How Marine Emissions of Bromoform Impact on the Remote Atmosphere

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10   **Abstract**

11   It is an open question how localized elevated emissions of CHBr$_3$ and other VSLHs, found in coastal and

12   upwelling regions and low background emissions, typically found over the open ocean, impact on the

13   atmospheric VSLH distribution. In this study, we use the Lagrangian dispersion model FLEXPART to

14   simulate atmospheric CHBr$_3$ resulting from assumed uniform background emissions, and from elevated

15   emissions consistent with those derived during three tropical cruise campaigns.

16   The simulations demonstrate that the atmospheric CHBr$_3$ distributions in the uniform background

17   emissions scenario are highly variable with high mixing ratios taking place in regions of convergence or

18   low wind speed. This relation holds on regional and global scales.

19   The impact of localized elevated emissions from the research cruises, on the atmospheric CHBr$_3$

20   distribution varies significantly from campaign to campaign. The estimated impact depends on the

21   strength of the emissions and the meteorological conditions. In the open waters of the western Pacific

22   and Indian Oceans, localized elevated emissions only slightly increase the background concentrations of

23   atmospheric CHBr$_3$, even when 1° wide source regions along the cruise tracks are assumed. Near the

24   coast, elevated emissions, including hotspots up to 100 times larger than the uniform background

25   emissions, can be strong enough to be distinguished from the atmospheric background. However, it is

26   not necessarily the highest hotspot emission that produces the largest enhancement, since the tug-of-war

27   between fast advective transport and local accumulation at the time of emission is also important.

28   Our results demonstrate that transport variations of the atmosphere itself are sufficient to produce highly

29   variable VSLH distributions, and elevated VSLHs in the atmosphere do not always reflect a strong

30   localized source. Localized elevated emissions can be obliterated by the highly variable atmospheric

31   background, even if they are orders of magnitude larger than the average open ocean emissions.

32

1. Introduction

Very short-lived halocarbons (VSLHs) with atmospheric lifetimes shorter than 6 months from natural oceanic sources are dominated by brominated and iodinated compounds (Carpenter and Liss, 2000; Quack et al., 2004; Law et al., 2006). VSLHs have drawn considerable interest due to their contribution to stratospheric ozone depletion and tropospheric chemistry (Solomon et al., 1994; Dvortsov et al., 1999; Salawitch et al., 2005; Feng et al., 2007; Tegtmeier et al., 2015; Hossaini et al., 2015). In this work, we focus on the VSLH bromoform ($CHBr_3$), since most organic oceanic bromine is released into the atmosphere in this form.

$CHBr_3$ concentrations measured in ocean waters are characterized by large spatial variability with elevated abundances in phytoplankton blooms (Baker et al., 2000, Liu et al., 2013) and equatorial and upwelling regions due to biological sources (Carpenter et al., 2009; Quack and Wallace, 2003; Quack et al., 2007; Fuhlbrügge et al., 2016). The open ocean generally shows homogeneous, low $CHBr_3$ concentrations, compared to higher concentrations and strong gradients found in coastal and shelf areas (Quack and Wallace, 2003). At the coast, high oceanic concentrations are related to macro algae (Klick and Abrahamsson, 1992) and anthropogenic sources (Boudjellaba et al., 2016) such as power plants (Yang, 2001) and desalination facilities (Agus et al., 2009).

Due to sparse measurements and limited process understanding, existing estimates of global air-sea flux distributions of $CHBr_3$ and other VSLHs are subject to large uncertainties (e.g. Warwick et al., 2006; Palmer and Reason, 2009; Liang et al., 2010; Ordóñez et al., 2012; Stemmler et al., 2013; Ziska et al., 2013; Carpenter et al., 2014). The spatial and temporal distribution of elevated emissions in coastal and upwelling regions is currently based on limited observations. Campaigns in these regions suggest that emissions generally increase near coastlines, and that sporadic peak emissions with extremely high values can be found (e.g. Butler et al., 2007; Liu et al., 2013; Fuhlbrügge et al., 2016; Fiehn et al., 2017). Analysis of the measurements suggests that such peak emissions are often of limited spatial extent and cover not more than a distance of 50-100 km along the cruise track.

There are two main approaches to derive the magnitude of VSLH emissions, i.e. the "bottom-up" approach (e.g. Quack and Wallace, 2003; Carpenter and Liss, 2000; Butler et al., 2007; Ziska et al., 2013) and the "top-down" approach (e.g. Warwick et al., 2006; Liang et al., 2010; Ordóñez et al., 2012). For the "bottom-up" method, measured surface sea water concentrations of VSLHs at the "bottom" (surface) are extrapolated to estimate global emissions. For the "top-down" method, the emissions of VSLHs are constrained by the measured abundances at the "top" (atmosphere) so that model simulations based on the constrained global emission estimates reproduce the observed atmospheric concentrations. These two approaches yield different estimates of the global VSLHs emissions, with the recent "top-down" approaches resulting in generally higher emissions than the recent "bottom-up" approaches.

In the tropical ocean waters of the Atlantic, the western Pacific and Indian Ocean, the existence of localized elevated $CHBr_3$ emissions and hotspots has been confirmed (Butler, et al, 2007; Liu et al., 2013;

Krüger and Quack, 2013; Fiehn et al., 2017). At the same time, these convectively active regions offer an efficient pathway for the vertical transport of short-lived oceanic compounds from the boundary layer to the stratosphere (e.g. Aschmann et al., 2009; Hossaini et al., 2012; Tegtmeier et al., 2012, 2013; Marandino et al., 2013; Liang et al., 2014). Moreover, the Asian monsoon has been recognized as an efficient transport pathway for short-lived pollutants and VSLHs (Randel et al., 2010; Hossaini et al., 2016; Fiehn et al., 2017). Given that elevated oceanic $CHBr_3$ emissions are expected to occur in the same regions as strong convection, it is of interest to analyze how these elevated emissions impact $CHBr_3$ in the atmospheric boundary layer, which feeds into the upward transport.

Measurements of $CHBr_3$ abundance in the atmospheric boundary layer show large spatial variability (e.g., Quack and Wallace, 2003; Montzka and Reimann, 2011; Lennartz et al., 2017). A compilation of available measurements by Ziska et al. (2013) suggests similar $CHBr_3$ distribution patterns in the atmospheric boundary layer as in the surface ocean, with higher mixing ratios in the equatorial, coastal and upwelling regions. However, given the sparse data base and the uncertainties in the spatial and temporal extent of oceanic emissions, the detailed distribution of boundary layer $CHBr_3$, cannot be well constrained (e.g., Hepach et al., 2014; Fuhlbrügge et al., 2013). On the one hand, the spatial and temporal extent of elevated localized emissions is usually unknown, leading to large uncertainties when estimating their overall magnitudes. On the other hand, the influence of meteorological conditions, distinctive transport patterns and variations of atmospheric sinks, such as the background OH field (e.g. Rex et al., 2014), can be expected to modulate the effect of elevated oceanic sources. Knowledge about the interplay between sources, transport and loss processes is relevant to understand the importance of localized elevated emissions for atmospheric abundances and to interpret existing atmospheric measurements with respect to potential sources and driving factors.

In this study, we use observational data from three tropical research cruises, one in the Indian Ocean (OASIS) and two in the western Pacific (TransBrom and SHIVA). We use the Lagrangian particle dispersion model FLEXPART to investigate the transport and atmospheric distribution of VSLHs. Taking bromoform ($CHBr_3$) as example, we compare the atmospheric signals estimated from the elevated and hotspot emissions measured during the ship campaigns to the distribution derived from only uniform background emissions. We use the term 'elevated emissions' when describing emissions that are on average up to a factor of 10 larger than the background and 'hotspot emissions' for sporadic emissions up to a factor of 100 larger than the background. The campaigns and the FLEXPART model are introduced in Sect. 2. In Section 3, we discuss the distributions and variability of atmospheric $CHBr_3$ based on uniform background emissions. We present the observed hotspots of $CHBr_3$ emissions in Section 4.1, and compare the simulated atmospheric mixing ratios resulting from elevated emissions during three campaigns with the background values (Section 4.2). Conclusions are given in Section 5.

2. Data and Methods

2.1 Background and in-situ CHBr$_3$ emissions

In this study, we distinguish between open ocean background and in-situ CHBr$_3$ emissions. Open ocean emissions are inferred to be around 100 pmol h^{-1} m^{-2} based on global bottom up scenarios (Quack and Wallace, 2003; Butler et al., 2007; Liu et al., 2013; Fiehn et al., 2017; Ziska et al., 2013). While emissions for individual regions and seasons can be higher or lower than this, including negative fluxes going from the atmosphere into the ocean, 100 pmol h^{-1} m^{-2} represents a typical mean value averaged over all oceanic basins between 60°S and 60°N. The background open ocean emissions exclude by design emissions from coastal, shelf and upwelling regions.

In-situ oceanic emissions of CHBr$_3$ have been calculated from the observational data collected during three tropical ship campaigns. During each campaign, surface air and water samples were collected simultaneously at regular intervals (every 3 to 6 hours). The emissions were calculated from these co-located data and the instantaneous wind speed (Ziska et al., 2013, Fuhlbrügge et al., 2016, Fiehn et al., 2017). The air-sea flux was obtained from the transfer coefficient (k_w) and the concentration gradient (Δc) between water concentration and the theoretical equilibrium water concentration (see details in Fiehn et al., 2017 and reference therein):

$$F = k_w \cdot \Delta c \qquad (1)$$

The two campaigns TransBrom (October 11th-23rd, 2009) and SHIVA (November 15th-28th, 2011) took place in the western Pacific, while the OASIS campaign (July 11th-August 6th, 2014) was conducted in the western Indian Ocean. The detailed cruise track and the magnitude of the oceanic CHBr$_3$ emissions of each campaign is given in Fig. 1. The in-situ emissions include both open-ocean emissions and elevated emissions from coastal, shelf, and upwelling regions.

2.2 Modeling

For the simulations of the atmospheric distribution and transport of CHBr$_3$, we used the Lagrangian particle dispersion model, FLEXPART (Stohl et al., 2005), which has been validated by previous comparisons with measurements (Stohl et al., 1998; Stohl and Trickl, 1999). Lagrangian particle models such as FLEXPART compute trajectories of a large number of so-called particles, presenting infinitesimally small air parcels, to describe the transport, diffusion and chemical decay of tracers in the atmosphere. The model includes turbulence in the boundary layer and free troposphere (Stohl and Thomson, 1999) and a moist convection scheme (Forster et al., 2007) following the parameterization by Emanuel and Živković-Rothman (1999). The representation of convection in FLEXPART simulations has been validated with tracer experiments and ^{222}Rn measurements (Forster et al., 2007). Chemical or radioactive decay of the transported tracer is accounted for by reducing the particle mass according to a prescribed lifetime of the tracer. Alternatively, the loss processes can be prescribed via OH reaction based on a monthly averaged 3 dimensional OH-field. In this study, we employ FLEXPART version 10.0, which is driven by 3-hourly meteorological fields from ECMWF (European Centre for Medium-

146 Range Weather Forecasts) reanalysis product ERA-Interim (Dee et al., 2011) with a horizontal resolution
147 of 1º x 1º and 61 vertical model levels.

148 We performed two kinds of simulations based on the different emission scenarios. The first one used the
149 uniform global background emission, and the second one used in situ emissions observed during
150 individual ship campaigns. Trajectories released from the global ocean surface or along the cruise track
151 carry the amount of CHBr$_3$ prescribed by the respective emission scenario. Chemical decay of CHBr$_3$
152 was accounted for by:

$$m(t + \Delta t) = m(t) \exp(-\Delta t/\beta) \tag{2}$$

154 where m is the particle mass, $\beta = T_{1/2}/ln(2)$ is the e-folding lifetime of CHBr$_3$, and $T_{1/2}$ is the half-life of
155 CHBr$_3$ (Stohl et al., 2005). In our study, a half-life of 17 days (e-folding lifetime of 24 days) is prescribed
156 to CHBr$_3$ during all runs (Montzka and Reimann, 2011). For the background runs, a uniform air-sea flux
157 of 100 pmol h^{-1} m^{-2} was prescribed over all ocean surface area between 60°S and 60°N. Three runs were
158 conducted covering the time period of the campaigns with a 1-month spin-up period in each case to reach
159 a stable background concentration in the atmosphere.

160 For the in-situ emissions of each campaign, simulations were based on the calculated CHBr$_3$ air-sea flux
161 (see Fig. 8), which was released along the cruise track. The periods of the corresponding background
162 simulations with emissions over the whole time period, were the same as the campaign simulations. For
163 each observational data point, an emission grid cell centered on the measurement location was created.
164 These grid cells were designed to be adjacent along the cruise track and, based on the density of the
165 measurements, were about 0.1 – 2.0° wide in cruise track direction. The grid cells were chosen to be of
166 a fixed width (0.5° or 1°) in the other direction and thus add up to the narrow band of 0.5° or 1° width
167 centered along the cruise track (Fig. 1). Our design of the emission grid cells assumes that the elevated
168 emissions can extend over a distance of 0.5°-1°. This choice has been motivated by the spatial variability
169 of the measurements along the cruise track (see also section 4.1 and Fig. 7). Elevated emissions larger
170 than 1000 pmol h^{-1} m^{-2} were found at 77 different locations along the three cruise tracks examined in this
171 paper. Out of the 77 measurements, only 11 corresponded to singular locations with no adjacent high
172 emissions at the neighboring points. The other 66 measurements clustered together at 18 different
173 locations with at least two adjacent observational points showing emissions larger than 1000 pmol h^{-1} m^{-}
174 ^{2}. We defined the length of such a location of elevated emissions as the distance between the first and
175 last data point with an air-sea flux exceeding 1000 pmol h^{-1} m^{-2}. Most of the 18 locations extended over
176 a distance larger than 0.5° (13 out of 18) and nearly half were larger than 1° (8 out of 18) supporting our
177 choice of the width of the emissions grid cells. Note that the spatial extent of the hotspots was comparable
178 to the wind field resolution that drove our trajectory simulations. The amount of CHBr$_3$ released from
179 each grid cell was determined by the observational air-sea flux of the corresponding data point and scaled
180 with the width of the narrow emission band described above. The specified CHBr$_3$ emission from each
181 cell was kept constant for the duration of the model run and distributed over a fixed number of trajectories.
182 In order to capture the small-scale processes (e.g. convection), the large number of 2000 trajectories were

183 released from each 1° x 1° area of background runs and 20000 from each emission grid of regional in-
184 situ runs.

185 Output data in form of CHBr$_3$ volume mixing ratios (*VMR*) available at a user-defined grid, were
186 calculated by:

$$VMR = \left(\frac{c_T}{\rho_a}\right) \cdot \left(\frac{m_a}{m_T}\right) \qquad (3)$$

188 where c_T is the CHBr$_3$ mass concentration, ρ_a is the density of the air, and m_a and m_T are the molecular
189 weight of air and CHBr$_3$, respectively.

190 For each grid cell, the CHBr$_3$ mass concentration is given by:

$$C_T = \frac{1}{V}\sum_{i=1}^{N} m_i f_i \qquad (4)$$

192 with m_i being the mass of CHBr$_3$ for particle i, f_i the mass fraction of CHBr$_3$ of particle i attributed to the
193 respective grid cell, N the total number of the particles, and V the volume of the grid cell (Stohl et al.,
194 2005). We run FLEXPART in the non-domain filling mode, therefore the particle distribution is not
195 correlated with air density. Particles, and thus bromoform, can accumulate in regions of low wind speeds
196 where the relatively long residence time allows that oceanic emissions constantly add new particles.
197 Similarly, particles can accumulate in regions of convergence where horizontal inflow pools marine
198 boundary layer air from different regions. The output files are recorded at a horizontal resolution of 1° x
199 1° and 0.5° x 0.5° for background runs and in-situ runs, respectively, at every 100 m from 100 m to 1 km,
200 and every 1 km from 1 km to 20 km every 3 hours.

201

202 **3. Atmospheric CHBr$_3$ based on open ocean background emissions**

203

204 In this section, we show the impact of atmospheric transport patterns on the atmospheric CHBr$_3$
205 distribution, with the uniform background CHBr$_3$ emission simulations. The CHBr$_3$ mixing ratios in the
206 lower atmosphere diagnosed from the uniform background emissions (referred to as CHBr$_3$ background
207 mixing ratios hereinafter) vary significantly from campaign to campaign and also within each campaign
208 region. Figures 2 to 4 present snapshots of the CHBr$_3$ background mixing ratios and the simultaneous
209 wind fields from ERA-Interim reanalysis for the three campaigns. For TransBrom (Fig. 2), high CHBr$_3$
210 mixing ratios appear south of 15° N with a maximum near the equator, where the wind is weak. In the
211 northern Pacific, which is dominated by an anticyclone centered around 165°E, 30°N, the background
212 values are much lower. On the 10th of October 2009, two bands of extremely low wind fields exist, one
213 directly south of the equator and one tilting from 15°N to 5°N, which both coincide with the highest
214 CHBr$_3$ abundances. On the 20th of October, these two bands collided into one with lowest winds centered
215 around 165°E, where we again find very high values of CHBr$_3$ of up to 0.8 ppt. For both case studies,
216 highest values are found in the region of the lowest wind speeds or slightly shifted towards the region of
217 strongest wind shear. Regions of high wind speeds, such as the northern Pacific anticyclone, are
218 characterized by very low CHBr$_3$.

219 For the SHIVA case (Fig. 3), the background $CHBr_3$ accumulates in a narrow region near Indonesia,

220 with corresponding wind fields smaller than 3 m/s. North of Indonesia, the strong easterly trade winds

221 generally above 10 m/s prevent the accumulation of higher background values within the region. Again,

222 the two case studies illustrate how changes of the wind patterns within a few days' drive changes of the

223 background $CHBr_3$ distribution. Another particular example is the northward extension of the low

224 equatorial winds around 90°E on 16th November 2011, which leads to higher $CHBr_3$ north of the equator

225 up to 15°N.

226 For OASIS (Fig. 4), the wind speed is higher than in the other two regions and these strong

227 southeast/southwest trade winds associated with the Asian monsoon extend over most of the Indian

228 Ocean. Consistent with the stronger winds, the background values for the OASIS case are significantly

229 lower than for the other two cases, although they also show accumulations in certain regions. These

230 relatively higher background mixing ratios appear partially in regions of low wind speeds (e.g., near the

231 equator between 70°E and 90°E on 17th of July) or in adjacent regions of high wind shear (e.g., north of

232 the equator between 70°E and 90°E for both case studies). For the latter case, the $CHBr_3$ accumulation

233 also extends into the region of high wind speeds, which is different from the distribution found for the

234 TransBrom and SHIVA regions. This difference occurs because the east coast of the Indian Subcontinent

235 offshore is a region with wind convergence (dotted region), which tends to accumulate air masses therein.

236 Given that the accumulation of $CHBr_3$ background mixing ratios follows the wind field patterns on a

237 regional scale in most cases, we hypothesize that the interplay between wind speed and convergence may

238 influences the $CHBr_3$ distribution.

239 In order to validate the hypothesis, we show a violin plot of regional background $CHBr_3$ mixing ratios

240 related to convergence/divergence, and to the wind speeds averaged over each simulation period in Fig.

241 5. For the TransBrom case, the averaged ranges of mixing ratios in regions of convergence and

242 divergence (Fig. 5a) go up to 0.7 ppt and 0.5 ppt, respectively, with interquartile ranges of 0.1-0.35 ppt

243 and 0.05-0.21 ppt. Probability of mixing ratios larger than 0.2 ppt is much higher for regions of

244 convergence compared to regions of divergence. Meanwhile, in the regions grouped by wind speed (Fig.

245 5b), higher $CHBr_3$ mixing ratios are more likely to occur in regions with lower wind speeds (i.e. in the

246 regions of 0.0-5.0 m/s, mixing ratio go up to 0.65 ppt, while in the regions of 10-15 m/s, mixing ratio go

247 up to 0.25 ppt). Similar distributions also occur for the SHIVA case. During the OASIS case, the $CHBr_3$

248 mixing ratios are much smaller than for the other two cases due to stronger winds. Highest mixing ratios

249 (~0.15 to ~0.2 ppt) are found in the regions of convergence (Fig. 5e). However, higher mixing ratios are

250 also generally found in the regions of higher wind speeds (Fig. 5f), as the regions of convergence locate

251 in the regions of high wind speed during the OASIS case. The distributions suggest that in general higher

252 $CHBr_3$ mixing ratios tend to occur in the regions of convergence or lower wind speed, with the exception

253 of the OASIS case where extremely high winds occurred and coincided with regions of convergence.

254 The relationship mentioned above also holds on a global scale. The global distributions of atmospheric

255 $CHBr_3$ based on background emissions and wind fields averaged over the time periods of the SHIVA

256   and OASIS campaigns are presented in Fig. 6 and 7. We omit the time period of the TransBrom case,

257   since the background $CHBr_3$ distribution diagnosed for this period is very similar to the background

258   found for the SHIVA period. The global $CHBr_3$ background mixing ratios (Fig. 6a, and 7a) display a

259   very heterogeneous distribution in spite of the uniform background emission used for the simulations.

260   High $CHBr_3$ mixing ratios are again generally located in the regions of convergence, which also generally

261   correspond to low wind speeds on a global scale. For the SHIVA period (November 2011), particularly

262   high $CHBr_3$ background values of 0.3 to 0.4 ppt are found along the equator over the Maritime continent,

263   West Pacific, Indian Ocean and at the West coast of Africa, all of which are characterized by particularly

264   low winds. In the Northern and Southeast Pacific, the wind speed is generally higher, and the

265   corresponding $CHBr_3$ values of less than 0.15 ppt are much lower than in the tropical region. For the

266   OASIS period (July/August 2014), the global $CHBr_3$ distribution is mostly reversed compared to the

267   SHIVA period and high winds over the Indian Ocean and Maritime continent lead to low $CHBr_3$

268   abundance in this region. The North Pacific on the other hand, with low wind speeds is now a region of

269   intense accumulation leading to 0.3-0.4 ppt of $CHBr_3$. The tropical West Pacific is the only region that

270   experiences relatively low winds during both seasons, and constantly shows high $CHBr_3$ for the SHIVA

271   and OASIS time periods.

272   The variations of the background $CHBr_3$ distribution can be generally explained by the seasonal

273   variations of the global wind field. The North Pacific and Northern Indian Ocean are dominated by the

274   East Asia Monsoon and the Monsoon of South Asia, respectively. The East Asia Monsoon is

275   characterized by strong northwesterly flow in boreal winter and weak southeasterly flow in boreal

276   summer due to the reverse of the thermal gradient between land and ocean (Webster, 1987; Ding and

277   Chan, 2005). Therefore, the accumulations of $CHBr_3$ in the North Pacific occurs during the boreal

278   summer months, rather than during boreal autumn/ early winter (TransBrom time period). The Monsoon

279   of South Asia, on the other hand, is characterized by weak northeasterly winds in boreal winter and

280   strong southwesterly winds in boreal summer (Webster, 1987; Webster et al., 1998). Thus background

281    $CHBr_3$ accumulation over the Northern Indian Ocean occurs mostly during boreal winter, while during

282   boreal summer (OASIS time period) a low $CHBr_3$ background can be expected. Because of the light

283   winds of the Inter Tropical Convergence Zone (ITCZ), a belt of relatively high $CHBr_3$ abundance exists

284   along the equator in the Northern Hemisphere, especially in the tropical Pacific and Atlantic. Strong

285   convection in the ITCZ enhances vertical transport of $CHBr_3$ out of the boundary layer, but overall the

286    $CHBr_3$ distribution is dominated by the horizontal wind fields and accompanying transport patterns. Due

287   to the more complex land-sea thermal difference, the seasonal variation of ITCZ in the West Pacific is

288   more significant than in the East Pacific (Waliser and Jiang, 2014). The relatively high accumulations of

289    $CHBr_3$ in the tropical East Pacific are confined to a narrow region near the equator for both seasons. As

290   for the tropical West Pacific, during boreal winter the ITCZ covers almost the whole Southeast Asia and

291   the high $CHBr_3$ abundances during SHIVA appear along the east coast of Malaysia. During boreal

292   summer, the ITCZ shifts northward and the high $CHBr_3$ abundances retreat northwestward.

293 In the above simulations, we assume a constant background emission in order to isolate the impact of
294 transport and loss processes on the atmospheric CHBr$_3$ distribution. Variations of the wind fields will
295 likely impact the oceanic air- sea flux and emissions. Such variations can change the background CHBr$_3$
296 distribution and may allow for increased mixing ratios in regions of strong winds. In addition to the wind
297 speed, variations in the atmospheric and, more importantly, the oceanic CHBr$_3$ concentrations can impact
298 the emission strength which can further change the complex atmospheric CHBr$_3$ distribution.

299

300 **4. Atmospheric CHBr$_3$ based on hotspot emissions**

301

302 Given the high variability of the atmospheric CHBr$_3$ background mixing ratios, resulting from
303 atmospheric transport processes (Section 3), it is of interest to analyze if and how much oceanic hotspot
304 emissions might impact this background distribution. In this section, the results of simulations based on
305 observed localized hotspot emissions will be compared to the background mixing ratios.

306

307 **4.1 Observed hotspot emission**

308

309 Oceanic CHBr$_3$ emissions, atmospheric CHBr$_3$ mixing ratios and the observed local surface wind speeds
310 are given in Fig. 8 for all three campaigns. The oceanic emissions of CHBr$_3$ vary substantially from
311 campaign to campaign with mean values of 261 pmol h^{-1} m^{-2} (TransBrom), 1228 pmol h^{-1} m^{-2} (SHIVA),
312 and 912 pmol h^{-1} m^{-2} (OASIS) with standard deviations of 600 pmol h^{-1} m^{-2} (TransBrom), 1460 pmol h^{-1}
313 m^{-2} (SHIVA) and 1159 pmol h^{-1} m^{-2} (OASIS), respectively (Tegtmeier et al., 2012; Ziska et al., 2013;
314 Fuhlbrügge et al., 2016; Fiehn et al., 2017).
315 All three campaigns show periods with localized elevated and hotspot emissions. For TransBrom, the
316 first two thirds of the campaign show negative (air-to-sea) or very low CHBr$_3$ fluxes, while the last third
317 was close to western Pacific islands and is characterized by overall elevated emissions with sporadic
318 hotspots of up to 4000 pmol h^{-1} m^{-2}. The SHIVA cruise track was mostly along the coastline, where
319 elevated emissions and hotspots occurred regularly. The OASIS cruise track alternated between open
320 ocean, upwelling and coastal areas, resulting in a large fluctuation between low background and localized
321 elevated emissions. The largest hotspot emissions were observed during this campaign, reaching values
322 of over 6000 pmol h^{-1} m^{-2}.
323 According to the flux parameterization applied here, the variability of air-sea flux is determined mostly
324 by the surface wind speed and the ocean-atmosphere concentration gradient. Highest emissions are
325 expected to occur during periods of high wind speed and large concentration gradients. During the
326 beginning of the TransBrom campaign (Fig. 8a), the wind speed peaks at over 15m/sec while the
327 corresponding CHBr$_3$ air-sea flux is low. Higher wind speeds co-occur with high air-sea fluxes at the
328 end of the campaign. For SHIVA (Fig. 8b) and OASIS (Fig. 8c), the relation between wind speed and
329 CHBr$_3$ emissions is more easily discernable.

330 All three campaigns demonstrate that high fluxes do not always lead to local high CHBr$_3$ mixing ratios

331 in the surface atmosphere. For example, several hotspots with oceanic emissions over 4000 pmol m^{-2} hr^{-}

332 ^{1} are found during OASIS, however, corresponding atmospheric mixing ratios are relatively low ($\sim$ 2

333 ppt). Vice versa, the highest atmospheric mixing ratios found during OASIS only coincide with high

334 fluxes during the last part of the campaign. These discrepancies underline the complex interplay of source,

335 transport and loss processes on the local atmospheric mixing ratios of short-lived compounds. A

336 relatively clear connection between elevated oceanic emissions and surface mixing ratios only occurs

337 during the SHIVA campaign and during the last part of the TransBrom campaign (Fig 8a and b).

338 The question arises how much of the atmospheric variability of short-lived compounds such as CHBr$_3$ is

339 impacted by the emission strengths and is addressed in the subsequent section based on the model results.

340

341 **4.2 Comparison of CHBr$_3$ from background and hotspot emissions**

342

343 In this section, we compare the concentrations of CHBr$_3$ due to background and localized elevated

344 emissions as simulated by FLEXPART. Atmospheric background and hotspot CHBr$_3$ at different

345 altitudes is simulated by FLEXPART, which is driven by the meteorological data from ECMWF. The

346 signatures of dynamical processes such as wind regimes, weather phenomena (e.g., typhoons) and

347 convection are captured by the model simulation and can be detected in the CHBr$_3$ distribution (Fig. 9).

348 For example, during the TransBrom campaign, the cruise encountered several tropical storms in the

349 western Pacific, one of which (Lupit, around October 14th, 2009) developed into a super typhoon within

350 several days (Krüger and Quack, 2013). As shown in Fig. 9, CHBr$_3$ accumulation representing the

351 structure of typhoon Lupit is clearly visible in the background distribution of CHBr$_3$ at 500 m (Fig. 9d)

352 altitude in the northern part of the western Pacific. This structure is still clear at 5 km altitude (Fig. 9a),

353 although with a weaker magnitude. Localized elevated sources of CHBr$_3$ (Figures 9b, c, e, and f) do not

354 add much due to the small spatial extent of the 0.5° or 1° emission cells and thus the limited amount of

355 overall released CHBr$_3$ is not discernible in the large scale structures. Higher abundances of atmospheric

356 CHBr$_3$ can be seen in the southern part of the western Pacific near Indonesia resulting from one of the

357 hotspot emissions observed during TransBrom (Fig. 1b). However, the background CHBr$_3$ in this area is

358 also high in this low-wind area, and thus the atmospheric signal of the up to 20 times stronger hotspot

359 emissions (Fig. 8a) is detectable for neither the 0.5° nor the 1° wide emission cells when compared to

360 the background. Note that the modelled atmospheric mixing ratios from both sources, hotspot and

361 background emissions, are smaller than the mixing ratios observed along the cruise track (Fig. 8)

362 suggesting stronger nearby emissions not covered in our scenarios and observations. The signature of

363 the hotspot emissions cannot be seen at 5 km altitude.

364 Fig. 10 shows the atmospheric CHBr$_3$ mixing ratios during the SHIVA campaign. The atmospheric signal

365 of the localized elevated emissions during SHIVA is much stronger than during TransBrom due to

366 stronger emissions and smaller background mixing ratios. First, for the 0.5° wide emission grids, two

367 highly localized, atmospheric CHBr$_3$ peaks appear close to the coast line near the equator around 105º E

368 with a maximum value around 0.4 ppt. These signals occur in a spot where the background is very low

369 (0.2 ppt). However, at the same time they are smaller than the maximum background values of up to 0.5

370 ppt in nearby regions (Fig. 10d). If the width of the emission grids is extended to 1°, the localized CHBr$_3$

371 peaks mentioned above grow into two distinct blobs near the equator of up to 0.8 ppt, which are

372 apparently larger than the regional background concentrations (Fig. 10f). Elevated emissions during the

373 second half of the campaign with several hotspot events, on the other hand, do not show such clear

374 atmospheric signals right above.

375 For the regions of localized elevated emissions, the convection is less effective and maximum mixing

376 ratios at 5 km are about 50% smaller compared to the values in the boundary layer. Thus only the signal

377 of the 1° wide emission cells can be detected at 5 km, while assuming that the emissions cover a smaller

378 region of 0.5° width will render their impact in the free troposphere negligible. Krystofiak et al. (2018)

379 calculated the fractions of convective-contributed trace gases from boundary layer to the upper

380 troposphere using airborne measurements during the SHIVA campaign and reported an even smaller

381 fraction of boundary layer CHBr$_3$ in the upper troposphere (about 15% due to convection).

382 Due to the dominant southwest monsoon over the Northern Indian Ocean in boreal summer, the resulting

383 atmospheric abundances of the OASIS case (Fig. 11) for both scenarios, background and localized

384 emissions, are much lower than for the other two campaigns. This is particularly surprising for the OASIS

385 hotspot emissions, which are in many cases larger than hotspot emissions during TransBrom or SHIVA.

386 In the open ocean, the atmospheric enhanced CHBr$_3$ mixing ratios resulting from the 0.5° (1°) wide

387 localized emission runs reach only 0.1 (0.2) ppt in a narrow belt near 60°E and are mostly smaller than

388 the background (around 0.15 ppt). An exception occurs near the coast of Madagascar, where both

389 background and hotspot emissions accumulate in the atmosphere. Maximum background values reach

390 up to 0.25 ppt and the hotspot signals peak with values of 0.3 ppt (0.5° wide emission cells) to 0.6 ppt

391 (1° wide emissions cells). These clear atmospheric signals of hotspot emissions are driven by the

392 enhanced coastal emissions near Madagascar. At 5 km altitude, atmospheric background values are very

393 low, and the hotspot contributions are close to zero.

394 In summary, although the observed emissions during the three cruises were significantly higher than the

395 background of 100 pmol m^{-2} hr^{-1}, our results show that such strong oceanic sources are not necessarily

396 detectable in the atmosphere, where transport processes can sometimes mask the impact of oceanic

397 emissions.

398

399 **5. Summary and Discussion**

400

401 In this study, we simulated atmospheric CHBr$_3$ abundances that result from uniform marine background

402 emissions compared to hotspot emissions using the Lagrangian dispersion model FLEXPART.

The simulations demonstrate that uniform background emissions from the ocean result in a highly variable atmospheric $CHBr_3$ distribution with high mixing ratios taking place in regions of convergence or low wind speed. This relation holds on regional and global scales underling the role of atmospheric transport processes as drivers of the distribution of short-lived trace gases with lifetimes in the range of days to weeks. The relation between atmospheric background and wind patterns described here will allow us to better predict the seasonal and regional characteristics of the tropospheric $CHBr_3$ distribution. Such knowledge will provide valuable information for analyzing and interpreting atmospheric data from ship and aircraft campaigns. For example, our results illustrate that elevated or low atmospheric $CHBr_3$ abundances cannot necessarily be used to draw conclusions about the oceanic source strength below.

Comparisons between atmospheric $CHBr_3$ resulting from background and peak emissions suggest that the impact of localized elevated emissions on the atmospheric $CHBr_3$ distribution depends on their relative strength, on their location, and on the time of emission. The "visibility" of elevated emissions in the atmospheric $CHBr_3$ distribution varies significantly between three cruises in the West Pacific and Indian Ocean. In the open ocean, signals of elevated emissions can hardly be distinguished from the background $CHBr_3$ distribution even for elevated sources extending over 1° wide source regions along the cruise tracks. Near the coast, however, signals of elevated emissions are often stronger to be distinguished from the background, in particular, hotspot emissions up to 100 times larger than the background. However, individual cases show that it is not necessarily the largest hotspot that gives a clear signal, but that the tug of war between fast advective transport and local accumulation at the time of emission is also important.

Our approach requires that we isolate uniform background $CHBr_3$ emission from coastal and shelf emissions, which can be significant (Fuhlbrügge et al., 2016; Fiehn et al., 2017) and would lead to higher atmospheric abundances. In consequence, we expect the background $CHBr_3$ mixing ratios inferred from our simulations to be smaller compared to observations and other modeling studies. In the Western Pacific (TransBrom), our simulated background mixing ratios at 5 km range from 0.0-0.4 ppt (Fig. 9-11). Measurements from aircraft campaigns in this region, CAST (Harris et al., 2016) and CONTRAST (Pan et al., 2016), show higher $CHBr_3$ mixing ratio of 0.03-0.79 ppt and 0.20-1.127 ppt between 4-6 km. Other model studies (e.g. Hossaini et al., 2016; Butler et al., 2018) based on $CHBr_3$ emission scenarios that include coastal and open ocean sources (e.g. Liang et al., 2010; Ordóñez et al., 2012; Ziska et al., 2013) also suggest the average $CHBr_3$ mixing ratio over 0.5 ppt in this region.

The constant background emissions of 100 pmol m^{-2} hr^{-1} used in our study are based on a simplified scenario and do not include coastal emissions. Nevertheless, our results demonstrate that atmospheric $CHBr_3$ signals, produced by localized elevated and even hotspot emissions, orders of magnitudes larger than the average open ocean emissions, can be obliterated by the highly variable atmospheric background. That is to say that transport variations of the atmosphere itself are sufficient to allow high concentrations in certain regions and that high concentrations of VSLH in the atmosphere do not always guarantee a strong local or regional source. For observational and modelling studies of VSLS and other short-lived

440 compounds, the impact of atmospheric transport patterns that are identified here can be used for the

441 interpretation of trace gas distributions and variability.

Data availability

The emission data of cruise campaigns are available at Pangaea (http://www.pangaea.de). FLEXPART output can be inquired from the authors.

Author contribution

Y. Jia and S. Tegtmeier designed the model experiments. Y. Jia carried out the FLEXPART calculations and produced the figures. Y. Jia and S. Tegtmeier wrote the manuscript with contributions and revisions from the co-authors B. Quack and E. Atlas.

Competing interests

The authors declare that they have no conflict of interest.

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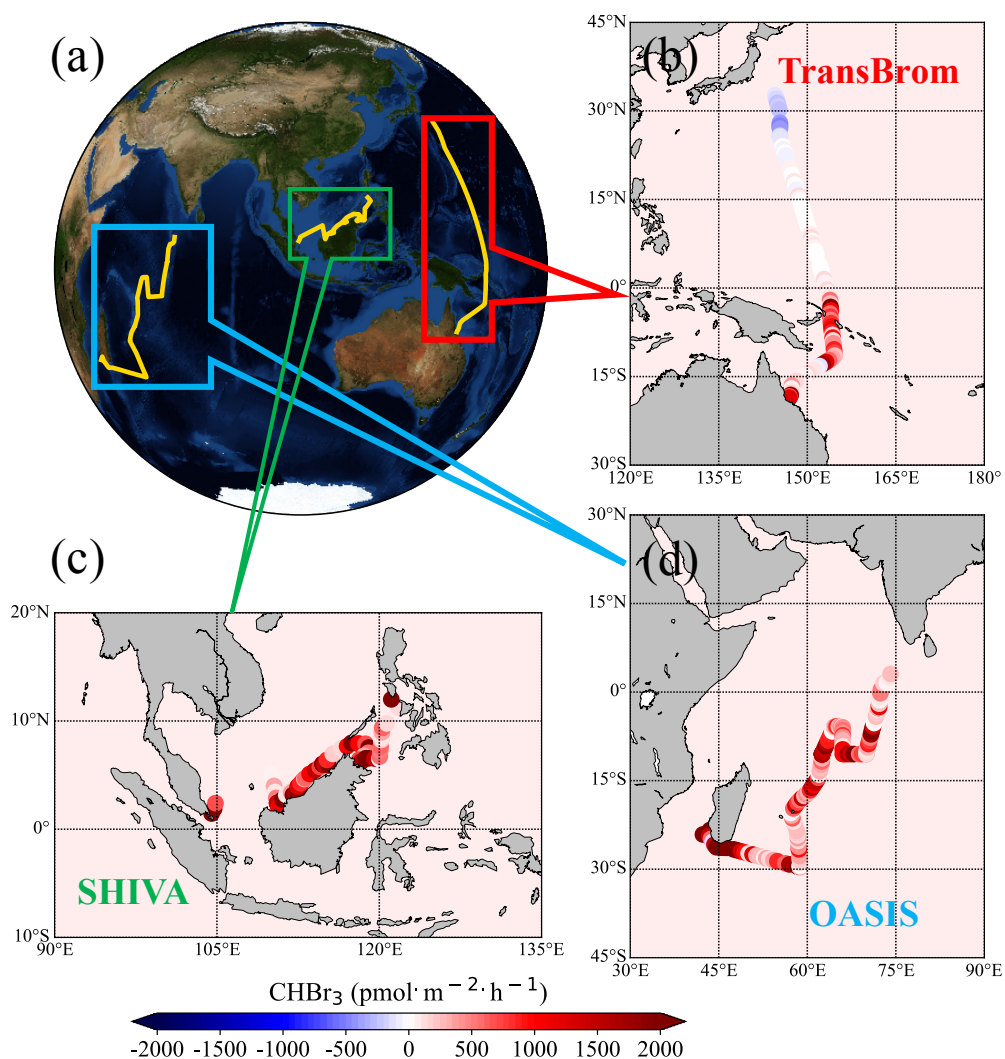
668 **Fig 1.** Cruise tracks of the three campaigns in the Indian Ocean and Western Pacific (a)

669 and CHBr$_3$ emissions (b, c, d) used in the model simulation. Global background

670 emissions (100 pmol m^{-2} hr^{-1}) and observed emissions along the tracks of the three

671 research cruises TransBrom (b), SHIVA (c), and OASIS (d).

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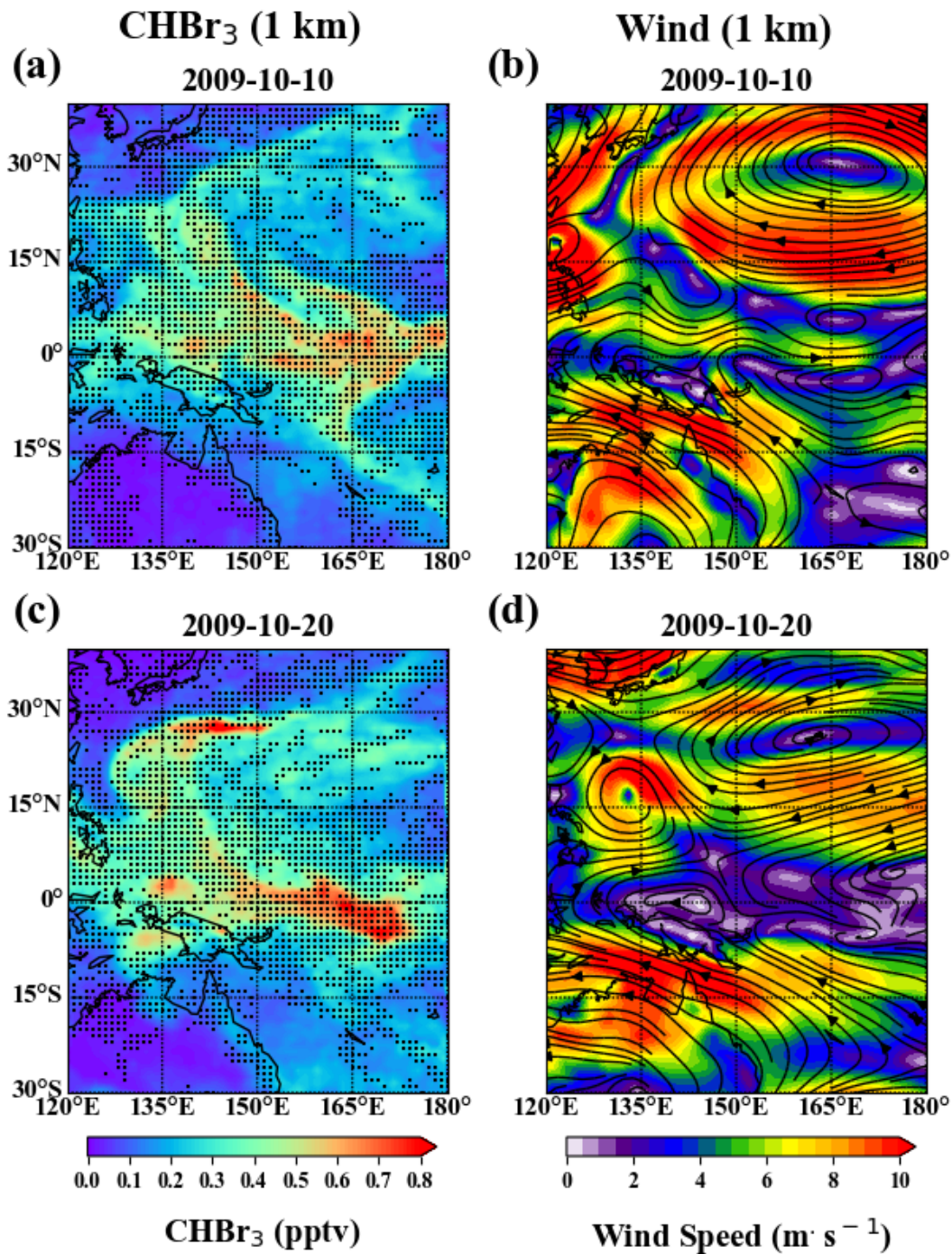
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678 **Fig. 2.** Two snapshots of spatial distributions of atmospheric $CHBr_3$, derived from
679 uniform oceanic background emissions of 100 pmol m^{-2} hr^{-1} (a, c), and ERA-Interim
680 reanalysis wind fields (b, d) at 1 km altitude during TransBrom. The wind speed is
681 denoted by color shades and the directions are denoted by the stream lines. The regions
682 of convergence are shaded in (a) and (c).
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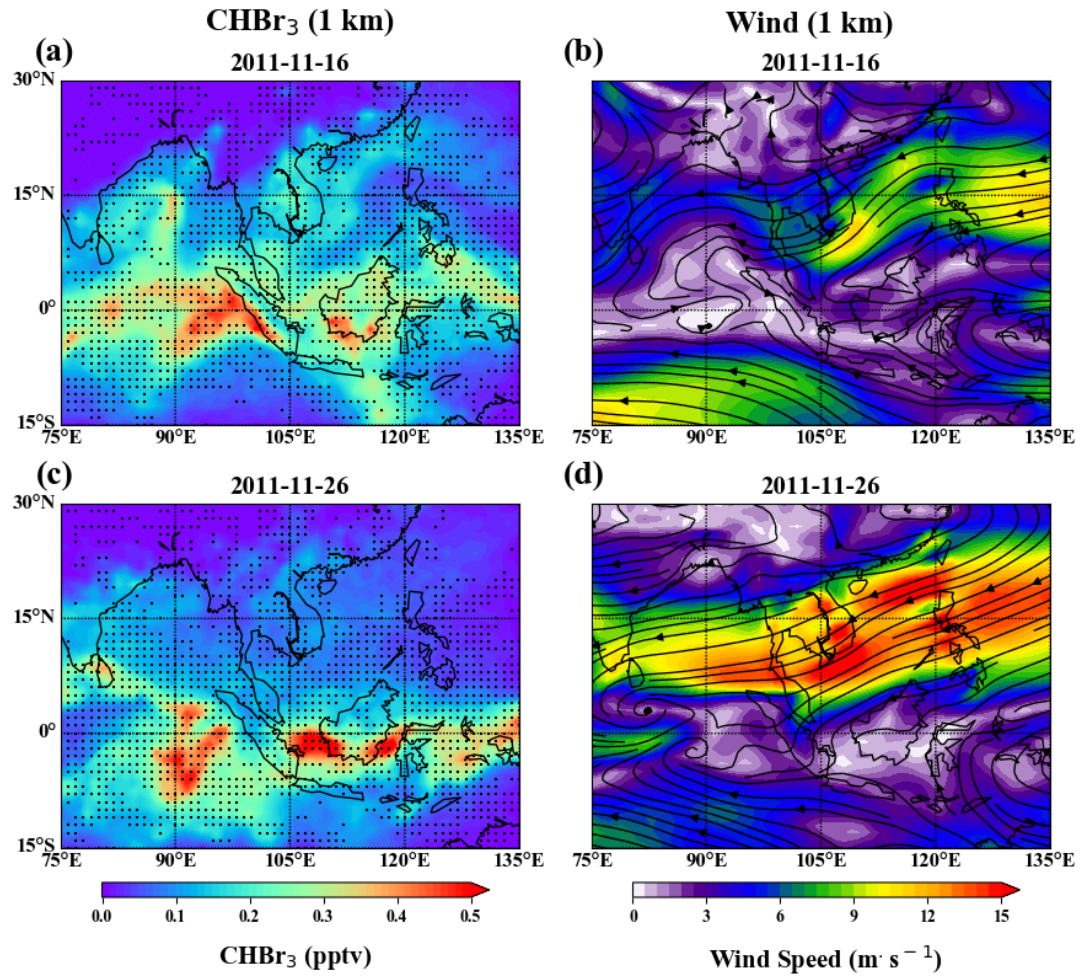
686   **Fig. 3.** Same as Fig. 2, but for SHIVA case.

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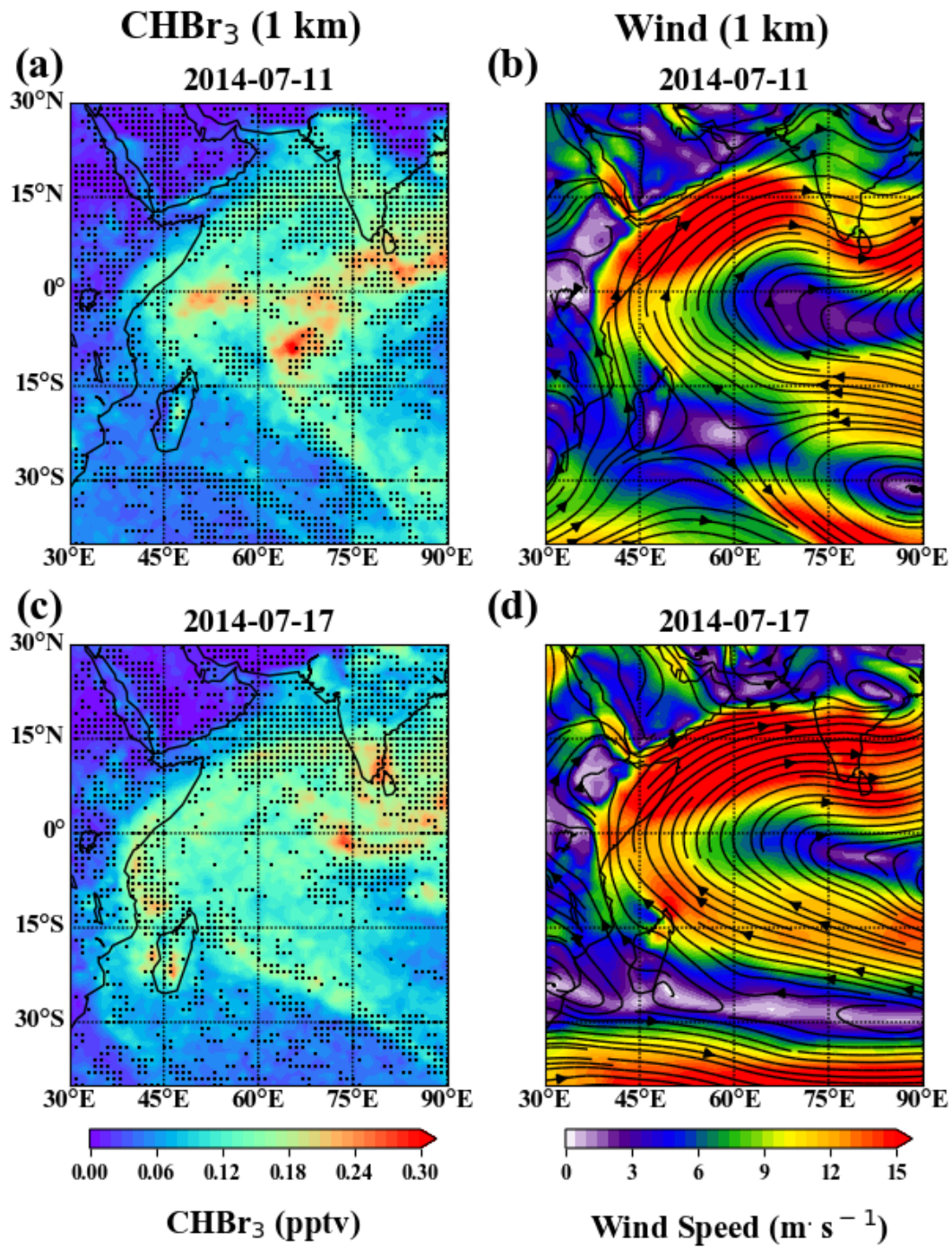
688   **Fig. 4.** Same as Fig. 2, but for OASIS case.

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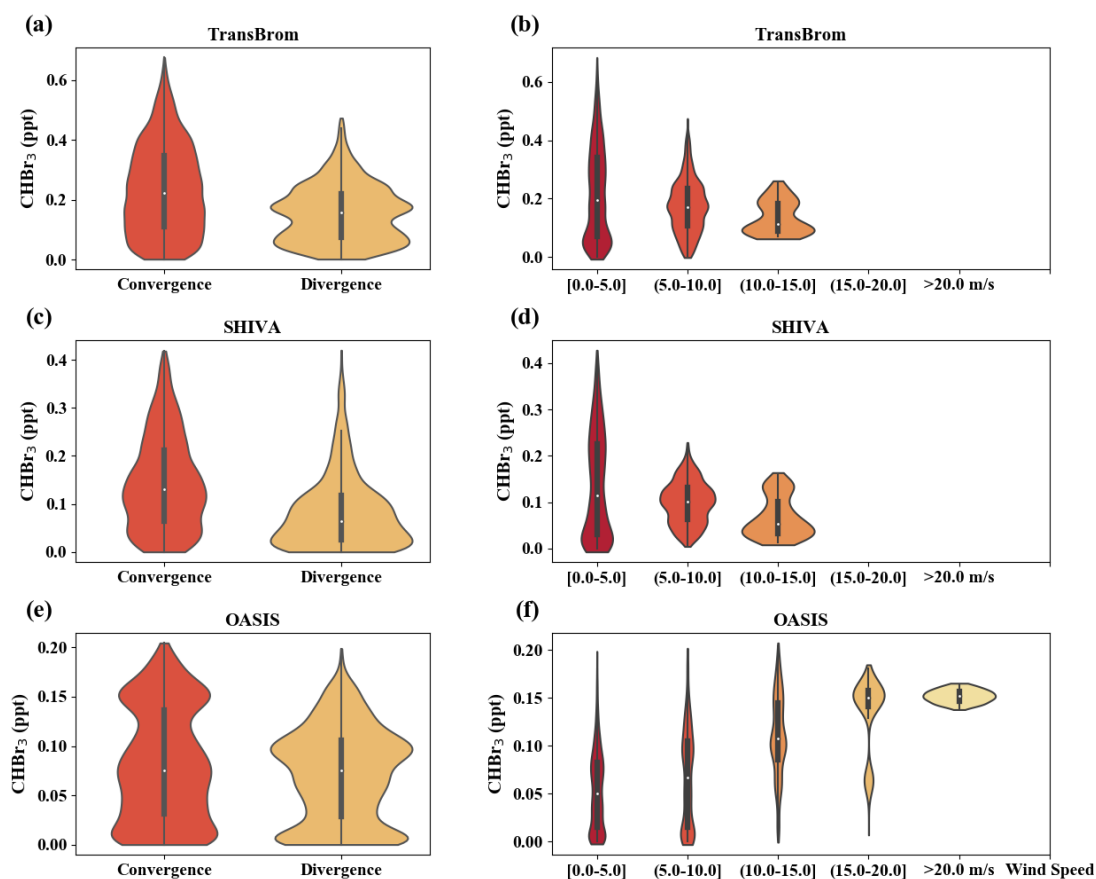
Fig.5. Violin plots of regional distributions of simulated background CHBr$_3$ mixing ratio by convergence and divergence (a, c, e), and by wind speed (b, d, f) at 1 km altitude averaged over TransBrom (a, b), SHIVA (c, d), and OASIS (e, f) period. The violin is a corresponding density plot with a boxplot inside. The white dots represent the medians. The thick black bars in the center represent the interquartile ranges. The thin black lines represent the rest of the distributions, except for the outliers.

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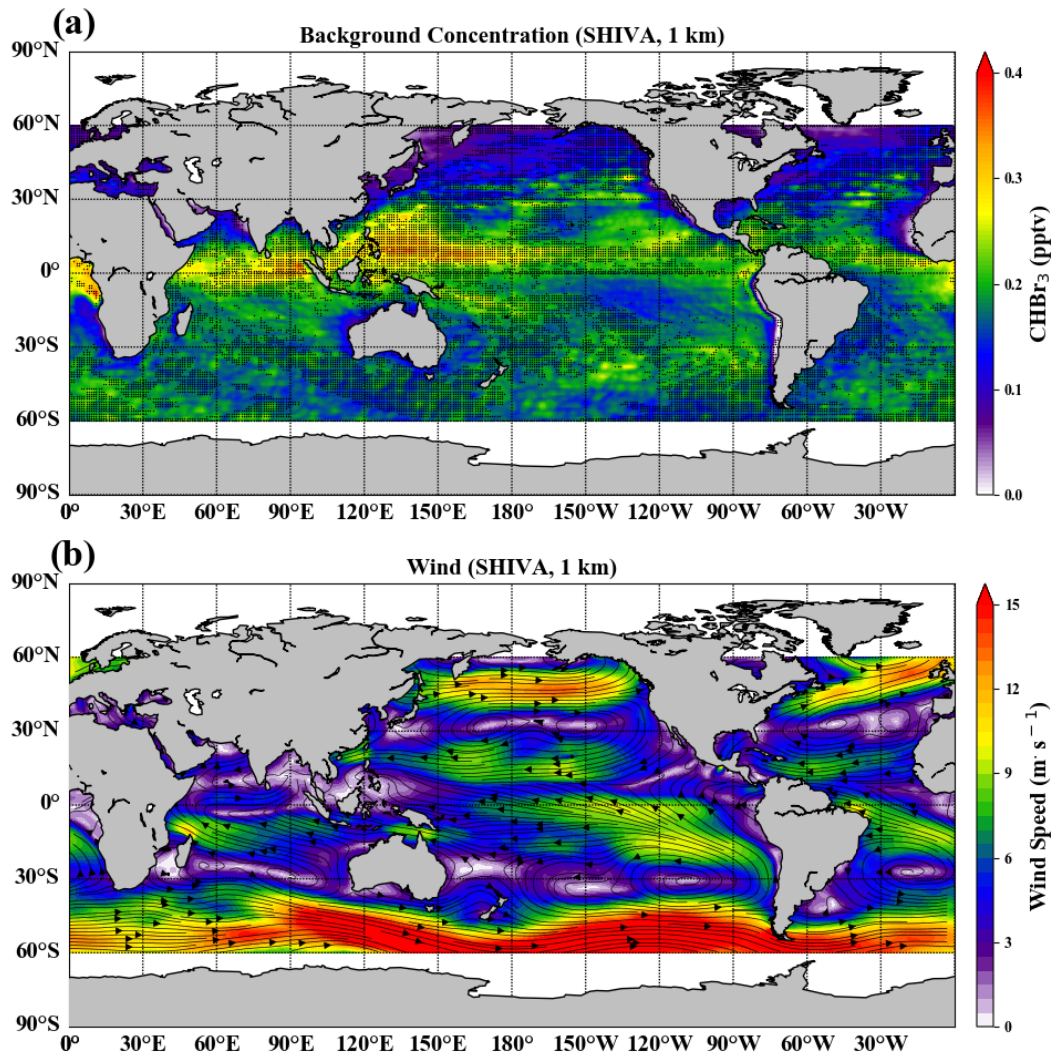
Fig. 6. Global distributions of CHBr$_3$ mixing ratios based on oceanic background
emissions (a), and ERA-Interim reanalysis wind fields (b) averaged during the time
period of the SHIVA cruise at 1 km. The wind speeds are denoted by color shades and
the directions are denoted by the stream lines. The regions of convergence are shaded
in (a).

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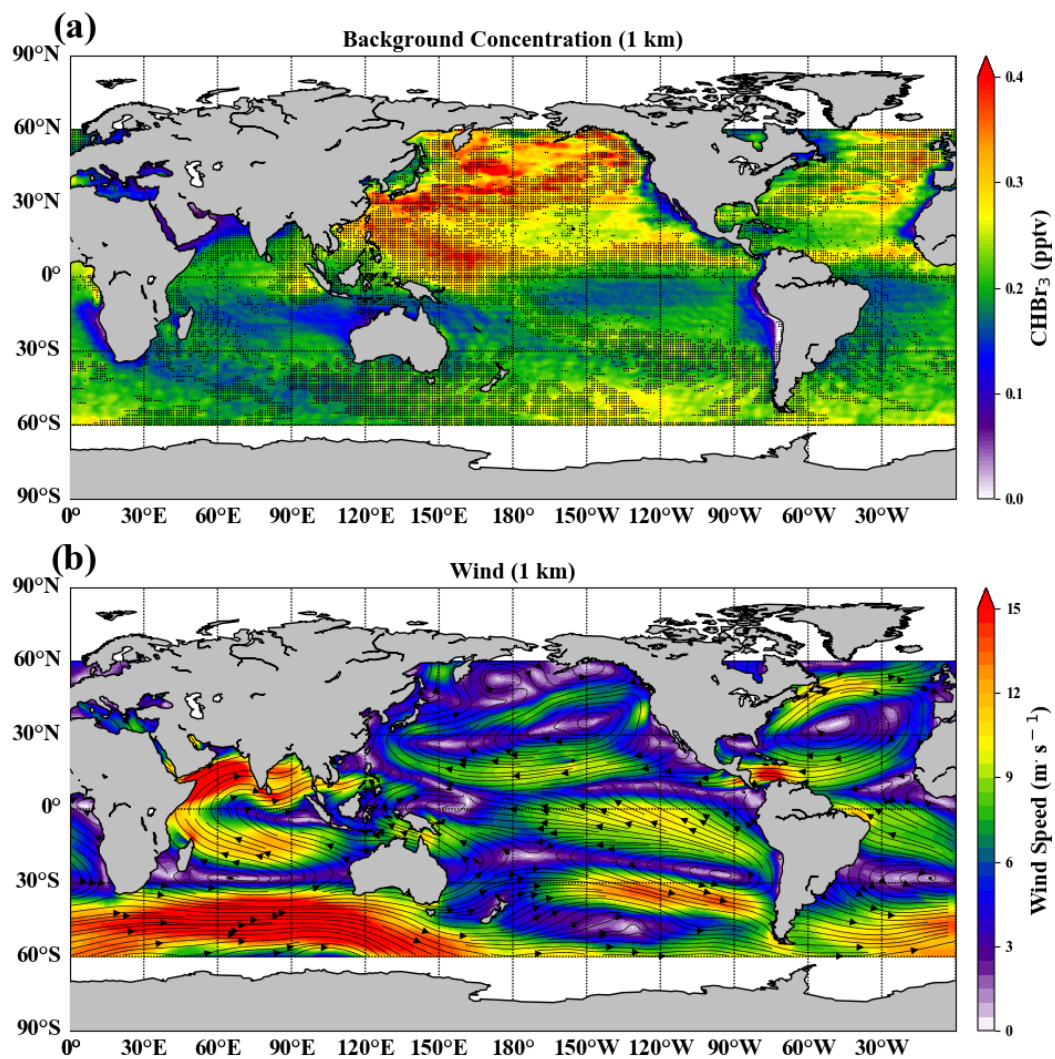
706   **Fig. 7.** Same as Fig. 6 but for OASIS case.

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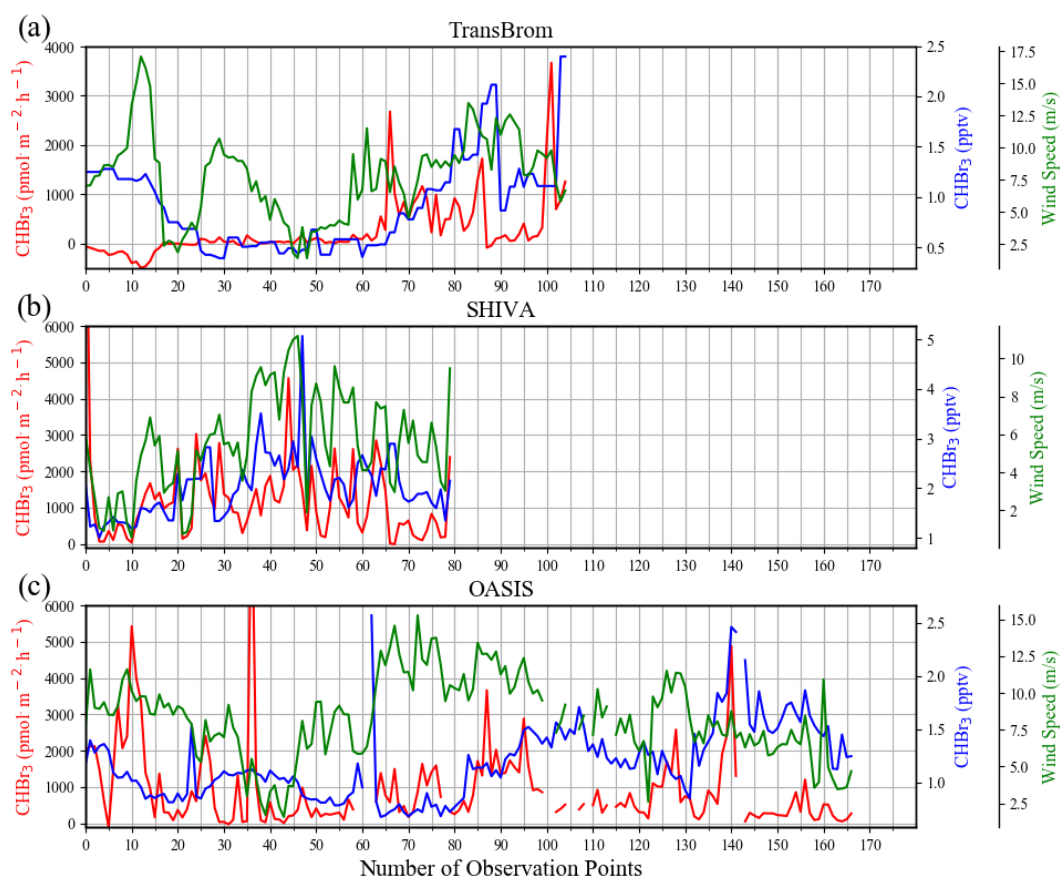
Fig. 8. Surface wind speeds (green), $CHBr_3$ air-sea flux (red), and atmospheric mixing

ratios of $CHBr_3$ near surface (blue) observed during TransBrom, SHIVA, and OASIS

campaigns.

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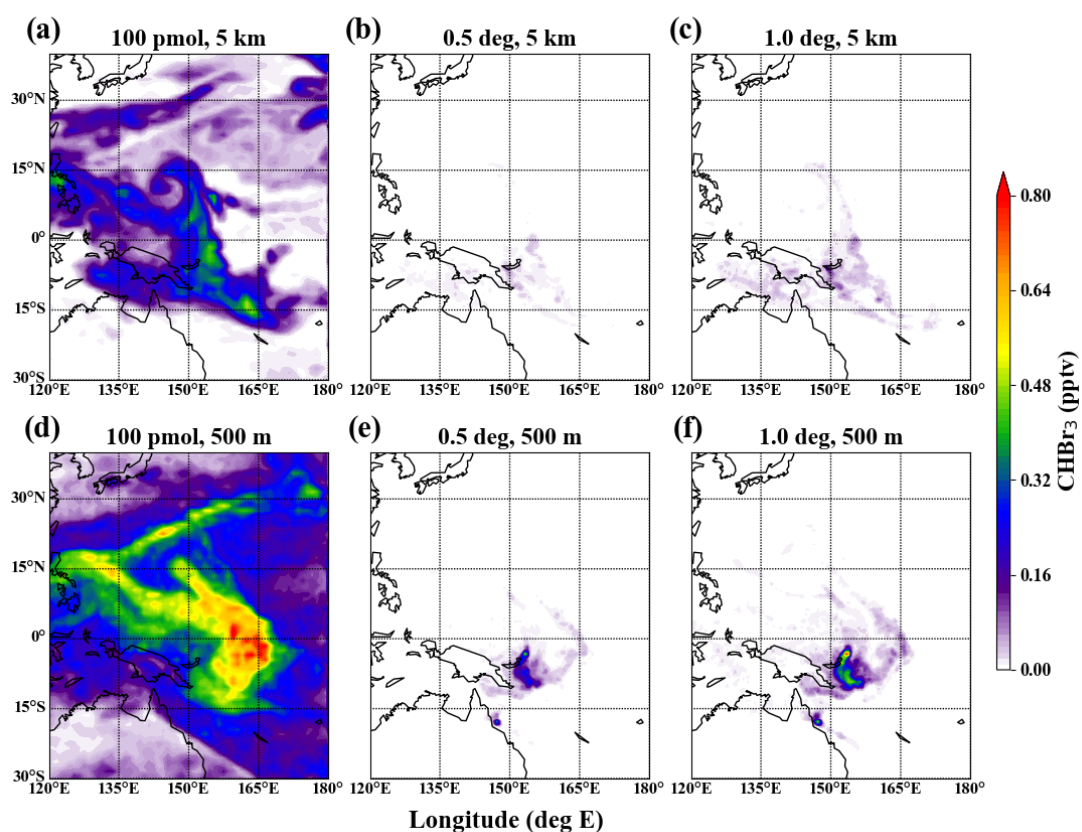
Fig. 9. Atmospheric CHBr$_3$ mixing ratios at different altitudes (500 m and 5 km) simulated for the time period of the TransBrom campaign. Simulations are based on background emissions (a, d), and elevated emissions observed during the campaign for 0.5° (b, e) and 1° (c, f) wide emission grid cells.

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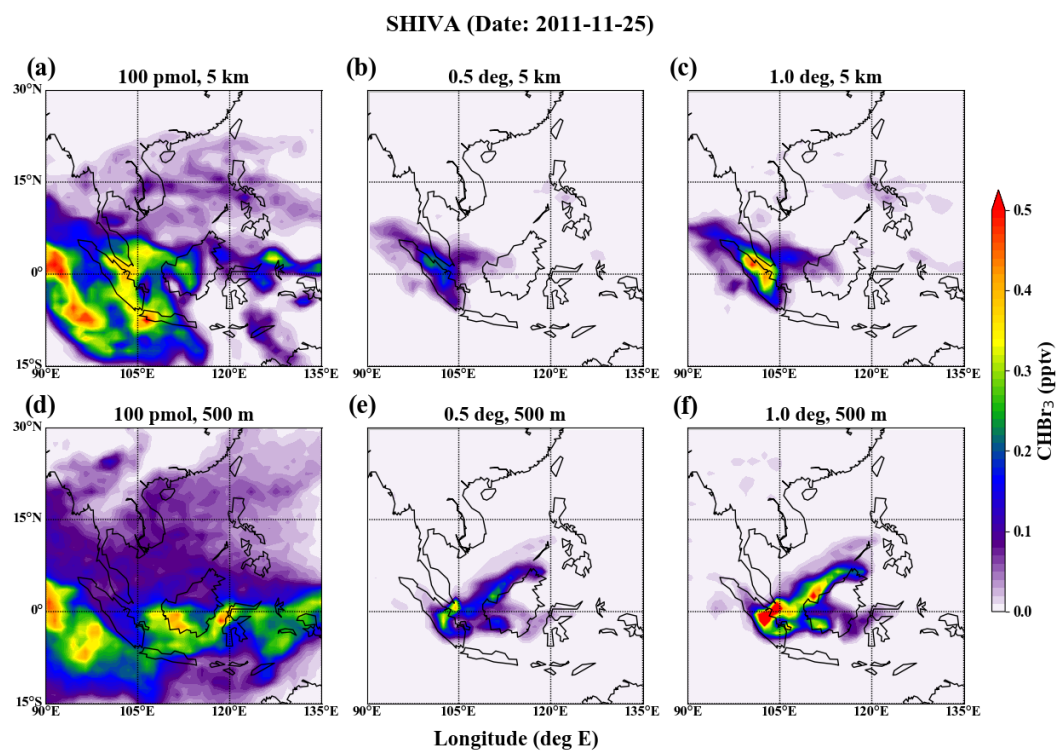
721 **Fig. 10.** Same as Fig. 9, but for the SHIVA campaign.

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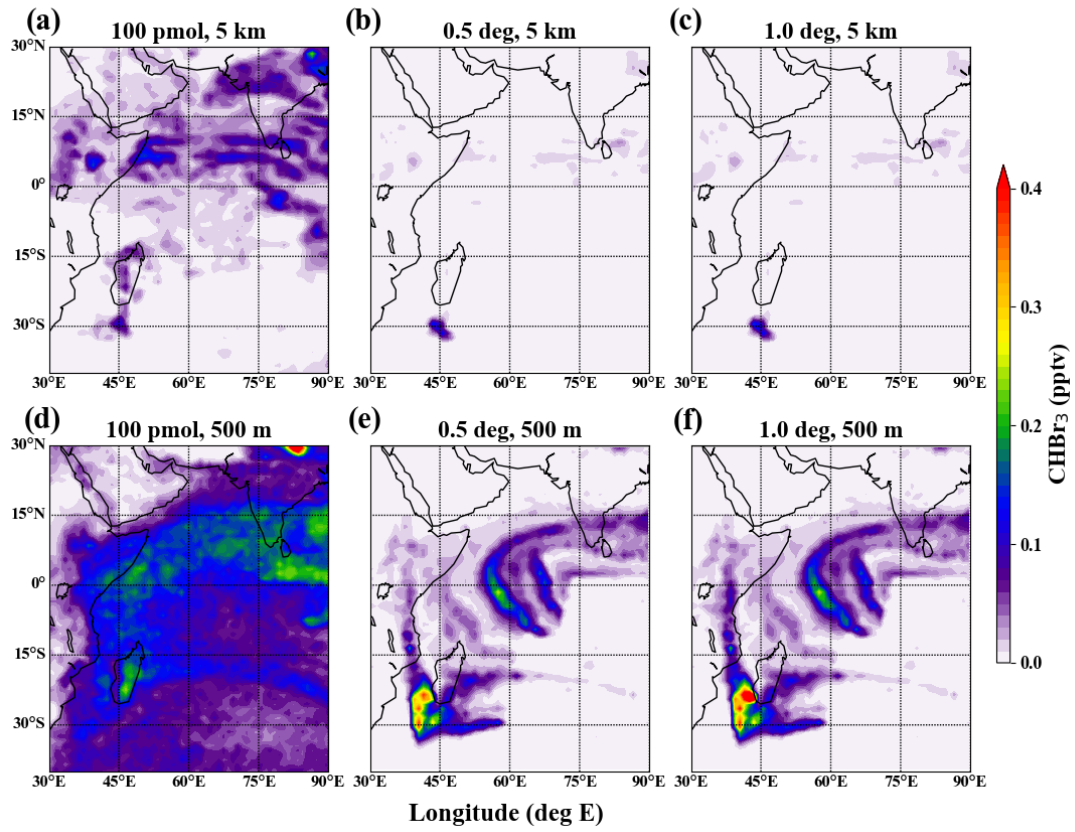
Fig. 11. Same as Fig. 9, but for the OASIS campaign.