

Referee comments on the revised version or on the author responses are in green type.

Referee comments on the original version are in black type.

Author responses are in blue type.

The revised manuscript includes a qualitative comparison with contemporaneous HNO<sub>3</sub> measurements by MLS. This analysis should have been taken further to provide a more useful quantitative comparison with the IASI HNO<sub>3</sub> column data. Additionally, incorporating CALIOP data would have enabled much better insight into the interannual variations of the drop temperatures and their relation to the distribution of PSC types. Overall, the revised manuscript does not put forward a compelling case for the scientific utility of the “drop temperatures”. I do not find that the authors have addressed previously raised concerns in the on-line referee comments sufficiently well to recommend publication. Detailed comments are given below.

Line numbers from here correspond to the revised version ...

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19 good vertical sensitivity in the mid-stratosphere (around 50 hPa),

The IASI HNO<sub>3</sub> nadir measurements can not be considered as having “good vertical sensitivity”. Prospective data users need to know the vertical resolution of the measurements and that is conveyed by the standard practice of quoting the full width at half maximum (FWHM) of the averaging kernel. There is no reason not to do this for a nadir sounder e.g. Maddy and Barnett, Vertical Resolution Estimates in Version 5 of AIRS Operational Retrievals, IEEE Trans. Geosci. Remote Sens., 46, 8, 2375-2384, 2008.

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29 nitric acid trihydrate (NAT) formation temperature

For reasons explained elsewhere the “formation temperature” can be better expressed as an “existence threshold”.

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26 Although the measured

27 HNO<sub>3</sub> total column does not allow differentiating the uptake of HNO<sub>3</sub> by different types of PSC particles  
28 along the vertical profile, an average drop temperature of  $\sim 194.2 \pm 3.8$  K, consistent with the nitric acid  
29 trihydrate (NAT) formation temperature (close to 195 K at 50 hPa)

The averaged “drop temperature” disregards the considerable interannual variability in the early stage formation of different types of Antarctic PSCs and the role played by the exposure of liquid PSCs to low temperatures in the formation of NAT i.e. many studies have shown that NAT is not uniquely constrained to nucleate at TNAT and some supersaturation is generally needed leading to a lower temperature for NAT formation (as in fact you discuss in the text L55-L75). Therefore, stating that the drop temperature is “consistent with TNAT”, which implies that PSCs are mainly NAT forming at TNAT, is invalid. On line 28 the “ ” sign should be deleted since a specific value and its uncertainty is quoted.

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30 potential vorticity lower than  $-10 \times 10^{-5}$  K.m<sup>2</sup>.kg<sup>-1</sup>.s<sup>-1</sup>

Some corresponding indication of the equivalent latitude range would be useful here.

30 vorticity lower than  $-10 \times 10^{-5} \text{ K.m}^2.\text{kg}^{-1}.\text{s}^{-1}$ . The spatial distribution and inter-annual variability of the  
31 drop temperature are investigated and discussed in the context of previous PSCs studies.

However, the study presented here does not include any observed data on PSCs and is therefore not a “PSC study”.

92 profile of  $\text{HNO}_3$  similar to the limb sounders, IASI provides reliable total column measurements of  
93  $\text{HNO}_3$  characterized by a maximum sensitivity in the low-middle stratosphere around 50 hPa (20 km)  
94 during the dark Antarctic winter (Ronsmans et al., 2016, 2018) where the PSCs cloud form (Voigt et al.,

Please give the fullwidth half-max (FWHM) of the vertical response in km and not just the height of maximum sensitivity.

126 In order to expand on the comparisons against FTIR measurements which  
127 is impossible during the polar night, Figure 1 (top panel) presents the time series of daily IASI total  
128  $\text{HNO}_3$  columns co-located with MLS VMR measurements within  $2.5 \times 5$  grid boxes at three pressure  
129 levels (at 30, 50 and 70 hPa), averaged in the 70 - 90 S equivalent latitude band. Similar variations in  
130  $\text{HNO}_3$  are captured by the two instruments with an excellent agreement for the timing of the strong  
131  $\text{HNO}_3$  depletion within the inner vortex core. IASI  $\text{HNO}_3$  variations generally match well those of MLS  
132  $\text{HNO}_3$  in each latitude band (see Figure 1 bottom panel for the 50 - 70 S equivalent latitude band; the  
133 other bands are not shown here).

251 The 50 hPa drop temperatures are detected between 189.2 K and 202.8 K, with an average of  
252  $194.2 \pm 3.8 \text{ K}$  (1 standard deviation) over the ten years. Knowing that TNAT can be higher or lower  
253 depending on the atmospheric conditions and that NAT starts to nucleate from 2–4 K below TNAT (Pitts  
254 et al., 2011; Hoyle et al., 2013; Lambert et al., 2016), the results here demonstrate the consistency  
255 between the 50 hPa drop temperature,

The software bug that was fixed in the revised version has changed the drop temperatures such that the year with 202.8K (previously 190.6K) is a significant outlier since it lies 8K higher than the 10-year mean drop temperature and is almost as much above the assumed 50hPa TNAT (195K). Therefore it does not support the statement on L365-366 that the 10-year range “demonstrated the good consistency between the 50 hPa drop temperature and the PSCs formation temperatures in that altitude region”.

295 Note that the high extremes in the drop temperature, which are found in some case above  
295 eastern Antarctica, should be considered with caution: they correspond to specific regions above ice  
296 surface with emissivity features that are known to yield errors in the IASI retrievals (Hurtmans et al.,  
297 2012; Ronsmans et al., 2016). Indeed, bright land surface such as ice might in some cases lead to poor  
298  $\text{HNO}_3$  retrievals. Although wavenumber-dependent surface emissivity atlases are used in FORLI  
299 (Hurtmans et al., 2012), this parameter remains critical and causes poorer retrievals that, in some  
300 instances, pass through the series of quality filters and could affect the drop temperature calculation.

...  
371 Except for extreme drop  
372 temperatures ( $\sim 210 \text{ K}$ ) that were found from year to year above eastern Antarctica and suspected to  
373 result from unfiltered poor quality retrievals in case of emissivity issues above ice

It is not clear why this data quality problem has not been addressed in the revised submission. Measurements that are known to be bad must be screened out.

=====  
332 Except above some parts of Antarctica which are prone to larger retrieval errors, the overall range in the  
333 drop 50 hPa temperature for total HNO<sub>3</sub> inside the isocontour for the averaged temperature of 195 K,  
334 typically extends from ~187 K to ~195 K, which falls within the range of PSCs nucleation temperature  
335 at 50 hPa: from slightly below TNAT to around 3-4 K below the ice frost point - Tice - depending on  
336 atmospheric conditions, on TTE and on the type of formation mechanisms (Pitts et al., 2011; Peter and  
337 Grooß, 2012; Hoyle et al., 2013).

...  
373 the range of drop  
374 temperatures is interestingly found in line with the PSCs nucleation temperature

Why is the discussion in L302-338 and L367-376 limited to nucleation of NAT and ice PSCs with no mention of STS? There is no nucleation barrier to STS formation and it generally forms in advance of ice nucleation except possibly under very fast cooling e.g. in mountain waves. STS is not even mentioned in the paper after the introduction in L55-72.

Line numbers from here correspond to the original version ...

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429 Major comments

430  
431 [The description of the polar HNO<sub>3</sub> variation presented in the paper is already well known from  
432 numerous other studies.]

433 The purpose of this paper is to demonstrate the interest of IASI for HNO<sub>3</sub> stratospheric studies  
434 (Ronsmans et al., 2018) after having undergone a rigorous validation exercise (Ronsmans et al., 2016).  
435 If limb measurements allows resolving the HNO<sub>3</sub> profile in the stratosphere, the potential of IASI lies in  
436 its exceptional spatial and temporal sampling. We demonstrate here that despite its limited vertical  
437 resolution forcing us to consider one total column, the information content that actually lies in the low  
438 and middle stratosphere offers potential to expand on previous polar stratospheric denitrification studies,  
439 usually performed using limb sounder measurements, and to continue the long-term records of HNO<sub>3</sub>  
440 started with the latter. We have tried in this paper not to repeat too much of our earlier work but some  
441 duplication was unavoidable; in particular, with respect to vertical sensitivity and errors (these are two  
442 aspects that referee1 finds in fact insufficiently described here).

=====  
444 [The lack of vertical resolution in the IASI HNO<sub>3</sub> measurements severely limits the interpretation of the  
445 results and precludes differentiation between denitrification and renitrification e.g. consider the effect of  
446 the vertical integration through depleted higher layers overlaying lower enhanced layers.]

447 We understand that the referee sees this as a limitation. However, despite the lack of vertical resolution,  
448 which is recognized in the paper and which forces us to consider total HNO<sub>3</sub> columns, IASI is  
449 characterized by a good sensitivity to HNO<sub>3</sub> at specific levels, in particular, in the range between ~70  
450 hPa to ~30 hPa in the southernmost latitude in winter and as such it provides an adequate means to

451 investigate the stratospheric processes in the polar nights.

452

453 In order to justify this further, we would like to refer to the figure 3 (top and bottom panels) of Ronsmans  
454 et al. (2016) that presents global distributions of the degrees of freedom for signal (DOFS, top panels)  
455 and of the altitude of maximum sensitivity of IASI to the HNO<sub>3</sub> profile (bottom panel), separately for  
456 January (left) and July (right) 2011, when the strong HNO<sub>3</sub> depletion occurs within the cold Antarctic  
457 winter. It shows clearly that the altitude of maximum sensitivity of the total columns is invariant at  
458 equatorial and tropical latitudes, whereas it varies with seasons at middle and polar latitudes. Above the  
459 Antarctic, the altitude of maximum sensitivity varies between ~9 km in summer and ~22 km in winter.  
460 The variations of the altitude of maximum sensitivity follow the altitude variations of maximum HNO<sub>3</sub>  
461 concentrations.

462

463 We agree that the IASI sensitivity was insufficiently put forward in the text. We made it more explicit  
464 at several places in the revised manuscript; e.g. in Section 1: "IASI provides reliable total column  
465 measurements of HNO<sub>3</sub> characterized by a maximum sensitivity in the low-middle stratosphere around  
466 50 hPa (20 km) during the dark Antarctic winter (Ronsmans et al., 2016; 2018) . . ." and in Section 2:  
467 ". . . the largest sensitivity of IASI in the region of interest, i.e. in the low and mid-stratosphere (from 70  
468 to 30 hPa), where the HNO<sub>3</sub> abundance is the highest (Ronsmans et al., 2016)."

The response does not address the specific case example of where IASI views HNO<sub>3</sub> depleted higher layers that overlay lower enhanced layers. How does the IASI column HNO<sub>3</sub> measurement change if the HNO<sub>3</sub> is redistributed in the vertical coordinate by denitrification and re-nitrification? A further question would be how does downwelling of higher values of HNO<sub>3</sub> affect the HNO<sub>3</sub> column?

=====  
471 [Although the IASI HNO<sub>3</sub> data has much better 2D horizontal resolution than any other measurement  
472 this has not been developed as a tool to provide information beyond that of satellite instruments that  
473 measure only along the orbit track.]

474 We do not fully agree. The determination of the drop temperature using the second derivative exploits  
475 the large dataset of daily IASI measurements. Furthermore, the spatial distributions of the drop  
476 temperature calculated at 50hPa, which are presented in the figure 5 of the manuscript, do actually take  
477 advantage of the excellent spatial/temporal resolution of IASI to provide information throughout the  
478 entire vortex and outside. This would probably not be feasible with other types of measurements.

As a further example of the 2D potential, could IASI be used to image the HNO<sub>3</sub> field to show depletion in the cold phases of mountains waves e.g. near the Palmer peninsula (similar to the wave structures seen in AIRS brightness temperatures) or is that defeated by the vertical integration caused by the poor vertical resolution?

=====  
480 [CALIOP PSC information is available for the same time frame, why was this not used? Certainly, PSC  
481 volumes vs time would be helpful in providing the underlying interannual variability of PSC types (NAT,  
482 STS, ice) to compare with the resulting drop temperatures derived from IASI. Similarly, at least some  
483 comparisons of the IASI HNO<sub>3</sub> column with integrated column calculated from Aura MLS are  
484 necessary to establish the validity of the measurements in the most severely depleted inner vortex core.]

485 Thank you for this comment. It is certainly a good idea to use the CALIOP measurements in support but  
486 this goes beyond the goal of this paper, which is to demonstrate the capability of IASI to measure HNO<sub>3</sub>  
487 columns that are relevant for stratospheric studies. Using CALIOP PSC information and, in particular,  
488 comparing the spatial distributions of IASI derived drop temperatures (Figure 5 of the revised paper)  
489 with maps of CALIOP PSC would be very interesting in order to go a step further in the analyses of the  
490 underlying HNO<sub>3</sub> condensation processes, but it will be challenging and add significant complexity  
491 given the high variability in the distribution of PSC types.

492

493 Regarding the second point on a comparison with MLS, we fully agree that this is highly relevant; it was  
494 also a request of referee #1. We provide here below a comparison with observations by MLS in three  
495 equivalent latitude bands (see Figure 1). We would like to point out that we here compare total columns  
496 measured by IASI with VMR measured by MLS at several pressure levels that cover the highest  
497 sensitivity of IASI (at ~50 hPa, ~70 hPa and ~30 hPa for the sake of the comparison). Hence, the  
498 comparison of IASI columns with MLS measurements is mostly qualitative at this stage and differences  
499 are expected for this reason. Note also that we have preferred comparing IASI HNO<sub>3</sub> columns with VMR  
500 measured by MLS at specific levels instead of integrated columns calculated from MLS, given the  
501 difference in the sensitivity profile between IASI and MLS, the non-negligible IASI sensitivity to HNO<sub>3</sub>  
502 in the troposphere where MLS does not measure HNO<sub>3</sub> etc, which makes the integrated columns from  
503 IASI vs MLS not directly comparable. It should be pointed out finally that part of the differences between  
504 IASI and MLS are likely due to the different number of co-located data within the 2.5°deg x 2.5°deg grid cells  
505 considered here for the comparison, with a much larger number of observations for IASI (through the  
506 quality filtering) than for MLS.

507

508 Despite this, the comparison shows similar spatial and seasonal variations between IASI total HNO<sub>3</sub>  
509 columns and MLS VMR between ~70 and 30 hPa in the different latitude bands, in particular, in the  
510 southernmost equivalent latitudes (see top panel). The strong HNO<sub>3</sub> depletion is well captured by both  
511 IASI and MLS measurements with a perfect match for the onset of the depletion. It further supports the  
512 good sensitivity of IASI to HNO<sub>3</sub> in the range of these pressure levels, justifying the methodology used  
513 in this study.

514

515 The cross-comparison with MLS is indeed insightful and gives further credit on the IASI observations  
516 during the polar night. That comparison figure between IASI and MLS has therefore been included in  
517 Section 2 of the revised manuscript and the text was changed to:

519 "In order to expand on the comparisons against FTIR measurements which is impossible during the polar  
520 night, Figure 1 (top panel) presents the time series of daily IASI total HNO<sub>3</sub> columns co-located with  
521 MLS VMR measurements within 2.5x2.5 grid boxes at three pressure levels (at 30, 50 and 70 hPa),  
522 averaged in the 70°deg S–90°deg S equivalent latitude band. Similar variations in HNO<sub>3</sub> are captured by the  
523 instruments with an excellent agreement for the timing of the strong HNO<sub>3</sub> depletion within the inner  
524 vortex core. IASI HNO<sub>3</sub> variations generally match well those of MLS HNO<sub>3</sub> in each latitude band (see  
525 Figure 1 bottom panel for the 50°deg S–70°deg S equivalent latitude band; the other bands are not shown h

"CALIOP measurements ... this goes beyond the goal of this paper, which is to demonstrate the capability of IASI to measure HNO<sub>3</sub> columns that are relevant for stratospheric studies". That goal was largely achieved already by Ronsmans et al (2016) and published in Atmos. Meas. Tech. This paper is under review for Atmos. Chem. Phys. and should relate more to a science investigation rather than a technical description. The comparisons with MLS are a welcome improvement, but unfortunately fall short of the analysis I was



expecting. Surely the tropospheric contribution of HNO<sub>3</sub> to the IASI column is not all that much (you could estimate the effect to confirm). I expected the MLS profile to be integrated with the IASI response function for a more direct comparison. That would facilitate a quantitative interpretation of the differences in the variation of the column data from the two instruments.

=====  
527 [Regarding the sensitivity of the IASI column HNO<sub>3</sub> measurements, I suggest presenting a few examples  
528 of vertical HNO<sub>3</sub> profiles (from a model or data), ranging from non-depleted to extreme depletion with  
529 calculations of the corresponding calculated integrated IASI column. This would help to indicate the  
530 sensitivity of the column measurement to changes in the vertical distribution of HNO<sub>3</sub> ... i.e. generate  
531 profiles of the change in the IASI column HNO<sub>3</sub> wrt the actual change in HNO<sub>3</sub> at a level, j,  
532  $d(\text{column})/d(\text{HNO}_3)_j$ .]

533 This is an example of information reported in earlier work and that we have tried not to repeat extensively  
534 here. To summarize, the validation study of Ronsmans et al. (2016) provides a complete characterization  
535 of the IASI HNO<sub>3</sub> retrievals: it shows example of vertical HNO<sub>3</sub> profiles along with the total retrieval  
536 error, the a priori profiles and associated averaging kernels profiles ( $d(\text{HNO}_3\text{ret})_i/d(\text{HNO}_3\text{true})_j$ ), along  
537 with the total column averaging kernel ( $d(\text{columnret})/d(\text{HNO}_3\text{true})_j$ ) and the sensitivity profile  
538 ( $d(\text{HNO}_3\text{ret})_i/d(\text{columntrue})_j$ ), were already given in Figures 1 and 2 of that study. Note that the averaging  
539 kernel profile describes how the true state changes the estimate at a specific altitude, i.e. how the retrieval  
540 smooths the true profile. The sum of the elements of an averaging kernel characterizing the retrieval at  
541 a specific altitude returns the sensitivity of the retrieval at that altitude, i.e. to which extent the retrieval  
542 at that specific altitude comes from the spectral measurement rather than the a priori, while the sum of  
543 the averaging kernels indicates how the true state at a specific altitude changes the retrieved total column,  
544 i.e. the altitude to which the retrieved total column is mainly sensitive/representative.

545

546 Figure 3 (top and bottom panels) of Ronsmans et al. (2016) further presents the global distributions of  
547 the degrees of freedom for signal (DOFS, top panels) and of the altitude of maximum sensitivity of the  
548 retrieval to the HNO<sub>3</sub> profile (bottom panel), separately for January (left) and July (right) 2011, when  
549 the strong HNO<sub>3</sub> depletion occurs within the cold Antarctic winter. It clearly shows that above the  
550 Antarctic, the altitude of maximum sensitivity varies between ~9 km in summer and ~22 km in winter  
551 (~ 50 hPa) on average.

552

553 To address the comment of the referee without repeating too much of the earlier results, we have  
554 carefully verified the manuscript with regard to unclear or incomplete statements about vertical  
555 sensitivity. The following has been added in Section 1: "IASI provides reliable total column  
556 measurements of HNO<sub>3</sub> with a maximum sensitivity in the low-middle stratosphere around 50 hPa (20  
557 km) during the dark Antarctic winter (Ronsmans et al., 2016; 2018) ... " and in Section 2: "... the largest  
558 sensitivity of IASI in the region of interest, i.e. in the low and mid-stratosphere (from 70 to 30 hPa),  
559 where the HNO<sub>3</sub> abundance is the highest (Ronsmans et al., 2016).

560

561 In order to convince the referee that IASI measurements capture the expected variations of HNO<sub>3</sub> within  
562 the polar night, we provide in Figure 1 below examples of vertical HNO<sub>3</sub> profiles retrieved within the  
563 dark Antarctic vortex (above Arrival Height) and outside the vortex (above Lauder). The retrieved  
564 profiles are shown along with their associated total retrieval error and averaging kernels (the total column  
565 AvK and the so-called "sensitivity profile" are also represented). Above Arrival Height during the dark  
566 Antarctic winter, we clearly see depleted HNO<sub>3</sub> levels in the low and mid-stratosphere and the altitude

567 of maximum sensitivity at around 30 hPa. At Lauder on the contrary, HNO<sub>3</sub> levels larger than the a priori  
568 are observed in the stratosphere with a larger range of maximum sensitivity.

I also wanted to see specific depleted vs non-depleted cases (one with a re-nitrification layer would be good  
also) generated along with the simulated IASI columns and the calculated columns. I suggest that the figure  
provided on the averaging kernels etc could be added to a supplemental material section with a description  
tailored to the cases studied here in addition to just referring readers to a prior publication.

=====

599 [L10: 191K is also consistent with STS temperatures (192 K) and is actually closer than TNAT (195 K)]

600 Indeed but as stated in the manuscript: "... recent observational and modelling studies have shown that  
601 HNO<sub>3</sub> starts to condense in early PSC season in liquid NAT mixtures well above Tice (~4 K below TNAT,  
602 close to TSTS)...". The NAT nucleation temperature at 50 hPa range from slightly below TNAT to around  
603 3-4 K below Tice, depending on atmospheric conditions, on TTE and on the type of formation  
604 mechanisms (Pitts et al., 2011; Peter and Groos, 2012; Hoyle et al., 2013).

605

606 Note that in replying to referee#1 we have identified a bug for the automatic detection of the drop  
607 temperature, as well as for the detection of the corresponding dates in the figure 2 of the manuscript. It  
608 has been corrected. The position of the drop temperatures does now perfectly match the yearly minima  
609 of the total HNO<sub>3</sub> second derivative. An average drop temperature over the ten years of IASI of 194.2  
610 +/- 3.8 K is now calculated, which is even closer to TNAT.

611

612 Finally, as requested by referee #1, we also now clearly mention in Section 4.1 of the manuscript the  
613 range of drop temperatures when calculated at two other pressure levels to better judge on the uncertainty  
614 of the drop temperature at 50 hPa (see Figure 3 here below):

615 ... Nevertheless, given the range of maximum IASI sensitivity to HNO<sub>3</sub> around 50 hPa, typically  
616 between 70 and 30 hPa (Ronsmans et al., 2016), the drop temperatures are also calculated at these two  
617 other pressure levels (not shown here) to estimate the uncertainty of the calculated drop temperature  
618 defined in this study at 50 hPa. The 30 hPa and 70 hPa drop temperatures range respectively over 185.7  
619 K – 194.9 K and over 194.8 K – 203.7 K, with an average of 192.0 +/- 2.9 K and 198.0 +/- 3.2 K (1  
620 standard deviation) over the ten years of IASI. The average values at 30 hPa and 70 hPa fall within the  
621 1 standard deviation associated with the average drop temperature at 50 hPa. It is also worth noting the  
622 agreement between the drop temperatures and the NAT formation threshold  
623 (TNAT ~193 K at 30 hPa and ~197 K at 70 hPa) (Lambert et al., 2016)."

CALIOP PSC data (Pitts et al 2013, doi:10.5194/acp-13-2975-2013) have been used to show that differ-  
ent PSC types exist in different temperature regimes, with ice PSCs detected close to the frost point, STS  
follows the expected equilibrium curve and NAT exhibits two preferred mode below the NAT existence tem-  
perature. The analysis presented here is not constrained by the simultaneous presence of known PSC types  
and in fact there may not even be any PSCs in the atmospheric path sampled. Therefore, it is too simplistic  
to compare the drop temperatures to TNAT. The proximity of the 10-year mean drop temperatures to TNAT  
does not constitute a validation as is claimed here. Individual years could be expected to show a variation  
in drop temperature because of interannual atmospheric differences. For instance, the years dominated by  
STS should necessarily show lower drop temperature than years dominated by NAT. The highest drop tem-  
peratures are far above PSC temperatures (e.g. 202.8K at 50hPa in one particular year) and deserve more  
scrutiny and should be investigated thoroughly. Interannual comparisons of the drop temperature may ben-  
efit from using (T-Tice) as the temperature coordinate (rather than absolute temperature) as this removes

variations due to changes in H<sub>2</sub>O partial pressure (see Fig 2 of Pitts et (2013)). There is a fundamental problem with making an assessment of the potential future scientific utility of the drop temperatures when they have only been evaluated in the absence of knowledge of the different types of PSCs present.

=====  
625 [L18: add more recent references e.g. Peter and Gross (2012). L28: Much more has been done in the  
626 past decade with MIPAS and CALIOP that should be referenced]

627 Thank you for this suggestion. Peter and GrooS (2012) was cited elsewhere in the manuscript but has  
628 been added here as well. Note that the goal of the introduction is not to provide an exhaustive list of all  
629 studies related to the PSC thermodynamics. Several general reference papers are cited and we have  
630 decided to put more focus here on HNO<sub>3</sub>.

=====  
632 [L59: This section should explain what is meant by "maximum sensitivity" etc.]

633 See our responses to the second major comment and specific comments above.

=====  
635 [L79: Information on the data quality for IASI HNO<sub>3</sub> is poor. Is the value of bias and uncertainty the  
636 same for depleted and non-depleted conditions?]

637 The reader is here invited to refer to the figure 4 of Ronsmans et al. (2016) which illustrates the global  
638 distribution of the total retrieval error for HNO<sub>3</sub> (integrated over 5 to 35 km) separately for January (left)  
639 and July (right) over the period of the IASI measurements. The mid- and polar latitudes are characterized  
640 by low total retrieval errors of around ~3-5% - which corresponds to a reduction by a factor of 18-30  
641 compared to the prior uncertainty (90%) and indicates a real gain of information – except above  
642 Antarctica during wintertime where the errors reach 25%. They are explained by (1) a weaker sensitivity  
643 (i.e. a larger smoothing error which represents in all cases the largest source of the retrieval error) above  
644 such cold surface (DOFS of ~0.95 within the dark Antarctic vortex – see figure 3 of Ronsmans et al.,  
645 2016) and by (2) a poor knowledge of the wavenumber-dependent surface emissivity above ice surface,  
646 which also varies in time (Hurtmans et al., 2012). ). This is made more explicit in Section 2 of the revised  
647 manuscript:

648

649 “The total columns are associated with a total retrieval error ranging from around 3% at mid- and polar  
650 latitudes to 25% above cold Antarctic surface during winter (due to a weaker sensitivity above very cold  
651 surface with a DOFS of ~0.95 and to an poor knowledge of the seasonally and wavenumber-dependent  
652 emissivity above ice surfaces which induces larger forward model errors), and a low bias (lower than  
653 12%) in polar regions over the altitude range where the IASI sensitivity is the largest, when compared  
654 to ground-based FTIR measurements (see Hurtmans et al., 2012; Ronsmans et al., 2016).”

The response does not address the specific case of whether there are differences in bias and uncertainty for depleted and non-depleted conditions.

=====  
656 [L82: Yet, problems with the retrievals because of cloud contamination seem to remain even after the  
657 25% cloud fraction filter is applied.]



658 We do not understand the referee comment here. In this section of the manuscript, we only describe the  
659 quality flags used in our analysis.

Even after all the quality controls are applied there are apparently still cases with poor retrievals that could  
be removed.

=====

661 [L83: Cloud contamination? Tropospheric cloud only or also thick ice PSCs?]

662 The clouds that have most impact are clearly tropospheric water clouds. Cirrus clouds or PSCs are mostly  
663 transparent in the IR; thick cirrus however show up in the longwave part of the IASI spectrum, below  
664 900 cm<sup>-1</sup>. We have added “tropospheric cloud contamination” in the text.

665

666 Note that the threshold of 25 % cloud cover was carefully chosen after a series of tests, which have  
667 shown that these scenes could be treated as cloud-free without significant impact on the retrievals  
668 (Hurtmans et al., 2012).

Thick ice PSCs have been detected by AIRS, TOVS HIRS2 and AVHRR (see  
Stajner et al. and refs therein,  
<https://doi.org/10.1029/2007GL029415>). Do these have an effect on the  
HNO<sub>3</sub> retrieved by IASI?

=====

670 [L102: Why was 2011 chosen?]

671 As expected from figure 1c, any other year could have been chosen instead of the year 2011 to illustrate  
672 the HNO<sub>3</sub> total columns versus temperatures (at 50 hPa) histogram in figure 1b. It is now clearly  
673 mentioned in the revised manuscript:

674

675 “Similar histograms are observed for the ten years of IASI measurements (not shown).”

=====

677 [L106: Heterogeneous hydrolysis of N<sub>2</sub>O<sub>5</sub> requires aerosol particles. So this process starts with cold  
678 binary aerosols (i.e. sulfates) before the formation of STS?]

679 Indeed, previous studies have shown enhanced HNO<sub>3</sub> columns during autumn in Antarctica and have  
680 attributed them to decreasing sunlight and conversion of N<sub>2</sub>O<sub>5</sub> to HNO<sub>3</sub> by the reaction of N<sub>2</sub>O<sub>5</sub> with  
681 background aerosols, before the formation of polar stratospheric clouds (e.g. Keys et al., Nature, 1993).  
682 At these temperatures, the conversion may occur on binary sulfuric aerosols.

683

684 The sentence has been rewritten as follows:

685

686 “These high HNO<sub>3</sub> levels result from low sunlight, preventing photodissociation, along with the  
687 heterogeneous hydrolysis of N<sub>2</sub>O<sub>5</sub> to HNO<sub>3</sub> during autumn before the formation of polar stratospheric  
688 clouds (Keys et al., 1993; Santee et al., 1999; Urban et al., 2009; DeZafra et al., 2001). This period also  
689 corresponds to the onset of the deployment of the southern polar vortex which is characterized by strong

690 diabatic descent with weak latitudinal mixing across its boundary, isolating polar HNO<sub>3</sub>-rich air from  
691 lower latitudinal airmasses.”

692

693 [L129: The onset of depletion seems to start when the temperatures fall substantially below 190K from  
694 inspection of Fig 1(c) and quite far below the red line marked at 195K.]

695 The onset of HNO<sub>3</sub> depletion starts in June at around 190K, which is in agreement with figure 1a.

=====  
697 [L136-137: Why are two temperatures (180 and 185 K) quoted for 30hPa? Why is the actual value from  
698 Fig1(c) (I estimate this as about 188K) for the 50hPa temperature not given in L129?]

699 The sentence has been rewritten for clarity:

700 “The results (not shown here) exhibit a similar HNO<sub>3</sub>-temperature behaviour at the different levels with,  
701 as expected, lower and larger temperatures in R2, respectively, at 30 hPa (down to 180 K) and at 70 hPa  
702 145 (down to 185K), but still below the NAT formation threshold at these pressure levels (TNAT =193 K  
703 at 30 hPa and 197 K at 70 hPa) (Lambert et al., 2016).”

=====  
705 [L138: ”characterized by” seems the wrong description for the chance occurrence that the maximum  
706 sensitivity of IASI HNO<sub>3</sub> falls in the same altitude range as the PSCs.]

707 Changed to: “. . . the altitude range of maximum IASI sensitivity to HNO<sub>3</sub> (see Section 2) is characterized  
708 by temperatures that are below the NAT formation threshold at these pressure levels, enabling the PSCs  
709 formation and the denitrification process.”

=====  
[L139-146: This section rather seems to belong in 711 the conclusions.]

712 L150-154 of the revised manuscript has been moved to the conclusions.

=====  
714 [L148: Clearly this does not ”go beyond the vertically integrated view” since the column HNO<sub>3</sub> is all  
715 that is available. It could be reworded as ”To identify the spatial and temporal variability of the column  
716 HNO<sub>3</sub> ...”]

717 Corrected as suggested.

718

719 [L165-169: Denitrification is the term used to describe the permanent removal of some HNO<sub>3</sub> from the  
720 gas phase by sedimentation of PSCs. Sequestration is the term used to describe the uptake of HNO<sub>3</sub>  
721 from the gas phase into PSCs. Denitrification by STS is a lengthy process compared to NAT since the  
722 smaller STS particles sediment slowly. STS can (and frequently does) form without the prior nucleation  
723 of NAT. IASI alone cannot discriminate between these processes and it should not be assumed that what  
724 is observed is the ”onset of HNO<sub>3</sub> denitrification”.]

725 We thank the referee for this remark. We are of course aware of the definition of the so-called  
726 ”denitrification”. We agree that, from IASI, we can only measure a ”removal from the gas phase”, caused

727 by sequestration into particles with or without sedimentation. Careful attention has now been made in  
728 the manuscript to avoid abusive use of the term “denitrification”. Hence, “onset of HNO<sub>3</sub> denitrification”  
729 has been changed to “the onset of HNO<sub>3</sub> depletion” in L.169 and where appropriated in the revised  
730 manuscript and the title has also been changed accordingly to:  
731 “Polar stratospheric HNO<sub>3</sub> depletion surveyed from a decadal dataset of IASI total columns”.

=====  
733 [L185-187: 210K is much too high for PSC formation, but could possibly be NAT that is in process of  
734 melting? If these are observed over ocean then they warrant further investigation. However, why are  
735 specific regions with emissivity features not flagged as such? They should be discarded rather than  
736 “used with caution”.]

737 Bright land surface such as desert or ice might in some cases lead to poor HNO<sub>3</sub> retrievals due to a poor  
738 knowledge of the wavenumber-dependent emissivity above such surfaces, which can alter the retrieval  
739 by compensation effects (Wespes et al., 2009). FORLI relies on the monthly climatology of surface  
740 emissivity built by Zhou et al. (2011) from several years of IASI measurements on a 0.5x0.5 grid and  
741 for each 8461 IASI spectral channels when available, or on the MODIS climatology that is unfortunately  
742 restricted to only 12 channels in the IASI spectral range; see Hurtmans et al. (2012) for more details.  
743 Although wavenumber-dependent surface emissivity atlases are used in FORLI, it is clear that this  
744 parameter remains critical and causes poorer retrievals that, in some instances, pass the posterior  
745 filtering. The total HNO<sub>3</sub> columns over eastern Antarctica which show drop temperatures much above  
746 195K might precisely be related to this. We have made this clear in Section 4.2 of the revised version:  
747

748 “. . . emissivity features that are known to yield errors in the IASI retrievals. Indeed, bright land surface  
749 such as ice might in some cases lead to poor HNO<sub>3</sub> retrievals. Although wavenumber-dependent surface  
750 emissivity atlases are used in FORLI (Hurtmans et al., 2012), this parameter remains critical and causes  
751 poorer retrievals that, in some instances, pass through the series of quality filters and could affect the  
752 drop temperature calculation.”

753  
754 We refer on the good agreement with MLS (suggested by the referee) to underline the potentiality of  
755 IASI to detect the HNO<sub>3</sub> variations as well within the Antarctic winter (see general comment and Figure  
756 1 here below).

What is the fraction of data that is affected by surface emissivity?

=====  
758 [L189: Modern reanalysis temperatures (e.g. ERA-I) do not “feature large uncertainties” large enough  
759 to account for a 195K to 210K shift. L195-L201: The limitations of the reanalysis temperatures seems  
760 to be an accuracy of better than 1K and clearly this in no way limits the derivation 760 of the “50hPa drop  
761 temperature” which simply necessitates finding the 50hPa reanalysis temperature that corresponds to  
762 the second derivative wrt time minimum in column HNO<sub>3</sub>.]

=====  
763 We agree with the referee’s comment; the discussion about the potential role of the uncertainty of the  
764 ECMWF reanalysis temperature on the drop temperature has been removed from the section, hence, this  
765 paragraph has been strongly revised accordingly:  
766

767 "... while biases in ECMWF reanalysis are too small for explaining the spatial variation in drop  
768 temperatures. Thanks to the assimilation of an advanced Tiros Operational Vertical Sounder (ATOVS)  
769 around 1998–2000 in reanalyses, to the better coverage of satellite instruments and to the use of global  
770 navigation satellite system (GNSS) radio occultation (RO) (Schreiner et al., 2007; Wang et al., 2007;  
771 Lambert and Santee, 2018; Lawrence et al., 2018), the uncertainties have been vastly reduced.  
772 Comparisons of the ECMWF ERA Interim dataset used in this work with the COSMIC data (Lambert  
773 and Santee, 2018) found a small warm bias, with median differences around 0.5 K, reaching 0–0.25 K  
774 in the southernmost regions of the globe at ~68–21 hPa where PSCs form."

=====  
776 [What is meant by "spatial variability"? The plots in Fig 5 show the spatial distribution of the drop  
777 temperature over a number of years but what variability is being considered? Interannual? Why have  
778 these spatial maps of drop temperatures not been compared with published maps of PSC types made by  
779 CALIOP or MIPAS. Wouldn't some correlation be expected according to the arguments made here? i.e.  
780 NAT PSCs at the higher temperature e.g. the highest temperatures (orange) appear downstream of the  
781 Palmar Peninsula in the "NAT ring" structure described by Hopfner et al (2006).]

782 Corrected: "Figure 5 shows the spatial variability" .. "Figure 5 shows the spatial distribution".

783

784 We do not understand the referee's comment here. Figure 5 of the manuscript shows the spatial  
785 distribution of the drop temperature calculated inside a region enclosed by an isocontour PV of  $-8 \times 10^{-5}$   
786  $\text{K} \cdot \text{m}^2 \cdot \text{kg}^{-1} \cdot \text{s}^{-1}$ , which, hence, encircles a region larger than the inner vortex core (see Figures 3 and 4 of  
787 the manuscript). The drop temperatures much above the NAT formation temperature, which are mostly  
788 found outside the averaged isocontour PV of  $-10 \times 10^{-5} \text{K} \cdot \text{m}^2 \cdot \text{kg}^{-1} \cdot \text{s}^{-1}$ , do not correspond to high minima  
789 ( $< -0.5 \times 10^{14} \text{ molec} \cdot \text{cm}^{-2} \cdot \text{d}^{-2}$ ) in the second derivative of  $\text{HNO}_3$  total column with respect to time. We  
790 cannot argue that it corresponds to the NAT belt of Höpfner et al. (2006) downstream of the Antarctic  
791 Peninsula, which was enclosed inside the region of the NAT threshold temperature; the highest drop  
792 temperatures from IASI are found on the contrary outside the isocontour of the NAT threshold  
793 temperature (see figure 5 of the revised manuscript). In addition, comparing the distributions of drop  
794 temperatures from IASI with PSC information from CALIPSO/MIPAS remain difficult given the  
795 difference in spatial coverage and, most importantly, the highly variable distribution of PSC types and  
796 of the NAT belt, temporally (daily) and spatially (Höpfner et al., 2006; Lambert et al., 2012).

797

798 Finally, in response to G. Manney and M. Santee, the contour of  $-10 \times 10^{-5} \text{K} \cdot \text{m}^2 \cdot \text{kg}^{-1} \cdot \text{s}^{-1}$  based on the  
799 minimum PV encountered at 50 hPa over the 10 May to 15 July period as well as the isocontours of 195  
800 K at 50 hPa for the averaged temperatures and the minima over the same period are also now represented  
801 in the revised Fig.5 and the distribution of the drop temperatures is much better described and explained  
802 in the revised version:

803 "The averaged isocontour of 195 K encircles well the area of  $\text{HNO}_3$  drop temperatures lower than 195  
804 K, which means that the bins inside that area characterize airmasses that experience the NAT threshold  
805 temperature during a long time over the 10 May – 15 July period. That area encompasses the inner vortex  
806 core (delimited by the isocontour of  $-10 \times 10^{-5} \text{K} \cdot \text{m}^2 \cdot \text{kg}^{-1} \cdot \text{s}^{-1}$  for the averaged PV) and show pronounced  
807 minima (lower than  $-0.5 \times 10^{14} \text{ molec} \cdot \text{cm}^{-2} \cdot \text{d}^{-2}$ ) in the second derivative of the  $\text{HNO}_3$  total column with  
808 respect to time (not shown here), which indicate a strong and rapid  $\text{HNO}_3$  depletion.

809 The area enclosed between the two isocontours of 195 K for the temperatures, the averaged one and the  
810 one for the minimum temperatures, show higher drop temperatures and weakest minima (larger than  $-$   
811  $0.5 \times 10^{14} \text{ molec} \cdot \text{cm}^{-2} \cdot \text{d}^{-2}$ ) in the second derivative of the  $\text{HNO}_3$  total column (not shown). That area is

812 also enclosed by the isocontour of  $-10 \times 10^{-5} \text{K.m}^2.\text{kg}^{-1}.\text{s}^{-1}$  for the minimum PV, meaning that the bins  
813 inside correspond, at least for one day over the 10 May - 15 July period, to airmasses located at the inner  
814 edge of the vortex and characterized by temperature lower than the NAT threshold temperature. The  
815 weakest minima in the second derivative of total HNO<sub>3</sub> (not shown) observed in that area indicate a  
816 weak and slow HNO<sub>3</sub> depletion and might be explained by a short period of the NAT threshold  
817 temperature experienced at the inner edge of the vortex. It could also reflect a mixing with strong HNO<sub>3</sub>-  
818 depleted and colder airmasses from the inner vortex core. The mixing with these “already” depleted  
819 airmasses could also explained the higher drop temperatures detected in those bins. Finally, note also  
820 that these high drop temperatures are generally detected later (after the HNO<sub>3</sub> depletion occurs, i.e. after  
821 the 10 May – 15 July period considered here – not shown), which supports the transport, in those bins,  
822 of earlier HNO<sub>3</sub>-depleted airmasses and the likely mixing at the edge of the vortex.”

=====  
824 [L205: Nothing has been presented that demonstrates PSC occurrence. For that you would need to  
825 compare to actual data on PSCs from CALIOP and/or MIPAS.]

826 Corrected: “PSCs occurrence” .. “NAT formation temperature”

=====  
828 [L224: Again, the suspect data should be discarded because of the detrimental impact on the scientific  
829 analysis. Also, if you cannot manage to work out and apply adequate quality control to your own data  
830 then you have no reason to expect anyone else to do so.]

831 See our response to comment [L185-187] above.

=====  
833 [L230: “To the best of our knowledge, it is the first time that such a large satellite observational data set  
834 of stratospheric HNO<sub>3</sub> concentrations is exploited to monitor the evolution HNO<sub>3</sub> versus temperatures”  
835 In fact you cite several papers that have done exactly this, but let’s take the one published over two  
836 decades ago by Santee et al (1999) titled “Six years of UARS Microwave Limb Sounder HNO<sub>3</sub>  
837 observations : Seasonal, interhemispheric, and interannual variations in the lower stratosphere”.  
838 <https://doi.org/10.1029/1998JD100089>. Not only does this paper compare HNO<sub>3</sub> with UKMO  
839 temperatures we are referred to a more complete paper on this topic on p8241 ... “The correlation of the  
840 HNO<sub>3</sub> behavior with temperature during this time period, and its implications for PSC phase and  
841 composition, is explored in detail by Santee et al (1998). I noticed that the outside edge of the “HNO<sub>3</sub>  
842 collar region” at 465K was defined by these authors as inside the  $0.25 \times 10^{-4} \text{K m}^2 \text{kg}^{-1} \text{s}^{-1}$  PV contour.  
843 This seems at odds with the  $1 \times 10^{-4}$  value that is used for the second derivative minimum calculation in  
844 this paper and seemingly places the boundary quite far equatorward. Santee et al (1998) also includes a  
845 description of the heterogeneous hydration of N<sub>2</sub>O<sub>5</sub> that would be helpful in response to the question  
846 above on L106.]

=====  
847 We here simply refer to the unprecedented potential of IASI in terms of its exceptional spatial and  
848 temporal sampling. Ronsmans et al. (2018) also referred to the IASI dataset and correlations with  
849 temperature were done but in a lesser extent. In order to avoid overselling, the sentence has been



850 rewritten:

851 “We show in this study that the IASI dataset allows capturing the variability of stratospheric HNO<sub>3</sub>  
852 throughout the year (including the polar night) in the Antarctic. In that respect, it offers a new  
853 observational means to monitor the relation of HNO<sub>3</sub> to temperature and the related formation of PSCs.”  
854

855 In this study, we use the PV fields taken from the ECMWF ERA Interim Reanalysis dataset at the  
856 potential temperature of 530 K (corresponding to ~20 km where the IASI sensitivity to HNO<sub>3</sub> is the  
857 highest), while Santee et al. (1998) considered 465K. We clearly see from Figures 3a and 4 of the  
858 manuscript that PV contours at  $-0.5 \times 10^{-4} \text{ K m}^2 \text{ kg}^{-1} \text{ s}^{-1}$  and at  $-0.8 \times 10^{-4} \text{ K m}^2 \text{ kg}^{-1} \text{ s}^{-1}$  encompass the so  
859 called HNO<sub>3</sub> collar region. The PV value of  $-1 \times 10^{-4} \text{ K m}^2 \text{ kg}^{-1} \text{ s}^{-1}$  is used in this study to calculate the  
860 drop temperature based on the second derivative minimum as it clearly encompass the regions inside the  
861 inner polar vortex (see Figure 3a and 4 of the manuscript).

=====  
863 [L231: “It could constitute a new accurate climatological parameter that could be inserted in the PSCs  
864 classification schemes.” The analysis presented does not support this statement. Specifically, how could  
865 the HNO<sub>3</sub> column amount be used in a classification scheme?]

866 This sentence has been removed.

The conclusion of the paper is that ability to monitor the polar atmosphere over several decades with current and planned IASI instruments “will provide an unprecedented long-term dataset of HNO<sub>3</sub> total columns”. The drop temperature is defined as the 50hPa temperature corresponding to the greatest rate of decline of the column HNO<sub>3</sub> with respect to time. However, even with a record now extending over a decade the scientific utility of this dataset has not been demonstrated.