Modeling the erythemal surface diffuse irradiance fraction for Badajoz, Spain

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Abstract.- Despite its important role on the human health and numerous biological processes, the diffuse component of the erythemal ultraviolet irradiance (UVER) is scarcely measured at standard radiometric stations and, therefore, needs to be estimated. This study proposes and compares ten empirical models to estimate the UVER diffuse fraction. These models are inspired from mathematical expressions originally used to estimate total diffuse fraction, but, in this study, they are applied
10   to the UVER case and tested against experimental measurements. In addition to adapting to the UVER range the various independent variables involved in these models, the total ozone column has been added in order to account for its strong impact on the attenuation of ultraviolet radiation. The proposed models are fitted to experimental measurements and validated against an independent subset. The best performing model (RAU3) is based on a model proposed by Ruiz-Arias et al. [2010] and shows values of r^2 equal to 0.91 and *rRMSE* equal to 6.1%. The performance achieved by this entirely
15   empirical model is better than those obtained by previous semi-empirical approaches, and therefore needs no additional information from other physically-based models. This study expands previous research to the ultraviolet range, and provides reliable empirical models to accurately estimate the UVER diffuse fraction.

1.- Introduction

20   Low doses of ultraviolet radiation are beneficial for human health, particularly for the synthesis of vitamin D3, critical in maintaining blood calcium levels [Webb et al., 1988; Glerup et al., 2000; Holick, 2004]. However, an excessive exposure has adverse consequences such as favoring skin cancer, immune suppression and eye disorders [Diffey, 2004; Heisler, 2010]. The effectiveness of ultraviolet radiation in producing erythema on human skin is usually quantified by the erythemal action spectrum [McKinlay and Diffey, 1987]. The ultraviolet radiation weighted by this action spectrum is named
25   erythemal ultraviolet radiation (UVER). Additionally, ultraviolet radiation may have a negative impact on ecosystems such as corals and phytoplankton communities and affect plant growth [Lesser and Farrell, 2004; Zepp et al. 2008; Häder et al., 2011, 2015]. It is also the main factor for degradation of paints and plastics exposed to outdoor conditions [Johnson and McIntyre, 1996; Verbeek et al., 2011].

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Recent studies have shown that, in addition to stratospheric ozone variability, changes in ultraviolet radiation in the last two decades have been influenced by variations in aerosols, clouds, and surface reflectivity [Arola et al., 2003; Herman, 2010]. Significant positive trends in ultraviolet radiation have been detected in different European countries and attributed to a decrease in cloud cover [Krzyscin et al., 2011; Smedley et al., 2012]. A significant positive trend of 2.1 % per decade in

UVER radiation has been detected in the Iberian Peninsula for the period 1985 – 2011 and attributed to aerosol reduction [Roman et al., 2015].

In the framework of the climate change, new variations in ultraviolet irradiance at the Earth's surface are expected for the next decades as a result of the predicted changes in clouds and aerosols [McKenzie et al., 2007; Bais et al., 2011; Williamson et al., 2014]. These variations in clouds and aerosols may affect not only the amount but also the diffuse/direct partitioning due to the stronger effectiveness of scattering at shorter wavelengths.

In contrast to the direct component, diffuse ultraviolet irradiance is difficult to block [Utrillas et al., 2010; Kudish et al., 2011]. For instance, diffuse UVER irradiance under a standard beach umbrella can reach 34% of global UVER irradiance [Utrillas et al., 2007] and up to 60% in a tree shade [Parisi, 2000]. This percentage increases notably with high load of aerosols and presence of clouds, especially in the case of broken-clouds [Alados et al., 2000; Calbó et al., 2005; Esteve et al., 2010]. However, very few studies focus on ultraviolet diffuse irradiance, mainly due to the scarcity in experimental measurements. While global ultraviolet irradiance is commonly registered worldwide, its diffuse component is seldom measured. Therefore, modeling is a good alternative to partly relieve this scarcity.

There are two main approaches to estimate solar radiation: using physically-based or empirical models. In general, the diffuse component of the radiation field is the magnitude most difficult to estimate, due to the high complexity of the processes involved. Thus, physically-based models, such as libRadtran [Mayer and Kylling, 2005], SBDART [Ricchiazzi et al., 1998] and TUV [Madronich and Flocke, 1997], require a very detailed and accurate description of the composition of the atmosphere, aerosols and clouds to reliably estimate the diffuse radiation. However, this detailed information is often unavailable and, therefore, an empirical approach is needed. Hence, in this paper the empirical approach was preferred because of its simplicity and modest requirements in terms of ancillary data. The empirical approach has been widely use by the scientific community to estimate the diffuse component in the total solar spectrum [Orgill and Hollands, 1977; Iqbal, 1983; Reindl et al., 1990; Gonzalez and Calbo, 1999; De Miguel et al., 2001; Boland et al., 2008; Ridley et al., 2010; Ruiz-Arias, 2010; Engerer, 2015].

In the particular case of the UV range, the complexity in modeling the diffuse component increases due to the higher effectiveness of the Rayleigh scattering. As far as we are aware, only a few studies have applied empirical models to estimate the diffuse solar irradiance in the ultraviolet range [Grant and Gao, 2003; Nuñez et al., 2012; Silva, 2015]. Moreover, the applicability of these studies is limited, since they rely on spectral measurements [Silva, 2015] or require information which is usually unavailable, such as cloud fraction and aerosols properties [Grant and Gao, 2003; Nuñez et al., 2012]. In this context, comprehensive studies focused on the proposal of reliable models based on commonly available data are needed.

In order to contribute to addressing this need, this study aims to propose empirical expressions for modeling hourly UVER diffuse fraction under different sky conditions and to compare their performance against experimental measurements. The

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proposed expressions will be inspired on the empirical formulae commonly used to estimate the diffuse fraction for total solar irradiance. Several radiometric and geometrical variables will be assessed in order to address their contribution to the UVER diffuse fraction. Additionally, the total ozone column will be included in the models which are proposed in this study, due to its essential role for the attenuation of ultraviolet radiation. Finally, the performance of the proposed expressions will be validated against experimental measurements.

2.- Instrumentation and data

Data presented here were collected at the radiometric station installed on the roof of the Physics building at the University Campus in Badajoz, Spain. This station is operated by the AIRE research group of the Physics Department of the University of Extremadura. This experimental site is located in south-western Spain (38.9º N; 7.01º W; 199 m a.s.l.). It is characterized by a very dry summer with prevailing cloud-free conditions, leading to noon irradiance values among the highest in Europe. Throughout the rest of the year very different cloud conditions can be found. Clouds are mainly associated with frontal systems coming from the Atlantic Ocean or to local convective systems. This region is also influenced by different aerosol types such as industrial/urban, mineral and forest fire particles. The mean aerosol optical depth at 440 nm measured at this station is 0.14 and the mean Angström exponent alpha is 1.2 [Obregón et al., 2012]. Extreme aerosol optical depth values higher than 0.3 can be occasionally reached as a result of desert dust intrusions from Sahara Desert (Northern Africa).

The period analyzed in this study comprises years 2011 and 2012, which ensures that a large variety of seasonal processes and meteorological conditions are sampled. The large variety of sun-geometry and meteorological situations that occur during a year guarantees the representativeness of the dataset for the proposal and assessment of empirical models for our location. However, it must be mentioned that snow and altitude are additional factors that have not been considered in this study. They are not represented in the dataset and the proposed models have not been tested for the processes they involve. These factors can significantly affect total UVER and the direct/diffuse ratio and, therefore, should be included for high and snowed locations.

The UVER irradiance data used in this study were recorded by two Kipp & Zonen UVS-E-T radiometers with serial numbers #000409 and #080017. The UVS-E-T radiometer measures erythemal ultraviolet irradiance between 280 and 400 nm, following the CIE action spectrum according to ISO 17166:1999 CIE S 007/E-1998 international standard [1998]. This action spectrum was originally proposed by McKinlay and Diffey [1987] to simulate the effectiveness of ultraviolet radiation in producing erythema on human skin. To ensure the reliability of the measurements, the pyranometers of our network are calibrated every two years at "El Arenosillo" Atmospheric Sounding Station of the National Institute for Aerospace Techniques (ESAt/ INTA) in Huelva, Spain (37.10º N, 7.06º W), according to the standard procedure recommended by the Working Group 4 of the COST Action 726 [Webb et al., 2006; Gröbner et al. 2007; Vilaplana et al., 2007]. This calibration procedure involves both laboratory characterization and outdoors intercomparison against a reference instrument, in our case, the QASUME unit belonging to the PMOD/WRC. The relative angular and spectral response functions are measured in the laboratory and integrated in a calibration matrix which depends on the solar zenith angle and the total ozone column (TOC). The uncertainty of UVER radiometers associated to this calibration procedure is

about 5 - 7% [Hülsen and Gröbner, 2007, Vilaplana et al., 2007]. In this study the calibration obtained during the intercomparison campaign held in July 2011 was applied to the period of study (2011-2012). The calibration matrix has been applied using the TOC values provided by the NASA Ozone Monitoring Instrument (OMI) for our location.

The data set consists of simultaneous measurements of horizontal global and diffuse UVER irradiance. Thus, while the UVS-E-T radiometer #000409 was installed on a table to measure global UVER irradiance, the UVS-E-T radiometer #080017 was installed on a Kipp & Zonen Solys 2 sun tracker to measure diffuse UVER irradiance. This device prevents the direct solar irradiance to reach the sensor by means of a small ball which continuously projects its shadow on the sensor. Since the portion of the sky obstructed by the shadow ball is negligible, no correction is required for these measurements [Ineichen et al., 1984].

Global and diffuse UVER measurements were recorded every minute by a Campbell Scientific CR-1000 data logger. Based on these data and the time of each measurement, a one-minute dataset consisting in the UVER diffuse fraction, UVER transmissivity, relative optical mass and cosine of the solar zenith angle was built. Subsequently these quantities were averaged hourly. In this study, hourly data has been used similarly to the majority of previous studies [Reindl et al., 1990; González and Calbó, 1999; Boland et al., 2001 and 2008; Ridley et al., 2010, Ruiz-Arias et al., 2010]. According to Ruiz-Arias [2010], while random errors are much lower than shorter intervals, it offers an appropriate agreement between data availability and the inherent solar radiation temporal variability. Thus, this temporal frequency is the one used by many applications, such as house energy ratings scheme software [Boland et al., 2001]. As a consequence, most of the statistical models are based on the hourly interval of the solar radiation data [Ruiz-Arias 2010, Gueymard and Ruiz-Arias, 2016].

Additionally, daily total ozone column (TOC) values as provided by the NASA Ozone Monitoring Instrument (OMI) through their website https://ozoneaq.gsfc.nasa.gov, were used in this study. Since only daily values were available, the ozone amount was assumed to be constant during each day.

3.- Methodology

The diffuse component of the solar radiation is usually quantified by the diffuse fraction (f) [Liu and Jordan, 1960; Iziomon and Aro, 1998; Utrillas et al., 2007], defined as the ratio between the diffuse (D) and the global (G) solar irradiances at the surface. The diffuse fraction describes the effectiveness of the atmosphere in scattering the incoming radiation. This magnitude is particularly interesting in the ultraviolet range since scattering is enhanced at shorter wavelengths [Kaskaoutis et al. 2006]. Using ratios of irradiances has the additional advantage of presenting an uncertainty much lower than beam or diffuse irradiances considered separately [Meloni et al., 2006; Badarinath et al., 2007]. Thus, the present study focuses on estimating the UVER diffuse fraction at the Earth's surface, f_{UVER}, defined as follows:

$$f_{UVER} = \frac{D_{UVER,0}}{G_{UVER,0}} \tag{1}$$

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where $D_{UVER,0}$ and $G_{UVER,0}$ stand for UVER diffuse and global irradiances at the Earth's surface respectively.

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Although there are very few models for estimating the ultraviolet diffuse fraction [Grant and Gao, 2003; Nuñez et al., 2012; Silva, 2015], several expressions proposed for modeling the diffuse fraction integrated along the complete solar wavelength interval (termed as total diffuse fraction), can be found in the literature (see, for example, compilations reported by Engerer [2015], and Gueymard and Ruiz-Arias [2016]). These models attempt to describe the absorption and scattering of solar

150 radiation when crossing the atmosphere. Since the mechanisms of absorption and scattering of ultraviolet solar radiation are qualitatively similar to those affecting other solar wavelengths, the models described in this study will be largely based on published models describing the total diffuse fraction. Towards this goal, a complete compilation of models for estimating total diffuse fraction was performed, the mathematical function and the variables involved were analyzed, and the most suitable models were adapted to the ultraviolet region.

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Regarding the independent variables to use, it must be noted that most empirical models for total diffuse fraction are primarily based on the total transmissivity (k_t), also named clearness index, as the main factor [Liu and Jordan, 1960; Iqbal,1980]. This quantity is defined as the ratio between the total irradiance at the Earth's surface ($G_{TOTAL,0}$) and the total irradiance at the top of the atmosphere ($G_{TOTAL,TOA}$). In this study, our proposed models will rely on the UVER transmissivity

160 (k_{UVER}), defined analogously to the total transmissivity but applied to the UVER irradiance as follows:

$$k_{UVER} = \frac{G_{UVER,0}}{G_{UVER,TOA}} \tag{2}$$

The UVER irradiance at the top of the atmosphere was calculated as follows [Iqbal, 1983]:

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$$G_{UVER,TOA} = S_{UVER} \left(\frac{r_0}{r} \right)^2 \cos(\theta) \tag{3}$$

where θ is the solar zenith angle, r_0 is the mean sun-earth distance, and r is the actual sun-earth distance for each date. The eccentricity correction-factor of the Earth's orbit, $(r_0/r)^2$, was calculated using the Spencer's formula [Spencer, 1971]. S_{UVER} is the erythemally-weighted solar constant, with an estimated value of 10.031 W/m². It must be mentioned that the solar

170 constant may vary over the solar cycle, mainly in the very short UV wavelengths [Lean et al, 1992; Kopp and Lean, 2011]. In the case of the erythemally-weighted irradiance at surface, the wavelength interval of interest starts at 290 nm and, therefore, the variation is lower than 1% [Floyd et al, 2002; DeLand and Cebula 2012; Yeo et al, 2015].

Figure 1 shows the relationship between UVER diffuse fraction (f_{UVER}) and UVER transmissivity (k_{UVER}), as derived from

175 Equations (1) and (2), respectively. The former is derived from measurements of the K&Z radiometers ($G_{UVER,0}$ and $D_{UVER,0}$) and the latter is derived from measurements of $G_{UVER,0}$ and from a computation (Eq. 3). A general dependence can be clearly

seen though the large scatter suggests the influence of other factors as well. In order to account for this variability, additional magnitudes directly related to the absorption and scattering of radiation in the atmosphere must be considered. This study focused on models relying on variables commonly available at standard radiometric stations, such as radiometric quantities and sun-geometry parameters directly related to the absorption and scattering of radiation in the atmosphere.

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Additionally, in the particular case of the ultraviolet wavelengths, the stratospheric ozone plays a very important role for modulating the radiation that arrives at the earth's surface. Therefore, in principle, the ozone amount must be included in the models. In order to test its impact on the UVER diffuse fraction (f_{UVER}), simulations with SBDART radiative transfer code [Ricchiazzi et al., 1998] were performed sampling different total ozone column values in the range 250 DU to 400 DU. This interval corresponds to the typical range of total ozone column reached in our location along a whole year. In order to analyze the effect due exclusively to changes in ozone, fixed cloud-free conditions and standard atmosphere profiles were considered. The results of the simulations showed that, in addition to showing a large impact on the beam and diffuse UVER irradiance separately, the total ozone column has a noteworthy impact on their ratio. Thus, for a solar zenith angle of 70º (the largest sampled by our measurements), changing the total ozone column from 250 DU to 400 DU yields a decrease in the UVER diffuse fraction of 4.6%. This clear variation should be considered in order to accurately estimate the UVER diffuse fraction and, therefore, the total ozone column has been included in the empirical models.

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3.1.- Models

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The approaches analyzed in this study correspond to models originally proposed for the total diffuse fraction f, but they are here applied for the ultraviolet range. This approach is justified by the fact that the physical processes of absorption and scattering of the UVER radiation are of equal nature to those affecting other solar wavelengths. Therefore, those models that succeed to describe the total diffuse fraction f are, in principle, good candidates for modelling the UVER diffuse fraction f_{UVER}.

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As mentioned above, total ozone column (TOC) is an essential attenuation factor for the UVER radiation and, therefore, it has been added to the models originally proposed for the total diffuse fraction. This new variable has been included by adding the term to each model's mathematical formula. It is to mention that a multiplicative approach consisting in the product of the model's original formula and a power function of TOC, has also been analyzed (not shown). However, the results were essentially the same as those achieved by simply adding a term and, therefore, this latter approach was preferred because of its higher simplicity and parsimony.

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3.1.1 Reindl et al.: Model REU

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The majority of empirical models for estimating the total diffuse fraction represents f as a piecewise function of k_t as the main factor [Orgill and Hollands, 1977; Erbs et al, 1982]. This relationship was first proposed by Liou and Jordan (1960)

when investigating the relationship between diffuse and global irradiances. Subsequent studies have included other variables [Bugler, 1977; Iqbal, 1980; Skartveit and Olseth, 1987; Reindl et al., 1990] in an attempt to improve the performance of the original functions.

In contrast, in the ultraviolet range, no piecewise behaviour is detected in the relationship between f_{UVER} and k_{UVER} (Figure 1) and, therefore, a single linear function was proposed for all the range of k_{UVER} values. On the other hand, Figure 1 shows a large spread in f_{UVER} for fixed values of k_{UVER}, suggesting to consider additional factors. Thus, Skartveit and Olseth (1987), and Reindl (1990) and Iqbal (1980) included the solar zenith angle in models for estimating the total diffuse fraction. This factor was added with the goal to account for the enhancement in the Rayleigh scattering as the solar zenith angle increases, mainly on clear days. This effect is even more relevant in the ultraviolet range due to the stronger effectiveness of Rayleigh scattering at shorter wavelengths. Therefore, our first model (named REU) corresponds to that originally proposed by Reindl et al. (1990) but, in this study, it is applied to the UVER case. Besides, as mentioned before, an additional term containing the total ozone column has been appended. Finally, the model REU proposed is:

$$f_{UVER}^{REU} = a + b \cdot k_{UVER} + c \cdot \cos(\theta) + d \cdot TOC \tag{4}$$

The name REU of the model stands for: "inspired by **Re**indl et al.' work but, in this study, applied to the **U**VER case". This nomenclature will be applied to the rest of models that are proposed hereinafter in this study.

3.1.2 Gonzalez and Calbo: Models GCU

The diffuse fraction shows further variability due to short-term changes in clouds or atmospheric turbidity. Gonzalez and Calbo [1999] proposed three variables Δ_1, Δ_2 and Δ_3 to account for this variability in the case of total diffuse fraction. In this study, these variables have been applied to the UVER case as follows:

$$\Delta_{UVER,1} = \ln\left(\frac{\sigma}{\overline{k_{UVER}}}\right) \tag{5}$$

$$\Delta_{UVER,2} = \ln\left(\frac{1}{(N-1)\overline{k_{UVER}}} \sum_{i=1}^{i=N} \left|k_{UVER(i+1)} - k_{UVER(i)}\right|\right) \tag{6}$$

$$\Delta_{UVER,3} = \ln\left(\frac{1}{(N-1)\overline{k_{UVER}}} \sum_{i=1}^{i=N} \left|k_{UVER,max} - k_{UVER,min}\right|\right) \tag{7}$$

where N is the number of cases, σ is the standard deviation of k_{UVER}, $\overline{k_{UVER}}$ is the mean value of the UVER transmissivity, $k_{UVER,max}$ is the maximum value and $k_{UVER,min}$ is the minimum value of k_{UVER} for each hour. Although looking similar, these variables mean different approaches to describe the short-term variability in the UVER diffuse fraction. Thus, $\Delta_{UVER,1}$ accounts for intermediate values between the minimum and the maximum, whereas $\Delta_{UVER,3}$ only depends on the extreme values. On the other hand, the fast variations between consecutive measurements are only addressed by variable $\Delta_{UVER,2}$. A logarithmic transformation is applied in Eq. 5, 6 and 7 to avoid values extending over several orders of magnitude and because this transformation increases the effect of these parameters on the diffuse fraction correlations [Gonzalez and Calbo, 1999]. It should be noted that measurements at a frequency higher than one per hour are needed to calculate these variables.

Similarly to the proposal of Gonzalez and Calbo [1999] for total diffuse fraction, variables $\Delta_{UVER,1}$, $\Delta_{UVER,2}$ and $\Delta_{UVER,3}$ were added to the model REU previously built for UVER diffuse fraction (Eq. 4), resulting in new models that are named GCU1, GCU2, GCU3:

$$f_{UVER}^{GCU1} = a + b \cdot k_{UVER} + c \cdot \cos(\theta) + d \cdot TOC + g \cdot \Delta_{UVER,1} \tag{8}$$

$$f_{UVER}^{GCU2} = a + b \cdot k_{UVER} + c \cdot \cos(\theta) + d \cdot TOC + g \cdot \Delta_{UVER,2} \tag{9}$$

$$f_{UVER}^{GCU3} = a + b \cdot k_{UVER} + c \cdot \cos(\theta) + d \cdot TOC + g \cdot \Delta_{UVER,3} \tag{10}$$

3.1.3 Boland et al. [2001]: Model BOU

Boland et al. [2001] proposed a logistic function to estimate the total diffuse fraction as a function of the total transmissivity. The logistic functions are S-shaped sigmoid curves where the increase is approximately exponential at the initial stage and, then, the growth slows as saturation begins. This behavior, but with decay, can be useful to describe the dependence of total diffuse fraction (f) on k_t [Boland et al., 2001, 2008]. Thus, f decreases as k_t increases, but with a saturation effect towards the clear sky value. This behaviour is also observed in the ultraviolet range and, therefore, the model proposed by Boland et al. has been applied to the UVER case. Additionally, a term including the total ozone column has been added to the exponent of the exponential function. The resulting model, named BOU, is:

$$f_{UVER}^{BOU} = \frac{1}{1 + \exp(a + b \cdot k_{UVER} + d \cdot TOC)} \tag{11}$$

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3.1.4 Ridley et al.: Model RIU

The original expression proposed by Boland et al. [2001] was later expanded by Ridley et al. [2010] to include four additional variables: 1) the solar zenith angle, 2) the apparent solar time *AST*, which accounts for differences in the atmosphere between morning and afternoon, 3) the daily clearness index *K* calculated as the ratio between the irradiation accumulated along the whole day at the Earth's surface and its value at the top of the atmosphere, and 4) a variable Ψ to account for the persistence at one hourly scale due to the very slow rate of change in the radiation under cloud-free or overcast skies.

Similarly, the UVER daily clearness index (K_{UVER}) and the persistence parameter (Ψ_{UVER}) have been calculated for the UVER case as follows:

$$\Psi_{UVER} = \frac{k_{UVER,i-1} + k_{UVER,i+1}}{2}$$

(12)

$$K_{UVER} = \frac{\sum\limits_{sunrise}^{sunset} G_{UVER}(0)}{\sum\limits_{sunrise}^{sunset} G_{UVER}(TOA)}$$

(13)

where $k_{UVER,i-1}$ and $k_{UVER,i+1}$ are the values before and after each hourly value of k_{UVER}. The apparent solar time was calculated according to the Spencer's formulae [Spencer, 1971; Iqbal, 1983].

It has to be noted that, in this study, the variable θ used by Ridley et al. [2010] has been replaced by $cos(\theta)$ in order to facilitate the comparison with the other models. Models with θ and $cos(\theta)$ were tested and showed nearly equal performance (not shown here). Following the Ridley et al.'s expansion, the new model, named RIU, is based on model BOU, where additional terms containing $cos(\theta)$, *AST*, Ψ_{UVER} and K_{UVER} have been included in the exponent of the exponential function:

$$f_{UVER}^{RIU} = \frac{1}{1 + \exp(a + b \cdot k_{UVER} + c \cdot \cos(\theta) + d \cdot TOC + g \cdot AST + h \cdot \Psi_{UVER} + j \cdot K_{UVER})}$$

(14)

3.1.5 Kuo et al.: Model KUU

Kuo et al. [2014] developed several correlation models aimed to estimate the hourly solar diffuse fraction in Taiwan. They compared four newly proposed models with fourteen models previously available in the literature. As a result of the comparison, they proposed a new model consisting of a multiple linear combination of the same independent variables included in Ridley et al.'s model. In this study, following Kuo et al.'s suggestion, a model named KUU was built for the

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UVER case, as follows:

$$f_{UVER}^{KUU} = a + b \cdot k_{UVER} + c \cdot \cos(\theta) + d \cdot TOC + g \cdot AST + h \cdot \Psi_{UVER} + j \cdot K_{UVER}$$

(15)

310 **3.1.6 Ruiz-Arias et al.: Models RAU**

Similarly to Ridley et al. [2010], Ruiz-Arias et al. [2010] proposed a model for the total diffuse fraction (*f*) based on a sigmoid function of the total transmissivity (k_t), but included also the Kasten and Young's relative optical mass [1989] as an additional predictor (*m*). In fact, Ruiz-Arias et al. [2010] proposed three versions of their model corresponding to combinations of k_t and *m* raised to various powers. Correspondly, in this study, these three models have been applied to the

315 UVER case and, additionally, a term including the ozone total column has been added to the exponent. Finally, three models named RAU1, RAU2 and RAU3 have been built as follows:

$$f_{UVER}^{RAU1} = A + B \cdot \exp(\exp(a + b \cdot k_{UVER} + d \cdot TOC))$$

(16)

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$$f_{UVER}^{RAU2} = A + B \cdot \exp(\exp(a + b \cdot k_{UVER} + c \cdot m + d \cdot TOC))$$

(17)

$$f_{UVER}^{RAU3} = A + B \cdot \exp(\exp(a + b \cdot k_{UVER} + c \cdot m + d \cdot TOC + g \cdot k_{UVER}^2 + h \cdot m^2))$$

(18)

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3.2 Fitting and comparison statistics

This study aims to fit the models to experimental data and subsequently compare their performance using an independent dataset. Towards that aim, the hourly dataset was randomly divided in two subsets: 1) the fitting subset, containing the 75% of data (3979 cases), for fitting the coefficients of the models, and 2) the validation subset, containing the remaining 25% of

330 data (1262 cases), for model validation and comparison. In principle, linear fitting is preferred since it requires no starting values of the fitting coefficients. Therefore, linear least squares fitting was applied whenever possible, that is, to models which are linear (REU, CGU1, CGU2, CGU3 and KUU) or linearizable, i.e. that can be reduced to a linear form with a change of variables (BOU and RIU). For the remaining cases (RAU1, RAU2 and RAU3) it was necessary to apply non linear fittings.

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The performance of the models proposed to estimate the UVER diffuse fraction was compared using both statistical and graphical tools. The coefficient of determination (r^2) and the relative root-mean-square error (*rRMSE*) were used to assess the goodness of fit of the models and their performance. The coefficient of determination is a measure of the proportion of

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total variance explained by the model, while the relative root-mean-square error quantifies the difference between modeled and measured values. They are defined as:

$$r^2 = 1 - \frac{\sum_{i=1}^{i=N} \left(x_i - x_i^* \right)^2}{\sum_{i=1}^{i=N} \left(x_i - \overline{x} \right)^2}$$

(19)

$$rRMSE\,(\%) = \frac{100}{\overline{x}} \sqrt{\frac{1}{N} \sum_{i=1}^{i=N} \left(x_i - x_i^* \right)^2}$$

(20)

where x_i^* and x_i are the modeled and measured values, respectively.

Additionally, the Taylor diagram [Taylor, 2001] and the relative differences were used for model comparison. The Taylor diagram provides a concise graphical summary of different aspects of the performance of a model such as the centered root-mean-square error, the correlation and the standard deviation. On the other hand, the relative residuals between modeled, x_i^* and measured, x_i, values are calculated as follows:

$$Residuals\,(\%) = 100 \cdot \frac{x_i - x_i^*}{x_i}$$

(21)

and can be analyzed as a function of the solar zenith angle, the UVER transmissivity, and the UVER diffuse fraction, providing interesting information about possible remaining dependences.

4. Results and Discussion

Main results of the fitting of each empirical model to the fitting subset are summarized in Table 1. Ordinary least squares fitting (also known as linear least squares) for models REU, GCU1, GCU2, GCU3, BOU, RIU and KUU, and non-linear fitting for models RAU1, RAU2 and RAU3 have been calculated. As mentioned in Section 3, some models involve parameters accounting for the short-term fluctuation. In particular, models based on Gonzalez and Calbo [1999], that is, GCU1, GCU2 and GCU3, include parameters $\Delta_{UVER,1}$, $\Delta_{UVER,2}$, $\Delta_{UVER,3}$, respectively. In the case of model RAU3, inspired from Ruiz-Arias et al. [2010], the parameters kt² and m² are introduced to account for the short-term variability. Results in Table 1 show that model GCU2, with a short-term variability parameter, presents a better performance than model REU, which has a same functional form without a short-term variability parameter. Conversely, models GCU1 and GCU3 do not show a better performance than REU despite to include short-term variability parameters. This primacy of $\Delta_{UVER,2}$ agrees with the case of total diffuse fraction as reported by Gonzalez and Calbó [1999]. In the case of models inspired from Ruiz-

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Most of the models performed notably well with r^2 higher or equal than 0.83 and *rRMSE* lower or equal than 8.6%, except for models BOU and RIU, which perform somewhat worse. This fact confirms the general suitability of different mathematical functional forms for estimating UVER diffuse fraction and emphasizes the need for comprehensive comparison studies like the present one.

Subsequently, the various models with their fitted coefficients were applied to the validation subset. The resulting r^2 and *rRMSE* values are shown also in Table 1. The values are very similar to those obtained for the fitting, indicating no overfitting effect. The best statistics are achieved by the three-variable model RAU3, with an excellent coefficient of determination of 0.91 and a low relative root mean squared error of only 6.4%. This model includes k_{UVER}, m and *TOC* as predictors.

Taylor diagram (Figure 2) confirms the generally good performance achieved by the proposed models, but also identifies two separate groups: on one hand, models BOU and RIU, and on the other hand, models REU, GCU2, KUU and RAU3, the latter performing moderately better. It is to note that the worst-performing models BOU and RIU are based on the same logistic function proposed by Boland et al. [2008]. It can be, therefore, concluded that such functional form is not as appropriate for the UVER case as those used by the remaining models. Moreover, that worse performance is not improved even when more variables are included such as in model RIU.

Models REU and KUU completely overlap, indicating that no improvement is achieved when variables *AST*, Ψ_{UVER} and K_{UVER} are added. Conversely, the variable $\Delta_{UVER,2}$, which was included with the aim to account for the short-term variability, means a substantial contribution to the better performance achieved by the model CGU2.

In addition to the regression statistics mentioned above, the relative residuals between measured and modeled values were calculated, and their variation with respect to solar zenith angle, UVER transmissivity and UVER diffuse fraction bands was analyzed. In order to clearly show the relationship with a particular variable, the relative residuals were averaged by intervals in that variable.

Figure 3 confirms the worse performance achieved by models BOU and RIU. The relative residuals for these two models are the largest among the models proposed in this study. These large residuals occur for low solar zenith angle, high k_{UVER} and low diffuse fraction, which correspond to cloud-free conditions near noontime. In addition, models BOU and RIU's relative residuals show a clear relationship with the three independent variables analyzed, suggesting that their functional

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form does not properly account for the relationship of f_{UVER} with the solar zenith angle and with k_{UVER}.

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In contrast, models REU, GCU2, KUU and RAU3 show much smoother patterns, with absolute relative residuals smaller than 5% for almost the entire range of θ and k_{UVER}. Concerning UVER diffuse fraction, these models tend to underestimate for intermediate values and overestimate for high values over 0.8. The model RAU3 is again the preferred model, with absolute relative residuals smaller than 3% except for the lowest values of the UVER diffuse fraction.

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Table 2 shows the fitting coefficients for each proposed models. It is important to note that the particular values of the coefficients are specific for our local conditions. Therefore, in order to apply the models to other locations, the coefficients should be calculated by fitting to local measurements.

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5.- Conclusions

This study aims to accurately estimate hourly UVER diffuse fraction at the earth's surface using empirical models. Towards this goal, ten mathematical expressions are proposed and their performance is compared to experimental measurements. All the empirical models analyzed are based on mathematical expressions originally suggested by Reindl. et al [1990], Gonzalez and Calbo [1999], Boland et al. [2008], Ridley et al. [2010], Kuo [2014], and Ruiz-Arias et al. [2010] for modeling the total diffuse fraction but, in this study, they are applied to the UVER case. Among a complete compilation of formulae used for estimating total diffuse fraction, those models that rely on variables commonly available at standard radiometric stations are selected. This criterion is applied in order to favor the general applicability of the results of the study. Additionally, a term including the total ozone column is added to account for the important role played by the stratospheric ozone in modulating the ultraviolet radiation that arrives at the earth's surface. As a result, the models REU, GCU1, GCU2, GCU3, BOU, RIU, KUU, RAU1, RAU2 and RAU3 are built, fitted against experimental data and finally validated.

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The fitting to experimental measurements revealed a generally good performance of all models except for models BOU and RIU, which perform somewhat worse. It can be said that the proposal of mathematical expressions and variables succeed to describe the variation in the UVER diffuse fraction. Results indicate that multiple linear combinations and the sigmoid function suggested by Ruiz-Arias [2010] are more suitable for the UVER case than the logistic function proposed by Boland et al. [2008]. In the case of total integrated radiation, logistic models proved to be useful since they reliably describe the abrupt change shown by the relationship between the total diffuse fraction and the total transmissivity. However, for the UVER measurements that relationship is much smoother and, therefore, the logistic models BOU and RIU provide no improvement with respect to more simple linear models REU, GCU2 and KUU. Conversely, the more complex sigmoid function proposed by Ruiz-Arias et al. [2010] achieves the best fitting statistics.

440 The fitting results are confirmed by the validation against an independent subset of measurements. The best performing model is RAU3 followed by GCU2, REU and KUU, and finally by RIU and BOU, which perform notably worse, with r^2 lower than 0.8 and $rRMSE$ higher than 9%. In particular, the model RAU3 achieves an excellent coefficient of determination of 0.91 and a low relative root mean squared error of only 6.4%. These are very good numbers compared to the only two approaches for UVER diffuse fraction that, to our knowledge, have been published up to date. Thus, Nunez et al.'s semi-

445 empirical approach applied to Valencia (Spain) achieved an r^2 equal to 0.84 [Nuñez et al., 2012], and Silva reported an r^2 of 0.79 for his study of Belo Horizonte (Brazil) [Silva, 2015]. It is important to notice that model RAU3 achieved better r^2 value while being the only entirely empirical model and, therefore, needing no additional information from physically-based models. This is an important advantage since the latter require detailed information which is often unavailable, limiting their applicability.

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Regarding the residuals, RAU3 is again the best model, with almost all absolute values smaller than 3% and no dependency with θ, k_{UVER}, nor f_{UVER}. Then, it can be concluded that the proposal of models have succeed in providing empirical models to accurately estimate the UVER diffuse fraction, with the RAU3 model being the preferred one.

455 This study positively contributes to estimate UVER diffuse irradiance and UVER diffuse fraction in locations where only UVER global irradiance measurements are available. Additionally, the models proposed here can be used to expand time series of UVER diffuse radiation to periods when global but not diffuse UVER irradiance was being measured. It should be mentioned that factors affecting the UVER diffuse fraction such as the altitude or surface albedo have not been tested in this study. Moreover, since only solar zenith angles below 70º have been considered, the models may not be suitable for the

460 large solar zenith angles encountered at high latitudes. Therefore, similar research must be conducted to assess the general validity of the proposed models and/or establish their possible adaptation to other locations. Although these results apply mainly to regions with similar characteristics than those analyzed in this study, the methodology and comparisons described in this paper can be used to develop similar analysis for other locations.

465 **6.- Data availability:** The data analyzed in this study are available from authors upon request (guadalupesh@unex.es).

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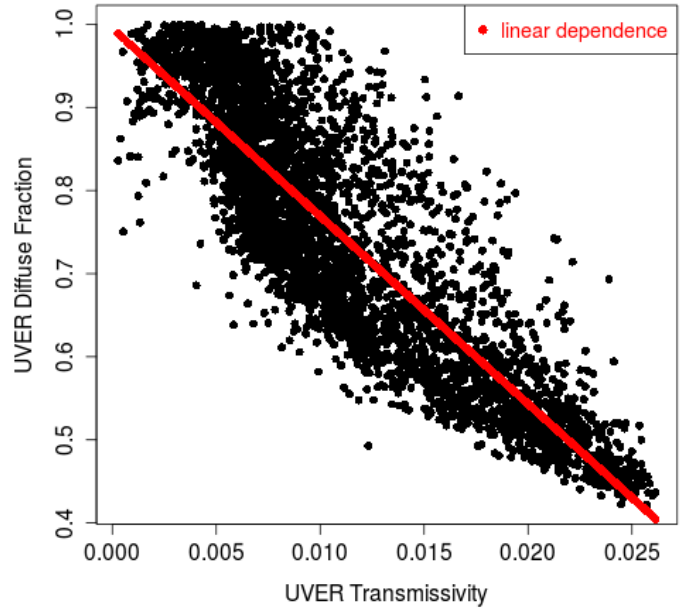
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Figure 1. UVER diffuse fraction (f_{UVER}) versus UVER transmissivity (k_{UVER}), and linear fitting.
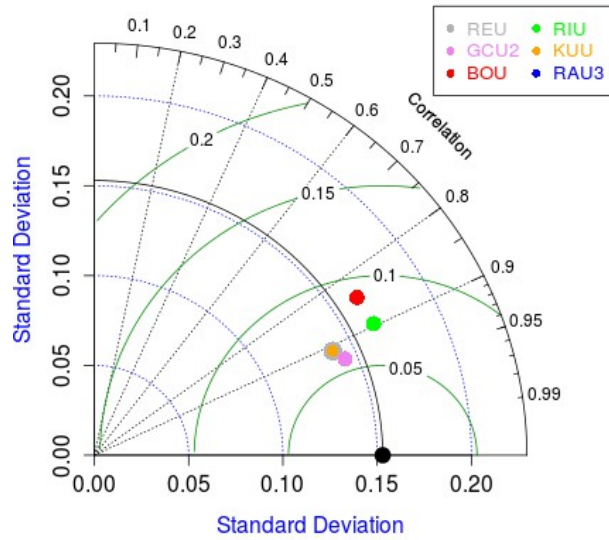
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745 Figure 2. Taylor diagram showing the performance of the models proposed to estimate the diffuse fraction, as compared to experimental measurements. This diagram summarizes different aspects of the performance of a model such as the centered root-mean-square error (green lines), the correlation and the standard deviation with respect to the reference data set (black dot).
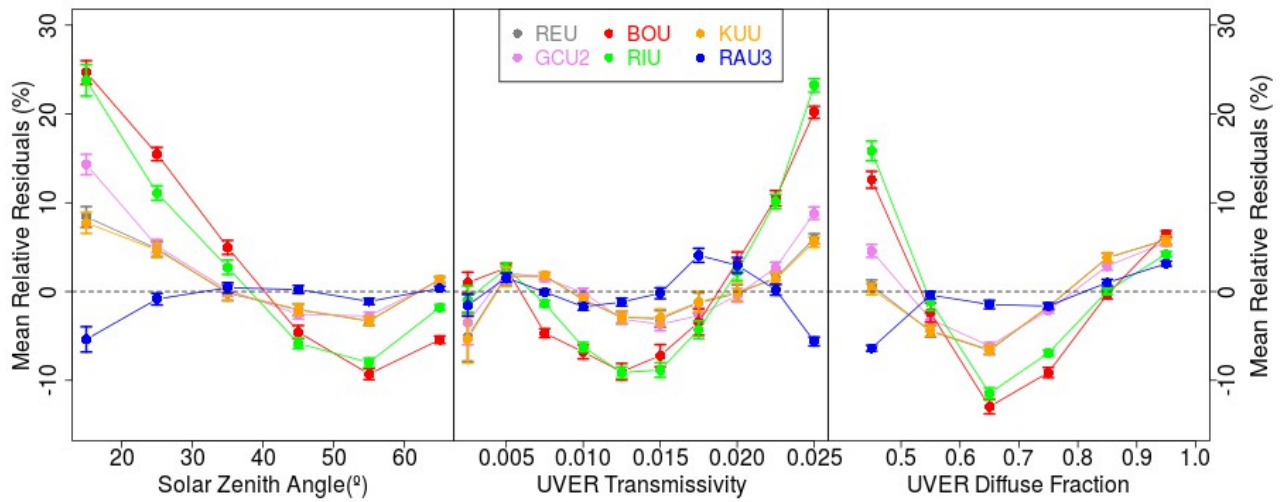
Figure 3. Mean relative residuals of each UVER diffuse fraction model versus a) solar zenith angle, b) UVER transmissivity, and c) predicted UVER diffuse fraction values. The range of the solar zenith angle (10º- 70º) has been divided into six equal bins while the ranges of the k_{UVER} (0, 0.025) and f_{UVER} (0,1) have been split into ten bins. The mean value for each bin (central point) and the standard deviation of the mean are shown (error bars)

Table 1. Coefficient of determination and relative root mean squared error corresponding to the fitting against experimental measurements and the validation of each model.

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Model	Fitting (3979 cases)		Validation (1262 cases)	
	R²	rRMSE (%)	R²	rRMSE (%)
REU	0.83	8.5	0.82	8.8
GCU1	0.83	8.5	0.83	8.8
GCU2	0.86	7.7	0.85	8.0
GCU3	0.83	8.4	0.84	8.6
BOU	0.66	12.0	0.68	11.8
RIU	0.78	9.8	0.78	9.7
KUU	0.83	8.6	0.82	8.8
RAU1	0.76	10.12	0.76	10.4
RAU2	0.82	8.8	0.81	9.2
RAU3	0.91	6.1	0.91	6.4

Table 2. Functional form of the models and empirical fitting coefficients with their corresponding standard error for Badajoz, Spain.

Model	Expression for f_{UVER}
REU	$(1.20\pm0.01)+(-35.4\pm0.4)\cdot k_{UVER}+(0.50\pm0.01)\cdot\cos(\theta)+(-1.12\pm0.04)\,x\,10^{3}\cdot TOC$
GCU2	$(1.34\pm0.01)+(-28.4\pm0.4)\cdot k_{UVER}+(0.32\pm0.01)\cdot\cos(\theta)+(-1.12\pm0.03)\,x10^{-3}\cdot TOC+(0.032\pm0.001)\cdot\Delta_{UVER,2}$
BOU	$\dfrac{1}{1+\exp\left((-3.7\pm0.1)+(146\pm2)\cdot k_{UVER}+(2.4\pm0.4)\,x\,10^{-3}\cdot TOC\right)}$
RIU	$\dfrac{1}{1+\exp\left((-5.0\pm0.1)+(71\pm14)\cdot k_{UVER}+(-4.2\pm0.1)\cdot\cos(\theta)+(9.9\pm0.4)\,x10^{-3}\cdot TOC+(-5.0\pm10)\,x10^{-3}\cdot AST+(-46\pm14)\cdot\Psi_{UVER}+(187\pm26)\cdot K_{UVER}\right)}$
KUU	$(1.23\pm0.01)+(-59\pm13)\cdot k_{UVER}+(0.47\pm0.01)\cdot\cos(\theta)+(-1.17\pm0.04)\,x10^{-3}\cdot TOC+(2.3\pm1.4)\,x10^{-3}\cdot AST+(23\pm13)\cdot\Psi_{UVER}+(-11\pm2)\cdot K_{UVER}$
RAU3	$(0.50\pm0.01)+(0.51\pm0.01)\cdot\exp\left(\exp\left((-23.4\pm0.6)+(788\pm20)\cdot k_{UVER}+(9.1\pm0.3)\cdot m+(-13.3\pm0.4)\,x10^{3}\cdot k_{UVER}^{2}+(-1.61\pm0.05)\cdot m^{2}+(1.76\pm0.05)\,x10^{(-2)}\cdot TOC\right)\right)$