



Multivariate analysis of Kelvin wave seasonal variability in ECMWF L91 analyses

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1 **Abstract.** The paper presents the seasonal variability of Kelvin waves (KWs) in 2007-2013 ECMWF analyses on 91 model
2 levels. The waves are filtered using the normal-mode function decomposition which simultaneously analyses wind and mass
3 field based on their relationships from linear wave theory. Both spectral as well as spatiotemporal features of the KWs are
4 examined in terms of their seasonal variability in comparison with background wind and stability. Furthermore, a differentiation
5 is made using spectral bandpass filtering between the slow horizontal barotropic KW response and the fast vertical projection
6 response observed as vertically-propagating KWs.

7 Results show a clear seasonal cycle in KW activity which is predominantly at the largest zonal scales (wavenumber 1-2)
8 where up to 50% more energy is observed during the solstice seasons in comparison with spring and autumn. The spatiotempo-
9 ral structure of the KW reveals the slow response as a robust "Gill-type" structure with its position determined by the location
10 of the dominant convective outflow winds throughout the seasons. Its maximum strength occurs during northern summer when
11 easterlies in the Eastern Hemisphere are strongest. The fast response in the form of free traveling KWs occur throughout the
12 year with seasonal variability mostly found in the wave amplitudes being dependent on background easterly winds.



13 1 Introduction

14 Atmospheric equatorial Kelvin waves (hereafter KWs), first discovered in the stratosphere (Wallace and Gousky, 1968), are
15 nowadays observed and studied over a broad range of spatial and temporal scales. A broad wavenumber-frequency spectrum
16 can be traced to the spatiotemporal nature of tropical convection which generates KWs along with a spectrum of other equatorial
17 waves. Atmospheric wave response to the stochastic nature of convection was studied by Garcia and Salby (1987) and Salby
18 and Garcia (1987) who made a distinction between (i) projection or vertical response to short-term heating fluctuations (e.g.
19 daily convection) and (ii) barotropic or horizontal response to seasonal convective heating. For KWs, the vertical response
20 gives rise to a broad frequency spectrum of vertically propagating KWs that radiate outward into the stratosphere where
21 they drive zonal-mean quasi-periodic flows such as the quasi-biennial oscillation (QBO, Holton and Lindzen, 1972). The
22 horizontal response to seasonal transitions in convective heating gives rise to planetary-scale disturbances with a half-sinusoidal
23 vertical structure confined to the troposphere. A part of this response remains stationary over the convective hotspot; its shape
24 resembling a classic "Gill-type" KW solution (Gill, 1980). The other part of the response intensifies and advances over the
25 Pacific, representing a transient component of the Walker circulation (Salby and Garcia, 1987).

26 Both components of the KW response received increased attention in the scientific community over the last decades in
27 terms of the role they play in the (intra)seasonal variability of the Tropical Tropopause Layer (hereafter TTL), defined as
28 a transition layer between the typical level of convective outflow at ~ 12 km where the Brunt-Väisälä frequency is at its
29 minimum, and the cold point tropopause at ~ 16 -17 km (Highwood and Hoskins, 1998; Fueglistaler et al., 2009). Within the
30 TTL, temperature variations play an important role in controlling the stratosphere-troposphere exchange of various species
31 such as ozone and water vapour thereby aiding in the dehydration process of air entering the stratosphere. The two parts of the
32 KW response alternate the TTL differently on different time scales (Highwood and Hoskins, 1998; Randel and Wu, 2005; Ryu
33 et al., 2008; Flannaghan and Fueglistaler, 2013); their relative contribution to TTL dynamics varies with season and is not yet
34 fully understood. The present study contributes to this topic by applying a novel multivariate analysis of Kelvin wave seasonal
35 variability in model-level analysis data.

36 Seasonal variations of Kelvin wave dynamics in the TTL have been previously studied using temperature data derived from
37 satellites such as SABER (Sounding of the Atmosphere using Broadband Emission Radiometry, Garcia et al., 2005; Ern et al.,
38 2008; Ern and Preusse, 2009), HIRDLS (High Resolution Dynamics Limb Sounder, Alexander and Ortland, 2010), and GPS-
39 RO (Global Positioning System Radio Occultation, Tsai et al., 2004; Randel and Wu, 2005; Ratnam et al., 2006). For example,
40 Alexander and Ortland (2010) reported a clear seasonal cycle around 16-17 km (~ 100 hPa) in KW temperature observed by
41 HIRDLS, coinciding closely with variations in background stability. A widely used method for the KW filtering from gridded
42 data is the space-time spectral analysis introduced by Hayashi (1982). It operates on single variable data and it has been widely
43 used to diagnose equatorial waves in the outgoing longwave radiation (OLR, e.g. Wheeler and Kiladis, 1999) and climate
44 model outputs (e.g. Lin and Coauthors, 2006). Based on 40-year ECMWF reanalysis (ERA-40) data, Suzuki and Shiotani
45 (2008) found that the temperature component of Kelvin waves tends to peak at 70 hPa while the zonal wind peaks at lower
46 altitudes, i.e. at 100 hPa (150 hPa) in Eastern (Western) hemisphere.



47 The zonal wind and geopotential height of the KW are closely related. For a single zonal wavenumber k , the geopotential,
48 Φ_{kw} , and the zonal wind U_{kw} of a zonally propagating KW are related according to the following equation:

$$49 \quad \Phi_{kw} = g h_{kw} = \frac{\nu}{k} U_{kw}, \quad \text{where} \quad U_{kw} = U_0 \exp\left(-\frac{\beta k y^2}{2\nu}\right). \quad (1)$$

50 Here, U_0 is the KW amplitude in zonal wind on the equator, $\beta = 2\Omega/a$ (Ω being the rotation rate and a the radius of Earth), ν
51 is the wave frequency, g is gravity and y is the distance from the equator. These expressions are obtained as a special solution
52 of the linearized shallow-water equations on the equatorial β -plane (e.g. Holton, 2004, Chapter 11). The KW e -folding
53 decay width, a_e , is given by $a_e = (c/2\beta)^{1/2}$, where the KW phase speed c is determined from the dispersion relation $\nu = kc$.
54 By prescribing the value of KW phase speed c (i.e. the equivalent depth of the shallow-water equation system), analytical
55 solutions from linear wave theory can be used to simultaneously analyze wind and height data of the KW wave on a single
56 horizontal level. Such multivariate analysis was carried out by Tindall et al. (2006) who analyzed several levels the ECMWF
57 15-year reanalysis dataset (ERA-15) in the lower stratosphere. They reported a maximum of Kelvin wave activity at 100 hPa
58 around the solstices when tropical cloud activity maximizes. For the ERA-15 data in 1981-93 period, their Kelvin wave analysis
59 explained approximately 1 K² of temperature variance on the equator at 100 hPa.

60 The present paper extends the use of linear wave theory from the equatorial β -plane to the three-dimensional (3D) spherical
61 coordinates in order to analyze KW wind and temperature fields in recent ECMWF operational analyses. We focus on seasonal
62 variability of KWs in the TTL layer in the ECMWF operational analyses during a period when the model employed 91 vertical
63 level (L91) between the surface and 1 Pa. The L91 model was in operations between 2006 and early summer 2013 when it
64 was replaced by 137 levels. This study thus explores most of information on the vertical structure of KWs available in the
65 L91 analysis data. We present a methodology for the simultaneous analysis of wind and temperature perturbations associated
66 with KWs with respect to the background state and apply it to quantify scale-dependent seasonal KW variability in several
67 frequency bands.

68 The paper consists of five sections. Methodology of the KW diagnosis and the data are presented in section 2. Section 3
69 presents the KW energetics in wavenumber space focusing on the seasonal cycle. Section 4 presents a 3D view on KWs in
70 L91 dataset, both for the horizontal as well as for the vertical projection KW response. Conclusions and outlook are given in
71 section 5.

72 2 Data and methodology

73 The Kelvin waves are filtered using the Normal-Mode Function (NMF) decomposition derived by Kasahara and Puri (1981) and
74 briefly summarized below. Input ECMWF operational analyses cover 6 years from January 2007 till June 2013, approximately
75 6.5 years. The dataset starts after two important updates in the ECMWF assimilation cycle: a resolution update on 1 February
76 2006 and the introduction of GPS-RO temperature profiles in the assimilation on 12 December 2006. The data ends at the
77 next update in vertical resolution from L91 to L137 on 25 June 2013. A case study of the large-scale KW in July 2007 (Žagar



78 et al., 2009) showed how the NMF method provides information on the horizontal and vertical wave structure and its vertical
79 propagation in the stratosphere.

80 2.1 Filtering of Kelvin waves by 3D normal-mode function expansion

81 The basic assumption behind the NMF expansion is that a global state of the atmosphere described by its mass and wind vari-
82 ables at any time can be considered as a superposition of the linear wave solutions upon a predefined background state. These
83 linear solutions describe two types of wave motions: Rossby waves and inertio-gravity waves which obey their corresponding
84 dispersion relationships. The associated eigensolutions in terms of the Hough harmonics define both mass and wind fields of
85 the waves. The linear wave theory approach has been successfully employed in many studies, especially for the large-scale
86 tropical circulation features (e.g. Gill, 1980; Salby and Garcia, 1987; Garcia and Salby, 1987)

87 The NMF decomposition derived by Kasahara and Puri (1981) uses the σ coordinates and a realistic vertical temperature
88 and stability stratification. 3D wave solutions of primitive equations linearized around the state of rest are represented as a
89 truncated time series of the Hough harmonic oscillations and the vertical structure functions. The expansion of a global input
90 data vector $\mathbf{X}(\lambda, \varphi, \sigma) = (u, v, h)^T$ can be represented by a discrete finite series as:

$$91 \begin{pmatrix} u(\lambda, \varphi, \sigma) \\ v(\lambda, \varphi, \sigma) \\ h(\lambda, \varphi, \sigma) \end{pmatrix} = \sum_{m=1}^M \mathbf{S}_m \left[\sum_{n=1}^R \sum_{k=-K}^K \chi_n^k(m) \mathbf{H}_n^k(\lambda, \varphi; m) \right] G_m(\sigma) \quad (2)$$

92 The zonal and vertical truncations (K and M , respectively) define maximum numbers of zonal waves at a single latitude
93 (wavenumber k) and a maximal number of vertical modes denoted m respectively. Parameter R is the total number of merid-
94 ional modes which is a sum of the eastward inertio-gravity waves (EIG), westward inertio-gravity waves (WIG) and Rossby
95 waves. Oscillations in the horizontal plane are given in terms of Hough harmonic functions, $\mathbf{H}_n^k(\lambda, \varphi)$ for every vertical struc-
96 ture eigenfunctions $G_m(\sigma)$. The horizontal and vertical solutions are connected by the equivalent depth parameter D_m , which
97 appears in Eq. (2) in the diagonal matrix \mathbf{S}_m with elements $(gD_m)^{1/2}$, $(gD_m)^{1/2}$ and D_m . Further details of the applied NMF
98 representation are given in Žagar et al. (2015).

99 The input data vector contains wind components u, v and the geopotential height h defined as $h = g^{-1}P$ where g is the
100 gravity and P is a modified geopotential given by: $P = \Phi + RT_0 \ln(p_s)$, i.e. the sum of the geopotential field Φ and a surface
101 pressure p_s term. Other two variables represent the specific gas constant for dry air (R) and the globally-averaged vertical
102 temperature profile (T_0). The nondimensional complex expansion coefficients $\chi_n^k(m)$ represent both geopotential height and
103 wind perturbations due to waves. The Kelvin mode is represented in (2) by the first eastward-propagating IG mode. Although
104 our meridional index starts from 1 (to follow otherwise used notation), we shall denote KW in the reminder of this study as the
105 $n = 0$ EIG mode, i.e. the KWs are given by coefficients $\chi_{kw} = \chi_0^k(m)$.

106 In our application to the L91 ECMWF dataset, we used data on the N64 Gaussian grid and 91 model levels with model
107 top located at 0.01 hPa (around 80 km). Data are analyzed 4 times per day, at 00, 06, 12 and 18 UTC. The pre-processing
108 step consists of the interpolation of winds and geopotential from the hybrid ($\sigma - p$) levels to σ levels after geopotential Φ is



109 computed on the hybrid levels. The truncation values are $K = 55$ and $M = 60$. Higher vertical modes were left out as their
110 contribution is negligible in the outputs in the TTL and the stratosphere. The relation between the truncation parameters and
111 the normal-mode projection quality is discussed in Žagar et al. (2015) and references therein.

112 Once the forward projection is carried out and coefficients $\chi_n^k(m)$ are produced, filtering of KWs in physical space can be
113 performed through (2) after setting all χ , except those representing the KWs, to zero. The result of filtering are fields u_{kw} ,
114 v_{kw} and h_{kw} which provide the KW zonal wind, meridional wind and geopotential height perturbations. Notice here that in
115 contrast to the equatorial β -plane, KWs on the sphere have a very small meridional wind component which is thus left out
116 from the discussion.

117 The KW temperature perturbation, T_{kw} can be derived from the h_{kw} fields on σ levels using the hydrostatic relation in σ
118 coordinates:

$$119 \quad T_{kw} = -\frac{g\sigma}{R} \frac{\partial h_{kw}}{\partial \sigma}. \quad (3)$$

120 The orthogonality of the normal-mode basis functions provides KW energy as a function of the zonal wavenumber and
121 vertical mode. After the forward projection, the energy spectrum of total (potential and kinetic) energy for each Kelvin wave
122 can be computed using the energy product for the k th and m th normal modes (Žagar et al., 2015) as:

$$123 \quad I_{kw}(k, m) = \frac{1}{2} g D_m \chi_{kw} [\chi_{kw}]^*. \quad (4)$$

124 The units are J kg^{-1} . The KW global energy spectrum as a function of the zonal wavenumber is obtained by summing energy
125 in all vertical modes:

$$126 \quad I_{kw}(k) = \frac{1}{2} \sum_{m=1}^M g D_m \chi_{kw} [\chi_{kw}]^*. \quad (5)$$

127 2.2 Examples of 3D structure of Kelvin waves in L91 analyses

128 Kelvin waves are shown in Fig. 1-2 for a few days in July 2010 to introduce and illustrate their properties as filtered by the
129 NMF methodology. The second part of July 2010 was characterized by an abundance in both vertically propagating as well as
130 quasi-stationary KW structures throughout the atmosphere.

131 Figure 1 illustrates the meridional structure of Kelvin waves on 25 July 2010 on 2 levels. KW activity was found largest in
132 the zonal wind component at 150 hPa over the Indian Ocean. The geopotential dipole structure is centered over the convective
133 hotspot over the Maritime continent. At 100 hPa, we find largest amplitude of KW temperature perturbations up to 4 K
134 positioned above the zonal wind maxima at 150 hPa. The meridional wind component of the KW is at most 0.22 ms^{-1} at 100
135 hPa which is negligible compared to the zonal wind component (maximum 12.5 ms^{-1}) making the KW wind field primarily
136 zonal. Note that the presented horizontal structure at a single level is a superposition of 60 vertical modes, i.e. 60 shallow water
137 models with equivalent depths from about 10 km to a couple of meters.

138 Figure 2 can be discussed in relation to Eq. (3). It states that the amplitude of the cold (warm) KW temperature perturbation
139 is proportional to the negative (positive) vertical gradient in geopotential, as well as in zonal wind since zonal wind and



140 geopotential components are in phase. Horizontally, the cold anomaly is always located between the westerly and the easterly
141 phase of the zonal wave component. Vertically, maximum positive temperatures are observed between easterly winds below and
142 westerly winds above. A rough estimation can be made of the vertical wavelength based on alternating zonal wind minima and
143 maxima. For example, on 31 July a quasi-stationary vertical wave structure with extension in the stratosphere located around
144 60°E has easterly winds located at 50 hPa (~ 21.5 km) and 150 hPa (~ 13.5 km), which makes a vertical wavelength of around
145 8 km. In the stratosphere, above 80 hPa, strong KW activity was present in the form of free waves propagating eastward and
146 downward, therefore with upward transport of KW energy (Andrews et al., 1987). KW amplitudes were largest over Eastern
147 hemisphere with temperatures up to 4 K and zonal winds up to 12 ms⁻¹. The large amount of KW activity occurred during
148 the easterly phase of the QBO with strong easterly winds present between 30 and 80 hPa (not shown), providing favourable
149 conditions for the waves to propagate upward.

150 Between 100 and 200 hPa during the second half of July, there was low-frequency KW activity present in the form of a
151 stationary and robust "wave-1" pattern with strong KW easterly winds up to 24 ms⁻¹ in Eastern Hemisphere and KW westerly
152 winds up to 10 ms⁻¹ in Western Hemisphere. The high vertical resolution within the TTL resolves shallow KW structures and
153 a typical slanted structure towards the east in KW easterlies as well. The appearance and strength of horizontal KW response
154 coincides with the presence of strong easterly winds in the TTL in the Eastern Hemisphere during this period (not shown).
155 Figure 2 also shows that below 300 hPa the KW activity decreases and we shall not discuss levels under 300 hPa in the
156 paper. More examples based on daily basis filtered from the 10-day deterministic forecast of the ECMWF can be found on the
157 MODES website¹.

158 2.3 Other data and impact of the background state

159 In addition to the outputs from modal decomposition, full zonal wind and temperature fields from ECMWF analyses are used
160 to compute the background fields based on the same N64 grid and over the same period (Jan 2007 - Jun 2013). Zonal wind U
161 and static stability N are latitudinally averaged in the belt 5°S-5°N on all model levels to produce their zonal structure.

162 Static stability profiles are estimated through

$$163 \quad N^2 = \frac{g^2}{\Theta} \frac{\partial \Theta}{\partial \phi} \quad (6)$$

164 in units of s⁻² and are defined on hybrid model levels on which the geopotential field ϕ and the potential temperature field Θ
165 are derived a priori from the input data. Both fields are shown in Fig. 3.

166 The zonal wind field has the largest values on average in the TTL around 150 hPa with westerly winds peaking in the
167 Western Hemisphere over the Pacific Ocean and easterly winds peaking in the Eastern Hemisphere over the Indian Ocean
168 and Indonesia. It represents a typical time-averaged outflow pattern in response to tropical convection (e.g. Fueglistaler et al.,
169 2009). Throughout the seasons there is a longitudinal shift of this pattern following the convective source which is most clearly
170 observed at 150 hPa. Such seasonal shift is visible up to 100 hPa in Fig. 3(b) where winds are weaker compared to 150 hPa.
171 In northern winter, zonal winds are strongest over Indonesia and Eastern Pacific with the zonal wind maxima position and

¹<http://meteo.fmf.uni-lj.si/MODES/>



172 strength similar compared to the longer ERA-40 dataset used by Suzuki and Shiotani (2008). During northern summer easterly
173 winds mainly prevail over the Indian Ocean, which is linked to the Indian Monsoon season.

174 At 100 hPa, the static stability illustrates the strongest seasonal cycle with values ranging from near-tropospheric values of
175 $3 \times 10^{-4} \text{ ms}^{-2}$ during northern winter towards stratospheric values of $5 - 6 \times 10^{-4} \text{ ms}^{-2}$ during northern summer. Note also
176 the resolved local maxima in static stability at 80 hPa above the warm pools, known as the Tropical Inversion Layer (TIL) and
177 which is possibly wave-driven (Grise et al., 2010; Kedzierski et al., 2016). Figure 3(b) suggests that the TIL descends down to
178 100 hPa during the summer months peaking over Western Pacific, in agreement with the cycle found in GPS-RO observations
179 by Grise et al. (2010).

180 Kelvin waves are subject to wave modulation in changing background environments. Along its trajectory, the potential
181 energy of the KW changes with varying background winds and stability which can be largely described by linear wave theory
182 as long as waves are not near their critical level involving breaking and dissipation (Andrews et al., 1987). For simplification,
183 KW modulation can be examined for the case of pure zonal as well as pure vertical wave propagation based on the wave
184 modulation analysis performed by Ryu et al. (2008). A few key points on their local wave action conservation principle are
185 summarized in the following.

186 In the tropical atmosphere, zonal modulation is the dominant process for KWs propagating in the stratosphere and in all non-
187 easterly winds in the TTL. Vertical modulation becomes important in the presence of easterly winds within the TTL. Zonal
188 modulation is found to affect both u_{kw} and T_{kw} components and their amplitudes are proportional to the Doppler-shifted phase
189 speed by $(c - U)^{1/2}$ in case of pure zonal propagation direction. This means that Kelvin waves diminish in amplitude over
190 regions with westerly winds and become more prone to dissipative processes, while amplify over regions with easterly winds².
191 In case of pure vertical modulation, the change in wave potential energy mainly resonates with the temperature component of
192 the Kelvin wave. Along the rays' vertical path, the waves amplitude is proportional to the Brunt-Väisälä frequency as $\propto N^{3/2}$,
193 and to the Doppler-shifted phase speed as $\propto (c - U)^{-1/2}$, such that N is expected to play a primary role above 120 hPa where
194 its value starts increasing rapidly (see Fig. 3).

195 Alexander and Orland (2010) showed through wave modulation principles that temporal variations in zonal-mean N indeed
196 are correlated with observed KW amplitudes at 16 km (approx. 100 hPa). A more extensive wave modulation analysis was
197 described by Flannaghan and Fueglistaler (2013) using the full ray tracing equations to demonstrate that zonal winds in the TTL
198 not only modulate Kelvin waves locally, but also create a lasting modulating effect on wave activity through ray convergence
199 in the stratosphere. In particular, the seasonal cycle of the upper tropospheric easterlies (on average located over the western
200 Pacific), that acts as an escape window for Kelvin waves throughout the year and largely explains the longitudinal structure of
201 Kelvin wave zonal wind and temperature climatology.

202 We shall present the seasonal variability of tropical convection by using the Outgoing Longwave Radiation (OLR) dataset
203 with daily outputs from the NOAA Interpolated OLR product (Liebmann and Smith, 1996). The OLR product, often used as
204 a proxy for convection, is extracted on a $2.5^\circ \times 2.5^\circ$ grid and interpolated on a N64 grid. Latitudinal averages are derived over

²Keeping in mind that vertical wave propagation and consequently modulation becomes increasingly important as well wherever easterly winds are strong.



205 larger domain, namely over 15°S-15°N since organized convection tend to happen more remote from the equator, especially
206 during the summer monsoon season over the Asian continent.

207 3 Kelvin wave energetics

208 We start with an overview of KW energy distribution among the zonal wavenumbers as given by (5), followed by the seasonal
209 cycle of KW energy as a function of zonal wavenumber.

210 3.1 Energy distribution of Kelvin wave

211 The seasonal cycle in the energy-zonal wavenumber spectra is shown in Fig. 4 after summing up over all vertical modes. On
212 average, energy decreases as the zonal wavenumber increases as typical for atmospheric energy spectra. As we deal with the
213 large scales, we show only the first six zonal wavenumbers with energy values shown separately for the annual mean and the
214 four seasons separately.

215 Figure 4 shows that largest seasonal variations in KW energy are found at the largest zonal scales. For all zonal wavenum-
216 bers, above annual-mean energy values are observed during winter and summer seasons while autumn and spring are below
217 annual-mean energy. In the zonal wavenumber 1, total KW energy varies between 200 Jkg⁻¹ in MAM season and somewhat
218 over 300 Jkg⁻¹ in JJA. In wavenumber 2, values do not exceed 100 Jkg⁻¹ and JJA still contains the largest energy. At higher
219 wavenumbers, DJF season becomes the most energetic. In $k > 4$, total KW energy is under 20 Jkg⁻¹ and continue to reduce
220 with k . The slope of the KW energy spectrum is between $-5/3$ and -1 at planetary scales (not shown), similar to the spectra
221 presented in Žagar et al. (2009) for July 2007 data. The summer spectra has on average the steepest slope compared to other
222 seasons, in particular the winter spectra. The energy distribution on planetary scales is mainly associated with large-scale trop-
223 ical circulation established in response to ongoing tropical convection. Therefore, the zonal distribution of tropical convection
224 may likely play a crucial role in explaining winter and summer season differences of KW energy, which will be explored in
225 next section.

226 3.2 Seasonal cycle of KW energy

227 Figure 5 illustrates more details on the seasonal cycle by showing KW energy time series at the largest scales represented by
228 zonal wavenumbers $k = 1$, $k = 2$ and remaining scales $k > 2$. During most summers and occasionally in winter (e.g. 2008)
229 the total amount of KW energy in $k = 1$ can reach up to 600 Jkg⁻¹, or twice the summer average. The minimum in $k = 1$
230 KW energy mainly occurs during October month followed by April with values dropping towards 100 Jkg⁻¹, or half the
231 autumn average. The temporal pattern in $k = 2$ is similar to the $k = 1$ pattern, but with a less pronounced semiannual cycle
232 with maximum values up to 200 Jkg⁻¹ and minimum values towards 30 Jkg⁻¹. On zonal scales $k > 2$, KWs still show a
233 semiannual cycle with highest vertically-integrated values of energy over winter seasons.

234 In particular, for zonal wavenumber $k = 1$ one can distinguish inter-monthly in addition to semiannual variability. Inter-
235 monthly variability is most clearly observed during northern summer, for example in July 2011 where one can distinguish



236 six separate peaks of over 400 Jkg^{-1} energy over a period of approximately 90 days resembling an average wave period of
237 about 18 days. These are typical periods for free propagating Kelvin waves as observed in the TTL and lower stratosphere (e.g.
238 Randel and Wu, 2005). Note here again that our KW energy is vertically integrated over the whole model depth. This means
239 that the observed intermonthly variability of KWs appears dominated by the cyclic process of free propagating KWs entering
240 the TTL, amplifying due to changing environmental conditions, followed by wave breaking or dissipation.

241 The dominant scales of temporal variability in KWs are illustrated by a frequency spectrum of $k = 1$ in Fig. 6. The spectrum
242 is produced by a Fourier transform of energy data time serie of 6.5 years to frequency space. The resulting power spectrum has
243 been smoothed by taking the Gaussian-shaped moving averages over the raw spectrum by using a Daniell kernel three times
244 (Shumway and Stoffer, 2010). The spectrum shows a clear peak at 1-day period representing tidal variability in KWs. After
245 that, a gradual increase of energy is seen towards the 16-day period with multiple individual periods standing out. For periods
246 longer than 20 days, individual peaks are found close to 25, 43 and 59 days. After that, most KW energy is contained by far in
247 the semiannual cycle. The frequency spectrum provides an useful starting point for the discussion in the next section when the
248 spatiotemporal patterns of KWs shall be examined in several spectral domains.

249 Returning to Fig. 5, a low-pass filter with 90 day cut-off has been applied on KW energy in order to keep only the two
250 main spectral peaks in Fig. 6. The result is visible as the thicker black line in Fig.5 for all three zonal wavenumber groups. A
251 semiannual cycle for all zonal wavenumbers is evident with most energy observed around January and July, while least energy
252 is observed approximately one month after the equinoxes. During the years 2007, 2010, 2011, and 2012, more $k = 1$ KW
253 energy is observed during summer compared to the follow-up winter. The winter of 2009-2010 was for example above average
254 with energy values for $k = 1$ above 350 Jkg^{-1} .

255 The year to year differences can be explained by many coupled factors: In general, one expects vertically-integrated KW
256 activity to increase when background wind conditions become favorable, i.e. in the presence of easterly winds. This occurs
257 in the TTL in relation to strong convective outflow (Garcia and Salby, 1987; Suzuki and Shiotani, 2008; Ryu et al., 2008;
258 Flannaghan and Fueglistaler, 2013) during winter and summer seasons mainly. Moreover, one can expect enhanced KW activity
259 whenever the easterly QBO cycle is present in the stratosphere (Baldwin and Coauthors, 2001; Alexander and Ortland, 2010)
260 or when the ENSO index is positive (Yang and Hoskins, 2013). The latter factor might explain partly the large difference in the
261 abundant amount of KW energy during the El Niño winter of 2009-2010 and the below-average amount of KW energy a year
262 after during the strong La Niña winter of 2010-2011. However, during the La Niña winter of 2007-2008, the amount of KW
263 energy is observed to be above normal. That winter was however characterized by favorable easterly QBO conditions in the
264 stratosphere while during the winter of 2010-2011 stratospheric winds were largely westerly of nature thereby prohibiting KW
265 activity. The role of these low-frequency atmospheric phenomena on KW seasonal variability is a topic of further research.

266 Finally, Fig. 5 also shows that July 2007, previously examined by Žagar et al. (2009), was an exceptionally energetic month.
267 A large part of that energy, approximately 400 Jkg^{-1} (52.7% of total KW energy), was projected on zonal wavenumber 1. In
268 spatiotemporal terms, it represented the presence of a strong dipole structure in the TTL (as in Fig.2), which is colocated with
269 favourable easterly wind conditions in the TTL as well as in the stratosphere (not shown).



270 4 A spatiotemporal view on Kelvin wave seasonal variability

271 4.1 Kelvin wave decomposition among wave periods

272 In this section, the spatiotemporal view of KWs shall be presented over three dominant ranges of wave periods in Fig. 6,
273 namely: (i) the (semi)annual cycle using a low-pass filter with cut-off period at 90 days, (ii) the intraseasonal period using
274 a bandpass filter over periods between 20-90 days, and finally (iii) the intramonthly period with bandpass filtered periods
275 between 3-20 days. The choice of ranges, especially the intramonthly periods is related to previous studies using observations.
276 For all three cases, mean 6 year fields as well as seasonal means shall be presented.

277 Both KW components u_{kw} and T_{kw} are Fourier-transformed to frequency space where the spectral expansion coefficients
278 χ_{kw} in domains outside the desired frequency ranges are put to zero. Case (i) results in KW components $u_{kw,l}$ and $T_{kw,l}$ where
279 l indicates the low-frequency component. Case (ii) results in $u_{kw,m}$ and $T_{kw,m}$ where m indicates the intramonthly period.
280 Case (iii) results in fields $u_{kw,h}$ and $T_{kw,h}$ where h stands for the high-frequency component. Previous studies have defined
281 free propagating Kelvin waves over similar ranges (3-20 days, Alexander and Ortland (2010); 4-23 days, Suzuki and Shiotani
282 (2008)) and similarly for intraseasonal periods (23-92 days, Suzuki and Shiotani (2008)). Next, seasonal averages will be taken
283 over the four seasons, resulting in variables $\overline{u_{kw,l}}^s$, $\overline{T_{kw,l}}^s$ for the low-frequency component and similarly for the other two
284 cases. The superscript s represents one of the four seasons: northern winter ($s = DJF$), spring ($s = MAM$), summer ($s = JJA$),
285 and autumn ($s = SON$).

286 Cases (ii) and (iii) contain purely subseasonal variability and therefore one can expect their mean 6-year fields to be zero-
287 valued since variability beyond 90 days has been put to zero. Similarly, mean fields for each of the four seasons results
288 in $\overline{u_{kw,h}}^s \ll \overline{u_{kw,l}}^s$ and $\overline{u_{kw,m}}^s \ll \overline{u_{kw,l}}^s$ and the same for the temperature component. This reflects the fact that positive
289 and negative phases of the fast KW responses average out to approximately zero on seasonal timescales (figure not shown).
290 Therefore, the seasonal mean over the absolute amplitudes for zonal wind and temperature are examined instead, i.e. $\overline{|u_{kw,h}|}^s$,
291 $\overline{|u_{kw,m}|}^s$ and similarly for temperature component, in order to study seasonal fluctuations in subseasonal KW amplitudes.

292 Figure 7 shows results for all three cases after taking mean over the whole period. The left panel resembles a dominant
293 "wave-1" structure with zonal wind maximized around 140 hPa. Easterly KW winds are strongest around 60°E and westerly
294 winds around the Date Line. Note that two stationary perturbations over African (30°E) and South American (80°W) orography
295 are the result of our terrain-following NMF analysis. If one compares the KW zonal wind pattern with the climatological zonal
296 wind pattern in Fig. 3(a) it can be observed that the zonal wind pattern is located around 20° west of the climatological pattern.
297 Wave temperature perturbations are largest where the vertical gradients in zonal wind are largest which explains the quadrupole
298 structure. Heating (cooling) by KWs is located at 100 hPa in Eastern (Western) Hemisphere and the other way around at
299 200-300 hPa.

300 The middle panel of Fig. 7 shows the average distribution of KW activity on intraseasonal timescales. The activity is largest
301 in the Eastern Hemisphere with average zonal wind maxima up to 3 ms⁻¹ and temperature maxima up to 0.7 K. Zonal wind
302 activity is largest over a broad area between 90 and 150 hPa over the Indian Ocean and the Maritime Continent. Temperature



303 activity occurs slightly higher around 90-100 hPa. Intraseasonal activity is locally somewhat increased also around 120°W,
304 west of the Andes mountain range.

305 Finally, Fig. 7c illustrates the average distribution of free propagating KWs. The Eastern Hemisphere again makes up for the
306 larger KW activity than the Western hemisphere, but the maximum is located more upward in comparison to the intraseasonal
307 scales, around 80 hPa. Zonal wind activity peaks up to 3 ms^{-1} over a broad range of 70-110 hPa and temperature peaks over
308 a more narrow area around 76 hPa (up to 0.75 K). The main area for KW activity is found over Indian Ocean region, while
309 least wave activity is above central Pacific. Towards the stratosphere KW activity reduces and becomes more uniform along in
310 longitudinal direction.

311 4.2 Low-frequency Kelvin wave variability

312 The seasonal patterns of the low-frequency components of the KW (from hereon referred to as the Gill-type KW response) is
313 presented as pressure-longitudinal cross-sections along the equator (at 0.7°N) of the KW seasonal means, given by $\overline{[u_{kw,t}]^s}$
314 and $\overline{[T_{kw,t}]^s}$ in Fig. 8.

315 The largest Gill-type KW response is found during NH summer. A strong dipole "wave-1" pattern is evident in the TTL. The
316 strongest zonal winds are found close to 150 hPa with easterlies up to -12 ms^{-1} centered over Indian Ocean and westerlies
317 up to 6 ms^{-1} over the Western Pacific. Negative temperature KW anomalies at 110 hPa are strongest as well during JJA with
318 values up to 1.5 K over Indian Ocean and annually averaged value of -0.5 K over Western Pacific.

319 During NH winter, the dipole pattern is shifted more eastward and upward compared to NH summer and has a more slanted
320 structure. Easterly (westerly) KW winds are located more east over the Maritime continent (central Pacific) and centered at 130
321 hPa. The upper temperature dipole pattern is found higher up at 90 hPa approximately. Values are somewhat weaker compared
322 to NH summer with easterlies up to -6 ms^{-1} and westerlies up to 5 ms^{-1} .

323 Finally, NH autumn and spring seasons are transition seasons with respect to the strength and position of the KW dipole as
324 it moves west- and downward towards summer and east- and upward towards winter. NH spring has the weakest KW dipole
325 with slightly stronger westerly winds up to 5 ms^{-1} .

326 The longitudinal position and the strength of the Gill-type KWs have been linked to the seasonal patterns of the background
327 winds in the TTL representing the upper level monsoon and Walker circulations (Flannaghan and Fueglistaler, 2013). The
328 average background winds maximize at 150 hPa as shown in Fig. 3(a). In Fig. 8, one can see how the KW easterlies in
329 Eastern Hemisphere are strongest during NH summer in relation to the Indian-South Asian monsoon circulation. Background
330 easterlies as strong as -30 ms^{-1} are located approximately 10° east of the KW maximum easterlies. NH winter has the strongest
331 background westerlies in relation to the upper-level circulation of the Western Pacific anticyclones. NH spring (autumn) shows
332 similar background wind patterns compared to NH winter (summer) but with weaker circulation.

333 Further details on longitudinal position and interannual variability of Gill-type KW response at its maximum value at 150
334 hPa are illustrated by the Hovmoller diagram in Fig. 9. For comparison, tropical convection is represented as well through
335 the OLR proxy variable averaged over 15°S - 15°N latitudes. All fields have been filtered with a 90 day cut-off low-pass filter
336 in order to highlight the seasonality. As a result, one can observe enhanced/reduced Gill-type KW activity during the same



337 individual seasons as seen from the timeseries in Fig. 5. Above average seasonal KW activity with stronger Gill-type structures
338 occurred during the summer of 2007 (mainly through its easterlies at 60°E) and during the winters of 2006-2007 and 2009-
339 2010. In these winters, El-Nino was active and a clear longitudinal eastward shift is observed in OLR, in the background
340 circulation (not shown), as well as in the Gill-type KW structure. The El-Nino winter of 2009-2010 was followed by a strong
341 La Nina winter with an increase in tropical convection over the Maritime continent (note: OLR values below 195 Wm⁻²).

342 The vertical seasonal movement of the KW dipole has been linked with the seasonal movement of the tropical tropopause
343 height (Flannaghan and Fueglistaler, 2013; Ryu et al., 2008). The position of the tropical tropopause height (represented by
344 a static stability value of $5 \times 10^{-4} \text{ s}^{-2}$ in Fig. 8) is found at approximately 85 hPa during winter and descends towards 100
345 hPa in summer, similar to values obtained from GPS-RO observations by Grise et al. (2010). In particular, during summer,
346 one can notice how the asymmetry in the tropical tropopause height over Indian Ocean around 60°E coincides with increasing
347 temperatures by the KW dipole up to 1.5 K. Such deformation of the tropical tropopause is also evident during winter and
348 autumn seasons.

349 Figures 10a and 10b illustrate seasonal-mean KW temperatures $\overline{T_{kw,l}^s}$ in relation to the tropical tropopause layer defined
350 by static stability N^2 . Seasonal variations in KW temperatures are colocated with the position of the tropopause, descending
351 down from its highest position during winter to its lowest position during summer. Temperature amplitudes are observed to
352 decline roughly above $N^2 = 5 - 6 \times 10^{-4} \text{ s}^{-2}$. Within this zonal-mean seasonal picture, zonal asymmetries in N^2 exist and
353 are found: (i) near the Date Line with values of $8 \times 10^{-4} \text{ s}^{-2}$ at 80 hPa during winter and $7 \times 10^{-4} \text{ s}^{-2}$ at 90 hPa during
354 summer and (ii) lower at 100 hPa over the Indian Ocean during summer. Particularly during NH summer, the deformation
355 of the zonal-mean static stability field collocates strongly with the position of a strong KW temperature anomaly over Indian
356 Ocean. A rough estimation is made on the contribution of the KW anomaly to the zonal deformation of the tropopause layer by
357 removing zonal-mean parts of both fields. First, static stability zonal anomalies, $\overline{N'^2}^s$, are derived by subtracting zonal-mean
358 values of N^2 from the full N^2 field per timestep and at every pressure level, followed by seasonal averaging. Next, we can
359 estimate the static stability change associated with the KW anomaly, using the relation: $N_{kw}^2 = \frac{g}{\theta} \frac{\partial \theta_{kw}}{\partial z}$, followed by seasonal
360 averaging as well, i.e. $\overline{N_{kw}^2}^s$.

361 As a result, Fig. 10c and 10d show how both static stability anomalies are overlapping. During winter, the structure of the
362 zonal anomaly $\overline{N'^2}^s$ has a positively-valued tilt eastward which stretches up to 80 hPa, while during summer a strong static
363 stability anomaly is found more localized over Indian ocean region with values in the TTL up to $\overline{N'^2}^{JJA} = \pm 0.8 \times 10^{-4} \text{ s}^{-2}$.
364 The anomaly associated with the KW temperature anomaly is found to peak up to $+0.6 \times 10^{-4} \text{ s}^{-2}$ during summer and up
365 to $+0.4 \times 10^{-4} \text{ s}^{-2}$ during winter. Finally, by dividing both fields with each other, the resulting contribution of the quasi-
366 stationary Kelvin wave to the observed deformation of the tropical tropopause layer is estimated up to 60% (80%) during NH
367 summer (winter).

368 4.3 Intraseasonal Kelvin wave variability

369 The seasonality of intraseasonal Kelvin wave variability is shown in Fig. 11 and shall be briefly discussed here. The NH
370 winter stands out as the most active season for KW activity, located mainly in the Eastern hemisphere centered at 100°E and



371 with maximum activity at 110 hPa for zonal wind and temperature with a second maximum in temperature at 90 hPa. Values
372 observed are up to 0.8 K for KW temperature and 5 ms^{-1} for KW zonal wind. During NH spring season, the KW activity fields
373 are weaker but spread over a larger area in the Eastern hemisphere and in the TTL with maximum activity centered at 120 hPa
374 (90 hPa) for the zonal wind (temperature) component. Both NH summer and autumn seasons have KW activity positioned
375 at lower altitudes and more westward. In both seasons, KW zonal wind activity is split up between two structures with an
376 eastward tilt with height; one with a maximum around 110°E and one pattern starting from 100 hPa and extending towards
377 60°E . Note also the increase in KW activity in the Western hemisphere below 150 hPa in the East Pacific. The maximum KW
378 activity in the temperature component for both seasons is positioned near 100 hPa approximately on the tropical tropopause
379 contour with value $5 \times 10^{-4} \text{ s}^{-2}$.

380 The eastward tilted structure is observed throughout all seasons except NH spring when background easterly winds are nearly
381 absent in the Eastern hemisphere. In all other seasons one can observe how the tilted structure is locked to the background
382 easterlies with maximum amplitudes located slightly above and west of it. Such eastward tilt with height has been frequently
383 observed, for example over radiosonde station Medan at 100°E during the early stage of MJO development (Kiladis et al.,
384 2005).

385 4.4 Free propagating Kelvin waves

386 The seasonal variability of free traveling Kelvin waves, represented by their absolute amplitudes $\overline{|u'_{kw,h}|^s}$ and $\overline{|T'_{kw,h}|^s}$, shall
387 be examined in relation to the background conditions. Figure 12 illustrates favorable regions for KW activity. In general, KW
388 activity increases upward from around 120 hPa towards its zonal-mean peak value at 76 hPa. The largest values are observed
389 in EH in region from 30°E till 150°E . The temperature component in particular has a constant maximum peak (up to 0.8 K
390 in EH) located around 76 hPa throughout the year, where also the largest increase in N^2 occurs as shown in Fig. 3. Above 70
391 hPa, KW activity continuously decreases in the stratosphere.

392 The longitudinal structure of the KW zonal wind shows two distinct peaks in the TTL, one consistently located at 76 hPa
393 and another around 100-110 hPa in the EH which is mainly present during solstice seasons. The first maximum coincides
394 with the temperature distribution which can be explained by their balance relationships and free horizontal propagation in the
395 stratosphere. Below the tropopause, KW activity is coupled to convective processes alternating the tropospheric vertical wave
396 structures as discussed by Flannaghan and Fueglistaler (2012).

397 The secondary maximum around 110 hPa in Fig. 12 is present mainly during solstice seasons in EH and it is associated
398 with the seasonal movement of the background wind. The maximum of KW wind and the background wind maximum move
399 eastward from winter to summer season similar to the low-frequency variability. A day-by-day comparison of the KW activity
400 and background wind confirms that propagating KWs amplify while approaching a region of strong easterlies, forming a
401 folding structure around it while the individual KWs dissipate towards the center of easterly winds. One can notice in Fig. 12 a
402 fast reduction of KW amplitudes eastward of its maximum towards the center of the background easterlies. It is likely related
403 to dissipation and wave breaking processes as observed over Indonesia (120°E) by Fujiwara et al. (2003). Within such regions,



404 the KW-background wind interaction becomes complex and the linearity assumption breaks (Ryu et al., 2008; Flannaghan and
405 Fueglistaler, 2013).

406 A comparison with the previous study by Suzuki and Shiotani (2008) using ERA-40 data shows that the L91 data contain
407 stronger KW activity in the vicinity of the background easterlies in the Eastern Hemisphere, and more fine-scale details which
408 can be explained by better analyses based on more observations and improved models including increased resolution. For
409 example, Suzuki and Shiotani (2008) used 5 levels of ERA-40 data between 50 and 200 hPa whereas the present study considers
410 25 model levels between 50 – 200 hPa. Maxima of the KW temperature signal appear in similar locations and strength except
411 for a small offset in vertical position (70 hPa in Suzuki and Shiotani (2008) versus 80 hPa in Fig. 12) and a larger zonal
412 asymmetry in our results.

413 Another view of the seasonal cycle of free propagating KWs is illustrated in Fig. 13 which focuses on the spatiotemporal
414 distribution of individual KW packets. Hovmoller diagrams of KW zonal wind and temperature at levels 110 and 200 hPa from
415 different years are shown along with the background zonal wind. In addition, the monthly-mean values of daily maximum KW
416 amplitudes occurring at a specific longitude along the equator are added next to each diagram.

417 The individual wave tracks at 110 hPa illustrate KWs with amplitudes exceeding 3 ms^{-1} and 0.6 K which are propagating
418 throughout the year in the Eastern Hemisphere, during June-October months only over the Pacific, and all except winter
419 months in most of the Western Hemisphere. Typical wave tracks start east of the 0° (30°W) meridian during winter (summer)
420 and largely disappear west of 120°E . The largest wave amplitudes are observed between 50°E and 100°E prior to regions
421 of easterly winds in agreement with Fig. 12. Here presented details show that most notable waves appear during the Asian
422 monsoon period with upper-level easterlies prevailing from June into September. The largest Kelvin wave amplitudes appear
423 confined to the June and July months followed by a rapid drop in August. In fact, a local minimum in the number of KWs as
424 well as in wave amplitudes occurs in August before the KW activity increases slightly during autumn.

425 At 200 hPa, the favorable area for KW propagation shifts to the Western Hemisphere and large KW activity is observed
426 west of the South American continent throughout the year (west of 80°W) with a westward extension over the Pacific during
427 northern summer. Another set of wave tracks starts over equatorial South America around 30°W (5°W) and continues till 60°E
428 (90°E) during northern summer (winter). The seasonal shifts of approximately 30° in KW tracks collocate with similar shifts
429 in the prevailing TTL winds.

430 The amplitude of KWs undergoes a clear annual cycle with a small secondary peak present during northern winter, as
431 represented by the monthly-means of daily maximum amplitudes along the equator on the rightside of Fig. 13. The largest
432 amplitudes are found at 110 hPa during NH summer with monthly-mean zonal wind (temperature) values up to 8.5 ms^{-1}
433 (1.8 K) in June. During the winter months Kelvin waves amplify more eastward with monthly-mean zonal wind (temperature)
434 values up to 7.8 ms^{-1} (1.6 K) in December. At 200 hPa, KW amplitudes are on average lower with a yearly-averaged amplitude
435 reduction around 55% in temperature and 35% in zonal wind. The semiannual cycle in maximum amplitudes remains visible up
436 till 70 hPa. Above 70 hPa, where KW activity remains large in Eastern Hemisphere (Fig. 12), the semiannual cycle is replaced
437 by an interannual cycle in line with the dominant impact of the QBO.



438 5 Discussion and Conclusions

439 We have applied the multivariate decomposition of the ECMWF operational analyses during the period 2007-2013 when
440 the operational data assimilation was performed on 91 levels. Model-level data were analyzed every 6 hours. The applied
441 decomposition provides simultaneously the wind components, geopotential height and temperature perturbations of the Kelvin
442 waves on the terrain-following levels. As the KW meridional wind component is very small it is not discussed. We focused
443 on the spatiotemporal features of the KW temperature and zonal wind components in four seasons. The Kelvin wave seasonal
444 cycle in the tropical tropopause layer (TTL) was compared with seasonal variability of Outgoing Longwave Radiation (OLR),
445 the background wind and stability fields, which are believed to play an important role for the KW variability. Our study of
446 the seasonal KW variability complements previous studies which applied different methods for the KW filtering and different
447 datasets. As KW is a normal mode of the global atmosphere, our filtering of the KW using the 3D-orthogonal normal-mode
448 function decomposition of global data is a useful approach to quantification of the KW variance. The KW is the most energetic
449 inertio-gravity mode of the global atmosphere (Žagar et al., 2009) and its representation in weather and climate models is
450 crucial for reliable simulations of the tropics and its impact on global circulation.

451 We have presented the total energy of the KWs in the L91 data extending between the surface and 1 Pa as a function of the
452 zonal wavenumber. Zonal wavenumber $k = 1$ contains a largest portion of KW energy in all seasons. Its energy varies between
453 ~ 300 in JJA in NH spring to over 400 J/kg in NH summer. In $k = 2$ there is 50% less energy than in $k = 1$ but the NH summer
454 is still the most energetic season. In all greater zonal wavenumbers, DJF season contains most energy.

455 Frequency spectrum has revealed a semiannual cycle as well as intraseasonal and intramonthly variability. Three ranges of
456 wave periods were analyzed: 3-20 days, 20-90 days and longer than 90 days. This choice was partly deliberate in order to
457 compare our results with several previous studies of KW variability. First we demonstrated that the seasonal-mean KW pattern
458 in the TTL, with (westerly) easterly winds in the (Western) Eastern hemisphere resembles a time-averaged Gill-type "wave-1"
459 pattern. The quadrature-shaped temperature component represents a thermally adjusted pattern with respect to the zonal wind
460 component, and contributes to seasonal (cooling) warming above 100 hPa in the (Eastern) Western hemisphere. The largest
461 KW amplitudes are observed during summer and winter seasons. From boreal summer towards winter, KW perturbation moves
462 eastward (from Indian Ocean basin towards Maritime Continent) and upward (e.g. zonal wind component moves up from 150
463 hPa towards 120 hPa). The KW zonal wind amplitude varies between 12 m/s strong easterlies over Indian ocean near 150 hPa
464 in JJA to 6 m/s over Western Pacific. Over Indian Ocean in JJA, the KW easterlies thus make almost half of the total wind
465 vector. The associated KW temperature perturbations are from 1.5 K over Indian ocean in JJA to -0.5 K over West Pacific.
466 The zonal modulation of Kelvin waves is found to be locked with respect to the seasonal movement of convection and the
467 convective outflow in the TTL. The modulation effect is strongest for Gill-type Kelvin waves during the summer monsoon
468 season, when strong easterly winds are present at 150 hPa, resulting in the largest KW zonal wind and temperature anomalies,
469 of which the latter results in deformation of the tropical tropopause over Indian Ocean.

470 Intraseasonal (periods 20-90 days) activity is strongest in NH winter with maxima up to 0.8 K for KW temperature and
471 up to 5 m/s for KW zonal wind centred at 120°E. Both temperature and zonal wind activities have eastward tilt with height.



472 In comparison to previous study by Suzuki and Shiotani (2008) using ERA-40 data, the slanted structure in the present data
473 continues to extend more upward and eastward which is likely due to the increased number of vertical model levels compared
474 to ERA-40. The importance of vertical model resolution for the KW wave structure and amplitude was demonstrated in Žagar
475 et al. (2012) and Podglajen et al. (2014).

476 For periods 3 – 20 days, the seasonal cycle of KWs is clearly seen in wave amplitude. The largest amplitudes are located
477 - from a zonal-mean perspective - between 70 and 100 hPa for both zonal wind and temperature as expected for the free-
478 propagating Kelvin waves but it is modulated by the seasonal movement of the TTL. A major zonal asymmetry was found in
479 KW activity: around 110 hPa Kelvin wave undergoes amplification mainly in Eastern Hemisphere during the solstice seasons,
480 while at 200 hPa a secondary region of KW amplification occurs in Western Hemisphere during boreal summer. Free propagat-
481 ing KWs show largest amplitudes in the vicinity of the strongest easterlies preferably west and above the center of easterlies.
482 The NMF methodology has made it possible to observe such dynamics on daily basis whenever easterlies are strong in the
483 TTL. Nearly real-time representation of the KW activity is available on <http://modes.fmf.uni-lj.si>.

484 In summary, our seasonal variability analysis shows that the background wind in the TTL linked with convective outflows,
485 play a dominant role in the longitudinal position where zonal modulation of Kelvin waves is preferred, while the tropical
486 tropopause and its seasonal vertical movement determines the vertical extent of KW modulation processes.

487 *Acknowledgements.* This study was funded by the European Research Council (ERC), Grant Agreement no. 280153 MODES, [http://meteo.fmf.uni-](http://meteo.fmf.uni-lj.si/MODES)
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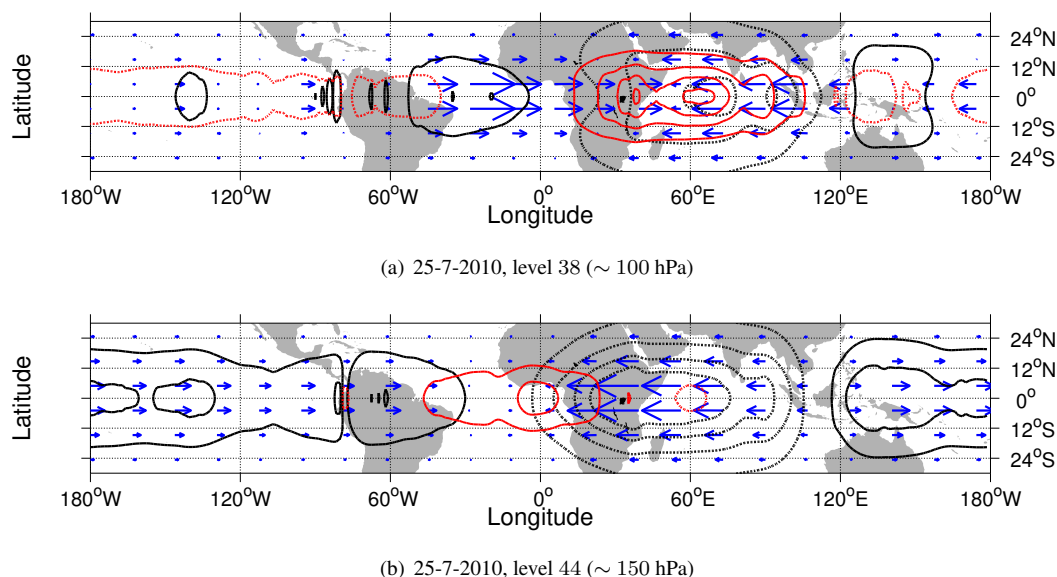


Figure 1. Horizontal map slices centered on the Tropical Belt on the 25th of July 2010 (panel b in Fig. 1) at (a) 100 hPa and (b) 150 hPa. Kelvin wave wind fields are represented by blue vectors. Contour fields indicate KW geopotential (h_{kw}) (black contours, every 20 m) and temperature (T_{kw}) (red contours, every 1 K). Dashed contours represent negative values for geopotential and temperature components.

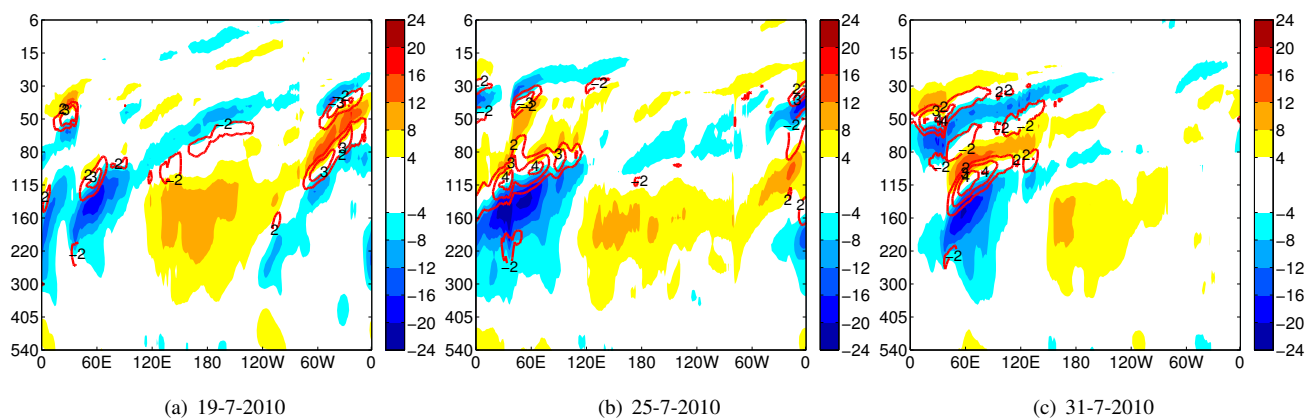


Figure 2. Longitude-pressure sections along 0.7°N of Kelvin wave zonal wind (red-blue shaded contours) and temperature (red contours) on the following days in July 2010: (a) 19, (b) 25 and (c) 31. Temperature is given every 1 K, contours starting at 2 K. Zonal wind values are given every 4 ms^{-1} .

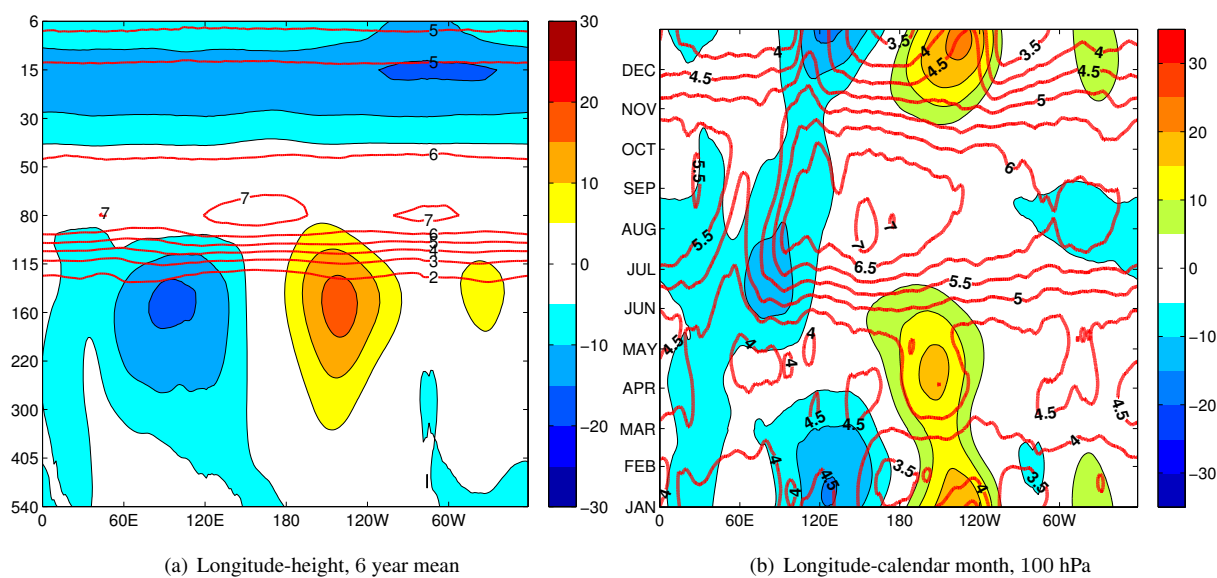


Figure 3. Zonal wind (blue-to-red contour fields, each 5 ms^{-1}) and static stability given by Eq. (6) (red contours each $1 \times 10^{-4} \text{ s}^{-2}$ for a) each $0.5 \times 10^{-4} \text{ s}^{-2}$ for b) panel). (a) Longitude-height section averaged over 6 years and (b) Longitude-time section at 100 hPa. Both fields are latitudinally averaged over 5°S - 5°N , and have been low-pass filtered a priori with a cut-off period of 90 days to highlight seasonal variability.

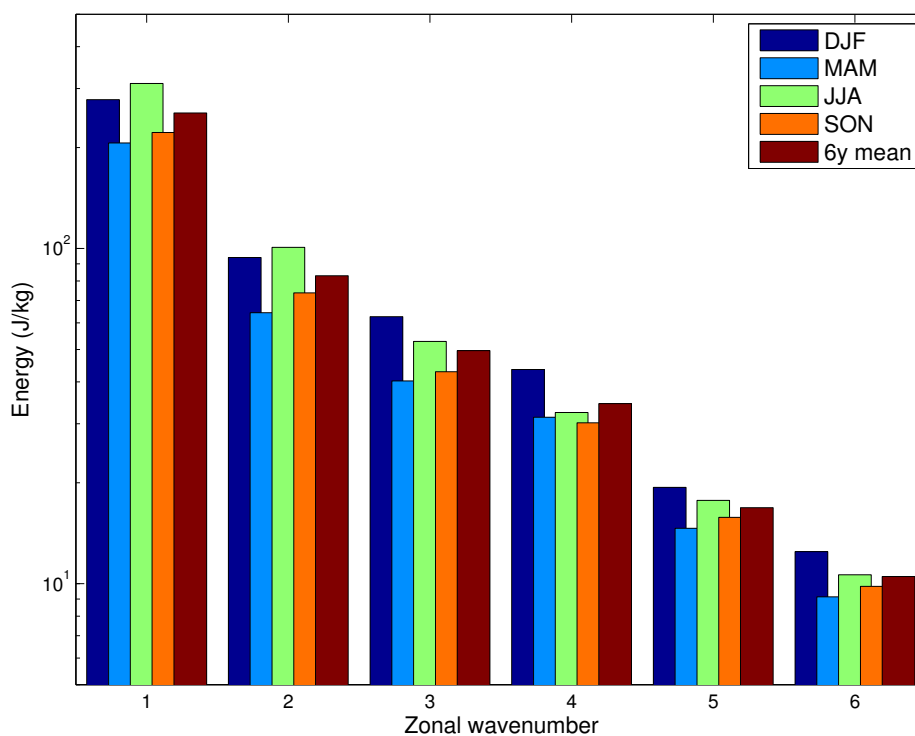
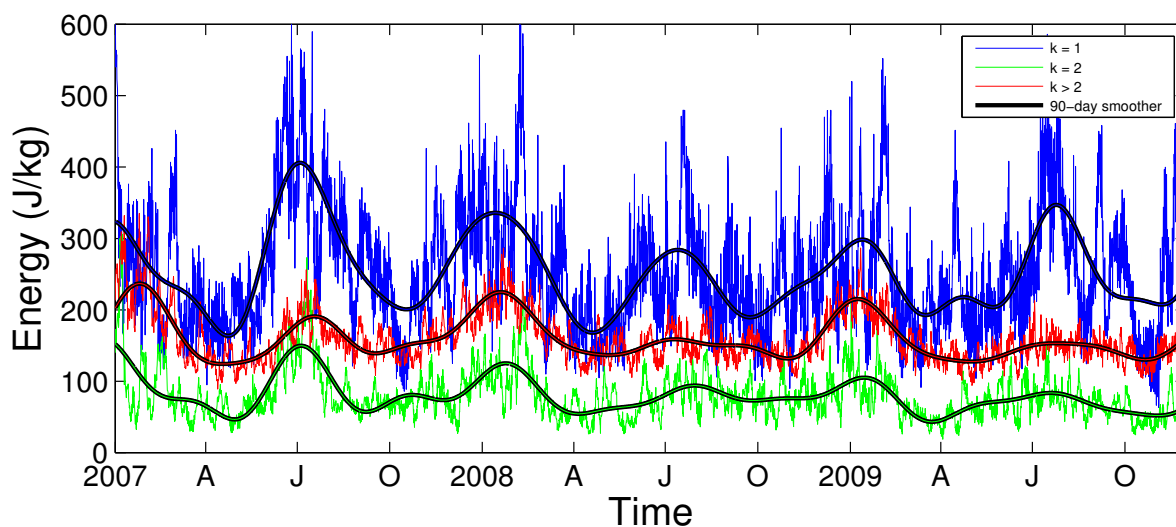
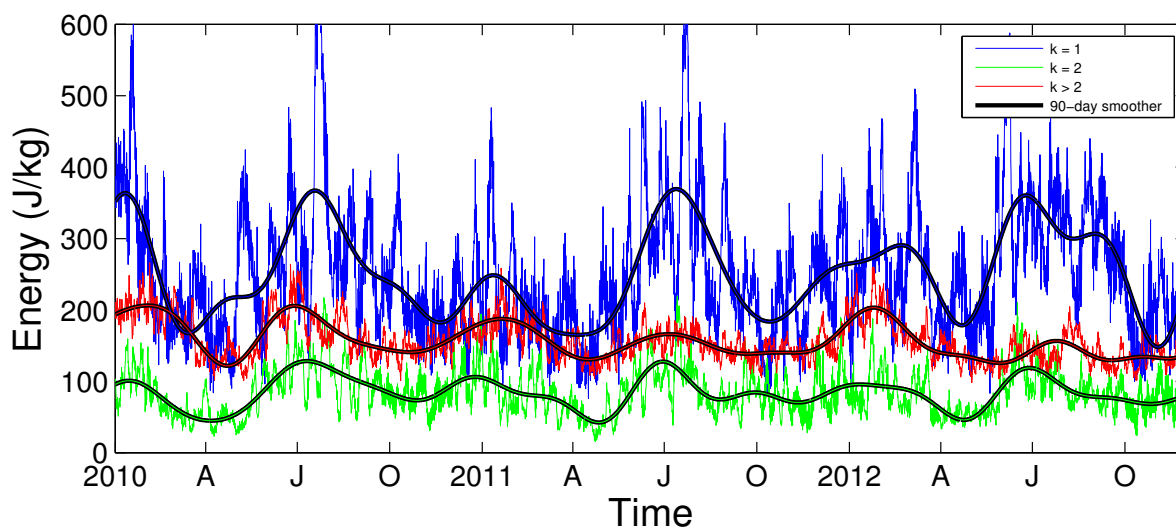


Figure 4. Kelvin wave energy (in J kg^{-1}) as function of zonal wavenumbers k for $k = 1 - 6$. For each k , energy values are averaged over the following periods from left (blue bars) to right (red bars): DJF (winter), MAM (spring), JJA (summer), SON (autumn) and for the full 6, 5 year period. Energy is summed over all vertical modes.



(a) 2007 – 2009



(b) 2010 – 2012

Figure 5. Timeseries of KW energy (in Jkg^{-1}) for various zonal wavenumbers over the following periods: (a) 2007 – 2009 and (b) 2010 – 2012 (incomplete year 2013 has been left out). 'A', 'J' and 'O' refer to the first day of April, July and October months. Energy values are summed over all vertical modes and for the following zonal wavenumbers: planetary scales $k = 1$ (blue) and $k = 2$ (green) and all smaller zonal scales, $k > 2$ (red). A 90-day low-pass filter has been applied (black lines) for each of the zonal wavenumber groups in order to filter out high-frequency variability and to highlight seasonal variability.

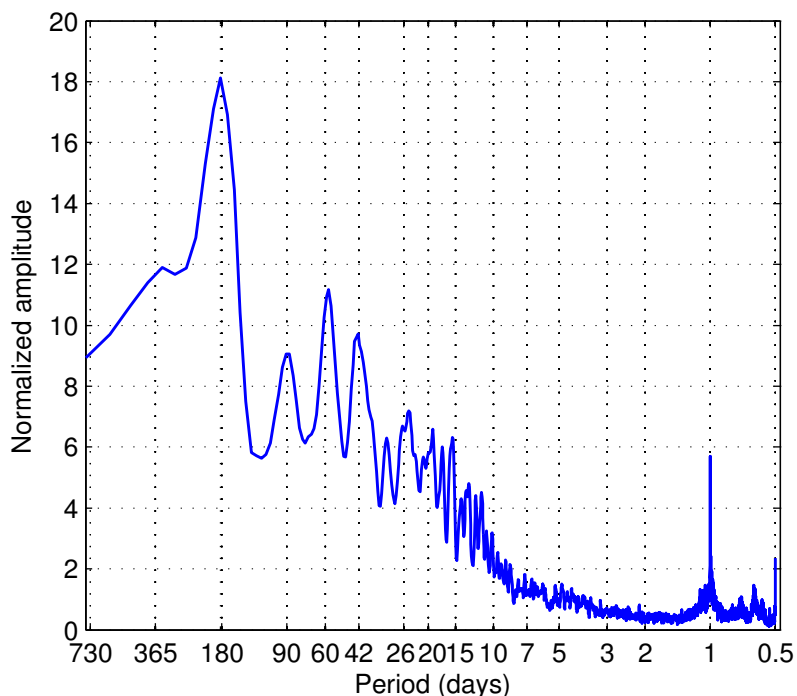


Figure 6. Frequency spectra of Kelvin wave energy for the zonal wavenumber $k = 1$ and summed over all vertical modes. A 1-2-1 filter with a Daniell kernel has been used to smooth the initial raw power spectra.

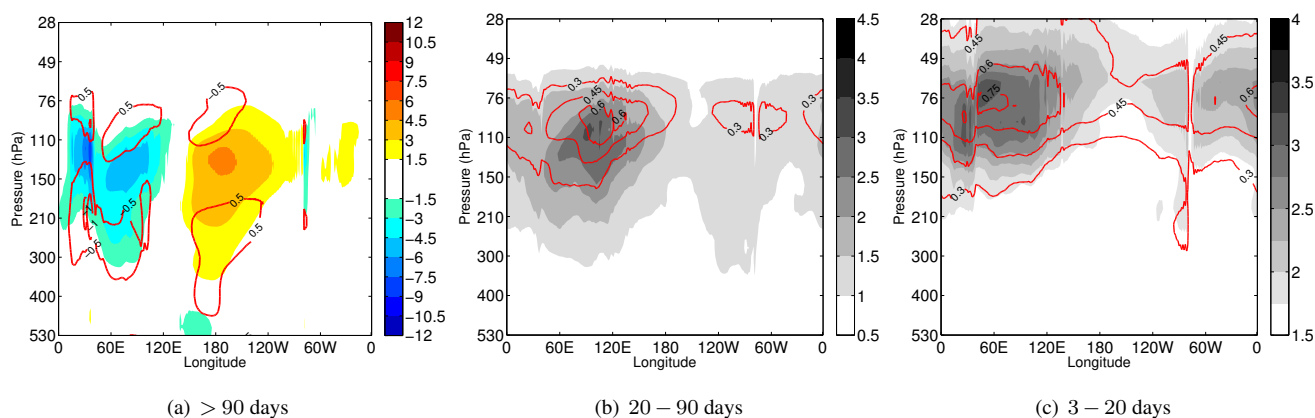


Figure 7. Longitude-pressure sections along 0.7°N of KW fields averaged over the full data period for: (a) low-frequency, (b) intraseasonal, and (c) intramonthly Kelvin wave variability with the associated filtered wave periods given in the panel captions. The contour scales are as follows: (a) mean KW zonal wind (colored fields, each 1.5 ms^{-1}) and KW temperature (red contours, each 0.5 K) components, (b) mean absolute amplitudes of KW zonal wind (grey fields, each 0.5 ms^{-1}) and temperature (red contours, each 0.15 K), and (c) mean absolute amplitudes of zonal wind (grey fields, each 0.25 ms^{-1}) and temperature (red contours, each 0.15 K).

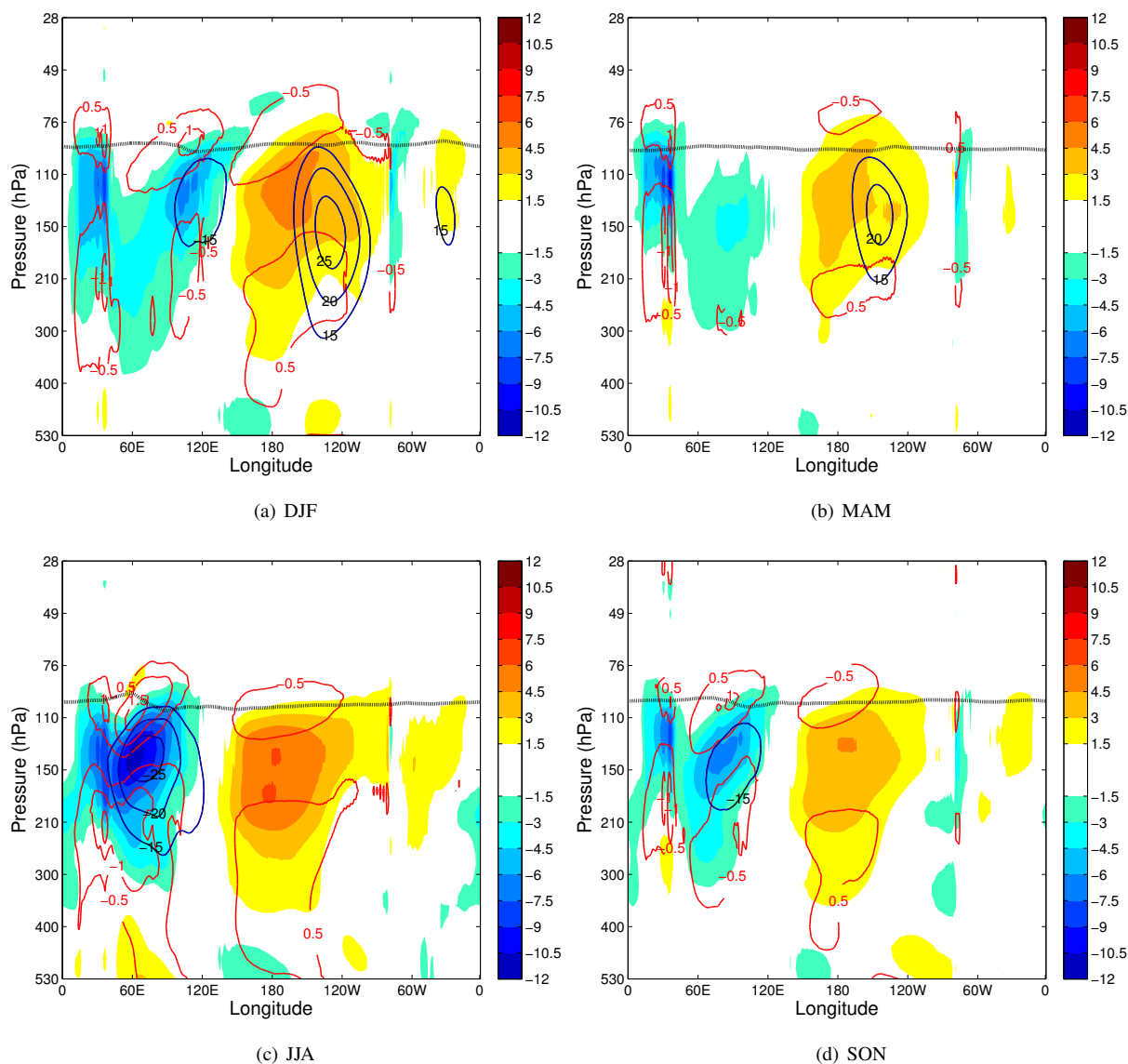


Figure 8. Longitude-pressure sections averaged over the following NH seasons: northern (a) winter, (b) spring, (c) summer, and (d) autumn. Kelvin wave zonal wind (blue-to-red colored contours) and temperature (red contours) fields on equator 0.7°N have same contour values as in Fig. 7(a). The background zonal wind is shown in blue contours (each 5 ms^{-1} , starting from 15 ms^{-1}). A single static stability contour of value $5 \times 10^{-4}\text{ s}^{-2}$ is shown as a thick dotted black line in order to represent the seasonal movement of the tropical tropopause height. Both background wind and stability fields are latitudinally-averaged over 5°S - 5°N . All fields are smoothed a priori using a low-pass filter with cut-off period of 90 days.

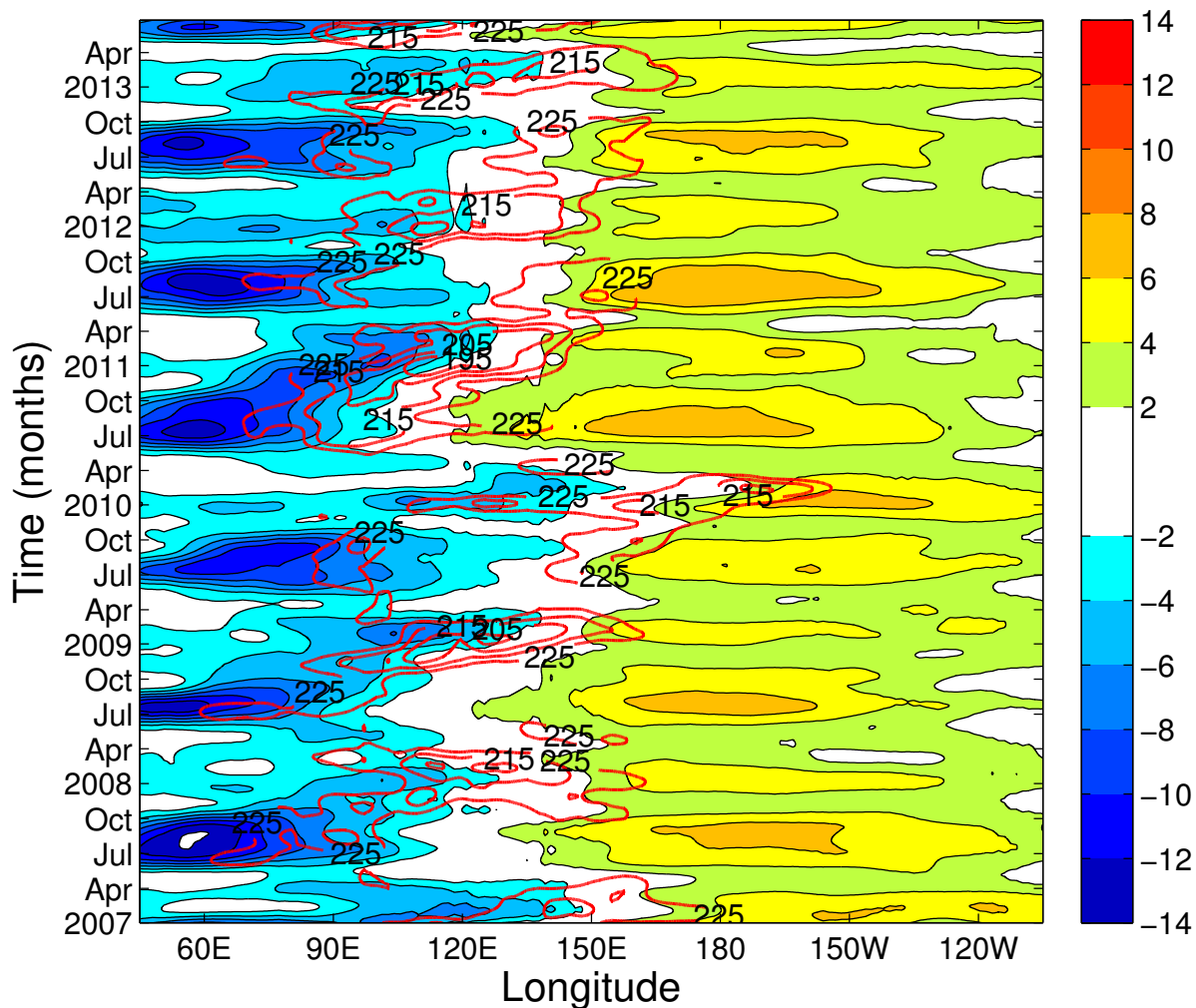


Figure 9. Longitude-time section at model level 45 (~ 153 hPa) of Kelvin wave zonal wind (blue to red shaded contours, each 2 ms^{-1}) using values at 0.7°N as well as Outgoing Longwave Radiation (red contours, each 10 Wm^{-2} starting at 225 Wm^{-2}) averaged over latitudes from 15°S till 15°N . Both fields have been filtered a priori using a low-pass filter with cut-off at 90 day period.

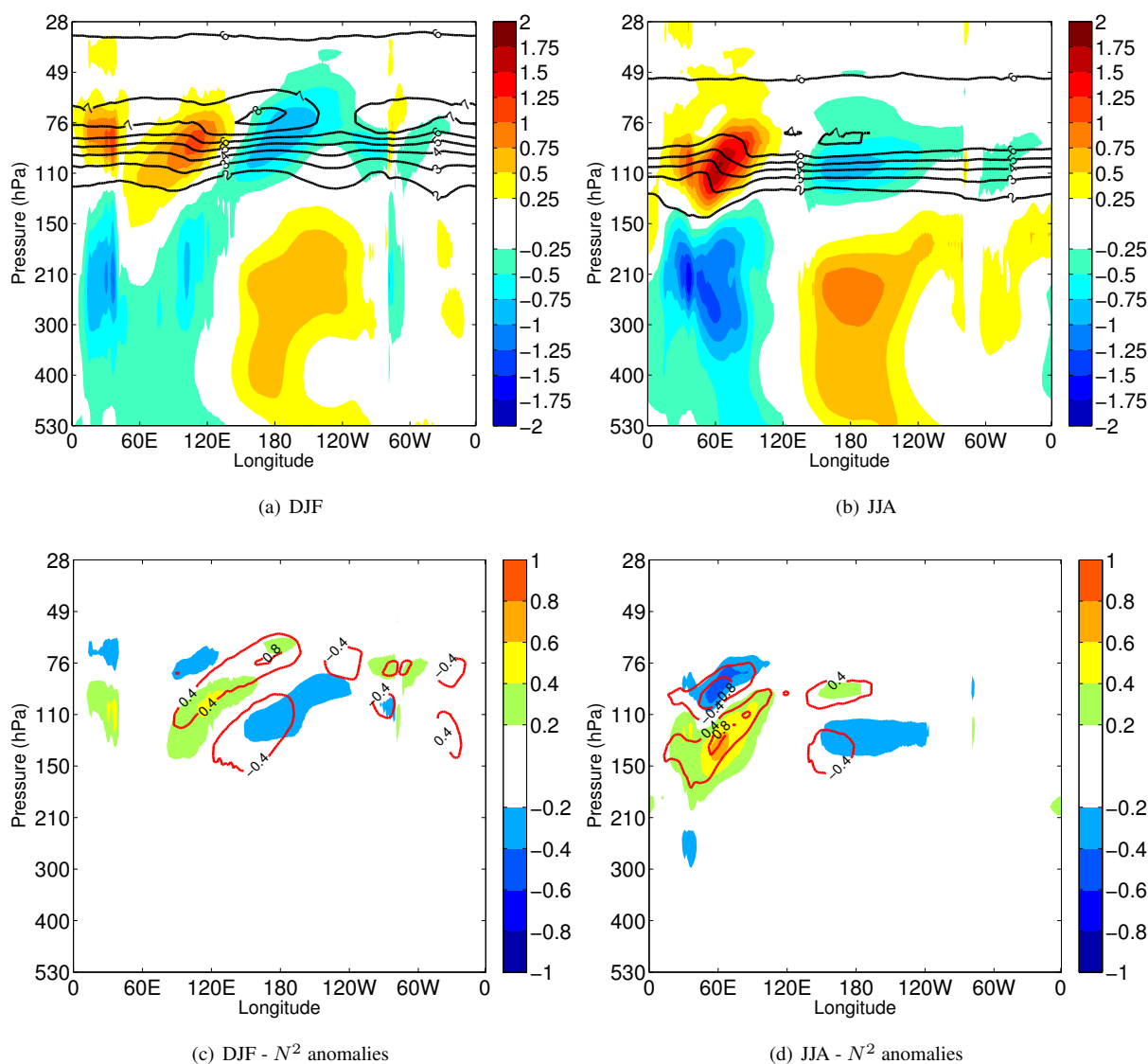


Figure 10. Seasonal-averaged view for (a and c) DJF and (b and d) JJA months as function of longitude and pressure for the following fields: (a-b panels) KW temperature, $\overline{T_{kw}^s}$, (blue-to-red, each 0.25 K) and static stability field, $\overline{N^{2s}}$ (black contours, each $1 \times 10^{-4} \text{ s}^{-2}$, starting at $2 \times 10^{-4} \text{ s}^{-2}$), (c-d panels) KW static stability anomaly, $\overline{N_{kw}^{2s}}$ (blue-to-red, each $0.2 \times 10^{-4} \text{ s}^{-2}$), and static stability anomaly with respect to the zonal mean, $\overline{N^{2s}}$ (red contours, each $0.4 \times 10^{-4} \text{ s}^{-2}$).

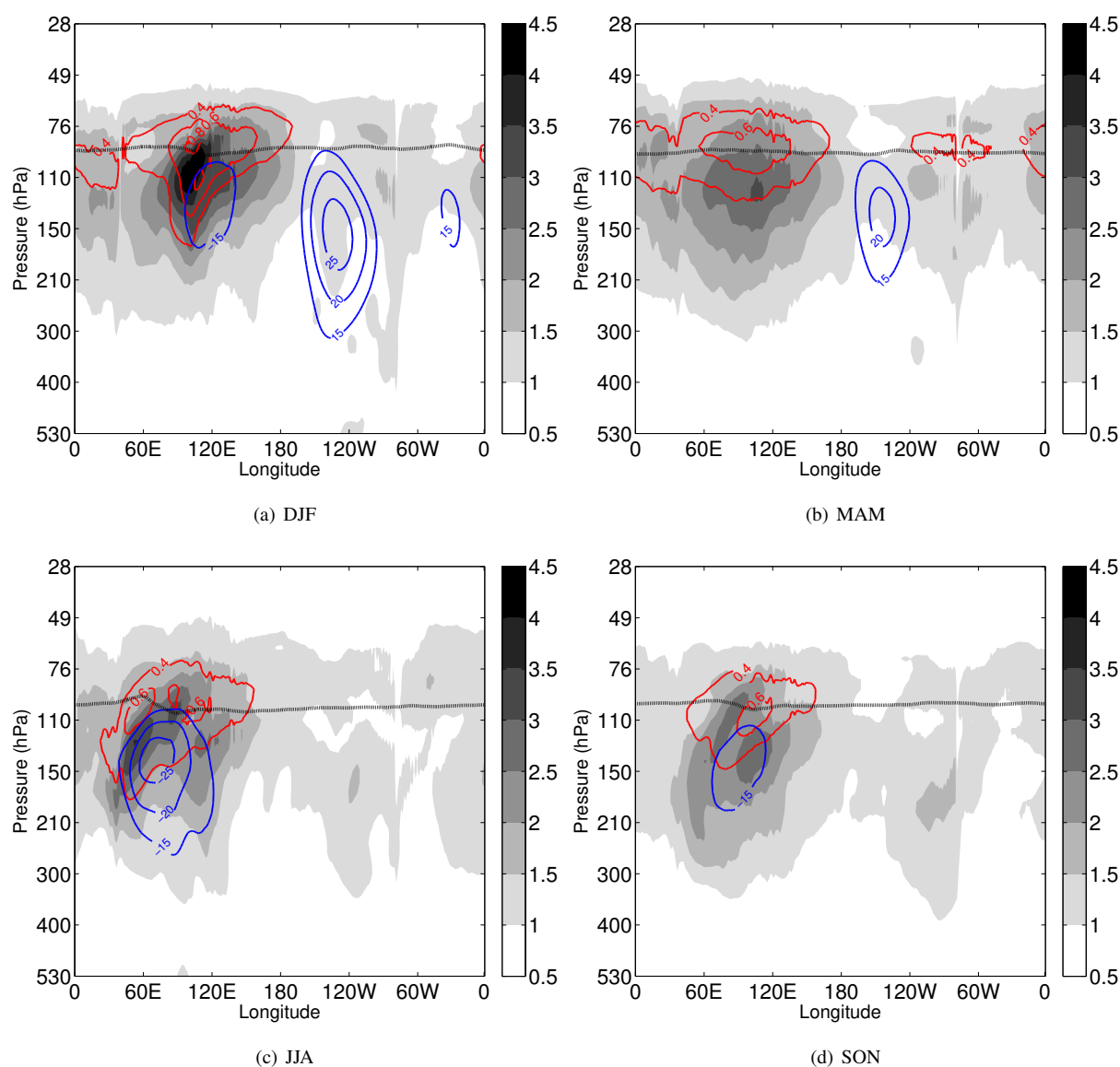


Figure 11. Longitude-pressure sections along the equator (0.7°N) of intraseasonal Kelvin wave zonal wind (white-to-black shades, each 0.5 ms^{-1}) and temperature (red contours, each 0.2 K) fields averaged over the following seasons: NH (a) winter, (b) spring, (c) summer, and (d) autumn. Seasonal-averaging is performed over the absolute values of KW zonal wind and temperature. The background zonal wind (blue contours) and the tropical tropopause height (single thick dotted contour) are defined as in Fig. 8.

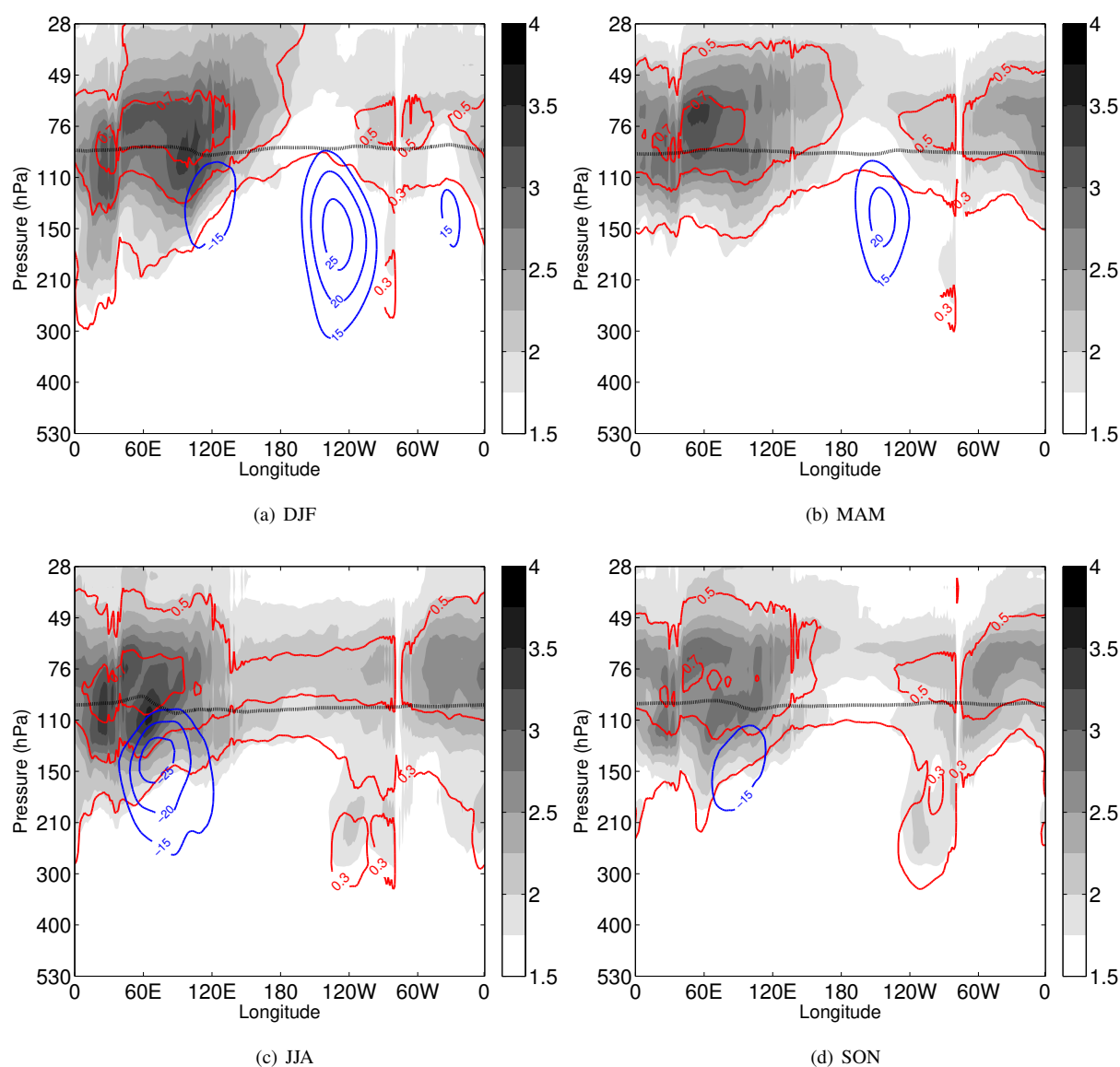


Figure 12. Longitude-pressure sections along the equator (0.7°N) of intramonthly Kelvin wave zonal wind (white-to-black shades, each 0.25 ms^{-1}) and temperature (red contours, each 0.2 K) fields averaged over the following seasons: NH (a) winter, (b) spring, (c) summer, and (d) autumn. Seasonal-averaging is performed over the absolute values of KW zonal wind and temperature. The background zonal wind (blue contours) and the tropical tropopause height (single thick dotted contour) are defined as in Fig. 8.

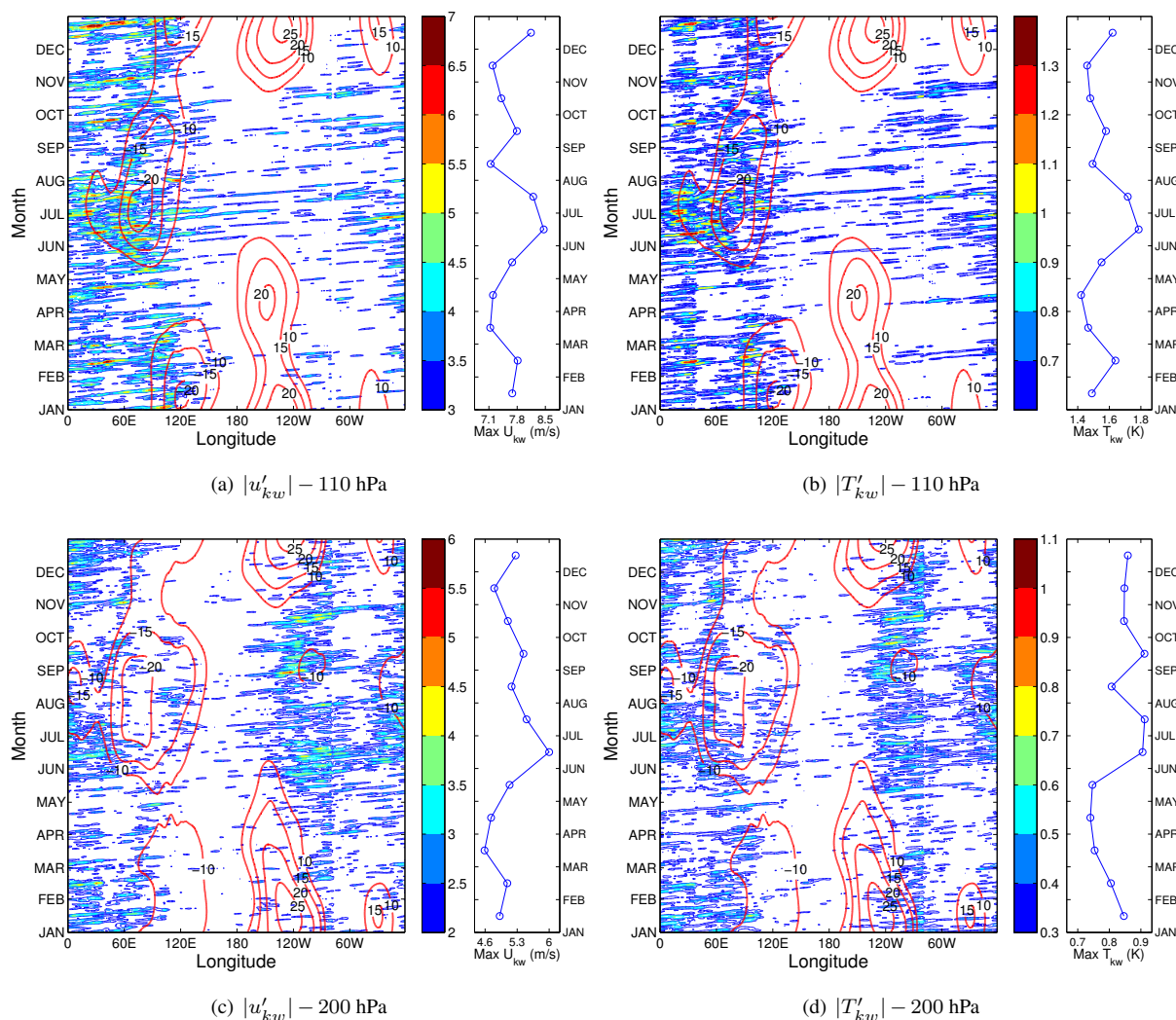


Figure 13. Climatological view of subseasonal KW components (a and c) zonal wind, $|u'_{kw}|$ (blue-to-red shading, each 0.5 ms^{-1}) and (b and d) temperature $|T'_{kw}|$ (blue-to-red shading, each 0.1 K) as function of month in a calendar year and longitude along the equator (0.7°N) at (a and b) model level 40 ($\sim 110 \text{ hPa}$) and (c and d) model level 49 ($\sim 200 \text{ hPa}$). This includes all traveling KWs with periods 3 – 20 days. For comparison, the background zonal wind field is illustrated as well (red contours, each 5 ms^{-1}) at corresponding model levels. On the right side, the circle-marked blue line represents monthly mean values (over all 6+ years) of daily maximum wave amplitudes occurring in longitude on the equator.