1	New insights into the vertical structure of the September 2015
2	dust storm employing 8 ceilometers over Israel
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37 Abstract

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39 On 7 September 2015 an unprecedented and unexceptional extreme dust storm struck the Eastern Mediterranean (EM) basin. Here, we provide an overview of the previous studies and describe the dust 40 plume evolution over a relatively small area, i.e., Israel. This study employs multiple tools including an 41 array of eight ceilometers, spectral radiometers (AERONET), ground particulate matter concentrations, 42 43 satellite images, global/diffuse/direct solar radiation measurements and radiosonde profiles. Main findings reveal that the dust plume penetrated Israel on the 7 September from the northeast in a downward 44 motion to southwest. On 8 September, the lower level of the dust plume reached 200 m above ground 45 level, generating aerosol optical depth of AOD>3, and extreme particulate matter concentration measured 46 on ground level up to ~10,000 μ m m⁻³. A most interesting feature on 8 September was the very high 47 variability in the surface solar radiation in the range of 200-600 W m⁻² (22 sites) over just a distance of 48 several hundred km in spite of the thick dust layer above. Furthermore, 8 September shows the lowest 49 radiation levels for this event. On the following day, 9 September, the surface solar radiation increased, 50 thus enabling a late (~ 11 UTC) sea breeze development mainly in the coastal zone along with 5 m s⁻¹ 51 surface winds associated with arc-shaped dust layers. On 10 September the AOD values started to drop 52 to ~ 1.5 , the surface concentrations of particulate matter decreased as well as the ceilometers aerosol 53 indications; Still, as indicated by CALIPSO a 2-4 km dust layer remained. 54

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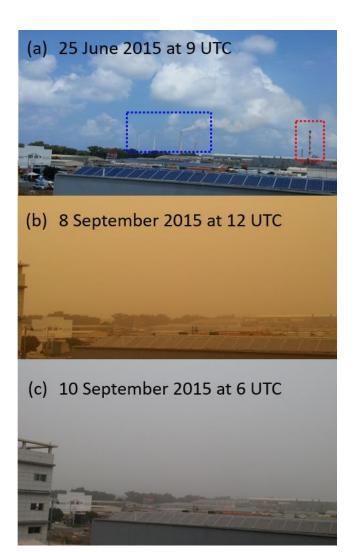
57 **1. Introduction**

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An exceptionally extreme dust storm prevailed over the Eastern Mediterranean (EM) on September 59 2015. The Israeli meteorological service (IMS) declared the dust storm to be extraordinary as it occurred 60 on early September (7-10 September), extended over a time span of 100 hours creating extreme ground 61 level particulate matter (PM) concentrations (e.g. 100 times above the hourly average of PM10 in 62 Jerusalem). On 7 September, prior to the penetration of the dust storm over Israel, IMS reported 63 (http://www.ims.gov.il/IMS/CLIMATE; in Hebrew) a heat wave which prevailed over Israel causing 64 harsh weather conditions of 80-90% relative humidity, 42 °C in valleys, 38 °C in mountains. On 8 65 September, visibility decreased below 3 km and consequently, inland aviation was prohibited until the 9 66 September (Fig.1). Concurrently, severe ground level PM concentrations resulted a public warning from 67 68 outdoor activities, issued by the environmental protection ministry. Finally, on 11 September, as visibility

increased, the IMS confirmed the dust storm ended, whereas the heat wave was over only two days later, on 13 September, subsequent to a profound change in weather conditions. The PM concentrations declined to values measured prior to the dust storm (<u>http://www.svivaaqm.net/Default.rtl.aspx;</u> in Hebrew) only on 14 September, though the AERONET measurements (https://aeronet.gsfc.nasa.gov) stationed in central and southern Israel reveal that the aerosol optical depth (AOD) resumed to values prior to the dust storm only on 17 September.

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Figure 1. Photographs taken from the central coast of Israel, adjacent to the Hadera ceilometer, 3.5 km southeast to the stacks of a power plant (indicated by a blue rectangle) and 600 m north to a factory stack (indicated by a red rectangle). The photographs were taken prior to the dust storm, on 25 July 2015 (a) and during the dust storm, on 8 September 2015 (b) and 10 September 2015 (c). Notice that during the dust event (b,c) the stacks clearly seen in (a) from the same spot, are invisible.

Investigation of the mechanisms leading the severe dust storm was performed by Gasch et al. (2017) 85 using a state of the art dust transport model ICOsahedral Nonhydrostatic (ICON) with the Aerosol and 86 Reactive Trace gases (ART) (Rieger, et al., 2015). The model concentrated on the EM with one global 87 domain (40 km grid spacing, and 90 vertical levels from 20 m to 75 km) and 4 nested grids (20, 10, 5 88 and 2.5 km grid spacing and 60 vertical levels from 20 m to 22.5 km). Simulations were done for three 89 consecutive days from 6-8 September. Model results delineated an unusual early incidence of an active 90 Red Sea Trough (Fig.2; Alpert et al, 2004) over Mesopotamia, followed by meso-scale convective 91 systems over the Syrian-Iraqi border generating three cold-pool outflows. On the night between 5 and 6 92 September, a convective system fueled by an inflow along the eastern side of the Red Sea Trough, moved 93 northeast over the Turkish-Syrian border region. The convective system intensified overnight and 94 95 generated a first weak cold pool outflow on the 6 September. After sun rise, as the nocturnal boundary layer dissipated, an increase of downward mixing lead to an increase of surface wind speeds consequently 96 97 causing dust to pick up over Syria. The high surface wind speeds sustained during the day due to a strong and shallow heat low whereas sea breeze transported dust to the south towards Jordan. At this point, the 98 99 atmospheric instability over the Syrian-Iraqi border created a second convective cool outflow from the Zagros mountain range west into Syria. The gust from the second cool pool outflow ignited a third cool 100 pool outflow at 20 UTC. The third outflow moved southerly along the eastern flank of the Red Sea 101 Trough and shifted warm and moist air masses along the way. Past midnight, on 7 September, the 102 intensified third cool pool outflow shifted to the west. At 10 UTC 7 September, rainfall and an increase 103 104 of surface wind speeds north-west of Syria strengthened the third cool pool outflow. Consequently, enormous dust emissions transported southwest up to 5 km. The aged second cool pool outbreak re-105 intensified over Jordan and southwestern Syria and merged with the third outflow with the nightfall of 7 106 September. After midnight, on 8 September, the dust transported over Israel and was mostly influenced 107 108 by local circulation systems in the EM. Model simulations were compared to in-situ measurements and 109 satellite images: visible electromagnetic spectrum from Moderate Resolution Imaging Spectroradiometer (MODIS: https://modis.gsfc.nasa.gov/) aboard the Aqua satellite; AOD from the Terra satellite; RGB 110 dust product from the Spinning Enhanced Visible and InfraRed Imager (SEVIRI) upon the Meteosat 111 112 Second Generation (MSG) satellite; total attenuated backscatter from Cloud-Aerosol Lidar upon the Infrared Pathfinder Satellite Observations (CALIPSO: https://www-calipso.larc.nasa.gov/). Ground level 113 114 meteorological measurements (3 sites) and PM concentration (3 sites) in Israel were employed. Results revealed the model lacked sufficient development of a super critical flow, which in effect produced the 115 116 excessive surface wind speeds. Eventually, this misled the forecast of the dust advection southwest into 117 Israel.

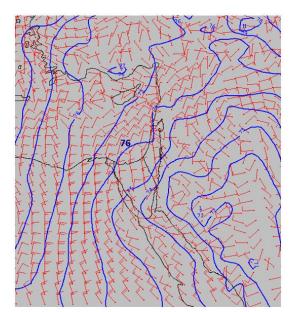


Figure 2. 925 mb map from 7 September 2015 12 UTC of the geopotential height with 1 dm interval (the 76 dm line passing over Israel; blue lines) and wind (10 KT each line; red arrows). Source: IMS 122 from UKMO British Met Office model. 123

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The fact that forecast models did not succeed in predicting this outstanding dust event motivated 126 Mamouri et al. (2016) to study its origin and development. Their research presented dust load 127 observations in the Cyprus region. Luckily, at the time of the dust storm, an EARLINET (European 128 Aerosol Research lidar Network: https://www.earlinet.org/) Raman lidar stationed in Limassol provided 129 vertical dust profiles and valuable optical dust properties of backscatter, extinction, lidar ratio and linear 130 depolarization ratio. They analyzed the optical thickness (AOT) and Angström exponent derived from 131 the MODIS Aqua satellite. MODIS Aqua AOT measurements were compared to the Limassol lidar 132 133 observations, AOD measurements from two AERONET sites (Cyprus and Israel) and ground level PM10 concentration from four Cyprus sites. On 7 September, EARLINET lidar observations measured two dust 134 layers (extending up to 1.7 km ASL and between 1.7-3.5 km ASL). The dust particle extinction 135 coefficient measured in Limassol had reached 1000 Mm⁻¹ followed by high PM10 concentration of 2000 136 μm m⁻³. Extreme values over Limassol, were reported on the 8 September as MODIS Aqua AOT 137 observations exceeded 5 (assuming overestimation up to 1.5) and hourly PM10 concentration of about 138 $8,000 \,\mu\text{m}\,\text{m}^{-3}$ (with uncertainties in the order of 50%). Unfortunately, on the 8 September, the lidar was 139 intentionally shut down to avoid potential damage to the instrument. Nevertheless, lidar observations 140 141 indicated another dense dust outbreak (1-3 km ASL) reaching Limassol on the 10 September, also visible by the MODIS Aqua AOT imagery. The researchers concluded the scale of the dust storm features was 142

too small for global and regional dust transport models. They presumed that the initiation of the dust plume was due to an intense dust storm (Habbob) in northeastern Syria and northern Iraq, leading to vigorous downbursts which consequently pushed huge amounts of dust and sand to the atmosphere. The lidar observations indicated a double layer structure of the dust, 1.5 and 4 km above sea level (ASL), pointing to multiple dust sources.

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Stavros et al., (2016) continued the investigation of the formation and mechanism of the dust storm 149 over Cyprus by a high regional atmospheric model of the integrated community limited area modeling 150 system (RAMS-ICLAMS). The model simulations focused on the generation of the dust storm on 6 and 151 7 September. Model results were fine-tuned by observations from EARLINET lidar stationed in 152 153 Limassol, radiosonde data from five sites (Cyprus, Israel, Jordan, and two from Turkey) and satellite imagery from MSG SEVIRI and CALIPSO CALIOP. The model was set to three grid space domains: 154 155 an external grid of 12X12 km, (over the EM) an inner set at 4X4 km (over northern EM) and 2X2 km grid for cloud resolving (over northeastern Syria). The vertical structure consisted of 50 terrain following 156 157 levels up to 18 km. The researchers assessed a strong thermal low over Syria followed by convection activity over the Iraq-Iran-Syria-Turkey borderline combined with land use changes (aftermath of the 158 war held in Syria) manufactured the extreme dust storm. The model succeeded to describe the dust 159 westward flow of a haboob containing the dust previously elevated over Syria as observed by MSG 160 SEVIRI and EARLINET lidar. However, there were some inaccuracies in the quantification of dust mass 161 profiles. They attributed the model discrepancies to the limited ability of the model to properly resolve 162 dust and atmospheric properties (e.g. change of land use and intense downward mixing). 163

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165 Evaluation whether the dust activation due to human perturbations to land use (such as in the Syria civil war) had an underlying effect on the dust storm formation or even to its increase were studied 166 by Pu and Ginoux, (2016). They examined the connection between the Pacific decadal oscillation and 167 the dust optical depth (DOD) in Syria. DOD, derived by the deep blue algorithm (Hsu et al., 2013) aerosol 168 product from MODIS Terra and MODIS Aqua satellite (10 km resolution picture) was combined with 169 170 monthly horizontal winds and geopotential heights generated by the European Centre for Medium-Range Weather Forecasts (ECMWF) reanalysis (horizontal resolution of 80 km and 37 vertical levels). The 171 172 dataset of DODs during the years 2003-2015 were compared to the Geophysical Fluid Dynamics Laboratory (GFDL) Atmospheric model (AM3) (Donner et al., 2011). The AM3 model produced AODs 173 174 and calculated the mass distribution and optical properties of aerosols, their chemical production, transport, and dry or wet deposition. Comparison of AOD model results, AOD AERONET measurements 175 176 and DOD from satellite observations revealed the model underestimated the AOD's particularly in the

EM. The authors assumed that the soil moisture parameter in the model were not set properly resulting in the AOD dissimilarities.

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The impact of the conflict in Syria on the aridity of the region and therefore, a possible direct 180 impact on the generation of the September dust storm was examined by Parolari et al., (2016). The 181 researchers conducted simulations using the Advanced Research Weather Research and Forecasting 182 (WRF-ARW) model from 30 August to 10 September over the EM. The model consisted of two nested 183 domains (9 and 3 km grid spacing and 35 vertical levels). Daily and monthly AOD data from MODIS 184 were computed by the deep blue algorithm over land. Anomalies of the September 2015 monthly average 185 AOD were compared to the monthly average of 2000-2015. The monthly average of September 2015 186 187 vegetation status in the region was estimated by MODIS normalized difference vegetation index (NDVI). Historical data was divided into two periods: none-drought (2001-2006) and drought (2007-2010). Wind 188 189 shear stress was calculated to estimate wind erosion. Their findings reveal that the enhanced dust uplift 190 and transportation of the September 2015 dust storm was due to meteorological conditions rather than 191 the land-use changes because of the civil conflict in Syria. WRF simulations revealed the well-known Shamal winds and cyclone associated with dust storms in the Middle East (Rao et al., 2003) were 192 193 characterized by northwesterly winds west of the low pressure zone in the Syrian-Iraqi border. However, the source of elevated dust concentrations over the EM coast on the 7and 8 September were attributed 194 cyclone front movement. On 6 September low level winds (700 hPa) were opposite to the northwesterly 195 high level (300 hPa) winds, consequently, generating enhanced surface shear stress and transported re-196 suspended PM westward. Furthermore, based on the past 20 years, the Israeli summer of 2015 was 197 unusually dry and hot and therefore enabled easier updraft of dust soil increasing the probability of dust 198 emissions. 199

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Jasmin (2016) compared the dust storm aerosol content provided by MSG SEVIRI observations, to the results of the open source Meteoinfo model (Wang, 2014). The Meteoinfo model was based on meteorological variables from ECMWF. The model meteorological conditions suggested a formation of two simultaneous dust storms, from northern Syria and from the Egyptian Sinai desert, resulting from updrafts created by low pressure systems.

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The aforementioned studies (summarized in Table 1) focused on the generation of the dust storm in the Syria region based on transport models, satellite imagery and in situ measurements. In our study we focus on the evolution of the dust plume over Israel in the lower atmosphere based on an array of 8 ceilometers, 52 in situ PM measurements, two AERONET sites and satellite imagery. We do not

investigate the reasons of the models frailer to predict this extraordinary storm, but rather attempt to present details of the evolution of the dust plume passage over Israel. The data presented here can be used as a tool to verify state of the art model simulations and provide a different point of view to the meteorological conditions governing the dust plume advection over the EM.

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In the following section we describe the measuring sites and instruments in this study. The list of instruments includes; ceilometers, PM measurements, AERONET, radiosonde, solar radiation and satellite imagery. Sect. 3 presents the results of the dust plume spatial and temporal scheme delineated by the ceilometer plots on 7 September through the 10 September 2015. We discuss and compare between the ceilometer plots and the aforementioned auxiliary measurements. In Sect. 4 we conclude our main findings of the dust plume advection in the lower atmosphere above Israel and the downward transport towards the ground.

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225 **2. Instruments**

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227 **2.1 Ceilometers**

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Ceilometers, initially intended for cloud level height detection, are automatic low cost lidars widespread 229 in airports and weather stations worldwide. As single wavelength lidars, ceilometers cannot produce the 230 information aerosol properties such as size distribution, scattering and absorption coefficients (Ansmann 231 et al., 2011; Papayannis et al., 2008). Nevertheless, with improvement of hardware and firmware over 232 the years, ceilometers have become a valuable tool in the study of the atmospheric boundary layer and 233 the vertical distribution of aerosols layers (Haeffelin and Angelini, 2012; Ansmann et al., 2003). 234 Furthermore, in 2013 ceilometers been assimilated in the EUMETNET (European Meteorological 235 Services network) Profiling Program (E-profile) to develop a homogeneous dataset from automatic lidars 236 237 and state-of-the-art ceilometers across Europe (http://eumetnet.eu/activities/observationsprogramme/current-activities/e-profile/alc-network/). 238

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Vaisala ceilometers type CL31, are commonly deployed worldwide, and the main research tool in this study as well. CL31 is a pulsed elastic micro lidar, employing an Indium Gallium Arsenide (InGaAs) laser diode transmitter of near infrared wavelength of 910 nm ± 10 nm at 25°C. In order to compensate the low pulse energy of the laser (defined "eye-safe") and to provide sufficient signal to noise ratio, the

pulse repetition rate is of 10 kHz every two seconds (Vaisala ceilometer CL31 user's guide: 244 http://www.vaisala.com). The backscatter signals are collected by an avalanche photodiode (APD) 245 receiver and designed into attenuated backscatter profiles within a reporting interval of 2-120 s 246 (determined by the user). The attenuated backscatter profiles are automatically corrected by an internal 247 calibration (resulting in a multiplication factor of 10^{-8} to convert the signal count to attenuated 248 backscatter), a cosmetic shift of the backscatter signal (to better visualize the clouds base), an obstruction 249 correction (when the ceilometers' window is blocked by a local obstacle) and an overlap correction (to 250 the height where the receiver field of view reaches complete overlap with the emitted laser beam). 251

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Kotthaus et al., (2016) examined the Vaisala CL31 ceilometer by comparing the attenuated 253 254 backscatter profiles from 5 units with different specification of senor hardware, firmware and operation 255 settings (noise, height and time reporting interval). They have concluded the instrument characteristics 256 that affect the quality and availability of the attenuated backscatter profiles in the following manner: At high altitudes, a discontinuity in the attenuated backscatter profile is evident at two height points, ~ 4949 257 and 7000 m. Background signals (instrument related) and cosmetic shift (firmware dependent) tend to be 258 either negative or positive up to 6000m and then switch signs above ~ 6000 m. Below 70 m an overlap 259 correction is applied internally by the ceilometer sensor as well as an obstruction correction (below 50 260 m). Between 50-80 m hardware related perturbation cause a slight offset in the attenuated backscatter 261 values. The authors advise the user-defined reporting interval should be no shorter than 30s to avoid 262 consecutive profiles partial overlap. However, they emphasize the internal calibration applied to convert 263 the signal count output to "attenuated backscatter" does not always fully represent the actual lidar 264 constant, therefore, it is not accurate enough for meteorological research. Nevertheless, since ceilometers 265 are not sensitive to molecular scattering and solar radiation contributes to the random noise, a background 266 correction can be derived by a 4 hr round of midnight attenuated backscatter profiles. Furthermore, a 267 range corrected attenuated backscatter can be derived by the attenuated backscatter profiles during an 268 existence of a stratocumulus cloud. 269

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Furthermore, Weigner et al, (2014) studied different retrieval methods to derive the aerosol backscatter coefficient from the ceilometers' attenuated backscatter profiles based on a comparison to auxiliary collocated instruments such as a sunphotmeter or a multiwavelength lidar. They focused on calibration methods, the rage detection limitations by the overlap function and the sensitivity of the attenuated backscatter signal to relative humidity. Although, the ceilometer wavelength range (given as 905 \pm 3 nm) is influenced by water vapour absorption, in the case of aerosol layer detection, water vapour distribution has a small effect on the signal change, indicating the mixed layer height (MLH) or an elevated mixed

layer, as the aerosol backscatter itself remains unchanged. Consequently, except for a case of a dry layer
in a humid MLH, the water vapour is unlikely to lead misinterpretation of the aerosol stratification.
Fortunately, most algorithms are based on a significant signal slope to define the aerosol layers, therefore,
can be determined from uncelebrated ceilometer attenuated backscatter profiles. The wavelet covariance
transform (WCT) was the method used in this study to evaluated the MLH (Uzan, et al, 2016), whether
determined by the creation of thermals or the subsidence of the dust plume.

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In this study, we address the aforementioned limitations as we refer to the ceilometer signal count 285 profiles between 100-1000 m AGL. The ceilometer array is comprised of 8 units in different sites (Fig 3 286 and Tables 2-3), 6 of which are owned by a governmental office. The ceilometers are CL31 type apart 287 for ceilometer CL51 stationed in the Weizmann Institute which has a higher backscatter profile range (up 288 to 15.4 km, Münkel et al., 2011). Unfortunately, calibration procedures were not held and maintenance 289 (cleaning of the ceilometer window) was done regularly only for the Beit Dagan ceilometer. Apart from 290 the Beit Dagan and Weizmann ceilometers (Table 4), we could not retrieve the technical information of 291 292 firmware and hardware type. However, we have been confirmed (personal communication) that the 293 combination of hardware and firmware had been done following Kotthaus et al (2016). The Beit Dagan ceilometer signal count were found to be weaker (up to 800 signal count compared to 10,000 in the other 294 295 CL31 ceilometers) due to different hardware definitions. Therefore, in order to present the Beit Dagan attenuated backscatter profiles aligned with the profiles of the other ceilometers (given in Fig. 17), the 296 Beit Dagan attenuated backscatter values were multiplied by 12.5 (10,000/800). 297

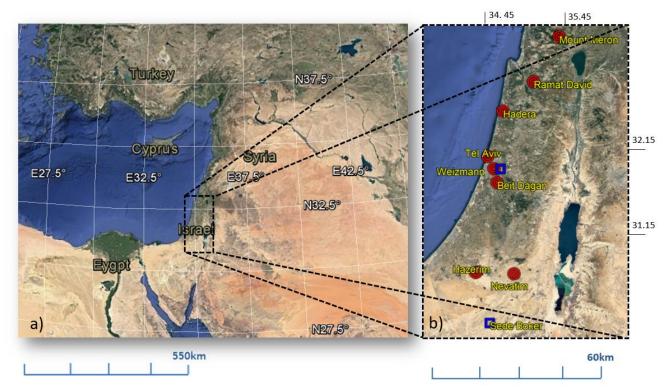


Figure 3. Google Earth map of the large domain (a) and Israel (b) with indications of 8 ceilometer sites (red circle; detail in Table 2) and two AERONET sites (blue square).

2.2 Radiosonde

Radiosonde (RS) type Vaisala RS41-SG is launched by the IMS twice a day at 00 UTC and 12 UTC from the Beit Dagan site adjacent to the Beit Dagan ceilometer. The radiosonde produces profiles of humidity, temperature, pressure, and wind speed and wind direction. The output files were downloaded from the University of Wyoming site (http://weather.uwyo.edu/upperair/sounding.html, station number 40172). With respect to Stull (1988), the MLH was defined by the RS profiles as the height where the following phenomena were identified: an inversion in the temperature, a significant drop in the relative humidity, strong wind shear and an increase in the virtual temperature (Uzan et al., 2016; Levi et al., 2011).

320 **2.3 Particulate matter monitors (PM10, PM2.5)**

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PM monitors are low-volume flow rate Thermo Fisher Scientific type FH 62 C14 (beta attenuation 322 323 method) and type 1405 TEOM (Tapered Element Oscillating Microbalance method). In the beta attenuation method (https://www3.epa.gov/ttnamti1/files/ambient/inorganic/overvw1.pdf) low-energy 324 325 beta rays are focused on deposits on a filter tape and attenuated according to the approximate exponential function of particulate mass (i.e., Beer's Law). These automated samples employ a continuous filter tape. 326 The attenuation through an unexposed portion of the filter tape is measured, the tape is then exposed to 327 the ambient sample flow where a deposit is accumulated. The beta attenuation is repeated, and the 328 difference in attenuation between the blank filter and the deposit is a measure of the accumulated 329 concentration. The weighing principle used TEOM 330 in the method (https://tools.thermofisher.com/content/sfs/manuals/EPM-TEOM1405-Manual.pdf) is based on a mass 331 change detected by the sensor as a result of the measurement of a change in frequency. The tapered 332 element at the heart of the mass detection system is a hollow tube, clamped on one end and free to 333 oscillate at the other. If additional mass is added, the frequency of the oscillation decreases. An electronic 334 335 control circuit senses oscillation adds sufficient energy to the system to overcome losses. An automatic 336 gain control circuit maintains the oscillation at a constant amplitude. A precision electronic counter measures the oscillation frequency with a 10-second sampling period. Both instruments report PM 337 338 concentration every 5 min. The location of PM measurement sites is given in Tables 5.

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341 **2.4 AERONET**

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343 AErosol RObotic NETwork (AERONET) is multiband photometer with an automatic sun tracking radiometer for direct sun measurements with a spectral range of 340 to 1640 nm wavelengths. The 344 345 photometer measures the solar extinction in each wavelength to compute aerosol optical depth (Holben et al., 1998). In Israel two AERONET units type CE318-N (https://aeronet.gsfc.nasa.gov) operate in Sede 346 347 Boker and the Weizmann Institute (Fig. 3). Unfortunately, the unit in Weizmann did not operate during 6-8 September 2015 due to power failure. In this study we used AERONET Level 2.0 data (cloud 348 screened and quality assured for instrument calibration) for AOD at 500 nm and the Ångstrom Exponent 349 defined by 440-870 nm. 350

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2.5 Global, direct and diffuse solar radiation measurements

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Global radiation measured in by IMS in 22 sites (Table. 6) by instrument type Kipp & Zonen pyranometer CMP-11. Integrated solar radiation (W m-2) from 300 to 3000 nm is produced every 10 min. Diffuse and direct radiation also measured in Beit Dagan (coastal region, 31 m ASL) and Beer Sheva (southern region, 71 m ASL). For diffuse radiation, a ring is installed over the pyranometer to shade direct solar radiation. Direct radiation is measured by a sun tracker pyrheliometer.

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362 **2.6 Satellite imagery**

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364 **2.6.1 SEVIRI (MSG satellite)**

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Meteosat Second Generation (MSG) is a new series of European geostationary satellites operated by 366 EUMETSAT (European Organization for the Exploitation of Meteorological Satellites). On board the 367 MSG is a 12-channel Spinning Enhanced Visible and Infrared Imager (SEVIRI) (Roebeling et al., 2006). 368 The combination of red, blue and green (RGB) channels (12-10.8 µm, green:10.8-8.7 µm, blue:10.8 µm, 369 respectively) produce imagery of dust in pink or magneta, dry land in pale blue at daytime and pale green 370 at nighttime, thick high-level clouds in red-brown tones and thin high-level clouds nearly black 371 372 (http://oiswww.eumetsat.int/). Access to EUMETSAT imagery is given in https://www.eumetsat.int/website/home/Images/RealTimeImages/index.html. Several studies (Romano 373 374 et al., 2013; Bennouna et al., 2009; Jolivet et al., 2008) compared AOD from MGS SEVIRI and AERONET measurements clarified the uncertainty of MSG SEVIRI AOD decreases as AOD rises. For 375 376 continental aerosol type, errors do not exceed 10 % in viewing zenith angles between 20° and 50°, probable for the study area considered here. Overall, the MSG SEVIRI AOD uncertainty it is expected 377 378 to be under 15% (Mei et al, 2012). North Africa Sand storm survey (NASCube: http://nascube.univlille1.fr) obtains AOD by temperature anomalies of SEVIRI RGB by the difference in emissivity of dust 379 and desert surfaces during daytime. 380

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386 **2.6.2 MODIS (Terra and Aqua satellites)**

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The MODerate resolution Imaging Spectrometer (MODIS) instrument flies aboard the Earth Observation 388 System's (EOS) Terra and Aqua polar-orbiting satellites, with Terra on a descending orbit (southward) 389 over the equator about 10:30 local sun time, and Aqua on an ascending orbit (northward) over the equator 390 391 about 13:30 local sun time. MODIS performs measurements by 36 channels between 412 to 14200 nm whereas the aerosol retrieval makes use of seven channels (646, 855, 466, 553, 1243, 1632 and 2119 nm 392 central wavelength), and a number of other wavelength bands associated with screening procedures. As 393 land signals (AERONET) and the atmospheric signals are comparable at ~ 550 nm, errors of 0.01 in 394 assumed surface reflectance will lead to errors on the order of 0.1 in AOD retrieval (Remer et al., 2006). 395

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398 2.6.3 CALIOP (CALIPSO satellite)

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The Cloud-Aerosol Lidar with Orthogonal Polarization (CALIOP) is a two-wavelength polarization lidar 400 (1064 and 532nm) that performs global profiling of aerosols and clouds in the troposphere and lower 401 402 stratosphere. CALIOP is the primary instrument on the Cloud-Aerosol Lidar and Infrared Pathfinder Satellite Observations (CALIPSO) satellite. CALIOP is required to accurately measure signal returns 403 from the aerosol-free region between 30 and 35 km as well as the strongest cloud returns. Samples 404 acquired below 40 km for the 532-nm channel and below 30 km for the 1064-nm channel are downlinked 405 406 as profile data. Data used here was based on level 2 version 4-10 CALIPSO product, spatial resolution of 5 km (20N-50N, 20E-50E) and vertical resolution of 60 m (limited up to 6 km). 407

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410 **3. Results and discussion**

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The following description of the dust event will proceed chronologically from 7 to 10 September and include main findings of the instruments (Sect. 2). The order of the instruments described follow the most interesting features revealed, not necessarily in the same order for each day.

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416 On 7 September, images from MODIS Aqua (Fig. 4a) and MODIS Terra (Fig. 4e) taken between 417 07:20-12:10 UTC show that the dust plume progressed from northeast in a near-circular motion over the 418 Mediterranean Sea. The penetration of the dust plume to Israel was indicated at ~ 05 UTC by an increase

in AOD along a decrease in the Angström exponent (Fig. 5) from AERONET Sede Boker site. The 419 relationship between the Angström exponent decrease and the mineral dust was pointed out by Mamouri 420 421 et al., (2016). They studied the dust layer particle linear depolarization by an EARLINET lidar stationed 422 in Limassol Cyprus and concluded that the linear depolarization ratio of 0.25-0.32 on 7,10 September, indicated the dominance of the mineral dust. In addition, an increase in the PM concentration started at 423 ~ 05 UTC (not shown) reaching the highest hourly values of 107 μ g m⁻³ PM2.5 (Table 5) and 491 μ g m⁻ 424 ³ PM10 (Table 6) only in the Jerusalem elevated sites and only at 22 UTC. This 17-hour gap is shown by 425 the ceilometers' plots (Fig. 6-12) of a downward motion of the dust plume from ~ 04 UTC in all 426 427 measuring sites except for the elevated Mount Meron site (1150 m ASL, Fig. 13). Following Gasch et al (2017) cold pool outflows concept, the exception of Mount Meron site is supported by the MSG-SEVIRI 428 429 picture (Fig. 14) showing that the first dust plume was fragmented (Fig.14, red arrow) and the second 430 dust plume (Fig.14, black arrow) had not passed over Israel before 12 UTC.

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The full picture of what happened on the 7 September can be further obtained from the MSG-432 433 SEVIRI picture at 12 UTC that shows AOD to be still under 1 in most parts of Israel (Fig. 15). At the same time, Beit Dagan radiosonde profiles show the MLH was still high (700 m ASL, Fig. 16). And, 434 ceilometers' profiles at 23 UTC (Fig. 17 a) show indications of typical cloud presence from 400 m ASL 435 in the shoreline site (Tel Aviv, 5 m ASL) and up to ~700 m ASL in the elevated southern site (Nevatim, 436 400 m ASL). Clouds are identified by the peak shape of the profiles (Uzan et al, 2016) and the high 437 attenuated backscatter of 10⁻¹ m⁻¹ sr⁻¹ which in this case is 4 orders of magnitude higher than the 438 attenuated backscatter of the dust plume (shown in Fig 17 b-c). In spite of the still high MLH and the 439 cloud presence, the solar radiation measurements clearly indicates the significance of the dust plume 440 effect through the decrease of the daily direct radiation along with an increase of the diffuse radiation 441 (Fig. 18, see 7 September daily variation) measured both in central Israel (Beit Dagan) and southern 442 Israel (Beer Sheva). 443

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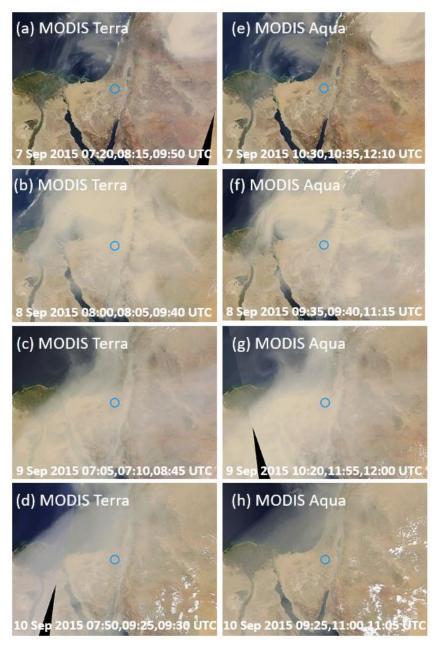


Figure 4. Pictures from MODIS terra (a-d) and MODIS Aqua (e-h). The date and time of overpass are indicated on each figure. Blue circle indicating the AERONET Sede Boker site. Source: https://aeronet.gsfc.nasa.gov.





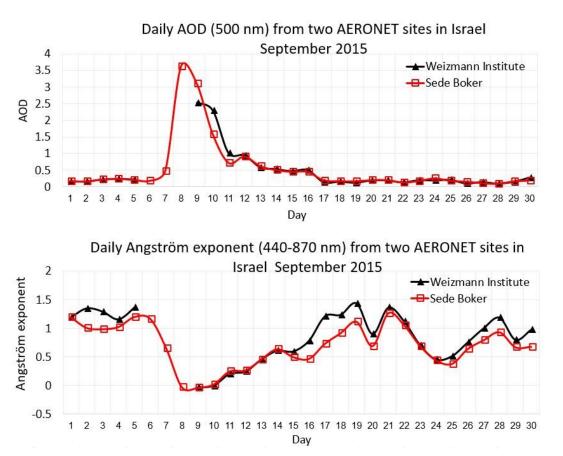


Figure 5. September 2015 daily average of AOD (top panel) and Angström exponent (bottom panel) from two AERONET sites in Israel (Sede Boker and Weizmann, see Fig.3). The Weizmann AERONET did not operate on 6-8 September due to power failure.

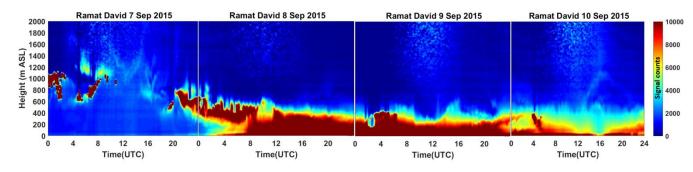


Figure 6. Ramat David ceilometer signal counts plots for 7-9 September 2015. Y-axis is the height up

to 2000 m ASL, X-axis is the time in UTC, signal count scale range between 0-10,000.

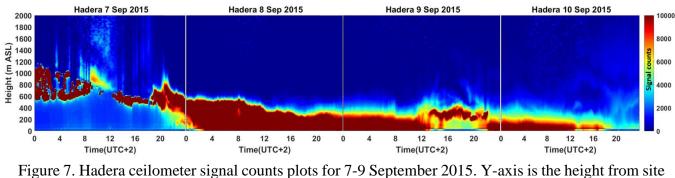


Figure 7. Hadera ceilometer signal counts plots for 7-9 September 2015. Y-axis is the height from site
deployment to 2000 m ASL, X-axis is the time in LST (UTC+2), signal count scale range between 010,000.

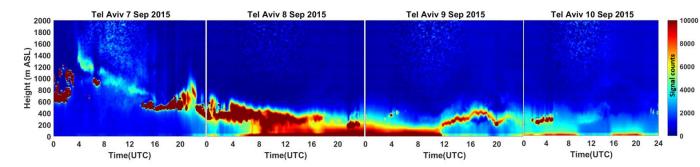


Figure 8. Tel Aviv ceilometer signal counts plots for 7-9 September 2015. Y-axis is the height from
site deployment to 2000 m ASL, X-axis is the time in UTC, signal count scale range between 010,000.

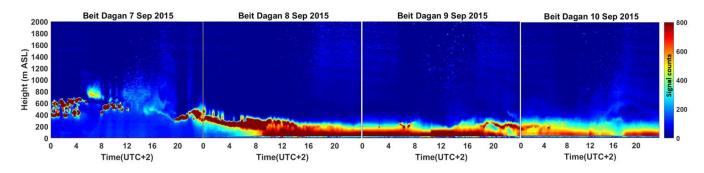


Figure 9. Beit Dagan ceilometer signal counts plots for 7-9 September 2015. Y-axis is the height from
site deployment to 2000 m ASL, X-axis is in LST (UTC+2), signal count scale range between 0-800.

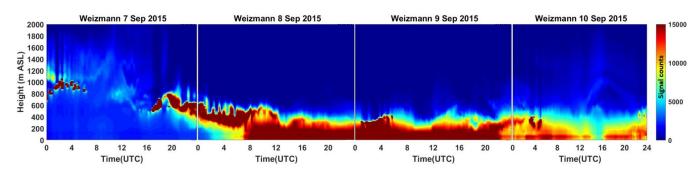


Figure 10. Weizmann ceilometer signal counts plots for 7-9 September 2015. Y-axis is the height from site deployment to 2000 m ASL, X-axis is in UTC, signal count scale range between 0-15,000.

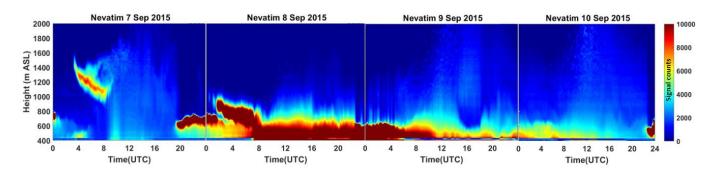


Figure 11. Nevatim ceilometer signal counts plots for 7-9 September 2015. Y-axis is the height from site deployment to 2000 m ASL, X-axis is the time in UTC, signal count scale range between 0-10,000.

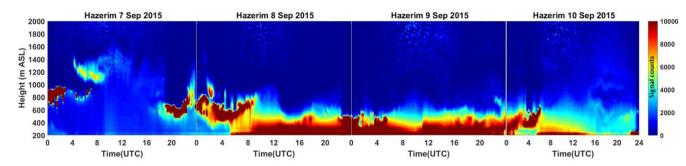


Figure 12. Hazerim ceilometer signal counts plots for 7-9 September 2015. Y-axis is the height from
site deployment to 2000 m ASL, X-axis is the time in UTC, signal count scale range between 010,000.

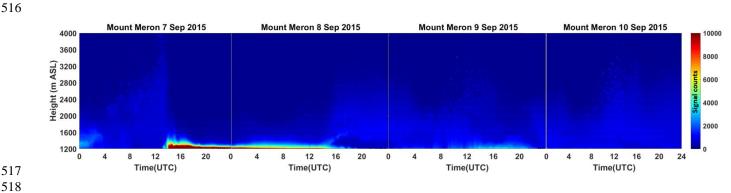
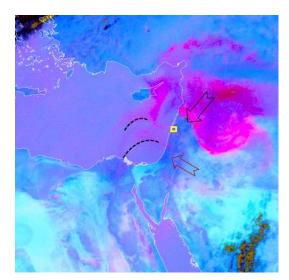


Figure 13. Mount Meron ceilometer signal counts plots for 7-9 September 2015. Y-axis is the height
from site deployment to 2000 m ASL, X-axis is the time in UTC, signal count scale range between 010,000.

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Figure 14. Picture from MSG-SEVIRI satellite of the dust RGB component (dust appears in pink colors) at 12 UTC 7 September 2015 with indications of Mount Meron ceilometer site (yellow square, Lon 33.0°, Lat 35.4°) and the dust plumes progression from east to west (red arrow and dashed lines) and from the northeast to southwest (black arrow).

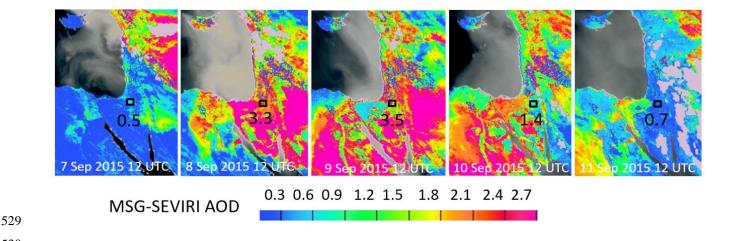
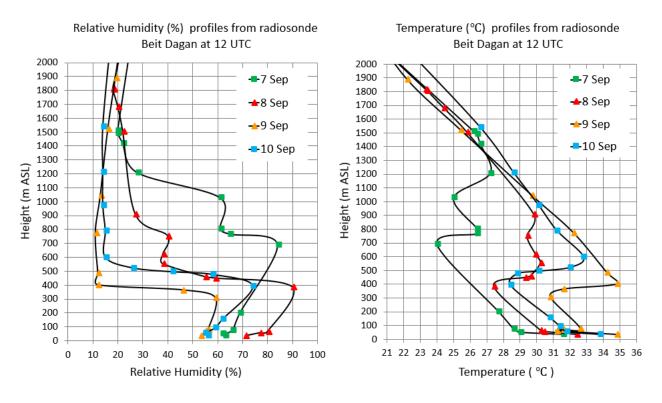


Figure 15. Aerosol Optical Depth (AOD) at 12 UTC 7-11 September 2015 analyzed by NAScube (Université de Lille) based on imagery from the MSG-SEVIRI satellite (by a combination of the SEVIRI IR8.7, IR10.8 and IR12.0 channels). The map includes indication of the Sede Boker AERONET site (black square) and its AOD value at 12 UTC.

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Figure 16. Radiosonde Beit Dagan profiles at 12 UTC between 7-10 September 2015 of relative
humidity (left panel) and temperature (right panel).

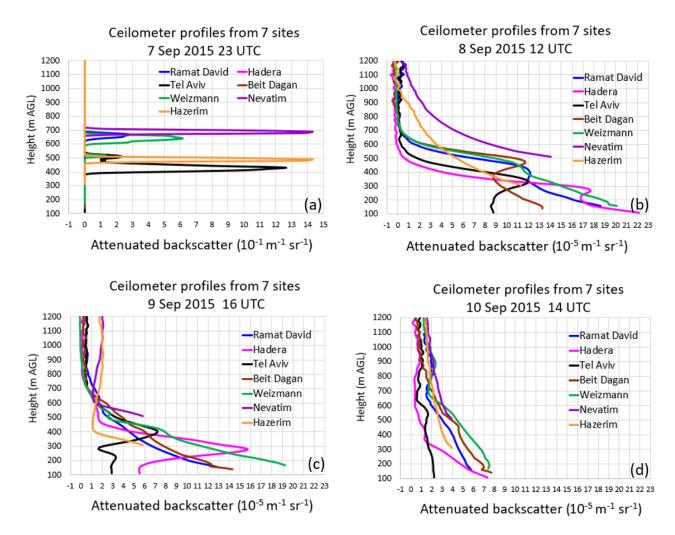


Figure 17. Ceilometer attenuated backscatter profiles from 7 sites (Ramat David, Hadera, Tel Aviv, Beit Dagan, Weizmann, Nevatim and Hazerim, Fig. 3) at 23 UTC 7 Sep 2015 (a), 8 September 2015 at 12 UTC (b), 9 September 2015 at 16 UTC (c) and 10 September 2015 at 14 UTC (d). Notice each profile begins relative to the height of its' measuring site (ASL) including a deletion of the first 100 m AGL due to inaccuracies in the first range gates of the CL31 ceilometers (for details see Sect. 2.1). Fig (a) shows cloud detection therefore it has a different scale $(10^{-1} \text{ m}^{-1} \text{ sr}^{-1})$ and a different x-axis range.

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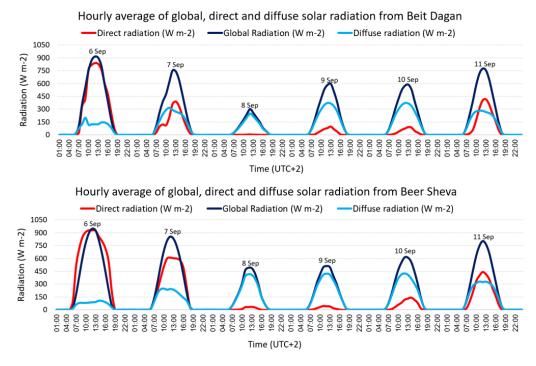


Figure 18. Hourly average of global, direct and diffuse solar radiation between 6-11 September 2015 from Beit Dagan and Beer Sheva.

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The main phase (the peak) of the dust storm occurred on 8 September. Images from MODIS Aqua (Fig. 4b) and MODIS Terra (Fig. 4f) taken between 08:00-11:15 UTC show the dust storm prevalence over Israel. Ceilometers' plots detect the descending motion of the dust plume reached ground level at ~ 08 UTC (Fig. 6-12). Simultaneously, Sede Boker AERONET AOD measurements increased up to ~4 along with a negative Angström exponent (not shown).

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563 An hour later, at ~ 09 UTC, extreme maximum PM hourly values were measured in the elevated 564 sites of Jerusalem Safra (10,280 μ g m⁻³ PM10) and Jerusalem Bar Ilan (3,063 μ g m⁻³ PM2.5). Whereby, 565 in the coast and the lower northern regions, maximum PM values were measured only 14 hours later at 566 ~23 UTC and were much lower (up to 3,459 μ g m⁻³ PM10 and 470 μ g m⁻³ PM2.5, see Tables 5-6). Fig. 567 19 illustrates the spatio-temporal variation of the PM10 extreme values, beginning at ~ 12 UTC in the 568 elevated Jerusalem sites and ending at midnight in the shoreline.

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570 Ceilometer plots from Tel Aviv, Hadera (with plot range of 0-15,000, not shown here), Beit 571 Dagan and Weizmann, reveal a two-layer shape, (beneath and above ~ 300 m ASL) starting from ~08 572 UTC. This two-layer shape later on combined into one dense layer. This pattern may explain the spatial 573 variation and time delay between the extreme PM measurements in the elevated vs. lower sites.

Referring to satellite imagery on 8 September, MSG-SEVIRI at 12 UTC (Fig. 15) underestimated 574 AOD to be 2.7 while Sede Boker AERONET measured a higher value of 3.3. Furthermore, MODIS 575 images (Fig. 4a, 4b) show a dominant dust plume over Israel, while solar global radiation measurements 576 (Fig.20a) show significant spatial variations in the reduction of the global radiation measured in the 577 different regions. Minimum values down to 200 W m⁻² were measured mainly in northern Israel. This 578 may infer the complex behavior of the dust dispersion in contrary to satellite imagery of a prevailing dust 579 storm over Israel. Additionally, in spite the extreme PM10 values of 9,031 µg m⁻³ measured in the 580 elevated southern site (Negev Mizrahi 577 m ASL, Table 6), the maximum global radiation in southern 581 Israel was still relatively high ($\sim 500 \text{ W m}^{-2}$). 582

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Overall, 8 September shows the highest PM concentrations and the lowest solar radiation levels for this dust storm event. The solar radiation was composed mainly of diffuse radiation (Fig.18) emphasizing the immense atmospheric dust loads preventing direct insolation. Surprisingly, the low solar radiation was still able to warm the ground and generate a late and weak sea breeze front (not shown). We assume the insufficient ground heating generated thermals that were too weak to create and inflate a MLH. Therefore, we assume the low MLH (300 m ASL) revealed by the Beit Dagan profiles from 8 September (Fig.16) which may indicate the dust plume base height.

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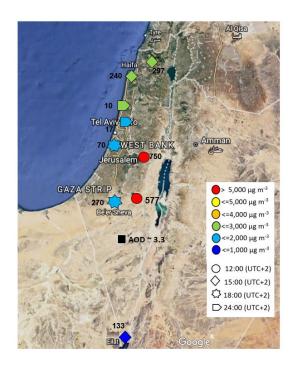


Figure 19. A map of PM10 maximum hourly concentration from 9 sites measured at 10 UTC on 8 September 2015. The map includes indications of the time of measurement (symbol shape), concentration range (symbol color), height of measurement site (numbers in black) and AERONET AOD from Sede Boker site.

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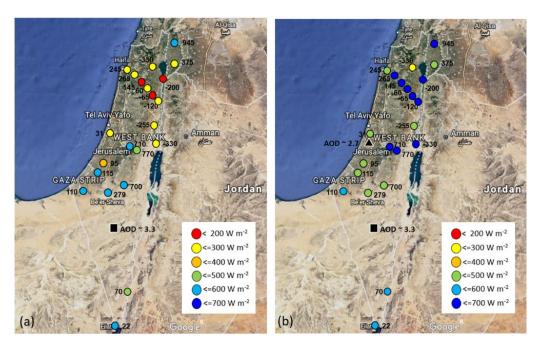


Figure 20. A map of maximum global solar radiation from 22 sites measured at 10 UTC (midday) on 8 September 2015 (a) and 9 September 2015 (b). The map includes indications of radiation range (see legend), height of measurement site (numbers in black) and AERONET AOD from Sede Boker site (black square) and Weizmann site (black triangle). On 8 September Weizmann AERONET did not operate.

609 On 9 September, MODIS images (Fig.4c and 4g) taken between 07:05-12:00 UTC show the dust plume progression southward to Egypt (Fig. 4), indicated by Sede Boker AERONET AOD >3 and a 610 negative Angström exponent (Fig. 5). Again, an underestimation of AOD MSG-SEVIRI at 12 UTC 611 shows lower AOD~ 2.7 compared to Sede Boker AERONET AOD ~ 3.5 (Fig.15). In contrary to the 612 high AOD measurements, and the descend of the MLH down to ~ 350 m ASL (Fig.16), PM values did 613 not crease but rather decrease below 900 μ g m⁻³ PM2.5 (Table 5) and 4050 μ g m⁻³ PM10 (Table 6). The 614 drop in PM concentration gave rise to an increase of solar radiation of up to 400 W m⁻² (Fig. 20 b). An 615 increase in solar radiation enables significant ground heating to temperature values measured prior to the 616 617 initiation of the dust storm (not shown). Thus, allowing generation of thermals and the creation of the sea breeze cycle which eventually produced an arc shape dust ascent visible in mainly in Tel Aviv and 618

- Hadera coastal ceilometers beginning at ~ 12 UTC (Fig.7-8). Interestingly on 9 September, compared to the peak of the dust storm on the day before, we do not see a significant difference in solar radiation in southern Israel, which continued to be relatively high ~500 W m⁻² (Fig. 20b).
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On 10 September, MODIS pictures from 7:50 -11:05 UTC (Fig. 4d and 4h) show the dust plume 623 over Israel transported southeast from Syria-Iraq to Sinai-Egypt. The CALIPSO single overpass Israel at 624 11:00-11:10 UTC revealed a dust layer between 2-4 km ASL (Fig.21). This corresponds with the 625 EARLINET lidar measurements in Limassol, Cyprus (Mamouri et al., 2016) detecting a dust plume 626 627 between 1-3 km ASL. We assume the CALIOP lidar did not produce data beneath 2 km ASL due to total attenuation. Fortunately, the ceilometers complement the dust profile (beneath 1~1 km ASL) showing a 628 629 reduction both in signal counts (Fig. 6-12) and in attenuated backscatter profiles (Fig.17d) pointing out a reduction in atmospheric dust loads. AOD from MSG-SEVIRI and Sede Boker AEONET show a 630 631 decrease down to (~1.5) and a low Angström exponent of ~0.5 indicating prevalence of mineral dust.

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633 Furthermore, a profound reduction in PM values, down to a third of the values from the day before (Table 6), was measured mainly in southern Israel. Therefore, an increase in direct radiation (therefore 634 an increase in global radiation as well) was measured in southern site as well (Fig.18). The reduction of 635 dust loads may also be denoted by the orange background color of the photograph on the 8 September 636 (Fig.1b) compared to the grey background visible on 10 September (Fig.1c). As the dust storm dissipated 637 cloud formation was visible from at ~4 UTC by ceilometers plots from Ramat David (Fig.6), Tel Aviv 638 (Fig.8), Weizmann (Fig.10) and Hazerim (Fig.11) (indicated by brown spots and evaluated by ceilometer 639 profiles -not shown), otherwise not evident by MODIS imagery (Fig 4d, 4h). 640

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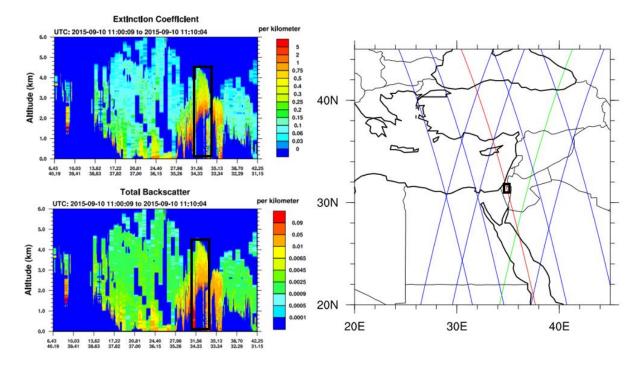


Figure 21. A map of CALIPSO satellite overpasses (right panel). The only overpass above Israel was on 10 September 2015 at 11:00-11:10 UTC (indicated by a red line). On the left panel are the CALIOP lidar products of total backscatter and extinction coefficient. Indications of the overpass over Israel is given by a black rectangle.

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As the dust storm continues to dissipate over Israel, we analyzed measurements from all instruments regarding values measured prior to the dust storm penetration. While AERONET AOD (Fig.5) and satellite imagery from MODIS and SEVIRI (not shown) detect clearance of the dust storm on 17 September. PM and ceilometer profiles indicate the dust storm ended 4 days earlier at 13 September (not shown). This postulates a scheme of two dust levels or multiple sources of the dust plumes, which may support similar conclusions from previous studies (Stavros et al, 2016; Mamouri et al, 2016, Gasch et al, 2017).

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669 **4. Conclusions**

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A very severe dust storm struck the EM on September 2015. Previous investigations including in situ and remote sensing measurements and models discussed the initiation of the dust storm in the Syrian-Iraqi border, the limitations of the models to forecast this unique event, and several aspects of its transport over the EM. The analysis concentrated mainly on the upper level of the atmosphere and analyzed specific time segments of the dust storm period. The benefit here is the provision of vertical profiles of the dust in the lower part of the troposphere and the continuous measurements at 8 sites in Israel.

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This study confirmed that the dust storm entered Israel on 7 September and showed the gradual 678 downfall of the dust plume from ~1000m ASL on 7 September down to ~400 m ASL on 8 September. 679 The detailed ceilometer profiles and auxiliary instruments enabled to separate the dust storm into two 680 dust layers (beneath and above 1 km) and show a complex dispersion which would have been challenging 681 for meso-scale model simulations. As the dust plume descended towards ground level on 8 September, 682 PM concentration increased in the elevated stations (up to 10,280 µg m⁻³ PM10) and radiation decreased 683 down to ~200 W m⁻² mainly in the northern region. On 9 September as the dust plume diluted, in spite 684 of the high AOD >3, the global radiation (mainly comprised of diffuse radiation) increased, thus enabling 685 the ground heating and the creation of a late sea breeze circulation (~12 UTC) visible as dust arcs near 686 687 the coast sites (Tel Aviv, Hadera and Beit Dagan). On 10 September, the dust plume motion continued southwest to Egypt, with indication of a dust layer between 2-4 km measured by CALIPSO overpass. 688 689 The progress of the dust storm from the Syria-Iraqi border (origin) southwest to Egypt over Israel, continued in two levels. The lower level (up to 1 km ASL) dissipated at 13 September while the level 690 691 aloft (above 1 km ASL) was observed until 17 September.

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5. Data availability

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696 PM10 measurements- Israeli Environmental ministry air quality monthly reports:

697 <u>http://www.svivaaqm.net</u>.

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699 Israeli Environmental ministry air quality monthly reports (in Hebrew):

700 <u>http://www.sviva.gov.il/subjectsEnv/SvivaAir/AirQualityData/NationalAirMonitoing/Pages/AirMoritor</u>

701 <u>ingReports.aspx</u>.

	Author contribution
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715	request.
714	Ceilometer profiles- the data is owned by governmental offices. The data is not online and provided by
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712	https://www.eumetsat.int/website/home/Images/RealTimeImages/index.html
711	http://nascube.univ-lille1.fr/cgi-bin/NAS3_v2.cgi.
710	Meteosat Second Generation Spacecraft pictures:
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708	AERONET data- https://aeronet.gsfc.nasa.gov.
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706	Radiosonde profiles –University of Wyoming: <u>http://weather.uwyo.edu/upperair/sounding.html</u> .
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704	http://www.ims.gov.il/IMS/CLIMATE/ClimateSummary/2015/hazesept+2015.htm
703	Weather reports- Israeli Meteorological Service September monthly report (in Hebrew):

718 Author contribution

Leenes Uzan carried out the research and prepared the manuscript under the careful guidance of SmadarEgert and Pinhas Alpert. The authors declare that they have no conflict of interest.

721

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Publications	Title	Main Tool	Main outcome
Pu, B. and Ginoux, P (2016)	The impact of the Pacific Decadal Oscillation on springtime dust activity in Syria	MODIS Terra MODIS Aqua DOD AOD, GFDL- AM3 model	Model underestimation in the EM due to inaccurate soil moisture
Parolari et al. (2016)	Climate, not conflict, explains extreme Middle East dust storm	WRF model	Unusual low level westerly wind spread to the EM, to reversely transport the previously eastward particles back to the EM.
Mamouri et al. (2016)	Extreme dust storm over the eastern Mediterranean in September 2015: satellite, lidar, and surface observations in the Cyprus region	MODIS, EARLINET profiles and PM10	Dust plumes from Syria entered the EM in a double layer structure, pointing to multiple dust sources
Stavros et al. (2016)	Remote sensing and modeling analysis of the extreme dust storm hitting Middle East and Eastern Mediterranean in September 2015	RAMS model EARLINET lidar, MSG and CALIPSO.	Low model ability to simulate the event, due to inaccuracies in model physical processes.
Jasim, F.H. (2016)	Investigation of the 6- 9 September 2015 Dust Storm over Middle East	Satellite MSG-SEVIRI, Meteoinfo model	Two dust storms simultaneously, from northern Syria and Sinai desert created by two low pressure systems
Gasch et al. (2017)	An analysis of the September 2015 severe dust event in the Eastern Mediterranean	ICON-ART model	An unusual early active Red Sea Trough with meso-scale convective systems generating cold-pool outflows producing the dust storm. Model lacked development of a super critical flow to produce excessive wind speeds

Table 1. The publications on the September 2015 dust event

930 Table 2. Ceilometers locations

Location	Site	Long/Lat	Distance from sho	oreline Height
			(km)	(m AGL)
Mount Meron	Northern	33.0/35.4	31	1,150
Ramat David	Northern	32.7/35.2	24	50
Hadera	Onshore	32.5/34.9	3.5	10
Tel Aviv	Onshore	32.1/34.8	0.05	5
Beit Dagan	Inland	32.0/34.8	7.5	33
Weizmann	Inland	31.9/34.8	11.5	60
Nevatim	Southern	31.2/34.9	44	400
Hazerim	Southern	31.2/34.7	70	200

- 931 *Ceilometer Weizmann is a CL51

936 Table 3. Ceilometers configurations

Location	Туре	Time resolution(sec)	Height resolution (m)	*Height range (km)	
Mount Meron	CL31	16	10	7.7	
Ramat David	CL31	16	10	7.7	
Hadera	CL31	16	10	7.7	
Tel Aviv	CL31	16	10	7.7	
Beit Dagan	CL31	15	10	7.7	
Weizmann	CL51	16	10	15.4	
Nevatim	CL31	16	10	7.7	
Hazerim	CL31	16	10	7.7	

937 * Height range dependents on sky conditions and is limited as AOD increases.

938 * In all ceilometers but in Beit Dagan site, data acquisition was limited to 4.5 km based on the BLview firmware

945 Table 4. Ceilometer technical information

Location	Туре	Engine board	Receiver	Transmitter	Firmware
Beit Dagan	CL31	CLE311	CLR311	CLT311	1.72
Weizmann	CL51	CLE321	CLRE321	CLT521	1.03

751 Table 5. Hourly maximum concentration of PM2.5, collected from 21 monitoring sites, between 7-10

952 September 2015. The values are ranked from low (dark green) to high (dark red) values.

				PM2.5 (μg m ⁻³)				
No.	Site	Height (m ASL)	Region	7-Sep-15	8-Sep-15	9-Sep-15	10-Sep-15	
1	Kefar Masarik	8	North	52	378	389	378	
2	Ahuza	280	North	36	743	650	419	
3	Newe Shaanan	240	North	43	400	466	525	
4	Nesher	90	North	43	564	496	349	
5	Kiryat Biyalic	25	North	53	424	703	447	
6	Kiryat Binyamin	5	North	40	223	412	256	
7	Kiryat Tivon	201	North	47	413	416	300	
8	Afula	57	North	44	836	550	405	
9	Raanana	54	Coast	38	173	291	229	
10	Antolonsky	34	Coast	32	470	626	386	
11	Ashdod	25	Coast	36	303	750	332	
12	Ironi D	12	Coast	34	424	507	327	
13	Tel aviv Central Station	29	Coast	41	716	803	451	
14	Ashkelon	25	Coast	61	182	537	119	
15	Jerusalem Efrata	749	Mountain	106	2285	434	403	
16	Jerusalem Bar Ilan	770	Mountain	107	3063	641	518	
17	Gedera	70	South	34	433	683	308	
18	Nir Israel	30	South	25	363	638	228	
19	Kiryat Gvaram	95	South	42	376	870	300	
20	Sede Yoav	105	South	45	323	245	228	
21	Negev Mizrahi	577	South	42	1748	526	317	

972 Table 6. Hourly maximum concentration of PM10, collected from 31 monitoring sites, between 7-10

973 September 2015. The values are ranked from low (dark green) to high (dark red) values.

				PM10 (μg m ⁻³)			
No.	Site	Height (m ASL)	Region	7-Sep-15	8-Sep-15	9-Sep-15	10-Sep-15
1	Galil Maaravi	297	North	114	3130	1987	1562
2	Karmelia	215	North	39	1120	1008	765
3	Newe Shaanan	240	North	104	3459	2471	1518
4	Haifa Port	0	North	78	1600	1965	1699
5	Nesher	90	North	117	3265	2746	1270
6	Kiryat Haim	0	North	82	1161	1625	1088
7	Afula	57	North	97	3239	2322	1961
8	Um El Kotof	0	Coast	99	2025	2028	1630
9	Orot Rabin	0	Coast	58	1152	1455	999
10	Barta	0	Coast	112	2540	2345	1612
11	Qysaria	19	Coast	54	1067	2116	1272
12	Rehuvot	70	Coast	88	2236	3045	1257
13	Givataim	0	Coast	112	1909	4014	1484
14	Yad Avner	77	Coast	61	1738	2902	1252
15	Ameil	20	Coast	96	2027	3472	1321
16	Shikun Lamed	17	Coast	51	1701	3244	1097
17	Station	29	Coast	87	1420	2176	998
18	Ashkelon	29	Coast	117	953	1692	551
19	Ariel	546	Mountain	128	2723	1481	1358
20	Jerusalem Efrata	770	Mountain	273	7820	1630	1437
21	Jerusalem Bar Ilan	749	Mountain	181	5588	1191	966
22	Jerusalem Safra	797	Mountain	491	10280	2389	1780
23	Gush Ezion	960	Mountain	310	6230	1679	1119
24	Erez	80	South	44	1000	1000	718
25	Beit Shemesh	350	South	115	2097	1943	1788
26	Carmy Yosef	260	South	85	1047	784	594
27	Modiin	267	South	185	2701	2245	1980
28	Bat Hadar	54	South	65	1342	2563	841
29	Nir Galim	0	South	94	1479	2292	1027
30	Negev Mizrahi	577	South	183	9031	2806	1730
31	Eilat	0	South	275	1867	1592	1684