# Multi-year record of atmospheric mercury at Dumont d'Urville, East Antarctic coast: continental outflow and oceanic influences

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# 13 Abstract

14 Under the framework of the Global Mercury Observation System (GMOS) project, a 3.5-year 15 record of atmospheric gaseous elemental mercury (Hg(0)) has been gathered at Dumont 16 d'Urville (DDU, 66°40'S, 140°01'E, 43 m above sea level) on the East Antarctic coast. Additionally, surface snow samples were collected in February 2009 during a traverse 17 18 between Concordia Station located on the East Antarctic plateau and DDU. The record of 19 atmospheric Hg(0) at DDU reveals particularities that are not seen at other coastal sites: a 20 gradual decrease of concentrations over the course of winter, and a daily maximum concentration around midday in summer. Additionally, total mercury concentrations in 21 surface snow samples were particularly elevated near DDU (up to 194.4 ng L<sup>-1</sup>) as compared 22 23 to measurements at other coastal Antarctic sites. These differences can be explained by the 24 more frequent arrival of inland air masses at DDU than at other coastal sites. This confirms 25 the influence of processes observed on the Antarctic plateau on the cycle of atmospheric 26 mercury at a continental scale, especially in areas subject to recurrent katabatic winds. DDU 27 is also influenced by oceanic air masses and our data suggest that the ocean plays a dual role 28 on Hg(0) concentrations. The open ocean may represent a source of atmospheric Hg(0) in 29 summer whereas the sea-ice surface may provide reactive halogens in spring that can oxidize

Hg(0). This paper also discusses implications for coastal Antarctic ecosystems and for the
 cycle of atmospheric mercury in high southern latitudes.

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# 33 **1** Introduction

The Antarctic continent is one of the last near-pristine environments on Earth and still 34 35 relatively unaffected by human activities. Except for pollutants released from Antarctic Research stations (e.g., Hale et al., 2008; Chen et al., 2015) and by marine and air-borne 36 37 traffic (Shirsat and Graf, 2009), only the long-lived atmospheric contaminants reach this 38 continent situated far from anthropogenic pollution sources. With an atmospheric lifetime on 39 the order of one year (Lindberg et al., 2007), gaseous elemental mercury (Hg(0)) is efficiently 40 transported worldwide. Hg(0) is the most abundant form of mercury in the atmosphere 41 (Lindberg and Stratton, 1998). It can be oxidized into highly-reactive and water-soluble 42 gaseous divalent species (Hg(II)) – that can bind to existing particles and form particulate 43 mercury (Hg(p)) – leading to the deposition of reactive mercury onto various environmental 44 surfaces through wet and dry processes (Lindqvist and Rodhe, 1985; Lin and Pehkonen, 45 1999). Upon deposition, Hg(II) can be reduced and reemitted back to the atmosphere as Hg(0) (Schroeder and Munthe, 1998). Assessing mercury deposition and reemission pathways 46 47 remains difficult due to an insufficient understanding of the involved physic-chemical 48 processes.

49 Only sparse measurements of atmospheric mercury have been performed in Antarctica and 50 there are still many gaps in our understanding of its cycle at the scale of this vast continent (~ 14 million km<sup>2</sup>) (Dommergue et al., 2010). To date, observations were made over one year at 51 52 the coastal site of Neumayer (NM, Ebinghaus et al., 2002; Temme et al., 2003) and during 53 summer campaigns at Terra Nova Bay (TNB, Sprovieri et al., 2002) and McMurdo (MM, 54 Brooks et al., 2008b). More recently, multi-year records have been obtained at Troll (TR) 55 situated approximately 220 km from the coast at 1275 m a.s.l. (Pfaffhuber et al., 2012) and 56 Concordia Station located at Dome C (denoted DC, 3220 m a.s.l.) (Angot et al., 2016). Under 57 the framework of the GMOS project (Global Mercury Observation System, www.gmos.eu), 58 atmospheric monitoring of Hg(0) has been implemented at Dumont d'Urville (DDU) located 59 in Adélie Land (Fig. 1) and we here report the obtained 3.5-year record of atmospheric Hg(0) 60 that represents the first multi-year record of Hg(0) available for the East Antarctic coast. In 61 this paper, the Hg(0) record from DDU is discussed in terms of influence of marine versus inland air masses, and compared to records available at other coastal (NM, TNB, MM) or
near-coastal (TR) stations. In parallel, total mercury was determined in surface snow samples
collected during a traverse between DC and DDU in February 2009. These results provide
new insight into the transport and deposition pathways of mercury species in East Antarctica.

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# 67 2 Experimental Section

#### 68 **2.1** Sampling site and prevailing meteorological conditions

69 From January 2012 to May 2015, Hg(0) measurements were performed at DDU station 70 located on a small island (Ile des Pétrels) about one km offshore from the Antarctic mainland. 71 A detailed description of the sampling site ("Labo 3") has been given by Preunkert et al. 72 (2013) while the climatology of this coastal station has been detailed by König-Langlo et al. 73 (1998). The average surface air temperature ranges from -1 °C in January to -17 °C in winter, 74 with a mean annual temperature of -12 °C. The annual mean surface wind speed is 10 m s<sup>-1</sup>, 75 with no clear seasonal variations. Due to the strong katabatic effects, the most frequent 76 surface wind direction is 120°E-160°E.

#### 77 2.2 Methods

#### 78 2.2.1 Hg(0) measurements

79 Hg(0) measurements were performed using a Tekran 2537B (Tekran Inc., Toronto, Canada). 80 The sampling resolution ranged from 10 to 15 minutes with a sampling flow rate of 1.0 L min 81 <sup>1</sup>. Concentrations are reported here as hourly averages and are expressed in nanograms per 82 cubic meter at standard temperature and pressure (273.15 K, 1013.25 hPa). Setting a 0.2 µm 83 PTFE filter and a 10 m long unheated sampling line on the front of the analyzer inlet, we 84 assume that mainly Hg(0) (instead of total gaseous mercury, defined as the sum of gaseous 85 mercury species) was efficiently collected and subsequently analyzed by the instrument 86 (Steffen et al., 2002; Temme et al., 2003; Steffen et al., 2008).

External calibrations were performed twice a year by injecting manually saturated mercury vapor taken from a temperature-controlled vessel, using a Tekran 2505 mercury vapor calibration unit and a Hamilton digital syringe, and following a strict procedure adapted from Dumarey et al. (1985). As described by Angot et al. (2014), fortnightly to monthly routine maintenance operations were performed. A software program was developed at the LGGE 92 (Laboratoire de Glaciologie et Géophysique de l'Environnement) following quality control
93 practice commonly applied in North American networks (Steffen et al., 2012). Based on
94 various flagging criteria (Munthe et al., 2011; D'Amore et al., 2015), it enabled rapid data
95 processing in order to produce clean time series of Hg(0). According to the instrument
96 manual, the detection limit is 0.10 ng m<sup>-3</sup> (Tekran, 2011).

#### 97 **2.2.2 Snow sampling and analysis**

Eleven surface snow samples (the upper 3 cm) were collected during a traverse between DC 98 99 and DDU conducted in February 2009. As described by Dommergue et al. (2012), samples 100 were collected using acid cleaned PTFE bottles and clean sampling procedures. After 101 sampling, samples were stored in the dark at -20 °C. Field blanks were made by opening and 102 closing a bottle containing mercury-free distilled water. Total mercury (Hgtot) in snow 103 samples was analyzed using a Tekran Model 2600. Hg<sub>tot</sub> includes species such as HgCl<sub>2</sub>, 104 Hg(OH)<sub>2</sub>, HgC<sub>2</sub>O<sub>4</sub>, stable complexes such as HgS and Hg(II) bound to sulfur in humic 105 compounds, or some organomercuric species (Lindqvist and Rodhe, 1985). The instrument 106 was calibrated with the NIST SRM-3133 mercury standard. Quality assurance and quality 107 control included the analysis of analytical blanks, replicates, and internal standards (Reference 108 Waters for mercury: HG102-2 at 22 ng/L from Environment Canada). The limit of 109 quantification - calculated as 10 times the standard deviation of a set of 3 analytical blanks was 0.3 ng L<sup>-1</sup> and the relative accuracy  $\pm 8\%$ . 110

111 Surface snow samples collected during traverses may have limited spatial and temporal 112 representativeness given the variability of chemical species deposition onto the snow surface, 113 and the occurrence of either fresh snowfall or blowing snow. The (in)homogeneity of surface 114 snow samples was investigated at MM by Brooks et al. (2008b). Surface (3-5 cm) snow samples were collected daily (n = 14) at different snow patches. Hg<sub>tot</sub> concentrations averaged 115  $67 \pm 21$  ng L<sup>-1</sup>. This result indicates that the spatial and temporal representativeness of surface 116 117 snow samples collected in Antarctica can be satisfactory and gives us confidence that spatial 118 differences in Hgtot concentrations reported in section 3.2.2 are not due to samples 119 inhomogeneity.

#### 120 **2.2.3 Ancillary parameters**

O<sub>3</sub> was continuously monitored with a UV absorption monitor (Thermo Electron Corporation
 model 49I, Franklin, Massachusetts) (Legrand et al., 2009). Collected at 15-s intervals, the
 data are reported here as hourly averages.

124 Back trajectories were computed using the HYSPLIT (Hybrid Single-Particle Lagrangian 125 Integrated Trajectory) model (Draxler and Rolph, 2013). Meteorological data from Global 126 Data Assimilation Process (available at ftp://arlftp.arlhq.noaa.gov/pub/archives/gdas1) were 127 used as input, and the model was run every hour in backward mode for 5 days at 0, 200, and 128 500 m above the model ground level. Three typical situations prevail at DDU: strong 129 katabatic winds flowing out from the Antarctic ice sheet situated south of the station, pure 130 marine air masses, or continental/marine mixed air masses with easterly winds due to the 131 arrival near the site of low-pressure systems (König-Langlo et al., 1998). Oceanic origin was 132 attributed to air masses having traveled at least 1 day over the ocean and less than 3 days out 133 of 5 over the high-altitude Antarctic plateau. Conversely, plateau origin refers to air masses 134 having traveled at least 3 days over the high-altitude Antarctic plateau and less than 1 day out 135 of 5 over the ocean. Finally, mixed origin refers to air masses having traveled less than 1 and 136 3 days out of 5 over the ocean and the high-altitude Antarctic plateau, respectively. It should 137 be noted that uncertainties associated with calculated backward trajectories arise from 138 possible errors in input meteorological fields and numerical methods (Yu et al., 2009), and 139 increase with time along the way (Stohl, 1998). According to Jaffe et al. (2005), back 140 trajectories only give a general indication of the source region. Despite these limitations, back 141 trajectories remained very similar at the three levels of altitude arrival at the site and we only 142 use here those arriving at the model ground level. This method also gave consistent results 143 with respect to the origin of various chemical species including  $O_3$  (Legrand et al., 2009), 144 HCHO (Preunkert et al., 2013), NO<sub>2</sub> (Grilli et al., 2013), and sea-salt aerosol (Legrand et al., 145 2016a).

#### 146 **2.3 Local contamination**

147 Pollution plumes due to the station activities (e.g., combustion, vehicular exhaust) 148 occasionally reached the sampling site. Such local pollution events can be easily identified for 149 instance by the fast decrease of O<sub>3</sub> or increase of HCHO mixing ratios (Legrand et al., 2009; 150 Preunkert et al., 2013). We used a criterion based on wind direction and sudden drops of  $O_3$ mixing ratios to filter the raw data (i.e., collected at 5 min intervals) and discard Hg(0) data 151 impacted by local pollution. Raw Hg(0) data above 1.60 ng m<sup>-3</sup>, corresponding to the mean + 152 153 3 standard deviation, obtained when the wind was blowing from 30°W to 70°E (i.e., the sector where main station activities are located), and accompanied by a drop of O<sub>3</sub> were discarded 154

155 from the data set. Using this criterion, only 0.1% of raw Hg(0) data was discarded, the Hg(0)
156 record being very weakly impacted by pollution plumes.

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# 158 **3 Results and Discussion**

The record of atmospheric Hg(0) from January 2012 to May 2015 is displayed in Fig. 2. 159 Hourly-averaged Hg(0) concentrations ranged from 0.10 to 3.61 ng  $m^{-3}$ , with an average value 160 of  $0.87 \pm 0.23$  ng m<sup>-3</sup> (mean  $\pm$  standard deviation). This mean annual Hg(0) concentration is 161 in good agreement with the value of  $0.93 \pm 0.19$  ng m<sup>-3</sup> (4-year average) reported by 162 Pfaffhuber et al. (2012) at TR, but lower than the concentration of  $1.06 \pm 0.24$  ng m<sup>-3</sup> (12-163 164 month average) reported by Ebinghaus et al. (2002) at NM. While the same device was used 165 at the three stations, the measurements may target different mercury species depending on 166 their configuration (e.g., heated/unheated sample line). The difference between total gaseous 167 mercury and Hg(0) data can be rather substantial since gaseous oxidized mercury (Hg(II)) concentrations of up to ~ 0.30 ng m<sup>-3</sup> were reported in spring/summer at several coastal 168 169 Antarctic stations (Sprovieri et al., 2002; Temme et al., 2003; Brooks et al., 2008b). To 170 allow a more accurate comparison of data available at the various Antarctic stations, more 171 harmonized sampling protocols are needed. Seasonal boundaries have been defined as 172 follows: summer refers to November-February, fall to March-April, winter to May-August, 173 and spring to September-October. Though being arbitrary, this dissection was done by 174 considering the time period over which the halogen chemistry (September-October) or the 175 OH/NO<sub>x</sub> chemistry (November-February) is dominant at DDU (see sections 3.1.2 and 3.2.2). The mechanisms which cause the seasonal variation of Hg(0) concentrations are discussed in 176 177 the following sections.

### 178 **3.1** From winter darkness to spring sunlight

#### 179 **3.1.1** Continental outflow and advection from lower latitudes in winter

A gradual 20% decrease in Hg(0) concentrations from  $0.89 \pm 0.09$  in average in May to 0.72  $\pm 0.10$  ng m<sup>-3</sup> in August (Fig. 3a) was observed at DDU. Conversely, concentrations remained rather stable at NM and TR in winter with mean values of  $1.15 \pm 0.08$  and  $1.00 \pm 0.07$  ng m<sup>-3</sup>, respectively (Ebinghaus et al., 2002; Pfaffhuber et al., 2012). Pfaffhuber et al. (2012) suggested that this stability of Hg(0) concentrations at TR is related to a lack of oxidation processes during the polar night.

186 A local reactivity at DDU – absent at other coastal stations – seems unlikely. Angot et al. 187 (2016) showed evidence of a gradual 30% decrease of Hg(0) concentrations at DC at the same 188 period of the year (Fig. 3a), probably due to a gas-phase oxidation, heterogeneous reactions, 189 or dry deposition of Hg(0) onto the snowpack. Since the decreasing trend observed in winter 190 is less pronounced at DDU than at DC, it most likely results from reactions occurring within 191 the shallow boundary layer on the Antarctic plateau, subsequently transported toward the 192 coastal margins by katabatic winds. This assumption is supported by the HYSPLIT model 193 simulations showing prevalence in winter  $(62 \pm 23\%)$  of air masses originating from the 194 Antarctic plateau reaching DDU (Fig. 4). The export of inland air masses towards the coastal 195 regions is not uniform across Antarctica and is concentrated in a few locations - "confluence 196 zones" - such as the Amery Ice Shelf region, the area near Adélie Land at 142°, the broad 197 region upslope from the Ross Ice Shelf, and the eastern side of the Antarctic Peninsula at ~ 198 60°W (Fig. 1) (Parish and Bromwich, 1987, 2007). Given its geographic location, DDU in 199 Adélie Land lies close to a confluence zone explaining the extent of the transport of air 200 masses from the Antarctic plateau. Conversely, several studies showed that stations such as 201 NM and HA are not significantly impacted by air masses originating from the Antarctic 202 plateau (Helmig et al., 2007; Legrand et al., 2016b), consistently explaining why Hg(0) 203 concentrations did not decrease at NM and TR throughout winter (Ebinghaus et al., 2002; 204 Pfaffhuber et al., 2012).

205 Despite the overall decreasing trend in winter, Hg(0) concentrations sporadically exhibited 206 abrupt increases when warm air masses from lower latitudes reached DDU. As illustrated by Fig. 5, Hg(0) concentration for example increased from 0.72 (8 June 2012) to 1.10 ng m<sup>-3</sup> (14 207 June 2012) with increasing temperature, and a significant positive correlation was found 208 209 between the two parameters (r = 0.88, p value < 0.0001, Spearman test). This result is 210 supported by an enhanced fraction of oceanic air masses reaching DDU at that time according 211 to the HYSPLIT model simulations (Fig. 5d). Consistently, aerosol data gained in the 212 framework of the French environmental observation service CESOA (http://www-213 lgge.obs.ujf-grenoble.fr/CESOA/spip.php?rubrique3) dedicated to the study of the sulfur 214 cycle at middle and high southern latitudes indicate a mean sodium concentration of 450 ng 215  $m^{-3}$  between 10 and 14 June 2012 (not shown) instead of  $112 \pm 62 \text{ ng } m^{-3}$  over the other days 216 of this month. It can be noted that the mean Hg(0) concentration in June 2012 was  $0.95 \pm 0.04$ ng m<sup>-3</sup> at TR (Slemr et al., 2015), and 1.02  $\pm$  0.04 ng m<sup>-3</sup> on Amsterdam Island (37°48'S, 217

218 77°34'E, Angot et al., 2014). These values are consistent with the increase seen at DDU in air
219 masses arriving from lower latitudes.

# 220 **3.1.2** The ice-covered ocean as a sink for Hg(0) in spring

221 First discovered in the Arctic in 1995 (Schroeder et al., 1998), Atmospheric Mercury 222 Depletion Events (AMDEs) have been subsequently observed after polar sunrise (mainly 223 from early September to the end of October) at coastal or near-coastal Antarctic stations at 224 NM (Ebinghaus et al., 2002), TNB (Sprovieri et al., 2002), MM (Brooks et al., 2008b), and TR (Pfaffhuber et al., 2012). These events, characterized by abrupt decreases of Hg(0) 225 concentrations below 1.00 ng m<sup>-3</sup> in the Arctic and 0.60 ng m<sup>-3</sup> in Antarctica (Pfaffhuber et 226 227 al., 2012), result from the oxidation of Hg(0) by reactive bromine species (e.g., Schroeder et 228 al., 1998; Lu et al., 2001; Brooks et al., 2006; Sommar et al., 2007). At DDU, Hg(0) data 229 covering the spring time period are scarce (Fig. 2) and we can just emphasize that the absence 230 of Hg(0) drops in October 2012 tends to suggest that AMDEs, if exist, are not very frequent at 231 DDU. Ozone Depletion Events (ODEs) are found to be less frequent and far less pronounced 232 at DDU compared to other coastal stations such as NM and HA (Legrand et al., 2009; 233 Legrand et al., 2016b). Based on the oxygen and nitrogen isotope composition of airborne 234 nitrate at DDU, Savarino et al. (2007) concluded to an absence of significant implication of 235 BrO in the formation of nitric acid at this site, contrarily to what is usually observed in the 236 Arctic where high levels of BrO are measured at polar sunrise (Morin et al., 2008). All these 237 observations are consistent with a less efficient bromine chemistry in East compared to West 238 Antarctica due to a less sea-ice coverage, as also supported by GOME-2 satellite observations 239 of the tropospheric BrO column (Theys et al., 2011; Legrand et al., 2016a). Additionally, air 240 masses originating from the Antarctic plateau prevailed ( $62 \pm 23$  %, Fig. 4) in spring at DDU 241 according to the HYSPLIT model simulations. This can also explain, to some extent, the lack 242 of AMDE-observations at DDU.

243 Despite the absence of large AMDEs at DDU, springtime oceanic air masses were associated 244 with low Hg(0) concentrations  $(0.71 \pm 0.11 \text{ ng m}^{-3}, \text{ see Fig. 3b})$ . A slight but significant 245 negative correlation was found between Hg(0) concentrations in spring and the daily-averaged 246 percentage of oceanic air masses reaching DDU (r = -0.38, *p* value = 0.01, Spearman test) 247 while a significant positive correlation was observed between springtime Hg(0) 248 concentrations and O<sub>3</sub> mixing ratios in these oceanic air masses (r up to 0.65, *p* value < 249 0.0001, Spearman test). Therefore, though being not as pronounced as AMDEs observed at other coastal stations, we cannot rule out that the rather low background Hg(0) levelsobserved in spring at DDU are due to a weak effect of the bromine chemistry.

# 252 **3.2** High variability in Hg(0) concentrations in summer

Hg(0) concentrations were highly variable during the sunlit period as compared to wintertime
(Fig. 2). Fig. 6 displays processes that may govern the atmospheric mercury budget at DDU in
summer, as discussed in the following sub-sections.

## 256 **3.2.1** Diurnal cycle of Hg(0) in ambient air

Fig. 7 displays the monthly mean diurnal cycle of Hg(0) concentrations at DDU. Undetected from March to October, a diurnal cycle characterized by a noon maximum was observed in summer (November to February). Interestingly, Pfaffhuber et al. (2012) did not observe any diurnal variation in Hg(0) concentrations at TR and there is no mention of a daily cycle at NM, TNB, and MM (Ebinghaus et al., 2002; Temme et al., 2003; Sprovieri et al., 2002; Brooks et al., 2008b).

Hg(0) concentrations at DDU were sorted according to wind speed and direction. With north at 0°, oceanic winds ranged from 270 to 110°E, coastal winds from 110 to 130°E, katabatic winds from 160 to 180°E, and continental winds from 130 to 160°E and from 180 to 270°E. Summertime Hg(0) concentrations exhibited a diurnal cycle regardless of wind speed and direction (Fig. 8). This result indicates that the observed diurnal cycle involves a local source of Hg(0) around midday which is, moreover, specific to DDU since the diurnal cycle is not observed at other coastal stations.

# 270 **3.2.1.1 Role of penguin emissions**

271 Large colonies of Adélie penguins nest on islands around DDU from the end of October to 272 late February, with a total population estimated at 60 000 individuals (Micol and Jouventin, 273 2001). Several studies highlighted that the presence of these large colonies at DDU in summer 274 significantly disturbs the atmospheric cycle of several species including ammonium and 275 oxalate (Legrand et al., 1998), carboxylic acids and other oxygenated volatile organic 276 compounds (Legrand et al., 2012), and HCHO (Preunkert et al., 2013). In a study 277 investigating sediment profiles excavated from ponds and catchments near penguin colonies 278 in the Ross Sea region, Nie et al. (2012) measured high mercury content in penguin excreta 279 (guano). Similarly, elevated total mercury concentrations were measured in ornithogenic soils 280 (i.e., formed by accumulation of guano) of the Fildes and Ardley peninsulas of King George 281 Island (De Andrade et al., 2012). When soil temperature rises above freezing in summer at 282 DDU, oxalate is produced together with ammonium following the bacterial decomposition of uric acid in ornithogenic soils (Legrand et al., 1998 and references therein). Dicarboxylic 283 284 acids such as oxalic acid were shown to promote the light-driven reduction of Hg(II) species 285 in aqueous systems and ice (Gårdfeldt and Jonsson, 2003; Si and Ariya, 2008; Bartels-286 Rausch et al., 2011). Emissions of Hg(0) from snow-covered ornithogenic soils are expected 287 to peak early and late summer – following the reduction of Hg(II) species in the upper layers 288 of the snowpack -, as also seen in the oxalate concentrations at DDU (Legrand et al., 1998). 289 Furthermore the rise of temperature at noon would strengthen Hg(0) emissions from 290 ornithogenic soils, possibly contributing to the observed diurnal cycle from November to 291 February.

### 292 **3.2.1.2** Possible role of the "sea breeze"

293 In summer, the surface wind direction sometimes changes from 120-160°E to North as 294 temperature rises over midday (Pettré et al., 1993; Gallée and Pettré, 1998), giving birth to an 295 apparent sea breeze. This phenomenon usually lasts half a day or less and air masses cannot 296 be referred to as oceanic (see section 2.2.3). Legrand et al. (2001) and Legrand et al. (2016b) 297 observed increasing atmospheric dimethylsulfide (DMS) and chloride concentrations, 298 respectively, during sea breeze events. However, our results indicate that Hg(0) 299 concentrations did not tend to increase systematically with the occurrence of a sea breeze 300 (e.g., Fig. 9).

# 301 **3.2.1.3 Role of snowpack emissions**

302 Angot et al. (2016) reported a daily cycle in summer at DC with maximal Hg(0)303 concentrations around midday. This daily cycle atop the East Antarctic ice sheet was 304 attributed to: i) an intense oxidation of Hg(0) in the atmospheric boundary layer due to the 305 high level of oxidants present there (Davis et al., 2001; Grannas et al., 2007; Eisele et al., 306 2008; Kukui et al., 2014), ii) Hg(II) dry deposition onto the snowpack, and iii) increased 307 emission of Hg(0) from the snowpack around midday as a response to daytime heating 308 following photoreduction of Hg(II) in the upper layers of the snowpack. Even if DDU is 309 located on snow free bedrock for most of the summer season, the same mechanism could 310 apply since the station is surrounded by vast snow-covered areas. However, such a dynamic 311 cycle of deposition/reemission at the air/snow interface requires the existence of a summertime atmospheric reservoir of Hg(II) species nearby DDU. This question is addressedin the following sub-section.

# **314 3.2.2** Transport of reactive air masses from the Antarctic plateau

315 Several previous studies pointed out that the major oxidants present in the summer 316 atmospheric boundary layer at coastal Antarctic sites differ in nature from site to site: 317 halogens chemistry prevails in the West, OH/NO<sub>x</sub> chemistry in the East (Legrand et al., 2009; 318 Grilli et al., 2013). Measurements made at HA in summer indicate a BrO mixing ratio of 3 319 pptv (Saiz-Lopez et al., 2007), a NO<sub>2</sub> mixing ratio of about 5 pptv (Bauguitte et al., 2012), and a 24 h average value of 3.9x10<sup>5</sup> radicals cm<sup>-3</sup> for OH (Bloss et al., 2007). Conversely, 320 321 BrO levels are at least lower by a factor of two at DDU (Legrand et al., 2016a) and Grilli et al. 322 (2013) reported a daily mean of 20 pptv for NO<sub>2</sub> in summer at DDU while Kukui et al. (2012) reported a 24 h average value of  $2.1 \times 10^6$  radicals cm<sup>-3</sup> for OH. Large OH/NO<sub>x</sub> concentrations 323 324 at DDU compared to HA were attributed to the arrival of air masses originating from the 325 Antarctic plateau where the OH/NO<sub>x</sub> chemistry is very efficient (Legrand et al., 2009; Kukui 326 et al., 2012).

327 Goodsite et al. (2004) and Wang et al. (2014) suggested a two-step oxidation mechanism for 328 Hg(0), favored at cold temperatures. The initial recombination of Hg(0) and Br is followed by 329 the addition of a second radical (e.g., I, Cl, BrO, ClO, OH, NO<sub>2</sub>, or HO<sub>2</sub>) in competition with 330 the thermal dissociation of the HgBr intermediate. Using the rate constants calculated by 331 Wang et al. (2014) for the reactions of BrO, NO<sub>2</sub>, and OH with the HgBr intermediate, we 332 found that BrO is the most efficient oxidant of HgBr at HA (lifetime of 1.9 min against 2.2 333 min with NO<sub>2</sub> and 11 days with OH). At DDU the situation is reversed with a lifetime of the 334 HgBr intermediate of 0.5 min with NO<sub>2</sub>, 3.9 min with BrO (assuming the presence of 1.5 pptv 335 of BrO in summer at DDU (Legrand et al., 2016a)), and 2 hours with OH. These results 336 suggest that the formation of Hg(II) species at DDU could be promoted by oxidants 337 transported from the Antarctic plateau towards the coast.

In addition to oxidants, inland air masses may transport mercury species. Low Hg(0) concentrations  $(0.76 \pm 0.30 \text{ ng m}^{-3})$  at DDU were associated with transport from the Antarctic plateau in summer (November to February, see Fig. 3b). A significant negative correlation was found in summer between Hg(0) concentrations and the daily-averaged percentage of air masses originating from the Antarctic plateau (r = -0.49, *p* value < 0.0001, Spearman test). Brooks et al. (2008a) reported elevated concentrations of oxidized mercury species at SP in summer  $(0.10 - 1.00 \text{ ng m}^{-3})$ . Similarly, Angot et al. (2016) observed low Hg(0) concentrations at the same period of the year at DC ( $0.69 \pm 0.35 \text{ ng m}^{-3}$ , i.e., ~ 25% lower than at NM, TNB and MM). Angot et al. (2016) also reported the occurrence of multi-day to weeklong Hg(0) depletion events (mean Hg(0) concentration ~ 0.40 ng m<sup>-3</sup>) likely due to a stagnation of air masses above the plateau triggering an accumulation of oxidants within the shallow boundary layer. These observations indicate that inland air masses reaching DDU in summer are depleted in Hg(0) and enriched in Hg(II).

#### 351 Transect from central to coastal Antarctica

The Hg<sub>tot</sub> concentration of snow samples collected in summer 2009 between DC and DDU 352 (see section 2.2.2) ranged from 4.2 to 194.4 ng L<sup>-1</sup> (Fig. 10). The closest sample from DC 353 exhibited a Hg<sub>tot</sub> concentration of 60.3  $\pm$  8.1 ng L<sup>-1</sup> (n = 3), in very good agreement with 354 concentrations found in surface snow samples collected in summer at DC (up to  $73.8 \pm 0.9$  ng 355 356 L<sup>-1</sup>, Angot et al., 2016). As illustrated by Fig. 10, Hg<sub>tot</sub> concentrations increased between 600-357 800 km and 1000-1100 km from DC in areas characterized by steeper slopes and higher snow 358 accumulation values. Several studies reported a gradual increase in snow accumulation from 359 DC toward the coast (Magand et al., 2007; Verfaillie et al., 2012; Favier et al., 2013), in 360 good agreement with a gradual increase in humidity (Bromwich et al., 2004). These results suggest that the wet deposition of Hg(II) species was enhanced near the coast, resulting in 361 elevated Hgtot concentrations in surface snow samples. Additionally, the presence of halides 362 such as chloride in snow can reduce the reduction rate of deposited Hg(II) species by 363 364 competing with the complexation of Hg(II) with dicarboxylic acids (Si and Ariya, 2008) resulting in higher Hgtot concentrations in coastal snowpacks (Steffen et al., 2014). It is worth 365 366 noting that the Hgtot concentrations between DC and DDU were higher than the values 367 measured in summer along other expedition routes in East Antarctica. Han et al. (2011) measured very low Hg<sub>tot</sub> concentrations ( $< 0.4 - 10.8 \text{ pg g}^{-1}$ ) along a ~ 1500 km transect in 368 east Queen Maud Land, and  $\mathrm{Hg}_{tot}$  concentrations ranged from 0.2 to 8.3 ng  $\mathrm{L}^{\text{-1}}$  along a 369 370 transect from ZG to DA (Fig. 1) (Li et al., 2014). Unfortunately none of the samples collected 371 during these two traverses were truly coastal – the most seaward samples were collected at altitudes of 948 and 622 m, respectively - preventing a direct comparison with the 372 concentration measured near DDU. The mean Hg<sub>tot</sub> concentration of  $67 \pm 21$  ng L<sup>-1</sup> reported 373 374 by Brooks et al. (2008b) at MM is the only truly coastal value available in Antarctica and is 375 lower than the value reported here near DDU.

The advection of inland air masses enriched in both oxidants and Hg(II) likely results in the build-up of an atmospheric reservoir of Hg(II) species at DDU – as confirmed by elevated Hg<sub>tot</sub> concentrations in surface snow samples –, confirming the hypothesis of a dynamic cycle of deposition/reemission at the air/snow interface.

# 380 **3.2.3** The ocean as a source of Hg(0)

381 DDU is located on a small island with open ocean immediately around from December to 382 February. It should be noted that during summers 2011/2012, 2012/2013, and 2013/2014, 383 areas of open waters were observed but with a significant unusual large amount of sea ice. 384 Sea ice be obtained from http://www.iup.unimaps can 385 bremen.de:8084/amsr2data/asi\_daygrid\_swath/s6250/ (Spreen et al., 2008).

386 According to Fig. 3b, Hg(0) concentrations in oceanic air masses were elevated from December to February (1.04  $\pm$  0.29 ng m<sup>-3</sup>), and a significant positive correlation was found 387 between Hg(0) concentrations and the daily-averaged percentage of oceanic air masses in 388 389 summer (r = 0.50, p value < 0.0001, Spearman test). While in winter the ice cover limited 390 mercury exchange at the air/sea interface (Andersson et al., 2008) leading to the build-up of 391 mercury-enriched waters, large emissions of Hg(0) from the ocean likely occurred in summer. 392 According to Cossa et al. (2011), total mercury concentrations can be one order of magnitude 393 higher in under-ice seawater than those measured in open ocean waters. The authors attributed 394 this build-up of mercury-enriched surface waters to the massive algal production at basal sea 395 ice in spring/summer triggering a large production of Hg(0), and to the mercury enrichment in 396 brine during the formation of sea ice. Elevated Hg(0) concentrations in oceanic air masses are 397 consistent with observations in the Arctic where Hg(0) concentrations in ambient air peak in 398 summer due to oceanic evasion and snowmelt revolatilization (Dastoor and Durnford, 2014). 399 Additionally, evasion from meltwater ponds formed on the remaining sea ice and observed 400 around the station may contribute to the increase in Hg(0) concentrations (Aspmo et al., 2006; 401 Durnford and Dastoor, 2011).

402

# 403 **4** Implications

#### 404 **4.1** For coastal Antarctic ecosystems

The reactivity of atmospheric mercury is unexpectedly significant in summer on the Antarctic plateau as evidenced by elevated Hg(II) and low Hg(0) concentrations (Brooks et al., 2008a;

Dommergue et al., 2012; Angot et al., 2016). This study shows that katabatic/continental 407 408 winds can transport this inland atmospheric reservoir toward the coastal margins where Hg(II) 409 species tend to deposit due to increasing wet deposition (Fig. 10). However, the 410 postdeposition dynamics of mercury and its ultimate fate in ecosystems remain unknown. 411 Bargagli et al. (1993) and Bargagli et al. (2005) showed evidence of enhanced 412 bioaccumulation of mercury in soils, mosses, and lichens collected in ice-free areas around 413 the Nansen Ice Sheet (Victoria Land, upslope from the Ross Ice Shelf), suggesting an 414 enhanced deposition of mercury species. Interestingly, four large glaciers join in the Nansen 415 Ice Sheet region and channel the downward flow of air masses from the Antarctic plateau 416 toward Terra Nova Bay, generating intense katabatic winds. The monthly mean wind speed is about 16 m s<sup>-1</sup> in this area (Bromwich, 1989). Along with an enhanced deposition of mercury 417 418 during AMDEs, the wind might as well be responsible for the advection of inland air masses 419 enriched in Hg(II) species as observed in our case study. As already pointed out by Bargagli 420 et al. (2005), coastal Antarctic ecosystems may become a sink for mercury, especially in view 421 of increasing anthropogenic emissions of mercury in Asia (Streets et al., 2009).

#### 422 **4.2** For the cycle of atmospheric mercury in high southern latitudes

423 The influence of the Antarctic continent on the global geochemical cycle of mercury remains 424 unclear (Dommergue et al., 2010). This study shows that the reactivity observed on the 425 Antarctic plateau (Brooks et al., 2008a; Dommergue et al., 2012; Angot et al., 2016) 426 influences the cycle of atmospheric mercury at a continental scale, especially downstream of 427 the main topographic confluence zones. The question is whether the katabatic airflow 428 propagation over the ocean is important. According to Mather and Miller (1967), the katabatic 429 flow draining from the Antarctic plateau merges with the coastal polar easterlies under the 430 action of the Coriolis force. The near-surface flow takes the form of an anticyclonic vortex 431 (King and Turner, 1997), limiting the propagation of katabatic flows over the ocean.

432

# 433 **5** Conclusion

We presented here a 3.5-year record of Hg(0) concentrations at DDU, first multi-year record on the East Antarctic coast. Our observations reveal a number of differences with other costal or near coastal Antarctic records. In winter, observations showed a gradual 20% decrease in Hg(0) concentrations from May to August, a trend never observed at other coastal sites. This is interpreted as a result of reactions occurring within the shallow boundary layer on the 439 Antarctic plateau, subsequently efficiently transported at that site by katabatic winds. In 440 summer, the advection of inland air masses enriched in oxidants and Hg(II) species likely 441 results in the build-up of an atmospheric reservoir of Hg(II) species at DDU, at least partly explaining the elevated (up to 194.4 ng L<sup>-1</sup>) Hg<sub>tot</sub> concentrations measured in surface snow 442 samples near the station during a traverse between DC and DDU. Additionally, Hg(0) 443 444 concentrations in ambient air exhibited a diurnal cycle in summer at DDU – phenomenon 445 never observed at other coastal Antarctic stations. Several processes may contribute to this 446 diurnal cycle, including a local chemical exchange at the air/snow interface in the presence of 447 elevated levels of Hg(II) species in ambient air, and emissions from ornithogenic soils present 448 at the site. Our data also highlight the fact that the Austral Ocean may be a net source for 449 mercury in the summer. Even though AMDEs are likely very rare at DDU compared to other 450 coastal stations, we cannot exclude that the sea-ice present offshore DDU at the end of winter 451 influenced springtime Hg(0) levels. Finally, having shown that the reactivity observed on the 452 Antarctic plateau influences the cycle of atmospheric mercury on the East Antarctic coast, this 453 study raises concern for coastal Antarctic ecosystems there.

454

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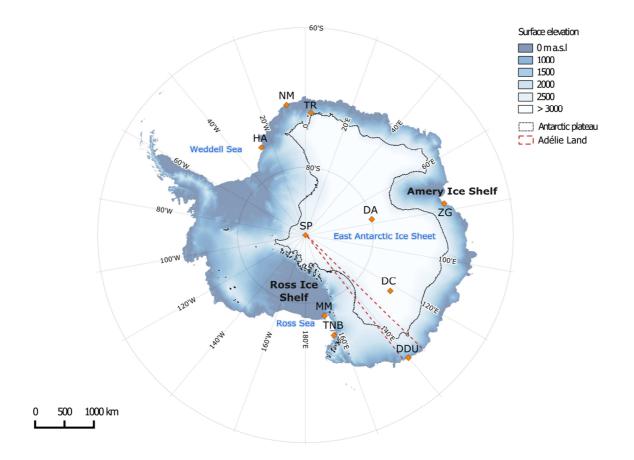
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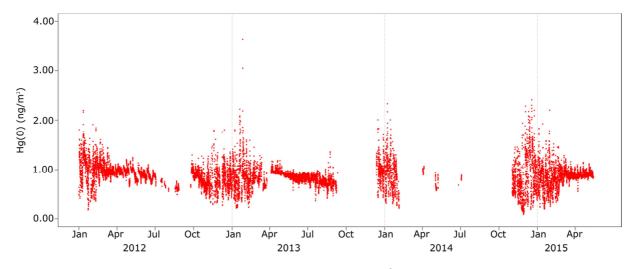
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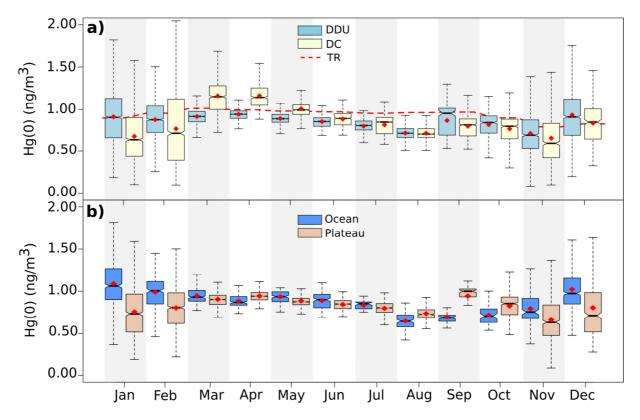
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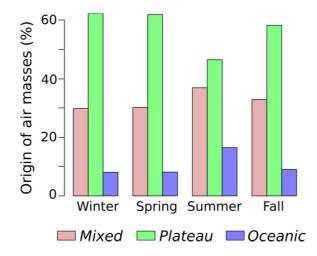
**Figure 1**: Map of Antarctica showing surface elevation (meters above sea level, m a.s.l) and the position of various stations: Halley (HA), Neumayer (NM), Troll (TR), Zhongshan Station (ZG), Dome A (DA), South Pole Station (SP), Concordia Station (DC), Dumont d'Urville (DDU), McMurdo (MM), and Terra Nova Bay (TNB). The black line delimits the high altitude plateau (> 2500 m a.s.l), and the red dotted line Adélie Land (from 136°E to 142°E).



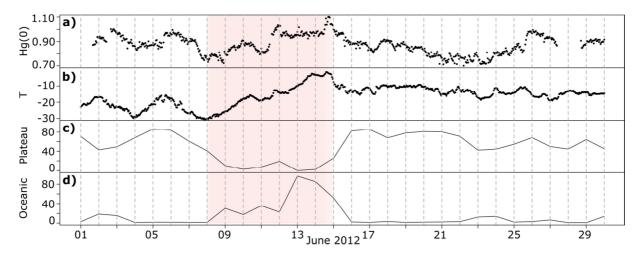
**Figure 2**: Hourly-averaged Hg(0) concentrations (ng/m<sup>3</sup>) measured at DDU from January 2012 to May 2015. Missing data are due to instrument failure or QA/QC invalidation. Hg(0) concentrations were highly variable during the sunlit period as compared to wintertime (May-August) suggesting a photochemically-induced reactivity at this period of the year.



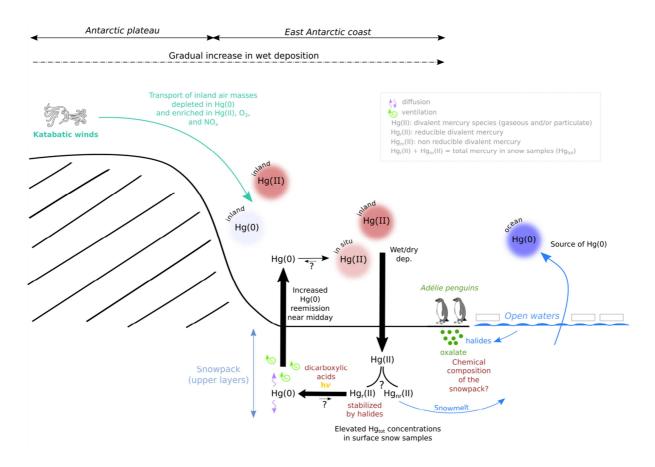
**Figure 3**: Box and whisker plot presenting the monthly Hg(0) concentration distribution **a**) from all the data collected at DDU and DC along with the monthly mean recorded at TR, and **b**) from all the data collected at DDU associated with air masses originating from the ocean or the Antarctic plateau according to the HYSPLIT simulations. • mean, bottom and top of the box: first and third quartiles, band inside the box: median, ends of the whiskers: lowest (highest) datum still within the 1.5 interquartile range of the lowest (upper) quartile. Outliers are not represented.



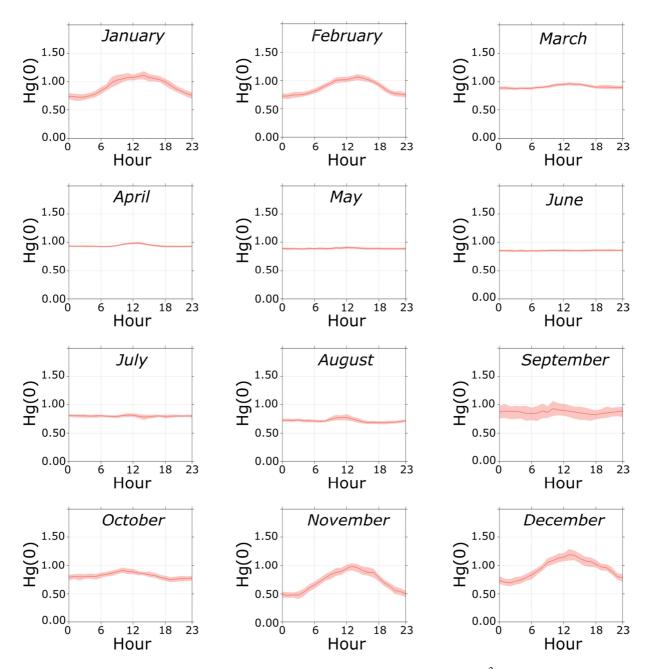
**Figure 4:** Mean percentage (%) of continental/oceanic mixed air masses (pink), and of air masses originating from the Antarctic plateau (green) or the ocean (blue) according to the HYSPLIT model simulations in winter (May to August), spring (September-October), summer (November to February), and fall (March-April).



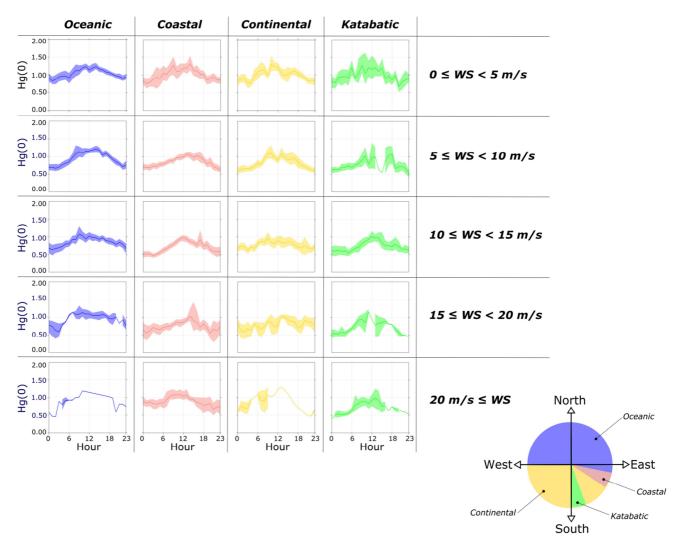
**Figure 5:** June 2012 variation of **a**) Hg(0) concentration  $(ng/m^3)$ , **b**) temperature (°C), **c**) daily-averaged percentage (%) of air masses originating from the Plateau (HYSPLIT model simulations), and **d**) daily-averaged percentage (%) of air masses originating from the ocean (HYSPLIT model simulations). From 8 to 14 June (period highlighted in red), both Hg(0) and temperature increased suggesting an advection of air masses from mid-latitudes, as confirmed by an elevated percentage of oceanic air masses.



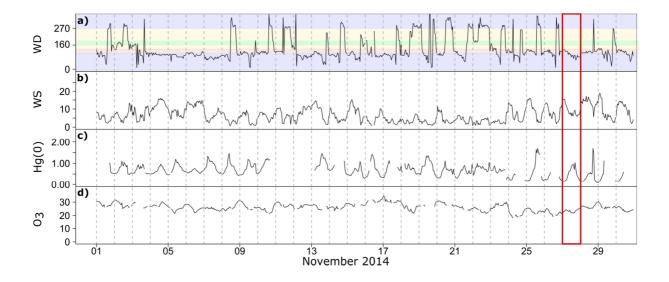
**Figure 6:** Schematic diagram illustrating the processes that may govern the mercury budget at DDU in summer. Katabatic winds transport inland air masses enriched in oxidants and Hg(II) toward the coastal margins. Hg(II) species deposit onto the snowpack by wet and dry processes leading to elevated concentrations of total mercury in surface snow samples. A fraction of deposited mercury can be reduced (the reducible pool, Hg<sub>r</sub>(II)) in the upper layers of the snowpack and subsequently reemitted to the atmosphere as Hg(0). Hg(0) emission from the snowpack maximizes near midday likely as a response to daytime heating. The chemical composition of the snowpack (halides, dicarboxylic acids) may influence the reduction rate of Hg(II) species within the snowpack. The ocean may be a net source of Hg(0) to the atmosphere.



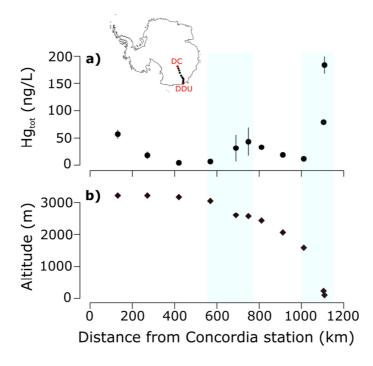
**Figure 7**: Monthly mean diurnal cycle of Hg(0) concentrations (in ng/m<sup>3</sup>) along with the 95% confidence interval for the mean, calculated from all the data collected at DDU (January 2012-May 2015). Hours are in local time (UTC+10). Hg(0) concentrations exhibit a strong diurnal cycle in summer (November to February).



**Figure 8:** Summertime (November to February) mean diurnal cycle of Hg(0) concentrations (in ng/m<sup>3</sup>), along with the 95% confidence interval for the mean, depending on wind direction and wind speed. With north at 0°, oceanic winds ranged from 270 to 110°, coastal winds from 110 to 130°, katabatic winds from 160 to 180°, and continental winds from 130 to 160° and from 180 to 270°. Hours are in local time (UTC+10). Hg(0) concentrations exhibit a diurnal cycle regardless of wind speed and direction.



**Figure 9:** November 2014 variation of **a**) wind direction (WD, in °), **b**) wind speed (WS, in m/s), **c**) Hg(0) concentration (in ng/m<sup>3</sup>), and **d**) O<sub>3</sub> mixing ratio (in ppbv). With north at 0°, oceanic winds ranged from 270 to 110° (purple), coastal winds from 110 to 130° (pink), katabatic winds from 160 to 180° (green), and continental winds from 130 to 160° and from 180 to 270° (yellow). On 27 November 2014 (period framed in red), a sea breeze is observed around midday: WD changes from ~ 120-130° to below 110° while WS decreases. Both Hg(0) concentrations and O<sub>3</sub> mixing ratios are not higher than during the previous days.



**Figure 10: a)** Total mercury concentration in surface snow samples (Hg<sub>tot</sub> in ng/L) along with standard deviation and **b**) altitude (m) vs. distance from Concordia station (DC) during the traverse from DC to DDU. Hg<sub>tot</sub> concentrations increased in areas highlighted in blue, characterized by steeper slopes and higher snow accumulation values.