Widespread polar stratospheric ice clouds in the 2015/2016 Arctic winter – Implications for ice nucleation

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Abstract. Low planetary wave activity led to a stable vortex with exceptionally cold temperatures in the 2015/2016 Arctic

15   winter. Extended areas with temperatures below the ice frost point temperature T_{ice} persisted over weeks in the Arctic

stratosphere as derived from the 36-years temperature climatology of the ERA-Interim reanalysis data set of the European

Center for Medium Range Weather Forecast ECMWF. These extreme conditions promoted the formation of widespread

polar stratospheric ice clouds (ice PSCs). The space-borne Cloud-Aerosol Lidar with Orthogonal Polarization CALIOP

instrument onboard the CALIPSO satellite (Cloud-Aerosol Lidar and Infrared Pathfinder Satellite Observation) continuously

20   measured ice PSCs for about a month with maximum extensions of up to 2×10^6 km^2 in the stratosphere.

On 22 January 2016, the WALES (Water Vapor Lidar Experiment in Space - airborne demonstrator) lidar onboard the High

Altitude and Long Range Research Aircraft HALO detected an ice PSC with a horizontal length of more than 1400 km. The

ice PSC extended between 18 and 24 km altitude and was surrounded by nitric acid trihydrate (NAT) particles, supercooled

ternary solution (STS) droplets and particle mixtures. The ice PSC occurrence histogram in the backscatter ratio to particle

25   depolarization ratio optical space exhibits two ice modes with high or low particle depolarization ratios. Domain filling 8-

days back-trajectories starting in the high-depol ice mode are continuously below the NAT equilibrium temperature T_{NAT} and

decrease below T_{ice} ~10 hours prior to the observation. Their matches with CALIPSO PSC curtain plots demonstrate the

presence of NAT PSCs prior to high-depol ice suggesting that the ice had nucleated on NAT. Vice versa, STS or no PSCs

were detected by CALIPSO prior to low-depol ice. In addition to ice nucleation in STS potentially with meteoric inclusions,

we find evidence for ice nucleation on NAT in the Arctic winter 2015/2016. The observation of widespread Arctic ice PSCs with high or low particle depolarization ratios advances our understanding of ice nucleation in polar latitudes. It further provides a new observational data base for the parameterization of ice nucleation schemes in atmospheric models.

1 Introduction

5   While synoptic-scale ice PSCs commonly occur in the Antarctic winter stratosphere (Solomon et al., 1986), widespread ice PSCs extending over several thousand km² have rarely been observed in the Arctic. Even since enhanced observational coverage of the polar regions by the CALIOP instrument (Pitts et al., 2009; Pitts et al., 2011) onboard the CALIPSO satellite, synoptic-scale ice PSCs were detected only occasionally in the Arctic (Pitts et al., 2013; Engel et al., 2013; Achtert and Tesche, 2014; Koshrawi et al., 2016). Sporadic evidence for synoptic scale ice PSCs in the Arctic is also derived from

10   infrared emission measurements from space with the Michelson Interferometer for Passive Atmospheric sounding (MIPAS) (Spang et al., 2017). Generally Arctic stratospheric temperatures are above the ice frost point temperature T_{ice} (Murphy and Koop, 2005) on synoptic scales and hence limit the formation of extended ice PSCs. The reason for the warmer temperatures in the Arctic compared to the Antarctic is a more alternated land-ocean contrast in the northern hemisphere, supporting the generation of planetary waves, which disturb and hence weaken the Arctic polar vortex due to inmixing of warmer mid-

15   latitude air (Solomon, 2004). In addition, radiative heating of the displaced and elongated vortex contributes to warmer Arctic vortex temperatures.

Ice PSCs exist at cold conditions with temperatures below T_{ice}, while other PSC types prevail at higher temperatures. These can contain NAT particles (Voigt et al., 2000a; Fahey et al., 2001), STS (Dye et al., 1992; Carslaw et al., 1994; Schreiner et al., 1999a; Voigt et al., 2000b) and particle mixtures. Other nitric acid containing solid particle types such as nitric acid

20   dihydrate NAD (Stetzer et al., 2006), or nitric acid water condensates (Thornberry et al., 2013; Gao et al., 2015) have been measured in laboratory. However, robust atmospheric evidence for NAD is missing so far (Höpfner et al. 2006a). Complementary to in-situ particle composition measurements (e.g. Schreiner et al., 1999b, Northway et al., 2002), lidar measurements led to a cloud type classification based on the optical properties of solid or liquid PSC particles (Toon et al., 2000). PSCs with depolarization below 0.04 at 532 nm wavelength were classified as STS (Pitts et al., 2009). PSCs with

25   higher depolarization were labelled Mix1 and Mix2, probably NAT clouds with lower or higher NAT particle number densities, respectively, and some amounts of STS. Finally depolarizing PSCs with backscatter ratios above 5 were classified as ice. Later studies (Pitts et al., 2011; Pitts et al., 2013; Achtert and Tesche, 2014) explored a more detailed differentiation of PSC types. In addition or in combination with lidar measurements, infrared emission measurements allow for advanced atmospheric PSC composition analysis and a classification in PSC type fractions (e.g. Höpfner et al., 2006a&b, Lambert et

30   al., 2012).

Large ice crystals in PSCs may sediment down and transport water vapor to lower altitudes (Fahey et al., 1990; Schiller et al., 2002). Further, due to their large surface areas, ice PSCs very efficiently process halogen compounds (Hanson and

Ravishankara, 1992) and hence contribute to ozone loss (Toon et al., 1989). Processing of halogenated reservoir gases on PSC particles leads to a release of unstable chlorine and bromine species, which become activated by sunlight in polar spring and effectively destroy ozone in catalytic cycles (Farman et al., 1985; Crutzen and Arnold, 1986). Ozone depletion is stopped, when the activated chlorine and bromine species react with nitrogen dioxide to reform stable reservoirs In the absence of ice at warmer temperatures, other PSC types can provide surfaces necessary for heterogeneous chlorine processing (Grooß et al., 2005; Manney et al., 2011; Drdla and Müller, 2012). E.g. in the Arctic winters 2009/10 and 2010/11 (Manney et al., 2011; Sinnhuber et al., 2011; von Hobe et al., 2013), polar ozone loss may have largely been driven by STS (Wohltmann et al., 2013) and NAT (Nakajima et al., 2016). In these cases, the role of ice as transporter for nitric acid enhancing denitrification and slowing down ozone loss may gain importance.

Ice PSCs may nucleate homogeneously in liquid STS aerosol or heterogeneously by the aid of solid particles (Toon et al., 1989; Peter, 1997; Zondlo et al., 2000). Homogeneous ice nucleation in STS may occur at high cooling rates as observed in localized mountain wave events (e.g. Dörnbrack et al., 2002). On synoptic scales in the Arctic, the ice saturation ratio S_{ice} rarely increases above the homogeneous ice nucleation threshold of ~1.6 for supercooled ternary solution droplets (Koop et al., 2000). Temperatures rarely decrease 3 to 4 K below T_{ice}, therefore homogeneous ice nucleation (Koop et al., 2000) might play a minor role in the Arctic on the synoptic scale. Heterogeneous ice nucleation was investigated in the laboratory by Hoose and Möhler (2012) for various ice nuclei down to temperatures of -60°C. In the polar stratosphere, heterogeneous ice nucleation in STS with meteoric dust (Cziczo et al., 2001; Curtius et al., 2005; Weigel et al., 2014) aided by small scale temperature fluctuations helped to explain the formation of a synoptic-scale ice PSC in the Arctic observed in January 2010 (Pitts et al., 2013; Engel et al., 2013). Here we present new measurements of large scale ice PSCs in the Arctic winter 2015/2016 and discuss ice nucleation pathways. In addition to ice nucleation in STS with meteoric dust inclusions as proposed by Engel et al. (2013), we suggest that ice nucleation on pre-existing NAT might be required to explain a second branch in the ice PSC occurrence histogram.

First, we describe the instrumentation and methods. Then we give an overview of the meteorological conditions, which led to a strong cooling of the polar vortex in the Arctic winter 2015/2016. Ice PSCs were observed over elongated periods by the spaceborne CALIOP lidar as described in section 4. On 22 January 2016, differential absorption lidar measurements of an elongated ice PSC were performed onboard the HALO research aircraft during the POLSTRACC (Polar Stratosphere in a Changing Climate) campaign. Based on the PSC occurrence histogram in the backscatter ratio to depolarization optical space, we define a threshold for the $1/R_{ice}$ threshold for ice and investigate the effect of different thresholds on ice PSC occurrence in sensitivity studies. The PSC histogram shows two branches in ice PSC occurrence. We calculate back-trajectories starting in each of the two branches and investigate possible ice formation pathways for each branch. Finally we discuss the implications of the suggested ice formation pathway on NAT for Arctic PSCs and tropical ice clouds and suggest a way forward for PSC modeling.

2 Instrumentation and methods

In this study, we use lidar measurements from the research aircraft HALO and from the CALIPSO spacecraft in combination with meteorological data from the European Centre for Medium-Range Weather Forecasts (ECMWF) numerical weather prediction model to investigate occurrence and formation of ice PSCs in the Arctic winter 2015/2016 with a special focus on
5 22 January 2016.

2.1 WALES lidar measurements on HALO during the POLSTRACC campaign

From 7 December 2015 till 20 March 2016, the international field campaign PGS (POLSTRACC/GW-CYCLE/SALSA) was conducted with the High Altitude and Long Range Gulfstream G550 research aircraft HALO out of Oberpfaffenhofen,
10 Germany (48°N, 11°E) and Kiruna, Sweden (68°N, 20°E). The POLSTRACC campaign particularly focused on investigating polar stratospheric chemistry and dynamics in the 2015/2016 Arctic winter. With a ceiling altitude of 15 km HALO is perfectly suited to study the evolution of the lower part of the polar vortex and to provide specific in-situ PSC information at these altitudes.

PSC abundance above HALO flight altitudes was detected with the WALES lidar instrument (Wirth et al., 2010; Groß et al.,
15 2014). The WALES lidar as configured during POLSTRACC has backscatter channels at 532 nm and 1064 nm wavelengths and additionally a high spectral resolution lidar (HSRL) channel and a depolarization channel at 532 nm wavelength for particle detection. The HSRL capability allows the retrieval of the extinction corrected backscatter coefficient of clouds at 532 nm without assumptions about the phase function of the particles (Esselborn, 2008). The backscatter ratio R is the ratio of the un-attenuated (extinction corrected) total (unpolarized) backscatter coefficient and the molecular backscatter. The
20 extinction correction is done by the HSRL channel using a molecular reference profile calculated from pressure and temperature data from ECMWF operational analyses at ~ 16 km horizontal resolution (6 h temporal resolution) and short term forecasts (1 h steps) to interpolate between the analyses. To further distinguish between particles of different type we use the depolarisation of the linear polarised laser light caused by scattering on non-spherical particles. The depolarisation caused by molecular scattering is removed from the signal, following the method outlined by Freudenthaler et al. (2009).
25 The quantity used further is therefore called linear *particle* depolarisation ratio. The relative sensitivity of the two polarized channels is recalibrated regularly during flight to guarantee reliable depolarization values. The particle depolarisation ratio is sensitive to the particle shape and size. Spherical particles do not depolarize and particles much smaller than the wavelength of the laser light also show unmeasurable low values. But in general there is no simple relation between depolarization and particle shape, size or composition, see for example Reichardt et al. (2002) for a more detailed discussion of this topic.
30 Nevertheless, cloud regions which show distinct depolarisation ratios point to a significantly different shape or size distribution. In combination with the backscatter ratio R, the particle depolarisation has been successfully used to

discriminate different PSC types from ground based, airborne and spaceborne lidar measurements (e.g. Pitts et al. (2009, 2011, 2013), Achtert and Tesche (2014) and references therein).

2.2 Spaceborne CALIOP lidar data

The Cloud-Aerosol Lidar with Orthogonal Polarization (CALIOP) lidar on board the Cloud-Aerosol Lidar and Infrared
5   Pathfinder Satellite Observations (CALIPSO) satellite measures backscatter at wavelengths of 1064 nm and 532 nm, with the 532 nm signal separated into parallel and perpendicular polarization components. The general performance of CALIOP and calibration of the CALIOP data are discussed in Hunt et al. (2009) and Powell et al. (2009). The PSC results in this paper are based on the Version 2.0 CALIOP PSC detection and composition discrimination algorithm (Pitts et al., 2018), which uses night time-only profiles from 8.2 to 30 km altitude of CALIOP V4.10 Lidar Level 1B 532 nm data smoothed to a
10   uniform 5 km horizontal (along track) by 180 m vertical resolution grid.

Following the methodology of Pitts et al. (2009; 2013), PSCs are detected as statistical outliers relative to the background stratospheric aerosol population in either 532 nm perpendicular backscatter (β_{perp}) or 532 nm scattering ratio R, which is the ratio of total backscatter to molecular backscatter. Successive horizontal averaging (5, 15, 45, and 135 km) is also used to ensure that strongly scattering PSCs (e.g., fully developed STS and ice) are found at the finest possible spatial resolution
15   while also enabling the detection of more tenuous PSCs (e.g., low number density liquid-NAT mixtures) through additional averaging. CALIOP PSC composition classification is based on comparing CALIOP data with temperature-dependent theoretical optical calculations of β_{perp} and R for non-equilibrium mixtures of liquid (binary H_2SO_4-H_2O or STS) droplets and NAT or ice particles. The assumption of NAT instead of nitric acid dihydrate (NAD) particles is based on Höpfner et al. (2006), who found no spectroscopic evidence for the presence of NAD from MIPAS observations of PSCs over Antarctica in
20   2003. The main improvement in PSC algorithm (Pitts et al., 2018) is that the threshold value of R separating ice and NAT mixture PSCs (R_{ice}) is calculated as a function of altitude and time based on the observed abundance of nitric acid and water as estimated from nearly coincident Aura MLS measurements (Manney et al., 2016). Thus R_{ice} (or $1/R_{ice}$) take into account the impact of denitrification/dehydration on the optical signature of ice and NAT mixture clouds. In mid-late January 2016, $1/R_{ice}$ range between 0.2 and 0.4 in the 18-24 km altitude region. The CALIPSO PSC areal coverage is estimated as the sum
25   of the occurrence frequency (number of PSC detections divided by the total number of observations) in ten equal area latitude bands spanning 90°-50° N, multiplied by the area of each band. We assume that the CALIOP observations from the approximately 15 daily orbits are representative of the PSC coverage within each latitude band. The data are aggregated on daily time scales and smoothed over 7 days to reduce noise.

2.3 Meteorological data sets

30   We use meteorological data from two operational analyses of the Integrated Forecasting System (IFS) of the European Centre for Medium-Range Weather Forecasts (ECMWF) to describe the meteorological conditions of the polar stratosphere in winter 2015/2016. From December 2015 to March 2016, the IFS produced the operational analyses cycle 41r1

(T_L1279L137) with a horizontal resolution of about 16 km ($0.14° \times 0.14°$ at the equator) and the experimental IFS cycle 41r2 (T_C1279L137) simultaneously with a higher resolution of about 8 km ($0.07° \times 0.07°$ at the equator). The later IFS cycle became operational after 8 March 2016 (Holm et al., 2016). The enhanced horizontal resolution was achieved by changing from linear to cubic spectral truncation and introducing an octahedral reduced Gaussian grid (Malardel and Wedi 2016).

5   Here, we show data in both resolutions for the 1 December 2015 to 8 March 2016 period to investigate the effects of a higher resolution on the meteorological data set, after 8 March 2016 the high resolution data are presented. We give 6-hourly operational analysis and use 1-hourly forecast data to interpolate between the time steps.

In addition, to compare to previous years, we use 6-hourly ERA-Interim reanalysis data (Dee et al. 2011) retrieved at a horizontal resolution of ($1° \times 1°$). ERA-Interim is a global atmospheric reanalysis from 1989 to today. The data assimilation

10   system used to produce ERA-Interim data is based on the release of the IFS cycle 31r2 (T_L255L60) in 2006. The system includes a 4-dimensional variational analysis with a 6-hour analysis window. The spatial resolution of the data set is ~ 80 km horizontally on 60 vertical levels from the surface up to 0.1 hPa. ERA-Interim data can be downloaded from the ECMWF public datasets web interface or from MARS archive. A detailed documentation of the ERA-Interim data archive is given by Berrisford et al. (2011).

15   We derive the minimum temperature T_{min} (K) between 65° and 90°N at the 30-hPa pressure surface from the ERA-Interim reanalysis and from the two operational analyses of the IFS. We further calculate the area A_{ice} with temperatures below the ice frost point T_{ice} using Murphy and Koop (2005) and the area A_{NAT} with temperatures below the NAT equilibrium temperature T_{NAT} using Hanson and Mauersberger (1988) for 4.6 ppmv H_2O and 7 ppbv HNO_3, as measured by MLS in the Arctic vortex in January 2016 (Manney and Lawrence, 2016).

20   **3 Temperature evolution of the Arctic stratosphere in the 2015/2016 winter**

The temperature evolution in the Arctic winter stratosphere is influenced by planetary wave activity. In early winter 2015 a strong tropical tropospheric temperature anomaly reinforced the meridional temperature gradient from the tropics to the poles which led to adverse conditions for the propagation of planetary waves (Matthias et al., 2016). Weak planetary wave activity measured as low meridional heat flux thus enforced the formation of a strong and stable polar vortex and caused

25   extremely low temperatures in the Arctic winter 2015/2016 (Dörnbrack et al., 2016).

Therefore, stratospheric temperatures decreased dramatically as derived from the meteorological data of the IFS of the ECMWF numerical weather prediction model. Figure 1A shows the evolution of T_{min} at latitudes > 65°N and at 30 hPa in the Arctic winter 2015/2016 in two resolutions. In December 2015, T_{min} decreased below T_{ice} and then remained below T_{ice} from late December 2015 till end of January 2016. Within the first cold phase, T_{min} down to 182 K were detected in the

30   operational IFS analysis with 16 km horizontal resolution. In February 2016, three minor stratospheric warmings influenced the vortex and led to warmer conditions in a coherent but slightly displaced polar vortex (Manney and Lawrence, 2016).

Then, the final stratospheric warming resulted in a split of the vortex by mid-March and the subsequent dissipation of the vortices associated with a temperature increase of more than 20 K at 30 hPa in a few days.

The exceptional coldness of the Arctic winter 2015/2016 also reflects in the fact that by the end of December 2015, T_{min} dropped even below the minimum temperatures $> 65°N$ at 30 hPa ever obtained by the ERA interim data record extending

5 from 1989 to 2016 (Dee et al., 2011). Throughout the Arctic winter 2015/2016, T_{min} was continuously lower than the mean temperature of the 36-years ERA interim data record (Voigt et al., 2016), see Figure 1.

To illustrate the effect of mesoscale temperature fluctuations on T_{min}, we also show T_{min} derived from the IFS cycle 41r2 at ~ 8 km horizontal resolution compared to the cycle 41r1 at ~ 16 km resolution at 30hPa for the December 2015 to 8 March 2016 time period where both data sets are available. Temperature deviations up to 7 K between the higher and lower

10 resolution data sets occurred during the first cold phase with $T_{min} < T_{ice}$ in December and early January as shown in Figure 1A. At higher resolution, minimum temperatures reach down to 179 K on the 30 hPa level. In the second cold period at the end of January 2016, temperature deviations up to 3 K are found in the higher resolution data set. Mesoscale gravity wave activity in December and January 2016 is better covered in the higher resolution data and could have caused this temperature difference (Dörnbrack et al., 2016). On 8 March 2016, the two data sets merge and from then on, ECMWF operational

15 analysis are given at ~ 8 km resolution.

At 30 hPa, the area A_{ice} with $T < T_{ice}$ extended over regions up to 3.6×10^6 km^2 (see Figure 1B), as derived from IFS cycle 41r1 analysis at 16 km resolution. From 18 January till the end of the month, A_{ice} was continuously more than one order of magnitude larger than the 36-years average and larger than the maximum of the 36-years ERA interim data record. In addition the area A_{NAT} with $T < T_{NAT}$ in January 2016 continuously reached the maximum of the 36-years ERA interim data

20 set (Figure 1C). Generally A_{ice} and A_{NAT} are lower than the mean Antarctic conditions as given by the dashed gray line in Figure 1B and C (shifted by 6 months to account for seasonality).

These extremely cold stratospheric winter conditions in the Arctic set the stage for synoptic-scale PSC formation in winter 2015/2016.

4 Occurrence of ice PSCs in the Arctic winter 2015/2016 derived from CALIOP observations

25 The CALIOP lidar onboard the CALIPSO spacecraft detected PSCs from December 2015 to January 2016 (Figure 2). On 28 January 2016 the CALIOP science data acquisition was suspended due to a spacecraft anomaly. The problem was subsequently resolved and data acquisition began again on 14 March 2016.

Ice PSCs were measured by CALIOP continuously for a month from late December 2015 to late January 2016 (Figure 2A). The ice PSCs were observed at altitudes between 15 and 26 km during the period with extremely cold temperatures inside

30 the Arctic vortex. The maximum extension of ice PSCs derived from CALIOP $A_{ice,max}$ of $(1.75 – 2.0) \times 10^6$ km^2 is reached on 30 December 2015. In this phase, the ice PSC formation was predominantly triggered by mountain wave activity and spread out to synoptic scales, see Dörnbrack et al. (2016). Considering the uncertainties in A_{ice} retrieval from the CALIOP

data set interpolated to latitude bands, the maximum A_{ice} derived from CALIOP observations agrees reasonably with maximum A_{ice} derived from weather forecast IFS data of 2.1×10^6 km^2 at 30 hPa (~21.6 km). Also the second peak in ice PSC occurrence A_{ice} on 24 January 2016 is captured by the CALIOP routine though with a smaller peak amplitude. A decrease of water vapor concentrations and dehydration due to falling ice crystals has been observed by MLS (Manney and Lawrence, 2016) at these altitudes throughout January 2016. This may explain lower A_{ice} derived from CALIOP data at the end of January compared to A_{ice} derived from in the IFS data set using a fixed H_2O mixing ratio of 4.6 ppmv to calculate A_{ice}.

NAT mixtures and STS PSCs were present in the CALIOP data set from early December 2015 till the end of the observation period (Figure 2B). Summarized, we find strong evidence for the unprecedented existence of widespread ice PSCs in the Arctic winter 2015/2016.

5 Optical properties of the PSC measured on 22 January 2016

PSCs were detected with the WALES lidar inside the vortex on all 6 HALO flights between 22 January and 29 February 2016. Before that date, the lidar on HALO was not operational. Due to the strong temperature increase at the end of January and measurement locations of HALO in warmer parts of the vortex, extended ice PSCs were observed by WALES solely on 22 January 2016.

5.1 Extension of the PSC on 22 January 2016

A large synoptic scale PSC was measured by WALES during a flight from Kiruna to the northern tip of Greenland on 22 January 2016 as shown in Figure 3. The PSC extended between 14 and 25 km altitude over a horizontal distance of 2200 km. It was continuously observed within in the Arctic vortex from 72°N to the outermost return point at 86°N. High backscatter ratios are indicators for the presence of large particle surfaces and high depolarization ratios suggest the presence of solid particles. The co-located measurements of the particle backscatter ratio and depolarization ratio thus allow for a cloud classification into different PSC types.

5.2 Classification of the PSC measured on 22 January 2016

Figure 4 shows the joint occurrence histogram of the inverse backscatter ratio 1/R and particle depolarization for the PSC measurements in Figure 3. The histogram bin size is 0.02×0.02 and the color scale indicates the number of cloud observations (4 km horizontal by 100 m vertical) falling within each bin. Overlaid are the regions which correspond to different PSC types following Pitts et al. (2011) with two modifications. First the sub-classification of Mix2 into Mix2-enhanced and normal Mix2 is dropped. Second, the $1/R_{ice}$ threshold for ice and NAT regions is set to 0.3. Without the latter change, a substantial part of the branch connecting STS and fully developed ice clouds would have been counted as NAT

Mix2 instead of ice, while ice is the more obvious interpretation of the lower branch in the ice class given the form of the joint histogram. The $1/R_{ice}$ threshold for ice and NAT regions is discussed in more details below.

We find two modes in the joint histogram of the ice class, a mode with high particle depolarization ratio (high-depol ice) connected to the NAT Mix2 regime and a mode with lower particle depolarization ratio at the same backscatter ratio (low-depol ice) connected to the STS class. Given several ten thousands individual measurement points within the ice class, these two different ice modes can clearly be distinguished. The NAT Mix2 regime with enhanced NAT particle concentrations is strongly populated. In contrast, no observation falls into the wave ice class forming in strong mountain waves.

We use the classification by Pitts et al. (2011) and the threshold of $1/R_{ice} = 0.3$ of ice versus NAT Mix2 to classify the PSC from Figure 3. A large ice PSC (red areas in Figure 5) extends in a cloud layer with a vertical thickness up to 6 km between 18 and 24 km altitude over distances of over 1400 km. The ice PSC is measured twice for more than 1.6 h during the outbound flight leg between 10:34 and 12:30 and during the inbound flight leg between 13:50 and 15:32 UTC. The synoptic-scale ice PSC is observed mainly at temperatures below T_{ice} as indicated by the dashed contour lines. Here T_{ice} is calculated based on temperature data from the integrated forecast system IFS (cycle 41r1) of ECMWF and measured water vapor mixing ratios. In the northern part of the cloud between 12:30 and 13:50 UTC, the ice PSC is surrounded by NAT Mix2 cloud layers (yellow area) extending above and next to the ice PSC to the North-East. Within the ice PSC, NAT can be masked by the higher optical signal from ice, but may be present. The southernmost part of the PSC observed between 10:15 and 10:34 and again between 15:15 and 16:00 UTC consists of non-depolarizing liquid STS droplets (blue region). The STS layer extends towards the South and below the ice layer.

The ice, NAT Mix and STS occurrence histogram of this PSC is plotted versus temperature difference to T_{ice} in Figure 6. The peak in ice occurrence is located 0.5 K below T_{ice}. The peak in NAT Mix2 (NAT Mix1) occurrence appears at 0.5 K (3 K) above T_{ice}, while STS occurrence peaks 0.5 to 1.5 K above T_{ice}. The occurrence of ice at temperatures below T_{ice} and of NAT Mix and STS at temperatures above T_{ice} is consistent with their expected temperature range, see also Pitts et al. (2013). The $1/R_{ice}$ threshold of 0.3 best matches the PSC measurements in the low-depol ice mode (see Figure 4). Also, the PSC area classified as ice agrees best with the ice area derived from numerical weather predictions. For this threshold, the peak of the NAT Mix2 occurrence is located above T_{ice}, while it is located near and below T_{ice} for $1/R_{ice}$ of 0.2. Therefore we use the threshold $1/R_{ice}$ of 0.3 for the WALES observations on 22 January 2016. The analysis of CALIPSO measurements throughout the winter 2015/16 with changing HNO_3 and H_2O concentrations requires a variable $1/R_{ice}$ threshold as used by Pitts et al. (2018). A detailed investigation of the $1/R_{ice}$ threshold is given in the supplementary material S1.

The PSC was also measured by CALIPSO on 22 January 2016 between 10:42 and 10:47 UTC. Figure 7 shows the CALIPSO curtain plot measured at the CALIPSO footprint near the HALO flight track shown in Figure 8. The classification by Pitts et al. (2018) is used to determine the PSC types from CALIPSO. PSC observations by WALES on the slower flying HALO aircraft lead to the time differences of up to 2 h between the collocated PSC measurements. Similar to the WALES observations, the CALIOP lidar detected an ice PSC between 18 and 24 km altitude with a horizontal extension of about 1200 km. An STS layer is located below the ice cloud and predominantly NAT Mix1 PSCs are measured to the North-West.

Few NAT Mix2 PSCs were measured by CALIPSO, mainly located at the edge of the ice cloud. Differences in the flight paths of HALO and the CALIPSO footprint explain the lower number of NAT Mix2 PSCs observed by CALIPSO compared to WALES. WALES measured NAT Mix2 mainly north-west of the CALIPSO footprint. Also a fraction of the low-depol ice mode was located predominantly north-west of the CALIPSO PSC curtain and therefore merely covered by CALIPSO.

6 Discussion of PSC formation

We now discuss PSC formation in two steps based on the WALES lidar and CALIPSO observations combined with trajectory analysis. First we investigate the formation of the ice, NAT and STS layers of the PSC observed on 22 January 2016 using 8 days back-trajectories starting in the different PSC layers at 21.5 km altitude every 2 minutes. Then we separate the ice PSC into a mode with high particle depolarization (high-depol ice mode) and a mode with low depolarization (low-depol ice mode) as derived from the backscatter ratio to depolarization occurrence histogram in Figure 4 and discuss ice formation pathways of each ice mode based on domain filling Lagrangian back-trajectory calculations and their matches with PSC measurements by CALIPSO up to 5 days prior to the ice. We start with the discussion of the formation of the ice, NAT and STS layers of the 22 January 2016 PSC.

6.1 Temperature history of trajectories starting in ice, NAT Mix2 and STS PSC layers

The trajectory calculations were performed using the Hybrid Single-Particle Lagrangian Integrated Trajectory dispersion model HYSPLIT (Draxler, 1998) with operational forecast of the deterministic IFS (cycle 41r1 interpolated to $0.25 \times 0.25°$) from ECMWF at 3-hourly time steps as meteorological input. We calculate 8 days backward-trajectories starting in the PSC observations every 2 min along the HALO flight track (corresponding to approximately 24 km horizontal spacing) at 21.5 km altitude (near 30 hPa) to investigate PSC formation. We further calculate T_{NAT} based on Hanson and Mauersberger (1998) for an altitude dependent climatological HNO_3 profile and use H_2O, T_{ice} and the ice saturation ratio S_{ice} from the meteorological data set to discuss ice nucleation pathways. Further we perform sensitivity studies every 2 minutes at higher and lower altitudes (19, 21 and 23 km) to account for particle sedimentation, which is not included in the simplified trajectory calculations shown here. As a rough estimate, a 10-µm sized ice crystal sediments about 1 km in a day (Fahey et al., 2001). For aspherical particles the sedimentation rates are even lower (Woiwode et al., 2014; Weigel et al., 2014; Woiwode et al., 2016), hence generally our simplified calculations cover the altitude range of sedimenting PSC particles.

In Figure 8 we show temperatures of backward-trajectories starting every 2 min in the PSC at 21.5 km altitude at the WALES cross section of the outbound flight leg to the North (corresponding to the section from 10:30 to 12:34 in Figure 5). The air mass trajectories circulate around the pole within these 8 days, the innermost trajectories (NAT Mix2 at the observation) stay near the Arctic cold pool and remain cold (T < 190 K) for 8 days while the outermost trajectories (STS at the time of observation) encounter higher temperatures (T > 198 K) outside of the cold pool. The innermost NAT Mix2 trajectory 4 bypasses Greenland and circulates within the vortex core, while the more southern trajectories are slowly lifted

above Greenland. The slow lift contributes to the synoptic cooling within the Arctic cold pool and induces a transient low amplitude mountain wave temperature disturbance in the lee of the Greenlandic island (dark blue areas in Figure 8). The synoptic cooling induces the formation of a large scale ice PSC above and to the East of Greenland within the cold pool of the polar vortex. Teitelbaum et al. (2003) previously investigated the important role of synoptic scale dynamics for PSC formation.

To investigate PSC formation, 4 trajectories representative for STS starting at 10:38 UTC (label 1, circle, blue line in panel B and C), low-depol ice at 11:36 UTC (label 2, filled diamond, red line), in high-depol ice at 12:14 UTC (labels 3, open diamond, orange line) and NAT Mix2 at 12:34 (label 4, square, green line) are selected to represent the temperature histories of the back trajectories of the different PSC types. The numbering of the trajectories is in chronological order during the outbound flight leg and the color coding in panel (B) and (C) corresponds to the PSC type classification in Figure 5, with a blue line for STS, a green line for NAT Mix2 and red line for low-depol ice at the time of the WALES PSC observation. The orange line is used to distinguish the trajectory starting in the high-depol ice mode from the low-depol ice mode (red line). Small symbols denote 48-h time steps along the 8 days backward trajectories.

Starting with the outermost STS trajectory, its temperature varies between $T_{NAT}-2K < T < T_{NAT}+11K$ for 6.5 days as the trajectories circulate mainly out of the Arctic cold pool. Then, the temperature decreases below T_{NAT} ~36 h and below T_{ice} ~16 h prior to the observation before increasing to $T_{ice} +2K$ at the time of STS detection. The temperature at the time of observation is within the STS temperature range.

Slightly further to the North, the low-depol ice mode back trajectory (red line) is located at the edge of the cold pool, with temperatures oscillating between $T_{NAT}-6K < T < T_{NAT}+6K$ for 7 days. The temperatures decrease below T_{NAT} ~36 h and finally below T_{ice} ~16 h prior to the observation. At the time of observation of the low-depol ice mode, the trajectories' temperatures are near and below T_{ice} explaining the detection of ice PSCs. The ice saturation ratio S_{ice} increases to 1.4 and varies between 1.4 and 0.92 in the last 16 h prior to observation. While $S_{ice} < 1.4$ does not allow for homogeneous ice nucleation (Koop et al., 2002; Murphy and Koop, 2006), heterogeneous nucleation (e.g. Hoose and Möhler, 2012) could explain ice formation in this case. Thereby meteoric material (Czicio et al., 2001; Voigt et al., 2005; Curtius et al., 2005) probabaly embedded in STS could serve as ice nuclei. This has been suggested by Engel et al. (2013) to explain the observation of an Arctic ice PSC in 2010. Small scale temperature fluctuations might be present in our case, but cannot be resolved in the meteorological model.

In contrast, the northernmost NAT Mix2 trajectory (4, green line) circulates within the cold pool at temperatures $T_{NAT}-8K < T < T_{NAT}-4K$ for 8 days without passing over Greenland. The temperatures remain above T_{ice} throughout that period and $T > T_{ice}$ at the time of observations, consistent with NAT. Homogeneous NAT nucleation rates (Knopf et al., 2002) are too low to support homogeneous NAT nucleation in this case. In contrast, NAT nucleation rates on meteoric dust (Voigt et al., 2005; Grooß et al., 2005; Hoyle et al., 2013; Grooß et al., 2014) may explain the formation of NAT within few days and are consistent with the observations of NAT.

Similarly, the trajectory of the high-depol ice mode (3, orange line) circulates within the polar vortex for 8 days at temperatures between T_{NAT}-10K $< T < T_{NAT}$. However, when passing over Greenland, the high-depol ice trajectories decrease below T_{ice} ~10 h prior to the observation. The temperature is T_{ice}-2K at the time of observation of high-depol ice and S_{ice} increases up to 1.2. This temperature history supports heterogeneous ice nucleation. Laboratory measurements of heterogeneous ice nucleation (Hoose and Möhler; 2012) were performed for a suite of different ice nuclei at temperatures down to 213 K, hence above PSC formation temperatures. Meteoric dust could potentially serve as ice nuclei in the stratosphere at S_{ice}~1.2, however surface area densities of meteoritic material are significantly smaller than those of NAT PSCs, therefore we investigate the possibility of ice nucleation on NAT and NAT serving as ice nuclei.

The temperature history of the high-depol ice trajectory also remains below the existence temperature of NAD (T_{NAD} ~ T_{NAT}-2.3 K, Voigt et al., 2005) for about a week. The formation of NAD from binary nitric acid water solutions has been observed in the laboratory (e.g. Knopf et al., 2002; Wagner et al., 2005; Stetzer et al., 2006), but homogeneous nucleation rates are too low to explain denitrification (Knopf et al., 2002). Also, pseudo-heterogeneous nucleation rates (Knopf, 2006) cannot explain atmospheric nitric acid hydrate particle number densities. Möhler et al. (2006) suggest that ambient supersaturations with respect to NAD of 8 to 9 are required over days to explain NAD particle nucleation at number densities, which can be detected with current instrumentations. Furthermore, a phase change from NAD into NAT or a nucleation of NAT on NAD under atmospheric conditions is not supported by laboratory experiments (Tizek et al., 2004; Stetzer et al., 2006). Observational evidence for the presence of NAD in the atmosphere is missing so far, e.g. Höpfner et al. (2006) found no spectroscopic evidence for the presence of NAD from MIPAS observations of PSCs over Antarctica. Therefore, we focus the discussion on ice nucleation on NAT in our study.

6.2 Ice nucleation pathways

The existence of the two ice modes with high and low particle depolarization ratios (Figure 4) in combination with the trajectory analysis suggests two different ice nucleation pathways. (1) Ice nucleation in STS may account for the low-depol ice mode. Inclusions of meteoric material in STS as suggested by Engel et al. (2013) are below the detection limit of the WALES lidar but may be present. (2) A second ice nucleation pathway is required to explain the formation of the high-depol ice mode. The long time period below T_{NAT} prior to the observation and the similarity of the NAT Mix2 and the high-depol ice trajectories (except for the last hours where the high-depol ice trajectories are ice supersaturated) points to ice nucleation on pre-existing NAT particles.

The hypothesis of ice nucleation on NAT ice is investigated in more detail using a Lagrangian match approach of 5 days back trajectories matched with CALIPSO PSC measurements similar to the method outlined by Santee et al. (2002). Domain filling trajectories starting in the WALES ice PSC observations are calculated every 2 minutes along the flight path from 10:30 UTC until 16:00 UTC (corresponding to ~24 km distance) at 15 different altitude levels with 500 m spacing between 17 and 24 km. Thus a total of 2490 trajectories are derived. We match the Lagrangian trajectories with all CALIPSO PSC curtain plots measured north of 60°N within 5 days prior to 22 January 2016. Pitts et al. (2018) is used to classify the PSC

types within the CALIPSO curtains. A match is defined as a point along a trajectory lying within a given horizontal distance limit from one of the CALIOP footprints at the same time. To yield a statistically significant number of matches we set a distance limit of 100 km. A threshold of this order is justified by the limited accuracy of the trajectory calculations, too. In addition, to make the spatial resolution comparable with the initial spacing of the trajectories' starting points, 15 CALIPSO

5   PSC data points with 3×180 m vertical and 5×5 km horizontal distance are grouped to smooth variations in the PSC type classification in nearby data points at the edge of PSC layers. A match point is classified as ice, NAT or STS, if > 50% of the CALIPSO pixel array is classified as the respective PSC type. The match method is selected to investigate whether a different PSC type was present before the high-depol and low-depol ice PSC observation by WALES on 22 January 2016. If so, the analysis indicates, which PSC type has been detected prior to high-depol or low-depol ice.

10   The trajectories are classified using the following criteria: (1) NAT has been measured in the CALIPSO curtain–trajectory match point prior to ice measured by WALES, or prior to the last CALIPSO match with ice on the same trajectory, (2) STS has been observed in the CALIPSO curtain–trajectory match point prior to ice, (3) no PSC (nil) has been observed in the match point prior to ice, or (4) no match (nom) has occurred along the trajectories.

The result of the match analysis is presented in Figure 9 overlaid on the ice PSC contour of the 22 January PSC. The result is

15   evident: NAT has been measured by CALIPSO on all match points directly prior to the high-depol ice mode. Thus the CALIPSO-trajectory match analysis delivers strong evidence for ice nucleation on NAT. Ice could have nucleated on pre-existing NAT particles, which could also explain the high particle depolarization ratio of the ice mode. In contrast, predominantly STS or no PSC has been observed on the match points prior to the low-depol ice mode. We note the possibility that due to fast evaporation times, STS might have existed along the trajectories classified is nil, but was not

20   present at the time of the CALIPSO overpass. Only few trajectories have no match points with CALIPSO.

Summarized, we find strong evidence for ice nucleation on NAT. Ice nucleated on NAT may produce ice particles with higher particle depolarization ratios. In contrast, the low-depol ice mode can be explained by ice nucleation in STS, possibly with solid meteoric inclusions as suggested by Engel et al. (2013). Ice nucleation in STS may lead to lower particle depolarization ratios of the ice mode. These two ice formation pathways would be consistent with the CALPSO and

25   WALES observations combined with results from domain filling trajectory analysis.

7 Conclusions and outlook

Extremely low temperatures existed in the Arctic stratospheric winter 2015/2016 because low planetary wave activity resulted in a stable vortex. Synoptic-scale ice PSCs formed in late December 2015 and persisted throughout January as observed by the CALIOP lidar onboard the CALIPSO satellite. The sedimentation of the ice PSC particles led to significant

30   dehydration (Koshrawi et al., 2017). From January to early March 2016, water vapor data MLS data show severe dehydration between 400 and 500 K potential temperatures (Manney and Lawrence, 2016). In addition, ice PSCs can serve as efficient transporters for nitric acid incorporated into ice. Large ice particles sediment faster than smaller NAT particles

and therefore ice can lead to efficient denitrification at PSC altitudes. Massive denitrification has been measured by MLS with an onset in mid-December 2015 throughout the Arctic winter (Manney and Lawrence, 2016; Khosrawi et al., 2017).

We use high resolution WALES lidar measurements of a large-scale ice PSC observed on 22 January 2016 in combination with domain filling back-trajectories matched to CALIPSO PSC observations to investigate ice nucleation pathways. Two

5 distinct modes with high and low particle depolarization in the ice PSC occurrence histogram suggest different particle size distributions and shapes in the two modes, possibly linked to two ice formation pathways. Ice nucleation in STS, possibly with meteoric inclusions as suggested by Engel et al. (2013), may lead to the ice mode with low particle depolarization ratios. In addition, NAT has been detected by CALIPSO on trajectory match points prior to high-depol ice. Thus, ice formation on NAT could lead to the ice mode with higher particle depolarization ratios. While the low-depol STS-ice branch

10 is frequently populated in space-borne CALIOP lidar data, the high-depol NAT Mix2-ice branch less frequently observed by CALIOP in Arctic or Antarctic PSC measurements (Pitts et al., 2013; Pitts et al., 2018). Larger noise in the satellite data with CALIPSO travelling at an orbit near 700 km and a higher detection limit of the CALIOP lidar data compared to the WALES measurements on aircraft cruising at 14 km altitude may contribute to this difference.

Lambert et al. (2012) investigate PSC occurrence in the Antarctic in early winter from 11 to 30 June 2008 based on CALIOP

15 and MLS observations. They detect a NAT Mix2 type PSC linked to the high-depol ice regime. They also note that the conservative estimate by Pitts et al. (2013) of the threshold $1/R_{ice} = 0.2$ may lead to a miss-classification of the ice PSC with respect to the NAT MIX2 regime. The threshold of 0.3 provides a better fit to their observation. Similarly to the Arctic case, the NAT Mix2 trajectories remain 3 to 5 K below T_{NAT} for 5 days, suggesting heterogeneous NAT nucleation. Ice formation is not investigated in detail for this Antarctic PSC, although ice is present. The low-depol ice mode is sparsely populated

20 supporting the hypothesis of ice nucleation on NAT in the Antarctic June 2008 PSC. Further high resolution lidar observations are required to estimate the global occurrence of the high-depol ice mode. These observations could then help to evaluate the importance of the suggested NAT-ice nucleation pathway in other regions of the atmosphere. In addition, combined lidar and in-situ observations from aircraft (e.g. Fahey et al., 2001; Northway et al., 2002) could help to answer the question on the abundance of the NAT Mix2, which might include aspherical NAT particles as observed by Molleker et al.

25 (2014) and Woiwode et al. (2016).

The NAT crystal has a stoichiometry of 1 $HNO_3 \times 3$ H_2O molecules and exists in $\alpha-$NAT or $\beta-$NAT crystal structure (Iannarelli and Rossi, 2015). Compared to liquid aerosol or meteoric particles, the NAT crystal structure is more similar to that of ice, and therefore NAT readily nucleates on ice as frequently observed in laboratory experiments (Hanson and Mauersberger, 1988; Iannarelli and Rossi, 2015; Gao et al., 2016; Weiss et al., 2016) and in mountain wave ice PSCs

30 (Carslaw et al., 2002; Fueglistaler et al., 2002; Luo et al., 2003; Voigt et al., 2003). Vice versa, we propose here that ice may nucleate on NAT. This pathway has been suggested in early PSC studies (e.g. Peter, 1997) and has later been neglected. Here, we find evidence for ice nucleation on NAT and elucidate its importance in the polar lower stratosphere. Direct laboratory measurements of the nucleation rate of ice on NAT are missing (Koop et al., 1997) and would be required to provide further evidence of the suggested ice nucleation process.

Ice nucleation on NAT may be important in the tropical tropopause region, where the existence of a NAT belt and cirrus has been detected by in-situ measurements (Popp et al., 2007, Voigt et al., 2008). Also CALIPSO measurements showed indications for the existence of NAT in the tropics (Chepfer and Noel, 2009), although Pitts et al. (2009b) argue that mixtures of liquid aerosols and thin cirrus instead might have been misinterpreted as NAT-like particles in the tropics. Based on the observational data set, the ice nucleation rate on NAT could be parameterized and implemented in a large scale model in order to assess the global relevance the different ice nucleation pathways in different regions of the atmosphere.

Data availability: The observational data are available at https://halo-db.pa.op.dlr.de. Operational meteorological analysis are achieved in the MARS archive at ECMWF (https://www.ecmwf.int/en/forecasts/documentation-and-support/changes-ecmwf-model/ifs-documentation), ERA-Interim data were taken from https://www.ecmwf.int/en/research/climate-reanalysis/era-interim.

Author contributions: CV performed the scientific study and wrote the paper. AD and BE analyzed the meteorological conditions for the Arctic winter 2016. MW and SG performed the HALO lidar measurements and data evaluation. MP and LP evaluated CALIPSO data. RB made the HYSPLIT trajectory calculations. BMS, WW and HO were the coordinators of the POLSTRACC campaign. All authors contributed to the manuscript. The authors declare that they have no conflict of interest.

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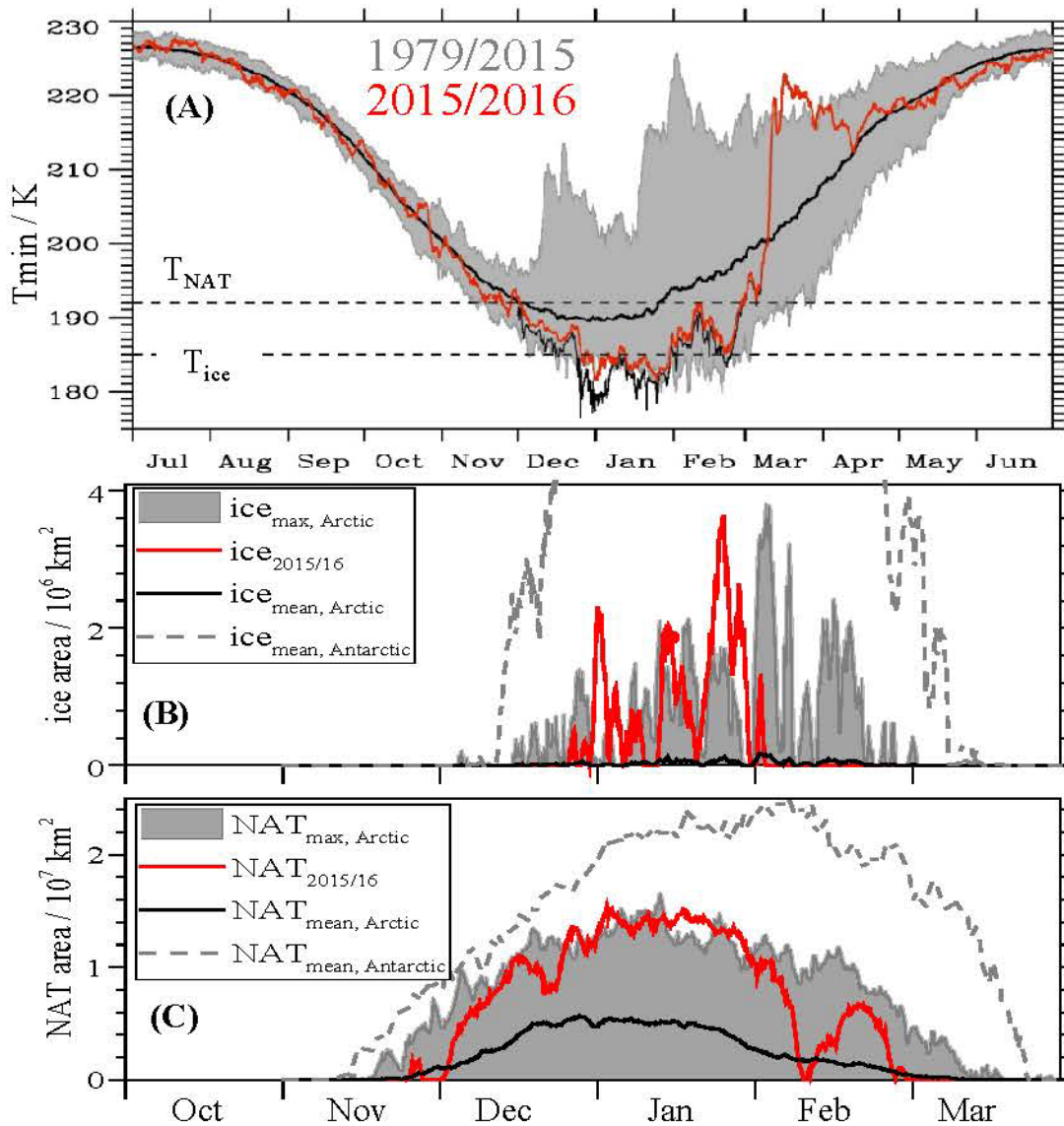
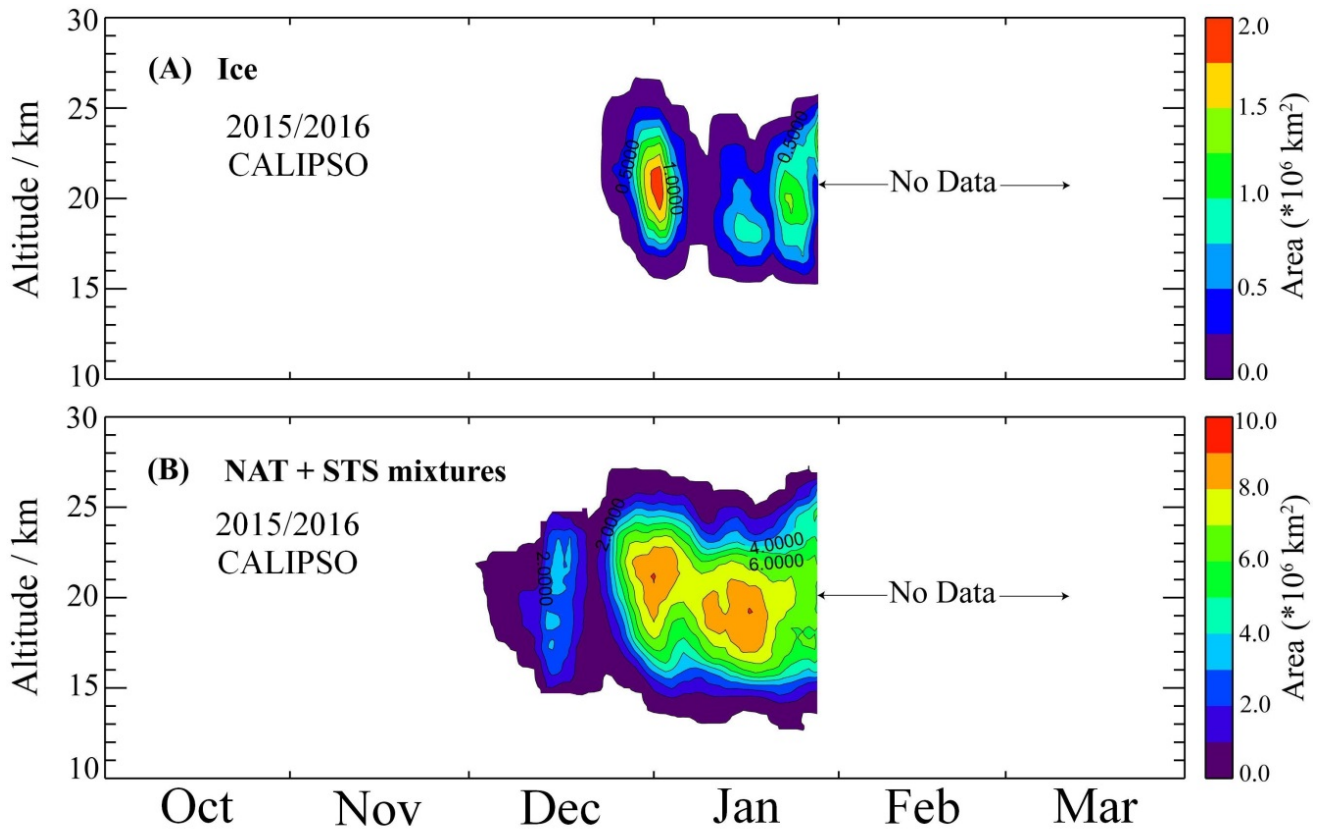


Figure 1: Temperature evolution of the 2015/2016 Arctic winter stratosphere (A) 6-hourly ECMWF ERA interim reanalysis (Dee et al. 2011) data retrieved at a horizontal resolution of 1°: minimum temperature T_{min} (K) between 65° and 90°N at the 30-hPa pressure surface. The thick black line denotes the mean values of T_{min} averaged during 1979 to 2015, and the shaded areas encompass the minimum and maximum values of T_{min} between 1979 and 2015. The red line marks the evolution of T_{min} from operational analyses of the IFS cycle 41r1 until 8 Mar 2016. The thin black line indicates T_{min} from the IFS cycle 41r2 in the preoperational phase from 1 December 2015 to 8 Mar 2016 retrieved at a resolution of 0.125°. After 8 March 2016, the black line continues as red curve of the operational IFS cycle 41r2. T_{ice} (Murphy and Koop, 2005) and T_{NAT} (Hanson and Mauersberger, 1988) are calculated using 4.6 ppmv H_2O and 7 ppbv HNO_3, relevant for the 2015/2016 Arctic vortex conditions (Manney and Lawrence, 2016). (B) Evolution of the vortex area with temperatures below T_{ice}. The black line marks the mean area below T_{ice} (A_{ice}) at 30 hPa pressure (~21.6 km) between 1979 and 2015. The gray shading indicates maximum and minimum A_{ice} in the same time period. The red line shows the evolution of A_{ice} at the 30 hPa pressure surface in the Arctic winter 2015/2016 of the IFS cycle 41r2. The gray dashed line gives the mean area below T_{ice} at 30 hPa south of 65°S from 1979 to 2015, shifted by 6 months to account for seasonality. (C) Same data for NAT.

5 **Figure 2: Curtain plot of the areal occurrence of PSCs in the winter 2015/2016 detected by the CALIOP lidar** onboard the CALIPSO satellite using the classification from Pitts et al. (2018). (A) Curtain plot of ice PSC occurrence. (B) Curtain plot of NAT and STS mixtures PSC occurrence. Ice PSCs were observed by CALIOP from end of December 2015 till end of January 2016.

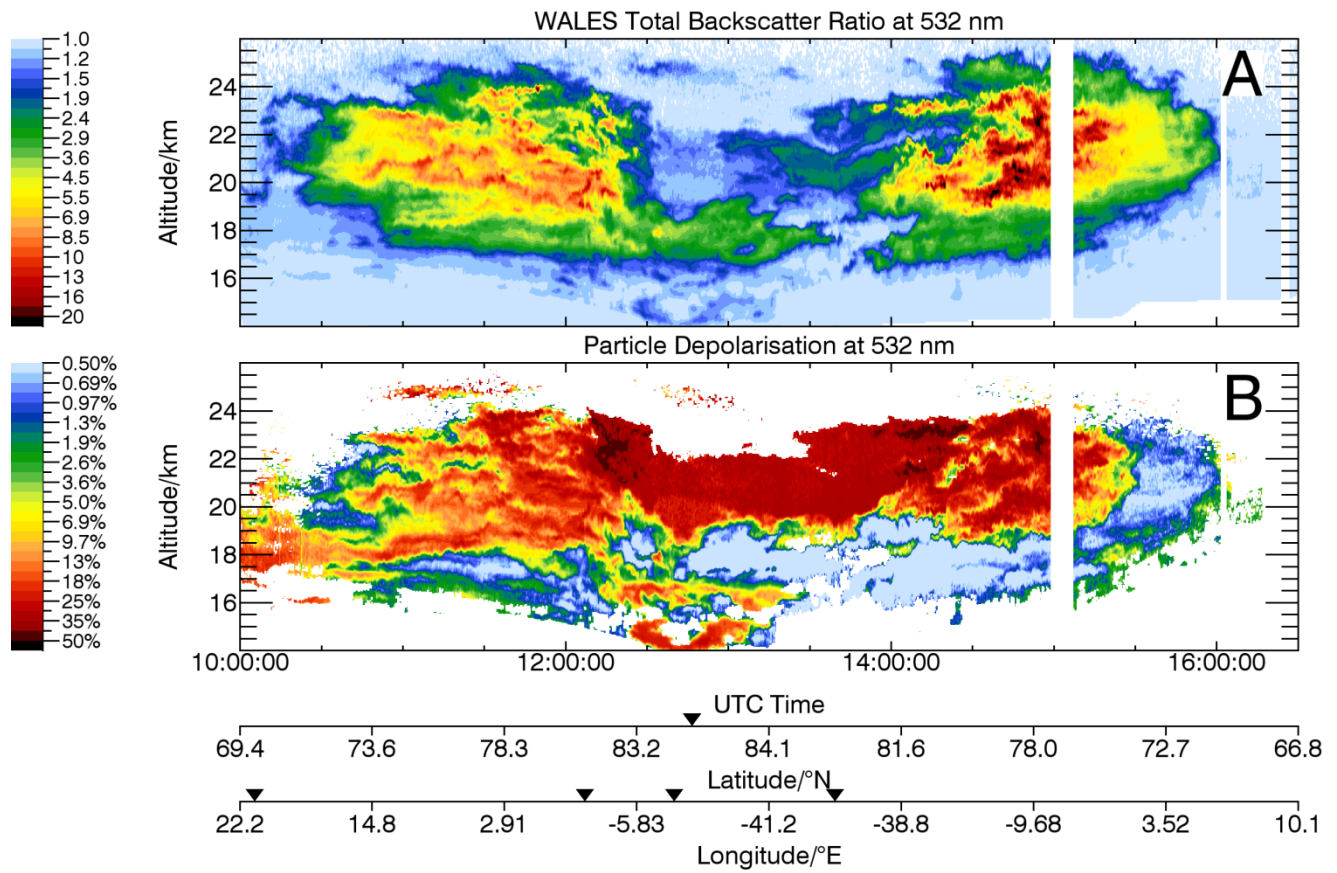



Figure 3 Lidar observation of a synoptic-scale polar stratospheric cloud observed on 22 January 2016 (A) Backscatter ratios from the WALES lidar (Wirth et al., 2009) at 532 nm wavelengths during a HALO flight into the Arctic vortex and (B) particle depolarization. The time of the flight, as well as latitude and longitude of the HALO flight path are indicated. Turning points of the HALO are marked by triangles.

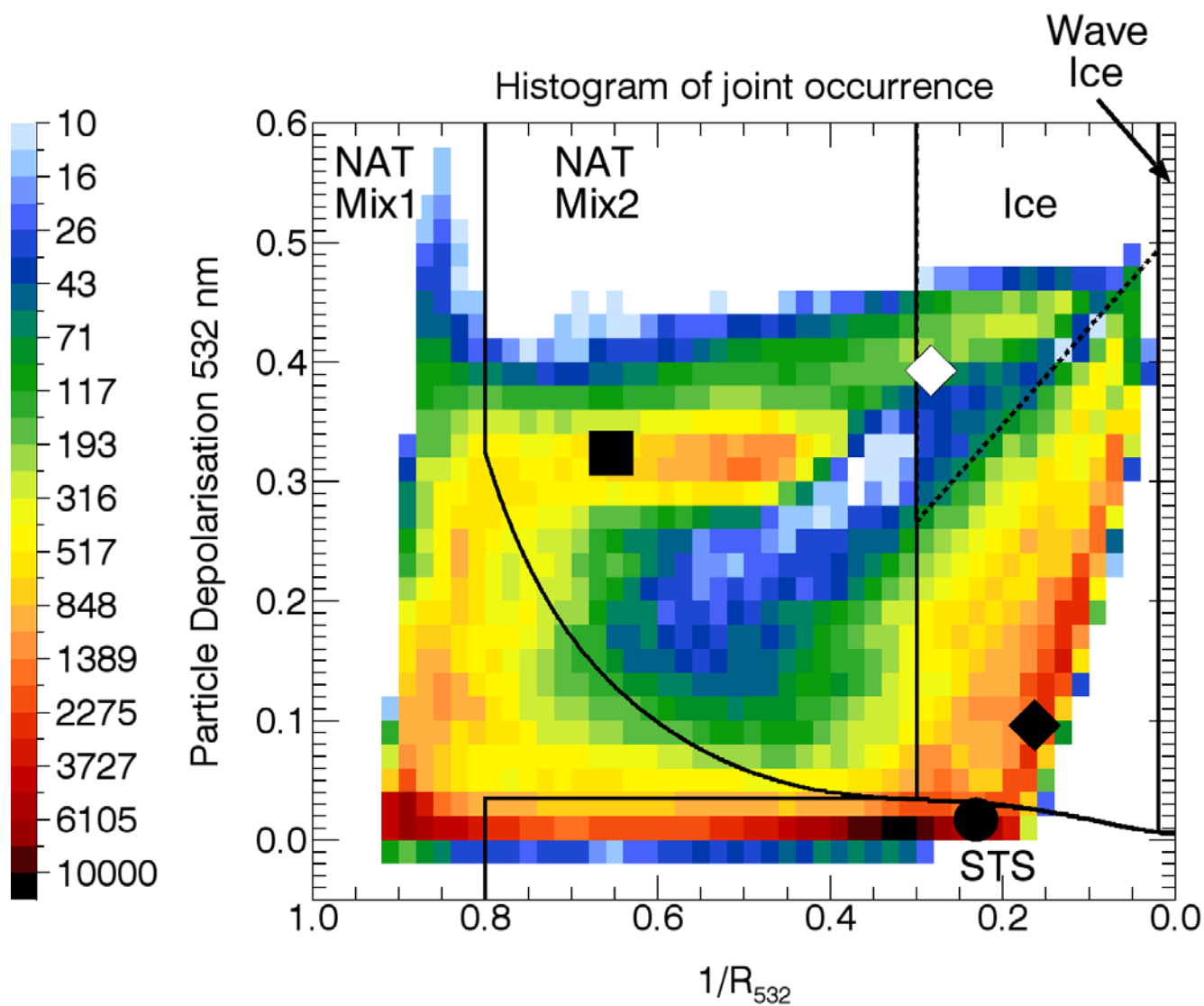



Figure 4 Composite 2-dimensional occurrence histogram of the PSC observed on 22 Janaury 2016 shown in Figure 3 in the 1/backscatter ratio versus particle depolarization ratio coordinate system. The solid black lines denote the boundaries of the PSC types defined by Pitts et al. (2011) with the threshold between ice and NAT Mix2 $1/R_{ice}= 0.3$. The dotted line separates the high-depol and low-depol ice modes. The symbols indicate the starting points of trajectories within different PSC types (square: NAT Mix2, hollow diamond: high-depol ice, filled diamond: low-depol ice and circle: STS).

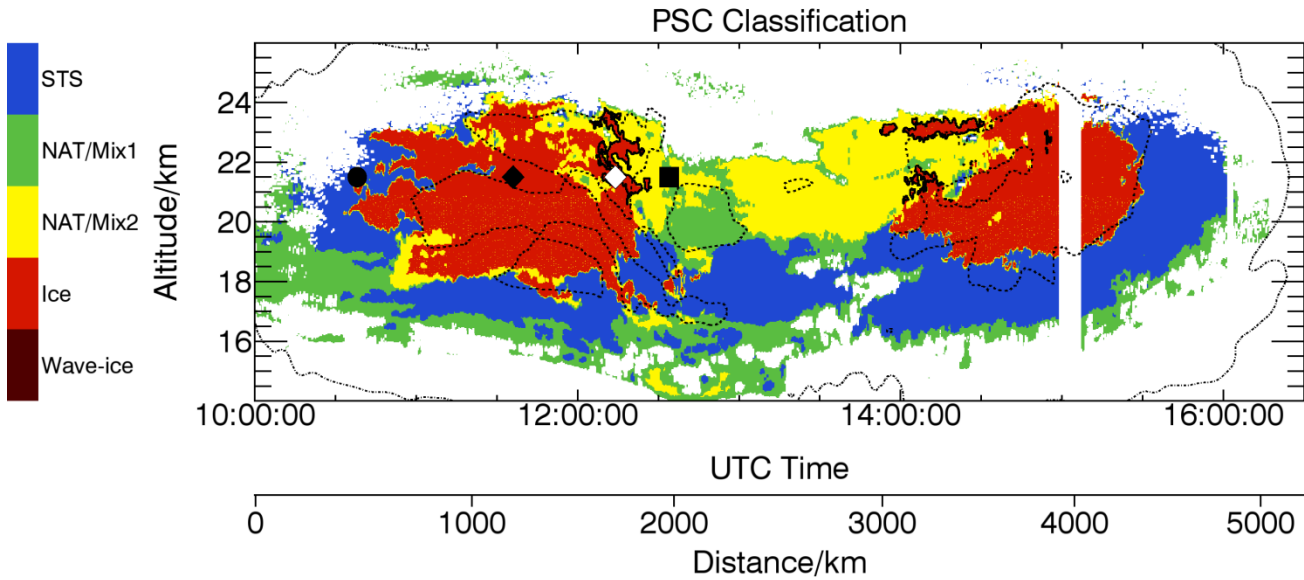



Figure 5 Classification of the synoptic-scale polar stratospheric ice cloud on 22 January 2016 using the classification given in Figure 3. A synoptic ice PSC (red area) extends over several hundred km above the HALO flight track. The thick black line encloses the low-depol ice mode (see Figure 4). To the North-West, the ice PSC is embedded in NAT layers (yellow: NAT Mix2, green: NAT Mix1). A liquid STS layer (blue) is located below and to the South-East of the ice PSC. The wave-ice class is not populated. The ending points of the trajectories for ice (high-depol ice (hollow diamond) and low-depol ice (filled diamond)), NAT (square) and STS (circle) given in Figure 8 are marked. The dashed line shows the T_{ice} contour and the dotted line the T_{NAT} contour lines derived from 6-hourly IFS operational weather analysis (cycle 41r2) interpolated to 1-hourly time steps using meteorological forecast data and the water vapor field measured by WALES as well as the HNO_3 field from MLS.

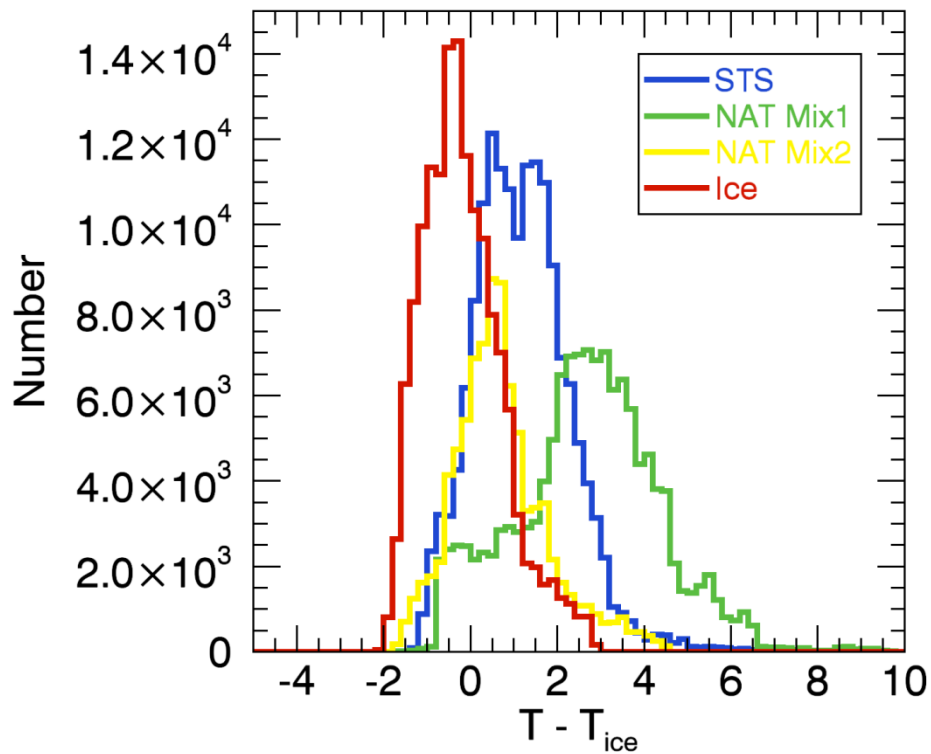



Figure 6 Occurrence histogram of Ice, NAT Mix2, NAT Mix1 and STS of the 22 January 2016 PSC (Figure 5) versus temperature difference to T$_{ice}$.

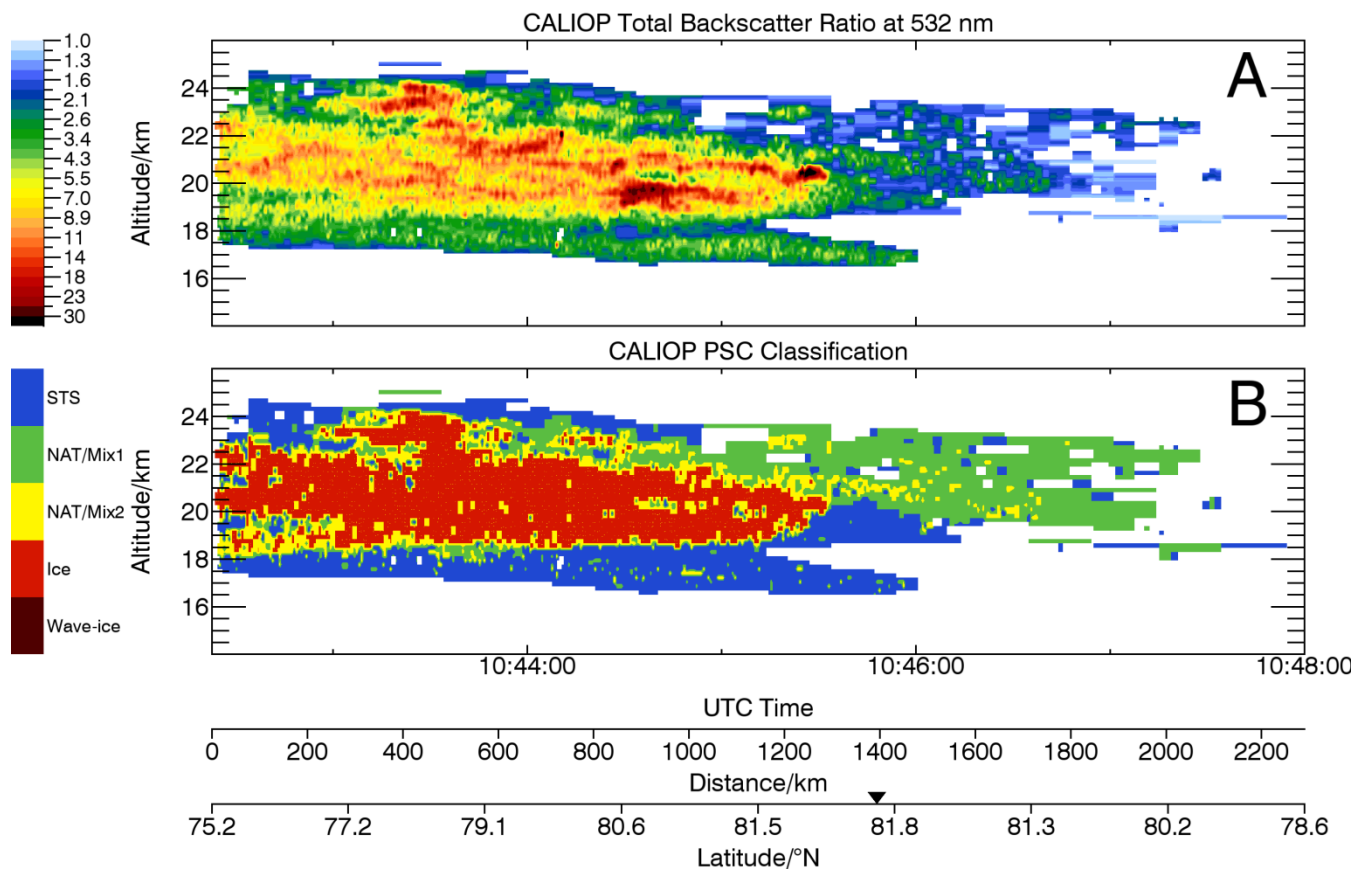



Figure 7 CALIPSO PSC measurements on 22 January 2016 near the HALO flight path (A) Backscatter ratios from CALIPSO at 532 nm wavelengths and (B) PSC type classification from Pitts et al. (2018). The time of the overpass between 10:42 and 10:47 UTC, as well as the horizontal distance and the latitude of the CALIPSO footprint are shown. The northernmost turning point of the CALIPSO footprint is marked by the triangle. The CALIPSO footprint is shown in Figure 8.

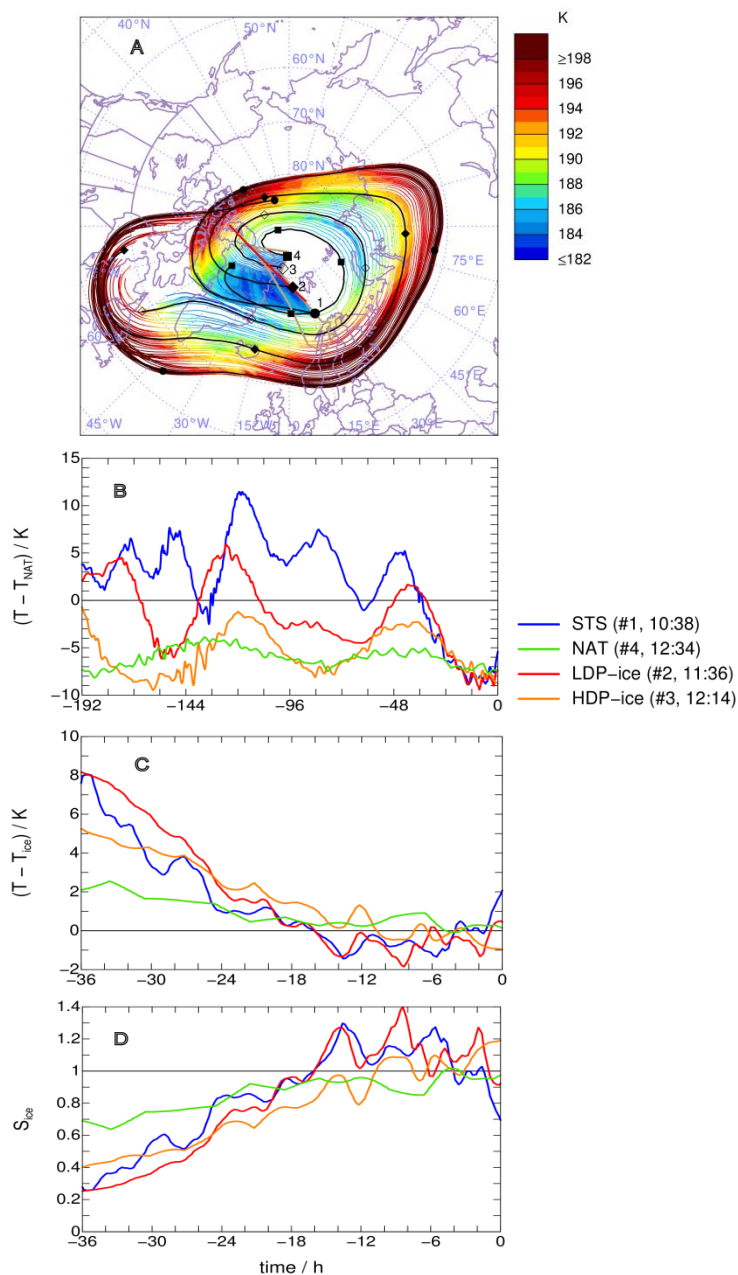



Figure 8 8-days back-trajectories starting in the different PSC layers measured on 22 January 2016 (A) Temperatures (colour coded) of the back-trajectories derived from the IFS operational analysis (cycle 41r1) starting every 2 min at 21.5 altitude in the PSC on 22 January 2016. Four back-trajectories starting in STS (1, circle), low-depol ice (2, LDP-ice, filled diamond), high-depol ice (3, HDP-ice, open diamond) and NAT Mix2 (4, square) are shown as black lines with small symbols marking every 48 hours. The HALO flight track is shown as brown line and the CALIPSO foot print as red line. The starting points of the 4 trajectories are also given in Figures 4 and 5. The temperature evolution along the 4 trajectories (STS: blue, NAT: green, LDP-ice: red and HDP-ice: orange) is shown in panel (B) as temperature difference to T_{NAT} for 8 days and (C) as the temperature difference to T_{ice} for 30 h prior to the observation. The ice saturation ratio S_{ice} is given in panel D. T_{NAT} is calculated from Hanson and Mauersberger (1988) for altitude dependent climatological HNO_3 profile, H_2O, T_{ice} and S_{ice} is calculated from the meteorological data set.

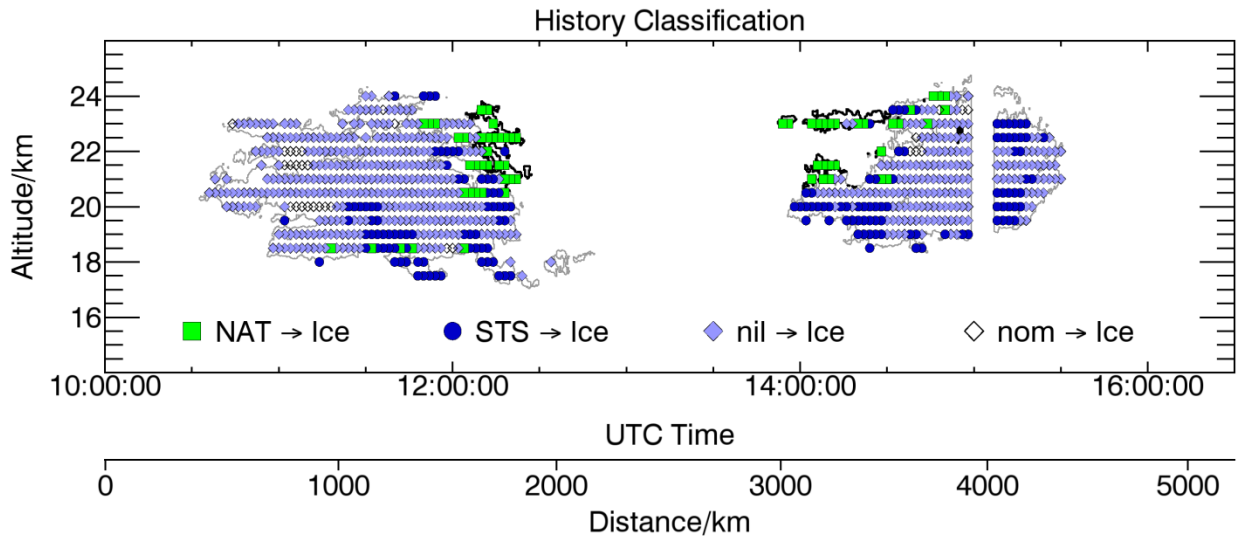



Figure 9 PSC type measured by CALIPSO prior to ice on match points with 5 days back-trajectories. Green squares indicate that NAT has been measured by CALIPSO prior to ice, blue circles indicate STS prior to ice, light blue diamonds (nil) indicate that no PSC has been observed and (nom) that no match point of the trajectories and the CALIPSO curtain exist within 5 days. NAT has been detected by CALIPSO prior to the high-depol ice mode measured by WALES (thick black contour line) and predominantly STS or no PSC have been detected prior to the low-depol ice mode measured by WALES (thin gray contour line). UTC times of the WALES ice PSC observations and the horizontal distance are given.