

1 **The Impact of Increasing Stratospheric Radiative Damping on the QBO Period**

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8

9 **Abstract.** Stratospheric radiative damping increases as atmospheric carbon dioxide concentration rises.

10 We use the one-dimensional mechanistic models of the QBO to conduct sensitivity experiments and
11 find that the simulated QBO period shortens due to the enhancing of radiative damping in the
12 stratosphere. This result suggests that increasing stratospheric radiative damping due to rising CO₂ may
13 play a role in determining the QBO period in a warming climate along with wave momentum flux
14 entering the stratosphere and tropical vertical residual velocity, both of which also respond to
15 increasing CO₂.

16

17 **1. Introduction**

18 The quasi-biennial oscillation (QBO) dominates the variability of the equatorial middle and lower
19 stratosphere and is characterized by a downward propagating zonal wind regime that regularly changes
20 from westerlies to easterlies. The QBO period ranges from 22 to 34 months with its average being
21 slightly longer than 28 months. The QBO not only manifests itself in the equatorial zonal winds, but also
22 leaves an imprint on the temperature in both the tropics and extratropics (Baldwin et al., 2001 and
23 references therein).

24 The QBO has far-reaching implications for global weather and climate systems. First of all, the QBO
25 exerts a marked influence on the distribution and transport of various chemical constituents such as
26 ozone (O₃) (e.g., Hasebe, 1994), water vapor (H₂O) (e.g., Kawatani et al., 2014), methane (CH₄), nitrous
27 oxide (N₂O), hydrogen fluoride (HF), hydrochloric acid (HCl), odd nitrogen species (NO_y) (e.g.,
28 Zawodny and McCormick, 1991), and volcanic aerosol (Trepte and Hitchman, 1992). Secondly, it is
29 well appreciated that the QBO influences the extratropical circulation in the winter stratosphere, which
30 is commonly known as the Holton–Tan effect (Holton and Tan, 1980; Labitzke, 1982). It has been noted
31 that the effect of the QBO on the extratropical winter stratosphere impacts the severity of stratospheric
32 ozone depletion (e.g., Lait et al., 1989). Furthermore, taking account of the QBO improves the simulation
33 and predictability of the extratropical troposphere (e.g., Marshall and Scaife, 2009). Finally, through its
34 modulation of temperature and vertical wind shear in the vicinity of the tropical tropopause, the QBO
35 influences tropical moist convection (Collimore et al., 2003; Liess and Geller, 2012), the El Niño–
36 Southern Oscillation (ENSO) (Gray et al., 1992; Huang et al., 2012; Hansen et al. 2016), the Hadley
37 circulation (Hitchman and Huesmann, 2009), the tropospheric subtropical jet (Garfinkel and Hartmann,
38 2011a, 2011b), the boreal summer monsoon (Giorgetta et al., 1999), and the Madden-Julian Oscillation
39 (Yoo and Son, 2016). Intriguingly, the QBO is also reported to influence the activities of tropical
40 cyclones (Gray et al., 1984; Ho et al., 2009), albeit this issue is still unsettled (Camargo and Sobel, 2010)
41 and needs further study.

42 Efforts to understand and simulate the QBO have been ongoing ever since its discovery by Ebdon
43 (1960) and Reed et al. (1961). Lindzen and Holton (1968) and Holton and Lindzen (1972) developed
44 the classical theory of the QBO. Namely, as waves propagate upward, they are attenuated by thermal
45 damping, encounter critical levels, and accelerate and decelerate the mean flow, providing momentum
46 sources for both the westerly and easterly phases of the QBO.

47 Holton and Lindzen's (1972) model (hereafter referred to as HL model) was further simplified by
48 Plumb (1977), the elegance of which made it a standard paradigm for the QBO. In Plumb's (1977)
49 Boussinesq formulation, the QBO period is inversely dependent upon both the momentum flux and
50 thermal dissipation rate. Hamilton (1981) further highlighted the role of the radiative damping rate on
51 both the realistic vertical structure and the realistic period of the QBO.

91 By adopting higher vertical resolutions and incorporating various gravity wave parameterization
92 schemes, many state-of-the-art climate models have shown the capability to self-consistently simulate
93 the QBO (Scaife et al., 2000; Giorgetta et al., 2002, 2006; Rind et al., 2014, 2020; Geller et al., 2016a;
94 Richter et al., 2020a, 2020b). Given the important implications of the QBO for the global climate system,
95 it is natural to ask how the QBO will change in a warming climate.

96 Giorgetta and Doege (2005) showed a shortening of the QBO period in their doubled CO₂
97 experiments. They reasoned that both the weakening of the tropical upwelling and the prescribed
98 increase of gravity wave sources lead to the reduction of the QBO period in a warming climate. However,
99 most climate models project a strengthening rather than weakening of tropical upwelling in a warmer
100 climate (Butchart et al., 2006; Butchart 2014; Li et al., 2008). Employing a model without any
101 parametrized non-orographic gravity waves, Kawatani et al. (2011) demonstrated that the intensifying
102 tropical upwelling in a warming climate dominates the counteracting effect of enhanced wave fluxes and
103 consequently projected a lengthening of the QBO period. Using fixed sources of parametrized gravity
104 waves, Watanabe and Kawatani (2012) also projected an elongation of the QBO period in a warming
105 climate and ascribed it to the stronger tropical upwelling. Analyzing four Coupled Model
106 Intercomparison Project phase 5 (CMIP5) models that could simulate a reasonable QBO, Kawatani and
107 Hamilton (2013) found that the projected trends of the QBO period were inconsistent in sign. They
108 further investigated the 60-year operational balloon-borne radiosonde observations provided by the Free

109 Berlin University and detected no significant trend in the QBO period. Richter et al. (2020b) investigated
110 the response of the QBO to doubled and quadrupled CO₂ climates among eleven models that participated
111 in Phase 1 of the Stratospheric-tropospheric Processes And their Role in Climate QBO-initiative (QBOi;
112 Butchart et al., 2018), and found no consensus on how the QBO period would respond to a changing
113 climate. Recently, Butchart et al. (2020) evaluated ten Coupled Model Intercomparison Project phase 6
114 (CMIP6) models with realistic QBO in two Shared Socioeconomic Pathways (SSPs, Gidden et al., 2019)
115 scenario simulations and surprisingly found that the QBO period shortens in seven of those ten models
116 in both in both SSP3-7.0 and SSP5-8.5 scenarios although only two and three models show a significant
117 shortening trend in the respective scenarios.

118 It is challenging to ascertain the trend of the QBO period in a warming climate. On one hand, a
119 speeding-up of the Brewer-Dobson circulation in a warming climate leads to a lengthening of the QBO
120 period in most climate models. On the other hand, there is a robust increase in the vertical component of
121 the EP flux for both eastward and westward propagating waves (Richter et al., 2020b; Butchart et al.,
122 2020), indicating that the QBO period shortens due to the enhanced wave driving in a warming climate.
123 The competing effects between enhanced wave driving and a faster Brewer-Dobson circulation suggests
124 that trends in the QBO period are likely to be small and difficult to detect due to the large cycle-to-cycle
125 variability that is reproduced by climate models (Butchart et al., 2020). In addition, uncertainty in the
126 representation of the parameterized gravity waves make it more elusive to detect the trend of the QBO
127 period in a warming climate (Schirber et al., 2015; Richter et al., 2020b).

128 Given the fact that the QBO period is influenced by the radiative damping (Plumb 1977; Hamilton
129 1981), a natural question to ask is whether it could play a role on the trend of the QBO in a warming
130 climate. ~~Fels (1985) estimated that the radiative damping time under a doubling of CO₂ would decrease~~

Deleted: Plumb (1956) showed that when the CO₂ concentration is increased from 330 ppmv to 660 ppmv, the cooling rate increases significantly in the middle and upper stratosphere while it is not changed below the 24 km height level. The cooling rate is increased by about 50% around the 40 km height level (see his Figure 8).

137 by about 23%. His estimate implies a shortening of the QBO period as the radiative damping rate
138 increases.

139 It is well-known that enhanced wave fluxes entering the stratosphere and stronger tropical upwelling
140 individually play a dominant role in determining the trends in the QBO period in a warming climate.
141 Does the competing effect between them leave some room for increasing stratospheric radiative damping
142 to exert an influence on the QBO period? In this paper, we use the HL model to isolate the effect of
143 radiative damping on the QBO period by assuming that the momentum flux entering the stratosphere
144 doesn't change in our experiments. Observational and modeling studies (Andrews et al., 1987; Kawatani
145 et al., 2009, 2010, 2011; Richter et al., 2020b; Holt et al., 2020) showed that the wave forcing spectrum
146 is similar to a discrete two-wave spectrum rather than red-noise or white-noise, all of which are
147 illustrated in Saravanan (1990). Accordingly, the QBO is indeed sensitive to stratospheric radiative
148 damping, and the HL model is suitable for us to conduct the sensitivity analysis.

149 The remainder of this paper is organized as follows. Section 2 investigates the sensitivity of the QBO
150 period to the radiative damping using HL's original model. Section 3 explores the sensitivity of the QBO
151 period to the radiative damping using a modified HL model where the semiannual forcing is removed.
152 Discussion and conclusions are presented in Sections 4 and 5 respectively.

153
154 **2. Sensitivity of the QBO period to enhanced stratospheric radiative damping in the original HL**
155 **model**

156 In the HL model the governing equation of mean flow emerges after the primitive momentum
157 equation is meridionally averaged over some suitable latitudinal belt over the equator.

158
$$\frac{\partial \bar{u}}{\partial t} = -\frac{1}{\rho_0} \frac{\partial}{\partial z} \left[\sum_{i=1}^n \bar{F}_i \right] + K_z \frac{\partial^2 \bar{u}}{\partial z^2} + G \quad (1)$$

159 where \bar{u} is mean zonal wind, ρ_0 is mean density, \bar{F}_i is the meridionally averaged vertical Eliassen-Palm
 160 flux associated with wave i , the index i refers to the individual waves, K_z is a vertical eddy diffusion
 161 coefficient, t is time, z is altitude, and G is semiannual forcing identical to that specified by HL.

162 The \bar{F}_i are evaluated with Lindzen's (1971) WKB formalism for equatorial waves in shear. When
 163 only infrared cooling acts to damp the waves the formulae for \bar{F}_i are

$$164 \quad \bar{F}_0(z) = A_0 \exp\left(-\int_{17\text{km}}^z \frac{\alpha N}{k(c-\bar{u})^2} dz\right) \quad (2)$$

165 for the Kelvin wave, and

$$166 \quad \bar{F}_1(z) = A_1 \exp\left[-\int_{17\text{km}}^z \frac{\alpha\beta N}{k^3(c-\bar{u})^3} \left(1 - \frac{k^2(\bar{u}-c)}{\beta}\right) dz\right] \quad (3)$$

167 for the mixed Rossby-gravity wave. As in HL, the wavenumber k , the phase speed c , and A_0 are chosen
 168 to be $2\pi/(40,000 \text{ km})$, 30 m s^{-1} , and $0.04 \text{ m}^2 \text{ s}^{-2} \rho_0(17 \text{ km})$, respectively for the Kelvin wave while
 169 they are equal to $-2\pi/(10,000 \text{ km})$, -30 m s^{-1} , and $-0.04 \text{ m}^2 \text{ s}^{-2} \rho_0(17 \text{ km})$, respectively for the
 170 mixed Rossby-gravity wave. In Eq. (1), $K_z = 0.3 \text{ m}^2 \text{ s}^{-1}$, which is also the same as in HL. In addition,
 171 $\beta = 2\Omega/a$, where Ω is earth's rotation rate, and a is earth's radius. HL's boundary conditions stipulated
 172 that $\bar{u} = 0$ at the lowest model level (17 km) and constrained \bar{u} to vary semiannually at the top level (35
 173 km).

174 In our control run that is used to depict the present-day QBO all the model parameters are identical
 175 to those used by HL in their original simulation. The Brunt-Väisälä frequency

$$176 \quad N = \sqrt{\frac{g}{T_0} \left(\frac{dT_0}{dz} + \frac{g}{c_p} \right)} \quad (4)$$

177 In Eq. (4), g is gravity, T_0 is mean temperature, and c_p is specific heat of dry air at constant pressure.
 178 HL set N in Eq. (4) to $2.16 \times 10^{-2} \text{ s}^{-1}$ with a scale height $H = 6 \text{ km}$. In addition, the Newtonian

179 cooling profile in our control run, i.e., $\alpha(z)$ in Eqs. (2) and (3), is also identical to that in the original
180 HL model and depicted in FIG. 1a as the black line. Namely, $\alpha(z)$ in the control run increases from
181 $(21 \text{ day})^{-1}$ at 17 km to $(7 \text{ day})^{-1}$ at 30 km and is kept at $(7 \text{ day})^{-1}$ between 30 km and 35 km. Fels
182 (1985) explained why the magnitude of this radiative damping rate is suitable for simulating the QBO
183 on the basis of the scale-dependent effect of radiative damping (Fels, 1982). Hamilton (1981)
184 demonstrated that the proper choice of $\alpha(z)$ is crucial in simulating a realistic vertical structure of the
185 QBO.

186 Eq. (1) was integrated for 100 years using the forward-backward scheme (Matsuno, 1966). The
187 vertical resolution was 250 m and identical to that in HL. The time step was 12 hr, i.e., one half of used
188 in HL, because the 24-hr time step resulted in numerical instability in our integration.

189 FIG. 2a shows the time–height section of the monthly averaged mean zonal wind simulated over the
190 first 20 years using the HL model. Both the QBO and the semiannual oscillation (SAO) are conspicuous.
191 The fast Fourier transform (FFT) method is used to calculate the frequency power spectra. In order to
192 more accurately derive the QBO period, the model was run for 100 years to increase the spectral
193 resolution. Frequency–height sections of the power spectral densities (PSD) over zero to the Nyquist
194 frequency, i.e., 0.5 cycle/month, depict two sharp lines (peaks) at $\frac{1}{30}$ and $\frac{1}{6}$ cycle/month, respectively
195 (not shown). In order to better visualize the magnitudes of the PSD, we show two truncated frequency–
196 height sections with FIG. 2b and FIG. 2c highlighting the QBO and the SAO respectively. FIG. 2b shows
197 that the QBO dominates over the model domain. The peak frequency corresponds to the period of 30
198 months. FIG. 2c shows the SAO dominates near the model top due to the fact a semiannual forcing was
199 imposed in the altitudes from 28 to 35 km.

200 It is worth mentioning that the QBO period shown here is longer than 26.5 months reported in the HL
201 paper (see their FIG. 1). Using the HL model parameters, the QBO period simulated by Plumb (1977)

202 was close to three years (refer to his FIG. 8a), which is longer than our simulated QBO period, i.e., 30.0
 203 months. Although we could not explain why our simulated QBO period is longer than that simulated by
 204 HL, we found that when the upper boundary condition is changed from $\bar{u} = 14 \sin(\omega_a t)$ and $\omega_a =$
 205 $\frac{2\pi}{180} \text{ day}^{-1}$ used in the HL's original model (refer to their Eqs. (2)) to $\frac{\partial \bar{u}}{\partial z} = 0$ used in Plumb (1977), the
 206 simulated QBO period becomes 34.3 month (figure not shown). In other words, when we adopted the
 207 stress-free upper boundary condition as in Plumb (1977), our simulated QBO period is comparable to
 208 that simulated by him, which lends credence to our reconstruction of the HL model.

209 In order to rigorously quantify the relationship of the Newtonian cooling coefficient at any altitude z
 210 between the reference and doubled CO₂, we follow Dickinson (1973) in using a radiative transfer model
 211 to calculate $Q_1(T)$ for a reference temperature profile $T(z)$ and $Q_1(T + \delta)$ for $T(z) + \delta$, where a small
 212 perturbation $\delta T = 0.1 \text{ K}$ with $T(z)$ being the 1976 U.S. standard atmosphere. Our radiative transfer
 213 computations use the MODTRAN gas absorption database with 0.1 cm^{-1} spectral resolution (Jin et al.
 214 2019; Berk et al. 2008). We then repeat the computations with the doubled CO₂ to yield $Q_2(T)$ and
 215 $Q_2(T + \delta)$. It follows that $\frac{\alpha_2}{\alpha_1} = \frac{Q_2(T+\delta) - Q_2(T)}{Q_1(T+\delta) - Q_1(T)}$, where $\alpha_1(z)$ and $\alpha_2(z)$ stand for the Newtonian cooling
 216 coefficient at any altitude z for the reference and doubled CO₂, respectively. In FIG. 1b the black line
 217 depicts the ratio for the broadband longwave radiation ($5 \mu\text{m} - 100 \mu\text{m}$) and the red line delineates the
 218 ratio for the CO₂ absorption band ($12 \mu\text{m} - 18 \mu\text{m}$). The ratio calculated over the broadband is
 219 conspicuously smaller than that for the CO₂ absorption band, because the changes in cooling rate from
 220 the temperature perturbation are larger over a wider spectral band. It is worth mentioning that the ratios
 221 calculated over the broadband in the middle stratosphere are close to 1.3 and comparable to what Fels
 222 (1985) estimated, i.e., about 23% decrease of the radiative damping time under a doubling of CO₂.

223 Returning to the 1D HL model, we synthesize those findings by prescribing $\alpha_2(z)$ in our
 224 experimental runs for the doubled CO₂ as follows: an increase of 30% between 30 and 35 km, no change

Deleted: As mentioned in Section 1, when the atmospheric carbon dioxide concentration is doubled the cooling rate increases significantly in the middle and upper stratosphere while it varies little below the 24 km height level. Accordingly, $\alpha(z)$ in the experimental run is kept the same as in the control run, i.e., increases from $(21 \text{ day})^{-1}$ at 17 km to $\frac{9}{91} \text{ day}^{-1}$ at 24 km. Since the cooling rate is increased by about 50% around the 40 km height level (Plass, 1956), the radiative damping rates are expected to also increase in the middle and upper stratosphere as the CO₂ concentration rises. However, the relative change of cooling rate Q in response to the increasing CO₂ is not identical to that of Newtonian cooling coefficient due to the facts that $Q = -\alpha(T - T_e)$ and the radiative equilibrium temperature T_e in the stratosphere decreases as the CO₂ concentrations increase (Manabe et al. 1967). In other words, for any given temperature profile as adopted by Plass (1956), the decreasing of the stratospheric T_e with increasing CO₂ concentration leads to $\alpha_2(z): \alpha_1(z) \neq Q_2(z): Q_1(z)$, where $\alpha_1(z)$ and $\alpha_2(z)$ represent the Newtonian cooling coefficient at any altitude z for the reference and doubled CO₂, respectively while $Q_1(z)$ and $Q_2(z)$ stand for the cooling rate likewise.

Deleted: $\alpha_2(z): \alpha_1(z)$,

Deleted: used by Plass (1956)

Deleted: For the CO₂ absorption band, the calculated ratio is evidently comparable to the ratio of cooling rates between the doubled CO₂ and the reference CO₂ shown in figure 8 of Plass (1956), with an additional small increase (<1.1) below the 24 km level.

255 below 24 km, and linear interpolation between 24 and 30 km. The resulting increase of radiative damping
256 rate from the control runs is depicted as the red line in FIG. 1a. This increase is reasonable based on our
257 results shown in FIG. 1b.

258 FIG. 3a shows the time–height section of the monthly averaged mean zonal wind simulated over the
259 first 20 years for the doubled CO₂ run, where the increased $\alpha(z)$ depicted as the red line in FIG. 1a was
260 employed while all other parameters are identical to those in the control run. Obviously, the QBO
261 dominates below 28 km while the semiannual oscillation (SAO) dominates above 31 km. Like FIG. 2b
262 and FIG. 2c, we only show two truncated frequency–height sections with FIG. 3b highlighting the QBO
263 and FIG. 3c highlighting the SAO. FIG. 3b also shows that the QBO prevails over the model domain. The
264 peak frequency corresponds to the period of 27.9 months. FIG. 3c shows the SAO dominates near the
265 model top due to the same imposed semiannual forcing as that in the control run.

266 In summary, using the original HL model we found that the increased radiative damping due to the
267 doubling of CO₂ shortens the QBO period by 7% (i.e., decreases from 30 months to 27.9 months).

268
269 **3. Sensitivity of the QBO period to enhanced stratospheric radiative damping in the modified HL**
270 **model without the semiannual forcing**

271 HL pointed out that the imposed semiannual oscillation was not essential for their QBO theory.
272 Applying $\frac{\partial \bar{u}}{\partial z} = 0$ as the upper boundary condition, Plumb (1977) showed a simulated QBO without
273 resorting to the semiannual momentum source (refer to his FIG. 8b). In the following control run, all
274 parameters are identical to those used in the previous control run in Section 2 except that G in Eq. (1) is
275 set to zero with $\frac{\partial \bar{u}}{\partial z}$ also being set to zero at $z = 35$ km. Hereafter we refer to it as the Plumb model¹. FIG.

¹ Strictly speaking, it is the HL model modified by Plumb (1977). In this paper, we don't use his eponymous model, i.e., the simplest possible model of the QBO, where Boussinesq fluids with uniform mean density were employed, because the HL model and its variant are considerably more realistic.

276 4a shows the time–height section of the monthly averaged mean zonal wind simulated over the first 20
277 years using the Plumb model. As expected, the QBO emerges without any trace of SAO since $G = 0$ in
278 Eq. (1). FIG. 4b shows that the QBO dominates over the whole model domain. The peak frequency
279 corresponds to the period of 37.5 months, which is comparable to that simulated by Plumb (1977) shown
280 in his FIG. 8b. Apparently, the QBO period from the Plumb model, i.e., 37.5 months shown in FIG. 4b,
281 is longer than that from the HL model, i.e., 30.0 months shown in FIG. 2b. This is partly because the
282 additional forcing G in Eq. (1) was removed in the Plumb model.

283 In the following experimental run, all parameters are identical to those used in the previous
284 experimental run in Section 2 except that G in Eq. (1) is set to zero with $\frac{\partial \bar{u}}{\partial z}$ also being set to zero at $z =$
285 35 km. In other words, the following experimental run using the Plumb model employed the same
286 parameters as the afore-mentioned control run using the Plumb model except that the increased $\alpha(z)$
287 shown as the red line in FIG. 1a was used in the following experimental run while $\alpha(z)$ shown as the
288 black line in FIG. 1a was used in the above control run. FIG. 5a shows the time–height section of the
289 monthly averaged mean zonal wind simulated over the first 20 years for the doubled CO₂ run. It is natural
290 that only the QBO emerges. A comparison of FIG. 4a and FIG. 5a shows that the QBO period shortens
291 when the infrared damping increases in response to the doubled CO₂. FIG. 5b shows that the QBO
292 dominates over the whole model domain. The peak frequency corresponds to the period of 31.6 months.

293 Using the Plumb model, we found that the increased radiative damping due to the doubling of CO₂
294 shortens the QBO period by 15.7% (i.e., decreases from 37.5 months to 31.6 months).

295

296 **4. Discussion**

297 Dunkerton (1997) showed that in the presence of tropical upwelling it was gravity waves rather than
298 large-scale Kelvin and mixed Rossby-gravity waves that contributed the bulk of QBO forcing.

299 Consequently, Geller et al. (2016a, 2016b) pointed out that enough gravity wave momentum flux is
300 required to model the QBO in a self-consistent manner in climate models and that the magnitude of the
301 subgrid-scale gravity wave momentum flux plays a crucial role in determining the QBO period. Since
302 there is no tropical upwelling in either the HL model or the Plumb model, it is natural that planetary-
303 scale Kelvin and mixed Rossby-gravity waves largely determine the QBO periods shown in Sections 2
304 and 3 due to the fact that the specified G is significantly weaker than that in the terrestrial stratosphere.
305 We conducted another sensitivity test where all parameters are identical to those in the HL model except
306 that G in both the control and experimental runs is twice as large as that used by HL. As the radiative
307 damping profile changes from the black line to the red line above 24 km shown in FIG. 1a, our simulated
308 QBO period decreases from 28.6 months to 27.3 months (figures not shown). This smaller percentage
309 decrease of 4.5% is not unexpected because G is not sensitive to the radiative damping at all and the
310 greater specified G reduces the fraction of the total wave forcing arising from the planetary waves.

311 We further conducted two sensitivity tests where all parameters are identical to those in the HL model
312 except that G in the first test is half as large as that used by HL and is equal to zero in the second test.
313 Surprisingly, as the radiative damping profile changes from the black line to the red line above 24 km
314 shown in FIG. 1a, our simulated QBO periods decreases from 30.0 months to 28.6 months both for G
315 being decreased by 50% and for $G = 0$ (figures not shown). This 4.7% decrease in the QBO period is
316 smaller than the 7% reduction obtained from the sensitivity test presented in Section 2 when G is the
317 same as that used by HL. It is surprising because the model atmosphere is expected to be more sensitive
318 to the changes in the radiative damping as G becomes smaller and smaller. Note that when our control
319 runs adopt the black radiative damping profile shown in FIG. 1a the simulated QBO periods are not
320 sensitive to the imposed semiannual forcing provided that G does not exceed the values employed by
321 HL. Similarly, when our experimental runs adopt the red radiative damping profile above 24 km shown

322 in FIG. 1a the simulated QBO periods are also not sensitive to the imposed semiannual forcing provided
323 that G does not exceed 50% of the values adopted in HL. The question naturally arises: what is
324 responsible for this unexpected behavior?

325 In Section 2, the simulated QBO periods are equal to 30 and 34.3 months when we adopted the no-
326 slip and stress-free upper boundary condition respectively with all other parameters being identical to
327 those used by HL. The results implicate the upper boundary conditions in the inconsistency. Plumb (1977)
328 pointed out that the upper boundary in HL was undesirably low and implied that raising the lid to an
329 additional 50% would be adequate for the robustness in his model. Here, we carry out a series of
330 sensitivity tests by raising the model lid gradually from 35 km to 55 km with the one-kilometer increment.
331 we will demonstrate how the behavior of the HL model with $G = 0$ converges with that of the Plumb
332 model. The modified HL model, i.e., the HL model with $G = 0$ is identical to the Plumb model except
333 that the former has the no-slip upper boundary condition while the latter has the stress-free upper
334 boundary condition. Both models share the same governing equation (5). Note that we set the radiative
335 damping rate above the 35 km level to its value at the 35 km level shown in FIG. 1a.

336 For the radiative damping profile corresponding to the reference CO₂, FIG. 6 shows that when the
337 model lid is placed at the 35 km level the simulated QBO period of 30.0 months with the no-slip upper
338 boundary condition (solid black line) is apparently shorter than that of 37.5 months with the stress-free
339 upper boundary condition (dashed black line). FIG. 6 also shows that as the model lid is raised
340 incrementally from the 35 km level to the 46 km level, the discrepancies between the simulated QBO
341 periods due to the different upper boundary conditions decrease monotonically. No matter whether we
342 adopt the no-slip or stress-free upper boundary condition, the simulated QBO period is 32.4 months for
343 the reference radiative damping profile provided that the model top is at or above the 46 km level.

344 Similarly, for the radiative damping profile corresponding to the doubled CO₂, FIG. 6 demonstrates
345 that when the model lid is placed at the 35 km level the simulated QBO period of 28.6 months with the
346 no-slip upper boundary condition (solid red line) is obviously shorter than that of 31.6 months with the
347 stress-free upper boundary condition (dashed red line). FIG. 6 also exhibits that as the model lid is raised
348 gradually from the 35 km level to the 40 km level, the discrepancies between the simulated QBO periods
349 due to the different upper boundary conditions decrease monotonically. No matter whether we adopt the
350 no-slip or stress-free upper boundary condition, the simulated QBO period for the enhanced infrared
351 cooling due to the doubled CO₂ is 30.0 months provided that the model top is at or above the 40 km level.
352 It is apparent that the required model top is lower when the radiative damping is augmented due to the
353 doubling of CO₂ because the planetary waves dissipate more steeply with height in presence of the
354 enhanced infrared cooling rates.

355 FIG. 6 suggests that when the model lid is sufficiently high the QBO period in response to the
356 enhanced radiative damping due to the increasing CO₂ will decrease from 32.4 to 30.0 months. This
357 7.4% decrease in the QBO period is independent of the upper boundary condition. Note that the relative
358 uncertainty in the ratio $\frac{\alpha_2}{\alpha_1}$ calculated over the broadband (refer to the black line shown in FIG. 1b) ranges
359 from 5% to 10% in the lower stratosphere and from 10 to 15% in the middle and upper stratosphere.
360 Thus, the relative uncertainty in the calculated ratio is 15% at a liberal estimate in the stratosphere. Using
361 the HL model with its top at the 48 km level, we further conducted two experiments by adopting $G = 0$
362 in Eq. (1) and increasing the radiative damping corresponding to the doubled CO₂ between 30 km and
363 48 km by $30\% - 30\% * 15\% = 25.5\%$ and $30\% + 30\% * 15\% = 34.5\%$ respectively relative to that
364 in the control run. The simulated QBO periods are 30.3 and 29.7 months respectively. Therefore, when
365 the model lid is sufficiently high the QBO period in response to the enhanced radiative damping due to
366 the doubled CO₂ will decrease by approximately $7.4\% \pm 0.9\%$.

Deleted: Plass (1956) indicated that the probable error of the cooling rate was about 10% below 20 km, increasing to 30% at 50 km and that the relative differences between the cooling rates should be considerably more accurate than their magnitude. In other words, the relative differences between the various cooling rates calculated by Plass (1956) should be considerably smaller 30%.

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376 Jonsson et al. (2004) showed that the doubled CO₂ induces a substantial cooling throughout most of
377 the middle atmosphere, which in turn increases the ozone mixing ratio by 15–20% in the upper
378 stratosphere and by 10–15% in the lower mesosphere (refer to their Figure 6). Incorporating this increase
379 into the ozone profile for the doubled CO₂, we recalculated the ratio of α_2 , the Newtonian cooling
380 coefficient for the doubled CO₂, to α_1 , the Newtonian cooling coefficient for the reference CO₂. Our
381 calculated α_2/α_1 is only slightly increased as compared with that shown FIG. 1b no matter whether the
382 CO₂ absorption band is $5\ \mu\text{m} - 100\ \mu\text{m}$ or $12\ \mu\text{m} - 18\ \mu\text{m}$ (figure not shown). It is not unexpected
383 because the infrared radiative cooling by ozone is significantly smaller than that by CO₂ (refer to Fig. 1
384 in Dickinson 1973) and, as a result, the 15–20% increases in the ozone mixing ratio will not make a

385 noticeable difference. Since the monthly and zonal mean 2-D ozone concentrations are specified in about
386 80% of CMIP5 models (Cionni et al., 2011) and the monthly mean 3-D ozone data are employed in
387 many CMIP6 models (Keeble et al., 2020), it is expected that the change in the radiative damping due
388 to the increase in ozone in response to the doubled CO₂ only marginally impacts the QBO periods in
389 those models that do not include an interactive chemistry. However, the 15–20% increases in the ozone
390 mixing ratio in response to the doubled CO₂ do contribute the shortening of the QBO period due to its
391 role in strengthening the tropical upwelling in the stratosphere (Bushell et al., 2010).

392 The real atmosphere involves the complex interactions among dynamics, chemistry, and radiation
393 (Andrews et al., 1987). First of all, the dynamical QBO goes hand in hand with the ozone QBO (Hasebe,
394 1994). Shibata and Deushi (2005) pointed out that the radiative heating related to the ozone QBO could
395 modify the secondary meridional circulation associated with the QBO, and consequently could modify
396 how the QBO westerlies move down faster than the QBO easterlies, leading to the elongation of the
397 QBO period in the chemically interactive models as compared with the chemically non-interactive
398 models. However, it is difficult to fathom how the intensity of the secondary meridional circulation

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403 associated with the QBO could change the QBO period due to the fact that any speedup or slowdown of
404 the descending westerly shear zones is roughly compensated by the concurrent slowdown or speedup of
405 the descending easterly shear zones. Furthermore, due to the increase of the zonal mean temperatures
406 and ozone concentrations with the altitude in the lower stratosphere, the motion, ozone, and thermal
407 waves are closely coupled. Pawson et al. (1992) found that the net linearized cooling coefficient from
408 CO₂ (Newtonian cooling) was more than compensated in the lower equatorial atmosphere arising from
409 absorption by the 9.6 μ m bands of ozone (see their figure 8). This reduced or negative radiative damping
410 in the lower stratosphere acts to lengthen the QBO period in presence of the increased ozone due to the
411 doubling of the CO₂ concentrations. Finally, taking into account the short-wave heating of eddy ozone,
412 Cordero et al. (1998) used a mechanistic model with a one-dimensional representation for mean flow
413 and a three-dimensional depiction for Kelvin and Rossby-gravity waves to demonstrate that the ozone
414 distribution, which maximizes in the middle stratosphere, leads to the local radiative damping decreases
415 by up to 15% below 35 km and its increases by up to 20%. They concluded that ozone feedbacks
416 lengthen the QBO period by about 2 months. Cordero and Nathan (2000) further developed a more
417 sophisticated mechanistic model with a two-dimensional representation for mean flow and a three-
418 dimensional depiction for Kelvin and Rossby-gravity waves, and surprisingly found that ozone
419 feedbacks had little influence on the QBO period. We believe that more studies should be conducted to
420 fully understand how to put together those various effects of interactive ozone on the QBO period.

421 Using the NASA Goddard Institute for Space Studies (GISS) Model E2.2-AP (Rind et al. 2020; Orbe
422 et al. 2020), DallaSanta et al. (2021), to the great extent, isolated the overall effect on the QBO period
423 of the increase in ozone due to the doubled CO₂. They first used a chemically non-interactive (NINT)
424 model to conduct two CMIP6 experiments: the preindustrial (pi) control run, and the doubled CO₂ (2X)
425 run. The two experiments only differ in the CO₂ concentration with the latter being two times the former.

426 Any other specification of these two experiments is the same, e.g., the ozone concentrations in the two
427 runs are identical. Their Figure 5 shows that the doubling of CO₂ in the NINT model shortens the QBO
428 period from 29.1 to 25 months. In other words, the doubling of CO₂ shortens the QBO period by 14%.
429 DallaSanta et al. (2021) then employed a chemically interactive CMIP6 model with a mass-based scheme
430 (Bauer et al., 2020), called One-Moment Aerosol (OMA), to conduct the pi control run and the 2X run.
431 The ozone concentrations simulated by the OMA model increase by 10-15% in response to the doubling
432 of CO₂ (figure not shown), which is consistent with the results of Jonsson et al. (2004). Figure 5 in
433 DallaSanta et al. (2021) indicates that the doubling of CO₂ in the OMA model shortens the QBO period
434 from 31.7 to 26.6 months. Namely, the doubling of CO₂ shortens the QBO period by 16%. In short, the
435 results of DallaSanta et al. (2021) appear to suggest that three-way interactions among dynamics,
436 chemistry, and radiation tend to slightly amplify the shortening of the QBO period in response to the
437 doubling of CO₂.

438 Note that N , the Brunt-Väisälä frequency, in Eqs. (2) and (3) also changes with increasing CO₂.
439 Richter et al. (2020b) showed that N^2 would be decreased by ~5% in the stratosphere when CO₂ is
440 doubled (refer to their Figure 2c). We used the HL model to conduct a sensitivity test by adopting $G = 0$
441 in Eq. (1) with the radiative damping profile corresponding to the doubled CO₂ and the top of the models
442 at the 48 km level. The rest of parameters in this sensitivity test are identical to those in all the previous
443 runs except that the Brunt-Väisälä frequency in this experimental run was 2.5% smaller than that in the
444 control run. The models were run for 1000 years to further increase the spectral resolution. We found
445 that when the Brunt-Väisälä frequency was decreased by 2.5%, the simulated QBO period was slightly
446 lengthened from 30 months to 30.2 months (figure not shown). In other words, the impact of decreasing
447 stratospheric buoyancy frequency on the QBO period is almost negligible.

448 Analyzing eleven CCM1-1 REF-C2 climate–chemistry simulations, Eichinger and Šácha (2020)
449 showed that the scale height in the stratosphere decreases by 2.3% per century. Accordingly, we used
450 the HL model to conduct another sensitivity test by adopting $G = 0$ in Eq. (1) with the radiative damping
451 profile corresponding to the doubled CO₂ and the top of the models at the 48 km level. The rest of
452 parameters in this sensitivity test are identical to those in all the previous control runs except that the
453 scale height in this experimental run was 2.3% smaller than that in the control run. The model was also
454 run for 1000 years for the sake of higher spectral resolution. We found that when the scale height was
455 decreased by 2.3%, the simulated QBO period was also shortened by about 2.3%, i.e., from 30 months
456 to 29.3 months (figure not shown). Apparently, the shortening of the QBO period due to the warming
457 climate is ascribed less to the shrinkage of the scale height in the stratosphere than to the enhancing of
458 the stratospheric radiative damping. Together, the shrinking scale height and the increasing radiative
459 damping shorten the QBO period by about 9.6%.

460

461 5. Conclusions

462 Plumb (1977) envisioned that stratospheric climate change would give rise to long-term changes in
463 the QBO period due to changes in radiative damping and the Brunt-Väisälä frequency. Using one-
464 dimensional (1D) models and taking into account the uncertainty due to the radiative damping rate, we
465 found that the enhanced radiative damping arising from the doubling of CO₂ leads to the shortening of
466 the QBO period by about $7.4\% \pm 0.9\%$ provided that the model top is higher than the 46 km level.
467 Furthermore, when we incorporated both the 2.3% shrinkage of the scale height and the enhanced
468 radiative damping, the QBO period is shortened by about 9.6%. In addition, the impact of decreasing
469 stratospheric buoyancy frequency is marginal. While the increased ozone in response to the doubling of
470 CO₂ appears to slightly further shorten the QBO period, more research needs to be done for the

Deleted: and increasing radiative damping due to

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473 ~~appreciation of the underlying mechanisms~~. Note that those models include neither gravity waves nor
474 tropical upwelling and assume that there are no changes in wave fluxes entering the equatorial
475 stratosphere.

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476 From a comprehensive model perspective, Richter et al. (2020b) showed that the changes in period
477 of the QBO in warming climate simulations varied quite significantly among these models. Some models
478 projected longer mean periods and some shorter mean periods for the QBO in a future warmer climate.
479 They argue that uncertainty in the representation of the parameterized gravity waves is the most likely
480 cause of the spread among the QBOi models in the QBO's response to climate change.

481 In addition, CO₂ increases in the NASA ~~GISS Model E2.2-AP~~ lead to a decrease of both QBO period
482 and QBO amplitude (DallaSanta et al., 2021). The period decrease is mostly associated with increases
483 in lower stratospheric momentum fluxes (related to parameterized convection), a finding consistent with
484 Geller et al. (2016a, 2016b) and Richter et al. (2020b). The amplitude decrease is mainly associated with
485 a strengthened residual mean circulation, also consistent with the literature, although the vertical
486 structure of the circulation response is nontrivial. It is worth mentioning that horizontal momentum flux
487 divergences could also play an important role in weakening the QBO (Match and Fueglistaler, 2019,
488 2020).

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489 Our 1D models only explored how the QBO period responds to the enhancing radiative damping of
490 planetary waves, the shrinking scale height in the stratosphere, and the decreasing stratospheric
491 buoyancy frequency due to the increasing CO₂ concentration. In order to investigate how those factors
492 affect gravity waves which play an even more important role in determining the QBO period than
493 planetary waves, high-resolution models such as those used by Kawatani et al. (2011, 2019) are desirable
494 to further our understanding. Ultimately, how the QBO period changes in response to the increasing CO₂
495 will be determined by the combined effects of the strengthening of tropical upwelling, the increasing of

499 wave fluxes entering the equatorial stratosphere, the enhancing of radiative damping, and the shrinking
500 of the scale height in the stratosphere, which warrants further research.

501

502 **Data availability**

503 Any data used in this paper can be made available from the corresponding author upon request.

504

505 **Author contributions**

506 All authors made equal contributions to this work.

507

508 **Competing interests**

509 The authors declare that they have no conflict of interest.

510

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517

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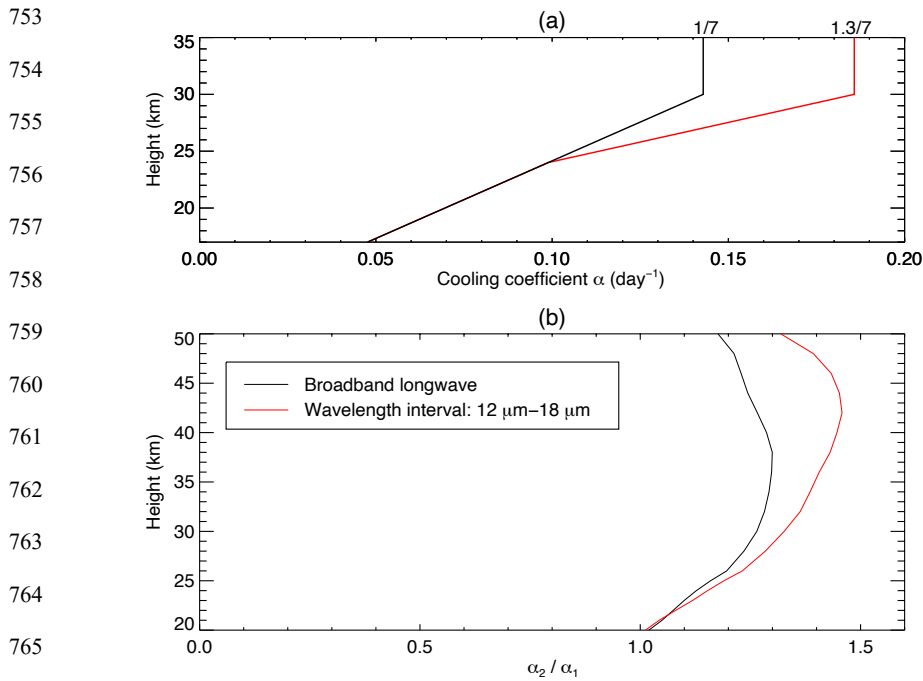
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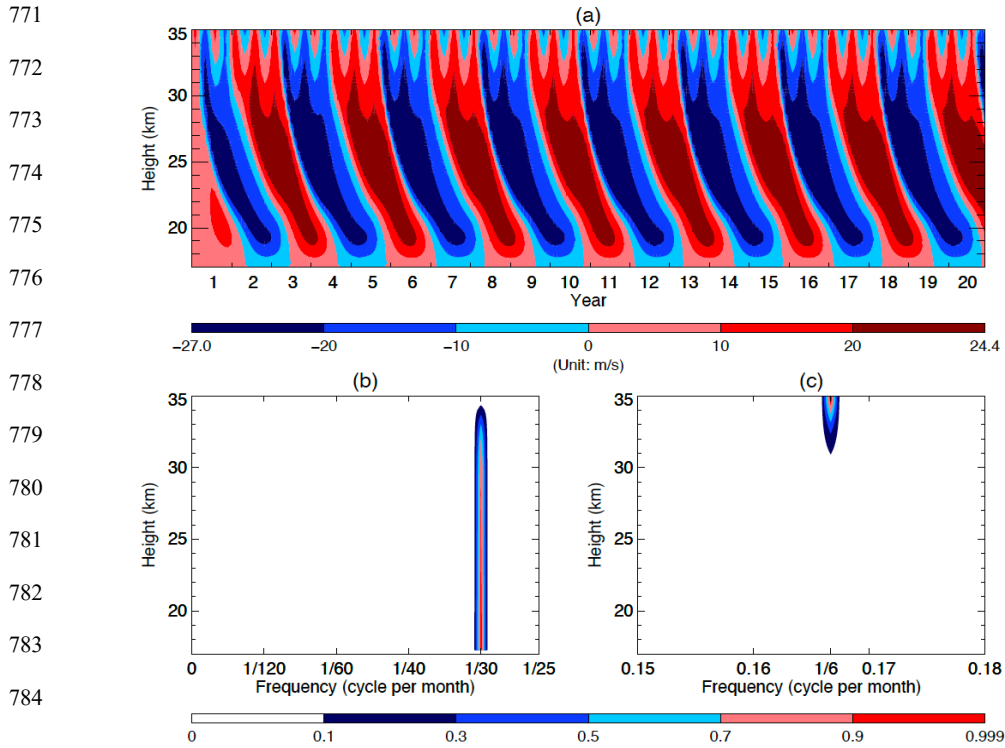
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766 **FIG. 1:** (a) Profiles of Newtonian cooling coefficients: The smaller values (black line) are used for the
 767 control runs while the larger values (red line) are used for the experimental runs. (b) Profiles of the
 768 ratio of α_2 to α_1 , where α_1 and α_2 denote the Newtonian cooling coefficient for the reference CO₂ and
 769 the doubled CO₂, respectively. The black line depicts the ratio for the broadband longwave (5 μm –
 770 100 μm) and the red line delineates that for the CO₂ absorption band (12 μm – 18 μm).



785 **FIG. 2:** (a) Time–height section of the monthly averaged mean zonal wind over the first 20 years from
786 the HL’s original model. (b) and (c) Frequency–height section of the power spectral densities (PSD) of
787 the standardized monthly averaged mean zonal wind of the 100 years. Note that in order to better
788 visualize the PSD in (b) and (c), we trimmed off the blank segments for the frequencies ranging from
789 $\frac{1}{25}$ to 0.15 cycle per month and those ranging from 0.18 to 0.5 cycle per month.

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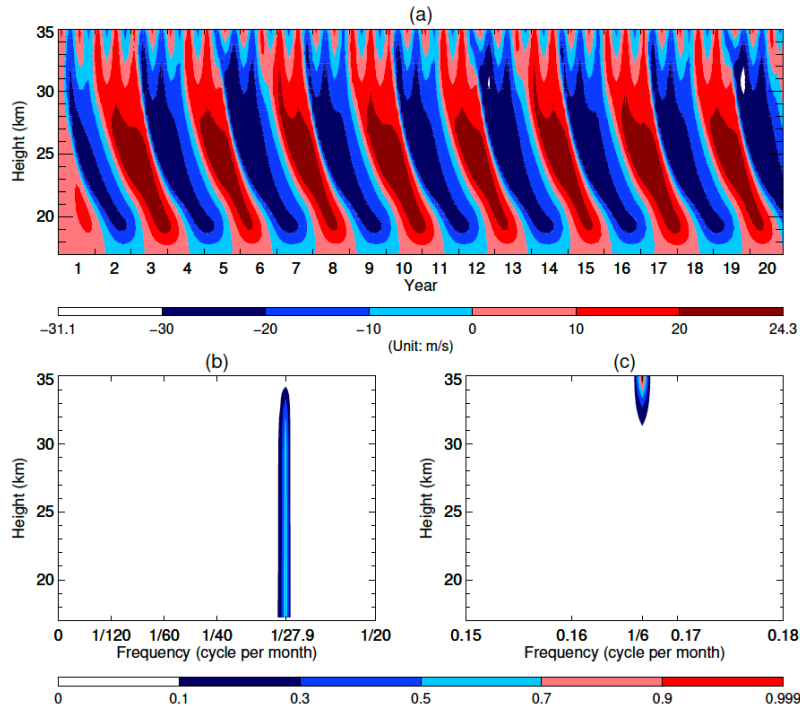
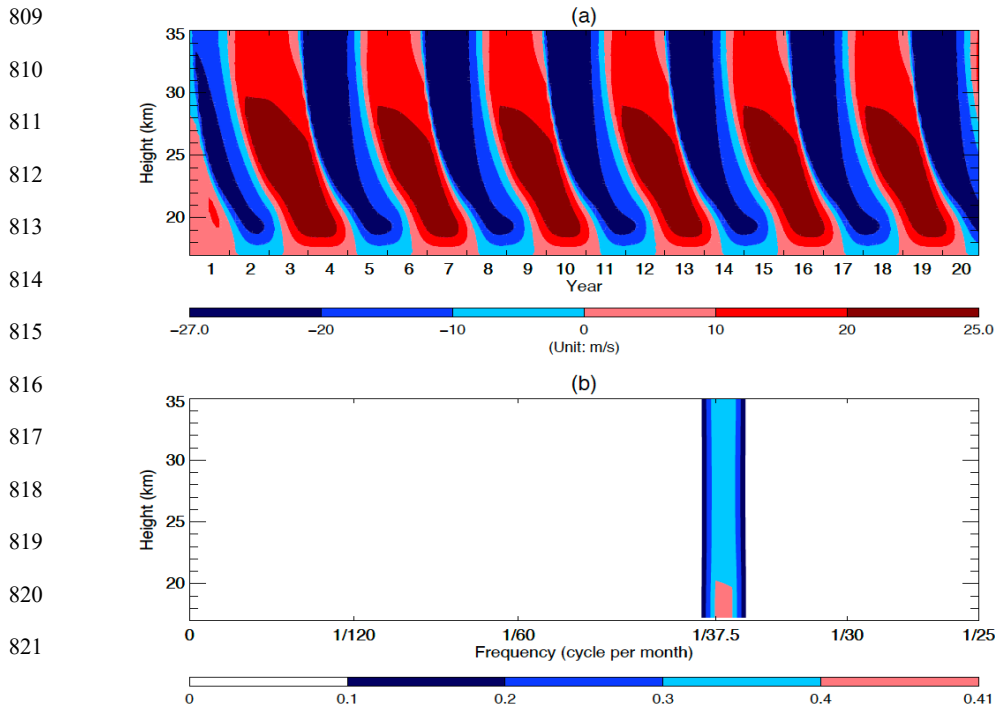


FIG. 3: (a) Same as FIG. 2a, but with the enhanced $\alpha(z)$ depicted as the red line in FIG. 1a. (b) and (c) Frequency–height section of the power spectral densities (PSD) of the standardized monthly averaged mean zonal wind of the 100 years. Note that in order to better visualize the PSD in (b) and (c), we trimmed off the blank segments for the frequencies ranging from $\frac{1}{20}$ to 0.15 cycle per month and those ranging from 0.18 to 0.5 cycle per month.



822 **FIG. 4:** (a) Time–height section of the monthly averaged mean zonal wind over the first 20 years from
 823 the HL’s model without the semiannual forcing. (b) Frequency–height section of the power spectral
 824 densities (PSD) of the standardized monthly averaged mean zonal wind of the 100 years. Note that in
 825 order to visualize the PSD, we trimmed off the blank segment for the frequencies ranging from $\frac{1}{25}$ to
 826 0.5 cycle per month.

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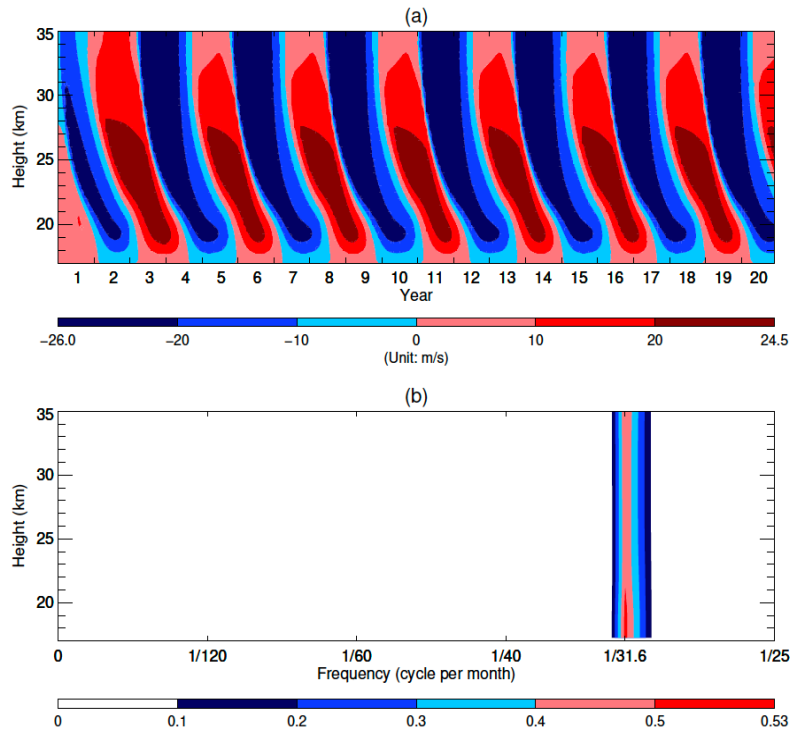
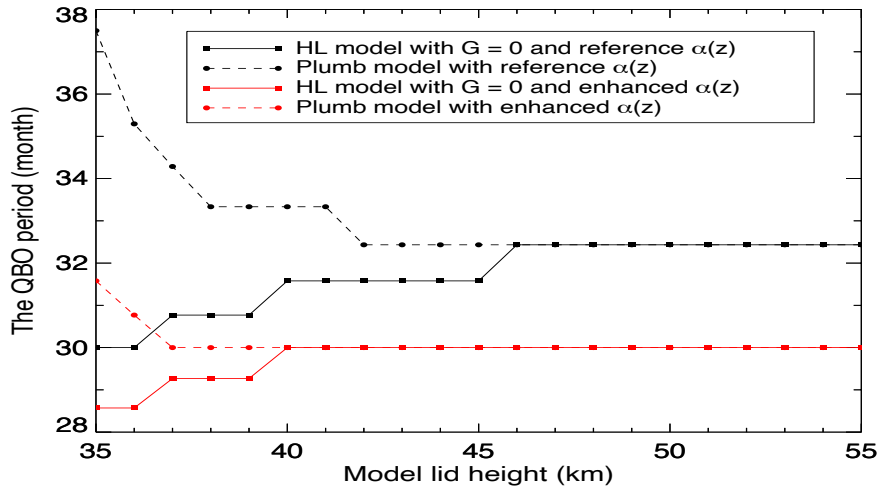


FIG. 5: (a) Same as FIG. 4a, but with the enhanced $\alpha(z)$ depicted as the red line in FIG. 1a. (b) Same as FIG. 4b, but for the doubled CO₂ Run.

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853 **FIG. 6:** The relationship between the simulated QBO period with the height of the model lid. Black and
854 red lines depict the results from using the reference radiative damping and the enhanced radiative
855 damping respectively while solid and dashed lines delineate the results from the HL model with $G = 0$
856 and the Plumb model respectively.