



1 2	Impact of light-absorbing particles on snow albedo darkening and associated radiative forcing over High Mountain Asia: High resolution WRF-Chem modeling and new satellite observations
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19 Abstract

20	Light-absorbing particles (LAPs), mainly dust and black carbon, can significantly impact
21	snowmelt and regional water availability over High Mountain Asia (HMA). In this study, for the
22	first time, online aerosol-snow interactions enabled and a fully coupled chemistry Weather
23	Research and Forecasting (WRF-Chem) regional model is used to simulate LAP-induced
24	radiative forcing on snow surfaces in HMA at relatively high spatial resolution (12 km, WRF-
25	HR) than previous studies. Simulated macro- and micro-physical properties of the snowpack and
26	LAP-induced snow darkening are evaluated against new spatially and temporally complete
27	datasets of snow covered area, grain size, and impurities-induced albedo reduction over HMA. A
28	WRF-Chem quasi-global simulation with the same configuration as WRF-HR but a coarser
29	spatial resolution (1 degree, WRF-CR) is also used to illustrate the impact of spatial resolution
30	on simulations of snow properties and aerosol distribution over HMA. Due to a more realistic
31	representation of terrain slopes over HMA, the higher resolution model (WRF-HR) shows
32	significantly better performance in simulating snow area cover, duration of snow cover, snow
33	albedo and snow grain size over HMA, as well as an evidently better atmospheric aerosol
34	loading and mean LAPs concentration in snow. However, the differences in albedo reduction
35	from model and satellite retrievals is large during winter due to associated overestimation in
36	simulated snow fraction. It is noteworthy that Himalayan snow cover have high magnitudes of
37	LAP-induced snow albedo reduction (4-8 %) in summer (both from WRF-HR and satellite
38	estimates), which, induces a snow-mediated radiative forcing of \sim 30-50 W/m ² . As a result,
39	Himalayas (specifically western Himalayas) hold the most vulnerable glaciers and mountain
40	snowpack to the LAP-induced snow darkening effect within HMA. In summary, coarse spatial
41	resolution and absence of snow-aerosol interactions over Himalaya cryosphere will result in
42	significant underestimation of aerosol effect on snow melting and regional hydroclimate.
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48 1. Introduction

49	Light-absorbing aerosol particles (LAPs; airborne dust and black carbon (BC) specks),
50	can impact on regional water availability over Asia in three ways. Firstly, LAPs can directly
51	interact with incoming solar radiation and induce thermo-dynamical modifications to synoptic
52	scale circulations (Hansen et al., 1997; Ramanathan et al., 2001; Bond et al., 2013; Lau et al.,
53	2006; Bollasina et al., 2011; Li et al., 2016). Secondly, acting as cloud condensation nuclei,
54	changes in concentrations of these particles can lead to microphysical modification of cloud
55	systems and precipitations (Fan et al., 2016; Li et al., 2016 ; Qian et al., 2009; Sarangi et al.,
56	2017). Finally, deposition of LAPs in the snowpack can also darken the snow, reduce its surface
57	albedo and accelerate snow warming and melting (Warren and Wiscombe, 1980; Qian et al.,
58	2015; Qian et al., 2011; Qian et al., 2009a; Lau et al., 2010; Xu et al., 2009; Hadley and
59	Kirchstetter, 2012; Dang et al., 2017). Modeling studies have suggested that the LAP-induced
60	snow darkening mechanism has warming and snow-melting efficacy even greater than that of
61	greenhouse gases (GHGs) (Hansen and Nazarenko, 2004; Flanner et al., 2007; Qian et al., 2011;
62	Skiles et al., 2012). To give a perspective, the concentration of just 100 ng of BC in 1 g of
63	snowpack will reduce the visible-wavelength albedo of grain radius 1000 μm by 10% (Fig. 1b of
64	Warren, 2013). A chain of positive feedback mechanisms results in such large impact of LAPs
65	(Qian et al., 2015). Initially, as snow starts to melt, the concentration of LAPs in snowpack
66	increases because a portion of LAPs accumulate at the surface of the snowpack instead of getting
67	washed away with meltwater (Conway et al., 1996; Flanner et al., 2007; Doherty et al., 2010).
68	This increase in LAP concentration leads to enhanced warming of the snowpack and thereby
69	increases the effective snow grain size, which further lowers snow albedo (Warren and
70	Wiscombe, 1980; Hadley and Kirchstetter, 2012). Nonetheless, at higher concentrations, grain





71 sizes can again be reduced due to the loss of mass from surface layers with the intense melting (Painter et al., 2013). As this process continues, sufficient snow melt occurs to expose the darker 72 73 underlying surface, leading to enhanced warming and snow ablation commonly-known as "snow 74 albedo feedback" (Warren and Wiscombe, 1980; Hansen and Nazarenko, 2004; Flanner et al., 2007; Qian et al., 2015). In turn, this earlier loss of snow cover induces surface warming and 75 perturbing regional circulations (Hansen and Nazarenko, 2004; Lau et al., 2010; Qian et al., 76 77 2011). This LAP-induced modification of snow albedo feedback is identified as one of major forcing agents affecting climate change with a high level of uncertainty (IPCC, 2013). 78

High Mountain Asia (HMA) includes the Tibetan plateau, central Asian mountains and 79 80 the Himalaya cryosphere. It holds the largest glacial cover (~9500 glaciers) outside the polar region (Dyurgerov, 2001). Observations revealed that a historical decadal increase in the surface 81 82 air temperature over HMA in a range of 0.6-1.8 °C (Shrestha et al., 1999; Wang et al., 2008), and the warming is faster over higher elevations (> 4000 m) in the last three decades (Xu et al., 83 84 2009b; Ghatak et al., 2014). The Himalaya glacier area has cumulatively decreased by $\sim 16\%$ during the period 1962 to 2004 (Kulkarni et al., 2010) and the spring snow cover is decreasing at 85 a decadal rate of ~ 0.8 million km² during the last 50 years (Brown and Robinson, 2011). The 86 87 average retreat rate on the north slope of Mount Everest is as high as $5.5-9.5 \text{ m y}^{-1}$ (Ren et al., 2006). The Himalaya cryosphere contributes to the stream flow in Indus and Ganges river 88 systems by ~ 50 % and $\sim 10-30$ %, respectively (Khan et al., 2017). Warming and glacier retreat 89 90 over the Himalaya cryosphere have a great potential to impact the fresh water availability for about 700 million people, modify regional hydrology, and disturb the agrarian economy of all 91 South Asian countries (Bolch et al., 2012; Immerzeel et al., 2010; Kaser et al., 2010; Singh and 92





Bengtsson, 2004;Barnett et al., 2005;Yao et al., 2007). Therefore, it is critical to disentangle the

- 94 factors contributing to glacier retreat and snow melt over HMA.
- 95 Regional warming due to increasing greenhouse gases (Ren and Karoly, 2006) has been reported
- as the primary cause of the high rate of warming and glacier retreat over HMA. However, in the
- 97 last decade, advancement in remote sensing and availability of measurements from several field
- 98 campaigns suggest that the contribution of LAP loading (in the atmosphere) to the warming and
- 99 glacier melting over HMA is probably greater than previously believed (Ramanathan et al.,
- 100 2007; Prasad et al., 2009; Menon et al., 2010). Continuous observations over the Nepal Climate
- 101 Observatory Pyramid (NCO-P) facility located at 5079 m a.s.l. in the southern foothills of Mt.
- 102 Everest revealed very high concentrations of black carbon (Marcq et al., 2010) and desert dust
- 103 (Bonasoni et al., 2010) especially in spring from Indo-Gangetic plains. Atmospheric LAPs are
- scavenged to the snow/ice surface by dry and wet deposition and cause measurable snow
- darkening and melting (Gautam et al., 2013; Yasunari et al., 2010b; Yasunari et al., 2013; Nair et
- 106 al., 2013; Ménégoz et al., 2014; Ming et al., 2008; Flanner and Zender, 2005). Thus, LAP
- 107 deposited in snow and associated snow darkening has been suggested as a key factor to the early
- snowmelt and rapid glacier retreat over HMA (Yasunari et al., 2010; Ming et al., 2008; Xu et al.,
- 109 2009a; Flanner et al., 2007; Qian et al., 2011).
- 110 While previous studies have underlined the significance of LAP-deposition in snow over
- 111 HMA, the estimation of LAP-induced snow darkening and associated radiative forcing is still
- highly uncertain (Qian et al., 2015). Many of these studies used online global model simulations
- at coarse spatial resolutions of ~50-150 km (Flanner and Zender, 2005; Ming et al., 2008; Qian
- et al., 2011). Others used offline simulation of the snow albedo effect using measured
- 115 concentrations of deposited LAP in surface snow or estimated from atmospheric loading and ice





116	cores (Yasunari et al., 2013; Nair et al., 2013; Wang et al., 2015). The complex terrain of HMA,
117	seasonal snowfall and near surface air circulation are not well resolved by coarse global climate
118	models (Kopacz et al., 2011; Ménégoz et al., 2014). Similarly, offline estimations are limited in
119	scope because they are site specific and are based on simplified assumptions about deposition
120	rates. Ideally, online high resolution simulations allowing for LAP-snow interactions should
121	facilitate a more realistic understanding of LAP deposition to snow and LAP-induced snow
122	darkening effect in terms of both magnitude and spatial variability over HMA.
123	In this study, a modified version of the online chemistry coupled with Weather Research
124	and Forecasting regional model (WRF-Chem v3.5.1), which, is then fully coupled with SNICAR
125	(SNow, ICe, and Aerosol Radiative) model, is used to perform first-ever high resolution (12 km)
126	simulation over the HMA region for the water year 2013-14 (October 1, 2013 to September 30,
127	2014). Satellite observations of snow properties like snow albedo, grain size, and LAP-induced
128	snow darkening from MODSCAG and MODDRFS retrievals are used for evaluation (Painter et
129	al, 2009; 2012). The main objective of this study is to evaluate the skill of high resolution WRF-
130	Chem model in simulating properties of snowpack, aerosol distribution, LAP in snow and LAP-
131	induced snow darkening over HMA using spatially and temporally complete (STC) remotely
132	sensed snow surface properties (SSP) from MODIS (Dozier et al, 2008; Rittger et al, 2016). Our
133	second objective is to demonstrate the benefit for aerosol and snow distributions in high
134	resolution runs by comparing to a coarse gridded quasi-global model simulation over HMA. This
135	quasi-global simulation is run with the same WRF-Chem configuration but at 1 degree spatial
136	resolution. Finally, the spatiotemporal variation of simulated LAP deposition, snow albedo
137	darkening and snow mediated LAP radiative forcing (LAPRF) over HMA are discussed. The





- 138 model details and datasets used are described in Section 2. Results and discussions are presented
- in Section 3 followed by conclusions in Section 4.
- 140 2: Model simulations and observational datasets
- 141 Below, we provide details on the aerosol module used in WRF-Chem, interactive
- 142 coupling with aerosol and SNICAR via land model, and the model setup for both 12 km and 1
- degree resolution runs. The details for the remote sensing observations that are used to evaluate
- the models are also provided.

145 2.1: Coupled WRF-Chem-CLM-SNICAR Model description

The WRF-Chem simulation is performed at $12 \text{ km} \times 12 \text{ km}$ horizontal resolution

147 (hereafter refereed as WRF-HR) with 210×150 grid cells (64–89°E, 23–40°N) (Figure 1) and

148 35 vertical layers. The simulation was conducted from 20th September, 2013 to 30th September,

149 2014, to provide one year of results (following a 10-day model spin-up). ERA-interim reanalysis

data at 0.7° horizontal resolution and 6 h temporal intervals are used for meteorological initial

and lateral boundary conditions. The simulation is re-initialized every 4th day to prevent the drift

152 of model meteorology. Model physics options used are the MYJ (Mellor–Yamada– Janjic)

153 planetary boundary layer scheme, Morrison 2-moment microphysics scheme, community land

154 model (CLM), Kain-Fritsch cumulus scheme and Rapid Radiative Transfer Model for GCMs

155 (RRTMG) for longwave and shortwave radiation schemes.

156 The CBM-Z (carbon bond mechanism) photochemical mechanism (Zaveri and Peters,

- 157 1999) coupled with eight bin MOSAIC (Model for Simulating Aerosol Interactions and
- 158 Chemistry) aerosol model (Zaveri et al., 2008) is used. This is the most sophisticated aerosol
- 159 module available for the WRF-Chem model. The sectional approach with eight discrete size bins





160	is used to represent the size distributions of all the major aerosol components (including sulfate,
161	nitrate, ammonium, black carbon (BC), organic carbon (OC), sea salt, and mineral dust) in the
162	model. The processes of nucleation, condensation, coagulation, aqueous phase chemistry, and
163	water uptake by aerosols in each bin size are included in the MOSAIC module. Dry deposition of
164	aerosol mass and number is simulated by including both diffusion and gravitational effects as per
165	Binkowski and Shankar (1995). Wet removal of aerosols follow Easter et al. (2004) and
166	Chapman et al. (2009) and includes grid resolved impaction and interception processes for both
167	in-cloud (rainout) and below-cloud (washout) aerosol removal. Processes involved in convective
168	transport and wet removal of aerosols by cumulus clouds are described in Zhao et al. (2013).
169	Anthropogenic emissions used in our study is at $0.5^{\circ} \times 0.5^{\circ}$ horizontal resolution and are
170	taken from the NASA INTEX-B mission Asian emission inventory for year 2006 (Zhang et al.,
171	2009). Biomass burning emissions at $0.5^{\circ} \times 0.5^{\circ}$ horizontal resolution for the water year 2013-14
172	are obtained from the Global Fire Emissions Database, Version 3 (GFEDv3) (Van Der Werf et
173	al., 2010), which are vertically distributed in our simulation using the injection heights
174	prescribed by Dentener et al. (2006) for the Aerosol Inter Comparison project (AeroCom). Sea
175	salt and dust emissions follow Zhao et al. (2014). Dust surface emission fluxes are calculated
176	with the Georgia Institute of Technology-Goddard Global Ozone Chemistry Aerosol Radiation
177	and Transport (GOCART) dust emission scheme (Ginoux et al., 2001), and emitted into the eight
178	MOSAIC size bins with respective mass fractions of 10^{-6} , 10^{-4} , 0.02, 0.2, 1.5, 6, 26, and 45%.
179	Aerosol optical properties are computed as a function of wavelength for each model grid
180	cell. The Optical Properties of Aerosols and Clouds (OPAC) data set (Hess et al., 1998) is used
181	for the shortwave (SW) and longwave (LW) refractive indices of aerosols and a complex
182	refractive index of aerosols (assuming internal mixture) is calculated by volume averaging for





each chemical constituent of aerosols for each bin. A spectrally-invariant value of $1.53 \pm 0.003i$ 183 is used for the SW complex refractive index of dust. Fast et al. (2006) and Barnard et al. (2010) 184 185 provide detailed descriptions of the computation of aerosol optical properties such as extinction 186 coefficient, single scattering albedo (SSA), and asymmetry factor in WRF-Chem. Following 187 Zhao et al. (2011) and Zhao et al. (2013a), aerosol radiative feedback is coupled with the Rapid Radiative Transfer Model (RRTMG) (Mlawer et al., 1997) and the direct radiative forcing of 188 189 individual aerosol species in the atmosphere are diagnosed. Aerosol-cloud interactions are included in the model following Gustafson et al. (2007). 190

The increasingly used Snow, Ice, and Aerosol Radiation (SNICAR) model simulates the 191 192 snow properties and associated radiative heating rates of multilayer snow packs (Flanner and 193 Zender, 2005; Flanner et al., 2009, 2012 and 2007). Fundamentally, it employs the snow albedo 194 theory (parameterization) based on Warren and Wiscombe (1980) and the two-stream radiative approximation for multilayers from Toon et al. (1989). SNICAR can also simulate aerosol 195 196 radiative effect in snow for studying the LAP heating and snow aging (Flanner et al., 2007). 197 Recently, laboratory and site measurements are used to validate the SNICAR simulated change 198 of snow albedo for a given BC concentration in snow (Hadley and Kirchstetter, 2012; Brandt et 199 al., 2011). For radiative transfer calculations, SNICAR defines layers matching with the five thermal layers in community land model (CLM) that vertically resolve the snow densification 200 and meltwater transport (Oleson et al., 2010). In WRF-Chem-SNICAR coupled model, BC and 201 202 dust deposition on snow is calculated in a prognostic approach through dry and wet deposition processes. BC in snow can be represented as externally and internally mixed with precipitation 203 hydrometeors depending on the removal mechanism involved, but dust is considered to only mix 204 externally with snow grains (following Flanner et al., 2012). SNICAR in WRF-Chem simulates 205





206	four tracers of dust based on size (with diameters of 0.1–1, 1–2.5, 2.5–5, and 5–10 μ m) and two
207	tracers of BC (externally and internally mixed BC with 0.2 μ m dry diameter) in snow. The
208	MOSAIC aerosol model simulates dust in the atmosphere with eight size bins (0.039–0.078,
209	0.078–0.156, 0.156–0.312, 0.312–0.625, 0.625 - 1.25, 1.25–2.5, 2.5–5.0, and 5.0–10.0 μm in dry
210	diameter). The first 4 bins are coupled with the smallest bin of dust particles in SNICAR. While
211	the next two MOSAIC bins (5 th and 6 th) map into the second bin of SNICAR, the 7 th and 8 th
212	MOSAIC dust bins correspond to the third and fourth SNICAR dust bins (Zhao et al., 2014),
213	respectively. Deposition of LAPs to snow in SNICAR are immediately mixed in the CLM
214	surface snow layer (< 3 cm). CLM adds excess water in the layer above to the layer beneath
215	during melting. The scavenging of aerosols in snow by meltwater is assumed to be proportional
216	to its mass mixing ratio of the meltwater multiplied by a scavenging factor. Scavenging factors
217	for externally mixed BC and internally mixed BC are assumed to be 0.03 and 0.2, respectively,
218	and 0.02, 0.02, 0.01, and 0.01 for the four dust bins (all externally mixed). Although these
219	scavenging factors are comparable to observations (Doherty et al., 2013), the scavenging ratios
220	can be highly heterogeneous and introduce high uncertainty into the estimation of LAP
221	concentrations in snow (Flanner et al., 2012; Qian et al., 2014). More detailed description about
222	the aerosol deposition and mixing processes, computation of optical properties of snow and
223	LAPs in WRF-Chem-CLM-SNICAR coupling can be found in Zhao et al.(2014) and Flanner et
224	al.(2012).

Configured in the way similar to the WRF-HR, a coarse $(1^{\circ} \times 1^{\circ})$ gridded WRF-Chem simulation is also performed using a quasi-global model (hereafter referred as WRF-CR) with 360×130 grid cells $(180^{\circ} \text{ W}-180^{\circ} \text{ E}, 60^{\circ} \text{ S}-70^{\circ} \text{ N})$. Periodic boundary conditions are used in the zonal direction. Reanalysis of the TROpospheric (RETRO) anthropogenic emissions for the





- 229 year 2010 (ftp://ftp.retro.enes.org/ pub/emissions/aggregated/anthro/0.5x0.5/) is used for
- anthropogenic aerosol and precursor gas emissions in the coarse gridded quasi-global WRF-
- 231 Chem simulation except for Asia and the United States. INTEX-B anthropogenic emissions
- 232 (Zhang et al., 2009) and US National Emission Inventory are used for Asia and the U.S.,
- respectively. Emissions of biomass burning aerosols, sea salt, and dust are treated in the same
- way as described above for the WRF-HR simulation. More details about the quasi-global WRF-
- 235 Chem simulation can be found in (Zhao et al., 2013b);Hu et al., 2016). Chemical initial and
- 236 boundary conditions to the WRF-HR simulation are provided by this quasi-global WRF-CR runs
- 237 for the same time period to include long-range transported chemical species.
- 238 2.2: Aerosol Optical Depth (AOD) dataset
- The aerosol robotic network (AERONET https://aeronet.gsfc.nasa.gov) is a global
- 240 network of ground based remote sensing stations that provides quality-controlled measurements
- of AOD with uncertainties ~0.01 under clearsky conditions over India (Holben et al., 1998;
- 242 Dubovik et al., 2000). CIMEL Sun scanning spectral radiometers are used to measure direct Sun
- radiance at eight spectral channels (340, 380, 440, 500, 675, 870, 940, and 1020 nm) and
- 244 measure spectral columnar AOD (Holben et al., 1998). AERONET provides measurements at
- ~ 15 min temporal resolution from sunrise to sunset.
- 246 Skyradiometer Network (Skynet) is another global network of ground based spectral
- 247 scanning radiometer (POM-01L, Prede, Japan) stations that provides quality-controlled
- 248 measurements of AOD (Nakajima et al., 1996). With an automatic sun scanner and sensor, it
- measures sky irradiance in five wavelengths i.e. 400, 500, 675, 870, and 1020 nm. The measured
- 250 monochromatic irradiance data is processed by using Skyrad.Pack version 4.2 software.
- 251 Calibration of the Sky radiometer is carried out on a monthly basis (http://atmos3.cr.chiba-





- 252 <u>u.jp/skynet/data.html</u>). Details of the instrumentation and software protocol can be found in
- 253 Campanelli et al. (2007) and Ningombam et al., (2015). In this study, we have also used AOD
- 254 measurements at 500 nm over MERAK, a high altitude Skynet station in Himalaya for water
- 255 year 2013-14.
- 256 The MODerate resolution Imaging SpectroRadiometer (MODIS) instrument onboard the
- 257 NASA AQUA satellite provides global coverage of daily radiance observations (at 1330 LT) in
- 258 36 spectral channels. Over North India, Tripathi et al., (2005) has shown that MODIS
- 259 observations correlate well with ground based measurements. For the evaluation of model
- simulated AOD, 1° gridded Level 3 AOD estimates (collection 6) at 0.55 µm wavelength
- obtained from the MODIS instrument are used during water year 2013-14. However, the MODIS
- land aerosol algorithm uses a dark target approach (Levy et al., 2007), which, is known to have
- large uncertainties over arid and mountainous surfaces (Levy et al., 2010).

264 2.3: Spatially and temporally complete MODSCAG and MODDRFS retrievals

- 265 Subpixel snow-covered area and snow grain size are retrieved from MODIS-observed
- surface reflectance data using the physically based MODIS Snow-Covered Area and Grain size
- 267 (MODSCAG) (Painter et al., 2009) algorithm. In each snow covered pixel, MODSCAG
- 268 attributes a fractional snow-covered area and grain size using spectral mixture analysis to
- determine proportion of the pixel that is snow and is not snow. MODSCAG is more accurately
- 270 identifies snow cover throughout the year than the widely used MODIS snow product:
- 271 MOD10A1 (Rittger et al., 2013). The MODSCAG snow-mapping algorithm for fraction of snow
- covered area has an uncertainty of ~ 5 % (Rittger et al., 2013). The current study incorporates
- pixel level snow cover area and snow grain size from MODSCAG over the HMA region to
- 274 evaluate snow pack simulation and LAP-induced albedo reduction. Further, MODIS Dust





275 Radiative Forcing in Snow (MODDRFS) model (Painter et al., 2012) is used to determine the LAP-induced albedo reduction over HMA. MODDRFS uses spectral reflectance differences 276 277 between the measured snow spectral albedo and the modeled clean snow spectral albedo. The 278 pixel level clean snow spectrum corresponding to MODSCAG retrieved snow grain sizes is 279 calculated using discrete ordinate radiative transfer solutions for visible wavelengths and solar zenith angles. Coupled, these products provide the determination of snow albedo for the 280 281 fractional snow cover with LAP inclusion. Reflectance inputs to MODSCAG and MODDRFS are degraded by cloud cover, off-nadir views, and data errors, but can be filtered in time and 282 space to improve data quality and consistency. Our method for spatially cleaning and filling 283 284 (Dozier et al., 2008;Rittger et al., 2016) combines noise filtering, snow/cloud interpolation and 285 smoothing to improve the daily estimates snow surface properties (SSP). Using remotely sensed forest height maps (Simard et al., 2011) and MODSCAG vegetation fraction, we adjust the 286 287 satellite viewable snow cover to account for snow under tree canopy (Rittger et al., 2016). We weight the observations based on satellite viewing angle that varies from 0 to 65 degree with 288 larger uncertainties in off-nadir views (Dozier et al., 2008). The result is a set of spatially and 289 290 temporally complete (STC) SSPs. Use of these products in an energy balance model to estimate snow water equivalent based on reconstruction produced more accurate snow cover than the 291 Snow Data Assimilation System (SNODAS) or an interpolation of observations from snow 292 293 pillows (Bair et al., 2016). In this study we use STC versions of MODSCAG and MODDRFS when comparing our WRF model output. The incomplete remotely sensed would be difficult to 294 295 use given the gaps in data and uncertainties related to viewing angle (Dozier et al., 2008). 296 Hereafter, the use of MODSCAG and MODDRFS terms will invariably refer to these STC-MODSCAG and STC-MODDRFS products. 297







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Figure 1: Panel A illustrates the terrain elevation at 1 km resolution from ETOPO1 dataset. Panel
B shows glacier classification over HMA region used in this study following Randolph Glacier
Inventory. Panels C and D illustrate terrain representation in WRF-HR run (12 km) and WRFCR (1 degree) runs, respectively. For reference, Mt. Everest (shown as black circle in Figure 1C)
is distinctly represented in Panel A and C, but not in Panel B.

304

305 2.4: Variation in terrain representation in WRF-HR and WRF-CR

Figure 1A illustrates the variations in terrain height over HMA at a resolution of 1 arc

307 using ETOPO1 Global Relief Model, a publicly available global topographical dataset (Amante

- and Eakins, 2009). It clearly shows the enormous relief in terrain as we move from the Indo
- 309 Gangetic plains (IGP) to the crest of the Himalaya and into the Tibetan Plateau (TP). The
- majority of HMA is above 4 km altitude with many Himalaya peaks at an altitude higher than 6
- km. Figure 1B illustrates the mountain ranges and glaciers classified as per the Randolph Glacier





- 312Inventory in the Fifth Assessment Report of the Intergovernmental Panel on Climate Change
- 313 (Pfeffer et al., 2014). Specifically, Pamirs, Hindu Kush, Karakoram, Kunlun, and Himalaya hold
- the most number of glaciers in HMA. Figures 1C and 1D illustrate the representation of terrain
- elevation in WRF-HR and WRF-CR, respectively. Compared to Figure 1A, location of mountain
- peaks (altitude > 5.5 km) are better represented in WRF-HR compared to WRF-CR, as is
- 317 particularly evident over the Karakoram, Kunlun, and Himalaya ranges. Moreover, the steep rise
- in elevation between IGP and TP is also well represented by WRF-HR, whereas it is more
- 319 gradual in WRF-CR.

320 2.5: Methodology

321 Simulation of the snow macro- and micro-physical properties, aerosol loading and LAP

322 in snow concentration from WRF-HR, WRF-CR and observational estimates (datasets described

above) over HMA are compared in Section 3.1 and Section 3.2. In Section 3.3, the WRF-HR

324 simulated LAP-induced snow albedo reduction values over HMA is compared with

325 corresponding MODIS satellite based STC-MODSCAG and STC-MODDRFS. Lastly, a

326 discussion on the high resolution model simulated LAP-induced radiative forcing estimates over

327 HMA is also presented in context to previous studies and other atmospheric forcing.

328 The simulated fractional snow covered area (fSCA), duration of snow cover over a grid

in terms of number of snow cover days (NSD), snow albedo (α) and snow grain sizes (SGS) and

LAP-induced snow albedo darkening ($\Delta \alpha$) for midday (1000 -1400 LT) conditions from both the

331 WRF models are compared with corresponding STC-MODSCAG and STC-MODDRFS

retrievals over HMA. The number of snow cover days (NSD: defined as days having fSCA

values ≥ 0.01) during water year 2013-14 is determined over each grid from STC-MODSCAG

and both model runs. They are compared with corresponding values from STC-MODSCAG





retrievals, which, are observed during Terra overpasses at 10:30 LT. We have used a window 335 from 10:00 LT to 16:00 LT for representing midday averages of modelled variables to 336 337 incorporate the variability due to differences in timing (between model and real scenario) of weather conditions like precipitation and clouds. In addition, the change in snow albedo during 338 10:00 - 14:00 LT is < 0.01 (Bair et al., 2017), which, is low compared to other model physics-339 and data retrieval related uncertainties. The WRF-CR simulated variables and STC-MODSCAG 340 341 and STC-MODDRFS retrievals are gridded to the resolution of WRF-HR (12 km) for ease of comparison. We have compared annual mean values as well as seasonal mean values for winter 342 (December - February) and summer (April - June) season, separately. We have not considered 343 344 the monsoon period in our analyses because the snow cover during the monsoon is negligible 345 (except in glaciated regions at high altitudes) relative to other months (Figure S1). To evaluate spatial heterogeneity in our model, the seasonal and annual distribution of these variables are 346 347 calculated separately for each sub region (shown in Figure 1) within HMA. In addition, to gain an understanding of the extent of temporal variability present in LAP-induced effects, we have 348 also presented daily midday variation in LAP-induced snow darkening and LAP-induced 349 350 radiative forcing at surface over Chotta Shingri glacier region (32.1-32.35 °N, 77.4-77.7 °E) located in the Chandra-Bhaga river basin of Lahaul valley, Pir Panjal range, in Western 351 Himalayas. It is an accessible and representative site for glacier mass balance studies in western 352 353 Himalayas. Chotta Shigri glacier has a cumulative glaciological mass loss of -6.72 m w.e. 354 between 2002 and 2014 (Azam et al., 2016). The simulated aerosol optical depth (AOD) is compared with available in-situ 355 356 observations (described in Section 2.2). Here, quality assured (Level 2) midday (1000 to 1400 LT) averages of AOD (550 nm) at seven AERONET stations (Lahore, Jaipur, Kanpur, Gandhi 357





372	3.1: Snow physical, microphysical and optical properties
371	3: Results and Discussions
370	NCO-pyramid station (Ginot et al., 2014), respectively.
369	Ata in eastern slopes of Pamirs (Wake et al., 1994), East Rongbuk (Ming et al., 2012) and near
368	LAP _{dust} are over Abramov glacier in western slopes of Pamirs (Schmale et al., 2017), Muztagh
367	et al., 2015; Ginot et al., 2014) is used. Similarly, the point measurements used for evaluating
366	NCO-pyramid site in Nepal at 5-6 km altitude (Kaspari et al., 2014; Yasunari et al., 2013; Jacobi
365	Xu et al., 2009a) and composite of recent in-situ measurements from various studies near the
364	(Svensson et al., 2018), East Rongbuk at 6.4 km altitude (Ming et al., 2012; Ming et al., 2008;
363	Muztagh Ata in eastern slopes of Pamirs (Xu et al., 2006), Uttaranchal region of W. Himalayas
362	over glaciated regions within our study domain. In this study, measurements of LAP_{BC} over
361	of BC (LAP _{BC}) and dust particles (LAP _{dust}) in the snow surface or the surface layer are available
360	few sites is compared with field measurements. Only a few field measurements of concentration
359	the simulated AOD values. Further, the simulated distribution of LAP concentration in snow at a
358	college, Kathmandu and CAS) and one SkyNet site within our study region are used to evaluate

373 The largest values of region-averaged annual mean fSCA within HMA are observed in both the

satellite retrievals and the model runs over the Karakoram region (mean=0.45) followed by

375 Pamirs, Himalayas and Hindu Kush in the HMA region (Figures 2A and 2B). In comparison, the

- 376 fSCA over Kunlun and TP are lower (<0.3), but, pockets of very high fSCA (~0.7) are visible
- over the west Kunlun ranges (Figure 2A). The annual mean fSCA values and the fine spatial
- variability are well simulated by WRF-HR (Figure 2B) over the entire HMA region. Of
- 379 exception are simulations over the Karakoram, where WRF-HR overestimates annual mean
- 380 fSCA, however, the distribution of annual mean fSCA from WRF-HR and STC-MODSCAG





- agree in all the sub regions (Figure 2D). This observation is largely valid also for summer
- months (Figure 2F). But, significant overestimation in distribution of fSCA (by >0.2) during
- 383 winter is present over Pamirs, Karakoram, W. Himalayas, TP and Kunlun region (Figures 2E).
- 384 STC-MODSCAG retrievals illustrate that the Pamirs (NSD=230 days) and Karakoram
- 385 (NSD=270 days) ranges remain snow covered for 7-9 months of the year (Figure 3A). Similarly,
- the grids in Hindu Kush (NSD=194), W. Himalayas (NSD=189 days) and C. Himalayas
- 387 (NSD=191 days) are snow covered for ~ 6-7 months. Mountains in E. Himalayas (NSD=142
- days) remain snow covered for only 4-5 months of the year. The distribution of annual NSD
- values simulated by WRF-HR in each sub region is close to STC-MODSCAG values (Figure
- 300 3D). Also, the spatial distribution and magnitude of simulated NSD by WRF-HR is similar to
- that from STC-MODSCAG for different seasons, separately (Figure S2). Thus, overestimation of
- annual mean fSCA in WRF-HR during winter is not due to mere averaging error associated with
- underestimation in simulated NSD during winter (Figure S2).







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Figure 2: Spatial distribution of annual mean snow cover fraction (fSCA) during midday (1000-396 1400 LT) for water year 2013-14 from A) STC-MODSCAG retrievals, B) simulated values from 397 WRF-HR and C) WRF-CR simulations. Panels D-F illustrate the distribution of midday mean 398 399 fSCA over each subregion identified by glacier classification following Randolph Glacier 400 Inventory (X-axis). The circle and vertical legs represent mean±standard deviation over each region for D) entire year, E) winter (December - February) and F) summer (April - June) season, 401 separately. Here, Hindu Kush, Karakoram, W.Himalayas, C.Himalayas, E.Himalayas, Tibetan 402 403 Plateau, West Kunlun and East Kunlun regions are abbreviated as HK, KK, WH, CH, EH, TP, 404 WKun and EKun, respectively.





406	We also calculated the number of days of snow cover with low fSCA (<0.5; Figure 3B)
407	and high fSCA (>0.5; Figure 3C) values, separately. The grids in Kunlun, Northern slope of
408	Karakoram, eastern slope of Pamirs and TP region are dominated by snow cover of relatively
409	low fSCA for most of their snow cover duration (Figure 3B). But, grids in Hindu Kush,
410	Himalayas and southern slopes of Karakoram are generally covered with high fSCA values
411	throughout the year (Figure 3C). The distribution of simulated NSD values over each sub region
412	for low and high fSCA scenario is shown in Figures 3E and 3F, respectively. WRF-HR can well
413	simulate the NSD over grids with high fSCA (Figure 3F) but significantly underestimates NSD
414	over grids with low snow cover (Figure 3E). Note that the regions dominated by low annual
415	fSCA in this water year are actually the same regions where WRF-HR simulated fSCA values
416	are being overestimated in winter (Figure 2E). Thus, simulation of fewer number of days with
417	low fSCA (and vice versa) in WRF-HR might also be contributing partially to the overestimation
418	of winter fSCA simulated in WRF-HR compared to STC-MODSCAG. Interestingly, WRF-CR
419	simulated NSD values for low fSCA case is in better agreement with STC-MODSCAG values
420	(Figure 3E). Winter mean distribution of WRF-CR simulated fSCA over Kunlun, W.Himalaya
421	and TP region better match STC-MODSCAG values than the corresponding WRF-HR simulated
422	fSCA values (Figure 2F). It is noteworthy that these subregions (which are dominated by low
423	fSCA grids) receive snowfall from western disturbances during winter months. The cloud cover
424	associated with the western disturbances over these sub regions are extensive in winter which
425	also introduces uncertainty in MODSCAG retrievals and STC processing and contributes to the
426	differences between WRF-HR and MODSCAG in fSCA.

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Figure 3: Spatial distribution of snow duration in terms of NSD from A) STC-MODSCAG 429 retrievals. Panel B and C are similar to Panel A, but, shows number of days when fSCA is below 430 431 and above 0.5 over each grid, respectively. Panels D illustrate the distribution of NSD over each sub region identified by glacier classification following Randolph Glacier Inventory. The circle 432 and vertical legs represent mean \pm standard deviation over each region for entire year. Here, 433 Hindu Kush, Karakoram, W.Himalayas, C.Himalayas, E.Himalayas, Tibetan Plateau, West 434 Kunlun and East Kunlun regions are abbreviated as HK, KK, WH, CH, EH, TP, WKun and 435 EKun, respectively. Panel E and F are similar to Panel D, but, for NSD corresponding to fSCA 436 437 values below and above 0.5, respectively.

438

439 Comparison between performance of WRF-CR and WRF-HR for fSCA clearly show







- ranges (Figure 2D). For instance, the simulated annual mean fSCA in WRF-CR around Mt.
- 442 Everest (shown as black circle in Figure 1) is less than 0.1 (Figure 2C). This is contrary to the
- 443 high fSCA values observed at Mt. Everest (0.7 in Figures 2A) and simulated by WRF-HR (0.7 in
- 444 Figures 2B). Moreover, the improvement is present in both winter and summer months
- 445 indicating it's independence from meteorological variations (Figures 2E and 2F). Analysis of
- 446 NSD values indicate that the snow cover duration in WRF-HR also improved significantly over
- these slopes (Figure S3) irrespective of the season. Note that WRF-CR underestimates the snow
- 448 duration over Hindu Kush and Himalayas by ~2-6 months (Figure S3) and the spatial location of
- 449 grids with very high annual mean fSCA values (mountain ranges) improved in WRF-HR
- 450 compared to the STC-MODSCAG data (Figures 2A-C). The observed improvement in fSCA and
- 451 NSD simulation over the slopes of Himalaya and Hindu Kush can be attributed to better terrain
- 452 representation in WRF-HR.

Next, the simulated microphysical properties of the snow pack are evaluated against the 453 454 remote sensing retrievals. Spatial patterns in annual mean SGS from STC-MODSCAG are similar to that seen in fSCA with highest values over the Karakoram and Himalayan ranges 455 (Figures 4A) corresponding to the highest elevations and likely the coldest temperatures 456 457 hindering snow grain growth. This spatial distribution of annual mean SGS values is well simulated in WRF-HR runs (Figure 4B). But, the annual mean values are largely overestimated 458 by 30-50 micron (Figure 4D) relative to STC-MODSCAG. The seasonal distribution of region-459 460 segregated SGS values from WRF-HR also compares well with that from STC-MODSCAG retrievals (Figure 4E and 4F). Simulated annual mean SGS from WRF-CR (Figure 4C) lack the 461 fine spatial variability seen in STC-MODSCAG and WRF-HR. Moreover, the SGS estimates are 462 largely underestimated (by up to 100 microns) by WRF-CR specifically over grids in central and 463





464	eastern Himalayas, TP and Kunlun ranges (Figure 4D). The large underestimation of SGS from
465	WRF-CR and overestimation of SGS from WRF-HR is present for both summer and winter
466	months (not shown). The overestimation of SGS from WRF-HR values corroborate well with the
467	finding that the simulated fSCA distribution from WRF-HR is largely skewed towards higher
468	values (Figure 3). Similarly, the unrealistically low mean values of SGS from WRF-CR over
469	Himalayas, TP and Kunlun ranges are consistent with the underestimation of fSCA and NSD
470	values over these regions (Figure 2 and 3). While, SGS retrievals from STC-MODSCAG are
471	based on observed surface reflectance, the modeled SGS is calculated from simulated snow mass
472	in top model layer in the grid. Hence, improvement in simulation of fSCA and NSD in high
473	resolution WRF-HR runs also caused the SGS values from WRF-HR to be closer to STC-
474	MODSCAG retrievals than SGS from WRF-CR runs. It is worth noting that the presence of
475	cloud cover influences STC-MODSCAG retrievals of SGS towards smaller grain sizes if clouds
476	are misidentified as snow. This systematic error could also contribute to the SGS differences
477	between WRF-HR and STC-MODSCAG estimates.

478







Figure 4: Spatial distribution of annual mean snow cover fraction (fSCA) during midday (1000-481 1400 LT) for water year 2013-14 from A) STC-MODSCAG retrievals and simulated values 482 483 from B) WRF-HR and C) WRF-CR runs is shown. Panel D illustrates the distribution of midday mean fSCA over each subregion identified by glacier classification following Randolph Glacier 484 Inventory. The circle and vertical legs represent mean±standard deviation over each region for 485 the entire year. Here, Hindu Kush, Karakoram, W.Himalayas, C.Himalayas, E.Himalayas, 486 Tibetan Plateau, West Kunlun and East Kunlun regions are abbreviated as HK, KK, WH, CH, 487 EH, TP, WKun and EKun, respectively. Panel E shows time-series of daily midday SGS (hashed 488 489 lines) and fSCA (solid lines) from MODSCAG (black), WRF-HR (blue) and WRF-CR (red) over a grid located near the Chotta Shingri glacier (marked by magenta diamond in Figure 3A) of 490 491 western Himalaya region.





492	Interestingly, the SGS values from WRF-CR, over grids comprising the Chotta Shingri
493	glacier (marked by magenta diamond in Figure 4A) of western Himalaya, are closer to STC-
494	MODSCAG observations compared to those from the high resolution model, WRF-HR. As a
495	sanity check of the above explanation of fSCA-SGS association, daily changes of SGS (hashed
496	lines in Figure 4E) and fSCA (solid lines in Figure 4E) from STC-MODSCAG (black), WRF-
497	HR (blue) and WRF-CR (red) over this glacier are compared. Fractional snow cover from STC-
498	MODSCAG gradually increase (from below 0.2) in November, 13 (to above 0.8) in February
499	and subsequently decrease back to 0.1 gradually by Septmeber, 14 at the glacier location.
500	Corresponding SGS values from STC-MODSCAG closely followed the seasonal trend in fSCA
501	varying around the values of 80-200 micron in winter. In comparison, simulated fSCA from
502	WRF-HR values drastically increased to 1 in starting of November, 13 (from no snow cover
503	before that), remain fully snow covered till mid-June, 14 and then steeply become snow free for
504	rest of period after June. Compared to satellite estimates, fSCA from WRF-HR are greater in
505	magnitude throughout the duration of snow cover indicating more snow mass simulated by
506	WRF-HR. Associated SGS values simulated by WRF-HR values are also greater in magnitude
507	(80-800 micron) than STC-MODSCAG estimates throughout the snow duration over the grid. In
508	addition, the day to day variation in simulated SGS values is much larger than the STC-
509	MODSCAG observations. However, fSCA variation from WRF-CR over this grid is very close
510	to the variation seen by STC-MODSCAG and the associated SGS values from WRF-CR of (50-
511	400 micron) are also closer to the estimated STC-MODSCAG values, supporting our argument
512	of fSCA-linked bias in SGS estimates between model and satellite retrievals.
513	The annual mean snow albedo (α) values and the distribution over each sub region from

satellite estimates (by combing grain sizes from STC-MODSCAG and decrease in albedo from





STC-MODDRFS (see Bair et al, 2016)) and simulated by both models are presented in Figure 4. 515 Highest annual mean α values (~ 0.65-0.75) are observed over mountain peaks in Karakoram, 516 517 Pamirs and W.Himalaya regions (Figure 5A). The location and magnitude of annual mean α over these grids are closely reproduced in WRF-HR with an underestimation of < 10% (Figure 5B). 518 WRF-CR simulated annual mean α values over these grids have a considerably larger 519 underestimate of ~50% (Figure 5C). Similar statistics are prevalent over all the sub regions of 520 521 HMA (Figure 5D). Specifically, the distribution of α values from WRF-HR nearly matches the observed distribution, but, the distribution of albedos from the coarser model, WRF-CR, are 522 523 generally 0.2-0.3 lower when compared to the observations. As above, cloud misidentified as 524 snow could increase grain sized leading to slightly higher albedos. Also, note the opposite bias direction for albedo simulated by WRF-HR compared to simulated SGS values. This is intuitive 525 as smaller snow particles cover greater surface area and therefore reflect more solar radiation 526 from the surface. A similar pattern in distribution of snow albedo from WRF-HR and WRF-CR 527 are also found over the sub regions for summer and winter months, separately (Figures 5E and 528 5F), indicating robust improvement in simulation of albedo values from WRF-HR throughout the 529 530 year. The differences in simulated α in WRF-HR with the observations increased during summer and were lower during winter. Here, it is worth mentioning that we are comparing instantaneous 531 estimates obtained from Terra overpass during 1000 LT with midday (1000-1400 LT) mean 532 533 model values. The inherent diurnally in $-\alpha$ values under clear sky conditions in summer season (Bair et al., 2017) might contribute partially to the observed enhancement in differences during 534 summer season. The improvement in α estimation from WRF-HR compared to WRF-CR can be 535 536 attributed to the relatively better simulation of the overall macro- and microphysical properties of the snowpack in high resolution runs. 537









Figure 5: Spatial distribution of annual mean snow cover fraction (fSCA) during midday (1000-539 1400 LT) for water year 2013-14 from A) STC-MODSCAG retrievals and simulated values from 540 B) WRF-HR and C) WRF-CR runs is shown. Panels D-F illustrate the distribution of midday 541 mean fSCA over each subregion identified by glacier classification following Randolph Glacier 542 543 Inventory. The circle and vertical legs represent mean±standard deviation over each region for 544 D) entire year, E) winter (December - February) and F) summer (April - June) season, separately. Here, Hindu Kush, Karakoram, W.Himalayas, C.Himalayas, E.Himalayas, Tibetan Plateau, West 545 Kunlun and East Kunlun regions are abbreviated as HK, KK, WH, CH, EH, TP, WKun and 546 EKun, respectively. 547





549 **3.2: Aerosol distribution and LAP in snow**

550	We used available in-situ and ground sun photometer measurements from seven different
551	sites across our study domain (location shown in Figure 6) to evaluate the simulated aerosol
552	optical depth (AOD). The annual mean midday AOD at each site is shown in Figure 6A. Three
553	sites (i.e. Merak, CAS and Kathmandu shown in Figures 6B-D) are located on the Himalaya
554	slopes and the other four sites (Lahore, Jaipur, Kanpur and Gandhi College shown in Figures 6E-
555	H) are located in the Indo-Gangetic Plains. In-situ measurements clearly illustrate a sharp
556	decrease (4-5 fold) in mean AOD as we traverse higher up the Himalayan slope. The annual
557	mean AOD for Lahore and Kanpur sites are 0.41 and 0.52, respectively, while the AOD over
558	high elevated sites i.e. Merak and CAS sites are 0.07 and 0.05, respectively. Also, MODIS-
559	observed AOD values prominently show the reduction in annual mean AOD from the Indo-
560	Gangetic Plains (MODIS-AOD ~ 0.4-0.7) to the Tibet region (MODIS-AOD ~ 0.1-0.2) (Figure
561	S5). Over the four sites in the Indo-Gangetic Plains, AOD simulated by both WRF-HR and
562	WRF-CR runs are well correlated with observations ($r= 0.5-0.6$, Figures 6E-6H). The biases in
563	modelled AOD are also similar (in the range of 0.2-0.4) in case of both WRF-HR and WRF-CR
564	runs (Figures 6E-6H). Thus, no significant improvement in AOD values are achieved over the
565	plain region with fine resolution. However, distinct and large improvement in simulated AOD is
566	seen over the high elevation sites due to the increase in spatial resolution. Note that AOD values
567	from WRF-CR are not strongly correlated with observations at these sites (Figures 6B-6D) and
568	also have very high positive biases in AOD values (even higher than annual mean values at
569	Merak and CAS stations). In contrast, the correlation between observations and WRF-HR is
570	reasonably good (r=0.5-0.8 at these sites) using fine spatial resolution in WRF-HR. The positive
571	biases in AOD from WRF-HR at Merak and CAS sites are lower than corresponding WRF-CR





- values by an order of magnitude. Presence of lower biases in AOD from WRF-HR over high
- 573 elevation sites indicates that the observed sharp decrease in AOD values across the Himalayan
- slope are better captured by the higher resolution WRF run (WRF-HR) than in the coarser run.
- 575 Greater annual mean AOD value is simulated by WRF-CR over the entire HMA region
- 576 compared to WRF-HR (Figure S5) supporting an overestimation of AOD from WRF-CR at
- higher elevation in addition to the few sites. The presence of high biases (0.3-0.4) over
- 578 Kathmandu valley even in WRF-HR runs indicate that model resolution even finer than 12 km is
- 579 likely needed to better resolve the AOD distribution in complex terrain around valleys in
- 580 Himalayan slope regions. Moreover, Jayarathne et al., 2018 shows that many local emissions are
- not accounted in global emissions which causes underestimation in simulated regional AOD
- values in these valleys. Temporal variability in monthly mean AOD (relatively higher AOD in
- summer) is simulated reasonably well by both the model versions (Figure S5).
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592 Figure 6: Comparison of midday (averaged over 1000-1400 LT) aerosol optical depth (AOD) measured by ground based sun photometer at seven sites within the study domain with 593 corresponding simulated AOD values from, both, WRF-HR and WRF-CR. The annual mean 594 595 AOD values over each site is shown by shade in topmost left Panel. The other panels illustrates the comparison over one of each stations shown by dots in topmost left Panel. The 'N' 596 mentioned in the legend in each panel is the total number of days when collocated data between 597 model and measurement is available over that site. These sample points are divided into 50 equal 598 bins of ascending AERONET-AOD values (2 percentile each) and averaged. The standard 599 deviation in each bin is shown by the vertical bars. The correlation coefficient values (r) are also 600 mentioned in the legend followed in brackets by the relative error values (Σ rmse/mean obs). 601

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Figure 7: Spatial distribution of annual mean LAP concentration in snow top layer for water year 608 2013-14 simulated by A) WRF-HR and B) WRF-CR. The black dots in Panel A denote the 609 locations where observations of BC in snow is available. Similarly, the magenta dots in Panel B 610 611 denote the locations from where observation of dust in snow is available. Panels C-D illustrate comparison of simulated LAP_{BC} (top) and LAP_{dust} (bottom), respectively, in topmost snow layer 612 with observed values over the marked locations in the Himalayan cryosphere. Annual mean 613 (circle) and distribution (box plot) is used as metric of comparison. The WRF-HR and WRF-CR 614 615 values are represented by blue and red, respectively. The pink lines in Panel A are the crosssections shown in Figure 8. 616





617	Significant differences in simulated AOD over high elevations of Himalaya slopes and
618	TP indicate that considerable differences might also be present in LAP concentrations in snow
619	between the two WRF simulations. Annual mean LAP concentrations in top snow layer from
620	WRF-HR and WRF-CR simulations are compared in Figures 7A-7B. The comparison shows that
621	LAP concentration in WRF-HR are significantly higher than WRF-CR simulated values.
622	Quantitatively, the WRF-HR simulated annual mean LAP concentrations over the Pamir (0.5
623	g/kg), Karakoram (0.45 g/kg), Hindu Kush (0.2 g/kg), W. Himalaya (0.3 g/kg), C. Himalaya (0.2
624	g/kg) and E. Himalaya (0.08 g/kg) ranges is 3-5 times higher than the same from WRF-CR runs.
625	In contrast, WRF-HR simulated LAP over TP (0.21 g/kg) and Kunlun ranges (0.8 g/kg) is similar
626	to the mean magnitude simulated by WRF-CR runs. As a sanity check, we evaluate the simulated
627	LAP concentrations against those previously reported in the literature. Figures 7C and 7D
628	illustrate the evaluation of mean annual LAP concentration from WRF-HR and WRF-CR
629	associated with BC (LAP _{BC}) and dust (LAP _{dust}), respectively, against the reported data (shown as
630	filled black circles). The locations of reported LAP_{BC} (black filled dots) and LAP_{dust} (magenta
631	filled dots) are shown in Figure 7C and Figure 7D, respectively. First, the difference in the
632	magnitude of LAP_{dust} and LAP_{BC} over HMA is striking. The LAP_{dust} is more than 1000 times
633	greater than LAP_{BC} both in observations and the models. Secondly, LAP_{BC} and LAP_{dust} values
634	from WRF-HR are much closer to reported values compared to WRF-CR values. The differences
635	in mean of reported LAP_{BC} and LAP_{dust} distribution to that simulated by WRF-HR at various
636	sites are in range of 5-30 μ g/kg and 5-20 mg/kg, respectively. WRF-CR well simulates the
637	concentrations of LAP_{BC} and LAP_{dust} over Pamirs (~ 10 mg/kg), but significantly underestimates
638	the LAP_{BC} and LAP_{dust} (by an order of magnitude) over the Himalayan ranges. Although the
639	reported data are not specific to water year 2013-14, it can be reasonably assumed that the inter-





- 640 annual variations of LAP concentration in snow is of the order of magnitude as uncertainty in the
- observations. Thus, the WRF-HR better simulates aerosol and LAP concentration than the WRF-
- 642 CR over the HMA region.
- 643 It is interesting to note that finer spatial resolution resulted in lower AOD but greater LAP values in snow over some places in HMA. For more insight, the vertical distribution of 644 aerosol concentration in altitude-latitude space (Figure 8) across two latitudinal cross-sections 645 646 (magenta colored lines in Figure 7A) is analyzed. Figure 8 illustrates the differences in simulated vertical distribution of mean aerosol number concentration along 78°E (row 1) and 87°E (row 2) 647 for WRF-HR (left column) and WRF-CR (right column) runs, respectively. Corresponding 648 terrain elevation (black solid line) and snow depth (magenta bars) are also overlaid in these plots. 649 The latitude-altitude plots clearly illustrate that improved representation of the terrain in WRF-650 651 HR shows the sharp change of elevation over Himalayan foothills and causes a steeper natural barrier to the transport of aerosols uphill from IGP to HMA region than in the WFR-CR model. 652 653 Also, high spatial resolution enhances snowfall in WRF-HR over the HMA region relative to the 654 WRF-CR model. While the former change increased annual dry deposition flux, more snowfall caused greater wet deposition annually in WRF-HR compared to WRF-CR (Figure S6). The 655 656 combination of these effects increases the deposition of aerosols and therefore LAP on the southern slopes of Himalaya in the WRF-HR run. This explains the coexistence of higher LAP 657 concentration/deposition and lower AOD across HMA in WRF-HR, compared to corresponding 658 659 WRF-CR results.







Figure 8: Longitudinally-averaged annual mean aerosol number concentration plotted in altitudelatitude space for two longitudinal traverses across the study domain, i.e. 78°N (Panels A and B)
and 87°N (Panels C and D) for both WRF-HR (left column) and WRF-CR (right column).
Corresponding terrain elevation is shown in solid black line. Corresponding to each latitude, the
longitudinally-averaged annual mean snow depth is also presented in magenta color bars (using
y-axis on the right).







673 3.3 LAP-induced Snow darkening and radiative forcing

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Figure 9: Spatial distribution of annual mean LAP-induced snow albedo darkening ($\Delta \alpha$) during 675 676 midday (1000-1400 LT) for water year 2013-14 from A) MODSCAG retrievals and B) simulated values from WRF-HR is shown. Panels C and D illustrate the distribution of midday mean $\Delta \alpha$ 677 over each sub region identified by glacier classification following Randolph Glacier Inventory. 678 The circle and vertical legs represent mean±standard deviation over each region for entire year. 679 Here, Hindu Kush, Karakoram, W.Himalayas, C.Himalayas, E.Himalayas, Tibetan Plateau, West 680 Kunlun and East Kunlun regions are abbreviated as HK, KK, WH, CH, EH, TP, WKun and 681 682 EKun, respectively. Panel E shows time-series of daily midday $\Delta \alpha$ from STC-MODSCAG (black) and WRF-HR (blue) over the grids located near the Chotta Shingri glacier (marked in 683 684 Figure A) of western Himalaya region. Also fractional snow cover and LAP concentrations from WRF-HR over the same grids are included. 685





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687	STC-MODDRFS retrievals illustrate that locations in Hindu Kush and W. Himalayas
688	have the highest annual mean LAP-induced reduction in snow albedo ($\Delta \alpha$ in %) followed by
689	Karakoram, C.Himalayas and Pamir regions (Figure 9A). WRF-HR simulated the spatial
690	variations in annual mean $\Delta \alpha$ reasonably well (Figure 9B), but, the magnitudes are
691	underestimated by \sim 20-40 % throughout the domain. Note that the biases in annual mean values
692	are lowest over grids in Himalayan ranges (where the underestimation is within 20%). Season
693	wise and region wise distribution plots show that the WRF-HR biases are higher in winter
694	months than the summer months (Figures 9C and 9D). While, WRF-HR simulated $\Delta \alpha$ values in
695	the winter span between 1-3 %, the corresponding STC-MODDRFS estimates of $\Delta \alpha$ are larger
696	with values ranging between 3-6% (thus no overlap with model values) over all the sub regions.
697	In summer months, the distribution of modeled WRF-HR $\Delta\alpha$ values over Karakoram and
698	Himalayan ranges are similar in magnitude, explaining the lower biases in annual mean values
699	over Himalayas. This spatiotemporal variability in differences between STC-MODDRFS
700	retrievals and simulated $\Delta \alpha$ values is consistent with the variability in biases of fractional snow
701	cover seen in WRF-HR (Figures2E and 2F). Specifically, the large underestimation [and
702	significant improvement relative to WRF-CR] in WRF-HR- $\Delta \alpha$ values (Figures 9C and 9D) in
703	winter [summer] over Karakoram, Hindu Kush and Himalayas is in agreement with
704	corresponding overestimation [improvement] of fSCA from WRF-HR over these regions (Figure
705	2).
706	For a closer look, the difference in daily midday mean $\Delta \alpha$ values from STC-MODDRFS

707 (black) and WRF-HR (blue) are compared (in Figure 9E) over the grids of Chotta Shingri glacier

708 (similar as Figure 4E). Corresponding, midday mean fSCA (light blue) and LAP (orange) from





709 WRF-HR are also plotted. $\Delta \alpha$ values from STC-MODDRFS are about 5% during winter months, but, increases in summer months until mid-June (peak value is 18%). Albedo reduction is closely 710 711 associated with the temporal progression in midday LAP concentration in snow over this region 712 at daily scale. In agreement, midday $\Delta \alpha$ values from WRF-HR are lower in winter months and 713 higher in summer months. Except occasional peaks, with magnitudes of 3-4 %, $\Delta \alpha$ values from WRF-HR largely remained below 3 % till late February. A steep increase in $\Delta \alpha$ values from 714 715 WRF-HR is seen in March (monthly mean $\sim 4\%$), April (9%), May (13%) and June (18%). As already discussed, the simulated fSCA values in WRF-HR are greater than observed fSCA from 716 STC-MODSCAG for most of the winter season (Figure 4E). STC-MODDRFS estimated $\Delta \alpha$ is 717 718 based on surface reflectance, while $\Delta \alpha$ calculated by model is for a surface layer of ~3 cm. The 719 surface snow layer in SNICAR/CLM continuously evolves as fresh snowfall is added or with snow melting, so the LAP concentrations in surface layer depend on new snowfall, meltwater 720 flushing, and layer combination/division (Flanner et al., 2007; Flanner et al., 2012; Oleson et al., 721 2010). Thus, more snow cover or thicker surface layers in winter results in lower values of 722 annual mean LAP concentration and thus underestimates associated with LAP-induced snow 723 724 darkening. In addition, the associated overestimation in modelled SGS during winter (Figure 4E) can also contribute to the lower WRF-HR- $\Delta \alpha$ values, because, bigger snow grains in WRF-HR 725 lead to lower clean albedo and thus smaller reduction in albedo compared to STC-MODDRFS. 726 Another notable point is the large $\Delta \alpha$ values (> 20 %) from WRF-HR that occur towards 727 728 the end of snow cover in June, which, is not seen in the STC-MODDRFS retrievals. The variations could be due to 2 reasons, either the snowpack is underestimated or the LAP 729 concentration is overestimated by the model. Some factors which can contribute to these 730 discrepancies in summer are 1) Larger LAP values may be simulated due to model uncertainties 731





in enhanced wet scavenging fluxes; 2) It is well known that with transport and deposition of 732 aerosols from IGP to W.Himalayas increases afternoon with evolution of boundary layer over the 733 734 IGP region (Dumka et al., 2015; Raatikainen et al., 2014). This feature is well simulated by the model (not shown). As STC-MODDRFS estimates are representative of 1000 LT, but, modelled 735 736 values are representative of midday mean (1000-1400 LT), more aerosol deposition might be resulting in higher $\Delta \alpha$ values WRF-HR during summer months; 3) At the same time, the 737 738 uncertainties associated with modelling aerosol-snow albedo microphysical feedbacks to snow melt may also be contributing to underestimation of snow packs in summer. However, more 739 740 certainty on these modelled values require evaluation against in-situ measurements. It is worth 741 mentioning here that no in-situ measurements are available for direct comparison of these high $\Delta \alpha$ values WRF-HR over W.Himalayas (Gertler et al., 2016). Nonetheless, the high values 742 simulated during summer end are in near-range of previously reported values over other HMA 743 regions. Kaspari et al., (2014) used the offline SNICAR model to report that BC concentrations 744 in spring snow/ice samples at Mera Glacier were large enough to reduce albedo by 6-10% but 745 that with the inclusion of dust, the reduction in albedo was 40-42% relative to clean snow. 746 747 Recently, Zhang et al., 2018 has combined a large dataset of LAP measurements in surface snow with offline SNICAR model to illustrate that $\Delta \alpha$ can be >35% over Tibetan Plateau. Moreover, 748 the composite effect of this discrepancy on seasonal/annual mean values is minimal as the 749 750 snowpack is at its minimum at summer end. Similar high daily variability and huge LAPRF values (~200 W/m²) in late summer as well as associated sudden decline in snow depth is also 751 reported in sites over upper Colorado river basin (Skiles et al., 2015; Skiles and Painter, 2017). 752

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Figure 10: A) Spatial distribution of annual mean snow mediated LAP-induced radiative forcing 757 (LAPRF) from WRF-HR and B) Spatial distribution of seasonal mean LAPRF over snow 758 covered for summer season (AMJ) of water year 2013-14. Panels C and D illustrate the 759 distribution of LAPRF (blue) and BC-only-LAPRF (black) over each sub region identified by 760 761 glacier classification following Randolph Glacier Inventory for winter and summer, respectively. The circle and vertical legs represent mean±standard deviation over each region for entire year. 762 Here, Hindu Kush, Karakoram, W.Himalayas, C.Himalayas, E.Himalayas, Tibetan Plateau, West 763 Kunlun and East Kunlun regions are abbreviated as HK, KK, WH, CH, EH, TP, W.Kun and 764 765 E.Kun, respectively. Panel E shows time-series of daily midday LAPRF (blue) and aerosol direct radiative forcing at surface (ADRF, blue) over the Chotta Shingri region (marked in Figure 9A) 766 767 of western Himalaya region. The simulated LAP-induced change in snow albedo is also shown





768 in hashed black line. Also snow depth and snowfall from WRF-HR over the same grid is included. 769 In agreement with the spatial variation in $\Delta \alpha$ values, the annual mean snow-mediated 770 LAP-induced radiative forcing (LAPRF) from WRF-HR over grids in W.Himalayas and C. 771 772 Himalayas are highest followed by Karakoram, Hindu Kush, E.Himalayas and Pamir regions (Figure 10A). The spatial distribution of summer mean LAPRF values from WRF-HR (Figure 773 774 10B) is similar to that of the annual mean LAPRF values, but, the summer magnitudes are higher 775 by an order of magnitude throughout most of the domain. This is mainly due to the large increase in $\Delta \alpha$ values in summer when LAPs aggregate on the surface compared with winter months 776 when LAPs are continuously covered by new snow (Figure 9). Spatio-temporal variability in 777 LAPRF is evident in the seasonal and regional distribution plots (Figure 10C and 10D). LAPRF 778 779 values over the edges of HMA are greater than the highland TP region in both winter and 780 summer months probably due to greater LAP deposition simulated over Hindu Kush, Karakoram 781 and Himalayas regions (Figure S7) from the close proximity of dust sources. Also, it is clearly 782 visible that the maximum LAPRF values within HMA region are present over grids in Himalayan ranges (during both winter and summer) with annual mean values > 50 W/m² (seen in 783 Figures10B and 10D) and maximum instantaneous values higher than 150 W/m² (not shown). 784 785 The time series of midday mean LAPRF values (Figure 10E) over the same grids in the Chotta Shingri Glacier region in western Himalaya is plotted to ascertain possible daily variability in 786 LAPRF for midday over the region. Corresponding LAP-induced albedo reduction, snow depth 787 788 and snowfall values are also plotted to show how LAPRF can affect local snow melting. The 789 daily snow depth increases to 3.6 m in winter (30-150 days) followed by gradual reduction and snow cover melting in summer (150-270 days). The mean midday LAPRF value is $\sim 10 \text{ W/m}^2$ in 790 winter season, but, the magnitude increased gradually during March (18 W/m²), April (44 791





 W/m^2), May (86 W/m^2) and June (123.5 W/m^2), eventually terminating with a peak value of 202 792 W/m² in mid-June. The temporal evolution in $\Delta \alpha$ values closely followed the LAPRF values with 793 794 a variation in range of 1-20%. The shortwave aerosol direct radiative forcing (ADRF) during 795 midday at the surface over the same grid is also shown (as blue curve) in Figure 10E. The 796 momentarily high values of ADRF during the period indicate dust storms or sudden increase in aerosol loading over the grid. A closer look illustrate that the large daily variations in LAPRF 797 798 and $\Delta \alpha$ are associated with the variability in ADRF and daily snowfall occurence over the grid. Fresh snowfall feedbacks result in subsequent reduction in LAPRF and enhancement in snow 799 depth (Figure 10E), while, higher aerosol loading over aged snow is followed by a clear increase 800 801 in LAPRF. During melting, our model considers that only a fraction of LAP is washed away with 802 meltwater, which, results in enhanced concentration of LAP during the end stages of the snowpack. This snow albedo feedback along with the momentary high aerosol loading (on 250th) 803 day) can explain the very high values of albedo reduction and LAPRF that were simulated during 804 the last days of the snow cover over this grid. Higher LAPRF values indicate more energy being 805 absorbed by the snow pack and thus more snow melting. Therefore, LAP-induced snow melting 806 807 effect over Himalaya is very significant and is the largest relative to other areas within HMA. 808 Snow-mediated radiative forcing only due to BC- deposition is shown in Figures10C and 10D. Although, similar spatiotemporal pattern in LAPRF and BC-only-LAPRF is simulated, 809 contrasting features in context to BC contribution to LAPRF are present in between the western 810 811 part and eastern parts of HMA. During winter, the magnitude of BC-only LAPRF values is similar to that of total LAPRF over the western part of HMA region (i.e. Pamirs, Karakoram, 812 Hindu Kush and W.Himalayas) suggesting that BC is a dominant contributor to net LAPRF 813 (Figures10C). But, large differences between LAPRF and BC-only LAPRF is present during 814





815 summer season over the western regions indicating that dust contribution is more or less equal to that of BC contribution over these regions (Figure 10D). In contrast, the dust contribution to 816 817 LAPRF values over the eastern domain (TP and Kunlun regions) is significant during winter 818 season (Figures10C and 10D). The spatiotemporal variability in dust and BC contribution is 819 reported previously and is mainly linked with the seasonal variability in meteorology and associated advection of South Asian emissions (Zhang et al., 2015; Wang et al., 2015a; Niu et 820 821 al., 2018). The ADRF-induced surface cooling effect may nullify the effects of LAPRF-induced warming effect on snowpack melting. But, the drastic increase in LAPRF values in April through 822 823 June causes the magnitude of LAPRF to be twice that of ADRF over W. Himalayas during snow 824 melting period, highlighting dominance of LAPRF (as also seen in Figure 10E).

825 The simulated annual mean, summer mean and BC-only-LAPRF values from WRF-HR 826 are in general, higher compared to previously reported estimates of LAPRF from model studies at coarser resolution. For example, Ménégoz et al., (2014) have reported annual mean LAPRF of 827 828 \sim 1-3 W/m² over Himalaya using an online simulation at 50 km resolution grid. Similarly, Qian 829 et al., (2011) used the Community Atmosphere Model version 3.1 at coarser spatial resolution to 830 show that simulated aerosol-induced snow albedo perturbations can generate LAPRF values of 831 5-25 W/m² during spring over HMA. Also, coarsely resolved GEOS-Chem runs simulated BConly-LAPRF can vary from 5 to 15 W/m² in the snow-covered regions over the TP (Kopacz et 832 al., 2011). Recently, a decade long simulation using the RegCM model at 50 km spatial 833 834 resolution also estimated maximum BC-only-LAPRF values of 5-6 W/m² over the Himalaya and southeastern TP averaged over non-monsoon season (Ji, 2016). However, the comparison of 835 WRF-HR and WRF-CR simulations provided in this study clearly show that the magnitudes of 836 snow macro- and micro-properties, aerosol loading and LAP-induced albedo darkening over 837





- 838 Himalayas improved significantly with finer spatial resolution. Thus, the global model simulated
- 839 LAPRF values are likely underestimated. In agreement, recently, Zhang et al., (2018) and has
- 840 estimated BC-only-LAPRF of 20-35 W/m² using offline SNICAR calculation forced with a
- greater coverage of measurements of surface snow content. He et al., (2018) have also reported
- similar high BC-only-LAPRF values after implementing a realistic snow grainsize
- 843 parameterizations in offline SNICAR calculations over HMA.

844 4. Summary and Implications

- 845 In this study, we use the SNICAR model coupled with WRF-Chem regional model at high
- spatial resolution (WRF-HR; 12 km) to simulate the transport, deposition, and radiative forcing
- of light absorbing particles (LAPs; mainly Black Carbon and dust) over the high mountains of
- Asia (HMA) during water year 2013-14. The snow grain sizes and LAP-induced snow darkening
- 849 was evaluated, for the first time, against comprehensive satellite retrievals (the MODSCAG and
- 850 MODDRFS spatial and temporally complete retrieved satellite observations) over HMA. The
- atmospheric aerosol loading is evaluated against satellite and ground-based AOD measurements
- over HMA region. Results from another simulation which employ the same model configuration
- but a coarser spatial resolution (WRF-CR; 1 degree) are also compared with WRF-HR to
- 854 illustrate the significance of a better representation of terrain on snow-pack and aerosol
- simulation over HMA. The main conclusions from our study are:
- 856a) The simulated macro- and micro-physical properties and the duration of snow packs over HMA
- 857 improve significantly due to the use of fine spatial resolution, especially over the southern slopes
- 858 of Hindu Kush and Himalayan ranges.
- 859b) Simulated aerosol loading over HMA is also more realistic in WRF-HR than in WRF-CR, which
- leads to a reduction in biases of annual mean LAP concentration in snow. This improvement is





- 861 attributed to a more realistic simulation of wet deposition (due to a better simulation of snow
- 862 pack) and dry deposition of LAPs (associated with a better representation of terrain) in WRF-
- 863 HR.
- 864c) WRF-HR captures the magnitude of LAP-induced snow albedo reduction ($\Delta \alpha$) over Himalayas
- and Hindu Kush region relatively well compared to the STC-MODDRFS retrievals during
- summer. However, during winter, large biases in modelled $\Delta \alpha$ values are present. This is
- 867 probably due to inherent uncertainties in model parameterizations and satellite retrievals
- associated with the cloud cover over HMA in winter period.
- 869d) The glaciers and snow cover regions located in the Himalaya have the highest LAPRF within
- HMA i.e. annual mean LAPRF ~ 20 W/m² and summer mean LAPRF ~ 40 W/m². This is
- 871 consistent with similar high values of $\Delta \alpha$ over Himalayan ranges i.e. annual mean $\Delta \alpha$ values ~ 2-
- 872 4 % and summer mean $\Delta \alpha$ values ~ 4-8 %. The annual mean LAP concentration in snow values
- 873 (200-300 mg/kg) are also high. Thus, the Himalaya (more specifically, western Himalayas) is

874 most vulnerable to LAP-induced snow melting within HMA.







884	LAPRF uncertainties in the model are warranted in the future by using in-situ observations (i.e.
885	field campaigns), specifically over the most affected western Himalayas, where relevant
886	measurements are largely absent (Gertler et al., 2016). Moreover, satellite retrievals will be
887	markedly improved in the coming decade with the NASA Decadal Survey Surface Biology and
888	Geology imaging spectrometer mission, which includes as a core measurement snow albedo and
889	its controls (National Academies of Science, 2018). These visible through shortwave infrared
890	imaging spectrometer retrievals have uncertainties an order of magnitude smaller (Painter et al.,
891	2013) than those from multispectral sensors such as MODIS and will provide a more accurate
892	constraint on the physically-based modeling pursued here.

893

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