

# EUBREWNET RBCC-E Huelva 2015 Ozone Brewer Intercomparison

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**Abstract.** From May 25<sup>th</sup> to June 5<sup>th</sup> 2015, the [\[..<sup>1</sup>\]](#) 10<sup>th</sup> Regional intercomparison campaign of the Brewer Calibration Center - Europe (RBCC-E) was held at El Arenosillo atmospheric sounding station of the Instituto Nacional de Técnica Aeroespacial (INTA). This campaign was [\[..<sup>2</sup>\]](#) jointly conducted by COST Action ES1207 EUBREWNET and the Area of Instrumentation and Atmospheric Research of INTA. Twenty one [\[..<sup>3</sup>\]](#) Brewers, 11 singles and 10 double monochromator instruments from 5 eleven countries participated and [\[..<sup>4</sup>\]](#) were calibrated for total column ozone and solar UV irradiance [\[..<sup>5</sup>\]](#).  
. Every RBCC-E campaign advances our understanding of the Brewer calibrations in general and specifically what areas still require attention. In this 2015 campaign we have introduced a formal approach to characterization of the internal instrumental stray-light, the filter non-linearity and the algorithm for correcting for its effects on the TOC calculations. This work shows a general overview of the ozone comparison [\[..<sup>6</sup>\]](#), the evaluation of the correction of the spectral stray  
10 light effect for the single-monochromator Brewer spectrophotometer, derived from the comparison with a reference double-

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monochromator Brewer instrument. At the beginning of the campaign, 16 out of the 21 participating Brewer instruments [<sup>7</sup>] agreed within better than  $\pm 1\%$ , and 10 instruments [<sup>8</sup>] agreed within better than  $\pm 0.5\%$  [<sup>9</sup>] considering data with Ozone Slant Column between 100 and 900 DU, which doesn't require instrumental stray light correction [<sup>10</sup>].

## 1 Introduction

- 5 The fully automated Brewer Spectrophotometer (Brewer, 1973; Kerr et al., 1985; Kerr, 2010) is together with the Dobson ozone spectrophotometer, the backbone of the World Meteorological Organization (WMO) ozone observation network providing high quality Total Ozone Column (TOC) data for more than 30 years and is now deployed at 200 ground based TOC monitoring stations worldwide. It is also capable of measurements of ozone vertical profiles (Umkehr method), spectral UV radiation and aerosol optical depth in the UV (AOD-UV), as well as columns of other trace constituents such  
10 as sulphur dioxide and nitrogen dioxide.

In November 2003 the WMO/GAW Regional Brewer Calibration Center for Europe (RBCC-E) was established at the Izaña Atmospheric Observatory (IZO) of the Agencia Estatal de Meteorología (AEMET) in Tenerife (Canary Islands, Spain). The RBCC-E [<sup>11</sup>] consists of calibration laboratory facilities and reference-maintenance equipment mainly composed [<sup>12</sup>] of three Brewer spectrophotometers, the denoted IZO Triad. This includes a Regional Primary Reference (Brewer #157), a  
15 Regional Secondary Reference (Brewer #183), and a Regional Traveling Reference (Brewer #185) which can be transported for calibration campaigns outside IZO. Initially, the RBCC-E transferred the calibration from the World Reference Triad in Toronto. However, due to uncertainties on the future maintenance of the World Triad, in 2011, the WMO scientific advisory group (WMO-SAG) authorized the RBCC-E to transfer its own calibration obtained by the Langley method.

RBCC-E regular intercomparisons are held annually, alternating between Arosa in Switzerland, and the El Arenosillo sound-  
20 ing station of the Instituto Nacional de Técnica Aeroespacial (INTA) at Huelva in the south of Spain. Since 2005, a total of 130 Brewer ozone spectrophotometer calibrations have been performed in these campaigns (see the campaign reports at the RBCC-E website, <http://rbcce.aemet.es>, and the GAW reports of the VII [<sup>13</sup>](Redondas et al., 2015), VIII [<sup>14</sup>](Redondas and Rodriguez-Franco, 2016), and IX [<sup>15</sup>](Redondas and Rodriguez-Franco, 2015b) intercomparison campaigns). In addition to the regular intercomparisons, the RBCC-E performs two types of campaigns supported by the ESA CalVal project:  
25 the NORDIC campaigns, with the objective to study the ozone measurements at high latitudes, and the Absolute calibration

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<sup>7</sup>removed: (76%)

<sup>8</sup>removed: (50%)

<sup>9</sup>removed: . After applying the final calibration that included the

<sup>10</sup>removed: , all working instruments agreed at the  $\pm 0.5\%$  level.

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campaigns performed at IZO with the participation of Brewer and Dobson reference instruments. [..<sup>16</sup> ]Figure 1 shows the number of Brewer instruments calibrated at these campaigns since 2003.

[..<sup>17</sup> ]

[..<sup>18</sup> ]

## 5 1.1 [..<sup>19</sup> ]

The aim of COST Action 1207 “EUBREWNET” is to establish a coherent network of European stations equipped with Brewer spectrophotometers for the monitoring of total ozone column (TOC), spectral UV radiation, and aerosol optical depth in the UV spectral range (AOD-UV), ensuring sustainable operation in the long-term [..<sup>20</sup> ](Rimmer et al., 2018). Among the primary aims of EUBREWNET is to [..<sup>21</sup> ]harmonise operations and develop approaches, practices and protocols to achieve consistency in quality control, quality assurance and coordinated operations, as well as to eliminate duplication of efforts at individual stations[..<sup>22</sup> ]. It also aims at establishing knowledge exchange and training, and at opening up a route to link with international agencies and other networks globally. [..<sup>23</sup> ]

[..<sup>24</sup> ]It was proposed by Fioletov et al. (2008) that the main problem of the Brewer network was the lack of QA/QC, the [..<sup>25</sup> ]non-uniformity of the data processing, and the lack of reprocessing of the observations after the calibration. This problem is also aggravated because there are more than 100 agencies providing Brewer data with different calibration practices, operational procedures and data processing, including many of these stations within Europe. The intercomparisons are a basic tool to achieve EUBREWNET objectives, in particular during RBCC-E calibration campaigns the instruments perform a common measurement schedule, the observations are uniformly processed and can be used to organise operator training on applying uniform operational procedures. During these campaigns we are able to address some important issues which affect the Brewer performance, these include instrumental parameters and the calibration methodology (Redondas and Rodriguez-Franco, 2012)

The EUBREWNET RBCC-E Huelva 2015 Ozone Brewer Intercomparison campaign had two main objectives: to establish the current status of the network with the participation of about the half of the participating Brewers in the network and to test the improvements introduced during the COST action on the total ozone processing (the stray light correction on single Brewer and the attenuation filter corrections).As well as ozone calibration, solar UV irradiance calibration was also

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<sup>17</sup>removed: Brewer instruments calibrated since 2003 by the RBCC-E in regular campaigns (at Huelva and Arosa), Nordic intercomparisons, and the Absolute calibrations performed at the Izaña Observatory.

<sup>18</sup>removed: Panoramic view of the 21 Brewer spectrophotometers on the terrace of the El Arenosillo sounding station, Huelva, coming from Canada (1), Netherlands (2), United Kingdom (3), Switzerland (1), Finland (1), Greece (1), Denmark (2), Russia (1), Algeria (1) and Spain(7).

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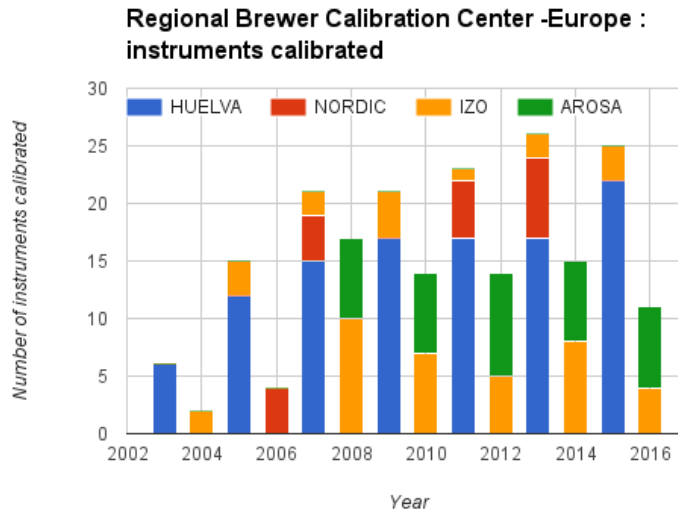
<sup>21</sup>removed: harmonize

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<sup>23</sup>removed: Close to 50 Brewer spectrophotometers are deployed in Europe, independently funded by national institutions.

<sup>24</sup>removed: In parallel to the campaign, COST Action 1207 organized several experiments to improve the quality of

<sup>25</sup>removed: Brewer data



**Figure 1.** Brewer instruments calibrated since 2003 by the RBCC-E in regular campaigns (at Huelva and Arosa), Nordic intercomparisons, and the Absolute calibrations performed at the Izaña Observatory.

**Figure 2.** Panoramic view of the 21 Brewer spectrophotometers on the terrace of the El Arenosillo sounding station, Huelva, coming from Canada (1), Netherlands (2), United Kingdom (3), Switzerland (1), Finland (1), Greece (1), Denmark (2), Russia (1), Algeria (1) and Spain(7).

performed by the travelling reference standard QASUME instrument of the World Calibration Center for UV (WCC-UV) and is described by the companion paper by (Lakkala et al., 2017), also it is remarkable that this campaign also provided the AOD-UV (López-Solano et al., 2017). [..<sup>26</sup> ]

5 The present work is organised as follows. The Brewer algorithm for TOC and the calibration methodology implemented during the RBCC-E campaigns is described in Sect 2, with focus on the improvements introduced by the stray light and the filter corrections. In Sect. 3 we describe and present results of the intercomparisons. In the blind days section we check the stability of the TOC for the approximately 2-year period between calibrations and the final days subsection shows the agreement of the instrument at the end of the campaign. In Sect. 4 we discuss the results and the implications of the correction introduced by the EUBREUNET processing and provide some closing remarks.

<sup>26</sup>removed: These included, studies on the Dead Time determination using the direct-sun measurements (Fountoulakis et al., 2016), characterization of the temperature dependence of the Brewer diffuser (Fountoulakis et al., 2017), cosine response measurements (?), investigation of the effects of polarization of the input window (Carreño et al., 2016), aerosol optical depth calibration (López-Solano et al., 2017), stray light characterization, comparison of total ozone between Phaethon and Brewer instruments, and an intercomparison of UV reference lamps

## 2 The calibration of the Brewer spectrophotometer

A brief description of the instrument and its operating principles in relation to the calibration is described in this section. A more detail description can be found in (Kerr, 2010; Savastiouk, 2006) .

The Brewer instrument measures the intensity of direct sunlight at six wavelengths ( $\lambda$ ) in the UV (303.2, 306.3, 310.1, 313.5, 5 316.8, and 320.1 nm) each covering a bandwidth of 0.5 nm (resolution power  $\lambda/\delta\lambda$  of approximately 600). The spectral measurement is achieved by a holographic grating in combination with a slit mask which selects the channel to be [..<sup>27</sup> ]analysed by a photomultiplier. The longest four wavelengths are used for the ozone calculation. The basic ozone measurement is the direct sun measurements. The sunlight enters the instrument through an inclined quartz window after which a right-angle prism directs the incoming light to the optical axis of the instrument. The light subsequently passes through the fore-optics, 10 which consist of a set of lenses to adequately focus the beam, an iris diaphragm, and two filter wheels. A ground quartz diffuser is located on the first filter wheel. The second filter wheel consists of a set of five neutral density filter attenuators which guarantee that the detector is working in its linear regime. After passing through the filter wheels, radiation is then focused onto the entrance slit of the monochromator. Presently, there are three commercially available installed Brewer models. The oldest model MKII is a single monochromator with a solar-blind filter a  $NiSO_4$  element sandwiched between 15 two UG-11 glass filters, located between the exit slit and the PMT to block radiation at wavelengths longer than about 325 nm. The model MKIV is capable of measuring  $NO_2$ , this instrument works with its grating in the third diffraction order for the ozone retrievals and the second diffraction order in  $NO_2$  retrievals. However, both instruments are affected by stray light, mainly, because of using a unique grating. In contrast, the the model MKIII is a double monochromator which provides enough stray light rejection to work in the first diffraction order for UV and ozone measurements.

20 Based on the Lambert-Beer law, the total ozone column in the Brewer algorithm can be expressed as [..<sup>28</sup> ](Kerr, 2010)

$$X = \frac{F - ETC}{\alpha\mu} \quad (1)$$

where  $F$  [..<sup>29</sup> ]is a linear combination of the logarithm of the measured spectral direct irradiances (also called double ratios) corrected for Rayleigh [..<sup>30</sup> ]molecular scattering,  $\alpha$  is the ozone absorption coefficient,  $\mu$  is the ozone air mass factor, and  $ETC$  is the extra-terrestrial constant. The  $F$ ,  $\alpha$  and  $ETC$  parameters are weighted functions at the operational wavelengths:

$$25 \quad F = \sum_i^4 w_i F_i - \frac{p}{p_0} \beta_i \mu \quad (2)$$

$$\alpha = \sum_i^4 w_i \alpha_i \quad (3)$$

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$$ETC = \sum_i^4 w_i F_{0i} \quad (4)$$

where  $\beta_i$  are the Rayleigh coefficients,  $p$  is the climatological pressure at the measurement site,  $p_0$  is the pressure at sea level, and  $F_i$  and  $F_{0i}$  are the individual measured and extra-terrestrial irradiances at each wavelength respectively. The weights  $w = [1, -0.5, -2.2, 1.7]$  have been chosen [..<sup>31</sup>] to minimise the influence of SO<sub>2</sub> and [..<sup>32</sup>] aerosol. To minimise effects of small shifts in wavelength (i.e.  $\delta F_o / \delta \lambda = 0$ ) the wavelength are chosen near a stationary point in the ozone absorption spectrum and as they verify:

$$\sum_i^4 w_i = 0 \quad (5)$$

$$\sum_i^4 w_i \lambda_i = 0 \quad (6)$$

This makes the measurement independent of any wavelength independent parameter like the absolute calibration [see Eq.(5)], and widely eliminates absorption features which depend, [..<sup>33</sup>] to first approximation, linearly on the wavelength, like for example the contribution from aerosols [..<sup>34</sup>] [see Eq.(6)] (Kerr, 2010).

[..<sup>35</sup>]

The Brewer retrieval of the TOC requires the knowledge of some instrument characteristics which are determined by calibration experiments during intercomparison campaigns (Redondas et al., 2015; Redondas and Rodriguez-Franco, 2015a, b, see e.g. the GAW reports of the Seventh, Eighth, and Ninth Intercomparison Campaigns of the RBCC-E). The instrumental calibration includes all the parameters that affect the [..<sup>36</sup>] counts measured by the spectrometer ( $F_i$ ), in particular the dead time correction [..<sup>37</sup>] (Fountoulakis et al., 2016), temperature coefficients (Berjón et al., 2017) [..<sup>38</sup>], and filter attenuations. The wavelength calibration [..<sup>39</sup>] determines the ozone absorption [..<sup>40</sup>] and Rayleigh scattering coefficients. The exact wavelengths measured by each Brewer spectrophotometer are slightly different from instrument to instrument. The so-called “dispersion test” [..<sup>41</sup>] is thus used to determine the exact wavelengths of each instrument and

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<sup>35</sup>removed: We can divide the calibration in instrumental, wavelength, and ETC transfer steps:

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its slit, or instrumental<sup>[.42]</sup> <sup>[.43]</sup>, function (Gröbner et al., 1998; Redondas et al., 2014a). An extra-terrestrial (calibration) constant is determined by the Langley method or by comparison with a reference instrument.

It is important to note that TOC It is then finally determined using ratios of measurements so there is no transfer of the radiometric scale. During the campaigns the transfer of the calibration to a network instrument is achieved by  
5 operating side by side with the reference <sup>[.44]</sup> Brewer. Once we have collected enough near-simultaneous direct sun ozone measurements, we calculate the new extra-terrestrial constant after imposing the condition that the measured ozone will be the same for simultaneous measurements for both instruments. In terms of Eq. 1, this leads to the following condition:

$$ETC_j = F_j + X_j^{\text{reference}} \alpha \mu_j \quad (7)$$

10 In the Langley determination of the ETC we assume that the ozone, as well other interfering absorbents, is constant through half of a day [see Eq. (7)] thus the regression of the double ratios  $F_j$  against the airmass will be a line with slope  $\alpha X$  and intercept  $ETC$  (Komhyr et al., 1989) . These assumptions are only achievable in clear environments and high altitude subtropical stations like Izaña and Mauna Loa. The Langley calibration of the RBCC-E instruments are described in detail in (Redondas et al., 2014b; León-Luis et al., 2018) and are outside the focus of this work.

15 For a correctly characterised network instrument, the determined ETC values show a Gaussian distribution and the mean value is used as the instrument's extra-terrestrial constant. One exception to this rule is the single monochromator Brewer models (MK-II and MK-IV) which are affected by stray light (Karppinen et al., 2015). In this case, the ETC distribution shows a tail at the lower ETC values for high Ozone Slant Column (OSC, the product of the total ozone content by the airmass). As we discuss in detail on the next section for this type of Brewer, only the stray-light-free region is used to  
20 determine the ETC, which generally ranges from 300 to 800 DU in the OSC, depending on the instrument.

The network Brewers were calibrated using the one parameter ETC transfer method, i.e., the ozone absorption coefficient was derived from the wavelength calibration (dispersion test) and only the ozone ETC constant was transferred from the reference instrument. The so-called "two parameters calibration method" (Staehelin et al., 2003), where the ozone absorption coefficient is also calculated from the reference, is also calculated and used as a quality indicator.

25 The calibration is an iterative process<sup>[.45]</sup>, changes during the instrumental and/or wavelength calibration will affect the final ETC<sup>[.46]</sup>, some instrumental characteristics are revealed during the ETC transfer by the comparison with the reference which requires full reprocessing of the calibration. For this reason the calibration campaigns are scheduled in three different periods:

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<sup>42</sup>removed: function, of each spectrophotometer, which differs slightly from instrumentto instrument (Redondas et al., 2014a).

<sup>43</sup>removed: Finally, the ETC transfer is performed by comparison

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<sup>46</sup>removed: and changes in the wavelength calibration will affect also to the final ETC

1. **Blind days:** the first days of the campaign are dedicated to determine the current status of the instrument **by comparison with the reference instrument**. During this period modifications of the instrument are not allowed.
2. **[..<sup>47</sup> ]Characterisation:** after the determination of how the instrument is measuring, the next days are dedicated to **[..<sup>48</sup> ]characterise** the instrument and perform the necessary adjustments and maintenance. The instrumental and wavelength calibration must be finished at the end of this period.
3. **Final days:** the period where the ETC transfer is performed, when the instrument is fully **[..<sup>49</sup> ]characterised** and stable.

The calibrations are maintained between transfers using the measurements of the internal quartz halogen lamp, the so called Standard Lamp (SL) test. A value, corresponding to a fictitious column density and often called R6, R5 or F-ratio (depending on whether the ozone, sulphur dioxide or nitrogen dioxide processing algorithm is applied), is obtained after processing. Slow variations of the SL test results with time may be representative of a change in the relative spectral (but not absolute, because of the way the retrieval algorithm works) sensitivity of the Brewer and may be even used to correct the final value of the solar measurements. This test is performed routinely to track the spectral response of the instrument and, therefore, the ozone calibration. A reference value for the SL, the so-called R6 ratio, is provided as part of the calibration of the instrument. The ozone is routinely corrected assuming that deviations of the R6 value from the reference value are the same as the changes in the ETC Extraterrestrial constant. This is then described by the Standard Lamp correction:

$$ETC_{new} = ETC_{old} - (SL_{ref} - SL_{measured}) \quad (8)$$

## 2.1 Stray light

There are two major sources of stray light on the Brewer: the “sky scattered” and the “out of band stray light, both of them produce an underestimation on the ozone calculation. The “sky scattered” stray light is due to the instrument having a field of view (FOV) larger than the apparent angle subtended by the sun in the sky, approximately 0.5° whereas the instrument FOV is about 1.5°, the measurement will thus include a “sky scattered” component (Josefsson, 2003, 2012). The measured radiation is a mixture of direct and sky scattered light with different spectral components resulting in relatively more radiation from shorter wavelengths resulting in an ozone underestimation. The contribution with high sun is negligible but for low sun or hazy conditions is highly significant.

The second source of stray light is due to the finite dimensions of the exit slits and imperfections on the gratings or other optical components such as the intensity measured contains radiation from other unwanted wavelengths. The higher

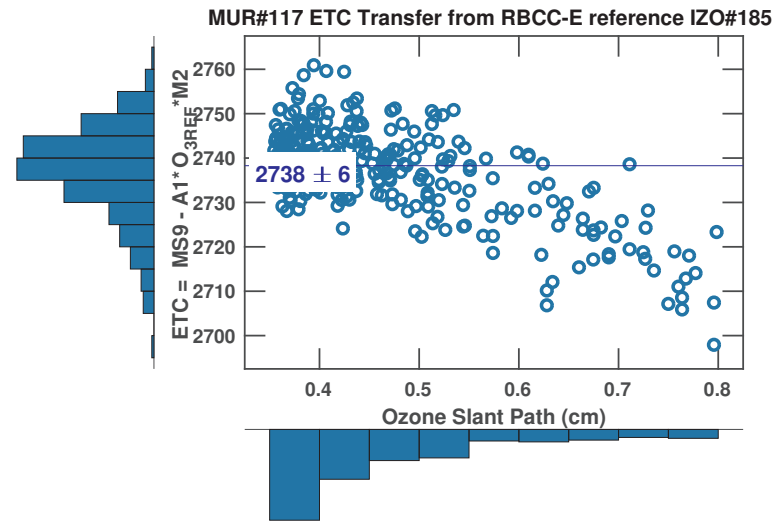
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[..<sup>50</sup>]

**Figure 3.** Distribution of individual ETC values determined by simultaneous measurements. In the horizontal axis, the ozone slant column (OSC). For this particular Brewer, the effect of the stray light is strong and clearly shown at values above 0.6 for the scaled OSC. The ETC is also calculated on the non-linear fit and compared with the 1 point calibration

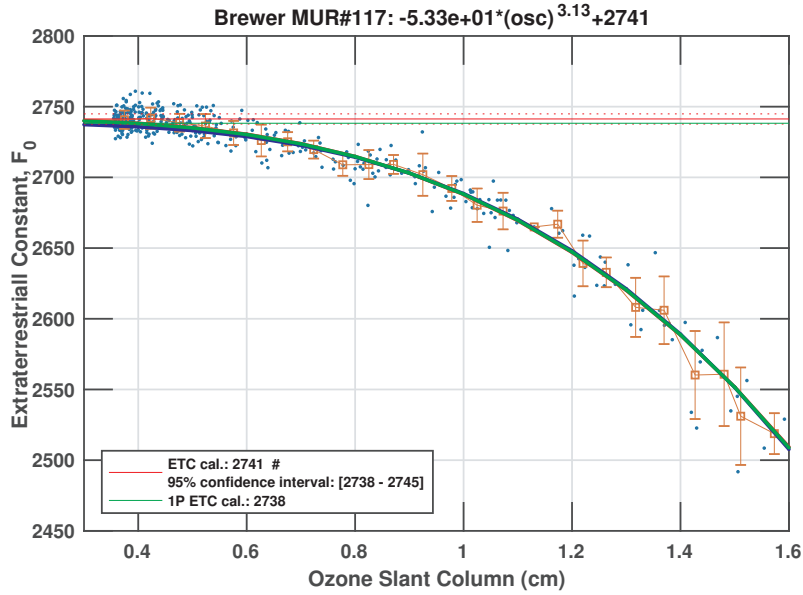
level of stray light in single monochromators increase the error of the measured UV radiation compared with the double-monochromators. This effect is largest at short UV wavelengths where the spectral irradiance level is low compared with the level of potential stray light. It is known to affect the Brewer spectral UV measurement (Kerr and McElroy, 1993; Bais et al., 1996) but only relatively recently has the effect on the ozone calculation been studied (Bojkov et al., 2008; Kiedron et al., 2008; Evans et al., 2009; Petropavlovskikh et al., 2011). The effect on the ozone calculation is enhanced when the short wavelength proportion is decreased due to a high slant path or high ozone content. The air mass in combination with the total ozone amount gives a measure of the absorption of the UV photons and is the key parameter for the stray light. There have been several attempts to model the stray light (e.g. Pulli et al. (2018) and the references within) but this requires the use of models or characterisation of the instruments that is not always available. Here we propose an empirical method based on the comparison of a double Brewer.

## 2.2 [..<sup>51</sup>]

The [..<sup>52</sup>]

<sup>51</sup>removed: ETC transfer and stray light

<sup>52</sup>removed: transfer of the calibration scale (namely, the ETC) to an instrument is done operating side by side with the reference Brewer. Once we have collected enough near-simultaneous direct sun ozone measurements, we calculate the new extraterrestrial constant after imposing the condition that the measured ozone will be the same for simultaneous measurements. In terms of Eq. 1, this leads to the following condition:



**Figure 4.** The stray light parameters  $k$  and  $s$  are determined by a nonlinear fit using the ETC determined from the stray-light free region as first guess parameters. The red horizontal line indicate the ETC constant retrieved from the fit, and the green one, the initial guess.

[..<sup>53</sup>]

[..<sup>54</sup>]

[..<sup>55</sup>] stray light effect can be corrected if the calibration is performed against a double monochromator instrument, assuming that it can be [..<sup>56</sup>] characterised following a power law of the ozone slant column:

$$5 \quad F = F[..<sup>57</sup>]<sub>m</sub> + k(X\mu)^s \quad (9)$$

where  $F$  are the true counts and [..<sup>58</sup>]  $F_m$ , the measured ones.

<sup>54</sup>removed: For a correctly characterized instrument, the determined ETC values show a Gaussian distribution and the mean value is used as the instrument's extraterrestrial constant. One exception to this rule is the single monochromator Brewer models (MK-II and MK-IV) which are affected by stray light (Karppinen et al., 2015). In this case, the ETC distribution shows (see Fig. 3) a tail at the lower ETC values for high Ozone Slant Column (OSC, the product of the total ozone content by the airmass). For this type of Brewer, only the stray-light-free region is used to determine the ETC, which generally ranges from 300 to 900 DU in the OSC, depending on the instrument.

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The extraterrestrial constant is

$$ETC_i = ETC[..<sup>59</sup>]_0 + k(X\mu)^s \quad (10)$$

where [<sup>60</sup>]  $ETC_0$  is the ETC for the stray-light-free OSC region and  $k$  and  $s$  are retrieved from the reference comparison (Figure 4). These parameters, determined in several campaigns, have been found to be stable and independent of the ozone calibration.

As the counts ( $F$ ) from the single Brewer instrument are affected by stray light, the ozone is calculated using an iterative process:

$$X_{i+1} = X_i + \frac{k(X_i\mu)^s}{\alpha\mu} \quad (11)$$

This empirical method was tested during the [NORDIC campaigns](#) and the counterpart campaigns at Izaña with the [MKII Brewer #037](#) operated by the [Finish Meteorological Institute](#) at Sodankyla (Finland) since 1988 (Karppinen et al., 2016). This instrument was calibrated four times at Izaña in 2009, 2011, and 2015 and at Sodankyla in 2011 (Rozen dael et al., 2013b, 2014). From these measurement campaigns we found the stray light correction obtained during the first campaign can be applied to the subsequent campaigns, obtaining a good agreement better than 0.5% on the 300-1800 OSC range. This is confirmed with the agreement of the determination of the stray light parameters  $k$  and  $s$  obtained during different campaigns (Table 3 of (Rozen dael et al., 2014)) at different locations with quite different sky conditions and even with changes on spectral response of the instrument. A similar experiment was done at Huelva with a good agreement using the previous campaign determination of the stray light parameters.

Figure 5 shows the ratio of the single Brewer participating in the campaign respect to reference instrument, where we can see that the OSC free region which we use for calibration starts for some instruments at 600 DU and is almost evident at 1000 DU, getting observations to 1600 DU is enough to determine the stray light constants. To help the determination of these parameters the measurement schedule is carefully defined to maximise the observations at high airmass, where around 30% of the simultaneous observations are performed with  $OSC > 600$  and 15% of  $>900$  with a significant number reaching the 1600 DU

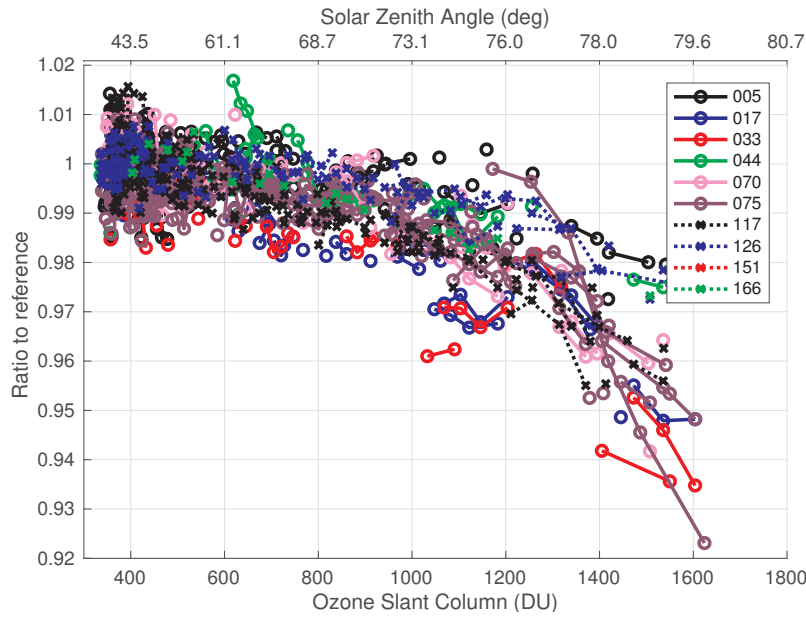
Usually just one iteration is needed for the atmospheric conditions at the intercomparisons carried out at El Arenosillo, with OSC values up to 1500 DU. For OSC measurements in the 1500–2000 DU range, two iterations are enough to correct the ozone (Figure 6). These stray light corrections are now implemented in the standard [<sup>61</sup>] [procedure](#) of EUBREWNET.

The effect of the stray-light has been studied in the Sodankyla Series, one of the longest record in the Arctic (Rozen dael et al., 2014). In this location even at noon the ozone slant column can be quite large. When the model from Karppinen et al. (2015) is applied to the series the effect on the monthly means is significant in spring but with a marginal effect on

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<sup>60</sup>removed:  $ETC_o$

<sup>61</sup>removed: processing



**Figure 5.** Percentage ozone differences with respect to the reference of the participating Single Brewers. On the upper x-axis the approximate solar zenith angle is indicated assuming TOC of 300 DU

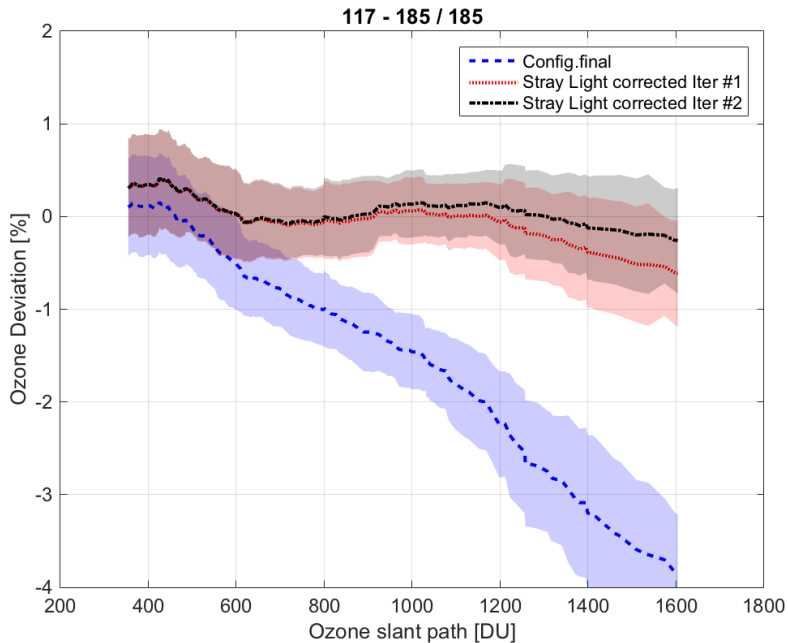
the trends. The comparison between the empirical correction and the model shows a good agreement (0.1 %) in most of the operative range but the model underestimate above 1600 DU reaching 1% at 2000 DU.

## 2.2 Correction of filter non-linearity

The Brewer spectrophotometer uses neutral density filters in order to optimise the intensity of the light which reaches the photomultiplier. The ozone is calculated using ratios so if the filters are neutral they have no effect on the ozone calculation (Eq5). In addition the weighing coefficients also verify (Eq6) so a linear attenuation with wavelength also does not affect the ozone calculation. In a real instrument however, some of the filters are not neutral, the ozone wavelengths vary from instrument to instrument, and the second condition is only an approximation. The effect on the calculated ozone will depends on the filter used (Savastiouk, 2006; Redondas et al., 2011). This dependence is clear when the instrument is compared with a reference and the comparison is seen to depend on the filter. Up to 20 ETC units (around 2% in ozone) has been observed during the campaign.

This error can be corrected if we know the spectral dependence of the filter  $AF_i$  and introduce a correction dependent of the filter used (f#)

$$F(f\#) = F_0 + \sum_{i=1}^4 w_i AF(f\#)_i$$



**Figure 6.** Percentage ozone differences with respect to the reference vs. Ozone Slant Path. In blue, using the final configuration constants, and in black and red, after the stray light correction has been applied, with one and two iterations, respectively. Data are averaged in  $[\pm 50 \text{ DU}]$  intervals, the shadow area represents one standard deviation.

The effect on ozone is only important at low air masses, where the ozone calculation is more dependent of the ETC, which implies high solar signal and in consequence high attenuation filters used.

Several methods have been developed to determine the wavelength dependence of the filter, using the internal lamp, the sun as source. However, as it affects high attenuation filters, there is not enough signal with the lamp and the non-linearity is difficult to determine with precision. Frequently the results are not significant with the number of test that can be performed during a campaign, and the comparison with a well characterised instrument is the preferred method during the intercomparisons. Also we can determine the correction directly by examining the record of the instrument and looking for the simultaneous measurements performed with consecutive filters or determine the ETC constant for every filter.

### 3 Intercomparison Results

#### 10 3.1 The X RBCC-E campaign

From May 25<sup>th</sup> to June 5<sup>th</sup> 2015, 21 Brewer spectrophotometers from 11 countries (see Table 2) took part in the X RBCC-E campaign held at the El Arenosillo atmospheric sounding station (Huelva, Spain). Besides the ozone calibration,

a solar UV irradiance calibration was performed by the traveling reference standard QASUME (Hülse et al., 2016) instrument of the World Calibration Center for UV (WCC-UV). The X RBCC-E campaign was the result of the collaboration between COST Action 1207 "EUBREWNET", <http://www.eubrewnet.org/cost1207>, and the Area of Instrumentation and Atmospheric Research of INTA (Redondas et al., 2016).

- 5 El Arenosillo Atmospheric Sounding Station, which belongs to the National Institute for Aerospace Technology (INTA), is located on the Atlantic Coast, in the province of Huelva of the region of Andalusia, in south-west Spain. Its surroundings correspond to the "Doñana" National Park, which guarantees its natural environment. Moreover, the climate at "El Arenosillo" is characterised by very frequent sunny conditions, around 280 clear sky days per year, being a really suitable site for intercomparison campaigns. The observatory has a big terrace with completely open horizon Fig(2). The
- 10 surroundings of the station consist of pine trees which provide a uniform albedo spatially and temporally throughout the whole year. This constant behaviour of the albedo allows the comparability of results obtained in different seasons and years.

- We collected during the campaign  $\approx 650$  direct sun ozone measurements with the reference instrument, most of them ( $\approx 65\%$ ) within the 300-600 DU and 18% on the 600-900 DU and 17% 900-1700 DU ozone slant path range, in larger
- 15 airmass. The mean number of near-simultaneous ozone measurements between the Brewers and the reference instrument was 350. Total ozone content values at El Arenosillo station during the intercomparison ranged between 320 to 380 DU. This campaign was characterised by high internal temperatures, with an average of  $32^{\circ}\text{C}$  and a standard deviation of  $5^{\circ}\text{C}$ , which is  $10^{\circ}\text{C}$  above the spectrometer normal operating temperatures (Berjón et al., 2017).

### 3.2 Reference Calibration

- 20 The RBCC-E triad is regularly calibrated, performing the instrumental [<sup>63</sup>]characterisation and wavelength calibration monthly. The three instruments are independently calibrated by the Langley plot method following the procedure described in [<sup>64</sup>]Redondas et al. (2014b) and León-Luis et al. (2018). Before and after the intercomparison campaigns, the [<sup>65</sup>]trav-
- elling instrument is compared with the two static instruments to verify that the calibration has not changed during transport (Figures 7 and 8), (Redondas et al., 2015; Redondas and Rodriguez-Franco, 2015a, b; León-Luis et al., 2018).

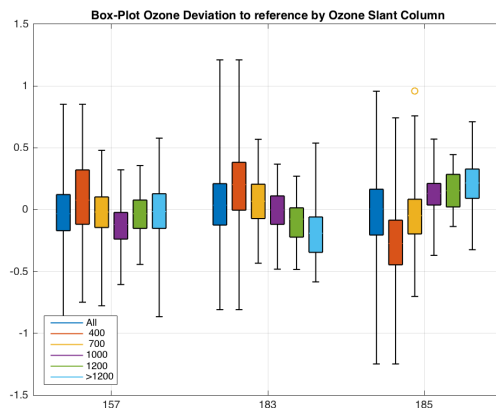
- 25 The campaign is a good opportunity to compare travelling reference instruments, that is instruments that are used to transfer calibrations. Brewers #017, managed by International Ozone Services (IOS) and directly calibrated to the Environment and Climate Change Canada, Toronto Triad (Fioletov et al., 2005; Fioletov and Natcheva, 2014; Natcheva, 2014), and #158, managed by Kipp & Zonen, manufacturer of the Brewer spectrophotometer, took part in the X RBCC-E campaign. Since
- 30 2007 the Brewer #158 is calibrated annually during the RBCC-E intercomparisons so it is already referenced to the RBCC-E triad. The calibration of IOS and RBCC-E Triads primary reference for the travelling instruments are discussed by León-Luis et al. (2018).

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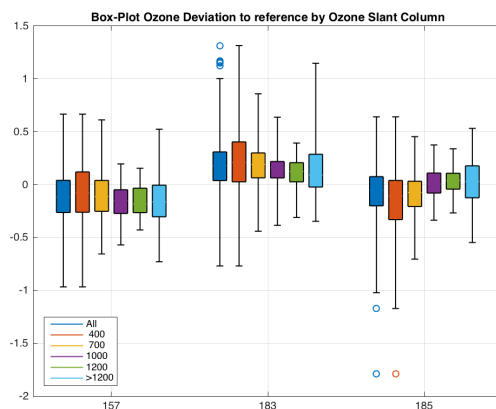
<sup>63</sup>removed: characterization

<sup>64</sup>removed: Ref. ?.

<sup>65</sup>removed: traveling



**Figure 7.** Box plot of the ozone percentage deviation from the mean before the X RBCC-E campaign at El Arenosillo in 2015.



**Figure 8.** Box plot of the ozone percentage deviation from the mean after the X RBCC-E campaign at El Arenosillo in 2015.

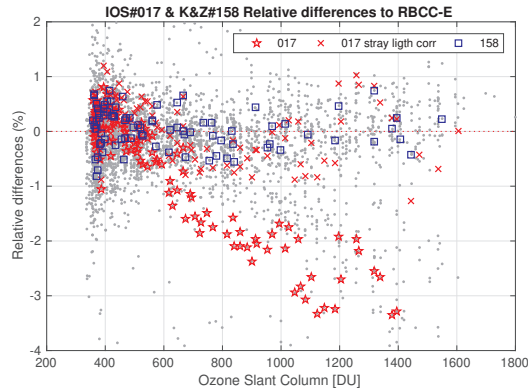
The agreement between the [\[..<sup>66</sup> \]travelling reference instrument during this campaign](#) was found to be quite good, with differences lower than 0.5% for OSC lower than 900 [\[..<sup>67</sup> \]DU](#) (see Table 1). Note that Brewer #017 is a single-monochromator instrument and is affected by stray light, thus underestimating the ozone at high OSC values above 600 [\[..<sup>68</sup> \]DU](#).

5 Table 3 shows the comparison of the reference instruments during the RBCC-E campaigns and the corresponding calibration report. The agreement is generally around +/- 0.5% but with exceptions. In contrast to the RBCC-E travelling instrument which is transported by boat/car to Huelva and as hand luggage using two extra-sites of the plane, to Arosa campaigns. The travelling instrument are usually transported by cargo and can have issues during transportation that are

<sup>66</sup>removed: reference instrument

<sup>67</sup>removed: DU

<sup>68</sup>removed: DU



**Figure 9.** Comparison of reference instruments during the X RBCC-E campaign: relative differences with respect to the IZO reference using the initial configuration during the campaign, in red for the IOS Brewer #017 (stars are used for the original observations and crosses for the stray light corrected ones), and in blue for the K&Z Brewer #158. The gray points are the relative differences to the IZO reference for all participating instruments

reflected in table 3 or require instrumental changes for example #158 had a new PMT and new electronics during Arosa 2014 and the SL tests do not reflect this change.

### 3.3 Blind Days

A blind comparison with the reference Brewer instrument is performed at the beginning of the campaign, thus providing information on the initial status of the instrument, i.e. how well the instrument performs using the original calibration constants (those operational at the instrument's station). Possible changes of the instrument response due to the travel can be detected through the analysis of internal tests performed before and after the travel.

[..<sup>69</sup>]

[..<sup>70</sup>]

10 The analysis of the SL historical record is one of the principal tools to establish the stability of the instrument calibration. Moreover the comparison with a reference during calibration campaigns is the most suitable tool to determine if the observed R6 changes are related or not with changes in the ETC constant.

<sup>69</sup>removed: The instruments are working during this period with their home calibration and the ozone is calculated using these calibration constants. The Standard Lamp (SL) test is an ozone measurement using the internal halogen lamp as a source. This test is performed routinely to track the spectral response of the instrument and, therefore, the ozone calibration. A reference value for the SL, the so-called R6 ratio, is provided as part of the calibration of the instrument. The ozone is routinely corrected assuming that deviations of the R6 value from the reference value are the same as the changes in the ETC Extraterrestrial constant. This then described by the Standard Lamp correction:



Fig/General\_008\_deviations\_O3\_blind\_-eps-converted-to.pdf

**Figure 10.** Percentage mean difference for the simultaneous direct sun measurements with the reference for all the participating instruments, with and without the standard lamp correction, in the stray-light-free OSC region (OSC<900). (The Brewer #151 is an not operative brewer , is not providing reliable data , and is outside the limits.)

[..<sup>71</sup> ]

During the El Arenosillo 2015 intercomparison campaign, most instruments agreed on average with the corresponding R6 reference value within  $\pm 10$  units, which is about 1% in ozone. [..<sup>72</sup> ]The stray light record tracks well small and slow changes on instrument responsivity but has issues when these changes are abrupt or huge. Generally, from our experience, the SL correction tracks the changes in the relative spectral sensitivity of the Brewers well unless the instruments characteristics change (whether intentionally or not) in ways that affect the ozone observations differently from the SL measurements, e.g. changes in the iris diaphragm, neutral density filters or zenith prism pointing towards the lamp or the sun. During the analysis we focus on the instruments that showed deviations of R6 values to the reference larger than 20 units (Fig. 11). Even in these cases for some instruments, for example Brewer#228, the SL correction improves the comparison, whereas for others like #163, where the change was produced by the modification of the fore optics the opposite happens. The comparison with a standard reference instrument is the only way to assess whether the SL correction properly tracks changes

<sup>71</sup>removed: Standard lamp R6 difference with respect to the R6 reference value from the last calibration during the blind days, before the maintenance. Variations within the  $\pm 10$  units range ( $\sim 1\%$  in ozone) are considered normal, whereas larger changes would require further analysis of the instrument performance.

<sup>72</sup>removed: Some instruments

on the calibration [..<sup>73</sup> ]or the changes observed are just due to changes of the lamp’s spectral emission (Fig. 10).[..<sup>74</sup> ][..<sup>75</sup> ][..<sup>76</sup> ]This analysis will determine if a re-evaluation of the ozone observations between calibrations are required after an analysis of the history of the instrument.

Table 1 shows the mean relative difference for the simultaneous direct sun measurements with the reference for all the participating instruments, with and without the standard lamp correction, in the stray-light-free OSC region. With the exception of Brewer #151, that can not be considered an operational instrument, the maximum difference found is 1.5%. This is a really good result considering that most of the instruments were calibrated two years [..<sup>210</sup> ]previously. The third column of the table shows the average of the best result for all the observation OSC range. This result is an estimation of the calibration agreement of the EUBREWNET network, with half of the instruments showing a perfect agreement within  $\pm 0.5\%$ , and 75% within the  $\pm 1\%$  level.

### 3.4 Final comparison

We define the final days as those available after the maintenance work has been finished for each participating instrument. These days are used to calculate the final calibration constants, so we [..<sup>211</sup> ]endeavour not to manipulate the instruments during this period. Furthermore, the SL R6 value recorded during the final days is normally adopted as the new reference value. It is also expected that this parameter will not vary more than 5 units during this period. We show in Fig. 12 the differences between the daily standard lamp R6 ratio and the proposed R6 reference value during the final days. As expected, the recorded SL values did not vary more than 5 units during this period.

Deviations of ozone values for all the participating instruments with respect to the RBCC-E travelling standard Brewer #185 are shown in Fig. 13 and summarized in Table 1. We have recalculated the ozone measurements using the final calibration constants and, in the case of single Brewer instruments, with and without the stray light correction as described in Sec. 2.

[..<sup>212</sup> ]The effect of the stray light correction is not large on the statistics, only 30% ( $sza > 60^\circ$ ) of the observations are affected on the single Brewer, and only for %15 ( $sza > 70^\circ$ ) is the effect bigger than 1% but taking into account the stray light allows for the instruments to be calibrated using the one parameter method. Not taking into account this correction using the two parameters method as in Bojkov et al. (2008) can cause a misleading calibration constant. For all the instruments, both the one parameter and the two parameters ETC transfer methods agreed to each other within the limit of  $\pm 5$  units for ETC constants and  $\pm 0.3\%$  for ozone absorption coefficients, which is an indication of the quality of the calibration provided.

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<sup>73</sup>removed: constants

<sup>74</sup>removed: In some instruments, for example Brewer

<sup>75</sup>removed: 075, the SL correction improves the comparison, whereas for others like

<sup>76</sup>removed: 165 the opposite happens. This

<sup>210</sup>removed: ago

<sup>211</sup>removed: tried

<sup>212</sup>removed: It should be noted that all Brewers were calibrated using the one parameter ETC transfer method, i.e., the ozone absorption coefficient was derived from the wavelength calibration (dispersion test) and only the ozone ETC constant was transferred from the reference instrument. The so-called “two parameters calibration method” ?, where the ozone absorption coefficient is also calculated from the reference, is also used as a quality indicator

We achieved a good agreement with the reference instrument Brewer #185 using the final calibration constants, see Fig. 13 and Table 1. With the application of the stray light correction to the single Brewer spectrophotometers, all instruments are within the  $\pm 0.5\%$  agreement range.

#### 4 Conclusions

5 <sup>213</sup> Briefly the Brewer network is calibrated with two main parameters: the ETC and the effective ozone absorption coefficient. The comparison with a reference instrument is used to establish the ETC but the effective ozone absorption coefficient can be derived also from the reference (two parameters method) as was done in the past (Staehelin et al., 2003) or directly from wavelength calibration as is performed at present (one parameter method) (Fioletov et al., 2005).  
10 adopted in the Brewer network. Although both methods give the same results in the 300-800 DU range (stray light free range), the two parameters calibration gives uniform results, smoothing the instrumental differences and reduces the stray light error on single Brewer (Bojkov et al., 2008). On the other hand, the one parameter calibration is more robust, it does not depend on reference wavelength calibration, and it highlights instrumental differences. An error in the ETC value yields an error that depends on the solar zenith angle, while an error in effective ozone absorption coefficient introduces a  
15 relative bias. The transition in the calibration methodology around the year 2000 can explain the change of the seasonal difference between ground base Brewer and satellites observed by Fioletov et al. (2008). During the RBCC-E campaigns we can show that the two point calibration can mask instrumental issues which are air-mass dependent. Moreover, both calibration methods give the same results on well characterised instruments and the difference between the calibration constants can be used as an indicator of the quality of the instrument calibration (Redondas and Rodriguez-Franco, 2012; 20 Roozendaal et al., 2014). For all participating instruments both the one parameter and the two parameters ETC transfer methods agreed to each other within the limit of  $\pm 5$  units for ETC constants and  $\pm 0.3\%$  for ozone absorption coefficients, indicating a high quality calibration.

To summarise the calibration results of the <sup>214</sup> 10th RBCC-E campaign, we found that during the blind days, using the two-year-old calibration issued in the previous campaign,

- 25
- 16 Brewer spectrophotometers ( $\sim 75\%$  of the participating instruments) were within the 1% agreement range.
  - 10 Brewer spectrophotometers ( $\sim 50\%$ ) were within the  $\pm 0.5\%$  range, i.e., show a perfect agreement.
  - The max average error was 1.5% for operational Brewer instruments within stray-light free conditions (OSC < 700 <sup>215</sup> DU).

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<sup>213</sup> removed: To summarize

<sup>214</sup> removed: X

<sup>215</sup> removed: DU

This results are in agreement with the RBCC-E campaigns celebrated in Huelva and Arosa from 2009 to 2015 (Figure 14), in this period 85 spectrometers [..<sup>216</sup> ]have been calibrated: 59 (69%) [..<sup>217</sup> ]show an agreement better than 1%, 32 (38%) within 0.5% and 7 (8%) [..<sup>218</sup> ]show a discrepancy greater than 2%.

After the new calibration was issued at the end of the X RBCC-E campaign,

- 5
  - All participating Brewer spectrophotometers were within the  $\pm 0.5\%$  agreement range.
  - Without the [..<sup>219</sup> ]stray light correction implemented large errors of up to 4% can be expected for single-monochromator Brewer instruments operating at OSC larger than 1000 [..<sup>220</sup> ]DU.
  - The implementation of the stray light correction in the calibration of single Brewer instruments improved their performance.

10 *Acknowledgements.* All this work would have not been possible without the participation, work and dedication of all the Brewer operators in the RBCCE intercomparison campaigns. [..<sup>224</sup> ]

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<sup>219</sup>removed: Stray Light

<sup>220</sup>removed: DU

<sup>224</sup>removed: Without their work and dedication this work would be not possible.

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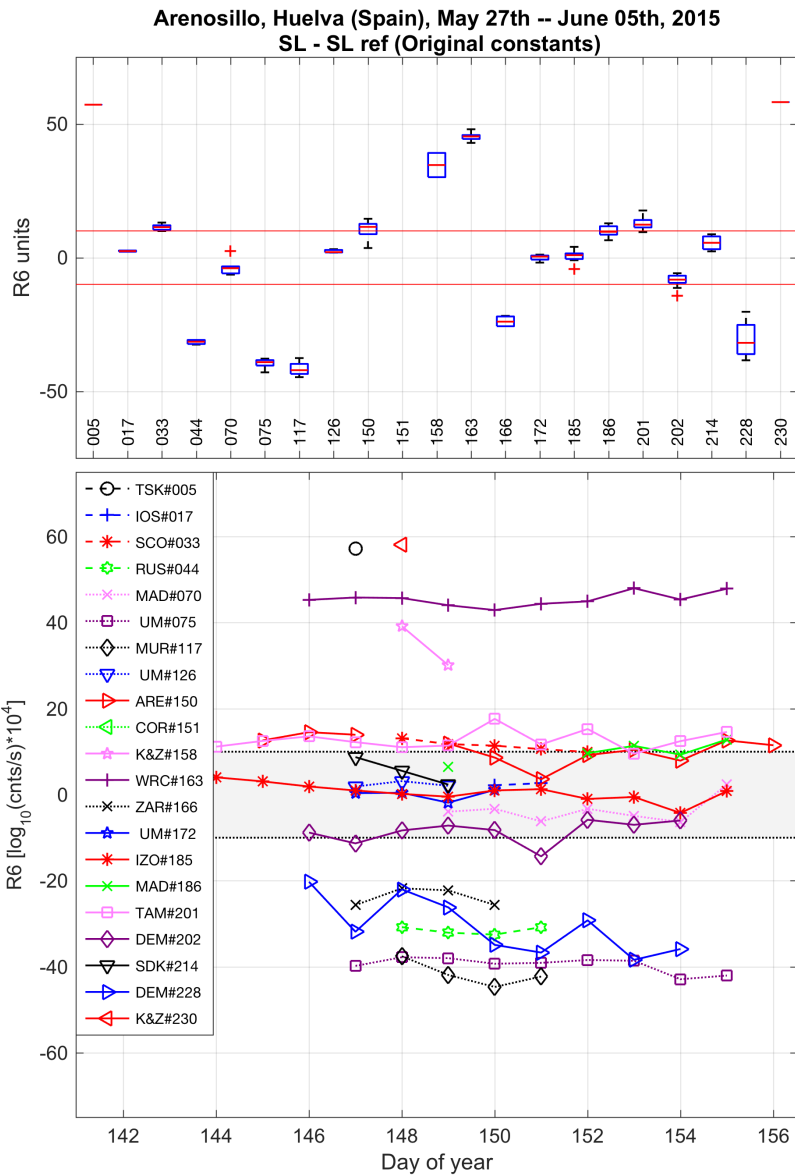
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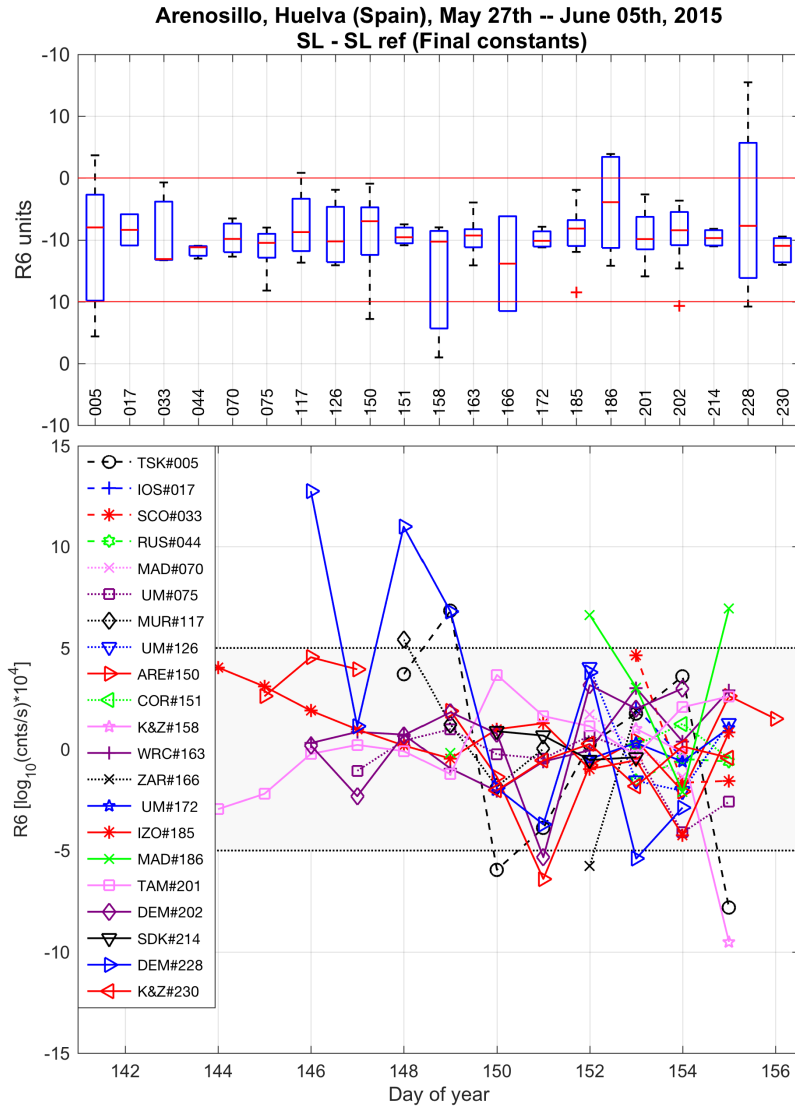
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**Figure 11.** [..<sup>77</sup>] Standard lamp R6 difference [..<sup>78</sup>] with respect to the R6 reference value from the last calibration during the blind days, [..<sup>79</sup>] before the maintenance, the upper panel shows the mean value for each instrument and [..<sup>80</sup>] the [..<sup>81</sup>] lower panel the [..<sup>82</sup>] daily mean during the [..<sup>83</sup>] campaigning. Variations within the  $\pm 10$  units range ( $\sim 1\%$  in ozone) from the reference value are considered normal, whereas larger changes would require further analysis of the instrument performance.

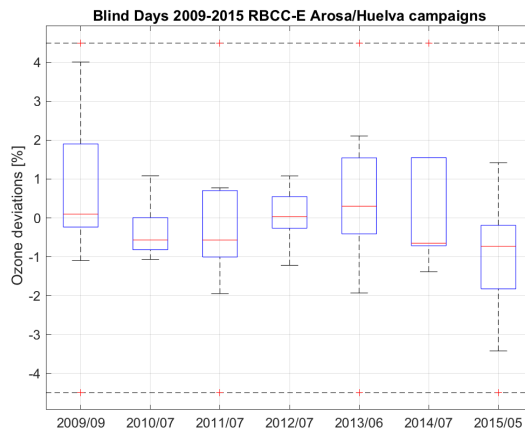


**Figure 12.** Differences between the daily standard lamp R6 ratio and the proposed R6 reference value during the final days, the upper panel shows the mean value for each instrument and the lower panel the daily mean during the campaign, is important to get an stable R6 value during the final days , within the  $\pm 5$  units range ( $\sim 5\%$  in ozone), to stabilise the reference value of the calibration.



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**Figure 13.** Final days mean percentage difference with respect to the reference Brewer for the simultaneous direct sun measurements for all the participating instruments, blue circles shows results without the stray light correction and red stars show results with the correction applied to single Brewer spectrophotometers.



**Figure 14.** Ozone deviations for the Blind Days with respect to the reference Brewer for the simultaneous direct sun measurements for all the participating instruments during the RBCC-E regular campaigns 2009-2015, the campaigns performed in odd years correspond to Arosa (Switzerland) and in [\[..<sup>221</sup>\]](#) even years in Huelva (Spain). This results correspond to the stray light free region  $OSC < 700$  [\[..<sup>222</sup>\]](#) DU, the outliers (red cross at +/- 4.5% levels) generally correspond to [\[..<sup>223</sup>\]](#) non-operating instruments.

**Table 1.** Summary of mean percentage difference before calibration, without and with Standard Lamp Correction, and after the calibration, on the last column with the stray light correction applied.

<b>Brewer ID</b>	<b>No corr.</b>	<b>SL corr.</b>	<b>Blind</b>	<b>Final</b>	<b>Stray</b>
<b>005</b>	-	-	-1.93	-0.2	-0.08
<b>017</b>	-0.31	-0.49	-0.98	-0.95	0.11
<b>033</b>	-0.8	-1.77	-1.09	-1.83	-0.48
<b>044</b>	-2.04	0.13	-0.21	-0.27	0.2
<b>070</b>	-0.73	-0.42	-0.71	-0.53	0.18
<b>075</b>	-3.42	-0.71	-1.2	-0.8	-0.2
<b>117</b>	-3.38	-0.45	-0.68	-0.6	0.04
<b>126</b>	-1.25	-1.41	-1.36	-0.29	-0.08
<b>150</b>	-0.45	-1.07	-0.45	-0.27	-
<b>151</b>	-17.36	9.94	7.95	0.67	0.83
<b>158</b>	-0.54	-2.45	-0.54	0.05	-
<b>163</b>	-1.5	-4.16	-1.5	-0.06	-
<b>166</b>	-0.15	1.45	-0.24	-0.58	-
<b>172</b>	-0.67	-0.67	-0.67	-0.01	-
<b>186</b>	0.13	-0.34	0.13	-0.05	-
<b>201</b>	1.21	0.52	0.52	0.09	-
<b>202</b>	-1.39	-0.95	-0.95	-0.06	-
<b>214</b>	1.42	1.19	1.19	-0.01	-
<b>228</b>	-1.93	-0.4	-0.4	-0.1	-
<b>230</b>	-0.15	-3.48	-0.15	0.36	-

**Table 2.** Principal Investigators and Instruments participating on the X RBCC-E campaign

Nr.	Country	Brewer	Participants	
1	Greece	005	Alkis Bais	Thessaloniki University
2	Canada	017	Volodia Savastiouk	International Ozone Services
3	Spain	033	Juan R. Moreta	AEMET, State Meteorological Agency from Spain
4	Russia Federation	044	Vadim Shirotov	Scientific and Production Association “Typhoon”
5	Spain	070	Juan R. Moreta	AEMET, State Meteorological Agency from Spain
6	United Kingdom	075	John Rimmer	Manchester University
7	Spain	117	Juan R. Moreta	State Meteorological Agency from Spain
8	United Kingdom	126	John Rimmer	Manchester University
9	Spain	150	J. M. Vilaplana	National Institute for Aerospace Technolog
10	Spain	151	Juan R. Moreta	State Meteorological Agency from Spain
11	Netherlands	158	Oleksii Marianenko	Kipp & Zonen
12	Switzerland	163	Julian [.. <sup>225</sup> ]Gröebner	Physikalisch-Meteorologisches Observatorium Davos
13	Spain	166	Juan R. Moreta	AEMET, State Meteorological Agency from Spain
14	United Kingdom	172	John Rimmer	Manchester University
15	Spain	185	Alberto Redondas	Izaña Atmospheric Research Center, AEMET
16	Spain	186	Juan R. Moreta	AEMET, State Meteorological Agency from Spain
17	Algeria	201	Bukelia Lamine	National Meteorological Office
18	Denmark	202	Paul Eriksen	Danish Meteorological Institute,
19	Finland	[.. <sup>226</sup> ]214	Tomi [.. <sup>227</sup> ]Karpprin	Finnish Meteorological Institute
20	Denmark	228	Niss Jepsen	Danish Meteorological Institute,
21	Netherlands	230	Keith M. Wilson	Kipp & Zonen

**Table 3.** Reference Comparison during RBCC-E campaigns

Location	year	#017	#158	#145	Report
Arosa	2008	-0.6			(Redondas and Rodriguez-Franco, 2008)
Huelva	2009	-0.6	0.8	-0.1	(Roosendaal et al., 2012)
Arosa	2010	-0.6			(Roosendaal et al., 2013b)
Huelva	2011	-0.1	-0.2	-0.6	(Roosendaal et al., 2013a)
Arosa	2012		-0.1		(Redondas et al., 2015)
Huelva	2013	-1.0	0.7		(Redondas and Rodriguez-Franco, 2015a)
Izaña	2014			-2.2	(Redondas et al., 2014b)
Arosa	2014	-1.2	1.5		(Redondas and Rodriguez-Franco, 2015b)
Huelva	2015	-0.5	-0.5		this work