



1 **Air–surface exchange of gaseous mercury over permafrost soil: an**
2 **investigation at a high-altitude (4700 m a.s.l.) and remote site in the central**
3 **Qinghai-Tibet Plateau**

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9 **Abstract.** The pattern of air–surface gaseous mercury (mainly Hg(0)) exchange in the Qinghai-Tibet
10 Plateau (QTP) may be unique because this region is characterized by low temperature, great
11 temperature variation, intensive solar radiation, and pronounced freeze-thaw process of permafrost soils.
12 However, air–surface Hg(0) flux in the QTP is poorly investigated. In this study, we performed field
13 measurements and controlled field experiments with dynamic flux chambers technique to examine the
14 flux, temporal variation and influencing factors of air–surface Hg(0) exchange at a high-altitude (4700
15 m a.s.l.) and remote site in the central QTP. The results of field measurements showed that surface soils
16 were net emission source of Hg(0) in the entire study. Hg(0) flux showed remarkable seasonality with
17 net high emission in the warm campaigns and net low deposition in winter campaign, and also showed
18 the diurnal pattern with emission in daytime and deposition in nighttime, especially on days without
19 precipitation. Rainfall events on the dry soils induced large and immediate increase in Hg(0) emission.
20 Snowfall events did not induce the pulse of Hg(0) emission, but snow melt resulted in the immediate
21 increase in Hg(0) emission. Daily Hg(0) fluxes on rainy or snowy days were higher than those of days
22 without precipitation. Controlled field experiments suggested that water addition to dry soils
23 significantly increased Hg(0) emission both in short and relatively long timescales, and also showed that
24 UV radiation was primarily attributed to Hg(0) emission in the daytime. Our findings imply that a warm
25 climate and environmental change could facilitate Hg release from the permafrost terrestrial ecosystem
26 in the QTP.



27 **1 Introduction**

28 Soils represent the largest Hg reservoirs in ecosystems and play a major role in the global Hg cycle
29 (Selin, 2009; Agnan et al., 2016). Background soils receive Hg input from atmospheric deposition,
30 which is mainly retained in organic-rich layers of upper soils (Schuster, 1991; Khwaja et al., 2006).
31 Under favorable conditions, Hg in soils can be reduced to Hg(0) and then emitted to the overlaying air
32 because of its high volatility (Schlüter, 2006). Therefore, soils can serve as both sources and sinks of
33 atmospheric Hg (Pirrone and Mason, 2009; Amos et al., 2013; Agnan et al., 2016).

34 In the past several decades, efforts have been made to improve the understanding of soil Hg
35 biogeochemistry (Zhang and Lindberg, 1999; Lin et al., 2010; Schlüter, 2006; Jiskra et al., 2015).
36 Measurements across various types of soils and climates show that air–soil Hg(0) exchange has highly
37 spatial and temporal variation and bidirectional exchange behavior (Agnan et al., 2016 and references
38 therein). Field measurements and laboratory experiments highlight that various factors and processes
39 influence air–surface Hg(0) exchange, including concentrations and species of soil Hg (Gustin et al.,
40 1999, 2002; Hintelmann et al., 2002; Bahlmann et al., 2006; Kocman and Horvat, 2010; Eckley et al.,
41 2011; Edwards and Howard, 2013; Mazur et al., 2015), solar radiation (Gustin et al., 2002; Moore and
42 Carpi, 2005; Gustin et al., 2006; Xin et al., 2007; Fu et al., 2008a; Kocman and Horvat, 2010; Park et al.,
43 2014), precipitation (Lindberg et al., 1999; Gustin and Stamenkovic, 2005; Gabriel et al., 2011), soil
44 temperature and moisture (Gustin et al., 1997; Gustin and Stamenkovic, 2005; Ericksen et al., 2006; Xin
45 et al., 2007; Briggs and Gustin, 2013; Park et al., 2014; Mazur et al., 2015), soil organic matter and pH
46 (Yang et al., 2007; Xin and Gustin, 2007; Mauclair et al., 2008); land cover (Dommergue et al., 2003;
47 Ericksen et al., 2005; Cobbett et al., 2007; Gabriel and Williamson, 2008; Zhu et al., 2011; Durnford et
48 al., 2012a, b; Toyota et al., 2014a, b); atmospheric Hg(0) concentrations and other chemical
49 compositions (Engle et al., 2004; Xin and Gustin, 2007; Fu et al., 2008a), biological activity (Choi and
50 Holsen, 2009), as well as atmospheric turbulence (Gustin et al., 1997; Poissant et al. 1999). Existing
51 studies on Hg(0) dynamics at air–surface interface are mainly performed in temperate regions (Agnan et
52 al., 2016 and reference therein). The seasonal frozen soils and permafrost widely distribute, accounting
53 for almost 70% of terrestrial area of Earth (NSIDC). However, the knowledge of air–surface Hg(0)



54 dynamics in cold region is limited (Cobbett et al., 2007; Durnford and Dastoor, 2011). Most current
55 parameters of air–soil Hg(0) exchange applied in Hg biogeochemical models are mainly derived from
56 temperate regions of North America and Europe (Zhu et al., 2016).

57 The Qinghai-Tibet Plateau (QTP) is located in the western China with the area of 2.5 million km²
58 and mean altitude of > 4000 m. Due to the high altitude and subsequent low temperature, a significant
59 portion (~ 1.5 million km²) of the QTP is underlain by permafrost (Kang et al., 2010). Because of the
60 harsh natural environment, limited research resources and difficulty of access and sampling logistics,
61 studies on Hg biogeochemistry in the QTP are limited. The role of QTP in the regional and global Hg
62 biogeochemical cycle is poorly understood (Ci et al., 2012; Agnan et al., 2016). At present, Hg studies
63 in the QTP mainly focused on the investigations of Hg concentration, speciation and distribution in
64 environmental samples, such as air (Fu et al., 2008b; 2012; Yin et al., 2015), snow and glacier (Loewen
65 et al., 2007; Wang et al., 2008; Zhang et al., 2012; Huang et al., 2012), and rain water (Huang et al.,
66 2013). The knowledge of Hg(0) dynamics at air–surface interface in the QTP is extremely poor. The
67 unique climatic condition, land cover and soil property suggest the need for the specific air–soil Hg(0)
68 flux data and mechanism representative of the environmental setting in the QTP to better constrain
69 global natural sources inventories (Ci et al., 2012; Agnan et al., 2016).

70 In this study, we applied dynamic flux chambers (DFCs) technique to investigate the flux,
71 temporal variation and influencing factors of air–surface Hg(0) exchange at a representative research
72 station in the central QTP. Meanwhile, controlled field experiments were performed to explore the
73 effect of rainfall and different wavebands of solar radiation on air–soil Hg(0) flux. Combining the result
74 of this study and other knowledge, we discuss the effect of future climatic and environmental change on
75 air–surface Hg(0) dynamics in the QTP.

76 **2 Methods**

77 **2.1 Study site**

78 The study was performed at the Beiluhe Permafrost Engineering and Environmental Research
79 Station affiliated to the Cold and Arid Regions Environmental and Engineering Research Institute,



80 Chinese Academy of Sciences (CAREER–CAS). The elevation of the Beiluhe region is about 4700 to
81 4800 m a.s.l. The station (34° 49' 45" N, 92° 56' 06" E) lies over the continuous permafrost zone in the
82 central QTP (Fig. 1). The terrain is undulation with sparse vegetation and surface fine sands or gravels.
83 The thickness of the active layer and permafrost around the station is 2.0–3.2 m and 60–200 m,
84 respectively; the active layer begins to freeze in September and thaw in May (Peng et al., 2015a). The
85 Beiluhe region experiences a continental climate with cold winter (up to –30 °C) and warm summer (up
86 to 25 °C). The magnitude of daily air temperature is up to 30 °C, and the annual mean surface air
87 temperature is about –2 to –3 °C (Peng et al., 2015a). The solar radiation is high and characterized by
88 intense UV radiation (Wei et al., 2006). The mean annual precipitation is about 300 mm and mostly
89 occurs during the May to October under the influence of the Southern Asian Monsoon; the annual
90 potential evapotranspiration (~1300 mm) greatly exceeds the precipitation (Peng et al., 2015a). As a
91 remote region, there is no direct human activity to influence the local Hg cycle.

92 2.2 Measurement of air–surface Hg(0) flux

93 The dynamic flux chambers (DFCs) technique was widely used to investigate Hg(0) flux between
94 air–surface interface because it is inexpensive, portable, easy to set up and operate (e.g., Kim and
95 Lindberg, 1995; Carpi and Lindberg, 1998; Gustin et al., 2006; Wang et al., 2006; Dommergue et al.,
96 2007; Fu et al., 2008a; Kocman and Horvat, 2010; Edwards and Howard, 2013). The principle for
97 measuring air–surface Hg(0) flux using the DFCs technique involves placing a chamber over a surface
98 and measuring the difference in air Hg(0) concentrations at inside and outside of the chamber. DFC
99 means continuously draw ambient air into the chamber through inlets at a set flushing flow rate. Air–
100 surface Hg(0) flux was calculated using Eq. (1),

$$101 \quad F = Q \frac{C_o - C_i}{A} \quad (\text{Eq. 1})$$

102 where F is the Hg(0) flux ($\text{ng m}^{-2} \text{h}^{-1}$), Q is the flushing flow through the chamber ($\text{m}^3 \text{h}^{-1}$), A is the
103 footprint of the chamber (m^2), C_o and C_i (ng m^{-3}) is air Hg(0) concentrations at outlet and inlet of the
104 chamber, respectively. Positive flux values indicate Hg(0) emission from the surface into the air;
105 negative flux values represent Hg(0) deposition to the surface from the air. The advantage and



106 limitation of DFCs technique for determining Hg(0) flux have been extensively discussed in previous
107 studies (e.g., Wallschlager et al., 1999; Gillis and Miller, 2000; Lindberg et al., 2002; Eckley et al.,
108 2010; Lin et al., 2012; Sommar et al., 2012; Zhu et al., 2015a, b).

109 In this study, quartz chambers were constructed for measuring Hg(0) flux and exploring the effect
110 of different rainfall depths and radiation condition on the Hg(0) flux. Quartz glass has many advantages
111 as construction material of chamber for determining Hg(0) flux in background soils. First, it has high
112 transmittance of the full spectrum of solar radiation, especially UV waveband (Fig. S1 in Supplement).
113 Therefore, quartz chamber is suitable to determine the more “actual” Hg(0) flux because the short
114 wavelength of solar radiation has been found to have important effect on Hg(0) dynamics at the air–soil
115 interface (Moore and Carpi, 2005; Bahlmann et al., 2006). Second, it has low potential for Hg(0)
116 adsorption and is easy to clean by heating to remove Hg bonding on the surface (Ci et al., 2016a). This
117 will decrease the systematic blank of measurement, which is critical for investigating Hg(0) flux over
118 background soils (Carpi and Lindberg, 1998).

119 Our semi-cylindrical quartz chamber was 8 cm high and 24 cm length with a footprint of 0.0384 m²
120 (0.16 m x 0.24 m) and an internal volume of 2.41 L, which is similar to previous studies (Eckley et al.,
121 2010 and reference therein). The chamber had nine inlets (8 mm in diameter) and three outlets which
122 were on the two opposite section of the chamber. The inlet sampling tube was placed near the ground
123 surface (3 cm above the surface) directly near the inlet of the chamber. A flushing ambient air was
124 drawn by vacuum pump (KNF, Inc. Germany) with 3.0 L min⁻¹ (0.18 m³ h⁻¹) through the chamber.

125 Since the harsh environment condition and the unstable power supply, the usage of the commercial
126 automatic Hg analyzer (such as Tekran 2537) to conduct filed measurements of Hg(0) flux is
127 challenging in the Beiluhe station. Therefore, air Hg(0) concentrations in both inlet and outlet of the
128 chamber were monitored manually by gold trap simultaneously with a 2–3 h intervals (Ci et al., 2016b).
129 The air was pumped through gold trap using air pump (KNF, Inc. Germany) with 0.50 L min⁻¹ (0.03 m³
130 h⁻¹). Hg(0) collected on gold traps was quantified on site by a cold vapor atomic fluorescence
131 spectrophotometer (CVAFS, Model III, Brooks Rand, USA) using two-stage gold amalgamation
132 method (Fitzgerald and Gill, 1979; Ci et al., 2011, 2016a). Gold trap efficiencies were determined in



133 laboratory (Hg(0) concentration: 3.2 to 13.4 ng m⁻³) and field (Hg(0) concentration: ~1 to 2 ng m⁻³ to
134 9.4 ng m⁻³), and multiple measurements using gold traps in series showed no breakthrough at the sample
135 flow rate of 0.50 L min⁻¹ for 5 hours. The method detection limit was 0.03 ng m⁻³ and precision was
136 3±2%.

137 The turnover time obtained from this protocol was 0.68 min, which is similar to previous studies
138 (Eckley et al., 2010 and reference therein). The flow rates of air both the inlet and outlet of chamber
139 were adjusted by a needle valve and controlled by a rotameter. Prior to the measurement, the rotameters
140 were calibrated by a mass flow meter and a volumetric gas meter. The accuracy of flow rate was ±3%.
141 The sampling control system was installed in the tent near the soil plot (<2 m).

142 In this study, a bare soil plot of 2 x 2 m was chosen and separated to four subplots (1 x 1 m, labeled
143 as Subplot A to D) to measure Hg(0) flux and further to investigate the effects of different wavelengths
144 of solar radiation on the Hg(0) exchange. Geologic unit of this soil plot is representative of major
145 lithologic units of the Beiluhe region. The study was conducted in four campaigns (May, September,
146 and December 2014 and May–June 2015), covering typical intra-annual meteorological condition in the
147 study region. It is well-known that the application of the chamber isolates the soil surface from air
148 turbulence and rain/dew/frost/snow, altering many environmental parameters that influence Hg(0)
149 production/consumption and deposition/emission at air–soil interface (Sommar et al., 2012 and
150 references therein). In the Beiluhe region, the rain/snow/dew/frost commonly occurs in short time scale
151 (10¹–10² minutes) because of the high-altitude location and great variation of temperature and unstable
152 weather conditions. To minimize the effect of great variation of soil surface condition on Hg(0) flux
153 measurements, when only to measure Hg(0) flux, we used the same chamber to measure Hg(0) flux in
154 four subplots (Subplot A to D) in turn. For the rainy and snowy days, if possible, the chamber was
155 moved to next subplot generally after the rain or snow because studies confirmed the significant
156 influence of precipitation on Hg(0) flux in short time scale (Lindberg et al., 1999; Gustin and
157 Stamenkovic, 2005; Johnson et al., 2003; Song and Van Heyst, 2005; Lalonde et al., 2001, 2003;
158 Dommergue et al., 2003, 2007, 2012; Fañ et al., 2007; Bartels-Rausch et al., 2008; Brooks et al., 2008;
159 Mann et al., 2014, 2015). For the days without precipitation, the chamber was moved to next subplot
160 before sunrise to capture the effect of frost or dew on air–surface Hg(0) flux in the light.



161 In this study, all materials in contact with Hg(0) were quartz, Teflon or borosilicate glass. The
162 chambers and tubing were rigorously acid washed (Ci et al., 2016a). The quartz chambers were heated
163 to 650 °C for 2 h prior the measurement to further remove all Hg (Ci et al., 2016a). System blanks of
164 the four chamber systems were systemically inspected on site by placing acid-cleaned Teflon filter
165 beneath the chamber and routinely inspected before and after the measurements under the field
166 condition. The overall blank results were taken as the average of all chamber blanks for each particular
167 day or from the entire period if continuous monitoring was conducted. The blank values of the four
168 chambers were found to be very low (mean \pm SD: -0.02 ± 0.03 ng m⁻² h⁻¹; range: -0.20 – 0.07 ng m⁻² h⁻¹)
169 and were not significantly different ($p > 0.05$) throughout the entire study. We also found that the mean
170 blank values were not significantly different ($p > 0.05$) from zero and were insignificant compared to the
171 measured Hg(0) fluxes (see below), then Hg(0) flux data was not blank-corrected in this study.

172 **2.3 Determination of soil Hg**

173 Surface soil samples (0–2 cm) were collected from four soil subplots during June 2014 campaign.
174 Soil samples were freeze-dried and homogenized for total Hg determination using a Milestone's DMA
175 direct Hg analyzer (detection limit: 0.01 ng Hg or 0.15 ug kg⁻¹) following the EPA Method 7473
176 (Briggs and Gustin, 2013), and the analytical accuracy was 3%.

177 **2.4 Measurements of environmental variables**

178 A meteorological station that located 60 m from the soil plot was used to collect the following
179 environmental variables: air temperature (°C), relative humidity (%), wind speed (m s⁻¹), precipitation
180 (mm), photosynthetically active radiation (PAR, μ mol m⁻² s⁻¹), surface soil temperature (°C). The
181 surface soil temperature was monitored by an ASI-111 Precision Infrared Radiometer (Campbell
182 Scientific Inc. USA). This sensor was installed on a bar 0.8 m above the soil surface and provided a
183 non-contact measurement of the surface temperature. We also measured the soil temperature at 1.0 cm
184 soil depth at the inside of chamber, and no significant difference was found between the soil
185 temperature in the outside and inside of chamber. Details on the measurements of environmental
186 variables are given in Peng et al. (2015a). The thickness of snowpack was measured manually.

187 **3 Results and Discussion**



188 3.1 Soil Hg

189 Soil Hg concentrations of four subplots varied from 13.11 ± 0.51 to $12.83 \pm 0.81 \mu\text{g kg}^{-1}$,
190 suggesting that the study region is a typical background soil for Hg. Soil Hg concentrations and air–soil
191 Hg(0) flux of four subplots were not statistically different ($p > 0.05$), indicating the properties of the soil
192 plot were homogeneous.

193 3.2 Hg(0) in ambient air

194 Figure 2 shows the temporal variation of air Hg(0) concentrations inside and outside of chamber,
195 air–surface Hg(0) flux and environmental variables during four campaigns in 2014–2015. Hg(0)
196 concentrations of ambient air ranged from 0.93 to 1.78 ng m^{-3} with a mean of $1.36 \pm 0.17 \text{ ng m}^{-3}$
197 ($n=361$), which were slightly lower than those in typical Northern Hemisphere background region (~ 1.5
198 ng m^{-3} , Sprovieri et al., 2007; Ebinghaus et al., 2011). To our knowledge, four measurements (including
199 this study) have been conducted to determine atmospheric Hg over the QTP (see Fig. 1 for the
200 locations). The gradient of atmospheric Hg(0) over the QTP was characterized by high concentrations in
201 Mt. Gongga ($3.98 \pm 1.62 \text{ ng m}^{-3}$, Fu et al., 2008b) and Mt. Waliguan ($1.98 \pm 0.98 \text{ ng m}^{-3}$, Fu et al.,
202 2012), moderate concentrations in the Beiluhe ($1.36 \pm 0.17 \text{ ng m}^{-3}$, this study), and low concentrations
203 in the Nam Co ($0.96 \pm 0.19 \text{ ng m}^{-3}$, Yin et al., 2015). It seems that the sampling station with relatively
204 long distance from the source region of atmospheric Hg(0) (i.e., the Central China, Streets et al., 2005)
205 had relatively low atmospheric Hg(0) concentrations (Fig. 1). Atmospheric Hg(0) concentrations were
206 high in three warm campaigns and low in winter campaign in the Beiluhe station (Fig. 3), which is
207 consistent with Mt. Gongga (Fu et al., 2008b) and Mt. Waliguan (Fu et al., 2012).

208 3.3 Air–surface Hg(0) flux and influencing factors

209 3.3.1 Temporal variation of Hg(0) flux

210 The mean of air–surface Hg(0) flux in the entire study period were $2.86 \text{ ng m}^{-2} \text{ h}^{-1}$ ($25.05 \mu\text{g m}^{-2} \text{ y}^{-1}$)
211 ¹), indicating that surface soils were net emission source of Hg(0). Hg(0) flux in this study is
212 comparable to those over background soils (-10 – $10 \text{ ng m}^{-2} \text{ h}^{-1}$, Wang et al., 2006; Erickson et al., 2006;
213 Fu et al., 2008a), but greatly lower than those over Hg-enriched soils (10^2 – $10^3 \text{ ng m}^{-2} \text{ h}^{-1}$, Gustin et al.,



1999; Wang et al., 2007; Edwards and Howard, 2013), indicating that soil Hg concentrations may be the dominant factor for controlling the magnitude of Hg(0) emission flux (Agnan et al., 2016 and reference therein).

Figure 2 shows that the Hg(0) flux was highly variable. The highest Hg(0) emission fluxes of $28.46 \text{ ng m}^{-2} \text{ h}^{-1}$ were observed at 13:00–15:00 on 13 September 2014 after a rain event on dry soils. The highest Hg(0) deposition fluxes of $-6.24 \text{ ng m}^{-2} \text{ h}^{-1}$ were observed at nighttime over the cold and dry surface soil during December 2014 campaign (03:30–06:30, 24 December 2014).

Hg(0) flux generally showed the diurnal pattern with high emission in daytime and remarkable deposition in nighttime, especially on days without precipitation (Fig. 2). Many studies have confirmed that solar radiation is one of the most important drivers for soil Hg(0) emission (Xin and Gustin, 2007; Choi and Holsen, 2009; Kocman and Horvat, 2010); high surface temperature also facilitates Hg(0) production and subsequent emission (Park et al., 2014). Therefore, the two environmental variables jointly regulate the diurnal pattern of Hg(0) flux. An in-depth discussion on synergistic effects of solar radiation and surface temperature on Hg(0) flux is provided in below. Interestingly, the diurnal pattern of Hg(0) flux of each day during December 2014 campaign was almost identical, which may be associated with very similar weather conditions throughout the entire campaign.

Hg(0) flux showed pronounced seasonality with high emission in three warm campaigns and net low deposition during winter campaign (Fig. 3). Similar seasonality has been reported in many other studies (e.g., Gabriel et al., 2006). As discussed below, in warm seasons, some environmental variables, such as high solar radiation, surface temperature and precipitation, facilitate the soil Hg(0) emission.

3.3.2 Effect of precipitation (rain and snow) on air–surface Hg(0) flux

Many studies addressed that precipitation greatly influences air–surface Hg(0) flux over different timescales (Lindberg et al., 1999). In previous studies, investigators mainly focused on the effect of rainfall/watering on air–soil Hg(0) flux (Lindberg et al., 1999; Johnson et al., 2003; Gustin and Stamenkovic, 2005; Song and Van Heyst, 2005; Corbett-Hains et al., 2012) or the fate and transport of Hg(0) at air–snow interface (Lalonde et al., 2001, 2003; Ferrari et al., 2005; Dommergue et al., 2003, 2007; Fañ et al., 2007; Bartels-Rausch et al., 2008; Brooks et al., 2008; Steen et al., 2009; Durnford et



241 al., 2012a, b; Mann et al., 2015). The field study on the effect of snowmelt on Hg(0) flux is very limited
242 (Cobbett et al., 2007). The precipitation at the Beiluhe station mainly occurs during May to October
243 (Peng et al., 2015a). Due to the high-altitude location of the Beiluhe station, snow event commonly
244 occurs in May to June and late September to October. Because of intensive solar radiation and surface
245 temperature, the snow melts or sublimates in short time scale (10^0 – 10^2 hour), i.e., little/no snow
246 accumulation occurs for long time (>3 day). Therefore, the Beiluhe region provides an unique
247 opportunity to investigate the different effects of rain, snow, and snowmelt on the air–surface Hg(0)
248 flux over different timescales.

249 Firstly, the effect of rain events on Hg(0) flux was investigated. Many studies reported that the
250 rainfall or watering of dry soils can promote soil Hg(0) emission immediately (Lindberg et al., 1999;
251 Gustin and Stamenkovic, 2005). We also found that the Hg(0) emission flux increased immediately
252 following the rainfall (Fig. 2). This finding is consistent with many other studies (Gustin and
253 Stamenkovic, 2005; Johnson et al., 2003; Lindberg et al., 1999; Song and Van Heyst, 2005).
254 Investigators supposed that the dramatic increases in Hg(0) emission may be attributed to the physical
255 displacement of Hg(0) present in soil air and desorption of loosely bound Hg(0) on soil particles by the
256 infiltrating water (Johnson et al., 2003; Gustin and Stamenkovic, 2005). Notably, Fig. 2 displays that the
257 pulse of Hg(0) emission after the rainfall was also observed at nighttime (such as 0:00 to 01:00 on 4
258 September 2014). Similar phenomenon was also documented by our controlled experiments (see below).
259 This indicates that the immediate increase in Hg(0) emission might not be controlled by photochemical
260 processes but by physical processes.

261 Soil moisture condition may also significantly regulate Hg(0) flux over relatively long timescales
262 (from hours to several days). Therefore, many experiments studied the effect of water addition on the
263 magnitude and pattern of air–soil Hg(0) flux over different timescales (Johnson et al., 2003; Gustin and
264 Stamenkovic, 2005). However, most of studies were performed in controlled laboratory or mesocosm
265 settings under certain well-defined, but not necessarily environment relevant conditions (Johnson et al.,
266 2003; Gustin et al., 2004; Gustin and Stamenkovic, 2005; Song and Van Heyst, 2005; Kocman and
267 Horvat, 2010; Corbett-Hains et al., 2012; Park et al., 2014). In this study, it was also challenging to
268 reveal the effect of rainfall on Hg(0) flux over relatively long timescales via field measurement since



269 the intermittent rain events occurred irregularly during June 2014 and September 2014 campaign (Fig.
270 2). For a better understanding of the effect of water addition, we chosen another similar soil plot with
271 homogeneous soil property to conduct controlled field experiments to explore the effect of different
272 rainfall depths on Hg(0) flux over different timescales (from minutes to hours). The controlled
273 experiments were performed during May–June 2015 campaign since this period had high surface
274 temperature and low precipitation (Fig. 2). In the Beiluhe station, hourly rainfall depth rarely exceeds
275 15 mm (Peng et al., 2015a). Therefore, we designed four different treatments of rainfall depth (0 mm, 1
276 mm, 5 mm, and 15 mm). The water addition to the dry soils commenced at night (01:40) on 30 May
277 2015 to exclude the effect of photochemical process in the first hours of experiments. We added the
278 Milli-Q water (Hg concentration $< 0.2 \text{ ng L}^{-1}$) to the inside and outside of the chamber by pre-clean
279 plastic syringe within 10 min to simulate the three different rain depths. Four chambers were used to
280 simultaneously measure Hg(0) flux over these four treatments for 22 hours (from 01:00 to 23:00) with
281 the same protocol described in the Methods section. Hg(0) flux was measured with 20 min intervals in
282 the first hours (from 01:00 to 04:20) of the experiments to investigate the temporal variation of Hg(0)
283 flux in the short timescale, and with 1 hour intervals for the rest period of the experiments.

284 The high-time resolution measurements captured the immediate and dramatic increases in Hg(0)
285 emission flux after the watering of dry soils (Fig. 4). The baseline Hg(0) flux of 0 mm treatment was
286 used as the benchmark for the different rainfall depth treatments to be compared against. Obviously, the
287 higher amount of water addition resulted in longer duration and higher accumulative flux of Hg(0)
288 emission pulse. The duration of Hg(0) emission pulse for 1 mm and 5 mm treatment was < 20 min
289 (from 01:40 to 02:00) and ~ 40 min (from 01:40 to 02:20), respectively, which was lower than that of 15
290 mm treatment (~ 80 min, from 01:40 to 03:00). The duration of Hg(0) emission for 15 mm treatment in
291 the daytime was also longer than that of 1 mm and 5 mm treatment (Fig. 4).

292 As shown in Fig. 4, the cumulative flux of Hg(0) emission during the entire study period mainly
293 included two fractions: the pulse of Hg(0) emission after the watering (i.e., emission flux by watering),
294 and the Hg(0) emission during the daytime (i.e., emission flux by radiation). Figure 5 displays that both
295 “emission flux by watering” and “emission flux by radiation” for 15 mm treatment were significantly
296 higher than those of 1 mm and 5 mm treatment. As mentioned above, the dramatic increase in Hg(0)



297 emission after the simulated rain can be explained by physical displacement of interstitial soil air by the
298 infiltrating water. The long emission duration and large immediate emission flux for soil plot with high
299 water addition can be explained by that the more water needed longer time to percolate the soil column
300 and displaced more soil Hg(0). Many previous studies suggested that the magnitude of Hg(0) emission
301 with a rainfall or stimulated rain depended on soil moisture condition, i.e., if the amounts of water
302 received by the soils was less than needed to saturate, the soil surface showed an immediate increase in
303 Hg(0) emission; after the soil became saturated, Hg(0) emission from surface soil was suppressed
304 (Klusman and Webster, 1981; Lindberg et al., 1999; Johnson et al., 2003; Gustin and Stamenkovic,
305 2005). In this study, the pulse of Hg(0) emission flux for 15 mm treatment was significantly higher than
306 that of 5 mm and 1 mm treatment (Fig. 5). The field water capacity and bulk density of soil in the
307 Beiluhe region is about 28% and 1 g cm^{-3} (Peng et al., 2015b), indicating that the 5 mm treatment may
308 induce the upper soil to saturate in short timescale since the duration of water addition was short (< 10
309 min). However, the pulse of Hg(0) emission for 15 mm treatment was significantly higher than that of 5
310 mm treatment. The surface soils with high sand content in the Beiluhe region have a high rate of water
311 infiltration and subsequently great infiltration depth. This process potentially increases the displacement
312 of soil Hg(0) and facilitates Hg(0) emission, as mentioned above. Therefore, in the field condition, the
313 duration and flux of pulse Hg(0) emission following water addition depends not only on how much
314 water received and soil moisture condition but also soil texture and soil water dynamics.

315 The water addition also increased the Hg(0) emission in the daytime, showing more water added,
316 longer duration of Hg(0) emission, and more Hg(0) emitted (Fig. 4 and Fig. 5). After the surface soil
317 was visibly dry, Hg(0) flux over soil plots with water addition had no significant difference from that of
318 the soil plot without water addition (i.e., 0 mm treatment). This result is consistent with many other
319 controlled studies. For example, Gustin and colleagues (Johnson et al., 2003; Gustin and Stamenkovic,
320 2005) found that once the soil water content became less than saturated, Hg(0) emission flux would be
321 significantly enhanced especially during the daytime, and once sufficient drying occurred, the magnitude
322 of Hg(0) emission flux tended to gradually decrease. Investigators suggested that as the water
323 evaporates and soil dries, capillary action drives the upward movement of water and chemicals
324 (including Hg components) and recharge the Hg pool in surface soils (i.e., the “wick effect”) and



325 subsequently favors the Hg(0) production and emission via photochemical processes in the light (Gustin
326 and Stamenkovic, 2005). In our study, even for the wettest soil plot (i.e., 15 mm treatment), the surface
327 soils were visually unsaturated in the daytime because of the low water retention, high infiltration rate
328 of local soils and intensive solar radiation. Therefore, the pattern of Hg(0) emission for soil plots with
329 high water addition is comparable to those of the above-mentioned studies.

330 Secondly, the effect of snow events on Hg(0) flux were investigated. One of the most significant
331 differences between the rainfall and snowfall on the effect of Hg(0) exchange was that the snowfall did
332 not induce the remarkable pulse of Hg(0) emission. For example, at 10:10 on 11 June 2014, a heavy
333 snowfall occurred and continued to 11:20. The great thickness of snowpack reached to ~12 cm.
334 However, no remarkable pulse of Hg(0) emission was observed during the snowfall. Instead, the Hg(0)
335 dynamics at air–snow interface showed the clear diurnal pattern with high emission in daytime and
336 deposition or emission albeit rather small in nighttime. This finding is consistent with previous studies
337 on air–snow interface (Cobbett et al., 2007). It can be seen that the pattern of Hg(0) dynamics at air–
338 snow interface was similar with that at air–soil interface, indicating that Hg(0) emission from surface
339 snow was also mainly regulated by photochemical processes (Ferrari et al., 2005; Dommergue et al.,
340 2003, 2007). However, it is well-known that the snowpack is a porous matrix, and gases are subjected to
341 diffusion in the snowpack. Therefore, our measurements of Hg(0) flux at air–snow interface did not
342 exclude the effect of Hg(0) dynamics at soil–snow interface, especially the low thickness (< 12 cm) of
343 snowpack in the study.

344 We found that the snow melting led to the remarkable peak of Hg(0) emission. For example,
345 during the sunrise of 12 June 2014, a precipitation with rain and snow induced the snowpack (12 cm) to
346 melt suddenly and completely (i.e., the bare soil with no surface snow), a pulse of Hg(0) emission (~8
347 $\text{ng m}^{-2} \text{h}^{-1}$) was observed, which was the largest Hg(0) emission flux during June 2014 campaign. We
348 supposed that the great increase in Hg(0) emission by snowmelt in this study was consistent with the
349 effect of rainfall, i.e., the displacement of soil Hg(0) during the snowmelt permeation of the soil column
350 resulted in the dramatic increase in Hg(0) emission. At present, the study on the effect of snowpack
351 melting on Hg(0) emission is limited. Cobbett et al. (2007) also found the remarkable increase in Hg(0)
352 emission in Canadian Arctic during the snow melt, although the Hg(0) flux was relatively small.



353 Finally, the effect of precipitation (including rain and snow) on daily Hg(0) flux was investigated.
354 The above-mentioned results and discussion suggest that the precipitation has great potential to
355 facilitate soil Hg(0) emission over different timescales via physical, chemical and biological processes.
356 The main processes include the displacement of soil Hg(0) by water, the “wick effect” to increase the
357 photo-reducible Hg(II) pool in surface soils, and the increased soil moisture to promote the biotic and
358 abiotic reduction of Hg(II). Another well-documented process is that the atmospheric wet deposition of
359 Hg will increase the Hg pool in surface soils and the newly deposited Hg is very active to reduce to
360 Hg(0) (Hintelmann et al., 2002), although our study did not focus on this issue. During June 2014
361 campaign, no precipitation occurred in the first two days (6–7 June 2014), but the rest days were
362 rainy/snowy days (Fig. 2). We tried to use the daily Hg(0) flux of the two sunny days as the benchmark
363 to compare with those of rainy/snowy days to investigate the effect of precipitation on the Hg(0) flux
364 over the timescale of one day. Figure 6 showed that the daily Hg(0) flux for rainy/snowy days were
365 higher (ranging from 16% to 154%) than the mean of the two sunny days. The result indicates that the
366 precipitation increased soil Hg(0) emission on the timescale of one day, although the low solar radiation
367 and temperature on rainy/snowy days would potentially decrease soil Hg(0) emission, as mentioned
368 above.

369 **3.3.3 Effect of solar radiation and soil temperature on air–surface Hg(0) flux**

370 Almost all laboratory experiments and field measurements, including this study, show that the high
371 solar radiation and elevated soil temperature synergistically facilitate the soil Hg(0) emission (Edwards
372 and Howard, 2013; Park et al., 2014). The following hypotheses have been proposed to explain the role
373 of solar radiation and temperature in promoting soil Hg(0) emission, including (1) solar radiation
374 promotes the photo-reduction of Hg(II) in surface soils to form Hg(0) in short time scale; (2) solar
375 radiation and high soil temperature reduce the apparent activation energy of Hg(0) desorption and
376 increase Hg(0) emission from surface soils; and (3) the high soil temperature favors the Hg(0)
377 production in soil column by biotic and abiotic processes (Carpi and Lindberg, 1998; Gustin et al.,
378 2002).



379 Many studies used the Arrhenius equation (Eq. 2) to quantitatively investigate the relationship
380 between soil temperature and Hg(0) flux.

381
$$F = Ae^{-Ea/RT} \text{ or } \ln(F) = \ln(A) - \frac{Ea}{RT} \text{ (Eq. 2)}$$

382 where F is the Hg(0) flux ($\text{ng m}^{-2} \text{ h}^{-1}$), R is the gas constant, T is the soil temperature (K), A is the pre-
383 exponential factor and Ea is the apparent activation energy. A plot of $\ln(F)$ versus $1/T$ obtains a straight
384 line with intercept equal to the log of the A , and the slope equal to $-Ea/R$. Theoretically, the concept of
385 the apparent Ea refers to the thermally controlled reaction. Therefore, Hg(0) flux induced by light and
386 precipitation should be excluded from the correlation analysis. However, in many previous studies,
387 especially for the field measurements, the bulk Hg(0) flux in the light was generally used to explore the
388 contribution of solar radiation or temperature to Hg(0) flux and did not isolate the respective effect of
389 the two factors (Fu et al., 2008a). This will systematically overestimate or underestimate the
390 contribution of solar radiation and temperature on soil Hg(0) emission depending on the source or sink
391 of soil for Hg(0) in the dark. Only in some controlled experiments, the separated data was used to
392 explore the respective role of radiation and temperature in soil Hg(0) emission (e.g., Kocman and
393 Horvat, 2010).

394 In this study, for respectively determining the contribution of solar radiation and temperature on
395 the Hg(0) flux, besides Hg(0) flux was measured in the natural light, Hg(0) flux in the dark was also
396 measured simultaneously with a foil-covered chamber. The temperature-corrected Hg(0) flux (i.e., bulk
397 Hg(0) flux in the light – Hg(0) flux in the dark) in daytime ($\text{PAR} > 0$) was considered to be the
398 contribution of the solar radiation. As mentioned above, the effect of precipitation should be excluded
399 from the data set, therefore we only collected Hg(0) flux data on days without precipitation during
400 December 2014 and May–June 2015 campaign.

401 Figure 7 displays the temporal variation of bulk Hg(0) flux in the light, Hg(0) flux in the dark, net
402 Hg(0) flux in the light (i.e., bulk Hg(0) flux in the light – Hg(0) flux in the dark) and the environmental
403 variables. Obviously, changes in solar radiation had a greater influence on soil Hg(0) flux than did
404 changes in soil temperature. The data showed that the soil served as a Hg(0) sink during all study days



405 in December 2014 campaign in the dark with high deposition flux in low soil temperature and low
406 deposition flux in high soil temperature. During study days of May–June 2015 campaign, the soils
407 served as a very low Hg(0) source in the midday with relatively high soil temperature. This finding is
408 consistent with many studies in background soils (Ericksen et al., 2006; Gustin et al., 2006; Fu et al.,
409 2008a; Edwards and Howard, 2013). It indicates that the soil temperature plays an important role in
410 Hg(0) dynamics at the air–soil interface, i.e., low soil temperature favors to absorb Hg(0) or reduce
411 Hg(0) emission.

412 After the temperature corrected, except for the midday of study days during May–June campaign,
413 the net Hg(0) flux in the light was higher than the bulk Hg(0) flux. The positive linear correlation was
414 found between cumulative PAR and cumulative Hg(0) flux in the daytime, although cumulative PAR
415 only explained ~28% of variation in cumulative Hg(0) flux in the daytime (Fig. S2 in Supplement).

416 We used the Hg(0) emission data set in the dark to calculate the E_a using the Arrhenius equation.
417 Since the soils in the dark was the sink of atmospheric Hg(0) in most of the study period, only limited
418 data set ($n=9$) can be used (Fig. S3 in Supplement). The E_a/R for Hg(0) emission from our remote soils
419 with extremely low Hg concentrations ($\sim 12 \text{ ug kg}^{-1}$) was 30.40. Table S1 in Supplement lists the E_a/R
420 for different soils with large variation of soil Hg concentrations, including this study, and shows that the
421 E_a/R of Hg(0) emission from soils with low Hg concentrations was higher than those of soils with high
422 Hg concentrations and significantly lower than that of theoretical value (7.31) of elemental Hg. It
423 indicates that surface soils with high Hg concentrations has great potential to emit Hg(0). This trend is
424 consistent with the laboratory study of Bahlmann et al. (2006), although the availability of Hg in soils
425 also significantly regulate Hg(0) emission (Bahlmann et al., 2006; Kocman and Horvat, 2010).

426 The QTP is characterized by high solar radiation with intense UV radiation. We further performed
427 the controlled experiment to quantify the role of different wavebands of solar radiation (UVB, UVA and
428 visible light) in Hg(0) flux. Figure 8 shows that UV radiation was the dominant waveband of solar
429 radiation for Hg(0) emission in the daytime, contributing >80% of Hg(0) emission in the light, and the
430 contribution of UV-B radiation accounted for >50% in all study days. This finding is consistent with
431 previous laboratory studies (Moore and Carpi, 2005; Bahlmann et al., 2006; Xin et al., 2007).



432 **4 Conclusions and implication**

433 In this study, we measured the Hg(0) flux between the air and surface permafrost soil in the QTP.
434 We also performed the controlled field experiments to explore the effect of precipitation and different
435 wavebands of solar radiation on the Hg(0) exchange air–soil interface. The result showed that the
436 environmental conditions, including solar radiation, soil temperature and precipitation, greatly
437 influenced the Hg(0) exchange between air and surface.

438 This study and other field measurements and laboratory experiments have clarified that the fate
439 and transport of soil Hg is very sensitive to the environmental variables (Krabbenhoft and Sunderland,
440 2013). Therefore, our results have several important implications to the Hg biogeochemical cycle in the
441 soils of QTP under the rapid climate warming and environmental change. Firstly, the increased surface
442 temperature in the QTP will potentially promote the remobilization of soil Hg. Field measurements and
443 modeling study have revealed that the surface temperature in the QTP is increasing, and the warming
444 trend exceeds those for the Northern Hemisphere and the same latitudinal zone (Kang et al., 2010).
445 Secondly, the increased UV radiation in the QTP may improve Hg(0) emission from surface soils. UV
446 radiation reaching the surface of the QTP is estimated to increase because of the decrease of
447 stratospheric O₃ (Zhou et al., 2013). Our result and many above-mentioned studies show that UV
448 radiation plays the primary role in promoting the surface Hg(0) emission in the daytime. Thirdly, the
449 temp-spatial pattern of precipitation in the QTP is also altering (Kang et al., 2010), which potentially
450 alters the flux and temp-spatial pattern of air–soil Hg(0) exchange in this region because of the
451 importance of precipitation on the Hg(0) exchange. However, this study was just the beginning to
452 explore the effect of climate change on the terrestrial Hg cycle in the QTP. The large uncertainties
453 highlight that more researches are needed in the future.

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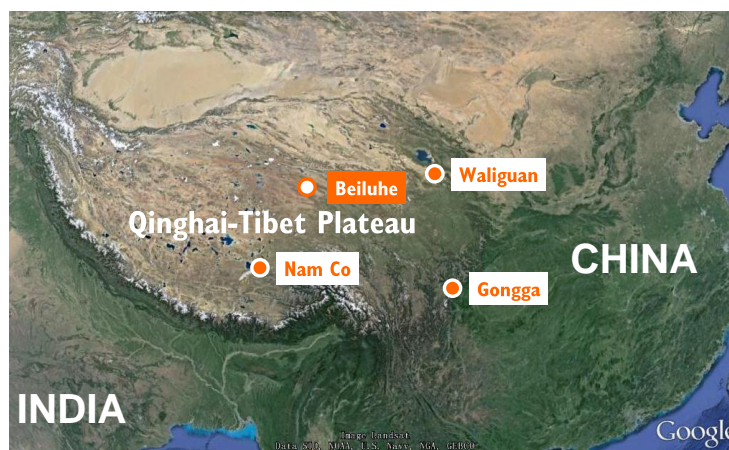


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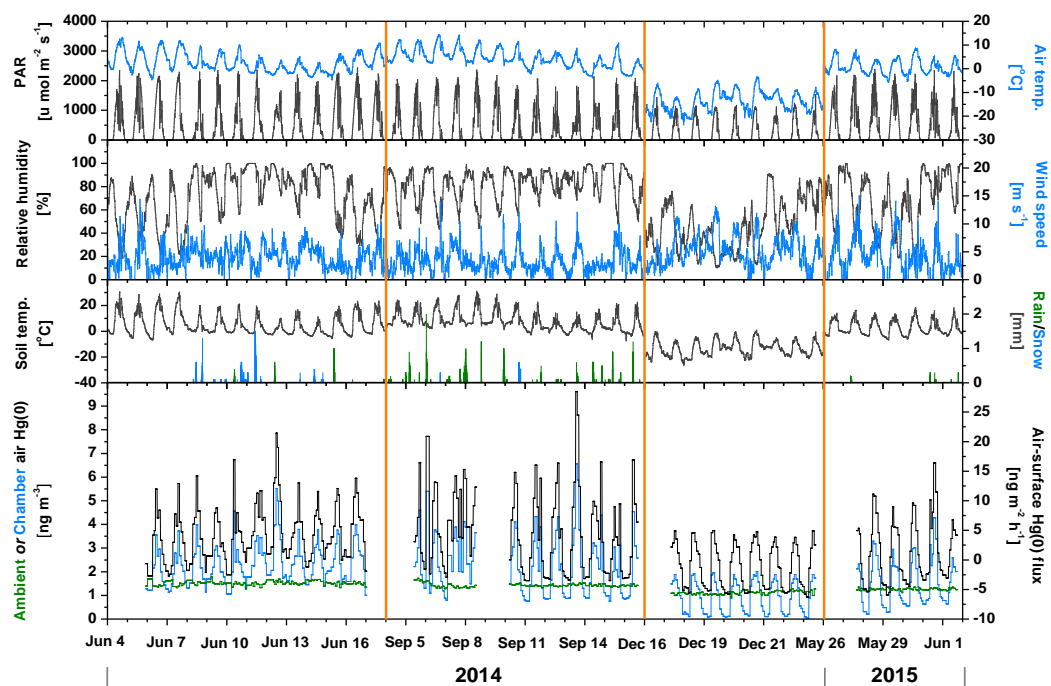
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728 Figure 1. Locations of the Beiluhe station (4760 m a.s.l., this study), Mt. Waliguan (3816 m a.s.l., Fu et al., 2012), Mt.
729 Gongga (1640 m a.s.l., Fu et al., 2008b) and Nam Co (4730 m a.s.l., Yin et al., 2015) where atmospheric Hg were
730 determined.



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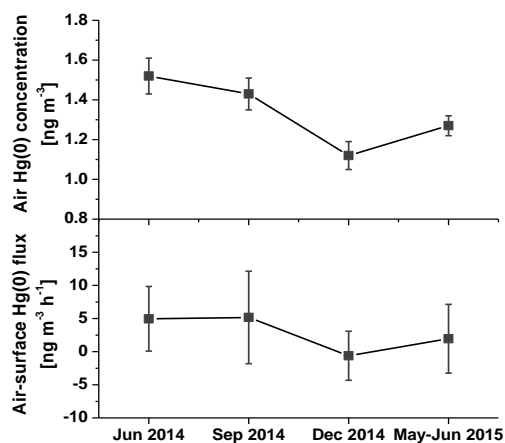
735 Figure 2. Temporal variation of environmental variables, air Hg(0) concentrations inside and outside of chamber, and air–
736 surface Hg(0) flux at the Beiluhe station in the central QTP during four campaigns in 2014–2015.

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742 Figure 3. Seasonal variation of Hg(0) concentration in ambient air and air–surface Hg(0) flux during four campaigns at the

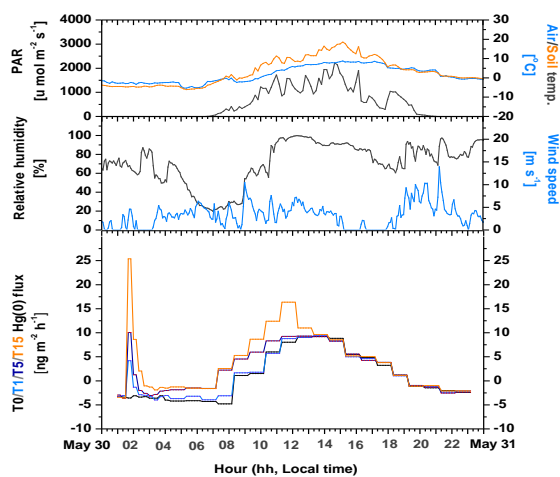
743 Beiluhe station in the central QTP in 2014–2015.

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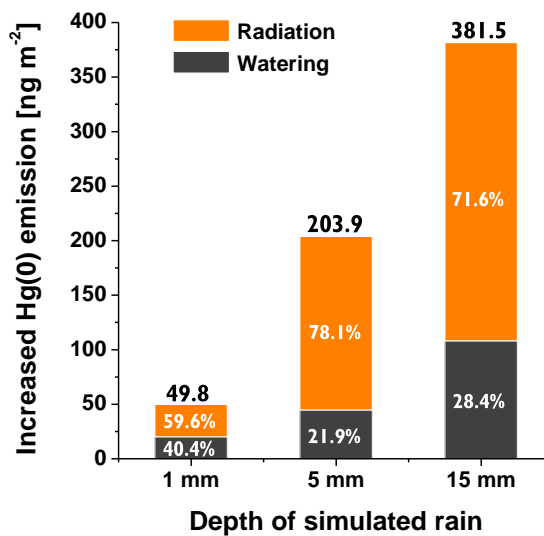
749 Figure 4. Temporal variation of Hg(0) flux over four soil plots with different treatment of water addition (T0: 0 mm
750 treatment, T1: 1 mm treatment, T5: 5 mm treatment and T15: 15 mm treatment) and the environmental variables.

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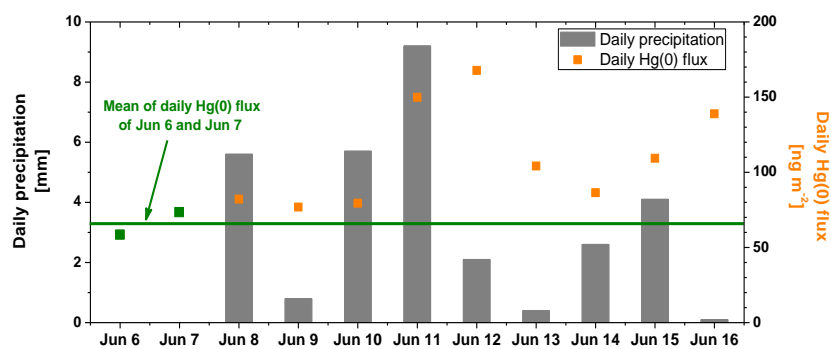
756 Figure 5. Increased Hg(0) emission for three different treatments (1 mm, 5 mm and 15 mm addition of water) compared with
757 the 0 mm treatment during the controlled experiment on 30 May 2015.

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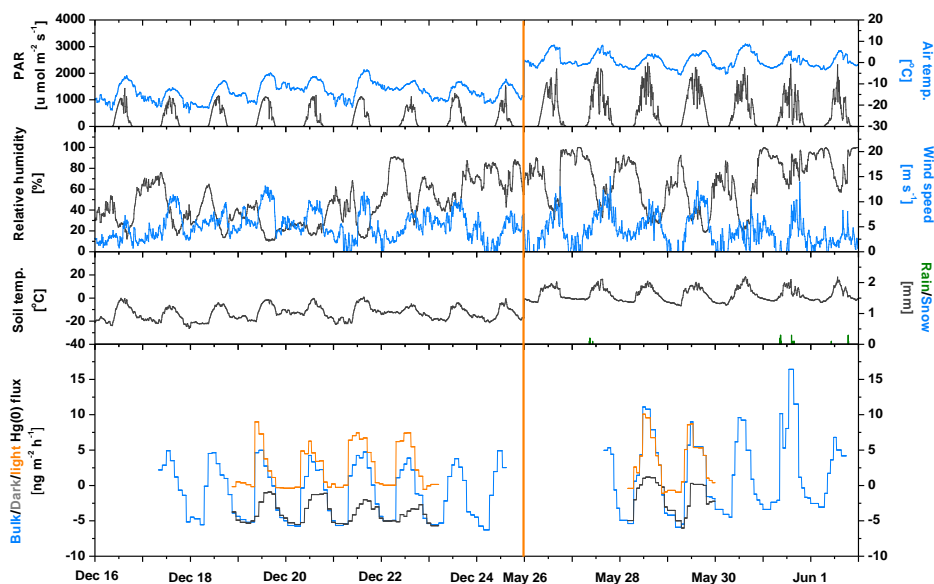
763 Figure 6. Daily Hg(0) flux and daily precipitation in June 2014 campaign.

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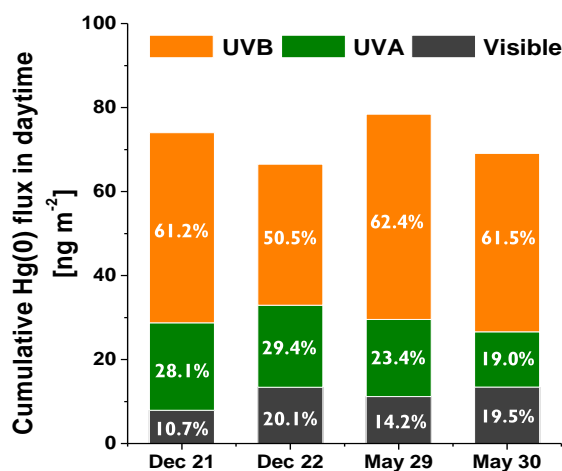
769 Figure 7. Temporal variation of bulk Hg(0) flux in the light, Hg(0) flux in the dark
770 Hg(0) flux in the light–Hg(0) flux in the dark) in six study days without precipitation during December 2014 campaign and
771 May–June 2015 campaign.

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777 Figure 8. Cumulative Hg(0) emission flux in daytime triggered by UVB, UVA and visible light in four study days during
778 December 2014 campaign and May–June 2015 campaign. Four chambers with different exposure treatments were used to
779 measure Hg(0) flux simultaneously in the daytime. Chamber-A was used to measure the Hg(0) flux in the natural light.
780 Chamber-B and Chamber-C were covered with UVB filter and UV filters to remove the corresponding wavebands from the
781 natural light, respectively. Chamber-D covered with foil was used to measure Hg(0) flux in the dark. The experiments were
782 performed in four days without precipitation (21–22 December 2014 and 29–30 May 2015) to exclude the effect of
783 precipitation. Hg(0) flux triggered by UVB, UVA and visible light was equal to difference of flux between Chamber-A and
784 Chamber-B, between Chamber-B and Chamber-C, and between Chamber-C and Chamber-D, respectively. The transmittance
785 of UVB filter and UV filter was shown in Fig. S1 in Supplement.