

# Variability of Lidar-Derived Particle Properties Over West Africa Due to Changes in Absorption: Towards an Understanding

Igor Veselovskii<sup>1</sup>, Qiaoyun Hu<sup>2</sup>, Philippe Goloub<sup>2</sup>, Thierry Podvin<sup>2</sup>, Mikhael Korenskiy<sup>1</sup>, Yevgeny Derimian<sup>2</sup>, Michel Legrand<sup>2</sup>, Patricia Castellanos<sup>3</sup>

<sup>1</sup>*General Physics Institute, Moscow, Russia.*

<sup>2</sup>*Univ. Lille, CNRS, UMR 8518 - LOA - Laboratoire d'Optique Atmospherique, Lille F-59000, France*

<sup>3</sup>*NASA Goddard Space Flight Center, Greenbelt, USA*

## Abstract

Measurements performed in Western Africa (Senegal) during the SHADOW-2 field campaign are analyzed to show that spectral dependence of the imaginary part of the complex refractive index (CRI) of dust can be revealed by lidar-measured particle parameters. Observations in April 2015 provide good opportunity for such study, because, due to high optical depth of the dust, exceeding 0.5, the extinction coefficient could be derived from lidar measurements with high accuracy and contribution of other aerosol types, such as biomass burning, was negligible. For instance, in the second half of April 2015, AERONET observations demonstrated a temporal decrease of the imaginary part of CRI at 440 nm from approximately 0.0045 to 0.0025. This decrease is in line with a change in the relationship between the lidar ratios (the extinction-to-backscattering ratio) at 355 nm and 532 nm ( $S_{355}$  and  $S_{532}$ ). For instance in the first half of April,  $S_{355}/S_{532}$  is as high as 1.5 and the backscatter Angstrom exponent  $A_\beta$ , is as low as -0.75, while after 15 April  $S_{355}/S_{532}=1.0$  and  $A_\beta$  is close to zero. The aerosol depolarization ratio  $\delta_{532}$  for the whole April exceeded 30% in the height range considered, implying that no other aerosol, except dust, occurred. The performed modeling confirmed that the observed  $S_{355}/S_{532}$  and  $A_\beta$  values match the spectrally dependent imaginary part of the refractive index as can be expected for mineral dust containing iron oxides. The second phase of the SHADOW-2 campaign was focused on evaluation of the lidar ratio of smoke and estimates of its dependence on relative humidity (RH). For five studied smoke episodes the lidar ratio increases from  $44\pm 5$  sr to  $66\pm 7$  sr at 532 nm and from  $62\pm 6$  sr to  $80\pm 8$  sr at 355 nm, when RH varied from 25% to 85%. Performed numerical simulations demonstrate, that observed ratio  $S_{355}/S_{532}$ , exceeding 1.0 in the smoke plumes, can indicate to increase of the imaginary part of the smoke particles in the ultraviolet (UV).

## 1. Introduction

Atmospheric dust provides significant impacts on the Earth's climate system and this impact remains highly uncertain (IPCC report, 2013). In modeling the direct aerosol effect, the

37 vertical profile of the aerosol extinction is one of the basic input parameters, and when this  
38 profile is derived from the elastic backscatter lidar observations, the knowledge of the extinction-  
39 to-backscatter ratio (so called lidar ratio) is essential (Klett, 1985). Although the desert dust in  
40 source regions is sometimes qualified as “pure dust”, it is always a mixture of various elements,  
41 e.g. iron oxides, clays, quartz and calcium-rich species, which proportions can vary (Sokolik and  
42 Toon, 1999; Wagner et al., 2012; Di Biagio et al., 2017, 2019 and references therein). Thus, the  
43 dust optical properties, and hence the lidar ratio ( $S$ ) can vary, depending on the relative  
44 abundance of various minerals in emission sources. The imaginary part of the complex refractive  
45 index (CRI) of different minerals can vary spectrally and often exhibits an increase in UV  
46 spectral region for dust, containing iron oxides. Therefore, the retrieval of the dust extinction  
47 profiles from elastic backscatter lidar observation should account for the spectral variation of the  
48 lidar ratio.

49 Raman and HSRL lidars are capable of providing independent profiling of aerosol  
50 backscattering and extinction coefficients (Ansmann et al., 1992), and therefore are widely used  
51 to measure the lidar ratios of dust from different origins (e.g. Sakai et al., 2003; Papayannis et  
52 al., 2008, 2012; Xie et al., 2008; Ansmann et al., 2011; Mamouri et al., 2013; Burton et al., 2014;  
53 Nisantzi et al., 2015; Giannakaki et al., 2016; Hofer et al., 2017, 2019; Soupiona et al., 2018,  
54 2019). The African deserts are the largest sources of mineral dust and numerous studies have  
55 been conducted for quantifying the particle intensive parameters (parameters independent of  
56 concentration) during dust transport from this source region to Europe and over the Atlantic  
57 Ocean (Mattis et al., 2002; Amiridis et al., 2005; Mona et al., 2006; Papayannis et al., 2008;  
58 Preißler et al., 2013; Groß et al., 2015; Rittmeister et al., 2017; Haarig et al. 2017). The dust  
59 properties are, however, modified during this transport, experiencing mixing and aging  
60 processes, thus the characterization of the dust properties near the source regions is highly  
61 important for the evaluation the parameters of “pure dust”.

62 The lidar ratios at 355 nm and 532 nm ( $S_{355}$  and  $S_{532}$ ) were measured during the  
63 SAMUM-1 and 2 experiments in Morocco and Capo Verde respectively (Esselborn et al., 2009;  
64 Tesche et al., 2009, 2011; Groß et al., 2011; Ansmann et al., 2011), as well as during the more  
65 recent SHADOW-2 experiment in Senegal (Veselovskii et al., 2016, 2018). The lidar ratios  $S_{355}$   
66 and  $S_{532}$  measured during SAMUM experiments did not present significant spectral dependence.  
67 For example, for SAMUM-2 campaign, the averaged values of  $S_{355}$  and  $S_{532}$  were  $53 \pm 10$  sr and  
68  $54 \pm 10$  sr respectively (Tesche et al., 2011). During SHADOW, however,  $S_{355}$  significantly  
69 exceeded  $S_{532}$  in many dust episodes, which was linked to an increase of the imaginary part of  
70 CRI of dust at 355 nm (Veselovskii et al., 2016).

71           The dust backscattering coefficient, in contrast to the extinction coefficient, is sensitive to  
72 the imaginary part of CRI (Perrone et al., 2004; Veselovskii et al., 2010; Gasteiger et al., 2011).  
73 Recall, that the particle extinction is the sum of absorption and scattering, and increase of  
74 absorption is accompanied by decrease of scattering, leading to weak dependence of the  
75 extinction on the imaginary part (Veselovskii et al. 2010, Fig.6). Thus, it is expected that  
76 enhanced absorption in the UV should increase the lidar ratio. In turn, the ratio  $S_{355}/S_{532}$  should  
77 characterize the spectral variation of the imaginary part of CRI. The latest version of AERONET  
78 products (3.0) provides inversions of the lidar related properties, including the lidar ratio, from  
79 almucantar scans with ground-based sun photometers. For these products, the shortest available  
80 wavelength is 440 nm. Despite the imaginary part at 440 nm ( $Im_{440}$ ) is lower than  $Im_{355}$ ,  
81 AERONET observations still show an increase of absorption at 440 nm in respect to 675 nm that  
82 yields a ratio of  $S_{440}/S_{675}$  close to 1.4 for Saharan dust (Shin et al., 2018). The goal of this work is  
83 to analyze the correlation of variations of  $Im_{440}$  from AERONET with measured values from  
84 lidar to reveal the effect of dust absorption on lidar-derived aerosol properties. We focus on  
85 height and day-to-day variations of the dust intensive properties, such as  $S_{355}$  and  $S_{532}$ ,  
86 depolarization ratio ( $\delta$ ), as well as the extinction and backscatter Ångström exponents ( $A_\alpha$  and  
87  $A_\beta$  respectively) measured during several strong dust episodes in April 2015 during the  
88 SHADOW-2 campaign.

89           The smoke aerosol particles, typically originated from biomass burning, can also have a  
90 pronounced spectral dependence of absorption (Nicolae et al., 2013). This is generally due to  
91 presence of carbonaceous particles with organic compounds, so-called brown carbon (BrC) (Sun  
92 et al., 2007; Kirchstetter, et al., 2004). The Sahel region is known for seasonal biomass burning  
93 caused by human activity on combustion of agricultural waste that can produce an abundant  
94 amount of BrC. The smoke can also be mixed with mineral dust during long-range transport or in  
95 the emission origin (Haywood et al., 2008). During the SHADOW-2 the observation period  
96 included the biomass burning season, thus an additional effort was dedicated to the examination  
97 of the spectral lidar ratio variability of transported biomass burning aerosol under different  
98 environmental conditions and presents a supplementary subject of the current study.

99           The paper is organized as follows. Section 2 describes the lidar system and provides the  
100 main expressions used for the data analysis. Several strong dust episodes, in April 2015, are  
101 analyzed in Section 3. In Section 4, the smoke episodes occurring from December 2015 to  
102 January 2016, are used to evaluate the variation of the smoke lidar ratio with relative humidity.  
103 The paper is finalized with conclusion.

104

105           **2. Experimental setup and data analysis**

106 The observations were performed with LILAS multiwavelength Raman lidar during  
 107 SHADOW-2 campaign at Mbour, Senegal. Information related to the SHADOW-2 and  
 108 observation site is presented in Veselovskii et al. (2016). The LILAS is based on a tripled  
 109 Nd:YAG laser with a 20 Hz repetition rate and pulse energy of 90/100/100 mJ at 355/532/1064  
 110 nm. The aperture of the receiving telescope is 400 mm. During the campaign, LILAS  
 111 configuration ( $3\beta+2\alpha+1\delta$ ) allowed the measurement of three particle backscattering ( $\beta_{355}$ ,  $\beta_{532}$ ,  
 112  $\beta_{1064}$ ), two extinction coefficients ( $\alpha_{355}$ ,  $\alpha_{532}$ ) and depolarization ratio at 532 nm ( $\delta_{532}$ ). To  
 113 improve the performance of the system at 532 nm the rotational Raman channel was used instead  
 114 of the vibrational one (Veselovskii et al, 2015). The measurements were performed at a 47  
 115 degrees angle to horizon. The backscattering coefficients and depolarization ratios were  
 116 calculated with a 7.5 m range resolution (corresponding to 5.5 m vertical resolution), while the  
 117 range resolution of extinction coefficient varied from 50 m (at 1000 m) to 125 m (at 7000 m).  
 118 Particle extinction and backscattering coefficients at 355 nm and 532 nm are calculated from  
 119 elastic and Raman backscatter signals, as described in Ansmann et al. (1992) and corresponding  
 120 uncertainties are shown on the plots. An additional Raman reception channel at 408 nm was  
 121 setup for profiling the water vapor mixing ratio (WVMR) (Whiteman et al., 1992).

122 The particle depolarization ratio  $\delta$ , determined as a ratio of cross- and co-polarized  
 123 components of the particle backscattering coefficient, was calculated and calibrated the same  
 124 way as described in Freudenthaler et al. (2009). The relative uncertainty of depolarization  
 125 measurements due to calibration is estimated as  $\pm 10\%$ . To analyze the complex aerosol mixtures,  
 126 containing dust (d) and smoke (s), we can write  $\beta = \beta^d + \beta^s$  and  $\alpha = \alpha^d + \alpha^s$ . The  
 127 depolarization ratio of such a mixture is therefore:

$$128 \quad \delta = \frac{\left(\frac{\delta^d}{1+\delta^d}\right)\beta^d + \left(\frac{\delta^s}{1+\delta^s}\right)\beta^s}{\frac{\beta^d}{1+\delta^d} + \frac{\beta^s}{1+\delta^s}} \quad (1)$$

129 Here  $\delta^d$  and  $\delta^s$  are the particle depolarization ratios of dust and smoke components respectively.

130 To characterize the spectral dependence of the extinction ( $\alpha$ ) and backscattering ( $\beta$ )  
 131 coefficients, the corresponding Ångström exponents are introduced as:

$$132 \quad A_\alpha = \frac{\ln\left(\frac{\alpha_{\lambda_1}}{\alpha_{\lambda_2}}\right)}{\ln\left(\frac{\lambda_2}{\lambda_1}\right)} \quad \text{and} \quad A_\beta = \frac{\ln\left(\frac{\beta_{\lambda_1}}{\beta_{\lambda_2}}\right)}{\ln\left(\frac{\lambda_2}{\lambda_1}\right)} \quad (2)$$

133 Where  $\alpha_{\lambda_1}$ ,  $\alpha_{\lambda_2}$ ,  $\beta_{\lambda_1}$ ,  $\beta_{\lambda_2}$  are the extinction and backscattering coefficients at wavelengths  $\lambda_1$  and  
 134  $\lambda_2$ . For the mixture of smoke and dust, the extinction Ångström exponent (EAE) can be  
 135 calculated from the ratio  $\frac{\alpha_{\lambda_1}}{\alpha_{\lambda_2}}$  :

$$136 \quad \frac{\alpha_{\lambda_1}}{\alpha_{\lambda_2}} = \frac{\alpha_{\lambda_1}^d + \alpha_{\lambda_1}^s}{\alpha_{\lambda_2}^d + \alpha_{\lambda_2}^s} = \frac{\alpha_{\lambda_1}^d \left(1 + \frac{\alpha_{\lambda_1}^s}{\alpha_{\lambda_1}^d}\right)}{\alpha_{\lambda_2}^d \left(1 + \frac{\alpha_{\lambda_2}^s}{\alpha_{\lambda_2}^d}\right)} = \frac{\alpha_{\lambda_1}^d \left(1 + \frac{\alpha_{\lambda_2}^s \left(\frac{\lambda_2}{\lambda_1}\right)^{A_{\alpha}^s}}{\alpha_{\lambda_2}^d \left(\frac{\lambda_2}{\lambda_1}\right)^{A_{\alpha}^d}}\right)}{\alpha_{\lambda_2}^d \left(1 + \frac{\alpha_{\lambda_2}^s}{\alpha_{\lambda_2}^d}\right)} = \frac{\alpha_{\lambda_1}^d \left(1 + \frac{\alpha_{\lambda_2}^s}{\alpha_{\lambda_2}^d} \left(\frac{\lambda_2}{\lambda_1}\right)^{(A_{\alpha}^s - A_{\alpha}^d)}\right)}{\alpha_{\lambda_2}^d \left(1 + \frac{\alpha_{\lambda_2}^s}{\alpha_{\lambda_2}^d}\right)} \quad (3)$$

137 Here  $A_{\alpha}^d$  and  $A_{\alpha}^s$  are the Ångström exponents of dust and smoke. The Ångström exponent of the  
 138 mixture is obtained from (3):

$$139 \quad A_{\alpha} = \frac{\ln \frac{\alpha_{\lambda_1}}{\alpha_{\lambda_2}}}{\ln \frac{\lambda_2}{\lambda_1}} = A_{\alpha}^d + \frac{1}{\ln \frac{\lambda_2}{\lambda_1}} \ln \left[ \frac{\left(1 + \frac{\alpha_{\lambda_2}^s}{\alpha_{\lambda_2}^d} \left(\frac{\lambda_2}{\lambda_1}\right)^{(A_{\alpha}^s - A_{\alpha}^d)}\right)}{\left(1 + \frac{\alpha_{\lambda_2}^s}{\alpha_{\lambda_2}^d}\right)} \right] \quad (4)$$

140 The backscattering Ångström exponent (BAE) can be calculated in a similar way. And finally,  
 141 the lidar ratio of the aerosol mixture is calculated as:

$$142 \quad S = \frac{S^d \beta^d + S^s \beta^s}{\beta^d + \beta^s} = S^d + \frac{\beta^s}{\beta} (S^s - S^d) \quad (5)$$

143 where  $S^d$  and  $S^s$  are the lidar ratios of dust and smoke.

144

### 145 **3. Dust observations in March and April 2015**

146 The aerosol over West Africa presents strong seasonal variations. The spring is  
 147 characterized by strong dust emission, while, during winter season, intense forest fires occurring  
 148 in the equatorial regions emit smoke particles that are transported over Senegal (Veselovskii et  
 149 al., 2018). The SHADOW-2 campaign included the following periods of measurements: 13  
 150 March – 25 April 2015, 8–25 December 2015 and 5-24 January 2016, so numerous dust and  
 151 smoke episodes were observed. In our analysis of lidar-derived aerosol properties, we considered  
 152 also aerosol columnar properties provided by AERONET (Holben et al. 1998) and aerosol  
 153 profiles predicted by the Modern-Era Retrospective analysis for Research and Applications,  
 154 Version 2 (MERRA-2) aerosol reanalysis (Gelaro et al., 2017; Randles et al., 2017). MERRA-2  
 155 is the first long-term global reanalysis to assimilate space-based aerosol observations and include

156 their radiative coupling with atmospheric dynamics. MERRA-2 is driven by the Goddard Earth  
157 Observing System (GEOS) model version 5 that includes the Goddard Chemistry, Aerosol,  
158 Radiation and Transport (GOCART) module. GOCART models the sources, sinks, and  
159 transformation of the following five aerosol species as external mixtures: dust, organic carbon  
160 (OC), black carbon (BC), sulfates (SU) and sea salt (SS). Dust and sea salt are represented by  
161 five non-interacting size bins, and have wind-speed dependent emissions. The MERRA-2  
162 reanalysis assimilates AOD observations from the twin Moderate Resolution Imaging  
163 Spectroradiometer (MODIS) instruments, MODIS-Terra and MODIS-Aqua, as well as the  
164 AERONET ground-based sun photometer network. In addition, the profiles of meteorological  
165 variables (P, T, RH), provided by radio-sondes at the Dakar airport, located ~70 km from the  
166 M'bour site, were also available. The relative humidity (RH) profiles over the M'bour site were  
167 calculated from the combination of lidar-derived WVMR and temperature profile from  
168 radiosounding.

169 Fig.1 shows the aerosol optical depth at 532 nm ( $AOD_{532}$ ) for March, April and  
170 December 2015 recalculated from AERONET AOD at 500 nm using the 440-675 nm Ångström  
171 exponent. The same figure shows the AODs for the five aerosol species used in MERRA-2  
172 model, such as dust, organic carbon (OC), black carbon (BC), sulfates (SU) and sea salt (SS).  
173 The optical depths provided by MERRA-2 and AERONET are in a good agreement. Dust is the  
174 predominant aerosol component for all three months with the highest values of AOD in April.  
175 The contribution of organic carbon (the main component of the biomass burning products) is  
176 significant in December, when the forest fire season starts in equatorial regions, though  
177 noticeable amount of OC is predicted also for March and for the beginning of April. The  
178 contribution of BC and SU to the total AOD is low: the sum of the corresponding AODs is  
179 below 0.1 for all three months.

180 The single scattering albedo (SSA) over the M'Bour site in 2015 provided by AERONET  
181 at 440 and 675 nm is shown in Fig.2. The  $SSA_{675}$  is above 0.97 for March – April period, but at  
182 440 nm dust absorption is stronger and, in March,  $SSA_{440}$  is about 0.9. However, in the middle of  
183 April,  $SSA_{440}$  increases up to 0.95, indicating that aerosols become less absorbing at shorter  
184 wavelengths. We can thus expect that variation of SSA at 355 nm between April and March  
185 should be even stronger. In our study we consider two groups of observations. The first group  
186 corresponds to the beginning of April, when SSA at 440 nm was lower. The second group covers  
187 the second half of April, when SSA at 440 nm increased. By analyzing these two groups we  
188 expect to reveal the effect of aerosol absorption, on lidar-derived aerosol properties.

189

190 **3.1. Dust episode on 1 – 4 April 2015**

191 In the beginning of April the dust was transported by Continental trades (Harmattan)  
192 from the northeastern/eastern drylands. For period 1 - 4 April, as follows from Fig.1b, the  
193 AOD<sub>532</sub> over Dakar increased up to 1.0. Fig.3 shows the spatio-temporal distributions of the  
194 aerosol backscattering coefficient  $\beta_{532}$ , particle depolarization ratio  $\delta_{532}$ , and water vapor mixing  
195 ratio for the nights 1-2, 2-3 and 3-4 April 2015. The corresponding air mass back-trajectories,  
196 shown in Fig.4, demonstrate that, on 1-2 and 2-3 April, air masses at all heights arrive from the  
197 North-East, whereas on 3-4 April the air masses above 2500 m are advected from the East. These  
198 air masses are characterized by higher humidity and may contain biomass-burning products.  
199 During these three nights, the linear particle depolarization ratio and WVMR present some  
200 evolution. On 1-2 April  $\delta_{532}$  exceeds 30% and does not change significantly within the dust  
201 layer, even if some decrease is observed above 2000 m after 03:00 UTC. By 3-4 April the  
202 depolarization ratio above 2500 m decreases below 25%, simultaneously with increase of the  
203 WVMR. During the dust episode, the relative humidity did not exceed 20% on 1-3 April, but on  
204 3-4 April it increased up to 40% above 2500 m.

205 Vertical profiles of dust particle properties such as aerosol extinction coefficients  $\alpha_{355}$ ,  
206  $\alpha_{532}$ , particle depolarization ratio  $\delta_{532}$  and lidar ratios  $S_{355}$ ,  $S_{532}$  are shown on Fig.5 for the three  
207 observation periods on 1, 2-3 and 3-4 April 2015. The profiles of backscattering coefficients for  
208 2-3 and 3-4 April are given by Fig.6, while the extinction and backscatter Ångström exponents,  
209 calculated at 355 and 532 nm wavelengths for three temporal intervals from Fig.5, are presented  
210 in Fig.7. During all three observation periods  $A_\alpha$  is slightly negative ( $A_\alpha = -0.1 \pm 0.1$ ) up to 2000  
211 m. For the dust component, MERRA-2 provides value of  $A_\alpha = -0.14$ , which agrees with  
212 observations. Above 2000 m,  $A_\alpha$  exhibits some increase, which is most significant on 3-4 April,  
213 when  $A_\alpha$  reaches  $0.3 \pm 0.1$  at 4000 m height. Simultaneous decrease of  $\delta_{532}$  indicates to the  
214 possible presence of smoke particles above 2000 m. The backscatter Ångström exponent  $A_\beta$ , in  
215 contrast with  $A_\alpha$ , is sensitive to the spectral dependence of the imaginary part of CRI, thus  
216 yielding complicated vertical variability of  $A_\beta$ . In particular, on 2-3 April  $A_\beta$  decreases from -0.5  
217 to -0.7 within 1500–2500 m height range, when  $A_\alpha$  remains stable.

218 As follows from Fig.5, on 1 April the lidar ratio  $S_{355} = 70 \pm 6$  sr does not change with  
219 height, while  $S_{532}$  gradually decreases from  $60 \pm 5$  sr at 1000 m to  $50 \pm 4$  sr at 3000 m height. On  
220 sessions that followed (Fig.5b,c) the lidar ratios at both 355 nm and 532 nm decreased. Thus, the  
221 range of lidar ratios variation for the dust episode on 1-4 April is 60-70 sr at 355 nm and 45-60 sr  
222 at 532 nm. The lidar ratios ( $S_{355}$  and  $S_{532}$ ) modeled by MERRA-2 for the dust component are  
223 also shown on Fig. 5. The corresponding lidar ratio values are of 70 sr and 42 sr respectively and  
224 do not vary with altitude as the model optical properties of all dust size bins based on spectral

225 complex refractive indices from the Optical Properties of Aerosols and Clouds (OPAC) tables  
226 (Hess et al. 1998) and the spheroidal shape models developed by Meng et al. (2010) are the same  
227 and fixed, as dust is treated as hydrophobic. Modeled value  $S_{355}$  is near the top of the range of  
228 observed values, while modeled  $S_{532}$  underestimates the observations.

229 The gradual decrease of  $S_{532}$  with height in Fig.5a,c is however unusual. There are, at  
230 least, two possible reasons to explain  $S_{532}$  height variation. The first one can be the presence of  
231 non-dust particles, for example, smoke. The second reason is that the properties (composition) of  
232 dust change with height. If non-dust particles are present, the particle intensive properties, such  
233 as  $S$ ,  $\delta$  and  $A_\alpha$  should vary with height in consistent way. The MERRA-2 modeling reported in  
234 Fig.1 shows that in the beginning of April the organic carbon is the second main contributor to  
235 the AOD, after dust. We should recall, however, that the model can provide a realistic range of  
236 OC variation, however not necessarily reproducing the exact spatio-temporal distribution of OC  
237 extinction coefficient.

238 In the dust episode considered, the most significant smoke contribution was observed on  
239 3-4 April. Fig.8a shows the profiles of measured  $\alpha_{355}$  and  $\alpha_{532}$  together with MERRA-2 modeled  
240 extinction coefficients at 532 nm for five aerosol components. The extinction Ångström  
241 exponents measured by lidar and modeled by MERRA-2 for dust component are given by  
242 Fig.8b. The same figure shows also the lidar derived water vapor mixing ratio profile together  
243 with the relative humidity. At low altitudes (below 2500 m), where aerosol is represented by  
244 pure dust, the measured and modeled values of extinction coefficients are close. Above 2500 m  
245 the measured value of  $\alpha_{355}$  exceeds that of  $\alpha_{532}$ , indicating the presence of smoke particles, while  
246 modeled contribution of OC to the total extinction is very low. The measured extinction  
247 Ångström exponent is about -0.1 below 2000 m, which well agrees with modeling results for  
248 pure dust. Increase of WVMR and RH above 2000 m coincides with growth of the  $A_\alpha$ . For the  
249 considered case, the model reproduces correctly the dust loading, but underestimates the smoke  
250 contribution. At 3500 m, the difference between measured and modeled  $\alpha_{532}$  is about  $0.045 \text{ km}^{-1}$   
251 which can be attributed to the smoke contribution.

252 Dust and smoke particles contributions to the total backscattering coefficient can be also  
253 separated on the basis of the depolarization measurements, assuming that depolarization ratios of  
254 these particles are known (Tesche et al., 2009). The results of such decomposition are presented  
255 in Fig.8c, assuming 35% and 7% for dust and smoke depolarization ratio. The  $\delta_{532}=7\%$  was the  
256 lowest value that we observed in elevated smoke layers during the SHADOW experiment  
257 (Veselovskii et al., 2018), however, due to large difference of smoke and dust depolarization  
258 ratios, the choice of exact value for the smoke did not influence significantly the results. The



259 contribution of smoke to the total  $\beta_{532}$  at 3500 m is  $0.0009 \text{ km}^{-1}\text{sr}^{-1}$ . For the smoke lidar ratio of  
260 50 sr at 532 nm (validity of this choice will be discussed in section 3.3), the smoke extinction  
261 coefficient is about  $0.045 \text{ km}^{-1}$ . This value agrees well with smoke contribution obtained from  
262 Fig.8a at 3500 m and thus can be used for estimating the smoke effect on the intensive aerosols  
263 properties derived from lidar measurements.

264 The depolarization ratio of the “dust-smoke” mixture, calculated with expression (1),  
265 matches the observed value since decomposition in Fig.8c is based on depolarization  
266 measurements. The Ångström exponent at 3500 m computed with (4) for  $\alpha_{532}^s = 0.045 \text{ km}^{-1}$ ,  $\alpha_{532}^d$   
267  $= 0.147 \text{ km}^{-1}$ ,  $A_\alpha^d = -0.1$  and  $A_\alpha^s = 0.9$  yields  $A_\alpha = 0.28$ , which is close to observed value  $0.26 \pm 0.08$ .  
268 Hence, the observed variation of  $A_\alpha$  above 2000 m on 3-4 April is well explained by smoke  
269 contribution. In a similar way, using (5) we can estimate the smoke lidar ratio ( $S_{532}^s$ ) that would  
270 match the observed decrease of  $S_{532}$ . To explain decrease of the lidar ratio at 3500 m from 50 sr  
271 to 45 sr, the smoke lidar ratio should be about 25 sr, which is unrealistically small (Burton et al.,  
272 2012). Such small lidar ratio could be attributed to the maritime aerosol, but then the lidar ratios  
273 at both wavelengths should decrease simultaneously. Recall that on 1-2 April smoke contribution  
274 was significantly lower, while decrease of  $S_{532}$  is about 10 sr. Thus, smoke particles presence  
275 cannot explain the observed decrease of  $S_{532}$  and it should be probably attributed to changes of  
276 dust composition (and so the imaginary part) with height.

277 Smoke lidar ratio is usually assumed to be higher than that of dust (Tesche et al., 2011;  
278 Burton et al., 2012), meanwhile in Fig 5c the lidar ratio  $S_{532}$  is not increased in presence of the  
279 smoke particles. It should however be noticed that our results were obtained at low RH. The  
280 smoke particles are hygroscopic and the lidar ratio should increase with RH. The way to  
281 characterize  $S_{532}^s$  over Dakar site can be based on the analysis of the lidar measurements during  
282 smoke episodes within height range where smoke contribution becomes predominant. The  
283 results of such analysis will be discussed later in section 3.3.

284

### 285 **3.2. Dust episodes on 14 and 24 April 2015.**

286 In the second part of April 2015, dust  $\text{AOD}_{532}$  exceeded 1.0 (Fig.1b) and contributions of  
287 other aerosol components were insignificant. Meanwhile, as follows from Fig.2,  $\text{SSA}_{440}$   
288 increased after 15 April, thus dust became less absorbing in the UV, which should influence the  
289 lidar-derived aerosol intensive properties. Fig.9 shows the values of the extinction coefficients  
290 and lidar ratios at 355 nm and 532 nm, together with depolarization ratio  $\delta_{532}$  and the Ångström  
291 exponents  $A_\alpha$  and  $A_\beta$  observed on 14 April (00:00 – 05:00 UTC) and 23-24 April (23:00-06:00

292 UTC). The first case is a “transition day” when  $SSA_{440}$  starts to increase. The aerosol extinction  
293 profiles presented in Fig.9a show that two dust layers can be distinguished. In the first layer  
294 (below 2.5 km), aerosol intensive properties are similar to that of 1-4 April with  $S_{355} > S_{532}$ ,  
295 slightly negative  $A_\alpha = -0.1$  and  $A_\beta$  as low as  $-0.35$ . In the second layer  $S_{355}$  and  $S_{532}$  coincide and  
296 both  $A_\alpha$  and  $A_\beta$  are close to zero. The depolarization ratio in the second layer is about 31%,  
297 slightly lower than in the first one. Thus, we can assume that increase of the imaginary part in  
298 UV in the first layer is more significant, than in the second one. From the analysis of airmass  
299 back-trajectories given in Fig.10, we can conclude that the air masses in the first layer originate  
300 from the Northeastern/Eastern drylands, while in the second layer the air masses arrive from the  
301 East. After 14 April,  $S_{355}$  and  $S_{532}$  coincided for the whole height range and results obtained on  
302 23-24 April (Fig.8 c, d) are the example of such observations. Airmass back-trajectories show  
303 that the air masses at both 2.0 and 3.0 km height are transported from East. The ratio  $S_{355}/S_{532}$  is  
304 close to 1.0 within the whole dust layer and both Ångström exponents  $A_\alpha$ ,  $A_\beta$  are close to zero.  
305 Thus, the results from Figs.9, 10 are indicating that lidar-derived aerosol properties depend on  
306 the dust source origin.

307

### 308 **3.3 Analysis of lidar ratio variations in March – April 2015**

309 Fig.11 summarizes the lidar ratio measurements for period from 29 March to 24 April  
310 2015 (first phase of SHADOW ended on 25 April). Here we focus on the properties of “pure  
311 dust”, thus do not show results before 29 March, when AOD was lower and the contribution of  
312 other aerosol types could be significant (Fig.1). For the Fig.11 we have chosen height intervals,  
313 where S value is stable and  $\delta$  exceeds 30%. For example, on 14 and 24 April lidar ratios are  
314 averaged inside 2.7-3.7 km and 2.0-4.0 km layers respectively. For the period considered,  $S_{355}$   
315 and  $S_{532}$  vary in the ranges 50 sr – 80 sr and 45 sr - 60 sr respectively with a mean values of 62 sr  
316 and 51 sr. Enhanced variability of  $S_{355}$  compared to  $S_{532}$  can be explained by variation of the  
317 imaginary part at 355 nm. At the beginning of the 29 March and 8 April dust episodes,  $S_{355}/S_{532}$   
318 ratio is as high as 1.5 and then gradually decreases. After 14 April,  $S_{355}/S_{532}$  ratio becomes close  
319 to 1.0, thus S presents no spectral dependence.

320 The day-to-day variation of the aerosol column properties, including the spectrally  
321 dependent complex refractive index, can be obtained from AERONET (Holben et al., 1998).  
322 Fig.12 shows the imaginary part of the aerosol refractive index at 440 nm and 675 nm ( $Im_{440}$ ,  
323  $Im_{675}$ ) provided by AERONET for the same period of time as in Fig.11. The  $Im_{440}$  strongly  
324 decreases after 14 April, correlating with the decrease of  $S_{355}/S_{532}$  ratio in Fig.11, which  
325 corroborates the suggestion, that variations of  $S_{355}/S_{532}$  ratio are related to variation of dust

326 absorption in the UV. The retrieved real part (Re) of the complex refractive index oscillates  
327 around  $Re=1.45$  and shows no significant spectral dependence. Correlation between  
328 enhancement of  $Im_{440}$ , with in respect to  $Im_{675}$ , and increase of lidar-derived  $S_{355}/S_{532}$  is clearly  
329 seen in Fig.13, showing time – series of difference  $Im_{440}-Im_{675}$  and  $S_{355}/S_{532}$  ratio.

330 To analyze the variations of the observed lidar ratios and the Ångström exponents, a  
331 simplified numerical simulation has been performed. For a realistic modeling of the dust lidar  
332 ratio, various mixtures of different mineral components and particles shapes should be  
333 considered. Sensitivity of the modeling results to the dust mixture parameters was demonstrated  
334 in study of Gasteiger et al. (2011). Such detailed modeling, however, is out of the scope of the  
335 present paper. Here we only intend to evaluate the main impact when the imaginary part of CRI  
336 is modified.

337 The lidar ratio depends not only on the complex refractive index but also on the dust  
338 particle size distribution (PSD). The PSDs provided by AERONET on 2 and 23 April 2015  
339 (three distributions for each day) are shown in Fig.14. The PSDs are similar and the effective  
340 radii for both days are about  $0.75 \mu m$ , thus, difference in  $S$  observed for 2 and 23 April should be  
341 related mainly to the complex refractive index. Fig.15a presents modeled  $S_{355}$  and  $S_{532}$  lidar  
342 ratios together with the extinction and backscattering Ångström exponents  $A_\alpha$ ,  $A_\beta$  as a function  
343 of the imaginary part. Computations were performed for the AERONET derived size distribution  
344 on 23 April from Fig.14 using the assembly of randomly oriented spheroids (Dubovik et al.,  
345 2006) with the real part  $Re=1.55$ .  $S_{355}$  and  $S_{532}$  increase with the imaginary part and the ratio  
346  $S_{355}/S_{532}$  is about 1.1. Extinction coefficient is slightly sensitive to the imaginary part, thus  
347 increase of  $S$  in Fig.15 is due to decrease of backscattering coefficient with  $Im$ . The modeled  $A_\alpha$   
348 is about  $A_\alpha=0.1$ , while  $A_\beta$  decreases with  $Im$  to  $A_\beta=-0.2$ . To estimate the influence of a spectrally  
349 dependent imaginary part  $Im(\lambda)$  on  $A_\beta$ , we have also performed computations assuming a fixed  
350  $Im_{532}=0.002$  and only  $Im_{355}$  is free to vary. Corresponding results are shown in Fig.15a with open  
351 stars. Spectral dependence of the imaginary part significantly decreases  $A_\beta$ : for  $Im_{355}=0.005$   
352 ( $Im_{355} - Im_{532}=0.003$ ),  $A_\beta$  decreases to  $-0.75$ .

353 We should recall however, that for the second half of April the observed ratio  $S_{355}/S_{532}$ ,  
354 was about 1.0, and both extinction and backscatter Ångström exponents were close to zero. To  
355 figure out the kind of PSD that would reproduce those observations, we retrieved the PSD from  
356  $3\beta+2\alpha$  measurements, as described in Veselovskii et al. (2002, 2010). For that purpose, data  
357 from 23-24 April (Fig.9), averaged within 2-3 km layer, were inverted and corresponding PSD is  
358 shown in Fig.14 with red line. Inversion was performed for the assembly of randomly oriented  
359 spheroids, in assumption of spectrally independent refractive index. Due to the limited number of

360 input data (five) we are able to reproduce only the main features of the PSD. The maximum of  
361 this lidar derived PSD is shifted towards larger radii, with respect to the AERONET size  
362 distribution, but at the same time, retrieved PSD contains significant contribution from the fine  
363 particles. The simulation results for this lidar derived PSD, are given by Fig.15b. The lidar ratios  
364  $S_{355}$ ,  $S_{532}$  for all values of the imaginary part are close. The backscatter and extinction Ångström  
365 exponents are close to zero, matching the observations of the second half of April 2015. Thus  
366 simulation results demonstrate dependence on the PSD chosen, but in both cases these lead to the  
367 same conclusion: observed low values of  $A_{\beta}$  can not be reproduced without accounting for  
368 spectral dependence of the imaginary part.

369 To compare computations and observations, information upon  $Im_{355}$  and  $Im_{532}$  values is  
370 needed. The recently measured refractive indices of dust, sampled at different regions of Africa,  
371 are presented by Di Biagio et al. (2019). In particular, for the countries located North and East of  
372 Senegal, the aerosol imaginary parts at 370, 470, 520, 660 nm are of 0.0043, 0.0033, 0.0026,  
373 0.0013 for Mauritania and 0.0048, 0.0038, 0.0030, 0.0024 for Mali respectively. The highest  
374 values of lidar ratios, observed in our measurements, are about 60 sr and 80 sr at 532 nm and 355  
375 nm respectively. Corresponding imaginary parts of CRI from Fig.15 can be estimated as  
376  $Im_{532}=0.002-0.003$  and  $Im_{355}=0.005-0.006$ , which agrees with results presented by Di Biagio et  
377 al. (2019). Assuming  $Im_{355}=0.005$  and  $Im_{532}=0.002$ , the modeled ratio  $S_{355}/S_{532}$  is about 1.44 and  
378  $A_{\beta}$  is about -0.75 for both AERONET and lidar derived PSDs, which again reasonably agrees  
379 with observations. The modeling performed is very simplified, still it confirms that the observed  
380 values of  $S_{355}/S_{532}$  ratio and  $A_{\beta}$  can be explained by the spectral dependence of the imaginary  
381 part of CRI.

382 Thus, based on our measurement results, two types of dust can be distinguished. The first  
383 type has high  $S_{355}/S_{532}$  ratio (up to 1.5). Such kind of dust is characterized by an increase of the  
384 imaginary part in the UV and it was observed, for example, during 29 March and 10 April  
385 episodes. For the second type, the ratio  $S_{355}/S_{532}\approx 1.0$ , so variation of the imaginary part of the  
386 refractive index between 532 and 355 nm wavelengths should be smaller than for the first type.  
387 Such dust was observed in the second half of April 2015. Both types are characterized by high  
388 depolarization ratio,  $\delta_{532}$  values, exceeding 30%, so depolarization measurements at 532 nm are  
389 not capable to discriminate between these two types of dust.

390 The difference in the observed dust properties is probably related to the mineralogical  
391 characteristics in the source region. From the back-trajectories analysis presented in Figs. 4 and  
392 10 one can suppose that the first type of dust was transported from the North–East, while the  
393 second type from the East. In order to verify if a difference in the dust emission source region  
394 and transport take place, we also analyzed the Infrared Difference Dust Index (IDDI) derived

395 from the Meteosat Second Generation (MSG) geostationary satellite imagery in the thermal  
396 infrared (TIR). The IDDI is developed by Legrand et al. (1985, 2001) originally for the Meteosat  
397 First Generation (MFG) and is based on impact of the airborne mineral dust on the TIR radiation  
398 emitted by the terrestrial surface. The physical principle of the IDDI derivation is in thermal  
399 contrast between terrestrial surface and atmosphere and the best sensitivity is found at around  
400 noon time when the surface temperature is maximal (Legrand et al., 1988). The IDDI product  
401 shows that brightness temperature of terrestrial surface observed by satellite can be reduced up to  
402 about 50°K in presence of airborne mineral dust, while reduction by about 10°K already  
403 indicates a major dust event (Legrand et al., 2001). A direct relationship between the IDDI and  
404 aerosol optical thickness in solar spectrum and visibility was also found (Legrand et al., 2001). It  
405 should be mentioned here that the IDDI was initially developed for MFG and the absolute  
406 consistence with the IDDI values from MSG should be examined due to differences in spatial  
407 and spectral resolutions between two sensors. However, the physical principles used for the IDDI  
408 determination are the same and a direct application of the MFG IDDI algorithm to MSG was  
409 found as possible. Moreover, tests showed that the absolute values of IDDI for a coincident  
410 overlapping period of MFG and MSG are very close. Nevertheless, employment of the IDDI  
411 from MSG is indeed applicable for the required in the current analysis purpose of solely dust  
412 spatial patterns detection.

413 The IDDI calculations, applied to the MSG images taken during the field campaign,  
414 clearly show a major dust event in northern and central Africa. The elevated IDDI values over  
415 Senegal are also visible. The IDDI images show distinguishable changes in the emission sources  
416 and transport features during the different phases of the observations. For instance, Fig. 16 shows  
417 that the dust emissions during the first phase of the event are originated in south Algeria,  
418 Mauritania and Mali (examples of images from 29 and 30 March 2015). Weeks later, spatial  
419 patterns of the elevated IDDI are shifted to south and show source regions in south of Niger  
420 (Fig.15c, d). Of course, attribution of emission sources mineralogy to aerosol spectral absorption  
421 is a complex task (Alfaro., et al 2004; Lafon et al., 2006; Di Biagio et al., 2017, 2019) and it is  
422 difficult to point to a specific source that could clearly explain the observed in this study change  
423 in the aerosol absorbing properties. However, the IDDI images clearly suggest a change in the  
424 dust transport regime that is consistent with the change in the dust optical properties.

425

#### 426 **4. Smoke episodes in December 2015 – January 2016**

427 During the SHADOW campaign, we had several strong smoke episodes in December  
428 2015 – January 2016, when air mass transported the products of biomass burning from the areas  
429 of intensive forest fires in equatorial region. The relative humidity in the advected smoke layers

430 varied from episode to episode, allowing evaluation of the RH influence on the smoke lidar  
 431 ratios  $S_{355}$ ,  $S_{532}$ . We should keep in mind, however, that for different days the smoke particles  
 432 could have different chemical composition, so evaluated RH dependence can be considered as  
 433 semi-quantitative only. The spatio-temporal evolution of the particle backscattering coefficient  
 434 and depolarization ratio at 532 nm, during the 14-15 December 2015 smoke episode, is given in  
 435 Fig.17. The same figure shows also the water vapor mixing ratio, a convenient tracer to identify  
 436 wet air mass arrived from the equatorial region. The smoke particles are usually contained in  
 437 elevated layers, being mixed with dust (Veselovskii et al., 2018). The height ranges where the  
 438 smoke particles are predominant can be identified by low depolarization ratio and enhanced  
 439 WVMR. For event considered, the smoke particles are predominant above 1500 m after  
 440 midnight.

441 The vertical profiles of  $\alpha_{355}$ ,  $\alpha_{532}$ ,  $S_{355}$ ,  $S_{532}$ ,  $A_\alpha$ ,  $A_\beta$  together with the water vapor mixing  
 442 ratio and the relative humidity, for 15 December (04:00 – 06:00 UTC), are shown in Fig.18. The  
 443 same figure presents decomposition of  $\beta_{532}$  to the dust and smoke contributions, based on  
 444 depolarization measurements (Tesche et al., 2011). The smoke episodes are characterized by  
 445 different relative humidity within the elevated layer. On 15 December, RH is about 40% in the  
 446 1500 – 2100 m range and the ratio  $\frac{\beta_{532}^s}{\beta_{532}}$  is about of 0.57 at 2000 m. The lidar ratio  $S_{532}$  decreases  
 447 from 50 sr to 44 sr in 1000 m - 2000 m range, while  $S_{355}$  rises from 58 sr to 67 sr, thus  $S_{355}$   
 448 significantly exceeds  $S_{532}$ . We should recall that lidar ratios presented in Fig.18 are attributed to  
 449 dust- smoke mixture. In principle, we can estimate  $S_{532}^s$  using Eq.5, because the ratio  $\frac{\beta_{532}^s}{\beta_{532}}$  is  
 450 available. Corresponding  $S_{532}^s$  profile obtained for assumed  $S_{532}^d=50$  sr is shown in Fig.18a  
 451 (black line).  $S_{532}^s$  is about 40 sr at 2000 m and it is close to measured  $S_{532}$  value. In the smoke  
 452 layer, the extinction Ångström exponent  $A_\alpha$ , can exceed  $A_\beta$ , due to negative contribution of  $A_\beta^d$ .  
 453 In particular, on 15 December  $A_\alpha$  is about 1.1, while  $A_\beta$  is close to zero.

454 To estimate the dependence of smoke lidar ratios  $S_{355}$  and  $S_{532}$  on RH, five smoke  
 455 episodes on 14-15, 15-16, 22-23, 24-25 December 2015 and 19-20 January 2016 were analyzed.  
 456  $S_{532}$  and  $S_{355}$ , together with relative humidity and the  $\frac{\beta_{532}^s}{\beta_{532}}$  ratio are summarized, for these  
 457 episodes, in Table 1. The heights chosen correspond to the values of relative humidity close to  
 458 maximum. The calculated values of RH are characterized by high uncertainties, because lidar  
 459 and sonde measurements are not collocated. Estimations of the corresponding uncertainties are

460 also given by Table 1. The lidar ratio values from Table 1 are plotted in Fig.19 as a function of  
461 RH. These plots, however, should be taken with care, because, due to variation of chemical  
462 composition and the aging processes, results may depend not only on RH. Moreover, the dust  
463 particles occurring in the elevated layers, as discussed, can introduce an additional ambiguity in  
464 the results. Nevertheless, Fig.19 demonstrates a clear increasing trend of  $S$  with RH, at both  
465 wavelengths. From this figure, one can also conclude that  $S_{355}$  always exceeds  $S_{532}$  and, that  $S_{532}$   
466 for smoke can be as small as  $44 \pm 5$  sr at low humidity. The small values of  $S_{532}$  for the “fresh  
467 smoke” (about 40 sr) were reported also by Burton et al. (2012).

468 To compare our observations with the lidar ratios used in the MERRA-2 model, we have  
469 also performed the simulation of  $S_{532}^{OC}(RH)$  and  $S_{355}^{OC}(RH)$  dependence for organic carbon (OC)  
470 based on the particle parameters and hygroscopic growth factor from MERRA-2 model. In  
471 MERRA-2 the organic carbon is the main component of the biomass burning products. The  
472 imaginary part of the OC increases in the UV due to the presence of “brown carbon” (BrC),  
473 which is a subset of organic carbon with strong absorption in the UV region (Bergstrom et al.,  
474 2007; Torres et al., 2007). The majority of BrC is emitted into the atmosphere through low-  
475 temperature, incomplete combustion of biomass. In the newest development of GEOS, biomass  
476 burning OC is now emitted as a new BrC tracer species that uses  $Im_{532}=0.009$  and  $Im_{355}=0.048$   
477 values (Hammer et al. 2016). Thus, the spectral behavior of the imaginary part of organic carbon  
478 refractive index depends on contribution of the BrC fraction to the primary organic carbon and  
479 on the physical-chemical processes in the smoke layer during its transportation. As a result, the  
480 spectral dependence of  $Im$  can present strong variations. In our study, the computations at 355  
481 nm were performed for four values of the imaginary part of dry particles  $Im_{355}=0.048, 0.03, 0.02,$   
482  $0.01$ . At 532 nm two values  $Im_{532}=0.005$  and  $0.009$  were considered. The parameters of the dry  
483 particle size distribution, the real part of CRI and the hygroscopic growth factor used in  
484 computations are given in Veselovskii et al. (2018). The particles are assumed to be  
485 homogeneous spheres and an increase of the volume for every RH value (calculated from the  
486 growth factor) occurs due to water uptake. Thus both the real and the imaginary part of CRI  
487 depend on RH.

488 The results of the simulations, shown in Fig.19, demonstrate strong dependence of the  
489 organic carbon lidar ratio on the imaginary part of dry particles and on the relative humidity. For  
490  $Im_{355}=0.048$ , for all RH,  $S_{355}$  is above 95 sr, which strongly exceeds the observed values. For  
491 lower  $Im_{355}$  the  $S_{355}(RH)$  dependence is more pronounced and for  $Im_{355}$  within the range 0.01-  
492 0.02, computed  $S_{355}$  are close to observed values. Computed  $S_{532}$  values at low RH exceed the

493 measured ones, but for  $RH > 70\%$  agreement between measurements and GEOS assumed optical  
494 properties for OC becomes reasonable.

495

## 496 **5. Summary and conclusion**

497 Our study shows the impact of the imaginary part variation on the lidar-derived dust  
498 properties. In contrast to extinction, the backscattering coefficient, and so the lidar ratio, are  
499 sensitive to the imaginary part of CRI. Hence,  $S_{355}/S_{532}$  ratio can be an indicator of the imaginary  
500 refractive index enhancement in the UV. Measurements performed during the SHADOW  
501 campaign, in dust conditions, show a correlation between the decrease of  $Im_{440}$ , derived from  
502 AERONET observations, and the decrease of lidar-derived  $S_{355}/S_{532}$  ratio. Namely, in the second  
503 half of April 2015,  $S_{355}/S_{532}$  decreased from 1.5 to 1.0, when  $Im_{440}$  decreased from 0.0045 to  
504 0.0025. Our numerical simulations confirm, that observed  $S_{355}/S_{532}$  (ratio close to 1.5) and  $A_{\beta}$   
505 (value close to -0.75) can be due to spectral variation of the imaginary part, attributed to iron  
506 oxides contained in dust particles. The simulations were performed for the model of randomly  
507 oriented spheroids, however, we should recall, that increase of the particle lidar ratio with the  
508 imaginary part should occur for any particle shape.

509 Thus, April 2015 observations suggest the presence of different dust types, characterized  
510 by distinct spectral dependence of  $Im(\lambda)$ . The analysis of backward trajectories and IDDI derived  
511 from MSG geostationary satellite confirms different air mass and dust particles transport features  
512 in the beginning and at the end of April. Hence, the observed variations of  $S_{355}/S_{532}$  can be  
513 related to the source region mineralogy. During April, the particle depolarization systematically  
514 exceeded 30%, therefore no discrimination between different types of dust was possible.  
515 Dependence of  $S_{355}/S_{532}$  ratio on dust origin, in particular, could explain, why during SAMUM  
516 experiments no significant spectral dependence of the lidar ratio was observed.

517 The results presented in this study demonstrate also that, for the selected temporal  
518 interval, the dust lidar ratios may present significant variation with height. Dust of different size  
519 and mineralogical composition can have different deposition rate, hence, complex refractive  
520 index can be height-dependent. For instance, on April 1<sup>st</sup>, the  $S_{532}$  decreased with height from 60  
521 sr to 50 sr within 1000–3000 m range, while depolarization ratio exceeded 30%. The analysis of  
522 this episode showed that variation of the lidar ratio is entirely attributed to variations of dust  
523 characteristics; the smoke aerosol contribution was insignificant. The data also demonstrate that  
524 a seemingly uniform dust layer may have quite a complex height variation. The results therefore  
525 suggest the relevance of including a varying mineralogy in radiative and climatic modeling of  
526 desert dust impacts. Dust mineralogy should be also taken into account, when possibility of the



527 particle microphysical parameters characterization on a base of multiwavelength lidar  
528 measurements is analyzed (Perez-Ramirez et al., 2019, 2020).

529         During December – January, the dry season in western Africa, our observations allowed  
530 in addition the analysis of biomass burning aerosol properties. These particles are a product of  
531 the seasonal forest fires and intensive agricultural waste combustion and can contain a  
532 substantial amount of organic compounds, characterized by an enhanced imaginary part in UV  
533 (so called BrC). For this aerosol type, the  $\text{Im}(\lambda)$  dependence should increase the lidar ratio at 355  
534 nm and influence  $S_{355}/S_{532}$ . The smoke particles can be also hydrophilic and the lidar ratio can  
535 therefore exhibit a strong dependence on RH. Several strong smoke episodes were observed  
536 during the SHADOW campaign. While we were able to evaluate the RH profiles, the  
537 dependence of the smoke lidar ratio with RH has been estimated. The results obtained should be  
538 taken as semi-qualitative only, due to possible variation of smoke particles composition from  
539 episode to episode and due to the presence of dust particles. Nevertheless, the results clearly  
540 demonstrate an increase of  $S_{532}$  from  $44\pm 5$  sr to  $66\pm 7$  sr and of  $S_{355}$  from  $62\pm 6$  sr to  $80\pm 8$  sr,  
541 when the RH increased from 25% to 85%.

542         We would like to conclude that the multi-wavelengths Raman and depolarization lidar  
543 measurements in western Africa enabled quite unique and comprehensive profiling of dust and  
544 smoke spectral absorption properties. The results demonstrated a high variability of the lidar  
545 ratio and the presence of its spectral dependence. Our study is one of the first attempts to track  
546 aerosol composition variability using lidar measurements and to understand the mechanism  
547 underlying the observed variations. However, the results presented were obtained for a single  
548 region in western Africa. It is important to repeat such studies at different locations around the  
549 world, including the Middle East, Central and East Asia, Australia, and North America in order  
550 to improve our knowledge on real-world dust optical properties, needed in climate relevant  
551 atmospheric modelling.

552

553 **Author contributions.** IV processed the data, and wrote the paper. QH and TP performed the  
554 measurements. PG supervised the project and helped with paper preparation. MK developed  
555 software for data analysis. YD and ML analyzed the satellite data and PC provided MERRA-2  
556 simulations.

557 **Data availability.** Lidar measurements are available upon request (philippe.goloub@univ-  
558 lille.fr).

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573

574 Table 1. Lidar ratios  $S_{355}$ ,  $S_{532}$  for five smoke episodes in December 2015 – January 2016 and  
575 corresponding the relative humidity RH. The table provides also the height and temporal interval

576 of observations. The contribution of the smoke particles to the total backscattering  $\frac{\beta_{532}^s}{\beta_{532}}$  is  
577 derived from depolarization measurements.

578

Date	Height, m	Time, UTC	$\frac{\beta_{532}^s}{\beta_{532}}$	RH, %	$S_{355}$ , sr	$S_{532}$ , sr
15 Dec	2000	04:00-06:00	0.57	42±8	67±7	44±5
15 Dec	1850	19:20-20:30	0.57	25±6	62±6	50±5
23 Dec	2250	05:00-07:00	0.65	65±13	76±8	56±6
24 Dec	3200	19:00-23:00	0.66	75±14	76±8	62±6
20 Jan	4500	01:00-07:00	0.8	85±15	80±8	66±7

579

580

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818 Captures to figures

819 Fig.1. The aerosol optical depth (AOD) at 532 nm (open circles) and AODs of the main aerosol  
820 components, such as dust, organic carbon (OC), black carbon (BC), sulfates (SU) and sea salt  
821 (SS) provided by the MERRA-2 for (a) March, (b) April and (c) December 2015 over Mbour.  
822 Open stars show AOD<sub>532</sub> provided by AERONET.

823

824 Fig.2. Aerosol single scattering albedo (SSA) at 675 nm and 440 nm provided by AERONET for  
825 M'bour site in 2015.

826

827 Fig.3. Tempo-spatial distributions of aerosol backscattering coefficient  $\beta_{532}$  (left column),  
828 particle depolarization ratio  $\delta_{532}$  (middle column) and water vapor mixing ratio (right column),  
829 for the nights 1-2 April (upper row), 2-3 April (middle row) and 3-4 April 2015 (bottom row).

830

831 Fig.4. Three-day backward trajectories for the air mass in M'bour on 2, 3, 4 April 2015 at 03:00  
832 UTC obtained with the HYSPLIT model.

833

834 Fig.5. Vertical profiles of extinction coefficients ( $\alpha_{355}$ ,  $\alpha_{532}$ ) and lidar ratios ( $S_{355}$ ,  $S_{532}$ ) at 355 nm  
835 and 532 nm together with particle depolarization ratio  $\delta_{532}$  measured on 1 April (20:40-22:20  
836 UTC), 2-3 April (23:40-04:30 UTC) and 3-4 April 2015 (23:00-02:00 UTC). Symbols show the  
837 lidar ratios of dust provided by MERRA-2 model ( $S_{355}M$ ,  $S_{532}M$ ).

838

839 Fig.6. Backscattering coefficients for observations presented in Fig.5 for 2-3 and 3-4 April.

840

841 Fig.7. Vertical profiles of the extinction and backscattering Ångström exponents ( $A_\alpha$  and  $A_\beta$ ) at  
842 355 – 532 nm for three temporal intervals from Fig.5.

843

844 Fig.8. Vertical profiles of (a) extinction coefficients at 355 nm and 532 nm ( $\alpha_{355}$ ,  $\alpha_{532}$ ) measured  
845 by lidar (lines) and modeled by MERRA-2 (line+symbol) for five aerosol components at 532  
846 nm; (b) extinction Ångström exponents at 355-532 nm obtained from lidar observations and  
847 modeled by MERRA-2 for pure dust (stars) together with water vapor mixing ratio (WVMR)  
848 and the relative humidity; (c) contribution of dust and smoke particles to  $\beta_{532}$  together with  
849 particle depolarization ratio  $\delta_{532}$ . Values of WVMR are multiplied by factor 10. Lidar  
850 measurements were performed on 3-4 April 2015 for period 23:00 – 02:00 UTC. Modeling  
851 results are given for 4 April 00:00 UTC.

852

853 Fig.9. (a, c) Vertical profiles of extinction coefficients ( $\alpha_{355}$ ,  $\alpha_{532}$ ) and lidar ratios ( $S_{355}$ ,  $S_{532}$ ) at  
854 355 nm and 532 nm; together with (b, d) particle depolarization ratio  $\delta_{532}$ , and extinction and  
855 backscattering Ångström exponents ( $A_\alpha$ ,  $A_\beta$ ) measured on (a, b) 14 April 2015 (00:00 – 05:00  
856 UTC) and (c, d) the night 23-24 April (23:00-06:00 UTC). Open symbols on plots (a, c) show  
857 the lidar ratios  $S_{355}M$  and  $S_{532}M$  provided by MERRA-2 model on 14 and 14 April at 00:00  
858 UTC.

859

860 Fig.10. Four-days backward trajectories for 14 April (03:00 UTC) and 24 April (00:00 UTC)  
861 2015 obtained with the HYSPLIT model.

862

863 Fig.11. Lidar ratios  $S_{355}$ ,  $S_{532}$  and the particle depolarization ratio  $\delta_{532}$  for dust episodes in March  
864 - April 2015. Open triangles show the ratio  $S_{355}/S_{532}$ .

865

866 Fig.12. Imaginary part of the refractive index at 440 nm and 675 nm provided by AERONET in  
867 March – April 2015

868

869 Fig.13. Difference  $Im_{440} - Im_{675}$  from Fig.12 together with lidar measured values  $S_{355}/S_{532}$  from  
870 Fig.11 for days in April 2015

871

872 Fig.14. The particle size distributions provided by AERONET on 2 and 23 April 2015 (three  
873 PSDs for each day). Red line shows the PSD derived from  $3\beta+2\alpha$  lidar measurements on 23-24  
874 April within 2.0 – 3.0 km height range.

875

876 Fig.15. Lidar ratios  $S_{355}$ ,  $S_{532}$  together with the extinction and backscattering Ångström  
877 exponents  $A_\alpha$  and  $A_\beta$  calculated for (a) AERONET PSD on 23 April from Fig.14 and (b) lidar  
878 derived PSD from Fig.14 as a function of the imaginary part. Open stars show  $A_\beta$  for spectrally  
879 dependent imaginary part  $Im(\lambda)$ , assuming that  $Im_{532}=0.002$  is fixed and only  $Im_{355}$  is free vary.  
880 Computations are performed for the assembly of randomly oriented spheroids with the real part  
881  $Re=1.55$ .

882

883 Fig.16. Infrared Difference Dust Index (IDDI) derived from MSG geostationary satellite at noon  
884 time. Panels (a), (b) show IDDI elevated values, representing airborne dust emission and  
885 transport, over central and northern Sahara on 29, 30 March 2015. The dust transport regime is  
886 visibly changed a few days later (17, 18 April 2015, panels (c), (d)); the elevated IDDI values are  
887 shifted to the south. The areas in white are cloud screened pixels; the IDDI is derived only over  
888 land due to the algorithm physical principle.

889

890 Fig.17. Tempo-spatial distributions of aerosol backscattering coefficient  $\beta_{532}$ , particle  
891 depolarization ratio  $\delta_{532}$  and water vapor mixing ratio during smoke episode on the night 14-15  
892 December 2015.

893

894 Fig.18. Vertical profiles of (a) extinction coefficients ( $\alpha_{355}$ ,  $\alpha_{532}$ ) and lidar ratios ( $S_{355}$ ,  $S_{532}$ ); (b)  
895 extinction, backscattering Ångström exponents ( $A_\alpha$ ,  $A_\beta$ ) at 355 – 532 nm and relative humidity  
896 RH; (c) contribution of dust and smoke to  $\beta_{532}$  together with particle depolarization ratio  $\delta_{532}$  on  
897 15 December (04:00 – 06:00 UTC). Black line in plot (a) shows the lidar ratio of smoke  $S_{532}^s$   
898 calculated from Eq.5.

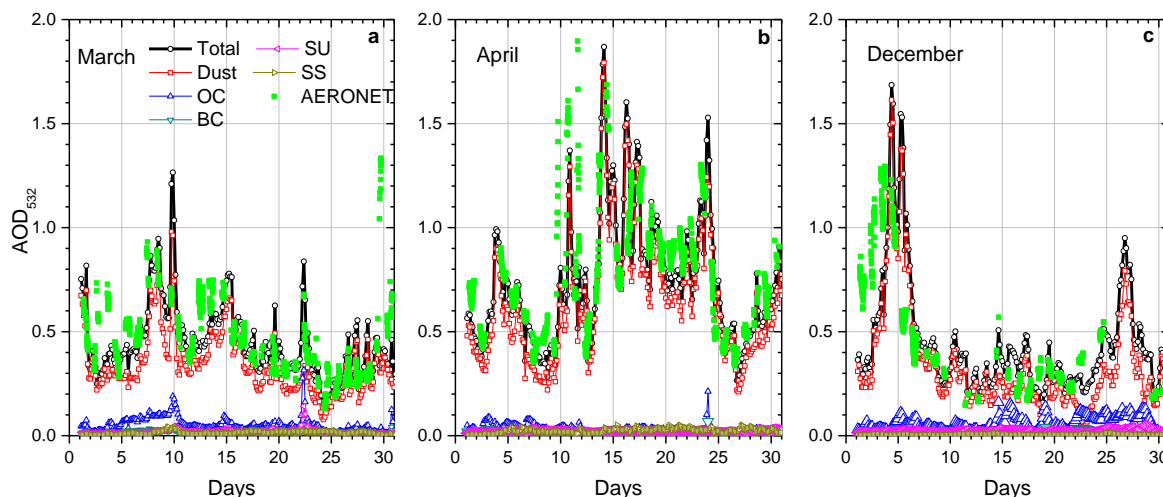
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900 Fig.19. Modeled lidar ratios of organic carbon at 355 nm and 532 nm (line + symbol) as a  
901 function of the relative humidity for the particle parameters used in the MERRA-2 model. At  
902 355 nm results are given for four values of the imaginary part of dry particles:  $Im_{355}= 0.048$ ,  
903 0.03, 0.02, 0.01. At 532 nm two values  $Im_{532}= 0.009$  and 0.005 are considered. The scattered  
904 symbols (circles) show the lidar ratios ( $S_{355}$ ,  $S_{532}$ ) observed during five smoke episodes from  
905 Table 1.

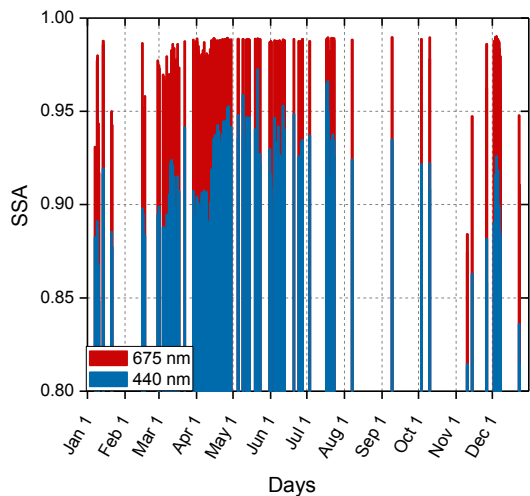
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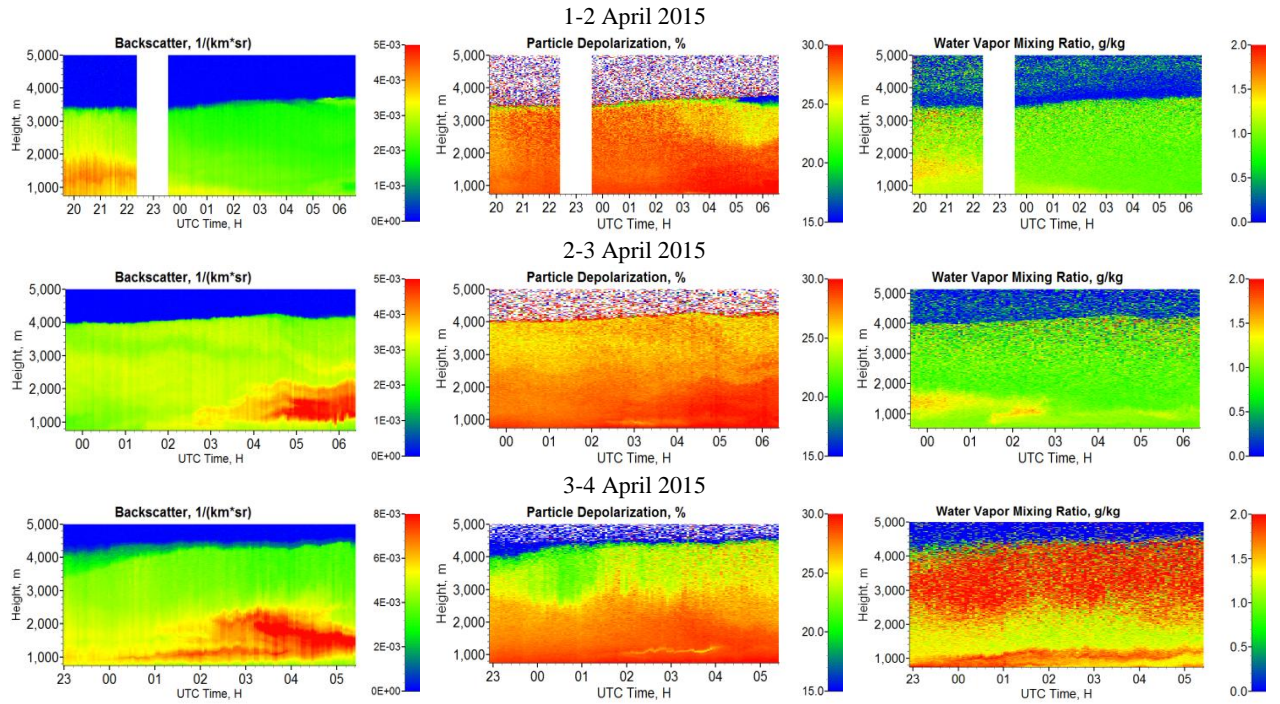


910 Fig.1. The aerosol optical depth (AOD) at 532 nm (open circles) and AODs of the main aerosol  
 911 components, such as dust, organic carbon (OC), black carbon (BC), sulfates (SU) and sea salt  
 912 (SS) provided by the MERRA-2 for (a) March, (b) April and (c) December 2015 over Mbour.  
 913 Open stars show AOD<sub>532</sub> provided by AERONET.  
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916 Fig.2. Aerosol single scattering albedo (SSA) at 675 nm and 440 nm provided by AERONET for  
 917 M'bour site in 2015.  
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922 Fig.3. Tempo-spatial distributions of aerosol backscattering coefficient  $\beta_{532}$  (left column),  
923 particle depolarization ratio  $\delta_{532}$  (middle column) and water vapor mixing ratio (right column)  
924 for the nights 1-2 April (upper row), 2-3 April (middle row) and 3-4 April 2015 (bottom row).  
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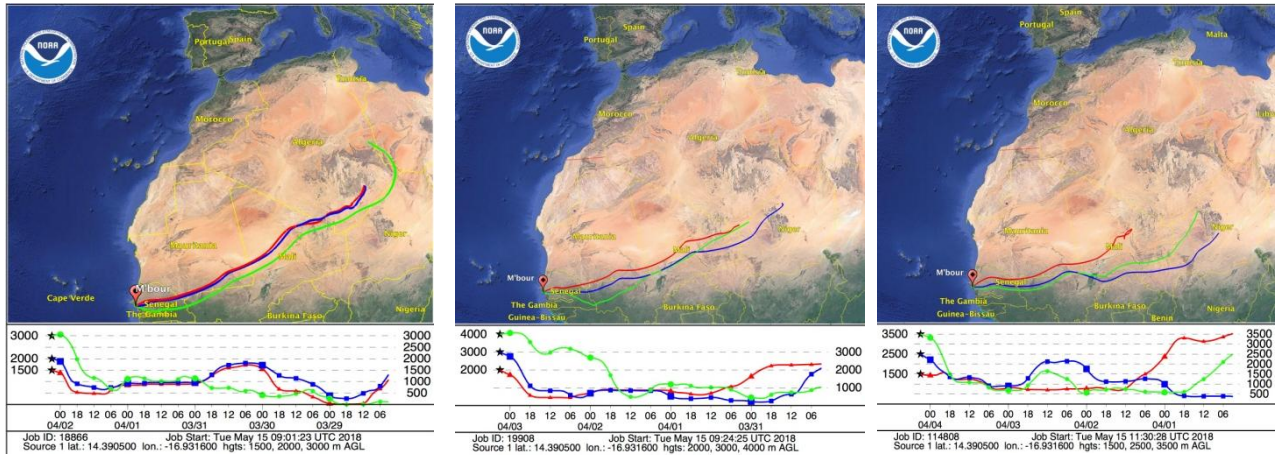
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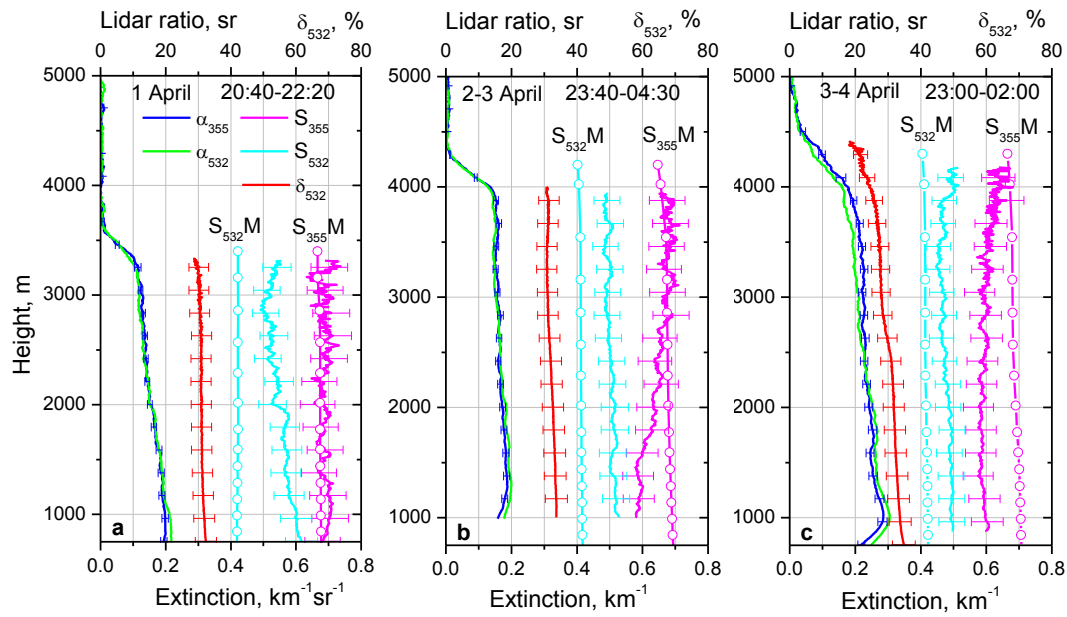
02 April 03:00

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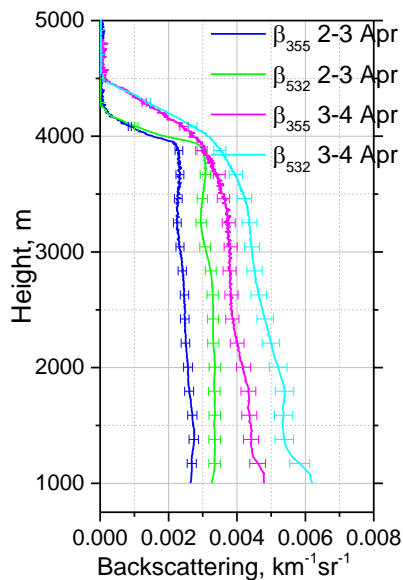


931 Fig.4. Three-day backward trajectories for the air mass in M'bour on 2, 3, 4 April 2015 at 03:00  
932 UTC obtained with the HYSPLIT model.  
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934 Fig.5. Vertical profiles of extinction coefficients ( $\alpha_{355}$ ,  $\alpha_{532}$ ) and lidar ratios ( $S_{355}$ ,  $S_{532}$ ) at 355 nm  
 935 and 532 nm together with particle depolarization ratio  $\delta_{532}$  measured on 1 April (20:40-22:20  
 936 UTC), 2-3 April (23:40-04:30 UTC) and 3-4 April 2015 (23:00-02:00 UTC). Symbols show the  
 937 lidar ratios of dust provided by MERRA-2 model ( $S_{355}M$ ,  $S_{532}M$ ).

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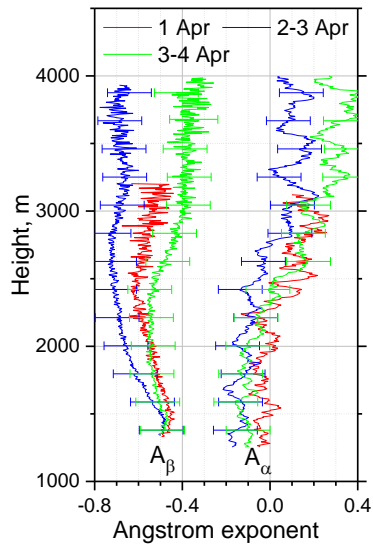
940 Fig.6. Backscattering coefficients for observations presented in Fig.5 for 2-3 and 3-4 April.

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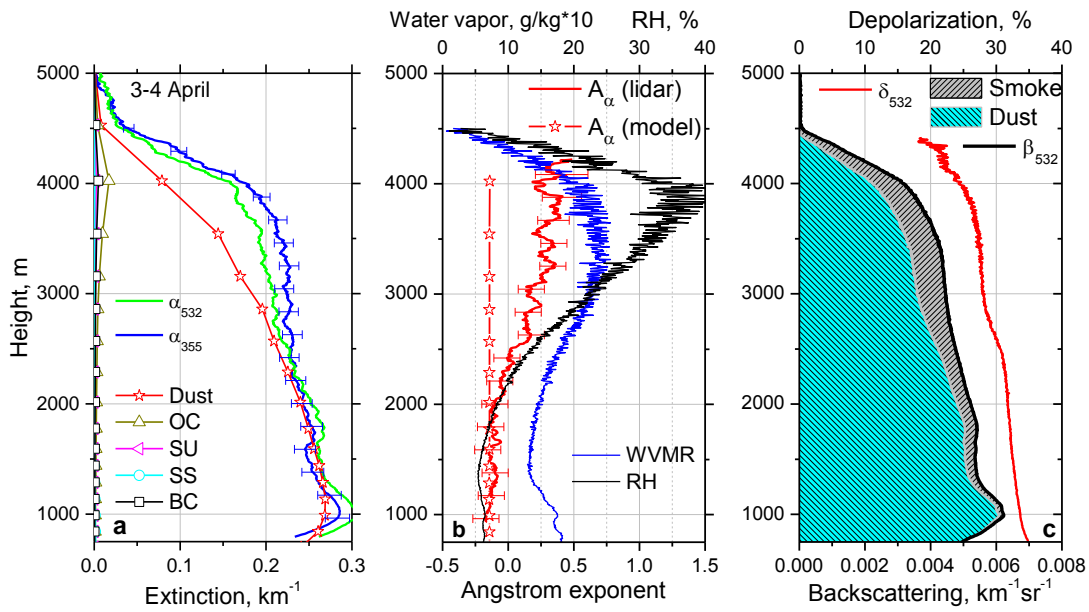
945 Fig.7. Vertical profiles of the extinction and backscattering Ångström exponents ( $A_\alpha$  and  $A_\beta$ ) at  
 946 355 – 532 nm for three temporal intervals from Fig.5.

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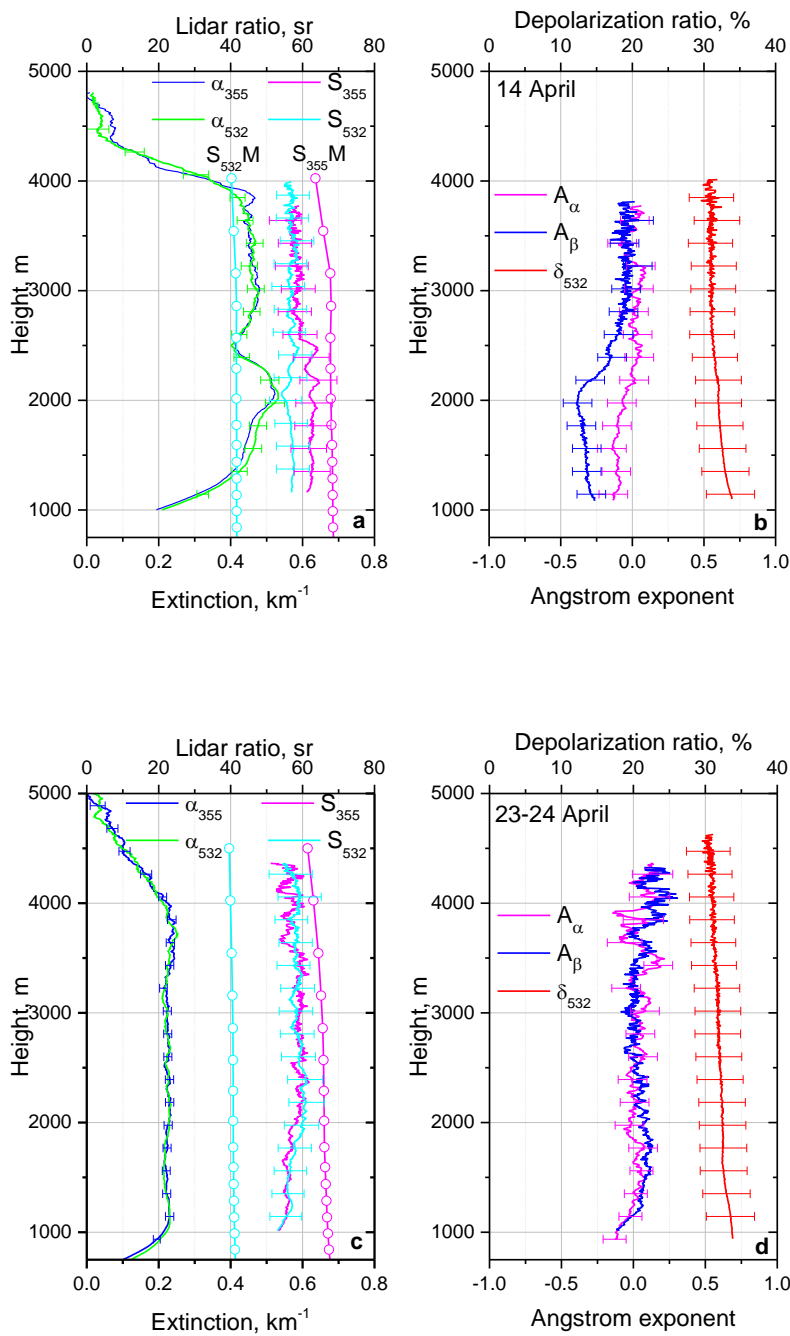


951 Fig.8. Vertical profiles of (a) extinction coefficients at 355 nm and 532 nm ( $\alpha_{355}$ ,  $\alpha_{532}$ ) measured  
 952 by lidar (lines) and modeled by MERRA-2 (line+symbol) for five aerosol components at 532  
 953 nm; (b) extinction Ångström exponents at 355-532 nm obtained from lidar observations and  
 954 modeled by MERRA-2 for pure dust (stars) together with water vapor mixing ratio (WVMR)  
 955 and the relative humidity; (c) contribution of dust and smoke particles to  $\beta_{532}$  together with  
 956 particle depolarization ratio  $\delta_{532}$ . Values of WVMR are multiplied by factor 10. Lidar  
 957 measurements were performed on 3-4 April 2015 for period 23:00 – 02:00 UTC. Modeling  
 958 results are given for 4 April 00:00 UTC.

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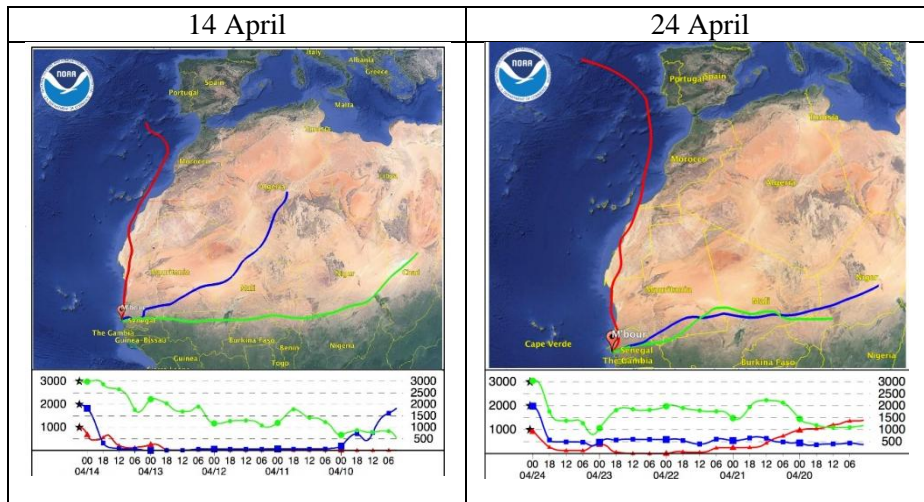
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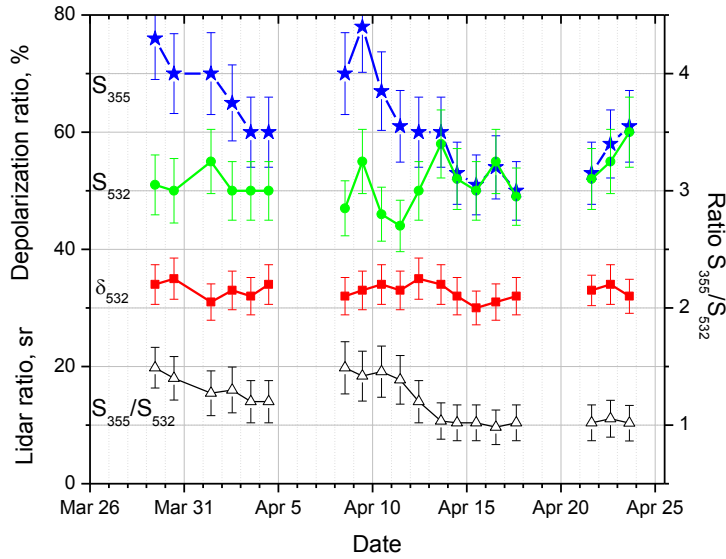
964 Fig.9. (a, c) Vertical profiles of extinction coefficients ( $\alpha_{355}$ ,  $\alpha_{532}$ ) and lidar ratios ( $S_{355}$ ,  $S_{532}$ ) at  
 965 355 nm and 532 nm; together with (b, d) particle depolarization ratio  $\delta_{532}$ , and extinction and  
 966 backscattering Ångström exponents ( $A_\alpha$ ,  $A_\beta$ ) measured on (a, b) 14 April 2015 (00:00 – 05:00  
 967 UTC) and (c, d) the night 23-24 April (23:00-06:00 UTC). Open symbols on plots (a, c) show  
 968 the lidar ratios  $S_{355M}$  and  $S_{532M}$  provided by MERRA-2 model on 14 and 14 April at 00:00  
 969 UTC.  
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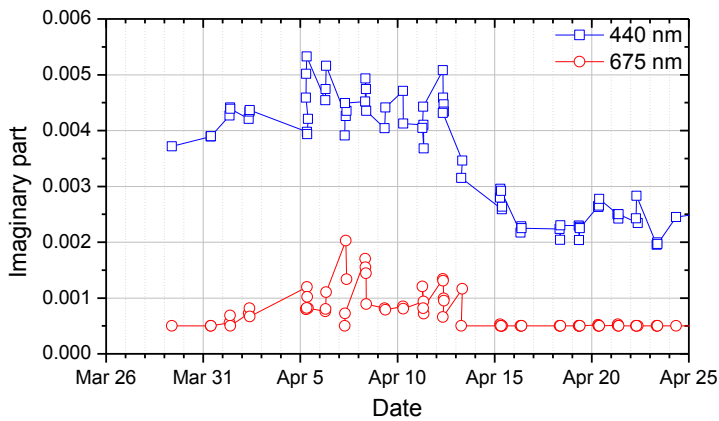
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Fig.10. Four-days backward trajectories for 14 April (03:00 UTC) and 24 April (00:00 UTC) 2015 obtained with the HYSPLIT model.



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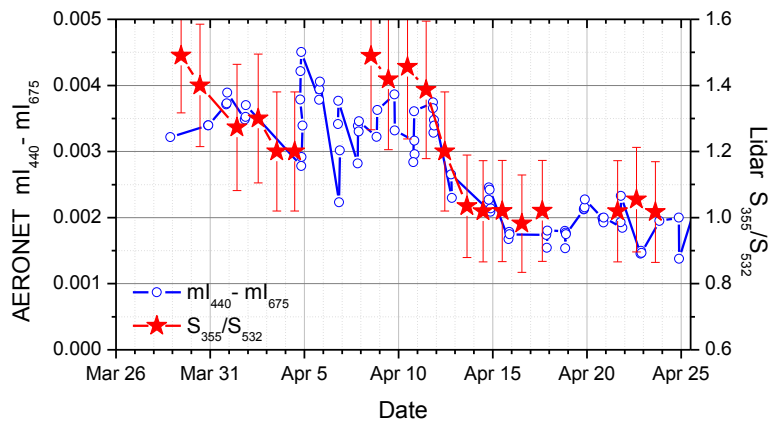
983 Fig.11. Lidar ratios  $S_{355}$ ,  $S_{532}$  and the particle depolarization ratio  $\delta_{532}$  for dust episodes in March  
 984 - April 2015. Open triangles show the ratio  $S_{355}/S_{532}$ .



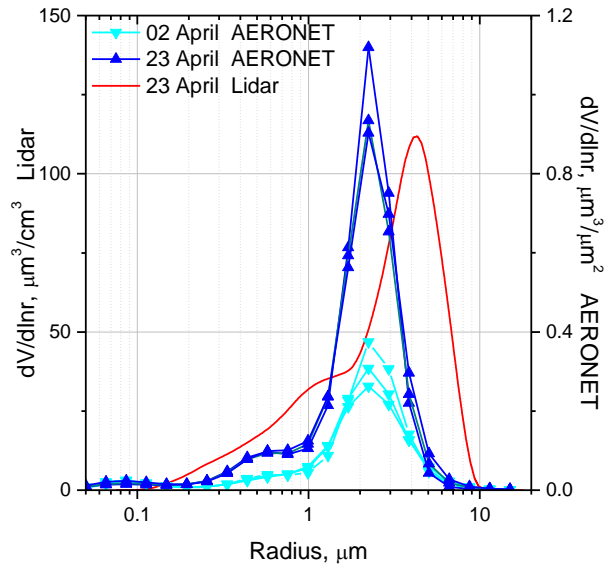
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986 Fig.12. Imaginary part of the refractive index at 440 nm and 675 nm provided by AERONET in  
 987 March – April 2015

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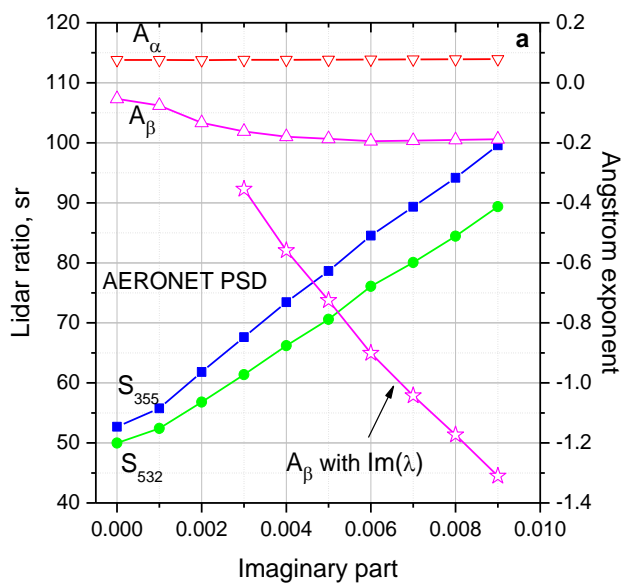
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991 Fig.13. Difference  $Im_{440} - Im_{675}$  from Fig.12 together with lidar measured values  $S_{355}/S_{532}$  from  
992 Fig.11 for days in April 2015.  
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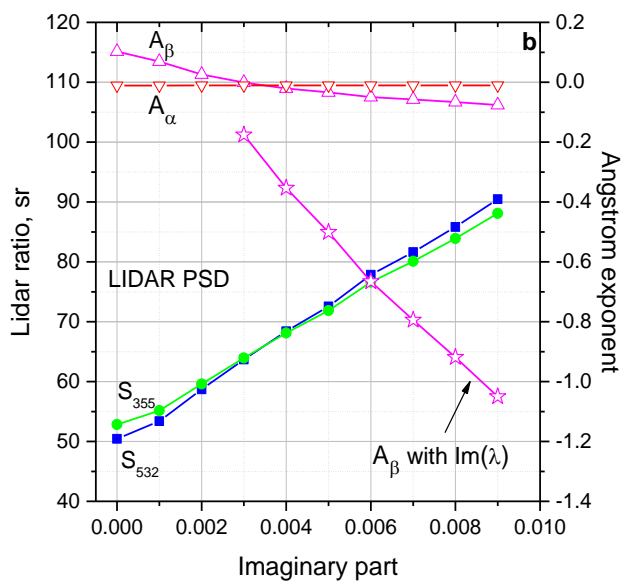
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Fig.14. The particle size distributions provided by AERONET on 2 and 23 April 2015 (three PSDs for each day). Red line shows the PSD derived from  $3\beta+2\alpha$  lidar measurements on 23-24 April within 2.0 – 3.0 km height range.

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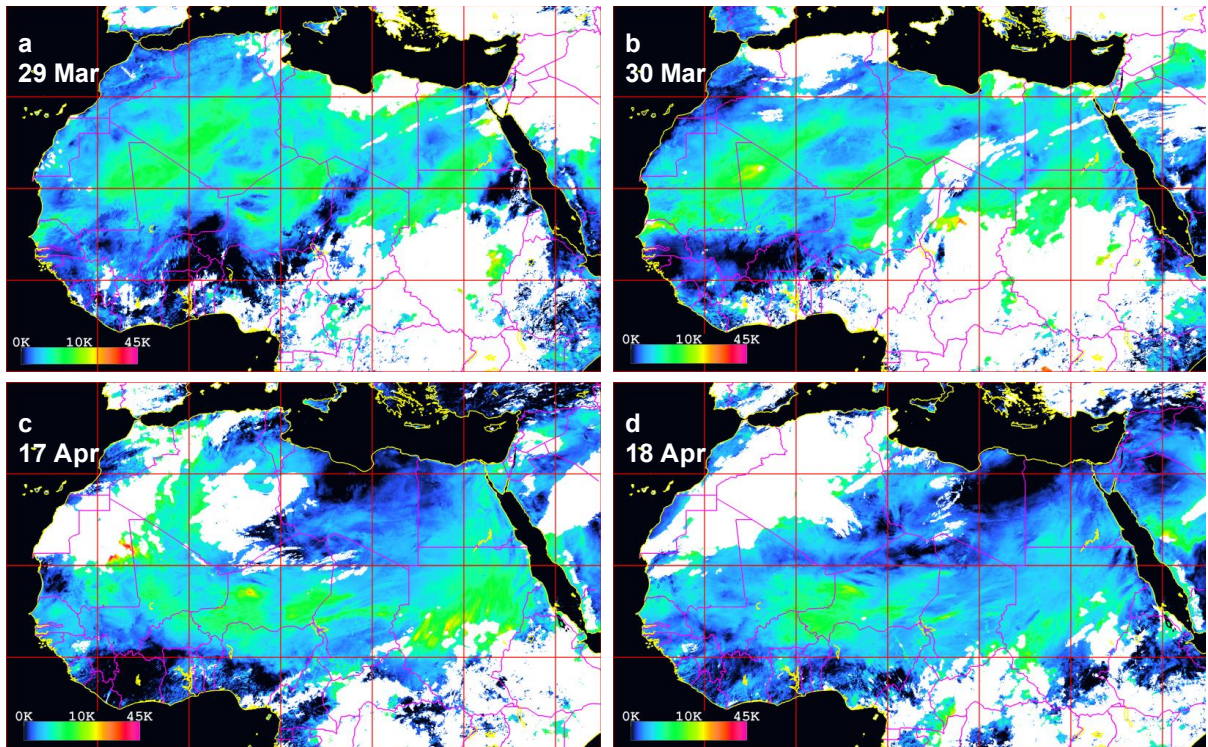
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Fig.15. Lidar ratios  $S_{355}$ ,  $S_{532}$  together with the extinction and backscattering Ångström exponents  $A_\alpha$  and  $A_\beta$  calculated for (a) AERONET PSD on 23 April from Fig.14 and (b) lidar derived PSD from Fig.14 as a function of the imaginary part. Open stars show  $A_\beta$  for spectrally dependent imaginary part  $\text{Im}(\lambda)$ , assuming that  $\text{Im}_{532}=0.002$  is fixed and only  $\text{Im}_{355}$  is free vary. Computations are performed for the assembly of randomly oriented spheroids with the real part  $\text{Re}=1.55$ .

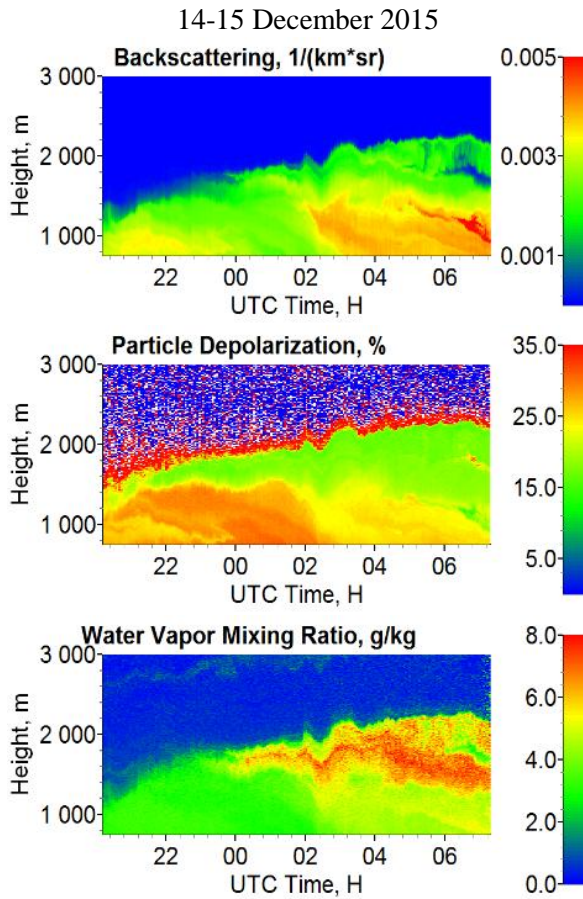




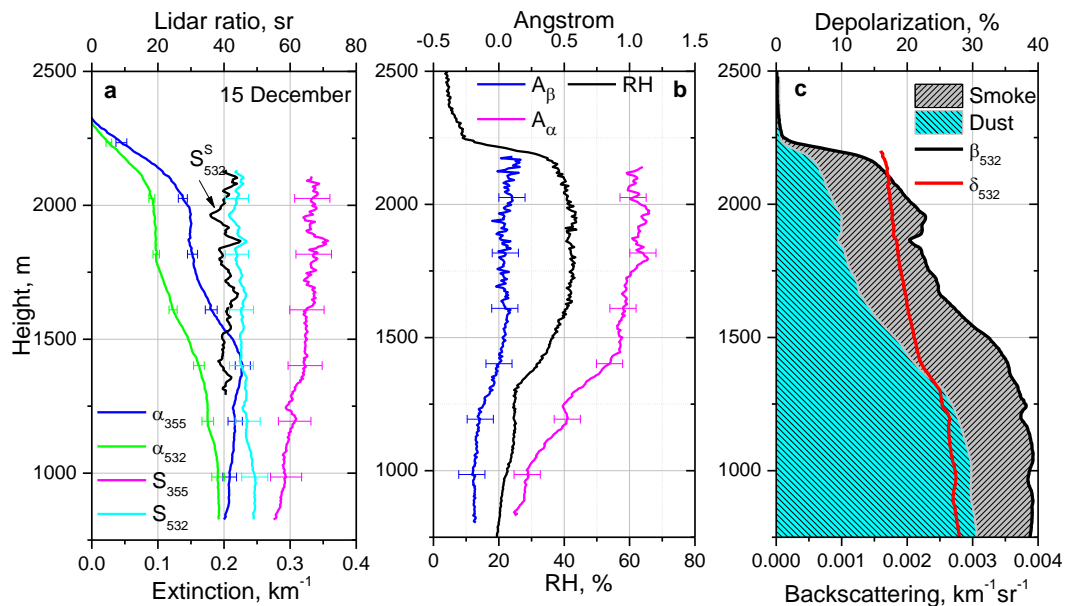
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Fig.16. Infrared Difference Dust Index (IDDI) derived from MSG geostationary satellite at noon time. Panels (a), (b) show IDDI elevated values, representing airborne dust emission and transport, over central and northern Sahara on 29, 30 March 2015. The dust transport regime is visibly changed a few days later (17, 18 April 2015, panels (c), (d)); the elevated IDDI values are shifted to the south. The areas in white are cloud screened pixels; the IDDI is derived only over land due to the algorithm physical principle.

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1025 Fig.17. Tempo-spatial distributions of aerosol backscattering coefficient  $\beta_{532}$ , particle  
1026 depolarization ratio  $\delta_{532}$  and water vapor mixing ratio during smoke episode on the night 14-15  
1027 December 2015.  
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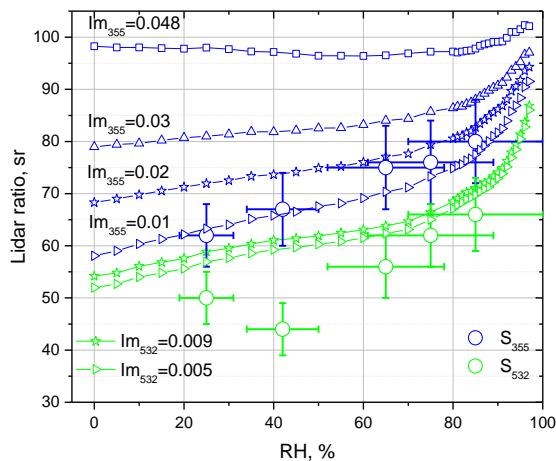


1031 Fig.18. Vertical profiles of (a) extinction coefficients ( $\alpha_{355}$ ,  $\alpha_{532}$ ) and lidar ratios ( $S_{355}$ ,  $S_{532}$ );  
 1032 extinction, backscattering Ångström exponents ( $A_\alpha$ ,  $A_\beta$ ) at 355 – 532 nm and relative humidity  
 1033 RH; (c) contribution of dust and smoke to  $\beta_{532}$  together with particle depolarization ratio  $\delta_{532}$  on  
 1034 15 December (04:00 – 06:00 UTC). Black line in plot (a) shows the lidar ratio of smoke  $S_{532}^s$   
 1035 calculated from Eq.5.

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Fig.19. Modeled lidar ratios of organic carbon at 355 nm and 532 nm (line + symbol) as a function of the relative humidity for the particle parameters used in the MERRA-2 model. At 355 nm results are given for four values of the imaginary part of dry particles:  $Im_{355} = 0.048, 0.03, 0.02, 0.01$ . At 532 nm two values  $Im_{532} = 0.009$  and  $0.005$  are considered. The scattered symbols (circles) show the lidar ratios ( $S_{355}, S_{532}$ ) observed during five smoke episodes from Table 1.