

Supplemental Material

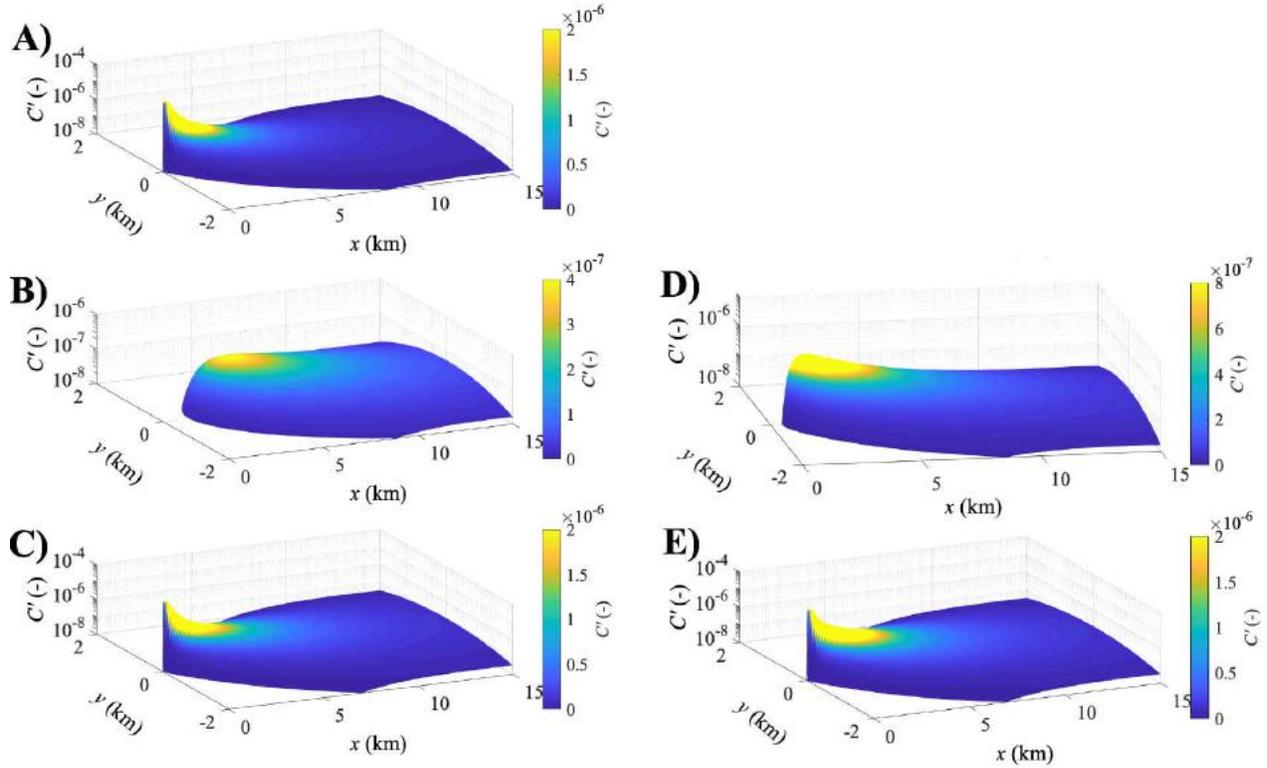
9 S1 The Gaussian Plume model theoretical basis

10 Sea surface emissions for focused bubble plumes can be described as a point source whose
 11 atmospheric plume can be modeled as a Gaussian plume (Hanna et al., 1982). The Gaussian plume
 12 model relates atmospheric emissions, E_A , to the concentration anomaly, C' , relative to ambient, C ,
 13 wind speed, u , and the atmospheric turbulence parameters, σ_y and σ_z , defined in a cartesian
 14 coordinate system where x is the downwind, i.e., wind direction, θ , y is the transverse direction, and
 15 z is the vertical coordinate,

$$\begin{aligned}
 16 \quad C'(x, y, z) = E_A \frac{\exp\left(\frac{-y^2}{2\sigma_y^2}\right)}{(2\pi u \sigma_z \sigma_y)} & \left(\exp\left(\frac{-(z-h)^2}{2\sigma_z^2}\right) + \exp\left(\frac{-(z+h)^2}{2\sigma_z^2}\right) + \exp\left(\frac{-(z-2(BL-h))^2}{2\sigma_z^2}\right) + \right. \\
 17 \quad \left. \exp\left(\frac{-(z+2(BL-h))^2}{2\sigma_z^2}\right) \right), & \tag{S1}
 \end{aligned}$$

18 where h is the measurement height. The second exponential term represents reflection off the sea
 19 surface and assumes a non-sticky molecule, the third term represents reflection off the marine
 20 boundary layer, at height BL , and the fourth term represents re-reflection off the surface (**Fig. S1**). A
 21 shallower boundary layer height, BL , contributes more significantly to this reflection, for example, a
 22 250-m BL and $u = 2.1 \text{ m s}^{-1}$ contributes 10-20% of C' from 2-km downwind (**Fig. S1**). C' is relative
 23 to the background concentration, C , which is based on concentrations outside the plume's edges and
 24 can include environmental gradients from large-scale transport processes.

25



26
 27 **Figure S1:** Surface concentration, C' , for Gaussian plumes for $u=2.1 \text{ m s}^{-1}$ **A)** with no planetary boundary layer, **BL**, **B)**
 28 reflection from a 250-m **BL** (Eqn. 1, terms 3+4) and **C)** Gaussian plume with 250-m **BL**. **D)** Reflection plume from a 100-
 29 m **BL** and **E)** Gaussian plume with 100-m **BL**. Arbitrary C' units. Data key on panel.

30
 31 Parameterization of σ_y and σ_z are:

$$32 \quad \sigma_y = a(1 + 10^{-4}x)^{-1/2}; \sigma_z = b(1 + cx)^n, \quad (\text{S2})$$

33 with a , b , and c depending on stability class (Briggs, 1973). Stability class (**Table S2**) is based on
 34 solar insolation, I , and u (Hanna et al., 1982). These turbulence parameterizations (**Eqn. S2**) are
 35 discrete (**Table S1**), which introduces uncertainty in σ_x and σ_y , particularly at stability class
 36 transitions. To remove discretization distortions, a 2nd-order polynomial was fit to the turbulence
 37 parameters a , b , and c , for each insolation class with respect to u (**Fig. S2**). The polynomial fit then
 38 was evaluated for u to determine a , b , and c for all three insolation classes. Strong, moderate, and
 39 weak insolation were fit for class values of $I = 800, 500, \text{ and } 175 \text{ W m}^{-2}$, respectively. The
 40 parameterizations are shown in **Fig. S2** and are for $1 < u < 6 \text{ m s}^{-1}$. Given the coarse nature of the
 41 Pasquill classes, no effort was made to optimize; simply, the center u for each class was used in the
 42 parameterization.

44 **Table S1.** Parameters for atmospheric turbulence parameterizations for different stability classes*

45 Class	<i>a</i>	<i>b</i>	<i>c</i>	<i>n</i>	Class Description	
46 A	<i>open</i>	0.22	0.20	0	-	Extremely unstable
47 B	<i>open</i>	0.16	0.12	0	-	Moderately unstable
48 C	<i>open</i>	0.11	0.08	2x10 ⁻⁴	-0.5	Slightly unstable
49 D	<i>open</i>	0.08	0.06	0.015	-0.5	Neutral
50 E	<i>open</i>	0.06	0.03	3x10 ⁻⁴	-1	Slightly stable
51 F	<i>open</i>	0.04	0.016	3x10 ⁻⁴	-1	Moderately stable

52 *Briggs (1973).

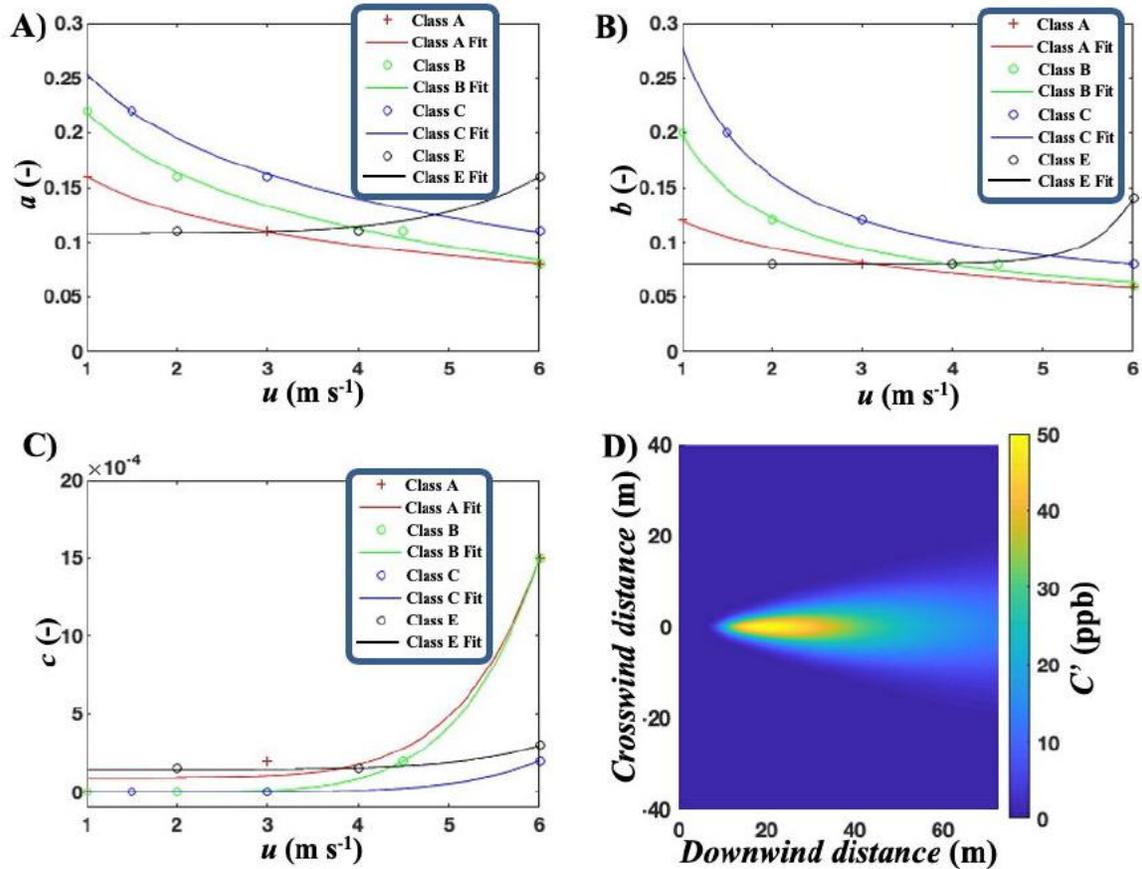
54 **Table S2.** Conditions defining Pasquill turbulence classes*

55 <i>u</i> m s ⁻¹	<i>I</i> _{Strong} >700 W m ⁻²	<i>I</i> _{Moderate} 350-700 W m ⁻²	<i>I</i> _{Weak} <350 W m ⁻²	Nighttime > ¹ / ₂ clouds	Nighttime < ³ / ₈ cloudiness
57 <2	A	A-B	B	E	F
58 2	A-B	B	C	E	F
59 4	B	B-C	C	D	E
60 6	C	C-D	D	D	D
61 >6	C	D	D	D	D

62 *from Hanna et al. (1982). *u* is wind speed. *I* is solar insolation.

63
64
65 The Gaussian plume model is for a passive dispersant and assumes that negligible along-wind
66 diffusion, that *u*, σ_y , and σ_z are vertically and horizontally uniform along its trajectory, that *u*
67 fluctuations are zero, and that *u* remains parallel to the *x* axis (no veering). These assumptions imply
68 an idealized, flat terrain of homogeneous roughness. Violations of these conditions are common in
69 terrestrial settings and can be addressed by model modification, e.g., Briggs (1973); Leifer et al.
70 (2016), the absence of topography and obstacles at sea implies that marine plumes generally satisfy
71 the Gaussian plume requirements.

72



73
 74 **Figure S2:** Pasquill turbulence parameters (Table S1) and 2nd-order polynomial least-squares linear-regression fits of A)
 75 a , B) b , and C) c for solar insolation stability classes (Table S1; class data key on figure) versus wind speed, u . See text
 76 for details. D) Gaussian plume surface concentration, C' , for 0.47 L min⁻¹ emissions, $u = 3.5$ m s⁻¹, and no boundary layer.

77
 78 Simulations were conducted for three conditions: infinite planetary boundary layer, i.e., no
 79 reflection, a 250-m BL , and a 100-m BL . The higher, 250-m BL only affects C' several kilometers
 80 downwind, whereas for $BL=100$ m, the reflection affects C' far closer to the source and far more
 81 strongly. The combined plume and *reflection* flatten the downwind plume near the axis, but with a
 82 sharper off-axis decrease in C' .

83

84 **S2 Focused seep areas**

85 **Table S3.** Location and direction from West Campus Station of informally-named seeps

86	<i>Seep Area</i>	θ	Latitude, Longitude
87	La Goleta Seep	152°	34° 23.503'N, 119° 51.193'W
88	Seep Tent Seep	198°	34° 23.063'N, 119° 53.428'W
89	Platform Holly	238°	34° 23.392'N, 119° 54.374' W
90	Trilogy Seep C	178°	34° 23.634'N, 119° 52.702'W
91	Trilogy Seep B	178°	34° 23.620'N, 119° 52.709'W
92	Trilogy Seep A	178°	34° 23.603'N, 119° 52.699'W
93	Patch Seep	140°	34° 21.850'N, 119° 49.755'W
94	Shane Seep	230°	34° 24.370'N, 119° 53.428'W
95	IV Super Seep	146°	34° 24.090'N, 119° 52.066'W
96	Tonya Seep	184°	34° 24.043'N, 119° 52.841'W
97	Horseshoe Seep	186°	34° 23.799'N, 119° 52.519'W
98	Rostocker Seep	99°	34° 24.230'N, 119° 50.438'W
99	Seadog Seep	240°	34° 24.172'N, 119° 54.374'W

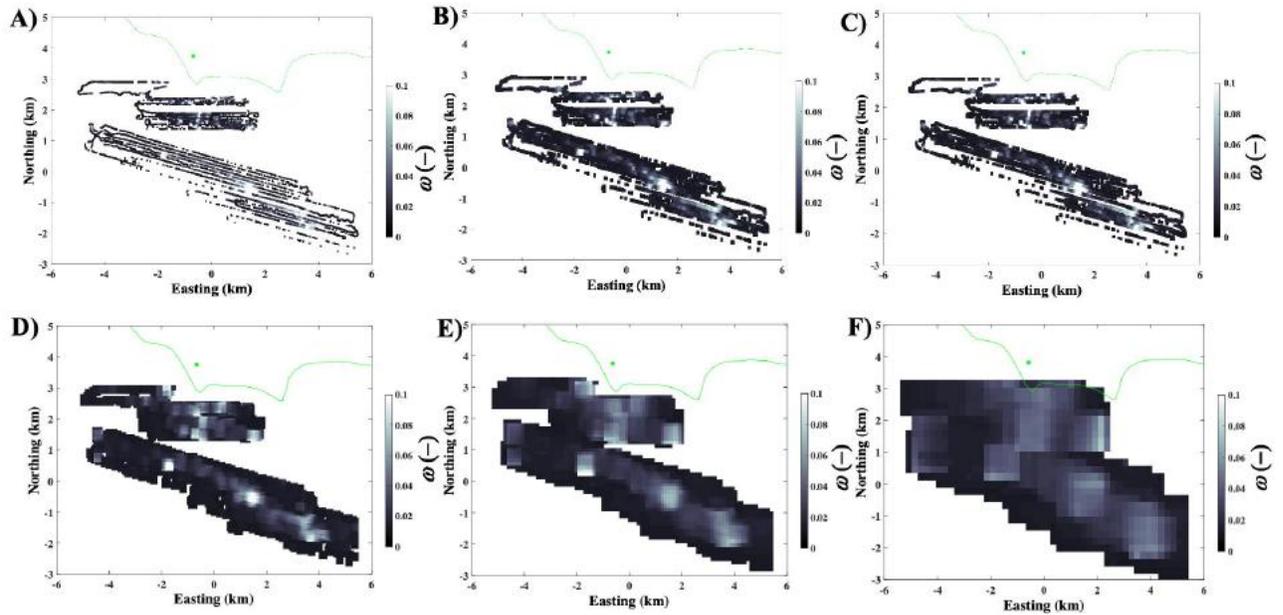
100 * θ – direction from West Campus Station (34.414949°N, 119.879690°W).

101

102 **S3 Sonar return**

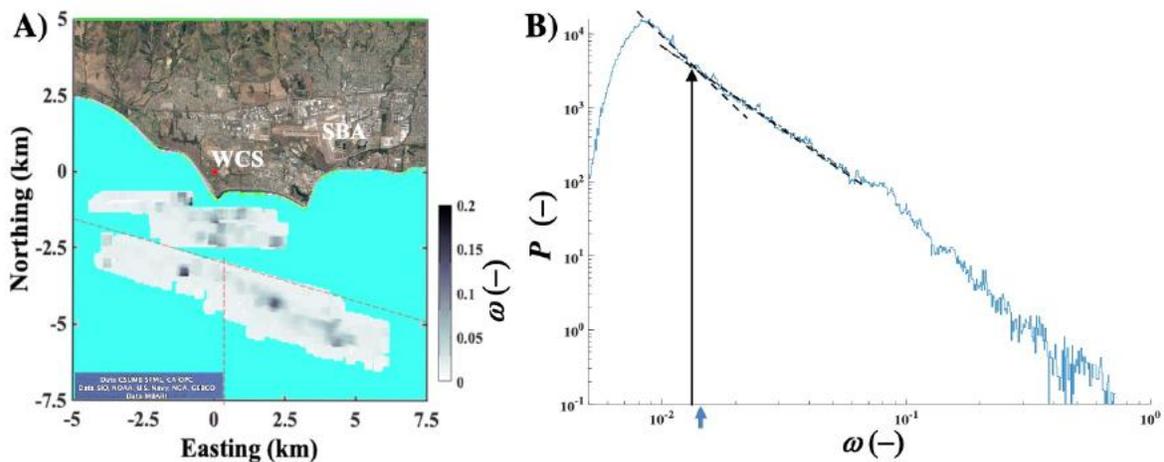
103 The Sept. 2005 seep field survey sonar return, ω , was gridded at spatial resolutions from 11 to
 104 225 m in a coordinate system with origin at West Campus Station, WCS, 34° 24.8969'N, 119°
 105 52.7814'W (**Fig. S3**). See Leifer et al. (2010) for details on the sonar survey data acquisition and
 106 analysis. Gridding involves averaging of all ω in each bin followed by a gap-filling low-pass filter.
 107 To gap fill, the center bin in a rolling 3x3-bin window which is empty, is replaced by the mean if
 108 there are more than 5 non-empty bins in the window. A hybrid 56/22-m gridding scheme gap filled
 109 the 22-m grid with the 56-m grid and was used in simulations (**Fig. S3C**).

110



111
 112 **Figure S3:** Sonar return, ω , maps for gridding at A) 11-m, B) 22 m, C) 22/56-m D), 56 m, E) 110 m, and F) 225 m. See
 113 text for details. Data key on panels.

114
 115 The noise level was 0.015 and was identified from a histogram of ω , where the ω probability
 116 distribution shows a power law shift and is the transition from seep distribution to noise domination
 117 (Fig. S4). Although there clearly is a seep contribution below the noise level, it cannot be segregated
 118 from noise and is neglected.
 119



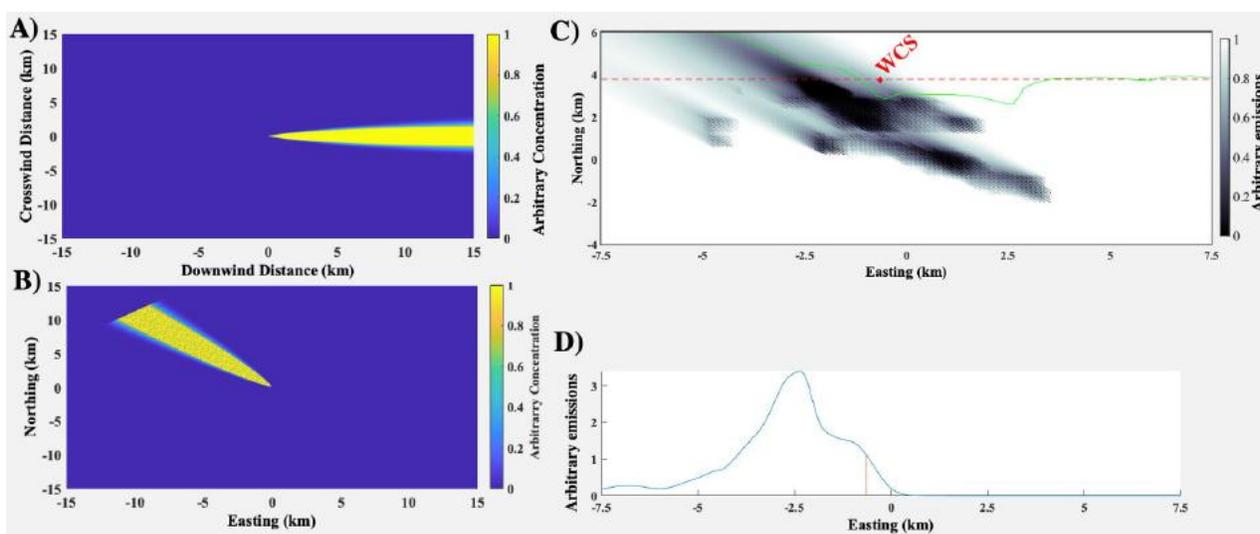
120
 121 **Figure S4:** A) 22/56-m gridded 2005 sonar return, ω . Green dot shows West Campus Station (WCS) location. Shown in
 122 the © Google Earth environment. B) Occurrence distribution, $P(\omega)$. Arrow shows noise level based on shift in power law
 123 fits to P at $\omega \sim 0.015$. Data key on panels. SBA - Santa Barbara Airport.

124

125 S4 Gaussian Plume model of an area plume

126 The model's core routines calculate a Gaussian plume (**Fig. S5A**) for the specified meteorological
127 conditions in a Cartesian coordinate system. Then, the plume is rotated to the wind direction, re-
128 gridded, and normalized to ensure mass conservation (**Fig. S5B**). The rotated plume is translated to
129 each grid cell, scaled to emissions for that grid cell, and finally, all plumes are added (**Fig. S5C**). C'
130 at WCS is calculated from the combined seep field emission plume map (**Fig. S5D**).

131



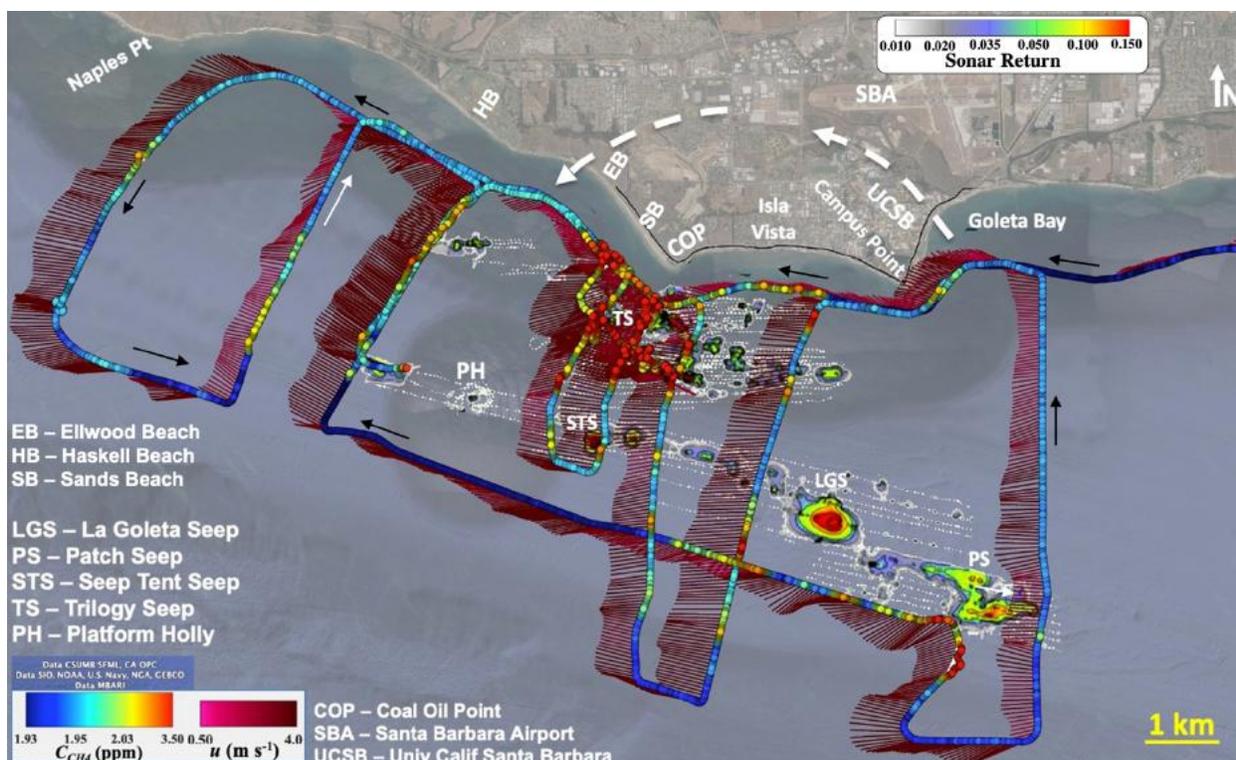
132 **Figure S5:** **A)** Gaussian plume concentration (arbitrary units) for 4 m s^{-1} wind speed and 250 m-thick boundary layer. **B)**
133 Plume rotated to a wind direction, θ , of 135° , re-gridded, and normalized. **C)** Assembled plumes from all grid cells. Dashed
134 red line shows a constant northing transect through West Campus Station, WCS. **D)** Concentration profile along WCS
135 transect (dashed red line in panel C). Red line shows location of WCS.
136

137

138 S5 Field atmospheric observations

139 Winds for 28 May 2016 were largely from the east with some modifications near the coast (**Fig.**
140 **S6**). Winds in Goleta Bay included a strong onshore component, which was absent offshore Isla Vista
141 where the shoreline features tall bluffs ($\sim 10\text{-}14 \text{ m}$). To the west of COP, winds had a strong offshore
142 component as far as halfway to Naples. Enhancements in the transect offshore west Sands Beach and
143 Haskell Beach are consistent with emissions from shallow waters, beaches, and/or coastal onshore
144 from seepage and/or abandoned oil wells.

145



146
 147 **Figure S6:** Methane, C_{CH_4} , and winds, u , for 28 May 2016. Sonar return map for reference in background. Dashed arrow
 148 shows proposed onshore winds; black arrows show travel direction. Shown in Google Earth environment. ©Google Earth.
 149 See **Table S3** for locations of informally named seeps.

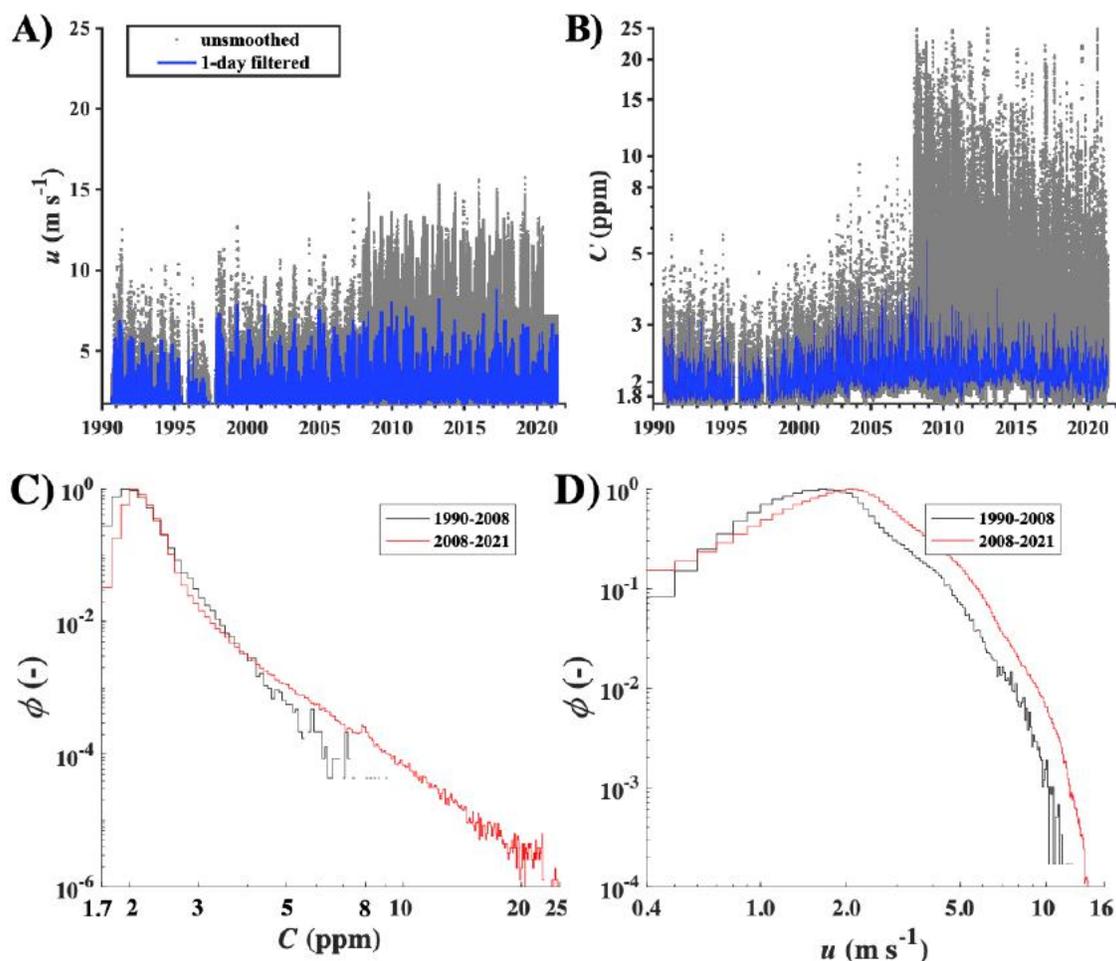
150
 151 On 28 May 2016 winds were strongly from the east, which is uncommon, and largely
 152 overwhelmed the sea breeze. The easterly winds flow onshore in Goleta Bay and then offshore around
 153 Ellwood and Sands Beach, to the west of COP. This likely relates to a combination of the sea breeze
 154 in Goleta Bay and the overall easterly flow which in turn drives an offshore flow (by continuity)
 155 around Ellwood and Haskell Beaches. This wind flow creates strong convergence and divergence air
 156 flows.

157
 158 **S6 West Campus Station data**

159 **S6.1 Trends**

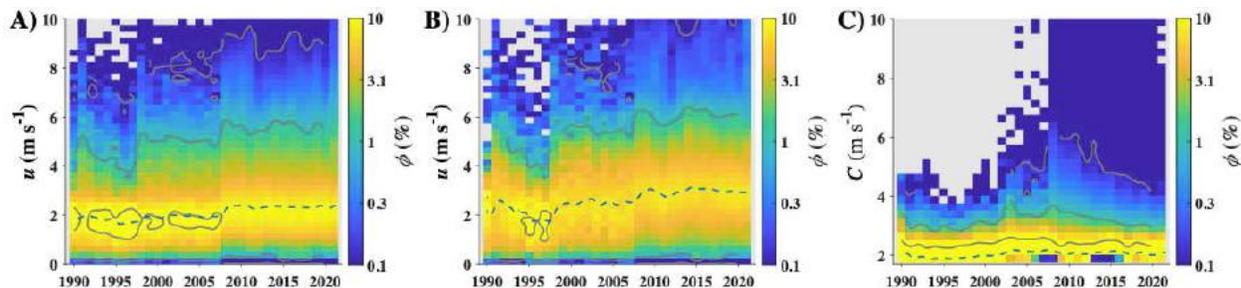
160 WCS data show a significant change in 2008 when the station was upgraded from 1-hour to 1-
 161 minute time resolution, allowing far higher values of total hydrocarbon concentration, THC, and u
 162 (**Fig. S7**). Hereafter, all concentrations and emissions are THC unless otherwise noted. To assess
 163 whether the station upgrade introduced biases, $\phi(u)$ prior and post 2008 were compared (**Fig. S7C**,

164 **7D)**. The slopes of $\phi(u)$ with respect to u for 2 to 4 m s⁻¹ are nearly identical for ϕ for the pre- and
 165 post-2008 data, the main change is for highest and weakest u , i.e., the range was extended. Moreover,
 166 several shifts between subsequent years for u and C have been documented (1997 and 2012 for u ,
 167 1997 and 2013 for C), i.e., such shifts in annualized values are not unique to 2008 (**Fig. S8**). This
 168 strongly argues that the changes in average, median, and baseline values post 2008 were from better
 169 measurement of gusts and weak winds and short positive C anomalies.
 170



171 **Figure S7:** WCS unfiltered and 24-hour smoothed data for **A)** wind speed, u , and **B)** concentration, C . WCS concentration
 172 probability distribution, ϕ , for before and after the 2008 upgrade for **C)** u and **D)** C . Data key on panel.
 173

174

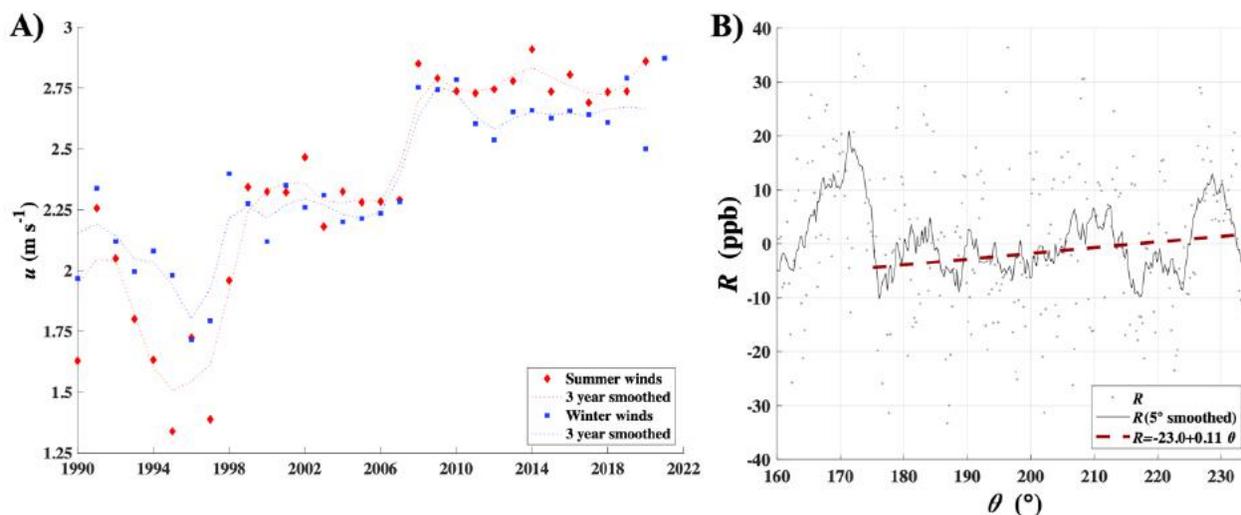


175
 176 **Figure S8:** Annual wind speed, u , probability, ϕ , for **A)** all wind directions, θ , **B)** seep directions ($135^\circ < \theta < 270^\circ$). Contours
 177 at $\phi=0.1, 1, 10\%$, calculated from 4-year smoothed data. **C)** ϕ for C for seep directions. Annual median (dashed line)
 178 shown on all panels. Data key for ϕ on figure.

179
 180 A linear regression analysis of the wind data showed increasing median u of $0.023 \text{ m s}^{-1} \text{ yr}^{-1}$ and
 181 $0.04 \text{ m s}^{-1} \text{ yr}^{-1}$ for all directions and for seep directions, respectively (**Fig. S9A**). This corresponds to
 182 an increasing sea breeze (from the seep direction). Summer winds have increased faster than winter
 183 winds, 0.043 versus 0.027 m s^{-1} , respectively.

184 The residual, R , to the Gaussian fit to $C'(\theta)$ showed an overall increasing linear spatial trend of
 185 $0.17 \text{ ppb degree}^{-1}$ from 175° to 210° (**Fig. S9B**). There is a peak at $R \sim 230^\circ$ that corresponds to Shane
 186 Seep and a peak at $R \sim 170^\circ$, which does not correspond to any named seeps (**Table S3**), but is in the
 187 general direction of the inshore seep trend to the east of the Trilogy Seeps (**Fig. 1**).

188



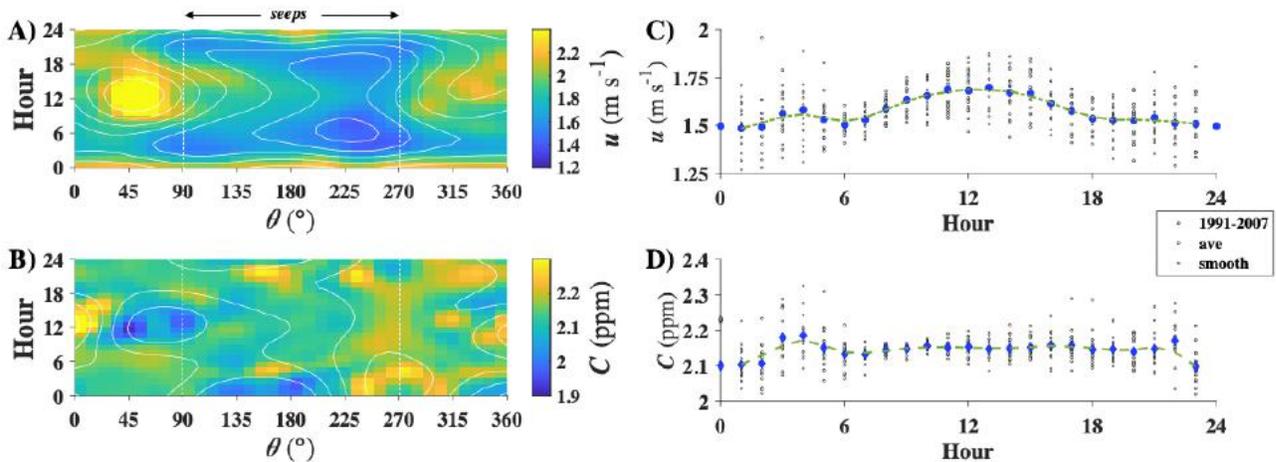
189
 190 **Figure S9:** **A)** Seasonally-segregated, annually-averaged wind speed, u for WCS. Summer is for Julian day 135-250,
 191 winter is Julian day 260-130, and three-year rolling-average. **B)** Residual, R , to dual Gaussian plume fit (**Fig. 4A**) to wind-
 192 direction resolved concentration, $C(\theta)$, for 1990-2019 for unsmoothed, 5° smoothed, and least-squares, linear-regression
 193 analysis fit. Data key on panel.

194

195 **S6.2 Diurnal**

196 Hourly-segregated averaged u and C for 1991-2007 (**Fig. S10**) show similar overall patterns to
197 2008-2021 for the seep directions (**Fig. 6**). The station improvements shifted averages, making
198 averaging the entire datasets inappropriate. One difference was the late-afternoon peak in C towards
199 the east-northeast, which was absent in the earlier period. The significant expansion of housing and
200 commercial developments in this direction in recent decades likely underlies this change.

201



202 **Figure S10:** 1991-2007 averaged wind direction, θ , and hourly-resolved **A)** wind speed, u , and **B)** concentration, C . **C)**
203 Seep-direction (90–270°), hourly-averaged wind speed, u_{seep} , and **D)** concentration, C_{seep} , averaged, individual years, and
204 3-year smoothed. Data key on figure. Midnight data missing due to daily calibration.
205

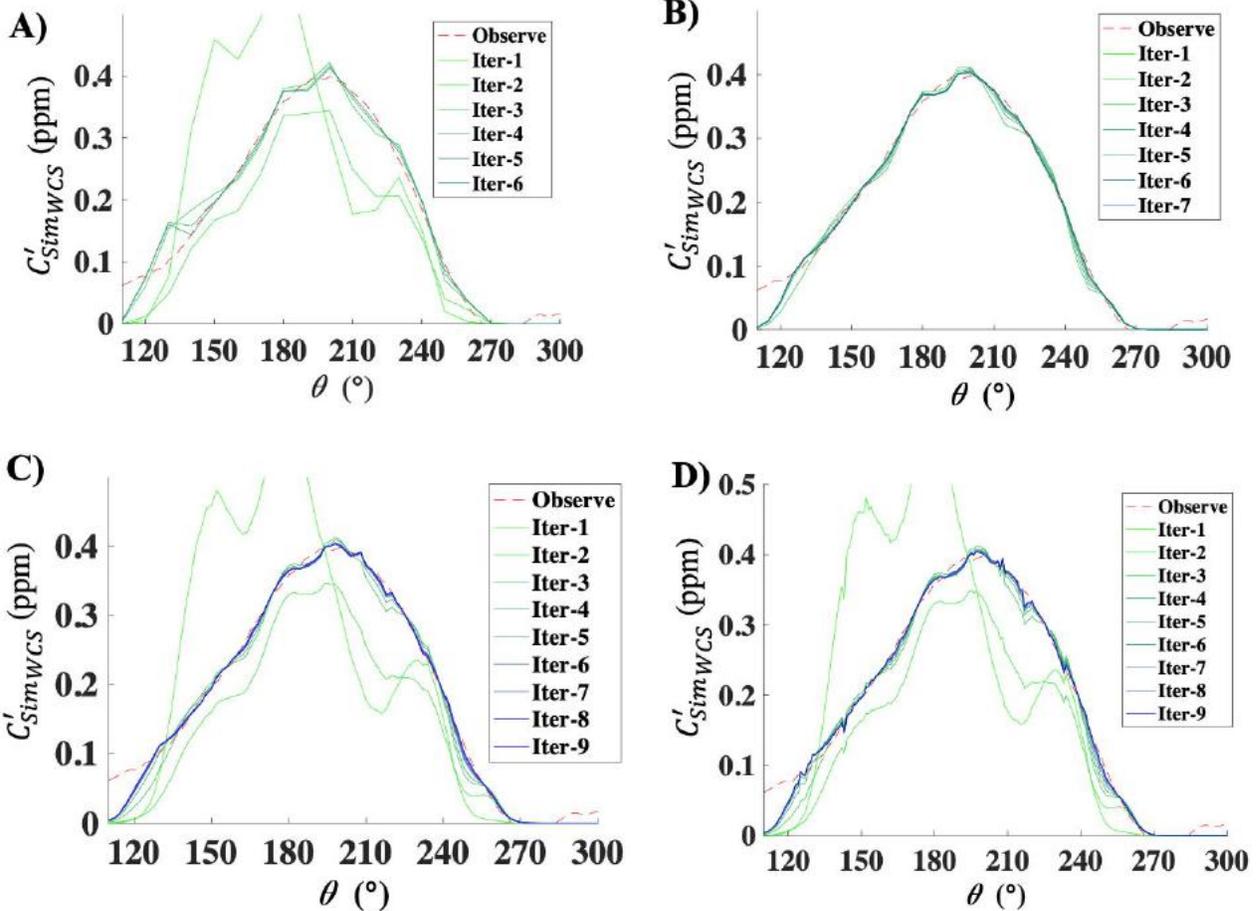
206

207 **S7 Seep field emissions**

208 **S7.1 Effect of model wind direction resolution**

209 Emissions simulations are run iteratively, initialized by the sonar distribution, which was
210 multiplied by a scaling factor, $K(\theta)$, initialized with $K=22,620$, that was applied to $E(x,y)$. Notably,
211 the θ -resolved C'_{sim} and C'_{obs} disagree significantly. This is unsurprising as the relationship between
212 ω and seabed emissions is complex (Leifer et al., 2017) with additional complexity arising from
213 transport across the water column. For sensitivity studies, the scaling factor, $K(\theta)$, see **Eqn. 3**, was
214 applied to $E(\theta)$ after each iteration until convergence (<1%), typically within 5 iterations (**Fig. S11**).

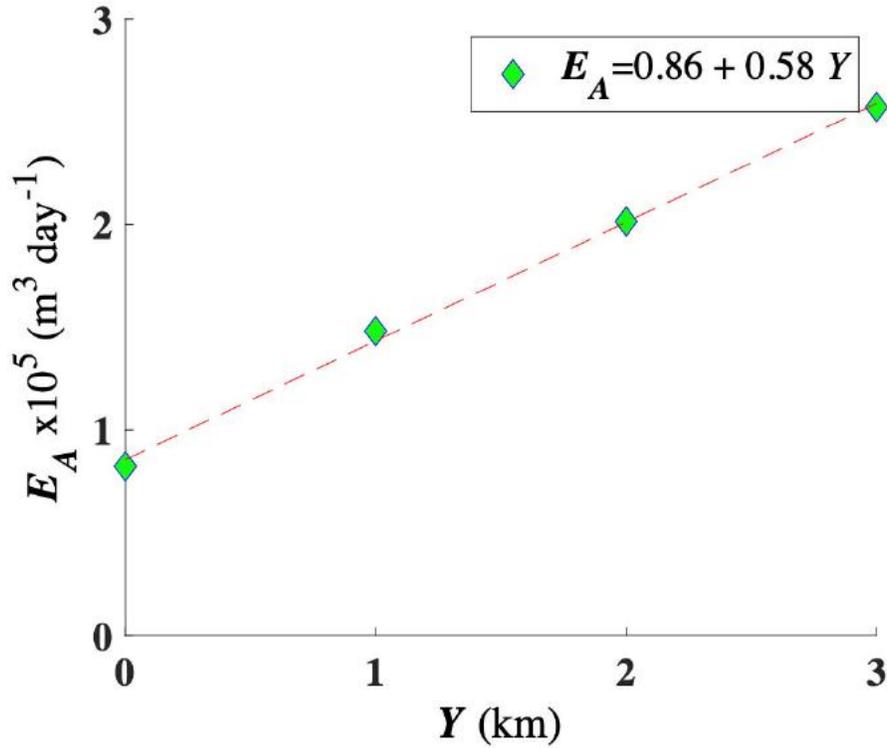
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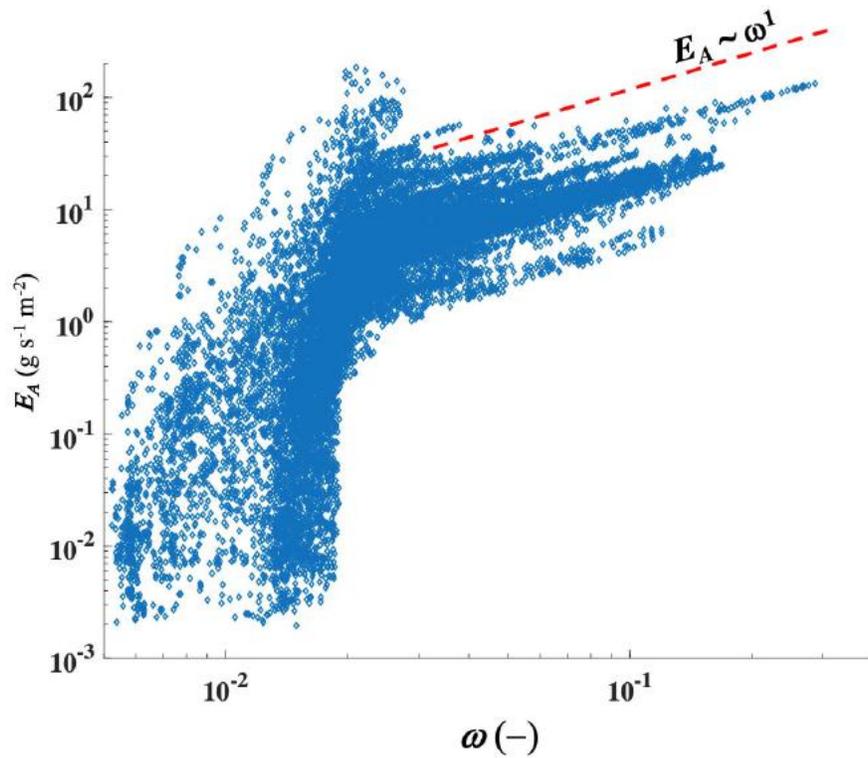
216
 217 **Figure S11:** Observed and simulated WCS concentration anomaly, C'_{Obs} and C'_{Sim} , respectively, versus wind direction, θ ,
 218 for all iterations for 22/56-m sonar map and angular, θ , resolutions, of **A)** 10°, **B)** 5°, **C)** 2°, and **D)** 1°. Data key on figure.

219
 220 Simulations were conducted with no distance correction, i.e., $K(r, \theta) = 1$ for north offsets of WCS
 221 location, Y , from 0 to 3 km. For reference, the inner and outer seep trends are 1.5 and 3.5 km from
 222 WCS at their closest direction. Simulations showed a near linear response in E_A to Y (**Fig. S12**), which
 223 was applied to as a distance scaling function, $K(r, \theta)$, in all further simulations.

224 S7.2 Effect of distance to seep field on emissions



225
226 **Figure S12:** Simulated atmospheric emissions, E_A , with respect to north offset, Y , modeled with no distance varying
227 correction factor, and linear fit shown on panel.



228
229 **Figure S13:** Sonar return, ω , versus derived atmospheric emissions, E_A . Line illustrates unity power law.

230

231 **S7.3 Sonar return versus atmospheric emissions**

232 Plotting all E_A versus ω shows a non-linear relationship (**Fig. S13**) with E_A increasing very steeply
233 with ω before rolling over for values of E_A in the range 1-10 g s⁻¹ m⁻². After rolling over, there were
234 several populations of ω -values that increased as a power law with an exponent slightly less than
235 unity. Emissions are the fraction of the seabed emissions that are transferred into the atmosphere,
236 including from adjacent pixels where emissions are from evasion – the simulation was initialized with
237 non-zero ω in neighboring cells to allow the simulation to infer sear-air gas evasion emissions from
238 these cells. This accounts for values of ω less than 0.02 in **Fig. S13**. The highest ω corresponds to the
239 strongest plumes where sonar tends to saturate (Leifer et al., 2017) but dissolved gas evasion
240 emissions to the atmosphere are more efficient due to the upwelling flow, which is more efficient in
241 strong plumes.

242

243 **S7.4 Emissions Uncertainty**

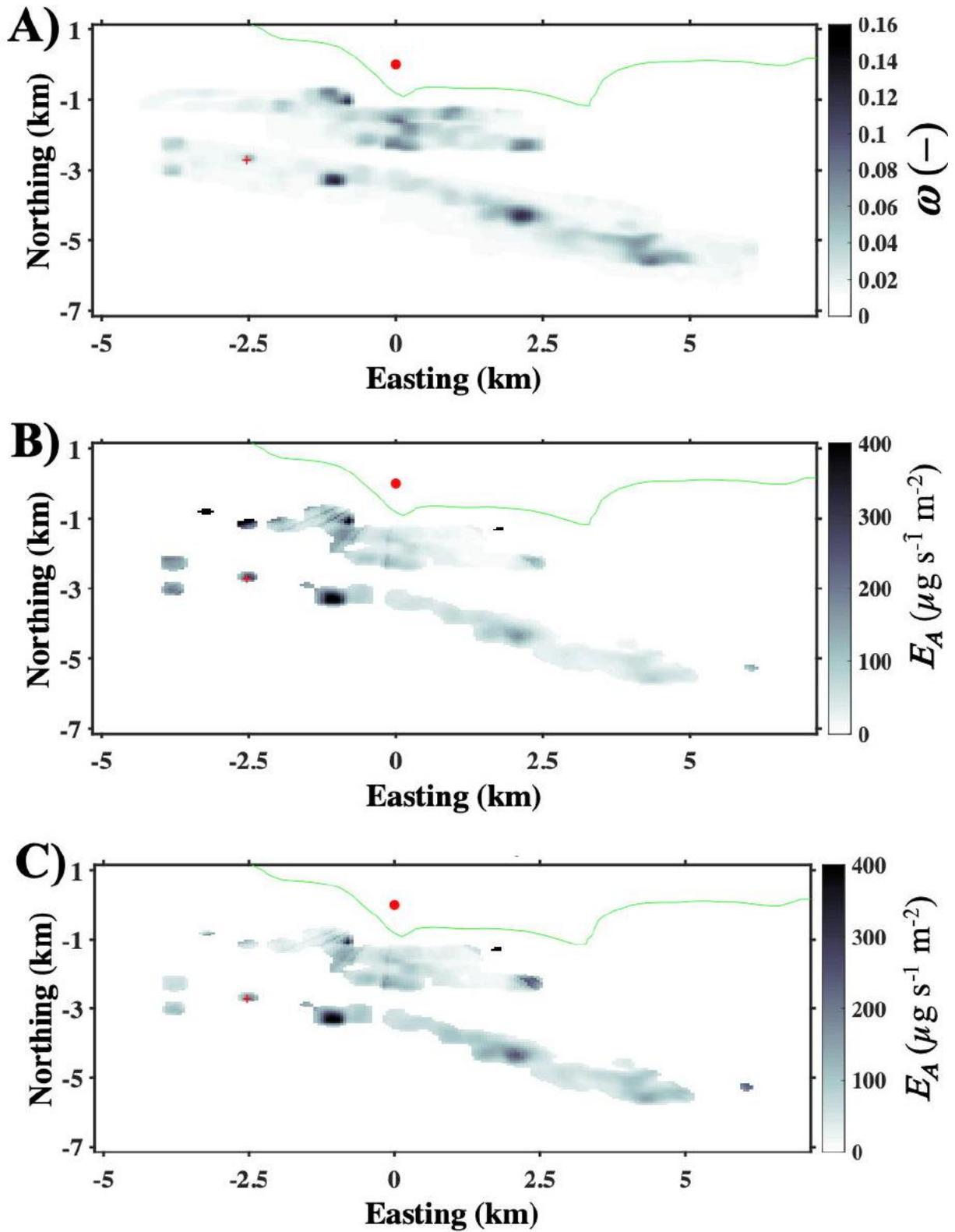
244 **S7.4.1 Angular resolution**

245 Sensitivity to angular resolution, $\delta\theta$, is very weak, <1% for reducing $\Delta\theta$ from 10° to 1°. For
246 example, a resolution of $\delta\theta=1^\circ$ produced artifacts and only changed E by 0.3% for reducing $\delta\theta$ from
247 2° to 1°. $\delta\theta$ resolution affects simulation resolution with coarse resolution blurring the fine-scale
248 sonar structure that is apparent in the 1° simulations (**Fig. S11D**); however, 1° simulations produced
249 overly-quantized results for the hybrid 22/56-m sonar maps including banding artifacts. Thus, θ
250 resolution was 2° outside of several sensitivity simulations.

251

252 **S7.4.2 Veering**

253 Veering, ψ , was implemented by adding ψ to θ in the calculation of $K(r, \theta)$, i.e., $K(r, \theta + \psi)$. Wind
254 veering emphasized emissions at the field's edges, clearly affecting the seep emissions map. $\psi=+10^\circ$
255 produced better agreement with the sonar map. Wind veering simulations (**Fig. S14**) overly
256 emphasized emissions at the seep field's edge towards which there was veering, although the no-wind
257 veering simulation also emphasized the western field's seepage. Overall, the $\psi = +10^\circ$ was an
258 improvement. Specifically, it improved the importance of La Goleta Seep compared to the no wind
259 veering simulation and also lessened the enhancement of the western seep field's edge (**Fig. 7B**).



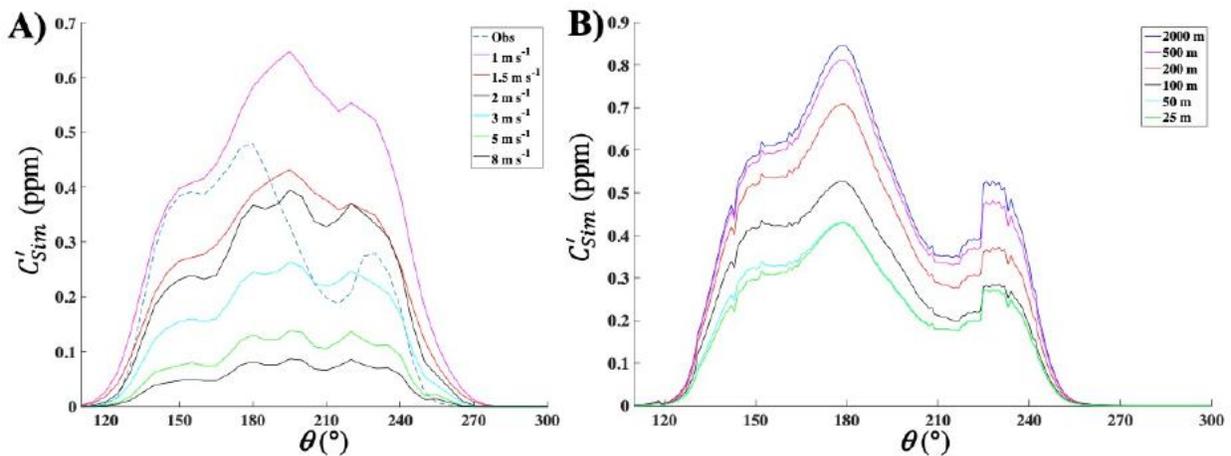
260
 261 **Figure S14:** A) Sonar return, ω , map gridded to 56-m resolution. B) Atmospheric emissions, E_A , map for wind veering,
 262 ψ , of B) $\psi=-10^\circ$ and C) $\psi=10^\circ$. Data key on figure.

263 E_A was extremely weakly sensitive to ψ , with $\pm 10^\circ$ veering corresponding to only a few percent
 264 in E_A . This low sensitivity arises because ψ simply shifts emissions to neighboring cells, only
 265 changing E_A when neighboring cells have no seepage forcing the model to increase emissions from
 266 more distant cells.

267
 268 **S7.4.3 Wind speed**

269 There is a strong sensitivity of E_A to u , particularly for small negative variations, although
 270 sensitivity saturated at $\pm 10\%$ with greater sensitivity for lower u . This sensitivity was shown by a
 271 series of non-iterative simulations of $C'(\theta)$ for a range of u , showing variations were most strong for
 272 very weak wind speeds. Observed u span the range of weak for the east and south to moderate for
 273 prevailing (**Fig. S15A**). Instrumentation uncertainty in u is a few percent, thus uncertainty in u arises
 274 if WCS winds are unrepresentative due to acceleration or deceleration between the seep field and
 275 WCS. C' is highest for $u = 1 \text{ m s}^{-1}$, which is weaker than actual winds, whereas C' for $u = 8 \text{ m s}^{-1}$ is
 276 about a tenth C' for $u = 1 \text{ m s}^{-1}$. At lower u , there is more “blurring” of plumes from different sources
 277 showing less defined structure than higher u . However, u and BL are not independent – by continuity,
 278 an increase in u – i.e., acceleration, requires a decrease in BL . Thus, BL sensitivity largely counters u
 279 sensitivity, discussed in **Sec. S8.5**.

280



281
 282 **Figure S15:** Simulated concentration anomaly, C'_{Sim} , versus wind direction, θ , for WCS for different **A)** wind speed, u ,
 283 including for the observed $u(\theta)$. **B)** Boundary layer height. $\delta\theta$ resolution is 1° . Data key on figure.

284

285 **S7.4.4 Boundary layer**

286 *BL* sensitivity is weakly non-linear and not strong - a 40% reduction in *BL* (250 to 150 m) only
287 decreased E_A by 17%, less so for a 100 m increase to 350 m in *BL* (**Fig. S15B**). The effect of large
288 *BL* changes saturates – for very high *BL*, it has no effect, and for very shallow *BL*, multiple reflections
289 homogenize the boundary layer leaving only dilution driving *BL* sensitivity. Thus, a factor of 40
290 variation in *BL* corresponds to a 200% change in $C'(\theta)$. Although the nocturnal *BL* is significantly
291 shallower, nocturnal winds largely are the (northerly) land breeze and thus uncommon for the seep
292 direction. *BL* values for the seep field and WCS are uncharacterized, thus, a 2nd-order polynomial was
293 fit to *BL* versus E_A and averaged (absolute) over possible ranges. For $200 < BL < 300$ m, the mean
294 absolute value was 3%, 6% for $150 < BL < 350$ m.

295

296 **S7.4.5 Concentration**

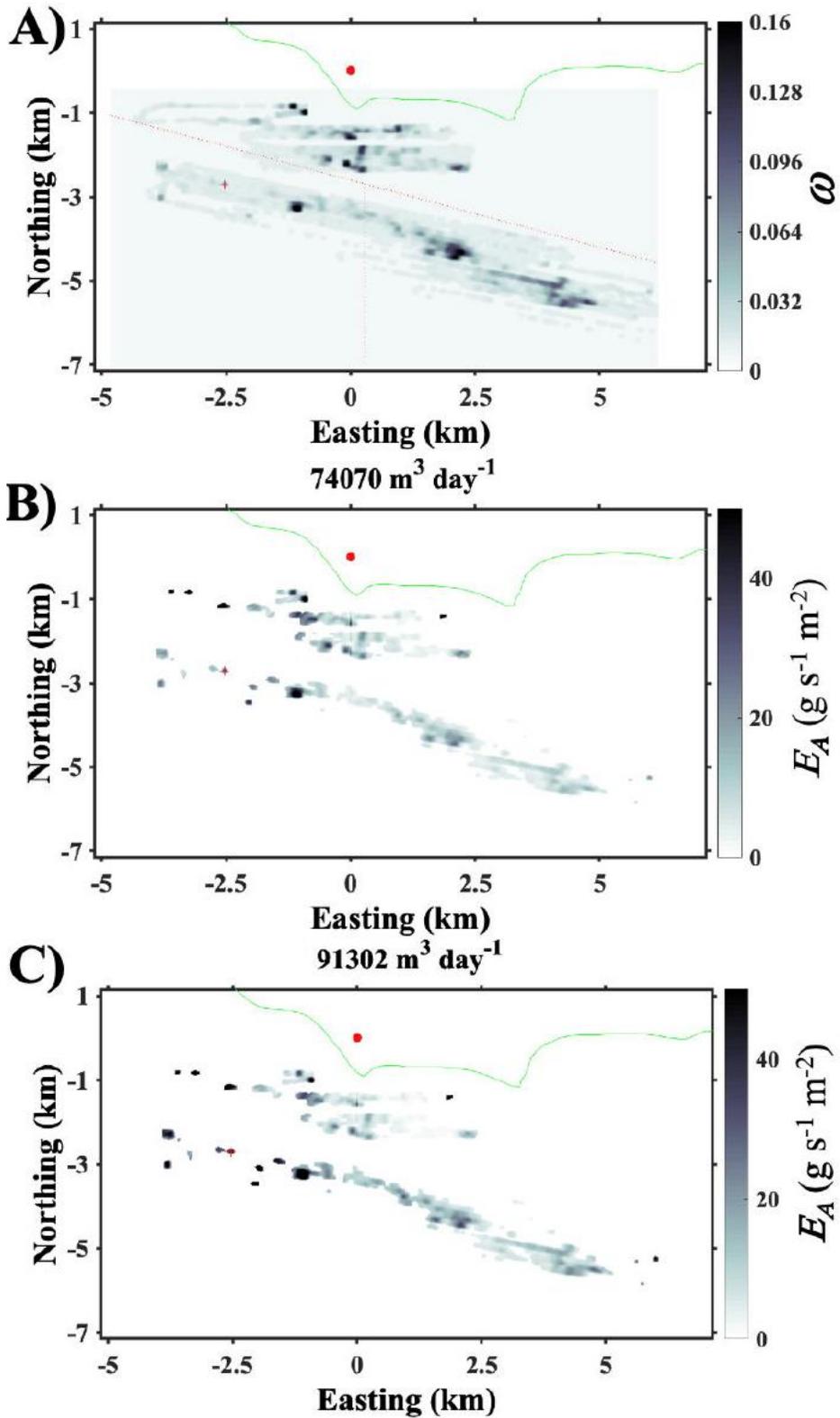
297 E_A sensitivity to C' is weaker than that of u , particularly for small variations and is linear (**Fig. 9**).
298 C uncertainty arises from drift in the flame ion detector between nightly calibrations. Thus, the
299 difference in concentration before and after calibration is the daily drift as well as any changes in
300 ambient, which should on average be zero, particularly for around midnight when the atmosphere is
301 stable. A 500-point histogram of the total daily drift was calculated and values $< \pm 0.2$ ppb were fit
302 with a Gaussian function (larger deviations exhibited a broader distribution which, likely resulted
303 from ambient trends) and found a half-width of 170 ppb, or ± 85 ppb. Thus, the average drift is ~ 40
304 ppb dy^{-1} , reducing for annualized values to ~ 2 ppb dy^{-1} ($1/\sqrt{365}$), i.e., $< 1\%$ of the C' value of 300
305 ppb. As such, uncertainty in C' is not a significant contributor to uncertainty in E_A .

306

307 **S7.4.6 Combined wind and boundary Layer**

308 E_A sensitivity to *BL* largely counters E_A sensitivity to u (**Fig. 9D**), even when mismatched – a
309 $\pm 40\%$ decrease in *BL* and $\pm 20\%$ increase in u corresponds to a 6% and 9% decrease in E_A , i.e., less
310 than half the mismatch. By continuity, this should be quite small, reflecting lesser changes in pressure,
311 temperature, and topographic forcing. Underlying this weak sensitivity is that the linear sensitivity of
312 E_A to u largely arises from dilution as for *BL*. Note, this weak sensitivity is for a change in *BL* from
313 the offshore to onshore and the resultant change on winds, not a change in the overall (regional) *BL*.

314



315
 316 **Figure S16:** A) Sonar return, ω , map gridded to 56/22-m resolution. Atmospheric emissions, E_A , for inshore/offshore
 317 trends relative enhancement, ζ , for **B)** $\zeta=-25\%$ and **C)** $\zeta=+25\%$ offshore enhancement.

318

319 **S7.4.7 Inshore / offshore partitioning**

320 The model cannot determine how to apportion emissions between the inshore and offshore seeps.
321 Simulations found a complex relationship between E_A and ζ (**Fig. 9**) because simulations re-assign
322 some emissions between the seep trends depending on the presence or absence of seepage for θ where
323 there is no seepage on one of the two trends (**Fig. S16**). As such, a potentially strong sensitivity arises
324 from the offshore and onshore seepage trend partitioning with respect to θ . Initial simulations that
325 used a uniform distance partitioning, unrealistically increased inshore seepage compared to ω . Given
326 the similarity in the ω probability distribution between the inshore and offshore trends (**Fig. 8**), i.e.,
327 whatever the functional relationship between E_A and ω is, similar ω implies similar E_A .

328 Sensitivity simulations were run for various shifts in the relative emissions from the inshore and
329 offshore trends – i.e., for +10% inshore -10% offshore, while maintaining a linear variation in $K(r, \theta)$.
330 Sensitivity is strong, and complex – interestingly, reversing between -10° and 10° . This behavior is
331 due to shifting emissions onto different seep areas in the inshore or offshore seep area. As such, for -
332 $30\% < \zeta < +25\%$, uncertainty ranges from -10% to +5%. Although higher changes in ζ are certainly
333 feasible, both trends source from the same formation, which means on a 30-year time scale (or much
334 shorter), changes in one seep trend likely reflect trends in all seep areas.

335 Additionally, seepage could shift within the trends – nearer or further - which was assessed by
336 the simulations of north shifts of WCS (**Fig. S12**). Total trend widths are ~ 300 m for the offshore and
337 600 m for the inshore, or an average distance deviation of ± 75 and 150 m, corresponding to E_A
338 variations of a few percent.

339

340 **S8 Future data needs and model improvements**

341 **S8.1 Key data needs**

342 The driver of uncertainty was whether emissions arose from the inshore or offshore seep trends,
343 which could not be determined from WCS data. A second air quality station located towards the seep
344 field east side would allow triangulation. If the station included a ceilometer, it also could address BL
345 uncertainty in E_A . Long-term BL data would characterize the relationship to solar insolation, diurnal
346 cycle, and winds, allowing application to historical data.

347 Additionally, THC is a less than ideal seep emission indicator as it can arise from other sources,
348 such as motor vehicles or fires – the station should include speciation, ideally a combination of CEAS
349 analyzers and a gas chromatograph.

350 The diurnal cycle is poorly constrained by WCS data given that the typical diurnal cycle of
351 strengthening winds in the afternoon also shifts towards the west, with seep field emissions bypassing
352 WCS. Given that an atmospheric plume survey can derive emissions in a couple of minutes, multiple
353 seep areas should be repeat surveyed from morning through early evening to characterize the
354 relationship between wind speed and emissions.

355 The study focus was marine seepage and neglected terrestrial seepage; however, the data analysis
356 demonstrated that there are notable terrestrial emissions. Further WCS data analysis should assess
357 emissions from Ellwood Field, with onshore surveys searching for seep emissions in neighborhoods
358 by measuring seep gases in the near downwind of abandoned wells and faults, particularly those that
359 intersect with the Ellwood Field.

360 Finally, the downcurrent dissolved plume's fate is poorly understood as there is an absence of
361 field data collected under typical afternoon winds. A combination of field data collected downwind
362 of the downcurrent plume, ideally in conjunction with water sampling, and further modeling of WCS
363 data can help constrain this emission source and better characterize the fate of COP seep field
364 emissions.

365

366 **S8.2 Model improvements**

367 The area plume inversion model was used to derive emissions on the field scale and sector scale.
368 Efforts to derive emissions at finer spatial scales were stymied by uncertainty in implementation of
369 wind veering. At the field scale, there was very low sensitivity of E to wind veering (**Fig. S14**);
370 however, simulations without wind veering shifted emissions off of important seeps like Trilogy Seep
371 to nearby seeps that are much smaller than Trilogy Seeps. Additionally, the model assigned emissions
372 to the field edges. Simulations suggested that wind veering that varies with wind direction is needed
373 – consistent with the field observations, i.e., wind veering for the southeast and southwest are
374 different. This is not surprising given the different orientations of the coast on the east and west sides
375 of COP, as well as the topography between WCS and the coast - bluffs (~20 m) on the east side, and
376 their lack to the west (Sands Beach). Additionally, further offshore, wind veering likely becomes less
377 varying across the field and also shifts more towards prevailing. Further field observations are needed

378 to correctly characterize and incorporate wind veering to allow simulations to characterize changes
379 to different areas of the field.

380 The model's enhancement of seepage at the seep field edges likely is from a combination of wind
381 veering and outgassing emissions from downcurrent plumes. Currently, outgassing in the seep field
382 area is assigned to the nearest seep area along the wind direction. This causes a problem for wind
383 directions in the seep field with no seepage in that direction, including beyond the seep field edges.

384 Attempts to uniformly assign emissions with distance – i.e., no preferences – led to an unrealistic
385 shift of emissions to the inshore seeps (based on a relative sonar return). A series of simulations where
386 WCS was “shifted” north showed a linear dependency on emissions, and thus the model assigned
387 emissions in any direction linearly with distance. One approach would be to prioritize major seepage
388 over minor seepage in the other seep trend. Such an approach should follow a series of field emissions
389 validation surveys.

390 There are a number of other model improvements that were not implemented due to
391 computational limits – current simulations take multiple days on a fast workstation (16-core
392 PowerMac, 96 GB RAM, 3.4 GHz). For example, rather than simulating the mean value of winds in
393 a direction, the probability distribution of u in each wind direction, e.g., **Fig. 5A**, could be used,
394 calculating essentially a weighted concentration for comparison with the observed concentration at
395 WCS. This would significantly increase computational load. An additional improvement would be to
396 include a boundary layer height that varied with wind speed and the diurnal cycle, and turbulence
397 parameterizations that are based on solar insolation. Values of boundary layer height and solar
398 insolation could be retrieved from NOAA weather models. It should be noted that the current model
399 complexity required several days to simulate at the highest spatial resolutions on a fast workstation.
400 Finally, higher angular resolution and sonar map resolution (doing one without the other causes non-
401 physical striation) improves the ability to capture changes at different seep areas, but the increased
402 angular resolution introduces more gaps in the seep emissions source function. This could be
403 addressed by adding a “dispersed” source of sea air gas evasion that follows currents.

404

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