



# Airborne measurements and large-eddy simulations of small-scale Gravity Waves at the tropopause inversion layer over Scandinavia

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Abstract. Coordinated airborne measurements were performed by the two research aircraft DLR Falcon and HALO (High Altitude and Long Range Aircraft) in Scandinavia during the GW-LCYCLE II (Investigation of the life cycle of gravity waves) campaign in 2016 to investigate gravity wave processes in the upper troposphere and lower stratosphere (UTLS) region. A mountain wave event was probed over Southern Scandinavia on 28 January 2016. The collected dataset constitutes a valuable

- 5 combination of in-situ measurements and horizontal- and altitude-resolved wind lidar and water vapour lidar measurements. In-situ data at different flight altitudes and downward pointing Doppler wind lidar measurements show pronounced changes of the horizontal scales in the vertical velocity field and of the leg-averaged momentum fluxes (MF) in the UTLS region. The vertical velocity field was dominated by small horizontal scales with a decrease from around 20 km to < 10 km in the vicinity of the tropopause inversion layer (TIL). These small scales were also found in the water vapour and reflectivity data. The
- 10 MF profile downstream of the main mountain ridge determined from the wind lidar data is characterized by negative fluxes below and positive fluxes above the TIL which show similar magnitudes. The combination of the observations and idealized large-eddy simulations (LES) revealed the occurrence of interfacial waves on the tropopause inversion during the MW event. Such interfacial waves have already been observed on boundary-layer inversions but their concept has not been applied to tropopause inversions so far. Our idealized simulations revealed that interfacial waves can occur also on tropopause inversions
- 15 and our analyses of the horizontal- and altitude-resolved airborne observations revealed that they actually do. As predicted by linear theory, the horizontal scale of those waves is determined by the wind and stability conditions above the inversion. They are found downstream of the main mountain peaks and their characteristic MF profile clearly distinguishes from the MF profile of a Kelvin-Helmholtz instability. Further, the idealized large-eddy simulations revealed that the presence of the TIL is crucial in producing this kind of trapped waves at tropopause altitude.

#### 20 1 Introduction

Gravity waves (GWs) are an important coupling mechanism between the lower and the middle and upper atmosphere. Propagating GWs transport momentum and energy and deposit them in regions where breaking and dissipation occurs. As such, GWs account for example for the well known upper mesospheric wind reversals as well as the cold polar summer mesopause and the warm winter stratopause (Dunkerton, 1978; Lindzen, 1981). So far, different sources for GWs in the troposphere have

25 been identified, e.g., flow over orography, convection, jets and fronts as well as secondary generation in the region of GW





breaking (Smith, 1979; Gill, 1982; Baines, 1995; Fritts and Alexander, 2003; Sutherland, 2010; Plougonven and Zhang, 2014; Vadas et al., 2003). Turbulence caused by breaking GWs and wave-induced modulations of the ambient wind can be hazards to aviation in the troposphere and lower stratosphere (Sharman et al., 2012; Bramberger et al., 2018). GWs are propagating from their sources in the troposphere and the tropopause region (Sato et al., 2009; Fritts et al., 2016). However, the atmospheric temperature and wind structures influence the propagation of GWs and alter their properties. The vertical distance from the

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temperature and wind structures influence the propagation of GWs and alter their properties. The vertical distance from the lower troposphere to the most prominent breaking region around the mesopause is about 70 km where changes in wind and temperature occur (Eckermann, 1995; Doyle and Jiang, 2006; Eckermann et al., 2007).

Since the works of Queney (1948) and Scorer (1949), different theoretical and numerical methods were used to study mountain wave (MW) propagation in an atmosphere of vertically varying wind and stability. Those methods are ray tracing,

- 35 determination of reflection and transmission coefficients, numerical solutions of the vertical structure equation for single wave components, analytical models, and asymptotic analysis (see Wurtele et al. (1987) and references therein, and more recently, Eckermann and Marks (1997); Gisinger et al. (2017); Achatz et al. (2010)). An important result of the theoretical investigations is that the stratospheric solution in a model taking into account a vertically varying background is not dominated by the classical solution of Queney (1948) but by reflected and downstream propagating (trapped) waves in the troposphere (Wurtele et al., 1987)
- 40 1987; Keller, 1994). This makes the wave spectrum (i.e. wavelengths) being determined by the vertical varying wind and stability and not by the topography spectrum which affects the relative amplitudes (Keller, 1994; Ralph et al., 1997). Fine scale structures in the atmosphere, such as sharp temperature inversions at the top of the boundary layer (Vosper, 2004; Sachsperger et al., 2015) or in the mesosphere (Fritts et al., 2018) can be wave guides leading to trapped waves which propagate horizontally along the inversions, i.e. interfacial waves. All those findings are in contrast to the fundamental characteristics of the hydrostatic
- 45 approximation. They are the absence of a mechanism which allows a wave to propagate horizontally and the consequent upward propagation of energy directly above the obstacle, regardless of the generating terrain (Wurtele et al., 1996).

Currently much activity using various ground-based, airborne, and satellite measurements is going on to get a complete picture of the GW activity and distribution around the globe and to enhance the understanding of source and propagation processes (e.g., Fritts et al., 2016; Podglajen et al., 2016; Wright et al., 2017; Shibuya et al., 2017; Kaifler et al., 2017; Krisch

- 50 et al., 2017). This knowledge is required to adequately model and parameterize atmospheric GWs in weather and climate models. So far, observational indications of GW behaviour in the tropopause region such as reflection and trapping are rare due to lack of horizontal- and altitude-resolved observations in the tropopause region. Using aircraft measurements, which were taken during the Terrain-Induced Rotor Experiment (T-REX, Grubišić et al. (2008)), Smith et al. (2008) were able use in-situ data to measure partial reflection of MWs at the tropopause for the first time. Using their linear model, they identified two levels
- of reflection, one at the altitude where the Scorer parameter changes due to changes in static stability and the other at the altitude with a discontinuity in wind speed but constant Scorer parameter. In addition, Smith et al. (2008) and Woods and Smith (2010) found signatures of trapped waves with a horizontal wavelength of about 15 km in the in-situ measurements in the tropopause inversion layer (TIL) during T-REX. They argue that the Sierra mountain range is unlikely to be the source for those 15-km waves. They do not expect those waves to reach the tropopause altitude from below because the background conditions caused
- 60 a considerable evanescent decay. Instead, they suggest that those waves were generated by a nonlinear steepening process.



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Follow-up model simulations lead to two different explanations. First, the short-wavelike fluctuations observed in the UTLS region are due to Kelvin-Helmholtz instability along shear lines locally induced by the primary mountain wave, i.e. they are not trapped GWs but instead small-scale wave motions resulting from Kelvin-Helmholtz instability (Mahalov et al., 2011). Second, the downward propagating GWs, which are created by MW breaking in the middle stratosphere, and their reflection at the tropppause can create kind of lee wave trapping in the lower stratosphere (Woods and Smith, 2011).

Coordinated airborne measurements were performed by the two research aircraft DLR Falcon and HALO (High Altitude and Long Range Aircraft) in Scandinavia during the GW-LCYCLE II (Investigation of the life cycle of gravity waves) campaign in 2016 to investigate GW processes in the upper troposphere and lower stratosphere (UTLS) region. A mountain wave event was probed over Southern Scandinavia on 28 January 2016. The collected dataset constitutes a valuable combination of in-

70 situ measurements and horizontal- and altitude-resolved wind lidar and water vapour lidar measurements in the UTLS. In-situ data at different flight altitudes and downward pointing Doppler wind lidar measurements revealed pronounced changes of the horizontal scales in the vertical velocity field and of the leg-averaged momentum fluxes (MF) in the UTLS region.

This paper examines the MW case over Scandinavia by means of ECMWF IFS meteorological analyses and the coordinated airborne measurements of the DLR Falcon and HALO which provide horizontal- and altitude-resolved data in the UTLS. The

- 75 wind data of the downward pointing Doppler lidar give the opportunity to calculate a continuous profile of MF over a twokilometer altitude range in the UTLS. In order to find out what determines the horizontal scales in the vertical velocity field and which process(es) can explain the observed characteristics, we investigate the possible existence of interfacial waves on the TIL (Vosper, 2004; Sachsperger et al., 2015, 2017) similar to their existence on an inversion in the troposphere (Cruette, 1976; Sachsperger et al., 2015; Chouza et al., 2015). The paper is organized as follows. Section 2 describes the models, data
- and methods used in this paper. Meteorological conditions and observations of the MW event on 28 January 2016 are analyzed in Section 3.1 and idealized large-eddy simulations of MW propagation in the presence of atmospheric inversions are presented in Section 3.2. Section 4 gives a tight discussion of the results and Section 5 concludes the paper.

# 2 Data and Methods

#### 2.1 ECMWF global analysis

Operational analyses of the ECMWF integrated forecast system (IFS) are used to analyse the meteorological conditions on 28 January 2016. These analyses (cycle 41r1<sup>1</sup>) have a horizontal resolution of about 16 km on the reduced linear Gaussian grid ( $T_L$ 1279). The highest of the 137 vertical levels (L137) is located at ~ 80 km (0.01 hPa). The layer thicknesses gradually increases from ~ 300 m at ~ 10-km altitude to ~ 400 m at ~ 20-km altitude and ~ 2 km at ~ 60-km altitude<sup>2</sup>.

<sup>&</sup>lt;sup>1</sup>https://www.ecmwf.int/en/forecasts/documentation-and-support/changes-ecmwf-model/cy41r1-summary-changes, last access Oct 2018 <sup>2</sup>https://www.ecmwf.int/en/forecasts/documentation-and-support/137-model-levels, last access Oct 2018





# 2.2 Airborne observations

#### 90 2.2.1 Coordinated research flights on 28 January 2016

The airborne observations took place during the intensive observing period 6 (IOP 6) on 28 January 2016 in the framework of the combined missions POLSTRACC (The Polar Stratosphere in a Changing Climate), GW-LCYCLE II and SALSA (Seasonality of Air mass transport and origin in the Lowermost Stratosphere using the HALO Aircraft). An overview of the performed HALO research flights can be found in Oelhaf et al. (2019). In January 2016, the DLR research aircraft Falcon and HALO

- 95 operated from the airport of Kiruna (67.82° N, 20.33° E), northern Sweden, to investigate chemical and dynamical processes in the UTLS region at high latitudes. The goal of IOP 6 was to measure a transient MW event over southern Scandinavia with coordinated cross-mountain flights of both aircraft. Figure 1a shows the operational area and the flight tracks of the research flights. The mountains were crossed at the same latitude two times by the DLR Falcon (flight legs RF07 FL2 and RF08 FL1) and three times by HALO (flight legs HL1, HL2 and HL4; Fig.1b). The limited range of the DLR Falcon required a refuel stop
- 100 at Karlstad airport. The DLR Falcon was flying close to the thermal tropopause on all flight legs and measured vertical winds and GW induced momentum fluxes with the in-situ sensor at flight altitude and with the downward pointing Doppler wind lidar below the aircraft. On the flight legs HL1 and HL2 HALO was flying in the troposphere (HL1 below the DLR Falcon) and measured wave structures at flight level and in the tropopause region with the upward pointing differential absorption lidar (DIAL) measuring water vapour (WALES) (Wirth et al., 2009).

#### 105 2.2.2 Wind lidar measurements

The DLR Falcon was among others equipped with a downward-looking coherent Doppler wind lidar (DWL) which has been operated by DLR since 1999 and which has been applied in a number of field campaigns (e.g., Chouza et al., 2017; Schäfler et al., 2018; Marksteiner et al., 2018; Lux et al., 2018). The DWL operates at a wavelength of 2  $\mu$ m and is equipped with a double-wedge scanner which enables to steer the laser beam to any position within a cone angle of 30°. A more detailed

110 description of the 2 μm DWL instrumental setup, the measurement principle, the applied retrieval algorithms and the accuracy and precision of the derived wind products is given by Chouza et al. (2015), Witschas et al. (2017), and more recently Witschas et al. (2020).

Usually, the 2 µm DWL is used to either measure the three dimensional wind vector by applying the velocity azimuth display (VAD) scan technique, or to measure vertical wind speeds by pointing the laser beam to nadir direction and compensate

- 115 any attitude changes of the aircraft by means of the double wedge scanner. As shown and discussed by Witschas et al. (2017), measurements of both horizontal and vertical wind profiles are very useful to characterize the spectral properties of mountain waves. In order to additionally gain knowledge of the momentum transport of mountain waves, horizontal wind speed (*u*) and vertical wind speed (*w*) need to be measured simultaneously. The leg-averaged momentum flux ( $\propto \overline{u'w'}$ ) can then be calculated like in Smith et al. (2016). Here, ' denotes perturbations of the respective quantity. For this purpose, the 2 µm DWL has been
- 120 operated with a modified scan pattern during the GW-LCYCLE II campaign for the first time. In particular, the laser beam has alternately been steered forth and back with an off-nadir angle of  $\pm 20^{\circ}$ . With that and the knowledge of the laser beam





pointing direction, u and w can be derived from a successive pair of line-of-sight (LOS) measurements. It is worth mentioning that u denotes the horizontal wind along flight direction here, which coincided well with the wind direction for the discussed flight (see section 3.1.1).

For the applied scan pattern, each LOS measurement took 2 s and the aircraft speed above ground was approximately  $200 \text{ m s}^{-1}$ . Thus, the horizontal resolution of the measured horizontal and vertical wind is  $\approx 800 \text{ m}$ . A more detailed explanation of the momentum-flux scan pattern of the 2 µm DWL will be presented in Witschas et al. (in preparation).

# 2.2.3 In-situ measurements

Horizontal and vertical velocity data at flight level are provided by the DLR facility Flight Experiments. For the DLR Falcon,
the velocity field is determined from data taken by a Rosemount model 858 flow angle sensor and a Honeywell Lasernav YG 1779 inertial reference system (IRS) (Bögel and Baumann, 1991). Measurements on HALO are conducted by the Basic HALO Measurement and Sensor System (BAHAMAS). Recent method and calibration details can be found in Mallaun et al. (2015) and Giez et al. (2017). For the horizontal wind the measurement uncertainties are smaller than 0.5 m s<sup>-1</sup> for HALO and 0.9 m s<sup>-1</sup> for Falcon, and smaller than 0.3 m s<sup>-1</sup> for the vertical wind (Heller et al., 2017; Bramberger et al., 2018). The data are used at a time resolution of 1 s. Perturbation quantities of the velocity data (u', v', w') are calculated by de-trending the data with a linear least square fit and subtracting the mean over the leg (Portele et al., 2018). Wavelet spectra of vertical velocity and MF cospectra of ρu'w' (Woods and Smith, 2010) with modifications of Portele et al. (2018) are computed based

# 2.3 Idealized numerical simulations

on Torrence and Compo (1998)<sup>3</sup>.

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140 EULAG<sup>4</sup> is a multi-scale computational model for the simulation of geophysical flows. It provides at least second-order accuracy in time and space (Prusa et al., 2008). EULAG solves the governing equations of motion either in an EUlerian or a LAGrangian form. Here, the non-hydrostatic equations of motion

$$\frac{D\mathbf{v}}{Dt} = -\nabla \frac{p'}{\bar{\rho}} + \mathbf{g}\frac{\theta}{\bar{\theta}} - \mathbf{f} \times \mathbf{v}' + \mathcal{M}',\tag{1}$$

$$\frac{D\theta}{dt} = 0, \tag{2}$$

$$Dt$$
 (7)

$$\nabla \cdot (\bar{\rho} \mathbf{v}) = 0, \tag{3}$$

are used in their Boussinesq approximated ( $\bar{\rho} = \rho_0$ ) form for the first set of simulations and decrease of density with altitude is taken into account for the second set of simulations (Smolarkiewicz et al., 2001; Prusa et al., 2008).  $\frac{D}{Dt}$  is the material derivative, v is the velocity vector, p is pressure,  $\rho$  is density,  $\theta$  is potential temperature,  $\mathcal{M}$  represents appropriate metric forces, and f and g symbolize the vectors of Coriolis parameter and gravity acceleration. Primes denote deviations from the ambient state and overbars refer to the horizontally homogeneous hydrostatic reference state of the Boussinesq expansion

<sup>&</sup>lt;sup>3</sup>Wavelet software was provided by C. Torrence and G. Compo, and is available at URL: http://atoc.colorado.edu/research/wavelets/, last access Nov 2019 <sup>4</sup>http://www2.mmm.ucar.edu/eulag/, last access Nov 2019





around a constant stability profile (Smolarkiewicz et al., 2001). EULAG has been applied for a broad range of topics in fluid dynamics including orographic GWs (e.g., Prusa et al., 1996; Grubišić and Smolarkiewicz, 1997). Detailed model setup is given in Section 3.2.

## 3 Results

# 155 3.1 MW event over Southern Scandinavia

#### 3.1.1 Meteorological situation

IOP 6 was a transient MW event over southern Scandinavia on 28 January 2016. Two synoptic low pressure systems over the tip of Greenland and over the Baltic sea caused moderate south-westerly to westerly winds (10 m s<sup>-1</sup> to 20 m s<sup>-1</sup>) in the troposphere and the excitation of MWs at the Southern Scandinavian mountain range (Fig. 2a). At tropopause level (300 hPa)
winds were westerly and below 30 m s<sup>-1</sup> over Southern Scandinavia as the polar front jet was located over the British Isles and northern Germany. A secondary jet streak occurred over the Norwegian Sea between Iceland and the Norwegian coast (Fig. 2b). The vertical cross section of horizontal wind speed interpolated in time and space along flight leg RF08 FL2 shows increasing wind speed up to 80 m s<sup>-1</sup> above 20-km altitude in the stratosphere (Fig 3a). The cross section of vertical wind shows vertically propagating MWs in the troposphere and increasing wave amplitudes in the stratosphere (Fig 3b). The resolved MWs in ECWMF are associated with the main mountain peaks of the IFS model topography.

165 MWs in ECWMF are associated with the main mountain peaks of the IFS model topography.
In Fig. 4, a time-height section of ECMWF horizontal wind speed located at the mountain ridge at point X<sub>0</sub> (Fig. 1a) is plotted. MWs were generated by moderate wind speeds in the lower troposphere on 28 January 2016. However, the MWs were prevented from propagating into the stratosphere until about 8 UTC due to weak winds close to 0 m s<sup>-1</sup> in the mid-troposphere. After 10 UTC, wind speeds above the tropopause and in the mid-troposphere increased which have allowed vertical propagation
170 of tropospheric GWs into the stratosphere. During the time of the research flights (red dots in Fig. 4), wind speeds below the

170 of tropospheric GWs into the stratosphere. During tropopause weakened again  $(10 \text{ m s}^{-1}\text{-}15 \text{ m s}^{-1})$ .

Vertical profiles of horizontal wind speed and Brunt-Vaisala frequency from an operational sounding<sup>5</sup> started from Stavanger (Fig. 1) at 12 UTC on 28 January 2016 are shown in Fig. 5a, b. This figure illustrates the moderate winds in the troposphere, the pronounced jump in static stability at the tropopause, that is typical for a TIL (Birner, 2006), and the increasing winds

175 with height in the stratosphere within the polar vortex. The critical horizontal wavelength which separates evanescent and propagating GWs was mainly larger than 10 km in the troposphere (i.e. only waves with horizontal wavelength > 10 km can propagate), smaller than 10 km in the vicinity of the TIL, and increasing towards 20 km above in the stratosphere (Fig. 5c).

# 3.1.2 Airborne observations

The coordinated flights of HALO and DLR Falcon provided simultaneous measurements of GW induced perturbations below and in the TIL. Figure 6a shows vertical velocities observed at flight level on all 5 cross mountain flight legs. Amplitudes of

<sup>&</sup>lt;sup>5</sup>sounding data from http://weather.uwyo.edu/upperair/

and PGS11 HL4 at 120 km distance.





 $2 \text{ m s}^{-1}$  in the troposphere and up to  $4 \text{ m s}^{-1}$  in the stratosphere are visible on all legs. Tropospheric measurements (HL1, HL2) show longer horizontal wavelengths compared to the observations at tropopause altitudes (RF07 FL2, RF08 FL1), which means that GW properties change in the vicinity of the tropopause. Wavelet power spectra of the observed vertical winds were computed to analyse the change in horizontal wavelengths (Fig. 6b). Wavelengths in the troposphere were in the order of 10 km to 20 h = (RCS11 HL 1) and HL2) and the order of 10 km to 20 h = (RCS11 HL 1) and HL2) and the order of 10 km to 20 h = (RCS11 HL 1) and HL2) are shown in the troposphere were in the order of 10 km to 20 h = (RCS11 HL 1) and HL2) are shown in the troposphere were in the order of 10 km to 20 h = (RCS11 HL 1) and HL2) are shown in the troposphere were in the order of 10 km to 20 h = (RCS11 HL 1) and HL2) are shown in the troposphere were in the order of 10 km to 20 h = (RCS11 HL 1) and HL2) are shown in the troposphere were in the order of 10 km to 20 h = (RCS11 HL 1) and HL2) are shown in the troposphere were in the order of 10 km to 20 h = (RCS11 HL 1) and HL2) are shown in the troposphere were in the order of 10 km to 20 h = (RCS11 HL 1) and HL2) are shown in the troposphere were in the order of 10 km to 20 h = (RCS11 HL 1) and HL2) are shown in the troposphere were in the order of 10 km to 20 h = (RCS11 HL 1) and HL2) are shown in the troposphere were in the order of 10 km to 20 h = (RCS11 HL 1) and HL2) are shown in the troposphere were in the order of 10 km to 20 h = (RCS11 HL 1) and HL2) are shown in the troposphere were in the order of 10 km to 20 h = (RCS11 HL 1) and HL2) are shown in the troposphere were in the order of 10 km to 20 h = (RCS11 HL 1) and HL2) are shown in the troposphere were in the order of 10 km to 20 h = (RCS11 HL 1) and HL2) are shown in the troposphere were in the order of 10 km to 20 h = (RCS11 HL 1) and HL2) are shown in the troposphere were in the order of 10 km to 20 h = (RCS1

to 30 km (PGS11 HL1 and HL2), while wavelet analysis shows that shorter wavelengths of 5 km to 9 km are dominating for RF07 FL2 and RF08 FL1 in downstream region. Longer waves with wavelengths of ≥ 10 km are found for the uppermost flight leg in the lower stratosphere (PGS11 HL4). Note that this was the last flight leg and it took place about two hours later than the other flight legs. The revealed wave signatures are not directly related to the topography spectrum (Fig. 6b) which was computed from the ASTER topography data (Schmugge et al., 2003) along the flight track (shown in Fig. 6a). The wave signatures are influenced by the background conditions.

Wavelet cospectra of MF were computed to study the propagation characteristics of the waves in more detail (Fig. 7). Alternating positive and negative MF at wavelengths of 10 km to 30 km were observed at distances of -100 km to 0 km and 100 km to 300 km below and in the tropopause region on the flight legs which took place at nearly the same time (PGS11 HL1 and HL2, RF08 FL1). This alternating pattern is an indication for reflected and trapped waves. Significant MF at shorter
195 wavelengths is found in the tropopause region and strongest alternating positive and negative signals occur downstream of the main mountain peaks. No significant positive or negative MF is found for the short scales < 10 km above the tropopause at 13 km altitude (PGS11 HL4). These findings indicate that the short waves are trapped in the tropopause region. Upward propagating longer waves (negative MF) with horizontal wavelengths of approximately 40 km to 50 km are found for RF08 FL1</li>

200 The DLR Falcon DWL measured in nadir mode on the first cross-mountain flight leg RF07 FL2 (Fig. 8). Measured vertical winds show fine scale up- and downdrafts over the mountains. The horizontal wavelengths of the GWs are smaller downstream of  $X_0$  on the lee side of the mountain and the phase lines are vertical. This again indicates wave trapping.

Lidar and in-situ measurements of the coordinated flight legs RF08 FL1 and PGS11 HL1 are shown in Fig. 9. As the DWL of DLR Falcon operated in flux mode on this leg, both the horizontal wind component in flight direction and the vertical wind component were measured. Contour lines of lidar measurements are overlaid by in-situ wind measurements of both aircraft. Wind measurements are complemented by water vapour measurements of the upward looking DIAL of HALO. Horizontal wind speeds in Fig. 9a show large scale wave structures with upstream tilted phase lines in the troposphere. In-situ measurements around the tropopause indicate similar wave structures but with stronger wind speeds compared to tropospheric values. In addition, large-scale wave structures are superimposed by small-scale waves with vertical phase lines. These small scale-

210 waves are more clearly visible in vertical wind measurements, which are more sensitive to smaller scale waves (Lane et al., 2003; Smith and Kruse, 2017), and show a clear change from longer to smaller wavelengths below and in the vicinity of the tropopause (Fig. 8 and 9b). This behaviour was already revealed by the wavelet analysis of the in-situ measurements (Fig. 6b). The short horizontal wavelengths are also visible in observations of water vapour mixing ratio (Fig. 9c) and the lidar reflectivity (Fig. 9d) between 8 km and 10 km altitude.





- Additional information regarding the wave propagation comes with the direct measurements of GW induced momentum fluxes by the new momentum flux method of the Falcon DWL. This was done on the coordinated flight leg RF08 FL1. Figure 10 illustrates vertical profiles of leg-averaged momentum fluxes along the cross-mountain flight legs RF08 FL1, PGS11 HL1 and HL4 obtained from lidar and in-situ measurements. The leg is split into an upstream part (segment 1), a part above the main mountain range (segment 2), and a part downstream of the main mountain range (segment 3). The profile of segment 3 indicates upward propagating (negative momentum fluxes) below 8 km altitude and downward propagating (positive momentum fluxes)
- around the tropopause. This suggests that waves were reflected and trapped in the vicinity of the TIL. Leg averaged in-situ momentum fluxes show upward propagating waves above the 9.5-km altitude over the main mountain range (segment 2). The measured MF profile in the downstream region (Figs. 10) does not match the typical profile of a Kelvin-Helmholtz instability that is characterised by one peak of positive MF (Mahalov et al., 2011).

#### 225 **3.2** Idealized simulations of MWs and the TIL

In this section, it is investigated if interfacial waves on an inversion found in the lower troposphere can also occur at tropopause altitudes and which conditions are necessary for their occurrence. It is tested if this wave trapping was possible on 28 January 2016 over southern Scandinavia by performing additional 2-dimensional simulations with the Scandinavian topography and background profiles which approximate the prevailing conditions on 28 January 2016. There is no intention to tune the simulations as close to the measurements as possible because the main goal in this study is to identify processes which could are the distributed with the UTE S

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explain the observed wave structure in the UTLS.

The computational parameters are chosen similar to Vosper (2004) for the first set of simulations. The 2-dimensional domain consists of 1032 and 2000 grid points in x- and z-direction, respectively, with grid increments of  $\Delta x = 100$  m and  $\Delta z = 10$  m (terrain following). This results in a total domain size of about 103 km × 20 km. The integration time step  $\Delta t$  is set to 1 s. Open boundaries are applied in x-direction. The model top is a rigid lid. The sponge layers at the horizontal edges of the domain are 8 km wide and the sponge layer at the top of the domain starts at 15-km altitude. As in Vosper (2004) an idealized ridge

$$h(x) = \begin{cases} h_0 [1 + \cos(Kx)]/2 & \text{for } |x| \le \pi/K \\ 0 & \text{for } |x| > \pi/K, \end{cases}$$
(4)

where  $K = 2\pi/L$ , and free slip lower boundary condition are used. Mountain height  $h_0$  is set to 400 m and width L to 10 km or 5 km. The usage of this idealized ridge is considered to be sufficient to investigate the occurrence and changes in the horizontal scale of the GWs in the vicinity of the TIL because it is known that the horizontal wavelength of interfacial waves is

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independent of the mountain half-width and height (Sachsperger et al., 2017). Four simulations are performed with a vertically constant horizontal velocity  $U = 8 \text{ m s}^{-1}$  and different profiles of potential temperature with the corresponding Brunt-Väisälä frequency as initial conditions. Table 1 summarizes the relevant initial parameters of the different model runs.

Figure 11a, b (RUN 1 and 2) shows that the EULAG model can reproduce the findings of Vosper (2004): an inversion at the top of an neutral boundary layer leads downstream propagating waves at the altitude of the inversion (Fig. 11b). The horizontal wavelength is approximately 5 km and the largest amplitudes are found in the vicinity of the inversion (Fig. 12c).





The signal is weak below the inversion (Fig. 12d) and absent if no inversion is present (Fig. 12a, b). If the stability above the inversion is increased to  $N_U = 0.02 \text{ s}^{-1}$  (stratospheric stability), the strength of the inversion  $\Delta \theta$  must be twice as large to allow for wave trapping and horizontal propagation on the inversion (RUN 3). The horizontal wavelength of the interfacial waves decreases with increasing stability above the inversion. The horizontal wavelength is approximately 2.5 km for  $U = 8 \text{ m s}^{-1}$ 250 and  $N_U = 0.02 \text{ s}^{-1}$  (Fig. 12f). This is because for  $N_U = 0.02 \text{ s}^{-1}$  waves with a horizontal wavelength of 5 km are no longer evanescent above the inversion and can propagate vertically (no trapping). The vertical profile of MF associated with the interfacial waves on the boundary layer inversion downstream of the idealized mountain (Fig. 11b, Fig. 12c) is shown in Fig. 13. Negative and positive MF of same magnitude are found below and above the inversion, respectively. This MF profile differs from the MF profile of a Kelvin-Helmholtz instability for which MF is zero below and above the instability (Mahalov 255 et al., 2011).

Figures 11d and 12h (RUN 4) show that interfacial waves can also exist on the TIL located between a stably stratified troposphere  $(N_L = 0.01 \text{ s}^{-1})$  and the stratosphere  $(N_U = 0.02 \text{ s}^{-1})$ . The horizontal wavelength is again approximately 2.5 km for  $U = 8 \text{ m s}^{-1}$ . Besides the upward propagating MWs and the interfacial waves, reflected waves with a horizontal wavelength

- of approximately 5 km exist downstream of the mountain in the troposphere (Figs. 11d and 12j) although the classical trapping 260 condition of a decreasing Scorer parameter with altitude in the troposphere (Scorer, 1949) is not fulfilled. The amplitudes of the reflected waves are found to be larger if an inversion is present at the tropopause than if there is just the jump from tropospheric to stratospheric stability (not shown). The results of these numerical simulations confirm that interfacial waves can exist on the TIL. The mechanism of their trapping (evanescence in the layer above the inversion), the resulting horizontal wavelength, and the occurring horizontal propagation match the published results for boundary-layer inversions (Vosper, 2004; Sachsperger
- 265
- et al., 2015).

The second set of simulations (RUN 5 and 6) uses a 2-dimensional domain with 2016 and 1000 grid points in x- and z-direction, respectively, with grid increments of  $\Delta x = 500$  m and  $\Delta z = 40$  m. This results in a total domain size of about 1008 km  $\times$  40 km. The integration time step  $\Delta t$  is set to 1 s. Open boundaries are applied in x-direction. The model top is a 270 rigid lid. The sponge layers at the horizontal edges of the domain are 40 km wide and the sponge layer at the top of the domain starts at 25 km altitude. The Scandinavian topography is interpolated on the 500 m grid from ASTER data. The initial profiles approximate the background conditions over Southern Scandinavia given by the Stavanger radiosonde on 28 January 2016 (Fig. 14a, b, d, e). Simulations without and with a TIL are performed. The simplified initial horizontal velocity profile does not contain negative shear above the tropopause but negative shear establishes in the course of the simulation (black dashed profiles shown in Fig. 14a, d are located at -150 km distance (Fig. 15f) 16 hours after start of the simulations). 275

Figures 14f and 15h (RUN 6) show that interfacial waves can also exist for the background conditions found on 28 January 2016 over Southern Scandinavia. They are found downstream of the main mountain peak in the vicinity of the TIL (Fig. 14f and Fig. 15f) and their horizontal wavelength is approximately 8 km (Fig. 15h). The horizontal band pattern in the wavelet spectrum resembles the one of the idealized mountain simulations presented above (Fig. 12c, f, h). Interfacial small

scale waves are absent in the troposphere below the TIL (Fig. 15i) and in the case of no TIL (RUN 5, Fig. 14c and Fig. 15a, c). 280 Reflected waves with horizontal wavelengths between 10 km and 30 km exist downstream of the main mountain peaks in the





troposphere (Figs. 14c, f and 15d, i). It was already mentioned that the horizontal wavelength of the interfacial waves is independent of the generating terrain and determined by the background wind and stability. These 2-dimensional simulations reveal the expected wavelength of the GWs over Southern Scandinavia downstream of the main mountain ridge on 28 January 2016,
i.e. approximately 8 km in the vicinity of the TIL and between 10 km and 30 km in the troposphere. However, the interfacial waves in the simulation are not as dominant as in the measurements (Fig. 15h vs. RF07 FL2 and RF08 FL1 in Fig. 6b). There is a stronger signal of the upward propagating MWs above the main mountain peaks (Fig. 15f, h).

## 4 Discussion

The atmospheric conditions during the MW case were characterized by moderate low level flow (~ 10 m s<sup>-1</sup>), comparatively
weak wind speed (~ 30 m s<sup>-1</sup>) around the TIL and increasing wind speed above (Sec. 3.1.1). The coordinated airborne measurements including the downward pointing Doppler wind lidar measurements revealed that the vertical velocity field was dominated by small horizontal scales with a decrease from around 20 km to < 10 km in the vicinity of the TIL. These small scales were also found in the water vapour and reflectivity data (Sec. 3.1.2). The corresponding MF indicates wave reflection and trapping at the TIL (Fig. 7 and Fig. 10). It is known that atmospheric inversions can be wave guides leading to wave</li>
trapping and downstream wave propagation (Vosper, 2004; Sachsperger et al., 2015; Chouza et al., 2015; Fritts et al., 2018) but observations of downstream wave propagation of small scale waves at tropopause inversions are rare (Smith et al., 2008; Woods and Smith, 2010).

In the course of investigating MW propagation in the UTLS with horizontal- and altitude-resolved airborne observations (Section 3.1.2), we found that the measured and simulated MF profiles downstream of the main mountain ridge (Figs. 10 and 13) do not match the typical profile of a Kelvin-Helmholtz instability (Mahalov et al., 2011). The shear was not strong enough to generate Kelvin-Helmholtz instabilities (Fig. 9a and Fig. 14a). Instead, the MF profile is characterized by negative fluxes below and positive fluxes above the TIL which show similar magnitudes (Fig. 10). Local values of the MF reach up to 0.034 kN m<sup>-1</sup> in magnitude which is slightly below the values found for an MW event over New Zealand during DEEPWAVE which where around 0.05 and 0.07 kN m<sup>-1</sup> in magnitude (Portele et al., 2018). The averaged MF in the stratosphere above the main mountain range (-0.16 Pa) is in the range of to the moderate MW cases of DEEPWAVE [Fig. 5b in (Smith et al., 2016)].

The observed horizontal wavelengths in the vicinity of the TIL were clearly evanescent in the stratosphere (Fig. 5c). This excludes their direct propagation from above followed by their trapping on the TIL similar to Woods and Smith (2011). The presence of interfacial waves that are trapped on inversions and that are likely generated by MWs coming across the inversions

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the propagation of the interfacial waves (Vosper, 2004; Sachsperger et al., 2015, 2017). However, the amplitudes depend on the energy source, which is better described by hydraulic theory than by traditional linear models (Sachsperger et al., 2017). The traditional linear models link the energy source to the topography which is inaccurate for interfacial waves (especially for large amplitudes) because of nonlinear effects. Sachsperger et al. (2017) included the nonlinear effects in their model for the amplitudes of interfacial waves on the boundary-layer inversion by assuming that the interfacial waves originate at the density

(Sachsperger et al., 2017) has not yet been observed at TILs. Linear theory is able to describe the horizontal wavelength and





315 interface further aloft in the interior of the fluid and the energy source for these non-hydrostatic lee wave train is the energy convergence at an internal jump between two fluid layers of different densities.

The determined wave properties (observations) match those of interfacial waves (simulations) for which stability and wind conditions above the inversion determine the horizontal scales of the waves (Vosper, 2004; Sachsperger et al., 2015). The performed simulations show that the presence of the TIL is crucial in producing the trapped waves at tropopause altitude and

- 320 vertical wind shear by the main mountain wave was not sufficient in this case (Fig. 14a). However, the amplitudes of the interfacial waves were underestimated compared to the upward propagating mountain waves in the simulations. This can have several reasons. The amplitudes of interfacial waves depend on the amount of energy provided by the main wave source at the interface and the acting nonlinear processes (Sachsperger et al., 2017). It was not yet investigated how the interaction and generation processes depend on the model resolution and if the amplitudes increase with increasing resolution. The simulations were only
- 325 2-dimensional so they cannot capture effects of the fully 3-dimensional mountain range. Moreover, potential additional energy input by downward propagating larger scale waves from stratospheric sources (e.g., polar night jet (Dörnbrack et al., 2018)) are not included in the simulations. The evaluation of these effects and the assessment of their sensitivities require additional extensive model simulations which are beyond the scope of this paper and can be addressed in a future study.

#### 5 Conclusions

330 The unique combination of observations from coordinated airborne in-situ and lidar measurements and idealized large-eddy simulations revealed the occurrence of interfacial waves on the tropopause inversion during an MW event over Southern Scandinavia on 28 January 2016. Such interfacial waves have already been observed on boundary-layer inversions but their concept has not been applied to tropopause inversions so far.

Strong shear induced by the main mountain wave can cause Kelvin-Helmholtz instabilities which results in similar patterns
in the vertical velocity field (Mahalov et al., 2011). Although the horizontal scales are around ten kilometers, which is similar to T-REX observations (Smith et al., 2008; Woods and Smith, 2010), neither Kelvin-Helmholtz instability nor downward propagation of secondary GWs generated by MW breaking in the middle stratosphere can explain our observations. The vertical shear was not as pronounced as during the T-REX case (Mahalov et al., 2011) because the tropopause jet was not well established over Southern Scandinavia on 28 January 2016. The wind speed influenced by the large scale mountain wave was
only between 10 m s<sup>-1</sup> and 40 m s<sup>-1</sup>. The stratospheric critical horizontal wavelengths calculated from co-located radiosonde measurements are larger than the observed scales in the UTLS region which would hinder their direct downward propagation from a breaking region located higher up.

Our idealized simulations revealed that interfacial waves can occur also on tropopause inversions similar to boundary layer inversions. Our analyses of the horizontal- and altitude-resolved airborne observations revealed that they actually do occur. As predicted by linear theory, the horizontal scale of those waves is determined by the wind and stability conditions above the inversion. They are found downstream of the main mountain peaks and their characteristic MF profile (negative fluxes below and positive fluxes above the inversion) clearly distinguishes from the MF profile of a Kelvin-Helmholtz instability.





Data availability. HALO and Falcon data is available on HALO-Database (https://halo-db.pa.op.dlr.de/mission/3). ECMWF data is available in the MARS archive (https://apps.ecmwf.int/archive-catalogue/). EULAG simulation data is available on the
 ftp-server (ftp://ftp.pa.op.dlr.de/pub/gisinger/ACP\_EULAG\_2020/).

*Author contributions.* S. Gisinger performed the large-eddy simulations and prepared the manuscript with contributions from all co-authors. J. Wagner did the analyses of the meteorological conditions and the aircraft insitu data. B. Witschas did the analyses of the lidar data.

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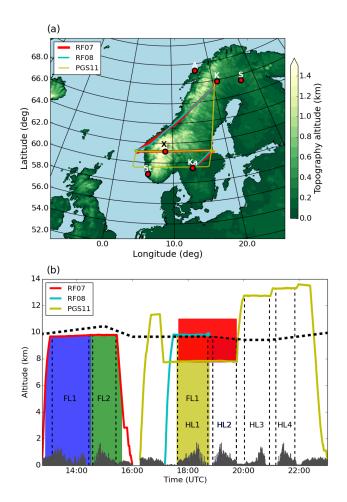
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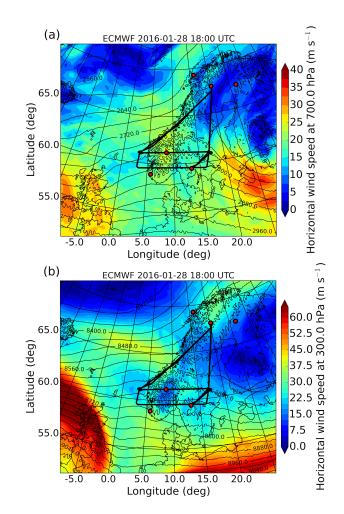




**Figure 1.** (a) Topographic map of Scandinavia and area of operation of IOP 6 during the GW-LCYCLE II campaign. The coloured lines indicate Falcon (RF07 and RF08) and HALO (PGS11) flight tracks. The red dots mark the position of Andenes (A), Kiruna (K), Sodankylä (S), Karlstad (Ka) and Stavanger (St). The location of the highest mountain peak on the cross-mountain flight legs is marked with  $X_0$ . Flight altitudes of Falcon and HALO are shown in (b). Falcon flight legs RF07 FL2 and RF08 FL1 and HALO flight legs PGS11 HL1, HL2 and HL4 are cross-mountain flights through  $X_0$ , which are analyzed in this study. Colour shaded areas mark regions covered by the upward looking HALO water vapour lidar WALES (red) and the downward looking Falcon Doppler wind lidar in scanning (blue), nadir (green) and flux mode (yellow). The temporal evolution of the ECMWF thermal tropopause height at point  $X_0$  is indicated with the thick dashed line.



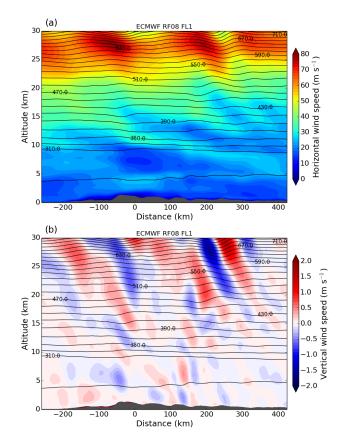




**Figure 2.** Meteorological maps of horizontal wind speed and geopotential height (black contour lines) at (a) 700 hPa and (b) 300 hPa at 18 UTC on 28 January 2016 obtained from the ECMWF model. Black lines indicate flight legs of the three research flights and red dots mark the same locations as in Figure 1.



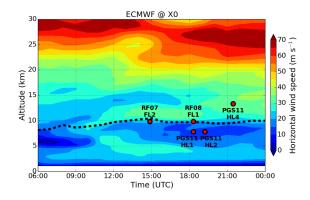




**Figure 3.** ECMWF vertical cross sections of (a) horizontal wind speed and (b) vertical wind speed interpolated in time and horizontal space along flight leg RF08 FL1. Black contour lines indicate potential temperature with an interval of 20 K. The cross section distance is centered at  $X_0$  (see Fig. 1).



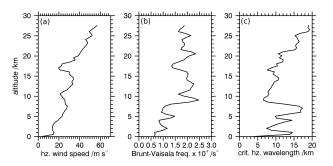




**Figure 4.** ECMWF time-height section of horizontal wind speed at point  $X_0$ . The black dashed line marks the height of the thermal tropopause. Red dots mark the altitudes of HALO and Falcon at point  $X_0$  of the respective flight legs (see also Fig. 1).



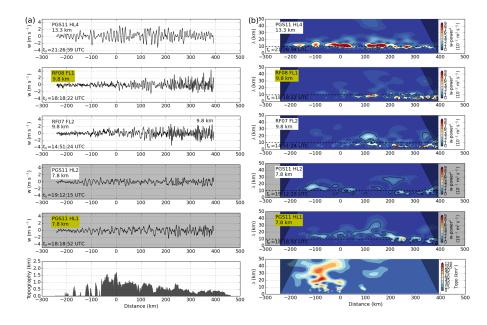




**Figure 5.** Vertical profiles of (a) horizontal wind speed, (b) Brunt-Vaisala frequency, and (c) critical horizontal wavelength of the radiosonde launched at Stavanger (St) southern Norway (Fig. 1) at 12 UTC on 28 January.







**Figure 6.** Cross mountain flight legs of Falcon and HALO for (a) in-situ vertical wind and topography and (b) corresponding wavelets of vertical wind. Black contour lines mark regions, which are significant on the 5%-level. The cone of influence is shaded in grey and flight legs, which were located below the tropopause (see labelled mean flight altitudes) are marked with grey background colour. Time t<sub>0</sub> indicates when the aircraft was located at  $X_0$  (see Fig. 1) and shows that PGS11 HL1 and RF08 FL1 (labelled with yellow boxes) took place nearly at the same time (HALO was flying 30 seconds behind Falcon).





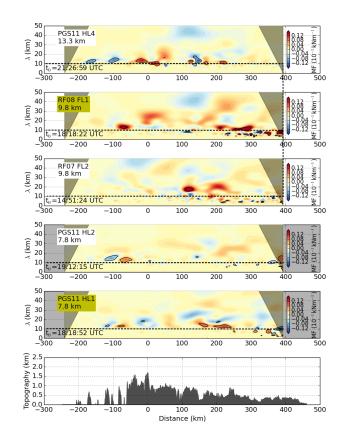


Figure 7. As in Fig. 6b but showing wavelet cospectra of MF ( $\rho u'w'$ ).





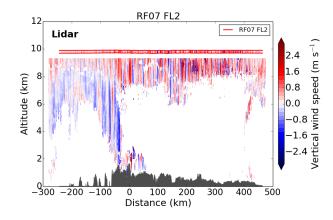
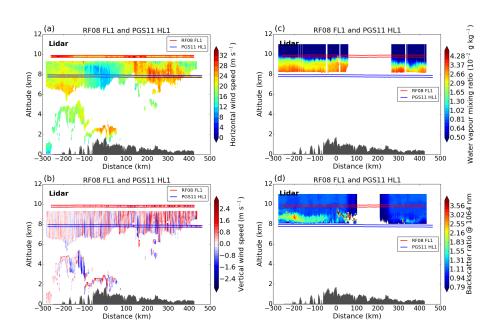


Figure 8. Vertical winds along flight leg RF07 FL2 measured by the DWL and in-situ instruments (marked by red horizontal lines) at flight level by the DLR Falcon.



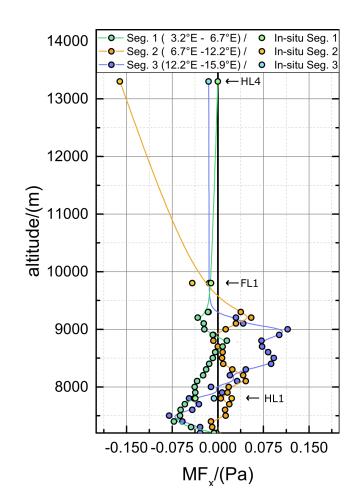




**Figure 9.** DWL measurements of (a) horizontal wind speed and (b) vertical wind speed and WALES measurements of (c) water vapour mixing ratio and (d) lidar reflectivity along flight leg RF08 FL1/PGS11 HL1 combined with corresponding in-situ measurements of HALO and DLR Falcon at flight level (marked by blue and red horizontal lines).







**Figure 10.** Leg-averaged momentum flux profiles along flight leg RF08 FL1 obtained from DWL and in-situ measurements which include also PGS11 HL1 and HL4. Segment 1 is the upstream part of the leg, segment 2 the part above the main mountain range, and segment 3 the downstream part of the leg. Bold black line separates positive and negative fluxes.





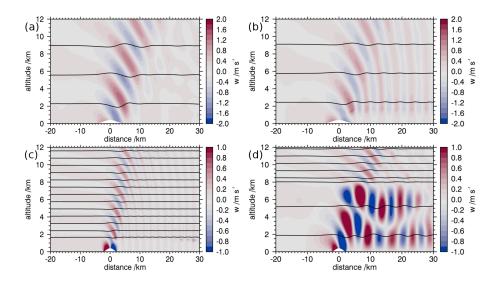


Figure 11. Vertical velocity of the idealized simulations of the cases with (a) a neutral boundary layer without inversion and  $N_U = 0.01 \text{ s}^{-1}$  (RUN 1), (b) a neutral boundary layer with an inversion of 3.3 K and  $N_U = 0.01 \text{ s}^{-1}$  (RUN 2), (c) a neutral boundary layer with an inversion of 6.6 K and  $N_U = 0.02 \text{ s}^{-1}$  (RUN 3), and (d) a stable troposphere ( $N = 0.01 \text{ s}^{-1}$ ) with a TIL of 6.6 K and  $N_U = 0.02 \text{ s}^{-1}$  (RUN 4).





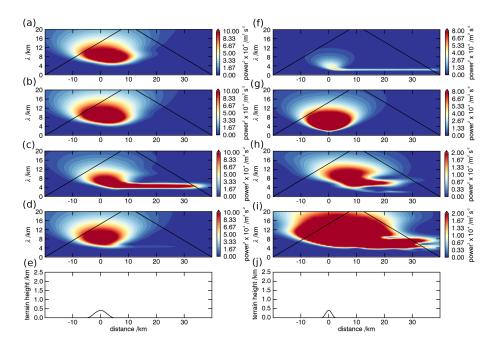


Figure 12. Wavelet spectra of the vertical velocity of the idealized simulations shown in Fig. 11 in the vicinity of (a, c, f, h) and below (b, d, g, i) the top of the boundary layer or the inversion layer. e and j show the idealized terrain. (a, b) is for a neutral boundary layer without inversion and  $N_U = 0.01 \text{ s}^{-1}$  (RUN 1), (c, d) for a neutral boundary layer with an inversion of 3.3 K and  $N_U = 0.01 \text{ s}^{-1}$  (RUN 2), (f, g) for a neutral boundary layer with an inversion of 6.6 K and  $N_U = 0.02 \text{ s}^{-1}$  (RUN 3), and (h, i) for a stable troposphere ( $N = 0.01 \text{ s}^{-1}$ ) with a TIL of 6.6 K and  $N_U = 0.02 \text{ s}^{-1}$  (RUN 4).





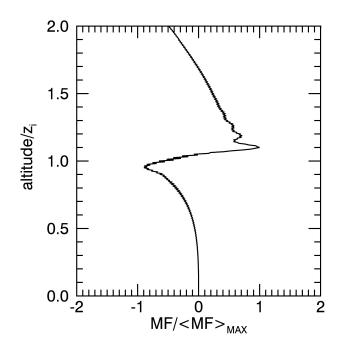
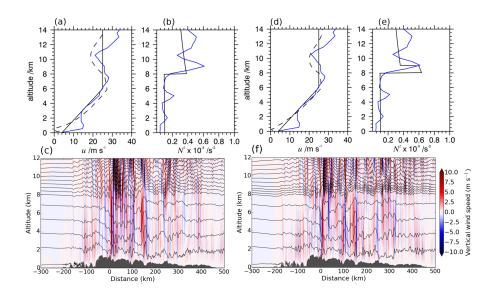


Figure 13. Averaged momentum flux of the interfacial waves on the boundary layer inversion (RUN 2: 3.3 K and  $N_U = 0.01 \text{ s}^{-1}$ , Fig. 11b, Fig. 12c) downstream of the idealized mountain (10-30 km horizontal distance). Values are normalized by the maximum flux in the vicinity of the inversion and given as a function of normalized altitude ( $z_i$  is the altitude of the inversion).



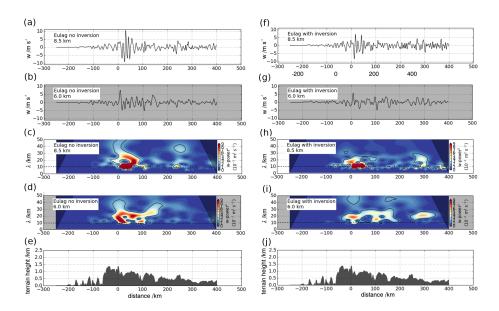




**Figure 14.** Initial profiles (black solid) and vertical velocity for the simulations with more realistic terrain without TIL (a-c, RUN 5) and with TIL (d-f, RUN 6). The initial profiles approximate the background conditions over southern Scandinavia on 28 January 2016 (blue profiles show the Stavanger radiosonde data). Negative shear above the tropopause establishes in the course of the simulations (a, d; black dashed, time = 16 h, distance = -150 km).







**Figure 15.** Vertical velocity and corresponding wavelet spectra of the simulations shown in Fig. 14 [left: without TIL (RUN 5), right: with TIL (RUN 6)]. a, c, f, h are at the altitude of the TIL and b, d, g, i below the tropopause. e and j show the terrain in the domain.





**Table 1.** Mountain width L, static stabilities of the lower and upper layer  $(N_L, N_U)$ , strength of the inversion  $(\Delta \theta)$  and inversion height  $(z_i)$ , upstream wind conditions, and background density of the EULAG simulations.

RUN	<i>L /</i> km	$N_L  / \mathrm{s}^{-1}$	$N_U / \mathrm{s}^{-1}$	$\Delta \theta$ /K	$ z_i$ /m	$U /\mathrm{ms^{-1}}$	$\bar{\rho} = const.$
1	10	0.00	0.01	0.0	400	8	yes
2	10	0.00	0.01	3.3	1600	8	yes
3	5	0.00	0.02	6.6	1600	8	yes
4	5	0.01	0.02	6.6	8000	8	yes
5	ASTER topo	0.01	0.02	0.0	8000	$4 \rightarrow 25$	no
6	ASTER topo	0.01	0.02	20	8000	$4 \rightarrow 25$	no