

# 1 **Characteristics of the summer atmospheric boundary layer height** 2 **over the Tibetan Plateau and influential factors**

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9 **Abstract** The important roles of the Tibetan Plateau (TP) atmospheric boundary layer (ABL) in climate, weather and air  
10 quality have long been recognized, but little is known about the TP ABL climatological features and their west-east  
11 discrepancies due to the scarce data in the western TP. Based on intensive sounding, surface sensible heat flux, solar  
12 radiation, and soil moisture observational datasets from the Third Tibetan Plateau Atmospheric Scientific Experiment and  
13 the routine meteorological operational sounding and ground-based cloud cover datasets in the Tibetan Plateau for the period  
14 2013-2015, we firstly investigate the west-east differences in summer ABL features over the TP and the associated  
15 influential factors. It is found that the heights of both the convective boundary layer (CBL) and the neutral boundary layer  
16 (NBL) exhibit a diurnal variation and a west-east difference in the TP, while these features are not remarkable for the stable  
17 boundary layer (SBL). Moreover, the ABL shows significant discrepancies in the amplitude of the diurnal variation and the  
18 persistent time of the development between the eastern and western TP. In the early morning (08:00 BJT), the ABL height  
19 distribution is narrow, with a mean height below 450 m above ground level (AGL) and a small west-east difference. The  
20 SBL observed at this moment accounts for 85% of the TP total ABL. There are a wide distribution in the ABL height up to  
21 4000 m AGL and a large west-east difference for the total ABL height at noon (14:00 BJT), with a mean height above 2000  
22 m AGL in the western TP and around 1500 m AGL in the eastern TP. The CBL accounts for 77% of the TP total ABL at this  
23 moment, with more than 50% of the CBL above 1900 m AGL. In the late afternoon (20:00 BJT), the CBL and SBL  
24 dominate the western and eastern TP, respectively, which results in a larger west-east difference of 1054.2 m between the  
25 western and eastern TP. The high ABL height in a cold environment over the western TP (relative to the plain areas) is  
26 similar to that in some extreme hot and arid areas such as Dunhuang and Taklimakan Deserts. In general, for the western  
27 (eastern) TP, there is low (high) total cloud coverage, with large (small) solar radiation at the surface and dry (wet) soil.  
28 These features lead to high (low) sensible heat flux and thus promotes (inhibits) the local ABL development. This study  
29 provides new insights for west-east structures of the summer ABL height, occurrence frequency, and diurnal amplitude over  
30 the TP region and the associated reasons.

## 31 **1 Introduction**

32 The atmospheric boundary layer (ABL) commonly refers to the bottom layer of the troposphere directly coupled with the  
33 earth's surface at a response time scale of about one hour or less, in which a variety of complex motions characterized by  
34 turbulence may be present (Stull, 1988). The turbulent motions in the ABL are responsible for the atmospheric mixing  
35 processes, which affects the vertical redistribution of water vapour, momentum, heat, and atmospheric pollutants (Stull, 1988;  
36 Garratt, 1992; Huang et al., 2007; Miao et al., 2015). The ABL height (ABLH) as a fundamental variable is critical to  
37 diagnose turbulent mixing, vertical disturbance, convective transport, pollutant dispersion, and atmospheric environmental  
38 and effective heat capacity (Garratt, 1993; Seibert et al., 2000; Guo et al., 2009; Esau and Zilitinkevich, 2010; Dai et al.,  
39 2014; Pal and Haefelin, 2015; Davy and Esau, 2016). Therefore, the accurate specification of the ABLH is essential to  
40 develop weather, climate, and air pollution prediction models.

41

42 The cloud-free ABL overland can be divided into three types, that is, the convective boundary layer (CBL), the stable  
43 boundary layer (SBL), and the neutral boundary layer (NBL) (Stull, 1988). The CBL usually has the strongest turbulence  
44 forced by surface buoyancy flux with or without wind shear, and is generally capped by a strong temperature inversion  
45 maintained through large-scale subsidence. The CBL height is a result of the balance of the turbulence-induced entrainment  
46 and the subsidence velocity (e.g. Driedonks and Tennekes, 1984). However, turbulence in the SBL is mainly driven by the  
47 mean wind shear against negative buoyancy flux from the stable thermal stratification within the nocturnal surface inversion  
48 (NSI). The SBL height is hence related to the boundary layer wind and wind shear, which sometimes are used to identify the  
49 SBL height. The NBL occurs in neutral conditions with the turbulence of almost the same intensity in all directions (Stull,  
50 1988; Blay-Carreras et al., 2014). It denotes the type of boundary layer with solely wind forcing and normally occurs during  
51 the transition from the daytime CBL to the night time SBL. It can also occur anytime when the buoyancy forcing is weak.  
52 The ABLH variability is dominated by its strong diurnal cycle (Stull, 1988; Garratt, 1992). In this diurnal cycle, the different  
53 manifestations of an ABLH are generated in response to the distinct forcing mechanisms that originate from mechanical  
54 (wind shear) and thermal (buoyancy) effects (Stull, 1988; Garratt, 1992). Over land and after sunrise, the surface is heated by  
55 solar radiation, resulting in upward heat flux that initiates strong updrafts of warm air. Such a mechanism generates a  
56 deepening of the CBL (Chen and Houze, 1997). At sunset, the surface cools more rapidly compared to the air above,  
57 resulting in negative heat flux that consumes turbulent kinetic energy. Consequently, shear-driven turbulence can only  
58 maintain a shallow SBL with the formation of the NSI (Zhang et al., 2011a; Miao et al., 2015). Above the NSI, the  
59 convective energy-containing eddies start to lose their strength and mixing capacity. This deep and near-adiabatic vertical  
60 region, which is the remnant of the daytime CBL, is known as the residual layer (RL). The use of precise information on the  
61 RL in numerical models is of fundamental importance when describing the evolution of the diurnal CBL (Blay-Carreras et  
62 al., 2014; Chen et al., 2016).

63

64 The ABLH can be calculated from temperature, humidity, and wind profiles (Holtslag and Boville, 1993; Seibert et al., 2000;  
65 Seidel et al., 2010; Bosveld et al., 2014; Davy, 2018). The CBL height is generally less than 2000–3000 m AGL and the SBL  
66 thickness is usually less than 400–500 m AGL (Garratt, 1992). The ABL height shows an obvious spatial variation due to  
67 differences in topography, thermal properties of the underlying surface, and weather conditions. For example, the CBL can  
68 grow to the height of 4700 m AGL in New Delhi before the outbreak of the South Asian monsoon, whereas it only reaches  
69 900 m AGL in Bangalore during the monsoon period (Raman et al., 1990). Seidel et al. (2010, 2012) pointed out that a large  
70 east-west spatial gradient of the ABLH at sunset in the United States spanning several time zones may be conflated with the  
71 diurnal variations of the ABL for the local solar time in the west earlier than in the east at fixed observation times. Guo et al.  
72 (2016) identified three large-scale ABLH spatial patterns in China, that is, a west-east gradient during sunrise, an east-west  
73 gradient during sunset, and a south-north gradient at noon. The reasons for the first two patterns are similar to those in the  
74 United States shown in Seidel et al. (2012), while the south-north gradient may be related to the local surface and  
75 hydrological processes (Guo et al., 2016; Zhang et al., 2017).

76  
77 The Tibetan Plateau (TP) with an average elevation exceeding 4000 m is characterized by complex land surface processes  
78 and boundary layer structures (Tao and Ding, 1981; Yanai and Li, 1994; Xu et al., 2002; Yang et al., 2004; Li and Gao, 2007;  
79 Sun et al., 2007; Zhao et al., 2019b). The ABLH in the TP can reach 2000–3000 m AGL, generally higher compared to some  
80 plains areas (with the ABLH of 1000–1500 m AGL) (Ye and Gao, 1979; Zhao and Miao, 1992; Xu et al., 2002; Zhang et al.,  
81 2003). The ABLH in the TP varies greatly with location and season. At Gaize station of the western TP, the super-thick  
82 ABLH may exceed 5000 m AGL during winter (Chen et al., 2013, 2016). In the central TP, the ABLH is lower, between 400  
83 and 1800 m AGL at Dangxiong station and 1750 m AGL at Namucuo Lake (Li et al., 2000; Liu et al., 2001, and Lü et al.,  
84 2008). Moreover, there is also a significant difference in the TP ABLH between dry and rainy seasons (Zuo et al., 2004). For  
85 instance, the ABLH at Naqu station is 2211–4430 m AGL in the dry season, while it is 1006–2212 m AGL in the rainy  
86 season (Li et al., 2011).

87  
88 Although observations and studies for the TP ABL features have made progress, routine meteorological operational  
89 sounding observations are scarce in the western TP due to the local high elevations, naturally harsh environmental conditions,  
90 and logistic challenges. The previous studies on the ABL in the western TP are usually based on observational data at  
91 Shiquanhe (during 15 days in one summer) and Gaize (during 22 days in one summer) stations (Song et al., 1984; Chen et al.,  
92 2013). Thus, the statistical representation of their results is limited. Moreover, there are significant differences in surface  
93 properties and general climate between the eastern and western TP (Wang et al., 2016). Few studies examined the west-east  
94 differences in the ABL features due to the scarce data in the western TP. To obtain a longer observational data in the western  
95 TP, the Third Tibetan Plateau Atmospheric Scientific Experiment (TIPEX-III) has made routine sounding launches at  
96 Shiquanhe, Gaize, and Shenzha stations of the western TP (Fig. 1) since 2013, which fills in the data gaps in the operational  
97 sounding network over the western TP (Zhao et al., 2018). Meanwhile, the TIPEX-III also carried out the intensive sounding

98 observations in the TP and adjacent stations at 14:00 Beijing Time (06:00 UTC) in June, July, and August (Zhao et al., 2018).  
99 Compared to the previous field experiments over the TP, the TIPEX-III has a wider and longer coverage of sounding  
100 observations over the western TP, providing valuable observational data for studying the ABL features in the western TP and  
101 the west-east differences of these features in the TP during summer.

102

103 This study utilizes the TIPEX-III sounding observational data to investigate the features of the ABLH in the TP and their  
104 differences between the western and eastern TP during summer, and analyzes the major factors affecting the ABLH in the  
105 TP. The remainder of this paper is organized as follows. Main features of data and methods are described in Section 2. In  
106 Section 3, the characteristics of the ABLH in the eastern and western TP and their regional differences are analyzed in detail.  
107 In Section 4, the major factors affecting the ABLH in the TP and the west-east differences are examined. Discussions and  
108 conclusions are given in Section 5.

## 109 **2 Data and analysis methods**

### 110 **2.1 Observation data**

111 The TIPEX-III carried out the intensive routine meteorological sounding observations at Shiquanhe (SQH), Gaize (GZ), and  
112 Shenza (SZ) stations of the western TP (marked by red dots in Fig. 1) since the 2013 summer (Zhao et al., 2018), which have  
113 been applied in research on the vertical structure of the upper troposphere and lower stratosphere at Gaize station during the  
114 rainy season and the effects of assimilating the intensive sounding data on downstream rainfall (Hong et al., 2016; Yu et al.,  
115 2018; Zhao et al., 2018; Zhao et al., 2019b). These intensive sounding data and the routine meteorological operational  
116 sounding data at 16 stations of the central-eastern TP from the China Meteorological Administration (marked by black dots  
117 in Fig. 1) are utilized in this study. The sounding observations at the above intensive and operational sounding stations were  
118 carried out at 08:00 Beijing Time (BJT; 00:00 UTC), 14:00 BJT (06:00 UTC), and 20:00 BJT (12:00 UTC) each day during  
119 summer (June, July, and August), including vertical profiles of temperature, humidity, and wind direction and speed. After  
120 the quality control of the sounding observational data, we select data from three time periods for this study: 15 June to 31  
121 July 2013, 15 June to 31 August 2014, and 1 June to 31 August 2015. There are 11,635 sounding profiles (Fig. 1a) from 19  
122 stations over the TP region consisting of 4745, 2049, and 4841 profiles at 08:00 BJT (Fig. 1b), 14:00 BJT (Fig. 1c), and  
123 20:00 BJT (Fig. 1d), respectively. It is evident that the observational sample size used in this study is much more compared  
124 to the previous studies. Meanwhile, it is noted that there is a large difference in the sample size between the intensive and  
125 operational observation records at 08:00 BJT and 20:00 BJT (Fig. 1b and d), which is called the original dataset for  
126 convenience. Consequently, we also select the test group dataset which contains the same intensive observation records as  
127 the operational ones at these two times to make sensitivity analysis (shown in Section 3.2), which shows that the difference  
128 in the sample size between the intensive and operational observation records does not change our conclusions.

129

130 To analyze the factors affecting the ABL in the TP, we use the TIPEX-III 30-min mean surface sensible heat flux (SHF),  
 131 downward solar radiation, and 5-cm soil volume moisture content at SQH (bare soil with less vegetation), Naqu (NQ; alpine  
 132 steppe), and Linzhi (LZ; alpine meadow with few shrubs and trees) stations in the 2014-2015 summers (Wang et al., 2016;  
 133 Zhao et al., 2018; Li et al., 2019, 2020). In addition, the manual operational ground-based cloud cover observations at 02:00,  
 134 08:00, 14:00, and 20:00 BJT from the China Meteorological Administration are also used in this study. These ground-based  
 135 could cover data have been utilized by Guo et al. (2016) and Zhang et al. (2017).

136

## 137 **2.2 Calculation method of ABLH**

138 The potential temperature gradient method, proposed by Liu and Liang (2010) and sketched in Fig. 2a, is utilized in  
 139 identifying the ABL type and calculating the ABL height. The CBL height is defined at the base of the overlying inversion  
 140 layer that caps the rising convective thermals. The SBL height is defined as the top of the underlying inversion layer, where  
 141 turbulence decreasing from the surface nearly ceases (Stull 1988). In the evening and morning transition periods when the  
 142 RL may occur, the neutral RL starting from the surface is identified with near-neutral conditions in the surface layer (that is  
 143 the NBL). Following Liu and Liang (2010), Zhang et al. (2017), and Zhao et al. (2019a), the original sounding observation  
 144 profiles with a fine vertical resolution of  $\sim 1$  hPa are interpolated to a vertical resolution of 5 hPa (corresponding to a  
 145 vertical interval around 50 m in the ABL) by the nearest neighbor interpolation method. On the basis of the near-surface  
 146 thermal gradient, such as a potential temperature ( $\theta$ ) difference ( $PTD$ ) between the fifth layer ( $\sim 250$  m;  $\theta_5$ ) and the second  
 147 layer ( $\sim 50$  m;  $\theta_2$ ), the ABL is classified as follows.

$$148 \quad PTD = \theta_5 - \theta_2 \begin{cases} < -\sigma, & \text{for CBL} \\ > +\sigma, & \text{for SBL.} \\ \text{else,} & \text{for NBL} \end{cases} \quad (1)$$

149 Here  $\sigma$  is the stability threshold of the near-surface potential temperature stratification. Since the neutral stratification  
 150 condition ( $\sigma = 0$ ) is rare in nature, consistent with Liu and Liang (2010),  $\sigma$  is set to 1.0 K. The threshold value of the NBL is  
 151 set to -1.0 to 1.0. Consequently, SBLs and CBLs with weak stable or unstable stratification are possibly identified as NBLs.

152

153 Once the boundary layer regime has been identified, we use the criteria defined by Liu and Liang (2010) to estimate the  
 154 ABLH for each regime. Since buoyancy is the dominant mechanism driving turbulence in the CBL, the ABLH is defined as  
 155 the height at which an air parcel rising adiabatically from the surface becomes neutrally buoyant (Stull 1988). First, we find  
 156 the lowest level ( $k_1$ ) (Fig. 2a) that meets the following condition

$$157 \quad \theta_{k_1} - \theta_1 \geq \sigma_u, \quad (2)$$

158 in which  $\sigma_u$  is the  $\theta$  increment that represents the minimum strength of the unstable layer. Once level  $k_1$  is determined,  
159 another upward scan is performed to find the lowest level at which the potential temperature gradient with height ( $\dot{\theta}_k$ ) meets  
160 the following criteria

$$161 \quad \dot{\theta}_k \equiv \frac{\partial \theta_k}{\partial z} \geq \dot{\theta}_r. \quad (3)$$

162 Here  $\dot{\theta}_r$  is the minimum strength for the overlying inversion layer and can be considered as the overshooting threshold of  
163 the rising parcel to define the scope of the entrainment zone for the CBL. The same procedure is adopted to determine the  
164 NBL height excluding the entrainment zone at the top (Fig. 2a). Various values of  $\sigma_u$  and  $\dot{\theta}_r$  will affect the determination of  
165 the boundary layer height and they are respectively set to 0.5 K and 4.0 K km<sup>-1</sup>, consistent with Liu and Liang (2010).  
166 Quantifying the uncertainty of the rawinsonde-based approach for identifying ABLH is important, which is closely related to  
167 the thermodynamic characteristics of the sounding profiles (Seidel et al., 2010, 2012; Davy, 2018; Lee and Pal, 2021). The  
168 ABLH determined by this potential temperature gradient method from soundings is highly consistency with that derived  
169 from lidar measurements, with a correlation coefficient of 0.96 and root-mean-square error of 211 m (Liu and Liang 2010).  
170 Moreover, the changes in ABLHs are  $\leq 177$  m when using 3.5, 4, and 4.5 K km<sup>-1</sup> as  $\dot{\theta}_r$ , respectively (Zhang et al., 2017). It  
171 is evident that the uncertainties of the above procedure can be negligible for both CBL and NBL, since most of their ABLHs  
172 are much higher. For the SBL, the turbulence in the ABL can result from either buoyancy forcing or wind shear. The SBL  
173 height is defined as the lower of the heights of both the thermal stable layer from the surface and the maximum wind in the  
174 low-level jet stream if present. More details of the definitions of the boundary layer regimes may be seen in Liu and Liang  
175 (2010). Figure 2c-d shows the typical profiles of potential temperature for CBL, NBL, and SBL at 20:00 BJT on June 10,  
176 2013, July 21, 2013, and August 11, 2013 at Lasa station, and the ABL heights calculated by the potential temperature  
177 gradient method are 3465, 1258, and 409 m AGL, respectively.

## 178 **3 Characteristics of the summer ABLH in the eastern and western TP**

### 179 **3.1 A general characteristic of the ABLH**

180 The diurnal variation is an important feature of the ABL, consisting of different periods of daytime, night-time, and  
181 day/night transitions (Liu and Liang, 2010). In the central TP (near 90 °E), 08:00 BJT, 14:00 BJT, and 20:00 BJT correspond  
182 to 06:00 (the early morning), 12:00 (noon), and 18:00 (the late afternoon) local solar time (LST) (Fig. 1b-c), respectively. To  
183 reveal a difference in ABLH between the eastern TP (ETP) and the western TP (WTP), we divide all sounding stations in the  
184 TP into two groups. One is for the WTP (to the west of 92.5 °E) with 8 stations, and the other is for the ETP (to the east of  
185 this longitude) with 11 stations.

186

187 Figure 3a-c shows the spatial distributions of the mean ABLH over the TP at 08:00 BJT, 14:00 BJT, and 20:00 BJT,  
188 respectively. In the early morning (08:00 BJT), the ABL is of the night-time property. The ABLH is generally low (<450 m  
189 AGL) over the TP and displays a relatively homogeneous feature (Fig. 3a). At this moment, the distribution of the ABLH is  
190 narrow, with a frequency peak of 35% at the ABLH of 300 m AGL (Fig. 3d) and 78.5% (99.6%) of the ABLH below 500  
191 (1000) m AGL (Fig. 3e). Figure 3f displays the zonal sections of the ABLH along 32°N, in which the cross section includes  
192 SQH, GZ, SZ, NQ, CD, GanZ, and HY stations. In this figure, the ABLH varies between 218.4 and 433.9 m AGL from east  
193 to west and presents a relatively homogeneous feature in the west-east direction.

194

195 At noon (14:00 BJT), with the well-developed daytime ABL (Fig. 3b), its height remarkably increases over the TP with an  
196 average of 1887.7 m AGL and exhibits a large west-east difference. There is a wide distribution of the ABLH up to 4000 m  
197 AGL, with a relatively flat peak between 900 and 2900 m AGL (Fig. 3d) and only 17.8% (more than 50%) of the ABLH  
198 below 1000 (above 1900) m AGL (Fig. 3e). At this moment, the regional mean ABLH is 2124.2 m AGL in the WTP and  
199 1693.5 m AGL in the ETP, with a mean difference of 430.7 m between the WTP and the ETP. Along 32°N, the ABLH  
200 remarkably increases from 1379.4 m AGL at GanZ station to 2504.2 m AGL at SQH station, with the west-east difference  
201 exceeding 1200 m (Fig. 3f). This regional difference in the TP ABLH could be likely related to the hydrologic factors such  
202 as air moisture and soil water (also see Section 4) that may modulate the spatial distribution of the daytime ABLH (Seidel et  
203 al., 2012).

204

205 In the late afternoon (20:00 BJT), the ABL begins to show the night-time features. The ABLH also starts to decrease in the  
206 ETP, with the regional mean height < 1000 m AGL, while it continues to increase at the west-most stations, with the regional  
207 mean height > 2000 m AGL (Fig. 3c). This result indicates a larger west-east difference (1054.2 m) between the WTP and  
208 the ETP. Especially, the ABLH is 602 m AGL at HY station and 2920.6 m AGL at SQH station, with a difference above  
209 2000 m between these two stations (Fig. 3f). At this moment, the frequency of the high ABLH decreases, with 12.8% of the  
210 frequency peak at the ABLH of 300 m AGL (Fig. 3d) and 50% of the ABL heights less than 1000 m AGL (Fig. 3e). It is  
211 evident that the west-east difference of the ABLH over the TP increases from noon to the late afternoon. During the evening  
212 transition, the daytime boundary layer undergoes a transition to the night-time boundary layer. Since the TP spans almost 1.5  
213 time zones from west to east (Fig. 1c), the local solar time is earlier in the west (where 20:00 BJT corresponds to 17:20 LST  
214 in the westernmost SQH station) compared to the east (where 20:00 BJT corresponds to 18:50 LST for the easternmost HY  
215 station), which supports an earlier transition from the daytime ABL to the night-time ABL in the east (Seidel et al., 2010,  
216 2012; Guo et al., 2016; Lee and Pal, 2017). Meanwhile, it is noted that this difference in the local time is less over TP than  
217 over China (Guo et al., 2016) and the United States (Seidel et al., 2010, 2012; Lee and Pal, 2017). Thus the contribution of  
218 the time zone difference to the regional difference of ABLH is relatively smaller in TP.

219

220 Figure 4 further shows the variations of the ABLH from 08:00 BJT to 14:00 BJT and from 14:00 BJT to 20:00 BJT,  
221 indicating varying rates in 6 h. It is seen from Fig. 4a that the ABLH in the TP increases substantially from 08:00 to 14:00  
222 BJT, with a mean growth rate of 1500 m/6 h. There is also a large west-east difference of the ABLH growth rate in this  
223 period, with the regional mean of 1800 m/6 h and 1300 m/6 h in the WTP and the ETP, respectively. From 14:00 to 20:00  
224 BJT (Fig. 4b), the growth rate of the ABLH is negative in the ETP, exhibiting an opposite trend to that in Fig. 4a, which  
225 indicates a significant decrease (around -600 m/6 h) of the ABLH after noon. In the WTP, the growth rate generally shows a  
226 weak increase (around 400 m /6 h) or decrease (around -140 m /6 h). It is evident that the growth rate from 08:00 to 14:00  
227 BJT may indicate the amplitude of the ABL diurnal variation over the TP. Compared to the ETP, the ABL in the WTP has  
228 the larger amplitude of the diurnal variation and the longer development time.  
229

### 230 **3.2 Characteristics of SBL, NBL, and CBL heights**

231 We further examine the characteristics of different ABL types. Figure 5 presents the spatial distribution of occurrence  
232 frequency of SBL, NBL and CBL at 08:00 BJT, 14:00 BJT, and 20:00 BJT. It is seen that the occurrence frequency exhibits  
233 significant discrepancies at different times for the SBL and CBL. At 08:00 BJT, the occurrence frequency of the SBL/CBL is  
234 large/little (Fig. 5a/Fig. 5g), with a mean value 84.9%/8.5% over the TP. At 14:00 BJT, the occurrence of the SBL/CBL  
235 remarkably decreases/increases, accounting for 3.1%/76.9% of the ABL (Fig. 5b/Fig. 5h). At 20:00 BJT, the SBL/CBL  
236 mainly occurs in the ETP/WTP (Fig. 5c/Fig. 5i), with a regional mean of 35.0%/65.0%. However, the NBL shows a  
237 relatively weaker temporal variation over the TP (Fig. 5d-f), with the mean occurrence frequency of 6.4%, 20.0%, and 25.5%  
238 at 08:00 BJT, 14:00 BJT, and 20:00 BJT, respectively. The above results are consistent with the diurnal development of the  
239 ABL structure including the SBL in the early morning, the CBL at noon, and different types of ABLs between the eastern  
240 and western TP in the late afternoon because of the latitudinal difference and the resultant difference in local solar times.  
241 Note that the observations were made simultaneously for all stations. Nevertheless, the daytime SBL and the night-time CBL  
242 may also occur with low frequencies in the TP, which is likely due to the ‘abnormal’ forcing associated with certain synoptic  
243 conditions or cloud coverage (Medeiros et al., 2005; Poulos et al., 2002; Stull, 1988).  
244

245 To analyse the temporal variations of the ABLH over the TP, the ABL height-occurrence frequency relationships for the  
246 SBL, NBL, and CBL at 08:00 BJT, 14:00 BJT, and 20:00 BJT are presented in Fig. 6a-f. For the SBL, the frequency  
247 distribution of the ABLH shows the similar feature at three measurement times (Fig. 6a-c) and is characterized by a narrow  
248 single mode, with the frequency peaks of 39.0%, 28.1%, and 36.6% at the ABLH of 200, 300, and 300 m AGL at 08:00,  
249 14:00, and 20:00 BJT, respectively, which indicates small temporal variations of the SBL height due to its turbulence  
250 inhabited. Moreover, the SBL height above 80% is < 600 m AGL and the cumulative frequency of the SBL height exceeding  
251 1000 m AGL is little (near zero) at 08:00, 14:00, and 20:00 BJT (Fig. 6d, e, and f). For the NBL and CBL, however, their  
252 heights vary strongly with time under the influence of surface heating in the daytime. At 08:00 BJT (Fig. 6a), the



253 distributions of the NBL and CBL heights are narrow, with the frequency peaks of 27.5% and 35.1% at the ABLH of 300 m  
254 AGL for NBL and CBL, respectively, similar to that of the SBL, which is possibly due to the initial development of the CBL  
255 and NBL in the early morning. At 14:00 BJT, the CBL and NBL have a wide distribution of the ABLH up to 4000 m AGL,  
256 with a relatively flat peak between 1000 and 3000 m AGL, which is remarkably different from a single peak of the SBL. The  
257 frequency of the NBL height between 500 and 3000 m AGL is generally less than 5% (Fig. 6b), with a frequency peak of 6.1%  
258 at 1000 m AGL, and more than 50% NBL height exceeds 1700 m AGL (Fig. 6e). The height of the CBL is higher, with a  
259 frequency peak near 4.5% between 1500 and 2500 m AGL (Fig. 6b) and more than 50% CBL height is above 2000 m AGL  
260 (Fig. 6e). These results show that the ABL develops well at noon. When the ABL begins turning to the night-time property at  
261 20:00 BJT (Fig. 6c and 6f), the distributions of the CBL and NBL heights are still wide but the frequency of the high ABL  
262 height decreases, with the frequency peak below 500 m AGL. It is obvious that the CBL and NBL heights show the similar  
263 results consistent with those from Zhang et al (2017). Stull (1988) and Blay-Carreras et al. (2014) revealed that the NBL  
264 often occurs in the transition periods between the CBL and the SBL. Since these transitions occur rapidly, the NBL may  
265 have the same characteristics in the state variables as the CBL prior to the transition although the dynamic forcing in the  
266 NBL maybe weaker compared to the CBL.

267

268 To reveal the spatial variations of the ABLH over the TP, the distributions of mean SBL, NBL, and CBL heights at 08:00  
269 BJT, 14:00 BJT, and 20:00 BJT are illustrated in Fig. 7. The SBL height is generally low and varies between 200 and 730 m  
270 AGL at these times, with a mean height of 336.0 m AGL at 08:00 BJT, 356.0 m AGL at 14:00 BJT, and 321.9 m AGL at  
271 20:00 BJT (Fig. 7a-c), which indicates the weak spatial differences of the SBL height over the TP at three observation times.  
272 For the NBL and CBL, their heights are still low in the early morning (Fig. 7d and 7g), with the ABLH < 450 m AGL, and  
273 have small spatial differences. At noon (Fig. 7e and 7h), the NBL and CBL heights rapidly increase, especially in the WTP,  
274 which leads to a remarkable east-west gradient in the ABL height. At this moment, there is a regional mean NBL/CBL  
275 height of 2074.6/2191.4 m AGL in the WTP and 1594.8/1788.0 m AGL in the ETP, with a difference of 479.8/403.4 m  
276 between the WTP and the ETP. In the late afternoon (Fig. 7f and 7i), the NBL/CBL height continues to increase in the WTP,  
277 with a regional mean of 2092.0/2192.2 m AGL, while the NBL/CBL height begins decreasing in the ETP, with a regional  
278 mean of 1423.1/1237.2 m AGL. This varying feature in the ETP and WTP results in the larger differences of 668.9/955.0 m  
279 in the NBL/CBL height between the WTP and ETP. Thus there is a significant difference in the frequency distribution of the  
280 ABL height between the ETP and the WTP in the daytime (Fig. 6g). The cumulative frequency contours gradually go  
281 upward from east to west (Fig. 6h). The eastern TP is dominated by a low CBL height, with the peak of 14.4% at the height  
282 of 350 m AGL (Fig. 6g) and the 50% (5%) CBL height below 1000 m AGL (above 2500 m AGL) (Fig. 6h). For the WTP,  
283 the strong peak of 4%-10% corresponds to the high CBL between 2500 and 3500 m AGL (Fig. 6g), especially at SQH  
284 station, and there are larger CBL heights, with almost 50% CBL extending upward to more than 2500 m AGL, almost 10%  
285 reaching 4000 m AGL or higher, and only 15% CBL below 1000 m AGL (Fig. 6h).

286

287 To investigate an effect of differences in the sample profiles shown in Fig. 1b and d, we use the test group dataset to repeat  
288 the above analyses. Figures 8a and b show the scatter plots of the occurrence frequency of the SBL, NBL, and CBL from the  
289 original and test group datasets at each of 19 stations at 08:00 BJT and 20:00 BJT, respectively. It is seen that the correlation  
290 coefficients between the two datasets are 0.92-0.99, with root-mean-square errors (RMSEs) of the occurrence frequency  
291 between 1.1% and 2.7%. The similar results are also seen in the SBL, NBL, and CBL heights at 08:00 BJT (Fig. 8c) and  
292 20:00 BJT (Fig. 8d). The correlation coefficients in the ABL height are 0.90-0.99. The RMSE of the SBL height is 14 m and  
293 25 m at 08:00 and 20:00 BJT, respectively. The RMSE of the CBL and NBL heights are 54-59 m at 08:00 BJT and 99-107m  
294 at 20:00 BJT. These high correlations and small errors show that the difference in the sample size does not change our  
295 conclusions.

296

297 From the foregoing analysis, the CBL and NBL heights in the TP show remarkable temporal variations and west-east spatial  
298 differences, while these features are not remarkable for the SBL. From noon to the late afternoon, the NBL and CBL are  
299 deeper in the WTP compared to the ETP, with the ABLH difference between the WTP and the ETP exceeding 600 m AGL  
300 at 20:00 BJT. Then, which factors contribute to this difference in the ABL between the WTP and ETP? In the following  
301 section, we examine some factors that may be responsible for the ABL height over the TP.

#### 302 **4 Factors responsible for the ABL height over the TP**

303 Previous studies have addressed effects of surface sensible heat flux (SHF), soil volume moisture content (VWC), downward  
304 solar radiation flux (DSR), and the cloud cover (CLD) on ABL height (Liu, et al., 2004; Zhao et al., 2011; Sanchez-Mejia  
305 and Papuga, 2014; Rihani et al., 2015; Lin et al., 2016; Zhang et al., 2017; Zhang et al., 2019; Qiao et al., 2019). However,  
306 these studies paid little attention to reasons for the west-east difference of the ABL between the eastern and western TP. To  
307 investigate a possible reason for this difference, we utilize the TIPEX-III SHF, DSR, and VWC at SQH, NQ, and LZ stations,  
308 and the corresponding meteorological operational CLD observations to analyze the relationships between these variables and  
309 the ABL height.

310

311 The driving force of turbulence in the ABL is the surface buoyancy flux as a result of surface and air temperature and  
312 humidity differences and the mean surface layer wind. The kinematic heat flux (KHF) and kinematic moisture flux (KMF) at  
313 the surface are the two directly factors responsible for the surface buoyancy flux (Brooks and Rogers, 2006). Since KMF is  
314 often small, KHF associated with SHF is examined as a major component of buoyancy flux in dry conditions over land.  
315 According to the method of Brooks and Rogers (2006), our calculation results show that the contribution from KMF to  
316 surface buoyancy flux is below 18% at SQH, NQ, and LZ stations. Moreover, the ABL may be largely affected by the effect  
317 of cumulative SHF in the daytime (Zhang et al., 2019). Thus we analyse the possible effect of SHF on the ABL. Figure 9a-c  
318 presents the scatter plots between the mean SHF over the past six hours and the ABL height at SQH, NQ, and LZ stations.

319 As shown in this figure, the correlation is 0.80, 0.81, and 0.71 (significant at the 99% confidence level) at these stations,  
320 respectively. When SHF is strong, the turbulent motion is strong and the ABL develops, which is consistent with the result  
321 of Zhang et al. (2011b). Their result shows a significant correlation of 0.78 in the arid area of Northwest China between the  
322 ABL thickness and the cumulative SHF. Figures 10a and b further present the features of the ABL height and SHF at SQH,  
323 NQ, and LZ stations. The mean value of SHF is 85 W/m<sup>2</sup>, 42 W/m<sup>2</sup>, and 33 W/m<sup>2</sup> at SQH, NQ and GZ stations, respectively,  
324 and has a large difference (52 W/m<sup>2</sup>) between SQH and NQ stations. This result indicates a decreasing trend of SHF from  
325 west to east in the TP, consistent with a reduction of the ABL height from SQH via NQ to LZ station (shown in Figs. 3 and  
326 10a). In addition, Fig. 11 demonstrates the diurnal variations of SHF and the ABL height at SQH, NQ, and LZ stations. The  
327 duration of positive SHF in a diurnal cycle at SQH, NQ and GZ stations is 14, 12 and 11 hours, respectively, and indicates a  
328 decreasing trend from west to east in the TP. It is clear that the peak of the SHF occurs earlier than the maximum ABLH in a  
329 diurnal cycle at SQH station. The maximum ABL height occurs near 20:00 BJT (approximately 17:20 LST), corresponding  
330 to a strong SHF. At LZ station, however, the SHF turns into a negative value at 20:00 BJT (18:20 LST) and then the ABL  
331 height decreases. Some past studies show that the development of ABL height generally lags the development of SHF, and  
332 ABL depth growth continues even after SHF attains the maximum daytime value until the time of early evening transitions  
333 (Chen et al., 2016; Zhang et al., 2019). Consequently, the difference in the ABL height between the WTP and ETP is closely  
334 associated with a west-east difference in SHF that is as a direct thermal factor for the ABL development in the TP.

335

336 The solar radiation at the surface is an important component of the surface energy budget, affecting surface temperature and  
337 SHF. We show the scatter plots between the 6-hour mean DSR and the ABL height at SQH, NQ, and LZ stations (Fig. 9d-f).  
338 The ABL height is highly correlated with the 6-hour average of DSR at these stations, with the correlation coefficients of  
339 0.86, 0.81, and 0.73, respectively, which is equivalent to those of SHF. The mean DSR shows a decreasing trend from SQH  
340 (510 W/m<sup>2</sup>) to LZ (200 W/m<sup>2</sup>) station. Since the solar irradiance at the surface is negatively associated with the local cloud  
341 cover (Guo et al., 2016; Lin et al., 2016; Li et al., 2017; Zhang et al., 2017), the cloud cover is also correlated to the ABL  
342 height. Figure 9g-i shows that the 6-hour mean CLD has significant correlations of -0.56, -0.65, and -0.54 with the ABL  
343 height at SQH, NQ, and LZ stations, respectively. A decrease of the mean ABL height from SQH to LZ station (Fig. 10a) is  
344 corresponded to an increase of cloud cover (Fig. 10d) and a decrease of DSR (Fig. 10c). When cloud cover is between 0 and  
345 20%, the mean ABL height for the NBL and CBL is 2019 m AGL/2732 m AGL in the ETP/WTP; and when cloud cover  
346 is >80%, the ABL height decreases to 741 m AGL/1626 m AGL in the ETP/WTP (Fig. 12). Therefore, the increased cloud  
347 cover inhibits the development of both the NBL and CBL. The difference in cloud cover between the WTP and ETP  
348 contributes to the west-east distribution of DSR and SHF, also finally contributing to the difference of the ABL development.  
349 Corresponding to more cloud cover in the ETP, the local ABL is more closely associated with atmospheric moisture  
350 processes.

351

352 Soil moisture is also an important factor affecting SHF. Low soil moisture generally coincides with a high surface sensible  
353 heat flux, which facilitates the ABL development (e.g., McCumber and Pielke, 1981; Sanchez-Mejia and Papuga, 2014;  
354 Rihani et al., 2015). Figure 9j-l shows that the relationship between the ABL height and the 6-hour mean VWC at SQH, NQ,  
355 and LZ stations. The ABL height at LZ station is negatively correlated to the local soil moisture, with a significant  
356 correlation coefficient of -0.45. This result indicates that the ABL height is lower when surface soil is moister. However, the  
357 negative correlation is weaker at SQH station, with a correlation coefficient of -0.21. This difference between the WTP and  
358 the ETP may be associated with the climatic feature of the local soil moisture. The surface type transitions from alpine  
359 meadow with few shrubs and trees or alpine steppe in the ETP to bare soil with few obstacles in the WTP (Wang et al.,  
360 2016). Accordingly, soil moisture decreases gradually from the ETP to the WTP (Fig. 10e), with a mean value of soil  
361 moisture below  $0.10 \text{ m}^3/\text{m}^3$  at SQH station and  $0.38 \text{ m}^3/\text{m}^3$  at LZ station. Little soil moisture in the WTP has a weak  
362 modulation to the local surface heat flux, which may lead to a weak correlation between the ABL height and soil moisture in  
363 the WTP.

## 364 **5 Summary and discussion**

365 Using the summer TIPEX-III intensive and meteorological operational observational datasets, we examine the ABL features  
366 and the relationships of the ABL height with surface sensible heat flux, solar radiation, cloud cover, and soil moisture in the  
367 TP region. The main conclusions are summarized as follows.

368

369 Generally speaking, the ABL height exhibits diurnal variations and regional differences in the TP, especially for the CBL  
370 and NBL. These features are weak for the SBL. Compared to the ETP, the ABL in the WTP has the larger amplitude of the  
371 diurnal variation and the longer development time. In the early morning, the ABL height is generally low over the TP, not  
372 showing a large west-east difference, and the distribution of the ABL height is narrow, with 78.5% of the ABL height < 500  
373 m AGL. At noon, the CBL and NBL heights remarkably increase and have a wide distribution in the ABL height up to 4000  
374 m AGL, with more than 50% of the ABL height exceeding 1900 m AGL. Their heights exhibit a large west-east difference.  
375 At this moment, the distribution of the SBL height is also narrow, with the cumulative frequency of 80% at the height of 500  
376 m AGL, and there is no remarkable west-east difference. In the late afternoon, there are a narrow distribution of the SBL  
377 height and wide distributions of both the NBL and CBL heights. At this moment, the ABL height continues to increase in the  
378 WTP, while it begins to decrease in the ETP. This feature results in a larger west-east difference in the ABL height. In spite  
379 of a cold environment in the TP (relative to plain areas), the WTP still has the ABL height above 2000 m AGL, which is  
380 similar to some extreme hot and arid areas such as Dunhuang and Taklimakan Deserts. In the ETP, the ABLH is similar to  
381 that in North China (1500-1900 m AGL) and is generally larger compared to the East Asian summer monsoon region (<  
382 1500 m AGL) such as the Yangtze River Delta and the Pearl River Delta (Zhang et al., 2011; Guo et al., 2016; Zhang et al.,  
383 2017; Qiao et al., 2019).

384

385 The occurrence frequency of the SBL and CBL in the TP shows remarkable temporal variations. Most (few) of the SBLs  
386 (CBLs) occur in the early morning and the occurrence frequency rapidly decreases (increases) at noon, accounting for 3.6%  
387 (76.9%) of the ABL in the TP. Possibly owing to a difference in the solar elevation angle with respect to longitude in the late  
388 afternoon, the SBL and CBL dominate the ETP and WTP, respectively. However, the NBL shows a relatively weak temporal  
389 variation over the TP, with the mean occurrence frequency of 6.4% in the early morning and around 20% at noon and in the  
390 late afternoon.

391

392 The ABL height is significantly correlated to SHF, DSR, and cloud cover in the TP and is also closely associated with soil  
393 moisture in the ETP. The decreasing trends in both SHF and DSR and the increasing trends in both cloud cover and soil  
394 moisture from west to east may cause the corresponding west-east reduction in the ABL height. In the WTP (ETP), with low  
395 (high) cloud cover, there is larger (smaller) downward solar radiation at the surface. Meanwhile, corresponding to bare soil  
396 (alpine meadow or steppe) in the WTP (ETP), there is a dry (wet) soil condition. These features cause high (low) sensible  
397 heat flux, thus promoting (inhibiting) the local ABL development. The above factors affecting the WET and ETP ABL  
398 heights are summarized in Fig. 13.

399

400 The Tibetan Plateau is an area very sensitive to global climate change, which exerts important thermal and dynamical effects  
401 on the general circulation and climate through the unique and complex land surface and boundary layer processes. Owing to  
402 new sounding observations in the WTP, our analysis firstly reveals remarkable west-east differences in the ABL height,  
403 occurrence frequency, and diurnal amplitude over the TP region during summer. It is noted that there is a big drop in the  
404 CBL height from 3000–4000 m AGL to 1000–2000 m AGL from the WTP to the ETP. Such a steep west-east  
405 inhomogeneity in the TP (with an East-West spatial scale of about 2000 km) is remarkably different from the regional  
406 variability of the ABLH on much larger scales (~4000 km) such as in the United States (Seidel et al., 2012) and in China  
407 (Guo et al., 2016). This unique inhomogeneity in the TP may trigger the local mesoscale circulation and precipitation (Segal  
408 et al., 1992; Goutorbe et al., 1997; Huang et al., 2009; Zhang et al., 2019; Qiao et al., 2019). Therefore, the influences of the  
409 west-east differences in the ABLH over the TP on the local weather and climate should be further studied in the future. In  
410 addition, this study merely investigates the characteristics of the summer ABLH in TP due to the limitation of the intensive  
411 sounding observations. More efforts should be made to expand the climatology of ABLH to other seasons in TP when more  
412 sounding data are available.

413

414 Code and data availability. All data used are available from the TIPEX-III on its homepages (<http://data.cma.cn/tipex>).

415

416 Author contributions. J.C. designed the study, analyzed the data and wrote the manuscript. P.Z. contributed to the study  
417 design, supervisor, and writing of the manuscript.

418

419 Competing interests. The authors declare that they have no conflict of interest.

420

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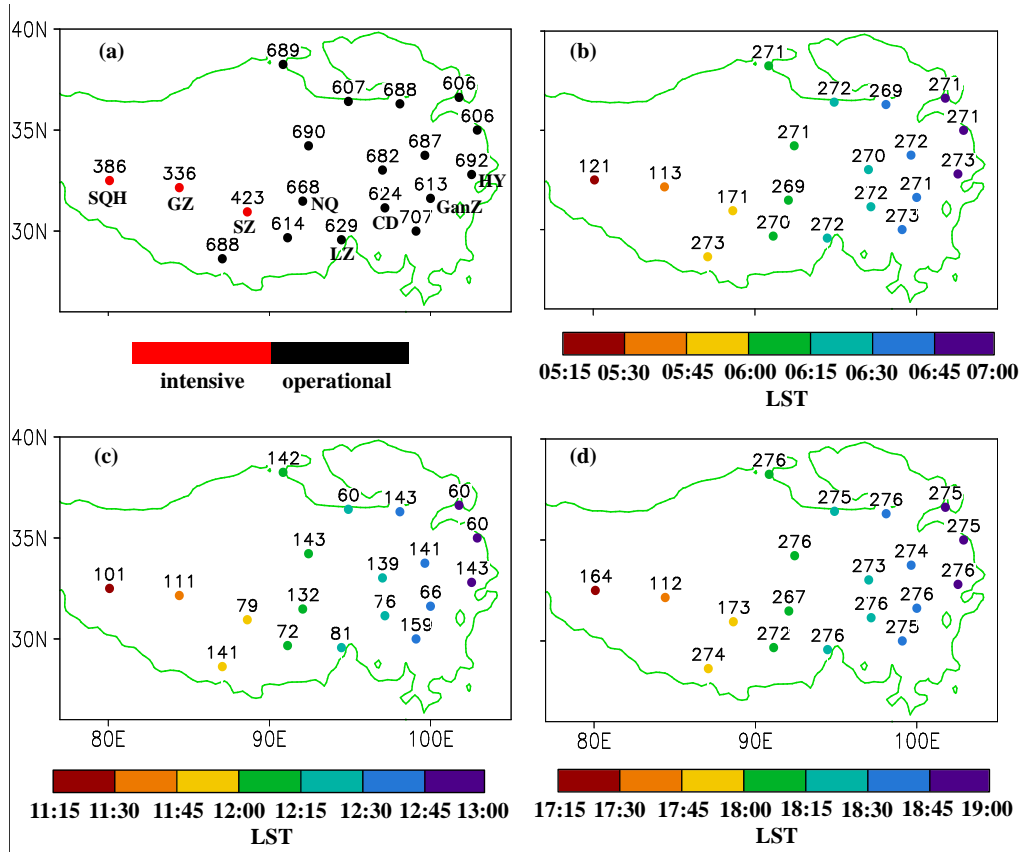
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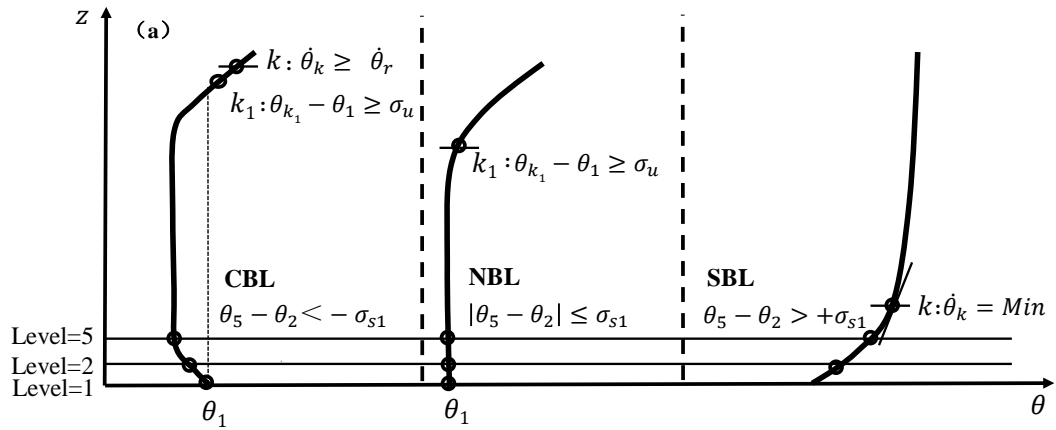
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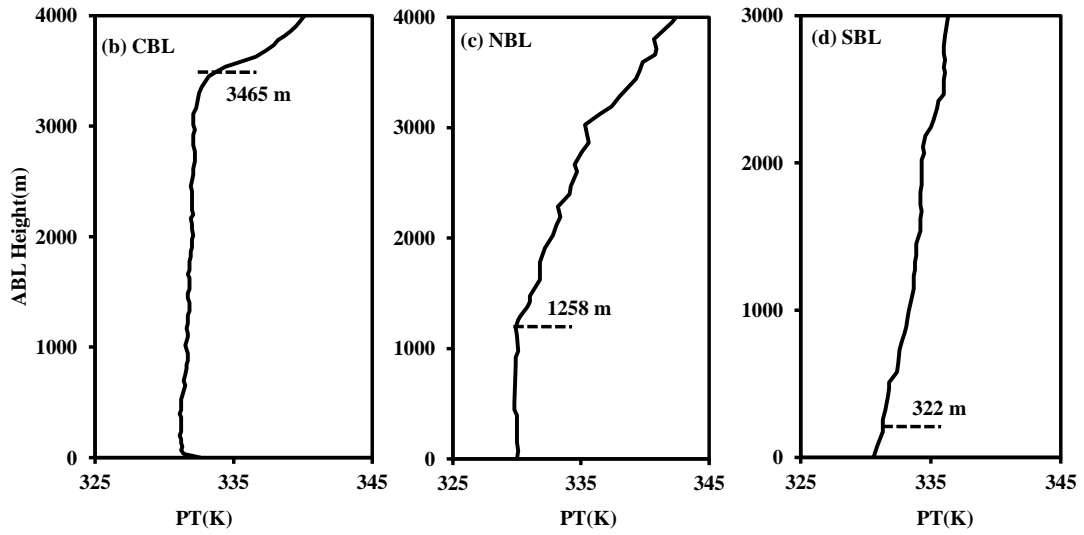
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 589 **Figure 1: Distribution of sounding stations, in which the number indicates sounding profiles at each station at (a) 08:00, 14:00, and**  
 590 **20:00 BJT, (b) 08:00 BJT, (c) 14:00 BJT, and (d) 20:00 BJT in the study period. Red (black) dots represent intensive (operational)**  
 591 **observations, and some observation station names are given as abbreviations in (a). Colored dots represent the local time of the**  
 592 **BJT time in (b), (c) and (d). The green line shows the 3000 m topography.**

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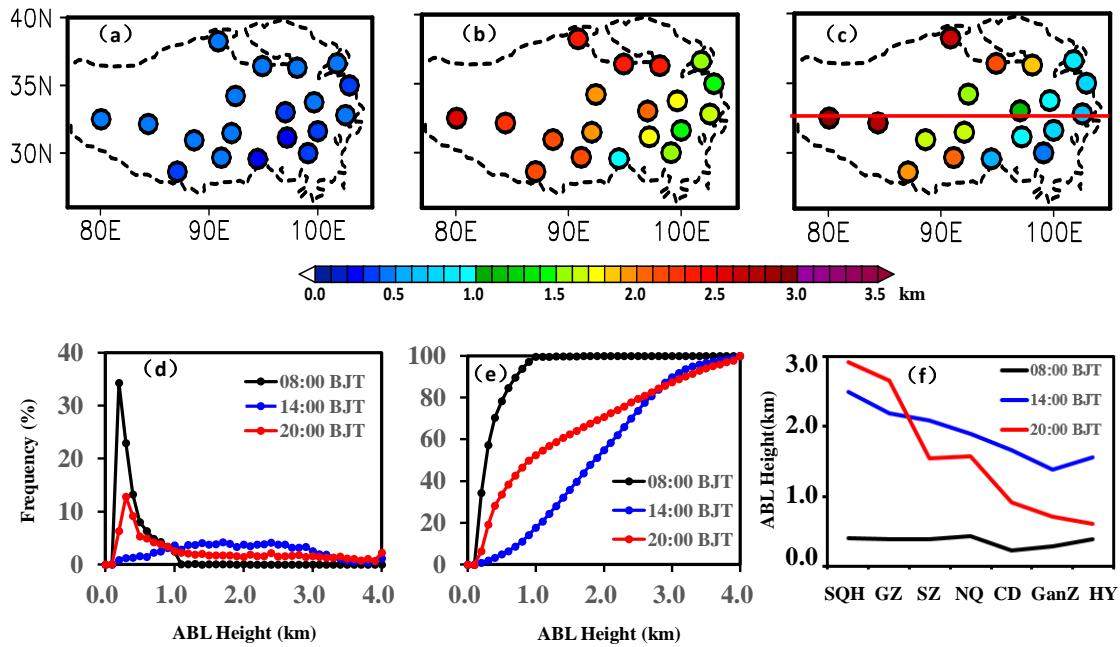
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597 **Figure 2:** (a) Illustration of the determination procedure for the convective boundary layer (CBL), neutral boundary layer (NBL),  
 598 and stable boundary layer (SBL) heights; and examples of the potential temperature (PT) profiles derived from sounding  
 599 observation at Lasa station at 20:00 BJT for (b) CBL on June 10, 2013, (c) NBL on July 21, 2013, and (d) SBL on August 11, 2013,  
 600 respectively. The dash line in (b)-(d) represents the ABL height identified using the algorithm described.

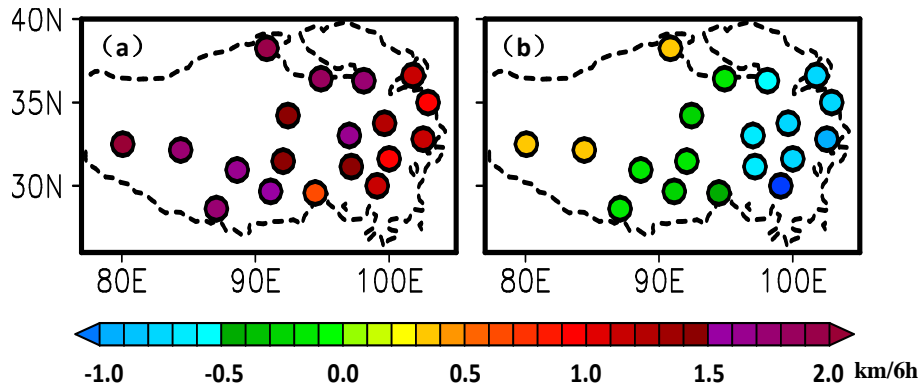
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603 **Figure 3: Spatial distribution of the mean ABL height (ABLH) at (a) 00:08 BJT, (b) 14:00 BJT, and (c) 20:00 BJT; (d) the regional**  
 604 **mean frequency and (e) cumulative frequency distributions of the ABLH in the TP at 08:00 BJT, 14:00 BJT, and 20:00 BJT; (f)**  
 605 **the west-east cross sections of the ABLH along 32°N (indicated by red line in (c)) at 08:00 BJT, 14:00 BJT, and 20:00 BJT.**

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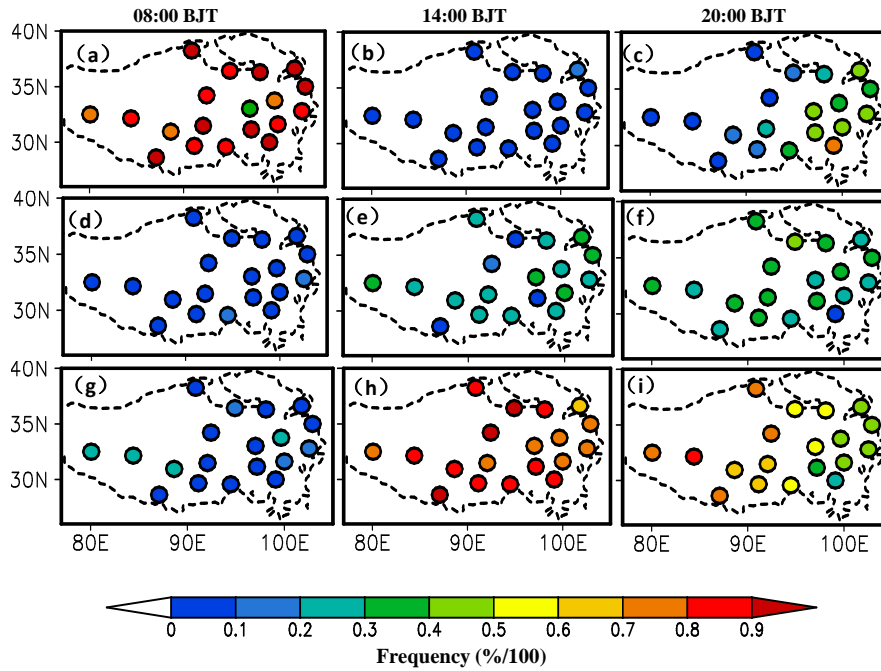


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609 **Figure 4: Spatial distribution of the ABLH growth rate from 08:00 BJT to 14:00 BJT (a) and from 14:00 BJT to 20:00 BJT (b).**

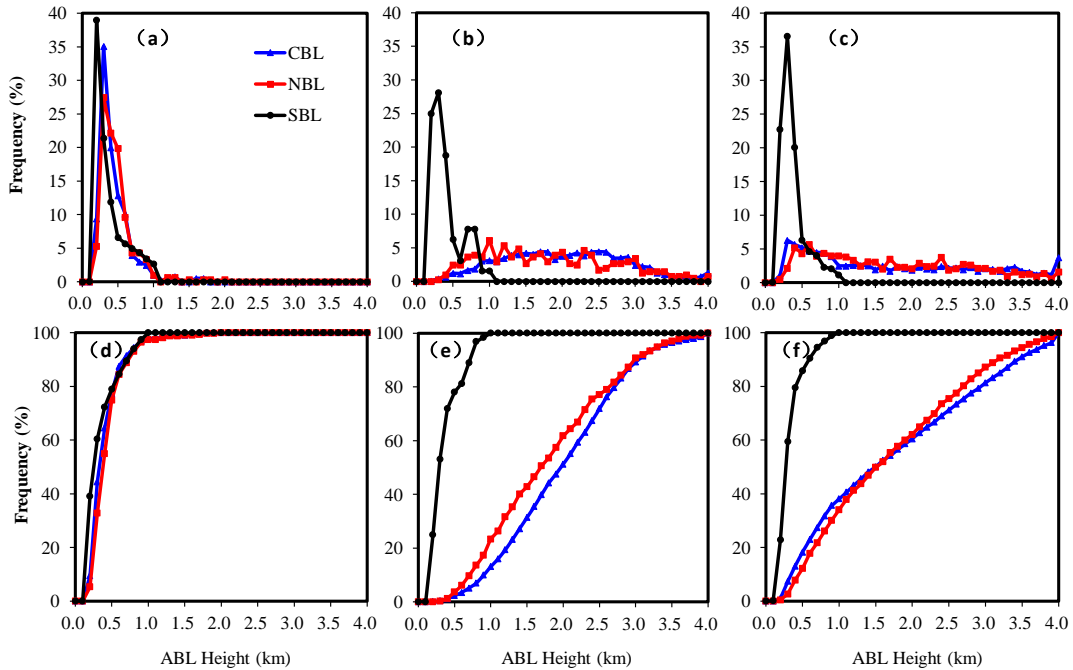
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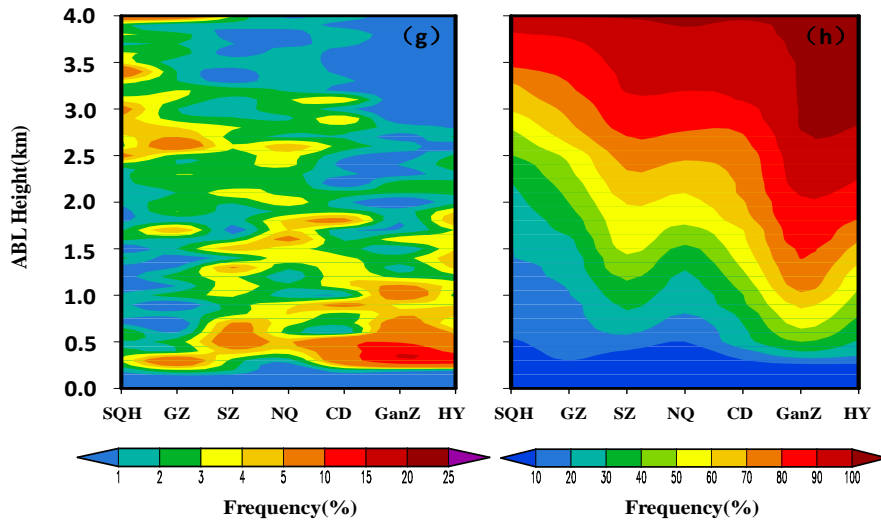
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613 **Figure 5: Spatial distribution of the occurrence frequency for the SBL (top), NBL (middle), and CBL (bottom) at 08:00 BJT, 14:00**  
 614 **BJT, and 20:00 BJT.**

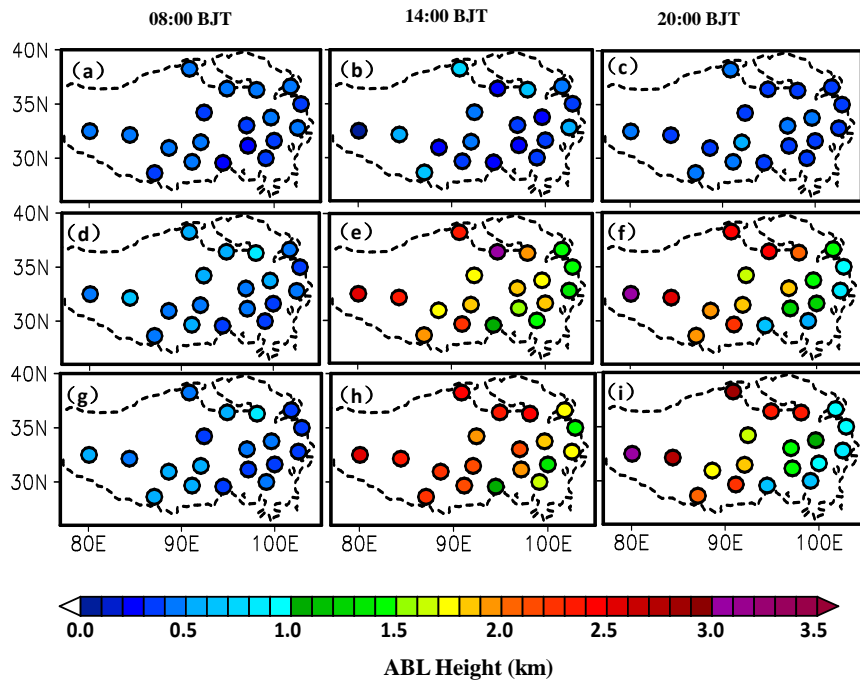


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617 **Figure 6: The regional mean frequency distributions of the ABLH over the TP for the CBL (blue), NBL (red), and SBL (black) at**  
 618 **(a) 08:00 BJT, (b) 14:00 BJT, and (c) 20:00 BJT; and (d)-(f) same as in (a)-(c) but for the cumulative frequency distributions; and**  
 619 **the west-east cross sections of frequency (g) and cumulative frequency (f) distributions of the CBL height along 32°N in the**  
 620 **daytime (14:00 and 20:00 BJT).**



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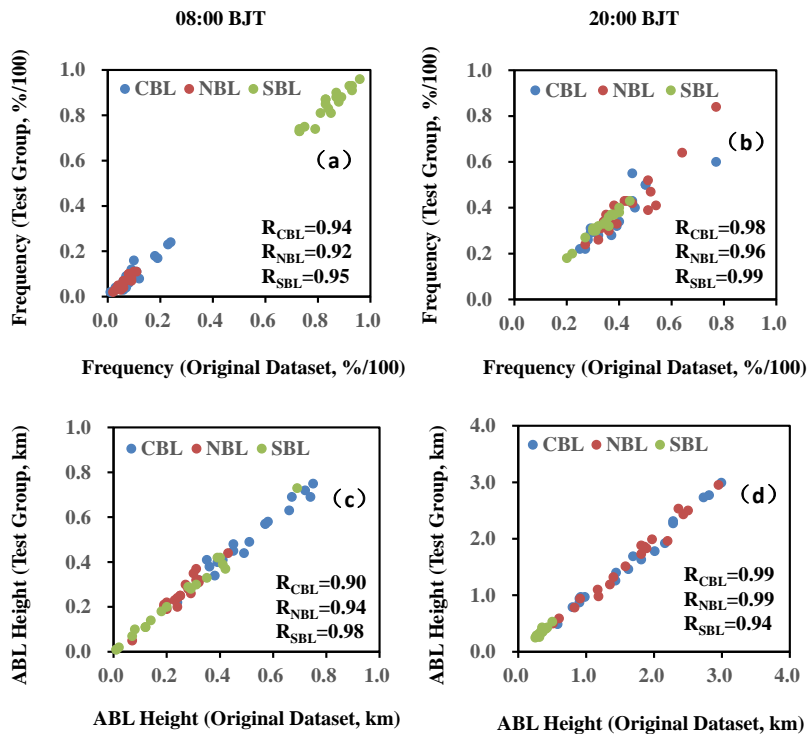
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623 **Figure 7: Spatial distributions of the mean ABLH for the SBL (top), NBL (middle), and CBL (bottom) at 08:00 BJT, 14:00 BJT,**  
 624 **and 20:00 BJT.**

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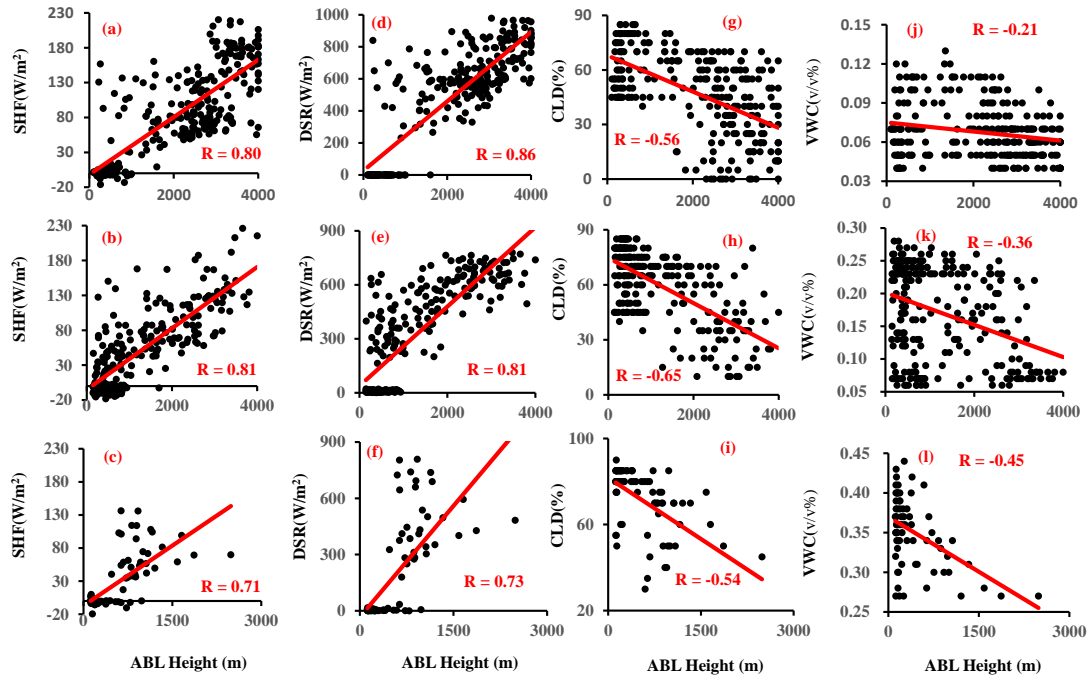
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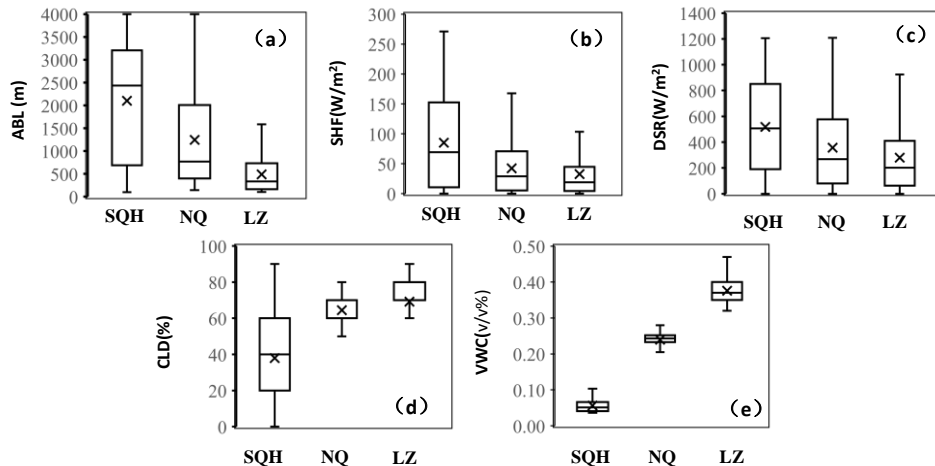
629 **Figure 8:** The scatter plots of occurrence frequency of the SBL, NBL, and CBL for the original and test group datasets at 19  
630 stations at (a) 08:00 BJT and (b) 20:00 BJT; and (c)-(d) same as in (a)-(b) but for the ABLH.





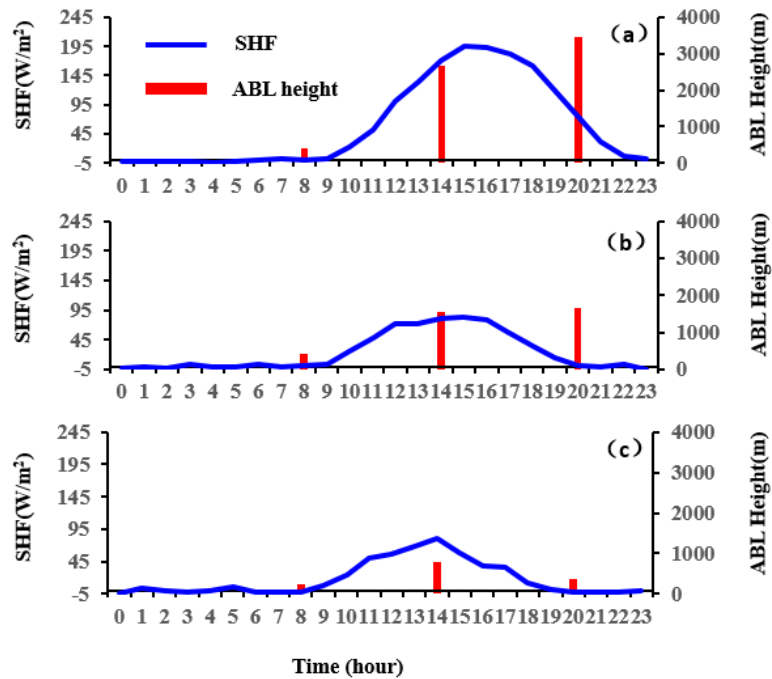
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632 Figure 9: Scatter plots of the ABLH and the 6-hour average of surface sensible heat flux (SHF) (a-c), surface downward solar  
 633 irradiance (DSR) (d-f), total cloud coverage (CLD) (g-i), and surface soil volume moisture content (VWC) (j-l) at 08:00 BJT, 14:00  
 634 BJT, and 20:00BJT at SQH (top), NQ (middle), and LZ (bottom) stations in the study period. The correlation coefficient (R) is  
 635 given in each panel.



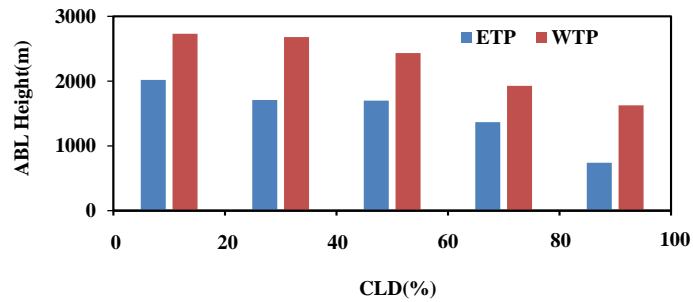
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637 Figure 10: (a) The ABLH, (b) SHF, (c) DSR, (d) CLD, and (e) VWC at SQH, NQ, and LZ stations in the study period. Horizontal  
 638 bars show the 5th, 25th, 50th, 75th, and 95th percentile values and "x" symbols show the corresponding mean value.



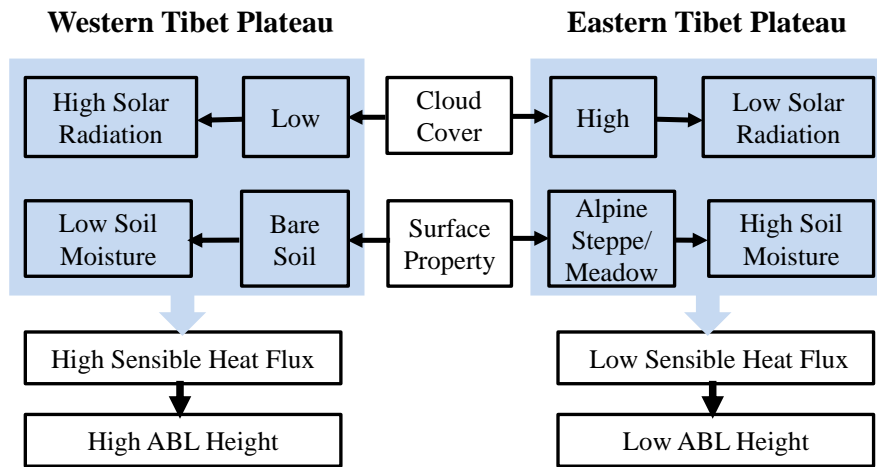
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640 **Figure 11: Diurnal variations of surface sensible heat flux (blue) and the ABLH (red) averaged over the study period at (a) SQH,**  
 641 **(b) NQ, and (c) LZ stations.**



642

643 **Figure 12: The mean ABLH (for the NBL and CBL) and CLD over the ETP (blue) and WTP (red) in the daytime (14:00 BJT and**  
 644 **20:00 BJT).**



645

646 **Figure 13: The schematic diagram for relationships between the ABLH and the influential factors in the ETP and the WTP.**