

1 ***Response to Anonymous Referee #1***

2 We thank the reviewer for the constructive suggestions/comments. Below we
3 provide a point-by-point response to individual comments (comments in italics,
4 responses in plain font; page numbers refer to the ACPD version; figures used in the
5 response are labeled as Fig. R1, Fig. R2,...).

7 ***Comments and suggestions:***

8 *NO₂ measurement: MAX-DOAS measures slant column densities which were*
9 *converted to vertical column densities as described in line 157 to 162. The authors*
10 *describe how they convert a vertical column density to local NO₂ concentrations in the*
11 *Supplement. They assume homogeneous concentrations within a 500 m thick boundary*
12 *layer (BL) irrespective of daytime. Constant height of a boundary layer (BL) over a*
13 *daytime is not realistic and will deliver a false diurnal variation of NO₂ concentrations.*
14 *Neither is a constant height of BL of 500 m applicable to different seasons.*

15 **Responses and Revisions:**

16 Differential Optical Absorption Spectroscopy (DOAS) is used to retrieve NO₂ and
17 O₄ differential slant column densities (DSCDs) from the measured scattered sunlight
18 spectra (Platt, 1994). In this study, each MAX-DOAS scanning cycle consists of eight
19 elevation viewing angles (2°, 3°, 6°, 8°, 10°, 20°, 30° and 90°) and lasts about 15 min.
20 The spectra are analyzed using the QDOAS spectral-fitting software suite developed at
21 BIRA-IASB (<http://uv-vis.aeronomie.be/software/QDOAS/>). Detail information about
22 the spectral fitting for NO₂ and O₄ is listed in Table S2 (in the revised supplement). As
23 pointed out by the reviewer, MAX-DOAS vertical column densities could not fully
24 represent the NO₂ surface concentration. So we used the HEIPRO (Heidelberg Profile,

25 developed by IUP Heidelberg) retrieval algorithm to retrieve NO₂ vertical profiles for
26 each MAX-DOAS scanning cycle. The purpose of calculating NO₂ profiles is to know
27 the NO₂ vertical distribution. More details about NO₂ profile retrieval are described in
28 the revised supplement. The NO₂ vertical profile was shown in Fig. R1, indicating that
29 the NO₂ is not homogeneously distributed vertically. We agree with the reviewer that
30 converting NO₂ DSCD_s to mixing ratios by assuming that the trace gases were
31 homogeneous within the 500 m height of the boundary is not suitable. We took the
32 suggestion from the reviewer and updated the NO₂ results in the revision. The retrieval
33 altitude grid is 80 layers of 50m thickness between 0.02 and 3.97 km. Thus, in the
34 revised manuscript we have revised our method by using the surface NO₂ concentration
35 (0.02 km) which from the NO₂ vertical profile (Fig. R1) to analyze. Due to the large
36 computational requirement, we were not able to complete the calculation of the NO₂
37 vertical profile for the whole year. However, in Figure R1 we showed the result of one
38 such NO₂ vertical profile (20th November, 2013), and the DFS and the errors of the
39 retrieval are to determine whether the retrieved method is reasonable or not. As shown
40 in Figure R2, the results suggest that the retrieved NO₂ vertical profile is reliable
41 according to the experience of other research (Wang et al., 2014).

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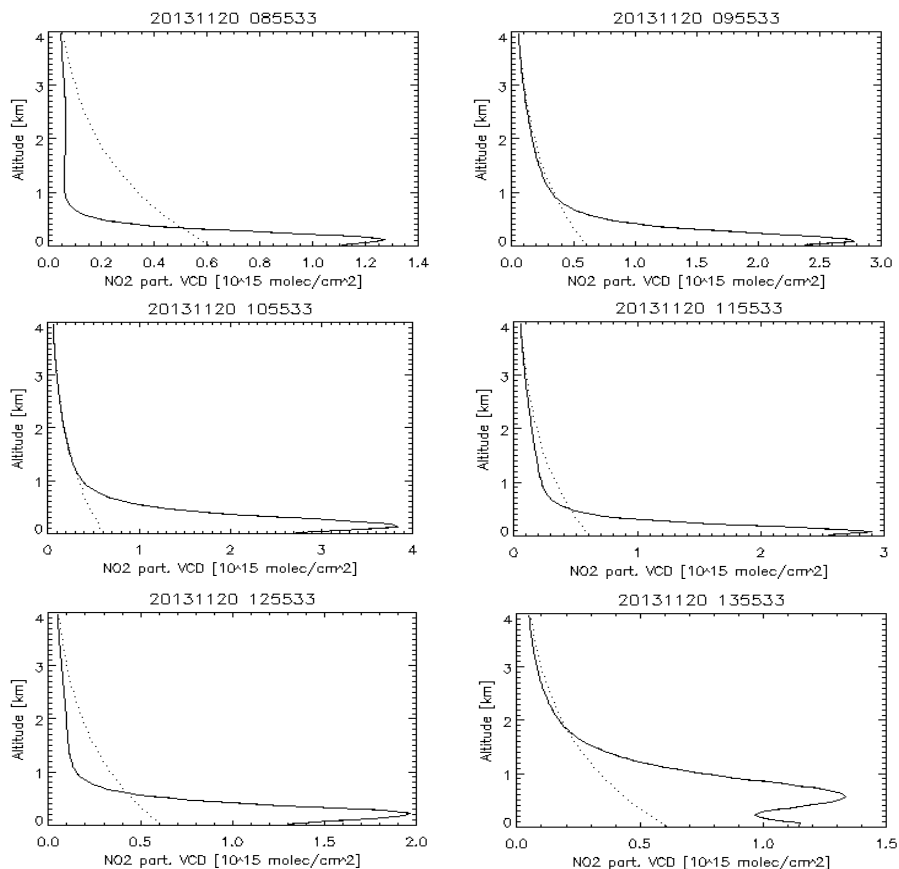
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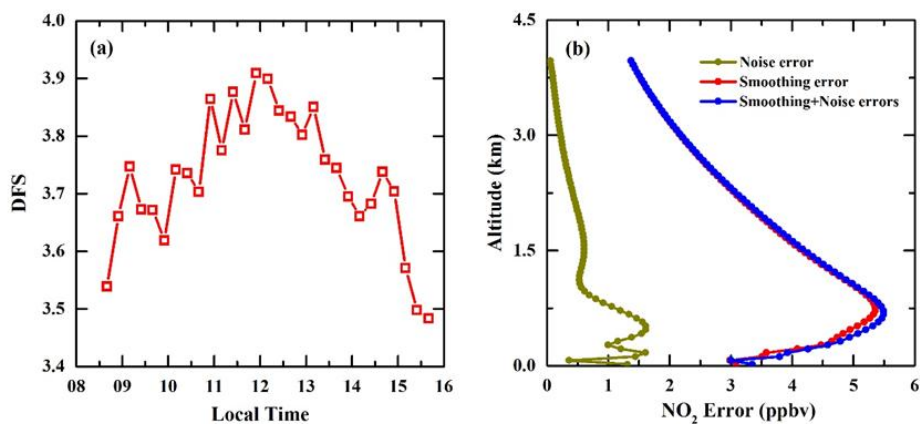


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Figure R1. Example NO₂ vertical profiles at six different times (shown on top of each graph as YYYYMMDD hhmmss) from MAX-DOAS measurements in Hefei (20 November, 2013).

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68 **Figure R2.** (a) DFS diurnal cycles corresponding to the NO₂ profile retrievals; (b)
 69 Errors of NO₂ vertical profile retrieval from MAX-DOAS measurements at Hefei (20
 70 November, 2013 at 10:25LT).

71 ***Comments and suggestions:***

72 *Section 4.1: The discussion of the PSCF results is difficult to follow. Figure 5 shows*
 73 *potential source areas of GEM during the haze events in December 2013 and January*
 74 *2014 but the equivalent figures for non-haze days in December 2013 and January 2014*
 75 *are shown only in supplementary information. It is their difference which can provide*
 76 *the information about the reason for higher GEM during the hazy days. Dtto about the*
 77 *Figure 6: two seasonal data sets should be presented, one for hazy days and one for*
 78 *non-hazy ones.*

79 **Responses and Revisions:**

80 We have updated and merged PSCF results for potential source areas analysis of
 81 GEM in the revised manuscript. Two seasonal data sets are now include, one for haze
 82 days and the one for non-haze days. Since the number of haze days accounts only for
 83 5.6% of the total days in spring and summer, we did not provide haze and non-haze
 84 PSCF results for spring and summer seasons. As autumn and winter are the prevalent

85 seasons for haze pollution, one PSCF result for haze days and another for non-haze
86 days are shown for autumn and winter, respectively. We have combined Figure 5 and
87 Figure 6 into a new figure (Figure 4) in the revised manuscript. The updated PSCF
88 results showed that higher GEM concentration was mainly influenced by local
89 emission sources during haze days. For non-haze days, the most important mercury
90 sources to the monitoring site were not only the local emission sources, but also those
91 from the neighboring region of Shandong, Henan and Jiangxi provinces. In summary,
92 the increase of GEM concentration during haze days was mainly caused by local
93 emission.

94 ***Comments and suggestions:***

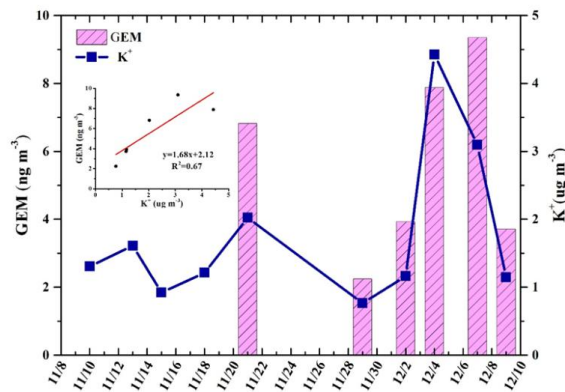
95 *The discussion of GEM vs CO correlations is deeply flawed. The low GEM/CO*
96 *slopes are interpreted as if biomass burning were the major source for both GEM and*
97 *CO in Hefei. To start with GEM/CO slopes represent their emission ratios if a) the*
98 *background concentrations do not change, b) the emissions remain constant, and c)*
99 *there is only dilution, no chemistry, on the way from the source to receptor during an*
100 *event. Using monthly or other “non-event” data would violate at least the condition a)*
101 *and b). In addition, whatever the sources of GEM might be, in a city of 7 million people*
102 *and some 1 million of vehicles most of the CO at the site within the city will almost*
103 *certainly come from local tailpipes rather than from distant isolated fire counts shown*
104 *in Figure S4. The authors present Figure S3 as additional evidence in favour of*
105 *biomass burning being the major source. The figure shows correlation between K+ and*
106 *an Air Quality Index, whose definition is not given in the paper. To be halfway credible,*
107 *K+ has to be correlated with CO. Even if K+ correlated with CO, it still will not prove*
108 *the biomass burning origin of the mercury. For that the density of the fire counts has to*

109 *be consistent with results of the PSCF analysis which it evidently is not. In addition*
110 *remote biomass burning would not yield highest GEM, RGM, and PBM concentration*
111 *at the lowest wind speeds-see section 4.2. In summary, the low GEM/CO ratio is*
112 *characteristic for the emissions of Hefei.*

113 **Responses and Revisions:**

114 Upon further examination of our data, we agree with the reviewer that our original
115 interpretation of the low GEM/CO was not fully supported. Therefore, we have
116 removed Figure S2 (in the ACPD supplement) and revised thoroughly section 4.1
117 regarding the GEM/CO ratio. The definition of mercury pollution events are not same
118 as haze days in this paper. These mercury pollution episodes were defined as a period
119 with hourly average GEM concentration higher than seasonal average GEM
120 concentration and the duration of elevated hourly GEM concentration lasted for over 10
121 hours (Kim et al., 2009). We discussed the correlation coefficients and slopes between
122 GEM concentration and CO concentration during pollution events (Table 3 in the
123 revised manuscript). In previous research, the Hg/CO slope and correlation between
124 GEM and CO concentrations has been used to identify long-range transport episodes or
125 local episodes: significant positive correlation for long-range transport episodes and
126 poor correlation for local episodes (Kim et al., 2009). According to the correlations
127 between GEM concentration and CO concentration, the mercury pollution episodes in
128 autumn and winter mainly belong to local episodes. Incomplete combustion like
129 residential coal and biomass burning combustion could lead to a lower Hg/CO ratio. We
130 agree with the reviewer's point. In summary, the low GEM/CO ratio may be
131 characteristic for the local emissions of Hefei.

132 As for water-soluble potassium (K^+) in 24-hr PM_{10} samples, the correlation
 133 between K^+ and Air Quality Index maybe not reliable. So we did the correlation
 134 between K^+ and GEM during the 24-hr PM_{10} sampling period (Fig. R3). Although the
 135 concentration of water-soluble potassium (K^+) in PM_{10} shows a good correlation
 136 ($R^2=0.67$) with GEM, due to the small number of compared samples ($n=6$), so it has
 137 great accidental and uncertainty. In addition, most pollutant concentration increased
 138 during this heavy pollution episodes (Nov-Dec, 2013). Good correlation might occur
 139 between K^+ and other pollutant, so it cannot fully prove that GEM come from the
 140 emission of biomass burning through good correlation between K^+ and GEM. Thus, we
 141 agree with the reviewer's comment and shortening this section. We have removed the
 142 discussion about K^+ and biomass burning altogether in the revised manuscript and
 143 supplement.



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 145 **Figure R3.** Correlation between water-soluble potassium (K^+) and GEM during
 146 heavy pollution periods (from 10 Nov to 9 Dec, 2013). Notes: water-soluble potassium
 147 (K^+) concentrations were analyzed from 24-hr PM_{10} (particulate matter less than 10 μm
 148 in diameter) samples, GEM concentrations were the average value during the 24-hr
 149 PM_{10} sampling period.

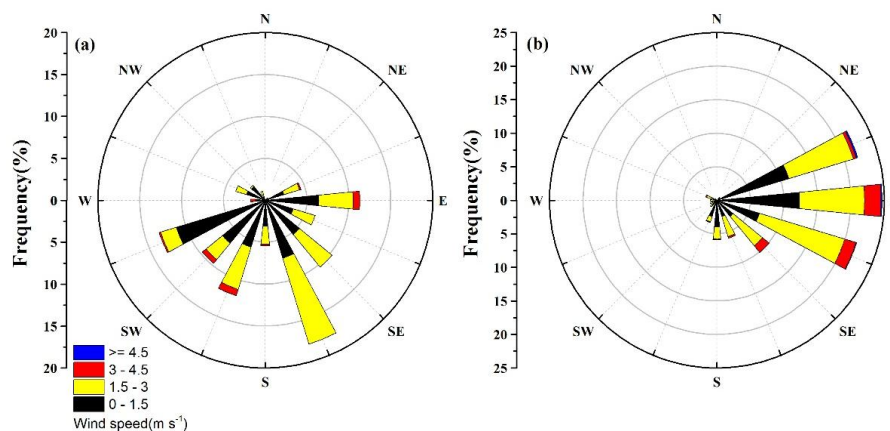
150 **Comments and suggestions:**

151 *Section 4.2: Highest PBM and PM2.5 concentrations in January are most likely due*
152 *to shallower boundary layer in January than in other months. That is probably meant*
153 *by “poor diffusion conditions in cold months”. The average PBM concentrations in*
154 *March differ hardly from other months except for January but their spread is larger. I*
155 *think that the precipitation and the frequency of change of air masses should be also*
156 *taken into account as driving forces for the PBM vs PM2.5 correlation.*

157 **Responses and Revisions:**

158 We agree with the reviewer’s suggestion. Unfortunately, we did not obtain the
159 precipitation data during our monitoring period, so we were not able to directly
160 examine the influence of precipitation on PBM. This is something we will investigate
161 in our future studies.

162 Although on average the PBM concentrations in March differ hardly from other
163 months (April-June), they fluctuated much greatly in March when compared to other
164 months. We have re-examined the wind rose diagrams for March and April (see Fig.
165 R4). The prevailing wind direction in March indeed varied much greatly than in April.
166 Thus, the larger fluctuation of PBM in March might be related to the frequency of
167 change in wind direction. We have thus removed our original interpretation that “higher
168 temperatures in the warmer months do not favor mercury adsorption”, and replaced it
169 with reference to changes in wind direction.



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171 Figure R4. The wind rose diagrams for (a) March and (b) April.

172 **Comments and suggestions:**

173 *Section 4.3: The interpretation of the diurnal variations here is almost certainly*
 174 *wrong. The authors interpret GEM and PBM diurnal variation in terms of changing*
 175 *height of boundary layer and declare that the opposite RGM diurnal variation must be*
 176 *of chemical origin. This must not be and probably is not true. RGM correlates with O3*
 177 *which is probably not formed in situ but admixed from the free troposphere (FT) as the*
 178 *height of BL increases during the morning. Higher RGM concentrations in FT than in*
 179 *BL have been reported by many researchers. Consequently, the RGM correlation with*
 180 *O3 and its anticorrelation with CO can be viewed as solely a transport phenomenon*
 181 *unrelated to any chemical process. The distinction between a transport and chemical*
 182 *processes is a general problem in the interpretation of diurnal variations. It can only be*
 183 *resolved by careful modeling using measured diurnal variation of the BL height and*
 184 *known concentrations in BL and FT or by using specific tracers for photochemical*
 185 *processes such as peroxy nitrates. In this particular case, diurnal variations of GEM,*
 186 *PBM, CO, NO_x, etc. emissions due to morning and evening rush hours, working times,*
 187 *etc. additionally complicate the interpretation of the diurnal variations. As mentioned*

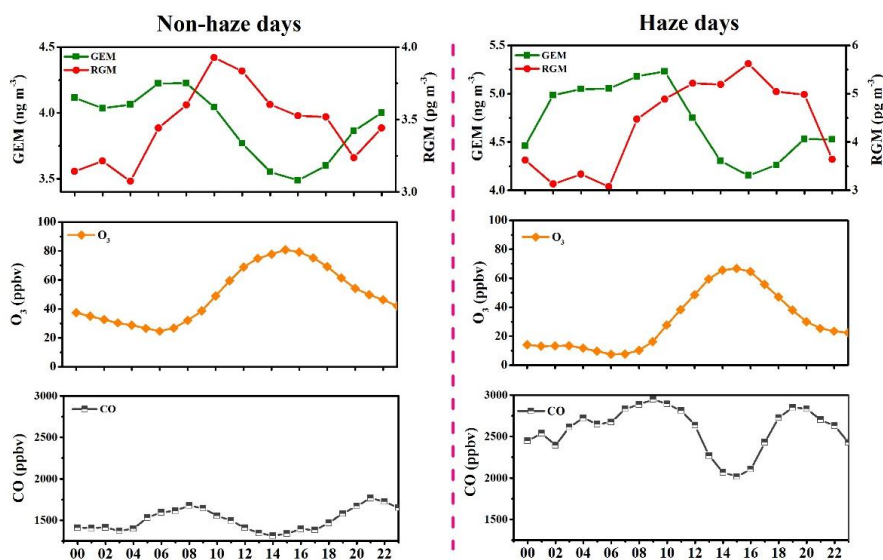
188 *before the diurnal variation of NO₂ is also flawed by the assumption of constant height*
189 *of boundary layer. In summary, the observed diurnal variation can be interpreted solely*
190 *as a transport phenomenon due to air exchange between BL and FT. As long as the*
191 *authors cannot rule out the transport hypothesis their chemical interpretation of the*
192 *diurnal variation and discussion of NO₂ kinetics are wishful thinking without any*
193 *evidential basis.*

194 **Responses and Revisions:**

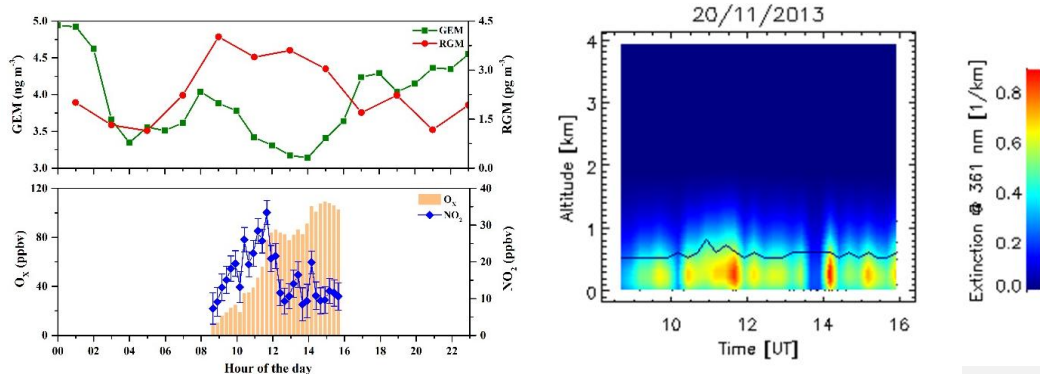
195 We agree with the reviewer that resolving transport and reaction processes of RGM is
196 not straightforward; the fact that we did not measure specific photochemical tracers
197 such as peroxy nitrates did not help. Two processes can affect the RGM concentrations
198 in the boundary layer air. The first is due to transport of RGM from the free
199 troposphere (FT). Diurnal variations of GEM, RGM, O₃ and CO concentrations
200 during non-haze and haze days are shown in Fig. R5 (Figure 7 in the revised
201 manuscript). For both non-haze and haze days, RGM concentrations remained at a
202 relatively constant level during night, and then increased suddenly prior to the sunrise.
203 We agree with the reviewer that such enhancement of RGM in early morning might
204 can be due, at least in part, to its transport from the free troposphere as the height of
205 BL increases. In summary, the observed RGM diurnal variation can be interpreted as a
206 transport phenomenon due to air exchange between BL and FT.

207 In addition, in situ photochemical oxidation of GEM could also increase the
208 concentration of RGM during daytime. To determine the relative importance of FT
209 transport and in situ photochemical oxidation, we examined the relationship between
210 RGM and the changes in the height of the atmospheric boundary layer and the odd
211 oxygen (O_x= O₃+NO₂) concentrations. Although we did not measure peroxy nitrates in

212 this study, we believe the concentration of odd oxygen ($O_X = O_3 + NO_2$ which is
 213 produced from the reaction between O_3 and NO) can be used as a tracer of the extent
 214 of photochemical processing in the urban atmosphere. Since NO_2 concentrations from
 215 MAX-DOAS were only available during daytime, we could only use O_X to be a
 216 indicator for daytime GEM oxidation. As per our manuscript, we selected 20th
 217 November 2013 as a case study to probe the importance of photochemical processes.
 218 Both RGM and O_X reached higher concentrations from 12:00 to 16:00, along with the
 219 lowest value of GEM. The height of atmospheric boundary layer changed very little
 220 over this period (12:00-16:00, see Fig. R6). This simple comparison suggests that the
 221 transport of FT RGM might be limited and that at least some of the RGM were
 222 formed from in situ oxidation of GEM. We further investigated the potential
 223 mechanism of the GEM oxidation to GOM.



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 225 **Figure R5.** Diurnal variations of GEM, RGM, O_3 , and CO concentrations during
 226 non-haze and haze days.
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229 **Figure R6.** A case study of diurnal variations of GEM, RGM, O_x, and NO₂ at Hefei
 230 (20th November, 2013, left panel). The right panel shows the retrieved aerosol
 231 extinction profile on the same day; the black line represents the height of the
 232 atmospheric boundary layer.

233 **Comments and suggestions:**

234 *Line 66-67: PBM is not highly surface reactive. “Affinity” might be better than*
 235 *“reactivity”.*

236 **Responses and Revisions:**

237 Corrected.

238 **Comments and suggestions:**

239 *Line 72: The most recent quoted reference is Pacyna et al. (2006). In 2016 and*
 240 *recent discussions about emissions this seems to be quite obsolete. Dtto line 82. Please*
 241 *quote more recent publications.*

242 **Responses and Revisions:**

243 We have updated the section by quoting two recent publications ((Pacyna et al.,
 244 2010) and (Zhang et al., 2015)).

245 **Comments and suggestions:**

246 *Line 322: The sentence is flawed both in content as in grammar. If taken at face*
 247 *value, the text insinuates emissions from power plants being “non-normal” although*

248 *they represent the largest GEM emissions in most inventories. Reference at line 584 is*
249 *incomplete.*

250 **Responses and Revisions:**

251 The sentence in Line 322 has been modified according to the revision in the revised
252 manuscript (second paragraph of section 4.1). We modified this sentence as follows:
253 “GEM and CO often share similar anthropogenic emission sources, such as industrial
254 coal combustion, domestic coal combustion, iron and steel production and cement
255 production (Wu et al., 2006;Wang et al., 2005). However, they also have their own
256 sources. For instance, power plants and nonferrous metal smelters emit mercury but
257 hardly any CO, while most of CO originates from vehicles which are not a major
258 emitter for mercury.”

259 We also corrected the reference at line 584 (Hu et al., 2014).

260 **Comments and suggestions:**

261 *Figure 8: Bottom plot: which symbol is PBM and which one PM2.5? The caption of*
262 *the Figure 8 seems to be inconsistent with the time scale of the bottom plot. The time*
263 *scale of the bottom plot has not equidistant intervals.*

264 **Responses and Revisions:**

265 We have rearranged the figures, added the symbols of PBM and PM_{2.5} and updated
266 the caption. The time scale of the bottom plot has corrected and had equidistant
267 intervals.

268 **Comments and suggestions:**

269 *Figure 9: The scales of the y-axes should be same for the haze and non-haze days to*
270 *facilitate a comparison. E.g. CO mixing ratios are much higher on hazy days.*

271 **Responses and Revisions:**

272 We have redrawn the figures so that the scales of the y-axes are same for the haze
273 and non-haze days.

274

275 **References**

276 Kim, S. H., Han, Y. J., Holsen, T. M., and Yi, S. M.: Characteristics of atmospheric
277 speciated mercury concentrations (TGM, Hg(II) and Hg(p)) in Seoul, Korea,
278 *Atmospheric Environment*, 43, 3267-3274, doi:10.1016/j.atmosenv.2009.02.038, 2009.

279 Pacyna, E. G., Pacyna, J., Sundseth, K., Munthe, J., Kindbom, K., Wilson, S.,
280 Steenhuisen, F., and Maxson, P.: Global emission of mercury to the atmosphere from
281 anthropogenic sources in 2005 and projections to 2020, *Atmospheric Environment*, 44,
282 2487-2499, 2010.

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284 spectroscopic technique, 127, 27-84, 1994.

285 Wang, L., Zhang, Q., Hao, J., and He, K.: Anthropogenic CO emission inventory of
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288 C., Gielen, C., and Müller, J.-F.: Evaluation of tropospheric SO₂ retrieved from
289 MAX-DOAS measurements in Xianghe, China, *Atmospheric Chemistry and Physics*,
290 14, 11149-11164, 2014.

291 Wu, Y., Wang, S., Streets, D. G., Hao, J., Chan, M., and Jiang, J.: Trends in
292 anthropogenic mercury emissions in China from 1995 to 2003, *Environmental science
293 & technology*, 40, 5312-5318, 2006.

294 Zhang, L., Wang, S., Wang, L., Wu, Y., Duan, L., Wu, Q., Wang, F., Yang, M., Yang,
295 H., and Hao, J.: Updated Emission Inventories for Speciated Atmospheric Mercury
296 from Anthropogenic Sources in China, *Environmental science & technology*, 49,
297 3185-3194, 2015.

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304 ***Response to Anonymous Referee #2***

305 We thank the reviewer for the constructive suggestions/comments. Below we
306 provide a point-by-point response to individual comments (comments in italics,
307 responses in plain font; page numbers refer to the ACPD version; figures used in the
308 response are labeled as Fig. R1, Fig. R2,...).

309 ***Comments and suggestions:***

310 *Minor issues Line 40: “and” should be placed between “spring, summer”.*

311 **Responses and Revisions:**

312 Corrected.

313 ***Comments and suggestions:***

314 *Line 55 and all other lines: Several years ago, the atmospheric mercury community*
315 *stopped using the term “reactive gaseous mercury”. This term was replaced with*
316 *“GOM”, gaseous oxidized mercury. I suggest you change all of your RGM to GOM.*

317 **Responses and Revisions:**

318 Agreed and corrections have been made in the text and figures.

319 ***Comments and suggestions:***

320 *Line 148: change “limits” to “limit”.*

321 **Responses and Revisions:**

322 Corrected.

323 ***Comments and suggestions:***

324 *Line 187: you mentioned an arbitrarily set criterion of 4 ng m⁻³. I really don't know*
325 *much about PSCF, but why do you use arbitrary criteria. How will different arbitrary*
326 *criteria affect your results?*

327 **Responses and Revisions:**

328 The “arbitrarily set criterion” was used in the definition of M_{ij} in the PSCF analysis
329 (see TrajStat_Help_v1.2). The definition of M_{ij} as follows: The number of endpoints
330 for the same cell having arrival times at the sampling site corresponding to pollutant
331 concentration higher than an arbitrarily set criterion is defined to be M_{ij} . In this study,
332 pollutant concentration refers to atmospheric mercury concentration (GEM
333 concentration). The words “arbitrarily set criterion” were used in PSCF method
334 introduction and could also be found in (Fu et al., 2012). However, in the actual
335 operation, we use a fixed GEM value as criterion. In this study, mean GEM
336 concentration of 4 ng m^{-3} during the whole study period was used as fixed criterion
337 (refer to Fu et al., 2012). We split the sentence in Lines 192-195 (ACPD version) to
338 describe clearly in the revised manuscript.

339 ***Comments and suggestions:***

340 *Line 241: “folds” should be fold or two-fold.*

341 **Responses and Revisions:**

342 Corrected

343 ***Comments and suggestions:***

344 *Line 296: change “sources region” to “source region”.*

345 **Responses and Revisions:**

346 Corrected.

347 ***Comments and suggestions:***

348 *Line 333: change “condition” to “conditions”.*

349 **Responses and Revisions:**

350 Corrected.

351 ***Comments and suggestions:***

352 *Line 391: remove (and) from you reference (Sommar et al. 2001).*

353 **Responses and Revisions:**

354 Corrected

355 **Comments and suggestions:**

356 *Line 454: change “may plays” to “may play”*

357 **Responses and Revisions:**

358 Corrected.

359 **Comments and suggestions:**

360 *Line 456: change “greatly” to “great”*

361 **Responses and Revisions:**

362 Corrected.

363 **Comments and suggestions:**

364 *Table 2 should be updated to include mercury speciation studies conducted in the*
365 *US over the past few years. There are several studies in the peer-reviewed literature*
366 *that can be cited in this table, bring it up to date.*

367 **Responses and Revisions:**

368 Agreed. We have added (Peterson et al., 2012) and (Ren et al., 2016) in Table 2.

369 **Comments and suggestions:**

370 *Figure 3. Can you eliminate this figure? These data on already in Table 1.*

371 **Responses and Revisions:**

372 We assume the reviewer meant Figure 4 in the ACPD version. The only overlap of
373 data between this figure and Table 1 are GEM, RGM and PBM mean concentrations.
374 This figure also provides further information such as various percentiles. We agree with
375 the reviewer’s comment and move this figure into the supplement.

376 **Comments and suggestions:**

377 *Figures 4 and 5. Can you merge these two figures into a single figure?*

378 **Responses and Revisions:**

379 We assume the reviewer meant Figures 5 and 6. We have merged PSCF results for
380 potential source areas analysis of GEM in the revised manuscript. Two seasonal data
381 sets have presented, one for haze days and another one for non-haze days.

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385 **References**

386 Fu, X. W., Feng, X., Shang, L. H., Wang, S. F., and Zhang, H.: Two years of
387 measurements of atmospheric total gaseous mercury (TGM) at a remote site in Mt.
388 Changbai area, Northeastern China, *Atmospheric Chemistry and Physics*, 12,
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392 Florida, USA, *Science of the Total Environment*, 424, 297-307, 2012.

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395 measurements at a suburban site in the Mid-Atlantic United States: Inter-annual,
396 seasonal and diurnal variations and source-receptor relationships, *Atmospheric
397 Environment*, 2016.

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404 **Speciated Atmospheric Mercury during Haze and Non-haze Days in**
405 **an Inland City in China**

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408 Qianqian Hong^{1,3}, Zhouqing Xie^{1,2,3*}, Cheng Liu^{1,2,3*}, Feiyue Wang⁴, Pinhua Xie^{2,3},
409 Hui Kang¹, Jin Xu², Jiancheng Wang¹, Fengcheng Wu², Pengzhen He¹, Fusheng Mou²,
410 Shidong Fan¹, Yunsheng Dong², Haicong Zhan¹, Xiawei Yu¹, Xiyuan Chi¹, Jianguo
411 Liu²

带格式的: 字体颜色: 自动设置, 英语(美国)

414 1. School of Earth and Space Sciences, University of Science and Technology of
415 China, Hefei, 230026, China

带格式的: 字体颜色: 自动设置

416 2. InnovationCAS, Center for Excellence in UrbanRegional Atmospheric Environment
417 of CAS, & Institute of Urban Environment of CAS, Xiamen, 361021, China

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418 3. Key Lab of Environmental Optics & Technology, Anhui Institute of Optics and
419 Fine Mechanics, Chinese Academy of Sciences, Hefei, 230031, China

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420 4. Centre for Earth Observation Science, and Department of Environment and
421 Geography, University of Manitoba, Winnipeg, MB R3T 2N2, Canada

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424 Correspondence author:

425 zqxie@ustc.edu.cn (Z.Q.X.); chliu81@ustc.edu.cn (C.L.)

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428 **Abstract.** Long-term continuous measurements of speciated atmospheric mercury
429 were conducted from July 2013 to June 2014 at Hefei, a mid-latitude inland city in
430 east central China, from July 2013 to June 2014, that experiences frequent haze
431 pollution. The mean concentrations (\pm standard deviation) of gaseous elemental
432 mercury (GEM), ~~reactive~~ gaseous oxidized mercury (RGMGOM) and particle-bound
433 mercury (PBM) were $3.95 \pm 1.93 \text{ ng m}^{-3}$, $2.49 \pm 2.41 \text{ pg m}^{-3}$ and $23.3 \pm 90.8 \text{ pg m}^{-3}$,
434 respectively, during non-haze days, and $4.74 \pm 1.62 \text{ ng m}^{-3}$, $4.32 \pm 8.36 \text{ pg m}^{-3}$ and
435 $60.2 \pm 131.4 \text{ pg m}^{-3}$, respectively, during haze days. Potential source contribution
436 function (PSCF) analysis suggested that the atmospheric mercury pollution during
437 haze days was caused primarily by local mercury emissions, instead of via long-range
438 mercury transport. ~~In addition, the~~ The disadvantageous ~~diffusion~~diffusion during
439 haze days ~~will also enhance~~favoured the ~~level~~accumulation of atmospheric mercury.
440 Compared to ~~the~~ GEM and ~~RGM, change in~~ GOM, PBM was found to be more
441 sensitive to ~~the~~ haze pollution. The mean PBM concentration during haze days was
442 2.5 times that during non-haze days due to elevated concentrations of particulate
443 matter. A remarkable seasonal trend in PBM was also observed with concentration
444 decreasing in the following order in response to the frequency of haze days: autumn,
445 winter, spring, ~~and~~ summer. ~~A distinct~~For both non-haze and haze days, GOM
446 concentrations remained at relatively constant during night, but increased rapidly
447 prior to sunrise. This GOM diurnal relationship was found variation could be due to
448 diurnal variation in air exchange between GEM and RGM during haze days, with the
449 peak values of RGM coinciding with boundary layer and free troposphere, but the
450 decline in GEM. Using HgOH as an intermediate product during GEM contribution
451 from photochemical oxidation, our could not be ruled out. This is supported by simple

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452 box-model calculations, which showed that oxidation of GEM to GOM does occur
453 and that the transport of free troposphere GOM alone is not large enough to account
454 for the observed increase in GOM. Our results suggest postulate that NO₂ aggregation
455 with HgOH could explain the HgOH intermediate may be a potential mechanism for
456 the enhanced production of RGMGOM during the daytime in haze days. Increasing
457 level of NO_x will potentially accelerate the oxidation of GEM despite the decrease of
458 solar radiation.

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461 1. Introduction

462 Mercury (Hg) is an environmental pollutant that has received much global
463 attention because of its toxicity and bioaccumulation via the aquatic food chain. The
464 most important transport pathway of mercury is via the atmosphere (~~Schroeder and~~
465 ~~Munthe, 1998~~(Schroeder and Munthe, 1998;Lindqvist and Rodhe, 1985).
466 Operationally, atmospheric mercury is commonly differentiated into three forms:
467 gaseous elemental mercury (GEM), reactive gaseous oxidized mercury (RGMGOM),
468 and particle-bound mercury (PBM). The sum of these three atmospheric speciated
469 mercury is defined as total atmospheric mercury (TAM = GEM + RGMGOM + PBM),
470 and the sum of GEM and RGMGOM is known as total gaseous mercury (TGM =
471 GEM + RGMGOM) (~~Gustin and Jaffe, 2010;Gustin et al., 2015~~(Gustin and Jaffe,
472 2010;Gustin et al., 2015). GEM is regarded as the predominant dominant form of
473 atmospheric mercury, accounting for over 95% of the total. GEM is stable in the
474 atmospheretroposphere, with a long residence time (0.5–2 yr) and can be transported
475 at the regional to global scale (~~Schroeder and Munthe, 1998;Lindberg et al.,~~

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476 ~~2007)(Schroeder and Munthe, 1998;Lindberg et al., 2007).~~ GEM can be oxidized to
477 RGM through photochemical processes, and further transformed to PBM on aerosol
478 surfaces. Although GEM is the predominant form of mercury in the air, trace amounts
479 of RGM and PBM control the mercury scavenged from the atmosphere (Lindberg and
480 Stratton, 1998). RGM. GEM can be oxidized to GOM through photochemical
481 processes, and further transformed to PBM on aerosol surfaces. GOM and PBM can
482 be readily removed from the air by wet and dry deposition as a result of their high
483 surface reactivity/affinity and water solubility (Lindqvist and Rodhe, 1985). Thus, the
484 chemical transformation/transformation between GEM, RGM/GOM, and PBM will
485 directly influence the atmospheric lifetime of mercury.

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486 As a result of the rapid industrial development and economic growth of recent
487 decades, China has become one of the major contributors to anthropogenic mercury
488 emissions to the environment (Wu et al., 2006;Pacyna et al., 2006;Streets et al.,
489 2005)(Wu et al., 2006;Pacyna et al., 2006;Pacyna et al., 2010;Zhang et al., 2015b).
490 Atmospheric mercury emissions from anthropogenic sources in China have been
491 estimated to be in the range of 500-700 tons/yr, accounting for 25-30% of the total
492 global anthropogenic mercury emissions (Streets et al., 2005;Wu et al., 2006)(Streets
493 et al., 2005;Wu et al., 2006). Research into atmospheric mercury in China is therefore
494 critical to the understanding of mercury cycling at both regional and global scales.
495 Long-term observation of atmospheric mercury has been conducted in different
496 regions in China, including both urban and remote areas. TGM concentrations
497 observed in urban and industrial regions of China were in the range of 2.7-35 ng m⁻³,
498 higher than the values reported for North America and Europe, and for the adjacent
499 Asian countries such as Korea and Japan (Stamenkovic et al., 2007;Dommergue et al.,

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500 ~~2002;Fang et al., 2009~~(Weigelt et al., 2013;Fang et al., 2009;Marumoto et al., 2015).
501 TGM and PBM concentrations in remote areas of China were also found to be higher
502 than those observed in North America and Europe ~~(Fu et al., 2008a;Fu et al.,~~
503 ~~2008b;Fu et al., 2012~~(Fu et al., 2008a;Fu et al., 2008b;Fu et al., 2012;Liu et al., 2010).

504 In recent years, haze pollution has become a major concern in China due to its
505 impacts on visibility, air quality, and climate. It is well known that haze formation is
506 mainly dependent on the atmospheric relative humidity (RH) and the concentration of
507 airborne particles (Chen et al., 2003;~~Sun et al., 2013~~;~~Sun et al., 2013~~). Most studies
508 on haze have focused on the measurements of airborne particulate matter; few
509 examined the influence of haze on the chemistry of atmospheric mercury, especially
510 PBM. In this study, we conducted one year synchronous observations of speciated
511 atmospheric mercury in Hefei, an inland city of China, which experiences frequent
512 haze events. The comparison of atmospheric mercury under haze days and non-haze
513 days during the study period allows us to examine the formation and deposition
514 mechanisms of mercury, as well as their temporal variations.

515

516 **2. Methods**

517 **2.1 Study site**

518 Continuous measurements of speciated atmospheric mercury were undertaken in
519 Hefei (31°52' N, 117°17' E) from July 2013 to June 2014. Hefei, the capital of Anhui
520 Province, is located in east central China, between the Changjiang (Yangtze River) and
521 the Huaihe (Huai River). Hefei has a humid subtropical climate with four distinct
522 seasons: June-August is considered summer, September-November autumn,
523 December-February winter and March-May spring. The prevailing wind is

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524 southeasterly in summer and northwesterly in winter. Like many Chinese cities, Hefei
525 has experienced rapid growth in the past 20 years. The total permanent population is
526 about 7.7 million. The city has also been witnessing an increasing frequency in haze
527 pollution, especially in winter months.

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528 The monitoring site was located on the Science Island, a small peninsula on the
529 Dongpu Reservoir in the northwestern outskirts of Hefei (Fig. 1). The sampling and
530 analytical instruments were installed 1.5 m above the rooftop (~ 20 m above the
531 ground) of the main building of Anhui Institute of Optics and Fine Mechanics. Further
532 information about the monitoring site can be found in a previous study (Hu et al.,
533 2014)(Hu et al., 2014). We chose this area as the monitoring site because it is one of
534 the cleanest areas in Hefei, not adjacent to any direct pollution sources such as power
535 plants, iron and steel works.

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537 2.2 Measurements of speciated atmospheric mercury

538 From July 2013 to June 2014, simultaneous measurements of speciated
539 atmospheric mercury concentrations were performed by an automated TekranTM
540 mercury speciation system. The system consisted of a Model 2537B mercury analyzer
541 combined with a Model 1130 RGMGOM unit and a Model 1135 PBM unit. The
542 system was configured to measure GEM every 5 min., and RGMGOM and PBM
543 every 2 ~~two hr.~~h.

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544 The details about the Tekran-based mercury speciation system can be found in
545 Landis et al. (2002). In general, the automated measurement process can be
546 summarized as sample collection, thermal desorption and determination. During the
547 collection period, ambient air was drawn to the system at a flow rate of 10 L/min.

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548 RGMGOM and PBM in the air were captured by a KCl-coated quartz annular
549 denuder in the 1130 unit and a quartz filter in the 1135 unit, respectively, whereas
550 GEM would pass through the denuder and filter and be quantified on the Tekran
551 2537B by cold-vapor atomic fluorescence spectroscopy (CVAFS). After an hour of
552 sampling, the 1135 quartz filter and the 1130 denuder would be switched to the thermal
553 decomposition mode at 800 °C and 500 °C, respectively, with the resulting Hg⁰
554 quantified by the 2537B unit in the next hour, while the 1135 and 1130 components
555 were flushed with zero-mercury gas for the next sampling.

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556 The instrument maintenance followed typical protocols used in similar studies
557 (Landis et al., 2002;Hu et al., 2014)(Landis et al., 2002;Hu et al., 2014). The quartz
558 annular denuder was recoated every two weeks, the quartz filter was replaced once a
559 month, and the Teflon filter (pore size 0.2 μm) in the sample inlet was changed every
560 two weeks. Automated recalibration of the Tekran 2537B was performed every 25 h
561 using an internal mercury permeation source. No calibration standards were available
562 for RGMGOM and PBM, but the 1σ precision for RGMGOM and PBM was about
563 15 % (Landis et al., 2002)(Landis et al., 2002). The detection limit in ambient air is
564 about 0.5 ng m⁻³ for GEM (or TGM) at a resolution of 5 min, and 1 pg m⁻³ for
565 RGMGOM and PBM at a resolution of 2 h (Gustin et al., 2015)(Gustin et al., 2015).

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566 Although the Tekran-based mercury speciation technique has been widely used
567 around the world, recent studies have shown that the technique does not efficiently
568 collect all gaseous oxidized mercury and thus may substantially underestimate the
569 concentration of reactive mercury (Huang et al., 2013;Gustin et al., 2013)(Huang et
570 al., 2013;Gustin et al., 2013). Therefore, the RGMGOM values reported in this study

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571 should be considered as the lower ~~limits~~limit of gaseous oxidized mercury in the air
572 ~~(Wang et al., 2014)~~(Wang et al., 2014).

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574 **2.3 Ancillary Data**

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575 Standard meteorological measurements including air temperature, air pressure,
576 RH, wind direction and speed were observed with a 5-min resolution. CO was
577 measured by an automated infrared carbon monoxide analyzer (Model EC9830T,
578 Ecotech Inc., Australia), with a detection limit of 40 ppbv. O₃ was measured every 5
579 min by an ozone analyzer (Model EC9810B, Ecotech Inc., Australia); its detection
580 limit and accuracy are 0.5 ppbv and 0.001 ppm, respectively. NO₂ was measured by a
581 Multi axis differential optical absorption spectroscopy (MAX-DOAS) instrument. The
582 collected spectra were analyzed using the QDOAS spectral fitting software suite
583 developed at BIRA-IASB (<http://uv-vis.aeronomie.be/software/QDOAS/>). ~~We used~~

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584 ~~the geometric approximation for conversion between slant column densities (SCDs)~~
585 ~~and vertical column densities (VCDs) (Ma et al., 2013). PM_{2.5} (particulate matter less~~
586 ~~than 2.5 μm in diameter) data are collected from China air quality online analysis~~
587 ~~platform (<http://www.aqistudy.cn/historydata/index.php>). In addition, 24 hr PM₁₀~~
588 ~~(particulate matter less than 10 μm in diameter) samples were collected on glass fiber~~
589 ~~filters by a high volume sampler during heavy pollution episodes (from 10 Nov to 9~~
590 ~~Dec 2013, n=11). Water soluble ions in the PM₁₀ samples were determined by ion~~
591 ~~chromatography (Model ICS 2100, Dionex). In order to identify the potential source~~
592 ~~of mercury, NASA's satellite hotspots/fire locations information were obtained from~~
593 ~~the Fire Information for Resource Management System (FIRMS)~~
594 ~~(<https://firms.modaps.eosdis.nasa.gov/firemap/>). PM_{2.5} (particulate matter less than 2.5~~

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595 [um in diameter\) data are collected from China air quality online analysis platform](#)
596 [\(http://www.aqistudy.cn/historydata/index.php\)](http://www.aqistudy.cn/historydata/index.php).

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598 **2.4 Potential Sources Contribution Function (PSCF) analysis**

599 To identify the possible influence of long-range transport on the distribution of
600 atmospheric mercury in Hefei, we calculated backward trajectories of air masses
601 using the HYSPLIT (Hybrid Single-particle Lagrangian Integrated Trajectory) model
602 with the Global Data Assimilation System (GDAS 1⁹) developed by the National
603 Oceanic and Atmospheric Administration (NOAA) (<http://www.ready.noaa.gov>)
604 (Draxler and Hess, 1998). Considering the atmospheric pollutants are mainly
605 concentrated in the low altitude during heavy pollution days, the trajectory arrival
606 heights were set at 500 m to represent the boundary layer where atmospheric
607 pollutants were well mixed. ~~In this study, 5-day back trajectories were calculated in~~
608 ~~ensemble forms which calculate 27 trajectories from the selected starting point (31°52'~~
609 ~~N, 117°17' E) (Fain et al., 2009).~~In this study, 3-day back-trajectories were generated
610 hourly by TrajStat, a software including HYSPLIT for trajectory calculation (Wang et
611 al., 2009).

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612 The contributions of other pollution source regions to the atmospheric mercury at
613 Hefei was identified by the Potential Sources Contribution Function (PSCF) analysis
614 with the TrajStat software ~~(Wang et al., 2009)~~. PSCF analysis has been shown to be
615 useful in spatially identifying pollution sources for pollutants with a long lifetime
616 such as elemental mercury and CO ~~(Xu and Akhtar, 2010)~~(Xu and Akhtar, 2010). The
617 PSCF values for the grid cells in the study domain were calculated by counting the
618 trajectory segment endpoints that terminate within each cell. The number of endpoints

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619 that fall in the ij_{th} cell is designated as N_{ij} . The number of endpoints for the same cell
 620 corresponding to the atmospheric mercury concentration higher than an arbitrarily set
 621 criterion (~~4 ng m⁻³ which was the~~ defined to be M_{ij} . In this study, mean GEM
 622 concentration of 4 ng m⁻³ during the whole study period) is defined to be M_{ij} . ~~was~~
 623 used as the mercury pollution criterion. The PSCF value for the ij_{th} cell is then defined
 624 as:

$$625 \text{PSCF}_{ij} = \frac{M_{ij}}{N_{ij}} W_{ij} \qquad \text{PSCF}_{ij} = \frac{M_{ij}}{N_{ij}} W_{ij}$$

626 (2)

627 W_{ij} is an arbitrary weight function to reduce the effect of small values of N_{ij} . The
 628 PSCF values were multiplied by W_{ij} to better reflect the uncertainty in the values for
 629 these cells (Polissar et al., 2001)(Polissar et al., 2001). The weight function reduces
 630 the PSCF values when the total number of endpoints in a particular cell is less than 3
 631 times the average value of the end points per cell:

$$632 W_{ij} = \begin{cases} 1.0 & N_{ij} \geq 3N_{ave} \\ 0.70 & 3N_{ave} > N_{ij} \geq 1.5N_{ave} \\ 0.40 & 1.5N_{ave} > N_{ij} \geq N_{ave} \\ 0.20 & N_{ave} > N_{ij} \end{cases} \qquad W_{ij} = \begin{cases} 1.0 & N_{ij} \geq 3N_{ave} \\ 0.70 & 3N_{ave} > N_{ij} \geq 1.5N_{ave} \\ 0.40 & 1.5N_{ave} > N_{ij} \geq N_{ave} \\ 0.20 & N_{ave} > N_{ij} \end{cases}$$

633 (3)

635 3. Results

636 We intended to continuously monitor speciated atmospheric mercury
 637 concentration over the course of a year; however, interruptions were inevitable due to
 638 instrument maintenance, which resulted in loss of data for the following four periods:
 639 (1) 25 September to 9 October 2013; (2) 5-14 November 2013; (3) 9-25 February
 640 2014; and (4) 1-14 April 2014. The rest of the data were grouped into haze days and

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641 non-haze days according to the China Meteorological Administration's haze standard
642 (QX/T 113-2010). Haze days refer to the days when the atmospheric visibility < 10
643 km and RH < 80% (Duan et al., 2016), and non-haze days refer to clear days with the
644 atmospheric visibility > 10 km. The visibility and RH information were collected
645 from the weather history data at the Luogang Airport of Hefei
646 (<http://www.wunderground.com/>). Through the study period of almost a year, a total
647 of 56 days were identified to be haze days, and 253 days to be non-haze days. All the
648 times reported herein are local time (UTC + 8 h).

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650 **3.1 Overall characteristics of speciated atmospheric mercury**

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651 The time series of GEM, **RGMGOM** and PBM concentrations at the study site
652 throughout the study period are shown in Fig. 2, and their frequency distributions are
653 shown in Fig. S1 (in the supporting information). The mean (\pm standard deviation)
654 GEM, **RGMGOM** and PBM concentrations during the whole study period were 4.07
655 ± 1.91 ng m⁻³, 3.67 ± 5.11 pg m⁻³, and 30.0 ± 100.3 pg m⁻³, respectively (Table 1).
656 The GEM concentrations in different seasons did not differ much. The highest GEM
657 concentration occurred in autumn (4.51 ± 2.10 ng m⁻³), while the lowest in spring
658 (3.89 ± 1.79 ng m⁻³). **RGMGOM** concentrations varied greatly during the study period
659 with much higher concentrations in autumn and the lowest in winter. A similar
660 seasonal variation in the **RGMGOM** concentration was observed at a remote site in
661 Mt. Gongga of southwest China (Fu et al., 2008b)(Fu et al., 2008b). The seasonal
662 trend in PBM was also observed in Hefei with its concentration decreasing in the
663 following order: autumn > winter > spring > summer. The mean PBM concentration
664 during the cold season was about 20 times that in summer, similar to the findings

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665 from many previous studies in China (~~Zhang et al., 2013;Fu et al., 2011;Fu et al.,~~
666 ~~2008b;Fang et al., 2001~~)(Zhang et al., 2013;Fu et al., 2011;Fu et al., 2008b;Fang et al.,
667 2001).

668 Comparisons of speciated atmospheric mercury concentrations with other urban
669 and rural areas in China and a few other countries are shown in Table 2. The mean
670 GEM concentration at Hefei is slightly higher than that in many remote areas in China
671 (~~Fu et al., 2008a;Fu et al., 2008b;Fu et al., 2012;Wan et al., 2009a;Wan et al.,~~
672 ~~2009b;Zhang et al., 2015a~~)(Fu et al., 2008a;Fu et al., 2008b;Fu et al., 2012;Wan et al.,
673 ~~2009a;Wan et al., 2009b;Zhang et al., 2015a~~), but is much lower than those in urban
674 areas of industrial cities such as Guiyang and Changchun where large point sources of
675 mercury exist (e.g., non-ferrous metal smelting, coal-fired power plants, and
676 residential coal burning) (~~Feng et al., 2004;Fu et al., 2011;Fang et al., 2004~~)(Feng et
677 ~~al., 2004;Fu et al., 2011;Fang et al., 2004~~). Although Hefei is geographically close to
678 Shanghai, a mega urban centre in China, it is interesting to note that the TGM
679 concentration of Shanghai is much lower than that of Hefei. This may be due to the
680 fact that Shanghai is a coastal city that is influenced more by cleaner marine air
681 masses (~~Friedli et al., 2011~~)(Friedli et al., 2011). Table 2 also shows that the average
682 concentration of GEM in Hefei is typically more than ~~fold or two folds fold~~ that in the
683 urban and rural areas in Europe and North America- (~~Liu et al., 2010;Li et al.,~~
684 ~~2008;Brooks et al., 2010~~),

685 686 3.2- Speciated atmospheric mercury during non-haze days

687 The frequency distribution of GEM, ~~RGMGOM~~ and PBM for the non-haze period
688 are shown in Fig. S1 (in blue). The mean concentration of GEM was 3.95 ± 1.93 ng

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689 m^{-3} . Its distribution was characterized by large fluctuations ranging from 0.2 to 23.8
690 ng m^{-3} , although more than half of the GEM values were in the narrow range 2-4 ng
691 m^{-3} . The mean concentration of **RGMGOM** was $2.49 \pm 2.41 \text{ pg m}^{-3}$ with a range of
692 0.5-33.5 pg m^{-3} , although most of the values were in the range of 1-4 pg m^{-3} . High
693 concentrations of **RGMGOM** (exceeding 10 pg m^{-3}) only accounted for 1.4% of the
694 total data. The mean **RGMGOM** concentration at the Hefei site is much smaller than
695 that reported from other study sites in China (Table 2), but is comparable to the values
696 observed from many European and North American sites ([Brooks et al., 2010](#);[Li et al.,](#)
697 [2008](#);[Liu et al., 2010](#);[Cheng et al., 2014](#))([Peterson et al., 2012](#);[Cheng et al., 2014](#);[Ren](#)
698 [et al., 2016](#)). The mean PBM concentraion at the Hefei site during the non-haze days
699 was $23.3 \pm 90.8 \text{ pg m}^{-3}$ with an exceptionally large range of 0.5-1827 pg m^{-3} . The
700 frequency distribution of PBM showed that high PBM concentrations (i.e., > 50 pg
701 m^{-3}) accounted for 6.4% of the total data. The PBM concentration under the non-haze
702 condition in Hefei is generally at a similar level to the remote areas, such as Mt.
703 Gongga, Mt. Waliguan and Shangri-[LiLa](#) in western China.

704 Diurnal variations of GEM, PBM and **RGMGOM** concentrations for non-haze
705 days are shown in Fig. 3. Both GEM and PBM concentrations exhibited similar
706 variations with elevated concentrations during night. The **RGMGOM** concentration
707 during the daytime was slightly higher than that in nighttime, typically peaking
708 between 10:00 and 12:00.

710 3.3 Speciated atmospheric mercury during haze days

711 Haze pollution mainly occurred in December and January at our monitoring site.
712 The four major haze pollution periods were identified in grey in Fig. 2. The mean

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713 concentrations of GEM, RGMGOM and PBM during these haze days were $4.74 \pm$
714 1.62 ng m^{-3} , $4.32 \pm 8.36 \text{ pg m}^{-3}$ and $60.2 \pm 131.4 \text{ pg m}^{-3}$, respectively (Table 1). The
715 frequency distributions of GEM, RGMGOM and PBM for the haze days are shown in
716 Fig. S1 (in gray). Comparison of GEM, RGMGOM and PBM concentrations during
717 haze and non-haze days is shown in Fig. 4.S2. GEM, RGMGOM and PBM
718 concentrations show significant differences between haze and non haze days
719 ($p < 0.001$, t-test). On average, the concentration of GEM in haze days was 1.2 times
720 that in non-haze days. Similarly, the concentration of RGMGOM in haze days was
721 about 1-1.7 times those in non-haze days. The largest impact of haze pollution is
722 however on PBM, with the mean PBM concentration in haze days about 2.5 times that
723 of non-haze days. High concentrations of RGMGOM (exceeding 10 pg m^{-3}) and PBM
724 concentrations (exceeding 50 pg m^{-3}) were also more frequently observed than in
725 non-haze days, accounting for 5.9% and 25%, respectively, of the total haze days.

726 Diurnal variations of GEM, PBM and RGMGOM concentrations for haze days
727 are shown in Fig. 3. GEM concentrations were higher during night, decreased during
728 daytime. The opposite pattern was observed for RGMGOM, which showed higher
729 concentrations during daytime than during night; the duration of the RGMGOM peak
730 also lasted longer for haze days. On the contrary, the PBM typically peaked just
731 before sunrise, with the lowest values occurred in the afternoon (14:00-16:00).

733 4. Discussion

734 4.1 Influence of atmospheric mercury emission source

735 ~~The statistically significant difference in the GEM concentration between~~
736 ~~non-haze days and haze days suggests that haze pollution could directly affect the~~
737 ~~concentration of elemental mercury.~~ In order to understand the mercury sources

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738 attribution ~~during haze days~~, the PSCF model analysis was conducted by using the
739 TrajStat software. ~~As shown in Fig. 5, the area south to the monitoring site and the~~
740 ~~neighboring provinces were the main sources region during haze days. Thus,~~
741 ~~atmospheric mercury in haze days were mainly affected by local or regional emission~~
742 ~~sources.~~

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743 The seasonal sources could ~~also~~ be inferred from the PSCF analysis with the
744 year-round data. Fig. ~~6(A)4a~~ showed the overall spatial contribution of mercury
745 emission sources in China. As Hefei is located in east-central China, its atmospheric
746 mercury concentration could be affected by both north and south emission sources,
747 including those from the North China Plain (especially Shandong Province) and the
748 neighboring provinces of Henan, Jiangsu, Jiangxi and Hubei. The total mercury
749 emissions from Henan and Shandong provinces were estimated to be over 50 and 45
750 tons in 2010, respectively, making them two largest Hg emitters in China (Zhang et al.,
751 2015b). Long-range transport could also impact the seasonal variations of
752 atmospheric mercury in Hefei. As shown in Figure ~~64~~, in spring, the major
753 contributors of atmospheric mercury to Hefei were from the southwestern region
754 including the local area and the Jiangxi and Hunan provinces. In summer, the main
755 contributors were from north of Anhui, as well as Henan and Jiangxi provinces, and
756 even from the Pearl River Delta region in the far south. ~~In autumn and winter, the~~
757 ~~prevalent seasons for haze pollution, the most important anthropogenic mercury~~
758 ~~sources to the monitoring site were the local emissions and those from the~~
759 ~~neighboring region of Shandong, Henan and the Yangtze River Delta region. The total~~
760 ~~mercury emissions from Henan and Shandong provinces were estimated to be over 50~~
761 ~~and 45 tons in 2010, respectively, making them two largest Hg emitters in China~~
762 ~~(Zhang et al., 2015b).~~ Since the number of haze days accounts only for 5.6% of the
763 total days in spring and summer, we did not provide haze and non-haze PSCF results

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764 for spring and summer seasons. As autumn and winter are the prevalent seasons for
765 haze pollution, one PSCF result for haze days and another for non-haze days are
766 shown for autumn and winter, respectively. The statistically significant difference
767 ($p < 0.001$) in the GEM concentration between non-haze days and haze days suggests
768 that haze pollution could directly affect the concentration of elemental mercury. As
769 shown in Figs. 4d and 4f, higher GEM concentration was mainly influenced by local
770 emission sources during haze days. For non-haze days, the most important mercury
771 sources to the monitoring site were not only the local emission sources, but also those
772 from the neighboring region of Shandong, Henan and Jiangxi provinces (see Figs. 4e
773 and 4g). In summary, the increase of GEM concentration during haze days was
774 mainly caused by local emission.

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775 ~~GEM and CO normally share anthropogenic emission sources, such as industrial~~
776 ~~and domestic coal combustion (Wu et al., 2006). However, they also have their own~~
777 ~~sources, vehicles are another kind of dominant sources for CO, while power plants are~~
778 ~~another type of mainly sources for GEM. The correlation between the concentrations~~
779 ~~of GEM and CO during non-haze and haze days is shown in Fig. S2. The slope of the~~
780 ~~trend line represents the Hg/CO ratio. Emissions from power plants typically have a~~
781 ~~higher Hg/CO ratio (Wu et al., 2006), whereas biomass burning and residential coal~~
782 ~~combustion have a lower Hg/CO ratio due to incomplete combustion (Weiss Penzias~~
783 ~~et al., 2007). The Hg/CO ratios from our study for both non-haze and haze days are in~~
784 ~~the range of $0.0003\text{--}0.0009\text{ ng m}^{-3}\text{--ppbv}^{-1}$, similar to the ratio reported for Alaska~~
785 ~~biomass burning ($0.0014 \pm 0.0006\text{ ng m}^{-3}\text{--ppbv}^{-1}$; Weiss Penzias et al., 2007),~~
786 ~~indicating that biomass burning might have played an important role in mercury~~
787 ~~emission in Hefei. This is further supported by the concentration of water soluble~~

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788 ~~potassium (K^+) in PM_{10} . K^+ is a typical component of biomass burning aerosol and~~
789 ~~has been used as a tracer element for qualitative identification of biomass burning~~
790 ~~(Cachier et al., 1991). As shown in Fig. S3, K^+ in PM_{10} shows a good correlation with~~
791 ~~air quality index (AQI) during the heavy pollution period of Nov-Dec, 2013. In~~
792 ~~addition, seven high GEM events were identified during the whole monitoring period~~
793 ~~(Table S1). 5-day backward trajectories for each GEM heavy pollution event for the~~
794 ~~time of at maximum GEM concentration are shown in Fig. S4. Air masses with~~
795 ~~elevated GEM concentration were mainly from NW, SW and East directions. In~~
796 ~~combination with the NASA's satellite hotspots/fire locations information from the~~
797 ~~Fire Information for Resource Management System (FIRMS), there were potential~~
798 ~~biomass burning occurred in these regions when air masses passed over (Fig. S4,~~
799 ~~Events 1-7). Therefore, biomass burning can contribute to the observed higher mercury~~
800 ~~concentrations, which not only came from local sources (Events 1 and 4), but could also~~
801 ~~be affect by other regions through long range transport processes (Events 2, 3, 6 and 7).~~

802 GEM and CO often share anthropogenic emission sources, such as industrial coal
803 combustion, domestic coal combustion, iron and steel production and cement
804 production (Wu et al., 2006; Wang et al., 2005). However, they also have their own
805 sources. For instance, power plants and nonferrous metal smelters emit mercury but
806 hardly any CO, while most of CO originates from vehicles which are not a major
807 emitter for mercury. The correlation coefficients and slopes between GEM
808 concentration and CO concentration during pollution events are shown in Table 3.
809 These mercury pollution episodes were defined as a period with hourly average GEM
810 concentration higher than seasonal average GEM concentration and the duration of
811 elevated hourly GEM concentration lasted for over 10 hours. The Hg/CO slope and

812 correlation between GEM and CO concentrations has been used to identify long-range
813 transport episodes and local episodes in previous research (Jaffe et al.,
814 2005;Weiss-Penzias et al., 2006;Kim et al., 2009). These episodes could classified
815 into long-range transport episodes or local episode by using the correlation
816 coefficients (R^2) of linear regression: significant positive correlation for long-range
817 transport episodes and poor correlation for local episodes (Kim et al., 2009). Six
818 episodes (events: 1-6) were found to have poor correlations between GEM and CO
819 concentrations (R^2 : 0.01-0.29) for local episodes, while four episodes (events: 7-10)
820 were found to have positive correlations between GEM and CO concentrations (R^2 :
821 0.51-0.79) for long-range transport episodes. These local episodes tend to occur in
822 autumn and winter. The slope of the trend line represents the Hg/CO ratio. Emissions
823 from power plants typically have a higher Hg/CO ratio (Wu et al., 2006), whereas
824 residential coal and biomass burning combustion have a lower Hg/CO ratio
825 (0.0013-0.0046 ng m⁻³ ppbv⁻¹) due to incomplete combustion (Weiss-Penzias et al.,
826 2007). The Hg/CO ratio for vehicles is nearly zero (Zhang et al., 2013). The Hg/CO
827 ratios from our study for pollution episodes are in the range of 0.0001-0.0050 ng m⁻³
828 ppbv⁻¹. In summary, the low GEM/CO ratio in Hefei might related to the local
829 incomplete combustion, like residential coal and biomass burning combustion.

830

831 **4.2 Impacts of meteorological factors for atmospheric mercury during haze days**

832 Meteorological ~~condition~~conditions, especially wind direction and speed, could
833 also impact the atmospheric mercury during haze days. The wind rose for the
834 monitoring site during the study period is shown in Fig. 75. Easterly and southeasterly
835 winds represented the prevailing wind directions at the study site. A wind rose

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836 diagram of GEM concentrations above the 90th percentile value is shown in Fig.
837 ~~7B5B~~. We found that 67% of the high GEM concentrations occurred at low wind
838 speed (below 1.5 m s⁻¹); however, wind speed below 1.5 m s⁻¹ accounted for only 1.7%
839 of total study. High ~~RGMGOM~~ and PBM concentrations appear not to be related to
840 high wind speed (wind speed: 3-5 m s⁻¹); only 1.4% and 2.6% of the high ~~RGMGOM~~
841 and PBM concentrations were observed under high wind-speed conditions,
842 respectively (Figs. ~~7C5C~~ and ~~7D5D~~). In general, most of the high atmospheric
843 mercury levels occurred in the low wind speed conditions. This slow wind speed
844 condition is not conducive to the spread of mercury and thus favours the
845 accumulation of atmospheric mercury, especially during haze days. This result is
846 further support that atmospheric mercury during haze days is mainly affected by local
847 emissions.

848 Both GEM and PBM concentrations exhibited great variations with elevated
849 concentration during night or early morning, regardless of the presence of haze. Such a
850 diurnal variation of GEM and PBM could be related to changes in the height of urban
851 boundary layer, which is typically low in the morning and night, and high during the
852 daytime (~~Yuan et al., 2005~~(Yuan et al., 2005;Mao et al., 2006). The maximum PBM
853 concentration (observed at 6:00) was more than 4 times higher than the minimum value
854 (observed at 16:00) both under non-haze and haze days, and about 76% PBM were
855 ~~removed~~declined during this period (6:00-16:00). However, the reductions of PBM as a
856 result of deposition during haze days was 62.7 pg m⁻³, which was about 2.4 times that in
857 non-haze days, suggesting that haze pollution could increase the removal of PBM and
858 thus reduce its atmospheric lifetime. Although PBM is not the major form of mercury
859 emitted to the atmosphere, it is crucial in atmospheric mercury transport and removal

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860 processes due to its short atmospheric lifetime. As shown in Fig. 6, highest PBM and
861 PM_{2.5} concentrations were observed in January, which most likely due to shallower
862 boundary layer in January than in other months. The co-variation in February is weaker,
863 possibly due to the loss of PBM data because of instrument maintenance (see Section
864 2.3). ~~8, the~~ The PBM concentration co-varied with the PM_{2.5} concentration, especially
865 in January when all the four PBM peak events were associated with increased PM_{2.5}
866 concentrations. (Fig. 6c). This result indicates that the PM_{2.5} concentration may play an
867 important role in the formation of PBM. Thus, elevated ~~The co-variation in February is~~
868 ~~weaker, possibly due to the loss of PBM data because of instrument maintenance (see~~
869 ~~Section 2.3).~~ Elevated PBM concentrations might be due to the poor diffusion
870 conditions in cold months and high PM pollution. ~~Although the concentrations of PM_{2.5}~~
871 ~~were similar from March to June, March showed higher PBM concentrations. This~~
872 ~~might be due to higher temperatures in the warmer months which do not favor mercury~~
873 ~~adsorption (Otani et al., 1986). These results indicate that both the PM_{2.5} concentration~~
874 ~~and temperature may play an important role in the formation of PBM.~~

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876 **4.3 ~~Chemical process for RGM~~Enhancements in GOM and the potential**

877 **oxidation mechanism**

878 ~~In contrast with the diurnal variations of GEM and PBM, RGM shows different~~
879 ~~diurnal trend. The RGM concentration during daytime was slightly higher than at~~
880 ~~night. As discussed before in 4.2 section, the diurnal variation of GEM and PBM~~
881 ~~could be related to changes in the height of urban boundary layer. So the role of~~
882 ~~boundary layer for the enhancement of RGM during daytime would be limited. The~~
883 ~~weak correlation ($r=0.164$, $p<0.001$) between RGM and CO suggests that regional~~
884 ~~anthropogenic sources are not a major source of RGM in the air. As shown in Fig. 9~~
885 ~~(haze days), the peak value of RGM coincided well with the lowest value of CO,~~
886 ~~suggesting that the production of RGM in haze days does not fully come from~~

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887 anthropogenic emission sources. Instead, RGM is more likely produced from the
888 situ oxidation of GEM; this is supported by the fact that the peak values of RGM
889 coincided well with the decline of GEM. We found that the RGM concentration did
890 not change much from 8:00 to 20:00 (local time) during haze days. This indicates that
891 photochemical reaction between GEM and RGM still takes place during haze days,
892 but the daytime change in the intensity of solar radiation has been greatly dampened
893 due to the light haze interaction.

894 Various atmospheric oxidants are capable of oxidizing GEM to RGM, including
895 halogen radicals, ozone, hydroxyl radicals (OH), among others (Holmes et al.,
896 2010; Wang et al., 2014). Halogen radicals, especially bromine atoms, are believed to
897 be the primary oxidant for GEM in the global troposphere (Holmes et al., 2010).
898 Unfortunately, we did not measure halogen radicals in this study. Ozone itself is not
899 an efficient oxidant for GEM oxidation due to low reaction rate (Hall, 1995; Holmes et
900 al., 2010). In the lower troposphere, ozone is produced from daytime photochemical
901 reactions involving volatile organic compounds (VOC) and nitrogen oxides (NO_x).
902 Due to fresh emissions of NO from vehicles can react with O₃ to form NO₂, so ozone
903 could not fully represent photochemical oxidation processes in urban area (Herndon et
904 al., 2008). Compared with ozone, odd oxygen (O_x = O₃ + NO₂) is a more conserved
905 tracer of the extent of photochemical processing in the urban atmosphere (Herndon et
906 al., 2008; Wood et al., 2010). Because of such NO₂ concentrations from MAX DOAS
907 are available only during the daytime, so we could only use O_x to be a indicator for
908 GEM oxidation occurred in the daytime. Diurnal variations of GEM, RGM and
909 Gas Phase Data (O_x, O₃, NO₂ and CO) concentrations during non-haze and haze days
910 are shown in Fig. 9. The increase of O_x is consistent with the increase of RGM during
911 haze days, but seems to lag behind the increase of RGM during non-haze days,
912 suggesting some oxidation processes might be at work during haze days. During haze
913 days, both the RGM and O_x reached highest values around 16:00, along with the
914 lowest value of GEM, indicating the chemical transformation between GEM and
915 RGM occurred. However, this phenomenon is not found in non-haze days.

916 The OH radical is also an important oxidant for mercury in the atmosphere.
917 Previous studies have shown that the major source of OH in the early morning is the
918 photolysis of HONO, which accumulates in the urban atmospheric boundary layer
919 during night (Kleffmann et al., 2005). The formation of HgOH as an intermediate
920 product of the Hg⁰(g) + OH oxidation reactions has been proposed by (Sommar et al.,
921 2001), although HgOH is highly unstable and could decompose back rapidly to Hg⁰
922 and OH (Goodsite et al., 2004). It has been proposed that secondary reactants such as

923 ~~NO₂, HO₂, RO, RO₂, and NO could assist the formation of Hg(II) from the initial~~
924 ~~HgOH intermediate, which outcompetes with the decomposition of HgOH (Calvert~~
925 ~~and Lindberg, 2005). As an example, we calculated the transformation between GEM~~
926 ~~and RGM under the influence of NO₂, using the reactions and rate constants shown in~~
927 ~~Table S2. As shown in Fig. S5, the production rate of NO₂HgOH, d[NO₂HgOH]/dt,~~
928 ~~increased almost linearly with increasing NO₂ under low NO₂ concentrations, and~~
929 ~~eventually reached a steady state when the NO₂ concentration is high enough.~~

930 Diurnal variations of GEM, GOM, O₃ and CO concentrations during non-haze
931 and haze days are shown in Fig. 7. The weak correlation (r=0.164, p<0.001) between
932 GOM and CO suggests that the CO-producing, primary emission is not a major source
933 of GOM in the air. This is clearly shown in Fig. 7 (haze days), where the peak value
934 of GOM coincided with the lowest value of CO. GOM also shows a diurnal trend that
935 is opposite to that of GEM and PBM; for both non-haze and haze days, GOM
936 concentrations remained relatively constant during night, but increased rapidly prior
937 to sunrise. Two processes can affect the GOM concentrations in the boundary layer air.
938 The first is due to transport of GOM from the free troposphere (FT). It is well
939 established that FT contains higher GOM concentrations than in the boundary
940 layer(e.g., (Murphy et al., 2006;Lyman and Jaffe, 2012;Timonen et al., 2013;Brooks
941 et al., 2014;Shah et al., 2016)). It is thus possible that higher GOM concentrations
942 observed prior to sunrise is due to admixing from the free troposphere as the height of
943 the boundary layer increases during the morning. In addition, in situ photochemical
944 oxidation of GEM could also increase the concentration of GOM during daytime.
945 Various atmospheric oxidants are capable of oxidizing GEM to GOM, including
946 halogen radicals, ozone, hydroxyl radicals (OH), among others (Holmes et al.,
947 2010;Wang et al., 2014).

948 To determine the relative importance of FT transport and in situ photochemical
949 oxidation, we examined the relationship between GOM and the changes in the height

950 of the atmospheric boundary layer and the odd oxygen ($O_x = O_3 + NO_2$) concentrations.
951 We used O_x because it is a more conserved tracer of the extent of photochemical
952 processes in the urban atmosphere (Herndon et al., 2008; Wood et al., 2010), as O_3
953 reacts with NO emitted from automobiles to form NO_2 . Example results are shown in
954 Fig. 8 for 20th November, 2013. As can be seen from the figure, both GOM and O_x
955 reached higher concentrations from 12:00 to 16:00, along with the lowest value of
956 GEM. The height of the atmospheric boundary layer changed very little (less than 0.1
957 km) over the same period (see Fig. S3). This simple comparison suggests that the
958 transport of FT GOM might be limited and that at least some of the GOM were
959 formed from in situ oxidation of GEM. Note that in our studies we could only
960 calculate daytime O_x concentrations, because NO_2 concentrations from MAX-DOAS
961 were only available during daytime.

962 We further investigated the mechanism of the GEM oxidation to GOM. Ozone
963 itself is not an efficient oxidant for GEM oxidation due to the low reaction rate (Hall,
964 1995; Holmes et al., 2010). Instead, halogen radicals (especially bromine atoms) and
965 halogen radicals, are believed to be the primary oxidants for GEM in the global
966 troposphere (Holmes et al., 2010). Unfortunately, we did not measure halogen radicals
967 in this study. OH radicals are known to be present in the early morning urban
968 boundary layer, primarily from the photolysis of HONO, which accumulates during
969 night (Kleffmann et al., 2005). Therefore, here we consider the oxidation of GEM by
970 OH radicals. The formation of HgOH as an intermediate product of the $Hg^0(g) + OH$
971 oxidation reactions has been proposed by Sommar et al., 2001, although HgOH is
972 highly unstable and could decompose back rapidly to Hg^0 and OH (Sommar et al.,
973 2001; Goodsite et al., 2004). It has been proposed that the presence of other gases X

974 (X = NO₂, HO₂, RO, RO₂, or NO) could assist the formation of Hg(II) by forming
975 X-HgOH, which outcompetes the decomposition of HgOH (Calvert and Lindberg,
976 2005;Dibble et al., 2012;Wang et al., 2014). As an example, we calculated the
977 transformation between GEM and GOM under the influence of NO₂, using the
978 reactions and rate constants shown in Table S1. As shown in Fig. S4, the production
979 rate of NO₂HgOH, d[NO₂HgOH]/dt, increased almost linearly with increasing NO₂
980 under low NO₂ concentrations, and eventually reached a steady state when the NO₂
981 concentration is high enough.

982 Based on the production rate of NO₂HgOH, we can estimate the production of
983 NO₂HgOH during the 1 hr sampling period when RGMGOM_A was captured by the
984 KCl-coated denuder in the Tekran 1130 unit. ~~As discussed earlier, a distinct diurnal~~
985 ~~relationship between GEM and RGM was observed both in non-haze and haze days~~
986 ~~(Fig. 9).~~ The production of NO₂HgOH and d[NO₂HgOH]/dt corresponding to
987 different NO₂ concentrations is shown in Table 3.4. With the increase of the NO₂
988 concentration, the contribution of the NO₂HgOH production to RGMGOM_A will
989 increase. If the NO₂ concentration is within 100 ppbv (from 0 to 100 ppbv), the
990 production of NO₂HgOH would be in range of 0.058-4.81 pg m⁻³ during the 1h
991 sampling period. ~~The mean NO₂ concentration during haze days is 24.3 ppbv (unit~~
992 ~~conversion see the supporting information), which is higher than in non-haze days~~
993 ~~(17.4 ppbv). The increments of RGM from sunrise to peak (6:00-12:00) are about~~
994 ~~2.14 pg m⁻³ (growth rate ~ 0.36 pg m⁻³ h⁻¹) and 0.41 pg m⁻³ (growth rate ~ 0.068 pg~~
995 ~~m⁻³ h⁻¹) during haze days and non-haze days, respectively. As illustrated in Table 3.4,~~
996 ~~the level of NO₂ observed in the urban atmosphere our study is high enough for to~~
997 ~~make increase, the GOM production of RGM during the 1h sampling period. In~~

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998 ~~addition, we found that the NO₂ concentrations increased rapidly after sunrise to reach~~
999 ~~peak values around noon, consistent with the increases of RGM during haze days. We~~
1000 ~~Our results, thus postulatesupport a recent study in the tropical equatorial Pacific~~
1001 ~~(Wang et al., 2014), that NO₂ aggregation with HgOH may beprovides provides a~~
1002 ~~possible mechanism to explain the enhanced production of RGM during the daytime~~
1003 ~~over the inland urban air. The NO₂ level in urban air GOM and the role of NO₂ might~~
1004 ~~have a be more important influence on the chemical transformations between the~~
1005 ~~GEM and RGM during haze days, rather than non haze days. But unfortunately, we~~
1006 ~~can not provide an adequate description about these oxidation processes. To solve this~~
1007 ~~problem, newin the urban air. More, laboratory and merecury modelmodeling studies~~
1008 ~~on mercury oxidation mechanism are neededin the presence of NO₂ and other gases~~
1009 ~~are thus warranted.~~

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1011 **5. Summary**

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1012 Continuous measurements of speciated atmospheric mercury were conducted at
1013 Hefei, a mid-latitude inland city in central China, from July 2013 to June 2014.
1014 Measurements of other trace gases (e.g. CO, O₃, NO₂) and meteorological parameters
1015 were employed to better understand the sources and oxidation pathways of
1016 atmospheric mercury. The mean GEM, ~~RGMGOM~~, and PBM concentrations during
1017 haze days were $4.74 \pm 1.62 \text{ ng m}^{-3}$, $4.32 \pm 8.36 \text{ pg m}^{-3}$ and $60.2 \pm 131.4 \text{ pg m}^{-3}$,
1018 respectively. Potential source contribution function (PSCF) analysis suggested that the
1019 local mercury emission rather than long-range transport is the most important
1020 contributor of atmospheric mercury pollution during haze days at our monitoring site.
1021 ~~The Hg/CO ratio and NASA's satellite fire locations information indicated that the~~

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1022 biomass burning may plays an important role in mercury emission. Haze pollution has
1023 considerable impact on PBM rather than on GEM and RGM. Both GEM and PBM
1024 concentrations exhibited greatly variations with elevated concentration during night.
1025 The diurnal variations of GEM and PBM might be related to the boundary layer depth;
1026 a lower boundary layer depth in the morning and night could elevate the mercury
1027 concentration. The slow wind speed condition is not conducive to the spread of
1028 mercury and thus favours the accumulation of atmospheric mercury, especially during
1029 haze days. We found that PBM concentrations co-varied with the PM_{2.5} concentration
1030 especially in January when all the four PBM peak events were associated with
1031 increased PM_{2.5} concentrations. In addition, The low GEM/CO ratio in Hefei could be
1032 a result of local incomplete combustion sources such as residential coal and biomass
1033 burning. Haze pollution has a more profound impact on PBM than on GEM and GOM,
1034 PBM showed a remarkable seasonal pattern, with higher concentrations in cold
1035 seasons and lower in warm seasons. Elevated PBM concentrations might be due to
1036 both the high loadings of particle matter and disadvantageous diffusion conditions
1037 during haze days especially in cold months. ~~The peaks of RGM were observed around~~
1038 ~~noon, which is probably due to the higher intensity of solar radiation.~~ Both GEM and
1039 photochemical oxidation processes at this time. Change in the odd oxygen ($O_x =$
1040 $O_3 + NO_2$) PBM concentrations exhibited great variations with elevated concentration is
1041 ~~normally applied to be an indicator for photochemical reaction. The increase of O_x is~~
1042 ~~consistent with the increase of RGM during haze days, but seems to lag behind the~~
1043 ~~increase of RGM during non-haze days, suggesting some oxidation processes.~~ night.
1044 The diurnal variations of GEM and PBM might be at work related to the boundary

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1045 layer depth; a lower boundary layer depth in the morning and night could elevate the
1046 mercury concentration.

1047 Unlike with the diurnal variations of GEM and PBM, GOM concentration
1048 remained relatively constant during night, and then increased rapidly prior to the
1049 sunrise. The enhancement of GOM during haze days daytime could be due to both the
1050 transport of GOM-enriched free troposphere air to the boundary layer and in situ
1051 oxidation of GEM in the boundary layer. Simple photochemical modeling supports
1052 the occurrence of daytime oxidation of GEM to GOM. Based on HgOH as an
1053 intermediate product, ~~we~~our calculations suggest that NO₂ aggregation with HgOH is
1054 a potential mechanism to explain the enhanced production of RGM during the
1055 daytime over the inland urban air. The NO₂ level in urban air might have a more
1056 important influence on the chemical transformations between the GEM and RGM
1057 during haze days, rather than non haze days GOM over the inland urban air.

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Acknowledgements

This research was supported by grants from the National Basic Research Program of
China (2013CB430000), the National Natural Science Foundation of China (Project
Nos. 91544103,41575021) and the External Cooperation Program of BIC, CAS
(Project No.211134KYSB20130012).

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1287 | Zhang, L., Wang, S., Wang, L., Wu, Y., Duan, L., Wu, Q., Wang, F., Yang, M., Yang, H.,
1288 | and Hao, J.: Updated Emission Inventories for Speciated Atmospheric Mercury
1289 | from Anthropogenic Sources in China, Environmental science & technology, 49,
1290 | 3185-3194, 2015b.

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1294 | **Table 1. Summary of GEM, RGMGOM and PBM concentrations measured in**
 1295 **Hefei from July 2013 to June 2014.**
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	GEM (ng m ⁻³)			<u>RGMGOM</u> (pg m ⁻³)			PBM (pg m ⁻³)		
	Mean ± σ	Range	N	Mean ± σ	Range	N	Mean ± σ	Range	N
Spring	3.89 ± 1.79	0.2-21.3	7890	4.49 ± 4.22	0.5-69.8	526	8.34 ± 8.97	1.6-130.1	542
Summer	4.08 ± 1.99	0.3-22.9	6050	3.66 ± 4.39	0.5-45.2	511	3.61 ± 4.38	0.5-41.9	570
Autumn	4.51 ± 2.10	0.4-23.8	3632	5.65 ± 8.93	0.5-78.9	274	59.9 ± 153.5	0.5-1615	339
Winter	4.05 ± 1.81	0.9-12.2	6381	2.59 ± 2.58	0.5-9.5	541	56.1 ± 134.9	0.5-1827	639
Total	4.07 ± 1.91	0.2-23.8	23953	3.67 ± 5.11	0.5-78.9	1852	30.02 ± 100.3	0.5-1827	2090
Non-haze	3.95 ± 1.93	0.2-23.8	20345	2.49 ± 2.41	0.5-33.5	1508	23.3 ± 90.76	0.5-1827	1708
Haze	4.74 ± 1.62	2.1-16.5	3608	4.32 ± 8.36	0.5-78.9	344	60.2 ± 131.4	1.6-1615	382

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1299 **Table 2. Speciated atmospheric mercury concentrations in Hefei and other urban**
 1300 **and rural areas.**

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Location	Classification	Time	TGM (ng m ⁻³)	GEM (ng m ⁻³)	RGM <u>GOM</u> (pg m ⁻³)	PBM (pg m ⁻³)	Reference
Hefei	Suburb	Jul 2013-Jun 2014	4.1	4.07	3.67	30	This study
Hefei	Suburb	Feb-May 2009	2.53	-	-	-	Hu et al. (2014)
Beijing	Rural	Dec 2008-Nov 2009	3.23	3.22	10.1	98.2	Zhang et al. (2013)
Shanghai	Urban	Aug-Sep 2009	2.7	-	-	-	Friedli et al. (2011)
Nanjing	Urban	Jan-Dec 2011	7.9	-	-	-	Zhu et al. (2012)
Guiyang	Urban	Nov 2001-Nov 2002	8.4	-	-	-	Feng et al. (2004)
Guiyang	Urban	Aug-Dec 2009	-	9.72	35.7	368	Fu et al. (2011)
Changchun	Urban	Jul 1999-Jan 2000	18.4	-	-	276	Fang et al. (2004)
Changchun	Suburb	Jul 1999-Jan 2000	11.7	-	-	109	Fang et al. (2004)
Mt.Changbai	Remote	Aug 2005-Jul 2006	3.58	-	65	77	Wan et al. (2009a, b)
Mt.Gongga	Remote	May 2005-July 2006	3.98	-	6.2	30.7	Fu et al. (2008a, b)
Mt.Waliguan	Remote	Sep 2007-Aug 2008	1.98	-	7.4	19.4	Fu et al. (2012a)
Mt.Leigong	Remote	May 2008-May 2009	2.8	-	-	-	Fu et al. (2010)
Shangri-La	Remote	Nov 2009-Nov 2010	2.55	-	8.22	38.82	Zhang (2015)
Detroit, USA	Urban	Jan-Dec 2004	-	2.5	15.5	18.1	Liu et al. (2010)
Dexter, USA	Rural	Jan-Dec 2004	-	1.6	3.8	6.1	Liu et al. (2010)
Houston, USA	Urban	Aug-Oct 2006	-	1.66	6.9	2.5	Brooks et al. (2010)
<u>Florida, USA</u>	<u>Urban</u>	<u>Jul 2009-Jul 2010</u>		<u>1.3</u>	<u>3</u>	<u>2</u>	<u>Peterson et al. (2012)</u>
<u>Maryland, USA</u>	<u>suburb</u>	<u>2007-2015</u>		<u>1.41</u>	<u>4.6</u>	<u>8.6</u>	<u>Ren et al. (2016)</u>
Göteborg, Sweden	Urban	Feb-Mar 2005	-	1.96	2.53	12.5	Li et al. (2008)
Nova Scotia, Canada	Urban	Jan 2010- Dec 2011		1.67	2.07	2.32	Cheng et al. (2014)
Northern Hemisphere background value					1.5-1.7		Lindberg et al. (2007)

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Table 3.
Table 3. Correlation coefficients and slopes between GEM concentration and CO
concentration during atmospheric mercury pollution events.

<u>Event</u>	<u>Start Time</u> (UTC + 8 hr)	<u>End Time</u> (UTC + 8 hr)	<u>Duration</u> (h)	<u>GEM</u> (ng m ⁻³)	<u>CO</u> (ppbv)	<u>GEM/CO</u> (slope, ng m ⁻³ ppbv ⁻¹)	<u>R²</u>
<u>1</u>	<u>2013/11/21 03:00</u>	<u>2013/11/22 02:00</u>	<u>23</u>	<u>8.37±2.42</u>	<u>4481.6±717.3</u>	<u>0.0018</u>	<u>0.29</u>
<u>2</u>	<u>2013/12/03 20:00</u>	<u>2013/12/04 09:00</u>	<u>13</u>	<u>7.51±0.67</u>	<u>5270.0±744.5</u>	<u>0.0001</u>	<u>0.02</u>
<u>3</u>	<u>2013/12/07 04:00</u>	<u>2013/12/09 04:00</u>	<u>48</u>	<u>9.21±1.16</u>	<u>5943.8±1394.1</u>	<u>0.0004</u>	<u>0.23</u>
<u>4</u>	<u>2013/12/19 09:00</u>	<u>2013/12/20 09:00</u>	<u>24</u>	<u>4.35±0.17</u>	<u>3907.6±353.0</u>	<u>0.0002</u>	<u>0.03</u>
<u>5</u>	<u>2013/12/24 19:00</u>	<u>2013/12/25 15:00</u>	<u>20</u>	<u>5.58±0.94</u>	<u>4930.8±919.7</u>	<u>0.0012</u>	<u>0.01</u>
<u>6</u>	<u>2014/01/17 22:00</u>	<u>2014/01/19 13:00</u>	<u>39</u>	<u>5.80±0.83</u>	<u>5746.3±1626.9</u>	<u>0.0003</u>	<u>0.28</u>
<u>7</u>	<u>2014/01/25 02:00</u>	<u>2014/01/25 22:00</u>	<u>20</u>	<u>6.03±0.50</u>	<u>8797.9±2244.3</u>	<u>0.0002</u>	<u>0.59</u>
<u>8</u>	<u>2014/03/16 05:00</u>	<u>2014/03/16 20:00</u>	<u>15</u>	<u>4.46±0.47</u>	<u>2261.7±440.2</u>	<u>0.0010</u>	<u>0.79</u>
<u>9</u>	<u>2014/03/17 06:00</u>	<u>2014/03/18 12:00</u>	<u>30</u>	<u>8.85±2.46</u>	<u>2697.1±590.3</u>	<u>0.0030</u>	<u>0.51</u>
<u>10</u>	<u>2014/05/21 00:00</u>	<u>2014/05/21 11:00</u>	<u>11</u>	<u>5.74±0.94</u>	<u>3676.7±1690.0</u>	<u>0.0050</u>	<u>0.79</u>

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Notes: these episodes were identified using the following criteria: (a) the duration of
elevated GEM concentration lasted for >10h; (b) the selected hourly average GEM
concentration higher than the seasonal average GEM concentration.

1312 | **Table 4.** The production of NO₂HgOH and d[NO₂HgOH]/dt at different NO₂
1313 concentrations

NO ₂ (ppbv)	10	20	30	40	50	60	70	80	90	100
d(NO ₂ HgOH)/dt (molecule cm ⁻³ s ⁻¹)	0.36	0.71	1.04	1.37	1.68	1.99	2.28	2.56	2.83	3.10
NO ₂ HgOH (pg m ⁻³ , 1hr)	0.56	1.10	1.63	2.13	2.61	3.08	3.54	3.97	4.40	4.81

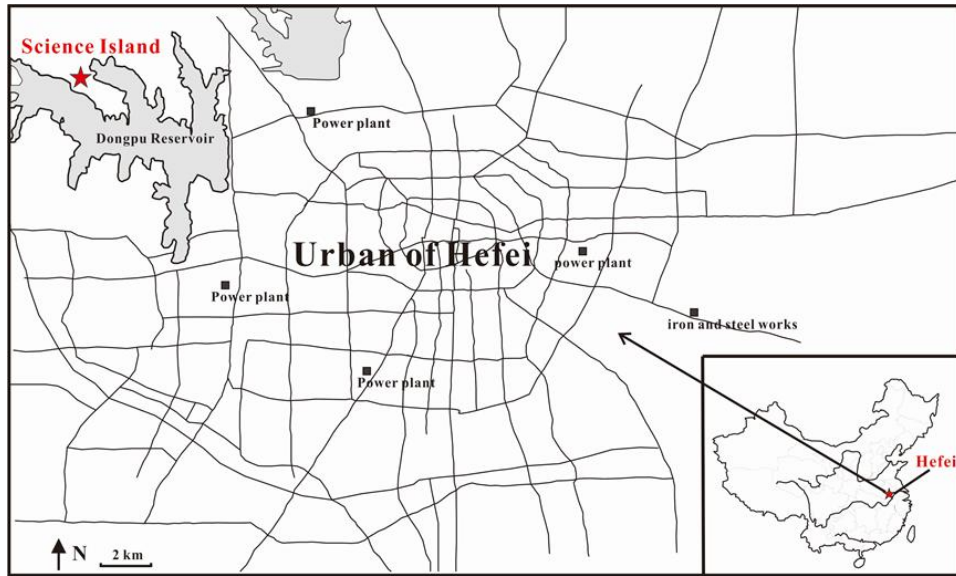
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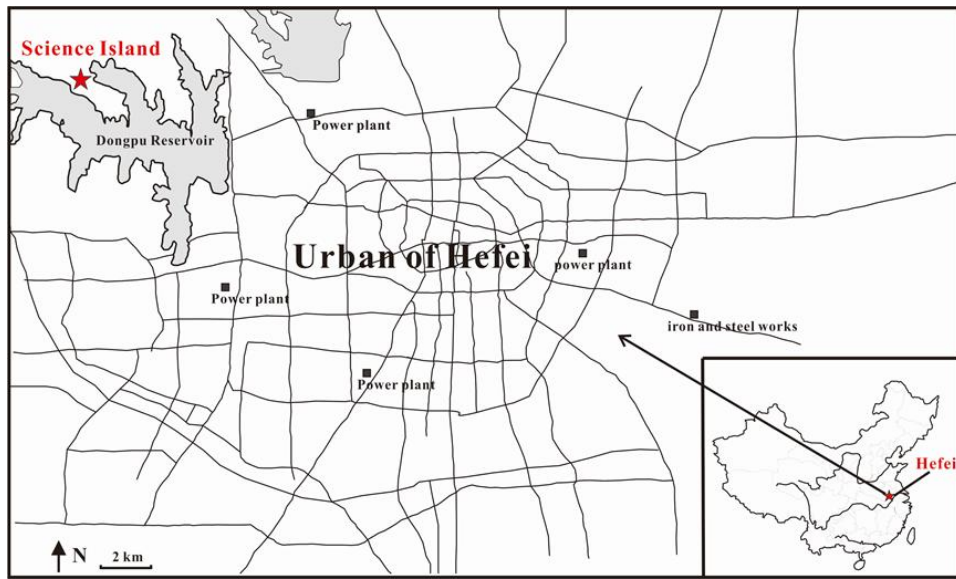
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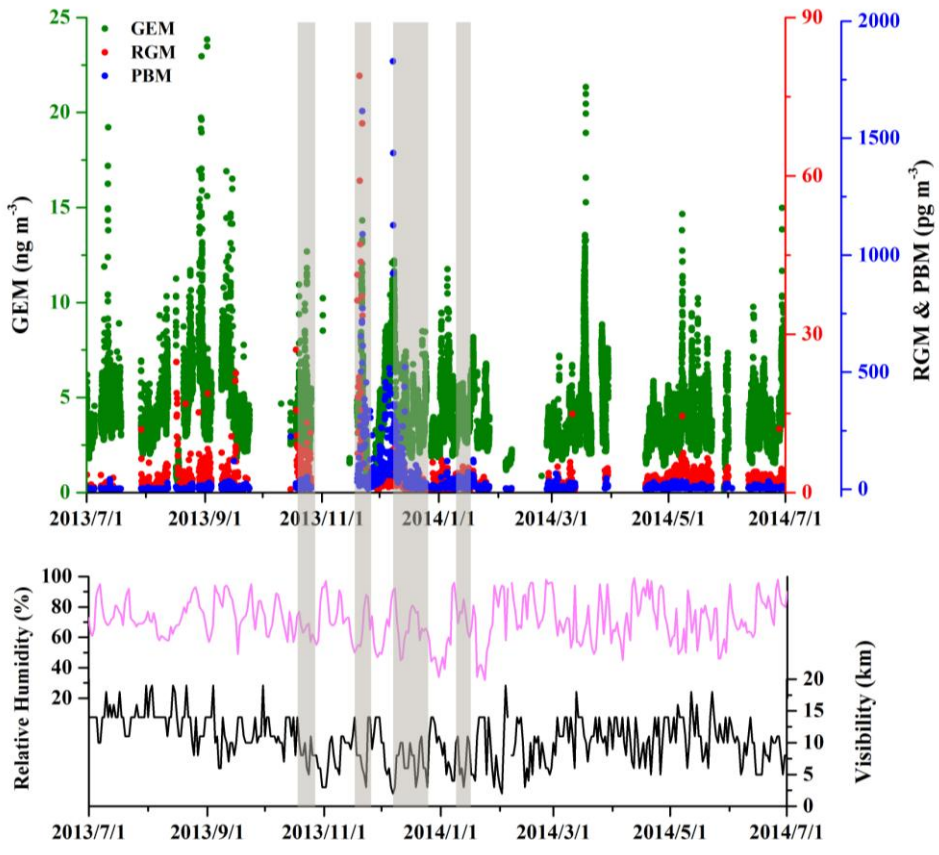
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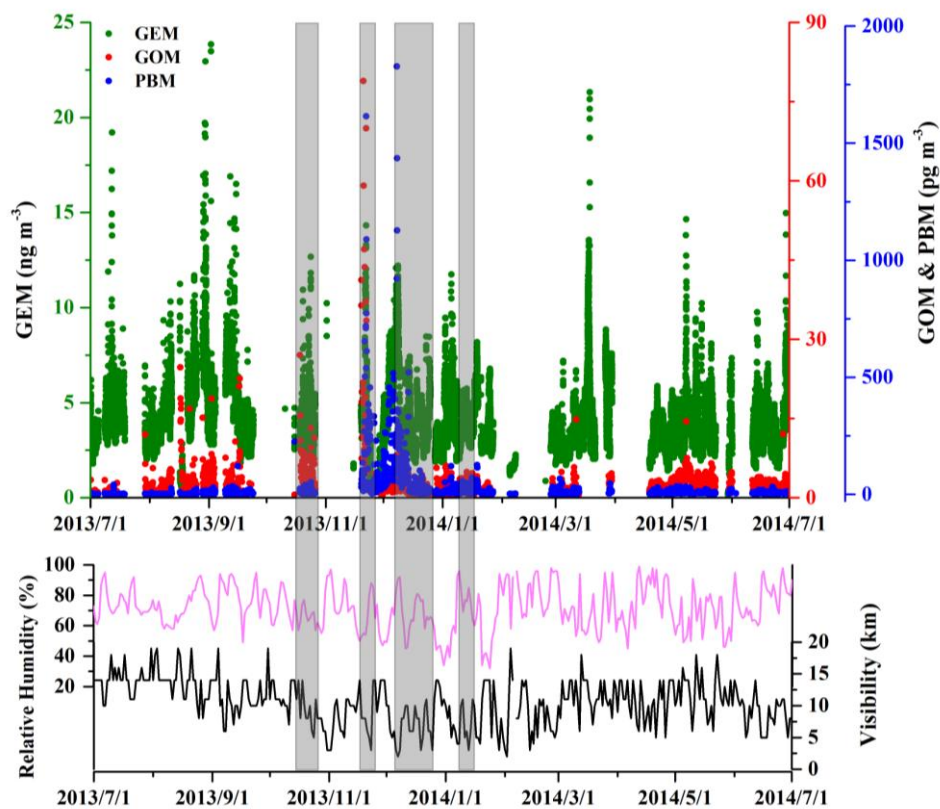
Fig. 1. Location of the monitoringstudy site in Hefei, China.

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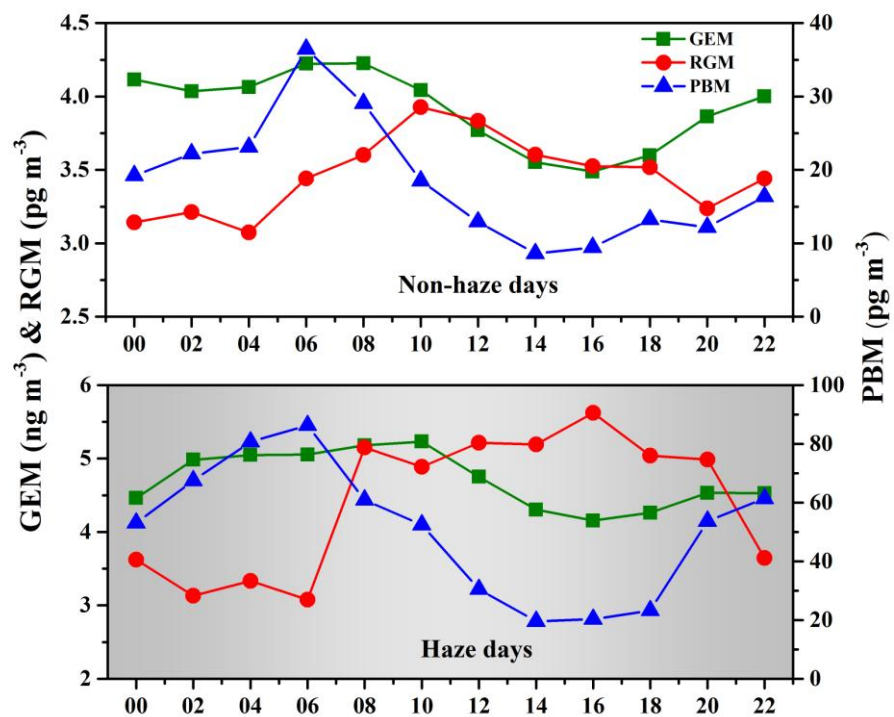
1324 **Fig. 2.** Time series of GEM, RGMGOM and PBM concentrations, along with visibility,
 1325 relatively humidity, at the monitoring site in Hefei from July 2013 to June 2014. The
 1326 GEM data were at a 5-min resolution, and the RGMGOM and PBM data were two-hour
 1327 averages. The gray columns show the major haze pollution episodes occurred during
 1328 the study period.

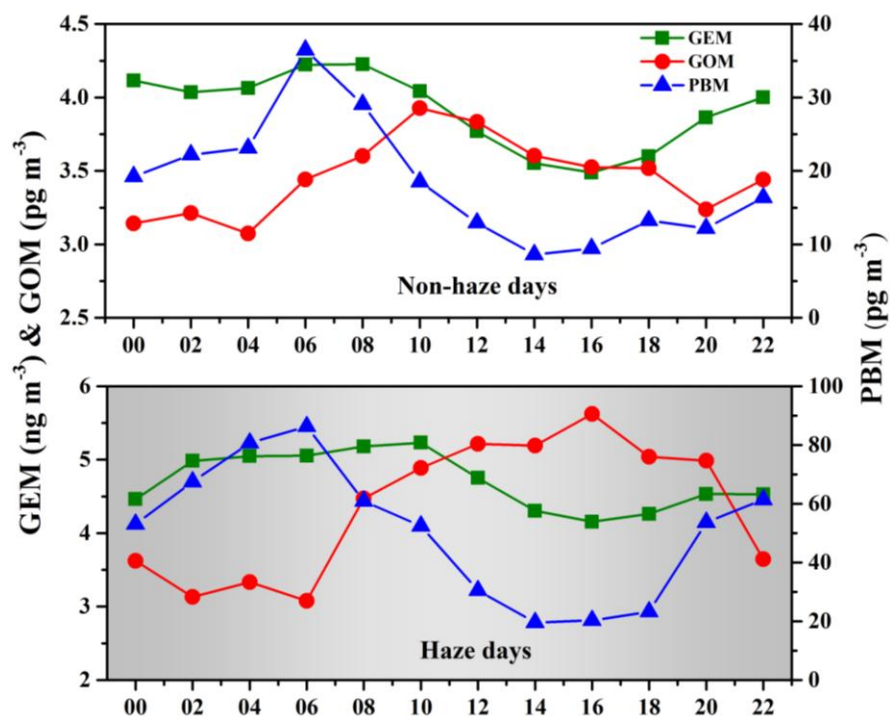
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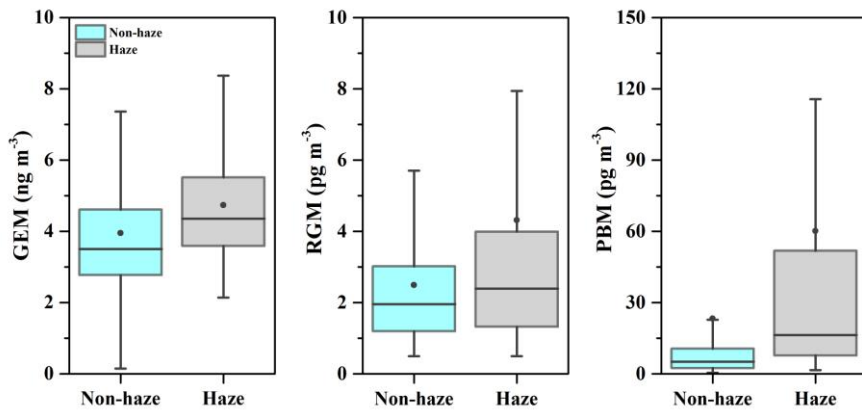
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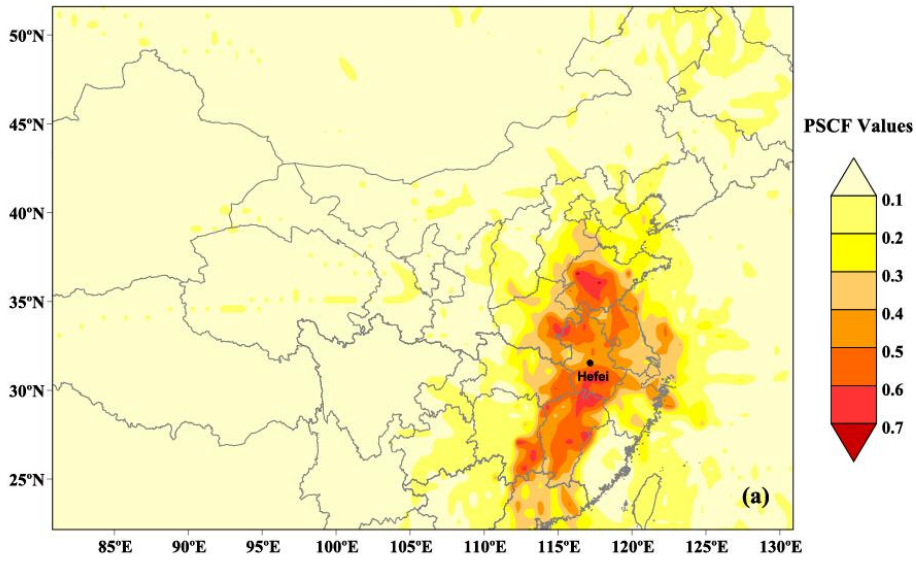
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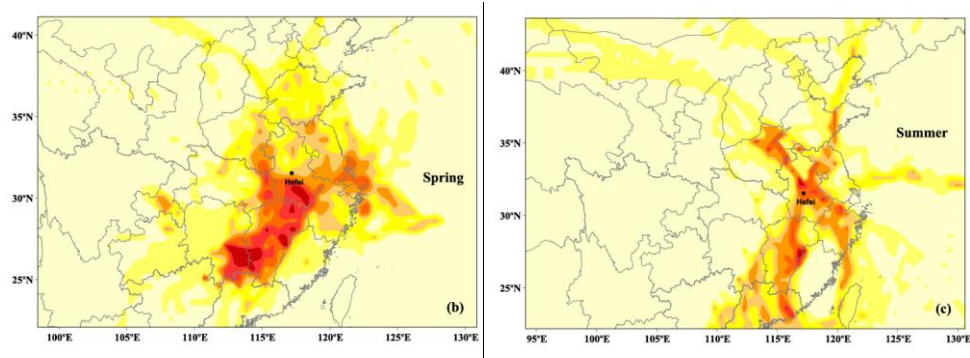
Fig. 3. Diurnal trends of GEM, RGMGOM and PBM concentrations in Hefei during non-haze and haze days (Local time = UTC + 8 hr). The data were two-hour averages.



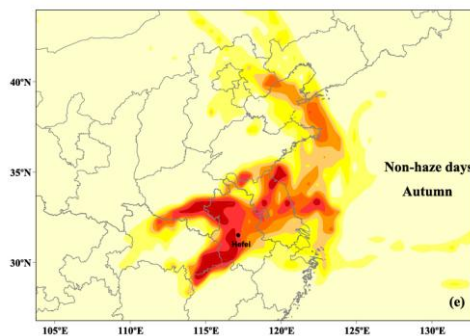
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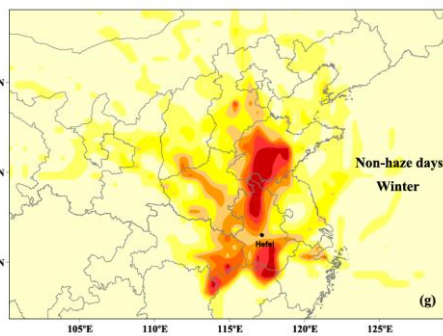
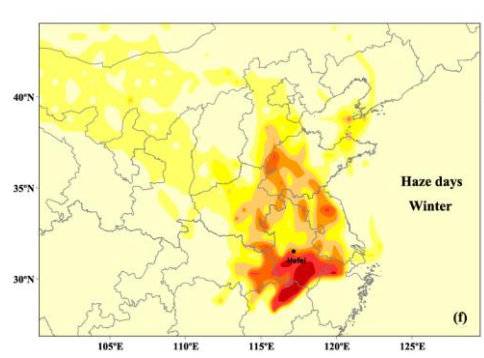
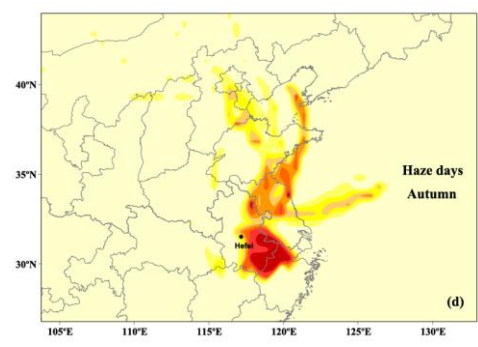


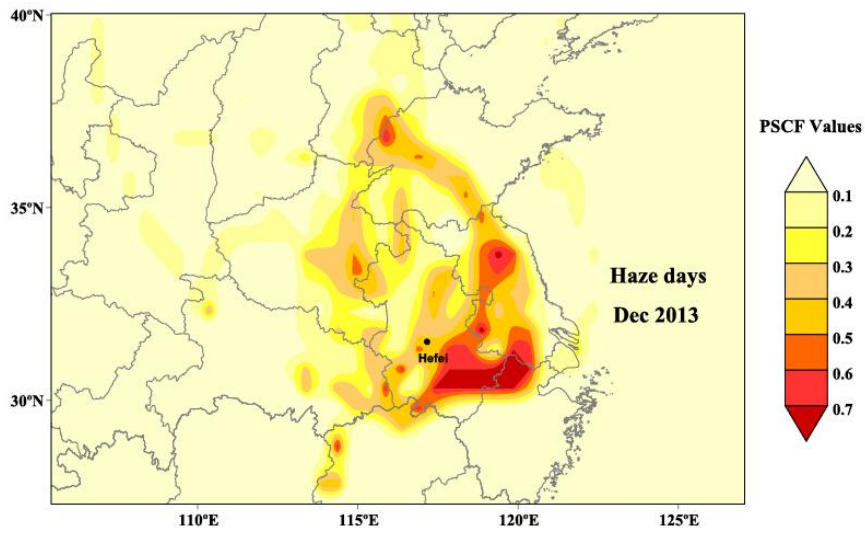
Fig. 4. GEM, RGM and PBM concentrations during non haze and haze days. The GEM data were at a 5 min resolution, the RGM and PBM data were two hour averages.

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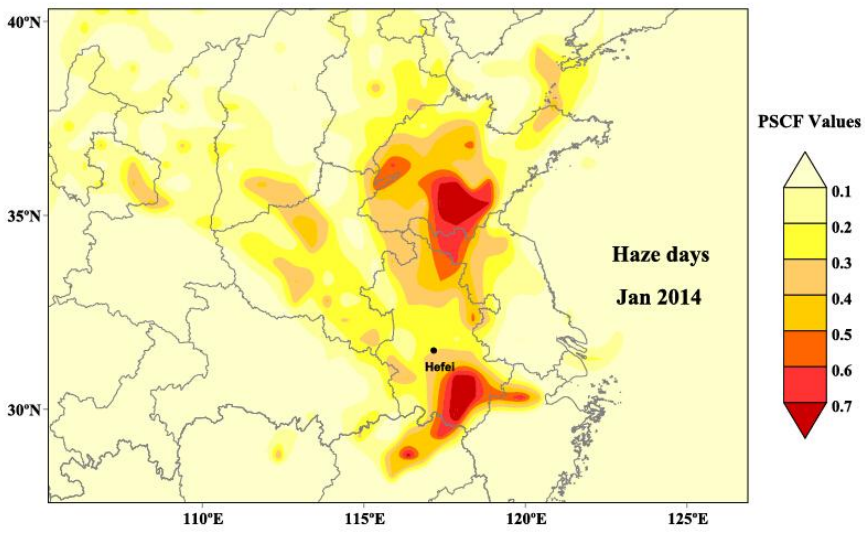
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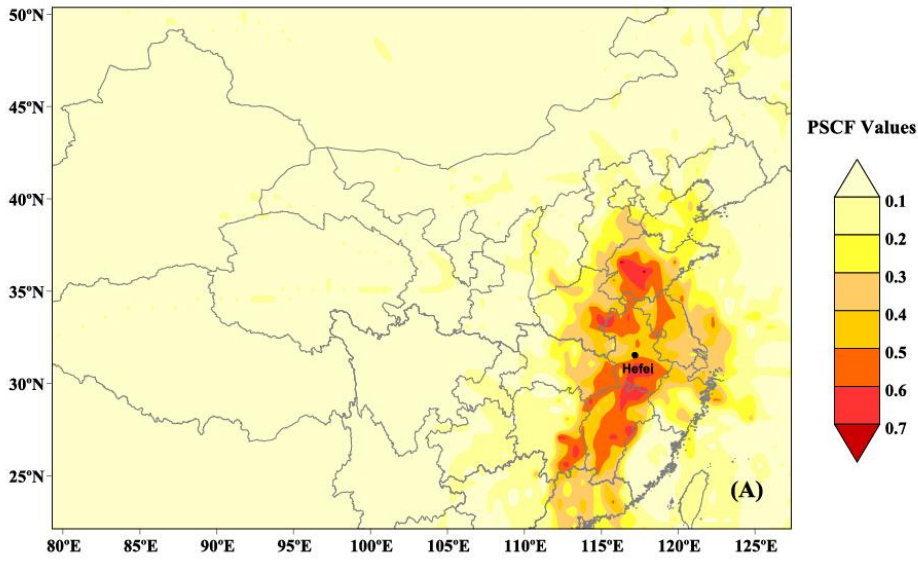
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Fig. 5. Likely sources areas of GEM during haze days identified by PSCF analysis.

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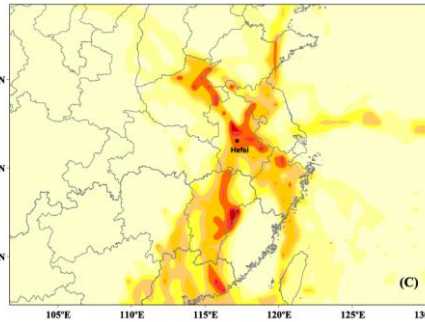
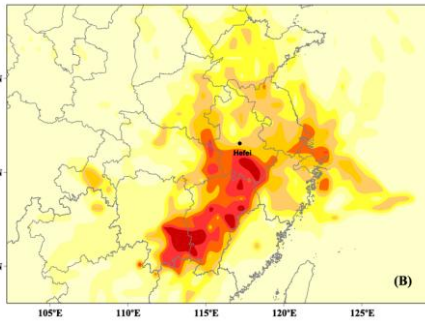
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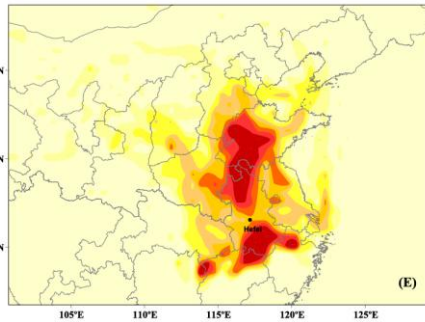
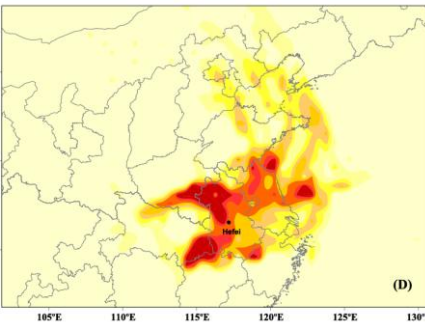
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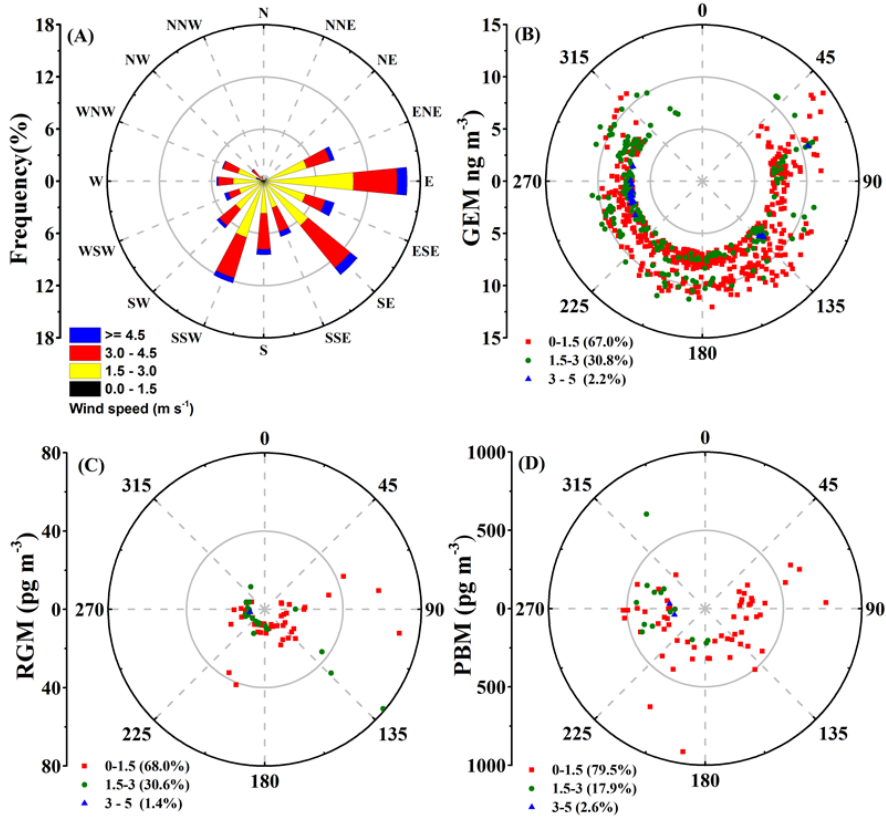
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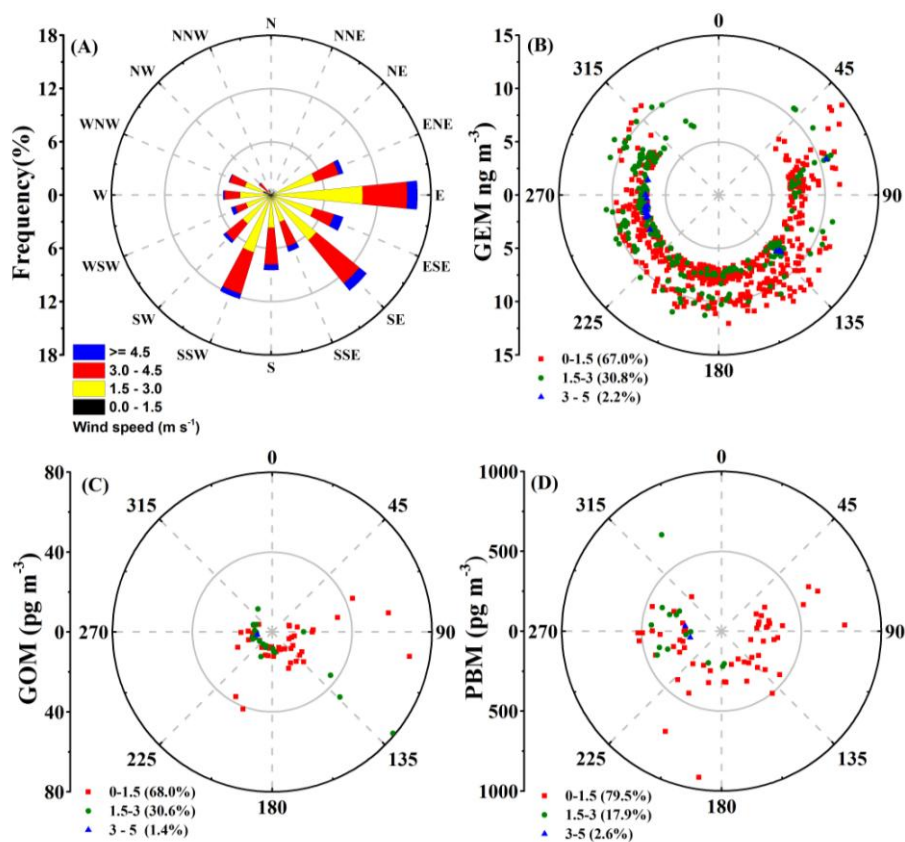
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Fig. 6. Likely emission source areas of GEM simulated by PSCF analysis. (A) overall (from July 2013 to June 2014), (B) spring, (C) summer, (D) haze days in autumn, (E) non-haze days in autumn, (f) haze days in winter, (g) non-haze days in winter.

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Fig. 75. Wind direction and speed at the Science Island Meteorological Station during

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the study period. (A) the wind rose for the whole study period; (B), (C) and (D) are

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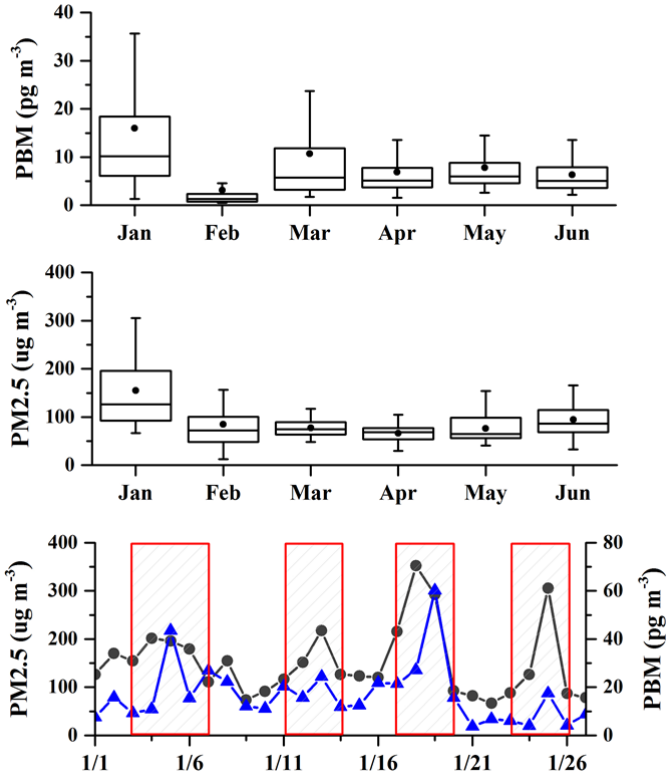
the wind rose diagrams for GEM, RGMGOM and PBM concentrations above the 90th

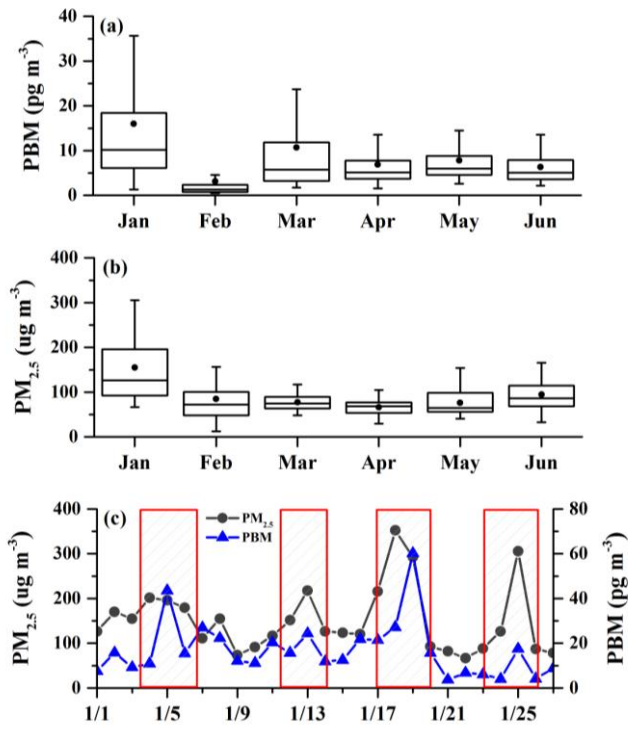
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percentile values, respectively.

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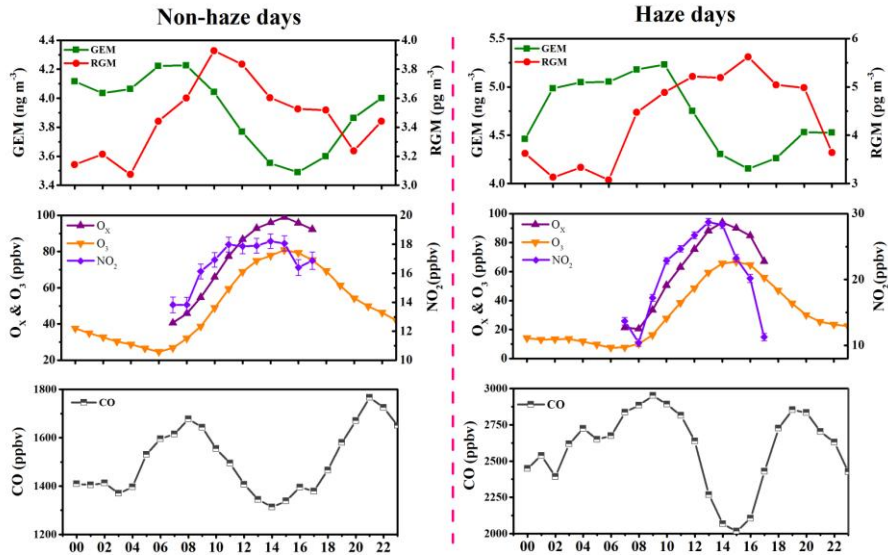


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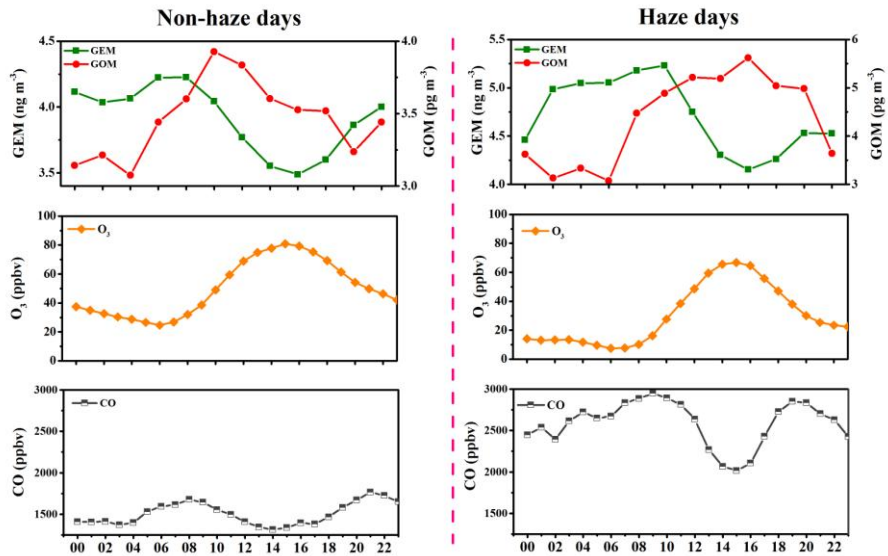
Fig. 8-6. Monthly variation of (a) PBM and (b) PM_{2.5} concentrations from January to June, 2014., (c) average daily PM_{2.5} and PBM concentrations in January, 2014.,

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Fig. 9.7. Diurnal variations of GEM, RGM and Gas Phase Data (O_3 , GOM, O_3 , NO_2

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and CO_2 concentrations during non-haze and haze days. Notes:-

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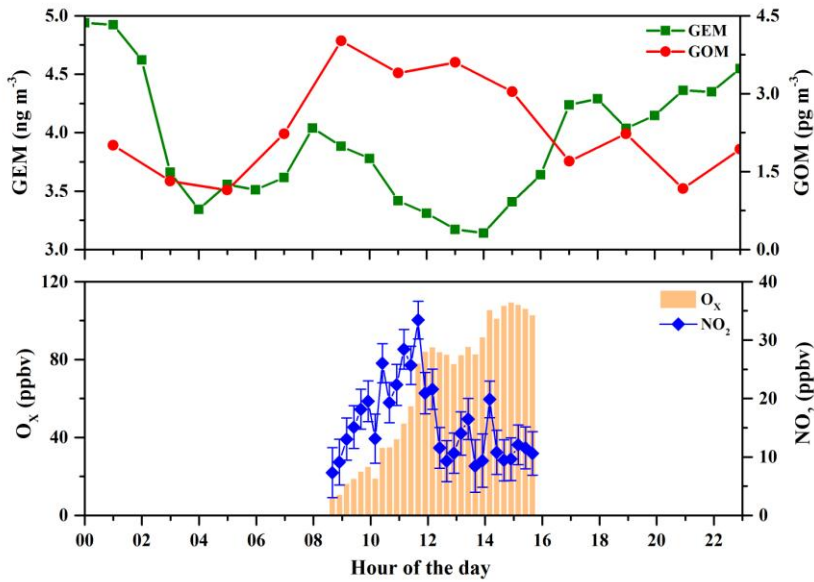
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Fig. 8. A case study of diurnal variations of GEM, GOM, O_x , and NO_2 at Hefei (20th November, 2013). The top panel shows the hourly averaged GEM and GOM concentrations, and the bottom is the carbon monoxide mixing ratio. Middle are the averaged $O_{3\tau}$ panel shows the O_x ($O_x = NO_2 + O_3$) and the NO_2 concentrations. Top are the hourly averaged GEM and RGM concentrations. The error bars for NO_2 refer to the NO_2 standard deviations.

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