



1 **Estimating Radiative Forcing Efficiency of Dust Aerosol** 2 **Based on Direct Satellite Observations: Case Studies over** 3 **the Sahara Desert and Taklimakan Desert**

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11 **Abstract.** The direct radiative forcing efficiency of the dust aerosol ($DRFE_{dust}$) is an important indicator
12 to measure the climate effect of the dust. The $DRFE_{dust}$ is determined by the microphysical properties
13 of the dust, which vary with the dust source regions. However, there are only sparse in-situ
14 measurements of them, such as the distribution of the dust aerosol particle size and the complex
15 refractive index in the main dust source regions. Furthermore, recent studies have shown that the
16 non-spherical effect of the dust particle is not negligible. The $DRFE_{dust}$ is often evaluated by estimating
17 given microphysical properties of the dust aerosols in the radiative transfer model (RTM). However,
18 considerable uncertainties exist due to the complex and variable dust properties, including the complex
19 refractive index and the shape of the dust. The $DRFE_{dust}$ over the Taklimakan Desert and the Sahara
20 Desert is derived from the satellite observations in this paper. The advantage of the proposed
21 satellite-based method is that there is no need to consider the microphysical properties of the dust
22 aerosols in estimating the $DRFE_{dust}$. For comparison, the observed $DRFE_{dust}$ is compared with that
23 simulated by the RTM. The differences in the dust microphysical properties in these two regions and
24 their impacts on $DRFE_{dust}$ are analyzed.

25 The $DRFE_{dust}$ derived from the satellite observation is $-39.6 \pm 10.0 \text{ Wm}^{-2}\tau^{-1}$ in March 2019 over
26 Tamanrasset and $-48.6 \pm 13.7 \text{ Wm}^{-2}\tau^{-1}$ in April 2019 over Kashi. According to the analyses of their
27 microphysical properties and optical properties, the dust aerosols from the Taklimakan desert (Kashi)
28 scatter strongly. The RTM simulated results (-41.5 to $-47.4 \text{ Wm}^{-2}\tau^{-1}$ in the Taklimakan Desert and
29 -32.2 to $-44.3 \text{ Wm}^{-2}\tau^{-1}$ in the Sahara Desert) are in good agreement with the results estimated by
30 satellite observations. According to previous studies, the results in this paper are proved to be



31 reasonable and reliable. The results also show that the microphysical properties of the dust can
32 significantly influence the $DRFE_{dust}$. The satellite-derived results can represent the influence of the dust
33 microphysical properties on the $DRFE_{dust}$, which can also validate the direct radiative effect of the dust
34 aerosol and the $DRFE_{dust}$ derived from numerical model more directly.

35 1 Introduction

36 Dust aerosols are considered to be one of the major components of the tropospheric aerosols
37 (Huneeus et al., 2012;Textor et al., 2007). The dust aerosols affect the radiation balance of the
38 earth-atmosphere system by scattering and absorbing solar radiation directly (Miller et al.,
39 2014;Satheesh, 2002). Estimating the direct radiation effect of the dust aerosol (DRE_{dust}) is crucial for
40 estimating climate forcing. The scattering of the dust influences the radiation in the shortwave (SW)
41 spectrum at the top of atmosphere (TOA), which causes stronger SW DRE_{dust} over dust source regions
42 (Slingo et al., 2006). Therefore, the evaluation of SW DRE_{dust} is important for climate modeling.

43 The variabilities of the mineral dust composition from soils in different source regions cause the
44 differences in dust microphysical properties (e.g., refractive index, size, and particle shapes). Anderson
45 et al. (2005) defined the Direct Radiative Forcing Efficiency of the dust aerosol ($DRFE_{dust}$) to quantify
46 the dust radiative effect (Anderson et al., 2005). The $DRFE_{dust}$ represents the DRE_{dust} of a certain
47 aerosol optical depth (AOD) at per unit area, which means the efficiency of the dust aerosol that affects
48 the net radiative flux of solar radiation. The $DRFE_{dust}$ is largely determined by the optical properties of
49 the dust aerosols (Shi et al., 2005), which are strictly controlled by the microphysical properties of the
50 particles (Di Biagio et al., 2014b;Di Biagio et al., 2017;Di Biagio et al., 2014a;Zhang et al., 2006).
51 Therefore, the $DRFE_{dust}$ is different concerning the dust aerosols from different source regions (Tanré
52 et al., 2001;Che et al., 2012). Without considering the influence of the aerosol loading on the DRE_{dust} ,
53 the $DRFE_{dust}$ has unique advantages in evaluating the differences of dust microphysical properties and
54 their impacts on the DRE_{dust} from different dust source regions (García et al., 2008).

55 The $DRFE_{dust}$ is often estimated by the General Circulation Model (GCM) and the Radiative
56 Transfer Model (RTM). Many studies have simulated the SW $DRFE_{dust}$ in different regions
57 (Valenzuela et al., 2012;Che et al., 2009;Bi et al., 2014). However, there are sparse in-situ
58 measurements of the dust microphysical properties in the main source regions. The large spatial



59 variability of aerosols and the lack of an adequate database on their properties makes DRE_{dust} and
60 $DRFE_{dust}$ much difficult to estimated (Satheesh and Srinivasan, 2006). To date, climate models
61 generally use temporal and spatial constant values to represent the dust microphysical properties (Di
62 Biagio et al., 2017; Di Biagio et al., 2014a; Bi et al., 2020). This may cause uncertainties in calculating
63 the dust radiative effect. Moreover, the shape of the dust particle in the model needs to be assumed.
64 Therefore, there are large uncertainties in estimating the $DRFE_{dust}$ with few measurements of the dust
65 microphysical properties from different source regions (Bi et al., 2020; Colarco et al., 2014; Zhao et al.,
66 2013).

67 Satellite observations can be used in estimating the $DRFE_{dust}$ because satellites can directly
68 observe the radiation budget of the earth in the TOA (Wielicki et al., 1998; Satheesh and Ramanathan,
69 2000), and the remote-sensing technique for the AOD has been developed (Remer et al., 2005; Hsu et
70 al., 2004). In the previous study, we developed a satellite-based method to estimate the $DRFE_{dust}$ over
71 land without any assumptions of the microphysical properties of dust aerosols (Tian et al., 2019). In
72 previous researches, performances of the models in simulating the dust radiative effect have been
73 indirectly validated by comparing the observations of the AOD, the single scattering albedo (SSA), the
74 distribution of the particle size, and the extinction profile of the aerosols with the simulated ones (Zhao et
75 al. 2010; Chen et al. 2014). Therefore, the satellite-based method provides a direct way to validate the
76 DRE_{dust} and the $DRFE_{dust}$.

77 The Taklimakan Desert and the Sahara Desert are the main dust source regions, which influence
78 many areas (Li et al., 2020; Mikami et al., 2006; Mbourou et al., 1997; Huang et al., 2014). Thus, the
79 assessment of the SW $DRFE_{dust}$ and microphysical properties of the dust over these regions is
80 important for evaluating regional and global climate changes.

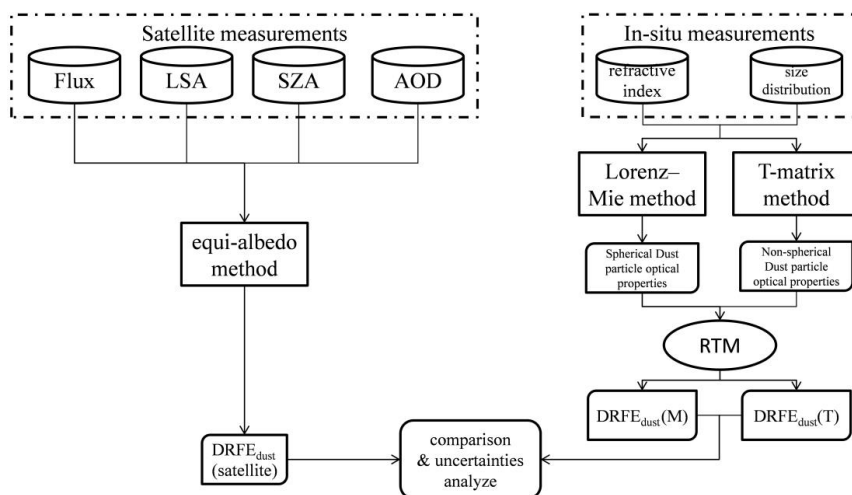
81 In this paper, the $DRFE_{dust}$ in dust storms over the Taklimakan Desert and the Sahara Desert is
82 evaluated based on satellite observations and the RTM, separately. With the comparison of the dust
83 microphysical properties and the $DRFE_{dust}$ in these two regions, the differences of the dust
84 microphysical properties are analyzed. Meanwhile, the influences of the dust microphysical properties
85 on the $DRFE_{dust}$ are investigated in this paper. The need for accurate information on the dust
86 microphysical properties and dust sources for simulating the $DRFE_{dust}$ is emphasized, and the
87 advantage of the satellite-based method in estimating the $DRFE_{dust}$ is revealed.



88 2 Methodology and data

89 In the previous study (Tian et al., 2019), the equi-albedo method has been proposed to estimate the
 90 DRE_{dust} and the $DRFE_{dust}$ over land based on satellite measurements directly. This method bases on the
 91 assumption that the SW radiative fluxes at the TOA of the clear sky (F_{clr}) are equal over the regions
 92 with similar land surface albedo (LSA) and solar zenith angle (SZA). Following this method, we
 93 estimated the $DRFE_{dust}$ based on the AOD and the SW radiative flux product from the same satellite
 94 platform.

95 Moreover, the $DRFE_{dust}$ in the RTM with dust aerosol microphysical properties is also evaluated.
 96 Based on the comparison between the $DRFE_{dust}$ results from the two methods, the differences in the dust
 97 microphysical properties over the Taklimakan Desert and the Sahara Desert are analyzed, and the
 98 differences in the $DRFE_{dust}$ are also discussed. The processing steps are shown in Fig. 1.



Flux: Radiative Flux observed by CERES;

LSA: Land Surface Albedo;

SZA: Solar zenith Angle;

AOD: Aerosol Optical Depth;

$DRFE_{dust}(\text{satellite})$: $DRFE_{dust}$ estimated from satellite measurements

$DRFE_{dust}(M)$: $DRFE_{dust}$ simulated from Lorenz-Mie method and RTM;

$DRFE_{dust}(T)$: $DRFE_{dust}$ simulated from T-matrix method and RTM;

99

100 **Figure 1: Processing flow chart of this paper.**



101 2.1 Methodology

102 2.1.1 The equi-albedo method

103 Previous studies have shown that F_{clr} is significantly influenced by the LSA and the SZA at the
104 TOA (Di Biagio et al., 2012; Tegen et al., 2010). It is hard to assess the SW DRE_{dust} and the $\text{DRFE}_{\text{dust}}$
105 over land derived from satellite observations due to the large dynamic range of the LSA (Satheesh,
106 2002). In the previous study (Tian et al., 2019), we proposed an equi-albedo method to minimize the
107 influence of the inhomogeneous LSA and SZA and directly derived the DRE_{dust} and the $\text{DRFE}_{\text{dust}}$ over
108 land from satellite observations based on the assumption that the F_{clr} is equal over the regions with
109 similar LSA and SZA.

110 DRE_{dust} was defined as the radiative fluxes difference between clear (F_{clr}) and dust loading (F_{dust})
111 conditions (Garrett and Zhao, 2006; Christopher et al., 2000; Ramanathan et al., 1989).

$$112 \quad \text{DRE}_{\text{dust}} = F_{\text{clr}} - F_{\text{dust}} \quad (1)$$

113 F_{dust} is the shortwave radiative flux at TOA in the cloud-free and dust aerosol loading condition
114 which is obtained directly from CERES data, and F_{clr} is the shortwave flux over the same region
115 without aerosol. F_{clr} cannot be observed directly, and the estimating of F_{clr} must be on the basis of
116 some realistic assumptions.

117 The equi-albedo method bases on the assumption that the SW radiative fluxes at the TOA of the
118 clear sky (F_{clr}) are equal over the regions with similar land surface albedo (LSA) and solar zenith angle
119 (SZA). Based on the assumption, the F_{clr} were estimated, then DRE_{dust} can be derived following Eq.
120 (1). According to the definition of $\text{DRFE}_{\text{dust}}$, it represents the net flux of solar radiation perturbed by
121 per unit dust AOD. Therefore, $\text{DRFE}_{\text{dust}}$ can be expressed as:

$$122 \quad \text{DRFE}_{\text{dust}} = \text{DRE}_{\text{dust}} / \tau_{\text{dust}} \quad (2)$$

123 where τ_{dust} is the AOD of dust aerosols, and τ_{dust} comes from the MODIS aerosol product.
124 Thus, $\text{DRFE}_{\text{dust}}$ was estimated based on the AOD and the SW radiative flux product from the same
125 satellite platform.

126 In the previous study (Tian et al., 2019), we have estimated the DRE_{dust} and the $\text{DRFE}_{\text{dust}}$ of two
127 dust storms in the Taklimakan Desert. The results were compared with the DRE_{dust} and the $\text{DRFE}_{\text{dust}}$
128 simulated by the RTM. The results indicated that the method is effective in estimating the SW $\text{DRFE}_{\text{dust}}$
129 over land. The microphysical properties of dust aerosols significantly influence on the DRE_{dust} and the
130 $\text{DRFE}_{\text{dust}}$ (Che et al., 2012; Li et al., 2018). The different microphysical properties of dust aerosols in



131 various dust source regions cause uncertainties in estimating the SW DRE_{dust} and $DRFE_{dust}$. Thus, the
132 equi-albedo method is used to estimate the SW $DRFE_{dust}$ directly using satellite observations in this
133 study. Based on the comparison of the $DRFE_{dust}$ in the Taklimakan Desert and the Sahara Desert, the
134 differences of dust microphysical properties in these two regions are analyzed and the influences of the
135 dust microphysical properties on estimating the $DRFE_{dust}$ are tested.

136 **2.1.2 Calculating method of dust optical properties**

137 Dust aerosols are often assumed as spherical particles in the GCM and the RTM (Wang et al.,
138 2013; Gao and Anderson, 2001). The Lorenz-Mie theory is used to calculate the optical properties of the
139 dust particles (Gouesbet and Gréhan, 2011). However, observations and researches have shown that
140 most dust aerosols are non-spherical in nature (Nakajima et al., 1989; Okada et al., 2001). Previous
141 researches also suggested that assuming particles as spherical or non-spherical has significant impacts
142 on calculating the dust optical properties (Kalashnikova and Sokolik, 2004; Borghese et al., 2007).
143 Therefore, the optical properties of dust aerosols are calculated using both the spherical and the
144 ellipsoidal methods for comparison to analyze the uncertainties caused by the assumption of dust
145 shapes in estimating the $DRFE_{dust}$ in this study.

146 To make it more accurate, the light scattering properties of spherical particles are generally
147 calculated based on the Mie and Lorenz theory (Mishchenko and Travis, 2008). Among several
148 methods for computing optical properties of non-spherical particles, the T-matrix method has been
149 extensively developed to many versions for various applications (Chylek et al., 1977; Mishchenko et al.,
150 1996). These versions of the available T-matrix code are accessed from the National Aeronautics and
151 Space Administration (NASA) Goddard Institute for Space Studies (GISS) group (Mishchenko and
152 Travis, 1998). Mie scattering method can be regarded as a special case of the T-matrix method. In this
153 study, the NASA-GISS code is used to calculate the optical properties of the spherical particles and the
154 ellipsoidal particles. The particle aspect ratio is set to 0.8.

155 **2.1.3 RTM**

156 Santa Barbara Disort Atmospheric Radiative Transfer (SBDART) is an RTM that calculates the
157 plane-parallel radiative transfer of the earth-atmosphere system (Ricchiazzi et al., 1998). The broadband
158 radiative flux at the TOA and the surface in clear-sky and dusty conditions can be obtained. It is

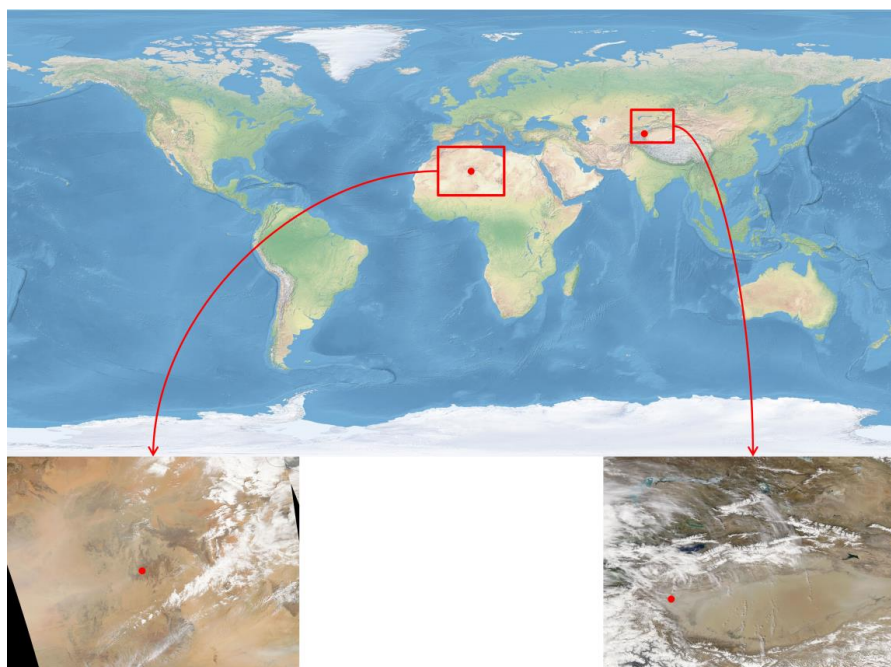


159 conducive to analyzing the radiative transfer theory in satellite remote sensing and atmospheric energy
160 budget studies. Furthermore, the model can flexibly set up aerosol properties, which is well suited to
161 calculate the radiative effect of different types of aerosols. The SBDART model has been widely used in
162 estimating the $DRFE_{dust}$ due to its design (Chen et al., 2011; Li et al., 2020; Iftikhar et al., 2018).

163 In this paper, the dust aerosol optical properties (the SSA and the ASYmmetry parameter,
164 abbreviated as ASY) are calculated using spherical and non-spherical methods. The Aerosol Robotic
165 Network (AERONET) inversion products, the LSA from Moderate Resolution Imaging
166 Spectroradiometer (MODIS) surface albedo product, and the default atmospheric profile of SBDART
167 (MID-LATITUDE WINTER) are used as the input parameters for the SBDART model in simulating the
168 $DRFE_{dust}$. Therefore, the DRE_{dust} changing with the AOD due to both dust aerosol microphysical
169 properties (including the complex refractive index and the distribution of the size) and optical properties
170 (including the SSA and the ASY) are simulated by the SBDART model. The impacts of the
171 microphysical properties and the optical properties of the dust aerosol on the DRE_{dust} are analyzed in
172 this study.

173 2.2 Data

174 This paper aims to analyze the differences in dust microphysical properties and the $DRFE_{dust}$ over
175 the Taklimakan Desert and the Sahara Desert to confirm the influences of dust aerosol microphysical
176 properties on simulating the $DRFE_{dust}$. Also, the advantages of the satellite-based method in estimating
177 the $DRFE_{dust}$ are analyzed. Therefore, the $DRFE_{dust}$ over the Taklimakan Desert and the Sahara Desert is
178 estimated by using both satellite observations and dust microphysical properties.



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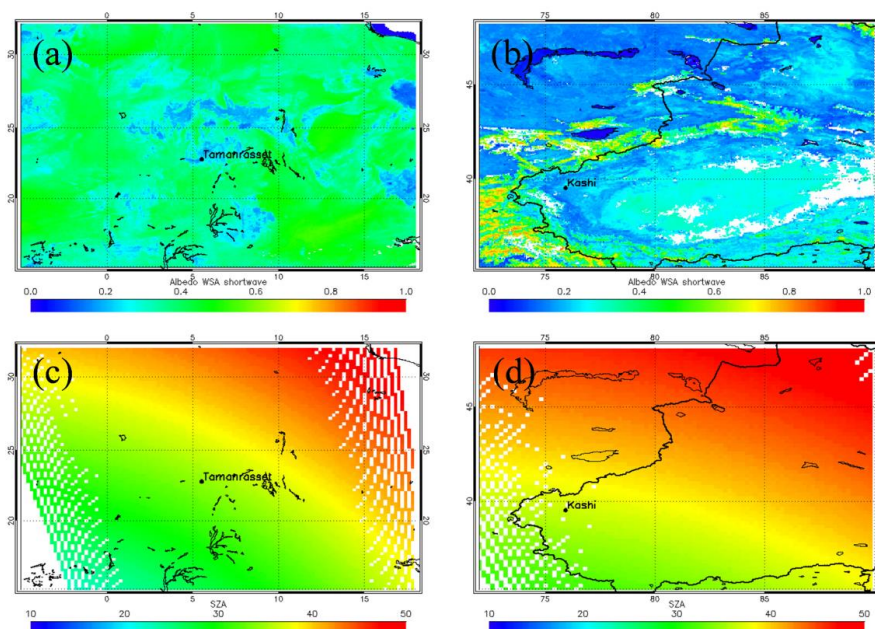
180 **Figure 2: The research regions and dust storms viewed by MODIS Aqua on 11 March and 9 April 2019.**

181 Fig. 2 shows the research regions (the red square areas) and the locations of in-situ sites
182 (Tamanrasset site and Kashi site, the red dots in the map and satellite images). Tamanrasset (22.79°N,
183 5.53°E, 1377 m above the mean sea level) locates in southern Algeria, which is free from the influence
184 of industrial activities. Thus, the aerosols measured in Tamanrasset can represent the pure dust aerosols
185 from the Sahara Desert (Guirado-Fuentes et al., 2014). Kashi (39.5°N, 75.9°E, 1320 m above the mean
186 sea level) locates in the vicinity of the Taklimakan Desert. Kashi represents a place affected by dust
187 aerosols transported from the Taklimakan Desert (Li et al., 2020). Thus, dust aerosols observed in
188 Tamanrasset and Kashi sites are typical samples of the dust aerosols from these two deserts. Moreover,
189 Tamanrasset and Kashi sites are similar in land surface type, altitude, and climate. As the LSA and the
190 SZA have a great impact on the SW radiative effect, the regions with similar LSA and SZA are chosen to
191 avoid the influence of different LSA and SZA on evaluating the differences of dust microphysical
192 properties and dust radiative effect from different dust source regions.

193 A dust storm occurred on 11 March 2019 in Tamanrasset. In Kashi, a dust storm occurred on 9
194 April 2019. These dust storms are shown visually by Aqua MODIS (Fig. 2). Fig. 3 shows the LSA and
195 the SZA observed by the AQUA satellite on 11 March 2019 in the Sahara Desert and on 25 April 2019

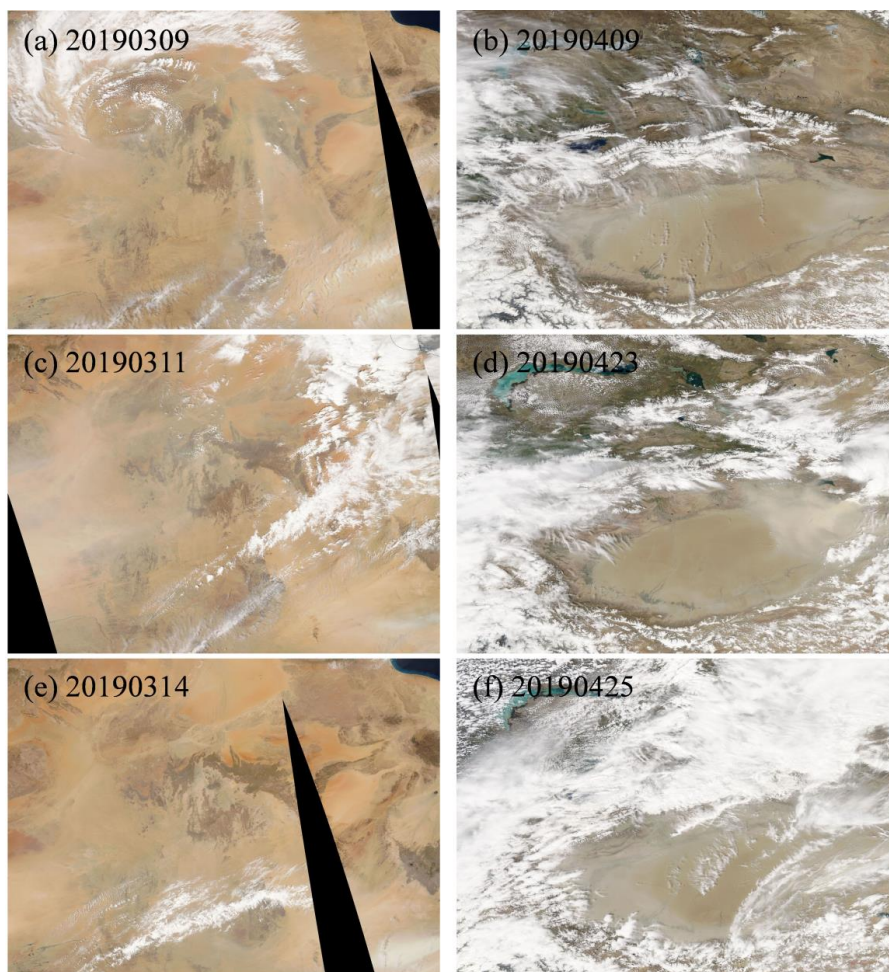


196 in Taklimakan Desert. In Fig. 3, the LSA and the SZA are similar in Tamanrasset and Kashi when the
197 satellite passes through. The data around Tamanrasset and Kashi in March and April are suitable for
198 analyzing the differences of dust microphysical properties and their influences on the $DRFE_{dust}$.



199
200 **Figure 3: (a) MODIS SW LSA and (c) SZA on 11 March 2019 over Tamanrasset; (b) MODIS SW LSA and**
201 **(d) SZA on 24 April 2010 over Kashi.**

202 The satellite-observed and dust microphysical properties data of the dust storms in March and
203 April 2019 in Tamanrasset and Kashi are collected to analyze the dust microphysical properties and
204 estimate the $DRFE_{dust}$ in the Taklimakan Desert and the Sahara Desert. Fig. 4 shows the satellite images
205 of these dust storms, which can be seen from the satellite images in cloud-free conditions over
206 Tamanrasset (left column of Fig. 4, Fig. 4(a), Fig. 4(c), Fig. 4(e)) and Kashi (right column of Fig. 4,
207 Fig. 4(b), Fig. 4(d), Fig. 4(f)). Both the satellite data and synergy dust microphysical properties data are
208 collected around Tamanrasset and Kashi sites for analyzing the differences in dust microphysical
209 properties and estimating the $DRFE_{dust}$.



210

211 **Figure 4: Dust storms viewed by AQUA/MODIS over target areas (Tamanrasset and Kashi).**

212 2.2.1 Satellite data

213 MODIS and CERES (Clouds and the Earth's Radiant Energy System) are the key instruments of
214 the AQUA and the TERRA satellite and are important in NASA's Earth Observing System (EOS). The
215 AOD products from MODIS and the radiative flux products at the TOA from CERES can be
216 synergistically used to estimate the $DRFE_{dust}$ directly.

217 Several algorithms have been developed for MODIS AOD remote-sensing products after MODIS
218 instruments were launched (Remer et al., 2005). Of these algorithms, the Deep Blue algorithm (Hsu et
219 al., 2004) solved the problems in aerosol retrieval by satellite remote-sensing for high reflectance land
220 surface types (such as arid, semi-arid, and desert areas), and retrieved the AOD over high reflectance land



221 surface types. In this paper, the deep blue AOD ($0.55\mu\text{m}$) data are used to discriminate the dust storm
222 regions. The LSA is also needed both in the satellite-based equi-albedo method and the RTM. The
223 MODIS Collection6 albedo product dataset (MCD43C3) (Schaaf et al., 2011;Schaaf et al., 2002;Schaaf
224 et al., 2008) provides high-quality land surface reflectance and albedo data over various types of land
225 surfaces by using anisotropy retrievals algorithm (Jin et al., 2003;Liang et al., 2002;Liu et al.,
226 2009;Román et al., 2010). The MCD43C3 product dataset is available from the Land Processes
227 Distributed Active Archive Center (LP DAAC) of NASA. The white-sky albedo from the MCD43C3
228 product is used to get the SW broadband LSA.

229 CERES single scanner footprint (SSF) level 2 dataset can provide the radiative flux at the TOA in
230 three broadband channels. Here the instantaneous SW channel ($0.3\text{--}5.0\mu\text{m}$) radiative flux at the TOA
231 from CERES SSF level 2 dataset is used. MODIS and CERES are onboard in the same satellite
232 platform (AQUA). The radiative flux derived from CERES is co-located with the MODIS scene. The
233 DRE_{dust} and the $\text{DRFE}_{\text{dust}}$ at the TOA are estimated by synergistically using MODIS and CERES
234 products.

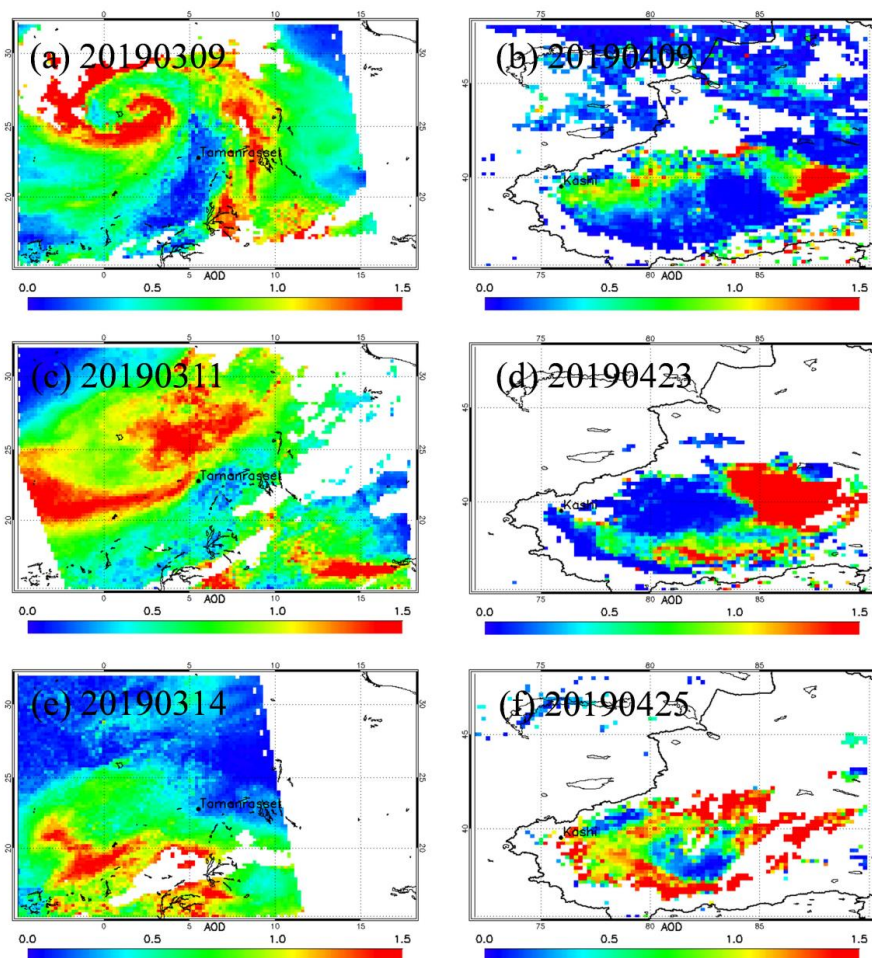
235 **2.2.2 Dust microphysical properties data**

236 The Aerosol Robotic Network (AERONET) is the largest ground-based network for measuring
237 aerosols with more than 400 sites installed.

238 The AERONET provides microphysical properties and optical properties of the aerosols at four
239 wavelengths (440, 675, 870, and 1020 nm). The AOD product is directly measured by the sun
240 photometer. The inversion algorithm retrieves the physical properties of aerosols such as volume size
241 distributions and the complex refractive index, and optical properties such as the SSA and the ASY
242 (Dubovik and King, 2000;Dubovik et al., 2006).



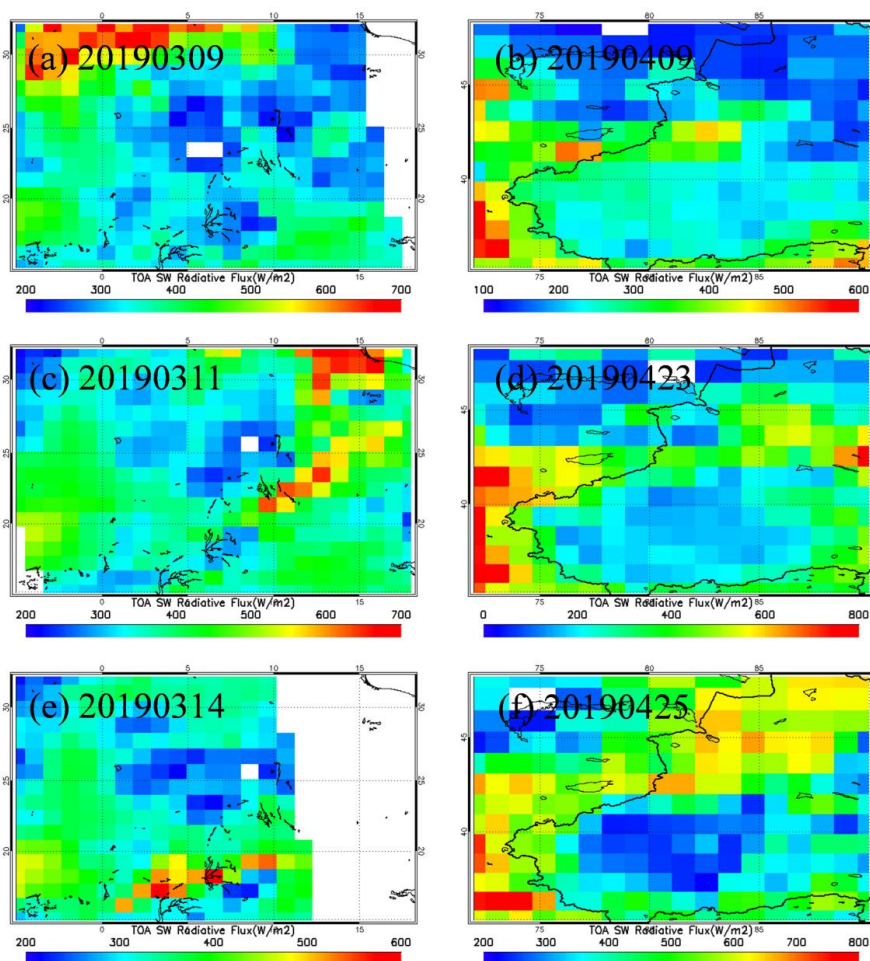
243 3 DRFE_{dust} estimated based on satellite observations



244

245 **Figure 5: AOD at 0.55 μm (τ_{550}) of the dust storm in March 2019 over Tamanrasset and that in April 2019**
246 **over Kashi.**

247 MODIS L2 deep blue AOD product of the dust storm in March 2019 over Tamanrasset and that in
248 April 2019 over Kashi are shown in Fig. 5. The missing data are shown in white; the high dust loading
249 regions are shown in red; the low dust loading regions are shown in blue. Fig. 5 shows that there are
250 heavy dust storms over Tamanrasset and Kashi with AOD great than 1.0 detected by MODIS.



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Figure 6: TOA SW radiative flux in March 2019 over Tamanrasset and that in April 2019 over Kashi.

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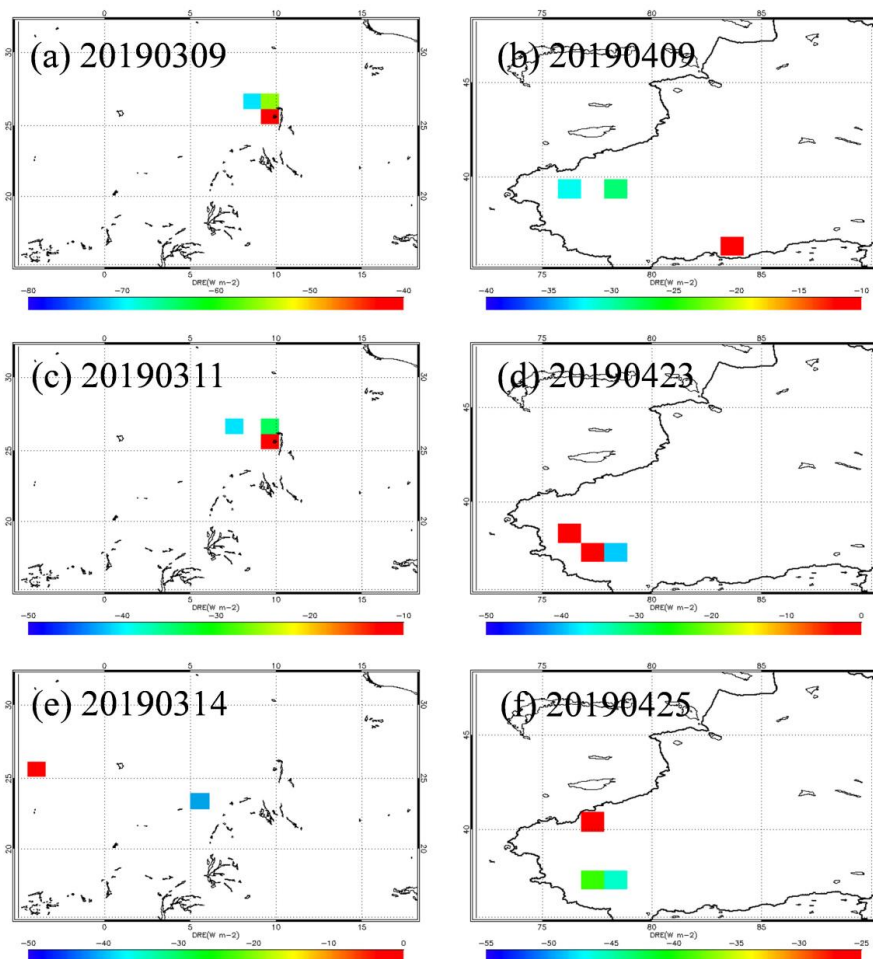
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Fig. 6 shows the TOA SW radiative flux measured by CERES in March 2019 over Tamanrasset and that in April 2019 over Kashi during the dust storms. The TOA SW radiative flux distribution shows the highest value over cloud conditions. The values in dust storm regions are higher than those in clear-sky regions. It is due to the fact that the SW albedo of the aerosols in the cloud and the dust is higher than those on the land surface. Thus, dust aerosols have a negative radiative effect in the SW spectrum. Following the equi-albedo method (Tian et al., 2019), the DRE_{dust} is estimated based on the measurements from MODIS and CERES both aboard on the AQUA satellite.

As Fig. 4, Fig. 5 and Fig. 6 shown, the spatial resolution of TOA flux from CERES/SSF product is $1^\circ \times 1^\circ$ grid, and LSA, SZA, AOD data from satellite have the different spatial resolution. In order to



262 match up LSA, SZA and AOD data with CERES TOA SW fluxes, we have resampled LSA, SZA and
263 AOD data into CERES SSF product horizontal spatial resolution. Then the F_{clr} and DRE_{dust} over
264 Tamanrasset and Kashi can be estimated following equi-albedo method.



265
266 **Figure 7: DRE_{dust} on March 2019 over Tamanrasset and on April 2019 over Kashi.**

267 Fig. 7 shows the distribution maps of the DRE_{dust} . The high dust aerosol loading regions show
268 significant negative radiative forcing. It indicates that the dust aerosol loading is negatively correlated
269 with the DRE_{dust} in these dust storm events. The distribution maps of the LSA and the SZA (Fig. 3)
270 show that the mean SW LSA measured by MODIS is around 0.18 and the mean SZA is around 35
271 degrees in Tamanrasset and Kashi. The distribution maps also show that the LSA and SZA vary greatly
272 in the same satellite scan image. To avoid the influence of the LSA and SZA in estimating the $\text{DRFE}_{\text{dust}}$,



273 pixels with LSA of 0.16–0.20 and SZA of 32–38 degrees are chosen to derive the $DRFE_{dust}$. Therefore,
 274 only few pixels having similar values of the LSA and the SZA over Tamanrasset and Kashi are picked
 275 for estimating the DRE_{dust} and the $DRFE_{dust}$. The influences of the dust microphysical properties on the
 276 $DRFE_{dust}$ are investigated. These pixels of the DRE_{dust} and its co-located AOD values are illustrated in
 277 Table 1.

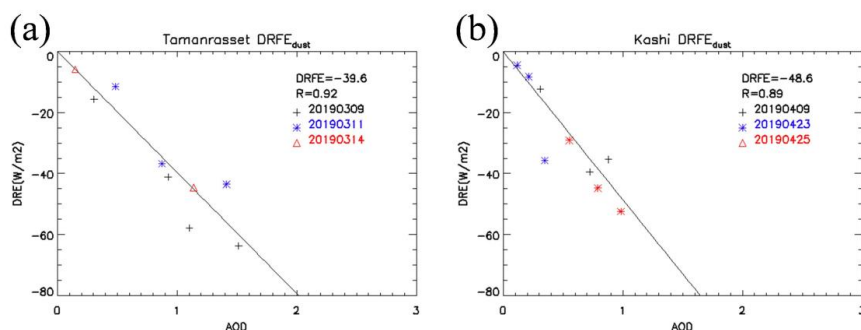
278 **Table 1: DRE_{dust} and AOD in March 2019 over Tamanrasset and that in April 2019 over Kashi during the**
 279 **dust storms.**

Regions & Dates		Properties	AOD	DRE_{dust}
Sahara Desert	20190309		0.92	-41.2
			1.51	-63.7
			1.11	-57.8
	20190311		0.31	-15.6
			0.48	-11.5
			1.41	-43.5
Taklimakan Desert	20190314		0.87	-36.7
			1.14	-44.6
	20190409		0.15	-5.8
			0.31	-12.3
20190423		0.72	-39.5	
		0.88	-35.4	
		0.21	-8.3	
		0.35	-35.7	
		0.11	-4.5	
20190425		0.79	-44.8	
		0.98	-52.4	
		0.55	-29.1	

280



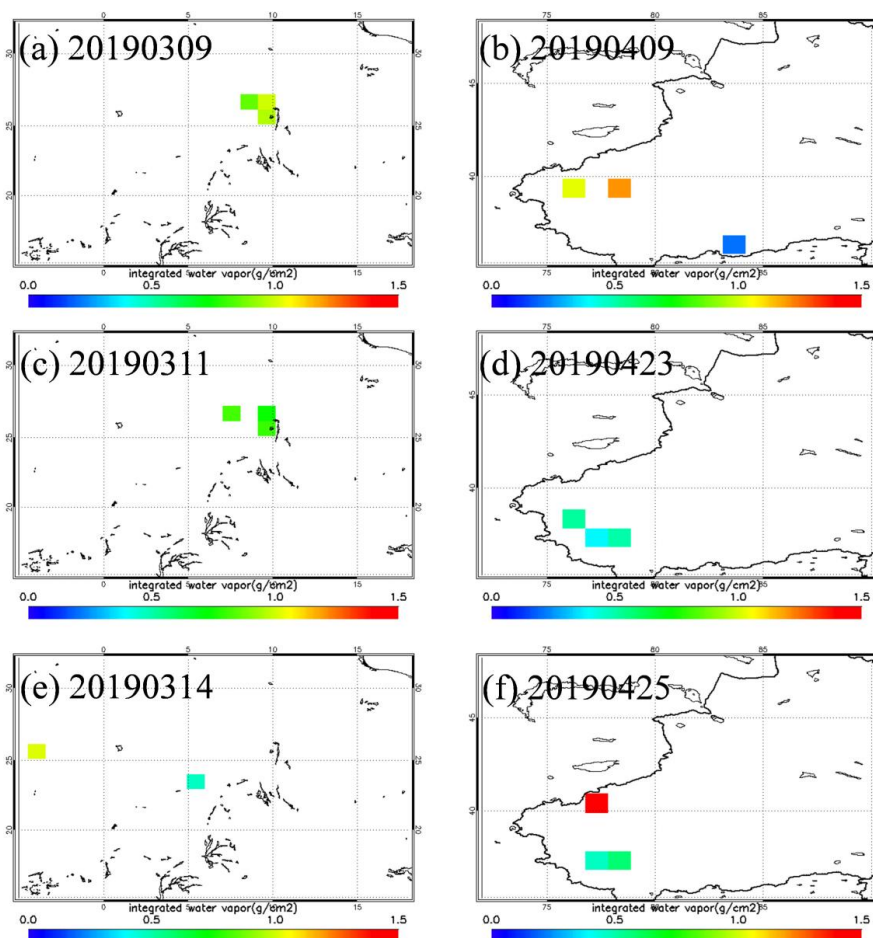
281 According to the definition, the $DRFE_{\text{dust}}$ represents the DRE_{dust} of a certain AOD at per unit area
282 during these storms in the desert dust source regions. Therefore, the $DRFE_{\text{dust}}$ can be estimated by
283 fitting the DRE_{dust} and the AOD.



284
285 **Figure 8: DRE_{dust} in (a) March 2019 over Tamanrasset and (b) April 2019 over Kashi.**

286 The linear relationship between the DRE_{dust} and the AOD can be found during dust storms around
287 Tamanrasset and Kashi, which is also investigated in previous studies (Kumar et al. 2015; Jose et al.
288 2016). Then, the $DRFE_{\text{dust}}$ can be estimated by regressing the DRE_{dust} and the AOD. In Fig. 8, the mean
289 $DRFE_{\text{dust}}$ of the dust storms is $-39.6 \text{ Wm}^{-2}\tau^{-1}$ over Tamanrasset and $-48.6 \text{ Wm}^{-2}\tau^{-1}$ over Tamanrasset.
290 The correlation coefficients are high with $R = 0.92$ in March 2019 over Tamanrasset and $R = 0.89$ in
291 April 2019 over Kashi. The AOD and DRE_{dust} values are well correlated. Positive dust AOD is
292 associated with negative DRE_{dust} .

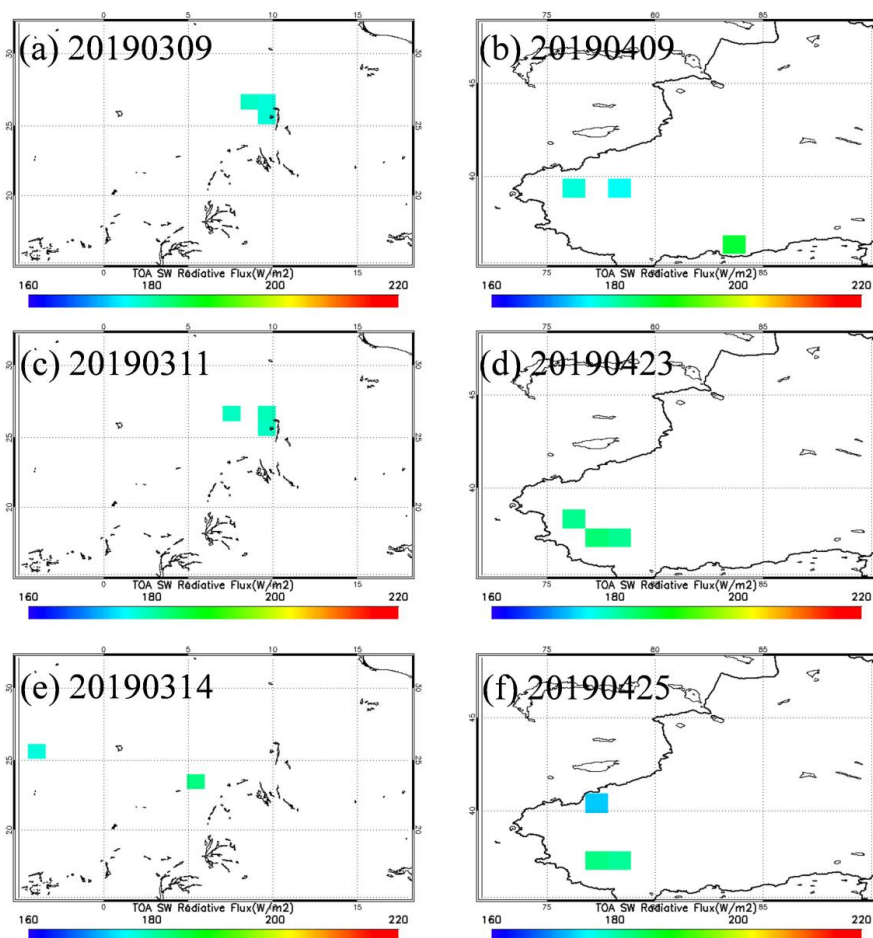
293 The equi-albedo method directly estimates the DRE_{dust} and the $DRFE_{\text{dust}}$ based on the satellite
294 observations. Therefore, the accuracy of the results (DRE_{dust} and $DRFE_{\text{dust}}$) derived from the
295 equi-albedo method is highly dependent on the accuracy of satellite observations. Therefore, the
296 uncertainties of the $DRFE_{\text{dust}}$ derived from the equi-albedo method mainly include the instantaneous
297 SW flux error from CERES measurements, the estimation uncertainties of the F_{clr} over the dust storm
298 region, and the uncertainty in the deep blue AOD product. Beside that, according to our sensitivity test
299 in the previous studies (Tian et al., 2019), the atmospheric profile, water vapor and height of dust layer
300 have insignificant influence on SW radiative flux at the TOA. It is reasonable to use same water vapor
301 and pre-defined vertical distribution for dust aerosols in one scene of satellite data. However, the
302 assumption of pixels has same water vapor and pre-defined aerosol vertical distribution over one scene
303 of satellite data still cause small uncertainty.



304

305 **Figure 9: Integrated water vapor (g/cm^2) from European Centre for Medium-range Weather**
306 **Forecasts (ECMWF) reanalyses dataset on March 2019 over Tamanrasset and on April 2019 over Kashi.**

307 Fig.9 shows integrated water vapor from ECMWF reanalyses dataset on March 2019 over
308 Tamanrasset and on April 2019 over Kashi. The integrated water vapor varies little over research areas,
309 the regional mean differences are $0.51\text{g}/\text{cm}^2$ and $0.18\text{g}/\text{cm}^2$ over Kashi and Tamanrasset, respectively.
310 In order to test the uncertainty caused by the varies of integrated water vapor over research areas, we
311 calculated SW radiative flux at the TOA in difference of integrated water vapor based on SBDART
312 model.

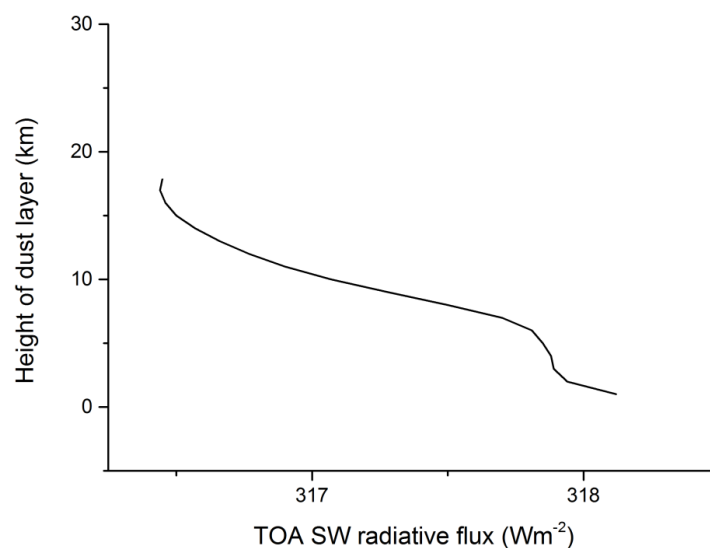


313

314 **Figure 10:** SBDART simulated clear-sky TOA radiative flux using integrated water vapor (g/cm^2) from
315 ECMWF reanalyses dataset on March 2019 over Tamanrasset and on April 2019 over Kashi.

316 Fig.10 shows SBDART simulated clear-sky TOA radiative flux using integrated water vapor from
317 ECMWF reanalyses dataset on March 2019 over Tamanrasset and on April 2019 over Kashi. The
318 regional mean differences of TOA radiative flux are 2.21% and 0.85% over Kashi and Tamanrasset,
319 respectively. This indicates the varies of integrated water vapor caused uncertainties of TOA radiative
320 flux are 2.21% and 0.85% over Kashi and Tamanrasset.

321 For the assumption of vertical profile for dust aerosols, we also tested the sensitivity of radiative
322 flux at the top of atmosphere to changes height of dust layer with SBDART model.



323

324 **Figure 11: The sensitivity test of SW radiative flux at the TOA to changes height of dust layer.**

325 As Fig.11 shown, the SW radiative flux at the TOA was decreased with the height of dust layer
326 was increased from 0km to 18km. However, the contents of the SW radiative flux change little with the
327 height of dust layer increased (within 1.5Wm^{-2} , 0.47%), which is little than CERES observation errors.

328 According to our previous study (Tian et al., 2019), the instantaneous SW flux error from CERES
329 measurements is about 3.13%, the estimation uncertainty of the F_{clr} is 3.15%, the uncertainty of the
330 deep blue AOD retrieved by MODIS is about 15% (Sayer et al., 2014), and the uncertainties of using
331 same water vapor (2.21% and 0.85% over Kashi and Tamanrasset, respectively) and pre-defined
332 aerosol vertical distribution (0.47%) over one scene of satellite data. Then, the total uncertainties of the
333 $\text{DRFE}_{\text{dust}}$ can be calculated by the equation Eq. (3) (Zhang et al., 2005).

$$334 U_t = \exp[\sum(\log U_s)^2]^{1/2} \quad (3).$$

335 U_s is the synthetical uncertainty factor of each source of the uncertainty (including the
336 instantaneous SW flux error from CERES measurements, the estimation uncertainty of the F_{clr} , and the
337 uncertainty of the deep blue AOD retrieved by MODIS). U_t is the total uncertainty of the $\text{DRFE}_{\text{dust}}$,
338 which is 25.37% and 28.19% ($10.0\text{Wm}^{-2}\tau^{-1}$ and $13.7\text{Wm}^{-2}\tau^{-1}$) in Tamanrasset and Kashi, respectively.
339 Therefore, the $\text{DRFE}_{\text{dust}}$ are $-39.6 \pm 10.0\text{Wm}^{-2}\tau^{-1}$ in March 2019 over Tamanrasset and -48.6 ± 13.7
340 $\text{Wm}^{-2}\tau^{-1}$ in April 2019 over Kashi.

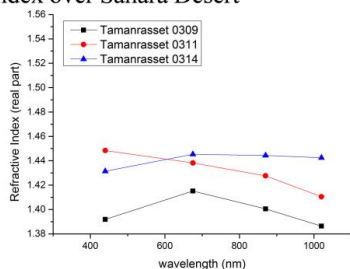


341 **4 Deriving $DRFE_{dust}$ from the RTM simulations**

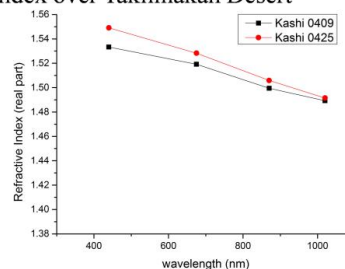
342 **4.1 Dust microphysical properties**

343 The focuses of this paper are the differences in the dust microphysical properties from different
344 dust source regions and the impacts of the dust microphysical properties on the $DRFE_{dust}$ simulation. As
345 important parameters concerning the radiative impacts, the volume size distribution and the refractive
346 index of the dust aerosol are compared in the dust storms over Tamanrasset and Kashi detected by
347 MODIS.

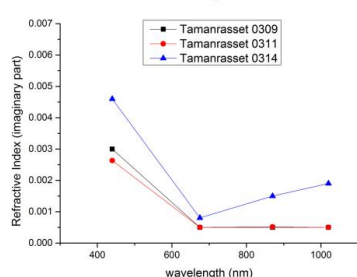
(a) Real parts of the complex refractive index over Sahara Desert



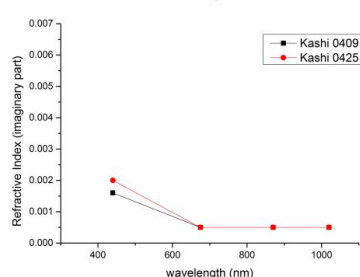
(b) Real parts of the complex refractive index over Taklimakan Desert



(c) Imaginary parts of the complex refractive index over Sahara Desert



(d) Imaginary parts of the complex refractive index over Sahara Desert



348
349 **Figure 12: Real and imaginary parts of the dust complex refractive index from the Sahara Desert and the**
350 **Taklimakan Desert.**

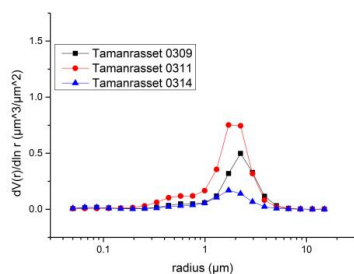
351 The refractive index is a measurement of the aerosol refraction and absorption efficiency. Aerosols
352 with high real parts of the complex refractive index values are indicated to be scattering types.
353 Conversely, aerosols with high imaginary parts are indicated to be absorbing types (Zhang et al., 2006).
354 Figure 12 shows the real and imaginary parts of the dust complex refractive index from the Taklimakan
355 Desert and the Sahara Desert during the dust storms. In Fig. 12, dust aerosols from the Taklimakan
356 desert (Fig. 12 (b)) have higher real parts and lower imaginary parts (Fig. 12 (d)) than the Sahara desert



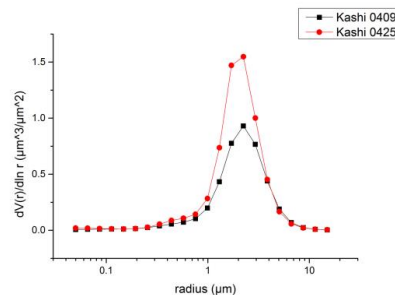
357 (Fig. 12 (a) and Fig. 12 (c)), showing that the dust aerosols from the Taklimakan Desert have stronger
358 scattering effects.

359 The volume size distribution of dust aerosols clearly shows the particle size difference between
360 dusty and clear-sky days. The moderate and coarse aerosol particles with a radius within 0.5–10 μm
361 show a significant increase under dusty conditions than those under non-dusty days. It is indicated that
362 the quantity of the coarse mineral dust particles increases because of the dust storms. Figure 13
363 illustrates the variation of the dust aerosol size distribution during the dust storms in March 2019 over
364 Tamanrasset and in April 2019 over Kashi. Most maximum dust aerosol size distribution peaks at the
365 radius of 1.71 in Tamanrasset and 2.24 in Kashi. Moreover, the peak values are higher in Kashi. It is
366 indicated that the dust storm is stronger in April 2019 over Kashi and the coarse mode aerosol particles
367 increase in particle volume compared with those in the dust storm in March 2019 over Tamanrasset.

(a) Dust aerosol size distribution over
Sahara Desert



(b) Dust aerosol size distribution over
Taklimakan Desert



368
369 **Figure 13: Dust aerosol size distribution over (a) the Sahara Desert and (b) the Taklimakan Desert.**

370 4.2 Dust optical properties

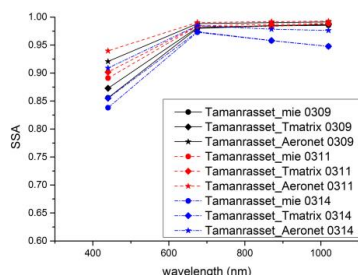
371 The dust optical properties can be calculated by synergistically using the real and imaginary parts of
372 the dust complex refractive index and the dust aerosol size distribution.

373 The SSA and the ASY are two key parameters determining the DRE_{dust} and the $\text{DRFE}_{\text{dust}}$. Accurate
374 measurements of the SSA and the ASY are important for the assessment of the direct effect of aerosols
375 on climate (Qie et al., 2019). The dust aerosol optical properties are calculated by using the Mie theory,
376 the T-matrix method, and the AERONET inversion products (Dubovik and King, 2000; Dubovik et al.,
377 2006).

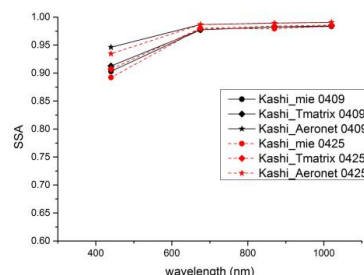


378 The SSA is presented as the ratio between the aerosol scattering and extinction coefficients. The
379 dust SSA describes the scattering properties of the dust aerosols. The SSA can largely determine the
380 magnitudes and signs of the DRE_{dust} and the $DRFE_{dust}$. Strongly scattering dust aerosols (i.e., $SSA = 1$)
381 always cause negative DRE_{dust} . By contrast, low SSA aerosols often cause positive DRE_{dust} , especially
382 over high LSA regions as the light absorbed by the aerosols can reduce the cooling effect. A high SSA is
383 correlated with low real parts of the complex refractive index, while a strong absorption is correlated
384 with a high imaginary part of the complex refractive index. Together with the size distribution, real parts
385 of the complex refractive index can determine the magnitude of the SSA.

(a) Single Scattering Albedo of dust aerosols over Sahara Desert



(b) Single Scattering Albedo of dust aerosols over Taklimakan Desert

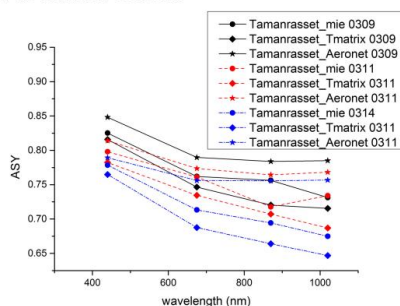


386
387 **Figure 14: Single scattering albedo from (a) the Sahara Desert and (b) the Taklimakan Desert..**

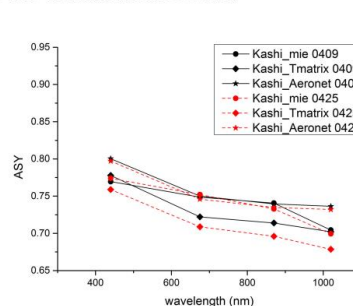
388 Figure 14 shows the variabilities of the dust aerosol SSA between different dust source regions and
389 different calculation methods. In Fig. 14, the maximum SSA value mostly occurs at the wavelength of
390 1020 nm, which indicates that the SSA is dependent on wavelength. Moreover, dust aerosols from the
391 Taklimakan desert (Kashi) in the figure have higher SSA value using both the Mie theory and the
392 T-matrix method. The higher value of SSA shows that dust aerosol particles scatter more
393 predominantly and strongly in the Taklimakan desert (Kashi), which may cause more significant
394 negative radiative forcing than the dust aerosols from the Sahara Desert (Tamanrasset).



(a) Asymmetry factor of dust aerosols over Sahara Desert



(b) Asymmetry factor of dust aerosols over Taklimakan Desert



395

396 **Figure 15: Asymmetry factor in (a) the Sahara Desert and (b) the Taklimakan Desert.**

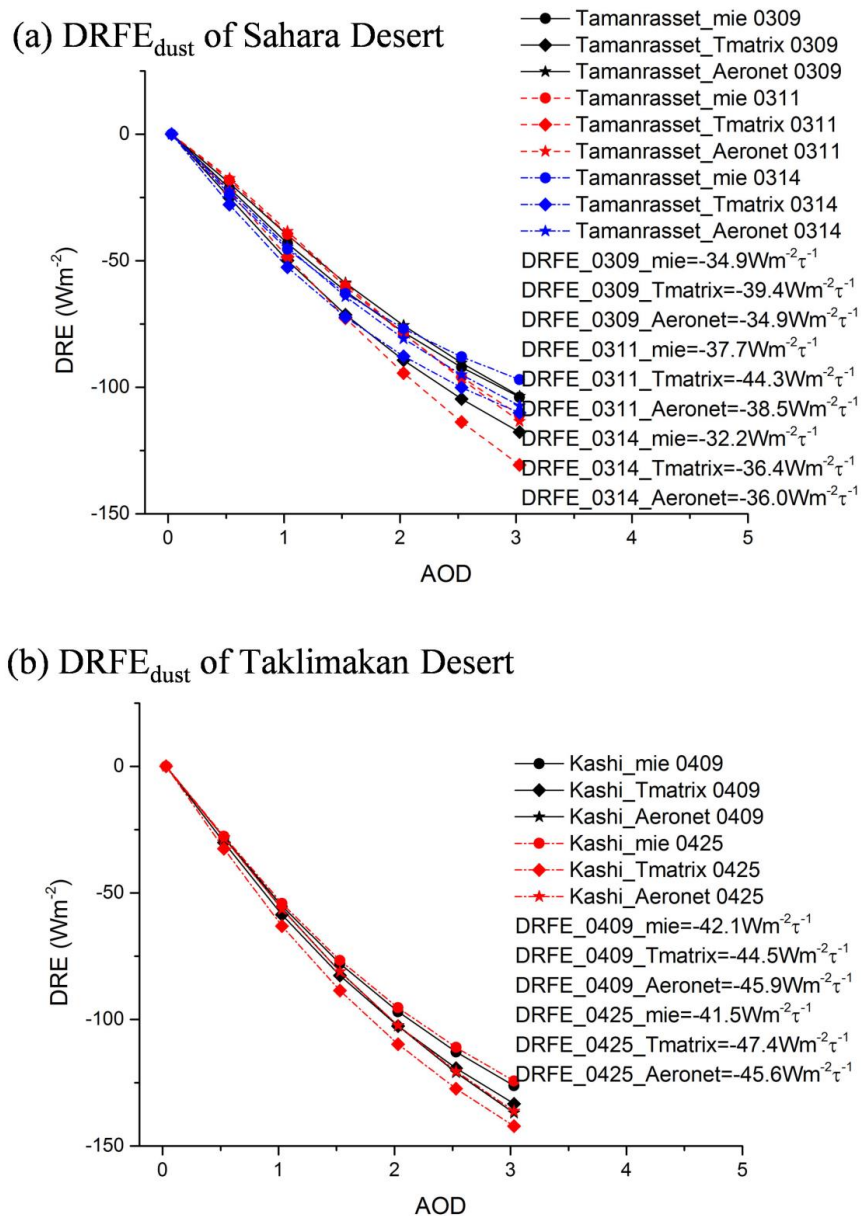
397 The ASY indicates the relative strength of the forward scattering, which determines the integrated
398 fractions of the energy that scatter backward and forward. The dust aerosol particles with sharp peaks in
399 the forward direction (0° scattering angle) have positive ASY. The ASY value increases with the
400 particle size.

401 The ASY in Fig. 15 shows marked spectral variation with higher values at shorter wavelengths. It
402 can be found that the dust aerosols from the Sahara Desert (Tamanrasset) have higher values of the ASY
403 than those from the Taklimakan desert (Kashi) in both the Mie theory and the T-matrix method. The
404 high values (over 0.80 at 440 nm) reflect the dominance of the absorbing of dust aerosols. The stronger
405 backward scattered energy may cause higher negative radiative forcing in the Taklimakan Desert
406 (Kashi).

407 According to the analyses of the microphysical properties and the optical properties, the dust
408 aerosols from the Taklimakan Desert (Kashi) scatter strongly. The negative $DRFE_{\text{dust}}$ from the
409 Taklimakan desert (Kashi) should be more significant than those from the Sahara Desert (Tamanrasset).
410 The results are in good agreement with those estimated by the satellite observations.



411 4.3 DRFE_{dust} derived from the RTM simulations



412

413 **Figure 16: DRFE_{dust} simulated by the SBDART in (a) Tamanrasset and (b) Kashi.**

414 The DRFE_{dust} estimated directly by the satellite observation is compared with that simulated by the

415 SBDART to verify the reliability. As shown in Fig. 16, with higher aerosol scattering (higher SSA) and



416 higher backward scattering coefficients (lower ASY), the negative $DRFE_{dust}$ from Kashi is more
417 significant. The mean $DRFE_{dust}$ from Kashi is $-37.1 \text{ W m}^{-2} \tau^{-1}$. The dust aerosols from Kashi have
418 stronger cooling effects than those from Tamanrasset, in which the mean $DRFE_{dust}$ is $-44.5 \text{ W m}^{-2} \tau^{-1}$.
419 The results are in good agreement with those estimated by the satellite observations. The $DRFE_{dust}$
420 estimated by the dust optical properties derived from the T-matrix method and the AERONET is closer
421 to those estimated by the satellite observations, which indicates that most dust aerosols are
422 non-spherical in the natural environment.

423 The results also show that the dust microphysical properties can significantly influence the
424 $DRFE_{dust}$. The mean difference of the $DRFE_{dust}$ between Tamanrasset and Kashi is 9.0% ($7.4 \text{ W m}^{-2} \tau^{-1}$).
425 Even for the same dust microphysical property, the $DRFE_{dust}$ varies significantly according to whether
426 the dust particles are considered spherical or non-spherical in different methods. For the differences of
427 the $DRFE_{dust}$ estimated using different methods, the mean standard deviations are 7.6% ($2.8 \text{ W m}^{-2} \tau^{-1}$)
428 in Tamanrasset and 6.8% ($3.0 \text{ W m}^{-2} \tau^{-1}$) in Kashi. Moreover, Li et al. (2020) pointed out that the
429 atmospheric profile, LSA and SZA, can also influence the simulation of the instantaneous $DRFE_{dust}$,
430 which agrees with our previous study (Tian et al., 2019). Additionally, it is difficult for climate models
431 or in-situ measurements to get the real distribution of the aerosol properties at a large spatial extent.
432 Also, it is hard to evaluate the uncertainties in radiative transfer simulations. It can cause significant
433 errors in evaluating the modulating effects of the mineral dust aerosols on climate (Huang et al.,
434 2009; Li et al., 2020).

435 **5 $DRFE_{dust}$ in the satellite-based observation and the simulation of the RTM**

436 According to the analyses of the dust aerosol microphysical properties and optical properties, the
437 dust aerosols from the Taklimakan Desert (Kashi) should scatter strongly. The RTM simulation results
438 are in good agreement with the results estimated by the satellite observation. Previous studies also
439 estimated the $DRFE_{dust}$ in the Sahara Desert and the Taklimakan Desert (Li et al., 2020; Li et al.,
440 2004; Garc ía et al., 2012; Xia and Zong, 2009), which validate our results.

441 **Table 2: SW $DRFE_{dust}$ from different studies.**

Dust source regions	Research	Model/Method	$DRFE_{dust} (\text{W m}^{-2} \tau^{-1})$	Description
Sahara	Garc ía et al	Ground+GAM	~ -35	AERONET



Desert	(2012)	E		DRFE _{dust} in December-January-February, with LSA<0.3.
	Li et al (2004)	Satellite+SBD ART	-35±3 (summer) -26±3 (winter)	Binned mean fitting TOA diurnal mean DRFE _{dust} over the Atlantic Ocean near the African coast.
	This paper	Satellite Satellite+SBD ART	-39.6±10.0 (Satellite) -32.2 ~ -44.3 (SBDART)	
Taklimakan Desert	Li et al (2020)	Ground+SBD ART	-45~-50	Instantaneous DRFE _{dust} at 04:08 UTC.
	García et al (2012)	Ground+GAM E	~ -45	AERONET DRFE _{dust} in March-April-May, with LSA<0.3.
	Xia and Zong (2009)	Satellite + SBDART	-48.1	Instantaneous DRFE _{dust} at about 05:00 UTC.
	This paper	Satellite Satellite+SBD ART	-48.6±13.7 (Satellite) -41.5 ~ -47.4 (SBDART)	

442

443 Table 2 illustrates the SW DRFE_{dust} of the Sahara Desert and the Taklimakan Desert in previous
 444 studies. García et al. (2012) evaluated the DRFE_{dust} based on the GAME model and the AERONET
 445 retrievals, which indicated that the mean DRFE_{dust} is around $-35 \text{ W m}^{-2} \tau^{-1}$ in the Sahara Desert and
 446 $-45 \text{ W m}^{-2} \tau^{-1}$ in the Taklimakan Desert in similar observational conditions (i.e., for the SZA between
 447 55° and 65° ; for the LSA below 0.3). Li et al. (2004) estimated the diurnal mean DRFE_{dust} at the TOA
 448 ($-35 \pm 3 \text{ W m}^{-2} \tau^{-1}$ in summer; $-26 \pm 3 \text{ W m}^{-2} \tau^{-1}$ in winter) over the Atlantic Ocean near the African
 449 coast. The results indicated that lower uncertainties are derived from the standard deviation of the
 450 best-fit curve around the observed points due to the binned mean fitting. For the Taklimakan Desert, Li
 451 et al. (2020) estimated the instantaneous SW DRFE_{dust} at 04:08 UTC (around $-43 \text{ W m}^{-2} \tau^{-1}$) at the
 452 TOA of Kashi on 25 April 2019. In this paper, the DRFE_{dust} of the Taklimakan Desert is estimated with
 453 the same dust properties referring to the works of Li et al. (2020). Furthermore, Xia and Zong (2009)



454 used both the satellite data and the SBDART model to represent the instantaneous (about 05:00 UTC)
455 SW $DRFE_{dust}$, which is $-48.1 \text{ W m}^{-2} \tau^{-1}$ at the TOA (Xia and Zong, 2009). Through comparison, it is
456 found that the satellite-based equi-albedo method and the SBDART model-derived SW $DRFE_{dust}$ are
457 $-39.6 \pm 10.0 \text{ W m}^{-2} \tau^{-1}$ and -32.2 to $-44.3 \text{ W m}^{-2} \tau^{-1}$ at the TOA over the Sahara Desert, respectively,
458 which are $-48.6 \pm 13.7 \text{ W m}^{-2} \tau^{-1}$ and -41.5 to $-47.4 \text{ W m}^{-2} \tau^{-1}$ at the TOA over the Taklimakan Desert,
459 respectively. The methods and results in these studies are comparable despite the differences. The
460 results show that the negative $DRFE_{dust}$ from the Taklimakan Desert is more significant than those from
461 the Sahara Desert. As the SZA and LSA variations are considered in these studies, the results in this
462 paper are reasonable and reliable. The compared results show that the $DRFE_{dust}$ derived from the
463 satellite-based equi-albedo method is closer to that in previous studies with lower uncertainty. The
464 $DRFE_{dust}$ estimated by the satellite-based equi-albedo method is obtained without the dust
465 microphysical properties being assumed. The uncertainties are mostly caused by observation errors.
466 Therefore, the uncertainties can be evaluated more reasonably. It provides a direct way to validate the
467 $DRFE_{dust}$ and the $DRFE_{dust}$.

468 **6 Discussion and conclusions**

469 This study analyzes the differences in the dust microphysical properties and the $DRFE_{dust}$ over the
470 Taklimakan Desert and the Sahara Desert during dust storms. The satellite-based equi-albedo method
471 and the RTM are both used to estimate the $DRFE_{dust}$ in this study. By comparing the results from
472 different methods and dust source regions, the $DRFE_{dust}$ differences caused by dust microphysical
473 properties and particle shapes are discussed.

474 The results show that the dust aerosols from the Taklimakan Desert have higher aerosol scattering
475 (higher SSA) and backward scattering coefficients (lower ASY), and it causes more significant
476 negative $DRFE_{dust}$ ($-48.6 \pm 13.7 \text{ W m}^{-2} \tau^{-1}$ by the satellite; -41.5 to $-47.4 \text{ W m}^{-2} \tau^{-1}$ by the SBDART)
477 than that in the Sahara Desert ($-39.6 \pm 10.0 \text{ W m}^{-2} \tau^{-1}$ by the satellite; -32.2 to $-44.3 \text{ W m}^{-2} \tau^{-1}$ by the
478 SBDART). It is indicated that the dust microphysical properties and particle shapes can significantly
479 influence on the $DRFE_{dust}$. The information on the accurate dust microphysical properties and dust
480 origins is highly required in the $DRFE_{dust}$ simulation. The scant measurements on dust microphysical
481 properties can cause large uncertainties in simulating the $DRFE_{dust}$. Previous studies proved that the



482 results in this paper are reasonable and reliable. The $DRFE_{dust}$ derived from the satellite-based
483 equi-albedo method is close to the results in previous studies.

484 However, there are still uncertainties in the simulation of the $DRFE_{dust}$. In contrast, the $DRFE_{dust}$
485 can be estimated directly from the satellite observation using the equi-albedo method without any
486 assumptions of the microphysical properties of dust aerosols. It has unique advantages in estimating the
487 $DRFE_{dust}$. Also, it can validate the DRE_{dust} and the $DRFE_{dust}$ derived from the numerical models more
488 directly.

489 **Data availability**

490 The CERES data can be accessed from the Atmospheric Sciences Data Center of NASA Langley
491 Research Center (https://ceres.larc.nasa.gov/order_data.php). The AQUA/MODIS aerosol Products
492 (MYD04_L2) can be accessed from the NASA Level-1 and Atmosphere Archive and Distribution
493 System (LAADS) Distributed Active Archive Center (DAAC) website
494 (<https://ladsweb.modaps.eosdis.nasa.gov/>). The MODIS albedo products (MCD43C3 Version 6) can be
495 accessed from the NASA LP DAAC website (<https://lpdaac.usgs.gov/tools/data-pool/>). The
496 AERONET data were obtained from the AERONET website (<http://aeronet.gsfc.nasa.gov>).

497 **Author contributions**

498 PZ and LC designed the study, and LT performed the study with suggestions from PZ and LC. LB
499 improved the scattering calculating method of dust particles. Both authors contributed to the writing of
500 this article.

501 **Competing interests**

502 The authors declare that they have no conflict of interest.

503 **Special issue statement**

504 This article is part of the special issue “Satellite and ground-based remote sensing of aerosol
505 optical, physical, and chemical properties over China”. It is not associated with a conference.



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