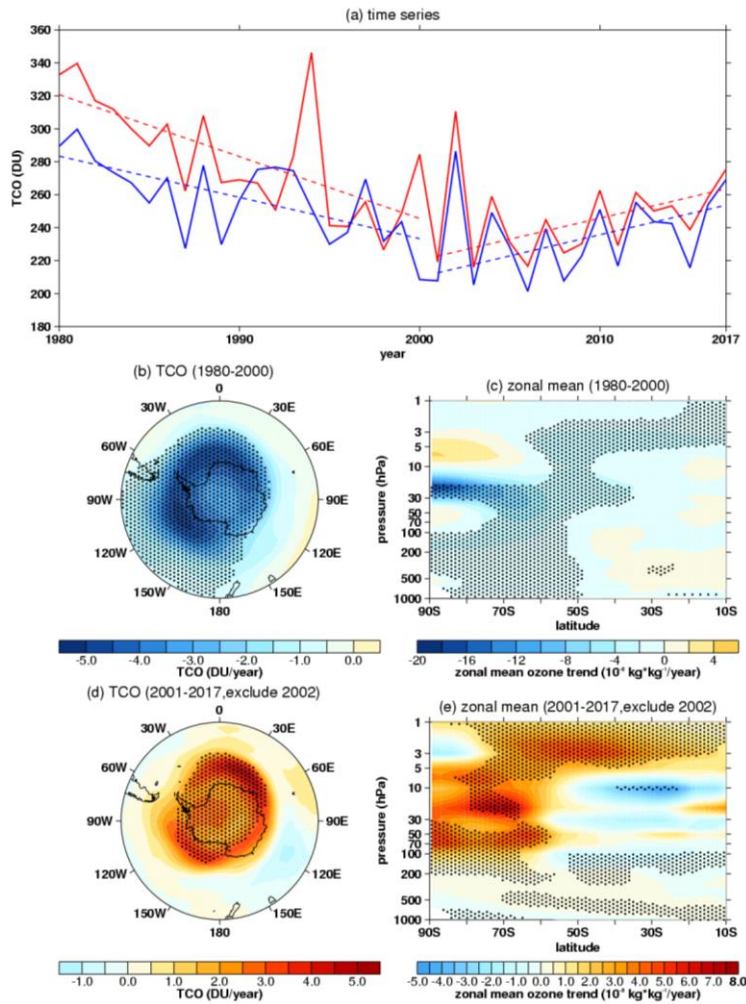
959
 960 **FIG. 2.** (a) The mean time series (solid lines) and piecewise (during 1980-2000 and
 961 2000-2018) linear regressions (dashed lines) of vertical E-P flux area-weighted from

962 100 hPa to 30 hPa over 70°S-50°S in September during 1980-2018 derived from ERA-
963 Interim (yellow), MERRA-2 (blue), JRA-55 (red) and NCEP-2 (green). Figure (b) is
964 the same as Figure (a), except for that the data in 2002 are removed. (c, d, e, f) The
965 trends (dots) and uncertainties (error bars) calculated during various periods using the
966 change-point method with different beginning and ending years (titles). Circles and
967 squares in Figures (c, d, e, f) represent positive trends from beginning years to change-
968 point years (x-axes) and negative trends from change-point years to ending years,
969 respectively. Different colors of dots and error bars in Figures (c, d, e, f) correspond to
970 colors in Figure (a), which represent trends and uncertainties derived from different
971 datasets. The long and short error bars in same color reflect the 95% and 90%
972 confidence intervals calculated by two-tailed t test. The error bar is omitted when the
973 significance of trend is lower than corresponding confidence level. Negative trends and
974 corresponding uncertainties with the beginning change-point years after 2005 are also
975 omitted, since the trend value shows large fluctuation with shortening of time series.
976 Figures (g, h, i, j) are the same as Figures (c, d, e, f), except that the data in 2002 are
977 removed when calculating trends and uncertainties.



978

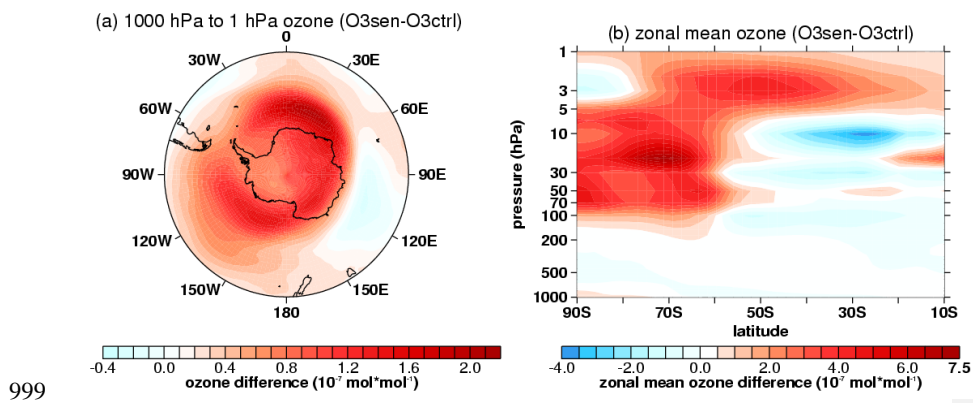
979 **FIG. 3.** Trends (shadings) and climatological distributions (contours with an interval
 980 of 20 gpm, positive and negative values are depicted by solid and dashed lines
 981 respectively, zeroes are depicted by thick solid lines) of southern hemispheric (a) 500
 982 hPa geopotential height zonal deviations with their (b) wave-1 component and (c)
 983 wave-2 component in September during 2000–2017 derived from MERRA-2 dataset.
 984 Trends of southern hemispheric (d) tropospheric E-P flux (arrows, units of horizontal
 985 and vertical components are 3×10^5 and 3×10^3 kg s⁻² per year, respectively) and its
 986 vertical component (shading) with their (e) wave-1 component and (f) wave-2
 987 component in September during 2000–2017 derived from MERRA-2 dataset. The
 988 stippled regions represent the trend significant at/above the 90% confidence level.



989

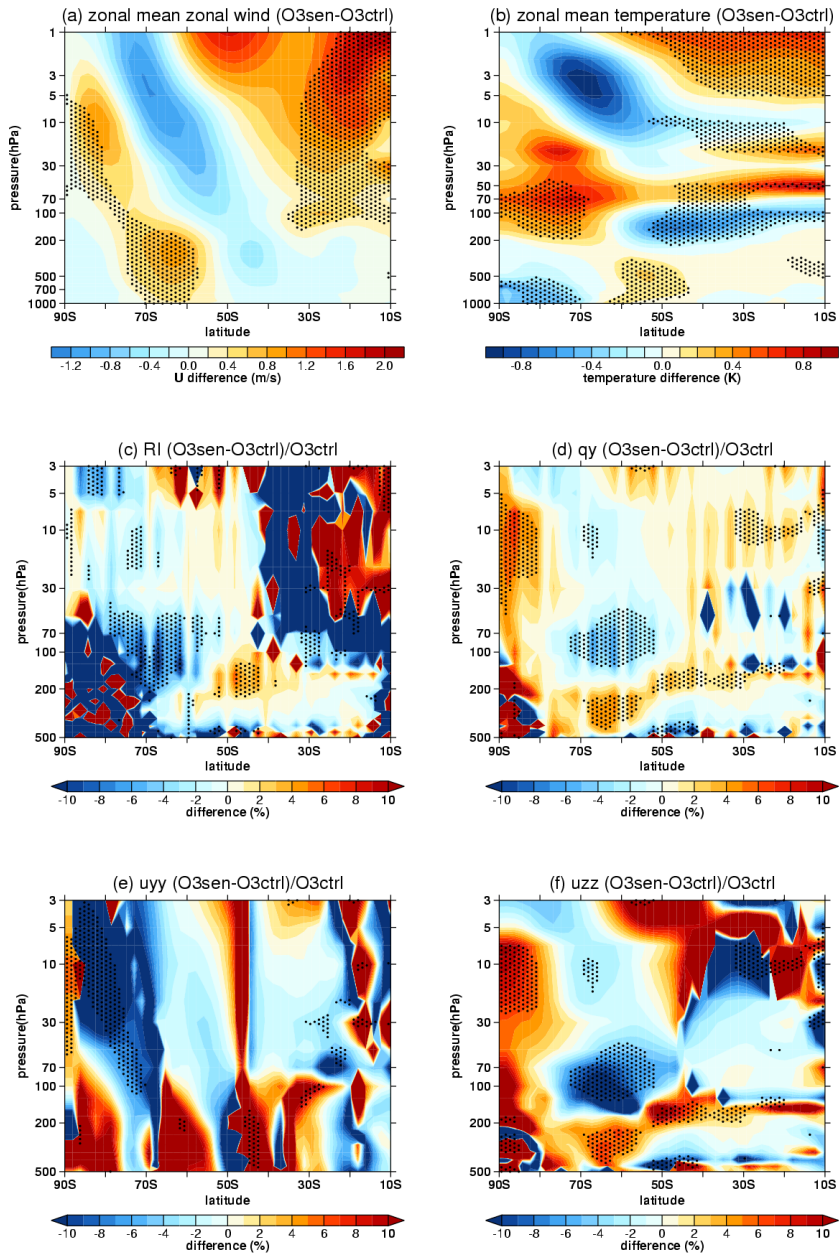
990 **FIG. 4.** (a) Time series (solid lines) of aera-weighted total column ozone (TCO) over
 991 60°S to 90°S derived from MERRA-2 (red) and SBUV (blue) datasets. The dashed lines
 992 represent linear regression of TCO. (b, d) The TCO trends in September during 1980-
 993 2000 (b) and 2001-2017 (d) derived from MERRA-2 dataset. The outermost latitudes
 994 in Figs. 4c, d are both 40°S. (c, e) The zonal mean ozone trends on latitude-pressure
 995 profile in September during 1980-2000 (c) and 2001-2017 (e) derived from MERRA-2

996 dataset. The stippled regions in Figs. 4b-e represent trends significant at/above the 90%
997 confidence level. Data in 2002 are removed when trends, regressions and significances
998 are calculated in Fig. 4.



999

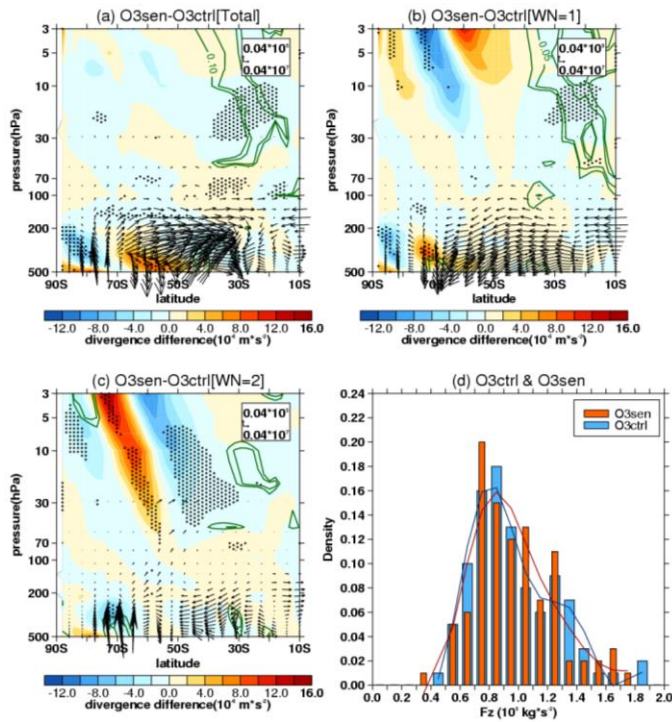
1000 **FIG. 5.** (a) Difference of horizontal ozone forcing field averaged from 1000 hPa to 1
1001 hPa between O3sen and O3ctrl. The outermost latitude in Fig. 5a is 40°S. (b) Zonal
1002 mean difference of ozone forcing fields on latitude-pressure profile in the southern
1003 hemisphere between O3sen and O3ctrl.



1004

1005 **FIG. 6.** Difference of (a) zonally averaged zonal wind, (b) zonally averaged

1006 temperature, (c) refractive index, (d) $a^2 \bar{q}_\varphi$, (e) $-\left[\frac{(\bar{u} \cos \varphi)}{\cos \varphi}\right]_\varphi$ (uyy term), (f)
 1007 $-\frac{a^2 f^2}{\rho_0} (\rho_0 \frac{\bar{u}_z}{N^2})_z$ (uzz term) between O3sen and O3ctrl. The stippled regions represent
 1008 the difference significant at/above 90% confidence level.

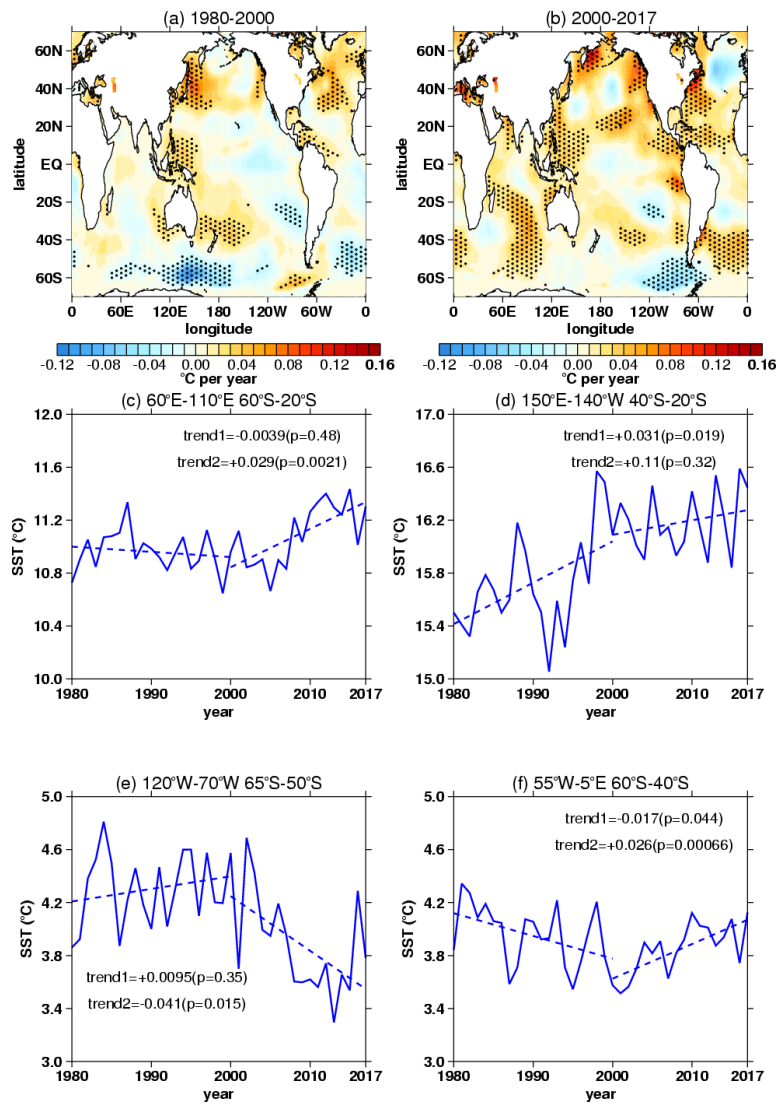


1009
 1010 **FIG. 7.** Differences of (a) stratospheric E-P flux (arrows, units in horizontal and vertical
 1011 components are 0.04×10^7 and $0.04 \times 10^5 \text{ kg s}^{-2}$, respectively) and its divergence
 1012 (shadings) with their (b) wave-1 component and (c) wave-2 component between the
 1013 sensitive experiment (O3sen) and the control experiment (O3ctrl). The stippled regions
 1014 represent the mean differences of E-P flux divergence significant at/above the 90%
 1015 confidence level. The green contours from outside to inside (corresponding to $p=0.1$,
 1016 0.05) represent the mean differences of vertical E-P flux significant at the 90% and 95%

1017 confidence levels, respectively. (d) Frequency distributions (pillars, blue for O3ctrl and
1018 orange for O3sen) of vertical E-P flux (F_z , area-weighted from 200 hPa to 10 hPa over
1019 70°S - 50°S) and its 5-point low-pass filtered fitting curves (solid lines, blue for O3ctrl
1020 and red for O3sen) derived from 100 ensemble members.

1021

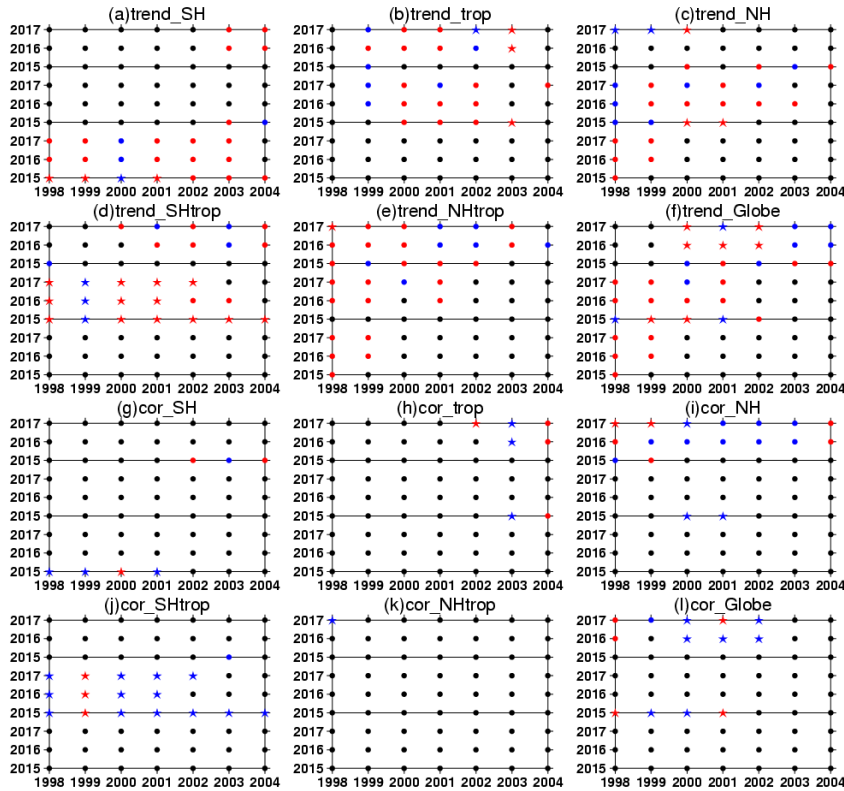
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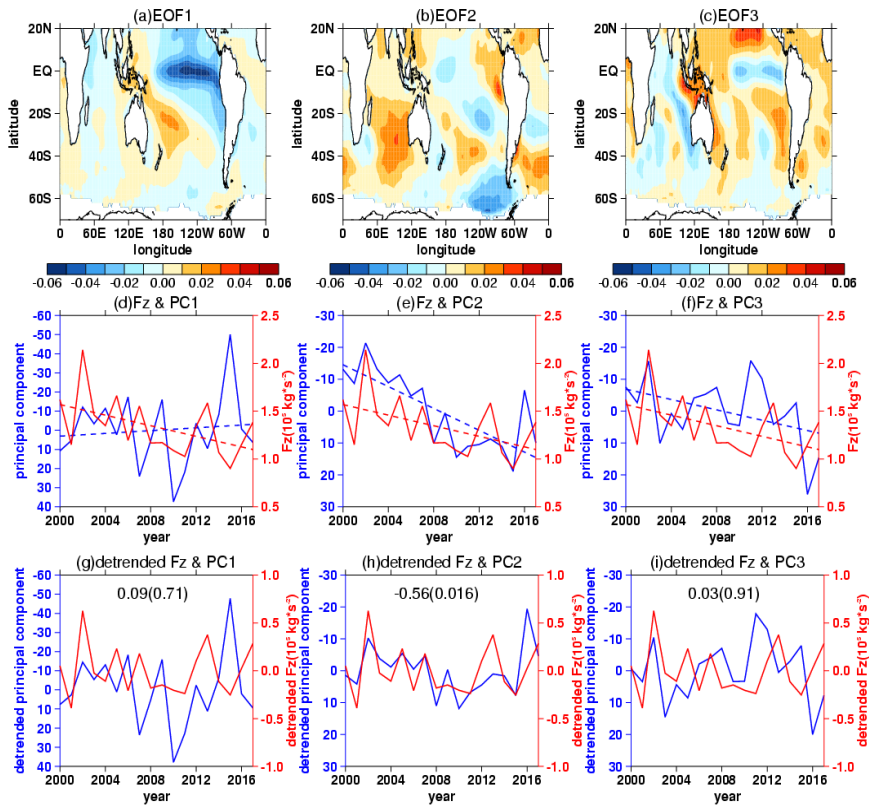
1023 **FIG. 8.** Trends of SST in September over (a) 1980-2000 and (b) 2000-2017 derived
 1024 from ERSST v5 dataset. The stippled regions represent the trends significant at/above
 1025 the 90% confidence level. (c-f) Time series (blue solid lines) of SST during 1980-2017
 1026 over different regions (titles). The dashed lines represent linear regressions of SST time
 1027 series on piecewise periods (1980-2000 and 2000-2017). The “trend1” and “trend2”

1028 labeled in Figs. 8c-f represent the trend coefficients and the corresponding significances
 1029 (bracketed) over 1980-2000 and 2000-2017, respectively.



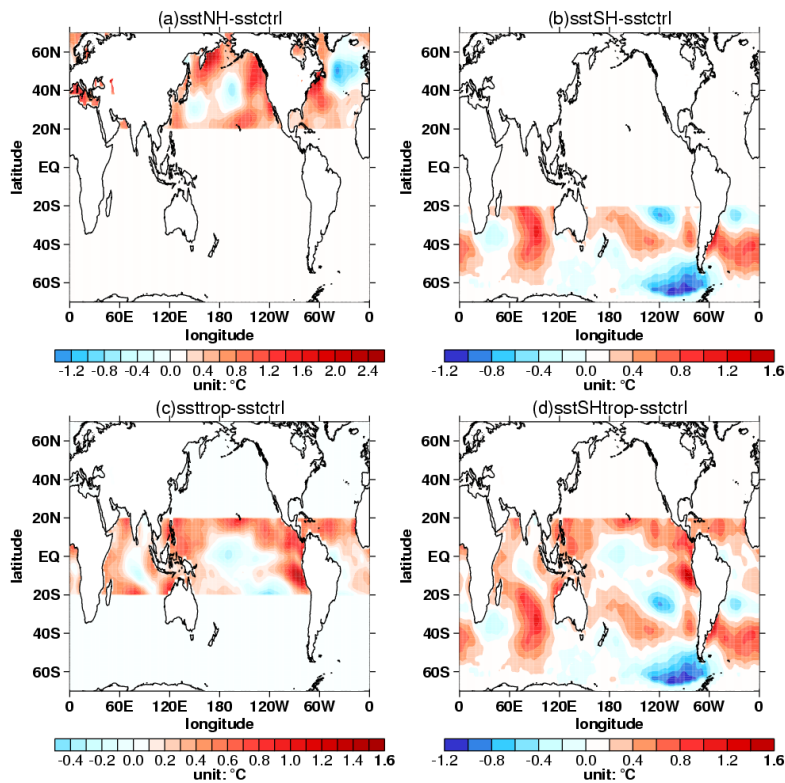
1030
 1031 **FIG. 9.** Trend significance of the first three SST principal components (PCs) in (a) the
 1032 extratropical southern hemisphere (SH, 70°S-20°S), (b) the tropics (TROP, 20°S-20°N),
 1033 (c) the extratropical northern hemisphere (NH, 20°N-70°N), (d) the extratropical
 1034 southern hemisphere and the tropics (SHtrop, 70°S-20°N), (e) the extratropical northern
 1035 hemisphere and the tropics (NHtrop, 20°S-70°N), (f) the globe (70°S-70°N) and the
 1036 corresponding (g, h, i, j, k, l) correlation significances between them and vertical E-P
 1037 flux (Fz, area-weighted from 100 hPa to 30 hPa over 70°S-50°S) during different
 1038 beginning years (x-axes) and ending years (y-axes). The red and blue dots indicate

1039 positive and negative trend or correlation coefficient are significant, respectively. The
 1040 black dots indicate the trends or correlation coefficients are not significant. The stars
 1041 indicate that the trends and the corresponding correlation coefficients are both
 1042 significant. Each panel is divided into three regions from bottom to top, corresponding
 1043 to the first, the second and the third principal components, respectively. The criterion
 1044 to distinguish whether the trends and correlations are significant or not is the 90%
 1045 confidence level.

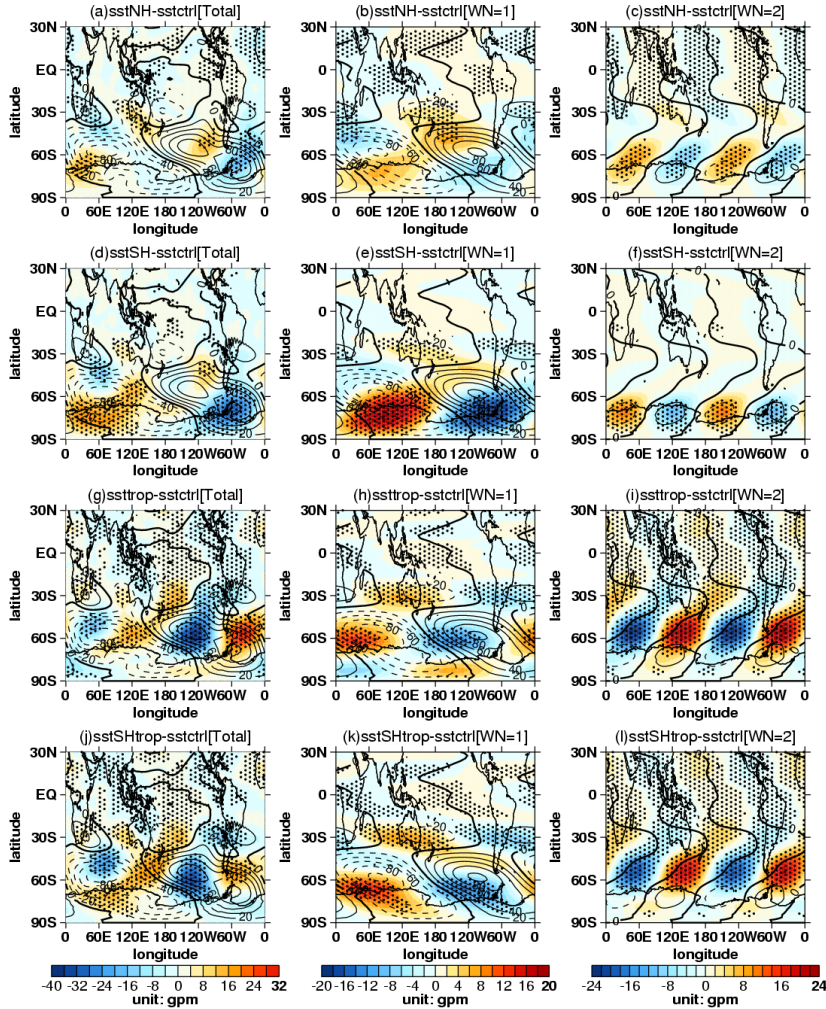


1046
 1047 **FIG. 10.** (a, b, c) The first three EOF patterns of SST in SHtrop region. (d, e, f) The
 1048 original time series of the first three principle components (PCs, blue solid lines
 1049 correspond to left inverted y-axes) and stratospheric vertical E-P flux (Fz, area-

1050 weighted from 100 hPa to 30 hPa over 70°S-50°S, red solid lines correspond to right y-
 1051 axes) in September during 2000-2017. The blue and red dashed lines in (d, e, f)
 1052 represent the linear regressions of PC time series and Fz time series, respectively. The
 1053 meaning of (g, h, i) are the same as (d, e, f) correspondingly, except the detrended time
 1054 series. The unbracketed and bracketed numbers in (g, h, i) represent the correlation
 1055 coefficients between detrended PC time series and Fz time series and the corresponding
 1056 p values calculated by two-tailed t test, respectively.



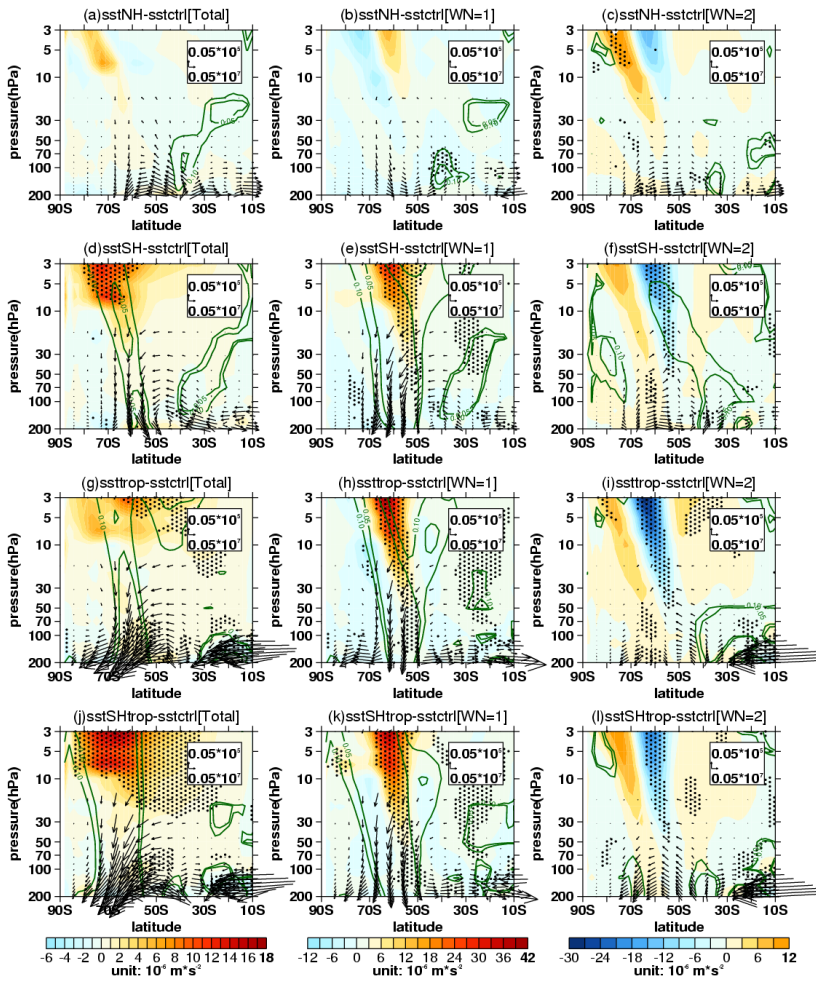
1057
 1058 **FIG. 11.** Differences in SST forcing field between sensitive experiments ((a) sstNH; (b)
 1059 sstSH; (c) ssttrop; (d) sstSHtrop) and the control experiment (sstctrl).



1060

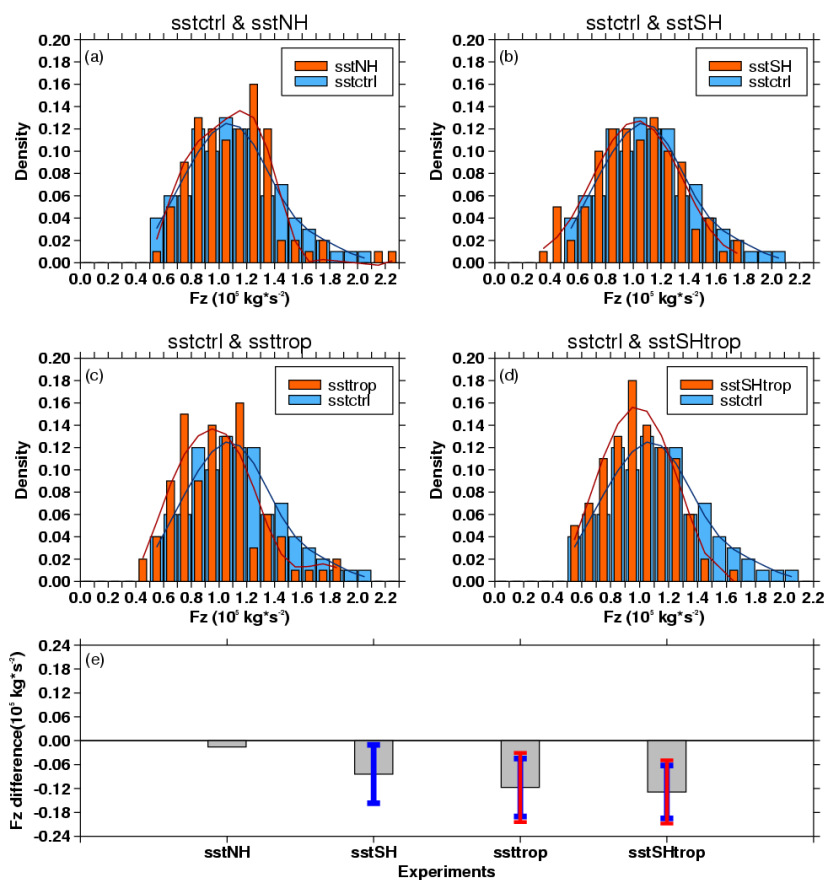
1061 **FIG. 12.** Differences (shadings) of (a, d, g, j) 500 hPa geopotential height zonal
 1062 deviations with their (b, e, h, k) wave-1 component and (c, f, i, l) wave-2 component
 1063 between sensitive experiments ((a, b, c) sstNH; (d, e, f) sstSH; (g, h, i) ssttrop; (j, k, l)
 1064 sstSHtrop) and the control experiment (sstctrl). The mean distributions (contours with
 1065 an interval of 20 gpm, positive and negative values are depicted by solid and dashed
 1066 lines respectively, zeroes are depicted by thick solid lines) of them are derived from the

1067 control experiment. The stippled regions represent the mean difference significant
 1068 at/above the 90% confidence level.



1069
 1070 **FIG. 13.** Differences of (a, d, g, j) stratospheric E-P flux (arrows, units in horizontal
 1071 and vertical components are 0.05×10^7 and 0.05×10^5 $\text{kg} \cdot \text{s}^{-2}$, respectively) and its
 1072 divergence (shadings) with their (b, e, h, k) wave-1 component and (c, f, i, l) wave-2
 1073 component between sensitive experiments ((a, b, c) sstNH; (d, e, f) sstSH; (g, h, i)
 1074 ssttrop; (j, k, l) sstSHtrop) and the control experiment (sstctrl). The stippled regions

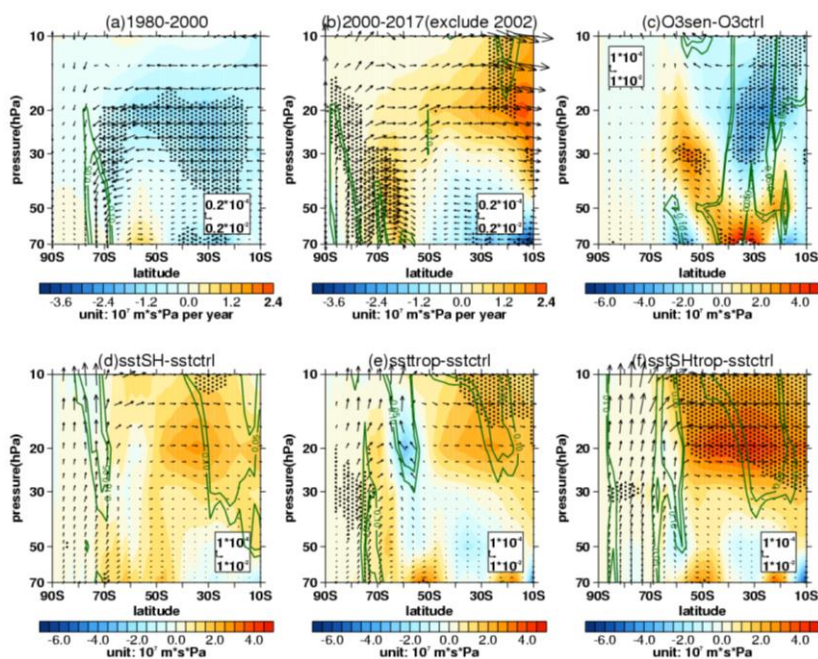
1075 represent the mean differences of E-P flux divergence significant at/above the 90%
 1076 confidence level. The green contours from outside to inside (corresponding to $p=0.1$,
 1077 0.05) represent the mean differences of vertical E-P flux significant at the 90% and 95%
 1078 confidence levels, respectively.



1079
 1080 **FIG. 14.** (a, b, c, d) Frequency distributions (pillars, blue for control experiment and
 1081 orange for sensitive experiments) of vertical E-P flux (F_z , area-weighted from 200 hPa
 1082 to 10 hPa over 70°S - 50°S) and its 5-point low-pass filtered fitting curves (solid lines,
 1083 blue for control experiment and red for sensitive experiments) derived from 100

1084 ensemble members of the control experiment (sstctrl) and sensitive experiments ((a)
 1085 sstNH; (b) sstSH; (c) ssttrop; (d) sstSHtrop), respectively. (e) Mean differences (grey
 1086 pillars) and corresponding uncertainties (error bars) of Fz between sensitive
 1087 experiments and the control experiment. The blue and red error bars reflect the 90%
 1088 and 95% confidence levels calculated by two-tailed t test, respectively. The error bar is
 1089 omitted when the significance of mean difference is lower than the corresponding
 1090 confidence level.

1091



1092

1093 **FIG. 15.** (a) Trends of southern hemispheric Brewer-Dobson circulation (arrows, units
 1094 in horizontal and vertical components are 0.2×10^{-2} and 0.2×10^{-4} $\text{m} \cdot \text{s}^{-1}$ per year,
 1095 respectively) and its stream function (shadings) in September during (a) 1980-2000 and
 1096 (b) 2000-2017 derived from MERRA-2 dataset. Data in 2002 are removed when trends

1097 are calculated in Figure (b). (c) Differences of Brewer-Dobson circulation (arrows,
1098 units in horizontal and vertical components are 10^{-2} and $10^{-4} \text{ m}\cdot\text{s}^{-1}$, respectively) and its
1099 stream function (shadings) between the O3ctrl and O3sen. (d, e, f) Differences of
1100 Brewer-Dobson circulation and its stream function between the control experiment
1101 (sstctrl) and sensitive experiments ((d) sstSH; (e) ssttrop; (f) sstSHtrop) with SST
1102 changes. The stippled regions represent the trends or differences of the stream function
1103 significant at/above the 90% confidence level. The green contours from outside to
1104 inside (corresponding to $p=0.1, 0.05$) represent the trends or differences of the vertical
1105 components significant at the 90% and 95% confidence levels, respectively.