

1 **The Impact of Increasing Stratospheric Radiative Damping on the QBO Period**

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10 **Abstract.** Stratospheric radiative damping increases as atmospheric carbon dioxide concentration rises.

11 We use the one-dimensional mechanistic models of the QBO to conduct sensitivity experiments and
12 find that when atmospheric carbon dioxide concentration increases, the simulated QBO period shortens
13 due to the enhancing of radiative damping in the stratosphere. This result suggests that increasing
14 stratospheric radiative damping due to rising CO₂ may play a role in determining the QBO period in a
15 warming climate along with wave momentum flux entering the stratosphere and tropical vertical
16 residual velocity, both of which also respond to increasing CO₂.

17

18 **1. Introduction**

19 The quasi-biennial oscillation (QBO) dominates the variability of the equatorial middle and lower
20 stratosphere and is characterized by a downward propagating zonal wind regime that regularly changes
21 from westerlies to easterlies. The QBO period ranges from 22 to 34 months with its average being
22 slightly longer than 28 months. The QBO not only manifests itself in the equatorial zonal winds, but also
23 leaves an imprint on the temperature in both the tropics and extratropics (Baldwin et al., 2001 and
24 references therein).

25 The QBO has far-reaching implications for global weather and climate systems. First of all, the QBO
26 exerts a marked influence on the distribution and transport of various chemical constituents such as

27 ozone (O_3) (e.g., Hasebe, 1994), water vapor (H_2O) (e.g., Kawatani et al., 2014), methane (CH_4), nitrous
28 oxide (N_2O), hydrogen fluoride (HF), hydrochloric acid (HCl), odd nitrogen species (NO_y) (e.g.,
29 Zawodny and McCormick, 1991), and volcanic aerosol (Trepte and Hitchman, 1992). Secondly, it is
30 well appreciated that the QBO influences the extratropical circulation in the winter stratosphere, which
31 is commonly known as the Holton–Tan effect (Holton and Tan, 1980; Labitzke, 1982). It has been noted
32 that the effect of the QBO on the extratropical winter stratosphere impacts the severity of stratospheric
33 ozone depletion (e.g., Lait et al., 1989). Furthermore, taking account of the QBO improves the simulation
34 and predictability of the extratropical troposphere (e.g., Marshall and Scaife, 2009). Finally, through its
35 modulation of temperature and vertical wind shear in the vicinity of the tropical tropopause, the QBO
36 influences tropical moist convection (Collimore et al., 2003; Liess and Geller, 2012), the El Niño–
37 Southern Oscillation (ENSO) (Gray et al., 1992; Huang et al., 2012; Hansen et al. 2016), the Hadley
38 circulation (Hitchman and Huesmann, 2009), the tropospheric subtropical jet (Garfinkel and Hartmann,
39 2011a, 2011b), the boreal summer monsoon (Giorgetta et al., 1999), and the Madden-Julian Oscillation
40 (Yoo and Son, 2016). Intriguingly, the QBO is also reported to influence the activities of tropical
41 cyclones (Gray et al., 1984; Ho et al., 2009), albeit this issue is still unsettled (Camargo and Sobel, 2010)
42 and needs further study.

43 Efforts to understand and simulate the QBO have been ongoing ever since its discovery by Ebdon
44 (1960) and Reed et al. (1961). Lindzen and Holton (1968) and Holton and Lindzen (1972) developed
45 the classical theory of the QBO. Namely, as waves propagate upward, they are attenuated by thermal
46 damping, encounter critical levels, and accelerate and decelerate the mean flow, providing momentum
47 sources for both the westerly and easterly phases of the QBO.

48 Holton and Lindzen’s (1972) model (hereafter referred to as HL model) was further simplified by
49 Plumb (1977), the elegance of which made it a standard paradigm for the QBO. In Plumb’s (1977)

50 Boussinesq formulation, the QBO period is inversely dependent upon both the momentum flux and
51 thermal dissipation rate. Hamilton (1981) further highlighted the role of the radiative damping rate on
52 both the realistic vertical structure and the realistic period of the QBO.

92 By adopting higher vertical resolutions and incorporating various gravity wave parameterization
93 schemes, many state-of-the-art climate models have shown the capability to self-consistently simulate
94 the QBO (Scaife et al., 2000; Giorgetta et al., 2002, 2006; Rind et al., 2014, 2020; Geller et al., 2016a;
95 Richter et al., 2020a, 2020b). Given the important implications of the QBO for the global climate system,
96 it is natural to ask how the QBO will change in a warming climate.

97 Giorgetta and Doege (2005) showed a shortening of the QBO period in their doubled CO₂
98 experiments. They reasoned that both the weakening of the tropical upwelling and the prescribed
99 increase of gravity wave sources lead to the reduction of the QBO period in a warming climate. However,
100 most climate models project a strengthening rather than weakening of tropical upwelling in a warmer
101 climate (Butchart et al., 2006; Butchart 2014; Li et al., 2008). Employing a model without any
102 parametrized non-orographic gravity waves, Kawatani et al. (2011) demonstrated that the intensifying
103 tropical upwelling in a warming climate dominates the counteracting effect of enhanced wave fluxes and
104 consequently projected a lengthening of the QBO period. Using fixed sources of parametrized gravity
105 waves, Watanabe and Kawatani (2012) also projected an elongation of the QBO period in a warming
106 climate and ascribed it to the stronger tropical upwelling. Analyzing four Coupled Model
107 Intercomparison Project phase 5 (CMIP5) models that could simulate a reasonable QBO, Kawatani and
108 Hamilton (2013) found that the projected trends of the QBO period were inconsistent in sign. They
109 further investigated the 60-year operational balloon-borne radiosonde observations provided by the Free
110 Berlin University and detected no significant trend in the QBO period. Richter et al. (2020b) investigated
111 the response of the QBO to doubled and quadrupled CO₂ climates among eleven models that participated

112 in Phase 1 of the Stratospheric-tropospheric Processes And their Role in Climate QBO-initiative (QBOi;
113 Butchart et al., 2018), and found no consensus on how the QBO period would respond to a changing
114 climate. Recently, Butchart et al. (2020) evaluated ten Coupled Model Intercomparison Project phase 6
115 (CMIP6) models with realistic QBO in two Shared Socioeconomic Pathways (SSPs, Gidden et al., 2019)
116 scenario simulations and surprisingly found that the QBO period shortens in seven of those ten models
117 in both in both SSP3-7.0 and SSP5-8.5 scenarios although only two and three models show a significant
118 shortening trend in the respective scenarios.

119 It is challenging to ascertain the trend of the QBO period in a warming climate. On one hand, a
120 speeding-up of the Brewer-Dobson circulation in a warming climate leads to a lengthening of the QBO
121 period in most climate models. On the other hand, there is a robust increase in the vertical component of
122 the EP flux for both eastward and westward propagating waves (Richter et al., 2020b; Butchart et al.,
123 2020), indicating that the QBO period shortens due to the enhanced wave driving in a warming climate.
124 The competing effects between enhanced wave driving and a faster Brewer-Dobson circulation suggests
125 that trends in the QBO period are likely to be small and difficult to detect due to the large cycle-to-cycle
126 variability that is reproduced by climate models (Butchart et al., 2020). In addition, uncertainty in the
127 representation of the parameterized gravity waves make it more elusive to detect the trend of the QBO
128 period in a warming climate (Schirber et al., 2015; Richter et al., 2020b).

129 Given the fact that the QBO period is influenced by the radiative damping (Plumb 1977; Hamilton
130 1981), a natural question to ask is whether it could play a role on the trend of the QBO in a warming
131 climate. Plass (1956) showed that when the CO₂ concentration is increased from 330 ppmv to 660 ppmv,
132 the cooling rate increases significantly in the middle and upper stratosphere while it is not changed below
133 the 24 km height level. The cooling rate is increased by more than 50% around the 40 km height level
134 (see his Figure 8).

135 It is well-known that enhanced wave fluxes entering the stratosphere and stronger tropical upwelling
 136 individually play a dominant role in determining the trends in the QBO period in a warming climate.
 137 Does the competing effect between them leave some room for increasing stratospheric radiative damping
 138 to exert an influence on the QBO period? In this paper, we use the HL model to isolate the effect of
 139 radiative damping on the QBO period by assuming that the momentum flux entering the stratosphere
 140 doesn't change in our experiments. Observational and modeling studies (Andrews et al., 1987; Kawatani
 141 et al., 2009, 2010, 2011; Richter et al., 2020b; Holt et al., 2020) showed that the wave forcing spectrum
 142 is similar to a discrete two-wave spectrum rather than red-noise or white-noise, all of which are
 143 illustrated in Saravanan (1990). Accordingly, the QBO is indeed sensitive to stratospheric radiative
 144 damping, and the HL model is suitable for us to conduct the sensitivity analysis.

145 The remainder of this paper is organized as follows. Section 2 investigates the sensitivity of the QBO
 146 period to the radiative damping using HL's original model. Section 3 explores the sensitivity of the QBO
 147 period to the radiative damping using a modified HL model where the semiannual forcing is removed.
 148 Discussion and conclusions are presented in Sections 4 and 5 respectively.

149

150 **2. Sensitivity of the QBO period to enhanced stratospheric radiative damping in the original HL** 151 **model**

152 In the HL model the governing equation of mean flow emerges after the primitive momentum
 153 equation is meridionally averaged over some suitable latitudinal belt over the equator.

$$154 \quad \frac{\partial \bar{u}}{\partial t} = -\frac{1}{\rho_0} \frac{\partial}{\partial z} \left[\sum_{i=0}^1 \bar{F}_i \right] + K_z \frac{\partial^2 \bar{u}}{\partial z^2} + G \quad (1)$$

155 where \bar{u} is mean zonal wind, ρ_0 is mean density, \bar{F}_i is the meridionally averaged vertical Eliassen-Palm
 156 flux associated with wave i , the index i refers to the individual waves, K_z is a vertical eddy diffusion
 157 coefficient, t is time, z is altitude, and G is semiannual forcing identical to that specified by HL.

158 The \bar{F}_i is are evaluated with Lindzen's (1971) WKB formalism for equatorial waves in shear. When
 159 only infrared cooling acts to damp the waves the formulae for \bar{F}_i are

$$160 \quad \bar{F}_0(z) = A_0 \exp\left(-\int_{17km}^z \frac{\alpha N}{k(c - \bar{u})^2} dz\right) \quad (2)$$

161 for the Kelvin wave, and

$$162 \quad \bar{F}_1(z) = A_1 \exp\left[-\int_{17km}^z \frac{\alpha\beta N}{k^3(c - \bar{u})^3} \left(1 - \frac{k^2(\bar{u} - c)}{\beta}\right) dz\right] \quad (3)$$

163 for the mixed Rossby-gravity wave. As in HL, the wavenumber k , the phase speed c , and A_0 are chosen
 164 to be $2\pi/(40,000 \text{ km})$, 30 m s^{-1} , and $0.04 \text{ m}^2 \text{ s}^{-2} \rho_0(17 \text{ km})$, respectively for the Kelvin wave while
 165 they are equal to $-2\pi/(10,000 \text{ km})$, -30 m s^{-1} , and $-0.04 \text{ m}^2 \text{ s}^{-2} \rho_0(17 \text{ km})$, respectively for the
 166 mixed Rossby-gravity wave. In Eq. (1), $K_z = 0.3 \text{ m}^2 \text{ s}^{-1}$, which is also the same as in HL. In addition,
 167 $\beta = 2\Omega/a$, where Ω is earth's rotation rate, and a is earth's radius. HL's boundary conditions stipulated
 168 that $\bar{u} = 0$ at the lowest model level (17 km) and constrained \bar{u} to vary semiannually at the top level (35
 169 km).

170 In our control run that is used to depict the present-day QBO all the model parameters are identical
 171 to those used by HL in their original simulation. The Brunt-Väisälä frequency

$$172 \quad N = \sqrt{\frac{g}{T_0} \left(\frac{dT_0}{dz} + \frac{g}{c_p} \right)} \quad (4)$$

173 In Eq. (4), g is gravity, T_0 is mean temperature, and c_p is specific heat of dry air at constant pressure.

174 HL set N in Eq. (4) to $2.16 \times 10^{-2} \text{ s}^{-1}$ with a scale height $H = 6 \text{ km}$. In addition, the Newtonian

175 cooling profile in our control run, i.e., $\alpha(z)$ in Eqs. (2) and (3), is also identical to that in the original
176 HL model and depicted in FIG. 1 as the black line. Namely, $\alpha(z)$ in the control run increases from
177 $(21 \text{ day})^{-1}$ at 17 km to $(7 \text{ day})^{-1}$ at 30 km and is kept at $(7 \text{ day})^{-1}$ between 30 km and 35 km. Fels
178 (1985) explicated why this cooling rate is suitable for simulating the QBO on the basis of the scale-
179 dependent effect of radiative damping (Fels 1982). Hamilton (1981) demonstrated that the proper choice
180 of $\alpha(z)$ is crucial in simulating a realistic vertical structure of the QBO.

181 Eq. (1) was integrated for 100 years using the forward-backward scheme (Matsuno, 1966). The
182 vertical resolution was 250 m and identical to that in HL. The time step was 12 hr, i.e., one half of used
183 in HL, because the 24-hr time step resulted in numerical instability in our integration.

184 FIG. 2a shows the time–height section of the monthly averaged mean zonal wind simulated over the
185 first 20 years using the HL model. Both the QBO and the semiannual oscillation (SAO) are conspicuous.
186 The fast Fourier transform (FFT) method is used to calculate the frequency power spectra. In order to
187 more accurately derive the QBO period, the model was run for 100 years to increase the spectral
188 resolution. Frequency–height sections of the power spectral densities (PSD) over zero to the Nyquist
189 frequency, i.e., 0.5 cycle/month, depict two sharp lines (peaks) at $\frac{1}{30}$ and $\frac{1}{6}$ cycle/month, respectively
190 (not shown). In order to better visualize the magnitudes of the PSD, we show two truncated frequency–
191 height sections with FIG. 2b and FIG. 2c highlighting the QBO and the SAO respectively. FIG. 2b shows
192 that the QBO dominates over the model domain. The peak frequency corresponds to the period of 30
193 months. FIG. 2c shows the SAO dominates near the model top due to the fact a semiannual forcing was
194 imposed in the altitudes from 28 to 35 km.

195 It is worth mentioning that the QBO period shown here is longer than 26.5 months reported in the HL
196 paper (see their FIG. 1). Using the HL model parameters, the QBO period simulated by Plumb (1977)
197 was close to three years (refer to his FIG. 8a), which is longer than our simulated QBO period, i.e., 30.0

198 months. Although we could not explain why our simulated QBO period is longer than that simulated by
 199 HL, we found that when the upper boundary condition is changed from $\bar{u} = 14 \sin(\omega_a t)$ and $\omega_a =$
 200 $\frac{2\pi}{180} \text{ day}^{-1}$ used in the HL's original model (refer to their Eqs. (2)) to $\frac{\partial \bar{u}}{\partial z} = 0$ used in Plumb (1977), the
 201 simulated QBO period becomes 34.3 month (figure not shown). In other words, when we adopted the
 202 stress-free upper boundary condition as in Plumb (1977), our simulated QBO period is comparable to
 203 that simulated by him, which lends credence to our reconstruction of the HL model.

204 As mentioned in Section 1, when the atmospheric carbon dioxide concentration is doubled the cooling
 205 rate increases significantly in the middle and upper stratosphere while it does not change below the 24
 206 km height level. The cooling rate is increased by more than 50% around the 40 km height level (Plass,
 207 1956). As implied in Dickinson (1973), the estimated cooling coefficient below the 0.2 hPa level is
 208 approximately proportional to the estimated cooling rate and is not sensitive to the chosen temperature
 209 profile. In other words, the relative increase in the cooling coefficient is roughly equal to the relative
 210 increase in the cooling rate as the CO₂ concentration is doubled. Accordingly, the Newtonian cooling
 211 profile in our experimental run, i.e., $\alpha(z)$ in Eqs. (2) and (3), is specified in FIG. 1 as the red line. Namely,
 212 $\alpha(z)$ in the experimental run increases from $(21 \text{ day})^{-1}$ at 17 km to $\frac{9}{91} \text{ day}^{-1}$ at 24 km, which is
 213 identical to that in the control run from 17 km to 24 km. We increased $\alpha(z)$ in the experimental run
 214 between 30 km and 35 km by 30% relative to that in the control run. In other words, $\alpha(z)$ is kept at
 215 $\frac{1.3}{7} \text{ day}^{-1}$ between 30 km and 35 km in the experimental run. This percentage increase of 30% in $\alpha(z)$
 216 for the doubled CO₂ above 30 km shown in FIG. 1 is somewhat less than that implied in the Figure 8 of
 217 Plass (1956), because we would like our estimated relative increase in $\alpha(z)$ to err on the conservative
 218 side given the inherent uncertainties mentioned by him. Between 24 km and 30 km, $\alpha(z)$ in the
 219 experimental run is formulated linearly with height from $\frac{9}{91} \text{ day}^{-1}$ at 24 km to at $\frac{1.3}{7} \text{ day}^{-1}$ at 30 km.

220 FIG. 3a shows the time–height section of the monthly averaged mean zonal wind simulated over the
221 first 20 years for the doubled CO₂ run, where the increased $\alpha(z)$ depicted as the red line in FIG. 1 was
222 employed while all other parameters are identical to those in the control run. Obviously, the QBO
223 dominates below 28 km while the semiannual oscillation (SAO) dominates above 31 km. Like FIG. 2b
224 and FIG. 2c, we only show two truncated frequency–height sections with FIG. 3b highlighting the QBO
225 and FIG. 3c highlighting the SAO. FIG. 3b also shows that the QBO prevails over the model domain. The
226 peak frequency corresponds to the period of 27.9 months. FIG. 3c shows the SAO dominates near the
227 model top due to the same imposed semiannual forcing as that in the control run.

228 In summary, using the original HL model we found that the increased radiative damping due to the
229 doubling of CO₂ shortens the QBO period by 7% (i.e., decreases from 30 months to 27.9 months).

230

231 **3. Sensitivity of the QBO period to enhanced stratospheric radiative damping in the HL model** 232 **without the semiannual forcing**

233 HL pointed out that the imposed semiannual oscillation was not essential for their QBO theory.
234 Applying $\frac{\partial \bar{u}}{\partial z} = 0$ as the upper boundary condition, Plumb (1977) showed a simulated QBO without
235 resorting to the semiannual momentum source (refer to his FIG. 8b). In the following control run, all
236 parameters are identical to those used in the previous control run in Section 2 except that G in Eq. (1) is
237 set to zero with $\frac{\partial \bar{u}}{\partial z}$ also being set to zero at $z = 35$ km. Hereafter we refer to it as the Plumb model¹. FIG.
238 4a shows the time–height section of the monthly averaged mean zonal wind simulated over the first 20
239 years using the Plumb model. As expected, the QBO emerges without any trace of SAO since $G = 0$ in
240 Eq. (1). FIG. 4b shows that the QBO dominates over the whole model domain. The peak frequency

¹ Strictly speaking, it is the HL model modified by Plumb (1977). In this paper, we don't use his eponymous model, i.e., the simplest possible model of the QBO, where Boussinesq fluids with uniform mean density were employed, because the HL model and its variant are considerably more realistic.

241 corresponds to the period of 37.5 months, which is comparable to that simulated by Plumb (1977) shown
242 in his FIG. 8b. Apparently, the QBO period from the Plumb model, i.e., 37.5 months shown in FIG. 4b,
243 is longer than that from the HL model, i.e., 30.0 months shown in FIG. 2b. This is partly because the
244 additional forcing G in Eq. (1) was removed in the Plumb model.

245 In the following experimental run, all parameters are identical to those used in the previous
246 experimental run in Section 2 except that G in Eq. (1) is set to zero with $\frac{\partial \bar{u}}{\partial z}$ also being set to zero at $z =$
247 35 km. In other words, the following experimental run using the Plumb model employed the same
248 parameters as the afore-mentioned control run using the Plumb model except that the increased $\alpha(z)$
249 shown as the red line in FIG. 1 was used in the following experimental run while $\alpha(z)$ shown as the
250 black line in FIG. 1 was used in the above control run. FIG. 5a shows the time–height section of the
251 monthly averaged mean zonal wind simulated over the first 20 years for the doubled CO₂ run. It is natural
252 that only the QBO emerges. A comparison of FIG. 4a and FIG. 5a shows that the QBO period shortens
253 when the infrared damping increases in response to the doubled CO₂. FIG. 5b shows that the QBO
254 dominates over the whole model domain. The peak frequency corresponds to the period of 31.6 months.

255 Using the Plumb model, we found that the increased radiative damping due to the doubling of CO₂
256 shortens the QBO period by 15.7% (i.e., decreases from 37.5 months to 31.6 months).

257

258 4. Discussion

259 The semiannual forcing, G in Eq. (1), in the HL model is imposed rather than results from the wave-
260 flow interaction. In other words, G in Eq. (1) is independent of mean flow, and is specified as $G =$
261 0 for $z \leq 28$ km, and $G = \omega_{sa} \bar{u}_{sa}$ for $z > 28$ km

262 where $\bar{u}_{sa} = 2(z - 28\text{km}) \text{ m s}^{-1} \text{ km}^{-1} \sin(\omega_{sa} t)$ and $\omega_{sa} = \frac{2\pi}{180} \text{ day}^{-1} \approx 4 \times 10^{-7} \text{ s}^{-1}$ (refer to
 263 Eqs. (2) in HL). Therefore, we have $\frac{\partial^2 \bar{u}_{sa}}{\partial z^2} = 0$ in the HL original model. We furthermore decompose \bar{u}
 264 into two components: \bar{u}_{QBO} and \bar{u}_{sa} . Combining Eq. (1), the decomposition of \bar{u} as $\bar{u} = \bar{u}_{QBO} + \bar{u}_{sa}$,
 265 the above-mentioned $\frac{\partial^2 \bar{u}_{sa}}{\partial z^2} = 0$, and $G = \omega_{sa} \bar{u}_{sa} = \frac{\partial \bar{u}_{sa}}{\partial t}$ for $z > 28 \text{ km}$, yields

$$266 \quad \frac{\partial \bar{u}_{QBO}}{\partial t} = -\frac{1}{\rho_0} \frac{\partial}{\partial z} \left[\sum_{i=0}^1 \bar{F}_i \right] + K_z \frac{\partial^2 \bar{u}_{QBO}}{\partial z^2} \quad (5)$$

267 for $z > 28 \text{ km}$.

268 Dunkerton (1997) showed that in the presence of tropical upwelling it was gravity waves rather than
 269 large-scale Kelvin and mixed Rossby-gravity waves that contributed the bulk of QBO forcing.
 270 Consequently, Geller et al. (2016a, 2016b) pointed out that enough gravity wave momentum flux is
 271 required to model the QBO in a self-consistent manner in climate models and that the magnitude of the
 272 subgrid-scale gravity wave momentum flux plays a crucial role in determining the QBO period. Since
 273 there is no tropical upwelling in either the HL model or the Plumb model, and the semiannual forcing,
 274 G , is dependent on neither \bar{u} in Eq. (1) nor \bar{u}_{QBO} in Eq. (5), it is natural that planetary-scale Kelvin and
 275 mixed Rossby-gravity waves largely determine the QBO periods shown in Sections 2 and 3 due to the
 276 fact that G only exerts a weak influence on the planetary wave forcing, i.e., $-\frac{1}{\rho_0} \frac{\partial}{\partial z} \left[\sum_{i=0}^1 \bar{F}_i \right]$ in Eqs. (1)
 277 and (5). We conducted another sensitivity test where all parameters are identical to those in the HL model
 278 except that G in both the control and experimental runs is twice as large as that used by HL. As the
 279 radiative damping profile changes from the black line to the red line above 24 km shown in FIG. 1, our
 280 simulated QBO period decreases from 28.6 months to 27.3 months (figures not shown). This smaller
 281 percentage decrease of 4.5% is not unexpected because the unrealistically larger G that is independent
 282 of \bar{u} makes the model atmosphere less sensitive to the changes in the radiative damping.

283 We further conducted two sensitivity tests where all parameters are identical to those in the HL model
284 except that G in the first test is half as large as that used by HL and is equal to zero in the second test.
285 Surprisingly, as the radiative damping profile changes from the black line to the red line above 24 km
286 shown in FIG. 1, our simulated QBO periods decreases from 30.0 months to 28.6 months both for G
287 being decreased by 50% and for $G = 0$ (figures not shown). This 4.7% decrease in the QBO period is
288 somewhat smaller than the 7% reduction obtained from the sensitivity test presented in Section 2 when
289 G is the same as that used by HL. It is surprising because the model atmosphere is expected to be more
290 sensitive to the changes in the radiative damping as G , which is independent of \bar{u} , becomes smaller and
291 smaller. Note that when our control runs adopt the black radiative damping profile shown in FIG. 1 the
292 simulated QBO periods are not sensitive to the imposed semiannual forcing provided that G does not
293 exceed the values employed by HL. Similarly, when our experimental runs adopt the red radiative
294 damping profile above 24 km shown in FIG. 1 the simulated QBO periods are also not sensitive to the
295 imposed semiannual forcing provided that G does not exceed 50% of the values adopted in HL. The
296 question naturally arises: what is responsible for this unphysical behavior?

297 In Section 2, the simulated QBO periods are equal to 30 and 34.3 months when we adopted the no-
298 slip and stress-free upper boundary condition respectively with all other parameters being identical to
299 those used by HL. The results implicate the upper boundary conditions in the inconsistency. Plumb (1977)
300 pointed out that the upper boundary in HL was undesirably low and implied that raising the lid to an
301 additional 50% would be adequate for the robustness in his model. Here, we carry out a series of
302 sensitivity tests by raising the model lid gradually from 35 km to 55 km with the one-kilometer increment.
303 we will demonstrate how the behavior of the HL model with $G = 0$ converges with that of the Plumb
304 model. The modified HL model, i.e., the HL model with $G = 0$ is identical to the Plumb model except
305 that the former has the no-slip upper boundary condition while the latter has the stress-free upper

306 boundary condition. Both models share the same governing equation (5). Note that we set the radiative
307 damping rate above the 35 km level to its value at the 35 km level shown in FIG. 1.

308 For the radiative damping profile corresponding to the reference CO₂, FIG. 6 shows that when the
309 model lid is placed at the 35 km level the simulated QBO period of 30.0 months with the no-slip upper
310 boundary condition (solid black line) is apparently shorter than that of 37.5 months with the stress-free
311 upper boundary condition (dashed black line). FIG. 6 also shows that as the model lid is raised
312 incrementally from the 35 km level to the 46 km level, the discrepancies between the simulated QBO
313 periods due to the different upper boundary conditions decrease monotonically. No matter whether we
314 adopt the no-slip or stress-free upper boundary condition, the simulated QBO period is 32.4 months for
315 the reference radiative damping profile provided that the model top is at or above the 46 km level.

316 Similarly, for the radiative damping profile corresponding to the doubled CO₂, FIG. 6 demonstrates
317 that when the model lid is placed at the 35 km level the simulated QBO period of 28.6 months with the
318 no-slip upper boundary condition (solid red line) is obviously shorter than that of 31.6 months with the
319 stress-free upper boundary condition (dashed red line). FIG. 6 also exhibits that as the model lid is raised
320 gradually from the 35 km level to the 40 km level, the discrepancies between the simulated QBO periods
321 due to the different upper boundary conditions decrease monotonically. No matter whether we adopt the
322 no-slip or stress-free upper boundary condition, the simulated QBO period for the enhanced infrared
323 cooling due to the doubled CO₂ is 30.0 months provided that the model top is at or above the 40 km level.
324 It is apparent that the required model top is lower when the radiative damping is augmented due to the
325 doubling of CO₂ because the planetary waves dissipate more steeply with height in presence of the
326 enhanced infrared cooling rates.

327 FIG. 6 suggests that when the model lid is sufficiently high the QBO period in response to the
328 enhanced radiative damping due to the increasing CO₂ will decrease from 32.4 to 30.0 months. This

329 7.4% decrease in the QBO period is independent of the upper boundary condition. Plass (1956) indicated
330 that the probable error of the cooling rate was about 10% below 20 km, increasing to 30% at 50 km and
331 that the relative differences between the cooling rates should be considerably more accurate than their
332 magnitude. In other words, the relative differences between the various cooling rates calculated by Plass
333 (1956) should be considerably smaller 30%. Using the HL model with its top at the 48 km level, we
334 further conducted two experiments by adopting $G = 0$ in Eq. (1) and increasing the radiative damping
335 corresponding to the doubled CO₂ between 30 km and 48 km by $30\% - 30\% * 30\% = 21\%$ and
336 $30\% + 30\% * 30\% = 39\%$ respectively relative to that in the control run. The simulated QBO periods
337 are 30.8 and 29.3 months respectively. Therefore, when the model lid is sufficiently high the QBO period
338 in response to the enhanced radiative damping due to the doubled CO₂ will decrease by approximately
339 $7.4\% \pm 2.5\%$.

340 Note that N , the Brunt-Väisälä frequency, in Eqs. (2) and (3) also changes with increasing CO₂.
341 Richter et al. (2020b) showed that N^2 would be decreased by $\sim 5\%$ in the stratosphere when CO₂ is
342 doubled (refer to their Figure 2c). We used the HL model to conduct a sensitivity test by adopting $G = 0$
343 in Eq. (1) with the radiative damping profile corresponding to the doubled CO₂ and the top of the models
344 at the 48 km level. The rest of parameters in this sensitivity test are identical to those in all the previous
345 runs except that the Brunt-Väisälä frequency in this experimental run was 2.5% smaller than that in the
346 control run. The models were run for 1000 years to further increase the spectral resolution. We found
347 that when the Brunt-Väisälä frequency was decreased by 2.5%, the simulated QBO period was slightly
348 lengthened from 30 months to 30.2 months (figure not shown). In other words, the impact of decreasing
349 stratospheric buoyancy frequency on the QBO period is marginal.

350 Analyzing eleven CCM1-1 REF-C2 climate–chemistry simulations, Eichinger and Šácha (2020)
351 showed that the scale height in the stratosphere decreases by 2.3% per century. Accordingly, we used

352 the HL model to conduct another sensitivity test by adopting $G = 0$ in Eq. (1) with the radiative damping
353 profile corresponding to the doubled CO₂ and the top of the models at the 48 km level. The rest of
354 parameters in this sensitivity test are identical to those in all the previous control runs except that the
355 scale height in this experimental run was 2.3% smaller than that in the control run. The model was also
356 run for 1000 years for the sake of higher spectral resolution. We found that when the scale height was
357 decreased by 2.3%, the simulated QBO period was also shortened by about 2.3%, i.e., from 30 months
358 to 29.3 months (figure not shown). Apparently, the shortening of the QBO period due to the warming
359 climate is ascribed less to the shrinkage of the scale height in the stratosphere than to the enhancing of
360 the stratospheric radiative damping. Together, the shrinking scale height and the increasing radiative
361 damping shorten the QBO period by about 9.6%.

362

363 **5. Conclusions**

364 Plumb (1977) envisioned that stratospheric climate change would give rise to long-term changes in
365 the QBO period due to changes in radiative damping and the Brunt-Väisälä frequency. Using one-
366 dimensional (1D) models we found that the enhanced radiative damping arising from the doubling of
367 CO₂ leads to the shortening of the QBO period by about $7.4\% \pm 2.5\%$ provided that the model top is
368 higher than the 46 km level. Furthermore, when we incorporated both the 2.3% shrinkage of the scale
369 height and the enhanced radiative damping, the QBO period is shortened by about 9.6%. In addition, the
370 impact of decreasing stratospheric buoyancy frequency on the QBO period is marginal. Note that those
371 models include neither gravity waves nor tropical upwelling and assume that there are no changes in
372 wave fluxes entering the equatorial stratosphere.

373 From a comprehensive model perspective, Richter et al. (2020b) showed that the changes in period
374 of the QBO in warming climate simulations varied quite significantly among these models. Some models

375 projected longer mean periods and some shorter mean periods for the QBO in a future warmer climate.
376 They argue that uncertainty in the representation of the parameterized gravity waves is the most likely
377 cause of the spread among the QBOi models in the QBO's response to climate change.

378 In addition, CO₂ increases in the NASA Goddard Institute for Space Studies Model E2.2-AP (Rind
379 et al. 2020) lead to a decrease of both QBO period and QBO amplitude (DallaSanta et al., in prep.). The
380 period decrease is associated with increases in lower stratospheric momentum fluxes (related to
381 parameterized convection), a finding consistent with Geller et al. (2016a, 2016b) and Richter et al.
382 (2020b). The amplitude decrease is associated with a strengthened residual mean circulation, also
383 consistent with the literature, although the vertical structure of the circulation response is nontrivial. It
384 is worth mentioning that horizontal momentum flux divergences could also play an important role in
385 weakening the QBO (Match and Fueglistaler, 2019, 2020).

386 Our 1D models only explored how the QBO period responds to the enhancing radiative damping of
387 planetary waves, the shrinking scale height in the stratosphere, and the decreasing stratospheric
388 buoyancy frequency due to the increasing CO₂ concentration. In order to investigate how those factors
389 affect gravity waves which play an even more important role in determining the QBO period than
390 planetary waves, high-resolution models such as those used by Kawatani et al. (2011, 2019) are desirable
391 to further our understanding. Ultimately, how the QBO period changes in response to the increasing CO₂
392 will be determined by the combined effects of the strengthening of tropical upwelling, the increasing of
393 wave fluxes entering the equatorial stratosphere, the enhancing of radiative damping, and the shrinking
394 of the scale height in the stratosphere, which warrants further research.

395

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401

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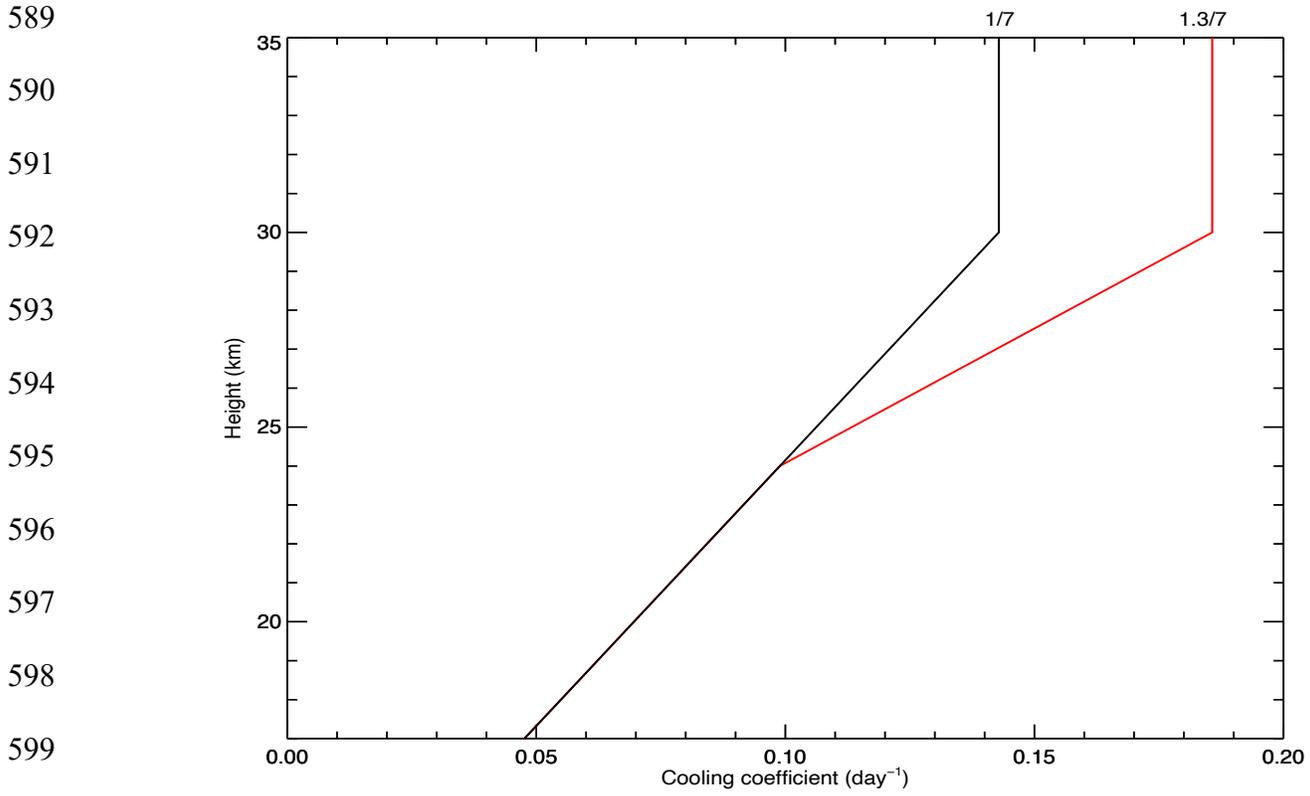
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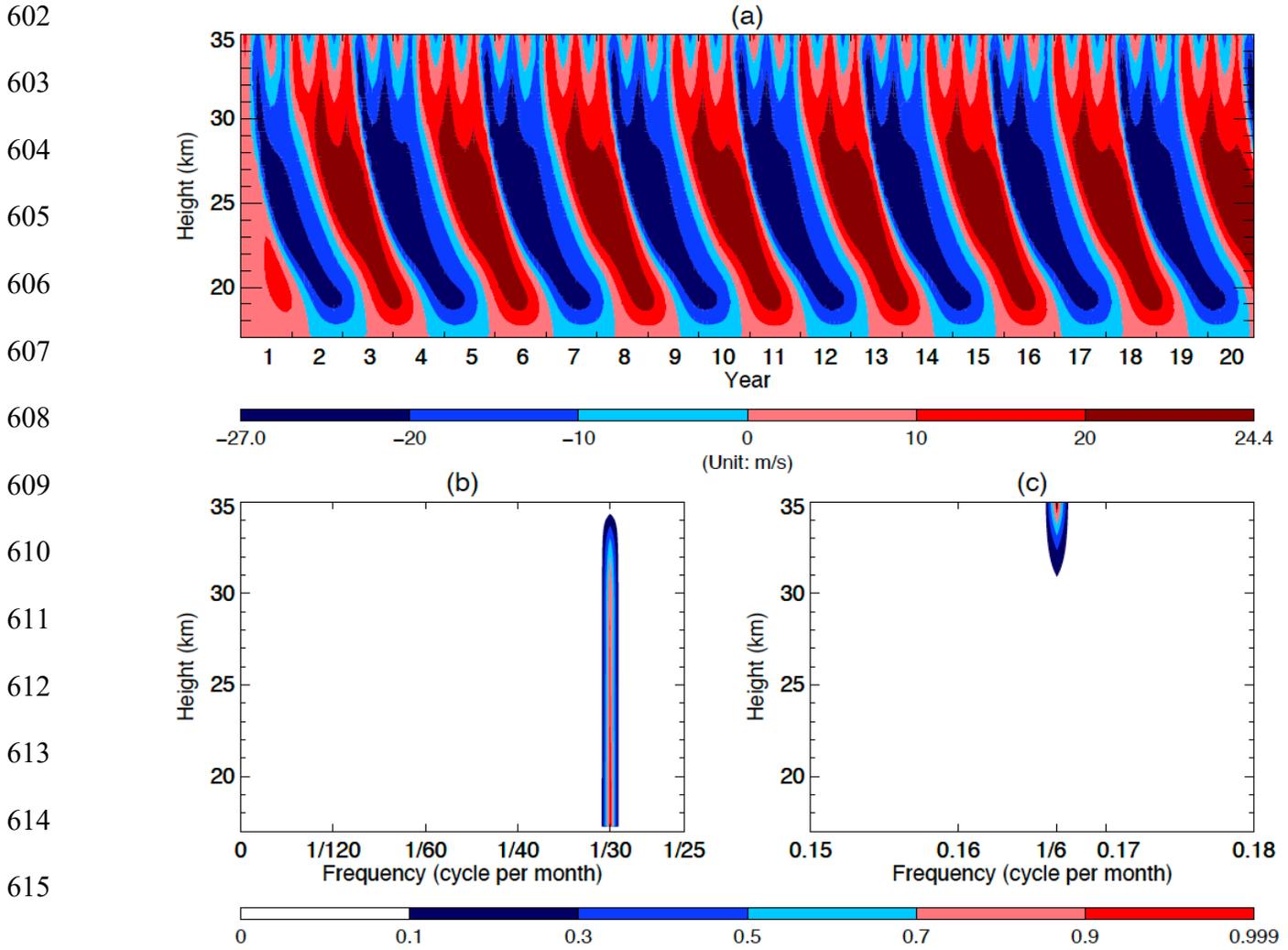
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600 **FIG. 1:** Newtonian cooling profiles: The smaller values (black line) are used for the control runs while
 601 the larger values (red line) are used for the experimental runs.



616 **FIG. 2:** (a) Time–height section of the monthly averaged mean zonal wind over the first 20 years from
 617 the HL’s original model. (b) and (c) Frequency–height section of the power spectral densities (PSD) of
 618 the standardized monthly averaged mean zonal wind of the 100 years. Note that in order to better
 619 visualize the PSD in (b) and (c), we trimmed off the blank segments for the frequencies ranging from
 620 $\frac{1}{25}$ to 0.15 cycle per month and those ranging from 0.18 to 0.5 cycle per month.

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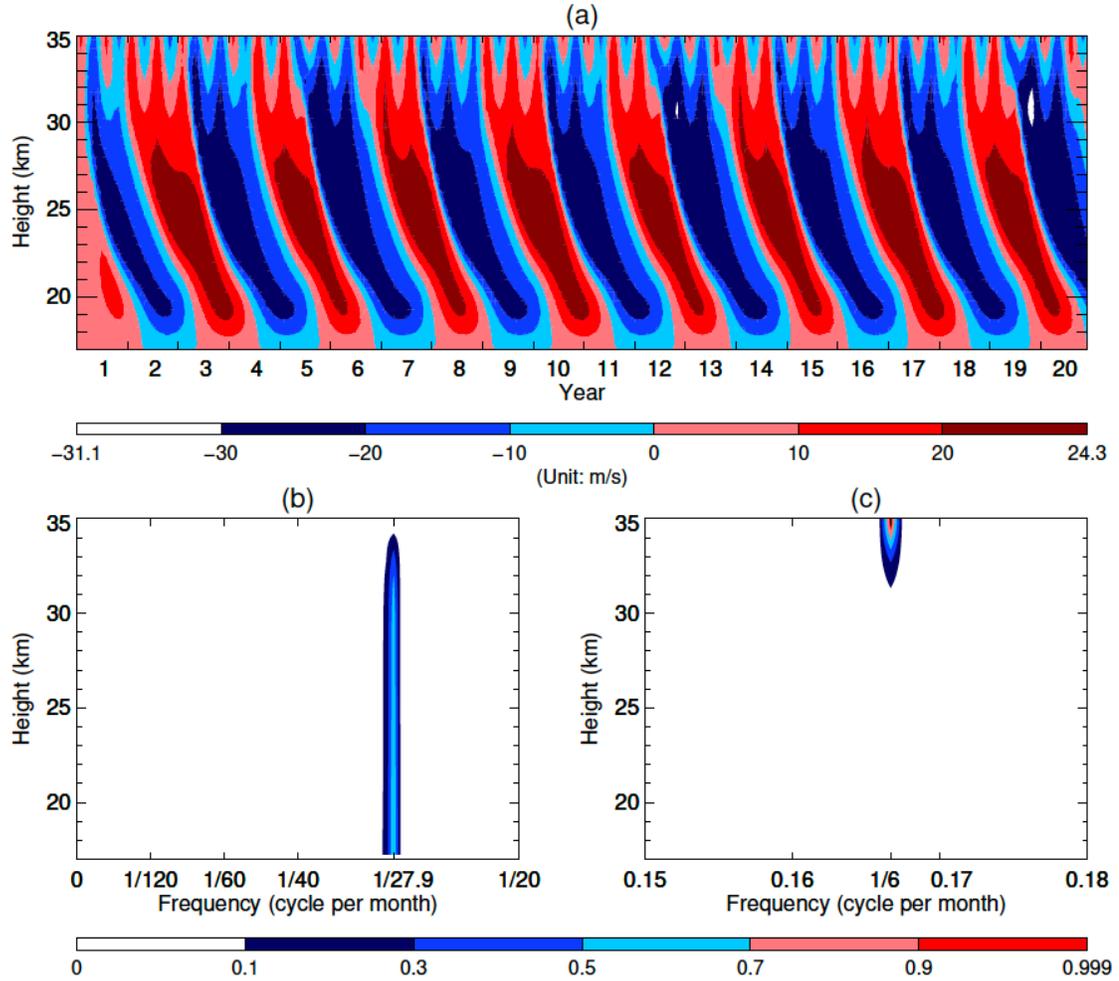
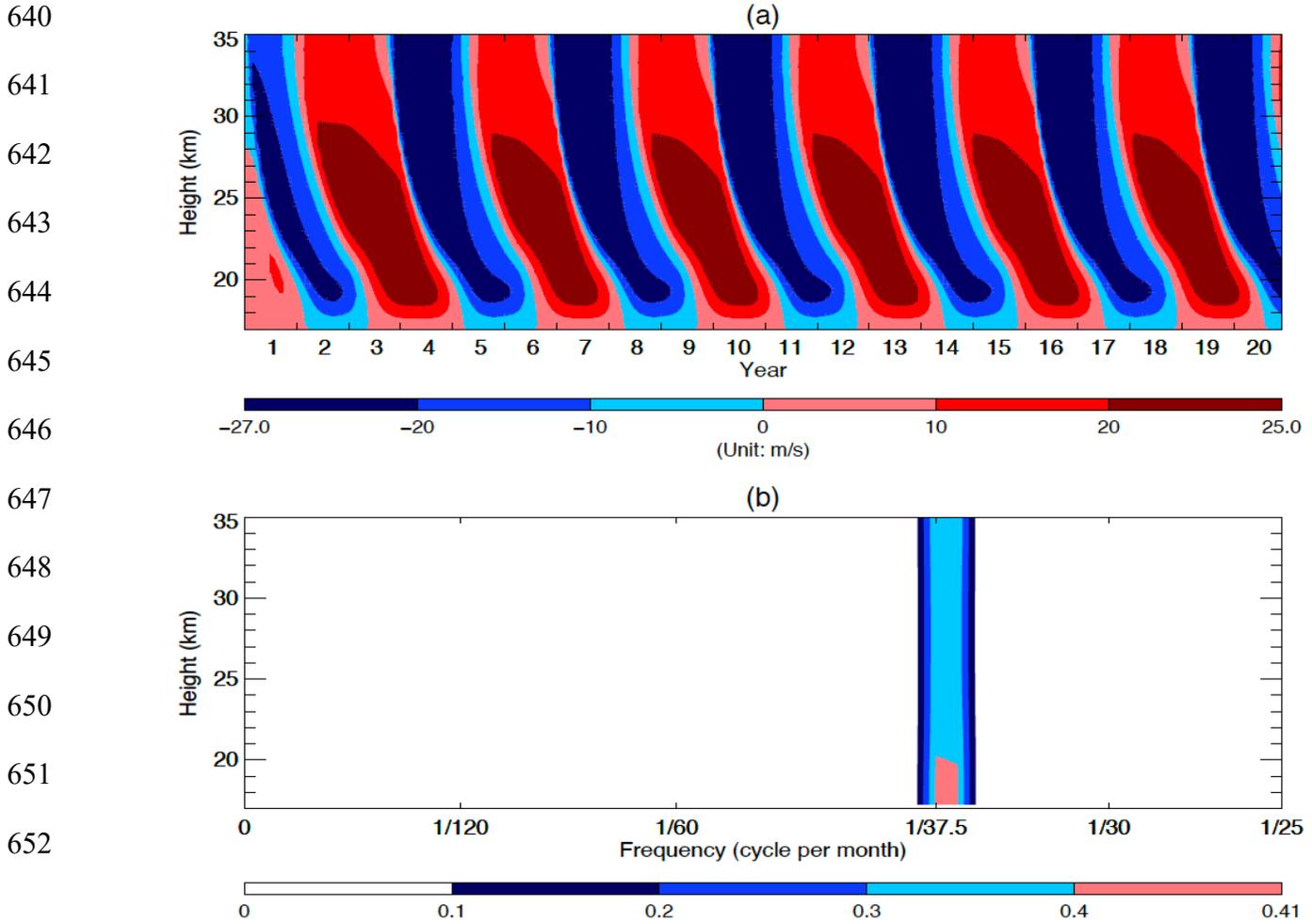


FIG. 3: (a) Same as FIG. 2a, but with the enhanced $\alpha(z)$ depicted as the red line in FIG. 1. (b) and (c) Frequency–height section of the power spectral densities (PSD) of the standardized monthly averaged mean zonal wind of the 100 years. Note that in order to better visualize the PSD in (b) and (c), we trimmed off the blank segments for the frequencies ranging from $\frac{1}{20}$ to 0.15 cycle per month and those ranging from 0.18 to 0.5 cycle per month.



653 **FIG. 4:** (a) Time–height section of the monthly averaged mean zonal wind over the first 20 years from
 654 the HL’s model without the semiannual forcing. (b) Frequency–height section of the power spectral
 655 densities (PSD) of the standardized monthly averaged mean zonal wind of the 100 years. Note that in
 656 order to visualize the PSD, we trimmed off the blank segment for the frequencies ranging from $\frac{1}{25}$
 657 to 0.5 cycle per month.

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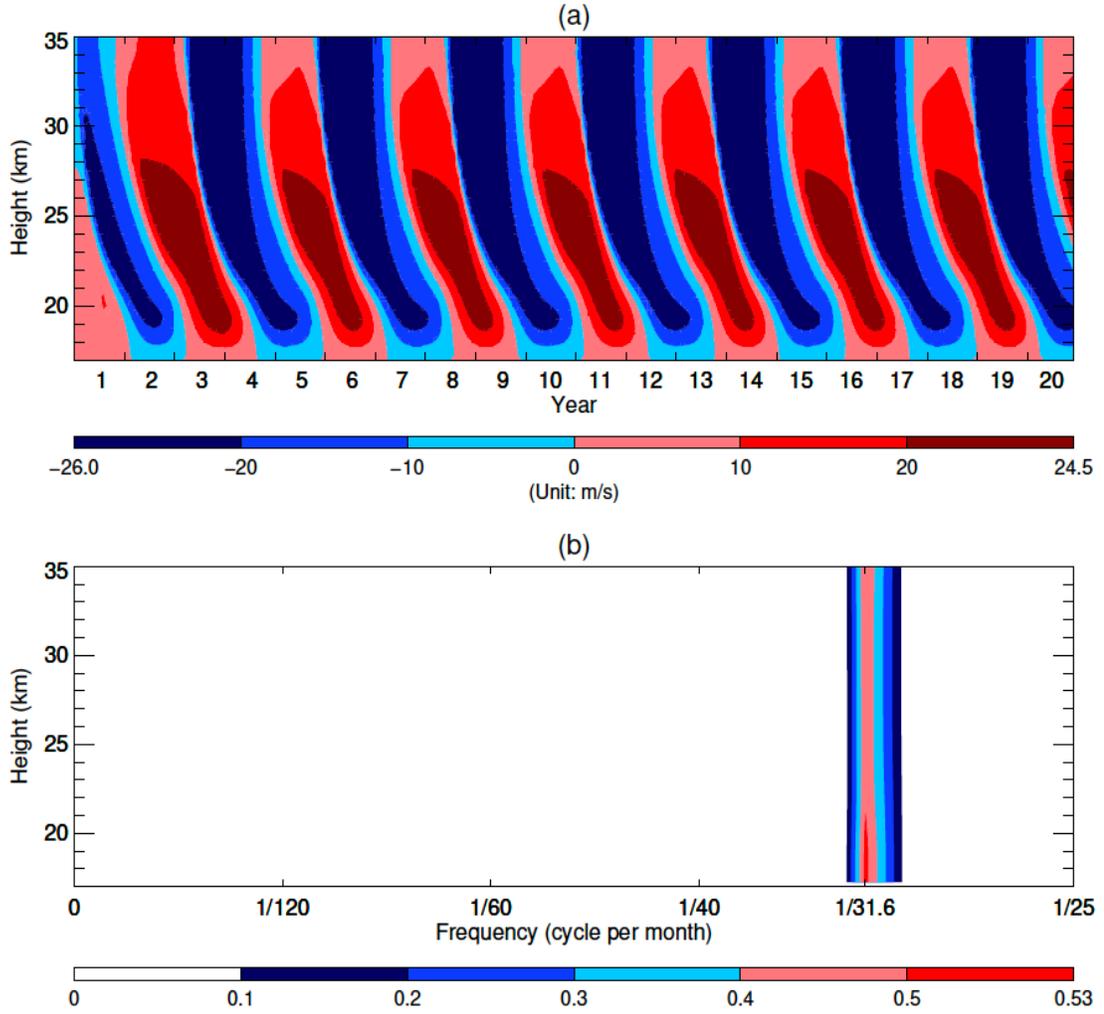
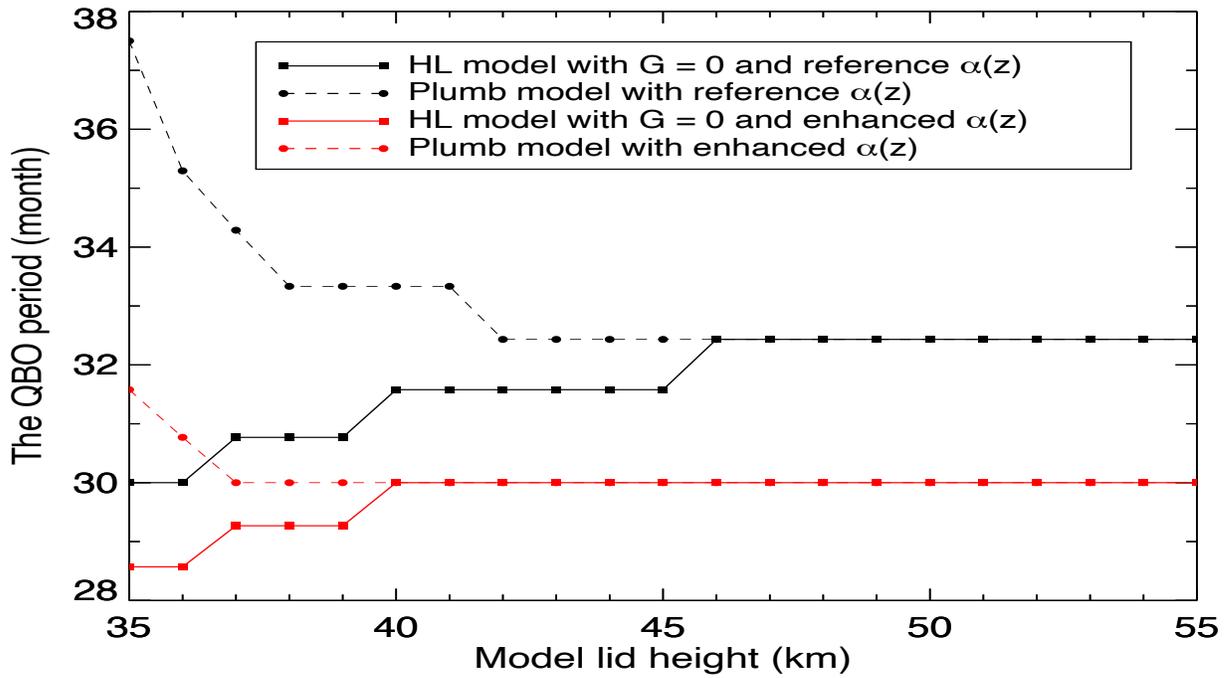


FIG. 5: (a) Same as FIG. 4a, but with the enhanced $\alpha(z)$ depicted as the red line in FIG. 1. (b) Same as FIG. 4b, but for the doubled CO₂ Run.

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684 **FIG. 6:** The relationship between the simulated QBO period with the height of the model lid. Black and
685 red lines depict the results from using the reference radiative damping and the enhanced radiative
686 damping respectively while solid and dashed lines delineate the results from the HL model with $G = 0$
687 and the Plumb model respectively.