

Direct Inversion of Circulation and Mixing from Tracer Measurements: I. Method

Thomas von Clarmann¹ and Udo Grabowski¹

¹Karlsruhe Institute of Technology, Institute of Meteorology and Climate Research,
Karlsruhe, Germany

Correspondence to: T. von Clarmann (thomas.clarmann@kit.edu)

Abstract. From a series of zonal mean global stratospheric tracer measurements sampled in altitude versus latitude, circulation and mixing patterns are inferred by the inverse solution of the continuity equation. As a first step, the continuity equation is written as a tendency equation, which is numerically integrated over time to predict a later atmospheric state, i.e. mixing ratio and air density. The integration is formally performed by multiplication of the initially measured atmospheric state vector by a linear prediction operator. Further, the derivative of the predicted atmospheric state with respect to the wind vector components and mixing coefficients is used to find the most likely wind vector components and mixing coefficients which minimize the residual between the predicted atmospheric state and the later measurement of the atmospheric state. Unless multiple tracers are used, this inversion problem is under-determined, and dispersive behaviour of the prediction further destabilizes the inversion. Both these problems are fought by regularization. For this purpose, a first order smoothness constraint has been chosen. The usefulness of this method is demonstrated by application to various tracer measurements recorded with the Michelson Interferometer for Passive Atmospheric Sounding (MIPAS). This method aims at a diagnosis of the Brewer-Dobson circulation without involving the concept of the mean age of stratospheric air, and related problems like the stratospheric tape recorder, or intrusions of mesospheric air into the stratosphere.

2006). Engel et al. (2009), however, found a weakly significant slow increase of the mean age of stratospheric air. The latter is defined as the mean time lag between the date of the transition of tropospheric air into the stratosphere and the date when the mixing ratio of a monotonically growing tracer was measured in the air volume under investigation, and its increase hints at a deceleration of the Brewer-Dobson circulation. These measurements have been challenged as not representative (Garcia et al., 2011), and global mean age of air measurements by Stiller et al. (2012) suggest that the true picture is not that one-dimensional. Instead, stratospheric age trends vary with altitude and with latitude. Determination of the age of air and its use as diagnostic of the intensity of the Brewer-Dobson circulation, however, has its own limitations: First, due to mixing processes, the age of a stratospheric air volume is not unique but characterized by an age spectrum, which has to be considered since the tropospheric growth of SF₆ mixing ratios is not strictly linear, and on which some *ad hoc* assumptions have to be made (Waugh and Hall, 2002). These include the adequacy of the Wald (inverse Gaussian) function for the representation of the age spectrum and the choice of its width parameter. Second, the most suited age tracer, SF₆, which has significant and monotonic growth rates in the troposphere, is not fully inert: It has a mesospheric sink (Hall and Waugh, 1998; Reddmann et al., 2001) and introduces some age uncertainty when mesospheric air subsides into the stratosphere in the polar winter vortex (Stiller et al., 2008). Third, the determination of the mean age relies upon a reference airmass where the age, by definition, is zero. When the age of air concept was introduced, the reference was simply the troposphere, which is well mixed and thus avoids any related complication (Solomon, 1990; Schmidt and Khedim, 1991). Since the age of air has become a model diagnostic, parts of the community have established the stratospheric entry point as a reference (Hall and Plumb, 1994), which makes

1 Introduction

In the context of climate change, possible changes of the intensity of the Brewer-Dobson circulation have become an important research topic. Climate models predict an intensification of the Brewer-Dobson circulation (Butchart et al.,

a difference due to the slow ascent of air through the tropical tropopause layer (Fueglistaler et al., 2009). For model validation, however, this redefined age of air is of limited use, because no long measured time series of tracer mixing ratios are available there.

Facing these difficulties, it is desirable to infer the atmospheric circulation directly from tracer measurements, without going back to the age of air concept. Multiple approaches have been developed to infer windfields from measured atmospheric state variables. Sequential data assimilation, and in its optimal form, the extended Kalman filter approach (e.g. Ghil and Malanotte-Rizzoli, 1991; Ghil, 1997), calculates the optimal average of the forecasted meteorological variables for the time of the observation and the observed meteorological variables themselves and uses this average to initialize the next forecast step. The wind field is calculated by a dynamical model. This method involves the generalized inversion of the observation operator where the forecast is used as a constraint. In contrast, so-called¹ variational data assimilation minimizes the residual between the forecasted and the measured atmospheric state variables by optimally adjusting the initialization of the forecast via inversion of an adjoint forecast model, constrained by some background state (Thompson, 1969). Both approaches rely on dynamical models² and are suited to infer the most probable atmospheric state variables rather than the windfield, which is a by-product of the assimilation. The windfield or atmospheric circulation can also be inferred directly by kinematic methods from tracer measurements. Such methods rely solely on the continuity equation, do not involve a dynamic model, and thus do not depend on any *ad hoc* parametrisation of effects which are either not resolved by the discrete model, computationally too expensive for explicit modeling, or simply not well understood. While this work is targeted primarily at an assessment of the Brewer-Dobson circulation, its applicability is much wider and includes stratospheric-tropospheric exchange, the mesospheric overturning circulation and others. Early approaches to infer the circulation from tracer measurements include Holton and Choi (1988) as well as Salby and Juckes (1994) who used approaches which share several ideas with ours. Direct inversion of wind speeds from tracer measurements in a volcano plume has, e.g., been suggested by Krueger et al. (2013), however without consideration of mixing. The continuity equation including diffusion terms has been exploited by Wofsy et al. (1994) for assessment of diffusion of stratospheric aircraft exhaust.

In this paper we present a method to infer two-dimensional (latitude/altitude) circulation and mixing coefficients from subsequent measurements of inert tracers. The application of

this method, i.e., the inference of the Brewer-Dobson circulation from global SF₆ distributions (Stiller et al., 2008, 2012) measured with the Michelson Interferometer for Passive Atmospheric Sounding (MIPAS), is presented in a companion paper. In order to avoid that the reader does not see the forest for the trees, we give a short overview of our method in Section 2. The prediction of pressure and tracer mixing ratio fields on the basis of the continuity equation and related error estimation is described in Section 3. The estimation of circulation and mixing coefficients by inversion of the continuity equation is presented in Section 4. In Section 5, the applicability of the method and the need of further refinements is critically discussed. The benefits of the method are discussed in Section 6. The paper concludes with recommendations how these results should be used and with an outlook on future work (Section 7). Changes of the Brewer-Dobson circulation during 2002–2012, i.e. the MIPAS mission, are currently investigated by means of this method and will be published in a subsequent paper.

2 General concept

Knowing the initial state of the atmosphere in terms of mixing ratio and air density distributions, wind speed and mixing coefficients at each gridpoint, a future atmospheric state can be predicted with respect to the distribution of any inert tracer. This procedure we call the forward problem. If no ideal tracers are available, source and sink terms of related species have to be included in the forward model. The goal of this work is to invert the forward model in order to infer the circulation and mixing coefficients from tracer measurements by minimization of the residual between the predicted and measured atmospheric state. This approach is complementary to free running climate models because it makes no assumptions about atmospheric dynamics except the validity of the continuity equation. It is further considered more robust than age-of-air analysis (Stiller et al., 2012) because it does not depend on a reference point where the age is assumed zero, nor does it require knowledge on the history of an air parcel.

Our concept involves the following operations. First, a general solution of the forward problem is formulated (Section 3). The forward problem is the solution of the prediction equation as a function of the initial atmospheric state for given winds and mixing coefficients. For our chosen solver, which involves the MacCormack (1969) integration scheme for the solution of the transport problem (Eqs. 5–10), the relevant dependencies of the final state on the initial state are reported in Section 3.2 (Eqs. 15–26). The formulation in matrix notation (Eqs. 27–28) allows the easy handling of multiple successive timesteps (Eq. 29) and an easy estimation of the prediction error via generalized Gaussian error estimation (Eq. 30). As a next step, the dependence of the predicted state on winds and mixing coefficients is estimated

¹The term ‘so-called’ is used here, because it is challenged that this method is really variational in the context of discrete variables (Wunsch, 1996)[p368].

²This statement refers to meteorological data assimilation. Chemical data assimilation uses chemistry transport models.

for a given initial state. This is achieved by differentiation of the solution of the prediction equation with respect to winds and mixing coefficients (Eqs. 34–76). These partial derivatives form the Jacobian matrix of the problem, with which the estimation of winds and mixing coefficients can be reduced to a constrained least square optimization problem where the inversely variance-weighted residual between the predicted atmospheric state and the respective measured atmospheric state is minimized. The latter step involves the generalized inverse of the Jacobian matrix (Eqs. 78–90).

3 The forward problem

The forward model reads the measured atmospheric state at time t and predicts the atmospheric state (number density of air, c , and volume mixing ratios, vmr) at time $t + \Delta t$ for given wind vectors and mixing coefficients representing the time interval $[t; t + \Delta t]$ by solving the continuity equation. The continuity equation allows to calculate the local tendencies of the number densities and volume mixing ratios. These local tendencies $\frac{\partial \rho}{\partial t}$ and $\frac{\partial vmr}{\partial t}$ are then integrated over time to give the new number densities and mixing ratios.

3.1 The continuity equation

The local change of number density ρ of air is in spherical coordinates (for all auxiliary calculations, see supplement):

$$\frac{\partial \rho}{\partial t} = -\frac{1}{r} \frac{\partial \rho v}{\partial \phi} + \frac{\rho v}{r} \tan(\phi) - \frac{\partial \rho w}{\partial z} - \frac{2\rho w}{r} - \frac{1}{r \cos(\phi)} \frac{\partial \rho u}{\partial \lambda} \quad (1)$$

where

- t = time
- λ = longitude
- ϕ = latitude
- z = altitude above surface
- r = $r_E + z$
- r_E = radius of Earth
- u = $(r_E + z) \cos \phi d\lambda/dt$
- v = $(r_E + z) d\phi/dt$
- w = dz/dt

Here the shallowness approximation (Hinkelmann (1951); Phillips (1966), quoted after Kasahara (1977)), which simplifies the equations using the assumption that z is much smaller than r_E and which is, often implicitly, used in the usual textbooks on atmospheric sciences (e.g. Brasseur and Solomon, 2005, their Eq. 3.46a), is intentionally not used for reasons which will become clear in Section 3.2.

The local change of the volume mixing ratio of gas g can be calculated from known velocities and mixing coefficients

as well as source/sink terms as

$$\begin{aligned} \frac{\partial vmr_g}{\partial t} = & \frac{S_g}{\rho} - \frac{u}{r \cos \phi} \frac{\partial vmr_g}{\partial \lambda} - \frac{v}{r} \frac{\partial vmr_g}{\partial \phi} \\ & - w \frac{\partial vmr_g}{\partial z} + \frac{1}{r^2} \frac{\partial}{\partial \lambda} \left[\frac{K_\lambda}{\cos^2 \phi} \frac{\partial vmr_g}{\partial \lambda} \right] \\ & + \frac{1}{r^2 \cos \phi} \frac{\partial}{\partial \phi} \left[K_\phi \cos \phi \frac{\partial vmr_g}{\partial \phi} \right] \\ & + \frac{1}{r^2} \frac{\partial}{\partial z} \left[r^2 K_z \frac{\partial vmr_g}{\partial z} \right] \end{aligned} \quad (2)$$

where

- vmr_g = volume mixing ratio of species g
- K_λ = zonal diffusion coefficient
- K_ϕ = meridional diffusion coefficient
- K_z = vertical diffusion coefficient
- S_g = the production/loss rate of species g in terms of number density over time

(Brasseur and Solomon (e.g. 2005, Eq. 3.46b) and Jones et al. (2007)).

Since we are only interested in a two-dimensional representation of the atmosphere in altitude and latitude coordinates, zonal advection and mixing terms are ignored in Eqs. (1–2). In this two-dimensional representation, all atmospheric state variables represent zonal mean values. The kinematic variables, *viz.* the velocities and mixing coefficients, have to be re-interpreted because they do not represent merely the zonally averaged velocities and mixing coefficients. Instead, they include also eddy transport and diffusion terms, and their interpretation is less unique than one might hope because it depends on the definition of the kinematic variables and the approximations used (see Appendix A for details of the interpretation of the kinematic quantities). The local change of number density ρ of air in a two-dimensional atmosphere thus is

$$\frac{\partial \rho}{\partial t} = -\frac{1}{r} \frac{\partial \rho v}{\partial \phi} + \frac{\rho v}{r} \tan(\phi) - \frac{\partial \rho w}{\partial z} - \frac{2\rho w}{r} \quad (3)$$

and the local change of vmr_g is calculated as

$$\begin{aligned} \frac{\partial vmr_g}{\partial t} = & \frac{S_g}{\rho} - \frac{v}{r} \frac{\partial vmr_g}{\partial \phi} - w \frac{\partial vmr_g}{\partial z} \\ & + \frac{1}{r^2 \cos \phi} \frac{\partial}{\partial \phi} \left[K_\phi \cos \phi \frac{\partial vmr_g}{\partial \phi} \right] \\ & + \frac{1}{r^2} \frac{\partial}{\partial z} \left[r^2 K_z \frac{\partial vmr_g}{\partial z} \right] \end{aligned} \quad (4)$$

3.2 Integration of tendencies

Integration of Eqs (3–4) is performed numerically for timesteps of Δt_p . For practical reasons, processes (advection, diffusion, sinks) are splitted, i.e. the tendencies due to these three classes of processes are integrated independently.

The timesteps Δt_p used for the integration are chosen smaller than the time difference Δt between two measurements, in order not to clash with the Courant limit (Courant et al., 1952). In the following we call Δt_p ‘micro time increment’ and the latter ‘macro time increment’. The atmospheric state after a macro time increment is predicted by successive prediction over the micro time increment. In the following, index i designates time t , $i+1$ designates the time $t + \Delta t_p$, etc, and I designates the time after the final micro time increment, i.e. the next macro time increment.

For the discrete integration of the advection part of the tendencies the MacCormack (1969) method is used in a generalized multidimensional version similar to the one described by (Perrin and Hu, 2006). This is a predictor-corrector method. For a general state variable $c(t, x, y) = c_i(x, y)$ at location (x, y) , and time t with $e(c)$ and $f(c)$ being functions of c , an equation of the form

$$\frac{\partial c}{\partial t} + \frac{\partial e(c)}{\partial x} + \frac{\partial f(c)}{\partial y} = 0 \quad (5)$$

is solved by preliminary predictions of the state variable as a first step: x is

$$\begin{aligned} c_{i+1}^*(x, y) &= c_i(x, y) - \\ &\quad \frac{\Delta t_p}{\Delta x} (e_i(x + \Delta x, y) - e_i(x, y)) \\ &\quad - \frac{\Delta t_p}{\Delta y} (f_i(x, y + \Delta y) - f_i(x, y)). \end{aligned} \quad (6)$$

These are then used in a subsequent correction step which gives the final prediction:

$$\begin{aligned} c_{i+1}(x, y) &= \\ &\quad \frac{1}{2} [c_i(x, y) + c_{i+1}^*(x, y) - \\ &\quad \frac{\Delta t_p}{\Delta x} (e(c_{i+1}^*, x, y) - e(c_{i+1}^*, x - \Delta x, y)) - \\ &\quad \frac{\Delta t_p}{\Delta y} (f(c_{i+1}^*, x, y) - f(c_{i+1}^*, x, y - \Delta y))] \end{aligned} \quad (7)$$

Application to the continuity equation in spherical coordinates requires reformulation of Eq. (3) (c.f., e.g., Chang and St.-Maurice, 1991):

$$\frac{\partial r^2 \rho \cos(\phi)}{\partial t} = -\frac{\partial r \rho v \cos(\phi)}{\partial \phi} - \frac{\partial r^2 \rho w \cos(\phi)}{\partial z} \quad (8)$$

The predictor of $r^2 \rho \cos(\phi)$ is then calculated as

$$\begin{aligned} [r^2 \rho_{i+1}(\phi, z) \cos(\phi)]^* &= \\ &\quad r^2 \rho_i(\phi, z) \cos(\phi) \\ &\quad - \frac{\Delta t_p}{\Delta \phi} (r \rho_i(\phi + \Delta \phi, z) \cos(\phi + \Delta \phi) v(\phi + \Delta \phi, z) \\ &\quad - r \rho_i(\phi, z) \cos(\phi) v(\phi, z)) \\ &\quad - \frac{\Delta t_p}{\Delta z} ((r + \Delta z)^2 \rho_i(\phi, z + \Delta z) w_{\phi, z + \Delta z} \cos(\phi) \\ &\quad - r^2 \rho_i(\phi, z) w_{\phi, z} \cos(\phi)) \end{aligned} \quad (9)$$

and the corrected prediction for ρ then gives

$$\begin{aligned} \rho_{i+1}(\phi, z) &= \frac{1}{2r^2 \cos(\phi)} \times (10) \\ &\quad \left[\rho_i(\phi, z) r^2 \cos(\phi) \right. \\ &\quad + [\rho_{i+1}(\phi, z) r^2 \cos(\phi)]^* \\ &\quad - \frac{\Delta t_p}{\Delta \phi} [[\rho_{i+1}(\phi, z) r v(\phi, z) \cos(\phi)]^* \\ &\quad - [\rho_{i+1}(\phi - \Delta \phi, z) r v(\phi - \Delta \phi, z) \cos(\phi - \Delta \phi)]^* \\ &\quad - \frac{\Delta t_p}{\Delta z} [[\rho_{i+1}(\phi, z) r^2 w(\phi, z) \cos(\phi)]^* \\ &\quad \left. - [\rho_{i+1}(\phi, z - \Delta z) (r - \Delta z)^2 w(\phi, z - \Delta z) \cos(\phi)]^* \right] \end{aligned}$$

For the local change of mixing ratio, operator splitting is performed. The horizontal and vertical advective parts of the continuity equation for mixing ratios in two dimensions are transformed into the following Mac-Cormack-integrable forms:

$$\left[\frac{\partial \frac{v \text{vmr}_g}{v}}{\partial t} \right]_{adv.horiz} = \frac{\partial \text{vmr}_g}{\partial \phi} \quad (11)$$

and

$$\left[\frac{\partial \frac{w \text{vmr}_g}{w}}{\partial t} \right]_{adv.vert} = \frac{\partial \text{vmr}_g}{\partial z}, \quad (12)$$

respectively.

For the diffusive component we use simple Eulerian integration:

$$\begin{aligned} [\text{vmr}_{g;i+1}(\phi, z) - \text{vmr}_{g;i}(\phi, z)]_{\text{diff}} &= \\ &\quad \frac{\Delta t_p}{2r^2 (\Delta \phi)^2 \cos(\phi)} \cdot \\ &\quad \left[(K_\phi(\phi, z) + K_\phi(\phi + \Delta \phi)) \cos(\phi + \frac{\Delta \phi}{2}) \right. \\ &\quad (v \text{vmr}_{g;i}(\phi + \Delta \phi, z) - v \text{vmr}_{g;i}(\phi, z)) \\ &\quad - (K_\phi(\phi, z) + K_\phi(\phi - \Delta \phi)) \cos(\phi - \frac{\Delta \phi}{2}) \cdot \\ &\quad (v \text{vmr}_{g;i}(\phi, z) - v \text{vmr}_{g;i}(\phi - \Delta \phi, z)) \left. \right] \\ &\quad + \frac{\Delta t_p}{2r^2 (\Delta z)^2} \left[(r + \frac{\Delta z}{2})^2 \cdot \right. \\ &\quad (K_z(\phi, z) + K_z(\phi, z + \Delta z)) \cdot \\ &\quad (v \text{vmr}_{g;i}(\phi, z + \Delta z) - v \text{vmr}_{g;i}(\phi, z)) - \\ &\quad (r - \frac{\Delta z}{2})^2 \cdot (K_z(\phi, z) + K_z(\phi, z - \Delta z)) \cdot \\ &\quad (v \text{vmr}_{g;i}(\phi, z) - v \text{vmr}_{g;i}(\phi, z - \Delta z)) \left. \right] \end{aligned} \quad (13)$$

Sinks of the species considered here are treated as unimolecular processes (c.f., e.g. Brasseur and Solomon, 2005, their Eq. 2.27d) and integrated as

$$\rho_{g;i+1} = \rho_{g;i} e^{-k_g \Delta t_p} \quad (14)$$

where k_g is the sink strength of the gas g .

The abundance of gas g after time-step Δt_p is then simply the sum of the increments due to horizontal and vertical advection, diffusion, and chemical losses.

Admittedly, there exist more elaborate advection schemes than the one used here. However, the need to provide the Jacobians needed in Sections 3.3–4 justifies a reasonable amount of simplicity. Further, numerical errors cannot easily accumulate, because after each timestep Δt , the system is re-initialized with measured data.

Since we do not have a closed system but have mass exchange and mixing with the atmosphere below the lowermost model altitude and above the uppermost altitude, the atmospheric state is not predicted for the lowermost and uppermost altitudes. Prediction is only possible from the second to one below the uppermost altitude. This restricted altitude range we henceforth call ‘nominal altitude range’. Instead, the atmospheric state of the uppermost and lowermost altitude is estimated by linear interpolation of measured values at times t and $t + \Delta t$ and used as boundary condition for prediction within the nominal altitude range. Alternatively, derivatives at the border can be approximated by asymmetric difference quotients.

We use the following convention: Atmospheric state variables are sampled on a regular latitude-altitude grid. For some gridpoints, no valid measurements may be available but we assume that, for each state variable, we have a contiguous subset of this grid with valid measurements. For state variable g , we have a total of J_g valid measurements within the ‘nominal altitude range’, each denoted by index j . A state variable in this context is either air density ρ ($g = 0$) or the mixing ratio of one species vmr_g . The nominal altitude range at latitude ϕ is the altitude range where, for each gridpoint, a valid measurement is available at the gridpoint itself, and for its northern, southern, upper, lower and diagonal neighbours. The use of asymmetric difference quotients can be emulated by extrapolation, generating artificial border values by, which guarantee that each gridpoint within the nominal altitude and latitude range has all required neighbours. The availability of neighbour-values is necessary to allow the calculation of numerical derivatives of the state variable with respect to latitude and altitude. We therefore have K_g border elements of each quantity g , each denoted by index k . This implies, for each state variable g , a total of $L_g = J_g + K_g$ gridpoints with indices l .

3.3 Integration in operator notation

For further steps (error propagation and the solution of the inverse problem) it is convenient to rewrite the prediction of

air density and mixing ratios in matrix notation. For this purpose, we differentiate the predicted air densities (Eq. 10) and mixing ratios (Eqs. 11–13) with respect to air density and mixing ratios of the gases under assessment at all relevant locations. The sensitivities of the densities of the first predictive step with respect to the initial densities at the same latitude and altitude are

$$\frac{\partial \rho_{i+1}(\phi, z)}{\partial \rho_i(\phi, z)} = \quad (15)$$

$$\begin{aligned} \frac{1}{2} \left[2 - \frac{\Delta t_p}{\Delta \phi} \left[\frac{v(\phi, z)}{r} \left(\frac{\Delta t_p}{\Delta \phi} \cdot \frac{v(\phi, z)}{r} + \frac{\Delta t_p}{\Delta z} w(\phi, z) \right) \right. \right. \\ \left. \left. + \frac{\Delta t_p}{\Delta \phi} \frac{v(\phi - \Delta \phi, z) v(\phi, z)}{r^2} \right] \right. \\ \left. - \frac{\Delta t_p}{\Delta z} \left[w(\phi, z) \left(\frac{\Delta t_p}{\Delta \phi} \cdot \frac{v(\phi, z)}{r} + \frac{\Delta t_p}{\Delta z} w(\phi, z) \right) \right. \right. \\ \left. \left. + \frac{\Delta t_p}{\Delta z} w(\phi, z - \Delta z) w(\phi, z) \right] \right] \end{aligned}$$

We further differentiate predicted air densities with respect to air densities at the adjacent southern latitude but the same altitude.

$$\begin{aligned} \frac{\partial \rho_{i+1}(\phi, z)}{\partial \rho_i(\phi - \Delta \phi, z)} = \frac{1}{2} \left[\frac{\Delta t_p}{\Delta \phi} \cdot \frac{v(\phi - \Delta \phi, z)}{r} \cdot \frac{\cos(\phi - \Delta \phi)}{\cos(\phi)} \right. \\ \left. \left(1 + \frac{\Delta t_p}{\Delta \phi} \cdot \frac{v(\phi - \Delta \phi, z)}{r} + \frac{\Delta t_p}{\Delta z} w(\phi - \Delta \phi, z) \right) \right] \end{aligned} \quad (16)$$

The derivative of the predicted air densities with respect to air densities at the adjacent northern latitude but the same altitude is

$$\begin{aligned} \frac{\partial \rho_{i+1}(\phi, z)}{\partial \rho_i(\phi + \Delta \phi, z)} = \frac{1}{2} \left[\frac{\Delta t_p}{\Delta \phi} \cdot \frac{v(\phi + \Delta \phi, z)}{r} \cdot \frac{\cos(\phi + \Delta \phi)}{\cos(\phi)} \right. \\ \left. \left(-1 + \frac{\Delta t_p}{\Delta \phi} \cdot \frac{v(\phi, z)}{r} + \frac{\Delta t_p}{\Delta z} w(\phi, z) \right) \right]. \end{aligned} \quad (17)$$

As a next step we differentiate predicted air densities with respect to the initial air densities at the next higher altitude but the same latitude.

$$\begin{aligned} \frac{\partial \rho_{i+1}(\phi, z)}{\partial \rho_i(\phi, z + \Delta z)} = \frac{1}{2} \left[\frac{\Delta t_p}{\Delta z} \cdot \frac{(r + \Delta z)^2}{r^2} w(\phi, z + \Delta z) \cdot \right. \\ \left. \left(-1 + \frac{\Delta t_p}{\Delta \phi} \cdot \frac{v(\phi, z)}{r} + \frac{\Delta t_p}{\Delta z} w(\phi, z) \right) \right]. \end{aligned} \quad (18)$$

The derivative of the predicted air densities with respect to the initial air densities at the next lower altitude but the same latitude is

$$\begin{aligned} \frac{\partial \rho_{i+1}(\phi, z)}{\partial \rho_i(\phi, z - \Delta z)} = \frac{1}{2} \left[\frac{\Delta t_p}{\Delta z} w(\phi, z - \Delta z) \frac{(r - \Delta z)^2}{r^2} \cdot \right. \\ \left. \left(1 + \frac{\Delta t_p}{\Delta \phi} \frac{v(\phi, z - \Delta z)}{r - \Delta z} + \frac{\Delta t_p}{\Delta z} w(\phi, z - \Delta z) \right) \right]. \end{aligned} \quad (19)$$

Finally we differentiate the predicted air densities with respect to the initial air densities at the adjacent southern latitude and higher altitude

$$\begin{aligned} & \frac{\partial \rho_{i+1}(\phi, z)}{\partial \rho_i(\phi - \Delta\phi, z + \Delta z)} = \\ & -\frac{1}{2} \left[\frac{v(\phi - \Delta\phi, z)}{r} \cdot \frac{\Delta t_p}{\Delta\phi} \cdot \frac{\Delta t_p}{\Delta z} \cdot \frac{(r + \Delta z)^2}{r^2} \cdot \right. \\ & \left. \cdot \frac{\cos(\phi - \Delta\phi)}{\cos(\phi)} w(\phi - \Delta\phi, z + \Delta z) \right] \end{aligned} \quad (20)$$

and vice versa

$$400 \quad -\frac{1}{2} \left[w(\phi, z - \Delta z) \frac{\Delta t_p}{\Delta z} \cdot \frac{\Delta t_p}{\Delta \phi} \cdot \frac{(r - \Delta z)^2}{r^2} \cdot \right. \\ \left. \cdot \frac{\cos(\phi + \Delta \phi)}{\cos(\phi)} \cdot \frac{v(\phi + \Delta \phi, z - \Delta z)}{r - \Delta z} \right] \quad 428 \quad (21)$$

where i is the index of the time increment, and where $\phi \pm \Delta\phi$ and $z \pm \Delta z$ refer to the adjacent model gridpoints in latitude and altitude, respectively.

405 For mixing ratios, the respective derivatives are:

$$\begin{aligned}
& \frac{\partial vmr_{i+1}(\phi, z)}{\partial vmr_i(\phi, z)} = \\
& 1 - \left(\frac{\Delta t_p}{\Delta \phi} \right)^2 \cdot \frac{v(\phi, z)}{r^2} \cdot \frac{1}{2} [v(\phi, z) + v(\phi - \Delta \phi, z)] \\
& - \left(\frac{\Delta t_p}{\Delta z} \right)^2 \cdot w(\phi, z) \cdot \frac{1}{2} [w(\phi, z) + w(\phi, z - \Delta z)] \\
& - \frac{\Delta t_p}{2r^2(\Delta \phi)^2 \cos(\phi)} \left[\left(K_\phi(\phi, z) + K_\phi(\phi + \Delta \phi, z) \right) \right. \\
& \cdot \cos\left(\phi + \frac{\Delta \phi}{2}\right) \\
& + \left. \left(K_\phi(\phi, z) + K_\phi(\phi - \Delta \phi, z) \right) \cos\left(\phi - \frac{\Delta \phi}{2}\right) \right] \\
& - \frac{\Delta t_p}{2r^2(\Delta z)^2} \left[\left(r + \frac{\Delta z}{2} \right)^2 \left(K_z(\phi, z) + K_z(\phi, z + \Delta z) \right) \right. \\
& + \left. \left(r - \frac{\Delta z}{2} \right)^2 \left(K_z(\phi, z) + K_z(\phi, z - \Delta z) \right) \right] \\
& - Loss(month, \phi, z) \Delta t_p;
\end{aligned} \tag{22}$$

$$\begin{aligned}
& \frac{\partial vmr_{i+1}(\phi, z)}{\partial vmr_i(\phi + \Delta\phi, z)} = \\
& -\frac{\Delta t_p}{\Delta\phi} \cdot \frac{v(\phi, z)}{2r} \left(1 - \frac{\Delta t_p}{\Delta\phi} \cdot \frac{v(\phi, z)}{r} \right) \\
& + \frac{\Delta t_p}{2r^2(\Delta\phi)^2 \cos(\phi)} \cdot \\
& \left(K_\phi(\phi, z) + K_\phi(\phi + \Delta\phi, z) \right) \cos\left(\phi + \frac{\Delta\phi}{2}\right); \tag{23}
\end{aligned}$$

$$\begin{aligned}
 & \frac{\partial v_{mr_{i+1}}(\phi, z)}{\partial v_{mr_i}(\phi - \Delta\phi, z)} = \\
 & \frac{v(\phi, z)}{2r} \cdot \frac{\Delta t_p}{\Delta\phi} \left(1 + \frac{\Delta t_p}{\Delta\phi} \cdot \frac{v(\phi - \Delta\phi, z)}{r} \right) \\
 & + \frac{\Delta t_p}{2r^2(\Delta\phi)^2 \cos(\phi)} \cdot \\
 & \left(K_\phi(\phi, z) + K_\phi(\phi - \Delta\phi, z) \right) \cos\left(\phi - \frac{\Delta\phi}{2}\right); \tag{24}
 \end{aligned}$$

$$\begin{aligned} \frac{\partial vmr_{i+1}(\phi, z)}{\partial vmr_i(\phi, z + \Delta z)} &= \\ -\frac{1}{2} \cdot w(\phi, z) \cdot \frac{\Delta t_p}{\Delta z} \left(1 - \frac{\Delta t_p}{\Delta z} w(\phi, z) \right) \\ + \frac{\Delta t_p}{2r^2(\Delta z)^2} \left(r + \frac{\Delta z}{2} \right)^2 \left(K_z(\phi, z) + K_z(\phi, z + \Delta z) \right); \end{aligned} \quad (25)$$

$$\begin{aligned} \frac{\partial v_{mr_{i+1}}(\phi, z)}{\partial v_{mr_i}(\phi, z - \Delta z)} &= \\ \frac{\Delta t_p}{\Delta z} \cdot \frac{1}{2} \cdot w(\phi, z) \left(1 + w(\phi, z - \Delta z) \frac{\Delta t_p}{\Delta z} \right) \\ + \frac{\Delta t_p}{2r^2(\Delta z)^2} \left(r - \frac{\Delta z}{2} \right)^2 \left(K_z(\phi, z) + K_z(\phi, z - \Delta z) \right), \end{aligned} \quad (26)$$

where $\text{Loss}(\text{month}, \phi, z)$ is the relative loss rate in the respective month at latitude ϕ and altitude z . These derivatives are simplifications in a sense that they do not consider the full chemical Jacobian but assume instead that the source strength depends on no other concentration than the actual concentration of the same species. For the typical long-lived so-called tropospheric source gases considered here, like SF₆ or CFCs, this assumption is appropriate. Pretabulated loss rates are used which have been calculated by locally integrating loss rates over an entire month at a time resolution adequate to resolve the diurnal cycle. From the monthly losses, the $\text{Loss}(\text{month}, \phi, z)$ values, which are the contribution of losses to the partial derivatives of the local mixing ratios with respect to the initial local mixing ratios, are calculated as the secant of the local decay curve.

With these expressions, the prediction of air density and volume mixing ratio can be rewritten in matrix notation for a single micro time increment. This notation simplifies the

estimation of the uncertainties of the predicted atmospheric state and the inversion of the prediction equation. In matrix notation, the prediction reads

$$\begin{aligned} \rho_{i+1} &= \begin{pmatrix} \rho_{I;k=1,K_0} \\ \rho_{i+1;k=1,K_0} \\ \rho_{i+1;j=K_0+1,L_0} \end{pmatrix} = \mathbf{D}_{\rho;i} \rho_i = \\ &= \begin{pmatrix} \mathbf{I}_K & \mathbf{0} & \mathbf{0} \\ \mathbf{W}_i & \mathbf{0} \\ \mathbf{0} & \mathbf{D}_{\rho,\text{nom}} \end{pmatrix} \begin{pmatrix} \rho_{I;k=1,K_0} \\ \rho_{i;k=1,K_0} \\ \rho_{i;j=K_0+1,L_0} \end{pmatrix} \end{aligned} \quad (27)$$

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where
 $\mathbf{D}_{\rho;i}$

is the $(L_0 + K_0) \times (L_0 + K_0)$ Jacobian matrix of air density for time increment i , i.e. the sensitivities of the prediction with respect to the initial state, $\frac{\partial c_{i+1,m}}{\partial c_{i,n}}$, here m and n run over the model gridpoints

is $K_0 \times K_0$ identity matrix;

are zero submatrices of the required dimensions;

\mathbf{W}_i is a $K_0 \times 2K_0$ -dimensional interpolation matrix;

$\mathbf{D}_{\rho,\text{nom}}$ is an $J_0 \times L_0$ Jacobian containing the partial derivatives $\partial \rho_{i+1;j} / \partial \rho_{i;l}$, applied to the nominal altitude range;

$\rho_{I;k=1,K_0}$ is the K_0 -dimensional vector of air densities in the border region after the final timestep, i.e. for the time of the next measurement;

$\rho_{i;k=1,K_0}$ is the K_0 -dimensional vector of air densities in the border region at the current timestep as resulting from interpolation in time;

$\rho_{i;j=K_0+1,L_0}$ is the J_0 -dimensional vector of air densities in the nominal region at the current timestep as resulting from integration according to the MacCormack scheme as described above.

The operation of these sub-matrices is illustrated in the uppermost three (violet, green, blue) blocks of Figure 1.

460 Since the source term depends on air density, the integration in matrix notation for vmr requires simultaneous treatment of vmr and air density, and we get, using notation ac-

cordant with air density:

$$\begin{aligned} \begin{pmatrix} \rho_{i+1} \\ \mathbf{vmr}_{i+1} \end{pmatrix} &= \begin{pmatrix} \rho_{I;k=1,K_0} \\ \rho_{i+1;k=1,K_0} \\ \rho_{i+1;j=K_0+1,L_0} \\ \mathbf{vmr}_{g;I;k=1,K_g} \\ \mathbf{vmr}_{g;i+1;k=1,K_g} \\ \mathbf{vmr}_{g;i+1;j=K_g+1,L_g} \end{pmatrix} = \\ &= \mathbf{D}_i \begin{pmatrix} \rho_i \\ \mathbf{vmr}_{g;i} \end{pmatrix} = \\ &= \begin{pmatrix} \mathbf{D}_{\rho;i} & \mathbf{0} & \mathbf{0} & \mathbf{0} \\ \mathbf{0} & \mathbf{I}_K & \mathbf{0} & \mathbf{0} \\ \mathbf{0} & \mathbf{W}_i & \mathbf{0} \\ \mathbf{0} & \mathbf{D}_{g,\text{nom}} \end{pmatrix} \begin{pmatrix} \rho_{I,k=1,K_0} \\ \rho_{i;k=1,K_0} \\ \rho_{i;j=K_0+1,L_0} \\ \mathbf{vmr}_{g;I,k=1,K_g} \\ \mathbf{vmr}_{g;i,k=1,K_g} \\ \mathbf{vmr}_{g;i,j=K_g+1,L_g} \end{pmatrix} \end{aligned} \quad (28)$$

where \mathbf{D}_i is the total Jacobian with respect to air densities and all involved gas mixing ratios, and where g runs over all gases. Note that

1. The Jacobian \mathbf{D}_i is time-dependent because it includes submatrices controlling the interpolation between the initial time and the end time. In the case of vmr, a further time dependence is introduced by the time-dependent source function.
2. the first ‘row’ of the Jacobian matrix includes identity \mathbf{I}_K because the prediction is not supposed to change the measured ρ_I and \mathbf{vmr}_I at the end of the macro time increment. This value is used to construct the boundary condition. Row is here written in quotes because the elements of this ‘row’ are matrices in themselves. Introduction of unity Jacobian elements is necessary because Eqs. (27–28) are autonomized, originally non-autonomous systems of differential equations.
3. \mathbf{W} is used to interpolate the boundary state between the initial time of the micro time interval, $t + (i-1)\Delta t_p$, and the time at the end of the time interval $t + \Delta t$ to give the atmospheric state at the border region at time $t + i\Delta t_p$.
4. The Jacobian submatrices $\mathbf{D}_{\rho,\text{nom}}$ and $\mathbf{D}_{g,\text{nom}}$ are used to predict the atmospheric state in the nominal range after one further micro time increment from the atmospheric state at the current time and the boundary condition. Its elements are described in Eqs. (15–21) and (22–26). The part of $\mathbf{D}_{g,\text{nom}}$ which refers to the border mixing ratios ($\mathbf{vmr}_{g;I,k=1,K_g}$) is zero. The dimension of $\mathbf{D}_{\rho,\text{nom}}$ is $\sum_g J_g \times (L_0 + K_0 + \sum_g (L_g + K_g))$
5. No simple mapping mechanism between the field of atmospheric state variables sampled at latitudes and altitudes and the vectors ρ and \mathbf{vmr}_g is provided because the fields are irregular in a sense that the number of relevant altitudes is latitude-dependent. Pointer variables have to be used instead.

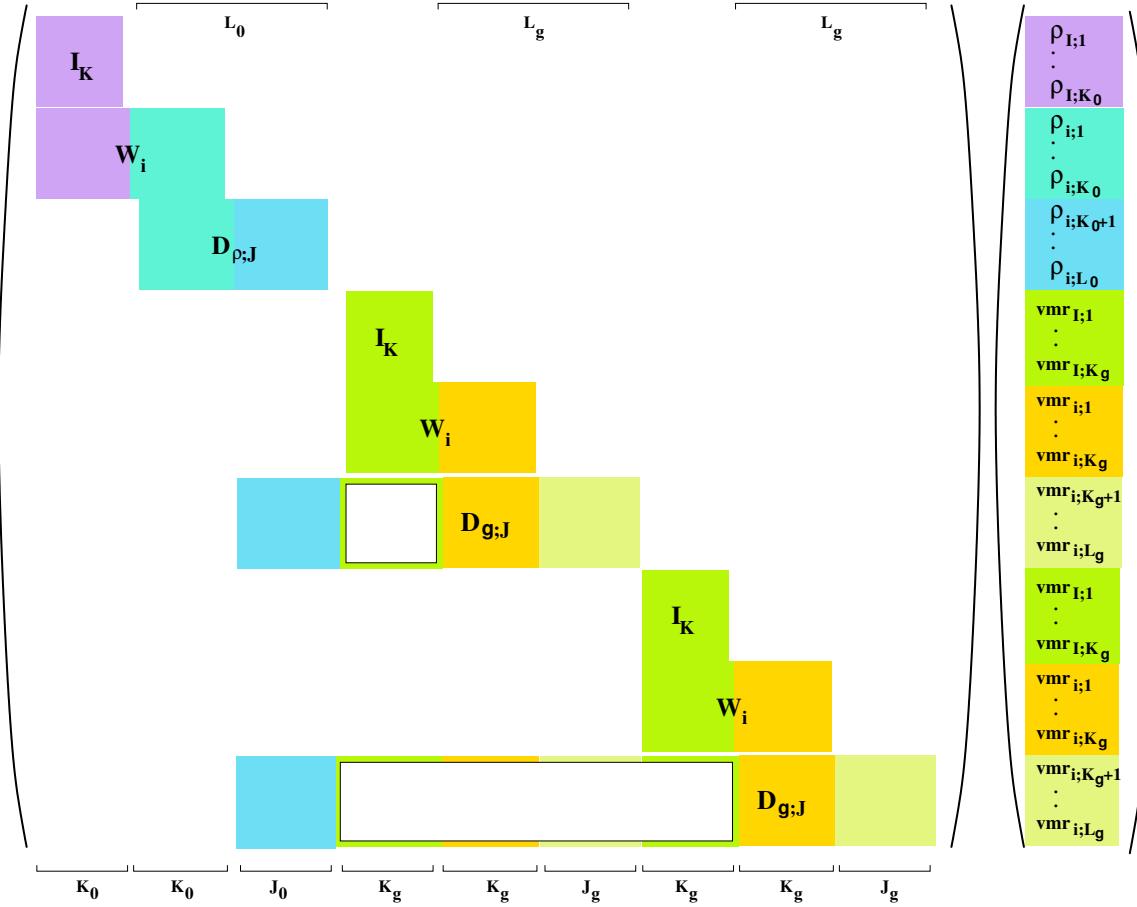


Figure 1. : Matrix structure of the right-hand part of Eq. 28. Colours indicate which matrix blocks operate on which part of the input vector. An example with two gases is shown.

The matrix structure is exemplified in Fig. 1. For the macro time increment Δt we get

$$\begin{pmatrix} \rho_I \\ vmr_I \end{pmatrix} = \left(\prod_{i=I}^1 \mathbf{D}_i \right) \begin{pmatrix} \rho_0 \\ vmr_0 \end{pmatrix} \quad (29)$$

505 3.4 Prediction errors

Let \mathbf{S}_0 be the $L \times L$ covariance matrix describing the uncertainties of all involved measurements ρ_0 and vmr_0 , with diagonal elements $s_{0,l,l} = \sigma_{0,l,l}^2$ and $L = \sum_0^{gases} L_g$. We assume that these measurement errors in the state variables used for the prediction are the only relevant error sources. With \mathbf{S}_0 and $\prod_{i=I}^1 \mathbf{D}_i$ available, generalized Gaussian error propagation for ρ_I and vmr_I can be easily formulated as:

$$\mathbf{S}_I = \left(\prod_{i=I}^1 \mathbf{D}_i \right) \mathbf{S}_0 \left(\prod_{i=I}^1 \mathbf{D}_i \right)^T. \quad (30)$$

Even if \mathbf{S}_0 is diagonal, i.e. the initial errors are assumed to be uncorrelated, error propagation through the forward model 525 will generate non-zero error covariances in \mathbf{S}_I representing

the atmospheric state at time $t + \Delta t$. \mathbf{S}_I will be needed in the inversion of circulation and mixing coefficients described in Section 4.

530 3.5 A note on finite resolution measurements

The measurements used are not a perfect image of the true atmospheric state but contain some prior information. In the case of the IMK data, a priori profiles are usually set zero, and the constraint is built with a Tikhonov-type first order finite differences smoothing constraint (c.f. von Clarmann et al. (2009)). That means that, besides the mapping of measurement and parameter errors, the only distortion of the truth via the retrieval is reduced altitude resolution; no other effect of the prior information is to be considered. Usually, any comparison between modelled and measured fields requires application of the averaging kernels of the retrieval to the model data in order to account for the smoothing by the constraint of the retrieval (assuming that the model grid is much finer than the resolution of the retrieval).

In our case, the situation is different: The model is initialized with measurements of reduced altitude resolution, and

the fields predicted by the model are then compared to measurements of the same altitude resolution. It is fair to assume that the model does not dramatically change the altitude resolution of the profiles, and thus comparable quantities are compared when the residuals between predicted and measured atmospheric state are evaluated.

3.6 A note on numerical mixing

Let the initial mixing ratio field be homogeneous except one point with delta-type excess mixing ratio. Assume further a homogeneous velocity field and zero mixing coefficients. If the velocity is such that the position of the excess mixing ratio is displaced during Δt by a distance which is not equal to an integer multiple of the gridwidth, then the resulting distribution will no longer be a delta-type distribution but will be smoothed. The widening of the delta peak we refer to as numerical mixing. The MacCormack transport scheme is designed to fight this diffusivity but some higher order effects may still survive. One might think that, during the inversion, the widening is misinterpreted as mixing, leading to too large mixing coefficients.

Again, in our case, the situation is different: The widening does not accumulate over the Δt_p timesteps, because we first calculate the operator $\prod_{i=1}^I \mathbf{D}_i$, which is applied only once to the initial field, which avoids accumulation of numerical mixing over timesteps. Still one widening process as described above can occur, when the forward model leads to a position of the new peak which cannot be represented in the grid chosen. However, since the gas distributions $vmr_{g;I}$ at the end of timestep Δt are sampled on the same grid, the maximum in the real atmosphere would be widened in the same way, and there would be no residual the inversion would try to get rid of by increasing the mixing coefficients. And the next time step Δt is initialized again with measured data, which also excludes accumulation of numerical mixing effects.

These considerations aside, there are other numerical artefacts. These are related to the numerical evaluation of partial derivatives of the state variables in our transport scheme chosen. Particularly in the case of delta functions in the state variable field, these cause side-wiggles behind and smearing in front of the transported structure. To keep these artefacts small, it is necessary to set the spatial grid fine enough that every structure is represented by multiple gridpoints.

4 The inverse problem

For convenience, we combine the variables of the initial atmospheric state and the predicted state at the end of the macro time interval, respectively, into the vectors

$$\tilde{\mathbf{x}}_0 = \begin{pmatrix} \rho_0 \\ vmr_0 \end{pmatrix} = (\tilde{x}_{0;1} \dots \tilde{x}_{0;L})^T, \quad (31)$$

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and

$$\tilde{\mathbf{x}}_I = \begin{pmatrix} \rho_I \\ vmr_I \end{pmatrix} = (\tilde{x}_{I;1} \dots \tilde{x}_{I;L})^T, \quad (32)$$

The related subsets of $\tilde{\mathbf{x}}_0$ and $\tilde{\mathbf{x}}_I$ which contain only state variables in the nominal altitude range but not those in the border region are $\mathbf{x}_0 = (x_{0;1} \dots x_{0;J})^T$ and $\mathbf{x}_I = (x_{I;1} \dots x_{I;J})^T$, respectively. The reason why the distinction between $\tilde{\mathbf{x}}$ and \mathbf{x} is made is that, contrary to the prediction step, for the inversion vector elements related to the interpolation of values in the border region are no longer needed. Further, we combine the fields of meridional and vertical wind components and mixing coefficients into the vector

$$\mathbf{q} = \begin{pmatrix} v \\ w \\ K_\phi \\ K_z \end{pmatrix}, \quad (33)$$

and assume constant velocities and mixing coefficients during the macro timestep. To infer circulation patterns and mixing coefficients from the measurements of air densities and mixing ratios, the Jacobian matrix \mathbf{F} ,

$$\mathbf{F} = (\mathbf{f}_1, \dots, \mathbf{f}_N) = (f_{j,n}) = \left(\frac{\partial x_{I,j}}{\partial q_n} \right) = \left(\frac{\partial \mathbf{x}_I}{\partial q_1}, \dots, \frac{\partial \mathbf{x}_I}{\partial q_N} \right), \quad (34)$$

is needed, where $N = 4J$ where $J = \sum_0^{gases} J_g$, because there are four unknown quantities, v_j , w_j , $K_{\phi;j}$, and $K_{z;j}$ at each gridpoint of the nominal region where these variables shall be inferred. The elements of \mathbf{F} are calculated from Eq. (28) by application of the product rule:

$$\tilde{\mathbf{f}}_n = \sum_{i=1}^I \left[\left(\prod_{k=1}^{i+1} \mathbf{D}_k \right) \left(\frac{\partial \mathbf{D}_i}{\partial q_n} \right) \left(\prod_{k=i-1}^1 \mathbf{D}_k \right) \tilde{\mathbf{x}}_0 \right], \quad (35)$$

where the tilde symbol in $\tilde{\mathbf{f}}_n$ indicates that the vectors resulting from Eq. (35) still include the border elements which have to be discarded to obtain \mathbf{f}_n . The quantity $\tilde{\mathbf{f}}_n$ is more efficiently computed using the following recursive scheme, where $\tilde{\mathbf{f}}_{l,i}$ is the respective column of the Jacobian after micro timestep i :

$$\tilde{\mathbf{f}}_{n,i} = \mathbf{D}_i \tilde{\mathbf{f}}_{n,i-1} + \frac{\partial \mathbf{D}_i}{\partial q_n} \left(\prod_{k=i-1}^1 \mathbf{D}_k \right) \tilde{\mathbf{x}}_0 \quad (36)$$

With the argument of \mathbf{D} specifying the column of the \mathbf{D} -matrix such that $D_{c,i}(\phi, z)$ relates $\rho_{i+1}(\phi, z)$ to $\rho_i(\phi, z)$, $D_{\rho,i}(\phi \pm \Delta\phi, z)$ relates $\rho_{i+1}(\phi, z)$ to $\rho_i(\phi \pm \Delta\phi, z)$, and $D_{\rho,i}(\phi, z \pm \Delta z)$ relates $\rho_{i+1}(\phi, z)$ to $\rho_i(\phi, z \pm \Delta z)$, and for vmr accordingly, the entries of \mathbf{D}_i relevant to \mathbf{v} are:

$$\begin{aligned} \frac{\partial \frac{\partial \rho_{i+1}(\phi, z)}{\partial \rho_i(\phi, z)}}{\partial v(\phi, z)} &= -\frac{\Delta t_p}{2\Delta\phi}. \\ &\cdot \left(\frac{\Delta t_p}{\Delta\phi} \cdot \frac{2v(\phi, z) + v(\phi - \Delta\phi, z)}{r^2} + 2\frac{\Delta t_p}{\Delta z} \cdot \frac{w(\phi, z)}{r} \right) \end{aligned} \quad (37)$$

$$\frac{\partial \frac{\partial \rho_{i+1}(\phi + \Delta\phi, z)}{\partial \rho_i(\phi + \Delta\phi, z)}}{\partial v(\phi, z)} = -\frac{1}{2} \cdot \left(\frac{\Delta t_p}{\Delta\phi} \right)^2 \cdot \frac{v(\phi + \Delta\phi, z)}{r^2} \quad (38)$$

$$\frac{\partial \frac{\partial \rho_{i+1}(\phi + \Delta\phi, z)}{\partial \rho_i(\phi, z)}}{\partial v(\phi, z)} = \frac{1}{2r} \cdot \frac{\Delta t_p}{\Delta\phi} \cdot \frac{\cos(\phi)}{\cos(\phi + \Delta\phi)} \cdot \left(1 + 2 \frac{\Delta t_p}{\Delta\phi} \cdot \frac{v(\phi, z)}{r} + \frac{\Delta t_p}{\Delta z} w(\phi, z) \right) \quad (39)$$

$$\frac{\partial \frac{\partial \rho_{i+1}(\phi, z)}{\partial \rho_i(\phi + \Delta\phi, z)}}{\partial v(\phi, z)} = \frac{1}{2} \cdot \left(\frac{\Delta t_p}{\Delta\phi} \right)^2 \cdot \frac{v(\phi + \Delta\phi, z)}{r^2} \cdot \frac{\cos(\phi + \Delta\phi)}{\cos(\phi)} \quad (40)$$

$$\frac{\partial \frac{\partial \rho_{i+1}(\phi - \Delta\phi, z)}{\partial \rho_i(\phi, z)}}{\partial v(\phi, z)} = \frac{1}{2r} \cdot \frac{\Delta t_p}{\Delta\phi} \cdot \frac{\cos(\phi)}{\cos(\phi - \Delta\phi)} \cdot \left(-1 + \frac{\Delta t_p}{\Delta\phi} \cdot \frac{v(\phi - \Delta\phi, z)}{r} + \frac{\Delta t_p}{\Delta z} w(\phi - \Delta\phi, z) \right) \quad (41)$$

$$\frac{\partial \frac{\partial \rho_{i+1}(\phi, z + \Delta z)}{\partial \rho_i(\phi, z)}}{\partial v(\phi, z)} = \frac{1}{2} \cdot \frac{\Delta t_p}{\Delta z} \cdot \frac{\Delta t_p}{\Delta\phi} \cdot \frac{r^2}{(r + \Delta z)^2} \frac{w(\phi, z)}{r} \quad (42)$$

$$\frac{\partial \frac{\partial \rho_{i+1}(\phi, z)}{\partial \rho_i(\phi, z + \Delta z)}}{\partial v(\phi, z)} = \frac{1}{2} \cdot \frac{\Delta t_p}{\Delta z} \cdot \frac{\Delta t_p}{\Delta\phi} \cdot \frac{(r + \Delta z)^2}{r^3} w(\phi, z + \Delta z) \quad (43)$$

$$\frac{\partial \frac{\partial \rho_{i+1}(\phi + \Delta\phi, z)}{\partial \rho_i(\phi, z + \Delta z)}}{\partial v(\phi, z)} = -\frac{1}{2} \cdot \frac{\Delta t_p}{\Delta\phi} \cdot \frac{\Delta t_p}{\Delta z} \cdot \frac{(r + \Delta z)^2}{r^3} \cdot \frac{\cos(\phi)}{\cos(\phi + \Delta\phi)} w(\phi, z + \Delta z) \quad (44)$$

$$\frac{\partial \frac{\partial \rho_{i+1}(\phi - \Delta\phi, z + \Delta z)}{\partial \rho_i(\phi, z)}}{\partial v(\phi, z)} = -\frac{1}{2} \cdot \frac{\Delta t_p}{\Delta\phi} \cdot \frac{\Delta t_p}{\Delta z} \cdot \frac{r}{(r + \Delta z)^2} \cdot \frac{\cos(\phi)}{\cos(\phi - \Delta\phi)} w(\phi - \Delta\phi, z) \quad (45)$$

$$\frac{\partial \frac{\partial vmr_{i+1}(\phi, z)}{\partial vmr_i(\phi, z)}}{\partial v(\phi, z)} = -\left(\frac{\Delta t_p}{\Delta\phi} \right)^2 \cdot \frac{1}{r^2} \cdot \left(v(\phi, z) + \frac{v(\phi - \Delta\phi, z)}{2} \right) \quad (46)$$

$$\frac{\partial \frac{\partial vmr_{i+1}(\phi + \Delta\phi, z)}{\partial vmr_i(\phi + \Delta\phi, z)}}{\partial v(\phi, z)} = -\left(\frac{\Delta t_p}{\Delta\phi} \right)^2 \cdot \frac{v(\phi + \Delta\phi, z)}{2r^2} \quad (47)$$

$$\frac{\partial \frac{\partial vmr_{i+1}(\phi, z)}{\partial vmr_i(\phi - \Delta\phi, z)}}{\partial v(\phi, z)} = \frac{1}{2r} \cdot \frac{\Delta t_p}{\Delta\phi} \cdot \left(1 + \frac{\Delta t_p}{\Delta\phi} \cdot \frac{v(\phi - \Delta\phi, z)}{r} \right) \quad (48)$$

$$\frac{\partial \frac{\partial vmr_{i+1}(\phi + \Delta\phi, z)}{\partial vmr_i(\phi, z)}}{\partial v(\phi, z)} = \frac{1}{2r} \left(\frac{\Delta t_p}{\Delta\phi} \right)^2 \frac{v(\phi + \Delta\phi, z)}{r} \quad (49)$$

$$\frac{\partial \frac{\partial vmr_{i+1}(\phi, z)}{\partial vmr_i(\phi + \Delta\phi, z)}}{\partial v(\phi, z)} = -\frac{\Delta t_p}{\Delta\phi} \cdot \frac{1}{r} \left(\frac{1}{2} - \frac{\Delta t_p}{\Delta\phi} \cdot \frac{v(\phi, z)}{r} \right) \quad (50)$$

Entries not mentioned here are zero. Entries relevant to w are:

$$\begin{aligned} \frac{\partial \frac{\partial \rho_{i+1}(\phi, z)}{\partial \rho_i(\phi, z)}}{\partial w(\phi, z)} &= -\frac{\Delta t_p}{\Delta\phi} \cdot \frac{\Delta t_p}{\Delta z} \cdot \frac{v(\phi, z)}{r} \\ &\quad - \left(\frac{\Delta t_p}{\Delta z} \right)^2 w(\phi, z) - \frac{1}{2} \left(\frac{\Delta t_p}{\Delta z} \right)^2 w(\phi, z - \Delta z) \end{aligned} \quad (51)$$

$$\frac{\partial \frac{\partial \rho_{i+1}(\phi, z + \Delta z)}{\partial \rho_i(\phi, z + \Delta z)}}{\partial w(\phi, z)} = -\frac{1}{2} \left(\frac{\Delta t_p}{\Delta z} \right)^2 w(\phi, z + \Delta z) \quad (52)$$

$$\begin{aligned} \frac{\partial \frac{\partial \rho_{i+1}(\phi, z)}{\partial \rho_i(\phi + \Delta\phi, z)}}{\partial w(\phi, z)} &= \frac{1}{2} \cdot \frac{\Delta t_p}{\Delta\phi} \cdot \frac{\Delta t_p}{\Delta z} \cdot \frac{v(\phi + \Delta\phi, z)}{r} \cdot \frac{\cos(\phi + \Delta\phi)}{\cos(\phi)} \\ &\quad - \left(\frac{\Delta t_p}{\Delta z} \right)^2 w(\phi, z + \Delta z) \end{aligned} \quad (53)$$

$$\begin{aligned} \frac{\partial \frac{\partial \rho_{i+1}(\phi + \Delta\phi, z)}{\partial \rho_i(\phi + \Delta\phi, z - \Delta z)}}{\partial w(\phi, z)} &= \frac{1}{2} \cdot \frac{\Delta t_p}{\Delta\phi} \cdot \frac{\Delta t_p}{\Delta z} \cdot \frac{v(\phi, z)}{r} \cdot \frac{\cos(\phi)}{\cos(\phi + \Delta\phi)} \\ &\quad - \left(\frac{\Delta t_p}{\Delta z} \right)^2 w(\phi, z - \Delta z) \end{aligned} \quad (54)$$

$$\begin{aligned} \frac{\partial \frac{\partial \rho_{i+1}(\phi, z + \Delta z)}{\partial \rho_i(\phi, z)}}{\partial w(\phi, z)} &= \frac{1}{2} \cdot \frac{\Delta t_p}{\Delta z} \cdot \frac{r^2}{(r + \Delta z)^2} \\ &\quad \cdot \left(1 + \frac{\Delta t_p}{\Delta\phi} \cdot \frac{v(\phi, z)}{r} + 2 \frac{\Delta t_p}{\Delta z} \cdot w(\phi, z) \right) \end{aligned} \quad (55)$$

$$\begin{aligned} \frac{\partial \frac{\partial \rho_{i+1}(\phi, z)}{\partial \rho_i(\phi, z + \Delta z)}}{\partial w(\phi, z)} &= \frac{1}{2} \cdot \left(\frac{\Delta t_p}{\Delta z} \right)^2 \cdot \frac{(r + \Delta z)^2}{r^2} \cdot w(\phi, z + \Delta z) \\ &\quad - \left(\frac{\Delta t_p}{\Delta z} \right)^2 w(\phi, z - \Delta z) \end{aligned} \quad (56)$$

$$\begin{aligned} \frac{\partial \frac{\partial \rho_{i+1}(\phi, z - \Delta z)}{\partial \rho_i(\phi, z)}}{\partial w(\phi, z)} &= \frac{1}{2} \cdot \frac{\Delta t_p}{\Delta z} \cdot \frac{r^2}{(r - \Delta z)^2} \\ &\quad \cdot \left(-1 + \frac{\Delta t_p}{\Delta\phi} \cdot \frac{v(\phi, z - \Delta z)}{r - \Delta z} + \frac{\Delta t_p}{\Delta z} \cdot w(\phi, z - \Delta z) \right) \end{aligned} \quad (57)$$

$$\frac{\partial \frac{\partial \rho_{i+1}(\phi + \Delta\phi, z - \Delta z)}{\partial \rho_i(\phi, z)}}{\partial w(\phi, z)} = -\frac{1}{2} \cdot \frac{\Delta t_p}{\Delta\phi} \cdot \frac{\Delta t_p}{\Delta z}. \quad (58)$$

$$\cdot \frac{r^2}{(r - \Delta z)^2} \cdot \frac{\cos(\phi)}{\cos(\phi + \Delta\phi)} \frac{v(\phi, z - \Delta z)}{r - \Delta z}$$

$$\frac{\partial \frac{\partial \rho_{i+1}(\phi, z + \Delta z)}{\partial \rho_i(\phi + \Delta\phi, z)}}{\partial w(\phi)} = -\frac{1}{2} \cdot \frac{\Delta t_p}{\Delta\phi} \cdot \frac{\Delta t_p}{\Delta z}. \quad (59)$$

$$\cdot \frac{r^2}{(r + \Delta z)^2} \cdot \frac{\cos(\phi + \Delta\phi)}{\cos(\phi)} \cdot \frac{v(\phi + \Delta\phi, z)}{r}$$

$$\frac{\partial \frac{\partial vmr_{i+1}(\phi, z)}{\partial vmr_i(\phi, z)}}{\partial w(\phi, z)} = \quad (60)$$

$$-\left(\frac{\Delta t_p}{\Delta z}\right)^2 \cdot \left(w(\phi, z) + \frac{w(\phi, z - \Delta z)}{2}\right)$$

$$\frac{\partial \frac{\partial vmr_{i+1}(\phi, z + \Delta z)}{\partial vmr_i(\phi, z + \Delta z)}}{\partial w(\phi, z)} = -\left(\frac{\Delta t_p}{\Delta z}\right)^2 \cdot \frac{w(\phi, z + \Delta z)}{2} \quad (61)$$

$$\frac{\partial \frac{\partial vmr_{i+1}(\phi, z)}{\partial vmr_i(\phi, z - \Delta z)}}{\partial w(\phi, z)} = \quad (62)$$

$$\frac{1}{2} \cdot \frac{\Delta t_p}{\Delta z} \cdot \left(1 + w(\phi, z - \Delta z) \cdot \frac{\Delta t_p}{\Delta z}\right)$$

$$\frac{\partial \frac{\partial vmr_{i+1}(\phi, z + \Delta z)}{\partial vmr_i(\phi, z)}}{\partial w(\phi, z)} = \frac{1}{2} \cdot \left(\frac{\Delta t_p}{\Delta z}\right)^2 \cdot w(\phi, z + \Delta z) \quad (63)$$

$$\frac{\partial \frac{\partial vmr_{i+1}(\phi, z)}{\partial vmr_i(\phi, z + \Delta z)}}{\partial w(\phi, z)} = \quad (64)$$

$$-\frac{1}{2} \cdot \frac{\Delta t_p}{\Delta z} \cdot \left(1 - 2 \frac{(\Delta t_p)}{(\Delta z)} \cdot w(\phi, z)\right)$$

Entries relevant to K_ϕ are:

$$\frac{\partial \frac{\partial vmr_{i+1}(\phi, z)}{\partial K_\phi(\phi, z)}}{\partial K_\phi(\phi, z)} = -\frac{\Delta t_p}{(\Delta\phi)^2} \cdot \frac{1}{2r^2}. \quad (65)$$

$$\cdot \frac{\cos\left(\phi + \frac{\Delta\phi}{2}\right) + \cos\left(\phi - \frac{\Delta\phi}{2}\right)}{\cos(\phi)}$$

$$\frac{\partial \frac{\partial vmr_{i+1}(\phi \mp \Delta\phi, z)}{\partial K_\phi(\phi, z)}}{\partial K_\phi(\phi, z)} = -\frac{\Delta t_p}{(\Delta\phi)^2} \cdot \frac{1}{2r^2} \cdot \frac{\cos\left(\phi \mp \frac{\Delta\phi}{2}\right)}{\cos(\phi \mp \Delta\phi)} \quad (66)$$

$$\frac{\partial \frac{\partial vmr_{i+1}(\phi, z)}{\partial K_\phi(\phi - \Delta\phi, z)}}{\partial K_\phi(\phi, z)} = \frac{1}{2r^2} \cdot \frac{\Delta t_p}{(\Delta\phi)^2} \cdot \frac{\cos(\phi - \frac{\Delta\phi}{2})}{\cos(\phi)} \quad (67)$$

$$\frac{\partial \frac{\partial vmr_{i+1}(\phi, z)}{\partial K_\phi(\phi, z)}}{\partial K_\phi(\phi, z)} = \frac{1}{2r^2} \cdot \frac{\Delta t_p}{(\Delta\phi)^2} \cdot \frac{\cos(\phi + \frac{\Delta\phi}{2})}{\cos(\phi)} \quad (68)$$

$$\frac{\partial \frac{\partial vmr_{i+1}(\phi + \Delta\phi, z)}{\partial K_\phi(\phi, z)}}{\partial K_\phi(\phi, z)} = \frac{1}{2r^2} \cdot \frac{\Delta t_p}{(\Delta\phi)^2} \cdot \frac{\cos(\phi + \frac{\Delta\phi}{2})}{\cos(\phi + \Delta\phi)} \quad (69)$$

$$\frac{\partial \frac{\partial vmr_{i+1}(\phi - \Delta\phi, z)}{\partial K_\phi(\phi, z)}}{\partial K_\phi(\phi, z)} = \frac{1}{2r^2} \cdot \frac{\Delta t_p}{(\Delta\phi)^2} \cdot \frac{\cos(\phi - \frac{\Delta\phi}{2})}{\cos(\phi - \Delta\phi)} \quad (70)$$

And finally, entries relevant to K_z are:

$$\frac{\partial \frac{\partial vmr_{i+1}(\phi, z)}{\partial K_z(\phi, z)}}{\partial K_z(\phi, z)} = -\frac{\Delta t_p}{(\Delta z)^2} \cdot \frac{(r + \frac{\Delta z}{2})^2 + (r - \frac{\Delta z}{2})^2}{2r^2} \quad (71)$$

$$\frac{\partial \frac{\partial vmr_{i+1}(\phi, z \mp \Delta z)}{\partial K_z(\phi, z)}}{\partial K_z(\phi, z)} = -\frac{\Delta t_p}{(\Delta z)^2} \cdot \frac{(r \mp \frac{\Delta z}{2})^2}{2(r \mp \Delta z)^2} \quad (72)$$

$$\frac{\partial \frac{\partial vmr_{i+1}(\phi, z + \Delta z)}{\partial K_z(\phi, z)}}{\partial K_z(\phi, z)} = \frac{1}{2} \cdot \frac{\Delta t_p}{(\Delta z)^2} \cdot \frac{(r + \frac{\Delta z}{2})^2}{(r + \Delta z)^2} \quad (73)$$

$$\frac{\partial \frac{\partial vmr_{i+1}(\phi, z - \Delta z)}{\partial K_z(\phi, z)}}{\partial K_z(\phi, z)} = \frac{1}{2} \cdot \frac{\Delta t_p}{(\Delta z)^2} \cdot \frac{(r - \frac{\Delta z}{2})^2}{r^2} \quad (74)$$

$$\frac{\partial \frac{\partial vmr_{i+1}(\phi, z)}{\partial K_z(\phi, z)}}{\partial K_z(\phi, z)} = \frac{1}{2} \cdot \frac{\Delta t_p}{(\Delta z)^2} \cdot \frac{(r + \frac{\Delta z}{2})^2}{r^2} \quad (75)$$

$$\frac{\partial \frac{\partial vmr_{i+1}(\phi, z - \Delta z)}{\partial K_z(\phi, z)}}{\partial K_z(\phi, z)} = \frac{1}{2} \cdot \frac{\Delta t_p}{(\Delta z)^2} \cdot \frac{(r - \frac{\Delta z}{2})^2}{(r - \Delta z)^2} \quad (76)$$

With these derivatives we linearize the prediction with respect to wind and mixing coefficients, i.e., we linearly predict the new atmospheric state for a given initial state \mathbf{q}_0 as a function of wind and mixing ratios.

$$\mathbf{x}_I = \mathbf{x}_0 + \mathbf{F}(\mathbf{q} - \mathbf{q}_0). \quad (77)$$

This formulation gives access to the winds and mixing ratios via inversion of \mathbf{F} .

Assuming linearity and Gaussian statistics, the most likely set \mathbf{q} of winds and mixing ratios during the macro time interval minimizes the following cost function:

$$\begin{aligned} \chi_1^2 &= (\mathbf{x}_{m;I} - \mathbf{x}_I)^T \mathbf{S}_r^{-1} (\mathbf{x}_{m;I} - \mathbf{x}_I) \\ &\approx (\mathbf{x}_{m;I} - \mathbf{x}_0 - \mathbf{F}(\mathbf{q} - \mathbf{q}_0))^T \mathbf{S}_r^{-1} \\ &\quad (\mathbf{x}_{m;I} - \mathbf{x}_0 - \mathbf{F}(\mathbf{q} - \mathbf{q}_0)) \end{aligned} \quad (78)$$

where $\mathbf{x}_{m;I}$ is the measured state at the end of the macro time step and \mathbf{S}_r is the error covariance matrix of the residual, which is, under the assumption that prediction error and

measurement errors are uncorrelated, the sum of the prediction covariance matrix and the measurement covariance matrix, both after the macro time step:

$$\mathbf{S}_r = \mathbf{S}_{m,I} + \mathbf{S}_p, \quad (79)$$

where \mathbf{S}_p is an $J \times J$ -matrix containing those elements of \mathbf{S}_I which are relevant to \mathbf{x}_I . $\mathbf{S}_{m,I}$ is the measurement error covariance matrix of the atmospheric state after the macro time step. The minimization of the cost function gives the following estimate $\hat{\mathbf{q}}$ of winds and mixing coefficients:

$$\hat{\mathbf{q}} = \mathbf{q}_0 + (\mathbf{F}^T \mathbf{S}_r^{-1} \mathbf{F})^{-1} \mathbf{F}^T \mathbf{S}_r^{-1} (\mathbf{x}_{m;I} - \mathbf{x}_I) \quad (80)$$

The matrix $\mathbf{F}^T \mathbf{S}_r^{-1} \mathbf{F}$ can be singular either because the related system of equations is under-determined or ill-posed due to nearly linearly dependent equations. Singularity is fought by adding the following constraint term to the cost function of Eq. (78):

$$\chi_{\text{con}}^2 = (\mathbf{q} - \mathbf{q}_a)^T \mathbf{R} (\mathbf{q} - \mathbf{q}_a) \quad (81)$$

$$\chi^2 = \chi_1^2 + \chi_{\text{con}}^2, \quad (82)$$

where \mathbf{q}_a is some prior assumption on velocities and mixing coefficients. \mathbf{R} is a $J \times J$ regularization matrix of which the choice is discussed below. From this, the constrained estimate of velocities and mixing coefficients can be inferred:

$$\hat{\mathbf{q}} = \mathbf{q}_a + (\mathbf{F}^T \mathbf{S}_r^{-1} \mathbf{F} + \mathbf{R})^{-1} \mathbf{F}^T \mathbf{S}_r^{-1} (\mathbf{x}_{m;I} - \mathbf{x}_I) \quad (83)$$

An equivalent formulation, which is more efficient if the dimension of \mathbf{q} is larger than that of \mathbf{x} (underdetermined problem), but which requires a non-singular regularization matrix and does not give easy access to diagnostics (see below), is (Rodgers, 2000):

$$\hat{\mathbf{q}} = \mathbf{q}_a + \mathbf{R}^{-1} \mathbf{F}^T (\mathbf{F} \mathbf{R}^{-1} \mathbf{F}^T + \mathbf{S}_{r,I})^{-1} (\mathbf{x}_{m;I} - \mathbf{x}_I). \quad (84)$$

The covariance matrix characterizing the uncertainty of estimated winds and mixing coefficients is

$$\mathbf{S}_q = (\mathbf{F}^T \mathbf{S}_r^{-1} \mathbf{F} + \mathbf{R})^{-1} \mathbf{F}^T \mathbf{S}_r^{-1} \mathbf{F} (\mathbf{F}^T \mathbf{S}_r^{-1} \mathbf{F} + \mathbf{R})^{-1}, \quad (85)$$

and the estimated winds and mixing coefficients are related to the true ones as

$$\mathbf{A} = \frac{\partial \hat{\mathbf{q}}}{\partial \mathbf{q}} = (\mathbf{F}^T \mathbf{S}_r^{-1} \mathbf{F} + \mathbf{R})^{-1} \mathbf{F}^T \mathbf{S}_r^{-1} \mathbf{F}, \quad (86)$$

which is unity in the case of unconstrained estimation of \mathbf{q} . In the case of Newtonian iteration, Eqs. (85–86) are evaluated using the Jacobian \mathbf{F} valid at the solution.

Due to the concentration-dependence of the source function and the q -dependence of \mathbf{F} , Eq. (29) is valid only in linear approximation. This is helped by putting the inversion

in the context of a Newtonian iteration (see, e.g., Rodgers (2000, p. 85). Eq. (80) becomes

$$\hat{\mathbf{q}}_{it+1} = \mathbf{q}_{it} + (\mathbf{F}_{it}^T \mathbf{S}_r^{-1} \mathbf{F}_{it})^{-1} \mathbf{F}_{it}^T \mathbf{S}_r^{-1} (\mathbf{x}_{m;I} - \mathbf{x}_{I,it}), \quad (87)$$

where it is the iteration index. Equation (83) becomes

$$\begin{aligned} \hat{\mathbf{q}}_{it+1} = & \mathbf{q}_{it} + (\mathbf{F}_{it}^T \mathbf{S}_r^{-1} \mathbf{F}_{it} + \mathbf{R})^{-1} \\ & (\mathbf{F}_{it}^T \mathbf{S}_r^{-1} (\mathbf{x}_{m;I} - \mathbf{x}_{I,it}) - \mathbf{R}(\mathbf{q}_{it} - \mathbf{q}_a)) \end{aligned} \quad (88)$$

or alternatively

$$\begin{aligned} \hat{\mathbf{q}}_{it+1} = & \mathbf{q}_{it} + (\mathbf{F}_{it}^T \mathbf{S}_r^{-1} \mathbf{F}_{it} + \mathbf{R})^{-1} \\ & \mathbf{F}_{it}^T \mathbf{S}_r^{-1} (\mathbf{x}_{m;I} - \mathbf{x}_{I,it} + \mathbf{F}_{it}(\mathbf{q}_{it} - \mathbf{q}_a)), \end{aligned} \quad (89)$$

and Eq. (84) becomes

$$\begin{aligned} \hat{\mathbf{q}}_{it+1} = & \mathbf{q}_a + \mathbf{R}^{-1} \mathbf{F}_{it}^T (\mathbf{F}_{it} \mathbf{R}^{-1} \mathbf{F}_{it}^T + \mathbf{S}_{r,I})^{-1} \\ & (\mathbf{x}_{m;I} - \mathbf{x}_{I,it} + \mathbf{F}_{it}(\mathbf{q}_{it} - \mathbf{q}_a)). \end{aligned} \quad (90)$$

With $\mathbf{q}_a = \mathbf{0}$ and diagonal $\mathbf{R} = \gamma \mathbf{I}$ we get the smallest possible velocities and mixing coefficients still consistent with the measurement, where tuning parameter γ will be set depending on how large fit residuals the user still considers to be ‘consistent’. With \mathbf{R} being diagonally blockwise composed of squared and scaled first order finite differences operators and $\mathbf{q}_a = \mathbf{0}$, smooth fields of wind vectors and mixing coefficients can be enforced. Setting \mathbf{q}_a the result of the previous macro time step corresponds to sequential data assimilation. In this application \mathbf{R} is set to the reciprocal uncertainty of \mathbf{q}_a plus some margin for allowance of variability of velocity and mixing coefficients in time. And finally, if prior knowledge is formed by independent measurements and their reciprocal uncertainties as constraint matrix, or within the debatable framework of Bayesian statistics, estimates $\hat{\mathbf{q}}$ would even be the most probable estimate of velocities and mixing ratios.

5 Proof of concept

5.1 Prediction of the atmospheric state

In a first step we test the predictive power of the formalism defined by Eqs. (3–29). Since the formalism itself is deductive and starts from a well established theoretical concept, the purpose of the test is solely to verify that the implementation of the formalism is correct and that involved numerical approximations are adequate. As a consequence of the Bonini paradox (c.f. Bonini (1963) and Starbuck (1975)), a model is the harder to understand the more complex it is. While the predictive power of a model usually increases with complexity, this does not necessarily hold for its explanatory power. Thus we have decided to test our model on the basis of very simple test cases, where major failure of the model is immediately obvious. Four test cases have been chosen, each dedicated to one kinematic variable (v , w , K_ϕ and K_z), while the other three were set to zero.

825 In the first case, v was set to a tenth of the Courant limit (Courant et al., 1928) (about $0.17/\cos(\phi)$ ms $^{-1}$) everywhere. As one would expect from the continuity equation applied to a spherical atmosphere, no changes in air 830 density except boundary effects at the poles were observed, and structures near the equator were transported by about 4° within a month, as expected from the equation of motion. A Gaussian-shaped perturbation of a halfwidth of one latitudinal gridwidth (4°) causes an upwind wiggle of less than 0.7% of the amplitude of the perturbation at a meridional velocity 835 of one gridpoint per month. There is no discernable change in the width of the transported structure. Similarly, for the second case a constant field of w of 1.1×10^{-3} ms $^{-1}$ lifts a structure upwards by about 3 km per month. Mixing coefficients were verified to smear out structures in the respective 840 direction while leaving air density and structures in the orthogonal direction unchanged. 890

5.2 Inversion of simulated measurements

845 Case studies based on real measurement data are inadequate as the sole proof of concept because the truth is unknown and the result thus is unverifiable. Instead we first test our 895 scheme on the basis of simulated atmospheric states and consider the scheme as verified if the velocities and mixing coefficients used to simulate the atmospheric states are sufficiently well reproduced. In the noise-free well conditioned case one might even expect, within the numerical precision of the system, the exact reproduction of the reference data; due to the – weak but non-zero – dispersivity of the numerical transport scheme the wiggles discussed in the previous 900 subsection cause, at some gridpoints, D-matrix entries of the wrong sign. In order to fight resulting convergence problems of the inversion, at least some small regularization is adequate, even if the system of equations to be solved is well or over-determined. Since the system in reality is, in tendency, 905 ill-conditioned and the constraint applied to the inversion prevents reproduction of the reference data, we use a variety of idealized tracers instead. After this initial test of functionality, more and more realistic test cases are constructed 910 in order to study the competing influence of constraint and measurement data on the solution.

865 During code development, a series of basic tests of increasing complexity were performed, including a variety of mixing ratio distributions transported with various velocity 915 fields. The main lesson learned was that, even when the rows of Equation (80) were independent and no formal ill-posedness could be diagnosed, the unphysical upwind wiggles in the vicinity of the mixing ratio peak as discussed in the previous subsection could trigger errors which are boosted during the iteration. This problem, which is associated 920 with sharp structures and large velocities (of the order of one gridwidth per macro timestep) can be solved by the use of a smoothing regularization matrix \mathbf{R} as discussed in the

last paragraph of Section 4, however at the cost of degraded spatial resolution of the result.

As an example we show the following test. An altitude-independent meridional velocity field (in degrees per month)

$$v = \frac{23.6562}{2} \cos \phi \quad (91)$$

was chosen while the vertical velocity was set to zero (Fig. 2a). This particular choice of the meridional velocity field keeps boundary problems at the poles due to divergence in a circulation which is not closed reasonably small. Initial mixing ratio distributions (in ppmv) of four idealized gases were constructed such that gas (a) was sugar-loaf shaped while those of the gases (b) to (d) were monotonical slopes.

$$vmr_a = \left(e^{-\left(\frac{z-36}{1.5}\right)^2/72} + 5.0 * 10^{-4} \right) e^{-\left(\frac{6(\phi+6)}{40}\right)^2/12} \quad (92)$$

$$vmr_b = \frac{(30 + 0.02z^2 - 0.2\phi)^3}{1000} \quad (93)$$

$$vmr_c = 20 * e^{0.02z+0.03\phi} \quad (94)$$

and

$$vmr_d = 30 * e^{-0.03z-0.002\phi}, \quad (95)$$

where z is altitude above surface in km and ϕ is latitude in degrees. The distribution of gas (a) is shown in Fig. 2b. The transport problem was solved in the forward mode using the velocity field as defined in Eq. (91) for integrating the tendency equation (Eq. 28) over one month. The resulting distribution of gas a after this macro-timestep is shown Fig. 2c. The “sugarloaf” is transported to the North by the expected distance. Due to the latitudinal gradients with lower velocities in the luff of the sugar-loaf and larger velocities in the lee, the shape of the “sugar-loaf” is slightly distorted, leading to a steeper slope in the luff and a flatter slope in the lee. No indication of problems with diffusity or dispersivity of the transport scheme is seen. At the poles boundary problems are visible which are unavoidable when such an unrealistic (but instructive) velocity field is used. The inversion nicely reproduces the initial velocity field (not shown because hardly discernable from Fig. 2a). Due to the smoothing constraint the peak velocity is decreased by 18%. This smoothing is the price to pay for the stabilization of the retrieval in the presence of the boundary problems discussed above. The residual between the simulated measurement of the distribution of gas (a) at $t_0 + \Delta t$ and the distribution predicted with the resulting velocity field is shown in (Fig. 2d).

5.3 Case Study with MIPAS measurements

The risk of case studies based on simulated data typically is that not all difficulties encountered with real data are fore-

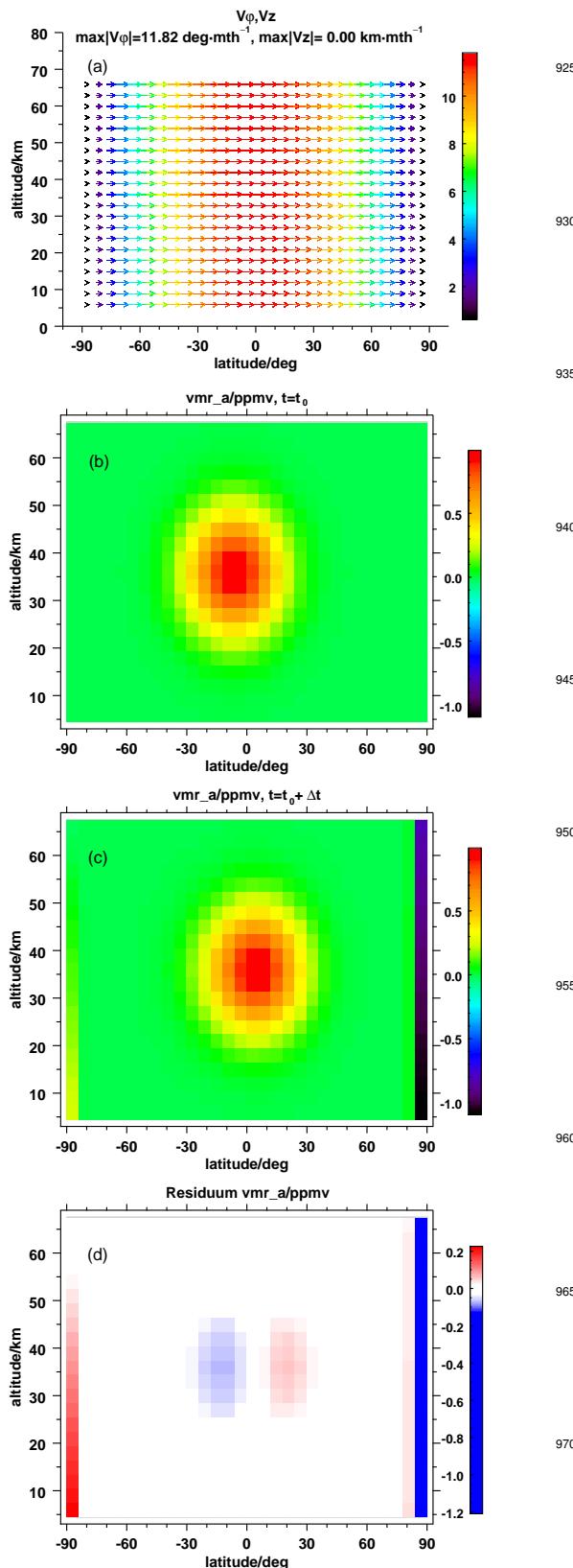


Figure 2. : Test case. (a) velocity field; (b) initial distribution of gas a; (c) distribution of gas a after one month; (d) residual between distribution of gas a after one month predicted with the resulting and the correct velocities. The units of the colour bar in panel (a) are degrees per month.

seen during theoretical studies. In order to demonstrate applicability to real data, global monthly latitude/altitude distributions of CFC-12, CH₄, N₂O and SF₆ (Kellmann et al., 2012; Plieninger et al., 2015; Haenel et al., 2015) measured with the Michelson Interferometer for Passive Atmospheric Sounding (MIPAS) (Fischer et al., 2008) were used. The purpose of these tests is demonstration of the feasibility of the method presented. An investigation of the atmospheric circulation on the basis of this method applied to MIPAS data is left for a companion paper. For this proof of concept, sinks of these long-lived tracers have been ignored but these will certainly be considered in scientific applications.

For this case study, zonal monthly mean distributions of air densities and mixing ratios of these four species from September and October 2010 were used. Figure 3 shows the measured distributions of these quantities in September (left column) and October (middle column), and the residuals between the measured and predicted contributions for October (right column). The residuals are reasonably small and show, except for methane in the polar upper stratosphere, no patterns which would hint at peculiarities with the inferred kinematic quantities.

The resulting circulation vectors which best explain the change of the mixing ratio distributions from September to October 2010 are shown in the upper left panel of Figure 4. Winter polar subsidence, summer polar upwelling, the mesospheric overturning circulation, the upper and lower branches of the Brewer-Dobson circulation and the tropical pipe are clearly visible. Details of the tropical pipe are visible in the right upper panel. As expected, the Brewer-Dobson circulation is much more pronounced in the northern (early) winter hemisphere. Velocities are roughly consistent with mean ages of stratospheric air as determined by Stiller et al. (2008) and Haenel et al. (2015) in a sense that the quotient of the typical circulation velocity and the distance between equator to pole gives an age estimate of the correct order of magnitude. While the inferred field of circulation vectors shows many detail features demanding scientific investigation in their own right (e.g., the latitude offset between the intertropical convergence zone and the stratospheric tropical pipe, or the interfacing between the stratospheric two-cell circulation and the overturning mesospheric circulation and the transition altitude between them), the reproduction of the expected features justifies confidence in the method proposed. Resulting mixing coefficients K_ϕ and K_z are shown in the left and right lower panels, respectively. Negative mixing coefficients indicate counter-gradient mixing, which seems to be most pronounced in the tropical upper stratosphere.

Jacobian elements with respect to v values and K values seem to form a null space. Thus the K -values were constrained to zero using diagonal components only in the \mathbf{R} matrix diagonal blocks associated with the K values. The strength of this constraint was adjusted such that the K values were as small as possible as long as this did not boost

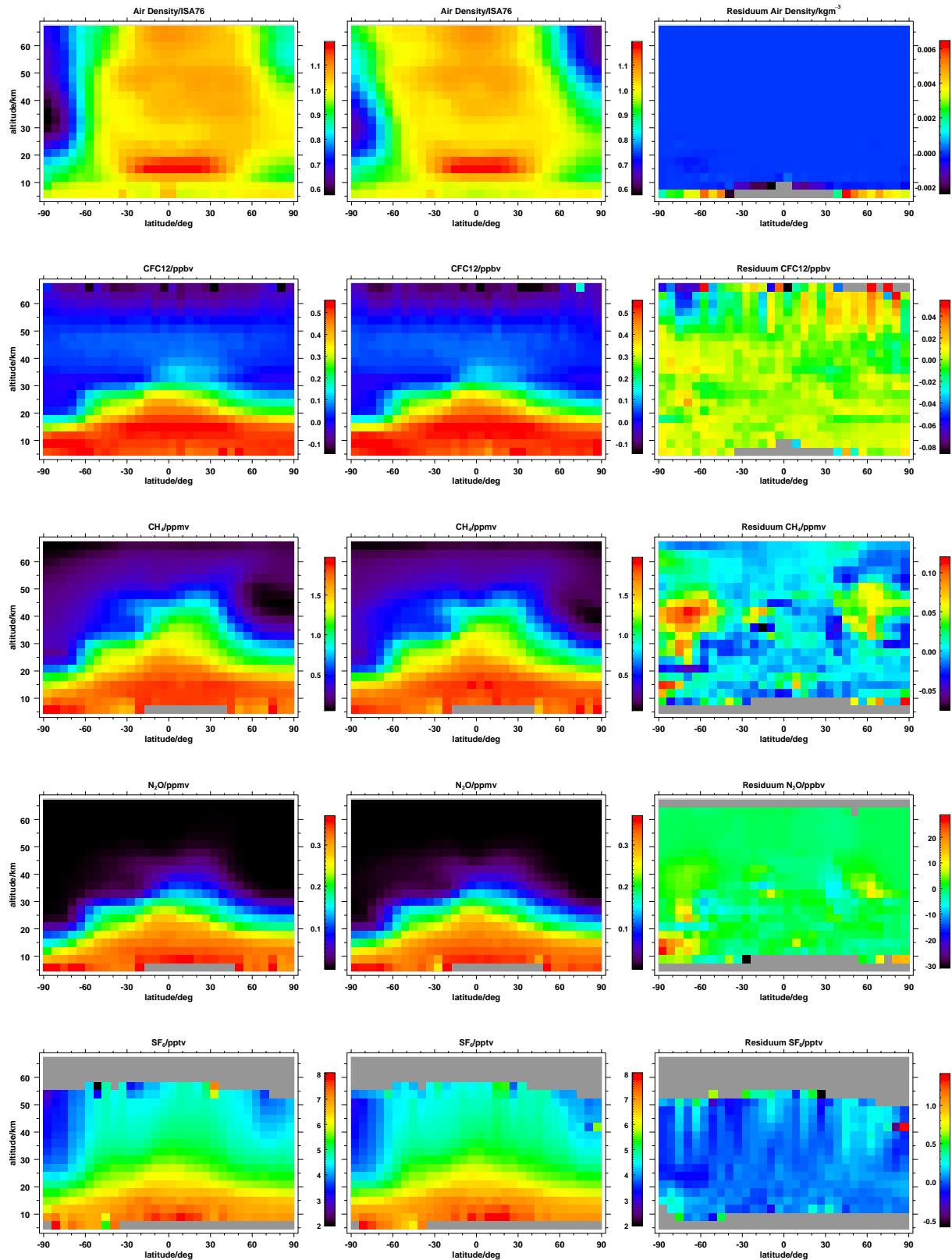


Figure 3. : Measured distributions in September (left column), October (middle column) and residual distributions between October measurements and predictions for October (right column) for air density and mixing ratios of CFC-12, CH₄, N₂O, and SF₆ (top to bottom). Grey gridboxes indicate non-availability of valid data.

the residual. Resulting K_ϕ and K_z distributions are shown in Figure 4.

The errors in the estimated transport velocities and mixing coefficients have been estimated according to Eq. (85) and are shown in Figure 5. The uncertainties in the transport velocities caused by the propagation of measurement errors are in the one percent range, indicating that the information contained in the measurements is adequate for the purpose of retrieving circulation parameters. It seems even possible to improve the time resolution of the circulation analysis and aim at weekly instead of monthly temporal sampling.

Larger errors above 65 km altitude and at the bins closest to the pole are border effects, resulting from the fact that no symmetric derivatives can be calculated there. The uncertainties in K_ϕ show the same patterns as the K_ϕ values themselves.

6 Discussion

The analysis of the age of stratospheric air can be understood as an integrated view at the equations of motion of stratospheric air, because the total travel time of the air parcel through the stratosphere is represented. The refinement of this method which analyzes the mean age just considers a weighted mean of the above, but it is still an integral method. Contrary to these integral methods, our direct inversion scheme supports a – in approximation, due to discrete sampling in the time domain – differential view of the same problem. The related advantages are: (a) independence of assumptions on the age spectrum, because during each time step mixing is explicitly considered; (b) insensitivity to SF_6 depletion in the mesosphere (c.f., e.g., Reddmann et al., 2001; Stiller et al., 2012), because the scheme uses the actual entry values of subsiding air as a reference; (c) applicability to non-ideal tracers in the stratosphere; since the atmospheric state is updated for each time step by measured value, depletion does not accumulate, even if no sink functions are considered; and (d) the logical circle that the lifetimes of non-ideal tracers depend on their trajectories (and thus atmospheric circulation), while the determination of the circulation requires knowledge of the lifetimes, can be solved. Our scheme requires knowledge only on the local, not the global, lifetimes; (e) the method is an empirical method which does not involve any dynamical model, i.e. the forces which cause the circulation are not required. The method only finds that kinematic state of the atmosphere which, according to the continuity equation, fits best to the measurements. These kinematic state values are provided as model diagnostics to assess the performance of dynamical models. Due to these advantages, the major problems in the empirical analysis of the Brewer-Dobson circulation as mentioned by Butchart (2014) are solved. Problems related to our method are (a) sensitivity of the inferred kinematic quantities to locally varying biases, (b) a tendency towards ill-posedness of

the inversion if distributions of too few tracers with too similar morphology are used, and (c) the usual artefacts arising if the numerical discretization is chosen too coarse. Results of the case study presented in Section 5.3 suggest that these problems are under control in the current application of the proposed scheme.

7 Conclusions and outlook

We have presented a method which infers mixing coefficients and effective velocities of a 2-D atmosphere by inversion of the continuity equation. The main steps of this procedure are (a) integration of the continuity equation over time to predict pressure and mixing ratios for given initial pressures and mixing ratios and initially guessed velocities and mixing coefficients; (b) propagation of errors of initial pressures and mixing ratios onto the predicted pressures and mixing ratios, by differentiation of the predicted state with respect to the initial state and generalized Gaussian error propagation; (c) estimation of the sensitivities of the predicted state with respect to the velocities and mixing coefficients; and (d) minimization of a quadratic cost function involving the residual between measured and predicted state at the end of the forecasting interval by inversion of the continuity equation. The inferred velocities are suggested to be used as a model diagnostic in order to avoid problems encountered with other model diagnostics like mean age of stratospheric air. It is important to note that the diagnostics inferred here are effective transport velocities and effective mixing coefficients in a sense that they include eddy transport and diffusion terms. Thus, they cannot simply be compared to zonal mean velocities and mixing coefficients of a 3D model but the eddy terms have to be considered when these diagnostics quantities are calculated. The application of this method on SF_6 distributions measured by MIPAS (Stiller et al., 2012) to diagnose the Brewer-Dobson circulation are discussed in a companion paper. Obvious future activities are the extension of this method to three dimensions and inclusion of sink functions of non-inert species to explore a larger number of tracers in order to better constrain the related inverse problem.

Appendix A: From 3D to 2D

The reduction of the transport problem from three to two dimensions involves Reynolds decomposition of the three-dimensional continuity equation and subsequent zonal averaging and gives rise to eddy mixing and transport terms. The inference of effective two-dimensional transport velocities and effective mixing coefficients from measurements discussed in the main paper relies on the fact that, within certain assumptions and approximations, all these eddy effects can be understood as additional pseudo-advection and pseudo-mixing terms according to the advection equation and the Fickian law, with gas-independent pseudo-velocities

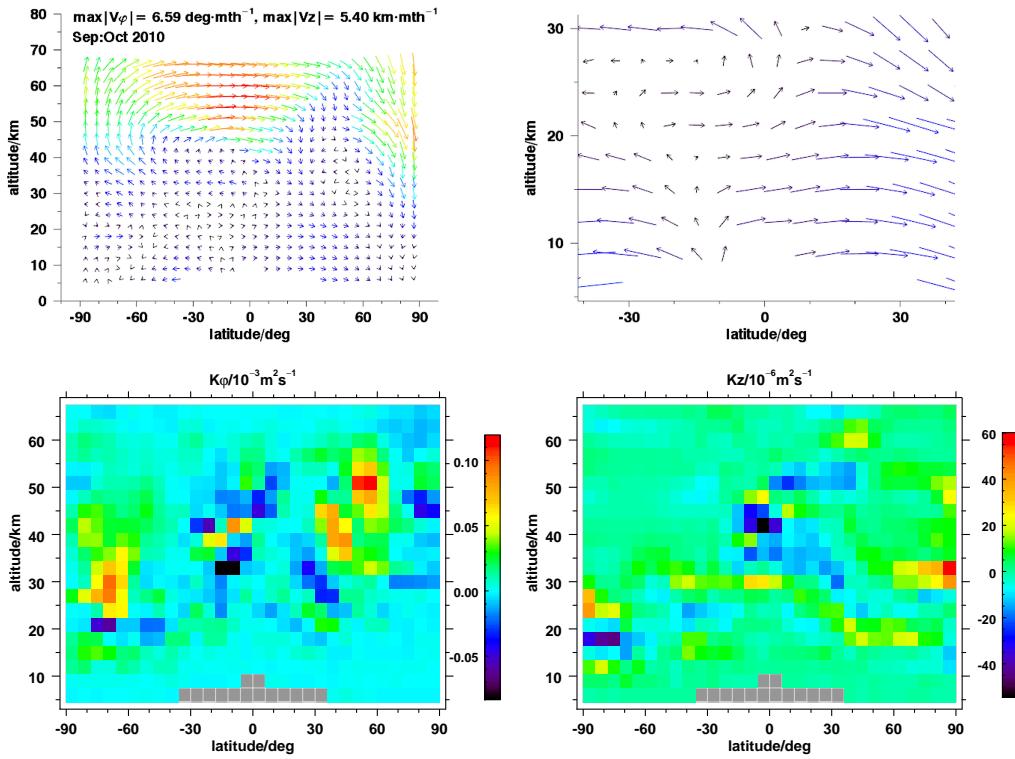


Figure 4. : Resulting circulation vectors ($v(z, \phi); w(z, \phi)$) (upper left panel), where colours on the red side of the rainbow colour scale represent higher velocities; a zoomed-in view of this (upper right panel); mixing coefficients K_ϕ (lower left panel) and K_z (lower right panel).

and pseudo-mixing coefficients. The exact interpretation of the two-dimensional velocities and mixing coefficients inferred from the measurements depends on the approximations made. We apply our scheme to zonally averaged mixing ratios (no mass-weighted averaging!). Contrary to the main text, where the symbols v , w , K_ϕ , and K_z are used in the 2-dimensional system, here zonal averages are indicated by a bar.

Assuming

- that the deviations from the zonal mean are small compared to the zonal mean itself such that linearization is justifiable,
- that meridional advection is negligibly small compared to zonal advection,
- that the time variation of the zonal mean quantities is assumed to be much slower than the time variation of the deviations from the zonal mean, which corresponds to the assumption of a quasi-steady state,

Tung (1982) derives a two-dimensional approximation to the continuity equation which, adjusted to our notation, written for geometric altitudes instead of potential temperatures, and using the shallowness approximation, reads

$$\begin{aligned}
 & \bar{\rho} \frac{\partial}{\partial t} \bar{v} \bar{m} \bar{r}_g + \left(\bar{\rho} \bar{v} - \frac{\partial}{\partial t} \bar{\rho}' \eta' \right) \frac{1}{r} \frac{\partial}{\partial \phi} \bar{v} \bar{m} \bar{r}_g + (A1) \\
 & \left(\bar{\rho} \bar{w} - \frac{\partial}{\partial t} \bar{\rho}' \Phi' \right) \frac{\partial}{\partial z} \bar{v} \bar{m} \bar{r}_g - \frac{1}{r} \frac{\partial}{\partial \phi} \left(\bar{\rho} K_{\phi\phi} \frac{1}{r} \frac{\partial}{\partial \phi} \bar{v} \bar{m} \bar{r}_g \right) - \\
 & \frac{1}{r} \frac{\partial}{\partial \phi} \left(\bar{\rho} K_{\phi z} \frac{\partial}{\partial z} \bar{v} \bar{m} \bar{r}_g \right) - \frac{\partial}{\partial z} \left(\bar{\rho} K_{z\phi} \frac{1}{r} \frac{\partial}{\partial \phi} \bar{v} \bar{m} \bar{r}_g \right) - \\
 & \frac{\partial}{\partial z} \left(\bar{\rho} K_{zz} \frac{\partial}{\partial z} \bar{v} \bar{m} \bar{r}_g \right) = \\
 & \bar{S} - \frac{1}{r} \frac{\partial}{\partial \phi} ((\bar{\rho} \bar{v})' \sigma') - \frac{\partial}{\partial z} ((\bar{\rho} \bar{z})' \sigma') - \frac{\partial}{\partial t} (\bar{\rho}' \sigma').
 \end{aligned}$$

where η' , Φ' and σ' are defined by

$$\left(\frac{\partial}{\partial t} + \frac{\bar{u}}{r \cos \phi} \frac{\partial}{\partial \lambda} \right) \eta' = v', \quad (A2)$$

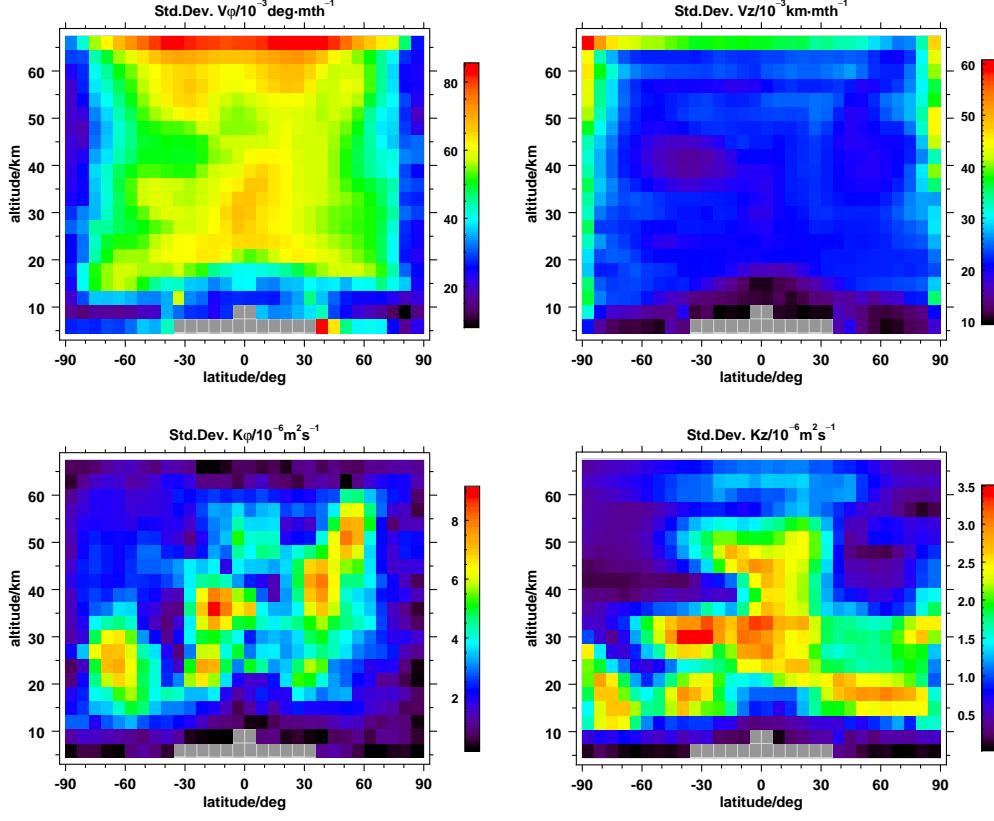


Figure 5. : Estimated uncertainties of $v(z, \phi)$, (upper left panel), $w(z, \phi)$ (upper right panel), K_ϕ (lower left panel) and K_z (lower right panel).

$$\left(\frac{\partial}{\partial t} + \frac{\bar{u}}{r \cos \phi} \frac{\partial}{\partial \lambda} \right) \Phi' = w' \quad (\text{A3})$$

and

$$\left(\frac{\partial}{\partial t} + \frac{\bar{u}}{r \cos \phi} \frac{\partial}{\partial \lambda} \right) \sigma' = S'. \quad (\text{A4})$$

Further,

$$K_{\phi\phi} = \frac{1}{\bar{\rho}} \overline{(\rho v)' \eta'} \quad (\text{A5})$$

$$K_{zz} = \frac{1}{\bar{\rho}} \overline{(\rho w)' \Phi'} \quad (\text{A6})_{1130}$$

$$K_{\phi z} = \frac{1}{\bar{\rho}} \overline{(\rho v)' \Phi'} \quad (\text{A7})$$

and

$$K_{z\phi} = \frac{1}{\bar{\rho}} \overline{(\rho w)' \eta'} \quad (\text{A8})$$

Assuming further that

– our scheme is applied only to long-lived species, such that chemical eddy terms can be ignored because chemical lifetimes are long compared to transport lifetimes (c.f. Pyle and Rogers (1980)),

– wave disturbances are dominated by steady or periodic terms, such that the terms with the mixed second derivative terms tend to disappear (Matsuno (1980); Clark and Rogers (1978); Pyle and Rogers (1980), quoted after Tung (1982)). Here an even weaker approximation is sufficient, namely

$$\frac{\partial}{\partial t} \overline{\eta' \Phi'} \approx 0. \quad (\text{A9})$$

This is equivalent with the assumption that $K_{z\phi} \approx -K_{\phi z}$

Eq. A1 can be rewritten as a tendency equation of the type

$$\frac{\partial}{\partial t} \overline{vmr_g} = -\frac{v^*}{r} \frac{\partial}{\partial \phi} \overline{vmr_g} - w^* \frac{\partial}{\partial z} \overline{vmr_g} + \frac{1}{r^2} \frac{\partial}{\partial \phi} \left[K_\phi^* \frac{\partial}{\partial \phi} \overline{vmr_g} \right] + \frac{\partial}{\partial z} \left[K_z^* \frac{\partial}{\partial z} \overline{vmr_g} \right] \quad (A10)$$

where

$$v^* = \bar{v} - \frac{1}{\bar{\rho}} \frac{\partial}{\partial t} \overline{\rho' \eta'} - \frac{1}{\bar{\rho} r} \frac{\partial \bar{\rho}}{\partial \phi} K_{\phi\phi} - \frac{\partial}{\partial z} K_{z\phi} - \frac{1}{\bar{\rho}} \frac{\partial \bar{\rho}}{\partial z} K_{z\phi}, \quad (A11)$$

$$w^* = \bar{w} - \frac{1}{\bar{\rho}} \frac{\partial}{\partial t} \overline{\rho' \Phi'} - \frac{1}{\bar{\rho}} \frac{\partial \bar{\rho}}{\partial z} K_{zz} - \frac{1}{r} \frac{\partial}{\partial \phi} K_{\phi z} - \frac{1}{\bar{\rho} r} \frac{\partial \bar{\rho}}{\partial \phi} K_{\phi z}, \quad (A12)$$

$$K_\phi^* = K_{\phi\phi} \quad (A13)$$

and

$$K_z^* = K_{zz} \quad (A14)$$

can be understood as virtual velocities and virtual mixing coefficients. All terms in Eq. (A10) have either the mathematical structure of an advection equation or a Fickian law. Thus, Eqs.(A11 - A14) provide the interpretation of the velocities and mixing coefficients inferred in the main part of the paper. These can be identified with the virtual velocities and virtual mixing coefficients inferred above. Obviously, this interpretation depends on the approximations made, and different approximations would lead to a different interpretation. If mixing coefficients describing subscale effects in the original 3-D model (the last two terms in Eq. (3)) are to be considered, then the effective mixing coefficient is the sum of the mixing coefficients describing these subscale effects plus the respective K^* term accounting for the eddy mixing.

We like to emphasize that none of the approximations and assumptions discussed above are used in our proposed method to infer velocities from zonal mean mixing ratio measurements. The discussion in this Appendix only tries to relate the resulting velocities to the velocities in a 3-D world. The ambiguities in the interpretation of the inferred 'effective' 2-D velocities suggests that it might be promising to switch from a theoretical to an empirist view and to conceive no longer the zonal mean of the 3-D velocities as the 'true' 2-D velocities, but those which satisfy the 2-D continuity equation. These can be – admittedly indirectly – observed with our suggested method and can be used to validate 2-D models, including their underlying concept of solving the 2-D transport problem. With this, also the adequacy of the assumptions made to approximate away the headache terms

which can be expressed neither as advection nor as Fickian law terms can be tested by means of comparison of the measured and 2-D modelled effective, i.e. transport relevant, velocities. This empiristic turn in argumentation might not fully solve all aspects of the problem of interpretation of the observed 2-D velocities from a 3-D perspective, but at least it moves the problem from the desk of the observation scientist onto the desk of the 2-D modeller.

Acknowledgements. The authors thank two anonymous reviewers for their thorough examination of the original manuscript and for helpful and important comments. Furthermore, TvC wishes to thank Hendrik Elbern, Richard Menard, Peter Braesicke, Björn-Martin Sinnhuber and Thomas Birner for drawing his attention to some important literature and for encouragement as well as Arne Babenhausenheide for helpful discussions.

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