

1 **Reply to reviewers' comments on "Core and margin in warm convective clouds.**
2 **Part I: core types and evolution during a cloud's lifetime"**

3 We would like to thank the reviewers for their insightful and helpful comments that
4 help up improve and clarify the manuscript. Before answering all the reviewers'
5 comments in details, we summarize shortly the main modifications done in the
6 manuscript for addressing the main comments:

- 7 1. Focused emphasis is now put on the novel parts of this study: the differences
8 between the three core types, their evolutions in time, and comparison to current
9 understanding of core size and location within a cloud.
- 10 2. The goals of the work are stated more clearly, and the importance of bin-
11 microphysics is discussed.
- 12 3. The introduction was revised significantly to include a more comprehensive
13 review of relevant previous works and the physical processes associated with
14 differences between the core types.
- 15 4. The theoretical section is presented as a summary of previous knowledge, with
16 purpose to gain intuitive understanding of the differences between the core
17 types.
- 18 5. One cloud field case study was replaced. The revised version presents a
19 continental shallow cumulus convection case study, based on long term
20 observations taken at the ARM Southern Great Plains site.

21
22 Please find below a point-by-point reply to all of the reviewers' comments.

23
24 **Reply to reviewer #1 – RC1**

25 **General Reviewer Comment:**

26 This article provides an analysis of the structure of cumulus clouds and the connection
27 between different regions in cumulus clouds. It defines three different regions in the
28 cloud: 1) a Relative Humidity core, where the air is fully saturated, 2) A buoyancy core,
29 here the air is positively buoyant and 3) a vertical velocity core, defined by upwards
30 motion. The authors argue that typically the buoyancy core is a subset of the relative

31 humidity core, which again is a subset of, or at least smaller than, the vertical velocity
32 core. They also consider the effect of mixing on buoyancy and the existence of
33 overshoots as an explanation for this result. I have three key concerns about the draft
34 as it stands, which I think would need to be addressed before publication. More detail
35 on each of these points can be found further below.

36 1) The main conclusions are not sufficiently novel.

37 2) There has been significant work on mixing in cumulus clouds in general and on the
38 role of processes at the cloud edge in particular that the current study does not refer to.

39 3) I have some major concerns about the analysis framework.

40 Despite the fact that I am critical of the theory and analysis as it stands, I think there is
41 value in analysing the simulations presented here in detail, particularly because many
42 previous studies that have worked on this topic have used an approach to condensation
43 based on immediate saturation adjustment in each grid cell. The use of a spectral bin
44 microphysics model in the context of this a study is therefore valuable, although the
45 difference between the two approaches is not made clear to a sufficient degree. I also
46 realise that there is a follow-up article which may include novel results that might be
47 partially supported by the present work.

48 I would encourage the authors to either fully revise the study, or to incorporate the parts
49 of the study that are needed to support part II into that article, but still make major
50 changes to these sections to address the concerns below. If the material is not
51 incorporated into part II, it would require a more significant overhaul of the draft than
52 would usually be considered a major revision. However, many previous studies have
53 looked into mixing in cumulus in further detail using SAM, so the authors should be
54 well-placed to improve their analysis. The draft is also generally well-written, so some
55 of the material could likely be reused.

56

57 **General Answer:** We thank the reviewer for the beneficial comments. We revised the
58 paper to be clearer and more complete according to the concerns raised in the review.
59 Please see all the details in the answers below.

60

61

62

63 **Main Comments:**

64 **MC1)** The conclusions are not sufficiently novel. The two main conclusions are not
65 really new results.

66 - The role of mixing in rapidly reducing cloud buoyancy has been established in
67 previous theoretical work: e.g. Morrisson (2017) looks into this in detail. This role of
68 mixing in rapidly reducing buoyancy is incorporated in at least three existing
69 parametrisations of cumulus convection (see Kain and Fritsch, 1990; De Rooy and
70 Siebesma, 2008 and Derbyshire et al., 2011). Moreover, the framework of critical
71 mixing fraction used in the studies of Kain and Fritsch and De Rooy and Siebesma
72 already provides a framework for understanding how mixing impacts on buoyancy, and
73 the result that mixing generally reduces buoyancy is well established. There are also a
74 number of existing diagrams that have been used to understand the thermodynamic
75 properties of cumulus clouds, which have also shown mixing generally leads to a region
76 of negative buoyancy (Paluch, 1979; De Roode, 2007). The fact that mixing between
77 core and margin parcels leads to intermediate buoyancy in a near-linear way can also
78 be deduced from previous work (e.g. Pauluis and Schumacher, 2010).

79

80 - Similarly, literature on the existence of convective overshoots dates back as far as at
81 least Betts (1973). In order for additional theory to be valuable, it should therefore
82 include predictions that can be directly and quantitatively compared against Large-
83 Eddy Simulation.

84

85 **MA1)** Thank you for this comment. We have revised the manuscript significantly so
86 that clear emphasis is put now on the novel parts of this work: the differences between
87 the three core types, their evolutions in time, and comparison to current understanding
88 of core size and location within a cloud. We are not aware of previous work which tried
89 to perform such a comprehensive analysis and comparison between the different cores.
90 In addition, we describe better in the revised version the previous relevant studies which
91 have dealt with positive vertical velocity and buoyancy in clouds and the effects of
92 entrainment on them.

93 The new abstract now focuses on the novel aspects of the work:

94 *“The properties of a warm convective cloud are determined by the competition*
95 *between the growth and dissipation processes occurring within it. One way to observe*

96 *and follow this competition is by partitioning the cloud to core and margin regions.*
97 *Here we look at three core definitions: positive vertical velocity (W_{core}),*
98 *supersaturation (RH_{core}), and positive buoyancy (B_{core}), and follow their evolution*
99 *throughout the lifetime of warm convective clouds.*

100 *Using single cloud and cloud field simulations with bin-microphysics schemes, we*
101 *show that the different core types tend to be subsets of one another in the following*
102 *order: $B_{core} \subseteq RH_{core} \subseteq W_{core}$. This property is seen for several different*
103 *thermodynamic profile initializations, and is generally maintained during the*
104 *growing and mature stages of a cloud's lifetime. This finding is in line with previous*
105 *works and theoretical predictions showing that cumulus clouds may be dominated by*
106 *negative buoyancy at certain stages of their lifetime.*

107 *During its mature growth stage, the cloud and its cores are centered at a similar*
108 *location. During cloud dissipation the cores show less overlap, typically reduce in*
109 *size, and migrate from the cloud centroid. In some cases, buoyancy cores can*
110 *reemerge and often reside at the cloud periphery. Thus, the core-shell model of a*
111 *positively buoyant center surrounded by negatively buoyant shell only applies to a*
112 *fraction of the cloud lifetime.”*

113

114 We see merit in including the rather simple theoretical derivations in the text for the
115 sake of completeness and ease of understanding for a reader who is not an expert in this
116 specific field. We have revised the theoretical section so that it better reflects previous
117 works (see the details in the next answer), but please note that none of those works
118 focused on the relative sizes of the different cores.

119

120 **MC2)** Lack of connection to the existing literature. I am citing a few key studies below
121 but this is by no means a comprehensive list.

122 - Over the past two decades there has been a large number of studies on the role of
123 negative buoyancy at the cloud edge (e.g., Zhao and Austin 2005; Jonker et al. 2008;
124 Heus and Jonker, 2008), as well as on the implications of the cloud edge for determining
125 effective mixing between clouds and their environment (e.g. Dawe and Austin 2011).

126 - Similarly, several studies have looked into the role of convective cores in simulations
127 with cloud tracking (e.g. Dawe and Austin 2012, Heus and Seifert 2013), and used this
128 to study the role of these cores throughout the life cycle of the cloud.

129

130 **MA2)** We thank the reviewer for including such an extensive list of relevant previous
131 literature that was indeed lacking in the original manuscript. We have included most of
132 these references (and others) in the revised manuscript, and now try to connect out
133 findings to previous literature wherever found to be of relevance. A few examples are
134 given here, from the introduction:

135 *“The common assumption when partitioning a convective cloud to its physical core
136 and margin is that that the cloud core is at its geometrical center and the peripheral
137 regions (i.e. edges) are the margin. Previous observational (Heus et al., 2009a; Rodts
138 et al., 2003; Wang et al., 2009) and numerical (Heus and Jonker, 2008; Jonker et al.,
139 2008; Seigel, 2014) works have studied the gradients of cloud thermodynamic
140 properties from cloud center to edge, and suggest that a cloud is best described by a
141 core-shell model. This model assumes a core with positive vertical velocity and
142 buoyancy, surrounded by a shell with negative vertical velocity and buoyancy. The
143 shell is the region where mixing between cloudy and environmental air parcels
144 occurs, leading to evaporative cooling → decrease in buoyancy → decrease in vertical
145 velocity.”*

146

147 Another part of the introduction:

148 *“Based on previous findings, here we explore the partition of clouds to core and
149 margin using three different objective core definitions where the cloud core threshold
150 is set to be a positive value (of buoyancy, vertical velocity, or supersaturation). Cloud
151 buoyancy (B) can be approximated by the following formula:*

152
$$B = g \cdot \left(\frac{\theta'}{\theta_0} + 0.61q'_v - q_l \right) \quad (1),$$

153 *Where θ_0 represents the reference state potential temperature, q_v is the water vapor
154 mixing ratio, and q_l is the liquid water content. The (') stands for the deviation from
155 the reference state per height (Wang et al., 2009). Buoyancy is a measure for the
156 vertical acceleration and its integral is the convective potential energy. Latent heat
157 release during moist adiabatic ascent fuels positive buoyancy and clouds' growth,*

158 *while evaporation and subsequent cooling drives cloud decay (de Roode, 2008; Betts,*
159 *1973). The prevalence of negatively buoyancy parcels at the cloud edges due to*
160 *mixing and evaporation is a well-known phenomenon (Morrison, 2017). Mixing*
161 *diagrams have been used to assess this effect (de Roode, 2008; Paluch, 1979; Taylor*
162 *and Baker, 1991), and are at the root of convective parameterization schemes*
163 *(Emanuel, 1991; Gregory and Rowntree, 1990; Kain and Fritsch, 1990) and*
164 *parameterizations of entrainment and detrainment in cumulus clouds (de Rooy and*
165 *Siebesma, 2008; Derbyshire et al., 2011). “*

166

167 And theoretical section:

168 *“Hence, for the adiabatic column case, B_{core} is always a proper subset of W_{core}*
169 *($B_{core} \subset W_{core}$). These effects are commonly seen in warm convective cloud fields*
170 *where permanent vertical layers of negative buoyancy (but with updrafts) within*
171 *clouds typically exist at the bottom and top regions of the cloudy layer (de Roode and*
172 *Bretherton, 2003; Betts, 1973; Garstang and Betts, 1974; Grant and Lock, 2004;*
173 *Heus et al., 2009b; Neggers et al., 2007).”*

174

175 **MC3)** I have some major concerns about the analysis framework.

176 **MC3.1)** One of my main concerns is that the theoretical arguments lean heavily on
177 analysis of adiabatic parcels, with mixing as an afterthought. Previous studies suggest
178 that adiabatic parcels do not occur in shallow and congestus cumulus (e.g. Romps and
179 Kuang, 2010), and this seems to be the case in the current study as well. Note that liquid
180 water path is plotted on a logarithmic scale, i.e. even the clouds that contain most liquid
181 water contain several times less liquid water than they would in the adiabatic case. An
182 adiabatic parcel model therefore offers only very limited insight into the dynamics of
183 cumulus convection. Several approaches exist that better represent the effects of mixing
184 throughout the cloud life cycle, e.g. continuous lateral entrainment (Lin and Arakawa
185 1997, Morrison 2017) or episodic mixing. In order for a theoretical framework to
186 provide quantitative predictions that can be tested against LES, one would likely need
187 such an approach. Moreover, it is clear from a number of recent publications (De Roode
188 et al. 2012 is cited already, but see also Romps and Charn, 2015; Morrison, 2016) that
189 theoretical models for the vertical velocity should incorporate the role of drag.

190

191 **MA3.1)** As expressed in MA1, the original manuscript was faulty in that it gave an
192 impression that it focused on entrainment effects. So we have revised the manuscript to
193 better explain that the theoretical estimations of entrainment effects only serve to give
194 an intuitive understanding of the results found using LES simulations. We use simple
195 dynamical concepts to explain the evolutions of the different core types in cumulus
196 clouds. As written now in the beginning of the theoretical section: “*Here we propose*
197 *simple physical considerations to evaluate the differences in cloud partition to core*
198 *and margin using different definitions. The arguments rely on key findings from*
199 *previous works (see Sect. 1) with aim to gain intuitive understanding of the potential*
200 *differences between the core types*”.

201 We are well aware that the use of an adiabatic cloud column (as done in the theoretical
202 part) is simplistic, nevertheless, we think it manages to easily convey theoretical ideas
203 and robust cloud field characteristics, such as convective overshooting. For the cloud
204 field analyses (using the CvM phase space), the adiabatic curve is taken as a reference
205 for the growth stage of clouds. We find that it makes it very helpful when trying to
206 extract temporal information from those figures, as was shown in previous publications
207 (Heiblum et al., 2016a, 2016b).

208 Regarding the importance of drag on vertical velocity, we address this point in the
209 introduction: “*Usually, the CAPE serves as a theoretical upper limit, and the vertical*
210 *velocity is smaller due to multiple effects (de Roode et al., 2012), most importantly the*
211 *perturbation pressure gradient force (which oppose the air motion) and mixing with*
212 *the environment (entrainment/detrainment) (de Roode et al., 2012; Morrison, 2016a;*
213 *Peters, 2016). Recent studies have shown that entrainment effects on vertical velocity*
214 *are of second order, and a rising thermal shows a balance between buoyancy and the*
215 *perturbation pressure gradient (Hernandez-Deckers and Sherwood, 2016; Romps*
216 *and Charn, 2015), the latter acting as a drag force on the updrafts. Nevertheless,*
217 *initial updraft magnitude and environmental conditions play a crucial role in*
218 *determining the magnitude of mixing effects on buoyancy, and thus also the vertical*
219 *velocity profile in the cloud (Morrison, 2016a, 2016b, 2017).*”.

220 And in the theoretical section:

221 “*Given an initial vertical velocity of ~ 0.5 m/s, the deceleration due to buoyancy (and*
222 *reversal to negative vertical velocity) should occur within a typical time range of 1 -*
223 *10 minutes. These timescales are much longer than the typical timescales of*
224 *entrainment (mixing and evaporation that eliminate the B_{core}) which range between*

225 *1 – 10 s (Lehmann et al., 2009). Moreover, the fact that a drag force typically balances*
226 *the buoyancy acceleration (Romps and Charn, 2015) can also contribute to a time*
227 *lag between effects on buoyancy and subsequent effects on vertical velocity.*
228 *Therefore, the switching of sign for vertical velocity should occur with substantial*
229 *delay compared to the reduction of buoyancy, and B_{core} should be a subset of W_{core}*
230 *(i.e. $B_{core} \subseteq W_{core}$) during the growing and mature stages of a cloud's lifetime.”*

231 Nevertheless, we note that our goal in the paper is not to develop a new parameterization
232 for vertical velocity or gain new insights on the different components of the vertical
233 velocity equation, but rather to get a general understanding of the processes affecting
234 each core type.

235

236 **MC3.2)** The use of the liquid water path as a measure of cloud mass also seems more
237 appropriate for stratocumulus than for cumulus convection, where clouds may be
238 slanted and mixing might lead to lateral growth. Referring to the mean liquid water path
239 as the cloud mass is confusing. Instead of liquid water path, a tracer concentration could
240 be used (e.g. Romps and Kuang, 2010) to provide a more robust measure of dilution.

241

242 **MA3.2)** The use of mean liquid water path (LWP) as a measure of cloud mass is an
243 inherent property of the CvM phase space. The center of gravity (COG) which is taken
244 as the vertical coordinate, can be easily linked to the LWP by using the theoretical case
245 of adiabatic cloud column. So there is a good reference case. In addition we note that
246 the mean LWP is taken by dividing the total mass by the mean cross-sectional area, so
247 that even if clouds are slanted the mean LWP will reflect a “slanted” column. Using
248 here the CvM phase space, as was done for cumulus cloud fields in previous works
249 (Heiblum et al., 2016a, 2016b) enables examination of all clouds, at all stages of
250 lifetime for the entire simulation. So we chose to use the CvM phase space due to its
251 suitability to our purposes while we explain its advantages and limitations in the text:

252 *“In this space, the Center-of-Gravity (COG) height and mass of each cloud in the*
253 *field at each output time step (taken here to be 1 min) are collected and projected in*
254 *the CvM phase space. This enables a compact view of all clouds in the simulation*
255 *during all stages of their lifetimes, with the main disadvantage being the loss of grid-*
256 *size resolution information on in-cloud dynamical processes”.*

257

258 **MC3.3)** The analysis in this study mostly considers convective elements as a whole, or
259 fractions of pixels within an element (figure 2 is an exception here). This provides
260 limited insight into the dynamics of the different cores. Previous studies have provided
261 a more detailed analysis into the circulation around rising cumulus clouds (see, e.g.
262 Blyth et al, 2005; Peters, 2016). Considering the different regions identified in these
263 previous studies would be another way to obtain novel results that go beyond the current
264 conclusions.

265

266 **MA3.3)** We agree with the reviewer that the insights into in-cloud core dynamics is
267 limited. However, as explained in depth above, this is not part of the objectives of this
268 work. We choose to perform a general comparison between the three core types for
269 large statistics of clouds rather than increase the understanding of core dynamics.

270

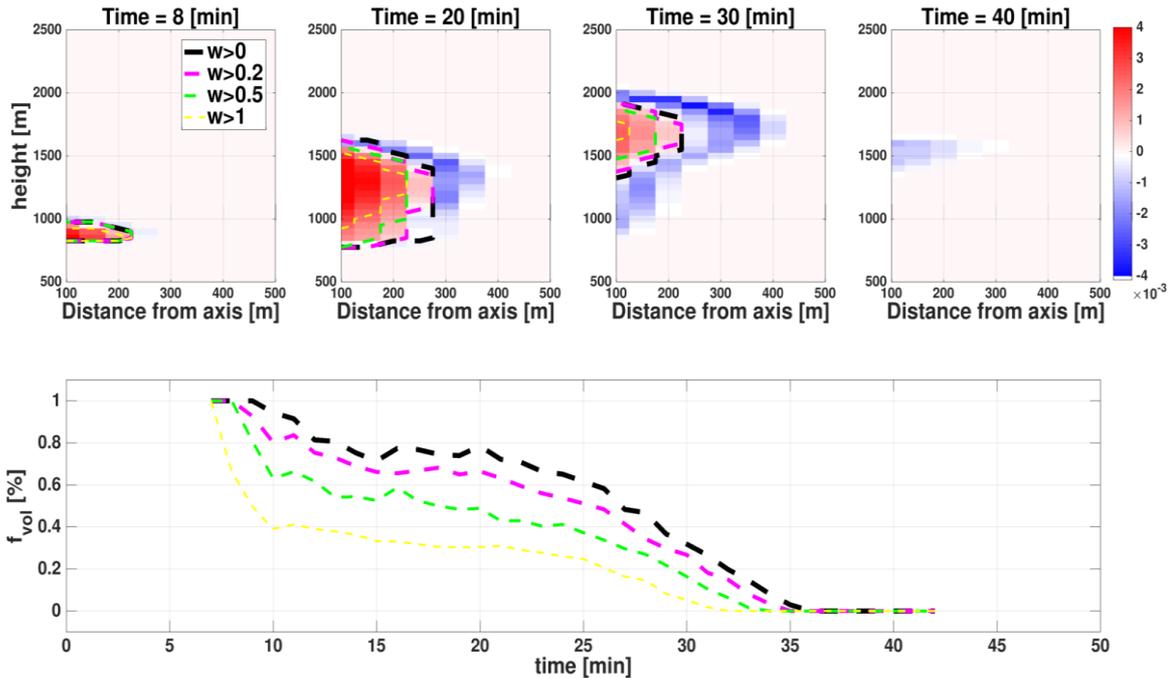
271 **MC3.4)** The thresholds used in the analysis should be discussed further. One of the
272 risks of only considering zero thresholds is that passive regions of the cloud with
273 marginal updraught velocities or regions where gravity waves lead to upwards motion
274 are included in the analysis. Some previous studies have addressed this issue by looking
275 into streamlines (e.g. Romp and Charn, 2015), however, this might not be very
276 straightforward to implement. Alternative approaches would be to determine
277 characteristic updraught values of buoyancy and vertical velocity, or considers multiple
278 thresholds.

279

280 **MA3.4)** The issue of which threshold to choose for the different core types occupied
281 us a great deal. Indeed, choosing a >0 threshold includes passive regions with marginal
282 updrafts increases the variance of the results for the small dissipating clouds. In Fig.
283 RA1 we show the sensitivity to the different threshold for the single cloud case.
284 Increasing the W_{core} threshold significantly affects its extent. We find that up to
285 $w > 0.15$, the W_{core} remains the largest, but for higher thresholds RH_{core} tends to be
286 the largest.

287 Nevertheless, we think the >0 threshold is the only one which is purely physical. All
288 other thresholds are case dependent and can change considerably from case to case. For
289 a study aiming to analyze specific cloud dynamical features it might make sense to
290 apply a strict threshold and limit the variance. But for a comparison we find that the
291 current threshold is the most general. This point is now discussed in the revised text:

292 “We note that setting the core thresholds to positive values (>0) may increase the
 293 amount of non-convective pixels which are classified as part of a physical core,
 294 especially for the W_{core} . Indeed, taking higher thresholds for the updrafts decreases
 295 the W_{core} extent and reduces the variance. Nevertheless, any threshold taken is
 296 subjective in nature, while the positive vertical velocity definition is process based
 297 and objective.”
 298



299
 300 Fig. RAI. Top: Four vertical cross-sections (at $t=8, 20, 30, 40$ minutes) during the
 301 single cloud simulation with aerosol concentration of 500 CCN. Y-axis represents
 302 height [m] and X-axis represents the distance from the axis [m]. The black, magenta,
 303 green and yellow dashed lines represent different vertical velocity core thresholds (see
 304 legend for values). The background represents the condensation (red) and evaporation
 305 rate (blue) [$g\ kg^{-1}\ s^{-1}$]. Bottom: Temporal evolution of vertical velocity core volume
 306 fractions (using different thresholds) from the total cloud volume (f_{vol}).

307
 308 **MC3.5)** The domain used may be too small for the Amazon simulations. In order to
 309 check this, one would need to check that the cloud top is sufficiently far removed from
 310 the domain top, and that convective cold pools are not dominating the spatial
 311 organization of convection.

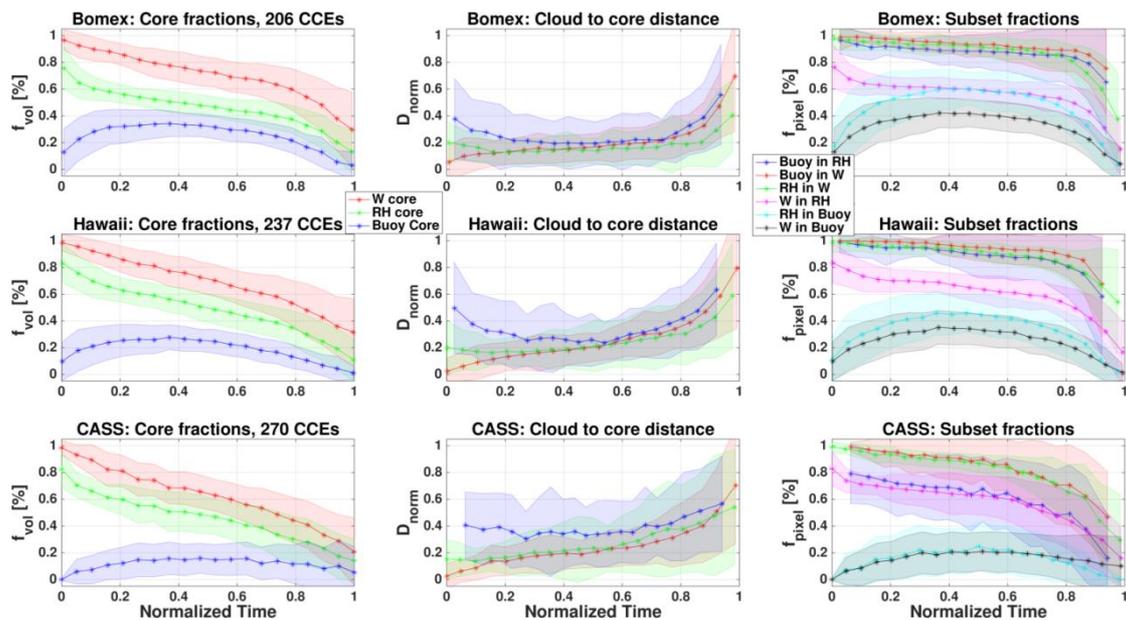
312

313 MA3.5) Thank you for this comment. After reexamination of the Amazon simulations
 314 we found that indeed cold pools play a dominant role in the organization of the cloud
 315 field. Therefore we decided to change the manuscript and show a newer and more
 316 documented case study (CASS - <http://portal.nersc.gov/project/capt/CASS/>).

317 The revised section about Cloud field model (section 2.2): *“To check the robustness
 318 of the cloud field results, two additional case studies are simulated: (1) The same
 319 Hawaiian profile used to initiate the single cloud model, and (2) a continental shallow
 320 cumulus convection cases study (named CASS), based on long term observations
 321 taken at the ARM Southern Great Plains (SGP) site (Zhang et al., 2017).”*

322 This continental case study produced similar results to the oceanic case studies in the
 323 paper as can be seen in the revised fig. 7 below.

324



325

326 Figure 7 (from the revised manuscript). Normalized time series of CCE averaged core
 327 fractions for the BOMEX (upper row), Hawaii (middle row), and CASS (bottom row)
 328 simulations. Both core volume fractions (f_{vol} , left column), normalized distances
 329 between cloud and core centroid locations (D_{norm} , middle column), and pixel fractions
 330 of one core within another (f_{pixel} , right column) are considered. Line colors indicated
 331 different core types (see legends), while corresponding shaded color regions indicate
 332 the standard deviation. Normalized time enables to average together CCEs with
 333 different lifetimes, from formation to dissipation. The number of CCEs averaged
 334 together for each simulation is included in the left column panel titles.

335

336

337

338

339 **Specific comments:**

340

341 **SC1)** Equation 2 seems to be based on a parcel that is not mixing with its environment
342 (see my main concern 1 above).

343

344 **SA1)** Equation 2 is added here to provide a non-expert with a basic understanding of
345 how supersaturation and vertical velocity are thermodynamically linked. For the sake
346 of accuracy, we have clarified that equation 2 (eq. 3 in new manuscript) refers to the
347 adiabatic case which neglects mixing: “*Neglecting mixing with the environment, S*
348 *and w can be linked as follows:*”

349
$$\frac{dS}{dt} = Q_1 w - Q_2 \frac{dq_l}{dt} \quad (3),$$

350 *where Q_1, Q_2 are thermodynamic factors (Rogers and Yau, 1989).”*

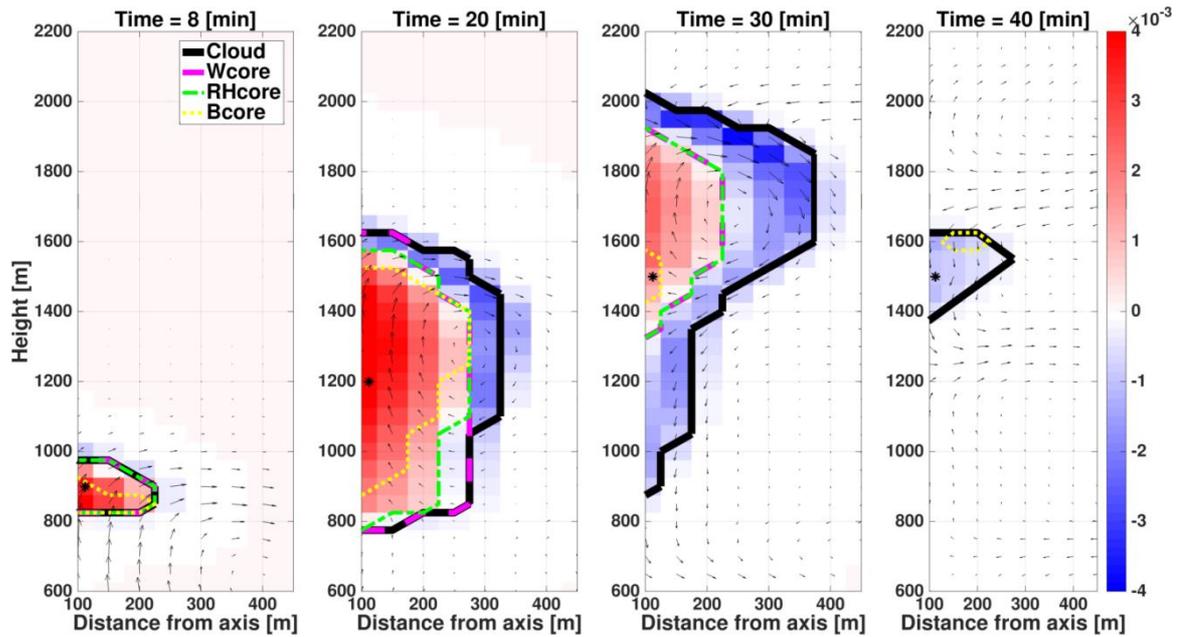
351

352 **SC2)** The presentation style of figures needs improvement. For example, in figure 2,
353 some of the contours overlap, which makes the current presentation confusing to the
354 reader. Some figures are also too small in my opinion (figure 2 is an example here).

355

356 **SA2)** Thank you for this comment. All of the figures were redone so that the texts are
357 larger, features are clearer, and cases where lines overlap can be distinguished. Figure
358 2 is added here for example:

359



360

361 *Figure 2 (from manuscript). Four vertical cross-sections (at $t=8, 20, 30, 40$ minutes)*
 362 *during the single cloud simulation. Y-axis represents height [m] and X-axis represents*
 363 *the distance from the axis [m]. The black, magenta, green and yellow lines represent*
 364 *the cloud, W_{core} , RH_{core} and B_{core} , respectively. The black arrows represent the wind,*
 365 *the background represents the condensation (red) and evaporation rate (blue) [g kg^{-1}*
 366 *s^{-1}], and the black asterisks indicate the vertical location of the cloud centroid. Note*
 367 *that in some cases the lines indicating core boundaries overlap (mainly seen for RH*
 368 *and W cores).*

369

370

371 **SC3)** Equations should avoid the use of acronyms, such as LWC (in any case, is this
 372 specific humidity or mixing ratio?)

373

374 **SA3)** Thank you for this comment, we have replaced LWC with q_l – liquid water mixing
 375 ratio in eq. 3 (see SA1 above), and eq. 1, as follows:

376
$$B = g \cdot \left(\frac{\theta'}{\theta_0} + 0.61q'_v - q_l \right) \quad (1).$$

377

378

379 **SC4)** Some of the terminology is unclear: are cloud growth/cloud suppression regions
 380 simply regions of net increase/decrease of supersaturation, or is something beyond this
 381 meant? It is important to point out that some of the regions that are subsaturated using

382 a bin microphysics scheme would be diagnosed as saturated in an approach to
383 condensation which performs immediate adjustment when defining the RH-core.
384 Similarly, buoyancy is best defined with respect to the surrounding environment, rather
385 than with respect to a reference profile.

386

387 **SA4)** Thank you for noticing this. We have added to the revised text exact definition of
388 cloud growth in the introduction: “...*partitioned to two main regions: i) a core region,*
389 *where mainly cloud growth processes occur (i.e. condensation – accumulation of*
390 *cloud mass), and...*”, the single cloud results: “*During cloud growth (i.e. (increase in*
391 *mass and size)...*“, and cloud field results: “...*fractions decrease with cloud growth*
392 *(increase in mass and COG height) while...*”.

393 In addition, we point out in the text the importance of bin-microphysics that enables
394 cases of sub-saturated cloudy pixels: “*It should be noted that the bin-microphysical*
395 *schemes used here calculate saturation explicitly, by solving the diffusion growth*
396 *equation, enabling super- and sub- saturation values in cloudy pixels. This is in*
397 *contrary to many other works that used bulk-microphysical schemes which rely on*
398 *saturation adjustment to 100% within the cloud (Khain et al., 2015). This difference*
399 *may produce significant differences on the evolution of clouds and their cores* “.

400 Finally, as suggested by the reviewer the buoyancy is taken with respect to the
401 surrounding environment. Citing from the text: “ *B_{core} : buoyancy (see definition in*
402 *Eq. (1)) above zero. The buoyancy is determined in each time step by comparing each*
403 *cloudy pixel with the mean thermodynamic conditions for all non-cloudy pixels per*
404 *vertical height*“.

405

406

407 **SC5)** Parentheses should only be used around the year when an author is cited and the
408 author name is part of the sentence.

409 **SA5)** Thank you, this issue has been addressed.

410

411 **SC6)** I could not find the reference to Dias et al. (2012). I have not done a
412 comprehensive check for other missing references at this point.

413

414 **SA6)** This reference is no longer relevant and has been removed from the text.

415

416 **Reply to reviewer #2 –RC2**

417 **General Reviewer Comment**

418 The work herein seeks to examine and compare three methods of defining convective
419 cores through analysis of buoyancy (B), relative humidity (RH), and vertical velocity
420 (W). The authors do a thorough job of comparing and contrasting the evolution of the
421 various core definitions and highlight the overlap or lack thereof among the 3 defining
422 core characteristics. They have performed their analysis via multiple methods including
423 a theoretical model, single column type model, and a couple of models at the LES scale
424 with bin microphysics and without saturation adjustment assumptions which can be
425 limiting. The results appear quite robust among all methods of representing convective
426 clouds and their cores and among various thermodynamic environments represented by
427 different initial soundings. The manuscript is well-written, clear and concise, but a few
428 questions and concerns, given below, should be addressed.

429

430 **General Answer:** We thank the reviewer for the beneficial comments and we were
431 happy to read that the reviewer found our results robust and the paper well written and
432 clear. The manuscript was revised according to all the comments.

433

434 **Main Comments:**

435

436 **MC1)** The motivation of the paper seems to lack its proper placement with respect to
437 previous published work regarding convective cores and entrainment. While the focus
438 of this work is specific to examining the relative differences between core definitions
439 and their evolution over time, the work should be more appropriately placed in context
440 and should emphasize what is novel in this work.

441

442 **MA1)** Thank you for this comment. In the revised manuscript clear emphasis is put on
443 the novel parts of this work: the differences between the three core types, their
444 evolutions in time, and comparison to previous understanding of core size and location
445 within a cloud. In addition, the introduction was changed significantly to include a
446 broader review of works and ideas from the past that are relevant to this work.

447 We have added a few sentences to the introduction that clarify the objectives of the
448 work: “*Specifically, we aim to answer questions such as:*

- 449 • *Which core type is largest? Which is smallest?*
- 450 • *How do the cores change during the lifetime of a cloud?*
- 451 • *Can different core types be used interchangeably without much effect on*
452 *analysis results?*
- 453 • *Are the cores centered at the cloud’s geometrical center, as expected from the*
454 *core-shell model?”*

455

456 **MC2)** Some aspects of this work regarding entrainment, dilution, and their impacts on
457 buoyancy are not new. However, the framework of comparing cores, core subsets, and
458 their evolution in multiple model frameworks is perhaps more unique. It may help to
459 better frame the paper in such a light.

460

461 **MA2)** Thank you for this comment. The abstract, introduction, and summary in the
462 revised manuscript now put more emphasis on the novelties of the work while referring
463 better to previous works when relevant. Although previous works have dealt with
464 positive vertical velocity and buoyancy in clouds and the effects of entrainment on
465 them, we do not know of a work which tries to perform a comprehensive comparison
466 between the different cores and tracks these cores throughout their lifetime. The new
467 abstract now focuses on these aspects of the work:

468 *“The properties of a warm convective cloud are determined by the competition*
469 *between the growth and dissipation processes occurring within it. One way to observe*
470 *and follow this competition is by partitioning the cloud to core and margin regions.*
471 *Here we look at three core definitions: positive vertical velocity (W_{core}),*
472 *supersaturation (RH_{core}), and positive buoyancy (B_{core}), and follow their evolution*
473 *throughout the lifetime of warm convective clouds.*

474 *Using single cloud and cloud field simulations with bin-microphysics schemes, we*
475 *show that the different core types tend to be subsets of one another in the following*
476 *order: $B_{core} \subseteq RH_{core} \subseteq W_{core}$. This property is seen for several different*
477 *thermodynamic profile initializations, and is generally maintained during the*

478 *growing and mature stages of a cloud's lifetime. This finding is in line with previous*
479 *works and theoretical predictions showing that cumulus clouds may be dominated by*
480 *negative buoyancy at certain stages of their lifetime.*

481 *During its mature growth stage, the cloud and its cores are centered at a similar*
482 *location. During cloud dissipation the cores show less overlap, typically reduce in*
483 *size, and migrate from the cloud centroid. In some cases, buoyancy cores can*
484 *reemerge and often reside at the cloud periphery. Thus, the core-shell model of a*
485 *positively buoyant center surrounded by negatively buoyant shell only applies to a*
486 *fraction of the cloud lifetime.”*

487

488 **Specific Comments:**

489

490 **SC1)** Line 40: Here you mention that negatively buoyant cloud may exist due to $W > 0$
491 and $S > 1$. You might specifically mention the other components of the W equation that
492 keep $W > 0$ and $S > 1$ and their relative contributions during stages of $B > 0$ and $B < 0$.
493 Perhaps this could also be addressed in the main text in greater detail. Once $B < 0$, the
494 other components of the W equation will begin to weaken since the “fuel” is missing.
495 What tends to weaken faster, and what implications does this have for the W core?

496

497 **SA1)** This part was removed from the revised abstract. The existence of negatively
498 buoyant clouds (as referred to in the previous version) can be attributed to inertia or
499 “leftover fuel from sub-cloudy layer buoyancy. The other components of the vertical
500 velocity equation (de Roode et al., 2012; Romps and Charn, 2015) can only decelerate
501 the buoyant updrafts and not actually create a cloud. An exception is large scale
502 advection and quasi-geostrophic ascent, which are irrelevant to the scope of this paper.
503 Many previous works have dealt with the relative importance and feedbacks of the W
504 equation components (de Roode et al., 2012; Morrison, 2016a, 2016b; Romps and
505 Charn, 2015) and we think it is a subject that requires a study on its own. However, we
506 revised the paper to better explain the W equation in the introduction, as follows:

507 *“Neglecting cases of air flow near obstacles or air mass fronts, buoyancy is the main*
508 *source for vertical momentum in the cloud. In its simplest form, the vertical velocity*
509 *(w) in the cloud can be approximated by the convective available potential energy*

510 *(CAPE) of the vertical column up to that height (Rennó and Ingersoll, 1996;*
511 *Williams and Stanfill, 2002; Yano et al., 2005):*

$$512 \quad 0.5w^2(\mathbf{h}) = \int_{h_0}^{\mathbf{h}} \mathbf{B}(z) dz = \mathbf{CAPE}(\mathbf{h}) \quad (2).$$

513 *Here we define CAPE to be the vertical integral of buoyancy from the lowest level of*
514 *positive buoyancy (h_0 , initiation of vertical velocity) to an arbitrary top height (h).*
515 *Usually, the CAPE serves as a theoretical upper limit, and the vertical velocity is*
516 *smaller due to multiple effects (de Roode et al., 2012), most importantly the*
517 *perturbation pressure gradient force (which oppose the air motion) and mixing with*
518 *the environment (entrainment/detrainment) (de Roode et al., 2012; Morrison, 2016a;*
519 *Peters, 2016). Recent studies have shown that entrainment effects on vertical velocity*
520 *are of second order, and a rising thermal shows a balance between buoyancy and the*
521 *perturbation pressure gradient (Hernandez-Deckers and Sherwood, 2016; Romps*
522 *and Charn, 2015), the latter acting as a drag force on the updrafts.”*

523 **SC2)** Lines 177: The potential initial temperature perturbation of 1C is rather large for
524 this type of shallow convection setup. Could such a large perturbation shock the initial
525 field and generate a sizeable convective pulse and gravity waves that impacts the rather
526 small domain size?

527 **SA2)** Thank you for this comment. There was a mistake in the text. The SAM model
528 was initialized with random $\pm 0.1^\circ\text{C}$ (instead of $\pm 0.1\text{-}1^\circ\text{C}$) perturbations throughout the
529 domain. It is corrected in the revised text.

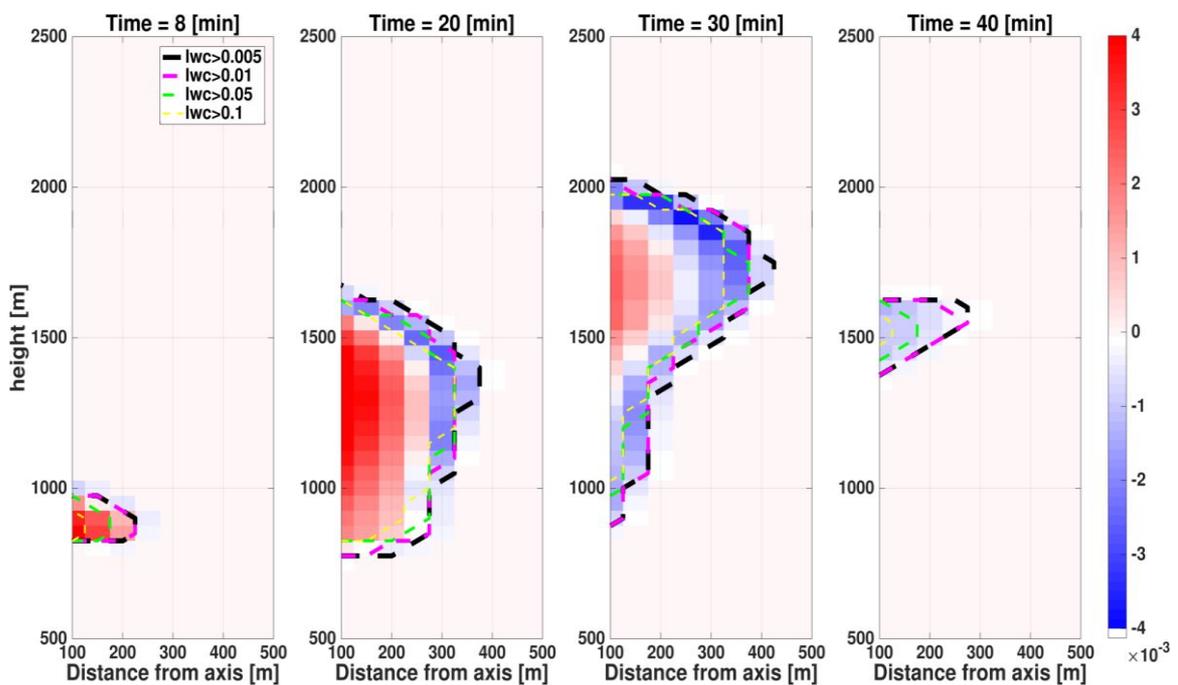
530

531 **SC3)** Line 182: The cloud pixel threshold here of 0.01g/kg seems rather small. What
532 could be deemed a visible cloud would likely be closer to 0.1g/kg. Including values
533 closer to 0.01g/kg would likely include very diffuse clouds at cloud edges that are
534 generated in models. Choosing a different threshold could seemingly have a great
535 impact on the definition of the cloud volume. Have you examined the impact of this
536 threshold choice? I am aware that many papers have used the 0.01 g/kg threshold; but
537 the choice here seems more critical given the examination of cloud volume and such.

538

539 **SA3)** The question of cloud pixel liquid water content (LWC) threshold is something
540 we have examined as part of this work. We started by taking an even lower threshold
541 of 0.005 g/kg (Cohen and Craig, 2006) but eventually raised the threshold to 0.01 g/kg

542 based on other works (Jiang et al., 2009; Xue and Feingold, 2006). The impact of
 543 threshold choice is shown in Fig. RB1 below. The 0.01 and 0.005 g/kg thresholds yield
 544 similar results with regards to cloud volume, while higher thresholds (0.05 and 0.1 g/kg)
 545 reduce cloud volume significantly. By taking areas of condensation and evaporation as
 546 indicators of cloudy regions, it can be seen that the higher values thresholds “miss”
 547 pixels with high evaporation rate (vapor diffusion), in both growing and dissipating
 548 stages of cloud lifetime. Hence, we find that the 0.01 g/kg threshold best reflects a
 549 cloudy volume, without the risk of including insignificant cloud debris as can be seen
 550 in some cases for the lower 0.005 g/kg threshold.
 551



552
 553

554 *Fig. RB1. Four vertical cross-sections (at t=8, 20, 30, 40 minutes) during the single*
 555 *cloud simulation with aerosol concentration of 500 CCN. Y-axis represents height [m]*
 556 *and X-axis represents the distance from the axis [m]. The black, magenta, green and*
 557 *yellow dashed lines represent different LWC thresholds for a cloudy pixel (see legend*
 558 *for values). The background represents the condensation (red) and evaporation rate*
 559 *(blue) [$g\ kg^{-1}\ s^{-1}$].*

560

561 **SC4)** Line 184-186: Here you state that buoyancy is determined relative to the mean
 562 thermodynamic conditions for non-cloud pixels. How is buoyancy computed and

563 applied in the dynamic core of the model? Are these the same or different, and what are
564 the implications if these are different?

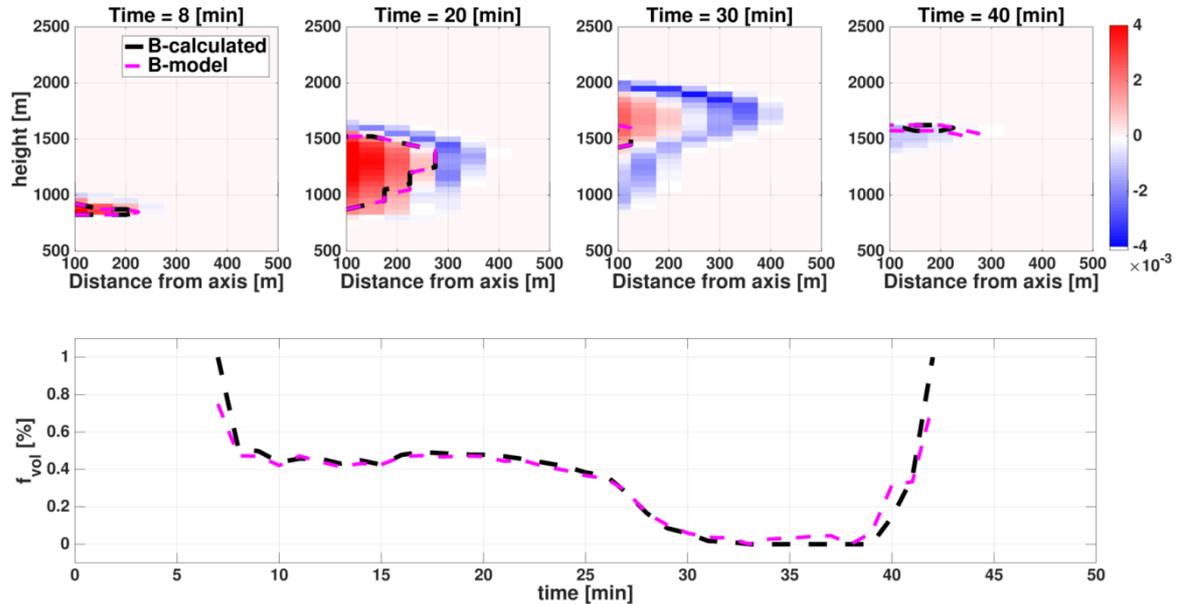
565

566 **SA4)** The buoyancy in the dynamical core of the axisymmetric model is calculated in
567 a similar way to the buoyancy calculations as described in the paper, with the sole
568 difference being the dynamical core buoyancy is calculated with respect to the mean
569 initial thermodynamic conditions while we take the mean instantaneous non-cloudy
570 thermodynamic conditions. Since the domain is sufficiently large and unaffected during
571 the simulation, the differences between the two buoyancy calculations is negligible, as
572 can be seen in Fig. RB2.

573 In the cloud field model (SAM) the dynamical core buoyancy is calculated with respect
574 to the mean horizontal thermodynamic conditions (cloudy and non-cloudy), which
575 gives almost identical results to our calculation in this work (i.e. with respect to only
576 non-cloudy).

577 Since each of the models' dynamical core calculates buoyancy a bit differently, we
578 chose one calculation that applies to both.

579



580

581 *Fig. RB2. Comparison of dynamical core buoyancy (magenta lines) with calculated*
582 *buoyancy (black lines). The top panels are similar to Fig. RB1, but with lines*
583 *representing core extent. The bottom panel shows the temporal evolution of buoyancy*
584 *core volume fraction from the total cloud volume (f_{vol}).*

585

586 **SC5)** Line 257: Here you state that the cloud top downdraft promotes adiabatic heating
587 that leads to the decay phase positive buoyancy. Is this definitive or supposition here?
588 Is this seen in other clouds? Is this adiabatic heating greater than any local evaporative
589 cooling?

590

591 **SA5)** A significant part of Part II of this work was devoted to the explanation of why
592 pockets of positive buoyancy appear in non-convective regions of dissipating clouds.
593 We show that if the evaporative cooling is weak enough (or no evaporation occurs), the
594 adiabatic heating is sufficient to create positive buoyancy in weak downdrafts. The
595 reader is referred to Part II within the text for the single cloud:

596 *“Further analysis (see Part II) shows that the entire dissipating cloud is colder and*
597 *more humid than the environment but downdrafts from the cloud top (see arrows in*
598 *Fig. 2) promote adiabatic heating, and by that increase the buoyancy in dissipating*
599 *cloudy pixels, sometimes reaching positive values. These buoyant pockets will be*
600 *discussed further in Part II.* “.

601 and cloud field:

602 *“The prevalence of cloud edge B_{core} pixels during dissipation can be explained by*
603 *adiabatic heating due to weak downdrafts (see Sect. 4.2, Part II) which are expected*
604 *at the cloud periphery.”.*

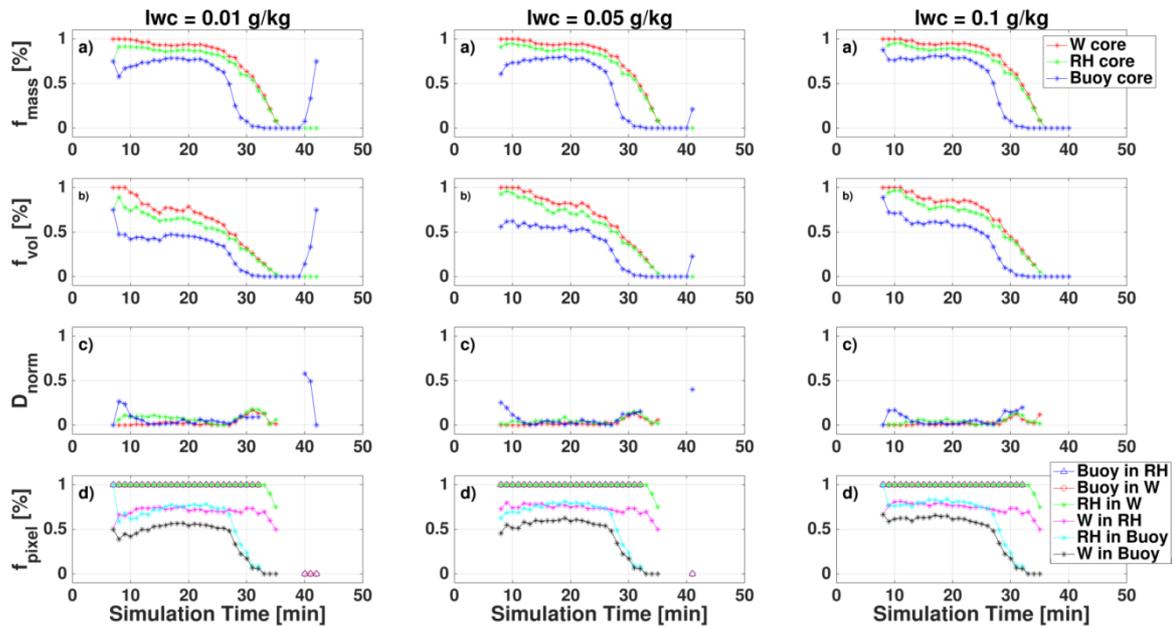
605

606 **SC6)** Line 268-270: Are the changes in cloud volume fraction susceptible to the choice
607 of cloud mass concentration used to define a cloud grid cell (0.01 g/kg)? How would
608 choosing a different threshold impact your analysis?

609

610 **SA6)** We have tested this question as part of this work and found that the main
611 conclusions would not have changed regardless of the LWC threshold chosen. This fact
612 is demonstrated in Fig. RB3 for the single cloud case, where it can be seen that the
613 subset properties of the three cores and their relative sizes are similar. The main
614 difference that arises is the positive buoyancy core that appears during dissipation only
615 for lower cloud LWC thresholds. However, we find this effect to be substantial in the
616 cloud field simulation and for other aerosol concentrations, and thus should not be
617 considered an outlier only seen for very low LWC pixels. An additional figure for a low
618 aerosol concentration of 25 CCN is also shown below (Fig. RB4), where an increase in
619 buoyancy core during dissipation is seen for all thresholds.

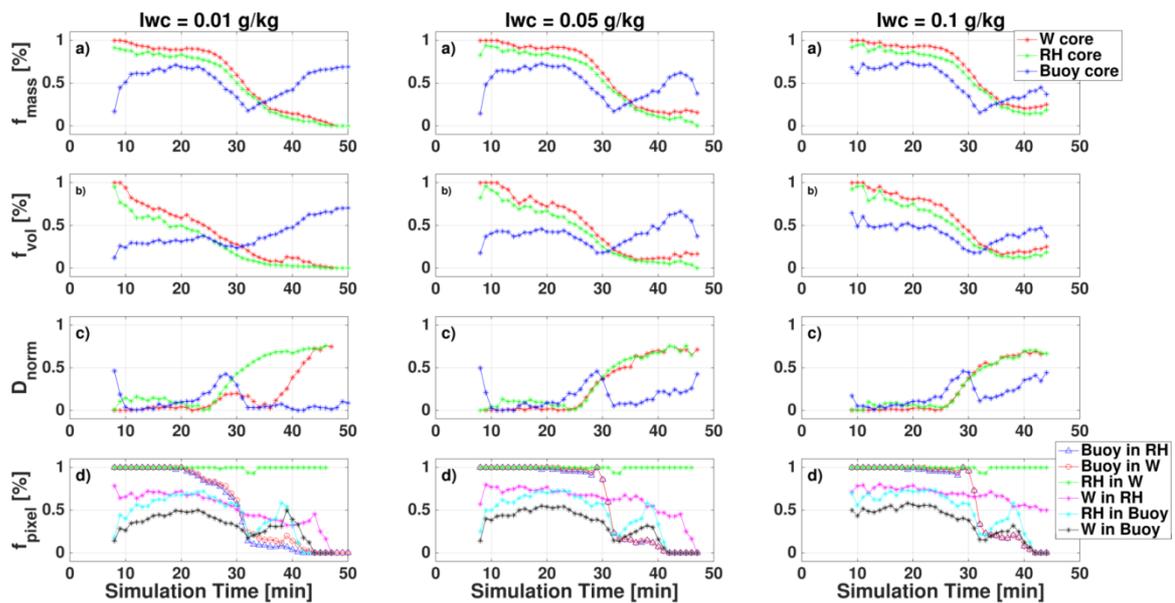
620



621

622 Fig. RB3. Same as figure 3 in the manuscript, but for three different cloudy pixel LWC
623 thresholds [g/kg]: 0.01 (left column), 0.05 (middle column), 0.1 (right column). Aerosol
624 concentration is 500 CCN.

625



626

627 Fig. RB4 Same as figure RB3, but for an aerosol concentration of 25 CCN.

628

629 **SC7)** Line 722: How valid is this non-changing temperature assumption to your
630 analysis?

631 This seems like a rather unrealistic and constricting assumption. The local dT could be
632 large which could greatly impact dB and mixing.

633 **SA7)** The non-changing temperature assumption only applies to the reference
634 environmental temperature. This assumption is based on the fact that the environment
635 is sufficiently large and its mean temperature is not affected by local evaporation. We
636 find this assumption to be standard practice for almost all models calculating buoyancy
637 (Khairoutdinov and Randall, 2003; Seigel, 2014), which take the horizontal mean
638 temperature as reference (which changes very slowly during the course of a simulation,
639 if at all), rather than a local temperature in the vicinity of a cloud.

640

641 **SC8)** Line 267: “expect” should be “except”.

642 **SA8)** Thank you, the change was carried out.

643

644 **SC9)** Line 371: “overweighs” should be “outweighs”.

645 **SA9)** Thank you, the change was carried out.

646

647 **SC10)** Line 591: “cloud’s” should be “clouds”.

648 **SA10)** Thank you, the typo was fixed

649

650 **SC11)** Line 619: “from precipitation” should be “by precipitation”.

651 **SA11)** Thank you, the change was carried out.

652

653 **SC12)** Line 634: This should read: “In cases where the: : :.”

654 **SA12)** Thank you, we added “where” to the sentence.

655

656 **SC13)** Figures: My main comment about the figures is that most of them need to be
657 larger, especially the fonts, so that they are easily readable. The time series plots need
658 to be much large in order to see overlap where it exists.

659 **SC13)** Thank you for this comment. All of the figures were redone so that the texts are
660 larger and cases where lines overlap can be distinguished.

661

662

663

664

665

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792 **Core and margin in warm convective clouds. Part I: core types and evolution**
793 **during a cloud's lifetime**

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807

808 **Abstract**

809 The properties of a warm convective cloud are determined by the competition between
810 the growth and dissipation processes occurring within it. One way to observe and follow
811 this competition is by partitioning the cloud to core and margin regions. Here we look
812 at three core definitions: positive vertical velocity (W_{core}), supersaturation (RH_{core}),
813 and positive buoyancy (B_{core}), and follow their evolution throughout the lifetime of
814 warm convective clouds.

815 Using single cloud and cloud field simulations with bin-microphysics schemes, we
816 show that the different core types tend to be ~~proper~~ subsets of one another in the
817 following order: $B_{core} \subseteq RH_{core} \subseteq W_{core}$. ~~Using single cloud and cloud field~~
818 ~~simulations, we find that~~ This property is seen for several different thermodynamic
819 profile initializations, and is generally maintained during the growing and mature stages
820 of a cloud's lifetime, ~~but can break down during the dissipation stage.~~ This finding is in
821 line with previous works and theoretical predictions showing that cumulus clouds may
822 be dominated by negative buoyancy at certain all stages of their lifetime.

823 During its mature and growth stage, ~~the~~ the cloud and its cores are centered at a similar
824 location, ~~while~~ During cloud dissipation the cores show less overlap, typically reduce
825 in size, and migrate from the cloud centroid. In some cases, buoyancy cores can
826 reemerge and often may reside at the cloud periphery. Thus, the core-shell model of a
827 positively buoyant center surrounded by negatively buoyant shell only applies to a
828 fraction of the cloud lifetime.

829 ~~A theoretical model is developed, showing that in both the adiabatic and non-adiabatic~~
830 ~~cases, B_{core} can be expected to be the smallest core, due to two main reasons: i)~~
831 ~~entrainment rapidly decreases the buoyancy core compared to the other core types, and~~
832 ~~ii) convective clouds may exist while being completely negatively buoyant (while~~
833 ~~maintaining positive vertical velocity and supersaturation).~~

834

835 1. Introduction

836 Clouds are important players in the climate system (Trenberth et al., 2009), and
837 currently constitute one of the largest uncertainties in climate and climate change
838 research (IPCC, 2013). One of the reasons for this large uncertainty is the complexity
839 created by opposing processes that occur at the same time but in different locations
840 within a cloud. Although a cloud is generally considered as a single entity, physically,
841 it can be partitioned to two main regions: i) a core region, where mainly cloud growth
842 processes occur ([i.e. condensation - accumulation of cloud mass](#)), and ii) a margin
843 region, where cloud suppression processes occur ([i.e. evaporation - loss of cloud mass](#)).
844 Changes in thermodynamic or microphysical (aerosol) conditions impact the processes
845 in both regions (sometimes in different ways), and thus the resultant total cloud
846 properties (Dagan et al., 2015). To better understand cloud properties and their
847 evolution in time, it is necessary to understand the interplay between physical processes
848 within the core and margin regions (and the way they are affected by perturbations in
849 the environmental conditions).

850 Considering convective clouds, there are several ~~parameters-objective measures~~ that ~~are~~
851 ~~commonly have been~~ used ~~in previous works~~ for separating a cloud's core from its
852 margins (will be referred to as physical cores hereafter). ~~Previous works have used these~~
853 ~~objective measures to define a cloud core (with the margins defined as the remaining~~
854 ~~regions of the cloud)~~. In deep convective cloud simulations the core is usually defined
855 by the updrafts' magnitude using a certain threshold, usually $W > 1 \text{ m}\cdot\text{s}^{-1}$ (Khairoutdinov
856 et al., 2009; Kumar et al., 2015; Lebo and Seinfeld, 2011; Morrison, 2012). ~~Studies on~~
857 ~~warm cumulus clouds studied the main parameters that affect warm cumulus clouds~~
858 ~~vertical velocity and have~~ defined the clouds' core as parts with positive buoyancy and
859 positive updrafts (de Roode et al., 2012; Dawe and Austin, 2012; Heus and Jonker,
860 2008; Siebesma and Cuijpers, 1995). ~~or solely regions with positively buoyancy~~ (Heus
861 and Seifert, 2013; Seigel, 2014). ~~More recently, cloud partition to regions of~~
862 ~~supersaturation and sub-saturation has been used to define the cloud core in single cloud~~
863 ~~simulations~~ (Dagan et al., 2015).

864 For simplicity, we focus ~~here~~ on warm convective clouds (only contain liquid water),
865 avoiding the additional complexity and uncertainties associated with mixed phase and
866 ice phase microphysics. ~~The common assumption when partitioning a convective cloud~~

867 to its physical core and margin is that the cloud core is at its geometrical center ~~is its~~
868 ~~core~~ and the peripheral regions (i.e. edges) are the margin. Previous observational
869 (Heus et al., 2009a; Rodts et al., 2003; Wang et al., 2009) and numerical (Heus and
870 Jonker, 2008; Jonker et al., 2008; Seigel, 2014) works have studied the gradients of
871 cloud thermodynamic properties from cloud center to edge, and suggest that a cloud is
872 best described by a core-shell model. This model assumes a core with positive vertical
873 velocity and buoyancy, surrounded by a shell with negative vertical velocity and
874 buoyancy. The shell is the region where mixing between cloudy and environmental air
875 parcels occurs, leading to evaporative cooling → decrease in buoyancy → decrease in
876 vertical velocity.

877 Based on previous ~~works~~ findings, here we explore the partition of clouds to core and
878 margin using three different objective core definitions where the cloud core threshold
879 is set to be a positive value (of buoyancy, vertical velocity, or supersaturation). Cloud
880 buoyancy (B) (~~which is the driving force for convection~~) is one of the intuitive
881 parameters used ~~and~~ can be approximated by the following formula:

$$882 \quad B = g \cdot \left(\frac{\theta'}{\theta_0} + 0.61q'_v - q_l \right) \quad (1),$$

883 Where θ_0 represents the reference state potential temperature, q_v is the water vapor
884 mixing ratio, and ~~LWC~~ q_l is the liquid water content. The (') stands for the deviation
885 from the reference state per height (Wang et al., 2009). Buoyancy is a measure for the
886 vertical acceleration and its integral is the convective potential energy, ~~or the fuel that~~
887 ~~drives cloud growth~~. Latent heat release during moist adiabatic ascent fuels positive
888 buoyancy and clouds' growth, while evaporation and subsequent cooling drives cloud
889 decay (de Roode, 2008; Betts, 1973). ~~The existene~~prevalence of negatively buoyancy
890 parcels at the cloud edges due to mixing and evaporation is a well-known phenomena
891 (Morrison, 2017). Mixing diagrams have been used to assess this effect (de Roode,
892 2008; Paluch, 1979; Taylor and Baker, 1991), and are at the root of convective
893 parameterization schemes (Emanuel, 1991; Gregory and Rowntree, 1990; Kain and
894 Fritsch, 1990) and parameterizations of entrainment and detrainment in cumulus clouds
895 (de Rooy and Siebesma, 2008; Derbyshire et al., 2011).

896 Neglecting cases of air flow near obstacles or air mass fronts, buoyancy is the main
897 source for vertical momentum in the cloud. In its simplest form, the vertical velocity

(w) in the cloud can be approximated by the convective available potential energy (CAPE) of the vertical column up to that height (Rennó and Ingersoll, 1996; Williams and Stanfill, 2002; Yano et al., 2005):

$$0.5w^2(h) = \int_{h_0}^h B(z) dz = CAPE(h) \quad (2).$$

Here we define CAPE to be the vertical integral of buoyancy from the lowest level of positive buoyancy (h_0 , initiation of vertical velocity) to an arbitrary top height (h). Usually, the CAPE serves as a theoretical upper limit, and the vertical velocity is smaller due to multiple effects (de Roode et al., 2012), most importantly the perturbation pressure gradient force (which oppose the air motion) and mixing with the environment (entrainment/detrainment) (de Roode et al., 2012; Morrison, 2016a; Peters, 2016). Recent studies have shown that entrainment effects on vertical velocity can be neglected are of second order, and a rising thermal shows a balance between buoyancy and the perturbation pressure gradient (Hernandez-Deckers and Sherwood, 2016; Romps and Charn, 2015), the latter acting as a drag force on the updrafts. Nevertheless, initial updraft and environmental conditions play a crucial role in determining the magnitude of mixing effects on buoyancy, and thus also the vertical velocity profile in the cloud (Morrison, 2016a, 2016b, 2017).

The supersaturation (S , where $S=1$ is 100% relative humidity) core definition ($S-1>0$ or $RH>100\%$) partitions the cloud core and margin to areas of condensation and evaporation. Since we consider convective clouds—here, the only driver of supersaturation during cloud growth is upward vertical motion of air. Thus, the vertical velocity core partitions the cloud to areas where the saturation ratio increases (upward motion) or decreases (downward motion). The vertical velocity (w) and the supersaturation (S , where $S=1$ is 100% relative humidity) can also be used for defining a cloud core, core. Neglecting mixing with the environment, they S and w can be and are linked as follows:

$$\frac{dS}{dt} = Q_1 w - Q_2 \frac{dLWCq_l}{dt} \quad (3),$$

where Q_1, Q_2 are thermodynamic factors (Rogers and Yau, 1989). The thermodynamic factors are nearly insensitive to pressure for temperature above 0°C , and both weakly decrease (less than 15% net change) with temperature increase between 0°C and 30°C

928 (Pinsky et al., 2013). The first term on the right-hand side is related to the change in the
929 supersaturation due to adiabatic cooling or heating of the moist air (due to vertical
930 motion). The second term is related to the change in the supersaturation due to
931 condensation/evaporation of water vapor/drops. Hence, the supersaturation in a rising
932 parcel depends on the magnitude of the updraft and on the condensation rate of vapor
933 to drops (a sink term). The latter is proportional to the concentration of aerosols in the
934 cloud (Reutter et al., 2009; Seiki and Nakajima, 2014), which serve as cloud
935 condensation nuclei (CCN) for cloud droplets. In Part II of this work we demonstrate
936 some of the insights gained by investigating differences between the different cores
937 properties and their time evolution when changing the aerosol loading.

938 The ~~goals-purpose~~ of this part of the work (part I) are to compare and understand the
939 differences between the three basic definitions of cloud core (i.e. W_{core} , RH_{core} , B_{core})
940 throughout a convective cloud's lifetime, using both theoretical arguments and
941 numerical simulations. It should be noted that the bin-microphysical schemes used here
942 calculate saturation explicitly, by solving the diffusion growth equation, enabling
943 super- and sub- saturation values in cloudy pixels. This is in contrary to many other
944 works that used bulk-microphysical schemes which rely on saturation adjustment to
945 100% within the cloud (Khain et al., 2015). This difference may produce significant
946 differences on the evolution of clouds and their cores. Specifically, we aim to answer
947 questions such as:

- 948 • Which core type is largest? Which is smallest?
- 949 • How do the cores change during the lifetime of a cloud?
- 950 • Can different core types be used interchangeably without much effect on
951 analysis results?
- 952 • Are the cores centered around the cloud' geometrical center, as expected from
953 the core-shell model?

954 The differences between the cores' evolution in time shed new light on the competition
955 of processes within a cloud in time and space. Moreover, such an understanding can
956 serve as a guideline to all studies that perform the partition to cloud core and margin,
957 and assist in determining the relevance of a given partition.

958 **2. Methods**

959 **2.1. Single cloud model**

960 For single cloud simulations we use the Tel-Aviv University axisymmetric, non-
961 hydrostatic, warm convective single cloud model (TAU-CM). It includes a detailed
962 (explicit) treatment of warm cloud microphysical processes solved by the multi-
963 moment bin method (Feingold et al., 1988, 1991; Tzivion (Tzitzvashvili) et al., 1989;
964 Tzivion et al., 1994). The warm microphysical processes included in the model are
965 nucleation, diffusion (i.e. condensation and evaporation), collisional coalescence,
966 breakup and sedimentation (for a more detailed description, see (Reisin et al., 1996)).

967 Convection was initiated using a thermal perturbation near the surface. A time step of
968 1 sec is chosen for dynamical computations, and 0.5 sec for the microphysical
969 computations (e.g. condensation-evaporation). The total simulation time is 80 min.
970 There are no radiation processes in the model. The domain size is 5x6 km, with an
971 isotropic 50 m resolution. The model is initialized using a Hawaiian thermodynamic
972 profile, based on the 91285 PHTO Hilo radiosonde at 00Z, 21 Aug, 2007. A typical
973 oceanic size distribution of aerosols is chosen (Altartatz et al., 2008; Jaenicke, 1988),
974 with a total concentration of 500 cm^{-3} . This concentration produced clouds that are non-
975 to weakly- precipitating. In Part II additional aerosol concentrations are considered,
976 including ones which produce heavy precipitation.

977

978 **2.2. Cloud field model**

979 Warm cumulus cloud fields are simulated using the System for Atmospheric Modeling
980 (SAM) Model (version 6.10.3, for details see webpage:
981 <http://rossby.msrc.sunysb.edu/~marat/SAM.html>) (Khairoutdinov and Randall, 2003)).
982 SAM is a non-hydrostatic, anelastic model. Cyclic horizontal boundary conditions are
983 used together with damping of gravity waves and maintaining temperature and moisture
984 gradients at the model top. An explicit Spectral Bin Microphysics (SBM) scheme
985 (Khain et al., 2004) is used. The scheme solves the same warm microphysical processes
986 as in the TAU-CM single cloud model, and uses an identical aerosol size distribution
987 and concentration (i.e. 500 cm^{-3}) for the droplet activation process.

988 We use the BOMEX case study as our benchmark for shallow warm cumulus fields.
989 This case simulates a trade-wind cumulus (TCu) cloud field based on observations
990 made near Barbados during June 1969 (Holland and Rasmusson, 1973). This case study
991 has a well-established initialization setup (sounding, surface fluxes, and surface
992 roughness) and large scale forcing setup (Siebesma et al., 2003). It has been thoroughly
993 tested in many previous studies (Grabowski and Jarecka, 2015; Heus et al., 2009b; Jiang
994 and Feingold, 2006; Xue and Feingold, 2006). To check the robustness of the cloud
995 field results, two additional case studies are simulated: (1) The same Hawaiian profile
996 used to initiate the single cloud model, and (2) an Amazonian warm cumulus case based
997 on the afternoon dry season mean profile for August 2001 obtained during the Large-
998 scale Biosphere-Atmosphere (LBA) experiment data at Belterra, Brazil (Dias et al.,
999 2012).

1000 All three soundings (BOMEX, Hawaiian, and Amazonian) and surface properties used
1001 to initialize the model are detailed in (Heiblum et al., 2016a) . The grid size is set to
1002 100 m in the horizontal direction and 40 m in the vertical direction for all simulations.
1003 The domain size is 12.8 km x 12.8 km x 4 km for the BOMEX simulation and extends
1004 to 5 km, 6 km in the vertical direction for the Hawaii and Amazon simulations,
1005 respectively. The time step for computation is 1 s for all simulations, with a total
1006 runtime of 8 hours. The initial temperature perturbations (randomly chosen within \pm
1007 0.1-1 °C) are applied near the surface, during the first time step.

1008

1009 **2.3. Physical and Geometrical Core definitions**

1010 A cloudy pixel is defined here as a grid-box with liquid water amount that exceeds 0.01
1011 g kg⁻¹. The physical core of the cloud is defined using three different definitions: 1)
1012 RH_{core} : all grid boxes for which the relative humidity (RH) exceeds 100%, 2) B_{core} :
1013 buoyancy (see definition in Eq. (1)) above zero. The buoyancy is determined in each
1014 time step by comparing each cloudy pixel with the mean thermodynamic conditions for
1015 all non-cloudy pixels per vertical height, and 3) W_{core} : vertical velocity above zero.
1016 These definitions apply for both the single cloud and cloud field model simulations
1017 used here. [We note that setting the core thresholds to positive values \(>0\) may increase](#)
1018 [the amount of non-convective pixels which are classified as part of a physical core,](#)
1019 [especially for the \$W_{core}\$. Indeed, taking higher thresholds for the updrafts decreases the](#)

1020 W_{core} extent and reduces the variance. Nevertheless, any threshold taken is subjective
1021 in nature, while the positive vertical velocity definition is process based and objective.
1022 Additional thresholds have also been checked for the updrafts or buoyancy definitions,
1023 yielding similar conclusions.

1024 The centroid (i.e. mean location in each of the axes) is used here to represent the
1025 geometrical location of the total cloud (i.e. cloud geometrical core) and its specific
1026 physical cores. The distances between the total cloud and its cores' centroids (D_{norm}),
1027 as presented here, are normalized to cloud size to reflect the relative distance between
1028 the two centroids, where $D_{norm} = 0$ indicates coincident physical and geometrical
1029 cores and $D_{norm} = 1$ indicates a physical core located at the cloud boundary. The
1030 single cloud simulations rely on an axisymmetric model and thus all centroids are
1031 horizontally located on the center axis while vertical deviations are permitted. For this
1032 model the distance is normalized by half the cloud's thickness. For the cloud field
1033 simulations both horizontal and vertical deviations are possible, therefore distances are
1034 normalized by the cloud's volume radius.

1035

1036 **2.4. Center of gravity vs. Mass (CvM) phase space**

1037 Recent studies (Heiblum et al., 2016a, 2016b) suggested the Center-of-Gravity vs. Mass
1038 (CvM) phase space as a useful approach to reduce the high dimensionality and to study
1039 results of large statistics of clouds during different stages of their lifetimes (such as seen
1040 in cloud fields). In this space, the Center-of-Gravity (COG) height and mass of each
1041 cloud in the field at each output time step (taken here to be 1 min) are collected and
1042 projected in the CvM phase space. This enables a compact view of all clouds in the
1043 simulation during all stages of their lifetimes. Although the scatter of clouds in the CvM
1044 is sensitive to the microphysical and thermodynamic settings of the cloud field, it was
1045 shown that the different subspaces in the CvM space correspond to different cloud
1046 processes and stages (Heiblum et al., 2016a, 2016b). The lifetime of a cloud can be
1047 described by a trajectory on this phase space.

1048 A schematic illustration of the CvM space is shown in Fig. 1. Most clouds are confined
1049 between the adiabat (curved dashed line) and the inversion layer base (horizontal
1050 dashed line). The adiabat curve corresponds to the theoretical evolution of a moist

1051 adiabat 1D cloud column in the CvM space. The large majority of clouds form within
1052 the growing branch (yellow shade) at the bottom left part of the space, adjacent to the
1053 adiabat. Clouds then follow the growing trajectory (grow in both COG and mass) to
1054 some maximal values. The growing branch deviates from the adiabat at large masses
1055 depending on the degree of sub-adiabaticity of the cloud field. After or during the
1056 growth stage of clouds, they may undergo the following processes: i) dissipate via a
1057 reverse trajectory along the growing one, ii) dissipate via a gradual dissipation
1058 trajectory (magenta shade), iii) shed off small mass cloud fragments (red shades), iv) in
1059 the case of precipitating clouds, they can shed off cloud fragments in the sub-cloudy
1060 layer (grey shade). The former two processes form continuous trajectories in the CvM
1061 space, while the latter two processes create disconnected subspaces.

1062

1063 **2.5. Cloud tracking**

1064 To follow the evolution of individual clouds within a cloud field we use an automated
1065 3D cloud tracking algorithm (see (Heiblum et al., 2016a) for details). It enables tracking
1066 of Continuous Cloud Entities (CCEs) from formation to dissipation, even if interactions
1067 between clouds (splitting or merging) occur during that lifetime. A CCE initiates as a
1068 new cloud forming in the field, and is tracked on the condition that it retains the majority
1069 (>50%) of its mass during an interaction event if occurs. Thus, a CCE can terminate
1070 due to either cloud dissipation or cloud interactions.

1071

1072 **3. Theoretical estimations for different core sizes considerations explaining the** 1073 **single-cloud-simulation-results**

1074 Here we propose simple physical considerations that predict to evaluate the simulated
1075 differences in cloud partition to core and margin using different definitions. The
1076 arguments rely on key findings from previous works (see Sect. 1) with aim to
1077 summarize our present understanding gain intuitive understanding of the potential
1078 differences between the core types. It is convenient to separate the analysis to an
1079 adiabatic case, and then add another layer of complexity and consider the effects of
1080 mixing of cloudy and non-cloudy air. In this theoretical derivation saturation

1081 [adjustment to RH=100% is assumed for both cases, while in the other models used in](#)
1082 [this study transient super- and sub-saturated cloudy parcels are treated \(more realistic\).](#)

1083

1084 **3.1. Adiabatic [case – no mixing](#)**

1085 ~~For the case of an adiabatic cloud column~~[Considering moist-adiabatic ascent](#), the
1086 excess vapor above saturation is instantaneously converted to liquid (saturation
1087 adjustment). Thus, the adiabatic cloud is saturated ($S=1$) throughout its vertical profile,
1088 and only W_{core} and B_{core} differences can be considered. It is assumed that the adiabatic
1089 convective cloud is initiated by positive buoyancy initiating from the sub-cloudy layer.
1090 As long as the cloud is growing it should have positive CAPE and will experience
1091 positive w throughout the column even if the local buoyancy at specific height is
1092 negative. Eventually the cloud must decelerate due to negative buoyancy and reach a
1093 top height, where $CAPE = 0$ and $w = 0$. [Hence, for the adiabatic column case, \$B_{core}\$ is](#)
1094 [always a proper subset of \$W_{core}\$. \(\$B_{core} \subset W_{core}\$ \).](#) [These effects are commonly seen in](#)
1095 [warm convective cloud fields where permanent vertical layers of negative buoyancy](#)
1096 [\(but with updrafts\) within clouds typically exist at the bottom and top regions of the](#)
1097 [cloudy layer](#) (de Roode and Bretherton, 2003; Betts, 1973; Garstang and Betts, 1974;
1098 Grant and Lock, 2004; Neggers et al., 2007).

1099

1100 **3.2. Cloud parcel entrainment model**

1101 [A mixing model between a saturated \(cloudy\) parcel and a dry \(environment\) parcel is](#)
1102 [used to illustrate the effects of mixing on the different core types.](#) The details of these
1103 theoretical calculations are shown in Appendix A. The initial cloudy parcel is assumed
1104 to be saturated (part of RH_{core}), have positive vertical velocity (part of W_{core}), and
1105 experience either positive or negative buoyancy (part of B_{core} or B_{margin}), as is seen
1106 for the adiabatic column case. Additionally, mixing is assumed to be isobaric, and in a
1107 steady environment where the average temperature of the environment per a given
1108 height does not change. The resultant mixed parcel will have lower humidity content
1109 and lower LWC as compared to the initial cloudy parcel, and a new temperature. In
1110 nearly all cases (beside in an extremely humid environment) the mixed parcel will be

1111 sub-saturated and evaporation of LWC will occur. Evaporation ceases when
1112 equilibrium is reached due to air saturation ($S=1$) or due to complete evaporation of the
1113 droplets (which means $S<1$, and the mixed parcel is no longer cloudy since it has no
1114 liquid water content).

1115 In addition to mixing between cloudy (core or margin) and non-cloudy parcels, mixing
1116 between core and margin parcels (within the cloud) also occurs. This mixing process
1117 can be considered as “entrainment-like” with respect to the cloud core. Considering the
1118 changes in the W_{core} and RH_{core} , there is no fundamental difference in the treatment of
1119 mixing of cloudy and non-cloudy parcels, or mixing between core and margin (because
1120 the margins and the environment are typically sub-saturated and experience negative
1121 vertical velocity). However, for the changes in the B_{core} after mixing, there exists a
1122 fundamental difference between mixing *with* the reference temperature/humidity state
1123 (in the case of mixing with the environment) and mixing *given* a reference
1124 temperature/humidity state (in mixing between B_{core} and B_{margin}). Thus, it is
1125 interesting to check the effects of mixing between B_{core} and B_{margin} parcels on the
1126 total extent of the B_{core} with respect to the other two core types. The details of this
1127 second case are shown in Appendix B.

1128

1129 **3.2.1. Effects of non-cloudy entrainment on buoyancy**

1130 When mixed with non-cloudy air, the change in buoyancy of the initial cloudy parcel
1131 (which is a part of W_{core} and RH_{core} and either B_{core} or B_{margin}) happens due to both
1132 mixing and evaporation processes. The theoretical calculations show that for all
1133 relevant temperatures ($\sim 0^\circ\text{C}$ to 30°C , representing warm Cu), the change in the parcel’s
1134 buoyancy due to evaporation alone will always be negative (see appendix A). It is
1135 because the negative effect of the temperature decrease ~~overweighs~~ outweighs the
1136 positive effects of the humidity increase and water loading decrease. Nevertheless, the
1137 total change in the buoyancy (due to both mixing and evaporation) depends on the
1138 initial temperature, relative humidity, and liquid water content of the cloudy and non-
1139 cloudy parcels.

1140 In Fig. A1 a wide range of non-cloudy environmental parcels, each with their own
1141 thermodynamic conditions, are mixed with a saturated cloud parcel with either positive

1142 or negative buoyancy. The main conclusions regarding the effects of such mixing on
1143 the buoyancy are as follows:

1144 i. To a first order, the initial buoyancy values are temperature dependent,
1145 where a cloudy parcel that is warmer (colder) by more than $\sim 0.2^\circ\text{C}$ than
1146 the environment will be positively (negatively) buoyant for common
1147 values of cloudy layer environment relative humidity ($\text{RH} > 80\%$).

1148 ii. Parcels that are initially part of B_{core} may only lower their buoyancy
1149 due to entrainment, either to positive or negative values depending on
1150 the environmental conditions.

1151 iii. The lower the environmental RH, the larger the probability for parcel
1152 transition from B_{core} to B_{margin} after entrainment.

1153 iv. Parcels that are initially part of B_{margin} can either increase or decrease
1154 their buoyancy value, but never become positively buoyant. The former
1155 case (buoyancy decrease) is expected to be more prevalent since it occurs
1156 for the smaller range of temperature differences with the environment.

1157 In summary, entrainment is expected to always have a net negative effect on B_{core}
1158 extent and B_{margin} values, while evaporation feedbacks serve to maintain RH_{core} in
1159 the cloud. Thus, we can predict that B_{core} should be a subset of RH_{core} (i.e. $B_{core} \subseteq$
1160 RH_{core}).

1161

1162 **3.2.2. Effects of core and margin mixing on buoyancy**

1163 We consider the case of mixing between the B_{core} and B_{margin} , meaning positively
1164 buoyant and negatively buoyant cloud parcels. For simplicity, we assume both parcels
1165 are saturated ($S=1$, both included in the RH_{core}). As seen above, such conditions exist
1166 in both the adiabatic case and in the case where an adiabatic cloud has undergone some
1167 entrainment with the environment. The buoyancy differences between the saturated
1168 parcels are mainly due to temperature differences, but also due to the increasing
1169 saturation vapor pressure with increasing temperature (see Appendix B for details).

1170 In Fig. B1 is it shown that the resultant mixed parcel's buoyancy can be either positive
1171 or negative, depending on the magnitude of temperature difference of each parcel (core
1172 or margin) from that of the environment. However, in all cases the mixed parcel is
1173 supersaturated. This result can be generalized: given two parcels with equal RH but
1174 different temperature, the RH of the mixed parcel is always equal or higher than the
1175 initial value. Hence, B_{core} can either increase or decrease in extent, while the RH_{core}
1176 can only increase due to mixing between saturated B_{core} and B_{margin} parcels. This
1177 again strengthens the assumption that B_{core} should be a subset of RH_{core} .

1178 We note that an alternative option for mixing between the core and margin parcels that
1179 exist here, where either or both of the parcels are subsaturated so that the mixed parcel
1180 is subsaturated as well. In this case evaporation will also occur. As seen in Appendix
1181 A, this should further reduce the buoyancy value of the mixed parcel (while increasing
1182 the RH).

1183

1184 3.2.3. Effects of entrainment on vertical velocity

1185 ~~We divide the entrainment effects on the W_{core} to two: i) a direct effect which includes~~
1186 ~~conservation of momentum of vertical velocity between the core and margin/non-~~
1187 ~~cloudy parcels, and ii) an indirect effect of vertical velocity changes due to buoyancy~~
1188 ~~changes caused by the entrainment. The vertical velocity equation dictates that~~
1189 ~~buoyancy is the main production term (de Roode et al., 2012; Romps and Charn, 2015),~~
1190 ~~and is balanced by perturbation pressure gradients and mixing (on grid and sub-grid~~
1191 ~~scales). Thus, all changes of magnitude (and sign) in vertical velocity should lag the~~
1192 ~~changes in buoyancy. This is the basis of convective overshooting and cumulus~~
1193 ~~formation in the transition layer (see Sect. 3.1). It is interesting to assess the magnitude~~
1194 ~~of this effect by quantifying the expected time lag between buoyancy and vertical~~
1195 ~~velocity changes. The direct effect can be considered to occur instantaneously. (de~~
1196 ~~Roode et al., 2012) Assuming homogeneous mixing of both parcels and a mixing~~
1197 ~~fraction of 0.5, the direct effect can be simplified to conservation of momentum before~~
1198 ~~and after mixing. Since both parcels are approximately of equal mass (in isobaric~~
1199 ~~mixing), the mixed parcel's vertical velocity will be the average of the initial velocities.~~
1200 ~~If the absolute value of the updraft in the W_{core} parcel is larger than that of the~~

1201 ~~downdraft in the margin/non-cloudy parcel, the resultant mixed parcel will remain part~~
1202 ~~of W_{core} . This is usually the case during the growing stages in clouds, where it can be~~
1203 ~~assumed that the surrounding air around W_{core} is at rest or with downdrafts weaker than~~
1204 ~~the updrafts within the W_{core} .~~

1205 ~~As opposed to the direct effect, the indirect effect is time dependent.~~ The calculations
1206 in Appendix A indicates negative buoyancy values reaching -0.1 m/s^2 due to
1207 entrainment. However, measurements from within clouds show that the temperature
1208 deficiency of cloudy parcels with respect to the environment is generally restricted to
1209 less than 1°C for cumulus clouds (Burnet and Brenguier, 2010; Malkus, 1957;
1210 Sinkevich and Lawson, 2005; Wei et al., 1998), and thus the negative buoyancy should
1211 be no more larger than -0.05 m/s^2 . This value is closer to current and previous
1212 simulations and also observations that show negative buoyancy values within clouds to
1213 be confined between -0.001 and -0.01 m/s^2 (de Roode et al., 2012; Ackerman, 1956).
1214 Given an initial vertical velocity of $\sim +0.5 \text{ m/s}$, the deceleration due to buoyancy (and
1215 reversal to negative vertical velocity) should occur within a typical time range of 1 - 10
1216 minutes. These timescales are much longer than the typical timescales of entrainment
1217 (mixing and evaporation that eliminate the B_{core}) which range between 1 – 10 s
1218 (Lehmann et al., 2009). Therefore, ~~even if entrainment acts to reduce the switching of~~
1219 ~~signs of vertical velocity, it does so with should occur with~~ substantial delay compared
1220 to the reduction of buoyancy, and B_{core} should be a subset of W_{core} (i.e. $B_{core} \subseteq$
1221 W_{core}) during the growing and mature stages of a cloud's lifetime.

1222

1223 **3.3. The relation between supersaturation and vertical velocity cores**

1224 ~~Here we revisit the terms in Eq. 3.~~ A rising parcel initially has no liquid water content,
1225 with its only source of supersaturation being the updraft w , and thus initially the RH_{core}
1226 should always be a ~~proper~~ subset of W_{core} . In general, since the sink term $\frac{dLWC}{dt}$
1227 becomes a source only when $S < 1$ (the condition for evaporation), the only way for a
1228 convective cloud to produce supersaturation (i.e. $S > 1$) is by updrafts during all stages
1229 of its lifetime. Once supersaturation is achieved, the sink term becomes positive $\frac{dLWC}{dt} >$
1230 0 and balances the updraft source term, so that supersaturation either increases or
1231 decreases. At any stage, if downdrafts replace the updrafts within a supersaturated

1232 parcel, the consequent change in supersaturation becomes strictly negative (i.e. $\frac{dS}{dt} <$
1233 0). This negative feedback limits the possibility to find supersaturated cloudy parcels
1234 with downdrafts. Hence, we can expect the RH_{core} to be smaller than W_{core} , ~~even~~
1235 though not necessarily a proper subset during the majority of a cloud's lifetime.

1236

1237

1238 4.3. Results - Single cloud simulation

1239 The differences between the three types of core definitions are examined during the
1240 lifetime of a single cloud (Fig. 2), based on the Hawaiian profile. The cloud's total
1241 lifetime is 36 minutes (between $t=7$ and $t=43$ min of simulation). Each panel in Fig. 2
1242 presents vertical cross-sections of the three cores (magenta - W_{core} , green - RH_{core} , and
1243 yellow - B_{core}) at four points in time (with 10-minute intervals). The cloud has an initial
1244 cloud base at 850m, and grows to a maximal top height of 2050 m. The condensation
1245 rates (red shades) increase toward the cloud center and the evaporation rates (blue
1246 shades) increase toward the cloud edges. Evaporation at the cloud top results in a large
1247 eddy below it that contributes to mixing and evaporation at the lateral boundaries of the
1248 cloud. Thus, a positive feedback is initiated which leads to cooling, negative buoyancy,
1249 and downdrafts. The dissipation of the cloud is accompanied with a rising cloud base
1250 and lowering of the cloud top.

1251 During the growing stage ($t=10, 20$ min), when substantial condensation still occurs
1252 within the cloud, all of the cores seem to be self-contained within one another, with
1253 B_{core} being the smallest and W_{core} being the largest. During the final dissipation stages,
1254 when the cloud shows only evaporation ($t=40$), W_{core} and RH_{core} disappear while there
1255 is still a small B_{core} near the cloud top. Further analysis shows that the entire
1256 dissipating cloud is colder and more humid than the environment but downdrafts from
1257 the cloud top (see arrows in Fig. 2) promote adiabatic heating, and by that increase the
1258 buoyancy in dissipating cloudy pixels, sometimes reaching positive values. These
1259 buoyant pockets will be discussed further in Part II. The results indicate that the three
1260 types of physical cores of the cloud are not located around the cloud's geometrical core
1261 along the whole cloud lifetime. During cloud growth (i.e. (increase in mass and size)

1262 the three types of cores surround the cloud's center, while during late dissipation the
1263 B_{core} is at offset from the cloud center.

1264 For a more complete view of the evolution of the three core types in the single cloud
1265 case, time series of core fractions are shown in Fig. 3. Panels a and b show the core
1266 mass (core mass / total mass $-f_{mass}$) and volume (core volume / total volume $-f_{vol}$)
1267 fractions out of the cloud's totals. The results are similar for both measures ~~expect~~
1268 ~~except~~ for the fact that core mass fractions are larger than core volume fractions. This
1269 is due to significantly higher LWC per pixel in the cores compared to the margins,
1270 which skews the core mass fraction to higher values. Core mass fractions during the
1271 main cloud growing stage (between $t=7$ and $t=27$ min simulation time) are around 0.7
1272 - 0.85 and core volume fractions are around 0.5 - 0.7. The time series show that as
1273 opposed to the W_{core} and RH_{core} fractions which decrease monotonically with time,
1274 B_{core} shows a slight increase during stages of cloud growth. In addition, for most of the
1275 cloud's lifetime the B_{core} fractions are the smallest and the W_{core} fractions are the
1276 largest, except for the final stage of the clouds dissipation where downdrafts from the
1277 cloud top creates pockets of positive buoyancy. These pockets are located at the cloud's
1278 peripheral regions rather than near the cloud's geometrical center as is typically
1279 expected for the cloud's core. In the cloud's center (the geometrical core) the B_{core} is
1280 the first one to terminate (at $t=32$ min) compared to both W_{core} and RH_{core} that decay
1281 together (at 36 min).

1282 For describing the locations of the physical cores, we examine the normalized distances
1283 (D_{norm}) between the cloud's centroid and the cores' centroids. The evolution of these
1284 distances is shown in Fig. 3c. At cloud initiation ($t=7$ min), when the cloud is very
1285 small, all cores' centroids coincide with the total cloud centroid location. The B_{core}
1286 (and RH_{core} to a much lesser degree) centroid then deviates from the cloud centroid to
1287 a normalized distance of 0.27 ($t=8$ min). As cloud growth proceeds, B_{core} grows and
1288 its centroid coincides with the cloud's centroid. All cores' centroids are located near the
1289 cloud centroid during the majority of the growing and mature stages of the cloud,
1290 showing normalized distances <0.1 . During dissipation ($t>27$ min), the cores' centroid
1291 locations start to distance away from the cloud's geometrical core followed by a
1292 reduction in distances due to the rapid loss of cloud volume. As mentioned above, it is

1293 shown that the regeneration of positive buoyancy at the end of cloud dissipation (t=40
1294 min) takes place at the cloud edges, with normalized distance >0.5.

1295 Finally, in Fig. 3d the fraction of pixels of each core contained within another core is
1296 shown. It can be seen that for the majority of cloud lifetime (up to t=33 min) B_{core} is
1297 subset (pixel fraction of 1) of RH_{core} , and the latter is a subset of W_{core} . As expected,
1298 the other three permutations of pixel fractions (e.g. W_{core} in B_{core}) show much lower
1299 values. The cloudy regions that are not included within B_{core} but are included within
1300 the two other cores are exclusively at the cloud's boundaries (see Fig. 2). The same
1301 pattern is seen for cloudy regions that are included within W_{core} but not in RH_{core} .
1302 During the dissipation stage of the cloud its self-containing property (i.e. $B_{core} \subseteq$
1303 $RH_{core} \subseteq W_{core}$) breaks down. Similar temporal evolutions as shown here are seen for
1304 the other simulated clouds (with various aerosol concentrations) in part II of this work.

1305

1306 **5. Results - Cloud field simulations**

1307 **5.1. Partition to different core types**

1308 To test the robustness of the observed behaviors seen for a single cloud ~~(and explained~~
1309 ~~in the theoretical part)~~, it is necessary to check whether they also apply to large statistics
1310 of clouds in a cloud field. The BOMEX simulation is taken for the analyses here. We
1311 discard the first 3 hours of cloud field data, during which the field spins-up and its mean
1312 properties are unstable. In Fig. 4 the volume (f_{vol}) and mass (f_{mass}) fractions of the
1313 three core types are compared for all clouds (at all output times – every 1 min) in the
1314 CvM space. As seen in Fig. 1, the location of specific clouds in the CvM space indicates
1315 their stage in evolution. Most clouds are confined to the region between the adiabat and
1316 the inversion layer base except for small precipitating (lower left region) and dissipating
1317 clouds (upper left region). The color shades of the clouds indicate whether a cloud is
1318 mostly core (red), mostly margin (blue), or equally divided to core and margin (white).

1319 As seen for the single cloud, the core mass fractions tend to be larger than core volume
1320 fractions, for all core types. This is due to the fact that LWC values in the cloud core
1321 regions are higher than in margin regions, so that a cloud might be core dominated in
1322 terms of mass while being margin dominated in terms of volume. Focusing on the

1323 differences between core types, the color patterns in the CvM space imply that B_{core}
1324 definition yields the lowest core fractions (for both mass and volume), followed by
1325 RH_{core} with higher values and W_{core} with the highest values. The absence of the B_{core}
1326 is especially noticeable for small clouds in their initial growth stages after formation
1327 (COG \sim 550 m and LWP $<$ 1 g m⁻²). Those same clouds show the highest core fractions
1328 for the other two core definitions. This large difference can be explained by the
1329 existence of the transition layer ([as discussed in Sect. 3](#)) near the lifting condensation
1330 level (LCL) in warm convective cloud fields which is the approximated height of a
1331 convective cloud base (Craven et al., 2002; Meerkötter and Bugliaro, 2009). Within
1332 this layer parcels rising from the sub-cloudy layer are generally colder than parcels
1333 subsiding from the cloudy layer. Thus, this transition layer clearly marks the lower edge
1334 of the buoyancy core as most convective clouds are initially negatively buoyant.

1335 Generally, the growing cloud branch (i.e. the CvM region closest to the adiabat) shows
1336 the highest core fractions. The RH_{core} and W_{core} fractions decrease with cloud growth
1337 (increase in mass and COG height) while the B_{core} initially increases, shows the highest
1338 fraction values around the middle region of the growing branch and then decreases for
1339 the largest clouds. The transition from the growing branch to the dissipation branch is
1340 manifested by a transition from core dominated to margin dominated clouds (i.e.
1341 transition from red to blue shades). Mixed within the margin dominated dissipating
1342 cloud branch, a scatter of W_{core} dominated small clouds can be seen as well. These
1343 represent cloud fragments which shed off large clouds during their growing stages with
1344 positive vertical velocity. They are sometimes RH_{core} dominated as well but are strictly
1345 negatively buoyant. The few precipitating cloud fragments seen for this simulation
1346 (cloud scatter located below the adiabat) tend to be margin dominated, especially for the
1347 RH_{core} .

1348

1349 **5.2. ~~Self-contained~~[Subset](#) properties of cores**

1350 From Fig. 4 it is clear that W_{core} tends to be the largest and B_{core} tends to be the
1351 smallest. To what degree however, are the cores [self-contained within subsets of](#) one
1352 another as was seen for the single cloud simulation? It is also interesting to check
1353 whether the different physical cores are centered near the cloud's geometrical core. In

1354 Fig. 5 the pixel fraction (f_{pixel}) of each core type within another core type is shown for
 1355 all clouds in the CvM space. A f_{pixel} pixel fraction of 1 (bright colors) indicates that the
 1356 pixels of the specific core in question (labeled in each panel title) ~~completely are a~~
 1357 ~~subset overlap with the pixels~~ of the other core (also labeled in the panel title) and a
 1358 f_{pixel} pixel fraction of 0 (dark colors) indicates ~~zero overlap no intersection~~ between the
 1359 two cores in the cloud. It is seen that B_{core} tends to be a subset of both other cores, with
 1360 f_{pixel} pixel fractions around 0.75-1 for most of the growing branch area and large mass
 1361 dissipating clouds which still have some positive buoyancy. The pixel fractions are
 1362 higher for B_{core} inside W_{core} compared with B_{core} inside RH_{core} , but both show
 1363 decrease with increase in growing branch cloud mass, meaning that chance for perfect
 1364 self-containing of the cores decreases in large clouds.

1365 The CvM space of RH_{core} inside W_{core} shows an even stronger relation between these
 1366 two core types. For almost all growing branch clouds, the RH_{core} is a subset of W_{core}
 1367 (i.e. $RH_{core} \subseteq W_{core}$). The decrease gradually with loss of cloud mass in the dissipation
 1368 branch. The other three permutations of f_{pixel} (W_{core} inside B_{core} , W_{core}
 1369 inside RH_{core} , and RH_{core} inside B_{core}) give an indication of cores sizes and of which
 1370 cloud types show no overlap between different cores. As stated above, growing
 1371 (dissipation) clouds show higher (lower) overlap between the different core types. The
 1372 W_{core} is almost twice as large as the B_{core} and 30%-40% larger than the RH_{core} along
 1373 most of the growing branch. In conclusion, we see a strong tendency for the ~~self-~~
 1374 ~~containingsubset~~ property of cores ($B_{core} \subseteq RH_{core} \subseteq W_{core}$) during the growth stages
 1375 of clouds. This property ceases for dissipating and precipitating clouds, especially for
 1376 the smaller clouds which show less overlap between core types.

1377 In Fig. 6 the normalized distances (D_{norm}) between the total cloud centroid and each
 1378 specific physical core centroid locations are evaluated. Along the growing branch the
 1379 cloud centroid and physical cores' centroids tend to be of close proximity, while during
 1380 cloud dissipation the cores' centroids tend to increase in distance from the cloud's
 1381 center. This type of evolution is most prominent for the W_{core} , which shows a clear
 1382 gradient of transition from small (dark colors) to large (bright colors) distances. The
 1383 B_{core} shows a more complex transition, from intermediate distance values (~ 0.5) at
 1384 cloud formation, to near zeros values along the mature part of the growing branch, back
 1385 to large values in the dissipation branch. Along the growing branch RH_{core} shows

1386 distances comparable to the W_{core} (except for large distances at cloud formation).
1387 However, compared to the other two core types, RH_{core} shows the smallest distances
1388 to the geometrical core during cloud dissipation. This is manifested by a relative
1389 absence of bright colors for dissipating clouds in Fig. 6.

1390 The prevalence of cloud edge B_{core} pixels during dissipation can be explained by
1391 adiabatic heating due to weak downdrafts (see Sect. 4.2, Part II) which are expected at
1392 the cloud periphery. The fact that there is little overlap between B_{core} and both W_{core}
1393 and RH_{core} pixels in dissipating clouds (see Fig. 5) serves to verify this assumption.
1394 The relative absence of isolated RH_{core} pixels at the cloud edges can be explained by
1395 the fact the pixels closest to the cloud's edge are most susceptible to mixing with non-
1396 cloudy air and evaporation, yielding subsaturation conditions. The innermost pixels are
1397 "protected" from such mixing and thus we can expect most RH_{core} pixels to be located
1398 near the geometrical core.

1399 The W_{core} case is less intuitive. During cloud dissipation complex patterns of updrafts
1400 and downdrafts within the cloud can create scenarios where the W_{core} centroid is
1401 located anywhere in the cloud. However, the results show that most small dissipating
1402 clouds tend to have their W_{core} pixels concentrated at the cloud edges. Comparing Fig.
1403 6 with Figs. 4 and 5, we can see that these pixels comprise only a tiny fraction of the
1404 already small clouds and do not overlap with RH_{core} and B_{core} pixels and thus are not
1405 related to significant convection processes. Further analysis shows that the maximum
1406 updrafts in these clouds rarely exceed 0.5 m/s (i.e. 90% of clouds with normalized
1407 distance > 0.9 have a maximum updraft of less than 0.5 m/s), and can thus be considered
1408 with near neutral vertical velocity.

1409

1410 **5.3. Consistency of the cloud partition to core types**

1411 The results for cloud fields are summarized in Fig. 7 that presents the evolution of core
1412 fractions of continuous cloud entities (CCEs, see Sect. 2.5 for details) from formation
1413 to dissipation. Only CCEs that undergo a complete life cycle are averaged here. These
1414 CCEs fulfill the following four conditions: i) form near the LCL, ii) live for at least 10
1415 minutes, ii) reach maximum cloud mean LWP values above 10 g m^{-2} , and iv) terminate
1416 with mass value below 10 g m^{-2} . As a test of generality, we performed this analysis for

1417 Hawaiian and Amazonian warm cumulus cloud field simulations in addition to the
 1418 BOMEX one. For each simulation, tens to hundreds of CCEs are collected (see panel
 1419 titles) and their core fractions are averaged according to their normalized lifetimes (τ).
 1420 Consistent results are seen for all three simulations. Clouds initiate with a W_{core}
 1421 fraction of ~ 1 , RH_{core} fraction of ~ 0.8 , and B_{core} fraction of ~ 0.1 . The former two
 1422 core types' volume fraction decreases monotonically with lifetime, while the latter core
 1423 type's volume fraction increases up to 0.3 at $\tau \sim 0.25$, and then monotonically decreases
 1424 for increasing τ . The fact that clouds end their life cycle with non-zero volume fractions
 1425 may indicate that some of the CCE terminate not because of full dissipation but rather
 1426 because of significant splitting or merging events.

1427 Normalized distances between core centroid and total cloud centroid (Fig. 7, middle
 1428 column) tend to monotonically increase for RH_{core} and W_{core} with CCE lifetime for all
 1429 simulations. The gradient of increase is larger at the later stages of CCE lifetime.
 1430 Initially the W_{core} is closer to the geometrical core but at later stages of CCE lifetime
 1431 (typically $\tau > 0.5$) this switches and RH_{core} remains the closest. As seen above, for the
 1432 first (second) half of CCE lifetime, the distance between B_{core} centroid and cloud
 1433 centroid decreases (increases), starting at normalized distances above 0.4 for all
 1434 simulations. The physical cores stay in proximity to the geometrical core for the
 1435 majority of their lifetimes for the three cases. Taking the value 0.5 as a threshold for
 1436 transition from centered physical cores to periphery physical cores, Bomex, Hawaii,
 1437 and Amazon simulation CCEs' W_{core} cross this threshold at $\tau = 0.94, 0.9$, and 0.86 ,
 1438 respectively. Thus, the assumption that a cloud's core (by any definition) is also
 1439 indicative of the cloud's centroid is true for the majority of a typical cloud's lifetime.

1440 The analysis of self-containing core properties (Fig. 7, right column) shows that the
 1441 assumption $B_{core} \subseteq RH_{core} \subseteq W_{core}$ is true for the initial formation stages of a cloud.
 1442 Although the corresponding pixel fractions decrease slightly during the lifetime of the
 1443 CCE, they remain above 0.9 (e.g. B_{core} is 90% contained within RH_{core}). A sharp
 1444 decrease in pixel fractions is seen for $\tau > 0.8$, as the overlaps between the different cores
 1445 is reduced during dissipation stages of the cloud. For all simulations, the highest pixel
 1446 fraction values are seen for the B_{core} inside W_{core} pair, followed by RH_{core}
 1447 inside W_{core} pair, and B_{core} inside RH_{core} pair showing slightly lower values. In
 1448 addition, it can be seen that the variance of average pixel fraction (per τ) increases with

1449 increase in τ . This is due to the fact the all CCEs initiate with almost identical
1450 characteristics but may terminate in very different ways. In part II of this work we show
1451 that this variance is highly influenced from precipitation which contributes to more
1452 significant interactions between clouds (Heiblum et al., 2016b). Indeed, the Amazon
1453 simulation shows the largest pixel fraction variance and produces the most precipitation
1454 out of the three simulations.

1455

1456 **6. Summary**

1457 In this paper we study the partition of warm convective clouds to core and margin
1458 according to three different definitions: i) positive vertical velocity (W_{core}), ii) relative
1459 humidity supersaturation (RH_{core}), and iii) positive buoyancy (B_{core}), with emphasis
1460 on the differences between those definitions. Using theoretical considerations of both
1461 an adiabatic cloud [column](#) and a simple two parcel mixing model (see appendix A and
1462 B), we support our simulated results as we show that the B_{core} ~~must~~ [is expected to](#) be
1463 the smallest of the three. [This finding is in line with previous works that showed that](#)
1464 [negative buoyancy is prevalent in cumulus clouds for a wide range of thermodynamic](#)
1465 [conditions](#) (de Roode, 2008; Paluch, 1979; Taylor and Baker, 1991). This is due to the
1466 fact that entrainment into the core (i.e. mixing with non-cloudy environment or mixing
1467 with the margin regions of the cloud) ~~acts instantaneously to reduce cloud buoyancy~~
1468 ~~values. In cases the mixed parcel is~~ [may result in sub-saturationed](#), [followed by](#)
1469 evaporation ~~that occurs and~~ always has a negative [net](#) effect on buoyancy. The same
1470 process has an opposing effect on the relative humidity of the mixed parcel and acts to
1471 reach saturation. Entrainment (or mixing) also acts to decrease vertical velocity, but at
1472 slower manner compared to the time scales of changes in the buoyancy and relative
1473 humidity. In addition, the supersaturation equation (Eq. (3)) predicts that it is unlikely
1474 to ~~attain~~ [maintain](#) supersaturation in a cloudy volume with negative vertical velocity.
1475 Hence, W_{core} ~~can be~~ [is](#) expected to be the largest of the three cores.

1476 Using numerical simulations of both a single cloud and cloud fields of warm cumulus
1477 clouds, we show that during most stages of clouds' lifetime, W_{core} is indeed the largest
1478 of the three and B_{core} the smallest. In addition to the differences in their sizes, the three
1479 cores tend to be subsets of one another (and located around the cloud geometrical

1480 center), in the following order: $B_{core} \subseteq RH_{core} \subseteq W_{core}$. This property is most valid
 1481 for a cloud at its initial stages and breaks down gradually during a cloud's lifetime. ~~The~~
 1482 ~~small B_{core} fractions (out of the total cloud) are due to two main reasons: i) buoyancy~~
 1483 ~~is strongly affected by mixing and evaporation, as the buoyant core is the first to~~
 1484 ~~disappear during the dissipation stages of a cloud, and ii) warm~~ The warm convective
 1485 cloud fields simulated here typically have a transition layer near the lifting condensation
 1486 level (LCL), ~~where~~ Thus, ascending parcels are colder than descending parcels so the
 1487 lower parts of the clouds are negatively buoyant or even lack a B_{core} at formation. After
 1488 cloud formation internal growth processes (i.e. condensation and latent heat release)
 1489 increase the B_{core} until dissipation processes become dominant and the ~~B_{core}~~
 1490 decreases quickly due to entrainment. In contrast, clouds are initially dominated by the
 1491 W_{core} and RH_{core} (fractions close to 1). The fractions of these cores then decrease
 1492 monotonically with cloud lifetime.

1493 During dissipation stages, the clouds are mostly margin dominated, such that most of
 1494 the small mass dissipation cloud fragments are entirely coreless. However, several
 1495 small mass dissipating cloud fragments which shed off large cloud entities (with large
 1496 COG height) may be core dominated, especially using the RH_{core} definition. The same
 1497 is observed for small precipitating cloud fragments which reside below the convective
 1498 cloud base. We note that the results here are similar for both volume and mass core
 1499 fractions out the cloud's totals, with the core mass fractions being larger due to a skewed
 1500 distribution of cloud LWC which favors the core regions. Moreover, we show that these
 1501 results are consistent for various levels of aerosol concentrations (will be seen in Part
 1502 II) and different thermodynamic profiles used to initialize the models.

1503 With respect to cloud morphology, it is shown that during cloud growth, which
 1504 comprises the majority of a warm cloud lifetime, the physical cores are centered near
 1505 the cloud's geometrical core, as is intuitively expected from a cloud's core. This
 1506 matches the convective cloud core-shell model. An exception to this is the initial growth
 1507 stages, where the B_{core} centroid can be located far from the cloud's centroid. During
 1508 dissipation, the core-shell model no longer applies to the clouds, as the cores decouple
 1509 from the geometrical core and often comprise just a few isolated pixels at the cloud's
 1510 edges. The W_{core} and B_{core} pixels tend to be more peripheral than RH_{core} during
 1511 dissipation (see Sect. 5.2). Downdraft induced adiabatic heating at the clouds' edge (see

1512 more in Part II) promote positive buoyancy while decreasing the chance for
 1513 supersaturation. During dissipation the overlap between different core types also
 1514 decreases rapidly, implying that minor local effects enable core existence rather than
 1515 cloud convection. Thus, only during mature growth stages can all three cores types can
 1516 be considered interchangeable. In Part II of this work we use the insights gained here
 1517 to understand aerosol effects on warm convective clouds, as are reflected by a cloud's
 1518 partition to its core and margin.

1519 **Acknowledgements**

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 1521 Technology, Israel (grant no. 3-14444).

1522

1523

1524 **Appendix A: Buoyancy changes due to mixing of cloudy and non-cloudy parcels**

1525 Here we present a simple model for entrainment mixing between a cloudy parcel (either
 1526 part of B_{core} or B_{margin}) and a dry environmental parcel. Entrainment mixes the
 1527 momentum, heat, and humidity of the two parcels. We consider the mixing of a unit
 1528 mass of cloud parcel which is defined by two criteria:

1529
$$S_1 \geq 1$$

$$B_1 > 0 \text{ or } B_1 < 0$$

1530 with a unit mass of dry environment parcel, defined by:

1531
$$S_2 < 1$$

1532 and explore the properties of the resulting mixed parcel.

1533 Assume that T_1, T_2, T_3 are the initial temperatures of the cloudy, environmental, and
 1534 resulting mixed parcel, respectively. $q_{v1}, q_{v2}, q_{v3}, \theta_1, \theta_2, \theta_3$, and q_{l1}, q_{l2}, q_{l3} are their
 1535 respective vapor mixing ratios, potential temperatures, and liquid water contents
 1536 (LWC).

1537 The change in buoyancy due to mixing will be:

1538
$$dB_{mix} = g * \left(\frac{\theta_3 - \theta_1}{\theta_2} + 0.61(q_{v3} - q_{v1}) - (q_{l3} - q_{l1}) \right) \quad (A1),$$

1539 with

$$1540 \quad T_3 = \mu_1 \cdot T_1 + \mu_2 \cdot T_2 \quad (\text{A2}),$$

$$1541 \quad q_{v3} = \mu_1 \cdot q_{v1} + \mu_2 \cdot q_{v2} \quad (\text{A3}),$$

$$1542 \quad q_{l3} = \mu_1 \cdot q_{l1} + \mu_2 \cdot q_{l2} \quad (\text{A4}),$$

1543 where μ_1 and μ_2 are the corresponding mixing fractions. We assume that the mixed
1544 parcel is at the same height as the cloudy and environmental parcels, and that the mean
1545 environmental temperature at that height stays the same after mixing. The potential
1546 temperature (θ) is calculated using its definition.

1547 After the mixing process, the resultant mixed parcel may be subsaturated ($S_3 < 1$), and
1548 cloud droplets start to evaporate. The evaporation process increases the humidity of the
1549 parcel. ((Korolev et al., 2016), Eq. (A8)) calculated the amount of the required liquid
1550 water for evaporation, in order to reach $S=1$ again:

$$1551 \quad \delta q = \frac{C_p R_v T_2^2}{L^2} \ln \left(\frac{1 + \frac{e_s(T_3) R_a L^2}{P C_p R_v^2 T_3^2}}{1 + S_3 \frac{e_s(T_3) R_a L^2}{P C_p R_v^2 T_3^2}} \right) \quad (\text{A5}),$$

1552 Where C_p is a specific heat at constant pressure, $e_s(T_3)$ is the saturated vapor pressure
1553 for the mixed temperature, P is pressure, L is latent heat, R_v, R_a are individual gas
1554 constants for water vapor and dry air, respectively. If the mixed parcel contains
1555 sufficient LWC to evaporate δq amount of water, the mixed parcel will reach
1556 saturation. We note that Eq. (A5) holds for cases where $|T_1 - T_2| < 10^\circ C$, which is
1557 well within the range seen in our simulations of warm clouds.

1558 Assuming the average environmental temperature stays the same after evaporation, the
1559 buoyancy after evaporation is calculated using the following formulas:

$$1560 \quad dB_{evap} = g \cdot \left(\frac{d\theta'_{evap}}{\theta_2} + 0.61 dq_{v_{evap}} - dq_{l_{evap}} \right) \quad (\text{A6}),$$

$$1561 \quad d\theta'_{evap} = dT_{evap} \quad (\text{A7}),$$

1562 From the first law of thermodynamics:

$$1563 \quad C_p \cdot dT_{evap} = -L \cdot dq_{v_{evap}} \quad (\text{A8}).$$

1564 The water vapor is the amount of liquid water lost by evaporation:

$$1565 \quad dq_{v_{evap}} = -dq_{l_{evap}} = \delta q \quad (A9),$$

1566 From the above we get:

$$1567 \quad dB_{evap} = g \cdot \delta q \left(1.61 - \frac{L}{c_p \theta_2} \right) \quad (A10).$$

1568 For a wide temperature range between $200 < \theta_2 < 300[K]$, dB_{evap} is always
 1569 negative. This result is not trivial because evaporation both decreases the T and
 1570 increases the q_v which have opposite effects. The total change in buoyancy is taken as
 1571 the sum of dB_{evap} and dB_{mix} .

1572 Figure A1 presents a phase space of possible changes in cloudy pixel buoyancy due to
 1573 mixing with outside air, for various thermodynamic conditions, and a mixing fraction
 1574 of 0.5. The initial cloudy parcel is chosen to be saturated ($S=1$) and includes a LWC of
 1575 1 g kg^{-1} . The pressure is assumed to be 850 mb, and the temperature 15°C . However,
 1576 we note that the conclusions here apply to all atmospherically relevant values of
 1577 pressure, temperature, supersaturation (values of $RH > 100\%$), and LWC in warm
 1578 clouds. The X-axis in Fig. A1 spans a range of non-cloudy environment relative
 1579 humidity values ($60\% < RH < 100\%$), and the Y-axis spans a temperature difference
 1580 range between the cloud and the environment parcels ($-3^\circ < dT < 3^\circ$). The initial (B_i)
 1581 and final (B_f , after entrainment) buoyancy values, and the differences between them
 1582 can be either positive or negative. The regions of $B_i > 0$ ($B_i < 0$) in fact illustrate the effects
 1583 of entrainment on B_{core} (B_{margin}) parcels.

1584 **Appendix B: Buoyancy changes due to mixing of core and margin parcels**

1585 Following the notations of appendix A, we now consider the mixing of two cloudy
 1586 parcels, one part of B_{core} and one part of B_{margin} . For simplicity, we choose the case
 1587 where both parcels are saturated and have the same LWC of 0.5 g kg^{-1} :

$$1588 \quad \begin{aligned} S_{core} &= S_{margin} = S_{cloud} = 1 \\ q_{l_{core}} &= q_{l_{margin}} = q_{l_{cloud}} = 0.5 \end{aligned} \quad (B1).$$

1589 The buoyancy of each cloudy parcel is determined in reference to the environmental
 1590 temperature and humidity, $T_{env}, q_{v_{env}}$, so that:

1591 $B_{cloud} = g * \left(\frac{\theta_{cloud} - \theta_{env}}{\theta_{env}} + 0.61(q_{v_{cloud}} - q_{v_{env}}) - q_{l_{cloud}} \right)$ (B2).

1592 As mentioned in the main text, we take a temperature range of $T_{env} - 3^\circ C < T_{cloud} <$
 1593 $T_{env} + 3^\circ C$. Each cloudy parcel's temperature also dictates its saturation vapor pressure
 1594 $e_s(T_{cloud})$ and therefore also its humidity content, $q_{v_{cloud}}$. Plugging these into Eq. (B2),
 1595 one can associate each temperature/humidity pair with the B_{core} or B_{margin} :

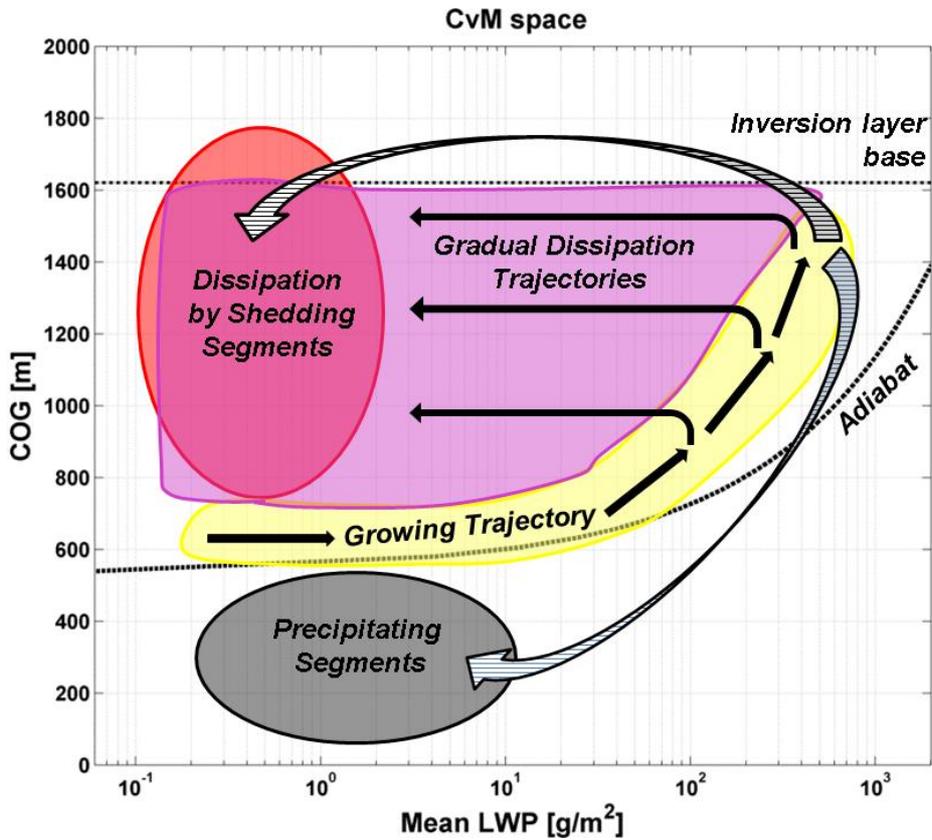
1596
$$\begin{aligned} T_{core} &= T_{cloud}(B_{cloud} > 0), q_{v_{core}} = q_{v_{cloud}}(B_{cloud} > 0) \\ T_{margin} &= T_{cloud}(B_{cloud} < 0), q_{v_{margin}} = q_{v_{cloud}}(B_{cloud} < 0) \end{aligned}$$
 (B3).

1597 The core and margin parcels can then be mixed (see appendix A) yielding a mixed
 1598 parcel temperature and humidity content, and thus a new relative humidity. The
 1599 buoyancy of the mixed parcel is obtained by inserting these parameters in Eq. (B2).

1600 In Fig. B1 the resultant buoyancy values and RH values after the mixing of B_{core}
 1601 parcels with B_{margin} parcels are shown. As defined in Appendix A, temperature
 1602 differences between the parcels and the environment are confined to $\pm 3^\circ C$. The
 1603 reference environmental temperature, pressure, and RH are taken to be $15^\circ C$, 850 mb,
 1604 and 90%, respectively. We note the main differences between this section and
 1605 Appendix A are the absence of evaporation and the fact that the core and margin
 1606 thermodynamic variables are the ones that vary while the reference environmental ones
 1607 are kept constant.

1608 It can be seen that all negatively buoyant parcels are colder than the environment and
 1609 nearly all positively buoyant parcels are warmer than the environment, except for a
 1610 small fraction that are slightly colder but positively buoyant due to the increased
 1611 humidity. The transition from $B_f > 0$ to $B_f < 0$ near the 1 to 1 line indicates that B_f is
 1612 approximately linearly dependent on the temperature differences with respect to the
 1613 environment. In other words, if $|T_{core} - T_{env}| > |T_{margin} - T_{env}|$, the mixed parcel is
 1614 expected to be part of the B_{core} (i.e. $B_f > 0$). The exponential increase in saturation vapor
 1615 pressure with temperature is demonstrated by the results of the mixed parcel final RH,
 1616 which all show supersaturation values. Additional sensitivity tests were performed for
 1617 this analysis, showing only weak dependencies on environmental parameter values,
 1618 while maintaining the main conclusions.

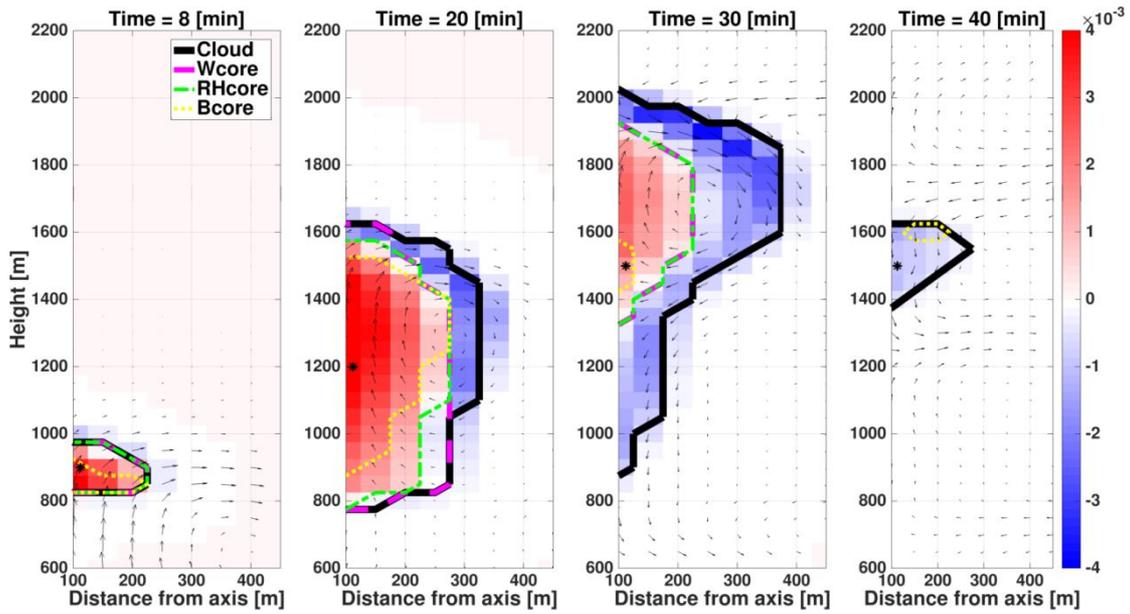
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1622 *Figure 1. A schematic representation of a cloud field Center-of-gravity height (Y-Axis)*
 1623 *vs. Mass (X-Axis) phase space (CvM in short). The majority of clouds are confined to*
 1624 *the region between the adiabatic approximation (curved dashed line) and the inversion*
 1625 *layer base height (horizontal dashed line). The yellow, magenta, red, and grey shaded*
 1626 *regions represent cloud growth, gradual dissipation, cloud fragments which shed off*
 1627 *large clouds, and cloud fragments which shed off precipitating clouds, respectively.*
 1628 *The black arrows represent continuous trajectories of cloud growth and dissipation.*
 1629 *The hatched arrows represent two possible discontinuous trajectories of cloud*
 1630 *dissipation where clouds shed segments.*

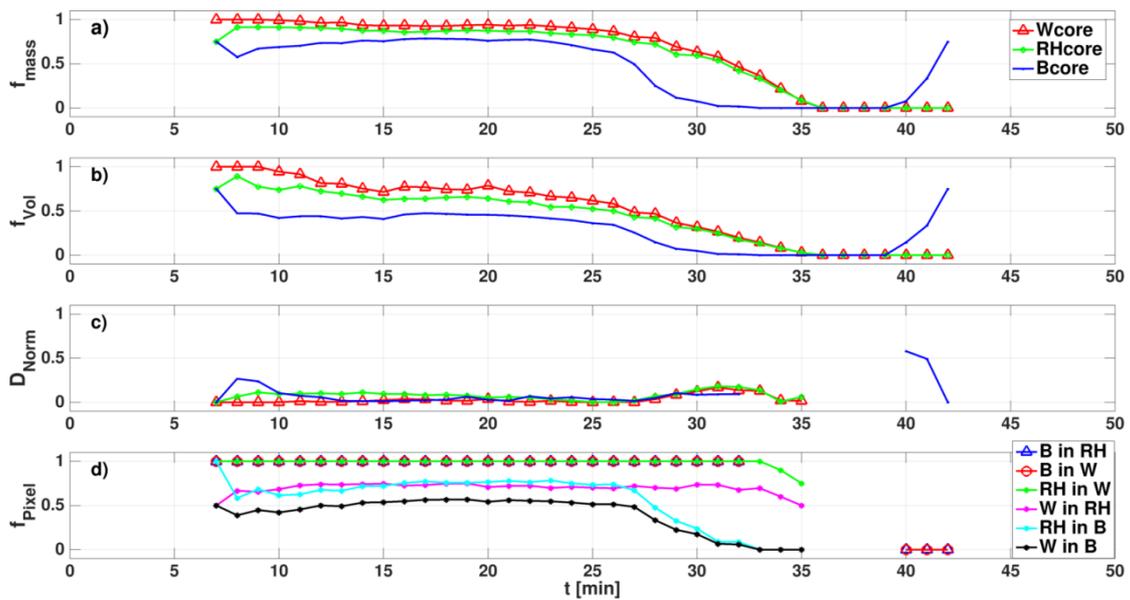
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1633 Figure 2. Four vertical cross-sections (at $t=8, 20, 30, 40$ minutes) during the single
 1634 cloud simulation. Y-axis represents height [m] and X-axis represents the distance from
 1635 the axis [m]. The black, magenta, green and yellow lines represent the cloud,
 1636 W_{core} , RH_{core} and B_{core} , respectively. The black arrows represent the wind, the
 1637 background represents the condensation (red) and evaporation rate (blue) [$g\ kg^{-1}\ s^{-1}$],
 1638 and the black asterisks indicate the vertical location of the cloud centroid. Note that in
 1639 some cases the lines indicating core boundaries overlap (mainly seen for RH and W
 1640 cores).

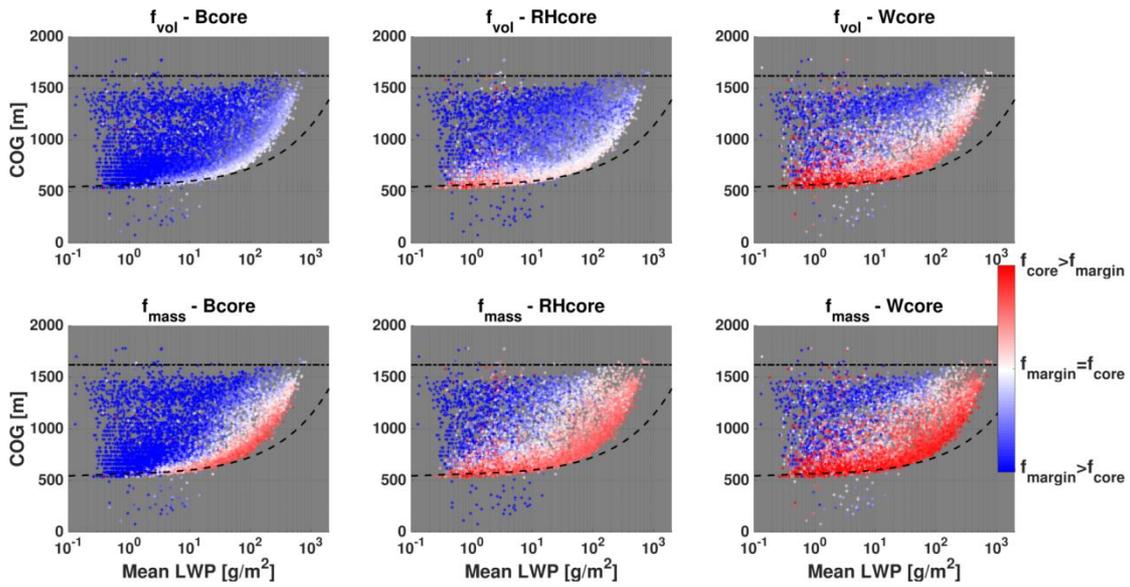
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1643 Figure 3. Temporal evolution of selected core properties, including: (a) The fraction of
 1644 the cores' mass from the total cloud mass (f_{mass}), (b) the fraction of the cores' volume
 1645 from the total cloud volume (f_{vol}), (c) the normalized distance between cloud centroid
 1646 and core centroid (D_{norm}), and (d) the fraction of cores' pixels contained within another
 1647 core (f_{pixel}), including all six permutations. See panel legends for descriptions of line
 1648 colors.

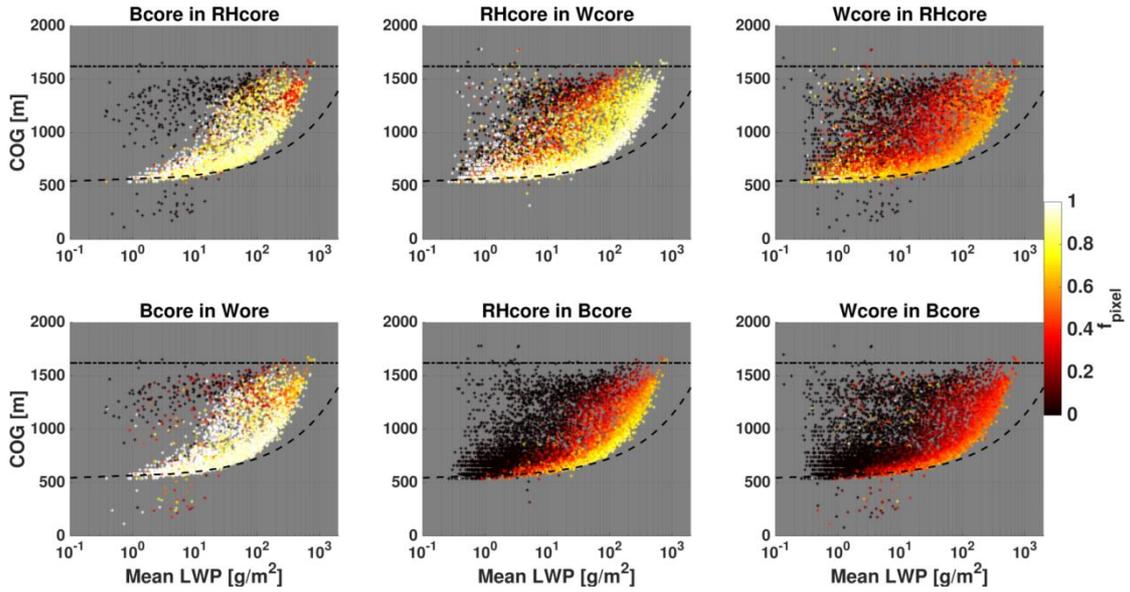
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1651 Figure 4. CvM phase space diagrams of B_{core} (left column), RH_{core} (middle column),
 1652 and W_{core} (right column) fractions for all clouds between 3 h and 8 h in the BOMEX
 1653 simulation. Both volume fractions (f_{vol} , upper panels) and mass fractions (f_{mass} , lower
 1654 panels) are shown. The red (blue) colors indicate a core fraction above (below) 0.5.
 1655 For a general description of CvM space characteristics the reader is referred to Sect.
 1656 2.4.

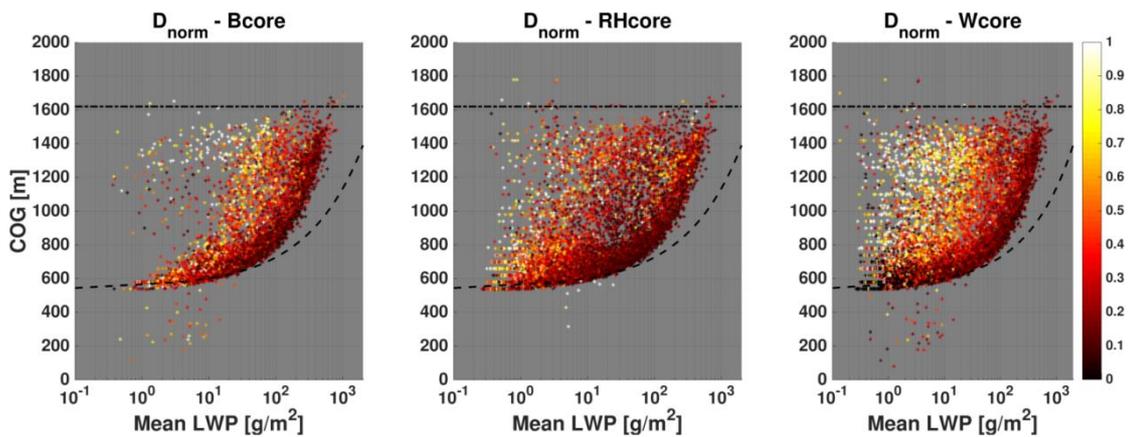
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1659 Figure 5. CvM phase space diagrams of pixel fractions (f_{pixel}) of each of the three
 1660 cores within another core, including six different permutations (as indicated in the
 1661 panel titles). Bright colors indicate high pixel fractions (large overlap between two core
 1662 types) while dark colors indicate low pixel fraction (little overlap between two core
 1663 types). The differences in the scatter density and location for different panels are due
 1664 to the fact that only clouds which contain a core fraction above zero (for the core in
 1665 question) are considered. For example, for the Buoy in RH panel (upper left), only
 1666 cloud that contain some pixels with positive buoyancy are considered.

1667

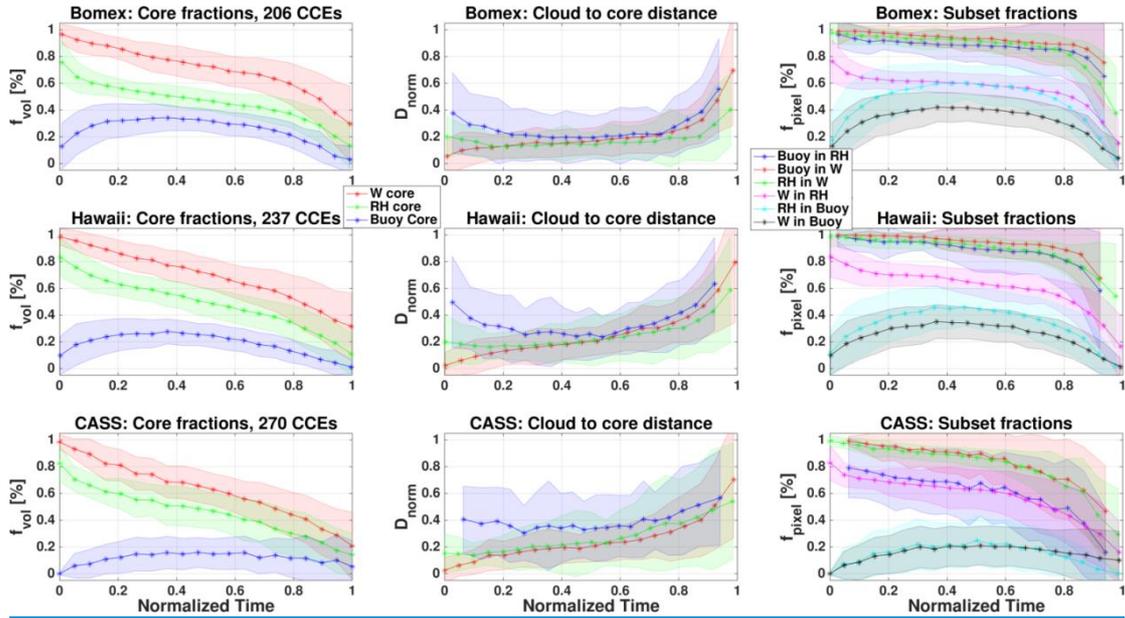


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1669 Figure 6. CvM phase space diagrams of distances between core centroid location and
 1670 cloud centroid location, for the three different physical core types. The distances are
 1671 normalized by the cloud volume radius (approximately the largest distance possible).

1672 Bright (dark) colors indicates large (small) distances. As seen in Fig. 5, only clouds
 1673 which contain a core fraction above zero (for the core in question) are considered.

1674



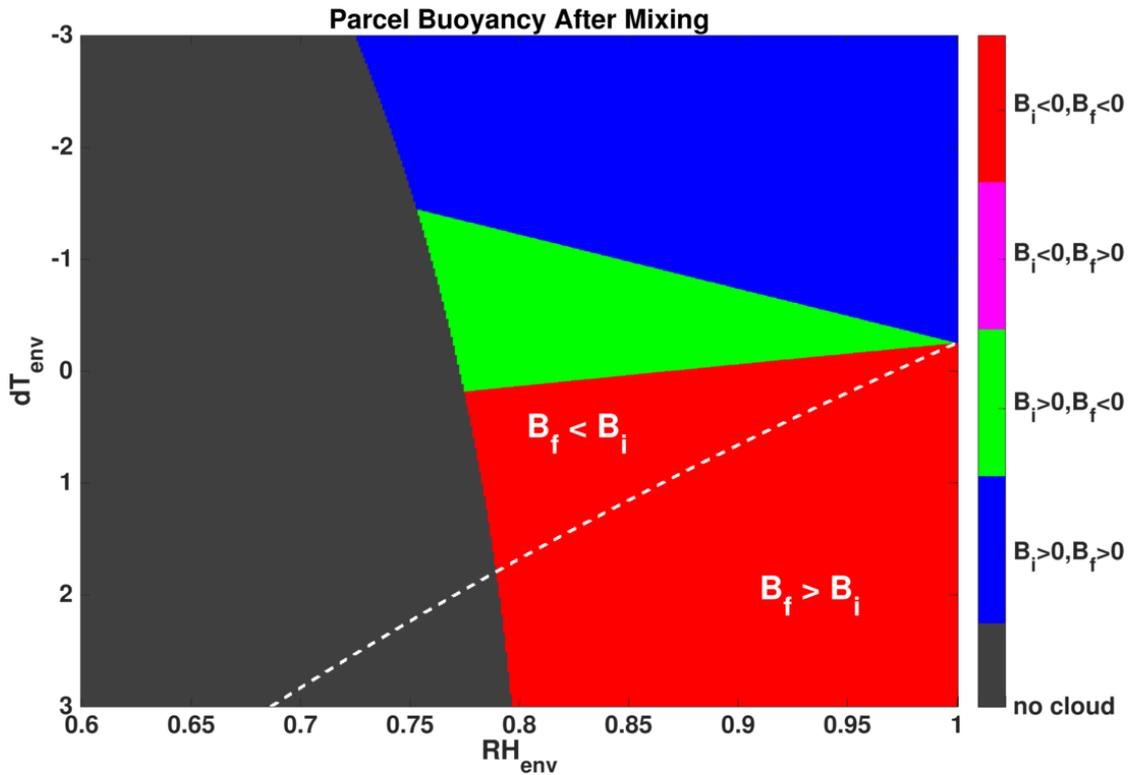
1675

1676

1677 Figure 7. Normalized time (τ) series of CCE averaged core fractions for the BOMEX
 1678 (upper row), Hawaii (middle row), and Amazon (bottom row) simulations. Both core
 1679 volume fractions (left column), normalized distances between cloud and core centroid
 1680 locations (middle column), and pixel fractions of one core within another (right
 1681 column) are considered. Line colors indicated different core types (see legends), while
 1682 corresponding shaded color regions indicate the standard deviation. Normalized time
 1683 enables to average together CCEs with different lifetimes, from formation to
 1684 dissipation. The number of CCEs averaged together for each simulation is included in
 1685 the left column panel titles.

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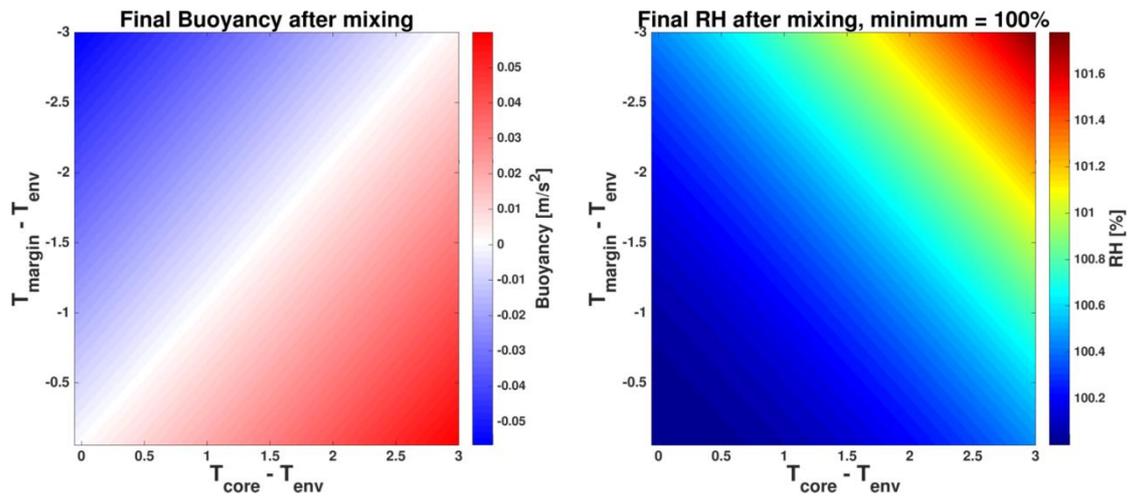
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1689 *Figure A1. Phase space presenting the effects of entrainment on cloud buoyancy, where*
 1690 *the initial cloudy parcel buoyancy (B_i) and final mixed parcel buoyancy (B_f) are*
 1691 *considered. A mixing fraction of 0.5 is chosen. The initial cloudy parcel is saturated*
 1692 *($S=1$), has a temperature of 15°C , pressure of 850 mb, and LWC of 1 g kg^{-1} . The X-axis*
 1693 *spans a range of environment relative humidity values (RH_{env}), and the Y-axis a*
 1694 *temperature difference ($dT_{env}=T_{env}-T_{cld}$) range between the cloud and the environment*
 1695 *parcels. Red color represents $B_i < 0$ & $B_f < 0$ (i.e. parcel stays negatively buoyant after*
 1696 *the mixing), magenta represents $B_i < 0$ & $B_f > 0$ (i.e. transition from negative to positive*
 1697 *buoyancy), green represents $B_i > 0$ & $B_f < 0$ (i.e. transition from positive to negative*
 1698 *buoyancy), and blue represents $B_i > 0$ & $B_f > 0$ (i.e. parcel stays positively buoyant).*
 1699 *The grey color represents mixed parcels that were depleted from water (LWC value*
 1700 *lower than 0.01 g kg^{-1}) after evaporation, and are considered non-cloudy. The white*
 1701 *line separates between areas where $B_f > B_i$ and $B_f < B_i$.*

1702



1703

1704 *Figure B1. Phase space presenting the resultant buoyancy (left panel) and relative*
 1705 *humidity (RH, right panel) when mixing B_{core} and B_{margin} parcels with equal RH but*
 1706 *different temperatures. A mixing fraction of 0.5 is chosen. Both parcels are initially*
 1707 *saturated (RH=100%), and have a LWC of 0.5 g kg^{-1} . The environment has a*
 1708 *temperature of 15°C and pressure of 850 mb . The X(Y)-axis spans the range of*
 1709 *temperature differences between the B_{core} (B_{margin}) parcel and the environment.*

1710

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