

1 **Enhanced light absorption and reduced snow albedo due to**
2 **internally mixed mineral dust in grains of snow**

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1 **Abstract.** Mineral dust is a major light-absorbing aerosol, which can significantly
2 reduce snow albedo and accelerate snow/glacier melting via wet and dry deposition on
3 snow. In this study, three scenarios of internal mixing of dust in ice grains were
4 analyzed theoretically by combining asymptotic radiative transfer theory and
5 (core/shell) Mie theory to evaluate the effects on absorption coefficient and ~~snow~~
6 albedo of the semi-infinite snowpack consisting of spherical snow grains. In general,
7 snow albedo was substantially reduced at wavelengths of $<1.0 \mu\text{m}$ by internal dust–
8 snow mixing, with stronger reductions at higher dust concentrations and larger snow
9 grain sizes. Moreover, calculations showed that a non-uniform distribution of dust in
10 snow grains can lead to significant differences in the values of the absorption coefficient
11 and ~~snowpack~~-albedo of dust-contaminated snowpack at visible wavelengths relative
12 to a uniform dust distribution in snow grains. Finally, using comprehensive in situ
13 measurements across the Northern Hemisphere, we found that broadband snow albedo
14 was further reduced by 5.2% and 9.1% due to the effects of internal dust–snow mixing
15 on the Tibetan Plateau and North American mountains. This was higher than the
16 reduction in snow albedo caused by black carbon in snow over most North American
17 and Arctic regions. Our results suggest that significant dust–snow internal mixing is
18 important for the melting and retreat of Tibetan glaciers and North American mountain
19 snowpack.
20

1 **1. Introduction**

2 Snow cover is one of the most reflective surfaces in the Earth system, and plays a crucial
3 role in the atmospheric solar radiation energy budget via snow albedo feedbacks (Di
4 Mauro et al., 2020; Flanner et al., 2011; Jacobson, 2004; Usha et al., 2020; Xie et al.,
5 2018). Previous studies have shown that light-absorbing particles (LAPs) effectively
6 reduce snow albedo and enhance the absorption of solar radiation after deposition, these
7 studies were based on in situ observations and model simulations (Casey et al., 2017;
8 Hadley and Kirchstetter, 2012; Shi et al., 2020; Warren and Wiscombe, 1980; Yasunari
9 et al., 2012; Yasunari et al., 2015). As a result, snow contaminated with LAPs shows
10 significant changes in morphology (Niwano et al., 2012; Rango et al., 1996), chemistry
11 (France et al., 2012; Reay et al., 2012), hydrology (Matt et al., 2018; Qian et al., 2015;
12 Rahimi et al., 2019), snowmelt rate (Kaspari et al., 2015; Warren, 1984;), and radiative
13 properties (Grenfell et al., 2002; Hansen and Nazarenko, 2004; Zhao et al., 2014).

14 Numerous studies have assessed the potential effects of LAPs, such as black carbon
15 (BC) and mineral dust, on snow albedo by assuming that LAPs mixed outside spherical
16 snow grains (i.e., external mixing) (Flanner et al., 2007; Kokhanovsky, 2013; Libois et
17 al., 2013; Wang et al., 2017; Warren and Wiscombe, 1980). For example, Warren and
18 Wiscombe (1980) calculated snow spectral albedo by solving a radiative transfer
19 equation using Mie theory and δ -Eddington approximations, and found that 10–100 ng
20 g^{-1} of BC and 1–10 $\mu\text{g g}^{-1}$ of dust in old snow decreased albedo by 1%–7% and 2%–
21 10% at 400 nm wavelength, respectively. Flanner et al. (2007) pointed out that the
22 reduction in snow albedo for 1000 ng g^{-1} BC in snow was 0.045 (0.17) with a 50 (1000)
23 μm snow grain radius (R_{ef}) based on the Snow, Ice, and Aerosol Radiation (SNICAR)
24 radiative transfer model. This model utilizes theory from Wiscombe and Warren (1980)
25 and the two-stream radiative transfer solution from Toon et al. (1989). Wang et al. (2017)
26 developed a Spectral Albedo Model for Dirty Snow (SAMDS) based on asymptotic
27 radiative transfer theory, which is a function of the snow grain radius, LAP (e.g., BC,
28 dust) mixing ratios, and the mass absorption coefficients (MACs) of LAPs. Their results

1 revealed that broadband snow albedo decreased 0.03 and 0.003 due to 200 ng g⁻¹ of BC
2 and 2 μg g⁻¹ of dust in snow, respectively, with a R_{ef} of 200 μm. Additionally, recent
3 studies found that snow nonsphericity can interact with LAP-snow mixing, which leads
4 to weaker LAP-induced albedo reductions for nonspherical snow shapes than snow
5 spheres (e.g., Dang et al., 2016; He et al., 2018, 2019).

6 Recently, direct snowpack observations have shown evidence for the existence of LAP–
7 snow internal mixing (e.g., Horhold et al., 2012; Spaulding et al., 2011). LAPs tend to
8 mix externally with snow grains through dry deposition and/or below cloud scavenging,
9 while internal LAP–snow mixing can be produced by nucleation, accretion, riming,
10 aggregation, and sintering during aerosol–cloud–precipitation processes (i.e., wet
11 deposition; Figure 1) (Flanner et al., 2012). Furthermore, Flanner et al. (2012) found
12 that internal BC/ice mixing (IBM) increased the absorption of snowpack by a factor
13 1.8–2.1 relative to external BC/ice mixing (EBM). He et al. (2018) indicated that IBM
14 enhanced the mean snow albedo reduction over the Tibetan plateau by 30%–60%
15 relative to EBM, based on the updated SNICAR model. Additionally, Dombrovsky and
16 Kokhanovsky (2020) demonstrated that non-uniform BC distribution in ice grains may
17 lead to significantly different absorption coefficients and snowpack albedo in visible
18 (VIS) wavelengths.

19 Numerous studies have addressed the role of IBM in enhancing the absorption of
20 snowpack due to its strong absorption effect relative to other LAPs (Dombrovsky and
21 Kokhanovsky, 2020; Flanner et al., 2012; He et al., 2018; Liou et al., 2011). In contrast,
22 few studies have considered the effects of internal dust/ice mixing (IDM) in snowpack.

23 Liou et al. (2014) is the pioneer to investigate the dust-snow internal mixing effects
24 based on the geometric-optics surface-wave approach. Subsequently, He et al. (2019)
25 used the same method to explicitly quantify the combined effects of dust-snow internal
26 mixing and snow grain nonsphericity on snow optical properties, thereafter, develop a
27 set of new dust-snow parameterizations for land/climate modeling applications for the
28 first time. Actually, dD dust particles are generally larger than BC, and act as more

1 efficient ice nuclei, showing a better ability to influence cloud formation and
2 precipitation (Creamean et al., 2013; Huang et al., 2014). Therefore, they are more
3 likely to mix internally with ice grains. Furthermore, dust can also dominate light
4 absorption and effectively decrease snow albedo because of its relatively high mass
5 abundance (ppm) in snowpack, especially in areas with seasonal and patchy snow cover
6 or mountainous regions (Di Mauro et al., 2015; Gabbi et al., 2015; Painter et al., 2012;
7 Reynolds et al., 2020; Xie et al., 2018).— Therefore, it is important to account for IDM
8 when estimating the impact of dust deposition on snow albedo.

9 In this study, we assess the effects of external/internal mixing of dust with ice grains on
10 the snowpack absorption coefficient and albedo using asymptotic radiative transfer
11 theory and Mie theory. In addition, the uniformity and nonuniformity of dust particle
12 distribution inside ice grains are considered for IDM based on the effective medium
13 approximation, and the combined effects of dust content and snow grain radius on snow
14 albedo are quantified. A schematic of various dust spatial distributions from this study
15 is presented in Figure 2. We further discuss snow albedo sensitivity to complex
16 refractive indices and dust particle size distribution. Based on a comprehensive set of
17 field measurements of dust concentrations, we estimate the reductions in snow albedo
18 by dust external/internal mixing with ice grains across the Northern Hemisphere.

19 **2 Methods**

20 **2.1 External mixing model**

21 For fairly pure snow, semi-infinite means semi-infinite is generally defined as
22 absorptions of about 20 cm in the VIS and 3 cm in the near-infrared (NIR) regions of
23 snow depths, respectively (Zhou et al., 2003). For a semi-infinite snow layer under
24 diffuse illumination conditions, albedo can be calculated using an asymptotic analysis
25 of radiative transfer theory, which is valid in small absorptions (Kokhanovsky and Zege,
26 2004; Zege et al., 1991):

$$27 \alpha_{\lambda} = \exp(-4S_{\lambda}) \quad (1)$$

1 where α_λ is the spectral snow albedo, λ is the wavelength, S_λ is the similarity
 2 parameter, and

$$3 \quad S_\lambda = \sqrt{\frac{\sigma_{abs}}{3\sigma_{ext}(1-g)}} \quad (2)$$

4 In Eq. (2), σ_{abs} and σ_{ext} are the absorption and extinction coefficients, respectively,
 5 and g is the asymmetry parameter (the average cosine of the phase function of the
 6 medium).

7 According to Eq. (18) and (25) in Kokhanovsky and Zege (2004), the extinction
 8 coefficients of particles can be expressed as:

$$9 \quad \sigma_{ext} = \frac{l_{tr}}{l_{tr}(1-g)} = \frac{3C_v}{2r_{ef}} \quad (3)$$

10 where l_{tr} is the photon transport path length, $C_v = \rho_{snow}/\rho_{ice}$ is the volumetric
 11 snow particle concentration, and the values $\rho_{ice} = 916.7 \text{ kg m}^{-3}$ and $\rho_{snow} = 300 \text{ kg}$
 12 m^{-3} are used in subsequent calculations. r_{ef} is the effective snow grain radius, which
 13 is equal to the radius of the volume-to-surface equivalent sphere ($r_{ef} = \frac{3\bar{V}}{4\bar{A}}$) where \bar{V}
 14 and \bar{A} are the average volume and cross-sectional (geometric shadow) area of snow
 15 grains, respectively.

16 For external dust/ice mixing (EDM) in a dust-contaminated snowpack, the total
 17 absorption coefficient (σ_{abs}) can be derived from the absorptions by snow (σ_{abs}^{snow}) and
 18 dust (σ_{abs}^{dust}):

$$19 \quad \sigma_{abs} = \sigma_{abs}^{snow} + \sigma_{abs}^{dust} \quad (4)$$

20 For example, consider a hypothetical case of snow composed of monodispersed,
 21 spherical grains of ice. Although non-spherical snow grains lead to a slight increase in
 22 snow albedo, Dang et al. (2016) found that the albedo of a snowpack consisting of non-
 23 spherical snow grains can be mimicked by using smaller, spherical grains; thus, we do
 24 not consider the effect of non-spherical snow grains in this study. Therefore, we used
 25 the following equation for the absorption coefficient of snow (Dombrovsky and Baillis,
 26 2010):

$$27 \quad \sigma_{abs}^{snow} = \frac{0.75C_v \cdot Q_{abs}^{ice}}{r_{ef}} \quad (5)$$

1 where Q_{abs}^{ice} is the efficiency factor of absorption for a single ice grain, and the value
 2 of Q_{abs}^{ice} can be calculated for homogeneous spherical ice grains considered in classical
 3 Mie theory.

4 The absorption coefficient of dust (Aoki et al., 2000; Marley et al., 2001; Warren et al.,
 5 2006) is expressed as:

$$6 \quad \sigma_{abs}^{dust} = \frac{Q_{abs}^{dust} \cdot \pi \cdot (r_{ef}^{dust})^2}{\frac{4}{3} \pi \cdot (r_{ef}^{dust})^3} \cdot C_{dust} = \frac{3Q_{abs}^{dust}}{4r_{ef}^{dust}} \cdot C_{dust} = MAC_{abs}^{dust} \cdot \rho_{dust} \cdot C_{dust} \quad (6)$$

7 where Q_{abs}^{dust} and MAC_{abs}^{dust} is the absorption efficiency and MAC of dust,
 8 respectively, that can be obtained via Mie theory, and ρ_{dust} and r_{ef}^{dust} represent the
 9 density and effective dust radius, respectively. In this study, ρ_{dust} was assumed to be
 10 2500 kg m^{-3} (Zender et al., 2003). We also assumed a log-normal dust size distribution
 11 with a geometric mean diameter of $0.65 \text{ }\mu\text{m}$ and standard deviation of 2.0 (equivalent
 12 to an effective radius of $1.1 \text{ }\mu\text{m}$), which represents dust from large-scale transport
 13 (Formenti et al., 2011; Maring et al., 2003) that is likely smaller in size than from local
 14 soil (Kok, 2011). The effects of dust size on snow optical properties and albedo were
 15 further quantified through sensitivity simulations (see section 3.4). Dust volumetric
 16 concentrations (C_{dust}) are expressed as:

$$17 \quad C_{dust} = \frac{\rho_{snow} \cdot C_{dust}^*}{\rho_{dust}} \quad (7)$$

18 where C_{dust}^* is the mass concentration of dust in snow (kg kg^{-1}). Thus, the spectral
 19 albedo of dust-contaminated snow for EDM can be easily calculated with Eq. (1) to (7).

20 **2.2 Internal mixing model**

21 For the IDM (Figure 2), we first determined the effective optical constants of ice
 22 containing small dust particles via an effective medium approximation (Maxwell-
 23 Garnett and Larmor, 1904). According to this approach, the complex permittivity of a
 24 composite medium in ice grains can be calculated in terms of particle polarizability by
 25 applying the Lorentz-Lorenz formula (Koledintseva et al., 2009; Markel, 2016). We
 26 used the following relationships to calculate effective complex refractive indices (RIs),

1 $m_{ef} = n_{ef} - ik_{ef}$ at known values of $m_{ice} = n_{ice} - ik_{ice}$ for pure ice and $m_{dust} = n_{dust} - ik_{dust}$
 2 for dust:

$$3 \quad m_{ef}^2 = m_{ice}^2 \frac{2\delta_{dust}(m_{dust}^2 - m_{ice}^2) + m_{dust}^2 + 2m_{ice}^2}{2m_{ice}^2 + m_{dust}^2 - \delta_{dust}(m_{dust}^2 - m_{ice}^2)} \quad (8)$$

4 where δ_{dust} is the local dust fraction volume in an ice grain. We obtained the spectral
 5 complex RIs of ice from Warren and Brandt (2008) and the spectral complex RIs of
 6 dust from Dang et al. (2015). The imaginary part of the complex RIs of ice (k_{ice}) and
 7 dust (k_{dust}) associated with absorption is shown in Figure 3. We also evaluated the effect
 8 of dust on the imaginary part of the effective complex RIs (k_{ef}) assuming dust mass
 9 concentrations (C_{dust}^*) of 1–100 ppm (or $\mu\text{g g}^{-1}$) in snow.

10 In all variations of the spatial distribution of dust particles in snow, the dust mass
 11 concentration was assumed to be constant, which means that the local dust fraction
 12 volume may differ. In an example of spherical ice grains with uniformly distributed
 13 dust, the dust fraction volume in an ice grain was determined as:

$$14 \quad \delta_{dust}^0 = \frac{C_{dust}^* \rho_{ice}}{C_v \rho_{dust}} \quad (9)$$

15 where the ratio of C_{dust}^*/C_v is the mass fraction of dust in the ice grain.

16 We considered two cases of non-uniform dust distributed in a spherical ice grain with
 17 radius r_{ef} : (1) We assumed that the same mass of dust is uniformly distributed in the
 18 central part of the ice grain ($r_{ef}^{dust} < r_c \leq r_{ef}$). (2) We assumed all of the dust was in
 19 the surface layer of the ice grain ($r_{ef}^{dust} < r_p \leq r_{ef}$) (Figure 2). In both cases the local
 20 value of δ_{dust} increases as:

$$21 \quad \delta_{dust} = \delta_{dust}^0 / \psi \quad \psi = \begin{cases} \bar{r}_c^3 & \text{central pollution} \\ 1 - (1 - \bar{r}_p)^3 & \text{peripheral pollution} \end{cases} \quad (10)$$

22 where $\bar{r}_c = r_c/r_{ef}$ and $\bar{r}_p = r_p/r_{ef}$. $\bar{r}_c = 1$ and $\bar{r}_p = 1$ correspond to uniformly
 23 distributed dust when $\psi = 1$ and $\delta_{dust} = \delta_{dust}^0$, and $\psi < 1$ and $\delta_{dust} > \delta_{dust}^0$ in
 24 other cases. Obviously, dust particles increase the imaginary part of the effective
 25 complex RIs in polluted ice grains (Figure 3).

26 In summary, the m_{ef} of a spherical ice grain with uniformly and non-uniformly

1 distributed ~~dust can be calculated~~, according to Eq. (8) to (10), then their corresponding
2 absorption efficiencies can be obtained using classical Mie theory and core/shell Mie
3 theory, respectively. The spectral snow albedo for IDM can be easily calculated using
4 Eq. (1) to (5).

5 **2.3 Broadband snow albedo calculations**

6 The spectral albedo (α_λ) is integrated over the solar spectrum ($\lambda = 300\text{--}2500$ nm) and
7 weighted by incoming solar irradiance (E_λ) to calculate broadband snow albedo (Hadley
8 and Kirchstetter, 2012):

$$9 \quad \alpha_{integrated} = \frac{\int \alpha_\lambda E_\lambda d\lambda}{\int E_\lambda d\lambda} \quad (11)$$

10 Following the study of Dang et al. (2017), the incoming solar irradiance we used is a
11 typical surface solar spectrum at mid to high latitudes from January to March under the
12 cloudy sky, calculated by the Santa Barbara DISORT Atmospheric Radiative Transfer
13 (SBDART) model (Pu et al., 2019). The SBDART model is a widely used atmospheric
14 radiation transfer model based on a collection of highly developed physical models,
15 including the Discrete Ordinate Radiative Transfer module (Stamnes et al., 1988), low-
16 resolution atmospheric transmission models, and Mie theory. The SBDART model can
17 be used to compute radiative transfer at different heights and directions under both clear
18 and cloudy sky conditions. Details on the SBDART model can be found in Ricchiazzi
19 et al. (1998).

20 **2.4 Dust concentration measurements**

21 To estimate the effect of dust on snow albedo in real snowpack, we collected a
22 comprehensive set of in situ dust concentration measurements during field campaigns
23 to Inner Mongolia, China (IMC) (Wang et al., 2013), the Tibetan Plateau (TP) (Ming et
24 al., 2016; Li et al., 2017, 2018; Li et al., 2016, 2019; Niu et al., 2017; Qu et al., 2014;
25 Zhang et al., 2017, 2018), Sapporo, Japan (SJ) (Kuchiki et al., 2015), the European Alps
26 (EP) (Di Mauro et al., 2015; Lim et al., 2014), and North American mountains (NAM)
27 (Painter et al., 2012; Reynolds et al., 2020). The field campaigns conducted in the TP

1 can be further grouped into three regions, as in previous $\delta^{18}\text{O}$ precipitation studies (Yao
2 et al., 2013). These three distinct domains were associated with the Indian monsoon
3 (Southern TP), westerlies (Northern TP), and transition (Central TP) (Yao et al., 2013).
4 It is worth noting that we only considered regions with higher dust concentrations (>1
5 ppm), such that polar regions are not included in this study. Measurements of dust in
6 the snow samples were generally obtained by weighing the filter before and after
7 filtration using a microbalance.

8 **3. Results**

9 **3.1 Impact on the imaginary part of the effective complex RIs**

10 We evaluated the effect of dust on k_{ef} , including k_{ice} and k_{dust} associated with absorption
11 (Figure 3). The k_{dust} was in a narrow range ($\sim 0.001\text{--}0.01$) and gradually decreased with
12 increasing wavelength in the ultraviolet (UV) and VIS regions, then remained stable in
13 the NIR band. The k_{ice} varied by eight orders of magnitude from the UV ($\sim 10^{-11}$) to NIR
14 ($\sim 10^{-3}$) bands, and increased with increasing wavelength, except at $1.03\ \mu\text{m}$ where k_{ice}
15 decreased slightly as a result of the presence of ice absorption features (Warren, 2019).
16 Figure 3 also shows the k_{ef} with dust mass concentrations ranging from 1 to 100 ppm
17 and wavelengths of 300–1500 nm. The k_{ef} clearly varied depending on the wavelength
18 and increased with dust mass concentrations. For a given dust mass concentration, k_{ef}
19 decreased with wavelengths from UV to VIS, then increased from VIS to NIR. For
20 example, the value of k_{ef} decreased from 4.26×10^{-8} at 300 nm to 1.36×10^{-8} at 500
21 nm, then rose to 1.73×10^{-6} at 1000 nm at a dust concentration of 10 ppm. Moreover,
22 the wavelength of ~~the minimum of the valley~~ k_{ef} ~~value~~ varied from ~ 500 nm to ~ 650 nm
23 depending on the dust mass concentrations (1 to 100 ppm). Additionally, it is worth
24 noting that k_{ef} was not sensitive to dust mass concentrations in wavelengths >1000 nm,
25 which was generally consistent with k_{ice} , because the difference between k_{dust} and k_{ice} is
26 more than compensated by the much larger difference in ice and dust concentration ~~the~~
27 ~~order of magnitude of k_{ice} was comparable to k_{dust}~~ at those wavelengths. Conversely, k_{ef}

1 showed significant differences relative to k_{ice} in the UV and VIS regions, with higher
2 dust mass concentrations demonstrating larger differences. For example, k_{ef} was
3 enhanced by 3, 21, and 205 times at 500 nm relative to k_{ice} for dust mass concentrations
4 of 1, 10, and 100 ppm, respectively.

5 **3.2 Impact on spectral snow absorption coefficient and albedo**

6 Dust in snow effectively enhances the snow absorption coefficient, but its effect on the
7 snow asymmetry factor and extinction efficiency is negligible (He et al., 2019).
8 Therefore, we mainly focused on the effects of EDM and three cases of IDM (uniform,
9 central, and peripheral) on the snow absorption coefficient (σ_{abs}). Figure 4a displays the
10 σ_{abs} for EDM and IDM (uniform) as a function of wavelength at different dust
11 concentrations. We used a snow grain radius of 200 μm (Figure 4), which is comparable
12 to previous observations of seasonal snow at mid to high latitudes in winter (Shi et al.,
13 2020; Wang et al., 2017). The results showed that EDM and IDM have distinct impacts
14 on σ_{abs} in UV and VIS, but small effects at wavelengths >1000 nm, which is due to
15 the optical properties of snow being affected by LAPs in UV and VIS and primarily
16 affected by snow itself at wavelengths >1000 nm. Additionally, σ_{abs} increased with
17 increased dust mass concentrations. For instance, σ_{abs} increased from 0.007 m^{-1} (pure
18 snow) to 0.03, 0.14, and 1.37 m^{-1} at 500 nm with 2, 10, and 100 ppm of dust with EDM,
19 respectively. For IDM (uniform), σ_{abs} increased to 0.06, 0.28, and 2.80 m^{-1} at 500 nm
20 with 2, 10, and 100 ppm of dust, respectively, with corresponding enhancement factors
21 of σ_{abs} ($E_{\sigma_{abs}}$, defined as the absorption coefficient of IDM divided by EDM) were
22 1.84, 2.00, and 2.05. Furthermore, the σ_{abs} for two cases of non-uniform dust
23 distribution in a spherical ice grain ($r_{ref} = 200 \mu\text{m}$) ~~can be regarded as a function~~
24 ~~of~~ depends on the wavelength, dust mass concentrations, \bar{r}_c , and \bar{r}_p (Figure 4b and 4c).
25 We note that \bar{r}_c and \bar{r}_p values of 1 correspond to uniformly distributed dust, and the
26 σ_{abs} increases and decreases (with the decrease of \bar{r}_c and \bar{r}_p) for IDM (central) and
27 IDM (peripheral), respectively. For example, σ_{abs} increased by 29%, 32%, and 33%
28 (500 nm) at dust mass concentrations of 2, 10, and 100 ppm, respectively, when \bar{r}_c

1 ~~decreased from 1 to 0.7 with \bar{r}_e values of 1 to 0.7~~ for IDM (central). However, σ_{abs}
2 decreased by 41%, 44%, and 44% (500 nm) at dust mass concentrations of 2, 10, and
3 100 ppm, respectively, ~~when \bar{r}_p decreased from 1 to 0.1 with \bar{r}_p of 1 to 0.1~~ for IDM
4 (peripheral). This indicates that the IDM (central) further enhanced snowpack light
5 absorption compared with the IDM (uniform), while the IDM (peripheral) reduced
6 snowpack light absorption with a corresponding σ_{abs} between the value of σ_{abs} for
7 IDM (uniform) and EDM. Figure 4d–f quantitatively shows the spectral snow
8 absorption coefficient enhancement for IDM ($E_{\sigma_{\text{abs}}}$). The enhancement decreased
9 sharply with increasing wavelengths, then reduced to 1.0 (i.e., no enhancement) at
10 wavelengths longer than $\sim 1.0 \mu\text{m}$ because of strong dust absorption and weak snow
11 absorption at shorter wavelengths. Obviously, $E_{\sigma_{\text{abs}}}$ was affected by dust mass
12 concentration, \bar{r}_c , and \bar{r}_p , but $E_{\sigma_{\text{abs}}}$ was insensitive to dust mass concentration at
13 wavelengths $< 450 \text{ nm}$.

14 Figure 5a–c shows the spectral snow albedo (α_λ) for EDM and IDM; ~~α_λ IDM~~ was
15 consistent with σ_{abs} , with the effects mainly present at wavelengths $< 1000 \text{ nm}$.
16 Generally, α_λ decreased with increased dust mass concentrations in UV and VIS, and
17 IDM was shown to further trigger the reduction of α_λ . For example, for EDM α_λ was
18 $\sim 0.97, 0.95, 0.85$ (at 500 nm) for dust concentrations of 2, 10, 100 ppm, respectively,
19 which was higher than the values for IDM (uniform) ($\sim 0.96, 0.93, 0.79$, respectively).
20 Compared with IDM (uniform), the α_λ for IDM (central) decreased by 0.5%, 1.1%,
21 and 3.5% (at 500 nm) for dust concentrations of 2, 10, and 100 ppm, respectively, ~~when~~
22 ~~\bar{r}_c decreased from 1 to 0.7 with \bar{r}_e values of 1 to 0.7~~. α_λ for IDM (peripheral)
23 increased by 0.8%, 1.9%, and 6.2% (at 500 nm) for the same dust concentrations, ~~when~~
24 ~~\bar{r}_p decreased from 1 to 0.1 with \bar{r}_p from 1 to 0.1~~. Moreover, the wavelength of the maximum
25 α_λ value varied from $\sim 500 \text{ nm}$ to $\sim 650 \text{ nm}$ depending on the dust mass concentrations,
26 which is consistent with changes of k_{ef} . Figure 5d–f shows the ratio (E_{α_λ}) of snow
27 spectral albedo for IDM to EDM where we observed that the E_{α_λ} increased with
28 increasing wavelengths and dust concentrations, and then became stable at 1.0. This is

1 because IDM can enhance the light absorption of snowpack more effectively at shorter
2 wavelengths and higher dust concentrations (Figure 4). Additionally, the values of \bar{r}_c
3 and \bar{r}_p also have non-negligible effects on $E_{\alpha\lambda}$, which can be decreased and increased
4 with decreasing \bar{r}_c and \bar{r}_p , respectively.

5 We found that the optical properties of an ice grain containing uniformly distributed
6 dust in its center, or concentric surface layer, can be affected by \bar{r}_c or \bar{r}_p . To better
7 understand this effect, Figure 6a–b displays the σ_{abs} at 500 nm as a function of \bar{r}_c and
8 \bar{r}_p with different dust concentrations and r_{ef} . This demonstrates r_{ef} has negligible
9 effects on σ_{abs} due to the geometric optical limits at $r_{\text{ef}} \approx 50 \mu\text{m}$, which shows the
10 universal (independent of r_{ef}) monotonic dependence of σ_{abs} on \bar{r}_c and \bar{r}_p for ice
11 grains with $r_{\text{ef}} \geq 50 \mu\text{m}$ (Velesco et al., 1997). As a result, the spectral absorption
12 coefficient of snow containing polydispersed ice grains can be obtained using our
13 results for a monodispersed model. Interestingly, σ_{abs} did not depend on \bar{r}_c when \bar{r}_c
14 < 0.75 and decreased almost linearly at higher \bar{r}_c values (Figure 6a); this phenomenon
15 can be explained by geometric optical effects (Mackowski et al., 1990). However, σ_{abs}
16 was significantly affected by the dust mass concentration; for example, σ_{abs} at 500 nm
17 was decreased by 28%, 32%, and 32% from its maximum value (0.08, 0.38, and 3.71
18 m^{-1}) for dust concentrations of 2, 10, and 100 ppm, respectively, ~~with~~ when \bar{r}_c
19 increased from < 0.75 to 1.0. The monotonic increase in σ_{abs} with the relative thickness
20 of the polluted ice grain surface layer (i.e., \bar{r}_p) was also noteworthy (Figure 6b). The
21 core/shell Mie theory calculations for ice grains with a thin surface layer ($\bar{r}_p = 0.01$)
22 gave almost the same σ_{abs} as that obtained for the EDM. As a result, the σ_{abs}
23 increased rapidly with $\bar{r}_p < 0.4$ and then increased more slowly until $\bar{r}_p = 1$, which
24 corresponds to IDM (uniform).

25 The α_λ at 500 nm as a function of \bar{r}_c and \bar{r}_p with different dust concentrations and
26 r_{ef} , is illustrated in Figure 6c and 6d. In general, α_λ at 500 nm decreased with
27 increasing dust mass concentration and r_{ef} ; the effect of grain radius can be explained
28 by the fact that the snow extinction coefficient is inversely proportional to r_{ef} , so that

1 for a given amount of dust, the single-scattering albedo of the snow-dust mixture is
2 smaller for large snow grains~~increasing the forward scattering with grain size~~ (Gardner
3 and Sharp, 2010). For a given dust mass concentration and r_{ef} , the α_λ at 500 nm
4 increased from its minimum value with $\bar{r}_c < 0.75$ to the maximum value with $\bar{r}_c = 1$,
5 corresponding to the findings of IDM (uniform). For example, at dust concentrations
6 of 2, 10, and 100 ppm, and a fixed r_{ef} of 100 μm , the α_λ at 500 nm increased by 0.2%,
7 0.7%, and 0.8%, respectively, when \bar{r}_c increased with \bar{r}_c from < 0.75 to 1. When the
8 r_{ef} was fixed at 500 μm , the α_λ at 500 nm increased by 1.9%, 2.5%, and 6.2% at dust
9 concentrations of 2, 10, and 100 ppm, respectively. Conversely, the α_λ at 500 nm
10 decreased from its maximum value when $\bar{r}_p = 0.01$ (similar to EDM) to the minimum
11 value with $\bar{r}_p = 1$, corresponding to the case of IDM (uniform). For example, the α_λ
12 at 500 nm decreased by 0.6%, 1.4%, and 1.4% when with \bar{r}_p increased from 0.01 to 1
13 for dust concentrations of 2, 10, and 100 ppm, respectively, and a fixed r_{ef} of 100 μm ,
14 whereas for a r_{ef} of 500 μm , α_λ decreased by 3.3% (2 ppm dust), 4.7% (10 ppm), and
15 10.1% (100 ppm). These results indicate that dust mass concentrations and r_{ef} can
16 amplify the influence of \bar{r}_c or \bar{r}_p on snow albedo. Moreover, the effect of dust mass
17 concentration on snow albedo is similar to r_{ef} . For example, dust mass concentrations
18 of 10 and 100 ppm and r_{ef} of 100 and 50 μm gave similar α_λ at 500 nm to dust mass
19 concentrations of 2 and 10 ppm and r_{ef} of 500 μm . According to this result, spectral
20 albedo measurements at a single wavelength are insufficient to obtain the mass fraction
21 of dust in snow cover because the same effect can also be explained by a combination
22 of different ice grain sizes and a non-uniform distribution of dust inside the grains. It
23 means that additional information is needed to determine accurate dust mass
24 concentrations. This may be a set of measurements at various wavelengths in the VIS
25 and NIR spectral ranges.

26 **3.3 Effects on broadband snow albedo**

27 Compared with the spectral optical properties, broadband results can provide more
28 general knowledge for the relevant research community. Figure 7 shows the spectrally

1 weighted α_λ ($\alpha_{\text{integrated}}$) over 300–2500 nm of a typical surface solar spectrum at mid
2 to high latitudes, which is comparable with previous studies (Dang et al., 2017; Wang
3 et al., 2017). Because the results of IDM (peripheral) effects on snow albedo fell
4 between results from EDM and IDM (central), we do not consider the case of IDM
5 (peripheral) in the following discussion. Instead, we focus on the effects of dust mass
6 concentration and r_{ef} on broadband snow albedo for EDM and IDM (uniform, central).
7 Similar to α_λ , $\alpha_{\text{integrated}}$ generally decreased with increasing dust mass
8 concentrations and r_{ef} such that $\alpha_{\text{integrated}}$ ~~declined more for~~ internal mixing ~~declined~~
9 ~~more~~ than external mixing. $\alpha_{\text{integrated}}$ showed ranges of 0.60–0.92, 0.54–0.92, and
10 0.51–0.92 for EDM, IDM (uniform), and IDM (central, $\bar{r}_c < 0.75$), respectively, with
11 dust mass concentrations of 0–100 ppm and r_{ef} of 50–1000 μm . For a given dust mass
12 concentration and r_{ef} , $\alpha_{\text{integrated}}$ for IDM (uniform) was smaller than EDM, which is
13 due to higher light absorption in the UV and VIS bands for IDM (uniform) relative to
14 EDM (Figure 4a). While $\alpha_{\text{integrated}}$ for IDM (uniform) was larger compared with
15 IDM (central, $\bar{r}_c < 0.75$), this can be attributed to the fact that radiation is focused near
16 the center of an ice grain with IDM (central) (Ackerman and Toon, 1981; Bohren, 1986),
17 enabling further absorptions from inclusions near the center of a grain due to the lensing
18 effect (Mackowski et al., 1990). For example, $\alpha_{\text{integrated}}$ (dust concentration of 20
19 ppm, r_{ef} of 500 μm) was 0.73 for IDM (uniform), less than EDM of 0.76, but higher
20 than IDM (central, $\bar{r}_c < 0.75$) of 0.72.

21 To quantify the effects of IDM on broadband snow albedo relative to EDM, we defined
22 a broadband snow albedo scaling factor ($E_{\alpha, \text{integrated}}$), which refers to the ratio of
23 $\alpha_{\text{integrated}}$ of IDM to EDM. Generally, for dust mass concentrations from 0 to 100 ppm
24 and r_{ef} of 50–1000 μm , $E_{\alpha, \text{integrated}}$ varied from 0.89 to ~ 1.00 for IDM (uniform) (Figure
25 8a) and from 0.85 to ~ 1.00 for IDM (central, $\bar{r}_c < 0.75$) (Figure 8b). $E_{\alpha, \text{integrated}}$
26 decreased significantly with increasing dust mass concentration and r_{ef} . In addition, $E_{\alpha, \text{integrated}}$
27 for IDM (central, $\bar{r}_c < 0.75$) was smaller than IDM (uniform). These results
28 have implications for the effects of IDM in real environments. For example, IMC has

1 typical dust concentrations of ~ 10 ppm (Wang et al., 2013), so $E_{\alpha, \text{integrated}}$ (r_{ef} of 50–
2 1000 μm) was 0.96–0.99 and 0.95–0.99 for IDM (uniform) and IDM (central, $\bar{r}_c < 0.75$),
3 respectively. In contrast, dust concentrations are typically ~ 100 ppm in the TP
4 (Ming et al., 2016; Li et al., 2017, 2018), so $E_{\alpha, \text{integrated}}$ ranged from 0.89 to 0.98 and
5 0.85 to 0.96 for IDM (uniform) and IDM (central, $\bar{r}_c < 0.75$), respectively. The results
6 show that IDM (uniform) and IDM (central, $\bar{r}_c < 0.75$) reduced broadband snow
7 albedo by $\sim 2.5\%$ and $\sim 3.0\%$, respectively, in clean snow and $\sim 6.5\%$ and $\sim 9.5\%$,
8 respectively, in polluted snow relative to EDM. Moreover, the sensitivity of $E_{\alpha, \text{integrated}}$
9 to mineral dust decreased with increasing dust concentrations. For example, the
10 difference in $E_{\alpha, \text{integrated}}$ (r_{ef} of 500 μm) between dust concentrations of 10 and 20 ppm
11 was 0.011 and 0.015 for IDM (uniform) and IDM (central, $\bar{r}_c < 0.75$), respectively,
12 while the corresponding differences between dust concentrations of 90 and 100 ppm
13 were only 0.004 and 0.005. when dust concentrations were 10–20 ppm, but only 0.004
14 and 0.005 for IDM (uniform) and IDM (central, $\bar{r}_c < 0.75$), respectively, when dust
15 concentrations were 90–100 ppm. These results provide a convenient method to
16 calculate the albedo of IDM when the albedo of EDM has been obtained for a given
17 dust mass concentration and r_{ef} .

18 3.4 Uncertainties

19 Although we calculated the imaginary RI values of dust using previous studies (section
20 2.2), there are still large variations which strongly depend on dust composition (e.g.,
21 hematite/iron content) (Balkanski et al., 2007; Wagner et al., 2012). To roughly account
22 for this, we estimated the influence of chosen imaginary RI values on spectrally
23 weighted snow albedo ($E_{\alpha, \text{integrated}}$) by increasing and decreasing the calculated
24 imaginary RI values by 50%. These changes in imaginary RIs are plausible because
25 they are consistent with other studies (McConnell et al., 2010; Wagner et al., 2012).
26 The results showed that $E_{\alpha, \text{integrated}}$ uncertainties attributed to the imaginary RIs of dust
27 were $\pm 3.9\%$ and $\pm 5.2\%$ for IDM (uniform) and IDM (central, $\bar{r}_c < 0.75$), respectively.
28 In contrast, observations have displayed large variations in the size distribution of dust

1 in the atmosphere and snow, and this variation is strongly affected by the dust source
2 and transport (Mahowald et al., 2014; Shao et al., 2011). In our standard simulation, we
3 assumed a log-normal dust size distribution with a geometric mean diameter of 0.65
4 μm and a standard deviation of 2.0 (equivalent to an effective radius of 1.1 μm), which
5 is typical for dust transported long-range (Formenti et al., 2011; Maring et al., 2003);
6 nearer sources of dust tend to be larger (Kok, 2011). Therefore, we investigated the
7 effects of dust particle size on our results by assuming another two log-normal size
8 distributions with effective radii of 2.5 μm and 5.0 μm , which were within the observed
9 size ranges in the atmosphere and snow and comparable with previously analyzed dust
10 particle sizes (Maring et al., 2003; Shao and Mao, 2016; Zhang et al., 2003). The results
11 showed that the uncertainty of $E_{\alpha, \text{integrated}}$ attributed to dust diameter was $\pm 6.1\%$ for both
12 IDM (uniform) and IDM (central, $\bar{r}_c < 0.75$), but it should be emphasized that the
13 uncertainty of $E_{\alpha, \text{integrated}}$ induced by dust sizes comes purely from the uncertainty of
14 broadband albedo of dust-snow external mixing due to different dust sizes. This is
15 because effective complex refractive indices of dust-snow internal mixture effective
16 optical constants of ice containing small dust particles (i.e., internal mixing) are
17 independent of dust particle size (Eq. 8). Overall, the total uncertainty of $E_{\alpha, \text{integrated}}$
18 from variations of imaginary RIs and dust diameter was $\pm 11.0\%$ and $\pm 11.2\%$ for IDM
19 (uniform) and IDM (central, $\bar{r}_c < 0.75$), respectively.

20 **3.5 Measurement-based estimate of the effects of dust on snow albedo**

21 Finally, widespread dust concentrations in snow across the Northern Hemisphere were
22 obtained to assess the effects of dust on snow albedo in real snowpack. Figure 9 shows
23 measured dust concentrations in snow in different regions; dust concentrations spanned
24 a broad range of values because of spatial and temporal variations in emissions,
25 transportation, and deposition among the different regions. Dust concentrations widely
26 varied from ~ 3 ppm to ~ 600 ppm, with the highest concentration in NAM and lowest
27 in the EP (Di Mauro et al., 2015; Lim et al., 2014; Painter et al., 2012; Reynolds et al.,
28 2020). However, snow samples collected in the days after a significant dust transport

1 event showed that dust concentrations in snow can be up to ~70 ppm in the EP (Di
2 Mauro et al., 2015). Additionally, the average dust concentrations in fresh snow were
3 18, 6, and ~~28~~17 ppm in the southern TP, central TP, and northern TP, respectively,
4 similar to the IMC (12 ppm) (Wang et al., 2013). However, dust concentrations in aged
5 snow (120, 300, and 140 ppm) were one to two orders of magnitude higher than in fresh
6 snow, indicating the important correlation between snow type and dust concentration
7 (Zhang et al., 2017, 2018).

8 We calculated the broadband snow albedo for EDM, IDM (uniform), and IDM (central,
9 $\bar{r}_c < 0.75$) based on the measured dust concentrations (Figure 10). The results showed
10 that broadband snow albedo decreased by 0.8%, 1.4%, and 1.6% in the EP for EDM,
11 IDM (uniform), and IDM (central, $\bar{r}_c < 0.75$), respectively, which was similar to SJ.
12 However, the broadband snow albedo decreased by up to 5.6%, 8.1%, and 9.4% in the
13 EP after a significant dust transport event, indicating strong snow albedo reduction
14 during these events. In addition, broadband snow albedo was reduced by 2.0%, 3.1%,
15 and 3.6% in IMC for EDM, IDM (uniform), and IDM (central, $\bar{r}_c < 0.75$), respectively.
16 Similar results were also found for the southern TP, central TP, and northern TP where
17 the broadband snow albedo for fresh snow was reduced by 2.5%, 1.4%, and ~~2.5~~3.3%,
18 respectively, for EDM, 3.9%, 2.1%, and ~~3.8~~5.1% for IDM (uniform), and 4.5%, 2.4%,
19 and ~~4.3~~5.7% for IDM (central, $\bar{r}_c < 0.75$). However, the broadband snow albedo was
20 more significantly reduced for aged snow: up to 6.0%, 8.1%, and 7.5% for EDM, 9.5%,
21 11.6%, and 10.5% for IDM (uniform), and 10.9%, 13.2%, and 12.3% for IDM (central,
22 $\bar{r}_c < 0.75$) in the southern, central, and northern TP, respectively. This indicates that the
23 effects of dust on snow albedo showed stronger reductions during snowmelt periods, it
24 is worth noting that the effect of dust on snow albedo for aged snow could be
25 underestimated due to the larger snow grain effective radius (r_{ef}) than fresh snow.
26 Moreover, the largest broadband snow albedo reductions were found in NAM with
27 ranges of 9.8%–17.6%, 13.9%–24.1%, and 15.9%–27.0% for EDM, IDM (uniform),
28 and IDM (central, $\bar{r}_c < 0.75$), respectively. These results suggest that the effects of

1 external or internal dust–snow mixing on snow albedo are particularly significant for
2 the TP and NAM regions, with stronger reductions in albedo. Therefore, these results
3 can have significant impacts on both local hydrological cycles and regional climate
4 change (Oaida et al., 2015; Xie et al., 2018).

5 **4. Discussion**

6 In this study, the application of the effective medium approximation greatly simplifies
7 the complexity of snow radiation transfer calculation for dust–snow internal mixing,
8 and the effect of non-uniform distribution of dust in snow grains on snowpack optical
9 properties are explicitly quantified for the first time. However, it is worth noting that
10 this method has its limitation when applying to large particles (e.g., dust) in snow
11 (Bohren 1986; Flanner et al., 2012), which can create some errors for the albedo
12 calculation of dust–snow internal mixing. To verify the credibility of our results, we
13 carefully make a comparison with the more rigorous calculations found in He et al.
14 (2019), which used the geometric-optics surface-wave approach (GOS) to consider the
15 impact of dust–snow uniform internal mixing on snow albedo reduction. As shown in
16 Figure 11, the results show that the enhancement ratio of snow albedo reduction (1.28)
17 due to dust–snow internal mixing (relative to external mixing) is slightly higher than
18 the value (1.16) reported by He et al. (2019), and this deviation is comparable to that
19 caused by snow nonsphericity (He et al., 2019). Therefore, we indicated that the
20 effective medium approximation used in this study is reasonable and reliable.

21 Over the past few decades, the effects of dust in snow on reductions in albedo ~~has~~ have
22 been widely demonstrated (Skiles et al., 2018; Zhang et al., 2018). However, the
23 magnitude of these effects has only been studied in a few regions, and uncertainties still
24 remain. Our study indicates that the albedo of dust-contaminated snowpack can be
25 affected by the dust–ice mixing state. In particular, IDM enhanced light absorption and
26 reduced snow albedo more significantly compared with EDM. For example, in IMC
27 and the TP, IDM reduced snow albedo by ~5% relative to EDM at a typical dust mass
28 concentration of 20 ppm and a snow grain radius of 500 μm . This exceeds the

1 contribution of BC to snow light absorption over most areas of North America and the
2 Arctic (Dang et al., 2017). In addition, the effects of IDM on snow albedo were
3 amplified by higher dust mass concentrations and larger snow grain sizes. We therefore
4 strongly suggest that IDM must be considered in future climate models, particularly to
5 more accurately evaluate the climate in areas where snowpack is heavily contaminated
6 with dust and is experiencing melting.

7 The mixed state between dust and snow gradually progresses from partial external
8 mixing to wholly internally mixed. Therefore, assuming a completely external mixing
9 of dust and snow grains will underestimate the effects of dust on snow albedo and
10 radiative forcing in numerical models (e.g., Dang et al., 2015; Nagorski et al., 2019).
11 Similarly, assuming completely internal mixing of dust and snow grains will
12 overestimate the effects of dust (e.g., He et al., 2019; Liou et al., 2014). ~~Therefore,~~
13 ~~information gained solely from the external mixing of dust and snow grains will~~
14 ~~underestimate the effects of dust on snow albedo and radiative forcing in numerical~~
15 ~~models (e.g., Dang et al., 2015; Nagorski et al., 2019).~~ ~~Similarly, only using information~~
16 ~~from internal mixing of dust and snow grains will overestimate the effects of dust on~~
17 ~~snow albedo and radiative forcing (e.g., He et al., 2019; Liou et al., 2014).~~ Zhao et al.
18 (2014) underestimated the effects of dust by treating wet-deposited dust as externally
19 mixed with snow grains. In future studies, we recommend the actual ratio between
20 external and internal mixing for dust in snow be examined with an environmental
21 scanning electron microscope equipped with a cold stage.

22 **5. Conclusions**

23 In this study, the effects of dust particles on absorption coefficients and snow albedo
24 were theoretically analyzed by combining asymptotic radiative transfer theory and
25 (core/shell) Mie theory. We initially considered external mixing – when dust is present
26 between ice grains – and variations of internal mixing of dust within ice grains. We
27 found that snow spectral absorption coefficients of IDM were larger than EDM across
28 UV to NIR wavelengths, but were negligible at wavelengths >1000 nm. The absorption

1 enhancement (relative to EDM) was wavelength-dependent and increased with
2 increased dust concentrations.

3 Compared with a uniform distribution of dust particles in ice grains, our calculations
4 showed that non-uniformly distributed dust particles may lead to significantly different
5 snow spectral absorption coefficients in the VIS band. Snow spectral absorption
6 coefficients were further increased when all of the dust was positioned in the central
7 part of ice grains, while the maximum absorption coefficient was found when the radius
8 of a dust-polluted core was <75% of the ice grain radius. In contrast, snow spectral
9 absorption coefficients decreased when all of the dust was positioned in the surface
10 layer of ice grains, and the minimum absorption coefficient was observed in the thin
11 surface layer of dust-polluted ice grains, which was similar for EDM. As a result,
12 broadband snow albedo decreased by up to 21%, 30%, and 33% for EDM, IDM
13 (uniform), and IDM (central, $\bar{r}_c < 0.75$), respectively, at dust concentrations of 100
14 ppm and r_{ef} of 1000 μm .

15 Based on comprehensive field measurements across the Northern Hemisphere, the
16 effect of dust on snow albedo in real snowpack was evaluated by assuming external and
17 internal dust–snow mixing. The largest reductions in broadband snow albedo were in
18 NAM because that region had the highest average dust concentrations; IDM (uniform)
19 and IDM (central, $\bar{r}_c < 0.75$) further decreased snow albedo by 4.6%–7.8% and 6.8%–
20 11.4%, respectively, compared with EDM. This implies an important influence of
21 internal dust–snow mixing in NAM.

22 **Data availability**

23 The code of (core/shell) Mie theory used in this study can be found at [http://gwest.gats-](http://gwest.gats-inc.com/software/software_page.html)
24 [inc.com/software/software_page.html](http://gwest.gats-inc.com/software/software_page.html).

25 **Author contributions**

26 WX designed the study and evolved the overarching research goals and aims. STL
27 wrote the first draft with contributions from all co-authors. STL and CJC applied formal

1 techniques such as statistical, mathematical, and computational to analyze study data.
2 CY and XXY collected the dust measurements across the Northern Hemisphere. ZY
3 and PW provided the majority of the methodology and software. CQL and ZXL
4 provided technical guidance. All authors contributed to the improvement of results and
5 revised the final paper.

6 **Competing interests**

7 The authors declare that they have no conflict of interest.

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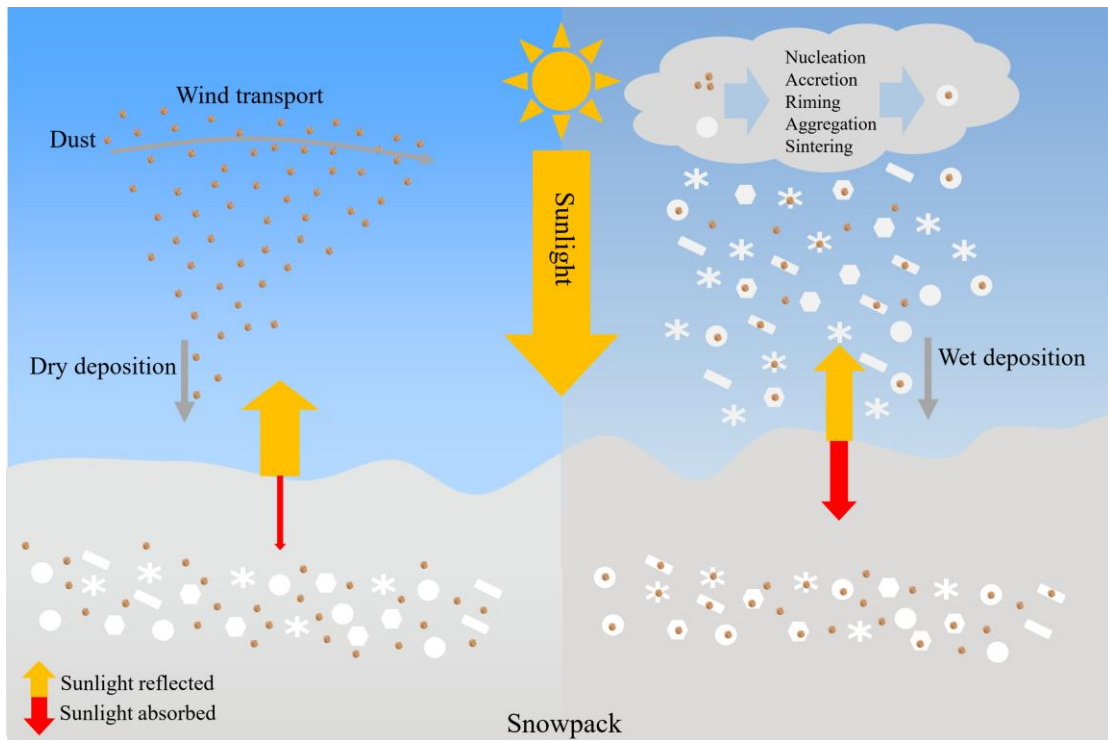
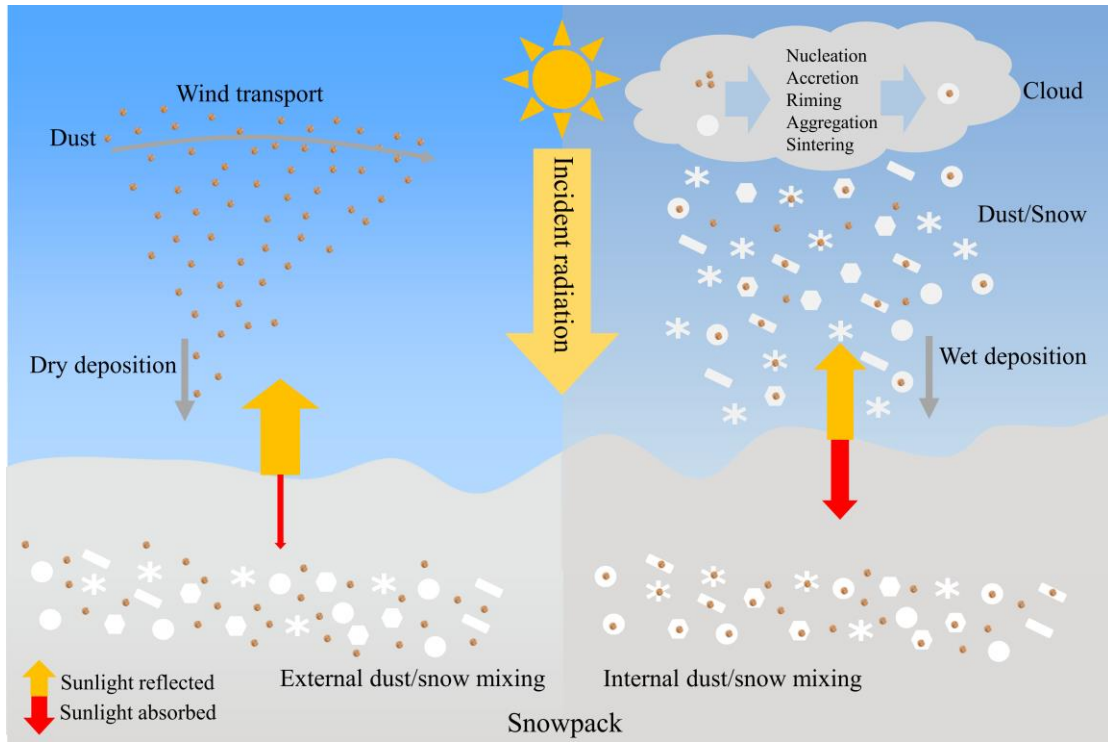
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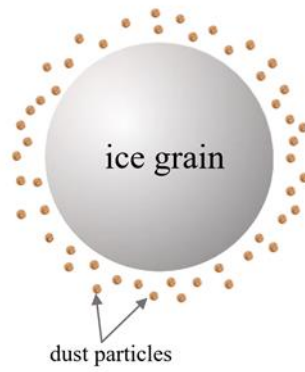
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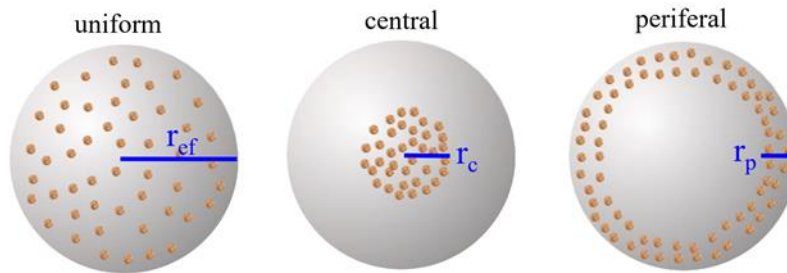
3 **Figure 1.** Schematic of dust mixing with snow grains internally. Dust tends to mix
 4 externally with snow grains through dry deposition and/or below cloud scavenging,
 5 while dust–snow internal mixtures can be produced by nucleation, accretion, riming,
 6 aggregation, and sintering during aerosol–cloud–precipitation processes known as wet
 7 deposition. Arrows represent how the absorption (red) and reflection (yellow) of
 8 incoming sunlight changes with dust–snow mixing state.

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External dust/ice mixing (EDM)

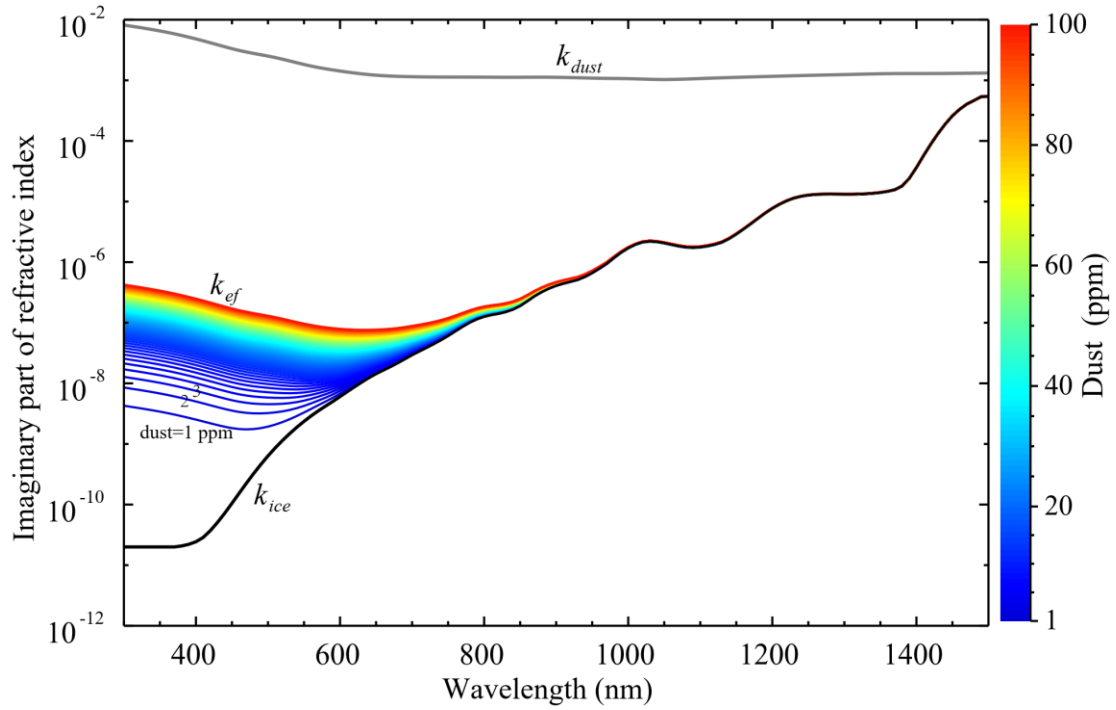


Internal dust/ice mixing (IDM)



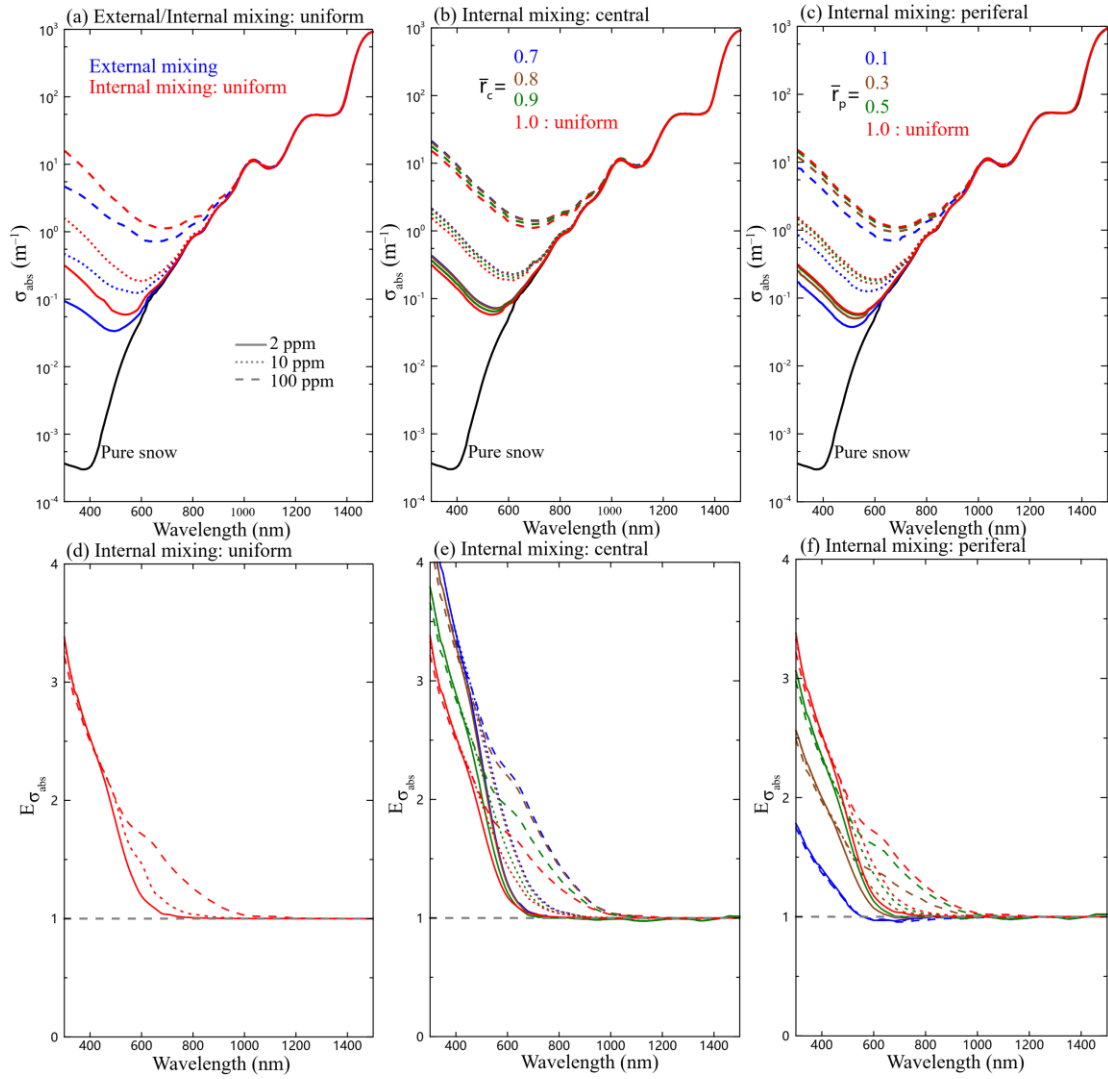
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Figure 2. Schematic depicting various mixing scenarios of snow grains and dust particles.



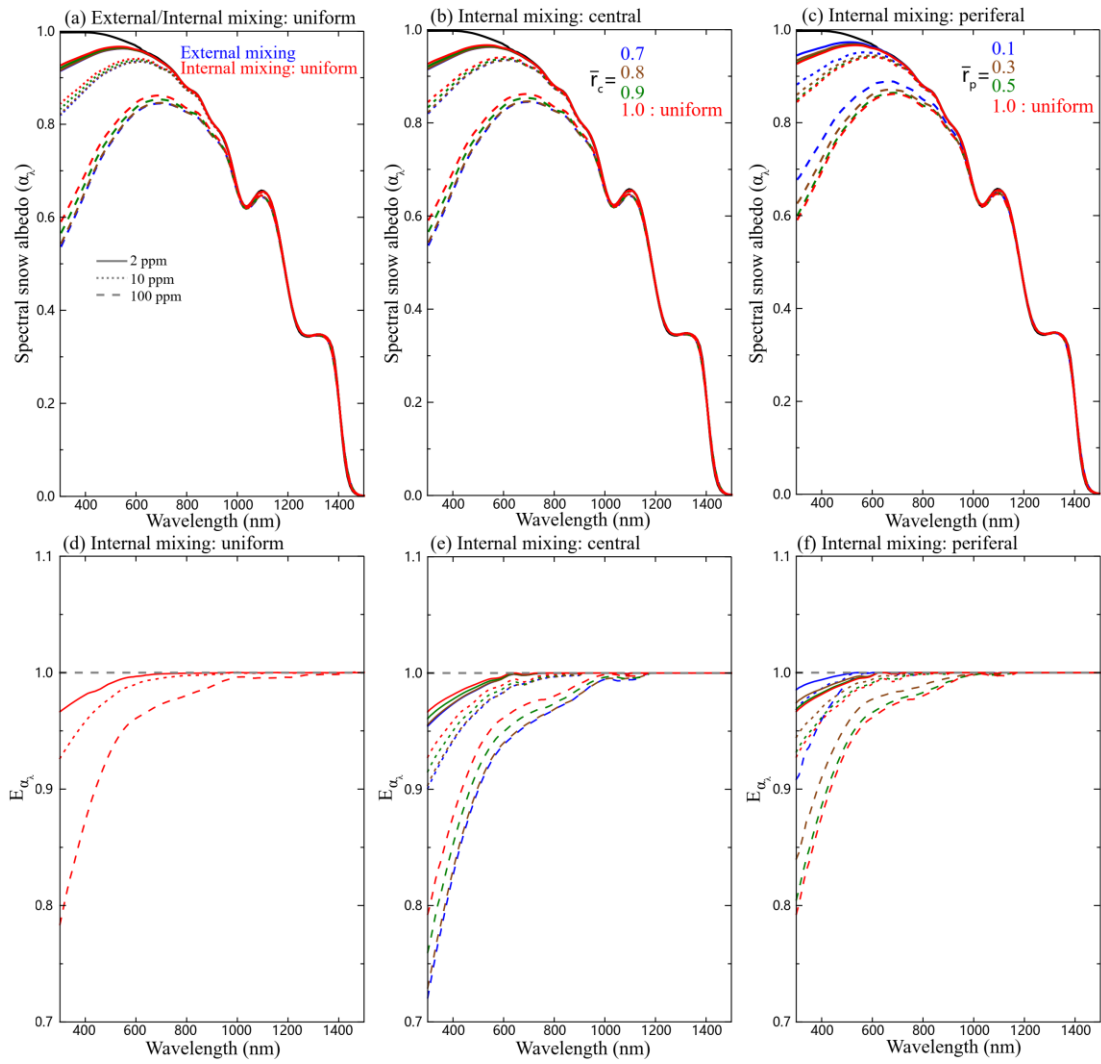
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Figure 3. Imaginary part of the spectral complex refractive indices of ice (k_{ice}) and dust (k_{dust}) (Warren and Brandt, 2008; Dang et al., 2015), with the imaginary part of the effective complex refractive indices (k_{ef}) as a function of wavelength, at dust mass concentrations (C_{dust}^*) of 1–100 ppm (or $\mu\text{g g}^{-1}$) in snow.



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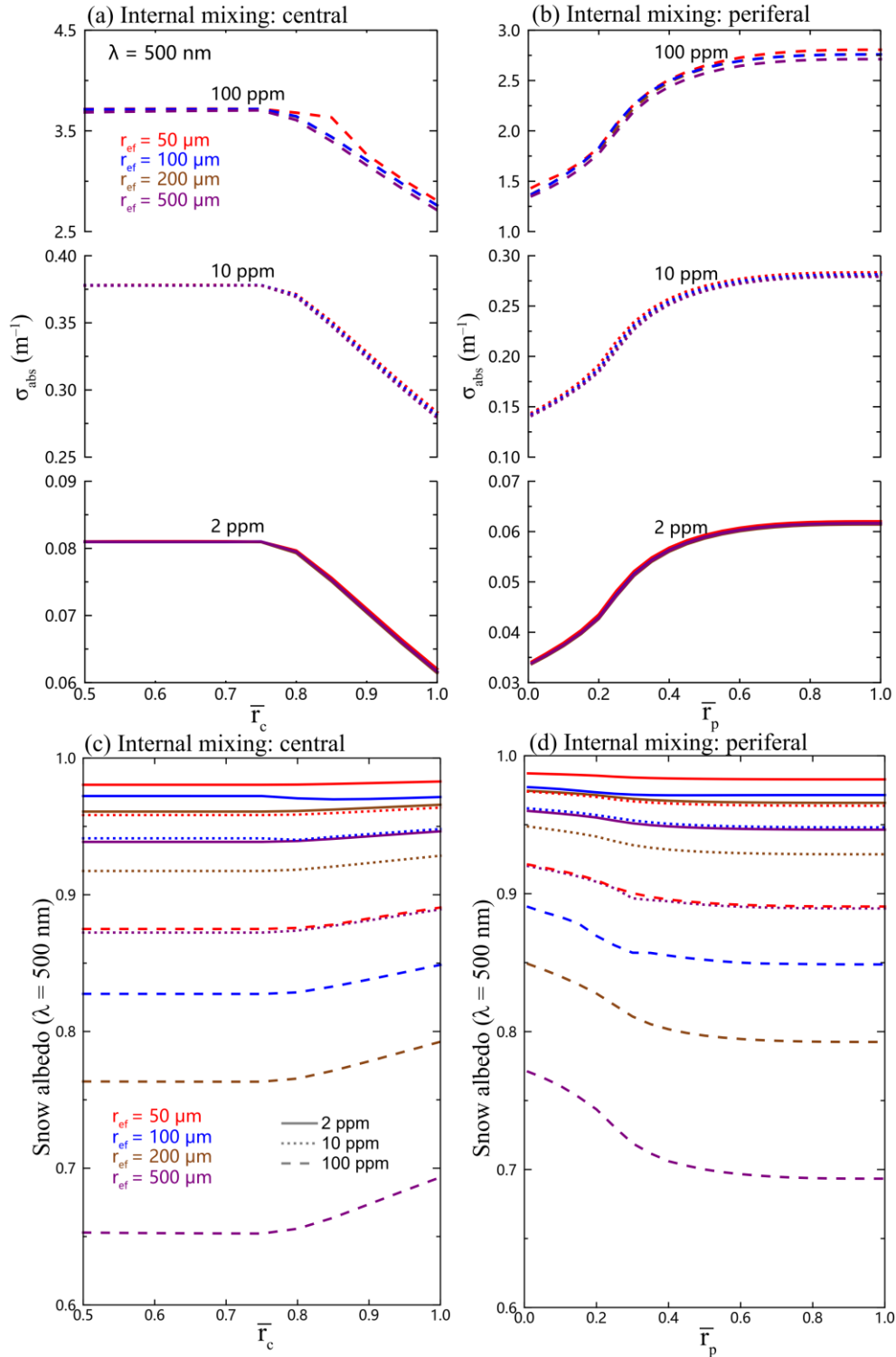
Figure 4. Snow absorption coefficients (σ_{abs}) for dust–snow (a) external and internal mixing (uniform), (b) internal mixing (central), and (c) internal mixing (peripheral), as a function of wavelength with different dust concentrations and \bar{r}_c and \bar{r}_p . The corresponding enhancement ($E_{\sigma_{\text{abs}}}$) caused by (d) internal mixing (uniform), (e) internal mixing (central), and (f) internal mixing (peripheral) relative to external mixing, is shown as a function of wavelength. The snow grain radius was assumed to be 200 μm .



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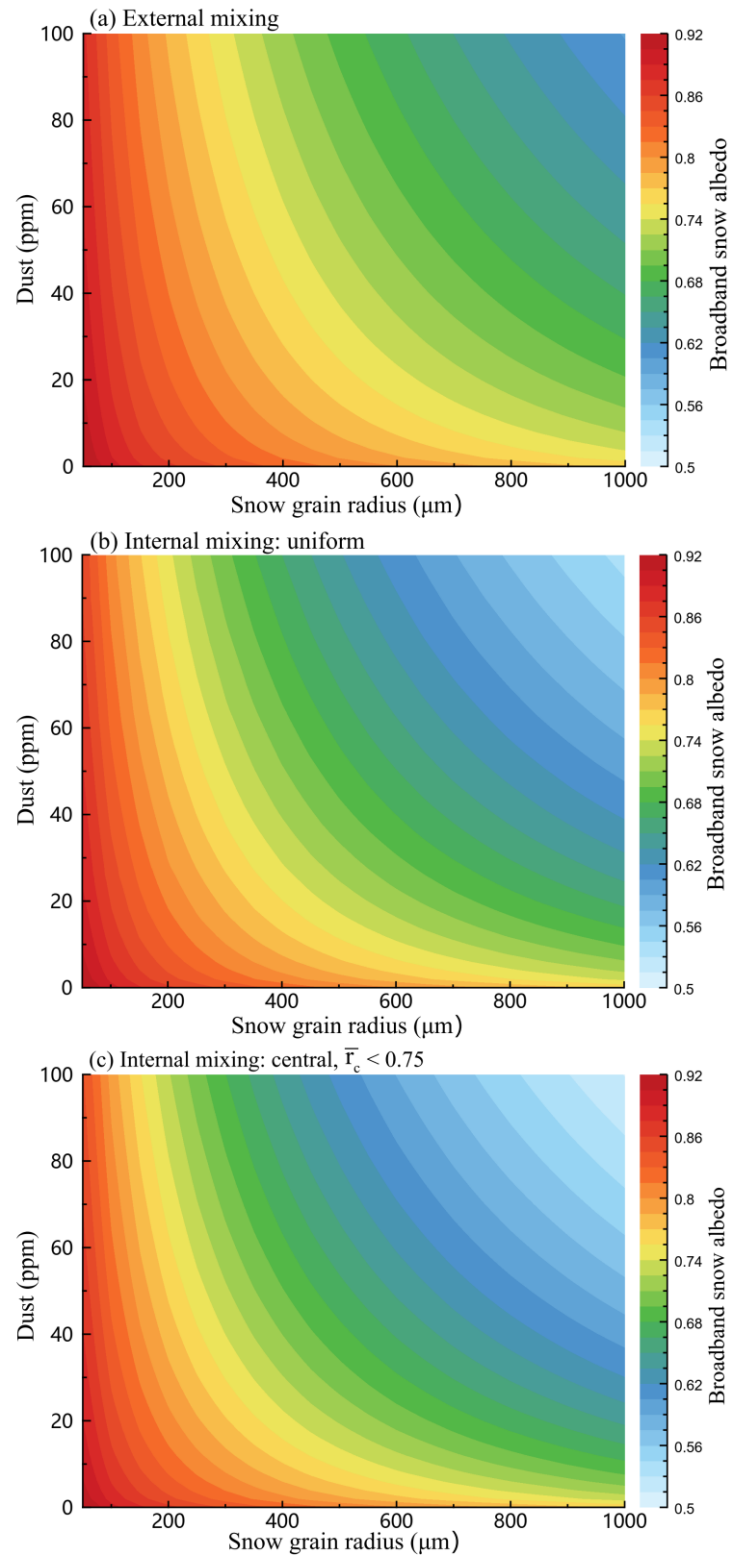
2 **Figure 5.** Same as Figure 4, but for spectral snow albedo (α_λ).

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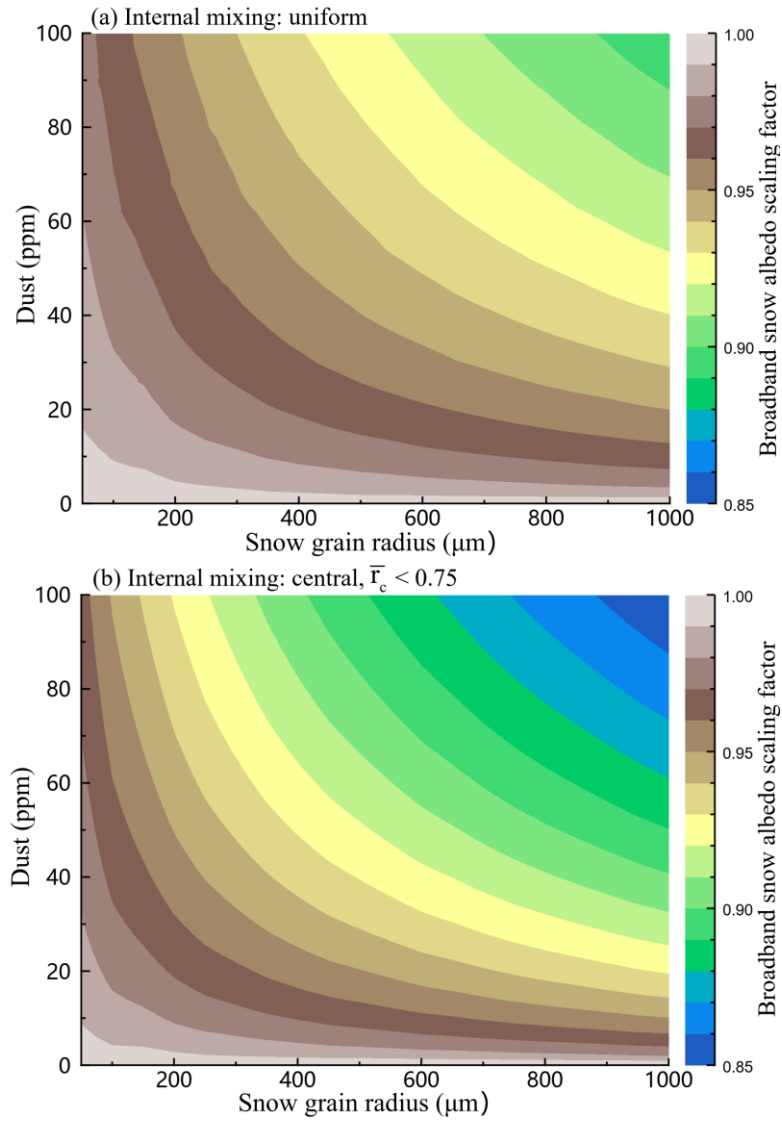


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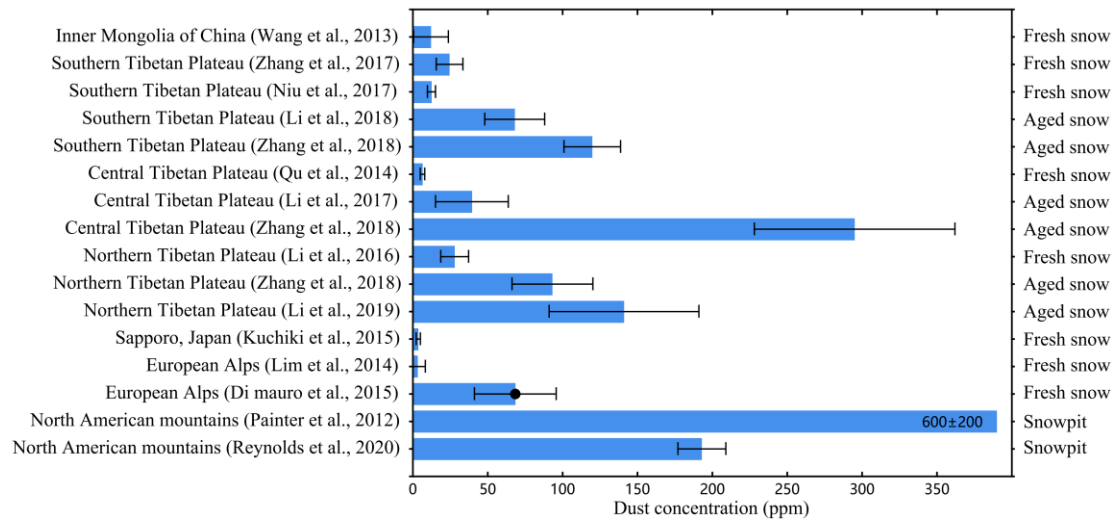
Figure 6. The snow absorption coefficient (σ_{abs}) at 500 nm wavelength as a function of \bar{r}_c and \bar{r}_p for (a) internal mixing (central) and (b) internal mixing (peripheral) with different snow grain radii and dust mass concentrations. (d) and (e) are the same as (a) and (b), but for snow albedo at 500 nm wavelength.



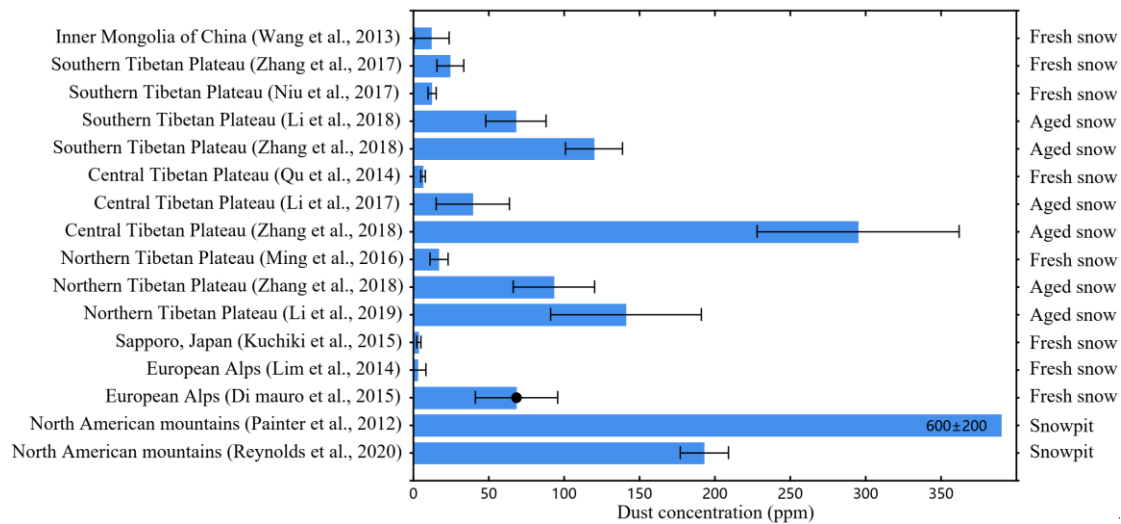
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2 **Figure 7.** Broadband snow albedo ($\alpha_{\text{integrated}}$) variations affected by different dust
3 mass concentrations and snow grain radii for (a) external mixing, (b) internal mixing
4 (uniform), and (c) internal mixing (central, $\bar{r}_c < 0.75$).
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 2 **Figure 8.** Variations in the broadband snow albedo scaling factor ($E_{\alpha, \text{integrated}}$, ratio of
 3 $\alpha_{\text{integrated}}$ for IDM to EDM) due to different dust mass concentrations and snow grain
 4 radii for (a) internal mixing (uniform) and (b) internal mixing (central, $\bar{r}_c < 0.75$).
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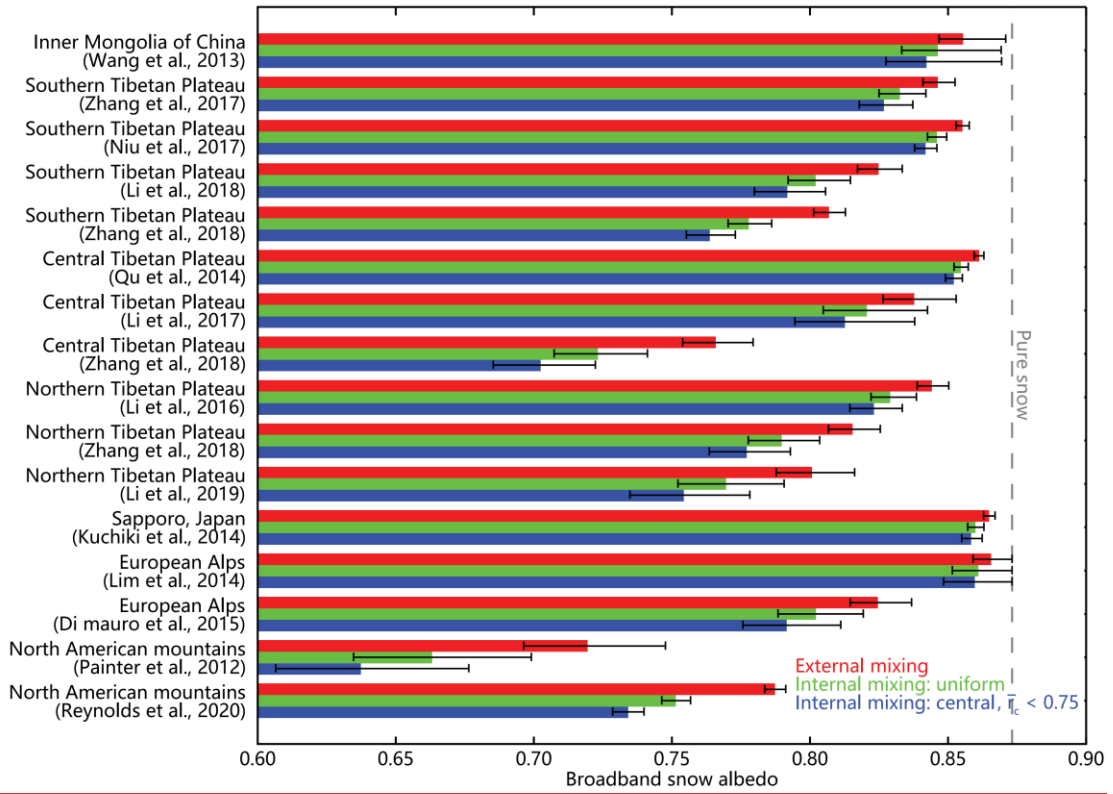


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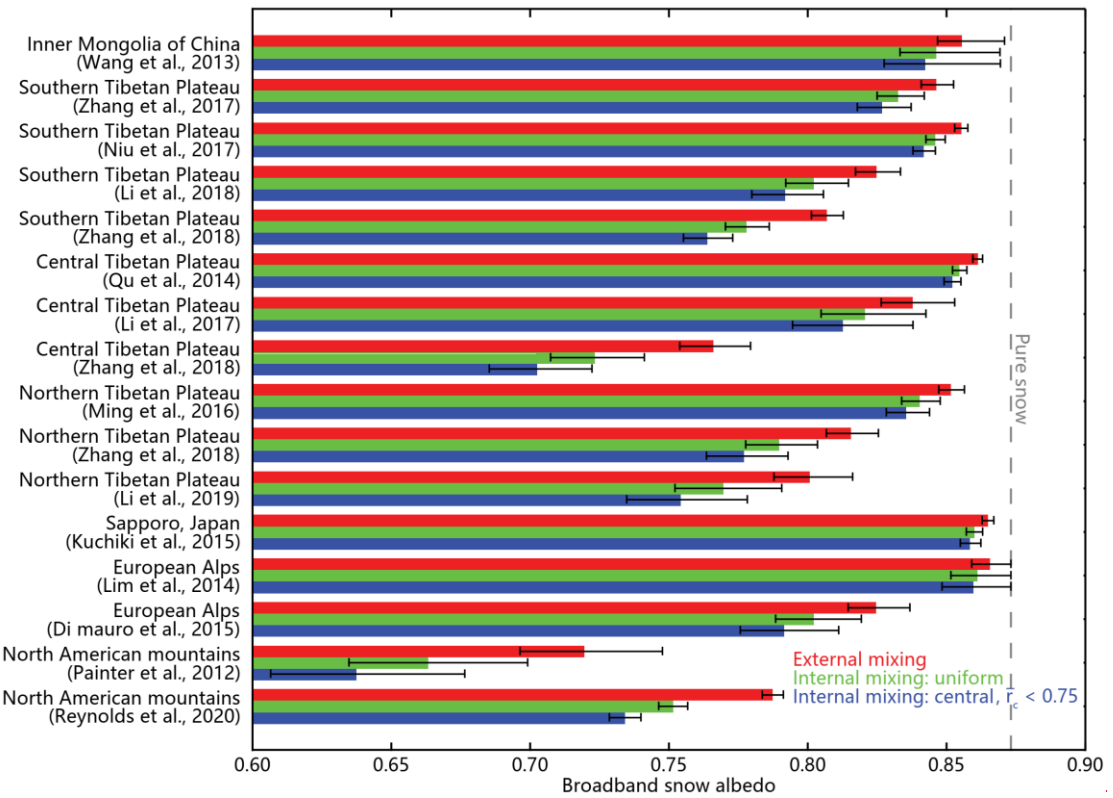
3 **Figure 9.** In situ measurements of dust concentrations in snow (fresh snow, aged snow,
 4 and snowpit from field sampling in different regions of the Northern Hemisphere. The
 5 solid black circle represents snow samples that were collected days after a significant
 6 dust transport event.

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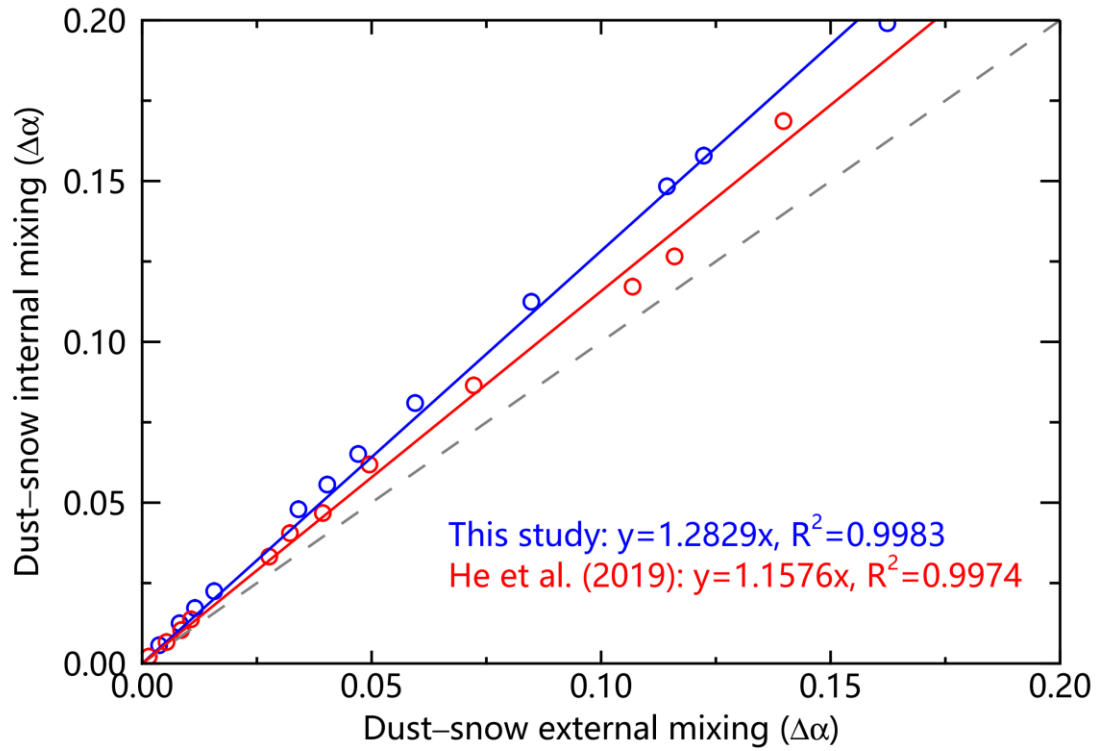
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3 **Figure 10.** Calculated broadband snow albedo based on dust concentration
4 measurements in different areas for dust–snow external mixing, internal mixing
5 (uniform), and internal mixing (central, $\bar{r}_c < 0.75$). The dashed line represents
6 broadband albedo of pure snow, and the snow grain radius was assumed to be 200 μm .



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Figure 11. Comparisons of snow albedo reduction ($\Delta\alpha$) under the cloudy sky caused by dust-snow uniform internal mixing (y-axis) and external mixing (x-axis) for this study (blue) and He et al. (2019) (red).