

Influence of temperature on the molecular composition of ions and charged clusters during pure biogenic nucleation

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45

46 **Abstract**

47 It was recently shown by the CERN CLOUD experiment that biogenic highly oxygenated
48 molecules (HOMs) form particles under atmospheric conditions in the absence of sulfuric
49 acid where ions enhance the nucleation rate by one to two orders of magnitude. The biogenic
50 HOMs were produced from ozonolysis of α -pinene at 5°C. Here we extend this study to
51 compare the molecular composition of positive and negative HOM clusters measured with
52 atmospheric pressure interface time-of-flight mass spectrometers (APi-TOFs), at three
53 different temperatures (25°C, 5°C and -25°C). Most negative HOM clusters include a nitrate
54 (NO_3^-) ion and the spectra are similar to those seen in the nighttime boreal forest. On the
55 other hand, most positive HOM clusters include an ammonium (NH_4^+) ion and the spectra are
56 characterized by mass bands that differ in their molecular weight by ~ 20 C atoms,
57 corresponding to HOM dimers. At lower temperatures the average oxygen to carbon (O:C)
58 ratio of the HOM clusters decreases for both polarities, reflecting an overall reduction of
59 HOM formation with decreasing temperature. This indicates a decrease in the rate of
60 autoxidation with temperature due to a rather high activation energy as has previously been
61 determined by quantum chemical calculations. Furthermore, at the lowest temperature
62 (-25°C) the presence of C_{30} clusters show that HOM monomers start to contribute to the
63 nucleation of positive clusters. These experimental findings are supported by quantum
64 chemical calculations of the binding energies of representative neutral and charged clusters.

65 **1. Introduction**

66 Atmospheric aerosol particles directly affect climate by influencing the transfer of radiant
67 energy through the atmosphere (Boucher et al., 2013). In addition, aerosol particles can

68 indirectly affect climate, by serving as cloud condensation nuclei (CCN) and ice nuclei (IN).
69 They are of natural or anthropogenic origin, and result from direct emissions (primary
70 particles) or from oxidation of gaseous precursors (secondary particles). Understanding
71 particle formation processes in the atmosphere is important since more than half of the
72 atmospheric aerosol particles may originate from nucleation (Dunne et al., 2016; Merikanto
73 et al., 2009).

74 Due to its widespread presence and low saturation vapor pressure, sulfuric acid is believed to
75 be the main vapor responsible for new particle formation (NPF) in the atmosphere. Indeed,
76 particle nucleation is dependent on its concentration, albeit with large variability (Kulmala et
77 al., 2004). The combination of sulfuric acid with ammonia and amines increases nucleation
78 rates due to a higher stability of the initial clusters (Almeida et al., 2013; Kirkby et al., 2011;
79 Kürten et al., 2016). However, these clusters alone cannot explain the particle formation rates
80 observed in the atmosphere. Nucleation rates are greatly enhanced when oxidized organics
81 are present together with sulfuric acid, resulting in NPF rates that closely match those
82 observed in the atmosphere (Metzger et al., 2010; Riccobono et al., 2014). An important
83 characteristic of the organic molecules participating in nucleation is their high oxygen content
84 and consequently low vapor pressure. The formation of these highly oxygenated molecules
85 (HOMs) has been described by Ehn et al. (2014), who found that, following the well-known
86 initial steps of α -pinene ozonolysis through a Criegee intermediate leading to the formation of
87 an $\text{RO}_2\cdot$ radical, several repeated cycles of intramolecular hydrogen abstractions and O_2
88 additions produce progressively more oxygenated $\text{RO}_2\cdot$ radicals, a mechanism called
89 autoxidation (Crouse et al., 2013). The (extremely) low volatility of the HOMs results in
90 efficient NPF and growth, even in the absence of sulfuric acid (Kirkby et al., 2016; Tröstl et
91 al., 2016). The chemical composition of HOMs during NPF has been identified from α -
92 pinene and pinanediol oxidation by Praplan et al. (2015) and Schobesberger et al. (2013),
93 respectively.

94 Charge has also been shown to enhance nucleation (Kirkby et al., 2011). Ions are produced
95 in the atmosphere mainly by galactic cosmic rays and radon. The primary ions are N^+ , N_2^+ ,
96 O^+ , O_2^+ , H_3O^+ , O^- and O_2^- (Shuman et al., 2015). These generally form clusters with water
97 (e.g. $(\text{H}_2\text{O})\text{H}_3\text{O}^+$) and after further collisions the positive and negative charges are transferred
98 to trace species with highest and lowest proton affinities, respectively (Ehn et al., 2010). Ions

99 are expected to promote NPF by increasing the cluster binding energy and reducing
100 evaporation rates (Hirsikko et al., 2011). Recent laboratory experiments showed that ions
101 increase the nucleation rates of HOMs from the oxidation of α -pinene by one to two orders of
102 magnitude compared to neutral conditions (Kirkby et al. 2016). This is due to two effects, of
103 which the first is more important: 1) an increase in cluster binding energy, which decreases
104 evaporation and 2) an enhanced collision probability, which increases the condensation of
105 polar vapors on the charged clusters (Lehtipalo et al., 2016; Nadykto, 2003).

106 Temperature plays an important role in nucleation, resulting in strong variations of NPF
107 at different altitudes. Kürten et al. (2016) studied the effect of temperature on nucleation for
108 the sulfuric acid - ammonia system, finding that low temperatures decrease the needed
109 concentration of H_2SO_4 to maintain a certain nucleation rate. Similar results have been found
110 for sulfuric acid – water binary nucleation (Duplissy et al., 2016; Merikanto et al., 2016),
111 where temperatures below 0°C were needed for NPF to occur at atmospheric concentrations.
112 Up to now, no studies have addressed the temperature effect on NPF driven by HOMs from
113 biogenic precursors such as α -pinene.

114 In this study we focus on the chemical characterization of the ions and the influence of
115 temperature on their chemical composition during organic nucleation in the absence of
116 sulfuric acid. The importance of such sulfuric acid-free clusters for NPF has been shown in
117 the laboratory (Kirkby et al., 2016; Tröstl et al., 2016) as well as in the field (Bianchi et al.,
118 2016). We present measurements of the NPF process from the detection of primary ions (e.g.
119 N_2^+ , O_2^+ , NO^+) to the formation of clusters in the size range of small particles, all under
120 atmospherically relevant conditions. The experiments were conducted at three different
121 temperatures (-25 , 5 and 25°C) enabling the simulation of pure biogenic NPF representative
122 of different tropospheric ~~conditions. altitudes.~~ This spans the temperature range where NPF
123 might occur in tropical or sub-tropical latitudes (25°C), high-latitude boreal regions (5°C)
124 and the free troposphere (-25°C). For example, NPF events were reported to occur in an
125 Australian Eucalypt forest (Suni et al., 2008) and the boreal station in Hyytiälä (Kulmala et
126 al., 2013). Nucleation by organic vapors was also observed at a high mountain station
127 (Bianchi et al, 2016). High aerosol particle concentrations were measured in the upper
128 troposphere over the Amazon Basin and tentatively attributed to the oxidation of biogenic
129 volatile organic compounds (Andreae et al., 2017).

130

131 2. Methods

132 The CLOUD chamber

2.1.

133 We conducted experiments at the CERN CLOUD chamber (Cosmics Leaving Outdoor
134 Droplets). With a volume of 26.1 m³, the chamber is built of electropolished stainless steel
135 and equipped with a precisely controlled gas system. The temperature inside the chamber is
136 measured with a string of six thermocouples (TC, type K) which were mounted horizontally
137 between the chamber wall and the center of the chamber at distances of 100, 170, 270, 400,
138 650, and 950 mm from the chamber wall (Hoyle et al., 2016). The temperature is controlled
139 accurately (with a precision of $\pm 0.1^\circ\text{C}$) at any tropospheric temperature between -65 and
140 30°C (in addition, the temperature can be raised to 100°C for cleaning). The chamber
141 enables atmospheric simulations under highly stable experimental conditions with low
142 particle wall loss and low contamination levels (more details of the CLOUD chamber can be
143 found in Kirkby et al. (2011) and Duplissy et al. (2016)). At the beginning of the campaign
144 the CLOUD chamber was cleaned by rinsing the walls with ultra-pure water, followed by
145 heating to 100°C and flushing at a high rate with humidified synthetic air and elevated ozone
146 (several ppmv) (Kirkby et al., 2016). This resulted in SO_2 and H_2SO_4 concentrations that
147 were below the detection limit (<15 pptv and $<5 \times 10^4 \text{ cm}^{-3}$, respectively), and total organics
148 (largely comprising high volatility $\text{C}_1\text{--C}_3$ compounds) that were below 150 pptv.

149 The air in the chamber is ionized by galactic cosmic rays (GCR); higher ion generation
150 rates can be induced by a pion beam (π^+) from the CERN Proton Synchrotron enabling
151 controlled simulation of galactic cosmic rays throughout the troposphere. Therefore, the total
152 ion pair production rate in the chamber is between 2 (no beam) and $100 \text{ cm}^{-3} \text{ s}^{-1}$ (maximum
153 available beam intensity, Franchin et al., 2015).

154 Instrumentation

155 The main instruments employed for this study were atmospheric pressure interface time-
156 of-flight (APi-TOF, Aerodyne Research Inc. & Tofwerk AG) mass spectrometers. The
157 instrument has two main parts. The first is the atmospheric pressure interface (APi) where
158 ions are transferred from atmospheric pressure to low pressures via three differentially

159 pumped vacuum stages. Ions are focused and guided by two quadrupoles and ion lenses. The
160 second is the time-of-flight mass analyzer (TOF), where the pressure is approximately 10^{-6}
161 mbar. The sample flow from the chamber was 10 L/min and the core-sampled flow into the
162 APi was 0.8 L/min, with the remaining flow being discarded.

163 There is no direct chemical ionization in front of the instrument. The APi-TOF measures
164 the positive or negative ions and cluster ions as they are present in the ambient atmosphere.
165 As described above, in the CLOUD chamber ions are formed by GCR or deliberately by π^+
166 beam, leading to ion concentrations of a few hundred to thousands per cm^3 , respectively. In
167 our chamber the dominant ionizing species are NH_4^+ and NO_3^- (see below). These ions
168 mainly form clusters with the organic molecules, which is driven by the cluster energies.
169 Therefore, the signals obtained do not provide a quantitative measure of the concentration of
170 the compounds. The higher the cluster energy with certain compounds the higher the ion
171 cluster concentration will be.

172 We calibrated the APi-TOF using trioctylmethylammonium bis (trifluoromethylsulfonyl)
173 imide (TBMA, $\text{C}_{27}\text{H}_{54}\text{F}_6\text{N}_2\text{O}_4\text{S}_2$) to facilitate the exact ion mass determination in both
174 positive and negative ion modes. We employed two calibration methods, the first one by
175 nebulizing TBMA and separating cluster ions with a high-resolution ultra-fine differential
176 mobility analyzer (UDMA) (see Steiner et al. (2014) for more information); the second one
177 by using electrospray ionization of a TBMA solution. The calibration with the electrospray
178 ionization was performed three times, one for each temperature. These calibrations enabled
179 mass/charge (m/z) measurements with high accuracy up to 1500 Th in the positive ion mode
180 and 900 Th in the negative ion mode.

181 Additionally, two peaks in the positive ion mode were identified as contaminants and also
182 used for calibration purposes at the three different temperatures: $\text{C}_{10}\text{H}_{14}\text{OH}^+$ and $\text{C}_{20}\text{H}_{28}\text{O}_2\text{H}^+$.
183 These peaks were present before the addition of ozone in the chamber (therefore being most
184 likely not products of α -pinene ozonolysis) and were also detected by a proton transfer
185 reaction time of flight mass spectrometer (PTR-TOF-MS). Both peaks appeared at the same
186 m/z at all three temperatures. Therefore, based on the calibrations with the UDMA, the
187 electrospray and the two organic calibration peaks, we expect an accurate mass calibration at
188 the three temperatures.

189 **Experimental conditions**

190 ^{2.3}All ambient ion composition data reported here were obtained during nucleation
191 experiments from pure α -pinene ozonolysis. The experiments were conducted under dark
192 conditions, at a relative humidity (RH) of 38% with an O₃ mixing ratio between 33 and 43
193 ppbv (Table 1). The APi-TOF measurements were made under both galactic cosmic ray
194 (GCR) and π^+ beam conditions, with ion-pair concentrations around 700 cm⁻³ and 4000 cm⁻³,
195 respectively.

196

197

198 Table 1. Experiments performed at the CLOUD chamber.

<i>Campaign</i>	<i>Experiment No.</i>	<i>Ionization</i>	<i>α-pinene (pptv)</i>	<i>O₃ (ppbv)</i>	<i>Mass spectrometer polarity</i>	<i>Temperature (°C)</i>
CLOUD 8	1211.02	GCR	258	33.8	Negative	5
CLOUD 10	1710.04	π^+ beam	618	41.5	Positive	5
CLOUD 10	1712.04	π^+ beam	511	40.3	Negative and positive	25
CLOUD 10	1727.04	π^+ beam	312	43.3	Negative and positive	-25

199 ^{2.4.}

200 **Quantum chemical calculations**

201 Quantum chemical calculations were performed on the cluster ion formation from the
202 oxidation products of α -pinene. The Gibbs free energies of formation of representative HOM
203 clusters were calculated using the MO62X functional (Zhao and Truhlar, 2008), and the 6-
204 31+G(d) basis set (Ditchfield, 1971) using the Gaussian09 program (Frisch et al., 2009). This
205 method has been previously applied for clusters containing large organic molecules (Kirkby
206 et al., 2016).

207

208 3. Results and discussion

209 Ion composition

210 3.1. Under ~~relatively~~ dry conditions ($RH = 0\%$) and GCR ionization, the main detected
211 positive ions were N_2H^+ and O_2^+ . With increasing RH up to $\sim 30\%$ we observed the water
212 clusters H_3O^+ , $(H_2O)\cdot H_3O^+$ and $(H_2O)_2\cdot H_3O^+$ as well as NH_4^+ , $C_5H_5NH^+$ (protonated
213 pyridine), Na^+ , and K^+ (Fig. 1A). The concentrations of the precursors of some of the latter
214 ions are expected to be very low: for example, NH_3 mixing ratios were previously found to be
215 in the range of 0.3 pptv (at $-25^\circ C$), 2 pptv (at $5^\circ C$) and 4.3 pptv (at $25^\circ C$) (Kürten et al.,
216 2016). However, in a freshly cleaned chamber we expect ammonia levels below 1 ppt also at
217 the higher temperatures. For the negative ions, NO_3^- was the main detected background
218 signal. Before adding any trace gas to the chamber the signal of HSO_4^- was at a level of 1%
219 of the NO_3^- signal (corresponding to $<5\cdot 10^{-4}$ molecules cm^{-3} , Kirkby et al., 2016), excluding
220 any contribution of sulfuric acid to nucleation in our experiments.

221 After initiating α -pinene ozonolysis, more than 460 different peaks from organic ions
222 were identified in the positive spectrum. The majority of peaks were clustered with NH_4^+ ,
223 while only 10.2 % of the identified peaks were composed of protonated organic molecules. In
224 both cases the organic core was of the type $C_{7-10}H_{10-16}O_{1-10}$ for the monomer region and C_{17-}
225 $_{20}H_{24-32}O_{5-19}$ for the dimer region.

226 In the negative spectrum we identified more than 530 HOMs, of which $\sim 62\%$
227 corresponded to organic clusters with NO_3^- or, to a lesser degree, $HNO_3\cdot NO_3^-$. The rest of
228 the peaks were negatively charged organic molecules. In general, the organic core of the
229 molecules was of the type $C_{7-10}H_{9-16}O_{3-12}$ in the monomer region and $C_{17-20}H_{19-32}O_{10-20}$ in the
230 dimer region. For brevity we refer to the monomer, dimer (and n-mer) as C_{10} , C_{20} and $C_{(10n)}$
231 respectively. Here, the subscript indicates the maximum number of carbon atoms in these
232 molecules, even though the bands include species with slightly fewer carbon atoms.

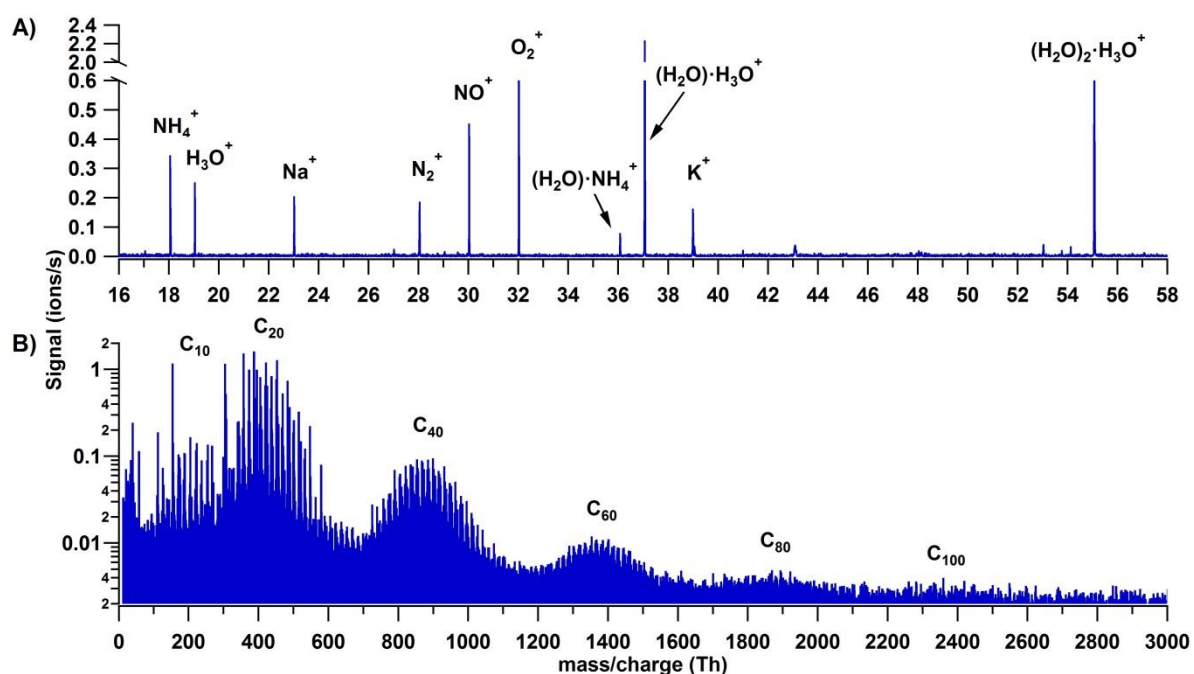
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234 3.1.1. Positive spectrum

235 The positive spectrum is characterized by bands of high intensity at C_{20} intervals, as
236 shown in B. Although we detected the monomer band (C_{10}), its integrated intensity was much

237 lower than the C₂₀ band; furthermore, the trimer and pentamer bands were almost completely
 238 absent. Based on chemical ionization mass spectrometry measurements, Kirkby et al. (2016)
 239 calculated that the HOM molar yield at 5°C was 3.2% for the ozonolysis of α-pinene, with a
 240 fractional yield of 10 to 20% for dimers. A combination reaction of two oxidized peroxy
 241 radicals has been previously reported to explain the rapid formation of dimers resulting in
 242 covalently bound molecules (see Section 3.3). The pronounced dimer signal with NH₄⁺
 243 indicates that (low-volatility) dimers are necessary for positive ion nucleation and initial
 244 growth. We observe growth by dimer steps up to C₈₀ and possibly even C₁₀₀. A cluster of two
 245 dimers, C₄₀, with a mass/charge in the range of ~ 700 - 1100 Th, has a mobility diameter
 246 around 1.5 nm (based on Ehn et al. (2011)).

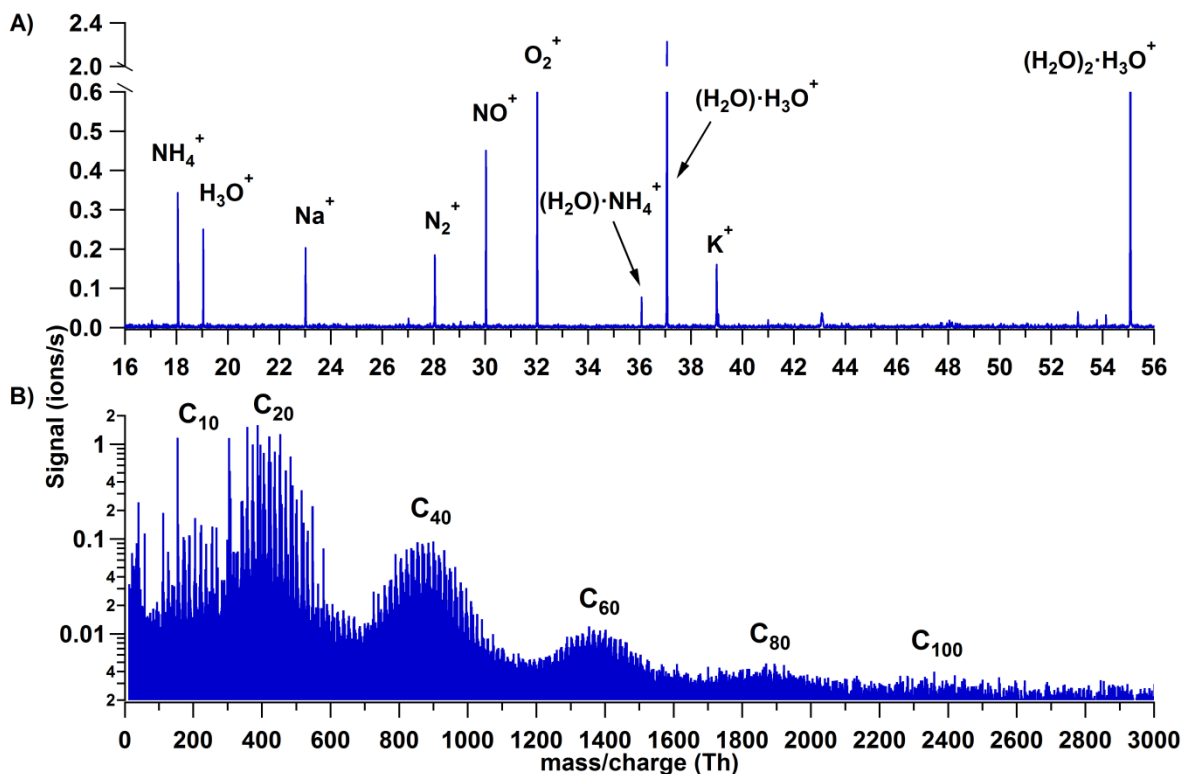
247 Our observation of HOMs-NH₄⁺ clusters implies strong hydrogen bonding between
 248 the two species. This is confirmed by quantum chemical calculations which shall be
 249 discussed in Section 3.3. Although hydrogen bonding could also be expected between HOMs
 250 and H₃O⁺, we do not observe such clusters. This probably arises from the higher proton
 251 affinity of NH₃, 203.6 kcal/mol, compared with H₂O, 164.8 kcal/mol (Hunter and Lias,
 252 1998). Thus, most H₃O⁺ ions in CLOUD will transfer their proton to NH₃ to form NH₄⁺.



253
 254 ~~Figure 1. Positive spectra at 5°C. A) Low mass region, where primary ions from galactic~~
 255 ~~cosmic ray are observed, as well as secondary ions such as NH₄⁺ which are formed by charge~~

256 ~~transfer to contaminants. B) Higher mass region during pure biogenic nucleation, which~~
257 ~~shows broad bands in steps of C₂₀. Most of the peaks represent clusters with NH₄⁺.~~

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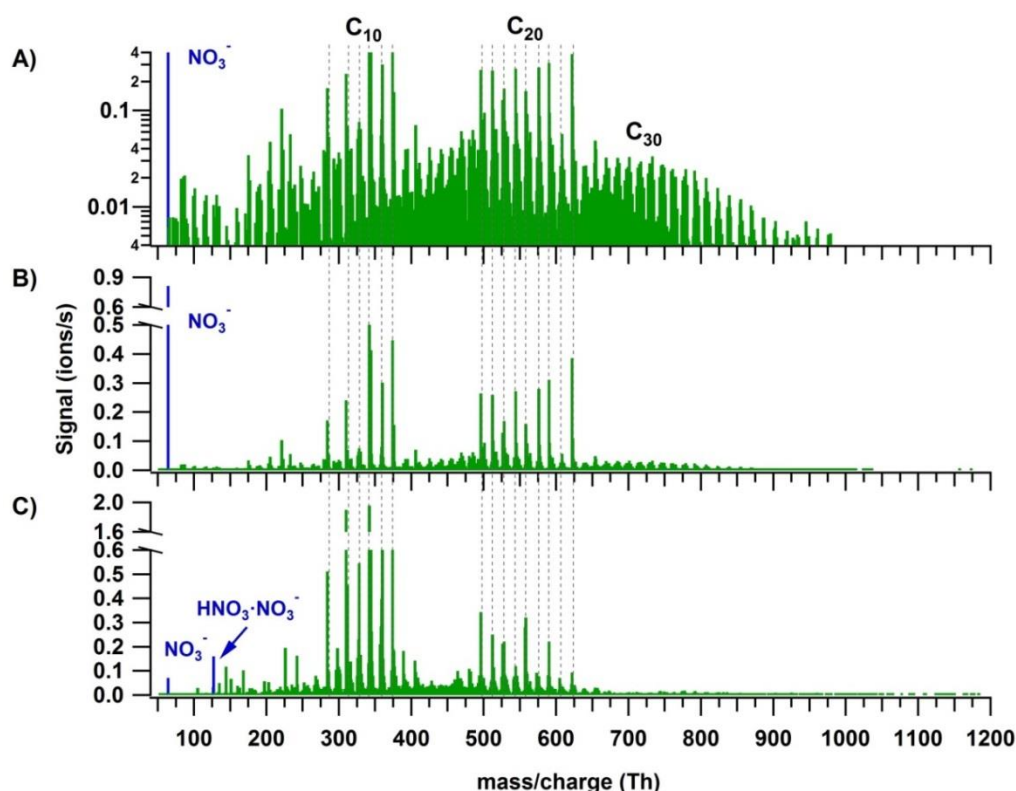
260 Figure 1. Positive spectra at 5°C. A) Low mass region, where primary ions from galactic
261 cosmic rays are observed, as well as secondary ions such as NH₄⁺ which are formed by
262 charge transfer to contaminants. B) Higher mass region during pure biogenic
263 nucleation, which shows broad bands in steps of C₂₀. Most of the peaks represent
264 clusters with NH₄⁺.

265

266 3.1.2. Negative spectrum

267 In the negative spectra, the monomer, dimer and trimer bands are observed during
268 nucleation (Fig. 2). Monomers and dimers have similar signal intensities, whereas the trimer
269 intensity is at least 10 times lower (Fig. 1 A and B). The trimer signal is reduced since it is a
270 cluster of two gas phase species (C₁₀+C₂₀). Additionally, a lower transmission in the APi-
271 TOF may also be a reason for the reduced signal.

272 In Fig. 2, we compare the CLOUD negative-ion spectrum with the one from nocturnal
 273 atmospheric measurements from the boreal forest at Hyytiälä as reported by Ehn et al. (2010).
 274 Panels 2A and 2B show the negative spectrum of α -pinene ozonolysis in the CLOUD
 275 chamber on logarithmic and linear scales, respectively. Panel 2C shows the Hyytiälä
 276 spectrum for comparison. Although the figure shows unit mass resolution, the high resolution
 277 analysis confirms the identical composition for the main peaks: $C_8H_{12}O_7 \cdot NO_3^-$,
 278 $C_{10}H_{14}O_7 \cdot NO_3^-$, $C_{10}H_{14}O_8 \cdot NO_3^-$, $C_{10}H_{14}O_9 \cdot NO_3^-$, $C_{10}H_{16}O_{10} \cdot NO_3^-$ and $C_{10}H_{14}O_{11} \cdot NO_3^-$
 279 (marked in the monomer region), and $C_{19}H_{28}O_{11} \cdot NO_3^-$, $C_{19}H_{28}O_{12} \cdot NO_3^-$, $C_{20}H_{30}O_{12} \cdot NO_3^-$,
 280 $C_{19}H_{28}O_{14} \cdot NO_3^-$, $C_{20}H_{30}O_{14} \cdot NO_3^-$, $C_{20}H_{32}O_{15} \cdot NO_3^-$, $C_{20}H_{30}O_{16} \cdot NO_3^-$, $C_{20}H_{30}O_{17} \cdot NO_3^-$ and
 281 $C_{20}H_{30}O_{18} \cdot NO_3^-$ (marked in the dimer region). The close correspondence in terms of
 282 composition of the main HOMs from the lab and the field both in the monomer and dimer
 283 region indicates a close reproduction of the atmospheric night-time conditions at Hyytiälä by
 284 the CLOUD experiment. In both cases the ion composition was dominated by HOMs
 285 clustered with NO_3^- . However, Ehn et al. (2010) did not report nocturnal nucleation, possibly
 286 because of a higher ambient condensation sink than in the CLOUD chamber.



287

288 | Figure 2. Comparison of the negative ion composition during α -pinene ozonolysis in
 289 CLOUD and during night time in the boreal forest at Hyytiälä (Finland). A) CLOUD

290 spectrum on a logarithmic scale. B) CLOUD spectrum on a linear scale. C) Typical night
291 time spectrum from the boreal forest at Hyytiälä (Finland), adapted from Ehn et al. (2010).

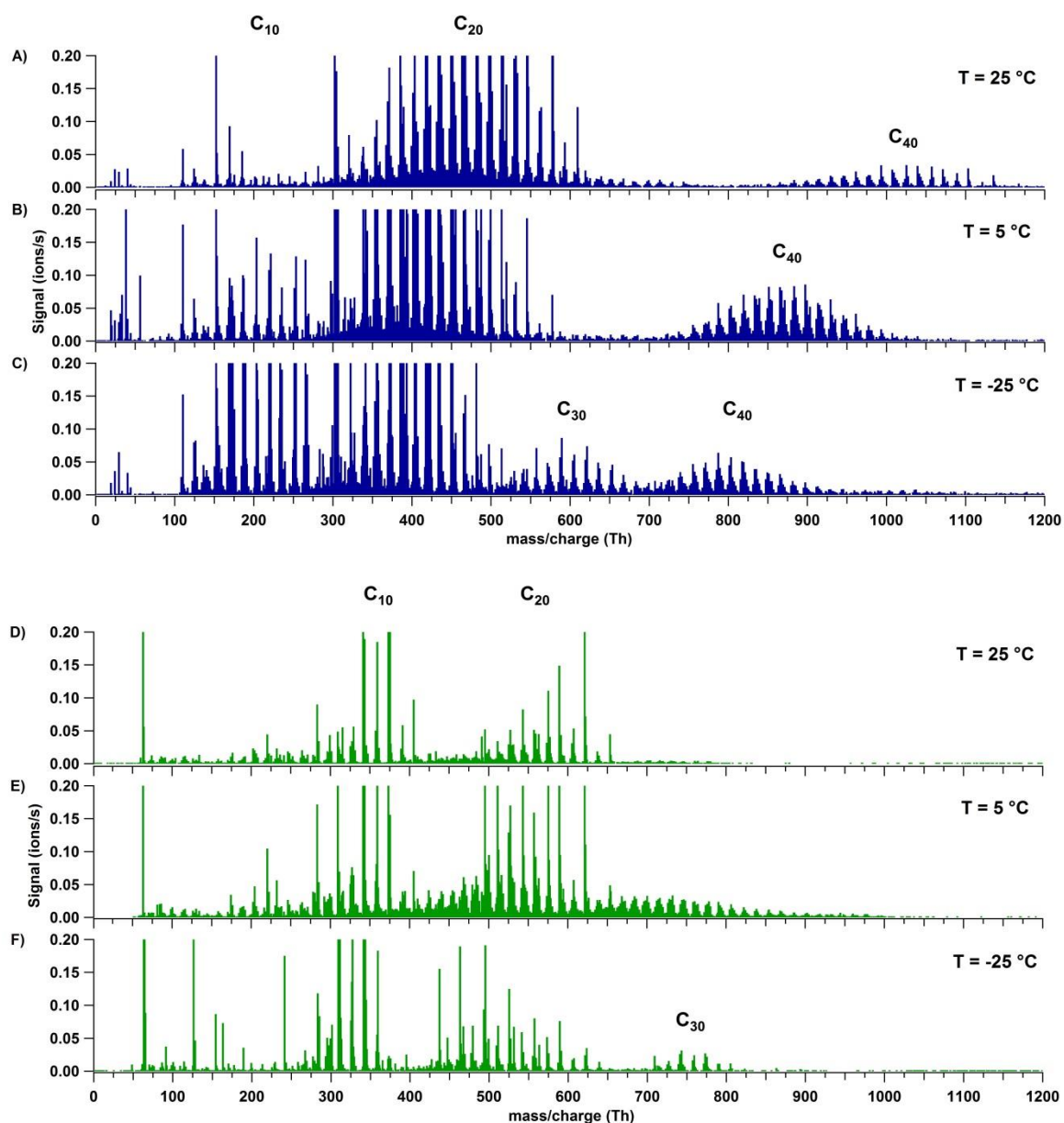
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293 **Temperature dependence**

3.2.

294 Experiments at three different temperatures (25 °C, 5°C and -25 °C) were conducted
295 at similar relative humidity and ozone mixing ratios (Table 1 and Fig. 2). Mass defect plots
296 are shown for the same data in Figure 4. The mass defect is the difference between the exact
297 and the integer mass and is shown on the y-axis versus the mass/charge on the x-axis. Each
298 point represents a distinct atomic composition of a molecule or cluster. Although the
299 observations described in the following are valid for both polarities, the trends at the three
300 temperatures are better seen in the positive mass spectra due to a higher sensitivity at high
301 m/z.

302 The first point to note is the change in the distribution of the signal intensity seen in
303 Figure 3 (height of the peaks) and in Figure 4 (size of the dots) with temperature. In the
304 positive ion mode, the dimer band has the highest intensity at 25 and 5°C (see also Fig. 1B),
305 while at -25°C the intensity of the monomer becomes comparable to that of the dimer. This
306 indicates a reduced rate of dimer formation at -25 °C, or that the intensity of the ion signal
307 depends on both the concentration of the neutral compound and on the stability of the ion
308 cluster. Although the monomer concentration is higher than that of the dimers (Tröstl et al.,
309 2016), the C₂₀ ions are the more stable ion clusters as they can form more easily two
310 hydrogen bonds with NH₄⁺ (see Section 3.3). Thus, positive clusters formed from monomers
311 may not be stable enough at higher temperatures. Moreover, charge transfer to dimers is also
312 favored.



313

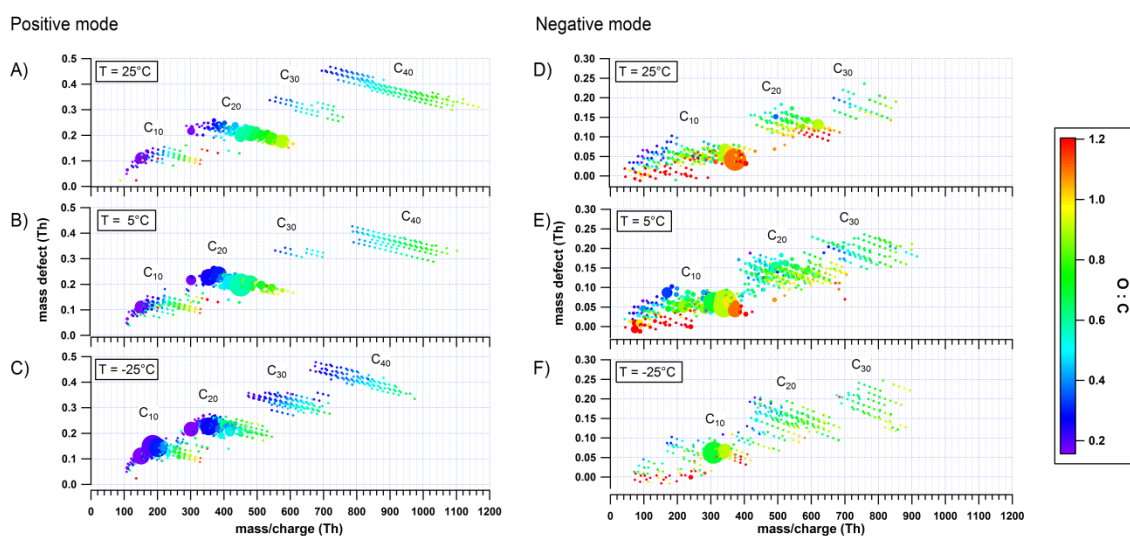
314 | Figure 32. Positive (A-C) and negative (D-F) mass spectra during pure biogenic nucleation
 315 | induced by ozonolysis of α -pinene) at three temperatures: 25 °C (A, D), 5 °C (B, E) and -25 °C
 316 | (C, F). A progressive shift towards a lower oxygen content and lower masses is observed in
 317 | all bands as the temperature decreases. Moreover, the appearance of C₃₀ species can be seen
 318 | in the positive spectrum at the lowest temperature (C).

319

320 The data also show a “shift” in all band distributions towards higher masses with
 321 increasing temperature, denoting a higher concentration of the more highly oxygenated
 322 molecules and the appearance of progressively more oxygenated compounds at higher
 323 temperatures. The shift is even more pronounced in the higher mass bands, as clearly seen in

324 the C₄₀ band of the positive ion mode in Figure 3 (A-C). In this case the combination of two
 325 HOM dimers to a C₄₀ cluster essentially doubles the shift of the band towards higher
 326 mass/charge at higher temperatures compared to the C₂₀ band. Moreover, the width of each
 327 band increases with temperature, as clearly seen in the positive ion mode in Figure 4,
 328 especially for the C₄₀ band. At high temperatures, the production of more highly oxygenated
 329 HOMs seems to increase the possible combinations of clusters, resulting in a wider band
 330 distribution.

331



332

333 ~~Figure 4. Mass-defect plots with the color code denoting the O:C ratio (of the organic core) at~~
 334 ~~25, 5 and -25°C for positive (A-C) and negative ion mode (D-F). A lower O:C ratio is~~
 335 ~~observed in the positive ion mode than in the negative ion mode. The intensity of the main~~
 336 ~~peaks (linearly proportional to the size of the dots) changes with temperature for both~~
 337 ~~polarities due to a lower degree of oxygenation at lower temperature.~~

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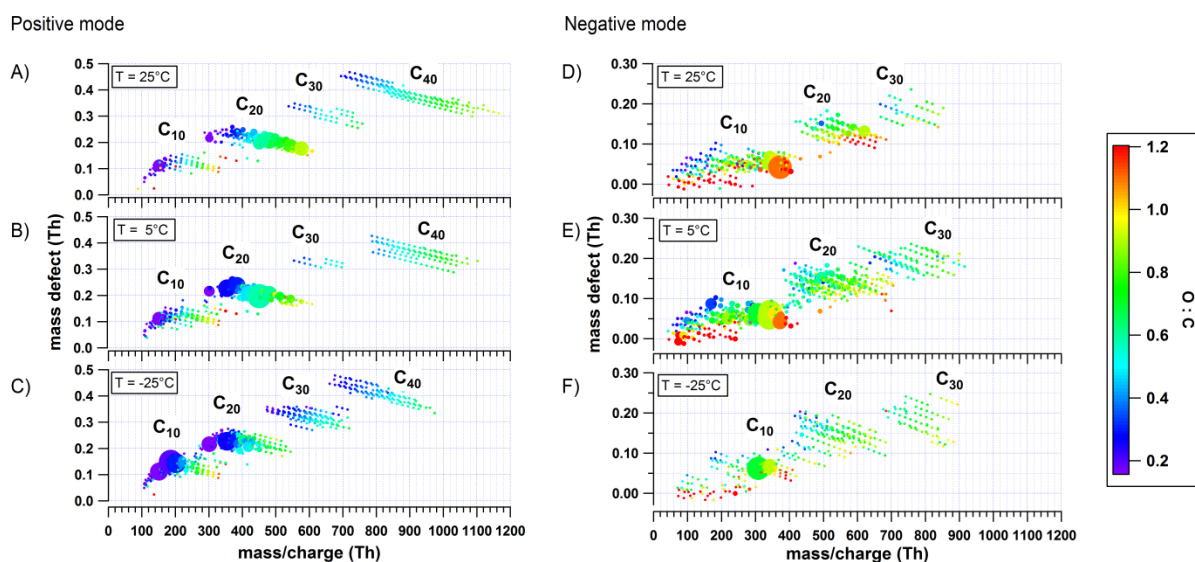
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344
 345 **Figure 4. Mass defect plots with the color code denoting the O:C ratio (of the organic**
 346 **core) at 25, 5 and -25°C for positive (A-C) and negative ion mode (D-F). A lower O:C**
 347 **ratio is observed in the positive ion mode than in the negative ion mode. The intensity of**
 348 **the main peaks (linearly proportional to the size of the dots) changes with temperature**
 349 **for both polarities due to a lower degree of oxygenation at lower temperature.**

350 This trend in the spectra indicates that the unimolecular autoxidation reaction
 351 accelerates at higher temperatures in competition to the bimolecular termination reactions
 352 with HO₂ and RO₂. This is expected. If unimolecular and bimolecular reactions are
 353 competitive, the unimolecular process will have a much higher barrier because the pre-
 354 exponential term for a unimolecular process is a vibrational frequency while the pre-
 355 exponential term for the bimolecular process is at most the bimolecular collision frequency,
 356 which is four orders of magnitude lower. Quantum chemical calculations determine
 357 activation energies between 22.56 and 29.46 kcal/mol for the autoxidation of different RO₂[•]
 358 RO₂ radicals from α-pinene (Rissanen et al., 2015). Thus, such a high barrier will strongly
 359 reduce the autoxidation rate at the low temperatures.

360 The change in the rate of autoxidation is also reflected in the O:C ratio, both in the
 361 positive ion mode (A-C), and the negative ion mode (D-F), showing a clear increase with
 362 increasing temperature. The average O:C ratios (weighted by the peak intensities) are
 363 presented in Table 2 for both polarities and the three temperatures, for all the identified peaks

364 (total) and separately for the monomer and dimer bands. For a temperature change from 25 to
 365 -25°C the O:C ratio decreases for monomers, dimers and total number of peaks. At high
 366 masses (e.g., for the C₃₀ and C₄₀ bands), the O:C ratio may be slightly biased since accurate
 367 identification of the molecules is less straightforward: as an example, C₃₉H₅₆O₂₅·NH₄⁺ has an
 368 exact mass of 942.34 Th (O/C = 0.64), which is very similar to C₄₀H₆₀O₂₄·NH₄⁺ at 942.38 Th
 369 (O/C = 0.60). However, such possible misidentification would not influence the calculated
 370 total O/C by more than 0.05, and the main conclusions presented here remain robust.

371

372 Table 2. Signal weighted average O:C ratios for positive and negative
 373 spectra at 25, 5 and -25 °C.

Temperature (°C)	O/C					
	Positive mode			Negative mode		
	Monomer	Dimer	Total	Monomer	Dimer	Total
25	0.37	0.57	0.54	0.94	0.81	0.90
5	0.34	0.51	0.49	0.88	0.66	0.75
-25	0.31	0.38	0.36	0.79	0.65	0.68

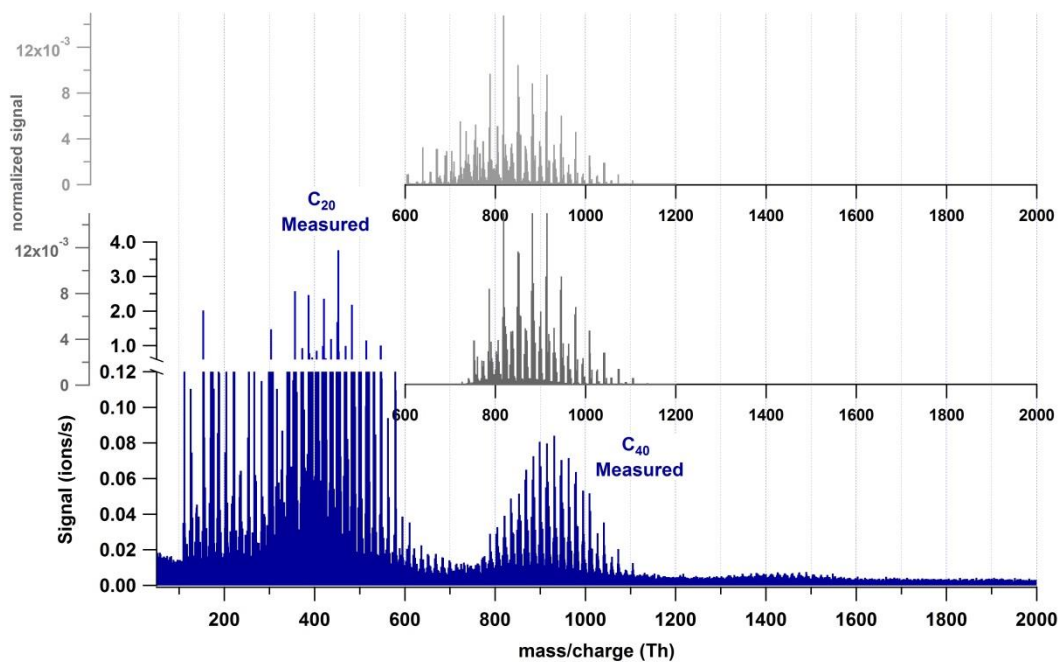
374

375 The O:C ratios are higher for the negative ions than for the positive ions at any of the
 376 three temperatures. Although some of the organic cores are the same in the positive and
 377 negative ion mode, the intensity of the peaks of the most oxygenated species is higher in the
 378 negative spectra. While the measured O:C ratio ranges between 0.4 and 1.2 in the negative
 379 ion mode, it is between 0.1 and 1.2 in the positive ion mode. An O:C ratio of 0.1, which was
 380 detected only in the positive ion mode, corresponds to monomers and dimers with 4–two
 381 oxygen atoms s-or dimers with two oxygen atoms. The presence of molecules with such low
 382 oxygen content was also confirmed with a proton transfer reaction time-of-flight mass
 383 spectrometer (PTR-TOF-MS), at least in the monomer region. Ions with O:C ratio less than
 384 0.3 are probably from the main known oxidation products like pinonaldehyde, pinonic acid,
 385 etc., but also from minor products like pinene oxide and other not identified compounds yet.
 386 It is likely that these molecules, which were detected only in the positive mode, contribute
 387 only to the growth of the newly formed particles (if at all) rather than to nucleation, owing to
 388 their high volatility (Tröstl et al., 2016). In this sense, the positive spectrum could reveal both

389 the molecules that participate in the new particle formation and those that contribute to
390 growth. The differences in the O:C ratios between the two polarities are a result of the
391 affinities of the organic molecules to form clusters either with NO_3^- or NH_4^+ , which, in turn,
392 depends on the molecular structure and the functional groups. Hyttinen et al. (2015) reported
393 the binding energies of selected highly oxygenated products of cyclohexene detected by a
394 nitrate CIMS, finding that the addition of OOH groups to the HOM strengthens the binding of
395 the organic core with NO_3^- . Even when the number of H-bonds between NO_3^- and HOM
396 remains the same, the addition of more oxygen atoms to the organic compound could
397 strengthen the bonding with the NO_3^- ion. Thus, the less oxygenated HOMs were not detected
398 in those experiments, neither in ours, in the negative mode. The binding energies were
399 calculated for the positive mode HOMs- NH_4^+ and are discussed in Section 3.3.

400 We also tested to which extent the formation of the C_{40} band could be reproduced by
401 permutation of the potential C_{20} clusters weighted by the dimer signal intensity. Figure 5
402 shows the measured spectrum (blue) and two types of modeled tetramers: one combining all
403 peaks from the C_{20} band (light gray) and one combining only those peaks with an organic
404 core with $\text{O}/\text{C} \geq 0.4$, i.e. likely non-volatile molecules (dark gray). The better agreement
405 consistency of the latter modeled mass spectrum of the tetramer band with the measured one
406 suggests that only the molecules with $\text{O}/\text{C} \geq 0.4$ are able to form the tetramer cluster. This
407 would mean that C_{20} molecules with 2-7 oxygen atoms are likely not to contribute to the
408 nucleation, but only to the growth of the newly formed particles. One has to note that the
409 comparison of modeled and measured spectrum relies on the assumption that the charge
410 distribution of dimers is also reflected in the tetramers.

411



412

413 | Figure 53. Comparison of the positive ion mode spectrum measured (blue), the C_{40} band
 414 obtained by the combination of all C_{20} molecules (light gray) and the C_{40} band obtained by
 415 combination of only the C_{20} molecules with $O/C \geq 0.4$ (dark gray). The low or absent signals
 416 at the lower masses obtained by permutation suggests that only the highly oxygenated dimers
 417 are able to cluster and form C_{40} .

418

419 These two observations (change in signal distribution and band “shift”) are not only
 420 valid for positive and negative ions, but also for the neutral molecules as observed by two
 421 nitrate chemical-ionization atmospheric-pressure-interface time-of-flight mass spectrometers
 422 (CI-APi-TOF; Aerodyne Research Inc. and Tofwerk AG). This confirms that there is indeed
 423 a change in the HOM composition with different temperature rather than a charge
 424 redistribution effect which would only be observed for the ions (APi-TOF). The detailed
 425 analysis of the neutral molecules detected by these CI-APi-TOFs will be subject of another
 426 paper and is not discussed here.

427 A third distinctive trend in the positive mode spectra at the three temperatures is the
 428 increase in signal intensity of the C_{30} band at -25°C . The increase in the signal of the trimer
 429 | also seems to occur in the negative ion mode when comparing panels D and F in Figure 3.
 430 For this polarity, data from two campaigns were combined (Table 1). To avoid a bias by
 431 possible differences in the APi-TOF settings, we only compare the temperatures from the
 432 same campaign, CLOUD 10, therefore experiments at 25°C and -25°C . The increase in the

433 trimer signal may be due to greater stability of the monomer-dimer clusters or even of three
434 C₁₀ molecules at low temperatures, as further discussed below.

435

436 **Quantum chemical calculations**

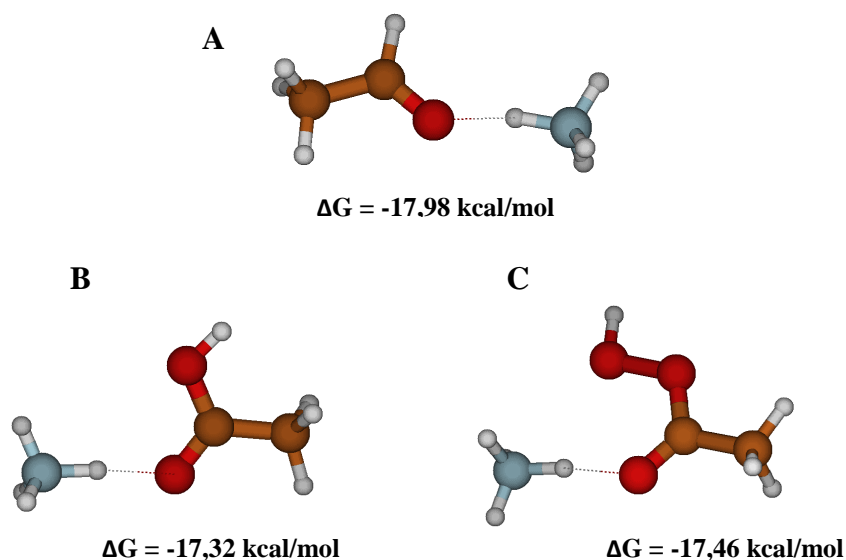
3.3.

437 Three points were addressed in the quantum chemical calculations to elucidate the most
438 likely formation pathway for the first clusters, and its temperature dependence. These
439 included (i) the stability of the organic cores with NO₃⁻ and NH₄⁺ depending on the binding
440 functional group, (ii) the difference between charged and neutral clusters in terms of
441 clustering energies, and finally (iii) the possible nature of clusters in the dimer and trimer
442 region.

443 The calculations showed that among the different functional groups the best interacting
444 groups with NO₃⁻ are in order of importance carboxylic acids (R-C(=O)-OH), hydroxyls (R-
445 OH), peroxy acids (R-C(=O)-O-OH), hydroperoxides (R-O-OH) and carbonyls (R-(R'-)
446 C=O). On the other hand, NH₄⁺ preferably forms a hydrogen bond with the carbonyl group
447 independent of which functional group the carbonyl group is linked to: Figure 6 shows
448 examples of NH₄⁺ clusters with corresponding free energies of formation for carbonyls ($\Delta G =$
449 -17.98 kcal/mol), carboxylic acid ($\Delta G = -17.32$ kcal/mol), and peroxy acid ($\Delta G = -17.46$
450 kcal/mol). For the three examples shown, the interaction of one hydrogen from NH₄⁺ with a
451 C=O group is already very stable with a free energy of cluster ion formation close to -18
452 kcal/mol.

453

454



455

456 | Figure 64. Quantum chemical calculations of the free energy related to the cluster formation
 457 between NH_4^+ and three structurally similar molecules with different functional groups: A)
 458 acetaldehyde, B) acetic acid and C) peracetic acid.

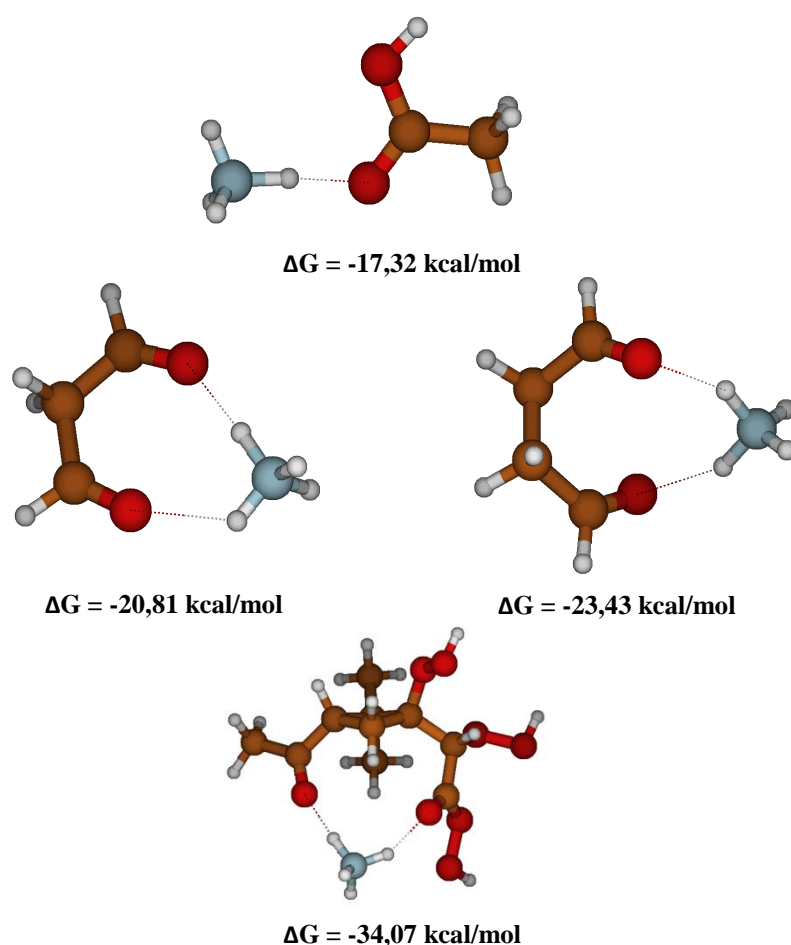
457

458 To evaluate the effect of the presence of a second C=O to the binding of the organic
 459 compound with NH_4^+ , we performed a series of calculations with a set of surrogates
 460 containing two C=O groups separated by a different number of atoms, as shown in Figure 7.
 461 The addition of a second functional group allows the formation of an additional hydrogen-
 462 bond, increasing the stability of the cluster considerably (almost two folds) from about -18
 463 kcal/mol to -34.07 kcal/mol, whereby the position of the second functional group to form an
 464 optimal hydrogen bond (with a 180° angle for N-H-O) strongly influences the stability of the
 465 cluster, as can be seen in Figure 7. Thus, optimal separation and conformational flexibility of
 466 functional groups is needed to enable an effective formation of two hydrogen bonds with
 467 NH_4^+ . This could be an explanation for the observation that the signal intensity is higher for
 468 dimers than for monomers, as dimers can more easily form two optimal hydrogen bonds with
 469 NH_4^+ .

470

471

472



474 | Figure 57. Quantum chemical calculations for different organic molecules with a
 475 carbonyl as the interacting functional group with NH_4^+ . Increasing the interacting
 476 groups from one to two increases the stability of the cluster. The distance between the
 interacting groups also influences the cluster stability.

477 As shown by Kirkby et al. (2016), ions increase the nucleation rates by one to two
 478 orders of magnitudes compared to neutral nucleation. This is expected due to the strong
 479 electrostatic interaction between charged clusters. To understand how the stability difference
 480 relates to the increase in the nucleation rate, the ΔG s of charged and neutral clusters were
 481 compared. For this, $\text{C}_{10}\text{H}_{14}\text{O}_7$ and $\text{C}_{20}\text{H}_{30}\text{O}_{14}$ were selected as representative molecules of the
 482 monomer and dimer region, respectively (Kirkby et al., 2016). Table 3 shows the calculated
 483 free energies of formation (ΔG) of neutral, positive and negative clusters from these C_{10} and
 484 C_{20} molecules at the three temperatures of the experiment. Results show that at 5°C , for

485 example, ΔG of the neutral dimer ($C_{10} + C_{10}$) is -5.76 kcal/mol while it decreases to -20.95
 486 kcal/mol when a neutral and a negative ion form a cluster ($C_{10} + C_{10}^-$). Similarly, trimers
 487 show a substantial increase in stability when they are charged, i.e., from -2.15 kcal/mol to
 488 -19.9 kcal/mol, for the neutral and negative cases, respectively. The reduced values of ΔG for
 489 the charged clusters (positive and negative) indicate a substantial decrease in the evaporation
 490 rate compared to that for neutral clusters, and, therefore, higher stability. Comparing the
 491 NH_4^+ and NO_3^- clusters, the energies of formation for the monomer are -22.5 kcal/mol and
 492 -25.99 kcal/mol, respectively, showing slightly higher stability for the negative cluster.
 493 Inversely, the covalently bound dimer showed greater stability for the positive ion (-30.9
 494 kcal/mol) compared to the negative ion (-25.65 kcal/mol).

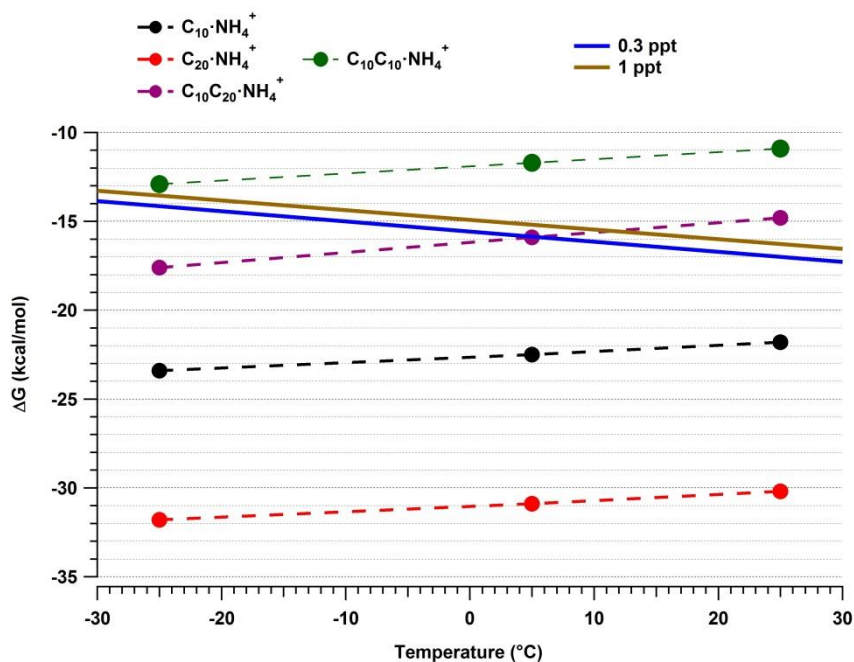
495

496 Table 3. Gibbs free energies of cluster formation ΔG at three different temperatures. ΔG for
 497 the molecules $C_{10}H_{14}O_7$ (C_{10}) and $C_{20}H_{30}O_{14}$ (C_{20}) forming neutral, as well as negative and
 498 positive ion clusters.

	<i>Cluster process</i>	$\Delta G_{-25^\circ C}$ (kcal/mol)	$\Delta G_{5^\circ C}$ (kcal/mol)	$\Delta G_{25^\circ C}$ (kcal/mol)
<i>Neutral</i>	$C_{10} + C_{10}$	-7.3	-5.7	-4.7
	$C_{10} + C_{20}$	-3.2	-2.1	-1.4
<i>Positive</i>	$C_{10} + NH_4^+$	-23.4	-22.5	-21.8
	$C_{20} + NH_4^+$	-31.8	-30.9	-30.2
	$C_{10} + C_{10} \cdot NH_4^+$	-12.9	-11.7	-10.9
	$C_{20} + C_{10} \cdot NH_4^+$	-26.0	-24.3	-23.3
	$C_{10} + C_{20} \cdot NH_4^+$	-17.6	-15.9	-14.8
<i>Negative</i>	$C_{10} + C_{10}^-$	-22.2	-20.9	-20.1
	$C_{20} + C_{10}^-$	-21.3	-19.9	-18.9
	$C_{10} + NO_3^-$	-27.3	-25.9	-25.1
	$C_{20} + NO_3^-$	-26.9	-25.6	-24.7
	$C_{10} + C_{10} \cdot NO_3^-$	-11.3	-10.1	-9.2

499

500 The temperature dependence of cluster formation is shown in Figure 8 for the positive ion
 501 clusters. The blue and brown solid lines represent the needed ΔG for evaporation-collision
 502 equilibrium at 0.3 pptv and 1 pptv HOM mixing ratio, respectively, calculated as described
 503 by Ortega et al. (2012). The markers show the calculated formation enthalpies ΔG for each
 504 of the possible clusters. For all cases, the trend shows an evident decrease in ΔG with
 505 decreasing temperature, with a correspondingly reduced evaporation rate.



506

507 | Figure 68. Quantum chemical calculations of Gibbs free energies for cluster formation at -25,
 508 5 and 25°C. Solid lines represent the required ΔG for equilibrium between evaporation and
 509 collision rates at 0.3 pptv and 1 pptv of the HOM mixing ratio, respectively. Markers show
 510 the ΔG for each cluster (organic core clustered with NH_4^+) at the three temperatures.
 511 $C_{10} \cdot NH_4^+$ (black circles) represent the monomer, $C_{20} \cdot NH_4^+$ (red circles) represent the
 512 covalently bound dimer, $C_{10}C_{10} \cdot NH_4^+$ (green circles) represent the dimer formed by the
 513 clustering of two monomers and $C_{10}C_{20} \cdot NH_4^+$ (purple circles) denote the preferential pathway
 514 for the trimer cluster (see Table 3).

515

516 At all three temperatures, the monomer cluster $C_{10} \cdot NH_4^+$ falls well below the equilibrium
 517 lines, indicating high stability. Even though the difference between -25°C and 25°C is just
 518 -1.6 kcal/mol in free energy, it is enough to produce a substantial difference in the intensity of
 519 the band, increasing the signal at least 8-fold at -25°C (as discussed in Section 3.2). In the
 520 case of the dimers, we consider the possibility of their formation by collision of a monomer
 521 $C_{10} \cdot NH_4^+$ with another C_{10} (resulting in a $C_{10}C_{10} \cdot NH_4^+$ cluster) or the dimer as $C_{20} \cdot NH_4^+$
 522 cluster. The calculations show clearly that the cluster $C_{10}C_{10} \cdot NH_4^+$ is not stable at any of the
 523 three temperatures (green line). In contrast, the covalently-bound C_{20} forms very stable
 524 positive and negative ion clusters (see Table 3). Trimers are mainly observed at lower
 525 temperatures. Since the $C_{10}C_{10} \cdot NH_4^+$ cluster is not very stable, we discard the possibility of a
 526 trimer formation of the type $C_{10}C_{10}C_{10} \cdot NH_4^+$. Thus, the trimer is likely the combination of a
 527 monomer and a covalently-bound dimer ($C_{20}C_{10} \cdot NH_4^+$). According to our calculations (Table

528 3) the preferred evaporation path for this cluster is the loss of C_{10} rather than the evaporation
529 of C_{20} . Therefore, we have chosen to represent only this path in Figure 8. The ΔG of this
530 cluster crosses the evaporation-condensation equilibrium around 5 °C and 14 °C for a HOM
531 mixing ratio of 0.3 pptv and 1 pptv, respectively, in good agreement with the observed signal
532 increase of the trimer at -25°C (Figure 3 A-C). It is important to note that, due to the
533 uncertainty in the calculations, estimated to be ≤ 2 kcal/mol, we do not consider the crossing
534 as an exact reference.

535 The ΔG of the negative ion clusters, which are also presented in Table 3, decrease
536 similarly to the positive ion clusters by around 2 kcal/mol between 25°C and -25°C. The
537 cluster formation energies of the monomer and the dimer with NO_3^- are in agreement with the
538 observed comparable signal intensity in the spectrum (Figure 2Figure) in a similar way as the
539 positive ion clusters. The covalently-bonded dimer ion $C_{20} \cdot NO_3^-$ is also more stable
540 compared to the dimer cluster $C_{10}C_{10} \cdot NO_3^-$, suggesting that the observed composition results
541 from covalently bonded dimers clustering with NO_3^- rather than two individual C_{10} clustering
542 to form a dimer.

543 The formation of a covalently bonded trimer seems unlikely, so the formation of highly
544 oxygenated molecules is restricted to the monomer and dimer region. The trimer could result
545 from the clustering of C_{10} and C_{20} species. Similarly, and based on the C_{20} pattern observed in
546 B, we believe that the formation of the tetramer corresponds to the collision of two dimers.
547 No calculations were done for this case due to the complexity related to the sizes of the
548 molecules, which prevents feasible high level quantum chemical calculations.

549 Finally, ~~a the comparison of between~~ the ΔG values ~~for charged and neutral clusters~~ as
550 presented in Table 3 confirms the expected higher stability of ~~charged ion~~ clusters compared
551 to neutral clusters, decreasing the evaporation rate of the nucleating clusters and enhancing
552 new particle formation.

553

554 4. Conclusions

555 Ions observed during pure biogenic ion-induced nucleation were comprised of mainly
556 organics clustered with NO_3^- and NH_4^+ and to a lesser extent charged organic molecules only

557 or organics clustered with $\text{HNO}_3\text{NO}_3^-$. We found good correspondence between the negative
558 ions measured in CLOUD with those observed in the boreal forest of Hyytiälä. The observed
559 similarity in the composition of the HOMs in the monomer and dimer region during new-
560 particle formation experiments at CLOUD suggests that pure biogenic nucleation might be
561 possible during night time if the condensation sink is sufficiently low, i.e., comparable to that
562 in the CLOUD chamber, where the wall loss rate for H_2SO_4 is $1.8 \cdot 10^{-3} \text{ s}^{-1}$ (Kirkby et al.,
563 2016). The positive mass spectrum showed a distinctive pattern corresponding to progressive
564 addition of dimers (C_{20}), up to cluster sizes in the range of stable small particles.

565 Temperature strongly influenced the composition of the detected molecules in several
566 ways. With increasing temperature, a higher oxygen content (O:C ratio) in the molecules was
567 observed in both the positive and the negative mode. This indicates an increase in the
568 autoxidation rate of peroxy radicals which is in competition with their bimolecular
569 termination reactions with HO_2 and RO_2 .

570 A broader range of organic molecules was found to form clusters with NH_4^+ than with
571 NO_3^- . Quantum chemical calculations using simplified molecules show that NH_4^+ preferably
572 forms a hydrogen bond with a carbonyl group independently of other functional groups
573 nearby. The addition of a second hydrogen bond was found to increase the cluster stability
574 substantially. Thus, the C_{20} -ions are the more stable ion clusters as they can form more easily
575 two hydrogen bonds with NH_4^+ . Although molecules with low oxygen content were
576 measured in the C_{20} band (1 - 4 oxygen atoms), only the molecules with $\text{O/C} \geq 0.4$ seem to be
577 able to combine to form larger clusters.

578 The quantum chemical calculations showed that the covalently-bonded dimer $\text{C}_{20} \cdot$
579 NO_3^- is also more stable than the dimer cluster $\text{C}_{10}\text{C}_{10} \cdot \text{NO}_3^-$, suggesting that the observed
580 composition results from covalently bonded molecules clustering with NO_3^- rather than C_{10}
581 clusters.

582 Temperature affected cluster formation by decreasing evaporation rates at lower
583 temperatures, despite of the lower O:C ratio. In the positive mode a pronounced growth of
584 clusters by addition of C_{20} -HOMs was observed. The formation of a C_{30} -cluster only
585 appeared at the lowest temperature, which was supported by quantum chemical calculations.
586 In the negative mode it appeared as well that the signal of the C_{30} -clusters became stronger

587 with lower temperature. The C₄₀- and higher clusters were probably not seen because of too
588 low sensitivity in this mass range due to the applied instrumental settings. More
589 measurements are needed to determine if the cluster growth of positive and negative ions
590 proceeds in a similar or different way.

591 Nucleation and early growth is driven by the extremely low volatility compounds, i.e.
592 dimers and monomers of high O:C ratios (Tröstl et al., 2016). Here, we observe a reduction
593 of the autoxidation rate leading to oxidation products with lower O:C ratios with decreasing
594 temperature. We expect that this is accompanied by a reduction of nucleation rates. However,
595 a lower temperature reduces evaporation rates of clusters and thereby supports nucleation.
596 The relative magnitude of these compensating effects will be subject of further investigations.

597

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623

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