1	Enhanced light absorption and reduced snow albedo due to
2	internally mixed mineral dust in grains of snow
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1 Abstract. Mineral dust is a major light-absorbing aerosol, which can significantly 2 reduce snow albedo and accelerate snow/glacier melting via wet and dry deposition on snow. In this study, three scenarios of internal mixing of dust in ice grains were 3 4 analyzed theoretically by combining asymptotic radiative transfer theory and 5 (core/shell) Mie theory to evaluate the effects on absorption coefficient and albedo of 6 the semi-infinite snowpack consisting of spherical snow grains. In general, snow albedo 7 was substantially reduced at wavelengths of $<1.0 \mu m$ by internal dust-snow mixing, 8 with stronger reductions at higher dust concentrations and larger snow grain sizes. 9 Moreover, calculations showed that a non-uniform distribution of dust in snow grains 10 can lead to significant differences in the values of the absorption coefficient and albedo 11 of dust-contaminated snowpack at visible wavelengths relative to a uniform dust 12 distribution in snow grains. Finally, using comprehensive in situ measurements across 13 the Northern Hemisphere, we found that broadband snow albedo was further reduced 14 by 5.2% and 9.1% due to the effects of internal dust-snow mixing on the Tibetan 15 Plateau and North American mountains. This was higher than the reduction in snow 16 albedo caused by black carbon in snow over most North American and Arctic regions. 17 Our results suggest that significant dust-snow internal mixing is important for the 18 melting and retreat of Tibetan glaciers and North American mountain snowpack.

1 1. Introduction

2 Snow cover is one of the most reflective surfaces in the Earth system, and plays a crucial 3 role in the atmospheric solar radiation energy budget via snow albedo feedbacks (Di 4 Mauro et al., 2020; Flanner et al., 2011; Jacobson, 2004; Usha et al., 2020; Xie et al., 5 2018). Previous studies have shown that light-absorbing particles (LAPs) effectively 6 reduce snow albedo and enhance the absorption of solar radiation after deposition, these 7 studies were based on in situ observations and model simulations (Casey et al., 2017; 8 Hadley and Kirchstetter, 2012; Shi et al., 2020; Warren and Wiscombe, 1980; Yasunari 9 et al., 2012; Yasunari et al., 2015). As a result, snow contaminated with LAPs shows 10 significant changes in morphology (Niwano et al., 2012; Rango et al., 1996), chemistry 11 (France et al., 2012; Reay et al., 2012), hydrology (Matt et al., 2018; Qian et al., 2015; 12 Rahimi et al., 2019), snowmelt rate (Kaspari et al., 2015; Warren, 1984;), and radiative 13 properties (Grenfell et al., 2002; Hansen and Nazarenko, 2004; Zhao et al., 2014).

14 Numerous studies have assessed the potential effects of LAPs, such as black carbon 15 (BC) and mineral dust, on snow albedo by assuming that LAPs mixed outside spherical 16 snow grains (i.e., external mixing) (Flanner et al., 2007; Kokhanovsky, 2013; Libois et 17 al., 2013; Wang et al., 2017; Warren and Wiscombe, 1980). For example, Warren and 18 Wiscombe (1980) calculated snow spectral albedo by solving a radiative transfer 19 equation using Mie theory and δ -Eddington approximations, and found that 10–100 ng g^{-1} of BC and 1–10 $\mu g g^{-1}$ of dust in old snow decreased albedo by 1%–7% and 2%– 20 10% at 400 nm wavelength, respectively. Flanner et al. (2007) pointed out that the 21 reduction in snow albedo for 1000 ng g^{-1} BC in snow was 0.045 (0.17) with a 50 (1000) 22 23 µm snow grain radius (Ref) based on the Snow, Ice, and Aerosol Radiation (SNICAR) 24 radiative transfer model. This model utilizes theory from Wiscombe and Warren (1980) 25 and the two-stream radiative transfer solution from Toon et al. (1989). Wang et al. (2017) 26 developed a Spectral Albedo Model for Dirty Snow (SAMDS) based on asymptotic 27 radiative transfer theory, which is a function of the snow grain radius, LAP (e.g., BC, 28 dust) mixing ratios, and the mass absorption coefficients (MACs) of LAPs. Their results

1 revealed that broadband snow albedo decreased 0.03 and 0.003 due to 200 ng g⁻¹ of BC 2 and 2 μ g g⁻¹ of dust in snow, respectively, with a R_{ef} of 200 μ m. Additionally, recent 3 studies found that snow nonsphericity can interact with LAP-snow mixing, which leads 4 to weaker LAP-induced albedo reductions for nonspherical snow shapes than snow 5 spheres (e.g., Dang et al., 2016; He et al., 2018, 2019).

6 Recently, direct snowpack observations have shown evidence for the existence of LAP-7 snow internal mixing (e.g., Horhold et al., 2012; Spaulding et al., 2011). LAPs tend to mix externally with snow grains through dry deposition and/or below cloud scavenging, 8 9 while internal LAP-snow mixing can be produced by nucleation, accretion, riming, 10 aggregation, and sintering during aerosol-cloud-precipitation processes (i.e., wet 11 deposition; Figure 1) (Flanner et al., 2012). Furthermore, Flanner et al. (2012) found 12 that internal BC/ice mixing (IBM) increased the absorption of snowpack by a factor 13 1.8-2.1 relative to external BC/ice mixing (EBM). He et al. (2018) indicated that IBM 14 enhanced the mean snow albedo reduction over the Tibetan plateau by 30%-60% 15 relative to EBM, based on the updated SNICAR model. Additionally, Dombrovsky and Kokhanovsky (2020) demonstrated that non-uniform BC distribution in ice grains may 16 17 lead to significantly different absorption coefficients and snowpack albedo in visible 18 (VIS) wavelengths.

19 Numerous studies have addressed the role of IBM in enhancing the absorption of 20 snowpack due to its strong absorption effect relative to other LAPs (Dombrovsky and 21 Kokhanovsky, 2020; Flanner et al., 2012; He et al., 2018; Liou et al., 2011). In contrast, 22 few studies have considered the effects of internal dust/ice mixing (IDM) in snowpack. 23 Liou et al. (2014) is the pioneer to investigate the dust-snow internal mixing effects 24 based on the geometric-optics surface-wave approach. Subsequently, He et al. (2019) 25 used the same method to explicitly quantify the combined effects of dust-snow internal 26 mixing and snow grain nonsphericity on snow optical properties, thereafter, develop a 27 set of new dust-snow parameterizations for land/climate modeling applications for the first time. Actually, dust particles are generally larger than BC, and act as more efficient 28

1 ice nuclei, showing a better ability to influence cloud formation and precipitation 2 (Creamean et al., 2013; Huang et al., 2014). Therefore, they are more likely to mix 3 internally with ice grains. Furthermore, dust can also dominate light absorption and 4 effectively decrease snow albedo because of its relatively high mass abundance (ppm) in snowpack, especially in areas with seasonal and patchy snow cover or mountainous 5 6 regions (Di Mauro et al., 2015; Gabbi et al., 2015; Painter et al., 2012; Reynolds et al., 7 2020; Xie et al., 2018). Therefore, it is important to account for IDM when estimating 8 the impact of dust deposition on snow albedo.

9 In this study, we assess the effects of external/internal mixing of dust with ice grains on 10 the snowpack absorption coefficient and albedo using asymptotic radiative transfer 11 theory and Mie theory. In addition, the uniformity and nonuniformity of dust particle 12 distribution inside ice grains are considered for IDM based on the effective medium 13 approximation, and the combined effects of dust content and snow grain radius on snow 14 albedo are quantified. A schematic of various dust spatial distributions from this study 15 is presented in Figure 2. We further discuss snow albedo sensitivity to complex 16 refractive indices and dust particle size distribution. Based on a comprehensive set of 17 field measurements of dust concentrations, we estimate the reductions in snow albedo 18 by dust external/internal mixing with ice grains across the Northern Hemisphere.

19 2 Methods

20 2.1 External mixing model

For fairly pure snow, semi-infinite means about 20 cm in the VIS and 3 cm in the nearinfrared (NIR) regions of snow depths, respectively (Zhou et al., 2003). For a semiinfinite snow layer under diffuse illumination conditions, albedo can be calculated using an asymptotic analysis of radiative transfer theory, which is valid in small absorptions (Kokhanovsky and Zege, 2004; Zege et al., 1991):

$$26 \quad \alpha_{\lambda} = \exp\left(-4S_{\lambda}\right) \tag{1}$$

27 where α_{λ} is the spectral snow albedo, λ is the wavelength, S_{λ} is the similarity

1 parameter, and

2
$$S_{\lambda} = \sqrt{\frac{\sigma_{abs}}{3\sigma_{ext}(1-g)}}$$
 (2)

In Eq. (2), σ_{abs} and σ_{ext} are the absorption and extinction coefficients, respectively, and g is the asymmetry parameter (the average cosine of the phase function of the medium).

According to Eq. (18) and (25) in Kokhanovsky and Zege (2004), the extinction
coefficients of particles can be expressed as:

8
$$\sigma_{ext} = \frac{1}{l_{tr}(1-g)} = \frac{3C_v}{2r_{ef}}$$
 (3)

9 where l_{tr} is the photon transport path length, $C_v = \rho_{snow}/\rho_{ice}$ is the volumetric 10 snow particle concentration, and the values $\rho_{ice} = 916.7 \text{ kg m}^{-3}$ and $\rho_{snow} = 300 \text{ kg}$ 11 m⁻³ are used in subsequent calculations. r_{ef} is the effective snow grain radius, which 12 is equal to the radius of the volume-to-surface equivalent sphere $(r_{ef} = \frac{3\overline{V}}{4\overline{A}})$ where \overline{V} 13 and \overline{A} are the average volume and cross-sectional (geometric shadow) area of snow 14 grains, respectively.

For external dust/ice mixing (EDM) in a dust-contaminated snowpack, the total absorption coefficient (σ_{abs}) can be derived from the absorptions by snow (σ_{abs}^{snow}) and dust (σ_{abs}^{dust}):

18
$$\sigma_{abs} = \sigma_{abs}^{snow} + \sigma_{abs}^{dust}$$
 (4)

For example, consider a hypothetical case of snow composed of monodispersed, spherical grains of ice. Although non-spherical snow grains lead to a slight increase in snow albedo, Dang et al. (2016) found that the albedo of a snowpack consisting of nonspherical snow grains can be mimicked by using smaller, spherical grains; thus, we do not consider the effect of non-spherical snow grains in this study. Therefore, we used the following equation for the absorption coefficient of snow (Dombrovsky and Baillis, 2010):

$$26 \qquad \sigma_{abs}^{snow} = \frac{0.75C_v \cdot Q_{abs}^{ice}}{r_{ef}} \tag{5}$$

27 where Q_{abs}^{ice} is the efficiency factor of absorption for a single ice grain, and the value

of Q^{ice}_{abs} can be calculated for homogeneous spherical ice grains considered in classical
 Mie theory.

The absorption coefficient of dust (Aoki et al., 2000; Marley et al., 2001; Warren et al.,
2006) is expressed as:

5
$$\sigma_{abs}^{dust} = \frac{Q_{abs}^{dust} \cdot \pi \cdot \left(r_{ef}^{dust}\right)^2}{\frac{4}{3}\pi \cdot \left(r_{ef}^{dust}\right)^3} \cdot C_{dust} = \frac{3Q_{abs}^{dust}}{4r_{ef}^{dust}} \cdot C_{dust} = MAC_{abs}^{dust} \cdot \rho_{dust} \cdot C_{dust}$$
(6)

where Q_{abs}^{dust} and MAC_{abs}^{dust} is the absorption efficiency and MAC of dust, 6 respectively, that can be obtained via Mie theory, and ρ_{dust} and r_{ef}^{dust} represent the 7 density and effective dust radius, respectively. In this study, ρ_{dust} was assumed to be 8 9 2500 kg m⁻³ (Zender et al., 2003). We also assumed a log-normal dust size distribution 10 with a geometric mean diameter of 0.65 µm and standard deviation of 2.0 (equivalent 11 to an effective radius of $1.1 \ \mu m$), which represents dust from large-scale transport 12 (Formenti et al., 2011; Maring et al., 2003) that is likely smaller in size than from local 13 soil (Kok, 2011). The effects of dust size on snow optical properties and albedo were 14 further quantified through sensitivity simulations (see section 3.4). Dust volumetric 15 concentrations (C_{dust}) are expressed as:

16
$$C_{dust} = \frac{\rho_{snow} \cdot C_{dust}^*}{\rho_{dust}}$$
 (7)

where C_{dust}^* is the mass concentration of dust in snow (kg kg⁻¹). Thus, the spectral albedo of dust-contaminated snow for EDM can be easily calculated with Eq. (1) to (7).

For the IDM (Figure 2), we first determined the effective optical constants of ice containing small dust particles via an effective medium approximation (Maxwell-Garnett and Larmor, 1904). According to this approach, the complex permittivity of a composite medium in ice grains can be calculated in terms of particle polarizability by applying the Lorentz-Lorenz formula (Koledintseva et al., 2009; Markel, 2016). We used the following relationships to calculate effective complex refractive indices (RIs), $m_{ef} = n_{ef} - ik_{ef}$ at known values of $m_{ice} = n_{ice} - ik_{ice}$ for pure ice and $m_{dust} = n_{dust} - ik_{dust}$ 1 for dust:

2
$$m_{ef}^2 = m_{ice}^2 \frac{2\delta_{dust}(m_{dust}^2 - m_{ice}^2) + m_{dust}^2 + 2m_{ice}^2}{2m_{ice}^2 + m_{dust}^2 - \delta_{dust}(m_{dust}^2 - m_{ice}^2)}$$
 (8)

3 where δ_{dust} is the local dust fraction volume in an ice grain. We obtained the spectral 4 complex RIs of ice from Warren and Brandt (2008) and the spectral complex RIs of 5 dust from Dang et al. (2015). The imaginary part of the complex RIs of ice (k_{ice}) and 6 dust (k_{dust}) associated with absorption is shown in Figure 3. We also evaluated the effect 7 of dust on the imaginary part of the effective complex RIs (k_{ef}) assuming dust mass 8 concentrations (C_{dust}^*) of 1–100 ppm (or µg g⁻¹) in snow.

9 In all variations of the spatial distribution of dust particles in snow, the dust mass
10 concentration was assumed to be constant, which means that the local dust fraction
11 volume may differ. In an example of spherical ice grains with uniformly distributed
12 dust, the dust fraction volume in an ice grain was determined as:

13
$$\delta_{dust}^{0} = \frac{C_{dust}^{*}}{C_{v}} \frac{\rho_{ice}}{\rho_{dust}}$$
(9)

14 where the ratio of C_{dust}^*/C_v is the mass fraction of dust in the ice grain.

We considered two cases of non-uniform dust distributed in a spherical ice grain with radius r_{ef} : (1) We assumed that the same mass of dust is uniformly distributed in the central part of the ice grain ($r_{ef}^{dust} < r_c \le r_{ef}$). (2) We assumed all of the dust was in the surface layer of the ice grain ($r_{ef}^{dust} < r_p \le r_{ef}$) (Figure 2). In both cases the local value of δ_{dust} increases as:

20
$$\delta_{dust} = \delta_{dust}^0 / \psi$$
 $\psi = \begin{cases} \overline{r_c}^3 & \text{central pollution} \\ 1 - (1 - \overline{r_p})^3 & \text{peripheral pollution} \end{cases}$ (10)

21 where $\bar{r}_c = r_c/r_{ef}$ and $\bar{r}_p = r_p/r_{ef}$. $\bar{r}_c = 1$ and $\bar{r}_p = 1$ correspond to uniformly 22 distributed dust when $\psi = 1$ and $\delta_{dust} = \delta^0_{dust}$, and $\psi < 1$ and $\delta_{dust} > \delta^0_{dust}$ in 23 other cases. Obviously, dust particles increase the imaginary part of the effective 24 complex RIs in polluted ice grains (Figure 3).

In summary, the m_{ef} of a spherical ice grain with uniformly and non-uniformly distributed, according to Eq. (8) to (10), then their corresponding absorption efficiencies can be obtained using classical Mie theory and core/shell Mie theory,
 respectively. The spectral snow albedo for IDM can be easily calculated using Eq. (1)
 to (5).

4 **2.3 Broadband snow albedo calculations**

5 The spectral albedo (α_{λ}) is integrated over the solar spectrum ($\lambda = 300-2500$ nm) and 6 weighted by incoming solar irradiance (E_{λ}) to calculate broadband snow albedo (Hadley 7 and Kirchstetter, 2012):

8
$$\alpha_{integrated} = \frac{\int \alpha_{\lambda} E_{\lambda} d\lambda}{\int E_{\lambda} d\lambda}$$
 (11)

9 Following the study of Dang et al. (2017), the incoming solar irradiance we used is a 10 typical surface solar spectrum at mid to high latitudes from January to March under the 11 cloudy sky, calculated by the Santa Barbara DISORT Atmospheric Radiative Transfer 12 (SBDART) model (Pu et al., 2019). The SBDART model is a widely used atmospheric 13 radiation transfer model based on a collection of highly developed physical models, 14 including the Discrete Ordinate Radiative Transfer module (Stamnes et al., 1988), low-15 resolution atmospheric transmission models, and Mie theory. The SBDART model can 16 be used to compute radiative transfer at different heights and directions under both clear 17 and cloudy sky conditions. Details on the SBDART model can be found in Ricchiazzi 18 et al. (1998).

19 **2.4 Dust concentration measurements**

20 To estimate the effect of dust on snow albedo in real snowpack, we collected a 21 comprehensive set of in situ dust concentration measurements during field campaigns 22 to Inner Mongolia, China (IMC) (Wang et al., 2013), the Tibetan Plateau (TP) (Li et al., 23 2017, 2018; Li et al., 2016, 2019; Niu et al., 2017; Qu et al., 2014; Zhang et al., 2017, 24 2018), Sapporo, Japan (SJ) (Kuchiki et al., 2015), the European Alps (EP) (Di Mauro 25 et al., 2015; Lim et al., 2014), and North American mountains (NAM) (Painter et al., 26 2012; Reynolds et al., 2020). The field campaigns conducted in the TP can be further grouped into three regions, as in previous δ^{18} O precipitation studies (Yao et al., 2013). 27

These three distinct domains were associated with the Indian monsoon (Southern TP), westerlies (Northern TP), and transition (Central TP) (Yao et al., 2013). It is worth noting that we only considered regions with higher dust concentrations (>1 ppm), such that polar regions are not included in this study. Measurements of dust in the snow samples were generally obtained by weighing the filter before and after filtration using a microbalance.

7 **3. Results**

8 **3.1 Impact on the imaginary part of the effective complex RIs**

9 We evaluated the effect of dust on k_{ef} , including k_{ice} and k_{dust} associated with absorption 10 (Figure 3). The k_{dust} was in a narrow range (~0.001–0.01) and gradually decreased with 11 increasing wavelength in the ultraviolet (UV) and VIS regions, then remained stable in the NIR band. The k_{ice} varied by eight orders of magnitude from the UV (~10⁻¹¹) to NIR 12 (~10⁻³) bands, and increased with increasing wavelength, except at 1.03 μ m where k_{ice} 13 14 decreased slightly as a result of the presence of ice absorption features (Warren, 2019). Figure 3 also shows the k_{ef} with dust mass concentrations ranging from 1 to 100 ppm 15 and wavelengths of 300–1500 nm. The k_{ef} clearly varied depending on the wavelength 16 and increased with dust mass concentrations. For a given dust mass concentration, k_{ef} 17 18 decreased with wavelengths from UV to VIS, then increased from VIS to NIR. For example, the value of k_{ef} decreased from 4.26×10^{-8} at 300 nm to 1.36×10^{-8} at 500 19 nm, then rose to 1.73×10^{-6} at 1000 nm at a dust concentration of 10 ppm. Moreover, 20 21 the wavelength of the minimum of k_{ef} varied from ~500 nm to ~650 nm depending on the dust mass concentrations (1 to 100 ppm). Additionally, it is worth noting that kef was 22 23 not sensitive to dust mass concentrations in wavelengths >1000 nm, which was generally consistent with k_{ice} , because the difference between k_{dust} and k_{ice} is more than 24 25 compensated by the much larger difference in ice and dust concentration at those 26 wavelengths. Conversely, k_{ef} showed significant differences relative to k_{ice} in the UV 27 and VIS regions, with higher dust mass concentrations demonstrating larger differences. For example, k_{ef} was enhanced by 3, 21, and 205 times at 500 nm relative to k_{ice} for dust mass concentrations of 1, 10, and 100 ppm, respectively.

3 3.2 Impact on spectral snow absorption coefficient and albedo

4 Dust in snow effectively enhances the snow absorption coefficient, but its effect on the 5 snow asymmetry factor and extinction efficiency is negligible (He et al., 2019). 6 Therefore, we mainly focused on the effects of EDM and three cases of IDM (uniform, 7 central, and peripheral) on the snow absorption coefficient (σ_{abs}). Figure 4a displays the 8 σ_{abs} for EDM and IDM (uniform) as a function of wavelength at different dust 9 concentrations. We used a snow grain radius of 200 µm (Figure 4), which is comparable 10 to previous observations of seasonal snow at mid to high latitudes in winter (Shi et al., 11 2020; Wang et al., 2017). The results showed that EDM and IDM have distinct impacts on $\sigma_{abs}\,$ in UV and VIS, but small effects at wavelengths >1000 nm, which is due to 12 13 the optical properties of snow being affected by LAPs in UV and VIS and primarily 14 affected by snow itself at wavelengths >1000 nm. Additionally, σ_{abs} increased with increased dust mass concentrations. For instance, σ_{abs} increased from 0.007 m⁻¹ (pure 15 snow) to 0.03, 0.14, and 1.37 m⁻¹ at 500 nm with 2, 10, and 100 ppm of dust with EDM, 16 respectively. For IDM (uniform), σ_{abs} increased to 0.06, 0.28, and 2.80 m⁻¹ at 500 nm 17 18 with 2, 10, and 100 ppm of dust, respectively, with corresponding enhancement factors 19 of σ_{abs} (E_{σ_{abs}}, defined as the absorption coefficient of IDM divided by EDM) were 20 1.84, 2.00, and 2.05. Furthermore, the σ_{abs} for two cases of non-uniform dust 21 distribution in a spherical ice grain ($r_{ef} = 200 \ \mu m$) depends on the wavelength, dust mass 22 concentrations, $\overline{r_c}$, and $\overline{r_p}$ (Figure 4b and 4c). We note that $\overline{r_c}$ and $\overline{r_p}$ values of 1 correspond to uniformly distributed dust, and the $\sigma_{abs}\,$ increases and decreases (with 23 24 the decrease of $\overline{r_c}$ and $\overline{r_p}$) for IDM (central) and IDM (peripheral), respectively. For 25 example, $\,\sigma_{abs}\,$ increased by 29%, 32%, and 33% (500 nm) at dust mass concentrations of 2, 10, and 100 ppm, respectively, when $\bar{r_c}$ decreased from 1 to 0.7 for IDM (central). 26 However, σ_{abs} decreased by 41%, 44%, and 44% (500 nm) at dust mass concentrations 27 of 2, 10, and 100 ppm, respectively, when $\bar{r_p}$ decreased from 1 to 0.1 for IDM 28

1 (peripheral). This indicates that the IDM (central) further enhanced snowpack light 2 absorption compared with the IDM (uniform), while the IDM (peripheral) reduced 3 snowpack light absorption with a corresponding σ_{abs} between the value of σ_{abs} for 4 IDM (uniform) and EDM. Figure 4d-f quantitatively shows the spectral snow absorption coefficient enhancement for IDM ($E_{\sigma_{abs}}$). The enhancement decreased 5 6 sharply with increasing wavelengths, then reduced to 1.0 (i.e., no enhancement) at 7 wavelengths longer than ~1.0 µm because of strong dust absorption and weak snow absorption at shorter wavelengths. Obviously, $E_{\sigma_{abs}}$ was affected by dust mass 8 9 concentration, $\overline{r_c}$, and $\overline{r_p}$, but $E_{\sigma_{abs}}$ was insensitive to dust mass concentration at 10 wavelengths <450 nm.

11 Figure 5a-c shows the spectral snow albedo (α_{λ}) for EDM and IDM; α_{λ} was 12 consistent with σ_{abs} , with the effects mainly present at wavelengths <1000 nm. 13 Generally, α_{λ} decreased with increased dust mass concentrations in UV and VIS, and 14 IDM was shown to further trigger the reduction of α_{λ} . For example, for EDM α_{λ} was 15 ~0.97, 0.95, 0.85 (at 500 nm) for dust concentrations of 2, 10, 100 ppm, respectively, 16 which was higher than the values for IDM (uniform) (~0.96, 0.93, 0.79, respectively). 17 Compared with IDM (uniform), the α_{λ} for IDM (central) decreased by 0.5%, 1.1%, 18 and 3.5% (at 500 nm) for dust concentrations of 2, 10, and 100 ppm, respectively, when 19 $\overline{r_c}$ decreased from 1 to 0.7. α_{λ} for IDM (peripheral) increased by 0.8%, 1.9%, and 6.2% 20 (at 500 nm) for the same dust concentrations, when $\bar{r_p}$ decreased from 1 to 0.1. 21 Moreover, the wavelength of the maximum α_{λ} value varied from ~500 nm to ~650 nm 22 depending on the dust mass concentrations, which is consistent with changes of k_{ef} . 23 Figure 5d–f shows the ratio $(E_{\alpha_{\lambda}})$ of snow spectral albedo for IDM to EDM where we observed that the $E_{\alpha_{\lambda}}$ increased with increasing wavelengths and dust concentrations, 24 25 and then became stable at 1.0. This is because IDM can enhance the light absorption of snowpack more effectively at shorter wavelengths and higher dust concentrations 26 (Figure 4). Additionally, the values of $\overline{r_c}$ and $\overline{r_p}$ also have non-negligible effects on 27 $E_{\alpha_{\lambda}}$, which can be decreased and increased with decreasing $\overline{r_c}$ and $\overline{r_p}$, respectively. 28

1 We found that the optical properties of an ice grain containing uniformly distributed 2 dust in its center, or concentric surface layer, can be affected by $\overline{r_c}$ or $\overline{r_p}$. To better understand this effect, Figure 6a–b displays the $\,\sigma_{abs}\,$ at 500 nm as a function of $\,\overline{r_c}\,$ and 3 4 $\overline{r_p}$ with different dust concentrations and r_{ef} . This demonstrates r_{ef} has negligible 5 effects on σ_{abs} due to the geometric optical limits at $r_{ef} \approx 50 \ \mu m$, which shows the universal (independent of r_{ef}) monotonic dependence of σ_{abs} on $\overline{r_c}$ and $\overline{r_p}$ for ice 6 7 grains with $r_{ef} \ge 50 \ \mu m$ (Velesco et al., 1997). As a result, the spectral absorption 8 coefficient of snow containing polydispersed ice grains can be obtained using our 9 results for a monodispersed model. Interestingly, σ_{abs} did not depend on $\overline{r_c}$ when $\overline{r_c}$ 10 <0.75 and decreased almost linearly at higher $\overline{r_c}$ values (Figure 6a); this phenomenon 11 can be explained by geometric optical effects (Mackowski et al., 1990). However, σ_{abs} 12 was significantly affected by the dust mass concentration; for example, σ_{abs} at 500 nm 13 was decreased by 28%, 32%, and 32% from its maximum value (0.08, 0.38, and 3.71 m⁻¹) for dust concentrations of 2, 10, and 100 ppm, respectively, when $\bar{r_c}$ increased 14 15 from <0.75 to 1.0. The monotonic increase in σ_{abs} with the relative thickness of the polluted ice grain surface layer (i.e., $\overline{r_p}$) was also noteworthy (Figure 6b). The 16 17 core/shell Mie theory calculations for ice grains with a thin surface layer ($\overline{r_p} = 0.01$) 18 gave almost the same σ_{abs} as that obtained for the EDM. As a result, the σ_{abs} 19 increased rapidly with $\bar{r_p} < 0.4$ and then increased more slowly until $\bar{r_p} = 1$, which 20 corresponds to IDM (uniform).

21 The α_{λ} at 500 nm as a function of $\overline{r_c}$ and $\overline{r_p}$ with different dust concentrations and r_{ef} , is illustrated in Figure 6c and 6d. In general, α_{λ} at 500 nm decreased with 22 23 increasing dust mass concentration and r_{ef} , the effect of grain radius can be explained 24 by the fact that the snow extinction coefficient is inversely proportional to r_{ef} , so that 25 for a given amount of dust, the single-scattering albedo of the snow-dust mixture is 26 smaller for large snow grains (Gardner and Sharp, 2010). For a given dust mass 27 concentration and r_{ef} , the α_{λ} at 500 nm increased from its minimum value with $\bar{r_c}$ 28 <0.75 to the maximum value with $\bar{r_c} = 1$, corresponding to the findings of IDM

1 (uniform). For example, at dust concentrations of 2, 10, and 100 ppm, and a fixed r_{ef} of 2 100 µm, the α_{λ} at 500 nm increased by 0.2%, 0.7%, and 0.8%, respectively, when $\bar{r_c}$ 3 increased from <0.75 to 1. When the r_{ef} was fixed at 500 µm, the α_{λ} at 500 nm 4 increased by 1.9%, 2.5%, and 6.2% at dust concentrations of 2, 10, and 100 ppm, 5 respectively. Conversely, the α_{λ} at 500 nm decreased from its maximum value when $\overline{r_p} = 0.01$ (similar to EDM) to the minimum value with $\overline{r_p} = 1$, corresponding to the 6 case of IDM (uniform). For example, the α_{λ} at 500 nm decreased by 0.6%, 1.4%, and 7 8 1.4% when $\bar{r_p}$ increased from 0.01 to 1 for dust concentrations of 2, 10, and 100 ppm, 9 respectively, and a fixed r_{ef} of 100 µm, whereas for a r_{ef} of 500 µm, α_{λ} decreased by 10 3.3% (2 ppm dust), 4.7% (10 ppm), and 10.1% (100 ppm). These results indicate that dust mass concentrations and r_{ef} can amplify the influence of $\overline{r_c}$ or $\overline{r_p}$ on snow albedo. 11 12 Moreover, the effect of dust mass concentration on snow albedo is similar to r_{ef} . For 13 example, dust mass concentrations of 10 and 100 ppm and r_{ef} of 100 and 50 µm gave similar α_{λ} at 500 nm to dust mass concentrations of 2 and 10 ppm and r_{ef} of 500 μ m. 14 15 According to this result, spectral albedo measurements at a single wavelength are 16 insufficient to obtain the mass fraction of dust in snow cover because the same effect 17 can also be explained by a combination of different ice grain sizes and a non-uniform 18 distribution of dust inside the grains. It means that additional information is needed to 19 determine accurate dust mass concentrations. This may be a set of measurements at 20 various wavelengths in the VIS and NIR spectral ranges.

21 **3**

3.3 Effects on broadband snow albedo

Compared with the spectral optical properties, broadband results can provide more general knowledge for the relevant research community. Figure 7 shows the spectrally weighted α_{λ} ($\alpha_{integrated}$) over 300–2500 nm of a typical surface solar spectrum at mid to high latitudes, which is comparable with previous studies (Dang et al., 2017; Wang et al., 2017). Because the results of IDM (peripheral) effects on snow albedo fell between results from EDM and IDM (central), we do not consider the case of IDM (peripheral) in the following discussion. Instead, we focus on the effects of dust mass

1 concentration and r_{ef} on broadband snow albedo for EDM and IDM (uniform, central). 2 Similar to α_{λ} , $\alpha_{integrated}$ generally decreased with increasing dust mass concentrations and r_{ef} such that $\alpha_{integrated}$ declined more for internal mixing than 3 4 external mixing. $\alpha_{integrated}$ showed ranges of 0.60–0.92, 0.54–0.92, and 0.51–0.92 for 5 EDM, IDM (uniform), and IDM (central, $\bar{r_c}$ <0.75), respectively, with dust mass 6 concentrations of 0–100 ppm and r_{ef} of 50–1000 µm. For a given dust mass 7 concentration and r_{ef} , $\alpha_{integrated}$ for IDM (uniform) was smaller than EDM, which is 8 due to higher light absorption in the UV and VIS bands for IDM (uniform) relative to 9 EDM (Figure 4a). While $\alpha_{integrated}$ for IDM (uniform) was larger compared with 10 IDM (central, $\bar{r_c} < 0.75$), this can be attributed to the fact that radiation is focused near 11 the center of an ice grain with IDM (central) (Ackerman and Toon, 1981; Bohren, 1986), 12 enabling further absorptions from inclusions near the center of a grain due to the lensing 13 effect (Mackowski et al., 1990). For example, $\alpha_{integrated}$ (dust concentration of 20 14 ppm, ref of 500 µm) was 0.73 for IDM (uniform), less than EDM of 0.76, but higher 15 than IDM (central, $\overline{r_c} < 0.75$) of 0.72.

To quantify the effects of IDM on broadband snow albedo relative to EDM, we defined 16 a broadband snow albedo scaling factor ($E_{\alpha, integrated}$), which refers to the ratio of 17 18 $\alpha_{\text{integrated}}$ of IDM to EDM. Generally, for dust mass concentrations from 0 to 100 ppm 19 and r_{ef} of 50–1000 µm, $E_{\alpha, \text{ integrated}}$ varied from 0.89 to ~1.00 for IDM (uniform) (Figure 20 8a) and from 0.85 to ~1.00 for IDM (central, $\overline{r_c}$ <0.75) (Figure 8b). $E_{\alpha, integrated}$ 21 decreased significantly with increasing dust mass concentration and r_{ef} . In addition, E_{α} , 22 integrated for IDM (central, $\bar{r_c} < 0.75$) was smaller than IDM (uniform). These results 23 have implications for the effects of IDM in real environments. For example, IMC has 24 typical dust concentrations of ~10 ppm (Wang et al., 2013), so E_{α} , integrated (*r_{ef}* of 50– 25 1000 μ m) was 0.96–0.99 and 0.95–0.99 for IDM (uniform) and IDM (central, $\bar{r_c}$ < 26 (0.75), respectively. In contrast, dust concentrations are typically ~100 ppm in the TP 27 (Ming et al., 2016; Li et al., 2017, 2018), so $E_{\alpha, \text{ integrated}}$ ranged from 0.89 to 0.98 and 0.85 to 0.96 for IDM (uniform) and IDM (central, $~\bar{r_c}~$ <0.75), respectively. The results 28

1 show that IDM (uniform) and IDM (central, $\bar{r_c}$ <0.75) reduced broadband snow 2 albedo by ~2.5% and ~3.0%, respectively, in clean snow and ~6.5% and ~9.5%, 3 respectively, in polluted snow relative to EDM. Moreover, the sensitivity of $E_{\alpha, integrated}$ 4 to mineral dust decreased with increasing dust concentrations. For example, the 5 difference in $E_{\alpha, \text{ integrated}}$ (*r_{ef}* of 500 µm) between dust concentrations of 10 and 20 ppm 6 was 0.011 and 0.015 for IDM (uniform) and IDM (central, $\bar{r_c}$ <0.75), respectively, 7 while the corresponding differences between dust concentrations of 90 and 100 ppm 8 were only 0.004 and 0.005. These results provide a convenient method to calculate the 9 albedo of IDM when the albedo of EDM has been obtained for a given dust mass 10 concentration and r_{ef} .

11 **3.4 Uncertainties**

12 Although we calculated the imaginary RI values of dust using previous studies (section 2.2), there are still large variations which strongly depend on dust composition (e.g., 13 14 hematite/iron content) (Balkanski et al., 2007; Wagner et al., 2012). To roughly account 15 for this, we estimated the influence of chosen imaginary RI values on spectrally 16 weighted snow albedo ($E_{\alpha, integrated}$) by increasing and decreasing the calculated 17 imaginary RI values by 50%. These changes in imaginary RIs are plausible because 18 they are consistent with other studies (McConnell et al., 2010; Wagner et al., 2012). 19 The results showed that $E_{\alpha, integrated}$ uncertainties attributed to the imaginary RIs of dust were $\pm 3.9\%$ and $\pm 5.2\%$ for IDM (uniform) and IDM (central, $~\overline{r_c}~<0.75),$ respectively. 20 21 In contrast, observations have displayed large variations in the size distribution of dust 22 in the atmosphere and snow, and this variation is strongly affected by the dust source 23 and transport (Mahowald et al., 2014; Shao et al., 2011). In our standard simulation, we 24 assumed a log-normal dust size distribution with a geometric mean diameter of 0.65 25 μ m and a standard deviation of 2.0 (equivalent to an effective radius of 1.1 μ m), which 26 is typical for dust transported long-range (Formenti et al., 2011; Maring et al., 2003); nearer sources of dust tend to be larger (Kok, 2011). Therefore, we investigated the 27 28 effects of dust particle size on our results by assuming another two log-normal size

1 distributions with effective radii of 2.5 µm and 5.0 µm, which were within the observed 2 size ranges in the atmosphere and snow and comparable with previously analyzed dust 3 particle sizes (Maring et al., 2003; Shao and Mao, 2016; Zhang et al., 2003). The results 4 showed that the uncertainty of $E_{\alpha, integrated}$ attributed to dust diameter was $\pm 6.1\%$ for both IDM (uniform) and IDM (central, $\bar{r_c}$ <0.75), but it should be emphasized that the 5 6 uncertainty of $E_{\alpha, integrated}$ induced by dust sizes comes purely from the uncertainty of 7 broadband albedo of dust-snow external mixing due to different dust sizes. This is 8 because effective complex refractive indices of dust-snow internal mixture are 9 independent of dust particle size (Eq. 8). Overall, the total uncertainty of $E_{\alpha, integrated}$ 10 from variations of imaginary RIs and dust diameter was $\pm 11.0\%$ and $\pm 11.2\%$ for IDM (uniform) and IDM (central, $\bar{r_c} < 0.75$), respectively. 11

12 **3.5 Measurement-based estimate of the effects of dust on snow albedo**

13 Finally, widespread dust concentrations in snow across the Northern Hemisphere were 14 obtained to assess the effects of dust on snow albedo in real snowpack. Figure 9 shows 15 measured dust concentrations in snow in different regions; dust concentrations spanned 16 a broad range of values because of spatial and temporal variations in emissions, 17 transportation, and deposition among the different regions. Dust concentrations widely 18 varied from ~3 ppm to ~600 ppm, with the highest concentration in NAM and lowest 19 in the EP (Di Mauro et al., 2015; Lim et al., 2014; Painter et al., 2012; Reynolds et al., 20 2020). However, snow samples collected in the days after a significant dust transport 21 event showed that dust concentrations in snow can be up to ~70 ppm in the EP (Di 22 Mauro et al., 2015). Additionally, the average dust concentrations in fresh snow were 23 18, 6, and 28 ppm in the southern TP, central TP, and northern TP, respectively, similar 24 to the IMC (12 ppm) (Wang et al., 2013). However, dust concentrations in aged snow 25 (120, 300, and 140 ppm) were one to two orders of magnitude higher than in fresh snow, 26 indicating the important correlation between snow type and dust concentration (Zhang 27 et al., 2017, 2018).

28 We calculated the broadband snow albedo for EDM, IDM (uniform), and IDM (central,

1 $\overline{r_c}$ <0.75) based on the measured dust concentrations (Figure 10). The results showed 2 that broadband snow albedo decreased by 0.8%, 1.4%, and 1.6% in the EP for EDM, IDM (uniform), and IDM (central, $\bar{r_c}$ <0.75), respectively, which was similar to SJ. 3 4 However, the broadband snow albedo decreased by up to 5.6%, 8.1%, and 9.4% in the 5 EP after a significant dust transport event, indicating strong snow albedo reduction 6 during these events. In addition, broadband snow albedo was reduced by 2.0%, 3.1%, and 3.6% in IMC for EDM, IDM (uniform), and IDM (central, $~\bar{r_c}~$ <0.75), respectively. 7 8 Similar results were also found for the southern TP, central TP, and northern TP where 9 the broadband snow albedo for fresh snow was reduced by 2.5%, 1.4%, and 3.3%, 10 respectively, for EDM, 3.9%, 2.1%, and 5.1% for IDM (uniform), and 4.5%, 2.4%, and 11 5.7% for IDM (central, $\overline{r_c}$ <0.75). However, the broadband snow albedo was more 12 significantly reduced for aged snow: up to 6.0%, 8.1%, and 7.5% for EDM, 9.5%, 13 11.6%, and 10.5% for IDM (uniform), and 10.9%, 13.2%, and 12.3% for IDM (central, $\overline{r_c}$ <0.75) in the southern, central, and northern TP, respectively. This indicates that the 14 15 effects of dust on snow albedo showed stronger reductions during snowmelt periods, it 16 is worth noting that the effect of dust on snow albedo for aged snow could be 17 underestimated due to the larger snow grain effective radius (r_{ef}) than fresh snow. 18 Moreover, the largest broadband snow albedo reductions were found in NAM with 19 ranges of 9.8%-17.6%, 13.9%-24.1%, and 15.9%-27.0% for EDM, IDM (uniform), and IDM (central, $\bar{r_c}$ <0.75), respectively. These results suggest that the effects of 20 21 external or internal dust-snow mixing on snow albedo are particularly significant for 22 the TP and NAM regions, with stronger reductions in albedo. Therefore, these results 23 can have significant impacts on both local hydrological cycles and regional climate 24 change (Oaida et al., 2015; Xie et al., 2018).

25 **4. Discussion**

In this study, the application of the effective medium approximation greatly simplifies the complexity of snow radiation transfer calculation for dust–snow internal mixing, and the effect of non-uniform distribution of dust in snow grains on snowpack optical

1 properties are explicitly quantified for the first time. However, it is worth noting that 2 this method has its limitation when applying to large particles (e.g., dust) in snow 3 (Bohren 1986; Flanner et al., 2012), which can create some errors for the albedo 4 calculation of dust-snow internal mixing. To verify the credibility of our results, we 5 carefully make a comparison with the more rigorous calculations found in He et al. 6 (2019), which used the geometric-optics surface-wave approach (GOS) to consider the 7 impact of dust-snow uniform internal mixing on snow albedo reduction. As shown in 8 Figure 11, the results show that the enhancement ratio of snow albedo reduction (1.28) 9 due to dust-snow internal mixing (relative to external mixing) is slightly higher than 10 the value (1.16) reported by He et al. (2019), and this deviation is comparable to that 11 caused by snow nonsphericity (He et al., 2019). Therefore, we indicated that the 12 effective medium approximation used in this study is reasonable and reliable.

Over the past few decades, the effects of dust in snow on reductions in albedo have 13 14 been widely demonstrated (Skiles et al., 2018; Zhang et al., 2018). However, the 15 magnitude of these effects has only been studied in a few regions, and uncertainties still remain. Our study indicates that the albedo of dust-contaminated snowpack can be 16 17 affected by the dust-ice mixing state. In particular, IDM enhanced light absorption and 18 reduced snow albedo more significantly compared with EDM. For example, in IMC 19 and the TP, IDM reduced snow albedo by \sim 5% relative to EDM at a typical dust mass 20 concentration of 20 ppm and a snow grain radius of 500 µm. This exceeds the 21 contribution of BC to snow light absorption over most areas of North America and the 22 Arctic (Dang et al., 2017). In addition, the effects of IDM on snow albedo were 23 amplified by higher dust mass concentrations and larger snow grain sizes. We therefore 24 strongly suggest that IDM must be considered in future climate models, particularly to 25 more accurately evaluate the climate in areas where snowpack is heavily contaminated 26 with dust and is experiencing melting.

The mixed state between dust and snow gradually progresses from partial externalmixing to wholly internally mixed. Therefore, assuming a completely external mixing

1 of dust and snow grains will underestimate the effects of dust on snow albedo and 2 radiative forcing in numerical models (e.g., Dang et al., 2015; Nagorski et al., 2019). 3 Similarly, assuming completely internal mixing of dust and snow grains will 4 overestimate the effects of dust (e.g., He et al., 2019; Liou et al., 2014). Zhao et al. 5 (2014) underestimated the effects of dust by treating wet-deposited dust as externally 6 mixed with snow grains. In future studies, we recommend the actual ratio between 7 external and internal mixing for dust in snow be examined with an environmental 8 scanning electron microscope equipped with a cold stage.

9 **5.** Conclusions

10 In this study, the effects of dust particles on absorption coefficients and snow albedo 11 were theoretically analyzed by combining asymptotic radiative transfer theory and 12 (core/shell) Mie theory. We initially considered external mixing – when dust is present between ice grains – and variations of internal mixing of dust within ice grains. We 13 14 found that snow spectral absorption coefficients of IDM were larger than EDM across 15 UV to NIR wavelengths, but were negligible at wavelengths >1000 nm. The absorption 16 enhancement (relative to EDM) was wavelength-dependent and increased with 17 increased dust concentrations.

18 Compared with a uniform distribution of dust particles in ice grains, our calculations 19 showed that non-uniformly distributed dust particles may lead to significantly different 20 snow spectral absorption coefficients in the VIS band. Snow spectral absorption 21 coefficients were further increased when all of the dust was positioned in the central 22 part of ice grains, while the maximum absorption coefficient was found when the radius 23 of a dust-polluted core was <75% of the ice grain radius. In contrast, snow spectral 24 absorption coefficients decreased when all of the dust was positioned in the surface 25 layer of ice grains, and the minimum absorption coefficient was observed in the thin 26 surface layer of dust-polluted ice grains, which was similar for EDM. As a result, 27 broadband snow albedo decreased by up to 21%, 30%, and 33% for EDM, IDM (uniform), and IDM (central, $\bar{r_c} < 0.75$), respectively, at dust concentrations of 100 28

1 ppm and r_{ef} of 1000 μ m.

Based on comprehensive field measurements across the Northern Hemisphere, the effect of dust on snow albedo in real snowpack was evaluated by assuming external and internal dust–snow mixing. The largest reductions in broadband snow albedo were in NAM because that region had the highest average dust concentrations; IDM (uniform) and IDM (central, $\bar{r_c} < 0.75$) further decreased snow albedo by 4.6%–7.8% and 6.8%– 11.4%, respectively, compared with EDM. This implies an important influence of internal dust–snow mixing in NAM.

9 Data availability

The code of (core/shell) Mie theory used in this study can be found at http://gwest.gatsinc.com/software/software page.html.

12 Author contributions

WX designed the study and evolved the overarching research goals and aims. STL wrote the first draft with contributions from all co-authors. STL and CJC applied formal techniques such as statistical, mathematical, and computational to analyze study data. CY and XXY collected the dust measurements across the Northern Hemisphere. ZY and PW provided the majority of the methodology and software. CQL and ZXL provided technical guidance. All authors contributed to the improvement of results and revised the final paper.

20 **Competing interests**

21 The authors declare that they have no conflict of interest.

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Figure 1. Schematic of dust mixing with snow grains internally. Dust tends to mix externally with snow grains through dry deposition and/or below cloud scavenging, while dust–snow internal mixtures can be produced by nucleation, accretion, riming, aggregation, and sintering during aerosol–cloud–precipitation processes known as wet deposition. Arrows represent how the absorption (red) and reflection (yellow) of incoming sunlight changes with dust–snow mixing state.



Figure 2. Schematic depicting various mixing scenarios of snow grains and dust

- particles.





Figure 3. Imaginary part of the spectral complex refractive indices of ice (k_{ice}) and dust (k_{dust}) (Warren and Brandt, 2008; Dang et al., 2015), with the imaginary part of the effective complex refractive indices (k_{ef}) as a function of wavelength, at dust mass concentrations (C_{dust}^*) of 1–100 ppm (or $\mu g g^{-1}$) in snow.



Figure 4. Snow absorption coefficients (σ_{abs}) for dust-snow (a) external and internal mixing (uniform), (b) internal mixing (central), and (c) internal mixing (peripheral), as a function of wavelength with different dust concentrations and \bar{r}_c and \bar{r}_p . The corresponding enhancement ($E_{\sigma_{abs}}$) caused by (d) internal mixing (uniform), (e) internal mixing (central), and (f) internal mixing (peripheral) relative to external mixing, is shown as a function of wavelength. The snow grain radius was assumed to be 200 µm.



2 Figure 5. Same as Figure 4, but for spectral snow albedo (α_{λ}) .





Figure 6. The snow absorption coefficient (σ_{abs}) at 500 nm wavelength as a function of $\overline{r_c}$ and $\overline{r_p}$ for (a) internal mixing (central) and (b) internal mixing (peripheral) with different snow grain radii and dust mass concentrations. (d) and (e) are the same as (a) and (b), but for snow albedo at 500 nm wavelength.



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Figure 7. Broadband snow albedo ($\alpha_{integrated}$) variations affected by different dust mass concentrations and snow grain radii for (a) external mixing, (b) internal mixing (uniform), and (c) internal mixing (central, $\bar{r_c} < 0.75$).



Figure 8. Variations in the broadband snow albedo scaling factor ($E_{\alpha, \text{ integrated}}$, ratio of $\alpha_{\text{integrated}}$ for IDM to EDM) due to different dust mass concentrations and snow grain radii for (a) internal mixing (uniform) and (b) internal mixing (central, $\bar{r_c} < 0.75$).



Figure 9. In situ measurements of dust concentrations in snow (fresh snow, aged snow,

- 3 and snowpit from field sampling in different regions of the Northern Hemisphere. The
- 4 solid black circle represents snow samples that were collected days after a significant
- 5 dust transport event.
- 6



Figure 10. Calculated broadband snow albedo based on dust concentration measurements in different areas for dust-snow external mixing, internal mixing (uniform), and internal mixing (central, $\bar{r_c} < 0.75$). The dashed line represents broadband albedo of pure snow, and the snow grain radius was assumed to be 200 µm.



Figure 11. Comparisons of snow albedo reduction (Δα) under the cloudy sky caused
by dust-snow uniform internal mixing (y-axis) and external mixing (x-axis) for this
study (blue) and He et al. (2019) (red).