

1 **Speciated Atmospheric Mercury during Haze and Non-haze Days in**  
2 **an Inland City in China**

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24

25 **Abstract.** Long-term continuous measurements of speciated atmospheric mercury  
26 were conducted from July 2013 to June 2014 at Hefei, a mid-latitude inland city in  
27 east central China that experiences frequent haze pollution. The mean concentrations  
28 ( $\pm$  standard deviation) of gaseous elemental mercury (GEM), gaseous oxidized  
29 mercury (GOM) and particle-bound mercury (PBM) were  $3.95 \pm 1.93 \text{ ng m}^{-3}$ ,  $2.49 \pm$   
30  $2.41 \text{ pg m}^{-3}$  and  $23.3 \pm 90.8 \text{ pg m}^{-3}$ , respectively, during non-haze days, and  $4.74$   
31  $\pm 1.62 \text{ ng m}^{-3}$ ,  $4.32 \pm 8.36 \text{ pg m}^{-3}$  and  $60.2 \pm 131.4 \text{ pg m}^{-3}$ , respectively, during haze  
32 days. Potential source contribution function (PSCF) analysis suggested that the  
33 atmospheric mercury pollution during haze days was caused primarily by local  
34 mercury emissions, instead of via long-range mercury transport. The disadvantageous  
35 diffusion during haze days also favoured the accumulation of atmospheric mercury.  
36 Compared to GEM and GOM, PBM was found to be more sensitive to haze pollution.  
37 The mean PBM concentration during haze days was 2.5 times that during non-haze  
38 days due to elevated concentrations of particulate matter. A remarkable seasonal trend  
39 in PBM was also observed with concentration decreasing in the following order in  
40 response to the frequency of haze days: autumn, winter, spring and summer. **For both**  
41 **non-haze and haze days, GOM concentrations remained at relatively constant during**  
42 **night, but increased rapidly prior to sunrise. This GOM diurnal variation could be due**  
43 **to diurnal variation in air exchange between the boundary layer and free troposphere,**  
44 **but the contribution from photochemical oxidation could not be ruled out. This is**  
45 **supported by simple box-model calculations, which showed that oxidation of GEM to**  
46 **GOM does occur and that the transport of free troposphere GOM alone is not large**  
47 **enough to account for the observed increase in GOM. Our results postulate that  $\text{NO}_2$**

48 aggregation with the HgOH intermediate may be a potential mechanism for the  
49 enhanced production of GOM during daytime.

50

## 51 1. Introduction

52 Mercury (Hg) is an environmental pollutant that has received much global  
53 attention because of its toxicity and bioaccumulation via the aquatic food chain. The  
54 most important transport pathway of mercury is via the atmosphere (Schroeder and  
55 Munthe, 1998;Lindqvist and Rodhe, 1985). Operationally, atmospheric mercury is  
56 commonly differentiated into three forms: gaseous elemental mercury (GEM),  
57 gaseous oxidized mercury (GOM) and particle-bound mercury (PBM). The sum of  
58 these three atmospheric speciated mercury is defined as total atmospheric mercury  
59 (TAM = GEM + GOM + PBM), and the sum of GEM and GOM is known as total  
60 gaseous mercury (TGM = GEM + GOM) (Gustin and Jaffe, 2010;Gustin et al., 2015).  
61 GEM is regarded as the dominant form of atmospheric mercury, accounting for over  
62 95% of the total. GEM is stable in the troposphere with a long residence time (0.5–2  
63 yr) and can be transported at the regional to global scale (Schroeder and Munthe,  
64 1998;Lindberg et al., 2007). GEM can be oxidized to GOM through photochemical  
65 processes, and further transformed to PBM on aerosol surfaces. GOM and PBM can  
66 be readily removed from the air by wet and dry deposition as a result of their high  
67 surface affinity and water solubility (Lindqvist and Rodhe, 1985). Thus, the chemical  
68 transformation between GEM, GOM and PBM will directly influence the atmospheric  
69 lifetime of mercury.

70 As a result of the rapid industrial development and economic growth of recent  
71 decades, China has become one of the major contributors to anthropogenic mercury

72 emissions to the environment (Wu et al., 2006;Pacyna et al., 2006;Pacyna et al.,  
73 2010;Zhang et al., 2015b). Atmospheric mercury emissions from anthropogenic  
74 sources in China have been estimated to be in the range of 500-700 tons/yr,  
75 accounting for 25-30% of the total global anthropogenic mercury emissions (Streets et  
76 al., 2005;Wu et al., 2006). Research into atmospheric mercury in China is therefore  
77 critical to the understanding of mercury cycling at both regional and global scales.  
78 Long-term observation of atmospheric mercury has been conducted in different  
79 regions in China, including both urban and remote areas. TGM concentrations  
80 observed in urban and industrial regions of China were in the range of 2.7–35 ng m<sup>-3</sup>,  
81 higher than the values reported for North America and Europe, and for the adjacent  
82 Asian countries such as Korea and Japan (Weigelt et al., 2013;Fang et al.,  
83 2009;Marumoto et al., 2015). TGM and PBM concentrations in remote areas of China  
84 were also found to be higher than those observed in North America and Europe (Fu et  
85 al., 2008a;Fu et al., 2008b;Fu et al., 2012;Liu et al., 2010).

86 In recent years, haze pollution has become a major concern in China due to its  
87 impacts on visibility, air quality, and climate. It is well known that haze formation is  
88 mainly dependent on the atmospheric relative humidity (RH) and the concentration of  
89 airborne particles (Chen et al., 2003;Sun et al., 2013). Most studies on haze have  
90 focused on the measurements of airborne particulate matter; few examined the  
91 influence of haze on the chemistry of atmospheric mercury, especially PBM. In this  
92 study, we conducted one year synchronous observations of speciated atmospheric  
93 mercury in Hefei, an inland city of China, which experiences frequent haze events.  
94 The comparison of atmospheric mercury under haze days and non-haze days during

95 the study period allows us to examine the formation and deposition mechanisms of  
96 mercury, as well as their temporal variations.

97

## 98 **2. Methods**

### 99 **2.1 Study site**

100 Continuous measurements of speciated atmospheric mercury were undertaken in  
101 Hefei (31°52' N, 117°17' E) from July 2013 to June 2014. Hefei, the capital of Anhui  
102 Province, is located in east central China, between the Changjiang (Yangtze River) and  
103 the Huaihe (Huai River). Hefei has a humid subtropical climate with four distinct  
104 seasons: June-August is considered summer, September-November autumn,  
105 December-February winter and March-May spring. The prevailing wind is  
106 southeasterly in summer and northwesterly in winter. Like many Chinese cities, Hefei  
107 has experienced rapid growth in the past 20 years. The total permanent population is  
108 about 7.7 million. **The city has also been witnessing an increasing frequency in haze**  
109 **pollution, especially in winter months.**

110 The monitoring site was located on the Science Island, a small peninsula on the  
111 Dongpu Reservoir in the northwestern outskirts of Hefei (Fig. 1). The sampling and  
112 analytical instruments were installed 1.5 m above the rooftop (~ 20 m above the  
113 ground) of the main building of Anhui Institute of Optics and Fine Mechanics. Further  
114 information about the monitoring site can be found in a previous study (Hu et al.,  
115 2014). We chose this area as the monitoring site because it is one of the cleanest areas  
116 in Hefei, not adjacent to any direct pollution sources such as power plants, iron and  
117 steel works.

118

## 119 **2.2 Measurements of speciated atmospheric mercury**

120 From July 2013 to June 2014, simultaneous measurements of speciated  
121 atmospheric mercury concentrations were performed by an automated Tekran™  
122 mercury speciation system. The system consisted of a Model 2537B mercury analyzer  
123 combined with a Model 1130 GOM unit and a Model 1135 PBM unit. The system  
124 was configured to measure GEM every 5 min., and GOM and PBM every 2 h.

125 The details about the Tekran-based mercury speciation system can be found in  
126 Landis et al. (2002). In general, the automated measurement process can be  
127 summarized as sample collection, thermal desorption and determination. During the  
128 collection period, ambient air was drawn to the system at a flow rate of 10 L/min.  
129 GOM and PBM in the air were captured by a KCl-coated quartz annular denuder in  
130 the 1130 unit and a quartz filter in the 1135 unit, respectively, whereas GEM would  
131 pass through the denuder and filter and be quantified on the Tekran 2537B by  
132 cold-vapor atomic fluorescence spectroscopy (CVAFS). After an hour of sampling, the  
133 1135 quartz filter and the 1130 denuder would be switched to the thermal  
134 decomposition mode at 800 °C and 500 °C, respectively, with the resulting Hg<sup>0</sup>  
135 quantified by the 2537B unit in the next hour, while the 1135 and 1130 components  
136 were flushed with zero-mercury gas for the next sampling.

137 The instrument maintenance followed typical protocols used in similar studies  
138 (Landis et al., 2002;Hu et al., 2014). The quartz annular denuder was recoated every  
139 two weeks, the quartz filter was replaced once a month, and the Teflon filter (pore size  
140 0.2 μm) in the sample inlet was changed every two weeks. Automated recalibration of  
141 the Tekran 2537B was performed every 25 h using an internal mercury permeation  
142 source. No calibration standards were available for GOM and PBM, but the 1σ

143 precision for GOM and PBM was about 15 % (Landis et al., 2002). The detection  
144 limit in ambient air is about  $0.5 \text{ ng m}^{-3}$  for GEM (or TGM) at a resolution of 5 min,  
145 and  $1 \text{ pg m}^{-3}$  for GOM and PBM at a resolution of 2 h (Gustin et al., 2015). Although  
146 the Tekran-based mercury speciation technique has been widely used around the  
147 world, recent studies have shown that the technique does not efficiently collect all  
148 gaseous oxidized mercury and thus may substantially underestimate the concentration  
149 of reactive mercury (Huang et al., 2013;Gustin et al., 2013). Therefore, the GOM  
150 values reported in this study should be considered as the lower limit of gaseous  
151 oxidized mercury in the air (Wang et al., 2014).

152

### 153 **2.3 Ancillary Data**

154 Standard meteorological measurements including air temperature, air pressure,  
155 RH, wind direction and speed were observed with a 5-min resolution. CO was  
156 measured by an automated infrared carbon monoxide analyzer (Model EC9830T,  
157 Ecotech Inc., Australia), with a detection limit of 40 ppbv. O<sub>3</sub> was measured every 5  
158 min by an ozone analyzer (Model EC9810B, Ecotech Inc., Australia); its detection  
159 limit and accuracy are 0.5 ppbv and 0.001 ppm, respectively. NO<sub>2</sub> was measured by a  
160 Multi axis differential optical absorption spectroscopy (MAX-DOAS) instrument. The  
161 collected spectra were analyzed using the QDOAS spectral fitting software suite  
162 developed at BIRA-IASB (<http://uv-vis.aeronomie.be/software/QDOAS/>). PM<sub>2.5</sub>  
163 (particulate matter less than 2.5  $\mu\text{m}$  in diameter) data are collected from China air  
164 quality online analysis platform (<http://www.aqistudy.cn/historydata/index.php>).

165

### 166 **2.4 Potential Sources Contribution Function (PSCF) analysis**

167 To identify the possible influence of long-range transport on the distribution of  
168 atmospheric mercury in Hefei, we calculated backward trajectories of air masses  
169 using the HYSPLIT (Hybrid Single-particle Lagrangian Integrated Trajectory) model  
170 with the Global Data Assimilation System (GDAS 1<sup>o</sup>) developed by the National  
171 Oceanic and Atmospheric Administration (NOAA) (<http://www.ready.noaa.gov>)  
172 (Draxler and Hess, 1998). Considering the atmospheric pollutants are mainly  
173 concentrated in the low altitude during heavy pollution days, the trajectory arrival  
174 heights were set at 500 m to represent the boundary layer where atmospheric  
175 pollutants were well mixed. In this study, 3-day back-trajectories were generated  
176 hourly by TrajStat, a software including HYSPLIT for trajectory calculation (Wang et  
177 al., 2009).

178 The contributions of other pollution source regions to the atmospheric mercury at  
179 Hefei was identified by the Potential Sources Contribution Function (PSCF) analysis  
180 with the TrajStat software. PSCF analysis has been shown to be useful in spatially  
181 identifying pollution sources for pollutants with a long lifetime such as elemental  
182 mercury and CO (Xu and Akhtar, 2010). The PSCF values for the grid cells in the  
183 study domain were calculated by counting the trajectory segment endpoints that  
184 terminate within each cell. The number of endpoints that fall in the  $ij_{th}$  cell is  
185 designated as  $N_{ij}$ . The number of endpoints for the same cell corresponding to the  
186 atmospheric mercury concentration higher than an arbitrarily set criterion is defined to  
187 be  $M_{ij}$ . In this study, mean GEM concentration of  $4 \text{ ng m}^{-3}$  during the whole study  
188 period was used as the mercury pollution criterion. The PSCF value for the  $ij_{th}$  cell is  
189 then defined as:

190 
$$PSCF_{ij} = \frac{M_{ij}}{N_{ij}} W_{ij} \quad (2)$$

191  $W_{ij}$  is an arbitrary weight function to reduce the effect of small values of  $N_{ij}$ . The  
 192 PSCF values were multiplied by  $W_{ij}$  to better reflect the uncertainty in the values for  
 193 these cells (Polissar et al., 2001). The weight function reduces the PSCF values when  
 194 the total number of endpoints in a particular cell is less than 3 times the average value  
 195 of the end points per cell:

$$196 \quad W_{ij} = \begin{cases} 1.0 & N_{ij} \geq 3N_{ave} \\ 0.70 & 3N_{ave} > N_{ij} \geq 1.5N_{ave} \\ 0.40 & 1.5N_{ave} > N_{ij} \geq N_{ave} \\ 0.20 & N_{ave} > N_{ij} \end{cases} \quad (3)$$

197

### 198 3. Results

199 We intended to continuously monitor speciated atmospheric mercury  
 200 concentration over the course of a year; however, interruptions were inevitable due to  
 201 instrument maintenance, which resulted in loss of data for the following four periods:  
 202 (1) 25 September to 9 October 2013; (2) 5-14 November 2013; (3) 9-25 February  
 203 2014; and (4) 1-14 April 2014. The rest of the data were grouped into haze days and  
 204 non-haze days according to the China Meteorological Administration's haze standard  
 205 (QX/T 113-2010). Haze days refer to the days when the atmospheric visibility < 10  
 206 km and RH < 80% (Duan et al., 2016), and non-haze days refer to clear days with the  
 207 atmospheric visibility > 10 km. The visibility and RH information were collected  
 208 from the weather history data at the Luogang Airport of Hefei  
 209 (<http://www.wunderground.com/>). Through the study period of almost a year, a total  
 210 of 56 days were identified to be haze days, and 253 days to be non-haze days. All the  
 211 times reported herein are local time (UTC + 8 h).

212

#### 213 3.1 Overall characteristics of speciated atmospheric mercury

214 The time series of GEM, GOM and PBM concentrations at the study site  
215 throughout the study period are shown in Fig. 2, and their frequency distributions are  
216 shown in Fig. S1 (in the supporting information). The mean ( $\pm$  standard deviation)  
217 GEM, GOM and PBM concentrations during the whole study period were  $4.07 \pm 1.91$   
218  $\text{ng m}^{-3}$ ,  $3.67 \pm 5.11 \text{ pg m}^{-3}$ , and  $30.0 \pm 100.3 \text{ pg m}^{-3}$ , respectively (Table 1). The GEM  
219 concentrations in different seasons did not differ much. The highest GEM  
220 concentration occurred in autumn ( $4.51 \pm 2.10 \text{ ng m}^{-3}$ ), while the lowest in spring  
221 ( $3.89 \pm 1.79 \text{ ng m}^{-3}$ ). GOM concentrations varied greatly during the study period with  
222 much higher concentrations in autumn and the lowest in winter. A similar seasonal  
223 variation in the GOM concentration was observed at a remote site in Mt. Gongga of  
224 southwest China (Fu et al., 2008b). The seasonal trend in PBM was also observed in  
225 Hefei with its concentration decreasing in the following order: autumn > winter >  
226 spring > summer. The mean PBM concentration during the cold season was about 20  
227 times that in summer, similar to the findings from many previous studies in China  
228 (Zhang et al., 2013;Fu et al., 2011;Fu et al., 2008b;Fang et al., 2001).

229 Comparisons of speciated atmospheric mercury concentrations with other urban  
230 and rural areas in China and a few other countries are shown in Table 2. The mean  
231 GEM concentration at Hefei is slightly higher than that in many remote areas in China  
232 (Fu et al., 2008a;Fu et al., 2008b;Fu et al., 2012;Wan et al., 2009a;Wan et al.,  
233 2009b;Zhang et al., 2015a), but is much lower than those in urban areas of industrial  
234 cities such as Guiyang and Changchun where large point sources of mercury exist  
235 (e.g., non-ferrous metal smelting, coal-fired power plants, and residential coal burning)  
236 (Feng et al., 2004;Fu et al., 2011;Fang et al., 2004). Although Hefei is geographically  
237 close to Shanghai, a mega urban centre in China, it is interesting to note that the TGM

238 concentration of Shanghai is much lower than that of Hefei. This may be due to the  
239 fact that Shanghai is a coastal city that is influenced more by cleaner marine air  
240 masses (Friedli et al., 2011). Table 2 also shows that the average concentration of  
241 GEM in Hefei is typically more than fold or two-fold that in the urban and rural areas  
242 in Europe and North America (Liu et al., 2010;Li et al., 2008;Brooks et al., 2010).

243

### 244 **3.2 Speciated atmospheric mercury during non-haze days**

245 The frequency distribution of GEM, GOM and PBM for the non-haze period are  
246 shown in Fig. S1 (in blue). The mean concentration of GEM was  $3.95 \pm 1.93 \text{ ng m}^{-3}$ .  
247 Its distribution was characterized by large fluctuations ranging from 0.2 to 23.8 ng  
248  $\text{m}^{-3}$ , although more than half of the GEM values were in the narrow range 2-4 ng  $\text{m}^{-3}$ .  
249 The mean concentration of GOM was  $2.49 \pm 2.41 \text{ pg m}^{-3}$  with a range of 0.5-33.5 pg  
250  $\text{m}^{-3}$ , although most of the values were in the range of 1-4 pg  $\text{m}^{-3}$ . High concentrations  
251 of GOM (exceeding 10 pg  $\text{m}^{-3}$ ) only accounted for 1.4% of the total data. The mean  
252 GOM concentration at the Hefei site is much smaller than that reported from other  
253 study sites in China (Table 2), but is comparable to the values observed from many  
254 European and North American sites (Peterson et al., 2012;Cheng et al., 2014;Ren et  
255 al., 2016). The mean PBM concentraion at the Hefei site during the non-haze days  
256 was  $23.3 \pm 90.8 \text{ pg m}^{-3}$  with an exceptionally large range of 0.5-1827 pg  $\text{m}^{-3}$ . The  
257 frequency distribution of PBM showed that high PBM concentrations (i.e., > 50 pg  
258  $\text{m}^{-3}$ ) accounted for 6.4% of the total data. The PBM concentration under the non-haze  
259 condition in Hefei is generally at a similar level to the remote areas, such as Mt.  
260 Gongga, Mt. Waliguan and Shangri-La in western China.

261 Diurnal variations of GEM, PBM and GOM concentrations for non-haze days are  
262 shown in Fig. 3. Both GEM and PBM concentrations exhibited similar variations with  
263 elevated concentrations during night. The GOM concentration during the daytime was  
264 slightly higher than that in nighttime, typically peaking between 10:00 and 12:00.

265

### 266 **3.3 Speciated atmospheric mercury during haze days**

267 Haze pollution mainly occurred in December and January at our monitoring site.  
268 The four major haze pollution periods were identified in grey in Fig. 2. The mean  
269 concentrations of GEM, GOM and PBM during these haze days were  $4.74 \pm 1.62 \text{ ng}$   
270  $\text{m}^{-3}$ ,  $4.32 \pm 8.36 \text{ pg m}^{-3}$  and  $60.2 \pm 131.4 \text{ pg m}^{-3}$ , respectively (Table 1). The frequency  
271 distributions of GEM, GOM and PBM for the haze days are shown in Fig. S1 (in  
272 gray). Comparison of GEM, GOM and PBM concentrations during haze and non-haze  
273 days is shown in Fig. S2. GEM, GOM and PBM concentrations show significant  
274 differences between haze and non haze days ( $p < 0.001$ , t-test). On average, the  
275 concentration of GEM in haze days was 1.2 times that in non-haze days. Similarly, the  
276 concentration of GOM in haze days was about 1-1.7 times those in non-haze days.  
277 The largest impact of haze pollution is however on PBM, with the mean PBM  
278 concentration in haze days about 2.5 times that of non-haze days. High concentrations  
279 of GOM (exceeding  $10 \text{ pg m}^{-3}$ ) and PBM concentrations (exceeding  $50 \text{ pg m}^{-3}$ ) were  
280 also more frequently observed than in non-haze days, accounting for 5.9% and 25%,  
281 respectively, of the total haze days.

282 Diurnal variations of GEM, PBM and GOM concentrations for haze days are  
283 shown in Fig. 3. GEM concentrations were higher during night, decreased during  
284 daytime. The opposite pattern was observed for GOM, which showed higher

285 concentrations during daytime than during night; the duration of the GOM peak also  
286 lasted longer for haze days. On the contrary, the PBM typically peaked just before  
287 sunrise, with the lowest values occurred in the afternoon (14:00-16:00).

288

## 289 **4. Discussion**

### 290 **4.1 Influence of atmospheric mercury emission source**

291 In order to understand the mercury sources attribution, the PSCF model analysis  
292 was conducted by using the TrajStat software. The seasonal sources could be inferred  
293 from the PSCF analysis with the year-round data. Fig. 4a showed the overall spatial  
294 contribution of mercury emission sources in China. As Hefei is located in east-central  
295 China, its atmospheric mercury concentration could be affected by both north and  
296 south emission sources, including those from the North China Plain (especially  
297 Shandong Province) and the neighboring provinces of Henan, Jiangsu, Jiangxi and  
298 Hubei. The total mercury emissions from Henan and Shandong provinces were  
299 estimated to be over 50 and 45 tons in 2010, respectively, making them two largest Hg  
300 emitters in China (Zhang et al., 2015b). Long-range transport could also impact the  
301 seasonal variations of atmospheric mercury in Hefei. As shown in Figure 4, in spring,  
302 the major contributors of atmospheric mercury to Hefei were from the southwestern  
303 region including the local area and the Jiangxi and Hunan provinces. In summer, the  
304 main contributors were from north of Anhui, as well as Henan and Jiangxi provinces,  
305 and even from the Pearl River Delta region in the far south. **Since the number of haze**  
306 **days accounts only for 5.6% of the total days in spring and summer, we did not**  
307 **provide haze and non-haze PSCF results for spring and summer seasons. As autumn**  
308 **and winter are the prevalent seasons for haze pollution, one PSCF result for haze days**

309 and another for non-haze days are shown for autumn and winter, respectively. The  
310 statistically significant difference ( $p < 0.001$ ) in the GEM concentration between  
311 non-haze days and haze days suggests that haze pollution could directly affect the  
312 concentration of elemental mercury. As shown in Figs. 4d and 4f, higher GEM  
313 concentration was mainly influenced by local emission sources during haze days. For  
314 non-haze days, the most important mercury sources to the monitoring site were not  
315 only the local emission sources, but also those from the neighboring region of  
316 Shandong, Henan and Jiangxi provinces (see Figs. 4e and 4g). In summary, the  
317 increase of GEM concentration during haze days was mainly caused by local  
318 emission.

319 GEM and CO often share anthropogenic emission sources, such as industrial coal  
320 combustion, domestic coal combustion, iron and steel production and cement  
321 production (Wu et al., 2006; Wang et al., 2005). However, they also have their own  
322 sources. For instance, power plants and nonferrous metal smelters emit mercury but  
323 hardly any CO, while most of CO originates from vehicles which are not a major  
324 emitter for mercury. The correlation coefficients and slopes between GEM  
325 concentration and CO concentration during pollution events are shown in Table 3.  
326 These mercury pollution episodes were defined as a period with hourly average GEM  
327 concentration higher than seasonal average GEM concentration and the duration of  
328 elevated hourly GEM concentration lasted for over 10 hours. The Hg/CO slope and  
329 correlation between GEM and CO concentrations has been used to identify long-range  
330 transport episodes and local episodes in previous research (Jaffe et al.,  
331 2005; Weiss-Penzias et al., 2006; Kim et al., 2009). These episodes could be classified  
332 into long-range transport episodes or local episode by using the correlation

333 coefficients ( $R^2$ ) of linear regression: significant positive correlation for long-range  
334 transport episodes and poor correlation for local episodes (Kim et al., 2009). Six  
335 episodes (events: 1-6) were found to have poor correlations between GEM and CO  
336 concentrations ( $R^2$ : 0.01-0.29) for local episodes, while four episodes (events: 7-10)  
337 were found to have positive correlations between GEM and CO concentrations ( $R^2$ :  
338 0.51-0.79) for long-range transport episodes. These local episodes tend to occur in  
339 autumn and winter. The slope of the trend line represents the Hg/CO ratio. Emissions  
340 from power plants typically have a higher Hg/CO ratio (Wu et al., 2006), whereas  
341 residential coal and biomass burning combustion have a lower Hg/CO ratio  
342 ( $0.0013\text{-}0.0046\text{ ng m}^{-3}\text{ ppbv}^{-1}$ ) due to incomplete combustion (Weiss-Penzias et al.,  
343 2007). The Hg/CO ratio for vehicles is nearly zero (Zhang et al., 2013). The Hg/CO  
344 ratios from our study for pollution episodes are in the range of  $0.0001\text{-}0.0050\text{ ng m}^{-3}$   
345  $\text{ppbv}^{-1}$ . In summary, the low GEM/CO ratio in Hefei might related to the local  
346 incomplete combustion, like residential coal and biomass burning combustion.

347

#### 348 **4.2 Impacts of meteorological factors for atmospheric mercury during haze days**

349 Meteorological conditions, especially wind direction and speed, could also  
350 impact the atmospheric mercury during haze days. The wind rose for the monitoring  
351 site during the study period is shown in Fig. 5. Easterly and southeasterly winds  
352 represented the prevailing wind directions at the study site. A wind rose diagram of  
353 GEM concentrations above the 90<sup>th</sup> percentile value is shown in Fig. 5B. We found  
354 that 67% of the high GEM concentrations occurred at low wind speed (below  $1.5\text{ m}$   
355  $\text{s}^{-1}$ ); however, wind speed below  $1.5\text{ m s}^{-1}$  accounted for only 1.7% of total study.  
356 High GOM and PBM concentrations appear not to be related to high wind speed

357 (wind speed: 3-5 m s<sup>-1</sup>); only 1.4% and 2.6% of the high GOM and PBM  
358 concentrations were observed under high wind-speed conditions, respectively (Figs.  
359 5C and 5D). In general, most of the high atmospheric mercury levels occurred in the  
360 low wind speed conditions. This slow wind speed condition is not conducive to the  
361 spread of mercury and thus favours the accumulation of atmospheric mercury,  
362 especially during haze days. This result is further support that atmospheric mercury  
363 during haze days is mainly affected by local emissions.

364 Both GEM and PBM concentrations exhibited great variations with elevated  
365 concentration during night or early morning, regardless of the presence of haze. Such a  
366 diurnal variation of GEM and PBM could be related to changes in the height of urban  
367 boundary layer, which is typically low in the morning and night, and high during the  
368 daytime (Yuan et al., 2005;Mao et al., 2006). The maximum PBM concentration  
369 (observed at 6:00) was more than 4 times higher than the minimum value (observed at  
370 16:00) both under non-haze and haze days, and about 76% PBM were declined during  
371 this period (6:00-16:00). However, the reductions of PBM as a result of deposition  
372 during haze days was 62.7 pg m<sup>-3</sup>, which was about 2.4 times that in non-haze days,  
373 suggesting that haze pollution could increase the removal of PBM and thus reduce its  
374 atmospheric lifetime. Although PBM is not the major form of mercury emitted to the  
375 atmosphere, it is crucial in atmospheric mercury transport and removal processes due to  
376 its short atmospheric lifetime. As shown in Fig. 6, highest PBM and PM<sub>2.5</sub>  
377 concentrations were observed in January, which most likely due to shallower boundary  
378 layer in January than in other months. The co-variation in February is weaker, possibly  
379 due to the loss of PBM data because of instrument maintenance (see Section 2.3). The  
380 PBM concentration co-varied with the PM<sub>2.5</sub> concentration, especially in January when

381 all the four PBM peak events were associated with increased  $PM_{2.5}$  concentrations (Fig.  
382 6c). This result indicates that the  $PM_{2.5}$  concentration may play an important role in the  
383 formation of PBM. Thus, elevated PBM concentrations might be due to the poor  
384 diffusion conditions in cold months and high PM pollution.

385

### 386 **4.3 Enhancements in GOM and the potential oxidation mechanism**

387 Diurnal variations of GEM, GOM,  $O_3$  and CO concentrations during non-haze  
388 and haze days are shown in Fig. 7. The weak correlation ( $r=0.164$ ,  $p<0.001$ ) between  
389 GOM and CO suggests that the CO-producing, primary emission is not a major source  
390 of GOM in the air. This is clearly shown in Fig. 7 (haze days), where the peak value  
391 of GOM coincided with the lowest value of CO. GOM also shows a diurnal trend that  
392 is opposite to that of GEM and PBM; for both non-haze and haze days, GOM  
393 concentrations remained relatively constant during night, but increased rapidly prior  
394 to sunrise. Two processes can affect the GOM concentrations in the boundary layer air.  
395 The first is due to transport of GOM from the free troposphere (FT). It is well  
396 established that FT contains higher GOM concentrations than in the boundary  
397 layer(e.g., (Murphy et al., 2006;Lyman and Jaffe, 2012;Timonen et al., 2013;Brooks  
398 et al., 2014;Shah et al., 2016)). It is thus possible that higher GOM concentrations  
399 observed prior to sunrise is due to admixing from the free troposphere as the height of  
400 the boundary layer increases during the morning. In addition, in situ photochemical  
401 oxidation of GEM could also increase the concentration of GOM during daytime.  
402 Various atmospheric oxidants are capable of oxidizing GEM to GOM, including  
403 halogen radicals, ozone, hydroxyl radicals (OH), among others (Holmes et al.,  
404 2010;Wang et al., 2014).

405 To determine the relative importance of FT transport and in situ photochemical  
406 oxidation, we examined the relationship between GOM and the changes in the height  
407 of the atmospheric boundary layer and the odd oxygen ( $O_x = O_3 + NO_2$ ) concentrations.  
408 We used  $O_x$  because it is a more conserved tracer of the extent of photochemical  
409 processes in the urban atmosphere (Herndon et al., 2008; Wood et al., 2010), as  $O_3$   
410 reacts with NO emitted from automobiles to form  $NO_2$ . Example results are shown in  
411 Fig. 8 for 20th November, 2013. As can be seen from the figure, both GOM and  $O_x$   
412 reached higher concentrations from 12:00 to 16:00, along with the lowest value of  
413 GEM. The height of the atmospheric boundary layer changed very little (less than 0.1  
414 km) over the same period (see Fig. S3). This simple comparison suggests that the  
415 transport of FT GOM might be limited and that at least some of the GOM were  
416 formed from in situ oxidation of GEM. Note that in our studies we could only  
417 calculate daytime  $O_x$  concentrations, because  $NO_2$  concentrations from MAX-DOAS  
418 were only available during daytime.

419 We further investigated the mechanism of the GEM oxidation to GOM. Ozone  
420 itself is not an efficient oxidant for GEM oxidation due to the low reaction rate (Hall,  
421 1995; Holmes et al., 2010). Instead, halogen radicals (especially bromine atoms) and  
422 halogen radicals, are believed to be the primary oxidants for GEM in the global  
423 troposphere (Holmes et al., 2010). Unfortunately, we did not measure halogen radicals  
424 in this study. OH radicals are known to be present in the early morning urban  
425 boundary layer, primarily from the photolysis of HONO, which accumulates during  
426 night (Kleffmann et al., 2005). Therefore, here we consider the oxidation of GEM by  
427 OH radicals. The formation of HgOH as an intermediate product of the  $Hg^0(g) + OH$   
428 oxidation reactions has been proposed by Sommar et al., 2001, although HgOH is

429 highly unstable and could decompose back rapidly to  $\text{Hg}^0$  and OH (Sommar et al.,  
430 2001;Goodsite et al., 2004). It has been proposed that the presence of other gases X  
431 ( $X = \text{NO}_2, \text{HO}_2, \text{RO}, \text{RO}_2, \text{or NO}$ ) could assist the formation of Hg(II) by forming  
432 X-HgOH, which outcompetes the decomposition of HgOH (Calvert and Lindberg,  
433 2005;Dibble et al., 2012;Wang et al., 2014). As an example, we calculated the  
434 transformation between GEM and GOM under the influence of  $\text{NO}_2$ , using the  
435 reactions and rate constants shown in Table S1. As shown in Fig. S4, the production  
436 rate of  $\text{NO}_2\text{HgOH}$ ,  $d[\text{NO}_2\text{HgOH}]/dt$ , increased almost linearly with increasing  $\text{NO}_2$   
437 under low  $\text{NO}_2$  concentrations, and eventually reached a steady state when the  $\text{NO}_2$   
438 concentration is high enough.

439 Based on the production rate of  $\text{NO}_2\text{HgOH}$ , we can estimate the production of  
440  $\text{NO}_2\text{HgOH}$  during the 1 hr sampling period when GOM was captured by the  
441 KCl-coated denuder in the Tekran 1130 unit. The production of  $\text{NO}_2\text{HgOH}$  and  
442  $d[\text{NO}_2\text{HgOH}]/dt$  corresponding to different  $\text{NO}_2$  concentrations is shown in Table 4.  
443 With the increase of the  $\text{NO}_2$  concentration, the contribution of the  $\text{NO}_2\text{HgOH}$   
444 production to GOM will increase. If the  $\text{NO}_2$  concentration is within 100 ppbv (from  
445 0 to 100 ppbv), the production of  $\text{NO}_2\text{HgOH}$  would be in range of  $0.058\text{-}4.81 \text{ pg m}^{-3}$   
446 during the 1h sampling period. As illustrated in Table 4, the level of  $\text{NO}_2$  observed in  
447 our study is high enough to make increase the GOM production. Our results thus  
448 support a recent study in the tropical equatorial Pacific (Wang et al., 2014) that  $\text{NO}_2$   
449 aggregation with HgOH provides provides a possible mechanism to explain the  
450 enhanced production of GOM and the role of  $\text{NO}_2$  might be more important in the  
451 urban air. More laboratory and modeling studies on mercury oxidation mechanism in  
452 the presence of  $\text{NO}_2$  and other gases are thus warranted.

## 453 5. Summary

454 Continuous measurements of speciated atmospheric mercury were conducted at  
455 Hefei, a mid-latitude inland city in central China, from July 2013 to June 2014.  
456 Measurements of other trace gases (e.g. CO, O<sub>3</sub>, NO<sub>2</sub>) and meteorological parameters  
457 were employed to better understand the sources and oxidation pathways of  
458 atmospheric mercury. The mean GEM, GOM and PBM concentrations during haze  
459 days were  $4.74 \pm 1.62 \text{ ng m}^{-3}$ ,  $4.32 \pm 8.36 \text{ pg m}^{-3}$  and  $60.2 \pm 131.4 \text{ pg m}^{-3}$ , respectively.  
460 Potential source contribution function (PSCF) analysis suggested that the local  
461 mercury emission rather than long-range transport is the most important contributor of  
462 atmospheric mercury pollution during haze days at our monitoring site. The low  
463 GEM/CO ratio in Hefei could be a result of local incomplete combustion sources such  
464 as residential coal and biomass burning. Haze pollution has a more profound impact on  
465 PBM than on GEM and GOM. PBM showed a remarkable seasonal pattern, with  
466 higher concentrations in cold seasons and lower in warm seasons. Elevated PBM  
467 concentrations might be due to both the high loadings of particle matter and  
468 disadvantageous diffusion conditions during haze days especially in cold months.  
469 Both GEM and PBM concentrations exhibited great variations with elevated  
470 concentration during night. The diurnal variations of GEM and PBM might be related  
471 to the boundary layer depth; a lower boundary layer depth in the morning and night  
472 could elevate the mercury concentration.

473 Unlike with the diurnal variations of GEM and PBM, GOM concentration  
474 remained relatively constant during night, and then increased rapidly prior to the  
475 sunrise. The enhancement of GOM during daytime could be due to both the transport  
476 of GOM-enriched free troposphere air to the boundary layer and in situ oxidation of

477 GEM in the boundary layer. Simple photochemical modeling supports the occurrence  
478 of daytime oxidation of GEM to GOM. Based on HgOH as an intermediate product,  
479 our calculations suggest that NO<sub>2</sub> aggregation with HgOH is a potential mechanism to  
480 explain the enhanced production of GOM over the inland urban air.

481

482

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691

692

693 **Table 1. Summary of GEM, GOM and PBM concentrations measured in Hefei**  
694 **from July 2013 to June 2014.**  
695

	GEM (ng m <sup>-3</sup> )			GOM (pg m <sup>-3</sup> )			PBM (pg m <sup>-3</sup> )		
	Mean ± σ	Range	N	Mean ± σ	Range	N	Mean ± σ	Range	N
Spring	3.89 ± 1.79	0.2-21.3	7890	4.49 ± 4.22	0.5-69.8	526	8.34 ± 8.97	1.6-130.1	542
Summer	4.08 ± 1.99	0.3-22.9	6050	3.66 ± 4.39	0.5-45.2	511	3.61 ± 4.38	0.5-41.9	570
Autumn	4.51 ± 2.10	0.4-23.8	3632	5.65 ± 8.93	0.5-78.9	274	59.9 ± 153.5	0.5-1615	339
Winter	4.05 ± 1.81	0.9-12.2	6381	2.59 ± 2.58	0.5-9.5	541	56.1 ± 134.9	0.5-1827	639
Total	4.07 ± 1.91	0.2-23.8	23953	3.67 ± 5.11	0.5-78.9	1852	30.02 ± 100.3	0.5-1827	2090
Non-haze	3.95 ± 1.93	0.2-23.8	20345	2.49 ± 2.41	0.5-33.5	1508	23.3 ± 90.76	0.5-1827	1708
Haze	4.74 ± 1.62	2.1-16.5	3608	4.32 ± 8.36	0.5-78.9	344	60.2 ± 131.4	1.6-1615	382

696

697

698 **Table 2. Speciated atmospheric mercury concentrations in Hefei and other urban**  
699 **and rural areas.**  
700

Location	Classification	Time	TGM (ng m <sup>-3</sup> )	GEM (ng m <sup>-3</sup> )	GOM (pg m <sup>-3</sup> )	PBM (pg m <sup>-3</sup> )	Reference
Hefei	Suburb	Jul 2013-Jun 2014	4.1	4.07	3.67	30	This study
Hefei	Suburb	Feb-May 2009	2.53	-	-	-	Hu et al. (2014)
Beijing	Rural	Dec 2008-Nov 2009	3.23	3.22	10.1	98.2	Zhang et al. (2013)
Shanghai	Urban	Aug-Sep 2009	2.7	-	-	-	Friedli et al. (2011)
Nanjing	Urban	Jan-Dec 2011	7.9	-	-	-	Zhu et al. (2012)
Guiyang	Urban	Nov 2001-Nov 2002	8.4	-	-	-	Feng et al. (2004)
Guiyang	Urban	Aug-Dec 2009	-	9.72	35.7	368	Fu et al. (2011)
Changchun	Urban	Jul 1999-Jan 2000	18.4	-	-	276	Fang et al. (2004)
Changchun	Suburb	Jul 1999-Jan 2000	11.7	-	-	109	Fang et al. (2004)
Mt.Changbai	Remote	Aug 2005-Jul 2006	3.58	-	65	77	Wan et al. (2009a, b)
Mt.Gongga	Remote	May 2005-July 2006	3.98	-	6.2	30.7	Fu et al. (2008a, b)
Mt.Waliguan	Remote	Sep 2007-Aug 2008	1.98	-	7.4	19.4	Fu et al. (2012a)
Mt.Leigong	Remote	May 2008-May 2009	2.8	-	-	-	Fu et al. (2010)
Shangri-La	Remote	Nov 2009-Nov 2010	2.55	-	8.22	38.82	Zhang (2015)
Detroit, USA	Urban	Jan-Dec 2004	-	2.5	15.5	18.1	Liu et al. (2010)
Dexter, USA	Rural	Jan-Dec 2004	-	1.6	3.8	6.1	Liu et al. (2010)
Houston, USA	Urban	Aug-Oct 2006	-	1.66	6.9	2.5	Brooks et al. (2010)
Florida, USA	Urban	Jul 2009-Jul 2010		1.3	3	2	Peterson et al. (2012)
Maryland, USA	suburb	2007-2015		1.41	4.6	8.6	Ren et al. (2016)
Göteborg, Sweden	Urban	Feb-Mar 2005	-	1.96	2.53	12.5	Li et al. (2008)
Nova Scotia, Canada	Urban	Jan 2010- Dec 2011		1.67	2.07	2.32	Cheng et al. (2014)
Northern Hemisphere background value					1.5-1.7		Lindberg et al. (2007)

701

702 **Table 3. Correlation coefficients and slopes between GEM concentration and CO**  
 703 **concentration during atmospheric mercury pollution events.**

704

Event	Start Time (UTC + 8 hr)	End Time (UTC + 8 hr)	Duration (h)	GEM (ng m <sup>-3</sup> )	CO (ppbv)	GEM/CO (slope, ng m <sup>-3</sup> ppbv <sup>-1</sup> )	R <sup>2</sup>
1	2013/11/21 03:00	2013/11/22 02:00	23	8.37±2.42	4481.6±717.3	0.0018	0.29
2	2013/12/03 20:00	2013/12/04 09:00	13	7.51±0.67	5270.0±744.5	0.0001	0.02
3	2013/12/07 04:00	2013/12/09 04:00	48	9.21±1.16	5943.8±1394.1	0.0004	0.23
4	2013/12/19 09:00	2013/12/20 09:00	24	4.35±0.17	3907.6±353.0	0.0002	0.03
5	2013/12/24 19:00	2013/12/25 15:00	20	5.58±0.94	4930.8±919.7	0.0012	0.01
6	2014/01/17 22:00	2014/01/19 13:00	39	5.80±0.83	5746.3±1626.9	0.0003	0.28
7	2014/01/25 02:00	2014/01/25 22:00	20	6.03±0.50	8797.9±2244.3	0.0002	0.59
8	2014/03/16 05:00	2014/03/16 20:00	15	4.46±0.47	2261.7±440.2	0.0010	0.79
9	2014/03/17 06:00	2014/03/18 12:00	30	8.85±2.46	2697.1±590.3	0.0030	0.51
10	2014/05/21 00:00	2014/05/21 11:00	11	5.74±0.94	3676.7±1690.0	0.0050	0.79

705 Notes: these episodes were identified using the following criteria: (a) the duration of  
 706 elevated GEM concentration lasted for >10h; (b) the selected hourly average GEM  
 707 concentration higher than the seasonal average GEM concentration.

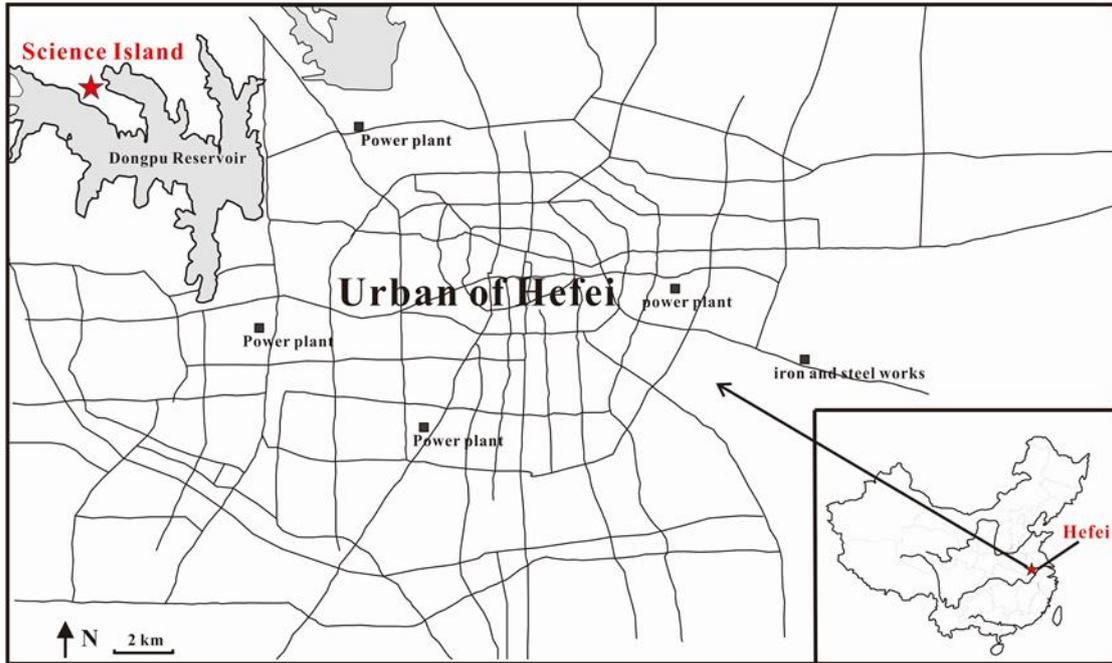
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710 **Table 4. The production of NO<sub>2</sub>HgOH and d[NO<sub>2</sub>HgOH]/dt at different NO<sub>2</sub>**  
 711 **concentrations**

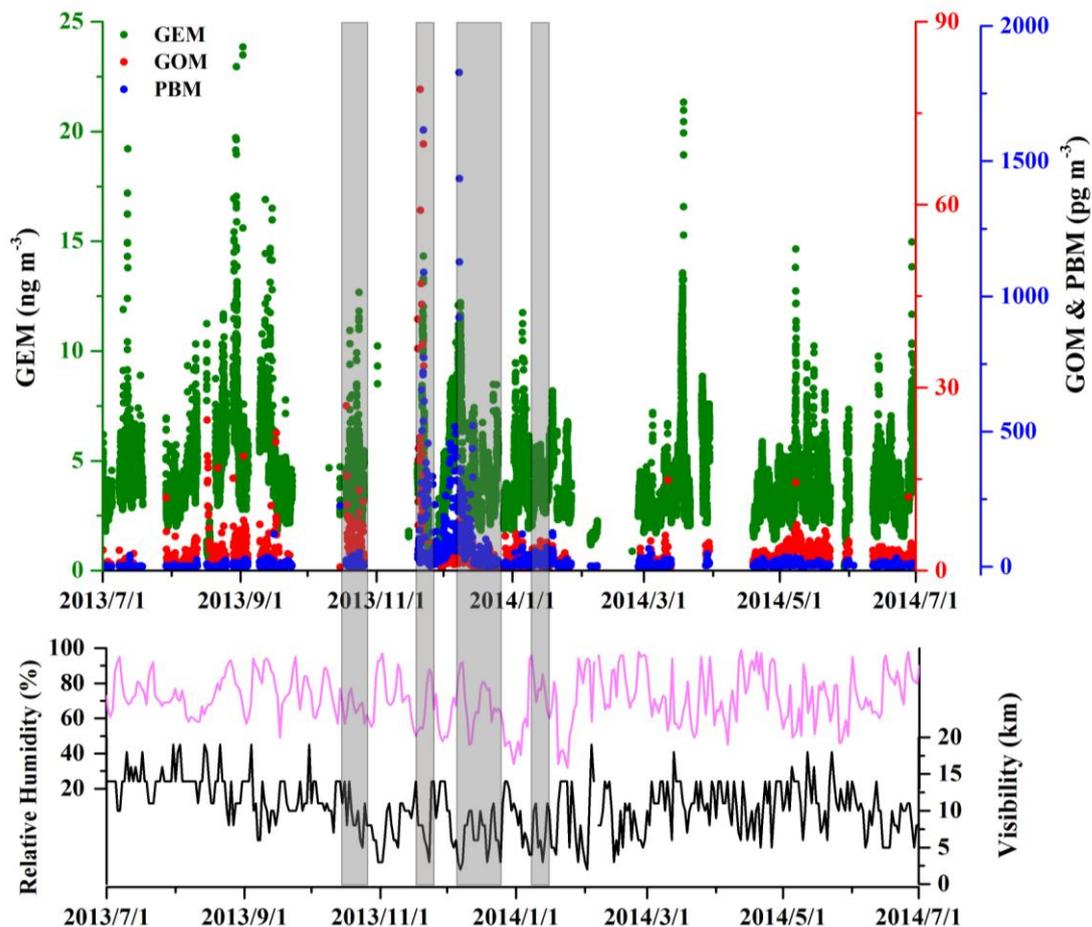
NO <sub>2</sub> (ppbv)	10	20	30	40	50	60	70	80	90	100
d(NO <sub>2</sub> HgOH)/dt (molecule cm <sup>-3</sup> s <sup>-1</sup> )	0.36	0.71	1.04	1.37	1.68	1.99	2.28	2.56	2.83	3.10
NO <sub>2</sub> HgOH (pg m <sup>-3</sup> , 1hr)	0.56	1.10	1.63	2.13	2.61	3.08	3.54	3.97	4.40	4.81

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**Fig. 1.** Location of the study site in Hefei, China.



716

717 **Fig. 2.** Time series of GEM, GOM and PBM concentrations, along with visibility,

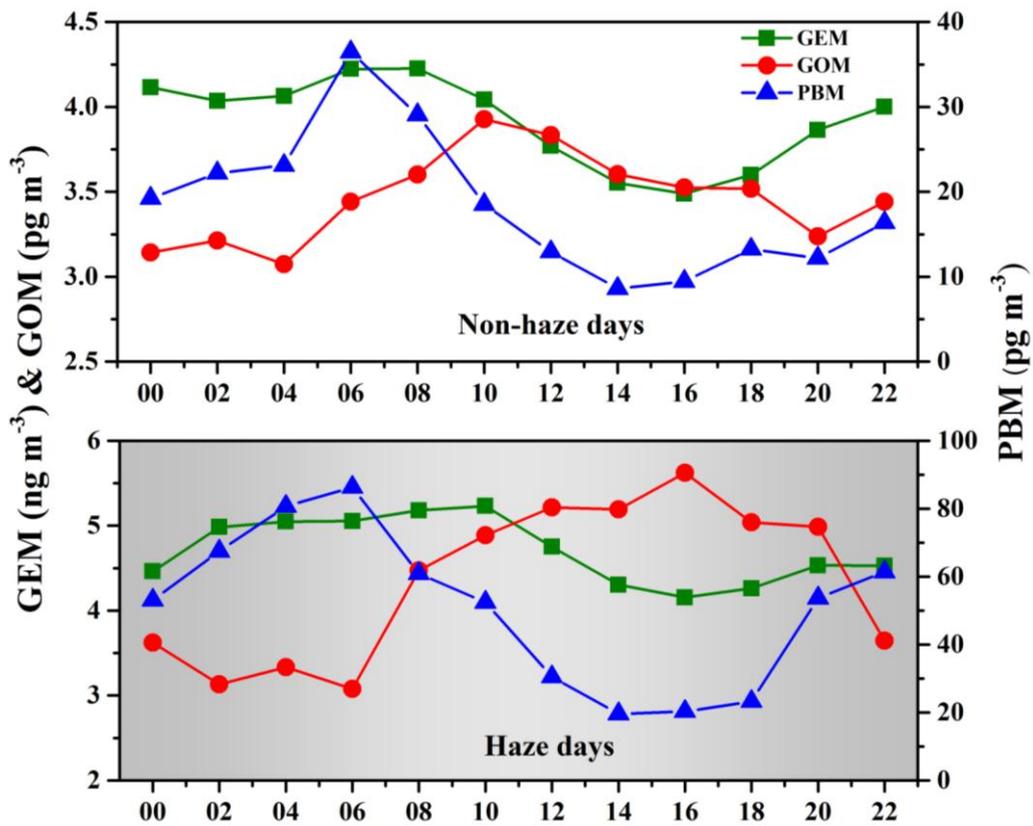
718 relative humidity, at the monitoring site in Hefei from July 2013 to June 2014. The

719 GEM data were at a 5-min resolution, and the GOM and PBM data were two-hour

720 averages. The gray columns show the major haze pollution episodes occurred during

721 the study period.

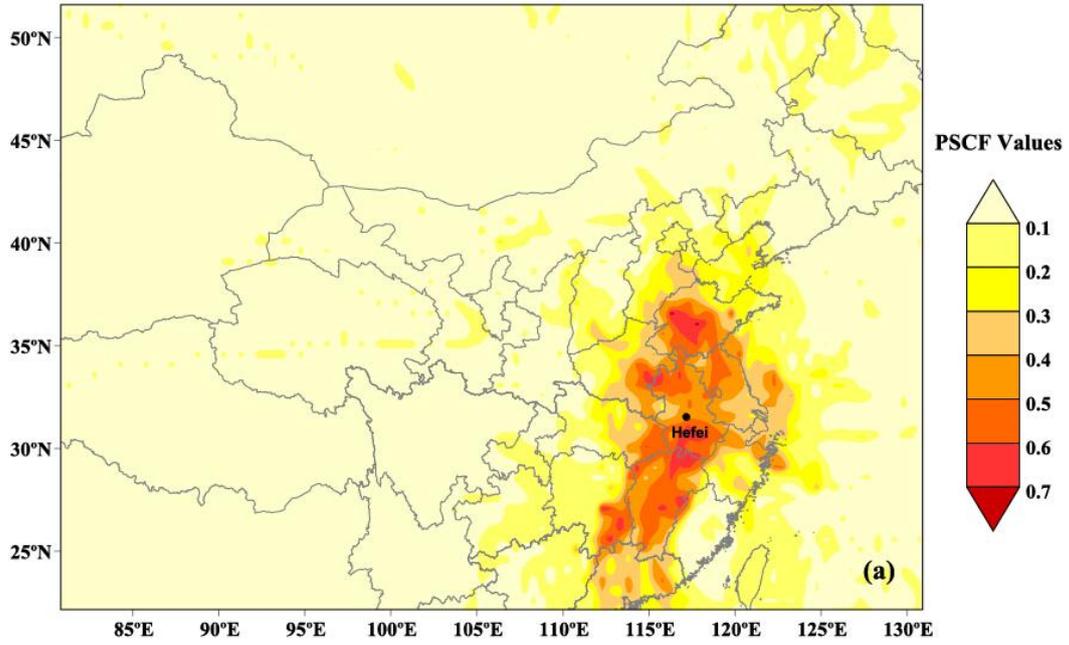
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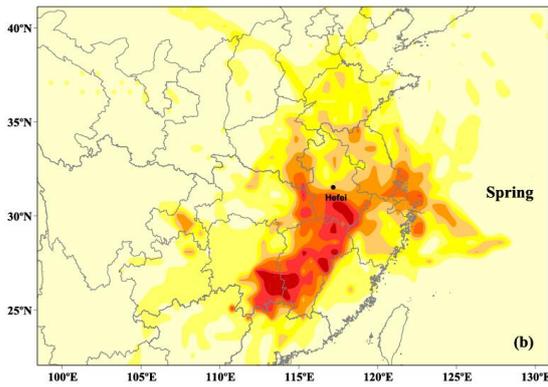
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724 **Fig. 3.** Diurnal trends of GEM, GOM and PBM concentrations in Hefei during  
 725 non-haze and haze days (Local time = UTC + 8 hr). The data were two-hour averages.

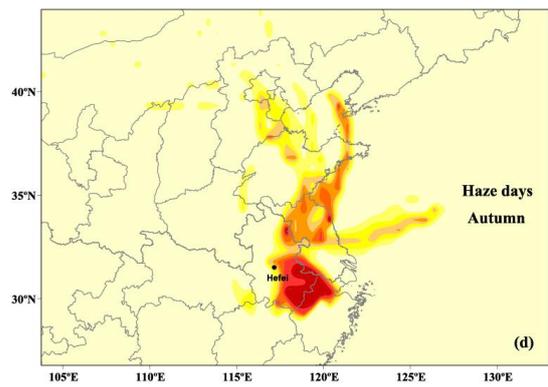
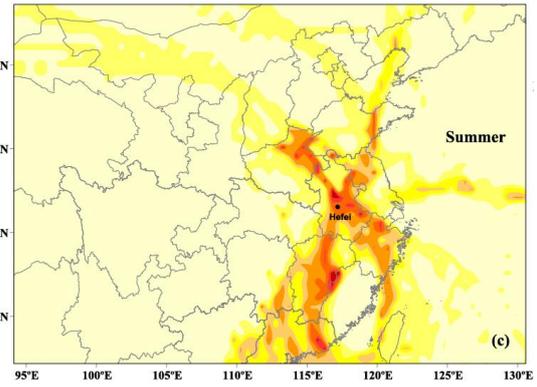
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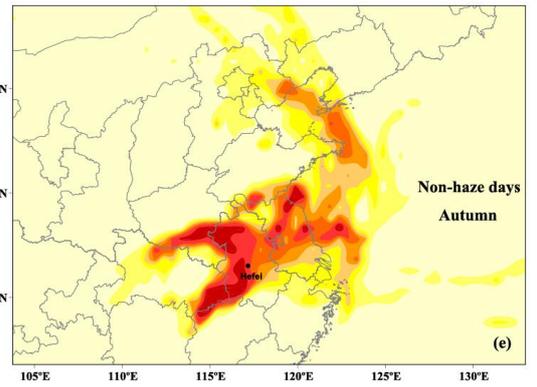
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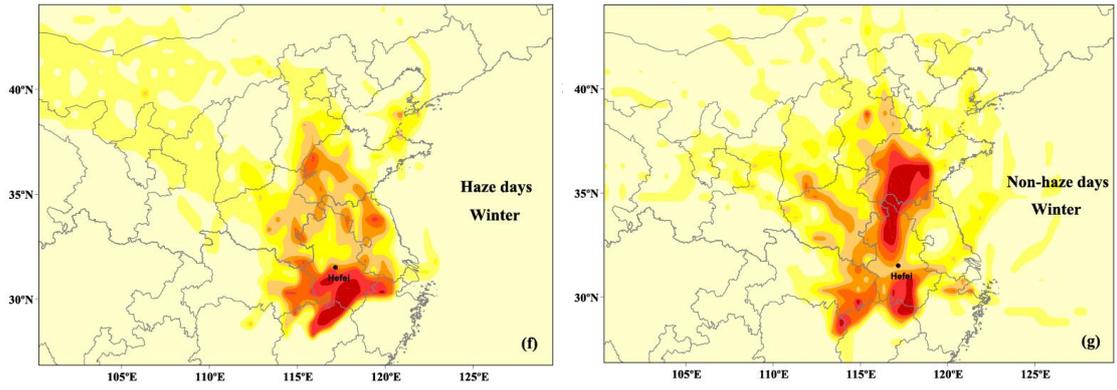


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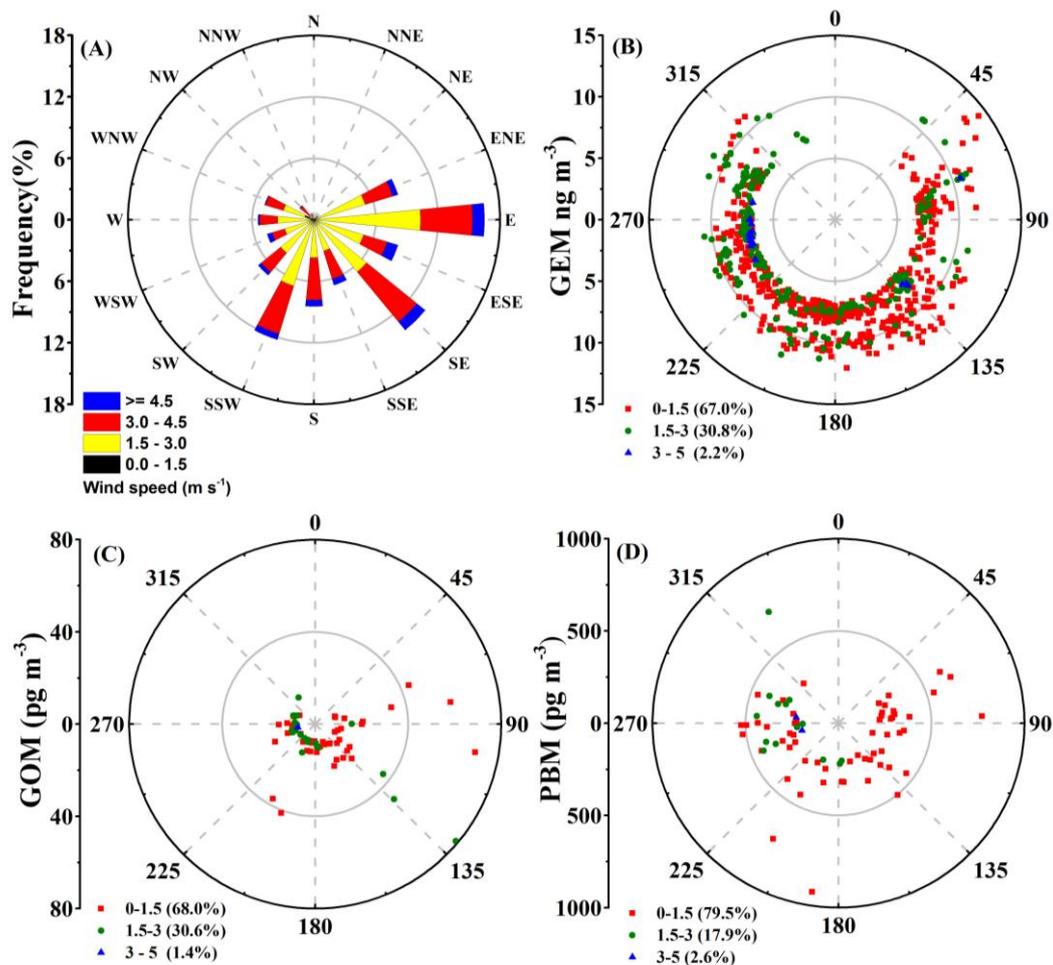
735

736 **Fig. 4.** Likely emission source areas of GEM simulated by PSCF analysis. (a) overall

737 (from July 2013 to June 2014), (b) spring, (c) summer, (d) haze days in autumn, (e)

738 non-haze days in autumn, (f) haze days in winter, (g) non-haze days in winter.

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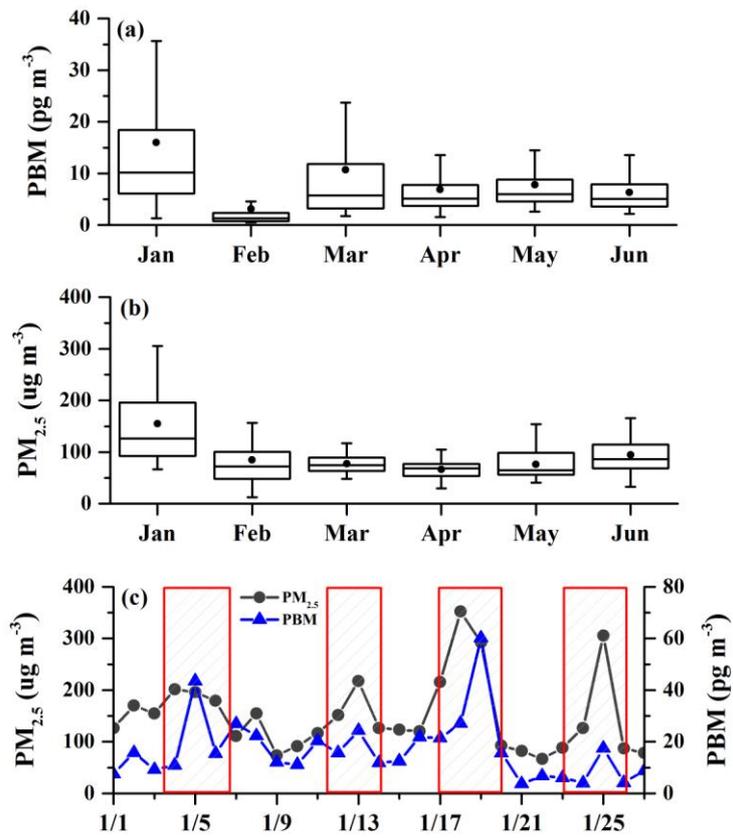
741 **Fig. 5.** Wind direction and speed at the Science Island Meteorological Station during

742 the study period. (A) the wind rose for the whole study period; (B), (C) and (D) are

743 the wind rose diagrams for GEM, GOM and PBM concentrations above the 90th

744 percentile values, respectively.

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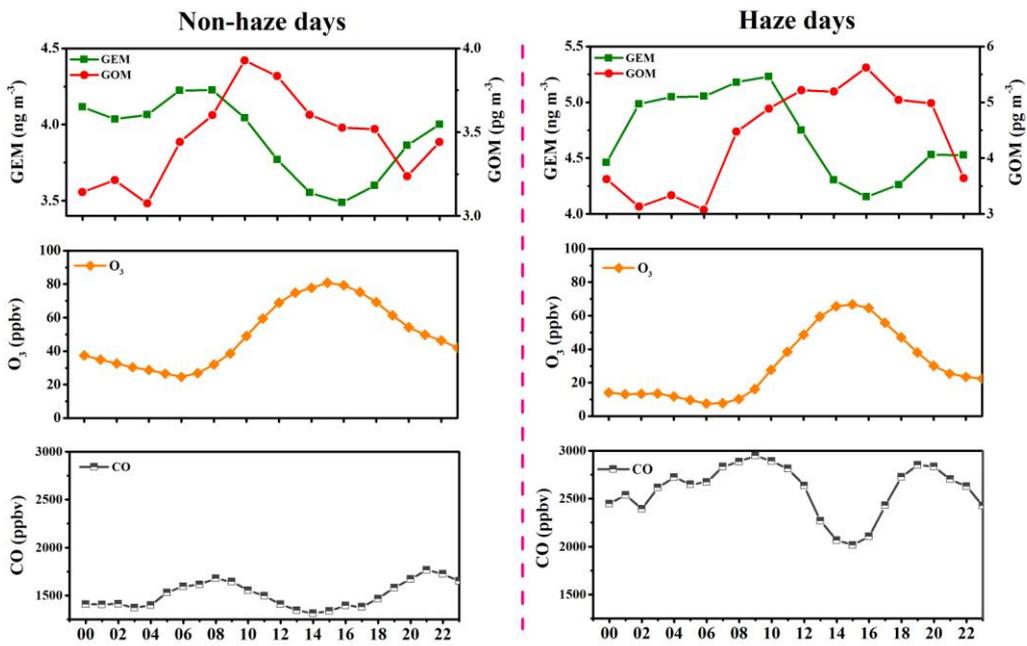


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747 **Fig. 6.** Monthly variation of (a) PBM and (b) PM<sub>2.5</sub> concentrations from January to

748 June, 2014, (c) average daily PM<sub>2.5</sub> and PBM concentrations in January, 2014.

749



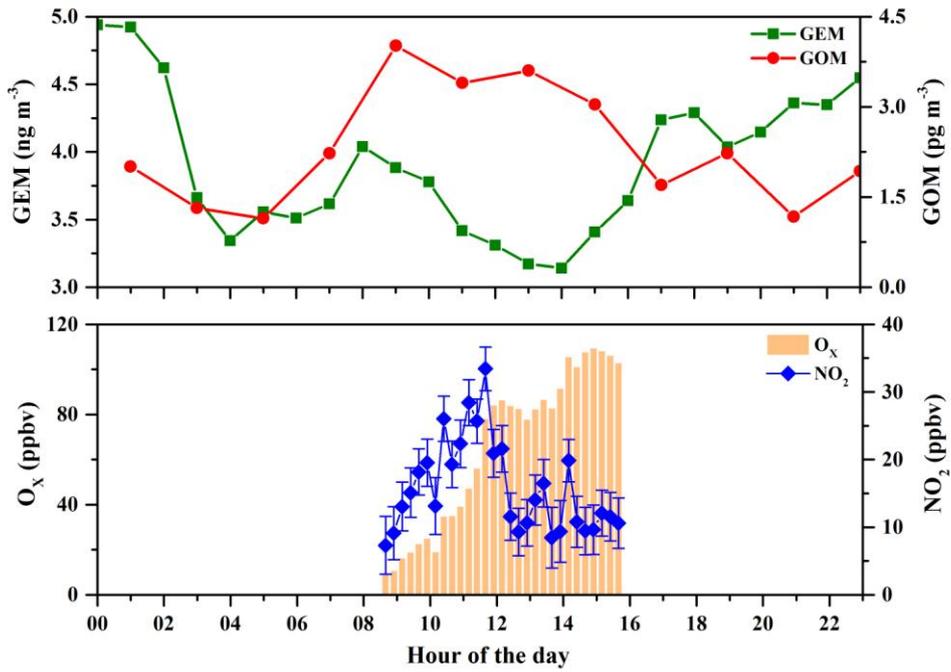
750

751 **Fig. 7.** Diurnal variations of GEM, GOM, O<sub>3</sub> and CO concentrations during non-haze

752 and haze days.

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754



755

756 **Fig. 8.** A case study of diurnal variations of GEM, GOM, O<sub>x</sub>, and NO<sub>2</sub> at Hefei (20th  
 757 November, 2013). The top panel shows the hourly averaged GEM and GOM  
 758 concentrations, and the bottom panel shows the O<sub>x</sub> (O<sub>x</sub>= NO<sub>2</sub> +O<sub>3</sub>) and the NO<sub>2</sub>  
 759 concentrations. The error bars for NO<sub>2</sub> refer to the NO<sub>2</sub> standard errors.

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