

1 **February 2017 extreme Saharan dust outbreak in the Iberian  
2 Peninsula: from lidar-derived optical properties to evaluation  
3 of forecast models**

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28

29 **Abstract**

30 An unprecedeted extreme Saharan dust event was registered in winter time from 20 to  
31 23 February 2017 over the Iberian Peninsula (IP). We report on aerosol optical

32 properties observed under this extreme dust outbreak through remote sensing (active  
33 and passive) techniques. For that, EARLINET (European Aerosol Research Lidar  
34 NETwork) lidar and AERONET (AErosol RObotic NETwork) Sun-photometer Cimel  
35 CE 318 measurements are used. The sites considered are: Barcelona (41.38°N, 2.17°E),  
36 Burjassot (39.51°N, 0.42°W), Cabo da Roca (38.78°N, 9.50°W), Évora (38.57°N,  
37 7.91°W), Granada (37.16°N, 3.61°W) and Madrid (40.45°N, 3.72°W).

38 In general, large aerosol optical depths (AOD) and low Ångström exponents (AE) are  
39 observed. An AOD of 2.0 at 675 nm is reached in several stations. Maximum values of  
40 AOD<sub>675</sub> of 2.5 are registered in Évora. During and around the peak of AOD<sub>675</sub>, AEs  
41 close to 0 are measured. With regard to vertically-resolved aerosol optical properties,  
42 particle backscatter coefficients as high as  $1.5 \cdot 10^{-5} \text{ m}^{-1} \text{ sr}^{-1}$  at 355 nm are recorded at  
43 every lidar stations. Mean lidar ratios are found in the range 40 - 55 sr at 355 nm and 34  
44 - 61 sr at 532 nm during the event inside the dust layer. Mean particle and volume  
45 depolarization ratios are found to be very consistent between lidar stations. They range  
46 0.19-0.31 and 0.12-0.26 respectively. The optical properties are also found very stable  
47 with height in the dust layer. Another remarkable aspect of the event is the limited  
48 height of the dust transport which is found between the ground and 5 km. Our  
49 vertically-resolved aerosol properties are also used to estimate the performances of two  
50 dust models, namely BSC-DREAM8b and NMMB/BSC-Dust, in order to evaluate their  
51 forecast skills in such intense dust outbreaks. We found that forecasts provided by the  
52 NMMB/BSC-Dust show a better agreement with observations than the ones from BSC-  
53 DREAM8b. The BSC-DREAM8b forecasts (24 h) present a large underestimation  
54 during the event. No clear degradation of the prognostics is appreciated in 24, 48, 72 h  
55 except for the Barcelona station.

57 **1 Introduction**

58 Mineral aerosols are usually originated over arid or semiarid regions as a consequence  
59 of continuous soil erosion produced by wind. The strong warming of desert areas during  
60 daytime produces vertical thermal turbulences that can reach altitudes of up to 5000 m,  
61 followed by periods of nocturnal stability (Santos et al. 2013). Massive resuspension of  
62 huge amounts of mineral aerosols are thus produced and can be transported long  
63 distances by different mechanisms. 40% of aerosol mass emitted into the troposphere is  
64 attributed to desert dust and it is considered as the second largest source of natural  
65 aerosols (Andreae 1995, Salvador et al. 2014). One of the main desert dust sources is  
66 the Sahara desert since it is responsible for more than half of the world atmospheric  
67 mineral dust (Prospero et al. 2002, Mahowald et al. 2005, Wagner et al. 2009, Salvador  
68 et al. 2016). Under specific synoptic meteorological situations, a large amount of  
69 Saharan dust is transported towards the Mediterranean basin (Lafontaine et al. 1990,  
70 Obregón et al. 2015, Cuevas et al. 2017).

71 Lately, the number of surveys which address the study of atmospheric mineral aerosols  
72 has been increased for several reasons. Firstly, from the climate change standpoint,  
73 mineral aerosols play an important role on atmospheric radiative budget through  
74 scattering and absorption of the incoming solar and outgoing infrared radiation, and  
75 acting as cloud condensation nuclei (Ansmann et al. 2005, Klein et al. 2010, IPCC  
76 2013). Currently, the large temporal and spatial variability is responsible for a high  
77 uncertainty degree in aerosol radiative forcing estimates (Boucher et al. 2013) (Forster  
78 et al. 2007). Furthermore, there is a lack of systematic statistical surveys during a long  
79 time period. Some of them, (Mona et al. 2006) (Salvador et al. 2013) (Pey et al. 2013),  
80 have indicated that the Mediterranean basin is affected by African dust outbreaks  
81 following a marked seasonal pattern. Summer prevalence has been detected in the

82 western side (Sicard et al. 2016), no seasonal trend has been observed in the central  
83 region and higher contributions of desert dust have been commonly produced in spring-  
84 early summer in the eastern side of this basin.

85 Winter is the season when these phenomena are less likely to occur across the whole  
86 Mediterranean basin (Querol et al. 2009). However, extreme dust outbreaks, as the one  
87 described in this paper or others that took place quite recently (Cazorla et al. 2017,  
88 Sorribas et al. 2017), occurred during the coldest season. According to the fifth IPCC  
89 (2013) report an extreme weather event can be defined as a rare phenomenon taking into  
90 account its historical statistical distribution for a particular place and/or time. Then 10th  
91 and 90th percentiles are usually considered as reference to define "rare". In the  
92 supplementary material: Fig. S1, S2, S3 and Table S1, percentile of  $AOD_{675}$  and  
93 Ångström exponent (440-870 nm) are presented having considered all data available at  
94 each station in the Iberian Peninsula. Along with it, data concerning this event is also  
95 represented in order to justify its extreme character. This is important to be highlighted  
96 as extreme weather events have been discussed and suggested to be connected to  
97 climate change. For instance some remaining questions concern whether or not such  
98 events take place earlier or later in the season or if their severity has been increased  
99 (World Meteorological Organization 2011).

100 What is more, it has been demonstrated that African dust is the main source contributing  
101 to the regional background levels of  $PM_{10}$  (particular matter with an aerodynamic  
102 diameter lower than 10  $\mu m$ ) across the Mediterranean (35-50% of  $PM_{10}$ ) with maximum  
103 contributions up to 80% of the total  $PM_{10}$  mass (Pey et al. 2013) . These sporadic but  
104 intense natural contributions of PM have been responsible of a high number of  
105 exceedances of the  $PM_{10}$  daily limit value (50  $\mu g/m^3$ , after the 2008/50/EC European  
106 Directive) as registered in different rural and urban monitoring sites across the

107 Mediterranean Basin (Querol et al. 2009, Salvador et al. 2013). Moreover, statistically  
108 significant evidences on the association between short-term exposure to desert dust and  
109 health outcomes have also been derived (Karanasiou et al., 2012).  $PM_{10}$  originating  
110 from the desert was positively associated with mortality and hospitalizations in 13  
111 Southern European cities for the period 2001-2010 (Stafoggia et al. 2016). A recent  
112 regional study carried out in Spain has associated  $PM_{10}$  levels with daily mortality  
113 during African dust outbreaks in most of the Spanish regions (Díaz et al. 2017).

114 In addition, massive aerosol emissions into the atmosphere can be an issue for aircraft  
115 operation. For instance, aircraft engines, that fly through atmospheres with significant  
116 mineral dust loads on a regular basis, usually undergo an accelerated aging, and as a  
117 result, an anticipated and unexpected overhaul and maintenance is required (Weinzierl  
118 et al. 2012). In addition, atmospheric mineral dust can cause a huge impact on aviation  
119 by reducing the visibility during the landing and takeoff of aircrafts (Weinzierl et al.  
120 2017).

121 For all these reasons, characterizing these events in detail is strictly necessary given the  
122 aforementioned implications on human society. In this article, we report on a record-  
123 breaking dust event that hit the Iberian Peninsula (IP) on 20 - 23 February 2017. The  
124 observational task has been carried out through remote sensing techniques at different  
125 sites located in the IP. Sun and sky scanning spectral radiometers and lidar  
126 measurements have provided observations concerning the spatial (vertical and  
127 horizontal) distribution of aerosol. In this sense, the lidar technique is indispensable  
128 since it can provide both temporally and vertically resolved dust layering structures. To  
129 give an idea of the magnitude of the extreme event it is noteworthy to state that the  
130 AOD was greater than 2 at 675 nm in several AERONET stations and for the most  
131 intense periods some lidar and sun-photometer retrievals could not be performed due to

132 high aerosol load, respectively, attenuating the lidar signal and blocking the sun. A  
133 previous work concerning such event at the IP found an AODs at 500 nm up to 1.5 in  
134 the south of Spain (Guerrero-Rascado et al. 2009). In this case, maximum values of  
135 particle backscatter coefficients ( $1.5 \cdot 10^{-5} \text{ m}^{-1} \text{ sr}^{-1}$  at 355 nm) were similar to those  
136 registered during this event, however it took place in September. Preissler et al. (2011)  
137 reported an aerosol optical thickness up to 2 in Portugal as a consequence of another  
138 extreme dust outbreak episode in April (Preissler et al. 2011).

139 Finally, having the capability to forecast such events is also very important. Comparison  
140 exercises between real and modeled data must be done in order to better comprehend  
141 extreme dust events but more importantly to provide accurate information to decision  
142 makers beforehand. Because of that, it has been checked if the results from dust models  
143 (BSC-DREAM8b and NMMB/BSC-Dust) are in agreement with observations as the  
144 relationship between certain meteorological patterns and extreme African dust events  
145 can provide useful information for human health, air traffic controllers, or to predict  
146 different climate change scenarios. However, dust models have proved to fail in certain  
147 occasions under extreme dust events (Mamouri et al. 2016) mainly because the scale  
148 used by models is not small enough to appreciate such phenomena.

149 The aim of this paper is to procure an overview of the available dust observations  
150 obtained from remote sensing techniques at different locations in the IP, to derive the  
151 aerosol optical property profiles from such observations and to compare them against  
152 the results computed from models. The paper is organized as follows. The instruments  
153 and methodology are briefly described in Sect. 2. Sect. 3 deals with the description of  
154 the synoptic situation and columnar aerosol optical properties from sun and sky spectral  
155 radiometers. In section 4, vertically-resolved optical properties are discussed. Section 5

156 presents the performance of the dust models. Finally, conclusions can be found in Sect.  
157 6.

158 **2 Instruments and methodology**

159 **2.1 AERONET CIMEL CE-318 Sun-photometers in the IP.**

160 The AErosol RObotic NETwork (AERONET) is a global ground-based network of  
161 sun/sky multi-wavelength CIMEL CE-318 sun-photometers that provides relatively  
162 long-term records of atmospheric columnar aerosol optical properties (Holben et al.  
163 1998). The CIMEL spectral sun-photometer measures the direct solar irradiances with a  
164 field of view of approximately  $1.2^\circ$  and the sky radiances (in the almucantar and  
165 principal plane scenarios), at several spectral channels (see table 1). The direct-sun  
166 measurements are used to obtain the spectral AOD, Ångström exponent at several  
167 wavelength pairs and precipitable water vapor, approximately every 15 min. The  
168 estimated AOD uncertainty (mainly due to the calibration) is between 0.01 and 0.02  
169 (Holben et al. 1998).

170 The sky radiance measurements can be inverted to estimate aerosol optical properties  
171 such as the size distribution, the percentage of spherical particles in the aerosol mixture,  
172 several microphysical parameters describing the total, fine and coarse aerosol modes  
173 and numerous spectral quantities: complex refractive index, single scattering albedo,  
174 phase function, asymmetry parameter, extinction and absorption optical depths. The  
175 aerosol properties retrieved are hence used for calculating the broad-band fluxes at the  
176 bottom and top of the atmosphere, the radiative forcing and forcing efficiencies are also  
177 provided. A detailed description of the version 2 AERONET inversion products is given  
178 by (Holben et al. 2001). Table 1 shows the six AERONET stations distributed in the IP  
179 that were considered in this study. In figure 2 these places are located on a map.

180

181

182 **Table 1 – Summary of the sites considered in the study, main characteristics of the**  
 183 **AERONET sun-photometers and EARLINET lidars used, and lidar measurement**  
 184 **time.**

Site	Long. (°)	Lat. (°)	Altitude (m a.s.l.)	AERONET Sun photometer channels for AOD (nm)	EARLINET Lidar channels (nm)			Lidar measurement time	
					Elastic	Raman	Vertical resol. (m)	Start time	Stop time
Barcelona	2.11° E	41.39° N	115	440, 675, 870, 1020	355, 532 total, 532 cross, 1064	387, 407, 607	3.75	08:11 UTC (23 Feb)	23:54 UTC (23 Feb)
Burjassot	0.42° W	39.51° N	60	340, 380, 440, 500, 675, 870, 1020, 1640	355 cross and parallel	387	15	-	-
Cabo da Roca	9.50° W	38.78° N	140	340, 380, 440, 500, 675, 870, 1020	-	-	-	-	-
Évora	7.91° W	38.57° N	293	340, 380, 440, 500, 675, 870, 1020	355, 532, 532 cross, 1064	387, 607	30	00:00 UTC (20 Feb)	23:59 UTC (23 Feb)
Granada	3.61° W	37.16° N	680	340, 380, 440, 500, 675, 870, 1020	355, 532 parallel, 532 cross, 1064	387, 407, 607	7.5	12:00h UTC (20 Feb)	18:00h UTC (20 Feb)
								19:00h UTC (20 Feb)	21:00h UTC (20 Feb)
								07:31h UTC (21 Feb)	14:21h UTC (21 Feb)
								07:31h UTC (22 Feb)	20:00h UTC (22 Feb)
Madrid	3.72° W	40.45° N	669	340, 380, 440, 500, 675, 870, 1020	355, 532, 1064	387, 407, 607	7.5 (elastic), 3.75 (Raman)	21:00h UTC (22 Feb)	23:36h UTC (22 Feb)
								05:00h UTC (23 Feb)	08:00h UTC (23 Feb)
								11:00h UTC (23 Feb)	11:52h UTC (23 Feb)

185

186 **2.2 EARLINET lidars in the IP**

187

188 The European Aerosol Research Lidar Network, EARLINET, aims at creating a  
 189 quantitative, comprehensive, and statistically significant database for the horizontal,  
 190 vertical, and temporal distribution of aerosols on a continental scale, providing the most

191 extensive collection of ground-based data for the aerosol vertical distribution over  
192 Europe (Pappalardo et al. 2014). In this work four Iberian EARLINET stations  
193 (Barcelona, Madrid, Évora and Granada) provided lidar data, all of them equipped with  
194 multi-wavelength lidars and some of them with depolarization capabilities (see Table 1).  
195 Burjassot lidar station was not available at this moment.

196 On a regular basis, the EARLINET protocol establishes that lidar measurements have to  
197 be carried out on Monday (at 14 UTC and at sunset) and on Thursday (sunset).  
198 However, under exceptional events, as the one described in this work, these stations  
199 perform additional measurements in order to register the phenomena as long as possible.  
200 Then, lidar signals were averaged over 30 or 60 minute periods in order to guarantee a  
201 proper signal-to-noise ratio throughout the vertical column. The criteria followed to  
202 choose such periods is based on the representation of the dust plume but also on the data  
203 availability at atmospheric levels where Rayleigh computation can be accomplished  
204 since the aerosol burden during this event was certainly high and produced a great  
205 radiation extinction, which hampered the Rayleigh retrieval. In this work, lidar  
206 measurements at each station were performed at the periods specified in Table 1.

207 Vertically resolved particle coefficients were derived by means of the Klett-Fernald  
208 algorithm (Klett 1981, Fernald 1984). This algorithm requires an assumption of the lidar  
209 ratio (LR), defined as the particle extinction ( $\alpha$ ) to particle backscatter ( $\beta$ ) coefficients  
210 ratio, and for mineral dust we have considered a value of 50 sr (Guerrero-Rascado et al.  
211 2008, Guerrero-Rascado et al. 2009, Muller et al. 2009, Muller et al. 2010, Preissler et  
212 al. 2011). If possible,  $\alpha$  and  $\beta$  coefficient profiles were retrieved independently  
213 (Ansmann et al. 1992), which in turn allow computing the vertically-resolved LR.  
214 Given the fact that the LR is an intensive parameter, it provides useful information for  
215 the analysis of aerosol optical properties. Another intensive variable is the Ångström

216 exponent (Ångström 1964). It is inversely related to the size of particles: the greater the  
217 exponent is, the smaller the particles are and vice versa (Amiridis et al. 2009). This is  
218 defined for the wavelength pair ( $\lambda_1$  and  $\lambda_2$ ) as:

$$\text{Å}_\alpha = -\frac{\log \left[ \frac{\alpha(\lambda_1)}{\alpha(\lambda_2)} \right]}{\log \left[ \frac{\lambda_1}{\lambda_2} \right]} \quad \square \quad (1)$$

219

220 Since extinction coefficients were not always available, Ångström exponent was only  
221 obtained for such cases. However, the three backscatter coefficients were always  
222 estimated, which allowed to retrieve, the backscatter-related Ångström exponent. For  
223 this reason this parameter is also estimated, and the relationship to the aerosol size is  
224 similar than the previous definition, although it is affected by other parameters such as  
225 refractive index so the relationship should not be straightforward. Last but not least,  
226 lidar systems equipped with depolarization channels procure relevant information about  
227 the aerosol type because backscatter signals related to the cross and parallel-polarized  
228 component varies depending on aerosol shape.

229 With regard to the errors associated to the measurements, we made use of the Monte-  
230 Carlo technique so as to estimate the uncertainties of the vertically-resolved backscatter  
231 and extinction coefficients. This technique is based on the random extraction of new  
232 lidar signals, each bin of which is considered a sample element of a given probability  
233 distribution with the experimentally observed mean value and standard deviation. The  
234 extracted lidar signals are then processed with the same algorithm to obtain a set of  
235 solutions from which the standard deviation is inferred as a function of height  
236 (Pappalardo et al. 2004).

237 2.3 Description of the models evaluated and methodology

238 The present analysis utilizes the operational 72-hour dust forecasts of the BSC-  
239 DREAM8b (Perez et al. 2006, Basart et al. 2012) and the NMMB/BSC-Dust (Perez et  
240 al. 2011) models (<http://www.bsc.es/ess/information/bsc-dust-daily-forecast>) for the  
241 period from 19 to 22 February 2017. Both models are developed and operated at the  
242 Barcelona Supercomputing Center (BSC). Table 2 summarizes the main parameters  
243 used in the configuration of the models.

244 **Table 2. Main parameters of the dust models used in this study.**

245

	BSC-DREAM8b	NMMB/BSC-Dust
Meteorological driver	Eta/NCEP	NMMB/NCEP
Model domain	North Africa-Middle East-Europe (25° W – 60° E and 0° – 65° N)	
Initial and boundary conditions	NCEP/GFS data (at 0.5° × 0.5° horizontal resolution) at 12 UT are used as initial conditions and boundary conditions at intervals of 6 hours	
Horizontal resolution	0.33° x 0.33°	
Vertical resolution	24 Eta-layers	40 $\sigma$ -hybrid layers
Time step	3h	
Dust size bins	8 (0.1–10 $\mu\text{m}$ )	
Radiation interactions	Yes	Yes
Dust initial condition	24 h forecast from the previous day's model run	

246

247 The modeled dust extinction values at 550 nm are directly compared with the observed  
248 particle extinction values at 532 nm because of the wavelength proximity and the low  
249 spectral extinction dependence of mineral dust (see Section 4). In order to have

250 continuous observations and to maximize their number, day and nighttime inversions of  
251 particle backscatter coefficients are used and converted to extinction by multiplying  
252 them by a constant lidar ratio of 50 sr. The vertical resolution of both dust models is  
253 much coarser than the lidar vertical resolution. In order to evaluate the models'  
254 capability to reproduce the vertical distribution of the dust extinction coefficient, the  
255 original lidar vertical resolution is downgraded to the resolution of the modeled profiles.

256 Given that the extinction value at a given height,  $h_i$ , of the models is the average  
257 extinction of the layer comprised between  $h_i - \frac{h_i - h_{i-1}}{2}$  and  $h_i + \frac{h_{i+1} - h_i}{2}$ , the extinction  
258 value of the lidar profile at height  $h_i$  is calculated as the mean value of the original lidar  
259 profile (at the lidar original vertical resolution) calculated in the exact same layer of  
260 each model. For the horizontal resolution, the lidar data can be considered as point  
261 observations, while the models represent uniform pixels of 0.33° resolution (~33 km).  
262 The temporal resolution is also different: while the models provide instantaneous  
263 profiles with a time steps of 3 hours, the lidar profiles are averaged over 30 or 60 min.  
264 Here we have compared each modeled profile at time  $t$  with a 30- (60-) min. averaged  
265 lidar-derived profile included in the interval  $[t - 30, t + 29 \text{ min.}]$  ( $[t - 60, t + 59 \text{ min.}]$ ).  
266 In case two consecutive measurements fulfil this criterion, the measurement which was  
267 running at time  $t$  is selected. The forecast skill analysis is performed in terms of two  
268 vertically integrated statistical indicators, namely the fractional bias ( $FB$ ), and the  
269 correlation coefficient ( $r$ ), as well as in terms of the center of mass (CoM).  $FB$  and  $r$   
270 are both calculated for the extinction coefficient. The fractional bias is a normalized  
271 measure of the mean bias and indicates only systematic errors, which lead to an  
272 under/overestimation of the estimated values. The linear correlation coefficient is a  
273 measure of the models' capability to reproduce the shape of the aerosol profile. The

274 vertical integration is made from the lowest pair of simultaneously available model and  
275 observed values up to 6 km. No lower limit was fixed because of the dust plume  
276 proximity to the ground surface. The upper limit was fixed to 6 km because nearly no  
277 dust was detected above that height. The CoM was approximated by the particle  
278 backscatter weighted altitude as defined in (Mona et al. 2006) who noted that this  
279 approximation “exactly coincides with the true center of mass if both composition and  
280 size distribution of the particles are constant with the altitude”.

281 In the following sections we evaluate the model performances for forecasts of 24 hours  
282 (Section 5.1) and then we compare these forecasts to longer ones of 48 and 72 h  
283 (Section 5.2) to see how the forecast skill behaves as the lead time increases. A forecast  
284 (or a lead time) of 24 h represents all forecasts in the range [0; 23h] since the model  
285 initialization. 48 and 72 h forecasts represent all forecasts in the range [24; 47h] and  
286 [48; 71h] since the model initialization, respectively.

### 287 **3 Synoptic situation and columnar properties**

#### 288 3.1 Synoptic situation

289 During the period from 20 to 23 February 2017, the synoptic situation in the IP was  
290 dominated by the influence of an anticyclone centered northwest from the Western  
291 coast, extending in ridge to South Central Europe and by the existence of a low pressure  
292 system, initially centered over Morocco, as illustrated in the ECMWF ERA5 reanalysis  
293 of the Geopotential height at 850 hPa at several hours (Fig.1). This low is very likely to  
294 be associated to Sharav cyclone (Alpert and Ziv, 1989). The plots presented in Figure 1  
295 also include the surface wind friction velocity ( $u_*$ ), which is a good indicator of  
296 possible dust emissions from deserts (Alfaro and Gomes, 2001; Darmenova et al., 2009  
297 and references therein). It is generally assumed that the dust flux from the surface

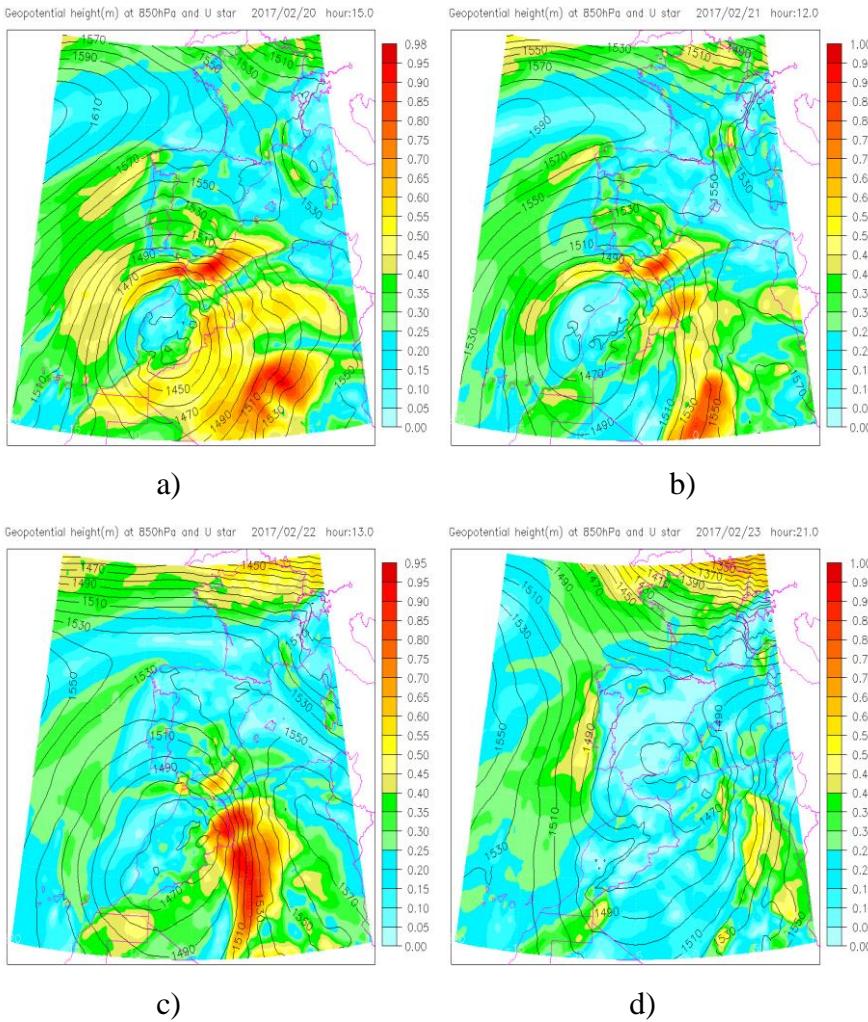
298 involves a power law of the wind friction velocity, as well as some parameters that  
299 characterize the surface, as the fraction of vegetation, the surface roughness and the soil  
300 texture and water content. Significant dust emissions are likely to occur for high friction  
301 velocities (above  $0.6\text{ms}^{-1}$ ), presenting lower sensitivity to land surface parameters  
302 (Darmenova et al., 2009).

303 The Geopotential field at 850 hPa (Fig.1) indicates the persistence of an atmospheric  
304 flow advecting air from the central North Africa (Algeria) crossing the IP. On 20  
305 February (Fig 1a) strong  $u_*$  values ( $> 6 \text{ ms}^{-1}$ ) represented over Algerian Sahara, a major  
306 dust source region (Ginoux et al., 2012) are suitable to force dust aerosol emissions  
307 (Darmenova et al., 2009). The well-shaped deep low, centered over central Morocco  
308 transported air from Algeria to southern Spain. Over the Central and Northern parts of  
309 the Peninsula, the dominant wind brings air from central Europe under the anticyclonic  
310 circulation. Wind vectors at 850 hPa are not represented in Fig.1 for clearness, though at  
311 this level it is reasonable to assume geostrophic wind. The situation maintains very  
312 similar in the next day and on the 22 February the low provokes high winds on the  
313 western side (central-northern Algeria), which may be seen by the proximity of the  
314 isopleths and by the strong values of  $u_*$  (Fig. 1c), which indicates strong dust  
315 emissions. On 23 February the northward shift of the Moroccan low originated weak  
316 precipitation events in several locations in the south of Portugal and Spain, but still  
317 transporting air from Algeria to Northeast Spain (~Catalonia). The  $u_*$  over the desert  
318 regions dropped significantly, hinting at the end of the significant dust emissions. The  
319 synoptic conditions changed sharply on 24 February with the passage of a frontal  
320 system that affected all the IP (not shown in Fig.1).

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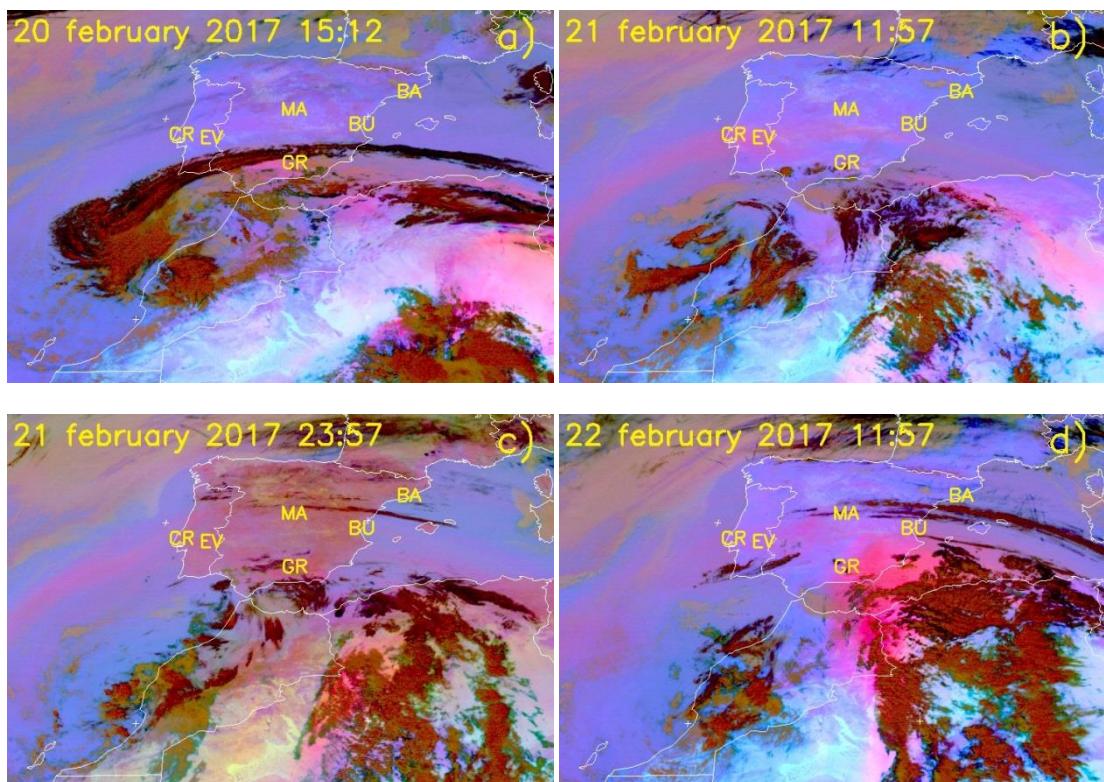
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327 **Fig. 1. European Centre for Medium-Range Weather Forecasts (ECMWF)**  
328 **reanalysis (ERA 5) of the Geopotential height at 850 hPa (black height contours)**  
329 **and surface wind friction velocity (color bar in  $\text{ms}^{-1}$ ) from 20 to 23 February 2017.**  
330 **Generated using Copernicus Atmosphere Monitoring Service information [2018].**

331 Fig. 2 presents RGB composites based upon the combination of infrared channels (8.7,  
332 10.8 and 12.0  $\mu\text{m}$ ) from the Spinning Enhanced Visible and InfraRed Imager (SEVIRI)  
333 on board Meteosat-10, showing the dust transport evolution (magenta) from 20 to 24  
334 February 2017. The dust was transported across the Alboran Sea (western  
335 Mediterranean Sea) and infiltrated in southern Iberian atmosphere on 20 February  
336 (Fig.2a), gradually transported towards west and north by the synoptic circulation,

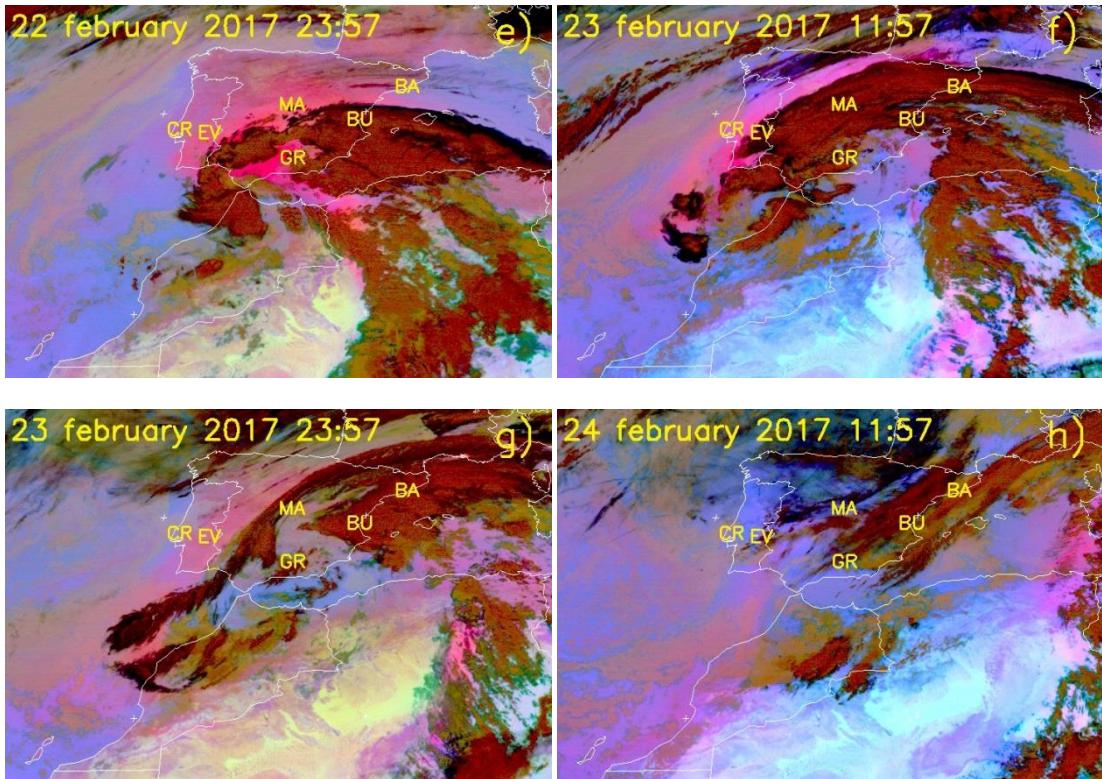
337 affecting the southern and western sites (CR, EV, GR) as illustrated by Figs.2b and 2c.  
338 On the 22 February the dust intrusion was reinforced by a thick plume that  
339 progressively entered the IP through the southeastern coast (Fig. 2d) extending north  
340 and westwards and affecting all sites represented in the images (Fig. 2e). This new  
341 intrusion was accompanied by the presence of high clouds that on the 23 February  
342 affected most of the IP, associated with the intensification and northward shift of the  
343 Moroccan low (Figs.2f and 2g). The arrival of a frontal system from northwest on the  
344 24 February interrupted the North African dust flow, pushing it towards the central  
345 Mediterranean regions (Fig. 2h).

346



347

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349

350

351 **Fig. 2. Meteosat RGB composites showing the evolution of the dust plume from 20**

352 **to 24 February 2017. The Iberian sites considered in the study are also represented**

353 **in the images: Barcelona (BA), Burjassot (BU), Cabo da Roca (CR), Évora (EV),**

354 **Granada (GR) and Madrid (MA).**

355 The temporal evolution of the back-trajectories, from 20 to 24 February 2017, arriving

356 over the six sites considered, at three atmospheric levels (2000, 3000 and 4000 m a.g.l.)

357 is represented in the supplementary material in Fig. S4. The back-trajectories were

358 calculated using the Hybrid Single-Particle Lagrangian Integrated Trajectory

359 (HYSPLIT) model (Stein et al., 2015; Rolph et al., 2017), available online at

360 <http://ready.arl.noaa.gov/HYSPLIT.php>. The sequence shows that the first sites

361 overpassed by air masses originating in northern Africa were: Granada (20 February;

362 Fig. S4a), followed by Évora and Cabo da Roca (21 February; Fig. S4b and c).

363 Burjassot and Madrid sites started to be influenced by North African air masses between

364 the 21 and 22 February (Fig. S4d and e) and finally also Barcelona remained under the

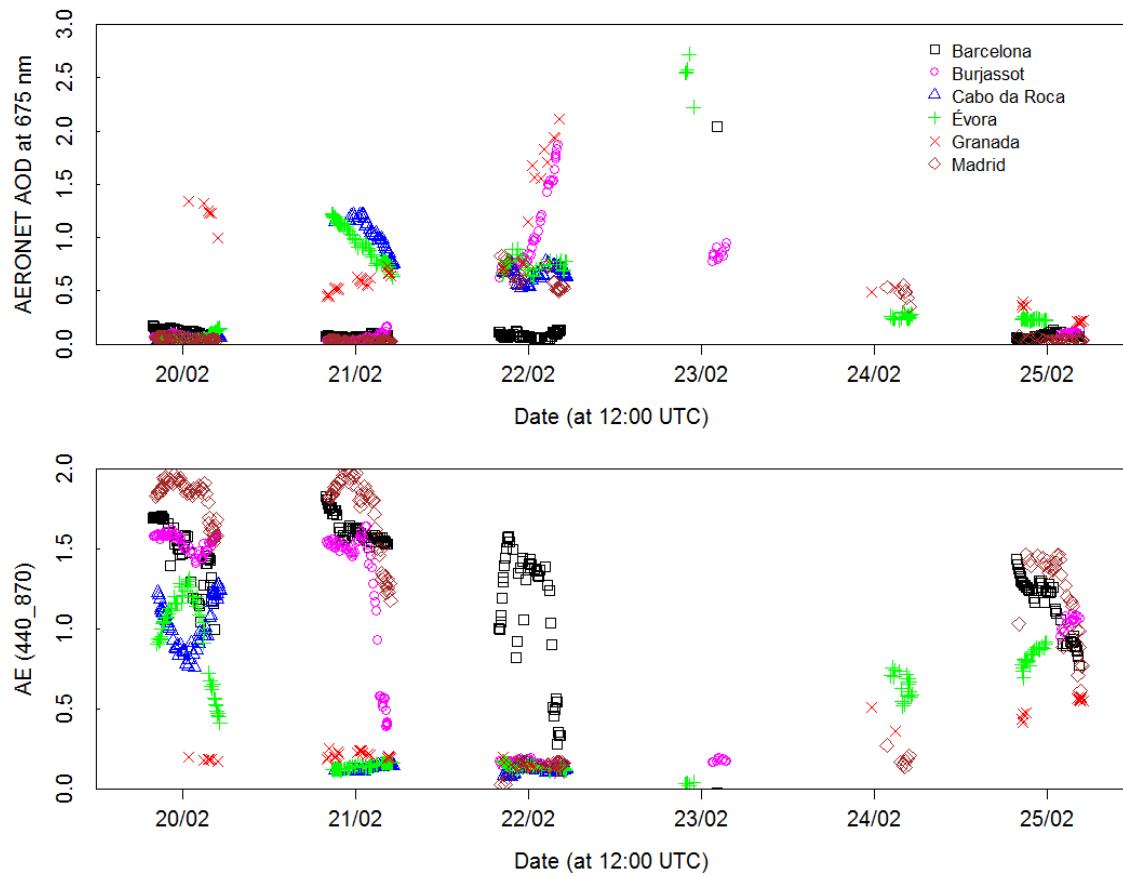
365 influence of the same air masses between the 23 and 24 February (Fig. S4f to and h).  
366 Information from Meteosat RGB composites (Fig.2) displaying the dust distribution  
367 over North African regions and back-trajectories (Fig. S4), hint at dust originating from  
368 central Algeria, which is a recognized major dust source region (Ginoux et al., 2012).  
369 This is also in agreement with the strong values of wind friction velocities found over  
370 the same region and shown in Fig. 1.

371

### 372 3.2 Columnar properties

373 The desert dust plume entered the IP from the South on the 20 February, and then it  
374 gradually reached the northwest and later on the eastern part of the IP. Fig. 3 shows the  
375 time series data of AOD at 675 nm and Ångström exponent (440 and 870 nm), from 20  
376 to 25 February 2017 in six sites distributed across the IP. An increase of the AOD was  
377 first noticed in Granada site on the 20 February, where the AOD values reach about 1.5,  
378 accompanied by very low values of AE, typical of desert dust intrusions, which is  
379 confirmed by the Meteosat composite in Fig. 2a. The dust plume maintains its influence  
380 over Granada and extends towards the western part of IP, affecting in the next day also  
381 Évora and Cabo da Roca sites, with AOD values ranging between about 0.6 and 1.2,  
382 once again with very low AE (<0.2). The dust transport continues and on the 22  
383 February, during daytime, desert dust is detected in all stations except for Barcelona  
384 where it is measured in the next day. Still on the 22 February, extremely high AOD  
385 values are reached in Granada and Burjassot (> 2.0) and moderately high in Madrid,  
386 Évora and Cabo da Roca (0.5<AOD<1.0), with AE values lower than 0.2 for all these  
387 stations. On the 23 February there are only a few AERONET measurements available  
388 due to the persistence of clouds over the region, nevertheless the AOD is still  
389 considerably high (>2.0) for Évora and Barcelona, with corresponding AE values

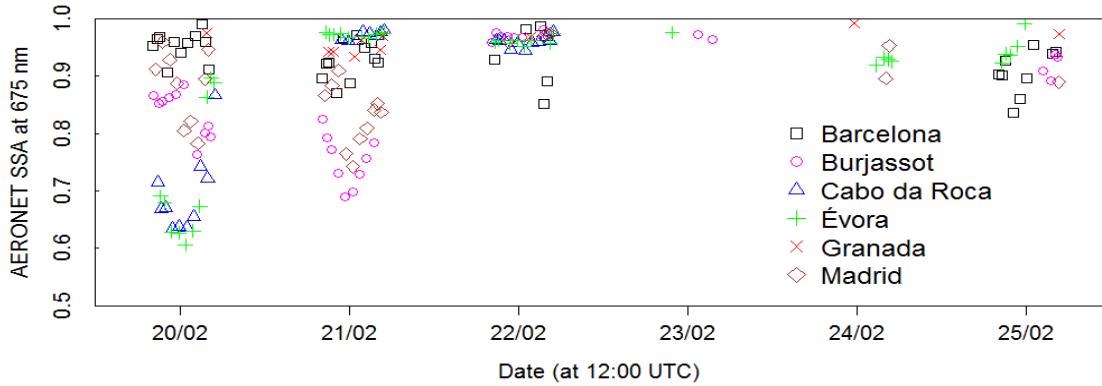
390 around zero in these sites, with the provenience of air masses from desert dust source  
391 regions supported by the back-trajectories presented in the supplementary material (Fig.  
392 S4). As mentioned before, the frontal system on the 24 February interrupted the dust  
393 transport and the AOD values on the 24 and 25 February show a consistent decrease  
394 with a corresponding increase of the AE.



395  
396

397 **Fig. 3. – AERONET AOD at 675 nm and AE (440 and 870 nm) from 20 to 25**  
398 **February 2017 in six sites distributed across the IP.**

399



400

401 **Fig. 4 - AERONET SSA at 675 nm from 20 to 25 February 2017 during the event**  
 402 **for six sites distributed across the IP.**

403 The single scattering albedo is characterized by relatively high values in all the stations  
 404 during the dust event, showing the predominant dispersive nature of these particles. The  
 405 lower SSA values in the first two days (greater absorption) in some of the sites (BU,  
 406 CR, EV, MA) depicted in Fig.4, are related with polluted air masses coming from  
 407 northwestern Europe (not shown here).

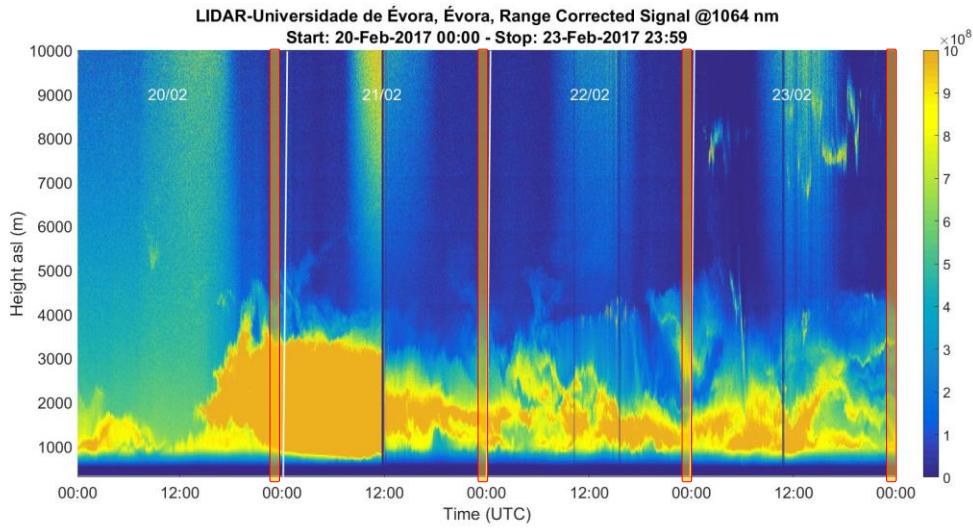
408 **4. Vertically-resolved optical properties**

409

410 **ÉVORA**

411 Fig. 5 represents the Range Corrected Signal (RCS) during 4 days, 24 hours per day,  
 412 which provides a very detailed overview of the phenomenon. It can be seen that the  
 413 African dust outbreak was especially intense at the beginning of the event, from 20  
 414 (12:00 UTC) to 21 (12:00 UTC) February. However, it must be noted that on 21 (12:00  
 415 UTC) February a change of the neutral-density filters in front of the detection channel  
 416 was necessary to be carried out in order to attenuate the received light. This obviously  
 417 reduced the RCS at this point but did not affect the retrieval of aerosol optical  
 418 properties. Four different periods have been selected so as to analyze aerosol optical

419 properties from the African plume observed in Évora (highlighted again in red in Fig.  
420 5). Nighttime measurements have been chosen for the analysis in order to estimate  
421 accurately such properties given the fact that independent extinction from Raman  
422 signals was available at this lidar station. The first period (21<sup>st</sup> Feb from 0:00-0:30  
423 UTC), presents the highest backscatter coefficient values out of all periods evaluated, so  
424 a special attention has been paid to this period (Fig. 6). Notwithstanding the other 3  
425 periods are also analyzed and they can be seen in the supplementary material Fig. S5,  
426 S6 and S7. Mean aerosol optical properties are exposed in this latter Table (3) for  
427 specific atmospheric layers where in principle the dust plume is representative. For  
428 instance, the first period analyzed presents an African dust plume that reaches also 5 km  
429 height asl, however maximum values of particle backscatter coefficient are reached at  
430 3222 m asl and from 4 to 5 km asl the presence of African dust is very small according  
431 to particle backscatter coefficient profiles. For this reason, it is considered more  
432 appropriate to evaluate the atmospheric layer detected between 1.5-3.5 km asl. At this  
433 atmospheric layer, backscatter-related Ångström exponent at the wavelength pairs:  
434 532/355, 1064/532 and 1064/355 were found to be  $0.08 \pm 0.33$ ,  $0.62 \pm 0.04$  and  $0.42 \pm 0.13$   
435 respectively and the extinction-related Ångström exponent at 532/355 nm was estimated  
436 to be  $0.16 \pm 0.45$ . These small values are typical for dust as previously reported during  
437 extreme African dust outbreaks (Mamouri et al. 2016) (Guerrero-Rascado et al. 2009,  
438 Preissler et al. 2011). The other periods also show relatively low backscatter-related  
439 Ångström exponents and Ångström exponent values, which in principle indicates a  
440 large particle size.



441

442 **Fig. 5. RCS at 1064 nm on 20-23 February 2017 for the period established between**

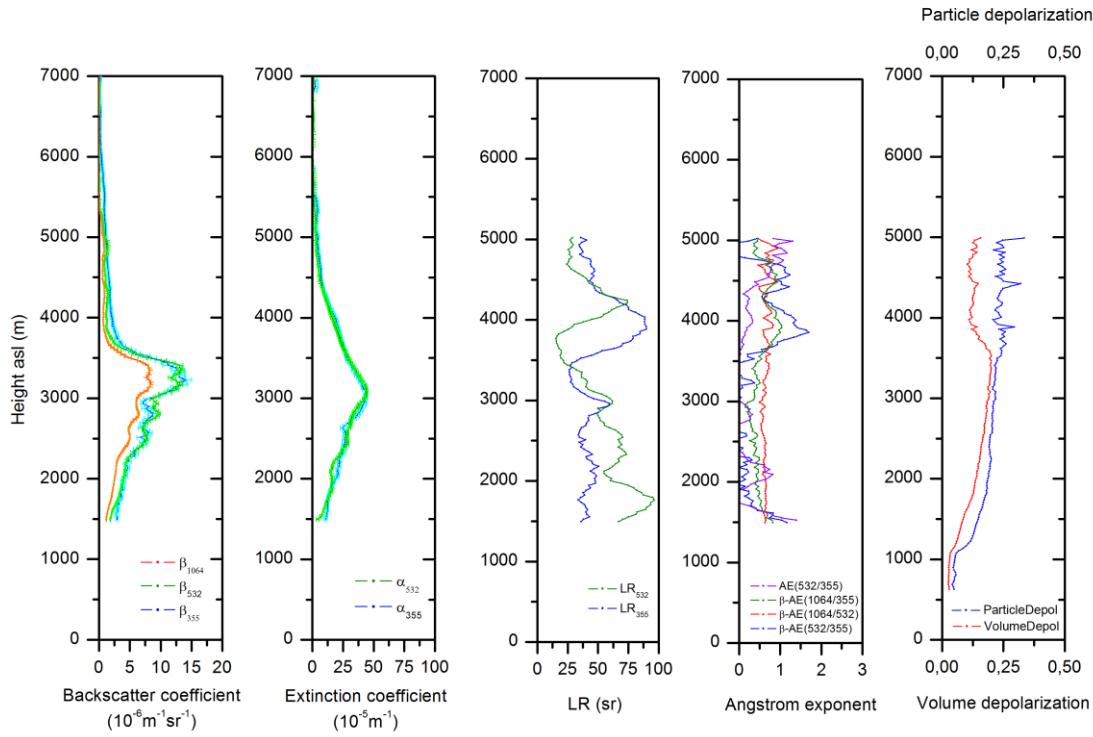
443 **0:00h-23:59 UTC respectively (Évora, 293 m asl).**

444 Since Raman signals were available and extinction coefficients were obtained  
 445 independently, particle lidar ratios were derived as well. The dust layer located between  
 446 1.5-3.5 km asl on 21Feb (00:00 UTC) presented a lidar ratio of  $40\pm8$  sr and  $61\pm18$  sr at  
 447 355 and 532 nm, respectively. Our estimates at 355 nm are in agreement with Mona et  
 448 al. that found a mean lidar ratio at 355 nm of  $38\pm15$  sr for three years of Raman lidar  
 449 measurements of Saharan dust (Mona et al. 2006). On the other hand, lidar ratio at 532  
 450 nm is found greater than the lidar ratio at 355 nm for the first period analyzed (21 Feb,  
 451 00:00 UTC), which is not usual for dust particles as it has been already pointed out by  
 452 other authors (Muller et al. 2010). Nevertheless, this trend is only observed in the first  
 453 period analyzed, the other three analyzed periods show a lidar ratio at 532 lower than  
 454 the lidar ratio at 355 nm. The reason behind this observation (high unexpected lidar  
 455 ratio values at 532 nm) can be attributed to non-accurate retrievals handicapped by the  
 456 high aerosol load, which produces great extinction and consequently a scarce lidar  
 457 signal to be evaluated. It is noteworthy to mention that the standard deviation of the  
 458 mean lidar ratio at 532 nm on 21Feb (00:00 UTC) is significantly higher compared to

459 the rest of the studied period. On another note, the lidar ratio at 355 nm on 23 Feb (at  
 460 00:00 and 23:39 UTC) seems a bit higher than values reported in literature (Mona et al.  
 461 2006) and it could be due to a decrease of the African dust outbreak intensity and  
 462 therefore a greater proportion of local aerosol might be present in the atmosphere. The  
 463 lidar ratio at 532 nm in all cases (apart from the first period) are consistent with  
 464 literature since typical values range 35-45 sr for typical desert dust (Mamouri et al.  
 465 2013, Nisantzi et al. 2015, Mamouri et al. 2016). In addition, the particle and volume  
 466 depolarization ratio were  $0.19 \pm 0.02$  and  $0.16 \pm 0.03$  for the aforementioned atmospheric  
 467 layer on 21Feb 00:00 UTC. These two latter parameters are constant with altitude,  
 468 which indicates that no changes in the aerosol type is observed within the atmospheric  
 469 layer of interest. They are also very similar for the four periods studied, however the last  
 470 period of study indicates lower particle and volume depolarization values that is  
 471 associated with the decrease of intensity of the Saharan dust outbreak and a greater  
 472 contribution of local aerosols.

473 **Table 3. Summary of mean aerosol optical properties retrieved for the 4 periods**  
 474 **analyzed from Raman lidar measurements (Évora).**

Atmospheric layer	LR <sub>355</sub> (sr)	LR <sub>532</sub> (sr)	β-AE 1064-532	β-AE 532-355	β-AE 1064-355	AE 532-355	δ-vol.	δ-part.
00:00 UTC-21Feb 1.5-3.5km asl	$40 \pm 8$	$61 \pm 18$	$0.62 \pm 0.04$	$0.08 \pm 0.33$	$0.42 \pm 0.13$	$0.16 \pm 0.45$	$0.16 \pm 0.03$	$0.19 \pm 0.02$
00:00 UTC-22Feb 1.5-4km asl	$45 \pm 4$	$38 \pm 8$	$0.76 \pm 0.12$	$-0.12 \pm 0.23$	$0.44 \pm 0.08$	$0.16 \pm 0.19$	$0.16 \pm 0.01$	$0.21 \pm 0.01$
00:00 UTC-23Feb 1.5-5km asl	$52 \pm 7$	$40 \pm 9$	$1.28 \pm 0.33$	$-0.62 \pm 0.48$	$0.58 \pm 0.19$	$0.01 \pm 0.27$	$0.16 \pm 0.02$	$0.19 \pm 0.01$
23:39 UTC-23Feb 1.5-4.5km asl	$55 \pm 12$	$34 \pm 8$	$1.00 \pm 0.18$	$-0.96 \pm 0.29$	$0.28 \pm 0.17$	$0.18 \pm 0.24$	$0.12 \pm 0.01$	$0.15 \pm 0.01$



476

477 **Fig. 6. Backscatter coefficient, extinction coefficient, Lidar ratio, Ångström**  
 478 **exponents, and particle and volume depolarization profiles at 00:00 UTC on 21,**  
 479 **February 2017 at Évora.**

## 480 GRANADA

481 In Granada, four lidar measurements were carried out during the extreme African dust  
 482 outbreak. In particular for the periods: 12:00-18:00 and 19:00-21:00 UTC on 20  
 483 February, 07:31-14:21 UTC on 21 February, and 07:31-20:00 UTC on 22 February.  
 484 Such measurements are represented in Fig. 7. The red highlights indicate as previously  
 485 the selected periods where vertically-resolved aerosol optical properties have been  
 486 derived. Such vertical profiles can be seen in the supplementary material in Fig. S7, S9,  
 487 S10 and S11. For a better comprehension of these data, mean aerosol optical properties  
 488 are presented in table 4 for the periods highlighted in red and for the atmospheric layer  
 489 where the dust plume is registered. In general terms, the maximum altitude of the dust

490 plume was registered at 4 km asl approximately and it was maintained relatively  
491 constant throughout the four lidar measurements. For certain periods (13:30-14:21 UTC  
492 on 21st Feb) intensification of the RCS is observed at the top of the dust plume, which  
493 may indicate cloud formation processes related to mineral dust.

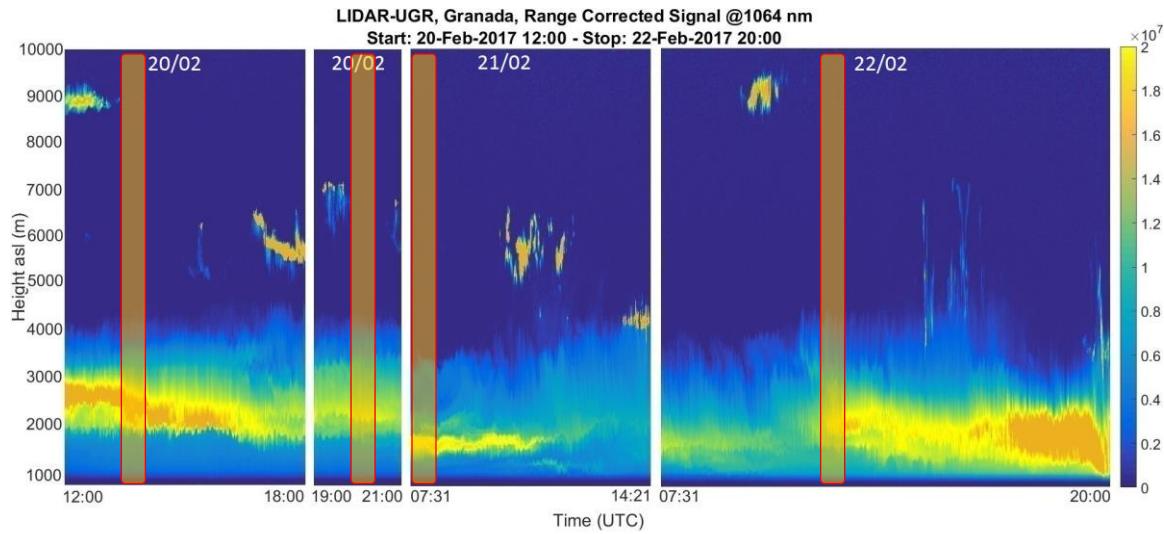
494 Concerning intensive aerosol optical properties, backscatter-related and extinction-  
495 related Ångström exponents were low, in accordance with previous lidar observations,  
496 which indicate a large aerosol size. The Raman retrieval could be performed only for  
497 the period 19:00-21:00 UTC on 20 February since it was not possible to perform during  
498 nighttime on other days. On 22 February, the African dust outbreak was so intense that  
499 it produced large extinction and hampered proper retrieval. So, lidar ratios obtained at  
500 Granada were  $52\pm7$  and  $53\pm6$  at 355 and 532 nm respectively. With regard to particle  
501 and volume depolarization ratios, these parameters show similar and consistent values  
502 to data obtained in the Évora station. Nevertheless, it is noteworthy to mention that the  
503 last analyzed period (12:30 UTC on 22nd Feb) exhibits the greatest particle and volume  
504 depolarization ratios observed in all lidar stations. These high values point out that a  
505 large backscatter signal related to the cross-polarized component is registered, which in  
506 turn is produced by non-spherical particles. This is associated to an enlargement on the  
507 contribution of mineral dust due to the reinforcement of the dust plume coming from  
508 Africa. Such reinforcement of the dust plume was observed on 22 Feb according to the  
509 synoptic meteorological situation (see section 3). In fact, it was not possible to retrieve  
510 proper lidar products for measurements carried out on 22 Feb from 17:30 UTC on,  
511 given the large extinction of radiation produced by the high contribution of mineral  
512 dust.

513 **Table 4. Summary of mean aerosol optical properties retrieved for the 4 periods**  
514 **analyzed from Raman lidar measurements (Granada).**

Atmospheric layer	LR <sub>355</sub> (sr)	LR <sub>532</sub> (sr)	β-AE 1064-532	β-AE 532-355	β-AE 1064-355	AE 532-355	δ-vol.	δ-part.
13:30 UTC-20Feb 2.0-4.0 km asl			0.27±0.12	0.19±0.30	0.24±0.04		0.19±0.03	0.22±0.04
20:00 UTC-20Feb 1.8-4.0 km asl	52±7	53±6	0.19±0.08	0.54±0.21	0.32±0.07	0.51±0.43	0.20±0.02	0.25±0.03
07:31 UTC-21Feb 1.5-3.4km asl			0.86±0.07	0.64±0.13	0.77±0.08		0.18±0.03	0.28±0.01
12:30 UTC-22Feb 1.5-4.0 km asl			0.39±0.12	0.32±0.17	0.36±0.07		0.26±0.01	0.31±0.02

515

516



517

518 **Fig. 7. RCS at 1064 nm on 20 February (12:00-18:00, 19:00-21:00 UTC), 21**  
 519 **February (07:31-14:21 UTC), 22 February (07:31-20:00 UTC) 2017 at Granada**  
 520 **(680 m asl).**

521 **MADRID**

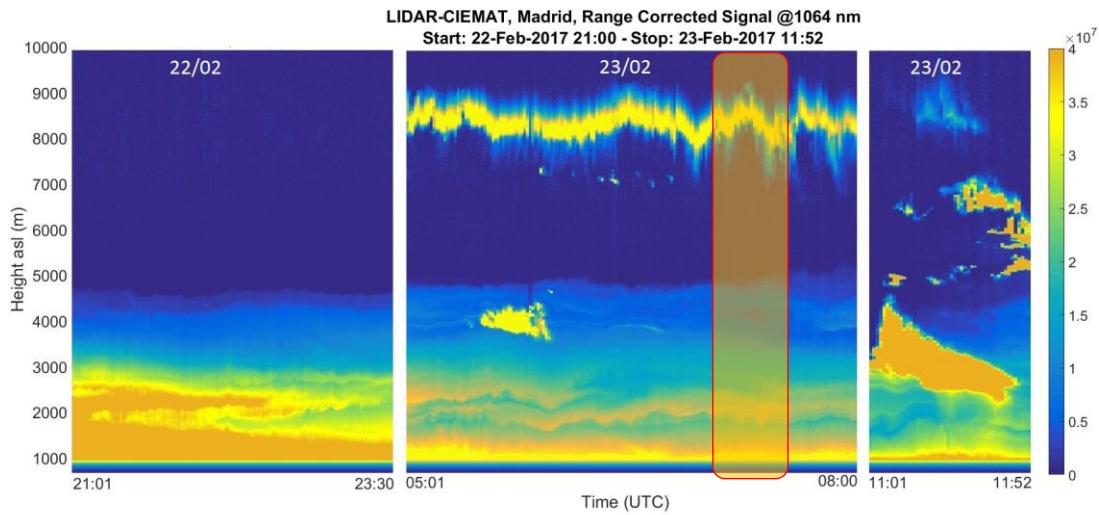
522 In Madrid, as it occurred in Barcelona, the African dust plume was only detected in the  
 523 last stage of the African event when the reinforcement of the dust intrusion was  
 524 produced by synoptic flows (from 22 February on). Lidar measurements on 20 February

525 (not shown) at Madrid still did no present any sign of this extraordinary plume. During  
526 this African event, three lidar measurements were available at this station: on 22 Feb  
527 (21:00-23:36 UTC) and 23 Feb (05:00-08:00 and 11:00-11:52 UTC). They are  
528 represented in Fig. 8. As it can be seen the thickness of the plume ranged from the  
529 ground to 5 km asl and in the last lidar measurement the plume was accompanied by  
530 thick clouds. Concerning the retrieval of vertically-resolved aerosol optical properties,  
531 only the period 05:00-08:00 UTC (23 Feb) was considered for this purpose. Such  
532 profiles are represented in Fig. 9, which concerns the period 06:59-07:29 UTC  
533 highlighted in Fig 8. Only one profile is presented given the fact that the extinction  
534 observed on the first and third lidar measurement was again excessive at low altitudes  
535 due to the dust plume, so Rayleigh extinction could not be appropriately computed. This  
536 is a problem we want to highlight as it appeared in several lidar stations when  
537 addressing this study and performing the retrievals under such extreme conditions (high  
538 aerosol load).

539 Finally, Fig. 9 presents 3 backscatter coefficient profiles at 1064, 532 and 355 nm and  
540 their respective backscatter-related Ångström exponents. No particle extinction  
541 coefficients could be obtained independently as Raman signal were too noisy due to the  
542 aforementioned reasons. Maximum values of particle backscatter coefficient are reached  
543 at 2200-2300 m asl. At this altitude  $\beta_{355}$  is  $(6.85\pm0.09)\cdot10^{-6}$ ,  $\beta_{532}$  is  $(6.35\pm0.13)\cdot10^{-6}$  and  
544  $\beta_{1064}$  is  $(5.75\pm0.01)\cdot10^{-6} \text{ m}^{-1}\text{sr}^{-1}$ . Mean backscatter-related Ångström exponents were  
545 found to be  $0.52\pm0.34$ ,  $0.28\pm0.17$ ,  $0.37\pm0.22$  at the wavelength pairs: 532/355,  
546 1064/532 and 1064/355 nm for the atmospheric layer established from lidar full overlap  
547 height to 4900 m. These low backscatter-related Ångström exponents are in accordance  
548 with previous lidar observations, which partially indicate a large aerosol size.

549

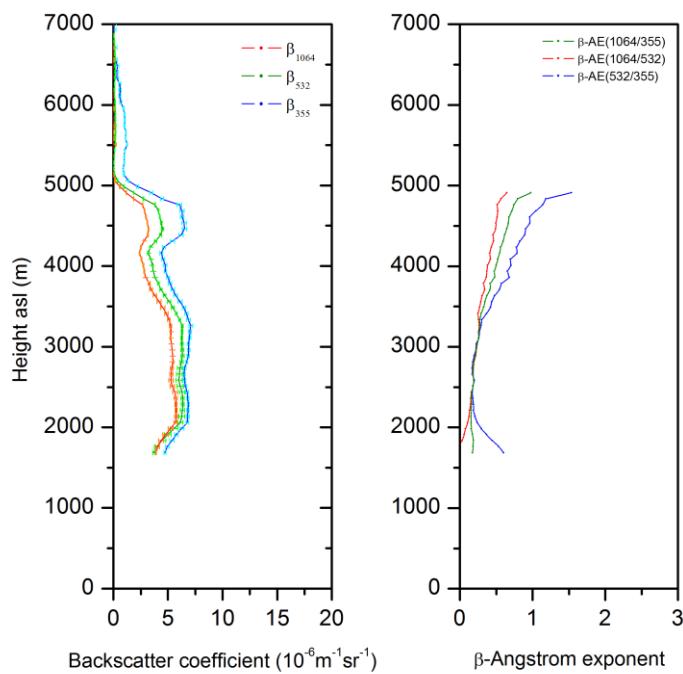
550



551

552 **Fig. 8. RCS at 1064 nm on 22 February (21:00-23:36), 23 February (05:01-08:00**  
553 **UTC), 23 February (11:00-11:52 UTC) 2017 at Madrid (669 m asl)**

554



555

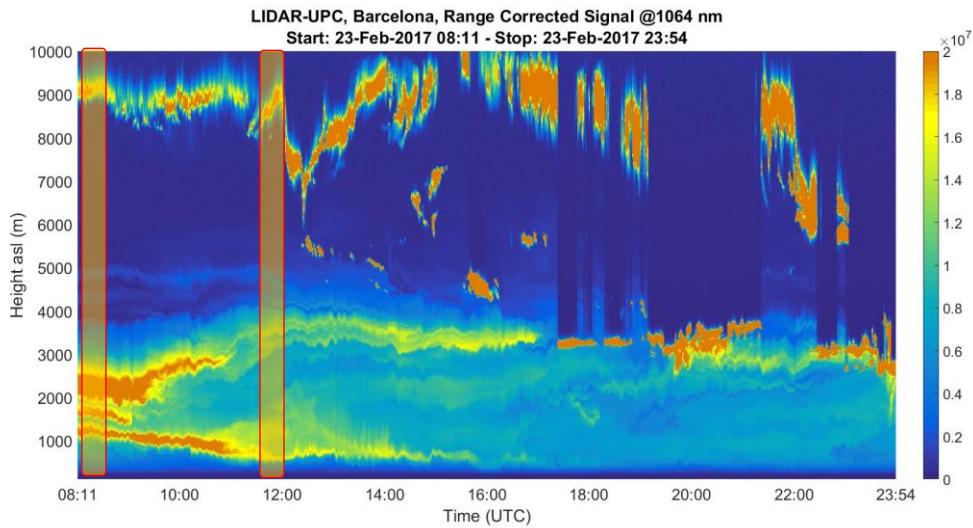
556 **Fig. 9. Backscatter coefficient and  $\beta$ -Ångström exponent profiles at 06:59 UTC on**  
557 **23 February 2017 at Madrid.**

558 **BARCELONA**

559

560 According to the meteorological overview, the Barcelona site was the latest place from  
561 the time standpoint that was hit by the extreme African dust outbreak. As it can be seen  
562 in Fig. 10 the African dust plume was registered throughout almost the entire 23  
563 February. At the beginning of the lidar measurement (from 08:11 to 12:00 UTC), the  
564 maximum altitude of the plume was detected at 5km asl approximately and after that it  
565 decreased gradually until it reached the value of 3-3.5 km at 23:54 UTC. Two periods of  
566 30 minutes have been selected (at 08:11 and 11:34 UTC) to show the retrieval of  
567 aerosol optical properties from the lidar measurement. The reason to choose these  
568 periods lies on the availability of a clear atmosphere to derive Rayleigh extinction. Both  
569 of them are highlighted in red on Fig 10. As indicated in the color bar, the RCS was  
570 considerably high for the atmospheric layer between 1 and 3 km during the period  
571 08:11-08:41 UTC. This is one of the reasons why this period of study was selected since  
572 in principle this variable is a proxy of the intensity of the African dust outbreak. The  
573 second period to be studied spans from 11:34 to 12:04 UTC. In this case, the dust plume  
574 is observed up to 5 km asl, although the structure is a bit different and the RCS is lower  
575 than in the first period. It must also be noted that from 12:00 UTC on the aerosol optical  
576 properties retrieval is quite complex since it is quite difficult to detect a clean  
577 atmospheric layer so as to derive the Rayleigh extinction, which is mandatory to infer  
578 the aforementioned aerosol optical properties. For the period 12:00-16:00 UTC  
579 dispersed clouds can be observed at 5-7 km and from 17:00-18:00 UTC on clouds are  
580 registered at the top of the dust plume layer (at 4 km), which prevents the Rayleigh

581 extinction computation. This latter observation is also interesting from the point of view  
 582 of cloud formation processes. Considering the evolution of the plume throughout the  
 583 entire lidar measurement at 4 km, it is plausible that African dust aerosol might act as  
 584 cloud nuclei (see RCS at 4 km from 18:00 to 23:54 UTC, the variable becomes more  
 585 intense than previously).

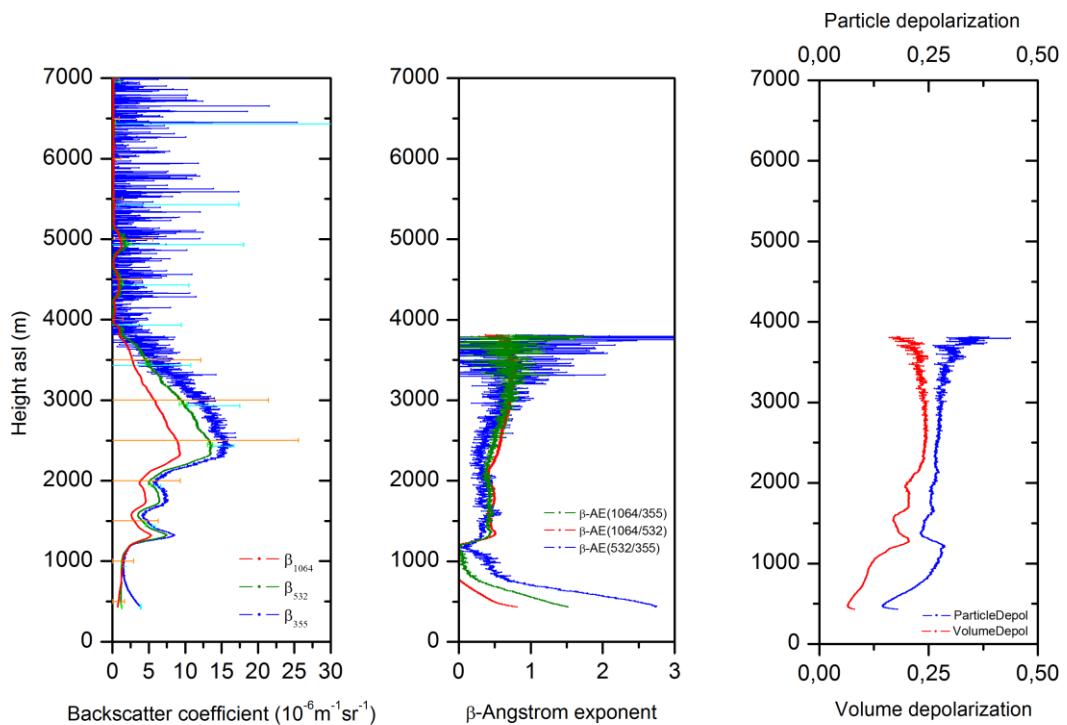


586

587 **Fig. 10. Range corrected signal (RCS) at 1064 nm on 23 February 2017 for the**  
 588 **period established between 08:11-23:54 UTC (Barcelona, 115 m asl).**

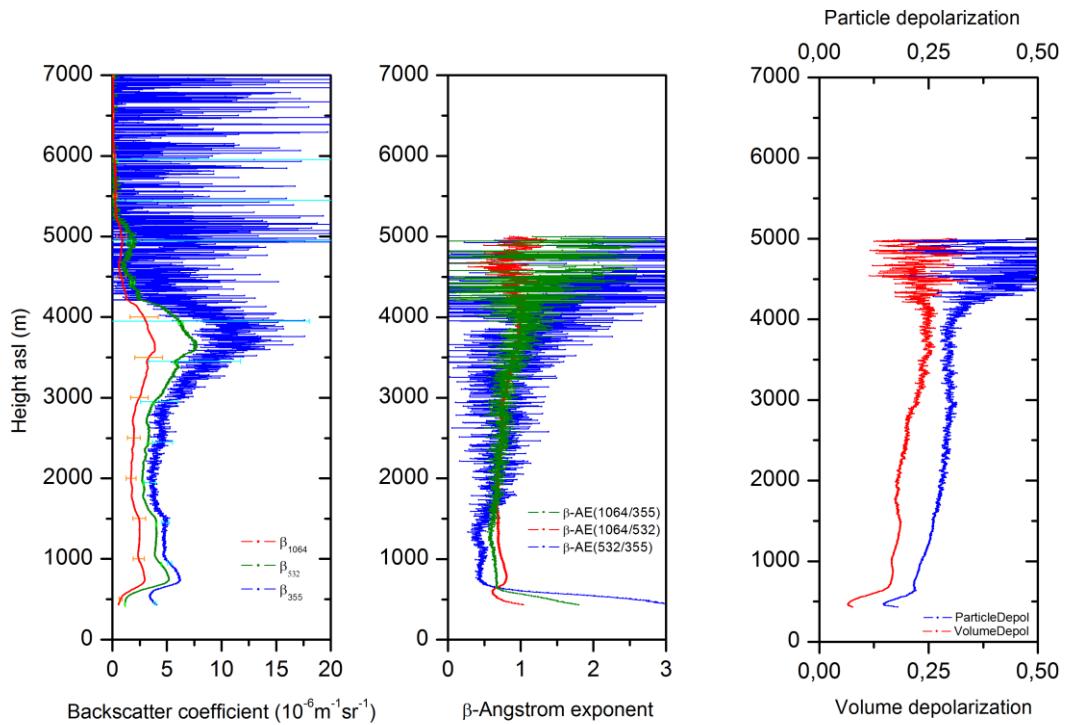
589 Fig. 11 shows aerosol optical properties obtained for the period 08:11-8:41 UTC. The  
 590 left panel represents the vertical profiles of particle backscatter coefficient at the three  
 591 wavelengths. The maximum values of this variable are reached at 2337 m asl. At this  
 592 altitude  $\beta_{355}$  is  $(1.53 \pm 0.14) \cdot 10^{-5}$ ,  $\beta_{532}$  is  $(1.35 \pm 0.04) \cdot 10^{-5}$  and  $\beta_{1064}$  is  $(0.9 \pm 1.6) \cdot 10^{-5} \text{ m}^{-1} \text{sr}^{-1}$ . The mean backscatter-related Ångström exponents are  $0.37 \pm 0.14$ ,  $0.45 \pm 0.22$ ,  
 594  $0.42 \pm 0.17$  respectively at the wavelength pairs: 532/355, 1064/532, 1064/355 for the  
 595 altitude range 1-3km asl. In general terms, the greater the aerosol size the lower the  
 596 Ångström exponent. In this case the variable used is the backscatter-related Ångström  
 597 exponent, which is similar to the previous one (Ångström exponent), so the relation is  
 598 affected by other parameters such as refractive index, etc. other than the aerosol size.

599 Nevertheless, these values are typical for African dust (Guerrero-Rascado et al. 2009),  
600 where aerosol size plays an important role on this parameter. It is noteworthy to  
601 mention that the vertical profile of the backscatter-related Ångström exponent is  
602 relatively constant through the atmospheric layer detected between 1-3 km asl. With  
603 regard to volume and particle depolarization ratio, we have found mean values of  
604  $0.21 \pm 0.03$  and  $0.26 \pm 0.01$  respectively for the aforementioned atmospheric layer. In  
605 addition, a slightly increase of depolarization ratio with altitude is observed. The reason  
606 behind it lies on the fact that non-spherical particles tend to produce a higher backscatter  
607 signal related to the cross-polarized component and higher depolarization ratios. African  
608 dust aerosols are well known as non-spherical particles. So this observation would  
609 suggest that at higher altitudes (from 1 to 3 km asl) the mineral dust is purer since  
610 depolarization ratios are greater. In relation to Fig. 12 (11:34-12:04 UTC), the aerosol  
611 dust plume is a bit weaker than in the previous period. The backscatter coefficient  
612 profiles are relatively lower and also the backscatter-related Ångström exponent profiles  
613 present higher values which should indicate partially a smaller aerosol size. In this  
614 sense, the contribution of the local aerosol may be greater. Considering these  
615 observations we can conclude that the intensity of the African dust for this period is  
616 lower than the previous one. Volume and particle depolarization ratios for the  
617 atmospheric layer situated at 1-3 km asl are similar than in the previous period. The  
618 mean values are  $0.19 \pm 0.01$  and  $0.28 \pm 0.02$  respectively.



619

620 **Fig. 11. Backscatter coefficient,  $\beta$ - $\text{\AA}$ ngström exponent, particle and volume**  
 621 **depolarization profiles at 08:11 UTC on 23 February 2017.**



622

623 **Fig. 12. Backscatter coefficient,  $\beta$ -Ångström exponent, particle and volume**  
624 **depolarization profiles at 11:34 UTC on 23 February 2017.**

625

626 **5 Performance of dust models during intense events**

627

628 This section aims at examining the performance of dust models to predict the 3D  
629 evolution of mineral dust during such intense outbreaks. The literature available on the  
630 evaluation of modelled dust vertical profiles usually inspects the behavior of such  
631 models on long time series or for a single moderate outbreak (Gobbi et al. 2013, Santos  
632 et al. 2013, Mona et al. 2014, Binietoglou et al. 2015, Sicard et al. 2015), and only  
633 rarely for intense outbreaks (Huneeus et al. 2016, Ansmann et al. 2017, Tsekeri et al.  
634 2017).

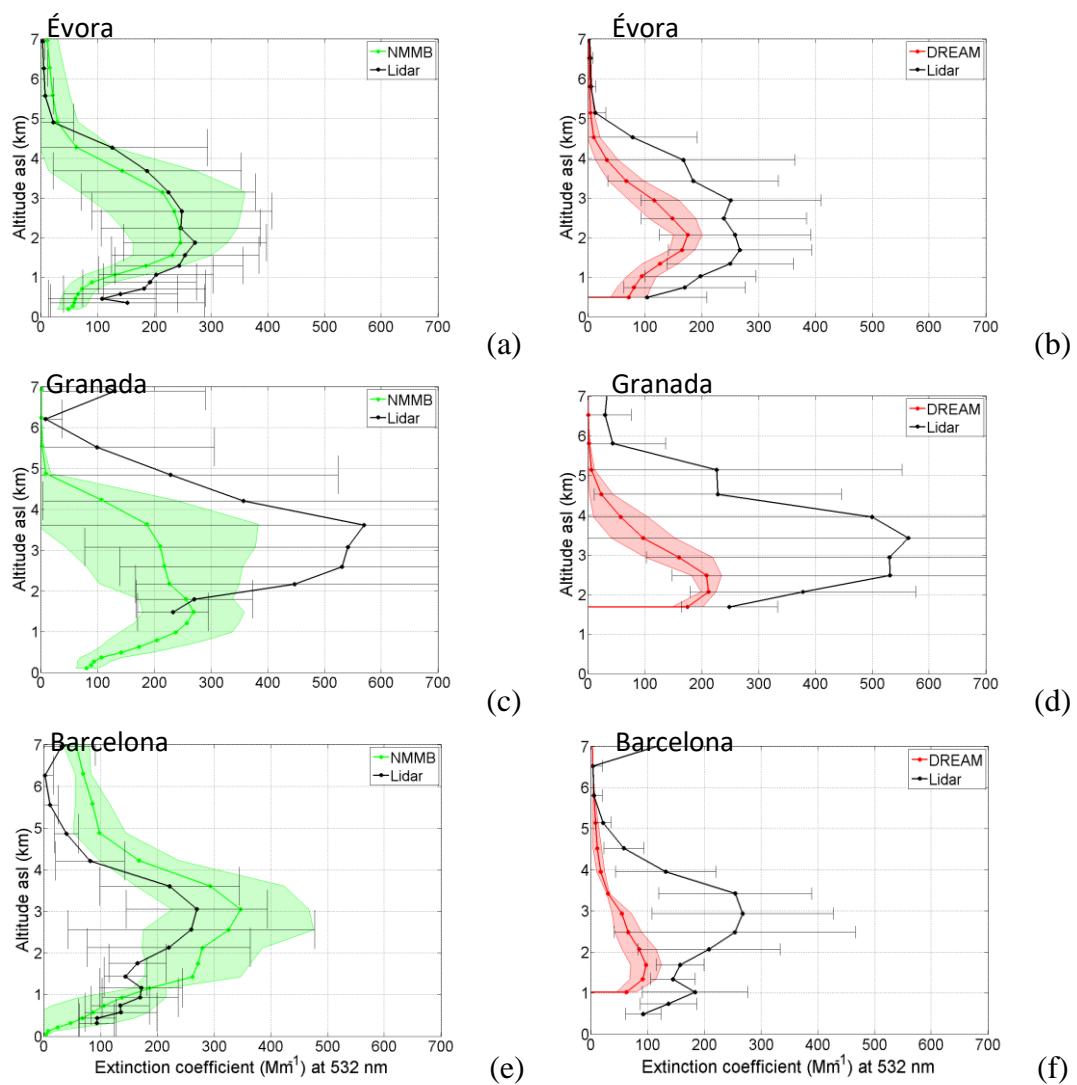
635 **5.1 Forecast skill for a lead time of 24 h**

636 The results are presented for the three sites of Évora, Granada and Barcelona. There are  
637 too few measured profiles in Madrid to allow for a statistical comparison. The  
638 comparison of the temporal mean profiles of extinction coefficient is made for  
639 NMMB/BSC-Dust and BSC-DREAM8b in Fig. 13. The temporal means are averaged  
640 over the whole period (see caption of Fig. 13). For each individual profile the  
641 correlation coefficient is plotted as a function of fractional bias (FB) in Fig. 14 and the  
642 temporal evolution of the latter two parameters is shown in Fig. 15. In the latter figure  
643 the time evolution of *FB* and *r* is also shown for lead times of 48 and 72 h and  
644 discussed in Section 5.2. The mean values of the fractional bias, the correlation  
645 coefficient and the center of mass for both models at each site are reported in Table 5.

646 Table 5 also contains these mean values for lead times of 48 and 72 h, which are  
647 discussed in Section 5.2.

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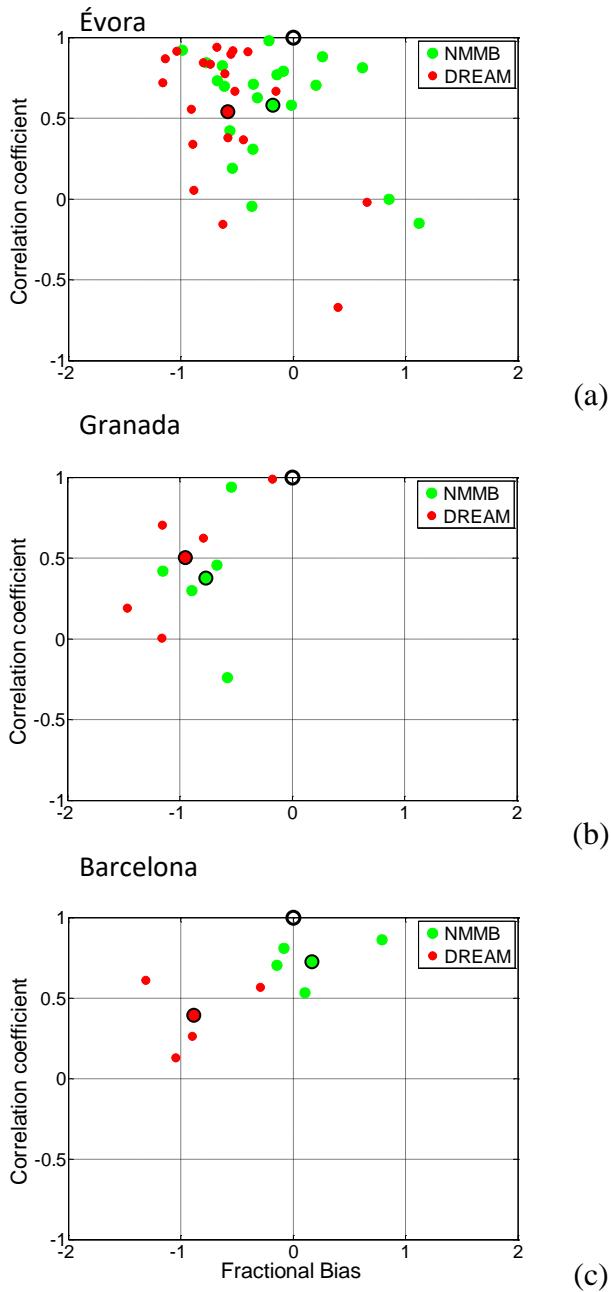
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652 **Fig. 13. Mean vertical distribution of mineral dust extinction coefficient estimated**  
653 **by NMMB/BSC-Dust in (a) Évora, (c) Granada and (e) Barcelona and by BSC-**  
654 **DREAM8b in (b) Évora, (d) Granada and (f) Barcelona. The period considered,**  
655 **not always continuous, are 21 Feb. 12UT – 23 Feb. 23UT, 21 Feb. 12UT – 22 Feb.**  
656 **19UT and 23 Feb. 08UT – 23 Feb. 21UT for Évora, Granada and Barcelona,**

657 respectively. The model shaded areas and the error bars of the lidar represent the  
658 standard deviations. All model forecasts are for a lead time of 24 h.

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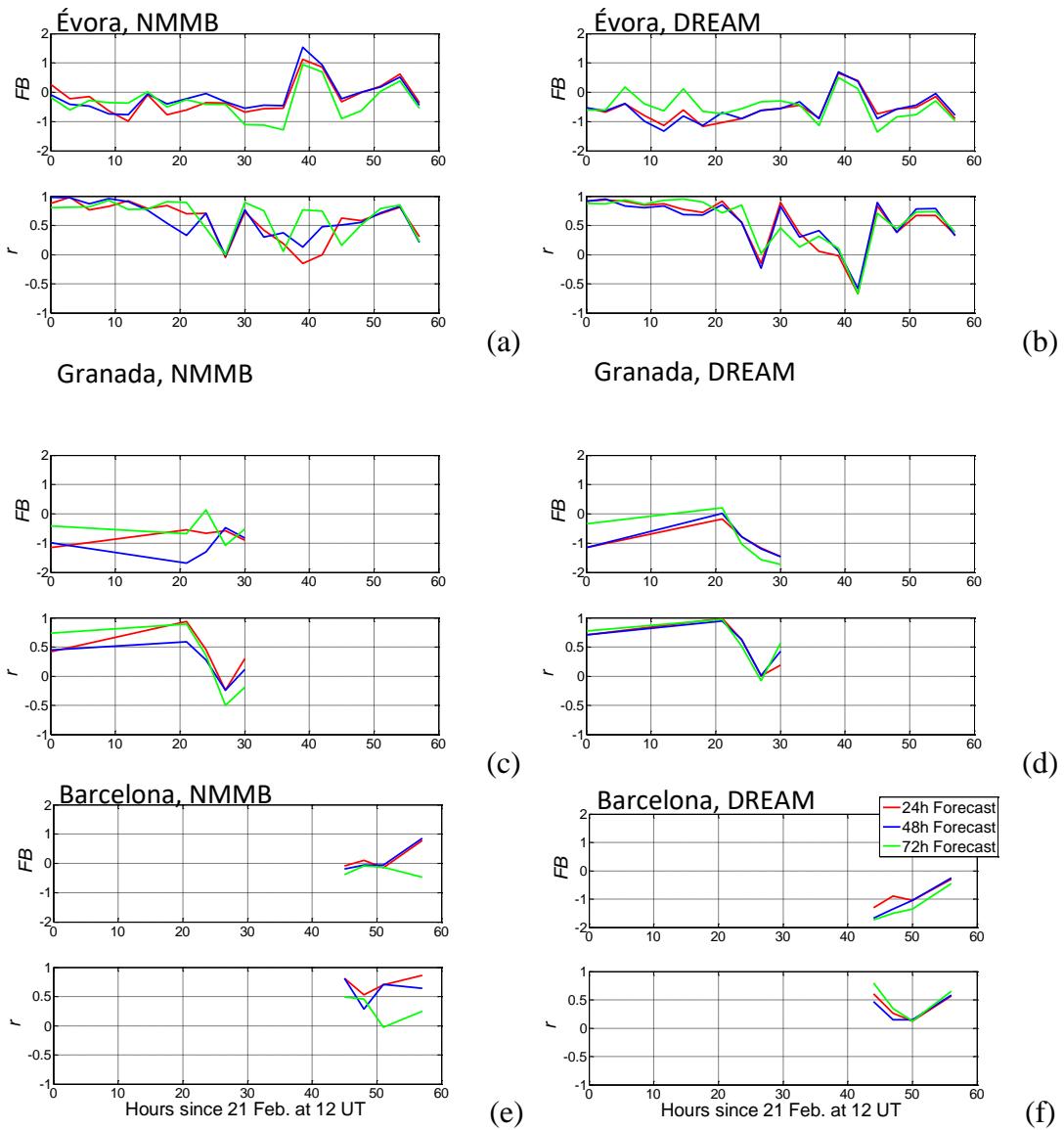
663

664 **Fig. 14. Correlation coefficient vs. fractional bias, both calculated for the extinction**  
665 **coefficient, calculated for each individual profile in (a) Évora, (b) Granada and (c)**  
666 **Barcelona. All model forecasts are for a lead time of 24 h. The mean values are**  
667 **represented by larger dots edged by a black line. The ideal (FB/r) pair, (0/1), is**  
668 **indicated by a black circle.**

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675 **Fig. 15.** Correlation coefficient and fractional bias, both calculated for the  
 676 extinction coefficient, vs. time for forecast lead times of 24, 48 and 72 h for  
 677 NMMB/BSC-Dust in (a) Évora, (c) Granada and (e) Barcelona and for BSC-  
 678 DREAM8b in (b) Évora, (d) Granada and (f) Barcelona. The legend shown in the  
 679 last plot applies to all plots.

680

681 **Table 5.** Main results of the comparison between models and observations. *FB* and  
 682 *r* are both calculated for the extinction coefficient.

Évora (21 Feb. 12UT – 23 Feb. 23UT)

	NMMB/BSC-Dust			BSC-DREAM8b		
Number of profiles				20		
Lead time (hours)	24	48	72	24	48	72
<i>FB</i> (%)	-18.0	-12.3	-34.5	-58.0	-57.5	-48.4
<i>r</i>	0.58	0.59	0.65	0.54	0.55	0.56
Model CoM (km)	2.70	2.82	3.04	2.21	2.26	2.38
Lidar CoM (km)	2.43			2.46		

Granada (21 Feb. 12UT – 22 Feb. 19UT)

	NMMB/BSC-Dust			BSC-DREAM8b		
Number of profiles				5		
Lead time (hours)	24	48	72	24	48	72
<i>FB</i> (%)	-76.7	-105.5	-50.8	-95.1	-91.6	-89.0
<i>r</i>	0.37	0.24	0.26	0.50	0.54	0.55
Model CoM (km)	2.26	2.38	2.14	2.70	2.79	2.95
Lidar CoM (km)	3.31			3.37		

Barcelona (23 Feb. 08UT – 23 Feb. 21UT)

	NMMB/BSC-Dust			BSC-DREAM8b		
Number of profiles				4		
Lead time (hours)	24	48	72	24	48	72
<i>FB</i> (%)	+16.8	+13.8	-26.9	-88.6	-108.0	-126.3
<i>r</i>	0.73	0.61	0.29	0.39	0.34	0.48
Model CoM (km)	3.61	3.72	4.37	2.51	2.53	2.69
Lidar CoM (km)	2.57			2.59		

683 When looking at the temporal mean profiles of extinction coefficient (Fig. 13), the most  
684 striking feature is the general large underestimation of BSC-DREAM8b at all heights  
685 independently of the site. This underestimation is smaller in Évora (mean  $FB = -48\%$  ,  
686 Table 5) than in Barcelona (mean  $FB = -88.6\%$  ), where the underestimation is larger. In  
687 Fig. 14 it is observed a horizontal spread of the variability of  $FB$  larger in Évora ([-115;  
688 +65 %]) than in Granada ([-150; -20 %]) and Barcelona ([-130; -30 %]) probably due to  
689 the smaller amount of vertical profiles available in Barcelona. The same tendency is  
690 observed on NMMB/BSC-Dust data. NMMB/BSC-Dust forecasts show a rather good  
691 agreement with the observations, especially in Évora and in Barcelona. While the  
692 model tends to underestimate the observations in Évora (especially below the CoM; the  
693 mean  $FB$  is -18.0 %) and in Granada (especially near the CoM; the mean  $FB$  is -76.7  
694 %), it tends to overestimate the observations in Barcelona (especially above 1 km; the  
695 mean  $FB$  is +16.8 %). The agreement between NMMB/BSC-Dust and the Évora lidar  
696 is remarkably good (Fig. 13a), taking into account the atmospheric variability  
697 represented by the lidar error bars and the rather long period considered (60 hours).  
698 While the NMMB/BSC-Dust profiles reach zero at an approximate height of 5 km in  
699 Évora and Granada (similarly to the observations), the profiles in Barcelona start  
700 decreasing linearly from  $\sim 100 \text{ Mm}^{-1}$  at 5 km height to  $\sim 50 \text{ Mm}^{-1}$  at 7 km (when the  
701 observations indicate an extinction coefficient lower than  $50 \text{ Mm}^{-1}$  above 4.5 km and  
702 reaching zero at 6 km). The increase of the lidar profile at 7 km is due to the presence of  
703 a cloud above 7 km (see Fig. 10) taken into account in the computation of the layer-  
704 average lidar extinction coefficient at the model height. Possible explanations of the  
705 differences observed between NMMB/BSC-Dust and the observation in Barcelona in  
706 the upper part of the profile are given in the next paragraph. Also in Barcelona the lidar  
707 profiles show a layer connected to the surface below 1.5 km, which is not reproduced by

708 either of the models. The main reason is probably the presence of non-dust type  
709 particles mixed with the dust detected in the observations but not taken into account in  
710 the models. It is also worth noting that BSC-DREAM8b reproduces less atmospheric  
711 variability than NMMB/BSC-Dust (in fig. 13, the red shaded areas are smaller than the  
712 green ones), whereas the atmospheric variability denoted by the lidar error bars is large  
713 at all sites. This seems to indicate that BSC-DREAM8b has less nervousness than  
714 NMMB/BSC-Dust.

715 The capacity of the models to reproduce the shape of the dust vertical distribution is  
716 estimated with the correlation coefficient calculated between individual modeled and  
717 observed profiles. NMMB/BSC-Dust seems to perform better in Barcelona ( $r$  values of  
718 0.73) and in Évora (0.58) than in Granada (0.37). BSC-DREAM8b seems to perform  
719 better in Évora ( $r$  values of 0.54) and in Granada (0.50) than in Barcelona (0.39). The  
720 low  $r$  value obtained with BSC-DREAM8b in Barcelona (0.39) is apparently due to a  
721 vertical downward transport forecast by the model and not visible from the observations  
722 (the peak of BSC-DREAM8b profile is approximately 2 km lower than the peak of the  
723 lidar, see Fig. 13f). (Huneeus et al. 2016), who compared NMMB/BSC-Dust and BSC-  
724 DREAM8b, among other models, to CALIOP (Cloud Aerosol Lidar with Orthogonal  
725 Polarization) profiles during an intense dust outbreak in April 2011 with an AOD  $\sim 0.8$ ,  
726 found a general underestimation of the dust layer height, that was attributed to an  
727 overestimation of the dust deposition near the source. The fact that the cloud of points  
728 along the  $r$ -axis is more spread in Évora (Fig. 14a) than in Granada or Barcelona (Fig.  
729 14b and c) is probably due to the longer time series available in Évora covering two and  
730 a half days of the event. Another indicator of the score of the models related to the  
731 vertical structure of the dust layer is the center of mass (CoM). In Évora both models  
732 retrieve well the center of mass of the dust layers (differences between modeled and

733 observed CoM are less than 0.27 km, see Table 5). In Granada, boths models reproduce  
734 smaller CoM values with discrepancies vs. the observations of 1.05 km (NMMB/BSC-  
735 Dust) and 0.67 km (BSC-DREAM8b). At this site both models predict a center of mass  
736 of the dust plume closer to the ground than it is in reality. In Barcelona BSC-  
737 DREAM8b predicts well the CoM with a discrepancy of 0.08 km, The overall  
738 performance of BSC-DREAM8b at all three stations are in relatively good agreement  
739 with the difference of  $0.3 \pm 1.0$  km found between the same model and the EARLINET  
740 station of Potenza, Italy, over a period of 12 years and for dust events with  $\text{AOD} < 0.9$   
741 (Mona et al. 2014). In Barcelona, the mean CoM forecasted by NMMB/BSC-Dust is  
742 3.61 km while the lidar measured a mean value of 2.57 km. This large difference is due  
743 to the mean NMMB/BSC-Dust profile of extinction in Barcelona which does not reach  
744 zero at  $\sim 5$  km, unlike at the other sites (Fig. 13e; see also the former paragraph). This  
745 finding suggests that one or several processes taken into account in NMMB/BSC-Dust  
746 and inducing vertical motion of the dust layers did actually not occur. One of these  
747 processes is the troposphere–stratosphere exchanges which in some cases has been  
748 found to be overestimated by the model because of a misrepresentation of the  
749 tropopause that normally limits the maximum altitude of dust transport (Janjic 1994).  
750 However, given the limited vertical extension of the dust plume ( $< 5$  km), such an  
751 explanation is very unlikely. In our case the vertical upward transport of the dust layers  
752 at high altitudes forecast in Barcelona but not in the southern sites is probably due to a  
753 too long aerosol lifetime in the upper layers and/or underestimated deposition processes  
754 (Mona et al. 2014). Interestingly this overestimation of NMMB/BSC-Dust in the upper  
755 layers was also observed by (Binietoglou et al. 2015) who found a slight overestimation  
756 of NMMB/BSC-Dust above 4.5-5 km when comparing the model with LIRIC  
757 (Lidar/Radiometer Inversion Code) profiles of mass concentration at several sites in

758 Europe and by (Sicard et al. 2015) who compared the model with profiles from  
759 EARLINET stations during a moderate dust event affecting the western Mediterranean  
760 Basin in July 2012.

761 5.2 Forecast skill temporal evolution and comparison for different lead times

762 The temporal evolution of the score of the models (in terms of  $FB$  and  $r$ ) for different  
763 lead times shown in Fig. 15 allows us to evaluate the forecast skill of each model as a  
764 function of time since the forecast initialization. The start of the time series is fixed on  
765 21 February, 2017, at 12 UTC, referred in the following as time  $T_0$ , when the first  
766 observations are available (in Évora and Granada). The observations available allow us  
767 to have 20 points of comparison (at a 3-hour time step) from the 21<sup>st</sup> at 12 UTC until the  
768 23<sup>rd</sup> at 23 UTC in Évora; 5 points of comparison on the 22<sup>nd</sup> between 07 and 19 UTC in  
769 Granada; and 4 points of comparison on the 23<sup>rd</sup> between 08 and 21 UTC in Barcelona.  
770 In all plots we have represented the temporal evolution of  $FB$  and  $r$  for lead times of  
771 24, 48 and 72 h. We first discuss the forecast skill temporal evolution for a lead time of  
772 24 h, and then compare it to the evolution at 48 and 72 hours.

773 In Évora during the first 20 hours (Fig. 15a and b, red lines) both models have similar  
774 and more or less stable correlation coefficients with values larger than 0.5. The  
775 fractional bias is negative and varies in the range [-100; 0 %]. It is larger (in absolute  
776 value) for BSC-DREAM8b than for NNMB/BSC-Dust. At  $T_0 + 20$  hours (the 22<sup>nd</sup> at 08  
777 UTC) the situation starts to degrade:  $FB$  variations are larger from one prognostic to  
778 the next, especially for NNMB/BSC-Dust, and  $r$  passes regularly below the value of  
779 0.5. A few hours before  $T_0 + 40$  hours (the 23<sup>rd</sup> at 04 UTC) and only for a period of 5-6  
780 hours both models overestimate the extinction coefficient ( $+50 < FB < +150$  %). During  
781 the first hours of the 23<sup>rd</sup> the AOD in Évora reached its highest values (~2.5; see Fig. 3).

782 In that sense, it seems that the peak of the event is well reproduced in time by the  
783 models but its intensity is overestimated. In Granada (Fig. 15c and d, red lines) the  
784 prognostic of NNMB/BSC-Dust is quantitatively better (smaller values of  $FB$ ) but  
785 qualitatively worst (smaller correlation coefficients) than for BSC-DREAM8b. Our  
786 findings in Granada are in the same line as those found by (Sicard et al. 2015) for a  
787 moderate dust event affecting the western Mediterranean Basin in July 2012 who also  
788 found that NNMB/BSC-Dust reproduced quantitatively better the profiles while BSC-  
789 DREAM8b reproduced better the shape of the profiles. The decrease of  $FB$  visible for  
790 both models in Granada and starting at  $T_0 + 20$  (the 22<sup>nd</sup> at 08 UTC) coincides with the  
791 increase of AOD from ~0.5 to values above 2.0 (see Fig. 3). While on the peak day in  
792 Évora (the 23<sup>rd</sup>) both prognostics show an overestimation for a short period of time, on  
793 the peak day in Granada (the 22<sup>nd</sup>) the general underestimation of both prognostics is  
794 accentuated, especially for BSC-DREAM8b. In Barcelona (Fig. 15 e and f, red lines)  
795 the comparison starts at  $T_0 + 44$  (the 23<sup>rd</sup> at 08 UTC) at the peak of the event in  
796 Barcelona (AOD>2.0, see Fig. 3). NNMB/BSC-Dust shows a very good quantitative  
797 agreement in the morning and an overestimation in the afternoon, while BSC-  
798 DREAM8b shows an underestimation, which decreases with time. The shape of the  
799 vertical profiles is better reproduced by NNMB/BSC-Dust ( $r > 0.5$ ) than by BSC-  
800 DREAM8b ( $r < 0.5$ ). In general the forecast skills of BSC-DREAM8b in Barcelona are  
801 not as good as those of the southernmost sites. This difference, also observed by  
802 (Huneeus et al. 2016) for dust northward transport, might be explained by the  
803 difficulties of the models in simulating horizontal winds and vertical dust propagation.

804 If we now look at the forecast skill as a function of lead time, i.e. at the differences  
805 between the red, blue and green lines in Fig. 15, corresponding, respectively, to lead  
806 times of 24, 48 and 72 hours, the most striking result is that, at first sight, no clear

807 degradation of the prognostics is clearly visible. There is a difference in the temporal  
808 evolution of the prognostics: the prognostics at 24 and 48 h are usually quite similar and  
809 the one at 72 h is the one that differs the most from the prognostic at 24 h. But, all in all,  
810 if one looks at the overall mean values in Table 6, no clear tendency appears neither in  
811 terms of  $FB$ , nor  $r$ , for Évora and Granada, the two stations closest to the source. In  
812 this sense these results are in agreement with those of (Huneeus et al. 2016) who found  
813 that the forecast skill of both models for AOD was independent of the forecasting lead  
814 time in the domain they defined as southern Europe. In Barcelona a slight degradation  
815 of the model scores occurs with increasing lead times: the fractional bias increases (in  
816 absolute value; BSC-DREAM8b) and the correlation coefficient decreases  
817 (NMMB/BSC-Dust) between the prognostics at 24 and 72 h. This deterioration of the  
818 forecast skills is not observed in (Huneeus et al. 2016) and may be due to the singularity  
819 and exceptionality of the event described in our study.

## 820 **6 Conclusions**

821 An extreme dust outbreak transported from Northern Africa to the western  
822 Mediterranean during 20-23 February 2017 has been reported and analyzed in the IP.  
823 By means of lidar and sun-photometer measurements, we have provided a  
824 representative picture of this extreme event by means of a detailed 4-D characterization  
825 of aerosol optical properties and their evolution during the African event. Furthermore,  
826 the combined use of active and passive remote sensing instruments along with dust  
827 models has provided useful information to better understand the complexity of dust  
828 long-range transport, its extreme character and also the capability of dust models to  
829 forecast such events.

830 The appearance of the Moroccan low reinforced by the Atlantic anticyclonic system was  
831 responsible for the tropospheric flow that advected atmospheric mineral dust over the IP

832 during this extreme event. From the photometry, we would like to remark two main  
833 ideas concerning the most intense stages of the event. Firstly, AOD at 675 nm were  
834 registered to be around and over 2, the Ångström Exponent (440/870 nm) was close to  
835 0, and SSA was close to 1 in most of AERONET stations, which indicates a high  
836 aerosol load, a large aerosol size and the dispersive nature of these particles,  
837 characteristics that are attributed to mineral dust. Secondly, the African dust outbreak  
838 was accompanied by the presence of clouds that hampered an adequate retrieval and  
839 consequently no sun-photometer observations were available at some AERONET  
840 stations.

841 From lidar measurements, the African dust plume could be observed in each lidar  
842 station. In general, the altitude range of the plume was observed from the ground until  
843 4-5 km asl approximately at every lidar station. Maximum values of backscatter  
844 coefficients at 532 nm were registered by each lidar system in the range  $1 - 1.5 \cdot 10^{-5} \text{ m}^{-1}$   
845  $\text{sr}^{-1}$ , where, during the most intense stages the high aerosol load prevented the retrieval,  
846 which could not be carried out. This is an issue that also complicated the retrieval in  
847 every site. Minimum backscatter-related Ångström exponents at these stages were  
848 monitored very close to 0, which are in agreement with the results provided by the  
849 sunphotometry. Lidar ratios were found in the range 40 - 55 sr at 355 nm and 34 - 61 sr  
850 at 532 nm during the event at Évora and Granada. Particle and volume depolarization  
851 ratios, registered at those stations where depolarizing channels were available, have  
852 shown an interesting consistency of these values given the fact they were very similar.  
853 In general, large particle and volume depolarization ratios are attributed to mineral dust  
854 since they are not spherical particles and produce a higher backscatter signal related to  
855 the cross-polarized component. The larger the particle and volume depolarization ratios,  
856 the purer mineral dust. Likewise, according to these depolarizing properties, lidar

857 systems equipped with this channel have indicated the different structures and aerosol  
858 layers throughout the vertical column to distinguish local aerosol from mineral aerosol  
859 (for instance in Granada). These findings suggest the need of use of combined  
860 instrumentation to characterize adequately aerosol optical properties during this kind of  
861 events.

862 When it comes to forecasting this extreme event, two dust models have been used:  
863 BSC-DREAM8b and NMMB/BSC-Dust. According to the fractional bias and the  
864 correlation coefficient analysis there is a large underestimation in the forecast of the  
865 extinction coefficient provided by BSC-DREAM8b at all heights independently of the  
866 site. By contrast, NMMB/BSC-Dust forecasts presented a better agreement with the  
867 observations, especially in Évora . However the NMMB/BSC-Dust reproduced a higher  
868 atmospheric variability than BSC-DREAM8b. Some discrepancies such as the forecast  
869 of dust by NMMB/BSC-Dust in layers well above 5 km are still not completely  
870 understood and further research is needed. Finally, with regard to the forecast skill as a  
871 function of lead time of each model, no clear degradation of the prognostic is  
872 appreciated at 24, 48 and 72h for Évora and Granada stations, however it does for  
873 Barcelona, which is in principle attributed to the singularity of the event.

874

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