



1 **Insights on Atmospheric Oxidation Processes by Performing Factor Analyses on**  
2 **Sub-ranges of Mass Spectra**

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13

14 **Abstract**

15 With the recent developments in mass spectrometry, combined with the strengths of factor analysis  
16 techniques, our understanding of atmospheric oxidation chemistry has improved significantly. The  
17 typical approach for using techniques like positive matrix factorization (PMF) is to input all measured  
18 data for the factorization in order to separate contributions from different sources and/or processes to  
19 the total measured signal. However, while this is a valid approach for assigning the total signal to  
20 factors, we have identified several cases where useful information can be lost if solely using this  
21 approach. For example, gaseous molecules emitted from the same source can show different temporal  
22 behaviors due differing loss terms, like condensation at different rates due to different molecular  
23 masses. This conflicts with one of PMF's basic assumptions of constant factor profiles. In addition,  
24 some ranges of a mass spectrum may contain useful information, despite contributing only minimal  
25 fraction to the total signal, in which case they are unlikely to have a significant impact on the  
26 factorization result. Finally, certain mass ranges may contain molecules formed via pathways not  
27 available to molecules in other mass ranges, e.g. dimeric species versus monomeric species. In this  
28 study, we attempted to address these challenges by dividing mass spectra into sub-ranges and  
29 applying the newly developed binPMF method to these ranges separately. We utilized a dataset from  
30 a chemical ionization atmospheric pressure interface time-of-flight (CI-API-TOF) mass spectrometer  
31 as an example. We compare the results from these three different ranges, each corresponding to  
32 molecules of different volatilities, with binPMF results from the combined range. Separate analysis  
33 showed clear benefits in dividing factors for molecules of different volatilities more accurately, in



34 resolving different chemical processes from different ranges, and in giving a chance for high-  
35 molecular-weight molecules with low signal intensities to be used to distinguish dimeric species with  
36 different formation pathways. In addition, daytime dimer formation (diurnal peak around noon) was  
37 identified, which may contribute to NPF in Hyytiälä. Also, dimers from  $\text{NO}_3$  oxidation were separated  
38 by the sub-range binPMF, which would not be identified otherwise. We recommend PMF users to try  
39 running their analyses on selected sub-ranges in order to further explore their datasets.

40

## 41 **1 Introduction**

42 Huge amounts of volatile organic compounds (VOC) are emitted to the atmosphere every year  
43 (Guenther et al., 1995;Lamarque et al., 2010), which play a significant role in atmospheric chemistry  
44 and affect the oxidative ability of the atmosphere. The oxidation products of VOC can contribute to  
45 the formation and growth of secondary organic aerosols (Kulmala et al., 2013;Ehn et al., 2014;Kirkby  
46 et al., 2016;Troestl et al., 2016), affecting air quality, human health, and climate radiative forcing  
47 (Pope III et al., 2009;Stocker et al., 2013;Zhang et al., 2016;Shiraiwa et al., 2017). Thanks to the  
48 advancement in mass spectrometric applications, like the aerosol mass spectrometer (AMS)  
49 (Canagaratna et al., 2007) and chemical ionization mass spectrometry (CIMS) (Bertram et al.,  
50 2011;Jokinen et al., 2012;Lee et al., 2014) our capability to detect these oxidized products, as well as  
51 our understanding of the complicated atmospheric oxidation pathways in which they take part, have  
52 been greatly enhanced.

53 Monoterpenes ( $\text{C}_{10}\text{H}_{16}$ ), one common group of VOC emitted in forested areas, have been shown to  
54 be a large source of atmospheric secondary organic aerosol (SOA). The oxidation of monoterpenes  
55 produces a wealth of different oxidation products (Oxygenated VOC, OVOC), including highly  
56 oxygenated organic molecules (HOM) with molar yields in the range of a few percent, depending on  
57 the specific monoterpene and oxidant (Ehn et al., 2014;Bianchi et al., 2019). Bianchi et al. (2019)  
58 summarized that HOM can be either Extremely Low Volatility Organic Compounds (ELVOC), Low  
59 Volatility Organic Compounds (LVOC), or Semi-volatile Organic Compounds (SVOC)  
60 (classifications by Donahue et al. 2012), depending on their exact structures. For less oxygenated  
61 products, the majority are likely to fall into the SVOC or the Intermediate VOC (IVOC) range. The  
62 volatility of the OVOC will determine their dynamics, including their ability to contribute to the  
63 formation of SOA and new particles (Bianchi et al., 2019;Buchholz et al., 2019).

64 The recent developments of CIMS techniques has allowed researchers to observe unprecedented  
65 numbers of OVOC, in real-time (Riva et al., 2019). This ability to measure thousands of compounds  
66 is a great benefit, but also a large challenge for the data analyst. For this reason, factor analytical  
67 techniques have often been applied to reduce the complexity of the data by finding co-varying signals



68 that can be grouped into common factors (Huang et al., 1999). For aerosol and gas-phase mass  
69 spectrometry, positive matrix factorization, PMF (Paatero and Tapper, 1994;Zhang et al., 2011) has  
70 been the most utilized tool. The factors have then been attributed to sources (e.g. biomass burning  
71 organic aerosol) or processes (e.g. monoterpene ozonolysis) depending on the application and ability  
72 to identify spectral signatures (Yan et al., 2016;Zhang et al., 2017). In the vast majority of these PMF  
73 applications to mass spectra, the mass range of ions has been maximized in order to provide as much  
74 input as possible for the algorithm. This approach was certainly motivated in early application of  
75 PMF on e.g. offline filters, with chemical information of metals, water-soluble ions, and organic and  
76 elemental carbon (OC and EC), where the number of variables is counted in tens, and the number of  
77 samples in tens or hundreds (Zhang et al., 2017). However, with gas-phase CIMS, we often have up  
78 to a thousand variables, with hundreds or even thousands of samples, meaning that the amount of data  
79 itself is unlikely to be a limitation for PMF calculation. In this work, we aimed to explore potential  
80 benefits of dividing the spectra into sub-ranges before applying factorization analysis.

81 An inherent requirement of factorization approaches is that the factor profiles, in this case the relative  
82 abundancies of ions in the mass spectra, of each factor stay nearly constant. Due to the complexity  
83 and number of atmospheric processes affecting the formation, transformation, and loss of VOC,  
84 OVOC and aerosol, this often does not hold, and is one of the main limitations of factorization  
85 approaches. Given the different volatilities of OVOC, it may even be expected that molecules from  
86 the same source may have very different loss time scales, which may affect the factor analysis.  
87 Volatility issue has been studied and reported for AMS data, with different volatilities of various OA  
88 types (Huffman et al., 2009;Crippa et al., 2014;Paciga et al., 2016;Äijälä et al., 2017). Semi-volatile  
89 oxygenated organic aerosol (SV-OOA) and Low-volatility oxygenated organic aerosol (LV-OOA)  
90 can both be mainly produced from biogenic sources, but get separated based on different volatilities  
91 by PMF (El Haddad et al., 2013). Sekimoto et al. (2018) found that the two profiles resolved with  
92 VOC emitted from biomass burning had different estimated volatilities. As the volatility of a molecule  
93 is linked to its molecular mass (Peräkylä et al., 2019), it may be beneficial to apply PMF separately  
94 to mass ranges where one can expect the loss processes to be similar, thereby resulting in more  
95 constant factor profiles. In this way, distinct sources are hopefully separated by PMF, with minimized  
96 influence of differing volatilities from one source.

97 The number of PMF or other factorization studies utilizing CIMS data remains very limited.  
98 “Traditional” PMF analyses have so far, to our knowledge, only been applied to nitrate-based  
99 chemical ionization atmospheric pressure interface time-of-flight (CI-APi-TOF) data (Yan et al.,  
100 2016;Massoli et al., 2018). One study has also utilized non-negative matrix factorization (NNMF) to  
101 look at diurnal trends of Iodide ToF-CIMS data (Lee et al., 2018). The lack of more studies utilizing



102 PMF, or other factorization techniques, on CIMS data is likely partly due to the complexity of the  
103 data, with multiple overlapping ions hampering HR peak fitting (Zhang et al., 2019). In addition,  
104 variable factor profiles may hamper PMF's ability to correctly separate the factors. The two CI-API-  
105 TOF studies utilized the nearly the entire measured spectrum (from around 200 Th to around 600 Th),  
106 either in unit mass resolution (UMR) or high resolution (HR) peak fitting data (Yan et al.,  
107 2016;Massoli et al., 2018). Massoli et al. (2018) estimated the volatility of the molecules they detected,  
108 finding that all the six extracted factors had notable contributions from IVOC, SVOC and (E)LVOC.  
109 These compound groups will have clearly different loss mechanisms, and thereby loss rates,  
110 suggesting that variation in factor profiles is inevitable, even if the source was identical for all  
111 molecules in the factor. We hypothesize that this effect further hampers the correct factorization, and  
112 further that this effect can be reduced by dividing the spectra into separate ranges, with each sub-  
113 range containing molecules with roughly similar loss mechanisms and rates.

114 As an additional motivation to separate different ranges from the mass spectrum, it is not only the  
115 loss mechanisms, but also the formation pathways that may differ. For example, atmospheric  
116 oxidation chemistry of organics is, to a large extent, the chemistry of peroxy radicals ( $\text{RO}_2$ ) (Orlando  
117 and Tyndall, 2012). These  $\text{RO}_2$  are initiated by VOC reacting with oxidants like ozone, or the hydroxyl  
118 ( $\text{OH}$ ) or nitrate ( $\text{NO}_3$ ) radicals, while their termination occurs mainly by bimolecular reactions with  
119  $\text{NO}$ ,  $\text{HO}_2$  and/or other  $\text{RO}_2$ . Some product molecules can be formed from many of the three  
120 termination pathways, while for example ROOR "dimers" can only be formed from  $\text{RO}_2+\text{RO}_2$   
121 reactions (Berndt et al., 2018a;Berndt et al., 2018b). This also means that there can be six different  
122 pathways to form dimers from the same precursors VOC, by combining  $\text{RO}_2$  formed from the same  
123 or different oxidants. As an example of the latter, an ROOR dimer can contain one moiety produced  
124 from ozone oxidation and another moiety from  $\text{NO}_3$  oxidation (Yan et al., 2016). Thus, their  
125 concentration is dependent on both the precursor VOC concentration, and the concentrations of both  
126 oxidants. Such a molecule will not have a direct equivalent in any of the monomer products,  
127 dependent on only one oxidant, which again may complicate the separation of such factors by PMF,  
128 if the entire spectrum is analyzed once. However, if separating the monomer and dimer products  
129 before PMF analysis, separation of different formation pathways can potentially become simpler.

130 Recently, we proposed a new PMF approach, binPMF, to simplify the analysis of mass spectral data  
131 (Zhang et al., 2019). This method divides the mass spectrum into narrow bins, typically some tens of  
132 bins per integer mass, depending on the mass resolving power of the instrument, before performing  
133 PMF analyses. In this way, binPMF does not require any time-consuming, and potentially subjective  
134 high resolution peak fitting, and can thus be utilized for data exploration at a very early stage of data  
135 analysis. Data preparation is nearly as simple as in the case of UMR analysis, yet it utilizes much



136 more spectral information as it does not sum up signal over all ions at each integer mass. In addition  
137 to saving time and effort in data analysis, the results are less sensitive to mass calibration fluctuations.  
138 Finally, the binning also greatly increases the number of input variables, which has the advantage that  
139 factor analysis with smaller mass ranges becomes more feasible, as more meaningful variation is  
140 supplied to the algorithm.

141 We designed this study to explore the benefits of separate analysis of different mass ranges from mass  
142 spectra. We used a previously published ambient dataset measured by a CI-API-TOF, and conducted  
143 binPMF analysis with three different mass ranges, roughly corresponding to different volatility ranges.  
144 We compared the results from the sub-range analyses with each other and with results from binPMF  
145 run on the combined ranges. We found that more meaningful factors are separated from our dataset  
146 by utilizing the sub-ranges, and believe that this study will provide new perspectives for future studies  
147 analyzing gas-phase CIMS data.

148

## 149 **2 Methodology**

150 The focus of this work is on retrieving new information from mass spectra by applying new analytical  
151 approaches. Therefore, we chose a dataset that has been presented earlier, though without PMF  
152 analysis, by Zha et al. (2018), and was also used in the first study describing the binPMF method  
153 (Zhang et al., 2019). The measurements are described in more details below in section 2.1, while the  
154 data analysis techniques used in this work are presented in section 2.2.

### 155 **2.1 Measurements**

#### 156 **2.1.1 Ambient site**

157 The ambient measurements were conducted at the Station for Measuring Ecosystem–Atmosphere  
158 Relations (SMEAR) II in Finland (Hari and Kulmala, 2005) as part of the Influence of Biosphere-  
159 Atmosphere Interactions on the Reactive Nitrogen budget (IBAIRN) campaign (Zha et al, 2018).  
160 Located in the boreal environment in Hyytiälä, SMEAR II is surrounded with coniferous forest and  
161 has limited anthropogenic emission sources nearby. Diverse measurements of meteorology, aerosol  
162 and gas phase properties are continuously conducted at the station. Details about the meteorological  
163 conditions and temporal variations of trace gases during IBAIRN campaign are presented by Zha et  
164 al. (2018) and Liebmann et al. (2018).

#### 165 **2.1.2 Instrument and data**

166 Data were collected with a nitrate ( $\text{NO}_3^-$ )-based chemical ionization atmospheric pressure interface  
167 time-of-flight mass spectrometer (CI-API-TOF, Jokinen et al., 2012) with about 4000 Th  $\text{Th}^{-1}$  mass  
168 resolving power, at ground level in September, 2016. In our study, the mass spectra were averaged to  
169 1 h time resolution from September 6<sup>th</sup> to 22<sup>nd</sup> for further analysis. We use the thomson (Th) as the



170 unit for mass/charge, with  $1 \text{ Th} = 1 \text{ Da}/e$ , where  $e$  is the elementary charge. As all the data discussed  
171 in this work are based on negative ion mass spectrometry, we will use the absolute value of the  
172 mass/charge, although the charge of each ion will be negative. The masses discussed in this work  
173 includes the contribution from the nitrate ion, 62, unless specifically mentioned. Furthermore, as the  
174 technique is based on soft ionization with  $\text{NO}_3^-$  ions, any multiple charging effects are unlikely, and  
175 therefore the reported mass/charge values in thomson can be considered equivalent to the mass of the  
176 ion in Da.

177 The forest site of Hyytiälä is dominated by monoterpene emissions (Hakola et al., 2006). The main  
178 feature of previous CI-API-TOF measurements in Hyytiälä (Ehn et al., 2014; Yan et al., 2016) has  
179 been a bimodal distributions of HOM, termed monomers and dimers, as they are formed of either one  
180 or two  $\text{RO}_2$  radicals, respectively. For the analysis in this study, we chose three mass/charge ( $m/z$ )  
181 ranges of 50 Th each (Figure 1), corresponding to regions between which we expect differences in  
182 formation or loss mechanisms. In addition to regions with HOM monomers and HOM dimers, one  
183 range was chosen at lower masses, in a region presumably mainly consisting of molecules that are  
184 less likely to condense onto aerosol particles (Peräkylä et al., 2019).

## 185 2.2 Positive matrix factorization (PMF)

186 After the model of PMF was developed (Paatero and Tapper, 1994), numerous applications have been  
187 conducted with different types of environmental data (Song et al., 2007; Ulbrich et al., 2009; Yan et  
188 al., 2016; Zhang et al., 2017). By reducing dimensionality of the measured dataset, PMF model greatly  
189 simplifies the data analysis process with no requirement for prior knowledge of sources or pathways  
190 as essential input. The main factors can be further interpreted with their unique/dominant markers  
191 (elements or masses).

192 The basic assumption for PMF modelling is mass balance, which assumes that ambient concentration  
193 of a chemical component is the sum of contributions from several sources or processes, as shown in  
194 equation (1).

$$195 \quad \mathbf{X} = \mathbf{TS} \times \mathbf{MS} + \mathbf{R} \quad (1)$$

196 In equation (1),  $\mathbf{X}$  stands for the time series of measured concentration of different variables ( $m/z$   
197 in our case),  $\mathbf{TS}$  represents the temporal variation of factor contributions,  $\mathbf{MS}$  stands for factor profiles  
198 (mass spectral profiles), and  $\mathbf{R}$  is the residual as the difference of the modelled and the observed data.  
199 The matrices  $\mathbf{TS}$  and  $\mathbf{MS}$  are iteratively calculated by a least-squares algorithm utilizing uncertainty  
200 estimates, to pursue minimized  $Q$  value as shown in equation (2), where  $S_{ij}$  is the estimated  
201 uncertainty, an essential input in PMF model.

$$202 \quad Q = \sum \sum \left( \frac{R_{ij}}{S_{ij}} \right)^2 \quad (2)$$



203 PMF model was conducted by multi-linear engine (ME-2) (Paatero, 1999) interfaced with Source  
204 Finder (SoFi, v6.3) (Canonaco et al., 2013). Signal-to-noise ratio (SNR) was calculated as  $SNR_{ij} =$   
205  $abs(X_{ij}) / abs(S_{ij})$ . When the Signal-to-noise ratio (SNR) is below 1, the signal of  $X_{ij}$  will be down-  
206 weighted by replacing the corresponding uncertainty  $S_{ij}$  by  $S_{ij}/SNR_{ij}$  (Visser et al., 2015). Future  
207 studies should pay attention to the potential risk when utilizing this method since down-weighting  
208 low signals element-wise will create a positive bias to the data. Robust mode was operated in the  
209 PMF modelling, where outliers ( $\left| \frac{R_{ij}}{S_{ij}} \right| > 4$ ) were significantly down-weighted (Paatero, 1997).

### 210 2.3 binPMF

211 As a newly developed application of PMF for mass spectral data, binPMF has no requirement for  
212 chemical composition information, while still taking advantage of the HR mass spectra, saving effort  
213 and time (Zhang et al., 2019). To explore the benefits of analyzing separated mass ranges, we applied  
214 binPMF to the three separated ranges. The three ranges were also later combined for binPMF analysis  
215 as comparison with the previous results. The PMF model requires both data matrix and error matrix  
216 as input, and details of the preparation of data and error matrices are described below.

#### 217 2.3.1 Data matrix

218 Different from normal UMR or HR peak fitting, in binPMF, the mass spectra are divided into small  
219 bins after baseline subtraction and mass axis calibration. Linear interpolation was first conducted to  
220 the mass spectra with a mass interval of 0.001 Th. Then the interpolated data was averaged into bins  
221 of 0.02 Th width. We selected three ranges for further analysis based on earlier studies (Ehn et al.,  
222 2014; Yan et al., 2016; Bianchi et al., 2019; Peräkylä et al., 2019).

- 223 - Range 1,  $m/z$  250 – 300 Th, 51 unit masses  $\times$  25 bins per unit mass = 1275 bins/variables,  
224 consisting mainly of molecules with five to nine carbon atoms and four to nine oxygen atoms  
225 in our dataset.
- 226 - Range 2,  $m/z$  300 – 350 Th,  $51 \times 25 = 1275$  bins, mainly corresponding to HOM monomer  
227 products, featured with nine to ten C- and seven to ten O-atoms.
- 228 - Range 3,  $m/z$  510 – 560 Th,  $51 \times 30 = 1530$  bins, mainly corresponding to HOM dimer products,  
229 with carbon numbers of sixteen to twenty and eleven to fifteen O-atoms.

230 For a nominal mass  $N$ , the signal region included in further analyses was between  $N-0.2$  Th and  $N+0.3$   
231 Th for Range 1 and 2, and between  $N-0.2$  Th and  $N+0.4$  Th for Range 3. The data were averaged into  
232 1-h time resolution and in total we had 384 time points in the data matrix.

#### 233 2.3.2 Error matrix



234 The error matrix represents the estimated uncertainty for each element of the data matrix and is crucial  
235 for iterative calculation of the  $Q$  minimum. Equation (3) is used for error estimation (Polissar et al.,  
236 1998),

$$237 \quad S_{ij} = \sigma_{ij} + \sigma_{\text{noise}} \quad (3)$$

238 where  $S_{ij}$  represents the uncertainty of  $m/z$   $j$  at time  $i$ ,  $\sigma_{ij}$  stands for counting statistics uncertainty  
239 and is estimated as follows,

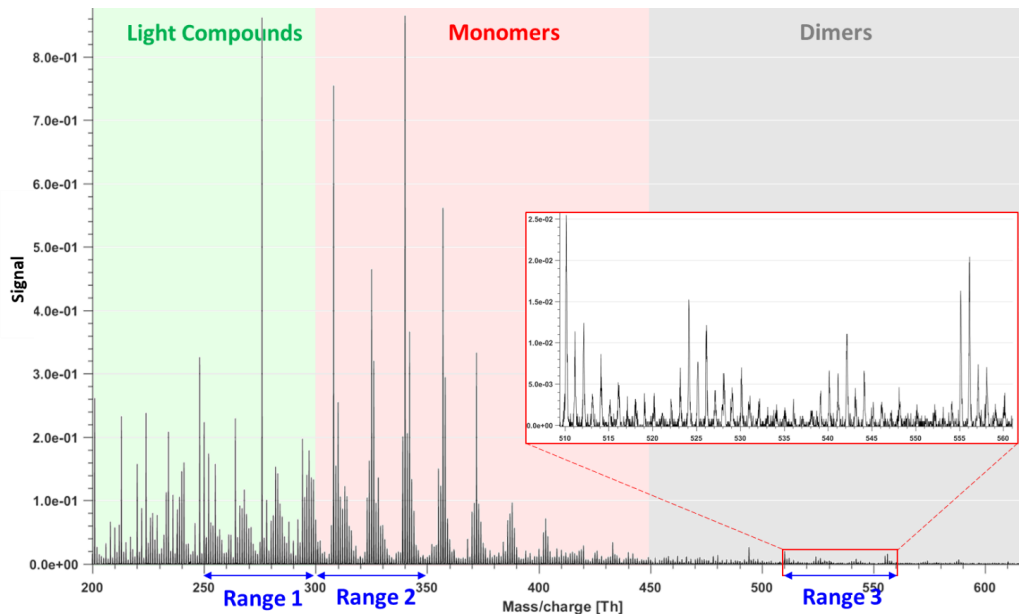
$$240 \quad \sigma_{ij} = a \times \frac{\sqrt{I_{ij}}}{\sqrt{t_s}} \quad (4)$$

241 where  $I$  is the signal intensity term, in unit of counts per second (cps),  $t_s$  stands for length of averaging  
242 in seconds, while  $a$  is an empirical coefficient to compensate for unaccounted uncertainties (Allan et  
243 al., 2003; Yan et al., 2016) and is 1.28 in our study as previously estimated from laboratory  
244 experiments (Yan et al., 2016). The  $\sigma_{\text{noise}}$  term was estimated as the median of the standard  
245 deviations from signals in the bins in the region between nominal masses, where no physically  
246 meaningful signals are expected.

## 247 **3 Results**

### 248 3.1 General overview of the dataset/spectrum

249 During the campaign, in autumn, 2016, the weather was overall sunny and humid with average  
250 temperature of 10.8 °C and relative humidity (RH) of 87% (Zha et al., 2019). The average  
251 concentration of  $\text{NO}_x$  and  $\text{O}_3$  was 0.4 ppbv and 21 ppbv, respectively. The average total HOM  
252 concentration was  $\sim 10^8$  molecules  $\text{cm}^{-3}$ .



253





254 Figure 1. Example of mass spectrum with 1-h time resolution measured from a boreal forest  
255 environment during the IBairn campaign (at 18:00, Finnish local time, UTC+2). The mass  
256 spectrum was divided into three parts and three sub-ranges were chosen from different parts for  
257 further analysis in our study. The nitrate ion (62 Th) is included in the mass.  
258 Figure 1 shows the 1 h averaged mass spectrum taken at 18:00 on September 12, as an example of  
259 the analyzed dataset. In addition to exploring the benefits of this type of sub-range analysis in relation  
260 to different formation or loss pathways, separating into sub-ranges may also aid factor identification  
261 for low-signal regions. As shown in Figure 1, there is a difference of 1-2 orders of magnitude in the  
262 signal intensity between Range 3 and Ranges 1-2. If all Ranges are run together, we would expect  
263 that the higher signals from Ranges 1 and 2 will drive the factorization. While if run separately,  
264 separating formation pathways of dimers in Range 3 will likely be easier. As dimers have been shown  
265 to be crucial for the formation of new aerosol particles from monoterpene oxidation (Kirkby et al.,  
266 2016; Troestl et al., 2016; Lehtipalo et al., 2018), this information may even be the most critical in  
267 some cases, despite the low contribution of these peaks to the total measured signal.  
268 binPMF was separately applied to Range 1, 2, 3, and a ‘Range combined’ which comprised all the  
269 three sub-ranges. All the PMF runs for the four ranges were conducted from two to ten factors and  
270 repeated three times for each factor number, to assure the consistency of the results. Factorization  
271 results and evolution with increasing factor number are briefly described in the following sections,  
272 separately for each Range (sections 3.2 – 3.5). More detailed discussion and comparison between the  
273 results are presented in Section 4.

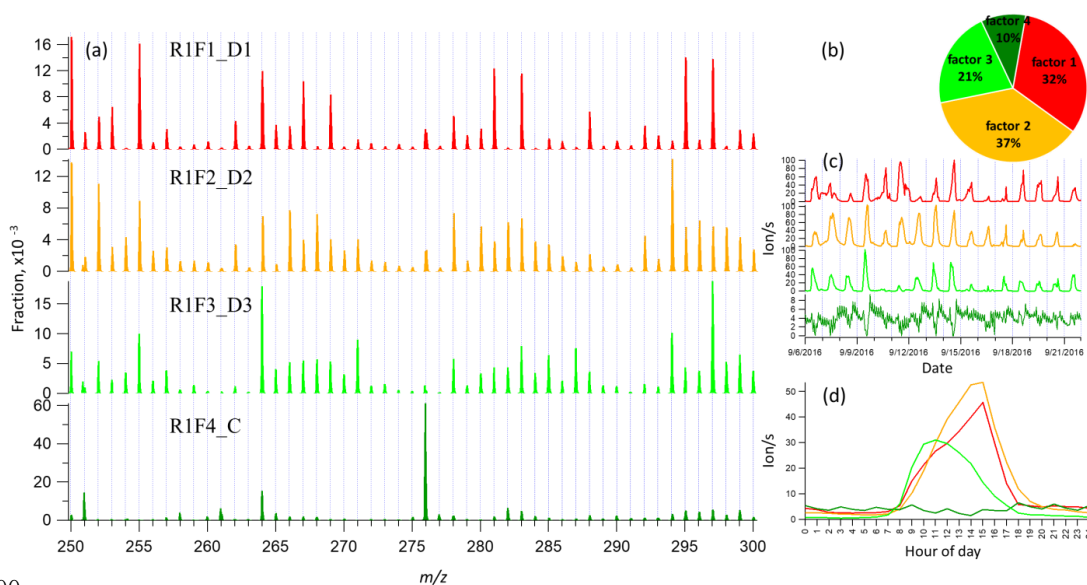
### 274 3.2 binPMF on Range 1 (250 – 300 Th)

275 As has become routine (Zhang et al., 2011; Craven et al., 2012), we first examined the mathematical  
276 parameters of our solutions. From two to ten factors,  $Q/Q_{exp}$  decreased from 2.8 to 0.7 (Fig S1 in  
277 supplementary information), and after three factors, the decreasing trend was gradually slowing down  
278 and approaching one, which is the ideal value for  $Q/Q_{exp}$  as a diagnostic parameter. The unexplained  
279 variation showed a decline from 18% to 8% from two to ten factors.

280 In the two-factor results, two daytime factors were separated, with peak time both at 14:00 - 15:00.  
281 One factor was characterized by large signals at 250 Th, 255 Th, 264 Th, 281 Th, 283 Th, 295 Th,  
282 297 Th. The other factor was characterized by large signals at 294 Th, 250 Th, 252 Th, 264 Th, 266  
283 Th, 268 Th, and 297 Th. In Hyytiälä, as reported in previous studies, odd masses observed by the  
284 nitrate CI-API-TOF are generally linked to monoterpene-derived organonitrates during the day (Ehn  
285 et al., 2014; Yan et al., 2016). When the factor number increased to three, the two earlier daytime  
286 factors remained similar with the previous result, while a new factor appeared with a distinct sawtooth  
287 shape in the diurnal cycle. The main marker in the spectral profile was 276 Th, with a clear negative



288 mass defect. When one more factor was added, the previous three factors remained similar as in the  
289 three-factor solution, and a new morning factor was resolved, with 264 Th and 297 Th dominant in  
290 the mass spectral profile, and a diurnal peak at 11:00.  
291 As the factor number was increased, more daytime factors were separated, with similar spectral  
292 profiles to existing daytime factors and various peak times. No nighttime factors were found in the  
293 analysis even when the factor number reached ten. We chose the four-factor result for further  
294 discussion, and Figure 2 shows the result of Range 1, with spectral profile, time series, diurnal cycle  
295 and factor contribution. As shown in Figure 2d, Factors 1-3 are all daytime factors, while Factor 4  
296 has a sawtooth shape, which is caused by contamination, mainly by perfluorinated acids, of the inlet's  
297 automated zeroing every three hours during the measurements (Zhang et al., 2019). The zeroing  
298 periods have been removed from the dataset before binPMF analysis, but the contamination factor  
299 was still resolved. This factor is discussed in more detail in sections 4.1.1 and 4.1.4.



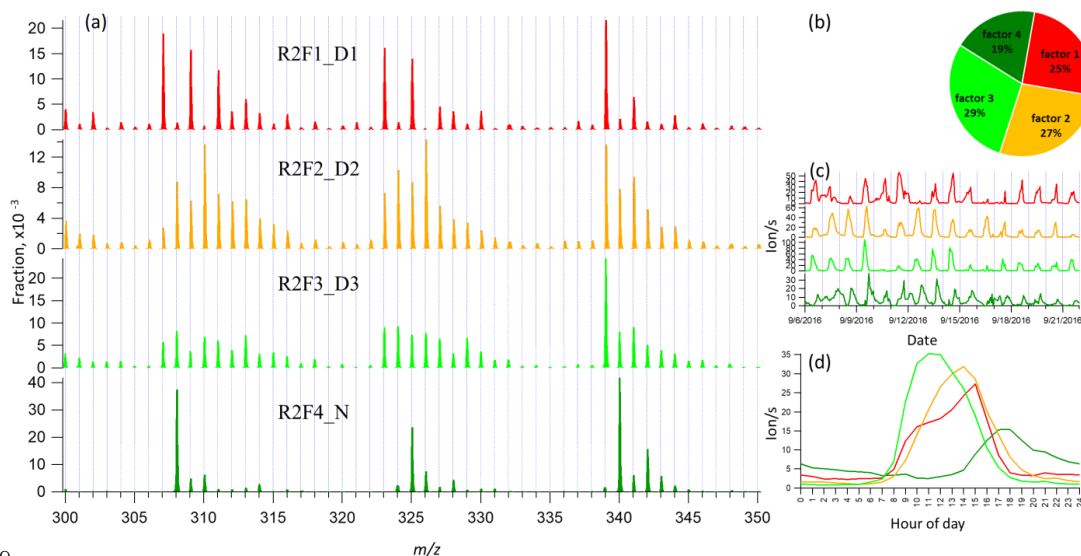
300  
301 Figure 2 Four-factor result for Range 1, for (a) factor spectral profiles, (b) factor contribution, (c)  
302 time series and (d) diurnal trend. Details on the factors' naming schemes are shown in Table 1.

### 303 3.3 binPMF on Range 2 (300-350 Th)

304 This range covers the monoterpene HOM monomer range, and binPMF results have already been  
305 discussed by Zhang et al. (2019) as a first example of the application of binPMF on ambient data.  
306 Our input data here is slightly different. In the previous study, the 10 min automatic zeroing every  
307 three hours was not removed before averaging to 1 hour time resolution while here, we have removed  
308 this data. Overall, the results are similar as in our earlier study, and therefore the result are just briefly



309 summarized below for further comparison and discussion in Section 4. Similar to Range 1, both the  
310  $Q/Q_{exp}$  (2.2 to 0.6) and unexplained variation (16% to 8%) declined with the increased factor number  
311 from two to ten.  
312 When the factor number was two, one daytime and one nighttime factor were separated, with diurnal  
313 peak times at 14:00 and 17:00, respectively. The nighttime factor was characterized by masses at 340  
314 Th, 308 Th and 325 Th (monoterpene ozonolysis HOM monomers (Ehn et al., 2014)) and remained  
315 stable throughout the factor evolution from two to ten factors. With the addition of more factors, no  
316 more nighttime factors got separated while the daytime factor was further separated and more daytime  
317 factors appeared, peaking at various times in the morning (10:00 am), at noon or in the early afternoon  
318 (around 14:00 pm and 15:00 pm). High contribution of 339 Th can be found in all the daytime factor  
319 profiles. As the factor number reached six, a contamination factor appeared, characterized by large  
320 signals at 339 Th and 324 Th, showing negative mass defects (Figure S2 in the Supplement). The  
321 factor profile is nearly identical to the contamination factor determined in Zhang et al. (2019), where  
322 the zeroing periods were not removed, causing larger signals for the contaminants. In our dataset,  
323 where the zeroing periods were removed, no sawtooth pattern was discernible in the diurnal trend,  
324 yet it could still be separated even though it only contributed 3% to Range 2. More about the  
325 contamination factors from different ranges will be discussed in Section 4.1.4. Since the aim of this  
326 study is mainly to explore the benefits of analyzing different ranges of the mass spectrum, we chose  
327 to show the four-factor result below, to simplify the later discussion and comparison. Figure 3 shows  
328 four-factor result of Range 2, with spectral profile, time series, diurnal cycle and factor contribution.



329

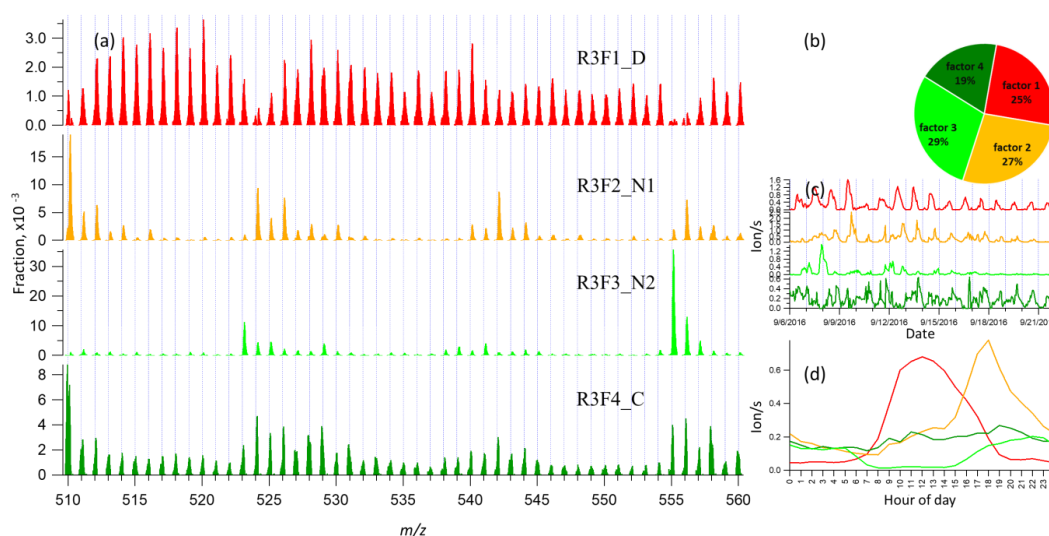


330 Figure 3 Four-factor result for Range 2, for (a) factor spectral profiles, (b) factor contribution, (c)  
331 time series and (d) diurnal trend. Details on the factors' naming schemes are shown in Table 1.

### 332 3.4 binPMF on Range 3 (510-560 Th)

333 Range 3 represents mainly the monoterpene HOM dimers (Ehn et al., 2014). Similar to Range 1 and  
334 2, both the  $Q/Q_{exp}$  (1.5 to 0.6) and unexplained variation (18% to 15%) showed decreasing trend with  
335 the increased factor number (2-10). As can be seen from Figure 1, data in Range 3 had much lower  
336 signals, compared to that of the Range 1 and 2, explaining the higher unexplained variation for Range  
337 3.

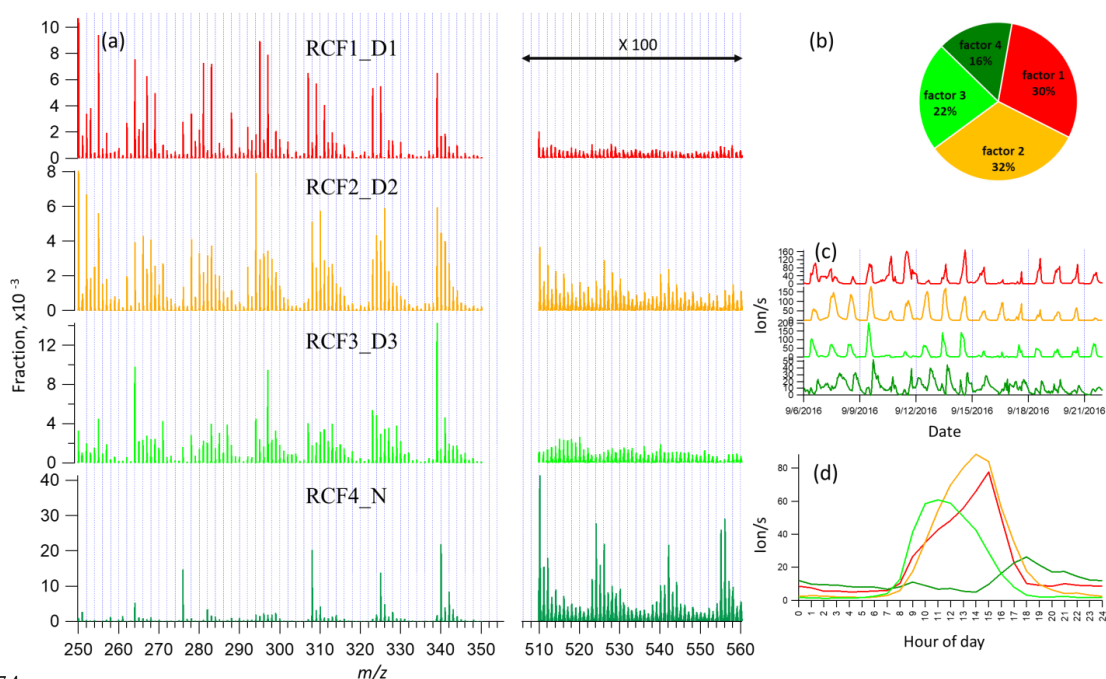
338 In the two-factor result for Range 3, one daytime and one nighttime factor appeared, with diurnal  
339 peak times at noon and 18:00, respectively. The nighttime factor was characterized by ions at 510 Th,  
340 524 Th, 526 Th, 542 Th, and 555 Th, 556 Th, while the daytime factor showed no dominant marker  
341 masses, yet with relatively high signals at 516 Th, 518 Th and 520 Th. As the number of factors  
342 increased to three, one factor with almost flat diurnal trend was separated, with dominant masses of  
343 510 Th, 529 Th, 558 Th. Most peaks in this factor had negative mass defects, and this factor was  
344 again linked to a contamination factor. The four-factor result resolved another nighttime factor with  
345 a dominant peak at 555 Th, and effectively zero contribution during daytime. As the factor number  
346 was further increased, the new factors seemed like splits from previous factors with similar spectral  
347 profiles. We therefore chose four-factor result also for Range 3 (results shown in Fig. 4) for further  
348 discussion.



349  
350 Figure 4 Four-factor result for Range 3, for (a) factor spectral profiles, (b) factor contribution, (c)  
351 time series and (d) diurnal trend. Details on the factors' naming schemes are shown in Table 1.



352 3.5 binPMF on Range Combined (250-350 Th & 510-560 Th)  
353 As comparison to the previous three ranges, we conducted the binPMF analysis on Range Combined,  
354 which is the combination of the three ranges. The results of this range are fairly similar to those of  
355 Ranges 1 and 2, as could be expected since the signal intensities in these ranges were much higher  
356 than in Range 3. As the number of factors increased (2-10), both the  $Q/Q_{exp}$  (1.3 to 0.6) and  
357 unexplained variation (16% to 8%) showed a decreasing trend.  
358 In the two-factor result, one daytime factor and one nighttime factor were separated. In the nighttime  
359 factor, most masses were found at even masses, and the fraction of masses in Range 3 was much  
360 higher than that in daytime factor. In contrast, in the daytime factor, most masses were observed at  
361 odd masses and the fraction of signal in Range 3 was much lower. During the day, photochemical  
362 reactions increase the concentration of NO, which serves as peroxy radical (RO<sub>2</sub>) terminator and often  
363 outcompetes RO<sub>2</sub> cross reactions in which dimers can be formed (Ehn et al., 2014). Thus, the  
364 production of dimers is suppressed during the day, yielding instead a larger fraction of organic nitrates,  
365 as has been shown also previously (Yan et al., 2016).  
366 With the increase of the number of factors, more daytime factors were resolved with different peak  
367 times. When the factor number reached seven, a clear sawtooth-shape diurnal cycle occurred, i.e. the  
368 contamination factor, caused by the zeroing. As more factors were added, no further nighttime factors  
369 were separated, and only more daytime factors appeared. To simplify the discussion and inter-range  
370 comparison, we also here chose the four-factor result for further analysis, as it already provided  
371 enough information for our main goal in this study. Figure 5 shows the four-factor result of Range  
372 Combined, with spectral profile, time series, diurnal cycle and factor contribution. The signals in  
373 range of 510-560 Th were enlarged 100-fold to be visible.



374

375 Figure 5 Four-factor result for Range Combined, for (a) factor spectral profiles, (b) factor  
376 contribution, (c) time series and (d) diurnal trend. Details on the factors' naming schemes are shown  
377 in Table 1.

#### 378 4 Discussion

379 In Section 3, results by binPMF analysis were shown for Ranges 1, 2, 3 and Combined. In this section,  
380 we discuss and compare the results from the different ranges. To simplify the inter-range comparison,  
381 we chose four-factor results for all the four ranges, with the abbreviations shown in Table 1. From  
382 Range 1, three daytime factors and a contaminations factor were separated. In Range 2, three daytime  
383 factors and one nighttime factor (abbreviated as R2F4\_N) were resolved. The R2F4\_N factor was  
384 characterized by signals at 308 Th ( $C_{10}H_{14}O_7 \cdot NO_3^-$ ), 325 Th ( $C_{10}H_{15}O_8 \cdot NO_3^-$ ), and 340 Th  
385 ( $C_{10}H_{14}O_9 \cdot NO_3^-$ ), and can be confirmed as monoterpene ozonolysis products (Ehn et al., 2014; Yan et  
386 al., 2016). With the increase of factor number to six, the contamination factor got separated also in  
387 this mass range. In Range 3, one daytime factor, two nighttime factors and a contamination factor  
388 were separated. The first nighttime factor (R3F2\_N1) had large peaks at 510 Th ( $C_{20}H_{32}O_{11} \cdot NO_3^-$ )  
389 and 556 Th ( $C_{20}H_{30}O_{14} \cdot NO_3^-$ ), dimer products that have been identified during chamber studies of  
390 monoterpene ozonolysis (Ehn et al., 2014). The molecule observed at 510 Th has 32 H-atoms,  
391 suggesting that one of the RO<sub>2</sub> involved would have been initiated by OH, which is formed during  
392 the ozonolysis of alkenes such as monoterpenes at nighttime (Atkinson et al., 1992; Paulson and



393 Orlando, 1996). The other nighttime factor (R3F3\_N2) was dominated by ions at 523 Th  
394 ( $C_{20}H_{31}O_8NO_3 \cdot NO_3^-$ ) and 555 Th ( $C_{20}H_{31}O_{10}NO_3 \cdot NO_3^-$ ), representing nighttime monoterpene  
395 oxidation involving  $NO_3$ . As these dimers contain only one N-atom, and 31 H-atoms, we can assume  
396 that they are formed from reactions between an  $RO_2$  formed from  $NO_3$  oxidation and another  $RO_2$   
397 formed by ozone oxidation. These results match well with the profiles in a previous study by Yan et  
398 al. (2016). The results of Range Combined are very similar to Range 2, with one nighttime factor and  
399 three daytime factors. The contamination factor was separated with increase of factor number to seven.  
400

401

Table 1. Summary of PMF results for the different mass ranges

Range	Factor number	Factor name <sup>a</sup>	Dominant peaks	Peak time
<b>1 (250 - 300 Th)</b>	1	R1F1_D1	250, 255, 295, 297	15:00
	2	R1F2_D2	250, 252, 294	15:00
	3	R1F3_D3	264, 297	11:00
	4	R1F4_C	276	- <sup>b</sup>
<b>2 (300 - 350 Th)</b>	1	R2F1_D1	307, 309, 323, 325, 339,	15:00
	2	R2F2_D2	310, 326, 339,	14:00
	3	R2F3_D3	339	11:00
	4	R2F4_N	308, 325, 340	18:00
<b>3 (510 – 560 Th)</b>	1	R3F1_D	516, 518, 520, 528, 540	12:00
	2	R3F2_N1	510, 524, 542, 556	18:00
	3	R3F3_N2	523, 555	22:00
	4	R3F4_C	510, 558	- <sup>b</sup>
<b>Combined (1, 2, 3)</b>	1	RCF1_D1	250, 255, 295, 339	15:00
	2	RCF2_D2	250, 252, 294, 339	14:00
	3	RCF3_D3	264, 297, 339	11:00
	4	RCF4_N	308, 340, 510, 524, 555, 556	18:00

402 <sup>a</sup> Factor name is defined with range name, factor number and name. For example, RxFy represents Factor y in Range x.  
403 RC stands for Range Combined. For the factor name, D is short for daytime, N for Nighttime, C for contamination.

404 <sup>b</sup> The contamination factor in Range 1 shows sawtooth pattern; while in Range 3 shows no diurnal pattern.  
405

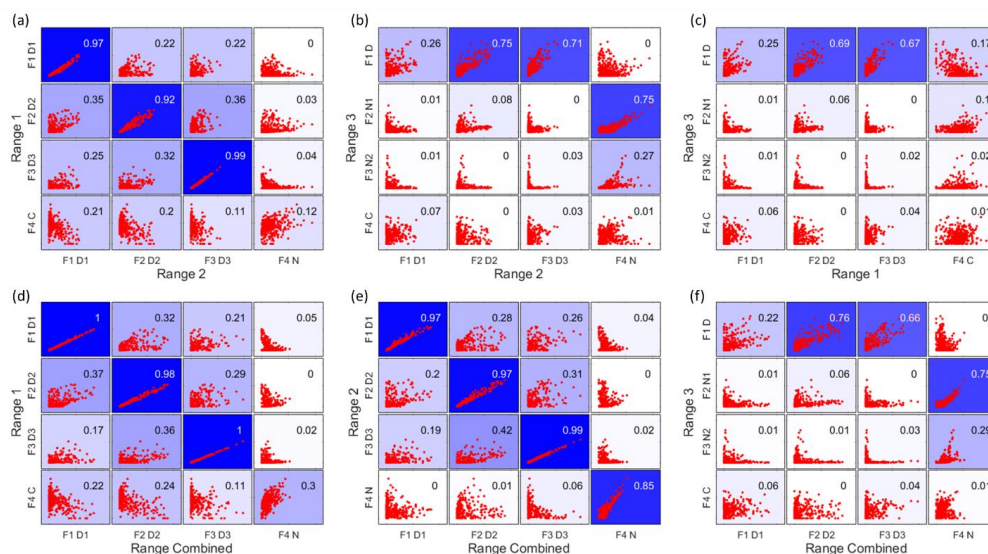
## 406 4.1 Comparison of different ranges

### 407 4.1.1 Time series correlation

408 In Figure 6, the upper panels show the time series correlations among the first three ranges. As  
409 expected based on the results above, generally the daytime factors, and the two nighttime  
410 monoterpene ozonolysis factors (R2F4\_N and R3F2\_N1) correlated well, respectively. However, the  
411 contamination factors did not show strong correlation between different ranges, even though  
412 undoubtedly from the same source. More about the contamination factors will be discussed in Section  
413 4.1.4. The lower panels in Figure 6 displays the correlations between the first three ranges and the  
414 Range Combined, and clearly demonstrates that the results of Range Combined is mainly controlled  
415 by high signals from Range 1 and 2. More detailed aspects of the comparison between factors in



416 different ranges is given in the following sections. The good agreements between factors from  
417 different subranges also help to verify the robustness of the solutions.



418  
419 Figure 6 Time series correlations among Range 1, 2, 3 (upper panels a-c), and between the first three  
420 ranges and the Range Combined (lower panels d-f). The abbreviations for different factors are the  
421 same in Table 1, with F for factor, D for daytime, N for nighttime and C for contamination, e.g. F1D1  
422 for Factor 1 daytime 1. The coefficient of determination,  $R^2$ , is marked in each subplot by a number  
423 shown in the right upper corners and by the blue colors, with stronger blue indicating higher  $R^2$ .

#### 424 4.1.2 Daytime factor comparison

425 As mentioned above, with increasing number of factors, usually more daytime factors will be resolved,  
426 reflecting the complicated daytime photochemistry. The three daytime factors between Range 1 and  
427 2 agreed with each other quite well (Figure 6a). However, R1F1\_D1 and R2F1\_D1 did not show  
428 strong correlation with the only daytime factor in Range 3 (R3F1\_D), while the other two daytime  
429 factors in both Range 1 and 2, i.e. R1F2\_D2, R1F3\_D3, R2F2\_D2, R2F3\_D3, correlated well with  
430 R3F1\_D from Range 3.

431 The 1<sup>st</sup> daytime factors from Range 1 and 2, R1F1\_D1 and R2F1\_D1, were mainly characterized by  
432 odd masses 255 Th, 281 Th, 283 Th, 295 Th, 297 Th, 307 Th, 309 Th, 311 Th, 323 Th, 325 Th, 339  
433 Th. The factors are dominated by organonitrates. Organic nitrate formation during daytime is  
434 generally associated with the termination of  $RO_2$  radicals by NO. This termination step is mutually  
435 exclusive with the termination of  $RO_2$  with other  $RO_2$ , leading to dimer formation. If the NO  
436 concentration is the limiting factor for the formation of these factors, the low correlations between  
437 the NO-terminated monomer factors, and the dimer factors, is to be expected. In contrast, if the other





438 daytime factors mainly depend on oxidant and monoterpene concentrations, some correlation  
439 between those, and the daytime dimer factor, is to be expected, as shown in Figure 6b, c.

440 All the spectral profiles resolved from Range Combined binPMF analysis inevitably contained mass  
441 contribution from 510 – 560 Th, even the daytime factor from Range Combined (RCF1\_D1) which  
442 did not show clear correlation with R3F1\_D from Range 3 (Figure 6e).

443 The 2<sup>nd</sup> and 3<sup>rd</sup> daytime factors in Range 1 and 2, R1F2\_D2, R1F3\_D3, R2F2\_D2, R2F3\_D3, had  
444 high correlations with R3F1\_D in Range 3. Daytime factors in Range Combined (RCF2\_D2 and  
445 RCF3\_D3) also showed good correlation with R3F1\_D in Range 3. However, if we compare R3F1\_D  
446 and the mass range of 510 – 560 Th of the daytime factors in Range Combined, just with a quick look,  
447 we can readily see the difference. The daytime factor separated in Range 3 (R3F1\_D) basically has  
448 no obvious markers in the profile, and as mentioned above, up to ten factors, there would only be  
449 more factors fragmented from the previous factor, with similar spectral profiles, but showed different  
450 profile pattern with 510 – 560 Th in RCF2\_D2 in Range Combined. The factorization of Range  
451 Combined was mainly controlled by Range 1 and 2 due to high signals, and the signals in Range 3  
452 are forced to be distributed according to the time series determined by Ranges 1 and 2. Ultimately,  
453 this will lead to failure in factor separation for this low-signal range.

#### 454 4.1.3 Nighttime factor comparison

455 Since high-mass dimers are more likely to form at night due to photochemical production of NO in  
456 daytime, which inhibits RO<sub>2</sub> + RO<sub>2</sub> reactions, Range 3 had the highest fraction of nighttime signals  
457 of all the sub-ranges. While Range 3 produced two nighttime factors, Ranges 2 and Combined showed  
458 one, and Range 1 had no nighttime factor. The difference between the two results also indicates the  
459 advantage of analyzing monomers and dimers separately.

460 The two nighttime factors in Range 3 can be clearly identified as arising from ozonolysis (R3F2\_N1)  
461 and a mix of ozonolysis and NO<sub>3</sub> oxidation (R3F2\_N2) based on the mass spectral profiles, as  
462 described above. The organonitrate at 555 Th, C<sub>20</sub>H<sub>31</sub>O<sub>10</sub>NO<sub>3</sub>·NO<sub>3</sub><sup>-</sup>, is a typical marker for NO<sub>3</sub>  
463 radical initiated monoterpene chemistry (Yan et al., 2016). However, several interesting features  
464 become evident when comparing to the results of Range 2 and Combined. Firstly, only one nighttime  
465 factor (R2F4\_N, RCF4\_N) was separated in each of these ranges, and that shows clear resemblance  
466 with ozonolysis of monoterpenes as measured in numerous studies, e.g. (Ehn et al., 2012; Ehn et al.,  
467 2014). Secondly, the high correlation found in Figure 6b between the ozonolysis factors (i.e.,  
468 R2F4\_N, R3F2\_N1, RCF4\_N), further supports the assignment. However, this factor is the only  
469 nighttime factor in the monomer range, suggesting that NO<sub>3</sub> radical chemistry of monoterpenes in  
470 Hyytiälä does not form substantial amounts of HOM monomers. The only way for the CI-APi-TOF  
471 to detect products of monoterpene-NO<sub>3</sub> radical chemistry may thus be through the dimers, where one



472 highly oxygenated RO<sub>2</sub> radical from ozonolysis reacts with a less oxygenated RO<sub>2</sub> radical from NO<sub>3</sub>  
473 oxidation.

474 In the results by Yan et al. (2016) the combined UMR-PMF of monomers and dimers did yield a  
475 considerable amount of compounds in the monomer range also for the NO<sub>3</sub> radical chemistry factor.  
476 There may be several reasons for this discrepancy. One major cause for differences between the spring  
477 dataset of Yan et al. (2016) and the autumn dataset presented here, is that nighttime concentrations  
478 of HOM was greatly reduced during our autumn campaign. The cause may have been fairly frequent  
479 fog formation during nights, and also the concentration of e.g. ozone decreased nearly to zero during  
480 several nights (Zha et al., 2018). It is also possible that the NO<sub>3</sub> radical-related factor by Yan et al.  
481 (2016) is probably a mixture of NO<sub>3</sub> and O<sub>3</sub> radical chemistry, while the monomer may thus be  
482 attributed to the O<sub>3</sub> part. Alternatively, the different conditions during the two measurement periods,  
483 as well as seasonal difference in monoterpene mixtures (Hakola et al., 2012), caused variations in the  
484 oxidation pathways.

#### 485 4.1.4 Contamination factor

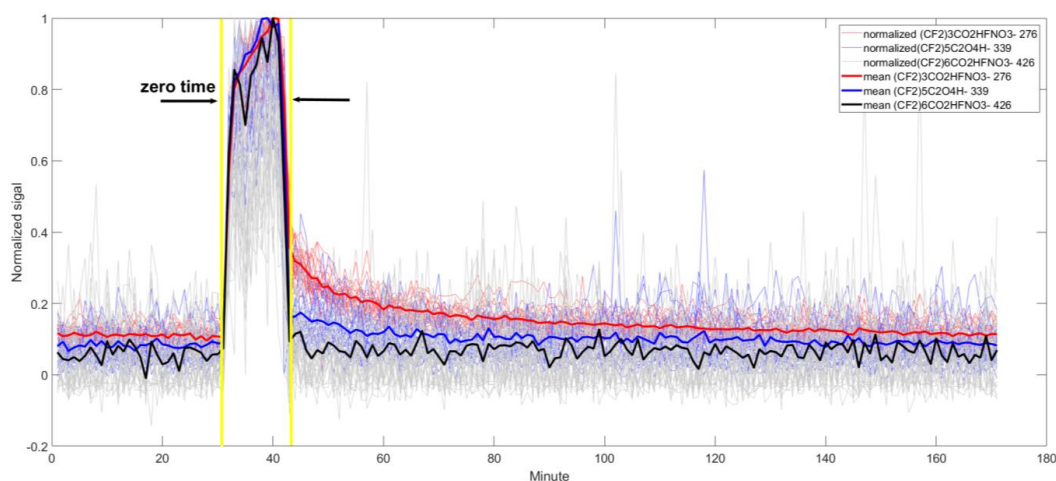
486 During the campaign, an automated instrument zeroing every three hours was conducted, by  
487 switching a valve to pass the air through a HEPA filter. Each zeroing process lasted for 10 min. While  
488 the zeroing successfully removed the low-volatile HOM and H<sub>2</sub>SO<sub>4</sub>, the zeroing process introduced  
489 contaminants into the inlet lines. The contaminants were primarily different types of perfluorinated  
490 organic acids, often off-gassing from e.g. Teflon tubing. For IVOC contaminants, these would be  
491 flushed through the inlet, while (E)LVOC would condense onto the inlet walls and not come off.  
492 However, SVOC contaminants may stick to the inlet tubing and slowly evaporate back into the  
493 sampled air. We removed all the 10-min zeroing periods, and averaged the data to 1-h time resolution,  
494 but contaminants were still identified in all ranges by binPMF.

495 Contamination contributed 10%, 3%, 19% and 4% to Range 1, 2, 3, and Combined, respectively, in  
496 the binPMF solutions where the contamination factor was first separated. This also explains why the  
497 contamination factor was separated much earlier in Ranges 1 and 3 than in Range 2. However, despite  
498 contributing slightly more to Range Combined than to Range 2, the contamination factor was  
499 separated when the factor number was increased by one in Range Combined. Here, the difference in  
500 volatility of the contaminants in the different sub-ranges may play a role, such that the contaminants  
501 in different sub-ranges behave differently. Thus, the behavior of the contamination factor across the  
502 combined range is not consistent. Therefore, we examined the zeroing effect with finer time  
503 resolution, i.e. 1 min, with three of the largest fluorinated compounds in each range of our mass  
504 spectrum, (CF<sub>2</sub>)<sub>3</sub>CO<sub>2</sub>HF·NO<sub>3</sub><sup>-</sup> (275.9748 Th), (CF<sub>2</sub>)<sub>5</sub>C<sub>2</sub>O<sub>4</sub>H (338.9721 Th), and (CF<sub>2</sub>)<sub>6</sub>CO<sub>2</sub>HF·NO<sub>3</sub><sup>-</sup>  
505 (425.9653 Th). Since the overall signal levels were very low for these compounds, the time series



506 became very noisy with such high time resolution. This made it impossible to perform HR fitting for  
507 the data, and instead we summed up the signal from the mass ranges where we expected unperturbed  
508 signal from these ions.

509 The time series with sawtooth pattern of the three fluorinated compounds is shown in Figure S3 in  
510 Supplement. From the time series, we selected a period of around three days of the 3-h cycles (25  
511 total), and in Figure 7 the cycles were aligned and superimposed on top of one another, normalized  
512 by the maximum during the zeroing. The normalized signals of the three compounds are shown in  
513 light colors, and the mean values shown in bold solid lines. This data includes also the zeroing periods  
514 to highlight the effect, but these periods were removed from the data used for our PMF analyses.



515  
516 Figure 7 Normalized signals for three fluorinated compounds during a 3-h cycle (180 minutes), with  
517  $(\text{CF}_2)_3\text{CO}_2\text{HF}\cdot\text{NO}_3^-$  (275.9748 Th) in red,  $(\text{CF}_2)_5\text{C}_2\text{O}_4\text{H}^-$  (338.9721 Th) in blue, and  
518  $(\text{CF}_2)_6\text{CO}_2\text{HF}\cdot\text{NO}_3^-$  (425.9653 Th) in black. We selected 25 cycles and normalized all the cycles by  
519 their individual maximum. The yellow window shows the zeroing time, for around 10 minutes, which  
520 has been removed from the data analysis. Light colors display the individual cycles, and the bold  
521 solid colors stand for the average for each compounds.

522 The signals of the three fluorinated compounds increased by 10 to 20 times during the zeros, due to  
523 off-gassing either in the filter or in the tubing in the zeroing setup. Immediately after the zeroing was  
524 stopped, signals of all three compounds dropped by about 60-90%, followed by a gradual decay. The  
525 decay period coincided with our ambient sampling, and therefore these signals are part of our dataset.  
526 It is evident that the three fluorinated compounds were from the same source (zeroing process), but  
527 due to their different volatilities, they were lost at different rates. This, in turn, means that the spectral  
528 signature of this source will change as a function of time, at odds with one of the basic assumptions



529 of PMF. Panels a and b in Figure S4 displays the temporal correlation with and without zeroing period  
 530 with 1 min time resolution. The correlation coefficients dropped greatly when the zero period was  
 531 removed, from 0.9 to 0.3 for  $R^2$  between 276 Th and 339 Th, and 0.8 to 0.1 between 276 Th and 426  
 532 Th. Similar effect is also found with the 1 h averaged data (Fig. S4c, d).  
 533 This detailed analysis of fluorinated contamination in our system was here merely used as an example  
 534 to show that volatility can impact source profiles over time. In this case, the contamination factor was  
 535 still identified both from the separate sub-ranges and from the combined data set using binPMF.  
 536 However, the contamination profile in the combined range is now averaged, compared to that from  
 537 separate ranges: the fractional contributions of contamination compounds to this profile, vary during  
 538 the process of each zeroing due to different volatility properties. In Figure S5, contamination factor  
 539 profiles from Range 3 and Range Combined were compared. It can be clearly seen that the profile of  
 540 Range Combined is more noisy than that of Range 3, probably due to the varied fractional  
 541 contributions of contamination compounds to the profile. In ambient data, products from different  
 542 sources can have undergone atmospheric processing, altering the product distribution. Our aim with  
 543 this analysis was to highlight the importance of differences in the sink terms due to different  
 544 volatilities of the products. This may be an important issue for gas phase mass spectrometry analysis,  
 545 potentially underestimated by many PMF users, as it is likely only a minor issue for aerosol data, for  
 546 which PMF has been applied much more routinely. If failing to achieve physically meaningful factors  
 547 using PMF on gas phase mass spectra, our recommendation is to try applying PMF to sub-ranges of  
 548 the spectrum, where IVOC, SVOC and (E)LVOC could be analyzed separately.

#### 549 4.2 Atmospheric insights

550 While the previous section discussed several findings with atmospheric implications, we highlight  
 551 two results below, which are particularly intriguing. We also include the correlation matrix of all  
 552 PMF and factors and selected meteorological parameters in Table 2.

553 Table 2 Correlation between factors and meteorological parameters and gases

	R1F1 _D1	R1F1 _D2	R1F1 _D3	R1F1 _C	R2F1 _D1	R2F2 _D2	R2F3 _D3	R2F4 _N	R3F1 _D	R3F2 _N1	R3F3 _N2	R3F4 _C	RCF1 _D1	RCF2 _D2	RCF3 _D3	RCF4 _N
O <sub>3</sub>	0.51	0.59	0.35	-0.18	0.47	0.57	0.36	0.43	0.55	0.33	0.27	0.22	0.49	0.57	0.33	0.34
NO	0.13	-0.01	0.24	-0.03	0.18	-0.02	0.24	-0.22	0.13	-0.19	-0.17	0.03	0.13	0.00	0.26	-0.18
NOx	-0.05	-0.22	-0.10	0.09	-0.01	-0.23	-0.11	-0.13	-0.16	-0.21	-0.04	0.04	-0.04	-0.22	-0.09	-0.11
RH	-0.46	-0.80	-0.63	0.30	-0.43	-0.82	-0.64	-0.27	-0.78	-0.39	-0.07	-0.07	-0.43	-0.82	-0.60	-0.21
T	0.66	0.72	0.40	-0.24	0.65	0.66	0.41	0.39	0.65	0.30	0.14	0.19	0.66	0.68	0.38	0.24
UVB	0.52	0.63	0.82	-0.40	0.52	0.68	0.84	-0.30	0.79	-0.08	-0.27	0.08	0.49	0.68	0.83	-0.29

554

555



#### 556 4.2.1 Daytime dimer formation

557 Dimers are primarily produced during nighttime, due to NO suppressing RO<sub>2</sub> + RO<sub>2</sub> reactions in  
558 daytime (Ehn et al., 2014; Yan et al., 2016). However, in this study, we found one clear daytime factor  
559 in Range 3 (R3F1\_D, peak at local time 12:00, UTC+2) by sub-range analysis. With high loadings  
560 from even masses including 516, 518, 520, 528, 540 Th, this only daytime factor in dimer range  
561 correlated very well with two daytime factors in Ranges 1 and 2 (R1F2\_D2, R1F3\_D3, R2F2\_D2,  
562 R2F3\_D3) (Figure 6b and c). Strong correlation between R3F1\_D with solar radiation was found,  
563 with R = 0.79 (Table 2). This may indicate involvement of OH oxidation in producing this factor.

564 As previous studies have shown, dimers greatly facilitate new particle formation (NPF) (Kirkby et  
565 al., 2016; Troestl et al., 2016; Lehtipalo et al., 2018), and this daytime dimer factor may represent a  
566 source of dimers that would impact the initial stages of NPF in Hyytiälä. Mohr et al. (2017) reported  
567 a clear diel pattern of dimers (sum of about 60 dimeric compounds of C<sub>16-20</sub>H<sub>13-33</sub>O<sub>6-9</sub>) during NPF  
568 events in 2013 in Hyytiälä, with minimum at night and maximum after noon, and estimated these  
569 dimers can contribute ~5% of the mass of sub-60 nm particles. The link between the dimers presented  
570 in that paper and those reported here will require further studies, as will the proper quantification of  
571 the dimer factor identified here.

#### 572 4.2.2 Dimers initiated by NO<sub>3</sub> radicals

573 Previous studies show that NO<sub>3</sub> oxidation of α-pinene, the most abundant monoterpene in Hyytiälä  
574 (Hakola et al., 2012), produces fairly little SOA mass (yields <4 %), while β-pinene shows yields of  
575 up to 53 % (Bonn and Moorgat, 2002; Nah et al., 2016). The NO<sub>3</sub>+β-pinene reaction results in low  
576 volatile organic nitrate compounds with carboxylic acid, alcohol, and peroxide functional groups (Fry  
577 et al., 2014; Boyd et al., 2015), while NO<sub>3</sub>+α-pinene reaction will typically lose the nitrate functional  
578 group and form oxidation products with high vapor pressures (Spittler et al., 2006; Perraud et al.,  
579 2010). Most monoterpene-derived HOM, including monomers, are low-volatile (Peräkylä et al.,  
580 2019) and thus a low SOA yield indicates a low HOM yield. Thus, while there are to our knowledge  
581 no laboratory studies on HOM formation from NO<sub>3</sub> oxidation of α-pinene, a low yield can be expected  
582 based on SOA studies.

583 As discussed in section 4.1.3, a dimer factor (R3F2\_N2) was identified as a mix of ozonolysis and  
584 NO<sub>3</sub> oxidation processes, dominated by the organonitrate at 555 Th, C<sub>20</sub>H<sub>31</sub>O<sub>10</sub>NO<sub>3</sub>·NO<sub>3</sub><sup>-</sup>. However,  
585 unlike the pure ozonolysis dimer factor which had a corresponding monomer factor (R = 0.86 between  
586 factor R2F4\_N and R3 F2\_N1), this NO<sub>3</sub>-related dimer factor did not have an equivalent monomer  
587 factor. This suggests that the NO<sub>3</sub> oxidation of the monoterpene mixture in Hyytiälä does not by itself  
588 form much HOM, but in the presence of RO<sub>2</sub> from ozonolysis, the RO<sub>2</sub> from NO<sub>3</sub> oxidation can take  
589 part in HOM dimer formation. This further implies that, different from previous knowledge based on



590 single-oxidant experiments in chambers, NO<sub>3</sub> oxidation may have a larger impact on SOA formation  
591 in the atmosphere where different oxidants exist concurrently. This highlights the need for future  
592 laboratory studies to consider systems with multiple oxidants during monoterpene oxidation  
593 experiments, to truly understand the role and contribution of different oxidants, and NO<sub>3</sub> in particular.

594

## 595 **5 Conclusions**

596 The recent development in mass spectrometry has greatly improved the detection of atmospheric  
597 vapors and their oxidation products. Factor analysis, such as PMF, can reduce dimensionality of the  
598 big datasets and extract factors relating to different atmospheric pathways/sources. Optimally, PMF  
599 can link laboratory-generated spectra with ambient observations, significantly improving our  
600 understanding of complicated atmospheric processes. However, one of PMF's basic assumptions is  
601 that factor profiles remain constant in time, yet for atmospheric gas-phase species, varying sources,  
602 reactions and sinks may violate this assumption. Some of these variations are likely not addressable  
603 in the data analysis stage, but others may be. For example, molecules formed from the same source  
604 can have different temporal behaviors due to varying volatilities, and thus condense at different rates.  
605 Performing PMF separately with smaller ranges may circumvent this problem. By utilizing the newly  
606 presented binPMF approach, more variables can be extracted from a narrow mass range compared to  
607 traditional UMR PMF, while preserving more information in the spectrum.

608 We conducted separate binPMF analysis on three different sub-ranges to explore the potential benefits  
609 of such an approach for producing more physically meaningful factors. We utilized ambient data  
610 measured by CI-API-TOF in a boreal forest environment, and selected sub-ranges from the mass  
611 spectrum that roughly corresponded to regions where we would expect the molecules to have similar  
612 volatilities and formation pathways. Selected ranges were Range 1 (250 – 300 Th), Range 2 (300 –  
613 350 Th), and Range 3 (510 – 560 Th). binPMF was separately applied to these ranges, as well as to  
614 the combination of all three for comparison.

615 The different sub-ranges produced some similar and some different factors. First of all, volatility of  
616 species indeed affect the PMF results. We could clearly prove the benefit of sub-range binPMF using  
617 the contamination factor as an example. We found that different compounds emitted from the same  
618 source showed different temporal trends, likely due to differences in volatilities. This increased the  
619 difficulties for PMF to separate this source in the combined data set, and the resolved profile was still  
620 less accurate than in the analysis for the sub-ranges. We recommend that future studies of gas-phase  
621 mass spectra should pay attention to this volatility effect on factor analysis.

622 Secondly, chemistry or sources contributing to the particular range can be better separated. Only the  
623 binPMF analysis on Range 3, where HOM dimers are typically observed, resolved two nighttime



624 factors, characterized by monoterpene oxidation related to  $\text{NO}_3$  and  $\text{O}_3$  oxidation. The monoterpene  
625 ozonolysis factor was separated from both Range 2 and 3, showing very good correlation between  
626 the ranges and mutually verifying the results.

627 Thirdly, peaks with smaller signal intensities can be correctly assigned. The signal intensities between  
628 different parts of the mass spectrum may vary by orders of magnitude. In the analysis of the combined  
629 range, the results were almost completely controlled by the higher signals from Range 1 and 2. The  
630 separate analysis on Range 3 allowed the low signals to provide important information, such as the  
631  $\text{NO}_3$  oxidation process. In addition, running binPMF on different separate mass ranges also allows us  
632 to compare the factors obtained from the different ranges and help to verify the results.

633 In addition, daytime dimer formation was identified, presumably initiated by  $\text{OH}/\text{O}_3$  with a diurnal  
634 peak at around noon, which may contribute to NPF in Hyytiälä. Also, based on the sub-range binPMF  
635 analysis, we successfully separated  $\text{NO}_3$ -related dimers which did not have an equivalent monomer  
636 factor. The  $\text{NO}_3$  related factor was consistent with earlier observations (Yan et al., 2016), but would  
637 not have been identified from this dataset without utilizing the different sub-ranges. In future  
638 laboratory experiments, more complex oxidation systems may be useful in order to understand the  
639 role  $\text{NO}_3$  oxidation in SOA formation.

640 In summary, we identified several reasons to recommend PMF users to try running their analysis on  
641 selected sub-ranges in addition to the whole spectra. Ultimately, the approach should be study-goal  
642 dependent. In some cases the researcher wants a quick factor analysis to explore different features of  
643 their data, while in others more accurate and quantitative separation of different sources and  
644 atmospheric processes are needed. As binPMF and UMR-PMF both require very little data  
645 preparation, we expect that it in most cases will be worth the time for the analyst to test how PMF  
646 results look for a few selected sub-ranges of their mass spectra.

647

653 **Data availability.** The data used in this study are available from the first author upon request: please  
654 contact Yanjun Zhang (yanjun.zhang@helsinki.fi).

662 Author contributions. ME and YZ designed the study. QZ and MR collected the data; data analysis  
663 and manuscript writing were done by YZ. All coauthors discussed the results and commented the  
664 manuscript.

665 **Competing interests.** The authors declare that they have no conflict of interest

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673

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