



1 **Global dust cycle and uncertainty in CMIP5 models**

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7

8 **Abstract**

9 Dust cycle is an important component of the Earth system and have been
10 implemented into climate models and Earth System Models (ESMs). An
11 assessment of the dust cycle in these models is vital to address the strengths and
12 weaknesses of these models in simulating dust aerosol and its interactions with the
13 Earth system and enhance the future model developments. This study presents a
14 comprehensive evaluation of global dust cycle in 15 models participating in the
15 fifth phase of the Coupled Model Intercomparison Project (CMIP5). The various
16 models are compared with each other and with an aerosol reanalysis as well as
17 station observations of dust deposition and concentrations. The results show that
18 the global dust emission in these models ranges from 735 to 8186 Tg yr⁻¹ and the
19 annual mean dust burden ranges from 2.5 to 41.9 Tg, both of which scatter by a
20 factor of about 10-20. The models generally agree with each other and
21 observations in reproducing the “dust belt” that extends from North Africa, Middle
22 East, Central and South Asia, to East Asia, although they differ largely in the



23 spatial extent of this dust belt. The models also differ in other dust source regions
24 such as North America and Australia, where the contributions of these sources to
25 global dust emissions vary by a factor of more than 500. We suggest that the
26 coupling of dust emission with dynamic vegetation can enlarge the range of
27 simulated dust emission.

28 For the removal process, all the models estimate that wet deposition is a
29 smaller sink than dry deposition and wet deposition accounts for 12-39 % of total
30 deposition. The models also estimate that most (77-91 %) of dust particles are
31 deposited onto continents and 9-23 % of them are deposited into oceans. A linear
32 relationship between dust burden, lifetime, and fraction of wet deposition to total
33 deposition from these models suggests a general consistency among the models.
34 Compared to the observations, most models reproduce the dust deposition and dust
35 concentrations within a factor of 10 at most stations, but larger biases by more
36 than a factor of 10 are also noted at specific regions and for certain models. These
37 results cast a doubt on the interpretation of the simulations of dust-affected fields
38 in climate models and highlight the need for further improvements of dust cycle
39 especially on dust emission in climate models.

40

41



42 **1. Introduction**

43 Dust cycle is an important component of the Earth system as it has strong impacts
44 on the Earth environment and climate system (Shao et al., 2011). Dust aerosol in the
45 atmosphere significantly impacts the climate systems via various pathways, such as
46 scattering and absorbing the solar and terrestrial radiation, modifying cloud radiative
47 forcing by acting as cloud condensation nuclei and ice nucleating particles, and reducing
48 the snow albedo when depositing onto snow (Boucher et al., 2013; Forster et al., 2007;
49 Liu, et al., 2012a; Mahowald et al., 2011; Wu et al., 2018a; Rahimi et al., 2019). Dust
50 affects the biogeochemical cycle by delivering the nutrients (e.g., mineral, nitrogen, and
51 phosphorus) from dust sources to the oceans/other continents (Jickells et al., 2005;
52 Mahowald et al., 2011). Dust aerosol is also one of the main contributors to air pollution
53 that is hazardous to human health (Bell et al., 2008; Lin et al., 2012).

54 To quantify the dust impacts on Earth system, dust cycle including dust emission,
55 transport, and dry and wet deposition has been incorporated in climate models and Earth
56 System Models (ESMs) since 1990s. These models have the capability to reproduce the
57 general patterns of global dust distribution (e.g., Ginoux et al., 2001; Zender et al., 2003;
58 Yue et al., 2009; Huneus et al., 2011; Liu et al., 2012b). However, large uncertainties
59 still exist in the simulated global dust budgets in these models, as revealed by a wide
60 range of model results. A comparison of 14 different models from the Aerosol
61 Comparison between Observations and Models (AeroCom) Phase I showed the estimated
62 global dust emission ranges from 514 to 4313 Tg yr⁻¹ and annual mean dust burden from
63 6.8 to 29.5 Tg (Huneus et al., 2011). Compared to the observations, these models from
64 AeroCom Phase I produce the dust deposition and surface concentration mostly within a



65 factor of 10 (Huneeus et al., 2011). Uncertainties of dust cycle have led to difficulty in
66 the interpretation of climate impacts of dust aerosol (Yue et al., 2010; Forster et al., 2007;
67 Boucher et al., 2013).

68 The Coupled Model Intercomparison Project Phase 5 (CMIP5) provides a
69 comprehensive dataset of meteorological variables and climate forcing agents such as
70 aerosols including dust during the period of 1850s to 2000s from a variety of climate
71 models and ESMs. Dust cycle is interactively calculated in some CMIP5 models for
72 historical climate simulations and future climate projections. Till now, only a few studies
73 have investigated dust simulations in CMIP5. Evan et al. (2014) evaluated African dust in
74 23 CMIP5 models and found the models underestimate dust emission, deposition, and
75 aerosol optical depth (AOD) and have low ability in reproducing the interannual
76 variations of dust burden. Pu and Ginoux (2018) compared the dust optical depth (DOD)
77 from 7 CMIP5 models with satellite observations from 2004 to 2016. They found that
78 these models can capture the global spatial patterns of DOD but with an underestimation
79 of DOD by 25.2% in the boreal spring, and some models cannot capture the seasonal
80 variations of DOD in several key regions such as Northern China and Australia. Wu et al.
81 (2018b) evaluated the dust emission in East Asia from 15 CMIP5 models and found that
82 none of the models can reproduce the observed decline trend of dust event frequency
83 from 1961 to 2005 over East Asia.

84 None of the above studies has investigated the global dust cycles including their
85 sources and sinks in the CMIP5 models. Therefore, this study is aimed at filling the gap
86 by presenting the strengths and weaknesses of CMIP5 models in simulating global dust
87 cycles. This study will also investigate the associated model uncertainties. As there are a



88 variety of complexities in the CMIP5 models (Flato et al., 2013), this study aims at
89 identifying the difference in simulated dust cycle as a result of these different
90 complexities. Of particular interest is that some models couple dust emission with
91 dynamic vegetation while the others calculate dust emission based on prescribed
92 vegetation conditions (Table 1), and thus the impacts of dynamic vegetation on dust
93 emission can be examined by comparing the results from these two group models, which
94 has been rarely studied previously.

95 The paper is organized as follows. Section 2 introduces the CMIP5 models,
96 including the dust emission parameterization. Section 3 describes the observation data
97 used for model validation. Section 4 presents the global dust budget and dust emission,
98 followed by evaluations of dust deposition flux and dust concentration with observations.
99 Discussion and conclusions are given in section 5.

100

101 **2. Model data**

102 Here we use the historical simulations from 15 CMIP5 models (Table 1). All the 15
103 models are fully-coupled models used for historical climate simulations and future
104 climate projections, which are included in the Fifth Assessment Report of
105 Intergovernmental Panel on Climate Change (Flato et al., 2013). A brief description of
106 these model is given in Table 1 and more detailed information can be found in the
107 references as listed.

108 An essential part of dust cycle is dust emission. The dust emission schemes used in
109 these models and the references are also listed in Table 1. Here we only provide a brief
110 summary of similarities and differences in these dust emission schemes. More details can



111 be found in the references (Cakmur et al., 2006; Ginoux et al., 2001, 2004; Marticorena
112 & Bergametti, 1995; Miller et al., 2006; Shao et al., 1996; Takemura et al., 2000, 2009;
113 Tanaka & Chiba, 2005, 2006; Woodward, 2001, 2011; Zender et al., 2003). In general,
114 these emission schemes similarly calculate dust emission based on near-surface wind
115 velocity (in terms of friction wind velocity or wind velocity at 10 m), soil wetness and
116 vegetation cover, and they mainly differ in how to account for these factors and
117 associated input parameters. Particularly, dust emission scheme is coupled to dynamic
118 vegetation in 5 models (GFDL-CM3, HadGEM2-CC, HadGEM2-ES, MIROC-ESM,
119 MIROC-ESM-CHEM). These models use prognostic vegetation to determine the dust
120 source regions. This introduces additional degrees of freedom and thus increases the
121 difficulty in simulating dust emission in these models compared to other models with
122 prescribed vegetation that is constructed from the observation. This will be discussed in
123 Section 4.

124 Another difference in dust emission scheme is the treatment of dust sizes including
125 the size range and mass partitioning in different sizes. 7 models (GFDL-CM3, MIROC4h,
126 MIROC5, MIROC-ESM, MIROC-ESM-CHEM, MRI-CGCM3, MRI-ESM1) have the
127 same dust size range of 0.2-20 μm in diameter. 5 of the other eight models (CanESM2,
128 CESM1-CAM5, CSIRO-Mk3-6-0, GISS-E2-H, GISS-E2-R) have smaller size ranges
129 (listed in Table 1), while the remaining 3 models (ACCESS1-0, HadGEM2-CC,
130 HadGEM2-ES) have the larger size range of 0.0632-63.2 μm . The impacts of dust size
131 distribution on the simulation of dust cycle will be discussed in later sections. However,
132 as only the total dust emission, deposition, and concentration are provided, we are unable
133 to investigate the difference in the mass partitioning among different dust sizes and its



134 evolution, which will be left for future studies.

135 Note that we select these models because they calculate dust emission interactively
136 by their dust emission schemes implemented, and meanwhile, model output of dust
137 emission flux and dust concentration are available from the CMIP5 archive. Also note
138 that not all the models have both dry and wet deposition archived and 8 models provide
139 only dry (GFDL-CM3) or wet deposition flux (CSIRO-Mk3-6-0, HadGEM2-CC,
140 HadGEM2-ES, MIROC4h, MIROC5, MIROC-ESM, MIROC-ESM-CHEM). Therefore,
141 for dust deposition, we derive the global total amount of dry (wet) deposition by
142 subtracting wet (dry) deposition from emission if only wet (dry) deposition is available.
143 For comparison with station observations, we will only use seven models with both dry
144 and wet deposition provided. If there are multiple ensemble simulations available for a
145 specific model, we will use the ensemble means from these simulations for this model
146 (Table 1). The historical simulations of CMIP5 cover the period of 1850-2005. However,
147 some model results prior to 1960 or 1950 are not provided in the CMIP5 archive (e.g.,
148 ensemble #2 and #3 from HadGEM2-CC prior to 1960 is not available; MIROC4h prior
149 to 1950 is not available). Therefore, we will focus on the period of 1960-2005 to include
150 as many models as possible and to include as many years as possible for the analysis of
151 present-day dust cycle.

152 To evaluate the CMIP5 model results, we also use the Modern-Era Retrospective
153 Analysis for Research and Applications, version 2 (MERRA-2). MERRA-2 is the latest
154 atmospheric reanalysis produced by NASA's Global Modeling and Assimilation Office
155 (Gelaro et al., 2017). MERRA-2 assimilates more observation types and have improved
156 significantly compared to its processor, MERRA. A major advancement of MERRA-2 is



157 that it includes the assimilation of AOD (Randles et al., 2017), which is not included in
158 MERRA and other commonly-used reanalysis datasets such as ECWMF Reanalysis
159 (ERA5) and NCEP/DOE Reanalysis II (R2). The aerosol fields (including dust) in
160 MERRA-2 are significantly improved compared to an identical control simulation that
161 does not include the AOD assimilation (Randles et al., 2017; Buchard et al., 2017). It
162 should be noted that as only AOD is taken into account in the aerosol assimilation, there
163 may be discrepancies in the related aerosol fields such as aerosol concentration and
164 deposition. In addition, dust emission is calculated directly from surface wind speed and
165 soil wetness based on the dust emission scheme of Ginoux et al. (2001), and there is no
166 direct impact on emission from aerosol assimilation. Therefore, there may be
167 inconsistency between dust emission, burden, and deposition. In fact, as shown in the
168 Section 4, there is imbalance between total dust emission and deposition globally and
169 adjustment of dust emission to fit the dust burden is still needed. Despite the limitation,
170 MERRA-2 provides a well-constrained global dust dataset, which is very useful for
171 model evaluations. We will use MERRA-2 as a referential data but with the knowledge
172 of its limitation. We will use the long-term means of dust-related variables during the
173 whole period when data is available (i.e., 1980-2018). Dust in MERRA-2 is treated by
174 five size bins spanning from 0.2 to 20 μm , which are summed to provide the total values.
175 MERRA-2 is provided at the resolution of $0.5^\circ \times 0.625^\circ$, which is similar to one CMIP5
176 model (MIROC4h) and finer than other CMIP5 models.

177

178 **3. Observations**



179 There are limited observational datasets that can be used for model evaluations.
180 There is no direct observation of dust emission flux, but satellite observations can provide
181 the locations of dust source regions where dust appears most frequently (e.g., Prospero et
182 al., 2002; Ginoux et al., 2012). Here we do not directly use these observations as they are
183 not available for our usage, but we will refer to the dust source map based on satellite
184 observations from previous studies (e.g., Prospero et al., 2002; Ginoux et al., 2012) and
185 qualitatively compare simulated dust emission regions with them.

186 Dust deposition is an important constraint on the global dust budget. Here we use
187 the dust deposition flux at 84 stations across the globe available from the AeroCom
188 project (Huneeus et al., 2011). The dataset is compiled from the Dust Indicators and
189 Records in Terrestrial and Marine Paleoenvironments (DIRTMAP) database (Kohfeld
190 and Harrison, 2001) and the data of Ginoux et al., (2001) and Mahowald et al. (1999,
191 2009). Dust deposition flux are recorded over a period of several to hundreds of years at
192 these stations. There are two types of deposition, dry deposition and wet deposition. To
193 evaluate the contribution of wet deposition to total deposition, we also use the fraction of
194 wet deposition to total deposition at 10 stations, which is compiled by Mahowald et al.
195 (2011). The fraction of wet deposition is obtained from the observations over several
196 years. Note as only minimum and maximum values of fraction of wet deposition are
197 provided for some stations, the average of the minimum and maximum values will be
198 plotted with the range provided when compared with the simulations.

199 Dust concentration is a key variable that reflects both the dust emission and
200 transport. We use the monthly surface dust concentrations at 20 sites managed by the
201 Rosenstiel School of Marine and Atmospheric Science at the University of Miami



202 (Prospero, 1996). We also use the monthly surface dust concentrations measured at 2
203 other stations: Rukomechi, Zimbabwe (Maenhaut et al., 2000a; Nyanganyura et al., 2007)
204 and Jabiru, Australia (Maenhaut et al., 2000b; Vanderzalm et al., 2003). In total, there are
205 22 stations globally. These stations are generally located in the downwind of dust source
206 regions and some of them are located in the remote regions (Table 2; Figure 1).
207 Measurements at these stations are taken over a period of two to tens of years. This
208 dataset has been widely used to evaluate global dust models (e.g., Ginoux et al., 2001;
209 Zender et al., 2003; Liu et al., 2012b) and also included in the AeroCom project
210 (Huneeus et al., 2011).

211 We consider the dataset above as a climatology although some of them did not cover
212 a long enough period such as tens of years. The distribution of these stations (for dust
213 deposition, fraction of wet deposition, surface dust concentration) are shown in Figure 1.
214 To compare model results with station observations, bi-linear interpolation is used to
215 generate the model results at the stations.

216

217 **4. Results**

218 **4.1 Global dust budget**

219 First, we present the global dust budgets in CMIP5 models. Table 3 lists the global
220 dust emission, wet deposition, burden, and lifetime in all the 15 models. The area fraction
221 of global dust emissions and ratio of wet deposition to total deposition are also given.
222 Overall, the models estimate the global dust emission in the range of 735-8196 Tg yr⁻¹,
223 with the MIROC4h having the lowest and two Hadley models (HadGEM2-CC and
224 HadGEM2-ES) having the highest emissions. The global dust emissions in CMIP5



225 models differ by about 11 times compared to about 8 times in the AeroCom models,
226 which give dust emissions in the range of 514-4313 Tg yr⁻¹ (Huneus et al., 2011). This
227 can be ascribed to a larger difference in the complexity of CMIP5 models compared to
228 AeroCom models (Section 2). In particular, HadGEM2-CC and HadGEM2-ES give about
229 twice of the largest emission estimated in the AeroCom models. The larger value in
230 HadGEM2-CC and HadGEM2-ES is mainly due to the overestimation of bare soil area
231 by the dynamic vegetation module in these models (Collins et al., 2011; Martin et al.,
232 2011). Additionally, the larger value may be also related to the larger dust size range in
233 the models (0.06 to 63 μm) with about 3300 Tg yr⁻¹ of dust emission for particles smaller
234 than 20 μm diameter (Bellouin et al, 2011). However, ACCESS1.0 with the same size
235 range as HadGEM2-CC and HadGEM2-ES produces 3-4 times smaller dust mission. As
236 shown in the evaluation of surface dust concentrations in Section 4.4, HadGEM2-CC and
237 HadGEM2-ES consistently overestimate the surface dust concentrations at the selected
238 stations (by 5 times on average). The MIROC4h model underestimates the surface dust
239 concentrations by more than 10 times (Section 4.4). If the estimations of MIROC4h,
240 HadGEM2-CC and HadGEM2-ES are not considered, global dust emissions in CMIP5
241 models are in the range of 1246-3698 Tg yr⁻¹, comparable to AeroCom results (Huneus
242 et al., 2011) and other estimations (e.g., Shao et al., 2011). The global dust emission in
243 MERRA-2 is 1620 Tg yr⁻¹, which is within the range of CMIP5 models.

244 For dust deposition, dust particles are deposited to the Earth's surface mainly by dry
245 deposition, and wet deposition accounts for 12-39% of total deposition in CMIP5 models.
246 The ratio of wet deposition to total deposition depends on several factors, for example,
247 dust size distribution, geographical locations of dust emission regions, and climate states



248 such as circulation and precipitation (e.g., Wu and Lin, 2013). The estimated global dust
249 burden ranges from 2.5 to 41.9 Tg, and from 8.1 to 36.1 Tg when MIROC4h and
250 HadGEM2-CC/ES are excluded. The lifetime of global dust particles ranges from 1.3 to
251 4.4 days. The dust burden (lifetime) in MERRA-2 is 20.3 Tg (4.1 days), which is larger
252 (longer) than most CMIP5 models. The fraction of wet deposition to total deposition in
253 MERRA-2 is 38.6%, which is in the upper end of CMIP5 results. There is a linear
254 relationship (with the correlation coefficient $R=0.67$, above the statistically significant
255 level of 0.01) between global dust burden and lifetime in CMIP5 models (excluding
256 HadGEM2-CC/ES; Figure 2a), indicating a longer lifetime of dust is generally associated
257 with a larger dust burden. Linear relationship ($R=0.46$, above the statistically significant
258 level of 0.05) is also found between lifetime and fraction of wet deposition (Figure 2b),
259 which indicates that a longer lifetime corresponds to a larger fraction of wet deposition in
260 the total deposition.

261

262 **4.2 Global dust emissions**

263 Dust emission is the first and the foremost process in the dust cycle and determines
264 the amount of dust entrained into the atmosphere. Figure 3 shows the spatial distribution
265 of dust emission fluxes from 15 CMIP5 models and MERRA-2 reanalysis. In general, all
266 the models can reproduce the main dust sources, known as the “dust belt” that extends
267 from North Africa, Middle East, Central Asia, South Asia, to East Asia and that can be
268 seen from satellite observations (Prospero et al., 2002; Ginoux et al., 2012). However, the
269 models differ significantly in the extent of this “dust belt”. Although a large group of
270 CMIP5 models (CSIRO-Mk3-6-0, GFDL-CM3, GISS-E2-H/S, MIROC5, MIROC-ESM,



271 MIROC-ESM-CHEM, MRI-CGCM3, and MRI-ESM1) simulate similarly the dust
272 emission regions mostly over deserts and adjacent arid/semi-arid regions, two of the
273 models (CESM1-CAM5 and MIROC4h) simulate much smaller areas of dust emission
274 and a few others (ACCESS1-0, CanESM2, HadGEM2-CC/ES) simulate more extended
275 dust emission regions. CESM1-CAM5 simulates isolated dust emission regions with “hot
276 spots” of dust emissions larger than $500 \text{ g m}^{-2} \text{ yr}^{-1}$, and dust emission in MIROC4h
277 concentrates only over the centers of deserts. In contrast, ACCESS1-0, CanESM2, and
278 HadGEM2-CC/ES not only simulate the dust emissions in deserts and adjacent regions,
279 but also produce a considerable amount of dust emissions over the Eastern Africa
280 (Somalia, Ethiopia, and Kenya), East India, and northern part of Indo China Peninsula,
281 which are rarely regarded as potential dust sources (Formenti et al., 2011; Shao, 2008).

282 Dust sources also exist in Australia, North America, South America, and South
283 Africa, as evident from surface observations (e.g., Shao, 2008) and satellite observations
284 (Prospero et al., 2002; Ginoux et al., 2012), although the emission fluxes are smaller than
285 those in the aforementioned “dust belt”. In these regions, most models produce a
286 considerable amount of dust emissions ($>5 \text{ g m}^{-2} \text{ yr}^{-1}$), while a small group of models
287 simulate much less or even negligible dust emissions. The models differ greatly in these
288 regions. For example, in Australia, two models (MIROC-ESM and MIROC-ESM-CHEM)
289 produces little dust emissions, while seven models (ACCESS1-0, CanESM2, CSIRO-
290 Mk3-6-0, GISS-E2-H/R, HadGEM2-CC/ES) produce much larger dust emissions with
291 emission fluxes higher than $10 \text{ g m}^{-2} \text{ yr}^{-1}$ in a large part of the region. In North America
292 which also has some dust sources (Wu et al., 2018a), five models (MIROC4h, MIROC-
293 ESM, MIROC-ESM-CHEM, MRI-CGCM3, MRI-ESM1) simulate little dust emissions,



294 while four models (ACCESS1-0, CanESM2, HadGEM2-CC/ES) simulate dust emission
295 fluxes exceeding $5 \text{ g m}^{-2} \text{ yr}^{-1}$ in a large part of the region. Note that ACCESS1-0 and
296 CanESMs also produce dust emissions in the high latitudes of Northern Hemisphere (>60
297 $^{\circ}\text{N}$) and eastern part of South America. The importance of high latitude dust is recognized
298 recently (Bullard et al., 2016), but the eastern part of South America has not been
299 regarded as a potential dust source (Formenti et al., 2011; Shao, 2008).

300 The contributions of dust emissions in nine different regions to global dust emission
301 is summarized in Table 4. The models consistently simulate the largest dust emission in
302 North Africa, which accounts for 36-79% of the global total dust emission. The models
303 also estimate large dust emissions in Middle East and East Asia, which account for 7-20%
304 and 4-19% of global dust emission, respectively. The contributions from Central Asia and
305 South Asia in CMIP5 models range from 1-14% and 0.9-10%, respectively. The
306 contributions from other sources (North America, South Africa, Australia, South America)
307 are much less consistent among the models, and the largest difference is in North
308 America (0.008-4.5%) and Australia (0.02-28%) by three orders of magnitude.

309 Particularly, HadGEM2-CC/ES simulate 25-28% of global dust emission from
310 Australia, which is comparable to that from sum of all Asian sources (Middle East,
311 Central Asia, South Asia, and East Asia). This estimate is unrealistically high, as will be
312 indicated by the comparison of surface dust concentrations in Section 4.4. The excessive
313 dust emission in Australia from HadGEM2-CC/ES may be related to the prognostic
314 vegetation used for dust emission, as the ACCESS1-0 model that uses the similar dust
315 emission parameterization but with the prescribed vegetation simulates a much lower
316 dust emission. The lowest dust emission in Australia is simulated by MIROC-ESM and



317 MIROC-ESM-CHEM, which contribute only 0.02-0.03% (1 Tg yr^{-1} or less) to the total
318 dust emission. This estimate is unrealistically low as Australia is an important dust
319 source (e.g., Shao et al., 2007) and is also much smaller than previous studies (e.g.,
320 Hunuees et al., 2011). The low dust emission in Australia from MIROC-ESM and
321 MIROC-ESM-CHEM may be related to the prognostic vegetation used for dust emission,
322 as the two other MIROC family models (MIROC4h and MIROC5) simulate significantly
323 higher dust emissions ($\sim 1\%$ of total dust emission).

324 The contributions from nine source regions in MERRA-2 to the total dust emission
325 are within the range of CMIP5 models. MERRA-2 estimates are obtained through the
326 assimilation of meteorology in model integrations and therefore uncertainties are reduced.

327 Since the amount of global dust emission differs substantially among different
328 models, the dust emission flux is further normalized by its global mean value in each
329 model for the comparison of dust emission area and intensity (Figure 4). Here the dust
330 emission area is defined as the region with normalized emission flux greater than 0.01.
331 Among the CMIP5 models, CESM-CAM5 and MIROC4h simulate the smallest dust
332 emission area, which are 2-3% of the global surface area, while CanESM2 simulates the
333 largest dust emission area (18% of the global surface area; Figure 4 and Table 3). The
334 maximum normalized dust emission flux is also the largest at 2682 and 3635 in CESM1-
335 CAM5 and MIROC4h, respectively, indicating the “hot spots” with extremely high dust
336 emission flux in the two models. The maximum normalized dust emission flux is
337 generally between 100 and 300 in other CMIP5 models and is approximately 200 in
338 MERRA-2 reanalysis.



339 The smallest dust emission area in CESM1-CAM5 is mainly because the model
340 adopts a geomorphic source erodibility with a threshold value of 0.1 for the dust emission
341 occurrences (Zender et al., 2003; Wu et al., 2016). Small dust emission area in MIROC4h
342 may be partly due to the higher horizontal resolution of the model (0.56°) than other
343 models (1°-3°) including MIROC5 (Table 1). The higher model resolution may change
344 the patterns of wind speeds and precipitation as well as the occurrence frequency of
345 strong winds and heavy precipitation and thus affect the dust emission regions. The
346 largest dust emission area in CanESM2 may be due to its prescribed land cover map,
347 and/or adoption of gustiness adjustment for wind friction velocity (von Salzen et al.,
348 2013). MERRA-2 gives a value of 7.4% for the dust emission area, which is in the
349 median of all the CMIP5 model results.

350 As normalized dust emission flux is comparable among the CMIP5 models, a global
351 map of multi-model mean and standard deviation of normalized dust emission flux are
352 thus constructed and shown in Figure 5. The multi-model mean represents the general
353 consensus among the CMIP5 models while the standard deviation indicates the
354 variability among models. The relative standard deviation is calculated by the ratio of
355 standard deviation to the mean, which is shown to illustrate the uncertainty among the
356 models. Mean normalized dust emission flux is large (>10) in the desert regions in North
357 Africa, Middle East, Central Asia, South Asia, East Asia, and Australia (Figure 5a). It
358 ranges from 1-10 in the desert adjacent regions and in small regions of South America,
359 North America, and South Africa (Figure 5a). The patterns of standard deviation of
360 multi-model results are generally similar to those of mean normalized dust emission flux
361 (Figure 5b). However, the relative standard deviation is quite different from the mean



362 normalized dust emission flux, and its pattern is nearly opposite (Figure 5c). The relative
363 standard deviation is mostly below 1 in the aforementioned desert regions with larger
364 mean normalized dust emission (>10) and increases to 1-4 in other regions with relative
365 smaller dust emission, indicating the large uncertainty of estimated dust emission flux in
366 the CMIP5 models.

367 Difference of dust emission uncertainty in different regions can be explained by two
368 reasons. First, in the deserts, soil is extremely dry (below the criteria for dust emission)
369 and surface is covered with little vegetation. In these regions, the models agree with each
370 other more easily in simulating the occurrence of dust emission. In the regions adjacent to
371 the deserts or with localized sandy lands, where soil is wetter and there is more
372 vegetation cover at the surface, the models differ significantly in the parameterizations of
373 dust emission, treatment of land cover, and simulated meteorology, and thus climate
374 models differ in their estimation of dust emission more strongly. Second, there are a
375 larger variety of complexities in the CMIP5 models compared to the models participating
376 in the AeroCom intercomparison (Section 2). Some models use the dynamic vegetation
377 for dust emission (e.g., HadGEM2-CC/ES, MIROC-ESM, MIROC-ESM-CHEM), and
378 deviate largely from other models over the regions with sparse vegetation cover such as
379 Australia. This further increases the differences in dust emission among the CMIP5
380 models.

381

382 **4.3 Dust deposition flux**

383 Dust deposition is a vital process in the dust cycle which removes dust particles
384 from the atmosphere and provides nutrients to the terrestrial and marine ecosystems.



385 Figure 6 shows the comparison of dust deposition flux at 84 selected stations between the
386 models and observations. Only seven CMIP5 models provide total dust deposition flux
387 (sum of dry and wet deposition), which are used here. The global dust emission in these
388 seven models ranges from 1600 to 3500 Tg yr⁻¹, which is at the medium level of all the
389 CMIP5 models. Observed annual mean dust deposition flux ranges from 10⁻⁴ to 10³ g m⁻²
390 yr⁻¹, indicating large spatial variabilities of dust deposition. In general, six of seven
391 CMIP5 models (excluding ACCESS1-0) reproduces the observed dust deposition flux
392 within a factor of 10 in most regions except over the Southern Ocean, Antarctica, and
393 Pacific. Over the Southern Ocean and in the Antarctica, all the models except CESM1-
394 CAM5 overestimate the dust deposition flux by more than a factor of 10 at two stations.
395 Over the Pacific Ocean, all the models except CanESM2 underestimate the dust
396 deposition flux by more than 10 times at several stations. In addition to the
397 overestimation over the Southern Ocean and Antarctica and the underestimation over the
398 Pacific Ocean, ACCESS1-0 mostly underestimate the dust deposition flux in other
399 regions with underestimation by more than a factor of 10 at several stations. Overall
400 ACCESS1-0 underestimates the dust deposition flux by approximately a factor of 2 on
401 average.

402 Similar to most of the CMIP5 models, MERRA-2 reproduces the observed dust
403 deposition flux within a factor of 10 at most stations except over the Southern Ocean and
404 Antarctica. Over the Southern Ocean and Antarctica, MERRA-2 tends to overestimate
405 the dust deposition flux by more than a factor of 10 at most stations. Compared to the
406 CMIP5 models, larger dust deposition over the Southern Ocean and Antarctica in
407 MERRA-2 may be related to the adoption of both meteorology and aerosol assimilation



408 in MERRA-2, which affects the dust transport and deposition. As mentioned in Section 2,
409 only AOD is taken into account in the aerosol assimilation for MERRA-2. Therefore the
410 large discrepancy of dust deposition at several stations in MERRA-2 may result from the
411 unrealistic representation of dust vertical profiles, size distribution, and deposition
412 process. Overall, the correlation coefficients between CMIP5 models and observations
413 (after taking the logarithms of both them; R_{\log}) range from 0.90 to 0.92 and are slightly
414 higher than that of MERRA-2 (0.87).

415 Dust deposition includes two mechanisms: dry and wet deposition. Figure 7 shows
416 the comparison of fraction of wet deposition in total deposition from models and
417 observations at 10 stations. These stations are located downwind of dust sources and can
418 be classified into two groups. One group are Bermuda (station #1) over the western
419 Atlantic Ocean, Amsterdam Island (station #2) over the southern Indian Ocean, Cape
420 Ferrat (station #3) in southern Europe, and New Zealand (station #6). For this group of
421 stations, fractions of wet deposition range from 17% to 70%. At these stations, all the
422 models simulate the fractions of wet deposition exceeding 75% and significantly
423 overestimate the fractions of wet deposition. MERRA-2 estimates smaller fractions of
424 wet deposition compared to the CMIP5 models but still significantly overestimates
425 fractions of wet deposition at these stations.

426 The other group includes Enewetak Atoll (station #4), Samoa (station #5) and
427 Fanning (station #8) over the tropical Pacific Ocean, Midway (station #7) over the
428 subtropical Pacific Ocean, Greenland (station #9) and Coastal Antarctica (station #10) in
429 the high latitudes. These stations are thousands of kilometers away from sources. At these
430 stations, observed fractions of wet deposition range from 65% to 90%, indicating the



431 dominance of wet deposition. Most of CMIP5 models except CanESM2 simulate the
432 fractions of wet deposition within 20% of observations. CanESM2 also simulates the
433 fraction of wet deposition comparable to observations except at Coastal Antarctica where
434 CanESM2 underestimates the fraction of wet deposition by up to 35%. MERRA-2
435 captures well the fraction of wet deposition over the tropical and subtropical Pacific
436 Ocean but significantly underestimate it by 40-45% in the high latitudes. The large
437 underestimation by CanESM2 and MERRA-2 may be related to the meteorology such as
438 precipitation and turbulent flux, or the parameterizations of dust deposition in the models,
439 which deserves future investigations.

440 Dust cycle can deliver nutrients from continents to oceans. Table 5 summarizes the
441 dust deposition and fraction of wet deposition onto the global surface, continents and
442 oceans, respectively in seven CMIP5 models and MERRA-2 reanalysis. Total deposition
443 in continents ranges from 1331 to 2850 Tg yr⁻¹ in seven CMIP5 models and accounts for
444 77-91 % of global total deposition. Total deposition in all the oceans ranges from 197 to
445 686 Tg yr⁻¹ and accounts for 9-23 % of global total deposition, indicating a considerable
446 uncertainty in dust deposition, which should be taken into account in modeling the
447 marine biogeochemistry with ESMs. MERRA-2 estimates 71% (29%) of dust deposited
448 in continents (oceans), and this estimation is smaller (larger) than all seven CMIP5
449 models, indicating MERRA-2 transport dust more efficiently to oceans. This is consistent
450 with the comparison of dust deposition flux shown in Figure 6 and may be related to the
451 assimilation of both meteorology and aerosols in MERRA-2. The fractions of wet
452 deposition (with respect to total deposition) in seven CMIP5 models are 8-33% and 49-71%
453 over continents and oceans, respectively. MERRA-2 estimates the fraction of wet



454 deposition (with respect to total deposition) 26% and 69% over the continents and oceans,
455 respectively, which lie within the range of CMIP5 models.

456

457 **4.4 Dust concentration**

458 Dust concentration is an important variable for its cycle. Figure 8 shows the
459 comparison of surface dust concentrations between models and observations at 22
460 selected stations. These stations are located in the downwind regions of dust sources, and
461 annual mean dust concentrations at these stations range from 10^{-1} to $10^2 \mu\text{g m}^{-3}$. In
462 general, the models reproduce observed surface dust concentrations within a factor of 10,
463 with the exceptions of HadGEM2-CC/ES and MIROC4h. Although HadGEM2-CC/ES
464 simulate well observed surface dust concentrations at the stations over the Atlantic Ocean
465 (stations #1-4) and slightly underestimate the observations in East Asia (stations #7-8),
466 the two models significantly overestimate surface dust concentrations at most of other
467 stations especially at the station located in Australia and downwind regions (stations
468 #15-21). This is consistent with their much higher dust emission in Australia compared to
469 other models (Table 3; Section 4.2). In contrast, MIROC4h largely underestimates
470 surface dust concentrations by 1-2 orders of magnitude at most stations. Although
471 compared to MIROC5, MIROC4h only simulates approximately 4 times lower global
472 dust emission, MIROC4h tends to concentrate all the dust emissions over smaller regions
473 of global surface (2.9% compared to 6.1%). Therefore, dust is less widely distributed in
474 the atmosphere and a smaller fraction of dust is transported to the downwind regions in
475 MIROC4h, as indicated by its almost 8 times smaller dust burden and only half the dust



476 lifetime compared to MIROC5. This difference can explain lower surface dust
477 concentrations in MIROC4h.

478 Although the CMIP5 models (excluding MIROC4h and HadGEM2-CC/ES) can
479 roughly reproduce the observed magnitudes of surface dust concentrations at most
480 stations, considerable discrepancy between models and observations can be found at
481 certain regions. Most models except CanESM2 significantly underestimate dust
482 concentrations at stations in Antarctica (stations #21 and #22), with the largest
483 underestimation by more than 2 orders of magnitude in MIROC-ESM/MIROC-ESM-
484 CHEM which also simulates much lower dust emissions in Australia, South Africa, and
485 southeastern South America. Eight models (ACCESS1-0, CESM-CAM5, CSIRO-Mk3-6-
486 0, GFDL-CM3, GISS-E2-H/R, MRI-CGCM3, MRI-ESM1) largely underestimate dust
487 concentrations by 1-2 orders of magnitude at station #6 in South Africa. Three MIROC
488 family models (MOROC5, MOROC-ESM, MIROC-ESM-CHEM) underestimate dust
489 concentrations by 1-2 orders of magnitude at several stations in the downwind regions of
490 Australia (stations #14, 15, and 17). Other noticeable discrepancies include
491 underestimations in East Asia by ACCESS1-0/MIROC5, underestimations over the
492 Tropical Pacific Ocean by CESM-CAM5/GISS-H2-H/GISS-E2-R, and overestimations
493 in Australia by CanESM2.

494 Overall the correlation coefficients and mean biases between CMIP5 models and
495 observations (after taking the logarithms of both of them; R_{\log} and MB_{\log}) ranges from
496 0.55 to 0.88 and from -5.59 to 1.52 for all CMIP5 models, respectively. If HadGEM2-
497 CC/ES and MIORC4h are excluded for the calculation, R_{\log} and MB_{\log} range from 0.60 to
498 0.88 and from -1.61 to 1.04, respectively. As a MB_{\log} of -0.7 (0.7) corresponds to a



499 general underestimation (overestimation) by a factor of 2, six models (CESM1-CAM5,
500 GISS-E2-H/R, MIROC5, MIROC-ESM, MIROC-ESM-CHEM) underestimate surface
501 dust concentrations by more than a factor of 2 on average, while CanESM2 overestimates
502 surface dust concentrations by the similar magnitude.

503 Compared to observations, MERRA-2 simulates well the dust concentrations at all
504 stations except station #6 in South Africa. This improvement by MERRA-2 compared to
505 the CMIP5 models may be due to the inclusion of both meteorology and aerosol
506 assimilation in MERRA-2. The correlation coefficients (R_{\log}) between MERRA-2 and
507 observations is 0.91, which is larger than all the CMIP5 models, and mean bias (MB_{\log}) is
508 close to zero (0.01).

509

510 **5. Discussion and Conclusions**

511 In this study we examine the present-day global dust cycle simulated by the 15
512 climate models participating in the CMIP5 project. The simulations are also compared
513 with a dataset MERRA-2 and observations of dust deposition and concentration. The
514 results show that the global dust emission in these models ranges from 735 to 8186 Tg yr⁻¹
515 and the global dust burden ranges from 2.5 to 41.9 Tg. The differences are larger than
516 those from models participating in the AeroCom project (Huneus et al., 2011), which is
517 a result of enhanced model complexities in modeling both climate and dust emission in
518 the CMIP5 models.

519 The simulated dust emission regions also differ greatly accounting for a global
520 surface area of 2.9%-18%. The models agree most with each other in reproducing the
521 “dust belt” that extends from North Africa, Middle East, Central Asia, South Asia, to East



522 Asia, but there are large uncertainties in the extent of this “dust belt” and other source
523 regions including Australia, North America, South America, and South Africa.
524 Particularly, some models simulate little dust emissions (<0.1% of global dust emission)
525 in Australia and North America, while some other models simulate larger dust emissions
526 there which account for 10-30% and 3-4% of global dust emission in Australia and North
527 America, respectively. It is also revealed that the increasing complexity of ESMs
528 (HadGEM2-CC/ES, MIROC-ESM, and MIROC-ESM-CHEM) by coupling dust
529 emission with dynamic vegetation can amplify the uncertainty associated with dust
530 emissions.

531 Removal of dust particles in the CMIP5 models is mainly through dry deposition,
532 and wet deposition only accounts for 12-39% of total deposition. The associated dust life
533 time is about 1.3-4.4 days. A clear linear relationship between dust burden, dust lifetime,
534 and fraction of wet deposition to total deposition is present in the CMIP5 models,
535 suggesting a general consistency among these models. The models also estimate that 77-
536 91% of emitted dust are deposited back to continents and 9-23% of them are deposited to
537 the oceans. The fraction of wet deposition is smaller in most CMIP5 models and dust
538 lifetime is shorter compared to MERRA-2 reanalysis, indicating a shorter distance for
539 dust transport from its sources in most CMIP5 models. Compared to the observations, the
540 CMIP5 models (except MIRCO4h) reproduce dust deposition flux and surface dust
541 concentration by a factor of 10 at most stations. Larger discrepancies are found in the
542 remote regions such as Antarctica and Tropical Pacific Ocean. In Australia and
543 downwind regions, four MIROC family models (MIROC4h, MIROC5, MIROC-ESM,
544 MIROC-ESM-CHEM) which simulate little dust emission in Australia largely



545 underestimate the dust concentrations at stations in the remote regions. Contrarily
546 HadGEM2-CC/ES overestimate dust concentrations. MIROC4h shows the largest
547 discrepancy by underestimating the surface dust concentrations by more than a factor of
548 100 in Australia and downwind regions. Overall, although MIROC4h simulates 4-5 times
549 lower global dust emission than other three MIROC family models, MIROC4h simulates
550 on average more than 50 times smaller surface dust concentrations at 22 stations. This
551 can be ascribed to the fact that most dust emissions in MIROC4h are concentrated over
552 the desert centers, which limits the long-range transport of dust particles to the remote
553 regions.

554 These results show large uncertainties of global dust cycle in ESMs. In fact, these
555 models are fully-coupled atmosphere-land-ocean models and some of them also include
556 the dynamic vegetation. As a result, uncertainties are larger compared to those in
557 previous models participating in the AeroCom intercomparison project where sea surface
558 temperature is prescribed, and more strictly, in some models, meteorological fields are
559 prescribed from reanalysis (Huneus et al., 2011). Larger uncertainties in the CMIP5
560 models with dynamic vegetation is expected, as a prognostic vegetation would depart
561 from the observed or constructed vegetation and may also lead to a large bias in soil
562 moisture, which may thus lead to an additional bias in dust emissions in these models.
563 Uncertainties of dust simulations also vary with regions, and a smaller uncertainty is
564 found in the deserts over the “dust belt” in the North Hemisphere, but a larger uncertainty
565 exists in other regions including Australia and North America. The large uncertainties of
566 global dust cycle in the CMIP5 models would cast a doubt on the reliability of dust
567 radiative forcing estimated in these models.



568 Because the dust lifecycle involves various processes with the scales from
569 micrometers to tens of thousands of kilometers and consists of lots of parameters, the
570 representation of dust cycle in climate models is a big challenge for the model
571 community. Dust emission is the first and foremost process for model improvements of
572 dust cycle (Shao, 2008; Shao et al., 2011). Improving dust emission not only lies in the
573 development of dust emission scheme but also in its implementation into climate models
574 (e.g., Shao, 2008; Wu et al., 2016; Wu et al., 2019). For example, different dust emission
575 schemes with specific land cover datasets and criteria for the occurrence of dust emission
576 are adopted in the models (Table 1 and references therein). Therefore, different results of
577 dust emission among the CMIP5 models reflect in many aspects the differences in
578 meteorology, land cover data, and dust emission parameterizations. A close look at these
579 factors in each model will help to unravel reasons behind the biases in these models. In
580 addition, the models are only evaluated with observed dust deposition and surface
581 concentrations. Although it is roughly acceptable, it is also desirable to collect the
582 observations of dust emission flux and use them for model evaluation. Particularly, for
583 dust deposition and dust concentration, some biases come from dust emission and others
584 from circulation and deposition parameterizations. It is only possible to separate the
585 contributions of different processes to the biases in dust deposition and concentration, if
586 observations of dust emission are also included in model comparison.

587 It should be mentioned that dust size distribution is an important parameter for dust
588 cycle (e.g., Shao, 2008; Mahowald et al., 2014), and it is not included in this study as the
589 model data are not available. Evolution of dust size distribution during dust transport and
590 deposition is critical to our understanding of the model bias in dust cycle. We suggest that



591 the size-resolved dust emission, concentration, and deposition should be outputted and
592 provided in the latest CMIP6 project (Eyring et al., 2016). Moreover, observations of
593 size-resolved dust concentration and deposition is urgently needed. A compile of
594 available observations of dust size distribution (e.g., Mahowald et al., 2014; Ryder et al.,
595 2018) are also required for model evaluation.

596

597 **Data availability**

598 CMIP5 results are available in <https://esgf-node.llnl.gov/search/cmip5/>. MERRA-2
599 is available in <https://disc.gsfc.nasa.gov/datasets?project=MERRA-2>. Observations of
600 dust deposition and fraction of wet deposition is provided in the literature led by N.
601 Huneus (<https://www.atmos-chem-phys.net/11/7781/2011/>). Observations of surface
602 dust concentrations are provided by Joseph M. Prospero from the Rosenstiel School of
603 Marine and Atmospheric Science at the University of Miami.

604

605 **Author contributions**

606 CW and ZL designed the study. CW did the data analyses with advices from ZL and
607 XL. CW wrote the manuscript with contributions from ZL and XL.

608

609 **Competing interests**

610 The authors declare that they have no conflict of interest.

611

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623

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945

946 **Table 1.** CMIP5 model used in this study. For comparison with CMIP5 models, MERRA-2 reanalysis is also included.

No.	Models ^a	Resolution	Ensemble number	Dust size (in diameter)	Vegetation cover	Dust emission scheme	Model reference
1	ACCESS1-0	1.3° × 1.9°	3	6 bins: 0.0632-0.2-0.632-2-6.32-20-63.2 μm	Prescribed	Woodward (2001, 2011)	Bi et al. (2013) Dix et al. (2013)
2	CanESM2	2.8° × 2.8°	5	2 modes: MMD= 0.78 μm (σ=2) and 3.8 μm (σ=2.15) ^b	Prescribed	Martcorena and Bergametti (1995)	Arora et al. (2011) von Salzen et al. (2013)
3	CESM1-CAM5	0.9° × 1.25°	2	2 modes: 0.1-1-10 μm ^c	Prescribed	Zender et al. (2003)	Hurrell et al. (2013)
4	CSIRO-Mk3-6-0	1.9° × 1.9°	10	4 bins: 0.2-2-4-6-12 μm	Prescribed	Ginoux et al. (2001, 2004)	Rotstayn et al. (2012)
5	GFDL-CM3	2° × 2.5°	5	5 bins: 0.2-2-3.6-6-12-20 μm	Prognostic	Ginoux et al. (2001)	Delworth et al. (2006) Donner et al. (2011)
6	GISS-E2-H	2° × 2.5°	12	4 bins: <2, 2-4-8-16 μm	Prescribed	Cakmur et al. (2006) Miller et al. (2006)	Schmidt et al. (2014)
7	GISS-E2-R	2° × 2.5°	12	4 bins: <2, 2-4-8-16 μm	Prescribed	Cakmur et al. (2006) Miller et al. (2006)	Schmidt et al. (2014)
8	HadGEM2-CC	1.3° × 1.9°	3	6 bins: 0.0632-0.2-0.632-2-6.32-20-63.2 μm	Prognostic	Woodward (2001, 2011)	Collins et al. (2011) Martin et al. (2011)
9	HadGEM2-ES	1.3° × 1.9°	4	As HadGEM2-CC	Prognostic	Woodward (2001, 2011)	Collins et al. (2011) Martin et al. (2011)
10	MIROC4h	0.56° × 0.56°	1	10 bins: 0.2-0.32-0.5-0.8-1.26-2-3.16-5.02-7.96-12.62-20 μm	Prescribed	Takemura et al. (2000)	Sakamoto et al. (2012)
11	MIROC5	1.4° × 1.4°	5	6 bins: 0.2-0.43-0.93-2-4.3-9.3-20 μm	Prescribed	Takemura et al. (2000, 2009)	Watanabe et al. (2010)
12	MIROC-ESM	2.8° × 2.8°	1	As MIROC4h	Prognostic	Takemura et al. (2000, 2009)	Watanabe et al. (2011)
13	MIROC-ESM-CHEM	2.8° × 2.8°	3	As MIROC4h	Prognostic	Takemura et al. (2000, 2009)	Watanabe et al. (2011)
14	MRI-CGCM3	1.1° × 1.1°	5	6 bins: 0.2-0.43-0.93-2-4.3-9.3-20 μm	Prescribed	Shao et al. (1996) Tanaka and Chiba (2005, 2006)	Yukimoto et al. (2011, 2012)



15	MRI-ESM1	1.1° × 1.1°	1	6 bins: 0.2-0.43-0.93-2-4.3-9.3-20 μm	Prescribed	Shao et al. (1996) Tanaka and Chiba (2005, 2006)	Yukimoto et al. (2011, 2012) Adachi et al. (2013)
16	MERRA-2	0.5° × 0.625°	1	5 bins: 0.2-2-3.6-6-12-20 μm	Prescribed	Ginoux et al. (2001)	Randles et al. (2017) Buchard et al. (2017)

947 ^a: Expansions of acronyms: ACCESS1-0, Australian Community Climate and Earth-System Simulator version 1.0; CanESM2, Second Generation Canadian Earth
 948 System Model; CESM1-CAM5, Community Earth System Model version 1-Community Atmosphere Model version 5; CSIRO-Mk3-6-0, Commonwealth Scientific
 949 and Industrial Research Organization Mark 3.6.0; GFDL-CM3, Geophysical Fluid Dynamics Laboratory Climate Model version 3; GISS-E2-H, Goddard Institute for
 950 Space Studies Model E2 coupled with HYCOM (Hybrid Coordinate Ocean Model); GISS-E2-R, Goddard Institute for Space Studies Model E2 coupled with the
 951 Russell ocean model; HadGEM2-CC, Hadley Centre Global Environment Model version 2 with Carbon Cycle configuration; HadGEM2-ES, Hadley Centre Global
 952 Environment Model version 2 with Earth System configuration; MIROC4h, Model for Interdisciplinary Research on Climate version 4 (high resolution); MIROC5,
 953 Model for Interdisciplinary Research on Climate version 5; MIROC-ESM, Model for Interdisciplinary Research on Climate-Earth System Model; MIROC-ESM-
 954 CHEM, Model for Interdisciplinary Research on Climate-Earth System Model with Chemistry Coupled; MRI-CGCM3, Meteorological Research Institute Coupled
 955 Atmosphere–Ocean General Circulation Model version 3; MRI-ESM1, Meteorological Research Institute Earth System Model version 1.
 956 ^b: MMD is the abbreviation of mass median diameter and σ is geometric standard deviation.
 957 ^c: Dust emission is calculated in the size range of 0.1-1 and 1-10 μm for accumulation and coarse modes, respectively.



958 **Table 2.** The location of observational stations for (a) surface dust concentration and
959 (b) fraction of wet deposition used in this study.

960 (a)

No.	Name	Latitude	Longitude	No.	Name	Latitude	Longitude
1	Miami	25.75°N	80.25°W	12	Fanning Island	3.92°N	159.33°W
2	Bermuda	32.27°N	64.87°W	13	Hawaii	21.33°N	157.7°W
3	Barbados	13.17°N	59.43°W	14	Jabirun	12.7°S	132.9°E
4	Izana Tenerife	28.3°N	16.5°W	15	Cape Grim	40.68°S	144.68°E
5	Mace Head	53.32°N	9.85°W	16	New Caledonia	22.15°S	167°E
6	Rukomechi	16°S	29.5°E	17	Norfolk Island	29.08°S	167.98°E
7	Cheju	33.52°N	126.48°E	18	Funafuti	8.5°S	179.2°W
8	Hedo	26.92°N	128.25°E	19	American Samoa	14.25°S	170.58°W
9	Enewetak Atoll	11.33°N	162.33°E	20	Cook Islands	21.25°S	159.75°W
10	Nauru	0.53°N	166.95°E	21	Palmer	64.77°S	64.05°W
11	Midway Island	28.22°N	177.35°W	22	Mawson	67.6°S	62.5°E

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962 (b)

No.	Name	Latitude	Longitude	No.	Name	Latitude	Longitude
1	Bermuda	32.27°N	64.87°W	6	New Zealand	34.55°S	172.75°E
2	Amsterdam Island	37.83°S	77.5°E	7	Midway	28.22°N	177.35°W
3	Cape Ferrat	43.68°N	7.33°E	8	Fanning	3.92°N	159.33°W
4	Enewetak Atoll	11.33°N	162.33°E	9	Greenland	65°N	44°W
5	Samoa	14.25°S	170.57°W	10	Coastal Antarctica	75.6°S	26.8°W

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967 **Table 3.** Global dust budgets in CMIP5 models.

Model	Emission ^a (Tg/yr)	Wet deposition ^b (Tg/yr)	Burden (Tg)	Life time (day)	Diameter (µm)
ACCESS1-0	2218 (13%)	261 (12%)	8.1	1.3	0.06 - 73
CanESM2	2964 (18%)	882 (30%)	35.8	4.4	Median (0.39, 2)
CESM1-CAM5	3454 (2.0%)	1243 (36%)	24.9	2.6	0.1 - 10
CSIRO-Mk3-6-0	3698 (8.9%)	1024 (28%)	36.1	3.6	0.2 - 12
GFDL-CM3	1246 (10%)	210 (17%)	13.5	4.0	0.1 - 10
GISS-E2-H	1699 (8.2%)	641 (38%)	17.5	3.8	<2 to 16
GISS-E2-R	1677 (8.2%)	625 (37%)	16.9	3.7	<2 to 16
HadGEM2-CC	8186 (11%)	1521 (19%)	41.9	1.9	0.06 - 63
HadGEM2-ES	7972 (10%)	1429 (18%)	41.4	1.9	0.06 - 63
MIROC4h	735 (2.9%)	179 (24%)	2.5	1.4	0.2 – 20
MIROC5	2716 (6.1%)	668 (25%)	19.0	3.0	0.2 – 20
MIROC-ESM	3339 (5.2%)	540 (16%)	15.5	2.0	0.2 – 20
MIROC-ESM- CHEM	3598 (5.2%)	591 (16%)	16.7	2.0	0.2 – 20
MRI-CGCM3	2107 (5.9%)	819 (39%)	14.3	2.5	0.2 – 20
MRI-ESM1	2052 (6.1%)	801 (39%)	13.9	2.5	0.2 – 20
MERRA-2 ^c	1620 (7.4%)	692 (38.6%)	20.3	4.1	0.2 – 20

968 ^a: The global dust emission area fraction is given in parenthesis next to the global dust
 969 emission. The dust emission area is defined as the region with the annual mean dust
 970 emission flux larger than 1% of global mean annual dust emission flux.

971 ^b: The ratio of wet deposition to total deposition is given in parenthesis next to wet
 972 deposition.

973 ^b: The global dust deposition is 1692 Tg, which is larger than dust emission because
 974 of no adjustment done with dust emission after aerosol assimilation (Section 2).



975 **Table 4.** Dust emission amount (Tg) in nine dust source regions. The contribution of each source region to global total dust emission is given in
 976 the parenthesis next to dust emission amount.

No.	Models	Global	North Africa	Middle East	Central Asia	South Asia	East Asia	Australia	North America	South America	South Africa
1	ACCESS1-0	2218	1097 (49.5%)	356 (16.1%)	95 (4.3%)	159 (7.2%)	132 (6.0%)	254 (11.4%)	49 (2.2%)	46 (2.1%)	21 (1.0%)
2	CanESM2	2964	1053 (35.5%)	415 (14.0%)	323 (10.9%)	99 (3.3%)	151 (5.1%)	218 (7.3%)	133 (4.5%)	365 (12.3%)	96 (3.2%)
3	CESM1-CAM5	3454	1609 (46.6%)	698 (20.2%)	495 (14.3%)	122 (3.5%)	329 (9.5%)	38 (1.1%)	35 (1.0%)	26 (0.7%)	101 (2.9%)
4	CSIRO-Mk3-6-0	3698	1863 (50.4%)	555 (15.0%)	122 (3.3%)	160 (4.3%)	589 (15.9%)	143 (3.9%)	23 (0.6%)	138 (3.7%)	106 (2.9%)
5	GFDL-CM3	1246	749 (60.1%)	150 (12.1%)	68 (5.4%)	41 (3.3%)	113 (9.1%)	52 (4.2%)	5 (0.4%)	44 (3.6%)	19 (1.5%)
6	GISS-E2-H	1699	1045 (61.5%)	252 (14.8%)	109 (6.4%)	96 (5.7%)	94 (5.5%)	71 (4.2%)	4 (0.3%)	22 (1.3%)	5 (0.3%)
7	GISS-E2-R	1678	1035 (61.7%)	238 (14.2%)	92 (5.5%)	90 (5.4%)	103 (6.1%)	86 (5.1%)	4 (0.2%)	23 (1.4%)	5 (0.3%)
8	HadGEM2-CC	8186	3124 (38.2%)	593 (7.2%)	403 (4.9%)	826 (10.1%)	359 (4.4%)	2278 (27.8%)	264 (3.2%)	196 (2.4%)	142 (1.7%)
9	HadGEM2-ES	7973	3221 (40.4%)	579 (7.3%)	418 (5.2%)	820 (10.3%)	321 (4.0%)	1988 (24.9%)	340 (4.3%)	144 (1.8%)	139 (1.7%)
10	MIROC4h	735	437 (59.4%)	71 (9.7%)	81 (11.1%)	45 (6.1%)	64 (8.8%)	9 (1.2%)	0.1 (0.02%)	3 (0.5%)	24 (3.2%)
11	MIROC5	2716	1762 (64.9%)	269 (9.9%)	175 (6.5%)	96 (3.5%)	243 (8.9%)	26 (1.0%)	4 (0.2%)	79 (2.9%)	61 (2.2%)
12	MIROC-ESM	3339	2627 (78.7%)	244 (7.3%)	72 (2.2%)	30 (0.9%)	273 (8.2%)	0.6 (0.02%)	0.3 (0.008%)	89 (2.6%)	6 (0.2%)
13	MIROC-ESM-CHEM	3598	2719 (75.6%)	274 (7.6%)	84 (2.3%)	44 (1.2%)	362 (10.1%)	1 (0.03%)	0.4 (0.01%)	100 (2.8%)	13 (0.4%)



14	MRI-CGCM3	2107	1146 (54.4%)	258 (12.2%)	22 (1.1%)	174 (8.3%)	390 (18.5%)	55 (2.6%)	2 (0.09%)	49 (2.3%)	11 (0.5%)
15	MRI-ESM1	2052	1108 (54.0%)	246 (12.0%)	21 (1.0%)	167 (8.1%)	392 (19.1%)	57 (2.8%)	2 (0.09%)	48 (2.3%)	10 (0.5%)
16	MERRA-2	1670	1104 (61.1%)	182 (16.2%)	56 (7.7%)	55 (3.1%)	162 (6.3%)	59 (2.6%)	8 (0.5%)	30 (1.7%)	15 (0.7%)



977 **Table 5.** Total dust deposition and wet deposition in the global surface, continents,
978 and oceans, respectively from CMIP5 models and MERRA-2 reanalysis. Only the
979 seven CMIP5 models with both dry and wet depositions provided are used here.

Model	Global		Continent		Ocean	
	Total	Wet ^a	Total ^b	Wet ^a	Total ^b	Wet ^a
ACCESS1-0	2216	261 (12%)	2019 (91%)	159 (8%)	197 (9%)	102 (52%)
CanESM2	2965	882 (30%)	2279 (77%)	513 (22%)	686 (23%)	369 (54%)
CESM1-CAM5	3454	1243 (36%)	2850 (83%)	945 (33%)	604 (17%)	298 (49%)
GISS-E2-H	1684	641 (38%)	1359 (81%)	410 (30%)	324 (19%)	231 (71%)
GISS-E2-R	1665	625 (37%)	1331 (80%)	392 (29%)	334 (20%)	232 (70%)
MRI-CGCM3	2109	819 (39%)	1649 (78%)	499 (30%)	460 (22%)	319 (69%)
MRI-ESM1	2054	801 (39%)	1609 (78%)	492 (30%)	445 (22%)	309 (69%)
MERRA-2	1792	692 (38.6%)	1272 (71%)	335 (26%)	520 (29%)	356 (69%)

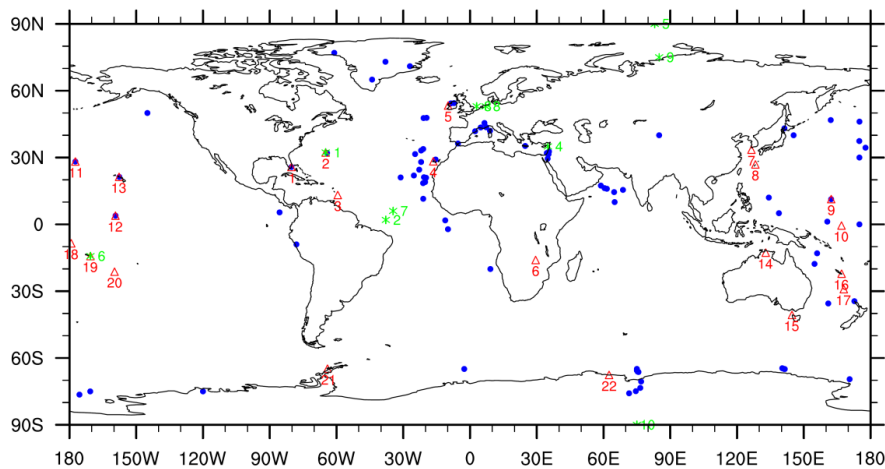
980 ^a: The ratio of wet deposition to total deposition is given in parenthesis next to wet
981 deposition.

982 ^b: The fraction of continental (or oceanic) deposition to global deposition is given in
983 next to continental (or oceanic) deposition.

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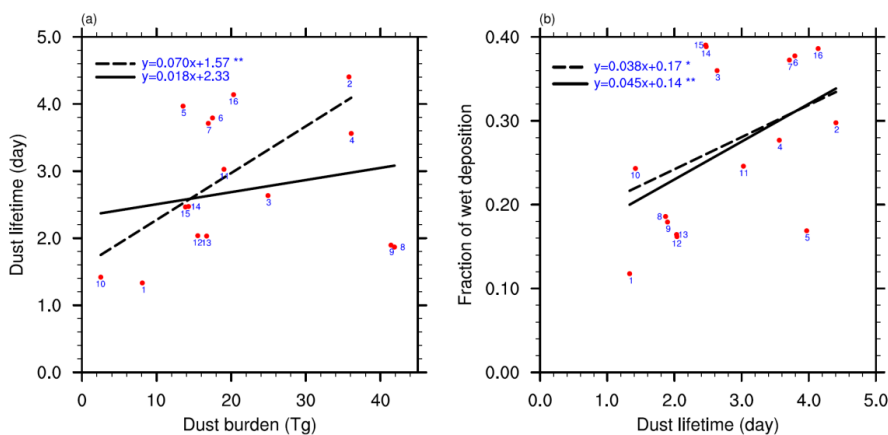
986

987 **Figure 1.** The distribution of observational stations used in this study: blue circles for
988 dust deposition, red triangles for surface dust concentrations, and green asterisks for
989 fraction of wet deposition. The descriptions of all these stations can be found in
990 Section 3.

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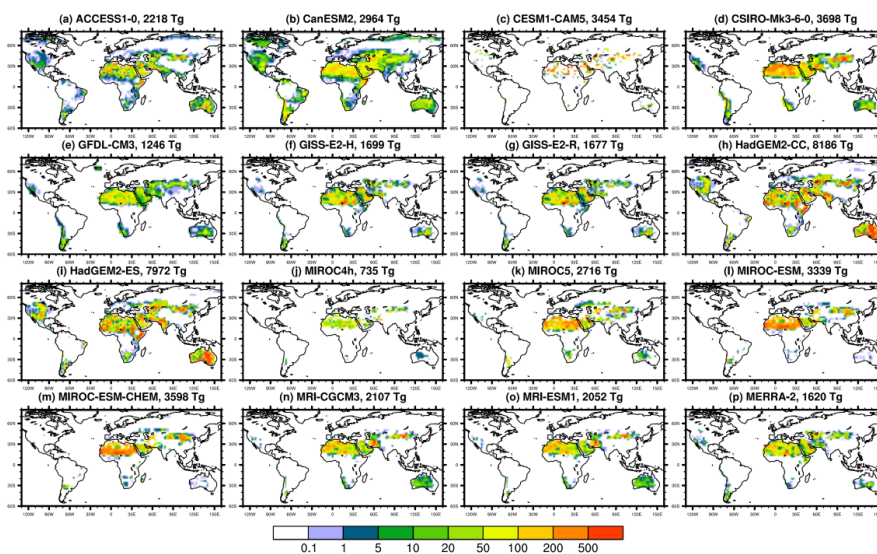
993

994 **Figure 2.** Scatter plot of (a) dust burden versus dust life time
995 versus fraction of wet deposition to total deposition in 15 CMIP5 models and in
996 MERRA-2 reanalysis. The models are indexed as Table 1. The regression lines from
997 all the CMIP5 models (solid) and the CMIP5 models excluding HadGEM2-CC/ES
998 models (dash) are also shown with the slopes and intercepts for the regression
999 equation. Significant test for each regression is denoted by one asterisk (*; above
1000 significant level of 0.1) and two asterisks (**; above significant level of 0.05) after
1001 each regression equation.

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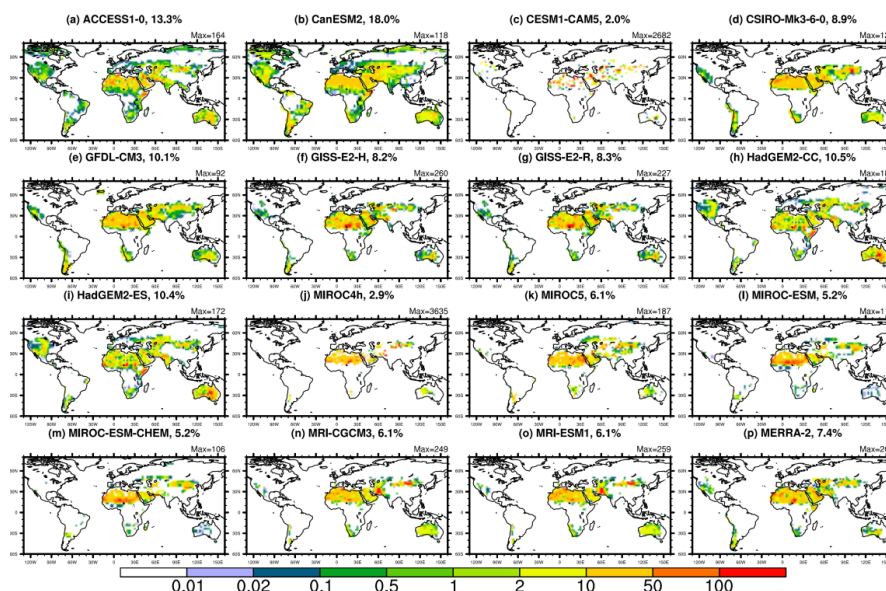
1004

1005 **Figure 3.** (a-o) Annual mean dust emission flux ($\text{g m}^{-2} \text{yr}^{-1}$) during 1960-2005 from
1006 15 CMIP5 models, and (p) annual mean dust emission ($\text{g m}^{-2} \text{yr}^{-1}$) during 1980-2018
1007 from MERRA-2 reanalysis. The total annual global dust emission is included in the
1008 title of each panel.

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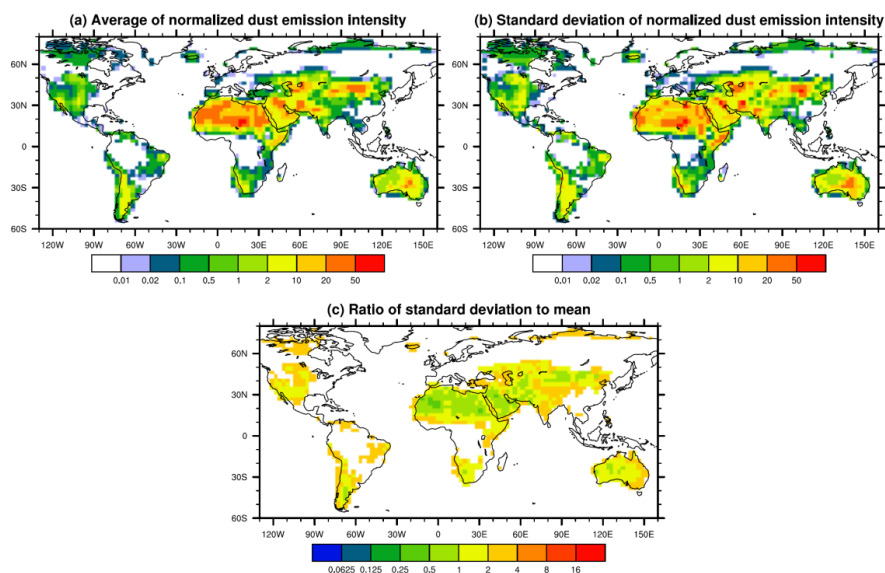
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1012 **Figure 4.** Normalized dust emission flux in 15 CMIP5 models and MERRA-2
1013 reanalysis. Normalized dust emission flux is calculated from dust emission flux
1014 divided by global mean for each model. The percentage of dust source area relative to
1015 global total surface area is given in the title of each panel. Dust source area is defined
1016 as the normalized dust emission flux greater than 0.01. The maximum normalized
1017 dust emission flux is also given in the top right corner of each panel.

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1020 **Figure 5.** Mean, standard deviation, and relative standard deviation (also known as
1021 coefficient of variation) of normalized dust emission flux from 15 CMIP5 models.

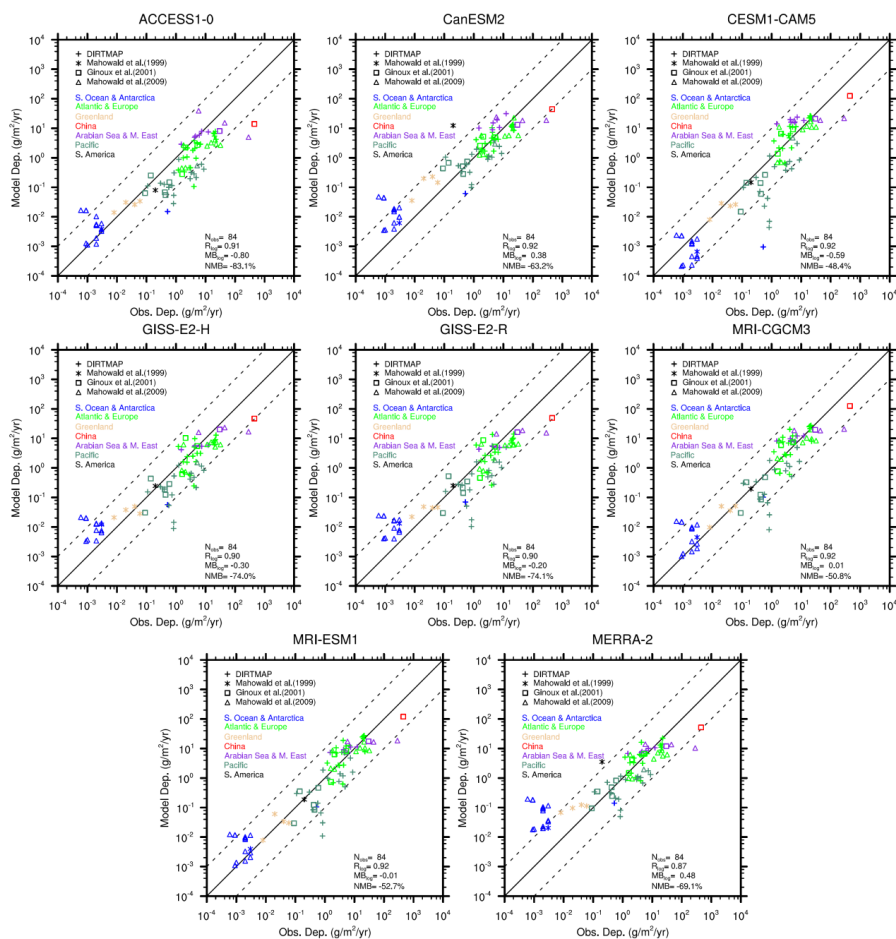
1022 Relative standard deviation is derived by calculating the ratio of standard deviation to

1023 mean.

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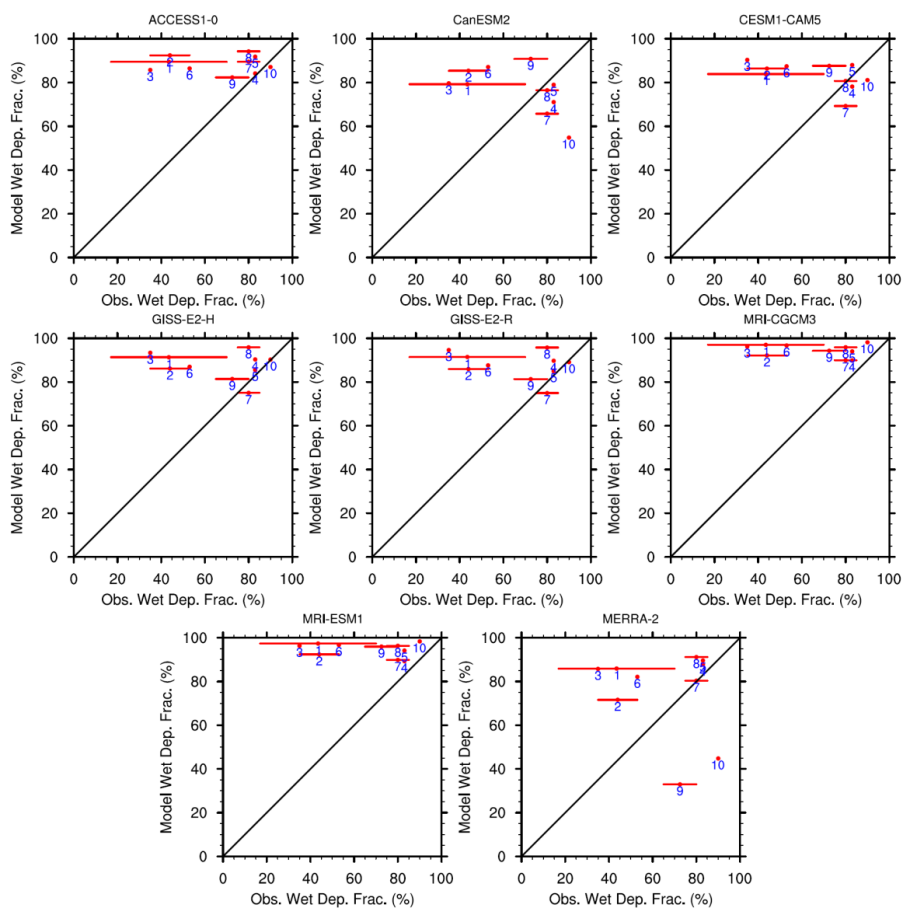
1026

1027 **Figure 6.** Scatterplot of dust deposition flux at 84 selected stations between models
 1028 and observations. The stations are marked with different styles according to the
 1029 sources of data and with different colors for different locations (Section 3). Also given
 1030 are the correlation coefficients and mean bias between models and observations (after
 1031 taking the logarithms; R_{\log} and MB_{\log} , respectively). The normalized mean bias (NMB)
 1032 that is calculated from the mean bias divided by mean observations is given as well.
 1033 The 1:1 (solid) and 1:10/10:1 (dash) lines are plotted for reference.

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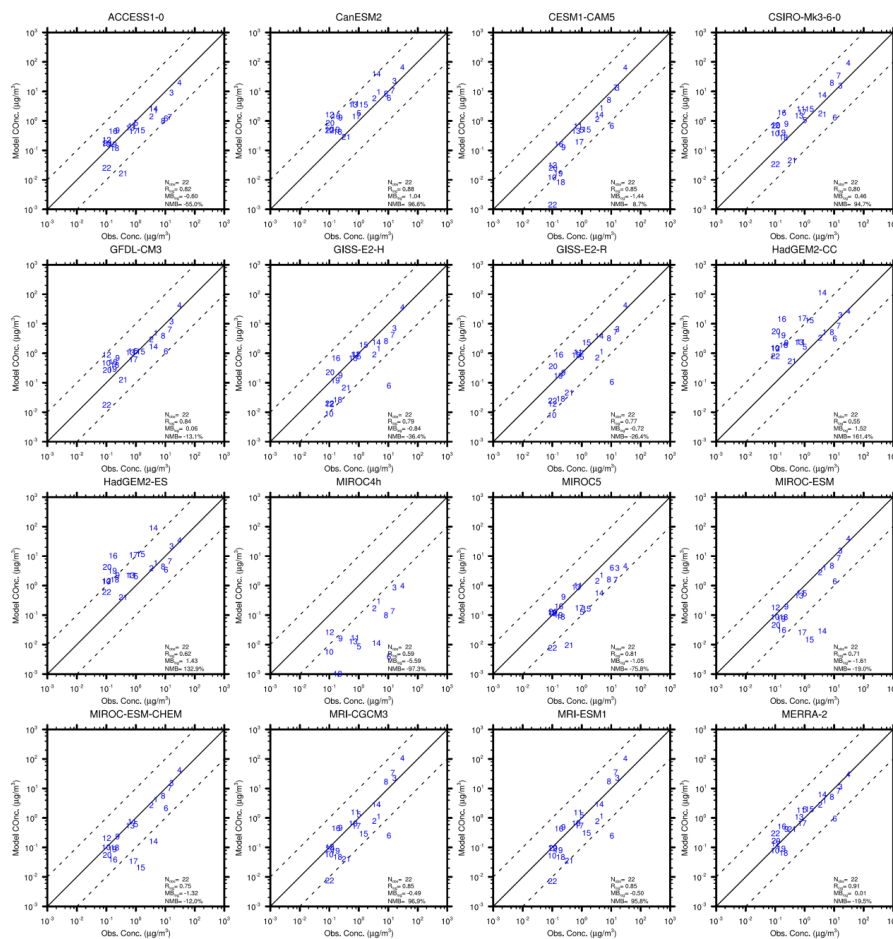
1036

1037 **Figure 7.** Scatterplot of fraction of wet deposition in total deposition between models
1038 and observations. For the observations that provide the minimum and maximum
1039 values, the mean of minimum and maximum values is used with the ranges indicated
1040 by a horizontal line. Station numbers are indexed following Table 2.

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1044 **Figure 8.** Scatterplot of surface dust concentration at 22 selected stations between
 1045 models and observations. The stations are indexed as Table 2 and their locations are
 1046 shown in Figure 1. Also given are the correlation coefficients and mean bias between
 1047 models and observations (after taking the logarithms; R_{\log} and MB_{\log} , respectively).
 1048 The normalized mean bias (NMB) that is calculated from the mean bias divided by
 1049 mean observations is given as well. The 1:1 (solid) and 1:10/10:1 (dash) lines are
 1050 plotted for reference. The comparison results for some stations (#15-17 and #19-22
 1051 for MIROC4h; #21 and #22 for MIROC-ESM and MIROC-ESM-CHEM) are not
 1052 shown as they are located too low and outside the frame.