1	XCO <sub>2</sub> in an emission hot-spot region: the COCCON Paris campaign 2015
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30	Abstract. Providing timely information on urban Greenhouse-Gas (GHG) emissions and their
31	trends to stakeholders relies on reliable measurements of atmospheric concentrations and the
32	understanding of how local emissions and atmospheric transport influence these
33	observations.
34	Portable Fourier Transform Infra-Red (FTIR) spectrometers were deployed at 5 stations in the
35	Paris metropolitan area to provide column-averaged concentrations of $CO_2$ (XCO <sub>2</sub> ) during a
36	field campaign in spring of 2015, as part of the Collaborative Carbon Column Observing
37	Network (COCCON). Here, we describe and analyze the variations of $XCO_2$ observed at

different sites and how they changed over time. We find that observations upwind and downwind of the city centre differ significantly in their XCO<sub>2</sub> concentrations, while the overall variability of the daily cycle is similar, i.e., increasing during night-time with a strong decrease (typically 2-3 ppm) during the afternoon.

42 An atmospheric transport model framework (CHIMERE-CAMS) was used to simulate XCO<sub>2</sub> and predict the same behaviour seen in the observations, which supports key findings, e.g. 43 44 that even in a densely populated region like Paris (over 12 Million people), biospheric uptake 45 of CO<sub>2</sub> can be of major influence on daily XCO<sub>2</sub> variations. Despite a general offset between modelled and observed XCO<sub>2</sub>, the model correctly predicts the impact of the meteorological 46 47 parameters (e.g. wind direction and speed) on the concentration gradients between different 48 stations. Looking at the local gradients of XCO<sub>2</sub> for upwind and downwind station pairs, is 49 found to be less sensitive to changes in XCO<sub>2</sub> boundary conditions and biogenic fluxes within 50 the domain and we find the model-data agreement further improves. Our modelling framework 51 indicates that the local XCO<sub>2</sub> gradient between the stations is dominated by the fossil fuel CO<sub>2</sub> 52 signal of the Paris metropolitan area. This further highlights the potential usefulness of XCO<sub>2</sub> 53 observations to help optimise future urban GHG emission estimates.

54

#### 55 **1 Introduction**

56 Atmospheric background concentrations of CO<sub>2</sub> measured since 1958 in Mauna Loa, USA, 57 have passed the symbolic milestone of 400 ppm (monthly mean) as of 2013 [Jones 2013]. 58 Properly quantifying fossil fuel CO<sub>2</sub> emissions (FFCO<sub>2</sub>) can contribute to define effective climate mitigation strategies. Focussing our attention on cities is a critical part of this 59 60 endeavour as emissions from urban areas are currently estimated to represent from 53 % to 61 87 % of global FFCO<sub>2</sub>, depending on the accounting method considered, and are predicted to 62 increase further [IPCC-WG3 2014, IEA 2008, Dhakal 2009]. As stated in the IPCC 5th 63 assessment report, "current and future urbanisations trends are significantly different from the past' and "no single factor explains variations in per-capita emissions across cities and there 64 are significant differences in per capita greenhouse gas (GHG) emissions between cities 65 within a single country" [IPCC-WG3 2014]. Therefore, findings in one city can often not be 66 67 simply extrapolated to other urban regions. Furthermore, the large uncertainty of the global 68 contribution of urban areas to CO<sub>2</sub> emissions today and in the future is why a new generation 69 of city-scale observing and modelling systems are needed.

In recent years, more and more atmospheric networks have emerged that observe GHG
concentrations using the atmosphere as a large-scale integrator, for example in Paris, France
(e.g., Bréon et al. 2015, Xueref-Remy et al, 2018), Indianapolis, USA (e.g. Turnbull et al. 2015,
Lauvaux et al. 2016), Salt Lake City, USA (Strong et al. 2011, Mitchell et al. 2018), Heidelberg,
Germany (e.g. Levin et al. 2011, Vogel et al. 2013) and Toronto, Canada (e.g. Vogel et al.

75 2012). The air measured at in-situ ground-based stations is considered to be representative 76 of surface CO<sub>2</sub> fluxes of a larger surrounding area (1 km<sup>2</sup>-10000 km<sup>2</sup>), i.e. the emissions of the Greater Paris Area dominate the airshed of the Ile-de-France (ca. 12'000 km<sup>2</sup>) (Staufer et 77 78 al. 2016). If  $CO_2$  measurements are performed both up-wind and downwind of a city, the 79 concentration gradient between the two locations is influenced by the local net emission 80 strength between both sites and atmospheric mixing [Xueref-Remy et al, 2018, Bréon et al. 81 2015, Turnbull et al. 2015]. To derive quantitative flux estimates, measured concentration data 82 are typically assimilated into numerical atmospheric transport models which calculate the 83 impact of atmospheric mixing on concentration gradients for a given flux space-time 84 distribution. Such a data assimilation framework implemented for Paris with three atmospheric CO<sub>2</sub> measurement sites [Xueref-Remy et al, 2018] previously allowed deriving quantitative 85 estimates of monthly emissions and their uncertainties over one year [Staufer et al. 2016]. 86

87 Space-borne measurements of the column-average dry air mole fraction of CO<sub>2</sub> (XCO<sub>2</sub>) are 88 increasingly considered for the monitoring of urban CO<sub>2</sub>. This potential was shown with OCO-89 2 and GOSAT XCO<sub>2</sub> measurements, even though the spatial coverage and temporal sampling 90 frequency of these two instruments were not optimized for FFCO<sub>2</sub> [Kort et al., 2012, 91 Janardanan et al. 2016, Schwandner et al. 2017], while other space-borne sensors dedicated 92 to FFCO<sub>2</sub> and with an imaging capability are in preparation [O'Brien et al, 2016, Broquet et al. 93 2017]. Important challenges of satellite measurements are that they are not as accurate as in-94 situ ones, having larger systematic errors, while the XCO<sub>2</sub> gradients in the column are typically 95 7-8 times smaller than in the boundary layer. Another difficulty of space-borne imagery with passive instruments is that they will only sample city XCO<sub>2</sub> plumes during clear sky conditions 96 97 for geostationary satellites and with an additional constraint to observations at around mid-98 day for low-earth orbiting satellites.

99 The recent development of a robust portable ground-based FTIR (Fourier Transform InfraRed) 100 spectrometer as described in Gisi et al. [2012] and Hase et al. [2015] (EM27/SUN, Bruker 101 Optik, Germany) greatly facilitates the measurement of  $XCO_2$  from the surface, with better 102 accuracy than from space and with the possibility of continuous daytime observation during 103 clear sky conditions. Typical compatibility (uncorrected bias) of the EM27/SUN retrievals of 104 the different instruments in a local network is better than 0.01 % (i.e. ` 0.04 ppm) after a careful 105 calibration procedure and a harmonized processing scheme for all spectrometers [Frey et al. 2015]. The Collaborative Carbon Column Observing Network (COCCON) [Frey et al. 2018] 106 107 intends to offer such a framework for operating the EM27/SUN. This type of spectrometer 108 therefore represents a remarkable opportunity to document XCO<sub>2</sub> variability in cities as a direct 109 way to estimate FFCO<sub>2</sub> [Hase et al. 2015] or in preparation of satellite missions. 110 When future low-Earth-orbit operational satellites with imaging passive spectrometers of

suitable capabilities to invert FFCO<sub>2</sub> will sample different cities, this will likely be limited to clear

sky conditions and at a time of the day close to local noon. Increasing the density of the COCCON network stations around cities will allow to evaluate those XCO<sub>2</sub> measurements and to monitor XCO<sub>2</sub> during the early morning and afternoon periods which will not be sampled with satellites low-earth orbit satellite., From geostationary orbit, which can also have other benefits, those time-periods can however be observed and could be compared to groundbased measurements [e.g. Butz et al., 2015, O'Brien et al. 2016].

118 This study focuses on the measurements of XCO<sub>2</sub> from ground based EM27/SUN 119 spectrometers deployed within the Paris metropolitan area during a field campaign in the 120 spring of 2015, and modelling results. This campaign can be seen as a demonstration of the 121 COCCON network concept applied to the quantification of an urban FFCO<sub>2</sub> source. Several 122 spectrometers were operated by different research groups, while closely following the 123 common procedures suggested by Frey at al. [2015]. The paper is organised as follows. After 124 the instrumental and modelling setup descriptions of section 2, the observations of the field 125 campaign and the modelling results will be presented in section 3. Results are discussed in 126 section 4 together with the study conclusions.

127

#### 128 2 Methods and materials

## 129 **2.1 Description of study area and field campaign design**

During the COCCON field campaign (April 28th to May 13th, 2015) five portable FTIR spectrometers (EM27/SUN, Bruker Optik, Karlsruhe, Germany) were deployed in the Parisian region (administratively known as *Île-de-France*) and within the city of Paris. The campaign was conducted in early spring as the cloud cover is typically low in April and May and the time between sunrise and sunset is more than 14 hours.

135 The Paris metropolitan area houses over 12 million people, with about 2.2 million inhabiting 136 the city of Paris. This urban region is the most densely populated in France with ~1000 inhabitants/km<sup>2</sup> and over 21000 inhabitants/km<sup>2</sup> for the city of Paris itself [INSEE 2016 -137 https://www.insee.fr/fr/statistiques]. The estimated CO<sub>2</sub> emissions from the metropolitan 138 139 region are 39 Mt/year, according to the air quality association (AIRPARIF), that monitors the 140 airshed of Greater Paris. On-road traffic emissions, residential and the tertiary (i.e. 141 commercial) sector are the main sources (accounting for over 75 %), and minor contributions 142 from other sectors such as industrial sources and airports [ [https://www.airparif.asso.fr/en/, 143 AIRPARIF 2016]. It was crucial to understand the spatial distribution of these CO<sub>2</sub> sources to 144 optimally deploy the COCCON spectrometers. To this end a 1 km emission model for France 145 by IER (Institut fuer Energiewirtschaft und Rationelle Energieanwendung, University of 146 Stuttgart, Germany) was used as a starting point [Latoska 2009]. This emission inventory is 147 based on the available activity data such as, e.g., traffic counts, housing statistics, or energy 148 use, and the temporal disaggregation was implemented according to Vogel et al. [2013]. In brief, the total emissions of the IER model were re-scaled to match the temporal factors forthe different emission sectors according to known national temporal emission profiles.

- 151 To quantify the impact of urban emissions on  $XCO_2$ , the FTIR instruments were deployed
- along the dominant wind directions in this region in spring, i.e., southwesterly [Staufer et al
- 153 2016], in order to maximize the likelihood to capture upwind and downwind air masses (see
- 154 Figure 1). The two southwesterly sites (GIF and RES) are located in a less densely populated
- area, where emissions are typically lower than in the city centre, where the station JUS is
  located. The data in Fig. 1 show that the densest FFCO<sub>2</sub> emission area extends northwards
  and eastwards. The two Northwesterly sites (PIS and MIT) were placed downwind of this area.
- All instruments were operated manually and typically started operations around 7-8 am local
- time from which they continuously observe XCO<sub>2</sub> until 5-6 pm.
- 160

## 161 2.2 Instrumentation, calibration, and data processing

162 The EM27/SUN is a portable FTIR spectrometer which has been described in detail in, e.g., 163 Gisi et al. [2012] and Frey et al. [2015]. Here, only a short overview is given. The centre piece 164 of the instrument is a Michelson interferometer which splits up the incoming solar radiation 165 into two beams. After inserting a path difference between the beams, the partial beams are 166 recombined. The modulated signal is detected by an InGaAs detector covering the spectral 167 domain from 5000 to 11000 cm<sup>-1</sup> and is called an interferogram. As the EM27/SUN analyzes 168 solar radiation, it can only operate in daylight sunny conditions. A Fourier transform of the 169 interferogram generates the spectrum and a DC correction is applied to remove the 170 background signal and only keep the AC signal (see Keppel-Aleks et al. [2007]). A numerical 171 fitting procedure (PROFFIT code) [Schneider and Hase et al., 2009] then retrieves column 172 abundances of the concentrations of the observed gases from the spectrum. The single-173 channel EM27/SUN is able to measure total columns of O<sub>2</sub>, CO<sub>2</sub>, CH<sub>4</sub> and H<sub>2</sub>O.The ratio over 174 the observed O<sub>2</sub> column, assumed to be known and constant, delivers the column-averaged 175 trace gas concentrations of XCO<sub>2</sub>, XCH<sub>4</sub> in µmol / mol dry air, with a temporal resolution of 176 one minute.  $XCO_2$  is the dry air mole fraction of  $CO_2$ , defined as  $XCO_2$  = Column[CO<sub>2</sub>] / 177 Column[Dry Air]. Applying the ratio over the observed oxygen (O<sub>2</sub>) column reduces the effect 178 of various possible systematic errors; see Wunch et al. (2011).

In order to correctly quantify small differences in XCO<sub>2</sub> columns between Paris city upstream and downstream locations, measurements were performed with the five FTIR instruments side by side before and after the campaign, as we expect small calibration differences between the different instruments due to slightly different alignment for each individual spectrometer. These differences are constant over time and can be easily accounted for by applying a calibration factor for each instrument. Previous studies showed that the instrument specific corrections are well below 0.1 % for XCO<sub>2</sub> [Frey et al. 2015, Chen et al. 2016] and are stable for individual devices. The 1-sigma precision for XCO<sub>2</sub> is in the order of 0.01 % - 0.02 % (< 0.08 ppm) e. g.</li>
[Gisi et al. 2012, Chen et al. 2016, Hedelius et al. 2016, Klappenbach et al. 2015]. The
calibration measurements for this campaign were performed in Karlsruhe w.r.t. the Total
Carbon Column Observing Network (TCCON) [Wunch et al. 2011] spectrometer at the
Karlsruhe Institute of Technology (KIT), Germany for 7 days before the Paris campaign
between April 9<sup>th</sup> and 23<sup>rd</sup>, and after the campaign on May 18<sup>th</sup> until 21<sup>st</sup>.

192 Figure S1 (left panel) shows the XCO<sub>2</sub> time series of the calibration campaign, where small 193 offsets between the instruments raw data are visible. As these offsets are constant over time, 194 a calibration factor for each instrument can be easily applied; actually these are the calibration 195 factors previously found for the Berlin campaign [Frey et al. 2015]. These factors are given in 196 Table 1, where all EM27/SUN instruments are scaled to match instrument No. 1. The 197 calibrated XCO<sub>2</sub> values for April 15<sup>th</sup> are shown in Fig. S1 None of the five instruments that participated in the Berlin campaign show any significant drift; in other words, the calibration 198 199 factors found one year before were still applicable. This is a good demonstration of the 200 instrument stability stated in section 2.2, especially as several instruments (Nos. 1, 3, 5) were 201 used in another campaign in Northern Germany in the meantime. The EM27/SUN XCO<sub>2</sub> 202 measurements can also be made traceable to the WMO international scale for in-situ 203 measurements by comparison with measurements of a collocated TCCON spectrometer 204 which are calibrated against in-situ standards by aircraft and aircore measurements [Wunch 205 et al. 2010, Messerschmidt et al. 2012] performed using the WMO scale.

During the campaign and for the calibration measurements we recorded double-sided interferograms with 0.5 cm<sup>-1</sup> spectral resolution. Each measurement of 58 s duration consisted of 10 scans using a scanner velocity of 10 kHz. For precise timekeeping, we used GPS sensors for each spectrometer.

210 In-situ surface pressure data used for the analysis of the calibration measurements performed 211 at KIT have been recorded at the co-located meteorological tall tower. During the campaign, 212 a MHD-382SD data-logger recorded local pressure, temperature and relative humidity at each 213 station. The analysis of the trace gases from the measured spectra for the calibration 214 measurements has been performed as described by Frey et al. [2015]. For the campaign 215 measurements we assume a common vertical pressure-temperature profile for all sites, 216 provided by the model, so that the surface pressure at each spectrometer only differs due to 217 different site altitudes. The 3-hourly temperature profile from the European Centre for Medium-218 Range Weather Forecasts (ECMWF) operational analyses interpolated for site JUS located in 219 the centre of the array was used for the spectra analysis at all sites. The individual ground-220 pressure was derived from site altitudes and pressure measurements performed at each site. 221 Before and after the Paris campaign, side by side comparison measurements were performed 222 with all 5 EM27/SUN spectrometers and the TCCON spectrometer operated in Karlsruhe at 223 KIT. All spectrometers were placed on the top of the IMK office building North of Karlsruhe. 224 The altitude is 133 m above sea level (a.s.l.), coordinates are 49.09° N and 8.43° E. The 225 processing of the Paris raw observations (measured interferograms) were performed as 226 described by Gisi et al. [2012] and Frey et al. [2015] for the Berlin campaign: spectra were 227 generated applying a DC correction, a Norton-Beer medium apodization function and a 228 spectral resampling of the sampling grid resulting from the FFT on a minimally sampled 229 spectral grid. PROFFWD was used as the radiative transfer model and PROFFIT as the 230 retrieval code.

231

## 232 **2.3 Atmospheric transport modelling framework**

233 We used the chemistry transport model CHIMERE (Menut et al., 2013) to simulate CO<sub>2</sub> 234 concentrations in the Paris area. More specifically, we used the CHIMERE configuration over 235 which the inversion system of Bréon et al. [2015] and Staufer et al. [2016] was built to derive 236 monthly to 6-hour mean estimates of the CO<sub>2</sub> Paris emissions. Its horizontal grid, and thus its 237 domain and its spatial resolution, are illustrated in Figure S2. It has a  $2 \times 2 \text{ km}^2$  spatial resolution for the Paris region, and  $2 \times 10 \text{ km}^2$  and  $10 \times 10 \text{ km}^2$  spatial resolutions for the 238 239 surroundings. It has 20 vertical hybrid pressure-sigma (terrain- following) layers that range 240 from the surface to the mid-troposphere, up to 500 hPa. It is driven by operational 241 meteorological analyses of the ECMWF Integrated Forecasting System, available at an 242 approximately  $15 \times 15$  km<sup>2</sup> spatial resolution and 3 h temporal resolution.

243 In this study the CO<sub>2</sub> simulations are based on a forward run over April 25<sup>th</sup> - May 12<sup>th</sup> 2015 244 with this model configuration; we do not assimilate atmospheric CO<sub>2</sub> data and so no inversion 245 for surface fluxes was conducted. In the Paris area (the Île-de-France administrative region), 246 hourly anthropogenic emissions are given by the IER inventory, see section 2.1. The 247 anthropogenic emissions in the rest of the domain are prescribed from the EDGAR V4.2 database for the year 2010 at 0.1° resolution [Olivier and Janssens-Maenhout et al., 2012]. In 248 249 the whole simulation domain, the natural fluxes (the Net Ecosystem Exchange: NEE) are 250 prescribed using simulations of C-TESSEL, which is the land-surface component of the 251 ECMWF forecasting system [Boussetta et al., 2013], at a 3 hourly and 15 × 15 km<sup>2</sup> resolution. 252 Finally, the CO<sub>2</sub> boundary conditions at the lateral and top boundaries of the simulation domain 253 and the simulation CO<sub>2</sub> initial conditions on April 25th 2015 are prescribed using the CO<sub>2</sub> 254 forecast issued by the Copernicus Atmosphere Monitoring Service (CAMS, 255 http://atmosphere.copernicus.eu/) at a ~15 km global resolution [Agusti-Panareda et al.,

256 2014].

The CHIMERE transport model is used to simulate the XCO<sub>2</sub> data. However, since the model does not cover the atmosphere up to its top, the CO<sub>2</sub> fields from CHIMERE are complemented with that of the CAMS CO<sub>2</sub> forecasts from 500 hPa to the top of the atmosphere to derive total column concentrations. The derivation of modelled XCO<sub>2</sub> at the sites, involves obtaining a kernel-smoothed CO<sub>2</sub> profile of CHIMERE and CAMS and vertical integration of these smoothed profiles, weighted by the pressure at the horizontal location of the sites.

The parametrisation used to smooth modelled  $CO_2$  profiles approximates the sensitivity of the EM27/sun  $CO_2$  retrieval is a function of pressure and sun elevation. Between 1000 hPa and 480 hPa, a linear dependency of the instrument averaging kernels on solar zenith angle ( $\Theta$ ) is assumed with boundary values following Frey et al. [2015]:

267

268 (1a) 
$$k(480 hPa) = 1.125$$

- 269 (1b)  $k(1000 hPa) = 1.0 + 0.45 s^3$
- 270

where  $s = \Theta/90^\circ$ . *k*. Approximate averaging kernels are obtained bylinear interpolation to the pressure levels of CHIMERE and CAMS, respectively. If *p* > 1000 hPa, *k* is linearly extrapolated. Above 480 hPa (*p* < 480 hPa), the averaging kernels can be approximated by

(2) 
$$k(u,s) = 1.125 - 0.6 u^3 - 0.4 u s^3$$

276

where *u* is (480 hPa - p) / 480. The kernel-smoothed CO<sub>2</sub> profile, CO<sub>2\_model</sub><sup>s</sup>, is obtained by 278

279 (3) 
$$CO_{2\_model}^{s} = K CO_{2\_model} + (I - K)CO_{2}^{a}$$

280 281

where  $CO_{2\_model}$  is the modelled  $CO_2$  profile by CHIMERE or CAMS, *I* the identity matrix and *K* is a diagonal matrix containing the averaging kernels *k*. The a priori  $CO_2$  profile,  $CO_2^a$ , is provided by the Whole Atmosphere Community Climate Model (WACCM) model (version 6) and interpolated to the pressure levels of CHIMERE and CAMS.  $CO_{2\_model}^s$  is the appropriate  $CO_2$  profile to calculate modelled XCO<sub>2</sub> at the location of the sites.

287

For a given site, the simulated XCO<sub>2</sub> data are thus computed from the vertical profile of this site as:

290

(3) 
$$XCO_{2\_CHIMERE} = \frac{1}{Psurf} \int_{Psurf}^{Ptop\_CHIM} CO_{2\_CHIM}^{s} dp + \int_{Ptop\_CHIM}^{P=0mbar} CO_{2\_CAMS}^{s} dp$$

292

where  $p_{surf}$  is the surface pressure,  $p_{top_CHIM} = 500$  hPa the pressure corresponding to the top boundary of the CHIMERE model, and  $CO_{2_CHIM}^{s}$  and  $CO_{2_CAMS}^{s}$  are the smoothed  $CO_2$  concentrations of CHIMERE and CAMS respectively. For comparison we also calculated
 XCO<sub>2</sub> at a lower spatial resolution with the CAMS data alone as:

- 297
- 298

$$XCO_{2\_CAMS} = \int_{Psurf}^{P=0mbar} CO_{2\_CAMS}^{s} dP$$

299

# 300 3 Results and discussion

#### 301 3.1 Observations

(4)

## 302 **3.1.1** Meteorological conditions and data coverage/instrument performance.

303 During the measurement campaign (April 28<sup>th</sup> until May 13<sup>th</sup>, 2015), meteorological conditions 304 were a major limitation for the availability of XCO<sub>2</sub> observations. Useful EM27/SUN 305 measurements require direct sunlight and low wind speeds typically yield higher local XCO2. 306 Most of the time during the campaign conditions were partly cloudy and turbid, and so 307 successful measurements at high solar zenith angle (SZA) were rare. Therefore, the data 308 coverage between April 28<sup>th</sup> and May 3<sup>rd</sup> is limited (see Table 2). As is typical for spring periods 309 in Paris, the temperature and the wind direction vary and display less synoptic variations than 310 in winter. The dominant wind directions were mostly northeasterly at the beginning of the 311 campaign and mostly southeasterly during the second half of the campaign. We find that the wind speeds during daytime nearly always surpass 3 m s<sup>-1</sup>, which has been identified by Breon 312 313 et al. [2015] and Staufer et al. [2016] as the cut-off wind speed above which the atmospheric 314 transport model CHIMERE performs best in modelling CO<sub>2</sub> concentration gradients in the 315 mixed layer.

Despite some periods with unfavourable conditions, more than 10,000 spectra were retrieved among the five deployed instruments. The quality of the spectra for each day was rated according to the overall data availability and consistent with Hase et al. (2015). The best measurement conditions prevailed for the period between May 7<sup>th</sup> and May 12<sup>th</sup>.

320

### 321 3.1.2 Observations of XCO<sub>2</sub> in Paris

322 The observed XCO<sub>2</sub> in the Paris region for all sites (10415 observations) ranges from 397.27 323 to 404.66 ppm with a mean of 401.26 ppm (a median of 401.15 ppm). The strong atmospheric 324 variability of XCO<sub>2</sub> across Paris and within the campaign period is reflected in the standard 325 deviation of 1.04 ppm for 1-minute averages. We find that all sites exhibit very similar diurnal 326 behaviours with a clear decrease of XCO<sub>2</sub> during daytime and a noticeable day-to-day 327 variability as seen in Figure 2. This is to be expected as they are all subject to very similar 328 atmospheric transport in the boundary layer height and to similar large-scale influences, i.e., 329 surrounding with stronger natural fluxes or air mass exchange with other regions at synoptic 330 time scales. However, observed XCO<sub>2</sub> concentrations at the downwind sites for our network 331 remain clearly higher from sites that are upwind of Paris (see Figure 2). The shifting dominant 332 wind conditions also explain why the site RES and GIF are lowest in the beginning of the campaign and higher on May 12<sup>th</sup> and 13<sup>th</sup> after meteorological conditions changed. This 333 334 indicates that the influence of urban emissions is detectable with this network configuration 335 under favourable meteorological conditions. By comparing the different daily variations in Fig. 336 3, it is apparent that the day-to-day variations observed at the two southwesterly (typically 337 upwind) sites GIF and RES are approximately 1 ppm, with both sites exhibiting similar diurnal 338 variations throughout the campaign period. This can be expected as their close vicinity would 339 suggest that they are sensitive to emissions from similar areas and to concentrations of air 340 masses arriving from the southwest.

The typical decrease in XCO<sub>2</sub> found over the course of a day is about 2 to 3 ppm. This decrease could be driven by (natural) sinks of CO<sub>2</sub>, which can be expected to be very strong as our campaign took place after the start of the growing season in Europe for most of southern and central Europe [Roetzer and Chmielewski 2001].

345 The observations at the site located in Paris (JUS) displays similarly low day-to-day variations 346 and a clear decrease in XCO<sub>2</sub> over the course of the day. The latter feature indicates that even 347 in the dense city centre, XCO<sub>2</sub> is primarily representative of a large footprint like in other areas 348 of the globe [Keppel-Aleks 2011] and supports the findings of Belikov et al. (2017) concerning 349 the footprints for the Paris and Orleans TCCON sites, Thus, our total column observations are 350 less critically affected by local emissions than in-situ measurements [Breon et al. 2015, 351 Ammoura et al. 2016]. It is also apparent that the decrease in XCO<sub>2</sub> (the slope) during the afternoon for April 28<sup>th</sup> and 29<sup>th</sup> as well as May 7<sup>th</sup> and 10<sup>th</sup> is noticeably smaller than at other 352 days during this campaign. As XCO<sub>2</sub> is not sensitive to vertical mixing, this has to be caused 353 354 by different CO<sub>2</sub> sources and sinks acting upon the total column arriving at JUS.

The two (typically downwind) sites PIS and MIT northeast of Paris show a markedly larger day-to-day spread in their general XCO<sub>2</sub> levels as well as strongly changing slopes for the diurnal XCO<sub>2</sub> decrease. For these sites the exact wind direction is critical as they can be downwind of the city centre that has a much higher emission density or less dense suburbs (see Fig. 1).

360

#### 361 3.1.3 Gradients in observed XCO<sub>2</sub>

In order to focus more on the impact of local emissions on atmospheric conditions and less on that of of  $CO_2$  fluxes from outside of our urban domain in our analysis of  $XCO_2$ , we choose to study the spatial gradients ( $\Delta$ ) between different sites. Fundamentally, this approach assumes that regional and large-scale fluxes have a similar impact on  $XCO_2$  for the sites within our network, due to the close proximity of sites and the smoothing of remote emission signals due to atmospheric transport by the time the air-mass arrives in our domain. Ideal conditions were

sampled during May 7<sup>th</sup>, with predominantly southwesterly winds, and on May 10<sup>th</sup> with 368 369 southerly winds. We can see in Fig. 4 that all sites were, on average, elevated compared to 370 RES, chosen as reference here as it was upwind of Paris during those days. The hodographs 371 for both days also indicate that the wind fields were consistent across Paris (see Figure S3). 372 The observations from GIF showed only minimal differences with RES, while the rest of the 373 sites (PIS, JUS and MIT) had  $\Delta$  values of 1 to 1.5 ppm. During southwesterly winds, MIT is 374 downwind of the densest part of the Paris urban area, and JUS is impacted by emissions of 375 neighborhoods to the southwest. The site of PIS is still noticeably influenced by the city centre 376 but, as can be seen in Fig. 1, we likely do not catch the plume of the most intense emissions 377 but rather from the suburbs. On May 10<sup>th</sup>, with its dominant southerly winds, the situation was 378 markedly different. While GIF was still only slightly elevated, the XCO<sub>2</sub> enhancement at MIT 379 was significantly lower and quite similar to JUS for large parts of the day. The highest  $\Delta XCO_2$ 380 can be observed at PIS, again typically ranging from 1 to 1.5 ppm. As seen in Fig. 1, PIS is 381 then directly downwind of the densest emission area, while MIT is only exposed to CO<sub>2</sub> 382 emissions from the eastern outskirts of Paris.

It is also important to note that the impact of the local biosphere that is assumed to cause the
strong decrease in XCO<sub>2</sub> during the day is not seen on both days for these spatial gradients.
For a more comprehensive interpretation of these observations the use of a transport model
(as described in section 2.3) is necessary.

387

## 388 3.2 Modelling

### 389 3.2.1 Model performance

390 Before interpreting the modelled XCO<sub>2</sub> we need to evaluate the performance of the chosen 391 atmospheric transport model framework as described in section 2.3. Comparing it to 392 meteorological observations (wind speed and wind direction) at GIF, we find that CHIMERE 393 predicts these variables well throughout the duration of the campaign (see Figure S4). 394 Changes in wind speed direction and speed are reproduced with a slight overestimation at low 395 wind-speeds (>1m/s). Besides the meteorological forcing, the model performance can also be 396 expected to depend on the chosen model resolution. Therefore, we compared XCO<sub>2</sub> at JUS 397 calculated based on the coarser resolution atmospheric transport and flux framework CAMS 398 (15 km), and the higher resolution emission modelling input for the framework based on 399 CHIMERE (2 km) for the inner domain and on CAMS boundary conditions (see Fig. S2). We 400 find that the coarser model displays similar inter-daily variations, but that the high-resolution 401 model modifies the modelling results on shorter time-scales. We find that the afternoon XCO<sub>2</sub> 402 decreases are often more pronounced in CHIMERE. Only the high-resolution will be 403 considered and referred to in the following. The impact of using different flux maps (fossil fuel 404 CO<sub>2</sub>) on the modelled XCO<sub>2</sub> can unfortunately not be explicitly investigated here as only one

high-resolution (1 km) emission product available for fossil fuel CO<sub>2</sub> was available for this
study region (see section 2.3) and other global emission products are usually not intended for
urban-scale studies.

408

# 409 **3.2.2 Modelled XCO<sub>2</sub> and its components**

410 The modelled XCO<sub>2</sub> for the five sites (Fig. 5) co-evolves over the period of the campaign with 411 occurrences of significant differences. This was already seen with the measurements, but the 412 model allows looking at the full time series. The model reveals clear daily cycles of XCO<sub>2</sub>, with 413 an accumulation during night-time and a decrease during daytime. Despite a good general 414 agreement of modelled XCO<sub>2</sub> at all sites for, e.g., the timing of daily minima and their synoptic 415 changes, differences in XCO<sub>2</sub> are observed between the sites for many days. Typically the 416 northeasterly sites (PIS, MIT) show an enhancement in modelled XCO<sub>2</sub> compared to the 417 southwesterly sites (GIF, RES).

418 To understand the synoptic and diurnal variations of the modelled XCO<sub>2</sub>, we analyzed the 419 contribution of different sources (and sinks) of CO<sub>2</sub>, namely the net ecosystem exchange (NEE), the fossil fuel CO<sub>2</sub> emissions (FFCO2), and the boundary conditions (BC), i.e., the 420 421 variations of CO<sub>2</sub> not caused by fluxes within our domain (the example of JUS is given in Fig. 422 6). The day-to-day variability of modelled  $XCO_2$  is dominated by changing boundary conditions 423 and coincides with synoptic weather changes. As the CO<sub>2</sub> emitted from the different sources 424 is transported in the model as independent tracers, the strong daily decrease in XCO<sub>2</sub> can be 425 directly linked to NEE, which leads to a decrease of  $\sim 1$  ppm (but up to 4 ppm) during the day, 426 but can also cause positive enhancements during nighttime driven by biogenic respiration. 427 The XCO<sub>2</sub> from fossil fuel emissions causes significant enhancements compared to the 428 background, but is often compensated by NEE. During short periods, fossil fuel emissions can however lead to enhancements of up to 4 ppm. 429

430

#### 431 **3.2.3 Modelled** Δ**XCO**<sub>2</sub> gradients and its components

432 To be able to assess the impact of local sources and reduce the influence of NEE and BC on 433 the modelled signals, we analyse the XCO<sub>2</sub> gradient (i.e. station-to-station difference) with 434 RES being taken as reference. In Fig. 7 we compare  $\Delta$ , in the top panel, and its components, 435 i.e. fossil fuel CO<sub>2</sub>, biogenic CO<sub>2</sub> and CO<sub>2</sub> transported across the boundary of the domain 436 (boundary conditions: BC), along a south-north direction. For the modelled  $\Delta$  we can see that 437 MIT shows a positive value during the campaign period whenever the predominant wind 438 direction was southwesterly. We also find that  $\Delta$  between JUS and RES was both negative 439 and positive during the campaign, and predominantly negative between MIT and JUS. When split into FFCO2, BC and NEE components, we can clearly see that the total  $\Delta$  is dominated 440

441 by FF causing XCO<sub>2</sub> offsets of up to 4 ppm, but more typically 1 ppm gradients are observed. 442 Gradients can also change rapidly (within a few hours) if the wind direction changes, for example on May 1<sup>st</sup> and May 12<sup>th</sup>. This highlights the fact that, during such conditions, we 443 444 cannot assume a simple upwind-downwind interpretation of our sites. As expected, the 445 contributions from BC and NEE are generally greatly reduced when analysing  $\Delta XCO_2$ . The 446 most important impact of NEE on the XCO<sub>2</sub> gradients of -1ppm and +1ppm can be seen on 447 May 8<sup>th</sup> and May 11<sup>th</sup>, respectively. This means that, despite greatly reducing the impact of 448 NEE on average, the contribution of NEE cannot be fully ignored. BC is an overall negligible 449 contribution to  $\Delta XCO_2$ , even though it reaches -0.4 ppm on May 11<sup>th</sup>.

450

### 451 **3.3 Model data and observations comparison**

452 **3.3.1 XCO**<sub>2</sub>

453 A comparison of modelled and observed XCO<sub>2</sub> is of course limited to the relatively short 454 periods when observations are available. Over these periods we can see a general issue in 455 reproducing the general XCO<sub>2</sub> for each day in the model as observed XCO<sub>2</sub> is significantly 456 lower revealing a fairly stable bias between 1 to 2 ppm. As our CO<sub>2</sub> boundary conditions were 457 from a forecast product, this is not unexpected, as already small issues in estimating carbon 458 uptake (or emissions) at the European scale can have such an impact on the boundary 459 conditions. However, we observe that the main features, like daily cycles and synoptic 460 changes of the modelled and observed XCO<sub>2</sub> are comparable as seen in Figure 8. The 461 daytime variations are well reproduced by the model and the general relative concentrations between sites are preserved, e.g., the highest values for XCO<sub>2</sub> at MIT are on May 9<sup>th</sup> and 462 highest XCO<sub>2</sub> for PIS are later on May 10<sup>th</sup> and May 11<sup>th</sup>. We also see that the timing of the 463 464 daily minima is not fully covered in the observed data as it typically happens after sunset and 465 cessation of biosphere uptake. To reduce the impact of uncertainties of the boundary 466 conditions on our analysis a gradient approach was tested.

467

#### 468 **3.3.2 ∆XCO**<sub>2</sub>

469 Due to the prevailing southeasterly wind conditions, we can compare XCO<sub>2</sub> at the typical 470 downwind sites (PIS, MIT) relative to the mostly upwind sites (RES, GIF) and expect elevated 471 XCO<sub>2</sub> downwind. Furthermore, we can expect to see negative gradients for opposing wind 472 conditions, i.e. northwesterly. For other wind conditions, the concentration difference is not 473 determined by emissions between the station pairs, but rather by the areas upwind of the sites, 474 (see Figure 1). We find that the  $\Delta XCO_2$  of PIS relative to RES generally falls along the 1:1 line with a slope of  $1.07\pm0.09$  with a Pearson's R of 0.8. Negative  $\Delta XCO_2$  values, seen in Fig. 9, 475 476 are associated with meteorological conditions when winds come from northerly directions,

477 i.e., the roles of normal upwind and downwind sites are reversed. For wind perpendicular to 478 the direct line of sight for (PIS, RES) the concentration enhancements are small and harder to 479 interpret and no slope was calculated. The gradient of XCO<sub>2</sub> MIT relative to RES has a 480 significantly lower range for modelled XCO<sub>2</sub> while the observed range of XCO<sub>2</sub> is similar to 481 PIS. The slope of observed to modelled  $\Delta XCO_2$  for upwind-downwind (or downwind-upwind 482 conditions) is 1.72±0.06 with a Pearson's R of 0.96. This points to a significant 483 underestimation of the impact of urban sources on the MIT-RES gradient, which is especially 484 visible in the more negative  $\Delta XCO_2$  during northerly wind conditions. This could indicate that 485 the spatial distribution of our emissions prior should be improved, i.e., emissions in the eastern 486 outskirts/suburbs are likely underestimated in the IER emissions model. The low modelled 487  $\Delta XCO_2$  could also be due to overestimated horizontal dispersion in the model, which seems 488 less likely. Again the model does not predict concentration differences well for perpendicular 489 wind conditions. When comparing the mean modelled daily cycle of the days with south-490 westerly wind conditions and when observations exist with the mean diurnal cycle for all days 491 within the field campaign period when MIT and PIS can be considered downwind of RES, we 492 find that the days with observations do not significantly differ from those without observations 493 (see Fig. 10). An investigation of typical diurnal variations of modelled  $\Delta XCO_2$  can only be 494 performed to a limited degree with the observational data available for suitable wind conditions. Within the large uncertainties, the modelled and observed  $\Delta XCO_2$  agree 495 496 throughout the day. When analysing the modelled  $\Delta XCO_2$  components we also find that the observed daytime increases of  $\Delta XCO_2$  are driven by  $CO_2$  added by urban FF  $CO_2$  burning 497 498 and that the impact of FF is significantly higher at PIS (up to 1 ppm) then at MIT site (0.5 ppm) 499 in the model, when both sites are downwind of Parisian emissions. Our observations indicate 500 that both sites have strong diurnal variations. Given that the most important biogenic sinks, in 501 our domain, can be expected to be found in the rural parts surrounding Paris we would expect 502 the biogenic contribution to be similar at both sites (as predicted by the model). This would 503 further point towards, that the impact of FF emissions on the MIT site is larger than predicted 504 by our modelling framework.

505 Different  $\Delta XCO_2$  diurnal variations can be found for other upwind-downwind site pairs, but they 506 are all systematically driven by the locally-added CO<sub>2</sub> from FFCO<sub>2</sub>.

507

#### 508 **5 Conclusion and Outlook**

509 For the two-week field campaign we demonstrated the ability of a network of five EM27/SUN 510 spectrometers, placed in the outskirts of Paris, to track the XCO<sub>2</sub> changes due to the urban 511 plume of the city. However, we also found that XCO<sub>2</sub> cannot be simply interpreted in the 512 context of local emissions as, even in such a densely populated area, XCO<sub>2</sub> is still significantly

influenced by natural CO<sub>2</sub> uptake during the growing season. Understanding the area 513 514 influencing XCO<sub>2</sub> and/or the use of suitable atmospheric transport models seems 515 indispensable to correctly interpret atmospheric XCO<sub>2</sub> variations. Using a gradient approach, 516 i.e., analysing the difference between XCO<sub>2</sub> measured at upwind and downwind stations, 517 greatly reduced the impact of CO<sub>2</sub> boundary condition, that reflect fluxes outside the domain 518 and biogenic fluxes within the domain. Overall, the XCO<sub>2</sub> variability modelled using our 519 ECMWF-CHIMERE system with IER (1 x 1 km<sup>2</sup>) emissions data was found to be comparable 520 with the observed variability and diurnal evolution of XCO<sub>2</sub>, despite a higher background for 521 modelled XCO<sub>2</sub>. Our modelling framework, run at a 2 x 2 km<sup>2</sup> resolution over Paris also 522 predicts that biogenic fluxes and boundary conditions (i.e. the influence of  $CO_2$  being 523 transported into our domain) have only very small impact on  $\Delta XCO_2$  only noticeably impacting 524 it during a few situations, specifically when meteorological conditions changes made the 525 concept of 'upwind' and 'downwind' not applicable. When comparing modelled and measured 526  $\Delta XCO_2$  we find strong correlations (Pearson's R) of 0.8 and 0.96 for PIS-RES and MIT-RES, 527 respectively. The offset between model and observations also diminished for  $\Delta XCO_2$  and the 528 slope found between observed and modeled PIS-RES gradient is statistically in accordance 529 with a 1:1 relationship (1.07±0.09). However, the slope of the MIT-RES XCO<sub>2</sub> gradient of 530 1.72±0.06 suggests that the emission model could potentially be improved, as it seems 531 unlikely that the general atmospheric transport in the model is the key issue as both site pairs 532 would be subject to very similar winds. Another potential source of error that needs to be 533 investigated is if such an underestimation of  $\Delta XCO_2$  could be caused by the limited model 534 resolution. It also seems rather likely that a 2x2km<sup>2</sup> model would cause a general spreading 535 of point source emissions and not systematically underestimate emissions impacts from less 536 densely populated, parts of Île-de-France. The data also confirm previous results by models 537 that XCO<sub>2</sub> gradients caused by a megacity do not exceed 2 ppm, which supports the previous 538 requirement for satellite observations of less than 1 ppm precision on individual soundings, 539 and biases lower than 0.5 ppm (Ciais et al. 2015). The gradients are mainly caused by the 540 transport of FFCO<sub>2</sub> emissions but, interestingly, during specific episodes, a noticeable 541 contribution comes from biogenic fluxes, suggesting that these fluxes cannot always be 542 neglected even when using gradients.

543 Unfortunately, the duration of the campaign was relatively short, so that an in-depth analysis 544 of mean daily cycles or the impact of ambient conditions (traffic conditions, temperature, solar 545 insolation, etc.) on the observed gradient and underlying fluxes could not be investigated here. 546 Hence, future studies in Paris and elsewhere should aim to perform longer-term observations 547 during different seasons, which will allow better understanding changes in biogenic and 548 anthropogenic CO<sub>2</sub> fluxes. A remotely-controllable shelter for the EM27/SUN instrument is

- 549 currently under development [Heinle and Chen, 2017]. This will considerably facilitate the 550 establishment of permanent spectrometer arrays around cities and other sources of interest. 551 Nevertheless, our study already indicates that such observations of urban XCO<sub>2</sub> and  $\Delta$ XCO<sub>2</sub> 552 contain original information to understand local sources and sinks and that the modelling 553 framework used here is a step forward to support their detailed interpretation in the future. An 554 improved model will also be able to adjust or better model the background conditions and 555 potentially use this type of observations to estimate local CO<sub>2</sub> fluxes using a Bayesian 556 inversion scheme similar to the existing system based on in-situ observations for Paris [Staufer 557 et al. 2016].
- 558 We expect that the previous successful collaboration in the framework of the Paris campaign 559 will mark the permanent implementation of COCCON as a common framework for a French-
- 560 Canadian-German collaboration on the EM27/SUN instrument. The acquisition of additional
- 561 spectrometers is planned by several partners.
- 562

# 563 Author contribution

- 564 FRV, MF, FH, IXR, MKS, PCh, PJ, YVT, CJ, TB, QT and JO, supported the field campaign 565 and contributed data to this study.
- 566 MF, FH, FRV, JS, GB and PCi planned the fieldwork and modelling activities for this study.
- 567 JS, GB, FC, and FRV performed the CHIMERE modelling, provided modelling data input 568 and/or analysed the output data.
- 569 MF, FH and FRV processed and analysed the EM27Sun data.
- 570 FRV, MF, JS, FH and PCi wrote sections of the manuscript and created figures and tables.
- 571 All authors reviewed, edited and approved the manuscript.
- 572

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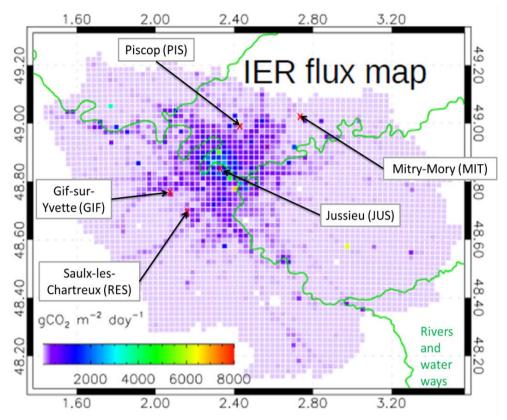
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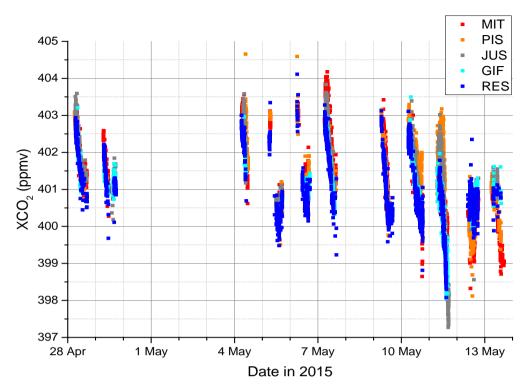
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Figure 1. CO<sub>2</sub> emissions in the Île-de-France region according to the IER emission inventory. Measurement sites are indicated by red crosses.



743 744 Figure 2. Time series of observed XCO<sub>2</sub> in the Parisian region for all five sites (all valid

745 data of 1 minute averages).

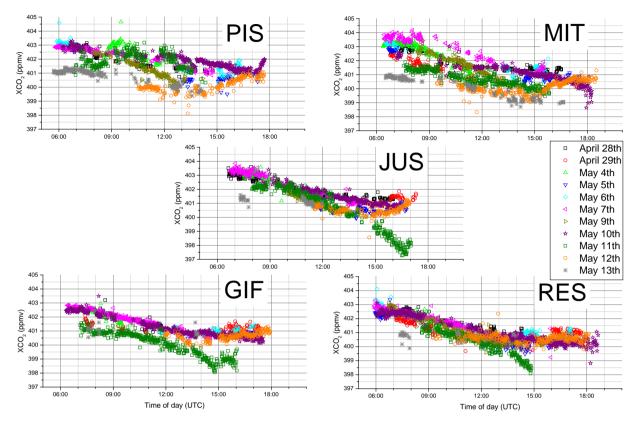


Figure 3. Time series of observed XCO<sub>2</sub> in the Parisian region sorted by station.

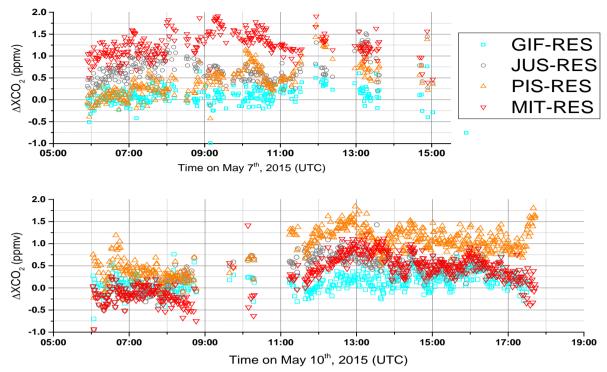
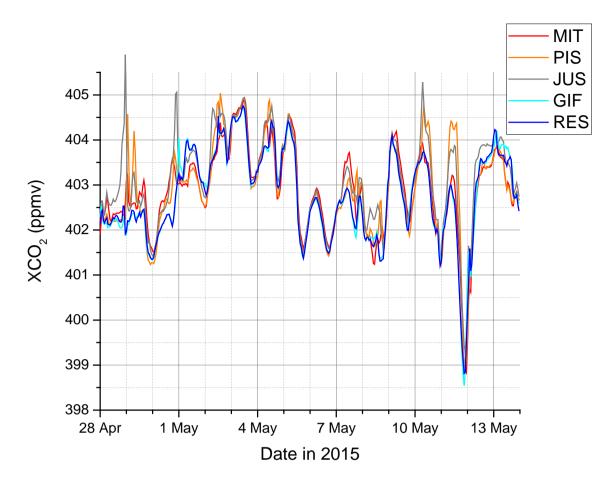
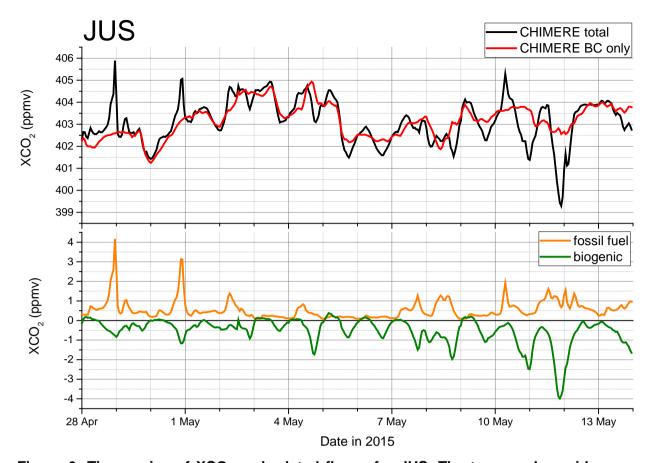


Figure 4. Observed spatial gradients of XCO<sub>2</sub> for May 7<sup>th</sup> (southwesterly winds) and May

**10<sup>th</sup> (southerly winds)**.



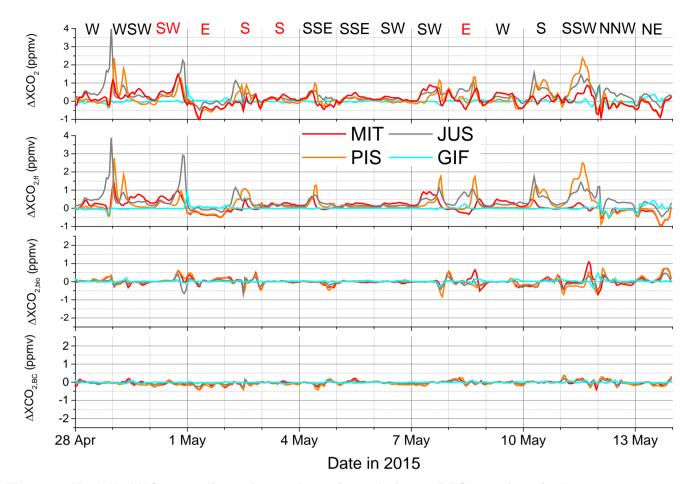
751752 Figure 5. Modelled XCO<sub>2</sub> for all stations.



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Figure 6: Time series of XCO<sub>2</sub> and related fluxes for JUS. The top panel provides a comparison of modelled total XCO<sub>2</sub> and XCO<sub>2</sub> variations due to changes in boundary conditions (BC only). The lower panel shows the contribution of the different flux

757 components, namely fossil fuel CO<sub>2</sub> emissions and biogenic fluxes.



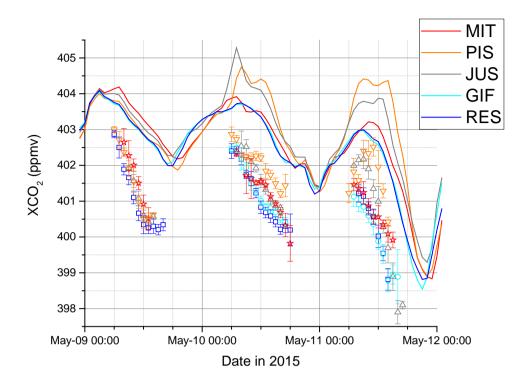
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Figure 7. Modelled XCO<sub>2</sub> gradients for each station relative to RES are given in the top panel with its contributing components in the panels below. Total  $\Delta$ XCO<sub>2</sub> (top), the

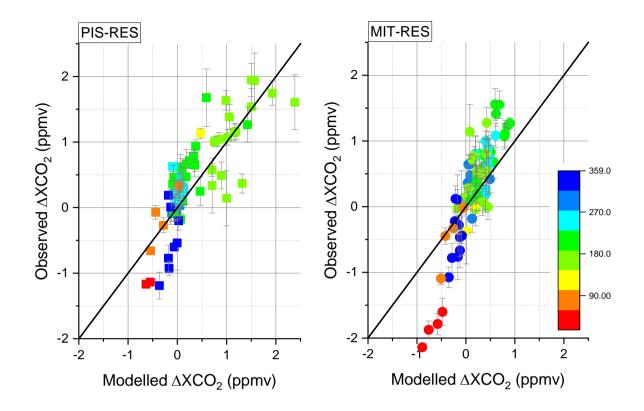
fossil fuel contribution  $\Delta$ XCO2,ff (second from top), the biogenic contribution  $\Delta$ XCO<sub>2,bio</sub> (third from top) and the influence of the boundary conditions,  $\Delta$ XCO<sub>2,BC</sub> (bottom), The

763 dominant wind conditions for each day given at the top of the figure and days without

764 observations due to precipitation are in red.



765 766 Figure 8. Comparison of modelled (solid lines) and observed hourly averaged XCO<sub>2</sub> (symbols) with standard deviations as error bars. 767



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Figure 9. Comparison of modelled and observed hourly averaged  $\Delta XCO_2$  for gradients between PIS and RES (left) and MIT and RES (right), with standard deviations of the minute values of the hourly mean as vertical bars and the points color coded by wind

772 direction from 0 to 359 degrees.

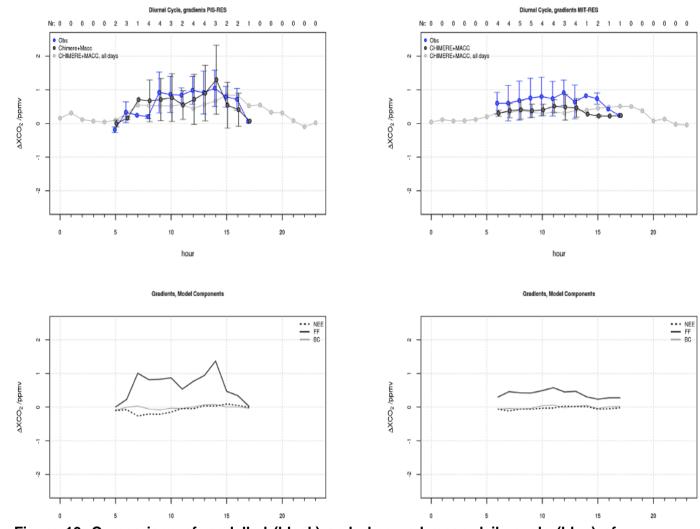


Figure 10. Comparison of modelled (black) and observed mean daily cycle (blue) of hourly averaged  $\Delta XCO_2$  of PIS (top left) and of MIT (top right) during the campaign when RES can be considered as upwind site. Labels on top of the upper figures denote the number of days contributing to the mean. The mean daily cycle for all days within the campaign period when PIS and MIT are downwind of RES is given in light grey The modelled contribution of different CO<sub>2</sub> sources/sinks to the mean daily cycle for days with observations for the two sites is given in the bottom panels.

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Instrument	XCO <sub>2</sub> factor Berlin	XCO <sub>2</sub> factor before	XCO <sub>2</sub> factor after	
		Paris	Paris	
1	1.0000 (0.0003)	1.0000 (0.0003)	1.0000 (0.0003)	
2	0.9992 (0.0003)	0.9991 (0.0003)	0.9992 (0.0003)	
3	1.0002 (0.0003)	1.0001 (0.0004)	1.0000 (0.0005)	
4	0.9999 (0.0003)	1.0000 (0.0004)	1.0000 (0.0004)	
5	0.9996 (0.0003)	0.9995 (0.0003)	0.9995 (0.0003)	

781 Table 1. Normalisation factors for the five EM27/SUN instruments derived during 782 measurements before and after the Paris field campaign. Values in parentheses are 783 standard deviations. Measurements of instrument 1 were arbitrarily chosen as reference from which the others were scaled. The calibration factors from a previous 784 785 field campaign in Berlin [Hase et al. 2015] are also shown. Calibration factors between

the two field campaigns agree well within 0.02 % (~0.08 ppm) for all instruments. 786

Date	No. of observations	Quality	Wind speed (m s <sup>-</sup> 1)	Wind direction
	MIT GIF PIS RES JUS		•	
28 Apr 2015 (Tu)	179 102 178 199 234	++	4	W
29 Apr 2015 (We)	110 124 0 161 53	+	5	SW-W
04 Mai 2015 (Mo)	194 85 96 163 83	+	6	S-SE
05 Mai 2015 (Tu)	77 27 85 185 92	+	8	S-SW
06 Mai 2015	81 88 87 139 0	+	8	SW
(We)				
07 Mai 2015 (Th)	169 313 252 286 238	+++	3	SW
09 Mai 2015 (Sa)	179 0 181 289 149	++	6	W
10 Mai 2015 (Su)	325 478 362 542 282	++++	3	S
11 Mai 2015	410 431 251 298 413	++++	3	SSW
(Mo)				
12 Mai 2015 (Tu)	324 222 230 326 203	+++	4	NNW
13 Mai 2015 (We)	159 18 182 28 56	+	4	NE

Table 2. Summary of all measurement days with the number of observations at each of
 the sites, Mitry Mory (MIT), Gif Sur Yvette (GIF), Piscop (PIS), Saulx-les-Chartreux
 (RES), Jussieu (JUS), the overall quality ranking of each day according to the number
 of available observations and temporal coverage (with classification from poor to great:
 +, ++, +++, ++++), and the ground-level wind speed and direction.