

Volcanic ash modeling with the on-line NMMB-MONARCH-ASH v1.0 model: model description, case simulation and evaluation

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Abstract

Traditionally, tephra transport and dispersal models have evolved decoupled (off-line) from numerical weather prediction models. There is a concern that inconsistencies and shortcomings associated to this coupling strategy might lead to errors in the ash cloud forecast. Despite this concern, and the significant progress to improve the accuracy of tephra dispersal models in the aftermath of the 2010 Eyjafjallajökull and 2011 Cordón Caulle eruptions, to date, no operational on-line dispersal model is available to forecast volcanic ash. Here, we describe and evaluate NMMB-MONARCH-ASH, a new on-line multiscale meteorological and transport model that attempts to pioneer the forecast of volcanic aerosols at operational level. The model forecasts volcanic ash cloud trajectories, concentration of ash at relevant flight levels, and the expected deposit thickness for both regional and global configurations. Its on-line coupling approach improves the current state-of-the-art of tephra dispersal models, especially in situations where meteorological conditions are changing rapidly in time, two-way feedbacks are significant, or distal ash cloud dispersal simulations are required. This work presents the model application for the first phases of the 2011 Cordón Caulle and 2001 Mt. Etna eruptions. The computational efficiency of NMMB-MONARCH-ASH and its application results compare favorably with other long-range tephra dispersal models, supporting its operational implementation.

Keywords: volcanic ash, on-line coupling, transport-meteorological modeling, operational forecast, NWPM, TTDM, Cordón Caulle, Mt. Etna.

1 **1 Introduction**

2 Explosive volcanic eruptions can eject large quantities of particulate matter (tephra) that, along with other
3 aerosol droplets and trace gases, are carried upwards into the atmosphere by the buoyant eruption column and
4 then dispersed by winds aloft (e.g. Sparks et al., 1997). Tephra particles smaller than 2 mm in diameter,
5 technically defined as volcanic ash (Schmid, 1981), can spread over large distances away from the source
6 forming ash clouds that jeopardize air-traffic (Casadevall, 1993), airports (Guffanti et al., 2009) and, for very
7 large eruptions, alter both atmospheric composition and chemistry (Myhre et al., 2013; Self, 2006). Tephra
8 Transport and Dispersal Models (TTDMs, e.g. Folch, 2012) are used to simulate the atmospheric transport,
9 dispersion and ground deposition of tephra, and to generate operational short-term forecasts to support civil
10 aviation and emergency management. The recent eruptions of Eyjafjallajökull (Iceland) in 2010 and Cordón
11 Caulle (Chile) in 2011 have reinforced the importance of tephra dispersal in the context of global aviation safety.
12 In addition to short-term forecast, other model applications include the reconstruction of past events, studying
13 the impact of volcano eruptions on climate, probabilistic tephra hazard assessments or simulation of recent
14 eruptions for model evaluation purposes. For any of those cases, TTDMs require a driving Numerical Weather
15 Prediction Model (NWPM) or a meteorological reanalysis dataset for the description of the atmospheric
16 conditions, and an emission or source model for the characterization of the eruption column (Fig. 1).

17 Traditionally, TTDMs have evolved decoupled (off-line) from NWPMs. In the off-line strategy, the
18 meteorological driver runs *a priori* and independently of the TTDM to produce the required meteorological
19 fields at regular time intervals. Meteorological data is then furnished to the TTDM, which commonly assumes
20 constant values for the meteorological fields during each time interval or, at most, performs a linear interpolation
21 in time. Although the off-line approach is operationally advantageous, there is a concern that it can lead to a
22 number of accuracy issues (e.g. inaccurate handling of atmospheric processes) and limitations (e.g. neglect of
23 feedback effects) that can be corrected by on-line approaches (Grell et al., 2004). These inconsistencies are
24 especially important when meteorological conditions change rapidly in time or for long-range transport.
25 However, uncertainties arising from off-line systems have received little attention, even if the experience from
26 other communities (e.g. air quality) highlights the importance of coupling on-line dispersal and meteorological
27 models (e.g. Baklanov et al., 2014; Grell and Baklanov, 2011). To date, only the Weather Research and
28 Forecasting model coupled to Chemistry (WRF-Chem; Grell et al., 2005) includes a coupled functionality that
29 allows simulating emission, transport, dispersion, transformation and sedimentation of pollutants released during
30 volcanic activities (Stuefer et al., 2013).

31 In this paper we describe and evaluate NMMB-MONARCH-ASH, a new on-line meteorological and
32 atmospheric transport model to simulate the emission, transport and deposition of ash (tephra) particles released
33 from volcanic eruptions. The model predicts ash cloud trajectories, concentration of ash at relevant flight levels,
34 and the expected deposit thickness for both regional and global domains. The novel on-line coupling in NMMB-
35 MONARCH-ASH allows solving both the meteorological and aerosol transport concurrently and interactively at
36 every time-step. This coupling strategy aims at improving the current state-of-the-art of tephra dispersal models,
37 especially in situations where meteorological conditions are changing rapidly in time, two-way feedbacks are
38 significant, or distal ash cloud dispersal simulations are required. The model builds on the Multiscale Online
39 Nonhydrostatic Atmosphere Chemistry model (NMMB-MONARCH; formerly known as NMMB/BSC-CTM)

1 (Badia et al., 2017; Jorba et al., 2012; Pérez et al., 2011) to represent the transport of volcanic particles. Its
2 meteorological core, the Non-hydrostatic Multiscale Model on a B grid (NMMB; Janjic and Black, 2007; Janjic
3 and Gall, 2012; Janjic, 2005; Janjic et al., 2011), allows for nested global-regional atmospheric simulations by
4 using consistent physics and dynamics formulations. The final objective in developing NMMB-MONARCH-
5 ASH is two-fold. On one hand, at a research level, we aim at studying the differences between the on-line/off-
6 line modeling strategies. Moreover, a second version of the model is projected to quantify the feedback effects of
7 dense volcanic ash clouds from large explosive eruptions on the radiative budget and local meteorology. On the
8 other hand, at an operational level, the low computational cost of the NMMB dynamic core presented in this
9 work suggests that NMMB-MONARCH-ASH could be applied for more accurate on-line operational forecasting
10 of volcanic ash clouds. Consequently, the focus on developing an on-line volcanic ash model is timely.

11 The remainder of the manuscript is arranged as follows: Section 2 summarizes the modeling background and the
12 standard physical schemes employed in NMMB-MONARCH-ASH; Section 3 provides a comprehensive
13 description of the ash related modules, including details about the emission, transport, and deposition of volcanic
14 particles; Section 4 validates the regional and global configurations of the model for the 2001 Mt. Etna and 2011
15 Cerdón Caulle long-lasting eruptions; Section 5 discusses the implementation and performance of the model for
16 its operational use and; finally, Section 6 provides a summary and conclusion of this work.

17

18 **2 Modeling background**

19 NMMB-MONARCH-ASH is a novel on-line multi-scale meteorological and atmospheric transport model
20 developed at the Barcelona Supercomputing Center (BSC). The model attempts to pioneer the forecast of
21 volcanic aerosols by embedding a series of new modules on the BSC's operational system for short/mid-term
22 chemical weather forecasts (NMMB-MONARCH) developed at the BSC in collaboration with the U.S National
23 Centers for Environmental Prediction (NCEP) and the NASA Goddard Institute for Space Studies. The
24 development of the volcanic ash module follows the implementation of the mineral dust (Pérez et al., 2011) and
25 sea-salt (Spada et al., 2013) modules in NMMB-MONARCH, and allows for a range of different physical
26 parameterizations for research and operational use. The system allows for feedback processes among gases,
27 aerosol particles and radiation, and includes a gas-phase module to simulate tropospheric gas-phase chemistry
28 (Badia et al., 2017; Jorba et al., 2012).

29 Its meteorological core, the Non-hydrostatic Multiscale Model on the B grid (NMMB), is a fully compressible
30 meteorological model with a non-hydrostatic option that allows for nested global-regional atmospheric
31 simulations by using consistent physics and dynamics formulations. The standard physical and numerical
32 schemes employed in NMMB are summarized in Table 1. The non-hydrostatic dynamics were designed to avoid
33 over-specification. The cost of the extra non-hydrostatic dynamics is about 20% of the cost of the hydrostatic
34 part, both in terms of computer time and memory (Janjic, 2001, 2003). The numerical schemes for the
35 hydrostatic and nonhydrostatic options available in the NMMB dynamic solver were designed following the
36 principles found in Janjic (1977) and developed and modified thereafter (Janjic, 1979, 1984, 2003), and are
37 summarized in Janjic and Gall (2012). The Arakawa B-grid horizontal staggering is applied in the horizontal
38 coordinate employing a rotated latitude-longitude coordinate for regional domains and latitude-longitude

1 coordinate (Janjic, 2003) with polar filtering for global domains. Rotated latitude-longitude grids are employed
2 for regional simulations in order to obtain more uniform grid distances. In this particular case, the Equator of the
3 rotated system runs through the middle of the integration domain, reducing the longitudinal grid-size as the
4 southern and the northern boundaries of the integration domain are approached (Janjic and Gall, 2012). In the
5 vertical, the Lorenz staggering vertical grid is used with a hybrid sigma-pressure coordinate. The general time
6 integration philosophy in NMMB uses explicit schemes when possible for accuracy, computational efficiency
7 and coding transparency (e.g., horizontal advection), and implicit for very fast processes that would otherwise
8 require a restrictively short time-step for numerical stability with explicit differencing (e.g., vertical advection
9 and diffusion, vertically propagating sound waves). The NMMB model became the North American Mesoscale
10 (NAM) operational meteorological model in October of 2011, and it has been computationally robust, efficient
11 and reliable in operational applications and pre-operational tests since then. In high-resolution NWP
12 applications, the efficiency of the model significantly exceeds those of several established state-of-the-art non-
13 hydrostatic models (e.g. Janjic and Gall, 2012).

15 **3 The volcanic ash module**

16 The ash module is embedded within the NMMB meteorological model and solves the mass balance equation for
17 volcanic ash taking into account: i) the characterization of the source term (emissions); ii) the transport of
18 volcanic particles (advection/diffusion); and iii) the particle removal mechanisms (sedimentation/deposition).
19 The coupling strategy of this module can be turned on or off, depending on the solution required (on-line vs. off-
20 line). The on-line version of the model solves both the meteorological and aerosol transport concurrently and
21 consistently (on-line coupling). This strategy allows the particle transport to be automatically tied to the model
22 resolution time and space scales, resulting in a more realistic representation of the meteorological conditions. In
23 contrast, the off-line approach uses an “effective wind field” in which, meteorological conditions (e.g. wind
24 velocity, mid-layer pressure, etc.) are set to constant, and are only updated at specific coupling intervals (i.e. time
25 for which meteorological fluctuations are not explicitly resolved). This strategy replicates the off-line coupling
26 effect of traditional dispersal models used at operational levels (e.g. coupling intervals of 1h or 6h). The
27 conservativeness of the model is evaluated to ensure that the ash transport scheme is consistent with the mass
28 conservation equation.

29 **3.1 Source term**

30 Explosive volcanic eruptions release large amounts of particles into the atmosphere. These particles, commonly
31 known as tephra, mix with ambient air to form an eruption column or volcanic plume. To forecast the ash cloud
32 movement and provide actual ashfall concentrations, tephra dispersal models require a complete characterization
33 of the parameters describing the source term. These parameters are generally referred to as Eruption Source
34 Parameters (ESPs) and include the eruption start and duration, column height, mass eruption rate (MER), vertical
35 distribution of mass and the particle grain size distribution (GSD). ESPs vary not only from one eruption to
36 another, but also during the different eruptive phases of a single event.

1 Typically, the eruption starting time, duration and column height are inferred/constrained from visual or satellite
 2 observations. However, other parameters such GSD, MER, or the vertical distribution of mass in the column are
 3 not available in real-time and must be inferred from previous events of similar characteristics (e.g. Mastin et al.,
 4 2009). Uncertainties in source parameter values are a key factor limiting the accuracy of ash-cloud model
 5 forecasts (Bonadonna et al., 2015a). The characterization of each ESP in NMMB-MONARCH-ASH is described
 6 in the following subsections.

7 **3.1.1 Mass eruption rate**

8 The Mass Eruption Rate (MER) gives the mass released by unit of time and defines the eruption intensity. Its
 9 characterization in NMMB-MONARCH-ASH is achieved by employing a series of empirical correlations
 10 between (observed) column height and eruption rate, which, according to plume similarity theory, scales roughly
 11 as the 4th power of height. Because of this strong dependence, uncertainties within 20% in the determination of
 12 column height can translate into uncertainties up to 70% for the MER (e.g., Biass and Bonadonna, 2011).
 13 Averaged column heights of eruptions that have not been directly observed are typically derived from
 14 characteristics of tephra deposits (e.g. Bonadonna and Costa, 2013; Carey and Sparks, 1986; Pyle, 1989), or
 15 derived from model inversion (e.g. Connor and Connor, 2006; Marti et al., 2016; Pfeiffer et al., 2005).

16 The empirical correlations to estimate MER in the model are described in Table 2, and are based either on fitting
 17 observations (e.g. Mastin et al., 2009), or more sophisticated fits accounting for wind bent-over effects (e.g.
 18 Degruyter and Bonadonna, 2012; Woodhouse et al., 2013). In addition, MER can also be derived using a more
 19 sophisticated 1-D plume model (see Sect. 3.1.5).

20 **3.1.2 Vertical distribution of mass**

21 The vertical distribution of mass in the column at the vent location is key when representing the plume,
 22 especially if wind shear exists with elevation at the volcano (Lin, 2012). To determine this distribution of mass,
 23 NMMB-MONARCH-ASH allows for the following geometrical distributions: i) point source, where mass is
 24 released as a single source point at a certain height above the vent, H_{plume} ; ii) top-hat, where mass is released
 25 along a umbrella-type slab of user-defined thickness, and iii) the so-called Suzuki distribution (Suzuki, 1983;
 26 Pfeiffer et al., 2005), which assumes a more complex vertical distribution of mass release along the eruption
 27 column;

28

$$S = MER \left\{ \left(1 - \frac{z}{H_{plume}} \right) \exp \left[A \left(\frac{z}{H_{plume}} - 1 \right) \right] \right\}^\lambda \quad (4)$$

29 where, S is the mass per unit of time (kg/s) released at a given height z above the vent, MER is the total mass
 30 eruption rate, H_{plume} is the column height above the vent, A and λ are the so-called Suzuki parameters. The
 31 parameter A dictates the height of the maximum particle release (concentration), whereas λ controls how closely
 32 mass distributes around this maximum. Any of the 3 options above can be combined independently with the

1 different options for MER estimation. In NMMB-MONARCH-ASH, the terrain following hybrid sigma-pressure
2 vertical levels of the model must be converted to elevations for each model integration time-step in order to
3 interpolate *MER* from the discrete source points into the nodes of the model grid.

4 **3.1.3 Grain size distribution**

5 The impact of explosive volcanic eruptions on climate and air traffic strongly depends on the concentration and
6 grain size distribution (GSD) of pyroclastic fragments injected into the atmosphere (e.g. Girault et al., 2014).
7 Grain size distribution is normally reconstructed by volcanologists from grain size data at individual outcrops,
8 ranging from basic unweighted average of the GSD at individual sparse outcrops, to various integration methods
9 of grain size data (e.g. Rose and Durant, 2009). The particle grain size distribution in NMMB-MONARCH-ASH
10 is specified through an input file, which defines the particle bin properties (bin mass fraction, diameter, density
11 and shape factor). In volcanology, grain size distributions are given in terms of the Φ , defined as $d = 2^{-\Phi}$,
12 where d is the particle diameter in mm. The granulometry file in the model can be furnished by the user
13 (typically derived from field data) or generated by an external utility program which produces Gaussian and Bi-
14 Gaussian distributions in Φ (log-normal in diameter d) (Costa et al., 2016; Folch et al., 2009).

15 **3.1.4 Particle aggregation**

16 The total grain size distribution (TGSD) erupted at the vent can be altered in case of particle aggregation, which
17 dramatically impacts particle transport dynamics thereby reducing the atmospheric residence time of aggregating
18 particles and promoting the premature fallout of fine ash. For computational purposes, particle aggregation in
19 NMMB-MONARCH-ASH is assumed to take place mainly in the eruption column, where particle concentration
20 and water contents are higher (the subsequent formation of aggregates downstream in the ash cloud under the
21 appropriate atmospheric conditions is not contemplated by the model). The model considers aggregates as
22 another particle class (bin), introduced as a standard source term by either solving: i) a series of simple analytical
23 expressions based on field observations or, ii) a more sophisticated wet aggregation model originally proposed
24 by Costa et al. (2010).

25 The analytical expressions available in the model modify the user-given particle grain size distribution by
26 assuming that a certain mass fraction of each granulometric class forms a new aggregate class added to the
27 TGSD. Despite the obvious limitations (obviates the physics of aggregation processes), these field-based
28 simplistic approaches are advantageous in that only the source term has to be modified in order to account for
29 aggregation. Table 3 provides an overview of these options. In addition to these empirical aggregation schemes,
30 NMMB-MONARCH-ASH also includes a wet aggregation model originally proposed by Costa et al. (2010).
31 This option allows for wet aggregation in the column providing an intermediate solution between the
32 unaffordable all-size class approach and the empirical solutions presented before. The model is based on a
33 solution of the classical Smoluchowski equation, obtained by introducing a similarity variable and a fractal
34 relationship for the number of primary particles in an aggregate. It also considers three different mechanisms for
35 particle collision: Brownian motion, ambient fluid shear, and differential sedimentation. Table 4 provides an
36 overview of the governing equations of this wet aggregation model.

1 **3.1.5 FPlume model**

2 A more sophisticated approach to obtain MER and the mass distribution in the column from the conditions at the
3 vent consists of solving a 1-D radially averaged BPT model for mass, momentum, and energy. These 1-D plume
4 models are more useful in operational roles and broad exploratory investigations (Costa et al., 2015; Devenish et
5 al., 2012). For that reason, NMMB-MONARCH-ASH is coupled with the 1-D FPlume model (Folch et al.,
6 2015); a 1-D cross-section averaged plume model which accounts for plume bent over, entrainment of ambient
7 moisture, effects of water phase changes on the energy budget, particle fallout and re-entrainment by turbulent
8 eddies, as well as variable entrainment coefficients fitted from experiments. The model also accounts for particle
9 aggregation in presence of liquid water or ice that depends on column dynamics, particle properties, and amount
10 of liquid water and ice existing in the column (Folch et al., 2010). This allows the plume model to predict an
11 “effective” grain size distribution depleted in fines with respect to that erupted at the vent. For a complete
12 definition of the governing equations of FPlume, refer to Folch et al. (2015). FPlume has two solving strategies
13 where the model: i) solves directly for column height for a given MER; or ii) solves iteratively for MER for a
14 given height. For any case, the following inputs need to be provided to the ash input file in NMMB-
15 MONARCH-ASH: eruption start and duration, vent coordinates and elevation, conditions at the vent (exit
16 velocity, temperature, magmatic water mass fraction, and total grain size distribution) and total column height or
17 mass eruption rate.

18 **3.2 Particle advection/diffusion**

19 Transport of volcanic ash by advection and turbulent diffusion is analogous to those of atmospheric tracers (e.g.
20 moisture) transport (Janjic et al., 2009) in NMMB. Tracer advection is Eulerian, positive-definite and
21 monotonic. The Adams-Bashforth scheme is used for horizontal advection and the Crank-Nicolson scheme for
22 vertical advection. For the horizontal diffusion, the model uses a second order scheme with two types of
23 parameterized dissipative processes: explicit lateral diffusion (often called horizontal diffusion, a 2nd order
24 nonlinear Smagorinsky-type approach; Janjic, 1990) and horizontal divergence damping (Janjic and Gall, 2012).

25 Plumes from high-intensity eruptions can be injected high into the stratosphere, reaching a maximum column
26 height and intruding laterally at the neutral buoyancy level (NBL) as a gravity current (Sparks et al., 1997). This
27 current can spread at velocities exceeding those of the surrounding winds, affecting tephra transport and
28 deposition near the source. As larger particles are removed by deposition and air is entrained, the plume density
29 decreases and momentum reduces such that, at a certain distance, atmospheric turbulence and wind advection
30 become the dominant atmospheric transport mechanisms (Baines and Sparks, 2005). Neglecting the gravitational
31 spreading of the umbrella cloud in tephra dispersal simulations could misrepresent the interaction of the volcanic
32 ash cloud and the atmospheric wind field for high-intensity eruptions and for proximal deposition of tephra
33 (Mastin et al., 2014). To account for the gravity-driven transport, NMMB-MONARCH-ASH is coupled with the
34 model of Costa et al. (2013) describing cloud spreading as a gravity current. This parameterization calculates an
35 effective radial velocity of the umbrella spreading as a function of time or cloud radius. The effective radial
36 velocity of the umbrella spreading is then combined with the wind field velocity centered above the vent in the
37 umbrella region to calculate the contribution of the gravitational spreading to the total cloud spreading. To
38 estimate the radial distance at which the critical transition between gravity-driven and passive transport occurs,

1 the umbrella front velocity is compared with the mean wind velocity at the NBL estimating the Richardson
2 number. Table 5 provides an overview of the governing equations of the gravity current model embedded in
3 NMMB-MONARCH-ASH.

4 **3.3 Particle sedimentation and dry deposition**

5 Particle sedimentation in NMMB-MONARCH-ASH is governed by the terminal velocity of settling particles.
6 This fall velocity is sensitive to particle size and atmospheric conditions, determining the residence time of ash
7 particles in the atmosphere. NMMB-MONARCH assumes that the settling velocities of aerosols (mineral dust,
8 sea salt, etc.) follow the Stokes law for spherical particles corrected by the Cunningham slip factor. The Stokes
9 law applies to the creeping or Stokes flow regime, in which the drag force is proportional to particle velocity,
10 and holds only for Reynolds numbers $Re \leq 0.1$. This regime is justified for small particles and aerosols ($< 20 \mu\text{m}$).
11 However, calculating fallout times based on settling according to Stokes Law is less adequate for coarse ash ($>$
12 $64 \mu\text{m}$), which sediments much faster. In addition, ash particles are not spherical, which complicates and further
13 slows fallout. In order to simulate properly a wider spectrum of particle sizes, NMMB-MONARCH-ASH adds a
14 new sedimentation module that covers the turbulent regime ($Re \geq 1000$) in which the drag force is proportional to
15 the square of the particle velocity. In this case, the gravitational particle settling velocity, v_s (in $\text{m} \cdot \text{s}^{-1}$), can be
16 expressed as:

17

$$v_s = \sqrt{\frac{4g(\rho_p - \rho_a)d}{3C_d\rho_a}} \quad (13)$$

18 where, ρ_a and ρ_p denote air and particle density, respectively, d is the particle equivalent diameter, and C_d is the
19 drag coefficient (depending on the Reynolds number). Strictly, the expression above is valid for spherical
20 particles in the turbulent regime but it is often generalized to the whole range of Re numbers and particle shapes
21 by defining the drag coefficient properly. Table 6 provides an overview of the different settling velocity models
22 available in NMMB-MONARCH-ASH, each relying on different empirical evaluations of drag coefficient.

23 Dry deposition, acting at the bottom layer of the model, is a complex process depending on physical and
24 chemical properties of the particle, the underlying surface characteristics and micro-meteorological conditions.
25 Dry deposition in the model is based on that originally proposed by Zhang et al. (2001). This parameterization
26 has been updated to account for the different settling velocities available for volcanic particles - Eq. (13). The
27 dry deposition velocity in the model, v_d (in $\text{m} \cdot \text{s}^{-1}$), is given by:

28

$$v_d = v_s + \frac{1}{(R_a - R_s)} \quad (18)$$

1 where, R_a is the aerodynamic resistance of the particle, and R_s is the surface resistance (both in $s \cdot m^{-1}$). These
2 terms take into account all the effects of the lowermost layer of the atmosphere, such as turbulence (R_a) and
3 Brownian diffusion, impaction and interception (R_s). It is worth mentioning that, for most of its resident time,
4 airborne volcanic ash lies above the near-surface atmospheric layers, where gravitation dominates, implying that,
5 in most cases, dry deposition has little influence on model results.

6 **3.4 Mass conservation**

7 Mass conservation is a critical requirement for any atmospheric transport algorithm. Non-conservative schemes
8 can significantly underestimate or overestimate concentrations, especially for long time integrations, in which it
9 is critical that the tracer advection scheme is consistent with the mass continuity equation (Jöckel et al., 2001).
10 Most mesoscale meteorological models use observation/analyzed fields or global model results as initial
11 conditions, and therefore they are not very sensitive to slowly accumulated mass inconsistencies as re-
12 initializations remove accumulations. However, dispersal models are usually very sensitive to mass
13 inconsistencies set in previous simulations or spin-up fields as initial conditions, thereby accumulating mass
14 inconsistencies. In addition to mass conservation, monotonicity and prevention of non-physical under and
15 overshoots in the solution are also a highly desirable characteristics in transport schemes (Rood, 1987). For these
16 reasons, the model includes a conservative, positive definite (i.e. tracer is a positive scalar) and monotone (i.e.
17 entirely increasing) Eulerian scheme for advection. The positive definiteness in the model is guaranteed by
18 advecting the square root of the tracer using a modified Adams-Bashforth scheme for the horizontal direction
19 and a Crank-Nicolson scheme for the vertical direction. The conservation of the tracer is achieved as a result of
20 the conservation of quadratic quantities by the advection scheme. Monotonization is applied *a posteriori* to
21 eliminate new extrema (Janjic et al., 2009). The conservative nature of NMMB-MONARCH-ASH is evaluated
22 by calculating the mass flux at the boundaries (for regional domains) of the computational domain, the airborne
23 mass, and the mass deposited on the ground to verify mass conservation at each time-step (e.g. $< 0.5\%$ mass
24 creation for a 30 day simulation).

25 **3.5 Numerical performance**

26 The high computational efficiency of the NMMB meteorological driver allows for the application of
27 nonhydrostatic dynamics at a global scale (Janjic et al., 2009), and supports that NMMB-MONARCH-ASH
28 could be used in an operational forecast of volcanic ash clouds. Model parallelization is based on the well-
29 established Message Passing Interface (MPI) library. The computational domain is decomposed into sub-
30 domains of nearly equal size in order to balance the computational load, where each processor is in charge to
31 solve the model equations in one sub-domain. The Eulerian schemes in the model require relatively narrow and
32 constant width halos (i.e. data points from the computational domain of neighboring sub-domains that are
33 replicated locally for computational convenience), which simplify and reduce communications.

34 To measure the time-to-solution required, we compute the parallel speed-up (computation speed) of the model;
35 that is, the performance gains of parallel processing in comparison to serial processing:

36

$$S_{(P)} = \frac{t_{(P=1)}}{t_{(P)}} \quad (19)$$

where S is the computed speed-up value, and t is the simulation run-time employing P processors instead of running it serially ($P = 1$).

To evaluate the efficiency of the model while using the computational resources, the parallel efficiency of the model is computed by looking at the ratio between the parallel speed-up over P :

$$E_{(P)} = \frac{S_{(P)}}{P} \quad (20)$$

Parallel efficiency is used as a metric to determine how far the model's speed-up is from the ideal. If the speed-up is ideal, the efficiency is 1, regardless of how many cores the program is running on. If the speed-up is less than ideal, the efficiency is less than 1.

4 Simulations and validation

The forecast skills of NMMB-MONARCH-ASH have been tested for several well-characterized eruptions, including the Pinatubo 1991 (Philippines), Etna 2001 (Italy), Chaitén 2008 (Chile) or Cordón Caulle 2011 (Chile) eruptions (e.g. Marti et al., 2013, 2014). Here, we present two applications of the model for the ash dispersal forecast of weak and strong long lasting eruptions. Section 4.1 summarizes the results of the regional and global simulations for the first days of the 2011 Cordón Caulle eruption. This event represents a suitable case study of strong long-lasting eruptions with changing winds, which is useful to evaluate the advantages of the on-line approach for operational forecast. In a parallel effort, Sect. 4.2 summarizes the results from the regional configuration of the model for the 2001 Etna eruption. This eruption is a good example of a weak, long-lasting eruption, useful when evaluating the sedimentation mechanisms of the model against well-characterized tephra deposits.

4.1 The 2011 Cordón Caulle eruption

The 2011 Cordón Caulle eruption was a typical mid-latitude Central and South Andean eruption, where dominating winds carried ash clouds over the Andes causing abundant ash fallout across the Argentine Patagonia. Besides the significant regional impacts on agriculture, livestock and water distribution systems, this eruption stranded thousands of passengers due to air traffic disruptions in the southern hemisphere, thereby causing important economic losses to airlines and society (e.g. Raga et al., 2013; Wilson et al., 2013). This event is evidence of the global nature of the volcanic ash dispersion phenomena and highlights the need for accurate real-time forecasts of ash clouds.

1 The Cordón Caulle volcanic complex (Chile, 40.5° S, 72.2° W, vent height 1420 m a.s.l.) reawakened on 4 June
2 2011 around 18:30 UTC after decades of quiescence. The initial explosive phase spanned more than two weeks,
3 generating ash clouds that dispersed over the Andes. The climactic phase (~27 h) (Jay et al., 2014) was
4 associated with a ~9 km (a.s.l.) high column (Osorio et al., 2014). For the period between 4 - 14 June, numerous
5 flights were disrupted in Paraguay, Uruguay, Chile, southern Argentina and Brazil. The two major airports
6 serving Buenos Aires and the international airport in Montevideo, Uruguay, were closed for several days, along
7 with airports in Patagonia (Wilson et al., 2013). A detailed chronology of the eruption can be found in Collini et
8 al. (2013) and Elissondo et al. (2016), the stratigraphy and characteristics of the resulting fallout deposit are
9 described in Pistolesi et al. (2015) and Bonadonna et al., (2015b), and a summary of the environmental impacts
10 of the eruption is discussed in Raga et al. (2013) and Wilson et al. (2013).

11 Here, we describe the synoptic meteorological situation during the first two weeks of eruptive activity (Fig. 2),
12 and give a brief chronology of the events in order to compare them with the predictions of the model. The
13 eruption developed as a long-lasting rhyolitic activity with plume heights above the vent between 9-10 km high
14 a.s.l. (4-6 June), 4 and 9 km during the following week (7-14 June) and < 6 km after 14 June (Global Volcanism
15 Program, GVP, <http://www.volcano.si.edu>; Siebert et al. 2010). The first major episode, on 4 June (18:45 UTC),
16 resulted in an ash cloud (9-10 km) that reached the Chile-Argentina border within the hour of the eruption. On
17 June 5, E-SE winds drove the plume to the Atlantic Ocean (1800 away from the source), leaving a large area of
18 Argentina territory affected by ash fall. On June 6, the plume changed its direction abruptly toward N-NE,
19 reaching the northern regions of the Argentine Patagonia, and then shifted direction again towards SE, threatening
20 the Buenos Aires air space. On June 7, a second episode resulted in a plume (4-9 km) dispersing ash further to
21 the north of Argentina leading to a more recognizable shift of winds over the E-SE. On June 8, the volcanic
22 cloud (9-10 km a.s.l.) dispersed towards NE with a bend toward SE 400 km from the source. On June 9, the
23 plume had a NE direction reaching the city of Buenos Aires and the northern boundary of Paraguay following a
24 frontal zone passing through Patagonia. This resulted in major air traffic disruption at the two international
25 airports that service the city: Aeroparque (AEP) and Ezeiza (EZE), which remained closed intermittently during
26 the following 15 days. Later during the day, the wind turned SE dispersing ash over Uruguay, Brazil and
27 Paraguay. Ash cloud continued to change in direction over the next 6 days, with clouds following the ridge
28 structure to the NE and SE, respectively.

29 **4.1.1. Regional simulation**

30 *Model set-up*

31 The model domain for the regional run is presented in Table 7 and consists of 268x268 grid points covering the
32 northern regions of Chile and Argentina using a rotated latitude–longitude grid with a horizontal resolution of
33 0.15° x 0.15° and 60 vertical layers. The top pressure of the model was set to 21 hPa (~34 km) with a mesh
34 refinement near the top (to capture the dispersion of ash) and the ground (to capture the characteristics of the
35 atmospheric boundary layer). The computational domain spans in longitude from 41° W to 81° W and in latitude
36 from 18° S to 58° S. Runs were performed with the on-line version of NMMB-MONARCH-ASH from 3 June
37 2011 at 00:00 UTC to 21 June 2011 at 00:00 UTC. The integration time-step for the meteorological core and

1 aerosol transport was set to 30 seconds. The dynamic time-steps for the long and short wave radiations were
2 computed every 120 time-steps. Feedback effects of ash particles on meteorology and radiation were not
3 included in this run. The meteorological driver was initialized with wind fields from the Era-Interim reanalysis at
4 $0.75^\circ \times 0.75^\circ$ resolution as initial and 6-h boundary conditions. In order to reduce the errors in meteorological
5 conditions, they were reinitialized every 24 h with a spin-up of 12 h. Daily eruption source parameters (ESP)
6 were obtained from Osoreo et al. (2014), who estimated column heights for each eruptive pulse using the Imager
7 Sensor data from the GOES-13 satellite, applying the cloud-top IR image technique (Kidder and VonderHaar,
8 1995). Mass flow rate released along the column was derived from column heights based on Mastin et al. (2009),
9 assuming a Suzuki vertical distribution of mass typical of explosive Plinian eruptions ($A=4$; $\lambda=5$). Grain size
10 distribution was obtained from Collini et al. (2013) and discretized in 10 bins ranging from -1Φ (2 mm) to 8Φ (4
11 μm) with a linear dependency of particle density on diameter ranging from 1.000 to 2.200 kg m^{-3} . Particle
12 sphericity was set to a constant standard value of 0.9 for all bins. The *percentage* aggregation model was used to
13 update the TGSD with a new bin for aggregates, resulting in a total of 11 bins.

14 *Validation of results against satellite imagery*

15 Model results for the airborne mass concentration of ash were validated using qualitative and quantitative
16 comparisons with data obtained using two different techniques. On one end, we performed a qualitative
17 comparison between the simulated column mass (g m^{-2}) from the model and the NOAA-AVHRR satellite
18 imagery provided by the high-resolution picture transmission (HRPT) division of the Argentinian National
19 Meteorological Service. Figure 3 shows how the model predictions for cloud trajectory and arrival times are in
20 agreement with observations, capturing the three major dispersion episodes. It should be noted that these types of
21 images are not directly comparable because the MODIS ash detection threshold and the reflectivity coefficients
22 of volcanic ash are not well constrained. However, the figure illustrates the capability of the model to predict the
23 variation of the cloud position with time.

24 Column mass simulations were also validated against ash mass loadings presented by Osoreo et al. (2015), who
25 retrieved ash-contaminated pixels detected on the basis of the concept of reverse absorption (Prata, 1989a,
26 1989b), i.e. those pixels with brightness temperature differences between 11 and 12 μm (BTD11-12 μm) that are
27 lower than 0.0 K. To minimize the presence of false positives, pixels with a BTD11-12 $\mu\text{m} > -0.6$ K and clear
28 sky pixels were removed. Mass loadings were mapped up to 15 g m^{-2} based on an approach which combines the
29 satellite data with look-up tables of brightness temperatures obtained with a radiative transfer model and optical
30 properties of andesite volcanic rocks (Prata, 2011). Figure 4 shows a good quantitative agreement between the
31 model results and the airborne ash mass loadings described above.

32 *Validation of results against fallout deposit*

33 Tephra was mostly deposited eastward from the source during the first 72 h of the event within an elongated area
34 between $40\text{-}42^\circ$ S and $64\text{-}72^\circ$ W. Results from the NMMB-MONARCH-ASH ash deposition forecast were
35 validated against: i) a detailed characterization of the proximal deposit for the first 72 h of the eruption, and ii) an
36 isopach map derived from measurements taken for the period beginning on 4 June until 30 June (Collini et al.,
37 2013).

1 To evaluate the simulated computed thicknesses (cm) by the model near the vent during the first 72 h of the
2 event, model results were compared against a comprehensive classification of the proximal deposit presented by
3 Pistolesi et al. (2015b), who constrained the stratigraphic sequence of the deposit in different units (phases).
4 Here, we constrain the deposit to the first three units of their work, corresponding to the first 72 h of the eruptive
5 even and including: i) Unit I, containing coarser-grained layers A-B, representing the very first stage of the
6 eruption within the first 50 km from the vent, and layers A–F associated to the first 24-30 h of the eruption
7 (afternoon of 4 to morning of 5 June); ii) Unit II, containing layer H, a fine pumice lapilli layer which was
8 emplaced starting on the night of 6 June; iii) Unit III, enclosing layer K2, the easiest to identify from several
9 coarser (fine-lapilli) grain-size layers, and being associated to the morning of 7 June. Figure 5 shows that
10 NMMB-MONARCH-ASH can reproduce the deposit presented by Pistolesi et al. (2015b) both in time and
11 space. Key sections located along the dispersal area (e.g. San Carlos de Bariloche – SCB, 90 km from the vent;
12 Ingeniero Jacobacci – IJ, 240 km east of the vent) were used as geographic references.

13 To evaluate the model performance at the end of our simulation, model results were also validated against an
14 isopach map derived from measurements taken from the 4 to 30 June presented by Collini et al. (2013). Deposit
15 load variations produced by remobilization were not considered in this analysis. Figure 6 shows good agreement
16 between the modeled deposit load (kg m^{-2}) at the end of the simulation and the measured ground deposit
17 isopachs (kg m^{-2}) at 30 June from Collini et al. (2013).

18 The model resulted in a cumulative mass of $\sim 4.2 \times 10^{11}$ kg. This value is in agreement with previous works,
19 where total mass was either modeled (Collini et al., 2013) or estimated by empirical fits (Bonadonna et al.,
20 2015b). Ashfall forecast with the NMMB-MONARCH-ASH model represented well the overall deposit load for
21 the 2011 Cautle eruption.

22 **4.1.2 Global simulation**

23 For this simulation, the global domain was configured using a regular latitude–longitude grid with a horizontal
24 resolution of $0.75^\circ \times 1^\circ$ and 60 vertical layers. The ash distribution is simulated between 3-21 June 2011 using
25 the Era-Interim reanalysis at $0.75^\circ \times 0.75^\circ$ resolution as initial and 6-h boundary conditions. Meteorological
26 conditions for the global runs were also reinitialized every 24h. The atmospheric model's fundamental time-step
27 was set to 180 s, while the rest of the model variables and grain size distribution remained the same as in the
28 regional simulation. Figure 7 shows the global dispersal of ash for the 2011 Cordon Cautle eruption at different
29 times of the simulation. As it can be inferred from this figure, by 10 June, the plume entered the Australian and
30 New Zealand airspace (Fig 7b) covering more than half of the southern hemisphere. At that point, the Civil
31 Aviation Authority of New Zealand warned pilots that the ash cloud was between 20,000 and 35,000 feet (6 to
32 11 kilometers), the average cruising level for many aircraft (Sommer, 2011). Before the end of our simulation,
33 on 13 June the ash cloud had completed its first circle around the globe. This is in agreement to satellite images
34 reported by the Darwin Volcanic Ash Advisory Centre (Darwin VAAC, 2011). Finally, results from the global
35 simulation are also in agreement with those from our regional run.

1 **4.1.3 Forecasting impacts on civil aviation**

2 NMMB-MONARCH-ASH can furnish values of airborne concentration at relevant flight levels (FL), defined as
3 the vertical altitude (expressed in hundred of feet) at standard pressure at which the ash concentration is
4 measured. This information is particularly important for air traffic management and can be used to decide
5 alternative routes to avoid an encounter with a volcanic cloud. Airborne concentration at FL050 (5,000 feet on
6 nominal pressure) is relevant for the determination of flight cancellations and airports closures, while
7 concentrations at FL300 (30,000 feet) are critical to assist flight dispatchers while planning flight paths and
8 designing alternative routes in the presence of a volcanic eruption. The model runs as if responding to an
9 eruptive event, i.e. we only used the semi-quantitative data available at that time as volcanological inputs.

10 Figure 8 shows the airspace contamination forecasted by the model during the 6-7 June at flight levels FL050
11 and FL300, within a latitude band between 20° S and 55° S. Model results show the volcanic cloud twisting in
12 different directions during that period of time, achieving critical concentration values within a wide area east of
13 the Andes range. On 6 June, simulation results show the volcanic cloud at high atmospheric pressure (~ 30,000
14 feet or 300 hPa) moving northwards, and the one at lower atmospheric pressure (~ 5,000 ft or 50 hPa)
15 threatening the main international airports that service the region of Buenos Aires (Fig. 8a). In the morning of 7
16 June, the ash cloud present at lower atmospheric pressure (~ 5,000 ft or 50 hPa) changed its direction towards the
17 SW, ultimately affecting part of the Patagonia and Chile (Fig. 8b), while higher ash clouds started their course
18 around the globe (Fig. 8c). These results suggest that the cancellation of multiple flights in several Argentinean
19 airports during this time was justified. It is important to point out that, for this work, our objective is not to
20 perform a detailed study of the Caulle eruption but to use it as a blind test to confront short-term model
21 predictions and semi-quantitative syn-eruptive observations.

22 **4.2 The 2001 Mt. Etna eruption**

23 Mt. Etna is the most active volcano in Europe and constitutes a continuous hazard for eastern Sicily. Since 1980,
24 Mt. Etna has injected large volumes of pyroclasts into the atmosphere (between 10^4 and 10^7 m³ per event) over
25 more than 160 eruptive episodes (Scollo et al., 2012). The explosive activity of Mt. Etna reached its climax in
26 2001 and 2002–03 when two major flank eruptions occurred; both characterized by long-lasting explosive
27 activity (Branca and Del Carlo, 2005). The 2001 event represents a good case to evaluate the deposition
28 mechanisms of the model against the well-characterized tephra deposit reported in Scollo et al. (2007). The
29 explosive activity at the 2570 m vent had three main phases characterized by phreatomagmatic, magmatic and
30 vulcanian explosions. The eruption started with a series of phreatomagmatic explosions during the first days of
31 the eruption. These explosions were followed by a second eruptive phase characterized by strombolian and
32 Hawaiian style explosions during 19-24 July. The explosive activity continued until 6 August with a series of
33 vulcanian explosions. Tephra fallout associated to the explosive activity during 21-24 July represented a major
34 source of hazard for eastern Sicily. Flight operations were cancelled at the Catania and Reggio Calabria airports
35 during the 22 and 23 July. A detailed chronology of the eruption can be found in Scollo et al. (2007). Volcanic
36 plumes were captured by the Multiangle Imaging Spectro Radiometer (MISR) on board NASA's Terra
37 spacecraft, and analyzed with stereo matching techniques to evaluate the height of the volcanic aerosol with a
38 precision of a few hundred meters (Scollo et al., 2012).

1 Here, we validate NMMB-MONARCH-ASH against the tephra deposit produced from the 2570 m vent for that
2 period of time, and compare the model performance against simulations results from the FALL3D model (Costa
3 et al., 2006; Folch et al., 2009) for the same event. FALL3D is an Eulerian model for transport and deposition of
4 volcanic ash particles solving a set of advection-diffusion-sedimentation equations (one equation for each
5 particle class) on a structured terrain-following grid using a second-order Finite Differences explicit scheme. The
6 FALL3D model is used at the Buenos Aires and Darwin Volcanic Ash Advisory Centers (VAAC) in operational
7 forecasts.

8 **4.2.1 Regional simulation**

9 *Model set-up*

10 Two regional domains were used to simulate the first phase of the 2001 eruption of Mt. Etna (Table 8). The first
11 domain (Regional 1), used to reconstruct the tephra deposit, consists of 101x101 grid points covering the SE
12 flank using a rotated latitude–longitude grid with a horizontal resolution of 0.05° x 0.05° and 60 vertical layers.
13 Similarly to the Cordón Caulle simulations, the top pressure of the model was set to 21 hPa (~34 km) with a
14 mesh refinement near the top and ground. The computational domain spans in longitude from 12.5° E to 17.5° E,
15 in latitude from 35.25° N to 40.25° N. Simulation runs were performed with the on-line version of NMMB-
16 MONARCH-ASH from 21 July 2001 at 00:00 UTC to 25 July 2001 at 00:00 UTC. The integration time-step for
17 the meteorological core was set to 10 seconds. The meteorological driver was initialized with Era-Interim
18 reanalysis meteorological data at 0.75° x 0.75° resolution as initial and 6-h boundary conditions. A spin-up of 12
19 h was used to prepare the meteorological conditions for run. Each daily model run was reinitialized with the
20 corresponding reanalysis of the model tracers' output from the previous day, and the associated eruption source
21 parameters. Meteorological conditions were reinitialized every 24 h. The grain size distribution and eruption
22 source parameters were obtained from Scollo et al. (2007), who assumed a Suzuki vertical mass distribution
23 located at the middle of the eruption column ($A=2$; $\lambda=1$), and employed the Mastin et al. (2009) empirical
24 relationship to characterize the MER and the Voronoi tessellation method to obtain the grain size distribution.
25 Finally, sensitivity analyses were performed against the different aggregation schemes available in the model. In
26 all cases, the TGSD was updated with a new bin for aggregates, resulting in a total of 8 bins.

27 A second regional domain (Regional 2) was used to evaluate tephra dispersal between 21 and 25 of July. In this
28 case, the domain consisted of 201x201 grid points covering a computational domain spanning in longitude from
29 41° E to 81° E, in latitude from 18° S to 58° S. This domain used a coarser horizontal resolution of 0.1° x 0.1° and
30 60 vertical layers. The integration time-step for the meteorological core was set to 30 seconds. The rest of model
31 set-up was kept the same as in the first regional domain (Regional 1).

32 *Validation of results against fallout deposit*

33 At the end of the second explosive phase, a continuous tephra layer covered Etna's flanks between Giarre and
34 Catania (from E to S). Ash deposition results from the model were validated against 47 samples collected
35 between 25 and 26 July from measured areas on flat open spaces, where the deposit did not show any reworking.
36 The computed tephra dispersal and deposition from NMMB-MONARCH-ASH was able to reproduce the

1 bilobate shape of the real deposit with the two axes oriented toward Acireale and Acicastello towns. Figure 9
2 compares the simulated deposit load (kg m^{-2}) at the end of the run against the isopachs map derived from
3 measurements taken from the 21-24 July (Scollo et al., 2007). The model resulted in a cumulative mass of
4 $\sim 1.18 \times 10^9$ kg. This value is in agreement with the results obtained from Scollo et al. (2007).

5 **4.2.2 Model intercomparison: NMMB-MONARCH-ASH vs. FALL3D**

6 To validate the model performance of NMMB-MONARCH-ASH for its operational implementation, we
7 compare the tephra deposition results of the model against those of the operational FALL3D model for the
8 reconstruction of the 2001 Mt. Etna eruption. For this comparison we ran both models using the same
9 meteorological and volcanological initial conditions (Table 8). Figure 10 shows the simulated thicknesses
10 (vertical axis) for both transport models against the observations (horizontal axis) presented in Scollo et al.
11 (2007). The model improved the tephra distribution results from FALL3D simulations for the same event (R^2 ;
12 0.80/0.62), reducing the RMSE (0.014/0.24) and bias (0.02/0.6) and the computational time by an order of
13 magnitude. In particular, all values simulated with NMMB-MONARCH-ASH plot inside the region between 5
14 and 1/5 (dashed orange line) times the observed mass at each station. The greatest differences perceived against
15 the observations for both models belong to those points located at distances less than 15 km from the vent
16 associated to the uncertainty in the ESPs. The mean value of the relative error between the computed values and
17 observed data is 64%, which improves those from FALL3d (91%), and are comparable with those of Scollo et
18 al., (2007), who obtained a 57% by deposit best-fitting using the HAZMAP dispersion model.

19

20 **5 Operational forecast with NMMB-MONARCH-ASH**

21 The Barcelona Supercomputing Center is currently working on a modeling integrated system to provide
22 operational forecast of volcanic ash with NMMB-MONARCH-ASH. The system includes a preprocessing tool
23 (prepares the model for real-data simulations), an executable file to run the model, and a user-based
24 postprocessing utility tool. Figure 11 shows a simple schematic representation of the operational implementation
25 of the model. The outcomes of this modeling system are currently being evaluated against two operational
26 models: i) the NOAA/ARL Hybrid Single Particle Lagrangian Integrated Trajectory Model (HYSPLIT; Draxler
27 and Hess, 1998) – used at the Washington VACC; and FALL3D (Costa et al., 2006; Folch et al., 2009) – used at
28 the Buenos Aires and Darwin VAACs. This section introduces the structure of the operational NMMB-
29 MONARCH-ASH system. Preliminary results for the model intercomparison against FALL3D are described in
30 Sect. 4.2.2.

31 **5.2 The preprocessing system**

32 The preprocessing utility system consists of a set of programs whose collective role is to prepare the model for
33 real-data simulations. Programs are grouped to preprocess geographical, meteorological and climatological
34 inputs and interpolate those to the model grid(s). The preprocessing system uses three main programs: *runfix*,
35 *degrib* and *runvariable*.

- 1 • *Runfix* defines the model domain(s) and interpolates static geographical data to the model grid(s). In
2 addition to computing the latitude and longitude of the rotated grid points, this program interpolates soil
3 categories, land use types, terrain height, annual mean deep soil temperature, monthly albedo,
4 maximum snow albedo, and slope category.
- 5 • *Degrrib* extracts the necessary meteorological fields from GRIB-formatted files, used as initial condition
6 for global simulations and as initial and boundary conditions for single regional domains (i.e. not nested
7 with a global domain). GRIB files contain time-varying meteorological fields obtained from another
8 regional or global NWPM. In addition to the available NCEP's North American Model (NAM) or
9 Global Forecast System (GFS) model, the program has been updated to include European Centre for
10 Medium-Range Weather Forecasts (ECMWF) ERA-Interim reanalysis data (Dee et al., 2011) as
11 forcing.
- 12 • *Runvariable* interpolates the meteorological fields extracted by *debgrib* to the model grid(s) defined by
13 *runfix* and prepares the climatological schemes. This program generates the initial and boundary
14 conditions that are ingested by NMMB using the NOAA Environmental Modeling System (NEMS;
15 Janjic, 2005; Janjic and Black, 2007), a high performance software superstructure and infrastructure
16 based on the Earth System Modeling Framework (ESMF) for use in operational prediction models at
17 NCEP.

18 **5.3 The ash module I/O files**

19 The model takes three run-specific input files:

- 20 • The model input file (*nmmmb.inp*), which defines the computational and physical schemes needed by the
21 meteorological core, the atmospheric model's fundamental time-step, and the parameterization for
22 chemical processes and radiative schemes for aerosol tracers (including ash), amongst other properties
23 of the model. For long-lasting eruptions, the model performs restart runs initializing the tracers from the
24 previous day's history file.
- 25 • The ash input file (*ash.inp*), which defines those parameters employed in the ash module. The user-
26 defined parameters include: i) the characterization of the source term: eruption source type, column
27 height and determination of the mass eruption rate, eruption duration, aggregation processes, and
28 particle settling velocity model. In the event of various eruptive phases, the respective ESPs for each
29 phase can be defined; ii) the settings to turn on/off the gravity current model altering the particle
30 transport in the umbrella cloud; and iii) the definition of the coupling strategy (on vs. off-line)
31 employed by the model.
- 32 • The granulometry input file (*ash.tgsd*), which specifies the diameter, density, sphericity, and relative
33 mass fraction of each particle bin. This information is typically obtained from field data or created by
34 external utility programs for idealized grain size distributions. If aggregation is active, a new bin class
35 for aggregates is added to the granulometry input file.

36 Once a simulation is concluded, NMMB-MONARCH-ASH writes the following output files:

- 1 • A log file (*ash.log*), containing information about the run, including a summary of the computed
2 volcanic ash source and mass balance statistics for each time-step, and errors and warnings if any.
- 3 • A forecast results file (*problemname.nc*) in NetCDF format containing, amongst other variables, the
4 total column mass concentration (g m^{-2}) and ground deposition (kg m^{-2}) for all bins, the concentration at
5 different Flight Levels (g m^{-3}) and the Aerosol Optical Depth. This information can be processed using
6 several open-source programs to generate plots and animations. Alternatively, the post-process utility
7 program *NMMB2GMT* has been developed to generate basic GMT scripts automatically.
- 8 • A restart file (*nmmmb.hst*) used to initiate a new run using the ash concentrations from a previous
9 simulation.

10 **5.4 The Postprocess system**

11 The postprocess utility tools are designed to interpolate outputs from the NMMB-MONARCH-ASH native
12 grid(s) to National Weather Service (NWS) standard levels (pressure, height, etc.) and standard output grids
13 (Lambert Conformal, polar-stereographic, etc.) in NetCDF format. The system also includes the *NMMB2GMT*
14 program, which uses the Generic Mapping Tools (GMT) software (Wessel and Smith, 1991) to produce similar
15 plots to the Volcanic Ash Graphics (VAG) used by Volcanic Ash Advisory Centers in operational forecasts.

16 **5.5 Scalability analysis**

17 To optimize a future operational implementation of the model, we aim to minimize the time-to-solution avoiding
18 communication overhead. In this context, we evaluate the model scalability (scaling efficiency) for its regional
19 and global configurations by performing a strong scalability test, in which the problem size of our simulation
20 (e.g., model domain and resolution) remains fixed while increasing the number of processing cores. Figure 12
21 shows the parallel speed-up (S ; Eq. 19), and efficiency (E ; Eq. 19) of the modeling system for a global
22 simulation of the climactic phase for the 2011 Cordón Caulle (Table 7). On the MareNostrum-III supercomputer,
23 maximum efficiency for the global simulation described in Table 7 is reached between 32-40 nodes (16 CPUs
24 each) with a parallel efficiency of 0.6.

25 The scalability analysis was performed on all the available source term and sedimentation schemes in the model.
26 The relative computational cost associated with the main processes in NMMB-MONARCH-ASH is presented in
27 Fig. 13. Processes represented include: meteorological prediction, volcanic ash transport and sedimentation
28 forecast, aggregation of particles, gravity current effects, and the restart phase. The restart phase represents the
29 CPU time employed to rerun the preprocess system every 24h of simulation. This figure suggests that the
30 computational increase (CPU time) associated to the ash module can vary from 5 to 55%, depending on the
31 number of computational nodes employed. It is important to note that, depending on the settling velocity model
32 employed, up to 60% of the time allocated to the ash module is assigned to the sedimentation term.

33 Results from the scalability analysis show that the model performance (in terms of speed-up) depends on the
34 problem size as well as on the domain partitioning topology. In that context, the relative computational cost of
35 the model's meteorological core (NMMB) is evaluated as a function of its domain decomposition (e.g.,
36 distribution of processing units for the horizontal domains – nodes i and j). For this analysis the bin-performance

1 dependency of the model is considered, therefore evaluating only the cost of one bin of ash. Results from this
2 analysis suggest that, for an optimal simulation using 32 nodes, the computational cost of the meteorological
3 core decreases over 10 % when the weight of the decomposition is focused on the j nodes (e.g., more
4 computational resources assigned for the Fast Fourier Transformation algorithm). The best domain
5 decomposition resulted in $6(i) \times 84(j) + 8(w)$; where i and j , are the number of processors employed in the
6 horizontal and vertical domains respectively, and w , the number of writing processors.

7 For operational purposes, the computational time employed to provide ash dispersal forecast using NMMB-
8 MONARCH-ASH is evaluated for the global simulation with 1 bin of ash. The maximum time required by the
9 model to perform a 24 h forecast, running all the available processes (e.g., advection, diffusion, sedimentation,
10 etc.) every time-step (180 seconds) is less than 3 minutes when using the best domain decomposition presented
11 before ($6 \times 84 + 8$). This time can be further optimized for operational purposes, i.e., calling the model physics less
12 frequently in order to save computational time. As a general rule of thumb, the adjustment time-step in seconds
13 for the meteorological core can be taken as 2.25 times the grid spacing in kilometers. For higher resolution
14 model runs made without parameterized convection, a time-step in seconds of about 1.9 to 2.0 times the grid
15 spacing may be more appropriate (Janjic and Gall, 2012).

16 **5.6 Cost-benefit analysis**

17 Employing on-line models for operational dispersal forecast requires larger computational resources and is not
18 always feasible at all operational institutes. Nevertheless, due to the increase in computing power of modern
19 systems, one can argue that such gradual migration towards stronger on-line coupling of NWPMs with TDMs
20 poses a challenging but attractive perspective from the scientific point of view for the sake of both high-quality
21 meteorological and volcanic ash forecasting.

22 The focus on volcanic aerosols integrated systems in operational forecast is timely. Experiences from other
23 communities (e.g. air quality) have shown the benefits from two-way online meteorology-chemistry modeling.
24 For example, the importance of the different feedback mechanisms for meteorological and atmospheric
25 composition processes have been previously discussed for models developed in the USA (Zhang, 2008) and
26 Europe (Baklanov et al., 2014). These benefits have been recently stressed by several studies covering the
27 analysis of the aerosol-transport and aerosol-radiation feedbacks onto meteorology from the air quality model
28 evaluation international initiative (AQMEII) in its phase 2 (Alapaty et al., 2012; Galmarini et al., 2015) and the
29 EuMetChem COST Action ES1004 (EuMetChem, <http://eumetchem.info>)

30 Demonstrating these benefits however, require running the on-line model with and without feedbacks over
31 extended periods of time. For the particular case of volcanic aerosols, further research is still required to quantify
32 the benefits posed by on-line couple models over traditional off-line TTDM on both atmospheric transport and
33 the radiative budget. The Barcelona Supercomputing Center is currently working to quantify these benefits with
34 NMMB-MONARCH-ASH model, and assess how the magnitude of the model forecast errors implicit in the off-
35 line approach compares with other better-constrained sources of forecast error, e.g. uncertainties in eruption
36 source parameters. Preliminary results from this study indicate that meteorology-transport inconsistencies from
37 off-line models can be, in some cases, in the same order of magnitude that those associated to the eruption source

1 parameters. In terms of computational cost, the computational efficiency of the model's meteorological core
2 allows for on-line integrated operational forecasts employing an equivalent computational time than FALL3D
3 for the same computational domain and number of processing cores.

4 5 **6 Summary and Conclusions**

6 We present NMMB-MONARCH-ASH, a new on-line multiscale meteorological and transport model developed
7 at the Barcelona Supercomputing Center (BSC) to forecast the dispersal and deposition of volcanic aerosols. The
8 objective of this model is to improve the current state-of-the-art of tephra dispersal models, especially in
9 situations where meteorological conditions are fluctuating rapidly in time, two-way feedbacks are significant, or
10 long-range ash cloud dispersal predictions are necessary. The model predicts ash cloud trajectories,
11 concentration of ash at relevant flight levels, and the expected deposit thickness for both regional and global
12 domains. NMMB-MONARCH-ASH solves the mass balance equation for volcanic ash by means of a new ash
13 module embedded in the BSC's operational system for short/mid-term chemical weather forecast (NMMB-
14 MONARCH). In addition to volcanic ash, the system is also capable to forecast the dispersion of other
15 atmospheric aerosols (e.g. dust, sea salt, black carbon, organic aerosol, sulfates, etc.). Its multiscale capability
16 allows for nested global-regional atmospheric transport simulations, taking into account the characterization of
17 the source term (emissions), the transport of volcanic particles (advection/diffusion), and the particle removal
18 mechanisms (sedimentation/deposition). The model has been shown to be robust and scalable to arbitrary
19 domain sizes (regional to global) and numbers of processors.

20 The forecast skills of NMMB-MONARCH-ASH have been validated against several well-characterized
21 eruptions, including the, Etna 2001 (Italy), Chaitén 2008 (Chile), Cordón Caulle 2011 (Chile) or Pinatubo 1991
22 (Philippines) eruptions (e.g. Martí et al., 2013, 2014). To evaluate the on-line coupling strategy and the
23 multiscale capability of the model, this paper summarizes the regional and global configurations of the model to
24 forecast the dispersal of ash for the first days of the 2011 Cordón Caulle eruption (strong long-lasting eruption
25 with rapid wind changes). In addition, to evaluate the sedimentation mechanisms of the model, this work also
26 includes the results from the regional configuration of the model for the first phase of the 2001 Etna eruption, a
27 good case study of weak long-lasting eruption with well-characterized tephra deposits. Simulation results
28 demonstrate that NMMB-MONARCH-ASH is capable to reproduce the spatial and temporal dispersal variability
29 of the ash cloud and tephra deposits.

30 31 **Software**

32 The work described in this paper is based on NMMB-MONARCH version 2.0.1 (released in April, 2014). The
33 code, written in FORTRAN-90, is portable and efficient on available parallel computing platforms. The figures
34 presented in this paper were generated using Gnuplot and NCAR Command Language (NCL).

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7 eruption dynamics. Numerical simulations were performed at the Barcelona Supercomputing Center with the
8 MareNostrum Supercomputer using 512 and 256 - 8x4 GB DDR3-1600 DIMMS (2GB/core) Intel SandyBridge
9 processors, iDataPlex Compute Racks, a Linux Operating System and an Infiniband interconnection.

10

11 **Competing interests**

12 The authors declare that they have no conflict of interest.

13

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1 **Tables**

2

3 **Table 1. Main characteristics of the NMMB-MONARCH-ASH meteorological solver.**

<i>Meteorological Solver</i>	<i>Scheme</i>	<i>Reference</i>
Spatial discretization		
Multi-scale domain ranging from large eddy simulations (LES) to global simulations		Janjic (2005)
Conservativeness		
Conservation of mass, momentum, energy, enstrophy and a number of other first order and quadratic quantities. Positive definiteness and monotonicity are preserved by tracer advection		Janjic (1984)
Coordinates /Grid		
Horizontal coordinate	Rotated latitude-longitude for regional domains, and latitude-longitude coordinate with polar filter for global domains	Janjic et al. (2009); Janjic and Gall, (2012)
Vertical coordinate	Terrain following hybrid sigma-pressure	Simmons and Burridge, (1981)
Horizontal grid	Arakawa B-grid staggering	Janjic, 2005; Janjic and Black, 2007)
Vertical grid	Lorenz staggering	Lorenz, (1960)
Time integration schemes		
Horizontally propagating fast-waves	Forward-backward scheme	Ames, (1969); Gadd, (1974); Mesinger, (1977); Janjic, 1979)
Vertically propagating sound waves	Implicit scheme	Janjic and Gall, (2012)
Horizontal advection & Coriolis terms	Modified (Stable) Adams-Bashforth scheme	
Vertical advection	Crank-Nicolson scheme	Janjic, (1977,1984)
TKE generation and dissipation	Iterative	
Advection terms		
Horizontal	Energy and enstrophy conserving, quadratic conservative, second order	Janjic and Gall, (2012)
Vertical	Quadratic conservative, second order	Janjic and Gall, (2012)
Diffusion terms		
Vertical	Surface layer scheme	Janjic (1994, 1996)
Lateral	Smagorinsky non-linear approach	Janjic (1990)
Physics Options		
Microphysics/Clouds	Ferrier (Eta)	Ferrier et al. (2002)
Short and Longwave Radiation	Rapid Radiative Transfer Model (RRTM)	Mlawer et al. (1997); Pérez et al. (2011)
Surface Layer	NMMB similarity theory scheme: Based on Monin-Obukhov similarity theory with Zilitinkevich thermal roughness length	Monin and Obukhov (1954); (Zilitinkevich, 1965); Janjic (1994, 1996)
Land Surface, Heat & moisture surface flux	LISS model	Vukovic et al. (2010)
Planetary Boundary layer / free atmosphere	Mellor-Yamada-Janjic scheme	Mellor and Yamada, (1982); Janjic (1996, 2001)
Convective adjustments	Betts-Miller-Janjic scheme	Betts and Miller, (1986); Janjic (1994, 2000).

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Table 2. Options implemented in NMMB-MONARCH-ASH to estimate the mass eruption rate from column height. Unless otherwise noted, the units for all parameters are in SI.

Reference	MER schemes	Eq.	Parameters
Mastin et al., (2009)	$MER = \bar{\rho} \left[\frac{0.5 H_{plume}}{10^3} \right]^{0.241}$	(1)	$\bar{\rho}$ = magma density (2500 kg m ⁻³) H_{plume} = column height above the vent (m).
Degruyter and Bonadonna (2012)	$MER = \pi \frac{\rho_{a0}}{g'} \left(\frac{\alpha^2 \bar{N}^3}{z_1^4 n} H_{plume}^4 + \frac{\beta^2 \bar{N}^2 \bar{v}}{6} H_{plume}^3 \right)$ $g' = g \left(\frac{c_0 \theta_0 - c_{a0} \theta_{a0}}{c_{a0} \theta_{a0}} \right)$ $\rho_{a0} = 1.105, \alpha = 0.1, \beta = 0.5, z_1 = 2.8, n = 0.177;$ $\theta_0 = 1200, \theta_{a0} = 268.7, c_0 = 1250, c_{a0} = 998$	(2)	ρ_{a0} = atmospheric density at the vent (kg m ⁻³) g' = reduced gravity \bar{N} = average buoyancy frequency (s ⁻¹) \bar{v} = average wind velocity across column height (m s ⁻¹) z_1 = Max. non-dimensional height α, β = radial and crossflow entrainment coefficients g = gravitational acceleration (9.81 m s ⁻²) c_0 = source specific heat capacity (J kg ⁻¹ K ⁻¹) c_{a0} = specific heat capacity of the atmosphere (J kg ⁻¹ K ⁻¹) θ_0 = source temperature (K) θ_{a0} = atmospheric temperature (K)
Woodhouse et al. (2013)	$MER = 0.35 \alpha^2 f(W_s)^4 \frac{\rho_{a0}}{g'} N^3 H_{plume}^4$ $f(W_s) = 1.44 \dot{\gamma} / \bar{N}$ $g' = g \left(\frac{c_p n_0 + c_s (1 - n_0) \theta_0 - c_a \theta_{a0}}{c_a \theta_{a0}} \right)$	(3)	Q = mass flux (kg s ⁻¹) W_s = dimensionless wind strength \bar{N} = average buoyancy frequency (s ⁻¹) $\dot{\gamma}$ = shear rate of atmospheric wind (s ⁻¹) c_s = specific heat of solids (J kg ⁻¹ K ⁻¹) c_a = specific heat of dry air (J kg ⁻¹ K ⁻¹) c_s = specific heat of water vapor (J kg ⁻¹ K ⁻¹)

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Table 3. Ash aggregation options in NMMB-MONARCH-ASH from analytical solutions based from field observations. Default aggregate properties can be modified by the user.

Name	New aggregate class	Default properties	Reference
NONE	No aggregation processes	n/a	n/a
CORNELL	50% of the 63–44 μm class aggregate 75% of the 44–31 μm class aggregate 100% of the < 31 μm class aggregate	Diameter = 250 μm Density = 350 kg m ⁻³ Sphericity = 0.9	Based on Cornell et al. (1983) Campanian Ignimbrite's deposit (Y5 ash layer)
PERCENTAGE	Takes a user-defined fixed percentage from each particle class	Diameter = 250 μm Density = 350 kg m ⁻³	Based on Sulpizio et al. (2012)

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1 **Table 4. Governing equations for NMMB-MONARCH-ASH wet aggregation model.**

<i>Wet aggregation scheme</i>	<i>Eq.</i>	<i>Parameters</i>
Number of particles of a class aggregated per unit volume	$\Delta n_f \approx \frac{\Delta n_{tot} N_j}{\sum_k N_k} \quad (k = k_{min}, \dots, k_{max}) \quad (5)$	Δn_{tot} = number of particles that aggregate per time interval N_j = number of particles of diameter j in an aggregate k = aggregation class N_k = number of particles of diameter k in an aggregate
Number of particles aggregated during Δt	$N_j = k_f \left(\frac{d_A}{d_j} \right)^{D_f} \quad (6)$	k_f = fractal prefactor ≈ 1 D_f = fractal exponent ≤ 3 d_A = aggregate diameter d_j = primary particle diameter
Total particle decay per unit volume during Δt	$\Delta n_{tot} = \alpha_m \left((A_B n_{tot}^2 + A_S \emptyset^{3/D_f} n_{tot}^{2-3/D_f} + A_{DS} \emptyset^{4/D_f} n_{tot}^{2-4/D_f}) \Delta t \right) \quad (7)$	α_m = mean sticky efficiency \emptyset = solid volume fraction
Number of particles available to aggregate	$n_{tot} = \sum_j \frac{6C_j}{\pi \rho_j d_j^3} \quad (8)$	n_{tot} = number of particles available to aggregate k_b = Boltzmann constant $1.38 \times 10^{-23} \text{ m}^2 \text{ kg s}^{-2} \text{ K}$ T = absolute temperature μ_o = dynamic viscosity of air Γ_s = fluid shear ξ = particle diameter to volume fractal relationship ρ_m = mean particle density ρ_a = air density
Kernels	<p>For Brownian motion: $A_B = -\frac{4k_b T}{3\mu_o}$</p> <p>Ambient fluid shear: $A_S = \frac{2\Gamma_s \xi^4}{3}$</p> <p>Differential sedimentation: $A_{DS} = -\frac{\pi(\rho_m - \rho_a)g\xi^4}{48\mu_o}$</p>	

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3 **Table 5. Governing equations for NMMB-MONARCH-ASH gravity current model.**

<i>Gravity current scheme</i>	<i>Eq.</i>	<i>Parameters</i>
Effective radial velocity of the umbrella spreading	$u_b(t) = \frac{2}{3} \left(\frac{3\lambda N q}{2\pi} \right)^{1/3} t^{1/3} \quad (9)$ $u_b(R) = \left(\frac{2\lambda N q}{3\pi} \right)^{1/2} \frac{1}{\sqrt{R}}$	u_b = effective radial velocity as a function of time (t) or cloud radius (R) λ = empirical constant ($\lambda \approx 0.2$) (Suzuki and Koyaguchi, 2009) N = Brunt-Väisälä frequency (atm. ambient stratification) q = volumetric flow rate into the umbrella region
Volumetric flow rate into the umbrella region	$q = C \sqrt{k} \frac{M^{3/4}}{N^{5/8}} \quad (10)$	M = efficiency of air entrainment k = mass eruption rate C = location base constant C_A : for tropical eruptions C_B : for mid latitude and polar eruptions
Contribution of the gravitational spreading	$ct = \left(\frac{u_b}{u_b + u_w} \right) \times 100 \quad (11)$	u_w = wind field velocity
Radial distance (gravity vs. passive transport)	$Ri = \frac{u_b^2}{u_w^2} = \frac{4}{9u_w^2} \left(\frac{3\lambda N q}{2\pi} \right)^{2/3} t^{-2/3} \quad (12)$	Ri = Richardson number $Ri > 1$: gravity-driven regime $Ri < 0.25$: passive transport regime

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1 **Table 6. Settling velocity models in NMMB-MONARCH-ASH.**

NAME/Reference	Drag coefficient	Eq.	Description/ Parameters
ARASTOPOUR (Arastoopour et al., 1982)	$C_d = \begin{cases} \frac{24}{Re} \{1 + 0.15Re^{0.687}\} & Re \leq 988.947 \\ 0.44 & Re > 988.947 \end{cases}$	(14)	For spherical particles only
GANSER (Ganser, 1993)	$C_d = \frac{24}{ReK_1} \{1 + 0.1118[Re(K_1K_2)]^{0.6567}\} + \frac{0.4305K_2}{1 + \frac{3305}{ReK_1K_2}}$	(15)	For spherical and non-spherical particles
	$K_1 = \frac{3}{[(d_n/d)]^{+2\psi-0.5}} ; K_2 = 10^{1.8148(-\text{Log}\psi)^{0.5743}}$		K_1, K_2 = shape factors d_n = average between the min and max. d axis
	$\psi_{\text{work}} = 12.8 \frac{(P^2Q)^{1/3}}{1 + P(1+Q) + 6\sqrt{1 + P^2(1+Q^2)}}$		ψ_{work} = sphere volume = particle sphericity ($\psi=1$ for spheres)
WILSON (Pfeiffer et al., 2005; Wilson and Huang, 1979)	$C_d = \begin{cases} \frac{24}{Re} \varphi^{-0.828} + 2\sqrt{1.07 - \varphi} & Re \leq 10^2 \\ 1 - \frac{1 - C_d _{Re=10^2}}{900} (10^3 - Re) & 10^2 \leq Re \leq 10^3 \\ 1 & Re \leq 10^3 \end{cases}$	(16)	ψ = particle aspect ratio $\psi = (b+c)2a^{-1}$ a, b, c = particle semi-axes
DELLINO (Dellino et al., 2005)	$v_s = 1.2605 \frac{v_a}{d} (Ar\varepsilon^{1.6})^{0.5206}$	(17)	For larger particles only
	$Ar = gd^3(\rho_p - \rho_a) \rho_a / \mu_a^2$		Ar = Archimedes number g = gravity acceleration ε = particle shape factor μ_a = dynamic viscosity d = particle equivalent diameter, ρ_p = particle density ρ_a = air density

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1 **Table 7. Model configuration for the 2011 Cordón Caulle regional and global runs. The regional run used a horizontal**
 2 **resolution of 0.15° x 0.15° with a 30s dynamic time-step, while the global domain used a horizontal resolution of 1° x**
 3 **0.75° with a 180s dynamic time-step.**

MODEL CONFIGURATION	
Dynamics	NMMB (30s/180s time-step)
Physics	Ferrier microphysics BMJ cumulus scheme MYJ PBL scheme LISS land surface model
Aerosols	11 ash bins (30s/180s time-step)
Source Term (emissions)	
Duration	20 days
Vertical distribution	Suzuki distribution
MER formulation	Mastin et al. (2009)
Aggregation model	Percentage
Sedimentation model	Ganser (1993)
Run Set-up	
Number of processors	512
Domain	Regional/Global
Horizontal resolution	0.15° x 0.15° / 1° x 0.75°
Vertical layers	60
Top of the atmosphere	21 hPa
Meteorology Boundary conditions (spatial resolutions)	ECMWF EraInterim Reanalysis (0.75° x 0.75°)

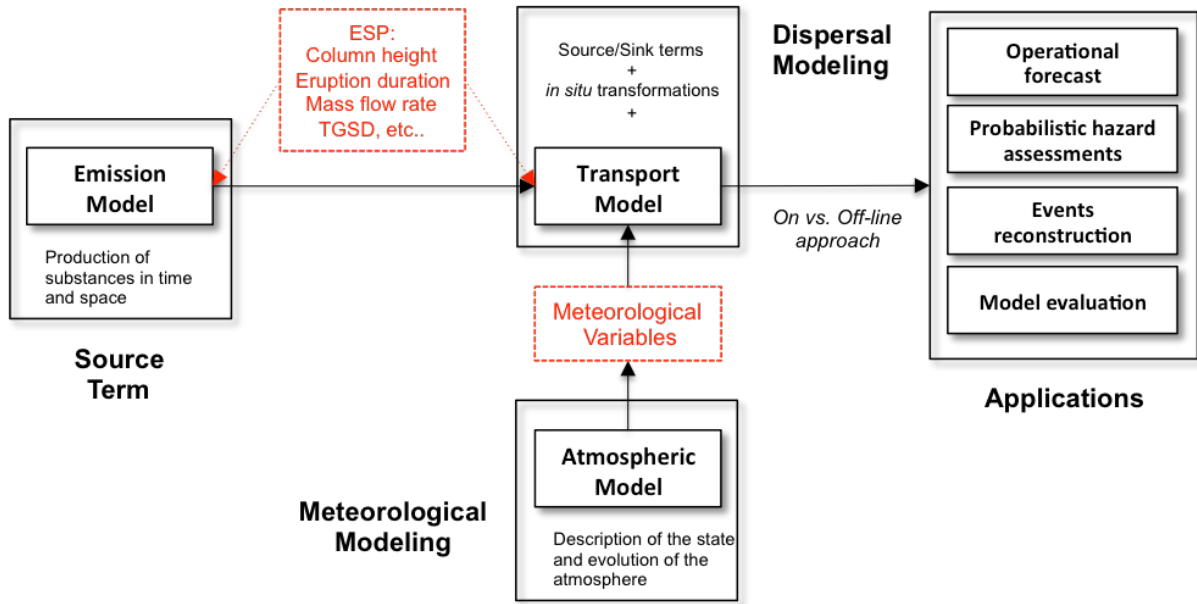
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1 **Table 8. Model configuration for the 2001 Mt. Etna regional simulations. Regional Run1 used a horizontal resolution**
 2 **of 0.1° x 0.1° with a 30s dynamic time-step, while Run2 used a finer horizontal resolution of 0.05° x 0.05° with a 10s**
 3 **dynamic time-step.**

Source Term (emissions)	
Duration	3 days
Vertical distribution	Suzuki distribution
MER formulation	Mastin
Column height above the vent	2570
Ash bins	8
Aggregation model	
Cornell et al. (1983)	
Sedimentation model	
Ganser (1993)	
Run Set-up	
Number of processors	256
Domain	Regional 1 / Regional 2
Horizontal resolution	0.1° x 0.1° / 0.05° x 0.05°
Vertical layers	60
Top of the atmosphere	21 hPa
Meteorology Boundary conditions (spatial resolutions)	ECMWF EraInterim Reanalysis (0.75° x 0.75°)

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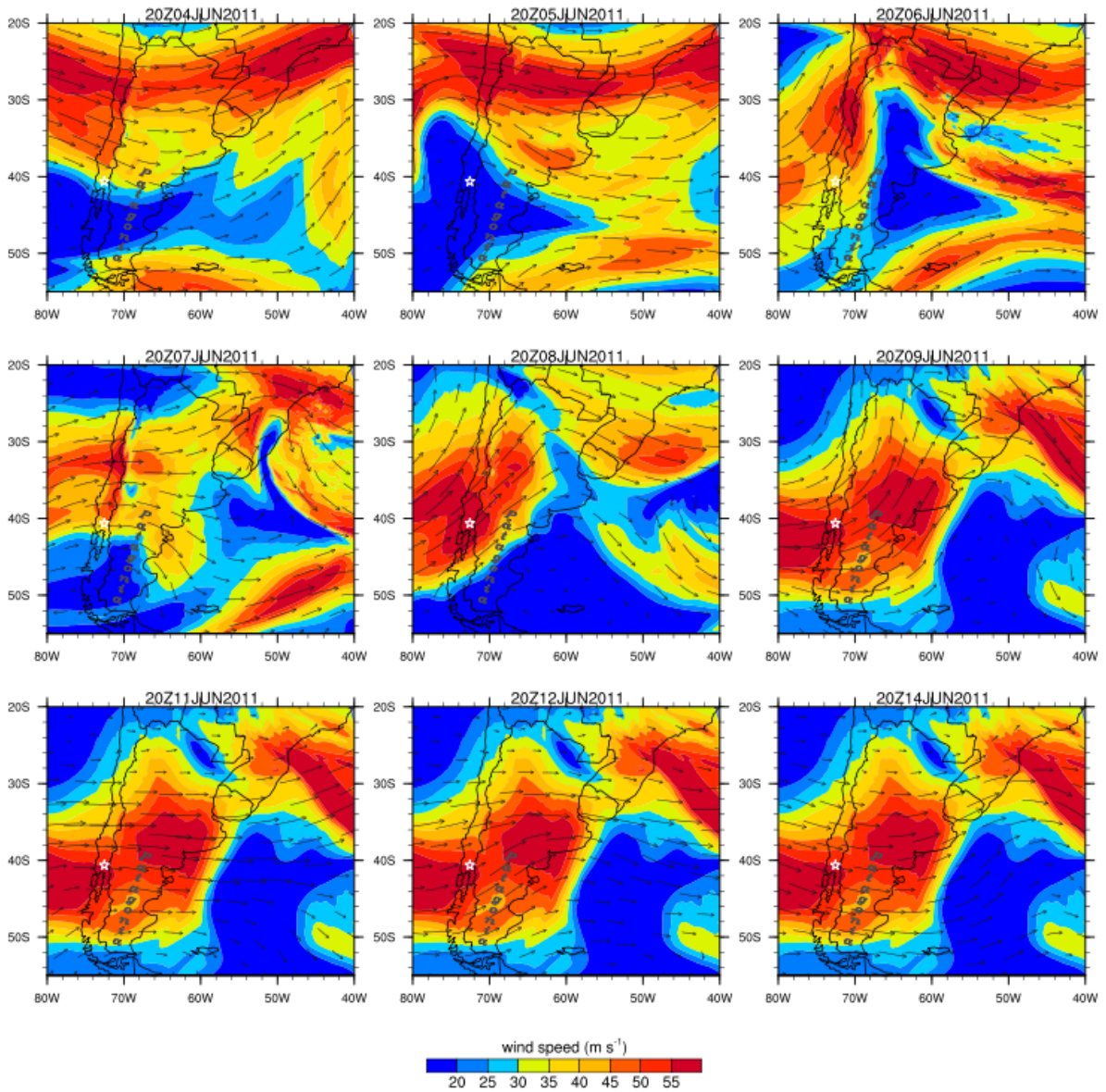
1 **Figure Captions**



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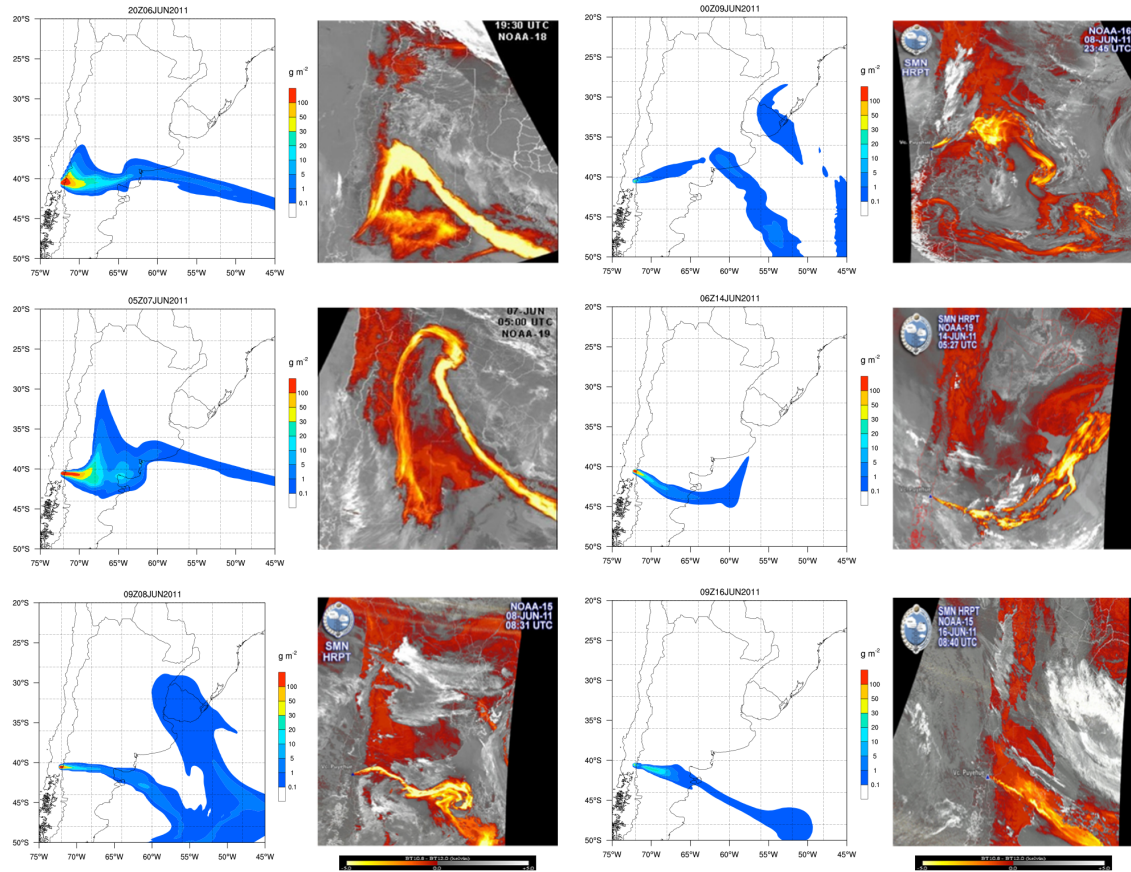
3 Figure 1. Schematic representation of the main components of an Atmospheric Transport Model. Red text shows

4 model specifications for the transport of volcanic ash.



1
 2 Figure 2. Meteorological synoptic situation during the first two weeks (4-14 June) of the 2011 Caulle (white star)
 3 activity over South America. Plots show the direction (vector) and velocity (contours m s^{-1}) of the wind at 9100
 4 m above ground level (300 hPa circa). Meteorological data obtained from the NMMB meteorological forecast
 5 driven with ERA-Interim reanalysis at 0.75° horizontal resolution.

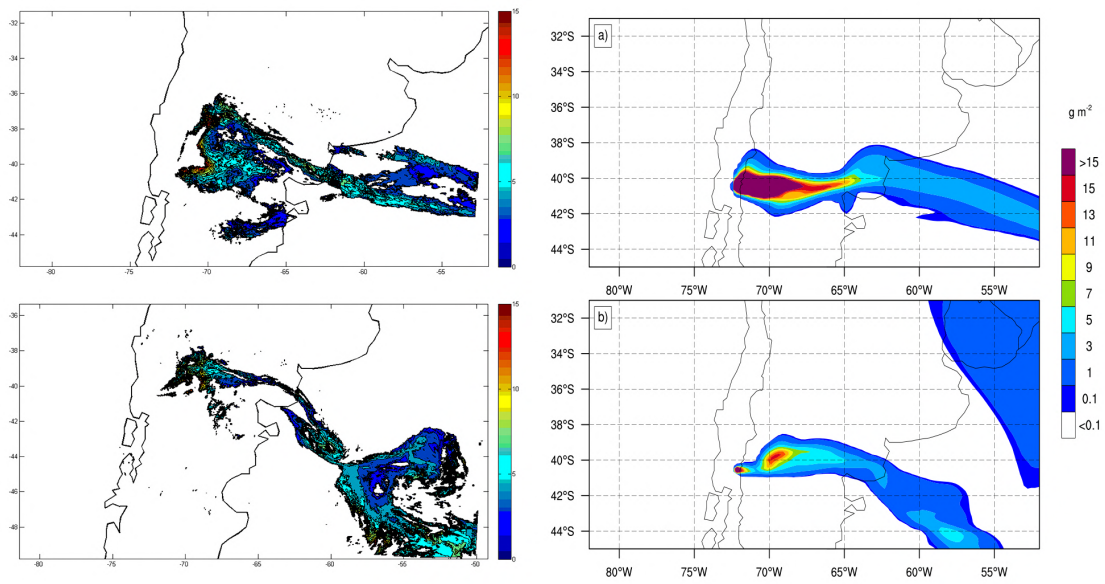
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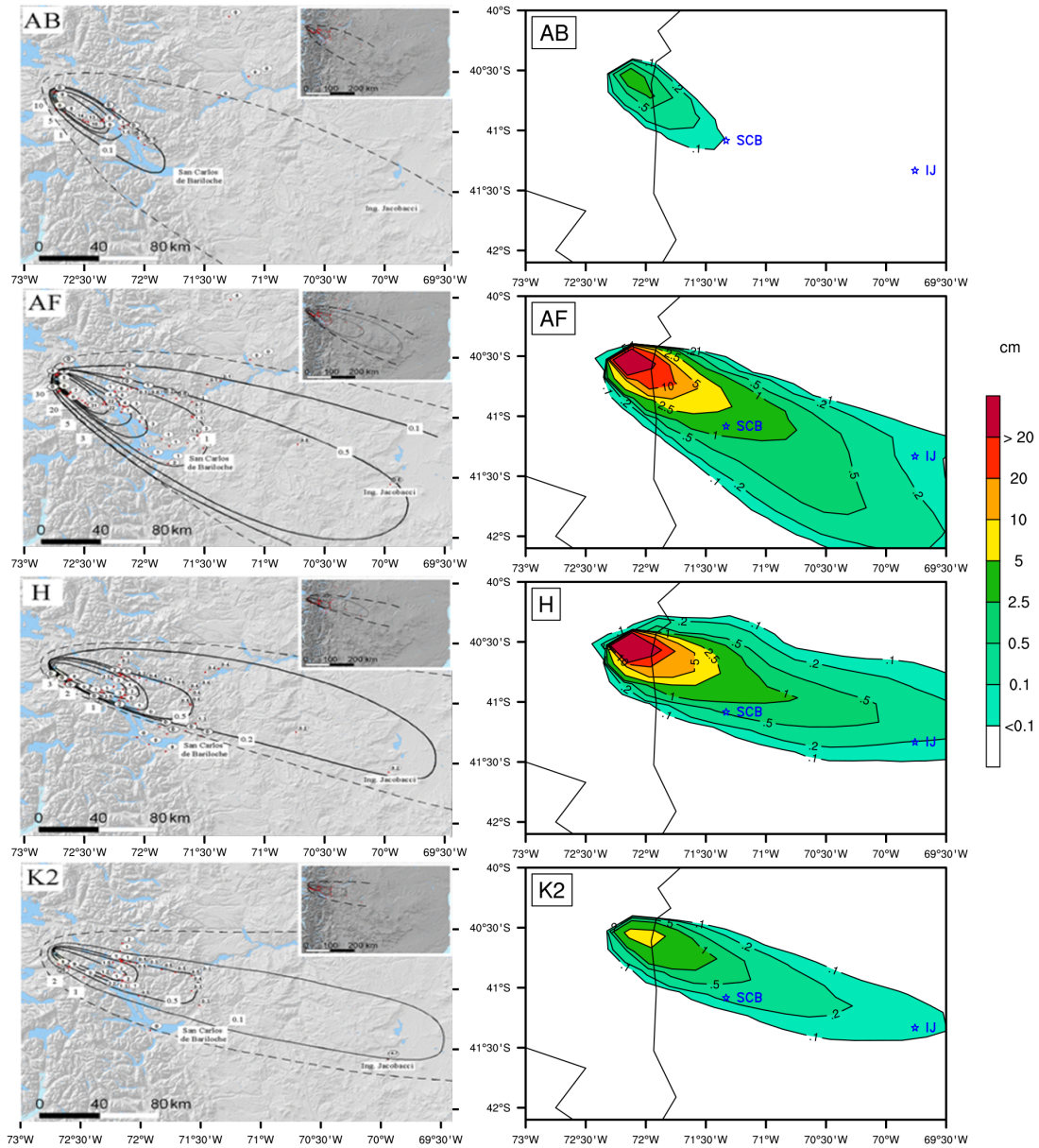
3 Figure 3. Composite image of NMMB-MONARCH-ASH results for dispersion of ash for the 2011 Cauller
4 eruption at different time slices. Simulation results are compared against split windows algorithm NOAA-
5 AVHRR satellite images (bands 11-12 microns). Contours indicate ash column load (g m^{-2}) resulting from
6 integrating the mass of the ash cloud along the atmospheric vertical levels.

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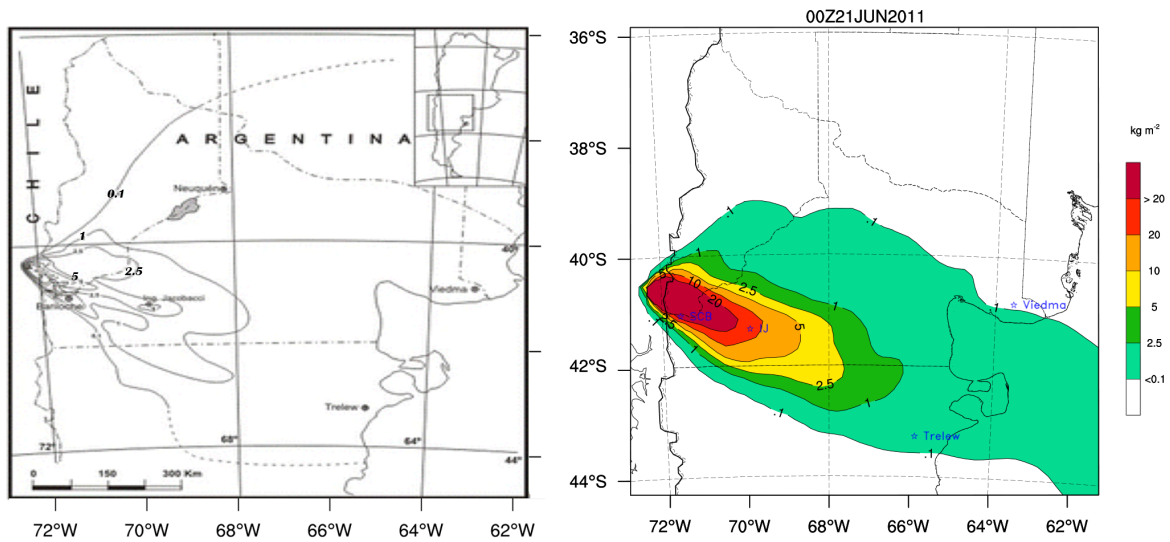
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3 Figure 4. Left: Mass loadings (g m^{-2}) of the 2011 Caule volcanic ash cloud from the MODIS-based retrievals
4 (Osoreo et al., 2015). Right: Predicted column mass (g m^{-2}) with NMMB-MONARCH-ASH for a) 6 June at
5 14:25 UTC and, b) 8 June at 14:15 UTC.



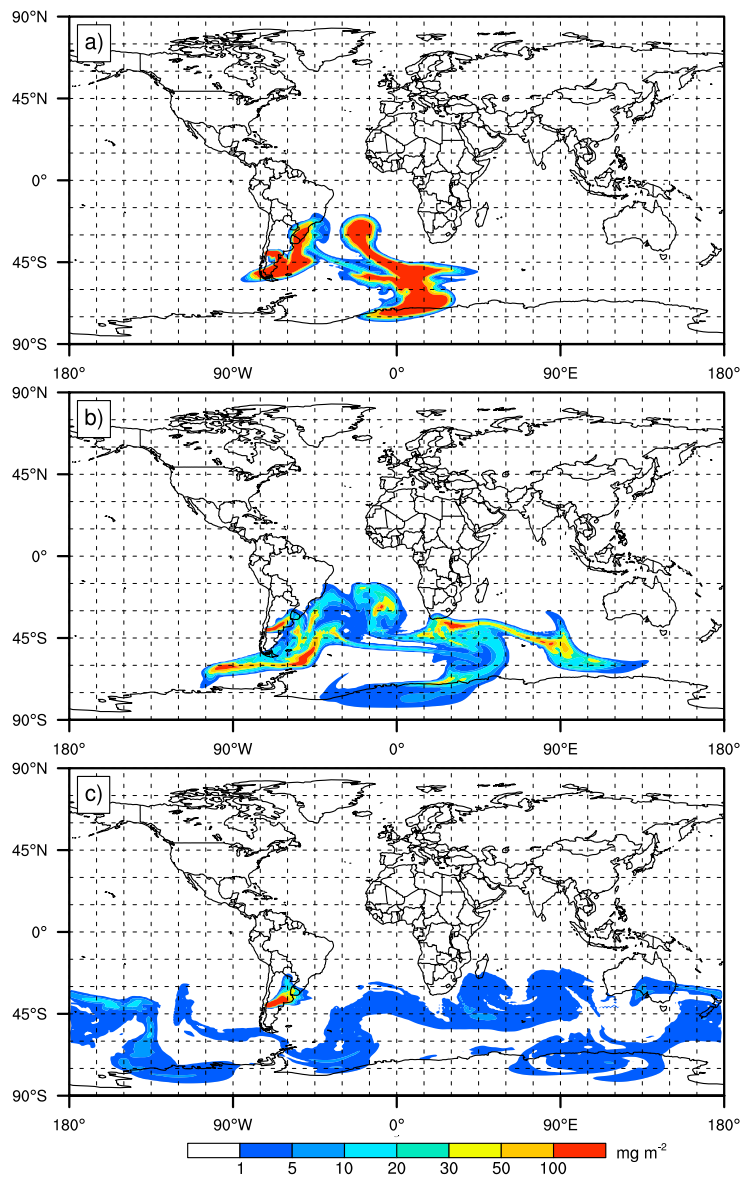
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3 Figure 5. Left: Isopach maps in centimeter of layers A-B, A-F, H, and K2. Dashed lines infer the limit of the
 4 deposits presented in Pistolesi et al. (2015b). Right: Corresponding NMMB-MONARCH-ASH computed
 5 thicknesses (cm). Key locations in blue include San Carlos de Bariloche (SCB) and Ingeniero Jacobacci (IJ), 90
 6 and 240 km east of the vent)



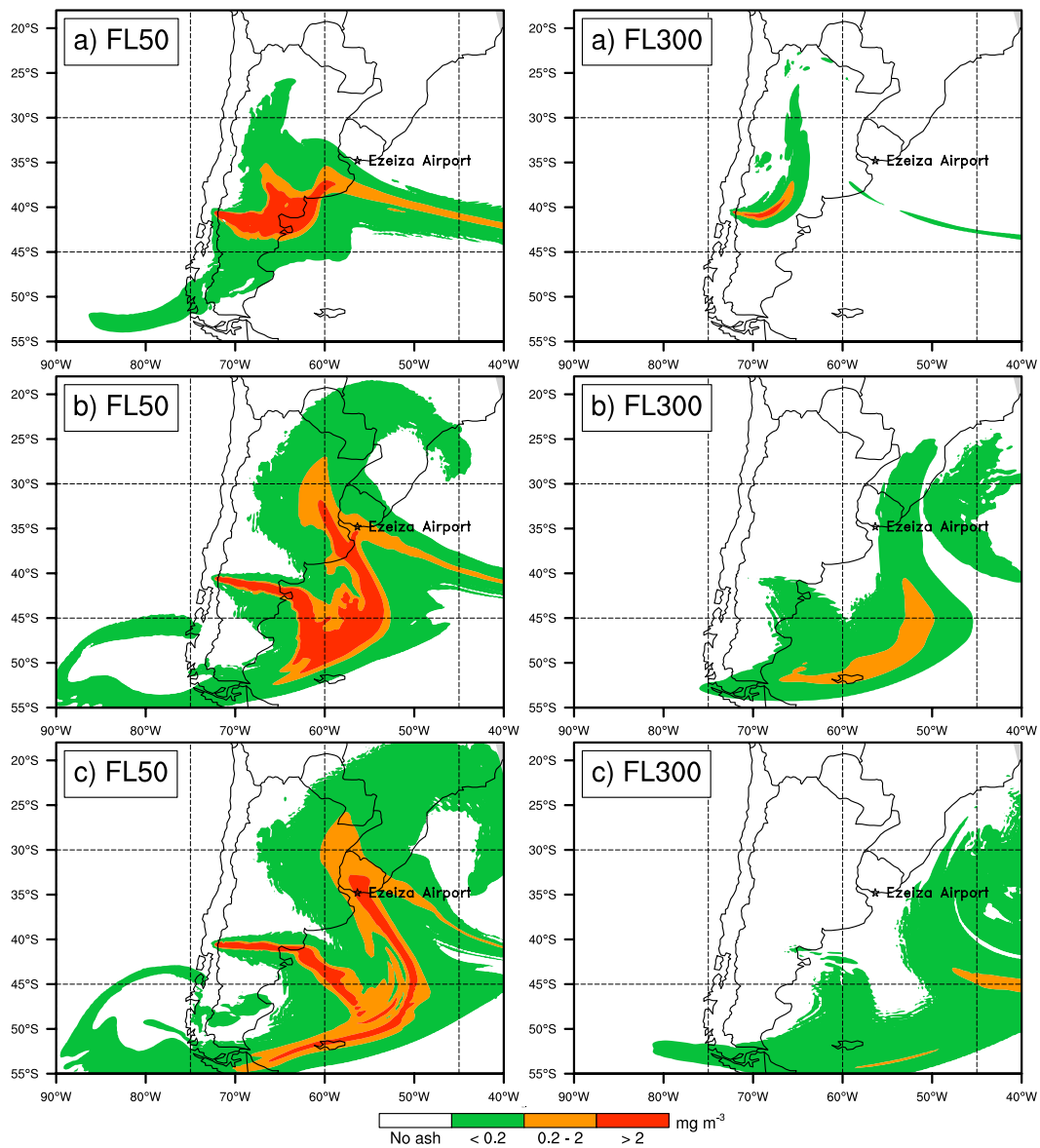
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2 Figure 6. Left: measured ground deposit isopachs (kg m^{-2}) for the period beginning on 4 June until 30 June.
 3 Dashed lines infer the limit of the deposits (modified from Collini et al., 2013). Right: Predicted deposit load (kg
 4 m^{-2}) with NMMB-MONARCH-ASH at the end of the simulation. Key locations in blue include San Carlos de
 5 Bariloche (SCB; 90 km from the vent), Ingeniero Jacobacci (IJ; 240 km east of the vent), and Trelew and
 6 Viedma (~ 600 km SE and NE of the vent, respectively).

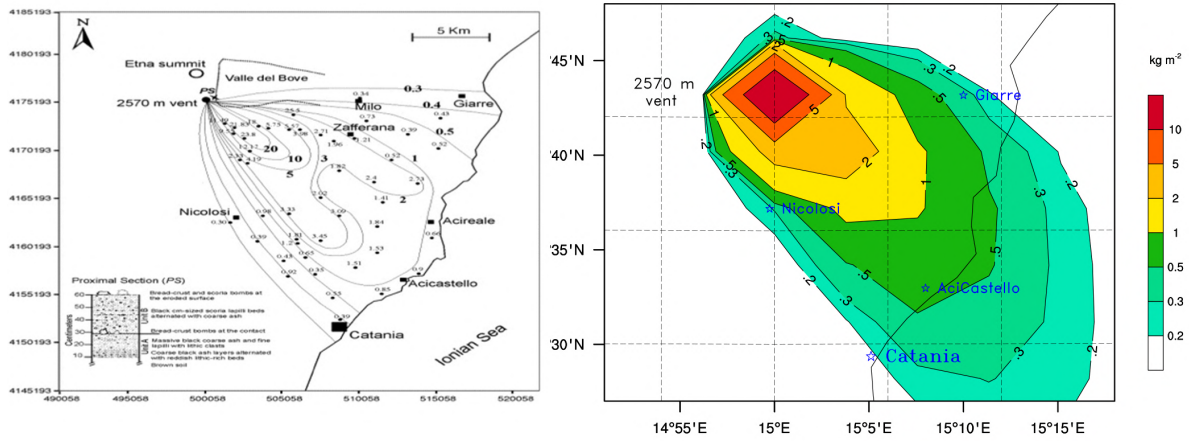


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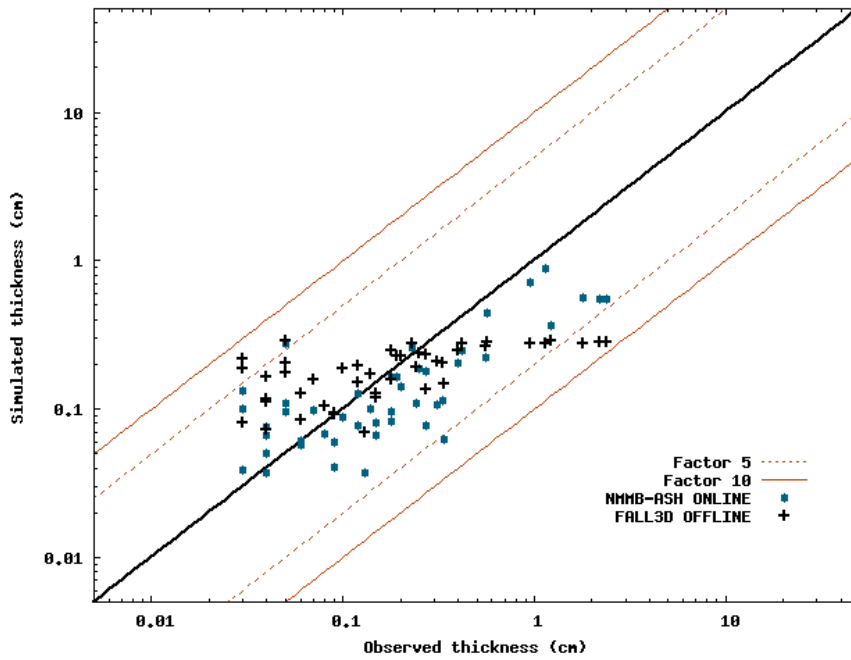
2 Figure 7. NMMB-MONARCH-ASH total column concentration (mass loading; mg m^{-2}) from our global
 3 simulation. Results for a) 8 June at 09:00 UTC, b) 10 June at 04:00 UTC, and c) 14 June at 06:00 UTC.



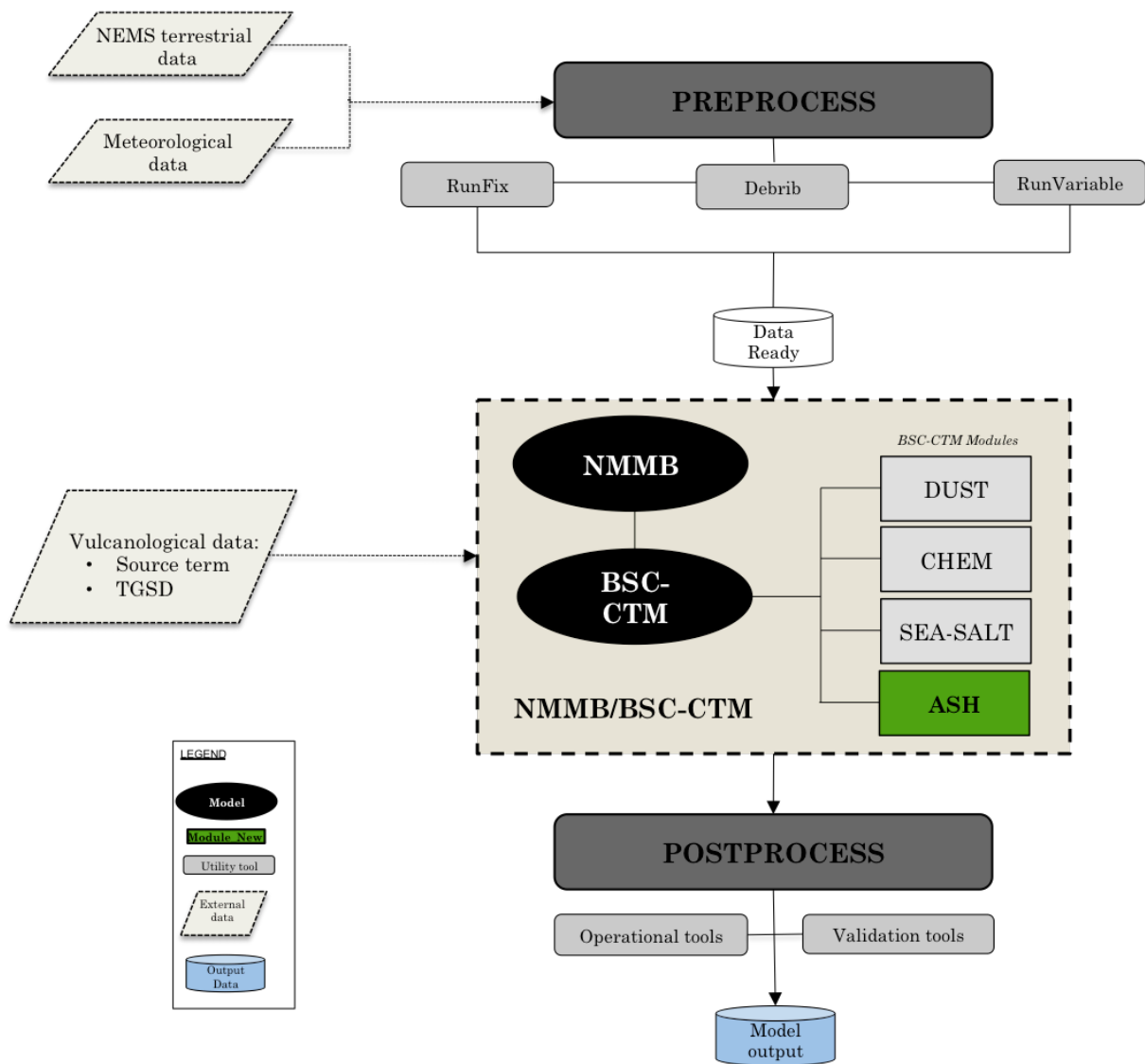
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 2 Figure 8. NMMB-MONARCH-ASH Flight level ash concentrations (mass loading; mg m⁻³) before and after
 3 closure of the Buenos Aires (Ezeiza) airport and air space. Results for FL50 (left) and FL300 (right) for a) 6 June
 4 at 11:00 UTC, b) 7 June at 04:00 UTC, and c) 7 June at 12:00 UTC. Safe ash concentration thresholds are shown
 5 (red contours illustrate “No Flying” zones).



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 2 Figure 9. Left: Isomass map of the tephra deposit formed between 21 and 24 July 2001. Curves are given in kg
 3 m⁻². Coordinates are given in UTM-Datum ED50 (Scollo et al., 2007). Right: Modeled deposit load (kg m⁻²)
 4 with NMMB-MONARCH-ASH at the end of the event.

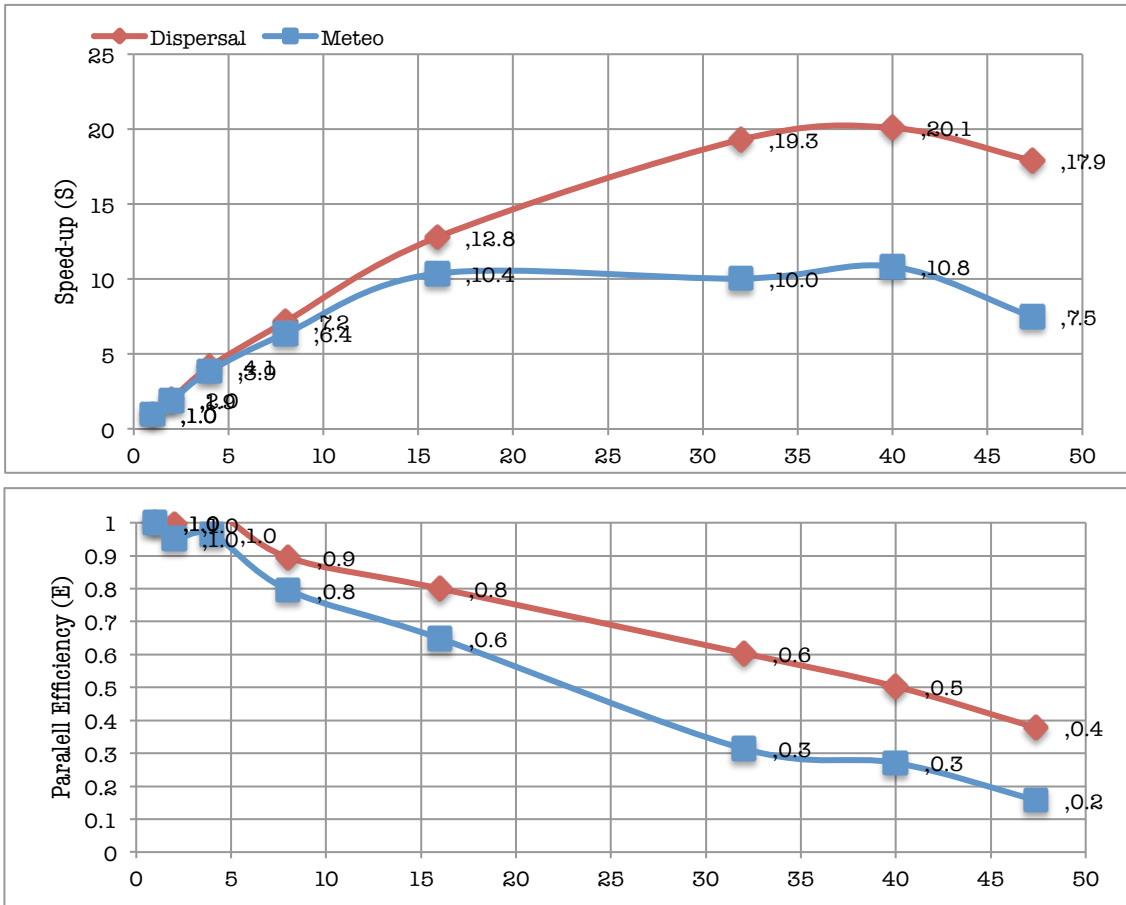


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 6 Figure 10. Simulated versus observed thicknesses for the reconstruction of the 2011 Etna eruption with NMMB-
 7 MONARCH-ASH (circles) and FALL3D (crosses). The solid bold line represents a perfect agreement, while the
 8 dashed and solid thin orange lines mark the region that is different from observed thicknesses by a factor 5 (1/5)
 9 and 10 (1/10), respectively.



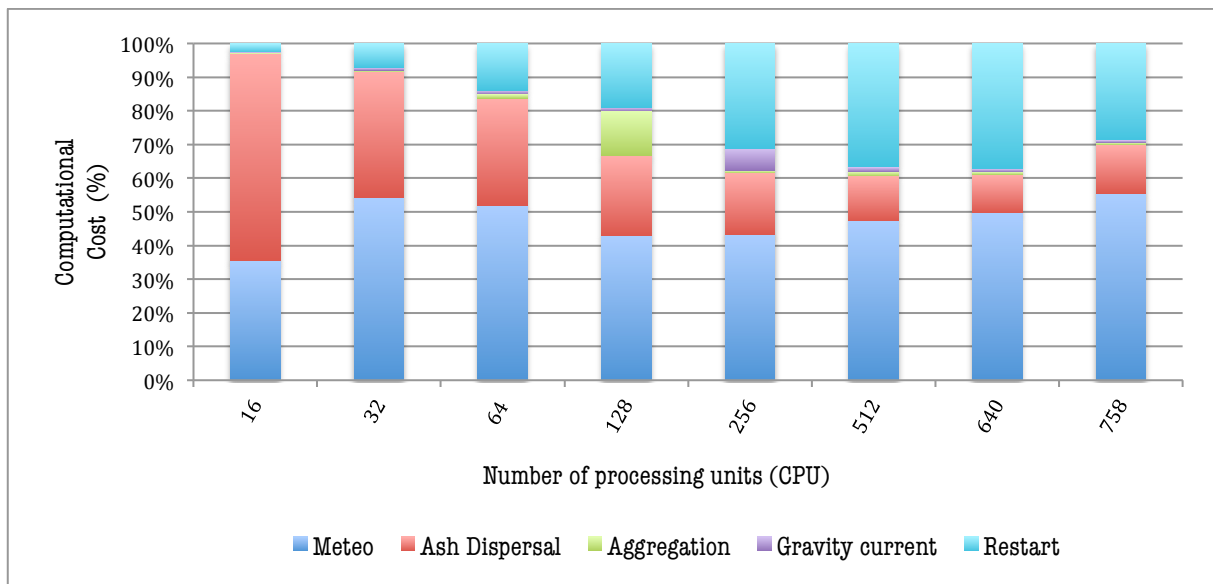
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2 Figure 11. Schematic representation of the operational implementation of NMMB-MONARCH-ASH.



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 2 Figure 12. NMMB-MONARCH-ASH scalability results. Top: parallel speed-up (S; computational speed) for
 3 meteorology only (blue) and for meteorology and dispersal combined (red). Bottom: parallel efficiency (E)
 4 versus number of computation nodes employed.

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2 Figure 13. NMMB-MONARCH-ASH relative computational cost (%) with increasing CPUs. Represented
 3 processes include: Meteorology (blue); Ash dispersal for 10 bins (red); Aggregation (green); Gravity current
 4 (purple) and; Restart (light blue).

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