



1 **Five satellite sensor study of the rapid decline of wildfire smoke in the**
2 **stratosphere**

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9

10 **Abstract**

11

12 Smoke from Western North American wildfires reached the stratosphere in large amounts in
13 August 2017. Limb-oriented satellite-based sensors are commonly used for studies of wildfire
14 aerosol injected into the stratosphere (OMPS-LP (Ozone Mapping and Profiler Suite Limb
15 Profiler) and SAGE III/ISS (Stratospheric Aerosol and Gas Experiment III on the International
16 Space Station)). We find that these methods are inadequate for studies the first 1 – 2 months after
17 such a strong fire event due to event termination (“saturation”). The nadir-viewing lidar CALIOP
18 (Cloud-Aerosol Lidar with Orthogonal Polarization) is less affected due to shorter path in the
19 smoke, and, further, provides means that we could use to develop a method to correct for strong
20 attenuation of the signal. After the initial phase, the aerosol optical depth (AOD) from OMPS-LP
21 and CALOP show very good agreement above the 380 K isentrope, whereas the OMPS-LP tends
22 to produce higher AOD than CALIOP in the lowermost stratosphere (LMS), probably due to
23 reduced sensitivity at altitudes below 17 km. Time series from CALIOP of attenuation-corrected
24 stratospheric AOD of wildfire smoke show an exponential decline during the first month after the
25 fire, which coincides with highly significant changes in the wildfire aerosol optical properties.
26 The AOD decline is verified by the evolution of the smoke layer composition, comparing the
27 aerosol scattering ratio (CALIOP) to the water vapor concentration from MLS (Microwave Limb
28 Sounder). Initially the stratospheric wildfire smoke AOD is comparable with the most important
29 volcanic eruptions during the last 25 years. Wildfire aerosol declines much faster, 80 – 90% of
30 the AOD is removed with a half-life of approximately 10 days. We hypothesize that this dramatic
31 decline is caused by photolytic loss. This process is rarely observed in the atmosphere. However,
32 in the stratosphere this process can be studied with practically no influence from wet deposition,
33 in contrast to the troposphere where this is the main removal path of sub-micron aerosol particles.
34 Despite the loss, the aerosol particles from wildfire smoke in the stratosphere are relevant for the
35 climate.

36

37 **1. Introduction**

38

39 Background stratospheric aerosol is composed of sulfuric acid, water, carbonaceous components,
40 and minor extraterrestrial and tropospheric components (Murphy et al., 2007; Kremser et al.,
41 2016; Martinsson et al., 2019). Volcanism is a strong source of the stratospheric sulfurous,
42 carbonaceous and ash aerosol (Martinsson et al., 2009; Andersson et al., 2013; Friberg et al.,
43 2014). Large eruptions, like that of Mt Pinatubo in 1991, affect the stratosphere for several years,
44 causing global cooling of several tenths of degrees Kelvin (Kremser et al., 2016). These eruptions
45 are scarce, only a few per century (Ammann et al., 2003; Stothers, 2007). Moderate eruptions are
46 more frequent contributors to the stratospheric aerosol (Vernier et al., 2011; Andersson et al.,
47 2015; Friberg et al., 2018), forming the persistently variable stratospheric background aerosol
48 (Solomon et al., 2011).

49

50 The stratospheric aerosol is also influenced by pyrocumulonimbus clouds (pyroCb) that form
51 during extreme weather conditions in connection with intense wildfires (Fromm et al., 2010). The



52 ongoing climate change is projected to increase the frequency of large wildfires (Kasischke et al.,
53 2006; Dennison et al., 2014). Interestingly, the two largest events have, in terms of stratospheric
54 impact, occurred during the last few years, in North America 2017 (Peterson et al., 2018) and
55 Australia 2019-2020 (Kablick et al., 2020). Here we investigate the great pyroCbs formed in
56 western North America on August 12 – 13, 2017. Figure 1a shows an example on the strong
57 impact on the stratospheric aerosol of the 2019 Raikoke volcanic eruption, one of the strongest
58 eruptions post Mt Pinatubo in 1991. In comparison, Figure 1b demonstrates the formidable early
59 impact of wildfire aerosol. The stratospheric impact of that fire has been described in terms of
60 light-backscatter reaching unprecedentedly high values for a non-volcanic aerosol layer (Khaykin
61 et al., 2018), light extinction about 20 times higher than after the Pinatubo volcanic eruption in
62 1991 (Ansmann et al., 2018), and mass of smoke comparable to that of a moderate sized volcanic
63 eruption (Peterson et al., 2018). The pyroCbs lifted smoke from the fire to the extratropical
64 tropopause region, where absorption of radiation by black carbon (BC) in the smoke induced
65 additional lift to 23 km altitude in 2 months (Yu et al., 2019).

66
67 Smoke particles from wildfires contain a dominating fraction of organic matter by mass
68 (Garofalo et al., 2019). Organic aerosol is susceptible to photochemical loss (Jimenez et al.,
69 2009), and laboratory studies have demonstrated that this phenomenon could be an important
70 sink of secondary aerosol mass (Molina et al., 2004; Sareen et al., 2013). The residence time of
71 stratospheric air spans months to years depending on its path in the Brewer-Dobson circulation
72 (Engel et al., 2009; Bönisch et al., 2009). Due to very low probability of clouds, fine aerosol
73 particles have considerably longer residence times in the stratosphere than in the troposphere,
74 which further emphasizes the importance of investigating photochemical loss in the stratosphere
75 (Martinsson et al., 2019).

76
77 The aim of this study is to further understand the stratospheric aerosol sources and its climate
78 impact. We develop methodology to correct for attenuation in dense smoke layers from wildfires
79 to properly deal with intense smoke injections into the stratosphere, with two main questions: 1)
80 does photochemical loss of wildfire smoke occur in the stratosphere, and 2) how does the AOD
81 of smoke from the wildfire studied here compare with volcanic aerosol?

82
83 The first decade of the 21st century was characterized by slower temperature evolution than
84 anticipated from CMIP5 models (Fyfe et al., 2016). The discrepancy was attributed to inter-
85 decadal Pacific oscillation (Medhaug et al., 2017), variations in solar forcing (Myhre et al., 2013)
86 and aerosol in the stratosphere from moderate volcanic eruptions (Santer et al., 2014). Should
87 wildfire smoke in the stratosphere be added to this list of phenomena that require more attention
88 in climate models?

89
90 Our investigation deals with the evolution of the wildfire AOD, and aerosol optical properties
91 obtained from the lidar CALIOP aboard the CALIPSO (Cloud-Aerosol Lidar and Infrared
92 Pathfinder Satellite Observation) satellite, OMPS-LP/Suomi and SAGE III/ISS in comparison
93 with volcanic injections to the stratosphere. Additionally, the water vapor concentrations of
94 individual smoke layers are investigated by the MLS, the spatial evolution of smoke layers is



95 investigated using OMPS-NM (Ozone Mapping and Profiler Suite Nadir Mapper), and the AODs
96 and extinction coefficients obtained from CALIOP are compared with that of OMPS-LP and
97 SAGE III/ISS.

98

99 **2. Methods**

100

101 This study of the dense stratospheric smoke layers from pyro-cumulonimbus formed over
102 Western North America in August 12 – 13, 2017 is based on five satellite sensors. For four of
103 them, OMPS-LP, SAGE III/ISS, MLS and OMPS-NM, high level products (Level 2) are used.
104 The CALIOP data evaluation is based on a Level 1 product. A method to correct for attenuation
105 of the CALIOP laser beam in the smoke layers is presented. For these reasons CALIOP requires
106 more space in this section compared to the other methods.

107

108 **2.1 CALIOP**

109

110 The evaluation of the CALIOP instrument carried by the CALIPSO satellite is based on version
111 4-10, level 1B data. CALIOP measures backscattering of laser light at two wavelengths, 532 and
112 1064 nm. For the shorter wavelength, scattered laser light is detected in parallel and
113 perpendicular polarizations relative to the outgoing beam. These almost nadir-viewing aerosol
114 and cloud measurements result in high resolution vertical profiles. For the altitude ranges <8.2,
115 8.2 – 20.2, 20.2 – 30.1 and 30.1 – 40 km the vertical resolutions are 30, 60, 180, and 300 m,
116 respectively. CALIPSO orbits between 82° S and 82° N, completing 14 – 15 orbits per day
117 (Winker et al., 2007; Winker et al., 2010).

118

119 **2.1.1 AOD**

120

121 Stratospheric AOD was obtained by integrating the backscattering intensity corrected for
122 attenuation (described below) from the tropopause to 35 km altitude. Figure 1b illustrates how
123 attenuation of the laser signal strongly reduced the signal below the dense smoke layer between
124 11 to 16 km altitude. We use the tropopause height according to MERRA-2 supplied with the
125 version 4.10 CALIOP data, which is a mixture of a dynamic and a thermal tropopause. The AOD
126 was averaged in the 20 - 80° N latitude range, where all nighttime swaths available from
127 CALIOP were included. The data were averaged over all longitudes in one-degree latitude bands,
128 and these latitude bands were averaged for the 20 - 80° N latitude range using area-weighting.
129 For dense layers, the lidar ratios estimated for the individual smoke layers were applied
130 (explained below). Apart from the first few days the lidar ratio shows no temporal evolution, it is
131 found to have geometrical mean of 48.9 sr with double-sided 95% confidence interval of 47.6 –
132 50.3 sr (Figure 2a), which is close to the typical background lidar ratio of 50 sr (Jäger and
133 Deshler, 2003). For layers that were not dense, the lidar ratio was held at this typical background
134 level. The volume depolarization ratio (δ_v) contains information that can be used to classify
135 aerosol layers. When the depolarization ratio is less than 0.05 the data is considered background
136 and the lidar ratio is set to 50 sr. Ice-clouds were removed in the lowest 3 km of the stratosphere
137 by identifying them in stratospheric layers where the backscattering was high (attenuated
138 backscattering larger than $0.0025 \text{ km}^{-1} \text{ sr}^{-1}$). Data in these layers were classified as probable
139 clouds if their δ_v was higher than 0.20, or smoke if δ_v was between 0.05-0.20, after which the data



140 within each swath were clustered depending on their location. Noise in the data led to some lone
141 pixels within layers of either ice or smoke. These were reclassified depending on the surrounding
142 pixels, making sure that no single pixel marked as aerosol occurred within the ice-cloud layers.
143 Layers of ice-clouds were then expanded upwards and horizontally to capture faint edges of the
144 clouds (Friberg et al., 2018). The classification was carried out on data at 8 km resolution along
145 each swath with their highest vertical resolution (30, 60, or 180 m, depending on altitude), after
146 which the tropospheric data were removed. Possible polar stratospheric cloud (PSC) signals north
147 of 45°N were excluded by classifying pixels with temperature below 195 K as possible PSC
148 occasions. Underlying pixels were also excluded, to prevent bias from attenuation of the lidar
149 signals or from settling ice-crystals (Friberg et al., 2018).

150

151 *2.1.2 Attenuation correction and radiative properties of individual smoke layers*

152

153 The evolution of the lidar, color and depolarization ratios were investigated using 32 separate
154 smoke layer measurements over the period 3 – 59 days after the fire. CALIOP has a statistical
155 disadvantage compared with lidars at the ground (Baars et al., 2019), because of small solid angle
156 due to long distance to the stratosphere (~700 km) and short measurement time. Optical
157 properties of old and faint individual smoke layers therefore could not be quantified with high
158 precision using CALIOP. The faint layers though still affect the AOD determinations described
159 above, where AOD elevation after the fire remains approximately one year. Out of the 32 smoke
160 layers studied, 29 were night-time measurements, whereas the remaining three are defined as
161 day-time measurements. These latter ones increased the number of early observations (day 3 – 5)
162 and were taken when the disturbance from solar radiation is small, i.e., shortly before the night.

163

164 The first weeks after the fire the smoke layers could be very dense with layer AODs exceeding 1,
165 causing strong attenuation of the CALIOP signals with two-way transmissions down to below
166 0.01. For the 532 nm wavelength the particle lidar ratio was estimated by aiming the scattering
167 ratio (R ; total-to-molecular backscattering ratio) below a smoke layer to a target value. The target
168 value was obtained from the background scattering ratio beside each smoke layer investigated,
169 which on average is $R = 1.08$, with standard deviation ± 0.05 . To reduce influence from noise, the
170 CALIOP data were averaged along the swath. The averaging range varied between the smoke
171 layers, due to its extension along the swath, the homogeneity of the layer, and avoidance of sub-
172 layer features.

173

174 The particle lidar ratio of an individual smoke layer was iterated until reaching the target value (R
175 = 1.08) described above from the combined effect of all altitude pixels. Pixels at altitudes outside
176 the smoke layer were set to the background lidar ratio of 50 sr (Jäger and Deshler, 2003). The
177 altitude resolution provided in the CALIOP data was used, where each altitude pixel (j) is
178 corrected for attenuation. The calculation starts at the highest altitude (40 km) and continues
179 downwards in two rounds. In the first round the star-marked quantities of equations 1-3 were
180 computed, correcting for attenuation from overlaying pixels. Before moving to the next altitude,
181 we account for self-attenuation from the pixel itself (equations to the right, without a star):

182



$$183 \quad \beta_j^* = \frac{\beta_j'}{\prod_{k=1}^{j-1} T_k^2}; \quad \beta_j = \frac{\beta_j^*}{\sqrt{T_j^{*2}}} \quad (1)$$

184
185 where β_j' is the attenuated backscattering and T^2 the two-way transmissions from both particles
186 and molecules. The two-way particle transmission is obtained by first computing the AOD:

$$187 \quad AOD_j^* = (\beta_j^* - \beta_{m,j})S_p\Delta z_j; \quad AOD_j = (\beta_j - \beta_{m,j})S_p\Delta z_j \quad (2)$$

189
190 where Δz_j is the height of the altitude pixel, $\beta_{m,j}$ is backscattering from air molecules, and S_p the
191 lidar ratio of the aerosol particles. The molecular lidar ratio, for computation of the molecular
192 extinction, was set to 8.70477 sr (Prata et al., 2017). CALIOP measurements are affected by
193 multiple scattering (Wandinger et al., 2010), causing overestimation of the backscattering. The
194 multiple scattering factor (η), the ratio of the apparent to the actual extinction coefficient, is not
195 known. Previous estimates are in the range 0.85 – 0.95 for layers thicker than 500 m (Prata et al.,
196 2017). Not correcting for multiple scattering results in determination of the effective lidar ratio,
197 which is lower than the actual lidar ratio by a factor η . In equation 2 thus the backscattering
198 inflated by multiple scattering is multiplied by an underestimated lidar ratio to, at least in part,
199 compensate for the effects of multiple scattering on the AOD. The two-way transmission of
200 altitude pixel j due to the particles present is obtained from:

$$201 \quad T_{p,j}^{*2} = \exp(-2AOD_j^*); \quad T_{p,j}^2 = \exp(-2AOD_j) \quad (3)$$

202
203
204 These calculations in equations 1 – 3 are carried out until the background layer between altitudes
205 a and b below the smoke layer reaches the target scattering ratio of 1.08 (Figure 3a):

$$206 \quad R = \frac{\sum_a^b \beta_j}{\sum_a^b \beta_{m,j}} \quad (4)$$

207
208
209 Error estimates of the effective lidar ratio were obtained by varying the target scattering ratio
210 from its average value ($R = 1.08$) mentioned above, to its ± 0.05 standard deviation range. The
211 fitting uncertainty in these estimates is strongly dependent on the light extinction in the smoke
212 layer. Dense layers result in very small uncertainties in the effective lidar ratio because of the
213 strong impact on R from a slight change in the extinction. Layers with lower extinction
214 progressively increase the uncertainties of the estimate. When the error estimate of the effective
215 lidar ratio fit exceeds 25% the result is excluded from the data analysis, which terminates
216 estimates of lidar ratios from day 22 after the fire.

217
218 The color ratio, the ratio between the backscattering at 1064 nm to 532 nm wavelength, is
219 affected by a difference in attenuation of the two wavelengths. This is clearly visible for dense
220 smoke layers in the CALIOP browse images by a gradual increase of the color ratio through the
221 layer because of the weaker attenuation for 1064 nm wavelength than for 532 nm (Figure 1d).



222 Therefore, estimations of the attenuation were undertaken also for the long wavelength. The
223 molecular backscattering is assumed to be 1/16 of that at 532 nm ($1/\lambda^4$ dependence of Rayleigh
224 scattering). Weak molecular scattering at 1064 nm prohibits lidar ratio estimation at that
225 wavelength by CALIOP. Instead, the lidar ratio was assumed to be 60 sr, inducing uncertainties
226 in the color ratio. The volume color ratio is obtained from:

$$227 \chi = \sum_{k=top}^{base} \beta_{1064,k} / \sum_{k=top}^{base} \beta_{532,k} \quad (5)$$

229 To limit influence from attenuation in the color ratio computations, the estimates were based on
230 the upper part of a smoke layer. Starting from the top of the smoke layer, the computations were
231 truncated when the two-way transmission of the 532 nm wavelength fell below 0.7. Varying the
232 1064 nm wavelength lidar ratio in the wide range of 60 ± 20 sr the uncertainty in the color ratio
233 becomes less than $\pm 5\%$ with this constraint applied. From the color ratio we define the particle
234 color ratio:

$$235 \chi_p = \sum_{k=top}^{base} (\beta_{1064,k} - \beta_{m,1064,k}) / \sum_{k=top}^{base} (\beta_{532,k} - \beta_{m,532,k}) = \frac{\chi R}{R-1} - \frac{1}{16(R-1)} \quad (6)$$

238 where we made use of the wavelength dependence of Rayleigh scattering for molecular
239 scattering, and the scattering ratio for the 532 nm wavelength was obtained from eqn. 4.

241 We also investigated the depolarization of the scattered laser beam at 532 nm by first forming the
242 volume depolarization ratio:

$$243 \delta_v = \sum_{k=top}^{base} \beta'_{532\perp,k} / \sum_{k=top}^{base} \beta'_{532,k} \quad (7)$$

246 where symbol \perp indicates scattered light polarized perpendicularly to the incident beam. Having
247 access to the volume depolarization and an estimate of the molecular depolarization ratio $\delta_m \approx$
248 0.003656 (Prata et al., 2017; Hostetler et al., 2006) the particle depolarization ratio is obtained
249 from:

$$250 \delta_p = \frac{\delta_v - \delta_m + \delta_v(1 + \delta_m)(R-1)}{\delta_m - \delta_v + (1 + \delta_m)(R-1)} \quad (8)$$

253 where R is obtained from eqn. 4.

254 **2.2 Extinction coefficients and AOD from OMPS-LP**

255
256 The aerosol data from OMPS-LP (Chen et al., 2018; Jaross et al., 2014; Loughman et al., 2018)
257 have lately been used extensively in the literature on volcanic and wildfire impact on the
258 stratospheric aerosol. Several data products are available, here we use the recently released Level
259 2 product: Suomi-NPP OMPS LP L2 AER Daily Product, version 2.0 (Taha et al., 2020). The
260 polar-orbiting Suomi satellite completes 14 - 15 laps per day. OMPS-LP is a limb-scattering
261
262



263 method that collects data looking backwards along the satellite orbit, and along two other
264 directions separated by 4.25° from the orbit, giving a cross-track separation of approximately 250
265 km at the tangent point. Measurements are undertaken in the wavelength and altitude ranges of
266 290 – 1000 nm and 10 – 80 km, respectively. The vertical resolution of OMPS-LP is 1.5 – 2 km
267 (Rault and Loughman, 2013). The aerosol product used here comprises 6 wavelengths (510, 500,
268 675, 745, 869 and 997 nm). The group responsible for the OMPS-LP version 2.0 data (Taha et
269 al., 2020) recommends caution when using data from altitudes below 17 km altitude due to loss
270 of sensitivity. This problem can be reduced by use of the 745 nm and longer wavelengths. Here
271 we will make use of two of wavelengths: 745 nm because of the reduced problem with
272 sensitivity, and 510 nm because it is the wavelength closest to that of CALIOP (532 nm).

273
274 The OMPS-LP aerosol extinction coefficients are provided on a grid with a vertical resolution of
275 1 km. To study the smoke from the August 2017 fire we compute the average AOD over all
276 longitudes in the latitude interval $20 - 80^\circ$ N for three layers, the LMS (tropopause to 380 K
277 isentrope), lower Brewer-Dobson branch (380 – 470 K) and the upper Brewer-Dobson branch
278 (470 K to 35 km altitude). The OMPS-LP version 2 dataset use a cloud detection algorithm (Chen
279 et al., 2016), and comes in two forms: one without filtering out signals from clouds, and the other
280 where signals affected by clouds and polar stratospheric clouds are removed. In Figure 4 we
281 show both these varieties for 745 nm wavelength, and, with and without flags regarding data
282 quality including profile retrieval errors (named RetrievalFlags in the OMPS-LP files), high root-
283 mean squares (ResidualFlags), and further errors from the South Atlantic anomaly, disturbances
284 from the Moon, solar eclipses, planets, and satellite maneuvers (SwathLevelQualityFlags). In the
285 two upper layers (Figures 4a and b) the differences are usually small between the varieties except
286 for some spikes, whereas the LMS data (Figure 4c) show large stochastic variability as well as
287 periods of clear differences between the varieties. Since this data is taken well below 17 km
288 altitude, sensitivity issues can be expected (Taha et al., 2020), see above. Days 130 – 190 (during
289 December 2017 to February 2018) several spikes appear in the two higher layers which likely are
290 caused by polar stratospheric clouds. The data set filtered for clouds and flagged stands out by
291 comparably small peaks, whereas the differences between the varieties usually are small
292 elsewhere. We therefore select the cloud-filtered and flagged data for further analysis in the
293 coming sections.

294

295 **2.3 Extinction coefficients from SAGE III/ISS**

296

297 SAGE III/ISS is a limb-viewing instrument based on solar occultation. Here we make use of
298 Level 2 aerosol extinction coefficients (SAGE III/ISS User's Guide, 2018), version 5.10,
299 supplied with a vertical resolution of 0.5 km. The upper limit of the slant path optical depth is
300 about 8, translating to a vertical optical depth of approximately 0.02 (SAGE III/ISS User's Guide,
301 2018). The orbiting of ISS differs markedly from the polar orbiting satellites CALIPSO
302 (CALIOP) and Soumi (OMPS-LP). This causes sporadic coverage by ISS of the latitudes of
303 interest here, resulting in that no average AODs over the $20 - 80^\circ$ N latitude range could be
304 formed with adequate time resolution. However, daily maximum extinction coefficients from
305 SAGE III/ISS could, when available, be included in a comparison with CALIOP and OMPS-LP.



306

307 **2.4 Water vapor measurements from MLS**

308

309 Water vapor concentrations (mixing ratio) in individual smoke layers was obtained from the MLS
310 instrument aboard the Aura satellite (Waters et al., 2006) in 12 vertical steps per decade of
311 pressure (version 5.0-1.0a, level 2). In nighttime measurements from days 6 – 59 after the fire,
312 the smoke layers studied by CALIOP were also investigated with MLS in almost simultaneous
313 measurements, both instruments being on satellites that are members of the A-train. Data in the
314 10 – 316 hPa atmospheric pressure range were used, with vertical resolution 1.3 – 3.2 km
315 (Livesley et al., 2020). Limited vertical resolution induces problems to obtain well defined
316 observation of H₂O concentration of smoke layers close to the strong H₂O concentration gradient
317 across the tropopause. Close to the tropopause, but in the stratosphere, no H₂O peak from a
318 smoke layer can be detected. As the distance to the tropopause increases, an H₂O peak from the
319 smoke layer becomes discernible. Further up from the tropopause, when the peak H₂O
320 concentration is well above the extratropical tropopause at atmospheric pressure of less than 110
321 hPa, a deep minimum appears between the tropopause gradient and the peak from the smoke
322 layer. All H₂O peaks were fitted with a Gaussian distribution operating on logarithmic pressure
323 and H₂O concentration to obtain estimates of the peak concentration and the corresponding
324 atmospheric pressure. To investigate a time dependence in the smoke layer composition the peak
325 H₂O concentration ($C_{\text{H}_2\text{O}}$) was compared with the attenuation-corrected aerosol scattering ratio
326 (R) from CALIOP, the optical equivalent of the mixing ratio, where the latter was obtained by
327 forming the geometrical mean over 900 m around the peak scattering ratio. The ratio of the of the
328 two quantities ($R/C_{\text{H}_2\text{O}}$) was formed, and its dependence on time from the fire was studied. Out of
329 the 13 smoke layers available with peak water vapor concentrations above the altitude of 110 hPa
330 atmospheric pressure, one was flagged as low quality in the MLS data set, leaving 12
331 observations for the study of the $R/C_{\text{H}_2\text{O}}$ evolution.

332

333 **2.5 UV aerosol index from OMPS-NM**

334

335 The UV aerosol index of OMPS-NM based on measurements at two wavelengths, 340 and 378.5
336 nm, is the official NASA aerosol index product according to OMPS-NM (NMMIEAI-L2 V2.1.1)
337 release notes (Torres, 2019). For strongly UV absorbing aerosols, like black carbon from
338 wildfires, the UV aerosol indexes strongly increases with altitude (Herman et al., 1997). Here the
339 OMPS-NM UV aerosol index was used to map the geographical evolution of the smoke layers,
340 that according to CALIOP measurements were distributed in both the troposphere and the
341 stratosphere.

342

343 **3. Results**

344

345 Here we use an approach based on five satellite sensors to study the influence on the stratosphere
346 of the great North American fire in August 2017. We start by briefly describing results from the
347 method to correct CALIOP data for attenuation of the backscattered laser light. Then follows a
348 comparison of AODs obtained from OMPS-LP and CALIOP. Absorption aerosol index from



349 OMPS-NM is used to describe the dispersion of the wildfire aerosol in the stratosphere. To
350 explain differences in AOD between OMPS-LP and CALIOP, a comparison of extinction
351 coefficients follows, where results from SAGE III/ISS also are included in the comparison. The
352 evolution of the optical properties of the wildfire aerosol is then described, before the North
353 American wildfire aerosol is compared with volcanic influence on the stratospheric AOD.
354 Finally, the fifth data set, water vapor from the MLS, is introduced in the discussion section,
355 where the evolution of the wildfire aerosol in the stratosphere is analyzed.

356

357 *3.1 Correction for attenuation*

358

359 The smoke layers usually were 1 – 3 km thick and could extend several degrees in longitude and
360 latitude. Measurements with the CALIOP lidar provide, in addition to short, nadir-viewing
361 measurement path in dense layers, the advantage that the signal is retrieved as a function of
362 position along the laser path with high resolution, which can be used to correct for attenuation of
363 the signal. Figure 3a shows the attenuated scattering ratio (R' ; the measured backscattering
364 divided by the calculated molecular backscattering) from an example-smoke-layer measured on
365 August 16, 2017. The scattering ratio should be close to 1 in air layers with low aerosol
366 concentration, whereas values below 1 is caused by attenuation from particles. As can be seen in
367 Figure 3a, the attenuated scattering ratio first increases (starting from above the layer). Then the
368 signal decreases and reaches well below unity from 11 km altitude and downwards, i.e., well
369 below the scattering ratio of particle-free air. By techniques described in the Methods section we
370 correct for attenuation and fit the lidar ratio (Figure 2a) to obtain an estimate of the
371 backscattering without attenuation, as illustrated by the scattering ratio (R) in Figure 3a.

372

373 The evolution of wildfire aerosol from day 3 to 59 after the North American PyroCbs on August
374 12, 2017, is first investigated by comparing 32 smoke layers from individual CALIOP swaths.
375 The influence from attenuation is shown in Figure 3b. Clear deviation from the 1:1 line appears
376 already at layer attenuated (uncorrected) AODs (AOD_{att}) of 0.12, and 50% reduction of the signal
377 appears at layer AOD_{att} of approximately 0.25. Reduction by more than 50% appears until day 10
378 after the fire, whereas those measurements close to the 1:1 line were taken after day 30. The
379 AOD, i.e., the AOD corrected for attenuation, exceeds the AOD_{att} by more than a factor of 5 in
380 the densest layers of this study (Figure 3b).

381

382 *3.2 Comparison of CALIOP and OMPS-LP*

383

384 To study the evolution of the stratospheric AOD, we form a 3-dimensional box in the
385 stratosphere extending over all longitudes in the 20 – 80° N latitude range. In this box we use all
386 daily profiles, 14 – 15 CALIOP and 42 – 45 OMPS-LP, to form the average AOD. We apply the
387 method to correct CALIOP data for attenuation, as described in the Methods section. AODs are
388 computed for three layers, the LMS, the lower Brewer-Dobson branch, and the upper Brewer-
389 Dobson branch, as shown in Figure 5.

390



391 When comparing AODs, the measurement wavelengths should be as close as possible, due to the
392 wavelength dependence of scattering. CALIOP AODs are shown for 532 nm wavelength, and the
393 OMPS-LP data are shown for the close wavelength of 510 nm. In addition, the 745 nm AODs
394 from OMPS-LP is shown. The response to the 2017 North American fire is weak in the upper
395 Brewer-Dobson branch (Figure 5a), whereas the two lower layers (Figures 5b-c) show clear
396 increase of the AOD. Comparing the two methods, they agree well in the upper Brewer-Dobson
397 branch. In the lower Brewer-Dobson branch we see good agreement between the two methods,
398 except for the first 1 – 2 months after the fire where much higher AODs are recorded by CALIOP
399 (Figure 5b). The latter is also true for the LMS, whereas the general agreement between the two
400 methods is poor (Figure 5c). The OMPS-LP documentation advise against using data from below
401 approximately 17 km altitude, approximately the upper limit of the LMS, due to loss of
402 sensitivity (Taha et al., 2020). We therefore do not perform any further comparisons in the LMS.
403 The stratosphere above the LMS (above the 380 K isentrope) shows good agreement between the
404 two methods, except for the first 1 – 2 months after the fire (Figure 5d).

405

406 **3.3 Early evolution of the smoke layers**

407

408 The daily AOD averages show large variability the first days after the fire because the lidar
409 measures narrow curtains through the atmosphere, Figure 5e. The variability remains until the
410 smoke layers become sufficiently dispersed, allowing several daily measurements of the smoke
411 layers. The nadir-viewing OMPS-NM provides UV (ultraviolet) absorbing aerosol index, where
412 strong signal for strongly UV light absorbing aerosol is obtained in the upper troposphere and the
413 stratosphere. Figure 6 shows the geographic evolution of the smoke layers from August 14 to 22,
414 2017 together with the orbits followed by the CALIOP measurements. Up to August 16 the
415 smoke is found in a rather confined area. From August 17 the smoke layers are stretched in
416 Easterly direction, and after that the smoke spreads rapidly to the East. The dispersion gradually
417 increases the number of daily CALIOP observations of the smoke. This can also be seen in
418 Figure 5e, where the variability in the daily AOD data becomes successively smaller. From day
419 10 (August 22) we see a clear pattern of decline of the AOD.

420

421 Figure 5e shows the total stratospheric AOD according to CALIOP from the tropopause to 35 km
422 altitude. We see a strong decline of the stratospheric AOD the first 1.5 months after the fire, and a
423 fitted exponential function has a half-life of 6.5 ± 0.9 days. Such a decline cannot be found in the
424 OMPS-LP AODs, which instead are increasing during the first month.

425

426 To further investigate this clear difference between the two methods, individual smoke layers are
427 investigated with respect to extinction coefficients. Figure 7a-d show the extinction coefficient of
428 strong smoke layers from four days in August and September 2017. From CALIOP we show the
429 attenuated extinction coefficients as well as the profiles corrected for attenuation. Together with
430 the CALIOP data the OMPS-LP data closest by are shown. It is obvious that OMPS-LP shows
431 very much smaller reaction to the smoke layers than CALIOP. However, we cannot be sure that
432 the two instruments viewed the same airmasses in these four examples, because the two
433 instruments do not belong to the same satellite constellation. To remove that obstacle, the daily



434 maximum stratospheric extinction coefficient from OMPS-LP was extracted and compared with
435 32 selected profiles' peak extinction coefficients from CALIOP. SAGE III/ISS was also included
436 in the comparison from day 19 after the fire. Unfortunately, the orbiting of ISS did not permit
437 measurements of the fire studied here before that day. The very strong signals from CALIOP are
438 not reflected in the OMPS-LP or SAGE III/ISS measurements, see Figure 7e. In part, this can be
439 explained by difference in vertical resolution, but as shown in Figures 7a-d, these high extinction
440 coefficients extend to broad vertical ranges that should allow detection of strong signals also by
441 OMPS-LP and SAGE III/ISS.

442

443 There is one principal difference between CALIOP on one hand and OMPS-LP and SAGE
444 III/ISS on the other hand: whereas the former is nadir-viewing (vertical) the latter two methods
445 operate in limb orientation (horizontal). This is important, because the horizontal extension of
446 smoke layers is much larger, e.g., the smoke layer in Figure 1b has a vertical extension of
447 approximately 2 km, whereas the horizontal extension is approximately 700 km. The vertical,
448 two-way transmission to the CALIOP sensor through this layer is approximately 0.01, which we
449 correct for. The horizontal path through this layer is 350 times longer, implying that the one-way
450 limb transmission becomes 10^{-350} for the same wavelength. Even if the horizontal extension
451 would be just one tenth the transmission is still as low as 10^{-35} . Obviously, the radiation used for
452 detection in OMPS-LP and SAGE III/ISS is rapidly eliminated in such smoke layers. Therefore,
453 these two methods are inadequate for studies of dense aerosol layers. The upper limit in terms of
454 vertical AOD is estimated to 0.02 (SAGE III/ISS Users Guide, 2018), corresponding to the
455 extinction coefficient of 0.02 km^{-1} for a 1 km thick layer. This problem is also acknowledged for
456 OMPS-LP (Chen et al., 2018; DeLand, 2019). Despite the clear limitation of OMPS-LP and
457 SAGE III/ISS in this respect, the large body of information on wildfires is based on these
458 methods, e.g., Bourassa et al., (2019), Das et al., (2021), Khaykin et al., (2020), Kloss et al.,
459 (2019), Torres et al., (2020) and Yu et al., (2019). By comparing with CALIOP we here show
460 that the limb-oriented techniques miss the dramatic events during the first 1 – 2 months after the
461 fire. The rapid decline of the wildfire smoke will be further analyzed below.

462

463 *3.4 Aerosol optical properties*

464

465 To further investigate the unusual evolution of the AOD, we turn to the optical properties of the
466 wildfire aerosol. The particle color and depolarization ratios are shown in Figure 2b and c. To test
467 the significance in the evolution the data were temporally divided into two equal halves by
468 number of data points, and geometric averages were formed (black lines in Figure 2). The particle
469 color ratio shows a highly significant decrease comparing the first to the last half of the data
470 points, whereas the particle depolarization ratio increases with high significance. The change in
471 the optical properties takes place up to 15 – 30 days after the fire. This coincides with the decline
472 of the AOD, thus connecting a change of the aerosol properties to the AOD decline.

473

474 *3.5 Stratospheric AOD variability caused by volcanism and wildfires*

475



476 The stratospheric AOD varies considerably over time mainly due to influence from explosive
477 volcanic eruptions as demonstrated in Figure 8, showing the period 2008 – 2018. In this time
478 span, nine volcanic eruptions clearly, but to varying degree, affected the stratospheric AOD. We
479 also identify two cases of influence from wildfires, the Victoria fire (Australia, 2009) and the fire
480 studied here (Western North America, 2017). The residence time in the stratosphere varies from
481 several years for tropical injections into the upper layer representing the upper branch in the
482 Brewer-Dobson circulation (BD) (Figure 8a), the order of a year in the shallow branch of the BD
483 circulation (Figure 8b), to months in the LMS (Figure 8c) (Friberg et al., 2018). The sum of the
484 three layers is shown in Figure 8d. The volcanic eruptions in these 11 years mainly affected the
485 two lower stratospheric layers, only the Kelut eruption (2014) clearly reached to the deep BD
486 branch. Fire aerosol contains black carbon, which absorbs radiation, heats surrounding air and
487 induces lifting, as observed after the fire studied here (Khaykin et al., 2018; Yu et al., 2019).
488 After both fires, we see weak AOD elevation in the deep BD branch (Figure 8a), but for the fire
489 studied here the two lower layers dominate the AOD, like most of the volcanic eruptions in the
490 eleven-year period.

491

492 Comparing the evolution of the AOD of the North American wildfire with the evolution of the
493 aerosol from two of the most important volcanic eruptions during the last 25 years (Figure 9), we
494 find that the maximum stratospheric AOD after the fire is similar to that after the 2011 Nabro and
495 2009 Sarychev eruptions. During the first couple of months after volcanic events the AOD grows
496 due to formation of condensable sulfuric acid from the emitted volcanic gas sulfur dioxide. In
497 contrast, the wildfire aerosol displays a rapid decline during the first few weeks, before the AOD
498 stabilizes (Figure 9). This is followed by a period of rather stable AOD of more than 6 months,
499 before the AOD evolution turns to a slower decline towards background conditions, with similar
500 seasonality as the aerosol from the volcanic eruptions discussed (Figure 9). This latter decline is
501 mainly caused by springtime transport out from the stratosphere at mid and high latitudes
502 (Bönisch et al., 2009; Martinsson et al., 2017).

503

504 **4. Discussion**

505

506 The smoke aerosol is distributed both in the LMS and in the lower BD branch like aerosol from
507 several volcanic eruptions (Figure 8). The rapid decline of the smoke aerosol during the first
508 month after the fire thus cannot be explained by transport out of the stratosphere. Measurements
509 with Raman lidars at three wavelengths indicate that the smoke from this North American fire
510 contain an accumulation mode but no coarse mode (Haarig et al., 2018; Hu et al., 2019). The
511 influence from sedimentation on submicron diameter particles is small (Martinsson et al., 2005).
512 Moreover, the change in the particle depolarization ratio (Figure 2c) indicates change of the
513 aerosol particle properties, and the particle color ratio decrease after the fire (Figure 2b) is the
514 expected outcome for reduced particle sizes. Based on these arguments we turn the attention to
515 loss of material from the aerosol particles to the gas phase to explain the rapid decrease in AOD
516 seen in Figure 5e.

517



518 Smoke layers contain water vapor that could induce hygroscopic growth/shrinkage. Water vapor
519 profiles for individual smoke layers from days 6 – 60 after the fire were obtained from the MLS.
520 Measurements close to the tropopause (Figure 10a) are affected by a steep gradient in H₂O
521 concentration. The profiles well above the gradient peaking at atmospheric pressure of less than
522 110 hPa are shown in Figure 10b. For the latter category the peak H₂O concentration is in the
523 range 7 – 14 ppmv, implying a maximum H₂O vapor pressure of 0.16 Pa. For typical conditions
524 in the extratropics that vapor pressure corresponds to a relative humidity of a few percent or less
525 (Murphy and Koop, 2005).
526

527 To further investigate the smoke layers, the temporal evolution of the composition is studied by
528 forming the ratio of the mixing ratios of two components: aerosol backscattering and H₂O at the
529 peak of respective vertical distribution. As pointed out above, the strong H₂O gradient around the
530 tropopause affects the MLS measurements. But for the smoke layers higher up, peaking above
531 110 hPa, we find a rapid decrease in the aerosol scattering ratio compared with the H₂O
532 concentration (Figure 10c). Fitting an exponential function ($\frac{R}{C_{H2O}} = a + be^{-t/\tau}$), the half-life
533 becomes 9.7±3.2 days, which is somewhat longer than that computed from the AOD (half-life
534 6.5±0.9 days). The rapid AOD decline (Figure 5e) is thus verified by relative concentrations of
535 aerosol and H₂O under well-controlled humidity conditions, whereas the low relative humidity
536 rules out hygroscopic growth and influence from clouds as the explanation of the AOD decline.
537

538 The near-field wildfire aerosol contains, besides black carbon (Bond et al., 2013; Ditas et al.,
539 2018), approximately 90% organic material (Garofalo et al., 2019). After emission, secondary
540 organic aerosol (SOA) is formed by oxidation of gas phase compounds (Shrivastava et al., 2017).
541 Knowledge of processes controlling formation and removal in the atmosphere is limited (Hodzic
542 et al., 2016). Global aerosol models usually remove SOA mainly by wet (90%) and, to a smaller
543 extent, by dry deposition (Tsigaridis et al., 2014). In contrast to the species dominating the
544 stratospheric aerosol and its precursor compounds during background conditions and volcanic
545 influence (sulfuric acid and sulfur dioxide), organic species are not the ultimate
546 thermodynamically stable compounds (Hallquist et al., 2009). Organic aerosol is an intermediate
547 state on routes, with little known rates, from emitted compositions to the highly oxidized gaseous
548 products CO and CO₂ (Jimenez et al., 2009). Modeling and numerous laboratory studies find
549 evidence for photolytic removal rates of organic aerosol similar to that of wet deposition in the
550 troposphere (Hodzic et al., 2016; Zawadowics et al., 2020). Recently, photolytic removal of
551 particulate SOA was included in the Whole Atmosphere Community Climate Model (WACCM6)
552 (Gettelman et al., 2019). Hodzic et al. (2015) estimate the photolytic loss over a 10-day period to
553 50% for most organic species at mid tropospheric conditions.
554

555 These high rates are disputed by Yu et al. (2019), claiming a lifetime of 150 days (half-life 104
556 days) of organic aerosol from the fire studied here, whereas Das et al. (2021) explain a similar
557 half-life of the same fire by large-scale circulation and particle sedimentation using OMPS-LP
558 and modeling. The experimental data used here cannot differentiate these two explanations,
559 although the slow part of the smoke decline is similar in seasonality to that of volcanic aerosol
560 (Figure 9) where photochemical loss is less important. The modeling study by Yu et al. (2019)



561 was based on mimicking the extinction according to SAGE III/ISS at 1020 nm wavelength at 18
562 km altitude. For three reasons their study misses the strong decline of the AOD during the first
563 month. Firstly, because the orbiting of ISS prohibits studies of the wildfire smoke the first 19
564 days after the fire, secondly because of the time required to transport the wildfire aerosol to 18
565 km altitude is approximately one month (Yu et al., 2019) and thirdly because problems with
566 event termination (“saturation”), see Figure 7e. We therefore conclude that that Yu et al. (2019)
567 could not observe the main decline of the aerosol taking place during the first 1 – 2 months after
568 the fire, see section 3.3 for further details.

569

570 Submicron aerosol particles have much longer residence time in the stratosphere than in the
571 troposphere due to sparsity of clouds, thus inhibiting the sink that traditionally is considered the
572 most important in the troposphere, i.e., wet deposition. This provides unique possibilities to study
573 photolytic loss without competition from other aerosol sinks. Interpreting the body of evidence
574 on the strong and rapid decline of the stratospheric AOD during the first month after the fire, we
575 find that photolytic loss of organic aerosol is a highly likely explanation. The rate of photolytic
576 loss is likely better described by the evolution of R/C_{H_2O} than by the AOD, because the latter
577 could to some degree be affected by transport across the tropopause. Our strong experimental
578 evidence leads us to the hypothesis that the rapid decline of the wildfire aerosol in the
579 stratosphere with a half-life of 10 days is caused by photochemical loss of organic material. This
580 should be further investigated by modeling, but that is outside the scope of the present study.

581

582 To further put the strong early decline of wildfire aerosol into context, we compare the AOD
583 during background conditions (years 2013 and 2014) with the year of the fire. When the
584 contribution of the exponential term is very small of the wildfire aerosol (after 7 half-lives), the
585 background is approximately 2/3 of the wildfire AOD (Figure 9). Taking the background into
586 account, the excess stratospheric aerosol due to the wildfire declines by 83% from the R/C_{H_2O}
587 value day 10 after the fire. The process starts before day 10, indicating that almost all the organic
588 aerosol constituting approximately 90% of the near-field wildfire aerosol mass (Garofalo et al.,
589 2019) could be lost by photolysis. Residual wildfire aerosol particles, likely stripped off by a
590 large fraction of its original organic content, remain in the stratosphere up to approximately one
591 year (Figure 9).

592

593 Finally, we investigate the stratospheric aerosol load from the wildfire by comparing with the
594 more studied volcanic impact (Table 1). The AOD growth, the average AOD over one year from
595 the fire/eruption subtracted by the average background AOD (2013 – 2014), is approximately 1/4
596 and 1/3 of that of two of the most important volcanic eruptions for the stratospheric aerosol in the
597 last 25 years (Sarychev 2009, Nabro 2011). The average excess aerosol during the year following
598 the fire corresponds to a radiative forcing of -0.06 W m^{-2} in the region 20 - 80° N, using standard
599 conversion as an approximation (Solomon et al., 2011).

600

601 **Conclusions**

602



603 In this study we investigate massive injections of smoke into the stratosphere from the August
604 2017 North American wildfires using five satellite sensors. Methodology was developed to
605 correct CALIOP data for attenuation of the laser signal. The CALIOP AOD and extinction
606 coefficients were compared with OMPS-LP and SAGE III/ISS. From 1 – 2 months after the fire
607 we find that OMPS-LP and CALIOP AOD agree very well at altitudes above the 380 K
608 isentrope, where the former demonstrates high sensitivity with small statistical fluctuations. The
609 methods differ dramatically during the first 1 – 2 months after the fire when the smoke layers are
610 dense, because the long optical path through the smoke of the limb-oriented instruments OMPS-
611 LP and SAGE III/ISS cause event termination (“saturation”). This is clearly demonstrated by the
612 low daily maximum extinction coefficients of the two instruments, being orders of magnitude
613 lower than the peak extinction coefficients of CALIOP. The nadir viewing CALIOP experiences
614 a much shorter optical path, because the vertical extension of smoke layers usually are orders of
615 magnitude shorter than for limb orientation. We find that CALIOP is an indispensable tool for
616 studies of dense smoke layers entering the stratosphere after intense wildfires, providing signal
617 along the laser path that can be used to correct for attenuation. Once the smoke layers are
618 sufficiently thin, the limb technique OMPS-LP provide sensitive measurements of the AOD that
619 can be used together with CALIOP.

620

621 The AOD from the wildfire declines exponentially with a half-life of 6.5 days. This decline is
622 further studied by the evolution of the ratio of the aerosol and water vapor mixing ratios of the
623 smoke layers, resulting in a massive decline of 80 – 90% of the wildfire aerosol with a half-life of
624 approximately 10 days. We find transport out of the stratosphere, sedimentation, influence from
625 clouds or hygroscopic growth/shrinkage to be highly unlikely explanations for the rapid decline
626 of wildfire aerosol in the stratosphere. Based on strong experimental evidence we hypothesize
627 that photochemical loss of organic aerosol causes the rapid decline, which would mean that
628 almost the entire organic fraction of the wildfire aerosol would be lost in the exponential decline.
629 The half-life according to this study agrees well with results from laboratory studies and global
630 modeling. Our unique result could be obtained because of the long residence time of aerosol
631 particles in the stratosphere, whereas tropospheric studies of photochemical loss are extremely
632 difficult because it is masked by SOA formation and wet deposition due to short residence time.
633 The residual aerosol leaves the stratosphere within a year in the Brewer-Dobson circulation.
634 Despite the initial loss, the long-term effects of wildfire smoke on the stratospheric AOD and
635 radiative forcing are considerable. The ongoing climate change is projected to increase the
636 frequency of wildfires, prompting the need for inclusion of wildfire impact on the stratospheric
637 aerosol load in the climate models.

638

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647

648 **Author Contributions**

649

650 B.G.M. designed the study, designed methodology, undertook part of the data analysis, and wrote
651 most of the paper. J.F. contributed to the design of the study, designed methodology, did part of
652 the data analysis, and wrote parts of the text. O.S.S. contributed to the data analysis and M.K.S.
653 contributed to the design of methodology. In addition, all authors participated in discussions and
654 commented on the manuscript.

655

656 **Data availability**

657

658 CALIOP V4.10 lidar data (<https://search.earthdata.nasa.gov/search?fp=CALIPSO>) are publicly
659 available.

660 OMPS-NM UV aerosol index was obtained from the publicly available site

661 <https://worldview.earthdata.nasa.gov/>.

662 OMPS-LP stratospheric aerosol optical depths were obtained from

663 https://disc.gsfc.nasa.gov/datasets/OMPS_NPP_LP_L2_AER_DAILY_2/summary

664 MLS water vapor concentrations were obtained from

665 https://disc.gsfc.nasa.gov/datasets?page=1&keywords=ML2H2O_005

666 SAGE III/ISS aerosol data were obtained from

667 https://asdc.larc.nasa.gov/project/SAGE%20III-ISS/g3bssp_51.

668

669 **Competing Interest**

670

671 The authors declare no competing interests.

672

673 **Additional Information**

674

675 Correspondence and requests for materials should be addressed to B.G.M.

676

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959 **Tables**

960

961 Table 1. Maximum and yearly average stratospheric AOD during background conditions and
962 during one year after the fire and after the two volcanic eruptions in Figure 9.

Year	Background 2013	Background 2014	Wildfire 2017	Sarychev 2009	Nabro 2011
AOD max	0.009	0.009	0.020	0.028	0.017
AOD	0.0075	0.0074	0.0097	0.0169	0.0138
AOD growth ^a	-	-	0.0023	0.0095	0.0064
RF ^b	-	-	-0.06	-0.24	-0.16

963 ^aGrowth of AOD due to influence from wildfire/volcanism obtained by subtracting the average of 2013
964 and 2014 AOD.

965 ^bRadiative forcing ($W m^{-2}$) of the background-subtracted AOD.

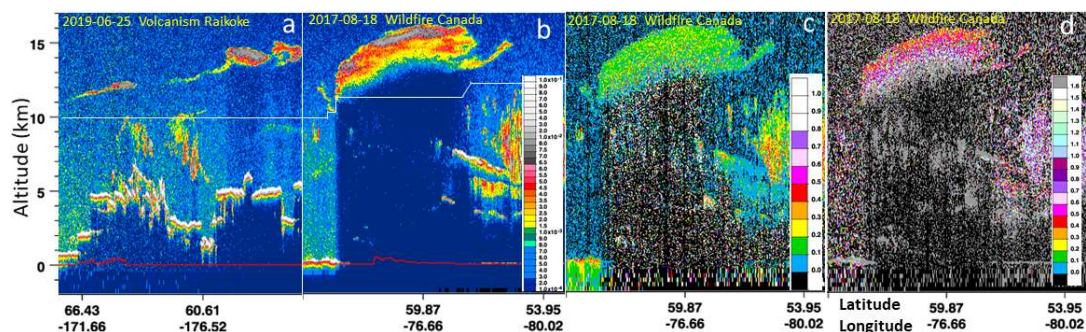
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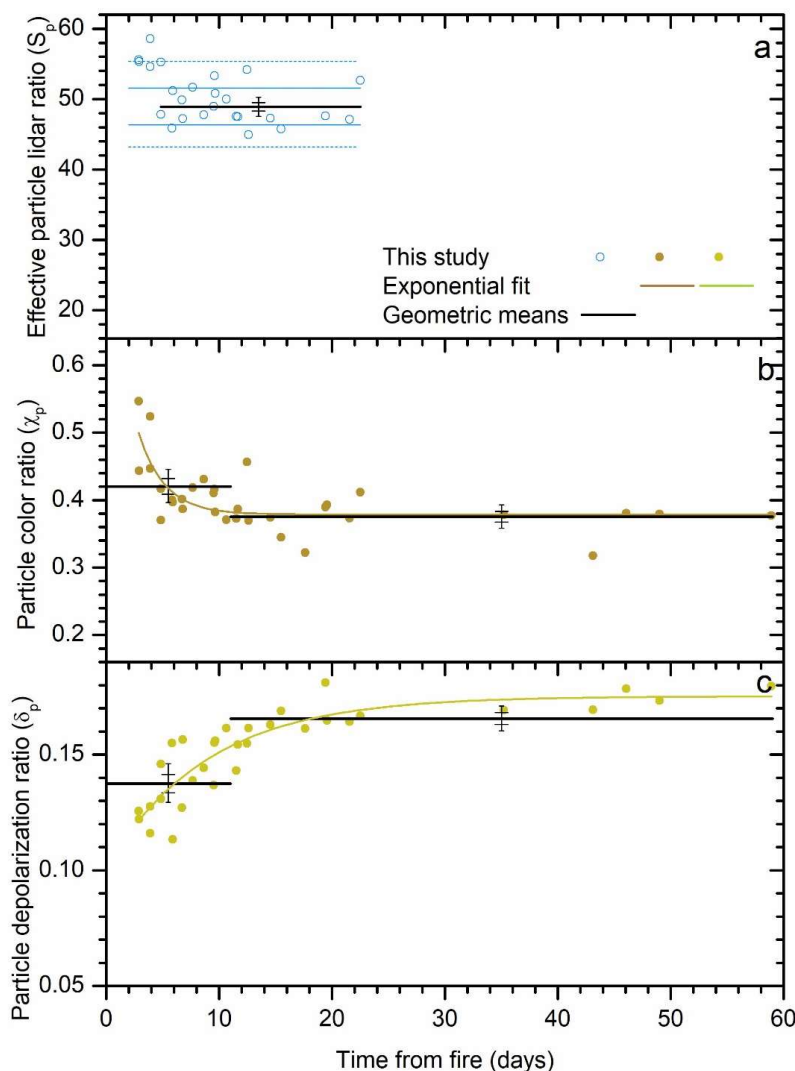
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Figures



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Figure 1. CALIOP curtains of total attenuated backscatter ($\text{km}^{-1} \text{sr}^{-1}$) at 532 nm from a) volcanic aerosol layers in the stratosphere three days after the 2019 Raikoke eruption and b) a stratospheric smoke layer from the August 12, 2017, North American wildfire. c) Volume depolarization ratio at 532 nm and d) attenuated color ratio (1064 to 532 nm) for the curtain in b). The white lines in a) and b) show the position of the tropopause.

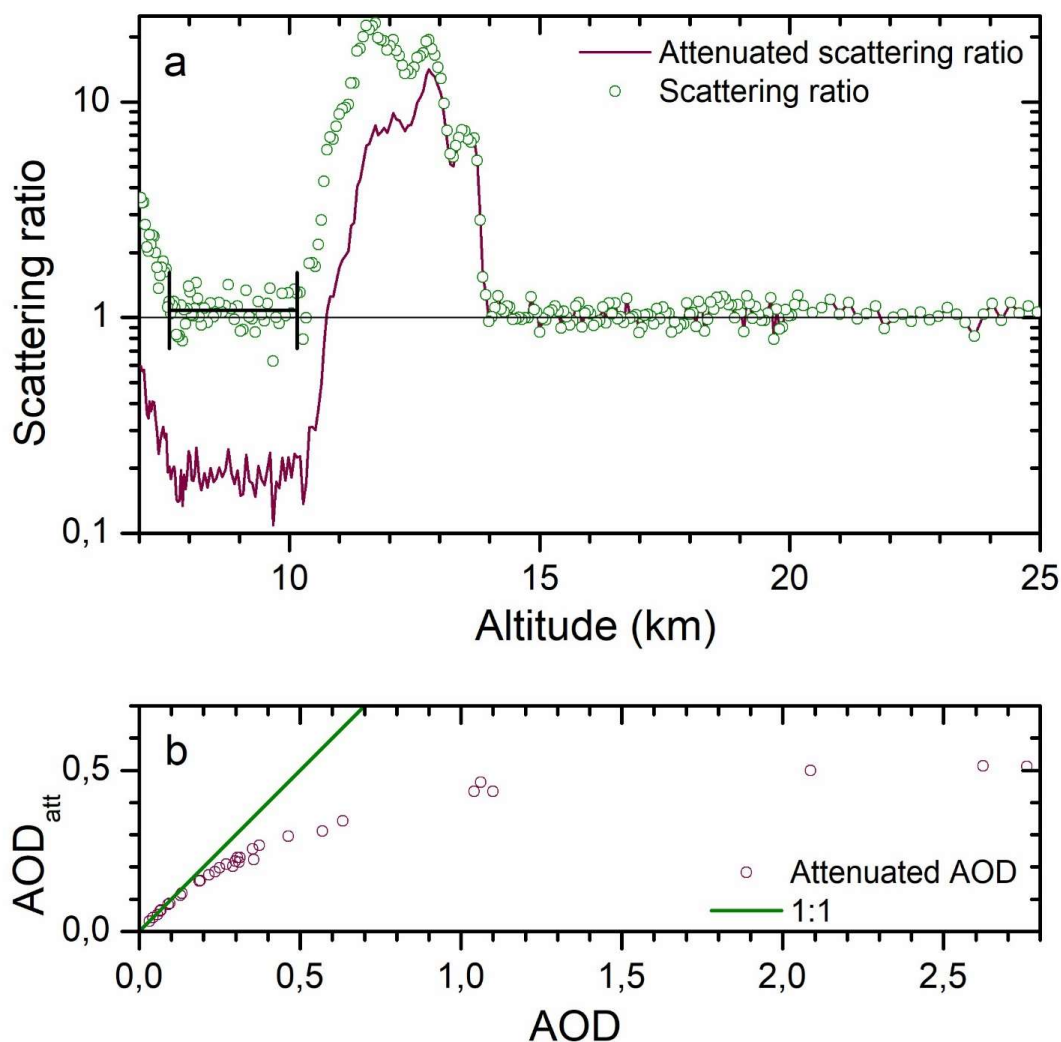


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980 Figure 2. Particle optical properties during the first 60 days after the fire. Black error bars show
981 standard error and the double-sided 95% probability range of the geometric means. a) Particle
982 lidar ratios for 532 nm where data points with fitting error exceeding 25% are discarded. The
983 black line shows the geometric mean after day 4, and the full and dotted blue lines show the
984 standard deviation and the double-sided 95% probability range of the distribution. b) Particle
985 color ratio (1064 nm divided by 532 nm wavelength backscattering) with exponential fit ($R^2 =$
986 0.48 , $P < 10^{-10}$), and c) particle depolarization ratio with exponential fit ($R^2 = 0.76$, $P < 10^{-10}$).
987 The color and depolarization ratios were divided in two equal groups by number of observations
988 to illustrate the highly significant changes with time of the optical properties, where the long and
989 short error bars are the standard error and the double-sided 95% probability range of the
990 geometric means.



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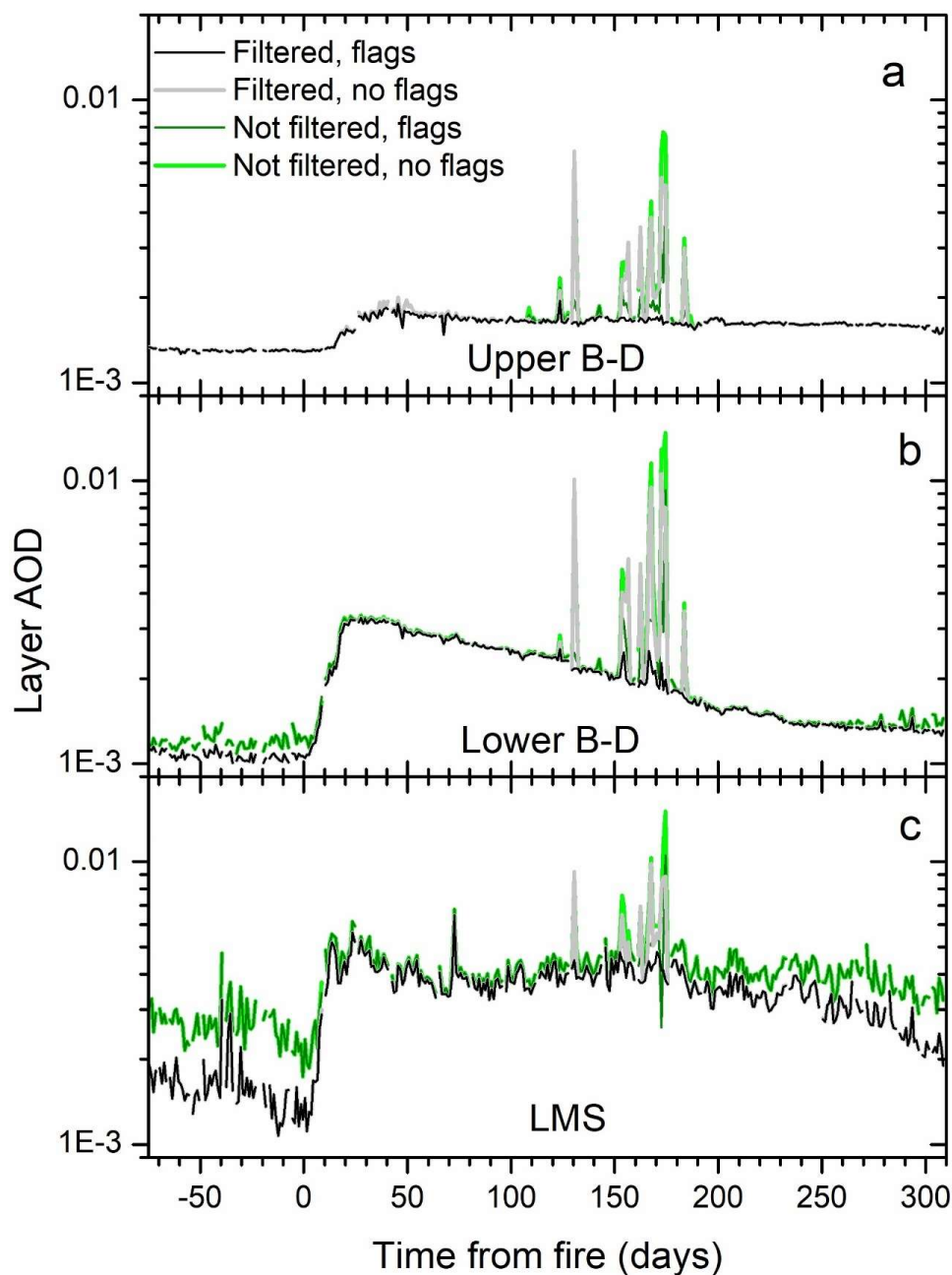
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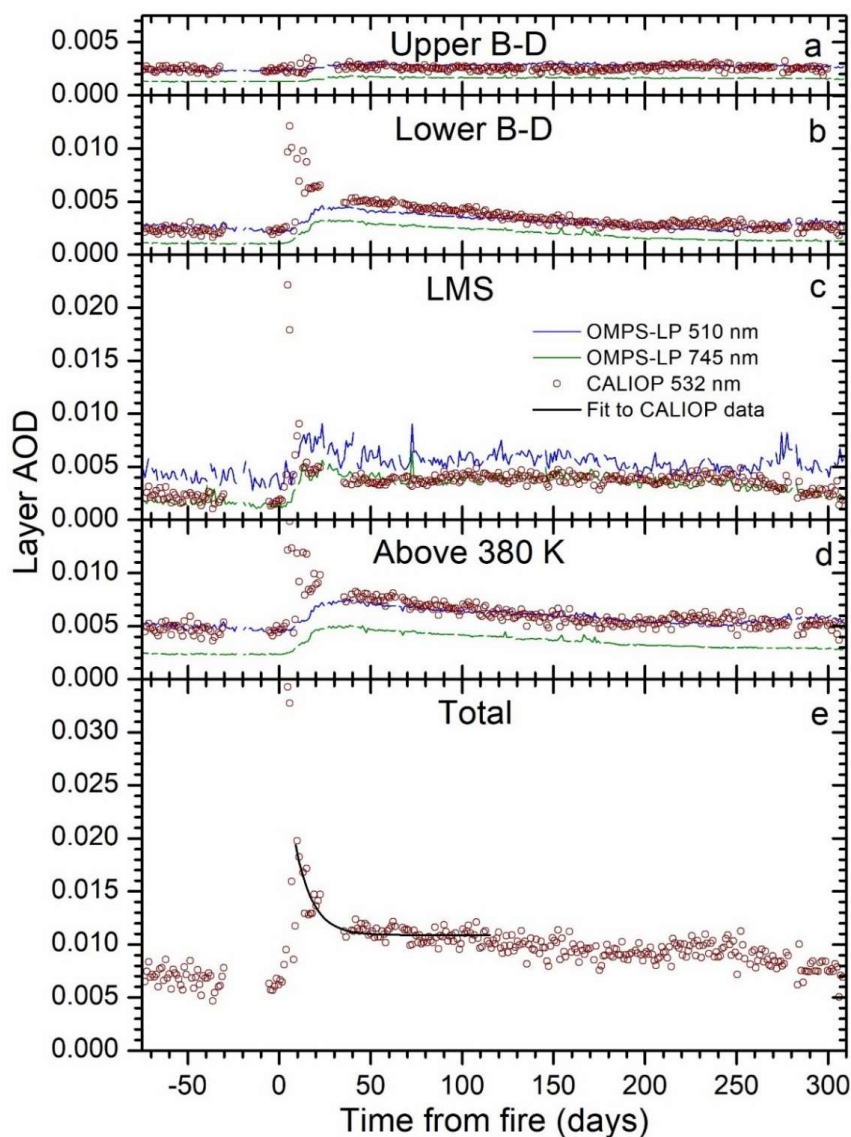
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Figure 3. Illustration of methodology and its effect. a) The attenuated and corrected scattering ratios as a function of altitude. Example of methodology for one smoke layer, where the scattering ratio between 7.5 – 10 km altitude, below the smoke layer at 10.5 – 14 km, is targeted to a value of 1.08 (explained in the method section) by iteratively fitting the lidar ratio for 532 nm wavelength. b) The attenuated layer AOD (AOD_{att}) related to the layer AOD corrected for attenuation. The 1:1 relation is shown by the full line.



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1001 Figure 4. OMPS-LP layer AODs averaged over 20 to 80° North for 745 nm wavelength using
1002 data filtered and not filtered from clouds and polar stratospheric clouds, and with and without
1003 data flagged for data quality. Layer AOD for a) the upper Brewer-Dobson branch (470 K
1004 isentrope – 35 km), b) the lower Brewer-Dobson branch (380 – 470 K) and c) the LMS
1005 (tropopause – 380 K) are shown.



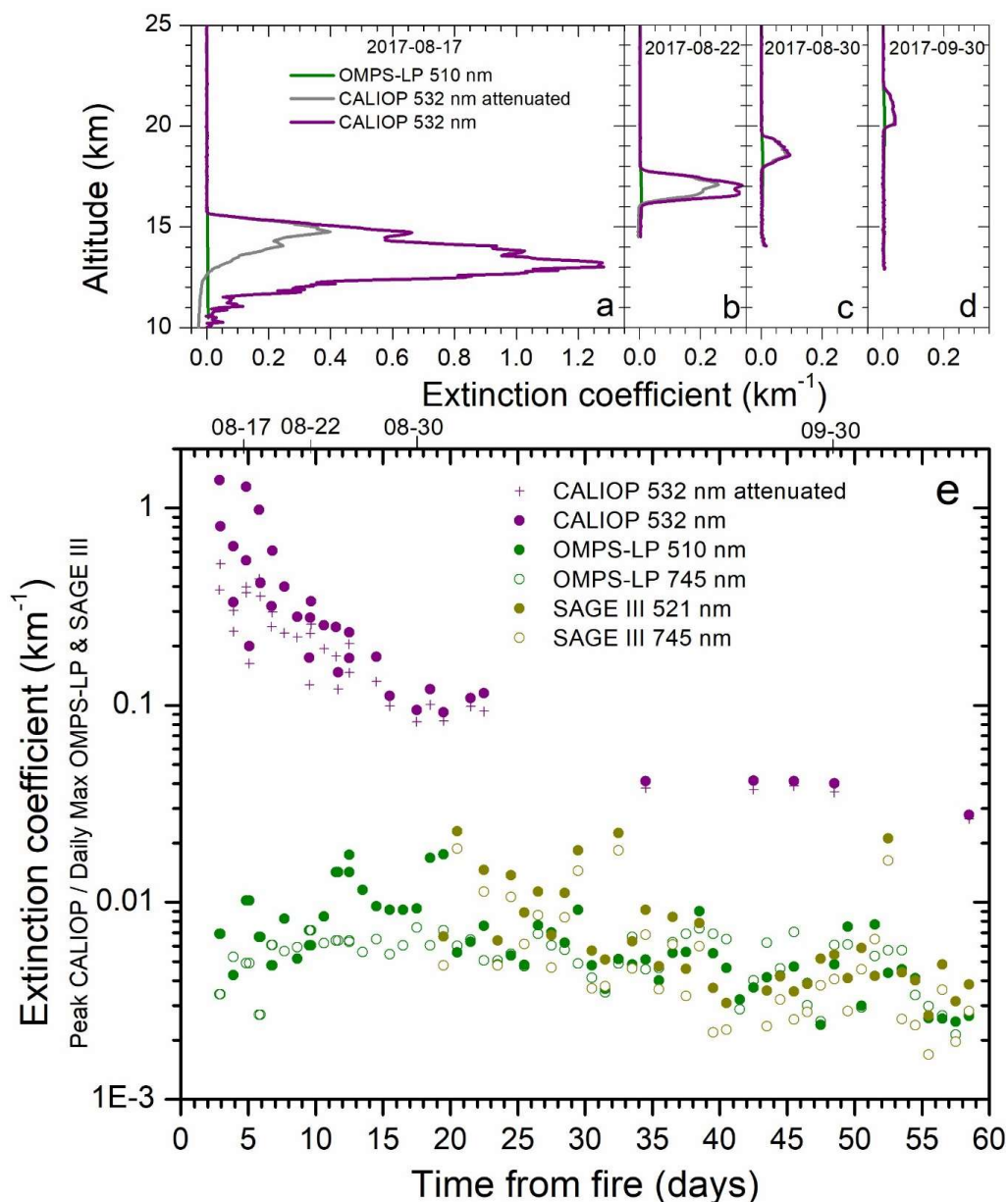
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1007 Figure 5. AOD evolution of the stratospheric AOD (daily average) from 75 days before to 310
1008 days after the 2017 western North American fires. Comparisons of AOD from CALIOP (532 nm)
1009 with OMPS-LP (510 and 745 nm) with cloud filtering and flags activated for a) the upper
1010 Brewer-Dobson branch (470 K isentrope – 35 km, b) the lower Brewer-Dobson branch (380 –
1011 470 K) c) the LMS (tropopause – 380 K), d) from 380 K to 35 km (sum of layers in a and b) and
1012 e) the stratosphere of CALIOP from the tropopause to 35 km (sum of layers in a, b and c). The
1013 black, full line is an exponential fit ($R^2 = 0.79$, $P < 10^{-10}$) to the AOD over days 10 – 115 after the
1014 fire. The total stratospheric AOD half-life of the fit is 6.5 ± 0.9 days.



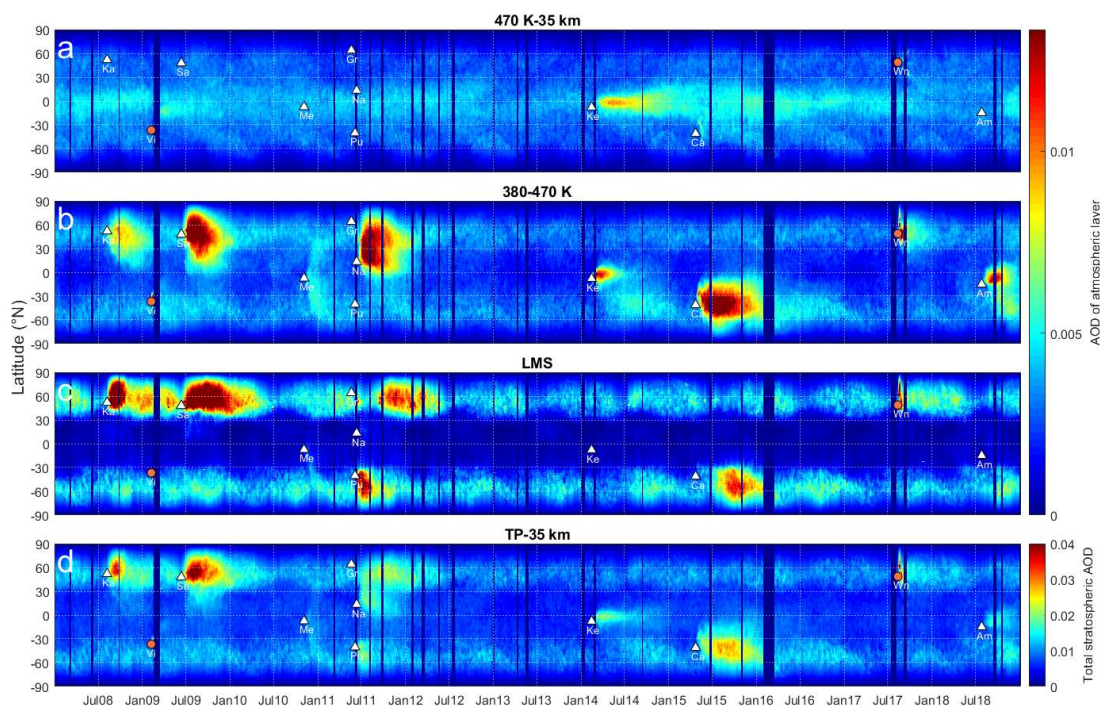
1015

1016 Figure 6. Daily OMPS-NM aerosol absorbing index (UV) August 14 – 22, 2017 over all
1017 longitudes and latitudes 20 - 80° N. This index is sensitive to UV absorbing aerosol particles in
1018 the upper troposphere and the stratosphere, where signals from tropospheric aerosol declines
1019 faster than from stratospheric due to short residence time. The yellow lines indicate nighttime
1020 swaths of the CALIPSO satellite, and the faint lines show CALIPSO daytime swaths.



1021

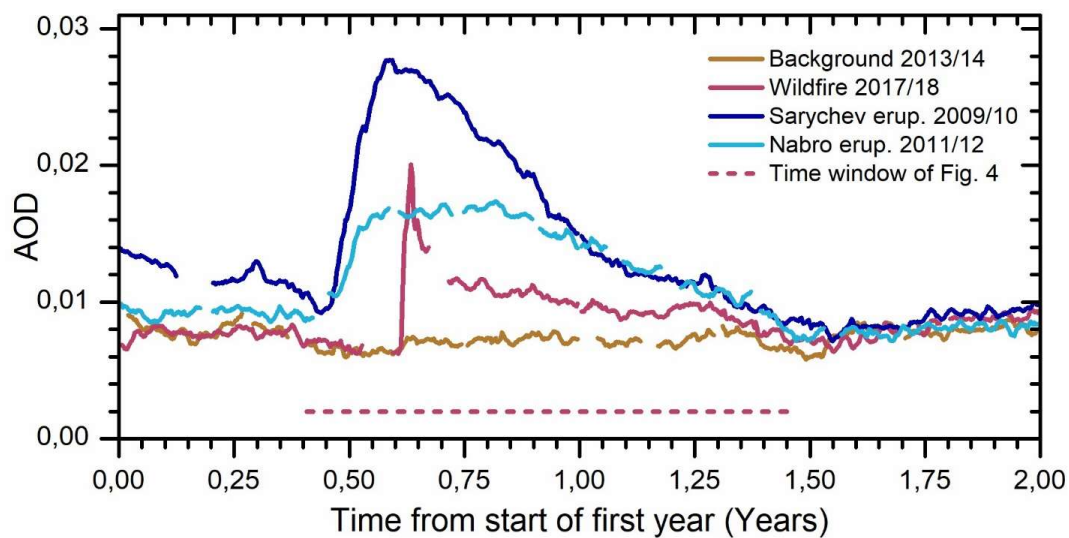
1022 Figure 7. Extinction coefficients according to CALIOP, OMPS-LP and SAGE III/ISS in the 20 -
 1023 80° North latitude range during the first 60 days following the North American fire. a - d)
 1024 selected profiles (attenuated and corrected for attenuation) from CALIOP compared with closest
 1025 profiles according to OMPS-LP. e) Peak extinction coefficient from selected CALIOP profiles
 1026 compared with daily maximum extinction coefficients from OMPS-LP and SAGE III/ISS. Note
 1027 that SAGE III/ISS data are missing the first 19 days because of irregular coverage of the latitude
 1028 range of interest.



1029

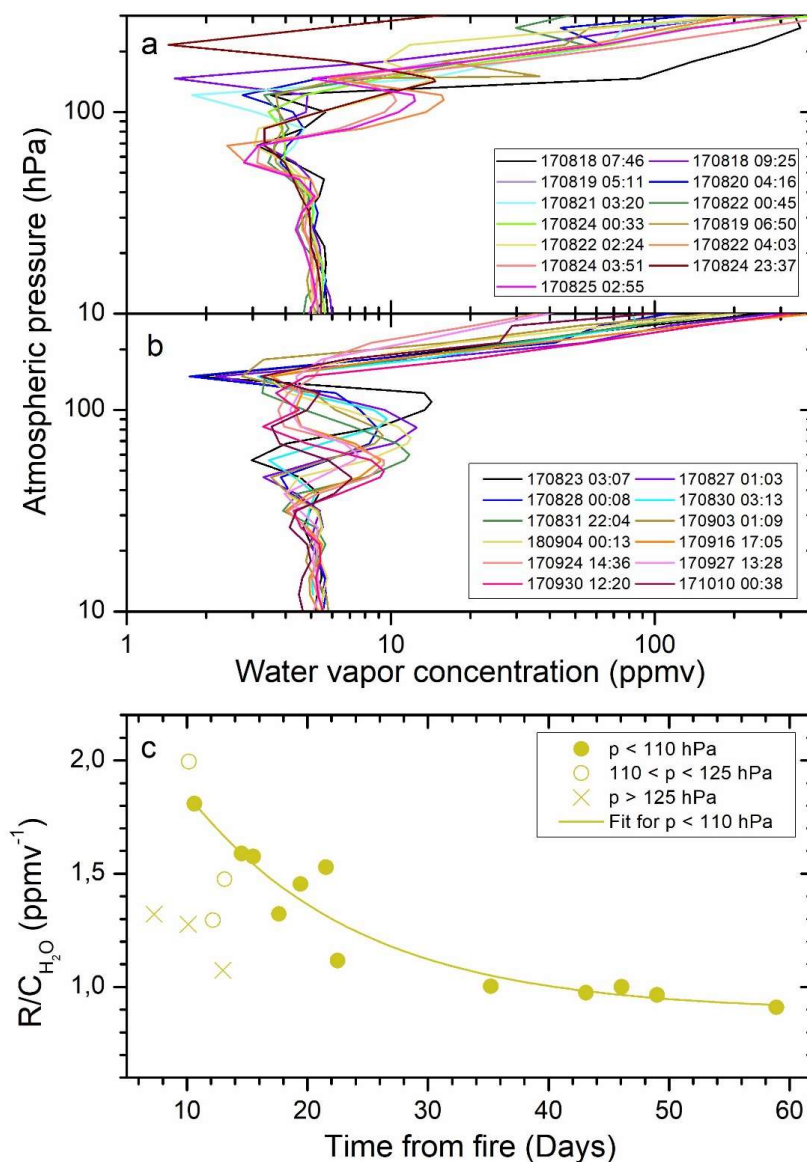
1030

1031 Figure 8. Zonally and eight-day moving average aerosol optical depth (AOD) of the stratosphere.
1032 a - c) AOD in three layers obtained from CALIOP data (level 1B): a) 470 K potential temperature
1033 to 35 km (deep Brewer-Dobson branch), b) 380 – 470 K (shallow Brewer-Dobson branch), c) the
1034 tropopause to 380 K (LMS). d) The total AOD from the tropopause to 35 km altitude. Volcanic
1035 eruptions marked by white triangles: Kasatochi (Ka), Sarychev (Sa), Merapi (Me), Grimsvötn
1036 (Gr), Puyehue-Cordón Caulle (Pu), Nabro (Na), Kelut (Ke), Calbuco (Ca), and Ambae (Am), and
1037 wildfires marked by orange circles: Victoria fire (Vi) and Western North American fires (Wn) at
1038 time and latitude of eruption/fire. The AODs are corrected for attenuation.



1039

1040 Figure 9. Evolution of the AOD in the 20 - 80° N interval (8-day moving average) over two
1041 years: close to background conditions in the latitude interval studied (2013 – 2014), the year and
1042 the following year of the August 12, 2017, fire (2017 – 2018), and the same for two volcanic
1043 eruptions, the June 12, 2009, Sarychev (2009 – 2010) and June 12, 2011, Nabro (2011 – 2012)
1044 eruptions.



1045

1046 Figure 10. Water vapor in the smoke layer. Microwave Limb Sounder (MLS) measurements of
 1047 water vapor concentrations (ppmv) Vs. atmospheric pressure for smoke layers a) close to the
 1048 tropopause and b) well above the tropopause (atmospheric pressure < 110 hPa at the H₂O peak)
 1049 for individual smoke layers available days 6 – 60 after the fire. c) The peak scattering ratio (R)
 1050 according to CALIOP divided by the peak water vapor concentration (C_{H₂O}) from MLS. The full
 1051 line is an exponential fit ($R^2 = 0.88$, $P < 3 \times 10^{-10}$) to smoke layers peaking in water vapor
 1052 concentration at a pressure less than 110 hPa. The half-life of the fit is 9.7 ± 3.2 days.