

1 **LIMS observations of lower stratospheric ozone in the southern polar springtime of 1978**

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24 **Abstract**

25 The Nimbus 7 limb infrared monitor of the stratosphere (LIMS) instrument operated from  
26 October 25, 1978, through May 28, 1979. This note focuses on its Version (V6) data and  
27 indications of ozone loss in the lower stratosphere of the southern hemisphere, subpolar region  
28 during the last week of October 1978. We provide profiles and maps that show V6 ozone values  
29 of only 2 to 3 ppmv at 46 hPa within the edge of the polar vortex near 60°S from late October  
30 through mid-November 1978. There are also low values of V6 nitric acid (~3 to 6 ppbv) and  
31 nitrogen dioxide (<1 ppbv) at the same locations, indicating that conditions were suitable for a  
32 chemical loss of Antarctic ozone some weeks earlier. These “first light” LIMS observations  
33 provide the earliest, space-based view of conditions within the lower stratospheric ozone layer of  
34 the southern polar region in springtime.

35

36 **1 Introduction and historical context**

37 The Nimbus 7 Total Ozone Mapping Spectrometer (TOMS) provided the first daily image of  
38 total ozone for the Southern Hemisphere (SH) on November 1, 1978. That image in Figure 1  
39 shows an equatorward extension of the region of low polar, total column ozone (TCO) between  
40 90°E and 135°E. Minimum TCO is of the order of 270 Dobson units (DU) at (75°S, 90°E) on  
41 this day. As a comparison, Farman et al. (1985) reported ground-based measurements of total  
42 ozone of about 225 DU on November 1 for 1980-1984 at Halley Bay (76°S, 333°E) and of about  
43 270 DU at Argentine Islands (65°S, 296°E) (see also TOMS total ozone values of Table 2 in  
44 Stolarski et al. (1986)). We note, however, that those values are higher than 220 DU, which is a  
45 threshold definition for “ozone hole” conditions (WMO, 2018).

46

47 There are very few observations of lower stratospheric ozone above Antarctica prior to  
48 November 1978, especially for the months of September and October when the seasonal loss of  
49 ozone is most significant (WMO, 2018). The historic Nimbus 7 Limb Infrared Monitor of the  
50 Stratosphere (LIMS) experiment (Gille and Russell, 1984) provided data for middle atmosphere  
51 temperature, geopotential height (GPH), ozone, water vapor (H<sub>2</sub>O), nitric acid vapor (HNO<sub>3</sub>),  
52 and nitrogen dioxide (NO<sub>2</sub>) from October 25, 1978, through May 28, 1979, for scientific analysis

53 and for comparisons with atmospheric models (e.g., Langematz et al., 2016). Remsberg et al.  
54 (2007) provide a description of its Version 6 (V6) ozone profiles. The mapping of the V6  
55 profiles to the LIMS Level 3 product employs a sequential estimation algorithm with a relaxation  
56 time of about 2.5 days for analyses of its zonal, 6-wavenumber Fourier coefficients at each of 28  
57 pressure levels of the middle atmosphere (Remsberg and Lingenfelter, 2010). We then  
58 generated daily, polar stereographic plots of V6 ozone and HNO<sub>3</sub> on pressure surfaces based on a  
59 gridding (2° latitude and 5.625° longitude) from those coefficients.

60

61 This note focuses on the character of the polar vortex and of the V6 ozone, HNO<sub>3</sub>, and NO<sub>2</sub> in  
62 that region of the lower stratosphere during the last week of October 1978. The LIMS  
63 measurements extend to only 64°S, due to the orbital inclination of Nimbus 7 and to the viewing  
64 geometry of the LIMS instrument (Gille and Russell, 1984). We will show that the profiles and  
65 pressure surface maps indicate that there was a loss of SH polar ozone during the springtime.  
66 Section 2 contains plots that show a loss of ozone inside the vortex in late October. Section 3  
67 reports on evidence for a denitrification of the air in the same region, indicating that there was a  
68 chemical loss of ozone some weeks earlier. Section 3 also presents time versus longitude or  
69 Hovmöller diagrams that reveal good correspondence for the low ozone and HNO<sub>3</sub> values within  
70 the vortex region well into November. Section 4 summarizes the findings.

71

## 72 **2 Antarctic ozone from late October to early November 1978**

73 Figure 2 shows SH polar plots of V6 ozone mixing ratios at 46.4 hPa for October 26 and for  
74 November 1, where the orbital measurements of LIMS extend only to 64°S. The plot at right  
75 shows that there are minimum ozone values of about 2.6 ppmv near 120°E and 315°E at 60°S on  
76 November 1, which agrees reasonably with the locations of low total ozone from the TOMS  
77 image of Fig. 1. Ozone is of order 3.5 to 4 ppmv at most other longitudes. Low ozone occurs  
78 within the edges of the polar vortex, based on the concurrent GPH field from the operational  
79 ECMWF Re-Analysis or ERA-40 products (Uppala et al., 2005). The bold contour in Fig. 2  
80 denotes the edge of the vortex, in the manner of Harvey et al. (2002). We define the vortex edge  
81 as the streamfunction contour coincident with maximum wind speed that also encloses a region

82 of rotation. Meek et al. (2017) showed that this definition of the vortex edge is in good  
83 agreement with the PV-gradient based definition of Nash et al. (1996). We note that daily plots  
84 of GPH are also available from LIMS V6. However, they exhibit a discontinuous anomaly at the  
85 46-hPa level for the vortex region between October 29 and 31, due to an interpolation of  
86 National Meteorological Center (NMC) GPH analyses supplied to the Nimbus 7 Project and used  
87 for the baseline pressure level of 50 hPa for the V6 GPH product (Remsberg et al., 2004). V6  
88 geometric height and GPH profiles above and below that level are the result of a hydrostatic  
89 integration of the LIMS-retrieved temperature versus pressure profiles of  $T(p)$ . Maps of V6  
90 GPS farther away from the 50-hPa level are very similar to those from ERA-40.

91

92 LIMS began its daily observations one week earlier than TOMS or on October 25, and the left  
93 plot of Fig. 2 shows that the ozone for October 26 at 31°E is about half of that at 119°E on  
94 November 1. The vortex on October 26 extends toward lower latitudes from about 60°S, 40°E.  
95 Both the vortex and region of low ozone deform and undergo a clockwise rotation from October  
96 26 onward, such that their low values extend equatorward at 120°E and at 315°E on November  
97 1. Bodeker et al. (2002) reported that the edge of the vortex often extends to near 60°S during  
98 October, and Stolarski et al (1986, their Fig. 1) and Hassler et al. (2011) reported on an  
99 analogous clockwise rotation of the vortex during October.

100

### 101 **3 Findings of denitrification of the vortex air in late October**

102 The location of the vortex edge is helpful in deciding which V6 species profiles one ought to  
103 examine with regard to any constraints from  $\text{HNO}_3$  and  $\text{NO}_2$  on the ozone chemistry. As an  
104 example, Fig. 3 shows V6 Level 2 ozone profile segments from 11.4 to 88 hPa for two locations  
105 on October 26, where ozone is now presented in units of partial pressure (in mPa) for a better  
106 delineation of its relative changes in the subpolar lower stratosphere. Estimates of accuracy for  
107 single V6 ozone profiles are 14%, 26%, and 34% for 10 hPa, 50 hPa, and 100 hPa, respectively  
108 (see row (g) of Table 1 in Remsberg et al., 2007). The V6 ozone profile (black solid) at 54.9°S,  
109 119°E is just outside the October 26 vortex, as shown by the black dot in Fig. 2, and its ozone  
110 values are nominal for subpolar latitudes. The largest contribution to total ozone from that

111 profile in Fig. 3 occurs at the 68-hPa level. A second V6 ozone profile (solid red) is from  
112 59.5S°, 31°E, and it is in a region of lower GPH as shown by the red dot in Fig. 2. Its ozone  
113 decreases rapidly from ~8.0 mPa at the 53-hPa level to 2.6 mPa at the 88-hPa level, indicating a  
114 significant loss of ozone in the lower stratosphere sometime prior to October 26. Komhyr et al.  
115 (1988, their Fig. 10) and Gerandt (1987) show from ozonesonde measurements that most of the  
116 observed losses of ozone for the mid-1980s occurred in the vortex in September and early  
117 October. Therefore, we also include in Fig. 3 an ozonesonde profile (solid green) from Syowa  
118 station (69°S, 40°E—the green dot in Fig. 2) for September 3, 1978, perhaps before there were  
119 any pronounced losses of ozone. Its ozone profile values are intermediate of those for the two  
120 V6 profiles of October 26.

121  
122 Loss of ozone due to reactive chlorine chemistry proceeds effectively in the presence of air that  
123 has undergone denitrification (Müller et al., 2008). Lambert et al. (2016) somewhat loosely set  
124 an HNO<sub>3</sub> threshold of < 5 ppbv for indicating denitrification at 46 hPa, based on Microwave  
125 Limb Sounder (MLS) data of 2008. Nitrous oxide is the source molecule for odd nitrogen  
126 (mainly HNO<sub>3</sub>) in the lower stratosphere, and its tropospheric values have grown by only a small  
127 amount from 1975 (~296 ppbv) through 2008 (~322 ppbv) (WMO, 2018); the HNO<sub>3</sub> threshold of  
128 5 ppbv should also be representative for 1978. Thus, in Figure 3 we also show the  
129 accompanying V6 profiles of HNO<sub>3</sub> and nighttime NO<sub>2</sub> for the same two locations on October  
130 26. HNO<sub>3</sub> and NO<sub>2</sub> at 31°E are a half (or 3 ppbv) and a third (or < 1 ppbv), respectively, of  
131 those at 119°E below about the 31-hPa level. Thus, both species indicate that there was a  
132 denitrification of the air in the vortex region and a likely loss of ozone due to reactive chlorine  
133 chemistry in the presence of polar stratospheric clouds (PSCs) several weeks earlier (Solomon,  
134 1999; WMO, 2018). Although the V6 temperature at 31°E on October 26 was 206 K (at 53  
135 hPa), it is normal to find temperatures in the Antarctic vortex that are below the chlorine  
136 activation threshold value of 195 K and in the presence of PSCs during September and early  
137 October (WMO, 2018).

138  
139 Figure 4 shows the corresponding V6 plots of HNO<sub>3</sub> at 46 hPa in terms of its mixing ratios,  
140 which have an estimated accuracy of ~9% (Remsberg et al., 2010, Table 10). There are very low

141 values of HNO<sub>3</sub> on October 26 poleward of 60°S and from 31°E to at least 90°E, indicating an  
142 earlier conversion of HNO<sub>3</sub> from vapor to condensed phase and the sedimentation of larger  
143 HNO<sub>3</sub> containing particles rather than an advection of low HNO<sub>3</sub> from lower latitudes. Low  
144 HNO<sub>3</sub> mixing ratios are also present within the vortex region on November 1. Analogous polar  
145 plots of the nighttime NO<sub>2</sub> fields are quite noisy (not shown) due to the large uncertainties for  
146 tangent layer NO<sub>2</sub> in the lower stratosphere. Nevertheless, most of the odd nitrogen reservoir at  
147 46 hPa comes from HNO<sub>3</sub>, not NO<sub>2</sub>. Together, they indicate the extent of denitrification of the  
148 air in the vortex region during late October 1978.

149  
150 We show in Figs. 5 and 6 the details of the changing ozone and nitric acid from late October  
151 through November. Figure 5 displays time/longitude or Hovmöller diagrams for both species at  
152 60°S; thick black contours indicate the vortex edge and dotted horizontal lines the vortex  
153 interior. The occurrence of lowest species mixing ratios shows clearly in the vortex region in  
154 late October. Figure 6 extends the findings of Fig. 5 through the end of November, and there is  
155 an eastward progression of the region of low values from late October to early November.  
156 Reduced mixing ratios of those species occur inside the vortex until about November 25, as  
157 expected for chemicals that are tracers of air motions in the lower stratosphere. The vortex  
158 distorts and then exhibits a stationary wave-1 pattern from November 5 onward, where height is  
159 lowest near 0°E. Mixing of air across the vortex edge appears slow for both ozone and HNO<sub>3</sub>  
160 during that time.

161

#### 162 **4 Summary and concluding remarks**

163 We find low V6 ozone mixing ratios of order 2 to 3 ppmv at 60°S within the edge of the polar  
164 vortex at 46 hPa during the last week of October and well into November 1978. There is good  
165 agreement between the V6 ozone map at 46 hPa and the TOMS image of total ozone in the  
166 region of the vortex on November 1. Low V6 HNO<sub>3</sub> mixing ratios of order 3 to 6 ppbv at the  
167 same locations indicate denitrification and conditions that were suitable for a chemical loss of  
168 Antarctic ozone some weeks earlier. We note that equivalent effective stratospheric chlorine  
169 (EESC) values used to predict conditions for the depletion of ozone in 1980 are about twice

170 those of 1950, while the 1980 values are only half those of 2000 (Newman et al., 2007). In  
171 hindsight and based on the LIMS V6 dataset, we conclude that there was very likely some  
172 halogen-catalyzed loss of ozone in the southern polar vortex in winter/spring of 1978. Yet, those  
173 ozone losses in the SH spring were not to the low level of a true “ozone hole” (<220 DU total  
174 ozone). We also conclude that the LIMS V6 Level 2 profiles and the daily-analyzed maps from  
175 their Level 3 zonal coefficients represent useful comparison data for model simulations of the  
176 changes in Antarctic ozone in spring 1978.

177

178 **Data Availability**

179 The LIMS V6 data archive is at the NASA EARTHDATA site of EOSDIS and its website:  
180 <https://search.earthdata.nasa.gov/search?q=LIMS> ). Nimbus 7 TOMS ozone is at  
181 [https://disc.gsfc.nasa.gov/datacollection/TOMSN7L2\\_008.html](https://disc.gsfc.nasa.gov/datacollection/TOMSN7L2_008.html). ECC ozonesonde ozone  
182 profiles are available from the World Ozone and Ultraviolet Radiation Data Centre or WOUDC  
183 at <https://woudc.org/data/explore.php>. ECMWF Re-Analysis (ERA-40) data are accessible  
184 through <https://climatedataguide.ucar.edu/climate-data/era40>.

185

186 *Author Contributions.* ER and VLH wrote the manuscript and prepared the figures with input  
187 from all the other co-authors. AK provided information about the TOMS ozone images. LG led  
188 the development of the LIMS version 6 algorithms. JCG and JMR are the Co-Principal  
189 Investigators of the LIMS experiment. They also commented on the new insight from the  
190 findings about ozone and nitric acid of October 1978.

191

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195 NASA Langley.

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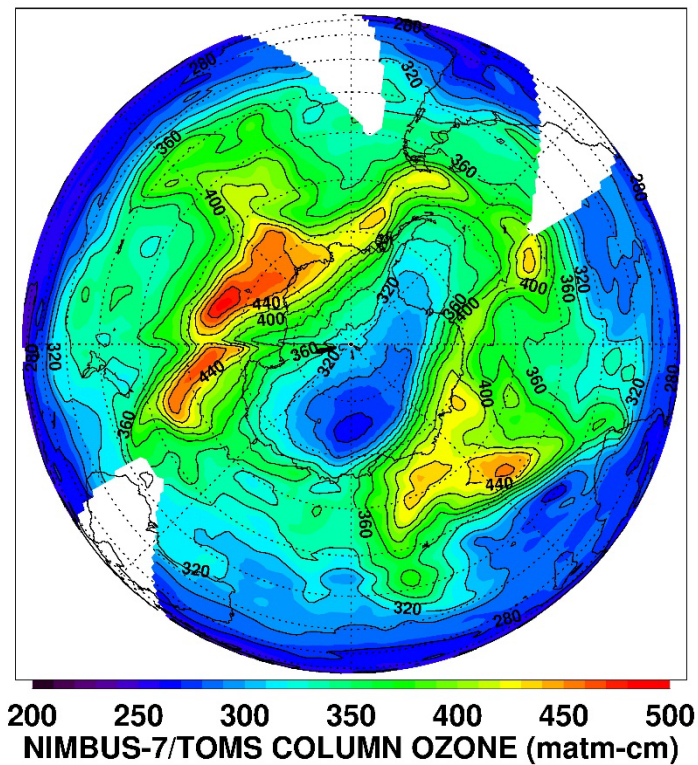
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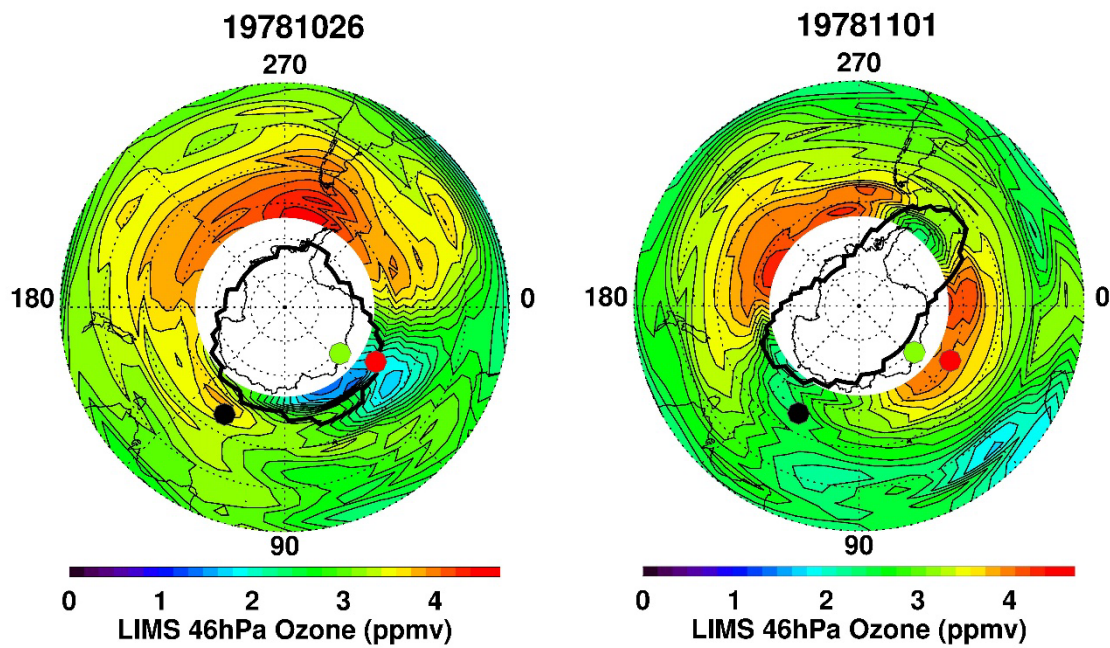
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285 Figure 1—Southern Hemisphere image of total column ozone (TCO) from TOMS for November  
286 1, 1978. Longitude orientation is 0°E to the right and 90°E at the bottom; latitude circles  
287 (dotted) have a spacing of 10 degrees. White areas indicate where there are discrete data voids  
288 or no measurements. Ozone units of matm-cm are equivalent to Dobson units (DU), where 1 DU  
289 is  $2.687 \times 10^{20}$  molecules-m<sup>-2</sup>. Black contours are TCO at intervals of 20 matm-cm.

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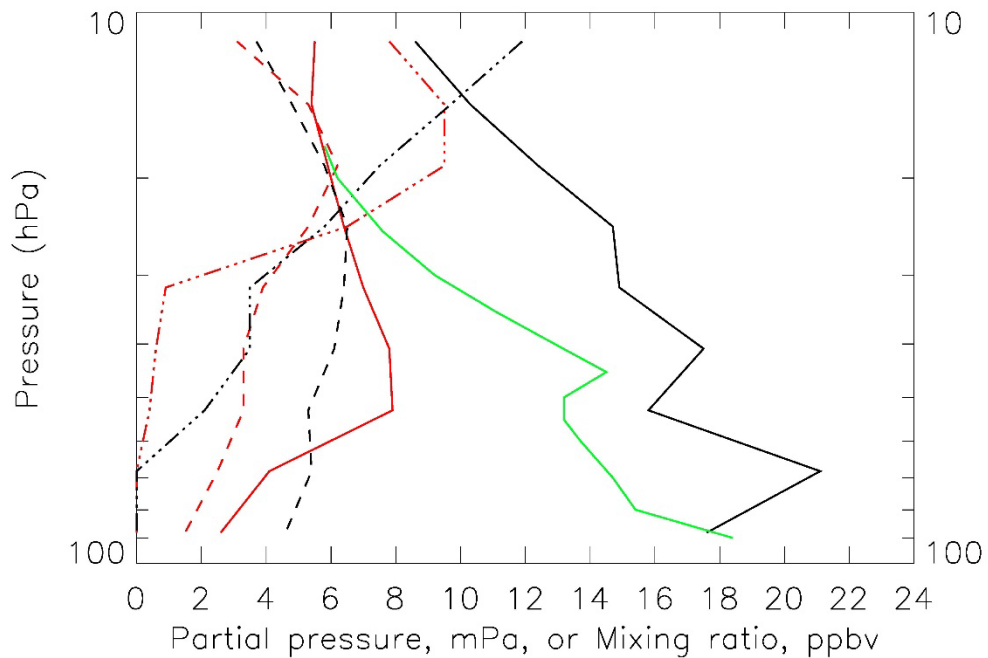


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293 Figure 2—V6 ozone mixing ratios at 46.4 hPa for October 26 and November 1, 1978. Polar  
294 plots extend from 30°S to the Pole and longitude is in °E with 0° at right. Bold contours denote  
295 the vortex edge from ERA-40. The superposed, three colored dots correspond to the locations of  
296 profiles on October 26 (black and red) and on September 3 (green) in Fig. 3.

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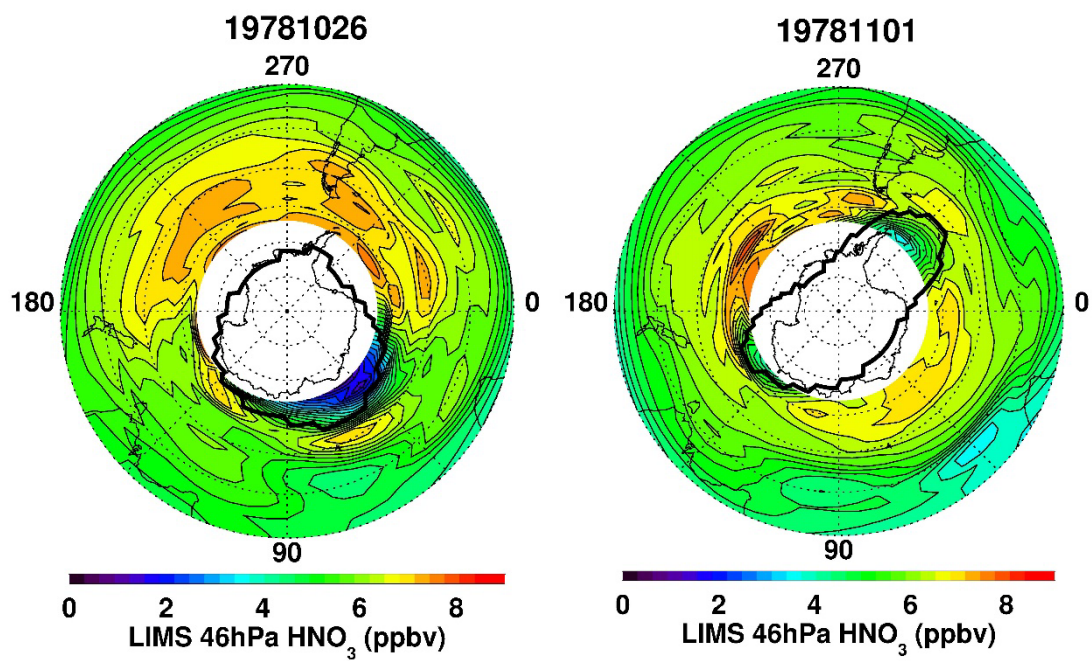


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301 Figure 3—V6 Level 2 species profiles for 59.5°S, 31°E (red) and 54.9°S, 119.4°E (black) on  
 302 October 26, 1978, and from an ozonesonde at Syowa (69°S, 40°E—green) on September 3,  
 303 1978. Ozone (solid) has units of millipascals (mPa), while HNO<sub>3</sub> (dashed) and NO<sub>2</sub> (dot-dashed)  
 304 have units of ppbv.

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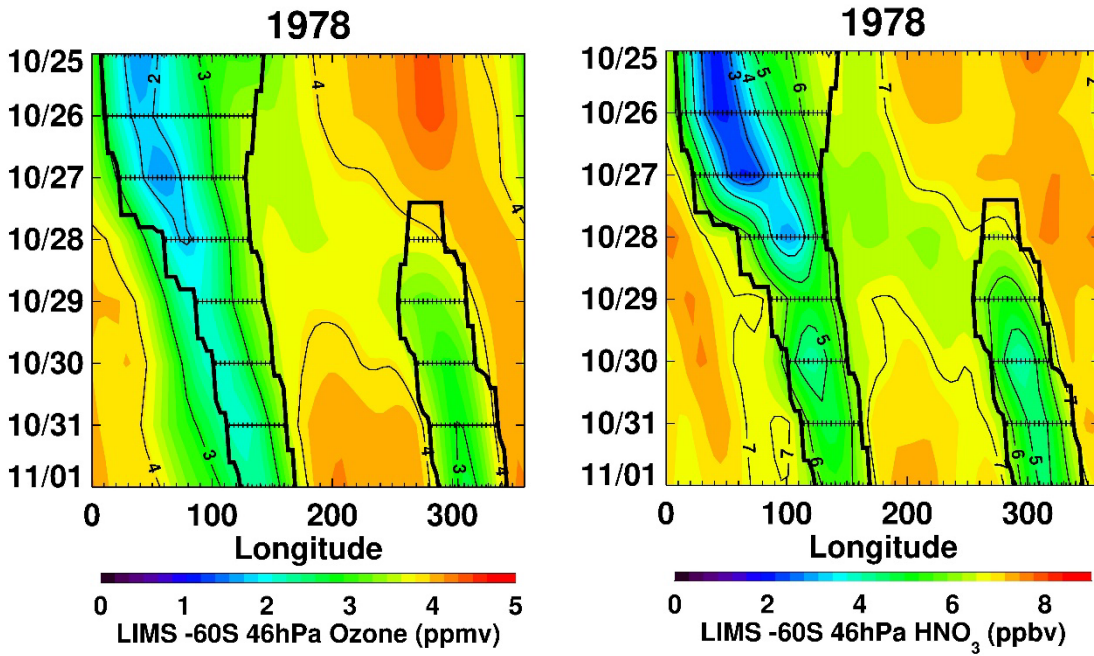
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308 Figure 4—As in Fig. 2, but for V6 HNO<sub>3</sub>.

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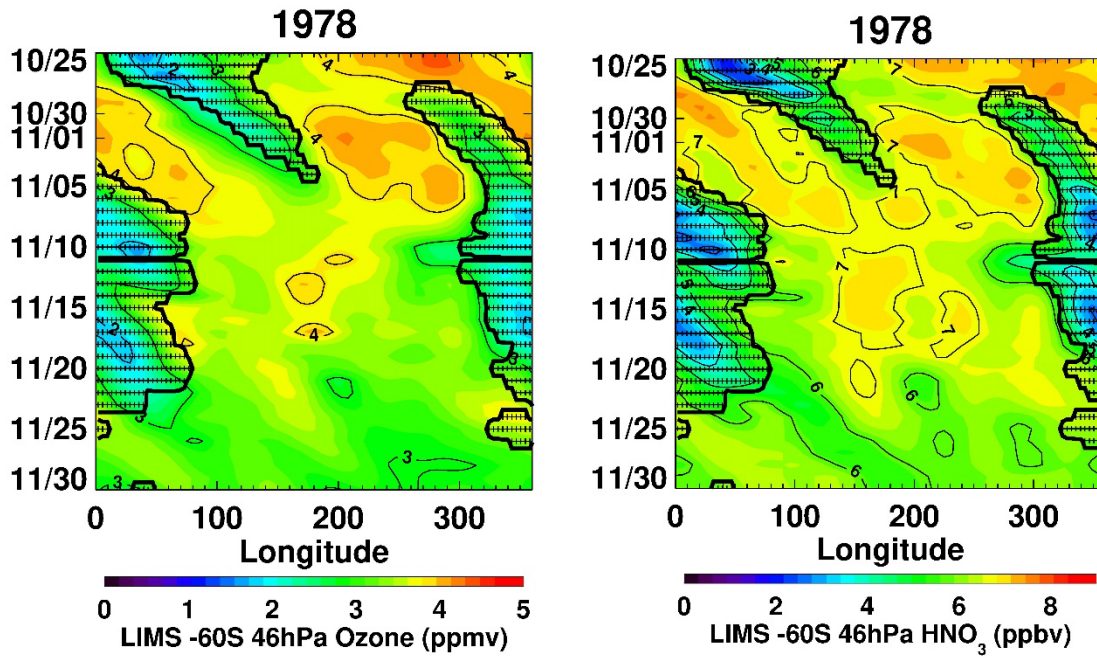


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312 Figure 5—Time/longitude or Hovmöller plots of LIMS ozone (left) and HNO<sub>3</sub> (right) for 60°S  
313 and 46 hPa. The ERA-40 vortex edge shows as thick black contours, and the vortex interior has  
314 horizontal dotted lines.

315

316



317

318 Figure 6—As in Fig. 5, but extended in time from October 25 to November 30, 1978.

319