

Howard et al. Response to Reviewer 1.

We thank the anonymous reviewer for the positive comments and for the very helpful suggestions. These have made many points in the paper much clearer. Please find below the comments and responses, numbered by individual comment. Within each entry we provide a) the reviewer comment, b) our response, and c) changes made to the manuscript.

1.
 - a. The authors declare night time deposition events as a strong evidence for a “multi-hop” mercury transport model. However, they present no evidence for day time reemissions necessary for being able to talk about “hops”.
 - b. This is an important point and we thank the reviewer for highlighting this. We agree that re-emission of GEM following the nocturnal depletion was not covered in the text and have revised it accordingly. Emission of GEM in the morning precedes the break-up of the nocturnal boundary layer (as evidenced by diurnal-component radon), suggesting emission from the surface. As further evidence of emission, in the early morning GEM concentrations overshoot what could be considered the background GEM concentration, observed in the mid-afternoon when boundary layer mixing is at its greatest. It is reasonable to hypothesise that this early morning emission of GEM is partly made up of the relatively volatile fraction of newly-deposited mercury, based on experience in Arctic AMDEs. Further, observed net GEM fluxes during similar depletion events at another Australian site showed that these depletions were not significant long-term sinks of GEM. We have reworded this section to make this clearer.
 - c. Page 10, Line 9: Reworded most of Section 3.2, through to Page 12, Line 16.
2.
 - a. Isn't there a discrepancy between a “multi-hop” model and an atmospheric GEM lifetime of 6 – 12 months mentioned in the “Introduction”?
 - b. We don't believe there is a discrepancy, as the prompt recycling process would not be resolved in the methods used to estimate GEM lifetime. We have reworded this section to define atmospheric lifetime as the mean time before which GEM is permanently removed from the atmosphere. As, under a “multi-hop” model, deposited GEM is rapidly re-emitted, this does not represent a permanent sink and would not affect calculation of mean atmospheric lifetime as determined from mass-balance approaches.
 - c. Page 2, Line 23: Removed “The low atmospheric ... inter-hemispheric transport processes.”
Page 2, Line 20: Inserted “The low atmospheric reactivity and low solubility of the elemental form (GEM) results in low wet/dry deposition rate and scavenging of GEM from the atmosphere. These attributes result in atmospheric transport being the dominant

distribution mechanism through the environment, with long-range transport possible across hemispheric scales. Differences in background atmospheric mercury concentrations between the hemispheres are hence dependent on emission rates, deposition rates, inter-hemispheric transport processes, and atmospheric mercury lifetimes. The atmospheric lifetime is defined here as the mean time after emission that GEM is removed from the atmosphere (Lindberg et al., 2007) and is estimated from mass–balance approaches utilising hemispheric background concentration and source/sink data (e.g. Slemr et al., 1985). The atmospheric lifetime of GEM is currently estimated at 5–12 months (Holmes et al., 2006; Selin et al., 2007; Holmes et al., 2010; Horowitz et al., 2017).”

3. a. Or between a “multi-hop” model and chapter 3.2.1 on long-range transport?
- b. We agree with the reviewer here that the wording used seemed to attempt to replace the model of long-range transport with one of continual deposition and reemission. This was not intended, as differing processes can take place within the planetary boundary layer and free troposphere. As re-emission is difficult to differentiate from emission in GEM flux studies it is possible that evidence for this prompt recycling process has not been clear in the past. Our intention is to highlight, through the evidence of the nocturnal depletion events, that surface deposition/re-emission may be more important to atmospheric mercury cycling than previously believed. We have reworded multiple sections to make this clearer.
- c. Page 10, Line 23: Removed “Such a phenomenon ... over long distances.”
Page 12, Line 12: Inserted “It is important to note that, due to inhibited mixing at the top of the nocturnal boundary layer, the extent of this depletion is limited to within tens to hundreds of metres above the surface. Beyond this, movement of free-tropospheric air continues to enable long-range transport of GEM. Nevertheless, were this phenomenon of rapid, bidirectional exchange with the surface to occur it would have a significant impact on our understanding of atmospheric mercury transport as it can impact the relative importance of intermediate and regional-scale sources, as well as expected time scales for observed decreases in environmental mercury following actions proposed under the Minamata Convention (Lindberg et al., 2007).”
Page 1, Line 16: Removed “These cycles provide ... over long distances.”
Page 1, Line 16: Inserted “These cycles provide strong further evidence supportive of a “multi-hop” model of GEM cycling, characterised by multiple surface depositions and re-emissions, in addition to long-range transport through the atmosphere.”
Page 14, Line 11: Removed “Such a phenomenon ... over long distances.”
Page 15, Line 14: Inserted “Analyses using diurnal-component radon suggest the rapid

increases around sunrise are partly due to volatilisation of newly-deposited mercury, such as seen in other NAMDEs and Arctic AMDEs. The extent of this multi-hop phenomenon may be widespread, which would have a significant impact on our understanding of atmospheric mercury transport, the delivery of atmospheric mercury to the environment, and the legacy of anthropogenic emissions of mercury.”

4. a. Section 2.2: Please state the standard conditions (pressure, temperature) at which the Hg concentrations are reported.
b. This should have been included and was overlooked.
c. Page 5, Line 5: Inserted “Reference volumes are reported at 1 atm and 0 °C.”
5. a. Section 3.1: Are the latitudinal differences statistically significant?
b. The values themselves are, however additional uncertainty as a result of the instrumentation and differences in sampling periods suggests that this is not the case.
c. Page 7, Line 10: Added “These differences are statistically significant (Student's t-test, $p < 0.0001$), though differences in the sampling periods introduces additional uncertainty due to seasonal variation at the sites. Further, an analysis of systematic instrument uncertainty for the Tekran 2537 by Slemr et al. (2015) showed this to be ~10 %.”
6. a. Page 2, line 17: “existence”?
b. The wording here has been altered.
c. Page 2, Line 17: Removed “resulting from its existence ... environmental conditions.”
7. a. Page 4, line 7: Reference “Köppen Aw” is missing in the list of references.
b. “Köppen Aw” was intended as the category type reported by Peel et al., though this was not made clear.
c. Page 4, line 7: Changed reference to “Köppen category Aw, as reported by Peel et al., 2007”.
8. a. Andreae and Merlet (2001) is a review article citing largely work by others. They did not determine the mercury emission factor from biomass burning. Please reword or cite the original work.
b. The use of “determined” was incorrect and unintended.
c. Page 8, Line 5: Changed “determined by Andreae and Merlet (2001)” to “reported by Andreae and Merlet (2001, and references within)”
1. a. Page 8, line 23: a net sink
b. In response to this comment and a similar one from Reviewer 2, we have altered the text to

better describe air–sea GEM exchange.

- c. Page 8, Line 23: Removed “Air-sea exchange of GEM ... from soil and vegetation.”

Page 9, Line 11: Inserted “Air-sea exchange of mercury is complex, with the ocean generally considered a net sink for atmospheric mercury (Mason and Sheu, 2002; Song et al., 2015).

Reduction of mercury within the photolytic zone can give rise to increased concentrations of elemental mercury and hence evasion of GEM to the atmosphere (Soerensen et al., 2014).

Terrestrial surfaces are also commonly sources of GEM; Nelson et al. (2012) modelled terrestrial mercury emission fluxes over Australia that were generally between 8 and 44 $\mu\text{g m}^{-2} \text{a}^{-1}$ from soil and vegetation. Figure 3 does not show a strong difference in concentration distributions between the two source regions.”

9. a. Page 9, line 23: Year of the reference “Howard et al.”?
b. This was an error; the correct reference should have been Howard and Edwards (2017).
c. See 1c.
10. a. Page 11, caption of Fig.4: Median values are usually points – please reword. What is IQR?
b. The shading shows nocturnal periods, hence the wording should have noted that the shading *edges* are sunset/sunrise times. IQR refers to inter-quartile range and was not defined.
c. Changed caption to “Edges of shading denote median sunset/sunrise times for each season. Data have been split into stability categories based on diurnal-component radon quartiles at sunrise (marked in top panels). Lines are median values and error bars indicate inter-quartile ranges.”
11. a. Page 11, line 5: Mean values should be given with their standard deviations or errors and the number of measurements.
b. These have been included.
c. Page 12, Line 30: Included “Mean values for each were $1.08 \pm 0.12 \text{ ng m}^{-3}$ ($n = 3048$), $0.97 \pm 0.13 \text{ ng m}^{-3}$ ($n = 81073$) and $0.90 \pm 0.10 \text{ ng m}^{-3}$ ($n = 46191$), respectively.”
12. a. Figure 5: Vertical line at bottom for NH wet season merges with the blue line. Please make it more distinct.
b. This has been fixed.
c. Altered Figure 5.

Howard et al. Response to Reviewer 2.

We are grateful to the anonymous reviewer for their thoughtful comments and for their positive comments regarding the manuscript. These comments have helped to sharpen a number of points within. Please find below the comments and responses, numbered by individual comment. Within each entry we provide a) the reviewer comment, b) our response, and c) changes made to the manuscript.

1.
 - a. Introduction, first paragraph: it should be made clear that the mercury emissions here refer to those emitted into the atmosphere
 - b. This has been fixed accordingly.
 - c. Page 2, Line 5: Inserted: "to the atmosphere"
2.
 - a. Introduction: Soerensen et al. 2014 ES&T Elemental Mercury Concentrations and Fluxes in the Tropical Atmosphere and Ocean analyzed GEM fluxes in the tropical ocean and is a relevant reference for the third paragraph
 - b. Included this reference in the introduction and in Section 3.1 (see 5c).
 - c. Page 3, Line 8: Inserted "Soerensen et al., 2014."
3.
 - a. Page 4, line 10: "relative" humidity
 - b. This has been changed.
 - c. Page 4, Line 11: Inserted "relative".
4.
 - a. Page 6, Section 3.1: Slemr et al. 2015 ACP Comparison of mercury concentrations measured at several sites in the Southern Hemisphere suggested that the systematic uncertainty among the GEM measurements at different stations is about 0.1 ng/m³. I am not sure whether the 9% difference between the ATARS station and Cape Grim really indicates a latitudinal gradient.
 - b. This is an important point raised by the reviewer. Although the values are statistically different, there are other uncertainties that bring into question this significance, including the systematic instrumental uncertainty raised by the reviewer. We have altered the text to highlight these uncertainties.
 - c. Page 7, Line 10: Added "These differences are statistically significant (Student's t-test, $p < 0.0001$), though differences in the sampling periods introduces additional uncertainty due to seasonal variation at the sites. Further, an analysis of systematic instrument uncertainty for the Tekran 2537 by Slemr et al. (2015) showed this to be ~10 %."

5.
 - a. Page 8, lines 22-24: The Ocean is commonly a sink of total atmospheric mercury, but is commonly a net source of GEM (the measured mercury species).
 - b. We thank the reviewer for pointing out this distinction and have adjusted the text accordingly.
 - c. Page 8, Line 23: Removed "Air-sea exchange of GEM ... from soil and vegetation."
Page 9, Line 11: Inserted "Air-sea exchange of mercury is complex, with the ocean generally considered a net sink for atmospheric mercury (Mason and Sheu, 2002; Song et al., 2015). Reduction of mercury within the photolytic zone can give rise to increased concentrations of elemental mercury and hence evasion of GEM to the atmosphere (Soerensen et al., 2014). Terrestrial surfaces are also commonly sources of GEM; Nelson et al. (2012) modelled terrestrial mercury emission fluxes over Australia that were generally between 8 and 44 $\mu\text{g m}^{-2} \text{a}^{-1}$ from soil and vegetation. Figure 3 does not show a strong difference in concentration distributions between the two source regions."
6.
 - a. Page 9, line 23, "Howard et al." typo
 - b. The typo has been edited.
 - c. Page 10, Line 4: Changed "Howard et al." to "Howard and Edwards (2017)".
7.
 - a. As this station also has data of other chemical species such as aerosol and ozone, including these data would highly facilitate the data analysis (for example on the effect of biomass burning on GEM).
 - b. The inclusion of such data would quite likely facilitate additional analyses, however they are not currently evaluated to the standard required for publication. When these data become available, future analyses will be undertaken.
 - c. No changes made.

Atmospheric mercury in the southern hemisphere tropics: seasonal and diurnal variations and influence of inter-hemispheric transport

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Abstract. Mercury is a toxic element of serious concern for human and environmental health. Understanding its natural cycling in the environment is an important goal towards assessing its impacts and the effectiveness of mitigation strategies. Due to the unique chemical and physical properties of mercury, the atmosphere is the dominant transport pathway for this heavy metal, with the consequence that regions far removed from sources can be impacted. However, there exists a dearth of long-term monitoring of atmospheric mercury, particularly in the tropics and southern hemisphere. This paper presents the first two years of gaseous elemental mercury (GEM) measurements taken at the Australian Tropical Atmospheric Research Station (ATARS) in northern Australia, as part of the Global Mercury Observation System (GMOS). Annual mean GEM concentrations determined at ATARS ($0.95 \pm 0.12 \text{ ng m}^{-3}$) are consistent with recent observations at other sites in the southern hemisphere. Comparison with GEM data from other Australian monitoring sites suggests a concentration gradient that decreases with increasing latitude. Seasonal analysis shows that GEM concentrations at ATARS are significantly lower in the distinct wet monsoon season than in the dry season. This result provides insight into alterations of natural mercury cycling processes as a result of changes in atmospheric humidity, oceanic/terrestrial fetch and convective mixing, and invites future investigation using wet mercury deposition measurements. Due to its location relative to the atmospheric equator, ATARS intermittently samples air originating from the northern hemisphere, allowing an opportunity to gain greater understanding of inter-hemispheric transport of mercury and other atmospheric species. Diurnal cycles of GEM at ATARS show distinct nocturnal depletion events that are attributed to dry deposition under stable boundary layer conditions. These cycles provide strong further evidence supportive of a “multi-hop” model of GEM cycling, characterised by multiple surface depositions and re-emissions, in addition to long-range transport through the atmosphere.

1 Introduction

Mercury (Hg) is a toxic element that has natural and anthropogenic sources, sinks and cycles within the environment. Human activities such as gold mining and biomass/fossil fuel combustion have perturbed the natural cycling of mercury through the addition of mercury emissions, which are re-deposited from the atmosphere to land, vegetation and water bodies. It is estimated that currently anthropogenic emissions to the atmosphere increase the global atmospheric mercury pool by 1960 tonnes annually, a value that represents 30 % of estimated mercury emissions, with the remainder emitted from natural geological sources (10 %) or re-emitted from stores of previously-deposited mercury (60 %). These mercury emission estimates are subject to large uncertainties (AMAP/UNEP, 2013; UNEP, 2013). That anthropogenic mercury sources now exceed those from natural sources on a global scale is of concern for both human and environmental health. Evidence suggests these additional sources are leading to increased concentrations of mercury in the oceans and in marine animals, with the consequence that bioaccumulation of toxic methylmercury within aquatic food chains has also increased (Mason et al., 2012; UNEP, 2013). There exists a significant pathway for methylmercury transfer to humans, as it is estimated that more than 100 million tonnes of fish are eaten world-wide each year and fish provide two and a half billion people with at least 20 % of their protein intake. Mercury in this latter form can seriously threaten human health through impacts on the development of foetuses and young children. In response to this threat, the United Nations Environment Programme (UNEP) has developed the Minamata Convention on Mercury, which is expected to be ratified in 2017.

The global cycling of mercury is unique amongst metals, as within Earth's atmosphere 90 to 99 % of mercury is found as gaseous elemental mercury (GEM), with the remaining portion composed of operationally-defined gaseous oxidised mercury (GOM) and particulate-bound mercury (PBM) — collectively known as reactive mercury (RM) (Gustin et al., 2013). The low atmospheric reactivity and low solubility of the elemental form (GEM) results in low wet/dry deposition rates and scavenging of GEM from the atmosphere. These attributes result in atmospheric transport being the dominant distribution mechanism through the environment, with long-range transport possible across hemispheric scales. Differences in background atmospheric mercury concentrations between the hemispheres are hence dependent on emission rates, deposition rates, inter-hemispheric transport processes, and atmospheric mercury lifetimes. The atmospheric lifetime is defined here as the mean time after emission that GEM is removed from the atmosphere (Lindberg et al., 2007) and is estimated from mass-balance approaches utilising hemispheric background concentration and source/sink data (e.g. Slemr et al., 1985). The atmospheric lifetime of GEM is currently estimated at 5–12 months (Holmes et al., 2006; Selin et al., 2007; Holmes et al., 2010; Horowitz et al., 2017).

With 68 % of the Earth's landmass and 88 % of the human population in the northern hemisphere, both natural and anthropogenic emissions of mercury are disproportionately distributed between the hemispheres. Towards the equator, the existence of the intertropical convergence zone (ITCZ) and the associated upward/poleward movement of the Hadley circulation leads to reduced tropospheric mixing across the atmospheric, or chemical equator (Bowman and Cohen, 1997; Hamilton et al., 2008; Holmes and Prather, in press) and hence a broad, hemispheric gradient of GEM concentrations (Slemr et al., 1985; Sprovieri et al., 2016). Stationary observations of GEM within the tropics are rare but those that are available report significant changes

in concentration as source regions shift across hemispheres with the drift of the atmospheric equator (Müller et al., 2012; Wang et al., 2014). The tropics also represent an important region for mercury cycling as they are home to around 40 % of the world's population, including over 50 % of people under the age of 15, a group at greater risk of adverse effects due to mercury exposure during early development (Bose-O'Reilly et al., 2010). Furthermore, this region hosts several large coastal communities within emerging and developing economies, in which environmental controls and advisories are not always well developed (Costa et al., 2012).

Characterisation of background GEM in the tropics and southern hemisphere (SH) has been hindered by a lack of observations and is based largely on intermittent ship voyages (Soerensen et al., 2012, 2014), along with a few long-term stationary records in South America, Africa, Antarctica, and islands in the Indian and eastern Pacific oceans (Sheu et al., 2010; Müller et al., 2012; Wang et al., 2014; Angot et al., 2014; Slemr et al., 2015; Angot et al., 2016). A recent comparison of interannual records from four mercury monitoring stations spanning a latitude range of 34° S to 72° S, of which the longest-running spans 7 years, suggests that background GEM concentrations in the southern hemisphere are between 0.85 and 1.05 ng m⁻³ (Slemr et al., 2015). Previous measurements of atmospheric mercury concentrations have also been reviewed by Sprovieri et al. (2010, 2016). The Australian continent, with its large non-Antarctic southern hemisphere landmass (22 %), a latitudinal distribution (11–44° S) spanning diverse climatic zones, and a mercury emission profile characterised by anthropogenic sources that are significantly smaller than natural and re-emitted sources (Nelson et al., 2012), presents unique opportunities for extending environmental mercury monitoring in a region that has largely been under-represented.

Initiated under the Global Mercury Observation System (GMOS) and considered for inclusion with the Asia Pacific Mercury Monitoring Network (APMMN), measurements of GEM are being undertaken at the Australian Tropical Atmospheric Research Station (ATARS), north-east of Darwin in Australia's Northern Territory. Of the six GMOS sites classed as tropical, ATARS is the southernmost and one of only two (along with Kodaikanal; 10.2314° N, 77.4652° E) situated in the eastern hemisphere. This site is therefore important in bridging the spatial gap in GEM measurements in equatorial regions around the globe. Originally an experimental radar site, ATARS was expanded in 2010 to incorporate greenhouse gas measurements as part of the Australian Greenhouse Gas Observation Network (Ziehn et al., 2016) and is operated jointly by the Australian Bureau of Meteorology (BoM) and the Commonwealth Science and Industrial Research Organisation (CSIRO). The Australian Nuclear Science and Technology Organisation (ANSTO) began continuous atmospheric radon measurements at the site in 2012 to aid in the determination of terrestrial influence on observed air masses (Chambers et al., 2016b). In June 2014, an additional expansion took place and now continuous aerosol, reactive gas (O₃, NO_x) and GEM measurements complement the suite of atmospheric measurements at the site (Mallet et al., 2016). This GEM dataset represents the first multi-year time series of atmospheric mercury monitoring in tropical Australia.

We present here the first two years of tropical GEM measurements from ATARS, examine their seasonal and diurnal variations, and evaluate the contribution of air masses transported from the northern hemisphere to the observed concentrations. These results add substantial new information to our understanding of mercury in the southern hemisphere and tropical atmosphere.

2 Methods

2.1 Site description

ATARS is situated on the Gunn Point peninsula (12.2491° S, 131.0447° E; Fig. 1), approximately 20 km north-east from the suburban edge of Darwin (2013 population 136,200; ABS, 2015) in Australia's Northern Territory. Between 2 and 9 km to the north and west of ATARS lies the edge of the peninsula that gives way to the Tiwi Islands and Timor Sea, whilst the land to the east and south is largely uninhabited and includes national parks and conservation areas.

The climate in the region is best described as tropical (Köppen category Aw, as reported by Peel et al., 2007) with mean monthly maximum temperatures between 30 °C and 33 °C (1941 to 2016 means; BoM, 2016) and a distinct monsoon (wet) season that coincides generally with the austral summer (December–February). The build-up to these monsoon seasons is characterised by steadily increasing minimum temperatures (19 °C in July to 25 °C in December) and associated increases in relative humidity (daily ranges of 37–60 % in July to 72–83 % in February). Mean annual rainfall is 1728 mm, with an average of 1604 mm (>90 %) of this falling in the period November–April. As the site is located on a peninsula, a sea/land breeze cycle is often experienced in the dry season, resulting in mostly south-easterly winds throughout the morning, tending northerly as the sea breeze circulation sets in from the nearby coast. In the wet season, shifting synoptic patterns result in an increased frequency of westerly winds.

The vegetation classification is savannah with coarse grasses and scattered tree growth immediately surrounding the site. Burning of the grassed areas occurs frequently, with a fire return interval of 1–2 years. Direct mercury analysis (see Edwards and Howard, 2013, for methodology) of vegetation within 500 m of the station gave total mercury concentrations of $7.23 \pm 0.37 \mu\text{g kg}^{-1}$ ($n = 18$) for grass and $21.09 \pm 3.79 \mu\text{g kg}^{-1}$ ($n = 9$) for tree litter. Sampling of soils in the same locations gave total mercury concentrations of $9.14 \pm 0.58 \mu\text{g kg}^{-1}$ ($n = 18$) in grassed areas and $26.49 \pm 3.31 \mu\text{g kg}^{-1}$ ($n = 9$) under forest canopy, confirming that soils in the area are categorised as background ($< 100 \mu\text{g kg}^{-1}$; Gustin et al., 2006). Sampling was undertaken in the early dry season, approximately 10–12 months after the last grass fire.

Anthropogenic emissions of mercury and its compounds to the atmosphere in and around Darwin are generally quite low. Australian National Pollutant Inventory (NPI) data for 2014–15 state that 6 sites situated between 20 km and 40 km from ATARS in the direction of Darwin (wind directions 190° to 240°) emitted a total of 0.12 kg Hg to the atmosphere (NPI, 2016). Other distributed anthropogenic mercury emissions in Darwin are estimated at less than 0.2 kg a^{-1} , based on 25 km x 25 km gridded population data (Nelson et al., 2012).

2.2 Measurements

Continuous (5-minute sample) GEM measurements were obtained using a Tekran 2537X Automated Ambient Air Analyser (2537X). This instrument is housed in an air-conditioned structure with internal temperature set at 25 °C. Air is sampled from a 10 m high tower through 7.95 mm I.D. perfluoroalkoxy tubing using a Thomas 2688 vacuum pump drawing approximately 50 l min^{-1} (residence time 0.6 s). The 2537X subsamples from this flow at 1 l min^{-1} through 6 m of heated polytetrafluoroethylene (PTFE) line maintained at 50 °C, and two 0.2 μm PTFE filters positioned before and after the heated line. The



Figure 1. Map of region surrounding ATARS. Composed in QGIS using Natural Earth dataset.

2537X operates on the principle of cold vapour atomic fluorescence spectroscopy (CVAFS) following gold amalgamation pre-concentration (see for example Ebinghaus et al., 1999; Munthe et al., 2001). This technique quantifies total gaseous mercury (TGM = GEM + GOM), however experience from other researchers suggests that the fraction of GOM in the atmosphere is generally small and removed upstream of the 2537X. As such we present the results here as GEM and not TGM, in line with reporting standards employed by other GMOS secondary sites (Sprovieri et al., 2016). Reference volumes are reported at 1 atm and 0 °C.

Quality assurance and quality control procedures were applied as per protocols derived for GMOS sites (Sprovieri et al., 2016). Calibration of the 2537X took place every 23 hours using an internal mercury permeation source maintained at 50 °C. Primary calibration of this source took place twice each year using manual injections of mercury vapour. No change in the internal source permeation rate was detected over this period. Furthermore, standard additions of mercury are automatically introduced to the 2537X from the internal permeation source every 35 samples (~3 hours) in order to verify GEM recovery performance.

Continuous hourly measurements of radon were sampled at 12 m using an ANSTO-designed and -built, 700 l dual-flow-loop two-filter radon detector (Whittlestone and Zahorowski, 1998; Chambers et al., 2011). This detector samples at 40 l min⁻¹ through 25 mm high-density polyethylene agricultural pipe and has a lower limit of detection of 40–50 mBq m⁻³. Calibrations are performed monthly by injecting radon from a 101.15 ± 4 % kBq ²²⁶Ra source (delivering 12.745 Bq ²²²Rn min⁻¹), traceable to NIST standards. Instrumental background is checked every 3 months. Radon measurements were corrected for

the response time of the instrument (Griffiths et al., 2016), although the main trends were not affected by this time correction. Time-corrected radon data were then split into “fetch” and “diurnal” components by interpolating between minimum afternoon (12:00 to 17:00) values when atmospheric mixing is greatest and subtracting these interpolated values (fetch component) from the original signal, leaving the diurnal component (see Chambers et al., 2016a, for details).

5 Meteorological measurements are collected at ATARS using a standard automated weather station (AWS) operated by the Australian Bureau of Meteorology. Precipitation data were collected using a 203 mm tipping bucket rain gauge and daily totals were summed to give cumulative season totals centred around a hydrologic year beginning 1st June. The temporal extents of what we define here as “wet seasons” were then determined using the method of Smith et al. (2008), whereby 15 % and 85 % of the total cumulative rainfall marked their onset and conclusion, respectively. The wet season of 2014–15 was further extended
10 to include two 100+ mm rain events that took place in November and March.

2.3 Modelling

As the atmospheric equator changes its position relative to the geographic equator, we employed a system of passive tracers within the GEOS-Chem chemical transport model to help assess the impact of air originating from the northern hemisphere (NH) on the site, based on the work of Holmes and Prather (in press). We use GEOS-Chem v10-01 driven by assimilated
15 meteorology from the NASA Goddard Earth Observing System Forward Processing (GEOS-FP) data product, run at $2^\circ \times 2.5^\circ$ horizontal resolution and 47 vertical levels from the surface to 0.01 hPa. Tracers with 90-day lifetimes were uniformly released from the surface in all model boxes poleward of 45° latitude within each hemisphere. The atmospheric equator is then defined as the point where mixing ratios of tracers from the two hemispheres are equal. Tracer concentrations in surface air over ATARS were saved as daily mean values in the model grid box containing the site (2° latitude by 2.5° longitude and an
20 approximate atmospheric depth of 130 m). Increasing the number of grid squares over which tracer values were averaged did not significantly affect the results.

The NOAA Hybrid Single Particle Lagrangian Integrated Trajectory (HYSPLIT) Model (Draxler, 1999; Draxler and Hess, 1998; Stein et al., 2015) was also employed to assess influences of air mass source regions. Global Data Assimilation System (GDAS) 0.5° meteorological reanalysis data were used to drive the model, and trajectories were initialised at 0.5 times the
25 mixed layer height as determined by HYSPLIT. To reduce the influence of local daily variation in GEM concentrations on this analysis, back trajectories were calculated for each hour of the day rather than as a daily or part-daily mean. For each trajectory, air parcel coordinates were calculated every two hours and weighted per the corresponding GEM concentration. These weighted values were then averaged over $0.5^\circ \times 0.5^\circ$ grid cells.

3 Results and discussion

3.1 Overall means and seasonal trends

Measurements of GEM at ATARS began on 5th June, 2014 and were still ongoing at the time of writing. Instrument maintenance/downtime plus application of QC protocols, including calibration and standard additions, resulted in 68.1 % temporal measurement coverage during the first two years of operation (Fig. 2, Table 1). Concentrations are normally distributed across this period with an overall mean of $0.95 \pm 0.12 \text{ ng m}^{-3}$ ($n = 130312$), which is within the range of long-term background GEM concentrations for the southern hemisphere as reported by Slemr et al. (2015). Mean GEM concentrations reported by Slemr et al. (2015) over 2012–13 at Cape Grim, Tasmania (40.6832° S , 144.6899° E) and by Morrison et al. (2015) over 2014–15 at Singleton, NSW (32.4777° S , 151.1018° E) were both 0.86 ng m^{-3} (9 % lower), suggesting a slight latitudinal gradient in GEM across the Australian continent. These differences are statistically significant (Student's t-test, $p < 0.0001$), though differences in the sampling periods introduces additional uncertainty due to seasonal variation at the sites. Further, an analysis of systematic instrument uncertainty for the Tekran 2537 by Slemr et al. (2015) showed this to be $\sim 10\%$. A latitudinal gradient within the southern hemisphere was more generally seen in median annual GEM concentrations for GMOS sites in 2013–14, based on data from 5 sites (Sprovieri et al., 2016). GEM measurements at ATARS were coincident with those reported by Sprovieri et al. (2016) for only the latter 6 months of 2014, a period spanning the late dry season and early wet season. Concentrations during this period were $1.02 \pm 0.10 \text{ ng m}^{-3}$ — higher than the overall mean at ATARS, though still lower than mean values reported for other tropical GMOS sites.

A seasonal trend is apparent in the GEM time series (Fig. 2), which shows higher concentrations during the dry season compared to the wet. Wind sector analysis also shows distinctly different wind patterns between wet and dry seasons (Fig. 3). During the wet season, $\sim 60\%$ of winds come to the site from a westerly direction, consistent with shifting of the ITCZ and associated low pressure systems towards northern Australia. In the dry season, south-easterly to north-easterly winds are more common ($\sim 65\%$ between 30° and 150°), although there is also a notable westerly element. Concentration distributions vary between seasons, with a larger fraction of values above 1 ng m^{-3} seen in the dry period. Within each season however, these distributions do not change significantly with wind direction. Furthermore, the small percentage of winds arriving from the south-west show no change in GEM distribution, implying that the low mercury emissions from Darwin are not significantly impacting measurements and that overall trends are indicative of influences from the global atmospheric mercury pool rather than local sources.

Figure 2 shows that the highest GEM values are concentrated into short peaks, clustered more heavily around the mid- to late-dry season. In the absence of local anthropogenic sources, this is considered consistent with biomass burning events and the associated release of mercury from volatilisation and thermal desorption from vegetation and soils (Melendez-Perez et al., 2014). These biomass burning events occur extensively in Northern Australia throughout the dry season as the result of natural and accidental lighting, as well as part of local land management practices (Russel-Smith et al., 2007). Nelson et al. (2009) concluded that burning in the northernmost part of Australia can contribute up to around $2 \text{ kg Hg km}^{-2} \text{ a}^{-1}$ to the atmosphere (2006 data, $25 \text{ km} \times 25 \text{ km}$ grid resolution).

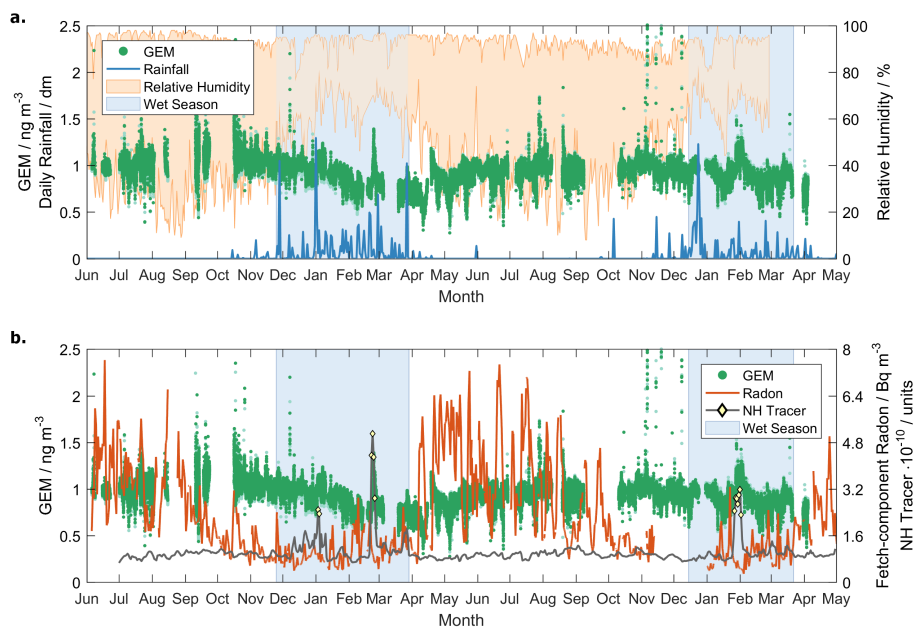


Figure 2. a. 5-minute GEM data, daily rainfall and daily min/max relative humidity values plus wet season ranges as defined by Smith et al. (2008). b. 5-minute GEM, hourly fetch-component radon and daily NH tracers. Days defined as NH-influenced are marked with diamonds.

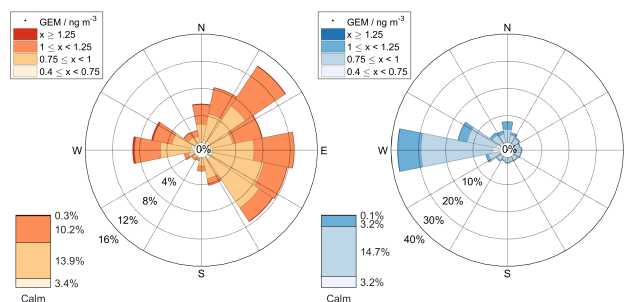


Figure 3. Directional GEM concentration distributions for (left) dry season and (right) all wet season half-hourly GEM data.

An intensive study of these biomass burning events undertaken at ATARS during the early dry season in 2014 also confirmed spikes in GEM concentration that were associated with biomass burning (Mallet et al., 2016; Desservettaz et al., 2017). The distance to the fire and atmospheric dispersion, as well as vegetation type and associated mercury loading, were all identified as factors influencing the strength of these biomass burning signals. Desservettaz et al. (2017) calculated emission factors for GEM between 0.0035 and 0.032 g Hg per kg dry fuel, around 2 orders of magnitude higher than that reported by Andreae and Merlet (2001, and references within) for savannah grasslands. The fires observed by Desservettaz et al. were shown to be from scrubland fires rather than grassland fires, excluding the possibility of direct comparison between the two results. With a full

suite of greenhouse gas and aerosol measurements taking place at ATARS, further identification of smoke plumes and precise calculation of emission factors is possible in a manner that is comparable with previous studies.

Wet season GEM concentrations in 2014–15 were characterised by a steady, gradual decrease that reversed abruptly in early April shortly after the onset of the dry season (Fig. 2). GEM concentrations during the 2015–16 wet season saw a similar, though
5 much less distinct decrease over a shorter and drier season. Figure 2 also shows that fetch-component radon concentrations begin to drop in both years around September–October, which HYSPLIT trajectories show is coincident with air mass origin shifting away from the Australian continent and towards the northern Arafura and Timor Seas. Throughout the wet season fetch-component radon remains low, though not at baseline levels (Zahorowski et al., 2013; Chambers et al., 2016b), implying that there is still some terrestrial influence on incoming air masses from the Australian continent or surrounding islands to the
10 north. Wet season wind data (Fig. 3) confirm that the predominant fetch during this period is from the west, where the Timor Sea lies less than 2 km from ATARS. Air-sea exchange of mercury is complex, with the ocean generally considered a net sink for atmospheric mercury (Mason and Sheu, 2002; Song et al., 2015). Reduction of mercury within the photolytic zone can give rise to increased concentrations of elemental mercury and hence evasion of GEM to the atmosphere (Soerensen et al., 2014). Terrestrial surfaces are also commonly sources of GEM; Nelson et al. (2012) modelled terrestrial mercury emission fluxes
15 over Australia that were generally between 8 and 44 $\mu\text{g m}^{-2} \text{a}^{-1}$ from soil and vegetation. Figure 3 does not show a strong difference in concentration distributions between the two source regions.

The increase in GEM concentrations in the early 2015 dry season was coincident with a shift to largely terrestrial-influenced fetch, as evidenced by a coincident increase in fetch-component radon, as well as with the conclusion of the monsoon season. The timing offset between decreases in GEM and fetch-component radon in the early wet and late dry seasons suggests that air
20 mass origin is not the only influence on wet season GEM decreases. Within tropical regions, wet deposition has been shown to be a significant pathway for mercury from the atmosphere to both oceanic and terrestrial ecosystems, even in relatively low-mercury air and despite the low solubility of mercury in its elemental form (Fostier et al., 2000; Costa et al., 2012; Hansen and Gay, 2013; Soerensen et al., 2014; Shanley et al., 2015). Mercury “rainout” — or the tendency for mercury rainwater loading to decrease with increasing precipitation — has also been demonstrated in Mercury Deposition Network (MDN) data in North
25 America (Glass and Sorensen, 1999; Prestbo and Gay, 2009) and positive correlations between GEM (TGM) and rainwater mercury have been reported in MDN data (GEM; Cole et al., 2014) and at Cape Point, South Africa (TGM; Brunke et al., 2016). Re-emission of any deposited mercury is likely to be inhibited throughout the wet season, as it has been shown that GEM emission from background mercury soils is suppressed when the soils are saturated (Briggs and Gustin, 2013). Mercury wet deposition is currently not being measured at ATARS; however, given the large differences in GEM trends between the
30 wet and dry seasons, these measurements could help to highlight differing processes between these periods.

3.2 Diurnal variation

Short, significant troughs in GEM values can be seen in Fig. 2, down to a minimum value of 0.28 ng m^{-3} . These are more pronounced in the dry season, though still common during the wet. GEM recoveries from standard additions during these periods were investigated and remained within 10 % of expected values with no evident pattern throughout the day, implying

the drops in observed GEM were due to natural phenomena and not a change in instrument GEM recovery. Atmospheric mercury depletion events (AMDEs) and the mechanisms behind them have been well-documented in polar regions (Steffen et al., 2008), though other similar events have been observed within the mid-latitudes (Mao et al., 2008; Brunke et al., 2010; Engle et al., 2010; Moore et al., 2013; Morrison et al., 2015; Fu et al., 2016; Howard and Edwards, 2017). The mechanisms
5 behind these mid-latitude depletion events are less clear and likely varied, with hypotheses such as chemical conversion of GEM to RM and subsequent deposition; transport of GEM-depleted air masses; or deposition of GEM from isolated atmospheric pools, being offered. Closer inspection of the dips in GEM observed at ATARS reveals that they occur overnight and are particularly pronounced in the early hours of the morning, with a marked rebound following sunrise.

The pattern of overnight GEM depletion is shown in diurnal composite data in Fig. 4, along with diurnal-component radon
10 and wind direction. Days have been defined from midday to midday, then sorted into groups according to quartiles of the diurnal-component radon value at sunrise (marked in the top figures). As radon fluxes are, across daily timescales, constant to first-order approximation, nocturnal build-up of radon is indicative of atmospheric stability, with highest radon values indicating the most stable atmospheres. This follows the radon-based stability categorisation method described by Chambers et al. (2016a) and Williams et al. (2016). In the dry season (left), it can clearly be seen that the magnitude of nocturnal GEM
15 depletion increases with increasing stability, and conversely, little-to-no depletion occurs under well-mixed boundary layers. Wind directions for the well-mixed category shift from coastal (westerly) in the early evening to terrestrial during the night. In contrast, wind directions for moderately-mixed to stable boundary layer categories are very similar to each other, shifting from a north-easterly to south-easterly direction shortly after sunset. Terrestrial fetches encompass this range of directions and the abrupt shift in wind direction at around 20:00 has little impact on the rates of GEM depletion or radon accumulation under
20 these stability categories. This shows that changes in advection of GEM from local source/sink regions are not responsible for observed depletion.

Wet season diurnal-component radon values (right) are lower than in the dry season, which fits with wind profile and fetch-
component radon data showing greater influence of oceanic fetch. Additionally, during the wet season rates of radon emission may be reduced in saturated soils, as reduction of pore space inhibits upward mobility to the point where radon within the
25 soil will undergo radioactive decay before reaching the surface (Griffiths et al., 2010). During the wet season, well-mixed and moderately-mixed categories are more indicative of the influence of ocean fetch than stability, as evidenced by wind directions of $273 \pm 8^\circ$ for these two categories. For weakly-mixed and stable categories, wind direction shifts southerly and easterly throughout the evening, from an oceanic fetch to a terrestrial fetch. It is not until this shift in wind direction occurs that GEM
depletion is observed, at a similar rate to that seen in the dry season under moderately-mixed to stable categories.

We suggest from these observations during the wet and dry seasons that the observed depletion results from deposition of
30 GEM over terrestrial surfaces. Under increasingly lower capping inversions associated with more stable boundary layers, a near-constant rate of surface deposition would result in greater concentration drops within the boundary layer, consistent with the observations at ATARS. Turbulent break-up of the nocturnal boundary layer at sunrise is also consistent with the rebound of GEM concentrations and drop in diurnal-component radon observed at this time. The rebound of GEM, however, begins
35 the hour before diurnal-component radon signals the break-up of the nocturnal boundary layer. In the absence of changes to

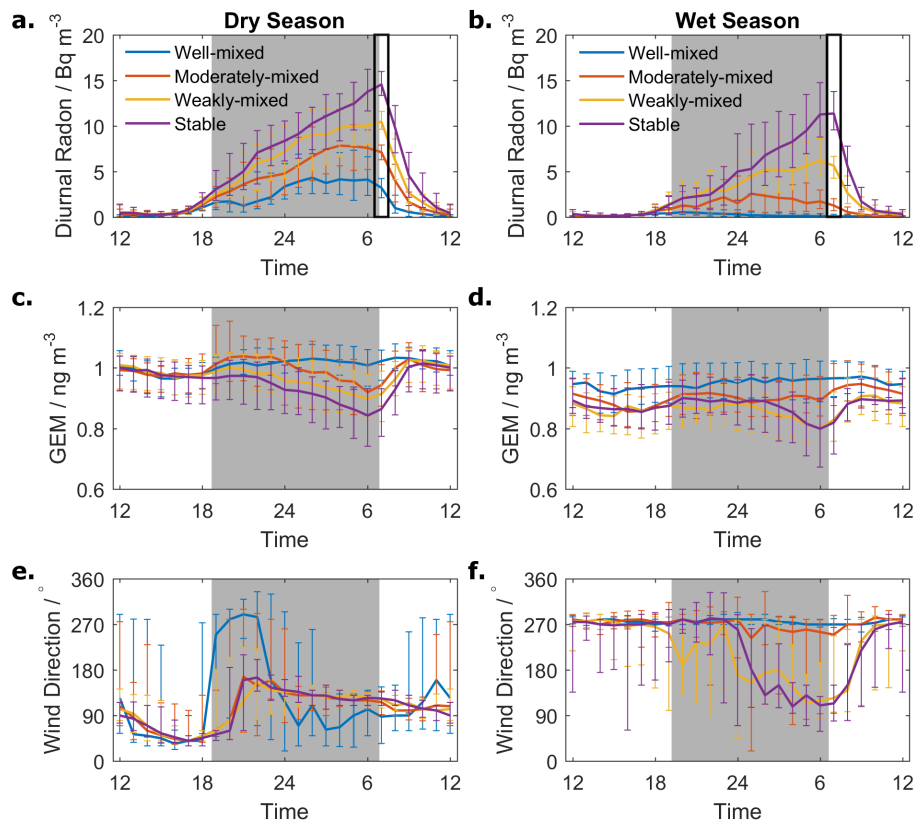


Figure 4. Diurnal composites of hourly radon (a,b), GEM (c,d) and wind direction (e,f) for (left) dry season data and (right) all wet season data. Edges of shading denote median sunset/sunrise times for each season. Data have been split into stability categories based on diurnal-component radon quartiles at sunrise (marked in top panels). Lines are median values and error bars indicate inter-quartile ranges.

advection or entrainment, this suggests emission of GEM from the surface. Furthermore, for stability categories where GEM depletion has taken place, daytime GEM concentration peaks at around 10:00 before decreasing to a minimum at around 15:00, where low radon values indicate the strongest turbulent mixing with free-tropospheric air. This “overshoot” of GEM in the early morning also cannot be explained by entrainment and, at least in the dry season, by changes to fetch. Early-morning GEM emission would likely be from the most readily volatile surface mercury, released under low-light conditions (shading denotes the period between geometric sunset/sunrise and so astronomical twilight will begin up to 75 minutes prior to the shaded edge). We propose that this initial release of GEM is volatilised from the reduction of mercury deposited overnight, as it has been shown that the most recently-deposited mercury during AMDEs is preferentially released due to photochemical reactions (Sherman et al., 2010).

10 Previous studies have shown that surface GEM fluxes over soils with mercury concentration at background levels are generally bi-directional, with little controlling influence from soil mercury concentration (Agnan et al., 2016, and references within). Correlations with solar radiation and air temperature tend to lead to emission fluxes throughout the day and deposition or near-

zero flux overnight. Howard and Edwards (2017), whilst undertaking micrometeorological measurements of surface GEM fluxes over a background mercury substrate grassland, observed nocturnal atmospheric mercury depletion events (NAMDEs) similar to the ones seen at ATARS. They attributed these events to enhanced nocturnal deposition of GEM under shallow, stable boundary layers. Enhancements in morning GEM emission were seen in days following the depletion events, similarly providing evidence for volatilisation of recently-deposited mercury. Further, cumulative GEM exchange over the 20-day study was near zero, highlighting the short-lived nature of this nocturnal GEM sink. This result, and the radon-based analyses presented earlier, provide strong evidence for a “multi-hop” process of atmospheric transport.

NAMDEs have also been observed in the northern hemisphere, in a range of ecosystems ranging from coastal to forested (Mao et al., 2008; Engle et al., 2010; Fu et al., 2016). Mao et al. (2008) attributed 70 % of their observed depletion to surface deposition and Fu et al. (2016) provided modelling evidence showing that stable boundary layers of height 100 m can be completely depleted of GEM due to deposition processes. The pervasiveness of NAMDEs across multiple ecosystems, and their pervasiveness throughout the ATARS time series across all seasons, suggests this multi-hop process is widespread. It is important to note that, due to inhibited mixing at the top of the nocturnal boundary layer, the extent of any nocturnal depletion is limited to within tens to hundreds of metres above the surface. Beyond this, movement of free-tropospheric air continues to enable long-range transport of GEM. Nevertheless, extensive and rapid bi-directional exchange with the surface would have a significant impact on our understanding of atmospheric mercury transport, impacting the relative importance of intermediate and regional-scale sources, as well as expected time scales for observed decreases in environmental mercury following actions proposed under the Minamata Convention (Lindberg et al., 2007).

3.2.1 Long-range transport

With seasonal changes in the latitudinal position of the ITCZ, ATARS is periodically located north of the atmospheric equator (Hamilton et al., 2008) and so the possibility of interhemispheric transport to the site was also of interest. Figure 2 shows the GEOS-Chem output for NH-released tracer concentrations at ATARS. Throughout most of the year — and consistently through the dry season — this value remains low, indicating that the site is far enough below the atmospheric equator to not be affected by transport of NH air. However, there are notable periods when this tracer value increases, along with coincident GEM increases. We arbitrarily defined air masses at the site to be significantly influenced by northern hemisphere air (herein termed “NH wet season”) when the ratio of NH tracers to SH tracers was greater than 0.5 (ratio not shown). Under this definition, ATARS saw 13 NH-influenced days over three distinct periods, all during the wet season and indicated in the lower panel of Fig. 2. Hereafter, wet season data that exclude these periods of NH influence are termed “SH wet season”.

The normalised frequency distribution of NH wet season GEM data is compared against those of dry season and SH wet season data in Fig. 5. Mean values for each were $1.08 \pm 0.12 \text{ ng m}^{-3}$ ($n = 3048$), $0.97 \pm 0.13 \text{ ng m}^{-3}$ ($n = 81073$) and $0.90 \pm 0.10 \text{ ng m}^{-3}$ ($n = 46191$), respectively. The differences between these means were small but significant; Student’s t-tests showed the minimum differences between the 95 % confidence interval of each mean to be 0.10 ng m^{-3} (NH wet – Dry) and 0.07 ng m^{-3} (Dry – SH wet). Comparison with log-normal probability density functions for other GMOS sites over the years 2013–14 (Fig. 4, Sprovieri et al., 2016) shows that GEM data sampled at ATARS are more closely related to those from other

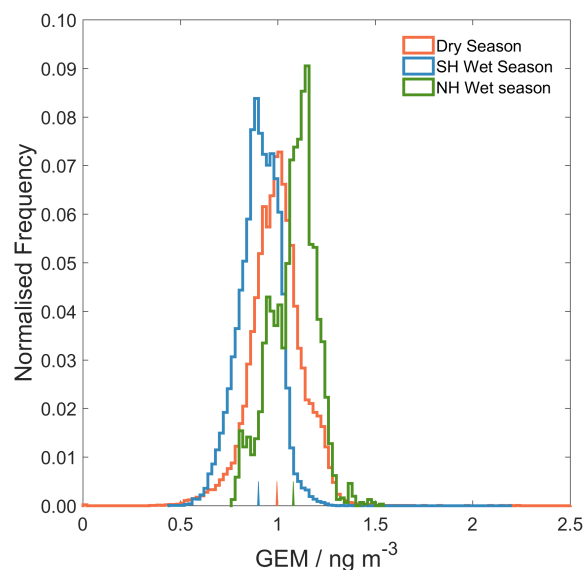


Figure 5. Normalised frequency for all 5-minute GEM data, split into dry season, SH wet season and NH wet season. Vertical lines at bottom of figure indicate mean values.

southern hemisphere sites, rather than tropical or northern hemisphere sites. This is likely due to the location of ATARS within the Maritime Continent — a region of high variability in the latitudinal position of the ITCZ — and its southerly latitude that places it outside this range and hence within the atmospheric southern hemisphere for most of the year.

- 5 Air mass source transport to ATARS across seasons was further investigated using 5-day HYSPLIT back trajectories. For NH-influenced air masses, use of 5-day trajectories and the geographic equator was found to be a poor predictor of NH influence at this site, with only 1.2 % of these trajectories originating from within the geographical NH. This is likely due to the significant disconnect between the geographical and meteorological equators over the Maritime Continent during the wet season. As such, 10-day back trajectories were calculated for these periods. Figure 6 shows median, 10th and 90th percentile
- 10 GEM-weighted trajectory coordinates for $0.5^\circ \times 0.5^\circ$ grid cells. During the dry season (top row), the influence of persistent high pressure cells across the Australian continent can be seen, with most air parcels flowing over central and north-eastern Australia. Changes to air mass source regions are seen with the southward movement of the ITCZ and associated low pressure cells that characterise the SH wet season (centre row). The differing GEM concentration distributions between the two seasons outlined earlier are further apparent in these two figures. For NH-influenced air masses (bottom row), this analysis shows that
- 15 most air masses — particularly those with the highest GEM concentrations — passed over the Indonesian archipelago. North of this, air masses moved over the South China Sea or Western Pacific Ocean, with little influence from terrestrial South East Asia. Given that Indonesia's population is greater than 250 million and its biomass burning season coincides with the Australian monsoon, it is likely that the observed increases in GEM concentrations in NH-influenced air masses are more indicative of anthropogenic or biomass GEM source influence from the Indonesian archipelago than the northern hemisphere background

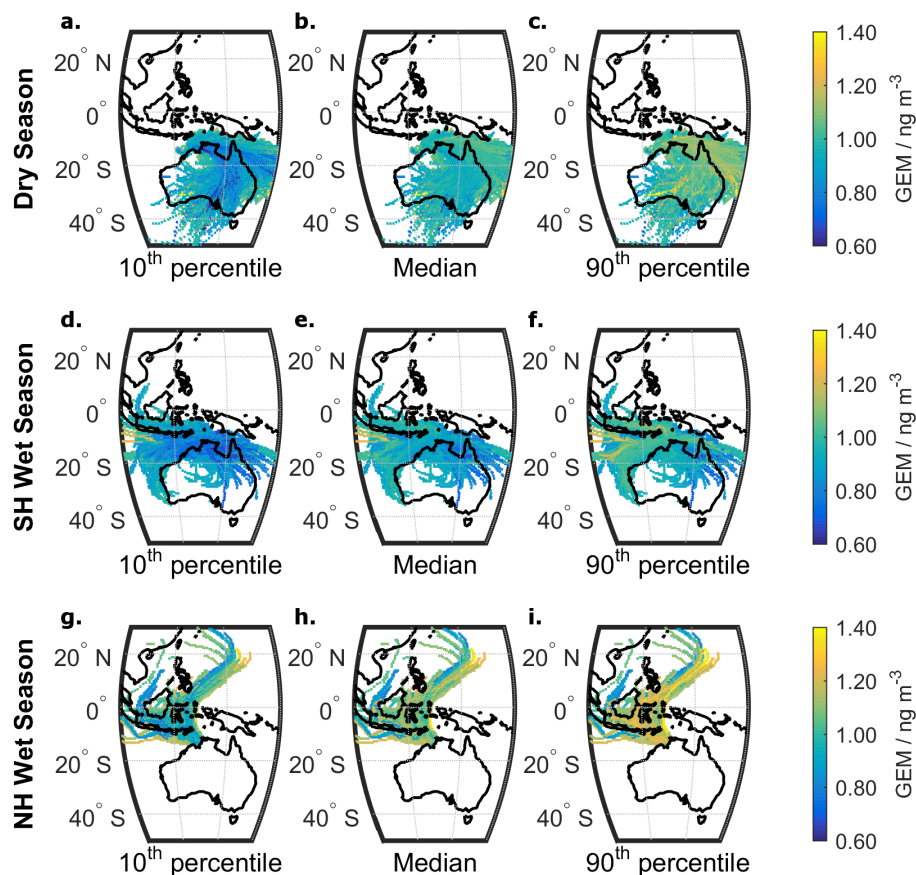


Figure 6. 10th percentile (left), median (centre) and 90th percentile (right) of hourly GEM-weighted HYSPLIT trajectories for $0.5^\circ \times 0.5^\circ$ grid squares. **a–c** are for dry season data, **d–f** for SH wet season data and **g–i** for NH wet season data. NH wet season map created using 10-day back trajectories, all others using 5-day trajectories.

source pool. Further investigation using chemical transport and mercury emission modelling is needed. Regardless, the current analysis shows that ATARS does observe air masses of northern hemisphere origin and that measurements of GEM and other atmospheric species during these periods may help to assess the effectiveness of transport models investigating hemispheric air exchange associated with movement of the atmospheric equator.

4 Conclusions

We present here the first two years of ongoing measurements of GEM taken in tropical Australia. Comparison with other Australian datasets suggests that a latitudinal gradient of GEM exists across the continent, with higher values towards the equator. Air masses from the northern hemisphere were shown to intermittently impact the tropical site ATARS, with associated

increases in GEM. Generally, the concentrations seen at ATARS were indicative of southern hemisphere rather than tropical air, as determined by comparison with other GMOS monitoring stations around the globe.

Seasonal variation in GEM was observed, with higher values observed in the tropical dry season compared to the wet. Spikes in GEM associated with biomass burning in the region were measured, taking place during the mid- to late-dry season. Wet
5 season GEM showed a decreasing trend throughout 2014–15; this was apparent though not as pronounced in the drier 2015–16 season. The cessation of this downward trend coincides with shifts of air mass source regions from oceanic to terrestrial, however the reverse is not the case for the onset of this trend. It is likely that precipitation rainout or aqueous-phase oxidation of GEM is responsible for this observed downward trend. Continued monitoring and wet deposition data may help to explain these seasonal features.

10 Daily cycles in GEM were observed at the site, characterised by nocturnal decreases in concentration followed by rapid increases around sunrise, then further decreases throughout the day. Differences in these daily trends between wet and dry seasons, along with associated changes in wind direction and stability, suggest that these nocturnal atmospheric mercury depletion events are related to dry deposition of GEM over terrestrial surfaces under increasingly stable boundary layers. Analyses using diurnal-component radon suggest the rapid increases around sunrise are partly due to volatilisation of newly-
15 deposited mercury, such as seen in other NAMDEs and in Arctic AMDEs. The extent of this multi-hop phenomenon may be widespread, which would have a significant impact on our understanding of atmospheric mercury transport, the delivery of atmospheric mercury to the environment, and the legacy of anthropogenic emissions of mercury.

Currently, multi-annual atmospheric mercury datasets for tropical and SH sites are rare, impacting the skill of regional and global models designed to further our understanding of the natural mercury cycle and its potential impacts on human and
20 environmental health. The value of measurements such as these is in comparisons with other similar measurements around the globe. As such, the addition of this site to monitoring networks such as the Global Mercury Observation System (GMOS) or the Asia Pacific Mercury Monitoring Network (APMMN) is important in achieving greater understanding of the mercury cycle, as it is currently only one of two monitoring sites located in the tropical eastern hemisphere.

Article 19 of the Minamata Convention commits parties to develop and improve anthropogenic mercury inventories; efforts
25 to monitor mercury and mercury compounds in environmental media; and modelling of mercury transport (including long-range transport and deposition), transformation and fate in a range of ecosystems. ATARS is uniquely positioned to enhance the information required for these monitoring and modelling activities.

5 Data availability

GEM data used for this publication are available from the GMOS data repository (<http://gmos.eu/sdi/>). Weather data are
30 collected and supplied by the Australian Bureau of Meteorology (<http://www.bom.gov.au/climate/data-services/>).

Competing interests. The authors declare that they have no conflict of interest.

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Table 1. Annual, seasonal and monthly mean, standard deviation and count for 5-minute GEM samples between June 2014 and June 2016. Wet season values calculated from hydrological years beginning in reported year.

	Year	Dry	Wet	Jan	Feb	Mar	Apr	May	Jun	Jul	Aug	Sep	Oct	Nov	Dec
2014	Mean	1.02	0.90	—	—	—	—	—	0.99	1.02	1.08	1.08	1.03	0.95	—
	Std. Dev.	0.10	0.12	—	—	—	—	—	0.07	0.09	0.10	0.16	0.13	0.07	0.08
	Count	34734	25060	—	—	—	—	—	2265	8421	1413	2592	3845	8238	7960
2015	Mean	0.93	0.94	0.92	0.79	0.76	0.82	0.89	0.95	0.96	1.00	0.99	1.01	0.99	0.96
	Std. Dev.	0.12	0.12	0.10	0.11	0.07	0.15	0.11	0.08	0.09	0.09	0.07	0.06	0.15	0.09
	Count	72060	54071	8184	4830	2447	5551	8848	8105	6964	5613	2076	5090	7636	6716
2016	Mean	0.92	—	0.94	0.91	0.88	—	—	—	—	—	—	—	—	—
	Std. Dev.	0.11	—	0.11	0.12	0.08	—	—	—	—	—	—	—	—	—
	Count	21576	—	8438	7631	5507	—	—	—	—	—	—	—	—	—