



1 **Supercooled Liquid Water Clouds observed and analysed at the**  
2 **Top of the Planetary Boundary Layer above Dome C, Antarctica**

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21 **Abstract**

22 A comprehensive analysis of the water budget over the Dome C (Concordia, Antarctica) station  
23 has been performed during the austral summer 2018-2019 as part of the Year of Polar Prediction  
24 (YOPP) international campaign. Thin (~100-m) supercooled liquid water (SLW) clouds have  
25 been detected and analysed using remotely sensed observations at the station (tropospheric  
26 depolarization LIDAR, microwave radiometer HAMSTRAD, net surface radiation from  
27 Baseline Surface Radiation Network, BSRN), radiosondes and using satellite observations  
28 (CALIOP/CALIPSO) combined with a specific configuration of the Numerical Weather  
29 Prediction model: ARPEGE-SH. Two case studies are used to illustrate this phenomenon. On  
30 24 December 2018, the atmospheric planetary boundary layer (PBL) evolved following a  
31 “typical” diurnal variation, that is to say with a warm and dry mixing layer at local noon thicker  
32 than the cold and dry stable layer at local midnight. Our study showed that the SLW clouds  
33 were observed at Dome C within the entrainment and the capping inversion zones at the top of  
34 the PBL. ARPEGE-SH was not able to correctly estimate the ratio between liquid and solid  
35 water inside the clouds. The SLW content was always strongly underestimated in the studied  
36 cases. The lack of simulated SLW in the model impacted the net surface radiation that was 20-  
37 30 W m<sup>-2</sup> higher in the BSRN observations than in the ARPEGE-SH calculations, mainly  
38 attributable to longwave downward surface radiation from BSRN being 50 W m<sup>-2</sup> greater than  
39 that of ARPEGE-SH. On 20 December 2018, a warm and wet episode impacted the PBL with  
40 no clear diurnal cycle of the PBL top. SLW cloud appearance coincided with the warm and wet  
41 event within the entrainment and capping inversion zones. The amount of liquid water measured  
42 by HAMSTRAD was ~20 times greater in this perturbed PBL than in the “typical” PBL. Since  
43 ARPEGE-SH was not able to accurately reproduce these SLW clouds, the discrepancy between  
44 the observed and calculated net surface radiation was even greater than in the “typical” PBL  
45 period, reaching +50 W m<sup>-2</sup>, mainly attributable to longwave downward surface radiation from



46 BSRN being  $100 \text{ W m}^{-2}$  greater than that of ARPEGE-SH. The absence of SLW clouds in  
47 NWP over Antarctica may indicate an incorrect simulation of the radiative budget of the polar  
48 atmosphere.

49



## 50 **1. Introduction**

51 Antarctic clouds play an important role in the climate system by influencing the Earth's  
52 radiation balance, both directly at high southern latitudes and, indirectly, at the global level  
53 through complex teleconnections (Lubin et al., 1998). In Antarctica, there are very few  
54 observational stations and most of them are located on the coast, a fact that limits the type and  
55 characteristics of clouds observed. Nevertheless, prior studies suggest that cloud properties vary  
56 geographically, with a fractional cloud cover around the South Pole of about 50 to 60% in all  
57 seasons, and a cloud cover of about 80 to 90% near the coast (Bromwich et al., 2012). Based  
58 on spaceborne observations (Adhikari et al., 2012), the Antarctic Plateau has the lowest cloud  
59 occurrence of the Antarctic continent (<30%). Furthermore, cloud parameters such as the  
60 hydrometeors size and the microphysical structure are also very difficult to retrieve in  
61 Antarctica. Nevertheless, some measurements exist showing that ice crystal clouds are mainly  
62 observed inland with crystal sizes ranging from 5 to 30  $\mu\text{m}$  (effective radius) in the core of the  
63 cloud; mixed-phase clouds are preferably observed near the coast with slightly larger ice  
64 crystals and water droplets (Lachlan-Cope, 2010).

65 The time and geographical distribution of tropospheric clouds above the whole Antarctic  
66 continent has been recently studied using the raDAR/liDAR-MASK (DARDAR) spaceborne  
67 products (Listowski et al., 2018). The authors determined that clouds are mainly constituted of  
68 ice above the continent. The presence of Supercooled Liquid Water (SLW, the water staying in  
69 liquid phase below 0°C) clouds shows variations according to temperature and sea ice fraction,  
70 decreasing sharply poleward, with an abundance two to three times less over the Eastern  
71 Antarctic Plateau than over the Western Antarctic. The difficulty of mesoscale high-resolution  
72 models and operational numerical weather prediction models to accurately calculate the net  
73 surface radiation due to the presence of clouds (particularly of SLW clouds) in Antarctica



74 causes biases of several tens of watt per square meters (King et al., 2006; Bromwich et al.,  
75 2013) impacting the Earth's radiative budget (Lawson and Gettelman, 2014).

76 With the support of the World Meteorological Organization (WMO) World Weather  
77 Research Programme (WWRP), the Polar Prediction Project (PPP) international programme  
78 has been dedicated to the development of improved weather and environmental prediction  
79 services for the polar regions, on time scales from hours to seasons  
80 (<https://www.polarprediction.net>). Within this project, the Year of Polar Prediction (YOPP),  
81 from 2018 to 2019, aims at enabling a significant improvement in environmental prediction  
82 capabilities for the polar regions and beyond, by coordinating a period of intensive observing,  
83 modelling, verification, user-engagement and educational activities. The Water Budget over  
84 Dome C (H<sub>2</sub>O-DC) project has been endorsed by YOPP for studying the water budget by means  
85 of ground-based measurements of water (vapour, solid and liquid) and clouds, by active  
86 (backscatter LIDAR) and passive (microwave radiometer) remote sensing, and operational  
87 meteorological analyses. The Dome C (Concordia) station is located in the Eastern Antarctic  
88 Plateau (75°06'S, 123°21'E, 3233 m above mean sea level, amsl).

89 H<sub>2</sub>O-DC concentrates on the Special Observing Period (SOP) of measurements in the  
90 Antarctic, from 16 November 2018 to 15 February 2019. During this time frame, several  
91 instruments have been employed.

92 1) The H<sub>2</sub>O Antarctica Microwave Stratospheric and Tropospheric Radiometer  
93 (HAMSTRAD) radiometer (Ricaud et al., 2010a) to obtain vertical profiles of temperature and  
94 water vapour, Integrated Water Content (IWC) or precipitable water, and Liquid Water Path  
95 (LWP), with an adjustable time resolution better than 7 minutes.

96 2) The tropospheric depolarization LIDAR (Tomasi et al., 2015) to obtain vertical profiles  
97 of backscattering and depolarization.



98        These two H<sub>2</sub>O-DC data sets have been complemented in the present analysis by the 3  
99 following observational datasets.

100        3) The Baseline Surface Radiation Network (BSRN) net surface radiances at the station.

101        4) The temperature profiles from radiosondes launched twice daily at the station during  
102 YOPP.

103        5) The spaceborne observations (backscatter and polarization) from the  
104 CALIOP/CALIPSO LIDAR in the vicinity of the station.

105        In addition, a specific configuration of the global ARPEGE model from Météo-France  
106 (Pailleux et al., 2015) is used to characterize the water budget above Dome C considering the  
107 gas, liquid and the solid phases to study the genesis of clouds (ice/liquid).

108        The aim of the present study is to combine all these observations and simulations in order  
109 to 1) detect the presence of SLW clouds above Dome C, 2) analyse the formation and evolution  
110 of such SLW clouds and 3) estimate the radiative impact of such clouds on the net surface  
111 radiation. We concentrate the analyses on two case studies observed during the YOPP  
112 campaign: one case when the Planetary Boundary Layer (PBL) exhibited a “typical” diurnal  
113 cycle (24 December 2018) and a second case when the diurnal cycle of the PBL was perturbed  
114 by a warm and wet episode (20 December 2018).

115        The data sets used in our study are presented in section 2. The analyses of the SLW clouds  
116 during the typical and the perturbed PBL periods are detailed in sections 3 and 4, respectively.

117 A conclusion synthesizes the study in section 5.

118

## 119 **2. Datasets**

### 120 **2.1. The HAMSTRAD Radiometer**

121        HAMSTRAD is a microwave radiometer to probe water vapour (H<sub>2</sub>O), liquid water and  
122 tropospheric temperature above Dome C. Measuring at 60 GHz (oxygen molecule line (O<sub>2</sub>) to



123 deduce the temperature) and at 183 GHz (H<sub>2</sub>O line), this unique, state-of-the-art radiometer was  
124 installed on site for the first time in January 2009 (Ricaud et al., 2010a). The measurements of  
125 the HAMSTRAD radiometer allow the retrieval of the vertical profiles of H<sub>2</sub>O and temperature  
126 from the ground to 10 km altitude with a temporal resolution of 7 minutes and vertical  
127 resolutions of 30 to 50 m in the atmospheric boundary layer, 100 m in the free troposphere and  
128 500 m in the upper troposphere-lower stratosphere. The temporal resolution on the instrument  
129 allows for detection and analysis of atmospheric processes such as the diurnal evolution of the  
130 PBL (Ricaud et al., 2012) and the presence of clouds and diamond dust (Ricaud et al., 2017).  
131 In addition, two other parameters can be estimated.

132 1) The Integrated Water Vapour (IWV) or precipitable water (kg m<sup>-2</sup>) obtained by  
133 integrating the absolute humidity profile from the surface to 10 km altitude.

134 2) The Liquid Water Path (kg m<sup>-2</sup>) that gives the amount of liquid water integrated along  
135 the vertical.

136 IWV has been validated against radiosondes at Dome C between 2010 and 2014 showing a  
137 5-10% wet bias of HAMSTRAD compared to the sondes (Ricaud et al., 2015). LWP has only  
138 been studied when the instrument was installed at the Pic du Midi station (2877 amsl, France)  
139 during the calibration/validation period in 2008 prior to its set up in Antarctica in 2009 (Ricaud  
140 et al., 2010b).

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## 142 **2.2. The tropospheric depolarization LIDAR**

143 A tropospheric depolarization LIDAR (532 nm) has been operating at Dome C since 2008  
144 (see [http://lidarmax.altervista.org/englidar/\\_Antarctic%20LIDAR.php](http://lidarmax.altervista.org/englidar/_Antarctic%20LIDAR.php)). The LIDAR provides  
145 5-min tropospheric profiles of aerosols and clouds continuously, from 20 to 7000 m amsl, with  
146 a resolution of 7.5 m. LIDAR depolarization (Mishchenko et al., 2000) is a robust indicator of  
147 non-spherical shape for randomly oriented cloud particles. A depolarization ratio below 10% is



148 characteristic of SLW clouds, while higher values are produced by ice particles. The possible  
149 ambiguity between SLW clouds and oriented ice plates is avoided at Dome C by operating the  
150 LIDAR 4° off-zenith (Hogan and Illingworth, 2003). The LIDAR observations at Dome C have  
151 already been used to study the radiative properties of water vapour and clouds in the far infrared  
152 (Palchetti et al., 2015). As a support to LIDAR data interpretation, time-lapse webcam videos  
153 of local sky conditions are also collected.

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### 155 **2.3. The BSRN Network**

156 At the Astroconcordia/Albedo-Rack sites, the upward and downward looking, heated and  
157 ventilated standard Kipp&Zonen CM22 pyranometers and infrared CG4 pyrgeometers provide  
158 measurements of hemispheric downward and upward broadband shortwave (SW, 0.3–3  $\mu\text{m}$ )  
159 and longwave (LW, 4–50  $\mu\text{m}$ ) fluxes at the surface, respectively. These data are used to retrieve  
160 values of net surface radiation (defined as the difference between the downward and upward  
161 fluxes). All these measurements follow the rules of acquisition, quality check and quality  
162 control of the BSRN as described in Driemel et al. (2018).

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### 164 **2.4. Radiosondes**

165 Vertical temperature and humidity profiles have been measured on a daily basis at Dome C  
166 since 2005, employing RS92 radiosondes using Vaisala standard evaluation programs. The  
167 sondes are known to have a cold bias of 1.2 K from the ground to about 4 km altitude (Tomasi  
168 et al., 2011 and 2012). During YOPP and the two case studies, launches were performed twice  
169 per day at 00:00 and 12:00 UTC.

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171 **2.5. CALIOP onboard CALIPSO**

172 Orbiting at 705-km altitude, the CALIPSO (Cloud Aerosol Lidar and Infrared Pathfinder  
173 Satellite Observations) mini-satellite has been observing clouds and aerosols since 2006 to  
174 better understand the role of clouds and aerosols in climate. To accomplish this mission, the  
175 CALIPSO satellite is equipped with a LIDAR, a camera and an infrared imager (Winker et al.,  
176 2009). CALIOP (Cloud-Aerosol Lidar with Orthogonal Polarization) is a dual-wavelength (532  
177 and 1064 nm) backscatter LIDAR. It provides high-resolution vertical profiles of clouds and  
178 aerosols along the orbit track (Young et al., 2009).

179

180 **2.6. The specific ARPEGE-SH Model**

181 A specific configuration of the operational global model ARPEGE was used for the YOPP  
182 Southern Hemisphere (YOPP-SH) period (15/11/2018–15/02/2019). This configuration named  
183 ARPEGE-SH is based on the operational global model used for Numerical Weather Prediction  
184 (NWP) ARPEGE (Pailleux et al., 2015), but with its highest horizontal resolution centred over  
185 Dome C instead of over France, as set up in ARPEGE. A 4D variational (4DVar) assimilation  
186 was performed every 6 h. The meteorological analyses were given by the ARPEGE-SH system  
187 together with the 24-hour forecasts at the node the closest to the location of Dome C. The  
188 horizontal and vertical resolution during the YOPP-SH period were 7.5 km at Dome C, with  
189 105 vertical levels, the first one being set at 10 m. Several ARPEGE-SH output parameters  
190 were selected for analysis: cloud fraction, ice, water vapour and liquid-water mixing ratio,  
191 temperature, Total Column Ice (TCI, ice integrated along the vertical), LWP, IWV, and net  
192 surface radiation.

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### 196 **3. Typical diurnal cycle of the PBL**

197 The first SLW cloud case study occurred on 24 December 2018 over a 24-hour period  
198 with a typical diurnal PBL cycle. All the results are presented in Universal Time Coordinated  
199 (UTC) with local time (LT) being eight hours ahead of UTC (LT = UTC + 8).

200

#### 201 **3.1. Clouds**

202 The presence of clouds is highlighted by the LIDAR backscatter and depolarization profiles  
203 shown in Figures 1a and b, respectively. High values of LIDAR cloud backscatter ( $\beta_c > 100$   
204  $\beta_{\text{mol}}$ , with  $\beta_{\text{mol}}$  the molecular backscatter), indicate that clouds and/or precipitation are present  
205 over the whole day with some significant differences. First, vertical “stripes” of high  
206 backscatter values are visible from 10 to 400 m height before 10:00 UTC and after 19:00 UTC,  
207 associated with high values of depolarization ratio ( $> 20\%$ ), characteristic of precipitating ice  
208 crystals. Second, high values of  $\beta_c$  associated with very low depolarization ratio ( $< 5\%$ ) occur  
209 within a thin layer of approximately 100-m depth around 500 m from 08:00 to 22:00 UTC, with  
210 some breaks around 11:00 and 19:00-21:00 UTC. From the LIDAR observations, this  
211 combination of high backscatter and low depolarization ratio signifies the presence of a SLW  
212 cloud (Figure 1c).

213 The NWP model ARPEGE-SH calculates cloud fraction, ice water and liquid water mixing  
214 ratios ( $\text{kg kg}^{-1}$ ) for 24 December 2018 (Figures 2a, b and c, respectively). We note that the  
215 outputs from ARPEGE-SH at 00:00 and 12:00 UTC are the analyses and, for the remaining  
216 time, the outputs are the hourly forecasts. ARPEGE-SH predicts the presence of clouds (cloud  
217 fraction  $> 0.95$ ) for most of the day except around 11:00 and 23:00 UTC (Fig. 2a). Before 12:00  
218 UTC, the clouds are mainly confined between 300 and 600-800 m whilst, after 12:00 UTC,  
219 they spread from the surface to 800 m. There are also high-level clouds at 2000-3000 m height  
220 but with a cloud fraction between 0.50 and 0.70. The majority of the clouds produced by



221 ARPEGE-SH are mainly composed of ice crystals (Fig. 2b) with some traces of droplets (Fig.  
222 2c). The ice/liquid water clouds derived from the LIDAR observations are superimposed over  
223 the SLW clouds calculated by ARPEGE-SH. It is obvious that ARPEGE-SH fails in estimating  
224 both the vertical distribution of liquid water (a thin layer is observed around 500 m whereas the  
225 modelled cloud layer extends from the surface to 800 m) and its temporal evolution (presence  
226 of SLW cloud almost all day long in ARPEGE-SH compared to SLW clouds from 08:00 to  
227 22:00 UTC in the observations).

228 The presence of clouds above the station can also be inferred from vertically-integrated  
229 variables such as: 1) TCI calculated by ARPEGE-SH, 2) LWP from HAMSTRAD and  
230 ARPEGE-SH, and 3) IWV from HAMSTRAD and ARPEGE-SH (Figures 3a, b and c,  
231 respectively). The ARPEGE-SH TCI on 24 December 2018 (Fig. 3a) oscillates between 10 and  
232  $30 \text{ g m}^{-2}$  except around 12:00 UTC when a clear minimum occurs ( $\sim 3 \text{ g m}^{-2}$ ), underlining the  
233 fact that ARPEGE-SH obtains ice clouds for the entire day, except at 12:00 UTC. The  
234 HAMSTRAD LWP shows an obvious increase from  $\sim 1.0$  to  $\sim 2.0\text{-}3.0 \text{ g m}^{-2}$  when the presence  
235 of SLW cloud is indicated by LIDAR observations. The ARPEGE-SH LWP is on average  $10^3$   
236 times lower than that observed by HAMSTRAD, underlining the fact that ARPEGE-SH  
237 misrepresents features of the SLW clouds over Dome C. HAMSTRAD and ARPEGE-SH IWV  
238 vary from  $0.65\text{-}1.05 \text{ kg m}^{-2}$  throughout the day on 24 December 2018, with an agreement to  
239 within  $0.1 \text{ kg m}^{-2}$  (i.e.  $\sim 10\text{-}15\%$ ), which is consistent with previous studies (Ricaud et al., 2017).  
240 Note that HAMSTRAD IWV is on average 10% higher than IWV inferred from radiosonde  
241 observations (Ricaud et al., 2015).

242 Observation of clouds from space-borne sensors has two main advantages: 1) to validate  
243 the nature of the cloud observed over Dome C (namely ice/liquid water), and 2) to estimate the  
244 vertical and horizontal extents of the detected cloud. Note that the CALIPSO spaceborne  
245 LIDAR operates at the same wavelength as the backscatter LIDAR at Dome C, with the same



246 method for discriminating ice from liquid water. Consequently, the two LIDARs should give  
247 consistent information for the detected phase of the clouds. The main difficulty with this  
248 approach is related to the time and space sampling of the spaceborne instrument, namely to find  
249 a satellite overpass coincident both in time and location with the cloud observed at Dome C.  
250 This, unfortunately, decreases the number of overpasses that is scientifically exploitable.  
251 Nevertheless, on 24 December 2018, 2 orbits of CALIOP/CALIPSO passed close to Dome C  
252 at times when SLW clouds were observed by ground-based instruments. We show the vertical  
253 feature mask and ice/water phase from the pass closest to the station, from 15:50 to 16:03 UTC  
254 (Figures 4a and b, respectively). Firstly, we note the presence of a cloud a few hundreds of  
255 meters deep near the surface in the vicinity of Dome C (Fig. 4a; note that the  
256 CALIOP/CALIPSO altitude is above sea level and Dome C is at an altitude of 3233 m amsl).  
257 Secondly, this cloud is composed of SLW (Fig. 4b), confirming the analysis based on the  
258 observations from the LIDAR and the HAMSTRAD radiometer. Furthermore, we can state that  
259 this SLW cloud is not a local phenomenon but is at least 2.5°-latitude wide, namely has a  
260 horizontal extent of ~280 km. The other orbit from 14:11 to 14:24 UTC (not shown) is slightly  
261 more distant than the one shown in Figure 4, but it exhibits a similar SLW cloud over an even  
262 greater horizontal extent of about 5° latitude (~550 km).

263

### 264 **3.2. Temperature and water vapour**

265 The presence of SLW clouds in the atmosphere is strongly dependent on the temperature  
266 field. From Fig. 2.33 of Pruppacher and Klett (2012), the percentage of clouds containing no  
267 ice becomes non-negligible at temperatures greater than -35°C, although SLW clouds have been  
268 observed at lower temperatures over Russia (-36°C) and the Rocky Mountains in the USA (-  
269 40.7°C). Recent laboratory measurements show that liquid water can exist down to -42.55°C  
270 (Goy et al., 2018).



271 On 24 December 2018, temperatures from both HAMSTRAD and ARPEGE-SH ranged  
272 from 240 to 250 K (-33 to -23°C) from the surface to 1-km altitude, compatible with the  
273 presence of SLW clouds. The diurnal variations of temperature and water vapour anomalies  
274 calculated by ARPEGE-SH and measured by HAMSTRAD are shown in Figure 5. For each  
275 altitude, the daily-averaged value has been subtracted. This has the advantages of highlighting  
276 areas of maximum and minimum changes along the vertical, and reduces biases when  
277 comparing the two data sets. Absolute anomalies (K) are presented for temperatures whilst  
278 relative anomalies (%) are shown for water vapour.

279 The diurnal variation of the ARPEGE-SH temperature (Fig. 5a) from the surface to 1 km  
280 amsl shows a warm atmosphere before 12:00 UTC and a fast cooling one afterward.  
281 HAMSTRAD shows a similar cooling (Fig. 5b), but the transition is not so abrupt and occurs  
282 later, around 15:00 UTC. The diurnal amplitude is greater in ARPEGE-SH (~5 K) than in  
283 HAMSTRAD (~3 K). The diurnal variation of the water vapour in ARPEGE-SH (Fig. 5c) from  
284 the surface to 1 km shows a wet atmosphere before 12:00 UTC and a drier atmosphere after,  
285 again with an abrupt transition. From HAMSTRAD, the diurnal variation of the water vapour  
286 (Fig. 5d) from the surface to 1 km is more complex, alternating wet and dry phases, which is  
287 particularly obvious at 500-m altitude: wet (00:00-03:00 UTC), dry (03:00-08:00 UTC), wet  
288 (08:00-09:00 UTC), dry (09:00-12:00 UTC), wet (12:00-22:00 UTC) and dry (22:00-24:00  
289 UTC). The time evolution of the SLW cloud (Fig. 1c) is superimposed on all the panels of  
290 Figure 5.

291 The diurnal variation of the top of the PBL calculated by ARPEGE-SH (defined as the  
292 level where the turbulence kinetic energy becomes lower than  $0.01 \text{ m}^2 \text{ s}^{-2}$ ) is also superimposed  
293 on all the panels of Figure 5. We note several key points.



294 1) The diurnal evolution of the top of the PBL is consistent with previous studies carried  
295 out at Dome C (e.g. Argentini et al., 2005; Ricaud et al., 2012; Casasanta et al., 2014), with a  
296 top higher when there is a relatively warm mixed layer than in cool stable conditions.

297 2) The SLW cloud appeared just below the ARPEGE-SH-estimated PBL top, around  
298 08:00 UTC, and persisted around the same altitude after 12:00 UTC even though the top of the  
299 PBL had dramatically decreased down to the surface.

300 3) The SLW cloud persisted after 12:00 UTC in a layer that is wetter and warmer than  
301 elsewhere in the surrounding environment, as demonstrated in both the ARPEGE-SH and  
302 HAMSTRAD data sets.

303

### 304 3.3. Potential Temperature Gradient

305 We now consider the mechanisms that allow the SLW cloud to persist in a thin layer (about  
306 100-m deep) around 500-600 m altitude. Even if the PBL gets thinner after 12:00 UTC, it is  
307 evident that a residual mixed layer remains above (see e.g. Figure 1.7 of Stull, 2012 and Figure  
308 12 top of Ricaud et al., 2012). At the time scale of a whole day, we can define the transition  
309 between the PBL and the free troposphere either at the top of the mixed layer, when this layer  
310 exists, or at the top of the residual layer, when a stable layer develops close to the surface and  
311 decouples the residual mixed layer from the surface. The transition is therefore characterized  
312 by a local maximum of the potential temperature ( $\theta$ ) vertical gradient ( $\partial\theta/\partial z$ ).

313 Figure 6 shows  $\partial\theta/\partial z$  computed from ARPEGE-SH, with the evolution of the PBL top and  
314 the SLW cloud superposed. The SLW cloud, once appeared at the top of the PBL around 08:00  
315 UTC, persists after 12:00 UTC in a layer around 500-600 m coinciding with the local maximum  
316 of  $\partial\theta/\partial z$ , even after the PBL collapses down to the surface.

317 Figures 7a, b and c show the vertical profiles of  $\theta$  (K) and  $\partial\theta/\partial z$  (K km<sup>-1</sup>) as calculated  
318 from temperature measured by the radiosondes and analysed by ARPEGE-SH at Dome C on



319 24 December 2018 at 00:00 and 12:00 UTC and on 25 December 2018 at 00:00 UTC,  
320 respectively. The atmosphere as analysed by ARPEGE-SH is about 3-5 K warmer than the  
321 observations. From 100 m upward, the maximum of  $\partial\theta/\partial z$  is measured at 400, 550 and 600 m  
322 on 24 December 2018 at 00:00 and 12:00 UTC and on 25 December 2018 at 00:00 UTC,  
323 respectively with an amplitude of 10, 12 and 40 K km<sup>-1</sup>, respectively. This is broadly consistent  
324 with the ARPEGE-SH analyses though calculated maxima of  $\partial\theta/\partial z$  (Fig. 7) are slightly higher  
325 (600, 700 and 600 m for the same dates, respectively) and less intense than those of radiosondes  
326 (8, 8 and 18 K km<sup>-1</sup>, respectively).

327 The collocation of the positive potential temperature gradient with the height of the SLW  
328 clouds is consistent with the schematic representation of the diurnal variation of the PBL  
329 illustrated by Stull (2012) and adapted by Ricaud et al. (2012) for the Eastern Antarctic Plateau.  
330 Figure 8 is a modified version of Figure 12 from Ricaud et al. (2012) employing the naming  
331 conventions of Stull (2012). The layer where the clouds develop over the mixed layer is named  
332 the “entrainment zone” and the layer where the ice-water cloud persists over the residual quasi-  
333 mixed layer is named the “residual capping inversion” zone. These two zones are characterized  
334 by a positive  $\partial\theta/\partial z$ . Considering both the potential temperature gradients and the vertical extent  
335 of the SLW cloud, these layers are quite thin, less than 100-m deep.

336 Considering that the SLW clouds are so thin, they resemble stratocumulus, as can be  
337 observed at middle latitudes. The diurnal cycle of the SLW cloud also evokes that of oceanic  
338 stratocumulus, with a trend to fragmentation and/or dissipation during the “day” (local noon)  
339 because of solar absorption and to a solid deck state during the “night” (local midnight) because  
340 of reversed buoyancy due to cloud top longwave cooling. We use here the “night” and “day”  
341 terms for convenience, though solar radiation remains positive 24-hr long at this period of the  
342 year. During the YOPP intensive observing period, SLW clouds were observed for several days  
343 but it is not yet evident whether they were formed during the “day” (local noon) when the mixed



344 layer becomes thick enough to reach the condensation level, and vertically broadened during  
345 the “night”, or created during the “night” (local midnight) and then dissipated during the coming  
346 “day”. Complementary observations would be needed, in particular turbulence profiles from  
347 the surface to above the top of boundary-layer clouds, to determine what is the  
348 coupling/decoupling diurnal cycle of these clouds.

349

### 350 **3.4. Net Surface Radiation**

351 The presence of clouds over Dome C has a strong impact on the net surface radiation as  
352 demonstrated by Ricaud et al. (2017). Figure 9 shows the time evolution of the net surface  
353 radiation as measured by the BSRN instruments and as calculated by ARPEGE-SH on 24  
354 December 2018, superimposed with SLW cloud height. We also show the time evolution of the  
355 difference between surface radiation ( $W m^{-2}$ ) observed by BSRN and calculated by ARPEGE-  
356 SH on 24 December 2018, in longwave downward ( $LW\downarrow$ ), longwave upward ( $LW\uparrow$ ),  
357 shortwave downward ( $SW\downarrow$ ) and shortwave upward ( $SW\uparrow$ ) components, superimposed with  
358 LWP. We highlight 4 periods with images taken from the webcam installed on the shelter  
359 hosting the LIDAR and HAMSTRAD: a) at 00:25 UTC (cirrus clouds, no SLW cloud), b) at  
360 03:56 UTC (cirrus clouds, no SLW cloud), c) at 09:46 UTC (SLW cloud) and d) at 17:20 UTC  
361 (SLW cloud). The net surface radiation shows maxima between 00:00 and 05:00 UTC (08:00-  
362 13:00 LT) and minima between 11:00 and 13:00 UTC (19:00-21:00 LT) in the ARPEGE-SH  
363 and BSRN time series. When SLW clouds are present in the observations (08:00-10:00, 12:00-  
364 19:00 and around 21:00 UTC), whilst absent in ARPEGE-SH, the measured net surface  
365 radiation is systematically greater than the simulated one by 20-30  $W m^{-2}$ . As the SLW  
366 horizontal extent is about 280 km and persists over more than 12 hours (section 3.1), this  
367 discrepancy in the net surface radiation between observation and NWP model may have a strong  
368 impact on the calculation of the radiation budget over Antarctica. In the presence of SLW clouds



369 after 12:00 UTC, this difference is mainly attributable to LW $\downarrow$  component, BSRN values being  
370 50 W m<sup>-2</sup> greater than those of ARPEGE-SH. Thus, SLW clouds tend to radiate LW radiation  
371 toward the ground (like greenhouse gases), at a level higher than more transparent clouds like  
372 cirrus. Note that there are differences from -30 to +60 W m<sup>-2</sup> between observed and calculated  
373 SW $\downarrow$  and SW $\uparrow$  components but this difference falls within  $\pm 10$  W m<sup>-2</sup> for the net SW surface  
374 radiation (SW $\downarrow$  - SW $\uparrow$ ).

375

#### 376 **4. Perturbed diurnal cycle of the PBL**

377 In this section, we focus on the second case study, 20 December 2018, when the diurnal  
378 cycle of the PBL was perturbed by the sudden arrival of very moist, warm air of oceanic origin  
379 (see Ricaud et al., 2017) on 20 December 2018. We analyse how this episode affected the  
380 presence and evolution of SLW clouds.

381

##### 382 **4.1. Clouds**

383 As in section 3.1, the high LIDAR backscatter ( $\beta_c > 100 \beta_{mol}$ ) and low depolarization  
384 ( $< 5\%$ ) showed the presence of SLW clouds (Figures 10a, b and c, respectively). Before 13:00  
385 UTC, there is no trace of any clouds above Dome C while from 13:00 to 23:00 UTC SLW  
386 clouds are detected between 200 and 600 m. On all panels, we also superimposed the PBL top  
387 calculated by the ARPEGE-SH model. We note that the PBL top does not drop to the surface  
388 after 12:00 UTC as on 24 December 2018 but rather remains between 100 and 200 m.  
389 Consistent with the conclusions derived from the observations of 24 December 2018, the SLW  
390 cloud, once present, stays just above the height of the PBL top.

391 The cloud fraction, ice water and liquid water mixing ratios (kg kg<sup>-1</sup>) calculated by  
392 ARPEGE-SH on 20 December 2018 are shown in Figures 11a, b and c, respectively. Contrary  
393 to the observations, the model simulates mixed-phase clouds (maximum cloud fraction of



394 ~30%), mainly composed of ice, with very little liquid-water, prior to 12:00 UTC; from 00:00  
395 to 06:00 UTC, the clouds are forecasted below the PBL top. After 12:00 UTC, clouds appear  
396 1-2 hours later in the model than in the observations, at 14:00-15:00 UTC, just below the PBL  
397 top (maximum cloud fraction of ~100%). The modelled cloud is mainly composed of ice with  
398 some traces of SLW above the PBL around 15:00-16:00 UTC. We note the presence of high  
399 altitude cirrus (ice) clouds calculated by ARPEGE-SH after 12:00 UTC around 3-4 km height,  
400 while not observed likely because of a lack of LIDAR sensitivity. As on 24 December 2018,  
401 the model fails to reproduce the presence of SLW clouds observed by the LIDAR near the PBL  
402 top.

403 The diurnal evolutions of the TCI calculated by ARPEGE-SH, the LWP from  
404 HAMSTRAD and ARPEGE-SH, and the IWV from HAMSTRAD and ARPEGE-SH on 20  
405 December 2018 are presented in Figures 12a, b and c, respectively, with the presence of SLW  
406 clouds derived from the LIDAR observations superimposed on Fig. 12b. Ice clouds are  
407 calculated by ARPEGE-SH mainly around 15:00-16:00 UTC, with TCI values comparable to  
408 those on 24 December 2018. SLW clouds are deduced from HAMSTRAD LWP between 13:00  
409 and 23:00 UTC which coincides well with the SLW clouds observed by the LIDAR. LWP  
410 values observed during this episode are much higher than on 24 December 2018, with a  
411 maximum amount of  $\sim 50 \text{ g m}^{-2}$  about 20 times greater than the one measured ( $\sim 2\text{-}3 \text{ g m}^{-2}$ ) on  
412 that day. Again, the ARPEGE-SH LWP is negligible ( $\sim 10^3$  times less than observations). In  
413 parallel with the rapid increase of LWP, IWV also jumps from  $\sim 0.5$  to  $\sim 2.3 \text{ kg m}^{-2}$  within one  
414 hour after 13:00 UTC. ARPEGE-SH also calculates an increase of IWV but lagged by one hour  
415 and much less intense ( $\sim 1.3 \text{ kg m}^{-2}$ ). Additionally, the model produces a systematically dryer  
416 atmosphere compared to HAMSTRAD by about  $0.5 \text{ kg m}^{-2}$  after 16:00 UTC, although before  
417 the cloudy period that starts at 12:00 UTC, ARPEGE-SH and HAMSTRAD IWV are consistent  
418 to within  $\pm 0.2 \text{ kg m}^{-2}$ .



419

## 420 **4.2. Temperature and water vapour**

421 The diurnal variations of the temperature and water vapour anomalies on 20 December  
422 2018 as calculated by ARPEGE-SH and measured by HAMSTRAD are shown in Figure 13. In  
423 ARPEGE-SH, a sharp transition between a warm and a cool atmosphere is evident at 12:00  
424 UTC below the top of the PBL. In HAMSTRAD, from 00:00 to 06:00 UTC, the atmosphere  
425 starts warming and then from 06:00 to 13:00 UTC, cools gradually to a minimum. After 13:00  
426 UTC, HAMSTRAD temperatures reveal a warming starting from the surface and progressively  
427 thickening until reaching the top of the PBL by the end of the day. Above the PBL, the  
428 HAMSTRAD-observed and ARPEGE-SH-calculated temporal evolution of temperature and  
429 water vapour are in an overall agreement. In the PBL, the model data show a moistening around  
430 05:00 UTC, but the most striking event is a sudden drying at 12:00 UTC. In HAMSTRAD,  
431 there is a continuous drying from 00:00 UTC, followed by an obvious transition at 13:00 UTC,  
432 opposite to that of ARPEGE-SH at 12:00 UTC. The warm and wet atmosphere observed after  
433 13:00 UTC develops a mixed layer, consequently the PBL top no longer collapses to a stable  
434 layer, in contrast to what was observed on 24 December. Furthermore, the SLW clouds present  
435 in the entrainment zone steadily remain at the PBL top until the end of the day.

436

## 437 **4.3. Potential Temperature Gradient**

438 Figure 14 shows  $\partial\theta/\partial z$  ( $\text{K km}^{-1}$ ) from ARPEGE-SH, with the evolution of the PBL top and  
439 the SLW cloud superimposed. In these perturbed conditions, the SLW clouds are present a few  
440 tens of meters above the top of the PBL after 12:00 UTC. The PBL top is located in a layer  
441 coinciding with the local maximum of  $\partial\theta/\partial z$ , around 100-300 m, and does not dramatically  
442 decrease to the surface for the rest of the day.



443 Figures 15a, b and c show the vertical profiles of  $\theta$  (K) and  $\partial\theta/\partial z$  (K km<sup>-1</sup>) as calculated  
444 from temperature measured by the radiosondes and analysed by ARPEGE-SH at Dome C on  
445 20 December 2018 at 00:00 and 12:00 UTC and on 21 December 2018 at 00:00 UTC,  
446 respectively. The ARPEGE-SH profiles are about 0-5 K warmer than the observations. From  
447 50 m upward, the maximum of  $\partial\theta/\partial z$  is measured at 75, 150 and 375 m on 20 December 2018  
448 at 00:00 and 12:00 UTC and on 21 December 2018 at 00:00 UTC, respectively, with a  
449 corresponding amplitude of 75, 40 and 55 K km<sup>-1</sup>. This is broadly consistent with the ARPEGE-  
450 SH calculations on 20 December 2018 prior to the warm and wet episode: at 00:00 UTC (Fig.  
451 15a), the calculated  $\partial\theta/\partial z$  is maximum at 75 m and reaches 100 K km<sup>-1</sup>, whereas at 12:00 UTC  
452 (Fig. 15b) it peaks at 200 m (slightly higher than observed) with a value of 50 K km<sup>-1</sup>. After the  
453 warm and wet episode on 21 December 2018 at 00:00 UTC (Fig. 15c), the vertical profile of  
454  $\partial\theta/\partial z$  calculated by ARPEGE-SH shows two maxima at 100 and 450 m with an amplitude of  
455 45 and 25 K km<sup>-1</sup>, respectively, that strongly differ from the observations.

456

#### 457 **4.4. Net Surface Radiation**

458 Figure 16 shows the net surface radiation as measured by the BSRN photometric  
459 instruments and as calculated by ARPEGE-SH for 20 December 2018, superimposed with the  
460 SLW clouds. We also show the time evolution of difference in surface radiation (W m<sup>-2</sup>)  
461 observed by BSRN and calculated by ARPEGE-SH on 20 December 2018 for LW $\downarrow$ , LW $\uparrow$ ,  
462 SW $\downarrow$  and SW $\uparrow$  components, superimposed with LWP. We highlight 4 periods with snapshots  
463 taken from the webcam: 1) 07:15 UTC (clear sky), 2) 12:35 UTC (clear sky), 3) 13:30 UTC  
464 (SLW cloud) and 4) 21:00 UTC (SLW cloud). Before 13:00 UTC, there are no clouds above  
465 Dome C whilst after 13:00 UTC clouds are present. The diurnal evolution of the modelled and  
466 observed net surface radiation shows a maximum of  $\sim +50$  W m<sup>-2</sup> in ARPEGE-SH and  $\sim +85$  W  
467 m<sup>-2</sup> in BSRN over the period 00:00-04:00 UTC and a minimum of about -50 W m<sup>-2</sup> around



468 12:00-13:00 UTC on both time series. Nevertheless, when SLW clouds are observed at 13:00  
469 UTC, the observed net surface radiation jumps to  $+10 \text{ W m}^{-2}$ , a feature not reproduced in the  
470 model. The difference between the BSRN-observed and ARPEGE-SH-modelled net surface  
471 radiation is larger than  $+30 \text{ W m}^{-2}$  when SLW clouds are present, reaching  $+60 \text{ W m}^{-2}$  when the  
472 LWP measured by HAMSTRAD is at its maximum ( $50 \text{ g m}^{-2}$  at 13:00 UTC). This is twice the  
473 difference observed in the non-perturbed PBL episode detailed in section 3.4. This underlines  
474 again the strong impact SLW clouds may have on the radiation budget over Antarctica. In the  
475 presence of SLW clouds after 13:00 UTC, the difference in net surface radiation is mainly  
476 attributable to  $\text{LW}\downarrow$  component, BSRN values being  $100 \text{ W m}^{-2}$  greater than those of ARPEGE-  
477 SH. Note that there are differences from  $-100$  to  $+60 \text{ W m}^{-2}$  between observed and calculated  
478  $\text{SW}\downarrow$  and  $\text{SW}\uparrow$  components but this difference falls below  $20 \text{ W m}^{-2}$  for the net SW surface  
479 radiation ( $\text{SW}\downarrow - \text{SW}\uparrow$ ).

480

## 481 **5. Conclusions**

482 A comprehensive water budget study has been performed during the YOPP international  
483 campaign held at Dome C (Concordia, Antarctica) from mid-November 2018 to mid-February  
484 2019. Supercooled liquid water (SLW) clouds were observed and analysed by means of remote-  
485 sensing ground-based instrumentation (tropospheric depolarization LIDAR, HAMSTRAD  
486 microwave radiometer, BSRN net surface radiation), radiosondes, spaceborne sensor  
487 (CALIOP/CALIPSO depolarization LIDAR) and the NWP ARPEGE-SH. The analyses  
488 focused on two periods showing 1) a “typical” diurnal cycle of the PBL on 24 December 2018  
489 (warm and dry, local mixing layer followed by a thinner cold and dry, local stable layer which  
490 develops when the surface has cooled down) and 2) a perturbed diurnal cycle of the PBL on 20  
491 December 2018 (a warm and wet episode prevented from a clear diurnal cycle of the PBL top).



492        Whatever the state of the diurnal cycle of the PBL top (typical or perturbed), thin (~100-m  
493 thick) SLW clouds have been observed by ground-based and spaceborne LIDARs developing  
494 within the entrainment and the capping inversion zones at the top of the PBL. Spaceborne lidar  
495 observations revealed horizontal extensions of these clouds as large as 280 and 550 km for the  
496 24 and 20 December cases, respectively. ARPEGE-SH was not able to correctly estimate the  
497 ratio between liquid and solid water inside the cloudy layers, with SLW always strongly  
498 underestimated in the studied cases. Consequently, the net surface radiation was affected by the  
499 presence of SLW clouds during these two episodes. The net surface radiation observed by  
500 BSRN was 20-30 W m<sup>-2</sup> higher than that modelled in ARPEGE-SH on 24 December 2018  
501 (typical diurnal cycle of the PBL), this difference reaching +50 W m<sup>-2</sup> on 20 December 2018  
502 (perturbed diurnal cycle of the PBL), consistent with the total observed liquid water being 20  
503 times greater in the perturbed PBL diurnal cycle than in the typical PBL diurnal cycle. The  
504 difference in the net surface radiation is mainly attributable to longwave downward surface  
505 radiation, BSRN values being 50 and 100 W m<sup>-2</sup> greater than those of ARPEGE-SH in the  
506 typical and perturbed cases, respectively.

507        Time coincident ground-based remote-sensed measurements of water (vapour, liquid and  
508 solid), temperature and net surface radiation are available at Dome C since 2015. Consequently,  
509 a comprehensive statistical analysis of the presence of SLW clouds will be performed in the  
510 near future. Coupled with modelling studies (NWP ARPEGE-SH, mesoscale models), an  
511 estimation of the radiative impact of these clouds on the local climate will then be performed.

512

### 513 **Data availability**

514        HAMSTRAD data are available at <http://www.cnrm.meteo.fr/spip.php?article961&lang=en>  
515 (last access: 28 August 2019). The CALIOP images are accessible at [http://www-  
516 calipso.larc.nasa.gov/](http://www-calipso.larc.nasa.gov/) (last access: 28 August 2019). The tropospheric depolarization LIDAR



517 data are reachable at [http://lidarmax.altervista.org/englidar/\\_Antarctic%20LIDAR.php](http://lidarmax.altervista.org/englidar/_Antarctic%20LIDAR.php) (last  
518 access: 28 August 2019). Radiosondes are available at <http://www.climantartide.it> (last access:  
519 28 August 2019). BSRN data can be obtained from the ftp server ([https://bsrn.awi.de/data/data-  
520 retrieval-via-ftp/](https://bsrn.awi.de/data/data-retrieval-via-ftp/)) (last access: 28 August 2019). The ARPEGE data and corresponding  
521 technical information are available from the YOPP Data Portal and from the ftp server (ftp.umr-  
522 cnrm.fr with user: yopp and password: Arpege) (last access: 28 August 2019).

523

#### 524 **Author contributions**

525 PR, MDG, AL, and PG provided the observational data while EB, NA and VG developed  
526 the model code and performed the simulations. PD, JLA and DV contributed to the data  
527 interpretation. All the co-authors participated in the data analysis. PR prepared the manuscript  
528 with contributions from all co-authors.

529

#### 530 **Competing interests**

531 The authors declare that they have no conflict of interest.

532

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545

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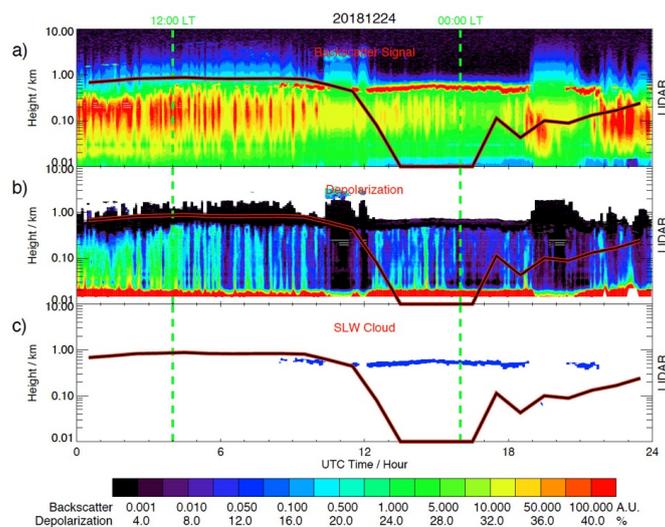


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628

## Figures



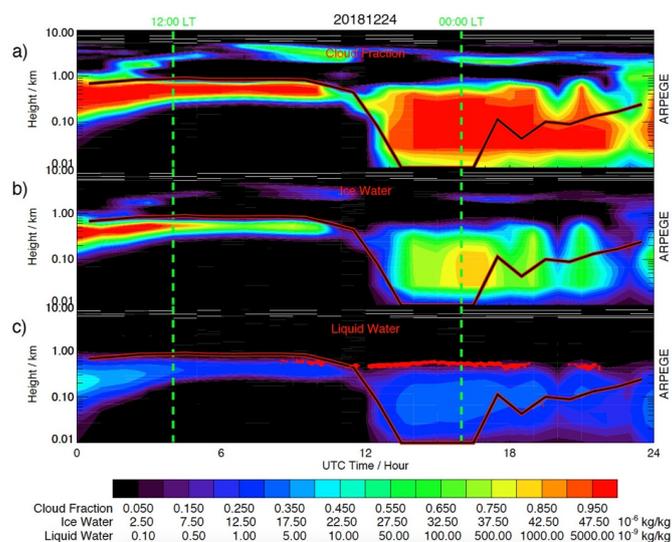
629

630 **Figure 1:** Diurnal variation on 24 December 2018 (UTC Time) along the vertical of: a) the  
631 backscatter signal (Arbitrary Unit, A.U.), b) the depolarization ratio (%) measured by the  
632 aerosol LIDAR, and c) the Supercooled Liquid Water (SLW) cloud height (blue) deduced from  
633 the aerosol LIDAR ( $\beta_c > 100 \beta_{mol}$ , depolarization  $< 5\%$ ). Superimposed to all the Figures is the  
634 top of the Planetary Boundary Layer calculated by the ARPEGE-SH model (black-red thick  
635 line). Two vertical green dashed lines indicate 12:00 and 00:00 LT.

636



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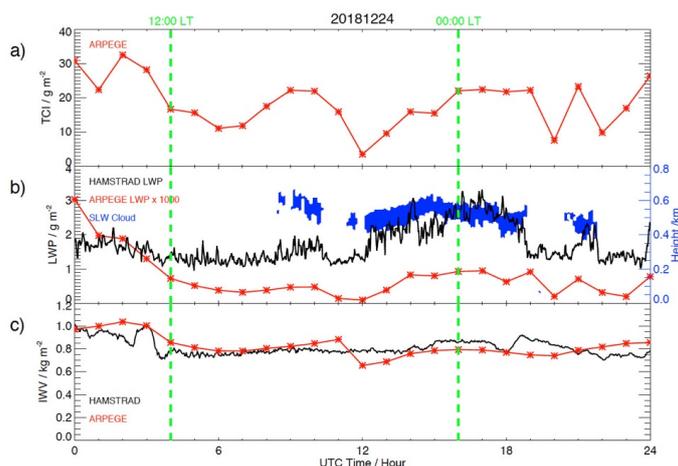


638

639 **Figure 2:** Time-height cross section on 24 December 2018 (UTC Time) of: a) the Cloud  
640 Fraction (0-1), b) the Ice Water mixing ratio ( $10^{-6}$  kg  $\text{kg}^{-1}$ ) and c) the Liquid Water mixing ratio  
641 ( $10^{-6}$  kg  $\text{kg}^{-1}$ ) calculated by the ARPEGE-SH model. Superimposed to all the panels is the top  
642 of the Planetary Boundary Layer calculated by the ARPEGE-SH model (black-red thick line).  
643 Superimposed in panel c is the SLW cloud (red area) height depth deduced from the LIDAR  
644 observations (see Fig. 1c). Two vertical green dashed lines indicate 12:00 and 00:00 LT.

645

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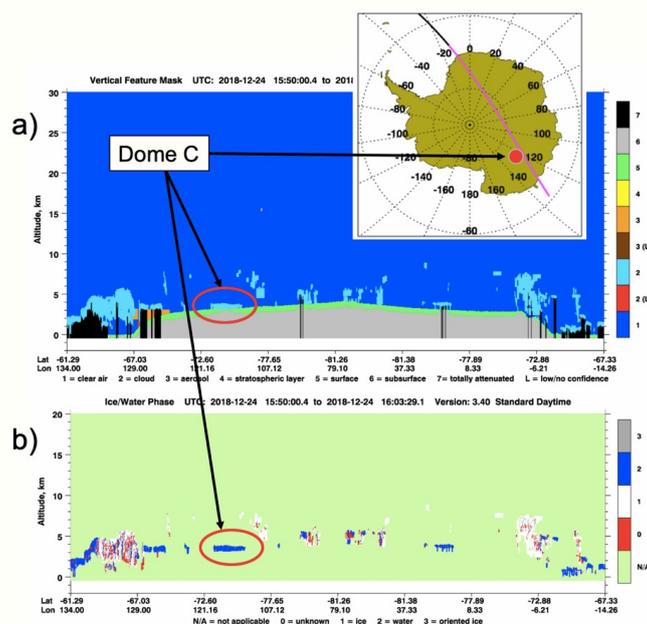


647

648 **Figure 3:** Diurnal variation on 24 December 2018 (UTC Time) of: a) the Total Column of Ice  
649 (TCI) ( $\text{g m}^{-2}$ ) calculated by ARPEGE-SH (red crossed line), b) the Liquid Water Path (LWP)  
650 measured by HAMSTRAD ( $\text{g m}^{-2}$ , black solid line) and calculated by ARPEGE-SH ( $\times 1000 \text{ g}$   
651  $\text{m}^{-2}$ , red crossed line) and c) the Integrated Water Vapour (IWV,  $\text{kg m}^{-2}$ ) measured by  
652 HAMSTRAD (black solid line) and calculated by ARPEGE-SH (red crossed line).  
653 Superimposed to panel b) is the SLW cloud thickness (blue area) deduced from the LIDAR  
654 observations (see Fig. 1c) (blue y-axis on the right of the Figure). Note LWP from ARPEGE-  
655 SH has been multiplied by a factor 1000. Two vertical green dashed lines indicate 12:00 and  
656 00:00 LT.

657

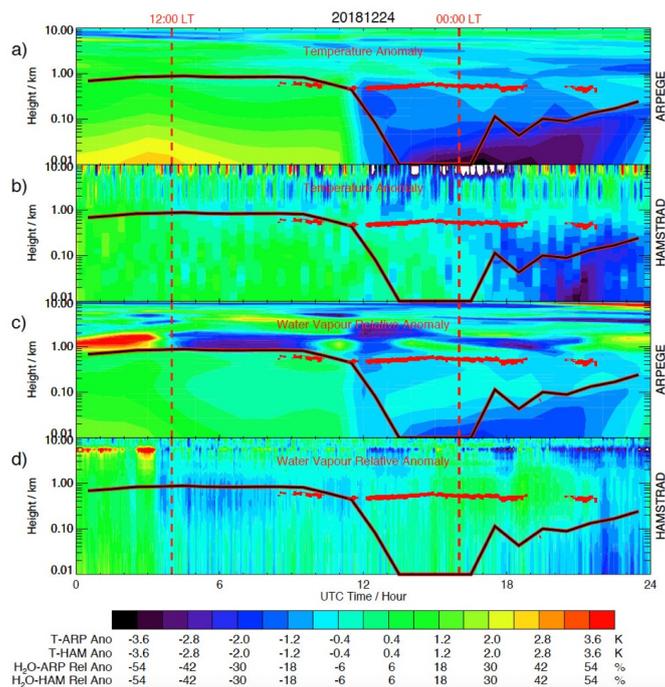
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659

660 **Figure 4:** CALIOP/CALIPSO spaceborne LIDAR observations along one orbit on 24  
661 December 2018 (15:50-16:00 UTC) in the vicinity of Dome C (75°S, 123°E): a) the Vertical  
662 Feature Mask highlighting a cloud (light blue) near the surface (red circle) and b) the Ice/Water  
663 Phase Mask highlighting a SLW (dark blue) cloud near the surface (red circle). The ground-  
664 track of the sensor (pink) has been embedded at the top of the Figure, with the location of Dome  
665 C marked (red filled circle). Note that the altitude is relative to the sea surface, with the height  
666 of surface of Dome C at an elevation of 3233 m amsl.

667

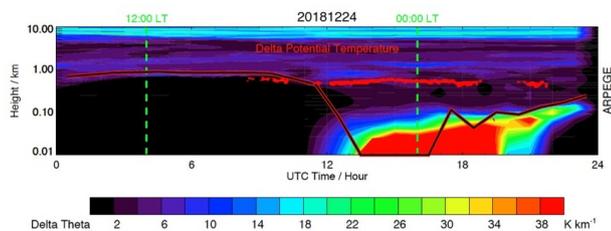


668

669 **Figure 5:** Time-height cross section on 24 December 2018 (UTC Time) of a) the temperature  
 670 anomaly (K) and b) the water vapour relative anomaly (%) calculated by ARPEGE-SH, c) the  
 671 temperature anomaly (K) and d) the water vapour relative anomaly (%) observed by  
 672 HAMSTRAD. Superimposed to all the Figures are the SLW cloud altitude (red area) deduced  
 673 from the LIDAR observations (see Fig. 1c) and the top of the Planetary Boundary Layer  
 674 calculated by the ARPEGE-SH model (black-red thick line). Two vertical red dashed lines  
 675 indicate 12:00 and 00:00 LT.

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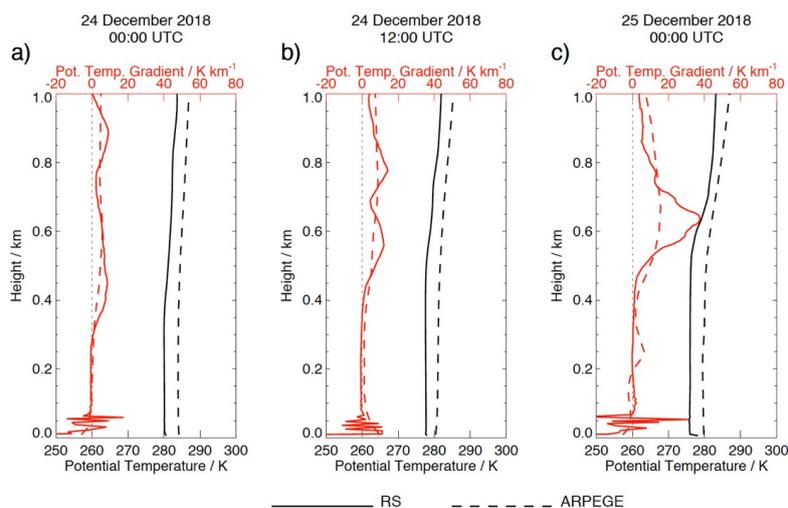
679 **Figure 6:** Time-height cross section of  $\partial\theta/\partial z$  ( $\text{K km}^{-1}$ ) calculated from ARPEGE-SH  
680 temperature on 24 December 2018 (UTC Time). Superimposed are the SLW cloud altitude (red  
681 area) deduced from the LIDAR observations (see Fig. 1) and the top of the Planetary Boundary  
682 Layer calculated by the ARPEGE-SH model (black-red thick line). Two vertical green dashed  
683 lines indicate 12:00 and 00:00 LT.

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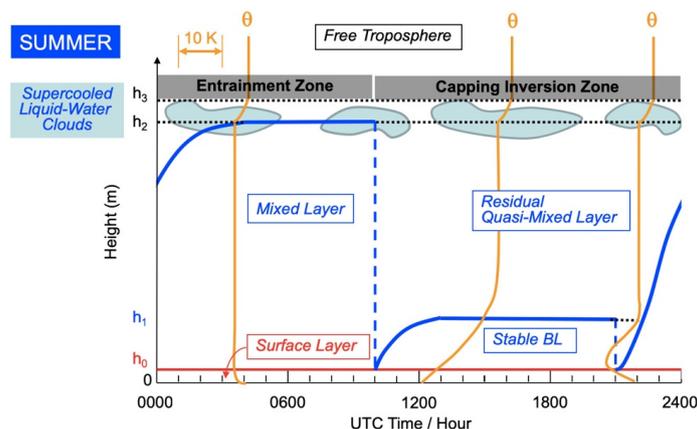
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687

688 **Figure 7:** Vertical profiles of potential temperature  $\theta$  (black) and the gradient in potential  
689 temperature  $\partial\theta/\partial z$  (red) as calculated from temperature measured by the radiosondes (solid line)  
690 and analysed by ARPEGE-SH (dashed line) at Dome C on 24 December 2018 at a) 00:00 and  
691 b) 12:00 UTC, and c) on 25 December 2018 at 00:00 UTC.

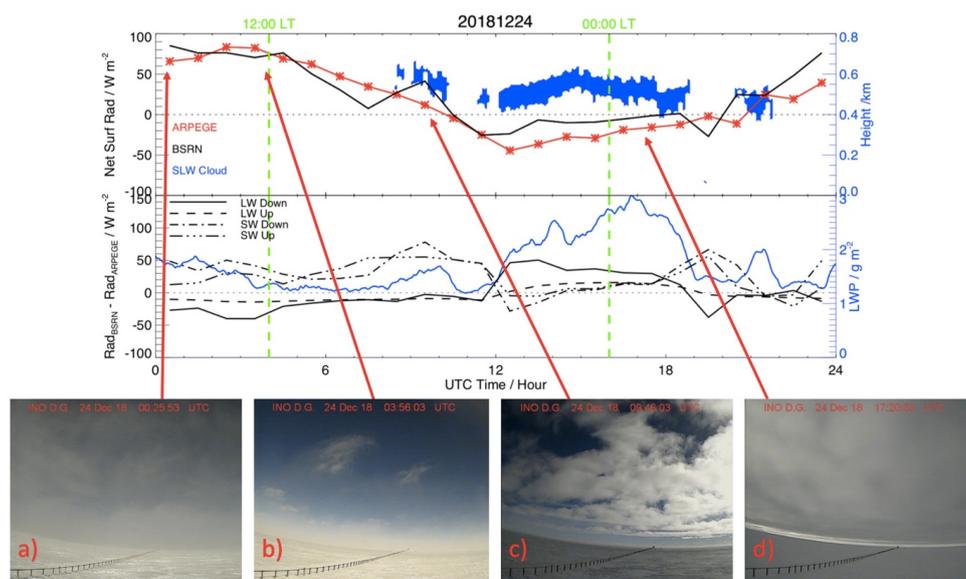
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694 **Figure 8:** Figure modified and updated from Fig. 12 of Ricaud et al. (2012) showing the diurnal  
695 evolution (UTC Time) of the different layers in the Planetary Boundary Layer (PBL) with  $h_0$   
696 the top of the surface layer,  $h_2$  the daily overall top of the PBL, and  $h_1$  the top of the intermediate  
697 stable layer within the PBL. The orange lines symbolize the vertical profiles of potential  
698 temperature  $\theta$ . The layer between  $h_2$  and  $h_3$  is named “entrainment zone” above the mixed layer  
699 and “capping inversion zone” elsewhere.  $h_3$  represents the top of this layer. SLW clouds are  
700 represented within this layer. Note that  $LT = UTC + 8 \text{ h}$ .

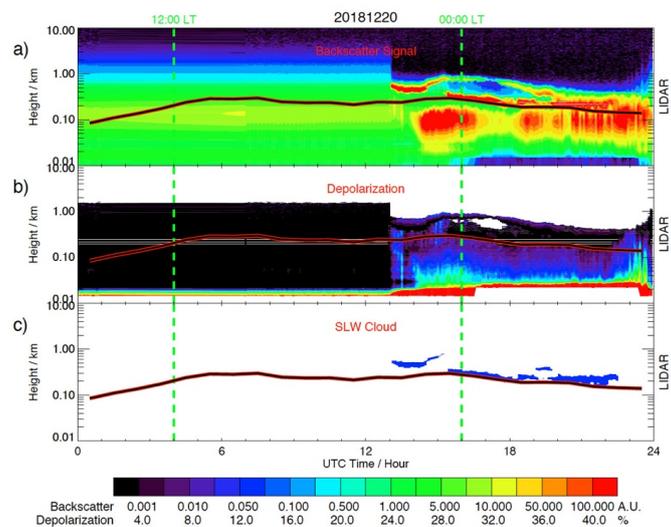
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703 **Figure 9:** (Top) Diurnal variation of the net surface radiation ( $\text{W m}^{-2}$ ) observed by BSRN (black  
704 solid line) and calculated by ARPEGE-SH (red crossed line) on 24 December 2018 in UTC  
705 Time. Superimposed is the SLW cloud altitude (blue) deduced from the LIDAR. (Middle)  
706 Diurnal variation of the difference between surface radiation ( $\text{W m}^{-2}$ ) observed by BSRN and  
707 calculated by ARPEGE-SH on 24 December 2018 for longwave downward (black solid),  
708 longwave upward (black dashed), shortwave downward (black dashed dotted) and shortwave  
709 upward (black dashed triple dotted) components. Superimposed is LWP (blue) measured by  
710 HAMSTRAD. (Bottom) Four webcam images showing the cloud coverage at: a) 00:25 UTC  
711 and b) 03:56 UTC (cirrus clouds, no SLW cloud), c) 09:46 UTC (SLW cloud) and d) 17:20  
712 UTC (SLW cloud). Two vertical green dashed lines indicate 12:00 and 00:00 LT.

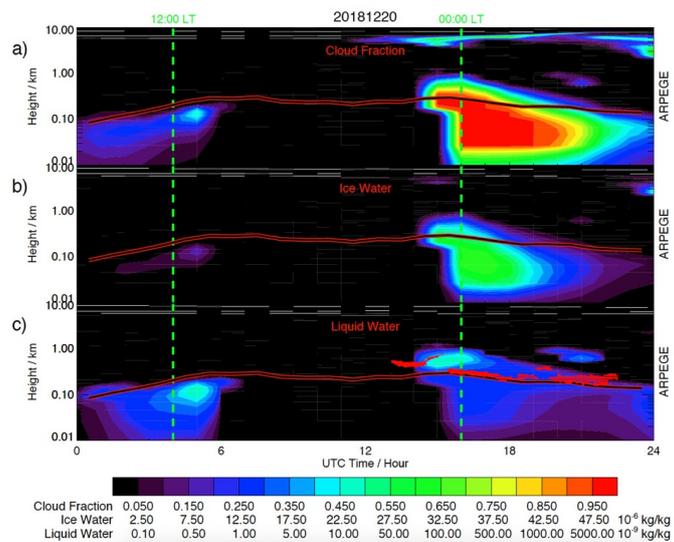
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715 **Figure 10:** Same as Figure 1 but for 20 December 2018.

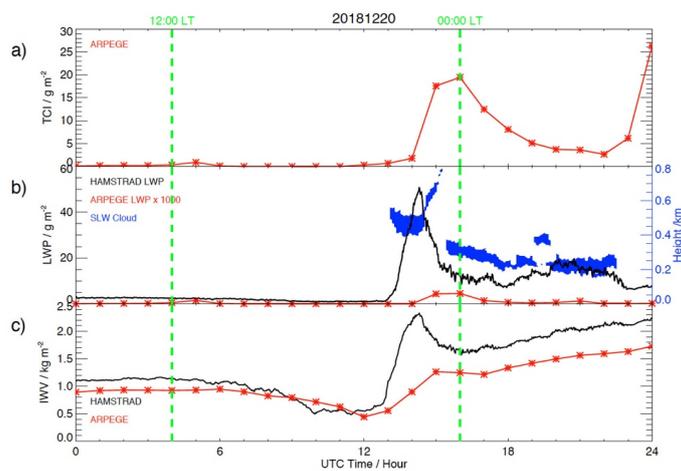
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718 **Figure 11:** Same as Figure 2 but for 20 December 2018.

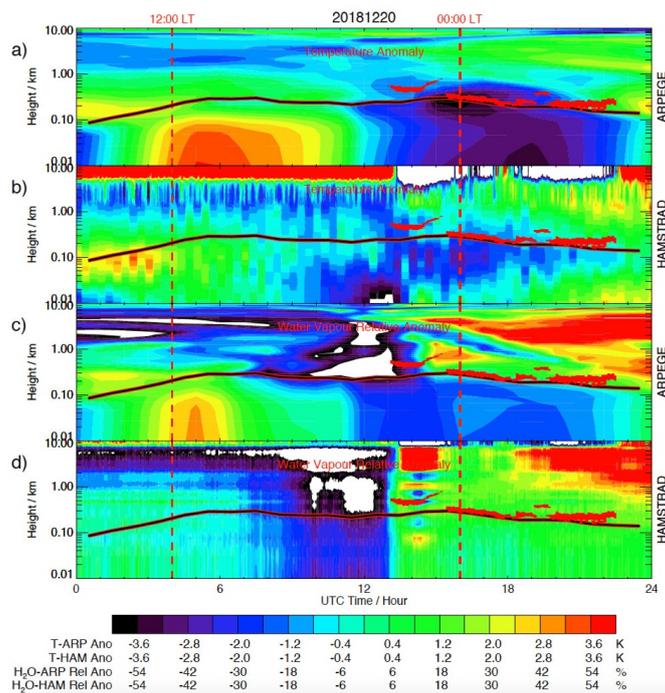
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720

721 **Figure 12:** Same as Figure 3 but for 20 December 2018.

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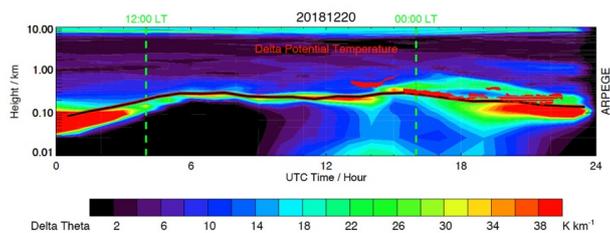
724 **Figure 13:** Same as Figure 5 but for 20 December 2018.

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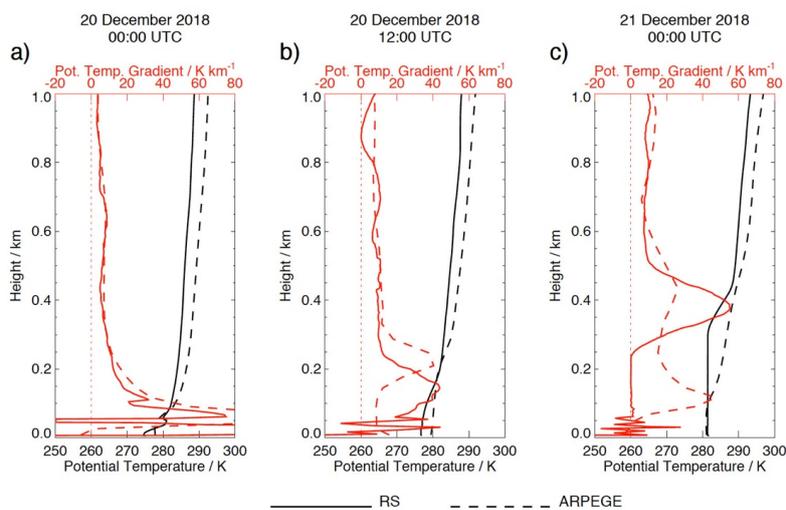


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729 **Figure 14:** Same as Figure 6 but for 20 December 2018.

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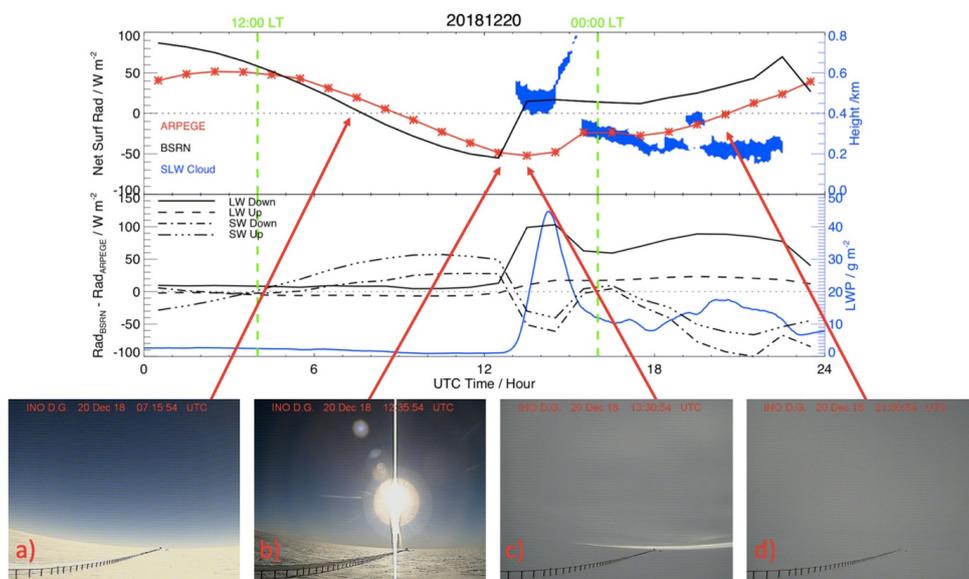
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733 **Figure 15:** Same as Figure 7 but on 20 December 2018 at a) 00:00 and b) 12:00 UTC, and c)

734 on 21 December 2018 at 00:00 UTC.

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738 **Figure 16:** Same as Figure 9 but for 20 December 2018 whilst the 4 webcam images were  
739 selected at: a) 07:15 and b) 12:35 UTC (clear sky), c) 13:30 and d) 21:00 UTC (SLW cloud).

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