



1 Dependency of Particle Size Distribution at Dust Emission on 2 Friction Velocity and Atmospheric Boundary-Layer Stability

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10 **Abstract.** Particle size distribution of dust at emission (dust PSD) is an essential quantity to be estimated in dust studies. It
11 has been recognized in earlier research that dust PSD is dependent on soil properties (e.g. whether soil is sand or clay) and
12 friction velocity, u_* , a surrogate for surface shear stress and descriptor for saltation bombardment intensity. This recognition
13 has been challenged in some recent papers, causing a debate on whether dust PSD is “invariant” and the search for
14 justification. In this paper, we analyze dust PSD measured in the Japan-Australian Dust Experiment and show that dust PSD
15 is dependent on u_* and on atmospheric boundary-layer stability. By simple theoretical and numerical analysis, we explain the
16 three reasons for the latter dependency. First, under similar mean wind conditions, the mean of u_* is larger for unstable than
17 for stable conditions. Second, u_* is stochastic and its probability distribution profoundly influences the magnitude of the
18 mean saltation flux due to the non-linear relationship between saltation flux and u_* . Third, in unstable conditions, turbulence
19 is usually stronger, which leads to higher saltation-bombardment intensity. This study confirms that dust PSD depends on u_* ,
20 and more precisely, on the probability distribution of u_* , which itself is stability dependent. We restate that for a given soil,
21 finer dust is released in case of stronger saltation.

22 1 Introduction

23 Gillette (1981) explained that dust emission can be produced by aerodynamic lift and saltation bombardment, but under
24 realistic wind, aerodynamic-lift emission is much weaker than saltation-bombardment emission. This hypothesis was
25 confirmed by Shao et al. (1993). It is recognized that saltation bombardment is the most important mechanism for dust
26 emission and dust emission rate, F , is proportional to streamwise saltation flux, Q .

27 Rice et al. (1995, 1996) visualized the process of saltation bombardment using wind-tunnel photos: a saltation particle at
28 impact on surface ejects a tiny amount of soil into air, leaving behind a crater. Models for estimating crater size have been
29 developed by, e.g., Lu and Shao (1999). The fraction of dust that gets emitted from the ejection is difficult to estimate,
30 because it depends both on inter-particle cohesion and bombardment intensity. Since inter-particle cohesion depends on



31 particle size, d , the fraction of dust emitted must also depend on d . Thus, for a given soil, the particle size distribution of dust
32 at emission (emission-dust PSD), $p_s(d)$, must depend on saltation bombardment or on friction velocity, u_* ($\sqrt{\tau}/\rho$ with τ being
33 surface shear stress and ρ air density). Alfaro et al. (1997) confirmed that $p_s(d)$ depends on u_* : as u_* increases, $p_s(d)$ shows a
34 higher fraction of dust of smaller d . Based on this result and the observations that different laboratory techniques for PSD
35 analysis yield profoundly different outcomes, depending on the disturbances applied to the samples (Figure 1), Shao (2001)
36 suggested to use a minimally-disturbed PSD, $p_m(d)$, as the limit of $p_s(d)$ for weak saltation, and a fully-disturbed PSD, $p_f(d)$,
37 as the limit of $p_s(d)$ for strong saltation. In this way, $p_s(d)$ is approximated as a weighted average of $p_m(d)$ and $p_f(d)$, namely,

$$38 \quad p_s(d) = \gamma p_m(d) + (1 - \gamma) p_f(d) \quad (1)$$

39 where $0 \leq \gamma \leq 1$ is an empirical function of $u_{*f}(d)$, the threshold friction velocity for particles of size d .

40 PSD of dust in air (airborne-dust PSD) has been collected from different places under different conditions. Airborne-dust
41 PSD and emission-dust PSD are not the same, unless airborne dust is observed close to the dust source and the dependency
42 of particle diffusivity on d is neglected. Available data of airborne-dust PSDs give the impression that they do not differ
43 much. It has thus been suggested that airborne-dust PSDs may be “not-so-different” and hence emission-dust PSDs may also
44 be “not-so-different”. Reid et al. (2008) stated that “on regional scales, common mode dust is not functionally impacted by
45 production wind speed, but rather influenced by soil properties such as geomorphology ...”. Kok (2001a, 2001b) proposed a
46 dust emission model by treating dust emission as a process of aggregate fragmentation by saltation bombardment. Since
47 aggregate fragmentation is similar to brittle fragmentation, the size distribution produced in the process is scale-invariant
48 (Astrom, 2006). Kok (2001a, 2001b) then proposed an emission-dust PSD and estimated its parameters from airborne-dust
49 PSDs. The proposed emission-dust PSD is frequently used in dust models (Pisso et al., 2019). However, whether the “not-so-
50 different” airborne-dust PSDs justify “brittle fragmentation” as the underlying process for dust emission requires scrutiny.

51 In comparison, the airborne-dust PSD measurements of Rosenberg et al. (2014) pointed to larger fraction of fine particles
52 than in earlier published data. Khalfallah et al. (2020) reported that airborne-dust PSD depends on atmospheric boundary-
53 layer (ABL) stability, and attributed this to the dependency of particle diffusivity on particle size. They stated that the
54 dependency of emission-dust PSD on u_* , as observed by Alfaro et al. (1997), may be of secondary importance in natural
55 conditions compared to its dependency on ABL stability.

56 The confusion surrounding emission-dust PSD prompted us to examine the data of Ishizuka et al. (2008) from the Japan-
57 Australian Dust Experiment (JADE). In JADE, airborne-dust PSD were measured at small height directly above the dust
58 source and well represents the emission-dust PSD. Hence, hereafter we no longer distinguish airborne- and emission-dust
59 PSD but simply refer emission-dust PSD as dust PSD. By composite analysis for different u_* and ABL stabilities, we show
60 that dust PSD depends on u_* , supporting the findings of Alfaro et al. (1997), and depends on ABL stability, supporting the
61 findings of Khalfallah et al. (2020). But in contrast to Khalfallah et al. (2020), we argue that these dependencies are not
62 mutually exclusive, but collectively point to the simple physics that dust PSD is dependent on saltation-bombardment
63 intensity and efficiency.



64 2 JADA Data

65 JADE was carried out during 23 Feb ~ 14 Mar 2006 on an Australian farm at (33°50'42.4"S, 142°44'9.0"E) (Ishizuka et al.,
66 2008, 2014). The 4 km² farmland was flat and homogeneous such that the JADE data are not affected by fetch. In JADE,
67 atmospheric variables, land surface properties, soil PSD and size-resolved sand and dust fluxes were measured. Three Sand
68 Particle Counters (SPCs) (Mikami et al., 2005) were used to measure the sand fluxes in the size range of 39 - 654 μm in 32
69 bins at 0.05, 0.1 and 0.3 m above ground at a sampling rate of 1 Hz. Using the sand fluxes, q_j ($j = 1, 32$), the PSD of saltation
70 particles (saltation PSD) is estimated for a particle size bin at d_j with bin size Δd_j as

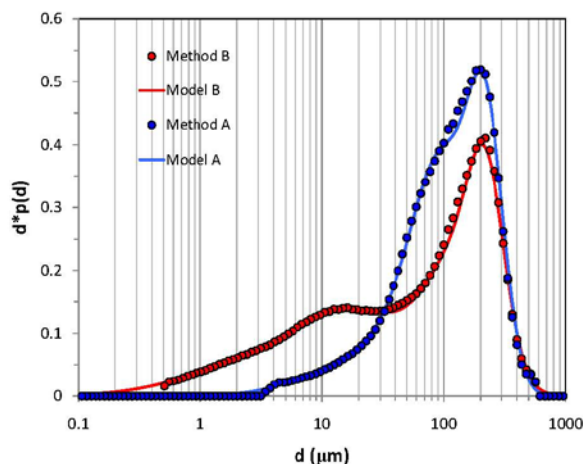
$$71 \quad p(d_j)\Delta d_j = q_j / \sum_{j=1}^{j=32} q_j \quad (2)$$

72 Dust concentration was measured using Optical Particle Counters (OPC) for 8 size groups: 0.3 – 0.6, 0.6 – 0.9, 0.9 – 1.4,
73 1.4 – 2.0, 2.0 – 3.5, 3.5 – 5.9, 5.9 – 8.4 and > 8.4 μm at 1, 2 and 3.5m above ground. The data for the > 8.4 μm bin were
74 excluded from analysis, as the upper size limit was not defined. Dust PSD is estimated as

$$75 \quad p(d_j)\Delta d_j = c_j / \sum_{j=1}^{j=7} c_j \quad (3)$$

76 where c_j denotes the dust concentration for size bin j .

77 Atmospheric variables, including wind speed, air temperature and humidity at various levels, radiation and precipitation
78 were measured using an automatic weather station. These quantities were sampled at 5-second intervals and their 1-minute
79 averages were recorded. Two anemometers mounted at 0.53 and 2.16m measured wind speed. From the atmospheric data,
80 the Obukhov length, L , sensible heat flux, H , and friction velocity, u_* , were derived. Also measured were soil temperature
81 and moisture.

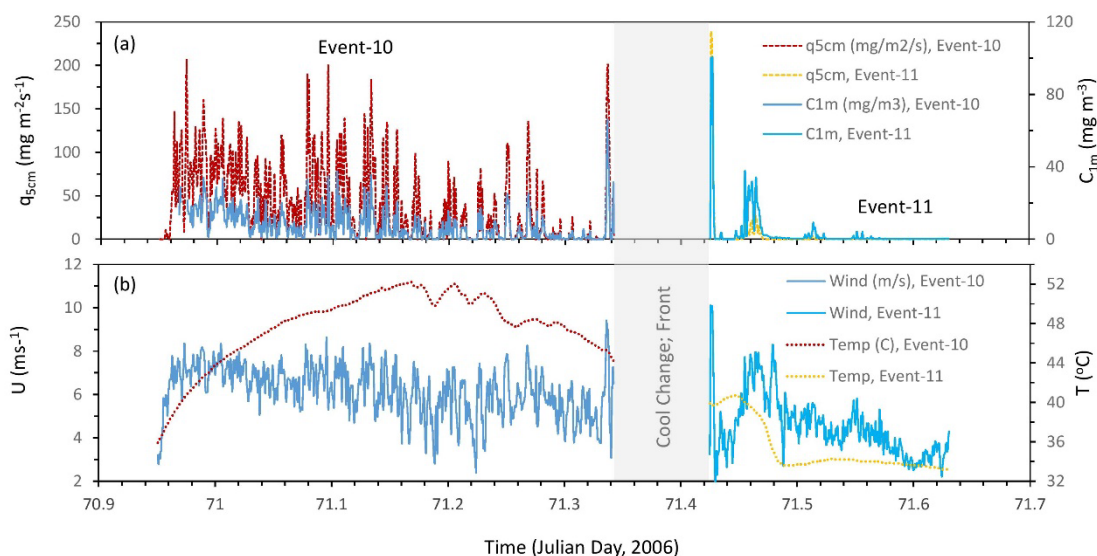


82
83 **Figure 1. Soil particle-size distribution obtained using Method A and Method B, together with the respective approximations**
84 **(Model A and Model B).**



85 Surface soil samples were taken and soil PSD was analysed in laboratory using Method A and B with a particle size
86 analyzer (Microtrac MT3300EX, Nikkiso). In Method A, water was used for sample dispersion with no ultrasonic action. In
87 Method B, sodium hexametaphosphate (HMP) 0.2% solution was used for sample dispersion and 1-minute ultrasonic action
88 of 40W was applied. The soil is loamy-sand based Method A, and sandy loam based on Method B. Figure 1 shows $p_A(d)$
89 (soil PSD from Method A) and $p_B(d)$ (soil PSD from Method B) and the corresponding approximations: p_A shows a larger
90 fraction of particles in the range of 30~300 μm , while p_B a larger fraction of particles in the range of 0.1~30 μm .

91 During JADE, 12 aeolian events were recorded. We select Event-10 (09:49~19:13 12 Mar 2006; Julian Day
92 70.9506940~71.3423611) and Event-11 (21:12 12 Mar ~ 02:08 13 Mar 2006, Julian Day 71.4250000~71.6305600) for the
93 analysis, because Event-10 occurred under daytime unstable, while Event-11 under night-time stable, conditions. Figure 2
94 shows the one-minute averages of wind speed at 0.53m, U , air temperature at 0.66m, T , saltation flux at 0.05m, $q_{5\text{cm}}$ and total
95 dust concentration at 1m, $C_{1\text{m}}$. Event-10 occurred on a hot day prior to a cool change (cold front causing temperature drop
96 but no rainfall), with near surface air temperature reaching 52°C and wind speed ~8ms⁻¹. Event-10 lasted ~10 hours. The cool
97 change occurred at ~19:00-21:00 13 Mar 2006 local time. The strong winds (probably also strong sand drift and dust
98 emission) accompanying the cool change caused the shutdown of the instruments and thus, unfortunately, this period was not
99 fully recorded. Event-11 occurred after the cool change in the nighttime of 12/13 Mar 2006, during which T was dropping
100 from ~40°C to ~33°C and U from ~8 ms⁻¹ to ~5ms⁻¹. Event-11 was much weaker than Event-10.

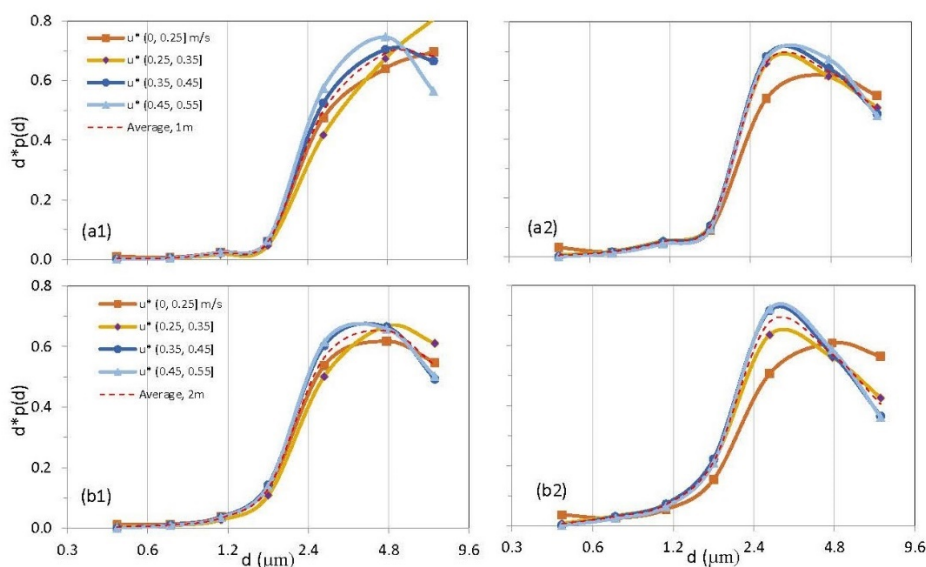


101
102 **Figure 2.** (a) one-minute averaged saltation flux at 0.05m, $q_{5\text{cm}}$, and dust concentration at 1m, $C_{1\text{m}}$, for Event-10 and Event-11; (b)
103 as (a) but for wind speed at 0.53m above ground, U , and air temperature at 0.66m, T . The cool change is marked.



104 3 Results

105 To examine dust-PSD dependency on friction velocity, we use u_* to denote the one-minute values of friction velocity, $p(u_*)$
 106 its probability density function (PDF), \bar{u}_* its mean and σ_{u_*} its standard deviation. The u_* values are divided into the
 107 categories of 0~0.25, 0.25~0.35, 0.35~0.45 and 0.45~0.55 ms^{-1} , and the corresponding dust PSDs and saltation PSDs are
 108 sorted accordingly. These PSDs are then composite averaged for the u_* categories. As Figure 3 shows that for both Event-10
 109 and -11, at both 1m and 2m height, as u_* increases, the mode of dust PSD shifts to finer particles. For Event-10, the most
 110 obvious shift occurs between the u_* categories 0~0.25 ms^{-1} and 0.25~0.35 ms^{-1} , while the shift between the 0.35~0.45 ms^{-1}
 111 and 0.45~0.55 ms^{-1} is less pronounced. The results shown in Figure 3 are consistent with the findings of Alfaro et al. (1997)
 112 and show that dust PSD is u_* dependent.



113

114 **Figure 3. Dust PSD for different u_* categories for Event-11 at levels 1m and 2m (a1 and b1), and for Event-10 (a2 and b2). Also**
 115 **shown are the PSDs averaged over all u_* values (red dashed line).**

116 Figure 3 shows also that the dust PSDs for Event-10 and -11 considerably differ. As said, Event-10 occurred under
 117 unstable, while Event-11 under stable conditions. Several quantities can be used as measure of ABL stability, but the one
 118 used in this study is the convective scaling velocity defined as

119

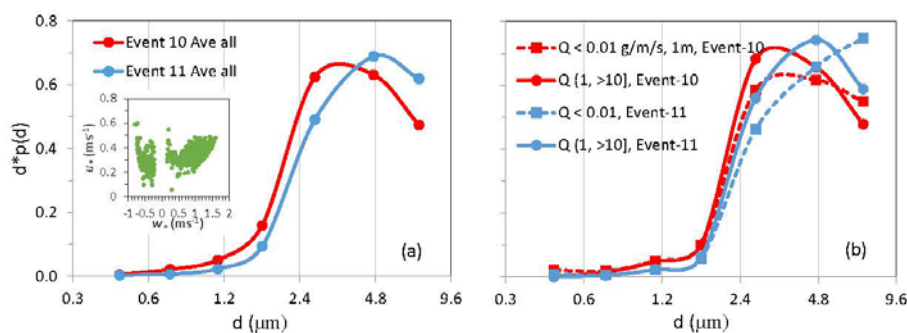
$$w_* = \left(\frac{g}{\bar{\theta}} H_0 z_l \right)^{\frac{1}{3}} \quad (4)$$

120 where $g/\bar{\theta}$ is the buoyancy parameter with g being the acceleration due to gravity and $\bar{\theta}$ the mean potential temperature; H_0 is
 121 surface kinematic heat flux (Kms^{-1}) and z_l a scaling length (set to the capping inversion height for convective ABL and 100m



122 for stable ABL). For unstable conditions, w_* is positive while for stable conditions negative. The reason for choosing w_* is
123 that it is a scaling parameter for the strength of turbulence in unstable ABL. Usually, w_* is not used for stable ABLs, but used
124 here as an indicator for the suppression of turbulence by negative buoyancy.

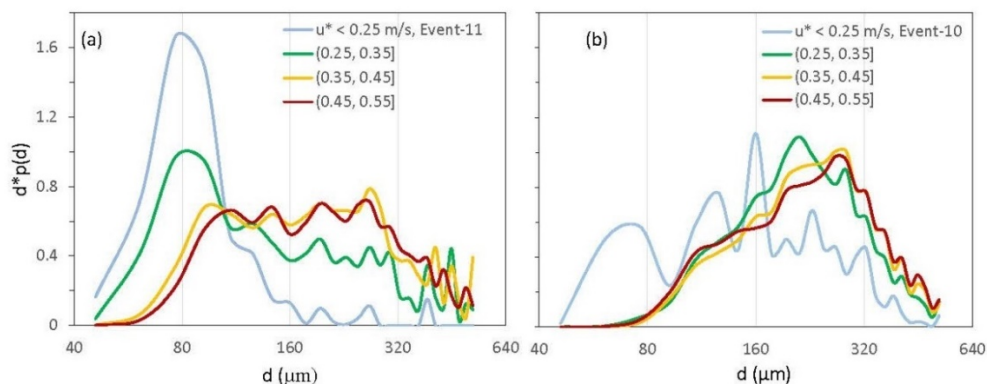
125 Figure 4a shows the dust PSD averaged over three (1, 2 and 3.5m) heights and all u_* values for Event-10 and -11. The
126 insert shows a scatter plot of u_* against w_* for Event-10 (right half) and -11 (left half). For Event-10, the mean and standard
127 deviation of u_* and w_* were respectively (0.36, 0.057) and (1.03, 0.29), all in ms^{-1} , and for Event-11 (0.28, 0.077) and (-0.41,
128 0.159). From Event-10 to -11, the dust PSD mode shifted from $3\mu\text{m}$ to $5\mu\text{m}$. During Event-10, a substantially higher fraction
129 of particles in the size range of $0.4 \sim 4\mu\text{m}$ was emitted. To further examine how dust PSD depends on saltation intensity, we
130 average the dust PSDs for different Q categories. Examples of the dust PSDs for $Q < 0.01 \text{ gm}^{-1}\text{s}^{-1}$ (weak saltation) and Q
131 (1, >10) $\text{gm}^{-1}\text{s}^{-1}$ (moderate to strong saltation) are shown in Figure 4b. Again, weak saltation corresponded to coarser dust
132 particles and strong saltation to finer dust particles, i.e., in Event-10 finer particles are emitted than in Event-11, a result that
133 can also be seen in Figure 3. However, the composite analysis of dust PSDs for the different Q categories shows that the
134 dust PSD dependency on w_* persisted (Figure 4b). The results shown in Figure 4a and 4b are consistent with those of
135 Khalfallah et al. (2020).



136

137 **Figure 4. (a) Dust PSD averaged over all u_* values and over the three levels 1, 2 and 3.5m for Event-10 and -11. The insert is a**
138 **scatter plot of u_* against w_* . (b) Dust PSDs averaged for the saltation flux categories $Q < 0.01 \text{ gm}^{-1}\text{s}^{-1}$ and $Q (1, > 10] \text{ gm}^{-1}\text{s}^{-1}$ for**
139 **Event-10 and -11.**

140 The reason for the dependency of dust PSD on u_* has been explained in Gillette et al. (1974), Gillette (1981), Shao et al.
141 (1993), Alfaro et al. (1997) and Shao (2001), because u_* is a descriptor of saltation bombardment intensity. But how is the
142 dependency of dust PSD on ABL stability, here w_* , explained? Khalfallah et al. (2020) attributed this to the different
143 diffusion of particles of different sizes in stable and unstable conditions. This interpretation does not seem to apply to the
144 JADE data, since the dust PSDs at 1, 2 and 3.5m levels do not substantially differ, and the dust particles considered here are
145 in a small size range ($0.38 \sim 8.3\mu\text{m}$) such that their diffusivities should be all almost identical to the eddy diffusivity.



146

147 **Figure 5. (a) Saltation PSD averaged for four different u_* categories for Event-11; (b) as (a), but for Event-10.**

148 The most conspicuous reason is the enhanced saltation bombardment in unstable conditions. Several observations can be
 149 made from the saltation PSD for Event-10 and -11 shown in Figure 5. First, for $u_* \leq 0.25 \text{ ms}^{-1}$ in Event-11, saltation PSD was
 150 confined to a narrow size range centered at $70\sim 80 \mu\text{m}$ where u_{*t} is minimum. This indicates that saltation
 151 splash/bombardment was weak to mobilize particles in other size ranges. In contrast, for $u_* \leq 0.25 \text{ ms}^{-1}$ in Event-10, saltation
 152 PSD covered a broader size range, implying that saltation splash was strong to entrain particles of other sizes. Second, for
 153 both Event-10 and -11, the peak values of saltation PSD were shifted to larger particles for larger u_* : for Event-10 the peak
 154 for $u_* = 0.35 \text{ ms}^{-1}$ was at $203.3 \mu\text{m}$, while for $u_* = 0.55 \text{ ms}^{-1}$ at $257.8 \mu\text{m}$. Clearly, since u_{*t} is particle size dependent, saltation
 155 PSD is a selective sample of the soil PSD by wind. Third, the saltation PSDs for given u_* categories (e.g., $0.35 < u_* \leq$
 156 0.45 ms^{-1} , Figure 5a and 5b) differed significantly between Event-10 and -11 as a consequence of ABL stability. In Event-11
 157 (Figure 5a), saltation was not fully developed, as the saltation PSD plateau in the size range $100\sim 300 \mu\text{m}$ suggests, implying
 158 again that saltation splash/bombardment was not efficient. In Event-10 (Figure 5b), saltation was more fully developed.

159 Based on Figures 3, 4 and 5, we conclude that dust PSD is not only u_* but also w_* dependent. We argue that the
 160 dependency on w_* can be attributed to saltation bombardment intensity from three perspectives. First, it is known from the
 161 ABL similarity theory that,

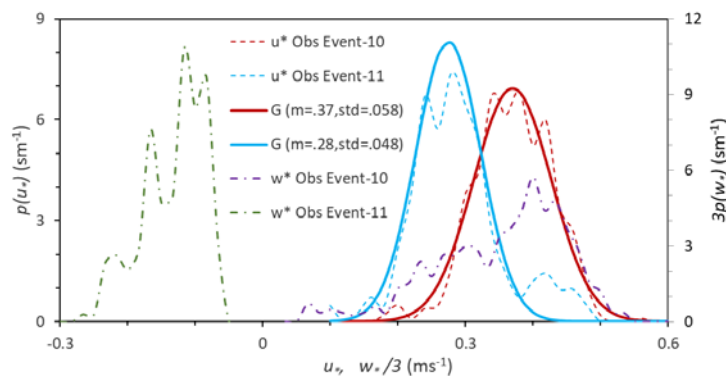
162
$$\bar{u}_* = \frac{\kappa z}{\phi_m} \frac{\partial \bar{u}}{\partial z} \quad (5)$$

163 where κ is the von Karman constant, z height and ϕ_m a similarity function (Stull, 1988):

164
$$\phi_m = \begin{cases} 1 + \beta_m \zeta & \zeta > 0 \text{ stable} \\ (1 - \gamma_m \zeta)^{-1/4} & \zeta < 0 \text{ unstable} \\ 1 & \zeta = 0 \text{ neutral} \end{cases} \quad (6)$$



165 where $\zeta = z/L$ (L is Obukhov length) and $\beta_m = 5$ and $\gamma_m = 16$ are empirical coefficients (Businger et al., 1971). For stable
 166 conditions, $\phi_m > 1$ and for unstable conditions $\phi_m < 1$. Figure 6 shows the PDFs of u_* and w_* for Event-10 and -11,
 167 together with the approximations for the PDFs of u_* . For Event-10, $\bar{u}_* = 0.37\text{ms}^{-1}$, while for Event-11, $\bar{u}_* = 0.28\text{ms}^{-1}$. The
 168 larger \bar{u}_* was partly responsible for the stronger saltation and dust emission during Event-10 than during Event-11.



169

170 **Figure 6. The probability density functions of u_* and w_* , $p(u_*)$ and $p(w_*)$, respectively, for Event-10 and -11, together with the**
 171 **Gaussian approximations for the $p(u_*)$ functions. The mean values (m) and standard deviations (std) for the Gaussian (G)**
 172 **distributions are given. Note that for $p(w_*)$, $3p(w_*)$ against $w_*/3$ is plotted to conveniently present the information in the same graph.**

173 Second, as Figure 6 shows, u_* is a stochastic variable. Li et al. (2020) suggested that $\tau = \rho u_*^2$ in neutral conditions is Gauss
 174 distributed. Klose et al. (2014) reported that τ in unstable conditions is Weibull distributed. The exact form of $p(\tau)$ requires
 175 further investigation, but the JADE data of u_* show that $p(u_*)$ is reasonably Gaussian. Hence,

176

$$p(\tau) = \frac{1}{2\rho u_*} p(u_*) \quad (7)$$

177 is skewed to smaller τ . Figure 6 shows that u_* in Event-10 not only had a larger mean value but also a larger variance than in
 178 Event-11. We emphasize that the variance of u_* strongly affects saltation, because saltation flux depends non-linearly on u_* .
 179 To illustrate this, we consider u_{*1} and u_{*2} , and assume that

- 180 • u_{*1} and u_{*2} are Gaussian distributed and have the same mean that equals u_{*t} (say 0.2ms^{-1})
- 181 • u_{*1} and u_{*2} have respectively standard deviation, σ_1 and σ_2 , with $\sigma_2 = \eta \sigma_1$ and $\eta > 1$; and
- 182 • Q satisfies the Owen's model (Owen, 1964),

183

$$Q_i = c u_{*i}^3 \left(1 - \frac{u_{*t}^2}{u_{*i}^2} \right) \quad \text{for } u_* > u_{*t};$$

184

$$\text{otherwise } 0; \quad \text{with } i = 1, 2 \quad (8)$$



185 where c is a dimensional constant. It follows that the ratio of the mean values of Q_2 and Q_1 is

$$186 \quad \eta_Q = \frac{\bar{Q}_2}{\bar{Q}_1} = \frac{\int_{u_{*t}}^{\infty} Q_2 p(u_{*2}) du_{*2}}{\int_{u_{*t}}^{\infty} Q_1 p(u_{*1}) du_{*1}} \quad (9)$$

187 Equation (9) can be evaluated numerically for different η (Table 1) and is approximately

$$188 \quad \eta_Q = 0.607 \eta^2 - 0.0028\eta + 0.4283 \quad (10)$$

189 This shows that $p(u_*)$ profoundly influences the magnitude of Q . For fixed \bar{u}_* , a large u_* variance corresponds to a larger \bar{Q} .

190 **Table1. Streamwise saltation flux ratios, η_Q , for different u_* std ratios, η (see text for details).**

η	1.2	1.4	1.6	1.8	2	3	4
η_Q	1.30	1.63	2.00	2.41	2.86	5.83	10.15

191

192 Third, in unstable conditions, turbulence is stronger due to buoyancy production, which leads to increased saltation
 193 bombardment intensity. We do not have independent evidence to verify this, but to illustrate the point, we use a two-
 194 dimensional (2-d, x_1 in mean wind direction and $x_3 \equiv z$ in vertical direction) saltation model (Supplement A) to simulate the
 195 impact kinetic energy of saltation sand grains. For given u_* and roughness length, z_0 , a 2-d turbulent flow is generated with
 196 the mean wind assumed to be logarithmic $\kappa \bar{u}_1 = \bar{u}_* \ln(z/z_0)$ and the velocity standard deviations satisfy

$$197 \quad \frac{\sigma_{u1}}{\bar{u}_*} = a \cdot \ln\left(\frac{z}{z_0}\right) \quad (11)$$

$$198 \quad \frac{\sigma_{u3}}{\bar{u}_*} = f_{u3}(\zeta) \cdot a \cdot \ln\left(\frac{z}{z_0}\right) \quad (12)$$

199 and the dissipation rate for turbulent kinetic energy, ε , satisfies

$$200 \quad \varepsilon \frac{\kappa z}{\bar{u}_*^3} = f_\varepsilon(\zeta) \quad (13)$$

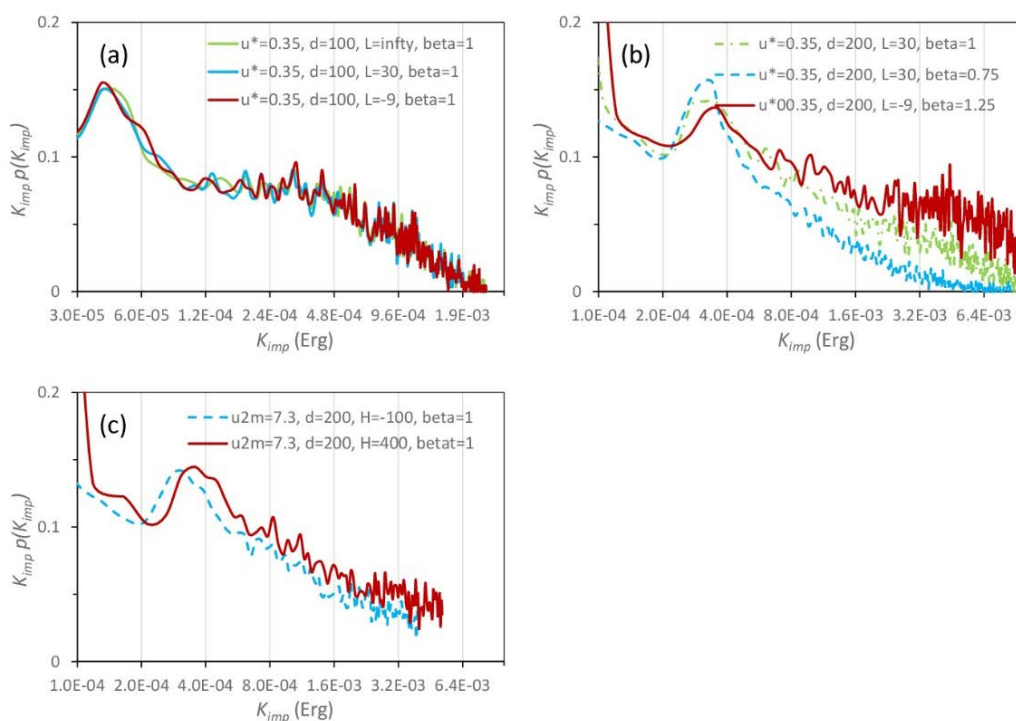
201 The similarity relationships $f_{u3}(\zeta)$ and $f_\varepsilon(\zeta)$ follow Kaimal and Finnigan (p16, 1995). As saltation takes place in the layer
 202 close to the surface, the vertical profiles of σ_{u1} and σ_{u3} are considered following Yahaya et al. (2003). The coefficient a
 203 ($=1.16\beta$) is varied by setting β to 0.75, 1.00 and 1.25 for weak, normal and strong turbulence, respectively.

204 In each numerical experiment, 20000 sand grains of identical size are released from the surface and their trajectories are
 205 computed. At impact on the surface, the particles rebound with a probability of 0.95 and a rebounding kinetic energy, K_{reb} ,
 206 0.5 times the impact kinetic energy, K_{imp} . The rebound angle is Gauss distributed with a mean of 40° and standard deviation
 207 5° . Splash entrainment is neglected. The PDF of K_{imp} , $p(K_{imp})$, is used as a measure bombardment intensity.



208 Many numerical experiments are carried out, but for our purpose, we show only the results of the ones listed in Table 2.
 209 The initial velocity components of sand grains (V_{1o} , V_{3o}) are generated stochastically. V_{1o} is Gauss distributed with a mean
 210 $\bar{V}_{1o} = \bar{u}_* \cos(55^\circ)$ and standard deviation, $\sigma_{V_{1o}} = 0.1\bar{u}_*$. V_{3o} is Weibull distributed with a shape parameter $A = 2$ and a scale
 211 parameter $B' = \bar{u}_* \sin(55^\circ) / \Gamma(1 + 1/A)$ where Γ is a Gamma function. To account for the influence of stability on V_{3o} , B'
 212 is modified such that the adjustment to $\sigma_{V_{3o}}$ is the same as that to $\sigma_{u_3}(10z_0)$, i.e., the modified scale parameter, B , is given
 213 by

$$214 \quad B = \beta f_{u_3} \left(\frac{30z_0}{L} \right) B' \quad (14)$$



215
 216 **Figure 7. Probability density function $p(K_{imp})$ (plotted in $K_{imp} p(K_{imp})$ against K_{imp} in logarithmic scale) for the numerical**
 217 **experiments. In (a), $p(K_{imp})$ is shown for $u_* = 0.35\text{ms}^{-1}$, $d = 100\mu\text{m}$ and $\beta = 1$ but for three different Obukhov lengths $L = \infty$, 30m and**
 218 **-9m. In (b), the effect of β on $p(K_{imp})$ is examined; and in (c) the effect of stability on $p(K_{imp})$ with given mean wind speed at $z = 2\text{m}$**
 219 **is examined.**

220 Figure 7a compares $p(K_{imp})$ for Exp1a, 1b and 1c and shows that $p(K_{imp})$ for these cases is very similar. The small
 221 differences in $p(K_{imp})$ between the cases suggest that the differences in particle trajectory arising from the stability
 222 modification to turbulence profile, with u_* fixed, are negligible. However, a small change in β , as Figure 7b shows for Exp2a,
 223 2b and 2c, can lead to significant changes in $p(K_{imp})$ with larger β corresponding to higher probability of larger K_{imp} , namely,



224 high saltation bombardment intensity. In Exp3a and 3b, u_{2m} (mean wind 2m height) is set to 7.3ms^{-1} and the surface sensible
225 heat flux, H , to -100 and 400Wm^{-2} . Figure 7c shows that $p(K_{imp})$ differs significantly with larger K_{imp} in unstable conditions.

226 **Table 2: Numerical experiments for saltation bombardment intensity. For all experiments, $z_0 = 0.48\text{mm}$, $C_0 = 5$, $C_1 = 2$ and**
227 **$\rho_p = 2650\text{kgm}^{-3}$.**

Exp	u_* (ms^{-1})	L (m)	d (μm)	β
Exp1a, 1b, 1c	0.35	$\infty, 30, -9$	100	1.0
Exp2a, 2b	0.35	30	200	0.75, 1
Exp2c	0.35	-9	200	1.25
Exp3a, 3b	$u_{2m}=7.3$	$H=-100; 400\text{Wm}^{-2}$	200	1

228

229 The experiments reveal that the PDF of particle initial velocity influences strongly saltation bombardment intensity, and
230 saltation in in unstable ABL intensifies saltation bombardment to cause finer dust-particle emission.

231 The numerical experiments reveal that the PDF of the particle initial velocity influence significantly on saltation
232 bombardment intensity, as the saltation particles in unstable ABL impact the surface with larger kinetic energy than in stable
233 ABL. This is also the result seen in Figure 5, i.e., saltation in Event-10 was more fully developed than in Event-11. The more
234 fully developed saltation in unstable ABL increases saltation bombardment intensity and hence the release of finer dust
235 particles, seen in Figure 4.

236 4 Conclusions

237 Using JADE data, we showed that dust PSD is dependent on friction velocity u_* . This finding is consistent with the wind-
238 tunnel study of Alfaro et al. (1997). The JADE data support the claim that dust PSD is saltation-bombardment dependent and
239 do not support the hypothesis that dust PSD is invariant.

240 The JADE data show that dust PSD, as well as saltation PSD, also depends on ABL stability. This finding is consistent
241 with the results of Khalfallah et al. (2020). Dust PSD is dependent on ABL stability for three reasons. First, under similar
242 mean wind conditions, the mean surface shear stress is larger in unstable than in stable conditions. Second, u_* is a stochastic
243 variable and the PDF of u_* profoundly influences the magnitude of saltation flux, Q , because of the non-linear relationship
244 between Q and u_* . With fixed u_* mean, a larger u_* variance corresponds to a larger Q . Unstable ABL has in general larger u_*
245 variances which generate stronger saltation bombardment and produce the emission of finer dust particles. Third, in unstable
246 ABL, turbulence is generally stronger and in strong turbulent flows, the proportion of saltation particles with large impacting
247 kinetic energy is larger than in weak turbulent flows. Consequently, saltation in unstable ABLs is more fully developed and
248 saltation bombardment has higher intensity.



249 The dependencies of dust PSD on u_* and ABL stability are ultimately attributed to the statistic behavior of u_* , i.e., its PDF
250 $p(u_*)$, or more simply its mean and variance. These dependencies point to the same fact that, for a given soil, saltation
251 bombardment plays the determining role for the dust PSD. Stronger saltation causes the emission of finer dust.

252
253 *Data availability.* Data can be accessed by contacting the corresponding authors.
254

255 *Author contributions.* Yaping Shao performed the data analyses and wrote the manuscript. Jie Zhang and Ning Huang
256 contributed to the conception of the study and helped perform the analysis with constructive discussions. Masahide Ishizuka,
257 Masao Mikami and John Leys Conceived, designed and performed the experiments.

258
259 *Competing interests.* The authors declare that they have no conflict of interest.
260

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