

# Polar stratospheric nitric acid depletion surveyed from a decadal dataset of IASI total columns

Gaetane Ronsmans<sup>1</sup>, Catherine Wespes<sup>1,\*</sup>, Lieven Clarisse<sup>1</sup>, Susan Solomon<sup>2</sup>, Daniel Hurtmans<sup>1</sup>, Cathy Clerbaux<sup>1,3</sup>, and Pierre-François Coheur<sup>1</sup>

<sup>1</sup>Université libre de Bruxelles (ULB), Spectroscopy, Quantum Chemistry and Atmospheric Remote Sensing (SQUARES), Brussels, Belgium

<sup>2</sup>Department of Earth, Atmospheric and Planetary Sciences, Massachusetts Institute of Technology, Cambridge, Massachusetts, USA

<sup>3</sup>LATMOS/IPSL, Sorbonne Université, UVSQ, CNRS, Paris, France

Correspondence: Catherine Wespes ([cwespes@ulb.ac.be](mailto:cwespes@ulb.ac.be))

## Abstract

In this paper, we exploit the first 10-year data-record (2008-2017) of nitric acid (HNO<sub>3</sub>) total columns measured by the IASI-A/Metop infrared sounder, characterized by an exceptional daily sampling and a good vertical sensitivity in the mid-stratosphere (around 50 hPa), to monitor the causal relationship between the temperature decrease and the observed HNO<sub>3</sub> loss that occurs each year in the Antarctic stratosphere during the polar night. Since the HNO<sub>3</sub> depletion results from the formation of polar stratospheric clouds (PSCs) which trigger the development of the ozone (O<sub>3</sub>) hole, its continuous monitoring is of high importance. We verify here, from the 10-year time evolution of the pair HNO<sub>3</sub>-temperature (taken from reanalysis at 50 hPa), the recurrence of specific regimes in the cycle of IASI HNO<sub>3</sub> and identify, for each year, the day and the 50 hPa temperature ("drop temperature") corresponding to the onset of strong HNO<sub>3</sub> depletion in the Antarctic winter. Although the measured HNO<sub>3</sub> total column does not allow differentiating the uptake of HNO<sub>3</sub> by different types of PSC particles along the vertical profile, an average drop temperature of  $-194.2 \pm 3.8$  K, ~~consistent with close to~~ the nitric acid trihydrate (NAT) ~~formation temperature existence threshold (close to~~  $\sim 195$  K at 50 hPa), is found in the region of potential vorticity lower than  $-10 \times 10^{-5}$  K.m<sup>2</sup>.kg<sup>-1</sup>.s<sup>-1</sup> (~~similar to the 70° – 90° S Eqlat region during winter~~). The spatial distribution and inter-annual variability of the drop temperature are investigated and discussed ~~in the context of previous PSCs studies~~. This paper highlights the capability of the IASI sounder to monitor the long-term evolution of the polar stratospheric composition and processes involved in the depletion of stratospheric O<sub>3</sub>.

## 1 Introduction

The cold and isolated air masses found within the polar vortex during winter are associated with a strong denitrification of the stratosphere due to the formation of PSCs (composed of HNO<sub>3</sub>, sulphuric acid (H<sub>2</sub>SO<sub>4</sub>) and water ice (H<sub>2</sub>O)) (Peter, 1997; Voigt et al., 2000; von König, 2002; Schreiner et al., 2003; Peter and Groß, 2012). These clouds strongly affect the polar chemistry by (1) acting as surfaces for the heterogeneous activation of chlorine and bromine compounds, in turn leading to enhanced O<sub>3</sub> destruction (Solomon, 1999; Wang and Michelangeli, 2006; Harris et al., 2010; Wegner et al., 2012) and by (2) removing gas-phase HNO<sub>3</sub> temporarily or permanently through uptake by PSCs and sedimentation of large PSC particles to lower altitudes. The denitrification of the polar stratosphere during winter delays the reformation of chlorine reservoirs and, hence, intensifies the O<sub>3</sub> hole (Solomon, 1999; Harris et al., 2010). The heterogeneous reaction rates on PSCs surface and the uptake of HNO<sub>3</sub> strongly depend on the temperature and on the PSCs particle type. The PSCs are classified into three

49 3 different types based on their composition and optical properties: type Ia solid nitric acid trihydrate -  
50 NAT ( $\text{HNO}_3 \cdot (\text{H}_2\text{O})_3$ ), type Ib liquid supercooled ternary solution - STS ( $\text{HNO}_3/\text{H}_2\text{SO}_4/\text{H}_2\text{O}$  with  
51 variable composition) and type II, crystalline water-ice particles (likely composed of a combination of  
52 different chemical phases) (Toon et al., 1986; Koop et al., 2000; Voigt et al., 2000; Lowe and  
53 MacKenzie, 2008). In the stratosphere, they mostly consist of mixtures of liquid/solid STS/NAT  
54 particles in varying number densities, with  $\text{HNO}_3$  being the major constituent of these particles. The  
55 large-size NAT particles of low number density are the principal cause of sedimentation (Lambert et al.,  
56 2012; Pitts et al., 2013; Molleker et al., 2014; Lambert et al., 2016). The formation temperature of STS  
57 ( $T_{STS}$ ) and the thermodynamic equilibrium temperatures of NAT ( $T_{NAT}$ ) and ice ( $T_{ice}$ ), have been  
58 determined, respectively, as:  $\sim 192$  K (Carslaw et al., 1995),  $\sim 195.7$  K (Hanson and Mauersberger, 1988)  
59 and  $\sim 188$  K (Murphy and Koop, 2005) for typical 50 hPa atmospheric conditions (5 ppmv  $\text{H}_2\text{O}$  and 10  
60 ppbv  $\text{HNO}_3$ ). While the NAT nucleation was thought to require temperatures below  $T_{ice}$  and pre-existing  
61 ice particles, recent observational and modelling studies have shown that  $\text{HNO}_3$  starts to condense in  
62 early PSC season in liquid NAT mixtures well above  $T_{ice}$  ( $\sim 4$  K below  $T_{NAT}$ , close to  $T_{STS}$ ) even after a  
63 very short temperature threshold exposure (TTE) to these temperatures but also slightly below  $T_{NAT}$  after  
64 a long TTE, whereas the NAT existence persists up to  $T_{NAT}$  (Pitts et al., 2013; Hoyle et al., 2013; Lambert  
65 et al., 2016; Pitts et al., 2018). It has been recently proposed that the higher temperature condensation  
66 results from heterogeneous nucleation of NAT on meteoritic dust in liquid aerosol (Hoyle et al., 2013;  
67 Groöb et al., 2014; James et al., 2018). Further cooling below  $T_{STS}$  and  $T_{ice}$  leads to nucleation of liquid  
68 STS, of solid NAT onto ice and of ice particles mainly from STS (type II PSCs) (Lowe and MacKenzie,  
69 2008). The formation of NAT and ice has also been shown to be triggered by stratospheric mountain-  
70 waves (Carslaw et al., 1998; Hoffmann et al., 2017). Although the formation mechanisms and  
71 composition of STS droplets in stratospheric conditions are well described (Toon et al., 1986; Carslaw  
72 et al., 1995; Lowe and MacKenzie, 2008), the NAT and ice nucleation processes still require further  
73 investigation. This could be important as the chemistry-climate models (CCMs) generally oversimplify  
74 the heterogeneous nucleation schemes for the PSCs formation (Zhu et al., 2015; Spang et al., 2018; Snels  
75 et al., 2019) preventing an accurate estimation of  $\text{O}_3$  levels. The influence of  $\text{HNO}_3$  in modulating  $\text{O}_3$   
76 abundances in the stratosphere is furthermore underrepresented in CCMs (Kvissel et al., 2012).

77  
78 Several satellite instruments measure stratospheric  $\text{HNO}_3$  (e.g. MLS/UARS (Santee et al., 1999),  
79 MLS/Aura (Santee et al., 2007), MIPAS/ENVISAT (Piccolo and Dudhia, 2007), ACE-FTS/SCISAT  
80 (Sheese et al., 2017) and SMR/Odin (Urban et al., 2009)). The spaceborne lidars CALIOP/CALIPSO  
81 and the infrared instrument MIPAS/Envisat) are capable to detect and classify the PSC types, and to  
82 follow their formation mechanisms (Lambert et al., 2016; Pitts et al., 2018; Spang et al., 2018) and  
83 references therein, which complements in situ measurements (Voigt et al., 2005) and ground-based lidar  
84 (Snels et al., 2019). From these available observational datasets, the  $\text{HNO}_3$  depletion has been linked to  
85 the PSCs formation and detected below the  $T_{NAT}$  threshold (Santee et al., 1999; Urban et al., 2009;  
86 Lambert et al., 2016; Ronsmans et al., 2018), but its relationship to PSCs still needs further investigation  
87 given the complexity of the nucleation mechanisms that depends on a series of parameters (e.g.  
88 atmospheric temperature, water and  $\text{HNO}_3$  vapour pressure, time exposure to temperatures, temperature  
89 history).

90  
91 In contrast to the limb satellite instruments mentioned above, the infrared nadir sounder IASI offers a  
92 dense spatial sampling of the entire globe, twice a day (Section 2). While it cannot provide a vertical  
93 profile of  $\text{HNO}_3$  similar to the limb sounders, IASI provides reliable total column measurements of  
94  $\text{HNO}_3$  characterized by a maximum sensitivity in the low-middle stratosphere around 50 hPa (20 km)  
95 during the dark Antarctic winter (Ronsmans et al., 2016, 2018) where the PSCs cloud form (Voigt et al.,  
96 2005; Lambert et al., 2012; Spang et al., 2016, 2018). This study aims to explore the 10-years continuous  
97  $\text{HNO}_3$  measurements from IASI for providing a long-term global picture of depletion and of its

dependence to temperatures during polar winter (Section 3). The temperature corresponding to the onset of the strong depletion in HNO<sub>3</sub> records (here referred to as ‘drop temperature’) is identified in Section 4 for each observed year and discussed in the context of previous studies.

## 2 Data

The HNO<sub>3</sub> data used in the present study are obtained from measurements of the Infrared Atmospheric Sounding Interferometer (IASI) embarked on the Metop-A satellite. IASI measures the Earth’s and atmosphere’s radiation in the thermal infrared spectral range (645 - 2760 cm<sup>-1</sup>), with a 0.5 cm<sup>-1</sup> apodized resolution and a low radiometric noise (Clerbaux et al., 2009; Hilton et al., 2012). Thanks to its polar sun-synchronous orbit with more than 14 orbits a day and a field of view of four simultaneous footprints of 12 km at nadir, IASI provides global coverage twice a day (9.30 AM and PM mean local solar time). That extensive spatial and temporal sampling in the polar regions is key to this study.

The HNO<sub>3</sub> vertical profiles are retrieved on a uniform vertical 1 km grid of 41 layers (from the surface to 40 km with an extra layer above to 60 km) in near-real-time by the Fast Optimal Retrieval on Layers for IASI (FORLI) software, using the optimal estimation method (Rodgers, 2000). Detailed information on the FORLI algorithm and retrieval parameters specific to HNO<sub>3</sub> can be found in previous papers (Hurtmans et al., 2012; Ronsmans et al., 2016). For this study, only the total columns (v20151001) are used, considering (1) the low vertical resolution of IASI with only one independent piece of information (FWHM of the averaging kernels of ~30 km), (2) the limited sensitivity of IASI to tropospheric HNO<sub>3</sub>, (3) the dominant contribution of the stratosphere to the HNO<sub>3</sub> total column and (4) the largest sensitivity of IASI in the region of interest, i.e. in the low and mid-stratosphere (from ~70 to ~30 hPa), where the HNO<sub>3</sub> abundance is the highest (Ronsmans et al., 2016). The IASI measurements capture the expected variations of HNO<sub>3</sub> within the polar night, as illustrated in Fig. 1 that shows examples of vertical HNO<sub>3</sub> profiles retrieved within the dark Antarctic vortex (above Arrival Height) and outside the vortex (above Lauder). The retrieved profiles are shown along with their associated total retrieval error and averaging kernels (the total column averaging kernels and the so-called “sensitivity profile” are also represented; see Ronsmans et al., 2016 for more details). Above Arrival Height during the dark Antarctic winter, we clearly see depleted HNO<sub>3</sub> levels in the low and mid-stratosphere and the altitude of maximum sensitivity (at around 30 hPa for this case). At Lauder on the contrary, HNO<sub>3</sub> levels larger than the a priori are observed in the stratosphere with a larger range of maximum sensitivity. The total columns are associated with a total retrieval error ranging from around 3% at mid- and polar latitudes to 25% above cold Antarctic surface during winter (due to a weaker sensitivity above very cold surface with a DOFS of 0.95 and to a poor knowledge of the seasonally and wavenumber-dependent emissivity above ice surfaces which induces larger forward model errors), and a low bias (lower than 12%) in polar regions over the altitude range where the IASI sensitivity is the largest, when compared to ground-based FTIR measurements (see Hurtmans et al., 2012 and Ronsmans et al., 2016 for more details). In order to expand on the comparisons against FTIR measurements which is not possible during the polar night, Fig. 2 (top panel) presents the time series of daily IASI total HNO<sub>3</sub> columns co-located with MLS measurements within 2.5x2.5 grid boxes, averaged in the 70°S–90°S equivalent latitude band. In order to account for the vertical sensitivity of IASI, the averaging kernels associated with each co-located IASI retrieved profiles were considered for this cross-comparison. The MLS profiles were first interpolated to the FORLI pressure grids, then converted into column profiles. They were also extended down to the surface by considering the FORLI-HNO<sub>3</sub> a priori profile. Similar variations in the HNO<sub>3</sub> column are captured by the two instruments, with an excellent agreement in particular for the timing of the strong HNO<sub>3</sub> depletion within the inner vortex core. Note that a similar good agreement between the two satellite datasets is obtained in other latitude bands (see Fig. 2 bottom panel for the 50°S–70°S equivalent latitude band; the other bands are not shown). In order to expand on the comparisons against FTIR measurements

147 which is impossible during the polar night, Figure 1 (top panel) presents the time series of daily IASI  
148 total HNO<sub>3</sub> columns co-located with MLS VMR measurements within 2.5×2.5 grid boxes at three  
149 pressure levels (at 30, 50 and 70 hPa), averaged in the 70°–90° S equivalent latitude band. Similar  
150 variations in HNO<sub>3</sub> are captured by the two instruments with an excellent agreement for the timing of  
151 the strong HNO<sub>3</sub> depletion within the inner vortex core. IASI HNO<sub>3</sub> variations generally match well  
152 those of MLS HNO<sub>3</sub> in each latitude band (see Figure 1 bottom panel for the 50°–70° S equivalent  
153 latitude band; the other bands are not shown here).

154  
155 Quality flags similar to those developed for O<sub>3</sub> in previous IASI studies (Wespes et al., 2017) were  
156 applied a posteriori to exclude data (i) with a corresponding poor spectral fit (e.g. based on quality flags  
157 rejecting biased or sloped residuals, fits with maximum number of iteration exceeded), (ii) with less  
158 reliability (e.g. based on quality flags rejecting suspect averaging kernels, data with less sensitivity  
159 characterized by a DOFS lower than 0.9) or (iii) with tropospheric cloud contamination (defined by a  
160 fractional cloud cover ≥ 25 %). Note that the HNO<sub>3</sub> total column distributions illustrated in sections  
161 below use the median as a statistical average since it is more robust against the outliers than the mean.

162  
163 Temperature and potential vorticity (PV) fields are taken from the ECMWF ERA Interim Reanalysis  
164 dataset, respectively at 50 hPa and at the potential temperature of 530 K (corresponding to ~20 km  
165 altitude where the IASI sensitivity to HNO<sub>3</sub> is the highest during the Southern Hemisphere (S.H.) winter  
166 (Ronsmans et al., 2016). Because the HNO<sub>3</sub> uptake by PSCs starts a few degrees or slightly below T<sub>NAT</sub>  
167 (~195.7 K at 50 hPa (Hanson and Mauersberger, 1988)) depending on the meteorological conditions  
168 (Pitts et al., 2013; Hoyle et al., 2013; Lambert et al., 2016; Pitts et al., 2018), a threshold temperature of  
169 195 K is considered in the sections below to identify the PSCs-containing regions. The potential vorticity  
170 is used to delimit dynamically consistent areas in the polar regions. In what follows, we use either the  
171 equivalent latitudes ("eqlat", calculated from PV fields at 530 K) or the PV values to characterize the  
172 relationship between HNO<sub>3</sub> and temperatures in the cold polar regions. Uncertainties in ERA-Interim  
173 temperatures will also be discussed below.

### 174 175 **3 Annual cycle of HNO<sub>3</sub> vs temperatures**

176  
177 Figure 32a shows the yearly HNO<sub>3</sub> cycle (solid lines, left axis) in the southernmost equivalent latitudes  
178 (70° - 90° S), as measured by IASI over the whole period of measurements (2008–2017). The total HNO<sub>3</sub>  
179 variability in such equivalent latitudes has already been discussed in a previous IASI study (Ronsmans  
180 et al., 2018) where the contribution of the PSCs into the HNO<sub>3</sub> variations was highlighted. The  
181 temperature time series, taken at 50 hPa, is here represented as well (dashed lines, right axis). From this  
182 figure, different regimes of HNO<sub>3</sub> total columns vs temperature can be observed throughout the year and  
183 from one year to another. In particular, we define here three main regimes (R1, R2 and R3) along the  
184 HNO<sub>3</sub>/temperature cycle. The full cycle and the main regimes in the 70° - 90° S eqlat region are further  
185 represented in Fig. 23b that shows a histogram of the HNO<sub>3</sub> total columns as a function of temperature  
186 for the year 2011. Similar histograms are observed for the ten years of IASI measurements (not shown).  
187 The red horizontal and vertical lines in Fig. 32a and Fig. 32b, respectively, represent the 195 K threshold  
188 temperature used to identify the onset of HNO<sub>3</sub> uptake by PSCs (see Section 2). The three identified  
189 regimes correspond to:

- 190  
191 - R1 is defined by the maxima in the total HNO<sub>3</sub> abundances covering the months of April and  
192 May (~3×10<sup>16</sup> molec.cm<sup>-2</sup>, R1 in Figures 32a and b), when the 50 hPa temperature strongly  
193 decreases (from ~220 to ~195 K). These high HNO<sub>3</sub> levels result from low sunlight, preventing  
194 photodissociation, along with the heterogeneous hydrolysis of N<sub>2</sub>O<sub>5</sub> to HNO<sub>3</sub> during autumn  
195 before the formation of polar stratospheric clouds (Keys et al., 1993; Santee et al., 1999; Urban

196 et al., 2009; de Zafra and Smyshlyaev, 2001). This period also corresponds to the onset of the  
197 deployment of the southern polar vortex which is characterized by strong diabatic descent with  
198 weak latitudinal mixing across its boundary, isolating polar HNO<sub>3</sub>-rich air from lower latitudinal  
199 airmasses.

- 200
- 201 - R2 which extends from June to October is characterized by the onset of the strong decrease in  
202 HNO<sub>3</sub> total columns at the beginning of June, when the temperatures fall below 195 K, followed  
203 by a plateau of total HNO<sub>3</sub> minima. In this regime, the HNO<sub>3</sub> total columns average below  $2 \times 10^{16}$   
204 molec.cm<sup>-2</sup> and the 50 hPa temperatures range mostly between 180 and 190 K.
  - 205
  - 206 - R3 starts in October when sunlight returns and the 50 hPa temperatures rise above 195 K. Despite  
207 the stratospheric warming with 50 hPa temperatures up to 240 K in summer, the HNO<sub>3</sub> total  
208 columns stagnate at the R2 plateau levels (around  $1.5 \times 10^{16}$  molec.cm<sup>-2</sup>). This regime likely  
209 reflects the photolysis of NO<sub>3</sub> and HNO<sub>3</sub> itself (Ronsmans et al., 2018) as well as the permanent  
210 denitrification of the mid-stratosphere, caused by the PSCs sedimentation. The likely  
211 renitrification of the lowermost stratosphere (Braun et al., 2019; Lambert et al., 2012) where the  
212 HNO<sub>3</sub> concentrations and the IASI sensitivity to HNO<sub>3</sub> are lower (Ronsmans et al., 2016) can  
213 hardly not be inferred from the IASI total column measurements. The plateau lasts until  
214 approximately February, where HNO<sub>3</sub> total column slowly starts increasing, reaching the April-  
215 May maximum in R1.

216

217 As illustrated in Fig. 32a, the three regimes are observed each year with, however, some interannual  
218 variations. For instance, the sudden stratospheric warming (SSW) that occurs in 2010 (see the  
219 temperature time series at 20 hPa for the year 2010; green dotted line) yielded higher HNO<sub>3</sub> total columns  
220 (see green solid line in July and August) (de Laat and van Weele, 2011; Klekociuk et al., 2011; WMO,  
221 2014; Ronsmans et al., 2018).

222

223 Figure 32c shows the evolution of the relationship between the daily averaged HNO<sub>3</sub> (calculated from a  
224 7-day moving average) with the highest occurrence (in bins of  $0.1 \times 10^{16}$  molec.cm<sup>-2</sup> and of 2K) and the  
225 50 hPa temperature, over the 10 years of IASI. The red vertical line represents the 195 K threshold  
226 temperature. Figure 32c clearly illustrates the slow increase in HNO<sub>3</sub> columns as the temperatures  
227 decrease (February to May, i.e. R3 to R1), the strong and rapid HNO<sub>3</sub> depletion occurring in June (R2),  
228 the plateau of low HNO<sub>3</sub> abundances in winter and spring (from August to November; R2 to R3). Figure  
229 32c also highlights the interannual variability in total HNO<sub>3</sub>, which is found to be the largest in R3, and  
230 shows a strong consistency in the onset of the depletion between each year (beginning of June when the  
231 temperatures fall below 195 K as indicated by the red vertical line). Given the span of PSCs formation  
232 over a large range of altitudes (typically between 10 and 30 km) (Höpfner et al., 2006, 2009; Spang  
233 et al., 2018; Pitts et al., 2018) and that of maximum IASI sensitivity to HNO<sub>3</sub> around 50 hPa (Hurtmans  
234 et al., 2012; Ronsmans et al., 2016), the temperatures at two other pressure levels, namely 70 and 30 hPa  
235 (i.e. ~15 and ~25 km), have also been tested to investigate the relationship between HNO<sub>3</sub> and  
236 temperature in the low and mid-stratosphere. The results (not shown here) exhibit a similar HNO<sub>3</sub>-  
237 temperature behavior at the different levels with, as expected, lower and larger temperatures in R2,  
238 respectively, at 30 hPa (down to ~180 K) and at 70 hPa (down to ~185 K), but still below the NAT  
239 formation threshold at these pressure levels ( $T_{NAT} \sim 193$  K at 30 hPa and ~197 K at 70 hPa) (Lambert et  
240 al., 2016). Therefore, the altitude range of maximum IASI sensitivity to HNO<sub>3</sub> (see Section 2) is  
241 characterized by temperatures that are below the NAT formation threshold at these pressure levels,  
242 enabling the PSCs formation and the denitrification process. Furthermore, the consistency between the  
243 195 K threshold temperature taken at 50 hPa and the onset of the strong total HNO<sub>3</sub> depletion seen in  
244 IASI data (see Fig. 32a and Fig. 32c) is in agreement with the largest NAT area that starts to develop in

245 June around 20 km (Spang et al., 2018), which justifies the use of the 195 K temperature at that single  
246 representative level in this study.

247

#### 248 **4 Onset of HNO<sub>3</sub> depletion and drop temperature detection**

249

250 To identify the spatial and temporal variability of the onset of the depletion phase, the daily time  
251 evolution of HNO<sub>3</sub> during the first 10 years of IASI measurements and the temperatures at 50 hPa are  
252 explored. In particular, the second derivative of HNO<sub>3</sub> total column with respect to time is calculated to  
253 detect the strongest rate of decrease seen in the HNO<sub>3</sub> time series and to identify its associated day and  
254 50 hPa temperature.

255

##### 256 **4.1 HNO<sub>3</sub> vs temperature time series**

257

258 Figure 34 shows the time series of the second derivative of HNO<sub>3</sub> total column with respect to time  
259 (blue) and of the temperature (red) averaged in the areas of potential vorticity smaller than  $-10 \times 10^{-5}$   
260  $\text{K} \cdot \text{m}^2 \cdot \text{kg}^{-1} \cdot \text{s}^{-1}$  to encompass the regions inside the inner polar vortex where the temperatures are the  
261 coldest and the total HNO<sub>3</sub> depletion occurs (Ronsmans et al., 2018). The use of that PV threshold value  
262 explains the gaps in the time series during the summer when the PV does not reach such low levels,  
263 while the time series averaged in the 70°- 90° S Eqlat band (dashed blue for the second derivative of  
264 HNO<sub>3</sub> and grey for the temperature) covers the full year. Note that the HNO<sub>3</sub> time series has been  
265 smoothed with a simple spline data interpolation function to avoid gaps in order to calculate the second  
266 derivative of HNO<sub>3</sub> total column with respect to time as the daily second-difference HNO<sub>3</sub> total column.  
267 The horizontal red line shows the 195 K threshold.

268

269 As already illustrated in Fig. 32a and Fig. 32c, the strongest rate of HNO<sub>3</sub> depletion (i.e. the second  
270 derivative minimum) is found around the 195 K threshold temperature, within some days (4 to 23 days)  
271 after total HNO<sub>3</sub> reaches its maximum, i.e. typically between the 11th of May (2013) and the 8th of June  
272 (2009). Except for the year 2014, the 50 hPa drop temperatures are detected between 189.2 K and 198.6  
273 K (194.1 K  $\pm$  2.8 K - 1 $\sigma$  standard deviation - on average over the 10 years, excluding 2014 that stands  
274 out with a drop temperature of 202.8 K) 202.8 K, with an average of 194.2  $\pm$  3.8 K (1 $\sigma$  standard  
275 deviation) over the ten years. Knowing that  $T_{\text{NAT}}$  can be higher or lower depending on the atmospheric  
276 conditions and that NAT starts to nucleate from  $\sim 2-4$  K below  $T_{\text{NAT}}$  (Pitts et al., 2011; Hoyle et al., 2013;  
277 Lambert et al., 2016), the results here tend to demonstrate the consistency between the 50 hPa drop  
278 temperature, i.e. the temperature associated with the strongest HNO<sub>3</sub> depletion detected from IASI, and  
279 the PSCs NAT formation temperature in that altitude region mid-stratosphere at polar latitudes. Note  
280 that the range observed in the 50 hPa drop temperature could reflect the preponderance by one type of  
281 PSCs over another from one year to the next. The results further justify the use of the single 50 hPa  
282 level for characterizing and investigating the onset of HNO<sub>3</sub> depletion from IASI. Nevertheless, given  
283 the range of maximum IASI sensitivity to HNO<sub>3</sub> around 50 hPa, typically between 70 and 30 hPa  
284 (Ronsmans et al., 2016), the drop temperatures are also calculated at these two other pressure levels (not  
285 shown here) to estimate the uncertainty of the calculated drop temperature defined in this study at 50  
286 hPa. The 30 hPa and 70 hPa drop temperatures range respectively over 185.7 K – 194.9 K and over 194.8  
287 K – 203.7 K, with an average of  $192.0 \pm 2.9$  K and  $198.0 \pm 3.2$  K (1 $\sigma$  standard deviation) over the ten  
288 years of IASI. The average values at 30 hPa and 70 hPa fall within the 1 $\sigma$  standard deviation associated  
289 with the average drop temperature at 50 hPa. It is also worth noting the agreement between the drop  
290 temperatures and the NAT formation threshold at these two pressure levels ( $T_{\text{NAT}} \sim 193$  K at 30 hPa and  
291  $\sim 197$  K at 70 hPa) (Lambert et al., 2016).

292

293 Figures 54a and b show the zonal distribution of HNO<sub>3</sub> total columns and of the temperature at 50 hPa,  
294 respectively, spanning 55° - 90° S over the whole IASI period, with, superimposed, three isocontour  
295 levels of potential vorticity ( $-10$ ,  $-8$  and  $-5 \times 10^{-5}$  K.m<sup>2</sup>.kg<sup>-1</sup>.s<sup>-1</sup> in blue, cyan and black, respectively) and  
296 one isocontour for the 50 hPa temperature. The PV isocontour of  $-10 \times 10^{-5}$  K.m<sup>2</sup>.kg<sup>-1</sup>.s<sup>-1</sup> is clearly shown  
297 to separate well the region of strong depletion in total HNO<sub>3</sub> according to the latitude and the time. The  
298 red vertical dashed lines indicates the average date for the drop temperatures calculated in the area of  $PV \leq -$   
299  $10 \times 10^{-5}$  K.m<sup>2</sup>.kg<sup>-1</sup>.s<sup>-1</sup> ( $194.2 \pm 3.8$  K; see Fig. 43) over the IASI period. It shows that the strongest rate  
300 in HNO<sub>3</sub> depletion occurs on average a few days before June. The delay of some days between the  
301 maximum in total HNO<sub>3</sub> and the start of the depletion (see Fig. 43) is also visible in Fig. 54a. The yearly  
302 zonally averaged time series over the ten years of IASI can be found in Fig. 65; it shows the  
303 reproducibility of the edge of the collar HNO<sub>3</sub> region and of the region of the strong HNO<sub>3</sub> depletion,  
304 respectively delimited by the PV isocontours of  $-5 \times 10^{-5}$  K.m<sup>2</sup>.kg<sup>-1</sup>.s<sup>-1</sup> and of  $-10 \times 10^{-5}$  K.m<sup>2</sup>.kg<sup>-1</sup>.s<sup>-1</sup> at 50  
305 hPa, measured by IASI from year to year.

306

## 307 4.2 Distribution of drop temperatures

308

309 To explore the capability of IASI to monitor the onset of HNO<sub>3</sub> depletion at a large scale from year to  
310 year, figure 76 shows the spatial distribution of the 50 hPa drop temperatures (based on the second  
311 derivative minima of total HNO<sub>3</sub> averaged in  $1^\circ \times 1^\circ$  grid cells) inside a region delimited by a PV value  
312 of  $-8 \times 10^{-5}$  K.m<sup>2</sup>.kg<sup>-1</sup>.s<sup>-1</sup>, for each year of the IASI period. The green contour represents the PV isocontour  
313 of  $-10 \times 10^{-5}$  K.m<sup>2</sup>.kg<sup>-1</sup>.s<sup>-1</sup>, averaged over the period 10 May – 15 July for each year, which delimits our  
314 region of interest. The isocontours of 195 K for the average temperatures and the minimum temperatures,  
315 as well as the isocontour of  $-10 \times 10^{-5}$  K.m<sup>2</sup>.kg<sup>-1</sup>.s<sup>-1</sup> for the minimum PV encountered at 50 hPa over the  
316 10 May to 15 July period are also represented. The calculated drop temperatures corresponding to the  
317 onset of HNO<sub>3</sub> depletion inside the averaged PV isocontour are found to vary between ~180 and ~210  
318 K and the corresponding dates range between ~mid-May and mid-July (not shown here). Note that the  
319 high extremes in the drop temperature, which are found in some cases above eastern Antarctica, should  
320 be considered with caution: they correspond to specific regions above ice surface with emissivity  
321 features that are known to yield errors in the IASI retrievals (Hurtmans et al., 2012; Ronsmans et al.,  
322 2016). Indeed, bright land surface such as ice might in some cases lead to poor HNO<sub>3</sub> retrievals.  
323 Although wavenumber-dependent surface emissivity atlases are used in FORLI (Hurtmans et al., 2012),  
324 this parameter remains critical and causes poorer retrievals that, in some instances, pass through the  
325 series of quality filters and could affect the drop temperature calculation.

326

327 The averaged isocontour of 195 K encircles well the area of HNO<sub>3</sub> drop temperatures lower than 195 K  
328 (typically from ~187 K to ~195 K), which means that the bins inside that area characterize airmasses  
329 that experience the NAT threshold temperature during a long time over the 10 May – 15 July period.  
330 That area encompasses the inner vortex core (delimited by the isocontour of  $-10 \times 10^{-5}$  K.m<sup>2</sup>.kg<sup>-1</sup>.s<sup>-1</sup> for  
331 the PV averaged over the 10 May – 15 July period) and show pronounced minima (lower than  $-0.5 \times 10^{14}$   
332 molec.cm<sup>-2</sup>.d<sup>-2</sup>) in the second derivative of the HNO<sub>3</sub> total column with respect to time (not shown here),  
333 which indicate a strong and rapid HNO<sub>3</sub> depletion. The area enclosed between the two isocontours of  
334 195 K for the temperatures, the averaged one and the one for the minimum temperatures, show generally  
335 higher drop temperatures and weakest minima (larger than  $-0.5 \times 10^{14}$  molec.cm<sup>-2</sup>.d<sup>-2</sup>) in the second  
336 derivative of the HNO<sub>3</sub> total column (not shown). That area is also enclosed by the isocontour of  $-10 \times 10^{-5}$   
337 K.m<sup>2</sup>.kg<sup>-1</sup>.s<sup>-1</sup> for the minimum PV, meaning that the bins inside correspond, at least for one day over  
338 the 10 May – 15 July period, to airmasses located at the inner edge of the vortex and characterized by  
339 temperature lower than the NAT threshold temperature. The weakest minima in the second derivative  
340 of total HNO<sub>3</sub> (not shown) observed in that area indicate a weak and slow HNO<sub>3</sub> depletion and might be  
341 explained by a short period of the NAT threshold temperature experienced at the inner edge of the vortex.

342 It could also reflect a mixing with strong HNO<sub>3</sub>-depleted and colder airmasses from the inner vortex  
343 core. The mixing with these already depleted airmasses could also explained the higher drop  
344 temperatures detected in those bins. These high drop temperatures are generally detected later (after the  
345 HNO<sub>3</sub> depletion occurs, i.e. after the 10 May – 15 July period considered here – not shown), which  
346 supports the transport, in those bins, of earlier HNO<sub>3</sub>-depleted airmasses and the likely mixing at the  
347 edge of the vortex. Finally, these spatial variations might also partly reflect the range of maximum  
348 sensitivity of IASI to HNO<sub>3</sub>, while biases in ECMWF reanalysis are too small for explaining the spatial  
349 variation in drop temperatures. Thanks to the assimilation of an advanced Tiros Operational Vertical  
350 Sounder (ATOVS) around 1998–2000 in reanalyses, to the better coverage of satellite instruments and  
351 to the use of global navigation satellite system (GNSS) radio occultation (RO) (Schreiner et al., 2007;  
352 Wang et al., 2007; Lambert and Santee, 2018; Lawrence et al., 2018), the uncertainties have been vastly  
353 reduced. Comparisons of the ECMWF ERA Interim dataset used in this work with the COSMIC data  
354 (Lambert and Santee, 2018) found a small warm bias, with median differences around 0.5 K, reaching  
355 0–0.25 K in the southernmost regions of the globe at ~68–21 hPa where PSCs form.

356  
357 Except above some parts of Antarctica which are prone to larger retrieval errors, the overall range in the  
358 drop 50 hPa temperature for total HNO<sub>3</sub> inside the isocontour for the averaged temperature of 195 K,  
359 typically extends from ~187 K to ~195 K, which falls within the range of PSCs nucleation temperature  
360 at 50 hPa: from slightly below  $T_{NAT}$  to around 3-4 K below the ice frost point -  $T_{ice}$  - depending on  
361 atmospheric conditions, on TTE and on the type of formation mechanisms (Pitts et al., 2011; Peter and  
362 GrooB, 2012; Hoyle et al., 2013). This underlines well the benefit of the excellent spatial and temporal  
363 coverage of IASI that allows capturing the rapid and critical depletion phase over a large scale.

## 364 365 **5 Conclusions**

366  
367 In this paper, we have explored the added value of the dense HNO<sub>3</sub> total columns dataset provided by  
368 the IASI/Metop satellite over a full decade (2008–2017) for monitoring the stratospheric depletion phase  
369 that occurs each year in the S.H. and for investigating its relationship to the NAT formation temperature.  
370 To that end, we focused on and delimited the coldest polar region of the S.H. using a specific PV value  
371 at 530 K (~50 hPa, PV of  $-10 \times 10^{-5} \text{ K} \cdot \text{m}^2 \cdot \text{kg}^{-1} \cdot \text{s}^{-1}$ ) and stratospheric temperatures at 50 hPa, taken from  
372 the ECMWF ERA Interim reanalysis. That single representative pressure level has been considered in  
373 this study given the maximum sensitivity of IASI to HNO<sub>3</sub> around that level over a range where the  
374 PSCs formation/denitrification process occur.

375  
376 The annual cycle of total HNO<sub>3</sub>, as observed from IASI, has first been characterized according to the  
377 temperature evolution. Three various regimes (R1 to R3) in the total HNO<sub>3</sub> - 50 hPa temperature  
378 relationship were highlighted from the time series over the S.H. polar region and described along the  
379 cycle: R1 is defined at play during April and May and characterized by a rapid decrease in 50 hPa  
380 temperatures while HNO<sub>3</sub> accumulates in the poles; R2, from June to September, shows the onset of the  
381 depletion when the 50 hPa temperatures fall below 195 K (considered here as the onset of PSCs  
382 nucleation phase at that level), with a strong consistency between each year; R3, defined from November  
383 until March when total HNO<sub>3</sub> remains at low R2 plateau levels, despite the return of sunlight and heat,  
384 characterizes the strong denitrification of the stratosphere, likely due to PSCs sedimentation at lower  
385 levels where the IASI sensitivity is low. For each year over the IASI period, the use of the second  
386 derivative of the HNO<sub>3</sub> column versus time was then found particularly valuable to detect the onset of  
387 the HNO<sub>3</sub> condensation to PSCs. It is captured, on average from IASI, a few days before June with a  
388 delay of 4–23 days after the maximum in total HNO<sub>3</sub>. Except for the year 2014, the corresponding  
389 temperatures ('drop temperatures') were detected between 189.2 K and 198.6 K ( $194.1 \pm 2.8$  on average  
390 over the 10 years, excluding the year 2014 that shows a drop temperature of 202.8 K). The corresponding



391 temperatures ('drop temperatures') were detected between 189.2 K and 202.8 K ( $194.2 \pm 3.8$  on average  
392 ~~over the 10 years~~), which tends to demonstrated the good consistency between the 50 hPa drop  
393 temperature and the PSCs formation temperatures in that altitude region. Finally, the annual and spatial  
394 variability (within  $1^\circ \times 1^\circ$ ) in the drop temperature was further explored from IASI total HNO<sub>3</sub>. Inside  
395 the isocontours of 195 K for the average temperatures and of  $-10 \times 10^{-5}$  K.m<sup>2</sup>.kg<sup>-1</sup>.s<sup>-1</sup> for the averaged PV  
396 at 50 hPa, the drop temperatures are detected between ~mid-May and mid-July, typically range between  
397 ~187 K to ~195 K and are associated with the highest minima (lower than  $-0.5 \times 10^{14}$  molec.cm<sup>-2</sup>.d<sup>-2</sup>) in  
398 the second derivative of the HNO<sub>3</sub> total column with respect to time, indicating a strong and rapid HNO<sub>3</sub>  
399 depletion. Except for extreme drop temperatures (~210 K) that were found from year to year above  
400 eastern Antarctica and suspected to result from unfiltered poor quality retrievals in case of emissivity  
401 issues above ice, the range of drop temperatures is interestingly found in line with the PSCs nucleation  
402 temperature that is known, from previous studies, to strongly depend on a series a factors (e.g.  
403 meteorological conditions, HNO<sub>3</sub> vapour pressure, temperature threshold exposure, presence of  
404 meteoritic dust). At the edge of the vortex, considering the isocontours of 195 K for the minimum  
405 temperatures or of  $-10 \times 10^{-5}$  K.m<sup>2</sup>.kg<sup>-1</sup>.s<sup>-1</sup> for the minimum PV, higher and later drop temperatures along  
406 with weakest minima in the second derivative of the HNO<sub>3</sub> total column with respect to time, indicating  
407 a slow HNO<sub>3</sub> depletion, are found. It likely results from a short temperature threshold exposure or a  
408 mixing with already depleted airmasses from the inner vortex core. The results of this study highlight  
409 the ability of IASI to measure the variations in total HNO<sub>3</sub> and, in particular, to capture and monitor the  
410 rapid depletion phase over the whole polar regions.

411  
412 We show in this study that the IASI dataset allows capturing the variability of stratospheric HNO<sub>3</sub>  
413 throughout the year (including the polar night) in the Antarctic. In that respect, it offers a new  
414 observational means to monitor the relation of HNO<sub>3</sub> to temperature and the related formation of PSCs.  
415 Despite the limited vertical resolution of IASI which does not allow to investigate the HNO<sub>3</sub> uptake by  
416 the different types of PSCs during their formation and growth along the vertical profile, the HNO<sub>3</sub> total  
417 column measurements from IASI constitute an important new dataset for exploring the strong polar  
418 depletion over the whole stratosphere. This is particularly relevant considering the mission continuity,  
419 which will span several decades with the planned follow-on missions. Indeed, thanks to the three  
420 successive instruments (IASI-A launched in 2006 and still operating, IASI-B in 2012, and IASI-C in  
421 2018) that demonstrate an excellent stability of the Level-1 radiances, the measurements will soon  
422 provide an unprecedented long-term dataset of HNO<sub>3</sub> total columns. Further work could also make use  
423 of this unique data set to investigate the relation between HNO<sub>3</sub>, O<sub>3</sub>, and meteorology in the changing  
424 climate.

#### 425 426 **Data availability**

427 The IASI HNO<sub>3</sub> data processed with FORLI-HNO<sub>3</sub> v0151001 are available upon request to the  
428 corresponding author.

#### 429 430 **Author contributions**

431 G.R. performed the analysis, wrote the manuscript and prepared the figures. C.W. and L.C. contributed  
432 to the analysis. C.W., S.S., P.-F. C. and L.C. contributed to the interpretation of the results. D.H. was  
433 responsible for the retrieval algorithm development and the processing of the IASI HNO<sub>3</sub> dataset. All  
434 authors contributed to the writing of the text and reviewed the manuscript.

#### 435 436 **Competing interests**

437 The authors declare no competing interests.

#### 438 439 **Acknowledgements**

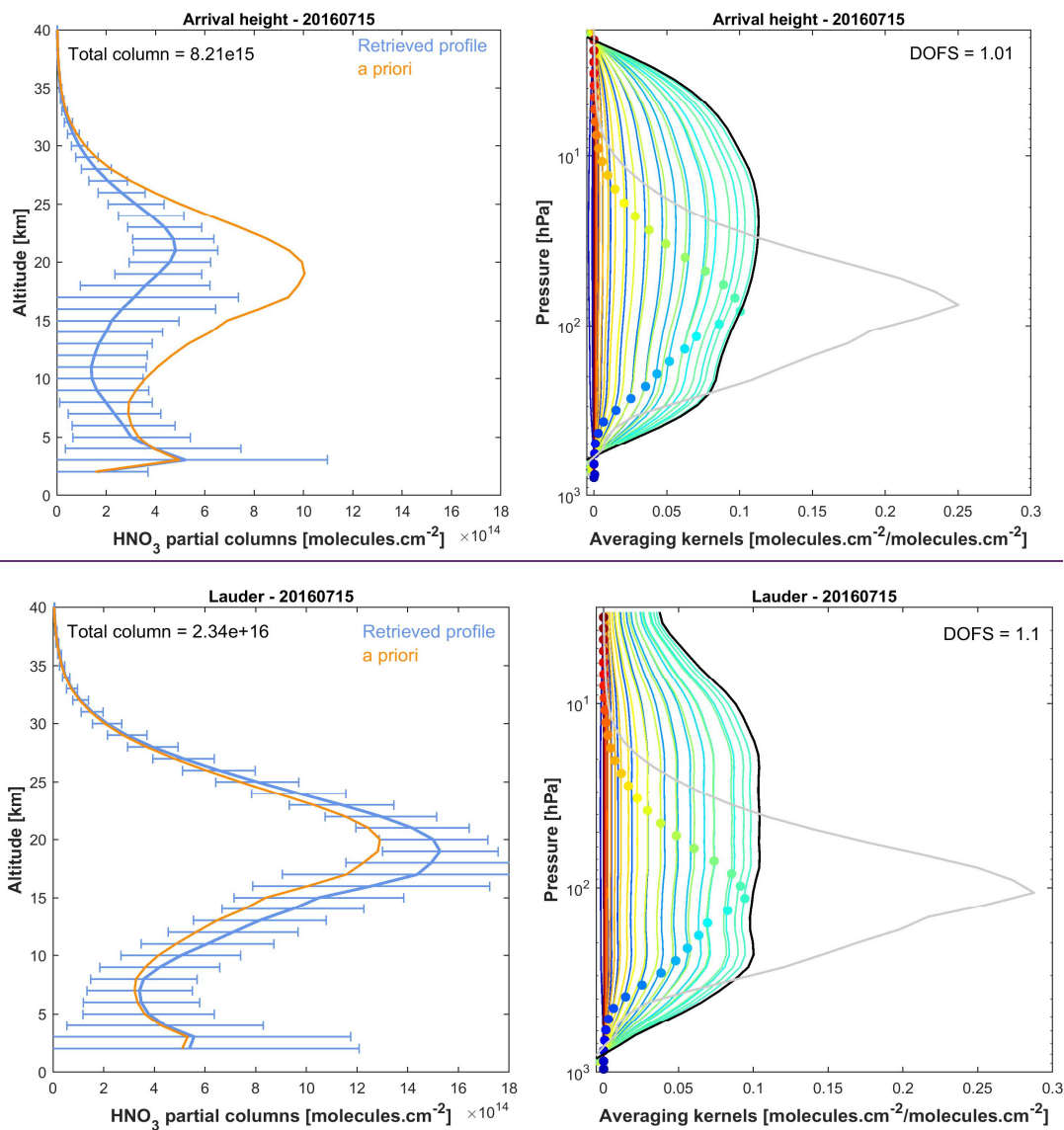
440 IASI has been developed and built under the responsibility of the Centre National d'Etudes Spatiales  
441 (CNES, France). It is flown on board the Metop satellites as part of the EUMETSAT Polar System. The  
442 IASIL1 data are received through the EUMETCast near-real-time data distribution service. The research  
443 was funded by the F.R.S.-FNRS, the Belgian State Federal Office for Scientific, Technical and Cultural  
444 Affairs (Prodex arrangement 4000111403 IASI.FLOW) and EUMETSAT through the Satellite  
445 Application Facility on Atmospheric Composition Monitoring (ACSAF). G. Ronsmans is grateful to the  
446 Fonds pour la Formation à la Recherche dans l'Industrie et dans l'Agriculture of Belgium for a PhD  
447 grant (Boursier FRIA). L. Clarisse is a research associate supported by the F.R.S.-FNRS. C. Clerbaux is  
448 grateful to CNES for financial support. S. Solomon is supported by the National Science Foundation  
449 (NSF-1539972).

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489 **Figure captions**

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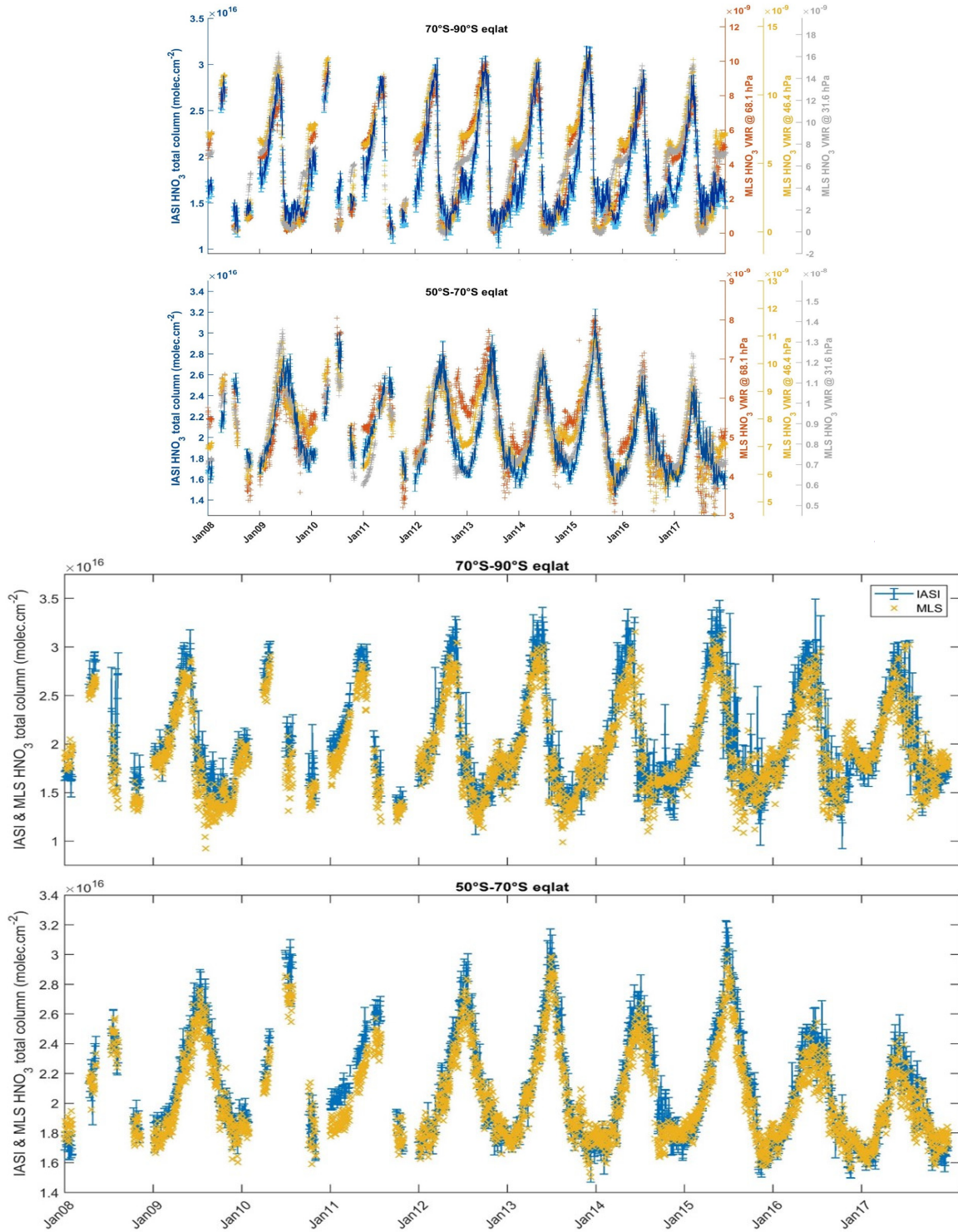
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Figure 1. Examples of IASI HNO<sub>3</sub> vertical profiles (in molec.cm<sup>-2</sup>) with corresponding averaging kernels (in molec.cm<sup>-2</sup>/molec.cm<sup>-2</sup>; with the total column averaging kernels (black) and the sensitivity profiles (grey) (divided by 10) above Arrival Height (77.49°S, 166.39°E; top panels) and Lauder (45.03°S, 169.40°E; bottom panels). The error bars associated with the HNO<sub>3</sub> vertical profile represent the total retrieval error. The a priori profile is also represented. The total column and the DOFS values are indicated.

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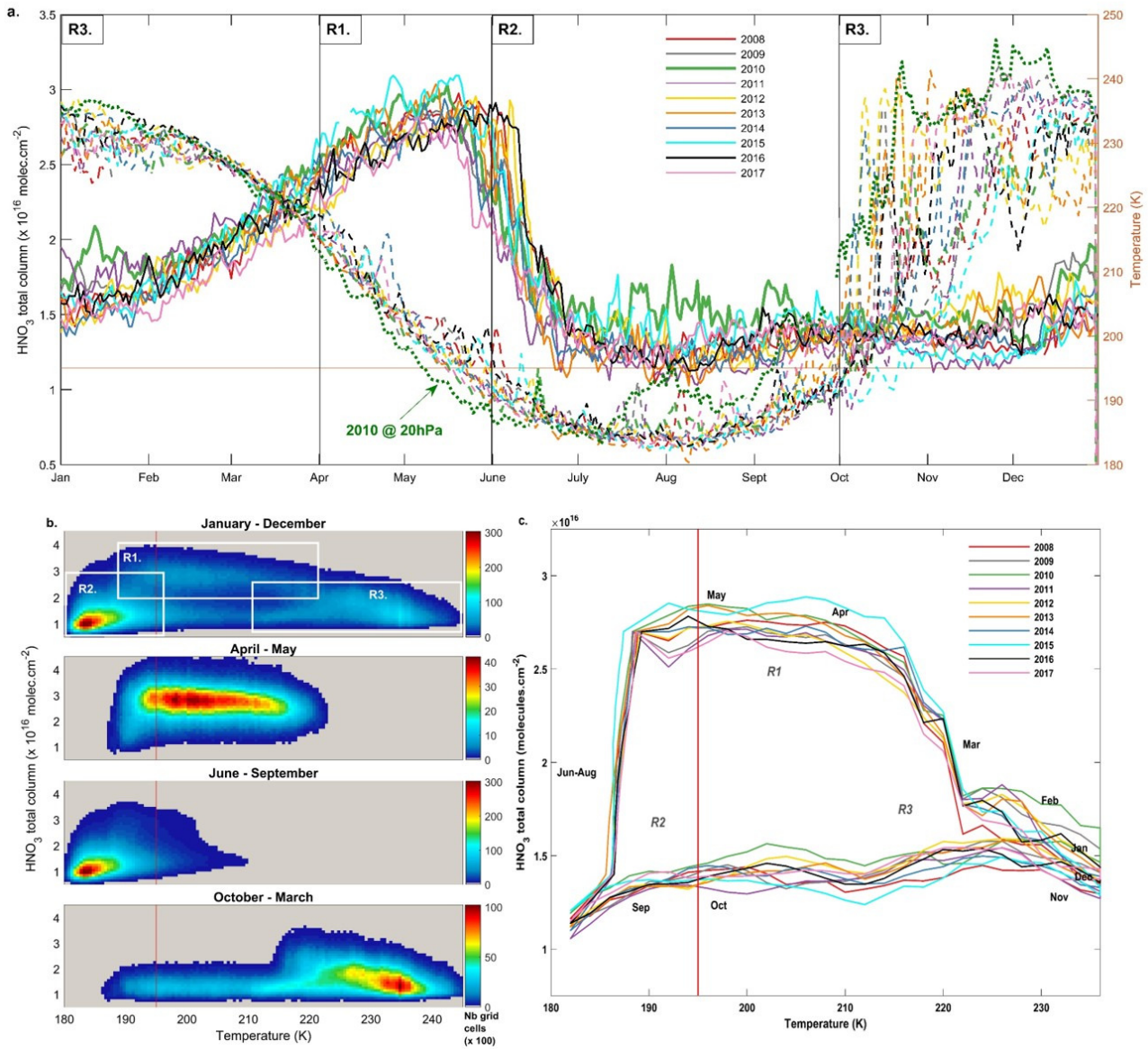
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**Figure 12.** Time series of daily IASI total  $\text{HNO}_3$  column (blue) co-located with MLS and of MLS total  $\text{HNO}_3$  columns (orange) within  $2.5 \times 2.5$  grid boxes, averaged in the  $70^\circ\text{S}$ – $90^\circ\text{S}$  (top panel), the  $50^\circ\text{S}$ – $70^\circ\text{S}$  (middle panel) and the  $30^\circ\text{S}$ – $50^\circ\text{S}$  (bottom panel) equivalent latitude bands. Time series of daily IASI total  $\text{HNO}_3$  column (blue, left y axis) co-located with MLS and of MLS VMR  $\text{HNO}_3$  within  $2.5 \times 2.5$  grid boxes at three pressure levels (at 30, 50 and 70 hPa; right y axis), averaged in the  $70^\circ\text{S}$ – $90^\circ\text{S}$  (top panel) and in the  $50^\circ\text{S}$ – $70^\circ\text{S}$  (bottom panel) equivalent latitude bands. The error bars (light-blue) represents  $3\sigma$ , where  $\sigma$  is the standard deviation around the IASI  $\text{HNO}_3$  daily average.

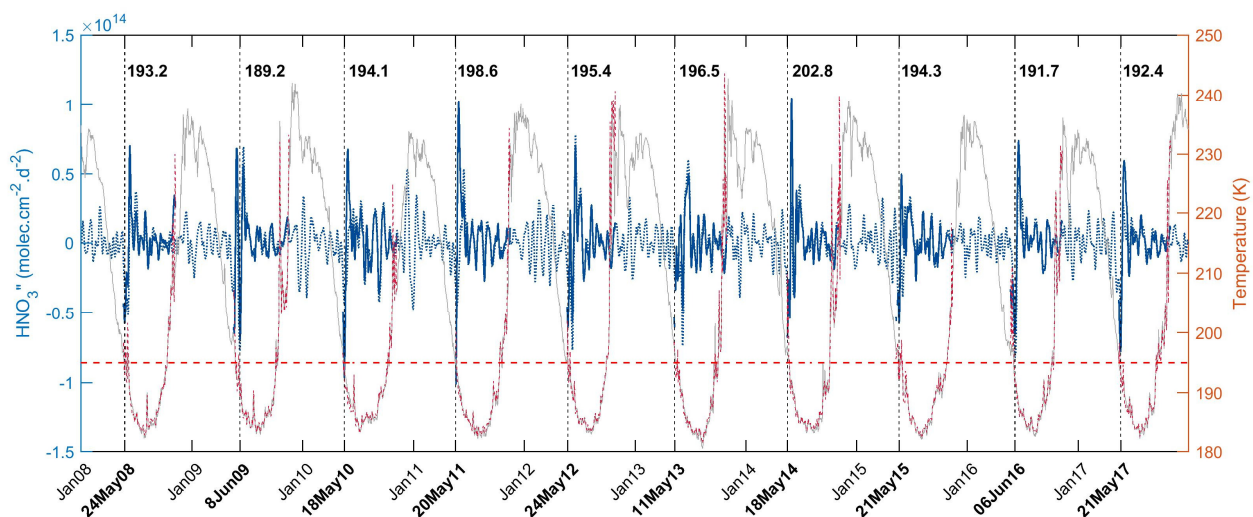
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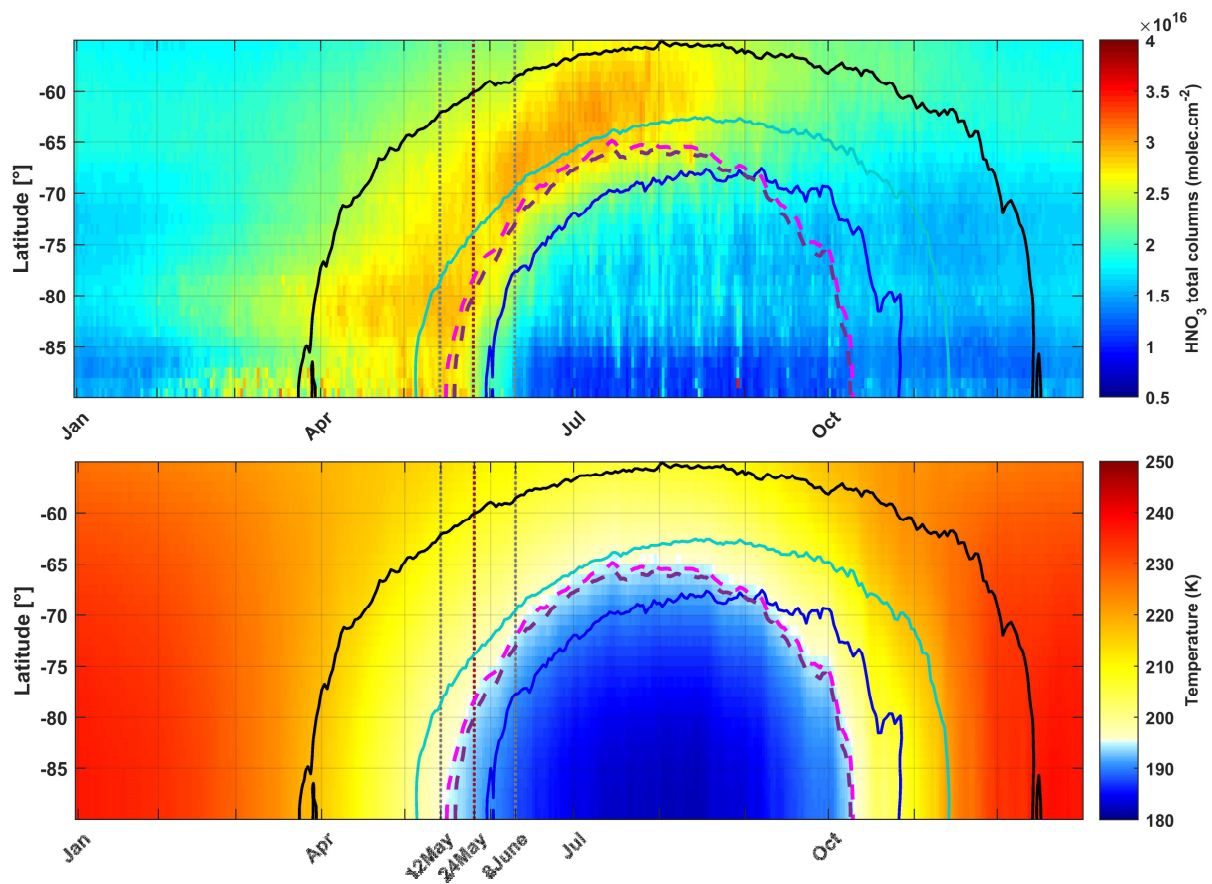
**Figure 23.** (a) Time series of daily averaged HNO<sub>3</sub> total columns (solid lines) and temperatures taken at 50 hPa (dashed lines) in the 70° - 90° S equivalent latitude band, for the years 2008 – 2017. The green dotted line represents the temperatures at 20 hPa for the year 2010. (b) HNO<sub>3</sub> total columns versus temperatures (at 50 hPa) histogram for the whole year (top) and for the 3 defined regimes (R1 - R3) separated in (a) for the year 2011. The colors refer to the number of gridded measurements in each cell. (c) Evolution of daily averaged HNO<sub>3</sub> total columns with the highest occurrence (in bins of 0.1×10<sup>16</sup> molec.cm<sup>-2</sup> and 2 K) as a function of the 50 hPa temperature for the years 2008 – 2017. The red horizontal or vertical lines represent the 195 K threshold temperature.

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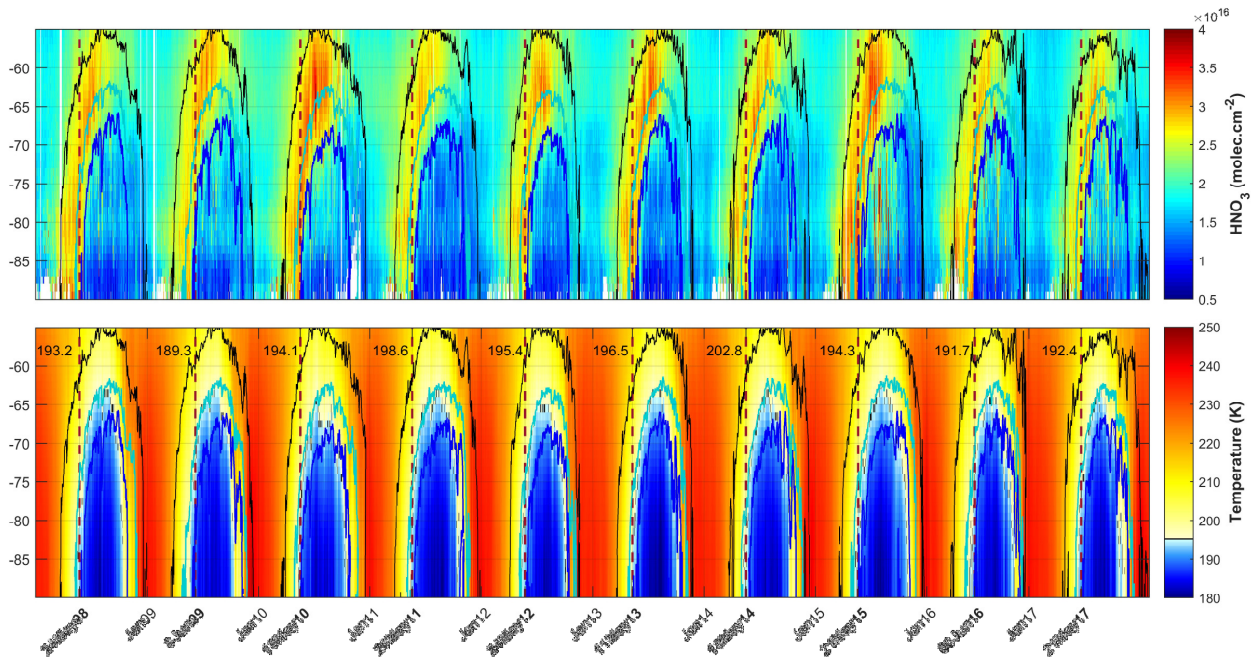
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**Figure 34.** Time series of total HNO<sub>3</sub> second derivative (blue, left y-axis) and of the temperature (red, right y-axis), in the region of potential vorticity at 530 K lower than  $-10 \times 10^{-5} \text{ K.m}^2.\text{kg}^{-1}.\text{s}^{-1}$ . The red horizontal line corresponds to the 195 K temperature. The vertical dashed lines indicate the second derivative minimum in HNO<sub>3</sub> for each year. The corresponding dates (in bold, on the x-axis) and temperatures are also indicated. The time series of total HNO<sub>3</sub> second derivative (dashed blue) and of temperature (grey) in the 70° – 90° S Eqlat band are also represented.



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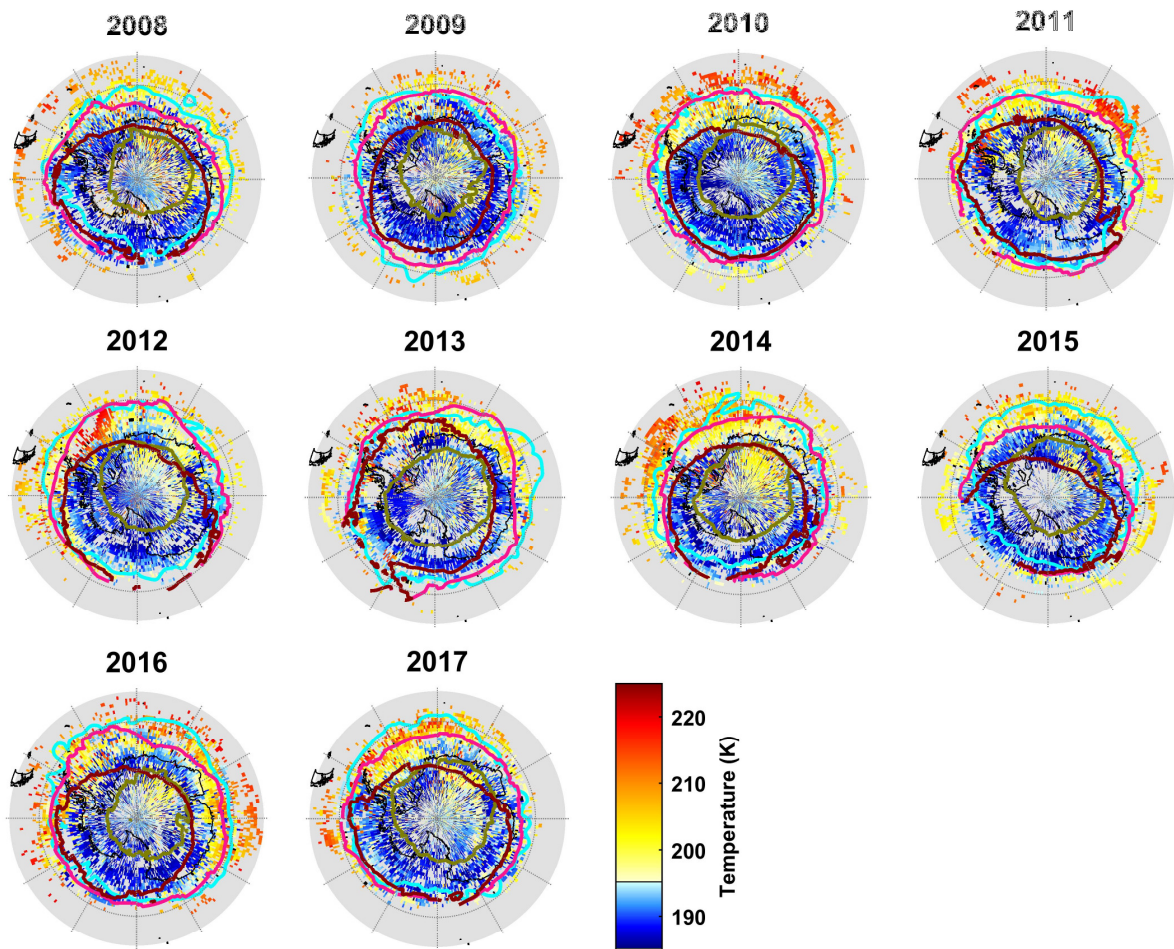
**Figure 45.** Zonal distributions of (a) HNO<sub>3</sub> total columns (in molec.cm<sup>-2</sup>) from IASI and (b) temperatures at 50 hPa from ERA Interim (in K) between 55° to 90° south and averaged over the years 2008 – 2017. Three isocontours for PV of -5 (black), -8 (cyan) and -10 (blue) ( $\times 10^{-5}$  K.m<sup>2</sup>.kg<sup>-1</sup>.s<sup>-1</sup>) at 530 K, the isocontours for the 195 K temperature (pink) and for the averaged 194.2 K drop temperature (purple) at 50 hPa are superimposed. The vertical grey dashed lines encompass the period of the second derivative minima and the red one indicates the average date for the drop temperatures calculated in the area delimited by a  $-10 \times 10^{-5}$  K.m<sup>2</sup>.kg<sup>-1</sup>.s<sup>-1</sup> PV contour.



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**Figure 56.** Zonally averaged distributions of (top) HNO<sub>3</sub> total columns (in molec.cm<sup>-2</sup>) from IASI and (bottom) temperatures at 50 hPa from ERA Interim (in K). The latitude range is from 55° to 90° south and the isocontours are PVs of -5 (black), -8 (cyan) and -10 (blue) (× 10<sup>-5</sup> K.m<sup>2</sup>.kg<sup>-1</sup>.s<sup>-1</sup> at 530 K). The vertical red dashed lines correspond to the second derivative minima each year in the area delimited by a -10×10<sup>-5</sup> K.m<sup>2</sup>.kg<sup>-1</sup>.s<sup>-1</sup> PV contour.





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**Figure 67.** Spatial distribution ( $1^\circ \times 1^\circ$ ) of the drop temperature at 50 hPa (K) (calculated from the total  $\text{HNO}_3$  second derivative minima) for each year of IASI (2008–2017), in a region defined by a PV of  $-8 \times 10^{-5} \text{ K.m}^2.\text{kg}^{-1}.\text{s}^{-1}$ . The isocontours of  $-10 \times 10^{-5} \text{ K.m}^2.\text{kg}^{-1}.\text{s}^{-1}$  at 530 K for the averaged PV (in green) and the minimum PV (in cyan) encountered over the period 10 May –15 June for each year and the isocontours of 195 K at 50 hPa for the averaged (in red) and the minimum (in pink) temperatures over the same period are represented.

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