We thank the three reviewers for their constructive comments, which helped improve the manuscript substantially. We have revised the manuscript according to the reviewers' comments. Listed below is our point-to-point response in blue to each comment that was offered by the reviewers.

## **Response to Reviewer #1**

## Comments:

The paper presents real-time continuous vertical measurements of particle extinction, gaseous NO2, and black carbon (BC) from ground level to 260 m during two severe winter haze episodes at an urban site in Beijing, China. This study is very interesting and important in helping understand the formation mechanisms and evolution of severe haze episodes in China. I have a few minor issues to be considered before accepting this paper for publication.

1, P1 L14-16, it states "there were four types of vertical profiles with different occurrence rate", but 37%+5%+29%+16%+14%=101% not 100%?

Thank the reviewer's carefulness. 5% here refers to the vertical differences rather than occurrence rate. The four types of vertical profiles account for 37%, 29%, 16%, and 15%, respectively of the total number of vertical profiles in this study, and they together account for 96% with the rest of 4% being other types.

2, The paper states "The travel height is 260m and the container travel at a constant speed of 8 m/s", so it takes 32.5 min to travel up and another 32.5 min to travel back down. But the time in Table S1 shows a very inconsistent travel time, some have travel time of 28.5 min, some have 26 min. Why is that?

Thank the reviewer's comments. In fact, the speed is approximately 9 m min<sup>-1</sup>. The time was not very consistent for each vertical profile because (1) the travel of the

container is controlled by a mechanical gear that failed several times in the middle of the experiments, (2) the container was left at different heights (sometimes 10 m high, and sometimes at ground level) after the "down" experiment. Following the reviewer's comments, we revised the sentence in the new version of the manuscript. It now reads:

"using a container that can travel on the BMT at a relatively constant speed of approximately 9 m min<sup>-1</sup>"

3, Because the measurements at different heights were not measured simultaneously (it had a 30 min lag), the sources and composition of aerosol may change in 30 min especially for local source, leading to biased vertical difference. How to address this?

Thank the reviewer for pointing this out. In this study, we analyzed each vertical profile, in particular those with significant changes as a function of height. To investigate the reasons for the vertical changes, we then analyzed the vertical profiles of meteorological parameters, and also compared the time series of aerosol species measured at ground level and 260 m. We found that the changes in vertical profiles were dominantly associated with meteorological conditions. We agree with the reviewer that the time-lag of local source emissions could affect the vertical profiles. According to previous studies (Sun et al., 2013;Sun et al., 2015), the dominant local source with dramatic changes in a short time is cooking emission. For example, during the period of V11, we observed a large increase in COA in 1 hour. However, such an increase appears not affect the vertical profiles of particle extinction and BC. The reason is that COA has minor impacts on light extinction, and cooking emission is not an important source of BC (Han et al., 2015; Wang et al., 2015; He et al., 2004). Because we didn't measure the vertical profiles of aerosol composition, the time-lag effect could not affect the vertical profiles of light extinction, BC, and NO<sub>2</sub> substantially.

We did observe a very few vertical profiles that were affected by the time-lag effect.

For example, a sudden increase in  $b_{ext}$  was observed in V19 when there was no *T* inversion, RH was constant across different heights, and WS was higher at higher altitudes. We found that such vertical profiles were mainly caused by the rapid increase in PM pollution within a short time (the concentration of NR-PM<sub>1</sub> increased from 30 to 127 µg m<sup>-3</sup> within approximately 20 min from 16:30 to 16:50 on January 12, 2015, Figure 2j). The V19 experiment started at 16:19 when the NR-PM<sub>1</sub> concentration was still low and the increase in the middle (~170 m) coincided with the rapid increase of PM pollution, leading to a significant change in  $b_{ext}$ .

Considering the reasons above, the time-lag effect would not affect our conclusions, and the vertical profiles are expected to representative.

4, P4, L12, it states "fossil OA (FOA) is predominantly from coal combustion emissions", why is this? s it not possible to resolve traffic HOA in Beijing using HR-AMS and ACSM-PMF?

Yes, it is very challenging to separate the traffic-related HOA from coal combustion OA (CCOA) through PMF analysis of unit mass resolution mass spectra of either HR-ToF-AMS or ACSM. The reasons include: (1) very similar spectral patterns between HOA and CCOA at m/z < 120; (2) very similar temporal variations and diurnal cycles; (3) limited sensitivity of the ACSM, particularly for m/z's > 50 with large uncertainties in ion transmission efficiencies. Sun et al. (2016) was able to separate HOA from CCOA by using PMF analysis of high resolution mass spectra and the UMR spectra to m/z = 350, while they were not separated in this study even extending the solution to 6 or 7 factors. Therefore, the two sources are combined into one factor, i.e., fossil fuel related OA (FFOA). Coal combustion OA was the most important primary OA in winter in Beijing, which is much higher than that of HOA (Hu et al., 2016;Sun et al., 2016;Sun et al., 2013). One of the reasons is due to the largely enhanced coal combustion emissions for residential heating while the diesel trucks and heavy-duty vehicles are only allowed inside the Beijing city between 23:00 - 6:00.

5, The ACTRIS ACSM intercomparison study (Crenn et al., 2015) shows that ACSM and HR-AMS measurement would vary (e.g., 36% for ammonium) to some degree even when measuring from the same inlet. Here in this study, the vertical difference between ACSM and HR-AMS would be less meaningful if they would already vary a lot. Do you have inter comparison study between ACSM and HR-AMS measuring from the same inlet?

Good point. We did a two-week inter-comparison between HR-ToF-AMS and ACSM measurements before this study. Although both ACSM and HR-ToF-AMS were calibrated, the ACSM measurements were further corrected using the regression slopes against HR-ToF-AMS measurements from the inter-comparisons to reduce the uncertainties in vertical comparisons.

Following the reviewer's suggestions, we added more details on the inter-comparisons in the revised manuscript. It now reads:

"Considering that the ACSM measurements can have uncertainties of 9 - 36% for different NR-PM<sub>1</sub> species (Crenn et al., 2015), we performed a two-week inter-comparison between ACSM and HR-ToF-AMS measurements at ground site. All submicron aerosol species measured by the ACSM were highly correlated with those measured by the HR-ToF-AMS ( $R^2 > 0.97$ ), and the regression slopes of ACSM against HR-AMS varied from 0.61 to 1.24 for different aerosol species (Sun et al., 2015). To reduce the uncertainties in vertical comparisons, the ACSM measurements were further corrected using the regression slopes determined from the inter-comparisons."

## **Response to Reviewer #2**

## Comments:

General comments:

This manuscript reports results obtained during two haze episodes at Beijing in November 2014 and January 2015. The authors deployed a set of instruments at ground level and on the top of the Beijing Meteorological Tower (260 m) to measure the vertical profile of a few selected parameters (light extinction coefficient, NO2, black carbon, non-refractory PM1, meteorological data). This manuscript is very well written, and is totally relevant for the readers of Atmospheric Chemistry and Physics. I think that the manuscript in its current version is already in a very good shape. However, I have a few minor comments that the authors may consider before final publication.

## Specific comments:

1) Given that the authors are comparing concentrations between ground level and 260 m altitude, I'm wondering whether they need to convert the concentrations in standard temperature and pressure (STP) before doing the comparisons. I know that when people compare aircraft measurements to ground level data, the conversion to STP volumes is very important. Here, between ground level and 260 m, I'm not sure whether the difference of pressure requires this conversion. Among all the parameters that were measured at both altitudes, pressure is the only one for which the vertical profile is not given in this manuscript. I would suggest that the authors include it in the supplementary material (for instance in Figures S1 and S2), and check whether it's worth adjusting the concentrations to standard conditions.



Figure R1 (a) Pressure (P), (b) T at 16 m and 280 m, and (c) the ratios of P/T at 280 m and 16 m.

We thank the reviewer for pointing this out. The pressure was measured at 16 m and 280 m in this study, which was used to evaluate the impacts of pressure on vertical differences. We calculated the ratios of pressure/temperature at the two heights (according to ideal gas law). As shown in Figure R1, the vertical difference caused by the pressure and temperature were both less than 4% during the two severe haze episodes, and the average differences are 2.8% and 2.4% in November and January. Such differences are much smaller than the measurement uncertainties of the ACSM and HR-ToF-AMS, we therefore did not convert the measurements at 260 m to those under standard temperature and pressure conditions.

2) Given that the chemical composition of non-refractory PM1 was measured with an HR-ToF-AMS at ground site and an ACSM on the top of the tower, I would suggest that the authors say a few words on their uncertainties. They can refer to the work of Crenn et al. (2015), who compared a set of 13 ACSMs with an HR-ToF-AMS. Then, the authors can check whether the differences in terms of concentrations and

compositions observed during their study are significant, or whether they are within the uncertainties of the instruments.

Thank the reviewer's comments. We did a two-week inter-comparison between HR-ToF-AMS and ACSM measurements before this study. Although both ACSM and HR-ToF-AMS were calibrated, the ACSM measurements were further corrected using the regression slopes against HR-ToF-AMS measurements from the inter-comparisons to reduce the uncertainties in vertical comparisons.

Following the reviewer's suggestions, we added more details on the inter-comparisons in the revised manuscript. It now reads:

"Considering that the ACSM measurements can have uncertainties of 9 - 36% for different NR-PM<sub>1</sub> species (Crenn et al., 2015), we performed a two-week inter-comparison between ACSM and HR-ToF-AMS measurements at ground site. All submicron aerosol species measured by the ACSM were highly correlated with those measured by the HR-ToF-AMS ( $R^2 > 0.97$ ), and the regression slopes of ACSM against HR-AMS varied from 0.61 to 1.24 for different aerosol species (Sun et al., 2015). To reduce the uncertainties in vertical comparisons, the ACSM measurements were further corrected using the regression slopes determined from the inter-comparisons."

3) Still concerning these two instruments, I'm curious to know how the authors performed the PMF analysis for the ACSM. Did they use the results obtained with the HR-ToF-AMS to choose the final PMF result for the ACSM (number of factors and fPeak)? I think that the authors need to add some evaluation plots for the 4-, 5-, and 6-factor solutions in the supplementary material, in order to justify the choice of the 5-factor solution for the two instruments. Here also, the authors can refer to the same inter-comparison between the HR-ToF-AMS and ACSMs (Fröhlich et al., 2015). In that work, the authors had noticed that some PMF factors can be quite difficult to separate in some ACSM datasets (especially the separation of COA from HOA).

We thank the reviewer's comments. We expanded the details on PMF and ME2 analysis substantially in the revised manuscript. The detailed evaluation of PMF results and the a-value based ME-2 solutions were given in Zhou et al. (2017).

"Positive matrix factorization (PMF) (Paatero and Tapper, 1994) was first performed to the unit mass resolution spectra of OA at ground level that were measured with HR-ToF-AMS during the same period as that of ACSM, and five factors including three primary OA (POA) factors, i.e., fossil fuel related OA (FFOA) predominantly from coal combustion emissions, cooking OA (COA), and biomass burning OA (BBOA), and two secondary OA (SOA) factors, i.e., less oxidized oxygenated OA (LO-OOA) and more oxidized OOA (MO-OOA) were identified. To better compare the OA factors between ground level and 260 m, the multi-linear engine 2 (ME-2) (Canonaco et al., 2013;Crippa et al., 2014) using the mass spectral profiles of three POA factors at ground level as constrains was performed to the ACSM OA spectra. In addition to the three POA factors, a LO-OOA and a MO-OOA were also resolved. It should be noted that such an approach could introduce some uncertainties for OA source apportionment at 260 m because POA factors are not exactly the same between ground level and 260 m. We also performed PMF analysis on ACSM OA spectra, and found that the BBOA factor cannot be resolved although biomass burning is a common source in winter. The detailed evaluation of PMF results and the a-value based ME-2 solutions were given in Zhou et al. (2017)."

4) Among the five PMF factors identified with the HR-ToF-AMS and ACSM, three factors correspond to primary particles directly emitted at ground level (FOA, COA, BBOA). I'm wondering whether the authors can do a comparison between their mass spectra (Figure S3), and check whether some specific signals changed significantly between ground level and 260 m (following photo-oxidation, for instance). I guess this comparison should be quite difficult, given that the instrument on the top of the tower was the ACSM (mass spectra in unit mass resolution).

Thank the reviewer's comments. The three POA factors at 260 m were determined

using a-value based ME-2 analysis. The mass spectra profiles of three POA factors resolved at ground site were used as constrains. In addition, PMF analysis was also performed to ACSM OA spectra, and only two primary OA factors were identified. The comparisons of mass spectral profiles for different a-values and also PMF results are shown in Figure R2. More detailed descriptions of the results are given in Zhou et al. (2017).





Figure R2: Mass spectra (left panel) and time series (right panel) of three POA factors resolved at 260 m by ACSM using multi-linear engine 2 (ME-2): (a) fossil fuel related OA (FFOA), (b) cooking OA (COA), and (c) biomass-burning OA (BBOA). The 4-factor solution of PMF results is also shown.

5) When I take a look at the vertical profiles of temperature during the 36 periods (Figures S1 and S2), I notice a strong decrease of the temperature at high altitude for at least 23 of these periods. This kind of vertical profile can have an incidence on the gas-particle partitioning of a few semi-volatile species (I'm especially thinking about

nitrate), which can condense more at high altitude. The authors can include discussion on this in the manuscript, for instance on page 11, lines 5-6, where the authors mention a higher contribution of nitrate at 260 m.

Thank the reviewer's careful review. We agree with the reviewer that gas-particle partitioning can have a significant impact on the formation of nitrate. We note that the highest temperature during the two severe haze episodes was approximately 15°C and 8°C, respectively. The evaporative loss of ammonium nitrate under such low temperature conditions could not be important. Therefore, gas-particle partitioning formation of nitrate is not expected to be as important as that observed in summer. Comparatively, the vertical differences in O<sub>3</sub>, NO<sub>2</sub>, and solar radiation could be more important for the nitrate differences via heterogeneous reactions of N<sub>2</sub>O<sub>5</sub> at nighttime and photochemical production during daytime. Unfortunately, we didn't have such measurements to evaluate such impacts. As the reviewer mentioned, the higher nitrate contribution at 260 m (page 11, lines 5-6) was mainly caused by the lower contribution of organics. In fact, the absolute nitrate concentration at 260 m was lower than that at ground site. The large increase in OA from local sources at ground site led to a decrease of relative contribution of nitrate in NR-PM<sub>1</sub>. Similarly, the sulfate contribution at ground site is also lower than that at 260 m. Therefore, it is very challenging to quantify the impact of gas-particle partitioning on the vertical differences of nitrate in this study. Such impacts should be investigated with more comprehensive measurements in the future studies.

## Technical comments:

6) Page 10, line 28: The SSA values are given in a certain range (i.e. min-max) for Fall 2014 and avg std dev for Winter 2015. Please choose one of the formats and use the same for the two periods, just to be consistent.

## It was revised.

7) Caption of Figure 5: "260 m (top panel). The".

Corrected.

### **Response to Prof. M. Wiegner**

I agree with the two anonymous reviewers that this is an interesting and well written paper. This concerns not only the "chemical" aspects mentioned but also the determination of the mixing layer height (MLH): it is estimated from extinction coefficient profiles (aethalometer onboard of a vertically moving container) and attenuated backscatter profiles (Vaisala CL51 ceilometer). Though both parameters are different they offer an excellent opportunity for intercomparisons, as both are related to aerosol optical properties. Consequently, the authors briefly cover this topic and conclude that the MLH tends to be overestimated when using the ceilometer. I suggest to discuss this interesting application of the data in more depth:

## Thank Prof. M. Wiegner for your comments.

• From the reference Tang et al. (2016) it can be inferred that the authors use BL-VIEW. It has been shown by Geiß et al. (2017, Atmos. Meas. Tech.), that depending on the different options the retrieved MLH can be different. Thus, it would be interesting to include a few additional information on how the authors determine the MLH. By the way: Geiß et al. also found that the different versions of BL-VIEW tend to (slightly) overestimate the MLH. The quantitative criteria underlying the determination of the "transition height" from  $b_{\text{ext}}$  should be outlined as well (or be stated, that it is from visual inspection).

We thank the reviewer's careful review. Following the reviewer's suggestions, we added more information about the retrieval method of CL51 data for MLH in our study.

## It now reads:

"The Vaisala software product BL-VIEW (version 2.0) was used to identify the MLH with the gradient method. The temporal and vertical attenuated backscatter coefficients were first smoothly averaged to avoid the effect of noise and interference

from the aerosol layering structure, and the maximum negative gradient value  $(-d\beta/dx)$  was then determined as the top of the mixing layer (Münkel et al., 2007;Geiß et al., 2017;Zhu et al., 2016)."

The transition height was determined mainly from the visual perspective, which was added in the revised manuscript. An uncertainty of  $\sim 5$  m is expected.

• A figure showing all coincident MLH-retrievals would be interesting. Fig. 11 – only shown in the conclusions – seems to provide this. I assume that "5", "6" etc. along with the MLH-curve correspond to Table S1? As the number of cases is relatively low (because quite often the MLH is larger than 260 m) this is not obvious and thus should be clearly emphasized. Maybe, the figure should be moved to the results-section. Fig. 7e is less suitable to demonstrate the differences as it does not cover the full set of measurements, and the reader might be confused from the two different vertical scales.

Thank the reviewer's comments. The numbers in Fig. 11 corresponded to those in Table S1 and Figure 1. We agree with the reviewer that the number of cases is relatively low because only vertical profiles with significant changes can be used to estimate the transition heights. Following the reviewer's suggestions, we clearly stated this point in the revised manuscript.

"Because of the relatively low number of cases and the limited height of the meteorological tower, future vertical measurements of particle extinction to a high altitude, e.g., using tethered balloon are needed to further validate the retrieval of MLH from CL51 measurements."

Fig. 7e is mainly used to demonstrate the evolution of vertical profiles during the severe haze episode. For clarity, we claimed "right axis" in the figure caption. Also, a clearer description "Note that only vertical profiles with significant changes were used for estimation of transition heights." was added in Figure 11.

• It is known that an overlap correction function is applied to ceilometer measurements, well below 260 m in case of the CL51. Consequently, comparisons with the independent extinction coefficient measurements could be a promising approach to check the plausibility of this correction, whenever adequate atmospheric conditions occur (e.g., no rapid changes of the aerosol distribution). Though a strict validation might be difficult and beyond the scope of this paper, it could be briefly discussed in section 4 whether or not this would be a possible extension of this study.

We totally agree with the reviewer. "future vertical measurements of particle extinction to a high altitude, e.g., using tethered balloon are needed to further validate the retrieval of MLH from CL51 measurements" was added in the revised text.

• A "definition" of "severe haze episodes" in terms of aerosol optical depth would be interesting: In case of very large AOD the ceilometer might not fully penetrate the mixing layer. From the extinction coefficient profiles (aethalometer) it seems that the AOD is however clearly below 1 (and thus not critical). Do conditions occur in Beijing when this is not true?

Yes, it is interesting to check the relationship between air pollution level and AOD. Indeed, several previous studies found the positive correlations between AOD and PM<sub>2.5</sub>, even at high PM<sub>2.5</sub> levels (> 200  $\mu$ g m<sup>-3</sup>), and the slopes are strongly regional dependent. Therefore, AOD could be used to judge the severe haze episodes. Concerning the penetration of mixing layer during severe haze episodes, the CL51 appears to work relatively well as indicated by Figure R3 when the NR-PM<sub>1</sub> was the highest in November. However, an accurate evaluation of such impacts needs vertical measurements to a high altitude, for example, > 1 km.





Technical comments:

• Page 5, line 32: include brackets around "2 mean".

# Corrected.

• Page 6, line 20: when referring to V3 and V24 etc. it would be useful to mention

Table S1 again, where the nomenclature is (more or less) explained.

Table S1 was added.

• Define the light green and the dark green curve in Fig. 7e.

The two green lines refer to the time series of NR-PM<sub>1</sub> at ground level and 260 m, which are described in the figure legend (above Fig. 7e).

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# Vertically-resolved Characteristicscharacteristics of Air Pollutionair pollution during Two Severe Winter Haze Episodestwo severe winter haze episodes in Urbanurban Beijing, China

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Abstract. We conducted the first real-time continuous vertical measurements of particle extinction ( $b_{ext}$ ), gaseous NO<sub>2</sub>, and black carbon (BC) from ground level to 260 m during two severe winter haze episodes at an urban site in Beijing, China. Our results

- 15 illustrated four distinct types of vertical profiles: 1) uniform vertical distributions (37% of the time) with vertical differences less than 5%; 2) higher values at lower altitudes (29%); 3) higher values at higher altitudes (16%), and 4) significant decreases at the heights of ~100 - 150 m (14%). Further analysis demonstrated that vertical convection as indicated by mixing layer height, temperature inversion, and local emissions are three major factors affecting the changes in vertical profiles. Particularly, the formation of Type 4 was strongly associated with the stratified layer that was formed due to the interactions of different air masses and temperature
- 20 inversions. Aerosol composition was substantially different below and above the transition heights with  $\sim 20 30\%$  higher contributions of local sources (e.g., biomass burning and cooking) at lower altitudes. A more detailed evolution of vertical profiles and their relationship with the changes in source emissions, mixing layer height, and aerosol chemistry was illustrated by a case study. BC showed overall similar vertical profiles as those of  $b_{\text{ext}}$  ( $R^2 = 0.92$  and 0.69 in November and January, respectively). While NO<sub>2</sub> was correlated with  $b_{\text{ext}}$  for most of the time, the vertical profiles of  $b_{\text{ext}}/NO_2$  varied differently for different profiles, indicating
- 25 the impact of chemical transformation on vertical profiles. Our results also showed that more comprehensive vertical measurements (e.g., more aerosol and gaseous species) at higher altitudes in the megacities are needed for a better understanding of the formation mechanisms and evolution of severe haze episodes in China.

#### **1** Introduction

Air pollution is a severe environmental problem in China (Chan and Yao, 2008;Zhang et al., 2015), which offers a great
 challenge for future air quality improvement and economic development. Severe haze episodes with surprisingly high concentration of PM<sub>2.5</sub> (particles with aerodynamic diameters less than 2.5 µm) dominantly occur in fall and winter seasons. However, current air

quality models often fail to accurately predict extreme haze episodes (Wang et al., 2014b;Zheng et al., 2015a;Wang et al., 2014a;Wang et al., 2014c). One reason is the incomplete understanding of the formation mechanisms of haze pollution. For example, the Weather Research and Forecasting - Community Multi-scale Air Quality (WRF-CMAQ) model showed significant improvements in simulating sulfate and nitrate concentrations and temporal variations in January 2013 episode by incorporating heterogeneous

- 5 reaction mechanisms (Zheng et al., 2015a), yet it failed to predict the haze peak on January 12–13 that was found to be mainly caused by regional transport (Ji et al., 2014;Sun et al., 2014). However, there are strong arguments on the role of regional transport in the haze pollution (Guo et al., 2014;Li et al., 2015b), especially during severe haze episodes with stagnant meteorological conditions and shallow boundary layer (Sun et al., 2014;Zheng et al., 2015b;Quan et al., 2013). Therefore, vertical characterization of air pollutants is critical for elucidating the formation and transport of regional haze events, especially for severe haze episodes with boundary layer
- 10 height less than 300 m (Quan et al., 2013; Tang et al., 2015; Wiegner et al., 2006).

Extensive measurements have been conducted at ground sites for characterization of the composition, sources and formation mechanisms of severe haze episodes in north China (Huang et al., 2014;Ji et al., 2014;Zhao et al., 2013;Sun et al., 2013b). However, vertical measurements are rather limited, particularly in the megacities in north China. Zhang et al. (2009) analyzed the vertical distributions of aerosol number and volume concentrations during 17 aircraft measurements in the spring of 2005 and 2006. Three

- 15 different types of vertical profiles were observed, which were mainly affected by meteorological conditions. Chen et al. (2009) further analyzed the vertical measurements of gaseous pollutants in Beijing in summer 2007 and found a significant impact of mountain-valley breeze on the vertical distributions of pollutants. Recent aircraft measurements for black carbon (BC) showed different vertical profiles between southern and northern air masses, and an enhanced regional transport between 0.5 and 1 km (Zhao et al., 2015). However, most aircraft measurements in the megacity of Beijing were conducted above ~300–500 m, while the vertical
- 20 distributions in the lower levels of the boundary layer (e.g., <300 m) are rarely characterized. In addition, how the vertical profiles evolve during an entire cycle of haze episode is poorly understood due to the limited aircraft measurements. Although recent measurements with tethered balloons provide more insights into the vertical characteristics of air pollutants (e.g., BC) (Ran et al., 2016;Li et al., 2015a), most of them were conducted at rural sites, which might be significantly different from urban areas with more complex pollution sources and land surfaces.</p>
- 25 The meteorological tower is a unique platform to study the vertical characteristics of air pollutants in the lower levels of the boundary layer. For example, the vertical profiles of trace gases and aerosol species from 3 to 270 m were comprehensively characterized using a 300 m tower at the Boulder Atmospheric Observatory (BAO) during late winter 2011 (Brown et al., 2013), which provides many new insights into the sources of chemical species (e.g., local emissions, regional transport, and point sources) and aerosol processing at different heights (Kim et al., 2014;Öztürk et al., 2013b). Because the BAO tower is located at a suburban area with low aerosol
- 30
  - mass loadings, the vertical profiles of aerosol and gas species (VandenBoer et al., 2013;Riedel et al., 2013;Öztürk et al., 2013a) could be substantially different from those in megacities in China. The Beijing 325 Meteorological tower (BMT) is located between the north 3<sup>rd</sup> and 4<sup>th</sup> ring road in the city center. Compared with the BAO tower, the land surface, sources emissions and atmospheric processes of aerosol particles at the BMT are far more complex. Early studies on the BMT were mainly focused on trace gases, PM<sub>2.5</sub>

and filter measurements (Meng et al., 2008;Sun et al., 2013a;Sun et al., 2009). The results showed that the mixing ratio of  $O_3$  often peaked at ~120 m and was maintained at higher levels in the residual layer at nighttime (Li et al., 2003), while  $SO_2$  was found to have the highest mixing ratio at ~50 m (Meng et al., 2008). Filter measurements on the BMT showed higher ratios of organic carbon to elemental carbon at higher altitudes, indicating the vertical differences in secondary formation (Chan et al., 2005). Tao et al. (2007)

- 5 further found enhanced gas-particle partitioning of polycyclic aromatic compounds at higher heights likely due to the lower temperature and higher relative humidity. However, most previous vertical measurements are either limited by the time resolution, e.g., 1–3 days for filter collection, or vertical resolution, e.g., typically 3 or 4 heights. Our understanding of the vertical profiles and evolution of air pollutants, and their interactions with the lower boundary layer during severe haze episodes is far from complete. Until recently, Sun et al. (2015;2016) and Zhao et al. (2017) conducted real-time simultaneous measurements of aerosol particle
- 10 composition at ground level and 260 m using two aerosol mass spectrometers. The results showed largely different characteristics of primary and secondary aerosols between ground level and 260 m, elucidating the different contributions of local emissions and regional transport to air pollution at different heights. In addition, the interactions between meteorological parameters and vertical differences were also illustrated by several case studies. Because these measurements were conducted at two fixed heights, i.e., ground level and 260 m, the vertical variations in the middle were never known, particularly during winter severe haze episodes with
- 15 the boundary layer height lower than 300 m (Tang et al., 2015).

In this study, we conducted the first continuous vertical measurements of particle extinction ( $b_{ext}$ ), NO<sub>2</sub> and BC between ground level to 260 m in the megacity of Beijing during two severe winter haze episodes along with synchronous measurements of aerosol particle composition at ground level and 260 m with two aerosol mass spectrometers. The types of vertical profiles are identified and their relationships with meteorological parameters and aerosol chemistry are elucidated. A more detailed evolution of vertical profiles and their interactions with boundary layer during the severe haze episode is further illustrated by a case study. Finally, the vertical profiles and relationships between  $b_{ext}$ , NO<sub>2</sub>, and BC are investigated.

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#### **2** Experimental Methods

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The vertically resolved measurements were conducted at the Tower Branch of Institute of Atmospheric Physics, Chinese Academy of Sciences using a container that can travel on the BMT at a <u>relatively</u> constant speed of <u>Sapproximately 9</u> m min<sup>-1</sup>. The instruments including a Cavity Attenuated Phase Shift Particulate Matter extinction monitor (CAPS-PM<sub>ext</sub>, Aerodyne Research Inc.) (Kebabian et al., 2007), a CAPS-NO<sub>2</sub> monitor (Kebabian et al., 2008;Ge et al., 2013), and a 7-wavelength Aethalometer (model: AE33; Magee Scientific Corp.) (Drinovec et al., 2015) were equipped in the container for real-time measurements of light extinction coefficient (*b<sub>ext</sub>*,  $\lambda = 630$  nm) of dry fine particles, gaseous NO<sub>2</sub>, and BC with a time resolution of 1 s, 1 s and 1 min, respectively. Because all instruments were powered by uninterruptible power supply, the vertical measurements were performed approximately every 4 h from ground level to 260 m during daytime and every 6 h from ground to 200 m during nighttime for safety, with the rest time being

ground measurements. In this study, the vertical experiments were conducted during two severe haze episodes that occurred during November 19–21, 2014, and January 12–16, 2015, and in total, 72 vertical profiles were obtained (Table S1, and Figures S1 and S2).

In addition, non-refractory submicron aerosol (NR-PM<sub>1</sub>) species including organics, sulfate, nitrate, ammonium and chloride, were measured simultaneously at ground level and 260 m on the tower with an Aerodyne High-Resolution Time-of-Flight Aerosol Mass Spectrometer (HR-ToF-AMS) and an Aerosol Chemical Speciation Monitor (ACSM), respectively. The detailed sampling setup and calibration of the two instruments, and data analysis have been described in Xu et al. (2015b) and Chen et al. (2015). Positive matrix 5 factorization (PMF) (2015) and Chen et al. (2015). Considering that the ACSM measurements can have uncertainties of 9 - 36% for different NR-PM<sub>1</sub> species (Crenn et al., 2015), we performed a two-week inter-comparison between ACSM and HR-ToF-AMS measurements at ground site. All submicron aerosol species measured by the ACSM were highly correlated with those measured by the HR-ToF-AMS ( $R^2 > 0.97$ ), and the regression slopes of ACSM against HR-AMS varied from 0.61 to 1.24 for different aerosol species (Sun et al., 2015). To reduce the uncertainties in vertical comparisons, the ACSM measurements were further corrected using 10 the regression slopes determined from the inter-comparisons. Positive matrix factorization (PMF) (Paatero and Tapper, 1994) (Paatero and Tapper, 1994) was also performed to the mass spectra of organic aerosols (OA). Five OA factors were resolved at both ground site and 260 m, which include three primary OA (POA) factors, i.e., fossil OA (FOA) predominantly from coal combustion cooking OA (COA), and biomass burning OA (BBOA), and two secondary OA (SOA) factors, i.e., less oxidized oxygenated OA (LO-OOA) and more oxidized OOA (MO-OOA). The mass spectral profiles of five OA factors are presented in Figure S3.- was first performed to the unit mass resolution spectra of OA at ground level that were measured with HR-ToF-AMS 15 during the same period as that of ACSM, and five factors including three primary OA (POA) factors, i.e., fossil fuel related OA (FFOA) predominantly from coal combustion emissions, cooking OA (COA), and biomass burning OA (BBOA), and two secondary OA (SOA) factors, i.e., less oxidized oxygenated OA (LO-OOA) and more oxidized OOA (MO-OOA) were identified. To better compare the OA factors between ground level and 260 m, the multi-linear engine 2 (ME-2) (Canonaco et al., 2013;Crippa et al., 20 2014) using the mass spectral profiles of three POA factors at ground level as constrains was performed to the ACSM OA spectra. In addition to the three POA factors, a LO-OOA and a MO-OOA were also resolved. It should be noted that such an approach could introduce some uncertainties for OA source apportionment at 260 m because POA factors are not exactly the same between ground level and 260 m. We also performed PMF analysis on ACSM OA spectra, and found that the BBOA factor cannot be resolved although biomass burning is a common source in winter. The detailed evaluation of PMF results and the a-value based ME-2 solutions were given in Zhou et al. (2017).

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Meteorological variables (temperature (T), relative humidity (RH), wind speed (WS), and wind direction (WD)) were measured at 15 heights on the BMT (8, 15, 32, 47, 63, 80, 102, 120, 140, 160, 180, 200, 240, 280, and 320 m) with four-cup anemometers (model O1OC, Met One Instruments), and standard meteoprobe (model HC2-S3, ROTRONIC). The wind field data, T, pressure, CO<sub>2</sub> and H<sub>2</sub>O were measured by ultrasonic anemometers (Gill Instruments Limited, Lymington, UK), and Li-7500CO<sub>2</sub>/H<sub>2</sub>O gas analyzers (LI-

COR, Inc., Nebraska, USA) at 7 heights (8, 15, 47, 80, 140, 200, and 280 m) on the BMT with a time resolution of 10 Hz(Liu et al., 30 2015)(Liu et al., 2015). The turbulent kinetic energy (TKE) of motion (e) along x, y, z directions is calculated as  $e_u = (u - \bar{u})^2$ ,  $e_v = (u - \bar{u})^2$  $(v - \bar{v})^2$ , and  $e_w = (w - \bar{w})^2$ , respectively, of which  $\bar{u}, \bar{v}$ , and  $\bar{w}$  are the 20 min averages of u, v, and w, respectively.  $E_u, E_v$  and  $E_w$ 

are then defined as the 20 min averages of  $e_u$ ,  $e_v$  and  $e_w$ , respectively. Virtual potential temperature (VPT) was calculated with meteorological variables measured at 7 heights.

In addition, a Doppler wind lidar (Windcube 200, Leosphere, Orsay, France) was deployed at the same location to measure wind profiles from 100 to 5000 m with a spatial resolution of 50 m and a time resolution of 10 min (Sun et al., 2016b)(Sun et al., 2016), and a single-lens Ceilometer (CL51, Vaisala, Finland) was used to measure the vertical attenuated backscatter coefficient (vertical resolution: 10 m). The mixing layer height (MLH) was then retrieved using the algorithm that has been detailed in Tang et al., 2016).(2015; 2016). The Vaisala software product BL-VIEW (version 2.0) was used to identify the MLH with the gradient method. The temporal and vertical attenuated backscatter coefficients were first smoothly averaged to avoid the effect of noise and interference from the aerosol layering structure, and the maximum negative gradient value ( $-d\beta/dx$ ) was then determined as the top of

10 the mixing layer (Münkel et al., 2007;Geiß et al., 2017;Zhu et al., 2016). The NR-PM<sub>1</sub> aerosol composition and the meteorological variables at ground level were input into the ISORROPIA-II (Nenes et al., 1998) model to predict liquid water content (LWC) associated with inorganic species.

#### **3** Results and discussion

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#### 3.1 Descriptions of the two severe haze episodes

- 15 The formation of the two haze episodes in Figure 1 was both initiated by a change in WD from the north-northwest to south-southwest, and then lasted approximately 3–4 days. However, the variations in NR-PM<sub>1</sub> were different between the two episodes. During the first episode in November (Ep1), the mass concentration of NR-PM<sub>1</sub> increased rapidly from less than 50 µg m<sup>-3</sup> to ~250 µg m<sup>-3</sup> in a day, and then remained at relatively constant levels until a further steep increase on November 21 due to a high RH and LWC event as indicated in Figures 1 and S4. The variations in NR-PM<sub>1</sub> were overall consistent with those of H<sub>2</sub>O which increased from 3 to 5.6 g m<sup>-3</sup> during this episode (Figure S5). In contrast, the variations of NR-PM<sub>1</sub> during the second episode in January (Ep2)
- were relatively more stable with the concentrations ranging from 100 to 150  $\mu$ g m<sup>-3</sup> until 12:00 on January 15 when steep increases in NR-PM<sub>1</sub> at both ground level and 260 m were observed. The average mass concentrations of NR-PM<sub>1</sub> during Ep1 were 171 and 125  $\mu$ g m<sup>-3</sup> at ground level and 260 m, respectively, which are 30% and 10% higher than those observed during Ep2, indicating that the PM pollution during Ep1 was more severe than during Ep2. A more detailed analysis showed that the higher PM during Ep1 was
- 25 mainly caused by the lower mixing layer height (mean:  $278 \pm 109$  m) than that during Ep2 (mean:  $498 \pm 314$  m) (Wang et al., 2010). As shown in Figure S4, the MLH in January was ubiquitously higher than that in November. For example, the MLH was below 400 m for most of the time during Ep1, while it was often above 600 m in daytime during Ep2. This is consistent with a more frequent occurrence of *T* inversions during Ep1 than Ep2, for example, *T* inversions were observed at every night from November 19 to 21 (Figure 1d), while only two short periods were observed on January 15 – 16 (Figure 1i).
- 30 The temporal variations of NR-PM<sub>1</sub> at ground level and 260 m were different between Ep1 and Ep2. While substantial vertical differences were observed for most of the time during Ep1, the variations of NR-PM<sub>1</sub> were overall similar between ground level and

260 m during Ep2. The average vertical difference of NR-PM<sub>1</sub> is 46  $\mu$ g m<sup>-3</sup> during Ep1, which is much higher than the 18  $\mu$ g m<sup>-3</sup> during Ep2. Figure 1 also shows that the vertical difference in NR-PM<sub>1</sub> was tightly related to *T* inversion, and larger differences typically occurred during periods with clear *T* inversions (Figure 1d). The average composition of NR-PM<sub>1</sub> was overall similar between ground level and 260 m during Ep1 and Ep2, except slightly higher contributions of organics and nitrate at 260 m. While

5 organics dominated NR-PM<sub>1</sub> both at ground level (43%) and 260 m (46–47%), the composition of inorganic aerosols was substantially different between Ep1 and Ep2. Particularly, the episode in January showed much higher contributions of sulfate (21–24%) than Ep1 (15–16%). Higher sulfate during Ep2 was likely due to higher RH that facilitated the aqueous-phase processing of SO<sub>2</sub> to form sulfate.

#### 3.2 Vertical profiles of pollutants

- 10 The vertical profiles of  $b_{\text{ext}}$ , NO<sub>2</sub> and BC varied substantially during the formation, evolution and clearing stages of the haze episodes, and strongly depended on the vertical changes in meteorological parameters. In this study, we analyzed each vertical profile by calculating the differences in vertical changes (= (max-min)/)/(2×mean<sub>7</sub>), Figure 2), performing linear fit analysis on the correlations between extinction and height, and checking the presence of *T* inversions (Figures 2 and 3). Overall, four distinctly different types of vertical profiles were categorized (Figure 4). Type 1 is well mixed as indicated by the small vertical differences (< 5%) and slopes (-
- 15  $0.3 0.3 \text{ Mm}^{-1} \text{ m}^{-1}$ ). Type 2 shows clear decreases as a function of heights, which is mainly categorized according to the large vertical differences (> 5%). In addition, all vertical profiles for Type 2 present negative slopes, and the slope values depend on the absolute extinction values. The vertical profiles of Type 3 are characterized by increases as a function of height, and all the slopes for this type are positive although the vertical differences can vary substantially. The vertical profiles of Type 4 showed significant decreases within a short height interval. Another major difference between Type 4 and Type 3 is the strong *T* inversion that occurred
- 20 at a height of less than 150 m. The four types together account for 96% of the total profiles, which can be considered representative during the two haze episodes. A more detailed summary of the four types of vertical profiles is presented in Tables S2 – S5.

(1) Uniform vertical distributions. As shown in Figure 4a, the vertical variations for the three vertical profiles of  $b_{\text{ext}}$  were small, and the vertical differences were all less than 5%, indicating that aerosol particles were relatively well mixed during these three periods. Such types of vertical profiles (37% of the total profiles) were typically observed during daytime, e.g., 10:00–16:00 when *T* and TKE

- are high (Figure S5), and MLH is above 260 m (370–1000 m here). Also, there were no *T* inversions, and the VPT showed small vertical variations (Figure S6) supporting relatively well mixed lower boundary. NO<sub>2</sub> showed remarkably similar vertical profiles to  $b_{\text{ext}}$ , which is indicated by the constant ratios of  $b_{\text{ext}}/\text{NO}_2$  across different heights (Figure 4a). The vertical profiles of *T* and RH during these three measurements were similar. While *T* showed a gradual decrease as the increase with increasing height, RH was relatively even across different heights. Note that the differences of meteorological variables in absolute values were significant, for example,
- 30 the temperature difference between V3 (vertical profile of 3, <u>Table S1</u>) and V24 (vertical profile of 24, <u>Table S1</u>) is  $\sim$ 6°C, and the wind speed of V24 ( $\sim$ 2.5 m s<sup>-1</sup>) is nearly 70% higher than that of V3 ( $\sim$ 1.5 m s<sup>-1</sup>). In addition, the TKE for V3 ( $\sim$ 0.2 m<sup>2</sup> s<sup>-2</sup>) is less than half of that of V24 ( $\sim$ 0.6 m<sup>2</sup> s<sup>-2</sup>) across different heights (Figure S5), indicating a more stagnant condition for V3. Such

meteorological differences might be one of the major reasons for the different  $b_{ext}$  levels, e.g., 454 Mm<sup>-1</sup> for V24 (vertical profile of 24) and 1482 Mm<sup>-1</sup> for V3, and also largely different  $b_{ext}/NO_2$  ratios among different vertical profiles (e.g., 8.8 Mm<sup>-1</sup> ppb<sup>-1</sup> for V24 and 16.5 Mm<sup>-1</sup> ppb<sup>-1</sup> for V3). We note slightly different aerosol composition between ground level and 260 m (Figure 5a) although the vertical differences in  $b_{ext}$  and the total mass concentrations of NR-PM<sub>1</sub> were small. For example, V24 showed 2% higher

5 contribution of sulfate and 4% lower organics in NR-PM<sub>1</sub> at ground level compared with that at 260 m. While secondary formation influenced by RH/T and solar radiation could be slightly different, the relative importance of local emissions versus regional transport at different heights might have also played a role.

(2) Higher values at lower altitudes. Such vertical profiles were frequently observed during this study (29% of the time), which were typically characterized by smooth decreases as a function of height, yet the vertical differences were larger than 5% (Figure 4b). The decreasing rates of b<sub>ext</sub> varied mostly between 0.4 – 1.4 Mm<sup>-1</sup> m<sup>-1</sup> except the profiles during clean periods (e.g., V17). Most of these profiles (60%) occurred at the time between 10:00 – 11:00 and 16:00 – 18:00 when the vertical convection was moderately high as indicated by E<sub>w</sub> in Figure S7. These results suggest that such vertical changes were mainly caused by the dilution effects associated with vertical convection. Another reason is the stronger local emissions at ground level, e.g., cooking emissions during lunch and dinner times (Sun et al., 2013b;Xu et al., 2015), yet it is slow to be mixed to a higher height, particularly at night with shallower
boundary layer. For example, Sun et al.(Sun et al., 2015b) found much larger ratios between 260 m and ground level for local miner times than regioned accorders accorder to result the intensities of lacel emissions are acfect the vertical primery energies than regioned accorders accorder to a night the the intensities of lacel emissions.

- primary species than regional secondary aerosols at nighttime, indicating that the intensities of local emissions can affect the vertical profiles substantially.(2015) found much larger ratios between 260 m and ground level for local primary species than regional secondary aerosols at nighttime, indicating that the intensities of local emissions can affect the vertical profiles substantially. Figure 4b shows similar vertical profiles of *T* and RH, yet the vertical differences of  $b_{ext}$  can vary substantially among different vertical
- 20 profiles. For example,  $b_{\text{ext}}$  in V29 decreased from 870 Mm<sup>-1</sup> at ground level to 720 Mm<sup>-1</sup> at 260 m, which is a change by ~17%. In comparison,  $b_{\text{ext}}$  in V34 showed a more dramatic decrease by ~34% from 1776 to 1170 Mm<sup>-1</sup>. Although V29 and V34 were measured at relatively similar time (11:00 and 13:00, respectively), the MLH was substantially different, which was 1106 and 546 m, respectively. The higher MLH associated with higher TKE and stronger vertical convection led to a much smaller vertical gradient for V29 than V34. The vertical profiles of NO<sub>2</sub> were similar to those of  $b_{\text{ext}}$  for most of the time. However, vertical variations in
- 25  $b_{\text{ext}}/\text{NO}_2$  ratios were also observed (Figure 4b). For example, the  $b_{\text{ext}}/\text{NO}_2$  ratio in V34 showed a slight decrease from 21 Mm<sup>-1</sup> ppb<sup>-1</sup> at ground level to ~20 Mm<sup>-1</sup> ppb<sup>-1</sup> at ~150 m, followed by a sudden decrease to 18 Mm<sup>-1</sup> ppb<sup>-1</sup>. Such a transition height in vertically-resolved  $b_{\text{ext}}/\text{NO}_2$  was consistent with that of  $b_{\text{ext}}$  and wind direction (Figure 4b), indicating a change of air masses with different pollution characteristics at approximately 150 m.

(3) Higher values at higher altitudes (16% of the time). Figure 4c shows three typical vertical profiles of  $b_{\text{ext}}$  that increase

30 substantially at approximately 120–150 m. A common meteorological feature for such vertical profiles is the *T* inversion (Figure 4c) and a change of wind direction at the same height. The heights of VPT with significant changes were generally similar to those of *T* (Figure S6), suggesting a stratification of the lower boundary layer at ~150 m. For example, wind direction showed a clear change from the northeast to south between 150–200 m and a further change to the southwest above 200 m during V33 (Figure 4c). The

change of WD was associated with an increase in RH, further indicating different air masses below and above the *T* inversion. As a result, the increase in  $b_{ext}$  above ~150 m was mainly caused by regional transport from the south-southwest. While the vertical profiles of NO<sub>2</sub> were all similar to those of  $b_{ext}$ , similar vertical transition points with the changes in  $b_{ext}/NO_2$  ratios were also observed (Figure 4c). As shown in Figure 5c, aerosol composition was very similar between ground level and 260 m before V33,

- 5 however, it showed increases in the contributions of organics and chloride at 260 m while it remained at small changes at ground level. These results further indicate a change of air masses at high altitude during V33. Compared with V33, V12 also showed an increase in  $b_{ext}$  between 120–200 m, which appears to be in contradiction to the much lower NR-PM<sub>1</sub> concentration at 260 m (177 µg m<sup>-3</sup>) than ground level (294 µg m<sup>-3</sup>) (Figure 1e). The vertical profile of RH showed consistently high values (~ 70–80%) below 200 m and then a large decrease above 200 m, leading to the presence of a stable layer below 200 m (Figure 1). TKE was ubiquitously less
- 10 than 0.3 m<sup>2</sup> s<sup>-2</sup>, indicating a stagnant atmosphere during this period. Considering the differences between ground level and 260 m, and the vertical changes in meteorological variables, we expect a rapid decrease of  $b_{ext}$  between 200–260 m although the data was not available. The increase of  $b_{ext}$  between 120 and 200 m was likely due to a decrease of WS from 2 to 1 m s<sup>-1</sup>.

(4) Significant decreases at the heights of  $\sim 100 - 150$  m (14% of time). Such vertical profiles were observed dominantly in November (9 of 10, Table S5) which were typically associated with rapid formation or cleaning of PM pollution and *T* inversions (Figures 1 and

- 15 4d). For example, V11 was conducted during the rapid formation stage of the pollution (20:25–20:46, Figure 1e), which was associated with the formation of the high RH/LWC event (Figures 1 and S4). RH and H<sub>2</sub>O were consistently higher at lower heights, while they showed sudden decreases at 100 m from ~70% to ~50% (300 m), and 5.5 to 4 g m<sup>-3</sup> (Figure S5b), respectively. *T* also showed a strong inversion between 100–200 m with the difference being approximately 1.3°C (Figure 4d). Figure 4d showed that the vertical changes in RH and *T*, and the formation of the higher RH event was likely due to the interaction of two different air masses
- 20 that changed from the northeast to northwest at 100 m. Such vertical profiles clearly indicate very different layers at different heights. This is further supported by the vertical profiles of VPT which showed significant increases at similar heights (Figure S6) and suggested a low stable stratification. As a result, the vertical profile of  $b_{ext}$  showed a sudden decrease at 100 m, from ~1800 to ~1200 Mm<sup>-1</sup> at 200 m. Consistently, the mass concentration of NR-PM<sub>1</sub> showed a significant increase by ~50 µg m<sup>-3</sup> at ground level, while it remained relatively constant at 260 m. The vertical profile of NO<sub>2</sub> also showed a similar decrease at the same transition height
- 25 (from 94 to 83 ppb) as  $b_{ext}$ , yet the change was much smaller than  $b_{ext}$ , leading to a similar decrease in vertical profile of  $b_{ext}/NO_2$ . These results indicate that the local accumulation effects caused by the high RH event are more significant on particles than gaseous species. The vertical changes also have a significant impact on aerosol composition. For example, the contribution of OA to NR-PM<sub>1</sub> was higher at ground level (48–51%) than 260 m (44%), mainly due to the accumulation of local pollutants, e.g., cooking aerosols during the high RH event (Figure 6). Compared with V11, V14 took place during the clearing stage of the pollution (Figure 1e). A
- 30 stable layer associated with a clear *T* inversion in the morning was observed at approximately 150 m. RH was significantly different below and above the layer. While RH was consistently high at ~70% below 150 m (Figure 4d), it showed a rapid decrease from 70% to 40% between 150–300 m, suggesting dryer air masses in higher altitudes. Such vertical profiles in RH and *T* led to a significant transition in  $b_{ext}$  at 150 m. While  $b_{ext}$  was relatively constant (~1350 Mm<sup>-1</sup>) below 150 m, it had a sudden decrease at 150 m and

reached the lowest value at 260 m (~770 Mm<sup>-1</sup>), which was nearly twice lower than that at low altitudes. Similarly, NO<sub>2</sub> showed a sudden decrease from 78 to 67 ppb at the same height, yet the change was much smaller than  $b_{ext}$ , leading to a significantly different  $b_{ext}/NO_2$  ratio below and above 150 m (~18 and ~13–14 Mm<sup>-1</sup> ppb<sup>-1</sup>, respectively, Figure 4d). The difference is consistent with that of NR-PM<sub>1</sub> between ground level and 260 m, which was 175 and 87 µg m<sup>-3</sup>, respectively. Owing to the different air masses and

5 meteorological conditions, aerosol composition was substantially different between ground level and 260 m, particularly when higher contributions of sulfate and lower contributions of organics were observed at ground level (Figure 5d).

#### 3.3 Case study of evolution of vertical profiles

Figure 7 shows a detailed evolution of vertical profiles of b<sub>ext</sub> and meteorological conditions from 00:00 on November 20 to 17:00 on November 21. The evolution can be classified into six stages, and the types of vertical profiles evolved routinely from Type 4 at nighttime to Type 2 / Type 1 during daytime, and then Type 3/Type 4 at nighttime again. The first stage (S1, 00:00–11:00) was characterized by consistent northerly winds and moderately high RH (60–80%). As a result, aerosol species and chemical composition were relatively stable at both ground level and 260 m during this stage. The vertical profile of b<sub>ext</sub> showed a clear transition height at approximately 120 m at 3:00 am, indicating the presence of a nocturnal stable boundary layer (Stull, 1988). The transition height started to increase with the increase of temperature and reached approximately 200 m at 11:00. Consequently, the
vertical profiles evolved from Type 4 to Type 2. The variation of the stable layer height was consistent with that of MLH which

- increased from 190 to 350 m. But note that the MLH was ubiquitously higher by  $\sim$ 50 m than the transition height that was determined from  $b_{\text{ext}}$ , which might indicate that the MLH retrieved from the ceilometer measurement was overestimated during severe haze episodes.
- The second stage (*S*2) was characterized by a transition of wind direction from north to south at 12:00, which remained unchanged 20 until 20:00. Due to the increasing *T* and vertical convection, the vertical differences between ground level and 260 m (e.g., NR-PM<sub>1</sub> in Figure 7e) were gradually reduced, and the vertical profiles were relatively even and evolved from Type 2 to Type 1, e.g., V9 between 14:00–15:00. This is also consistent with the relatively high MLH, which was generally above 350 m. These results indicate that aerosol particles were relatively well mixed during this stage. Increases in secondary nitrate and SOA were also observed at both ground level and 260 m, which is likely due to the enhanced photochemical production and/or regional transport from the south. The concentration of NR-PM<sub>1</sub> at 260 m showed a significant decrease from ~150 to ~80 µg m<sup>-3</sup> during stage 3 (*S*3, 16:00–20:00), while it remained relatively constant at ground level. Such changes were likely caused by a significant increase in wind speed between 200– 300 m that diluted the pollutants substantially (Figure 7b). This is further supported by significant increases in TKE (*E<sub>u</sub>*, *E<sub>v</sub>*, *E<sub>w</sub>*) above 200 m (Figure S5). Indeed, we observed a significant decrease in NO<sub>2</sub> level from ~70 ppb at 200 m to 30 ppb at 260 m, and
- 30 ground level and 260 m was observed during the later stage of S3 (18:00–20:00), which was primarily caused by significant increase in primary OA (mainly local cooking OA) from ~25 to 130  $\mu$ g m<sup>-3</sup> at the ground site (Figure 6b).

also a large decrease in BC from 22 to 14  $\mu$ g m<sup>-3</sup> for V10 ( $b_{ext}$  was not available). A further increase in vertical difference between

After *S*3, *T* showed a significant decrease while RH showed a rapid increase from 55% to ~80% below 100 m, and LWC at ground surface reached >50 mg m<sup>-3</sup>. The high RH event was mainly caused by the interaction between the warmer southwestern air mass and the colder northeastern one (Figure 7b). The height of the high RH (> 70%) layer was consistently below 200 m during *S*4 (20:00–22:30). Therefore, aerosol species showed much smaller variations at 260 m compared with the large increases at ground level

- 5 (Figure 6b). Not surprising,  $b_{ext}$  showed a large vertical gradient (Type 4), for example,  $b_{ext}$  for V11 decreased from ~2000 Mm<sup>-1</sup> at ground level to 1200 Mm<sup>-1</sup> above 100 m. Such a vertical profile was also consistent with a dramatic decrease in MLH from 220 to 135 m. Note that the most significant increases in organics, sulfate, and chloride during S4 likely indicates the rapid accumulation and/or transformation of coal combustion emissions (Wang et al., 2015). The high RH layer height gradually evolved to ~240 m during stage 5 (S5, 22:30–9:00) according to the vertical variations in RH, and correspondingly, the MLH increased from 130 to 260
- 10 m. Figures 7b and 7c show consistently low WS and high RH at the early stage of S5. Although NR-PM<sub>1</sub> at ground level showed a gradual decrease, the contributions of secondary species to NR-PM<sub>1</sub> increased. These results likely indicate an enhanced formation of secondary aerosol species due to aqueous phase processing while the scavenging compensated the increases in the total mass. As a comparison, we didn't observe similar increases in secondary aerosol species at 260 m because of low RH (<70%). The vertical profile of  $b_{\text{ext}}$  showed a transition point at the height of ~100–130 m, overall consistent with the vertical variations in meteorology.
- 15 After 1:00, an increase in WS above 100 m was observed, leading to a faster decrease of NR-PM<sub>1</sub> at 260 m than ground level. The height of the high RH layer was also decreased from ~240 to 150 m. Such changes in meteorological conditions again led to a strong vertical gradient of  $b_{ext}$ . For example,  $b_{ext}$  for V14 was ~1400 Mm<sup>-1</sup> at ground level, while it was nearly a factor of two lower at 260 m (~750 Mm<sup>-1</sup>), consistent with the difference in NR-PM<sub>1</sub> concentration (170 vs. 90 µg m<sup>-3</sup>). After 9:00, RH quickly dropped to <20% at both ground level and 260 m. Because of the increasing *T*, MLH rapidly increased from <200 to ~550 m (13:00). Consistently, the vertical differences between ground level and 260 m were largely reduced, and the vertical profiles of  $b_{ext}$  showed relatively even
- distributions (Type 1). These results indicate that aerosol particles during *S*6 became relatively well mixed.

#### 3.4 Vertical co-variations of pollutants

Figure 8 shows the correlations of BC and NO<sub>2</sub> with  $b_{ext}$  for all vertical profiles during the two haze episodes. The vertical profiles of BC were similar to  $b_{ext}$  for most of the time ( $R^2 = 0.92$  and 0.69 in November and January, respectively). By performing a linear

25 regression of BC to three POA factors (FOA, COA, BBOA) and secondary aerosol species (SA = SOA + sulfate + nitrate + ammonium) (Sun et al., 2016), we found that ~30% of BC was associated with SA during the severe haze episode in November, while it was only ~11% during the one in January. This explains the better correlation between  $b_{ext}$  and BC in November than in January. The average single-scattering albedo (SSA) calculated with a constant mass absorption efficiency for BC (7.3 m<sup>2</sup> g<sup>-1</sup>) (Han et al. 2015) using Equation (1) was 0.84 and 0.86 in November and January, respectively.

$$30 \quad SSA = (b_{ext} - BC \times 7.3) \div b_{ext} \tag{1}$$

The SSA values were generally consistent with those observed in previous studies (Li et al., 2015c;Lee et al., 2007;Han et al., 2015;Han et al., 2017), e.g.,  $0.85 - (\pm 0.9106)$  during the fall of 2014 (Han et al., 2015) and  $0.85 (\pm 0.04)$  in the winter of 2015 (Han

et al., 2017). We also observed slightly lower values at lower altitudes, indicating an enhanced BC absorption due to the influences of local emissions (Sun et al., 2015;Sun et al., 2010). However, for several specific events, we observed very significantly different correlations between BC and  $b_{ext}$ . For example, V11 showed strong vertical gradients for both BC and  $b_{ext}$  with sudden decreases at approximately 100 m (Figure S8). As indicated in Figure 4d, such changes were mainly associated with corresponding changes in

- 5 meteorological parameters, which were characterized by a *T* inversion, a decrease in RH, and a wind direction change from northeast to northwest. The vertical profile of SSA first showed a gradual decrease from 0.82 at ground level to 0.78 at 100 m, followed by a large increase to 0.86 at 200 m (Figure S8). As shown in Figure 9, aerosol composition was substantially different below (ground level) and above (260 m) the transition height. While organics showed higher contribution at ground site than at 260 m (48 51% vs. 44%), the nitrate contribution was relatively lower (18–20% vs. 23 24%). Also, OA at the ground site was dominated by primary
- 10

OA from cooking and biomass burning emissions (70 - 74%), while it was mainly composed of secondary OA (LO-OOA + MO-OOA, ~66%) at 260 m. These results indicate that aerosol particles below the transition height were largely influenced by local source emissions, while those above the height were more affected by secondary aerosols.

NO<sub>2</sub> showed overall similar vertical profiles as  $b_{ext}$  (Figure S9) during the severe haze episode in November as indicated by the tight correlations in Figure 10 ( $R^2 > 0.75$  for 63% of the time). The slopes of  $b_{ext}$  versus NO<sub>2</sub> were relatively stable before November 21 by varying from 16 Mm<sup>-1</sup> ppb<sup>-1</sup> to 20 Mm<sup>-1</sup> ppb<sup>-1</sup>. However, we observed large decreases in  $b_{ext}/NO_2$  from 26 Mm<sup>-1</sup> ppb<sup>-1</sup> to 2-3 Mm<sup>-1</sup> ppb<sup>-1</sup> on November 21. This is consistent with the fact that the NO<sub>2</sub> concentration slowly decreased from ~ 60 to 20 ppb, while  $b_{ext}$ rapidly decreased from ~2000 to 60 Mm<sup>-1</sup>. Compared with November, the vertical profiles of NO<sub>2</sub> showed more differences from those of  $b_{ext}$  during the haze episode in January. The similar vertical profiles with  $R^2 > 0.75$  accounted for 32% of the total profiles (Figure 10). The average  $b_{ext}/NO_2$  ratios also varied more significantly (~10 – 20 Mm<sup>-1</sup> ppb<sup>-1</sup> for most of the time) than those in

- 20 November. Also note that the  $b_{ext}/NO_2$  ratios in January appeared to show a clear diurnal pattern with higher values at nighttime, which can be explained by the different primary and secondary sources between daytime and nighttime. The types of vertical profiles of  $b_{ext}/NO_2$  are relatively similar to those of  $b_{ext}$ . As shown in Figure 4, the vertical profiles of  $b_{ext}/NO_2$  showed either uniform distributions (Type 1), decreases/or increases as a function of height (Types 2 and 3), or significant decreases at ~100 m (V11) or 180-200 m (V14) (Type 4).
- Figures 8 c and d also show that the correlations between  $b_{ext}$  and NO<sub>2</sub> were not linear, and the ratio of  $b_{ext}/NO_2$  appeared to increase as a function of pollution level. For example, the ratio of  $b_{ext}/NO_2$  decreased from ~26 to ~17 Mm<sup>-1</sup> ppb<sup>-1</sup> during the high RH period of 00:00–7:00 on November 21 as the mass concentration of NR-PM<sub>1</sub> decreased from ~300 to 170 µg m<sup>-3</sup> (Dan et al., 1999;Dan et al., 2003;Mcmurry et al., 2004;Wayne et al., 1991). One explanation is the enhanced formation of secondary aerosol species during severe haze episodes that were associated with the oxidation of precursors (e.g., SO<sub>2</sub> and NO<sub>2</sub>). This is particularly important for SO<sub>2</sub>
- 30 that can be rapidly oxidized to form sulfate via aqueous-phase and/or fog processing (Sun et al., 2013c;Quan et al., 2015), while the role of  $NO_2$  oxidation is generally small due to its much slower aqueous-phase processing rates (Seinfeld and Pandis, 2006), which is also supported by the shorter lifetime of  $SO_2$  compared with that of  $NO_2$  in the atmosphere. In addition, the scavenging rates of gaseous species and particles can be substantially different during the high RH event (Wayne et al., 1991;Dan et al., 1999;Dan et al.,

2003;Mcmurry et al., 2004). While most aerosol species can be efficiently scavenged (e.g., > 50% during fog events with high RH) (Gilardoni et al., 2014), the scavenging of NO<sub>2</sub> is much slower due to its low solubility. This could be another important factor affecting the changes in  $b_{ext}$ /NO<sub>2</sub> ratios.

#### 4 Conclusions and implications

- 5 The vertical profiles of particle extinction, gaseous NO<sub>2</sub>, and BC from ground level to 260 m measured during two severe winter haze episodes in the megacity of Beijing varied very dynamically and interacted closely with boundary layer and meteorological conditions. Four types of vertical profiles (96% of the time) were sorted out to elucidate the vertical evolution characteristics of air pollutants during the two severe haze episodes. The diurnal evolution of vertical profiles of  $b_{ext}$  highlighted the presence of a nocturnal stable boundary layer between 100–150 m during severe haze episodes and its impact on the vertical characteristics of air
- 10 pollution. However, the transition heights determined from the vertical changes in  $b_{ext}$  through visual perspective were consistently lower than the MLH retrieved from the ceilometer measurements. As shown in FigureFig. 11, the differences between the two methods can be as large as 100 - 150 m during the periods with the highest  $b_{ext}$  (vertical profiles of 11 - 13) and  $\sim 40 - 100$  m during other periods with nocturnal stable boundary layer. These results indicate that the ceilometer measurements might often overestimate the MLH, particularly at nighttime during severe haze episodes, and the vertically resolved measurements on the BMT can be an
- 15 essential supplement, particularly for the lidar measurements with a blind zone below ~200-300 m. Because of the relatively low number of cases and the limited height of the meteorological tower, future vertical measurements of particle extinction to a high altitude, e.g., using tethered balloon are needed to further validate the retrieval of MLH from CL51 measurements. Traditional air quality forecast models often underestimate severe haze episodes substantially. In addition to the incomplete understanding of the formation mechanisms, the very complex and dynamic vertical variations could also be one of the major reasons. Our results also
- 20 highlight that more comprehensive vertical measurements (e.g., more aerosol and gaseous species) at a higher altitude in the megacities are urgently needed for a better understanding of the formation mechanisms and evolution of severe haze episodes.

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#### **Figure Captions:**

Figure 1: Vertical evolution of (a, f) wind direction, (b, g) wind speed, (c, h) relative humidity, and (d, i) temperature during two severe haze episodes in November and January. (e) and (j) show the time series of NR-PM<sub>1</sub> mass concentrations at ground level and 260 m. The two pie charts in each panel present the average composition of NR-PM<sub>1</sub> for each episode at ground level (right) and 260 m (left), respectively, and the numbers on the top of pie charts are the average mass concentrations. The number of vertical profile experiments is also marked in (e, j) as green (up) and gray (down) vertical lines, and the time for each vertical profile is detailed in Table S1.

Figure 2: Time series of (a) vertical difference (= max - min) and slope of extinction versus height; (b) extinction and vertical difference (=  $(max-min)/(2^*avg)$ ) for each vertical profile during the haze episode in November. The circles refer to the down profiles while boxes indicate the up profiles.

Figure 3: Time series of (a) vertical difference (= max – min) and slope of extinction versus height; (b) extinction and vertical difference (= (max-min)/(2\*avg)) for each vertical profile during the haze episode in January. The circles refer to the down profiles while boxes indicate the up profiles.

Figure 4: Four selected types of vertical profiles of  $b_{ext}$ ,  $b_{ext}/NO_2$ , *T*, RH, and WS. "U" in the legend indicates the "up" experiments, the others are the "down" ones. The vertical profile 9 (V9) for  $b_{ext}/NO_2$  is missing because NO<sub>2</sub> data was not available. The rightmost panels show four selected vertical profiles of WD, i.e., V3, V25, V33, and V11. For clarity, some vertical profiles were offset by certain values as indicated by "+" or "-" in the figure. Note that all vertical profiles reaching 260 m were measured at daytime, while those of 200 m were measured at nighttime (Table S1).

Figure 5: Average chemical composition of NR-PM<sub>1</sub> for selected vertical profiles (a) V24, (b) V29, (c) V33, and (d) V14 at ground level (bottom panel) and 260 m (top panel-). The numbers on bar charts are the mass fractions of sulfate and organics. "U" and "D" represent "up" and "down" experiment, respectively.

Figure 6: Evolution of non-refractory submicron aerosol species and NR-PM<sub>1</sub> composition at (a) 260 m and (b) ground level. POA (=
 FOAFFOA+COA+BBOA) and SOA (= LO-OOA + MO-OOA) were from positive matrix factorization of OA. RH and WD are also shown for a comparison.

Figure 7: Evolution of vertical profiles of  $b_{ext}$  and meteorological parameters from 00:00 on November 20 to 17:00 on November 21. Also shown are time series of NR-PM<sub>1</sub> mass concentrations at ground level and 260 m, and MLH<sub>7</sub> (right axis). Note that the vertical profiles of  $b_{ext}$  between 16:00–17:30 were not available due to a malfunction of the CAPS PM<sub>ext</sub>, therefore, the two vertical profiles of NO<sub>2</sub> were used as surrogates. The shaded areas in (e) indicate the time periods for the vertical measurements.

Figure 8: Correlations between BC and  $b_{ext}$  NO<sub>2</sub> and  $b_{ext}$  for all vertical profiles in (a, b) November and (c, d) January. The data points are color coded by the heights. The solid circles in (a) and (b) are data points for V11.

Figure 9: Average chemical composition of NR-PM<sub>1</sub> and OA at ground level and 260 m for vertical profile 11 (V11). Right panels are the mass fractions of NR-PM<sub>1</sub> and OA species. "U" and "D" represent "up" and "down" experiment, respectively.

Figure 10. Time series of correlation coefficients ( $R^2$ ) and slopes between  $b_{ext}$  and NO<sub>2</sub> for each vertical profile in (a) November and (b) January. The circles refer to the down profiles while other are the up profiles.

Figure 11: A comparison of mixing layer height (MLH) with that estimated from the transition heights of  $b_{ext}$ . The numbers of vertical profiles as marked in Fig. 1 and the differences of the two heights are also shown in the plot. Note that only vertical profiles with significant changes were used for estimation of transition heights.

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Figure 1: Vertical evolution of (a, f) wind direction, (b, g) wind speed, (c, h) relative humidity, and (d, i) temperature during two severe haze episodes in November and January. (e) and (j) show the time series of NR-PM<sub>1</sub> mass concentrations at ground level and 260 m. The two pie charts in each panel present the average composition of NR-PM<sub>1</sub> for each episode at ground level (right) and 260 m (left), respectively, and the numbers on the top of pie charts are the average mass concentrations. The number of vertical profile experiments is also marked in (e, j) as green (up) and gray (down) vertical lines, and the time for each vertical profile is detailed in Table S1.



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